

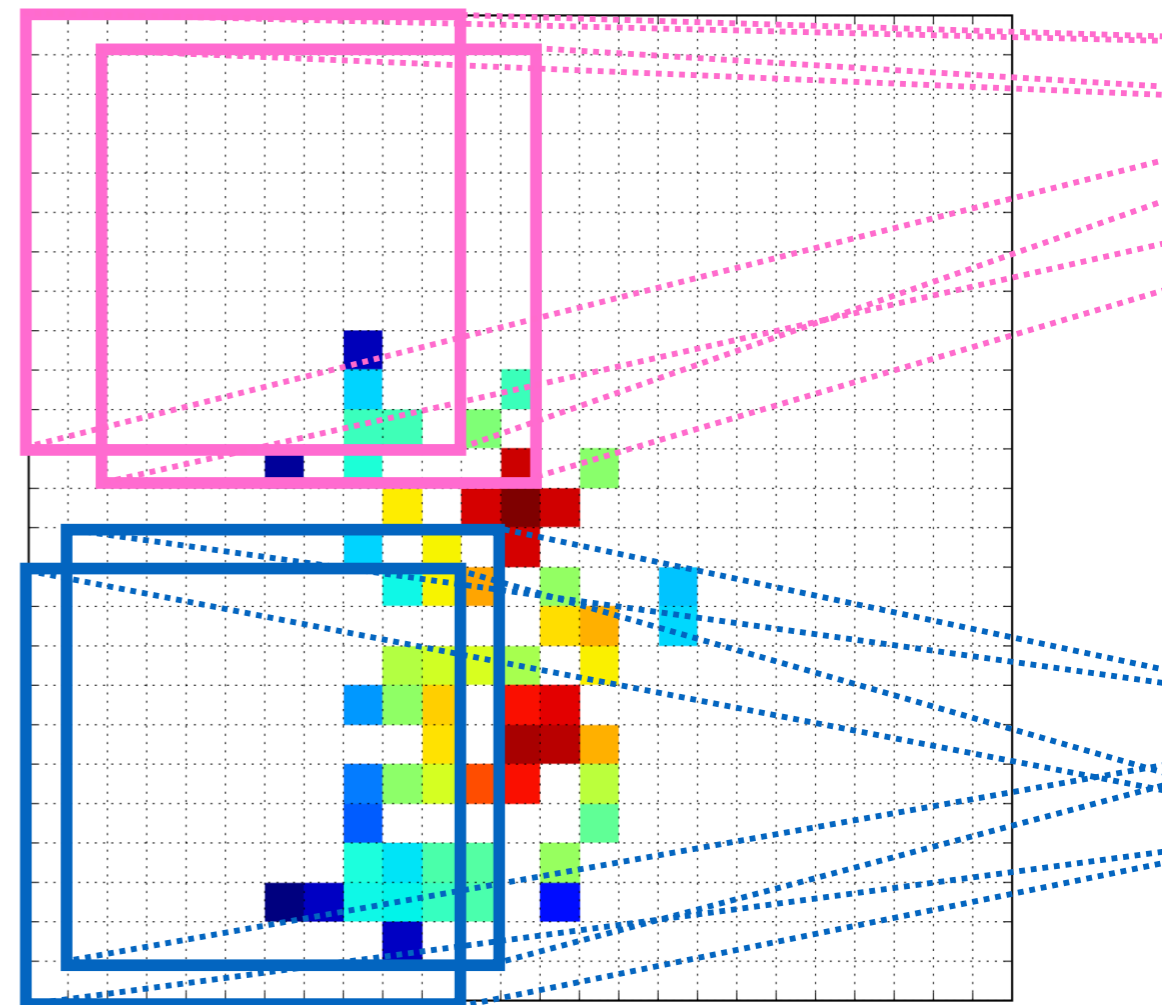
Advanced Machine Learning in High Energy Physics



Benjamin Nachman

*Lawrence Berkeley
National Laboratory*

*EFI Data Analytics Workshop
October 26, 2017*



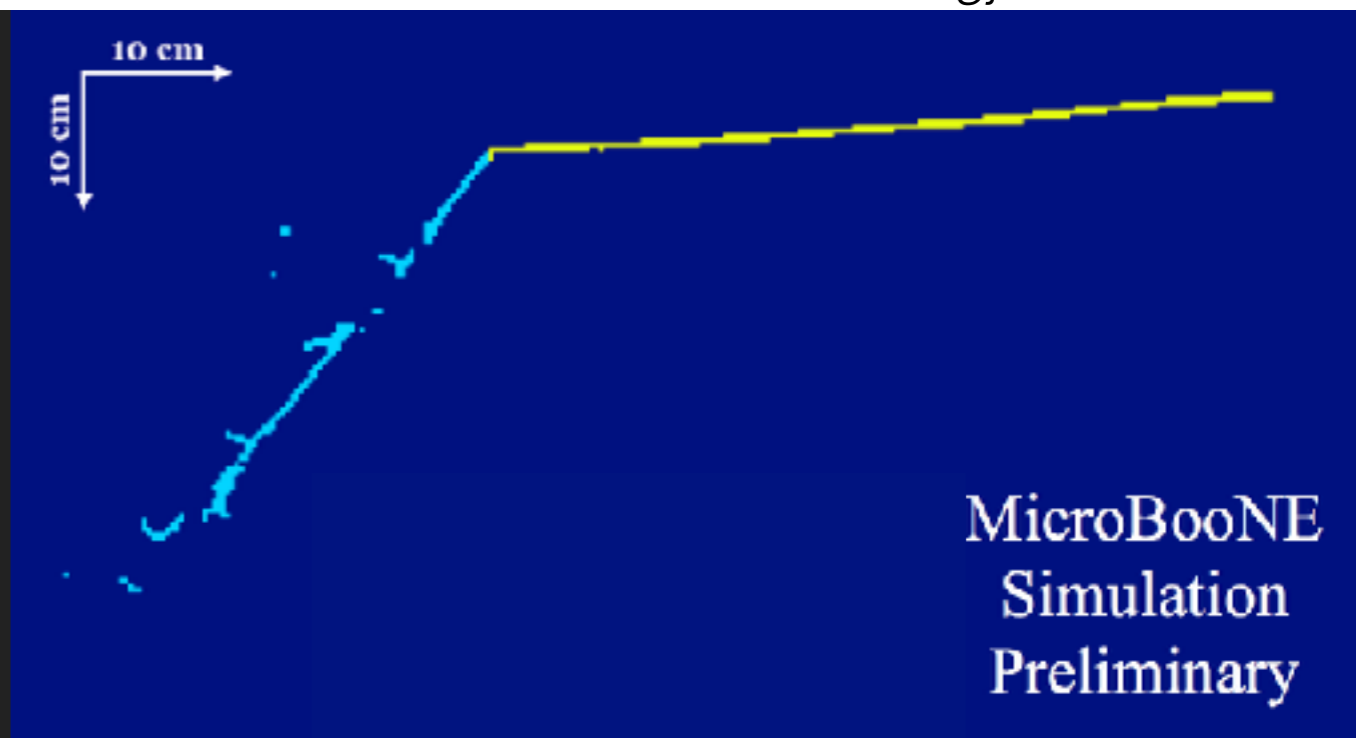
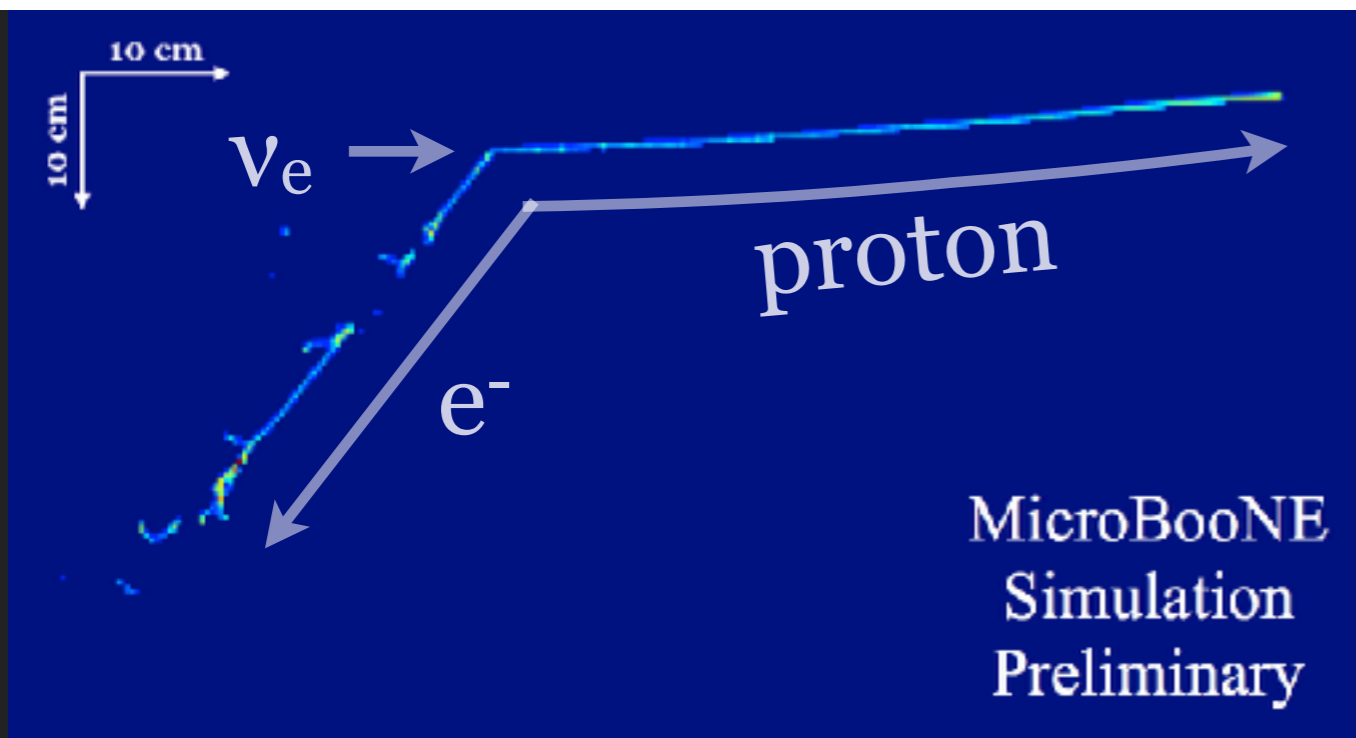
Disclaimer: I'm not going to talk about:

- deep learning “simply” replacing shallow learning
- interesting work from the large neutrino experiments
 - non-image based classification at the LHC

a lot of the content still applies

Instead, I'll use hadronic final states at the LHC to illustrate DNN classification, regression, & generation

from T. Wongjirad's DPF talk



Hadronic final states at the LHC

Center-of-mass energy = 13 TeV



Run: 302347

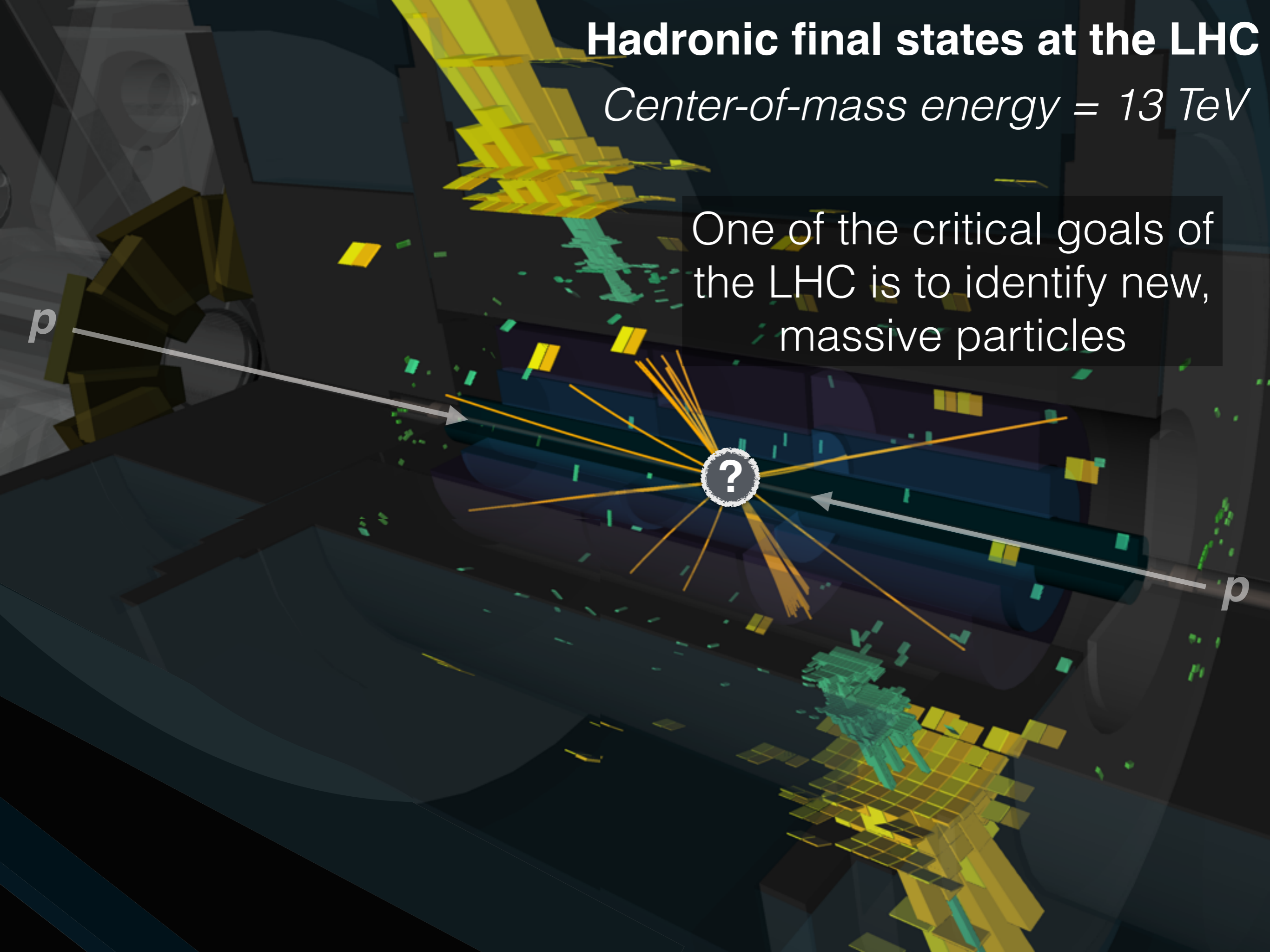
Event: 753275626

2016-06-18 18:41:48 CEST

Hadronic final states at the LHC

Center-of-mass energy = 13 TeV

One of the critical goals of the LHC is to identify new, massive particles



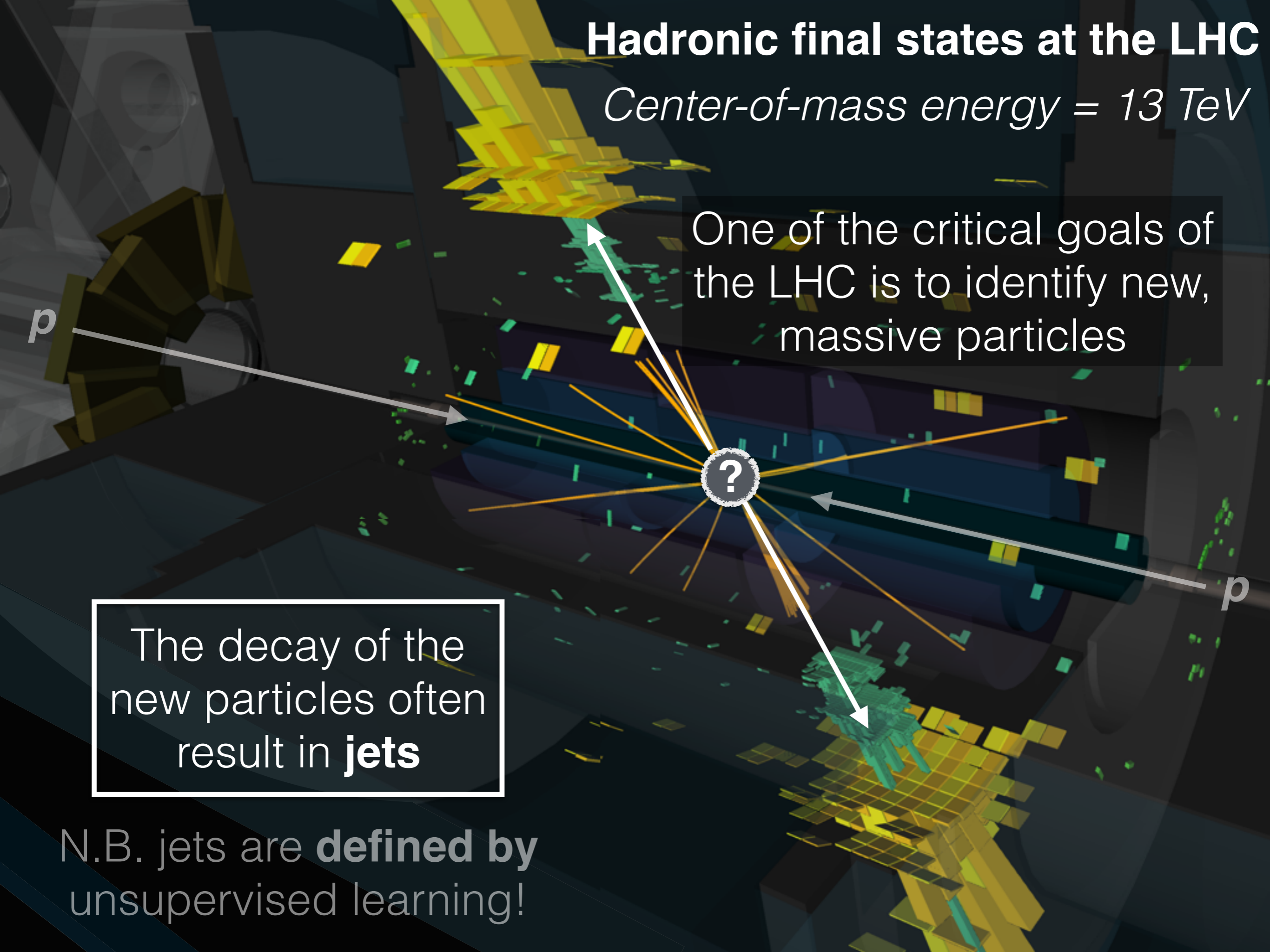
Hadronic final states at the LHC

Center-of-mass energy = 13 TeV

One of the critical goals of the LHC is to identify new, massive particles

The decay of the new particles often result in **jets**

N.B. jets are **defined by** unsupervised learning!

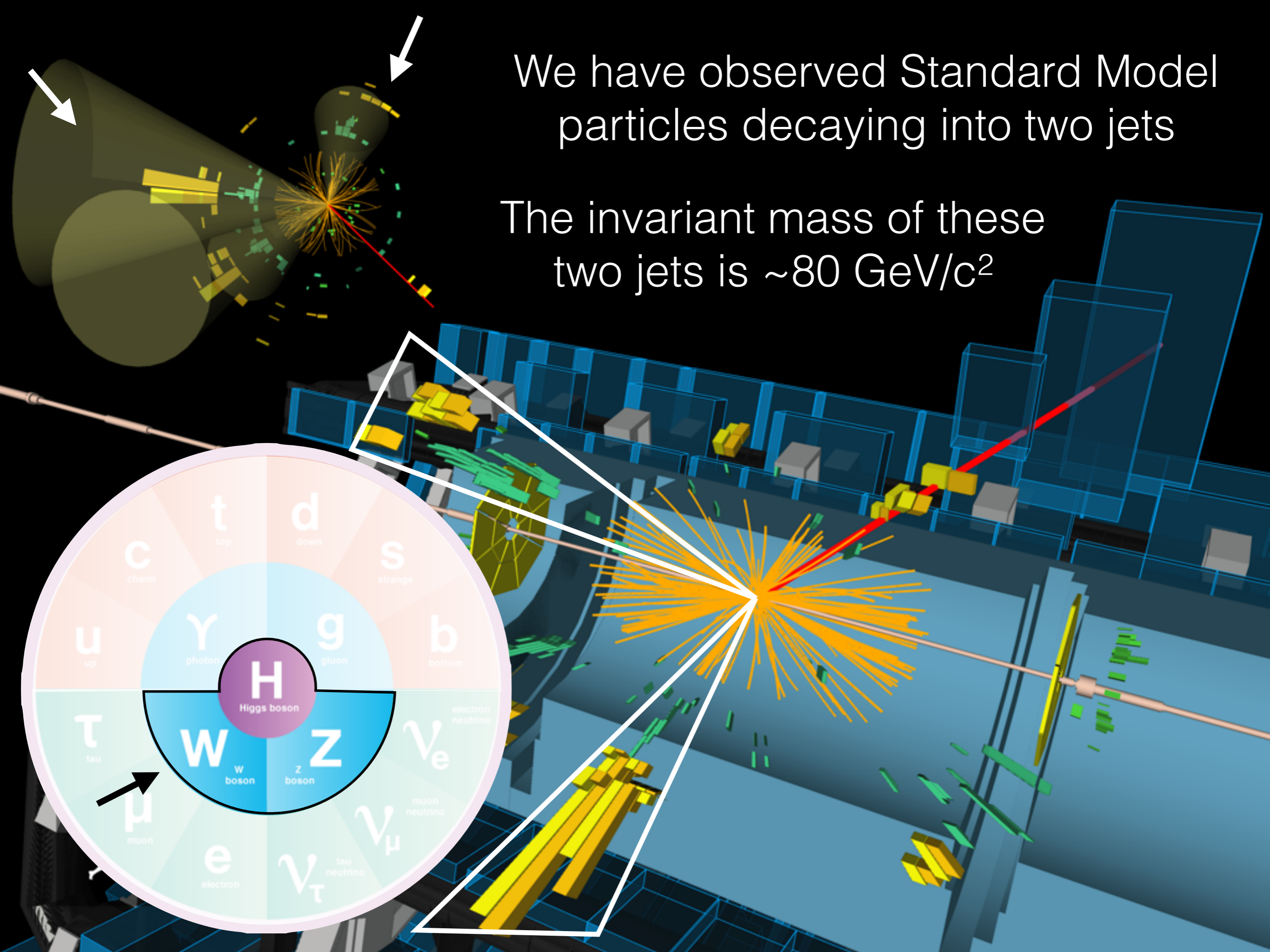
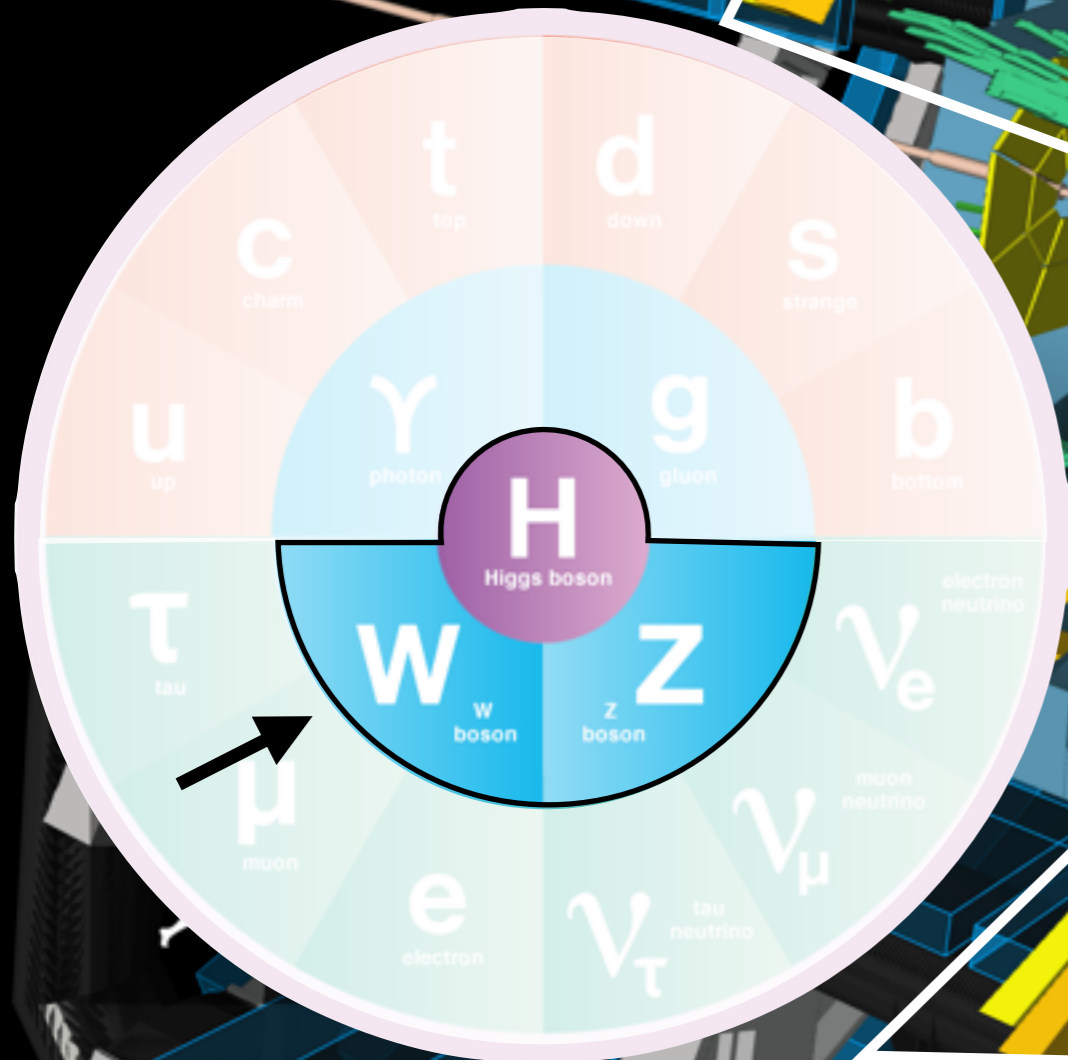


We have observed Standard Model particles decaying into two jets

The invariant mass of these two jets is $\sim 80 \text{ GeV}/c^2$

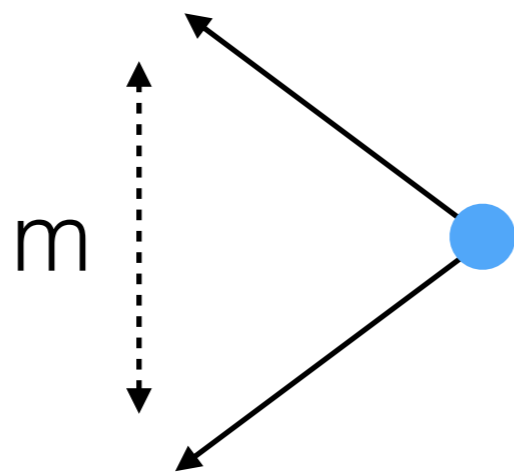
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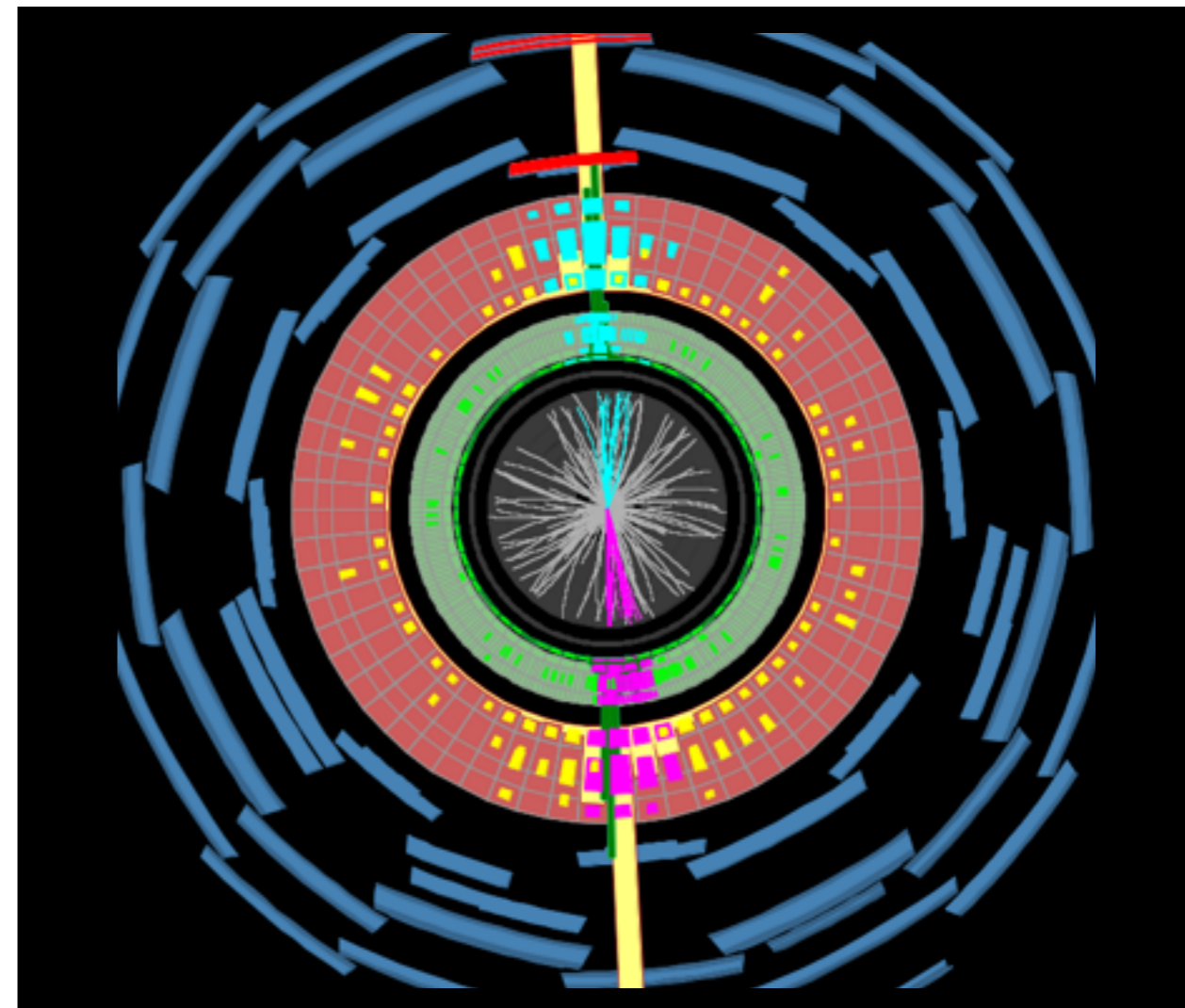
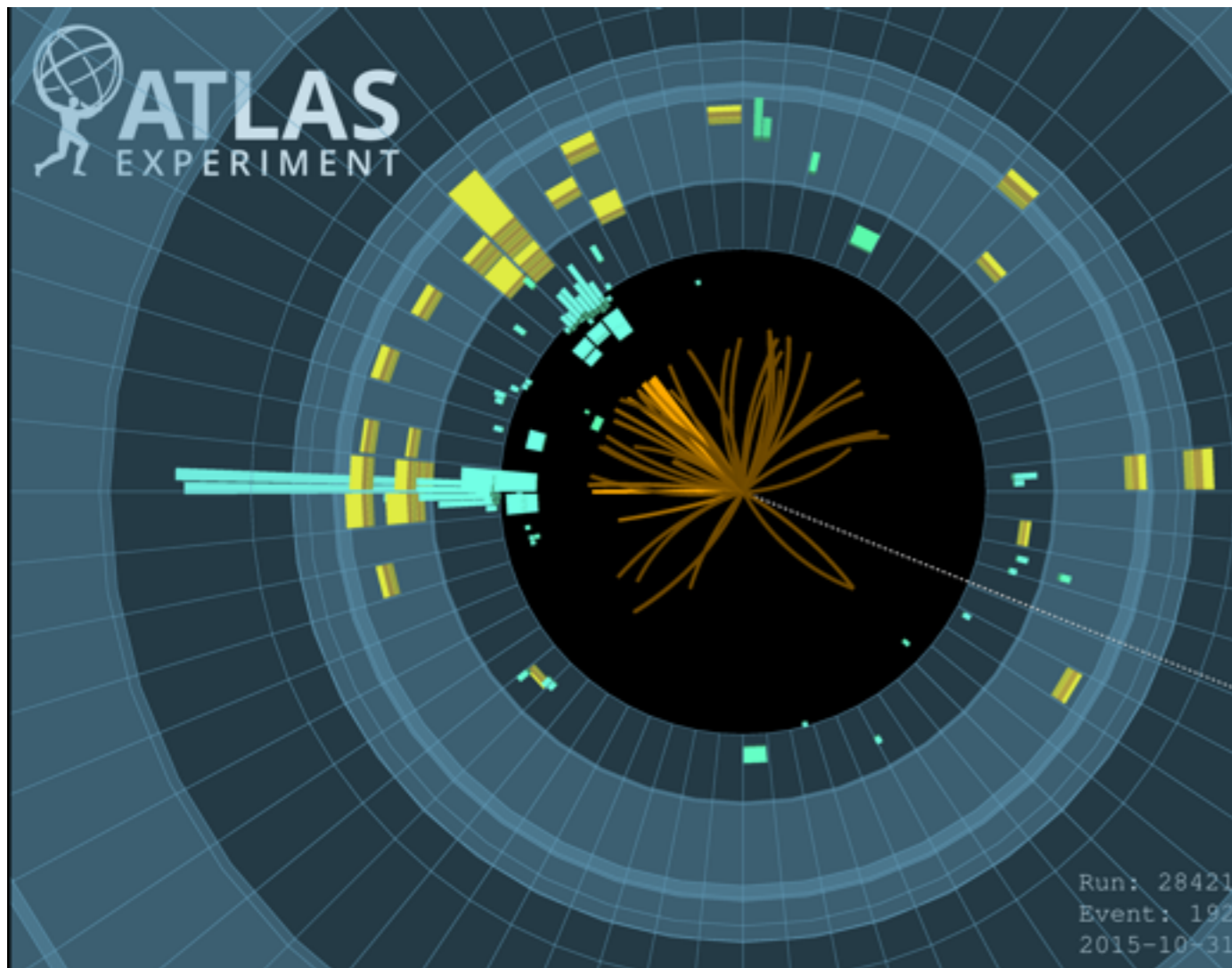
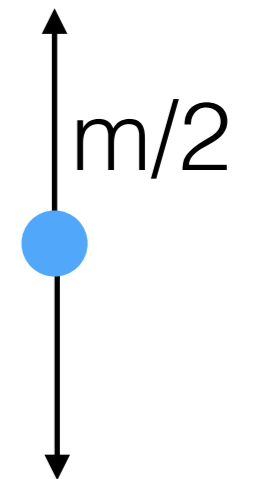


What if you take one of those SM dijet resonances and Lorentz boost it?

