

The Search for Neutrinoless Double-Beta Decay

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Neutrinoless Double-Beta Decay ($0\nu\beta\beta$)

Small, non-zero neutrino masses motivate the possibility that neutrinos are Majorana particles, indistinguishable from antineutrinos.

Searches for neutrinoless double-beta decay ($0\nu\beta\beta$) are the only current practical probes of Majorana nature of neutrinos.

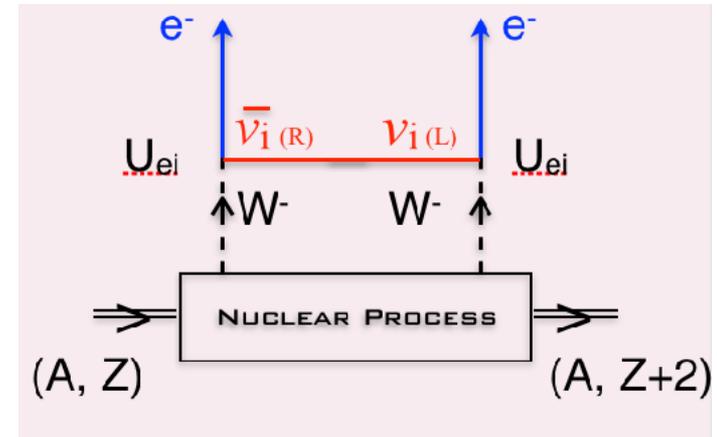
$0\nu\beta\beta$ requires:

Non-zero neutrino masses:

- “wrong-handed” helicity admixture $\sim m_i/E_{\nu_i}$
- *Any process that allows $0\nu\beta\beta$ to occur requires Majorana neutrinos with non-zero mass.* -Schechter and Valle, 1982

Lepton Number non-conservation

- No experimental evidence that Lepton number must be conserved
- Allowed based on general SM principles, such as electroweak-isospin conservation and renormalizability



Implications of a $0\nu\beta\beta$ Measurement

Observation of $0\nu\beta\beta$:

Immediately implies neutrinos are Majorana particles, regardless of mechanism.

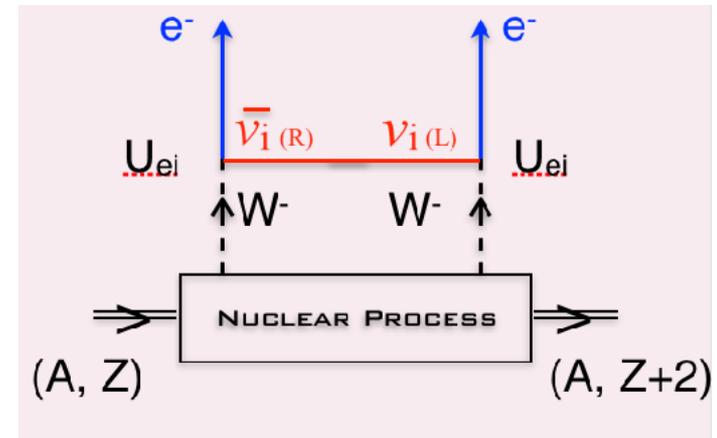
Demonstrates Lepton Number non-conservation.

Provides a model-dependent measurement of absolute neutrino mass.

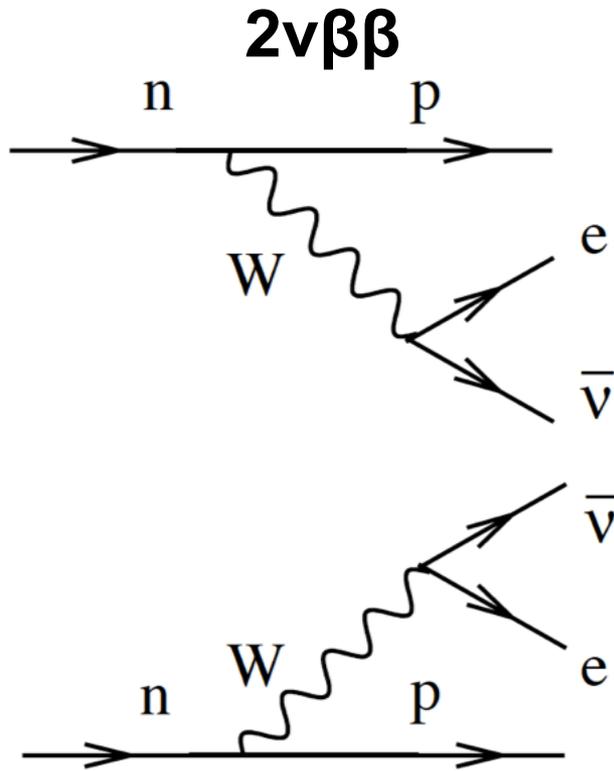
Allows access to CP-violating Majorana phases in the PMNS neutrino mixing matrix.

Offers an explanation for the lightness of neutrinos as compared to the other standard model fermions.

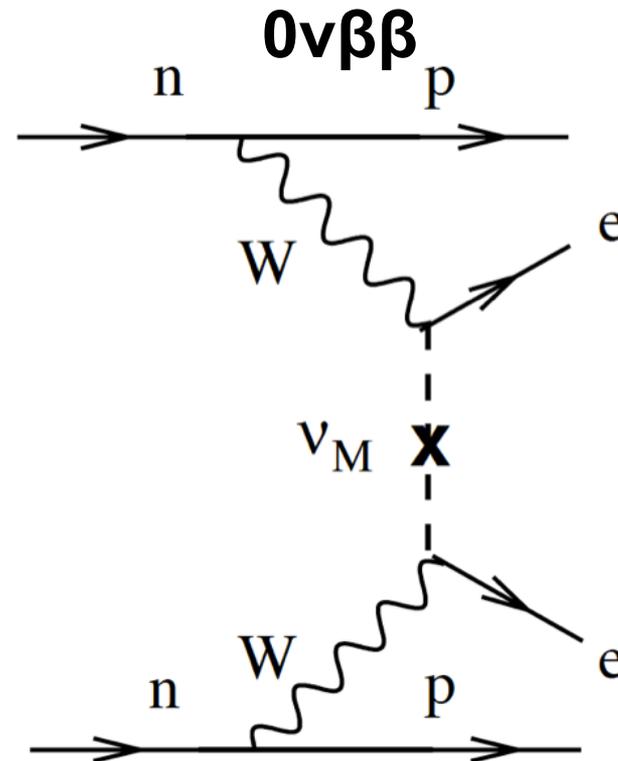
Motivates plausible mechanisms for the matter/antimatter asymmetry in the Universe.



Two Modes of Double-Beta Decay

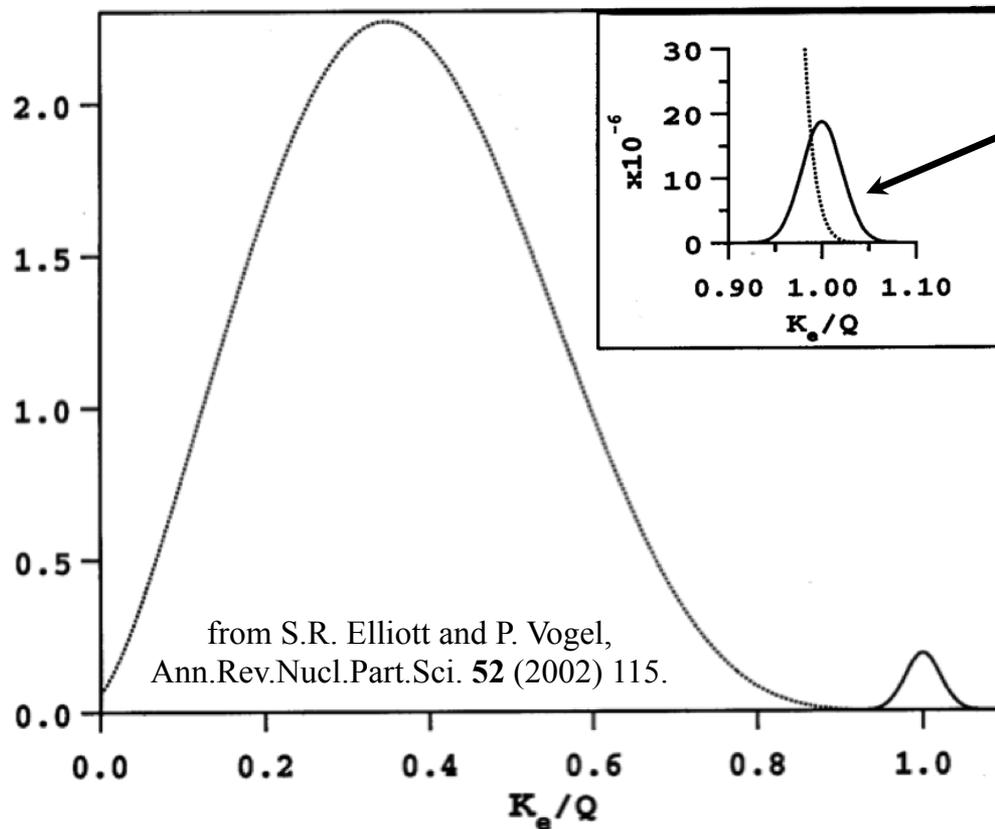


Standard-Model process
 Observed in several nuclei
 Lifetimes $\sim 10^{21}$ years



Requires Majorana neutrinos
 Lepton-Number non-conserving
 Lifetimes $\geq 10^{26}$ years

Experimental Signature



$0\nu\beta\beta$ peak
(5% FWHM)
(normalized to 10^{-6})

Summed betas yield
monochromatic peak
at Q-value.

Energy resolution
separates $0\nu\beta\beta$ signal
from irreducible $2\nu\beta\beta$
background.

Effective Majorana Masses

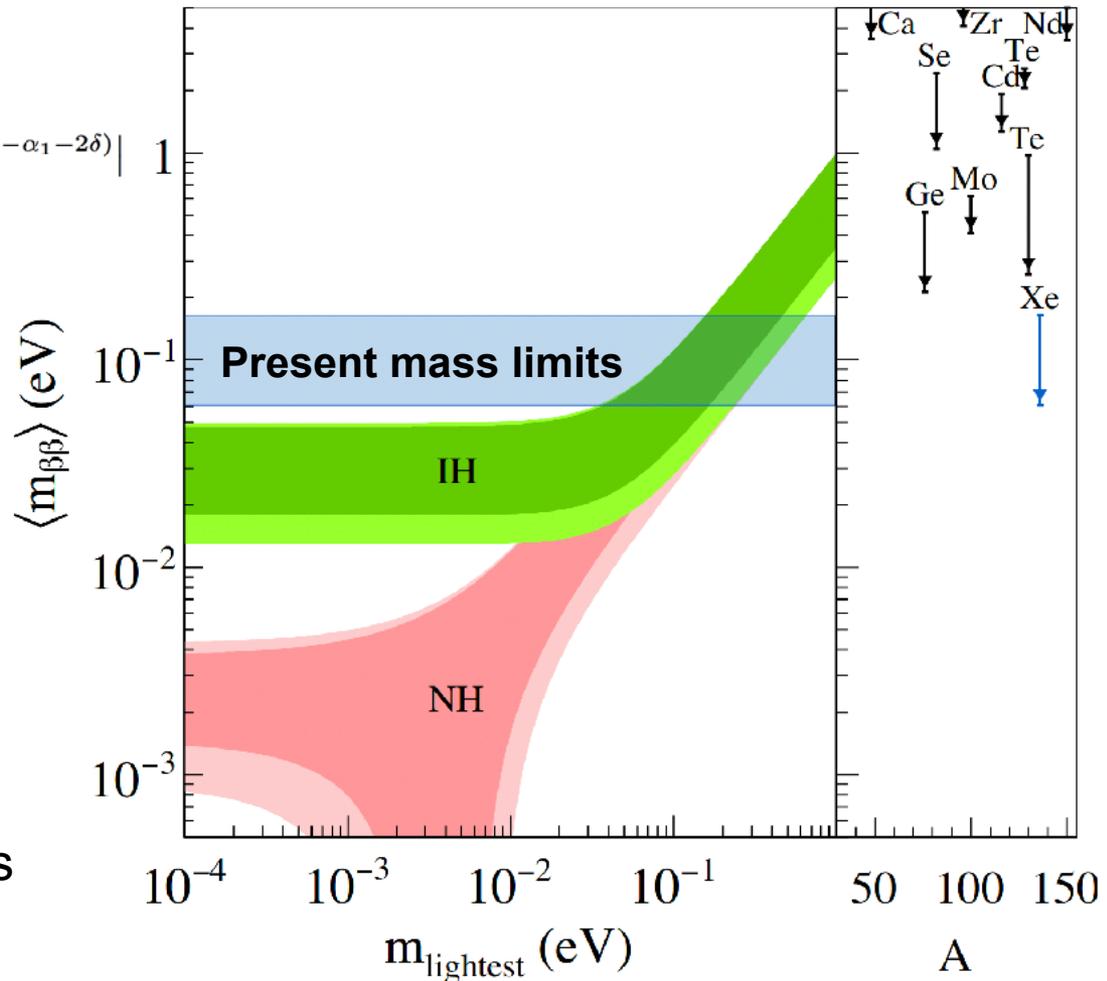
$$m_{\beta\beta} \equiv \left| \sum_k m_k U_{ek}^2 \right|$$

$$= |m_1 |U_{e1}|^2 + m_2 |U_{e2}|^2 e^{i(\alpha_2 - \alpha_1)} + m_3 |U_{e3}|^2 e^{i(-\alpha_1 - 2\delta)}|$$

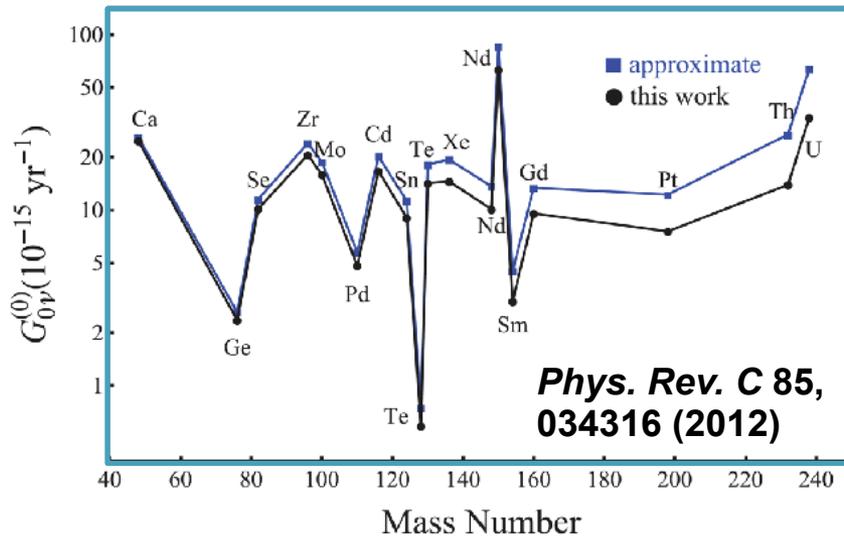
Allowed effective Majorana neutrino masses determined by mass ordering, mixing parameters, m_{lightest} .

$0\nu\beta\beta$ mass measurements complimentary to direct mass measurements (^3H -decay) and cosmological inferences.

Coverage of Inverted Ordering region motivates sensitivity aims of next-generation experiments.



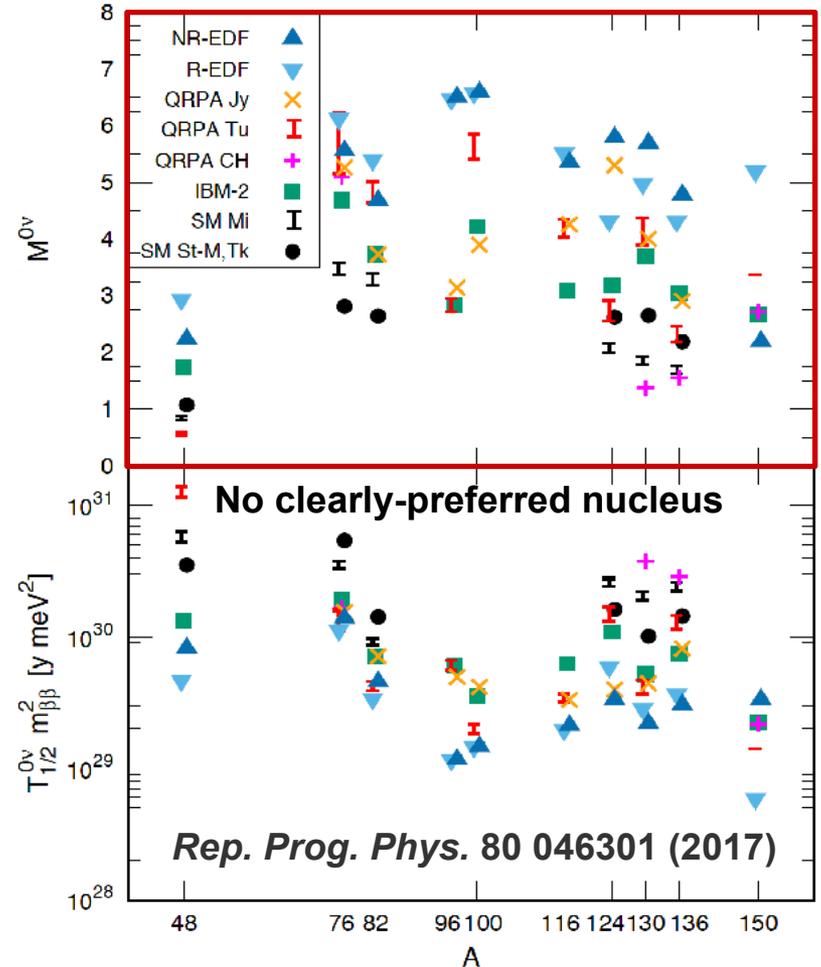
Theoretical Considerations



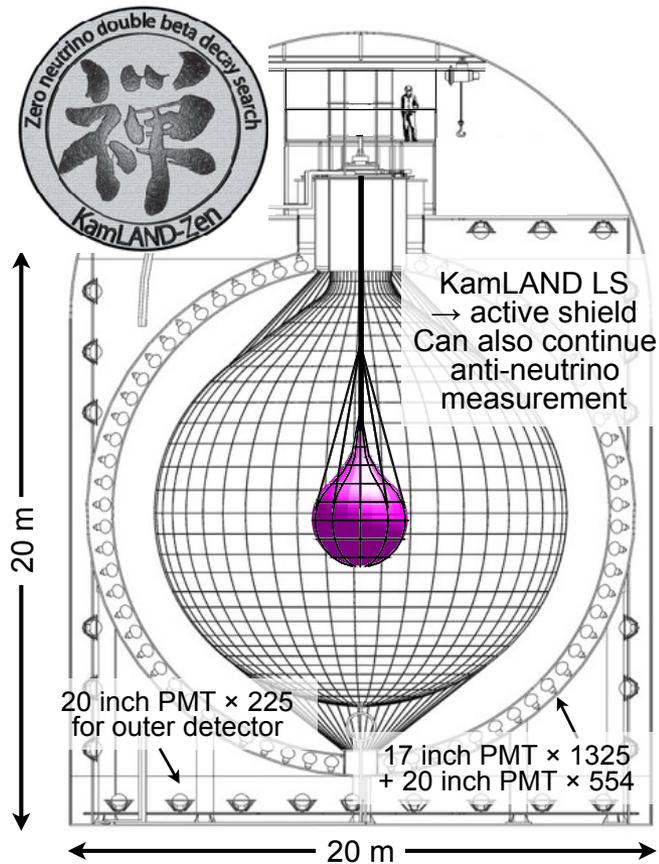
$$[T_{1/2}^{0\nu}]^{-1} = G_{0\nu}(Q, Z) |M_{0\nu}|^2 m_{\beta\beta}^2$$

