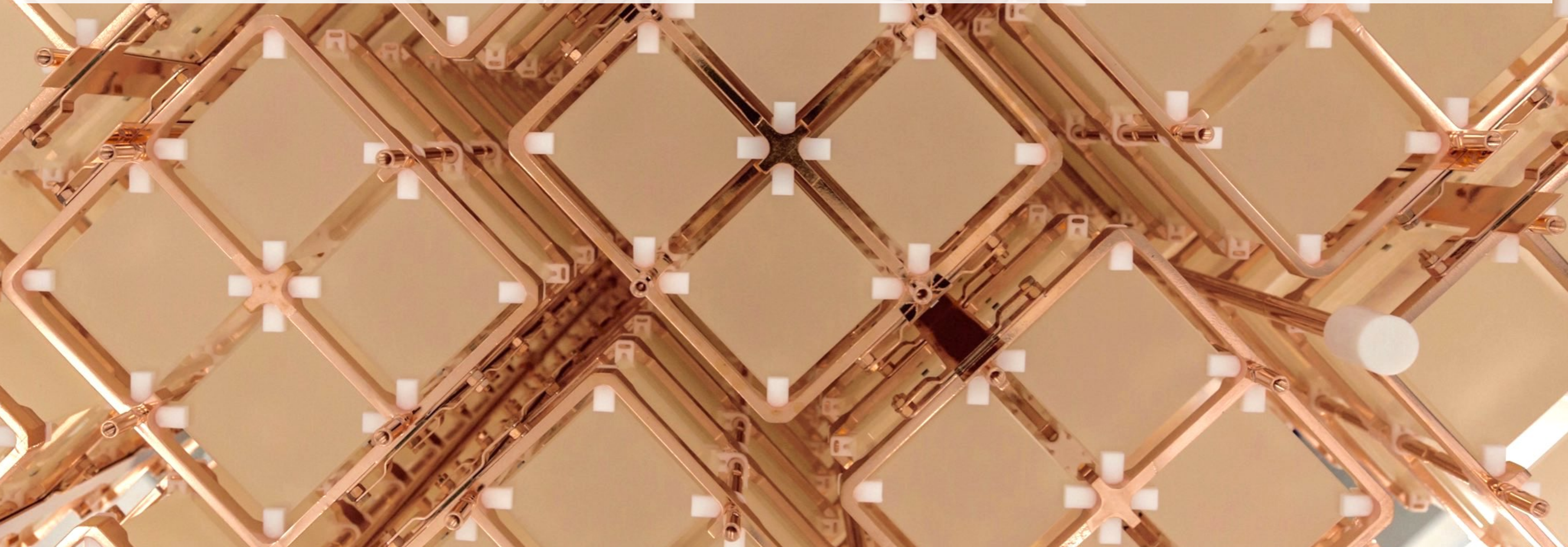


# Results from the CUORE experiment

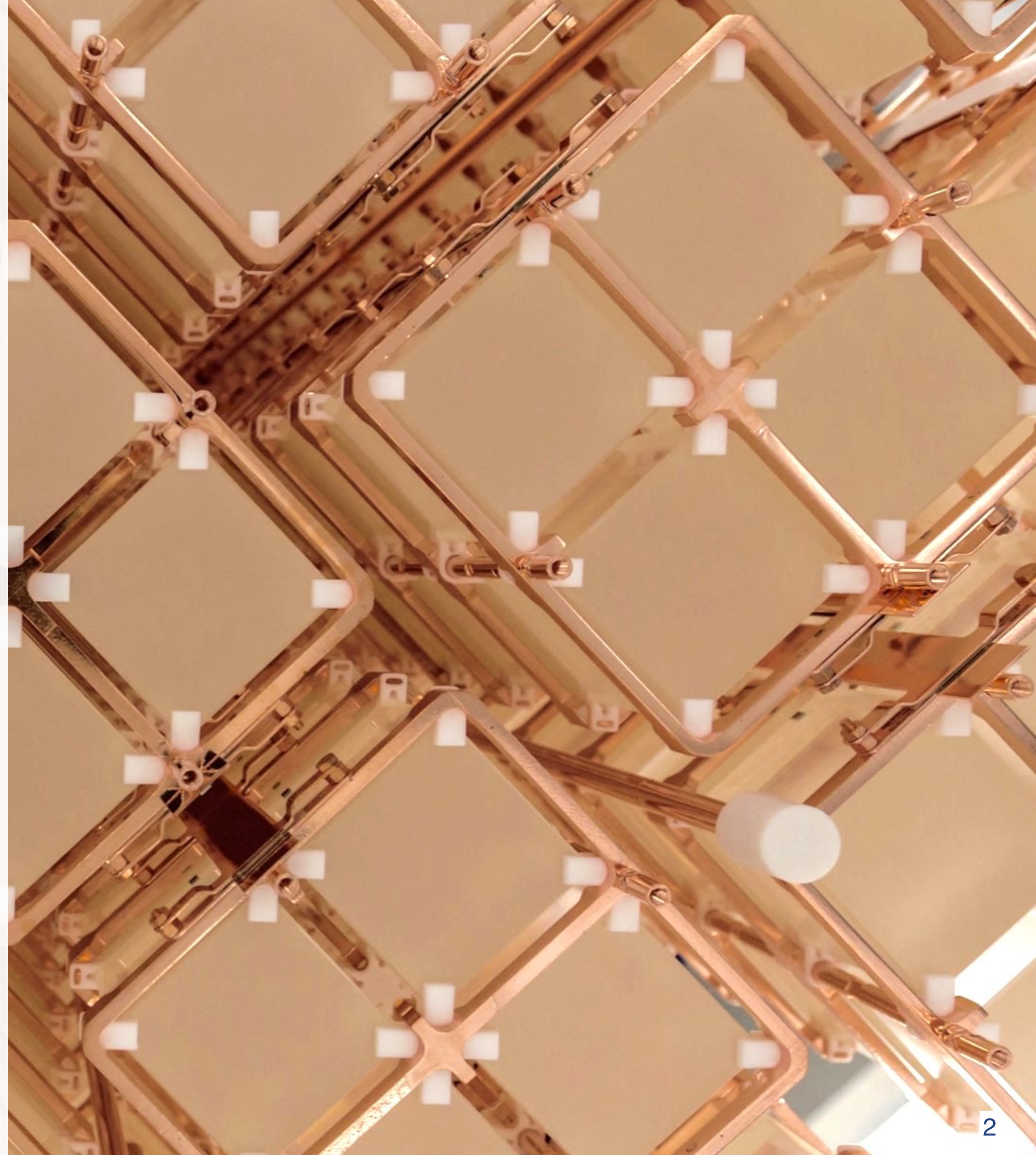


Matteo Biassoni on behalf of the **CUORE** Collaboration  
INFN - Sez. Milano Bicocca



