

Beyond the Standard Model: Where do we go from here?



**Flavour Overview**  
or  
**The (Excessive) Success of the Standard Model**

Gaia Lanfranchi – LNF-INFN

John Hopkins Workshop – GGI, Florence – October 2018

## The Common Lore

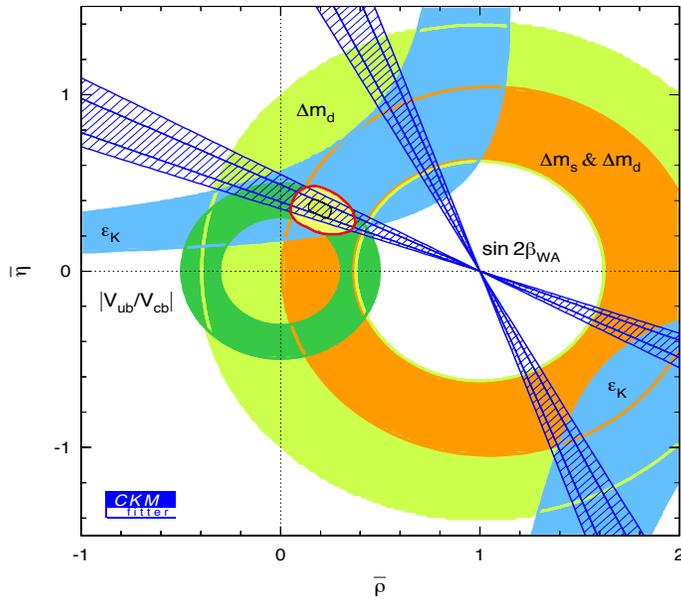
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The SM is considered the low-energy effective theory of a more complete theory which holds at a higher scale. New particles are expected to show up at the EW scale to solve the hierarchy problem and have sizeable couplings to SM particles, so they can be detected either via direct detection or via their indirect effects in loops.

Flavor physics = test of the CKM paradigm

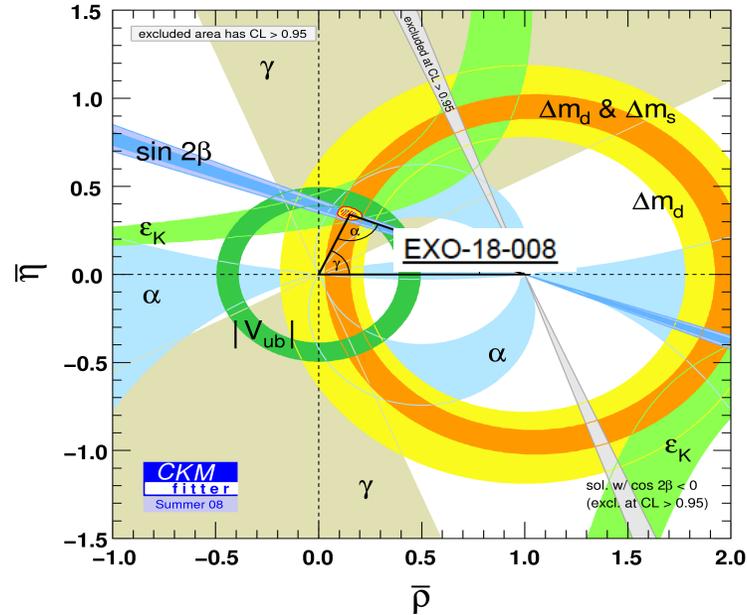
# The impressive (excessive) success of the SM

EPS 2001



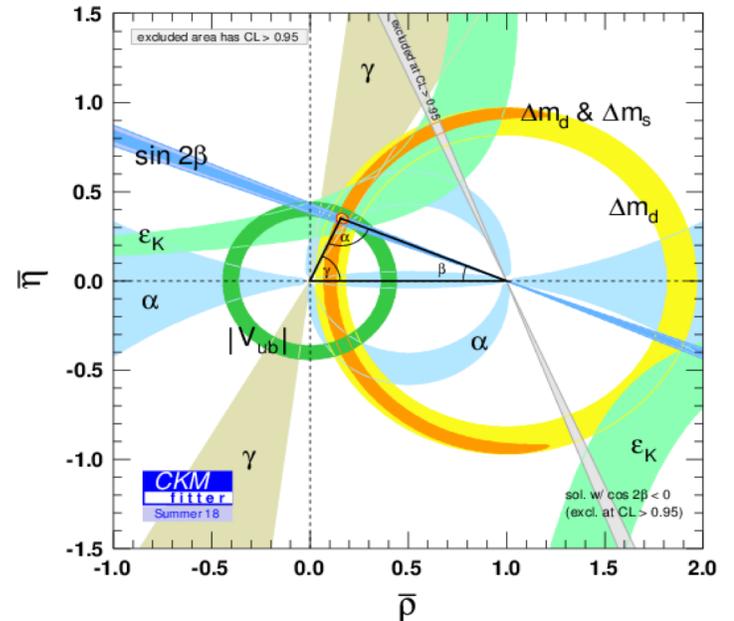
~ Pre B factories

ICHEP 2008



Pre-LHC

ICHEP 2018



LHC and B factories 2018

Dramatic progress in the last ~20 years.

# Flavor data sets

Experiment	$\int L dt$	$\sigma(bb)$	$\sigma(cc)$	$\sigma(ss)$	operation
BaBar	530 fb <sup>-1</sup>	1.1 nb	1.6 nb	0.4 nb	1999-2008
Belle	1040 fb <sup>-1</sup>	1.1 nb	1.6 nb	0.4 nb	1999-2010
BESIII	16 fb <sup>-1</sup>	-	6 nb (3770 MeV)	-	2008 -
KLOE2	5.5 fb <sup>-1</sup>	-	-	$\sim 3 \mu\text{b}$ (1020 MeV)	2014-2018
ATLAS	$> 100 \text{fb}^{-1}$	250-500 $\mu\text{b}$	-	-	2009 -
CMS	$> 100 \text{fb}^{-1}$	250-500 $\mu\text{b}$	-	-	2009 -
LHCb	1 + 2 + $> 5 \text{fb}^{-1}$	250 - 500 $\mu\text{b}$	1200 - 2400 $\mu\text{b}$	$\sim 10^{13} K_S/\text{fb}^{-1}$	2009 -
NA62	-	-	-	$> 5 \cdot 10^{12} K^+$ decays in FV	2016 -
Belle-II	$> 0.5 \text{fb}^{-1}$ (50 ab <sup>-1</sup> )	1.1 nb	1.6 nb	0.4 nb	2018 -

A wealth of experimental results in the last decade is constraining the CKM matrix with unprecedented level of precision.

# The CKM Matrix

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$$V_{\text{CKM}} \propto \begin{pmatrix} |V_{ud}| & |V_{us}| & |V_{ub}| e^{-i\gamma} \\ -|V_{cd}| & |V_{cs}| & |V_{cb}| \\ |V_{td}| e^{-i\beta} & -|V_{ts}| e^{-i\beta_s} & |V_{tb}| \end{pmatrix}$$

The SM describes the mixing of quarks of different generations through the weak force.  
3 Generations, 1 Phase: single source of CPV in the SM.

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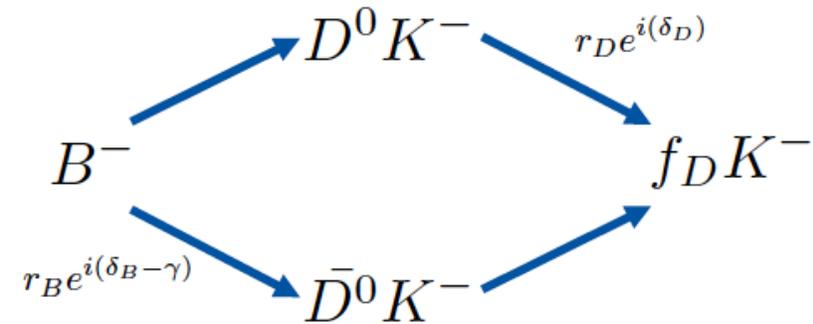
$B \rightarrow \pi\pi, \rho\rho$	$\alpha / \Phi_2$	$B \rightarrow D^* l\nu / b \rightarrow c l\nu$	$ V_{cb} $ via Form factor / OPE
$B \rightarrow D^{(*)} K^{(*)}$	$\gamma / \Phi_3$	$B \rightarrow \pi l\nu / b \rightarrow u l\nu$	$ V_{ub} $ via Form factor / OPE
$B \rightarrow J/\psi K_s$	$\beta / \Phi_1$	$M \rightarrow l\nu (\gamma)$	$ V_{ud} $ via Decay constant $f_M$
$B_s \rightarrow J/\psi \Phi$	$\beta_s$	$\epsilon_K$	$(\rho, \eta)$ via $B_K$
$K \rightarrow \pi \nu \text{ anti-}\nu$	$\rho, \eta$	$\Delta m_d, \Delta m_s$	$ V_{tb} V_{t\{d,s\}} $ via Bag factor $B_B$
		$B_{(s)} \rightarrow \mu^+ \mu^-$	$ V_{t\{d,s\}} $ via Decay constant $f_B$

# The CKM Matrix: the angle $\gamma$

$$V_{\text{CKM}} \propto \begin{pmatrix} |V_{ud}| & |V_{us}| & |V_{ub}| e^{-i\gamma} \\ -|V_{cd}| & |V_{cs}| & |V_{cb}| \\ |V_{td}| e^{-i\beta} & -|V_{ts}| e^{-i\beta_s} & |V_{tb}| \end{pmatrix}$$

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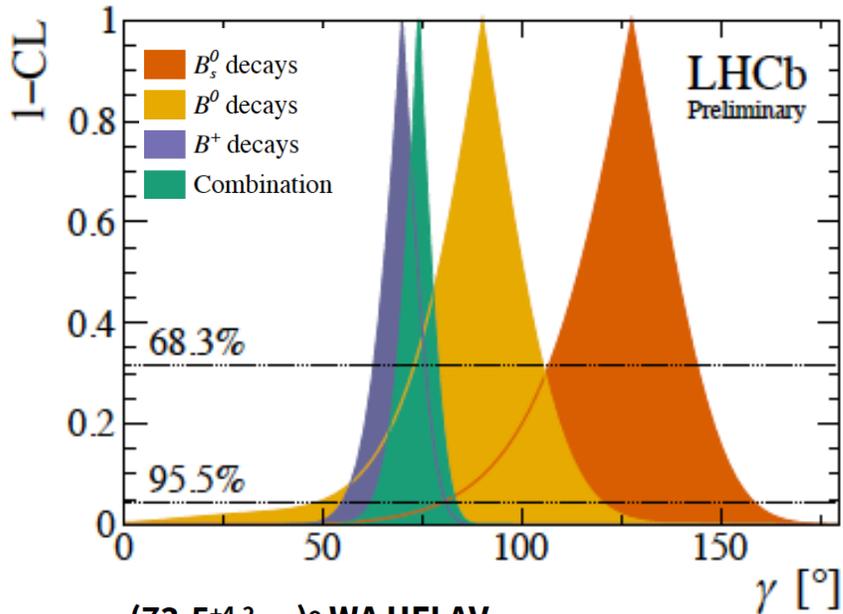


Sensitivity from interference of  $b \rightarrow c$  and  $b \rightarrow u$  amplitudes

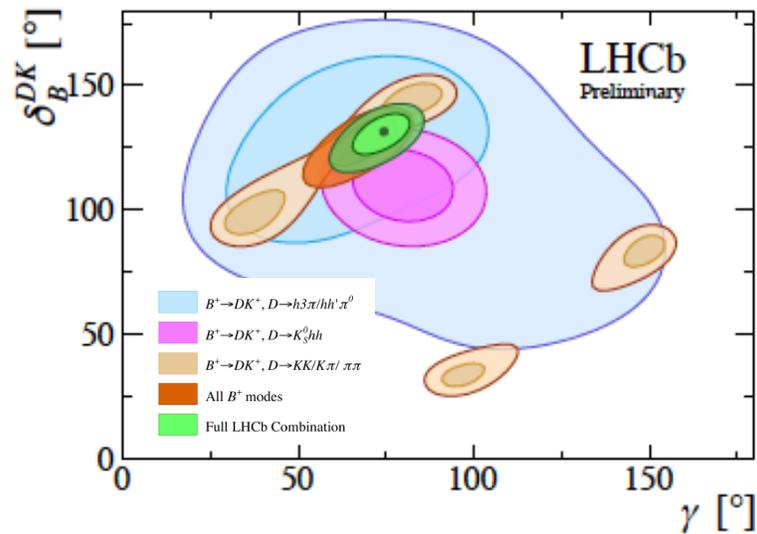
# The angle $\gamma$

Can be measured from tree-level processes only, standard candle in the SM.  
 Theory is “pristine” in this approach, theory uncertainty  $\mathcal{O}(10^{-7})$ .  
 Triumph of LHCb with  $B^0$ ,  $B^+$ ,  $B_s$  combination with D phase input from Cleo-c (BESIII data coming)  
 Now reaching an accuracy comparable to  $\alpha$ .

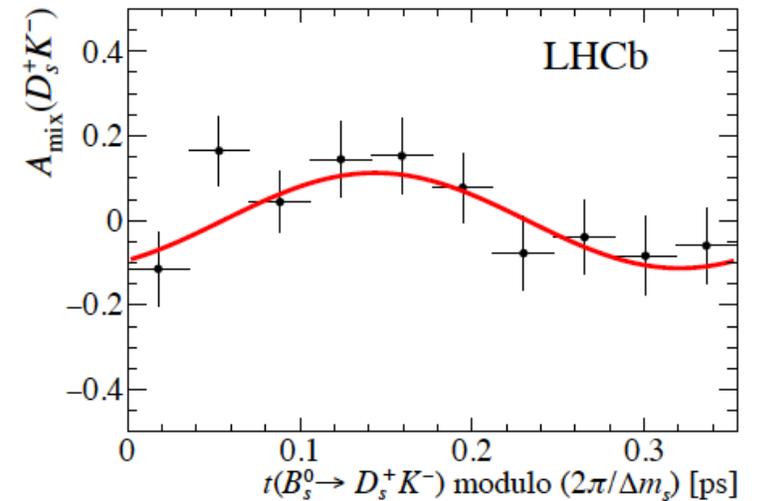
LHCb arXiv:1806.01202  
 LHCb JHEP 1803 (2018) 059  
 LHCb PLB777 (2018) 16-30  
 LHCb JHEP 1806 (2018) 084



$(73.5^{+4.2}_{-5.1})^\circ$  WA HFLAV  
 $(73.5^{+5.1}_{-5.7})^\circ$  LHCb  
 $(65.3^{+1.0}_{-2.5})^\circ$  Indirect CKMFitter



Measurement of the CKM angle  $\gamma$  using  
 $B^\pm \rightarrow DK^\pm$  with  $D \rightarrow K^0_S \pi^+ \pi^-$ ,  $K^0_S K^+ K^-$  decays



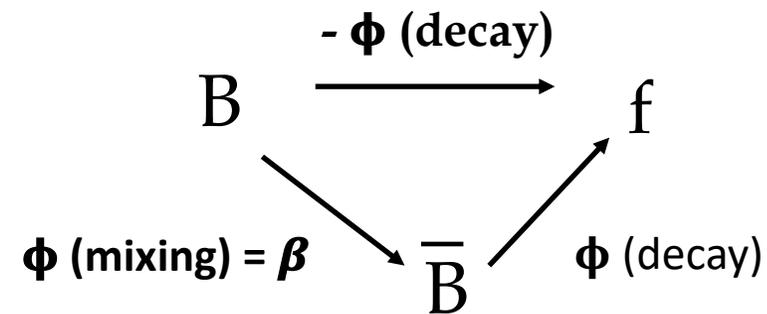
New time-dependent measurements in  
 $B \rightarrow D\pi$  ( $2\Phi_1 + \Phi_3$ ) and  $B_s \rightarrow D_s K$  ( $\Phi_5 - \Phi_3$ )

# The CKM Matrix: the angle $\beta$

$$V_{\text{CKM}} \propto \begin{pmatrix} |V_{ud}| & |V_{us}| & |V_{ub}| e^{-i\gamma} \\ -|V_{cd}| & |V_{cs}| & |V_{cb}| \\ |V_{td}| e^{-i\beta} & -|V_{ts}| e^{-i\beta_s} & |V_{tb}| \end{pmatrix}$$

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$K \rightarrow \pi \nu \text{ anti-}\nu$	$\rho, \eta$

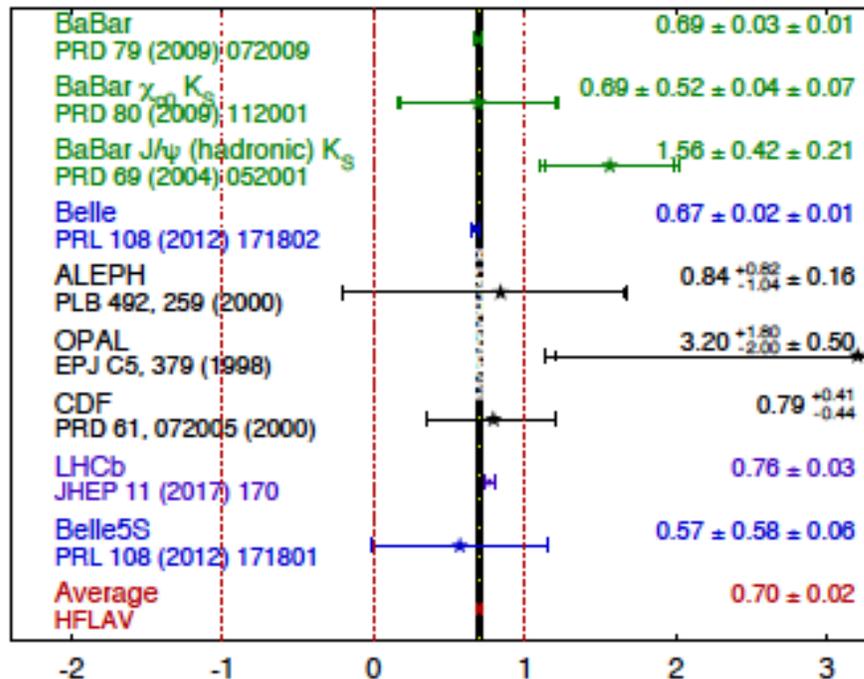


Sensitivity from interference between decay with and without mixing

# The angle $\beta$

Belle+Babar arXiv: 1804.06153  
 Belle+Babar arXiv: 1804.06152  
 LHCb JHEP 1711 (2017) 170  
 LHCb PRD95 5, 052005 (2017)

$\sin(2\beta) \equiv \sin(2\phi_1)$  **HFLAV**  
 Moriond 2018  
 PRELIMINARY



$0.70 \pm 0.02$  WA HFLAV

$0.740^{+0.020}_{-0.025}$  Indirect CKMfitter

Ambiguity in the sign of  $\cos(2\beta)$  solved by BaBar+Belle:  
 $\cos(2\beta) > 0$  at  $7\sigma$

Penguin pollution constrained by the measurement of the  $B^+ \rightarrow J/\psi K^+$  asymmetry well below the current experimental uncertainty on  $\sin(2\beta)$ :

$$\phi^{\text{eff}}(\text{obs}) = \phi(\text{tree}) + \Delta\phi(\text{penguin}) + \phi(\text{NP})$$

$$\Delta\phi(\text{penguin}) = (-1.10^{+0.70}_{-0.85})^\circ$$

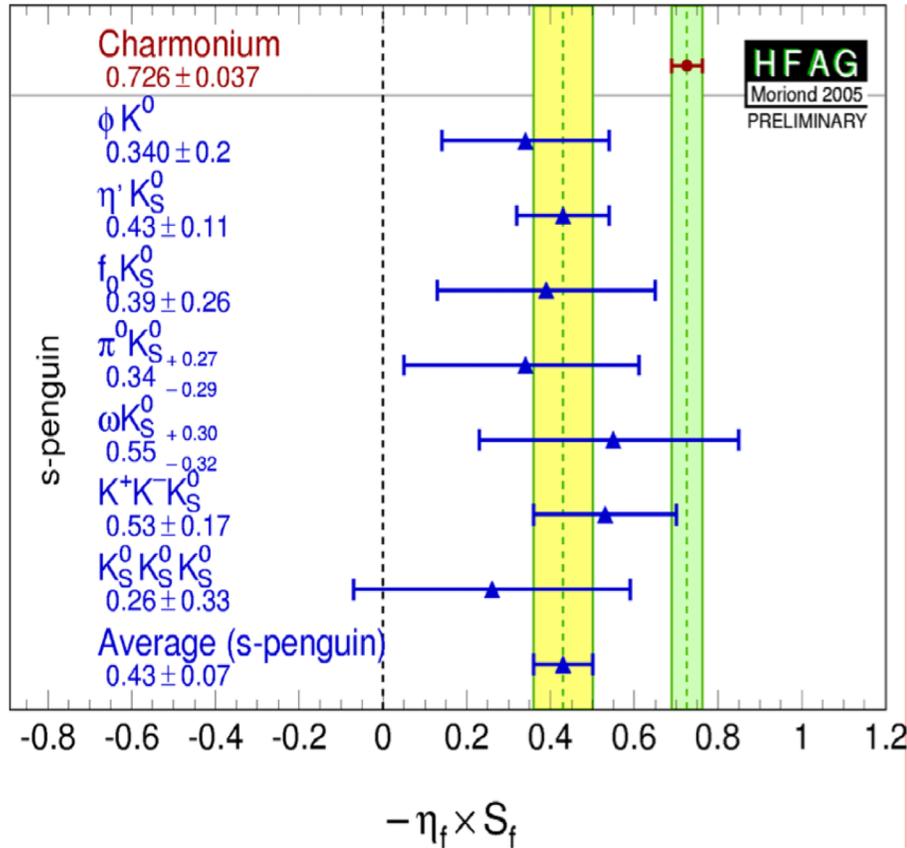
$$\Delta\phi(\text{obs}) = 1.6^\circ$$

Accuracy from LHCb rapidly approaching that of B-factories. Very good agreement with indirect  $\sin(2\beta)$  determination, no hints of NP.

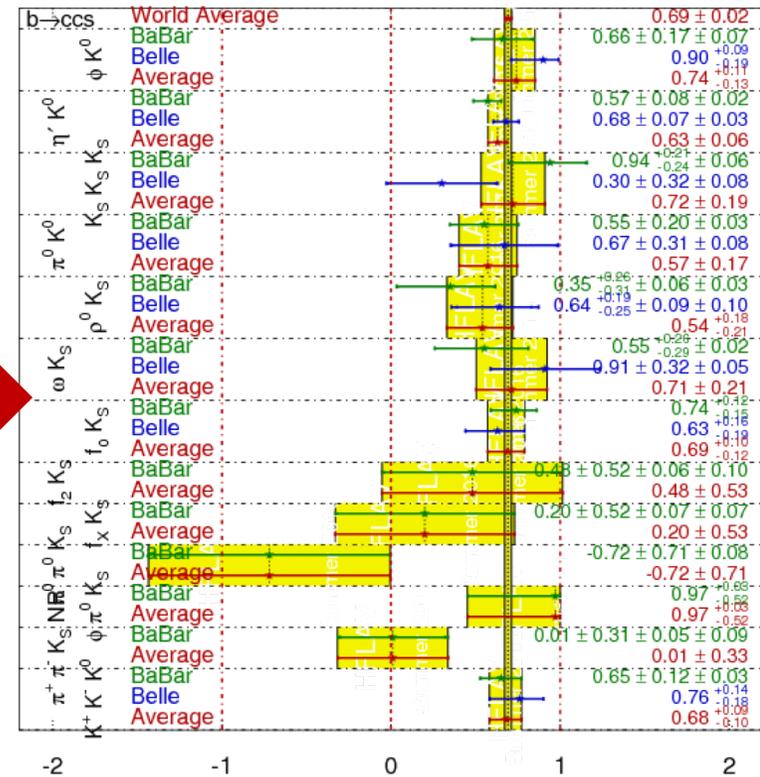
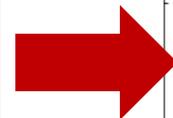
.... and this has not always been the case:

CKM2005 –  $3.7\sigma$  between  $b \rightarrow c\bar{c}s$  and s-penguins

$\sin(2\beta^{\text{eff}})$  in 2005



$\sin(2\beta^{\text{eff}}) \equiv \sin(2\phi_1^{\text{eff}})$  **HFLAV** Summer 2016



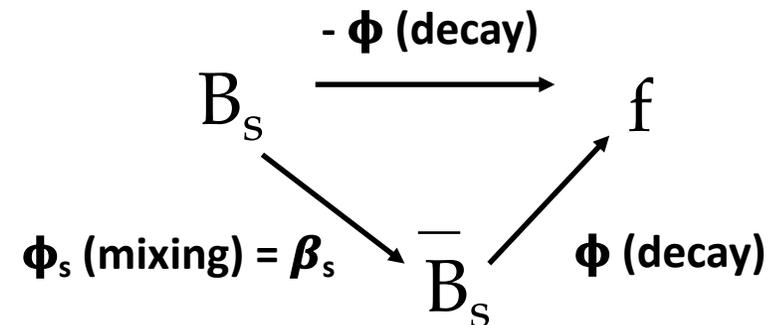
.... Anomalies come and go ...

# The CKM Matrix: the angle $\beta_s$

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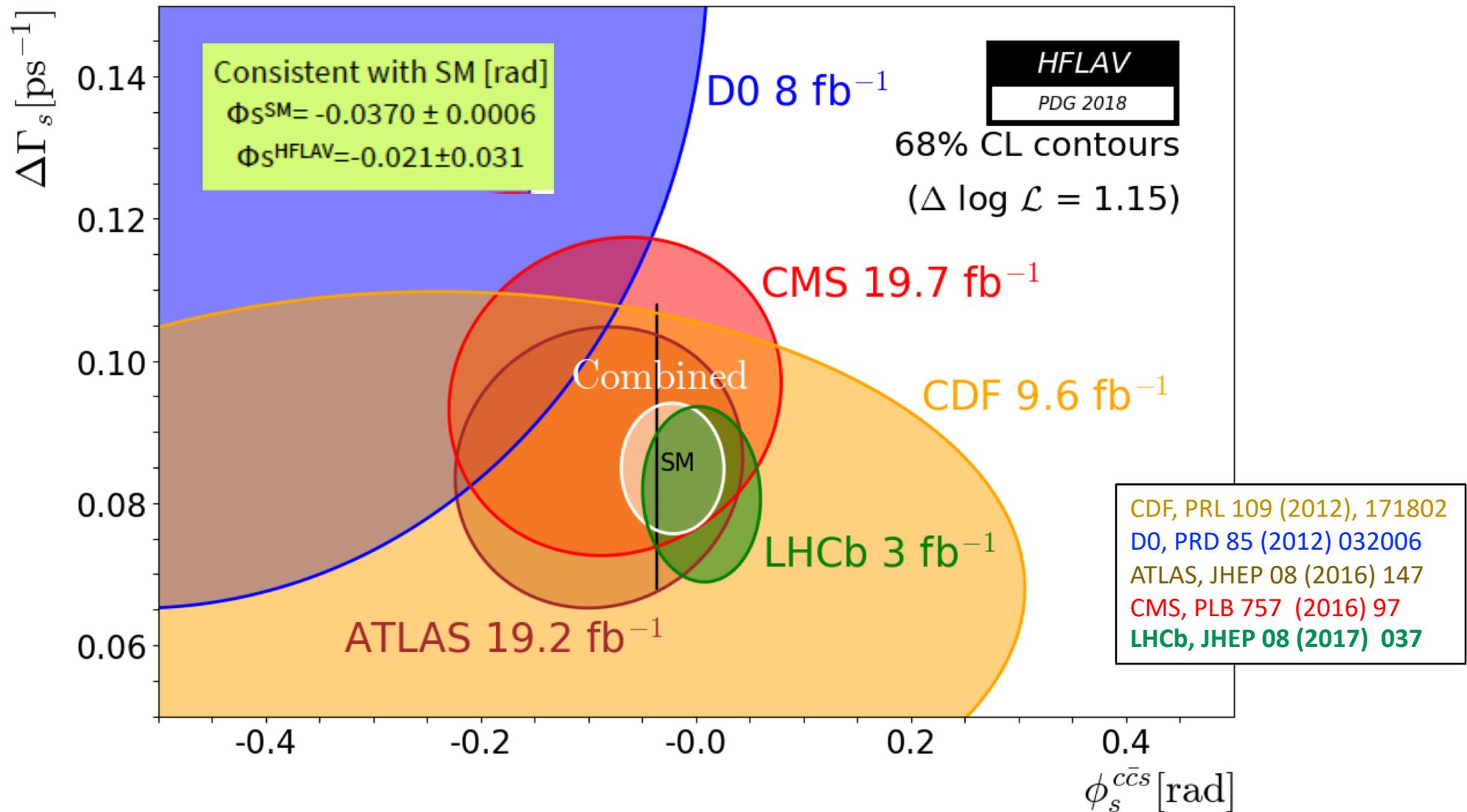
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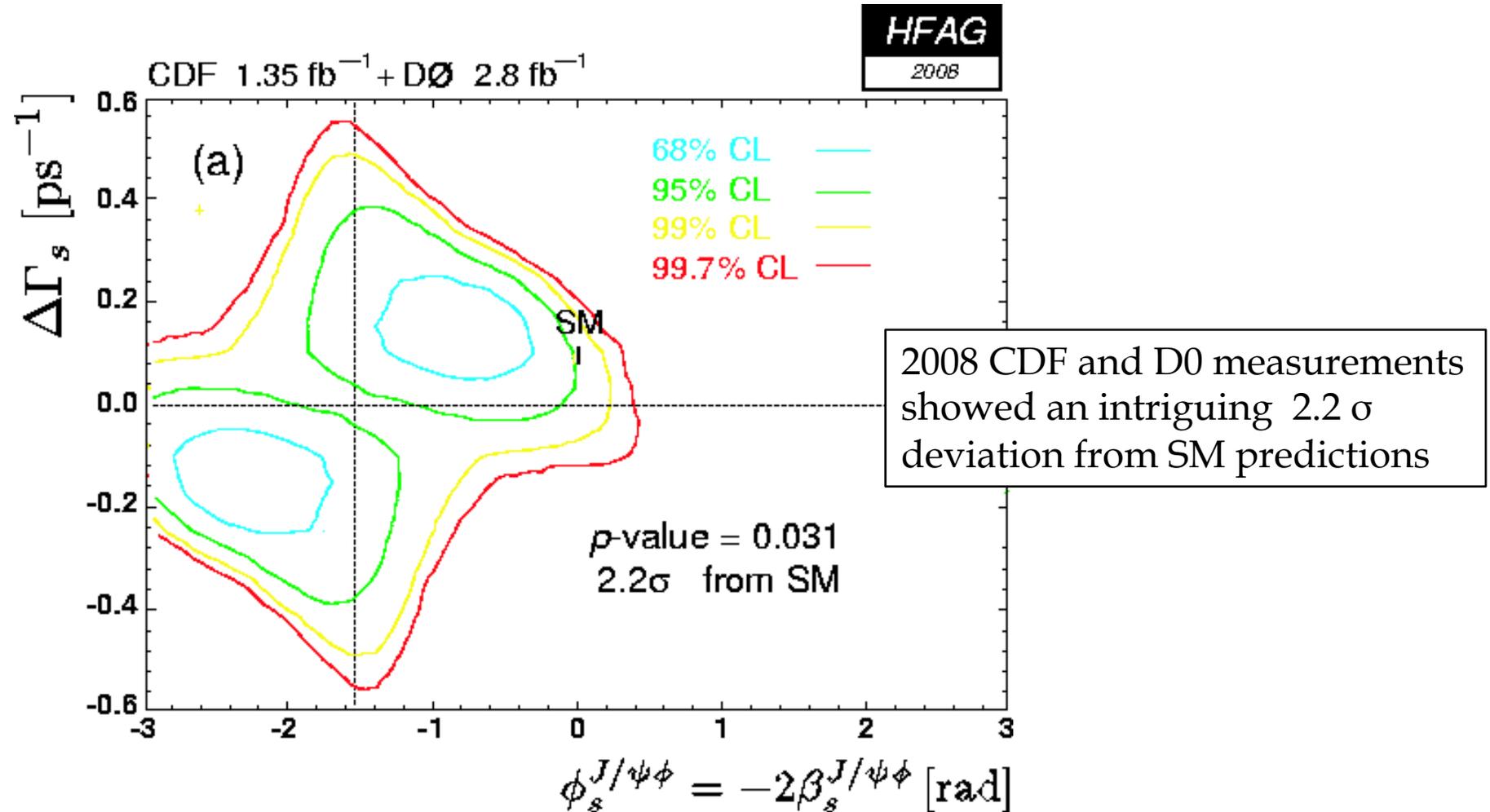
Sensitivity from interference between decay with and without mixing

# The angle $\phi_s = -2\beta_s$



Remarkable agreement between measurements and SM predictions

.... and also in this case this was not always the case:



.... Anomalies come and go .....