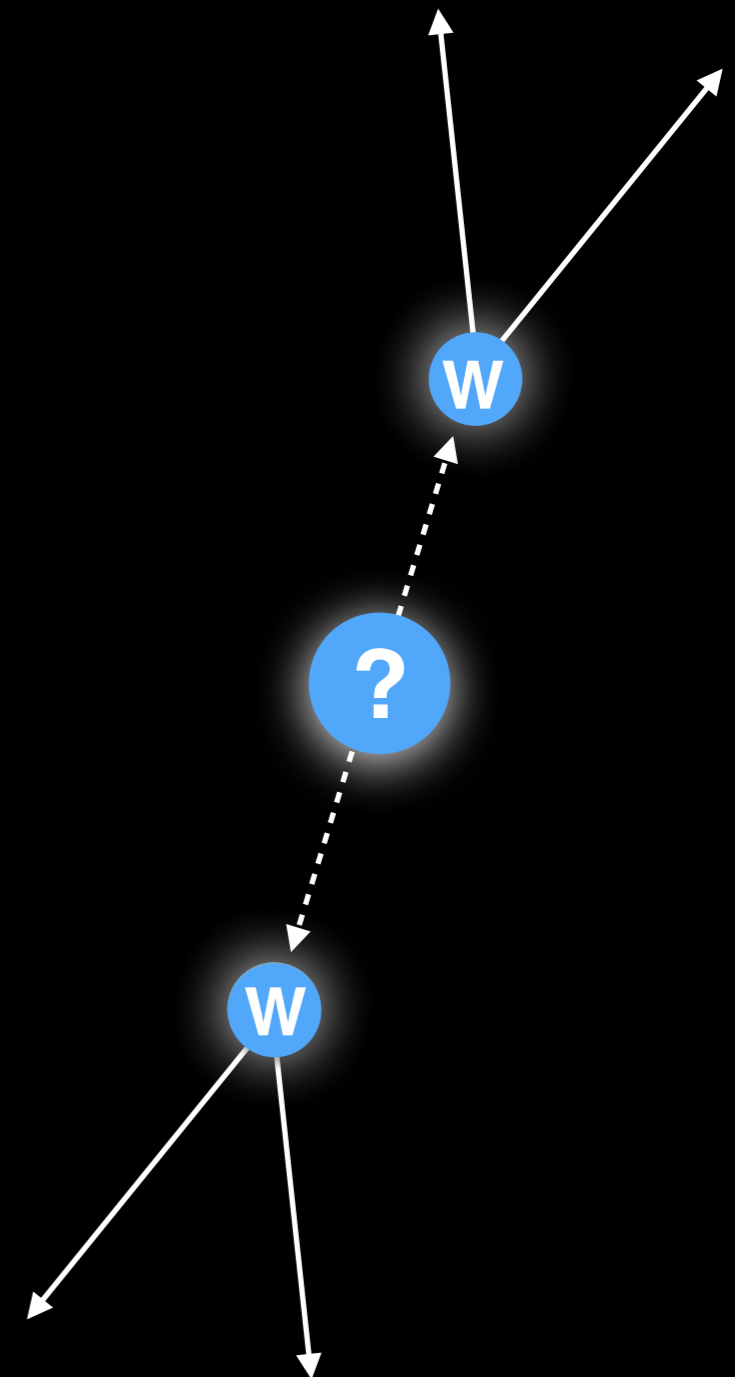
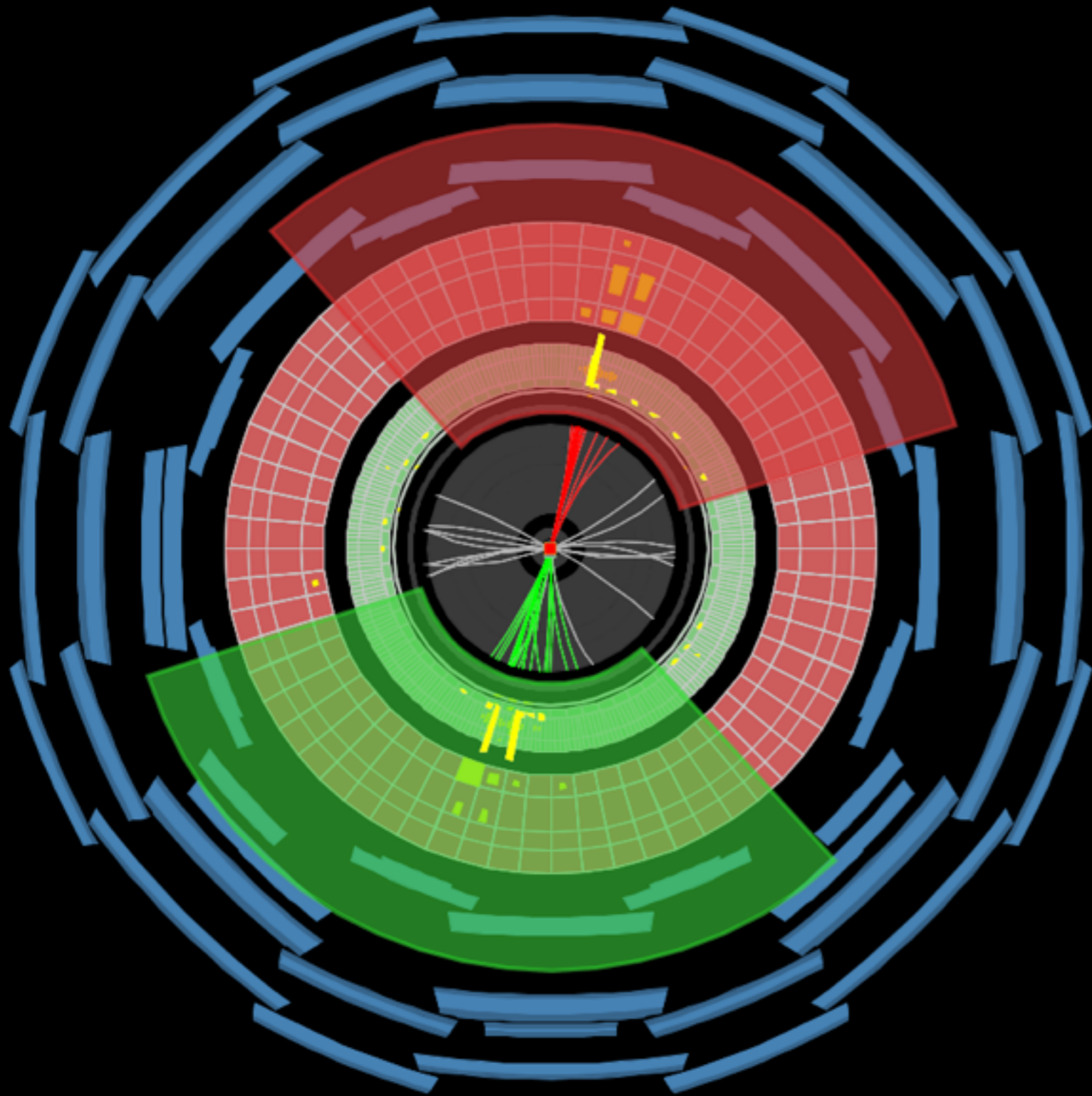
$$\phi \sim 1/\gamma = m/E$$



$$\gamma = E/m$$



W bosons are naturally boosted if they result from the decay of something even heavier

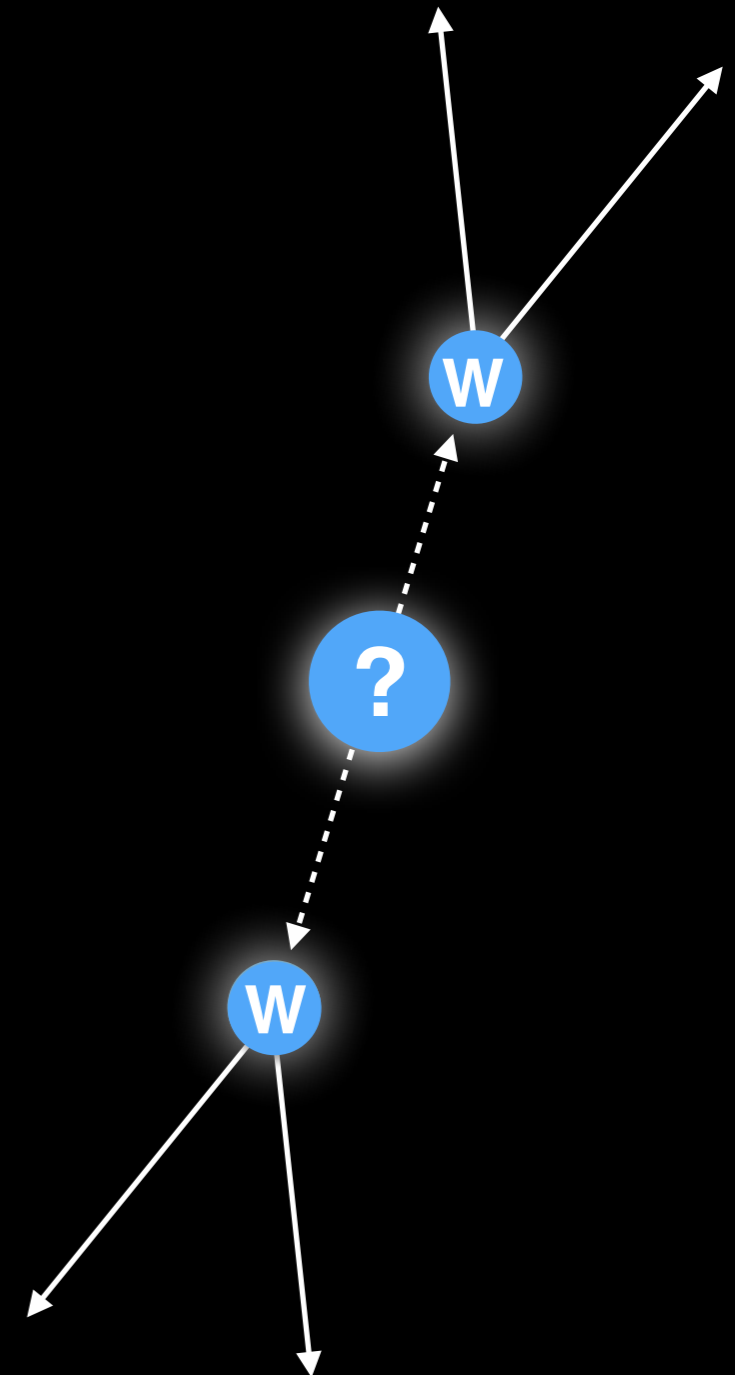


W bosons are naturally boosted if they result from the decay of something even heavier

Goal: Find W jets in an enormous sea of generic q/g jets

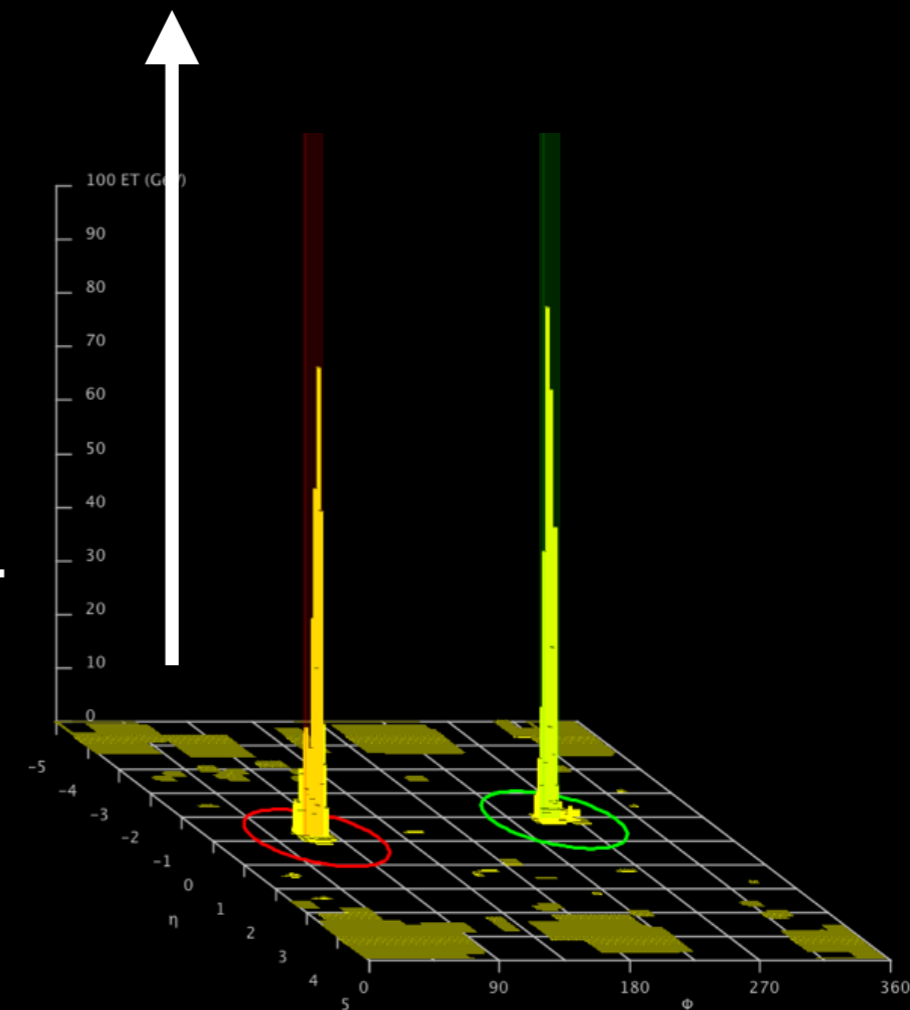
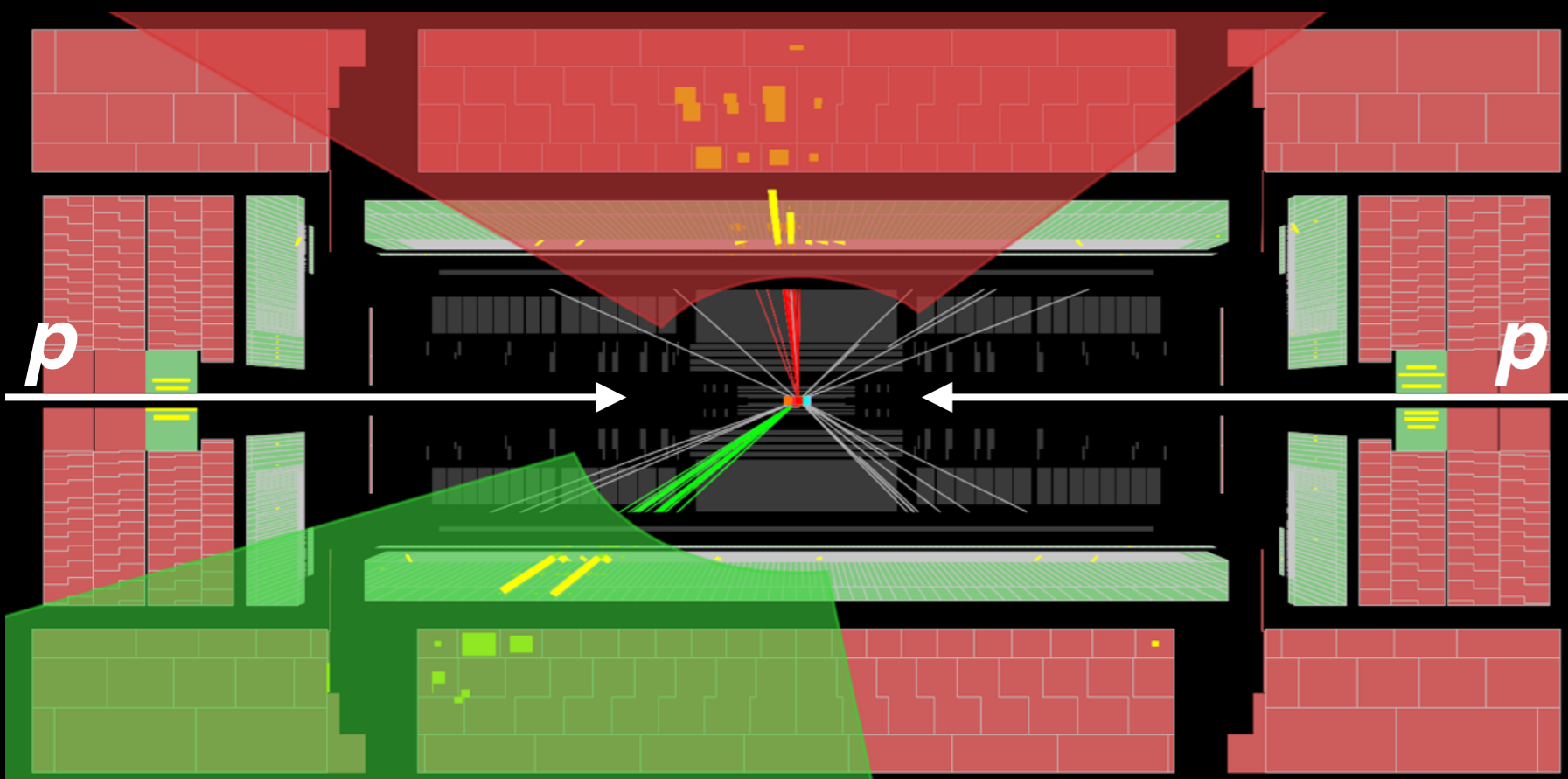
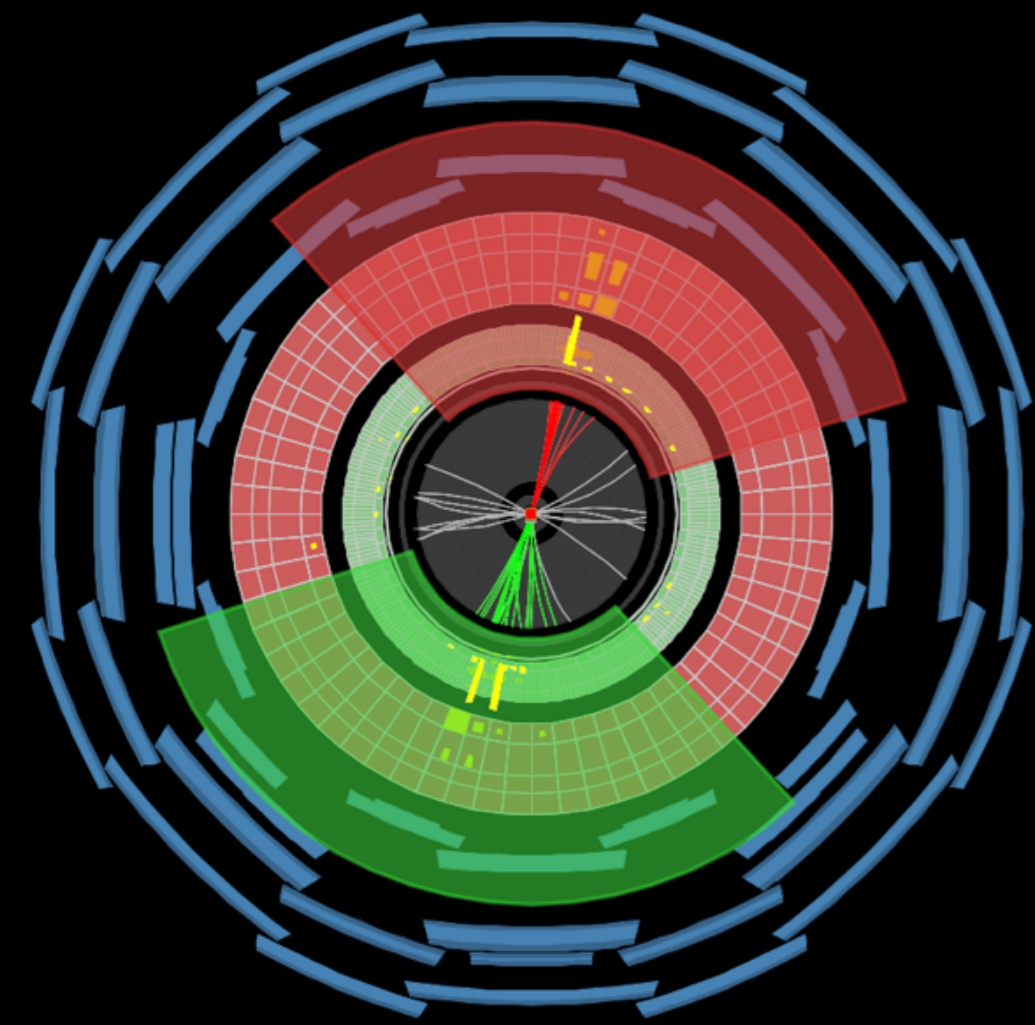


These jets have a non-trivial structure!

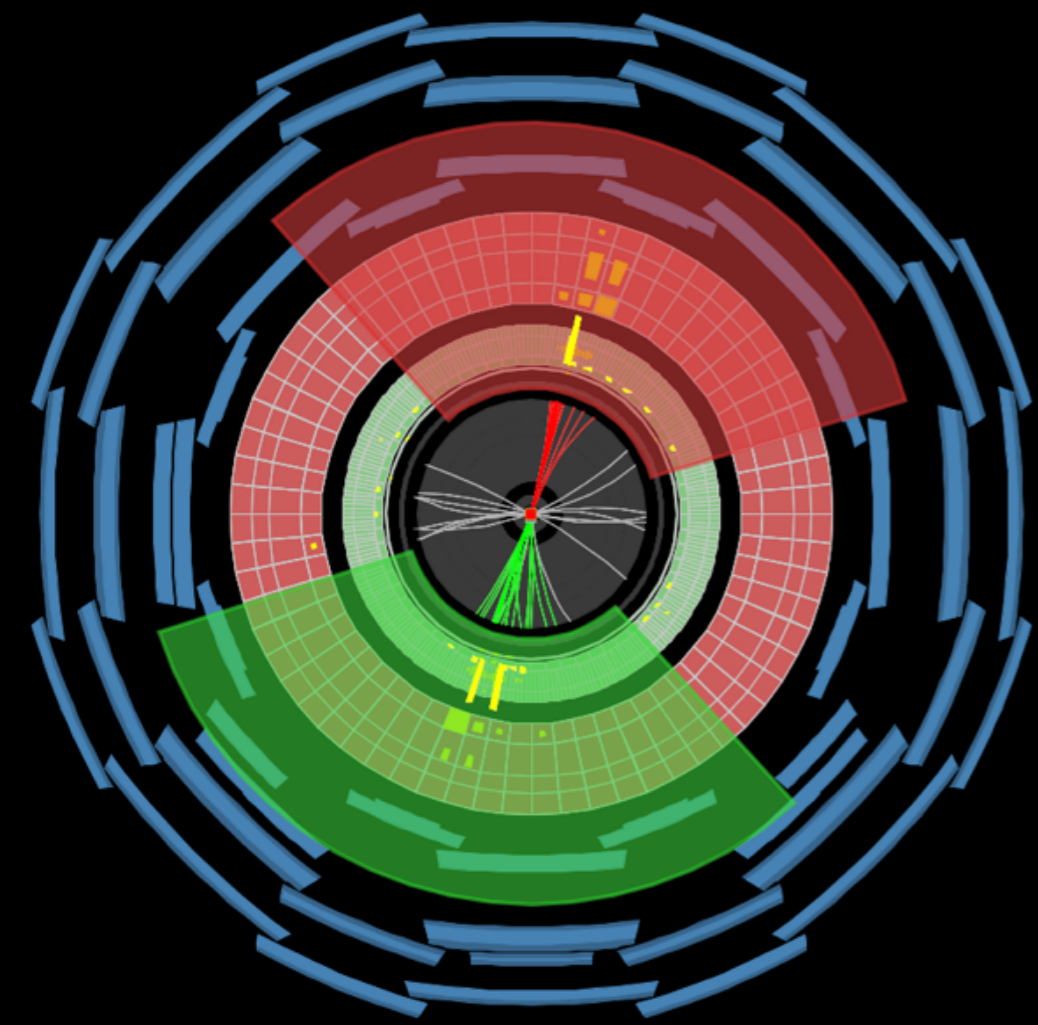


Searching for new particles decaying into boosted W bosons requires **looking at the radiation pattern inside jets**

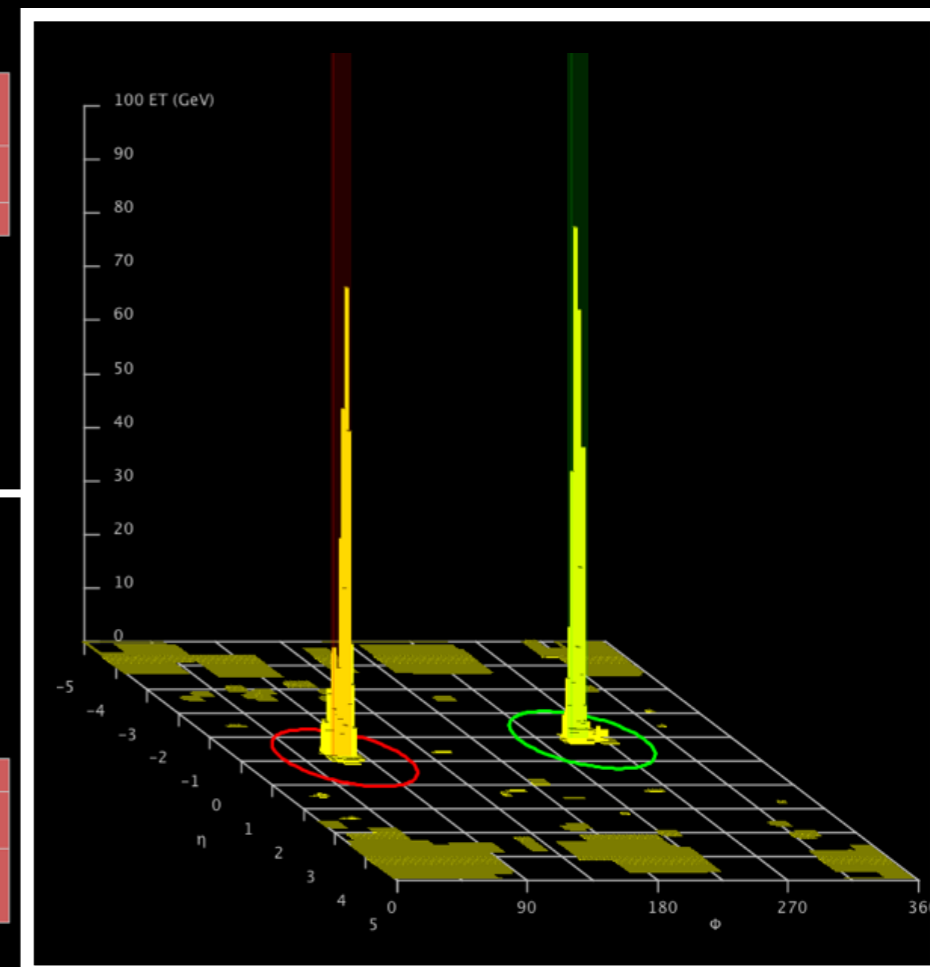
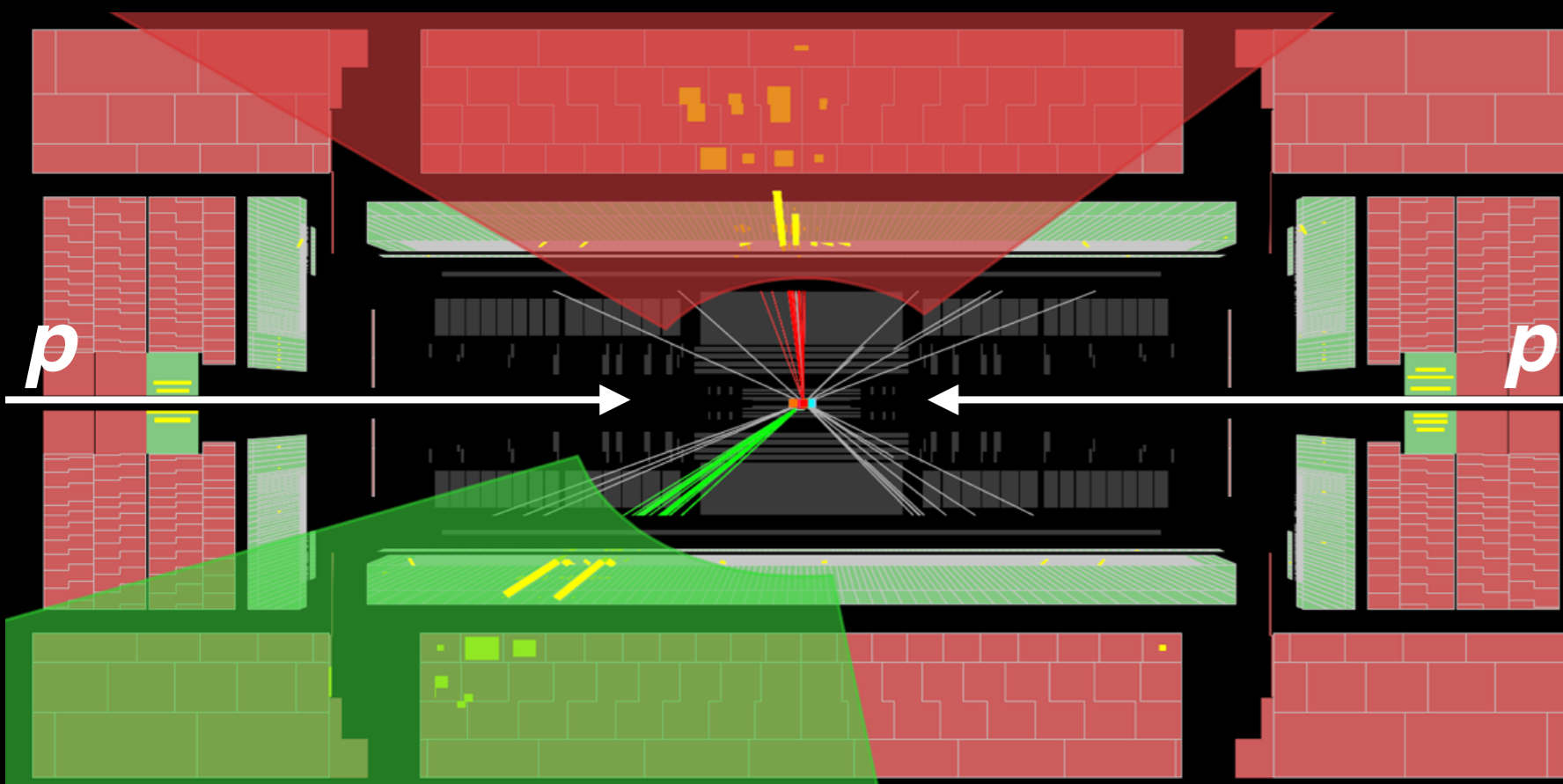
momentum transverse to the beam (p_T)



