

## **(n,cp) reactions study at the n\_TOF facility at CERN: results for the Cosmological Lithium problem**

M. Barbagallo<sup>1,2</sup>, on behalf of the n\_TOF Collaboration<sup>2</sup>

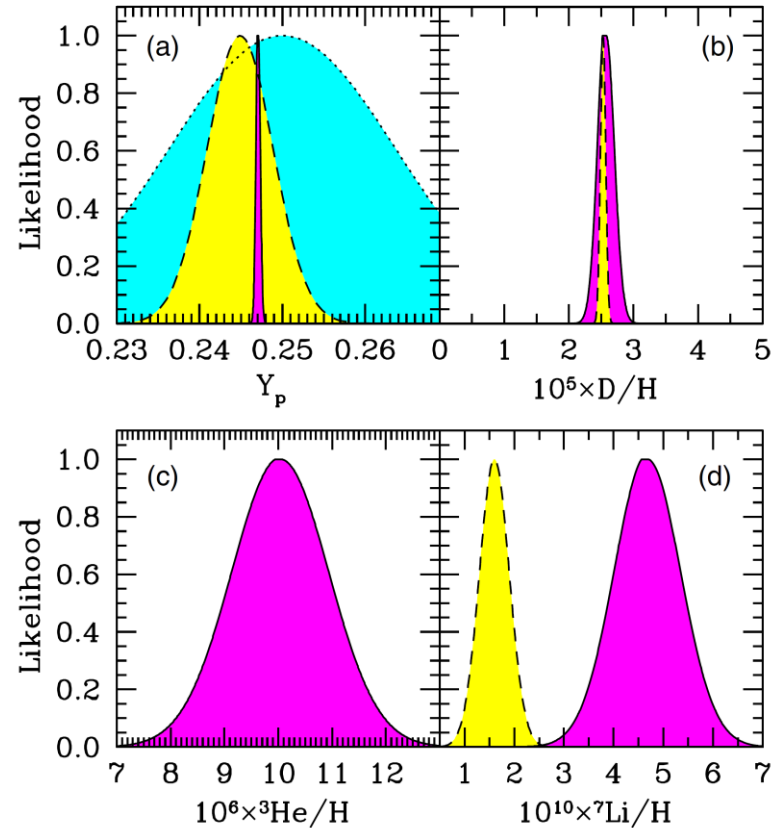
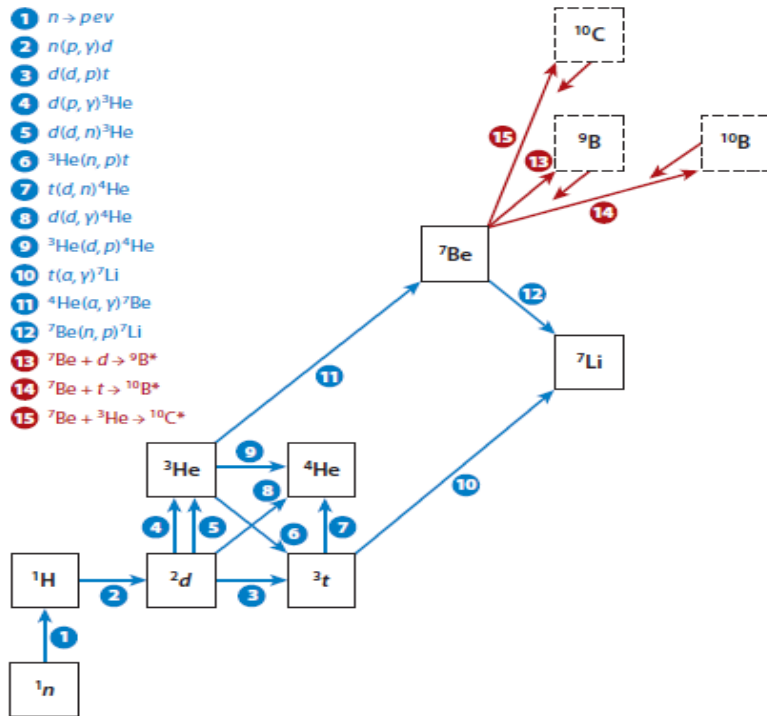
1-Istituto Nazionale di Fisica Nucleare, sez. di Bari

2-CERN

- The Cosmological Lithium Problem
- The  ${}^7\text{Be}(n,\alpha)$  and the  ${}^7\text{Be}(n,p)$  cross section measurements
- Implications for Nuclear Astrophysics
- Conclusions and Perspectives

# Cosmological Lithium Problem

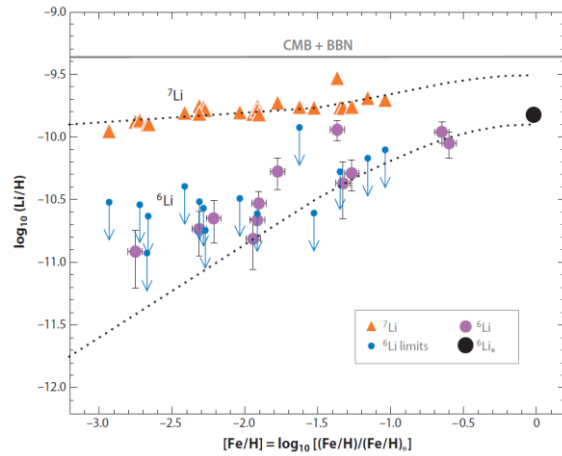
Big Bang Nucleosynthesis successfully predicts the abundances of light elements, i.e. D and  $^4\text{He}$ , but...



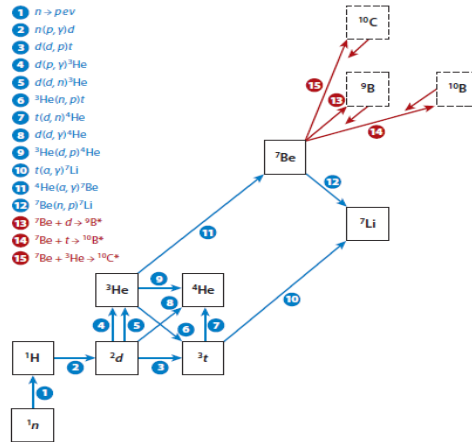
**Serious discrepancy** between the predicted abundance of  $^7\text{Li}$  and the value inferred by measurements (Spite et al.)  $\longrightarrow$  **Cosmological Lithium problem (CLiP)**

(At least) Three classes of solutions for this longstanding problem:

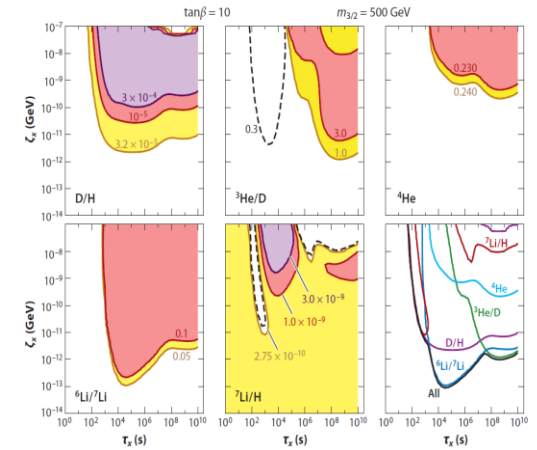
## Astrophysical



## Nuclear Physics

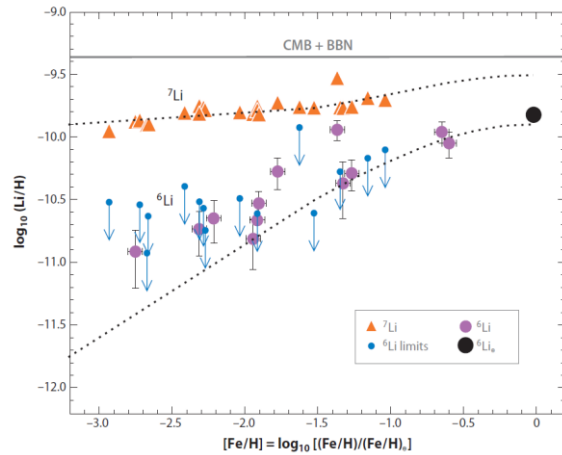


## Non Standard Physics

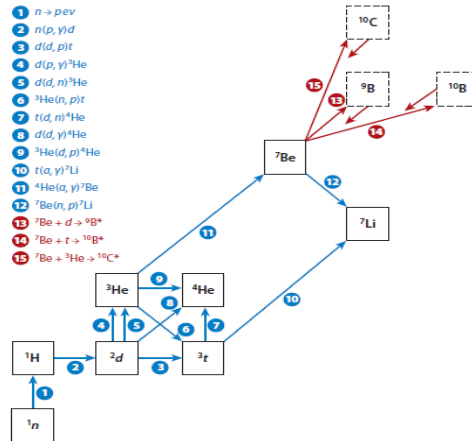


(At least) Three classes of solutions for this longstanding problem:

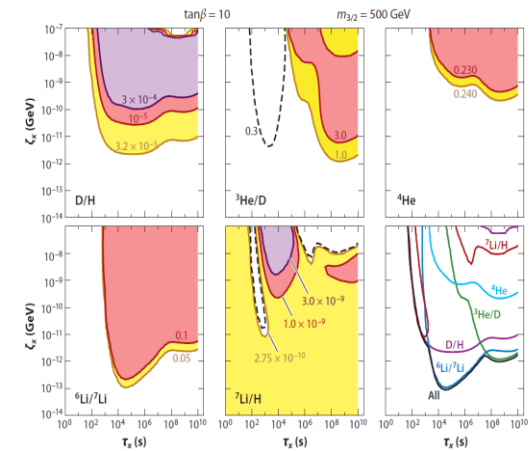
## Astrophysical



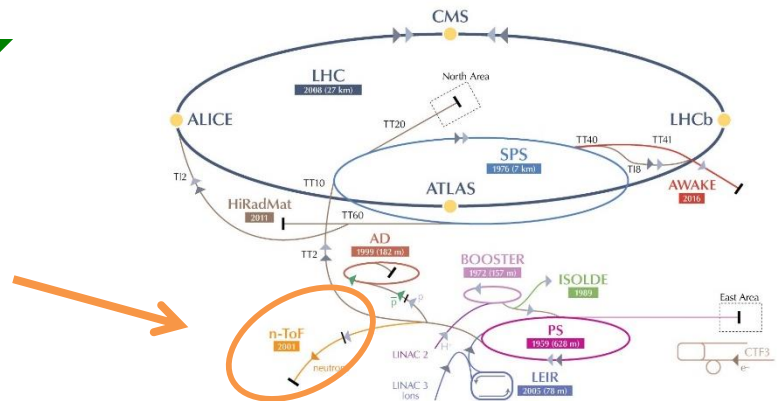
## Nuclear Physics



## Non Standard Physics



Neutron Time Of Flight facility: **n\_TOF**





*"Quand on était jeune.."*



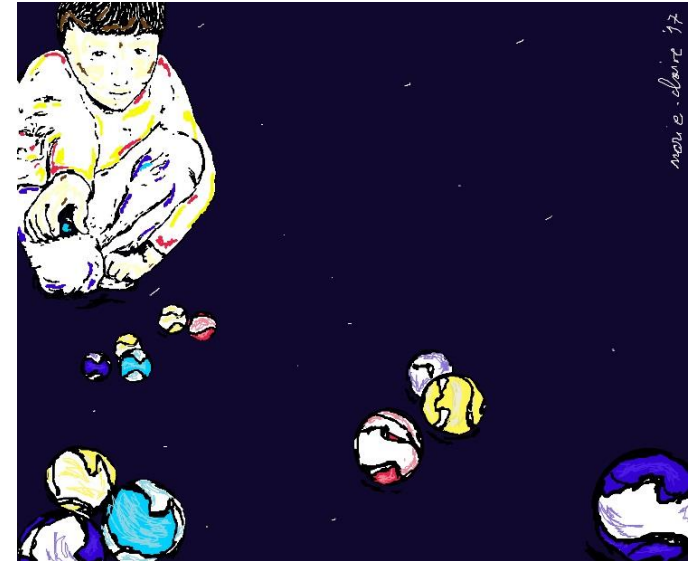
Approximately 95% of primordial  ${}^7\text{Li}$  is produced from the **electron capture decay** of  ${}^7\text{Be}$  ( $T_{1/2}=53.2$  d).

${}^7\text{Be}$  decay rate in plasma(?)

${}^7\text{Be}$  production channels have been widely investigated and they are known with good accuracy.

${}^7\text{Be}$  is destroyed via **(n,p)** and (p,x), (d,x), ( ${}^3\text{He}$ ,x), ... reactions. Small contribution of the **(n, $\alpha$ )** reactions according to **estimated** cross section.