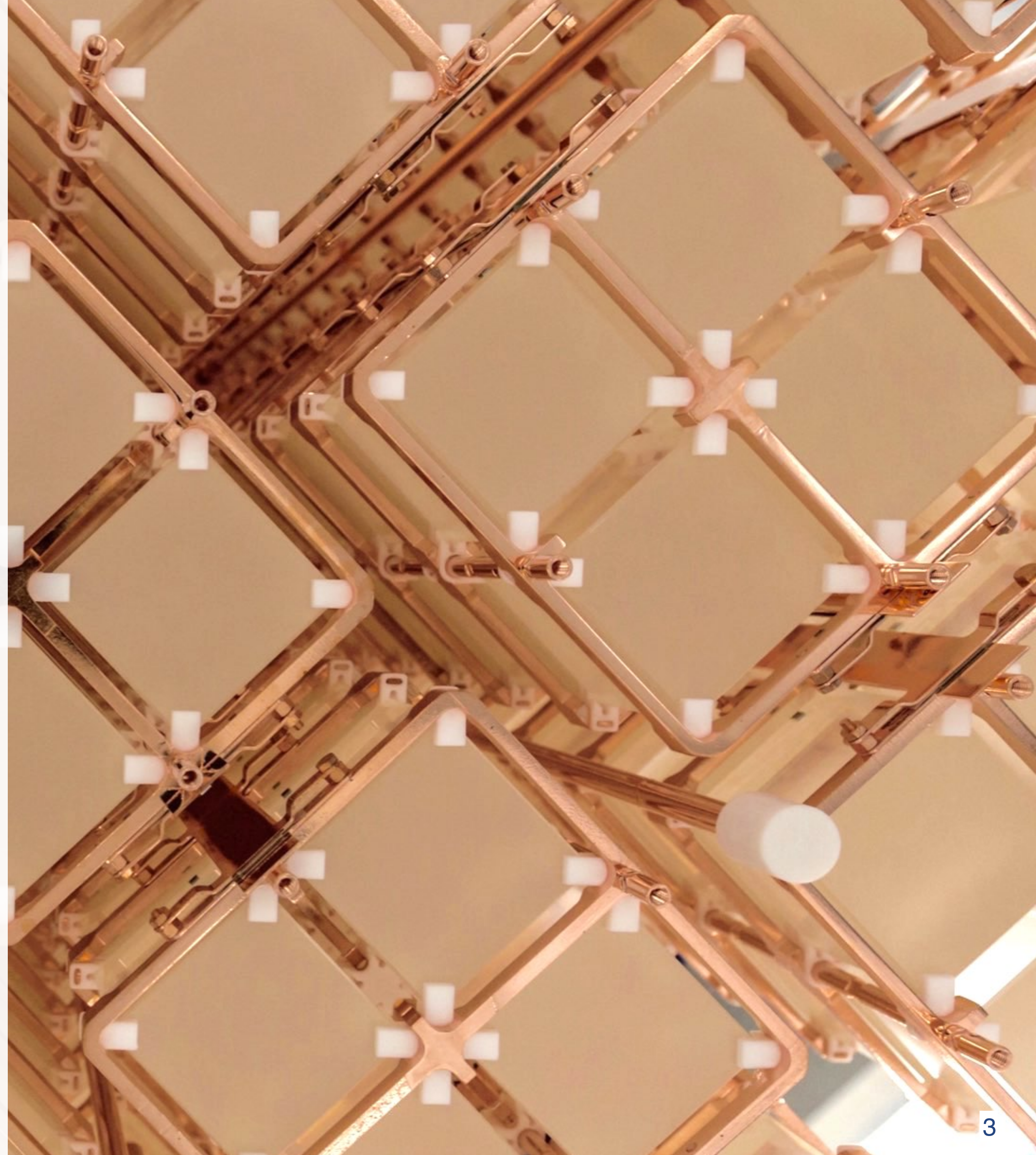
# Outline

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  - $2\nu\beta\beta$
- Conclusions and outlook



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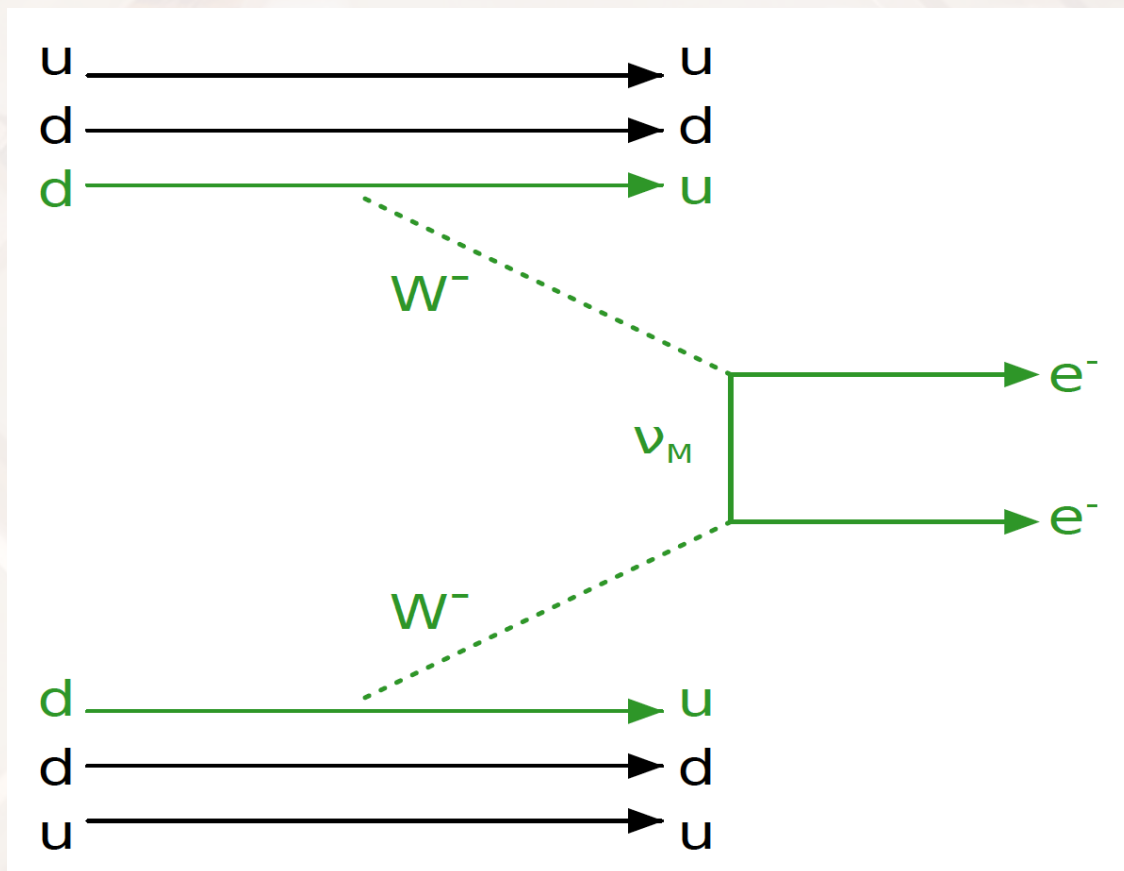


# TeO<sub>2</sub> and thermal detectors for 0νDBD

Second order nuclear process, alternative to beta decay forbidden by mass difference for some even-even nuclei

$$(A, Z) \rightarrow (A, Z + 2) + 2e^- + 2\bar{\nu}_e \quad \text{2nd order SM process, } T_{1/2} \sim 10^{18\sim 24} \text{ years}$$

$$(A, Z) \rightarrow (A, Z + 2) + 2e^-$$



- SM forbidden,  $\Delta L = 2$
- **if observed, then** neutrino is a Majorana particle
- underlying mechanism can give insight into beyond SM physics
  - light neutrino mass scale and hierarchy
  - heavy neutrino
  - ...

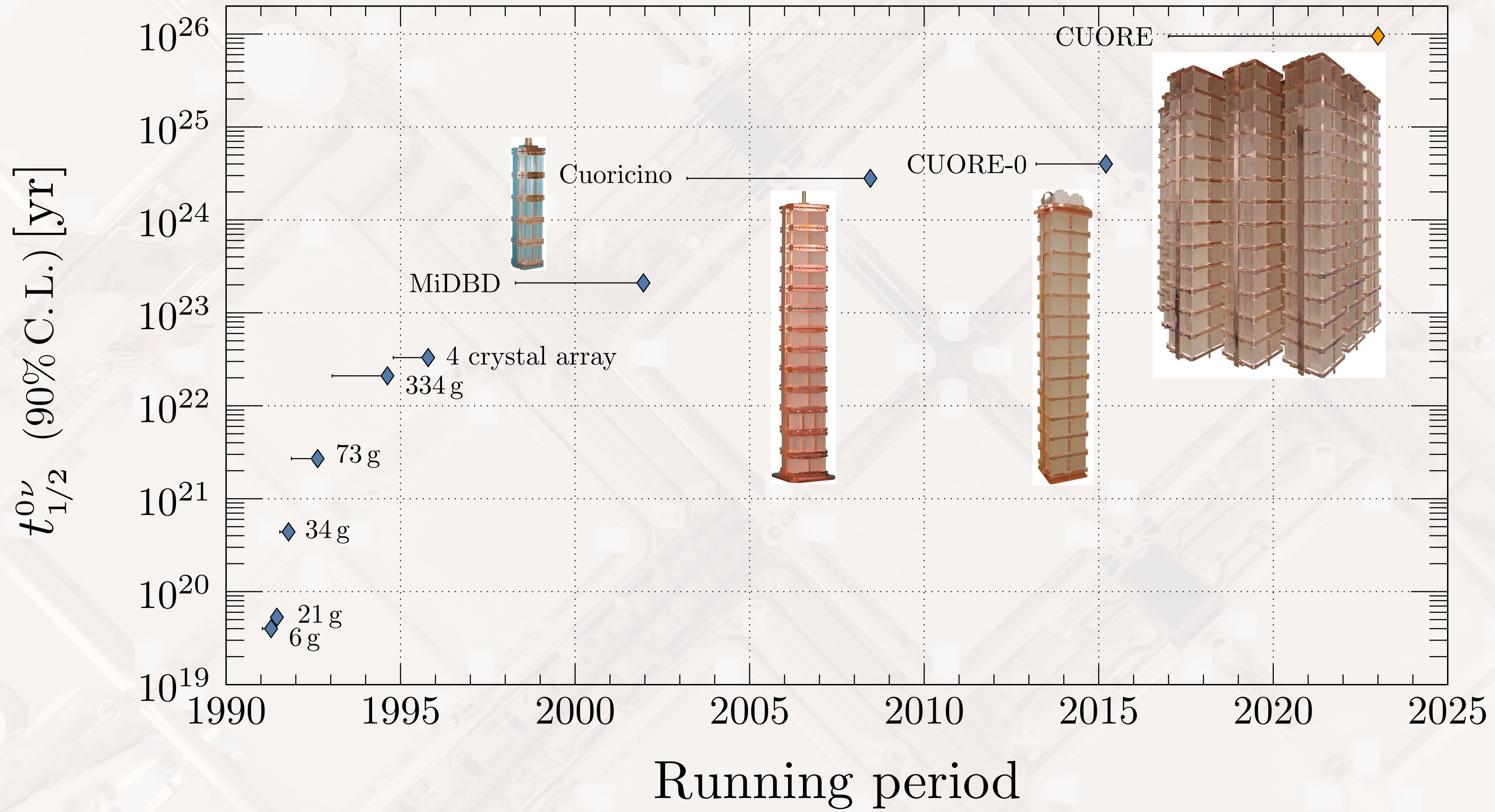
<sup>130</sup>Te is a good candidate source for 0νDBD search:

- high natural isotopic abundance (~34%)
- NME and phase space on average
- Q-value (2528 keV) above most of the natural radioactivity
- easy to mix in convenient chemical compounds (TeO<sub>2</sub>)

**Thermal detectors** are a good choice for 0νDBD search:

- excellent energy resolution
- large active mass and efficiency/unit cost
- fully active source and sensitive volume, no dead-layer

# TeO<sub>2</sub> arrays: state of the art

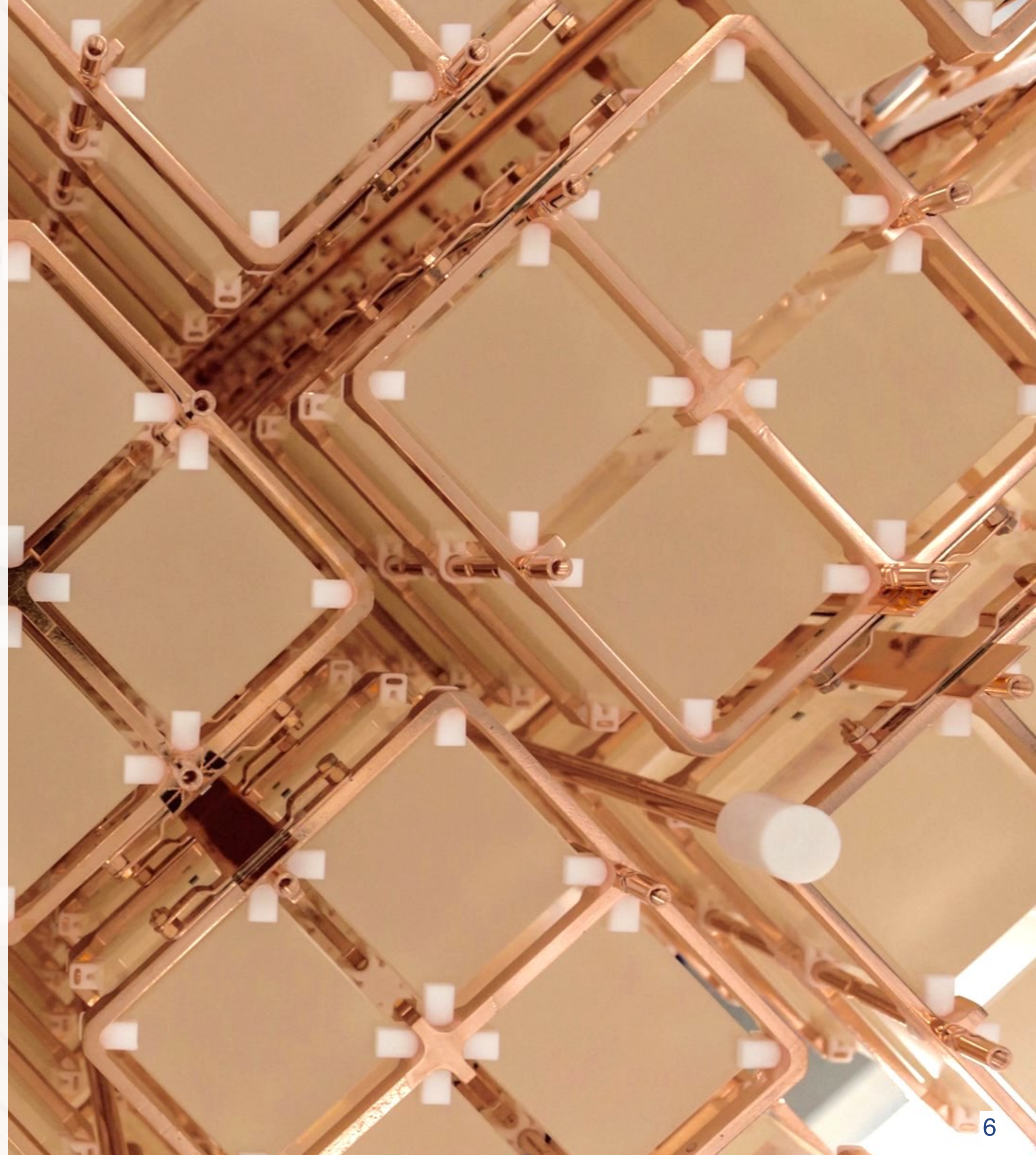


CUORE is the latest in a long progression of TeO<sub>2</sub> detectors which included two large demonstrators:

- Cuoricino (2.8x10<sup>24</sup> y)
- CUORE-0 (4.0x10<sup>24</sup> y combined)

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Cryogenic Underground Observatory for Rare Events

**Primary goal:** search for  $0\nu\beta\beta$  decay in  $^{130}\text{Te}$

**Detector design:**

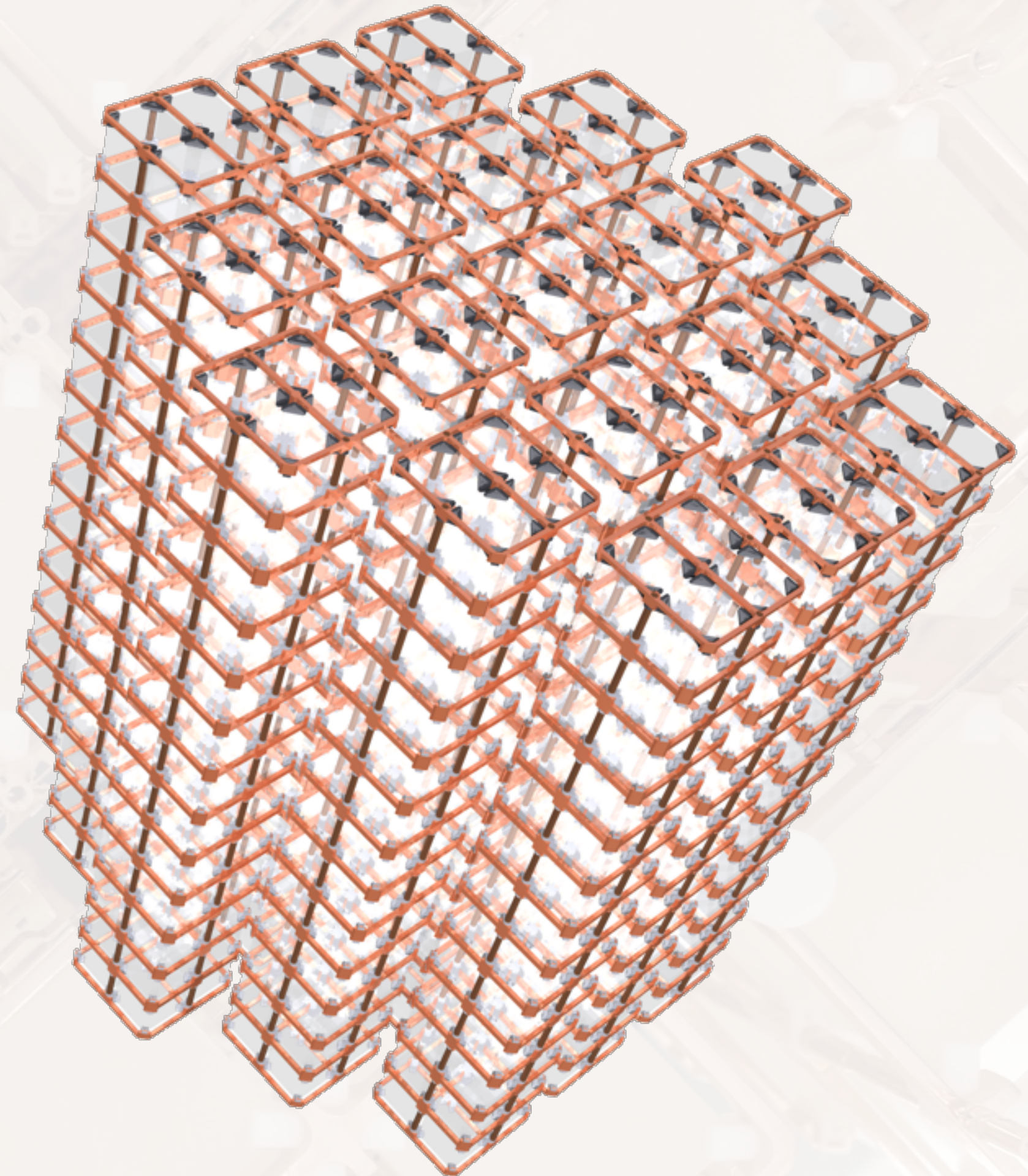
closely packed array of 988  $\text{TeO}_2$  crystals arranged in 19 towers