# Scale of New Physics from the analysis of FCNC $\Delta F=2$ transitions

The SM is considered the low-energy effective theory of a more complete theory which holds at a higher scale. New particles are expected to show up at the EW scale to solve the hierarchy problem and they are expected to have sizeable couplings to SM particles, hence they can be detected either via direct detection or via indirect effects in loops.

$$C_{B_q} e^{2i\phi_{B_q}} = \frac{\langle B_q | H_{\text{eff}}^{\text{full}} | \bar{B}_q \rangle}{\langle B_q | H_{\text{eff}}^{\text{SM}} | \bar{B}_q \rangle} = \frac{F_i L_i}{\Lambda^2} = \frac{C_i^{NP}}{\Lambda^2} e^{i\phi_{NP}} \quad \text{Similar for K and D}$$

$F_i$  is a function of the (complex) NP flavor coupling,  
 $L_i$  is a loop factor that is present in models with no tree-level FCNC  
 $\Lambda$  is the scale of NP, typical mass of new particles mediating  $\Delta F = 2$  transitions

Scale of  
New Physics 

$$C_i^{NP} = 1, \phi_i^{NP} = \text{any value (generic flavour),}$$

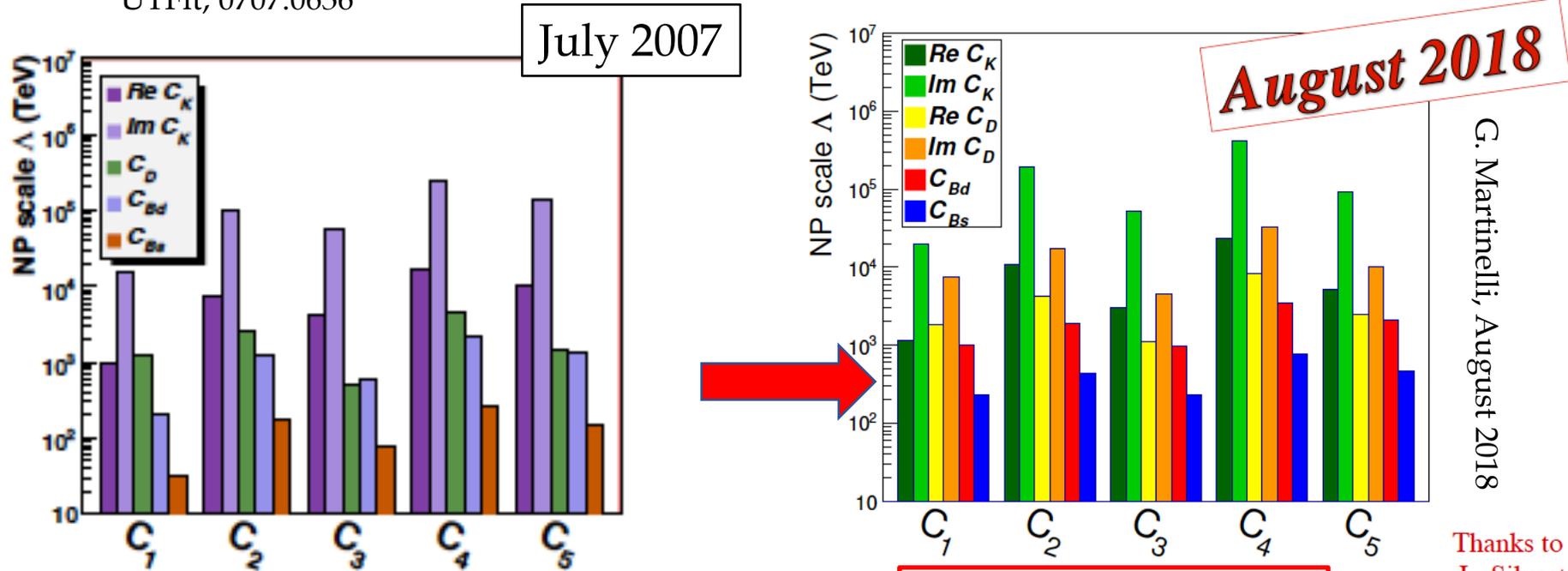
$$C_i^{NP} = \text{CKM}, \phi_i^{NP} = \text{any value (NMFV),}$$

$$C_i^{NP} = \text{CKM}, \phi_i^{NP} = \text{SM (MFV)}$$

# Limits on the $\Lambda$ scale for strong coupling, $C_i$ (NP) $\sim 1$ , any phase

Dramatic increase of the  $\Lambda$  scale from low-energy measurements in the last ten years  
 Constraints from Kaon physics still dominate ( $\Lambda \sim 10^5$ ), followed by charm ( $\Lambda \sim 10^4$ ), beauty ( $\Lambda \sim 10^3$ )

UTFit, 0707.0636



Possible NP operators:

$$Q_1^{q_i q_j} = \bar{q}_{jL}^\alpha \gamma_\mu q_{iL}^\alpha \bar{q}_{jL}^\beta \gamma^\mu q_{iL}^\beta,$$

$$Q_2^{q_i q_j} = \bar{q}_{jR}^\alpha q_{iL}^\alpha \bar{q}_{jR}^\beta q_{iL}^\beta,$$

$$Q_3^{q_i q_j} = \bar{q}_{jR}^\alpha q_{iL}^\beta \bar{q}_{jR}^\beta q_{iL}^\alpha,$$

$$Q_4^{q_i q_j} = \bar{q}_{jR}^\alpha q_{iL}^\alpha \bar{q}_{jL}^\beta q_{iR}^\beta,$$

$$Q_5^{q_i q_j} = \bar{q}_{jR}^\alpha q_{iL}^\beta \bar{q}_{jL}^\beta q_{iR}^\alpha.$$

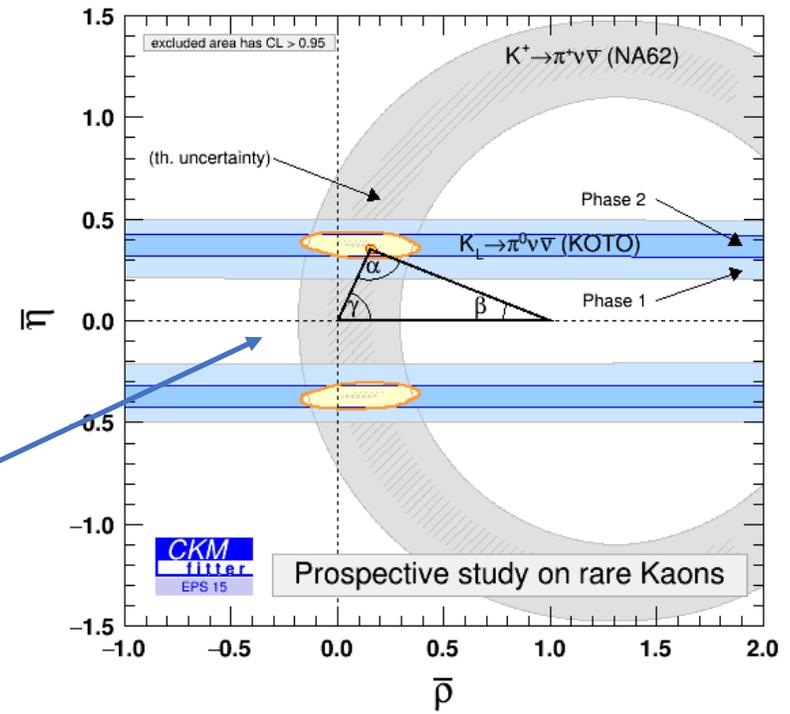
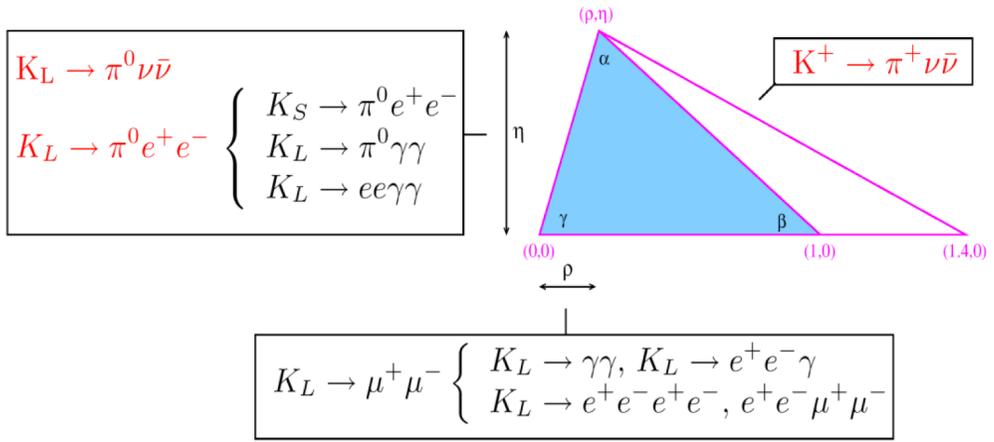
$\varepsilon_K$	$\Lambda = 5 \cdot 10^5 \text{ TeV}$
D	$\Lambda = 5 \cdot 10^4 \text{ TeV}$
$B_d$	$\Lambda = 5 \cdot 10^3 \text{ TeV}$
$B_s$	$\Lambda = 10^3 \text{ TeV}$

Thanks to  
L. Silvestrini

New Physics either is very heavy or it mimics the SM in its flavor-breaking pattern...

# A completely independent way of constraining CKM: rare kaon decays

K physics alone can fully constrain the CKM unitarity triangle.  
Comparison with B physics can provide description of NP flavour dynamics



## Theory predictions

$$BR(K^0 \rightarrow \pi^0 \nu \bar{\nu}) = (3.4 \pm 0.6) \times 10^{-11}$$

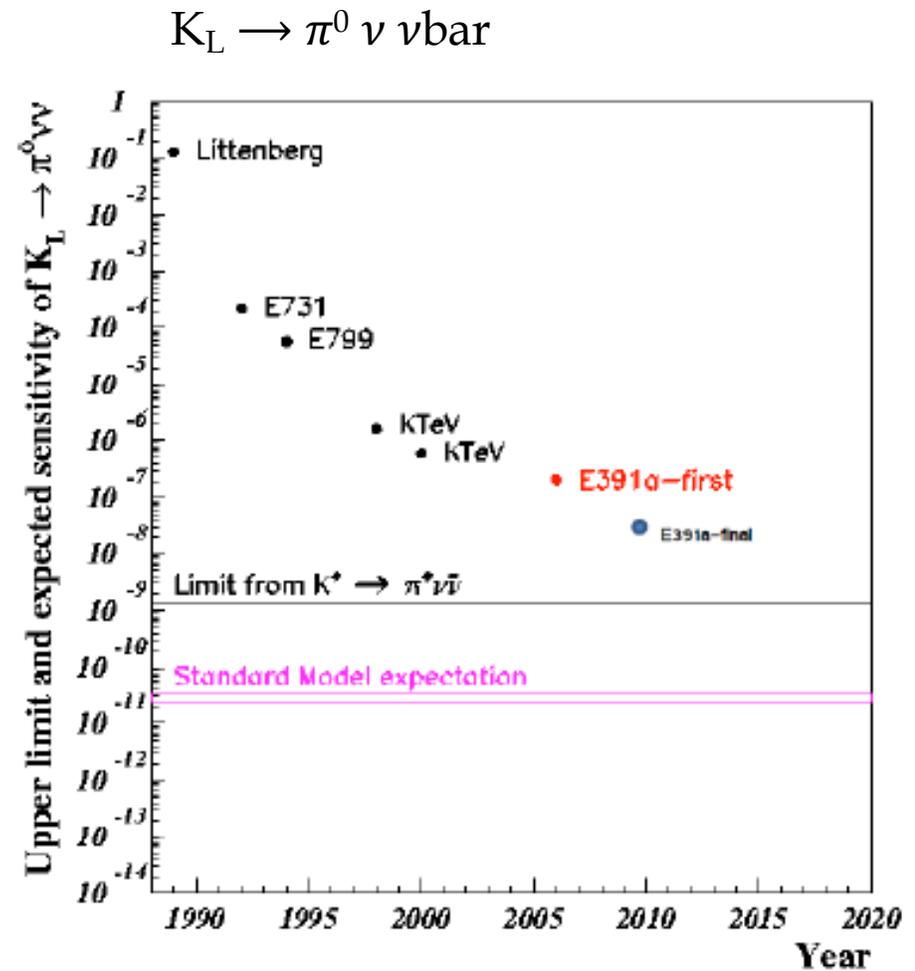
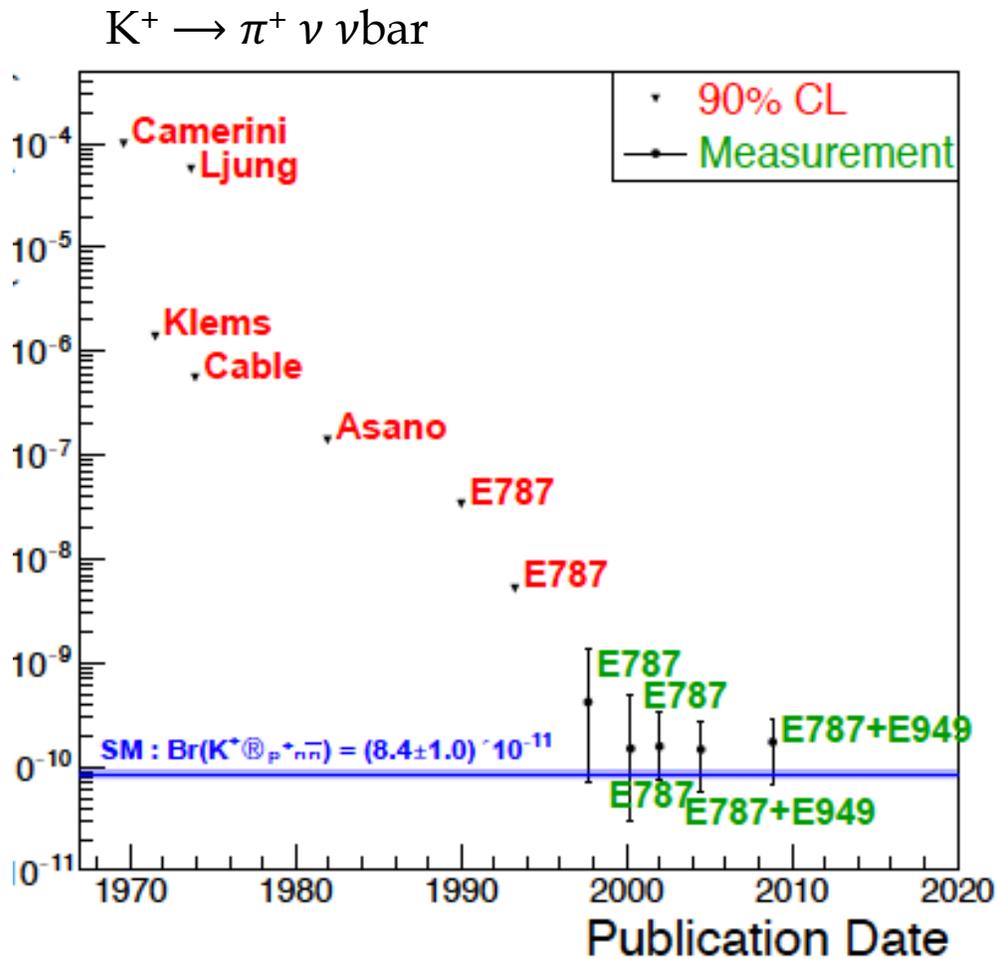
$$BR(K^+ \rightarrow \pi^+ \nu \bar{\nu}) = (8.4 \pm 1.0) \times 10^{-11}$$

[Buras et al. JHEP 1511 (2015) 33]

## Constraints in the $(\bar{\rho}, \bar{\eta})$ plane, using a prospective scenario for the NA62 and KOTO experiments.

- For NA62, a measurement of  $BR(K^+ \rightarrow \pi^+ \nu \bar{\nu})$  with a 10% accuracy is assumed;
- For KOTO, a two-step prospective scenario is assumed : first, a  $3\sigma$  evidence for the  $K_L \rightarrow \pi^0 \nu \bar{\nu}$  decay (Phase 1, lighter blue), followed by a later measurement of  $BR(K_L \rightarrow \pi^0 \nu \bar{\nu})$  with 10% accuracy (Phase 2, darker blue).

# $K \rightarrow \pi \nu \bar{\nu}$ : experimental status - the past



$$\text{BR}(K^+ \rightarrow \pi^+ \nu \bar{\nu}) = (17.3_{-10.5}^{+11.5}) \times 10^{-11} \quad \text{Phys. Rev. D 77, 052003 (2008), Phys. Rev. D 79, 092004 (2009)}$$

$$\text{BR}(K_L \rightarrow \pi^0 \nu \bar{\nu}) < 2.6 \times 10^{-8} \quad (90\% \text{ C. L.}) \quad \text{Phys. Rev. D 81, 072004 (2010)}$$

# The NA62 experiment @ K12 in EHN3

<https://na62.web.cern.ch/>



NA62 is taking data with the aim of measuring  $\text{BR}(\text{K}^+ \rightarrow \pi^+ \nu \bar{\nu})$  with 10% accuracy

# $K \rightarrow \pi \nu \nu$ experimental status: the present/future

□ **NA62 is taking data with the aim at measuring  $BR(K^+ \rightarrow \pi^+ \nu \nu \text{ bar})$  with 10% accuracy:**

- Current limit based of 2016 data ( $10^{11}$  K decays,  $\sim 1\%$  of the total data sample)

$$BR_{2016}(K^+ \rightarrow \pi^+ \nu \bar{\nu}) < 14 \times 10^{-10} @ 95\% CL$$

- 20 SM events expected with the 2017-2018 dataset.

- data taking after 2018 to be approved:

- expected ultimate sensitivity under evaluation.

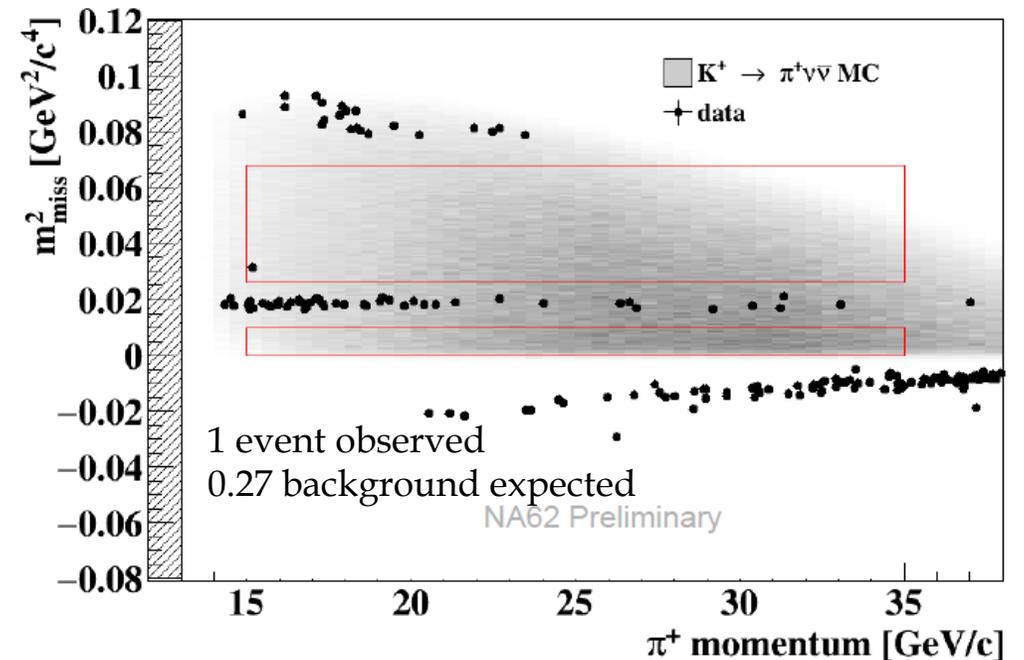
□ **KOTO has presented at ICHEP 2018 a limit based on the 2015 dataset:**

$$BR(K_L \rightarrow \pi^0 \nu \nu) < 3.0 \times 10^{-9} @ 90\% CL$$

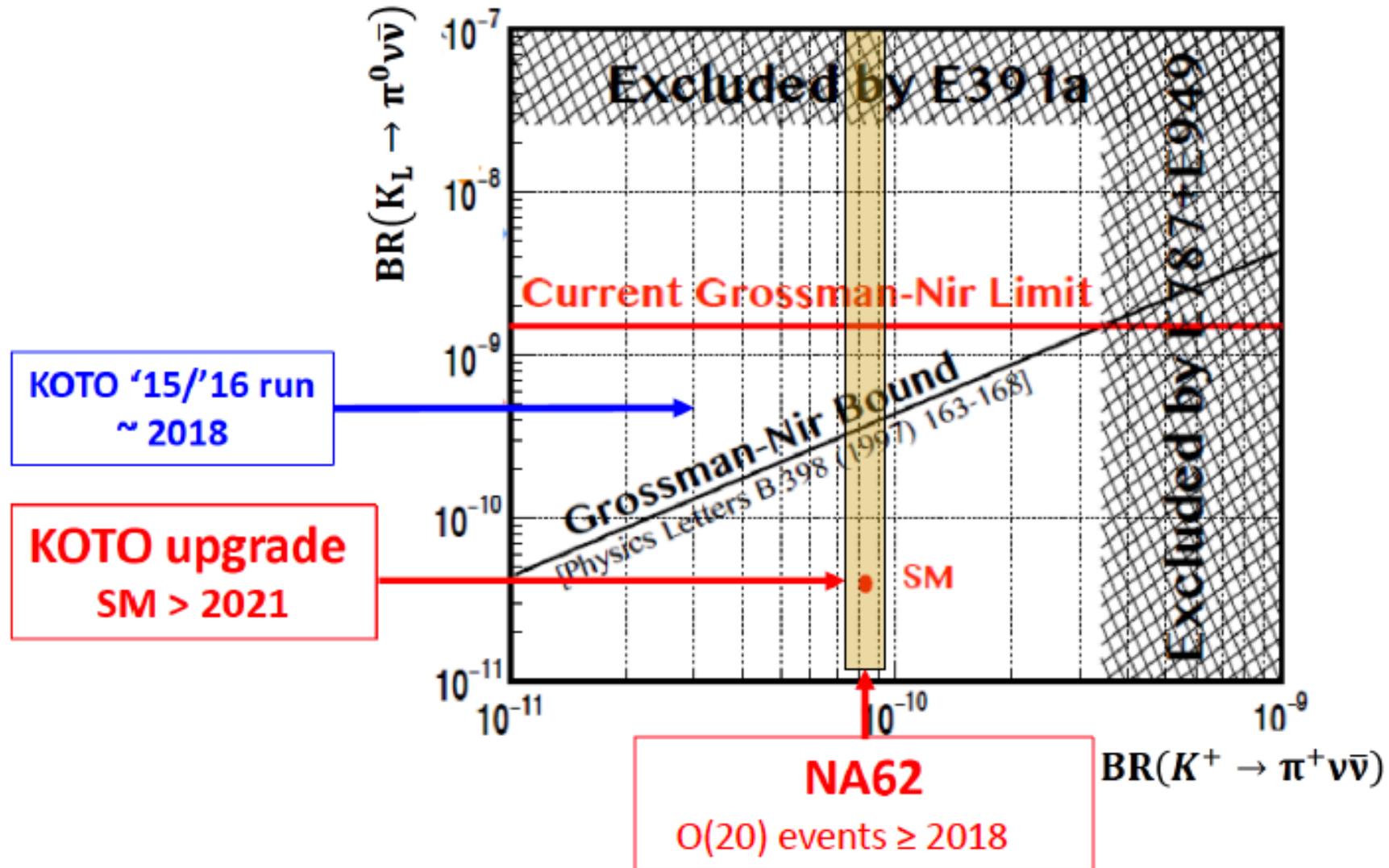
and aims to reach the SM sensitivity after 2021.