Up next: jet images

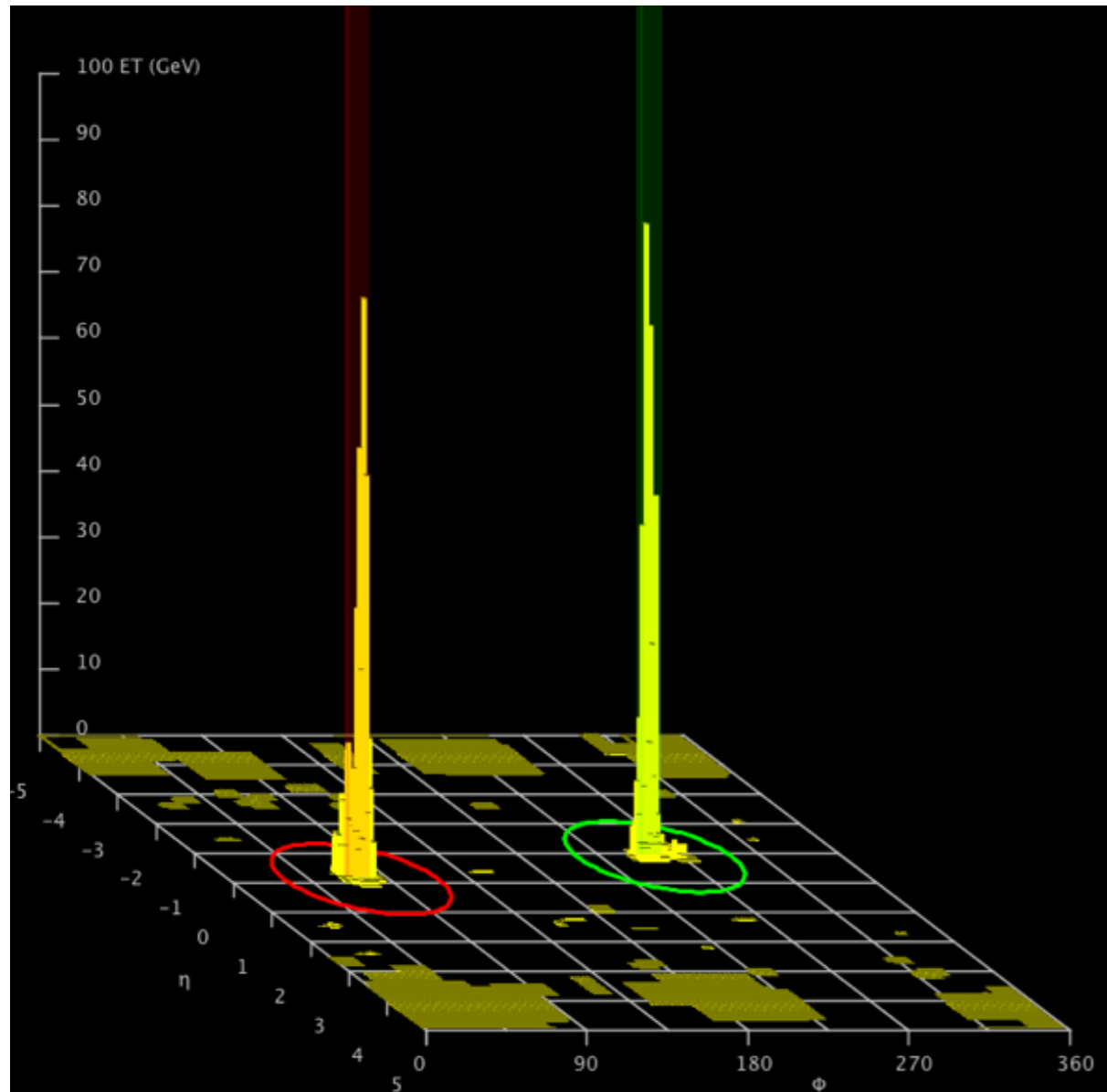


like a digital image!

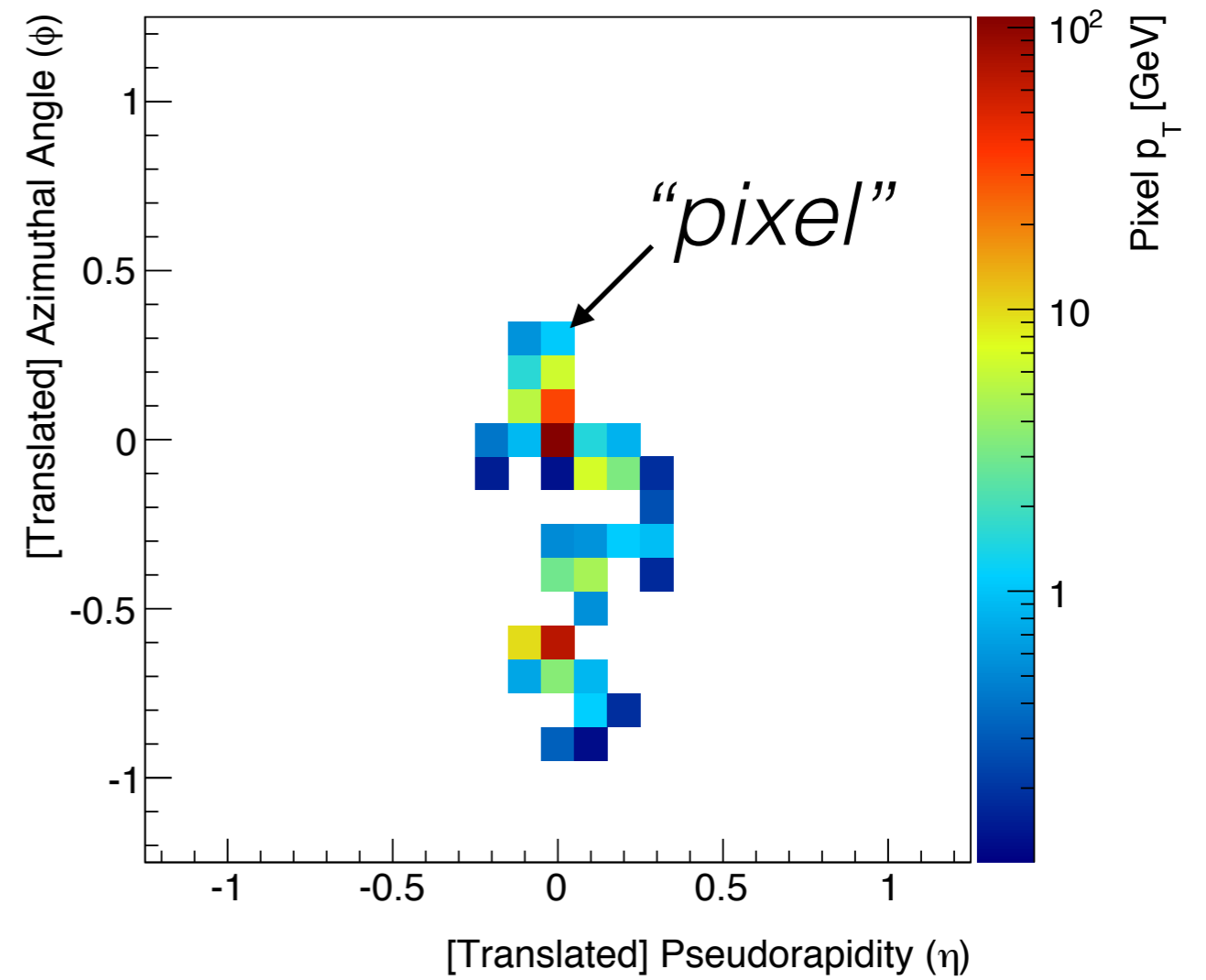


the Jet Image

J. Cogan et al. JHEP 02 (2015) 118



Boosted W

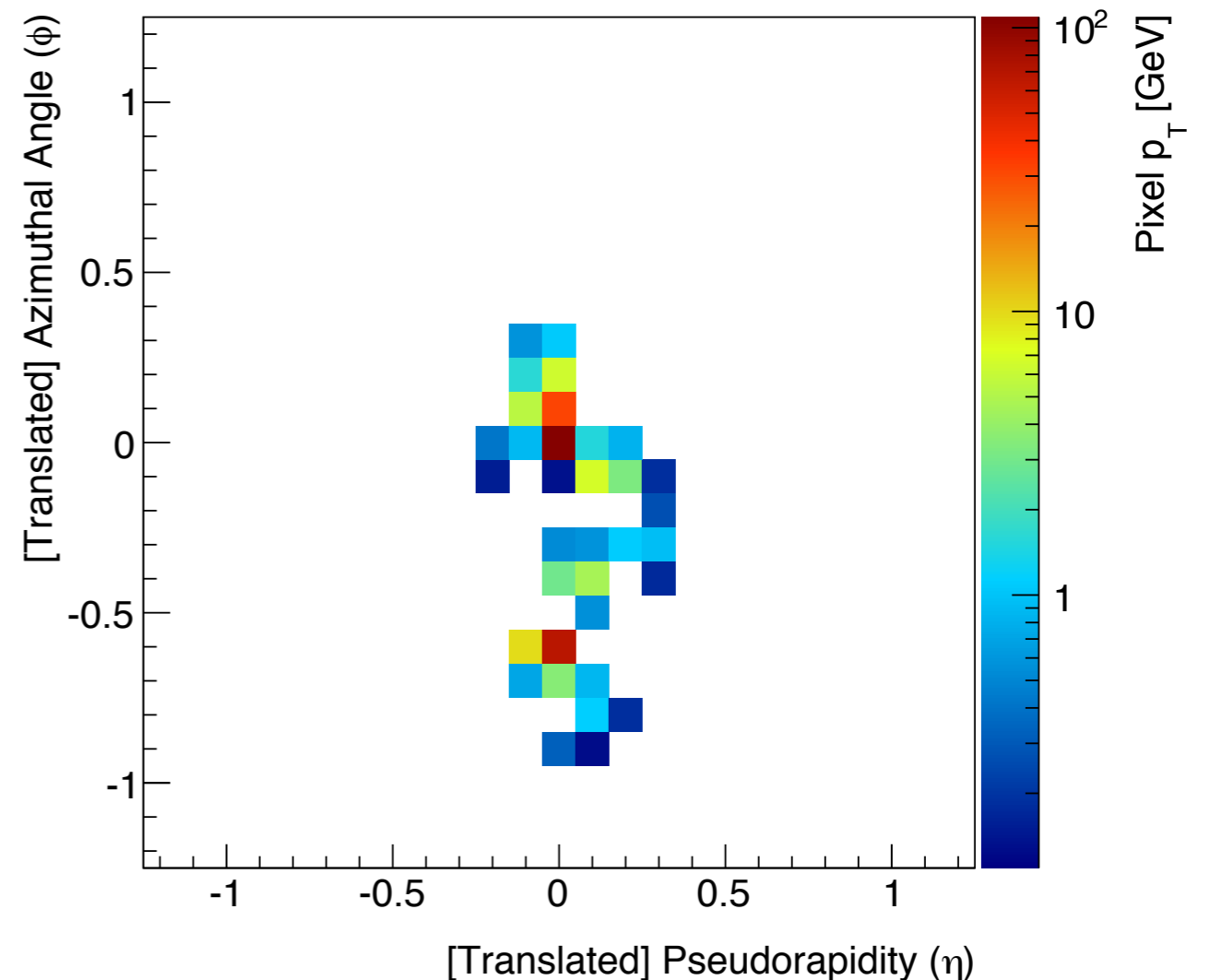


*nothing like a
'natural' image!*

the Jet Image

J. Cogan et al. JHEP 02 (2015) 118

Boosted W

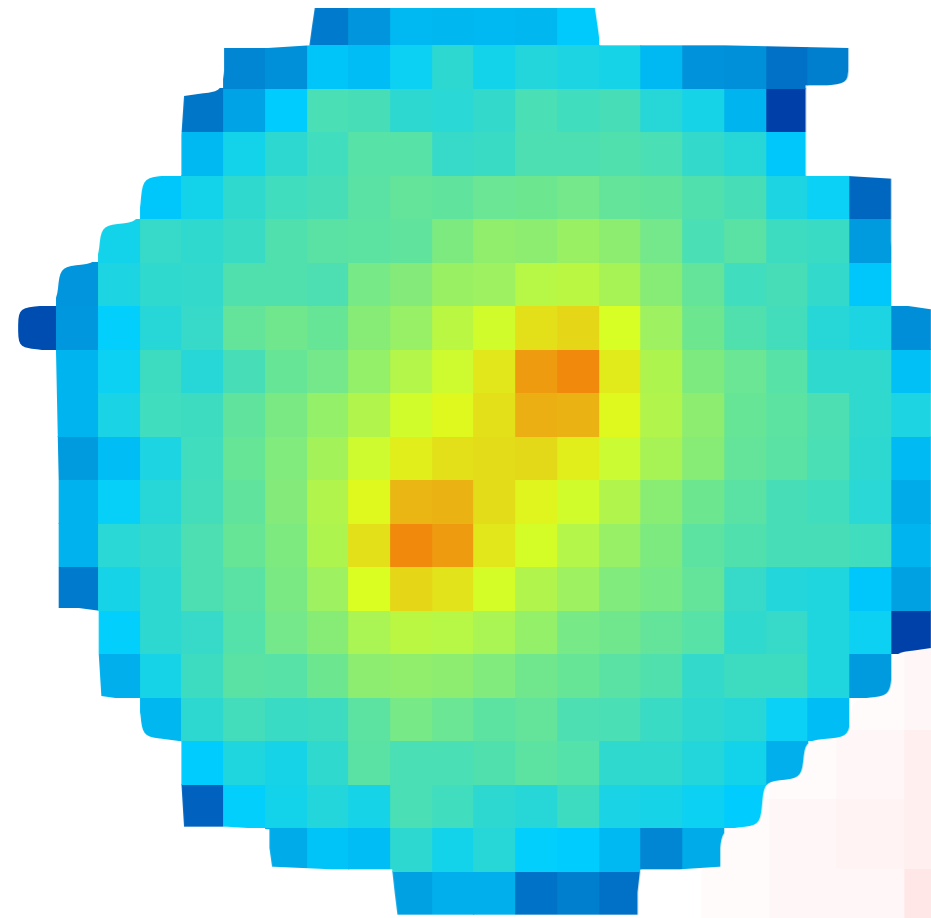


Credit: Peter G Trimming (Wikipedia)

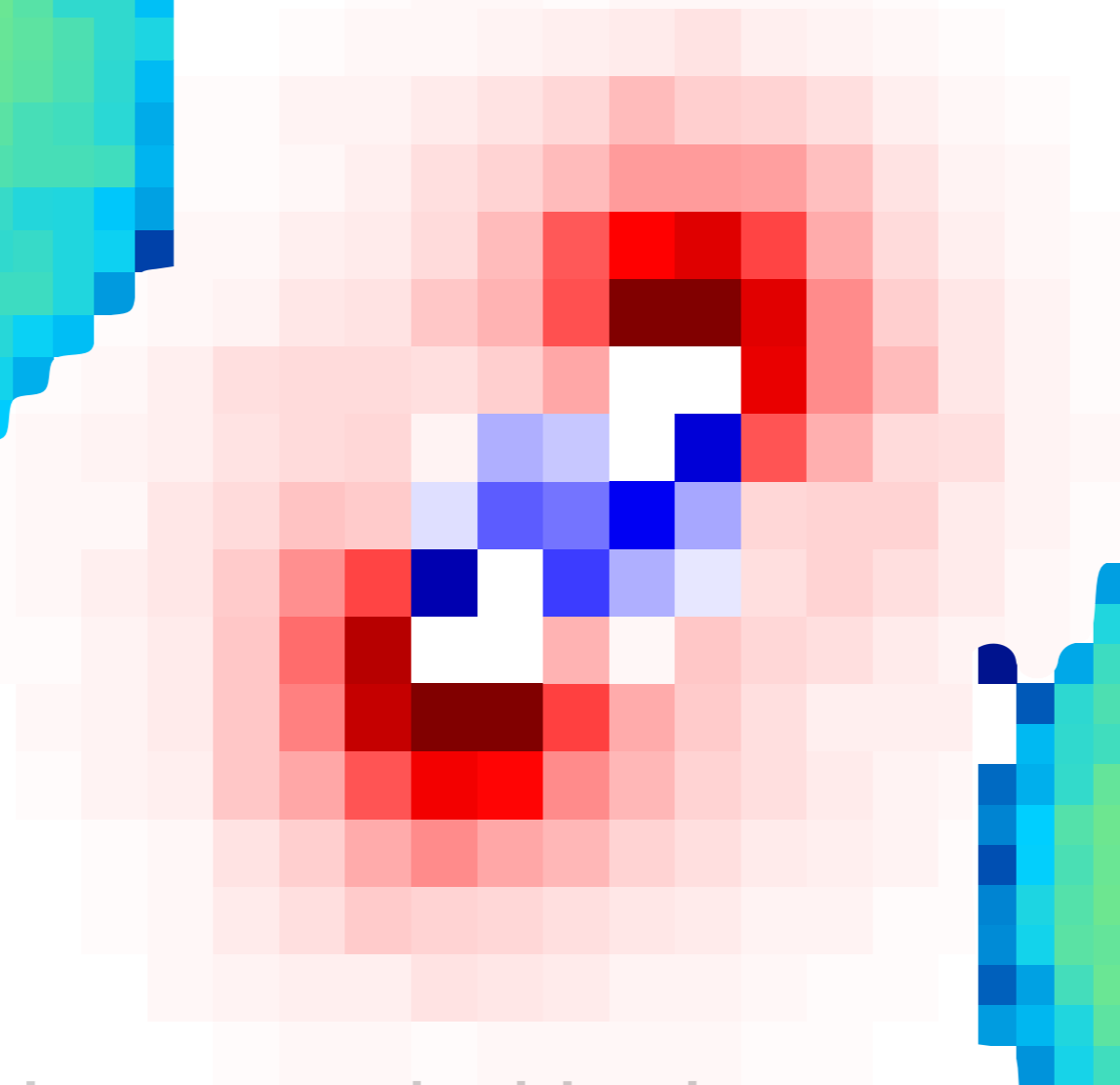
*no smooth edges, clear features, low
occupancy (number of hit pixels)*

Can directly visualize physics

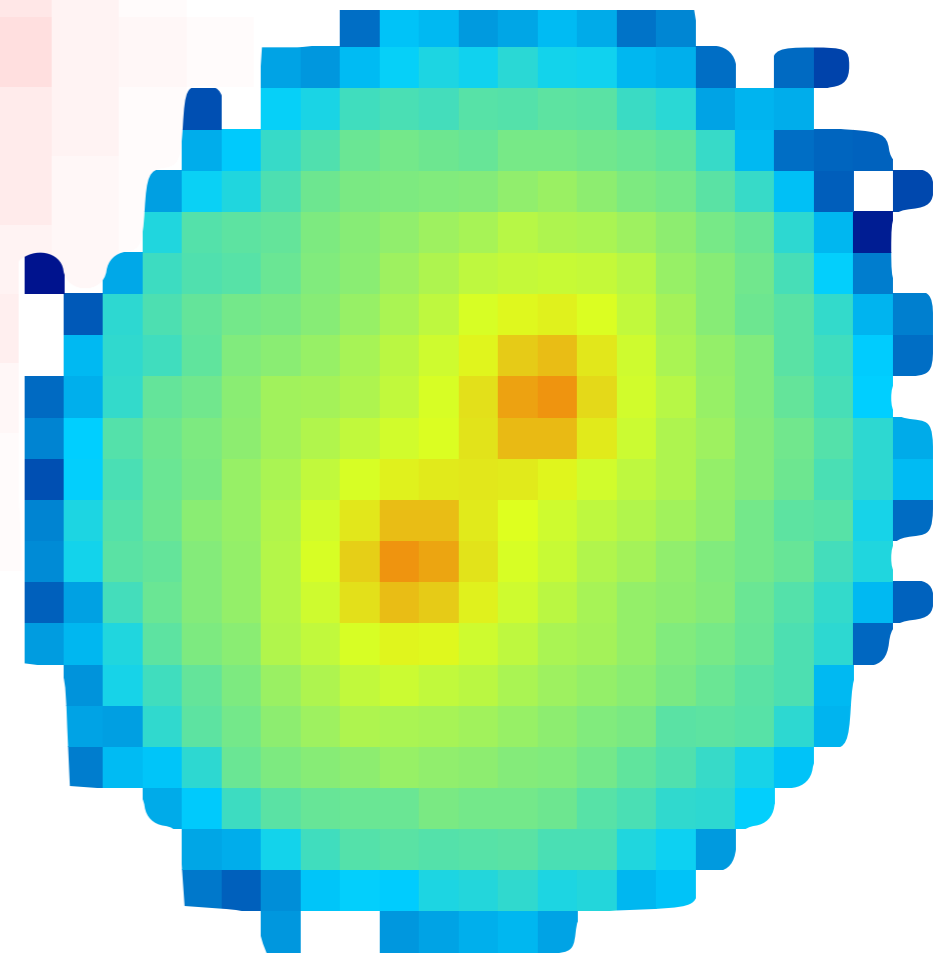
and we can benefit from the extensive image processing literature



$W \rightarrow q\bar{q}$



$g \rightarrow q\bar{q}$



there is information encoded in the physical distance between pixels

Can directly visualize physics

and we can benefit from the
extensive image processing literature

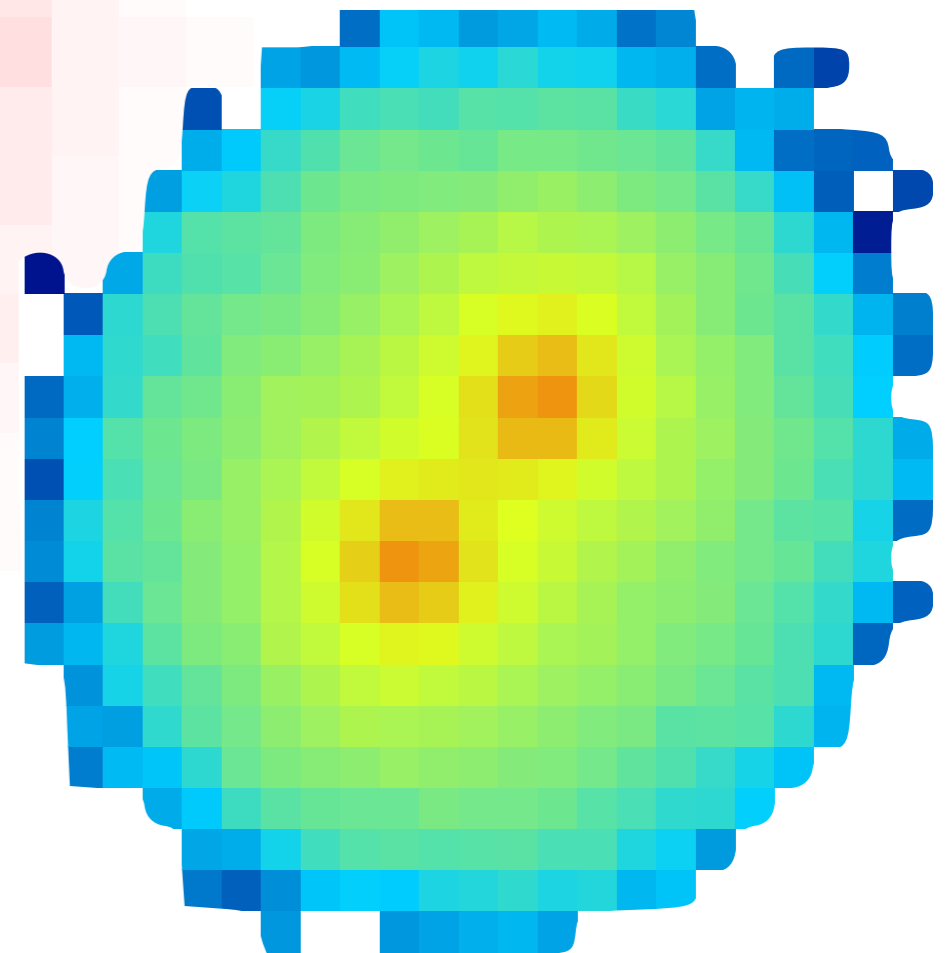
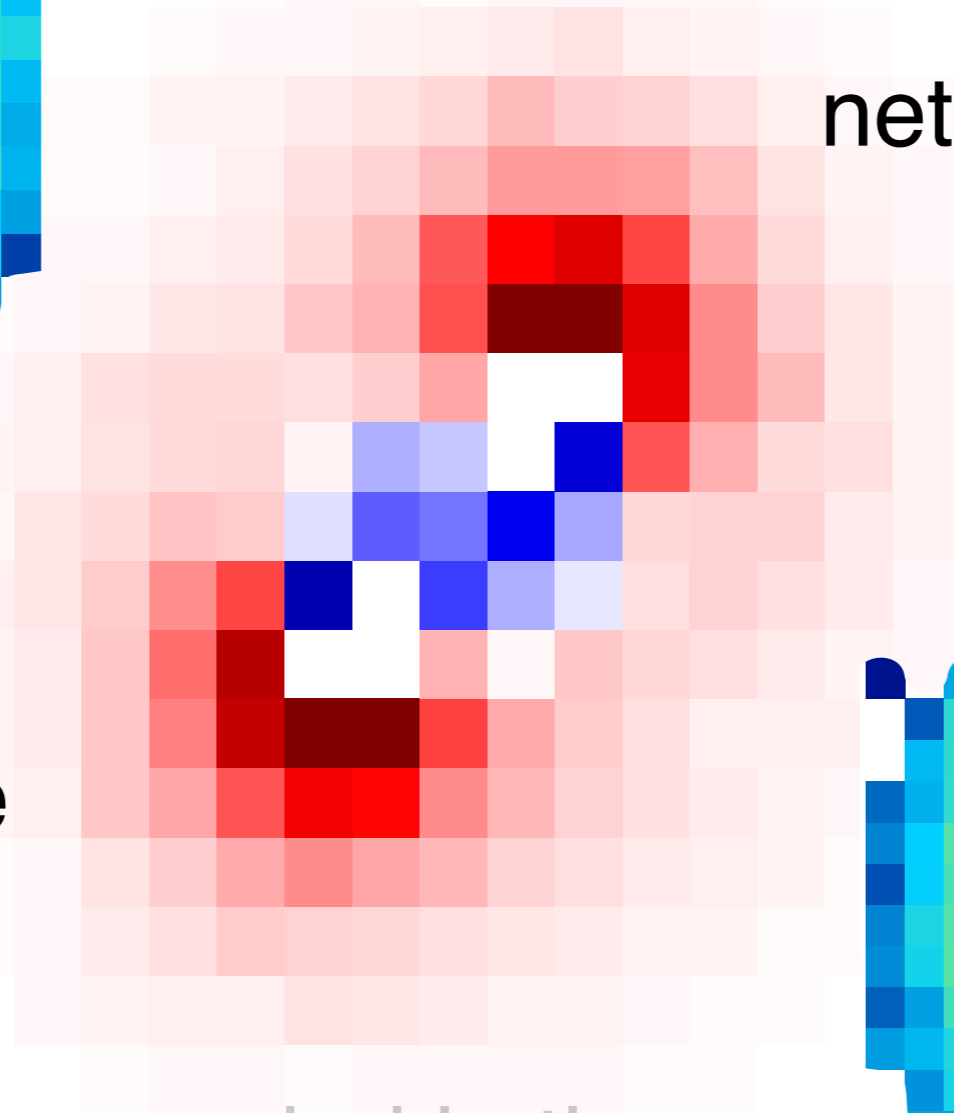
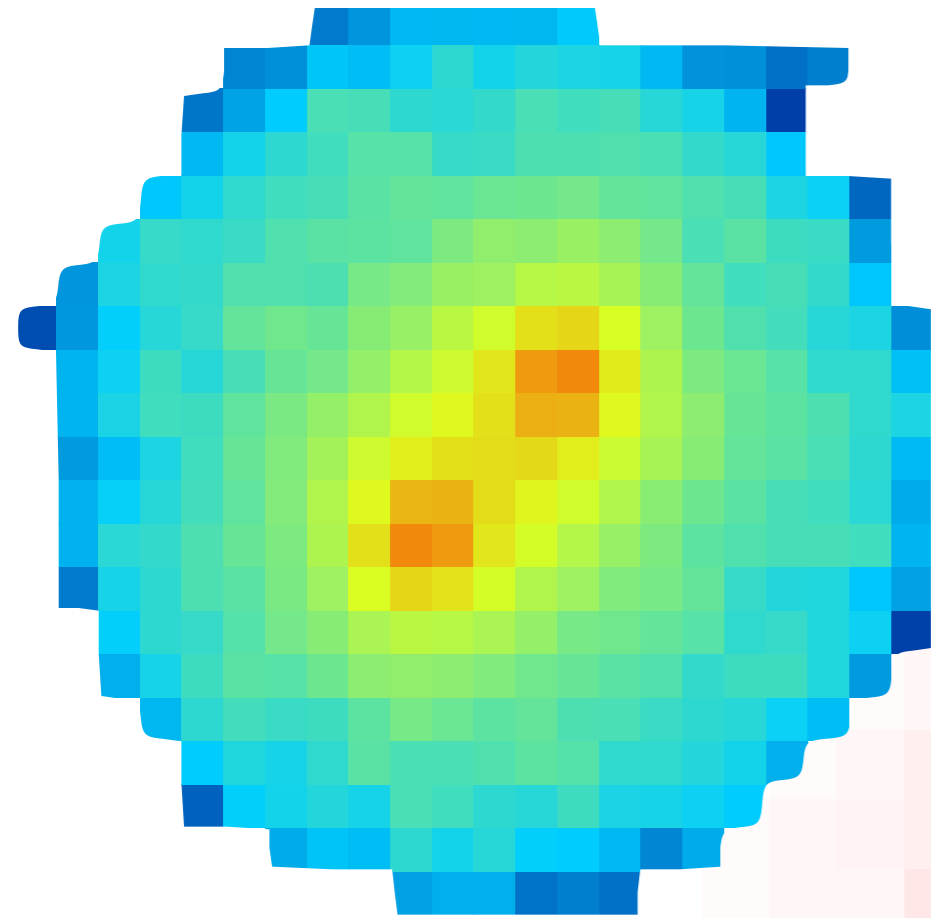
net strong-force charge

