



GGC

Grupo de Gravitação e Cosmologia

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LISHEP2018 - Sessão A
Salvador, 04.09.2018

Matéria Escura

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Instituto de Física
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- **Evidências observacionais**
 - **Astrofísica**
 - **Cosmologia**
- **Candidatos**
- **Experimentos/Sondas**

Qual a estrutura interna da matéria?

Quais interações regem os fenômenos na natureza?

$< 10^{-18}$ metros



$> 10^{26}$ metros

$< 10^{-18}$ metros

$$\frac{1}{1\,000\,000\,000\,000\,000\,000\,000}$$



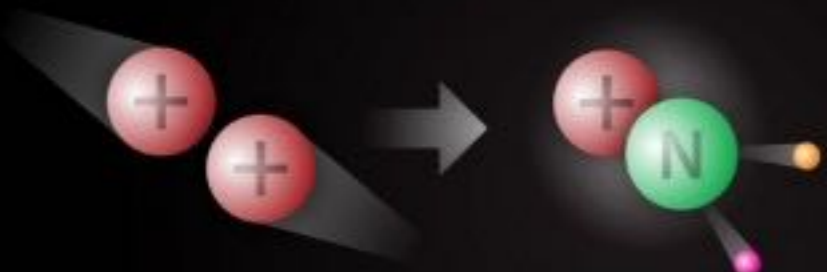
100 000 000 000 000 000 000 000 000 000

$> 10^{26}$ metros

1 H Hydrogen																	2 He Helium																														
3 Li Lithium	4 Be Beryllium											5 B Boron	6 C Carbon	7 N Nitrogen	8 O Oxygen	9 F Fluorine	10 Ne Neon																														
11 Na Sodium	12 Mg Magnesium											13 Al Aluminum	14 Si Silicon	15 P Phosphorus	16 S Sulfur	17 Cl Chlorine	18 Ar Argon																														
19 K Potassium	20 Ca Calcium	21 Sc Scandium	22 Ti Titanium	23 V Vanadium	24 Cr Chromium	25 Mn Manganese	26 Fe Iron	27 Co Cobalt	28 Ni Nickel	29 Cu Copper	30 Zn Zinc	31 Ga Gallium	32 Ge Germanium	33 As Arsenic	34 Se Selenium	35 Br Bromine	36 Kr Krypton																														
37 Rb Rubidium	38 Sr Strontium	39 Y Yttrium	40 Zr Zirconium	41 Nb Niobium	42 Mo Molybdenum	43 Tc Technetium	44 Ru Ruthenium	45 Rh Rhodium	46 Pd Palladium	47 Ag Silver	48 Cd Cadmium	49 In Indium	50 Sn Tin	51 Sb Antimony	52 Te Tellurium	53 I Iodine	54 Xe Xenon																														
55 Cs Cesium	56 Ba Barium	57-71 Lanthanides	72 Hf Hafnium	73 Ta Tantalum	74 W Tungsten	75 Re Rhenium	76 Os Osmium	77 Ir Iridium	78 Pt Platinum	79 Au Gold	80 Hg Mercury	81 Tl Thallium	82 Pb Lead	83 Bi Bismuth	84 Po Polonium	85 At Astatine	86 Rn Radon																														
87 Fr Francium	88 Ra Radium	89-103 Actinides	104 Rf Rutherfordium	105 Db Dubnium	106 Sg Seaborgium	107 Bh Bohrium	108 Hs Hassium	109 Mt Meitnerium	110 Ds Darmstadtium	111 Rg Roentgenium	112 Cn Copernicium	113 Nh Nihonium	114 Fl Flerovium	115 Mc Moscovium	116 Lv Livermorium	117 Ts Tennessine	118 Og Oganesson																														
<table border="1"> <tbody> <tr> <td>57 La Lanthanum</td> <td>58 Ce Cerium</td> <td>59 Pr Praseodymium</td> <td>60 Nd Neodymium</td> <td>61 Pm Promethium</td> <td>62 Sm Samarium</td> <td>63 Eu Europium</td> <td>64 Gd Gadolinium</td> <td>65 Tb Terbium</td> <td>66 Dy Dysprosium</td> <td>67 Ho Holmium</td> <td>68 Er Erbium</td> <td>69 Tm Thulium</td> <td>70 Yb Ytterbium</td> <td>71 Lu Lutetium</td> </tr> <tr> <td>89 Ac Actinium</td> <td>90 Th Thorium</td> <td>91 Pa Protactinium</td> <td>92 U Uranium</td> <td>93 Np Neptunium</td> <td>94 Pu Plutonium</td> <td>95 Am Americium</td> <td>96 Cm Curium</td> <td>97 Bk Berkelium</td> <td>98 Cf Californium</td> <td>99 Es Einsteinium</td> <td>100 Fm Fermium</td> <td>101 Md Mendelevium</td> <td>102 No Nobelium</td> <td>103 Lr Lawrencium</td> </tr> </tbody> </table>																		57 La Lanthanum	58 Ce Cerium	59 Pr Praseodymium	60 Nd Neodymium	61 Pm Promethium	62 Sm Samarium	63 Eu Europium	64 Gd Gadolinium	65 Tb Terbium	66 Dy Dysprosium	67 Ho Holmium	68 Er Erbium	69 Tm Thulium	70 Yb Ytterbium	71 Lu Lutetium	89 Ac Actinium	90 Th Thorium	91 Pa Protactinium	92 U Uranium	93 Np Neptunium	94 Pu Plutonium	95 Am Americium	96 Cm Curium	97 Bk Berkelium	98 Cf Californium	99 Es Einsteinium	100 Fm Fermium	101 Md Mendelevium	102 No Nobelium	103 Lr Lawrencium
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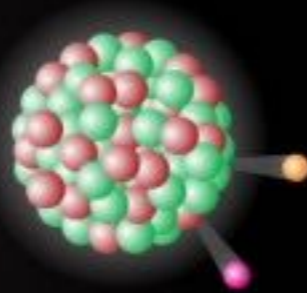
4 Interações Fundamentais

Weak Nuclear Force



Converting protons into neutrons

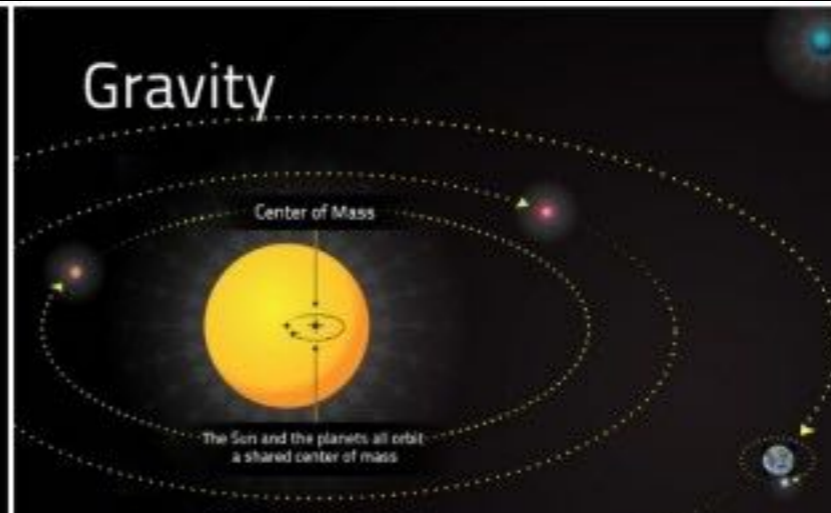
When two protons collide and fuse, a disruption in the weak nuclear force emits a positron and neutrino, which converts one of the positively charged proton to a neutrally charged Neutron. Without the weak nuclear force converting protons into neutrons, certain complex nuclei cannot form.



Releasing radiation

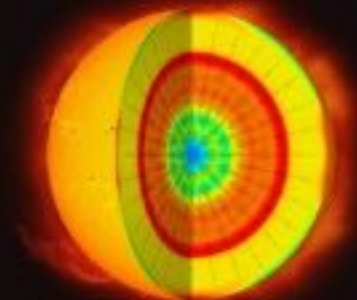
Heavy atoms have an imbalance of protons and neutrons, so the weak nuclear force converts protons to neutrons releasing radiation.

Gravity



Adding motion to the Universe

Gravity forms stars, planets, and moons, and forces these objects to spin on an axis and move along an orbital path. The planets appear to be orbiting the center of the Sun, but the Sun and planets all orbit a shared center of mass. Planets with enough mass can develop orbiting moons or rings of debris.



Creating energy

Gravity is the force that creates pressure and fusion energy in the core of stars allowing them to burn for millions of years.

Electromagnetic Force



Forming atoms and molecules

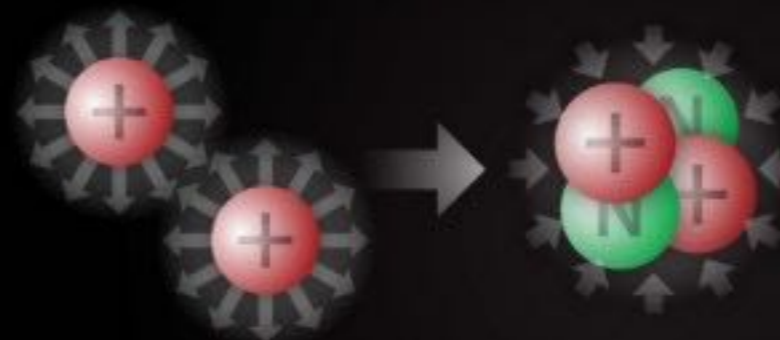
The electromagnetic force pulls negatively charged electrons into bound orbits around positively charged nuclei to form atoms and molecules. As a gas cools, electrons will find their way into the presence of atomic nuclei. Larger nuclei with a greater positive charge pull in more electrons until atoms and molecules have a balance of charges.



Generating light

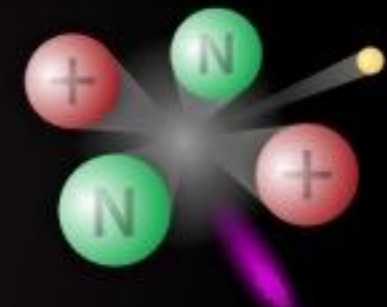
When a negative electron interacts with a positive proton, the electromagnetic force adds energy to the electron generating a photon.

Strong Nuclear Force



Binding protons in atomic nuclei

Positively charged particles naturally repel each other, it takes an extreme amount of force to hold protons together. The strong nuclear force overcomes the repulsion between protons to hold together atomic nuclei. Without the strong nuclear force, complex nuclei cannot form.

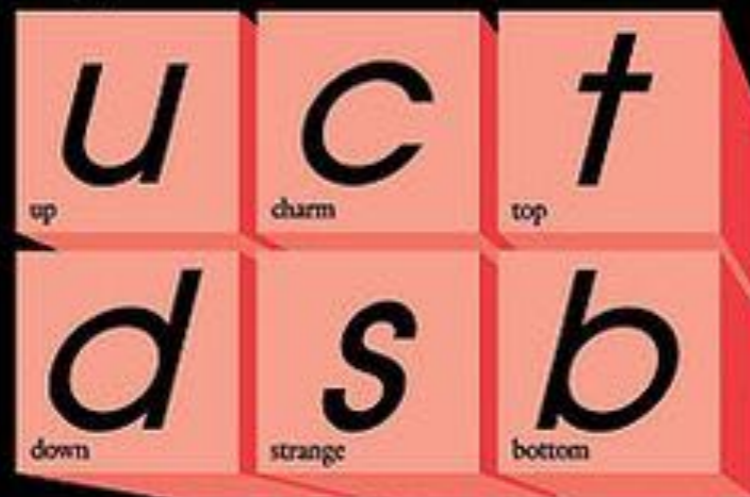


Breaking the bond

Enormous energy is released as gamma rays and neutrinos when the strong nuclear force is broken between protons and neutrons.

Modelo Padrão da Física de Partículas

Quarks



Forces









Leptons



Zoom até o aglomerado de galáxias Abell 315

Crédito: ESO/ Digitized Sky Survey 2

The Origin of the Solar System Elements

1 H	big bang fusion 					cosmic ray fission 					2 He						
3 Li	4 Be	merging neutron stars 					exploding massive stars 					5 B	6 C	7 N	8 O	9 F	10 Ne
11 Na	12 Mg	dying low mass stars 					exploding white dwarfs 					13 Al	14 Si	15 P	16 S	17 Cl	18 Ar
19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr
37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe
55 Cs	56 Ba		72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn
87 Fr	88 Ra																
			57 La	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu
			89 Ac	90 Th	91 Pa	92 U											

Crédito: Jennifer A. Johnson/The Ohio State University; NASA; ESA

A visualization of the cosmic web, showing a complex network of dark matter filaments and nodes. The filaments are represented by thin, dark lines, and the nodes are represented by clusters of bright, yellowish-orange points. The background is a dark, textured surface with a mottled appearance.

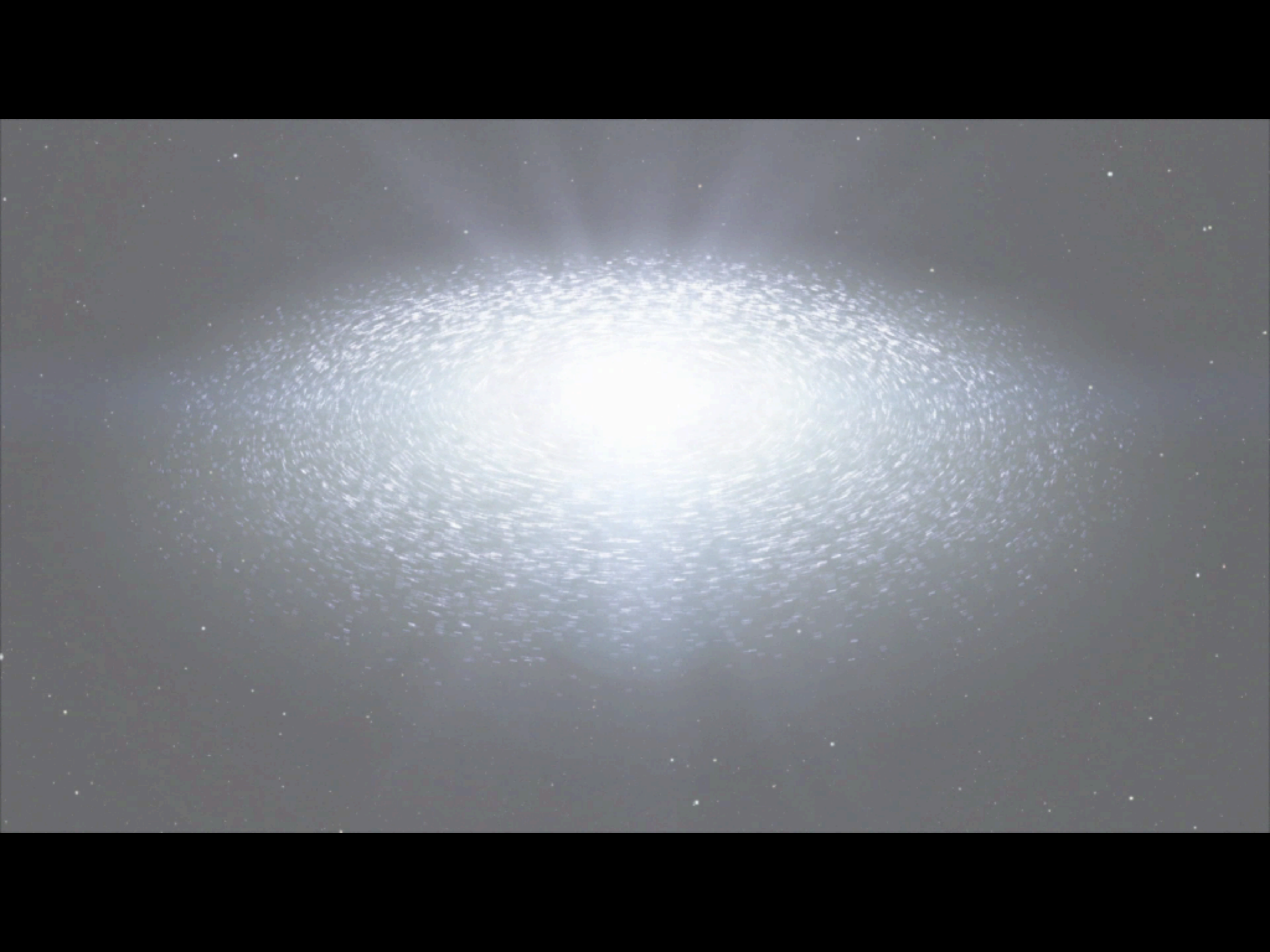
Evidências Observacionais de Matéria Escura

The background of the slide is a dark, textured field of stars, likely representing a star field or a galaxy's core. The stars are small, bright points of light, some appearing as short trails, set against a dark, grainy background. The overall color palette is dark, with shades of black, grey, and blue, punctuated by the white and yellowish-gold of the stars.

