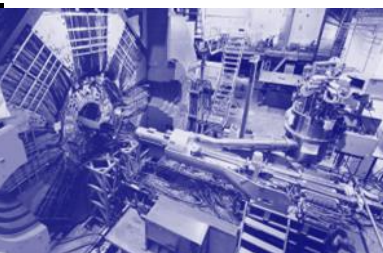


Particle and Astro-Particle Physics in China

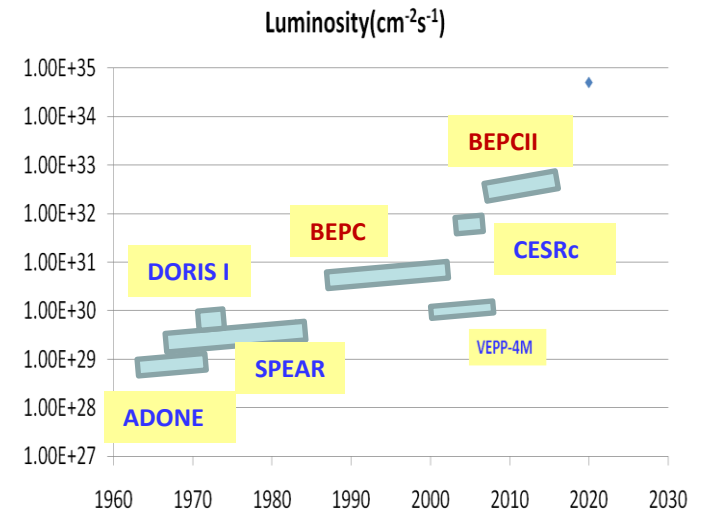
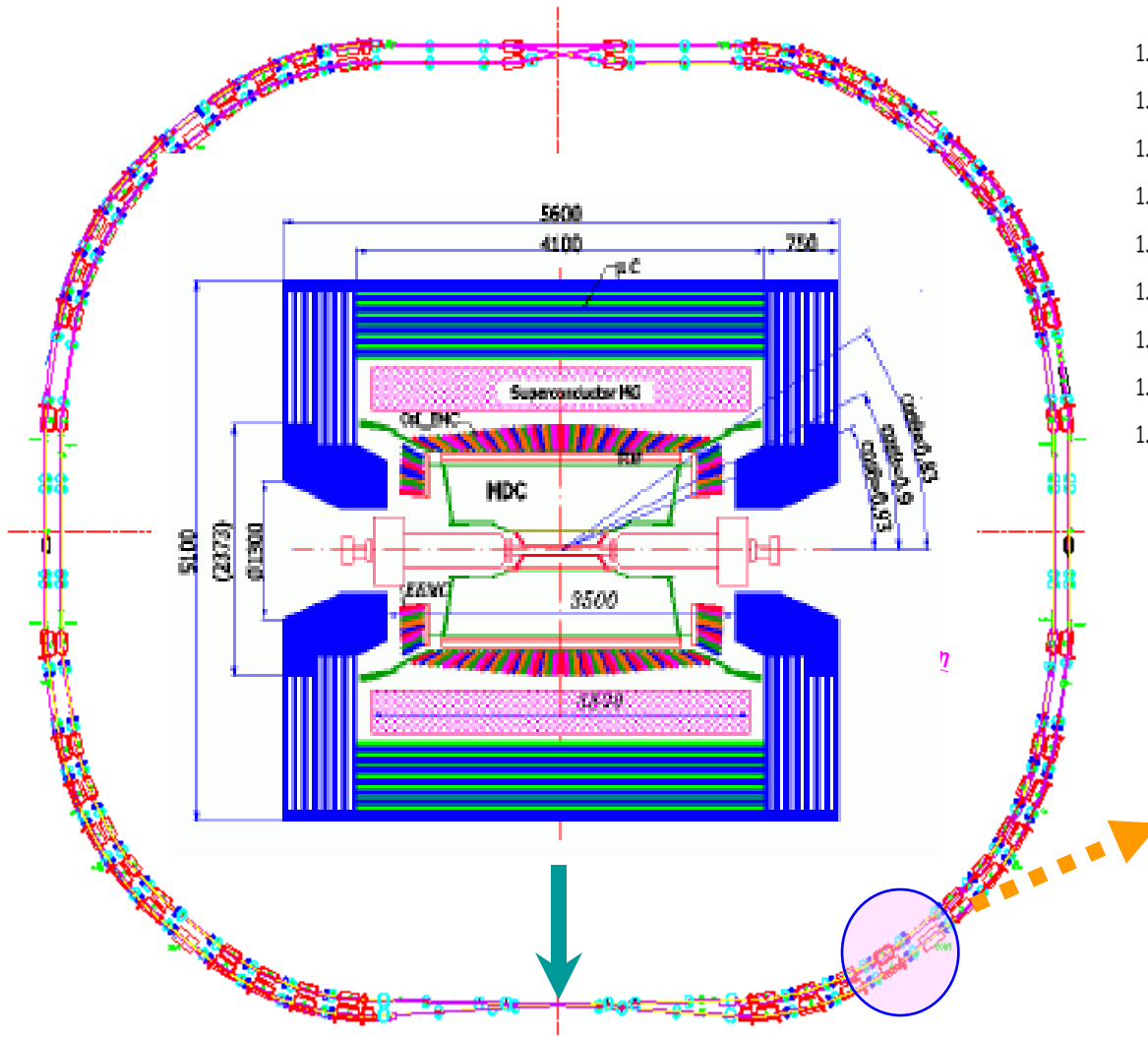
Yifang Wang

Institute of High Energy Physics

LISHEP 2018, Sep. 12, 2018



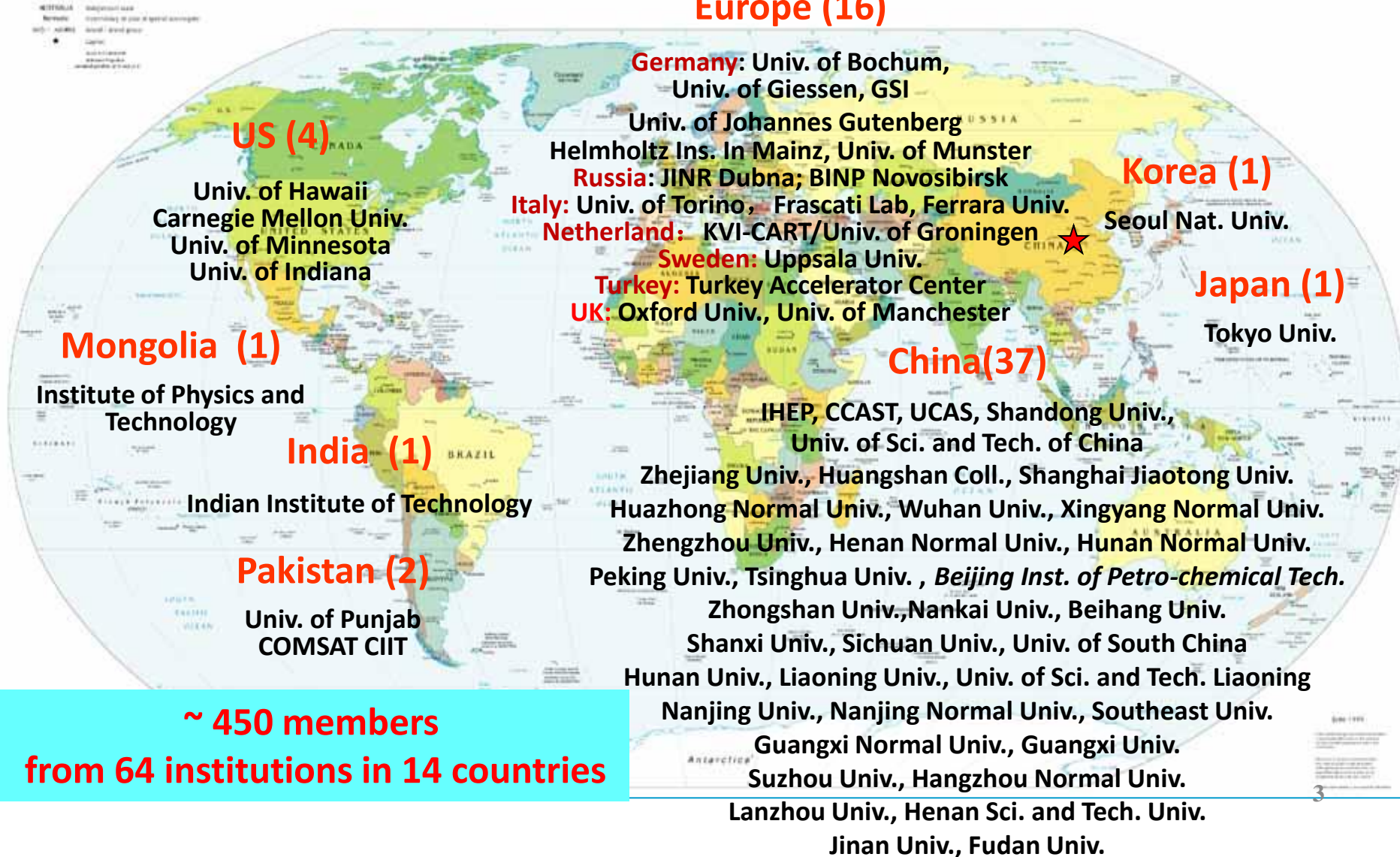
HEP in China Started from BEPC



BEPCII/BESIII Upgrade: 2004-2008

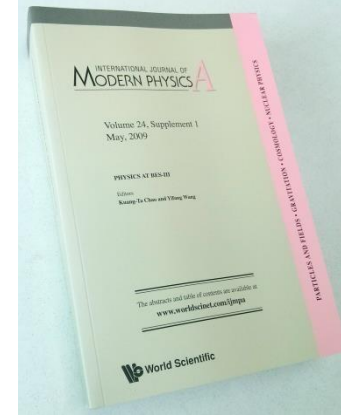
BESIII Collaboration

Political Map of the World, June 1999

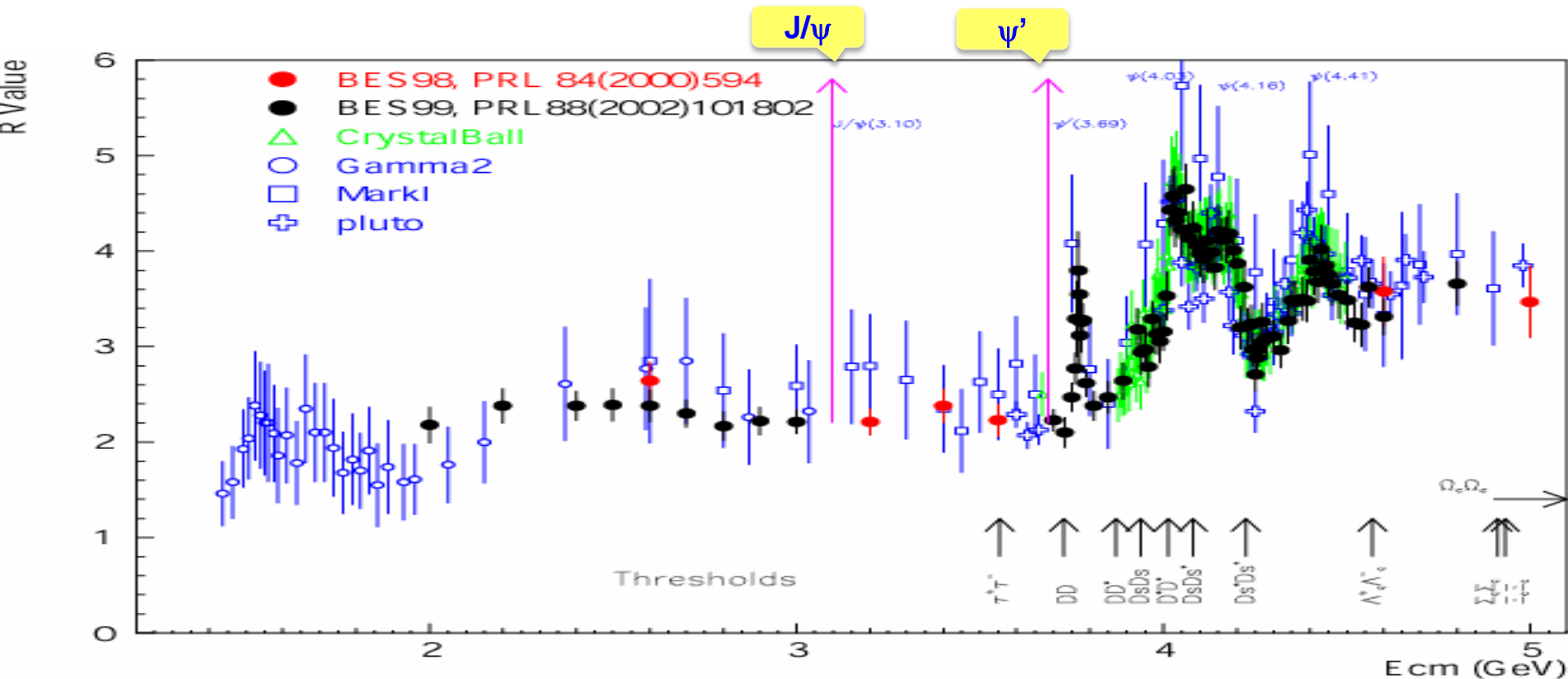


Physics at BESIII

- ◆ Hadron spectroscopy and non-conventional hadron searches
- ◆ Charmonium production and decays: **XYZ states**
- ◆ Precision study of R, form factors, tau mass, CKM matrix, etc.