□ **KLEVER proposal at CERN**

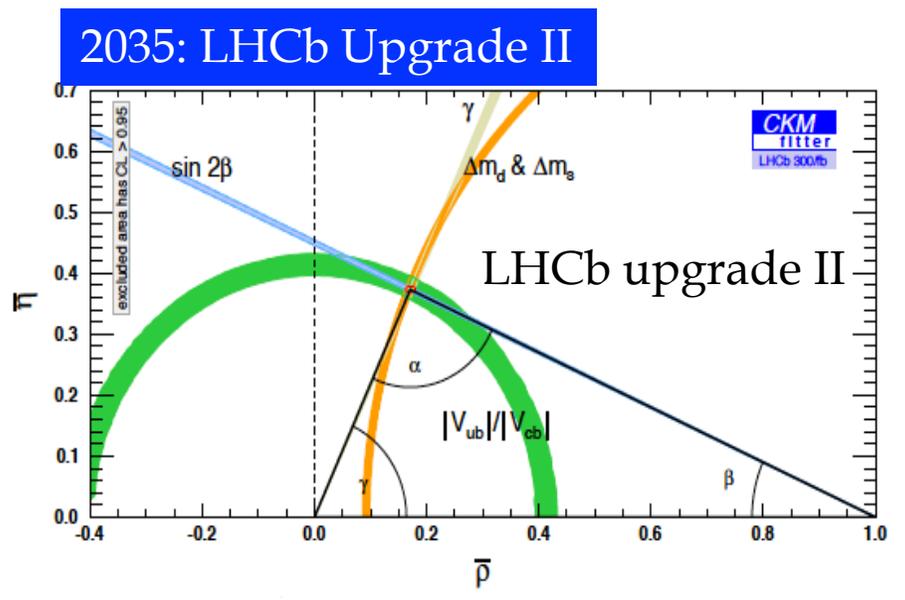
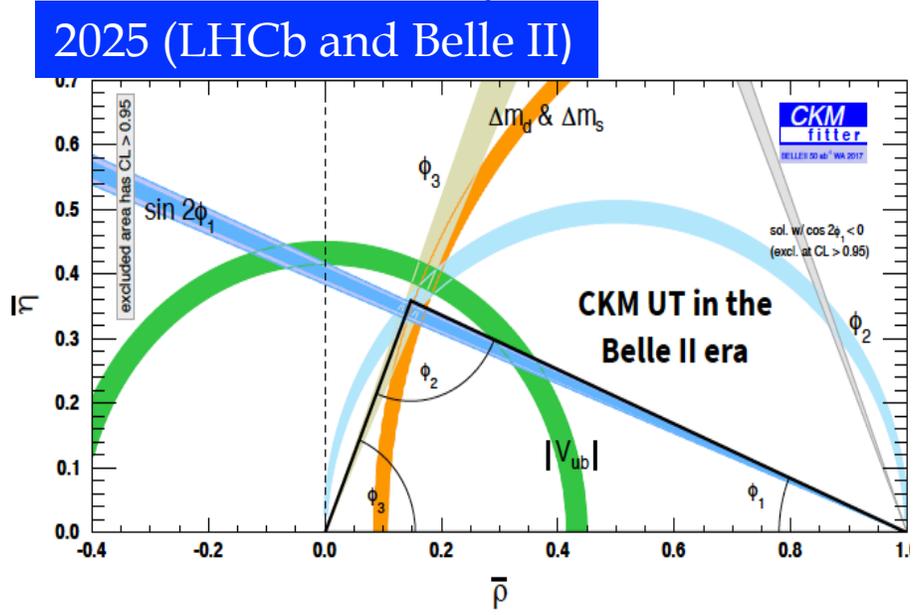
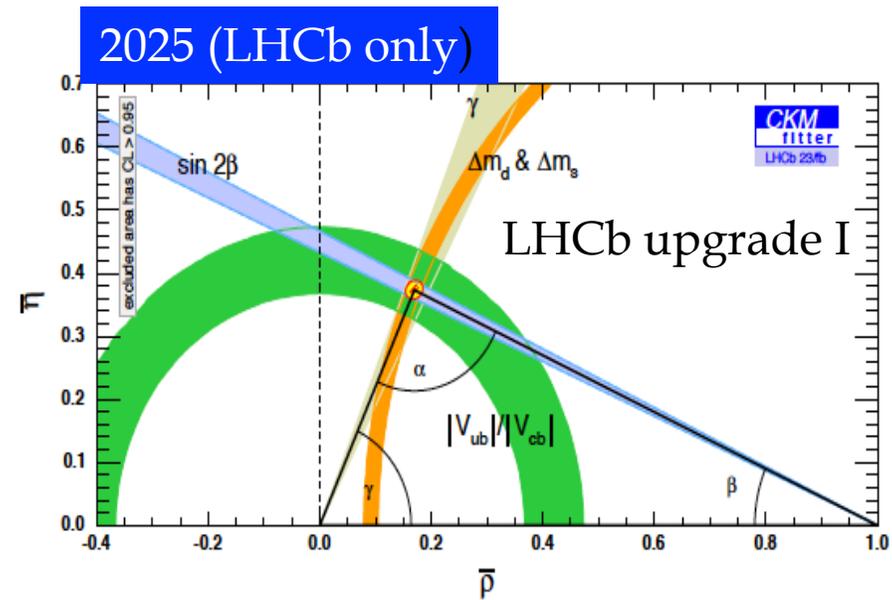
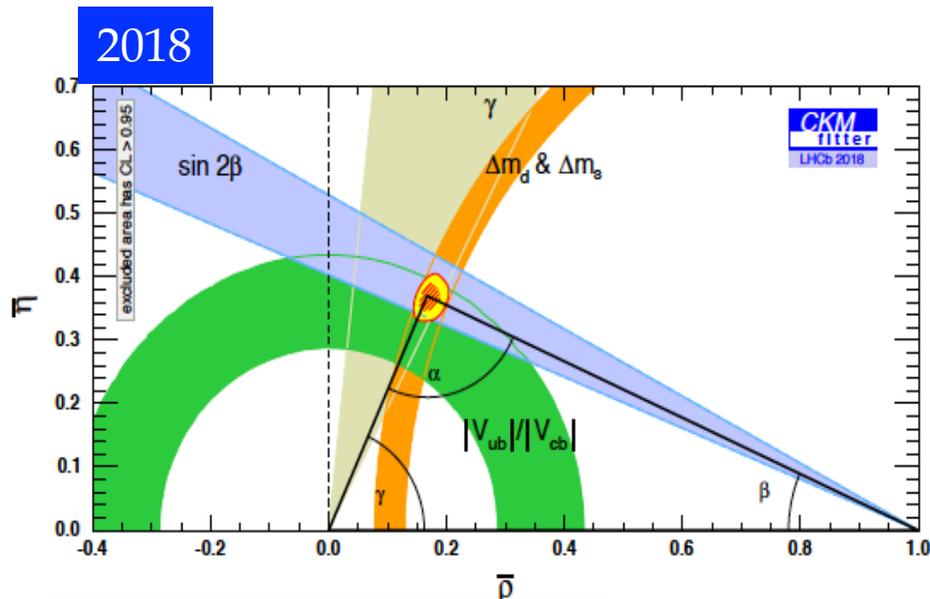
aims to measure the  $BR(K_L \rightarrow \pi^0 \nu \nu \text{ bar})$  with 20% accuracy by 2030++



# $K \rightarrow \pi \nu \bar{\nu}$ : near prospects



# Progress on the CKM knowledge expected in the next ~10-20 years



Rare B decays  
(or: the famous B anomalies)

# Lepton Flavor Universality

Lepton Flavor Universality (LFU) is a basic property of the SM: universal coupling with the three lepton generations.

$$\mathcal{L}_f = \bar{f} i D_\mu \gamma^\mu f \quad f = l_L^i, q_L^i, \quad i = 1, 2, 3 \quad \text{for each of three generations in weak interactions}$$

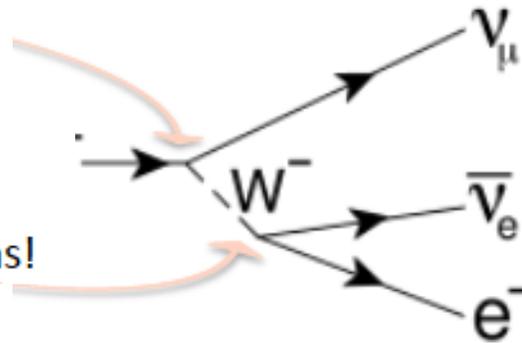
$$D_\mu = \partial_\mu + ig \frac{1}{2} \vec{\tau} \cdot \vec{W}_\mu + ig' \frac{1}{2} Y_W B_\mu$$

$$\mathcal{L}_{eff} = -\frac{G_F}{\sqrt{2}} J_\mu^\dagger J^\mu$$

$$\frac{g^2}{8m_W^2} = \frac{G_F}{\sqrt{2}}$$

the same for all SM fermions

the same coupling of lepton and its neutrino with  $W$  for all three lepton generations!



# Lepton Flavor Universality

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the same for all SM fermions

$$\mathcal{L}_{eff} = -\frac{G_F}{\sqrt{2}} J_\mu^\dagger J^\mu$$

$$\frac{g^2}{8m_W^2} = \frac{G_F}{\sqrt{2}}$$

However LFU is not a fundamental symmetry of the SM:

it is accidental in the gauge sector, where the gauge bosons equally couple to the different lepton flavors.

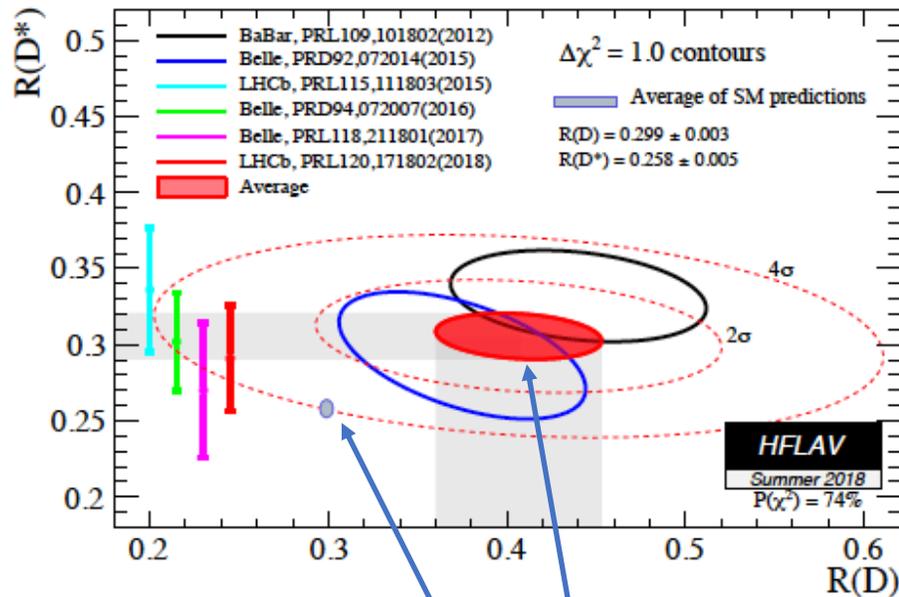
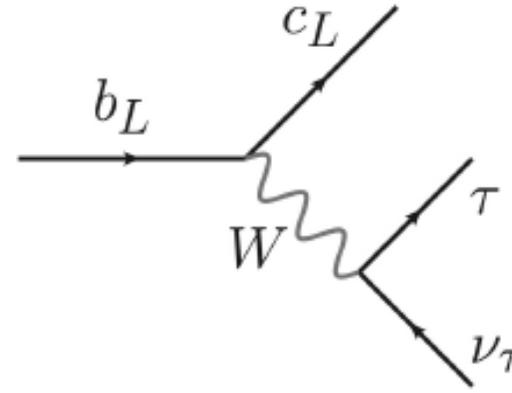
LFU is broken by the Higgs couplings to masses, which are flavor-specific,

but they have a negligible effect on the partial widths of the decays.

**LFU tested up to percent precision in  $Z \rightarrow l^+ l^-$  and  $J/\psi \rightarrow l^+ l^-$  decays (PRD 98 (2018) 030001)**

# The first B-anomaly: $b \rightarrow$ charm charged currents, tree level

$$R_{D^{(*)}} = \frac{BR(B \rightarrow D^{(*)} \tau \nu_\tau)}{BR(B \rightarrow D^{(*)} \mu \nu_\mu)}$$



World average near  $3.9 \sigma$  away from SM predictions

HFLAV averages:

$$R(D^*) = 0.304 \pm 0.013 \pm 0.007 \quad (3.4 \sigma)$$

$$R(D) = 0.407 \pm 0.039 \pm 0.024 \quad (2.3 \sigma)$$

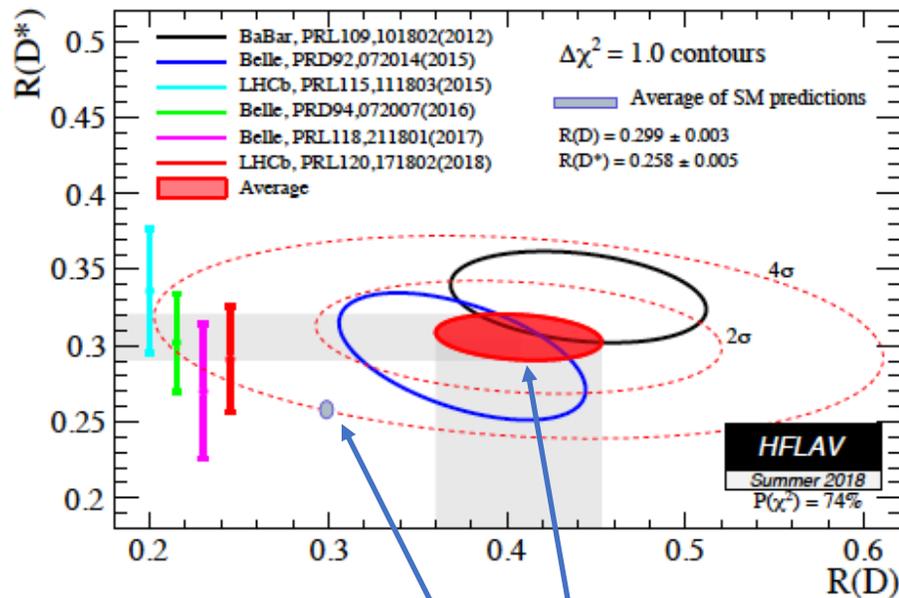
Theory predictions:

$$R(D^*) = 0.252 \pm 0.003 \quad \text{Fajfer, 2012}$$

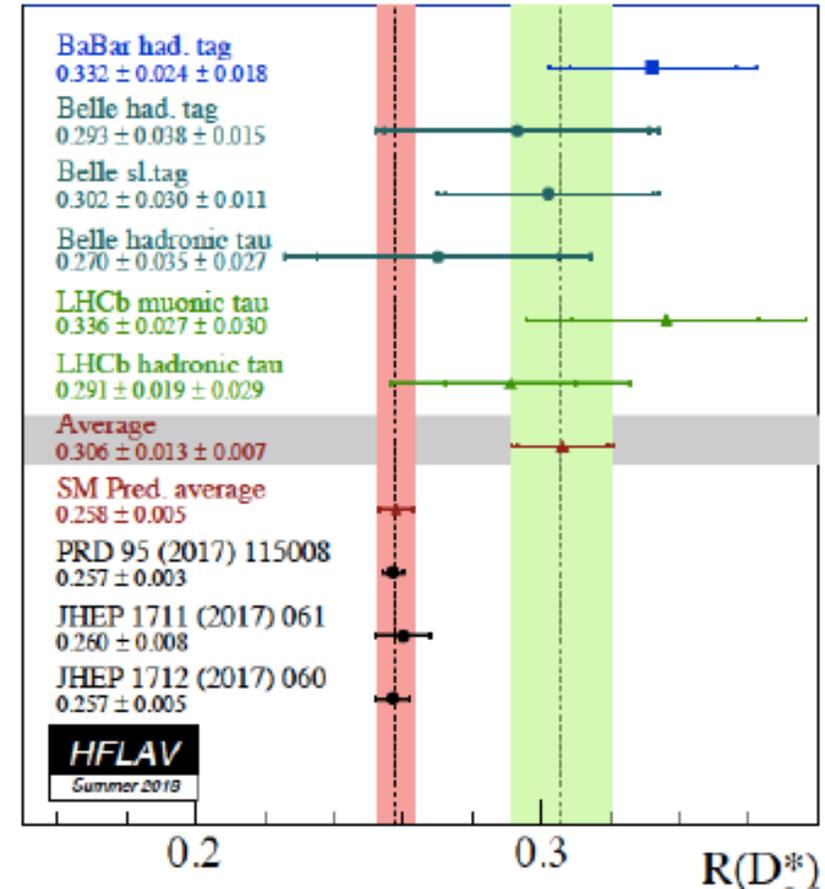
$$R(D) = 0.300 \pm 0.008 \quad \text{HFLAV average}$$

# The first B-anomaly: $b \rightarrow$ charm charged currents, tree level

$$R_{D^{(*)}} = \frac{BR(B \rightarrow D^{(*)} \tau \nu_\tau)}{BR(B \rightarrow D^{(*)} \mu \nu_\mu)}$$



World average near  $3.8 \sigma$  away from SM predictions

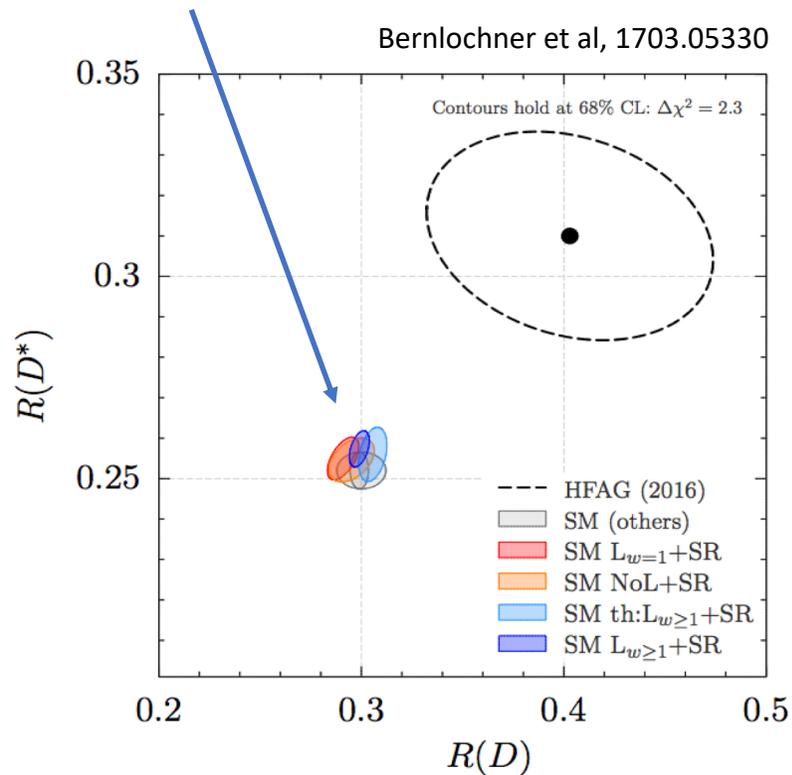


However:

- $R(D^*)$  with hadronic modes are  $1 \sigma$  consistent with SM
- Belle  $R(D^*)$  combination  $< 2 \sigma$  from SM

# The first B-anomaly: $b \rightarrow$ charm charged currents, tree level

Theory well under control:



**Theory predictions:**

Fajfer, Kamenik, Nisandzic, 1203.2654

Bailey et al, 1206.4992

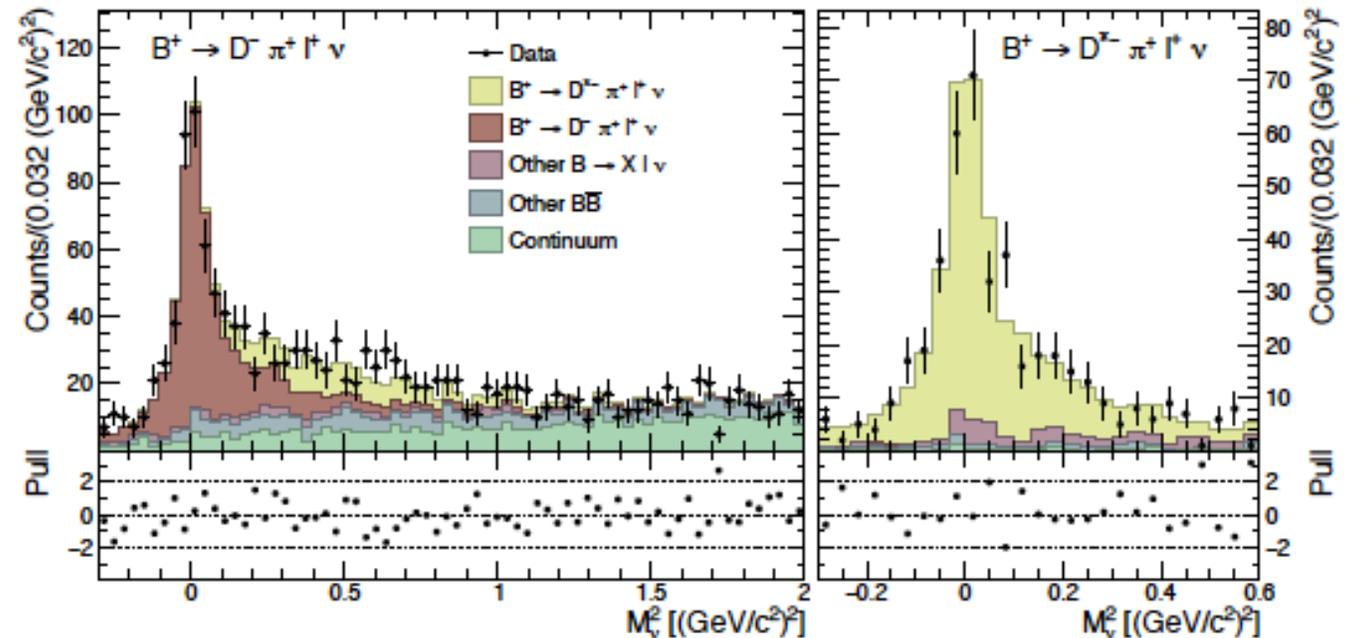
Beciveric et al., 1206.4977

Bernlochner, Ligeti, Papucci, Robinson, 1703.05330

Bigi, Gambino, Schacht, 1707.09509

Experimental caveat:

$\tau$  are mimicked by background (enhancement).  
more studies of  $B \rightarrow D^{**}(\rightarrow D n \pi) l \nu$  necessary

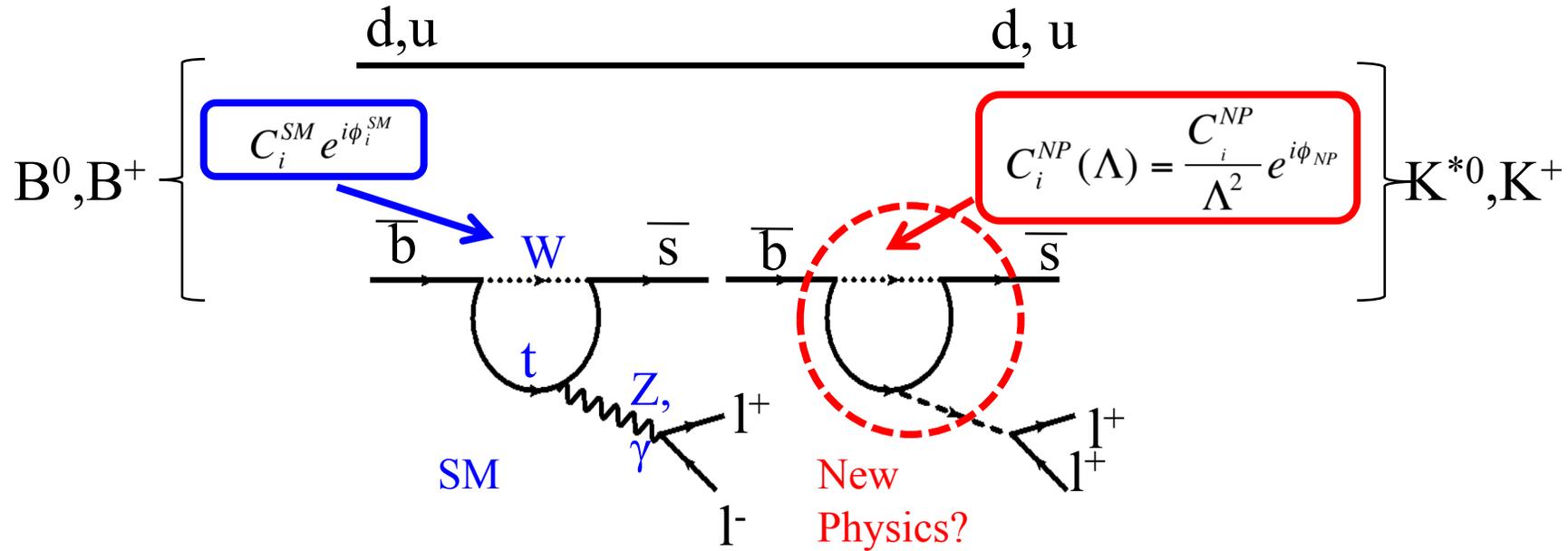


Measurement of the branching fraction of  $B \rightarrow D^{(*)} \pi l \nu$  at Belle using hadronic tagging in fully reconstructed events

P. Urquijo, ICHEP 2018

Belle, PRD 98 (2018) 012005, arXiv 1803.06444

The second set of b-anomalies:  
FCNC semi-leptonic transitions  $b \rightarrow s l^+ l^-$



$$H_{\text{eff}} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{tq}^* \frac{e^2}{16\pi^2} \sum_i (C_i Q_i + C'_i Q'_i) + h.c.$$

- $i = 1, 2$  Tree
- $i = 3 - 6, 8$  Gluon penguin
- $i = 7$  Photon penguin
- $i = 9, 10$  Electroweak penguin
- $i = S$  Higgs (scalar) penguin
- $i = P$  Pseudoscalar penguin

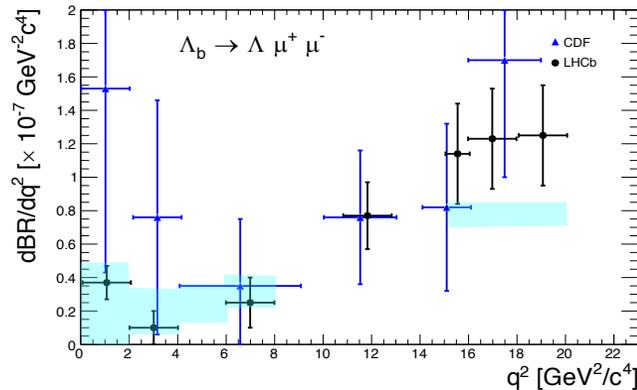
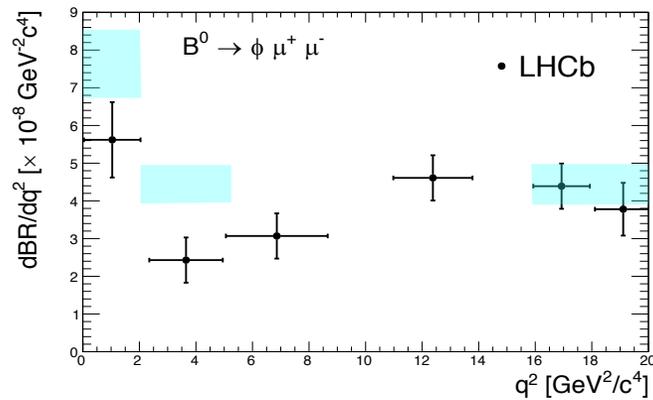
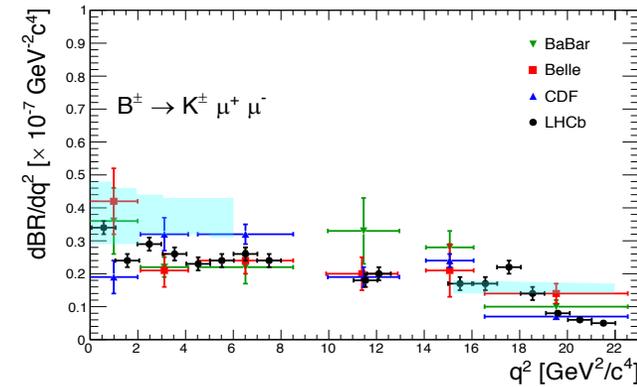
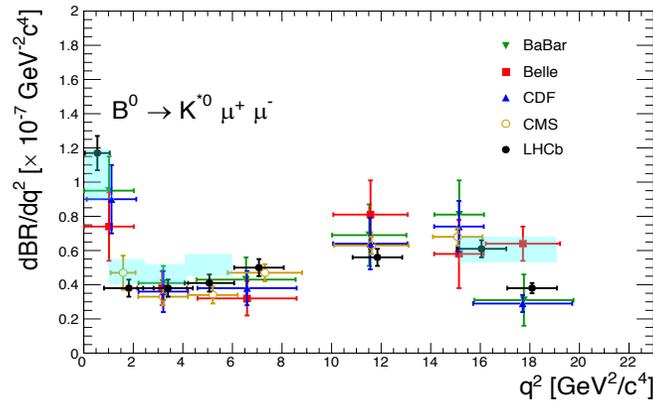
Scale of  
New Physics



- $C_i^{NP} = 1, \phi_i^{NP} = \text{any value (generic flavour),}$
- $C_i^{NP} = \text{CKM}, \phi_i^{NP} = \text{any value (NMFV),}$
- $C_i^{NP} = \text{CKM}, \phi_i^{NP} = \text{SM (MFV)}$

# FCNC semileptonic $b \rightarrow s l^+ l^-$ decays : Decay Rates

The experimental measurements of the BR tend to lie below the SM expectations across the full  $q^2$  range. The discrepancy is largest for the  $B_s \rightarrow \phi \mu \mu$  decay in the large recoil region ( $1 < q^2 < 6 \text{ GeV}^2/c^4$ ),  $3\sigma$  discrepancy.



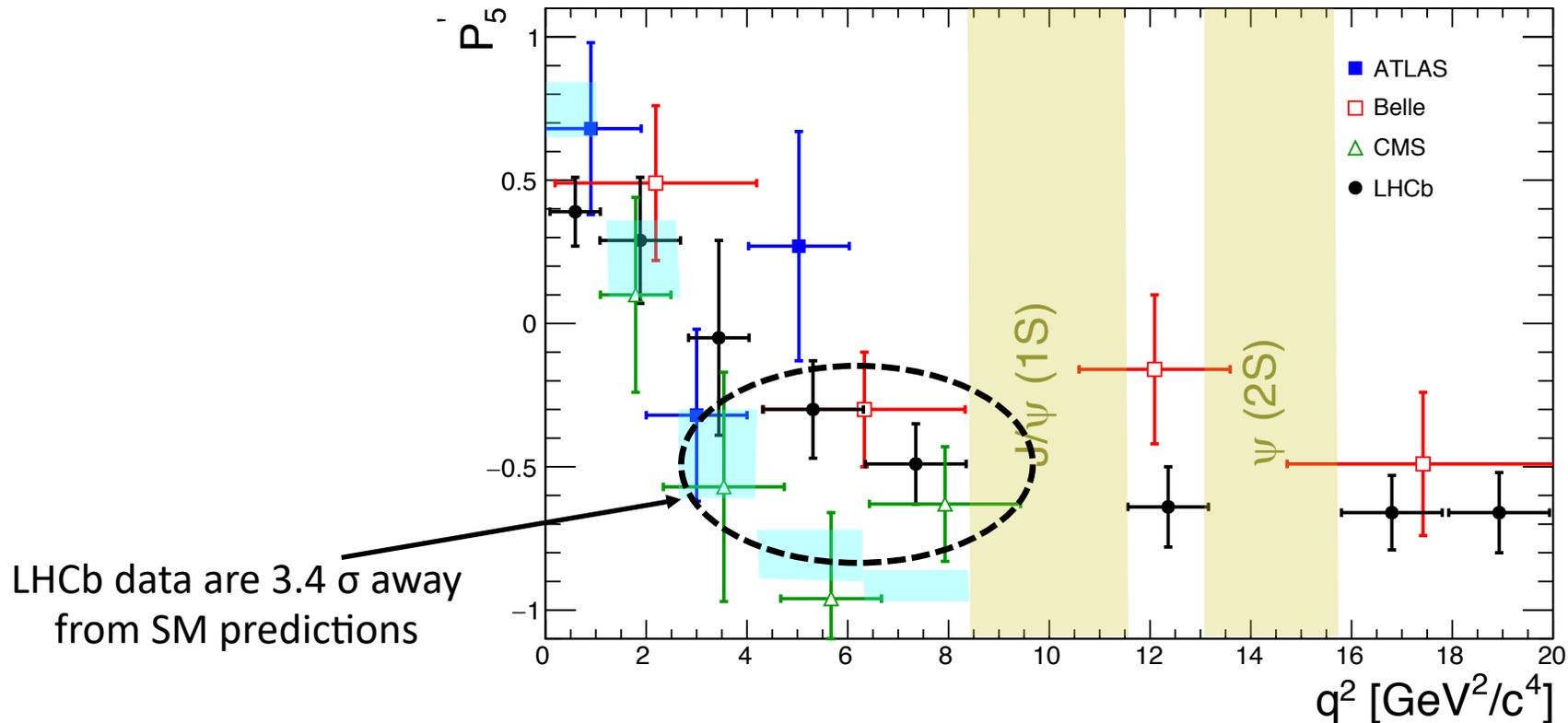
## Experimental results:

- BaBar, PRD 86, 032012 (2012)
- Belle, PRL 103, 171801 (2009)
- CDF, PRL 107, 201802 (2011)
- CMS, PLB 753, 424 (2016)
- LHCb, JHEP 06, 133 (2014)
- LHCb, JHEP 11, 047 (2016)
- LHCb, JHEP 09, 179 (2015)
- LHCb, JHEP 06, 115 (2015)

**Theory caveat:** exclusive decays  $B \rightarrow M l^+ l^-$  to a meson  $M$  require the knowledge of the  $B \rightarrow M$  form factors in the full kinematic range  $4m^2 < q^2 < (m_B - m_M)^2$ . Data today are more precise than theory.

