

# Leptogenesis and light DM in the Scotogenic model

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PRISMA

#DMLHC2018



# Scotogenic model spectrum I

E. Ma [hep-ph/0601225]

	$SU(2)_L$	$U(1)_Y$	$Z_2$
$\Sigma$	2	1/2	-
$N_i$	1	0	-
$\Phi$	2	1/2	+
$L$	2	-1/2	+

$Z_2$  symmetry:

→ Forbids decays  $N_i \rightarrow \nu\gamma$

→ No vev for  $\Sigma$

→ Radiative seesaw

→ Stable particle:  $N_1 \rightarrow DM$

$$\mathcal{L} \supset h_{ki} \bar{N}_k \tilde{\Sigma}^\dagger L_i - \frac{1}{2} \bar{N}_k^c M_k N_k + \text{h.c.}$$

$M_1 = \mathcal{O}(\text{keV}) \longrightarrow \textcolor{red}{\textit{Freeze-In}}$

$M_{2,3} = \mathcal{O}(\text{few 100 GeV}) \longrightarrow \textcolor{blue}{\textit{Freeze-Out}}$

# Scotogenic model spectrum II

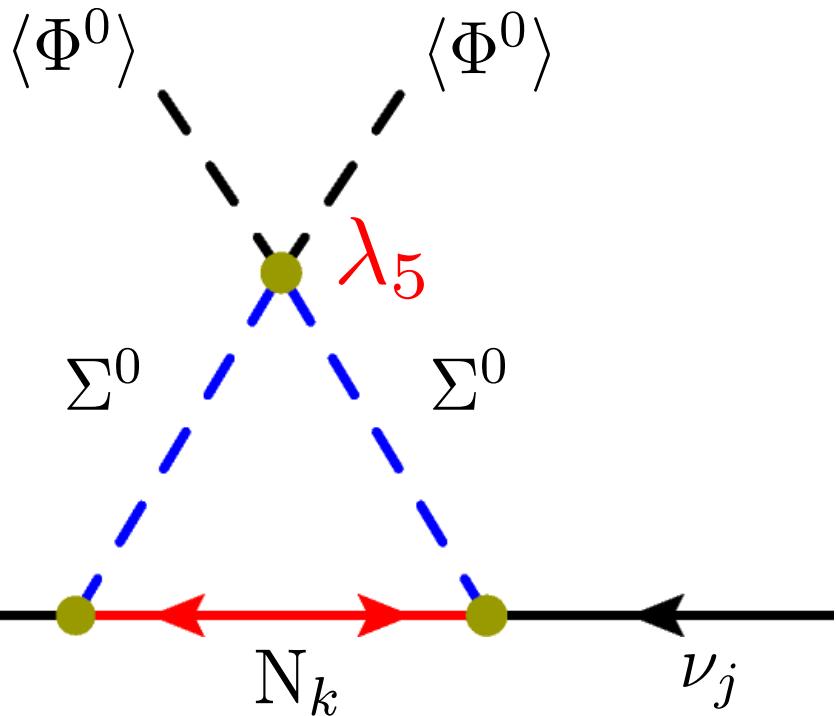
Scalar sector:

$$V(\Phi, \Sigma) = \mu_1^2 \Phi^\dagger \Phi + \mu_2^2 \Sigma^\dagger \Sigma + \frac{\lambda_1}{2} (\Phi^\dagger \Phi)^2 + \frac{\lambda_2}{2} (\Sigma^\dagger \Sigma)^2 \\ + \lambda_3 (\Phi^\dagger \Phi) (\Sigma^\dagger \Sigma) + \lambda_4 (\Phi^\dagger \Sigma) (\Sigma^\dagger \Phi) + \frac{\lambda_5}{2} ((\Phi^\dagger \Sigma)^2 + \text{h.c.})$$

$$\Sigma = \begin{pmatrix} \sigma^+ \\ \frac{1}{\sqrt{2}} (S^0 + iA^0) \end{pmatrix}$$

$$\left. \begin{aligned} m_\pm^2 &= \mu_2^2 + \frac{\lambda_3 v^2}{2} \\ m_A^2 &= m_\pm^2 + \frac{(\lambda_4 - \lambda_5)v^2}{2} \\ m_S^2 &= m_A^2 + \lambda_5 v^2 \end{aligned} \right\} \leq 1 \text{ TeV}$$

# Radiative neutrino masses



$$m_0^2 \equiv \frac{m_S^2 + m_A^2}{2} \gg M_k^2$$

$$(m_\nu)_{ij} \approx \frac{\lambda_5 v^2}{8\pi^2} \frac{h_{ki} h_{kj}}{m_0^2} M_k$$

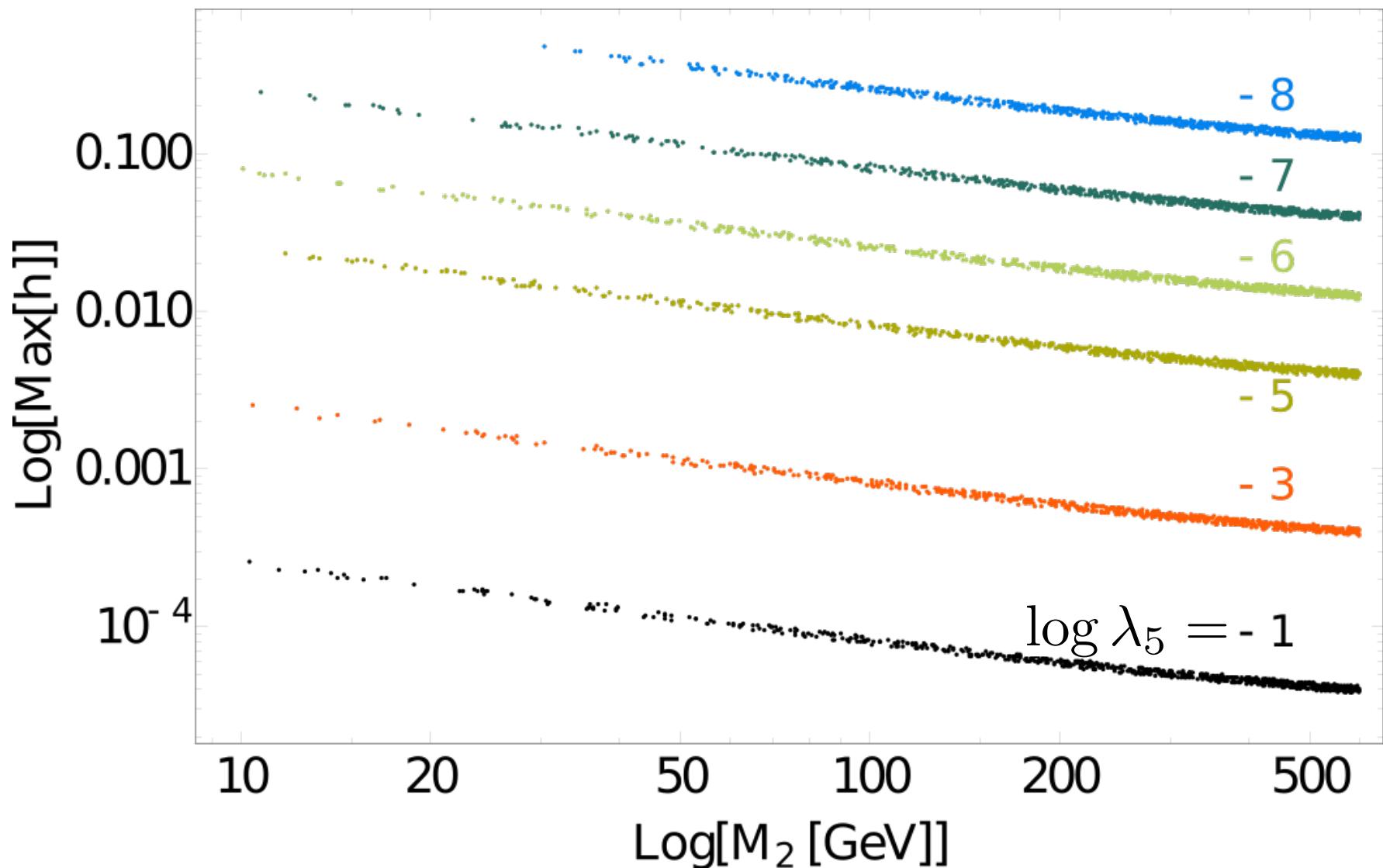
- Casas-Ibarra Parametrization:

$$h = \sqrt{\Lambda^{-1}} R \sqrt{m_\nu} U_{\text{PMNS}}^\dagger$$

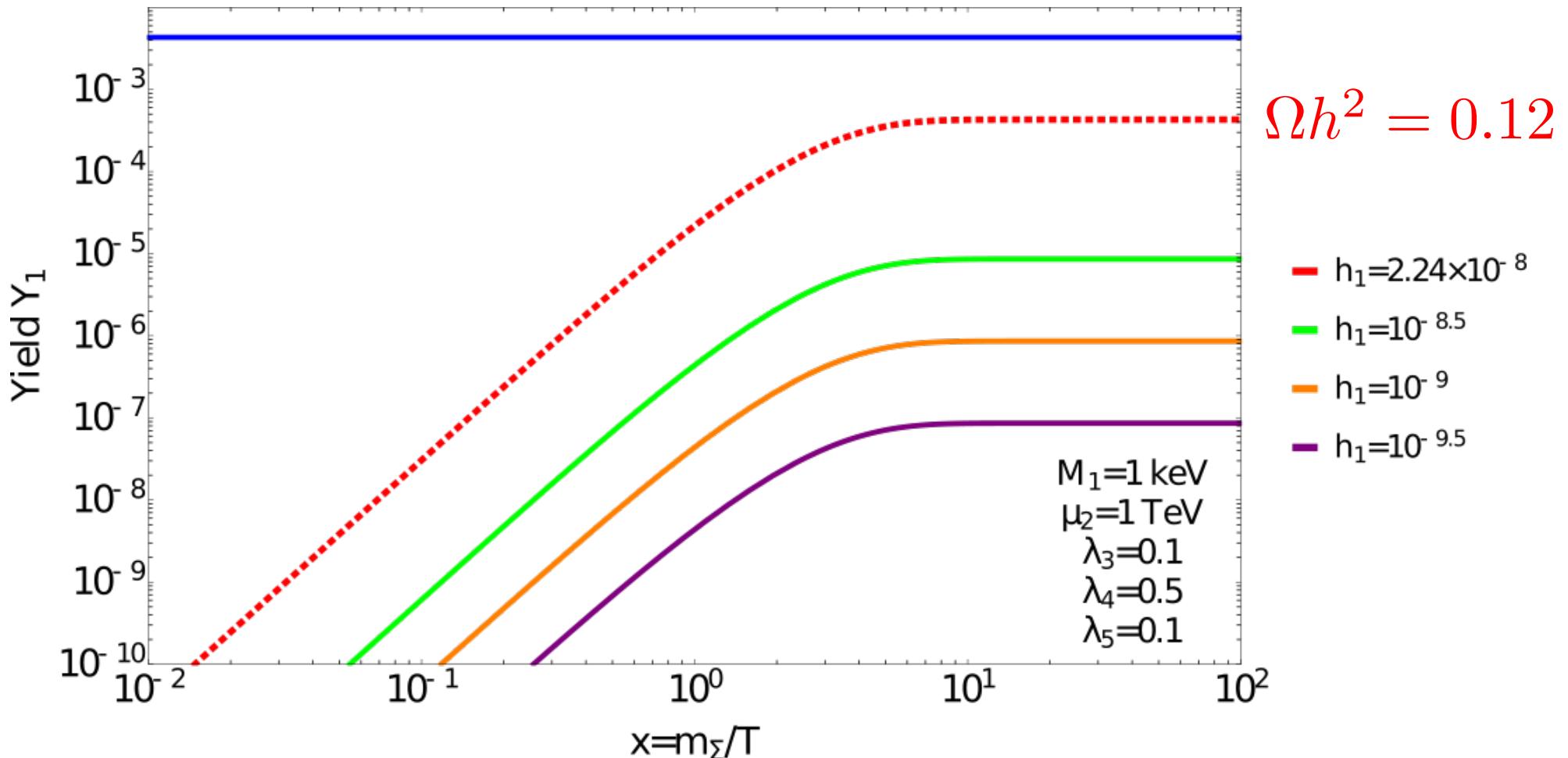
$$m_\nu = \text{Diag}(0, \sqrt{\Delta_{\text{sol}}^2}, \sqrt{\Delta_{\text{sol}}^2 + \Delta_{\text{atm}}^2})$$

$$R = \begin{pmatrix} 0 & \cos(w + i\gamma) & \sin(w + i\gamma) \\ 0 & -\sin(w + i\gamma) & \cos(w + i\gamma) \end{pmatrix}$$