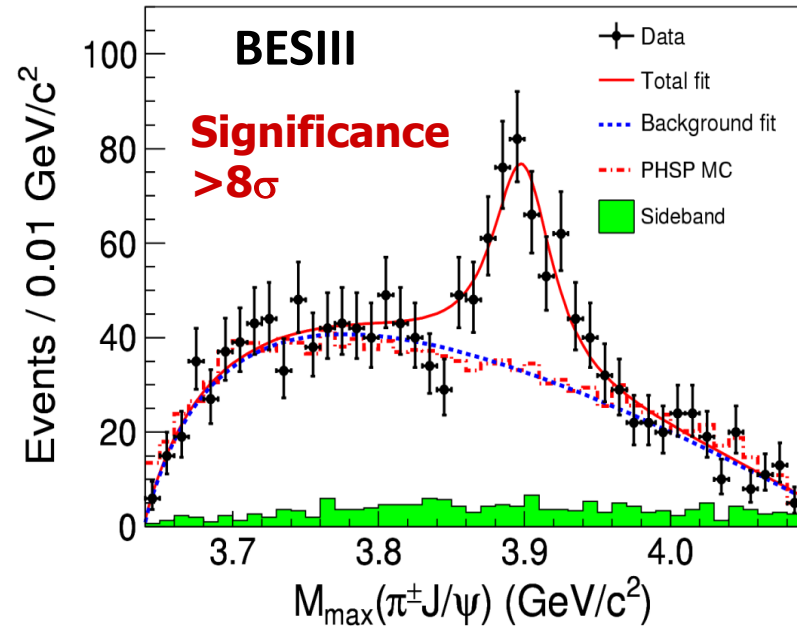


Int. J. Mod. Phys. A, Vol. 24 (2009)



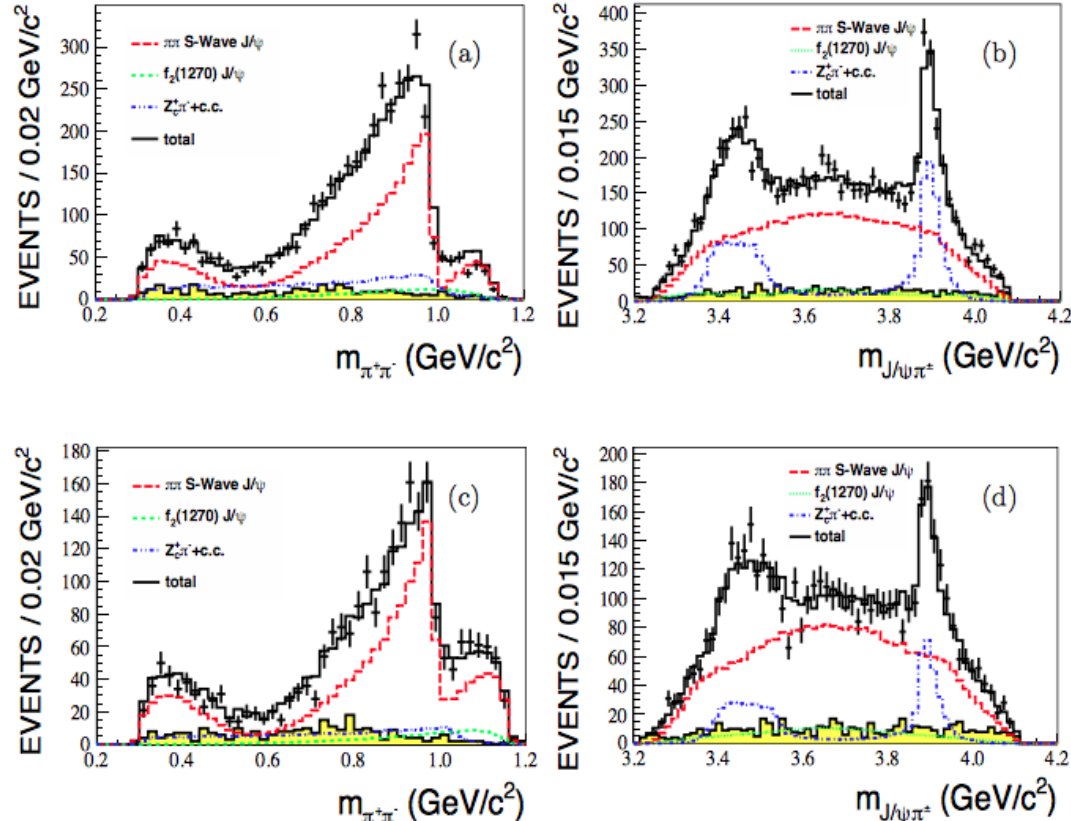
$Z_c^+(3900)$ in $e^+e^- \rightarrow \Upsilon(4260) \rightarrow \pi^+\pi^-J/\psi$

BESIII: PRL110, 252001 (2013)



- $M = 3899.0 \pm 3.6 \pm 4.9$ MeV
- $\Gamma = 46 \pm 10 \pm 20$ MeV
- 307 ± 48 events

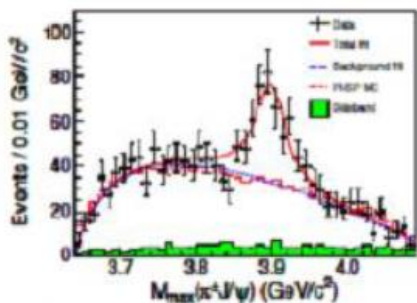
PRL 119, 072001 (2017)



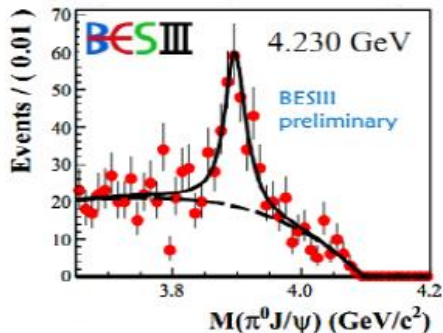
J^P of Z_c favor 1^+ @ 7.3σ over others

- Amplitude analysis with helicity formalism
- Simultaneous fit to data at 4.23 & 4.26 GeV

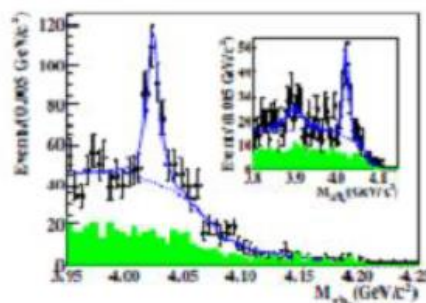
Many Z_c 's at BESIII



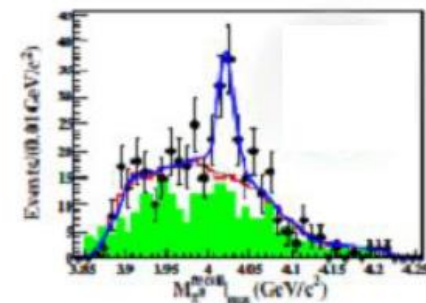
$$e^+e^- \rightarrow \pi^- \pi^+ J/\psi$$



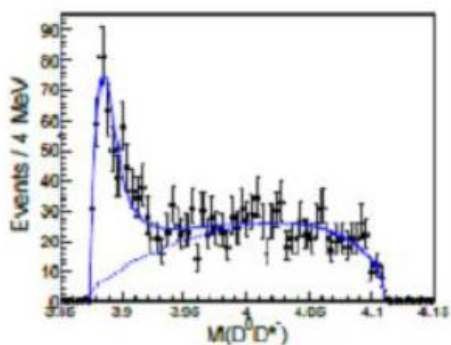
$$e^+e^- \rightarrow \pi^0 \pi^0 J/\psi$$



$$e^+e^- \rightarrow \pi^- \pi^+ h_c$$

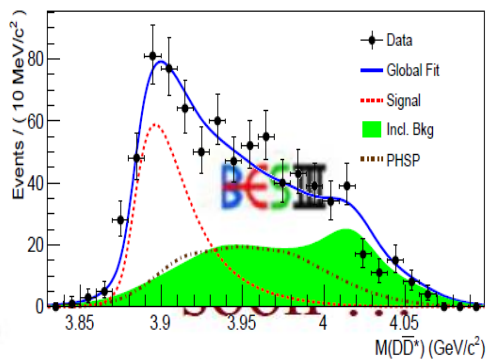


$$e^+e^- \rightarrow \pi^0 \pi^0 h_c$$



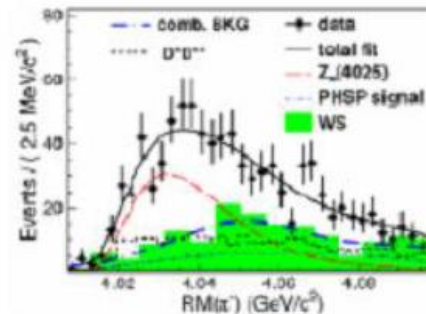
$$e^+e^- \rightarrow \pi^- (D\bar{D}^*)^+$$

$$Z_c(3900)^+?$$



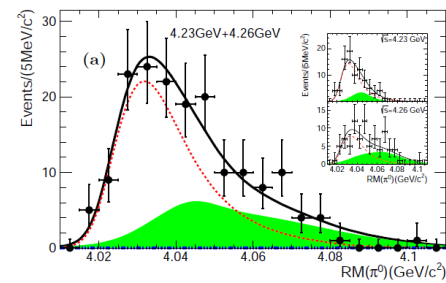
$$e^+e^- \rightarrow (D\bar{D}^*)^0 \pi^0$$

$$Z_c(3900)^0?$$



$$e^+e^- \rightarrow \pi^- (D^* \bar{D}^*)^+$$

$$Z_c(4020)^+?$$

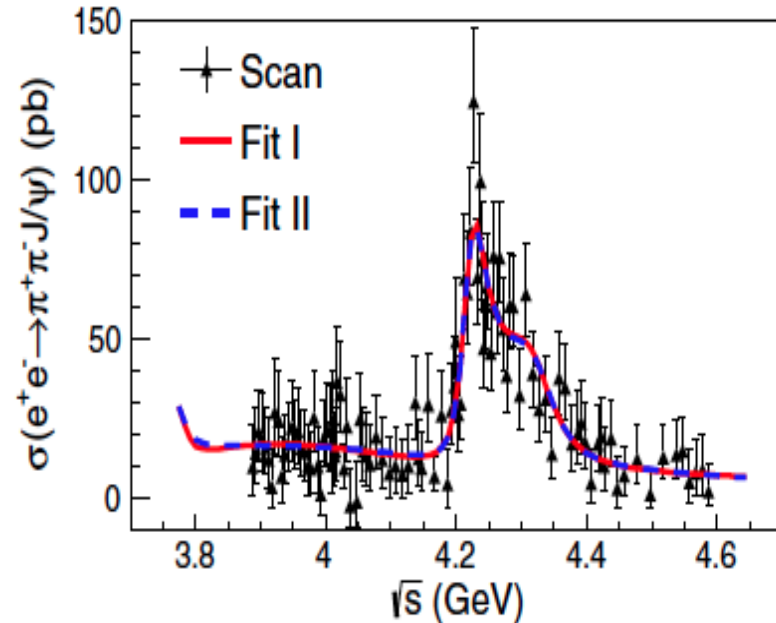
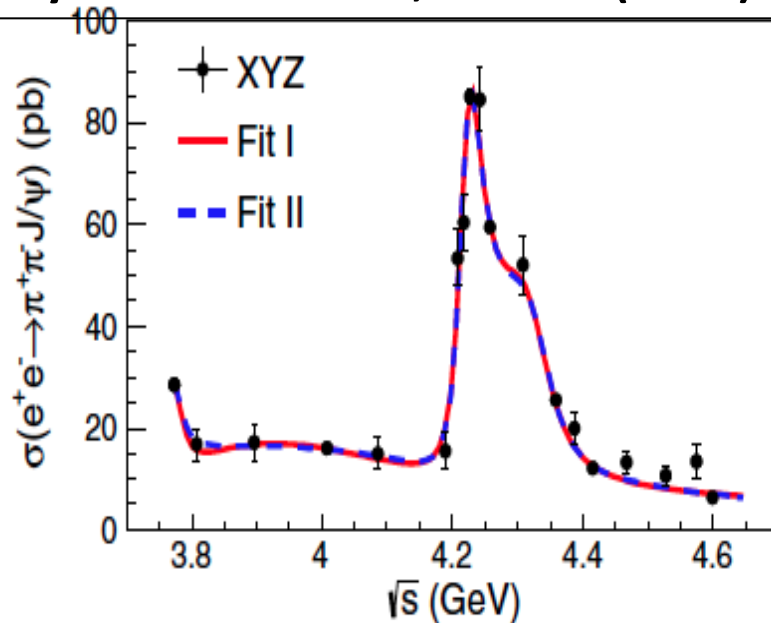


$$e^+e^- \rightarrow \pi^0 (D^* \bar{D}^*)^0$$

$$Z_c(4020)^0?$$

Y(4260): with structures

Phys. Rev. Lett. 118, 092001 (2017)



□ Coherent sum of two BW-like structures + one incoherent $\psi(3770)$

➤ $M = (4222.0 \pm 3.1 \pm 1.4) \text{ MeV}$, $\Gamma = (44.1 \pm 4.3 \pm 2.0) \text{ MeV}$

Lower and narrower than previous Y(4260) PDG values

➤ $M = (4320.0 \pm 10.4 \pm 7) \text{ MeV}$, $\Gamma = (101.4 \pm 25 \pm 10) \text{ MeV}$

a little bit lower than Y(4360) PDG value

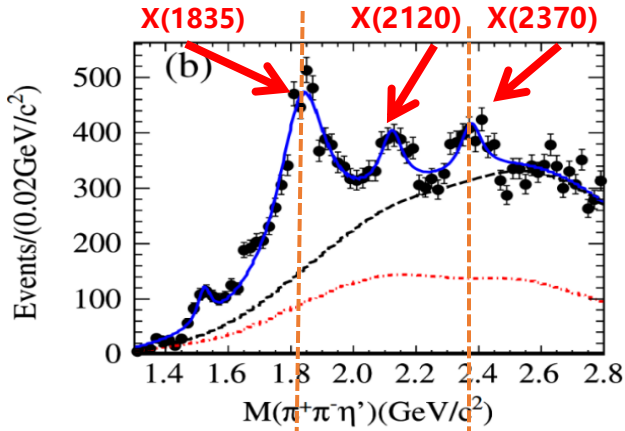
□ Compared with one BW fit, the sig. of the second BW is 7.6σ

□ Y(4260) + Y(4360) ? The first observation of $Y(4360) \rightarrow \pi^+\pi^- J/\psi$? 7

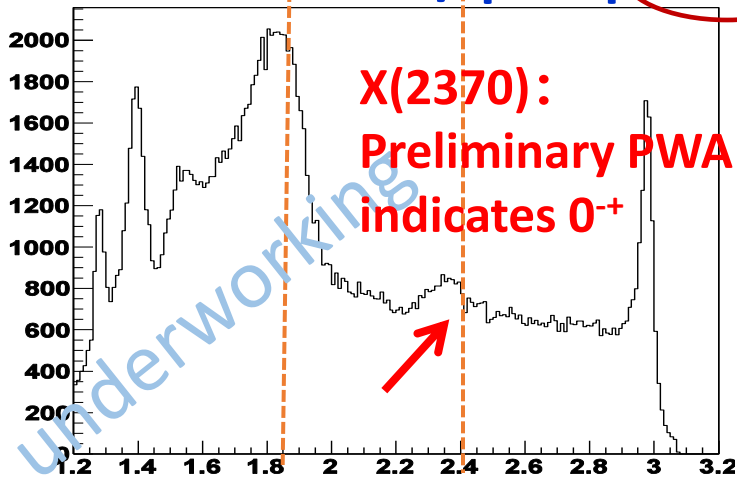
Systematic study of glueballs at BESIII

-- X(2370): a 0^{-+} glueball candidate

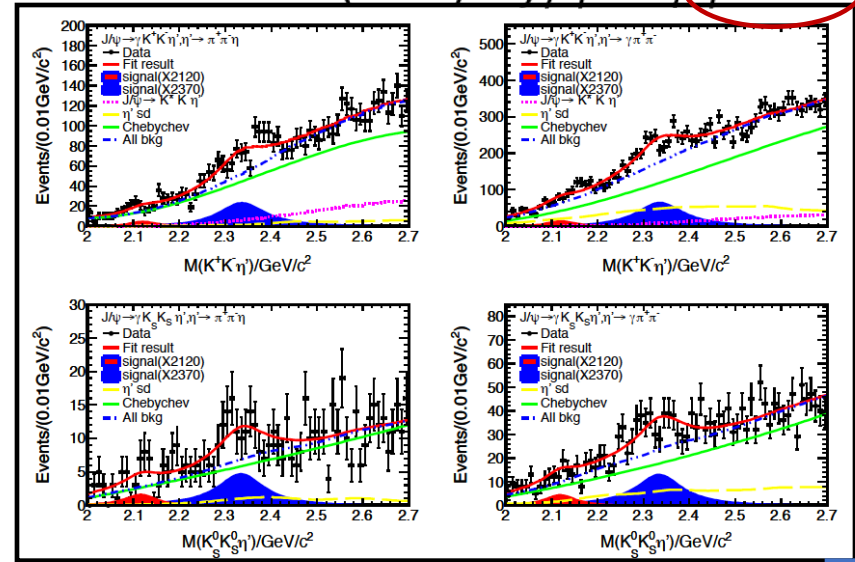
PRL 106 072002 $J/\psi \rightarrow \gamma \eta' \pi^+ \pi^-$