$$|M_{0\nu}|^2 \propto g_a^2$$

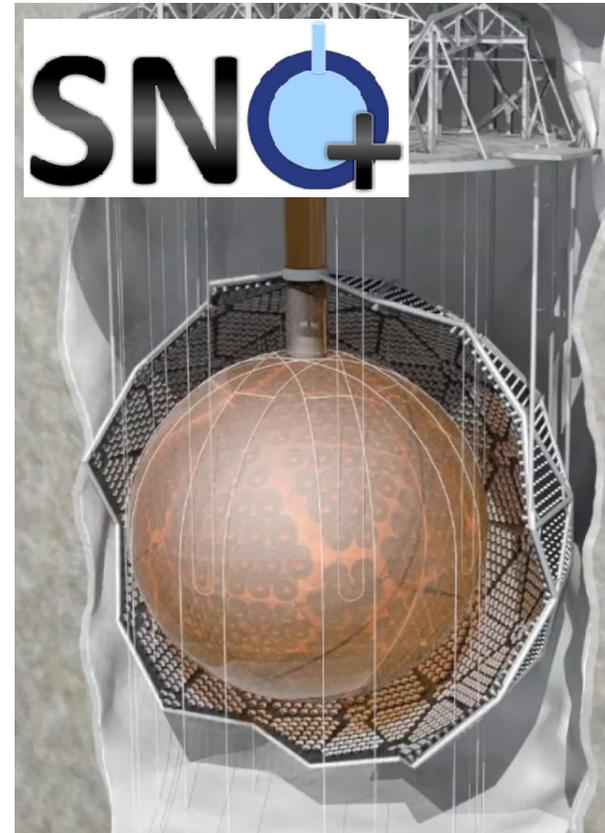
Unknown quenching of g_a , uncertainty of NME and dominant LNV mechanism can impact mass interpretation.



Large Liquid Scintillator Experiments: Making use of existing large detectors



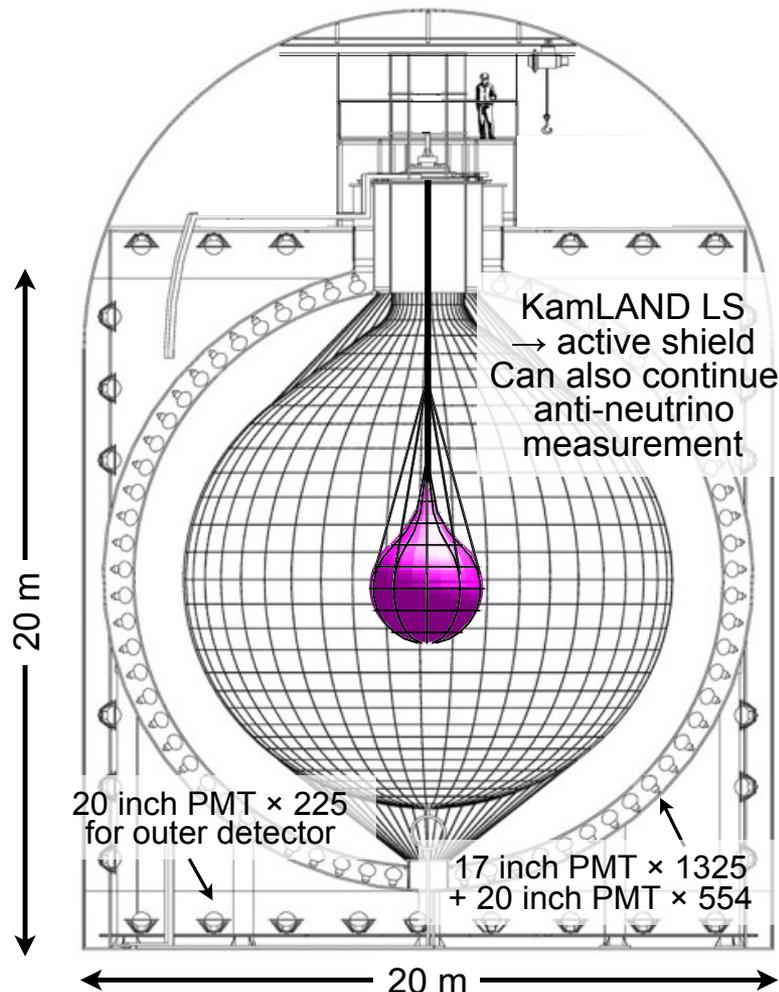
KamLAND-Zen
Xe-doped LS



SNO+
Te-doped LS

KamLAND-Zen

^{136}Xe loaded LS → into KamLAND center with inner balloon.



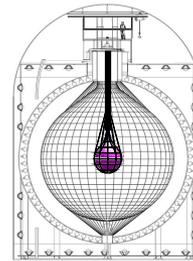
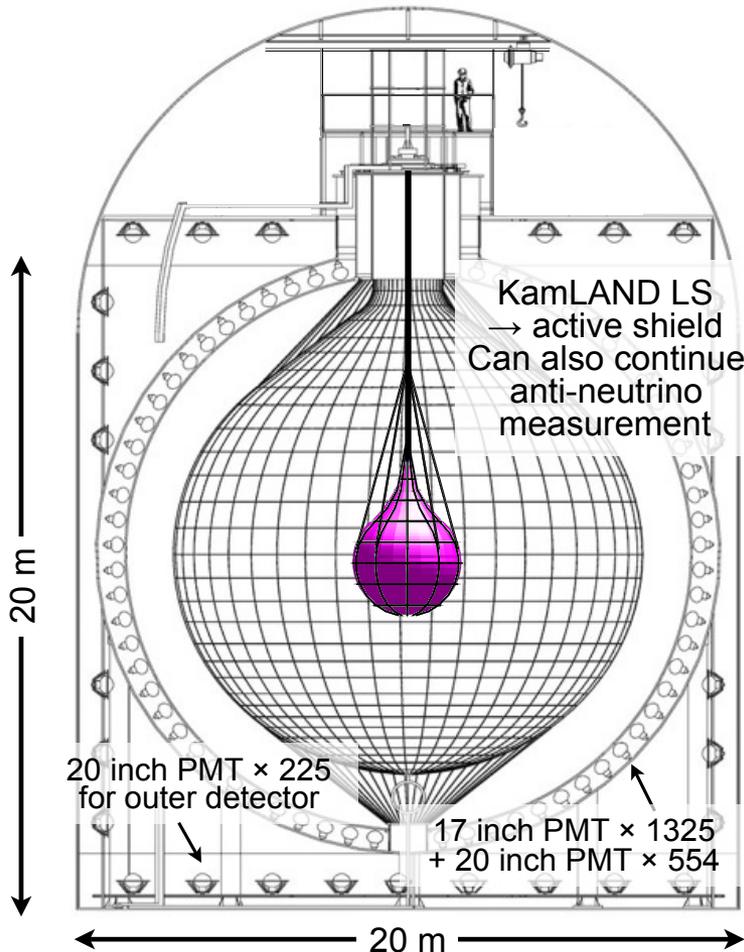
Double beta decay isotope: ^{136}Xe

- Q-value 2.458 MeV
- Dissolved into LS ~3% by weight
- Extract from LS (blank measurement)
- Noble gas → chemical stability
- Enrichment ~90%
- Purification method established
- Half life of $2\nu\beta\beta$ decay is long ($\sim 10^{21}$ yr)

Inner balloon

- 25 μm thickness clean nylon
- ^{238}U , ^{232}Th \sim a few $\times 10^{-12}$ g/g
- Enough strength & transparency
- Heat welding (no glue)
- Production in class-1 clean room

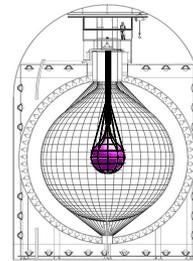
Three Phases of KamLAND-Zen



Past

KamLAND-Zen 400

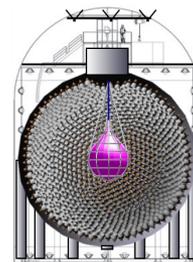
320-380 kg of Xenon
Data taking 2011 ~ 2015



Present

KamLAND-Zen 800

~750 kg of Xenon
DAQ to start in this year



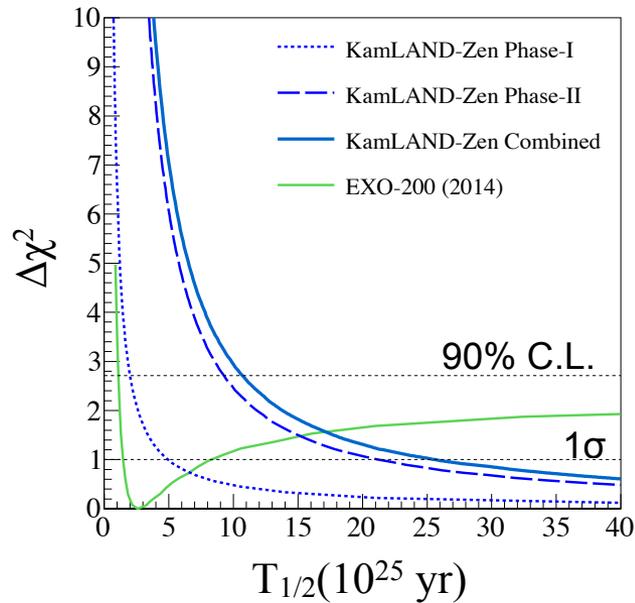
Future

KamLAND2-Zen

~1 ton of ^{136}Xe
Better energy resolution

KamLAND-Zen $0\nu\beta\beta$ Search Results

^{136}Xe Half-life limit



Combined result (90% C.L.)

$$T^{1/2} > 1.07 \times 10^{26} \text{ yr}$$

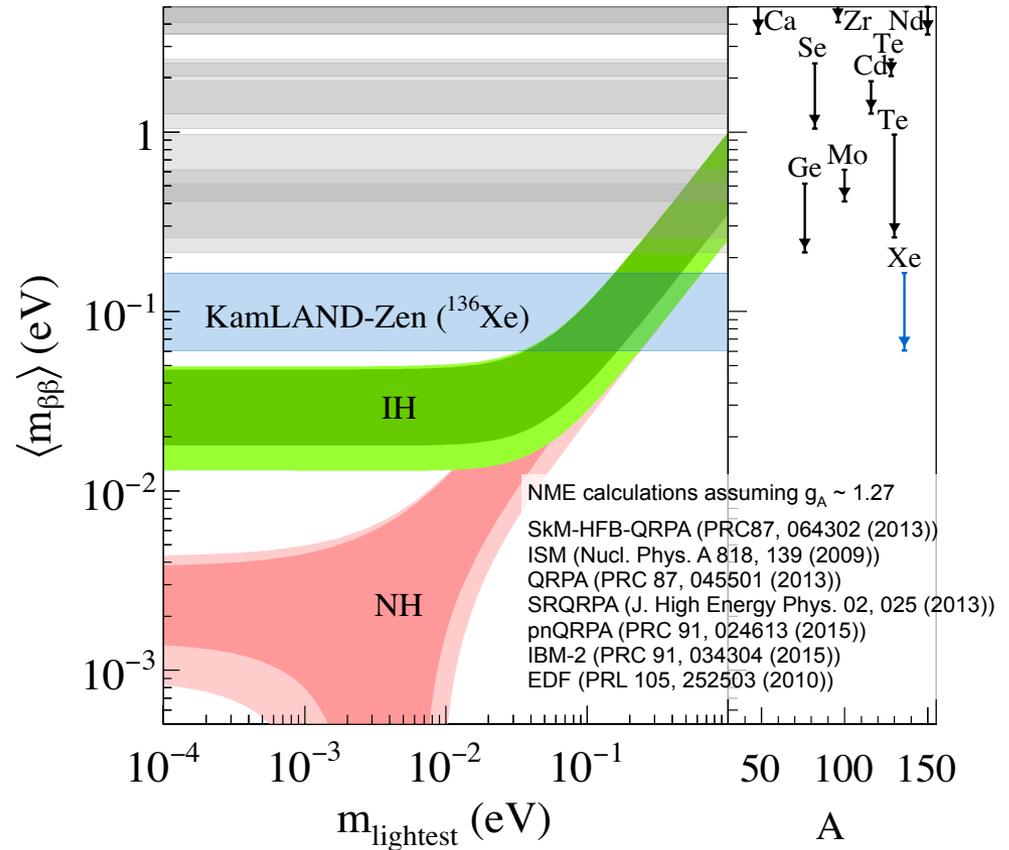
Phase-I: $T^{1/2} > 1.9 \times 10^{25} \text{ yr}$

Phase-II: $T^{1/2} > 9.2 \times 10^{25} \text{ yr}$

PRL 117, 082503 (2016)

Gando, Azusa - doi.org/10.5281/zenodo.1286895

Limit for effective neutrino mass



$$\langle m_{\beta\beta} \rangle < 61-165 \text{ meV}$$

It reaches below 100 meV!

KamLAND2-Zen

- Enlarge opening
- General use: accommodate various devices such as CdWO₄, NaI, CaF₂ detectors

R&D to improve the energy resolution

Winstone cone & High QE PMT

Improve light collection efficiency and photo coverage

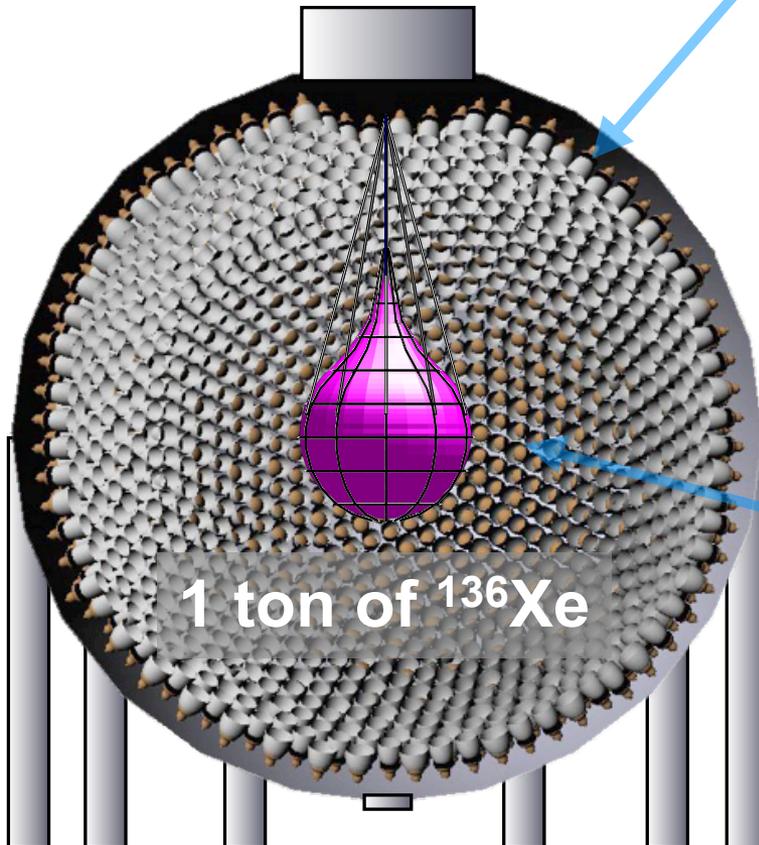
×1.9



17" PMT



20" PMT



1 ton of ¹³⁶Xe

Brighter LS

×1.4

Current LS ~8,000 photon/MeV

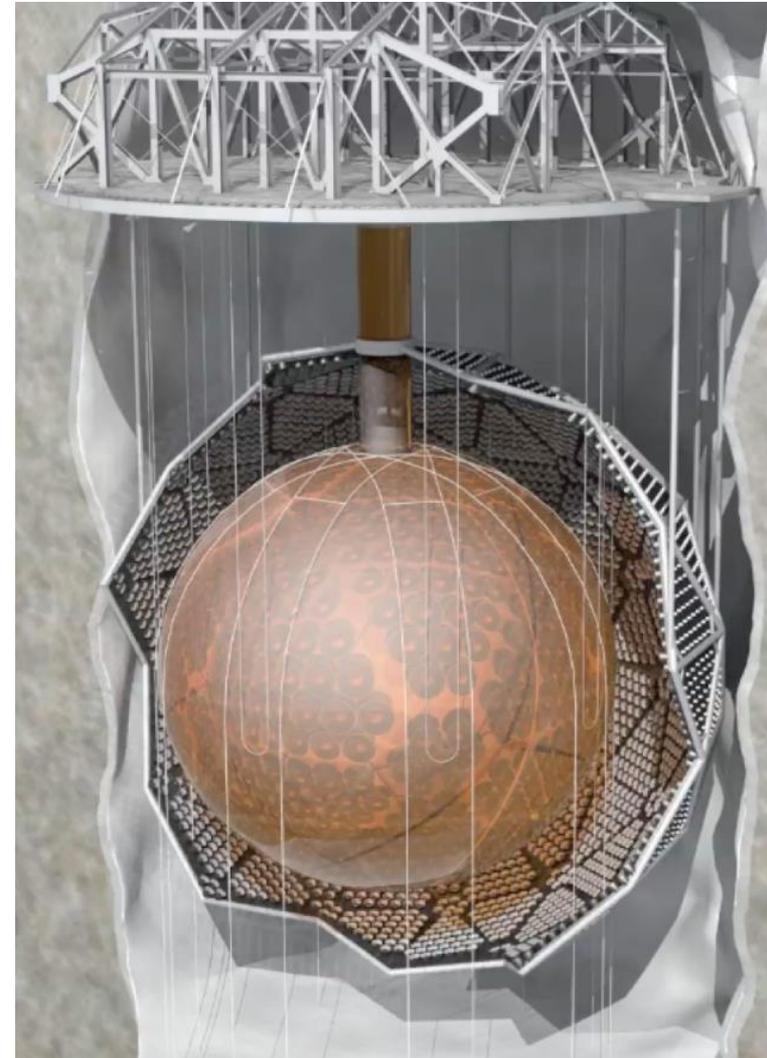
LAB based new LS ~12,000 photon/MeV

$\sigma(2.6\text{MeV})=4\% \rightarrow < 2.5\%$
Target $\langle m_{\beta\beta} \rangle \sim 20\text{meV}$ in 5 yrs