**Design parameters:**

- mass of  $\text{TeO}_2$ : **742 kg** (206 kg of  $^{130}\text{Te}$ )
- low background aim:  **$10^{-2}$  c/(keV·kg·yr)**
- target energy resolution: **5 keV FWHM** in the Region Of Interest (ROI)
- high granularity
- deep underground location
- strict radio-purity controls on materials and assembly

**CUORE projected sensitivity (5 years, 90% C.L.):**

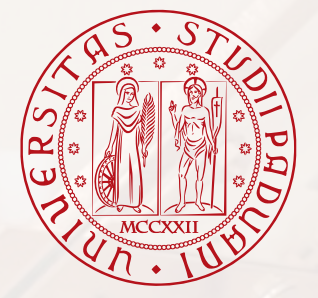
$$T_{1/2} > 9 \times 10^{25} \text{ yr}$$



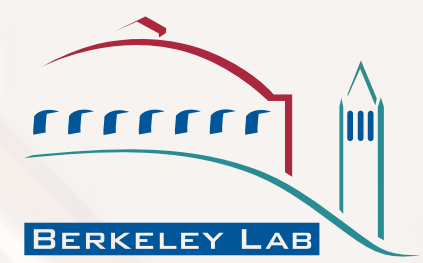
# The CUORE Collaboration



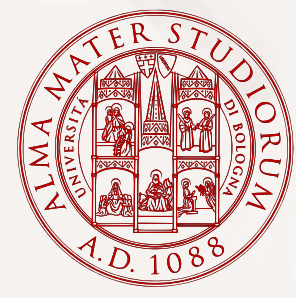
UCLA



Yale



CAL POLY  
SAN LUIS OBISPO



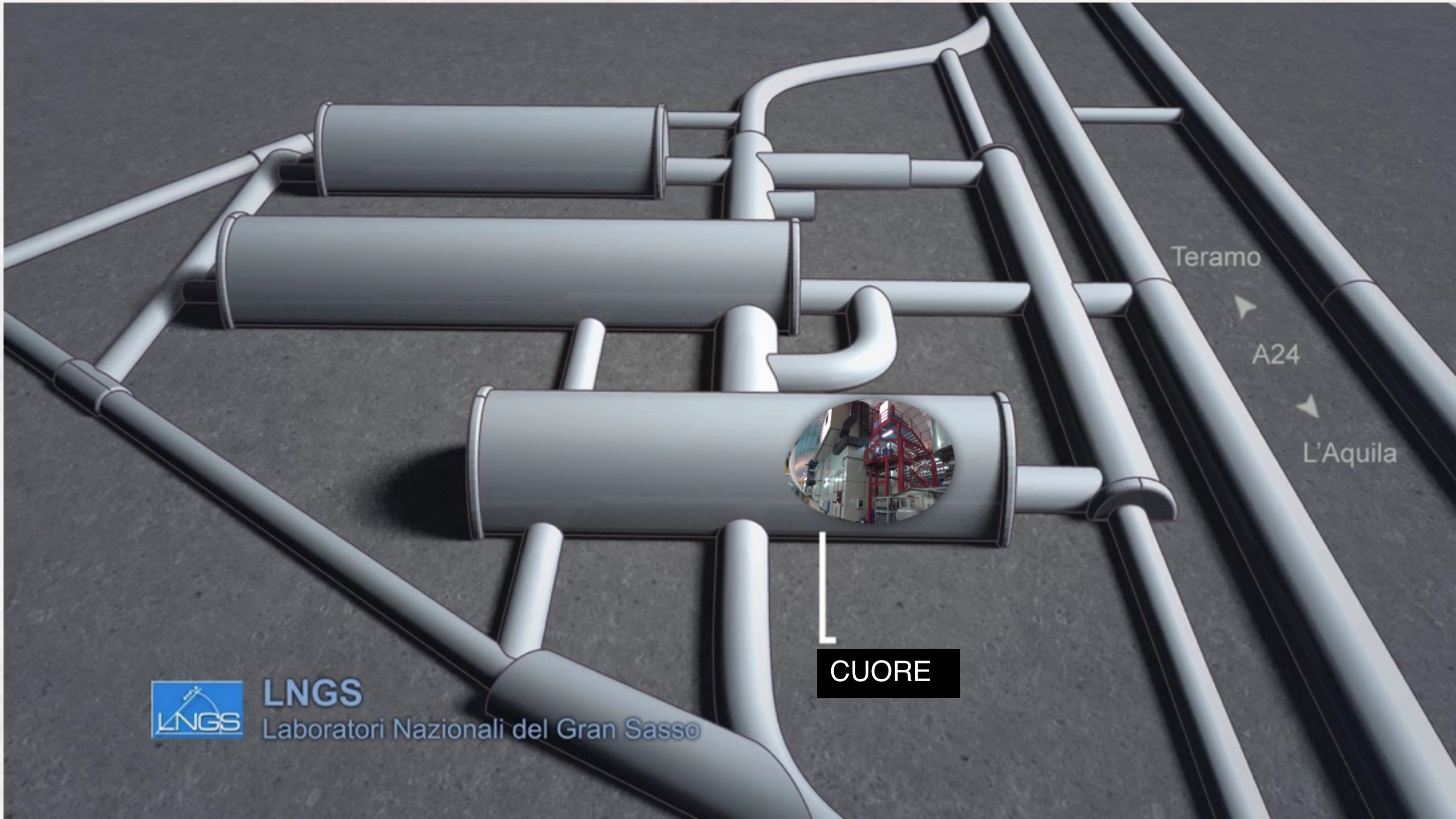




1400 m of rock ( $\sim 3600$  m.w.e.) deep

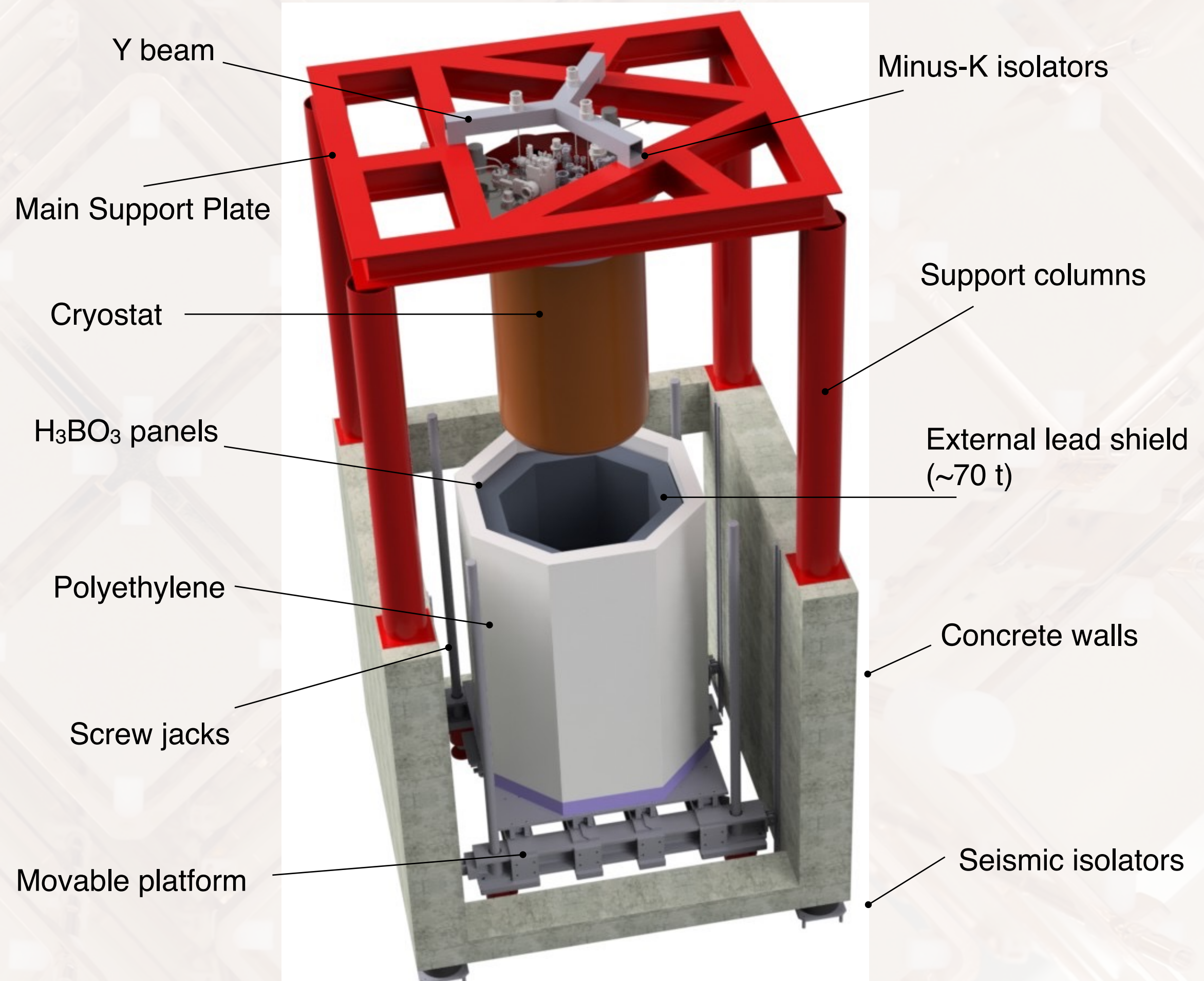
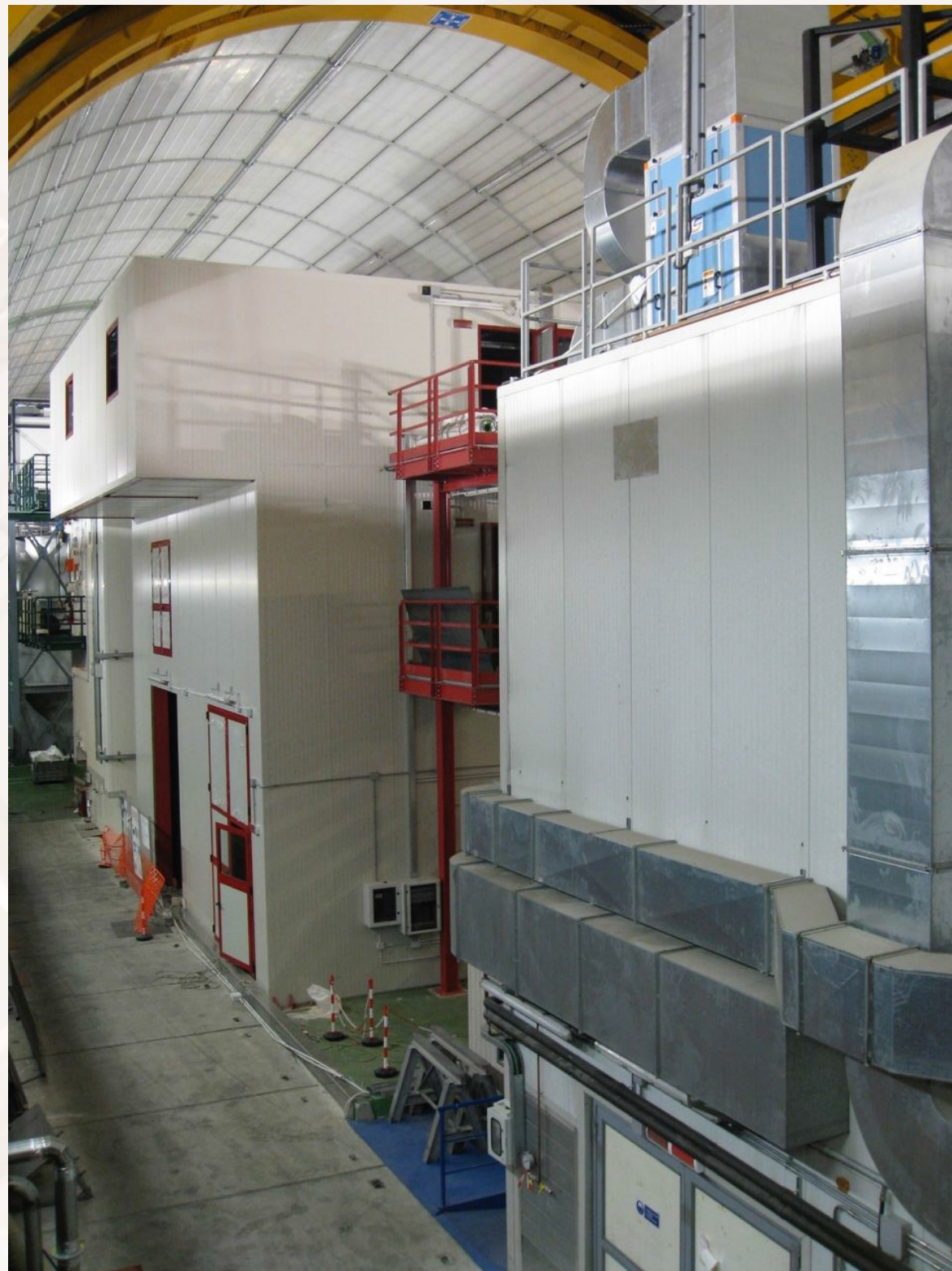
- $\mu$ 's:  $\sim 3 \times 10^{-8} / (\text{s} \cdot \text{cm}^2)$
- $\gamma$ 's:  $\sim 0.73 / (\text{s} \cdot \text{cm}^2)$
- neutrons:  $4 \times 10^{-6} \text{ n}/(\text{s} \cdot \text{cm}^2)$  below 10 MeV