Evidências Observacionais de Matéria Escura:

As galáxias espirais e suas rotações





Década de 1970

Vera Rubin

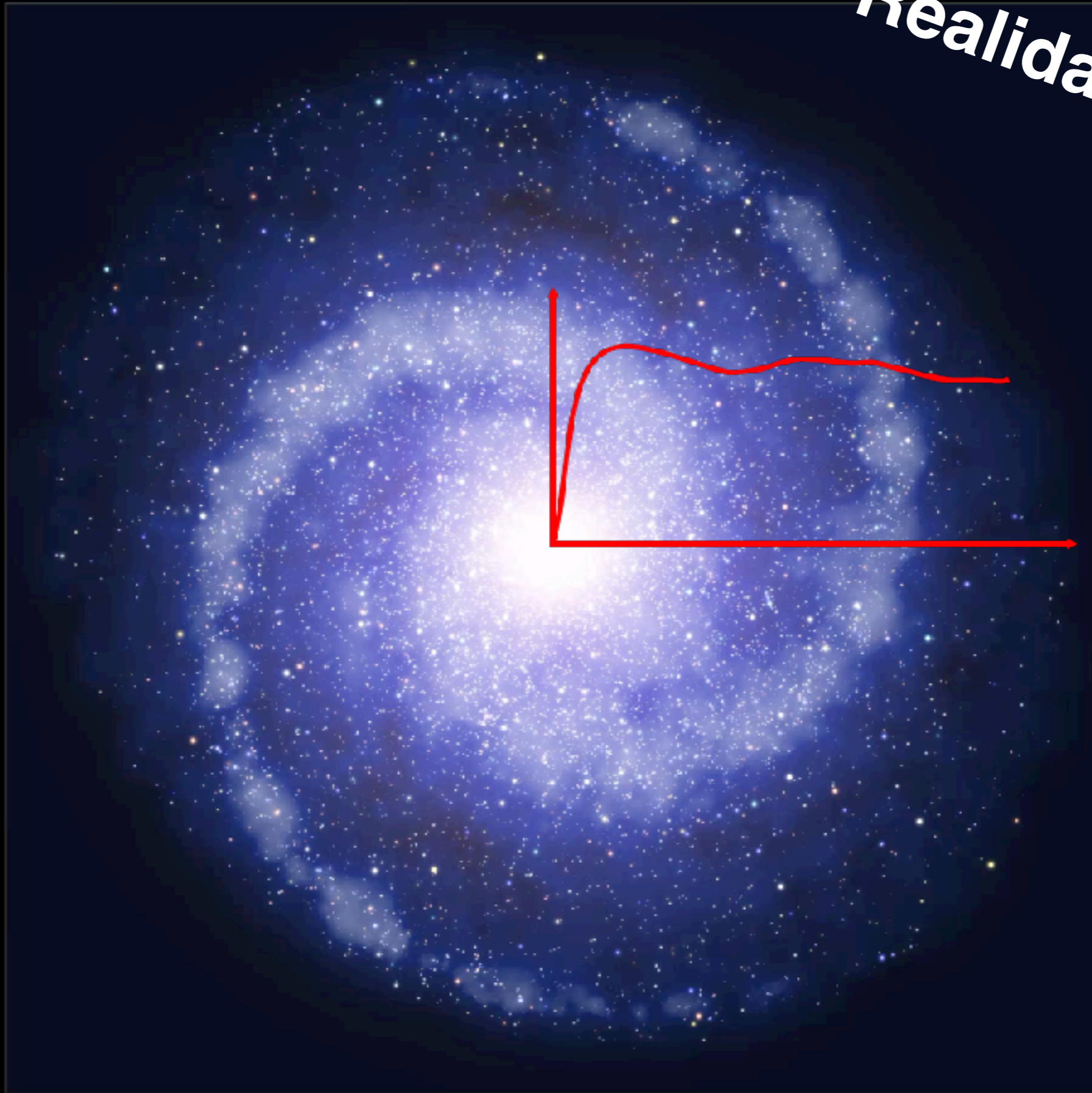
Astrônoma norte-americana



Expectativa



Realidade



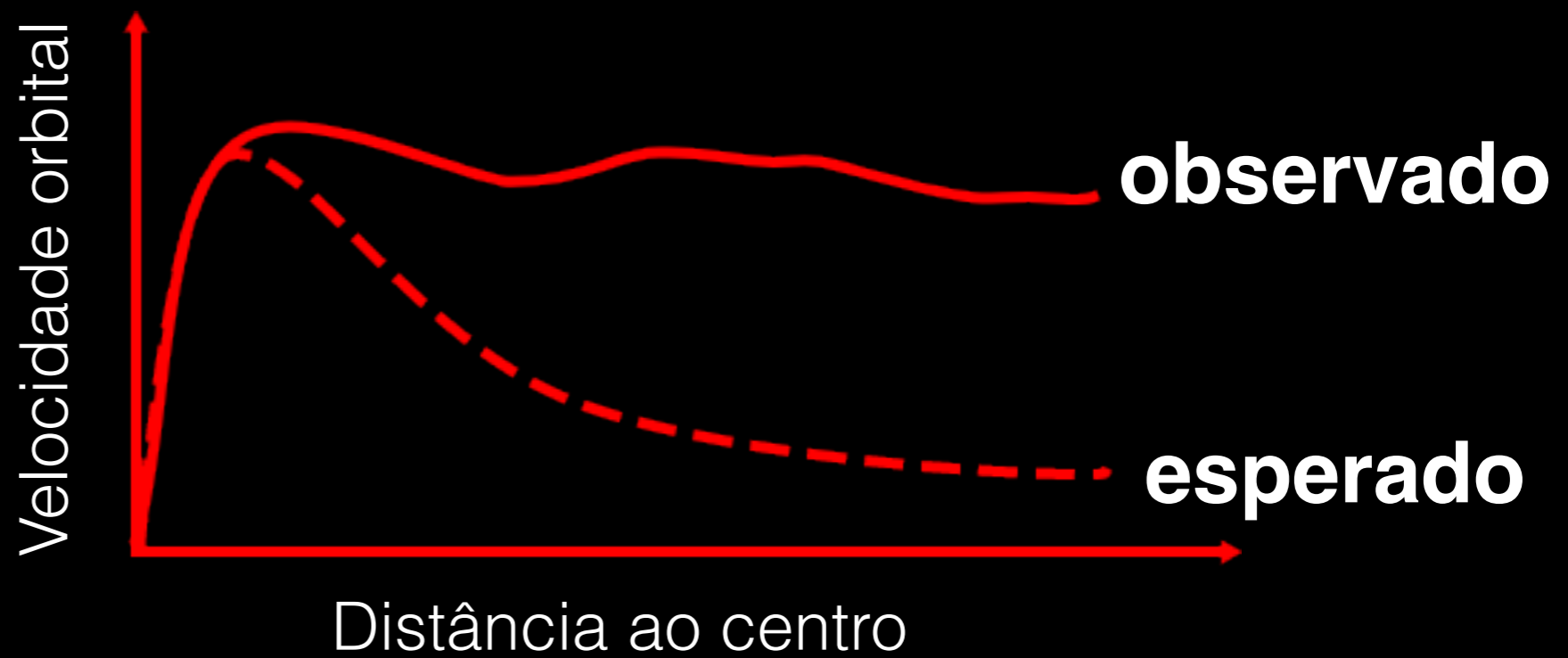
Realidade



Expectativa



$$v = \sqrt{\frac{GM(R)}{R}}$$



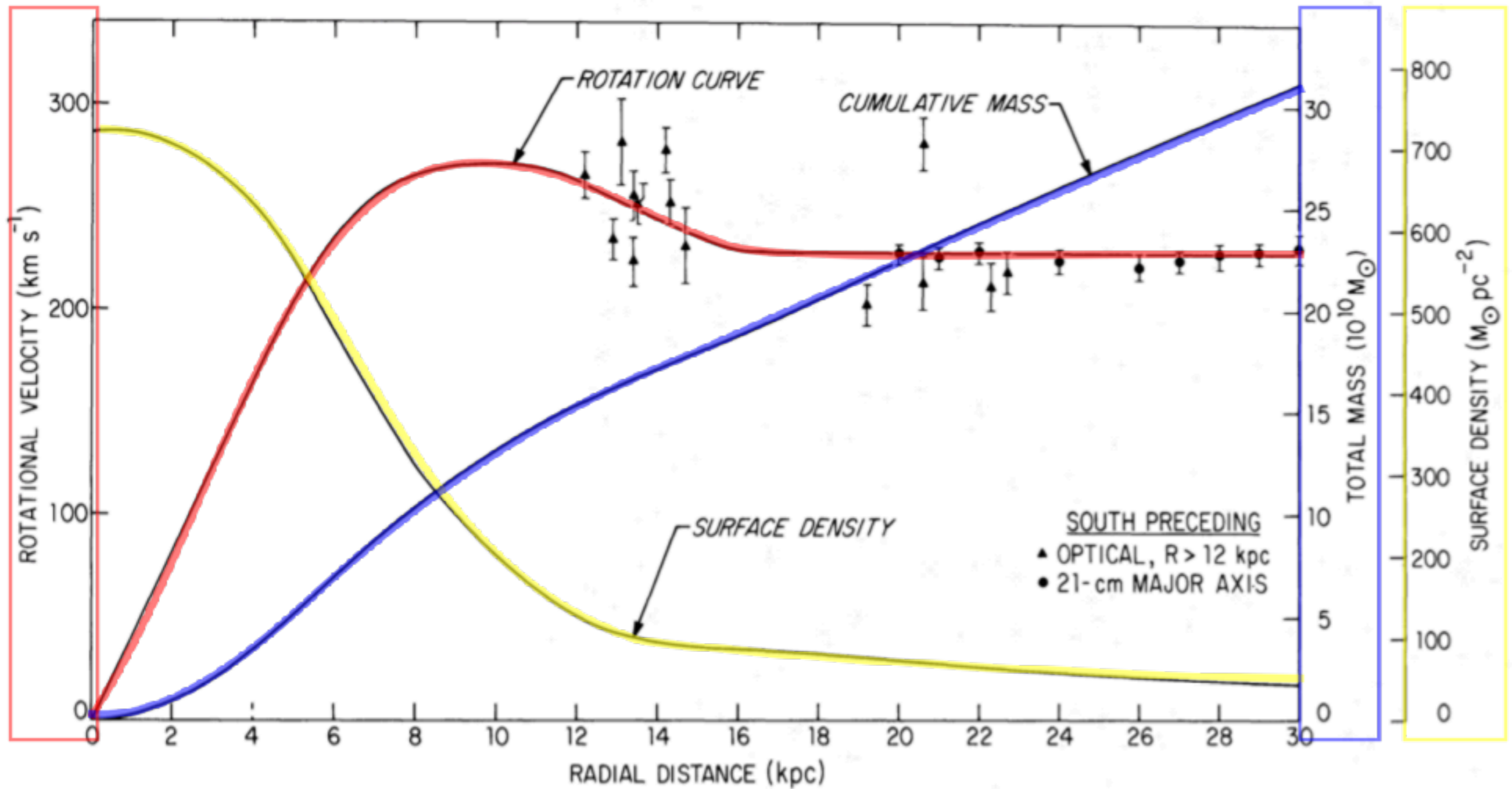
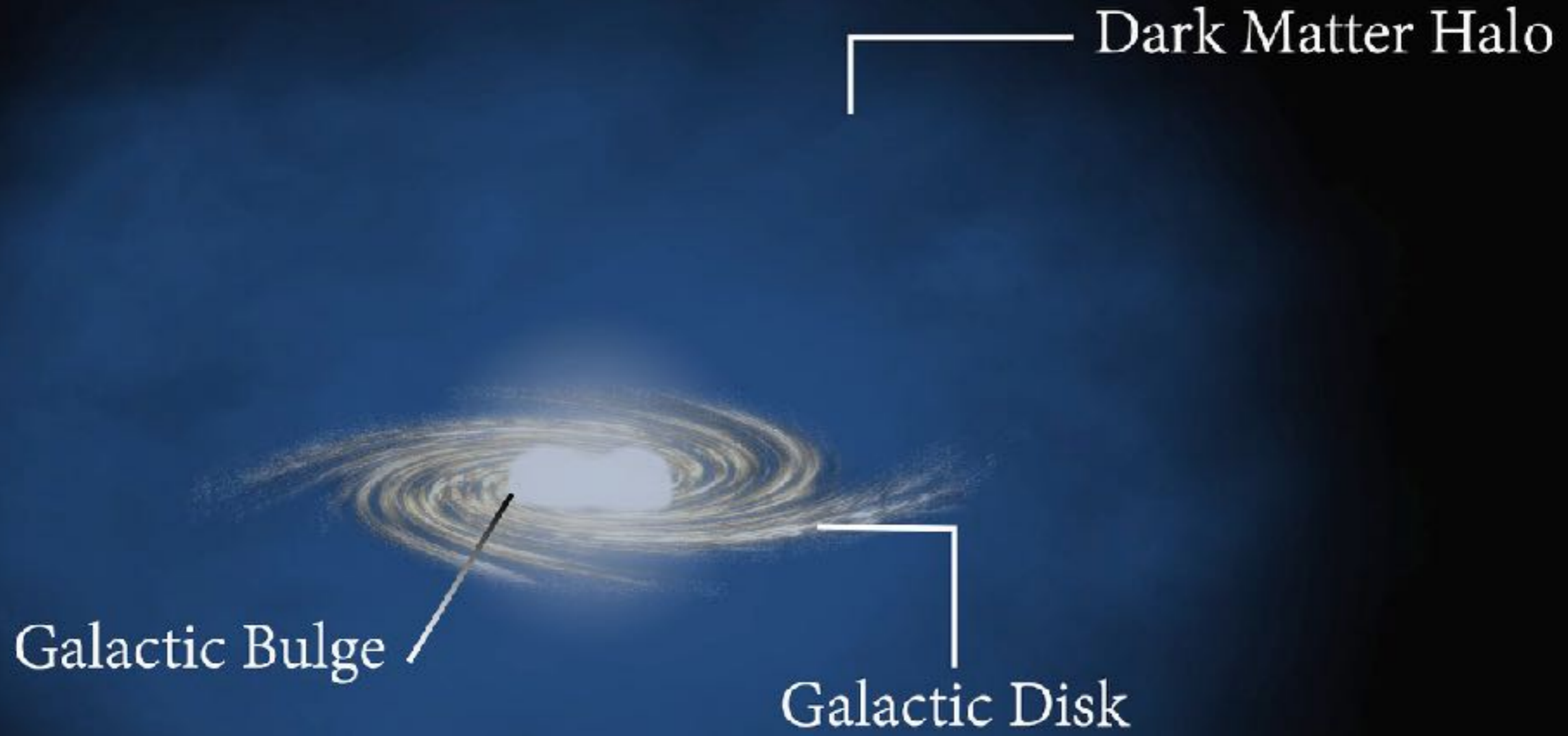


FIG. 16.—The adopted rotation curve, a composite of optical (Rubin and Ford 1970) data and 21-cm major axis measurements. The surface density and cumulative mass curves are for a highly flattened model.

The rotation curve and geometry of M31 at large galactocentric distances.
 Roberts, M. S.; Whitehurst, R. N. Astrophysical Journal, Vol. 201, p.327 - 10/1975



The background of the slide is a Cosmic Microwave Background (CMB) fluctuation map, showing a complex pattern of dark and light regions representing temperature variations in the early universe. The text is overlaid on this background.

Evidências Observacionais de Matéria Escura:

Aglomerados de galáxias

Aglomerado de galáxias Coma

Cerca de 10 000 galáxias,
cada uma com bilhões de estrelas



Teorema de Virial: estado estacionário

$$2E_K + E_U = 0$$

$$M_{Coma} = \frac{\langle v^2 \rangle r_h}{\alpha G}$$

~300 vezes mais massa que o observado



Fritz Zwicky (1898 - 1974).

Fritz Zwicky

1933: calculou a massa total do aglomerado de galáxias Coma baseado no movimento das galáxias que se situavam próximas à sua borda.

The background of the slide is a map of the Cosmic Microwave Background (CMB) radiation, showing a complex pattern of dark and light spots representing temperature fluctuations in the early universe. The text is overlaid on this background.