SNO+

- SNOLAB, Ontario
- 780 ton LAB/PPO (2g/L) in 6m radius acrylic vessel (AV)
- ~9400 PMTs at 8.5m
- Inherited SNO detector
 - Upgraded electronics & DAQ
 - New hold-down rope net
 - New underground LS & Te plants

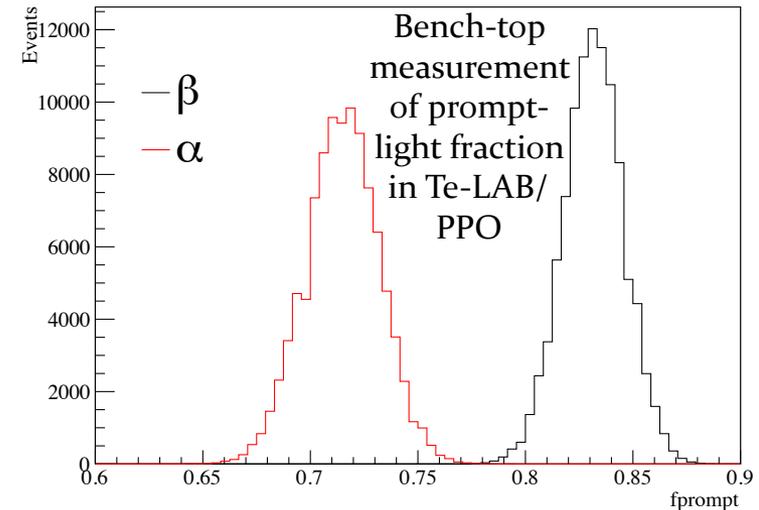
Orebi Gann, Gabriel - doi.org/10.5281/zenodo.1286908



SNO+

Method: Load LAB/PPO with 0.5% ^{130}Te

- High nat abundance of ^{130}Te
→ *enrichment unnecessary*
- High intrinsic photon yield, low absn
- Favourable $2\nu\beta\beta:0\nu\beta\beta$ NME
- Isotope-out background measurement
- High detection efficiency

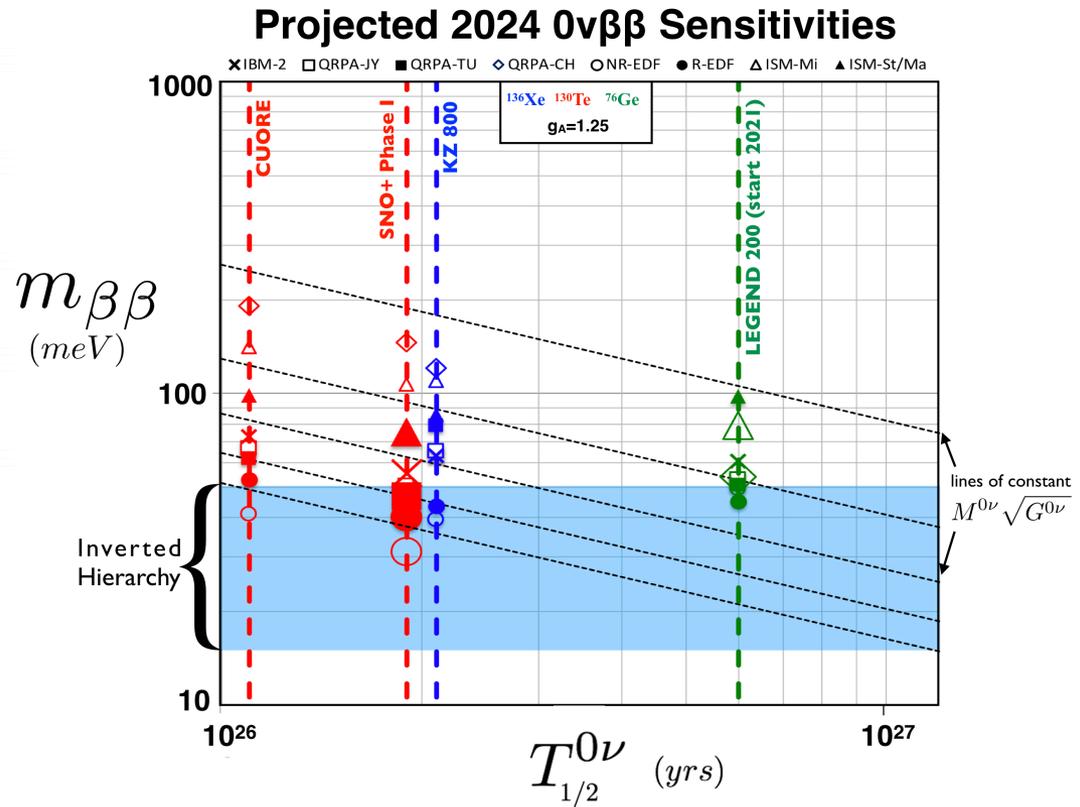
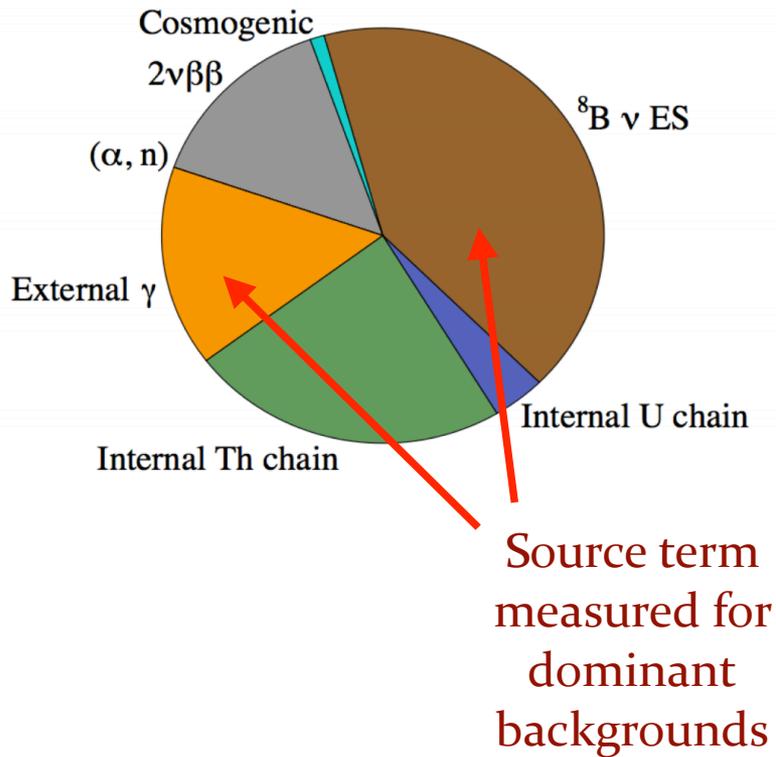


- Low backgrounds (dominated by $^8\text{B } \nu_e$)
 - ▶ Fiducialisation \Rightarrow self-shielding
 - ▶ Particle ID and coincident timing
BiPo rejection > 99.99% in ROI
 - ▶ Deep location (6000 m.w.e.)
- Large target mass, easy scaling

Orebi Gann, Gabriel - doi.org/10.5281/zenodo.1286908

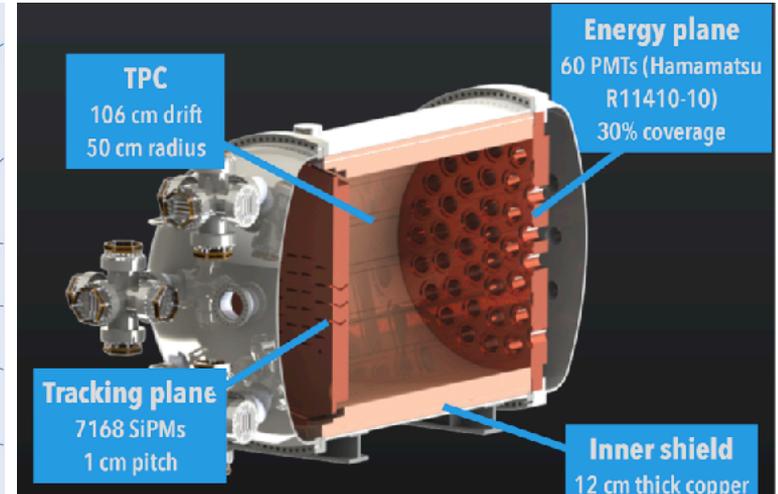
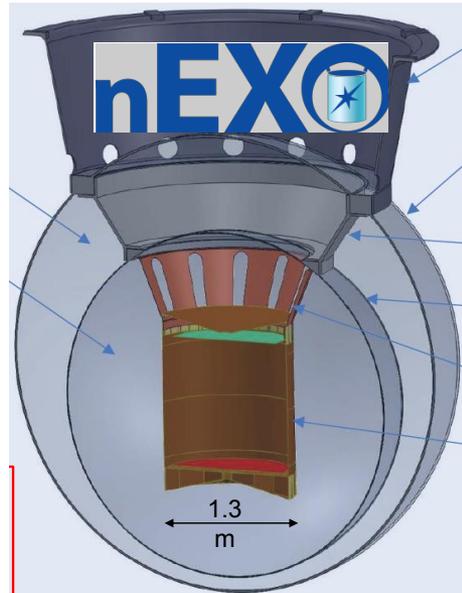
SNO+ Towards $0\nu\beta\beta$

Background levels consistent with or below “nominal” value used for sensitivity projections



Orebi Gann, Gabriel - doi.org/10.5281/zenodo.1286908

Xenon Time Projection Chambers



Monolithic TPCs with little/no background-inducing material in central volume, possibility of tagging Ba daughter.

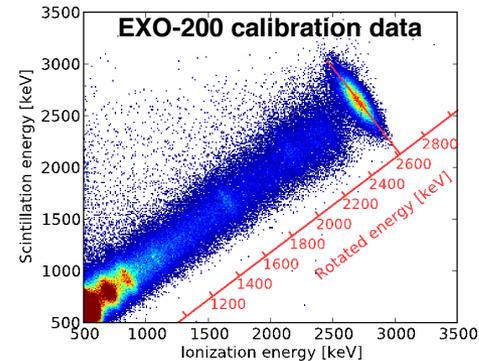
EXO-200, nEXO: Liquid Xe

NEXT: High-pressure gaseous Xe

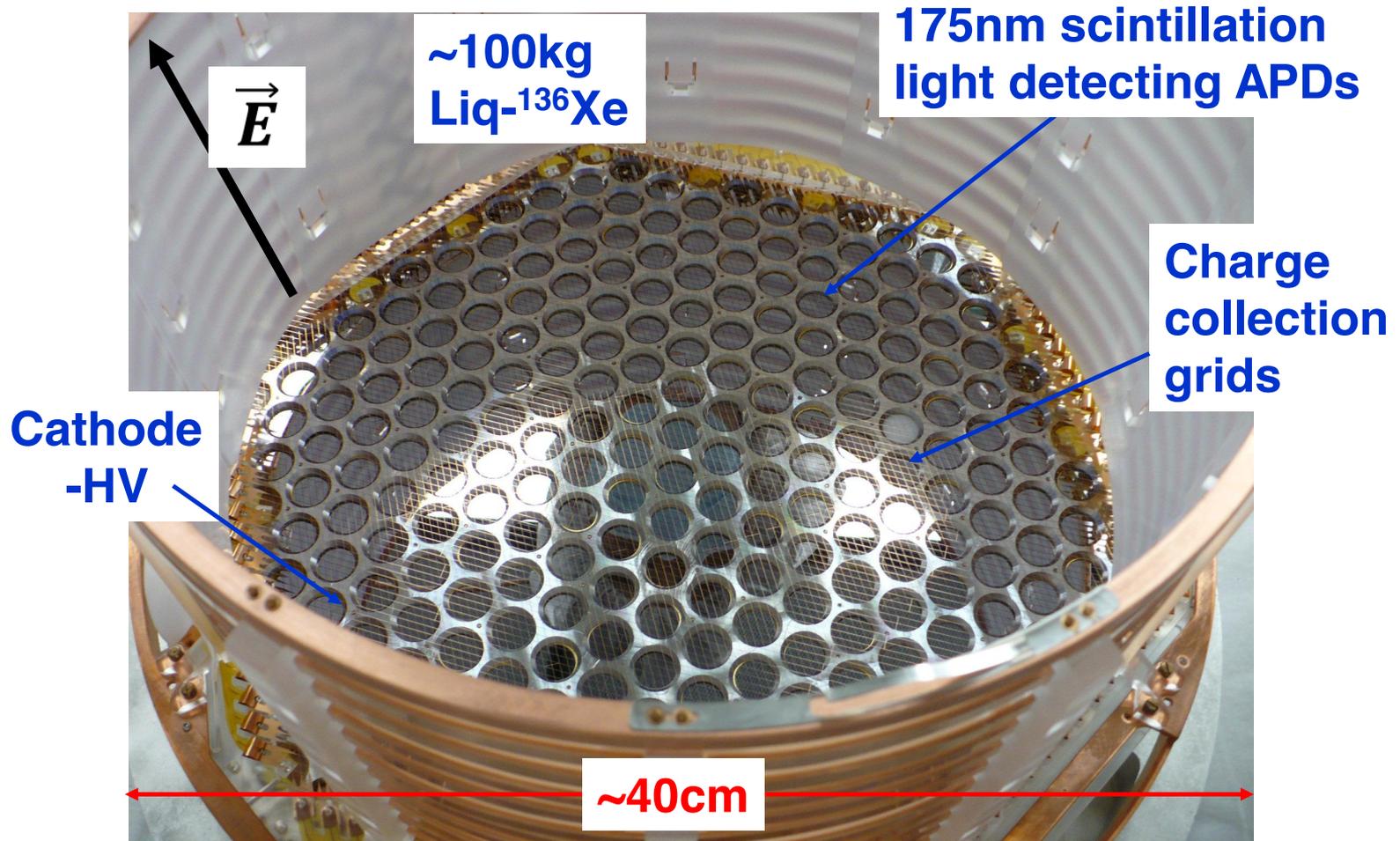
EXO: Enriched Xenon Observatory

The EXO program

- Use ^{136}Xe in liquid phase
- Initial R&D on energy resolution using scintillation-ionization correlation
- Build EXO-200, first 100kg-class experiment to produce results. Phase II in progress, will end in Dec 2018
- Build the 5-tonne nEXO, reaching $T_{1/2} \sim 10^{28}$ yr and entirely covering the Inverted Hierarchy
- Develop a technique for tagging the final state Ba as a possibility to further upgrade nEXO and substantially exceed $T_{1/2} = 10^{28}$ yr