Nuclear Physics  
solution to CLiP

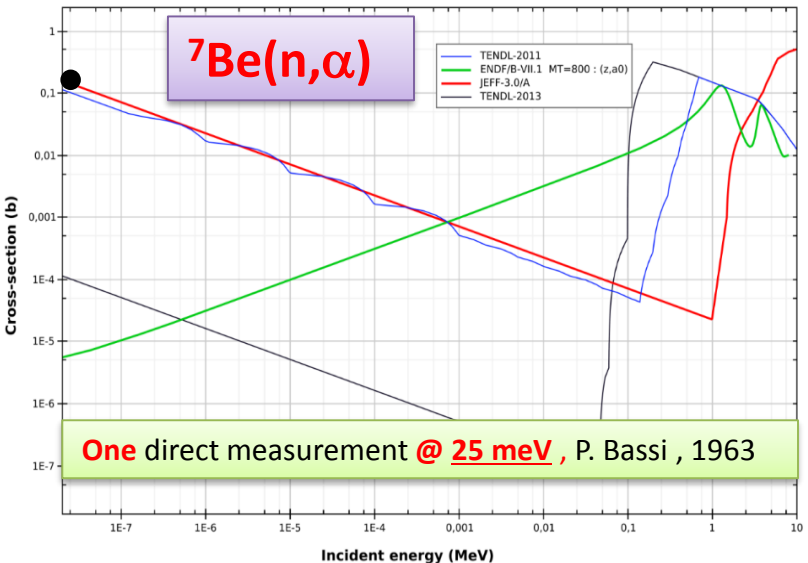
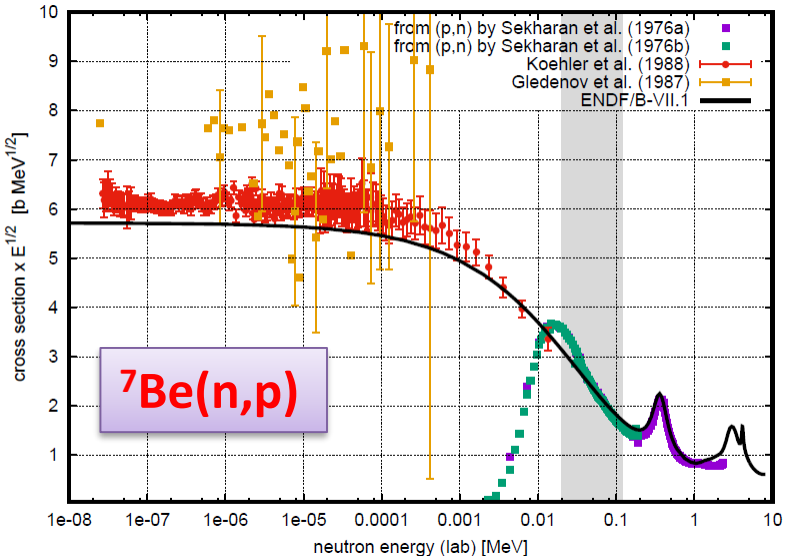
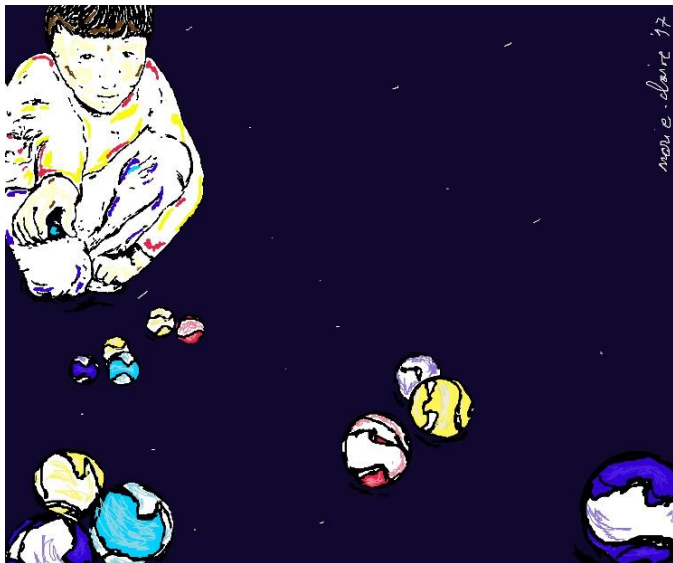


# Cosmological Lithium Problem and $^7\text{Be}$

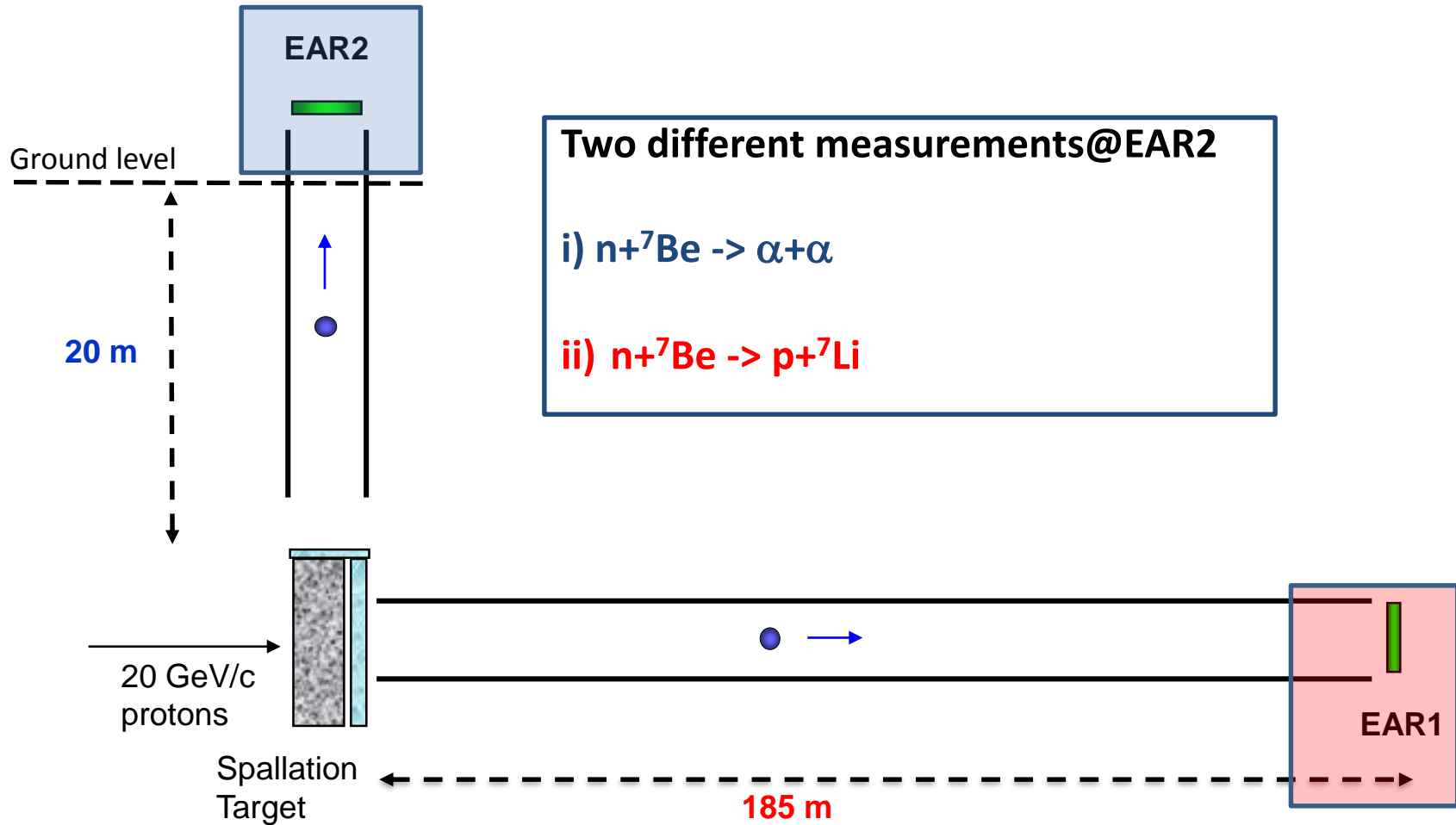
Approximately 95% of primordial  $^7\text{Li}$  is produced from the **electron capture decay** of  $^7\text{Be}$  ( $T_{1/2}=53.2$  d).

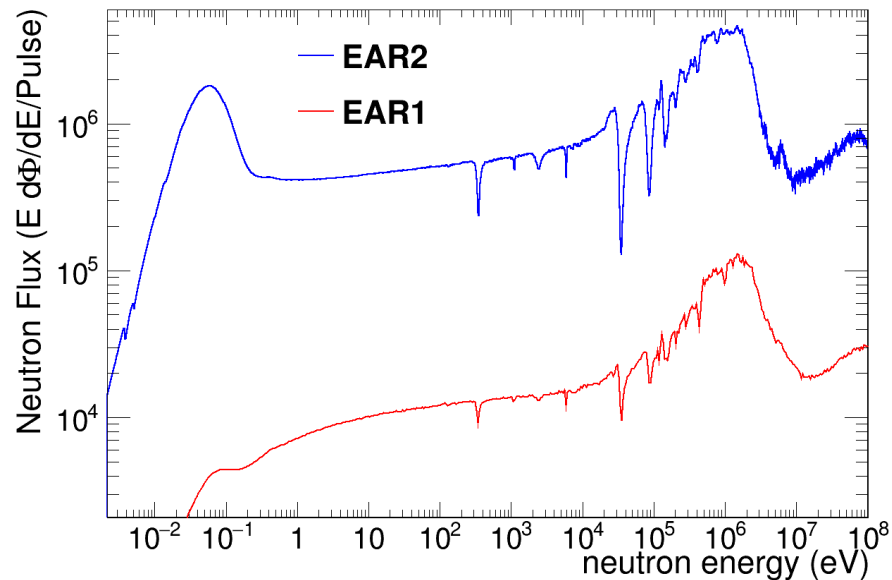
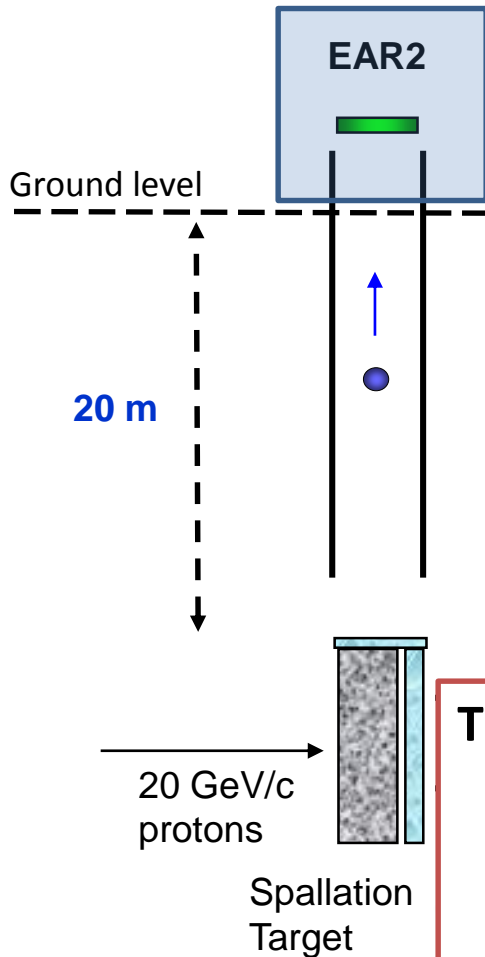
Nuclear Physics solution to CLiP

- $^7\text{Be}$  decay rate in plasma(?)
- $^7\text{Be}$  production channels have been widely investigated and they are known with good accuracy.
- $^7\text{Be}$  is destroyed via **(n,p)** and **(p,x)**, **(d,x)**, **( $^3\text{He}$ ,x)**, ... reactions. Small contribution of the **(n, $\alpha$ )** reactions according to **estimated** cross section.







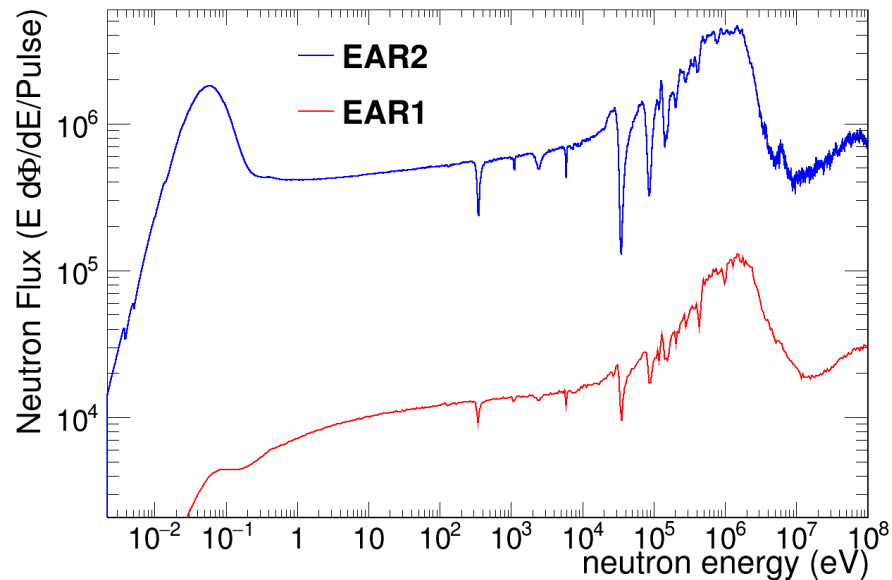
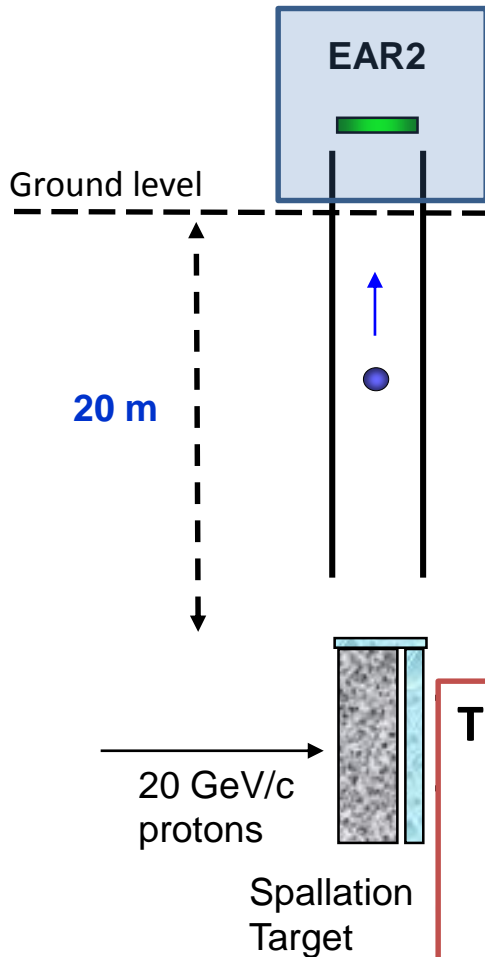