# FCNC semileptonic $b \rightarrow s l^+ l^-$ decays : Angular Distributions

ATLAS, CMS, LHCb and Belle have performed a full angular analysis of the  $B_d \rightarrow K^* \mu\mu^-$  channel. While the large majority of the angular observables is consistent with SM expectations, the LHCb measurement of the  $P'_5$  variable shows hints of deviations from SM predictions in the region  $4 < q^2 < 8 \text{ GeV}^2$



## Theory predictions:

Descotes-Genon et al. JHEP 12, 125 (2014).

## Data:

LHCb, JHEP 02 104 (2016)

Belle, PRL 118, 11801 (2017)

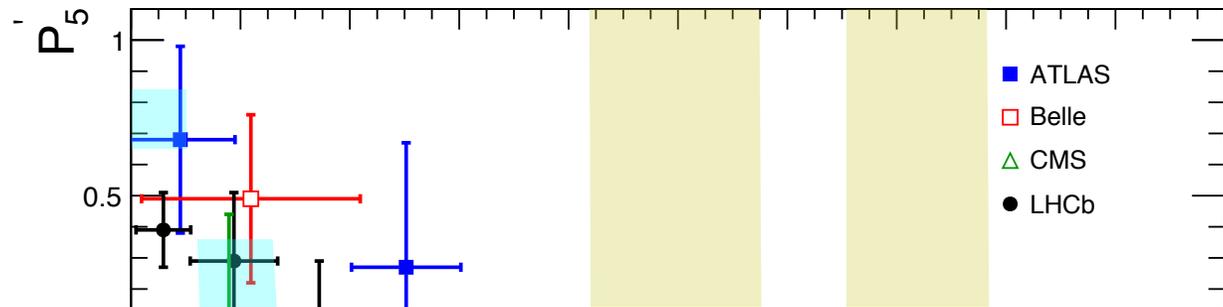
ATLAS-CONF-2017-023

CMS, PLB 81, 517 (2018)

The main challenge on the theory side is represented by soft gluon corrections to the charm loop, that have been estimated in LCSR but remain the largest source of uncertainty

# FCNC semileptonic $b \rightarrow s l^+ l^-$ decays : Angular Distributions

ATLAS, CMS, LHCb and Belle have performed a full angular analysis of the  $B_d \rightarrow K^* \mu\mu^-$  channel. While the large majority of the angular observables is consistent with SM expectations, the LHCb measurement of the  $P'_5$  variable shows hints of deviations from SM predictions in the region  $4 < q^2 < 8 \text{ GeV}^2$

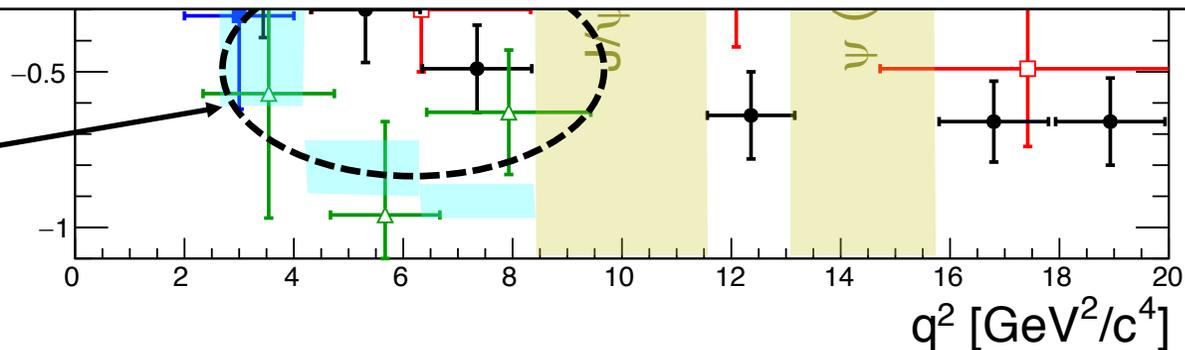


**Theory predictions:**  
Descotes-Genon et al. JHEP 12, 125 (2014).

**Data:**  
LHCb, JHEP 02 104 (2016)

**Clarifying the role of the charm-loops will remain a mandatory task before NP can be considered the responsible of the angular anomalies of the  $B_d \rightarrow K^* \mu\mu^-$  mode**

LHCb data are  $3.4 \sigma$  away from SM predictions



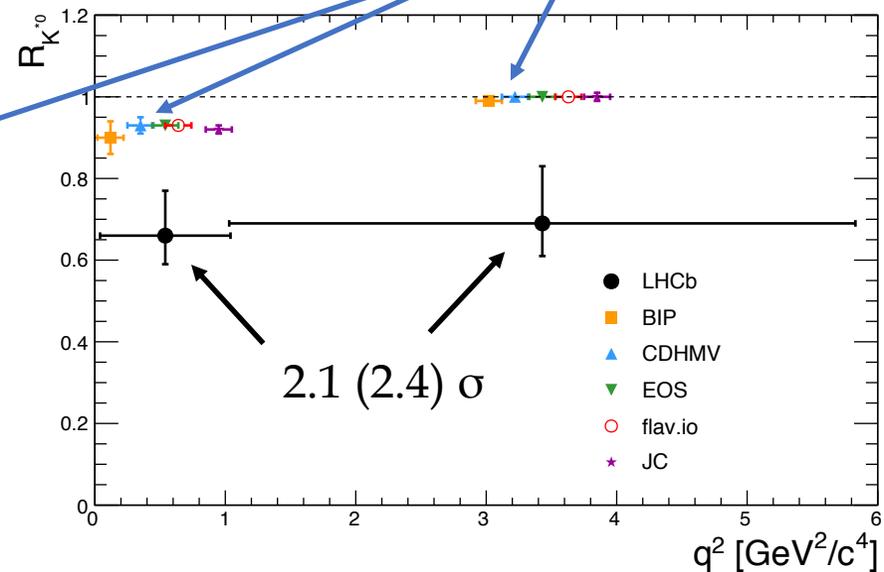
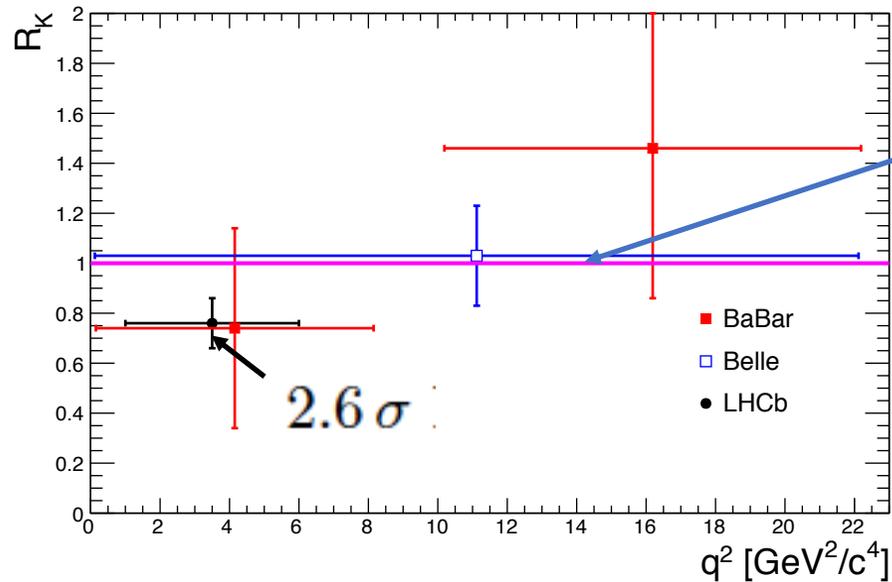
The main challenge on the theory side is represented by soft gluon corrections to the charm loop, that have been estimated in LCSR but remain the largest source of uncertainty

# Lepton Flavor Universality in FCNC

LFU violation in neutral currents has been measured in  $b \rightarrow sl+l-$  (with  $l = e, \mu$ ) transitions involving ratios of branching fractions:

$$R_{K^{(*)}} = \frac{BR(B \rightarrow K^{(*)} \mu\mu)}{BR(B \rightarrow K^{(*)} ee)} \quad q^2 \in [q_{min}^2, q_{max}^2]$$

SM values known at o(1%) level  
See Bordone, Isidori et al., EPJC 76 (2016) 440  
and References defined in arXiv:1705..05802



LHCb, PRL 113 (2014) 151601  
 BaBar, PRD 86, (2012) 032012  
 Belle, PRL 103 (2009) 171801

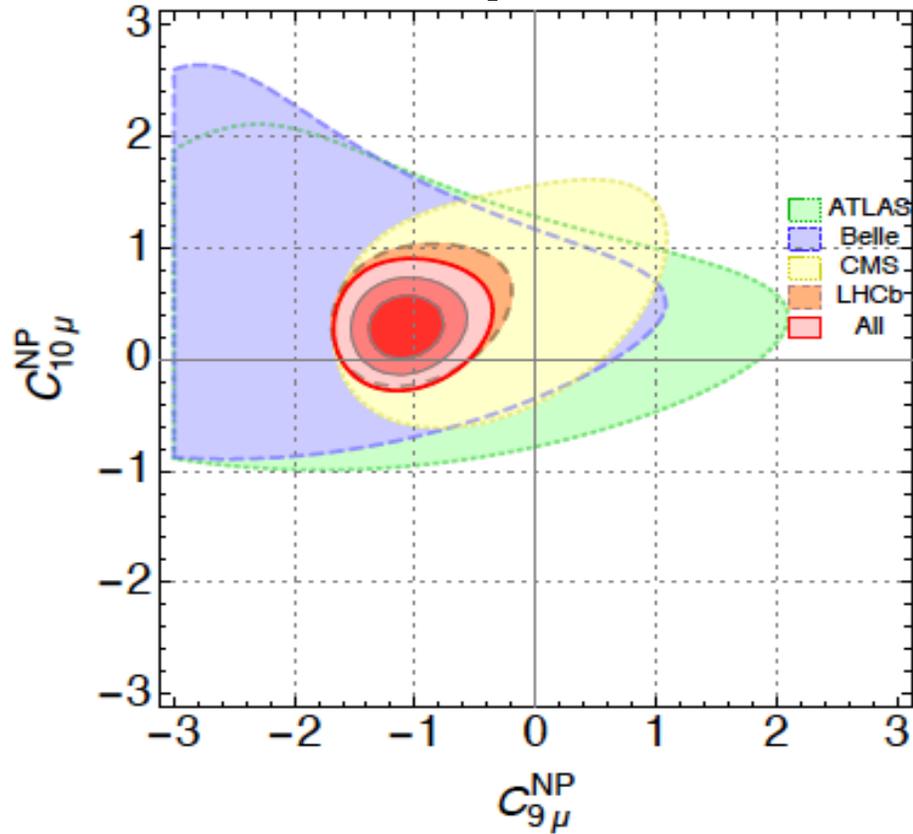
LHCb, JHEP 08 (2017) 055

# Model independent fits

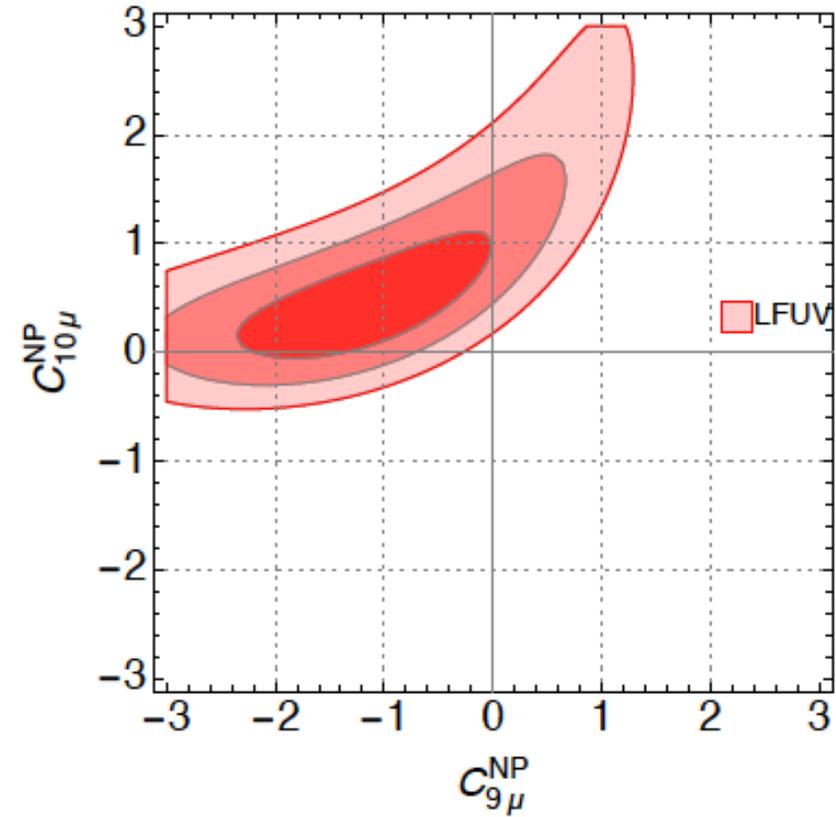
$C_9(\text{NP})$  deviates from 0 by  $>4\sigma$

Independent fits by many groups favour:  $C_9(\text{NP}) \sim -1$  or  $C_9(\text{NP}) = -C_{10}(\text{NP})$

Capdevila, Crivellin, Descotes-Genon, Matias 1704.05340



All  $b \rightarrow s l^+ l^-$  observables  
(decay rates, angular observables,  $R_K, R_{K^*}$ )



Only  $R_K, R_{K^*}$  from LHCb

More updated fit results in David Straub's talk

# NP explaining both $R_{D^{(*)}}$ and $R_{K^{(*)}}$ anomalies

S Fajfer, ICHEP 2018

$$R_{D^{(*)}}^{exp} > R_{D^{(*)}}^{SM}$$

$$\mathcal{L}_{NP} = \frac{1}{(\Lambda^D)^2} 2 \bar{c}_L \gamma_\mu b_L \bar{\tau} \gamma^\mu \nu_L$$

$$\Lambda^D \simeq 3 \text{ TeV}$$

$$\Lambda^D = \Lambda$$

NP in FCNC  $B \rightarrow K^{(*)} \mu^+ \mu^-$   
has to be suppressed

$$R_{K^{(*)}}^{exp} < R_{K^{(*)}}^{SM}$$

$$\mathcal{L}_{NP} = \frac{1}{(\Lambda^K)^2} \bar{s}_L \gamma_\mu b_L \bar{\mu}_L \gamma^\mu \mu_L$$

$$\Lambda^K \simeq 30 \text{ TeV}$$

If the scale is the same  $\Lambda^D \sim \Lambda^K$

$$\frac{1}{(\Lambda^K)^2} = \frac{C_K}{\Lambda^2} \quad C_K \simeq 0.01$$

suppression factor

# Models at the TeV scale explaining both $R_{D^{(*)}}$ and $R_{K^{(*)}}$ anomalies

S Fajfer, ICHEP 2018

## Scalar LQ as pseudo-Nambu-Goldstone boson

Gripaios et al, 1010.3962,  
Gripaios et al., 1412.1791,  
Marzocca 1803.10972...

## Models with scalar LQs

Hiller & Schmaltz, 1408.1627,  
Becirevic et al. 1608.08501, SF and Kosnik,  
1511.06024, Becirevic et al., 1503.09024,  
Dorsner et al, 1706.07779,  
Cox et al., 1612.03923,  
Crivellin et al.,1703.09226...

## $W'$ , $Z'$ in warped space

Megias et al.,1707.08014

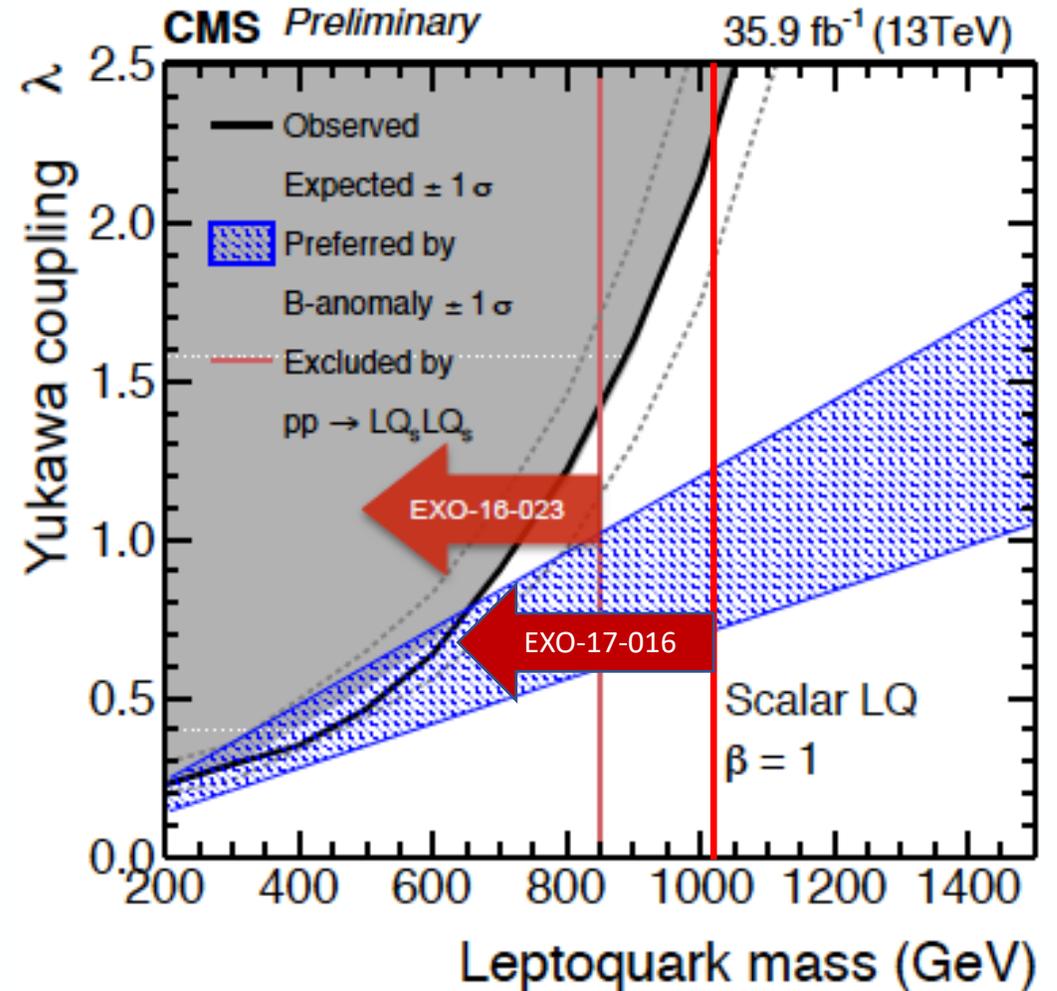
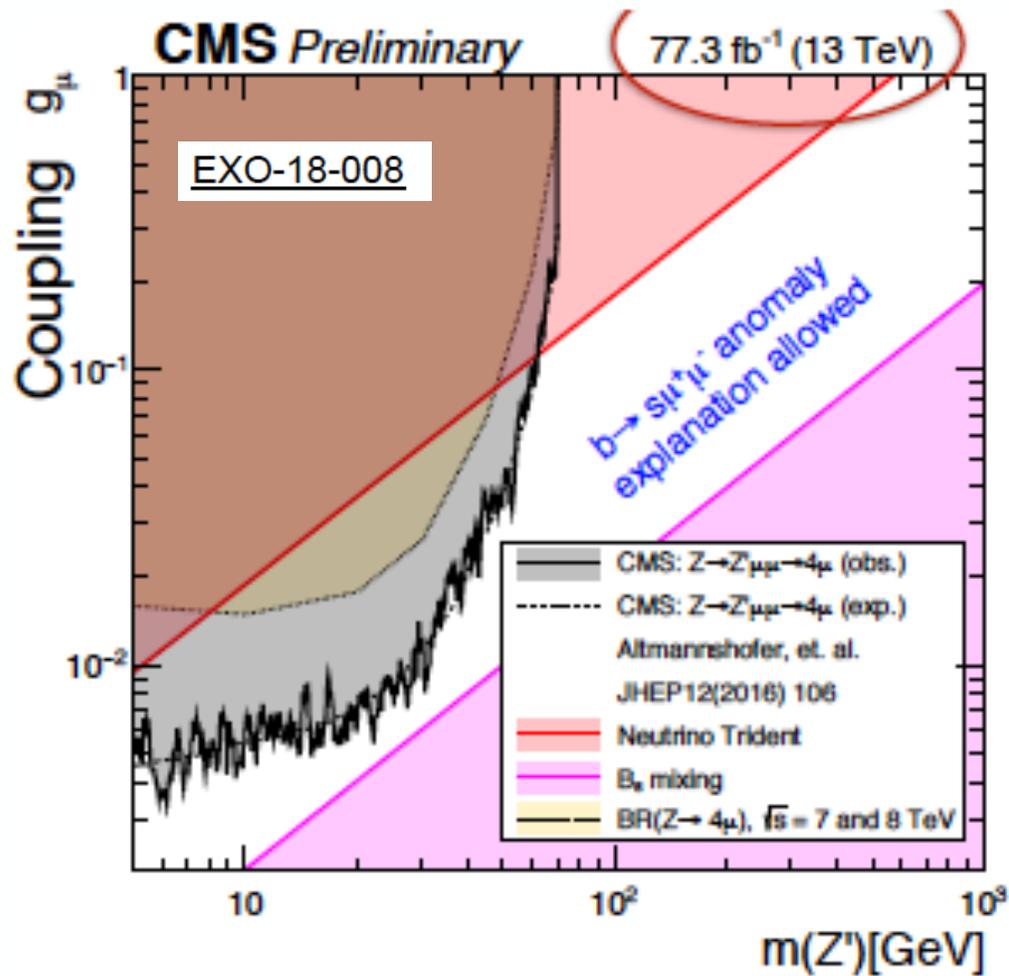
## Vector resonances (from techni-fermions)

Barbieri et al.,1506.09201, Buttazzo et al.  
1604.03940,  
Barbieri et al., 1611.04930  
Blanke & Crivellin, 1801.07256,...

## Gauge bosons

Greljo et al., 1804.04642  
Cline, Camalich, 1706.08510  
Calibbi et al.,1709.00692  
Assad et al., 1708.06350  
Di Luzio et al.,1708.08450  
Bordone et al.,1712.01368, 1805.09328...

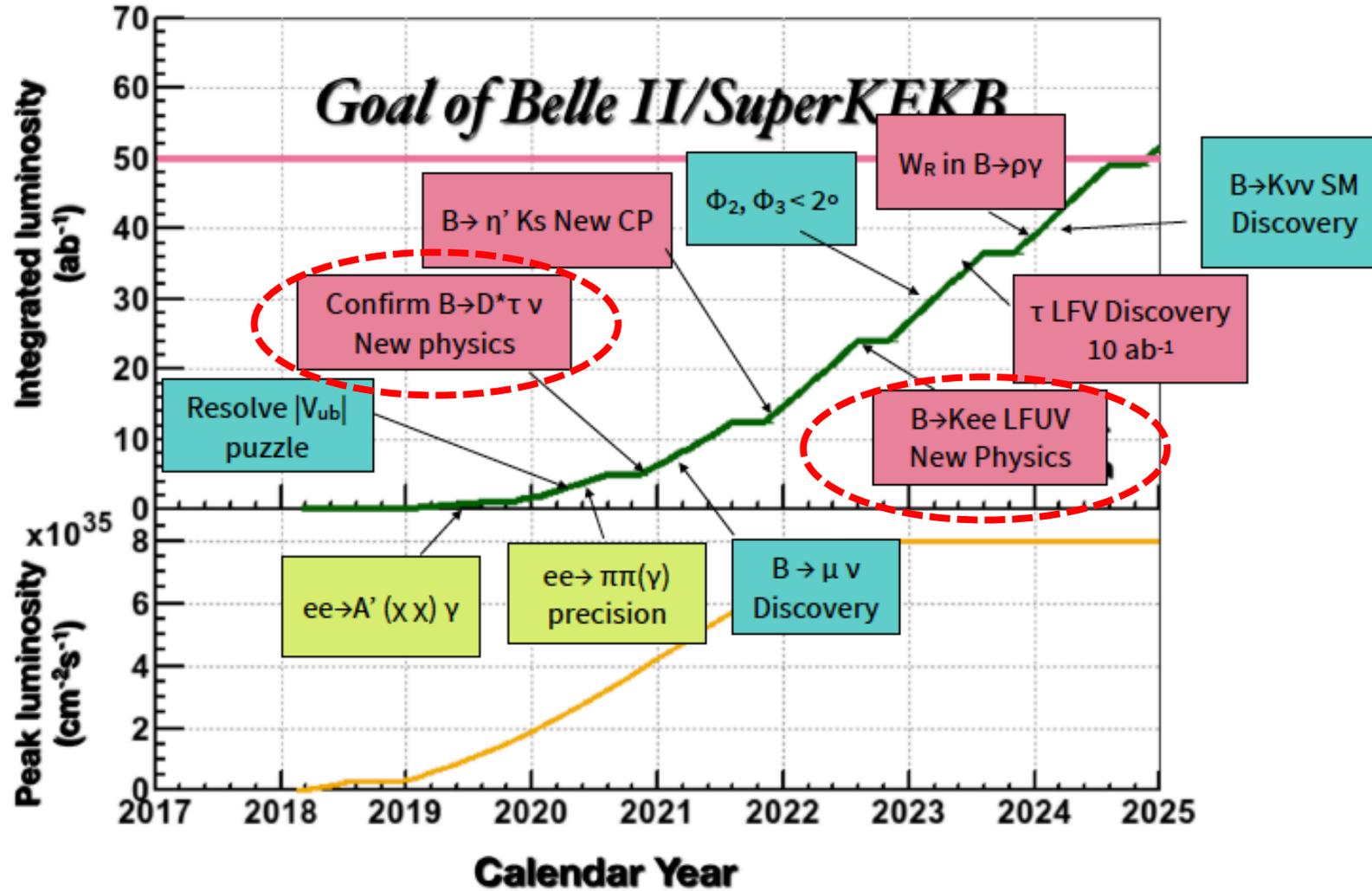
# Direct searches for $Z'$ and scalar Lepto-Quarks



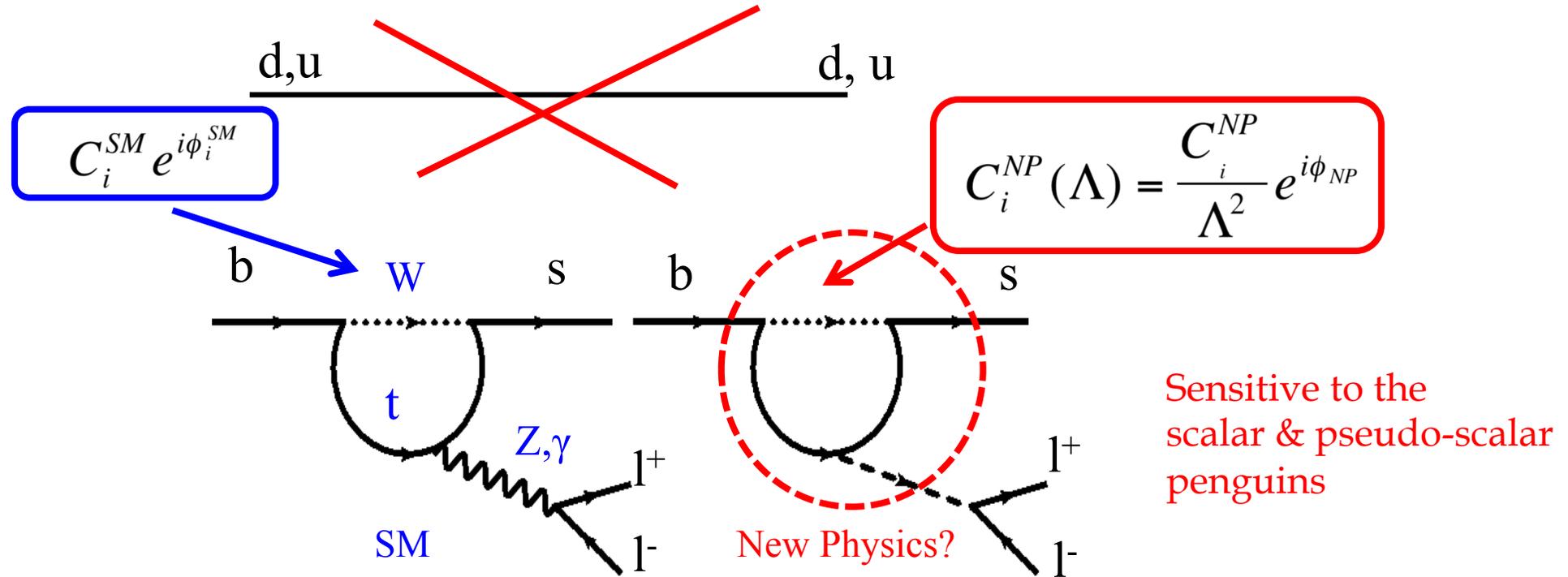
ATLAS and CMS are exploring the region favoured by the LHCb anomalies  
(sensitivity so far driven by CMS results)