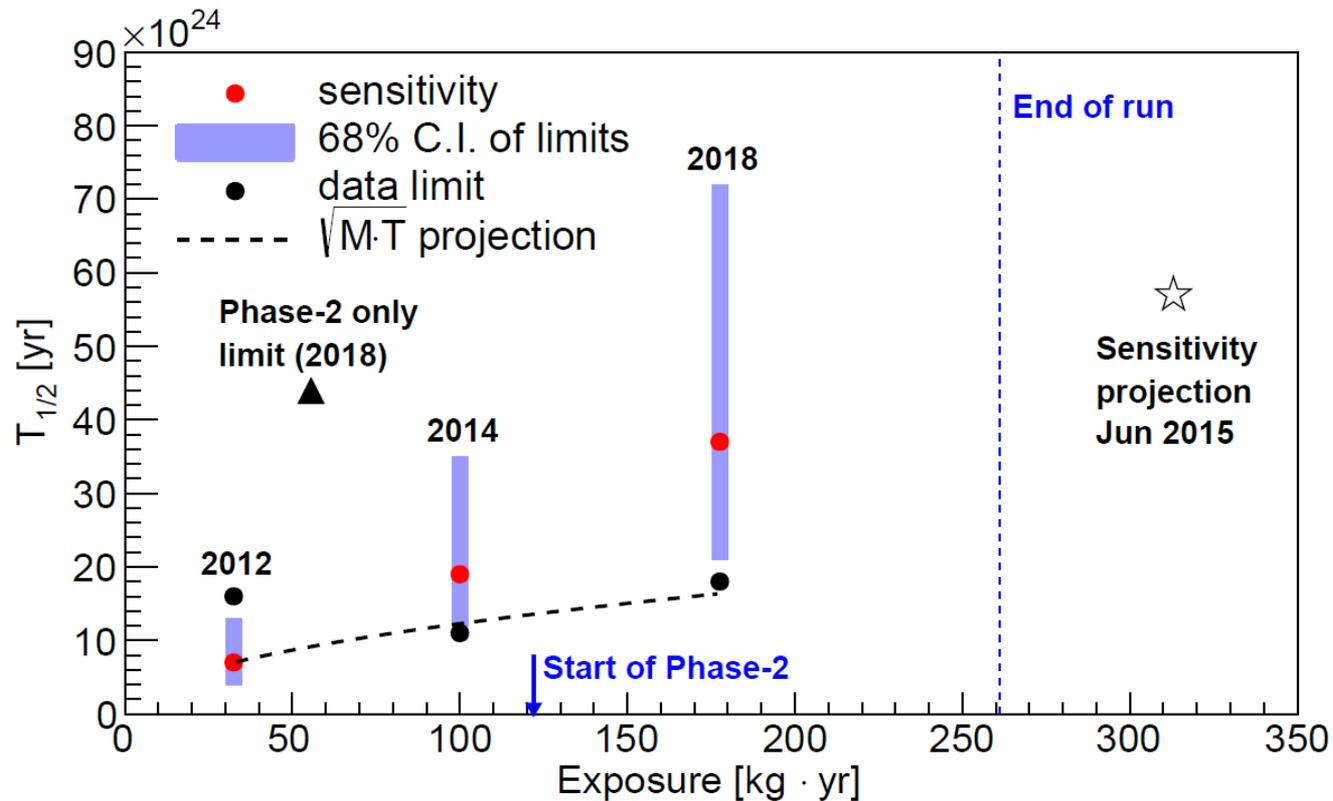


EXO-200 Time Projection Chamber

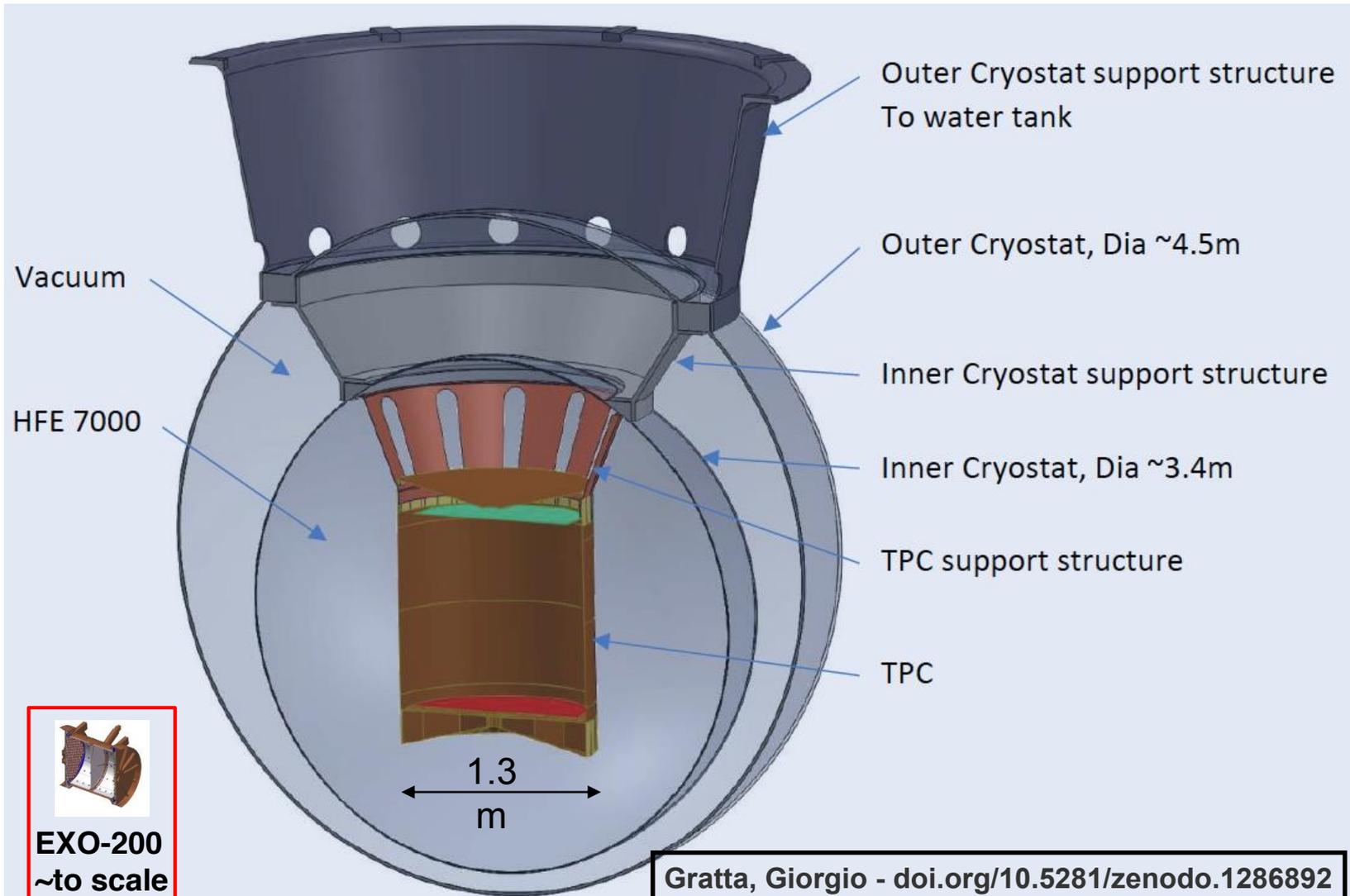


EXO-200 $0\nu\beta\beta$ Results

	Sensitivity (yr)	90% CL Limit (yr)	$\langle m_{\beta\beta} \rangle$ (meV)
PRL 109, 032505 (2012)	0.7×10^{25}	1.6×10^{25}	
Nature 510, 229 (2014)	1.9×10^{25}	1.1×10^{25}	
PRL 120 072701 (2018)	3.8×10^{25}	1.8×10^{25}	147-398

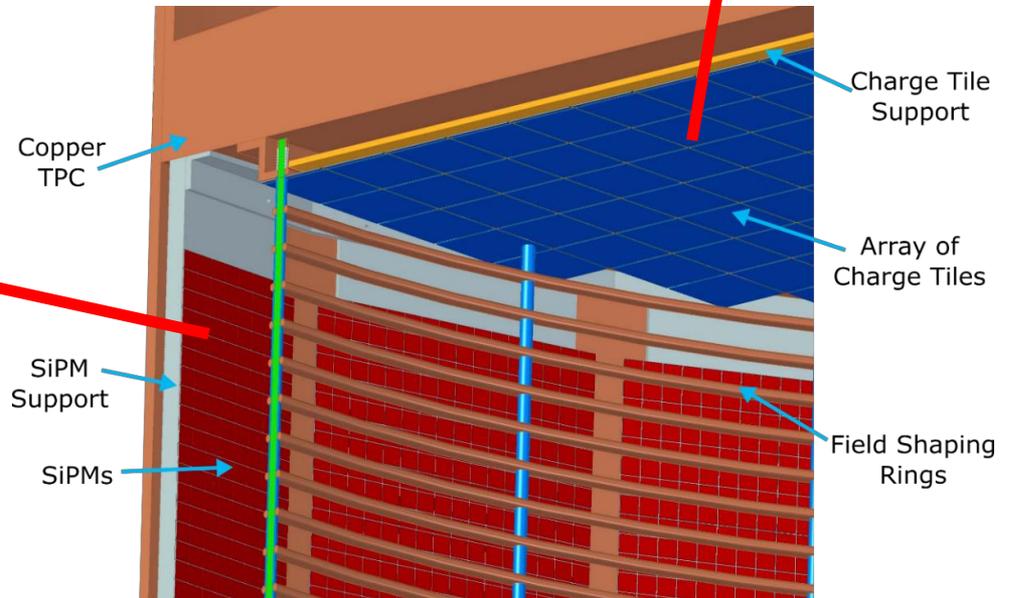
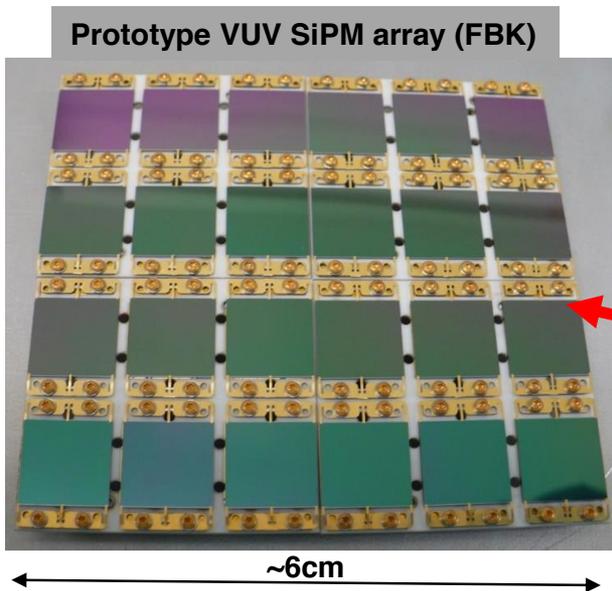
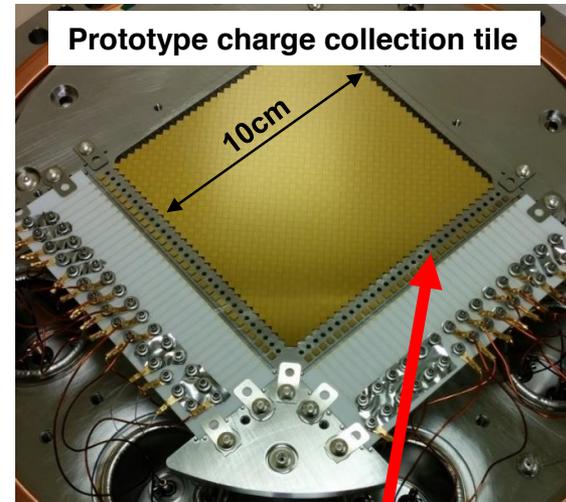


nEXO

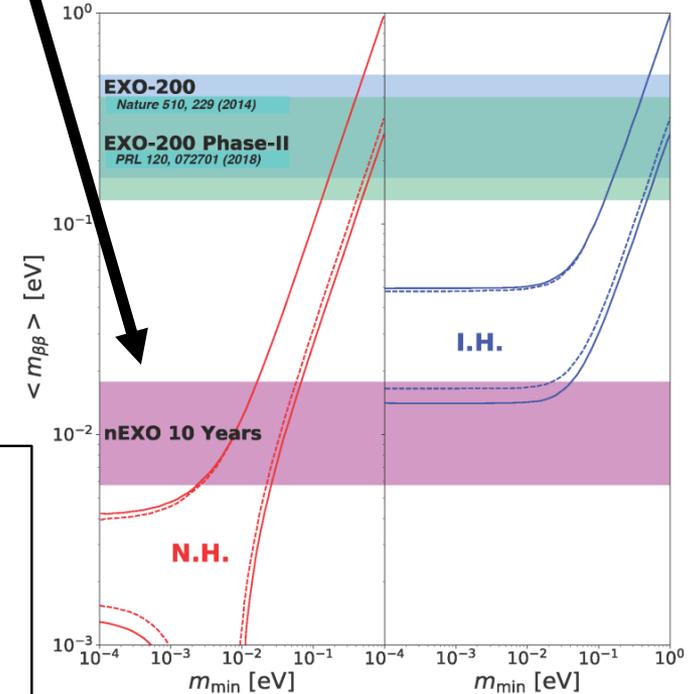
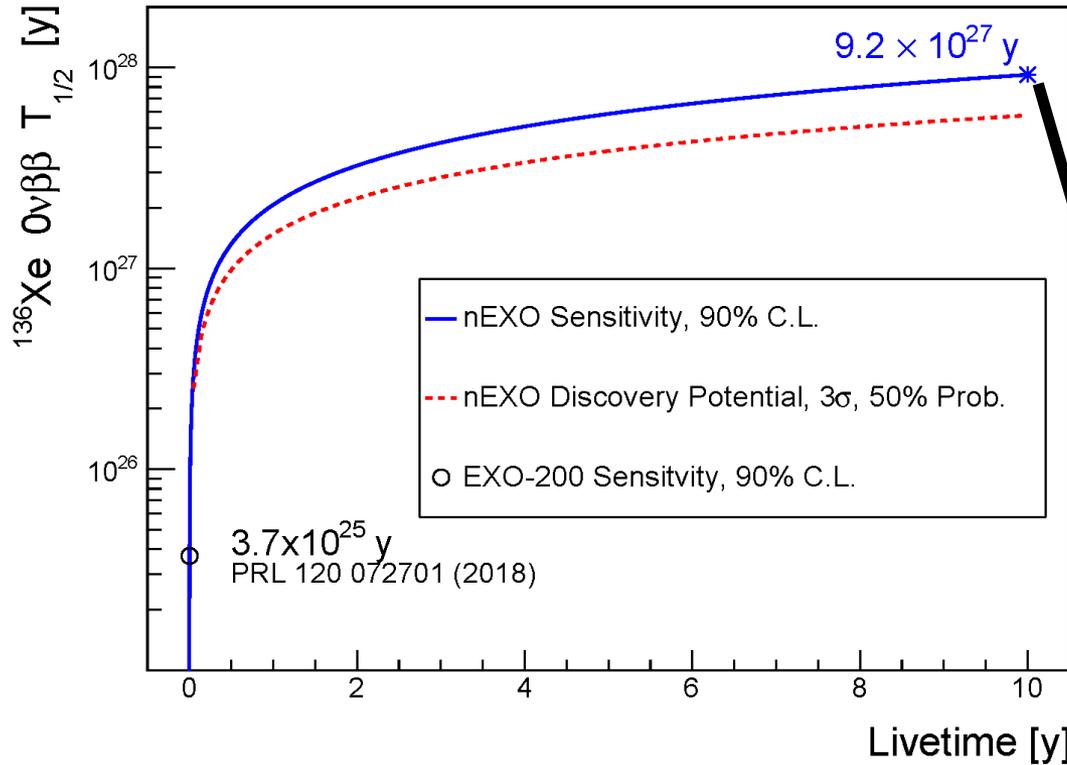


Major changes on the EXO-200 theme

- Only one drift volume
- ASIC electronics in LXe
- Silica substrate charge collection tiles
- VUV SiPMs ($\sim 4.5\text{m}^2$)
- Little plastics in the TPC (Sapphire, Silica)



nEXO Sensitivity



- $g_A = g_A^{\text{free}} = -1.2723$

- Band is the envelope of NME:

EDF: T.R. Rodríguez and G. Martínez-Pinedo, PRL 105, 252503 (2010)

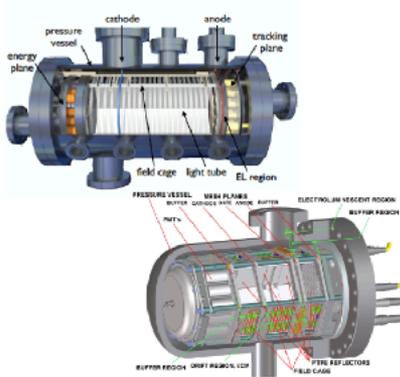
ISM: J. Menendez et al., Nucl Phys A 818, 139 (2009)

IBM-2: J. Barea, J. Kotila, and F. Iachello, PRC 91, 034304 (2015)

QRPA: F. Šimković et al., PRC 87 045501 (2013)

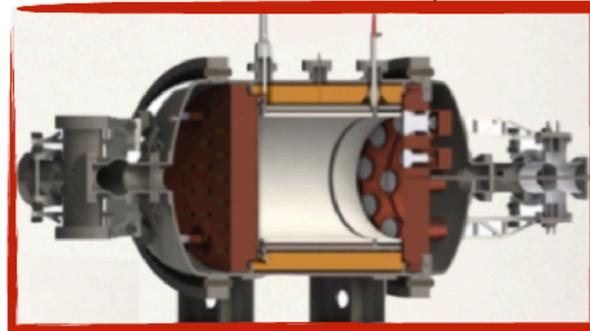
SkyrmeQRPA: M.T. Mustonen and J. Engel PRC 87 064302 (2013)

NEXT: Neutrino Experiment with a Xenon TPC



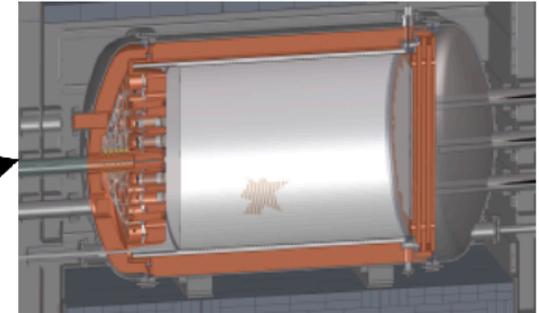
Prototypes (~1 kg)
[2009 - 2014]

Demonstration of detector concept
[<1% FWHM, tracking]



NEXT-NEW (~5 kg)
[2015 - 2018]