*M. Barbagallo et al (n\_TOF Coll.), Eur. Phys. J. A (2013) 49: 156*

*M. Sabate'-Gilarte et al (n\_TOF Coll.), Eur. Phys. J. A (2017) 53: 210*

## The much higher flux in EAR2 allows to:

- measure samples of **very small mass** ( $\ll 1$  mg)
- measure **short-lived radioisotopes** (i.e. **53.2 d!**)
- collect data on a much **shorter time**
- **measure (n,charged particle) reactions with thin samples**



*M. Barbagallo et al (n\_TOF Coll.), Eur. Phys. J. A (2013) 49: 156*

*M. Sabate'-Gilarte et al (n\_TOF Coll.), Eur. Phys. J. A (2017) 53: 210*

**Be-7**



**13 GBq/ $\mu\text{g}$  !!!**

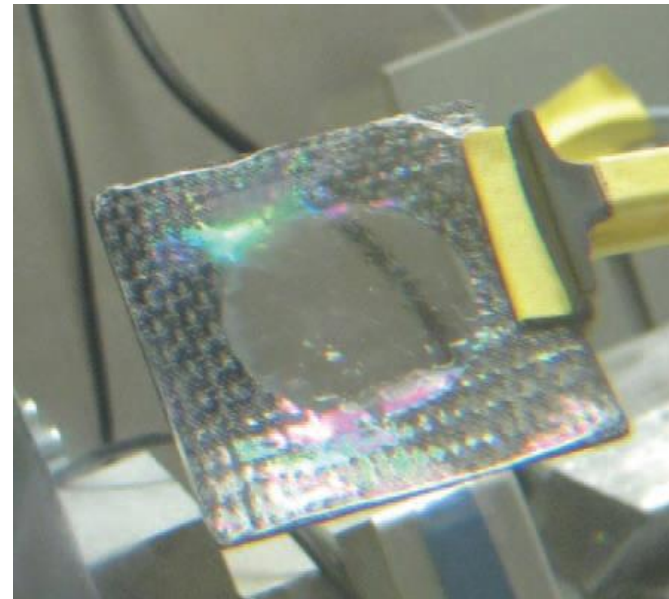
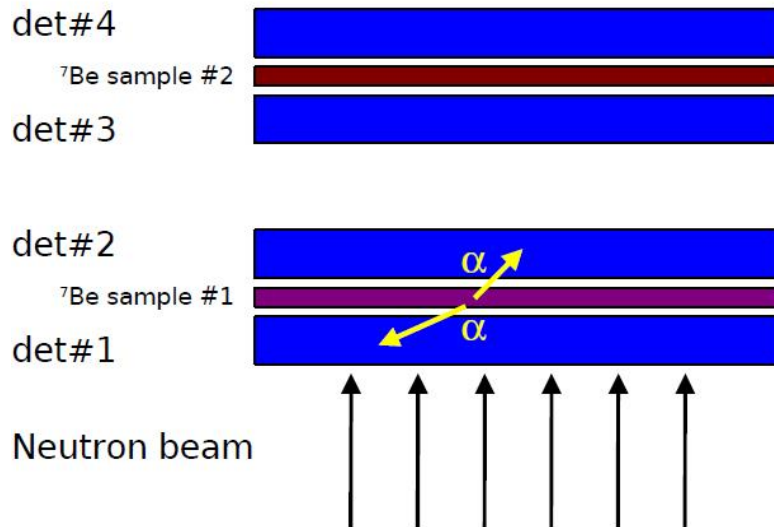
## The much higher flux in EAR2 allows to:

- measure samples of **very small mass** ( $\ll 1$  mg)
- measure **short-lived radioisotopes** (i.e. **53.2 d!**)
- collect data on a much **shorter time**
- **measure (n,charged particle) reactions with thin samples**

# ${}^7\text{Be}(n,\gamma\alpha){}^4\text{He}$ measurement: the making of



- Silicon detectors **directly inserted in the beam**
- Two different samples, 40 GBq total activity (prepared with different and independent techniques)



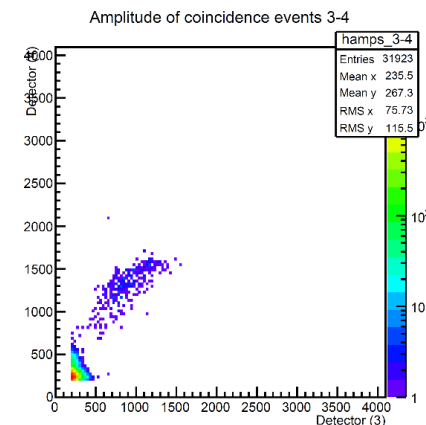
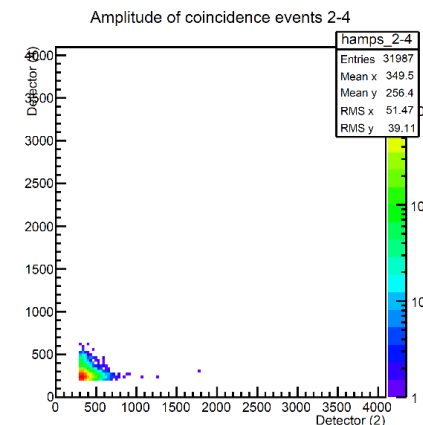
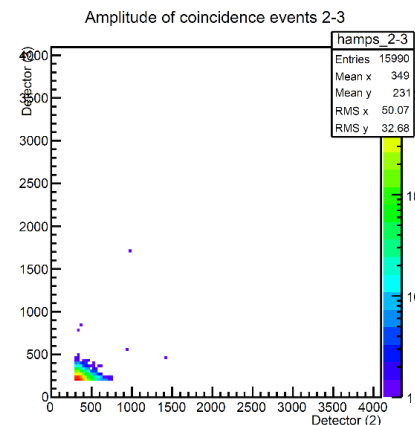
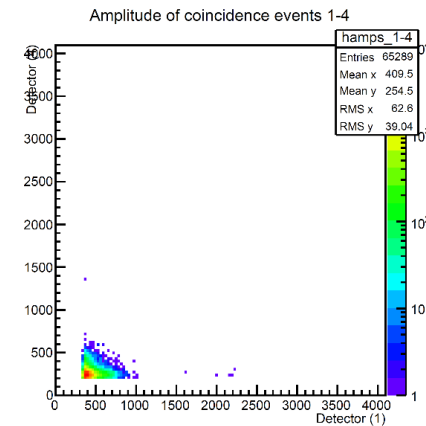
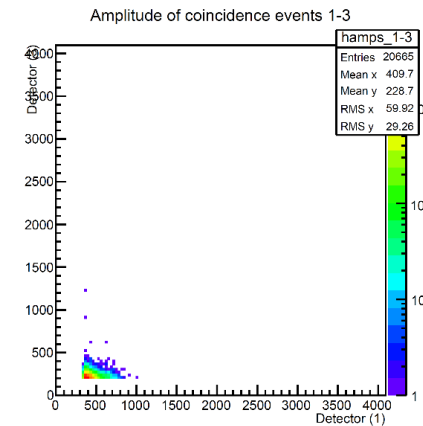
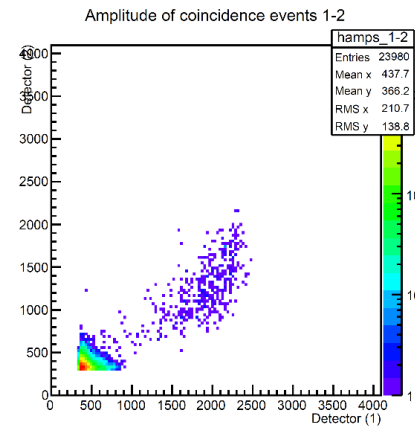
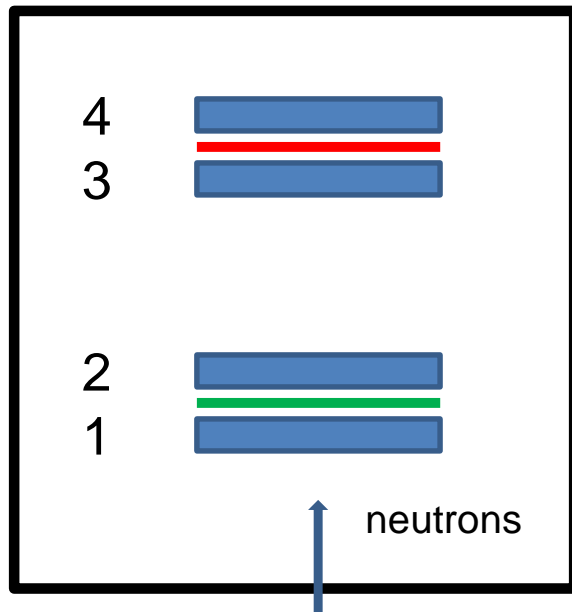
L. Cosentino et al. (*n\_TOF Coll.*), *NIM A* 830 (2016) 197-205

E. Maugeri et al. (*n\_TOF Coll.*), *Journ. of Instr.*, 12, P02016, (2017)

**Such a setup offered, among other features, redundancy, allowing to reduce systematic uncertainties.**

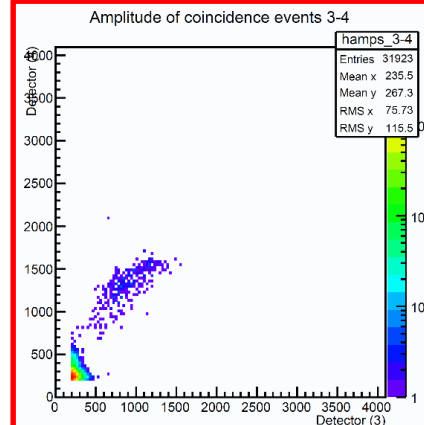
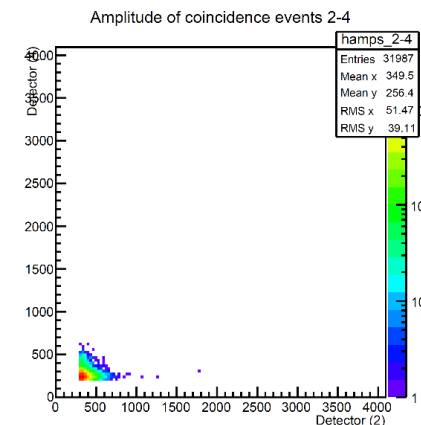
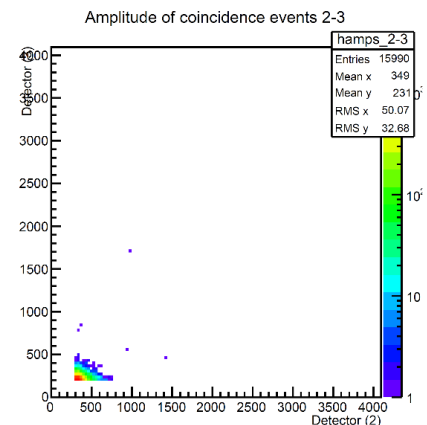
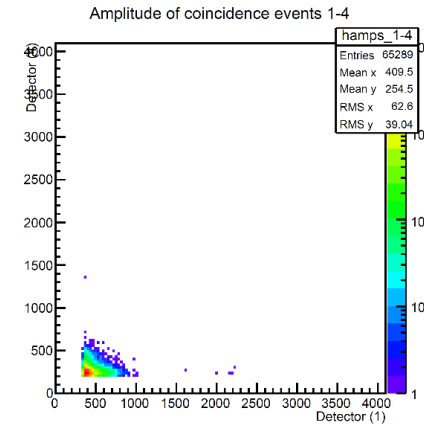
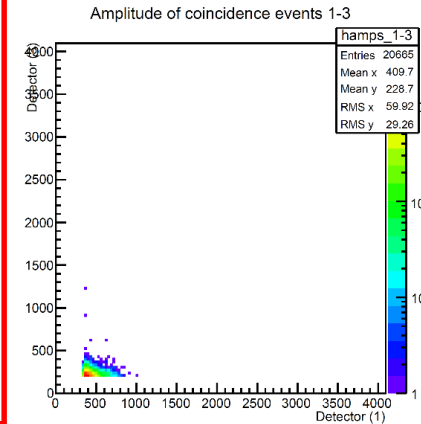
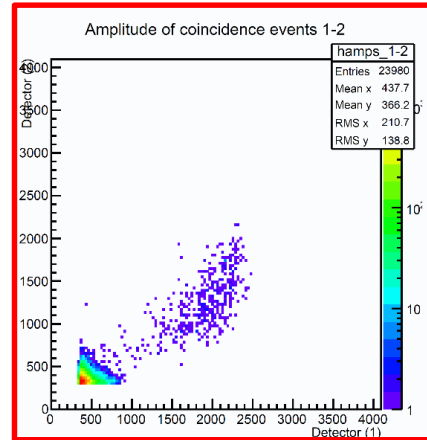
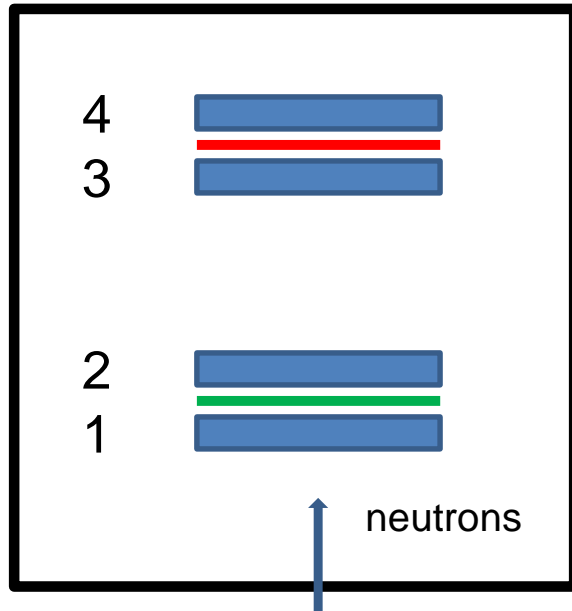
# $^7\text{Be}(n,\gamma\alpha)$ cross-section measurement

Two different sandwiches of silicon detectors.



