

LATTES: a new detector concept for a large field-of-view gamma-ray experiment in the Southern hemisphere

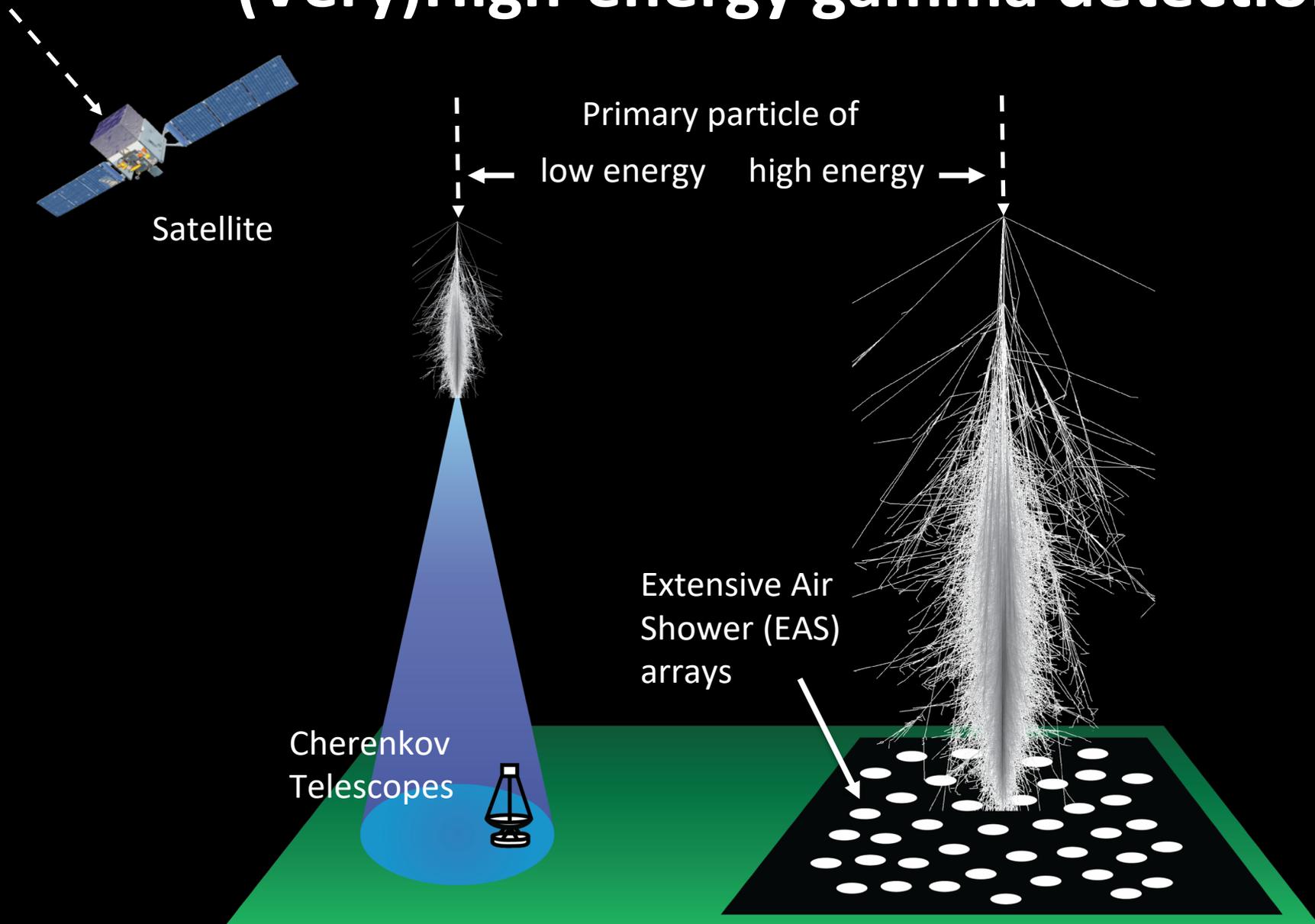
Bernardo Tomé

on behalf of the LATTES team

www.lip.pt/experiments/lattes/



(Very)High-energy gamma detection



Extensive air shower arrays at high-altitude = large field of view + large duty cycle + low energy