# DM I: impact of $\lambda_5$



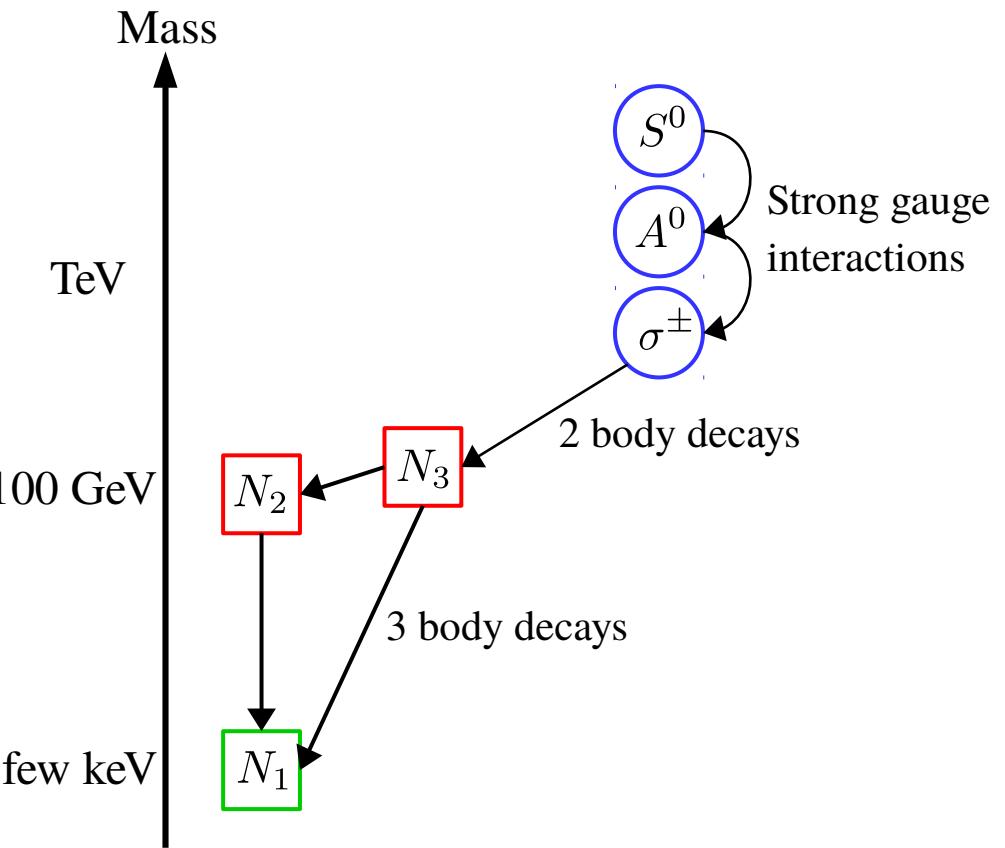
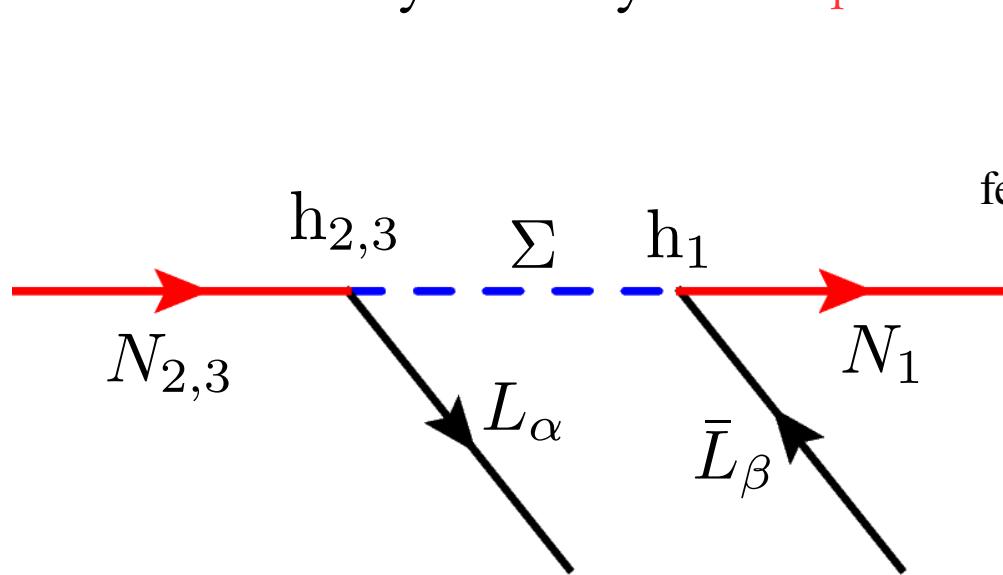
# DM II: N<sub>1</sub> Freeze-In



$$\Omega_{FI} h^2 = 2.75 \cdot 10^8 \frac{m_1}{\text{GeV}} Y_1 \approx 0.12 \left( \frac{y_1}{10^{-7.86}} \right)^2 \left( \frac{m_1}{1 \text{ keV}} \right) \left( \frac{1 \text{ TeV}}{m_\pm} \right)$$