Possible to evaluate random coincidences comparing uncorrelated couples of detectors.

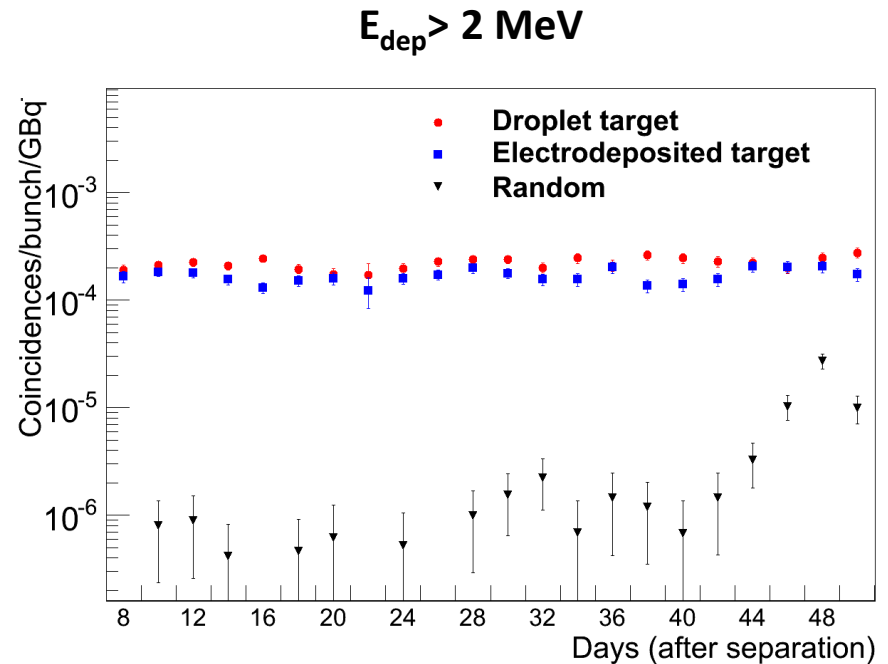
Two different sandwiches of silicon detectors.



Possible to evaluate random coincidences comparing uncorrelated couples of detectors.

**Strong rejection of background: coincidence signals, low duty cycle beam, Time-of-Flight.**

- Protons from  ${}^7\text{Be}(n,p)$  reactions
- $\gamma$  from  ${}^7\text{Be}$  decay
- $n+{}^7\text{Li} \longrightarrow {}^8\text{Li} \xrightarrow{\beta\text{-decay}} {}^8\text{Be}^* (800 \text{ ms})$   
 $\downarrow$   
 $\alpha+\alpha$
- ${}^9\text{Be}(n,2n)$ ,  ${}^7\text{Li}(p,\gamma)$ ,  ${}^7\text{Be}(p,\gamma)$

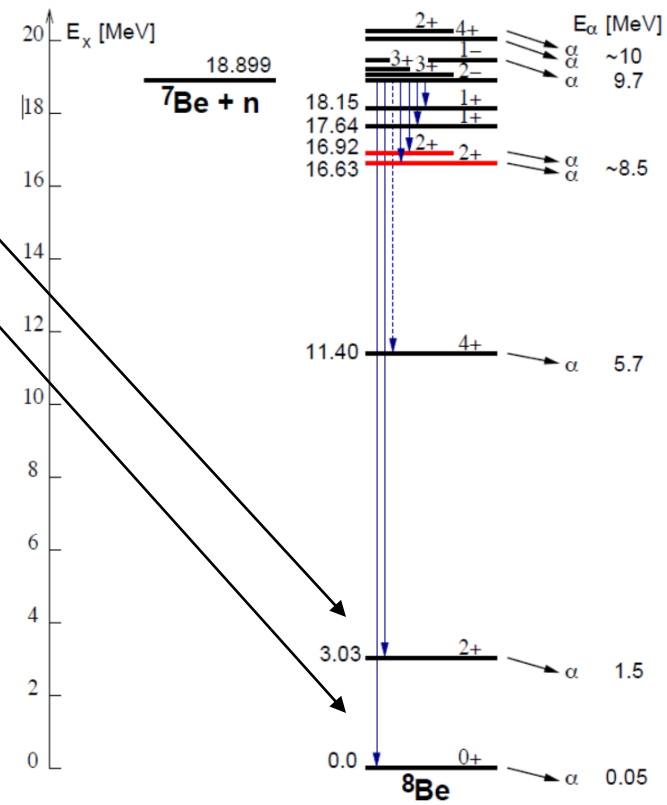


**Strong rejection of background: coincidence signals, low duty cycle beam, Time-of-Flight.**

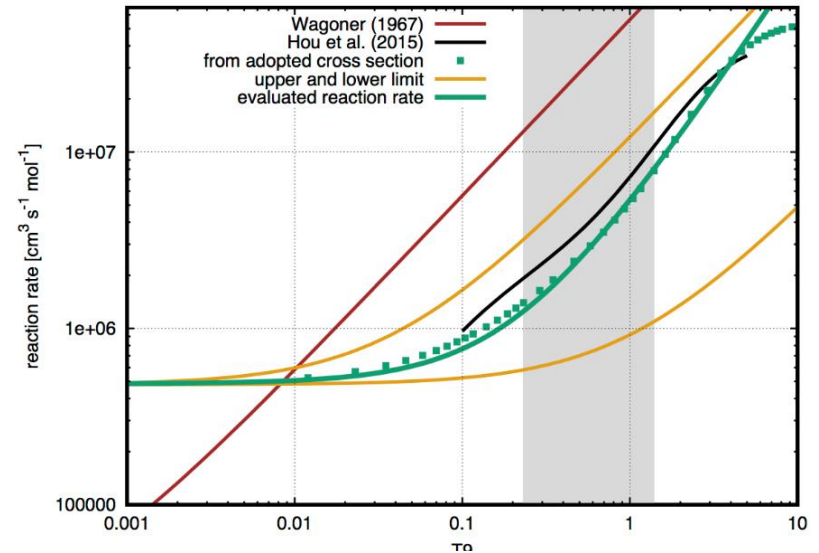
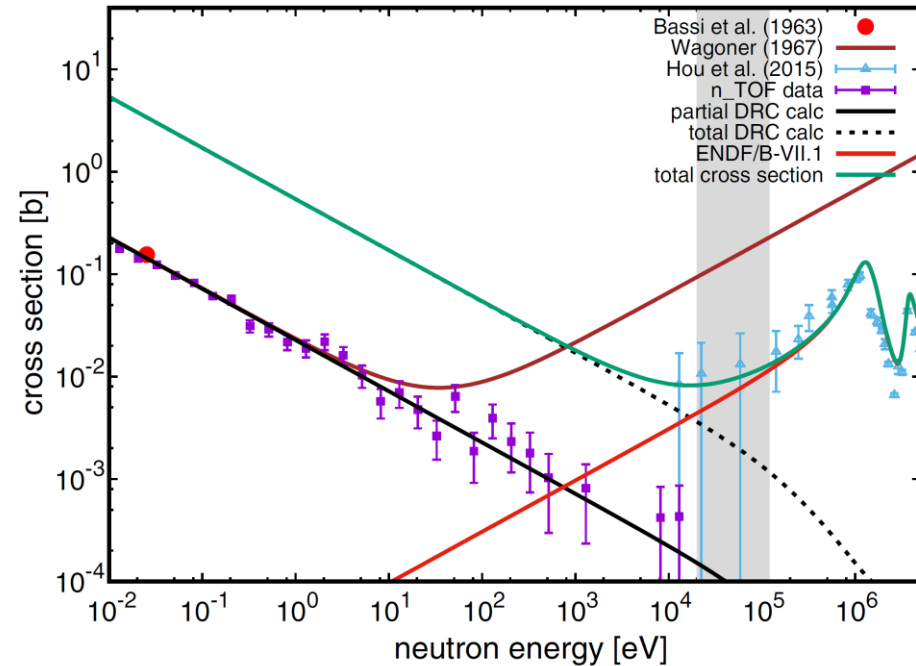
$E_{\text{dep}} > 2 \text{ MeV}$

Two low energy states of  ${}^8\text{Be}$  not accessible experimentally.

Missing states fractional contributions have been calculated.





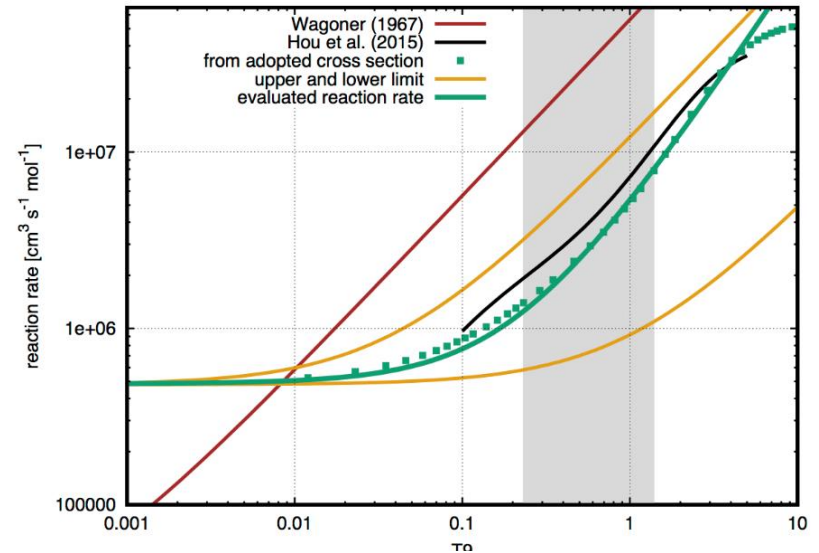
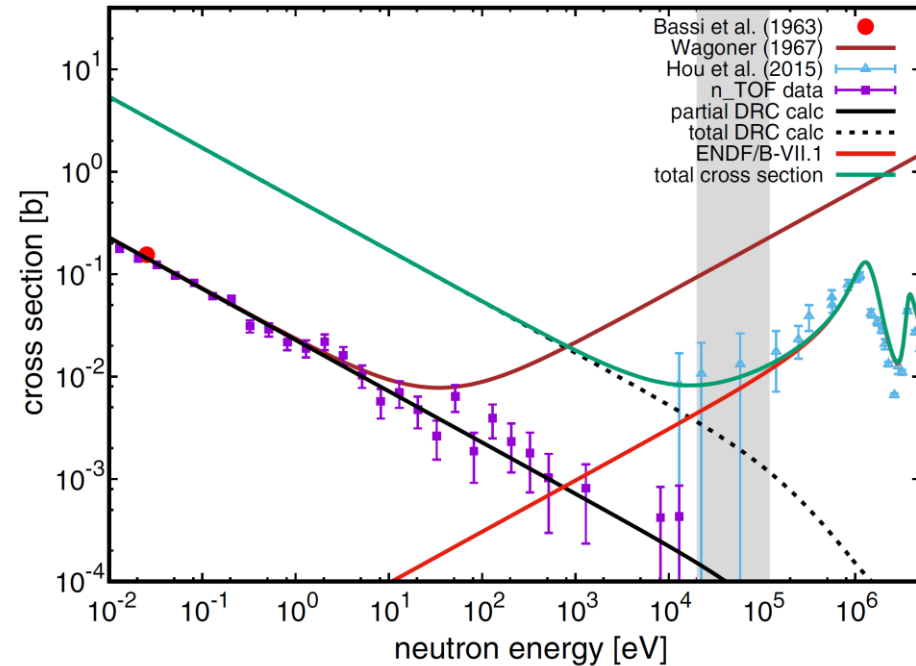


$$N_A \langle \sigma v \rangle = 4.81 \times 10^5 + 1.84 \times 10^6 T_9 + 3.03 \times 10^6 T_9^{3/2}$$