VERITAS



MAGIC



RPCs

ARGO

RPCs

LHASSO



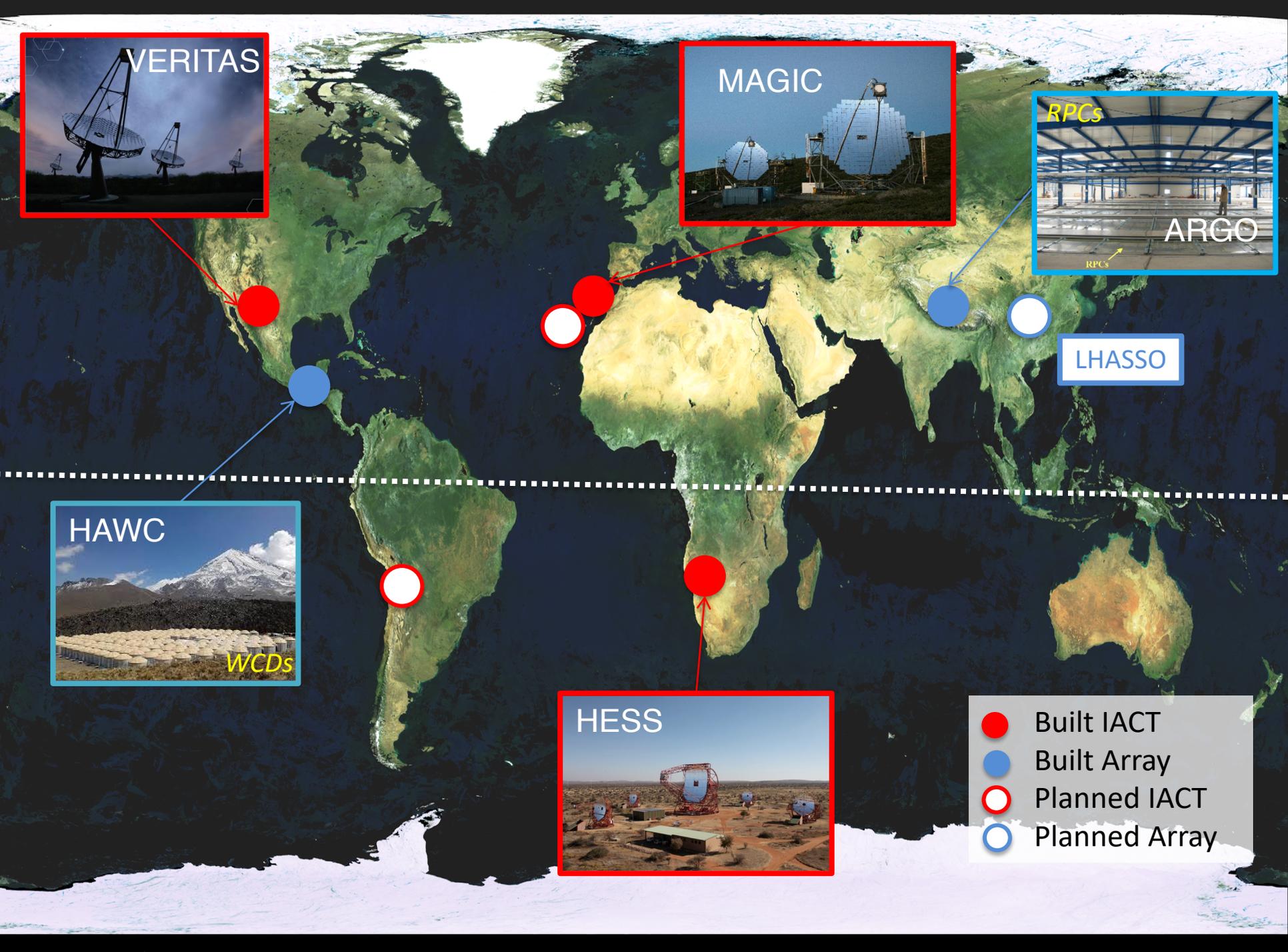
HAWC

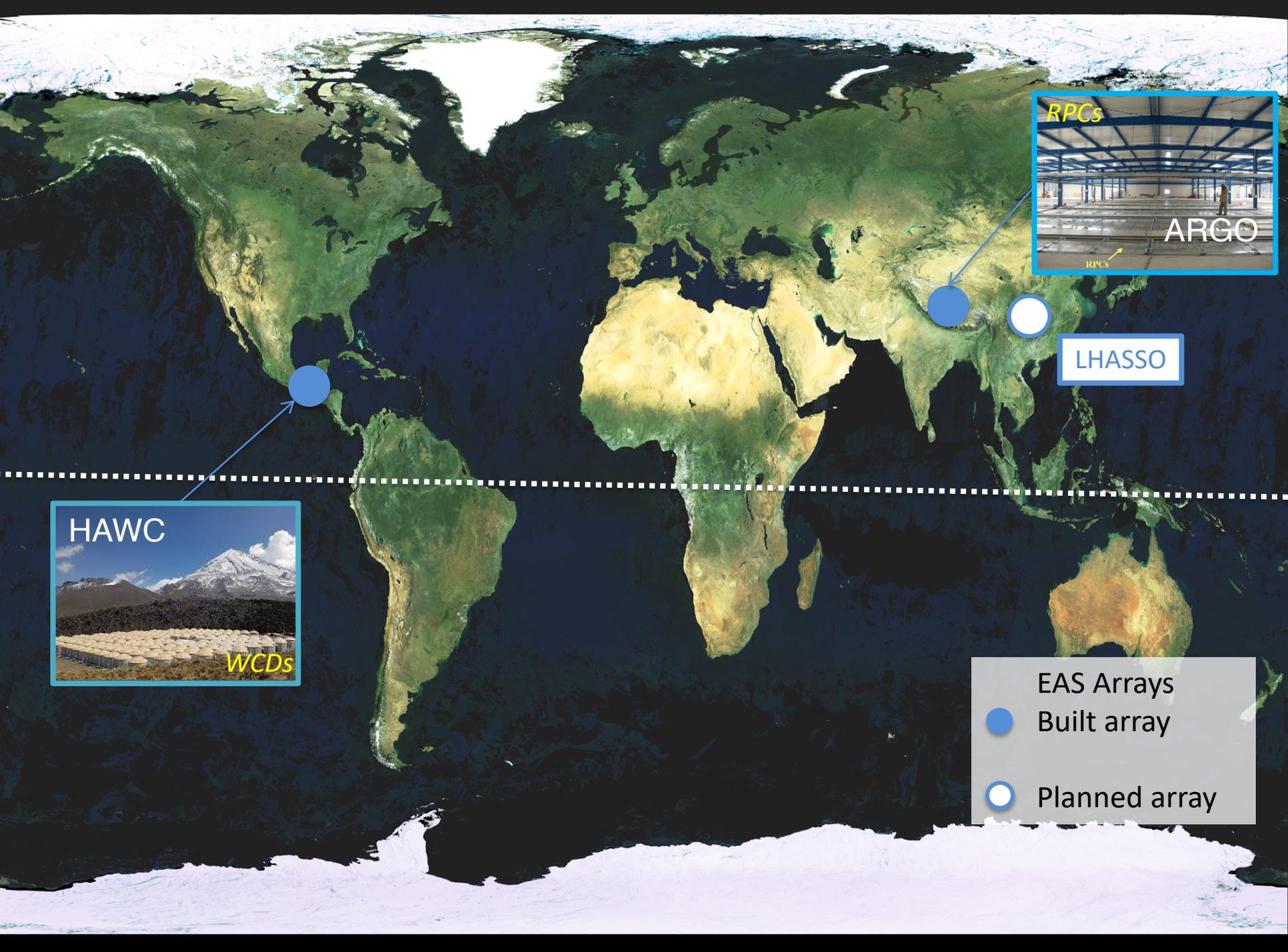
WCDs



HESS

- Built IACT
- Built Array
- Planned IACT
- Planned Array



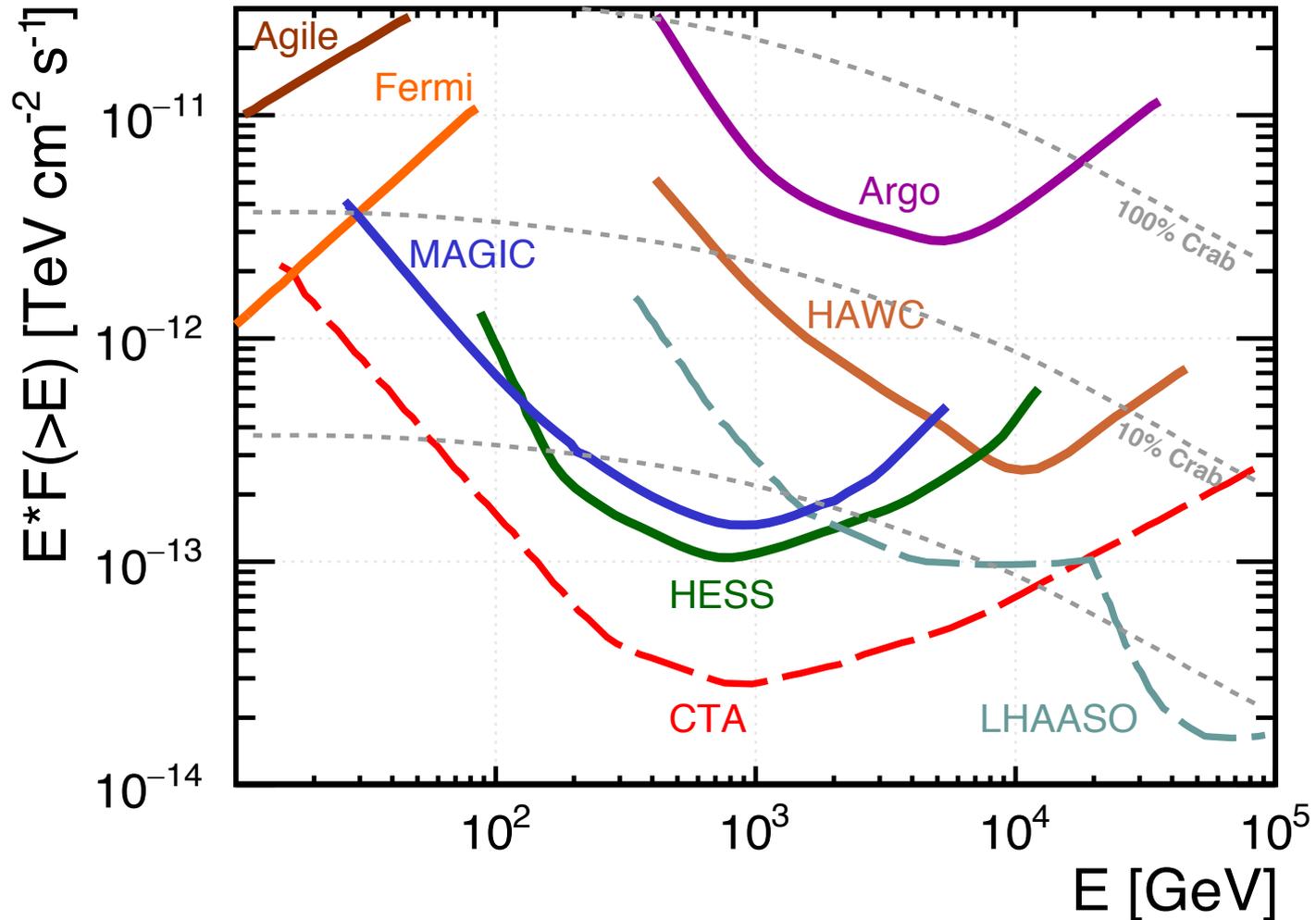


LHASO

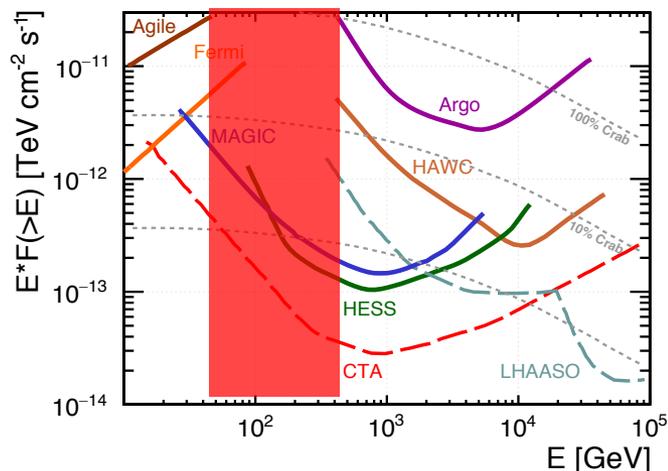


- EAS Arrays
- Built array
- Planned array

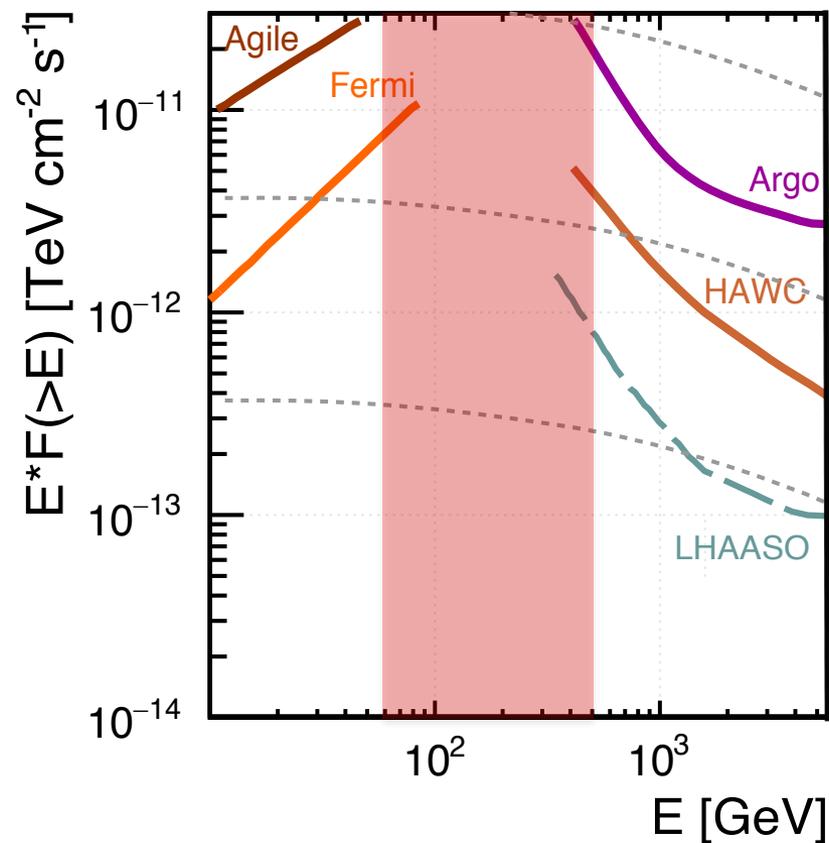
Sensitivity to high-energy gamma-rays



Wide FoV experiments



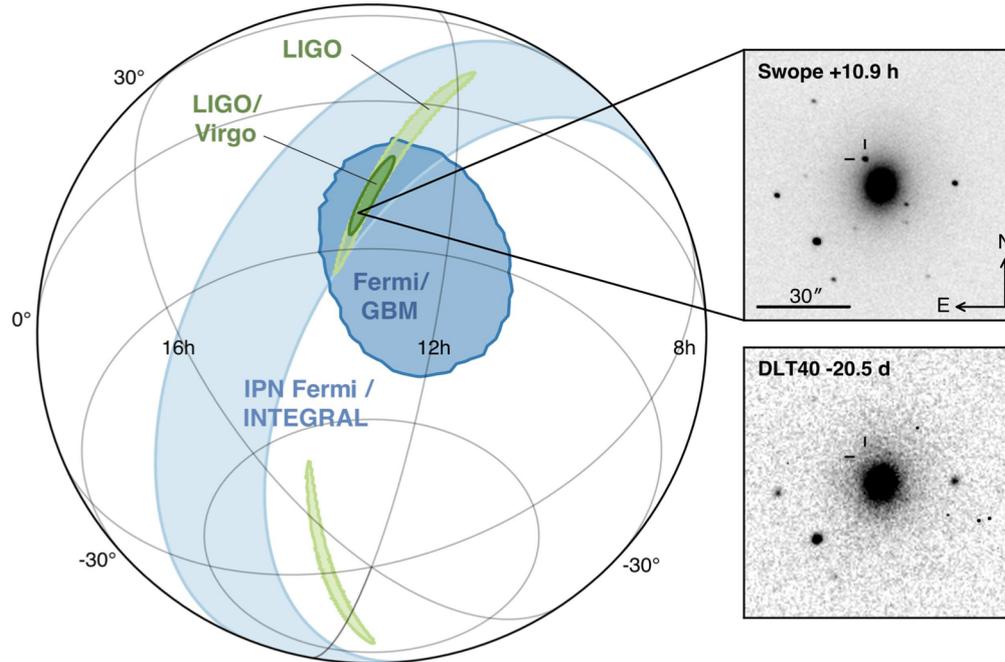
- No wide field-of-view experiment covering the **southern hemisphere** sky
- Gap around **100 GeV** between satellite and ground based observations



The era of **multi-messenger** observations



Joint publication of LIGO, VIRGO, INTEGRAL, Fermi, IceCube, Pierre Auger ...



- Simultaneous observation of a Gravitational Wave + electromagnetic counterparts
- Study of **transient phenomena in all energy windows** is one of the main ingredients
- **Large sky coverage** to maximize chances for multi-detection

Large Array Telescope for Tracking Energetic Sources

Astroparticle Physics 99 (2018) 34–42

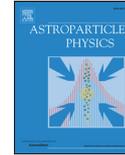


ELSEVIER

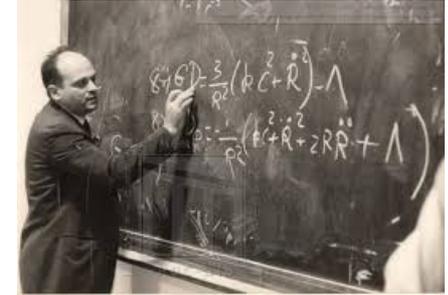
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Astroparticle Physics

journal homepage: www.elsevier.com/locate/astropartphys



Astropart.Phys. 99 (2018) 34-42



Cesar Lattes

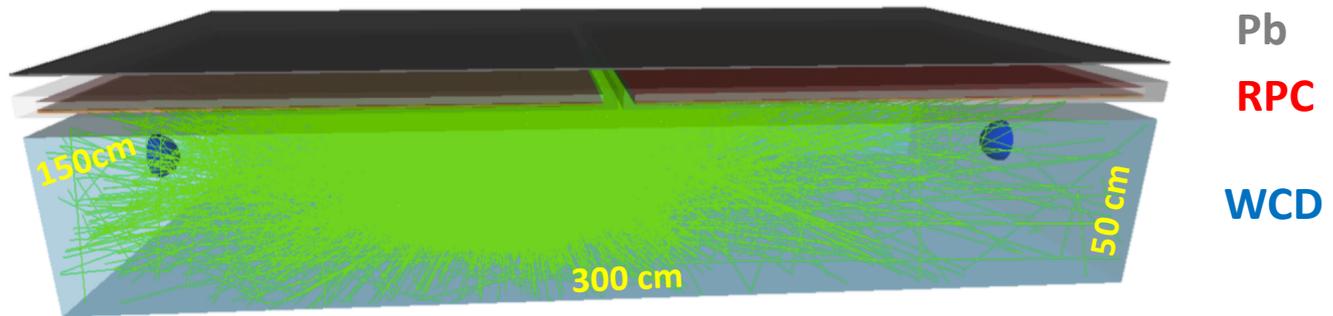


Design and expected performance of a novel hybrid detector for very-high-energy gamma-ray astrophysics

P. Assis^{a,b}, U. Barres de Almeida^c, A. Blanco^d, R. Conceição^{a,b,*}, B. D'Ettorre Piazzoli^e,
A. De Angelis^{f,g,b,a}, M. Doro^{h,f}, P. Fonte^d, L. Lopes^d, G. Matthiaeⁱ, M. Pimenta^{b,a}, R. Shellard^c,
B. Tomé^{a,b}

- A next generation large FoV gamma-ray observatory :
 - Located in the South Hemisphere
 - Low energy threshold:
 - High altitude
 - Hybrid detector concept

LATTES hybrid detector concept

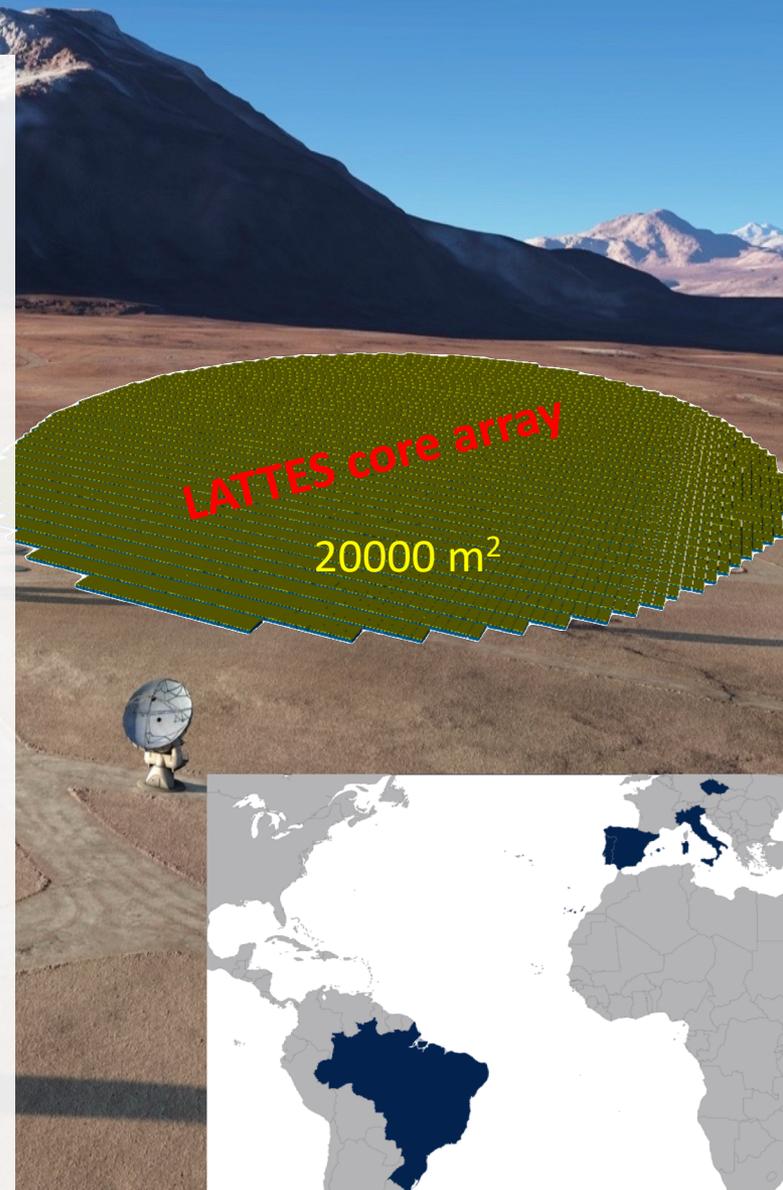


RPCs : time and spatial resolution
WCDs: e.m. energy, g/h discrimination
and trigger

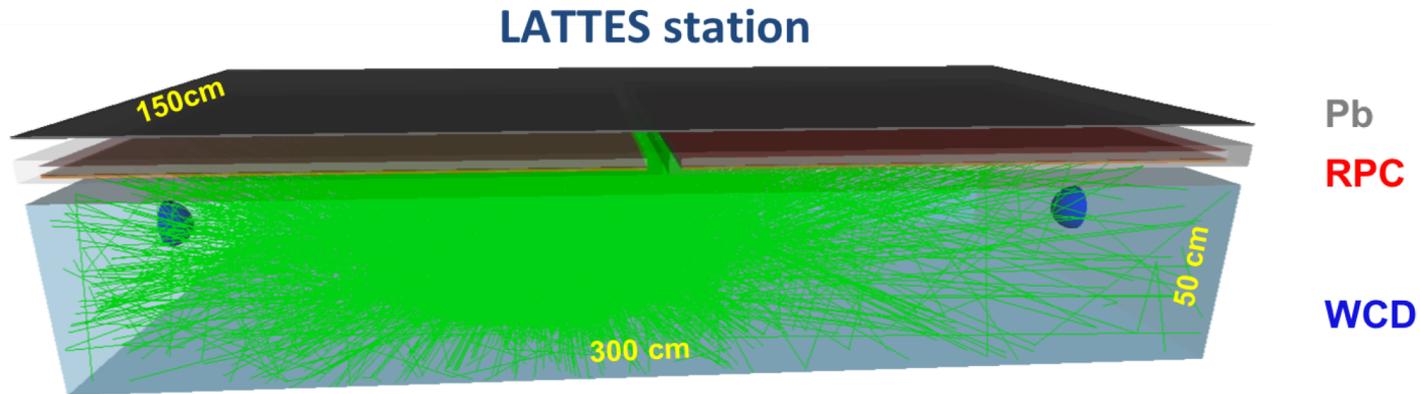
LATTES

Large Array Telescope for Tracking Energetic Sources

- ❖ Joint Brazil/Italy/Portugal initiative
- ❖ Czech group joined recently
- ❖ Possible sites:
 - ❖ Atacama Large Millimeter Array site - Chajnantor plateau (5200 m)
 - ❖ North of Argentina (~ 5000 m)
 - ❖ ...
- ❖ LATTES array baseline
 - ❖ Compact core array
 - ❖ Area: 20 000 m²
 - ❖ Target lowest energies ($E_{\min} \sim 100$ GeV)
 - ❖ Sparse array
 - ❖ Area: 100 000 m²
 - ❖ Cover energies up to 100 TeV



LATTES hybrid concept baseline design



Thin lead plate (Pb)

- 5.6 mm (one radiation length)

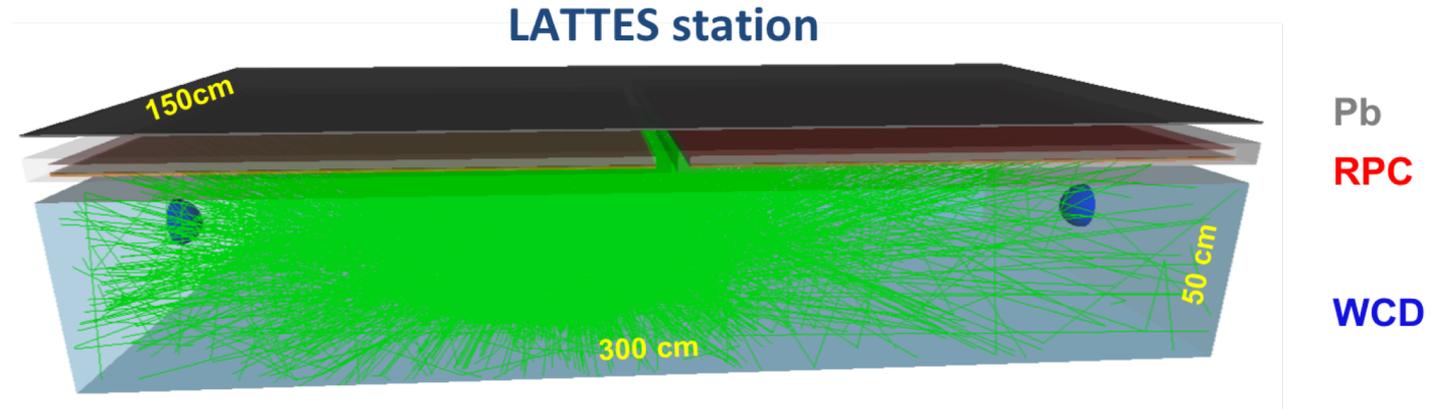
Resistive Plate Chambers (RPC)