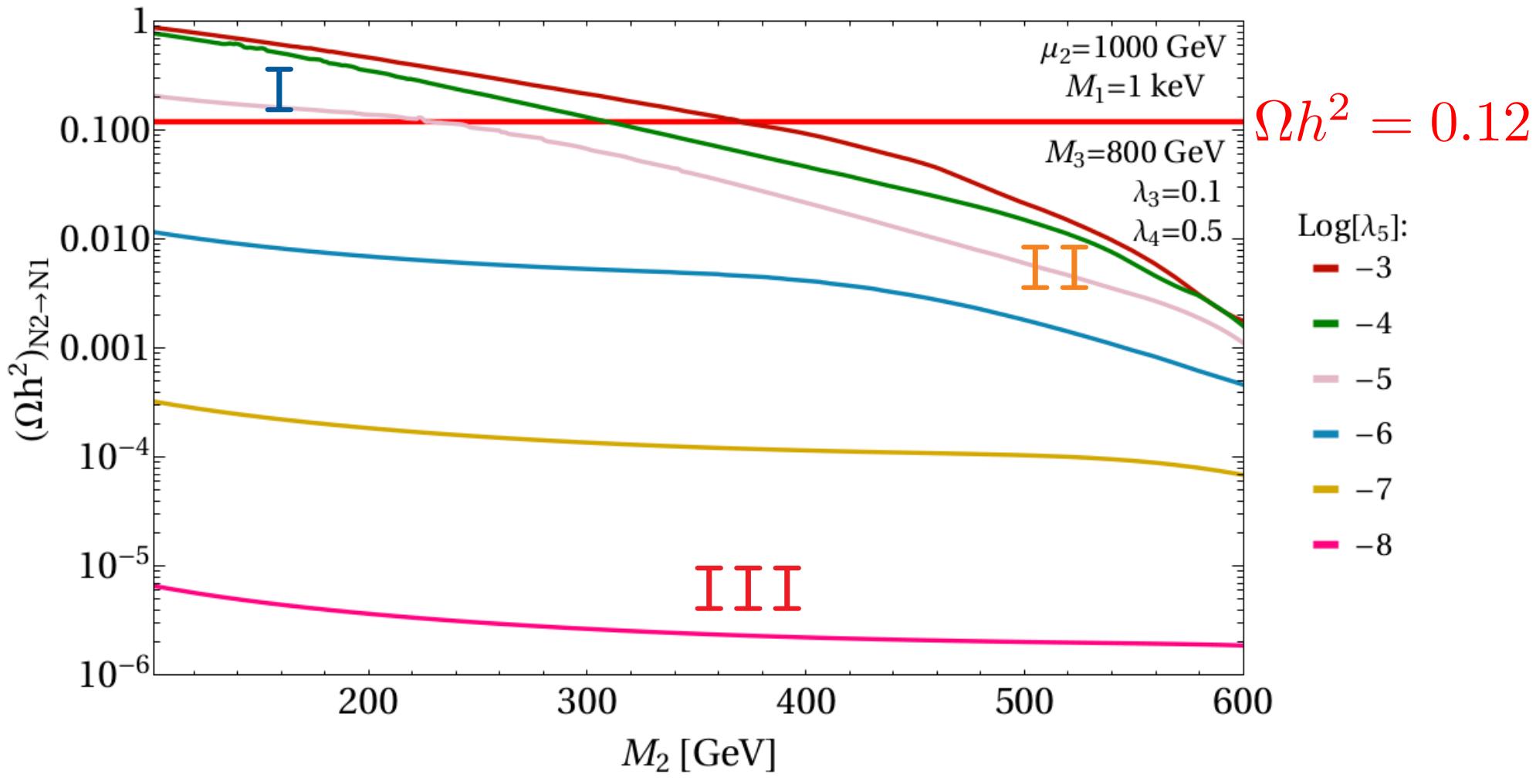
# DM III: $N_2$ decay contributions

- $h_1 \ll 1$ : Decouple  $N_1$
- Decay of  $Z_2$  odd particles into  $N_2$
- Three body decay to  $N_1$



$$\Omega h^2 = \Omega h_{\text{FI}}^2 + \Omega h_{N2 \rightarrow N1}^2$$

# DM III: $N_2$ decay contributions



I + III: Only  $N_2$  annihilation

II: Additional scalar annihilation

# Leptogenesis I: Fundamentals

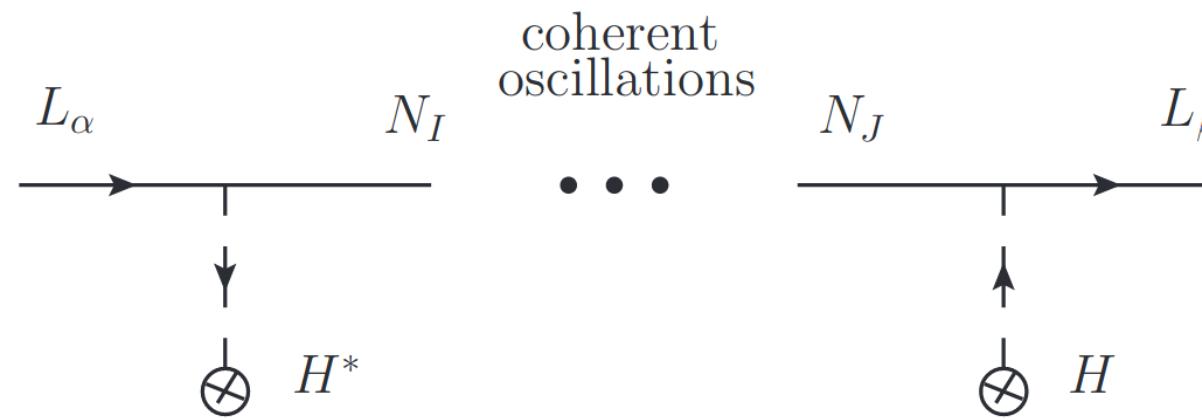
Goal: *CP violation for lepton asymmetry*

Akhmedov, Rubinov, Smirnov [hep-ph/9803255]

T. Hambye, D.Teresi [1606.00017]

## 1) ARS:

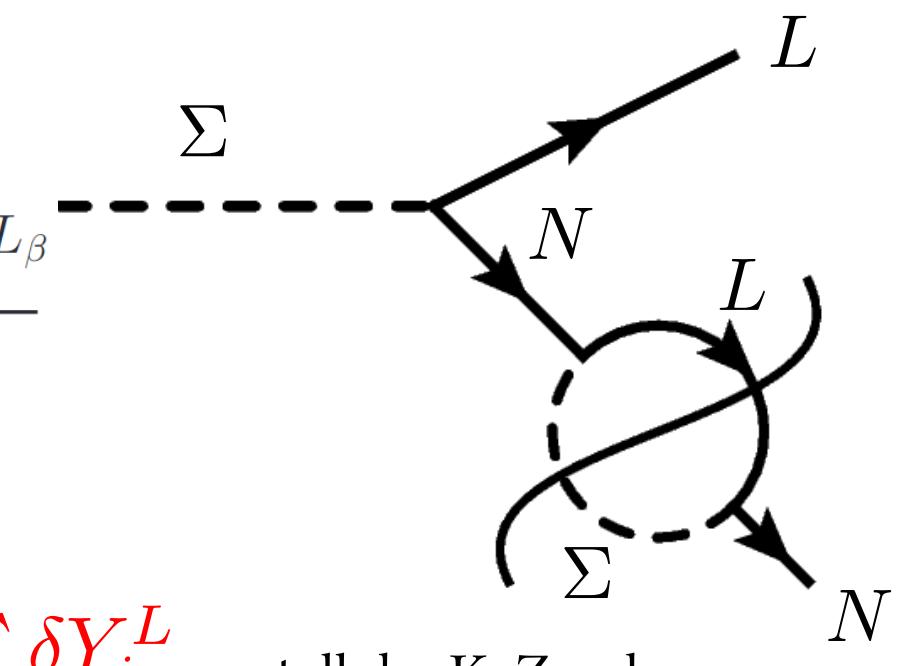
- Oscillations among  $N_i$
- Asymmetry generation in each flavor