$g \rightsquigarrow q\bar{q}$

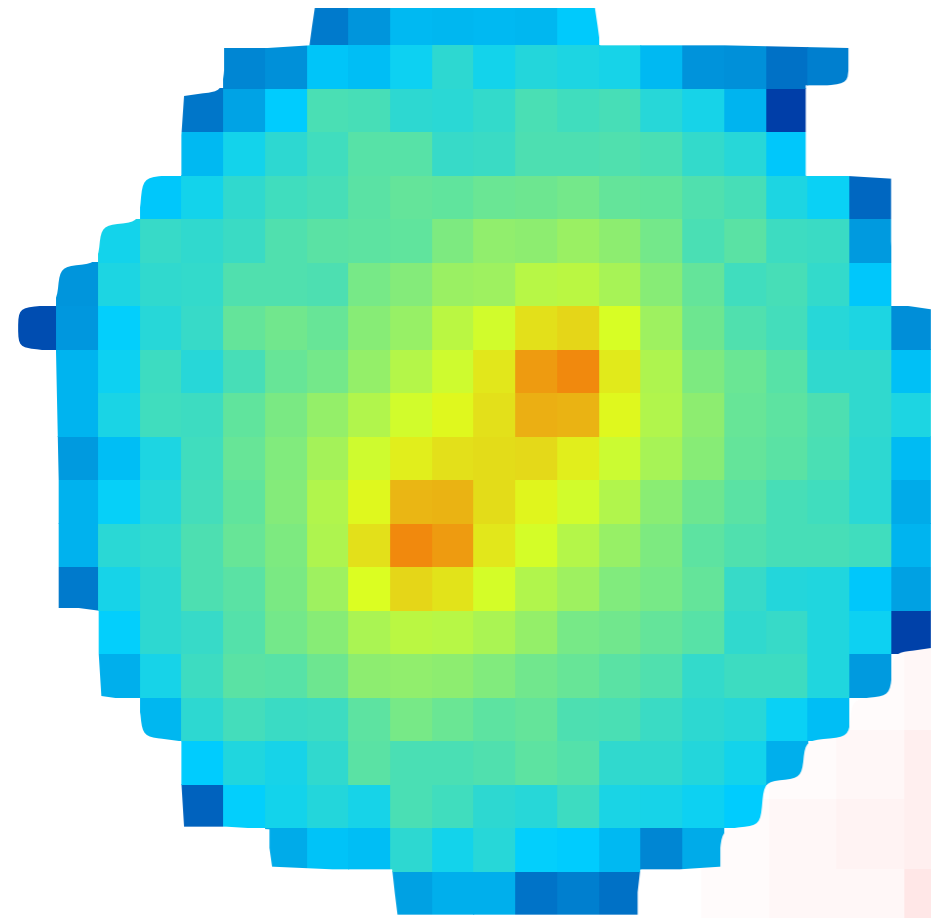
$W \rightsquigarrow q\bar{q}$

radiates like a dipole
(no net charge)

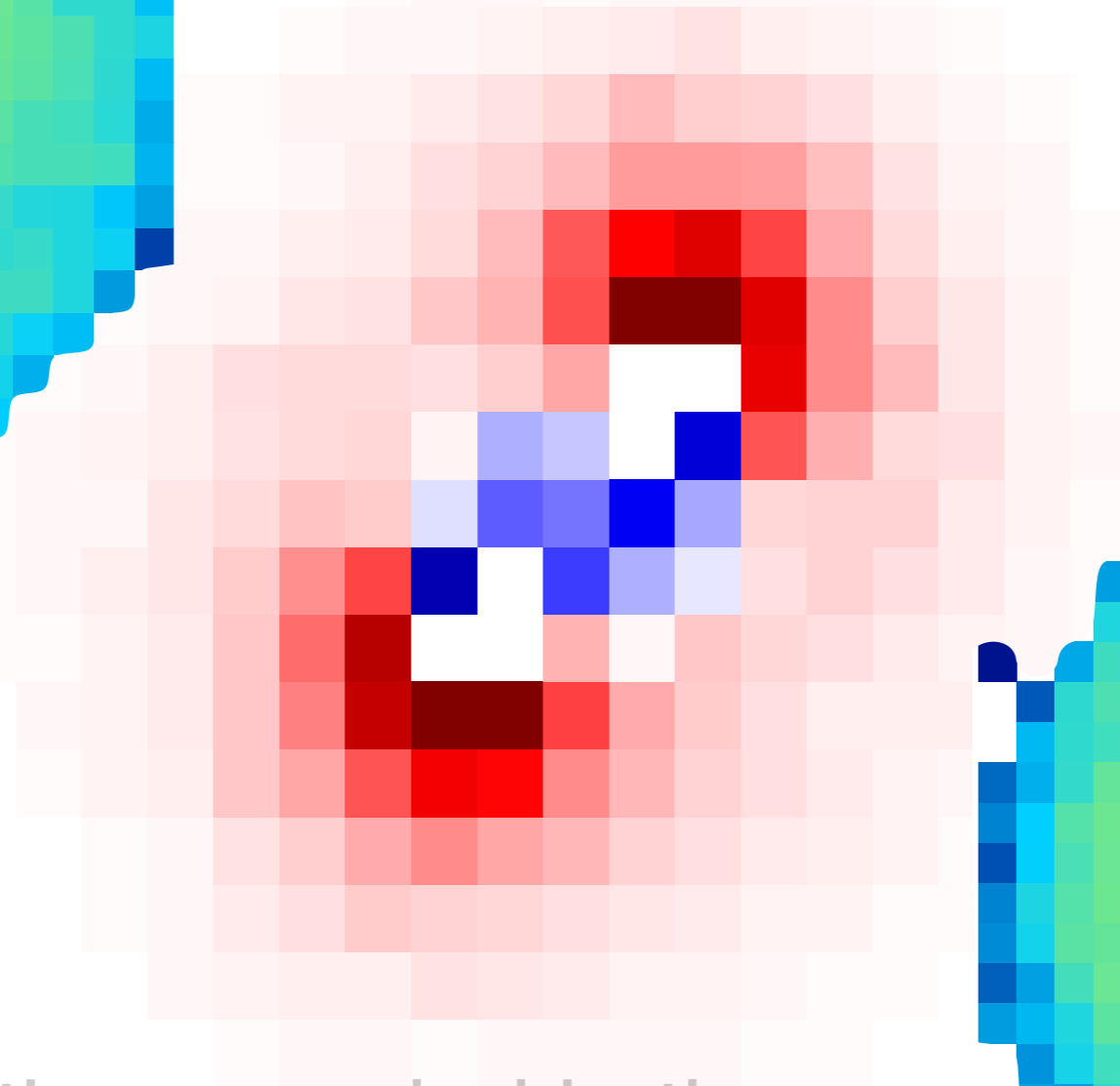
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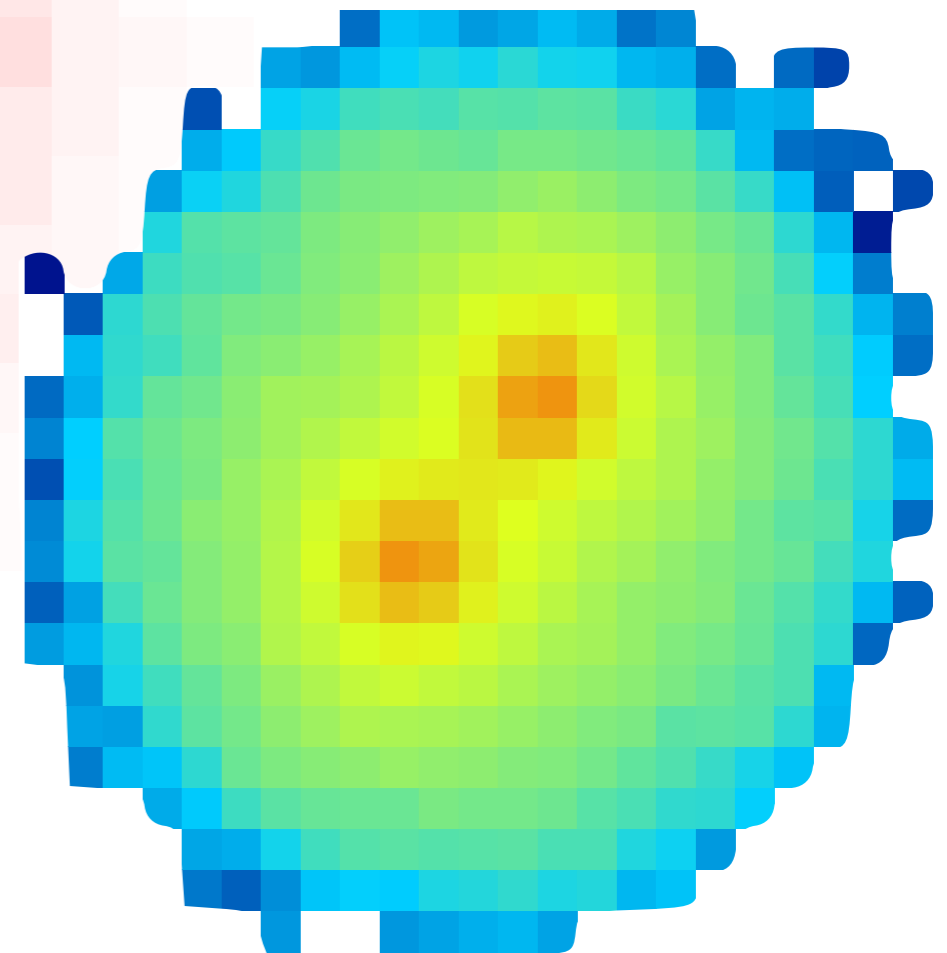
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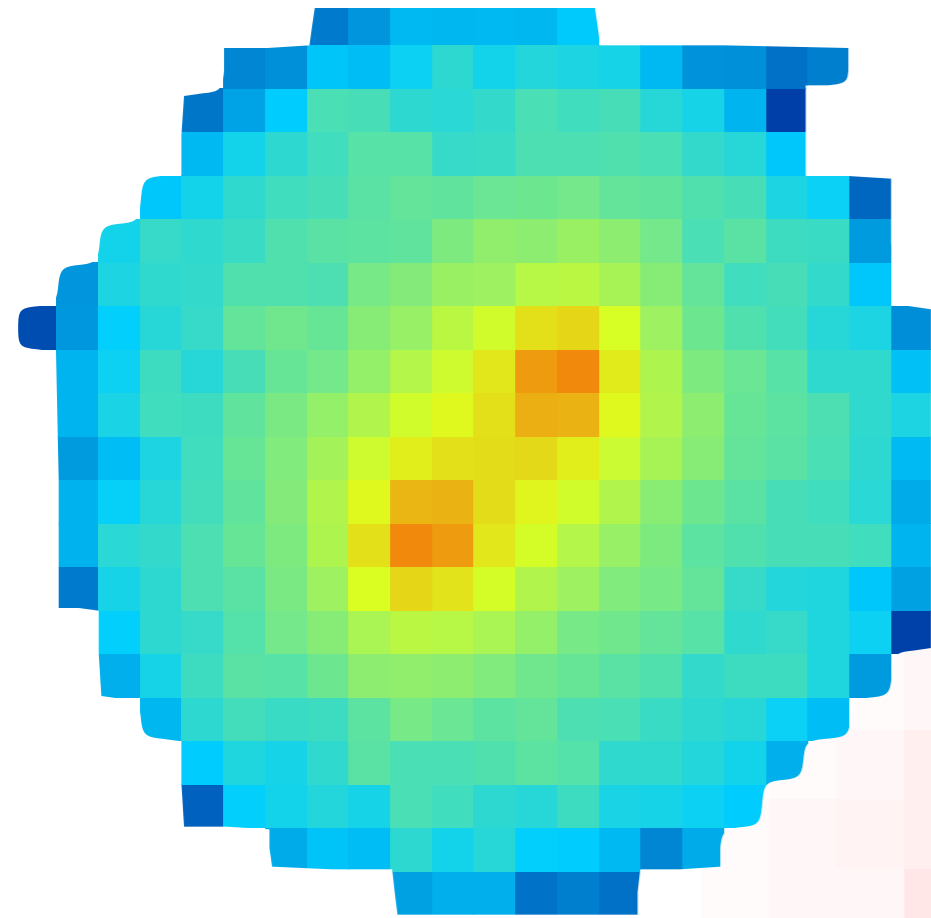


$g \rightarrow q\bar{q}$

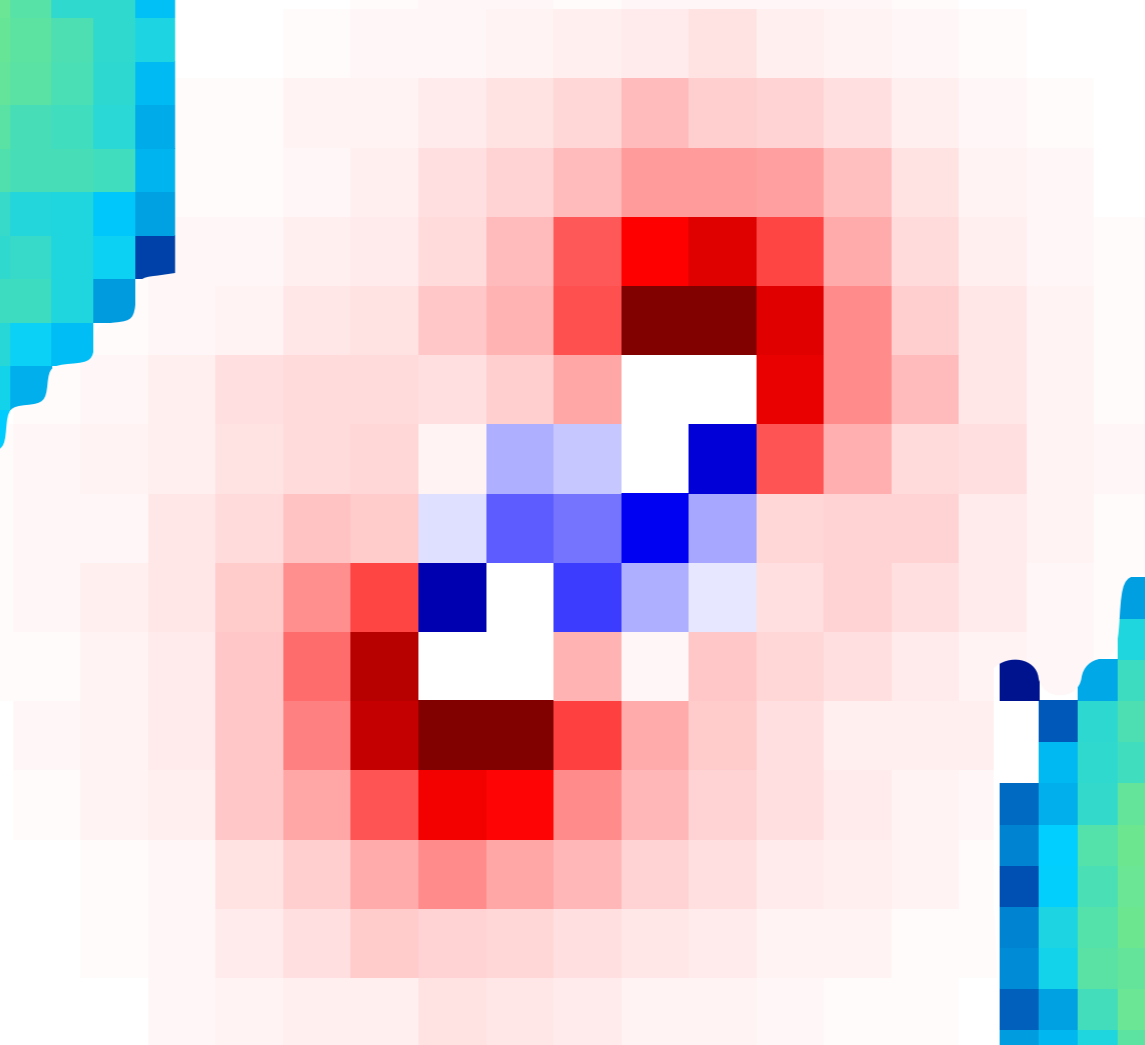


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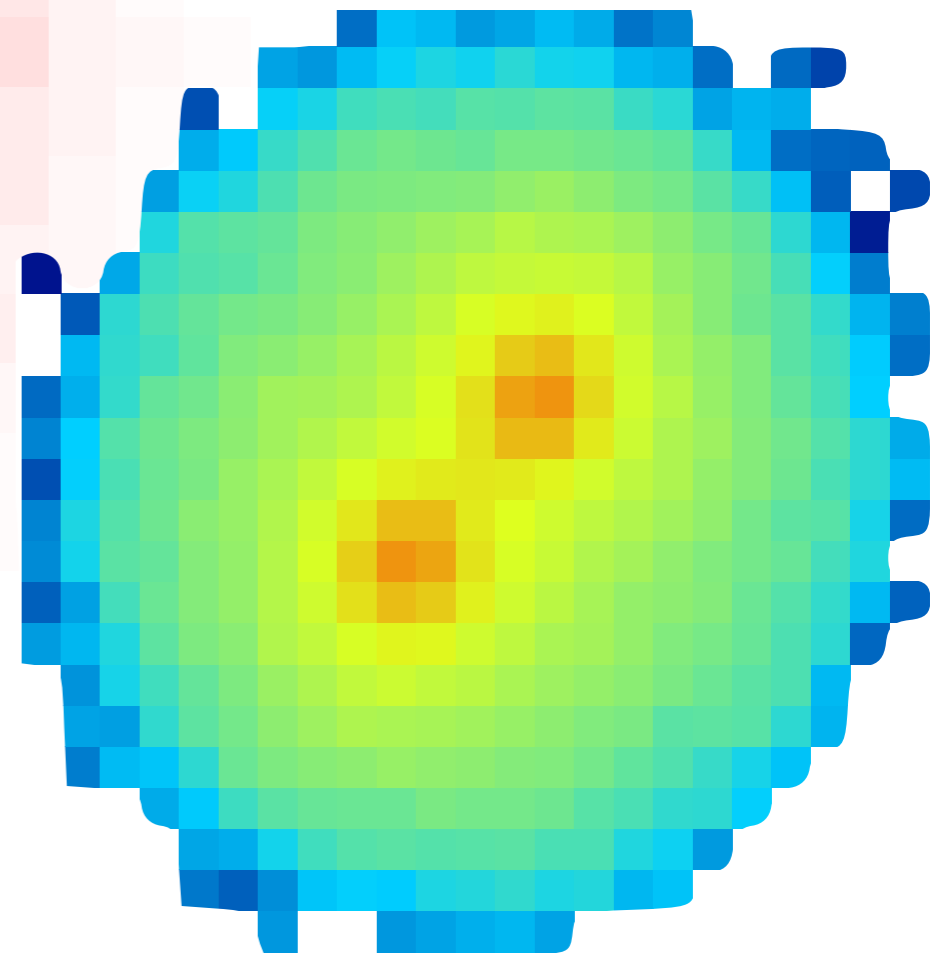
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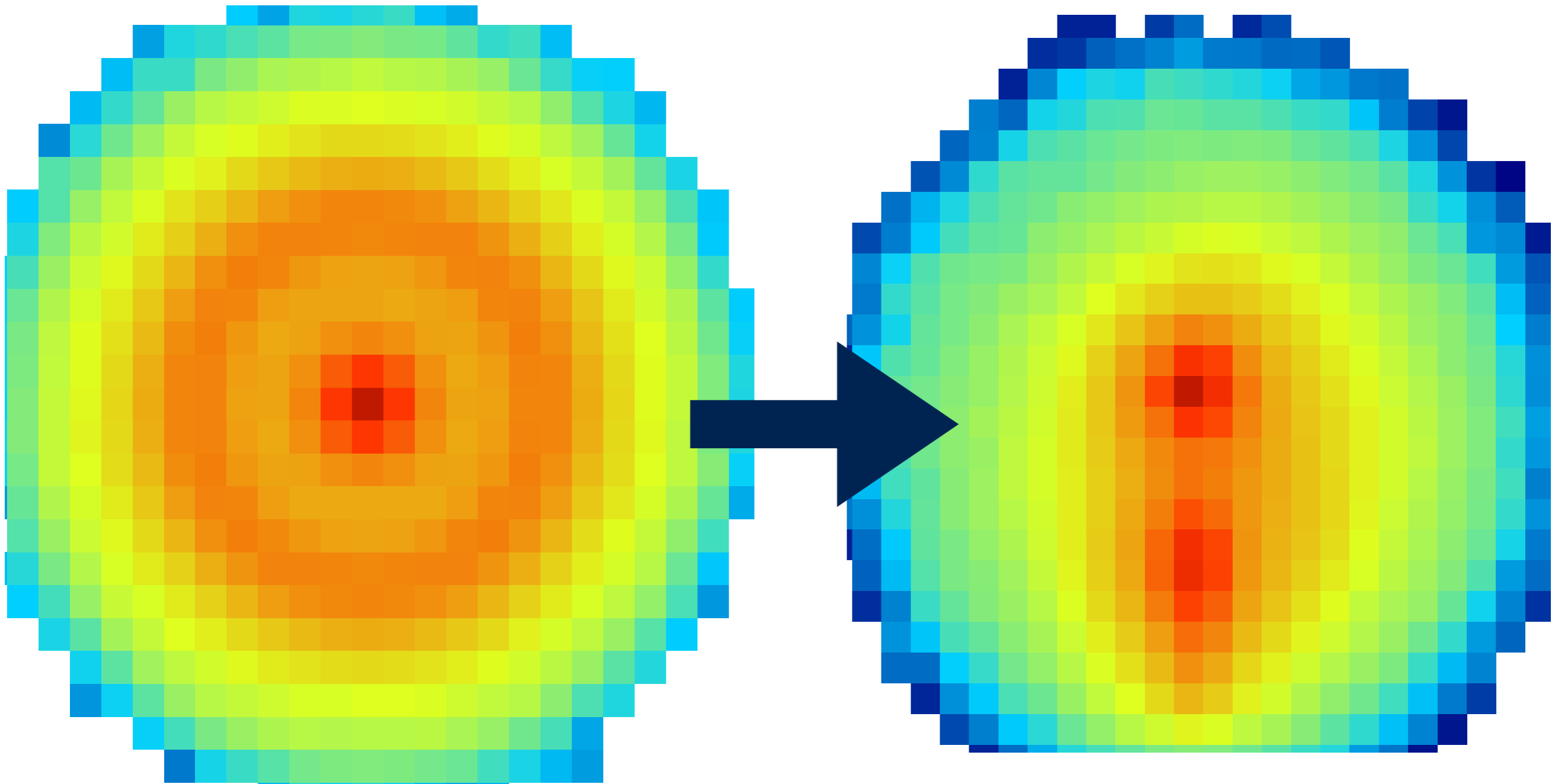


$g \rightarrow q\bar{q}$



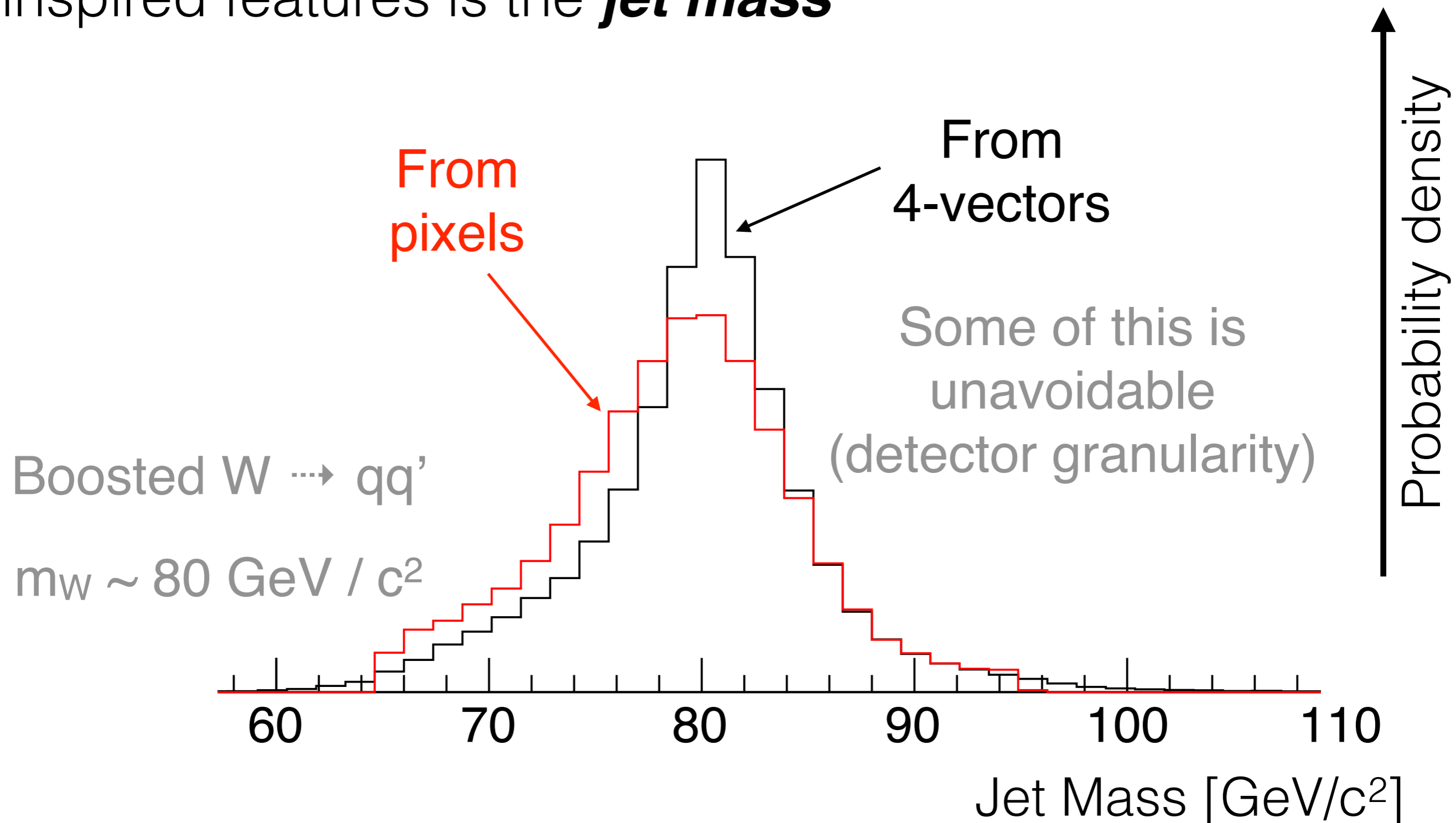
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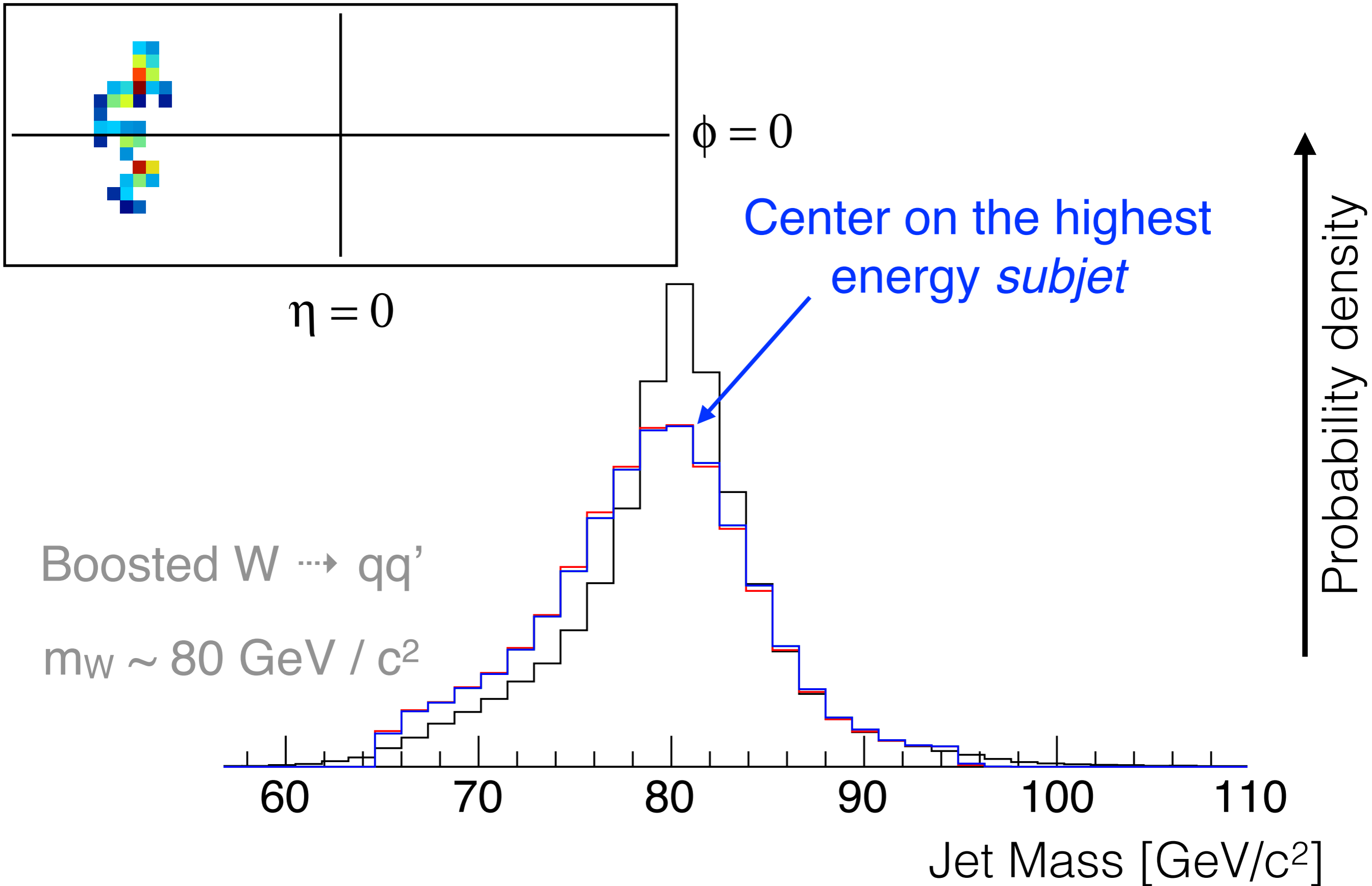
One of the first typical steps is pre-processing

