

STERILE NEUTRINO SEARCHES

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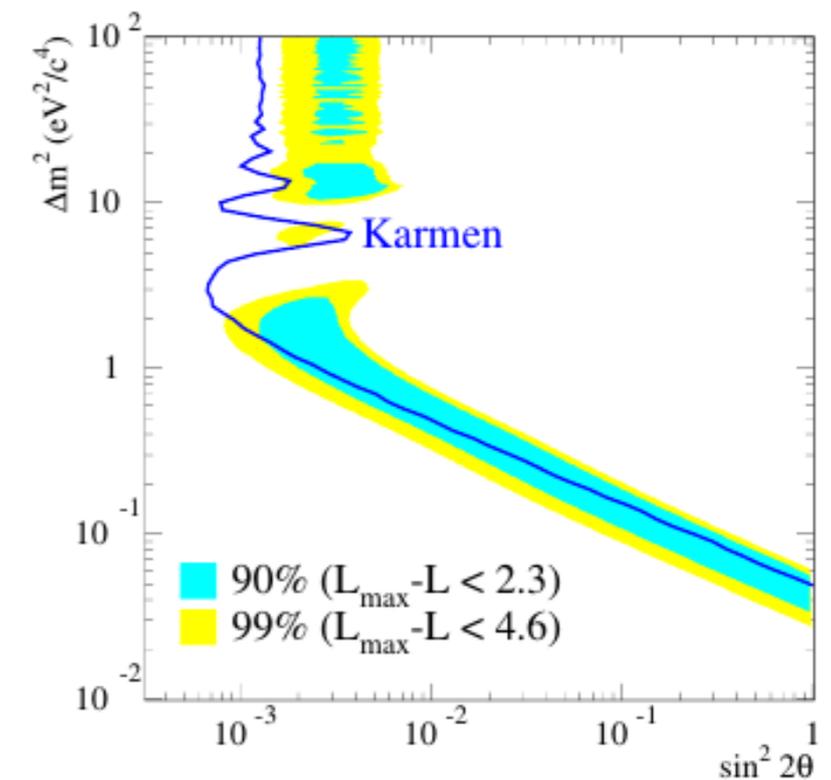
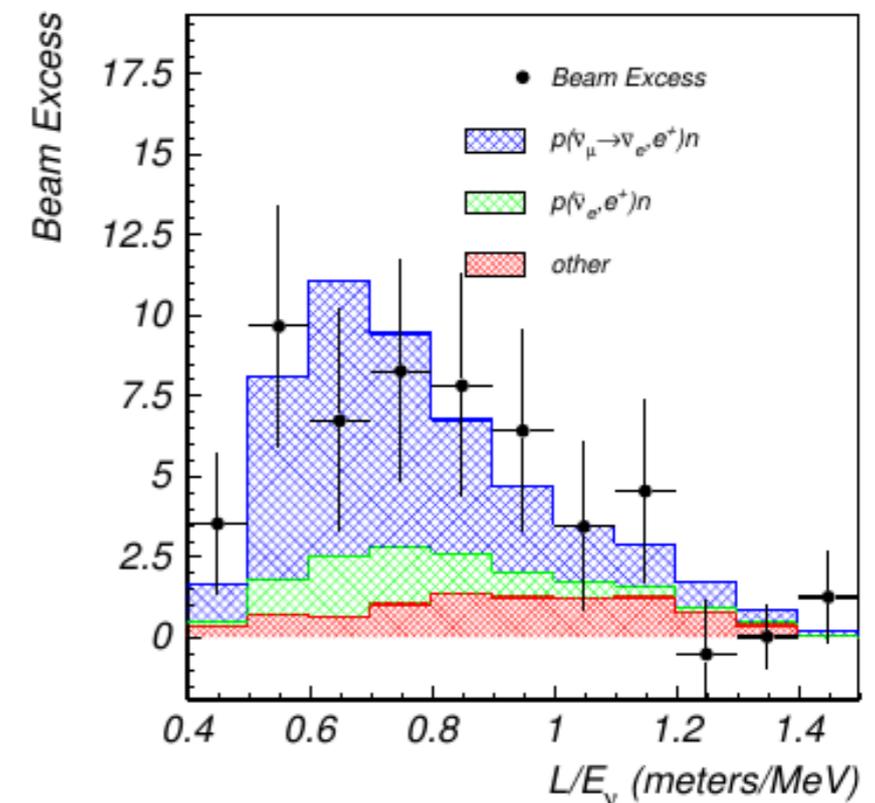
Center for
Underground Physics 

PIC2018: XXXVIII International Symposium on
Physics in Collision, Bogota, Columbia



- Anomalies
- Reactor antineutrino experiments
- MiniBooNE latest results & future experiments
- IceCube, Nova and MINOS
- Summary

- ▶ The LSND experiment observed an excess of events in a beam.
- ▶ The Karmen collaboration did not confirm the claim, but couldn't fully exclude it either.
- ▶ The signal is compatible with neutrino oscillation $\Delta m^2 \geq 0.1 \text{eV}^2$.
- ▶ To explain LSND result, one needs new neutrino mass eigenstates.

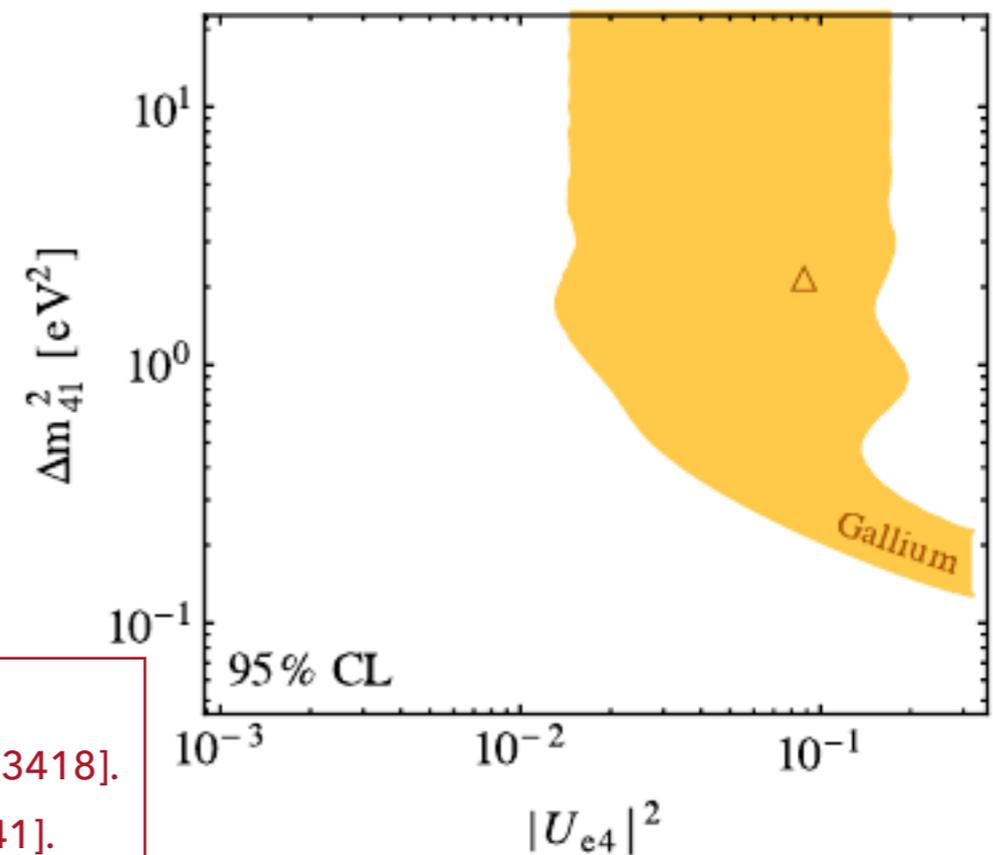
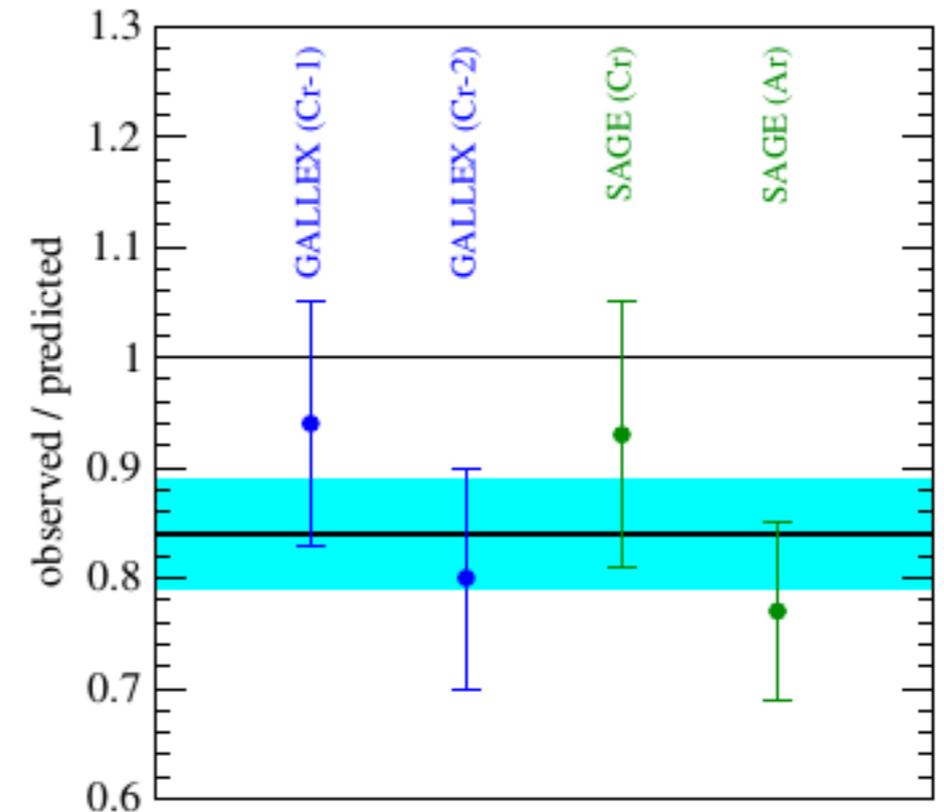


A. Aguilar-Arevalo *et al.* [LSND collab], Phys. Rev. D 64 (2001) 112007 [hep-ex/0104049].
 B. Armbruster *et al.* [KARMEN collab], Phys. Rev. D 65 (2002) 112001 [hep-ex/0203021].

- ▶ The **GALLEX** and **SAGE** solar neutrino experiments were calibrated with intense neutrino sources ^{51}Cr and ^{37}Ar .
- ▶ These measurements show a **significant deficit** with respect to the predicted values:

$$R = 0.84 \pm 0.05$$

- ▶ 3σ deficit can be interpreted by neutrino oscillation, $\Delta m^2 \geq 1\text{eV}^2$

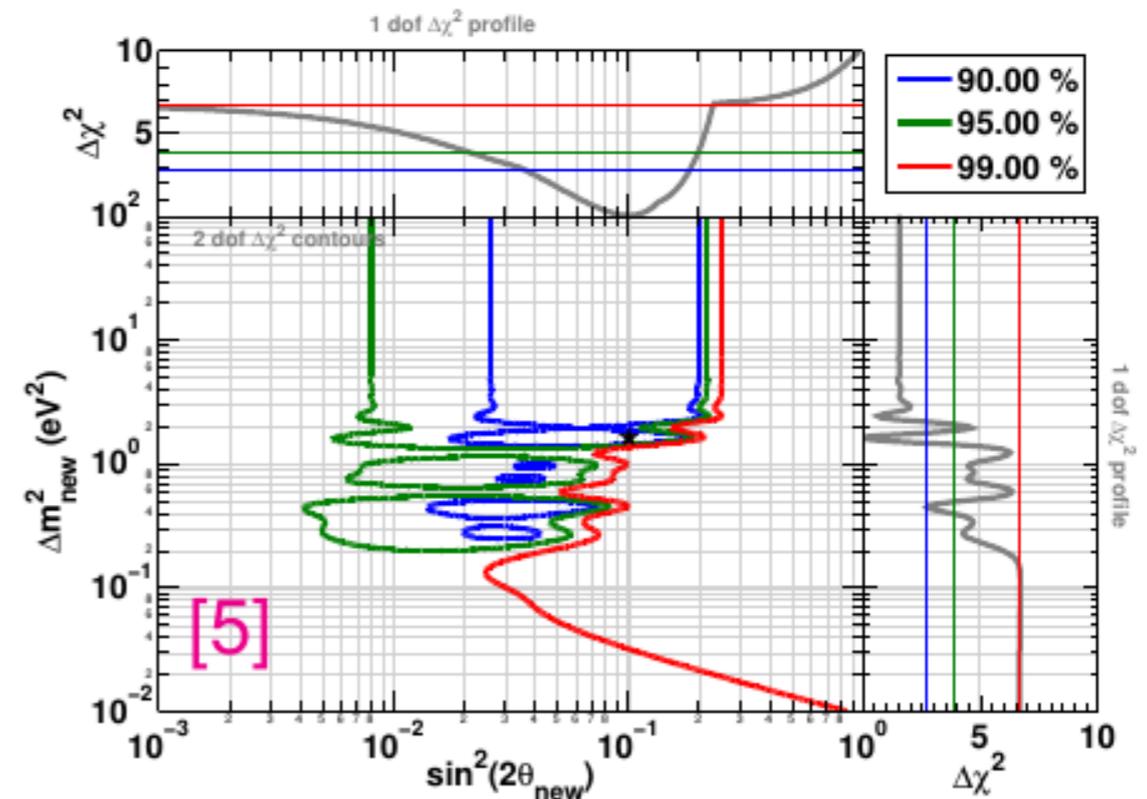
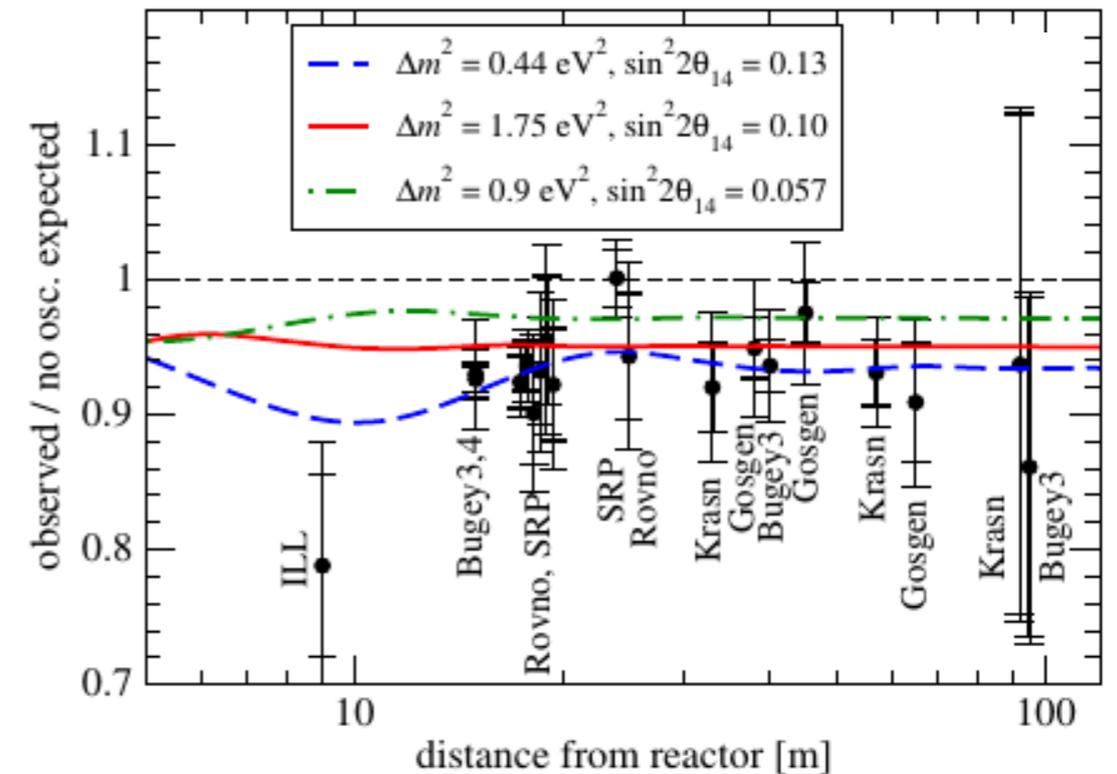


F. Kaether et al., Phys. Lett. B685 (2010) 47-54 [arXiv:1001.2731].
 J. Abdurashitov et al. [SAGE collab], Phys. Rev. C59 (1999) 2246-2263 [hep-ph/9803418].
 J. Abdurashitov et al. [SAGE collab], Phys. Rev. C73 (2006) 045805 [nucl-ex/0512041].

- ▶ The reactor neutrino flux was reevaluated, this calculation result in a small increase of the flux by about 3.5%.
- ▶ All reactor short-baseline experiment are actually observing a deficit.

$$R = 0.927 \pm 0.023$$

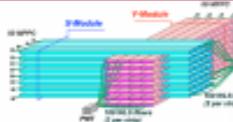
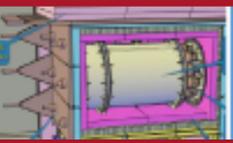
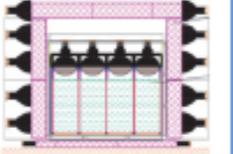
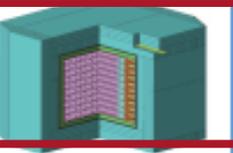
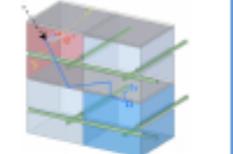
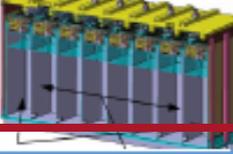
- ▶ These anomaly could be interpreted as being due to SBL neutrino oscillation $\Delta m^2 \geq 1\text{eV}^2$ or incorrect flux.



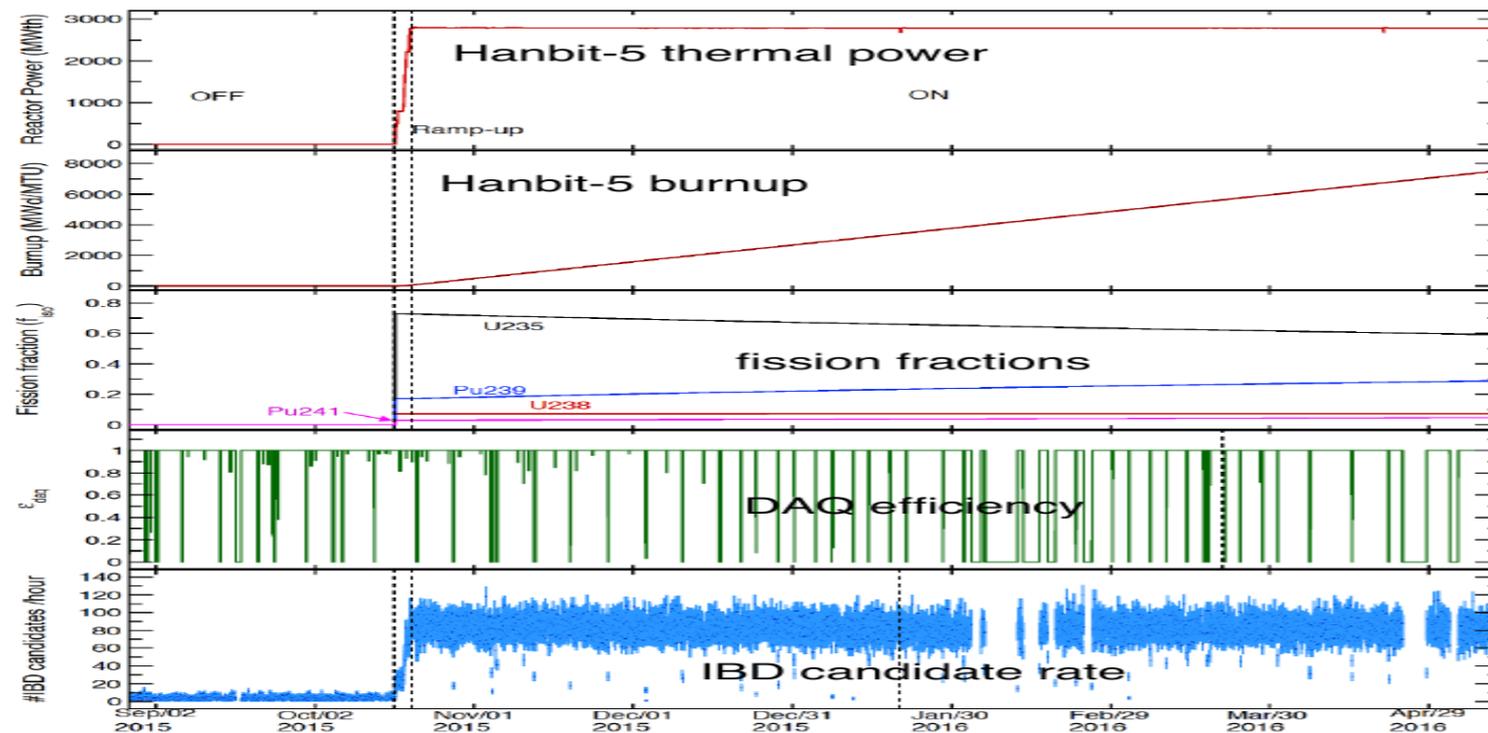
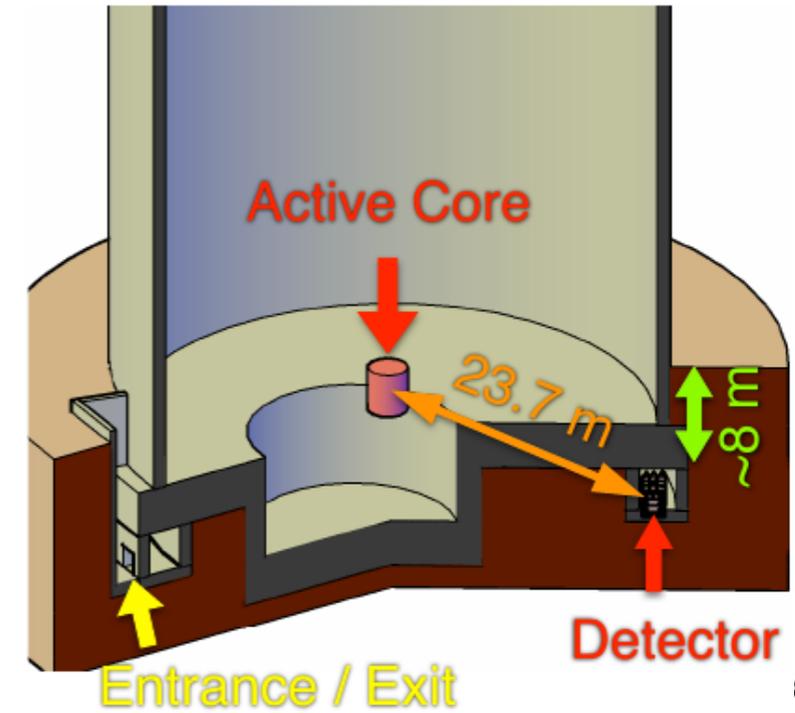
T.A. Mueller et al., Phys. Rev. C83 (2011) 054615 [arXiv:1101.2663].