...and Belle-II in a few years will help to clarify these patterns:

P. Urquijo, ICHEP 2018

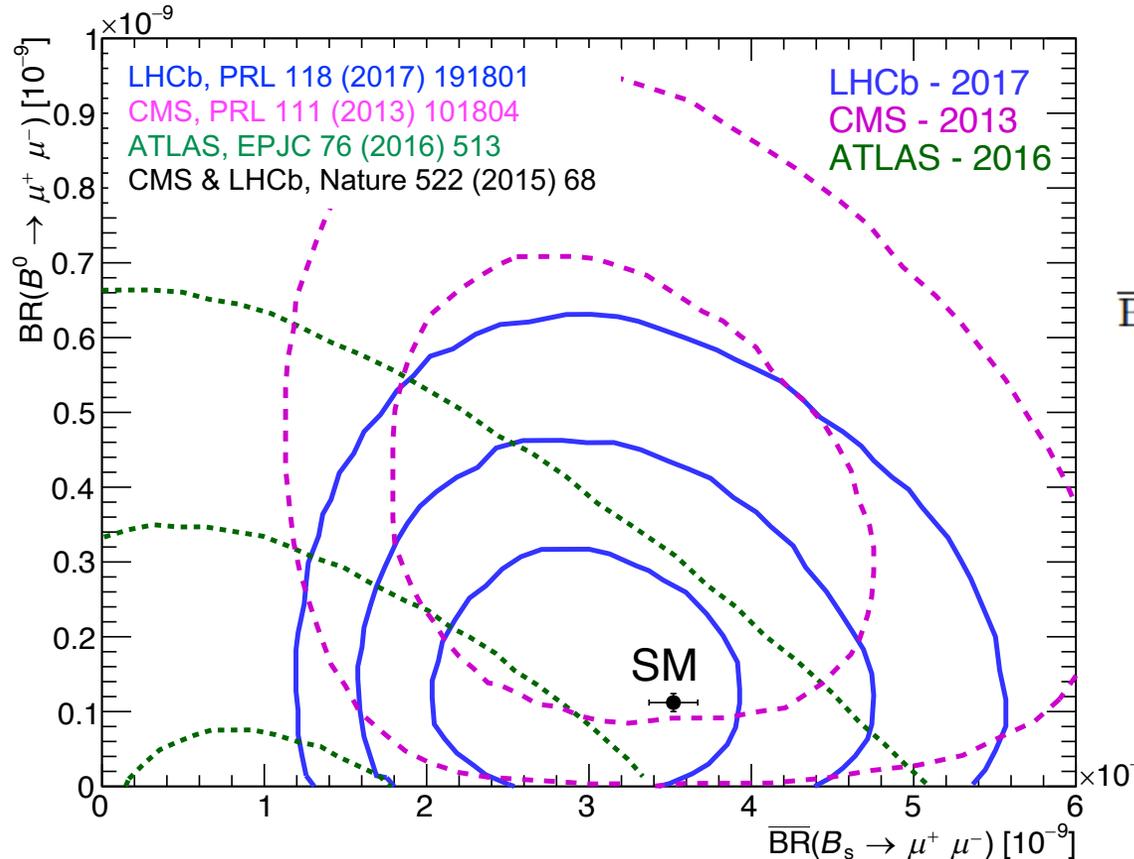


Back to the main diagram:



You can recognize here the main diagram that drives the  $B^0_{(s)} \rightarrow \mu^+ \mu^-$  decays...

# The $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$ and $\text{BR}(B^0 \rightarrow \mu^+ \mu^-)$



## Experimental Values:

$$\overline{\text{BR}}(B_s^0 \rightarrow \mu^+ \mu^-)_{\text{CMS}} = (3.0_{-0.9}^{+1.0}) \times 10^{-9},$$

$$\overline{\text{BR}}(B_s^0 \rightarrow \mu^+ \mu^-)_{\text{LHCb 2017}} = (3.0 \pm 0.6_{-0.2}^{+0.3}) \times 10^{-9},$$

$$\overline{\text{BR}}(B_s^0 \rightarrow \mu^+ \mu^-)_{\text{ATLAS}} = (0.9_{-0.8}^{+1.1}) \times 10^{-9}$$

$$\text{BR}(B^0 \rightarrow \mu^+ \mu^-) < 3.4 \times 10^{-10} \text{ @ 95\% CL (LHCb)}$$

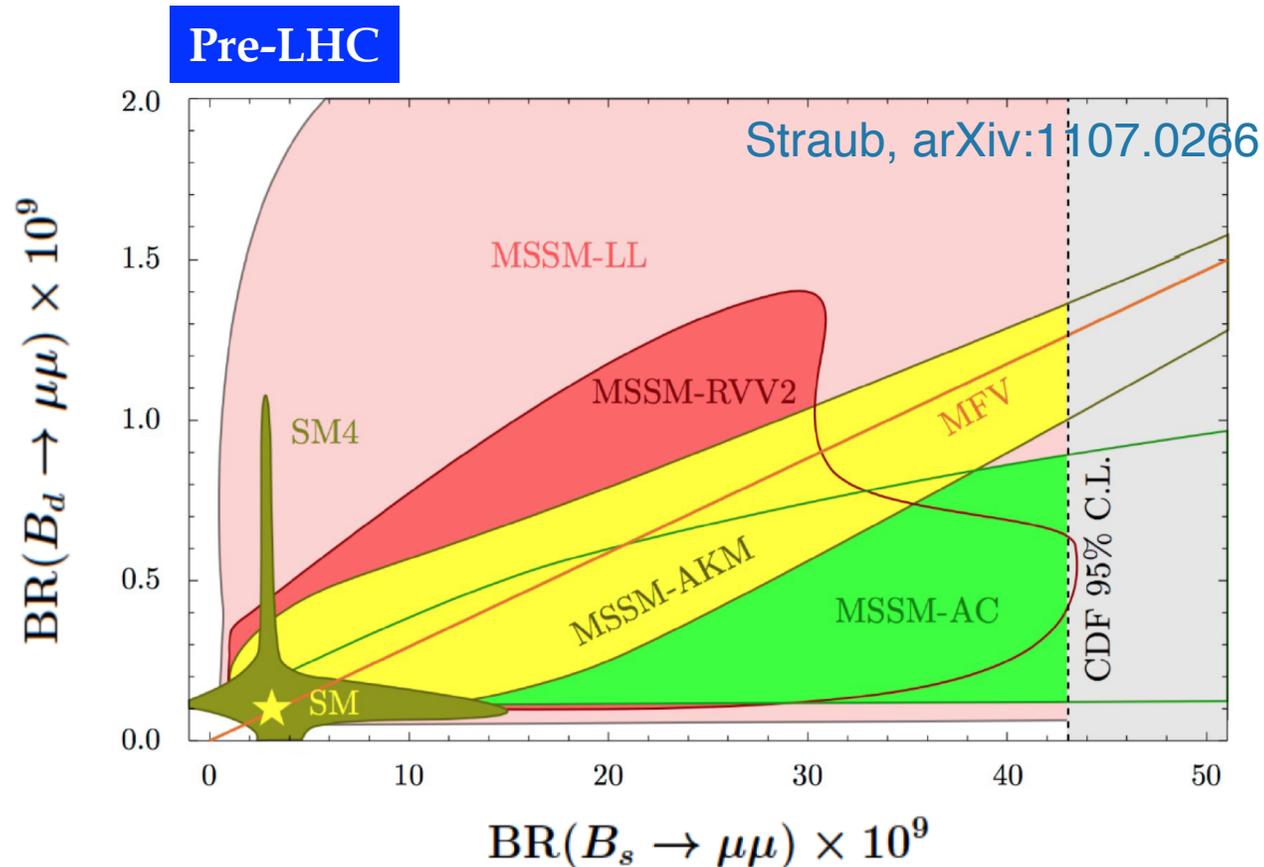
## Theory predictions:

$$\text{BR}(B_s^0 \rightarrow \mu^+ \mu^-) = (3.66 \pm 0.23) \times 10^{-9}$$

$$\text{BR}(B^0 \rightarrow \mu^+ \mu^-) = (1.06 \pm 0.09) \times 10^{-10}$$

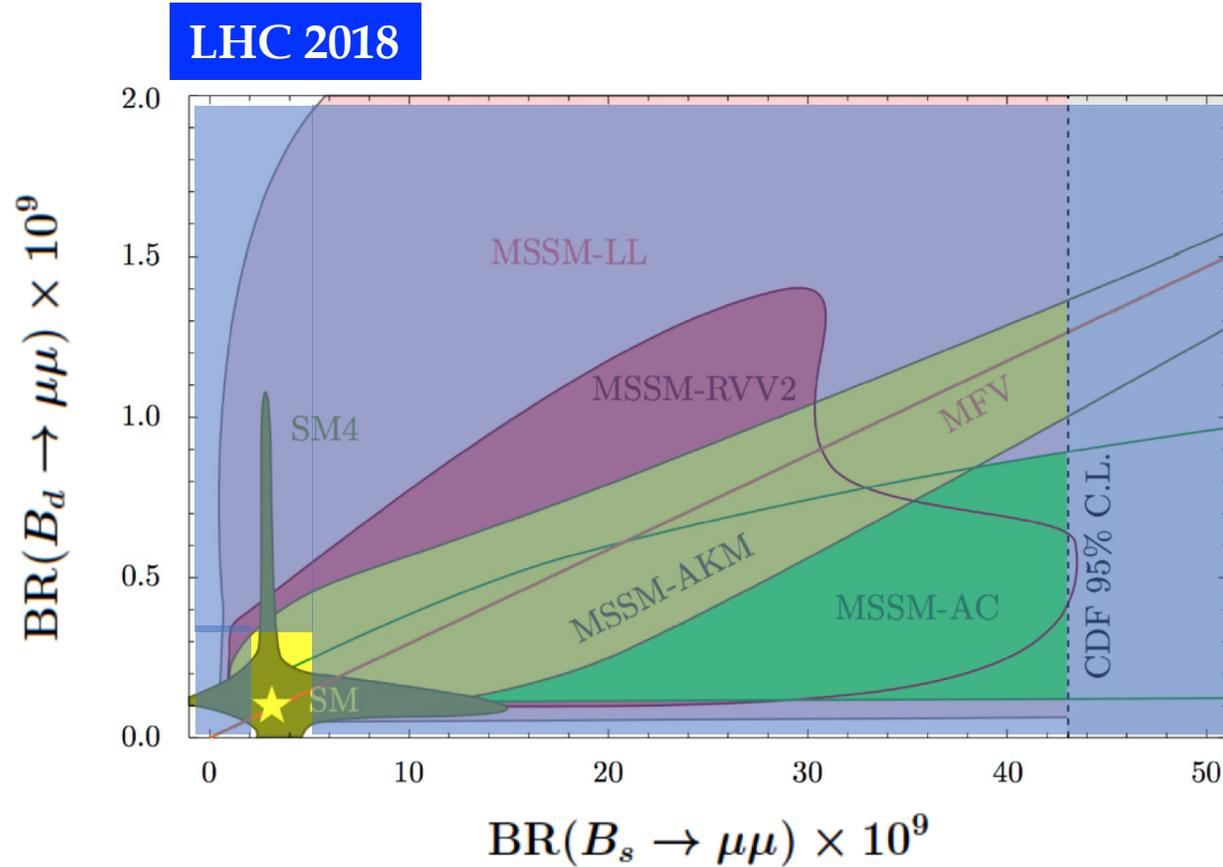
Excellent agreement between measurements and SM predictions

# The $BR(B_s \rightarrow \mu^+ \mu^-)$ and $BR(B^0 \rightarrow \mu^+ \mu^-)$ and New Physics



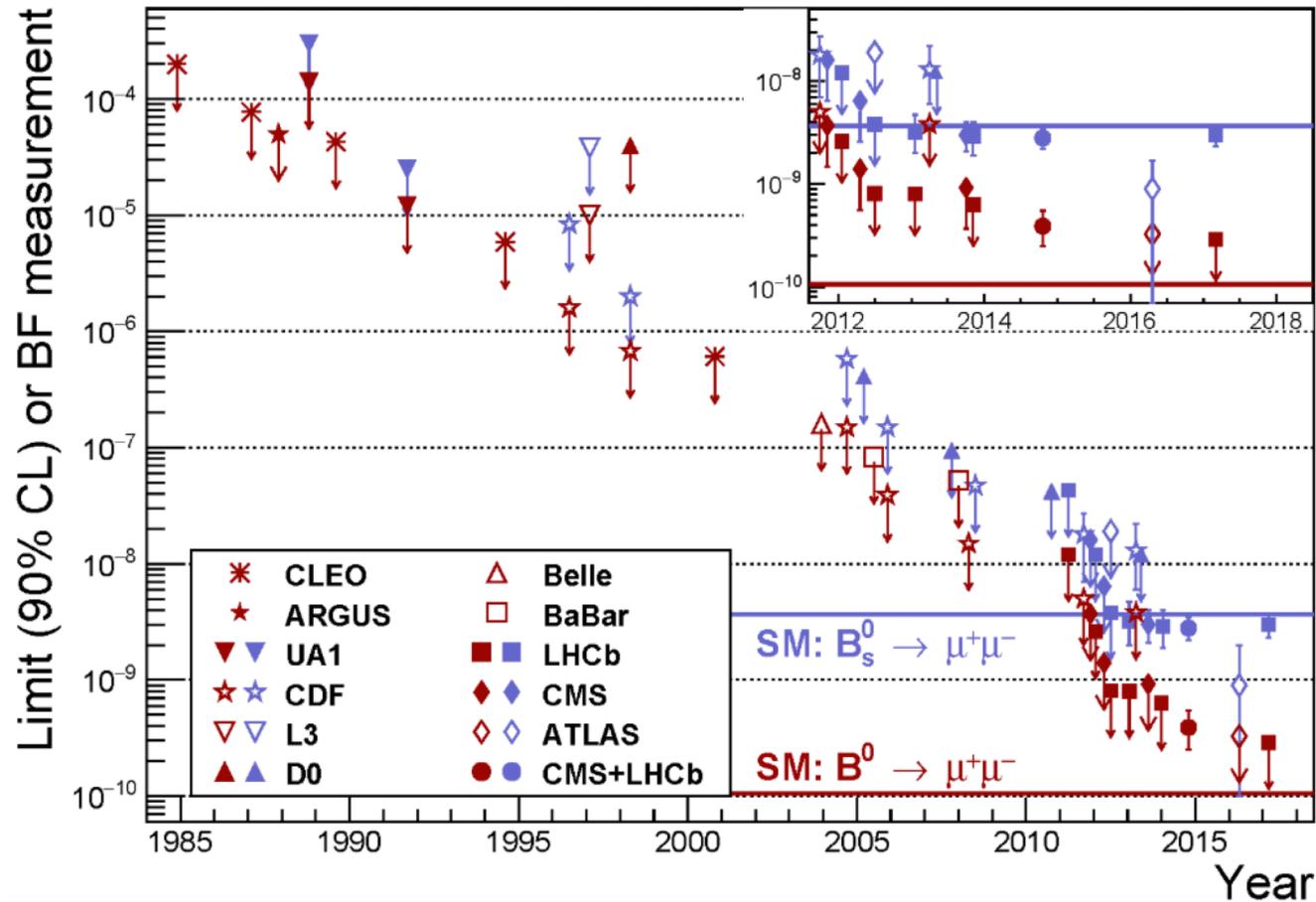
NP contributions arising from scalar or pseudo-scalar operators could lift the helicity suppression possibly enhancing the value of the branching fractions. These modes are particularly sensitive to models with an extended Higgs sector, as in MSSM with two Higgs doublets or in supersymmetric models with non-universal Higgs masses, etc (infinite literature on this topic).

# The $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$ and $\text{BR}(B^0 \rightarrow \mu^+ \mu^-)$ and New Physics



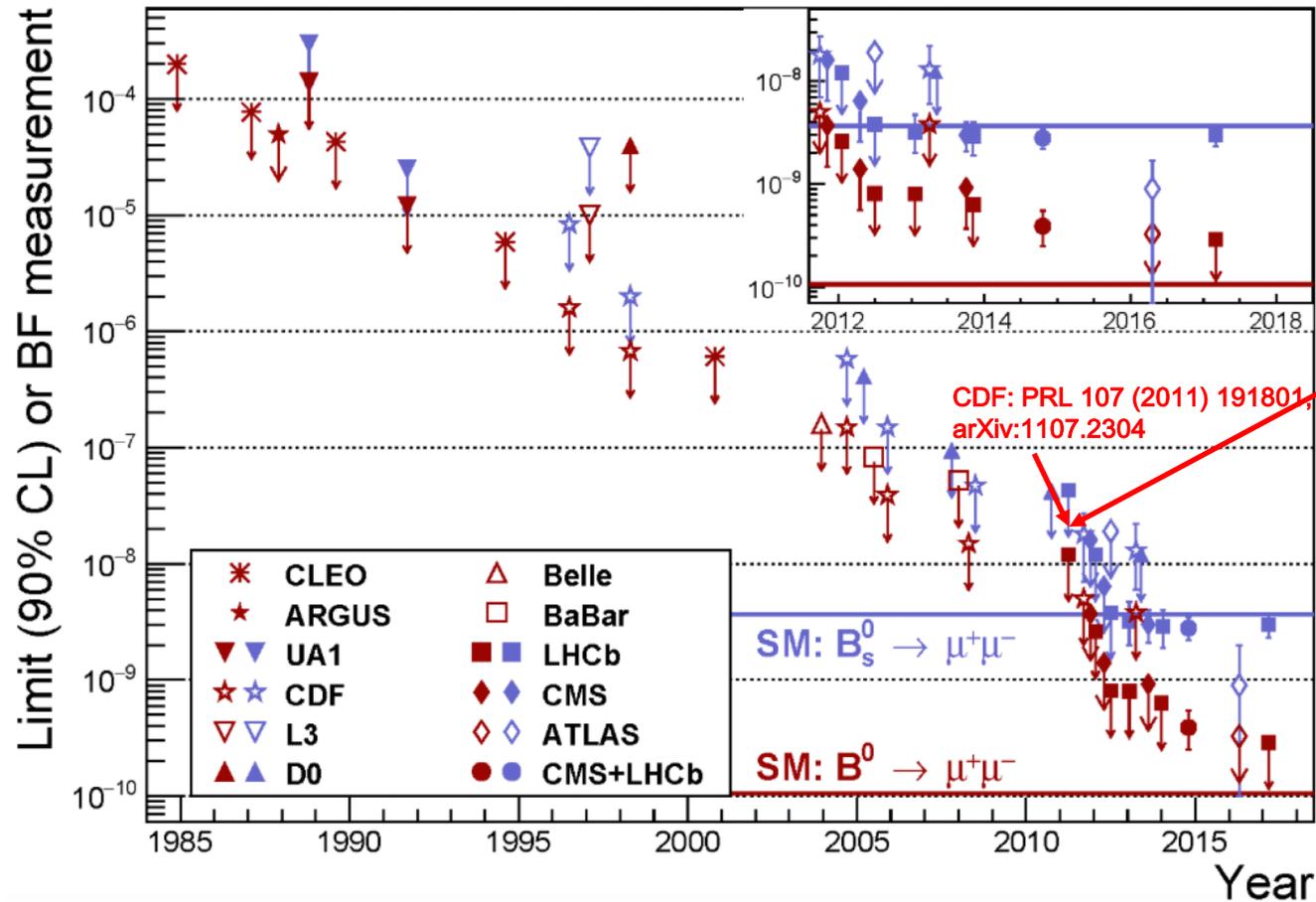
A killer for BSM theories with extended Higgs sector.

# The $\text{BR}(B_s \rightarrow \mu^+\mu^-)$ and $\text{BR}(B^0 \rightarrow \mu^+\mu^-)$ : a story 30 years long



*.. but was not always plain sailing...*

# The $\text{BR}(B_s \rightarrow \mu^+\mu^-)$ and $\text{BR}(B^0 \rightarrow \mu^+\mu^-)$ : a story 30 years long

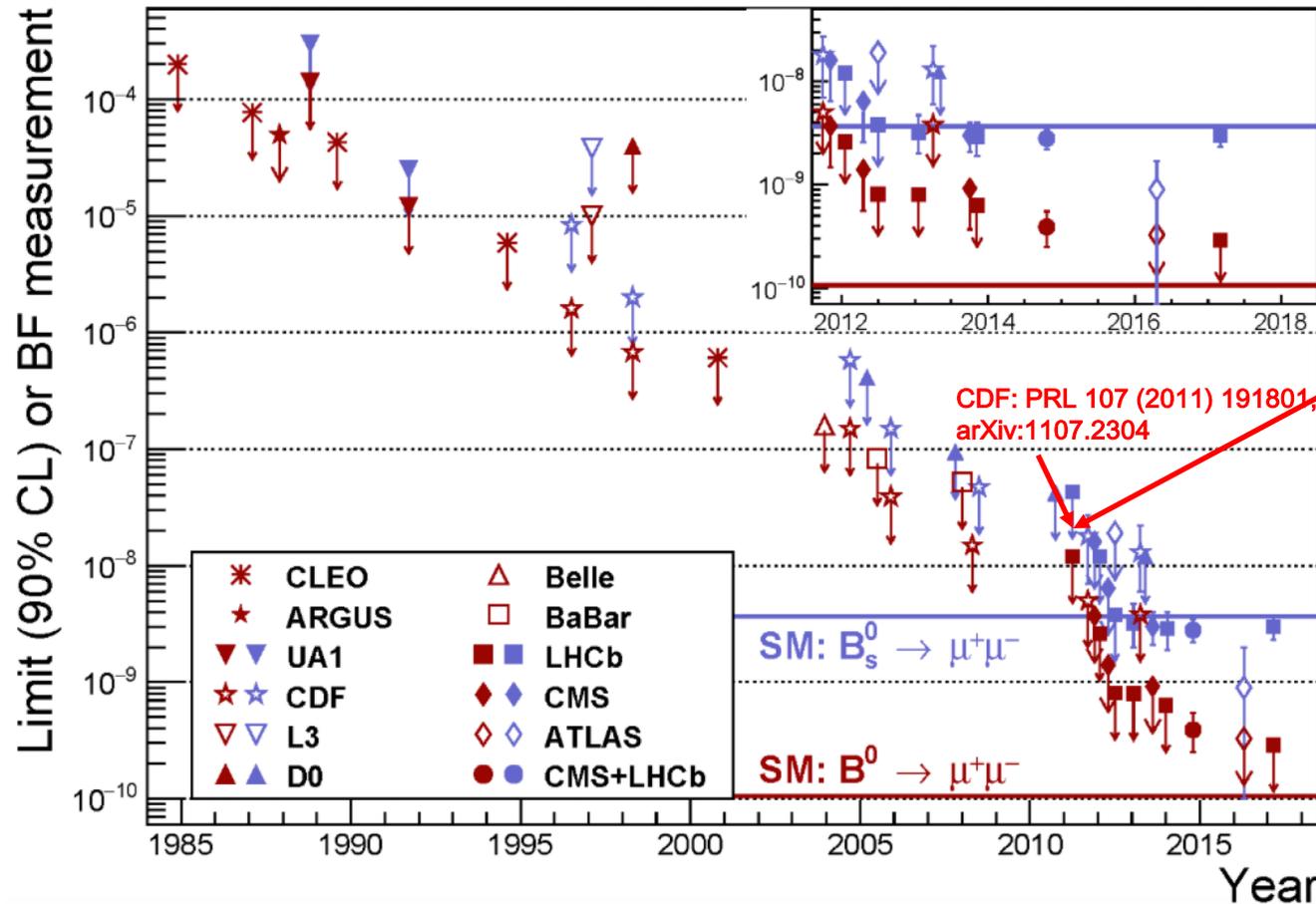


**July 2011:**

CDF publishes the  $3\sigma$  evidence for:

$$\text{BR}(B_s \rightarrow \mu^+\mu^-) = (1.8^{+1.1}_{-0.9}) 10^{-8} \sim 5 \text{ BR}(\text{SM})$$

# The BR( $B_s \rightarrow \mu^+ \mu^-$ ) and BR( $B^0 \rightarrow \mu^+ \mu^-$ ): a story 30 years long



**July 2011:**

CDF publishes the  $3\sigma$  evidence for:

$$BR(B_s \rightarrow \mu^+ \mu^-) = (1.8^{+1.1}_{-0.9}) 10^{-8} \sim 5 BR(SM)$$

**EPS 2011 (few weeks later):**

Guy Wilkinson presents at EPS (Grenoble) the first LHCb + CMS combination:

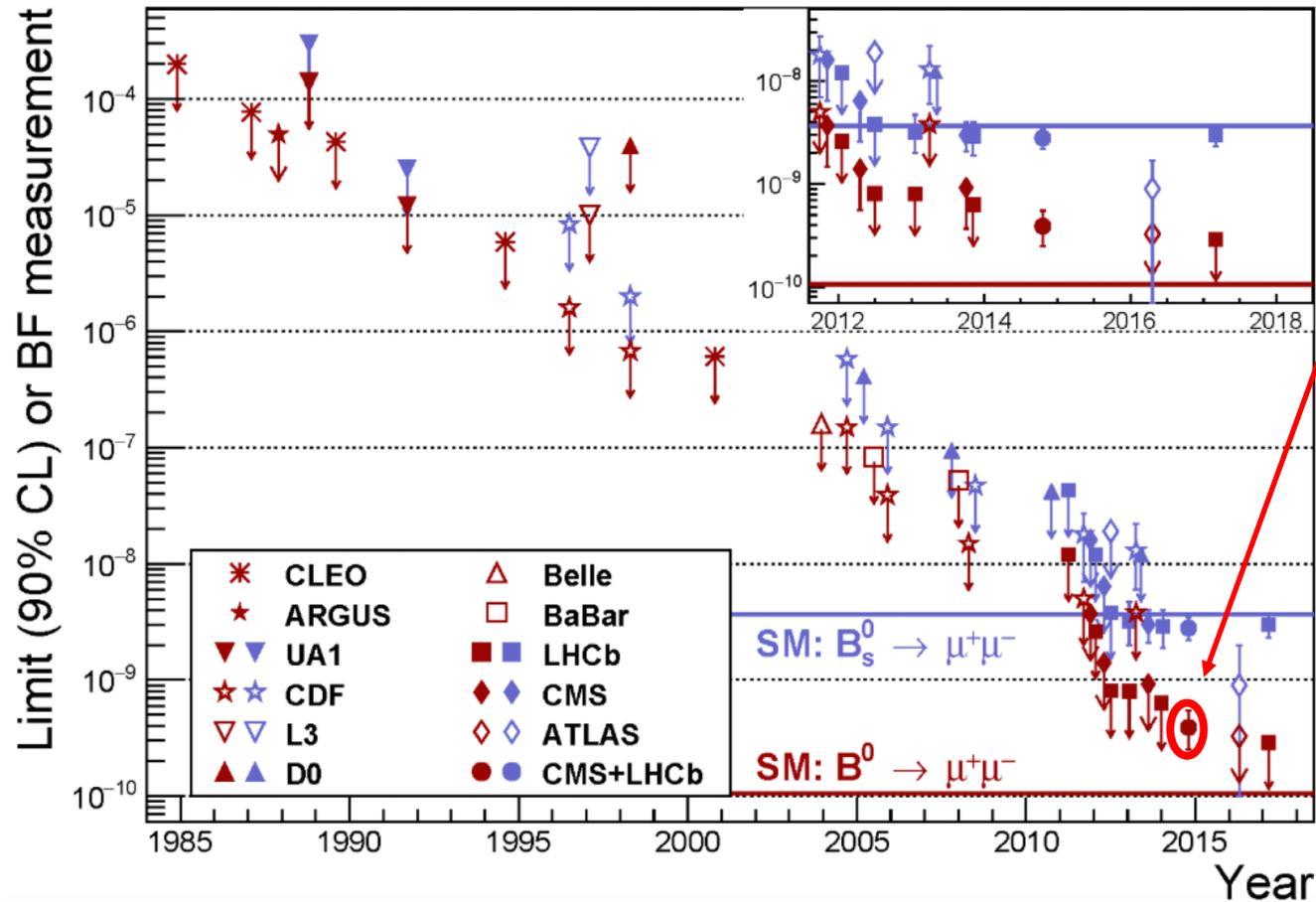
$$BR(B_s \rightarrow \mu^+ \mu^-) < 1.1 \cdot 10^{-8} @ 95\% CL$$

and excludes the CDF central value at 99.7% CL

*...Anomalies come and go...*

(or about the importance of having independent measurements from independent experiments)

# The $\text{BR}(B_s \rightarrow \mu^+\mu^-)$ and $\text{BR}(B^0 \rightarrow \mu^+\mu^-)$ : a story 30 years long



**2015:**

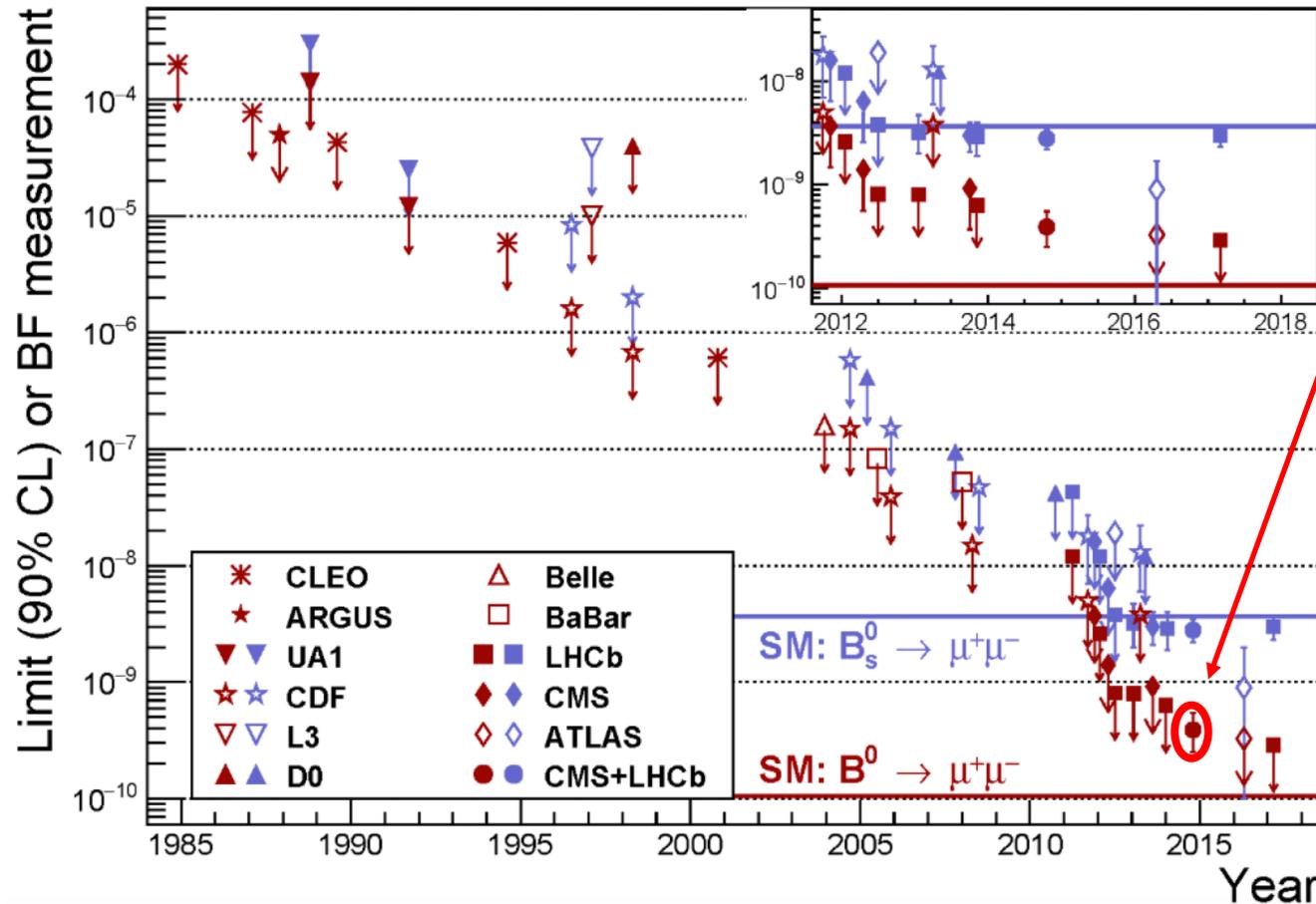
LHCb and CMS see  $2\sigma$  excess each  
in the decay  $B^0 \rightarrow \mu^+\mu^-$ .