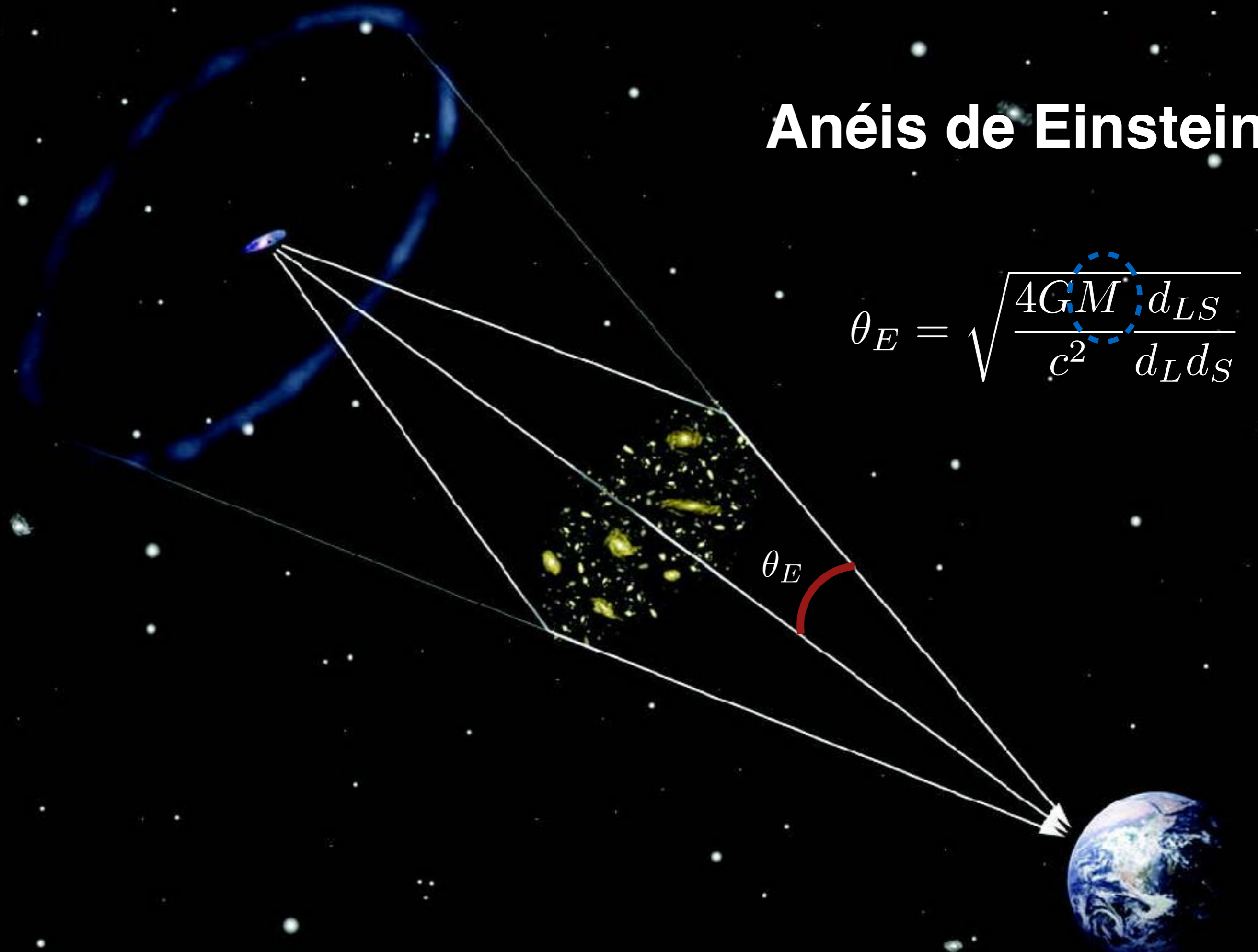
Evidências Observacionais de Matéria Escura:

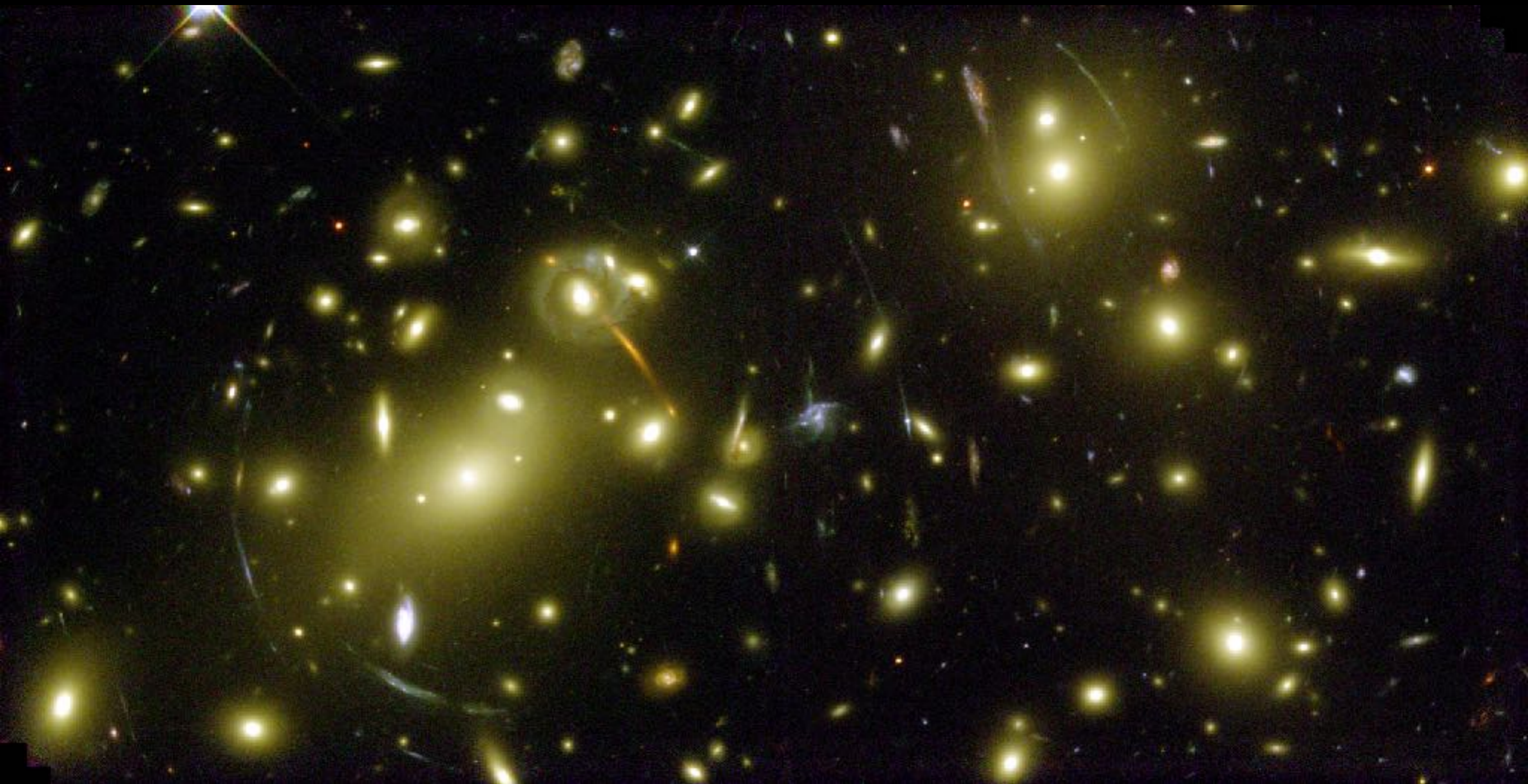
Lentes gravitacionais



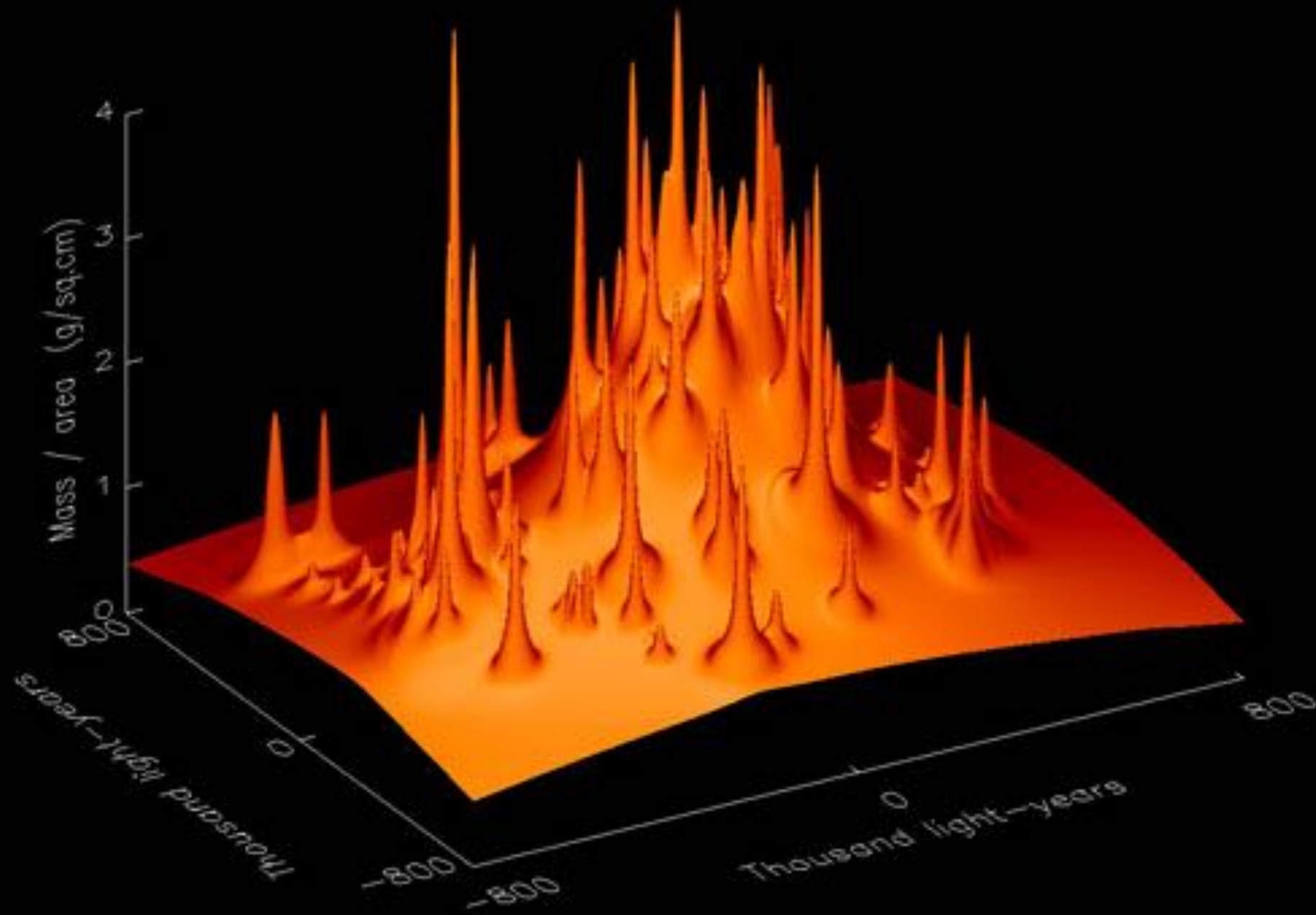
Anéis de Einstein

$$\theta_E = \sqrt{\frac{4GM}{c^2} \frac{d_{LS}}{d_L d_S}}$$





Crédito: NASA, Andrew Fruchter and the ERO Team



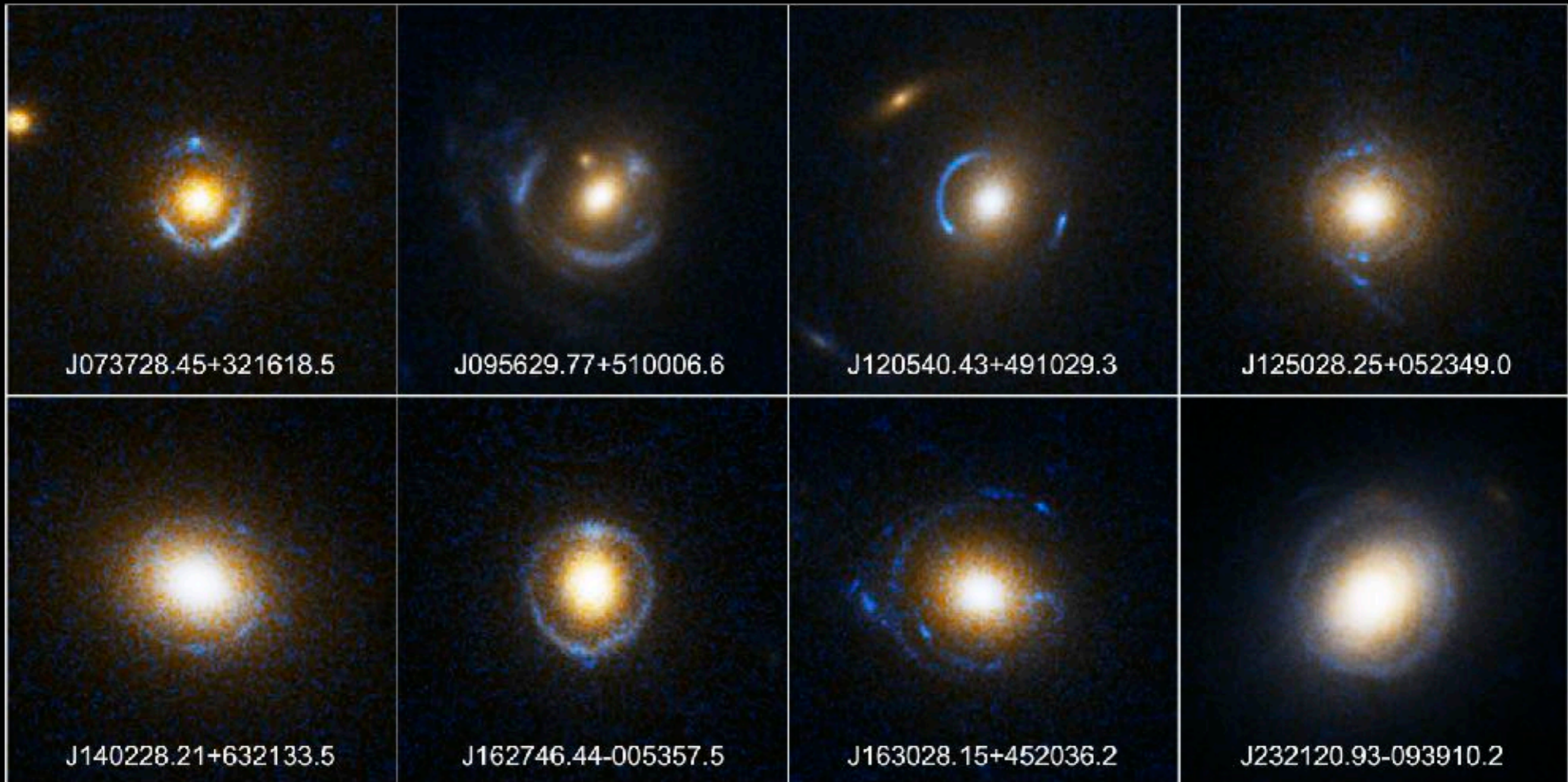
Mapeamento de distribuição de massa em aglomerado de galáxias via lentes gravitacionais.



Mapas de matéria escura em aglomerados de galáxias.