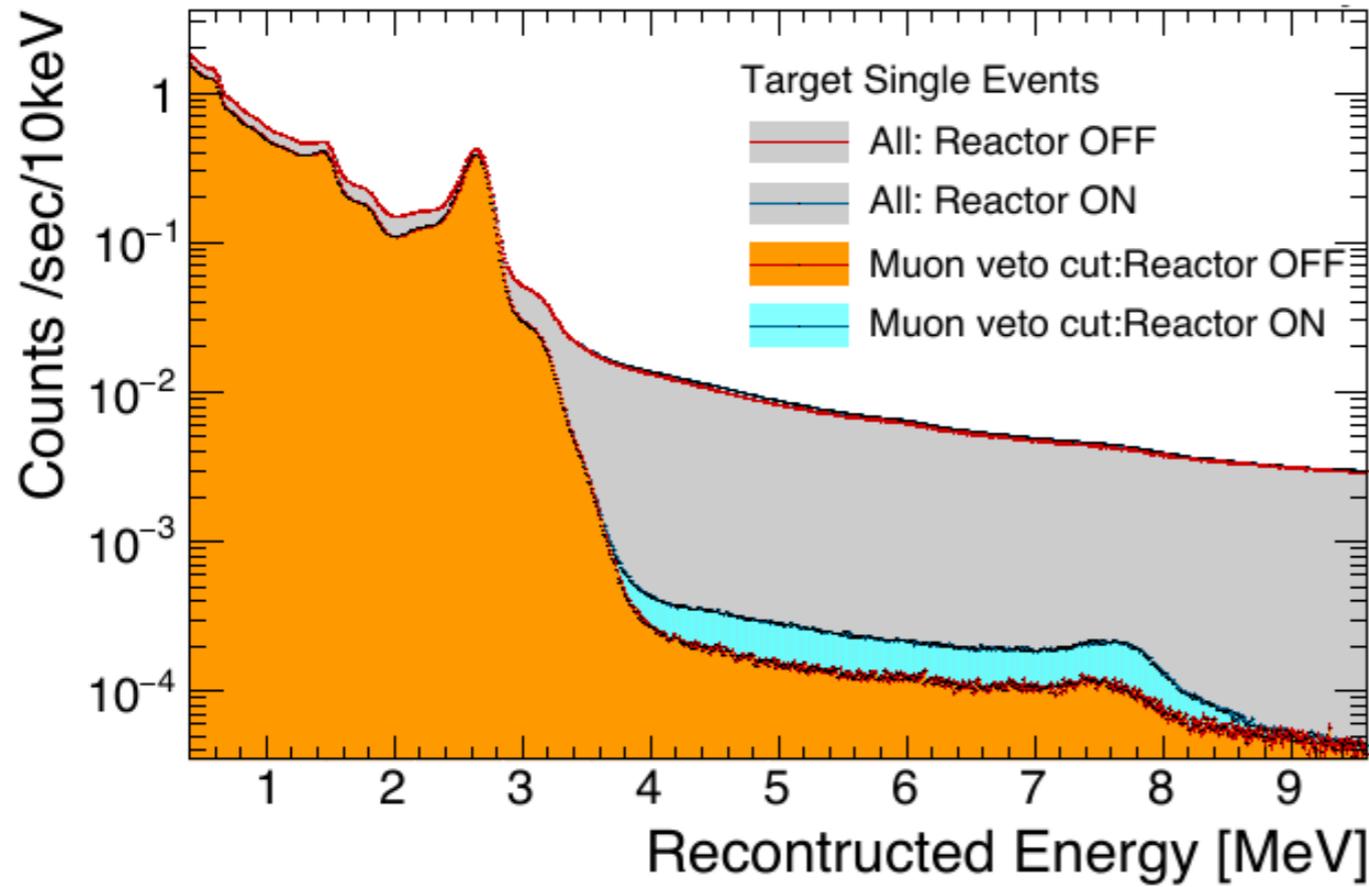
P. Huber, Phys. Rev. C 84 (2011) 024617 [arXiv:1106.0687].

G. Mention et al., Phys. Rev. D83 (2011) 073006 [arXiv:1101.2755].

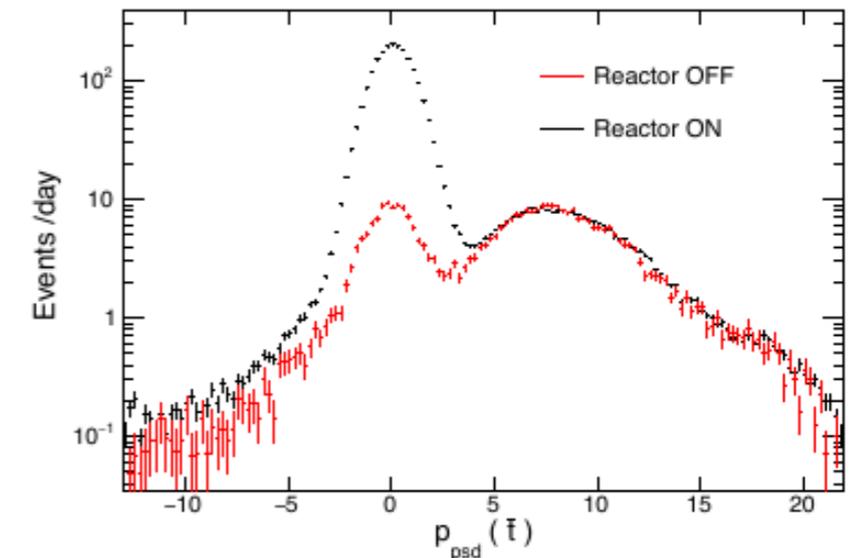
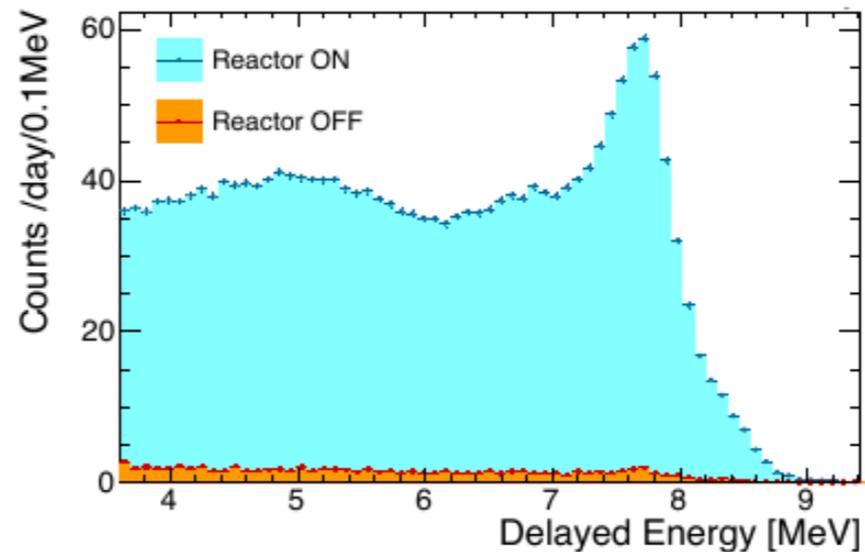
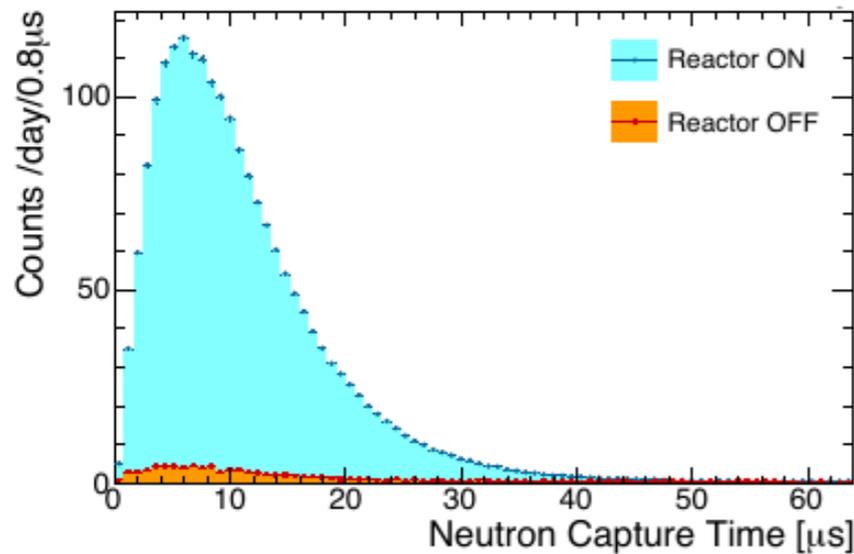
Experiment	Reactor Power/Fuel	Overburden (mwe)	Detection Material	Segmentation	Optical Readout	Particle ID Capability
DANSS (Russia) 	3000 MW LEU fuel	~50	Inhomogeneous PS & Gd sheets	2D, ~5mm	WLS fibers.	Topology only
NEOS (South Korea) 	2800 MW LEU fuel	~20	Homogeneous Gd-doped LS	none	Direct double ended PMT	recoil PSD only
nuLat (USA) 	40 MW ^{235}U fuel	few	Homogeneous ^6Li doped PS	Quasi-3D, 5cm, 3-axis Opt. Latt	Direct PMT	Topology, recoil & capture PSD
Neutrino4 (Russia) 	100 MW ^{235}U fuel	~10	Homogeneous Gd-doped LS	2D, ~10cm	Direct single ended PMT	Topology only
PROSPECT (USA) 	85 MW ^{235}U fuel	few	Homogeneous ^6Li -doped LS	2D, 15cm	Direct double ended PMT	Topology, recoil & capture PSD
SoLid (UK Fr Bel US) 	72 MW ^{235}U fuel	~10	Inhomogeneous $^6\text{LiZnS}$ & PS	Quasi-3D, 5cm multiplex	WLS fibers	topology, capture PSD
Chandler (USA) 	72 MW ^{235}U fuel	~10	Inhomogeneous $^6\text{LiZnS}$ & PS	Quasi-3D, 5cm, 2-axis Opt. Latt	Direct PMT/ WLS Scint.	topology, capture PSD
Stereo (France) 	57 MW ^{235}U fuel	~15	Homogeneous Gd-doped LS	1D, 25cm	Direct single ended PMT	recoil PSD

- ▶ Using commercial reactor (2.8 GWth)
- ▶ Baseline = 23.7 m
- ▶ Overburden ~ 20 m.e.w.
- ▶ Homogenous GdLS (1008L) neutrino target
 - ▶ energy resolution: 4.8% @ 1 MeV
- ▶ Physics data taking: Aug. 2015 ~ May 2016
 - ▶ 46 live days of reactor off
 - ▶ 180 live days of reactor on
- ▶ 1977 (85) IBD candidates/day during on (off), S/B ~ 22.

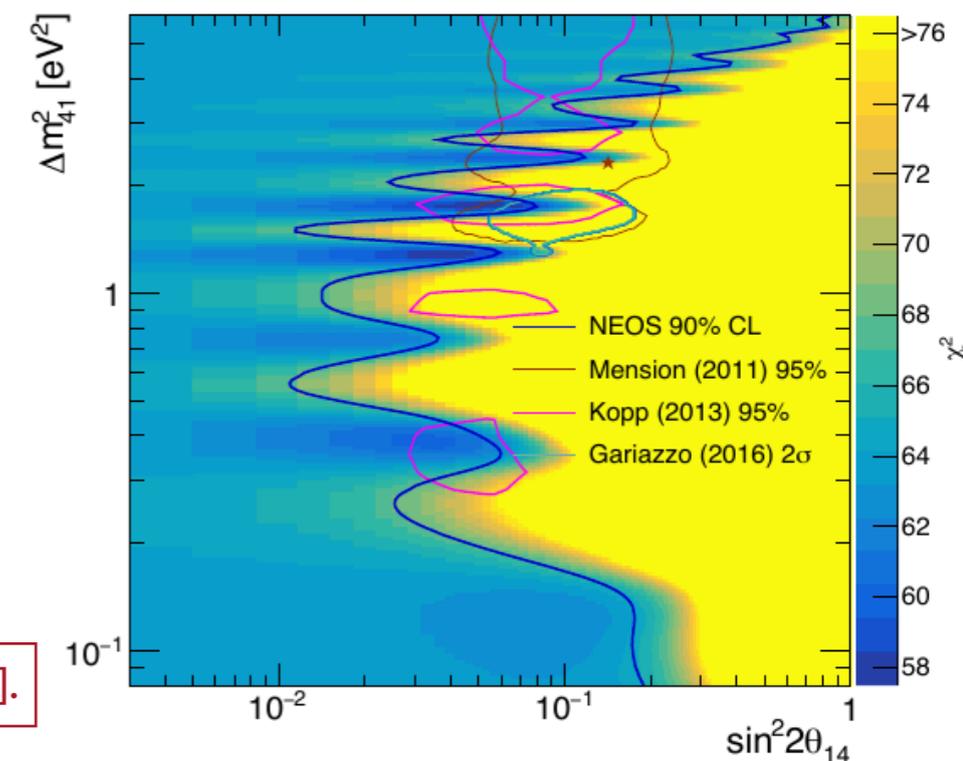
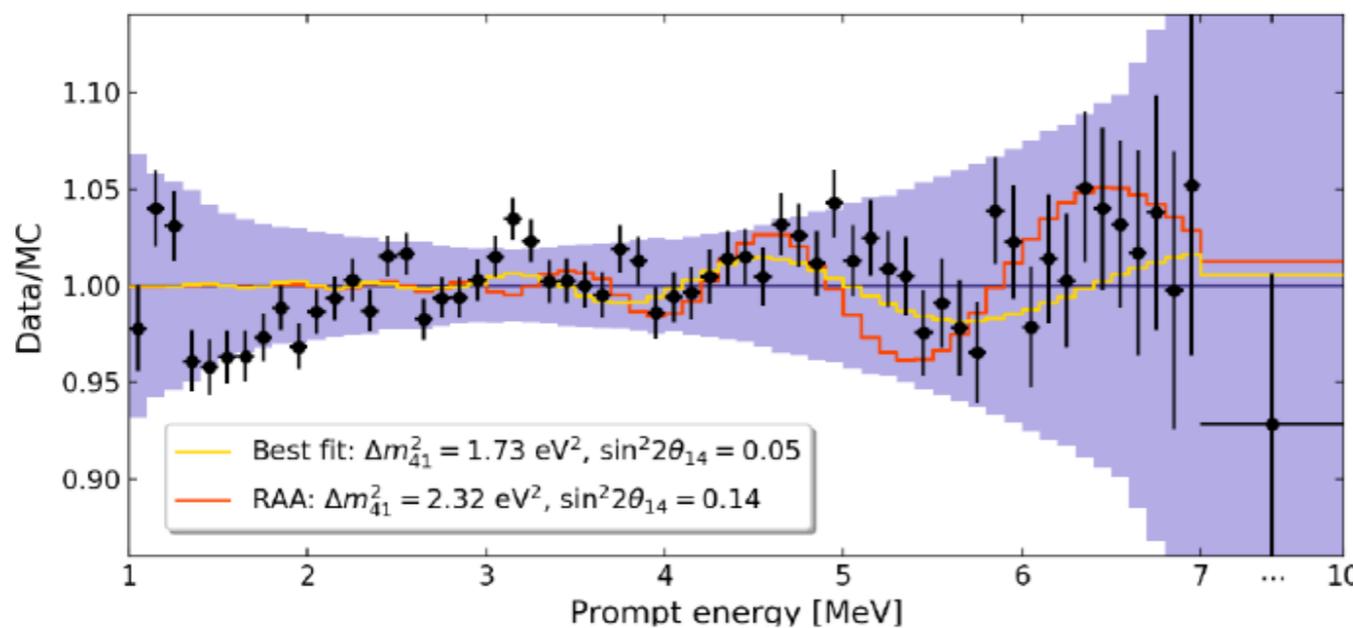
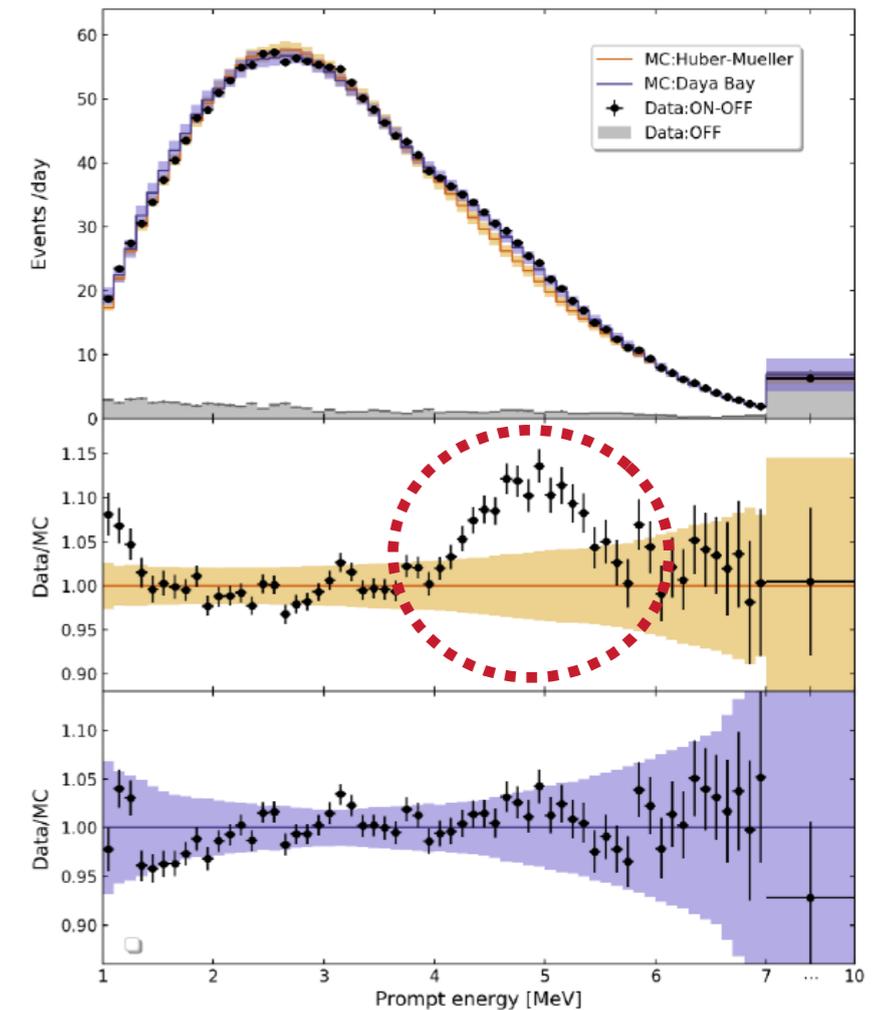




- neutron capture time: 1~30 μs
- delayed energy: 4~10 MeV
- muon veto time: 150 μs
- multiplicity: no other event within 30 μs before / 150 μs after.
- pulse shape discrimination reduces 70% of background.