LHCb+CMS combination, Nature 522 (2015) 68:

$B^0 \rightarrow \mu^+\mu^-$  observed at  $3\sigma$

$\text{BR}(B^0 \rightarrow \mu^+\mu^-) = (3.9^{+1.6}_{-1.4}) \times 10^{-10} \sim 4 \text{ BR(SM)}$

# The $BR(B_s \rightarrow \mu^+ \mu^-)$ and $BR(B^0 \rightarrow \mu^+ \mu^-)$ : a story 30 years long



**2015:**

LHCb and CMS see  $2\sigma$  excess each in the decay  $B^0 \rightarrow \mu^+ \mu^-$ .

LHCb+CMS combination, Nature 522 (2015) 68:

$B^0 \rightarrow \mu^+ \mu^-$  observed at  $3\sigma$

$BR(B^0 \rightarrow \mu^+ \mu^-) = (3.9^{+1.6}_{-1.4}) \times 10^{-10} \sim 4 BR(\text{SM})$

**2017:**

LHCb excludes this value at  $> 95\%$  CL:

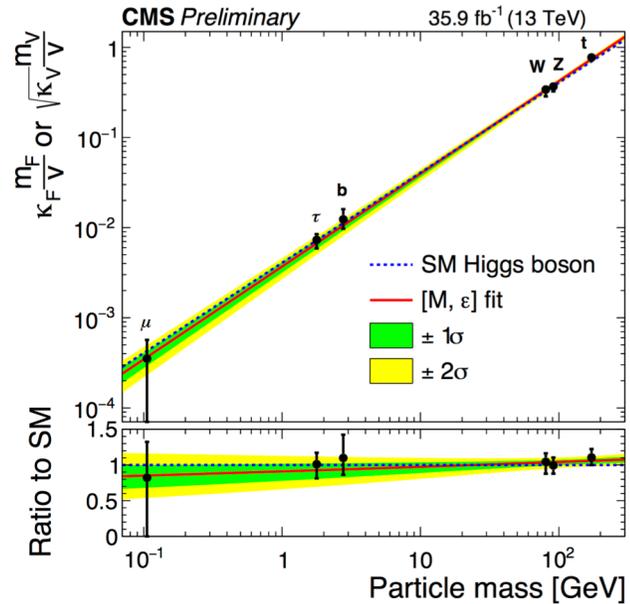
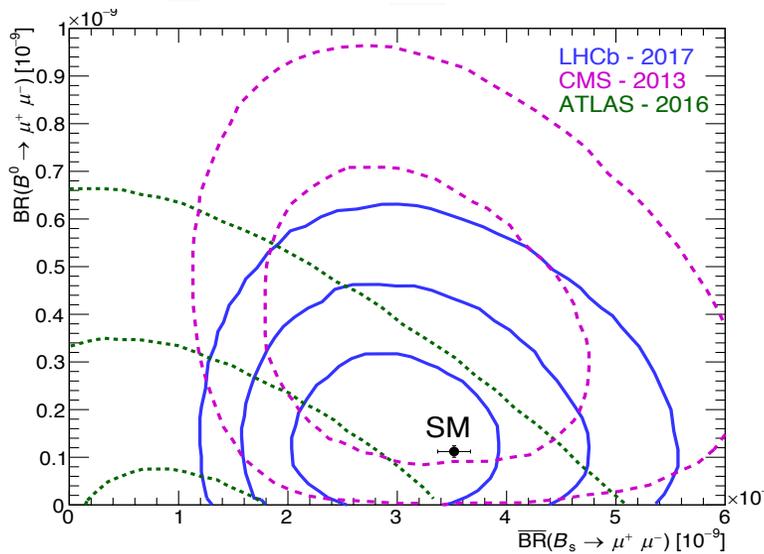
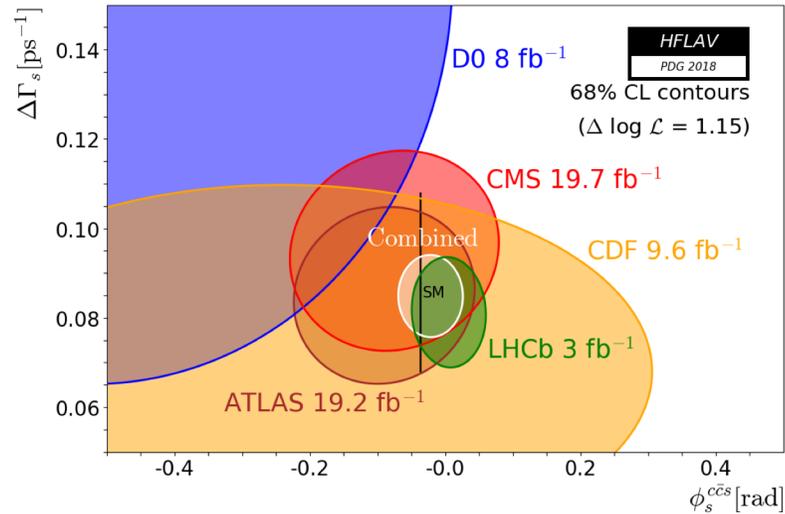
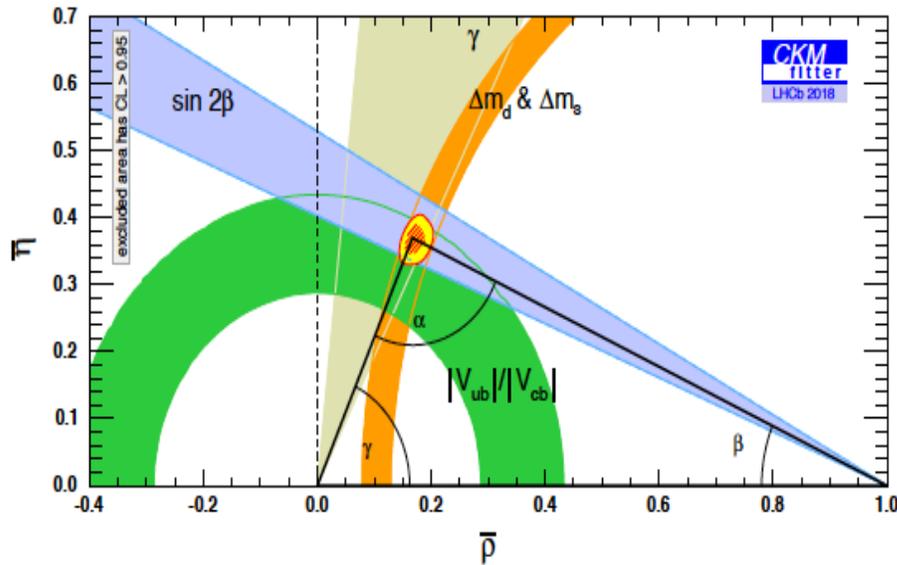
PRL 118 (2017) 191801

$BR(B^0 \rightarrow \mu^+ \mu^-) < 3.4 \times 10^{-10}$  @ 95% CL

*...Anomalies come and go ...*

*(or about how the combination of  $2\sigma$  effects can produce a (fake) hint of New Physics....)*

# The (excessive) Success of the Standard Model



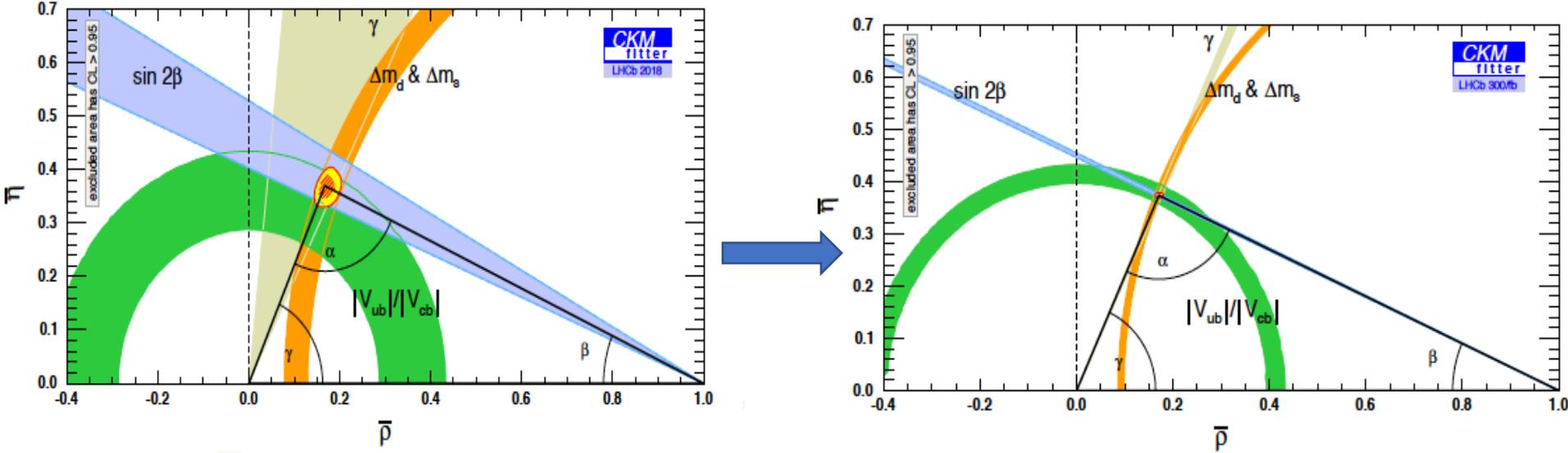
Where do we go from here ?

*“While the absence of NP appears as a paradox to us, still the picture repeatedly suggested by the data in the last 20 years is simple and clear: **the SM, extended to include some form of Dark Matter and Majorana’s neutrinos**, which can explain the active neutrino masses and oscillations via the see-saw mechanism and the baryogenesis through leptogenesis, can be valid up to some very high energy, possibly up to the Planck scale.”*



*Guido Altarelli,  
Proceedings of Vulcano Workshop 2014  
Frontier Objects in Astrophysics and Particle Physics*

Where do we go from here ?



A viable possibility... but is it the only one?

## Where do we go from here ?

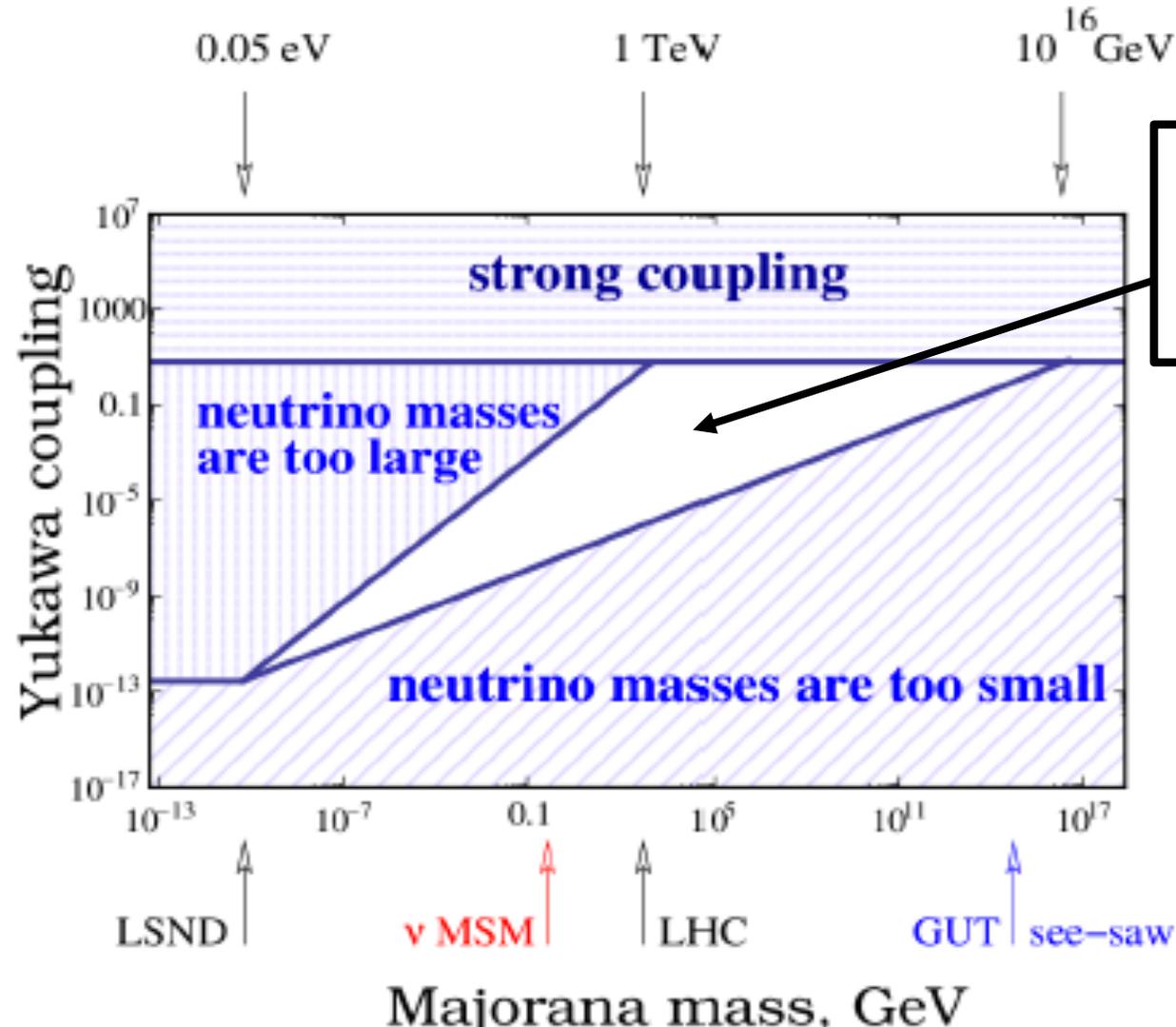
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- So far the experimental efforts in the accelerators' domain have been concentrated on the discovery of NP with masses at the TeV scale (or above) and sizeable couplings to SM particles.
- We did not observe so far unambiguous deviations from SM predictions, hence:
  - either NP is very heavy and/or it mimics the SM in its flavor-breaking pattern
  - **or it is below the Fermi scale and couples very feebly (hence they are very long-lived) to SM particles and so far escaped detection.**

**An attractive possibility and a change of paradigm in the accelerators' domain.**

# Heavy neutrinos as explanation of the neutrino masses and oscillations

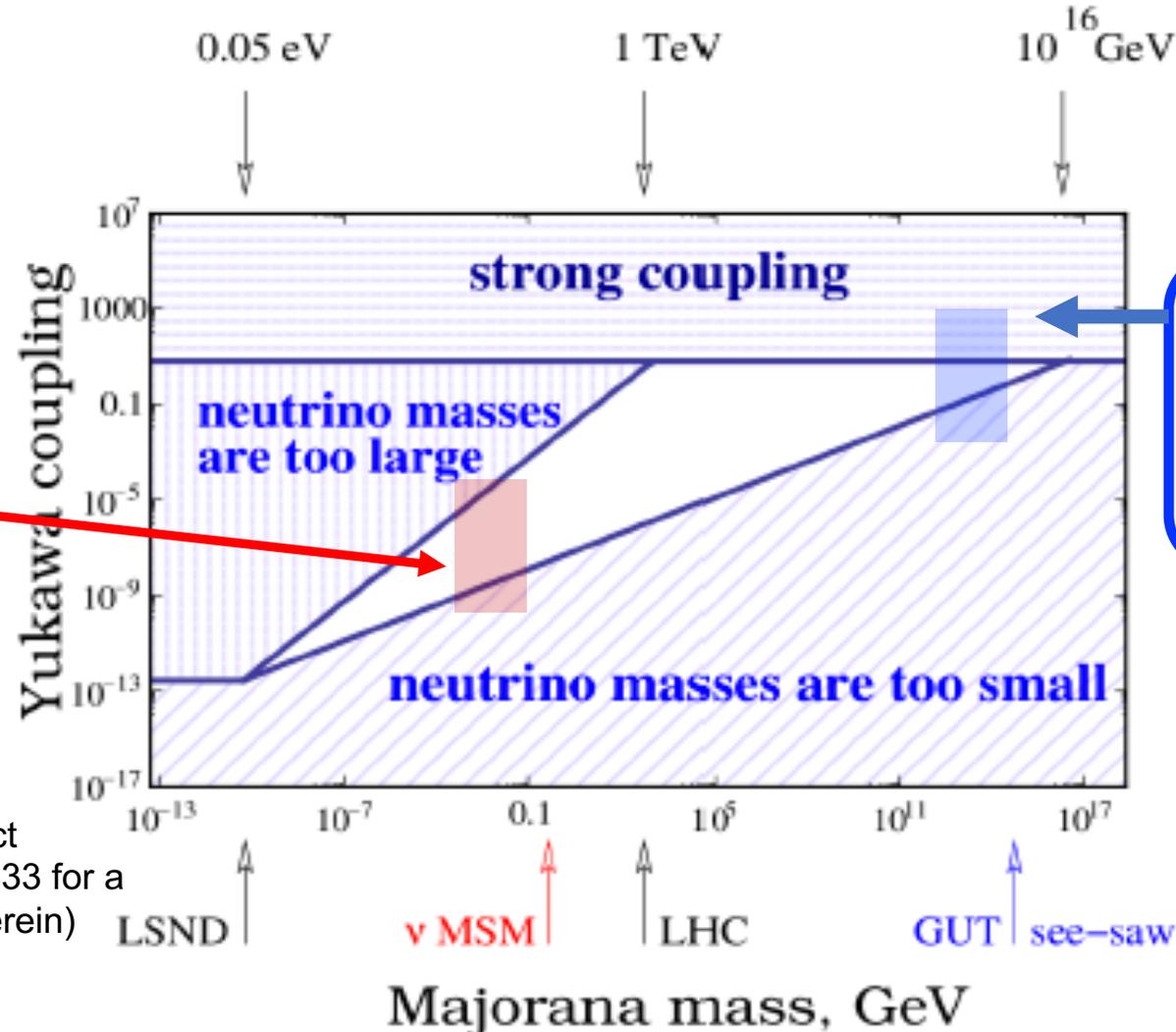
See-saw mechanism with RH neutrinos with Yukawa couplings to the Higgs and SM leptons. RH neutrinos can have masses from  $10^{-9}$  to  $10^{15}$  GeV.



Any number in the white area works to be consistent with the observed pattern of neutrino masses and oscillations

# Heavy neutrinos as explanation of the neutrino masses and oscillations

See-saw mechanism with RH neutrinos with Yukawa couplings to the Higgs and SM leptons. RH neutrinos can have masses from  $10^{-9}$  to  $10^{15}$  GeV.



## Alternative choice: EW see-saw ( $\nu$ MSM)

It is "natural" to assume that the masses of the RH neutrinos are at EW scale

## Heavy neutrinos below the Fermi scale

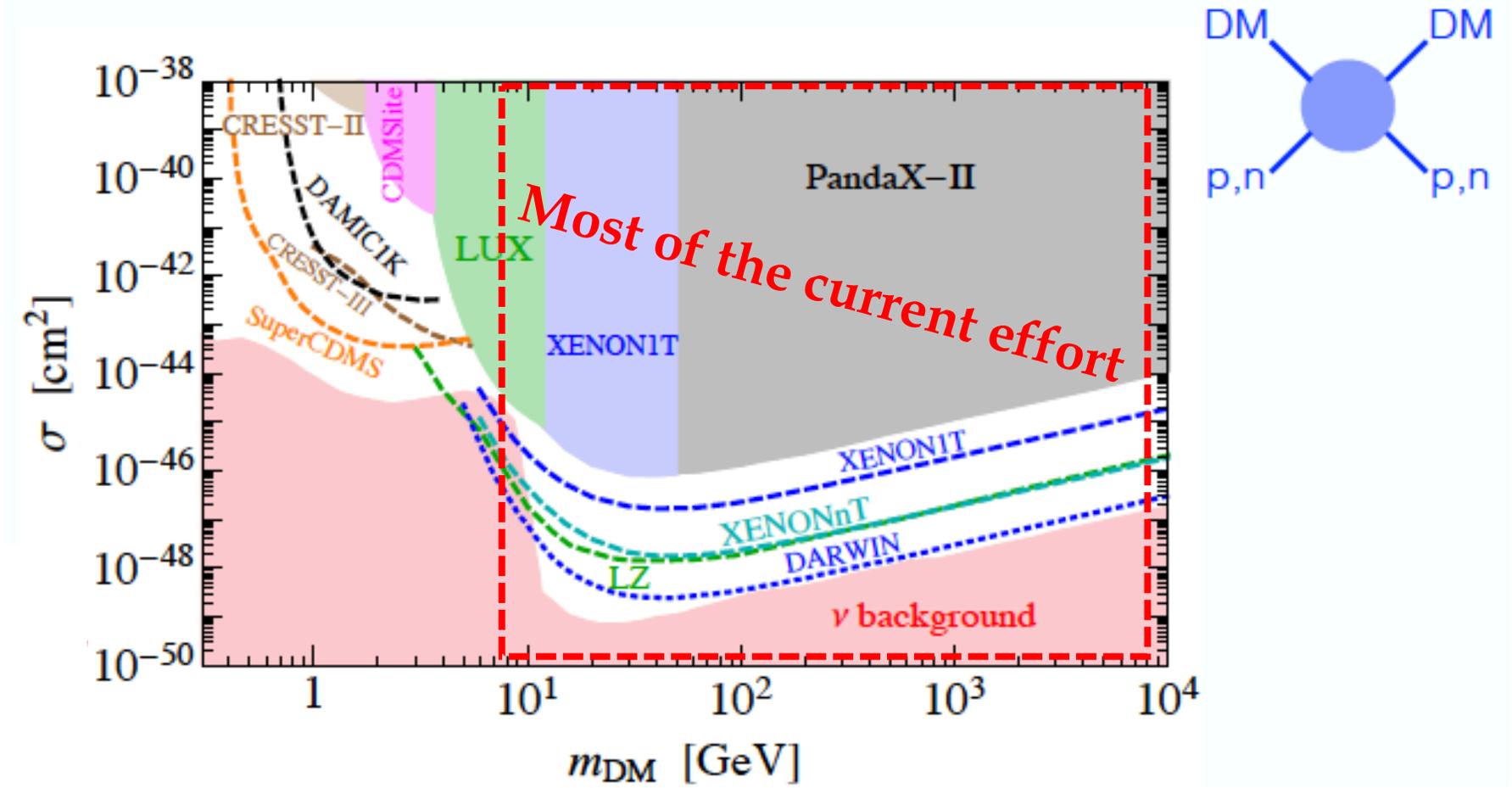
Infinite literature on the subject  
(see Shaposhnikov et al., 1808.10833 for a recent study and references therein)

## Popular choice: GUT see-saw

It "natural" to assume that Yukawa couplings of the RH neutrinos are similar to SM Yukawa.

# Search for Dark Matter with thermal origin

DM candidates with thermal origin can have mass between 10 keV and 100 TeV.



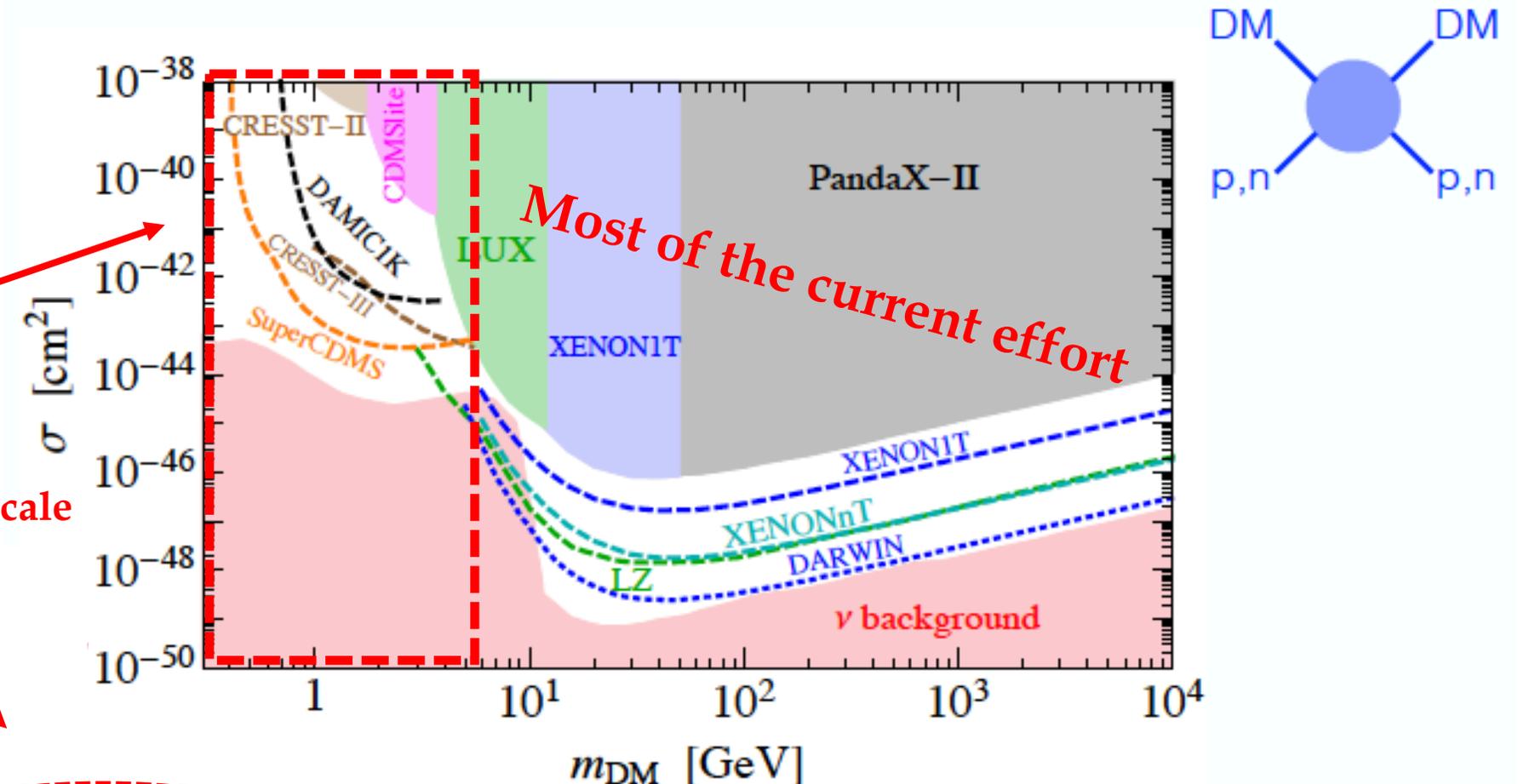
< 10 keV  
DM too hot, spoils  
structure formation



> 100 TeV  
DM overproduced

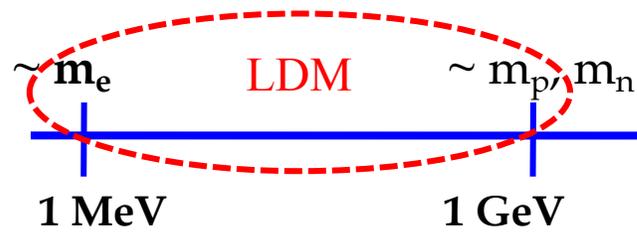
# Search for Dark Matter with thermal origin

DM candidates with thermal origin can have mass between 10 keV and 100 TeV.