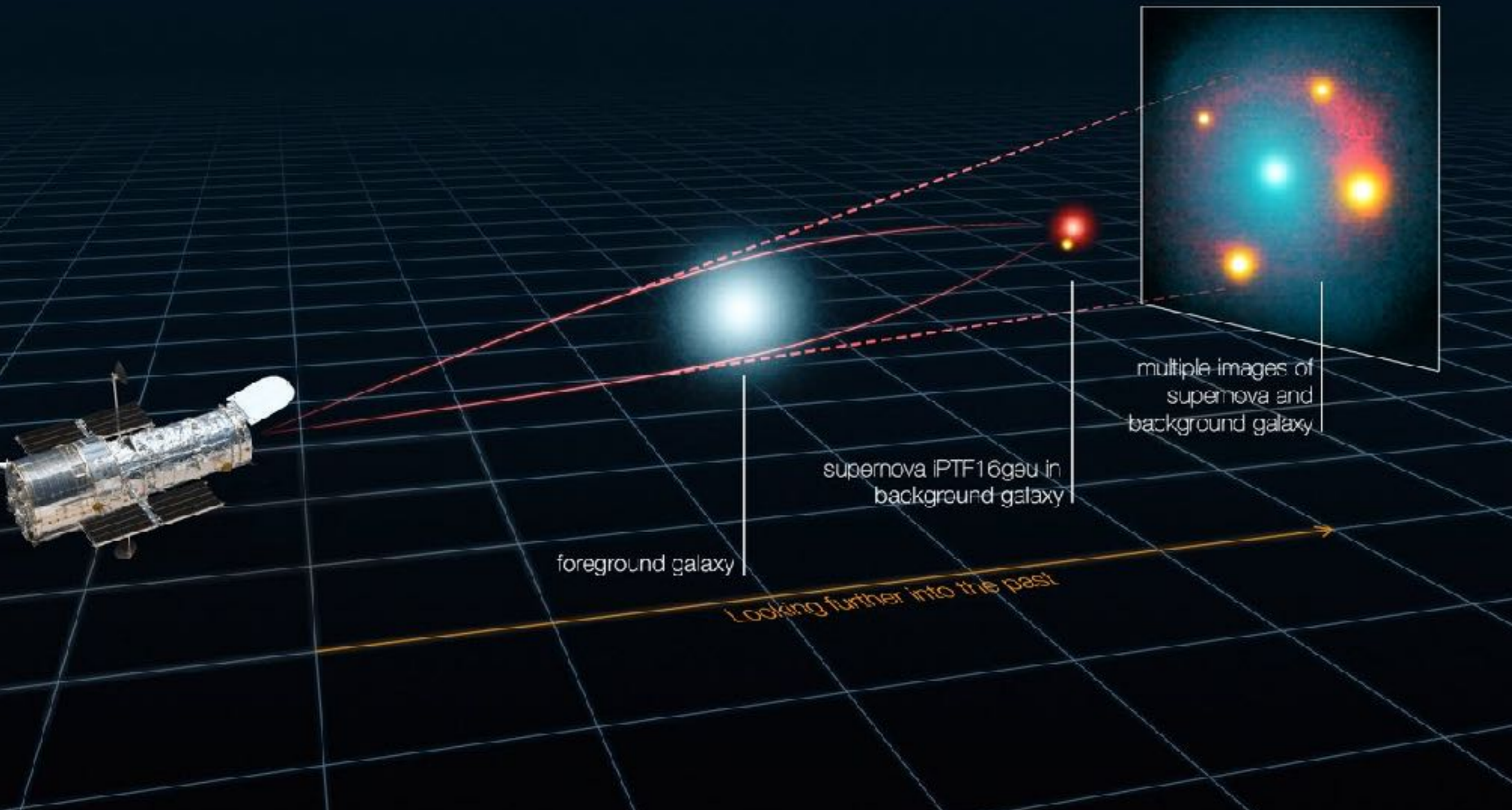
Credit: NASA, ESA, D. Harvey, R. Massey

Lenteamento gravitacional por Matéria Escura em Galáxias Elípticas



Einstein Ring Gravitational Lenses
Hubble Space Telescope • Advanced Camera for Surveys

Imagem da Supernova sob efeito de lente gravitacional



Original image by ALMA (ESO/NRAO/NAOJ), L. Calçada (ESO), Y. Hezaveh et al., edited and modified by Joel Johansson.

The background of the slide is a dark, textured field of stars, likely representing a star cluster or galaxy. The stars are small, bright points of light in various colors, including yellow, orange, and white, scattered across a dark, almost black background. The overall appearance is that of a deep space or astronomical image.

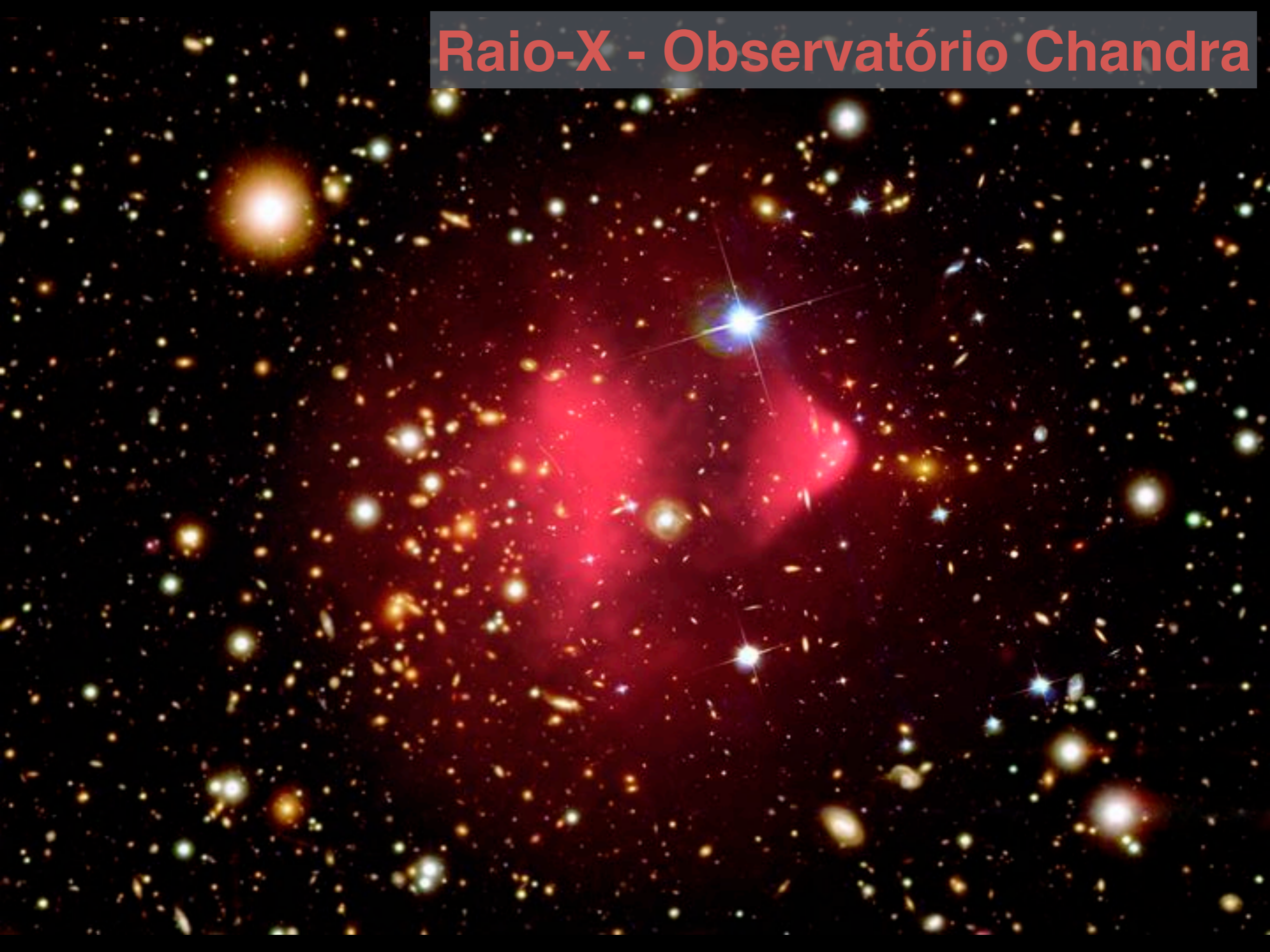
Evidências Observacionais de Matéria Escura:

Aglomerado da Bala (Bullet Cluster)

Luz visível - HST



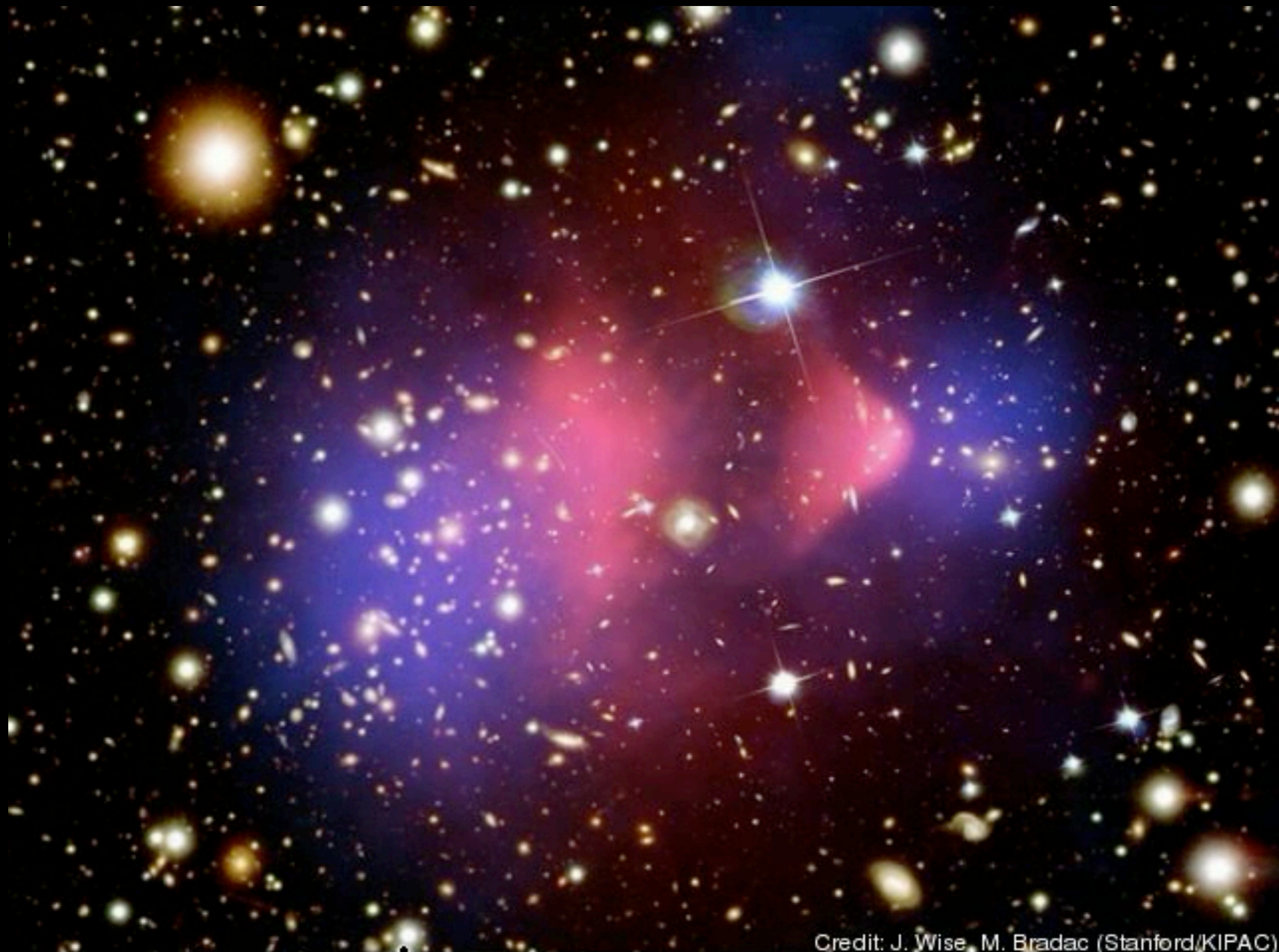
Raio-X - Observatório Chandra



Mapeamento de matéria escura: Lentes Gravitacionais







Credit: J. Wise, M. Bradac (Stanford/KIPAC)



Evidências Observacionais de Matéria Escura:

**Galáxia ultra difusa
NGC 1052-DF2**



Zoom através da constelação Cetus (Baleia) até observação da galáxia ultra difusa NGC 1052-DF2.

Crédito: ESA/Hubble, Digitized Sky Survey, Nick Risinger

(março 2018)

NGC 1052-DF2

~ tamanho da Via Láctea

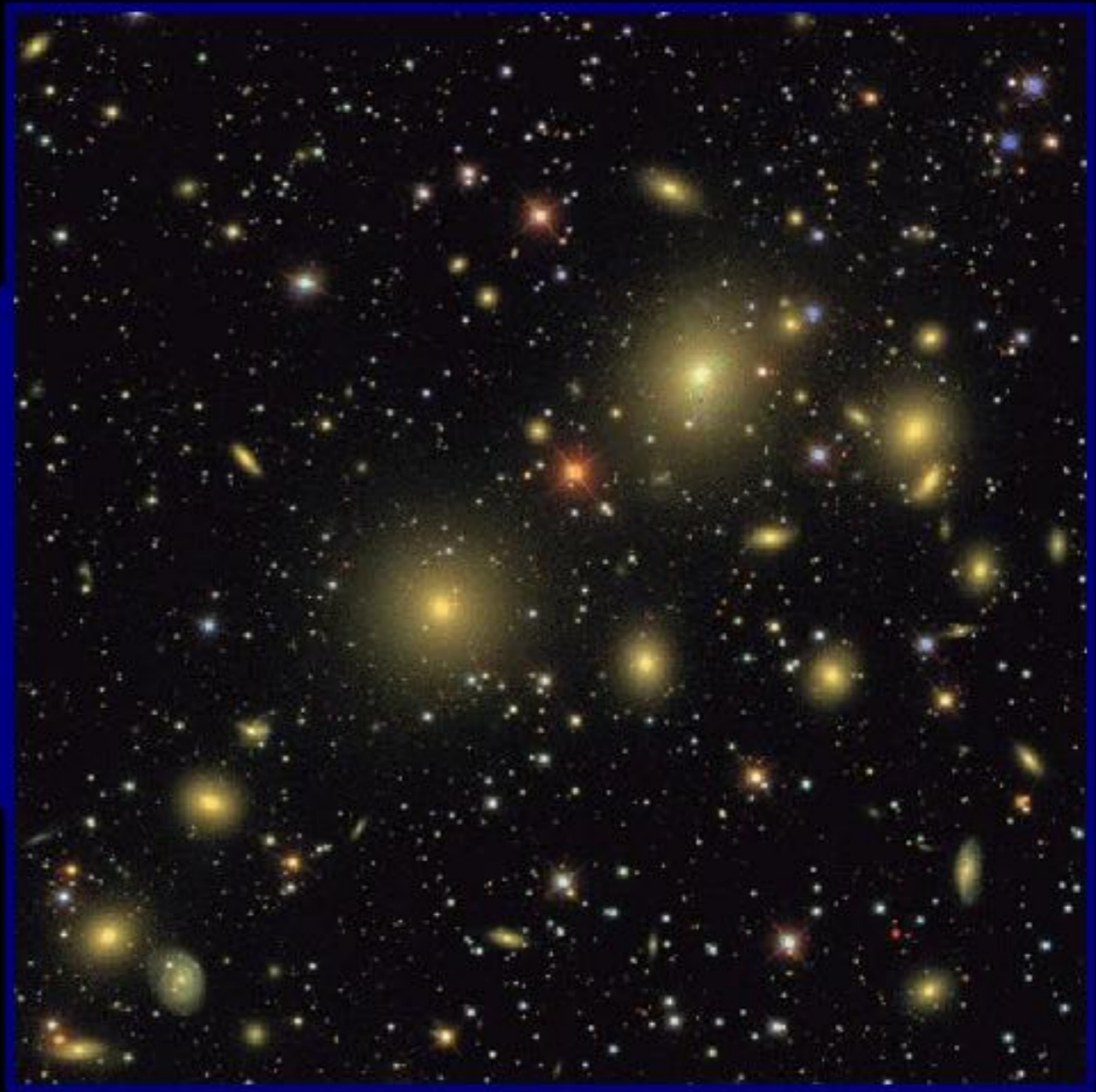
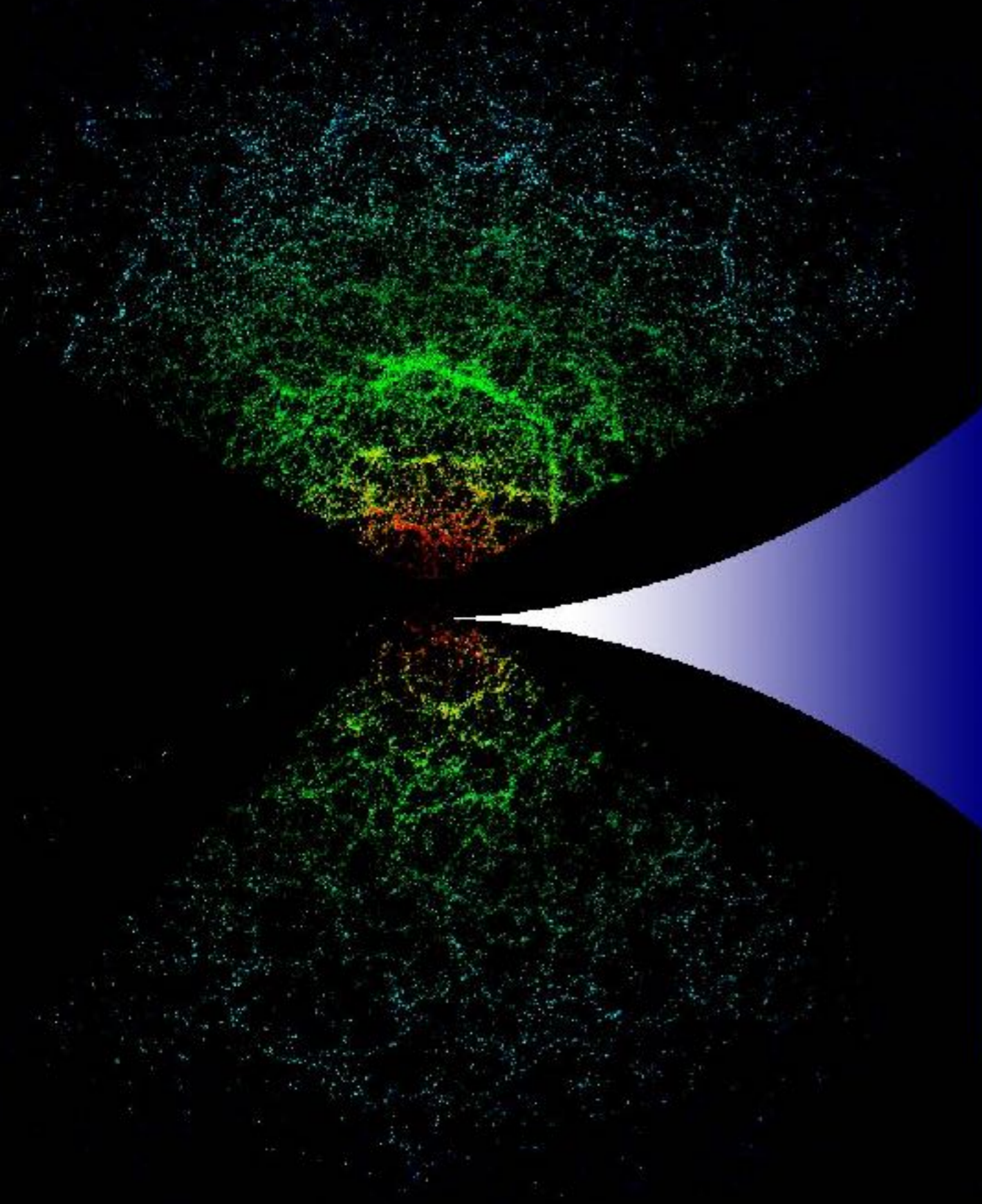
0,5% do número de estrelas da Via Láctea