## 2) Decays of $\Sigma \rightarrow N L$

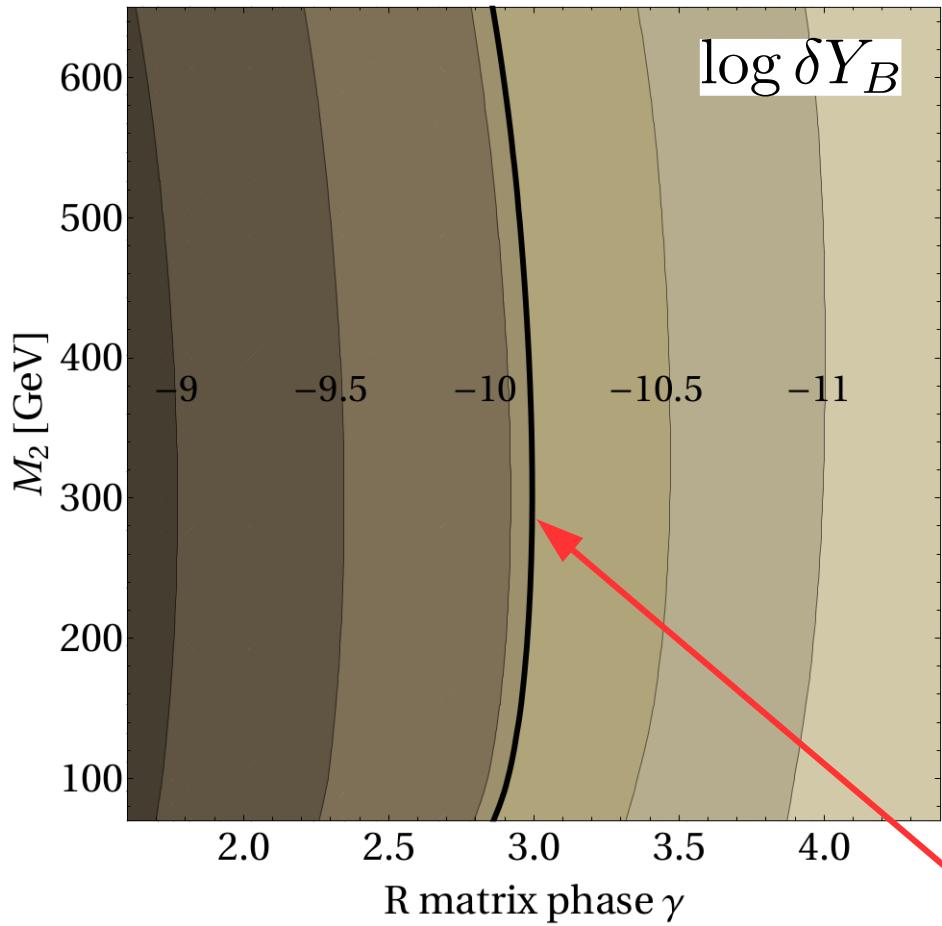
- Thermally induced on-shell particles

$$\delta Y_B = -\frac{2}{3} \sum_i \delta Y_i^L \rightarrow \text{talk by K. Zurek}$$