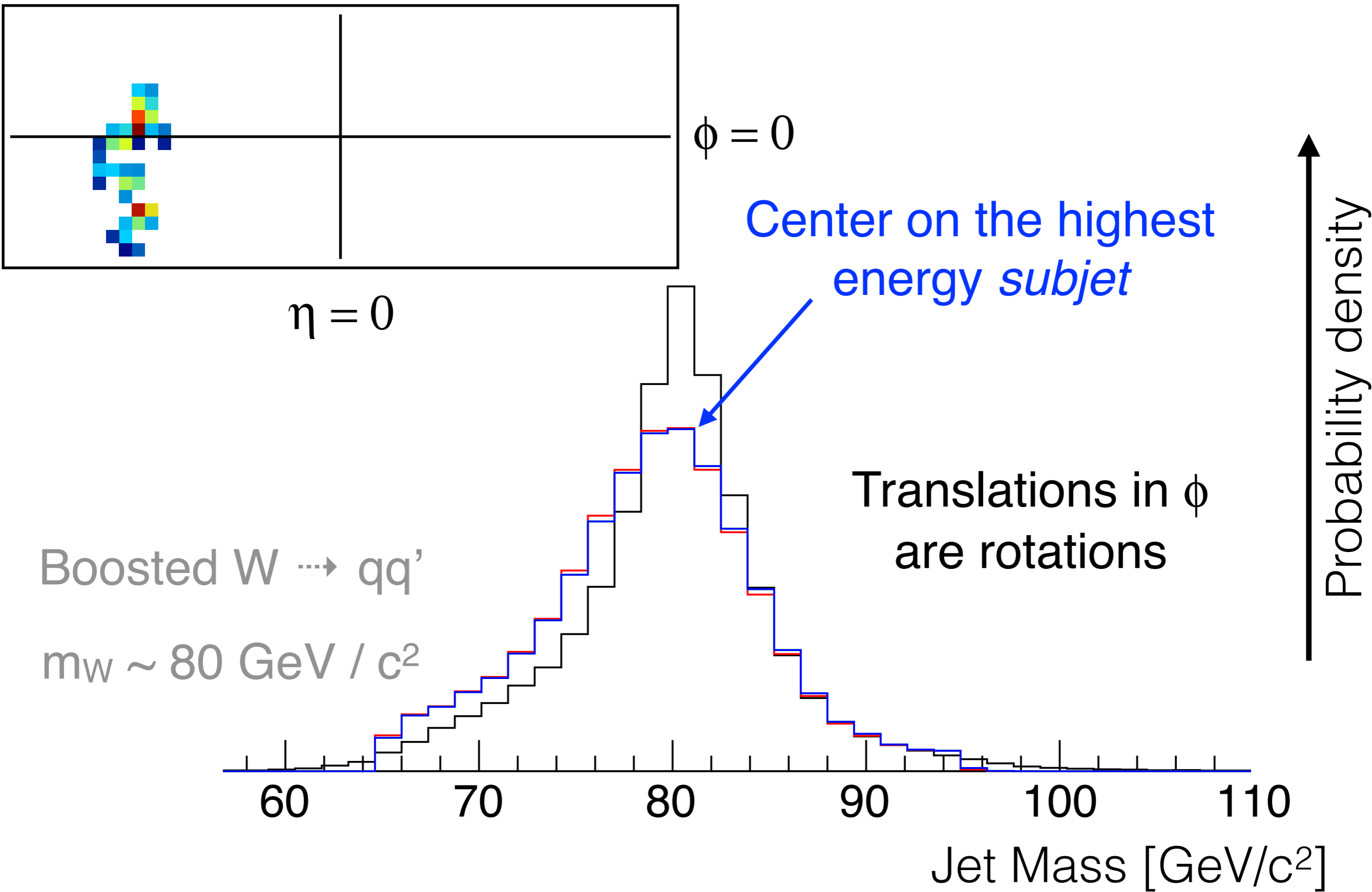


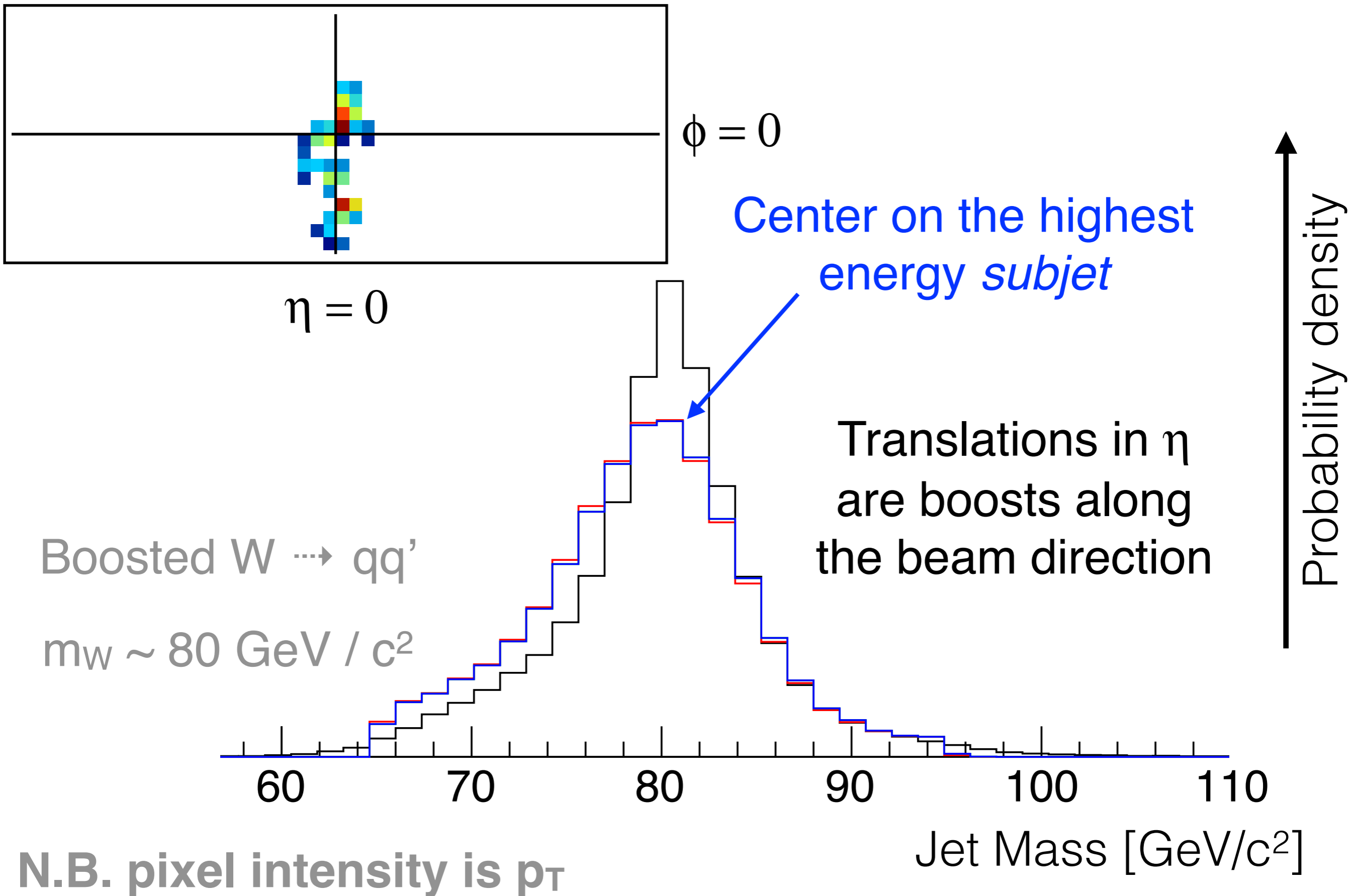
Can help to learn faster & smarter; but must be careful!

One of the most useful physics-inspired features is the **jet mass**

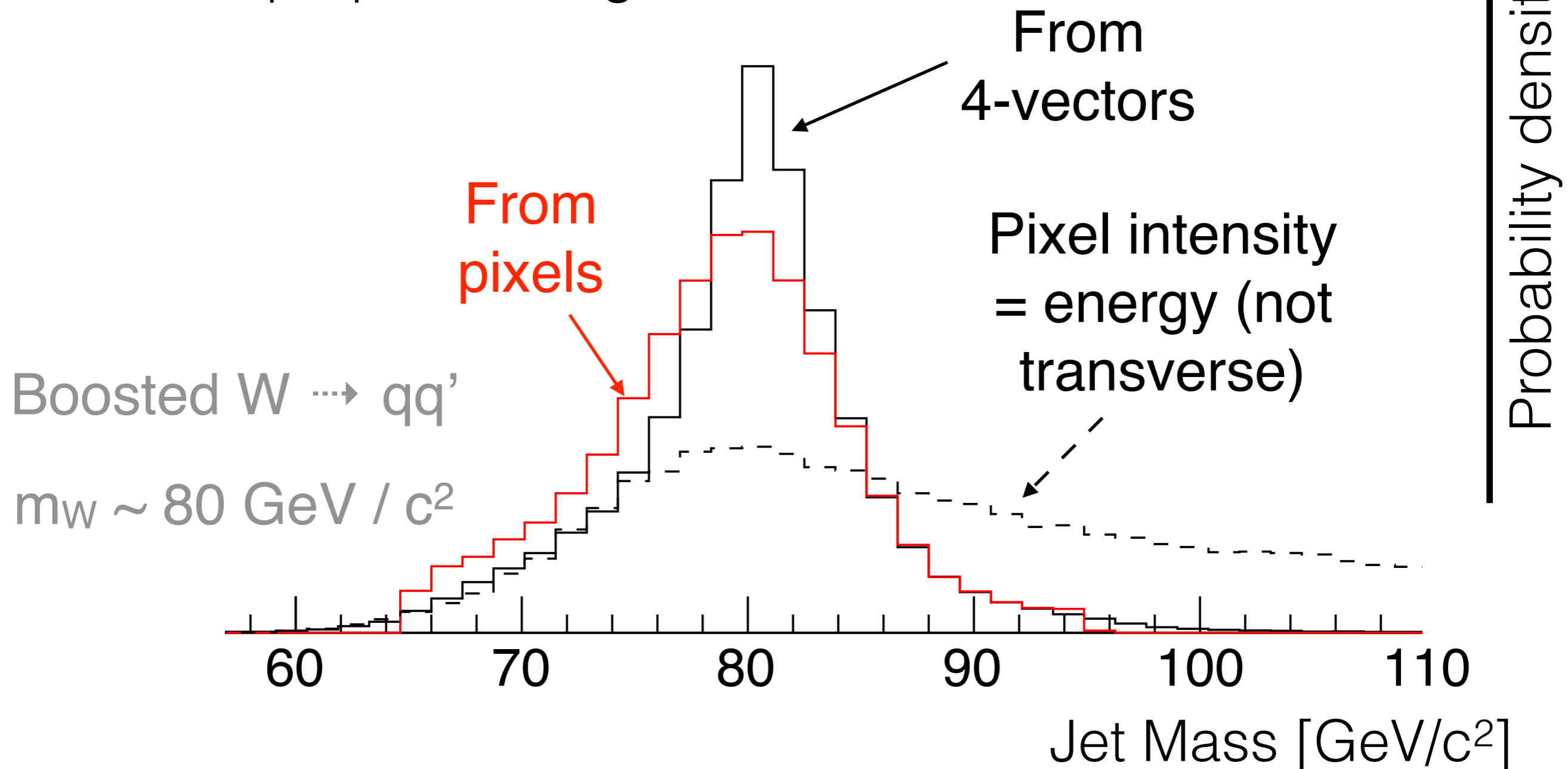




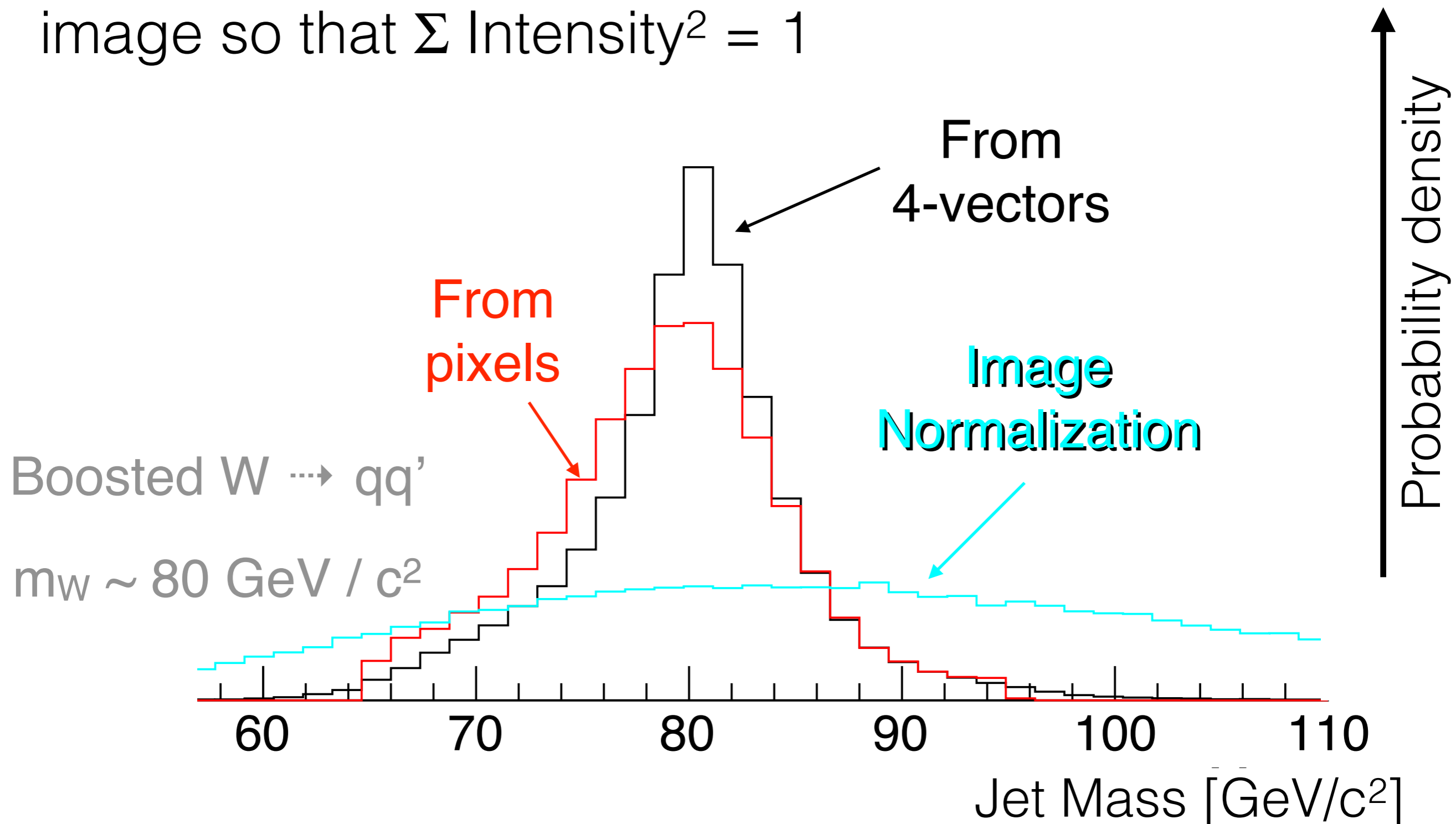




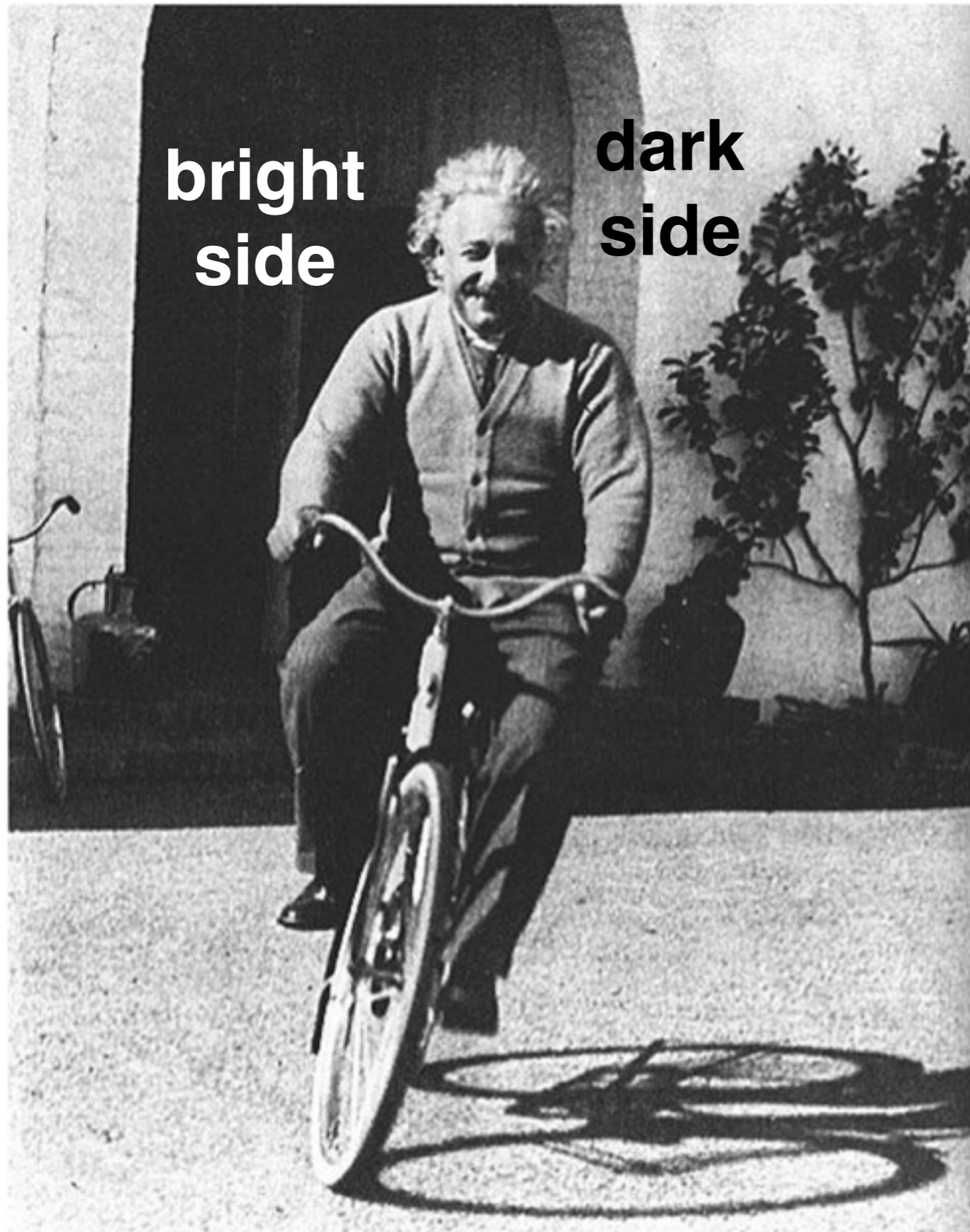
Information can be **washed out** without care in preprocessing



It is common to normalize each image so that $\sum \text{Intensity}^2 = 1$



Intuition via analogy *why normalization can hurt*

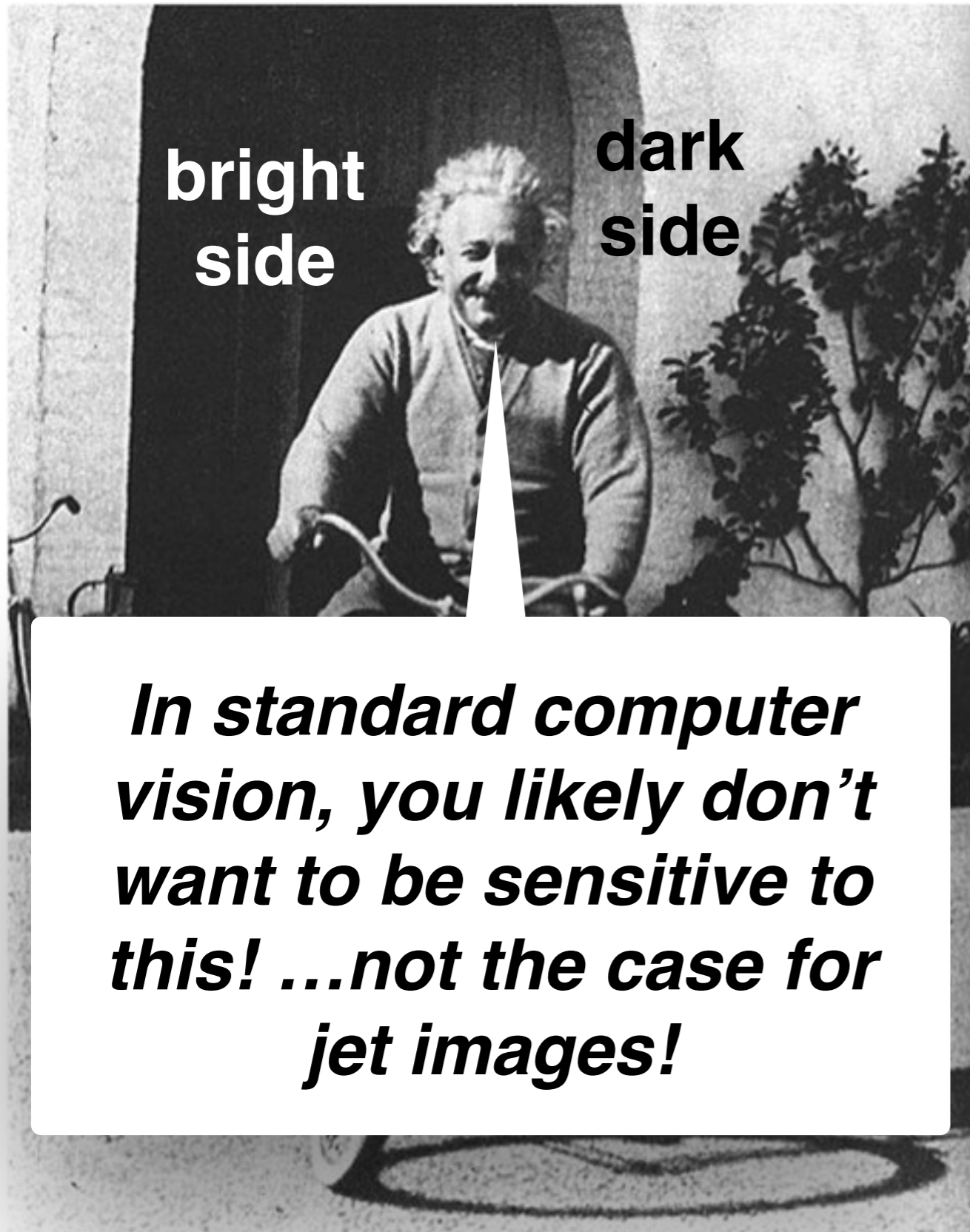


In both pictures, total intensity of Einstein's face is about the same.



However, his face's **image mass** is quite different!

Intuition via analogy *why normalization can hurt*

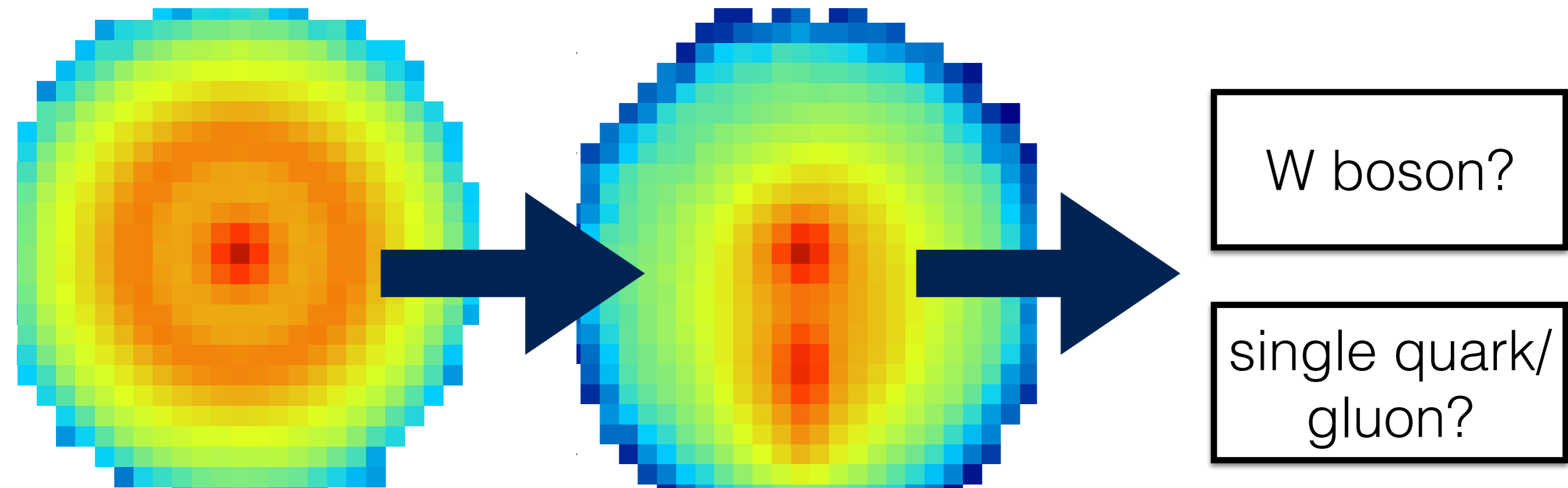


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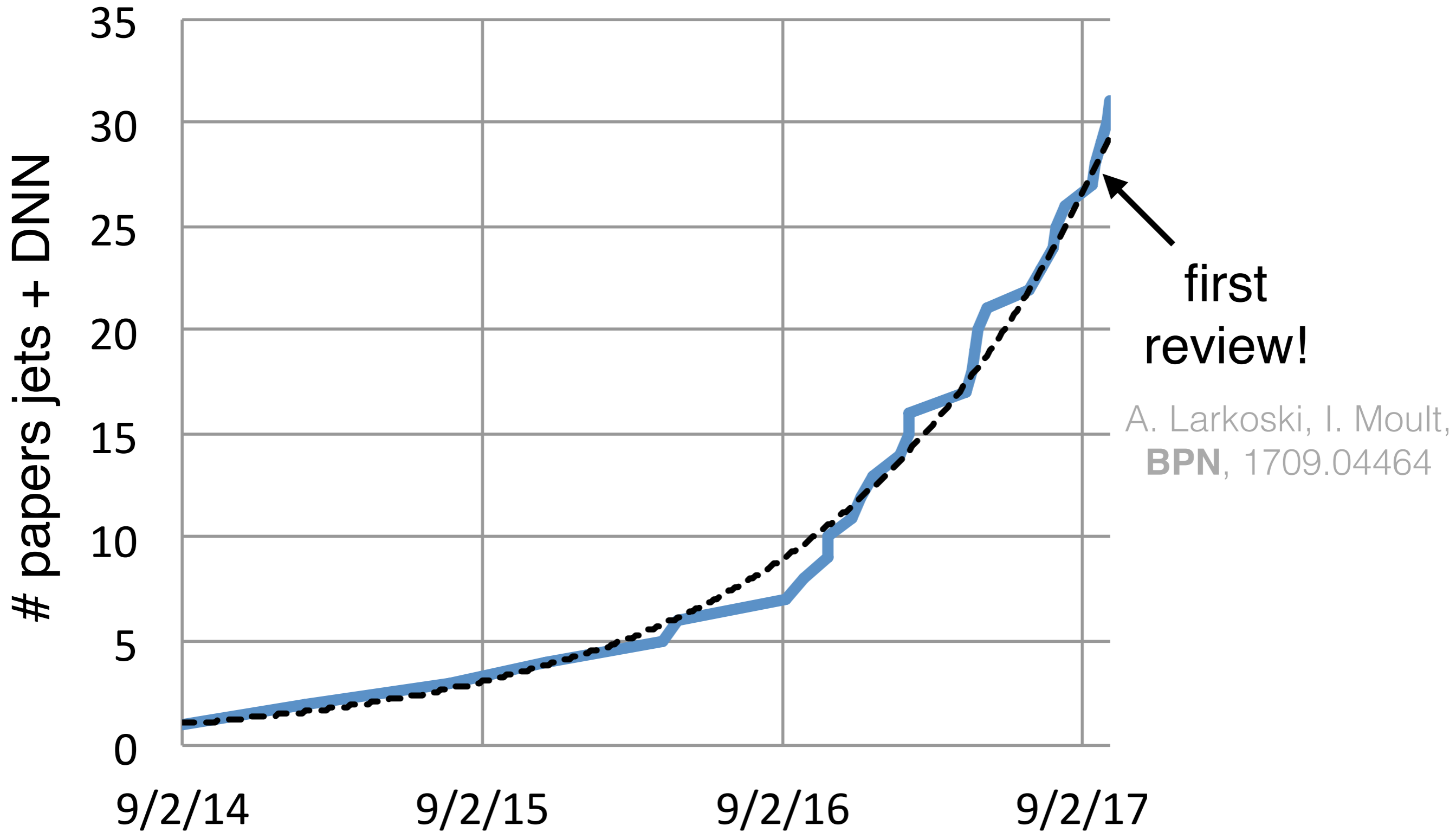


However, his face's **image mass** is quite different!

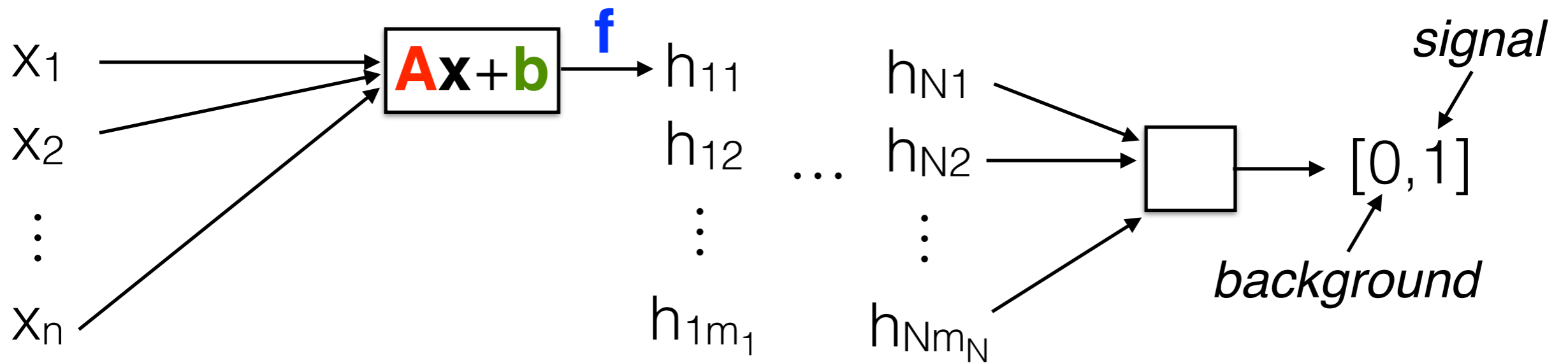
Now, with a carefully processed image, we can ask: where did this jet come from?