1/400 da quantidade de matéria escura esperada!!

The background of the slide is a Cosmic Microwave Background (CMB) fluctuation map, showing a complex pattern of dark and light spots representing temperature variations in the early universe. The text is overlaid on this background.

Evidências Observacionais De Matéria Escura:

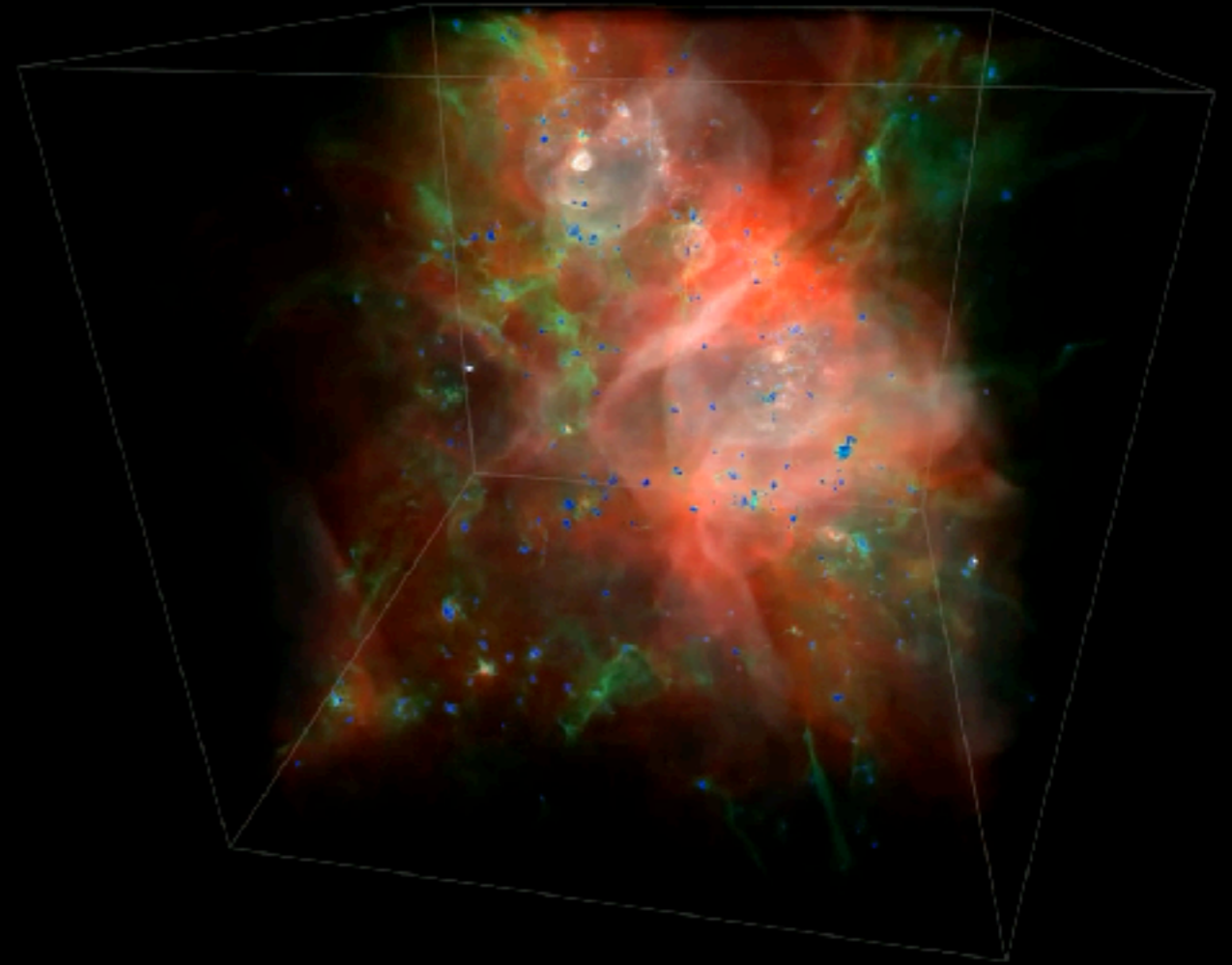
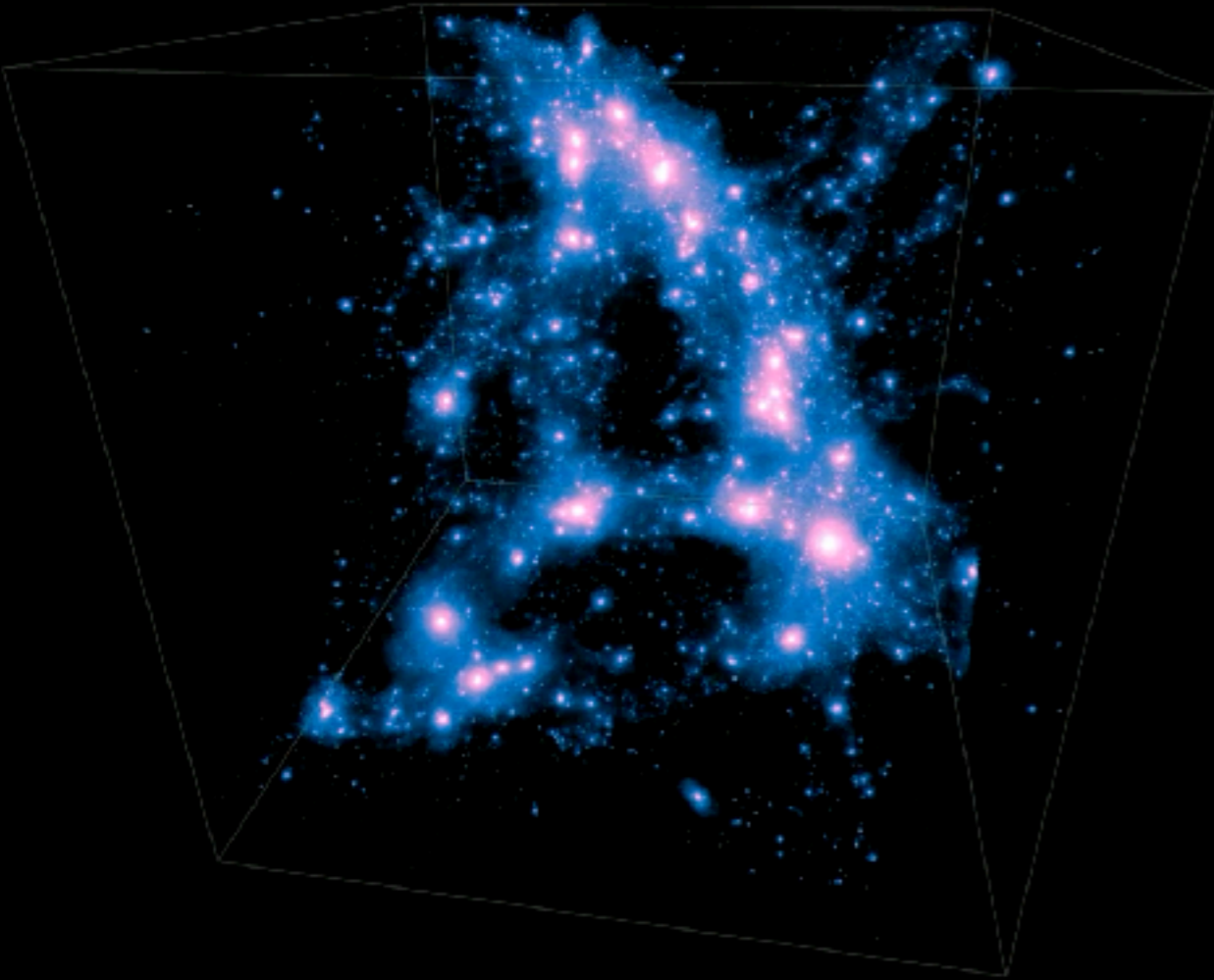
Formação de estruturas em larga escala



66 976 galáxias do SDSS

Dark Matter

Gas Temperature



redshift : 0.30
Time since the Big Bang: 10.4 billion years

stellar mass : 72.4 billion solar masses

ILLUSTRIS



Crédit : Illustris Project / Illustris Project

<http://www.illustris-project.org/media/>

Foto do Telescópio Hubble



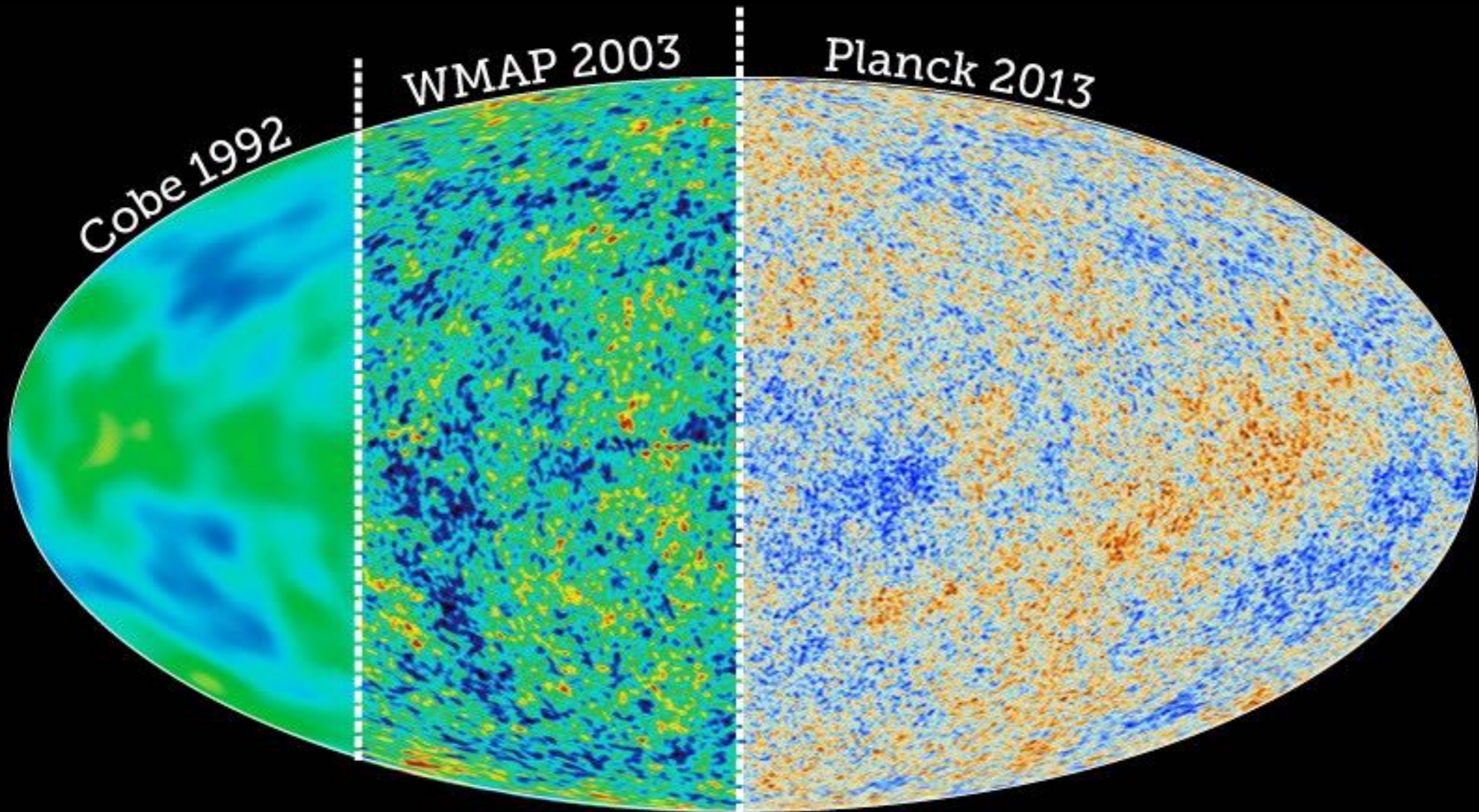
Simulação do Illustris



The background of the slide is a Cosmic Microwave Background (CMB) radiation map, showing a complex pattern of dark and light spots representing temperature fluctuations in the early universe. The text is overlaid on this map.