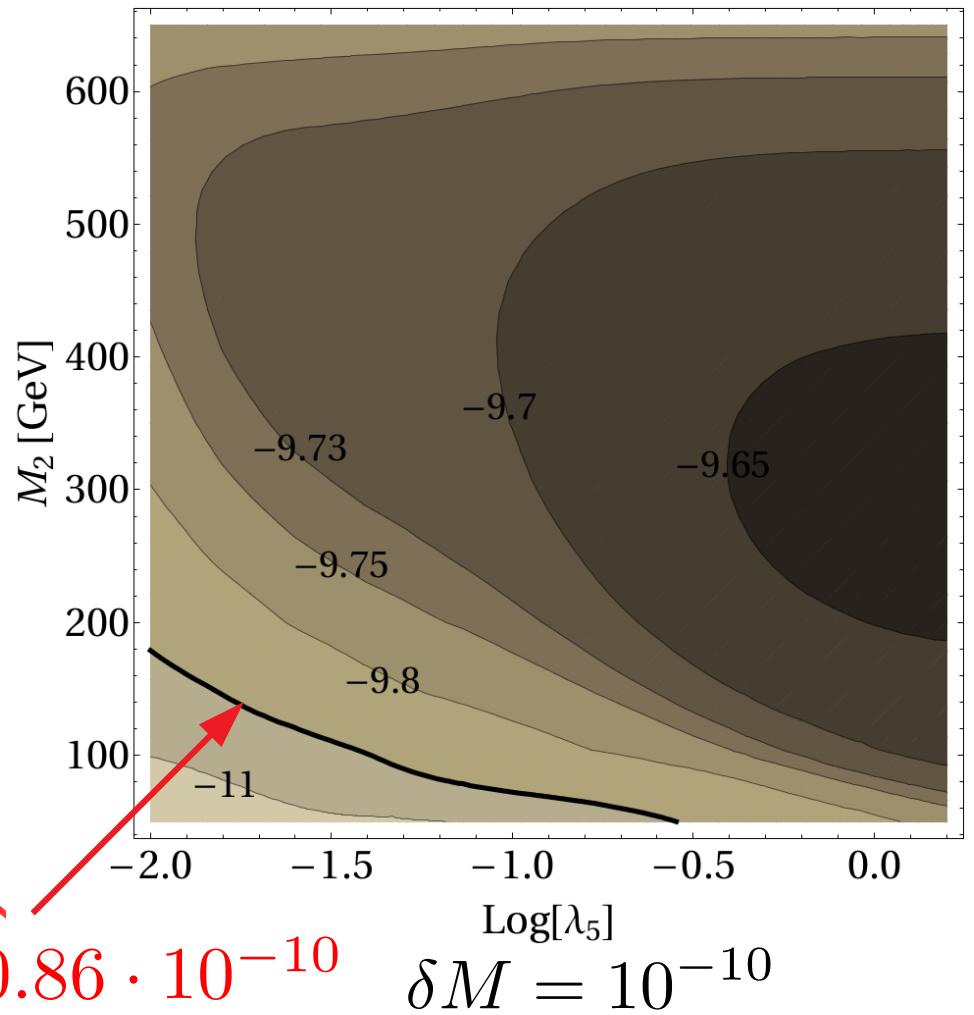
# Leptogenesis II: results

$$\frac{d \delta Y_i^L}{dz} \approx S - \delta Y_i^L W$$



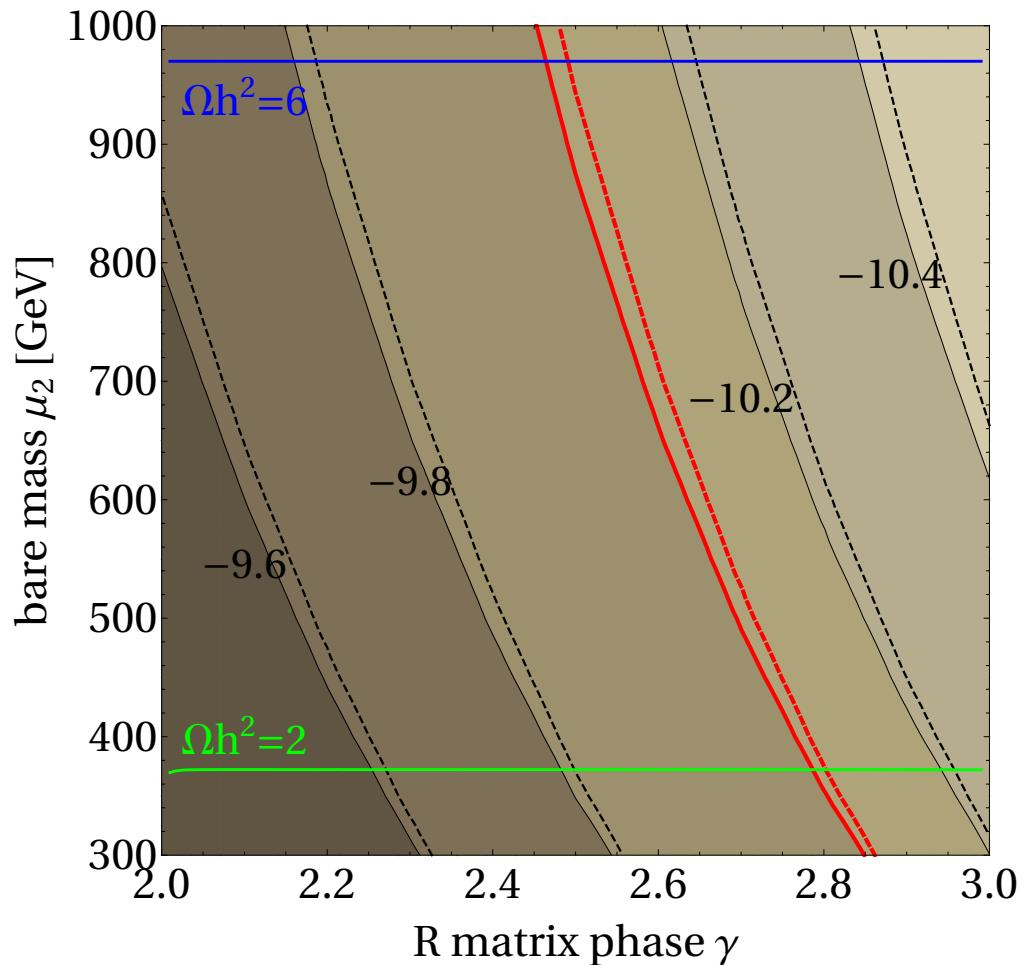
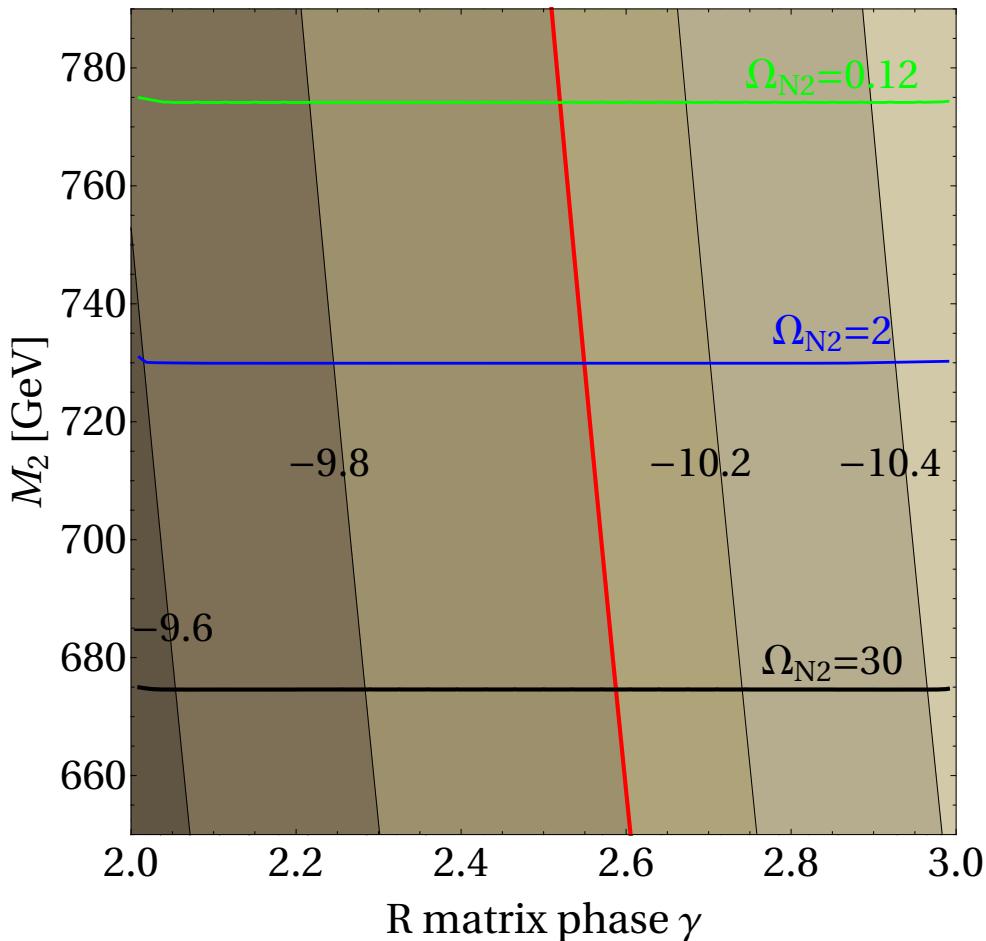
$$\delta Y_B = 0.86 \cdot 10^{-10}$$

$\rightarrow$  *Small* Yukawas required



# Merging DM + Leptogenesis

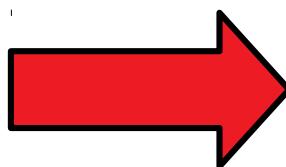
Additional degeneracy between  $\Sigma$  and  $N_2$  masses



Preliminary results!