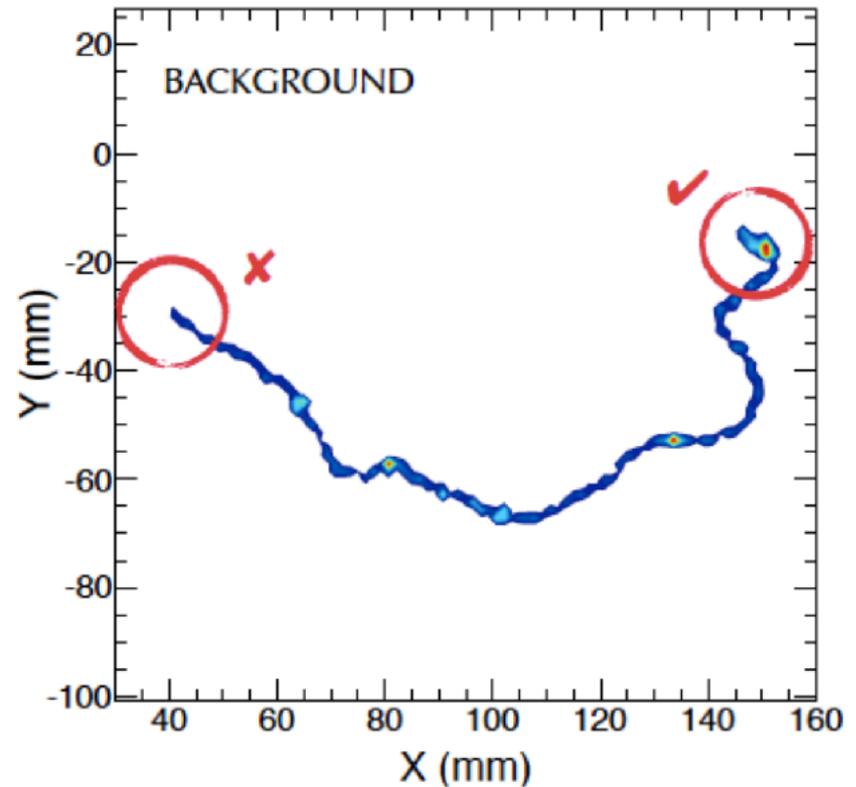
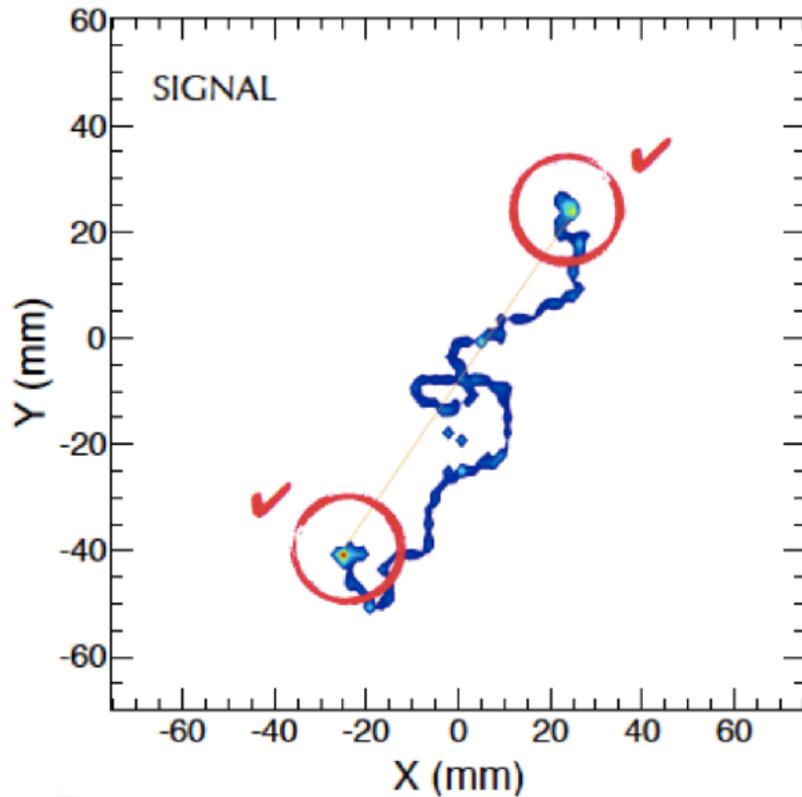
Underground and radio-pure operations, background, $2\nu\beta\beta$



NEXT-100 (~100 kg)
sensitivity: 6×10^{25} yr
[2018 - 2020's]

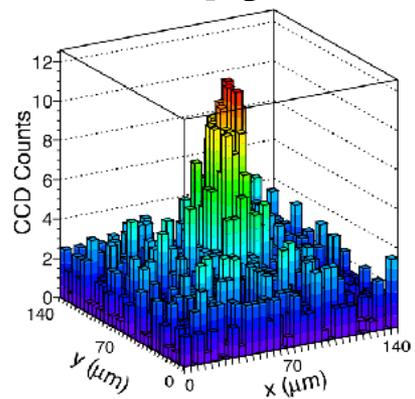
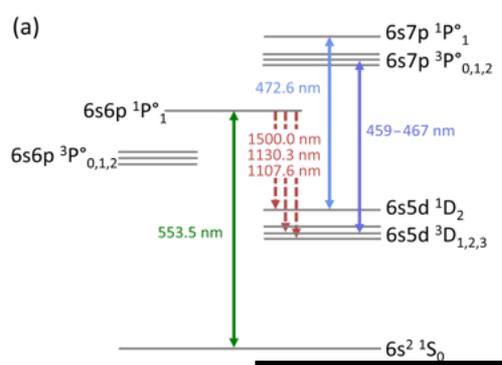
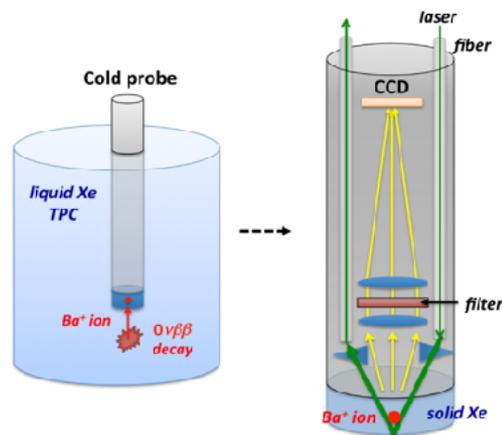
Neutrinoless double beta decay searches

NEXT: Background Suppression through Event Topology



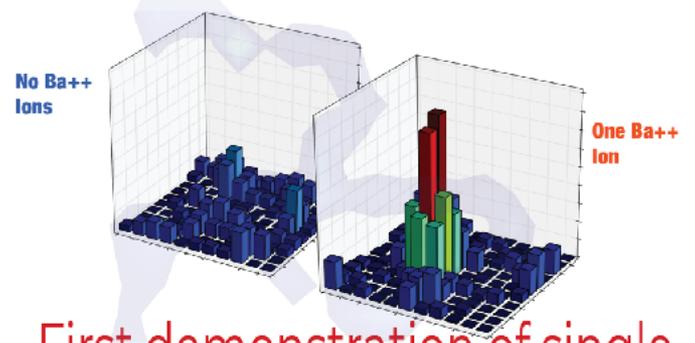
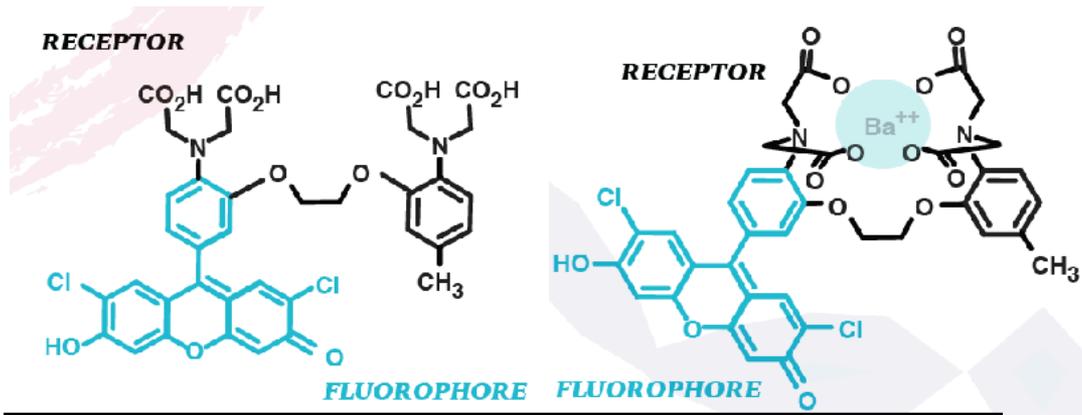
Tagging of the Ba Daughter

EXO: Laser-Induced Fluorescence Spectroscopy



B. Mong, et al. - doi.org/10.1103/PhysRevA.91.022505

NEXT: Single-Molecule Fluorescence Imaging



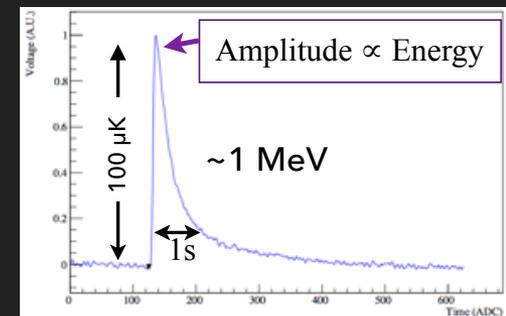
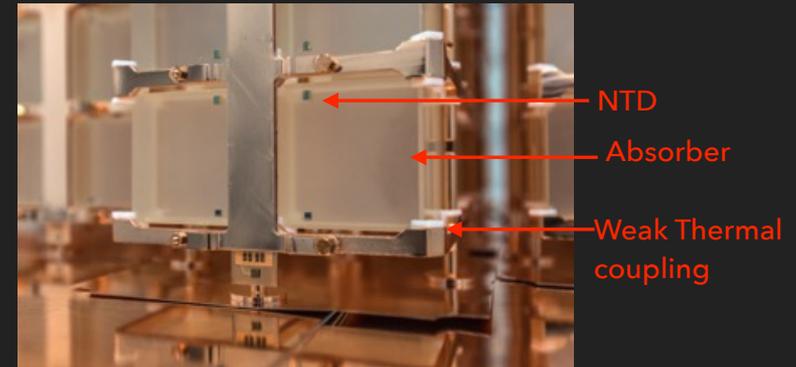
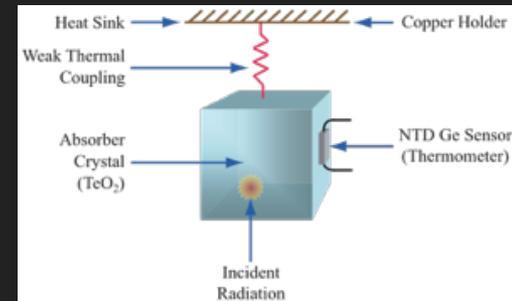
First demonstration of single Ba⁺⁺ ion sensitivity.

A. McDonald & F. Psihas - doi.org/10.5281/zenodo.1300986

CUORE / CUPID: Te Bolometers

Cryogenic Bolometers

- ▶ Solid state detectors operating at low temperatures: $\sim 10\text{mK}$
- ▶ 1 MeV energy deposition causes $\sim 100\ \mu\text{K}$ increase in temperature
 - ▶ Read out with an Ge Neutron Transmutation Doped thermometer
- ▶ Extremely good energy resolution: 5 keV FWHM at 2.5 MeV
- ▶ Detector is made out of Te and contains the candidate isotope inside
- ▶ Flexible choice of candidate isotope

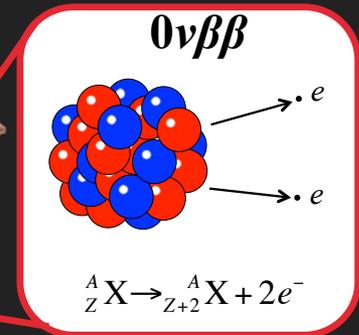
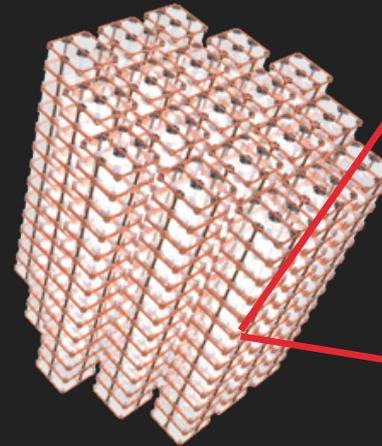


Ouellet, Jonathan - doi.org/10.5281/zenodo.1286904

CUORE / CUPID: Te Bolometers

The CUORE Experiment

- ▶ Primary physics goal is the search for $0\nu\beta\beta$ decay of ^{130}Te
- ▶ Array of 988 TeO_2 bolometers
- ▶ Located at the Gran Sasso National Lab in Italy (3600 m.w.e. overburden)
 - ▶ **Large mass** of 741 kg (206 kg of candidate isotope ^{130}Te)
 - ▶ **Energy resolution** goal of 5 keV FWHM at $Q_{\beta\beta}$ (2527 keV)
 - ▶ **Low background** goal of 10^{-2} cnts/(keV·kg·yr) at the ROI
 - ▶ High uptime and signal efficiency
- ▶ Sensitivity to $T_{1/2} = 9 \times 10^{25}$ yr in 5 years of live time



Sensitivity

$$T_{1/2} \sim \varepsilon \sqrt{\frac{Mt}{b\Delta E}}$$

Ouellet, Jonathan - doi.org/10.5281/zenodo.1286904

CUORE: Recent Results

Phys. Rev. Lett. 120, 132501 (2018)

Search for $0\nu\beta\beta$ Decay with CUORE

- ▶ Background index is consistent with expectations

$$(1.4 \pm 0.2) \times 10^{-2} \text{ cnts}/(\text{keV}\cdot\text{kg}\cdot\text{yr})$$

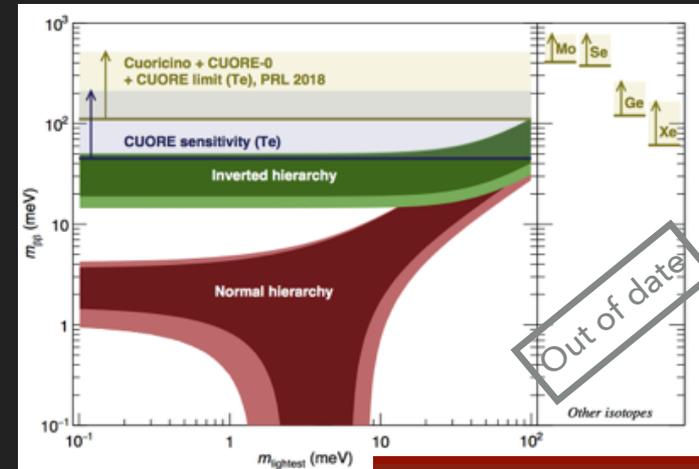
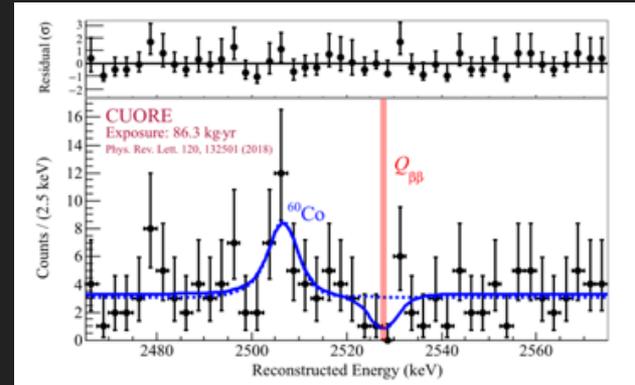
- ▶ Median expected sensitivity

$$T_{1/2}^{0\nu} = 7.0 \times 10^{24} \text{ yr}$$

- ▶ Combined limit with CUORE-0 and Cuoricino:

$$T_{1/2}^{0\nu} > 1.5 \times 10^{25} \text{ yr (90\%C.L.)}$$

$$m_{\beta\beta} < 110 - 520 \text{ meV}$$



Out of date

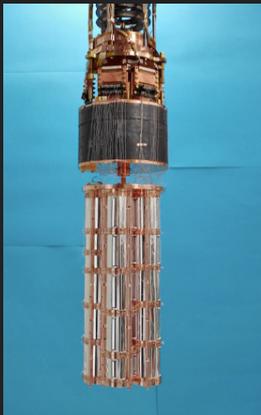
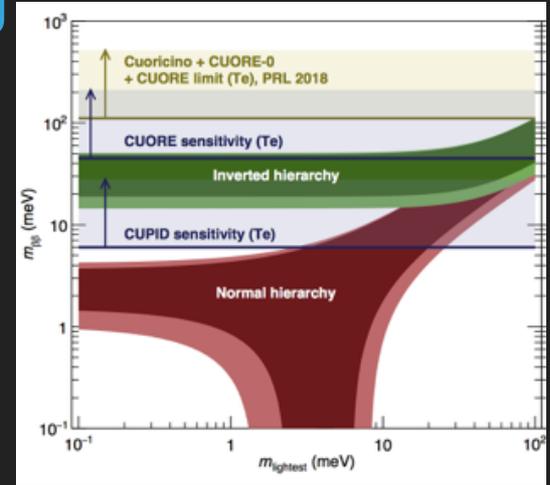
$$m_{\beta\beta} \equiv \left| \sum_i U_{ei}^2 m_i \right|$$

Ouellet, Jonathan - doi.org/10.5281/zenodo.1286904

Next Steps: CUPID

CUORE Upgrade with Particle ID (CUPID) R&D

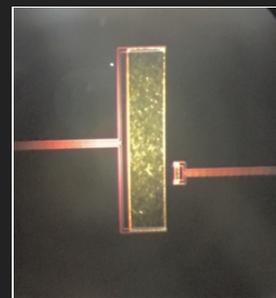
- ▶ Next generation of $0\nu\beta\beta$ decay experiments seek to be sensitive to the full IH region ($m_{\beta\beta} \sim 6 - 20$ meV, $T_{1/2} \sim 10^{27}$ yr)
- ▶ ~1000 enriched light emitting bolometers mounted in the CUORE cryostat
- ▶ Nearly zero background goal of ~ 0.1 cnts/(ROI·yr)
- ▶ Worldwide effort focused on demonstrating readiness to construct a tonne-scale bolometric experiment