Evidências Observacionais De Matéria Escura:

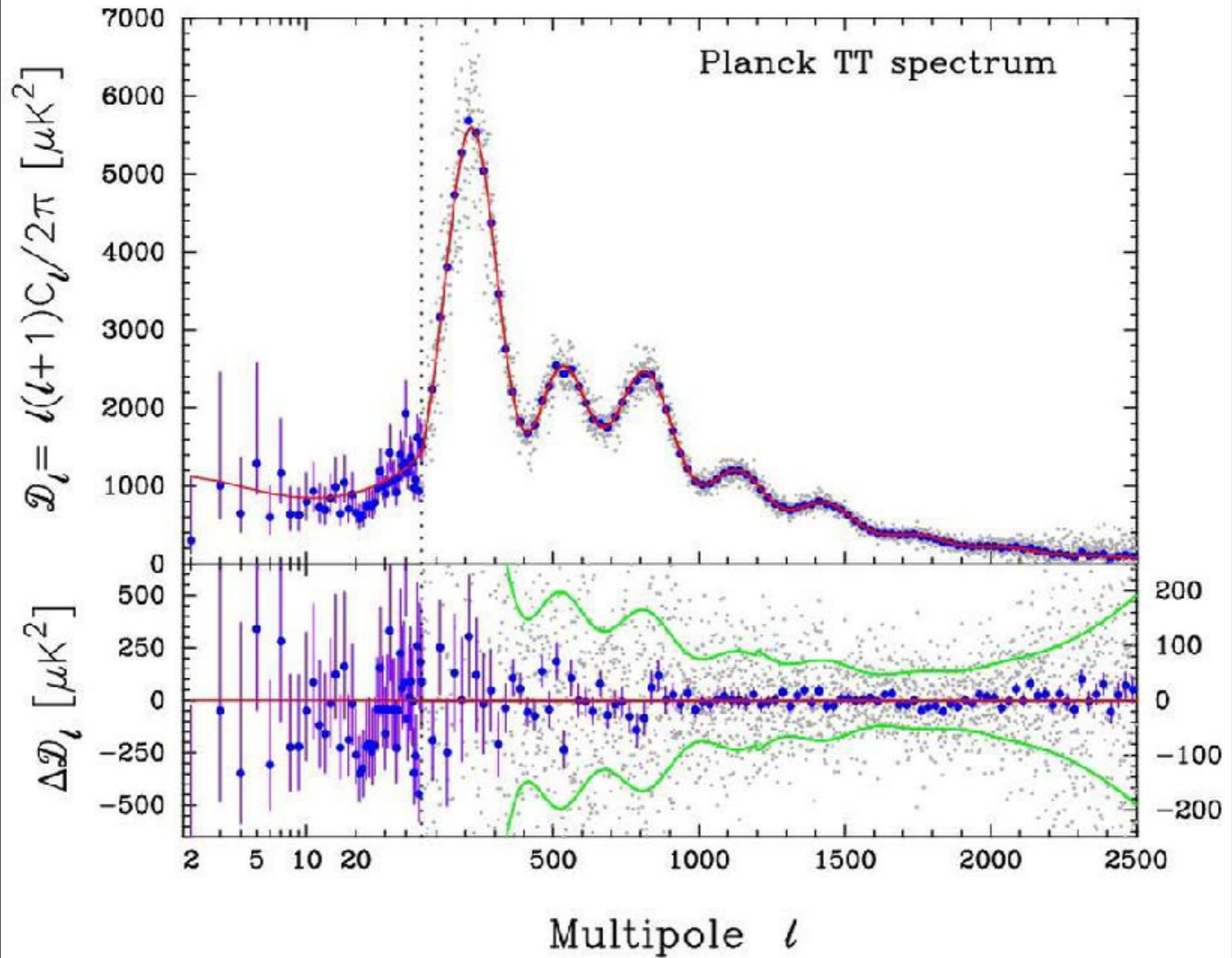
Radiação Cósmica de Fundo em Microondas



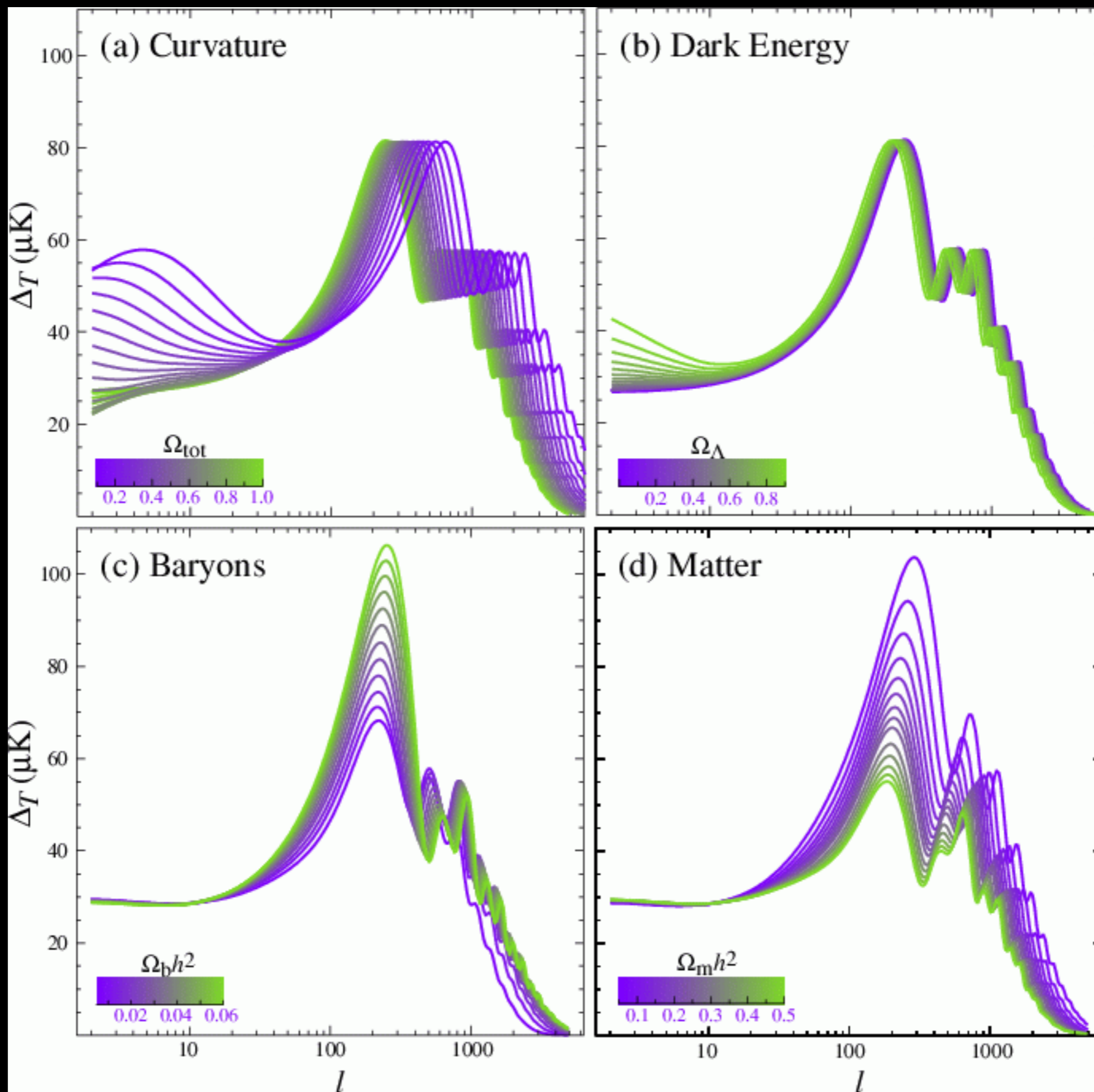
Cobe 1992

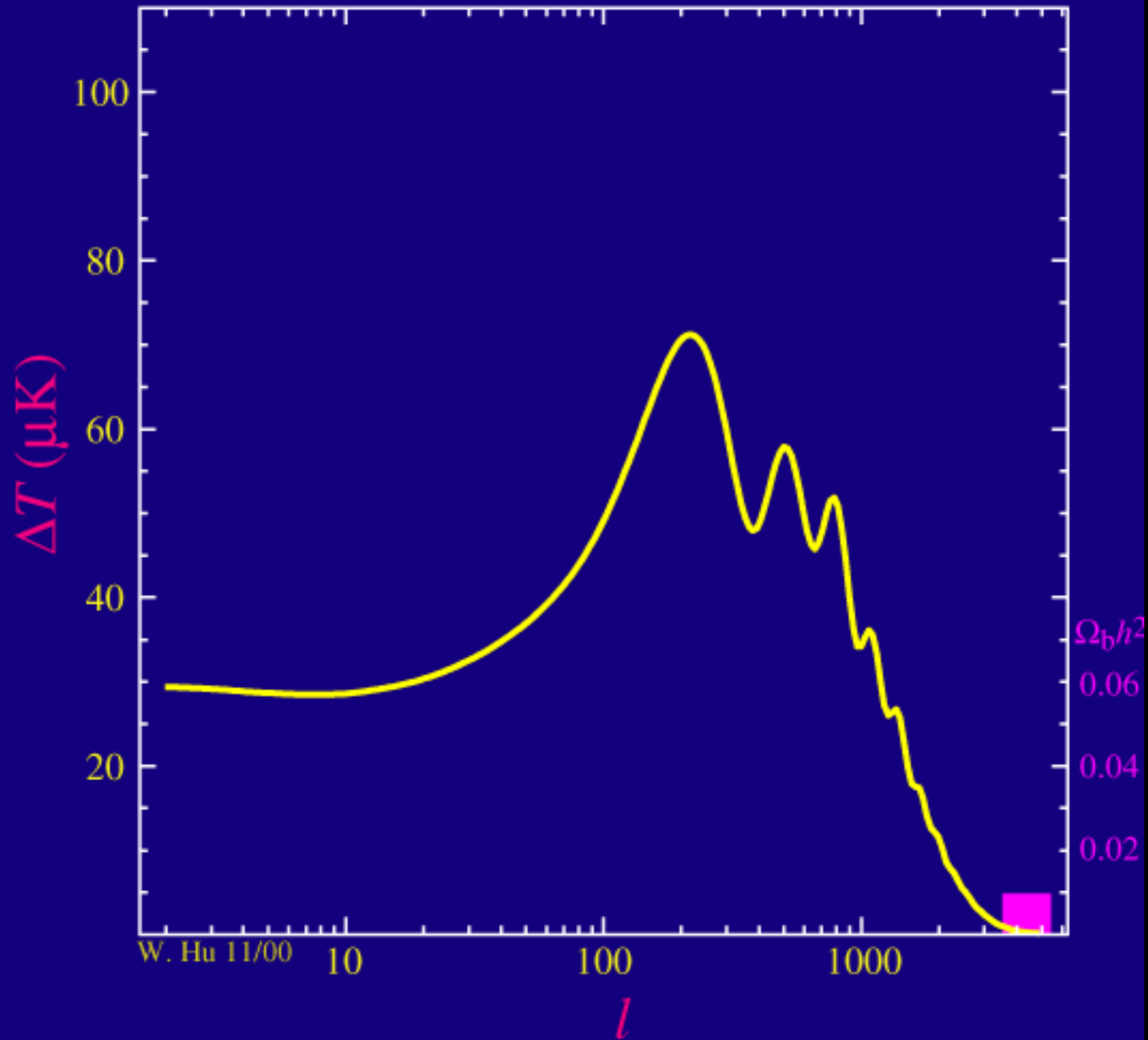
WMAP 2003

Planck 2013

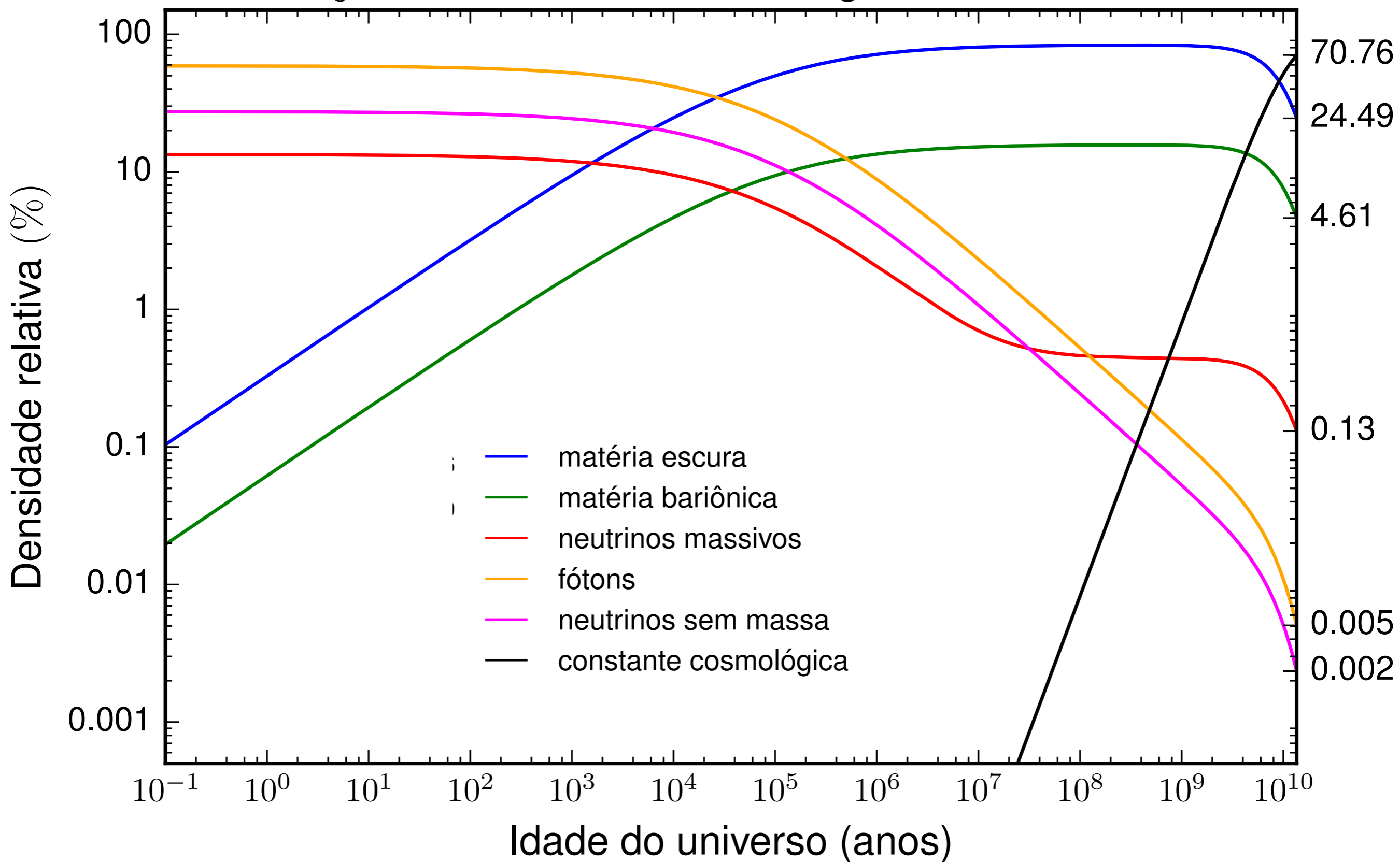


Crédito: Planck Collaboration: P. A. R. Ade et al., 2014, A&A.





Evolução das densidades de energia no Modelo Padrão

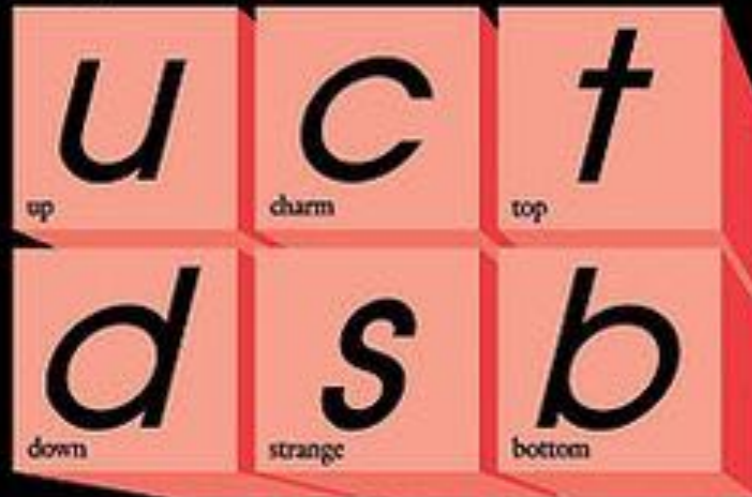


A visualization of the cosmic web, showing a complex network of dark matter filaments and nodes. The filaments are represented by thin, dark lines, and the nodes are represented by clusters of small, bright yellow and orange dots. The background is a dark, textured grey, suggesting the overall structure of the universe.

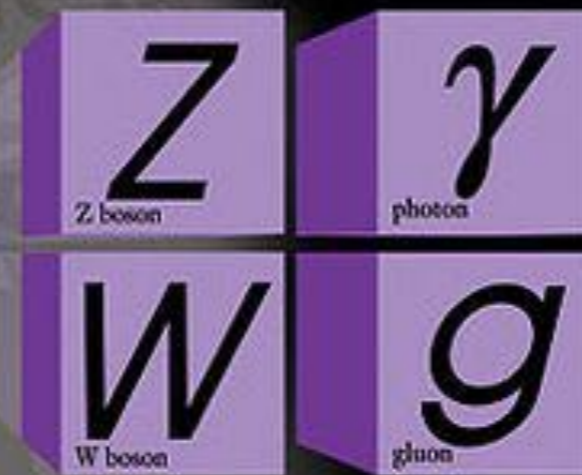
Natureza da Matéria Escura

É composta pela matéria ordinária?