- 2 RPCs per station
- Each RPC with 4x4 readout pads

Water Cherenkov Detector (WCD)

- 2 PMTs; 15 cm diameter
- inner walls covered with white diffusing Tyvek

LATTES hybrid concept



Thin lead plate (Pb)

- Convert the shower photons;
- Improve angular reconstruction

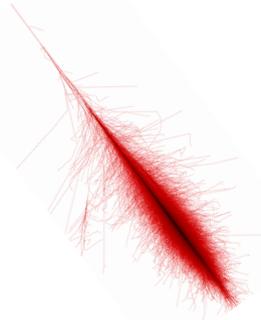
Resistive Plate Chambers (RPC)

- Sensitive to charged particles
- Very good time (1 ns) and spatial resolution (tens of cm)
- Improve geometric reconstruction
- Explore shower particle patterns at ground

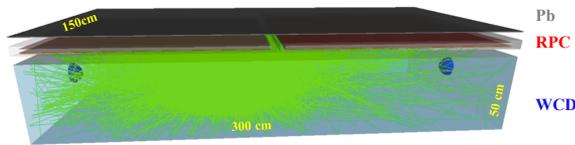
Water Cherenkov Detector (WCD)

- Sensitive to shower photons and charged particles
- Measure energy flow at ground
- Improve trigger capability
- Improve gamma/hadron discrimination

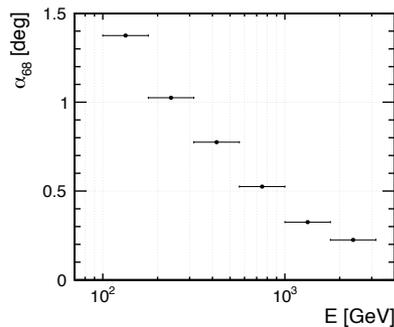
Simulating LATTES performance



Shower simulation
(CORSIKA)



Detector simulation
(Geant4)



Shower reconstruction
(LATTESrec)

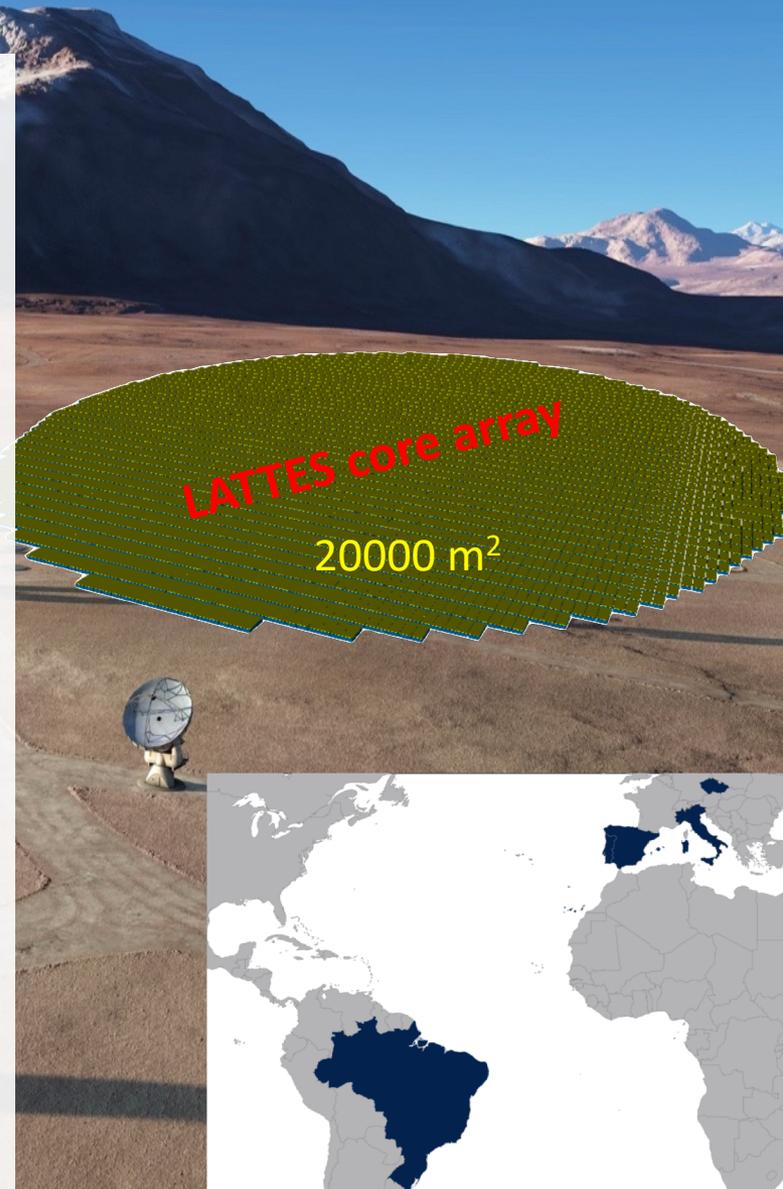
LATTES

Large Array Telescope for Tracking Energetic Sources

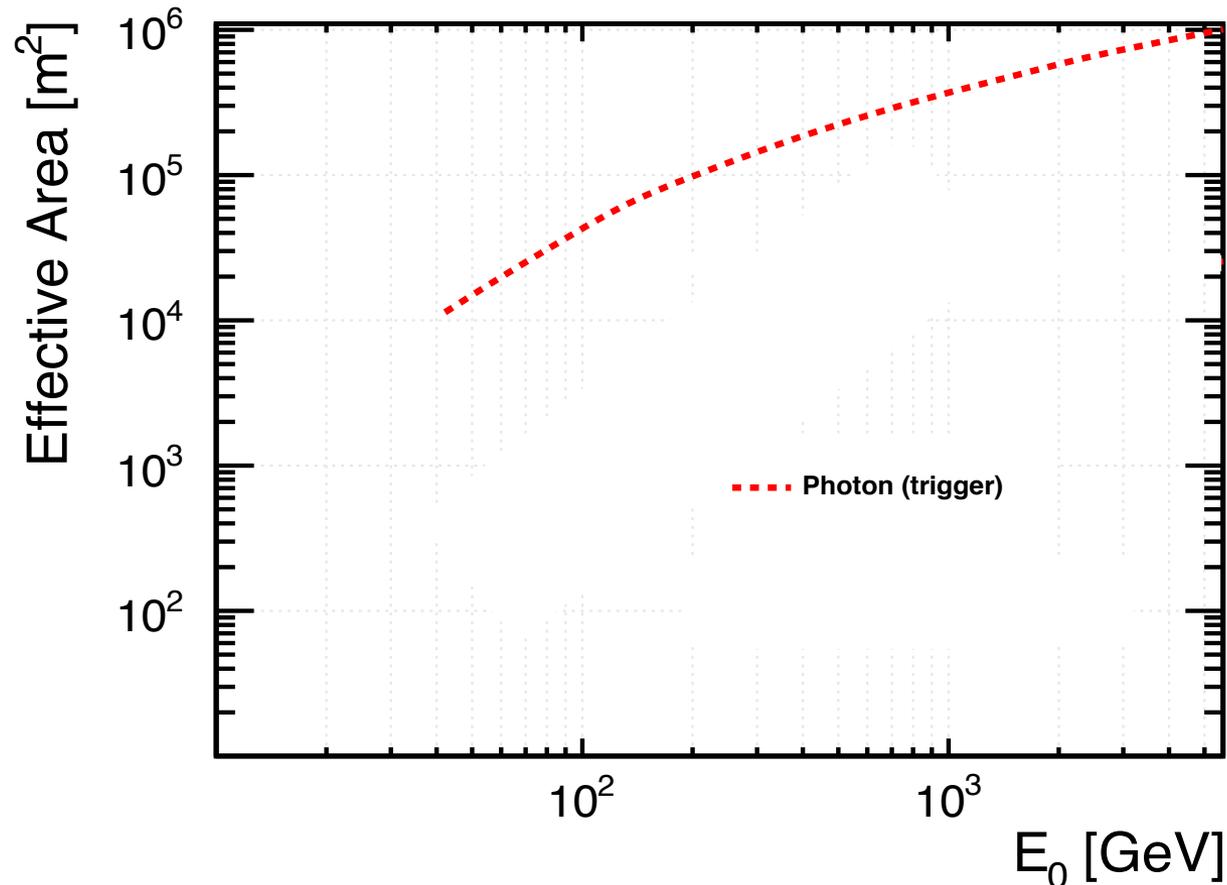
- ❖ Joint Brazil/Italy/Portugal initiative
- ❖ Czech group joined recently
- ❖ Possible sites:
 - ❖ Atacama Large Millimeter Array site - Chajnantor plateau (5200 m)
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 - ❖ ...

❖ LATTES array baseline

- ❖ Compact core array
 - ❖ Area: 20 000 m²
 - ❖ Target lowest energies ($E_{\min} \sim 100 \text{ GeV}$)
- ❖ Sparse array
 - ❖ Area: 100 000 m²
 - ❖ Cover energies up to 100 TeV



Trigger efficiency



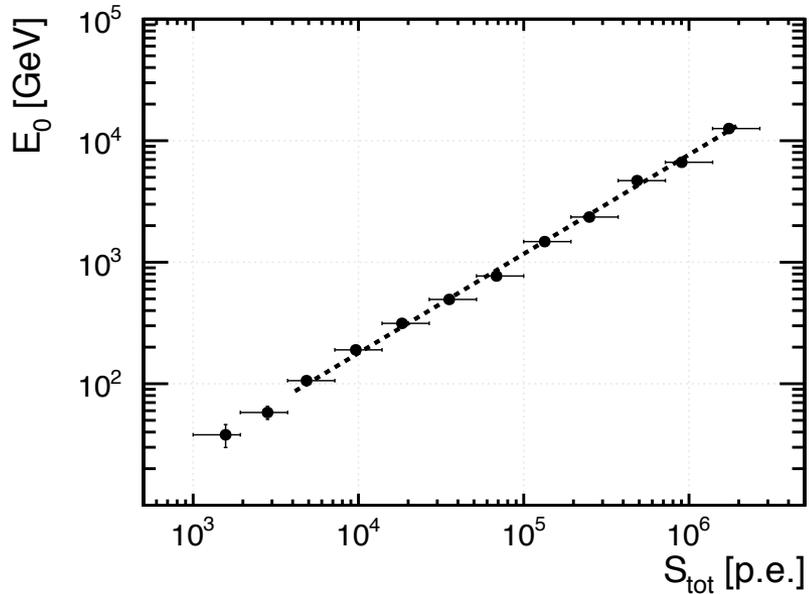
- Use **WCD stations to improve trigger** at low energies
 - Trigger condition
 - Station: require more than 5 p.e. in each PMT
 - Event: require 3 triggered stations
- Trigger efficiency of the order of 10% @ 50 GeV

Energy reconstruction

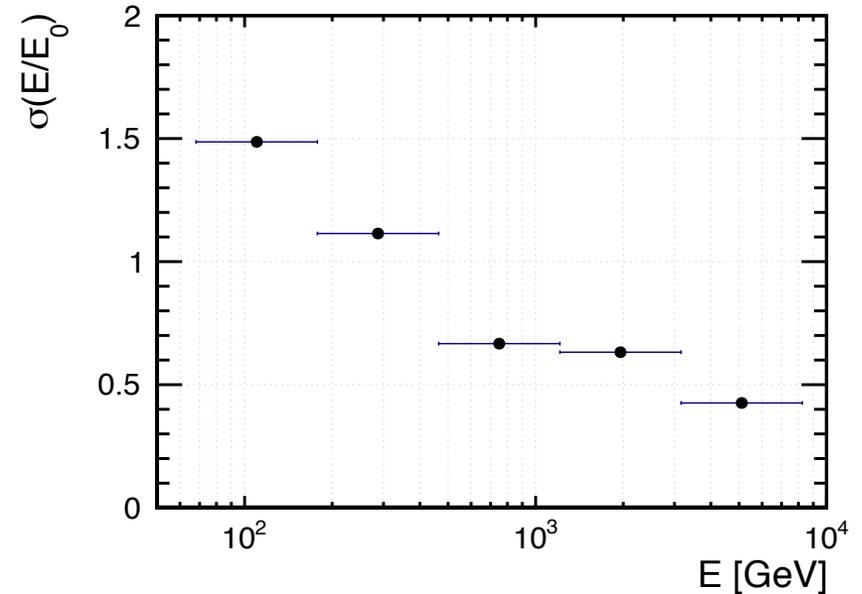
E_0 → Simulated energy

E → Reconstructed energy

Energy Calibration



Energy Resolution

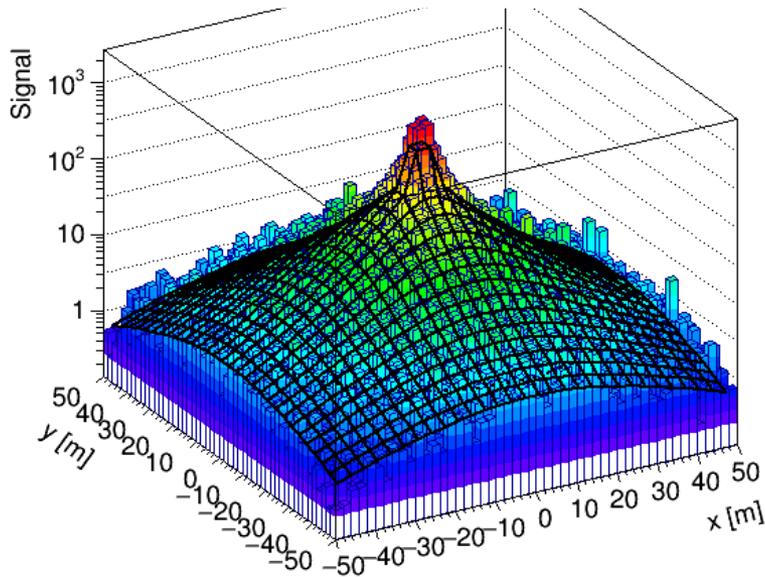


- Use as **energy estimator** the **total signal** recorded by **WCDs**
 - Use only shower cores reconstructed inside array
 - Not accounting for shower fraction outside the array
- Energy resolution at low energy dominated by shower to shower fluctuations

Shower geometry reconstruction

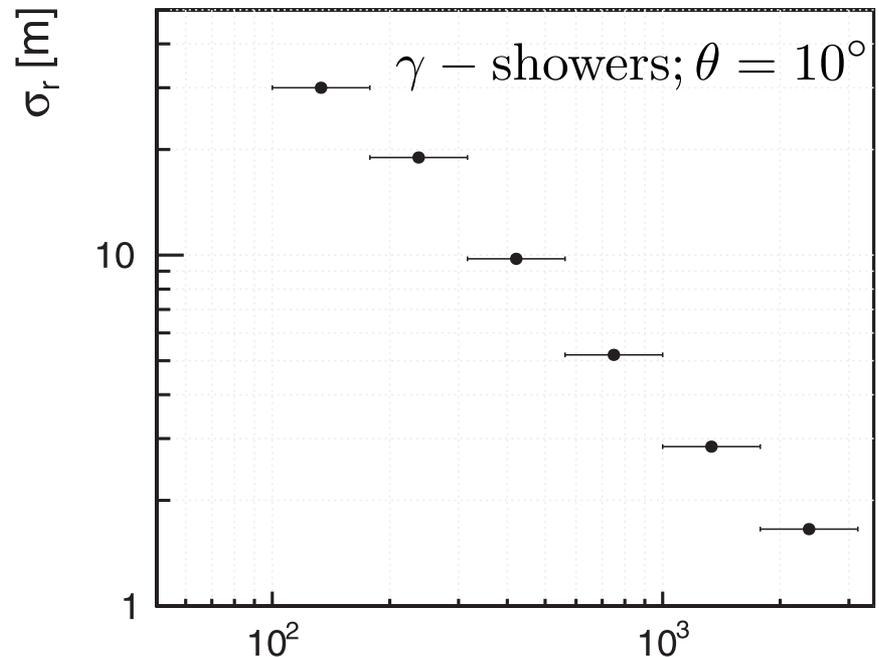
core position

$$S_i = S(A, \vec{x}, \vec{x}_i) = A \left(\frac{1}{2\pi\sigma^2} e^{-|\vec{x}_i - \vec{x}|^2 / 2\sigma^2} + \frac{N}{(0.5 + |\vec{x}_i - \vec{x}|/R_m)^3} \right)$$



- Resolution better than 10 m for $E > 300$ GeV.