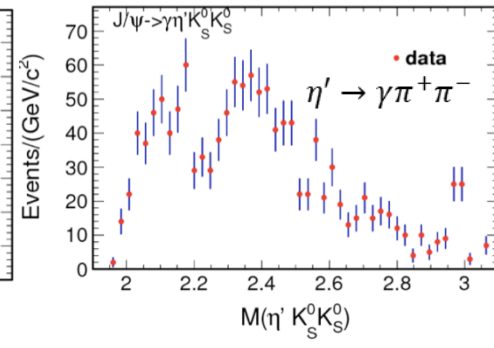
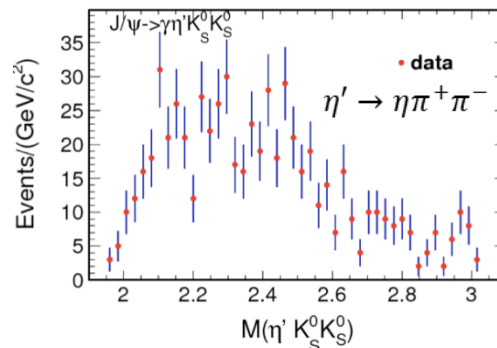
$J/\psi \rightarrow \gamma \pi^0 \pi^0 \eta$



Observation of X(2370) in $J/\psi \rightarrow \gamma \eta' KK$



With new J/ψ data taken in 2018, clear structures in $J/\psi \rightarrow \gamma \eta' Ks Ks$



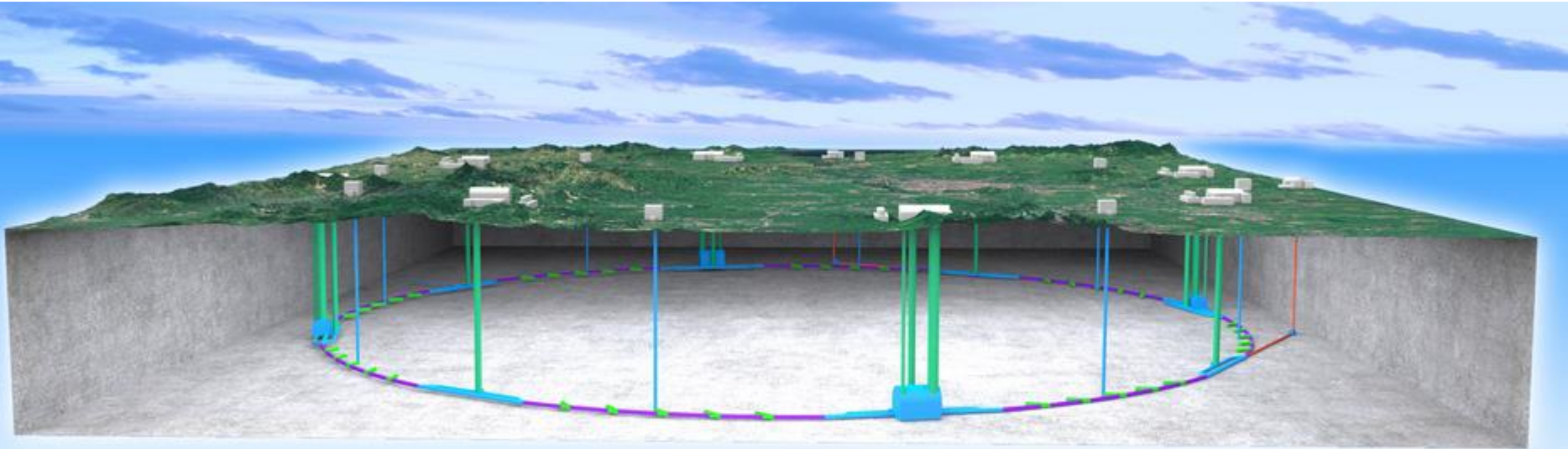
Future Plan of BEPCII/BESIII

- Up to now, ~ 20-30 papers/year, ~ 200 papers in total
- Minor upgrades completed, more under study
- BEPCII/BESIII will continue to operate for another ~8 years.

	Previous data	BESIII present	Goal
J/ψ	BESII 58M	6 B 120* BESII	10 B
ψ'	CLEO: 28 M	0.5 B 20* CLEOc	3B
ψ''	CLEO: 0.8/fb	2.9/fb 3.5*CLEOc	20 /fb
Above open charm threshold	CLEO: 0.6/fb @ ψ(4160)	0.5/fb @ ψ(4040) 2.3/fb@~4260, 0.5/fb@4360 0.5/fb@4600, 1/fb@4420 Scan from 4.19 – 4.28, 10 MeV step, 500 pb ⁻¹ /point, 7 points	5-10 /fb
R scan & Tau	BESII	3.8-4.6 GeV at 105 energy points 2.0-3.1 GeV at 20 energy points	
Υ(2175)		100 pb ⁻¹	
ψ(4160)		3 fb ⁻¹	

Next step: CEPC — Higgs factory

- Since 2005, we were discussing the next machine after BEPC/BEPCII
- Thanks to the low mass Higgs, there is the possibility to build a Higgs Factory: Circular e+e- Collider(CEPC)
 - Looking for Hints (from Higgs) → direct searches
 - The tunnel can allow us to build pp, AA, ep colliders in the far future: Super proton-proton Collider(SppC)



Science of CEPC-SPPC

- **Electron-positron collider(90, 250 GeV)**
 - **Higgs Factory (10^6 Higgs) :**
 - Precision study of Higgs(m_H , J^{PC} , couplings), Similar & complementary to ILC
 - Looking for hints of new physics
 - **Z & W factory (10^{10} Z^0) :**
 - precision test of SM
 - Rare decays ?
 - **Flavor factory: b, c, τ and QCD studies**
- **Proton-proton collider(~ 100 TeV)**
 - **Directly search for new physics beyond SM**
 - **Precision test of SM**
 - e.g., h^3 & h^4 couplings

**Precision measurement + searches:
Complementary with each other !**

Precision Higgs Physics by CEPC

$$\mathcal{L} = \mathcal{L}_{SM} + \sum_i \frac{c_i}{M^2} \mathcal{O}_{6,i} \quad \delta \sim c_i \frac{v^2}{M^2}$$

A total of 10^6 Higgs

% precision \rightarrow $M \sim 1$ TeV

to new physics $\rightarrow \sim \times 10$ over LHC

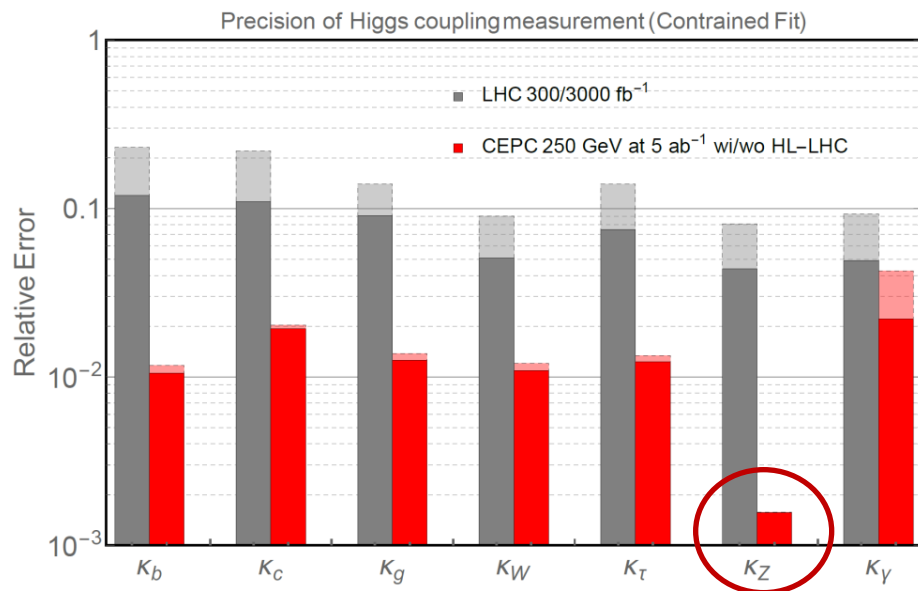
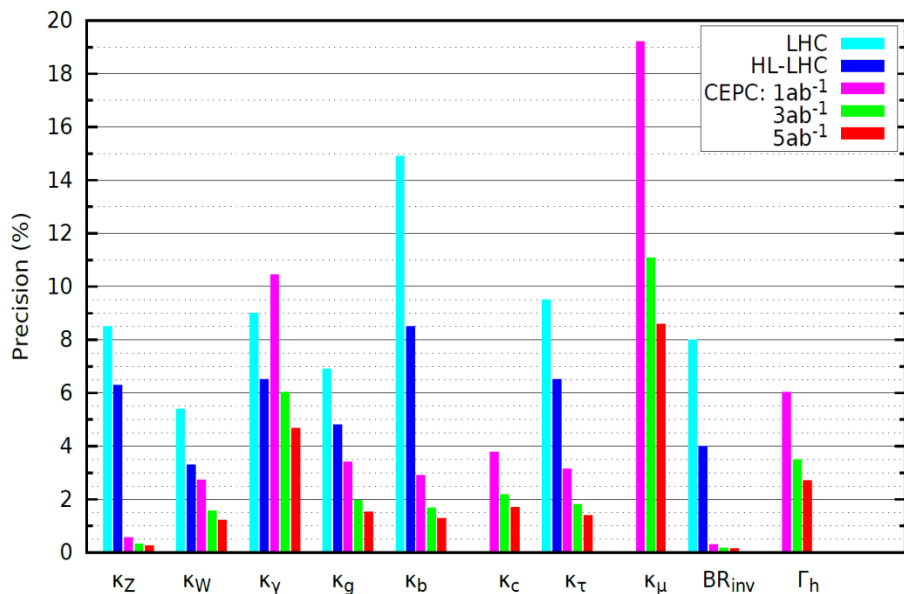
Only elementary particle

➤ with spin 0

➤ with non-gauge interactions

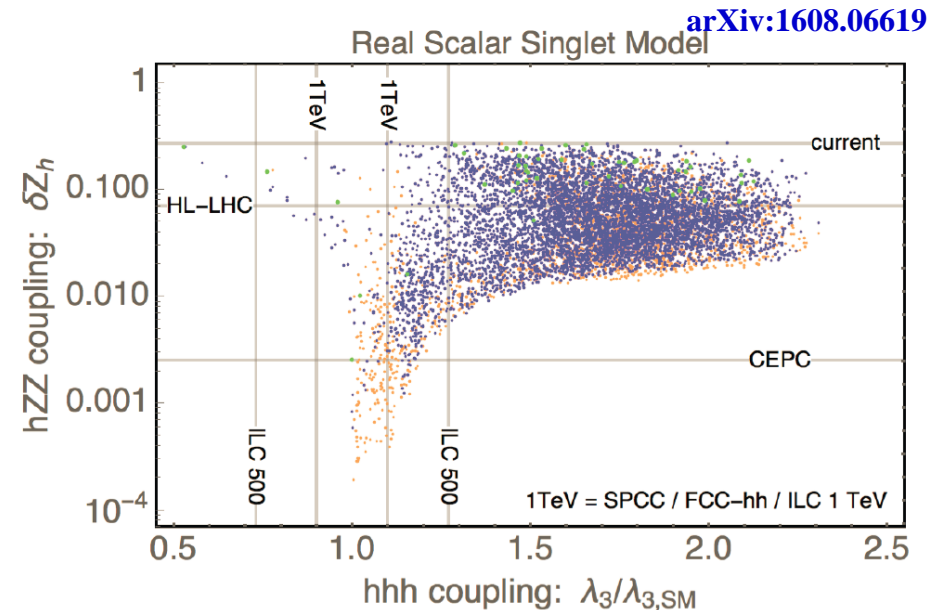
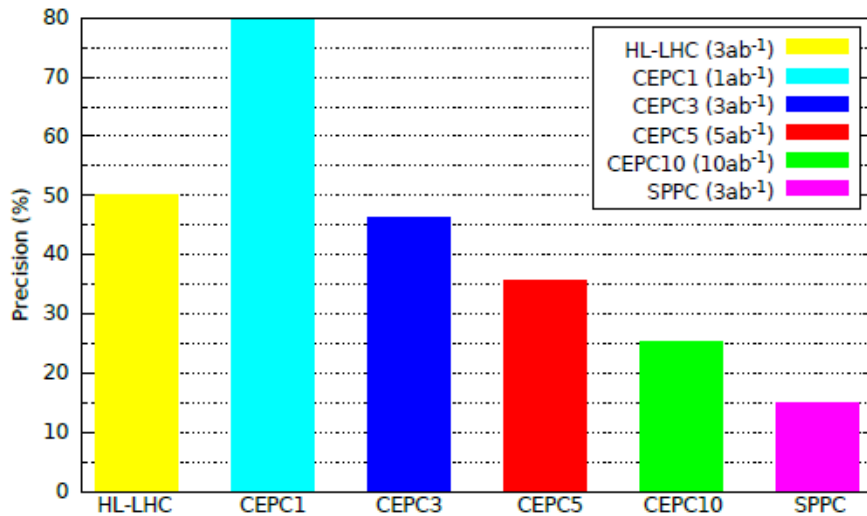
Related to “problems” of SM

A Step can not be skipped



Nature of EW Phase Transition ?

- 1st or 2nd order → Huge implications
 - O(1) deviations in h^3 coupling
 - O(1%) shift in h -Z coupling
- CEPC can determine it:
 - h^3 coupling at CEPC: 20-30%
 - h -Z coupling at CEPC: < 0.2%

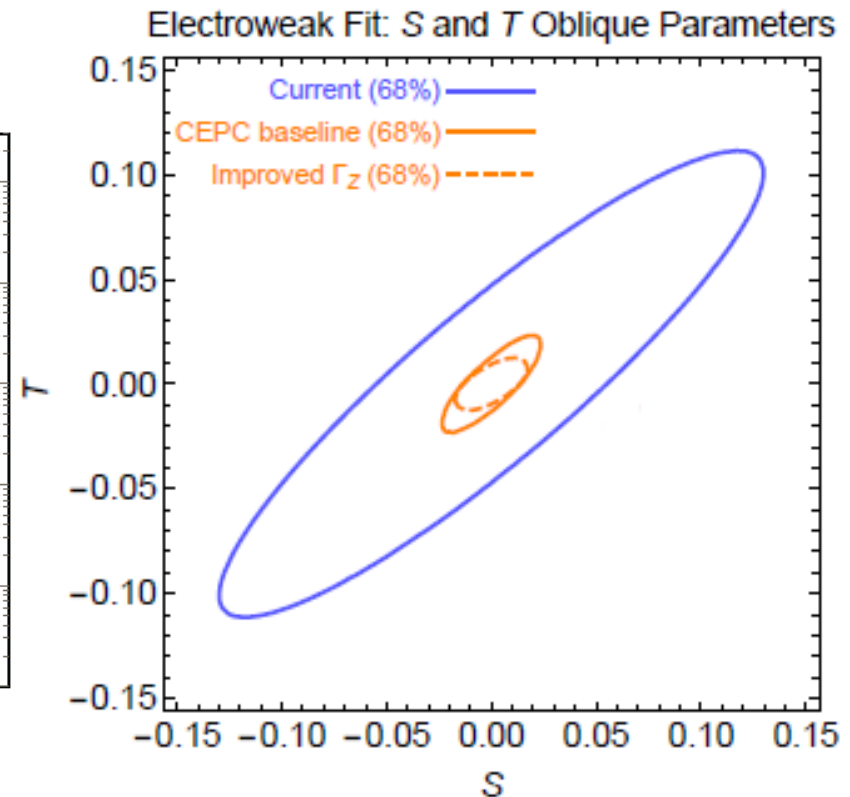
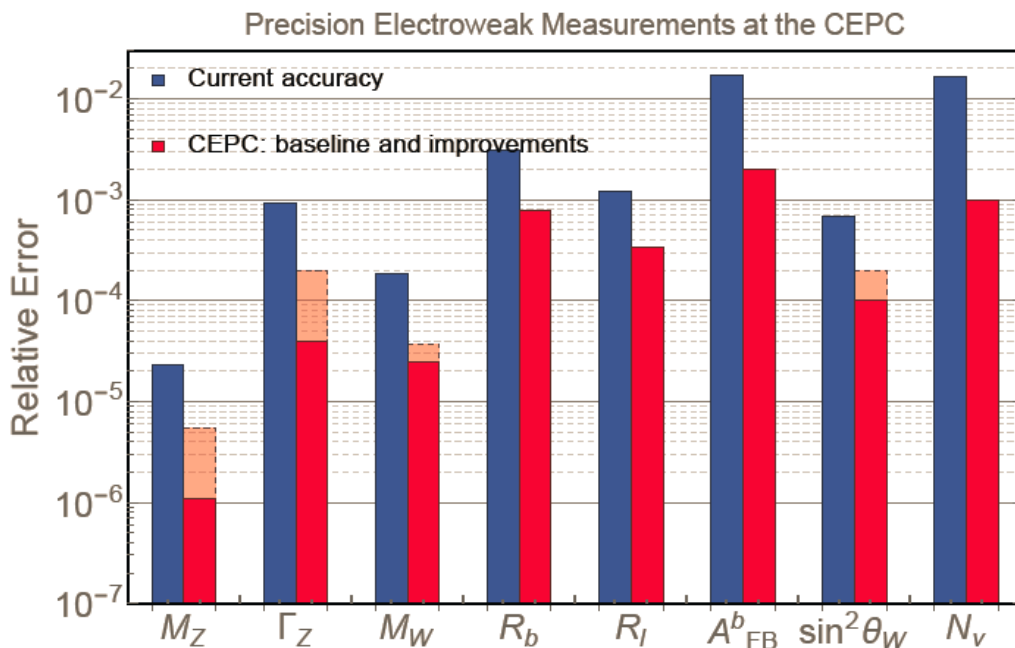