M. Barbagallo et al. (*n\_TOF Coll.*), *Phys. Rev. Lett.* 117, 152701, 2016

- <http://home.cern/about/updates/2016/10/ntof-plays-hide-and-seek-cosmological-lithium>
- <http://home.infn.it/it/comunicazione/news/1999-il-mistero-nascosto-nei-primi-tre-minuti-di-vita-dell-universo>

# ${}^7\text{Be}(n,\alpha){}^4\text{He}$ n\_TOF results and CLiP

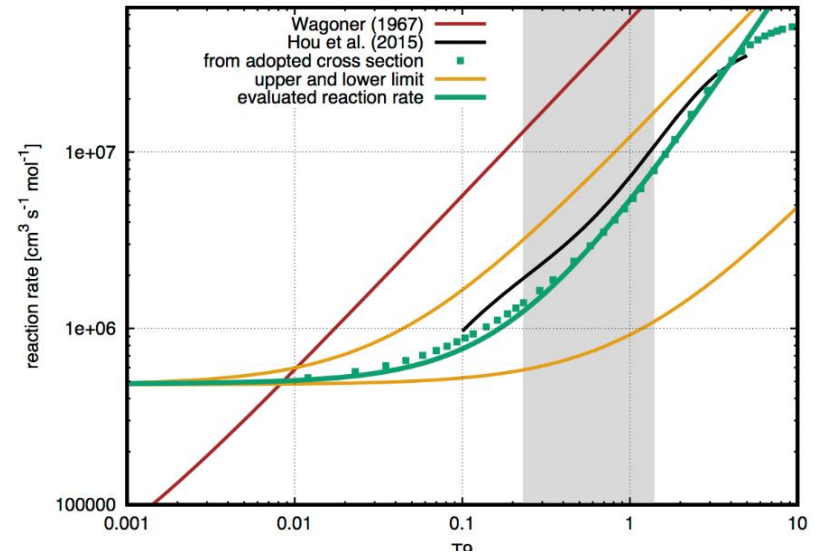
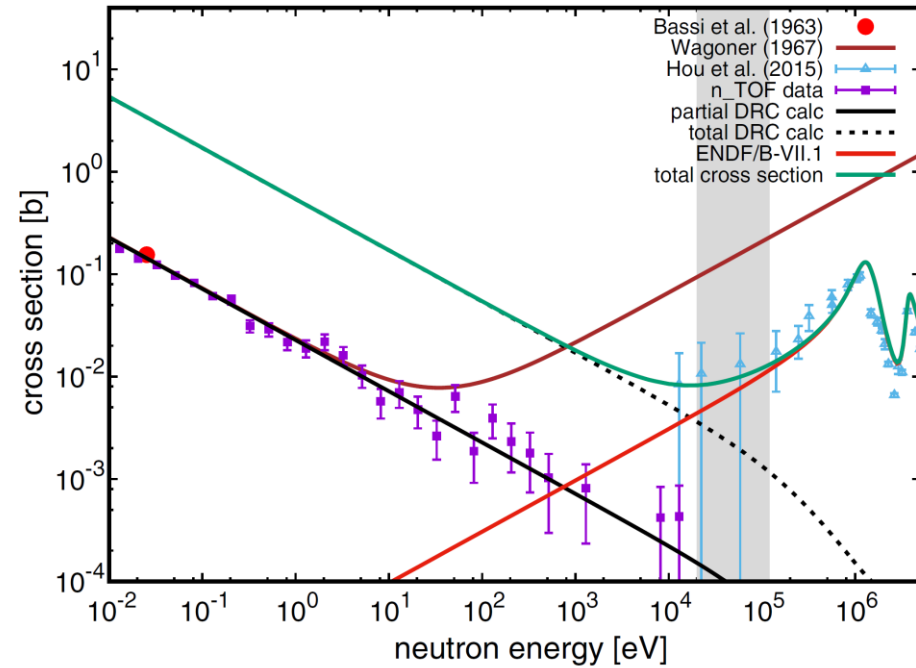


$$N_A \langle \sigma v \rangle = 4.81 \times 10^5 + 1.84 \times 10^6 T_9 + 3.03 \times 10^6 T_9^{3/2}$$

M. Barbagallo et al. (*n\_TOF Coll.*), *Phys. Rev. Lett.* 117, 152701, 2016

- <http://home.cern/about/updates/2016/10/ntof-plays-hide-and-seek-cosmological-lithium>
- <http://home.infn.it/it/comunicazione/news/1999-il-mistero-nascosto-nei-primi-tre-minuti-di-vita-dell-universo>

**As for (n, $\alpha$ ) measurement, the Cosmological Lithium Problem gets worse!**



$$N_A \langle \sigma v \rangle = 4.81 \times 10^5 + 1.84 \times 10^6 T_9 + 3.03 \times 10^6 T_9^{3/2}$$

M. Barbagallo et al. (*n\_TOF Coll.*), *Phys. Rev. Lett.* 117, 152701, 2016

- <http://home.cern/about/updates/2016/10/ntof-plays-hide-and-seek-cosmological-lithium>
- <http://home.infn.it/it/comunicazione/news/1999-il-mistero-nascosto-nei-primi-tre-minuti-di-vita-dell-universo>

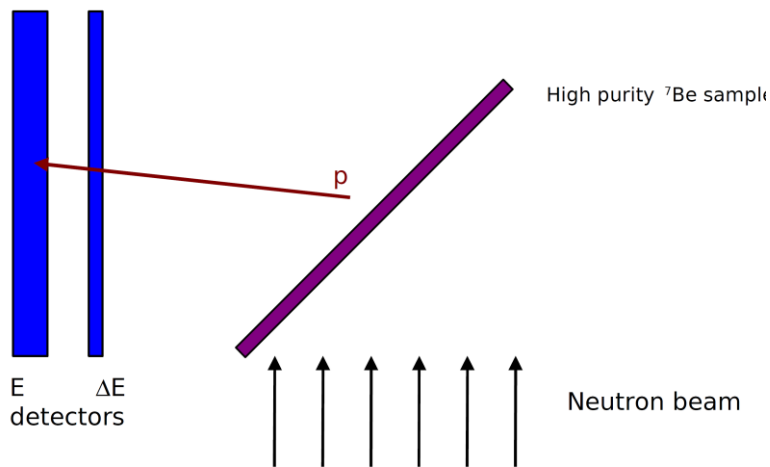
**As for (n,α) measurement, the Cosmological Lithium Problem gets worse!**

T. Kawabata et al., *Phys. Rev. Lett.* 118, 052701, 2017



Detection and identification of protons of 1.4 MeV and 1 MeV

Silicon telescope @n\_TOF-EAR2.

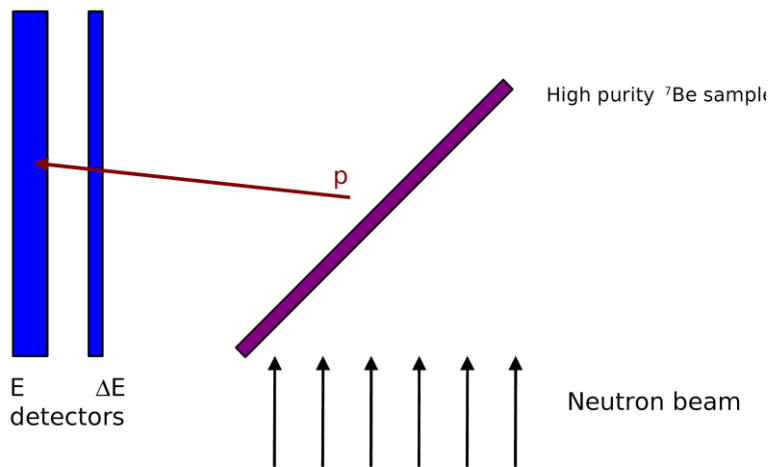


# ${}^7\text{Be}(n,p){}^7\text{Li}$ measurement



Detection and identification of protons of 1.4 MeV and 1 MeV

Silicon telescope @n\_TOF-EAR2.



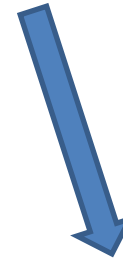
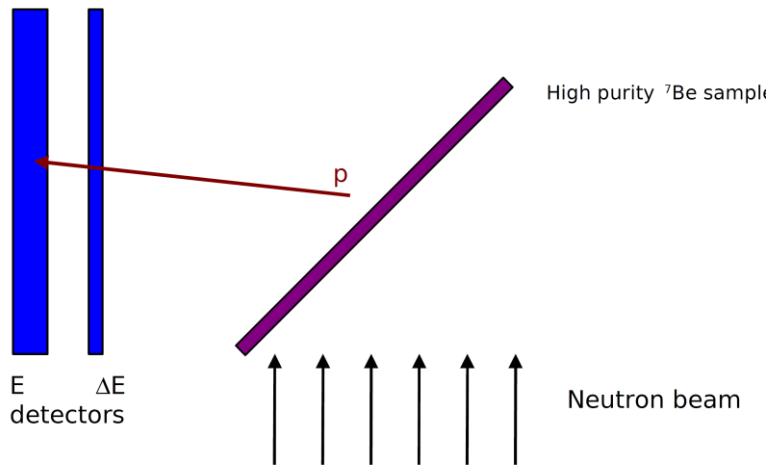
1 GBq high purity sample needed

(Chemical separation not sufficient)



Detection and identification of protons of 1.4 MeV and 1 MeV

Silicon telescope @n\_TOF-EAR2.



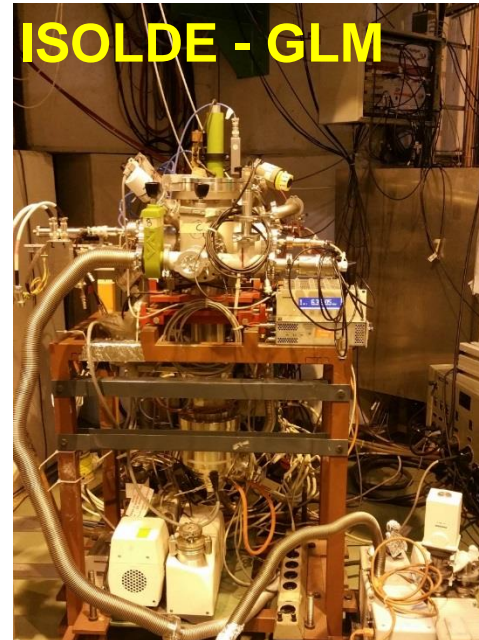
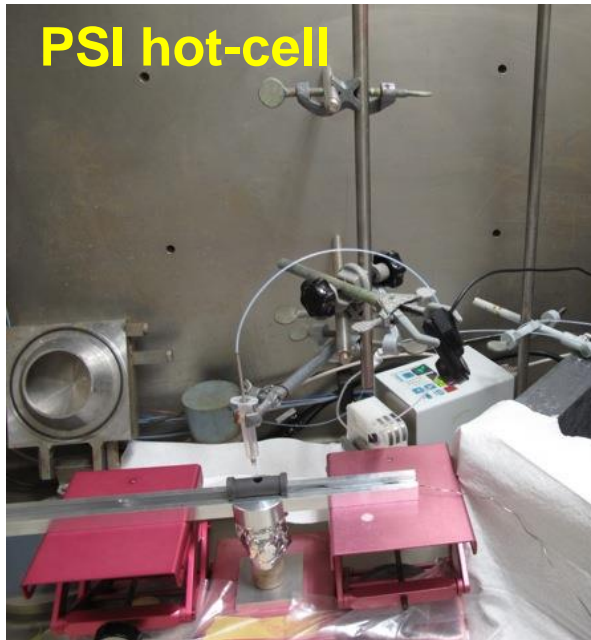
1 GBq high purity sample needed

(Chemical separation not sufficient)

- **First joint n\_TOF-ISOLDE experiment**
- **First time ever measurement of a neutron induced reaction cross-section using a target produced with a radioactive beam.**

A three steps experiment:

- Extraction of 200 GBq from water cooling of SINQ spallation source at PSI.
- Implantation of 30 keV ( $\sim 45$  nA)  ${}^7\text{Be}$  beam on suited backing using ISOLDE-GPS separator and RILIS.
- Measurement at n\_TOF-EAR2 using a silicon telescope (20 and 300  $\mu\text{m}$ , 5x5  $\text{cm}^2$  strip device).