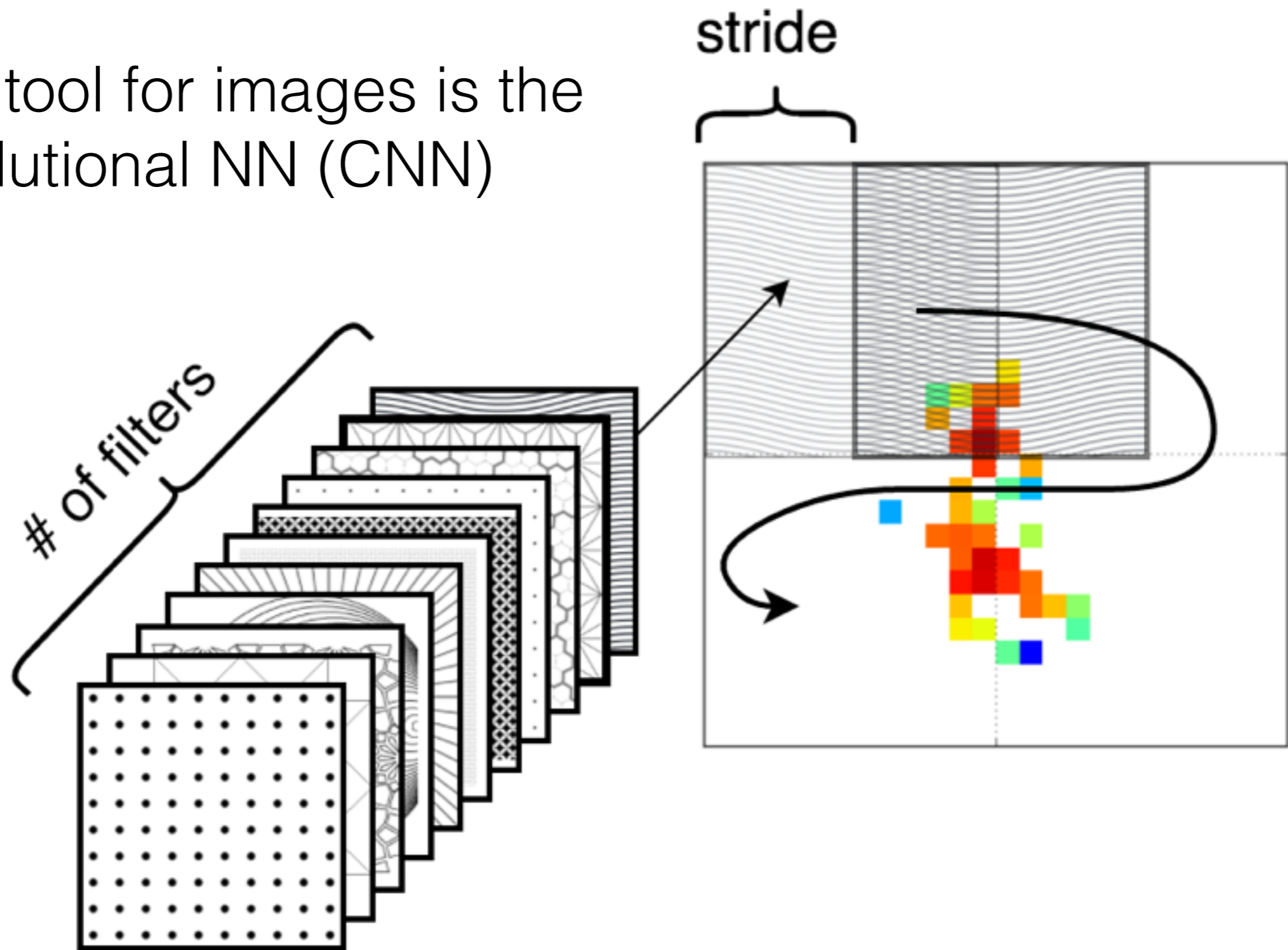
ultimate classification is achieved with modern machine learning using **all pixels as input!**



Typical 'fully connected' network:

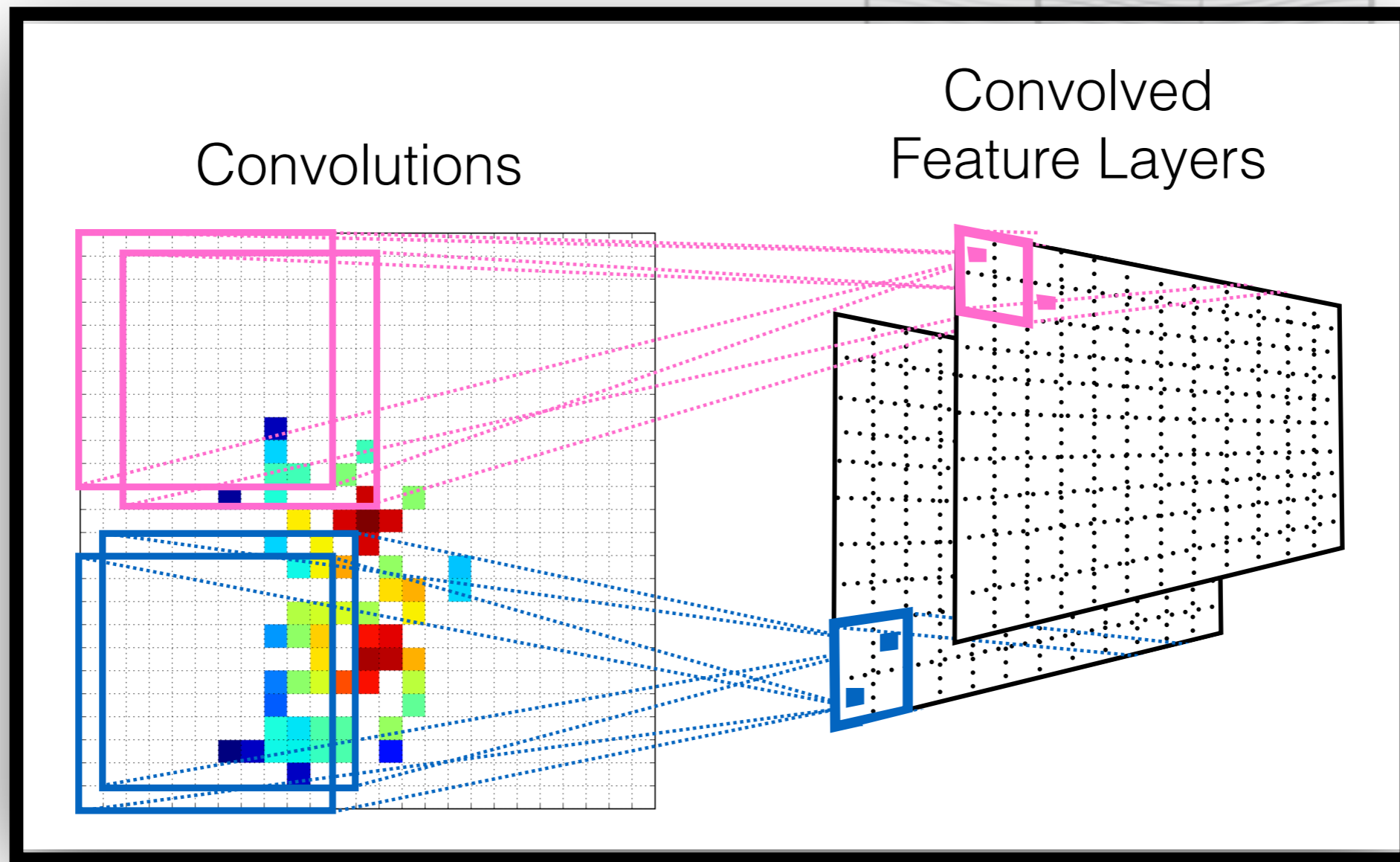


Common tool for images is the convolutional NN (CNN)

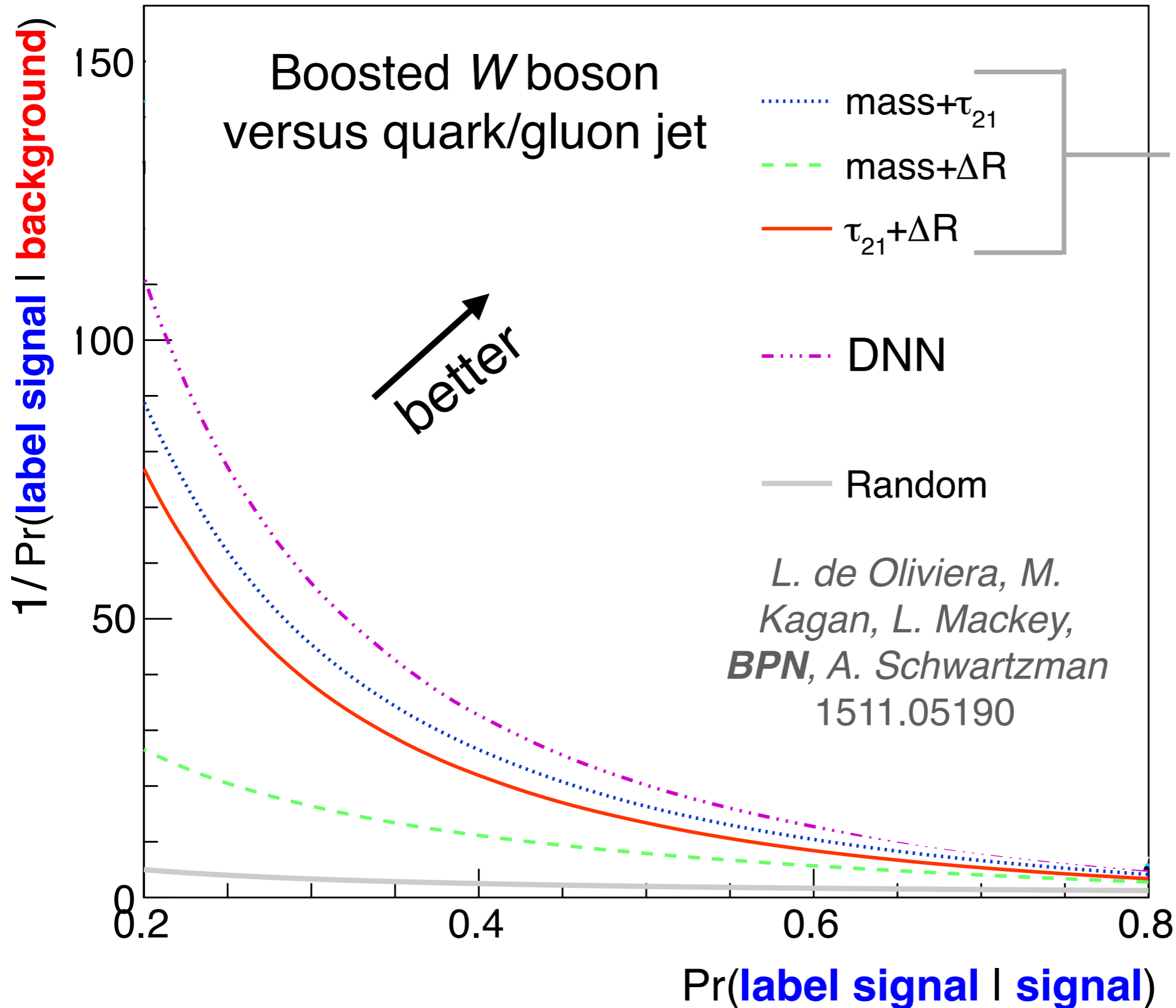


The filter is like the **A**, only the dimensionality is now the filter size ($\ll n$) and not the image size (n).

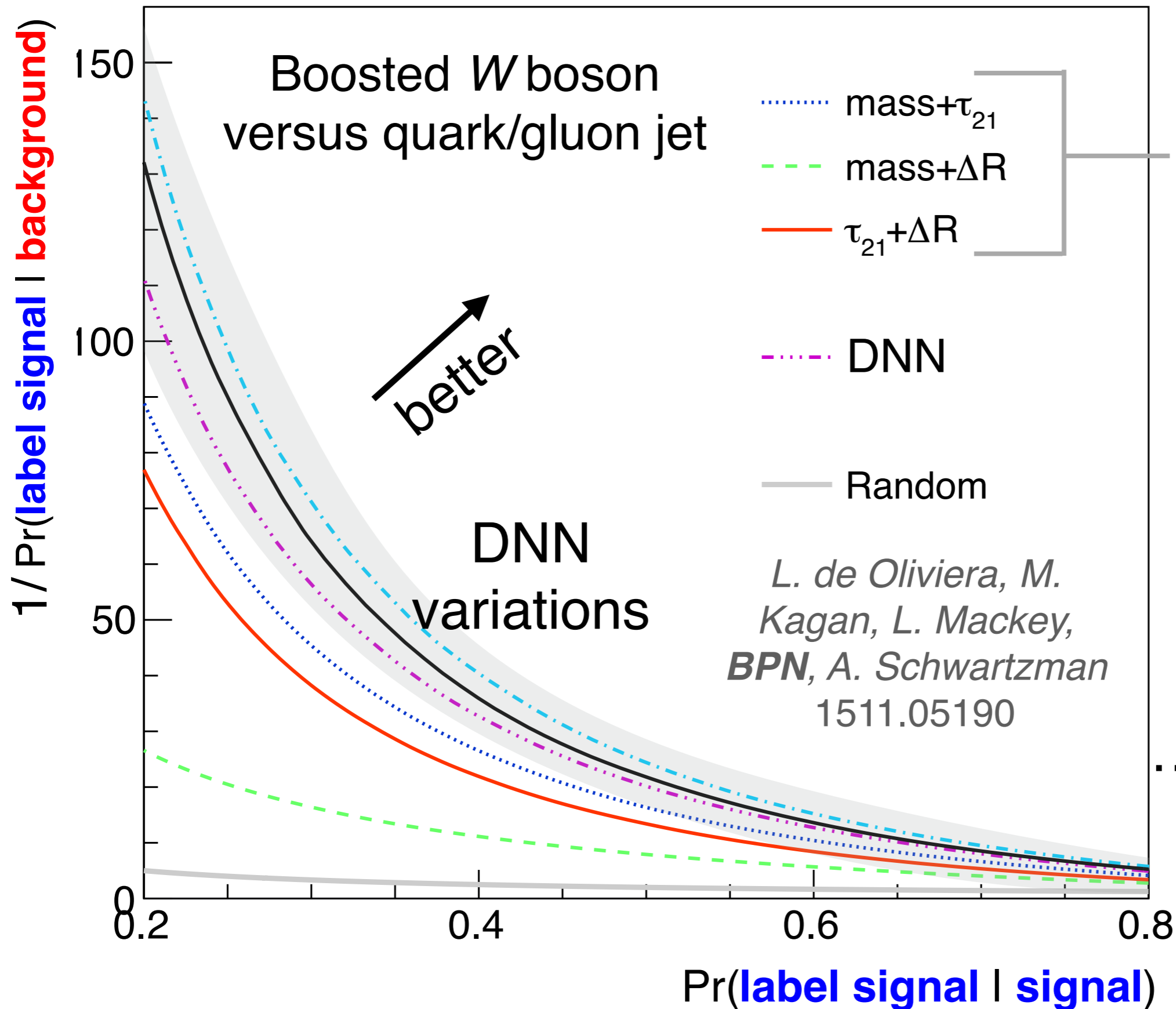
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mass, τ_{21} , ΔR are all simple functions of the image

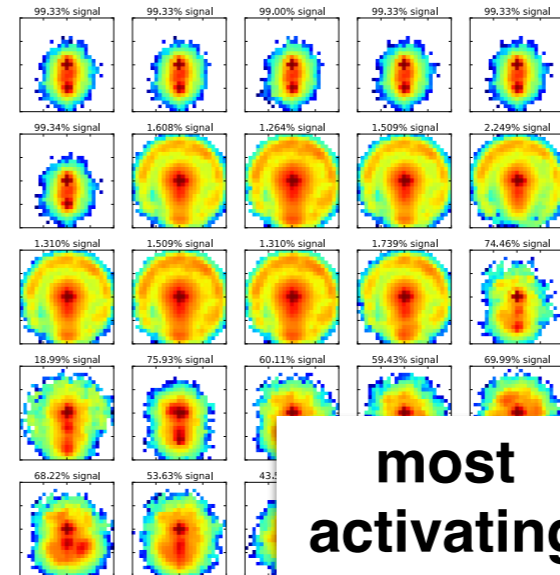
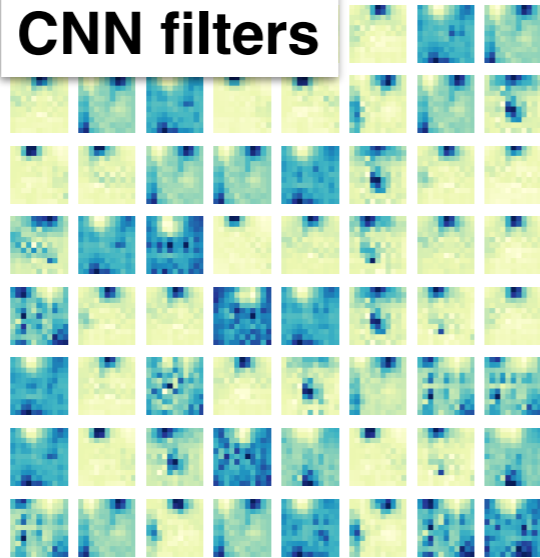
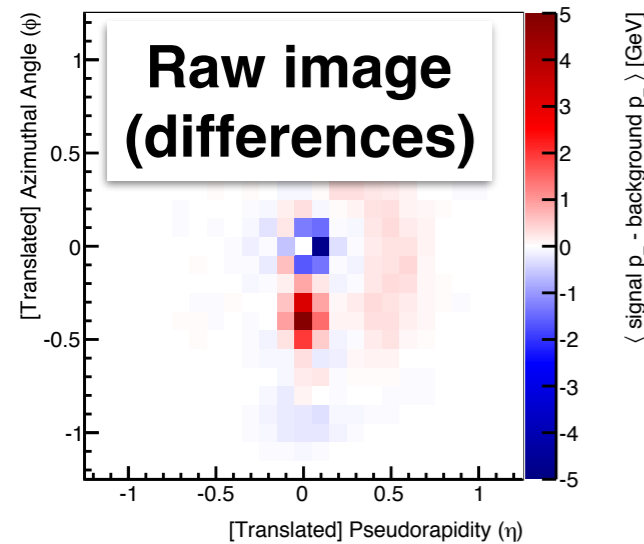


mass, τ_{21} , ΔR are all simple functions of the image

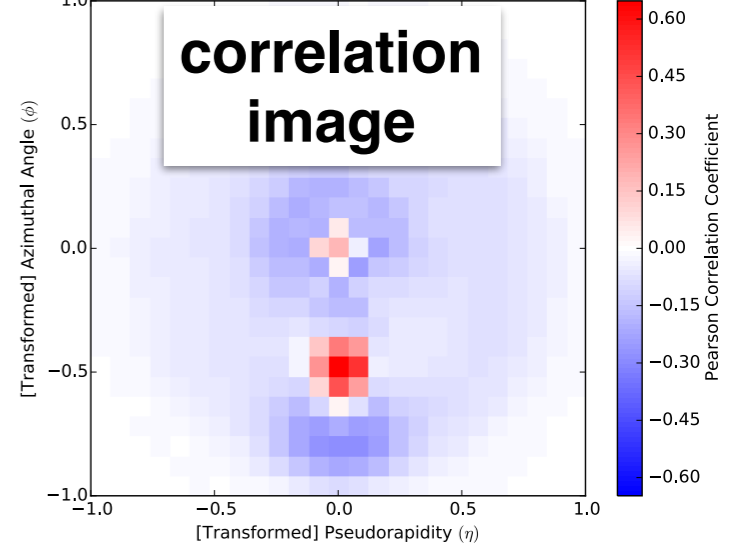
...what the DNN is learning is active R&D!

Opening the **box** is critical for improving robustness

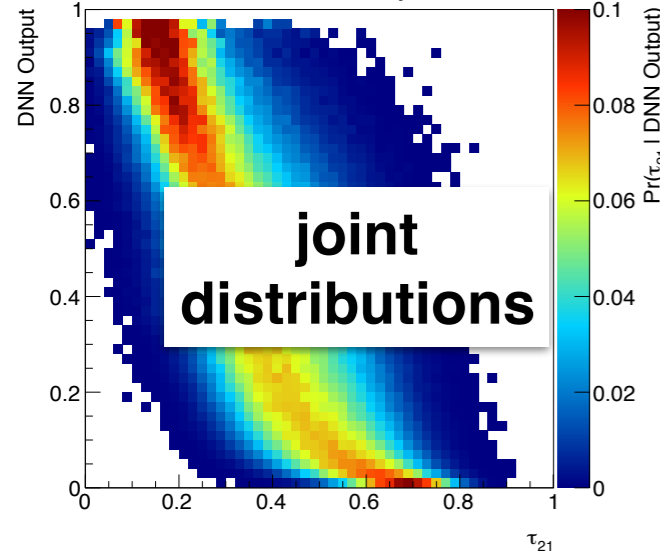
$250 < p_T/\text{GeV} < 260 \text{ GeV}, 0.59 < \tau_{21} < 0.61, 79 < \text{mass}/\text{GeV} < 81$
 $\sqrt{s} = 13 \text{ TeV}, \text{Pythia 8}$



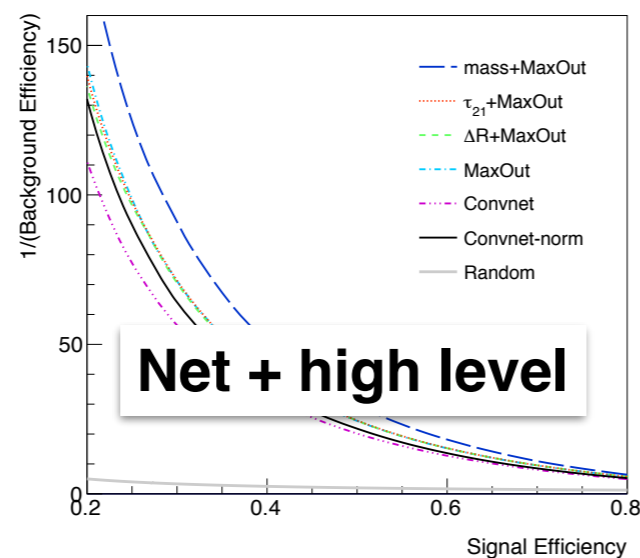
Correlation of Deep Network output with pixel activations.
 $p_T^W \in [250, 300]$ matched to QCD, $m_W \in [65, 95]$ GeV



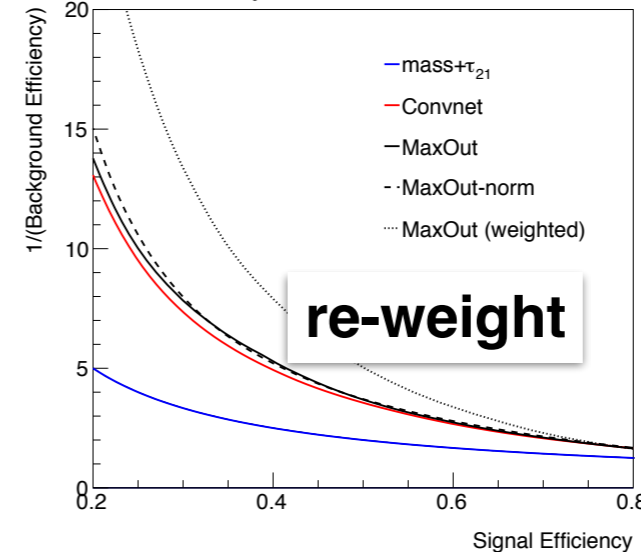
$250 < p_T/\text{GeV} < 300 \text{ GeV}, 65 < \text{mass}/\text{GeV} < 95$
 QCD, $\sqrt{s} = 13 \text{ TeV}, \text{Pythia 8}$



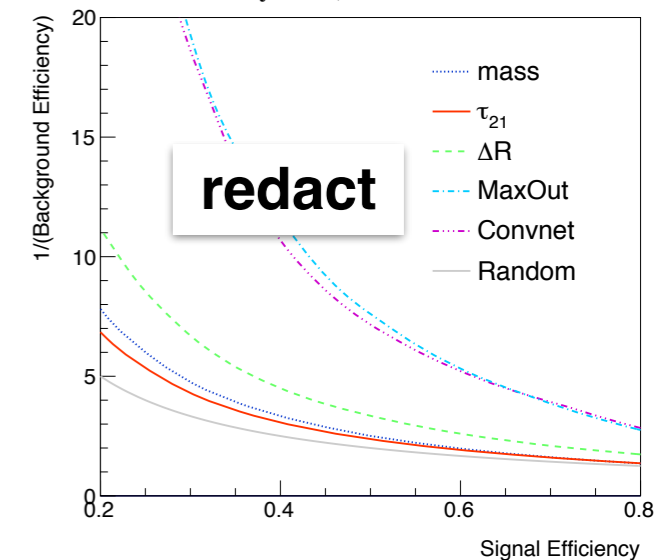
$240 < p_T/\text{GeV} < 260 \text{ GeV}, 0.19 < \tau_{21} < 0.21, 79 < \text{mass}/\text{GeV} < 81$
 $\sqrt{s} = 13 \text{ TeV}, \text{Pythia 8}$



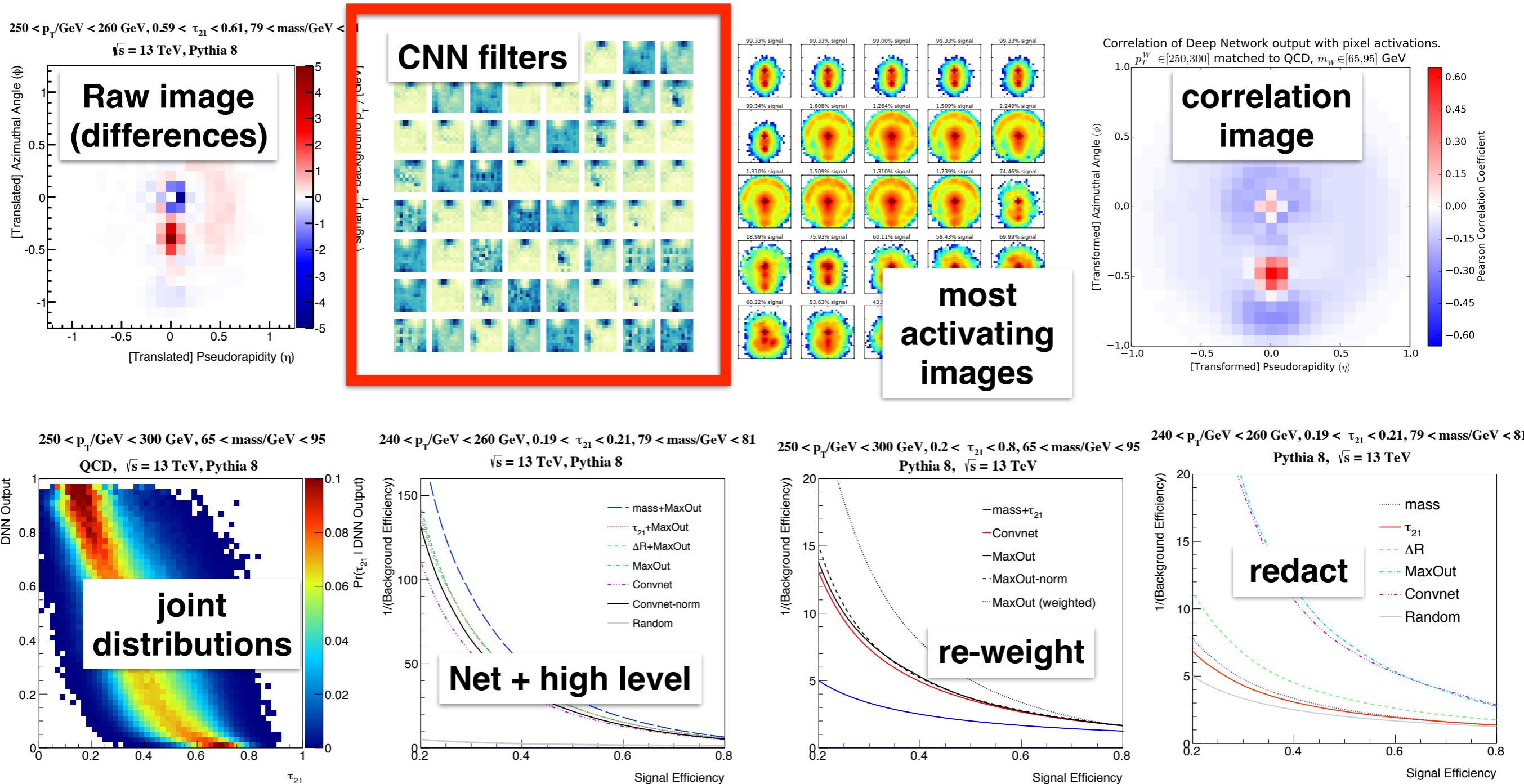
$250 < p_T/\text{GeV} < 300 \text{ GeV}, 0.2 < \tau_{21} < 0.8, 65 < \text{mass}/\text{GeV} < 95$
 Pythia 8, $\sqrt{s} = 13 \text{ TeV}$



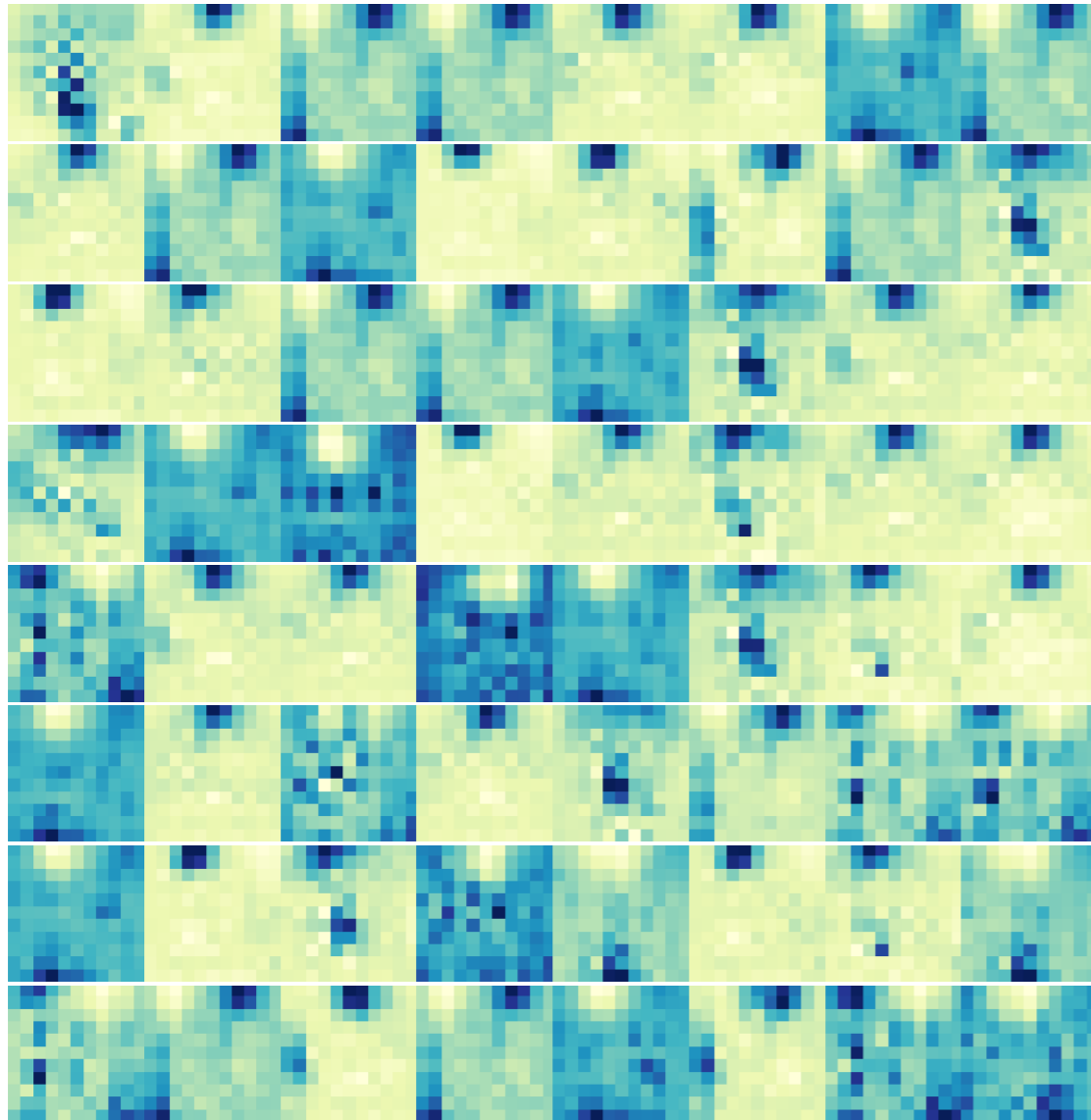
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 Pythia 8, $\sqrt{s} = 13 \text{ TeV}$



Opening the **box** is critical for improving robustness

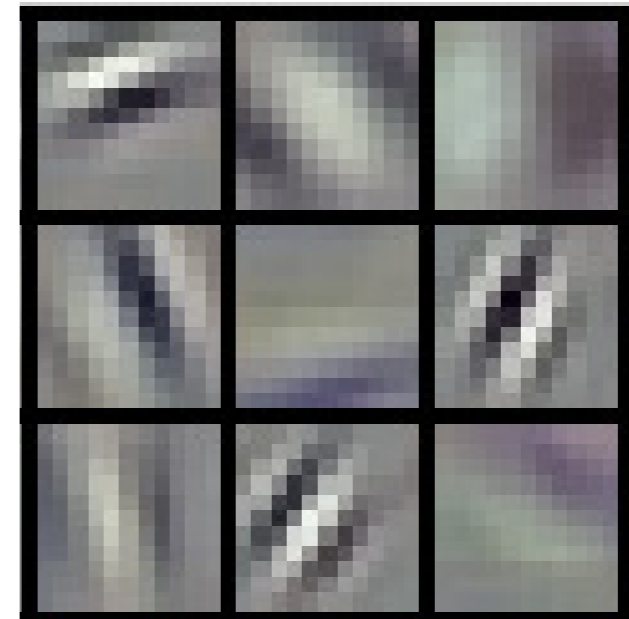


Filters are images! Can visualize ‘higher-level features’ learned by the network



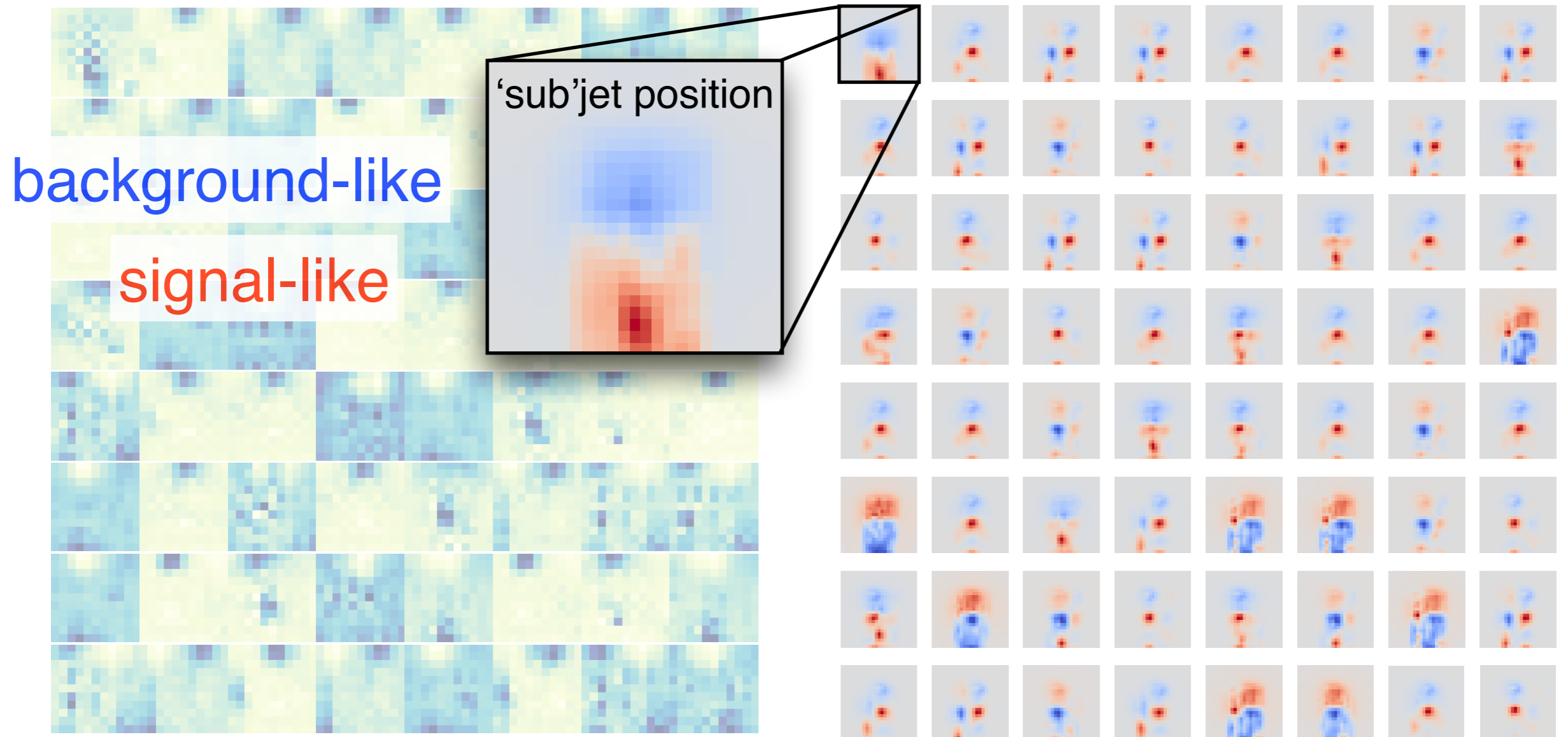
Jet Images

learned edge detection



“Natural” Images

Filters are images! Can visualize 'higher-level features' learned by the network



Jet Images Layer 1 Filters

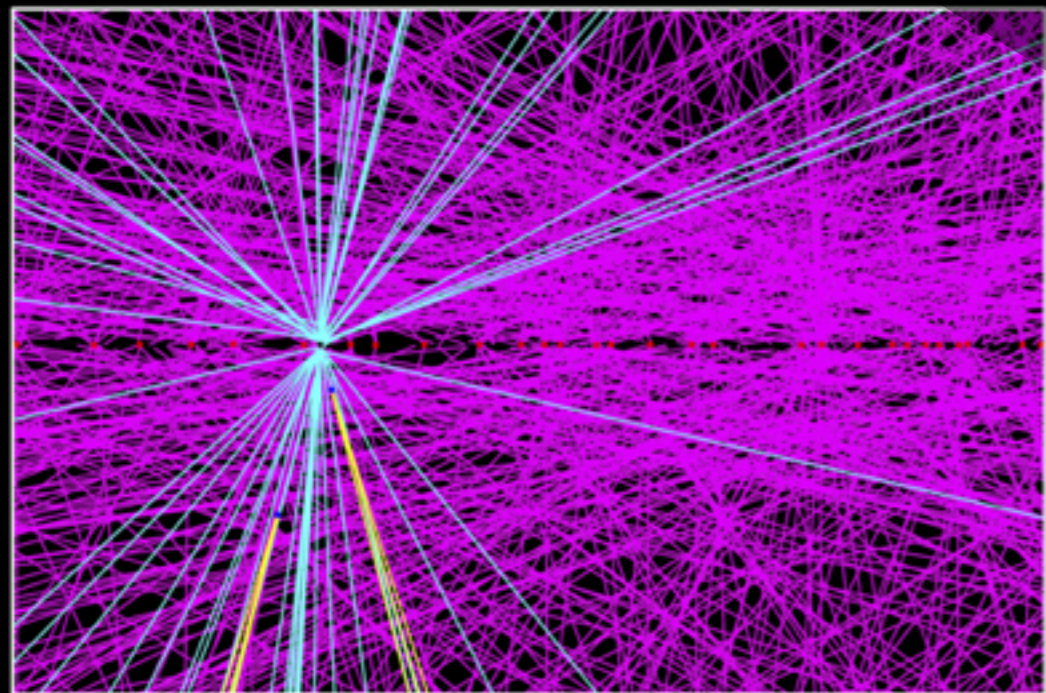
Filters convolved with
signal - background

Beyond Classification I: Removing Noise

pp collisions at the LHC
don't happen one at a time!