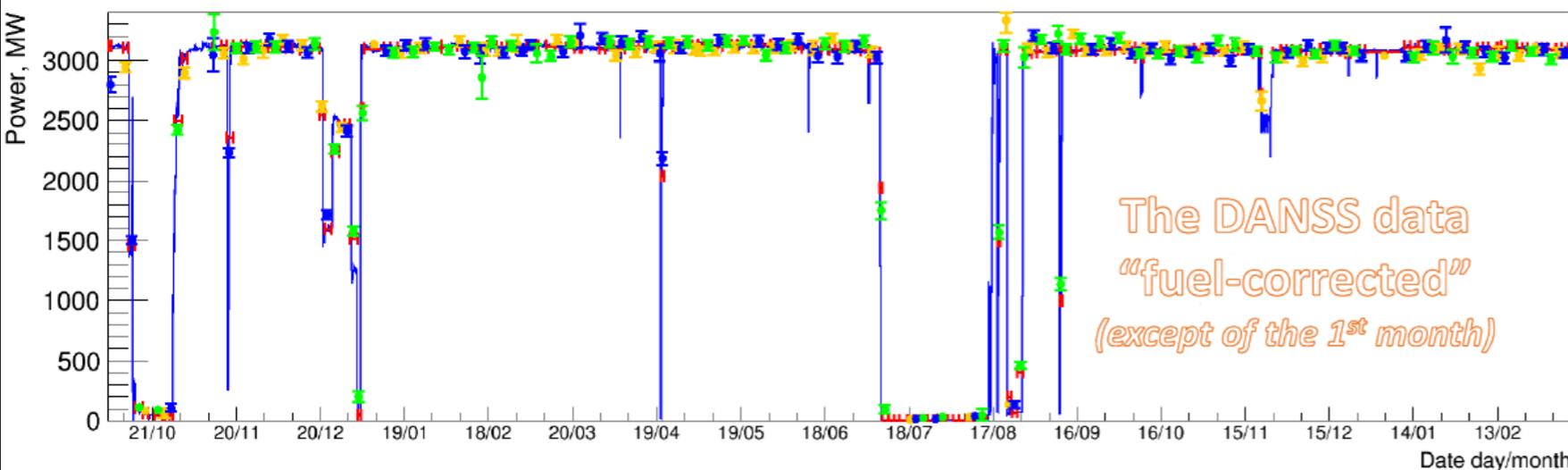
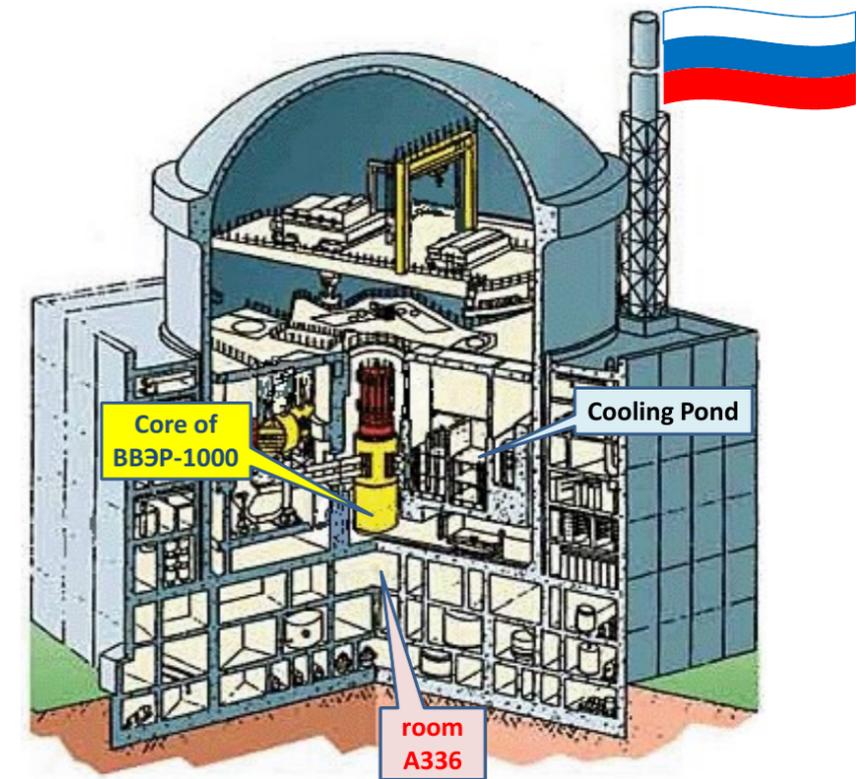


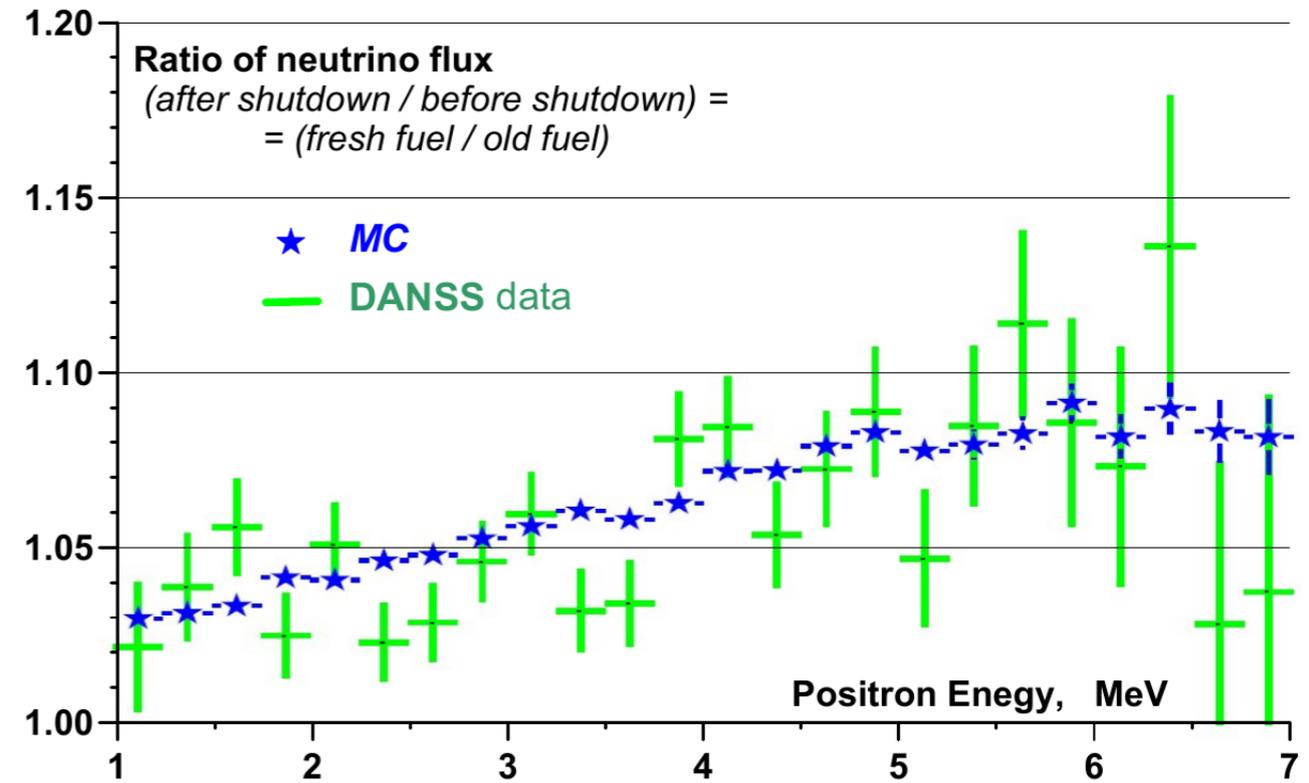
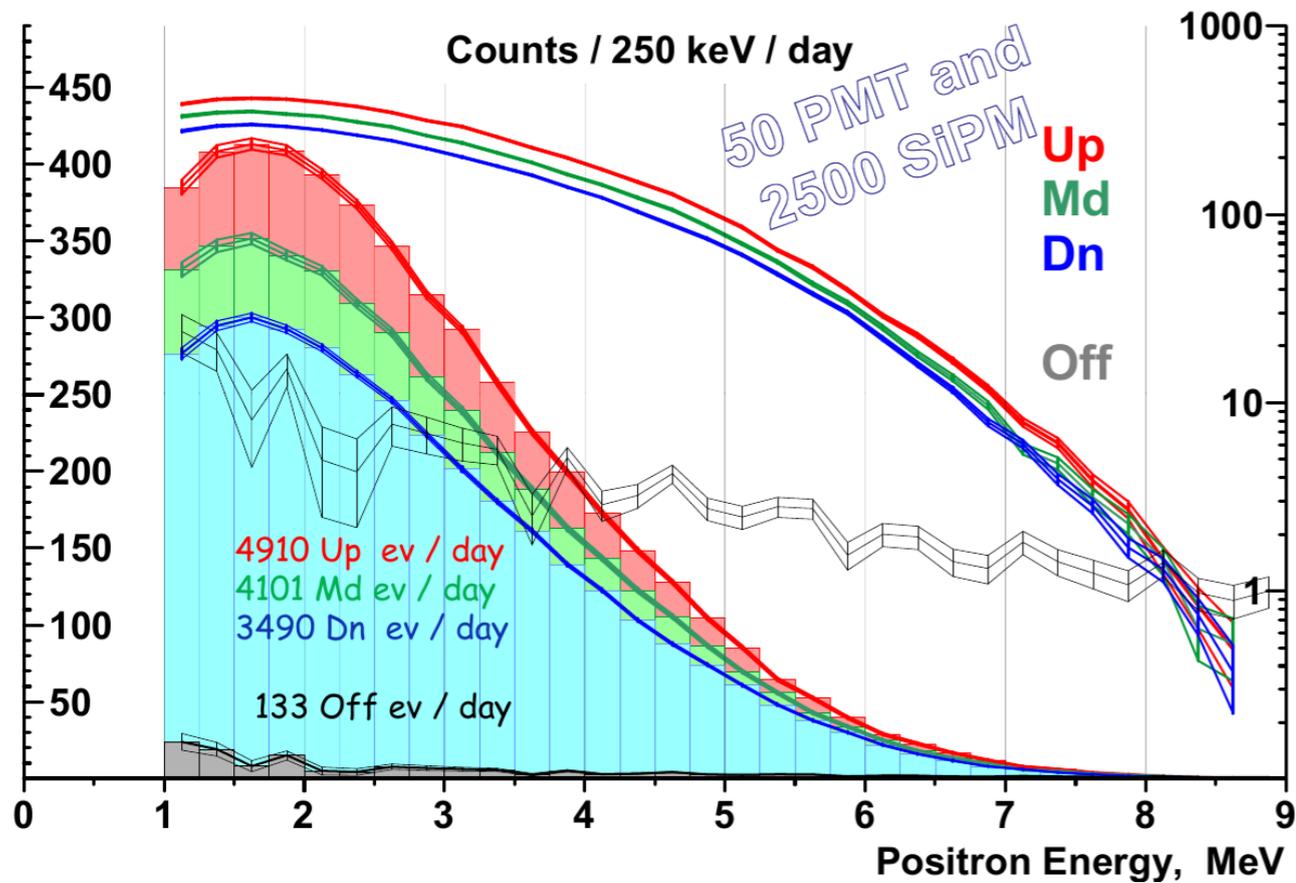
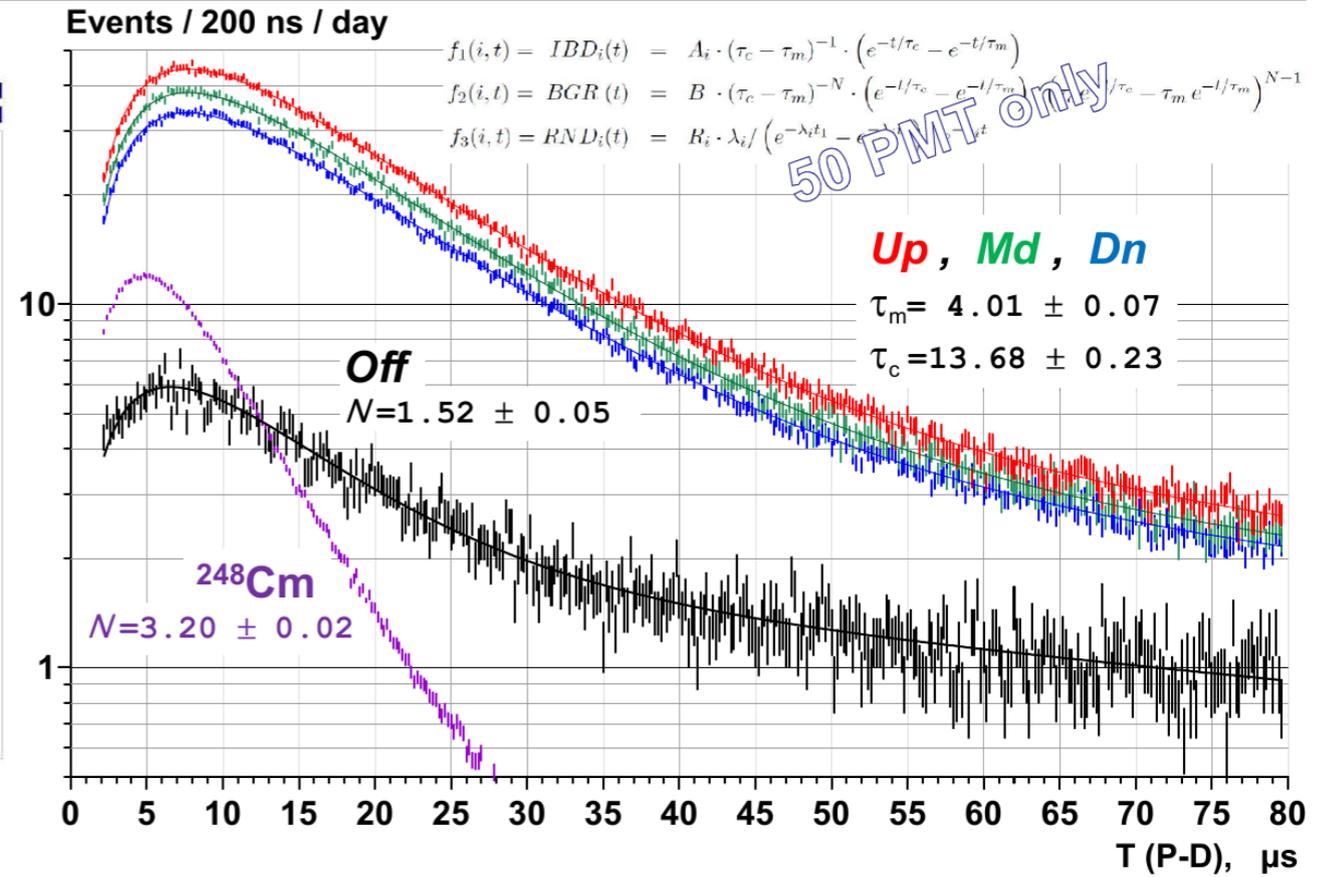
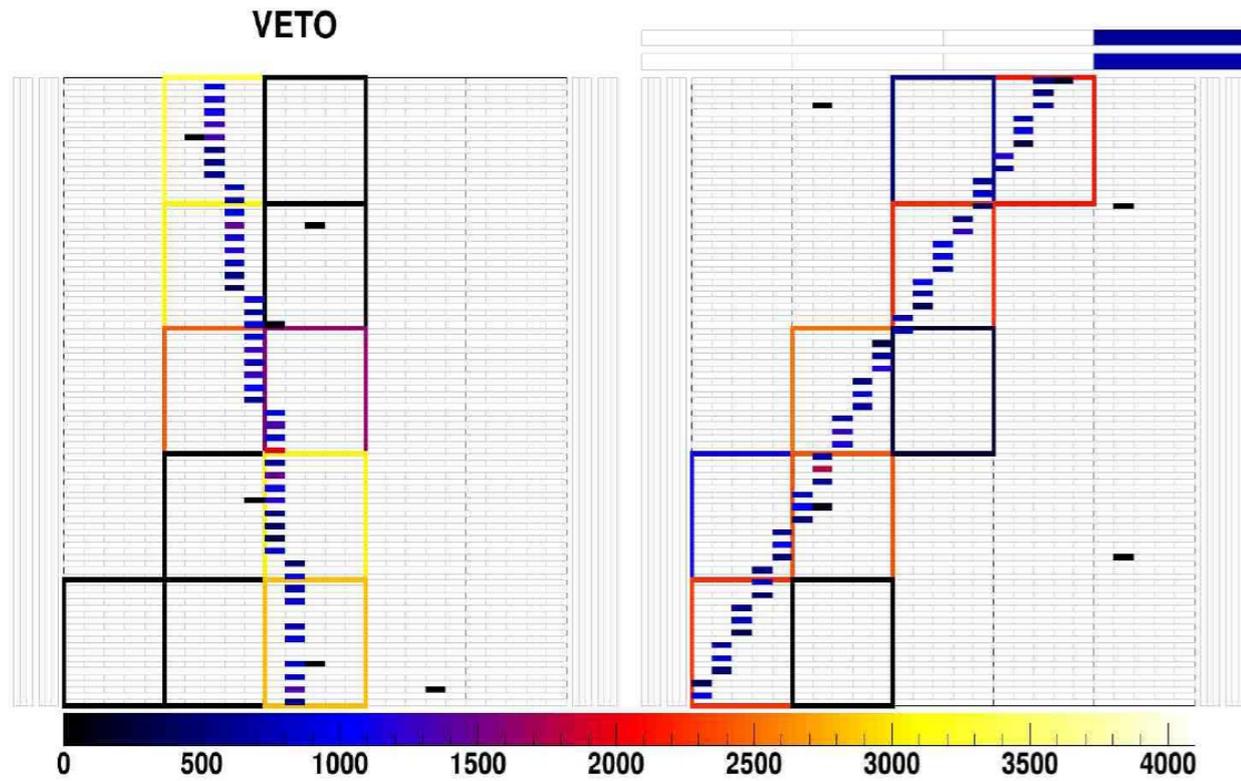
- ▶ Spectral anomaly observed (vs Huber-Mueller normalization)
- ▶ Fit with the Daya Bay shape
 - ▶ Best fit: $\Delta m_{41}^2 = 1.73 \text{eV}^2$, $\sin^2 2\theta_{14} = 0.05$
 $\Delta m_{41}^2 = 1.3 \text{eV}^2$, $\sin^2 2\theta_{14} = 0.04$
- ▶ No strong sign of active-to-sterile neutrino oscillation
- ▶ Fine structures in neutrino spectrum or oscillation?