Quarks



Forces



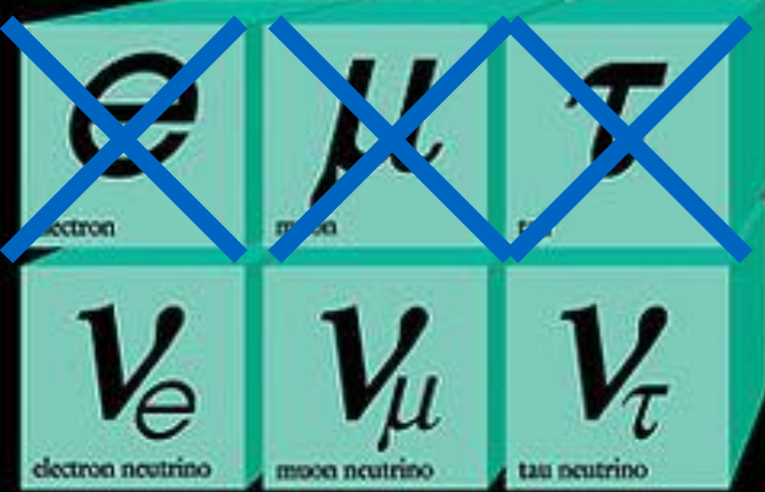
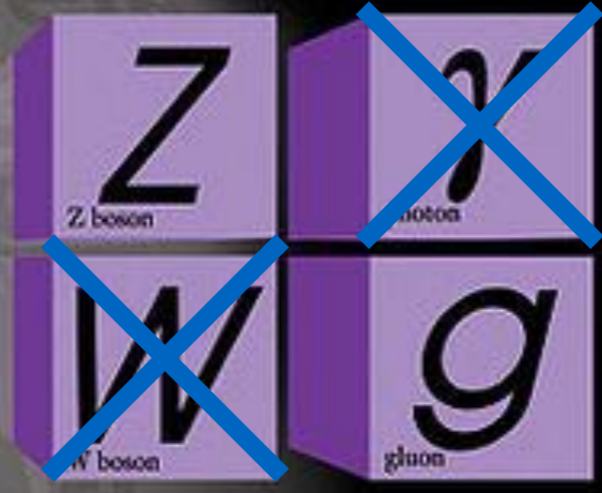
Leptons

Força eletromagnética

Quarks

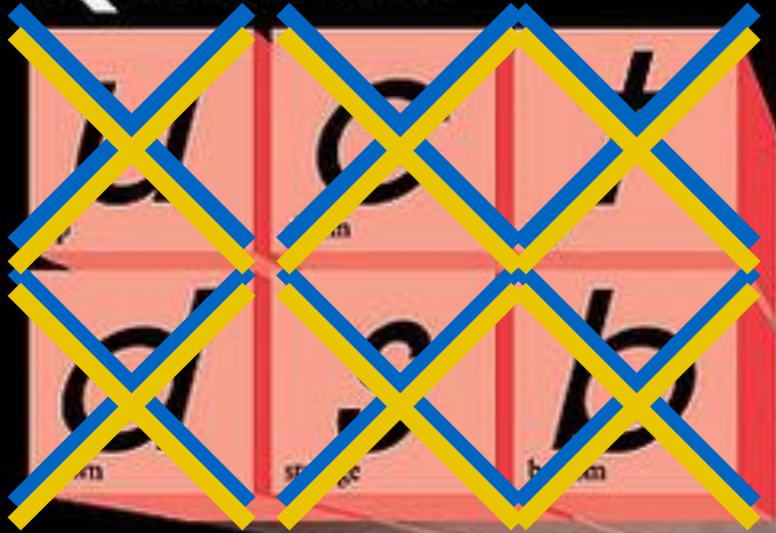


Forces

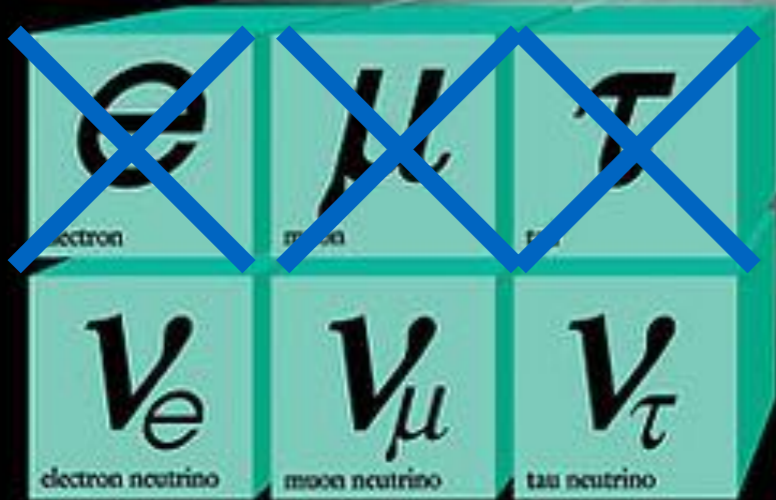


Leptons

Quarks

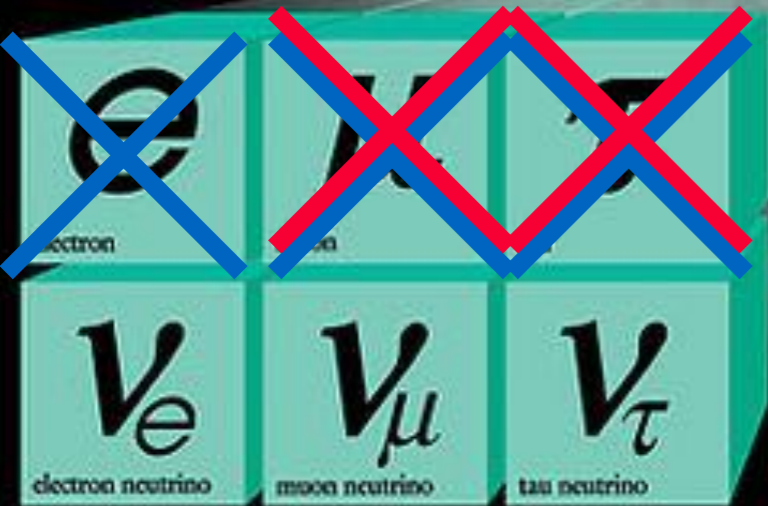


Forces



Leptons

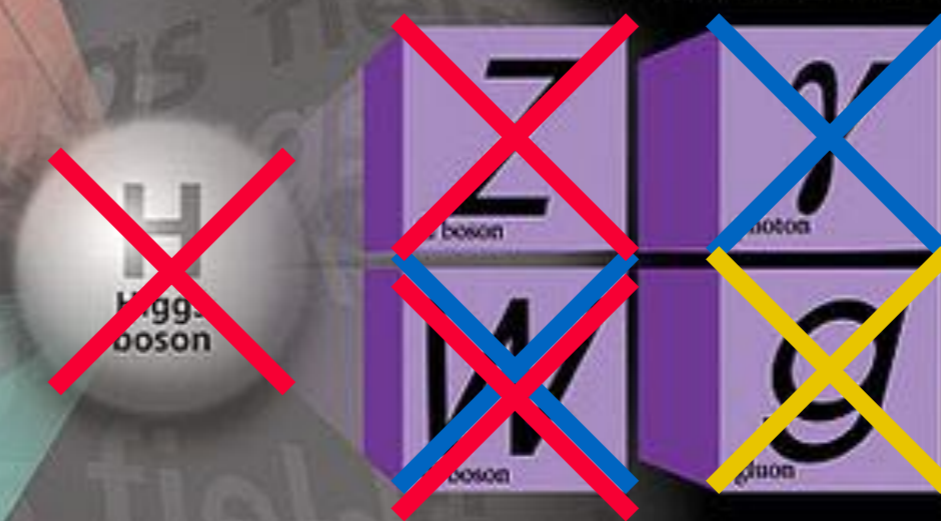
Quarks



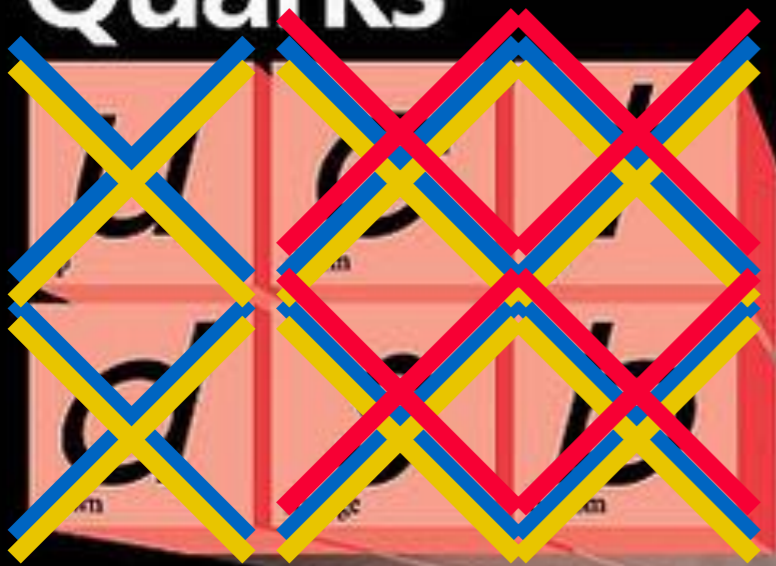
Leptons

- Força eletromagnética
- Força forte
- Instável

Forces

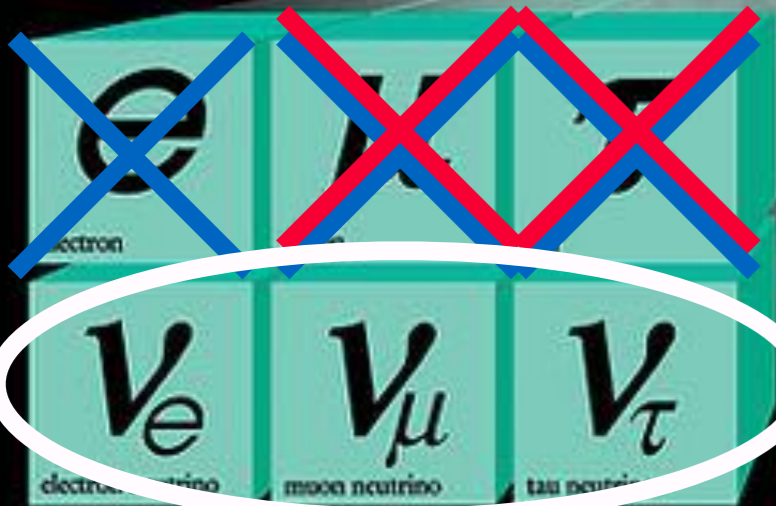
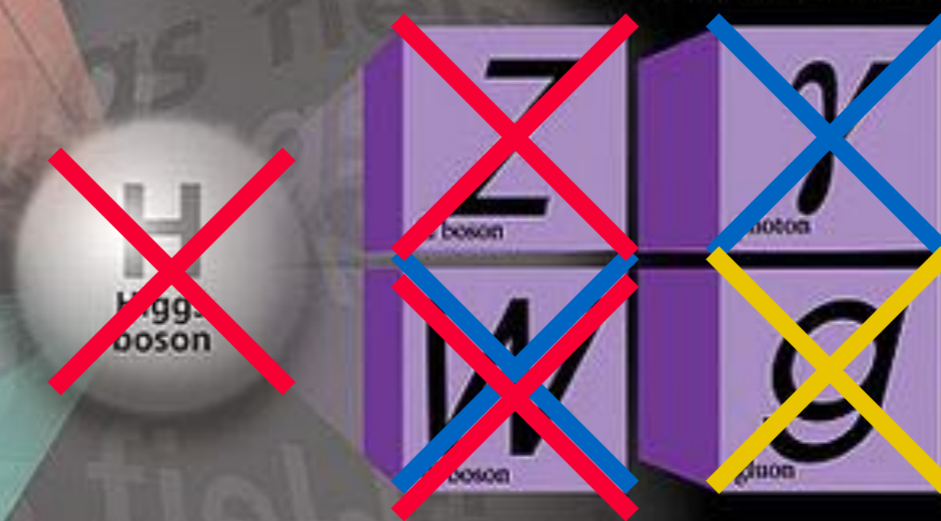


Quarks



- Força eletromagnética
- Força forte
- Instável

Forces



Leptons

muito rápidos
muito "quentes"

Antimatéria

Neutrinos

Gás

MACHOs

“MAssive Compact Halo Objects”

**Anãs marrons, anãs brancas,
estrelas de nêutrons e buracos negros.**

Áxions

$m_{\max} \sim 2 \times 10^{-41} \text{ kg}$

50 bilhões de áxions teriam
massa equivalente a 1 elétron

Buracos negros primordiais

$m_{\max} \sim 2 \times 10^{+35} \text{ kg}$

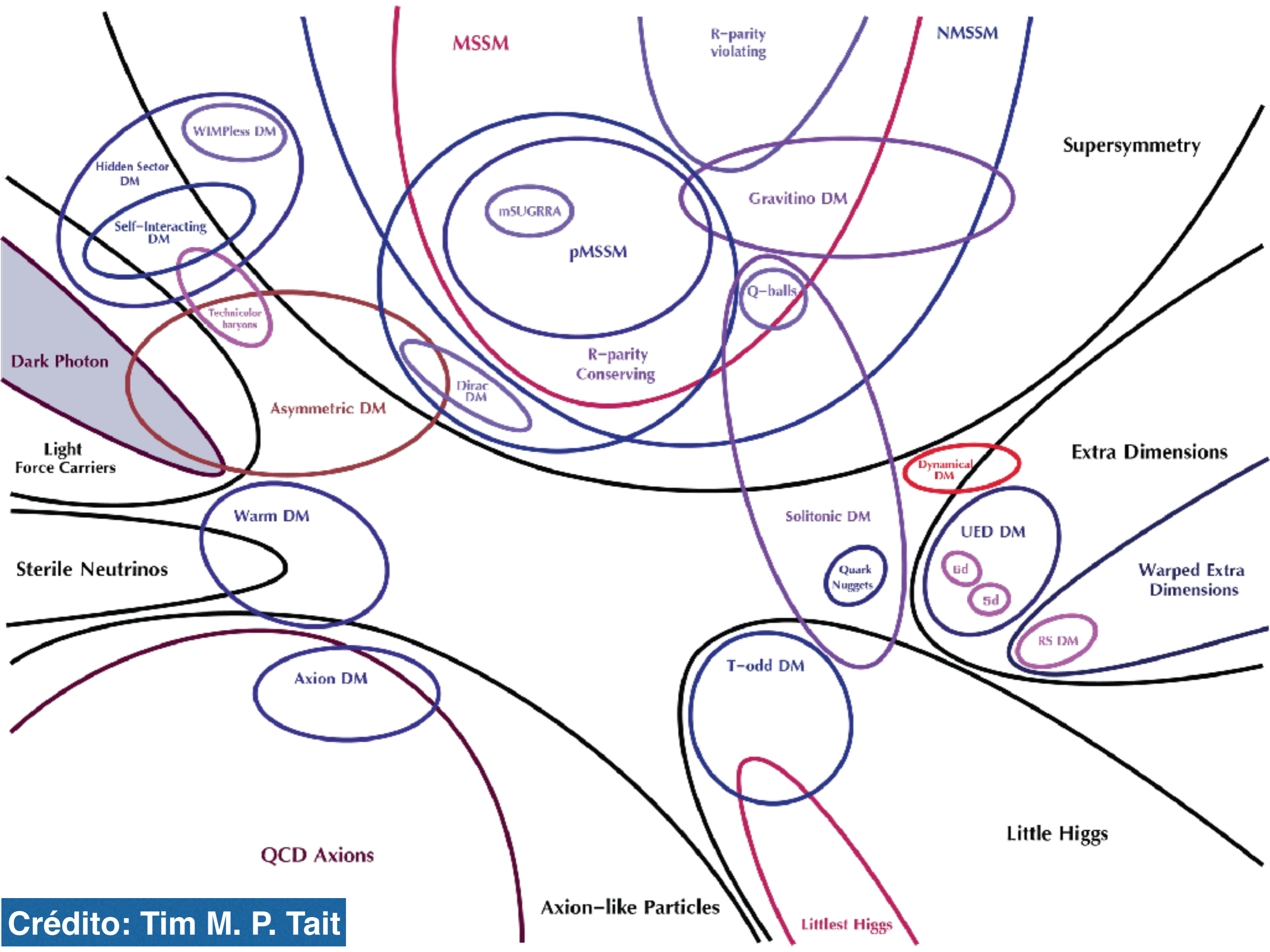
30 bilhões de Terras teriam
massa equivalente a 1 BH primordial

Partículas Supersimétricas

photinos, gravitinos, axinos, sneutrinos, gluinos...