2018-9-12

M. McCullough, PRD 90(2014)015001

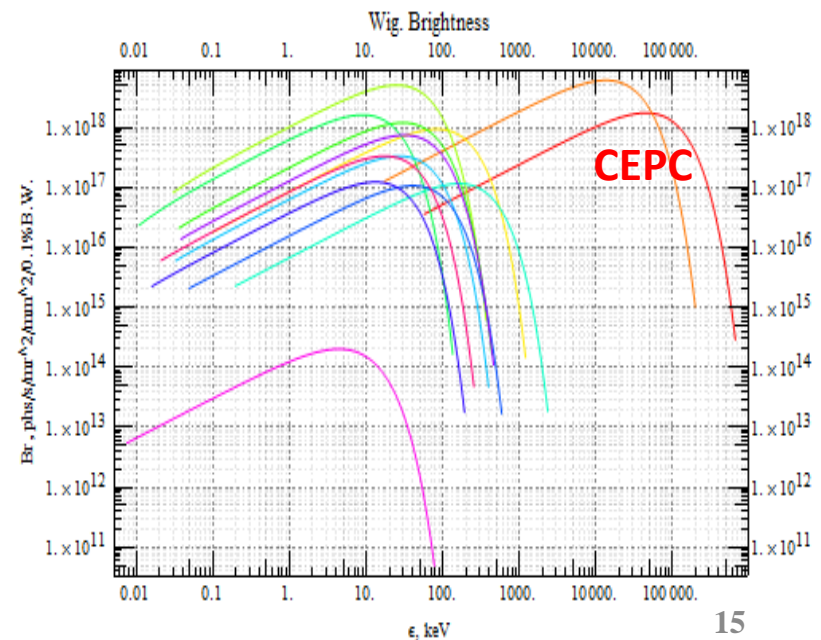
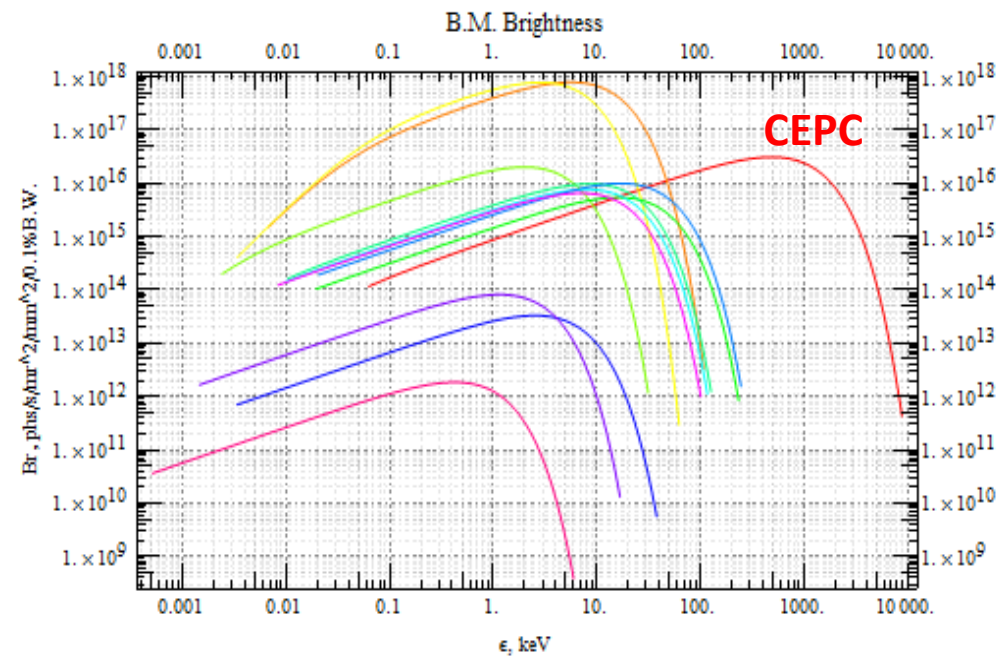
Improvement in Electroweak Precision

- A total of 10^{10} Z
- A detailed study of Z & W to look for deviations from the Standard Model
- Can probe new physics up to \sim TeV, better than HL-LHC by a factor of 3

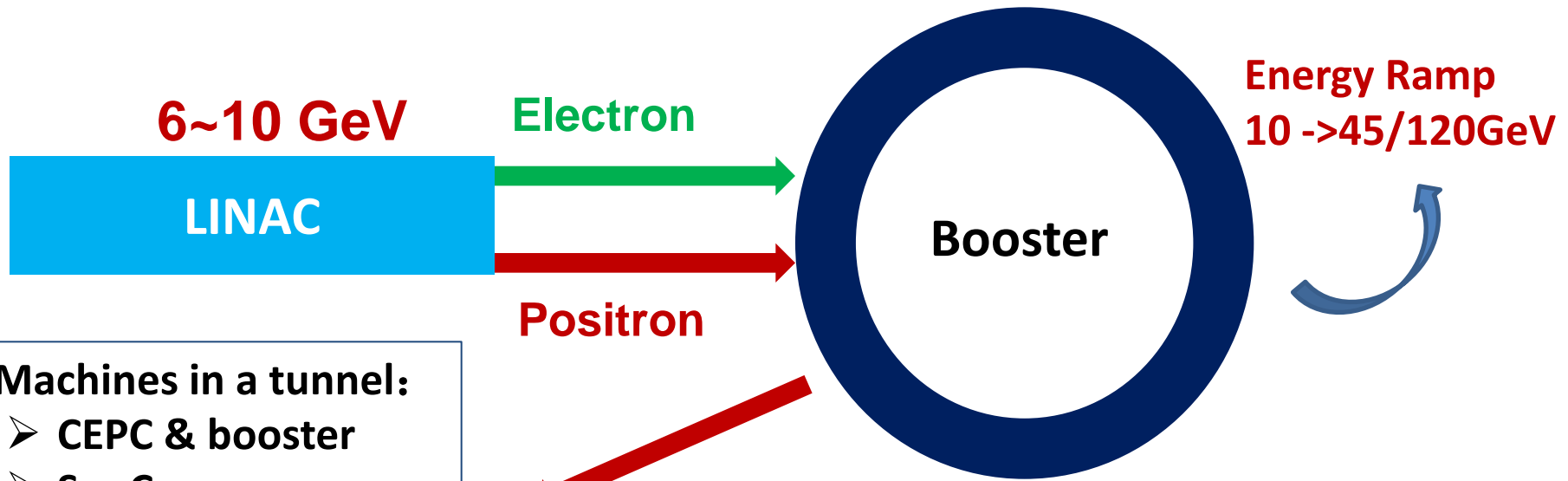


CEPC is also a great light source

- From dipole magnet, the photon energy can reach **628keV**.
- From Wiggler or undulator, the photon energy can reach **100MeV**
- Incredible application on nuclear physics, material science, micro-processing, etc.



CEPC Accelerator Design

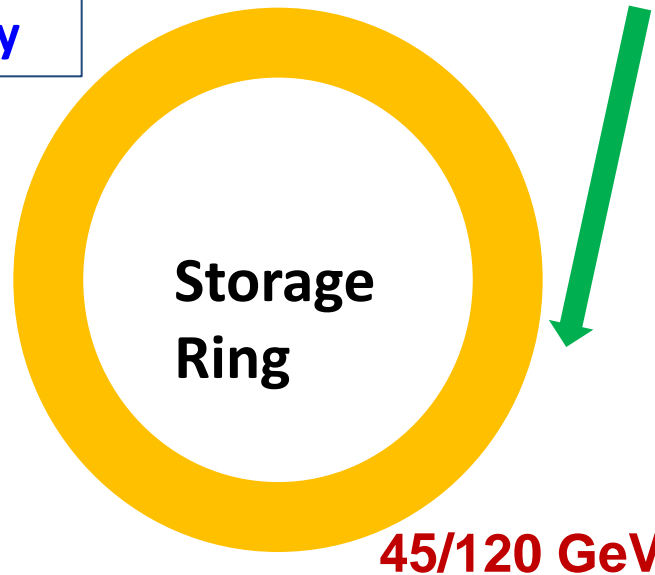
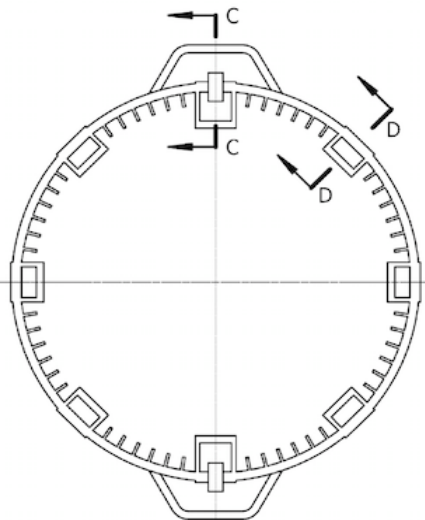


3 Machines in a tunnel:

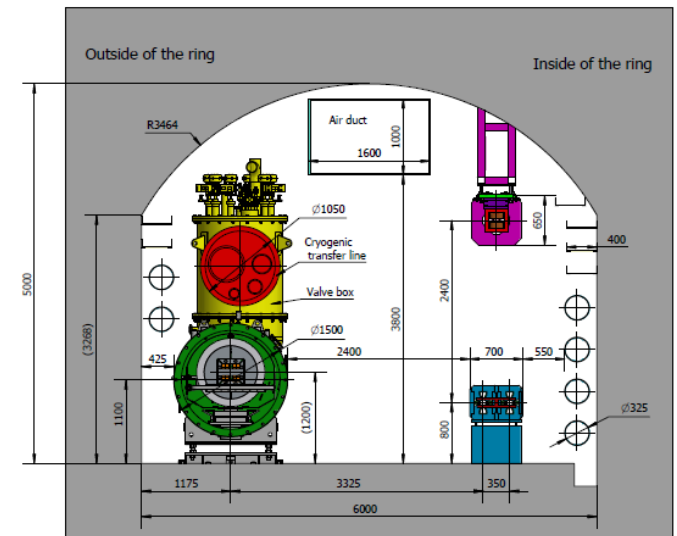
- CEPC & booster
- SppC

Compatibility is the key

隧道俯视图示意图



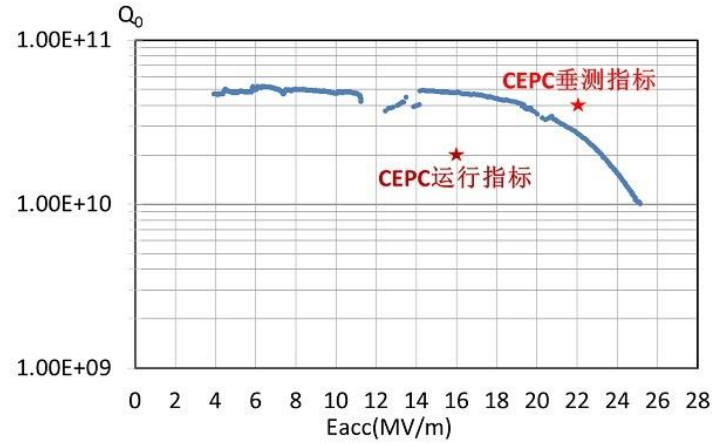
TUNNEL CROSS SECTION OF THE ARC AREA



Main Parameters

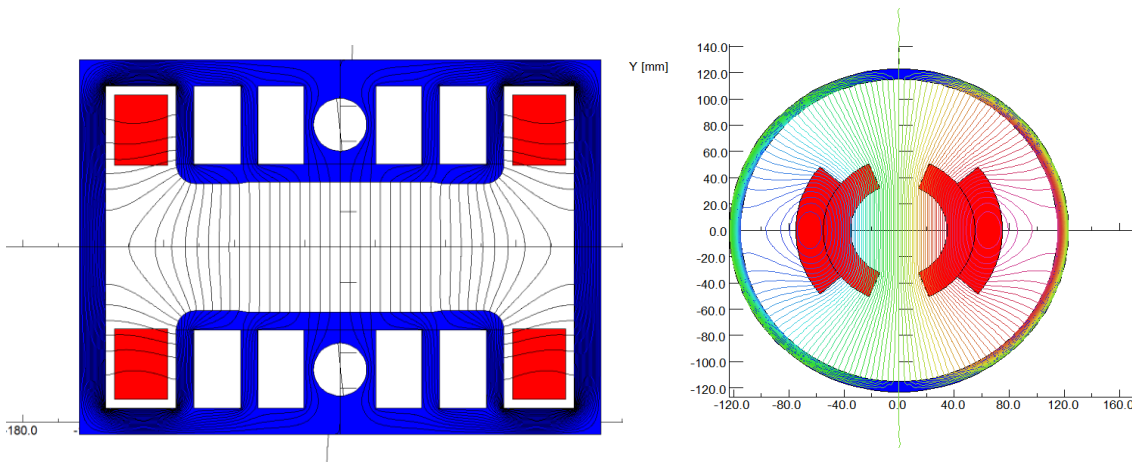
	<i>Higgs</i>	<i>W</i>	<i>Z (3T)</i>	<i>Z (2T)</i>
Number of IPs	2			
Beam energy (GeV)	120	80	45.5	
Circumference (km)	100			
Synchrotron radiation loss/turn (GeV)	1.73	0.34	0.036	
Crossing angle at IP (mrad)	16.5×2			
Piwinski angle	2.58	7.0	23.8	
Number of particles/bunch N_e (10^{10})	15.0	12.0	8.0	
Bunch number (bunch spacing)	242 (0.68 μ s)	1524 (0.21 μ s)	12000 (25ns+10% gap)	
Beam current (mA)	17.4	87.9	461.0	
Synchrotron radiation power /beam (MW)	30	30	16.5	
Bending radius (km)	10.7			
Momentum compact (10^{-5})	1.11			
β function at IP β_x^*/β_y^* (m)	0.36/0.0015	0.36/0.0015	0.2/0.0015	0.2/0.001
Emittance $\varepsilon_x/\varepsilon_y$ (nm)	1.21/0.0031	0.54/0.0016	0.18/0.004	0.18/0.0016
Beam size at IP σ_x/σ_y (μ m)	20.9/0.068	13.9/0.049	6.0/0.078	6.0/0.04
Beam-beam parameters ξ_x/ξ_y	0.031/0.109	0.013/0.106	0.0041/0.056	0.0041/0.072
RF voltage V_{RF} (GV)	2.17	0.47	0.10	
RF frequency f_{RF} (MHz) (harmonic)	650 (216816)			
Natural bunch length σ_z (mm)	2.72	2.98	2.42	
Bunch length σ_z (mm)	3.26	5.9	8.5	
Betatron tune ν_x/ν_y	363.10 / 365.22			
Synchrotron tune ν_s	0.065	0.0395	0.028	
HOM power/cavity (2 cell) (kw)	0.54	0.75	1.94	
Natural energy spread (%)	0.1	0.066	0.038	
Energy acceptance requirement (%)	1.35	0.4	0.23	
Energy acceptance by RF (%)	2.06	1.47	1.7	
Photon number due to beamstrahlung	0.29	0.35	0.55	
Lifetime _simulation (min)	100			
Lifetime (hour)	0.67	1.4	4.0	2.1
F (hour glass)	0.89	0.94	0.99	
Luminosity/IP L (10^{34} cm $^{-2}$ s $^{-1}$)	2.93	10.1	16.6	32.1