Y. J. Ko et al. [NEOS collab], Phys. Rev. Lett. 118 (2017) 121802 [arXiv:1610.05134].

- ▶ Pressurized water reactor (3.1 GW_{th})
- ▶ Baseline: 10.7(up), 11.7(md), and 12.7 m(dn)
- ▶ Overburden: ~ 50 m.w.e.
- ▶ Polystyrene-base plastic scintillator strip with Gd-loaded reflector (1.6 mg/cm²)
 - ▶ highly segmented with 2500 individual strips
 - ▶ 2500 SiPMs and 50 PMTs
- ▶ Data taking since Apr. 2016
- ▶ 4910(133) IBD events/day

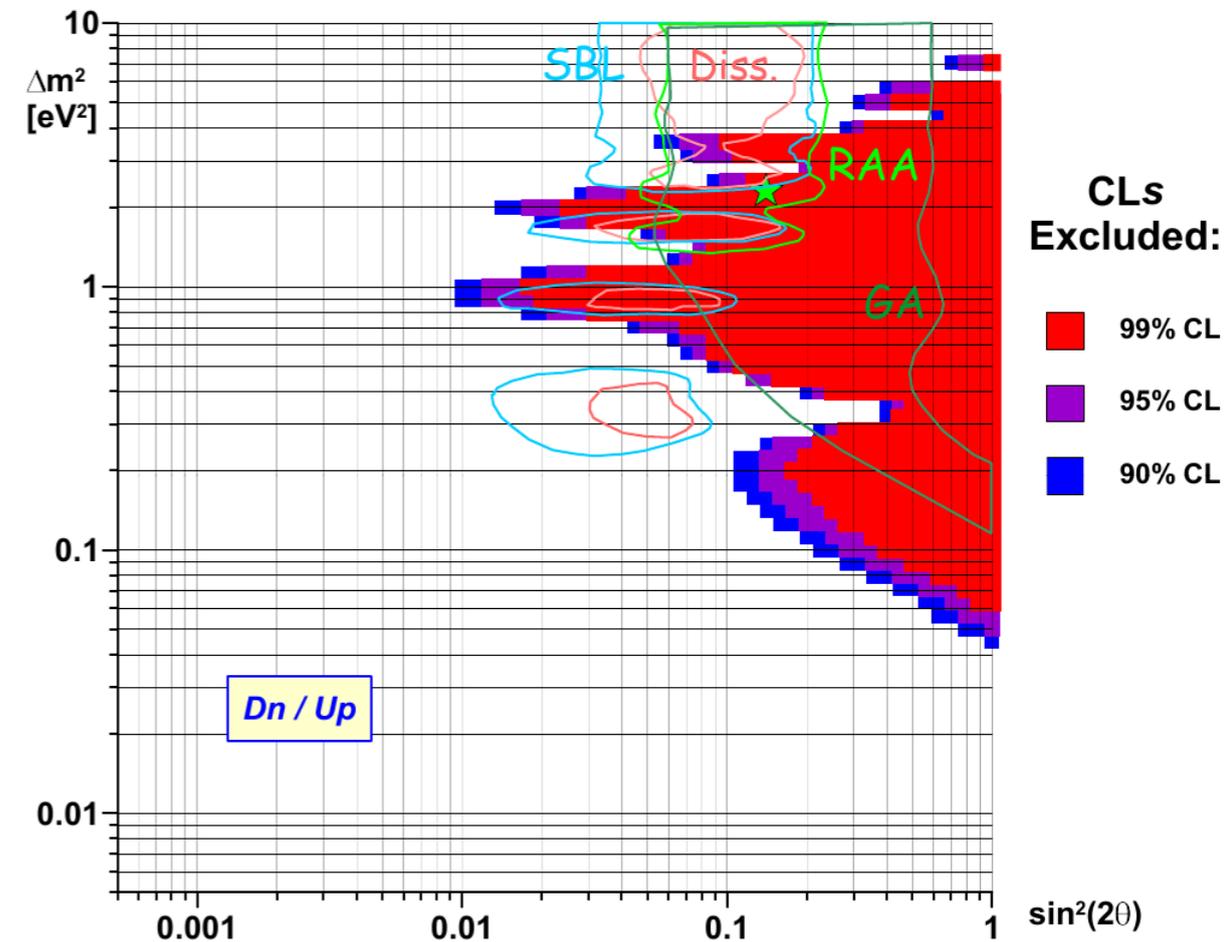
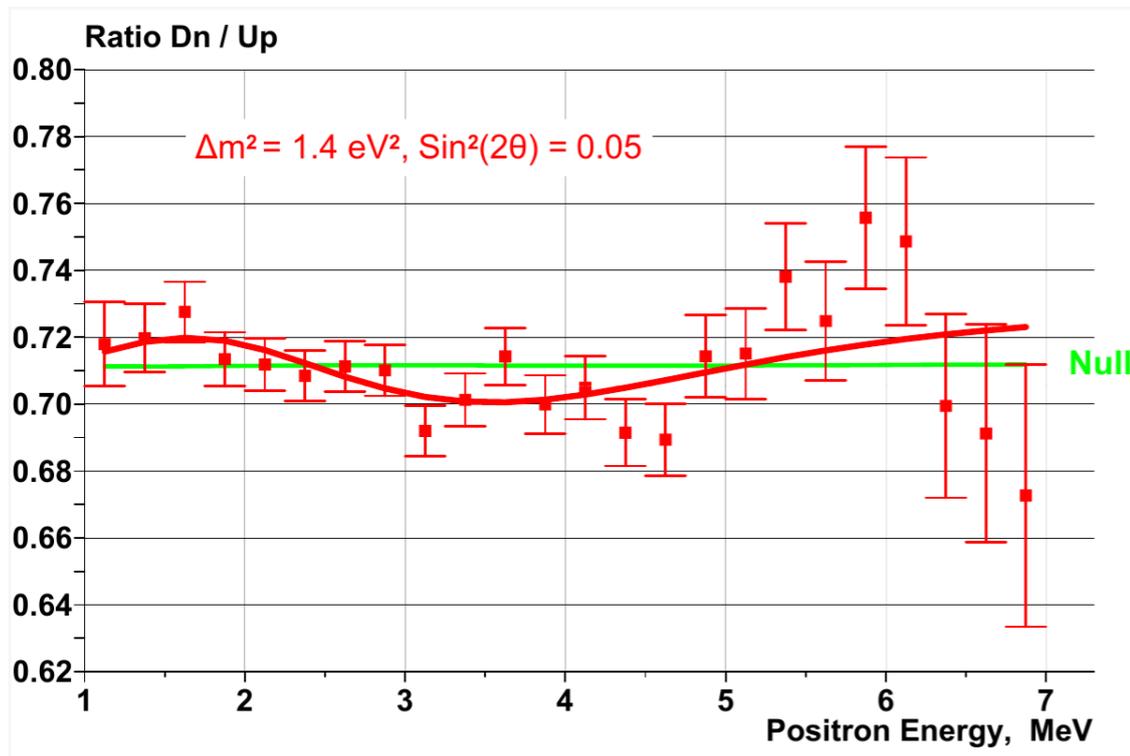
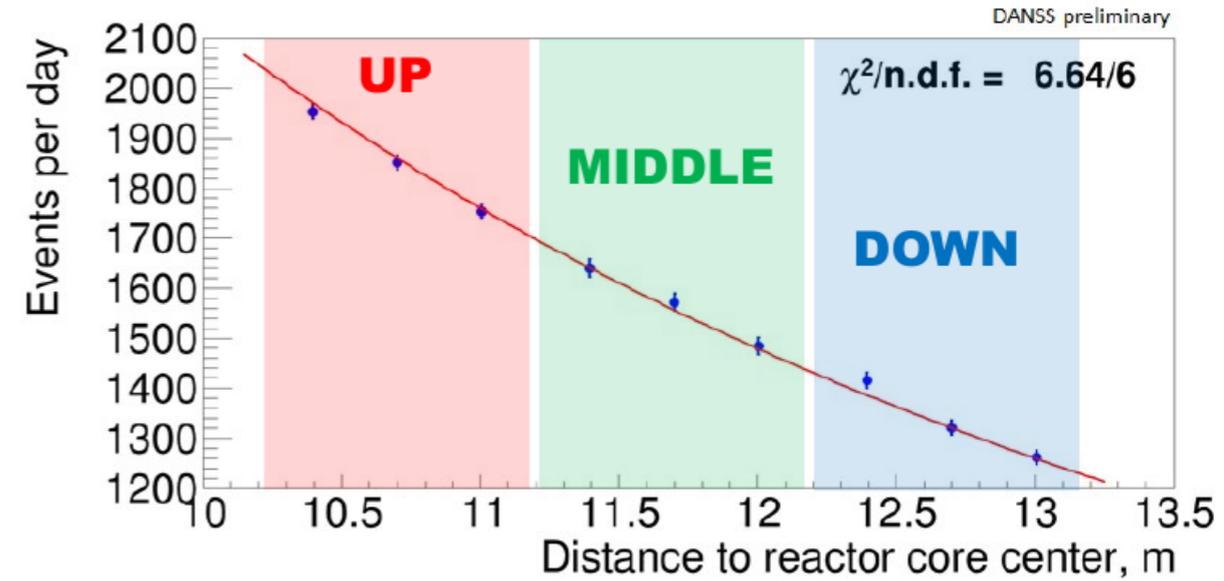




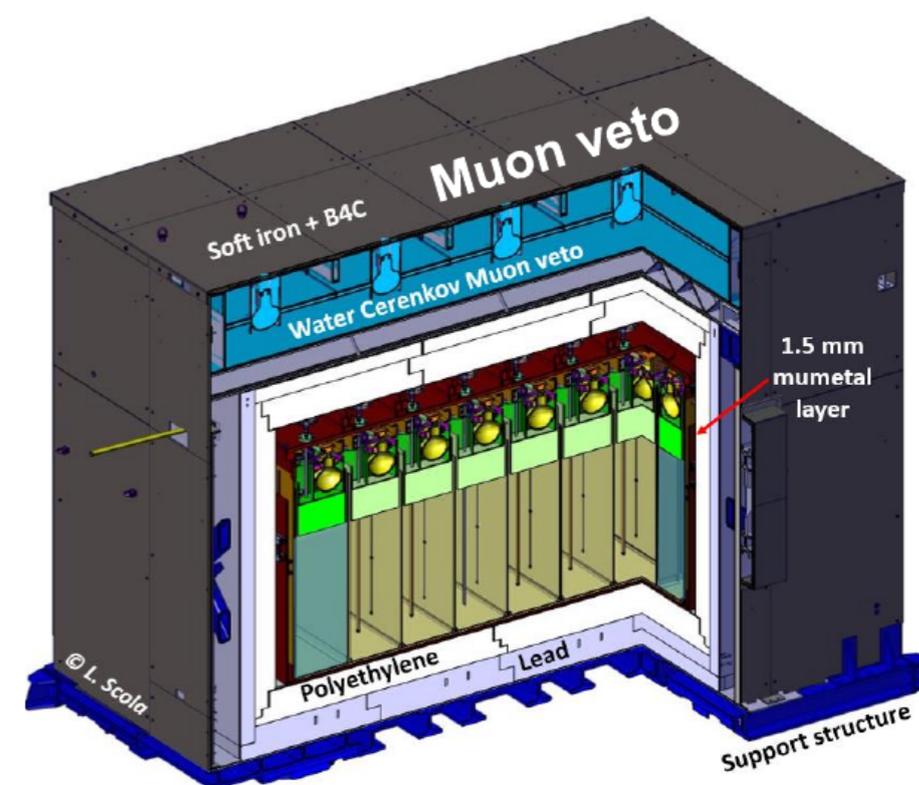
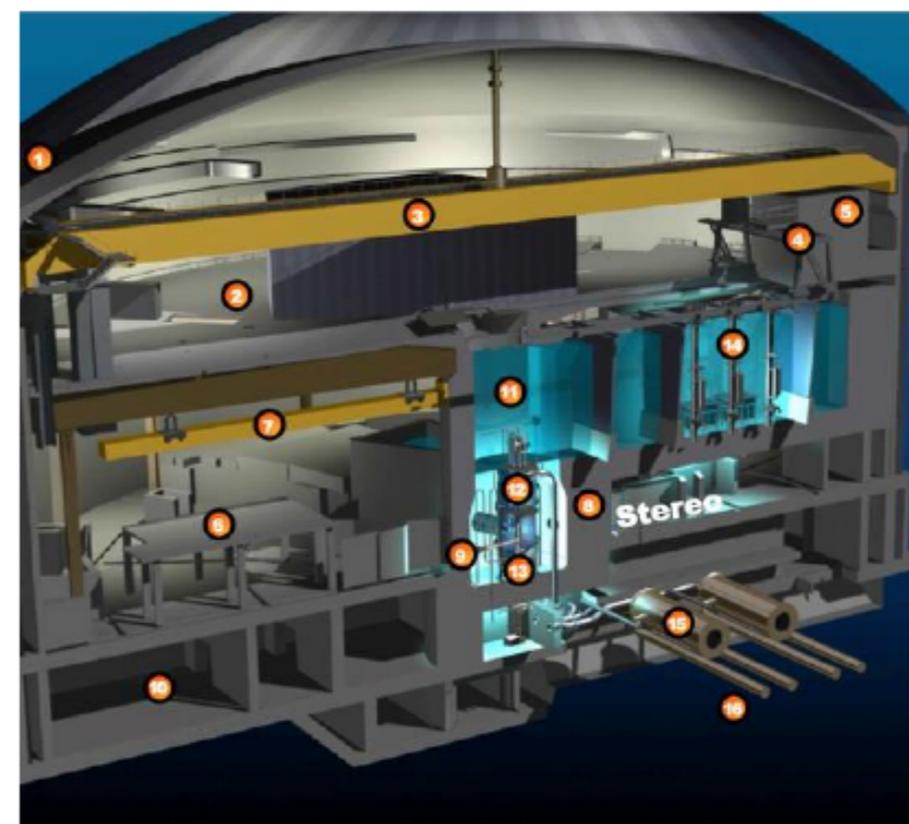
- ▶ IBD intensity follows the $1/L^2$ dependence.
- ▶ Fit with m_d/up and d_n/up (model independent)

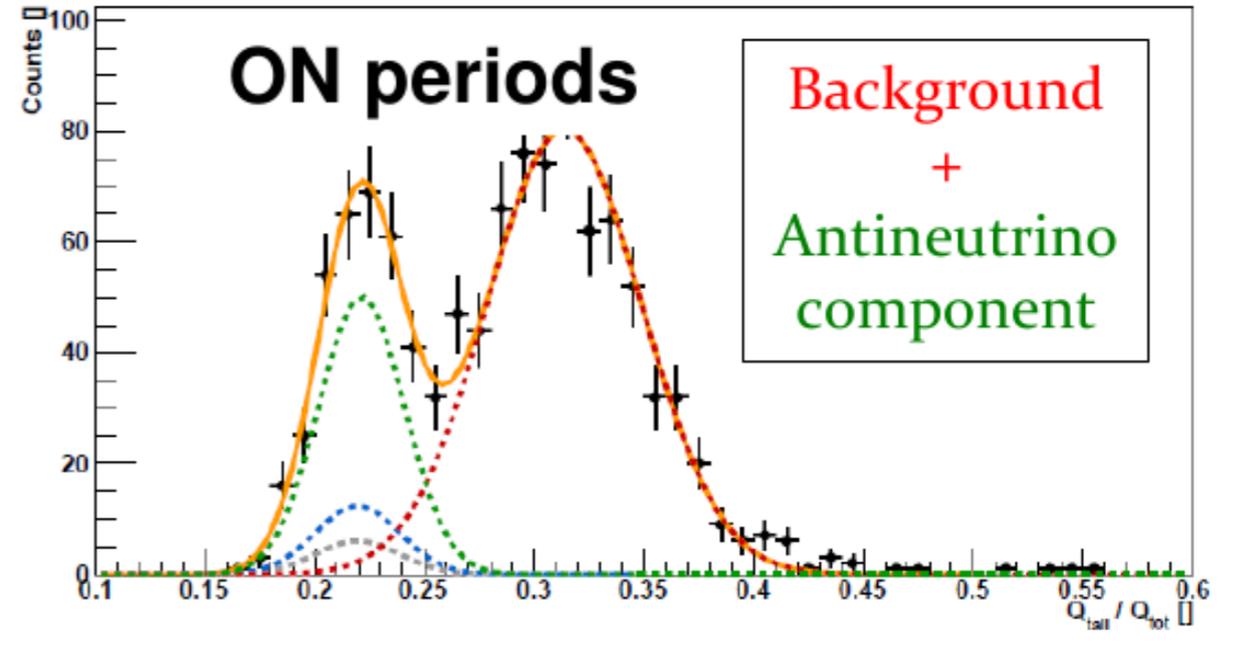
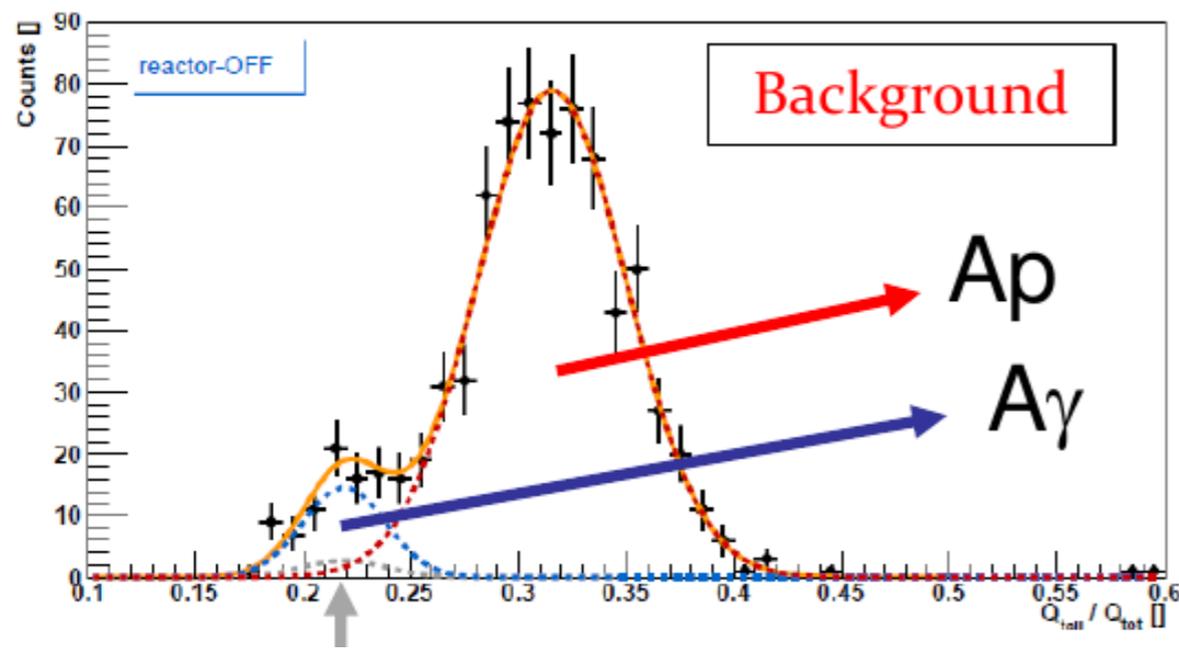
$$\Delta m_{41}^2 = 1.4 \text{eV}^2, \quad \sin^2 2\theta_{14} = 0.05$$

- ▶ DANSS data taking and analyses on-going

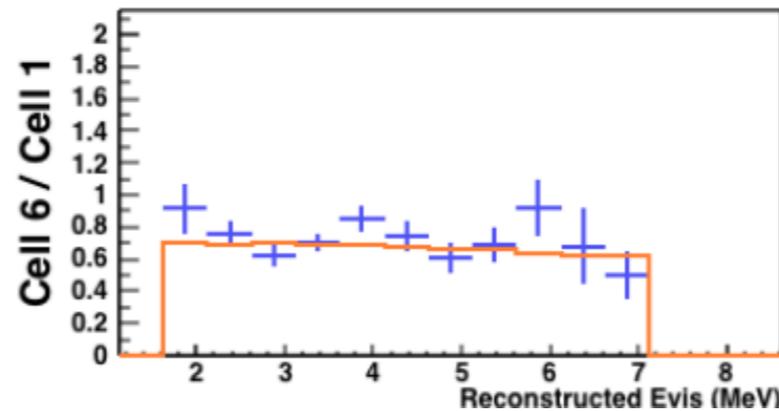
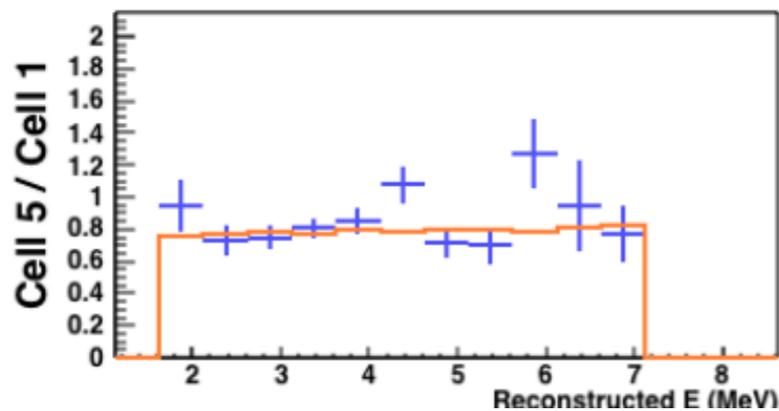
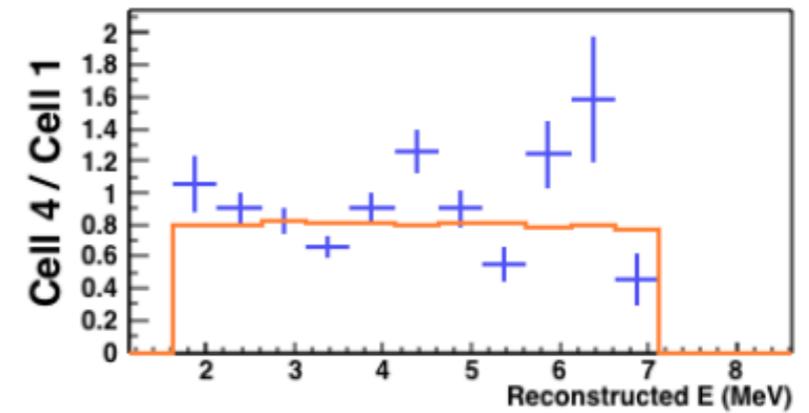
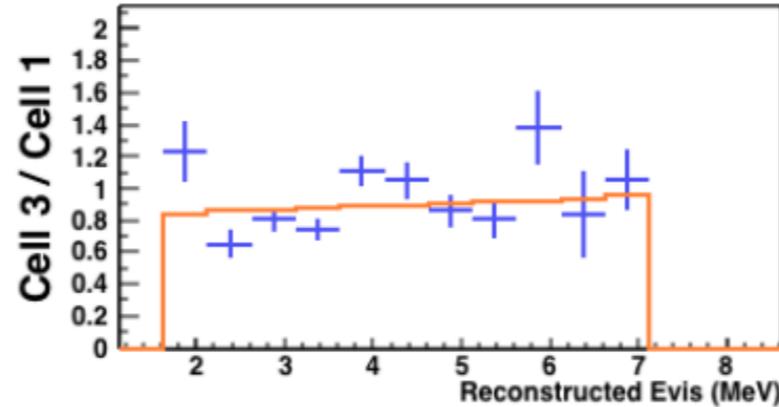
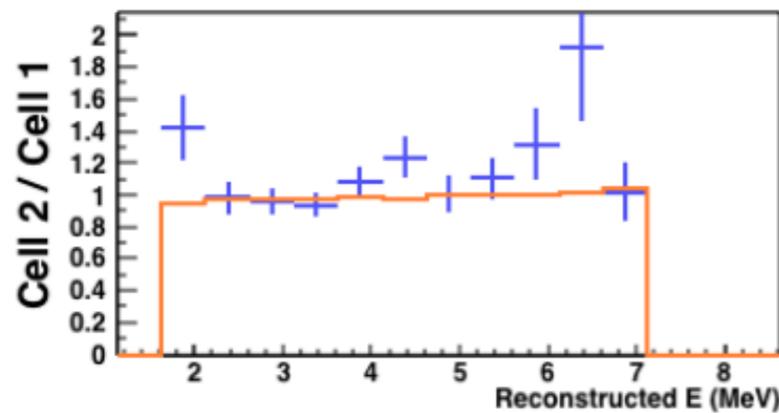