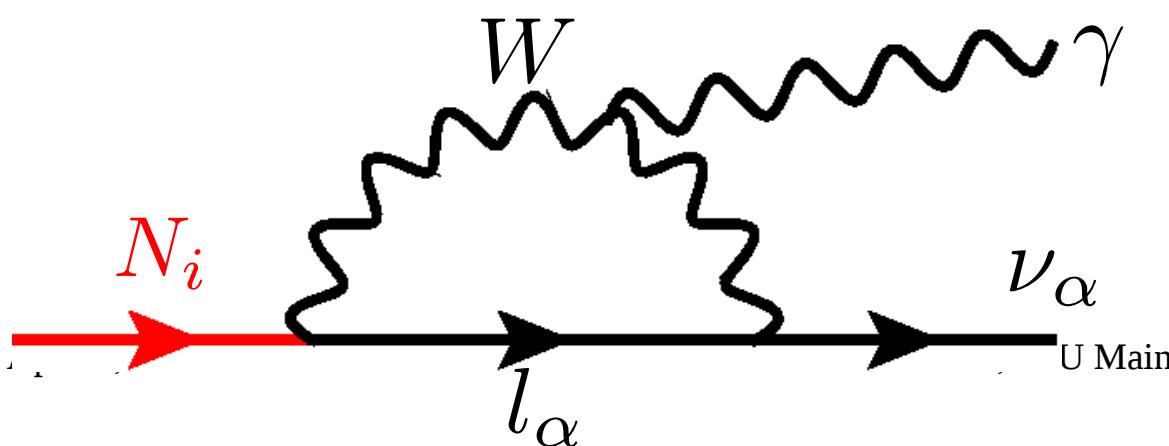
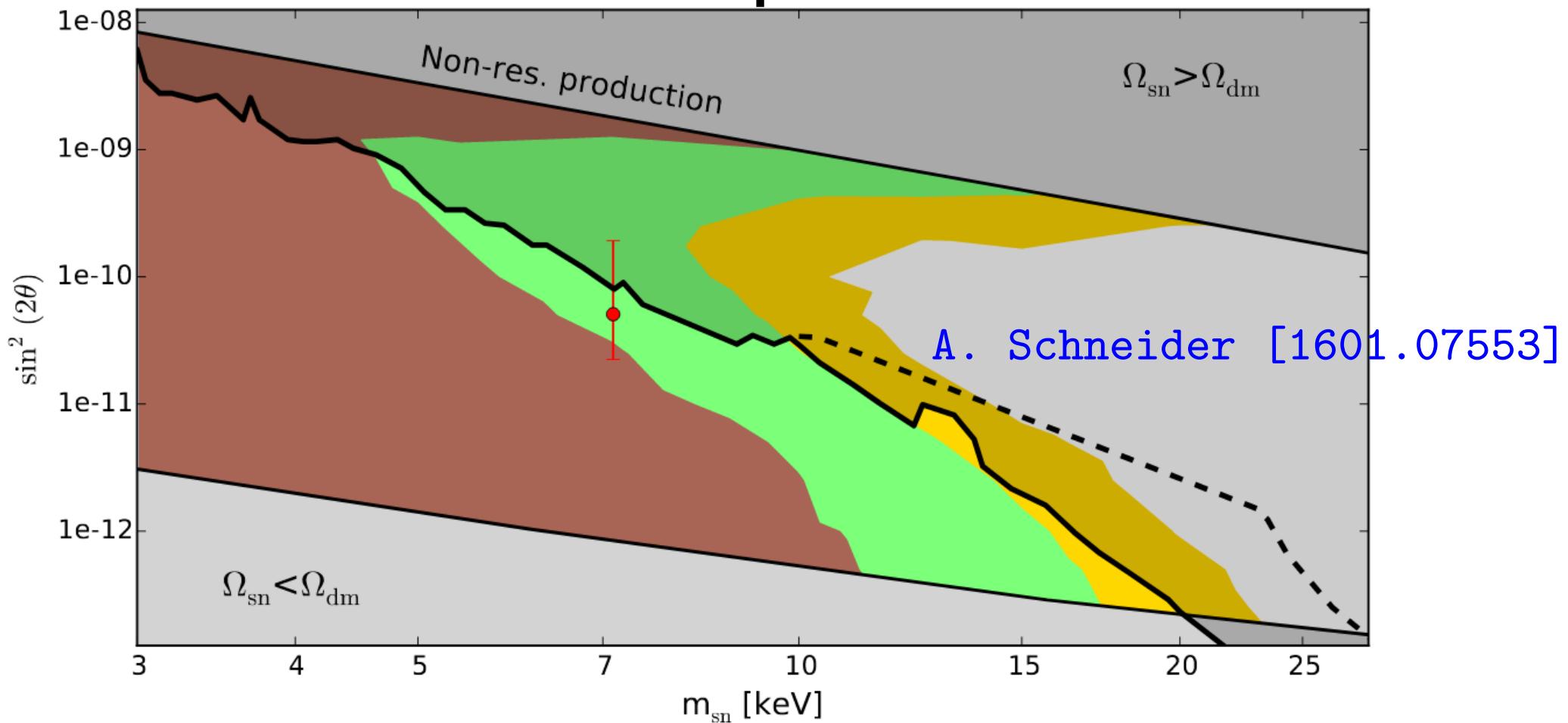
# Summary

- Sub TeV realization of Scotogenic model
- Successful DM density:
  - Freeze-In of  $N_1$
  - $N_2$  decay contribution
- Significant baryon asymmetry from Leptogenesis



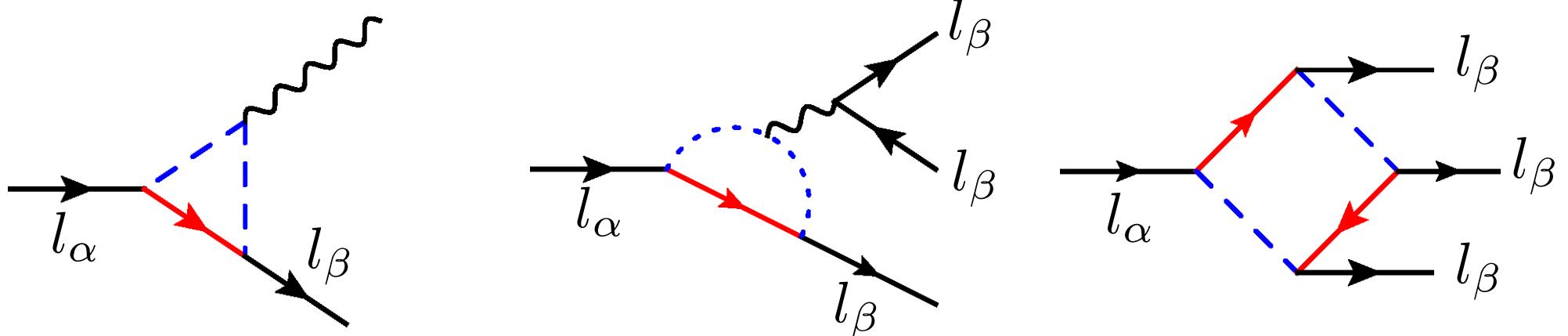
Parameter region for combination!