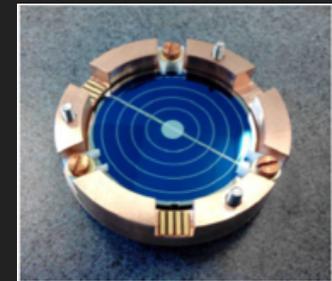
CUPID-0/Se at LNGS
 $Zn^{82}Se$



CUPID-Mo at Modane
 $Li_2^{100}MoO_4$



Neganov-Luke and
TES-based light
detectors for TeO_2

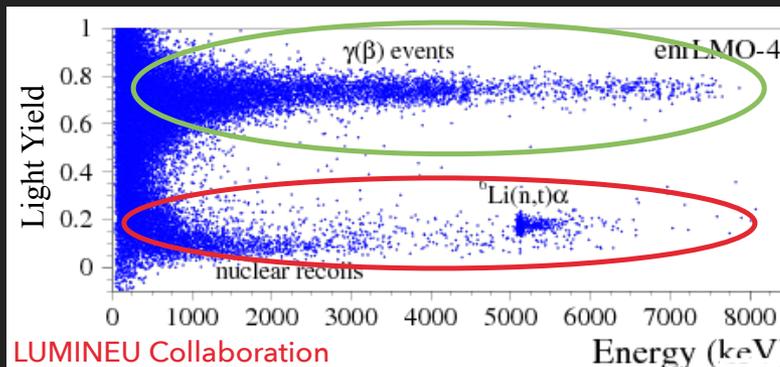
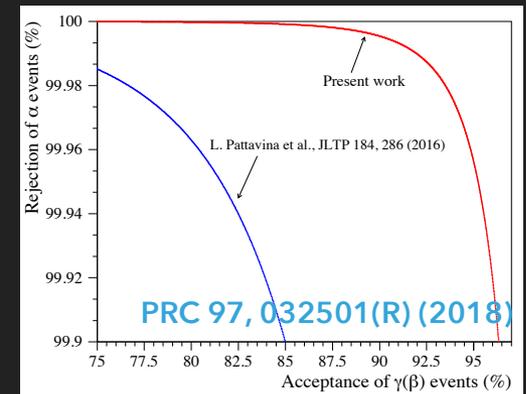
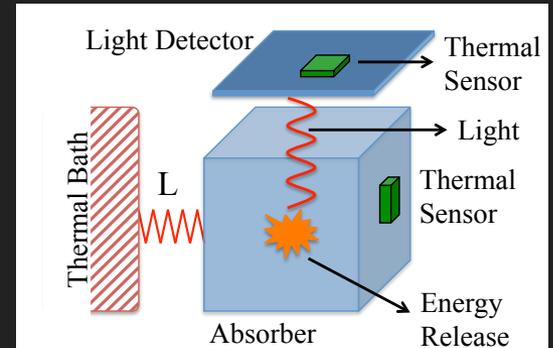


Next Steps: CUPID

CUORE Upgrade with Particle ID (CUPID) R&D

Goal of reducing the background in the ROI by rejecting all α events with particle ID

- ▶ Add the ability to read out the light emitted in a particle interaction (scintillation/Cherenkov)
- ➔ Combine the energy resolution of bolometers with the background discrimination of a dual channel detector



LUMINEU Collaboration



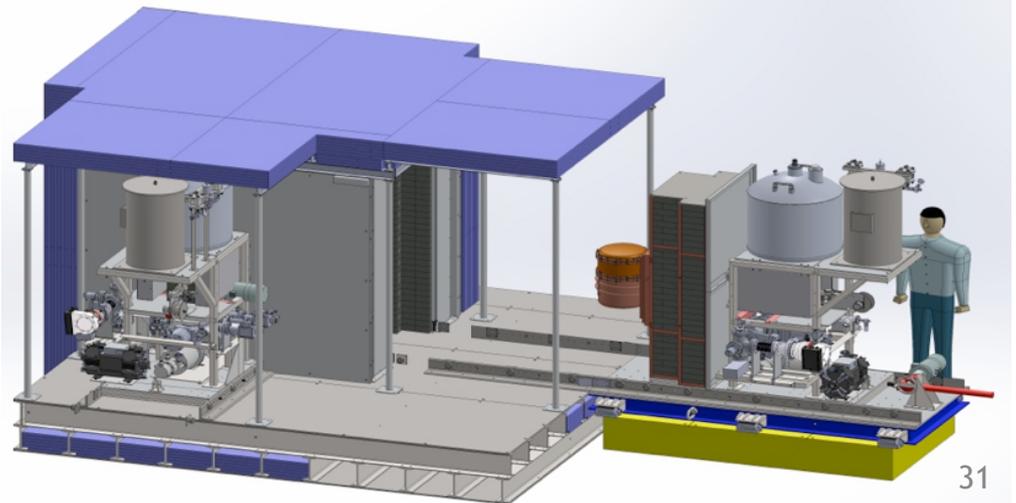
Germanium Semiconductor Detectors

GERDA:

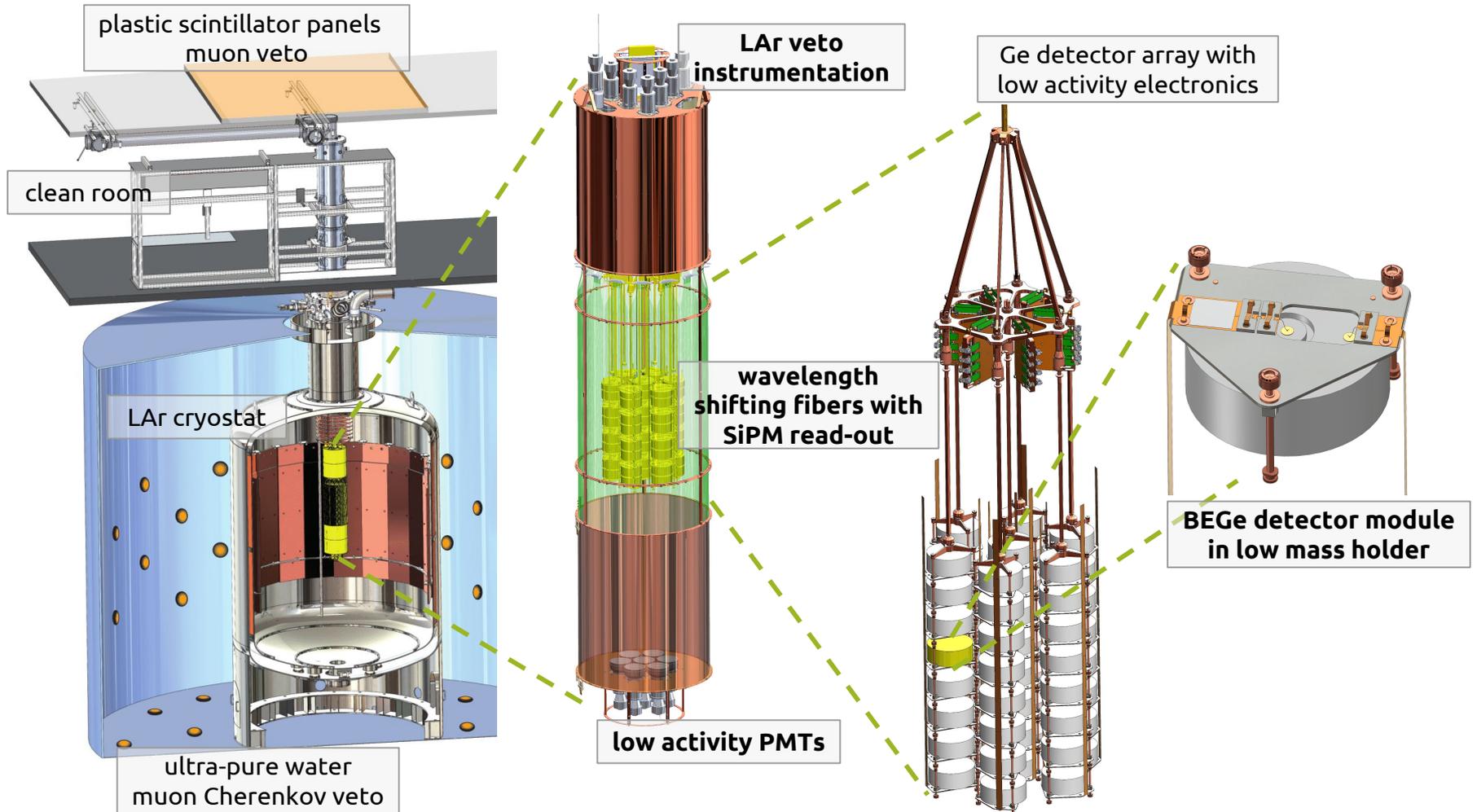
Detector immersion in Argon cryogenic bath and active shielding

MAJORANA DEMONSTRATOR:

Detectors deployed in ultra-low-background vacuum cryostats, compact shield.



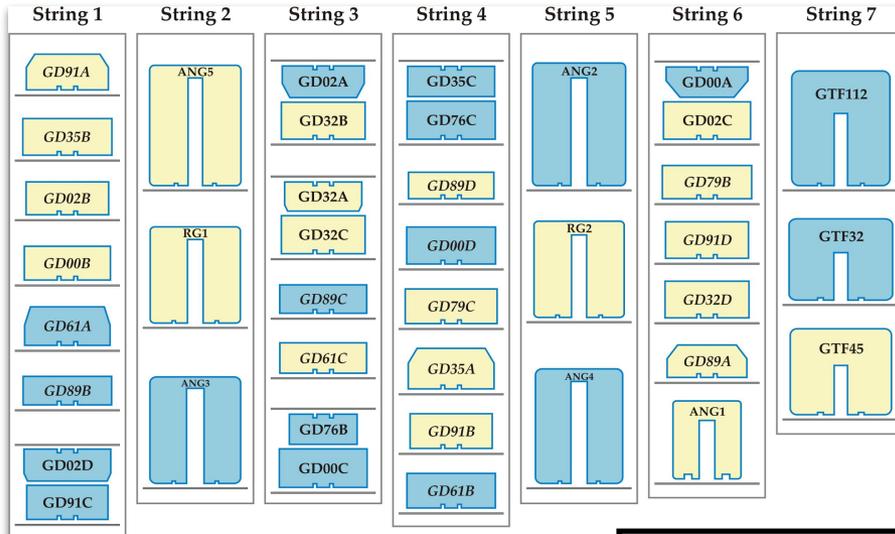
GERDA Phase II



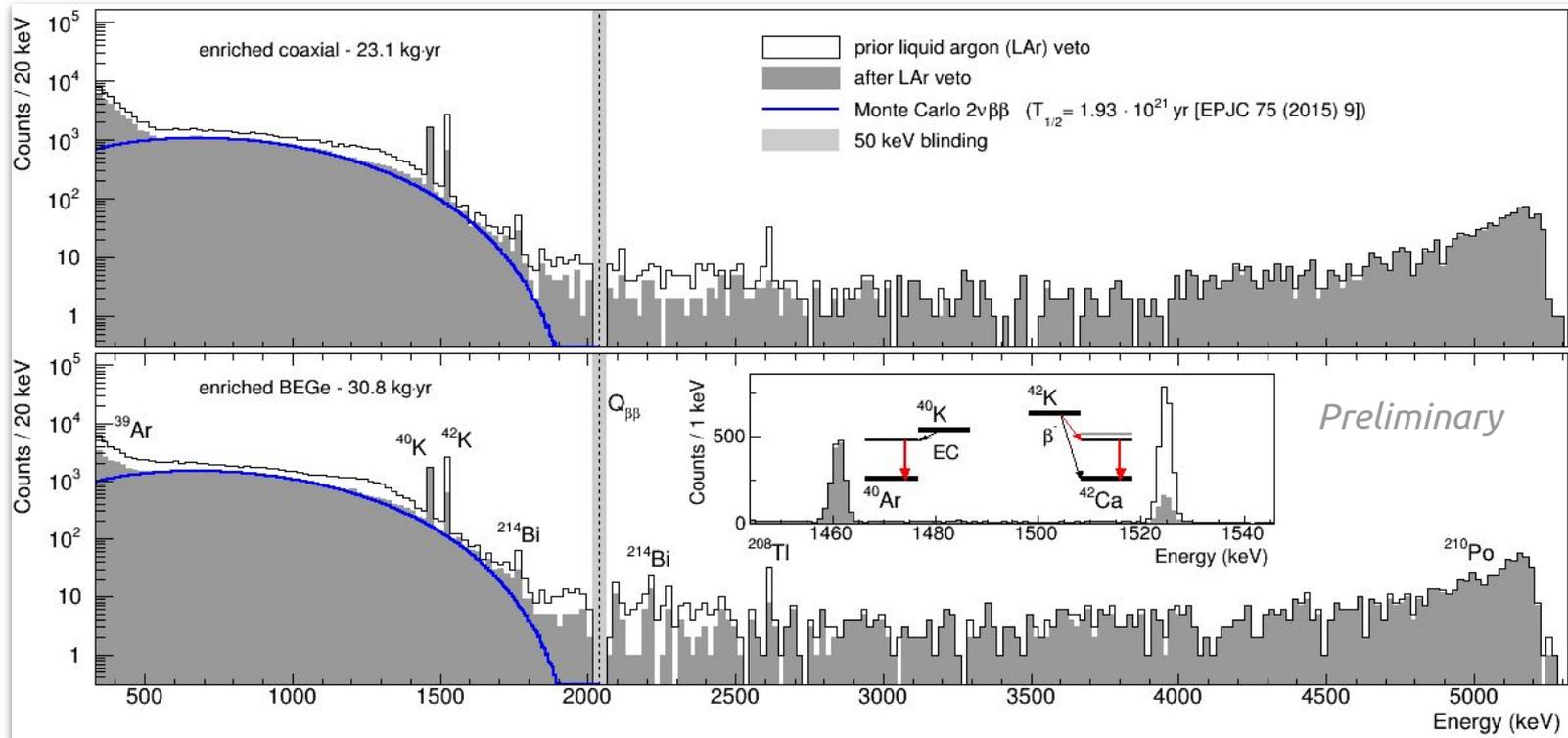
GERDA Ge Detectors

7 strings with 40 detectors in total

- 7 enriched semi-coaxial (15.6 kg)
- 30 enriched thick window BEGe (20.0 kg)
- 3 natural semi-coaxial (7.6 kg)

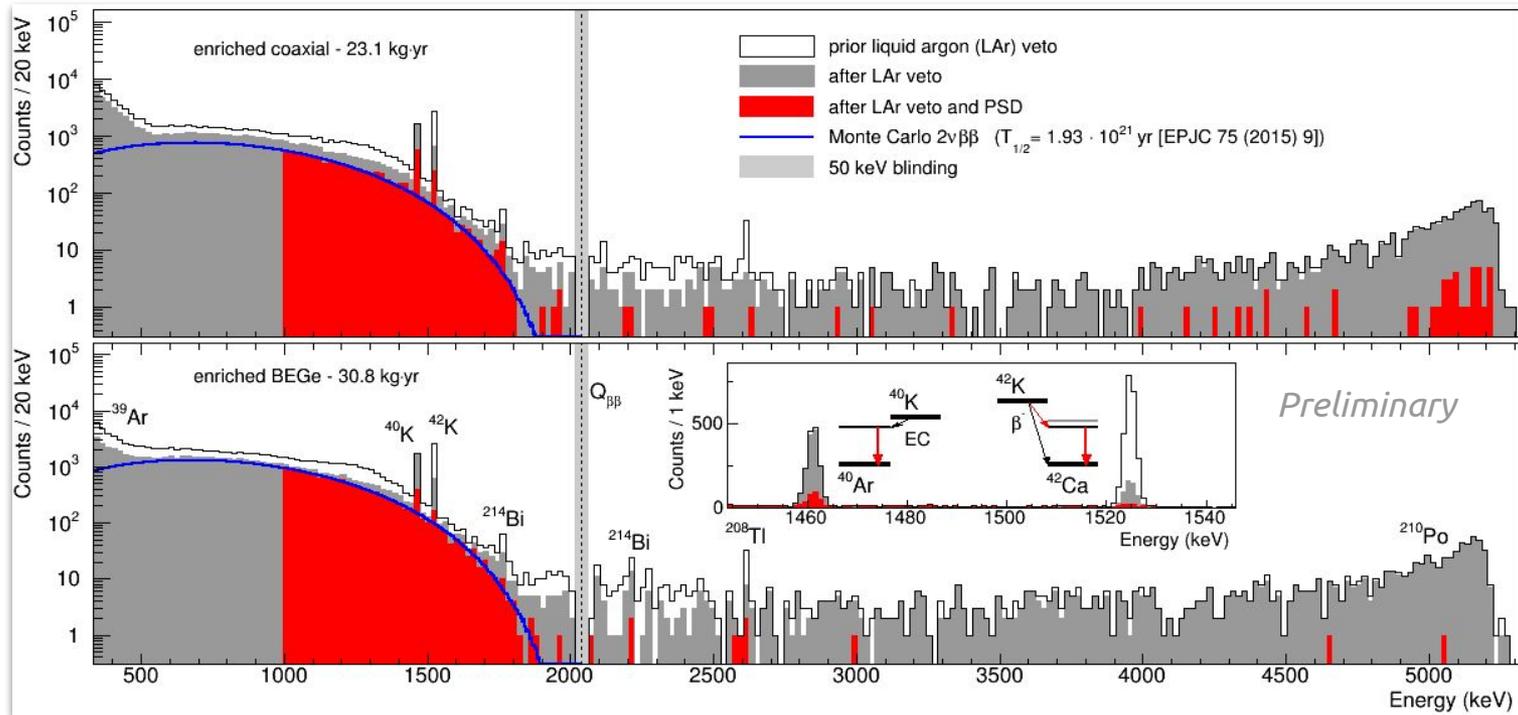


GERDA: Background Suppression with LAr Veto



- Almost pure $2\nu\beta\beta$ spectrum after LAr veto cut (600-1300 keV)
- LAr veto cut signal acceptance 97.7(1)%

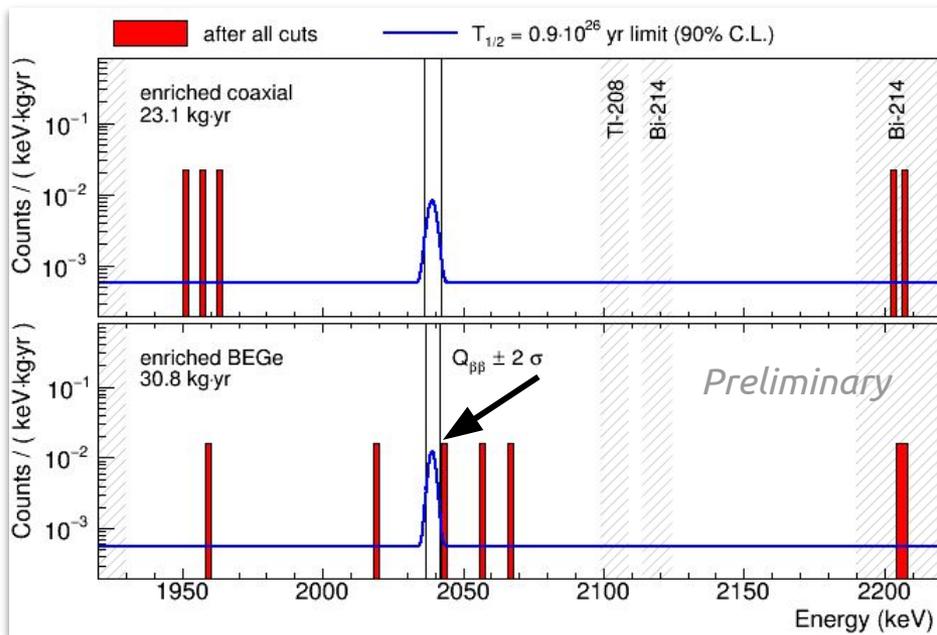
GERDA: Background Suppression with PSD



- Both K lines and high energy α events strongly suppressed
- High $0\nu\beta\beta$ signal efficiency
(71.2 ± 4.3)% for Coax and (87.6 ± 2.5)% for BEGe detectors

GERDA: Background Index

Background index window: 1930-2190 keV, excl. ± 5 keV around two known γ lines and around $Q_{\beta\beta}$



$$\text{Coax}^*: 5.7^{+4.1}_{-2.6} \cdot 10^{-4} \text{ cts}/(\text{keV}\cdot\text{kg}\cdot\text{yr})$$

$$\text{BEGe}: 5.6^{+3.4}_{-2.4} \cdot 10^{-4} \text{ cts}/(\text{keV}\cdot\text{kg}\cdot\text{yr})$$

One new event in the BEGe dataset with energy 2042 keV

The MAJORANA DEMONSTRATOR

Operating underground at the 4850' Sanford Underground Research Facility

- Demonstrating backgrounds low enough to justify building a tonne scale experiment.