- ▶ 58 MW research reactor (ILL) with highly enriched fuel (^{235}U , 93%)
- ▶ Compact core ($\Phi 40 \times 80$ cm)
- ▶ Baseline: [9-11] m from core
- ▶ Shallow depth and gamma and neutron background due to experimental lines
- ▶ 6 cells filled with GdLS
- ▶ 66 days ON (396 IBD events/day)





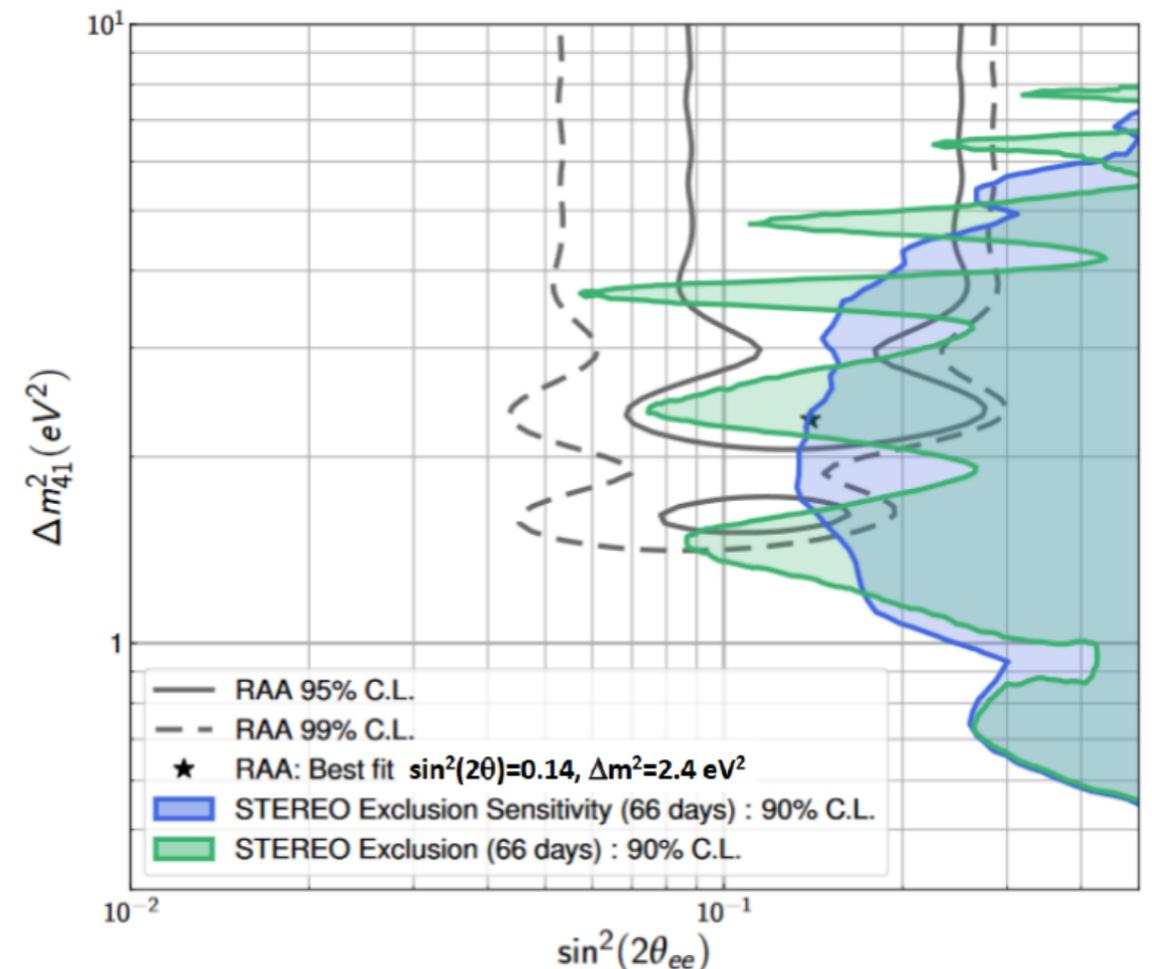
66 days ON data



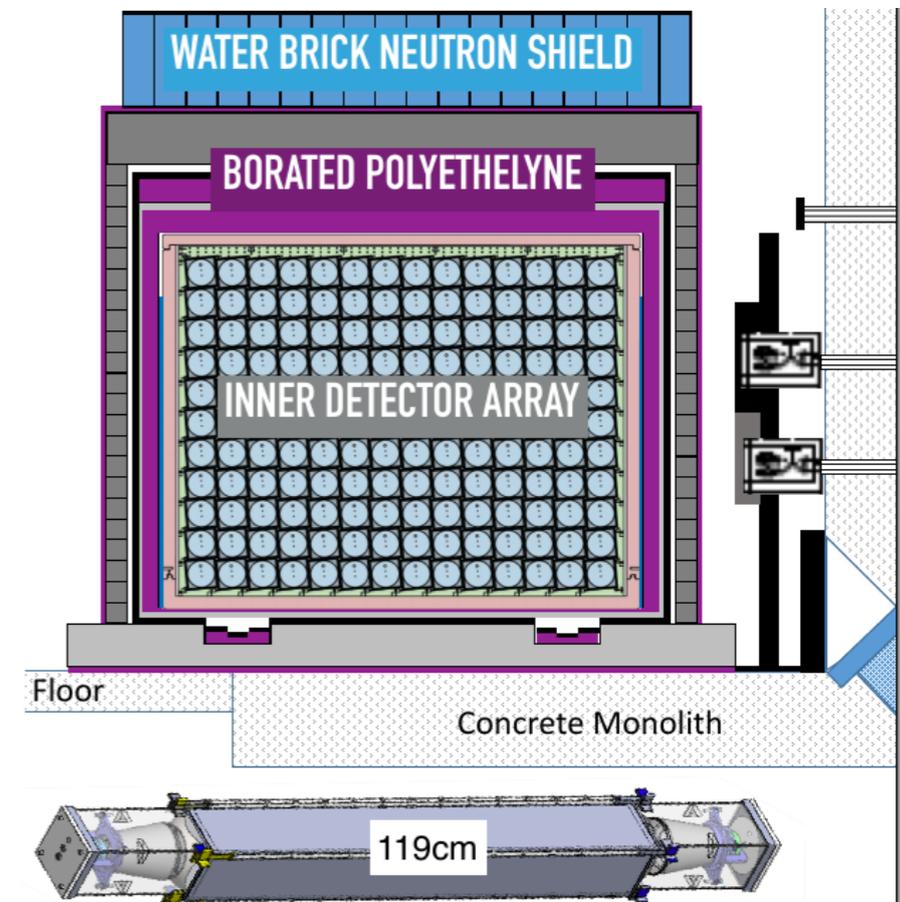
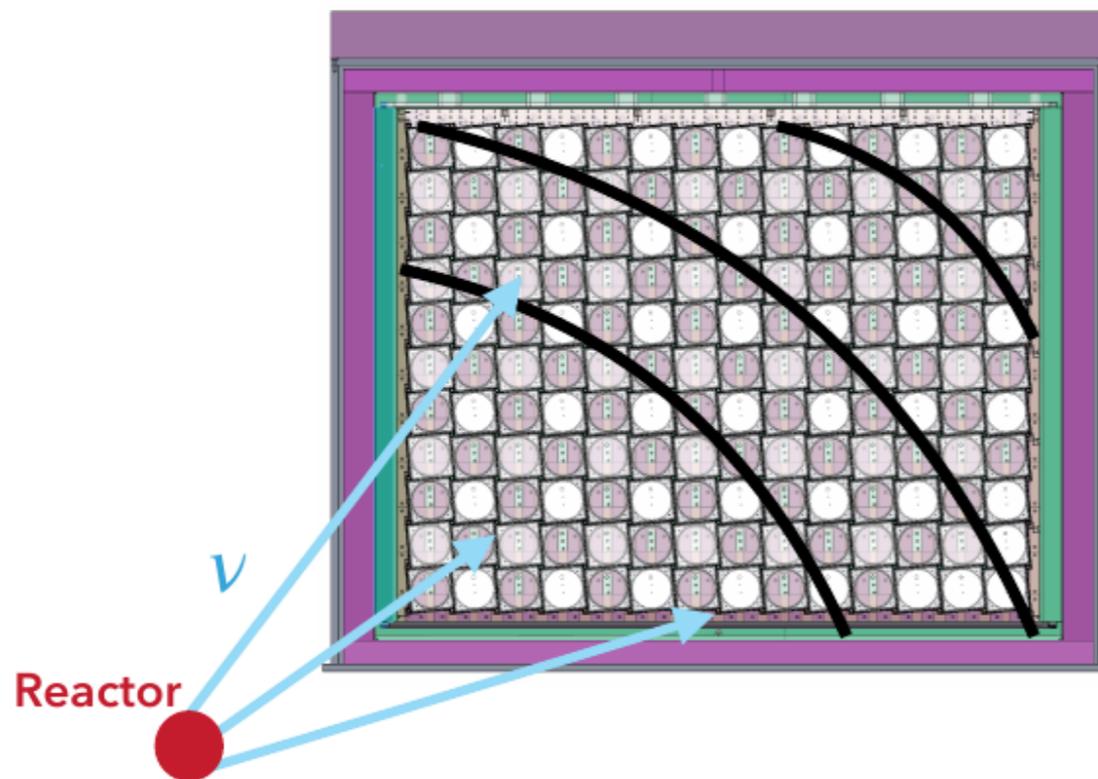
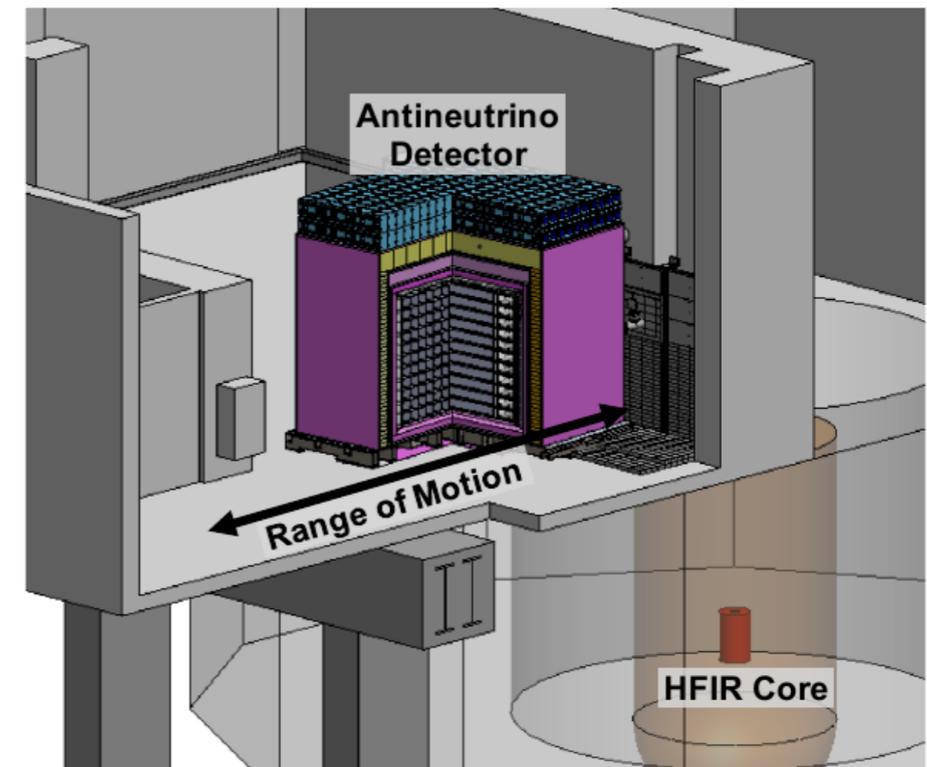
+ Measured ratios

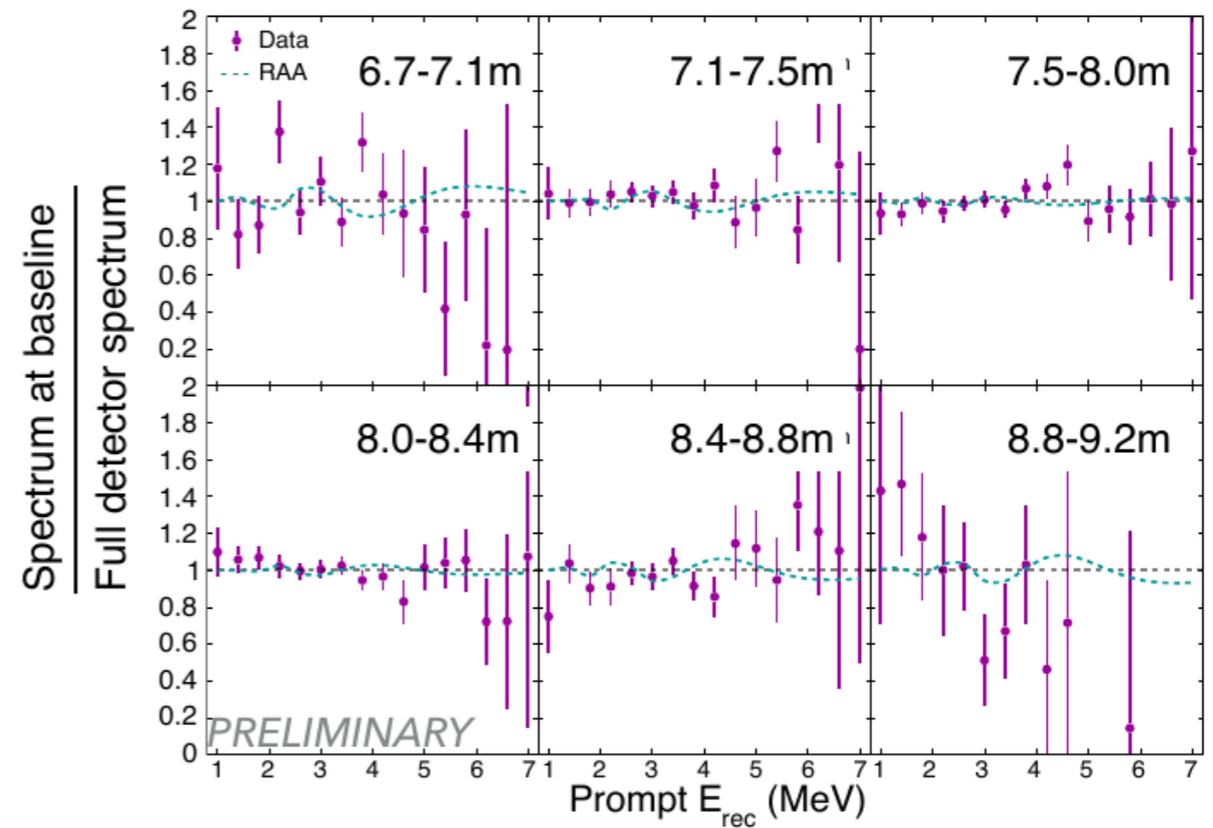
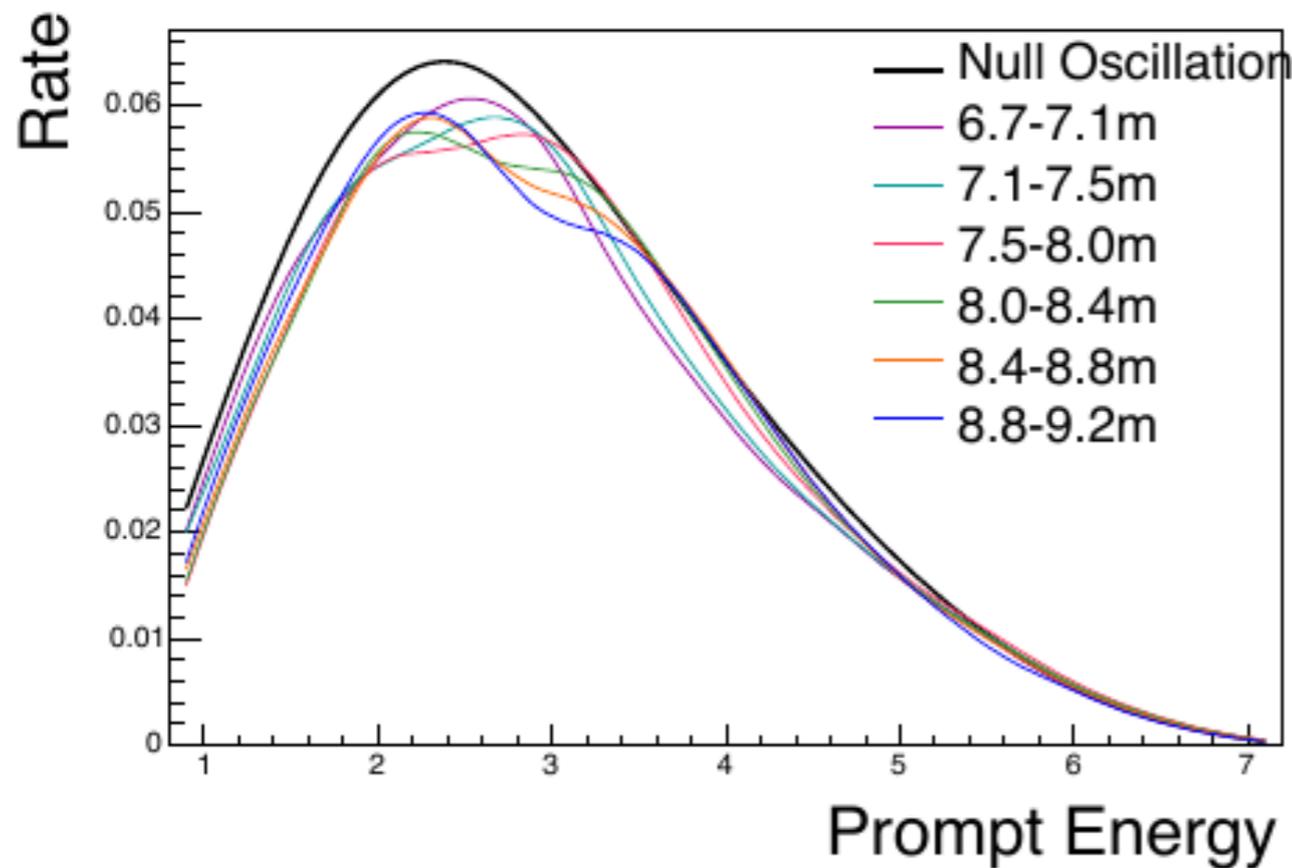
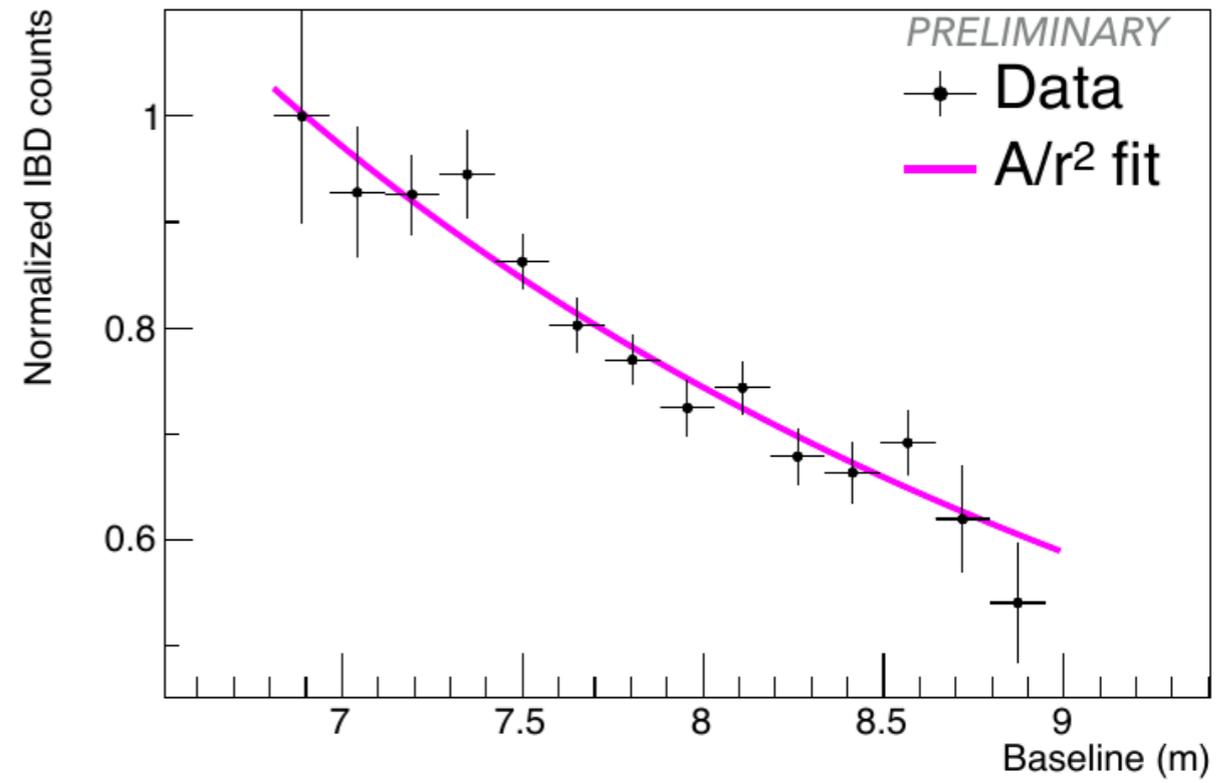
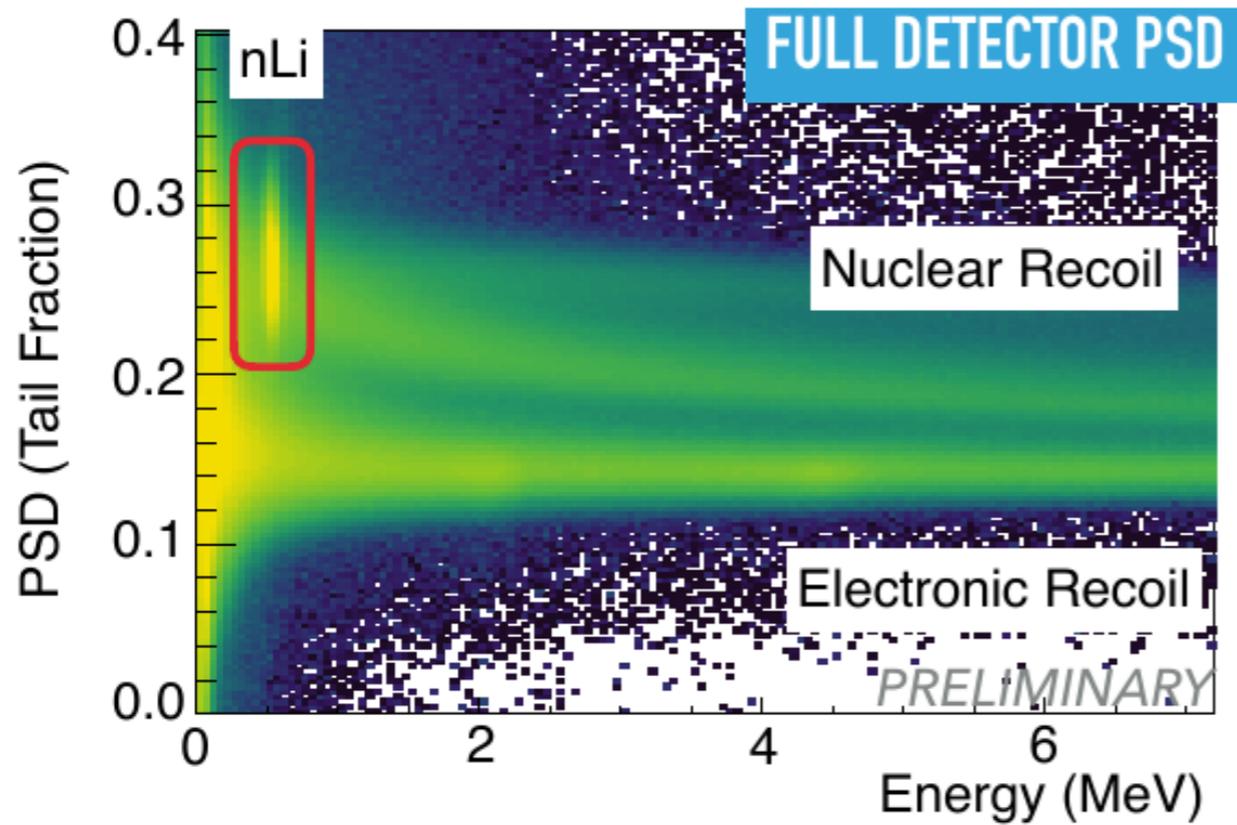
- Non oscillated predictions