# Backup slides



$$\Gamma \propto \sin(2\theta)^2 m_s^5$$

# Backup slides



T. Toma, A. Vicente [1312.2840]

Process	BR bound
$\mu \rightarrow e \gamma$	$4.2 \cdot 10^{-13}$
$\tau \rightarrow e \gamma$	$3.3 \cdot 10^{-8}$
$\tau \rightarrow \mu e \gamma$	$4.4 \cdot 10^{-8}$
$\mu \rightarrow 3e$	$1.0 \cdot 10^{-12}$
$\tau \rightarrow 3e$	$2.7 \cdot 10^{-8}$
$\tau \rightarrow 3\mu$	$2.1 \cdot 10^{-8}$

→ *Upper bound* on Yukawa strength

Particle Data Group +  
1605.05081

# Backup slides

- Boltzmann equation for  $N_1$

$$Y'_1(x) = 6 \sum_{i=\pm, S} \frac{45 M_{\text{Pl}} y_1^2}{32\pi^4 \cdot 1.66(g_*)^{3/2}} \frac{m_i^3}{m_1^4} x^3 \cdot K_1 \left( \frac{m_i}{m_1} x \right)$$

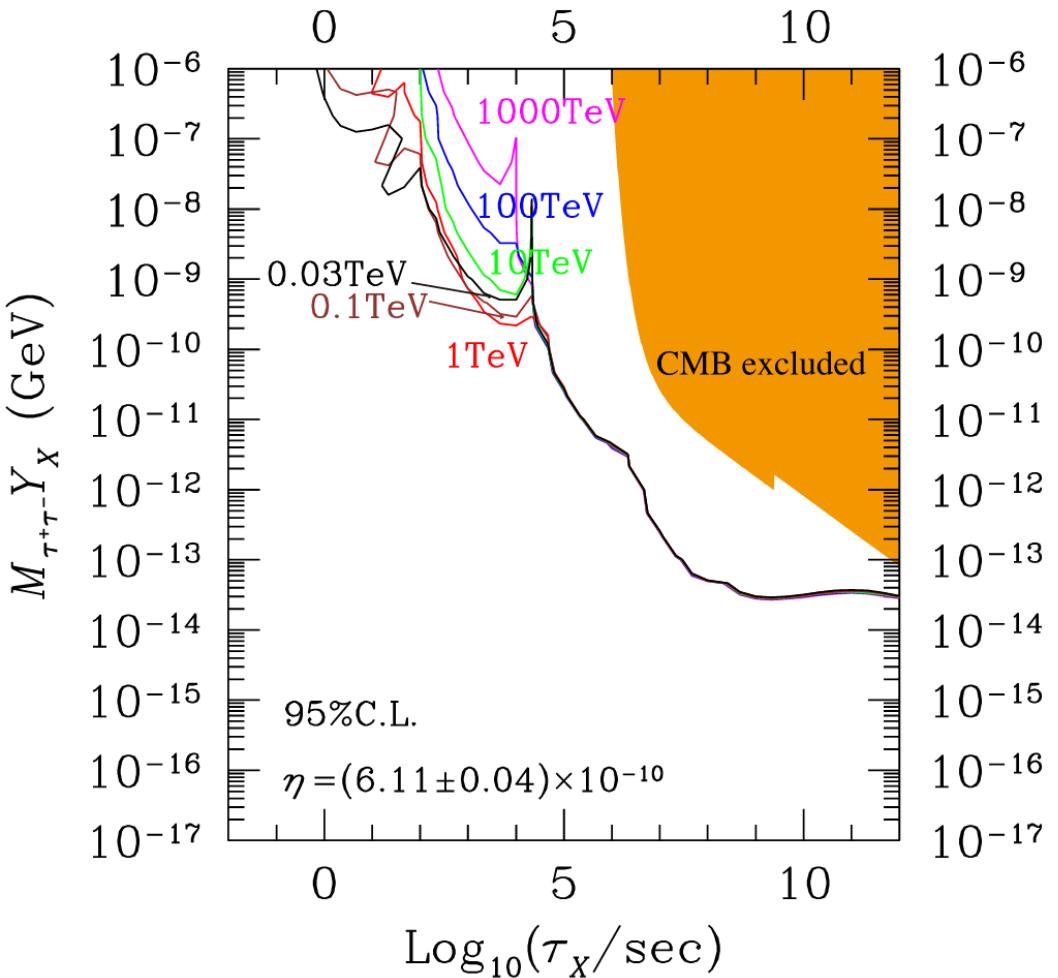
- Decay width  $N_{2,3} \rightarrow N_1$ :

$$\Gamma(N_{2,3} \rightarrow N_1 \bar{l}_\alpha l_\beta) = \frac{M_{2,3}^5}{6144\pi^3 m_S^4} (|h_{1\beta}|^2 |h_{2,3\alpha}|^2 + |h_{1\alpha}|^2 |h_{2,3\beta}|)$$

E. Molinaro et. al. [1405.1259]

# BBN bounds

$$(h_2)^2 \gtrsim 6.3 \cdot 10^{-7} \left( \frac{m_\Sigma}{1 \text{ TeV}} \right)^4 \left( \frac{1 \text{ TeV}}{M_2} \right)^5 \left( \frac{10^{-8}}{|h_1|} \right)^2 \left( \frac{1 \text{ sec}}{\tau} \right)$$



- Naively:  $t_{\text{BBN}} \sim 1 \text{ sec}$
  - Dominantly decays into  $\tau$  pairs
- Small yield required

M. Kawasaki et.al. [1709.01211]

# Backup slides

- Boltzmann equation for Leptogenesis:

$$\frac{dn_{\alpha\beta}^N}{dt} = -i[E_N, n^N(\mathbf{k})]_{\alpha\beta} - \frac{1}{2} \left\{ \gamma^{LC} + \gamma^{LV}, \frac{n^N}{n_{\text{eq}}^N} - I \right\}_{\alpha\beta} \\ + \frac{\delta n_l^L}{2n_{\text{eq}}^L} \left( (\gamma_{WQ,l}^{LC} - \gamma_{WQ,l}^{LV}) + \frac{1}{2} \left\{ \gamma_{WC,l}^{LC} - \gamma_{WC,l}^{LV}, \frac{n^N}{n_{\text{eq}}^N} \right\} \right)_{\alpha\beta}$$

$$\frac{d\bar{n}_{\alpha\beta}^N}{dt} = \frac{dn_{\alpha\beta}^N}{dt} (n \rightarrow \bar{n}, \gamma \rightarrow \gamma^*, \delta n_l^L \rightarrow -\delta n_l^L)$$

$$\frac{d\delta n_l^L}{dt} = \frac{1}{n_{\text{eq}}^N} \text{tr} \left\{ (\gamma_l^{LC} - \gamma_l^{LV}) n^N \right\} - (\gamma \rightarrow \gamma^*, n \rightarrow \bar{n})$$

$$- \frac{\delta n_l^L}{n_{\text{eq}}^L} \text{tr} \left\{ \gamma_{WQ,l}^{LC} + \gamma_{WQ,l}^{LV} \right\} \qquad \qquad \qquad \delta Y_B = -\frac{2}{3} \sum_i \delta Y_i^L$$

$$- \frac{\delta n_l^L}{2n_{\text{eq}}^L} \frac{1}{n_{\text{eq}}^N} \text{tr} \left\{ n^N (\gamma_{WC,l}^{LC} + \gamma_{WC,l}^{LV}) \right\} - (\gamma \rightarrow \gamma^*, n \rightarrow \bar{n})$$