WIMPs

“Weakly Interacting Massive Particles”

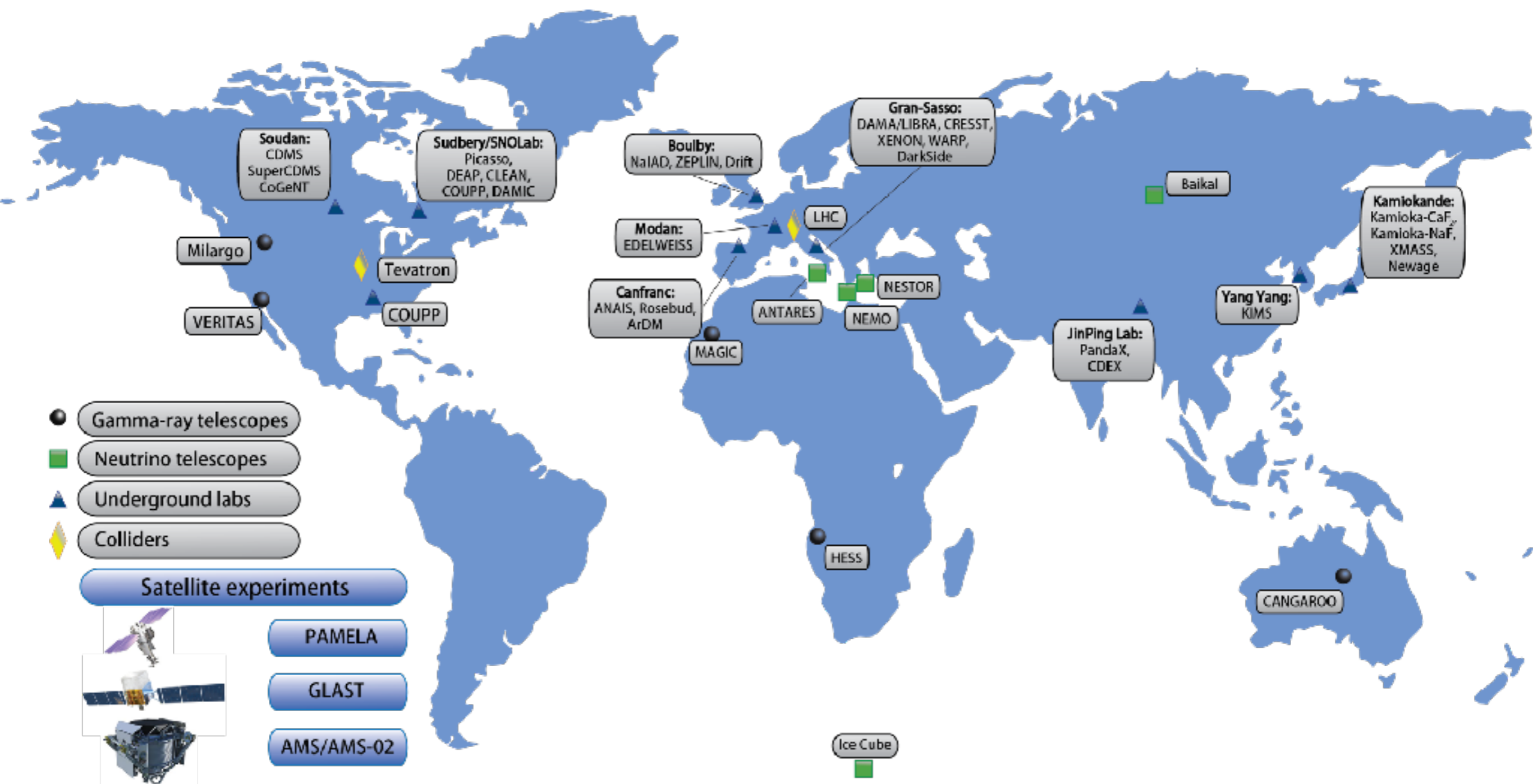


Crédito: Tim M. P. Tait

Gravidade Escura?

A visualization of the cosmic web, showing a complex network of dark matter filaments and galaxy clusters. The filaments are represented by thin, dark lines, and the galaxy clusters are represented by bright, yellowish-orange points of light. The background is a dark, textured grey, suggesting the distribution of dark matter in the universe.

Esforços para encontrar a partícula de Matéria Escura



Recent Progress in Search for Dark Sector Signatures
 M.A. Deliyergiyev. Open Phys. 14 (2016) no.1, 281-303

Cryogenic Dark Matter Search

Large Hadron Collider

FermiLab

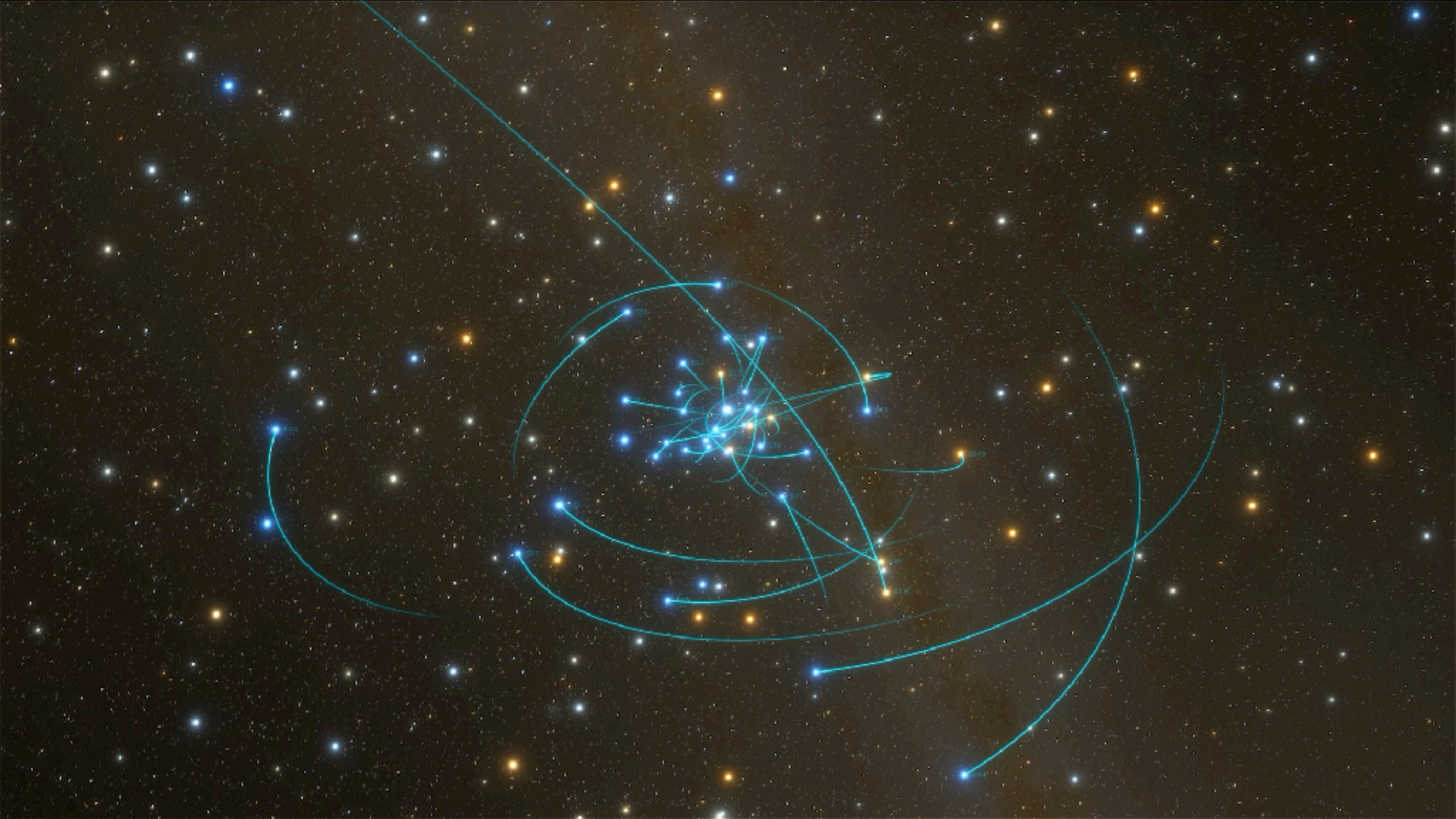
CoGeNT / XENON / DAMA/LIBRA

NASA's Fermi gamma-ray Satellite

Gamma Ray Large Array Space Telescope

LZ Dark Matter Experiment

(...)



Crédito: ESO/GRAVITY Collaboration

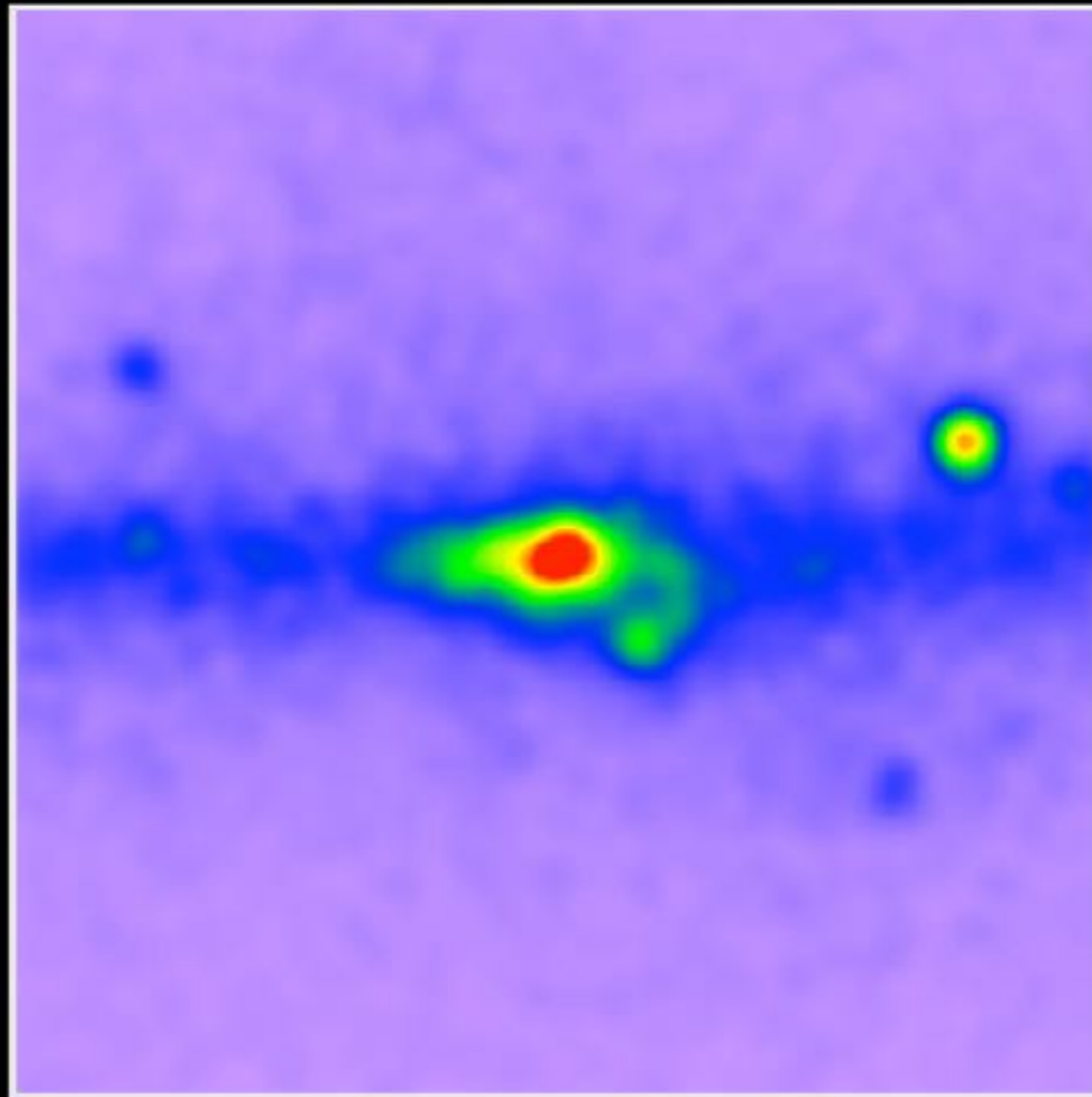
Comportamento de estrelas próximas ao buraco negro no centro da Via Láctea



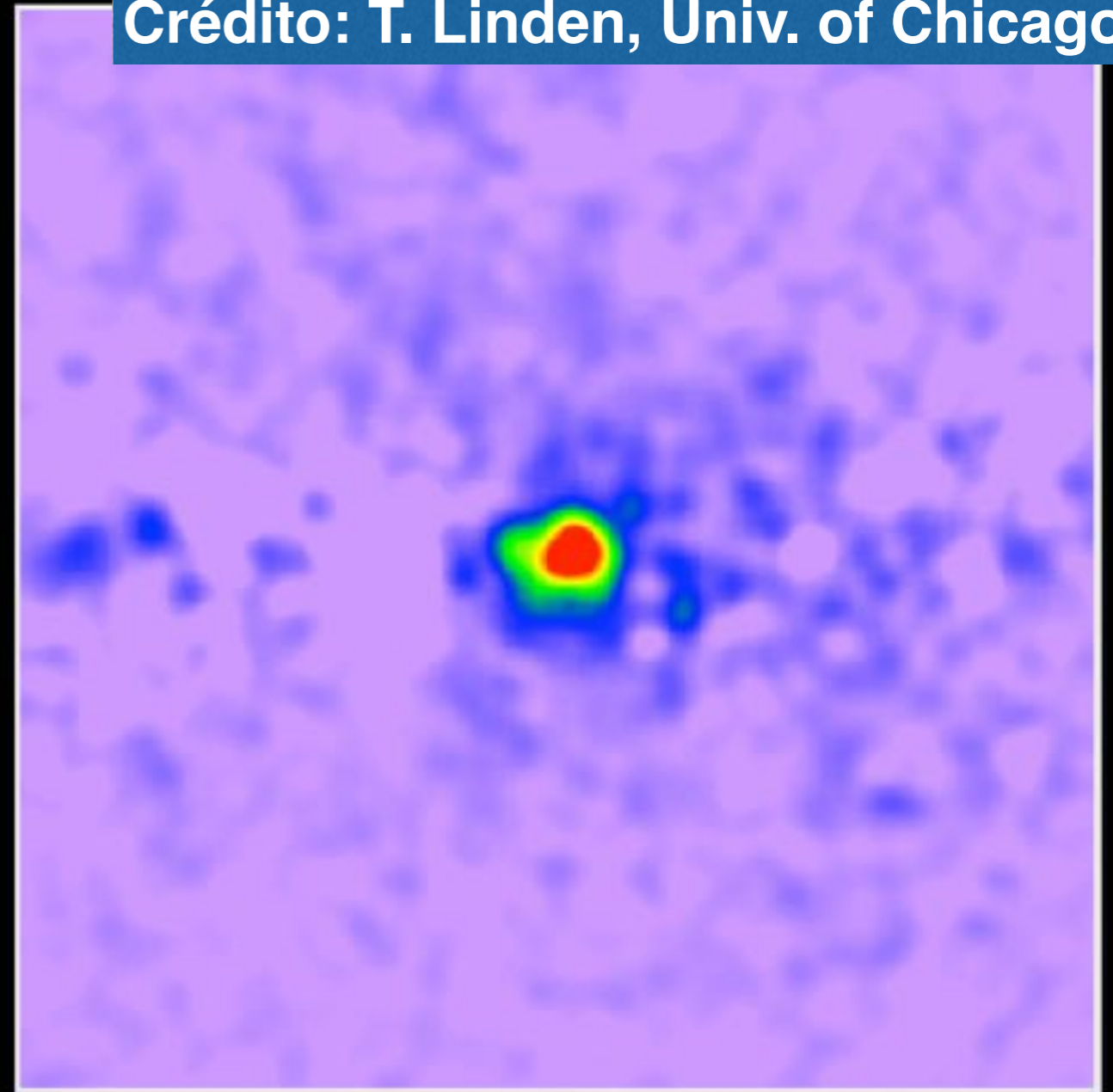
Crédito: ESO/MPE/M. Schartmann/L. Calçada

Uncovering a gamma-ray excess at the galactic center

Crédito: T. Linden, Univ. of Chicago



Unprocessed map of 1.0 to 3.16 GeV gamma rays



Known sources removed

Mapa de raios gama detectados no centro galáctico pelo LAT/Fermi. À direita, raios gama após remoção de fontes conhecidas, podendo ter como fonte a aniquilação de matéria escura.



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