E. Maugeri et al., *Nucl. Instr. and Meth. A* 889 (2018) 138-144.

M. Barbagallo et al., *Nucl. Instr. and Meth. A* 887 (2018) 27-3

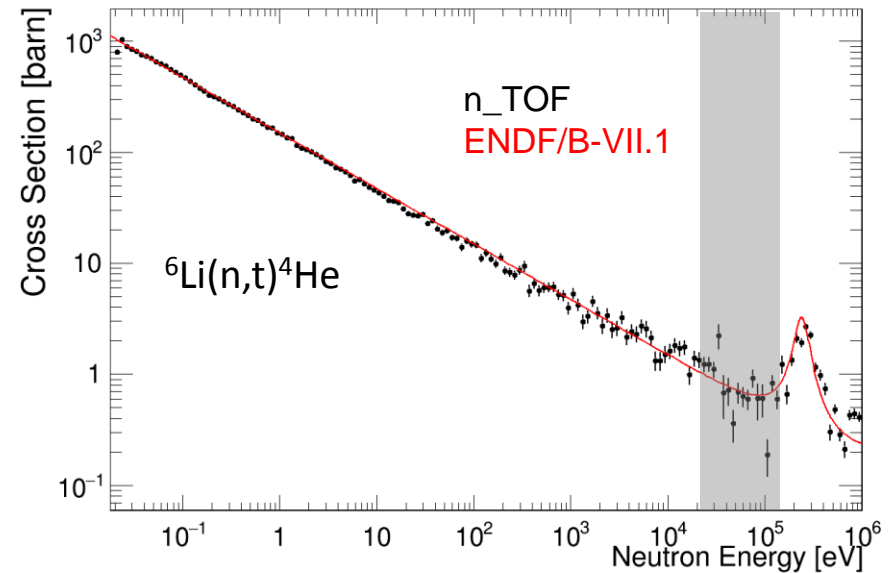
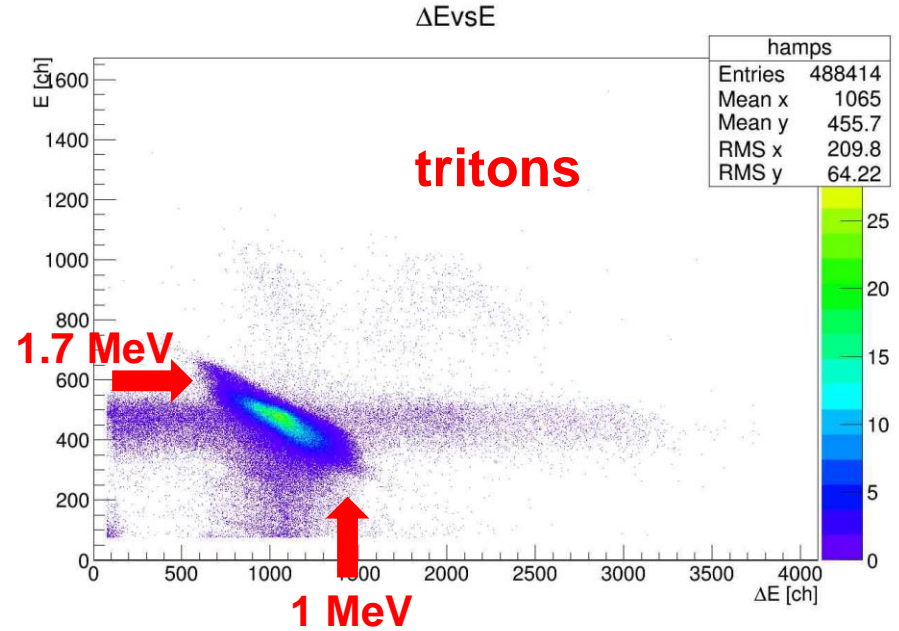
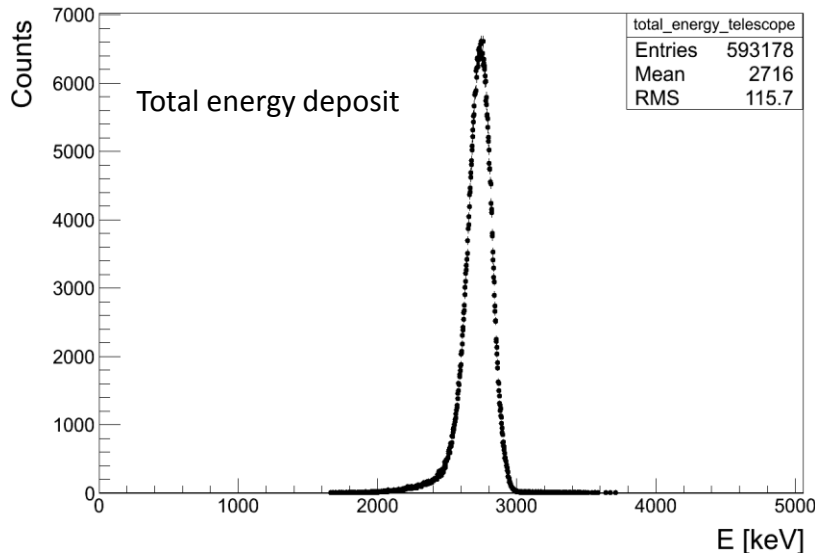


# ${}^7\text{Be}(n,p){}^7\text{Li}$ measurement

The detection system was characterized using  $\alpha$ -source and the well-known  ${}^6\text{Li}(n,t){}^4\text{He}$  reaction.

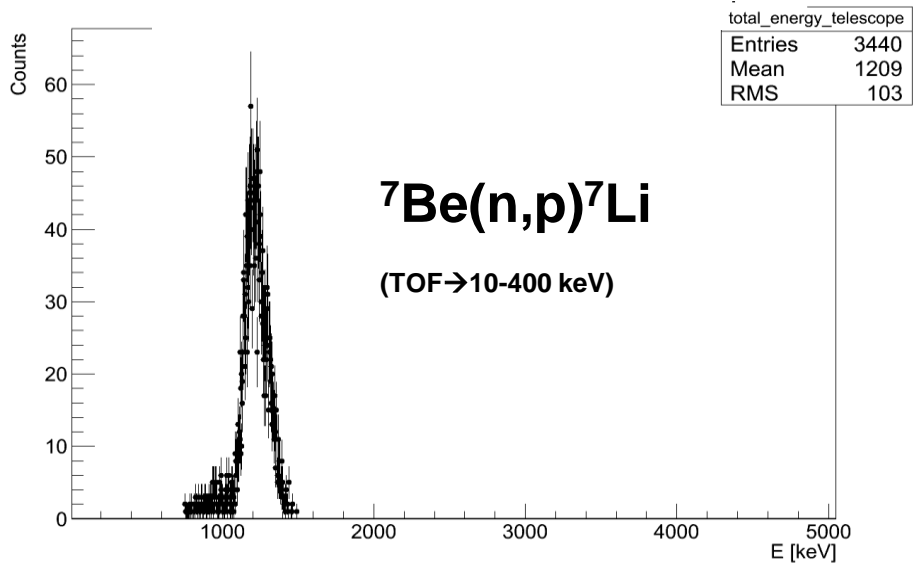
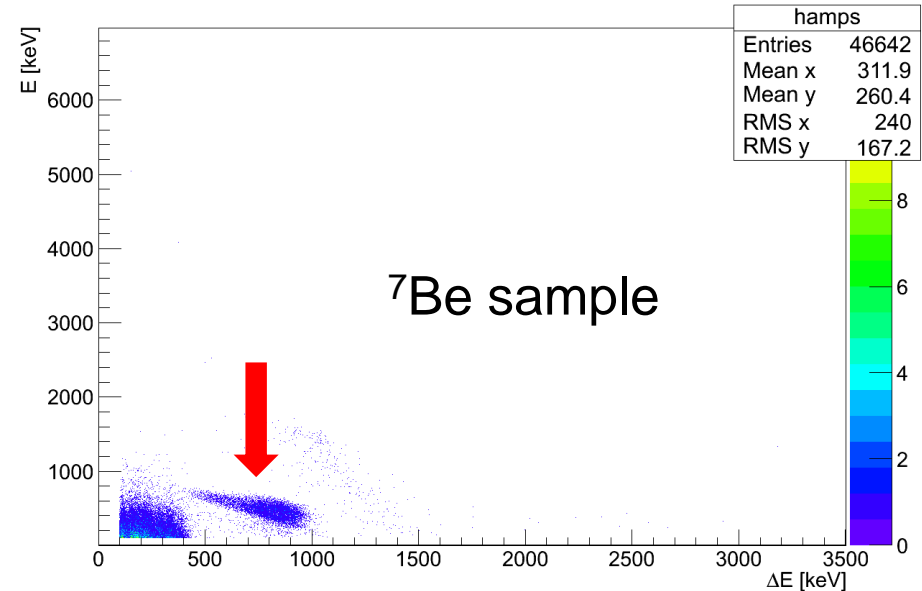
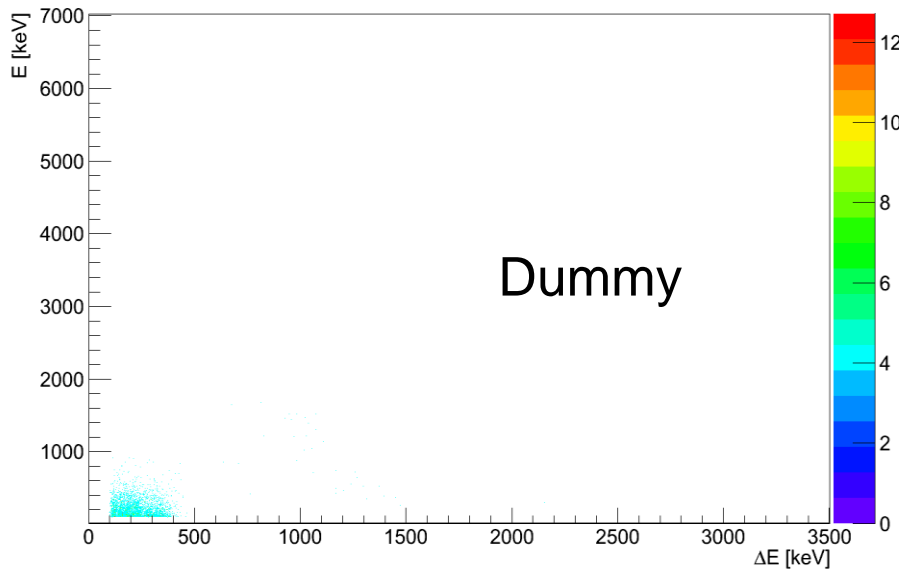


$Q = 4.78 \text{ MeV}$





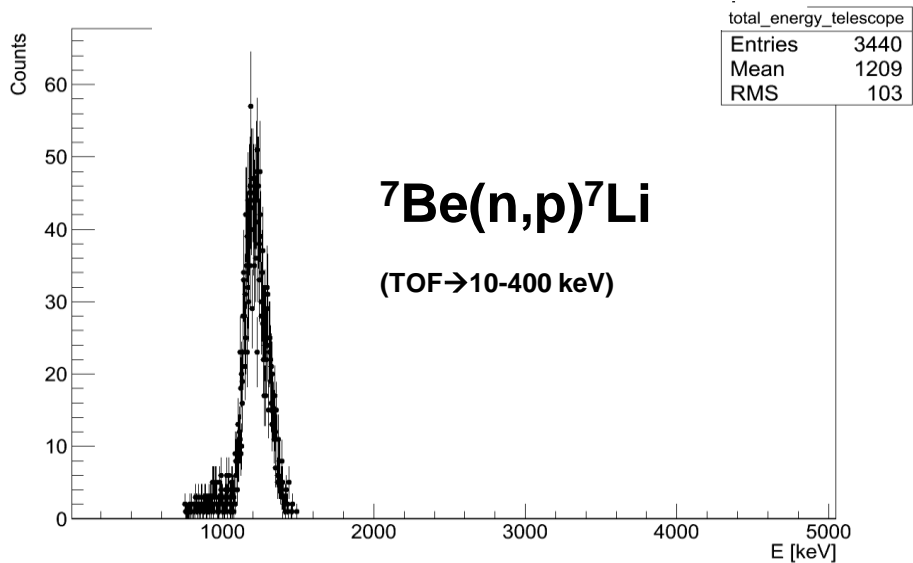
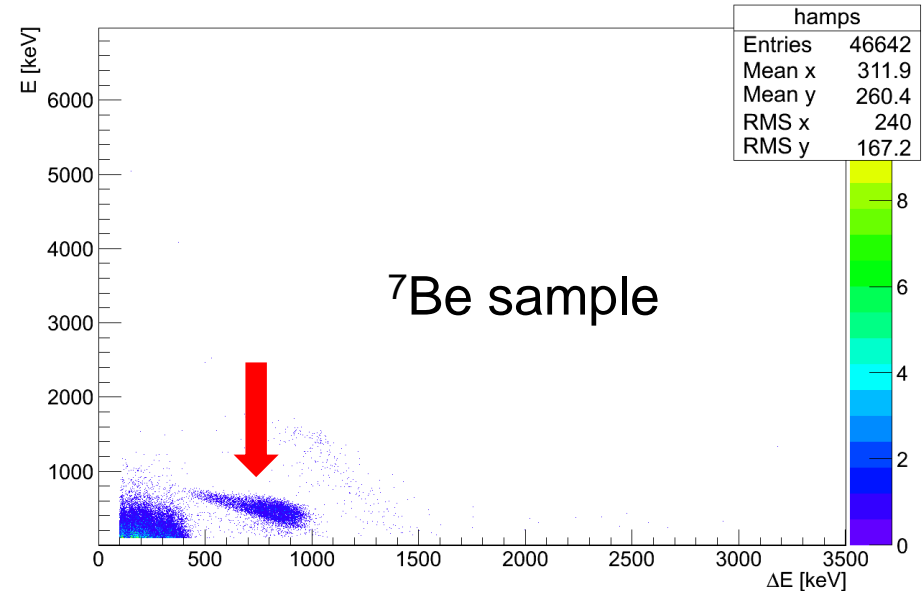
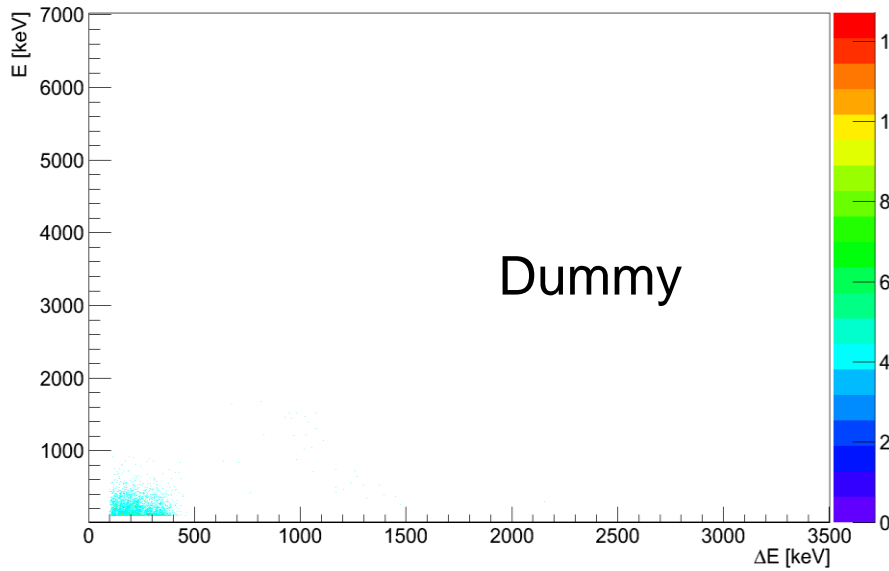
# ${}^7\text{Be}(n,p){}^7\text{Li}$ measurement results



First time ever direct measurement of  ${}^7\text{Be}(n,p)$  reaction in the range of interest for Big Bang Nucleosynthesis.