Goals: - Establishing feasibility to construct & field modular arrays of Ge detectors.
- Searching for additional physics beyond the standard model.

Energy resolution of 2.5 keV FWHM @ 2039 keV is the best of any $\beta\beta$ -decay experiment

Background Goal in the $0\nu\beta\beta$ peak after analysis cuts:

- With the achieved resolution: 2.5 counts/(FWHM t yr)
- Projected backgrounds based on assay results ≤ 2.2 counts/(FWHM t yr)

44.1-kg of Ge detectors

- 29.7 kg of 88% enriched ^{76}Ge crystals
- 14.4 kg of natGe
- Detector Technology: P-type, point-contact.

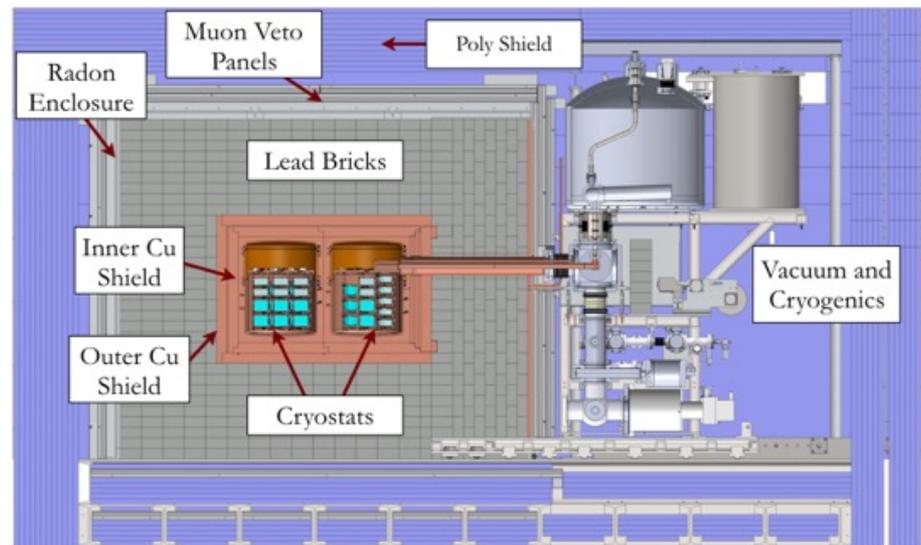
2 independent cryostats

- Ultra-clean, electroformed Cu
- 22 kg of detectors per cryostat
- Naturally scalable

Compact Shield

- Low-background passive Cu and Pb shield with active muon veto

Funded by DOE Office of Nuclear Physics, NSF Particle Astrophysics, NSF Nuclear Physics with additional contributions from international collaborators.



[N. Abgrall et al. Adv. High Energy Phys **2014**, 365432 (2014)]

MAJORANA DEMONSTRATOR Clean Materials

Ultra-pure materials

- Low mass design
- Custom cable connectors and front-end boards
- Selected plastics & fine Cu coax cables
- Underground Electro-formed Cu
 - Th decay chain (ave) $\leq 0.1 \mu\text{Bq/kg}$
 - U decay chain (ave) $\leq 0.1 \mu\text{Bq/kg}$



Machining and Cleaning

- Cu machining in an underground clean room
- Cleaning of Cu parts by acid etching and passivation
- Nitric leaching of plastic parts

Detector assembly

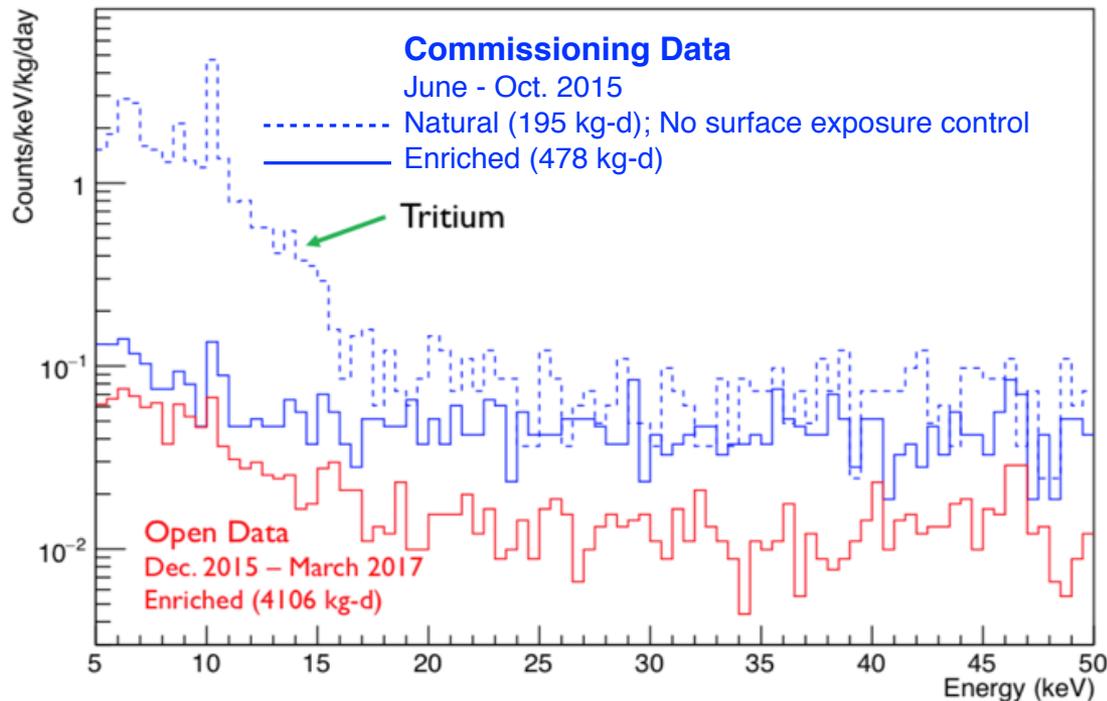
- Dedicated glove boxes with a purged N_2 environment



MAJORANA DEMONSTRATOR Beyond-the-Standard-Model Physics Searches

The low backgrounds at low energy allows additional searches

- Controlled surface exposure of enriched material to minimize cosmogenics
- Sub-keV energy thresholds possible (< 500 eV)
Efficiency below 5 keV is under study.
- Excellent energy resolution (0.4 keV FWHM at 10.4 keV)



Permits low-energy physics

pseudoscalar dark matter

vector dark matter

14.4-keV solar axion

$e^- \Rightarrow 3\nu$

Pauli Exclusion Principle

Phys. Rev. Lett. **118** 161801 (2017)

Other exotic physics

Lightly ionizing particles

Phys. Rev. Lett. **120** 211804 (2018)

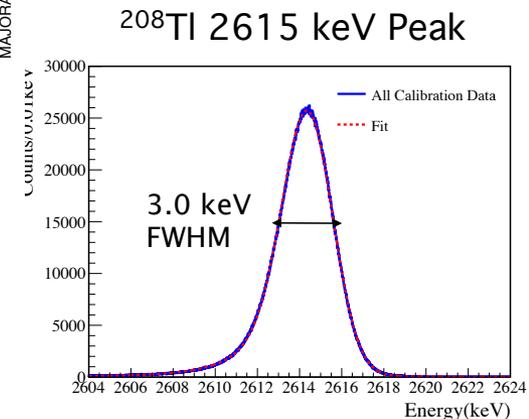
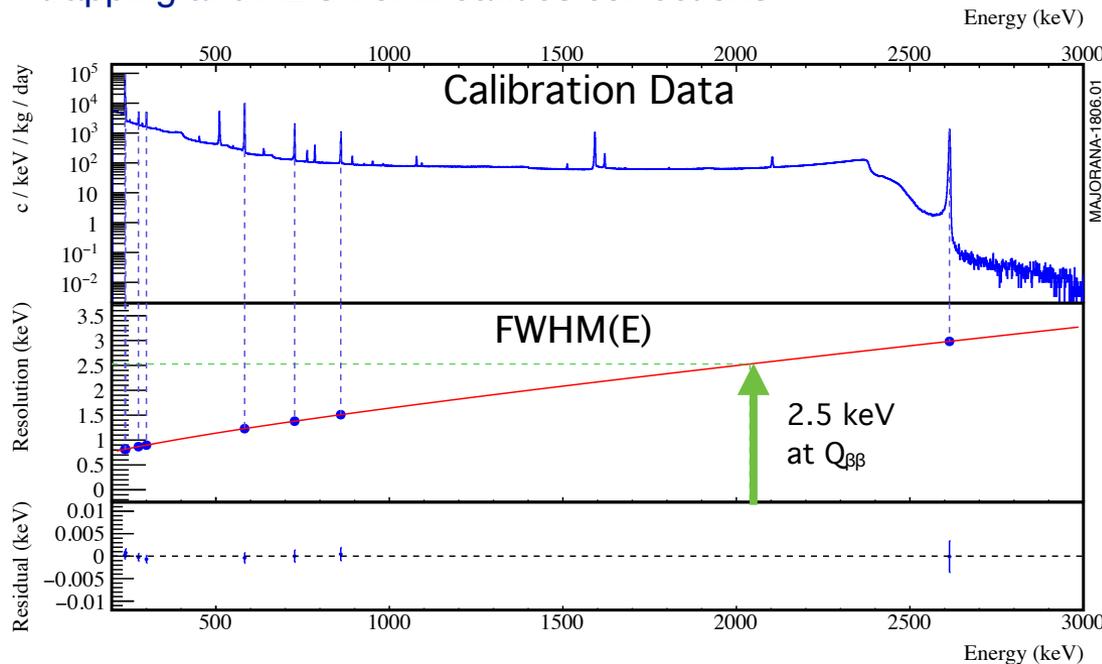
MAJORANA DEMONSTRATOR: Energy Resolution

Calibration of the detector array with a ^{228}Th line source

[NIMA 872,16 (2017) arXiv:1702.02466]

- Source is inserted and retracted for scheduled calibrations
- Provides energy calibration, gain stability checks, and tuning of single-site (DEP) and multi-site (SEP) cuts

Excellent energy resolution attained improved by charge trapping and ADC nonlinearities corrections



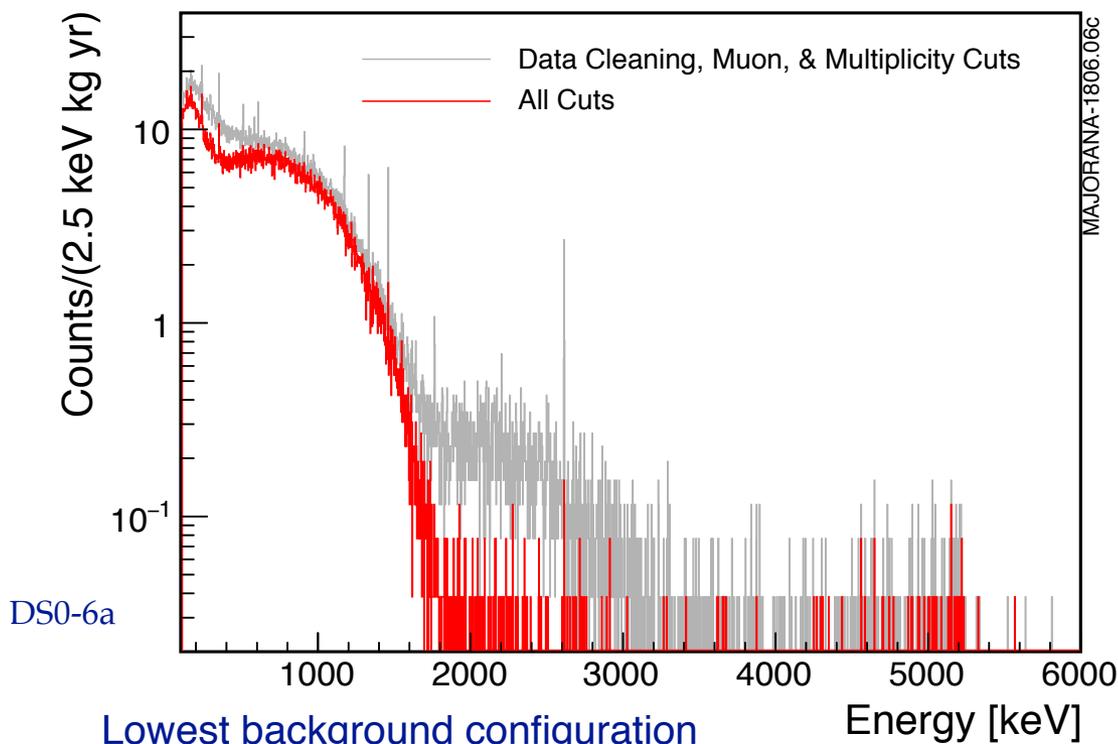
Best achieved for $0\nu\beta\beta$ searches!

MAJORANA DEMONSTRATOR: Recent $0\nu\beta\beta$ Results

All data (all open and previously blinded) up until 15 April 2018

- All data: 26 kg-yr (^{enr}Ge)

PRELIMINARY



Full Exposure Background

15.4 ± 2.0 cts/(FWHM t yr)

The expected counts in the optimal ROI (4.13 keV): 0.66

After final unblinding:
1 event at 2040 keV

Lowest background configuration

- Active Exposure: 21.3 kg yr (^{enr}Ge)

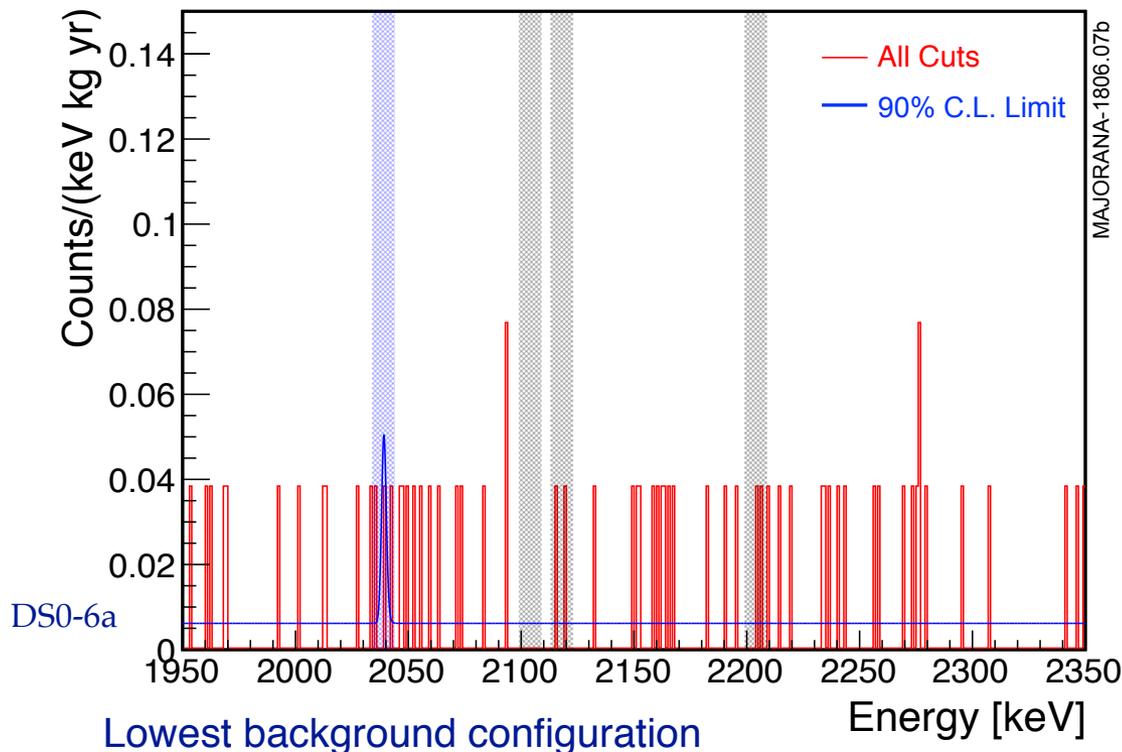
- Background rate: 11.9 ± 2.0 cts/(FWHM t yr); $4.7 \pm 0.8 \times 10^{-3}$ cts/(keV kg yr)

MAJORANA DEMONSTRATOR: Recent $0\nu\beta\beta$ Results

All data (all open and previously blinded) up until 15 April 2018

- All data: 26 kg-yr (^{enr}Ge)

PRELIMINARY



Median Sensitivity:

$$4.8 \times 10^{25} \text{ yr (90\% CL)}$$

Full Exposure Limit

$$T_{1/2}^{0\nu} > 2.7 \times 10^{25} \text{ yr (90\% CL)}$$

Full Exposure Background

$$15.4 \pm 2.0 \text{ cts}/(\text{FWHM t yr})$$

The expected counts in the optimal ROI (4.13 keV): 0.66

After final unblinding:

1 event at 2040 keV

Lowest background configuration

- Active Exposure: 21.3 kg yr (^{enr}Ge)

- Background rate: $11.9 \pm 2.0 \text{ cts}/(\text{FWHM t yr})$; $4.7 \pm 0.8 \times 10^{-3} \text{ cts}/(\text{keV kg yr})$

LEGEND

“The collaboration aims to develop a phased, Ge-76 based double-beta decay experimental program with discovery potential at a half-life significantly longer than 10 years, using existing resources as appropriate to expedite physics results.”