# CUORE @ LNGS



# Underground Laboratory

- Three-story building
- Hosting the cryostat supporting structure

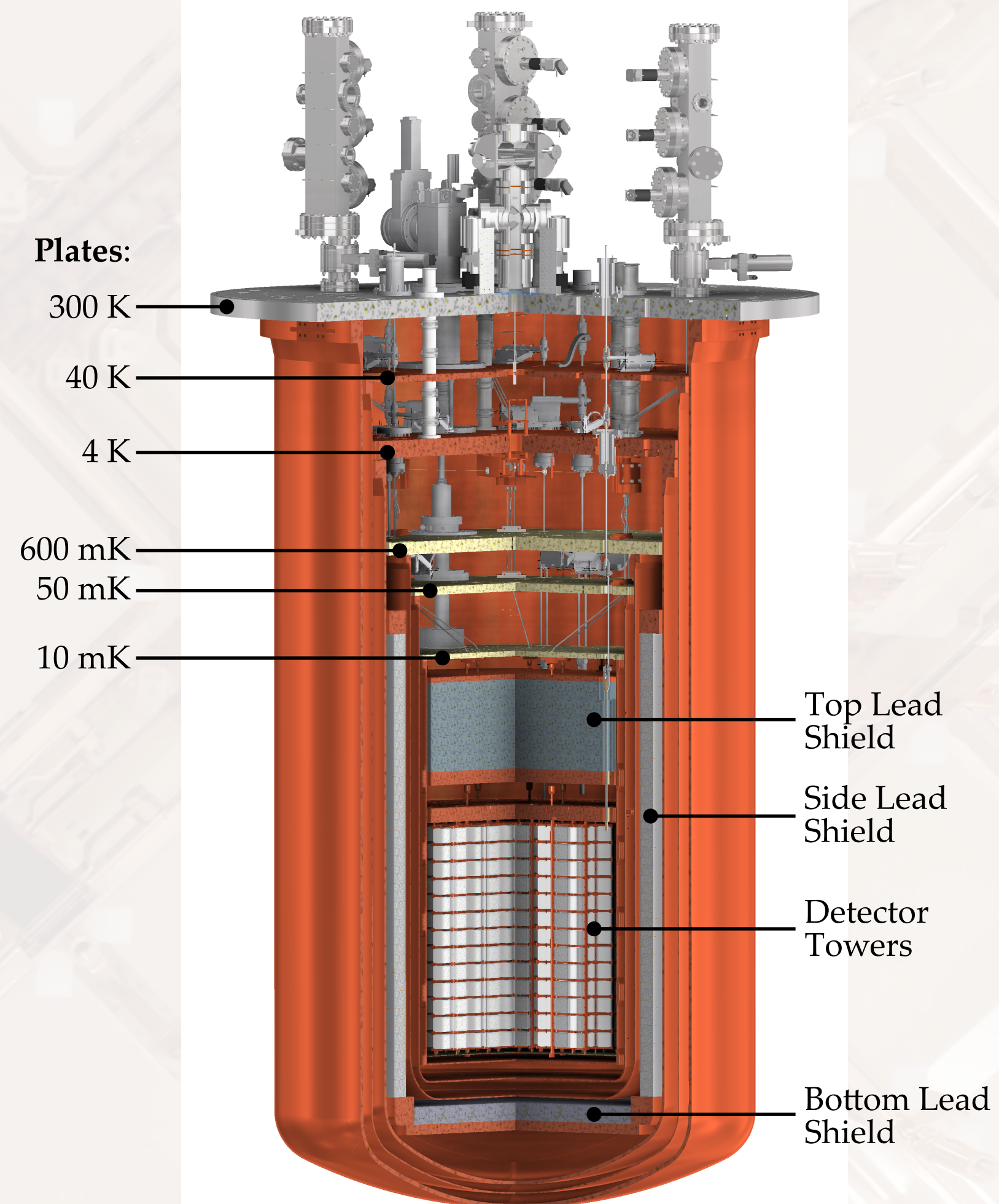


# The CUORE cryostat

## Challenges:

- Cool down ~1 ton detector to ~10 mK
- Mechanically decoupled for extremely low vibrations
- Low background environment

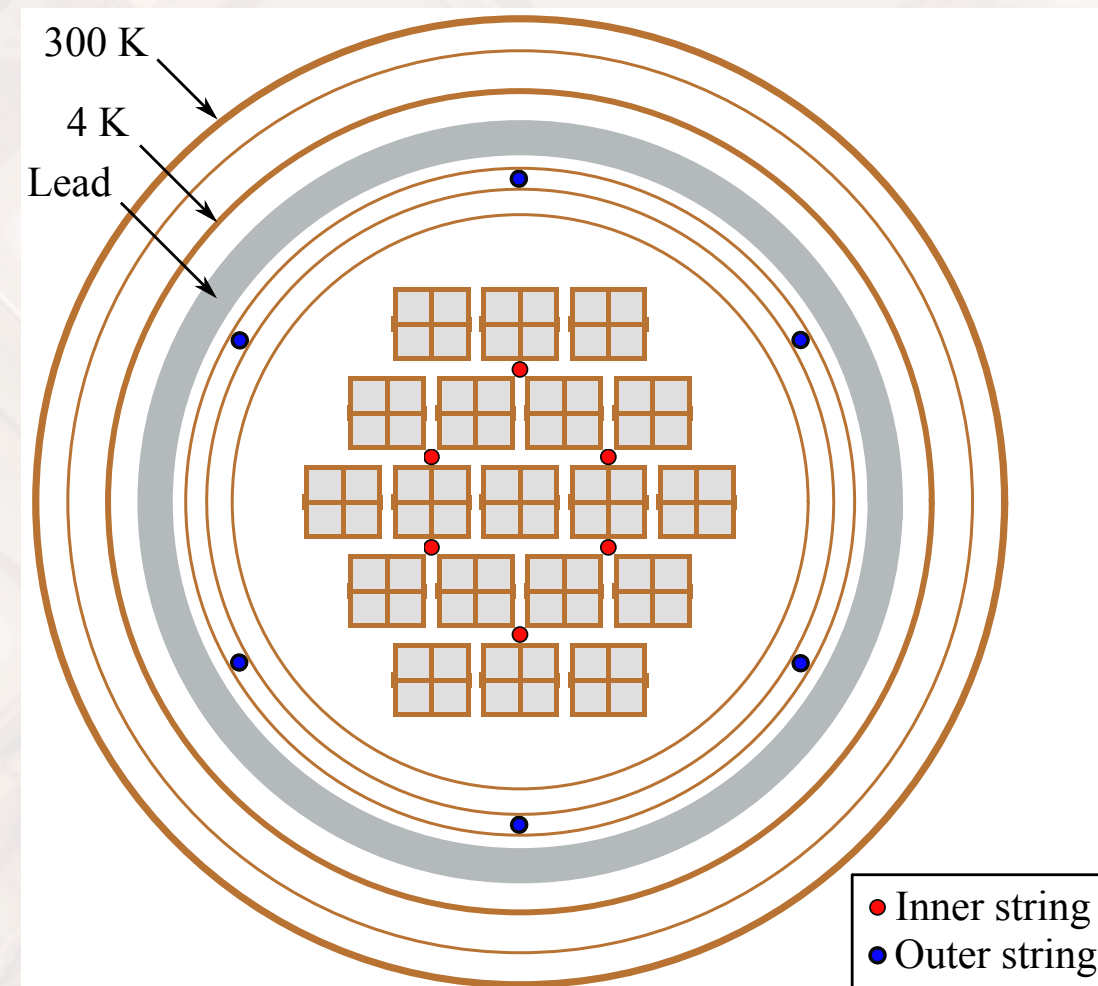
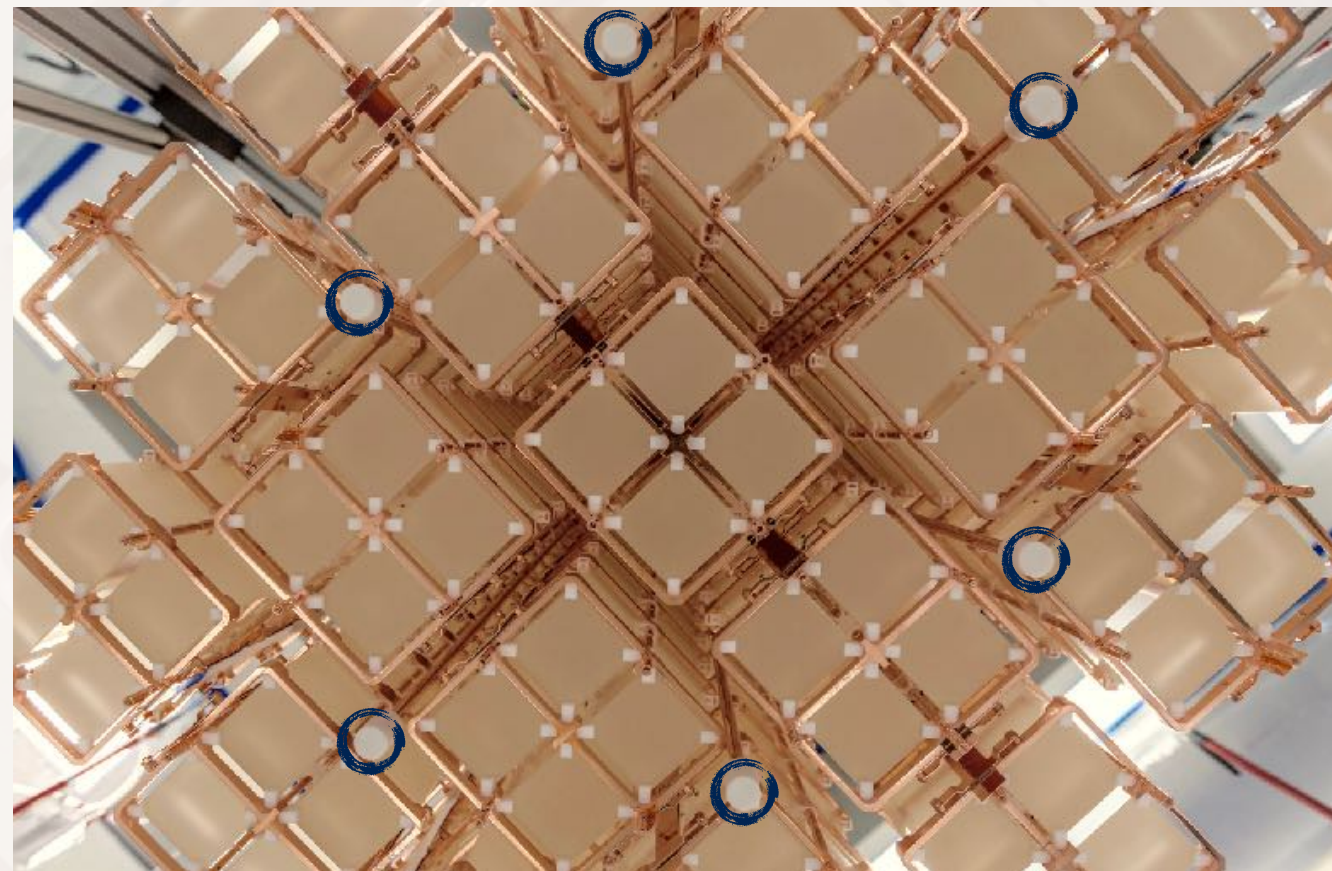
- Cryogen-free cryostat
- Fast Cooling System ( $^4\text{He}$  gas) down to ~50K
- 5 pulse tubes cryocooler down to ~4K
- Dilution refrigerator down to operating temperature ~10 mK
- Nominal cooling power:  $3 \mu\text{W}$  @ 10mK
- Cryostat total mass ~30 tons
- Mass to be cooled < 4K: ~15 tons
- Mass to be cooled < 50 mK: ~3 tons (Pb, Cu and  $\text{TeO}_2$ )