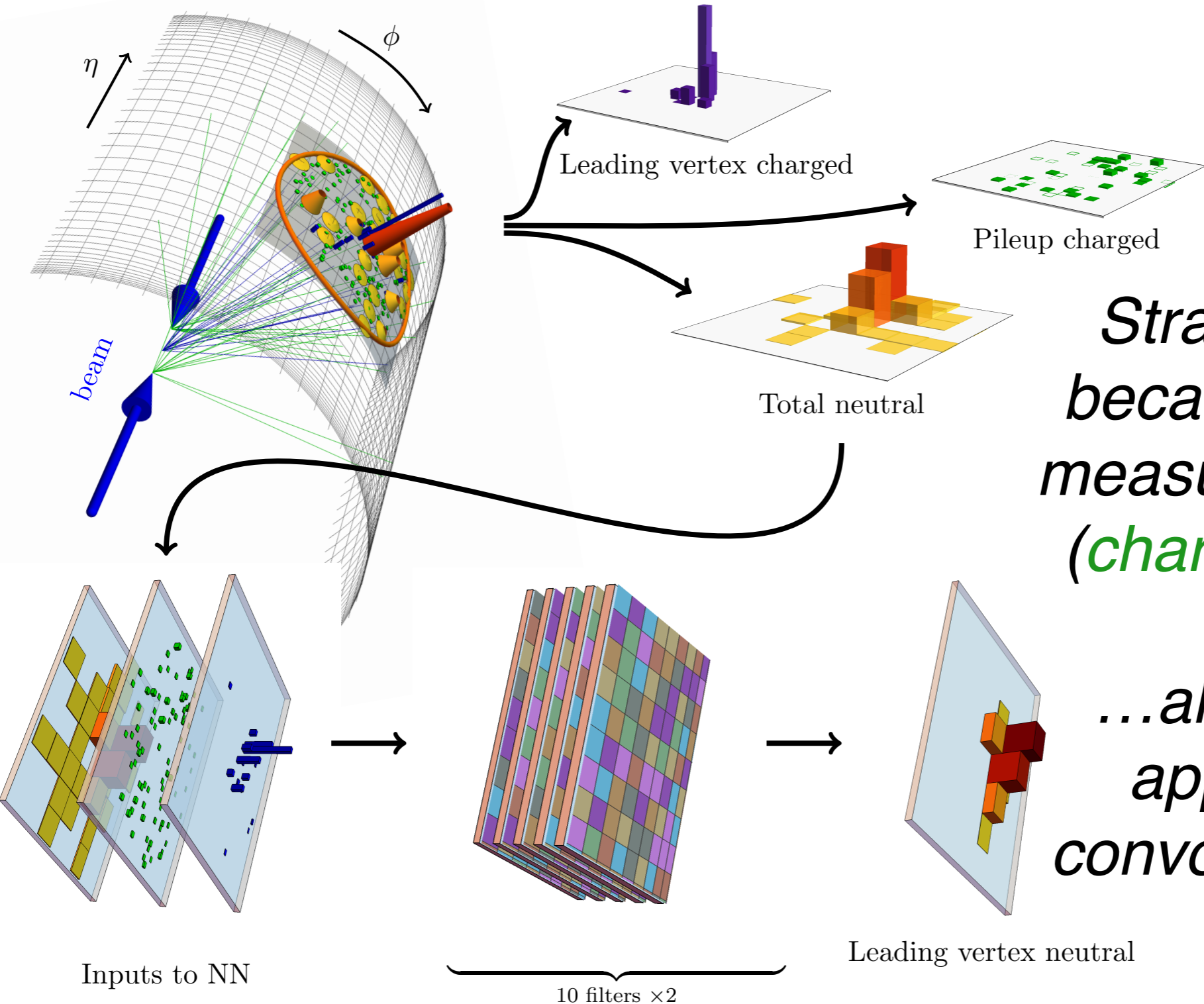


the extra collisions are called **pileup**
and add soft radiation on top of our jets



this is akin to image
de-noising - we can
use ML for that!

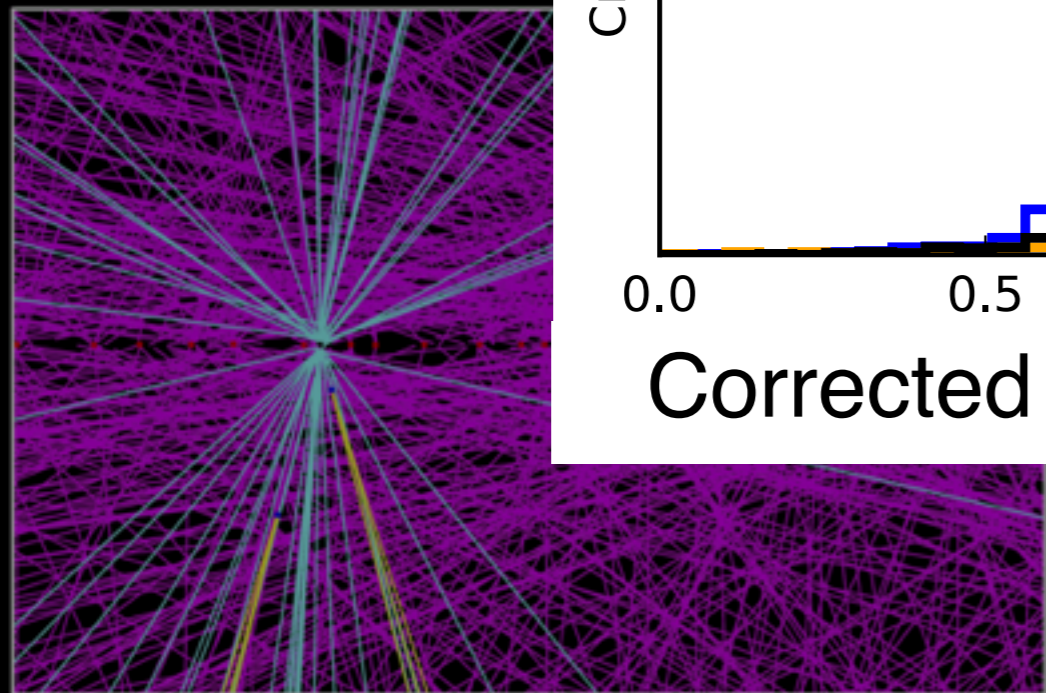
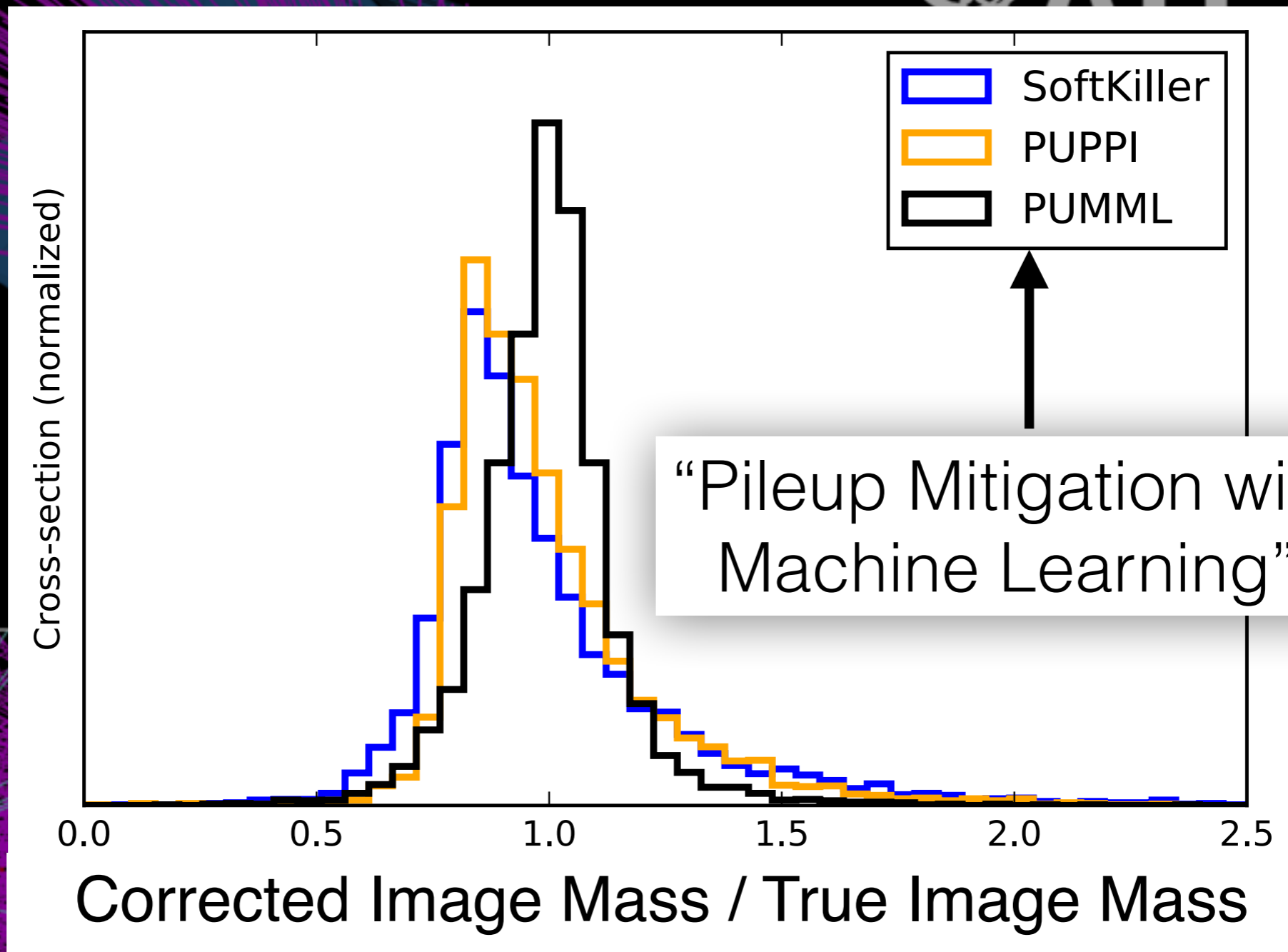
Beyond Classification I: Removing Noise



Strange noise because we can measure ~2/3 of it (charged pileup)

...also a natural application of convolutional NNs!

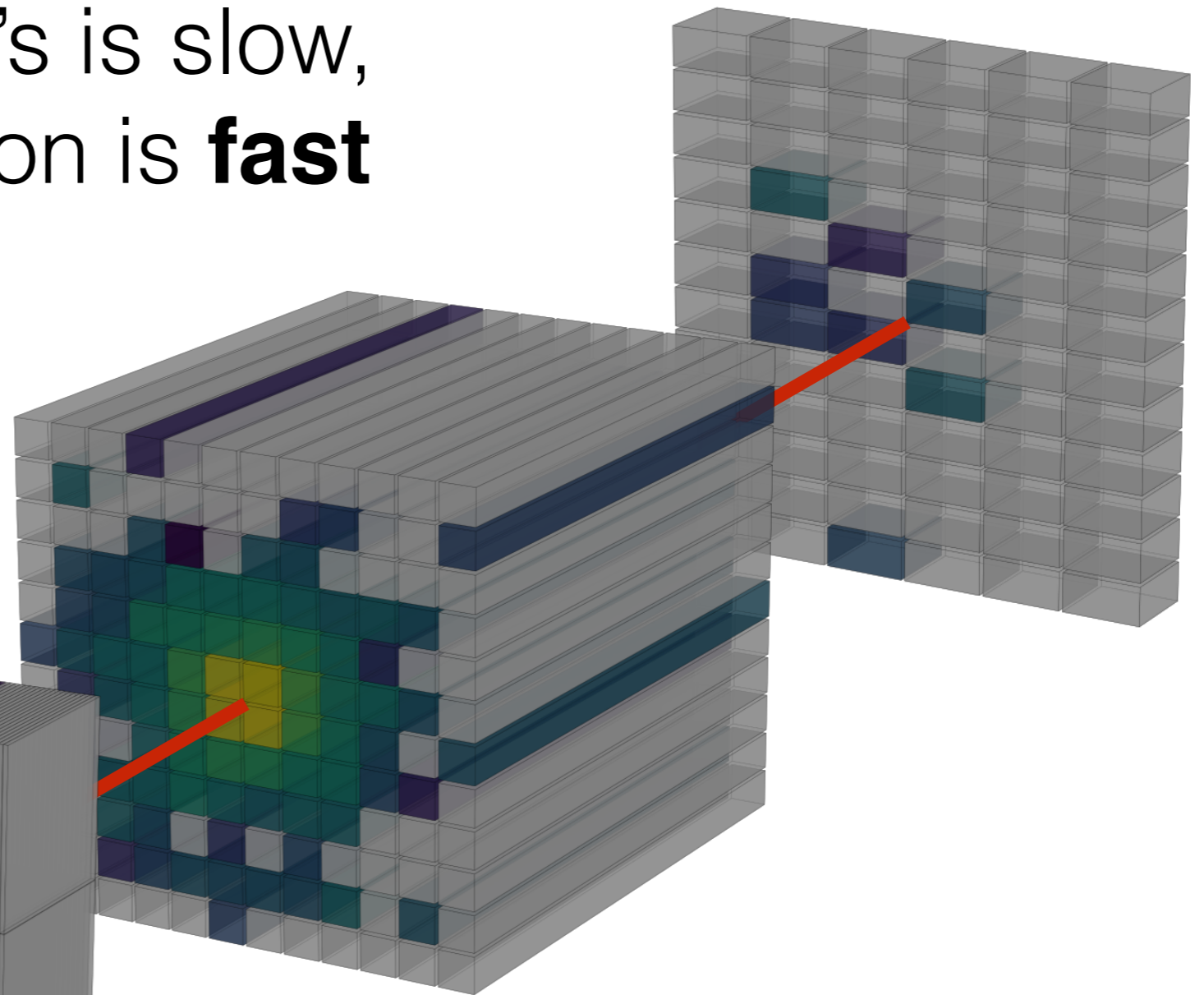
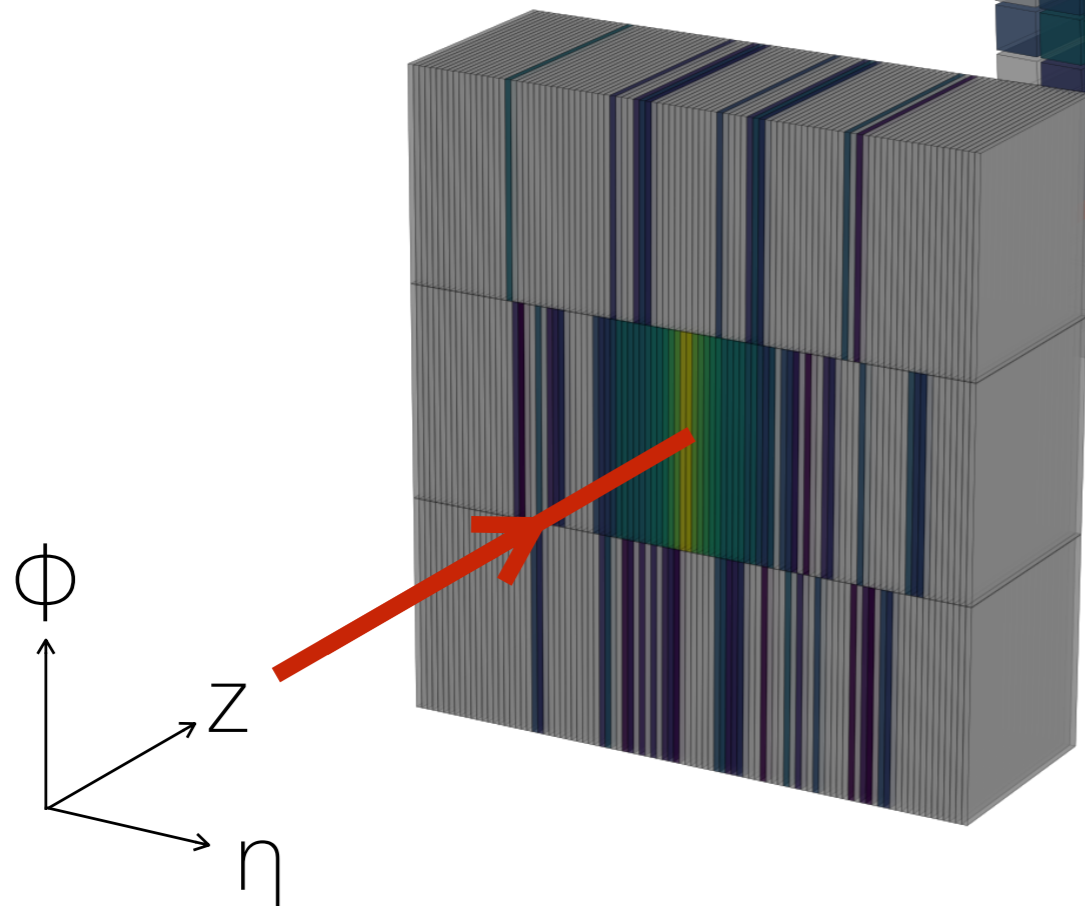
Beyond Classification I: Removing Noise



Beyond Classification II: Simulation NN

Training NN's is slow,
but evaluation is **fast**

Physics-based
simulations of
jets are **slow**



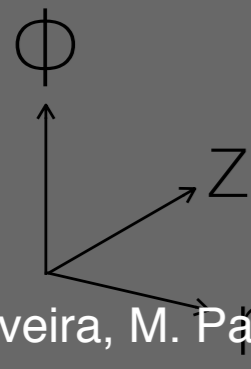
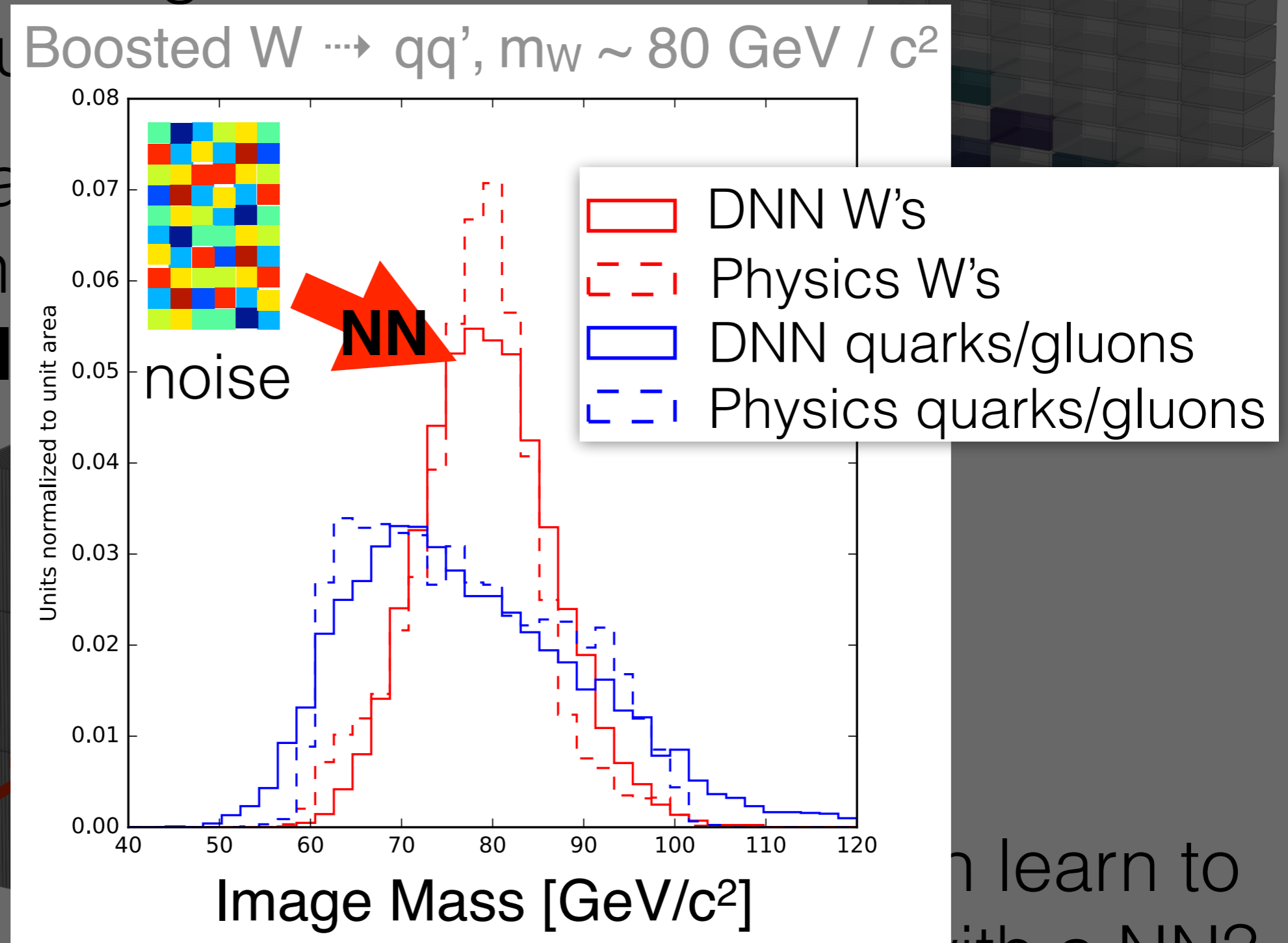
What if we can learn to
simulate jets with a NN?

Beyond Classification II: Simulation NN

Training NN's is slow,

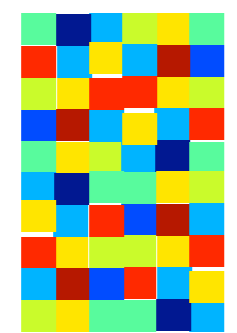
but Boosted $W \rightarrow qq'$, $m_W \sim 80 \text{ GeV} / c^2$

Physics-based simulation jets are slow

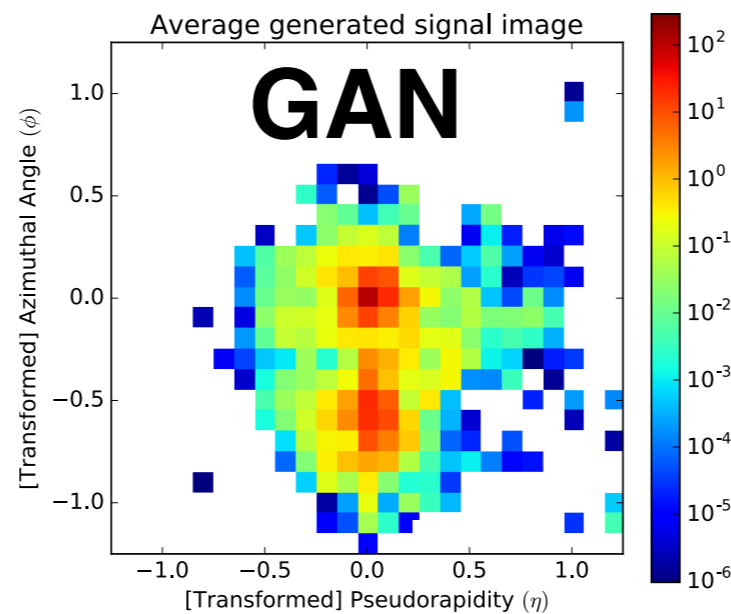


Can we learn to simulate jets with a NN?

Generative Adversarial Networks (GAN):
*A two-network game where one **maps noise to images** and one **classifies images as fake or real**.*

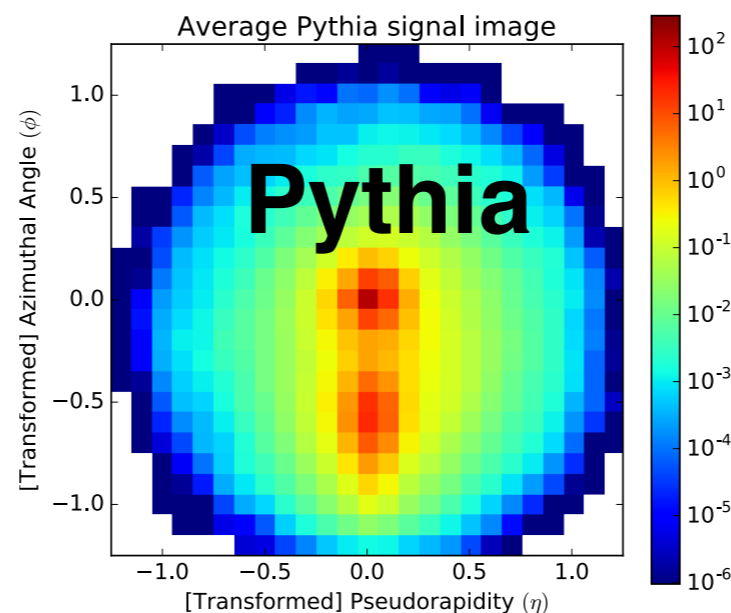


noise



{real, fake}

When **D** is maximally confused, **G** will be a good generator

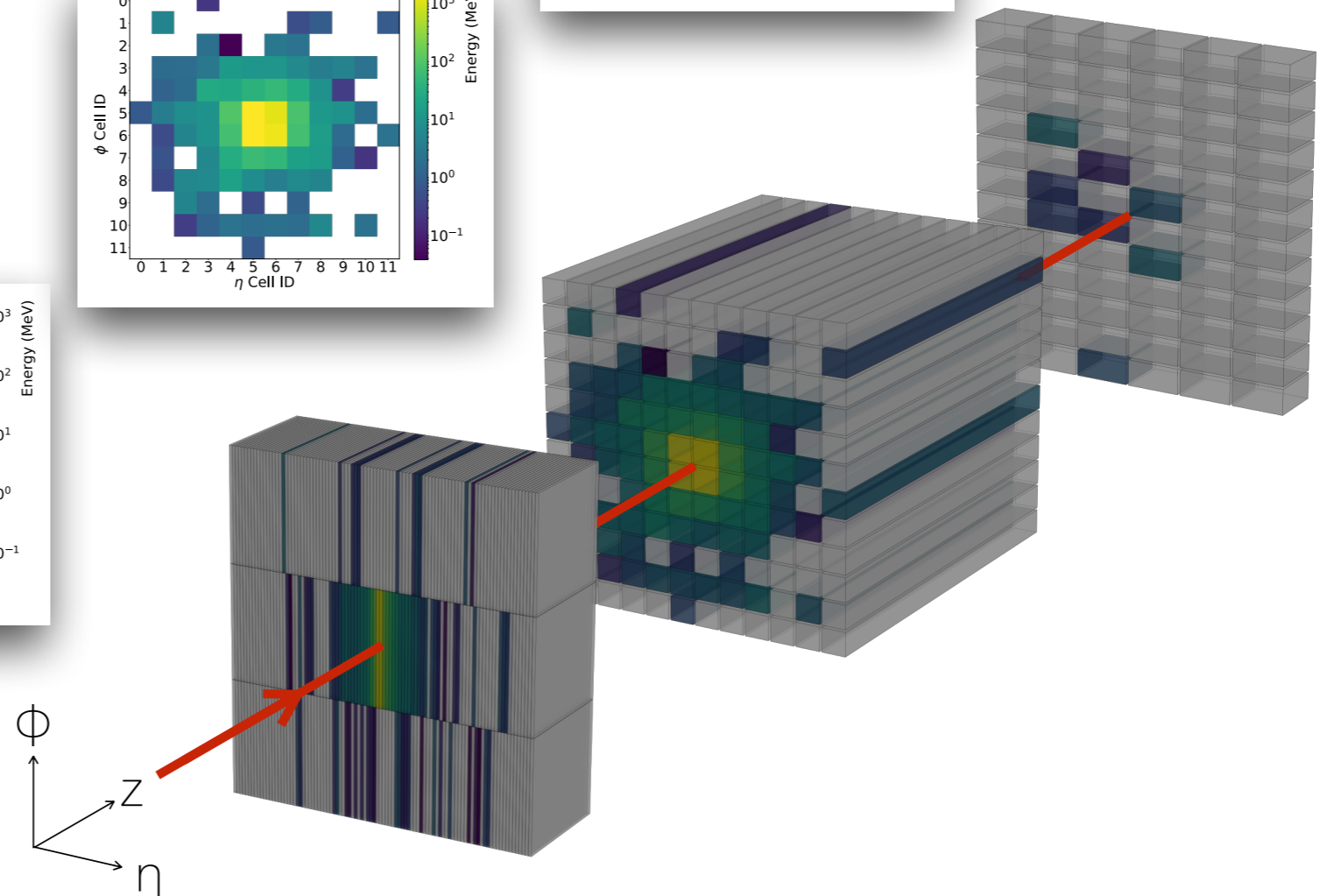
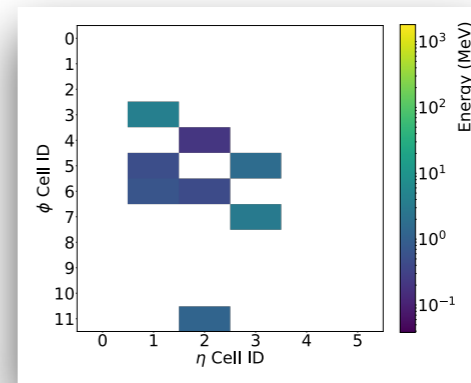
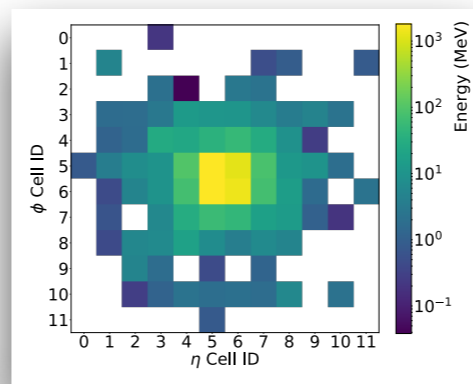
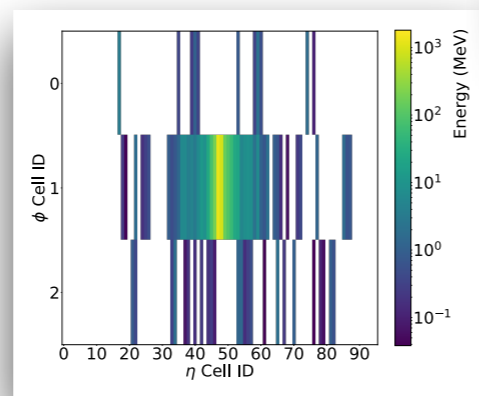


Physics-based simulator

+ More Layers for Generation

What about **multiple layers** with **non-uniform granularity** and a **causal relationship**?