R&D and Prototypes

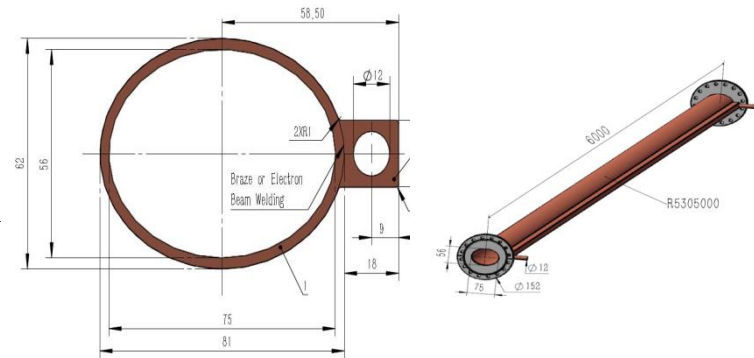


Collaborating on Klystrons

Superconducting RF Cavities



High precision, low field dipole magnet



6m long vacuum pipes(Al & Cu)

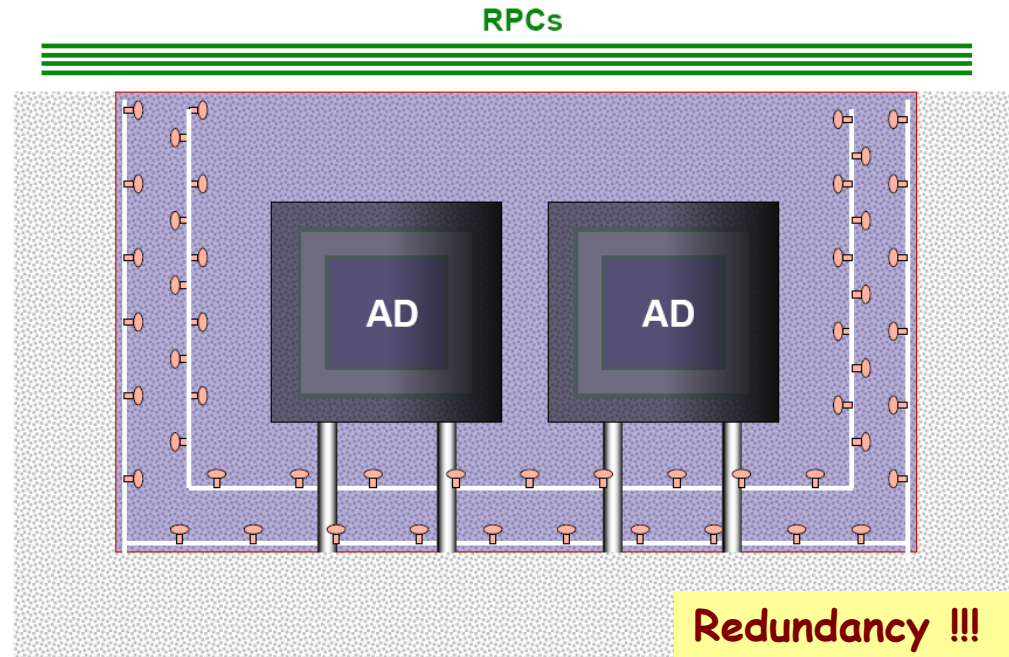
CEPC Schedule (ideal)



Major Projects: Current and Future

		Current	Future
Accelerator-based	Precision frontier	BESIII LHCb, Belle II, PANDA, COMET	ILC, FCC ? CEPC → SppC ?
	Energy frontier	CMS, ATLAS	
Non-accelerator-based	Underground	Daya Bay	JUNO
		Jinping: PANDAx, CDEX	Exp. ?
	Surface	ARGO/AS γ	LHASSO
	Space	AMS	HERD ?
HXMT, Polar, DAMPE		XTP ?	

Daya Bay Experiment



◆ Successful construction(2007-2011):

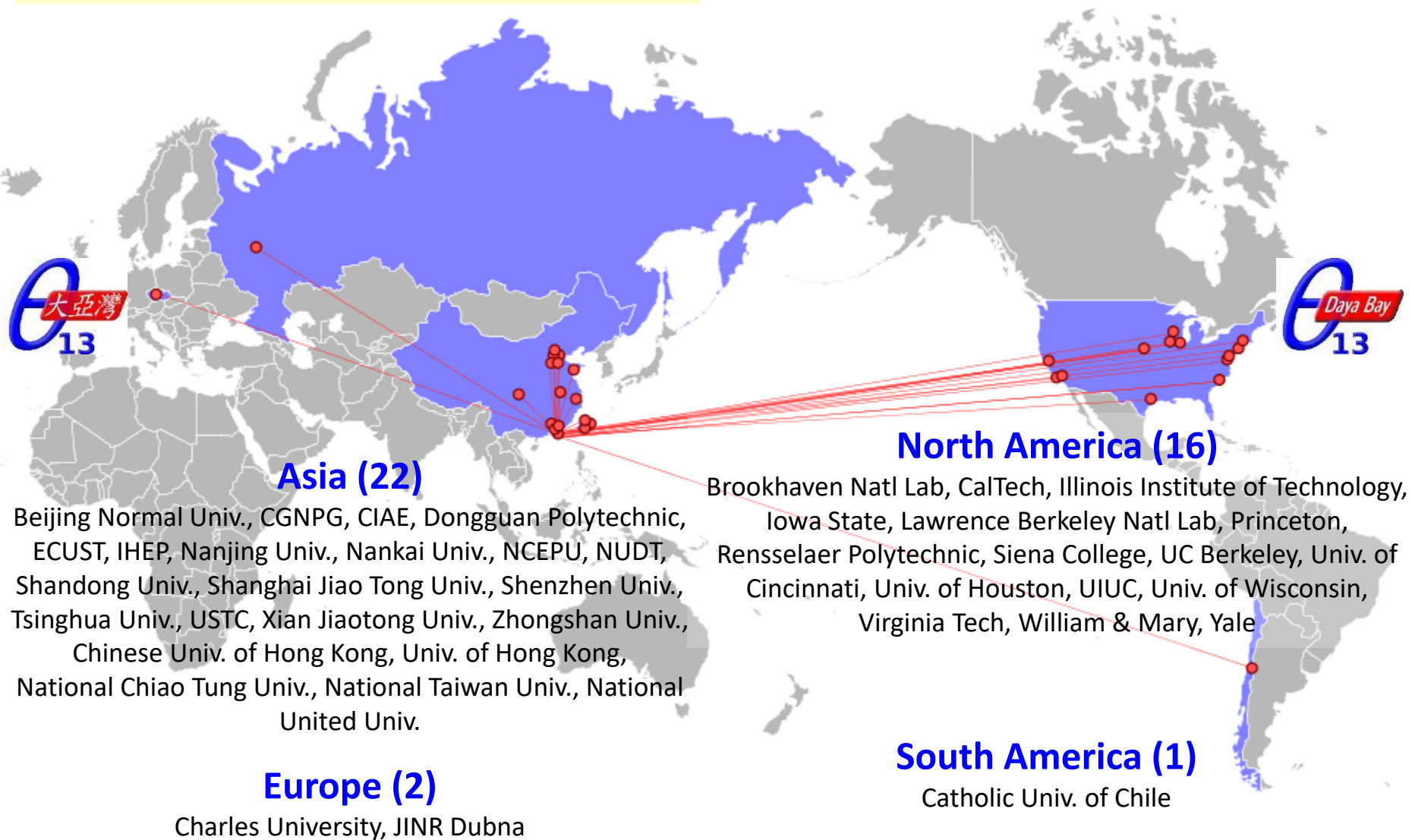
- ⇒ 3 km tunnel, 5 halls
- ⇒ 200t Gd-loaded LS, 8 identical detector modules
- ⇒ 3 water Č detectors
- ⇒ 3200 m² RPC detectors
- ⇒ 8000 readout channels
- ⇒ Calibration system



Daya Bay Collaboration



41 institutions, 193 collaborators

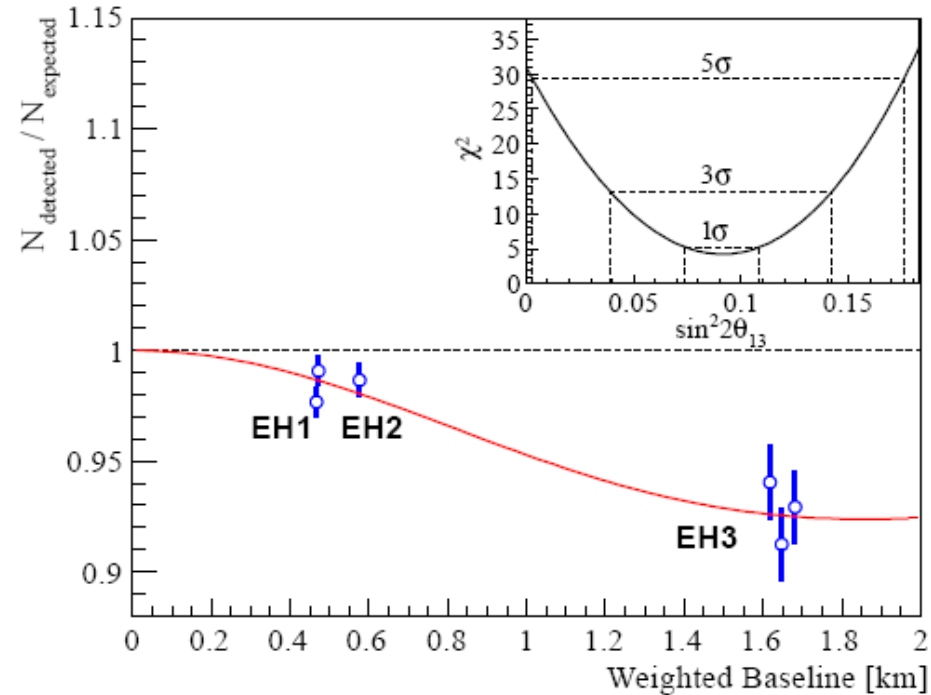
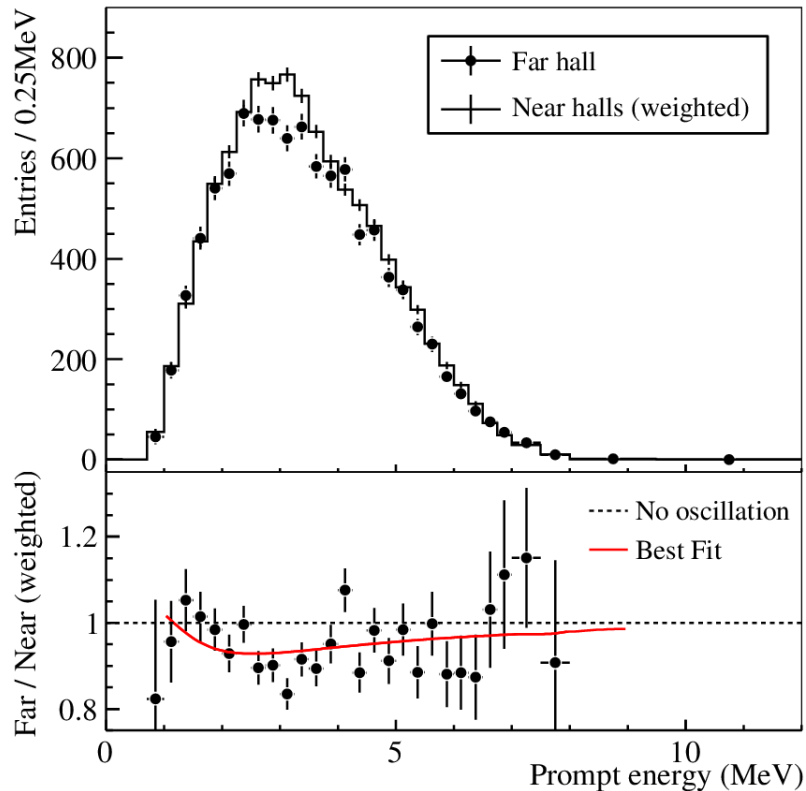