- ▶ An exclusion contour has been obtained using the ratio method and raster scan.
- ▶ Best fit value of the RAA rejected at 97.5% C.L.
- ▶ Data taking will continue until end of 2019 (>300 days of reactor-on data)

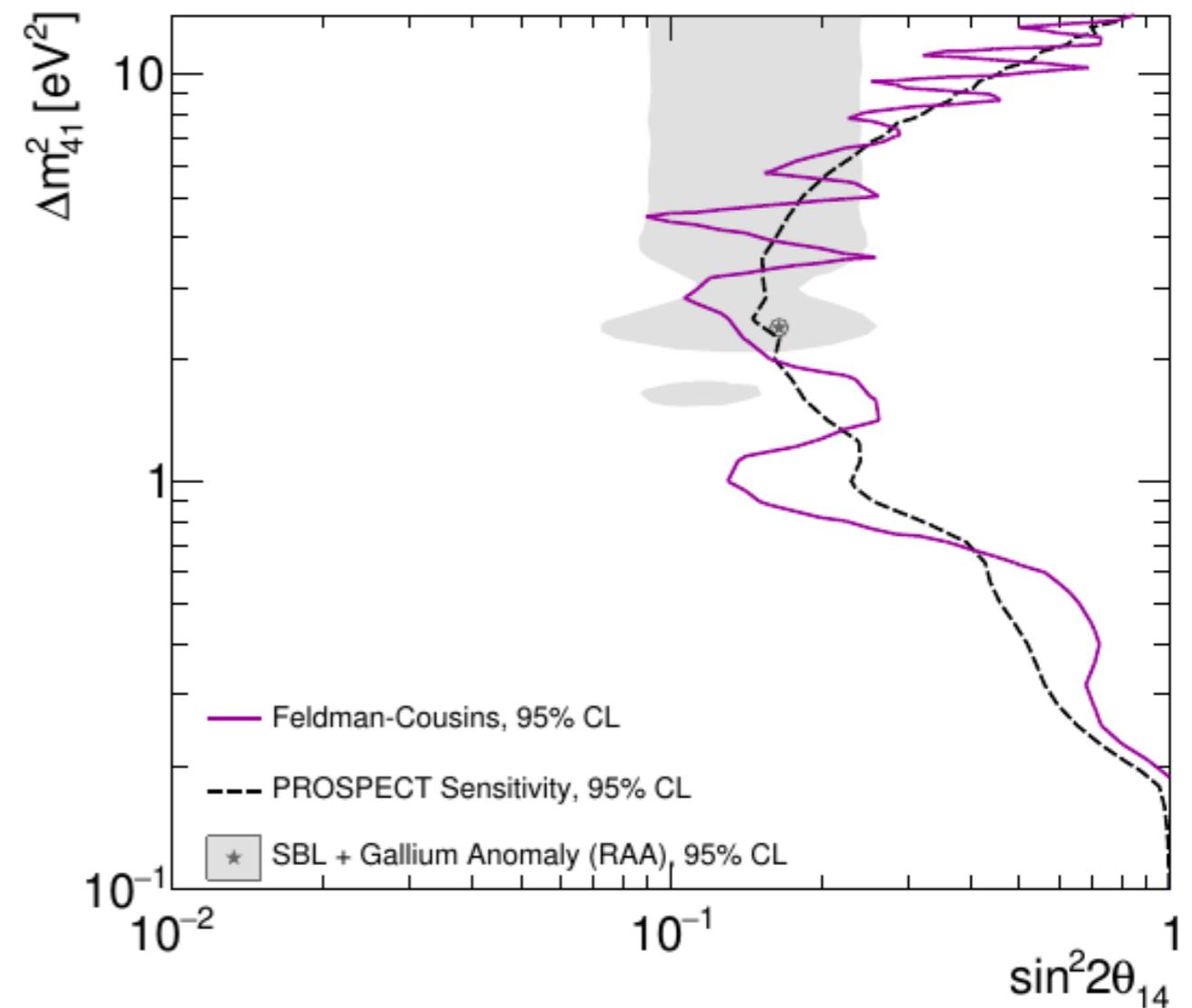


- ▶ 85MW highly enriched uranium reactor (>99% ^{235}U fission)
- ▶ Compact core ($\Phi 44 \times 51$ cm)
- ▶ 154 segments with $^6\text{LiLS}$ (total mass 4ton)
- ▶ Measure neutrino spectrum for each baseline and compare shape to the detected full volume -> **reactor model-independent search for sterile neutrino**
- ▶ 33 (28) days of reactor On(Off): 750 IBDs/day

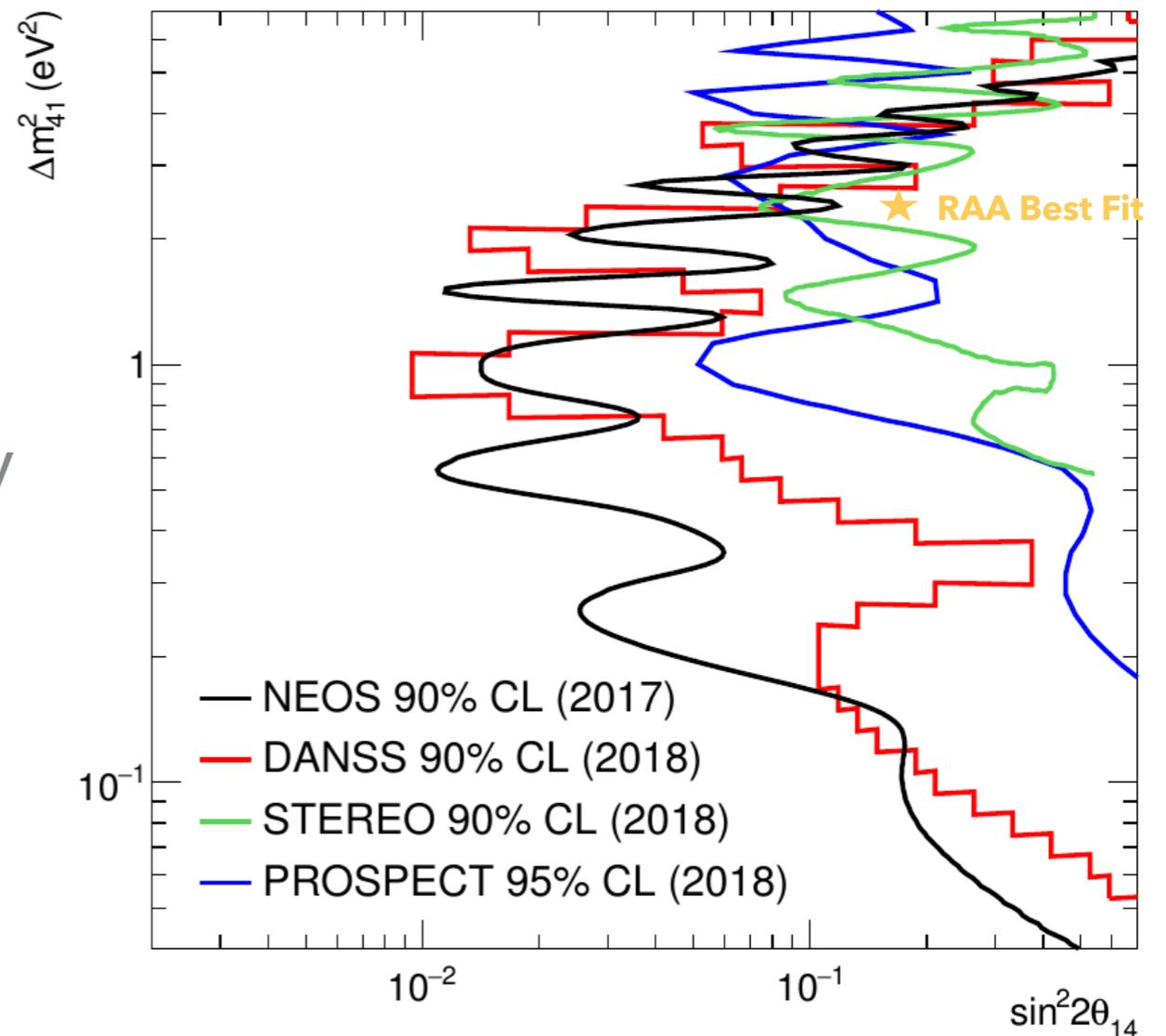




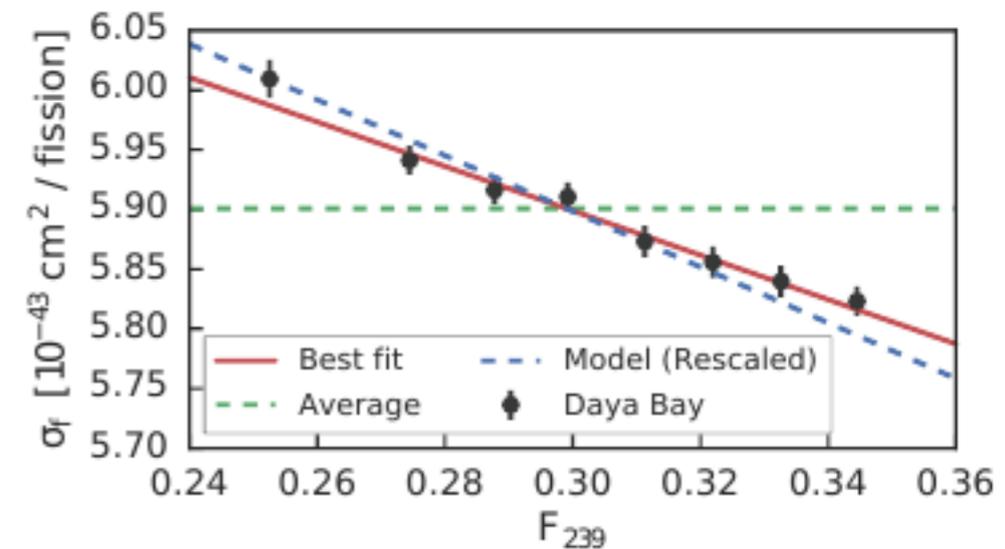
- ▶ 95% exclusion curve based on 33 days reactor on operation
- ▶ Disfavor RAA best fit at $> 95\%$ (2.3σ)



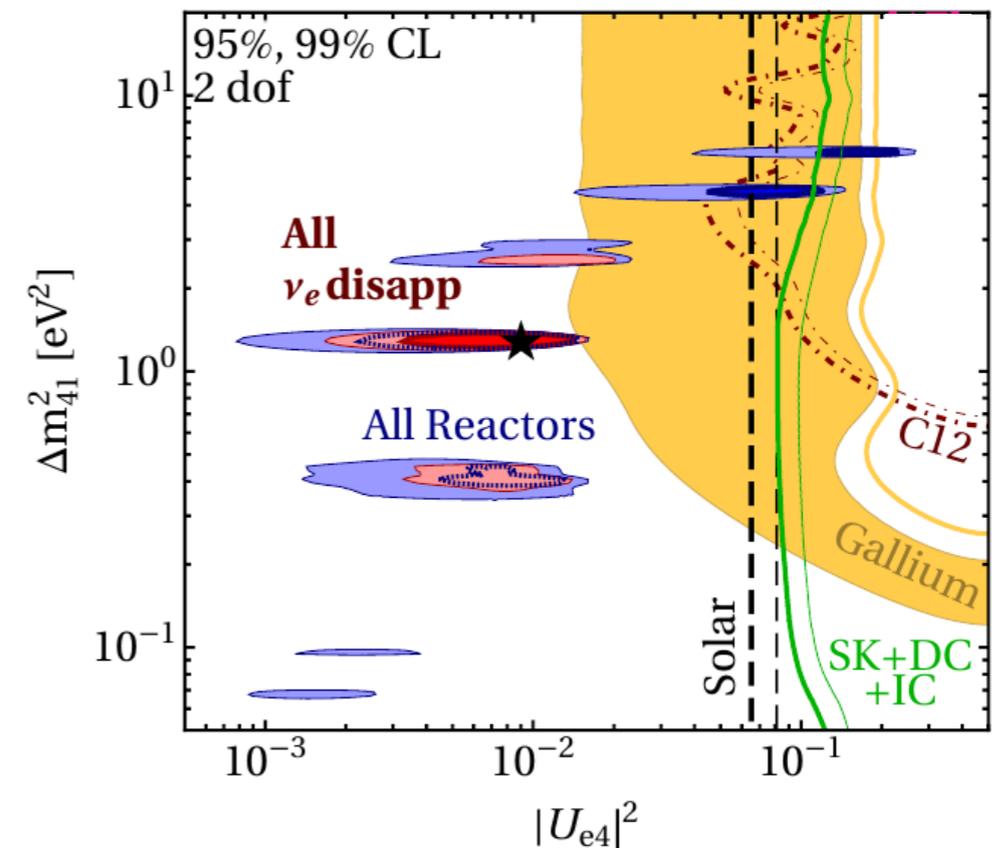
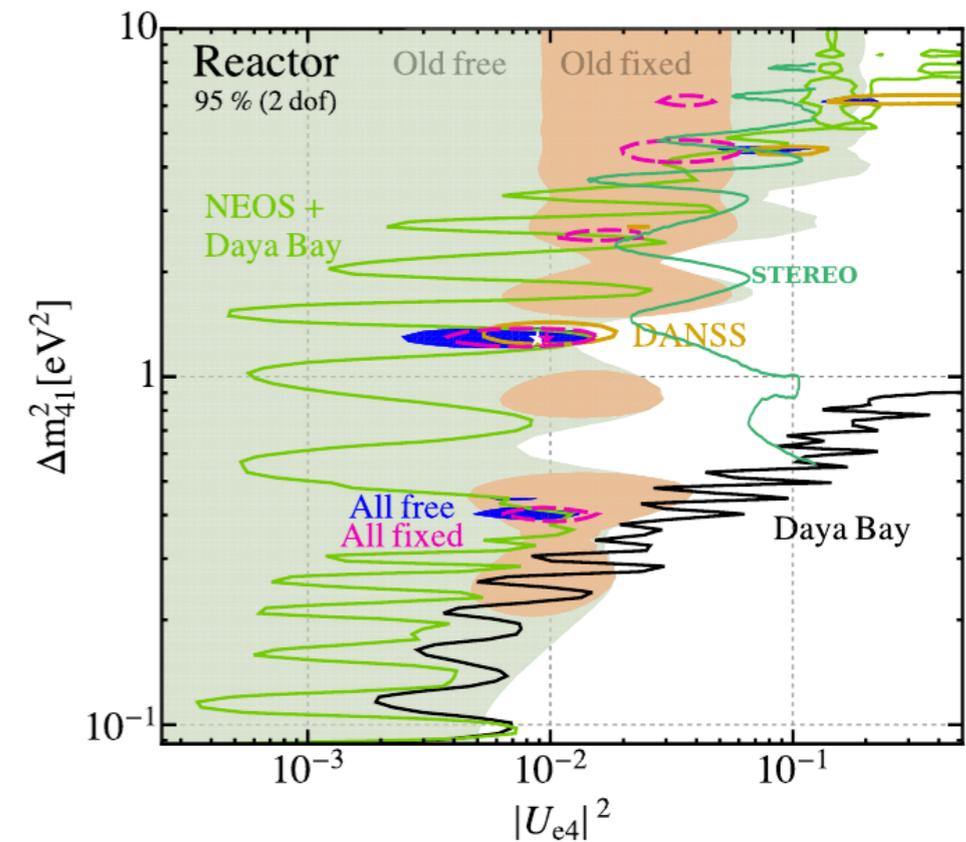
- ▶ Four independent experiments exclude the RAA best fit.
- ▶ NEOS and DANSS oscillation agree with each other.
- ▶ NEOS and DANSS show small energy modulation. -> **hint for SBL oscillation?**
- ▶ NEOS has 5 MeV bump at prompt energy spectrum.