Light Dark Matter and corresponding light mediators below the Fermi scale

< 10 keV  
DM too hot, spoils structure formation



WIMPs paradigm

> 100 TeV  
DM overproduced

# Using existing/future (flavor) facilities to study the Dark Sector

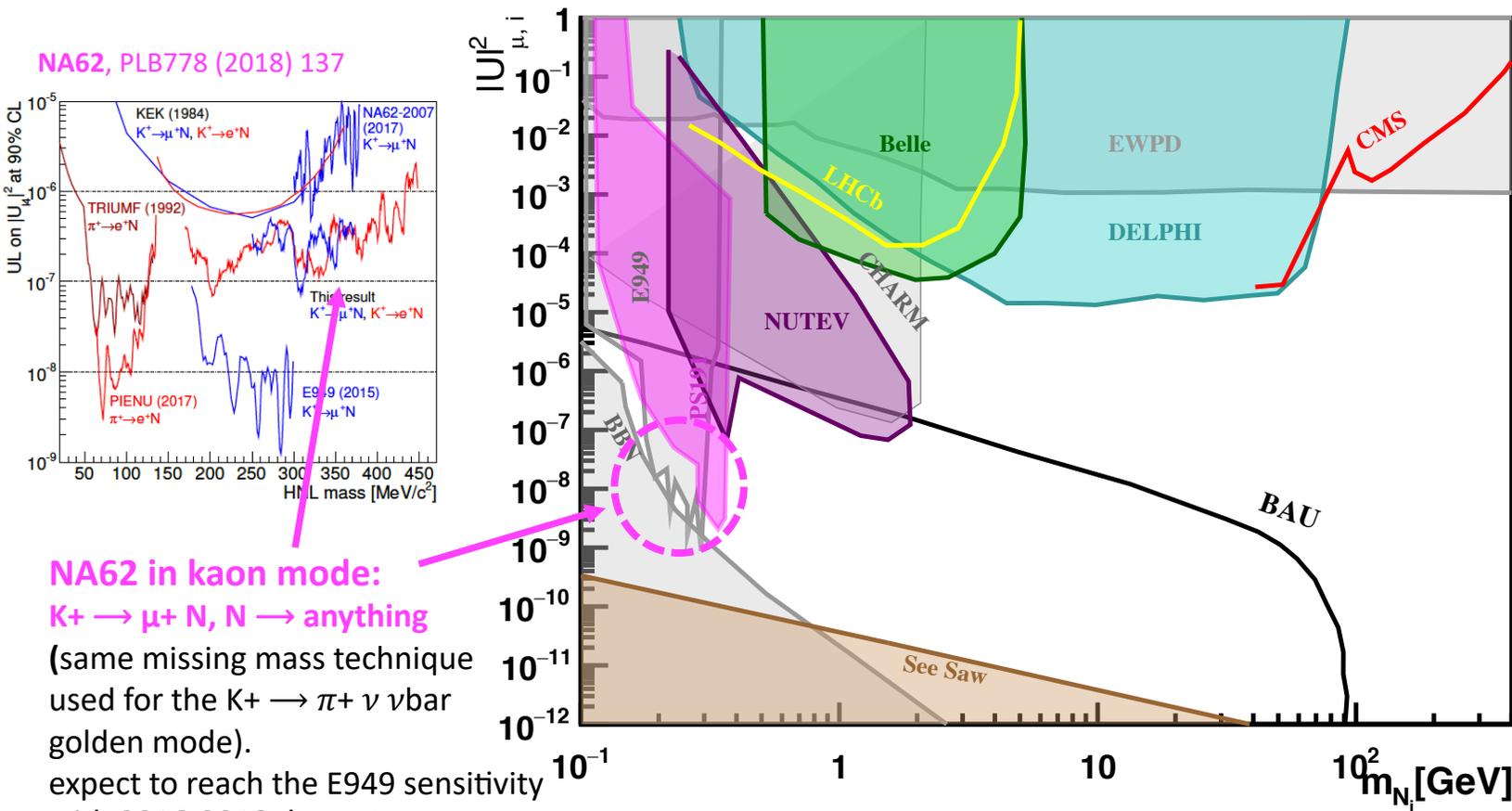


Kaon, D, B physics

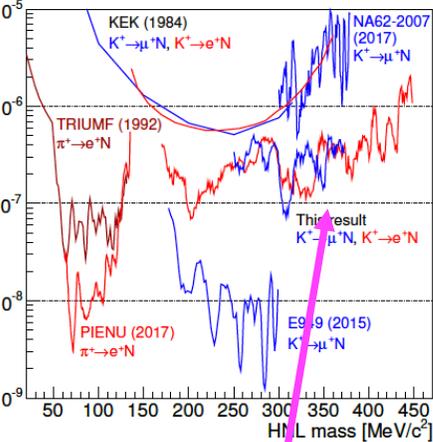
Hidden Sector physics

# Using existing/future (flavor) facilities to study the Dark Sector

## Heavy neutrinos below the EW scale

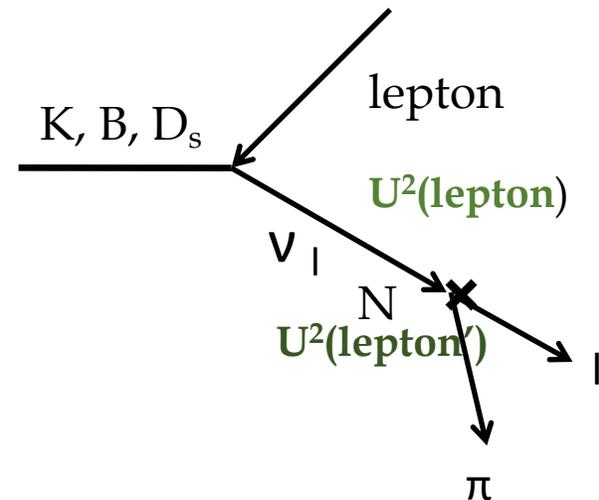


NA62, PLB778 (2018) 137



**NA62 in kaon mode:**  
 $K^+ \rightarrow \mu^+ N, N \rightarrow \text{anything}$   
 (same missing mass technique used for the  $K^+ \rightarrow \pi^+ \nu \bar{\nu}$  golden mode).

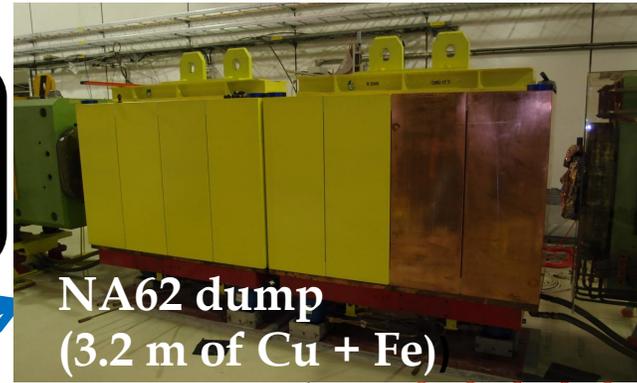
expect to reach the E949 sensitivity with 2016-2018 dataset.



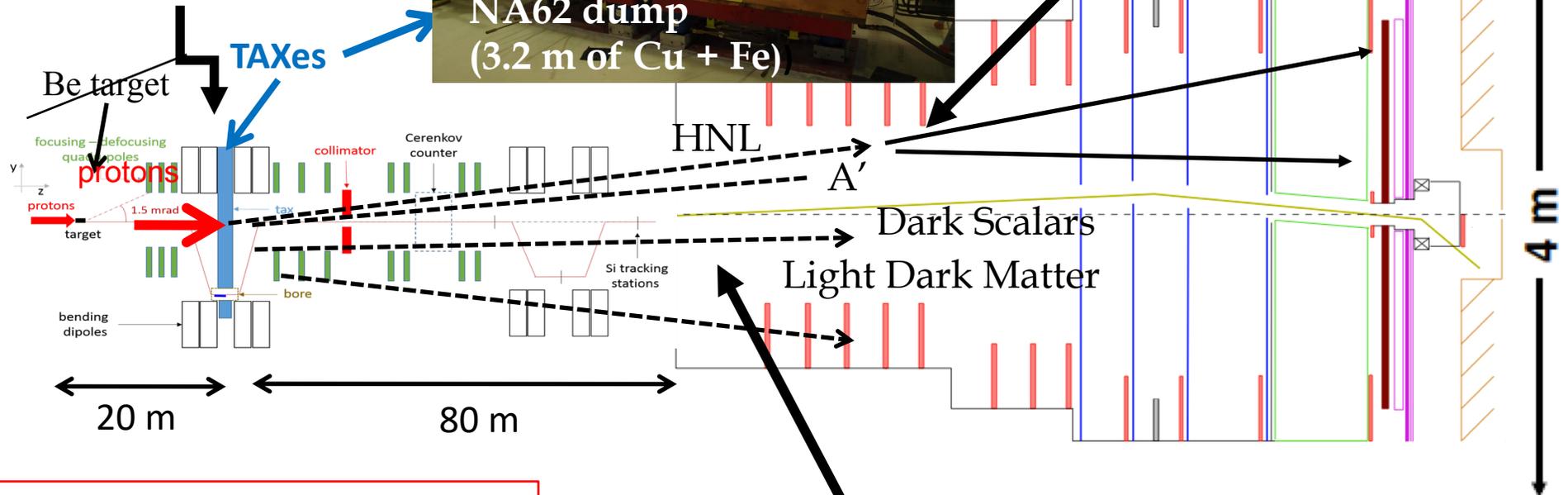
CMS, PRL 120 (2018) 221801  
 Belle, PRD87 (2013) 071102  
 LHCb, PRD 85 (2012) 112004  
 Delphi, Zphys C 74 (1997) 57  
 CHARM Phys. Lett. 166B (1986) 473  
 NuTeV, PRL 83 (1999), 4943  
 E949: PRD 91 (2015) 052001  
 PS191: PLB 2013 (1988) 332

# NA62 in “dump” operation mode

**Dump mode:** the Be target can be moved away from the beam and the beam let impinging on the collimators ( $22 \lambda_1$ ):



**Signal signature:**  
a vertex appearing in the decay volume and nothing else



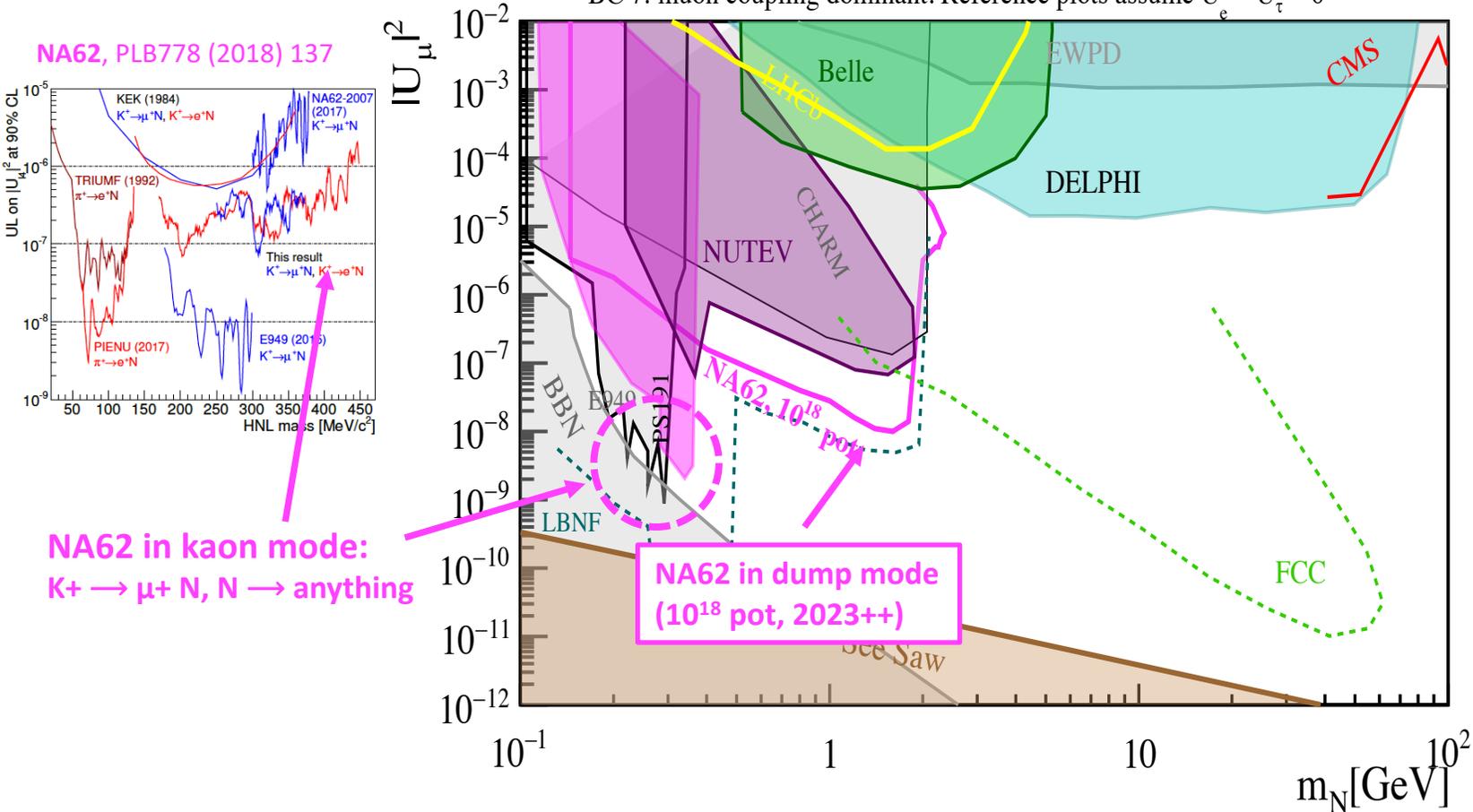
**Proposal to dedicate o(1) year in dump mode at the end of the kaon mode period (2023<sup>++</sup>).**

Heavy Neutral Leptons, Dark Photons, Dark scalars, and ALPS can be originated by charm, beauty and photons produced in the interaction of protons with the dump.

# Using existing/future (flavor) facilities to study the Dark Sector

## Heavy neutrinos below the EW scale

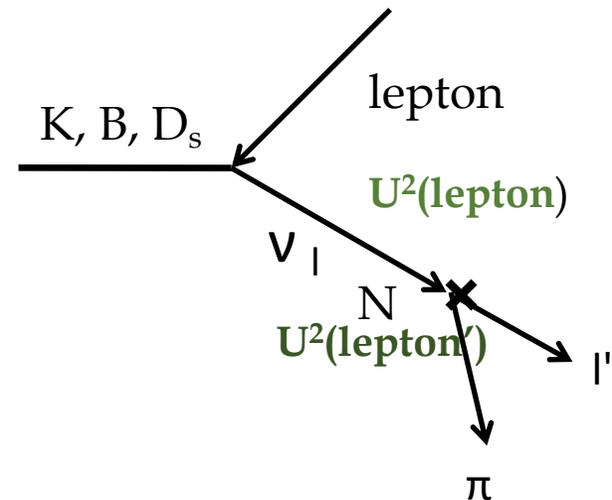
BC 7: muon coupling dominant. Reference plots assume  $U_e = U_\tau = 0$



NA62, PLB778 (2018) 137

NA62 in kaon mode:  
 $K^+ \rightarrow \mu^+ N, N \rightarrow \text{anything}$

NA62 in dump mode  
( $10^{18}$  pot, 2023++)

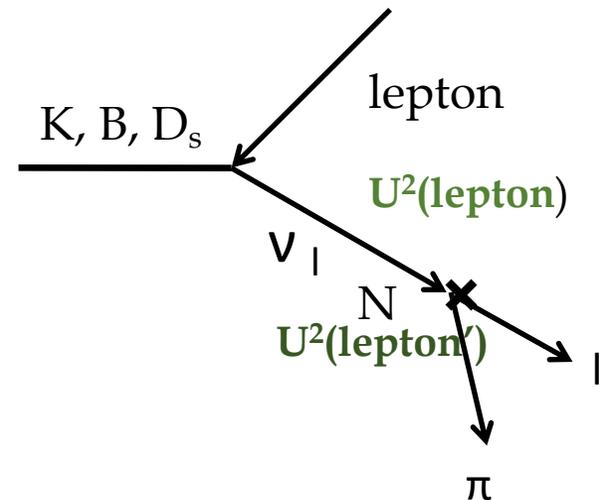
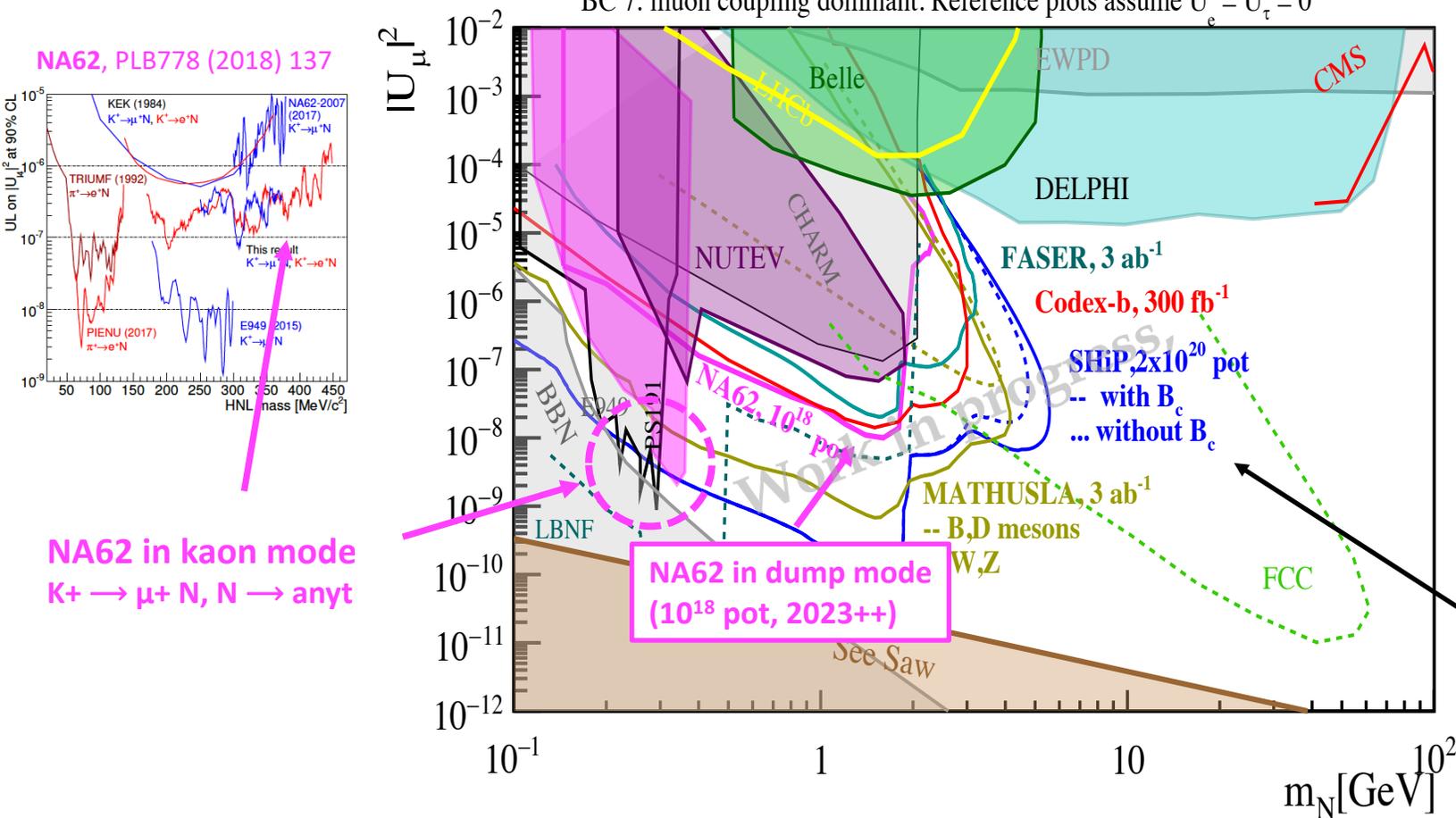


CMS, PRL 120 (2018) 221801  
 Belle, PRD87 (2013) 071102  
 LHCb, PRD 85 (2012) 112004  
 Delphi, Zphys C 74 (1997) 57  
 CHARM Phys. Lett. 166B (1986) 473  
 NuTeV, PRL 83 (1999), 4943  
 E949: PRD 91 (2015) 052001  
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# Using existing/future (flavor) facilities to study the Dark Sector

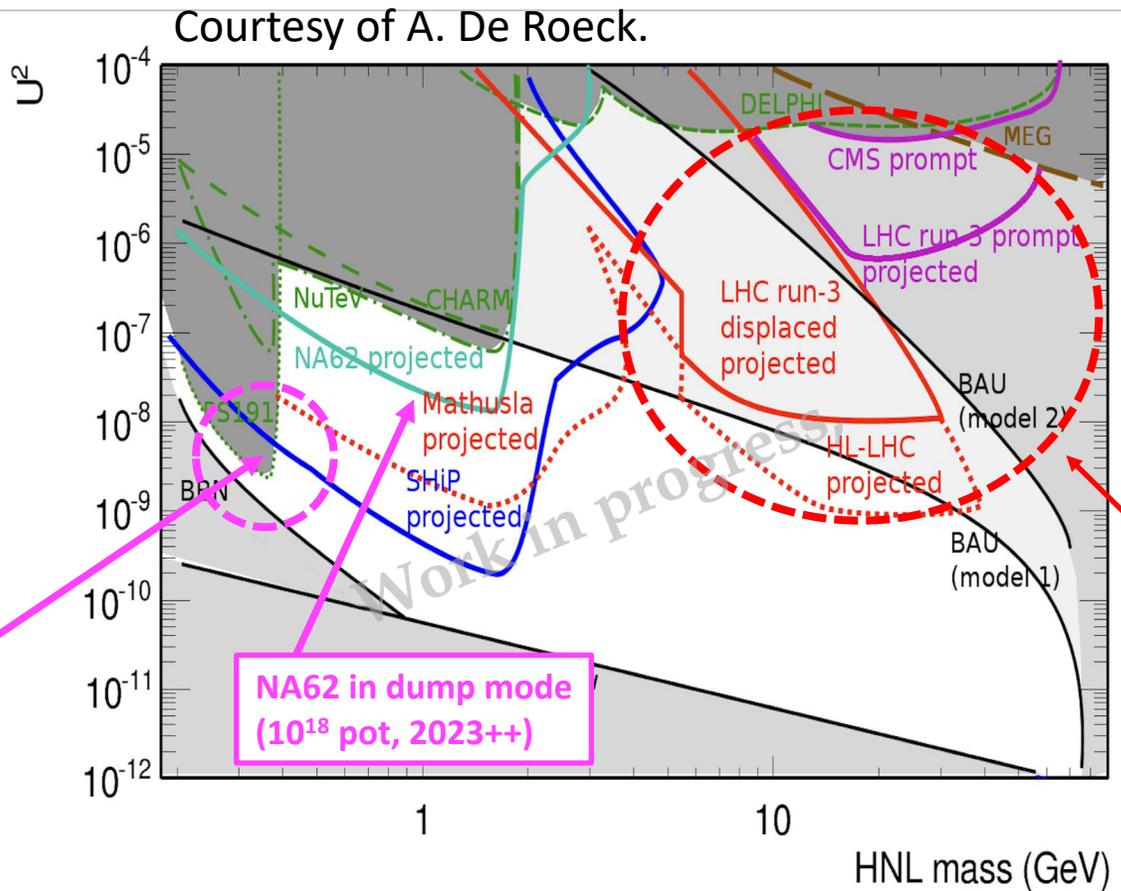
## Heavy neutrinos below the EW scale

BC 7: muon coupling dominant. Reference plots assume  $U_e = U_\tau = 0$

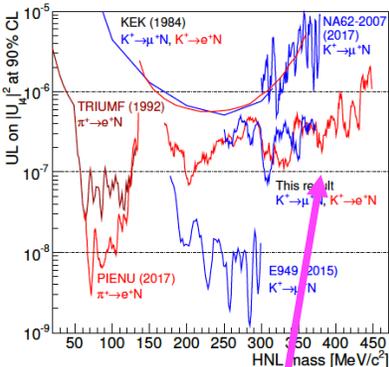


Projects presented in the Physics Beyond Colliders Study group (documents in preparation for ESPP) See J. Jaeckel talk tomorrow

# Using existing/future (flavor) facilities to study the Dark Sector

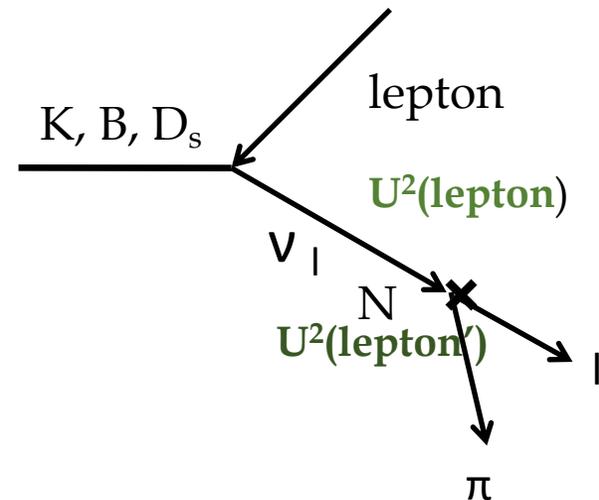


NA62, PLB778 (2018) 137



NA62 in kaon mode:  
 $K^+ \rightarrow \mu^+ N, N \rightarrow \text{anything}$

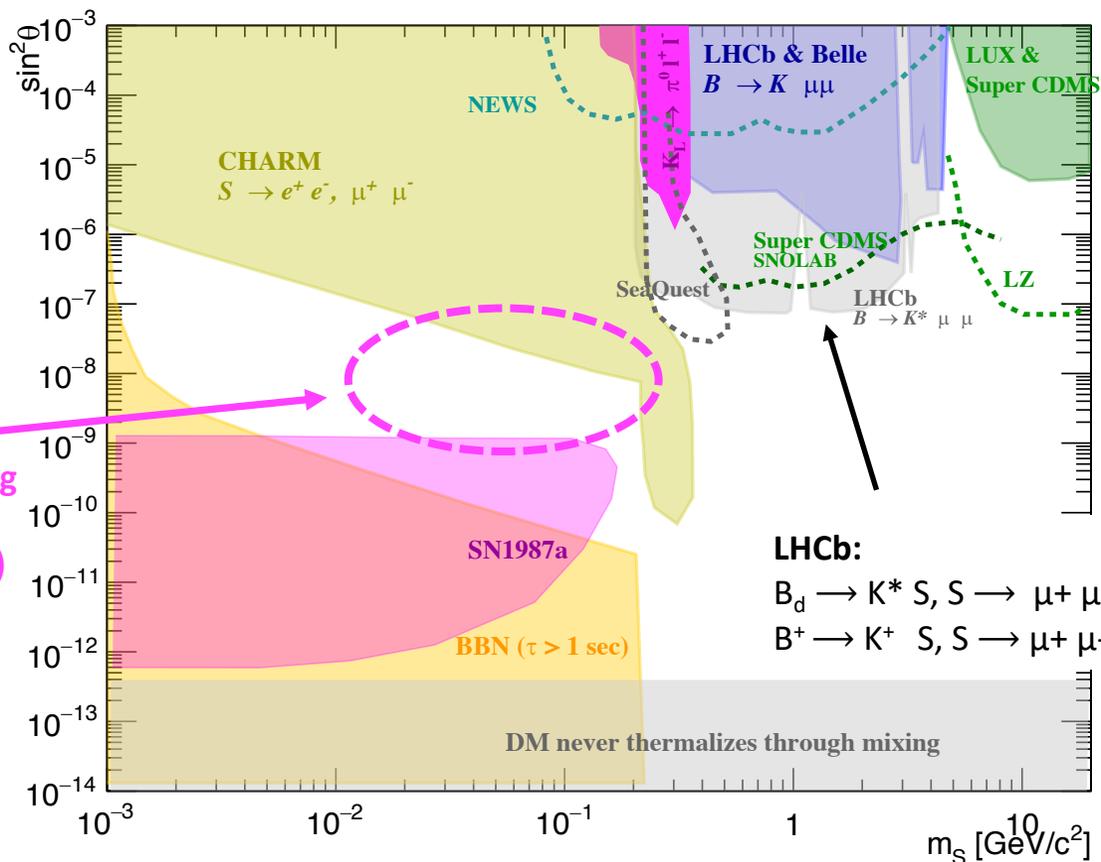
NA62 in dump mode  
 ( $10^{18}$  pot, 2023++)