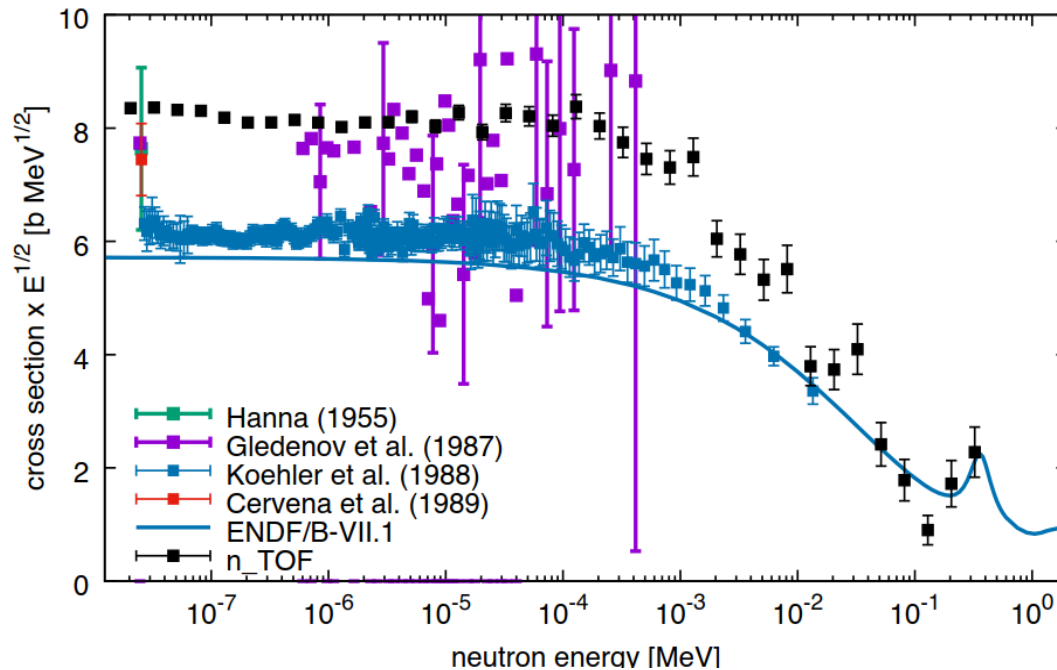
Stay tuned...

# ${}^7\text{Be}(n,p){}^7\text{Li}$ measurement results



First time ever direct measurement of  ${}^7\text{Be}(n,p)$  reaction in the range of interest for Big Bang Nucleosynthesis.

~~Stay tuned...~~

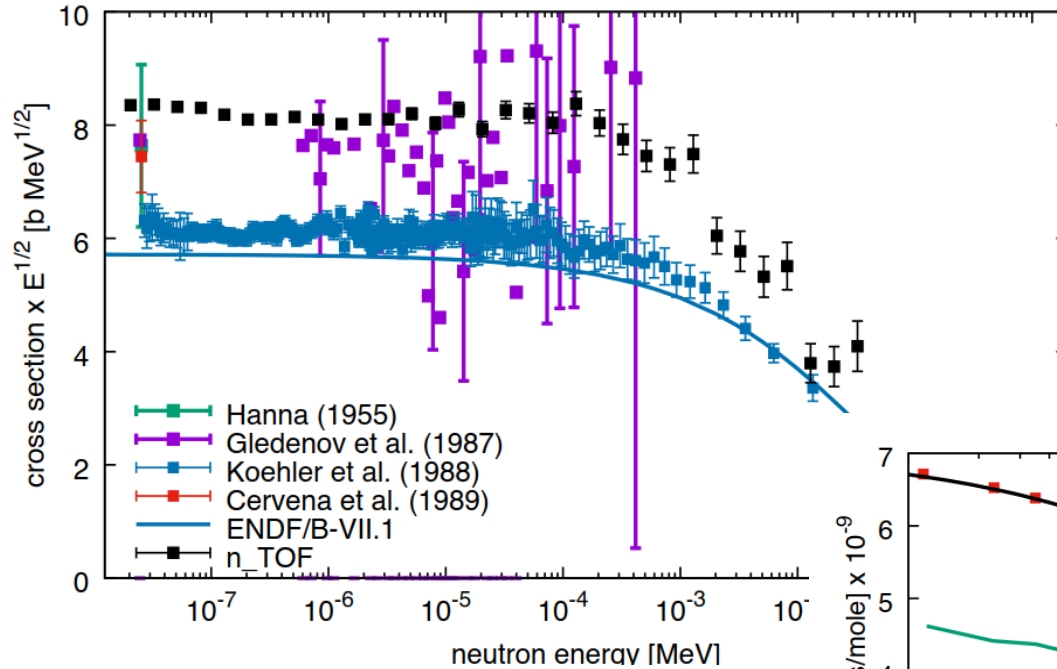


*L. Damone et al. (n\_TOF Coll.), submitted to Phys. Rev. Lett.*

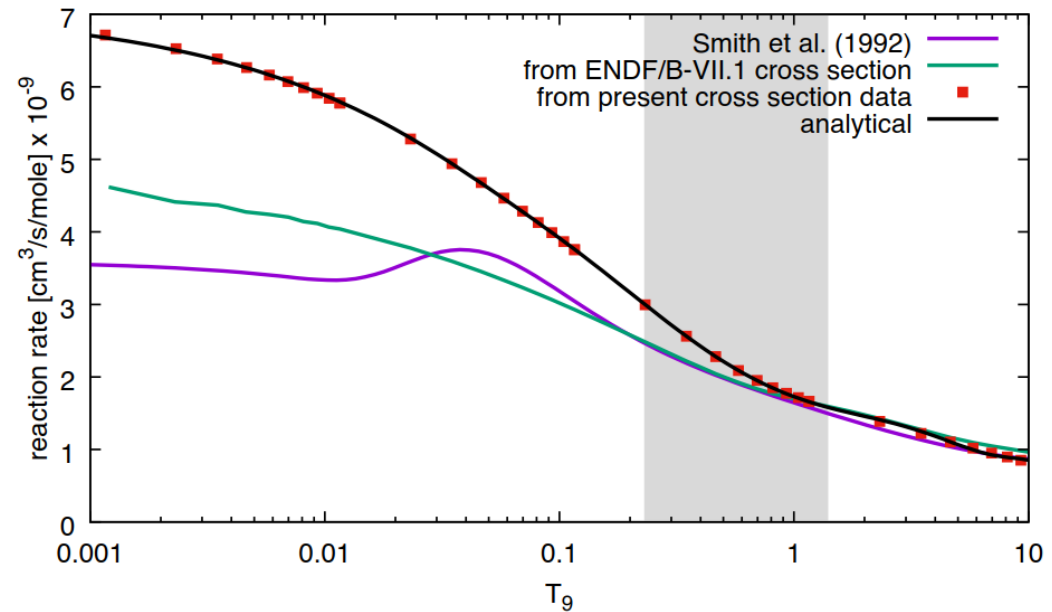
**First time ever direct measurement of  ${}^7\text{Be}(n,p){}^7\text{Li}$  reaction in the range of interest for Big Bang Nucleosynthesis.**

~~Stay tuned...~~

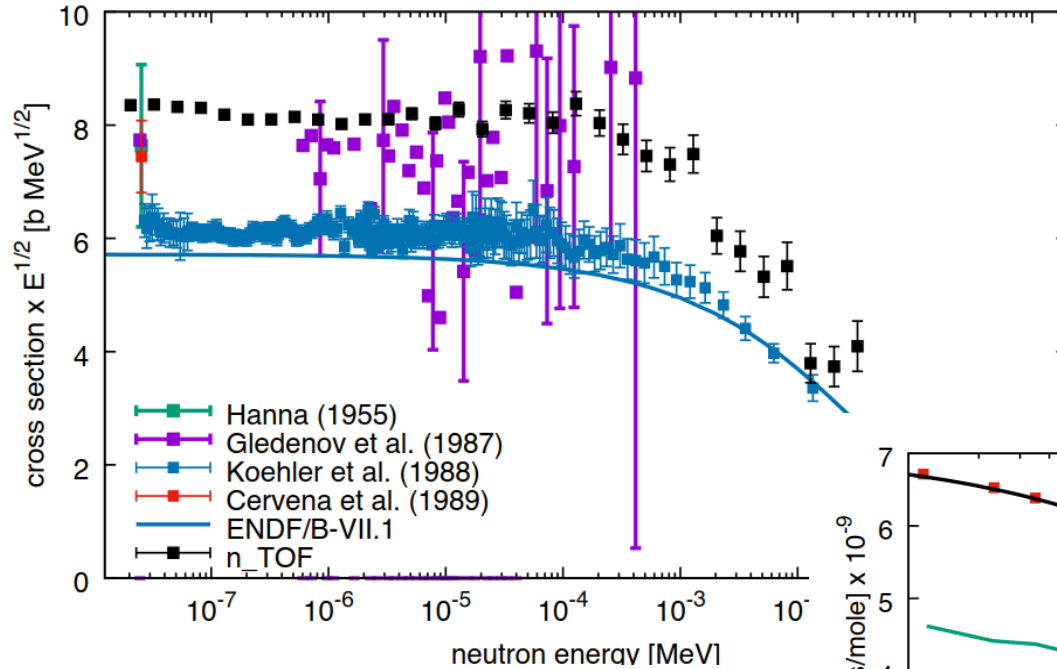
# ${}^7\text{Be}(n,p){}^7\text{Li}$ measurement results



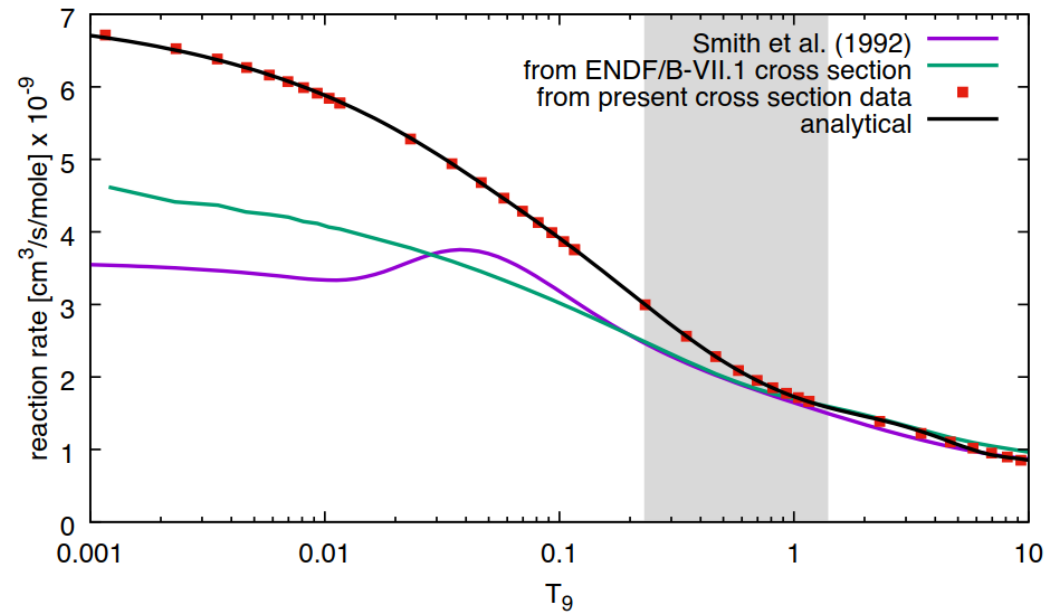
*L. Damone et al. (n\_TOF Coll.), submitted to Phys. Rev. Lett.*