- ▶ Results on the overall IBD yield from Daya Bay:
 - ▶ 1.7σ below expectation
 - ▶ dependence on fuel composition is 3.1σ away from model

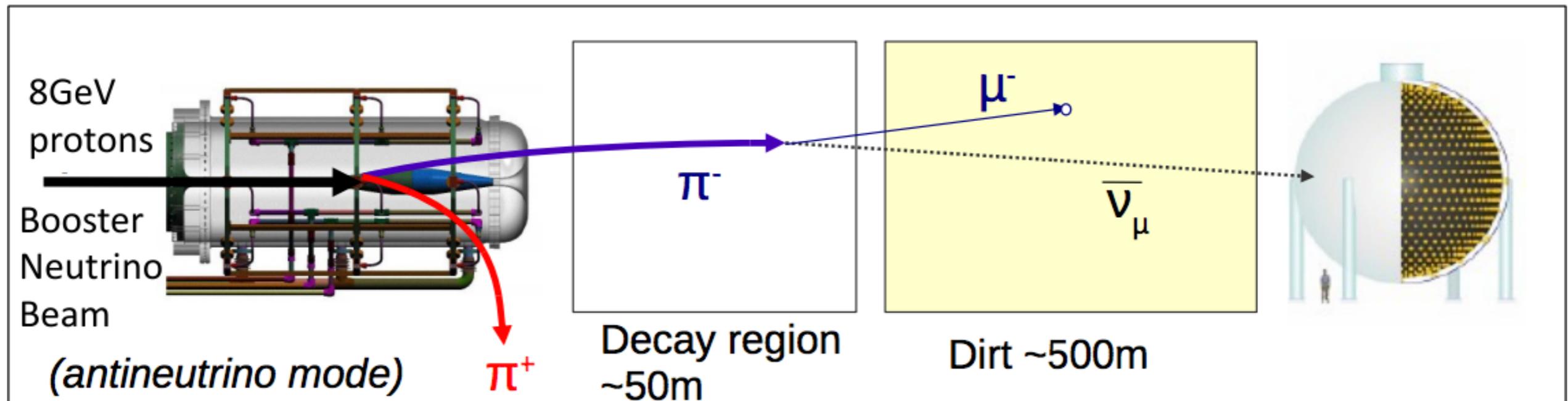


- ▶ Global analysis of all reactor data
 - ▶ two models:
 - ▶ Free: unconstrained normalizations for [235U, 238U, 239Pu, 241Pu]
 - ▶ Fixed: assumes Huber fluxes
 - ▶ **2.9σ(3.5σ)** hint for sterile neutrino from analysis with free(fixed) fluxes
 - ▶ fit dominated by DANSS+NEOS
- ▶ Reactor, gallium, Karmen, LSND, atmospheric, and solar neutrino
 - ▶ small tension between gallium and reactor data
 - ▶ hint for sterile slightly increase to **3.2σ(3.8σ)**
 - ▶ global best fit: $\Delta m_{41}^2 = 1.3 \text{eV}^2$

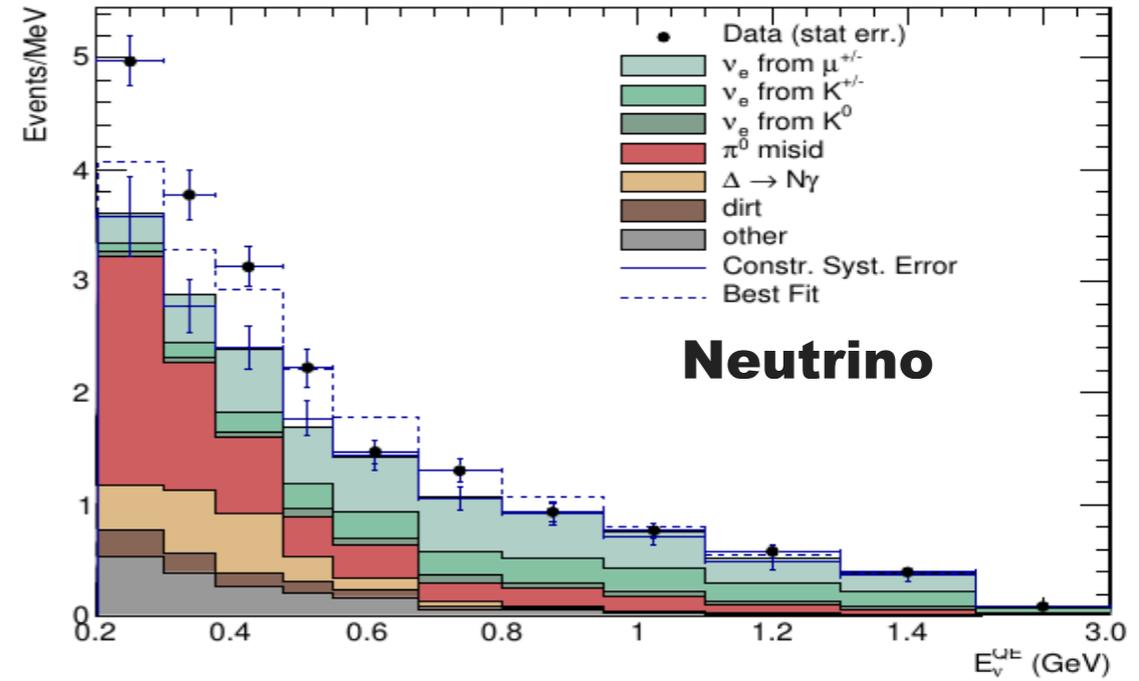
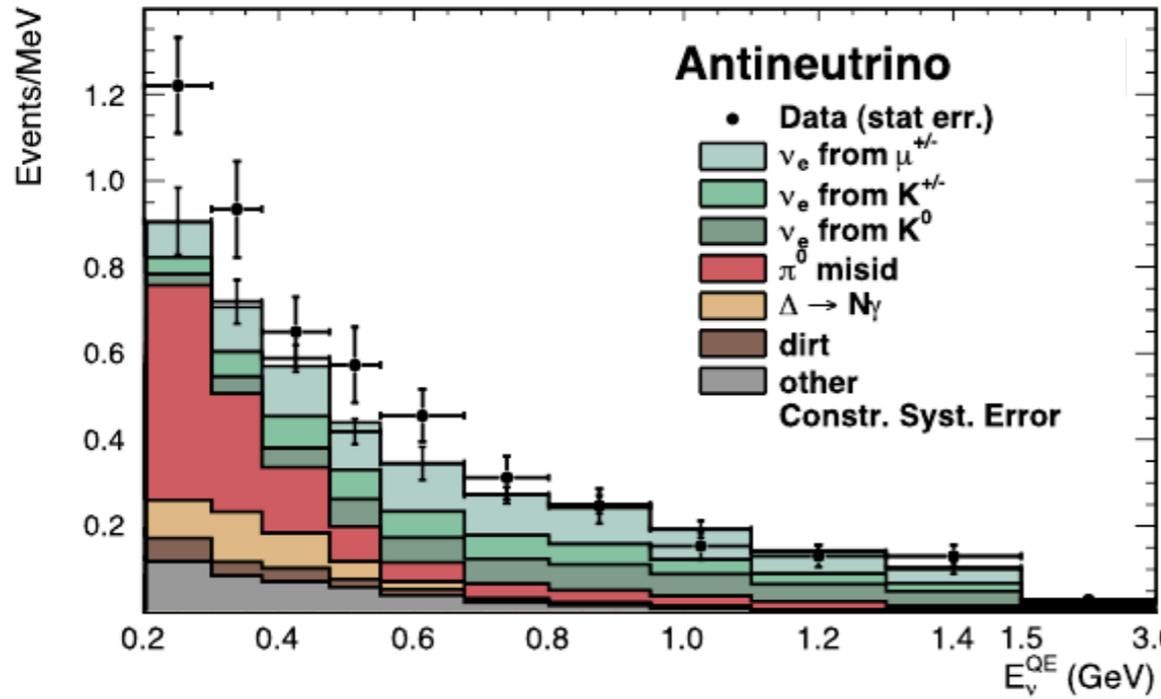


M. Dentler et al., arXiv:1803.10661.

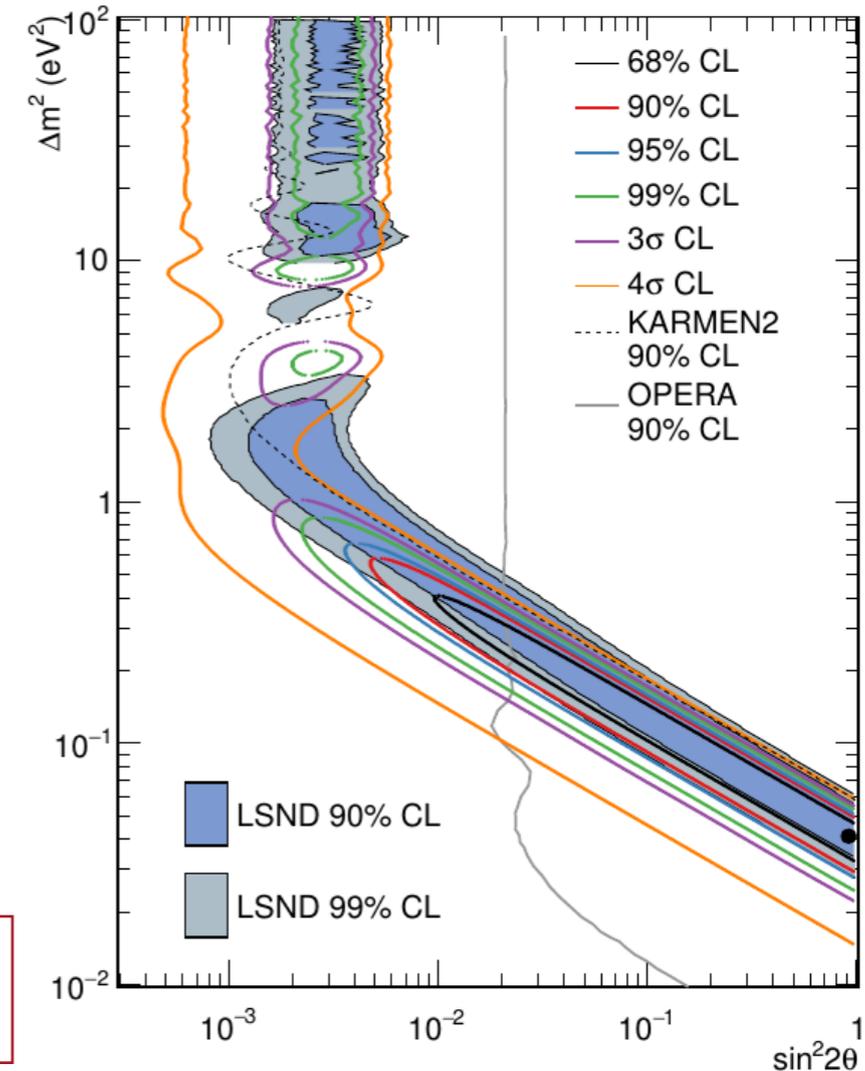
- ▶ MiniBoone results
- ▶ Future experiments
 - ▶ SBN (Short Baseline Neutrino)
 - ▶ JSNS2 (J-PARK Sterile Neutrino Search at J-PARK Spallation Neutron Source)



- ▶ MiniBooNE was approved in 1998, with the goal of addressing the LSND anomaly
- ▶ Similar L/E
 - ▶ LSND ~ 30m/30MeV
 - ▶ MiniBooNE ~ 500m/500MeV
- ▶ 800-ton mineral oil Cherenkov detector
- ▶ 11.27×10^{20} POT in ν mode + 12.84×10^{20} POT in $\bar{\nu}$ mode

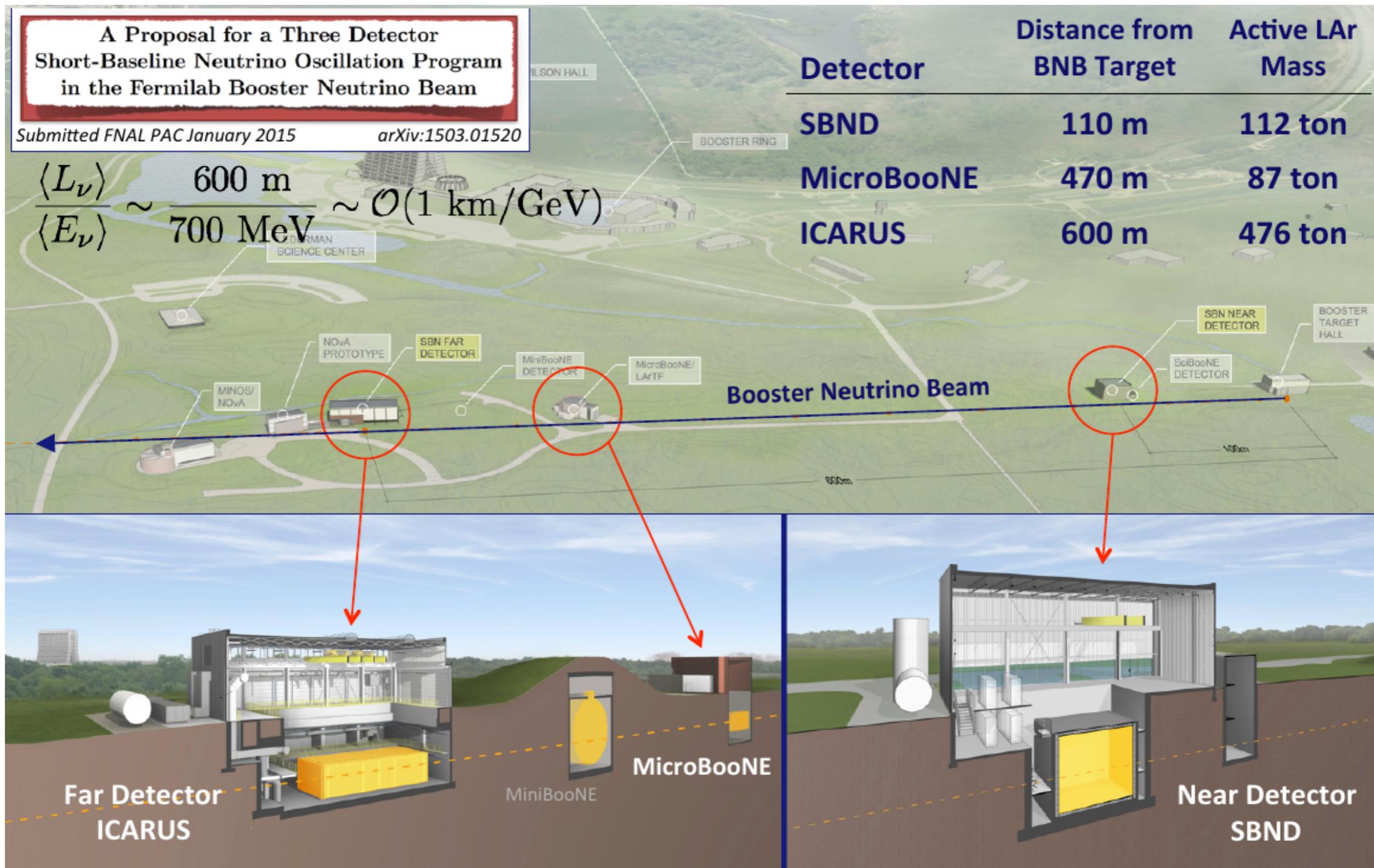


- ▶ Antineutrino data (2012)
 - ▶ antineutrino signal is compatible with 2ν oscillation.
 - ▶ antineutrino signal is compatible with LSND.
- ▶ Neutrino data (2018)
 - ▶ neutrino signal does not rule out LSND.
- ▶ Combined data
 - ▶ combined fit -> **4.8σ** evidence for sterile neutrino
 - ▶ with LSND (3.8σ), global preference for sterile neutrino is **6.1σ**.