Sensitivity to HNL from LHC, HL-LHC: plot being prepared for the LLP white paper (P. Mermod)

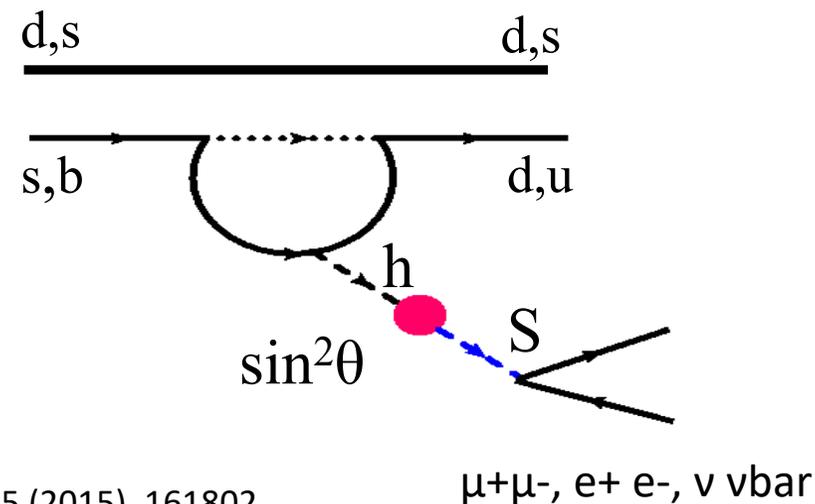
# Using existing/future (flavor) facilities to study the Dark Sector

## Light Dark Matter and corresponding light mediators: the Dark Scalar case



NA62:  
 $K^+ \rightarrow \pi^+ S$ ,  $S \rightarrow \text{anything}$   
 (by product of the main  
 $K^+ \rightarrow \pi^+ \nu \bar{\nu}$  analysis)

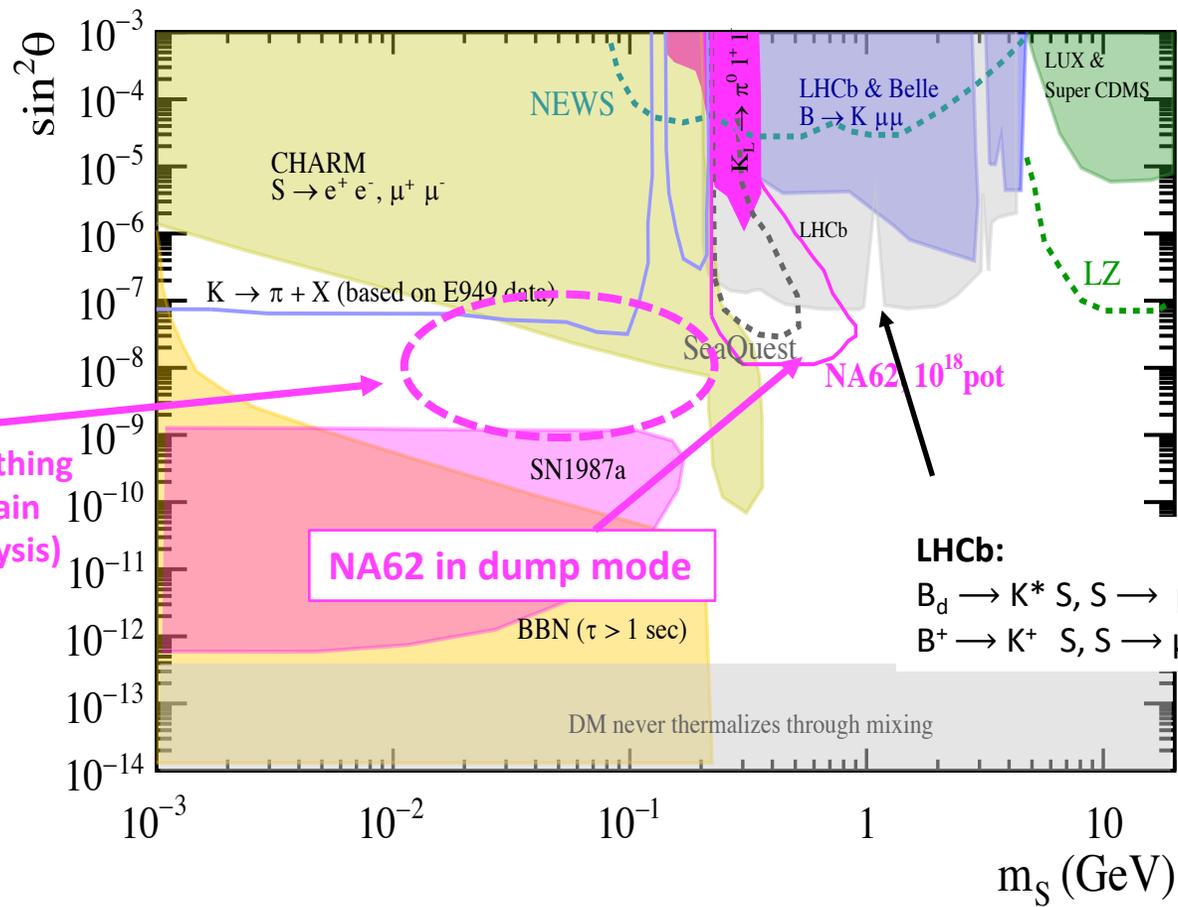
LHCb:  
 $B_d \rightarrow K^* S$ ,  $S \rightarrow \mu^+ \mu^-$ , PRL 15 (2015), 161802  
 $B^+ \rightarrow K^+ S$ ,  $S \rightarrow \mu^+ \mu^-$ , PRD 95 (2017) 071101



See for example, Rept.Prog.Phys. 79 (2016) no.12, 124201 and references therein  
 and G.Perez et al, JHEP 1706 (2017) 050

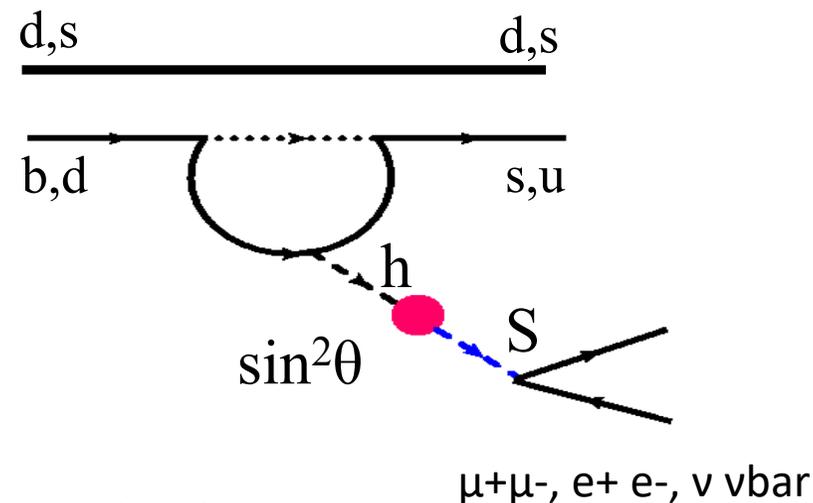
# Using existing/future (flavor) facilities to study the Dark Sector

Light Dark Matter and corresponding light mediators: the Dark Scalar case



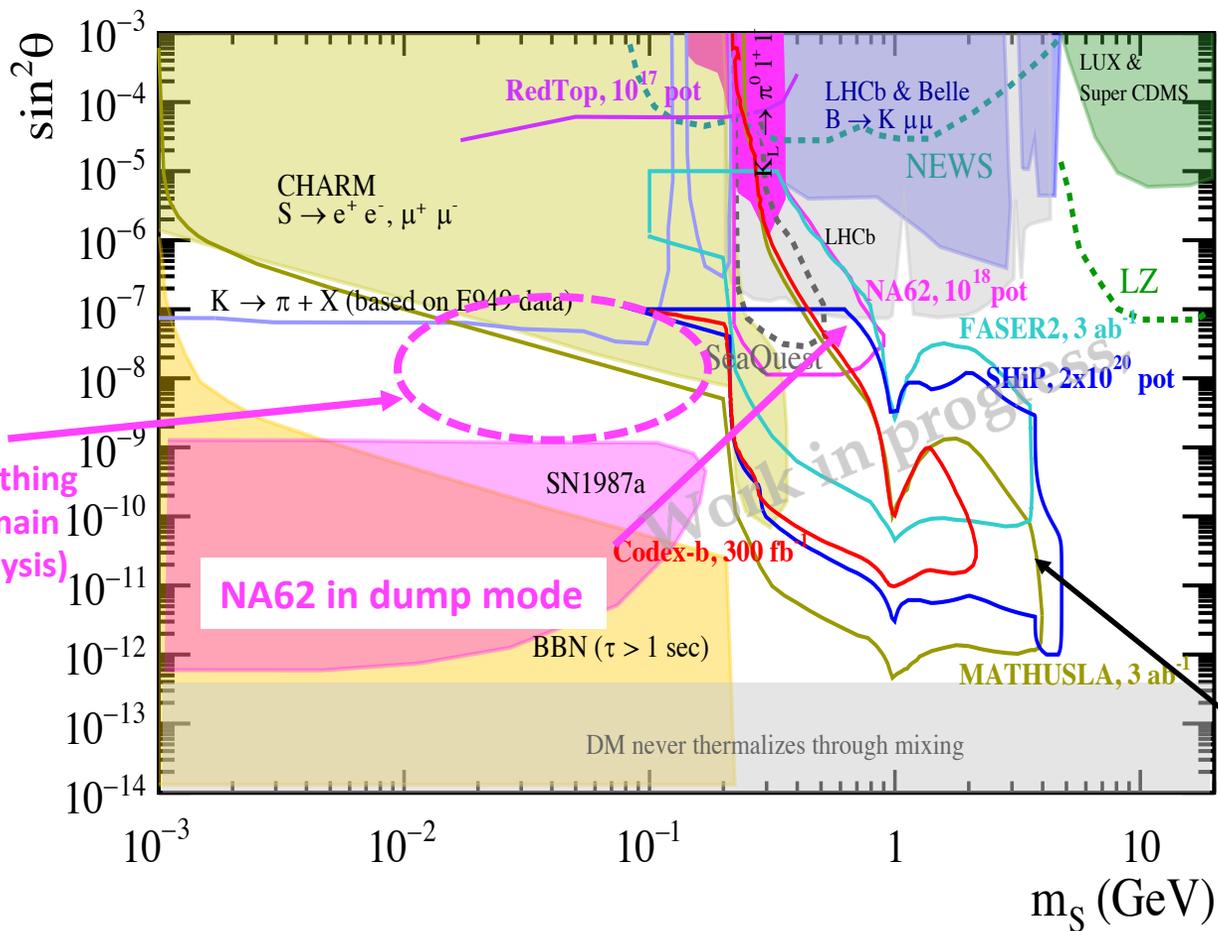
**NA62:**  
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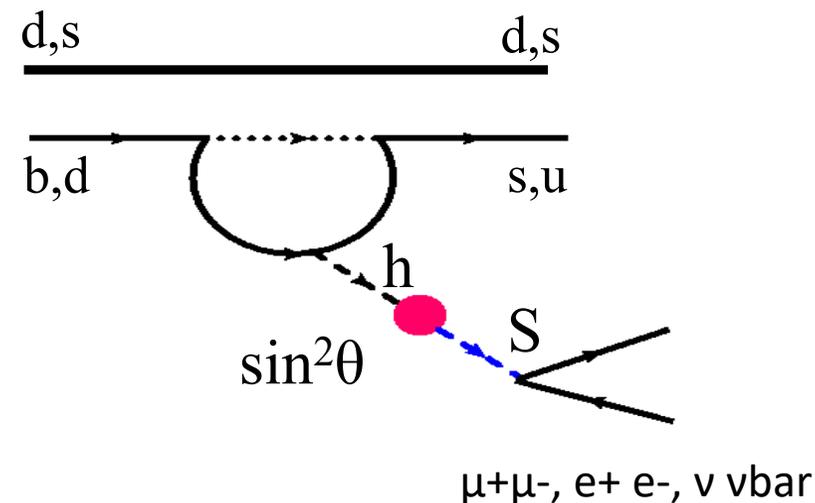


# Using existing/future (flavor) facilities to study the Dark Sector

Light Dark Matter and corresponding light mediators: the Dark Scalar case



NA62:  
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Projects presented in the  
 Physics Beyond Colliders Study group  
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 See J. Jaeckel talk tomorrow

# Conclusions

# Conclusions

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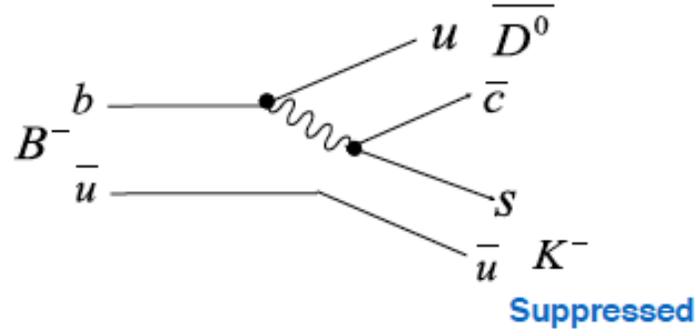
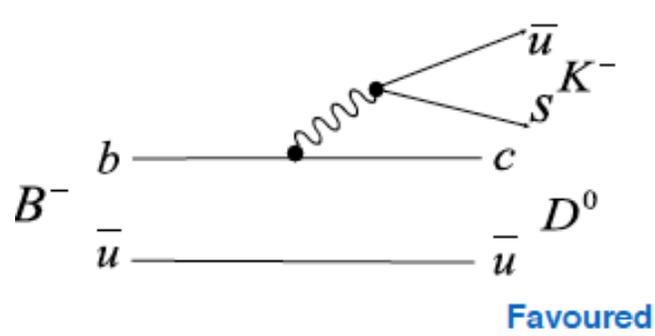
- A wealth of experimental results in the last few years are constraining NP contributions in the CKM paradigm, in particular in  $b \rightarrow s l^+ l^-$  transitions, with unprecedented level of precision.
- We did not observe so far unambiguous deviations from SM predictions, hence:
  - either NP is very heavy and/or mimic the SM in its flavor-breaking pattern:
    - LHCb upgrade(s), Belle-II, NA62, ATLAS/CMS, etc. will keep testing this paradigm during the HL-LHC era;
  - or it is below the Fermi scale and couples very feebly to SM particles and so far escaped detection:
    - new proposals at existing and/or future (flavor) facilities (but not only) will possibly test this new paradigm in the next 5-15 years.

.....Where do we go from here?

SPARES

# $\Phi_3/\gamma$ (phase of $V_{ub}$ ) Determination

- Theory is “pristine” in these approaches,  $\ll 1\%$  on  $\Phi_3$



$$r_B = \frac{|A_{\text{suppressed}}|}{|A_{\text{favoured}}|} \approx \frac{V_{ub}V_{cs}^*}{V_{cb}V_{us}^*} \times [\text{colour supp.}] = 0.1 - 0.2$$

Relative weak phase is  $\Phi_3$ , Relative strong phase is  $\delta_R$

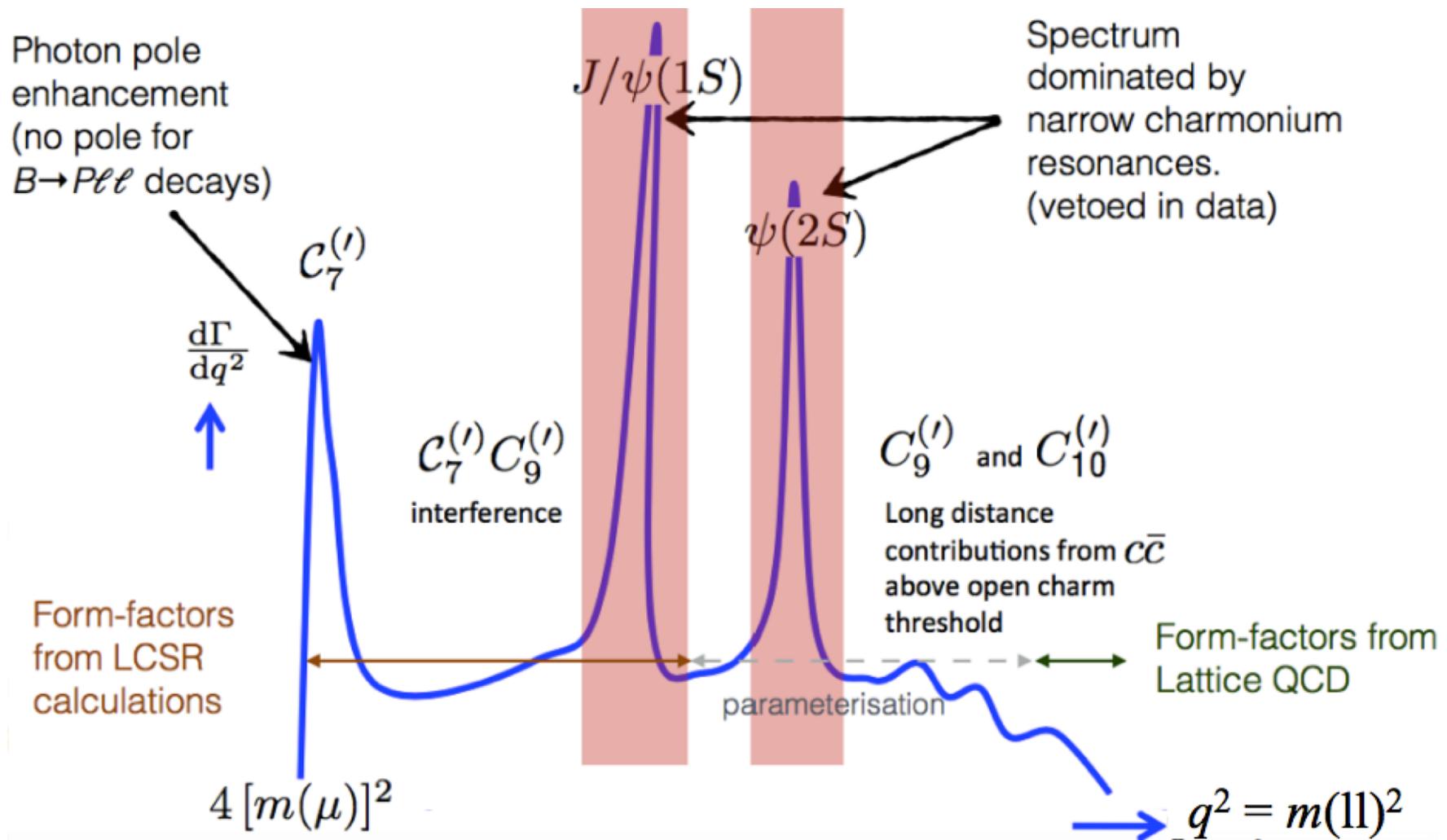
A dream of Belle & Babar: difficult due to  $V_{ub}$  and colour suppression. Many Direct CPV techniques developed at the B-factories.

### 3 $D^0$ mode categories:

- $D_{CP}$ , CP eigenstates [GLW]
- $D_{\text{sup}}$ , Doubly cabibbo suppressed [ADS]
- 3-Body [GGSZ]

- LHCb  $\Phi_3$  from 98 observables

B decay	D decay	Method
$B^+ \rightarrow DK^+$	$D \rightarrow h^+h^-$	GLW
$B^+ \rightarrow DK^+$	$D \rightarrow h^+h^-$	ADS
$B^+ \rightarrow DK^+$	$D \rightarrow h^+\pi^-\pi^+\pi^-$	GLW/ADS
$B^+ \rightarrow DK^+$	$D \rightarrow h^+h^-\pi^0$	GLW/ADS
$B^+ \rightarrow DK^+$	$D \rightarrow K_s^0 h^+ h^-$	GGSZ
$B^+ \rightarrow DK^+$	$D \rightarrow K_s^0 h^+ h^-$	GGSZ
$B^+ \rightarrow DK^+$	$D \rightarrow K_s^0 K^+ \pi^-$	GLS
$B^+ \rightarrow D^* K^+$	$D \rightarrow h^+ h^-$	GLW
$B^+ \rightarrow DK^{*+}$	$D \rightarrow h^+ h^-$	GLW/ADS
$B^+ \rightarrow DK^{*+}$	$D \rightarrow h^+ \pi^- \pi^+ \pi^-$	GLW/ADS
$B^+ \rightarrow DK^+ \pi^+ \pi^-$	$D \rightarrow h^+ h^-$	GLW/ADS
$B^0 \rightarrow DK^{*0}$	$D \rightarrow K^+ \pi^-$	ADS
$B^0 \rightarrow DK^+ \pi^-$	$D \rightarrow h^+ h^-$	GLW-Dalitz
$B^0 \rightarrow DK^{*0}$	$D \rightarrow K_s^0 \pi^+ \pi^-$	GGSZ
$B_s^0 \rightarrow D_s^\mp K^\pm$	$D_s^\pm \rightarrow h^+ h^- \pi^\pm$	TD
$B^0 \rightarrow D^\mp \pi^\pm$	$D^\pm \rightarrow K^+ \pi^- \pi^\pm$	TD



# PBC-BSM: physics targets

Light thermal DM and corresponding light mediators, RH neutrinos at the GeV scale, Axions and axion-like particles, EDM in proton/deuteron and in long-lived charmed hadrons; Ultra-TeV New Physics in extremely rare processes.

Light mediators must be SM singlets, hence options limited by SM gauge invariance:

Portal	Coupling
Dark Photon, $A_\mu$	$-\frac{\epsilon}{2 \cos \theta_W} F'_{\mu\nu} B^{\mu\nu}$
Dark Higgs, $S$	$(\mu S + \lambda S^2) H^\dagger H$
Axion, $a$	$\frac{a}{f_a} F_{\mu\nu} \tilde{F}^{\mu\nu}, \frac{a}{f_a} G_{i,\mu\nu} \tilde{G}_i^{\mu\nu}, \frac{\delta_{\mu a}}{f_a} \bar{\psi} \gamma^\mu \gamma^5 \psi$
Sterile Neutrino, $N$	$y_N L H N$

This is the set of the simplest fields and renormalizable interactions that can be added to the SM to answer the three fundamental questions.

# Aerial picture of the North Area

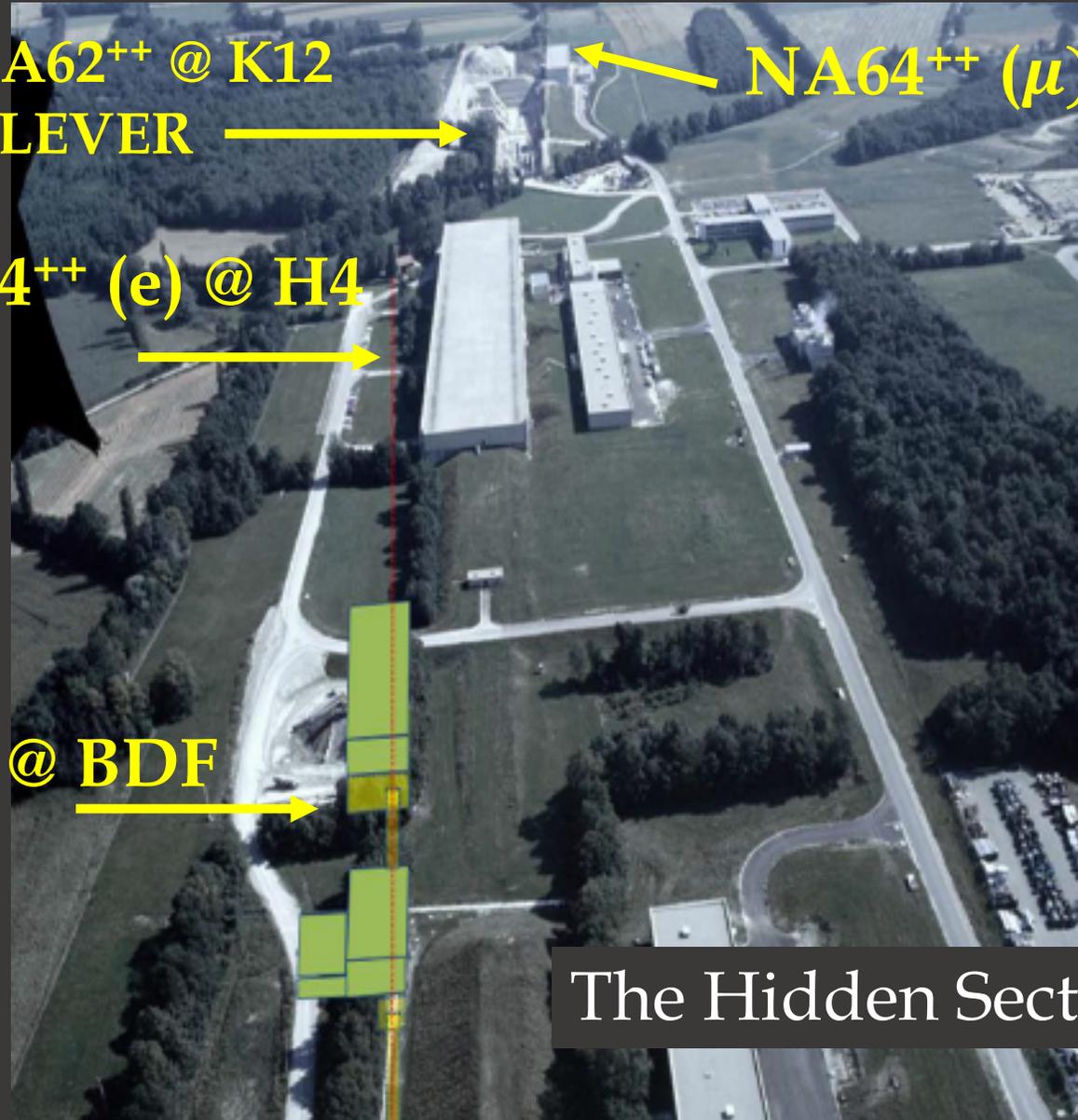
NA62<sup>++</sup> @ K12

KLEVER

NA64<sup>++</sup> ( $\mu$ ) @ M2

NA64<sup>++</sup> (e) @ H4

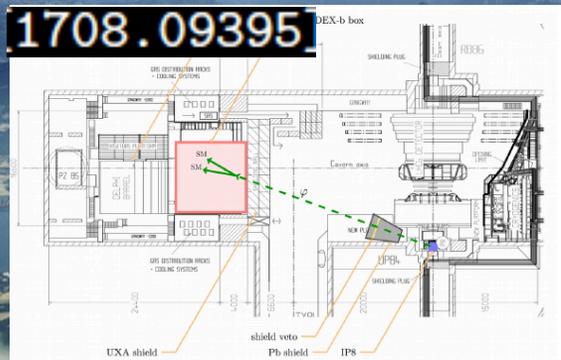
SHiP, tauFV @ BDF



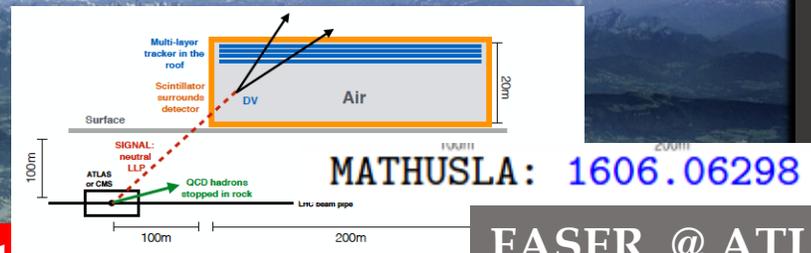
The Hidden Sector "Campus" (HSC)

# MilliQan, MATHUSLA, FASER, Codex-b @ the LHC IPs

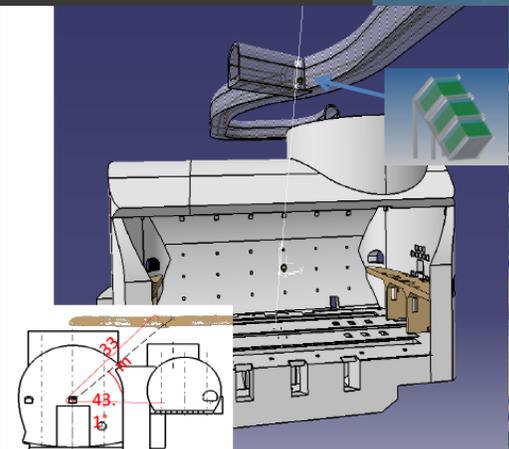
Codex-b @ LHCb IP



MATHUSLA @ ATLAS or CMS IPs



MilliQan @ CMS IP



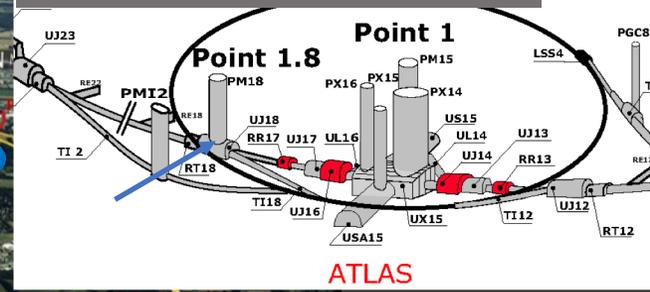
LHCb

ATLAS

CMS

SPS

FASER @ ATLAS IP



FASER: 1708.09389

LHC

LHC 27 km

MilliQan: 1607.04669

+ an extremely active LLP community inside ATLAS, CMS and LHCb collaborations