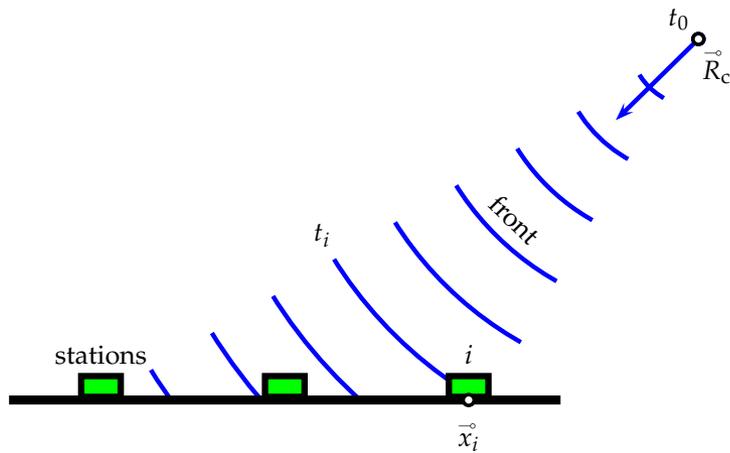
- Fit the LDF of the WCD signal:
 - Photon average LDF to fix the shape
 - Function inspired in HAWC
 - Nearly no evolution with energy
 - Position of maximum gives the shower core



Shower geometry reconstruction

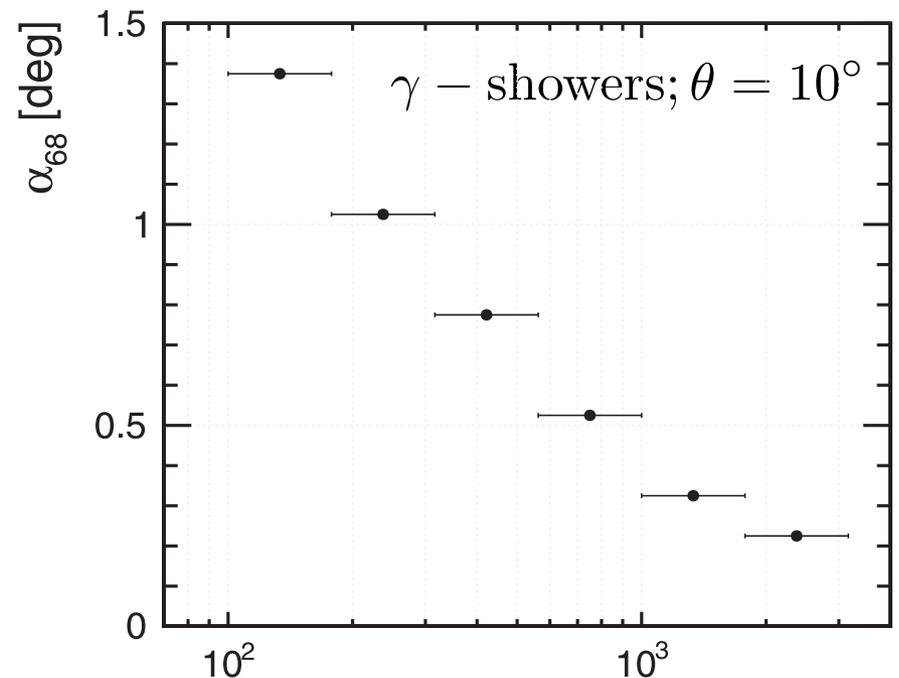
direction

$$\chi^2 = \sum (c \cdot (T_n - T_0) - X_n \cdot l - Y_n \cdot m - R_n \cdot \alpha)^2$$

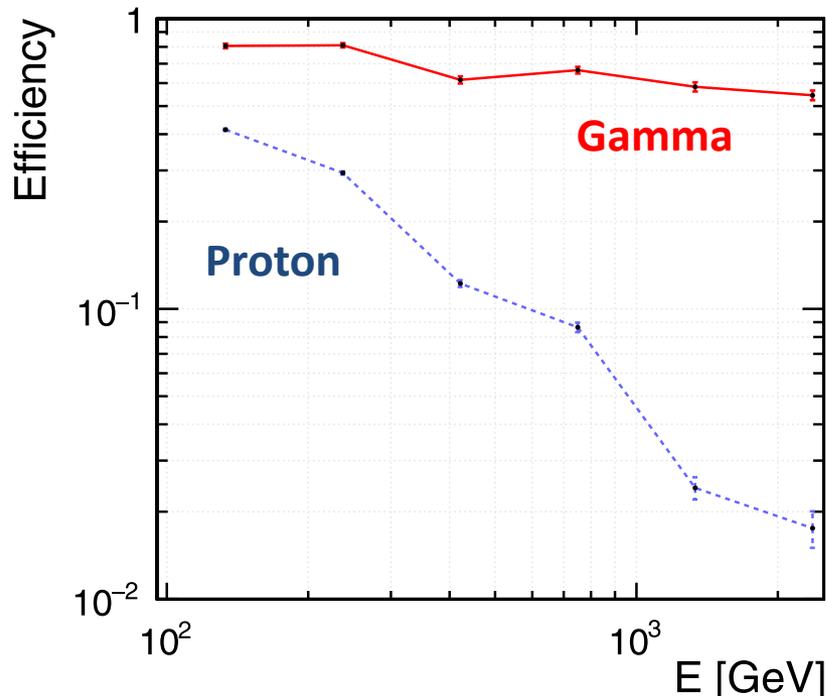
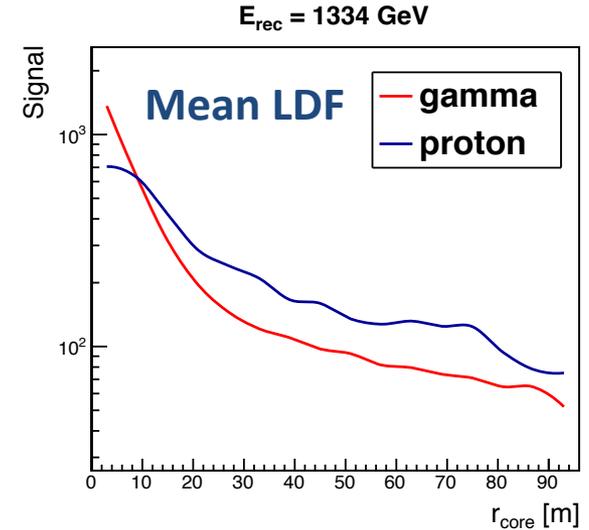
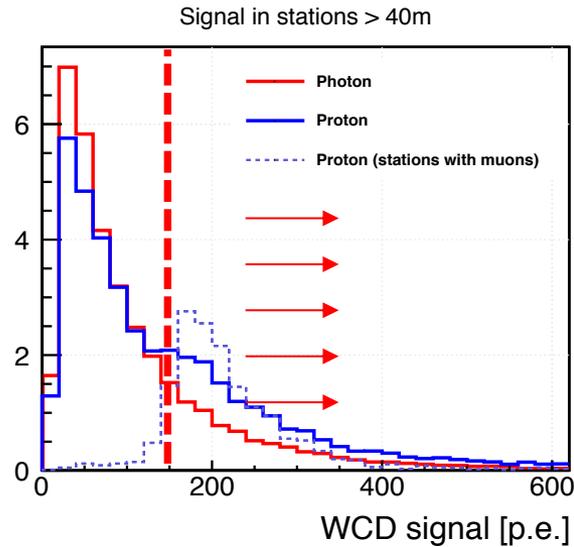
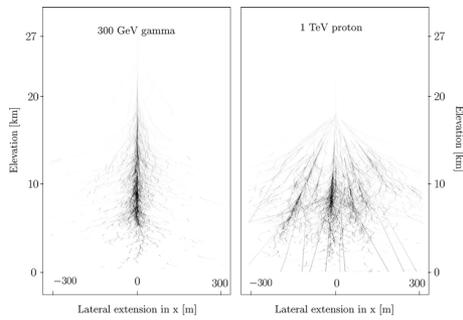


- Fit the RPC hit times:
 - Take advantage of RPCs **high spatial and time resolution**
 - Fit the shower geometry using a shower conic front model
 - Good angular resolution for all events reconstructed inside the array;

- Resolution better than 1.5° for $E > 100$ GeV

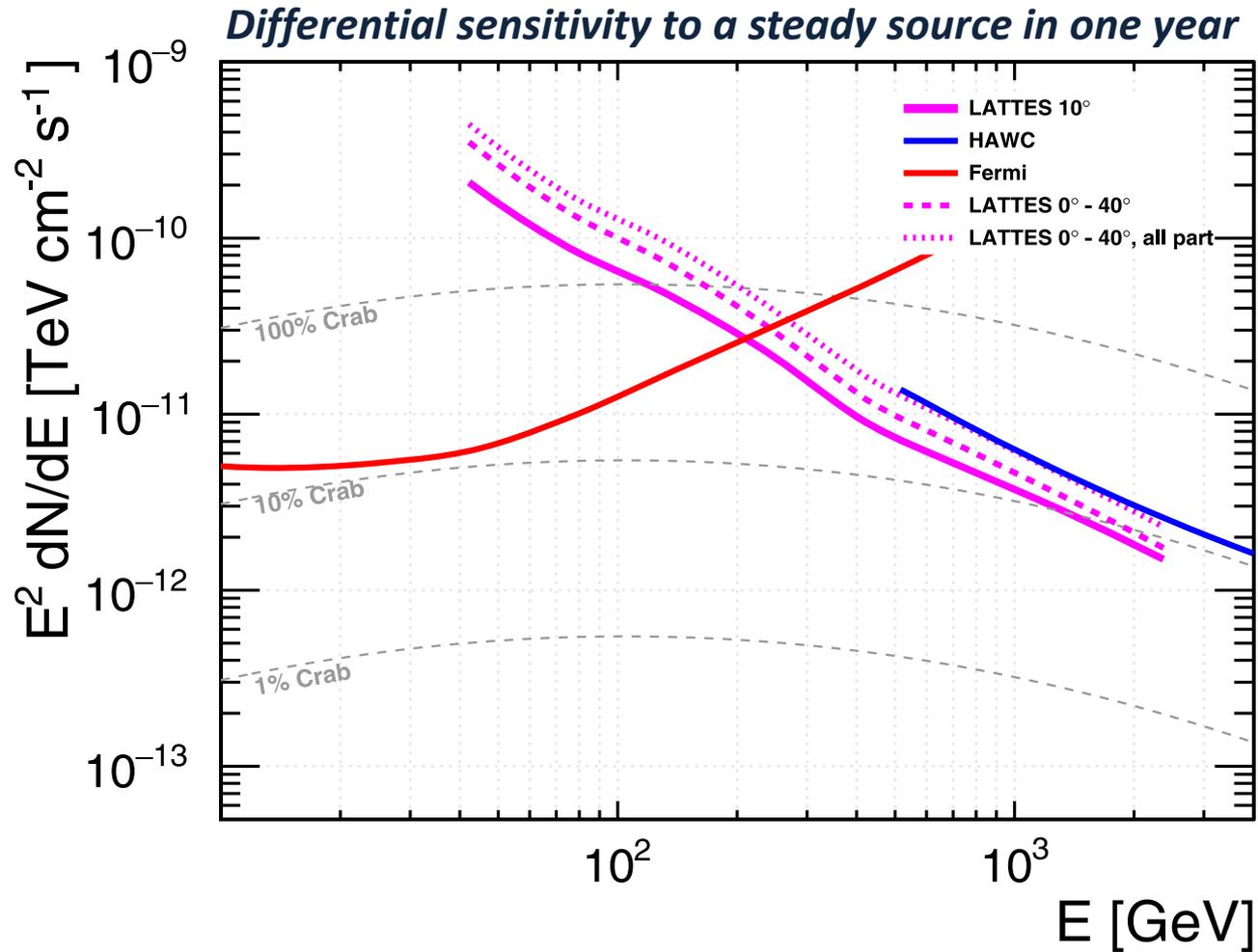


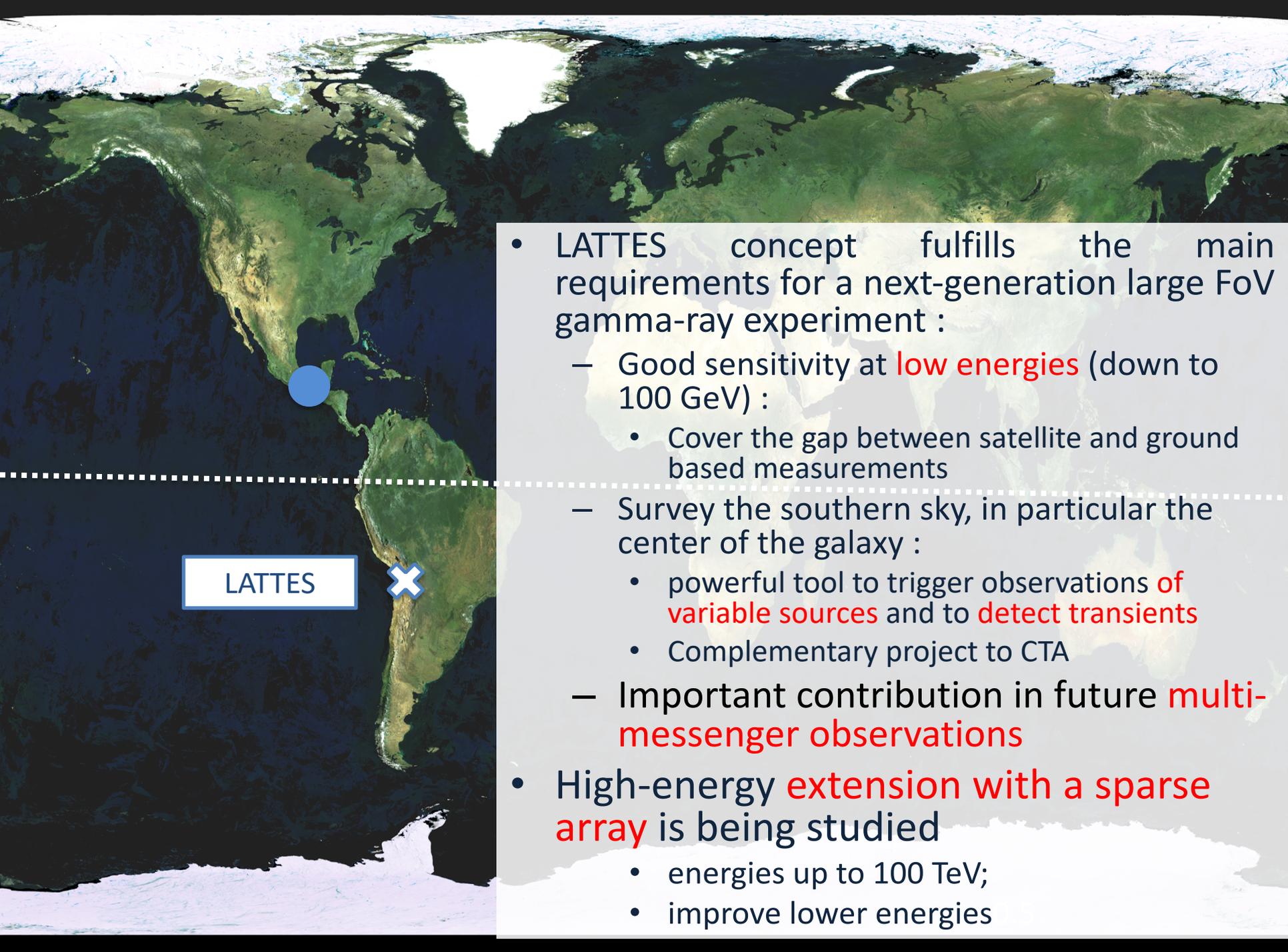
Gamma/hadron discrimination



- Explore shower footprint on ground:
 - Use info from WCD only;
 - Look for high p_t sub-showers and/or muons;
 - Shower **compactness** discrimination variable
 - LDF of gamma showers is more steep than the LDF of hadron showers
 - Fisher discriminant analysis to combine variables
- Room for improvements