A New Type of Oscillation

◆ Electron anti-neutrino disappearance:

$$R = 0.940 \pm 0.011 \text{ (stat)} \pm 0.004 \text{ (syst)}$$

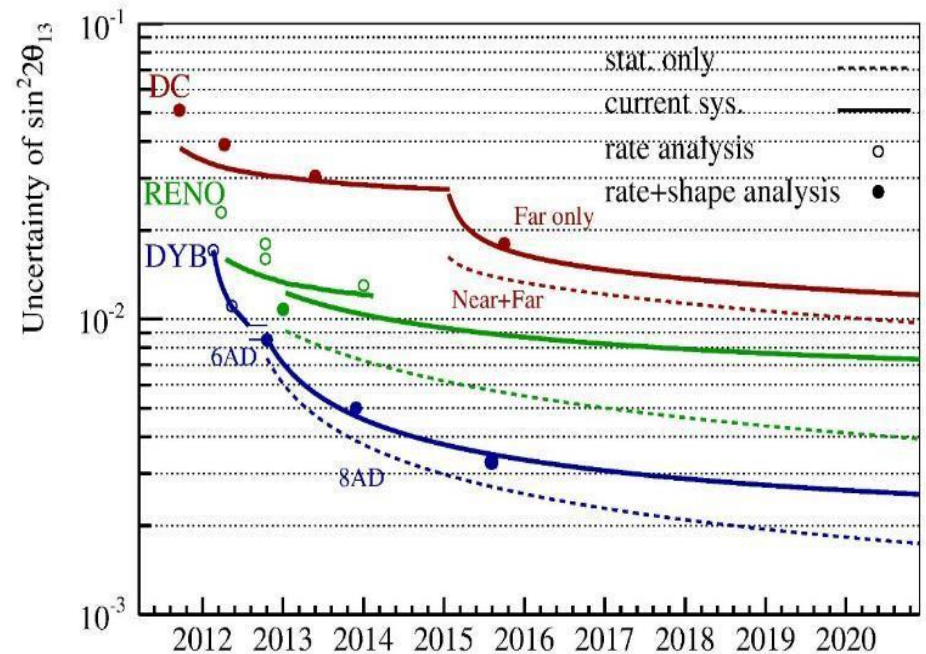
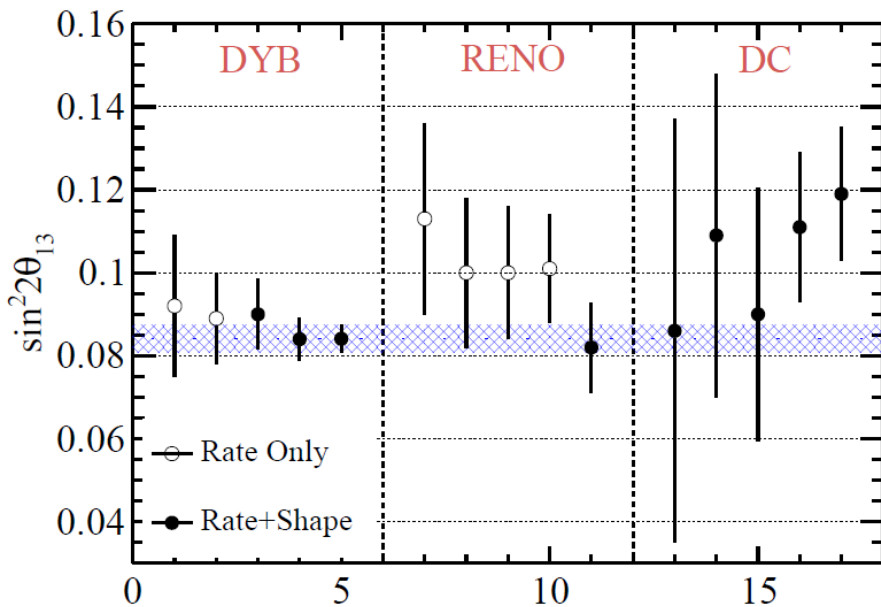
F.P. An et al., Phys.
Rev. Lett. 108, (2012)
171803



$$\sin^2 2\theta_{13} = 0.092 \pm 0.016 \text{ (stat)} \pm 0.005 \text{ (syst)}$$
$$\chi^2/\text{NDF} = 4.26/4, \quad 5.2 \sigma \text{ for non-zero } \theta_{13}$$

Oscillation Results and Prospects

- ◆ $\sin^2 2\theta_{13}$ Precision improved from $\sim 20\%$ to $\sim 4\%$
- ◆ Continue to operate until 2020, to measure both $\sin^2 2\theta_{13}$ and Δm^2_{ee} to $< 3\%$

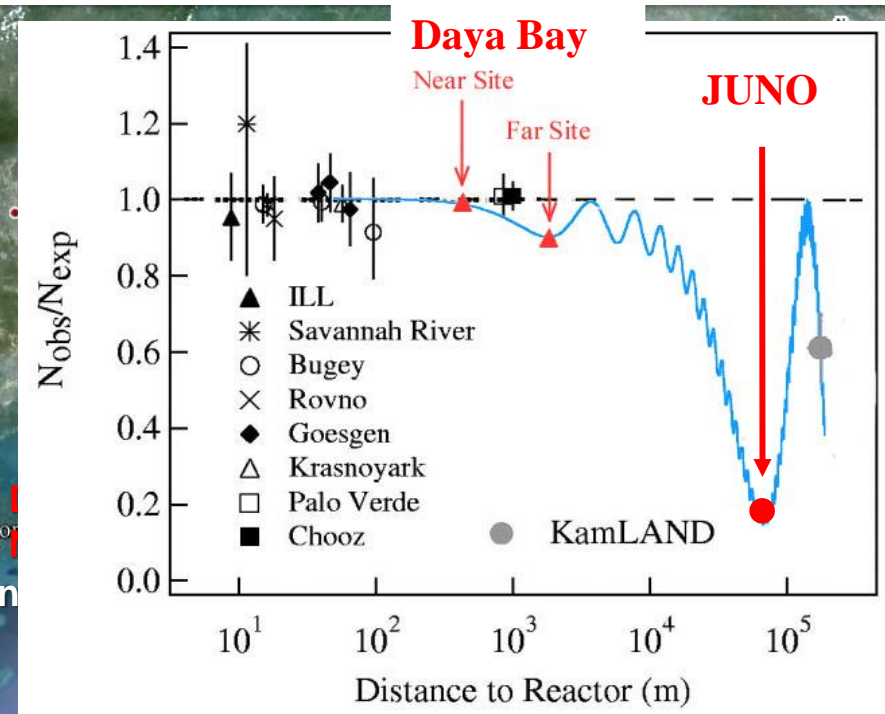


Release time	Data	Config	$\sin^2 2\theta_{13}$	Δm^2_{ee}
2012/3/8 [17]	55 days	6 ADs	$0.092 \pm 0.016 \pm 0.005$	-
2012/10/23 [18]	139 days	6 ADs	$0.089 \pm 0.010 \pm 0.005$	-
2013/10/24 [20]	217 days	6 ADs	$0.090^{+0.008}_{-0.009}$	$2.59^{+0.19}_{-0.20}$
2014/6/24 [19]	217 days	6 ADs (nH)	0.083 ± 0.018	-
2015/5/13 [21]	621 days	6+8 ADs	0.084 ± 0.005	2.42 ± 0.11
2016/3/14 [22]	621 days	6+8 ADs (nH)	0.071 ± 0.011	-
2016/7/5 [23]	1230 days	6+8 ADs	0.0841 ± 0.0033	2.50 ± 0.08

Next Step: JUNO Experiment

NPP	Daya Bay	Huizhou	Lufeng	Yangjiang	Taishan
Status	Operational	Planned	Planned	Under construction	Under construction
Power	17.4 GW	17.4 GW	17.4 GW	17.4 GW	18.4 GW

Overburden ~ 700 m



by 2020: 26.6 GW

Physics at JUNO

Mass Hierarchy: For 6 years, JUNO can determine the MH independent to CP phase with a significance:

	Relative Meas.	Use absolute Δm^2
Ideal case	4σ	5σ
Realistic case	3σ	4σ

Other Physics topics:

- Supernova neutrinos
- Geo-neutrinos
- Solar neutrinos
- Double beta decays

Precision Measurement of Mixing Parameters: Probing the unitarity of U_{PMNS} to $\sim 1\%$ level !

	Current	JUNO
Δm^2_{12}	3%	0.6%
Δm^2_{23}	3%	0.6%
$\sin^2\theta_{12}$	4%	0.7%
$\sin^2\theta_{23}$	11%	N/A
$\sin^2\theta_{13}$	10%	-

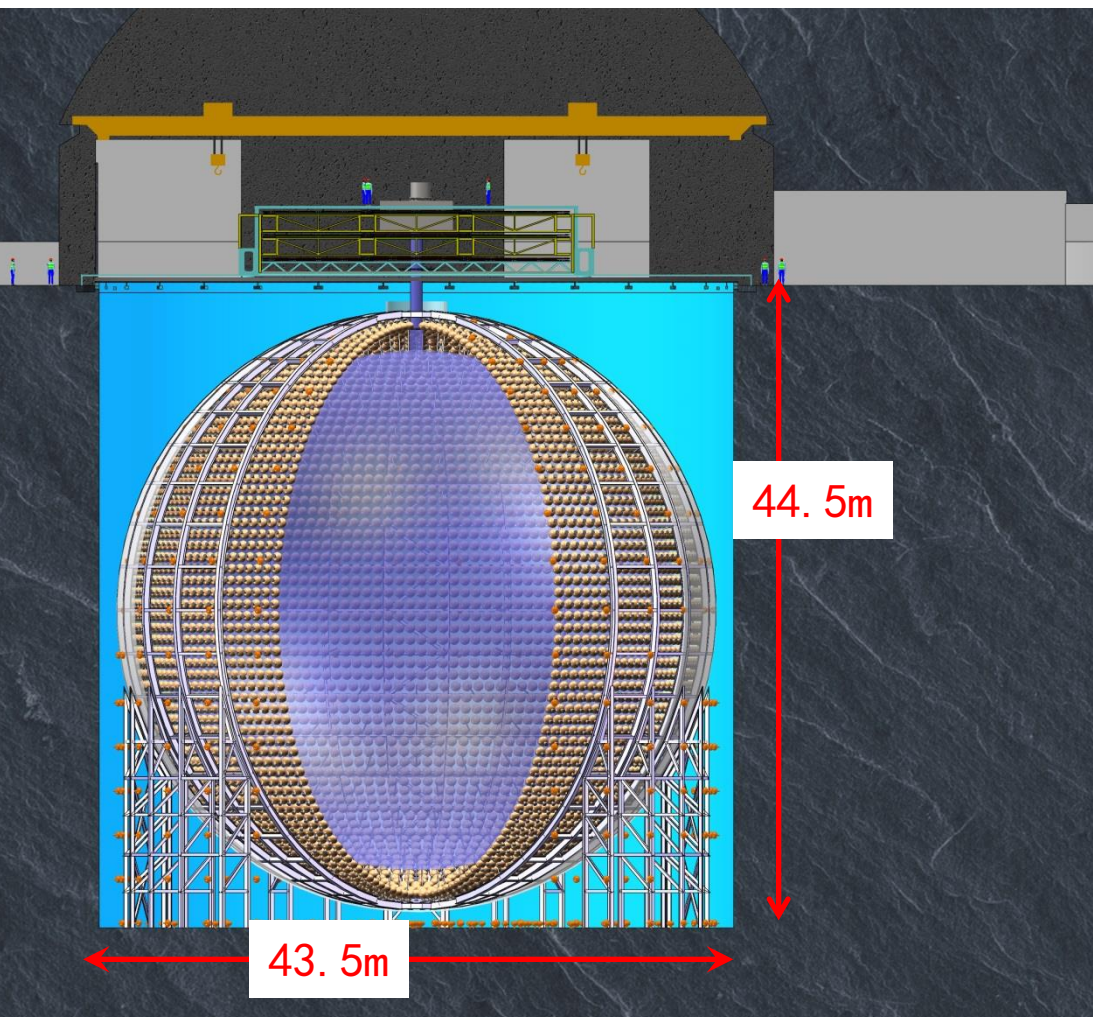
More precise than CKM matrix elements !

J. Phys. G 43: 030401 (2016)
(arXiv:1507.05613)



JUNO Detector and Challenges

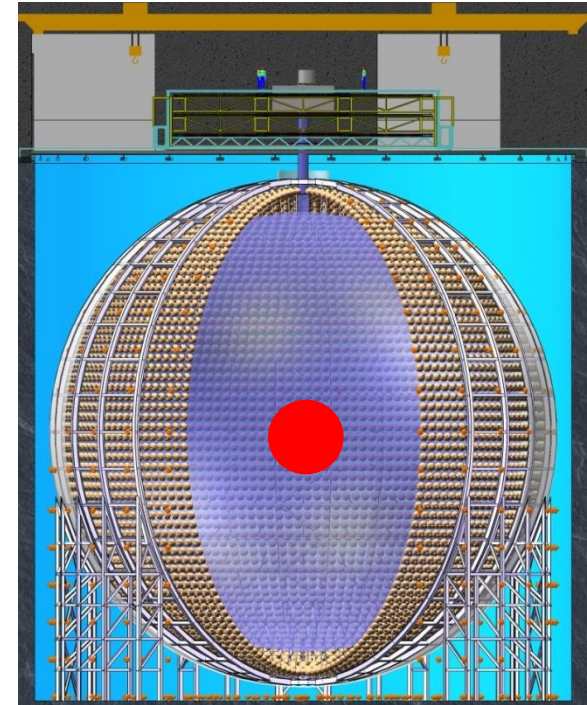
- Largest LS detector → × 20 KamLAND, × 40 Borexino
- Highest light yield → × 2 Borexino, × 5 KamLAND



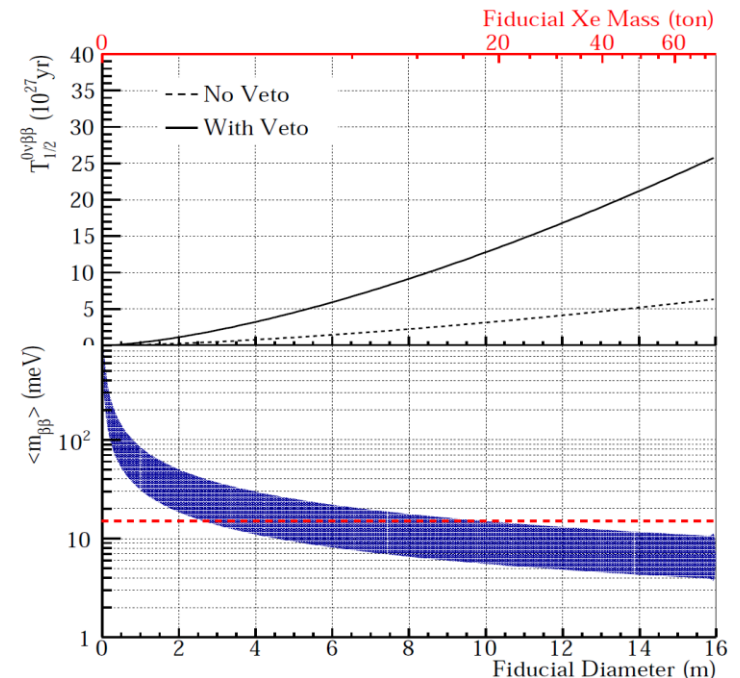
- Hugh cavern:
 - ~ 48m × 70m
- Largest Acrylic tank:
 - Φ 35.4米 (13m @ SNO)
- 20 kt LS
 - Best attenuation length:
25m (15m @ Daya Bay)
- 20000 20" PMT
 - Highest photon detection efficiency : $30\% * 100\% = 30\%$ ($25\% * 60\% = 15\%$ @ SuperK)

JUNO- $\beta\beta$

- ◆ Insert a balloon filled with ^{136}Xe -loaded LS (or ^{130}Te) into the JUNO detector
- ◆ Cosmic-induced backgrounds are removed by cutting a volume around the muon track
- ◆ Yes, sensitivity does scale with the mass



	Isotopes	Mass(t)	$\langle m_{\beta\beta} \rangle, \text{meV}$
nEXO	^{136}Xe	5	7-22
GERDA	^{76}Ge	1	10-40
Majorana	^{76}Ge	1	10-40
SNO+	^{130}Te	8	19-46
KamLAND -Zen	^{136}Xe	1	~ 20
JUNO-$\beta\beta$	^{136}Xe	50	4-12



JUNO Collaboration



17 Countries & regions, 72 institution, 580 members



Europe (28)

Belgium(1)

ULB

Czech(1)

Charles U

Latvia(1)

IECS

Finland(1)

U.Oulu

France(5)

APC Paris

CPPM Marseille

IPHC Strasbourg

Subatech Nantes

CENBG-IN2P3

Italy(8)

INFN-Catania

INFN-Frascati

INFN-Ferrara

INFN-Milano

INFN-Mi-Bicocca

INFN-Padova

INFN-Perugia

INFN-Roma 3

Germany(7)

FZ Jülich

RWTH Aachen

TUM

U.Hamburg

IKP FZI Jülich

U.Mainz

U.Tuebingen

Russia(3)

INR Moscow

JINR

MSU

Slovakia (1)

FMPICU

America(6)

US(2)

UMD

UMD-Geo

Chile(2)

PCUC

UTFSM

Brazil (2)

PUC-Rio

UEL

China (33)

BJ Nor. U.

CAGS

Chongqing U.

Shanghai JT U.

DGUT

ECUST

Guangxi U.

HIT

IHEP

U. Of South China

Ninan U.

Nanjing U.

Natl. Chiao-Tung U.

Natl. Taiwan U.

Natl. United U.

Asia (38)

Nankai U.

NCEPU

Pekin U.

SDU

Sichuan U.