# ${}^7\text{Be}(n,p){}^7\text{Li}$ measurement results



*L. Damone et al. (n\_TOF Coll.), submitted to Phys. Rev. Lett.*

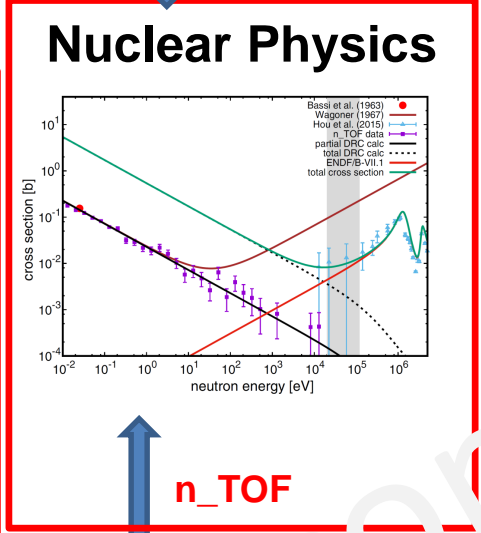
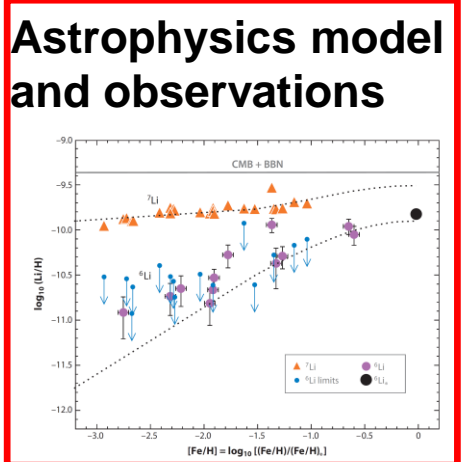


**Solution to Cosmological Lithium Problem has to be sought in other Physics scenarios !!!**

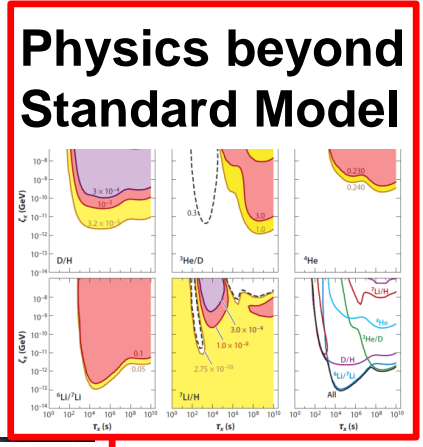
- Uncertainties in nuclear data strongly affect the Big Bang Nucleosynthesis calculations for the abundance of  ${}^7\text{Li}$  and could possibly explain (at least shade new light on) the **Cosmological Lithium Problem**.
- ${}^7\text{Be}(n,\alpha){}^4\text{He}$  cross-section has been measured for the first time in a wide energy range, using **n\_TOF-EAR2** neutron beam and two samples prepared at **PSI**.
- The  ${}^7\text{Be}(n,p){}^7\text{Li}$  cross-section measurement has been performed at **n\_TOF-EAR2**, using a **1.1 GBq** pure sample implanted at **ISOLDE** from 0.02 eV to 500 keV (**first time at BBN energy window**).
- The new estimate of the  ${}^7\text{Be}$  destruction rate **based on the new n\_TOF results** yields to a slight decrease of the predicted cosmological Lithium abundance, **insufficient to provide a viable solution to the Cosmological Lithium Problem**.
- **Solution to CLiP is somewhere else!**

# Perspectives

Log(Present knowledge) [a.u.]



resonant states in  ${}^7\text{Be}(cp,X)$ , lifetime in plasma..



Fascination [a.u.]

- Uncertainties in nuclear data strongly affect the Big Bang Nucleosynthesis calculations for the abundance of  ${}^7\text{Li}$  and could possibly explain (at least shade new light on) the **Cosmological Lithium Problem**.
- ${}^7\text{Be}(n,\alpha){}^4\text{He}$  cross-section has been measured for the first time in a wide energy range, using **n\_TOF-EAR2** neutron beam and two samples prepared at **PSI**.
- The  ${}^7\text{Be}(n,p){}^7\text{Li}$  cross-section measurement has been performed at **n\_TOF-EAR2**, using a **1.1 GBq** pure sample implanted at **ISOLDE** from 0.02 eV to 500 keV (**first time at BBN energy window**).
- The new estimate of the  ${}^7\text{Be}$  destruction rate **based on the new n\_TOF results** yields to a slight decrease of the predicted cosmological Lithium abundance, **insufficient to provide a viable solution to the Cosmological Lithium Problem**.
- **Solution to CLiP is somewhere else!**
- *The new data can be used for more accurate calculation of the reaction yield and neutron spectrum in the near-threshold  ${}^7\text{Li}(p,n){}^7\text{Be}$  reaction, important for neutron sources and Nuclear Astrophysics.*

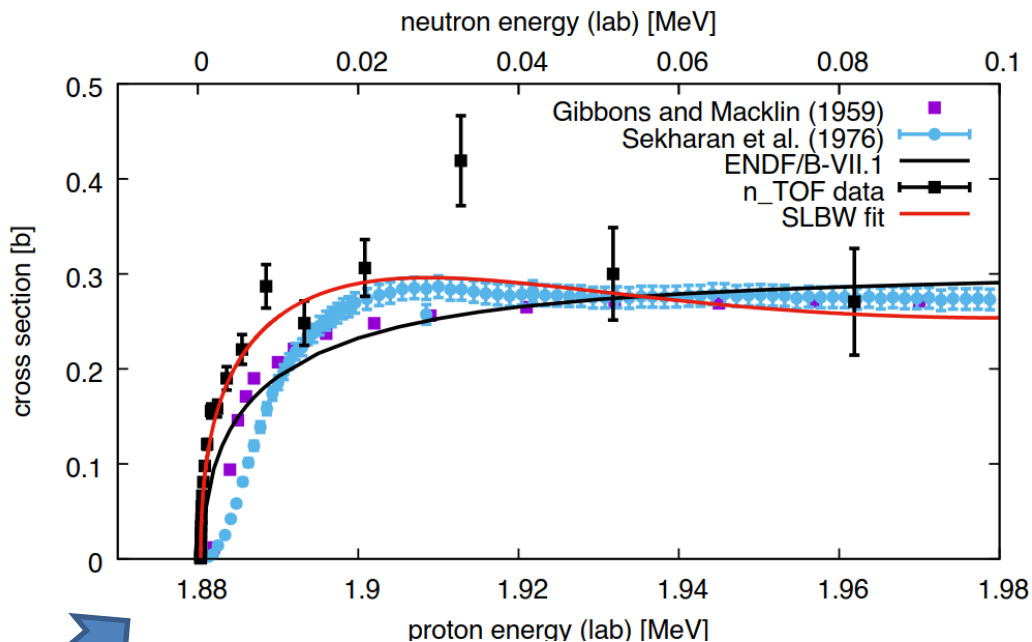
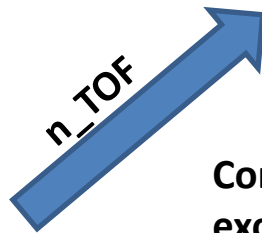


The **present data** can also provide information on cross-section of the  ${}^7\text{Li}(p,n){}^7\text{Be}$  reaction, one of the most important reactions for **neutron production at low-energy accelerators**.

Need to know **excitation function** for  ${}^7\text{Li}(p,n){}^7\text{Be}$  reaction @25-30 keV

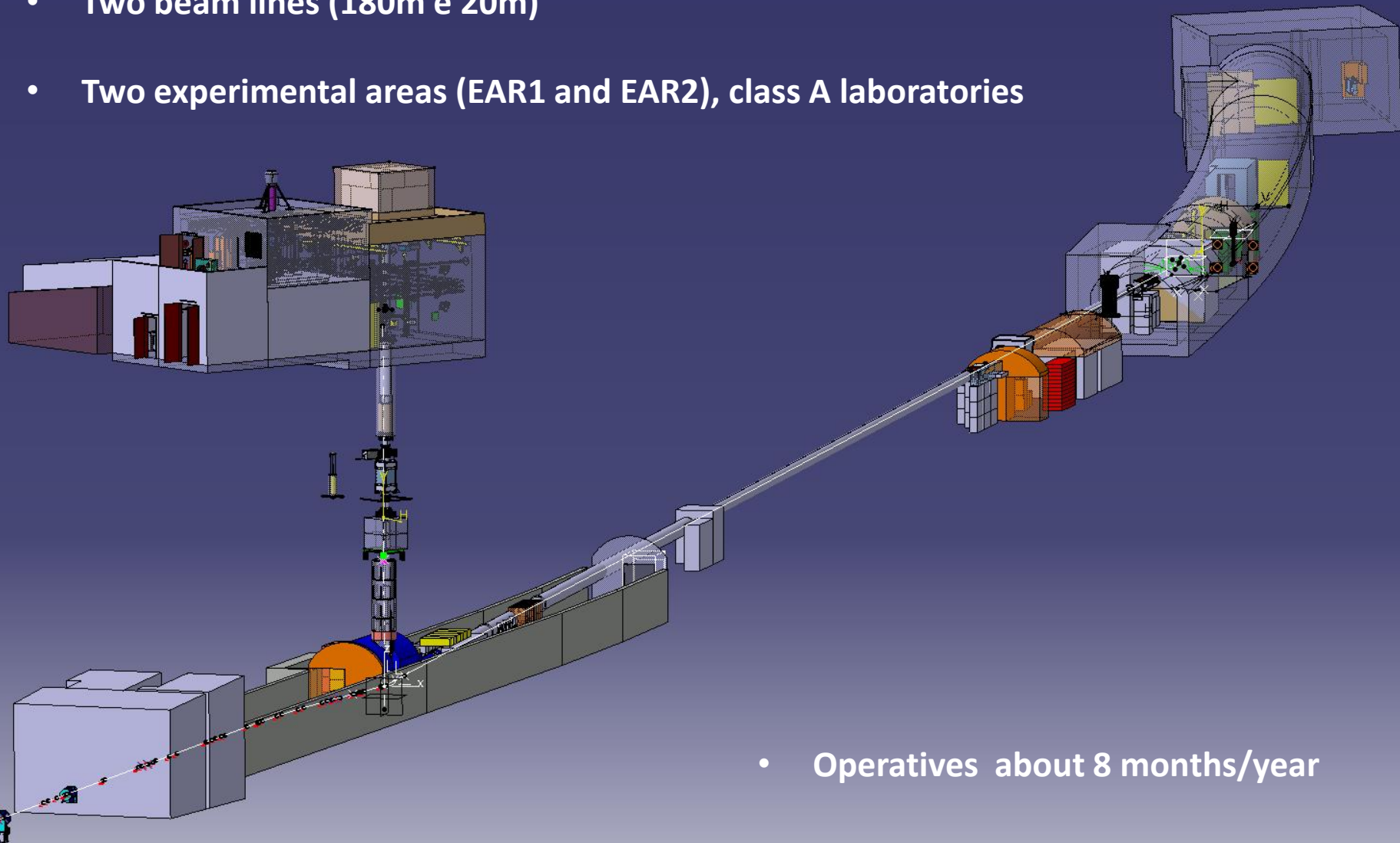
- low energy of emitted neutrons
- well calibrated stable proton beam
- poor energy resolution

BUT..



Compared with direct measurements, the extracted excitation function shows a much faster rise above the threshold.

- Two beam lines (180m e 20m)
- Two experimental areas (EAR1 and EAR2), class A laboratories



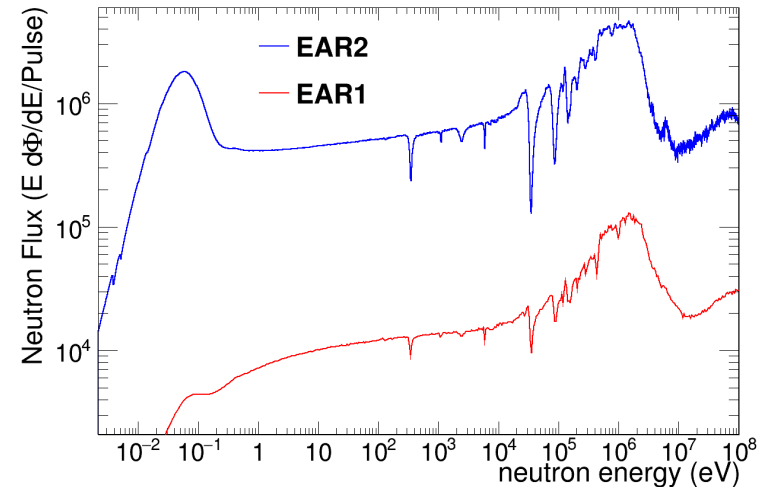
- Operatives about 8 months/year

Advantages of the **Proton Synchrotron beam: high energy, high peak current (7e12 ppp/7 ns)**

High instantaneous neutron flux  
( $10^5$  n/cm<sup>2</sup>/pulse and  $10^7$  n/cm<sup>2</sup>/pulse).

Very convenient for measurements of:

- radioactive isotopes,
- low cross sections,
- isotopes available in small quantities



**Other features of the neutron beams:**

- High **resolution in energy** ( $\Delta E/E = 10^{-4}$  in EAR1 and  $\Delta E/E = 10^{-3}$  in EAR2)
- Wide **energy range** ( $1 \text{ meV} < E_n < 1 \text{ GeV}$ )
- Low **repetition** rate ( $< 0.8 \text{ Hz}$ ) (no wrap-around)