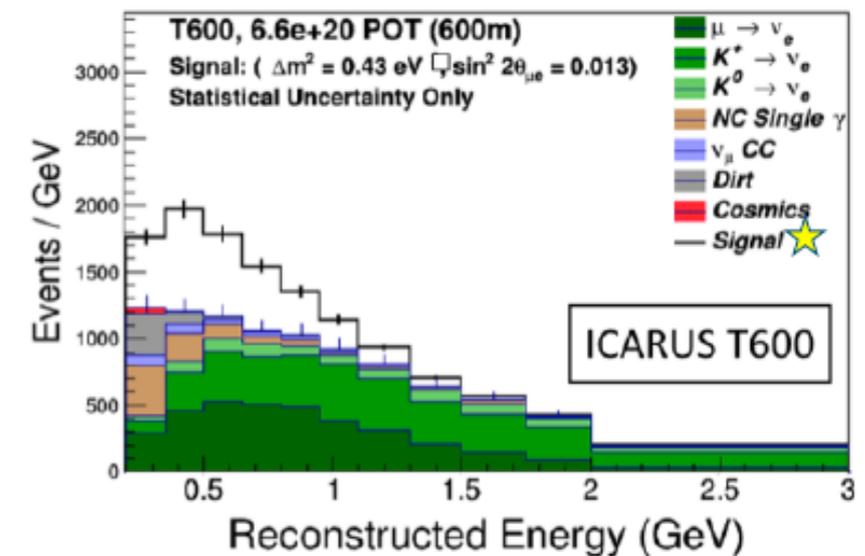
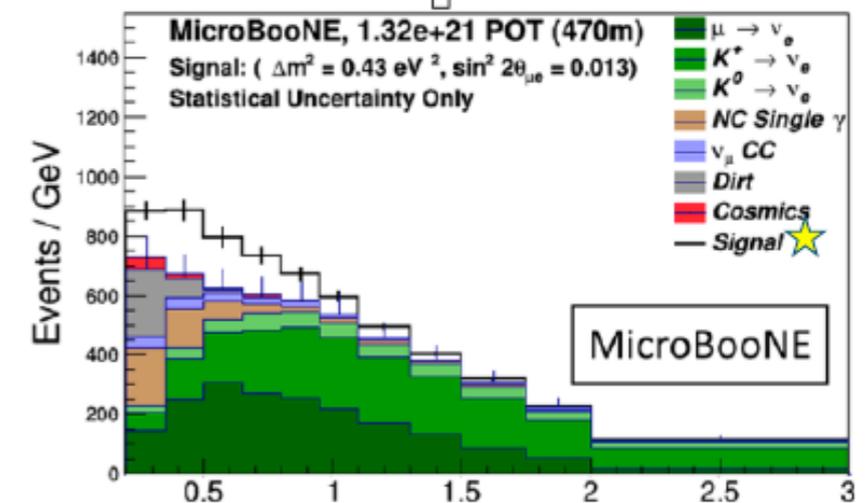
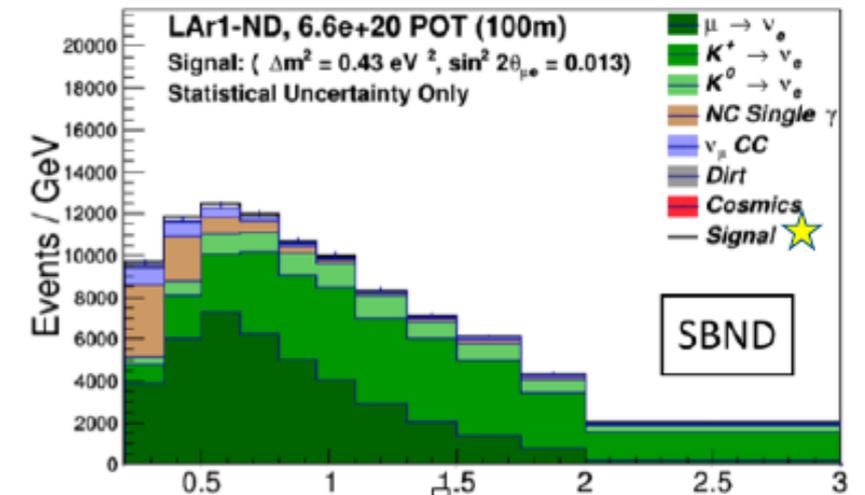
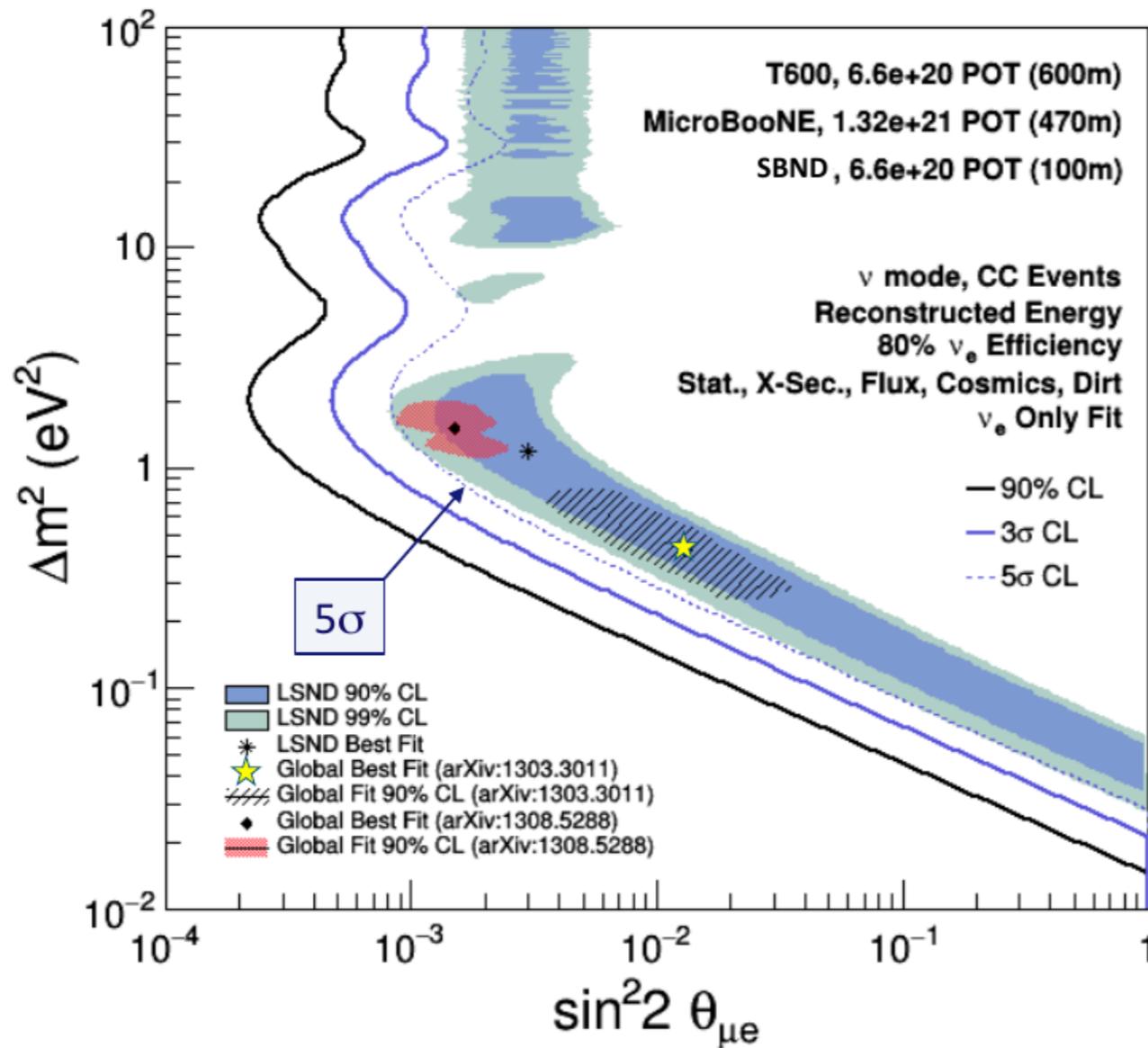
A.A. Aguilar-Arevalo et al. [MiniBooNE collab], PRL 110 (2013) 161801 [arXiv:1303.2588].
 A.A. Aguilar-Arevalo et al. [MiniBooNE collab], arXiv:1805.12028.

SBN PROGRAM @ FERMILAB

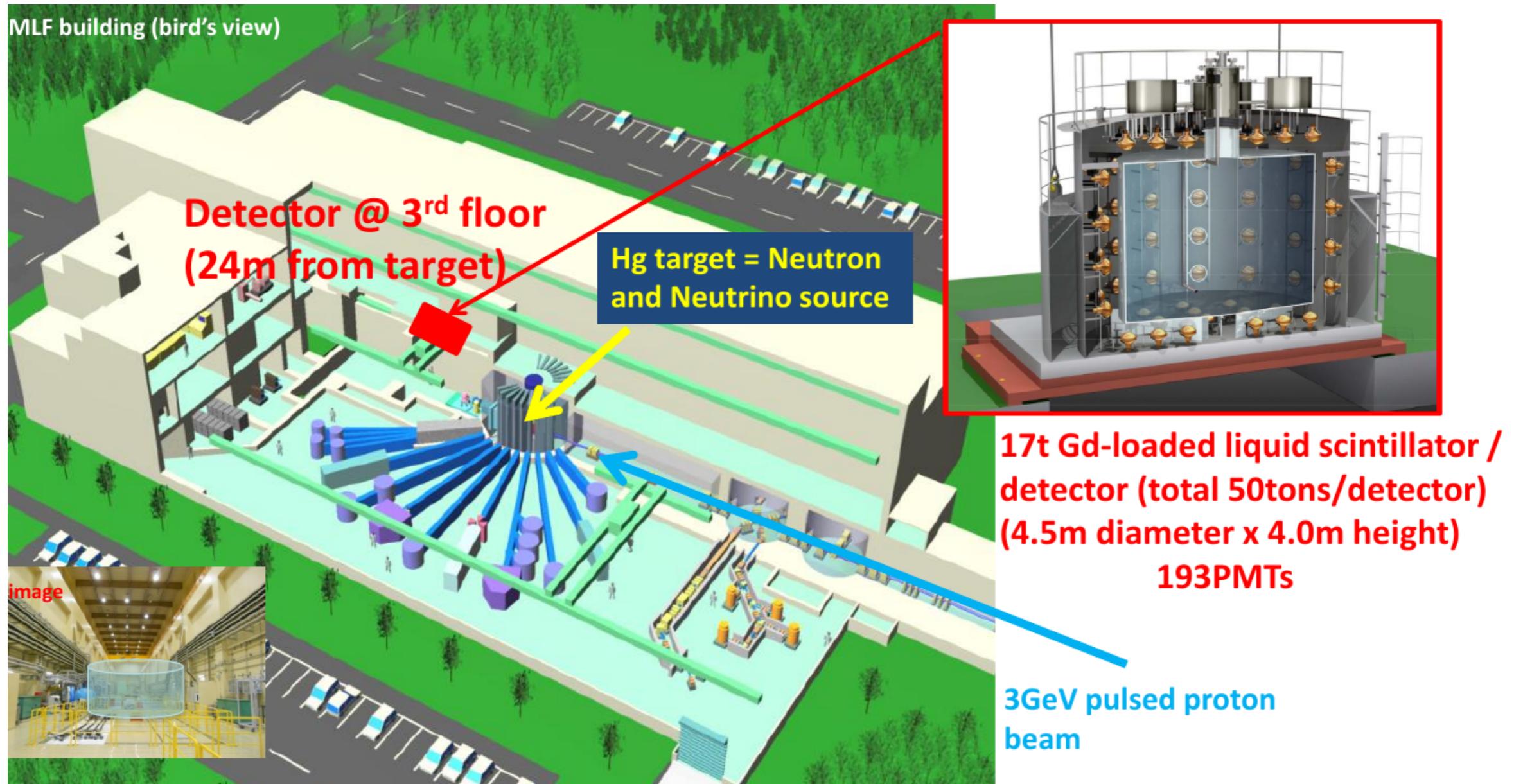


SBN PROGRAM @ FERMILAB

Definitive program to address LSND/MiniBoone anomalies in next ~5 years.



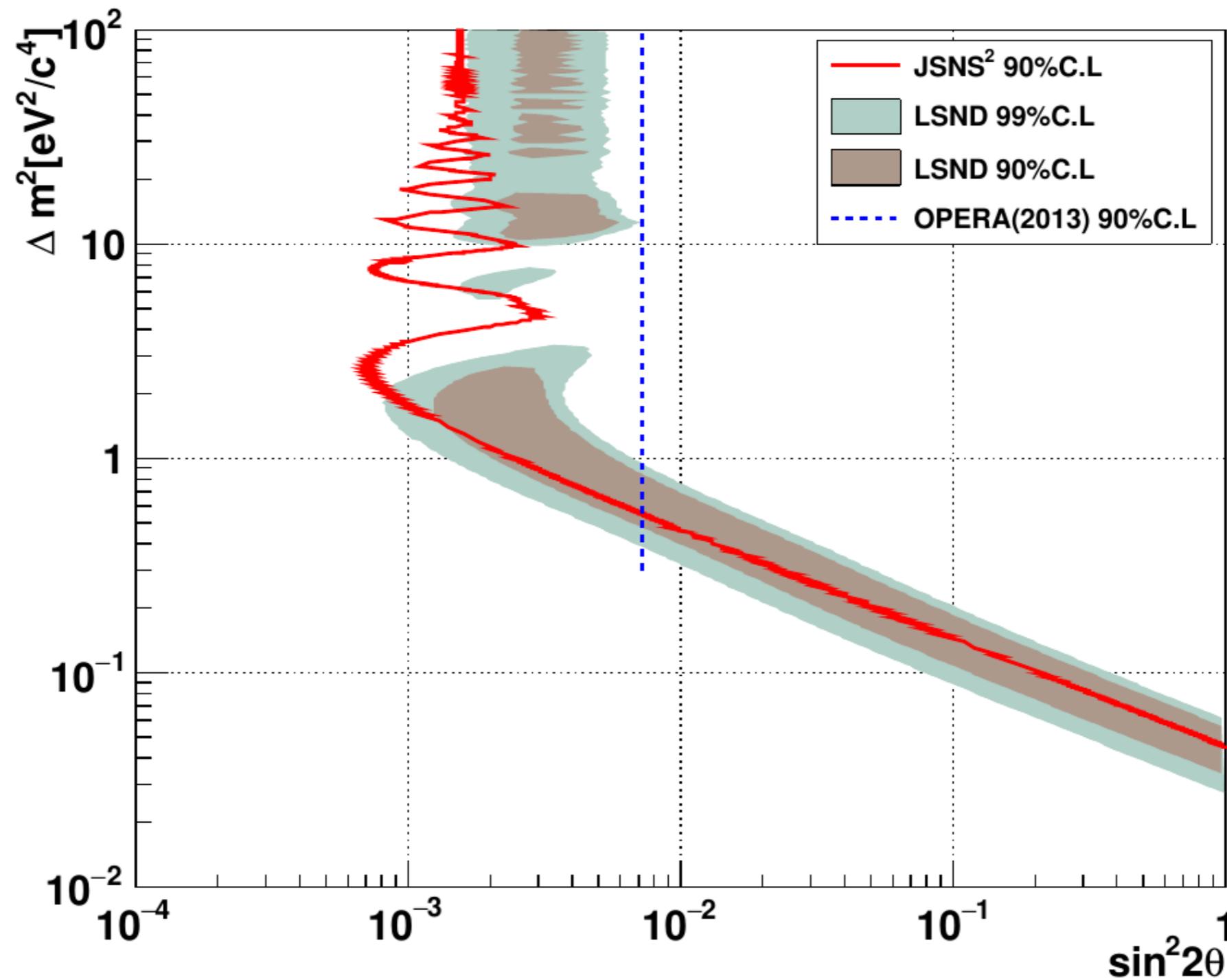
JSNS2 @ J-PARK



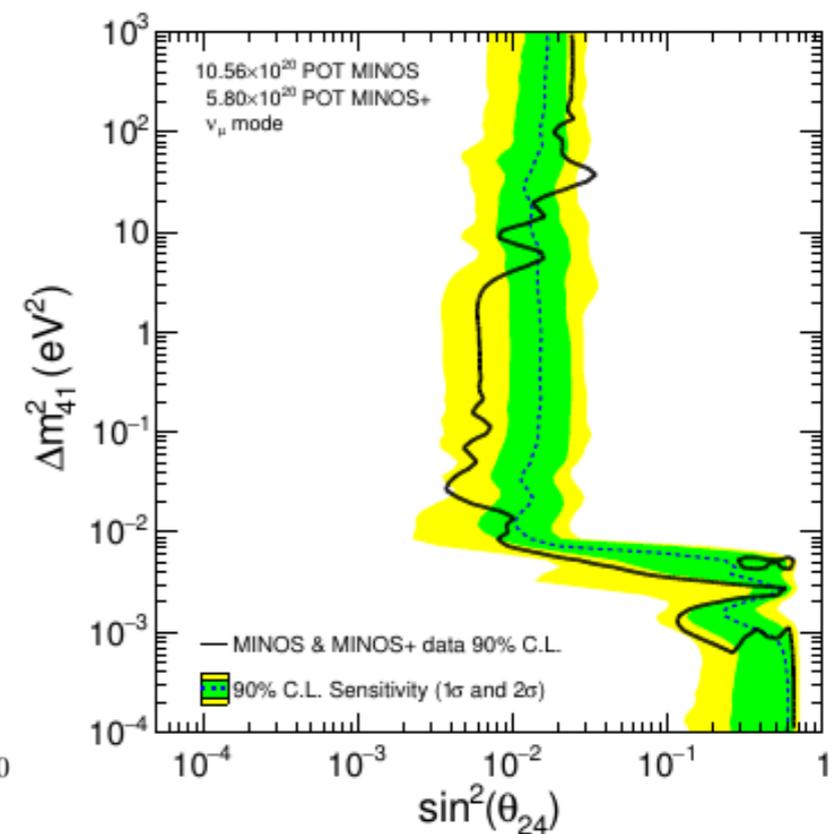
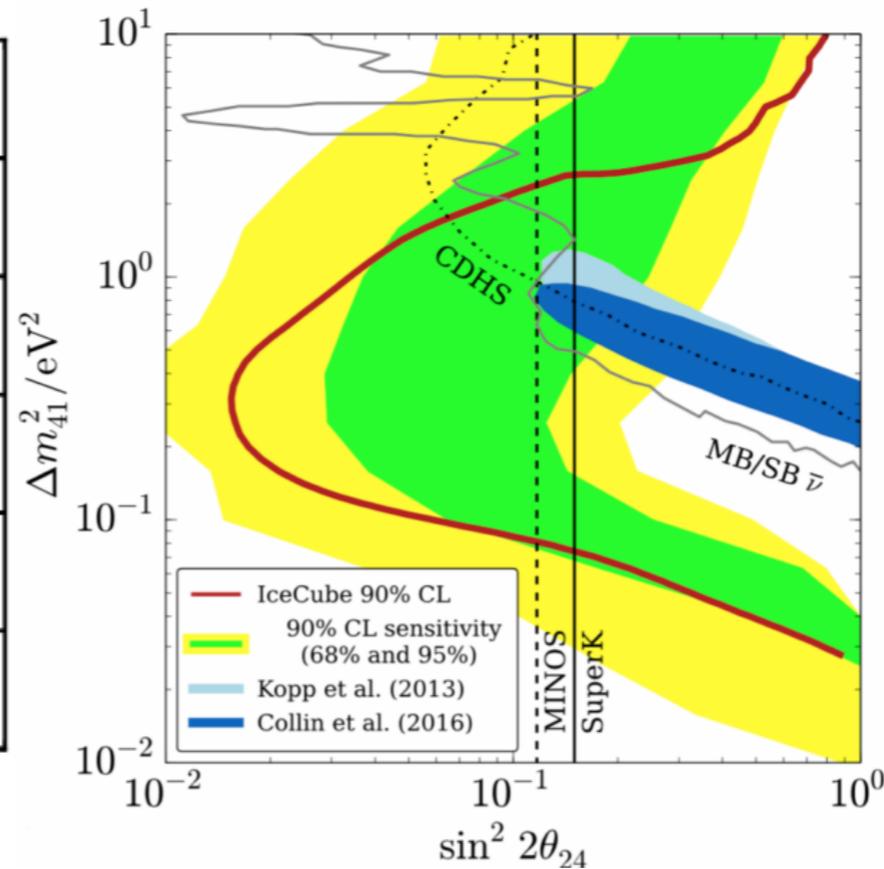
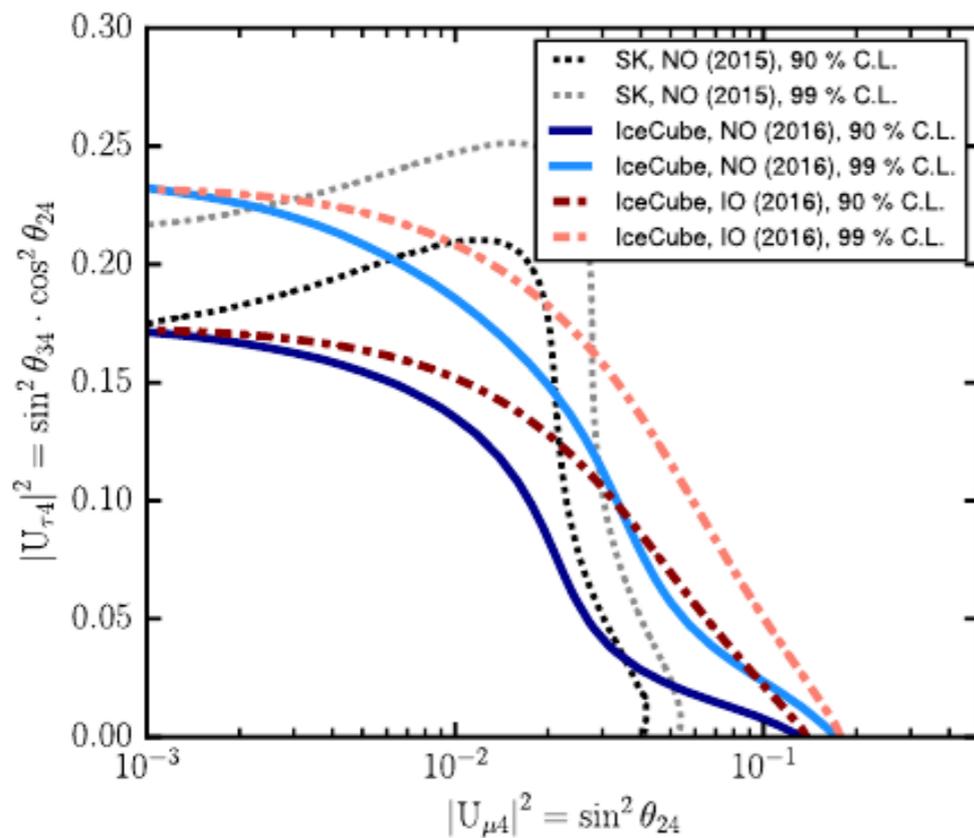
Searching for neutrino oscillation : $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$ with baseline of 24m.
no new beamline, no new buildings are needed \rightarrow quick start-up

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Sensitivity of the JSNS2 experiment with the latest configuration (1 MW \times 3years \times 1 detector).



- ▶ IceCube atmospheric data probe neutrino mixing with heavy state.
- ▶ Sterile neutrino oscillation also studied by NOvA using neutral-current data.
- ▶ Recent MINOS/MINOS+ analysis improves bound on neutrino disappearance.



M.G. Aartsen et al. [IceCube collab], Phys. Rev. D 95 (2017) 112002 [arXiv:1702.05160].

M.G. Aartsen et al. [IceCube collab], Phys. Rev. Lett. 117 (2016) 071801 [arXiv:1605.01990].

P. Adamson et al. [NOvA collab], Phys. Rev. D 96 (2017) 072006 [arXiv:1706.04592].

P. Adamson et al. [MINOS collab], [arXiv:1710.06488].

- ▶ Anomalies in $\nu_e \rightarrow \nu_e$ disappearance and $\nu_\mu \rightarrow \nu_e$ appearance experiments still exist. Each of these anomalies can be individually explained by sterile neutrinos. But
 - ▶ reactor neutrino data face issues with **flux normalization** and the 5 MeV bump, as well as small tensions in reactor vs gallium and “rates” vs DANSS/NEOS
 - ▶ MiniBooNE data show an excess in low-E neutrino data, which is not so manifest in antineutrino data.
- ▶ No anomaly is found in $\nu_\mu \rightarrow \nu_\mu$ disappearance experiments.
- ▶ Many experiments are in progress or are being prepared. We expect these anomalies will be solved in the near future.