CIAE

SYSU

Tsinghua U.

UCAS

USTC

Jilin U.

Wuhan U.

Wuyi U.

Xi'an JT U.

Xiamen U.
NUU.

Armenia(1)

Yerevan Phys. Inst.

Thailand(3)

SUT

PPRLCU

NARIT

Parkistan(1)

PINSTECH

Detector R&D and Construction

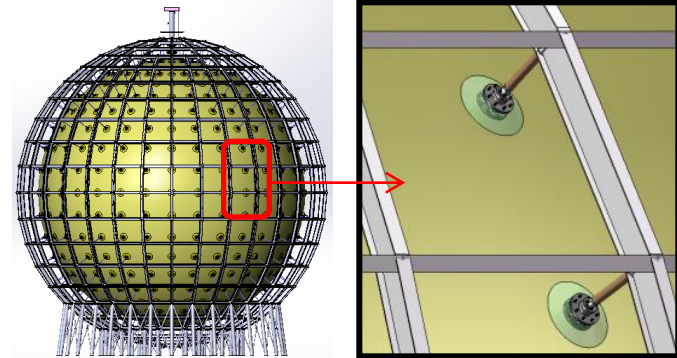
◆ A huge acrylic tank:

- ⇒ R&D and design completed, procedures of manufacturing understood
- ⇒ Prototypes underway
- ⇒ Contract signed



◆ A SS structure to hold the acrylic tank and to mount PMTs

- ⇒ Design completed, contract signed
- ⇒ Issues like tem. change, earth quake, Assembly & installation understood



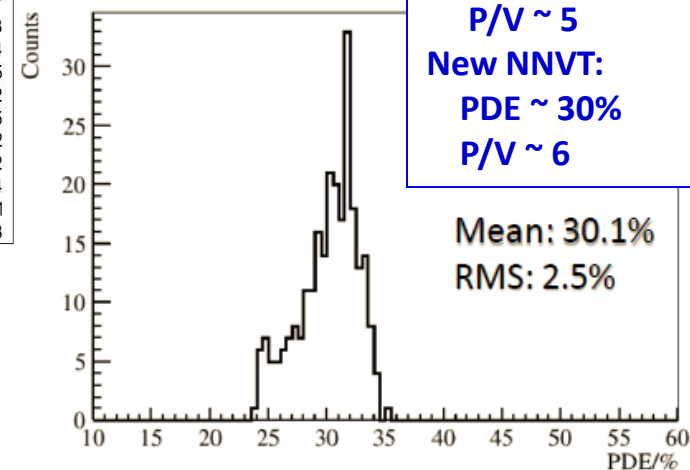
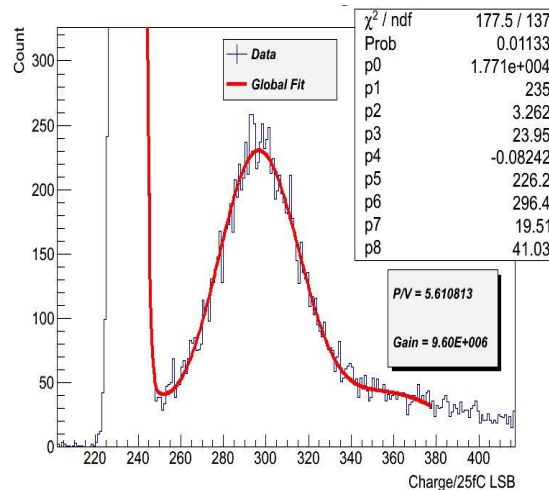
◆ 20 kt liquid scintillator

- ⇒ Completed the optimization of LS for the best light yield and attenuation length
- ⇒ Completed the 20t test of purification: attenuation length > 20 m.
- ⇒ Radio-purity under testing



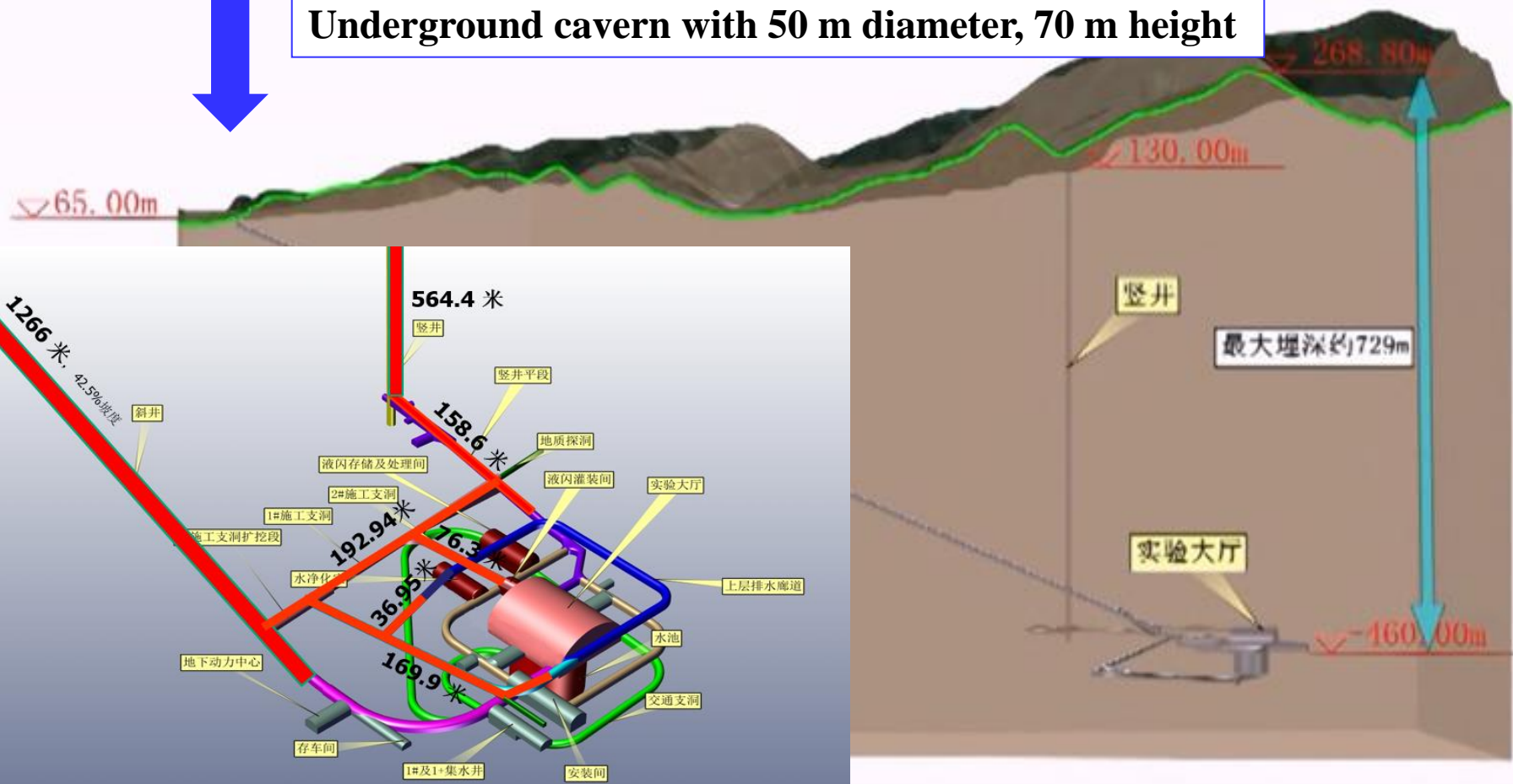
Photomultipliers

- ◆ Large, High QE PMT's is the key to obtain sufficient photons
- ◆ Efforts started in 2009 to develop MCP-PMT, with a goal of QE > 35%
- ◆ Partnered with companies and research institutes: NNVC, XIOPM, etc.
- ◆ Successful development:
 - ⇒ NNVC: $QE(27\%)*DE(100\%) > 27\%$
 - ⇒ Hamamatsu: $QE(30\%)*DE(90\%) > 27\%$
- ◆ Contract signed(based on quality, availability, cost, risk et al.)
 - ⇒ 15000 from NNVC, 5000 from Hamamatsu
- ◆ More than 6000 MCP-PMT from NNVT and 4000 PMT from Hamamatsu
 - ⇒ Tubes are better than spec. of the contract



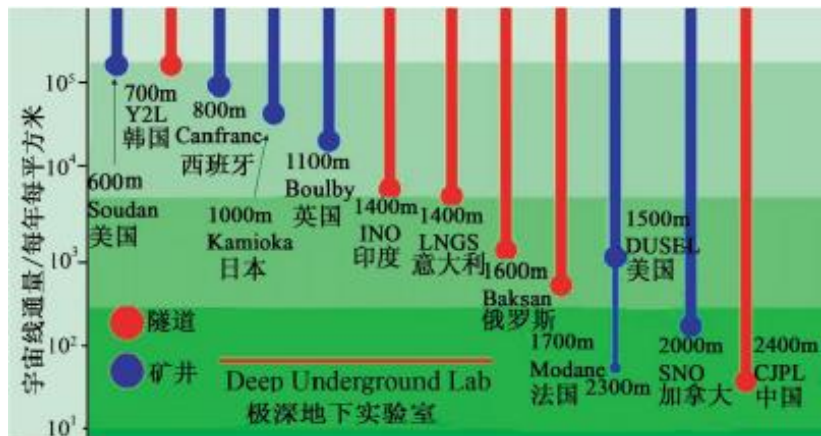


~ 600m vertical shaft, ~1300m sloped tunnel
 Underground cavern with 50 m diameter, 70 m height

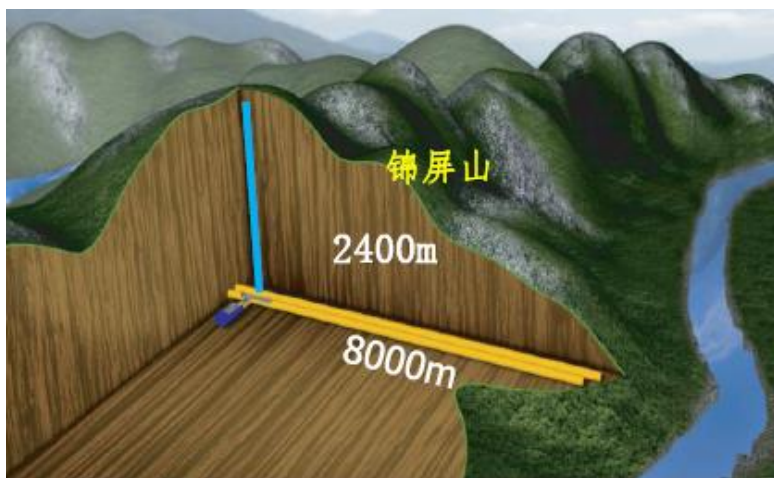
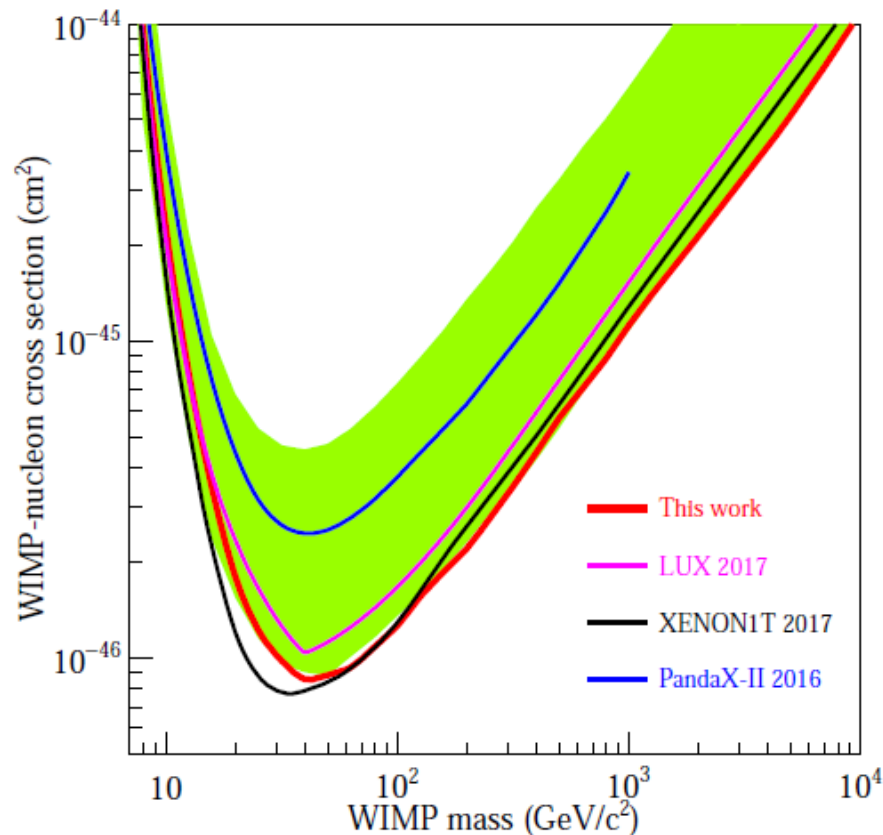


JinPin Underground Laboratory

- The deepest underground laboratory in the world: 2400 m
- Current experiments: dark matter searches
 - Xe-based PandaX
 - Ge-based CDEX

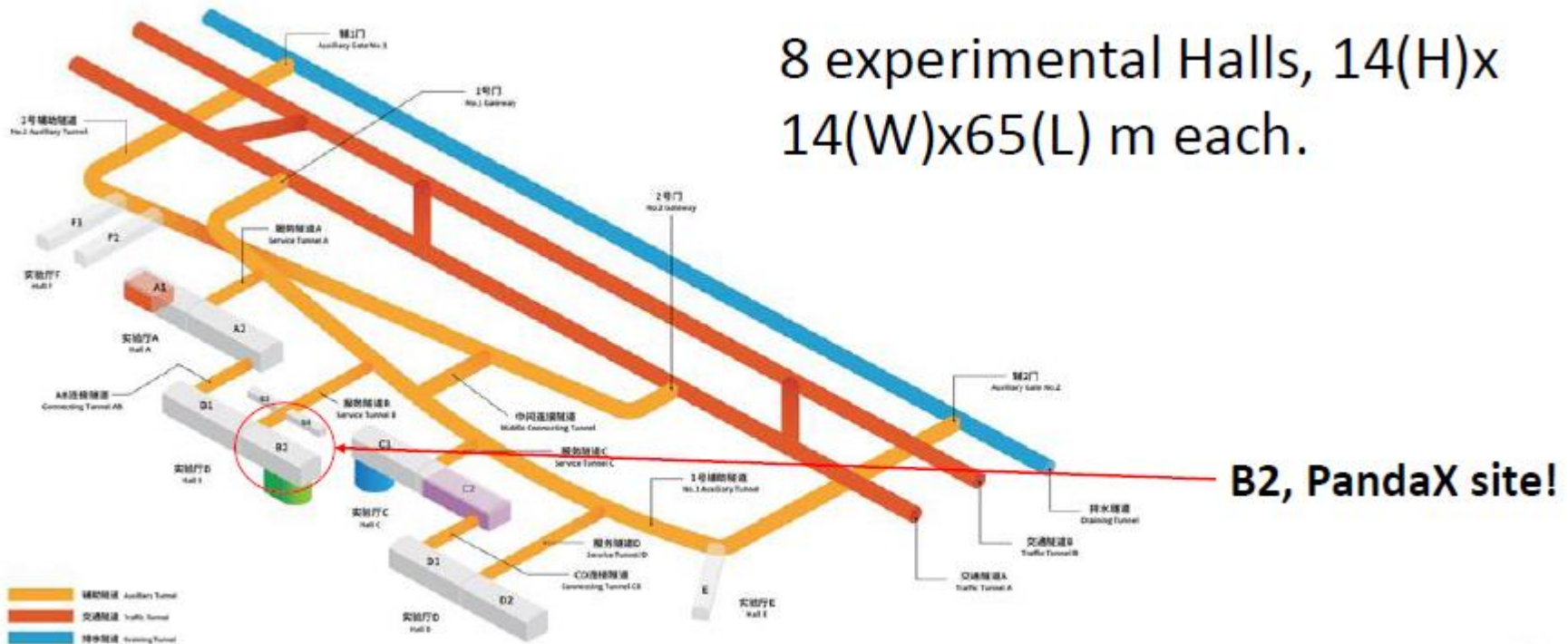


PANDAX-II: PRL 119(2017)181302



Future

- Approved for the infrastructure construction
- Next generation Xe- & Ge-based dark matter & $\beta\beta$ decay searches
- Other possibilities: Solar neutrinos, geoneutrinos, etc.