LATTES sensitivity





LATTES

- LATTES concept fulfills the main requirements for a next-generation large FoV gamma-ray experiment :
 - Good sensitivity at **low energies** (down to 100 GeV) :
 - Cover the gap between satellite and ground based measurements
 - Survey the southern sky, in particular the center of the galaxy :
 - powerful tool to trigger observations of **variable sources** and to **detect transients**
 - Complementary project to CTA
 - Important contribution in future **multi-messenger observations**
- High-energy **extension with a sparse array** is being studied
 - energies up to 100 TeV;
 - improve lower energies

Acknowledgements



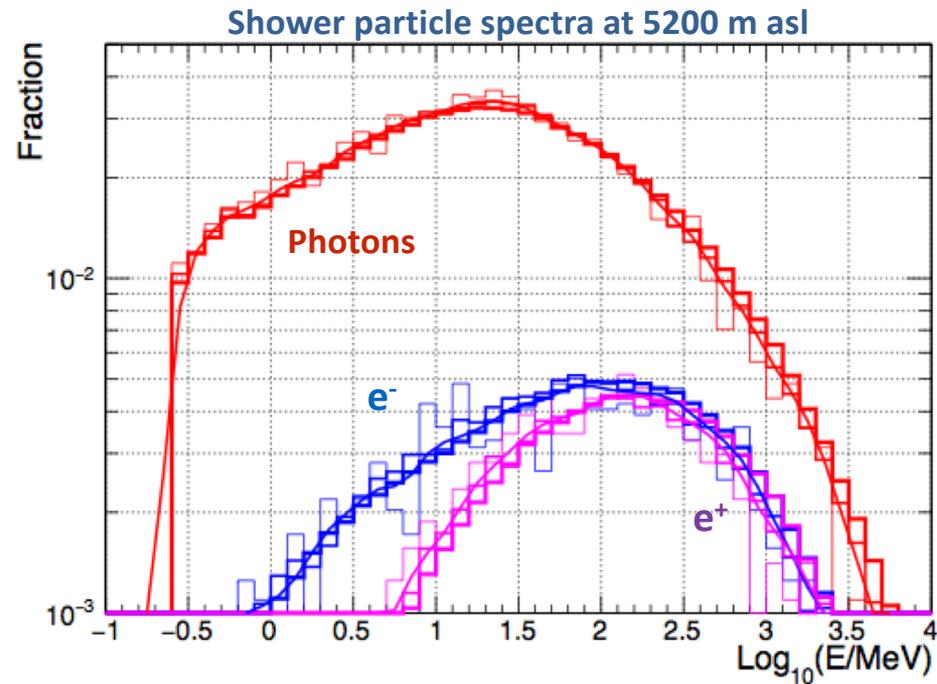
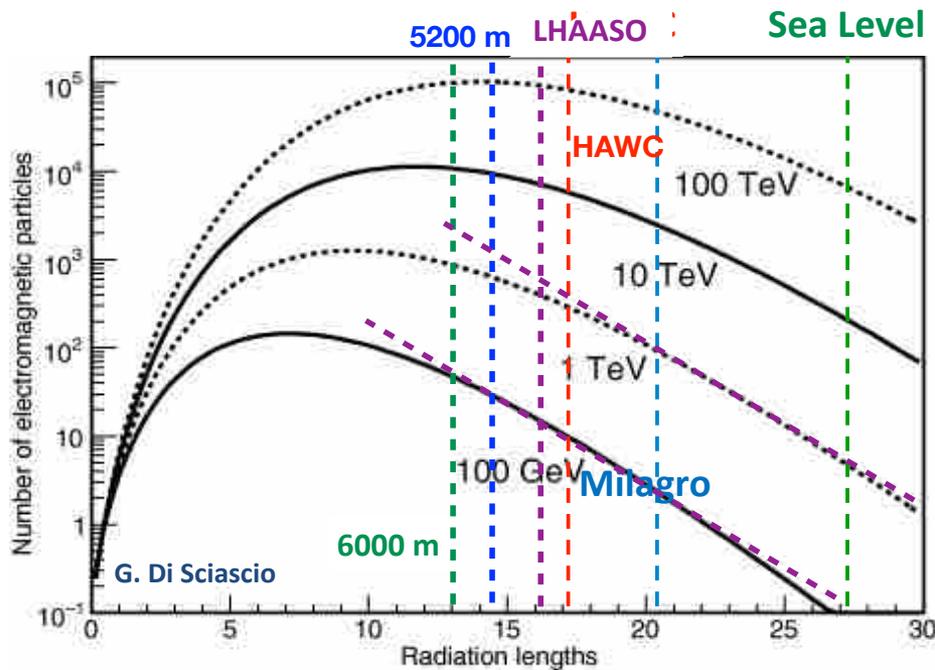
REPÚBLICA
PORTUGUESA



TÉCNICO
LISBOA

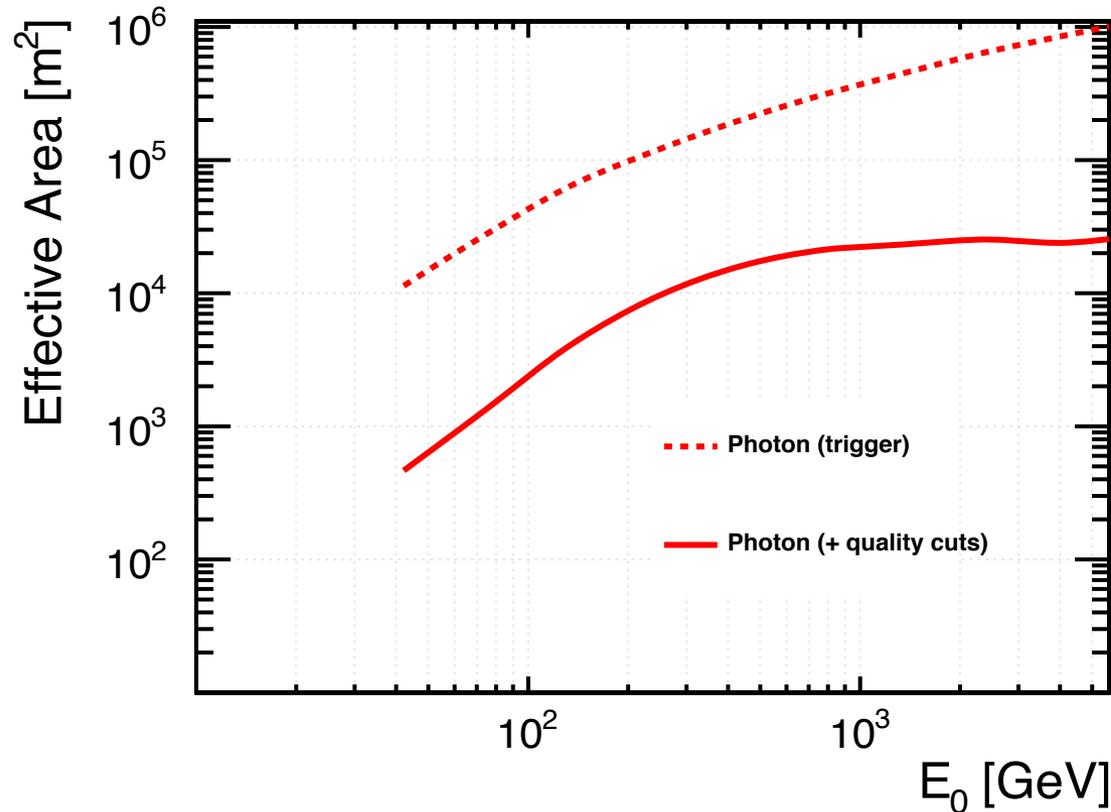
BACKUP SLIDES

The low energy challenge



- Need to:
 - Go to high altitude;
 - Convert the shower photons;
 - Measure energy flow.

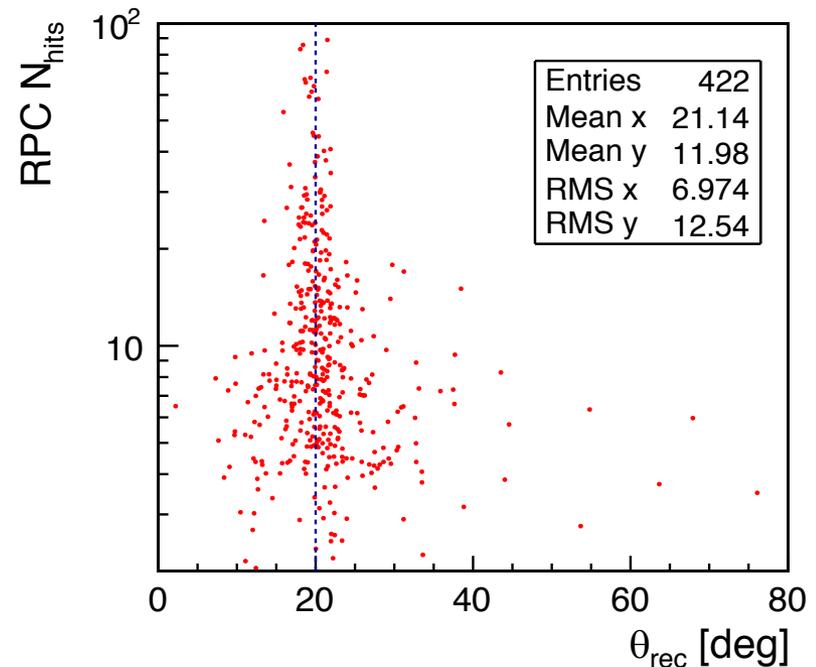
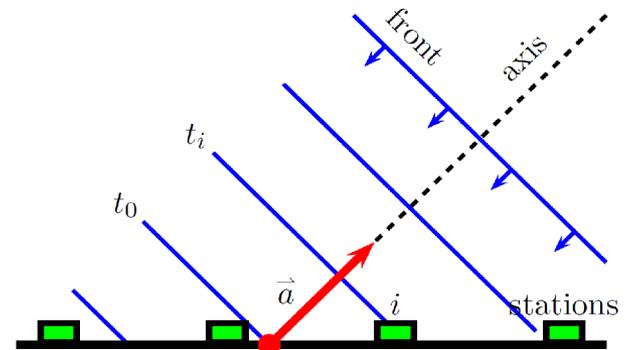
Effective Area



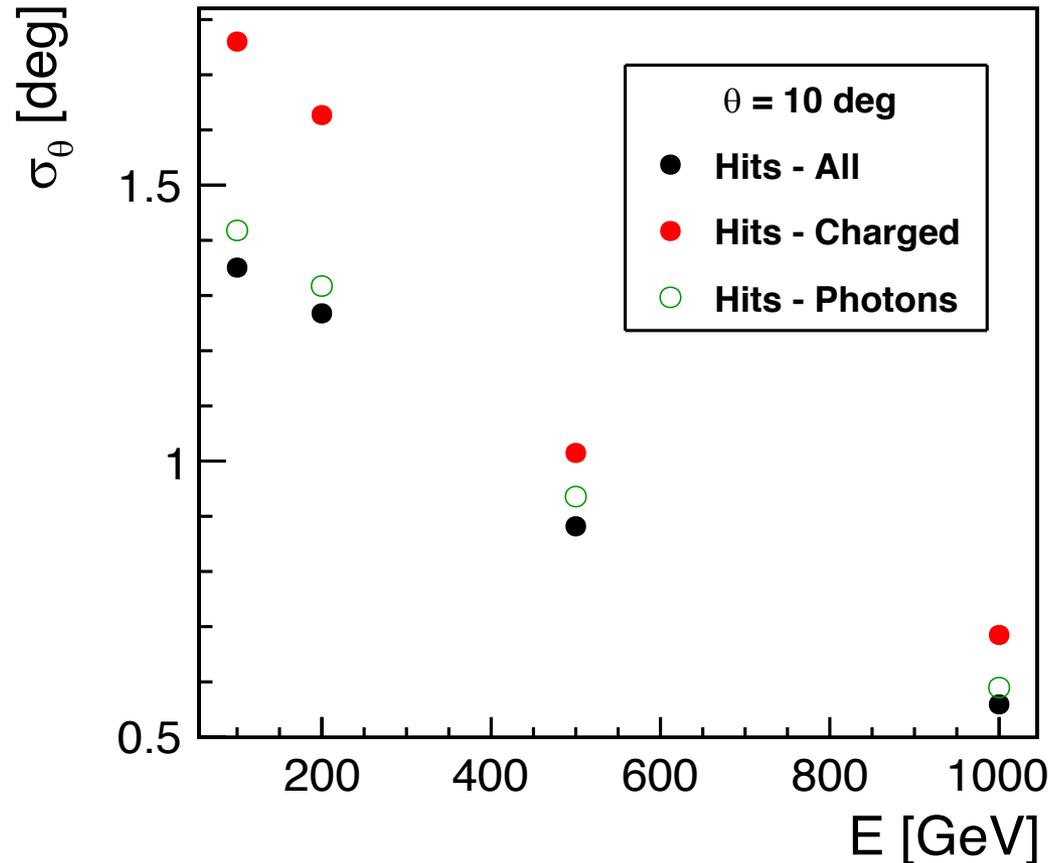
Even applying all quality cuts LATTES
gets an effective area of $\sim 1000 \text{ m}^2$ for $E = 100 \text{ GeV}$

Reconstruction of shower geometry

- **Use RPC hit time** information to reconstruct the shower
 - Take advantage of **high spatial and time resolution**
- Shower geometry reconstruction:
 - Use **shower front plane approximation**
 - Analytical procedure
 - Apply trigger conditions
 - Apply **cut** on the **number of registered hits** by the RPCs

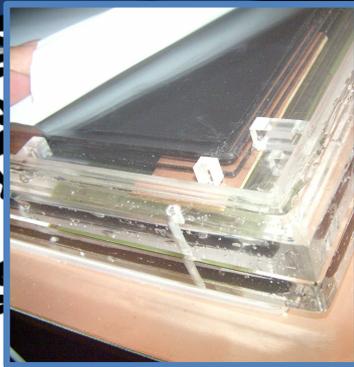


Contributions to the geometric reconstruction



- **Photons** retain a **higher correlation** with the **shower geometry** than charged particles
- Could we measure photons with the RPC instead?