Not jet images per se,
but the technology is
more general than jets!

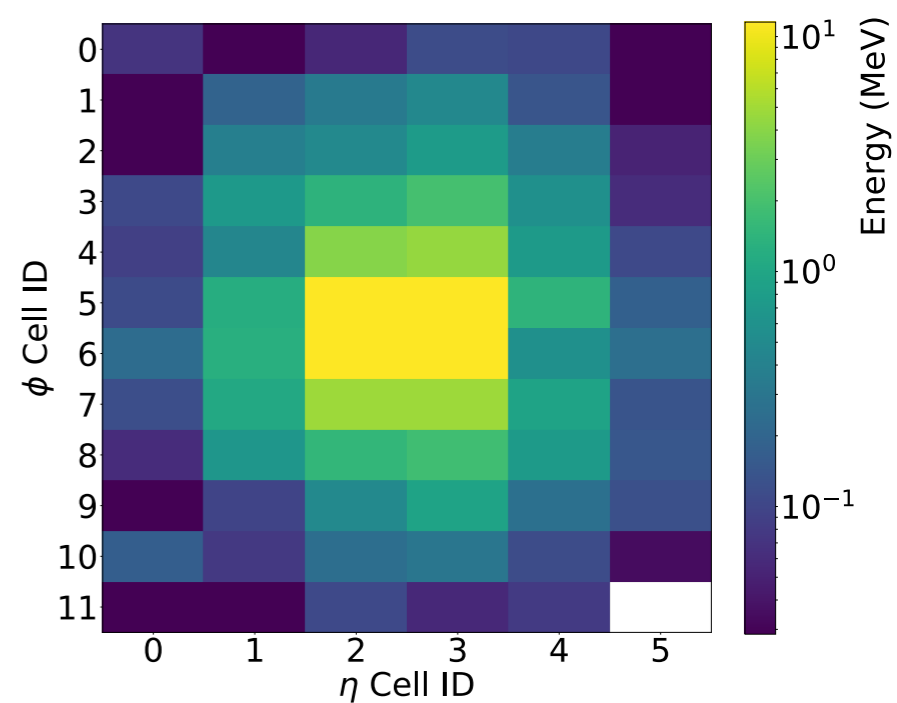
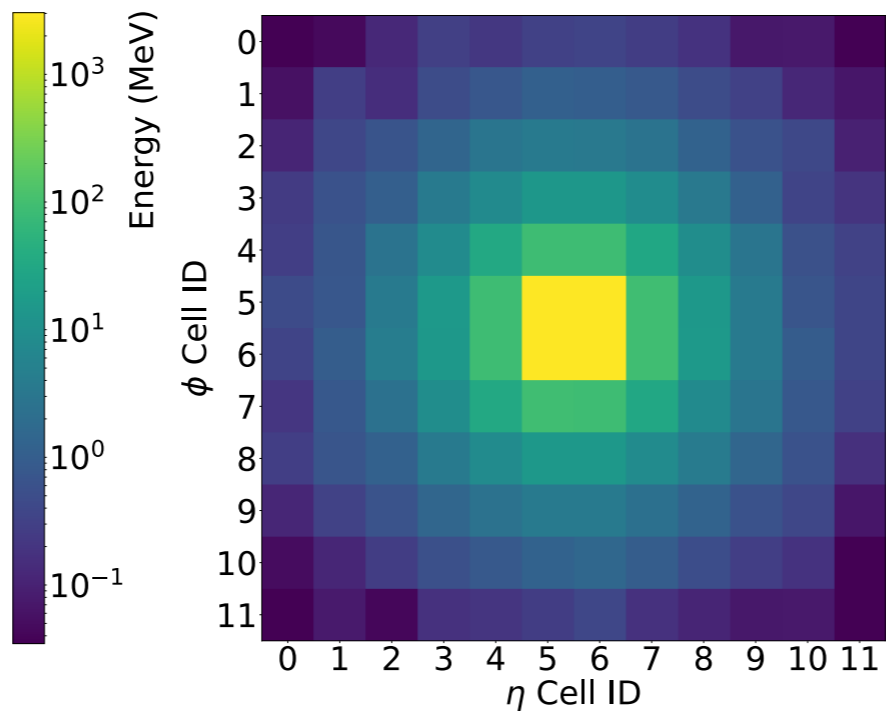
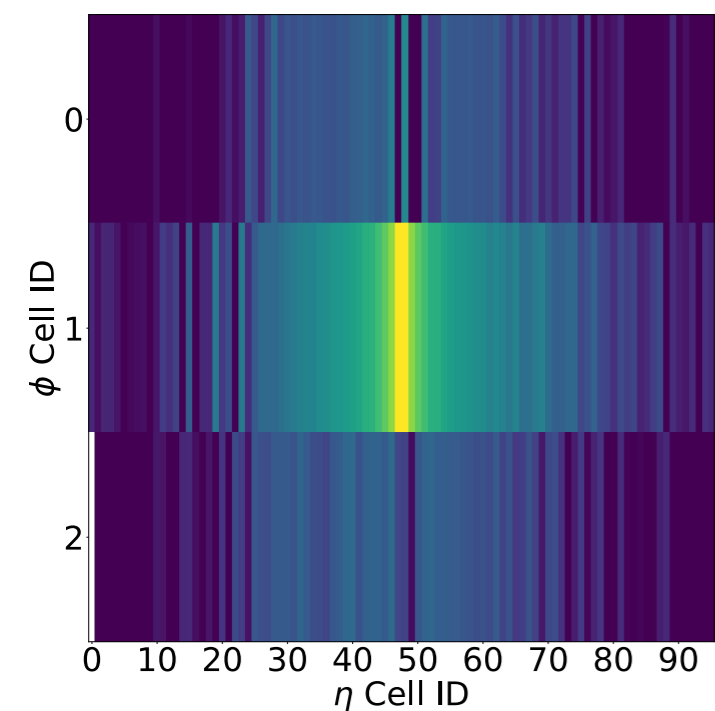
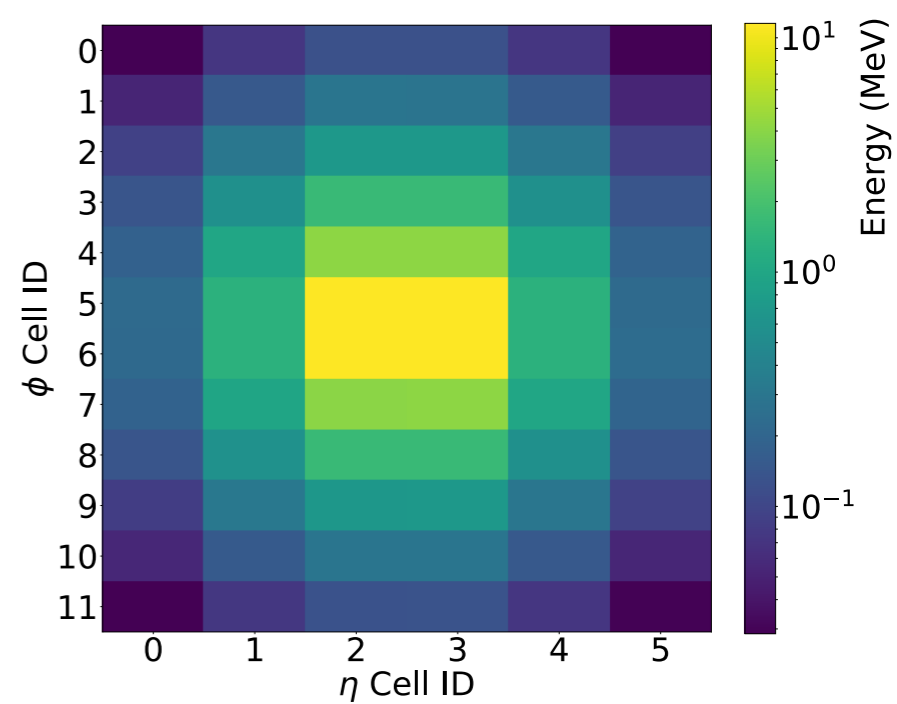
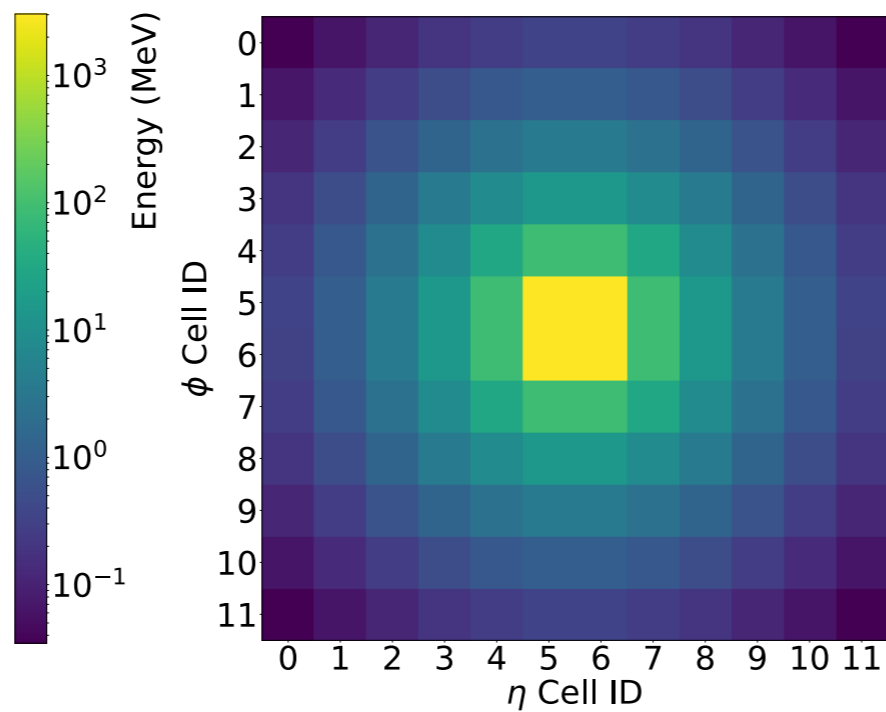
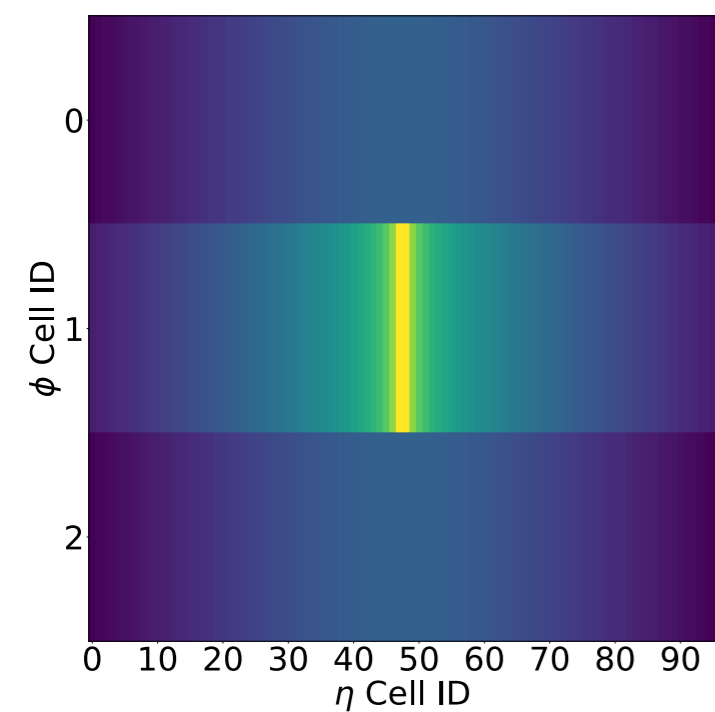


M. Paganini, L. de Oliveira,
and **BPN** 1705.02355

Average Images

Geant4

*M. Paganini, L. de Oliveira, and **BPN** 1705.02355*



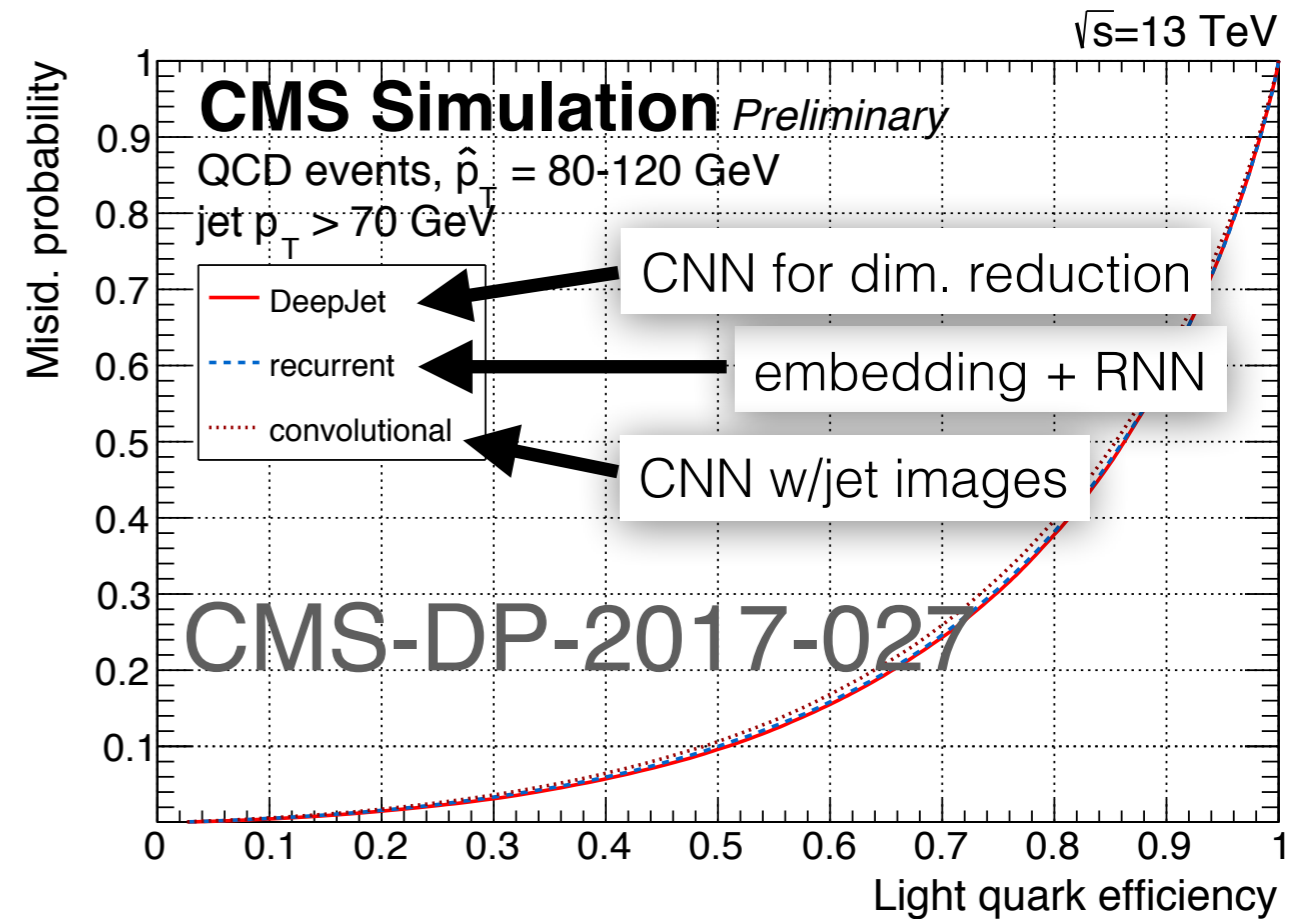
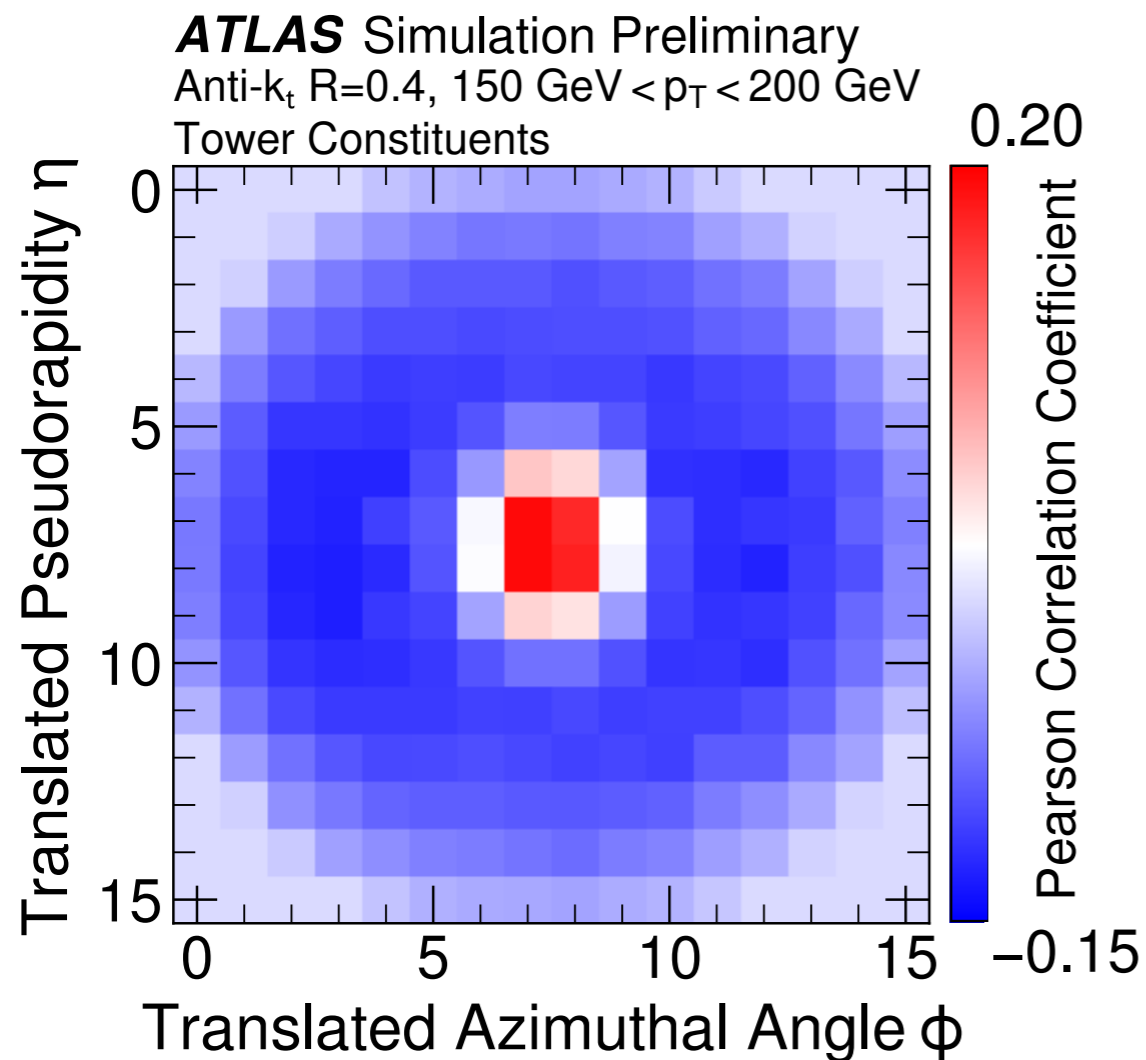
CaloGAN

*M. Paganini, L. de Oliveira, and **BPN** 1705.02355*

| Generation Method | Hardware | Batch Size | milliseconds/shower |
|-------------------|----------|------------|---------------------|
| GEANT4 | CPU | N/A | 1772 ← |
| CALOGAN | CPU | 1 | 13.1 |
| | | 10 | 5.11 |
| | | 128 | 2.19 |
| | | 1024 | 2.03 |
| | GPU | 1 | 14.5 |
| | | 4 | 3.68 |
| | | 128 | 0.021 |
| | | 512 | 0.014 |
| | | 1024 | 0.012 ← |

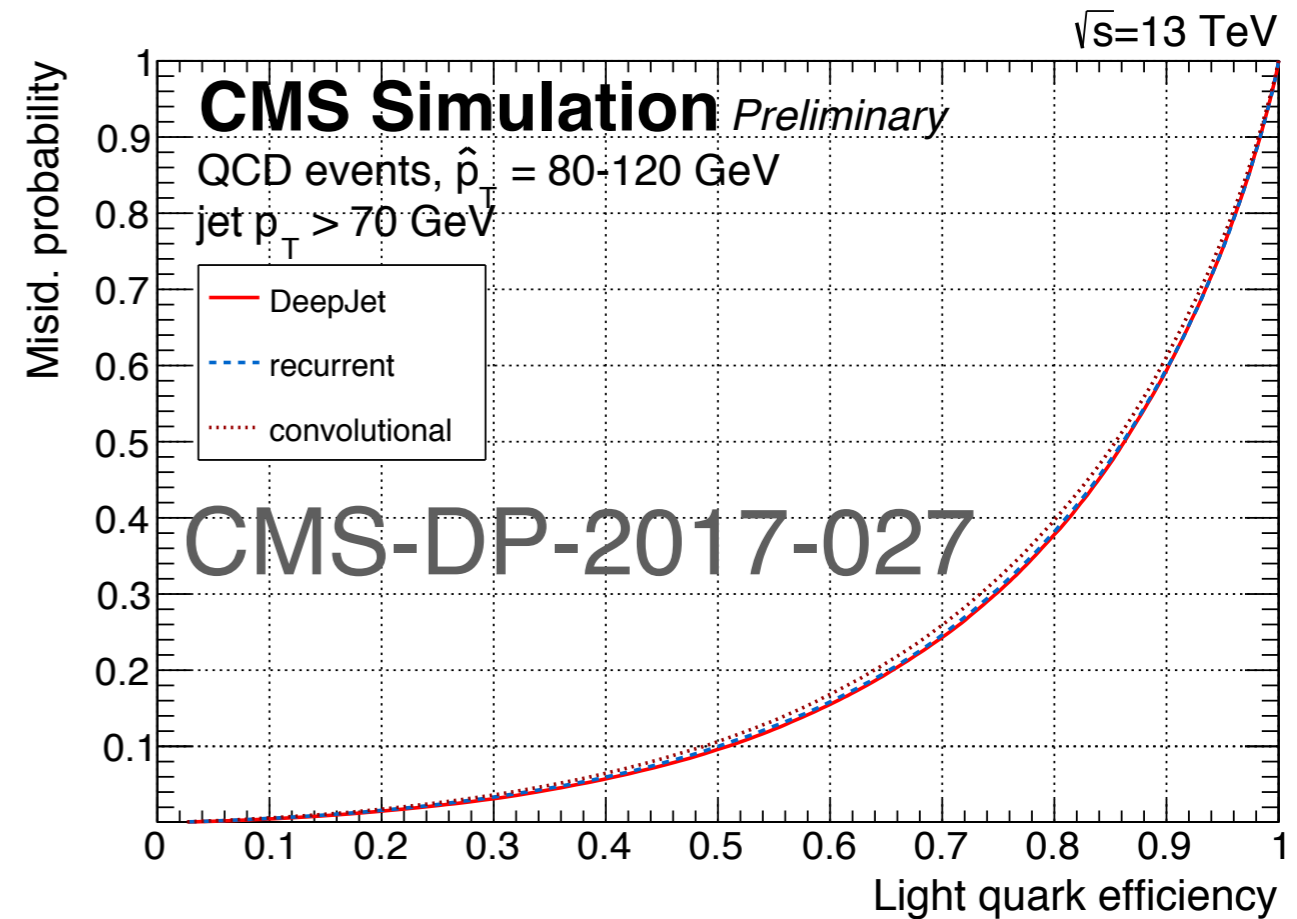
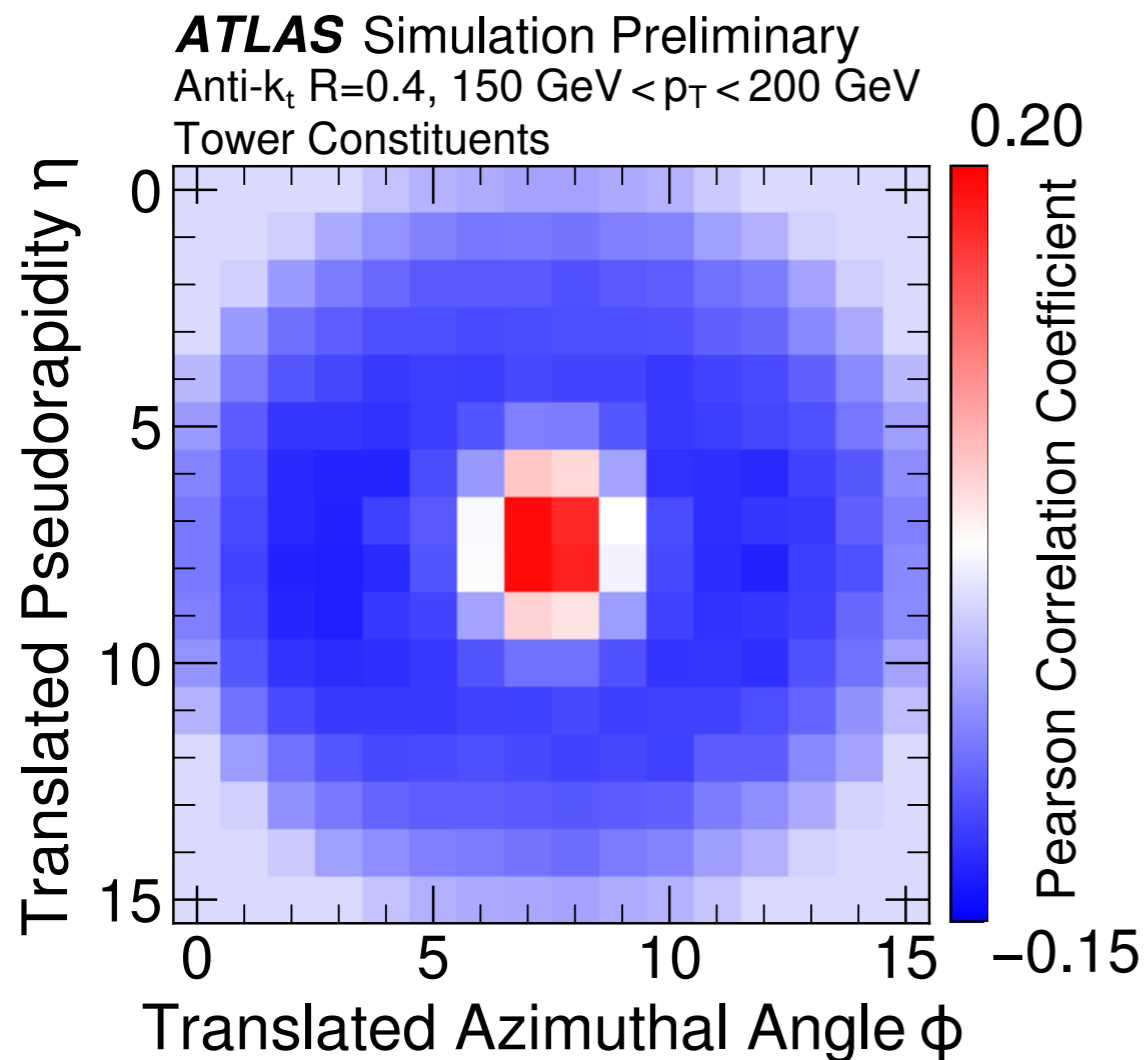
R&D cycle is fast; sometimes integration can be slow

There may be multiple ways to get to the same solution



CMS, ATLAS, LHCb,
MicroBooNE, NOvA, DUNE, etc.
are increasing their use of
DNNs for many applications

DNN classification, regression, and generation are powerful tools to fully exploit the physics program at the LHC & beyond



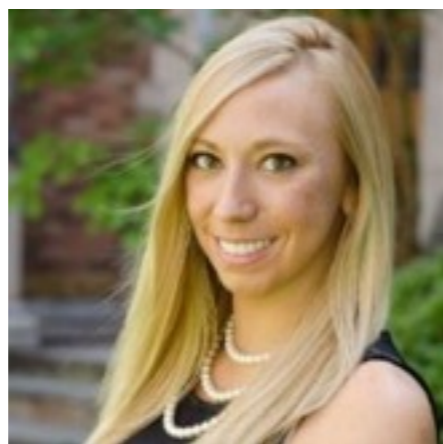
The key to robustness is to study what is being learned; this may even help us to learn something new!

Collaborators



Lucio
Dery

Stanford



Michela
Paganini

Yale



Eric
Metodiev

MIT



Patrick
Komiske

MIT



Zihao
Jiang

Stanford



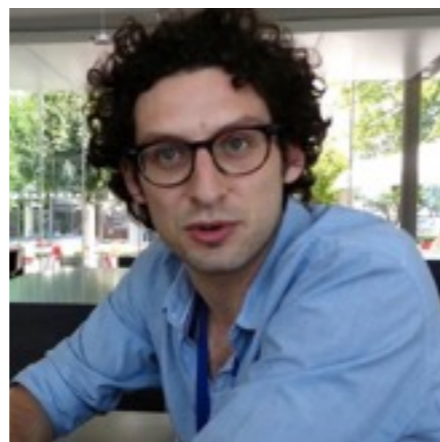
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Luke
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Kagan

SLAC



Jesse
Thaler

MIT



Matt
Schwartz

Harvard



Ariel
Schwartzman

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