Yangbajing Cosmic-ray Observatory: Since 90's

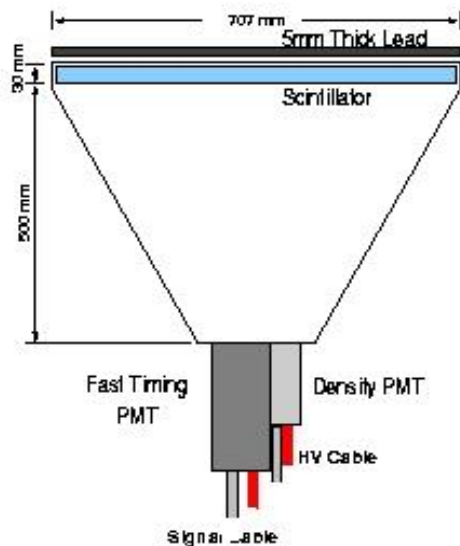


Sino-Japanese AS γ experiment (scintillation detector array)

Sino-Italian ARGO experiment (RPC hall)

ASy scintillation detector

Sino-Italian ARGO experiment (part of RPC carpet)

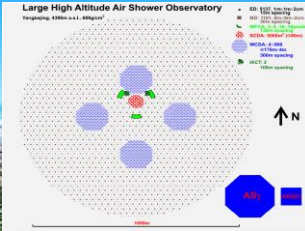
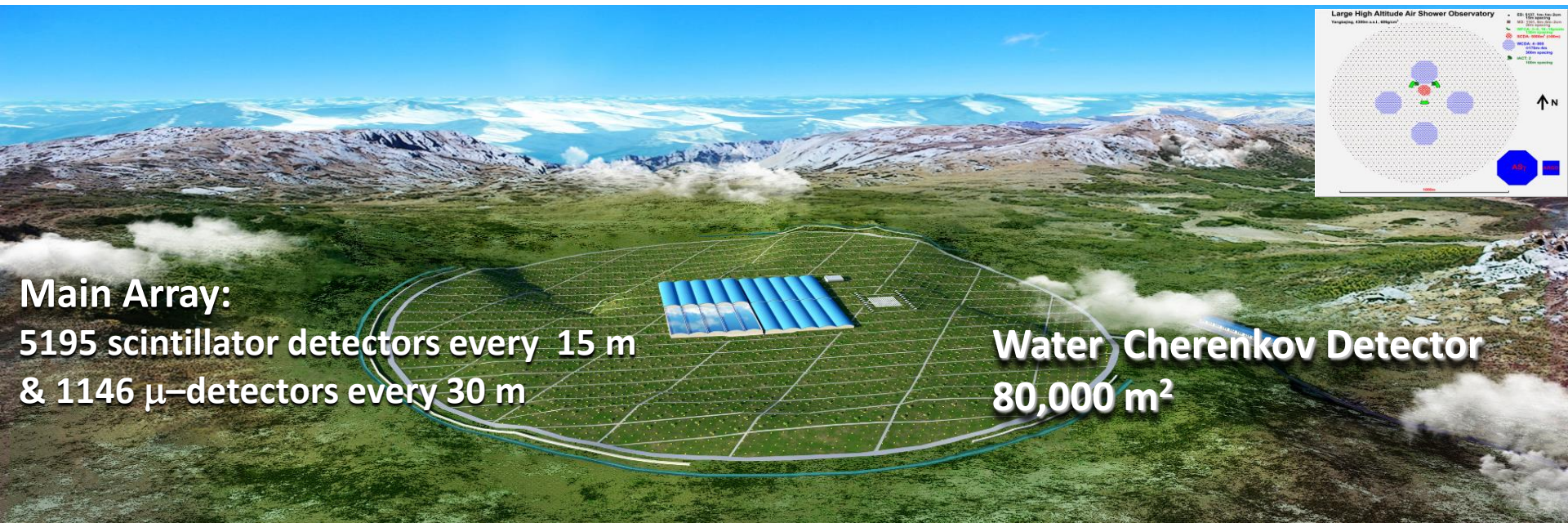


A new Phase: LHAASO

— Large High Altitude Air Shower Observatory



Sichuan, 4300 m a.s.l.



Main Array:

5195 scintillator detectors every 15 m
& 1146 μ -detectors every 30 m

Water Cherenkov Detector
80,000 m²

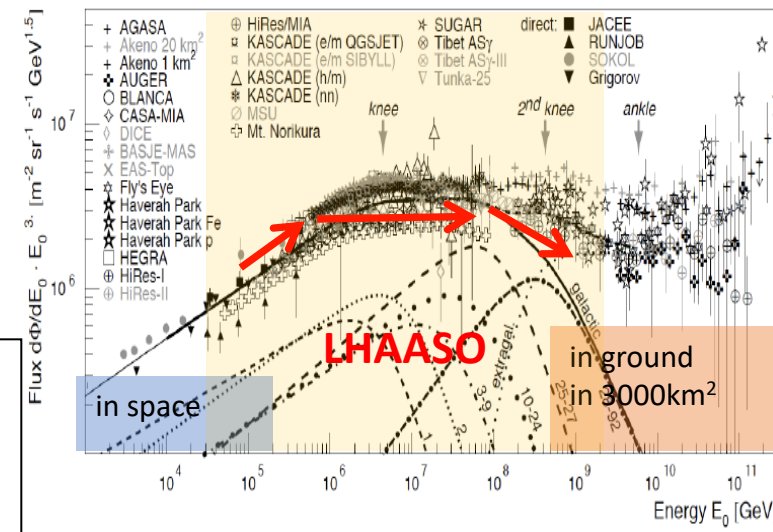
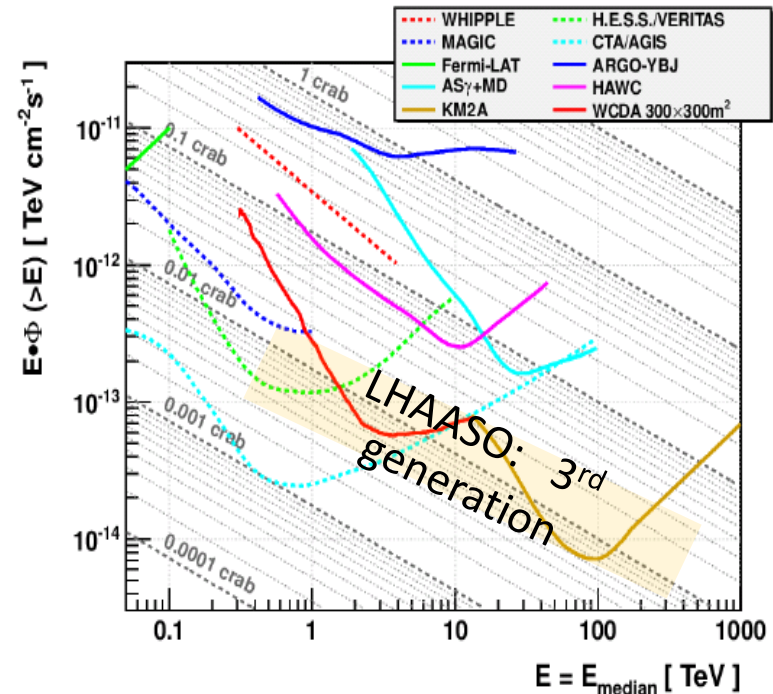
Science at LHAASO

- Unique for 10 TeV gamma astronomy with the highest sensitivity in the world
- Window for discovering the hadronic origins of cosmic rays
- Crucial CR data covering a very wide energy region of knees
- Exploring for new physics, such as DM or quantum gravity

Complementary to CTA:

- All the time
- All the sky
- Time-variant sources
- Extended sources
- Fast indication for CTA

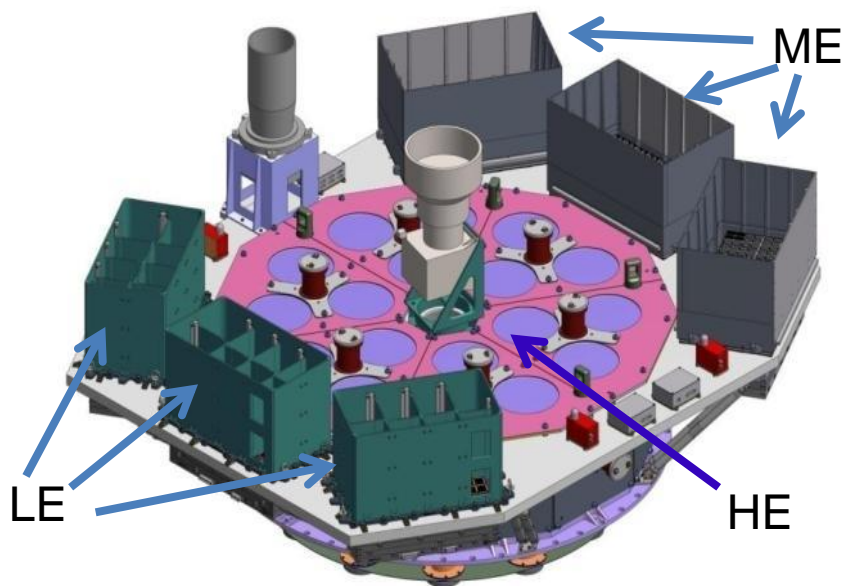
LHAASO is funded for 2016-2020, 1/4 of the detector array will be operational by end of this year



Two Scientific Satellites on orbit now

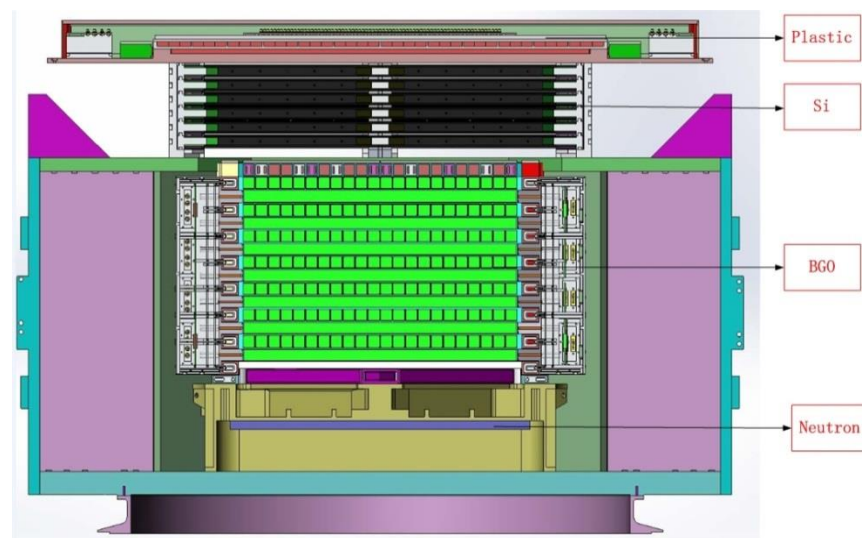
Hard X-ray modulated telescope (HXMT)

- Full sky survey with good angular resolution and sensitivity
- Launched in June, 2017



DARk Matter Particle Explorer (DAMPE)

- Cosmic-ray & gamma detector up to 10 TeV with good resolution
- Launched in Dec., 2015



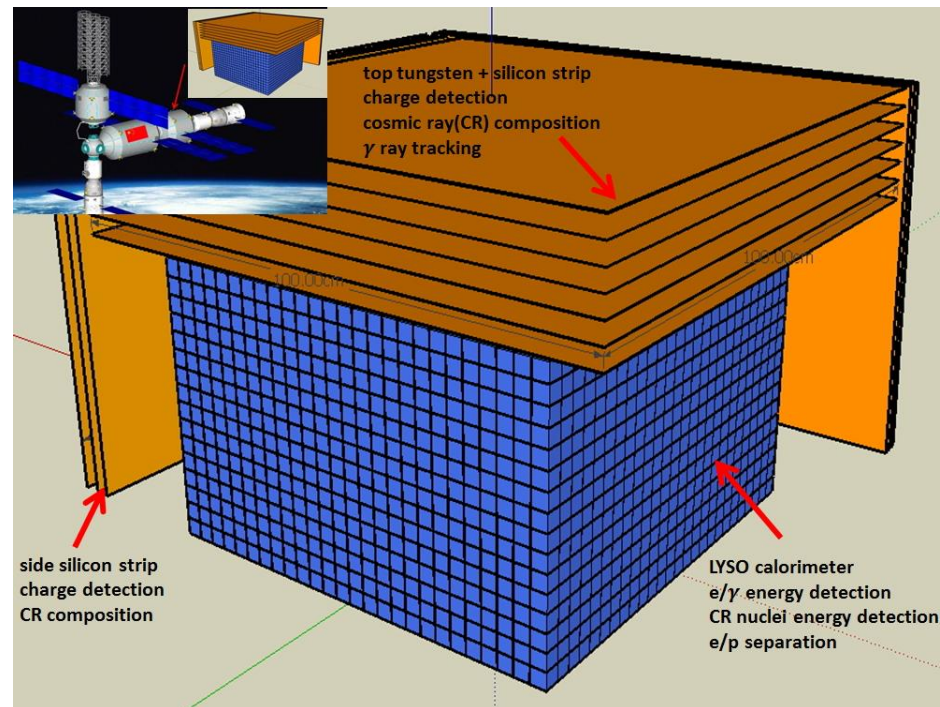
HERD @ the China's Space Station

• Science

- Dark matter search: γ from 100 – 10,000 GeV
- γ -ray astronomy: GRBs, microquasars, Blazars and other transients down to 100 MeV
- Spectral and composition measurements of CRs between 300 GeV to PeV with a large geometrical factor
- Complementary to LHAASO: directly measured composition & spectrum in space

• Status

- Groups from China, Italy, Switzerland, Sweden,...
- Launch in ~2023



	$X0(\lambda)$	$\Delta E/E$ for e	e/p sep	e GF $m^2sr@$ 200GeV	p GF $m^2sr@10$ 0TeV
HERD (2020)	55(3)	1%	10^{-6}	3.1	2.3
Fermi (2008)	10	12%	10^{-3}	0.9	--
AMS02 (2011)	17	2%	10^{-6}	0.12	--
DAMPE (2015)	31	1%	10^{-4}	0.3	--
CREAM (2015)	20(1.5)	--	--	--	0.2

Acceptance & H-energy > 10X all others

Summary

- **Particle Physics is a great field**
 - Incredible success in the past
 - More to come in the future
- **China is catching up, thanks to its economic growth**
 - Great success on BESIII, Daya Bay and PandaX
 - Bright prospects for JUNO and **CEPC**
 - More active in international projects: LHC(all 4 exp.s), BELLE/BELLEII, COMET, EXO, LBNF, Panda, GlueX, ...
 - More to come in Astro-particle physics: DAMPE, HXMT, LHAASO, **HERD**,...
- **Looking forward to collaborate with the world**