Ongoing developments and tests on RPCs, electronics and read-out systems



RPC developments



DAQ Engineering prototype

RPC based muon hodoscope for precise studies of the Auger WCD



Top RPC

Gianni Navarra WCD

Bottom RPC

Construction and Assembling



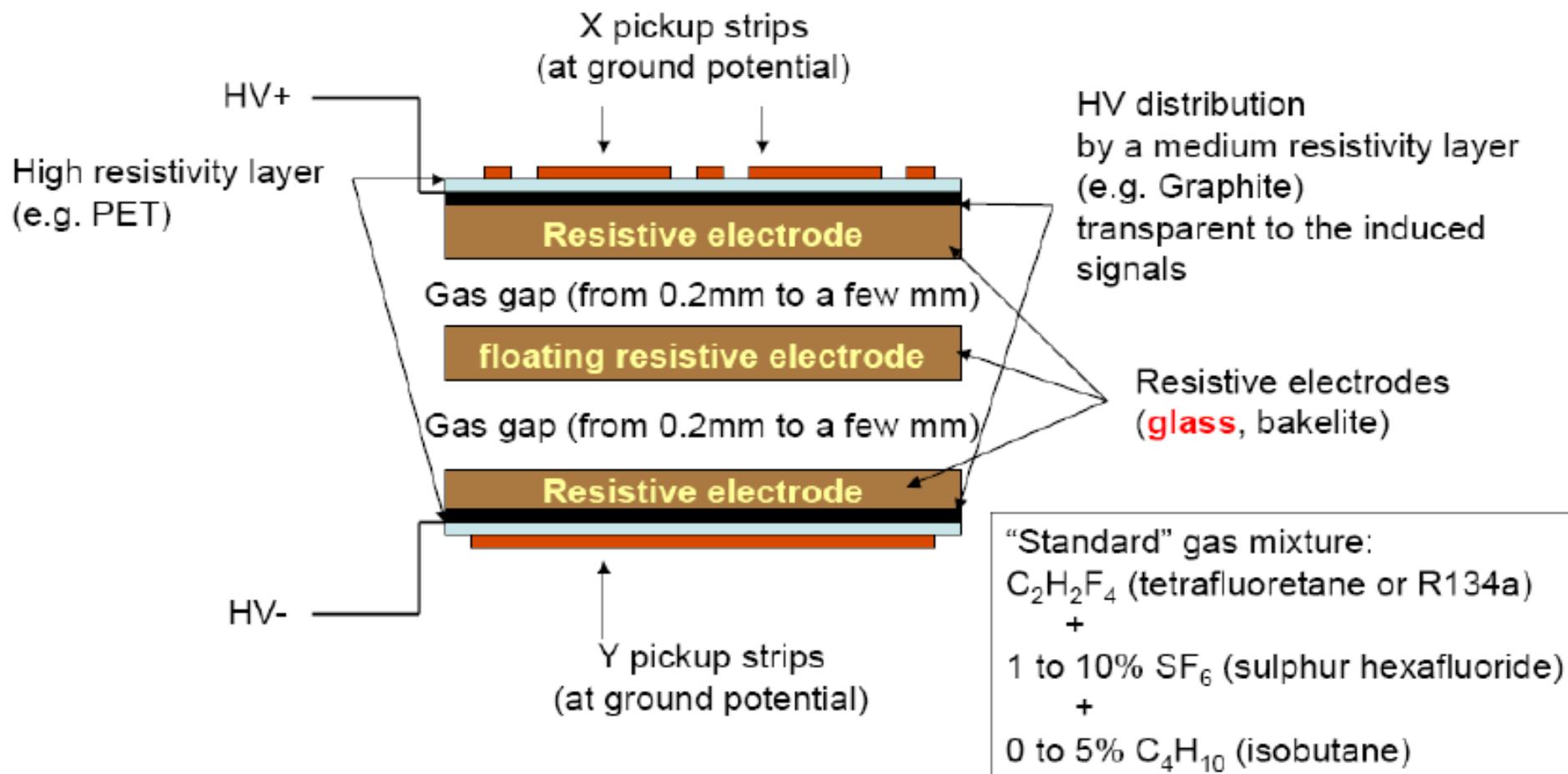
RPC hodoscope

RPCs in the field @ Auger



RPCs – basic structure

Many variations allowed



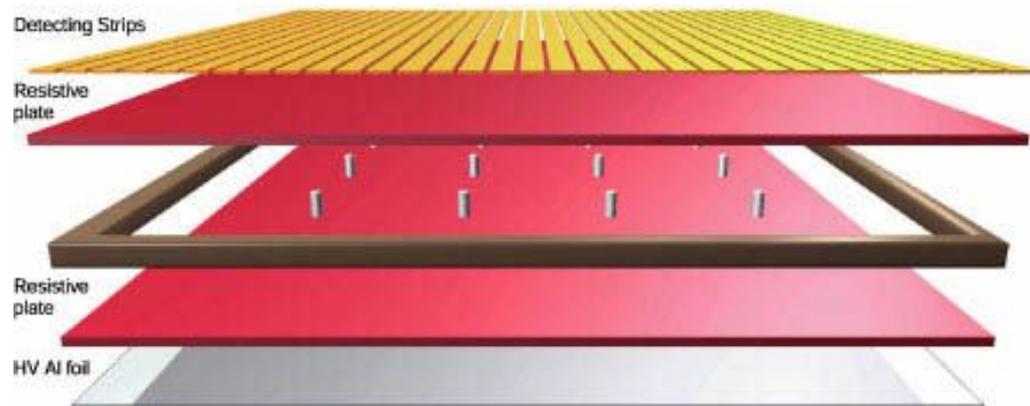
The current is limited by the resistive electrodes: no sparks by construction

↳ **very safe detector, although limited to low particle rates ($\sim 2\text{kHz}/\text{cm}^2$)**

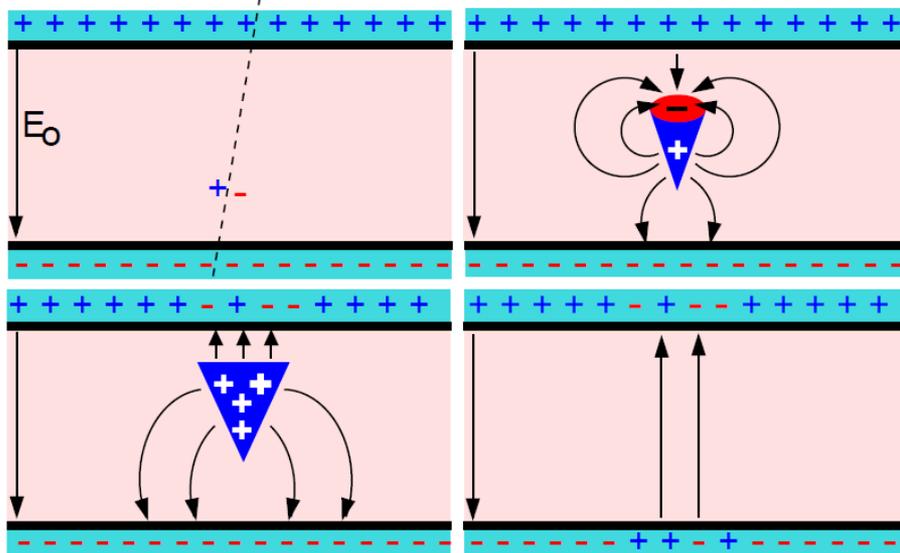
↳ **excellent efficiency (99%), time ($\sim 50\text{ ps}$) and position resolution ($\sim 100\mu\text{m}$)**

RPCs Resistive Plate Chamber

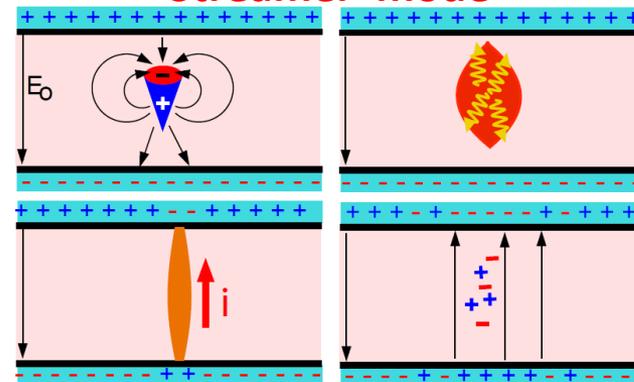
- Gaseous detector
- Planar geometry
- uniform electrical field imposed.
- High resistive plates in between the electrodes limit the avalanche current.
- Signal is picked up by the induction of the avalanche in the readout pads.



Avalanche mode

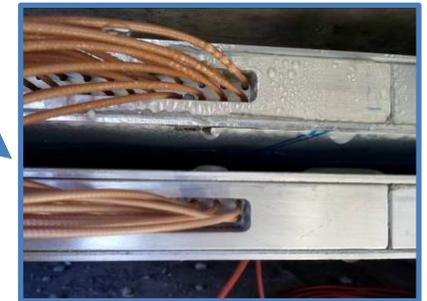
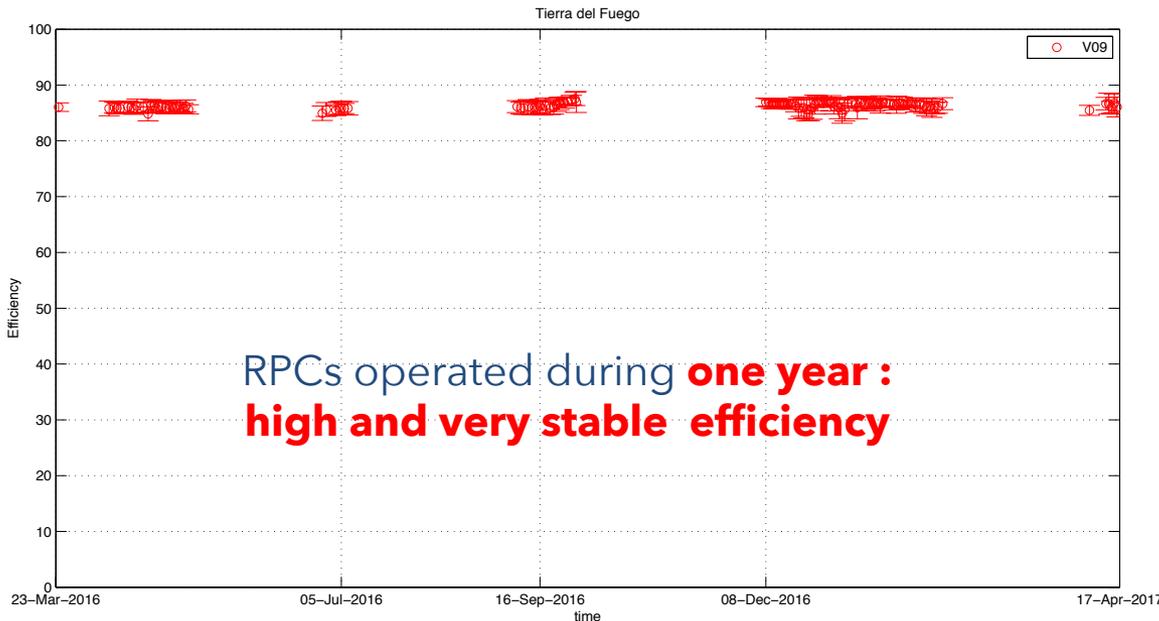
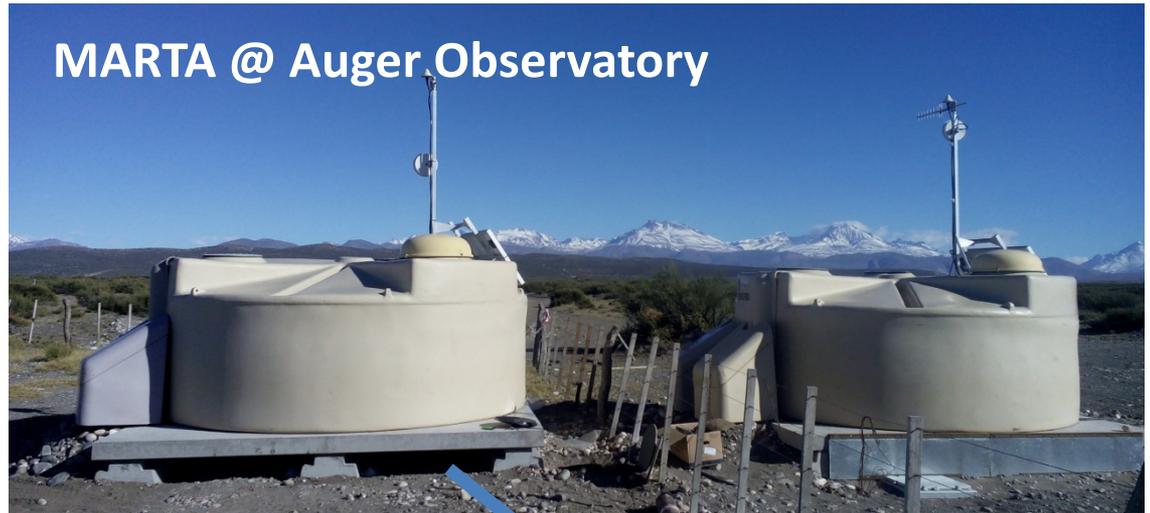


Streamer mode



Outdoor field operation

- Several RPCs operating in the Pampa for more than one year

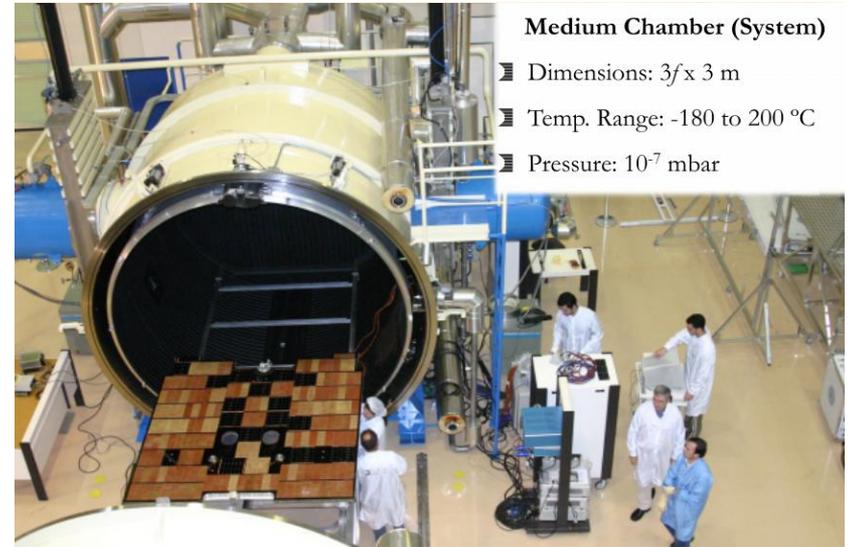


RPCs under the WCD

Low pressure tests

São Paulo, Brazil

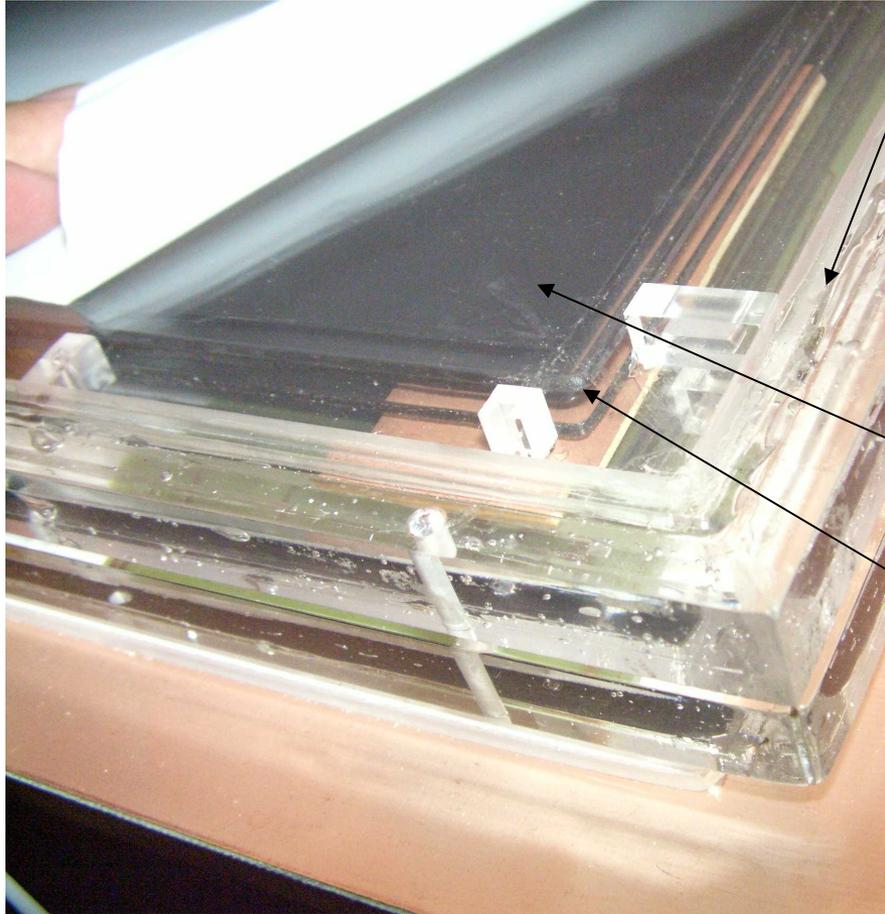
National Institute for Spatial Research



Medium Chamber (System)