Combining successes of MJD clean materials development, GERDA active shield, and new initiatives.

LEGEND-200:

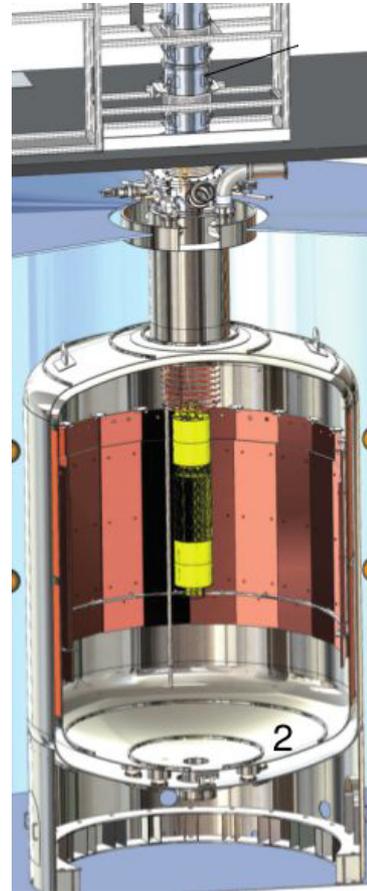
- ~200-kg array deployment in existing GERDA cryostat at LNGS
- BG goal (5x lower) 0.6 c/(FWHM-tn-yr)
- Start by 2021

LEGEND-1k:

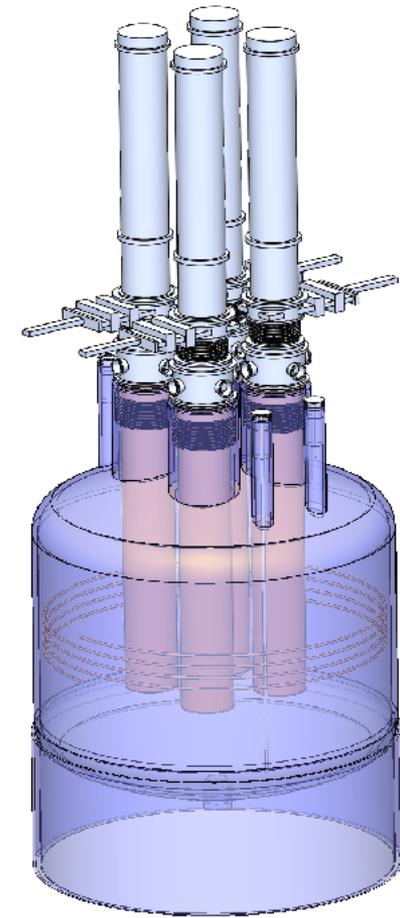
- 1000-kg, ~4-module array with phased deployment.
- BG goal (30x lower) 0.1 c/(FWHM-tn-yr)
- Timeline connected to US DOE down-select
- Location TBD - Required depth under investigation

Active Shield: Investigating reduction of Ar backgrounds:

- Depleted Argon
- Detector encapsulation
- Frozen scintillator
- Liquid Neon



LEGEND-200

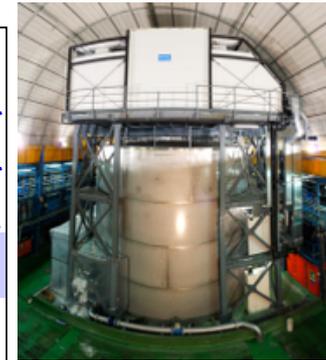
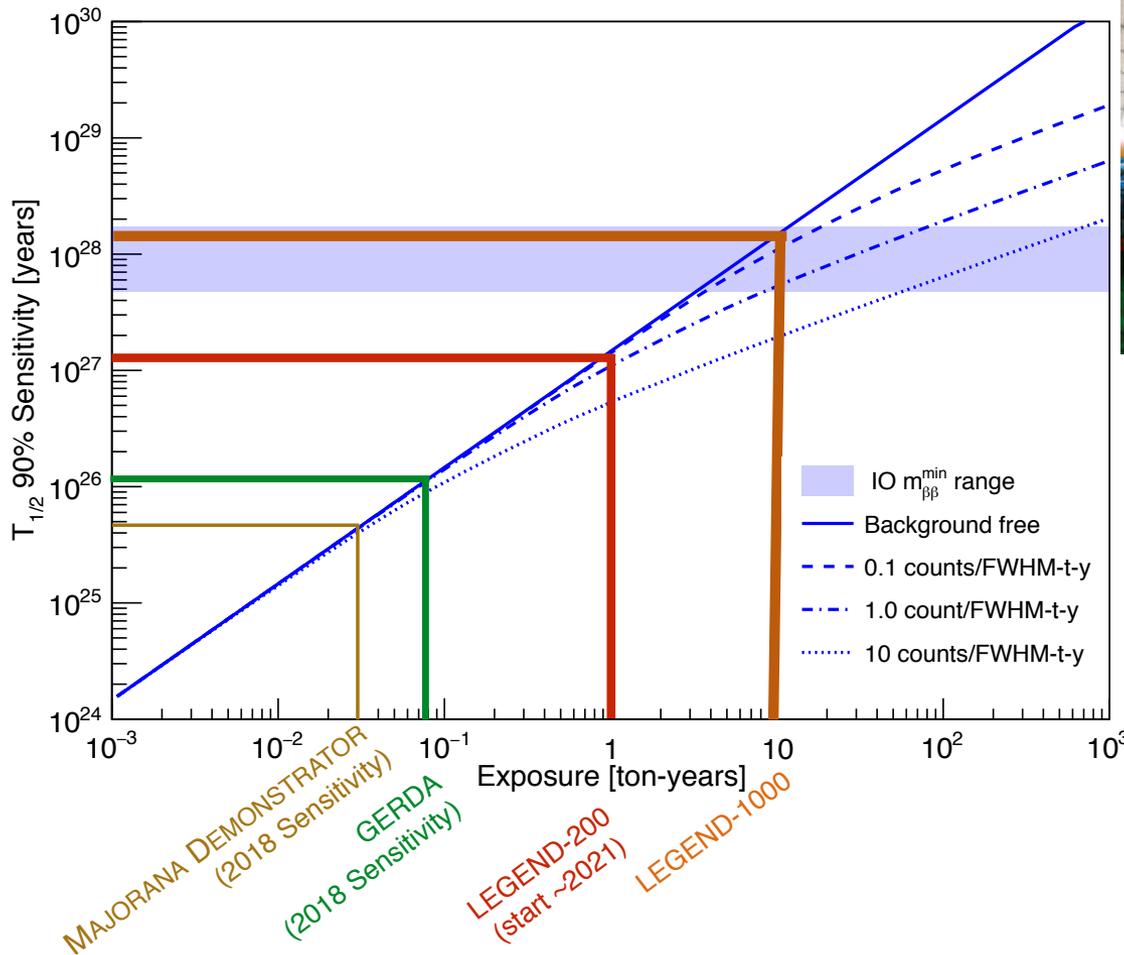


LEGEND-1k

LEGEND: Future Sensitivity

Next Generation ^{76}Ge : LEGEND — Large Enriched Germanium Experiment for Neutrinoless $\beta\beta$ Decay (52 Institutions, ~250 Members)

^{76}Ge (87% enr.)



First Stage:

- (up to) 200 kg ^{76}Ge in upgrade of existing infrastructure at LNGS
- BG goal 0.6 cts/(FWHM t yr)
- Data start ~2021
- Will use existing MAJORANA & GERDA detectors (65 kg), plus new detectors (135 kg)

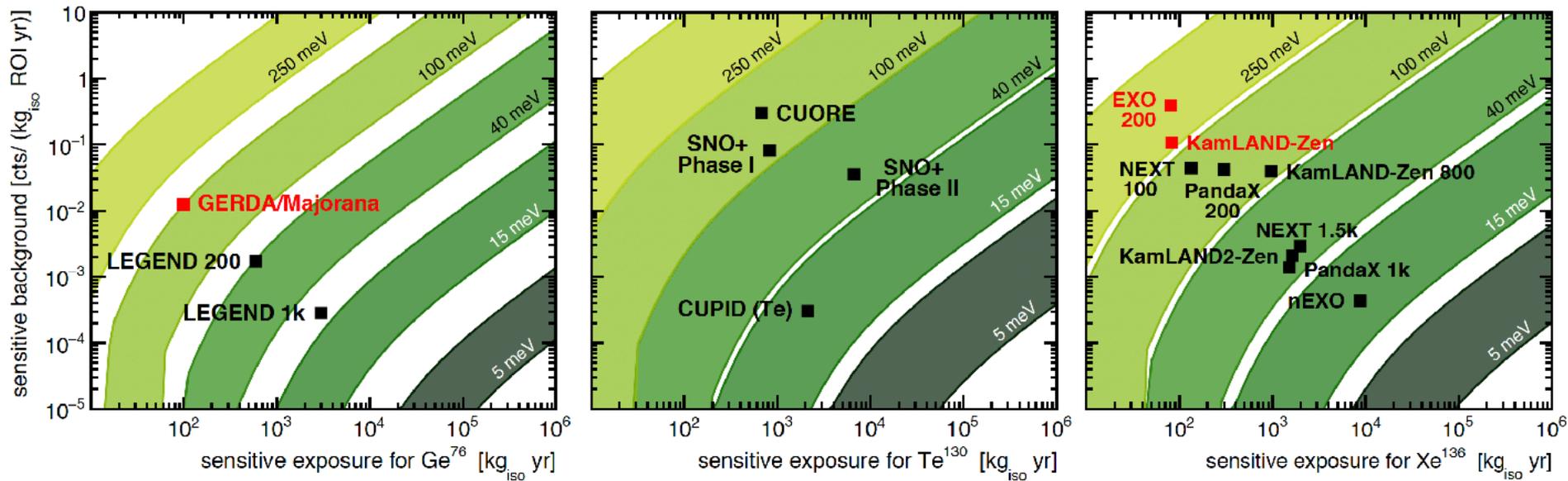
Subsequent Stages:

- 1000 kg ^{76}Ge (staged)
- BG goal: 0.1 cts/(FWHM t yr)
- Location: TBD

The Global Picture

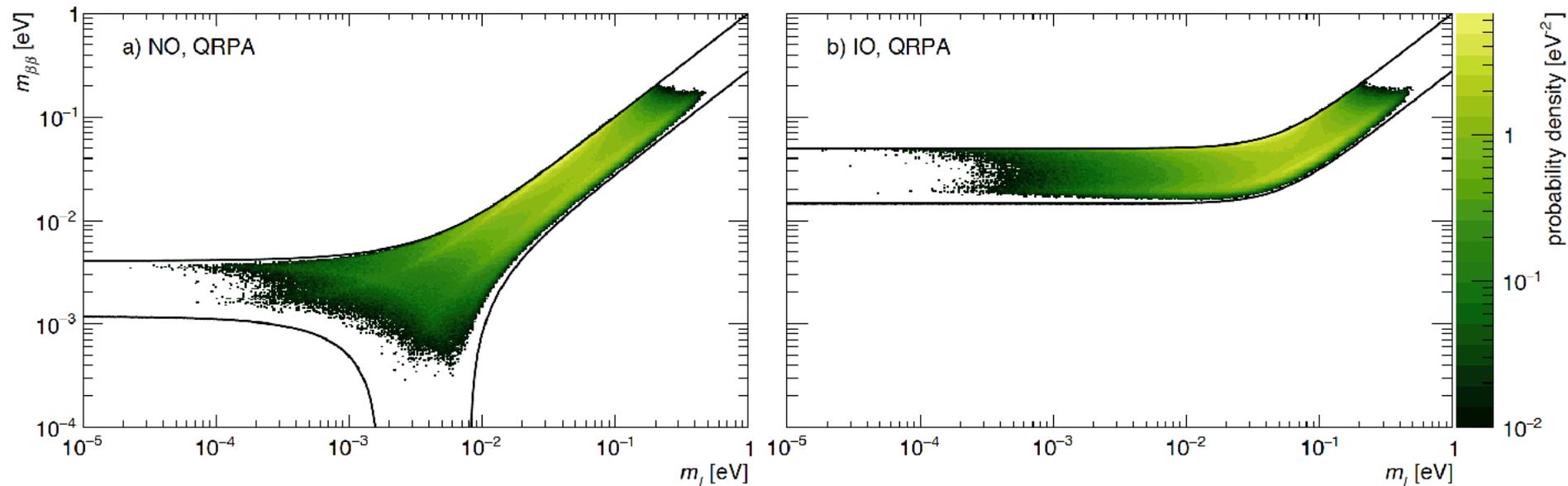
Experiment	Iso.	Iso. Mass [kg _{iso}]	σ [keV]	ROI [σ]	ϵ_{FV} [%]	ϵ_{sig} [%]	\mathcal{E} [$\frac{\text{kg}_{iso} \text{ yr}}{\text{yr}}$]	\mathcal{B} [$\frac{\text{cts}}{\text{kg}_{iso} \text{ ROI yr}}$]	3σ disc. $\hat{T}_{1/2}$ [yr]	sens. $\hat{m}_{\beta\beta}$ [meV]	Required Improvement		
											Bkg	σ	Iso. Mass
LEGEND 200 [62, 63]	⁷⁶ Ge	175	1.3	[-2, 2]	93	77	119	$1.7 \cdot 10^{-3}$	$8.4 \cdot 10^{26}$	40–73	3	1	5.7
LEGEND 1k [62, 63]	⁷⁶ Ge	873	1.3	[-2, 2]	93	77	593	$2.8 \cdot 10^{-4}$	$4.5 \cdot 10^{27}$	17–31	18	1	29
SuperNEMO [69, 70]	⁸² Se	100	51	[-4, 2]	100	16	16.5	$4.9 \cdot 10^{-2}$	$6.1 \cdot 10^{25}$	82–138	49	2	14
CUPID [59, 60, 71]	⁸² Se	336	2.1	[-2, 2]	100	69	221	$5.2 \cdot 10^{-4}$	$1.8 \cdot 10^{27}$	15–25	n/a	6	n/a
CUORE [53, 54]	¹³⁰ Te	206	2.1	[-1.4, 1.4]	100	81	141	$3.1 \cdot 10^{-1}$	$5.4 \cdot 10^{25}$	66–164	6	1	19
CUPID [59, 60, 71]	¹³⁰ Te	543	2.1	[-2, 2]	100	81	422	$3.0 \cdot 10^{-4}$	$2.1 \cdot 10^{27}$	11–26	3000	1	50
SNO+ Phase I [67, 72]	¹³⁰ Te	1357	82	[-0.5, 1.5]	20	97	164	$8.2 \cdot 10^{-2}$	$1.1 \cdot 10^{26}$	46–115	n/a	n/a	n/a
SNO+ Phase II [68]	¹³⁰ Te	7960	57	[-0.5, 1.5]	28	97	1326	$3.6 \cdot 10^{-2}$	$4.8 \cdot 10^{26}$	22–54	n/a	n/a	n/a
KamLAND-Zen 800 [61]	¹³⁶ Xe	750	114	[0, 1.4]	64	97	194	$3.9 \cdot 10^{-2}$	$1.6 \cdot 10^{26}$	47–108	1.5	1	2.1
KamLAND2-Zen [61]	¹³⁶ Xe	1000	60	[0, 1.4]	80	97	325	$2.1 \cdot 10^{-3}$	$8.0 \cdot 10^{26}$	21–49	15	2	2.9
nEXO [73]	¹³⁶ Xe	4507	25	[-1.2, 1.2]	60	85	1741	$4.4 \cdot 10^{-4}$	$4.1 \cdot 10^{27}$	9–22	400	1.2	30
NEXT 100 [65, 74]	¹³⁶ Xe	91	7.8	[-1.3, 2.4]	88	37	26.5	$4.4 \cdot 10^{-2}$	$5.3 \cdot 10^{25}$	82–189	n/a	1	20
NEXT 1.5k [75]	¹³⁶ Xe	1367	5.2	[-1.3, 2.4]	88	37	398	$2.9 \cdot 10^{-3}$	$7.9 \cdot 10^{26}$	21 49	n/a	1	300
PandaX-III 200 [66]	¹³⁶ Xe	180	31	[-2, 2]	100	35	60.2	$4.2 \cdot 10^{-2}$	$8.3 \cdot 10^{25}$	65–150	n/a	n/a	n/a
PandaX-III 1k [66]	¹³⁶ Xe	901	10	[-2, 2]	100	35	301	$1.4 \cdot 10^{-3}$	$9.0 \cdot 10^{26}$	20–46	n/a	n/a	n/a

The Global Picture



- Colored bands represent NME uncertainties.
- Assumed $g_a = 1.27$ (no quenching).

Discovery Potential



- Increasingly significant evidence is suggesting a preference for Normal Ordering.
- While it is essentially impossible to completely rule out NO Majorana Neutrinos with $0\nu\beta\beta$ null results, there is significant discovery potential for next-generation experiments even in the case of Normal Ordering.

Thank you for your attention...

**and thank you for your help,
Alfredo Galindo-Uribarri and Andrea Pocar!**