- ▮ Dimensions: 3f x 3 m
- ▮ Temp. Range: -180 to 200 °C
- ▮ Pressure: 10^{-7} mbar

CONSTRUCTION DETAILS



Signal-transparent and nice-looking acrylic box, 1mm thick covers

Permanently glued

RPC fits tightly inside

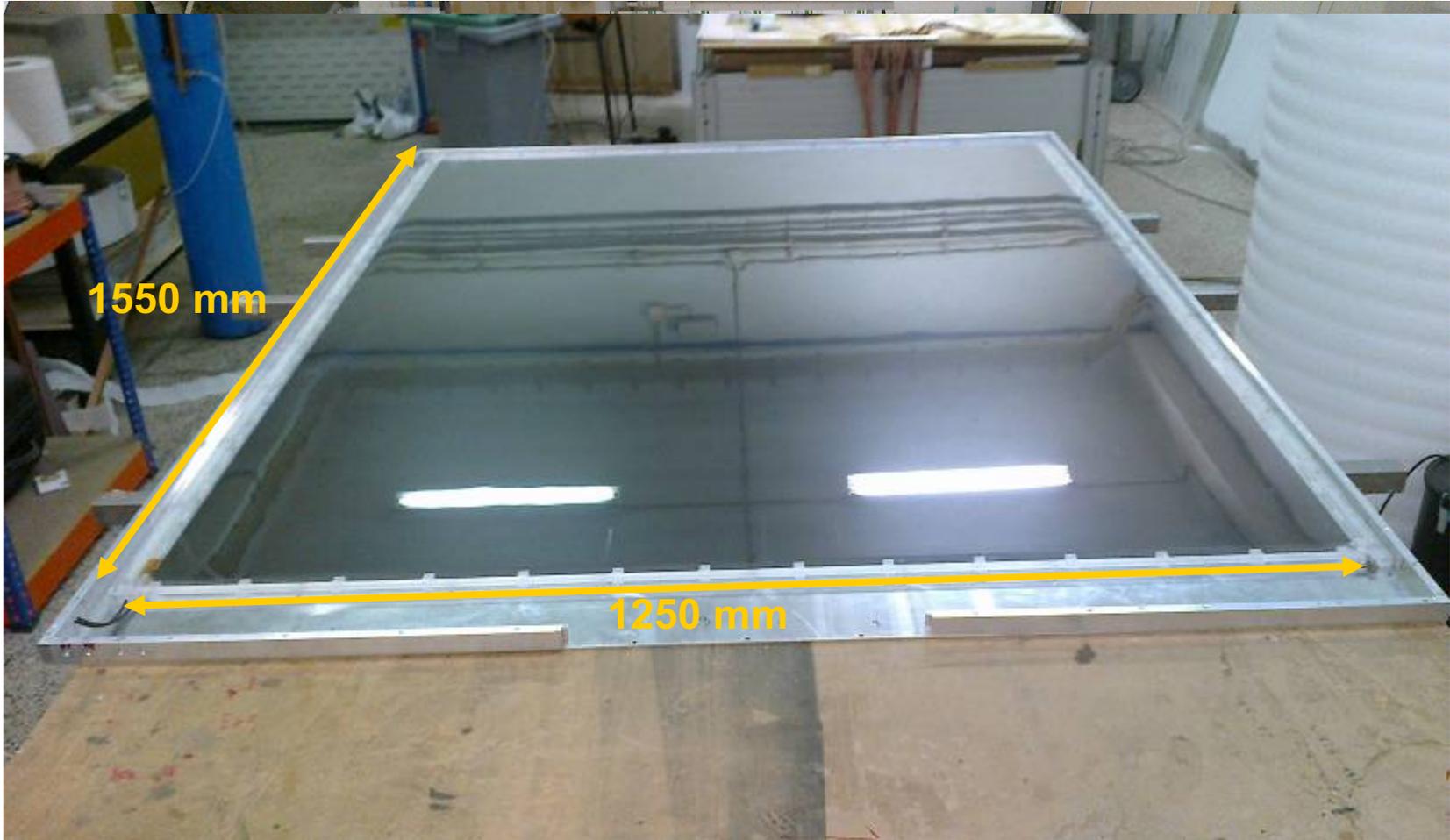
- ✓ good electrode support mechanics
- ✓ excellent HV insulation
- ✓ excellent gas tightness

HV layer, also signal-transparent

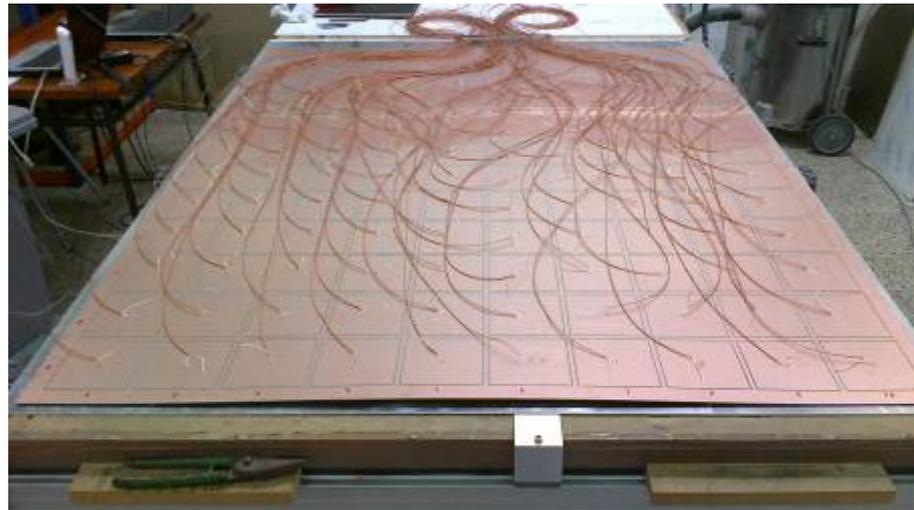
3 RPC glasses (2mm soda-lime)

External pickup electrodes

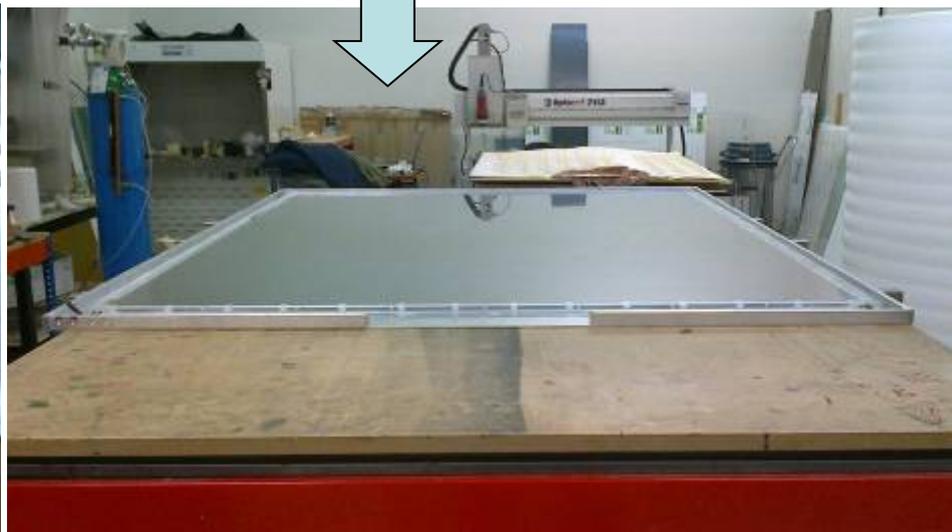
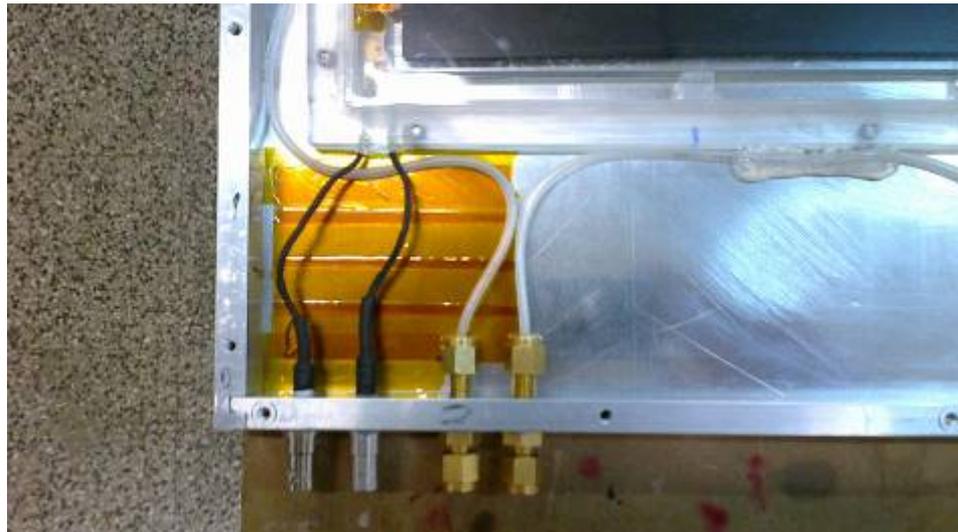
CONSTRUCTION DETAILS. RPC and gas volume



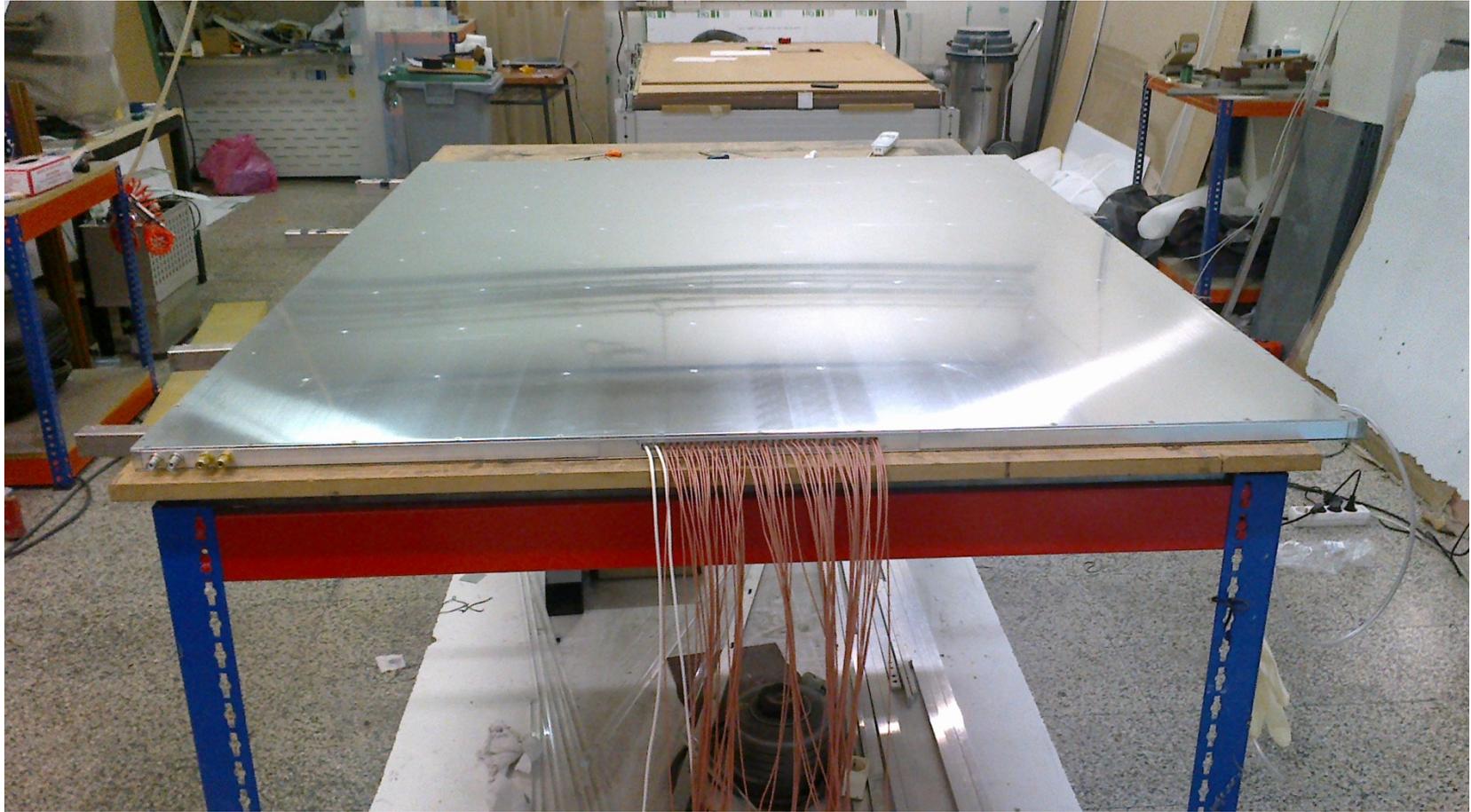
CONSTRUCTION DETAILS. Readout 8x8 pads



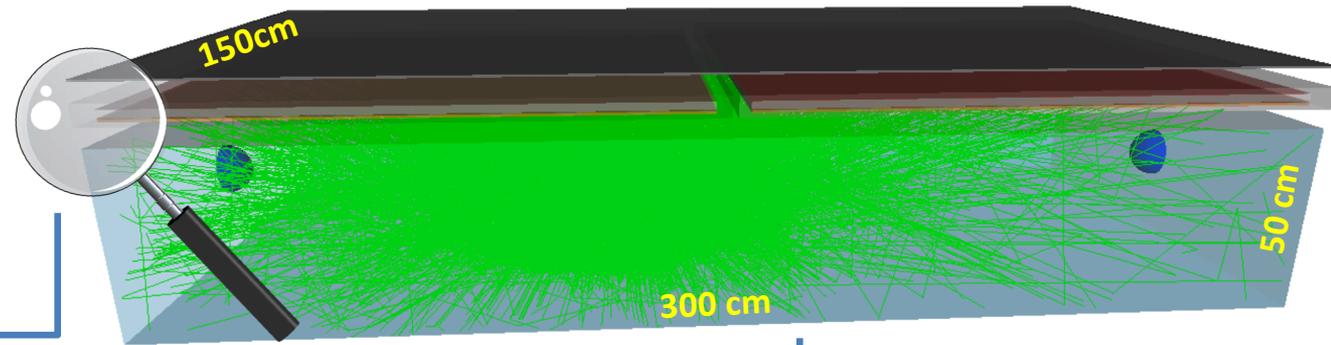
CONSTRUCTION DETAILS. Assembly



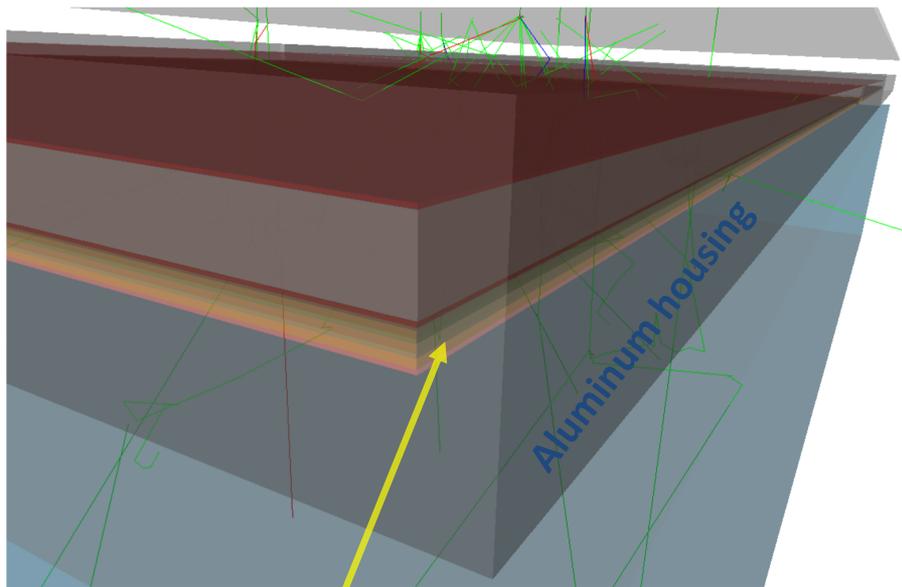
CONSTRUCTION DETAILS. RPC enclosed



LATTES station in Geant4



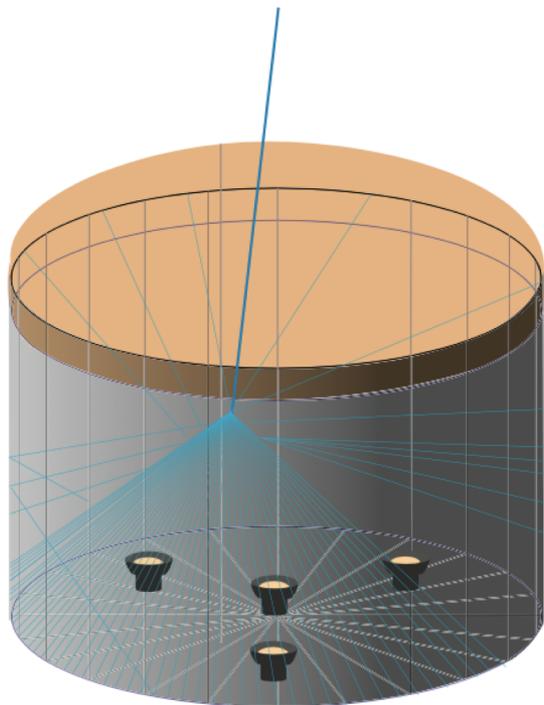
Detailed RPC structure



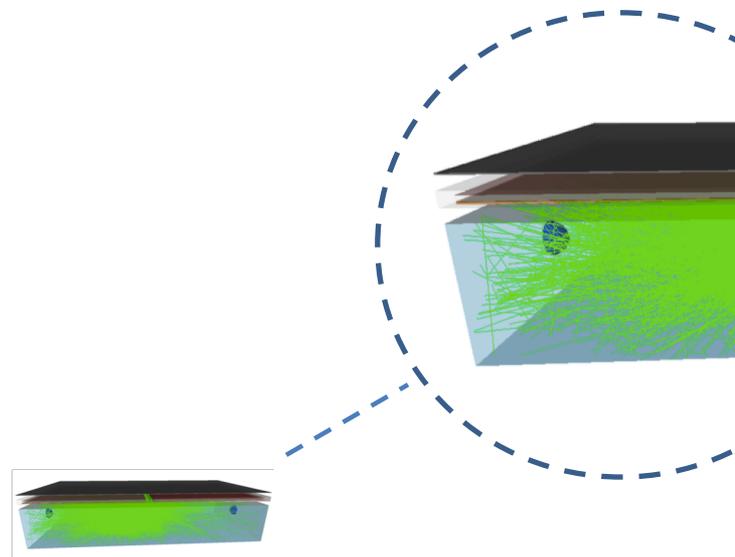
Acrylic box with glass electrodes and 1 mm gas gaps

- Explore **Geant4 capabilities** to simulate **optical photon propagation**;
- λ dependence of all relevant processes/materials taken into account;
- **Water**
 - Attenuation length ~ 80 m @ $\lambda = 400$ nm
- **PMT**
 - $Q.E._{\max} \sim 30\%$ @ $\lambda = 420$ nm
- **Tyvek**
 - Described using the **G4 UNIFIED optical model**;
 - Specular and diffusive properties;
 - $R \sim 95\%$, for $\lambda > 450$ nm

Station: HAWC vs LATTES



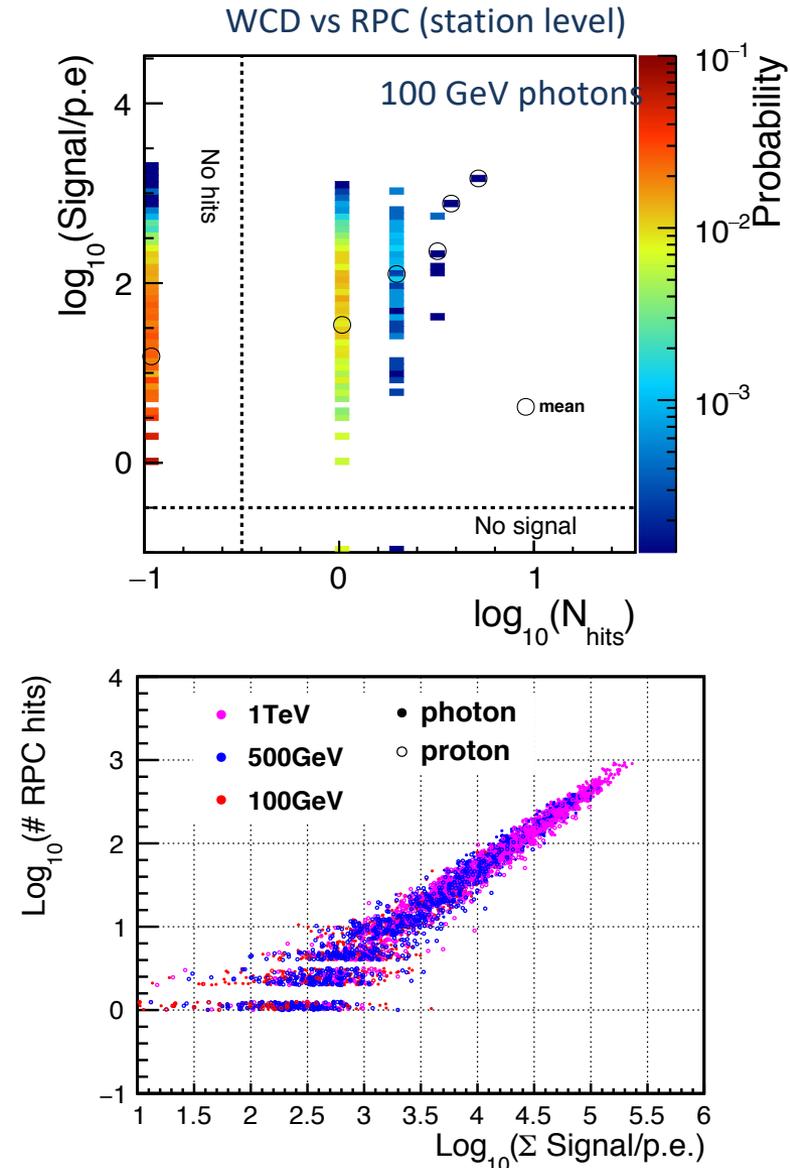
HAWC
(present detector)



LATTES
(next generation)

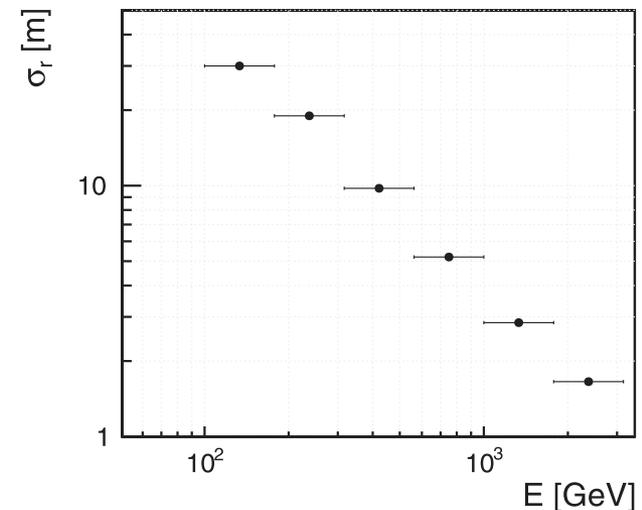
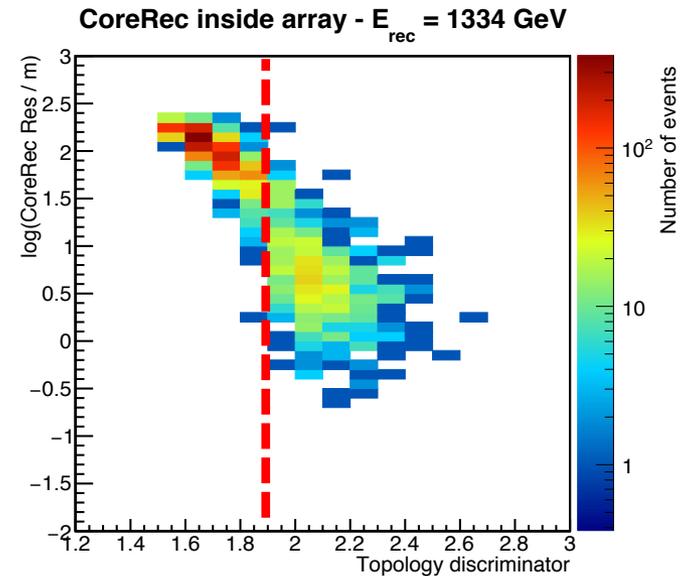
LATTES: complementarity

- Combined detection:
 - Lower the **energy threshold**
 - Improve the trigger conditions (WCD)
 - Enable detector **inter-calibrations**
 - Energy calibration can be used to **control detector systematic uncertainties**
 - Check Monte Carlo simulations performance
 - Enhance **gamma/hadron discrimination**
 - Explore shower characteristics
 - Access to combined Argo/HAWC discrimination techniques

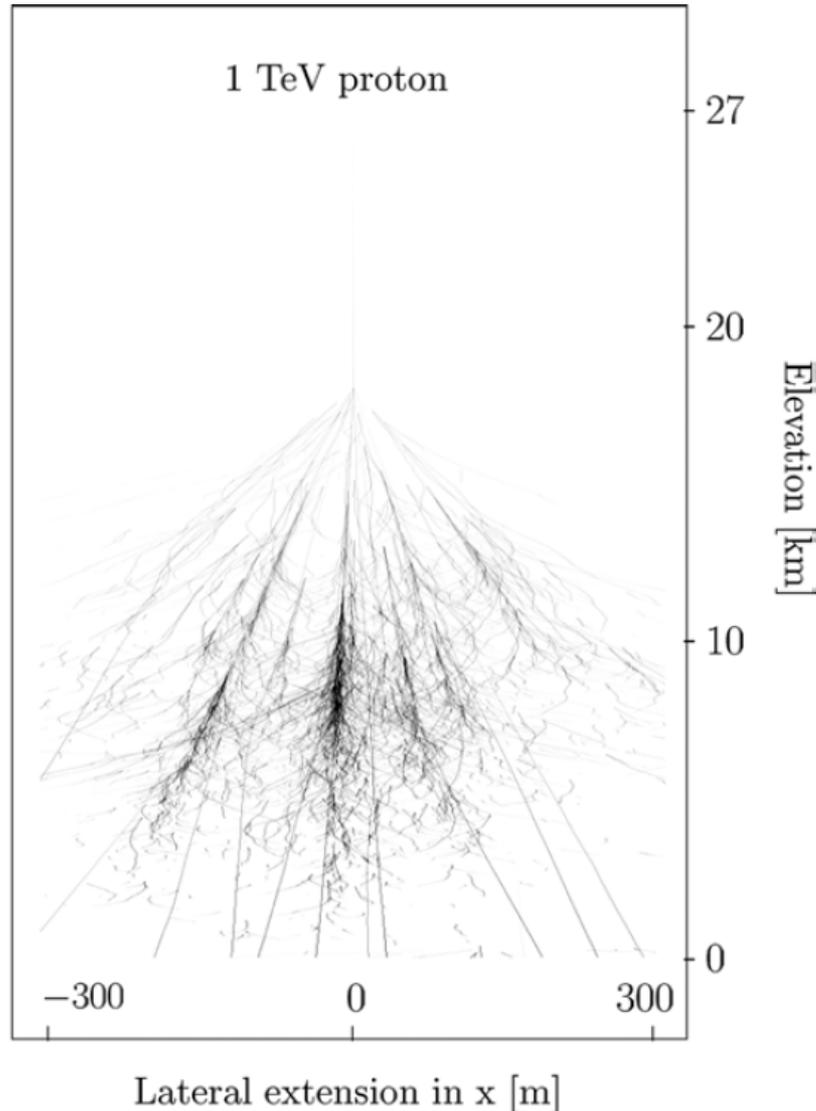


Shower core reconstruction

- Test whether the shower is inside/outside the array
 - Explore LDF topology
 - Is maximum observed inside of array?
 - Currently exploring the quality of the fit
 - Fixed cut for all energies
- Resolution better than 10 meters for showers above 300 GeV



Strategies for primary discrimination



- **Hit pattern at ground**
 - Hits from hadronic showers are more sparse than in gamma induced showers
 - RPC detectors
 - Explored by the ARGO collaboration
- **Search for energetic clusters far from the shower core**
 - Present only in hadronic showers
 - Water Cherenkov Detectors
 - Explored by the HAWC collaboration
- ***Combine both strategies using an hybrid detector: LATTES***
 - *Work on-going...*

High-energy discrimination strategy

- Get the photon average LDF for each reconstructed energy bin
- Fit the average LDF to each single event
 - Absorb the normalization factor
- Compute the shower **compactness**
 - Event LDF “distance” to the photon average LDF

$$\text{Compactness} = \log_{10} \left(\sum_i^n |\langle LDF \rangle (x_i) - y(x_i)| \right)$$