# Detector calibration system

## Challenges:

- Provide a uniform calibration of all the CUORE detectors
- Deployment of thoriated strings through the cryostat, from room temperature into the detector core



## Plates:

300 K

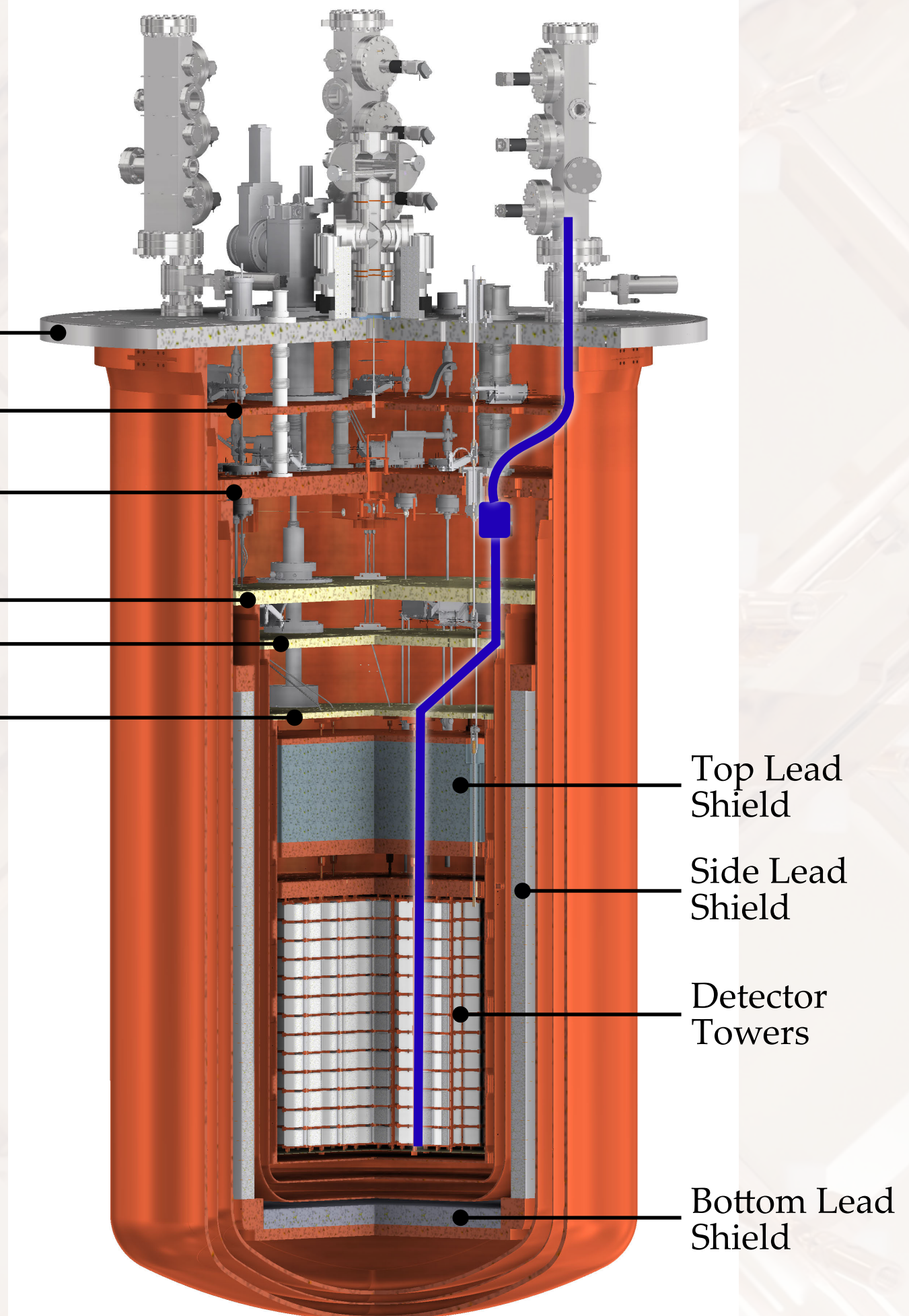
40 K

4 K

600 mK

50 mK

10 mK



J. S. Cushman et al. The detector calibration system for the CUORE cryogenic bolometer array. Nuclear Instruments and Methods A 844, 32-44 (2017). arxiv:1608.01607

# Passive shielding

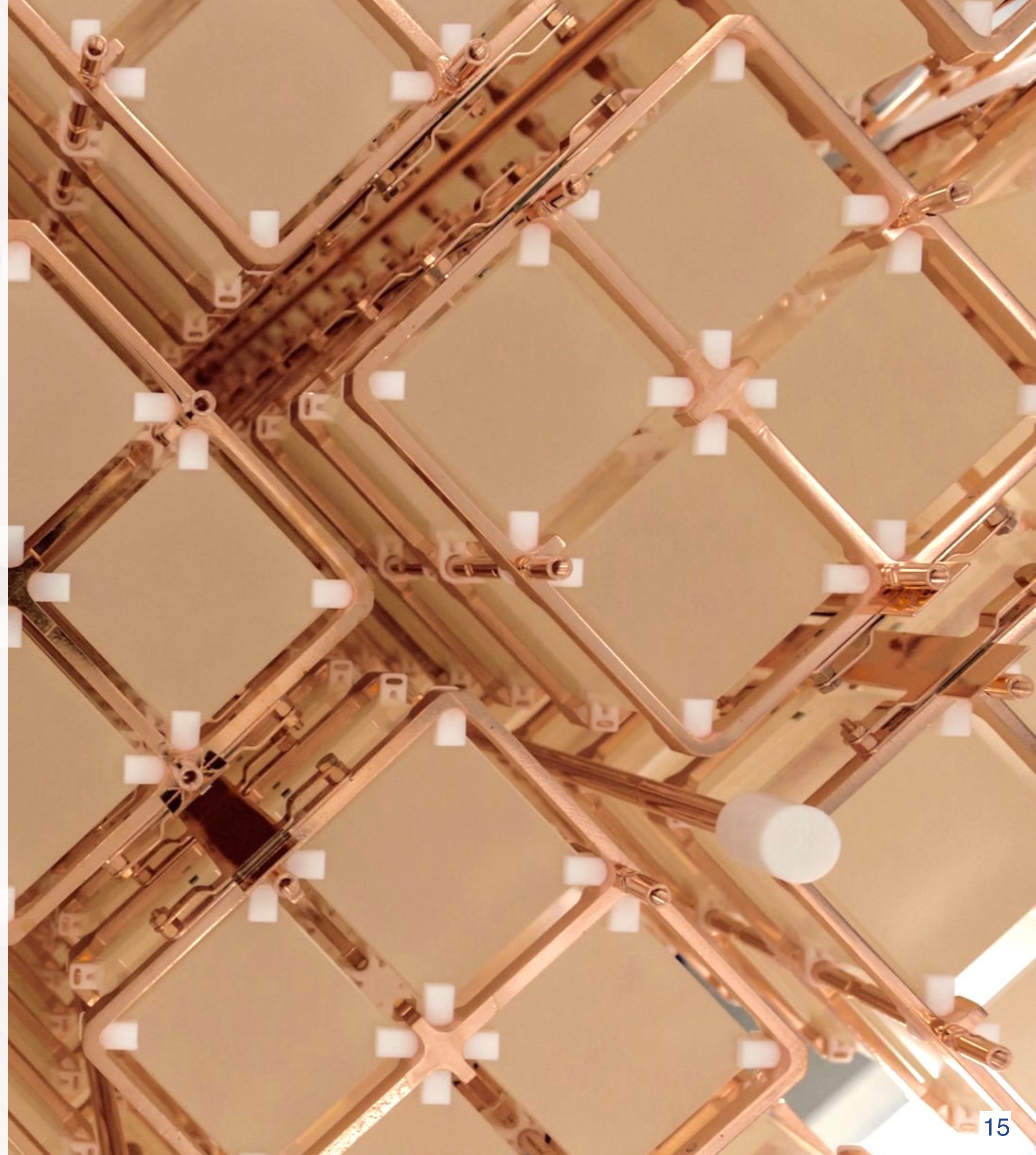
## Challenges:

- Protect the detectors with a heavy shield against gamma and neutron activity from external sources ( $\sim 70$  tonnes lead +  $\text{H}_3\text{BO}_3$ )
- Select materials that don't contribute themselves to the background level (ancient roman lead and selected NOSV copper)
- Cool down inner layers of the shielding to the correct temperature (2.5 tonnes @ 50mK + 5.5 tonnes @ 4K)



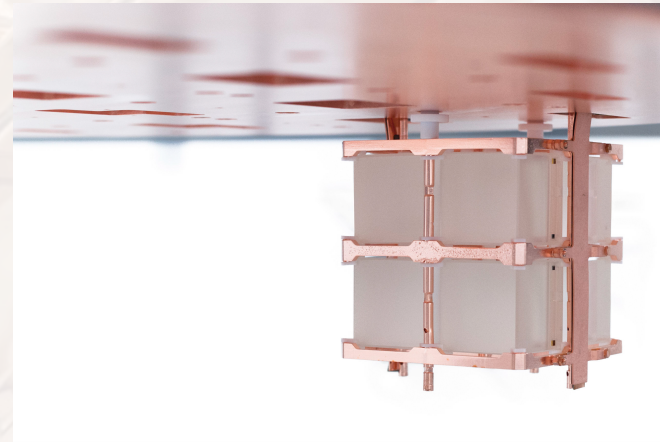
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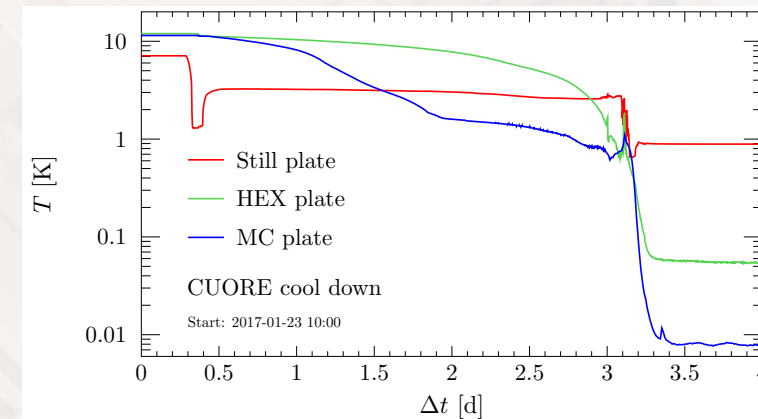
## Cryostat Commissioning

- 8-channels functioning mock-up detector
- Noise study and mitigation
- Stable base temperature < 7 mK
- Calibration sources deployment



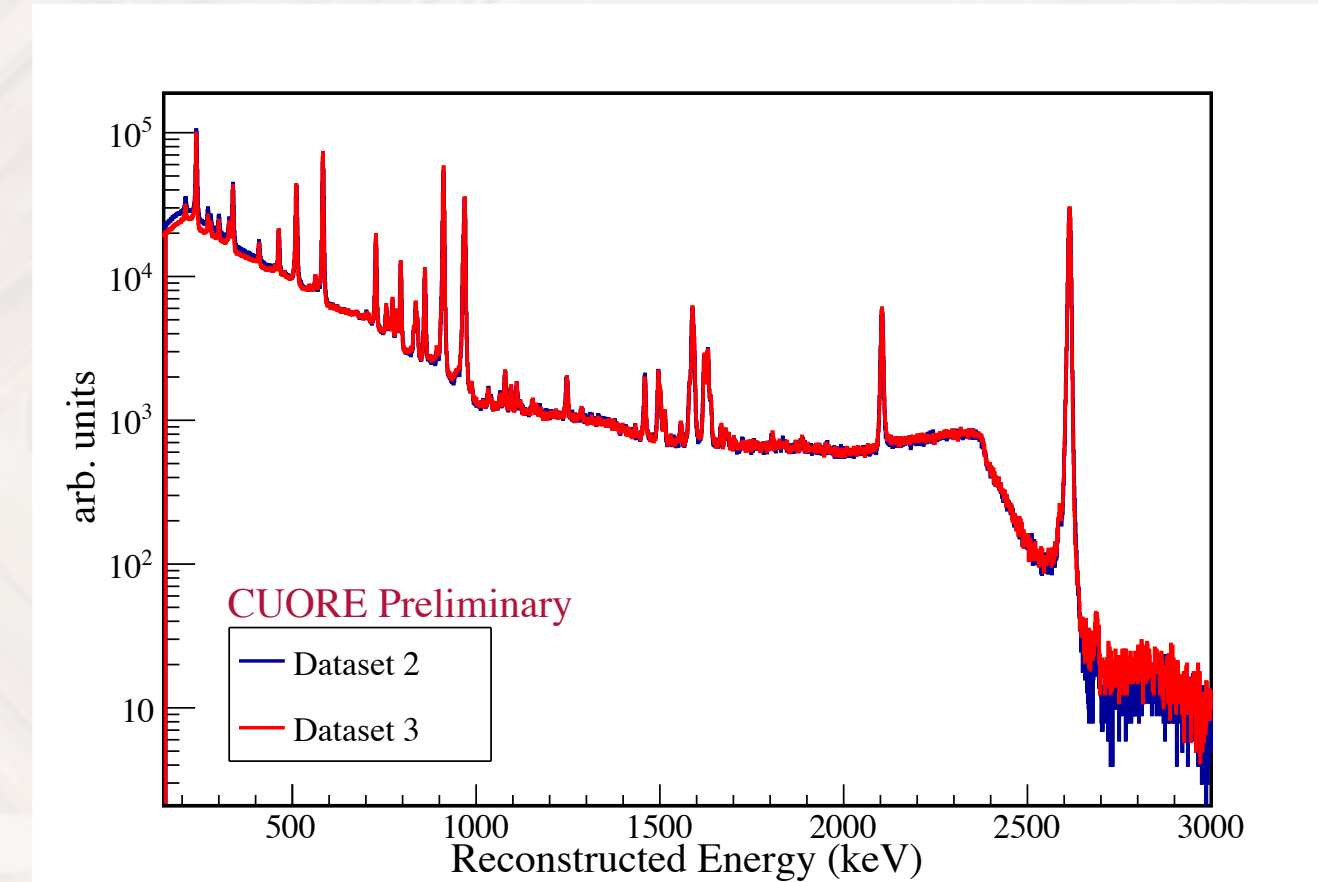
## Detector Cool down

- 300 K  $\rightarrow$  4 K in 22 days
- 4 K  $\rightarrow$  7 mK in 3.5 days
- First pulses seen just after cool-down



## Second period of Science Data taking

- New optimised working temperature
- New working points
- PT noise reduction
- Back to stable data taking:
  - calibration runs completed
  - physics runs ongoing
- Resolution in new data unchanged after maintenance



Feb 2016

Aug - Oct 2016

Jan 2017

Mar - Apr 2017

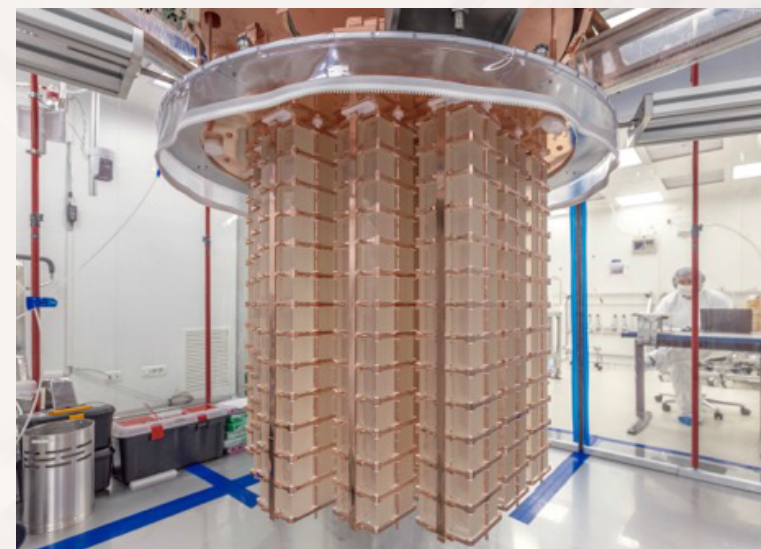
May - Sep 2017

Nov - Mar 2017

Mar 2018

## Detector Installation

- Radon-free environment
- 1 tower/day, 3 operators
- Read-out testing
- Cryostat interfaces
- Inner radiation shields



## Detector pre-operation

- Optimisation of all sub-systems
- Working temperature and working point selection
- Noise reduction

## First period of Science Data taking

- **Dataset 1:** physics data bracketed by calibrations (May - Jun 2017)
- Short optimisation campaign
- **Dataset 2:** Aug - Sep 2017
- 984/988 operational channels
- 86.3 kg\*y exposure

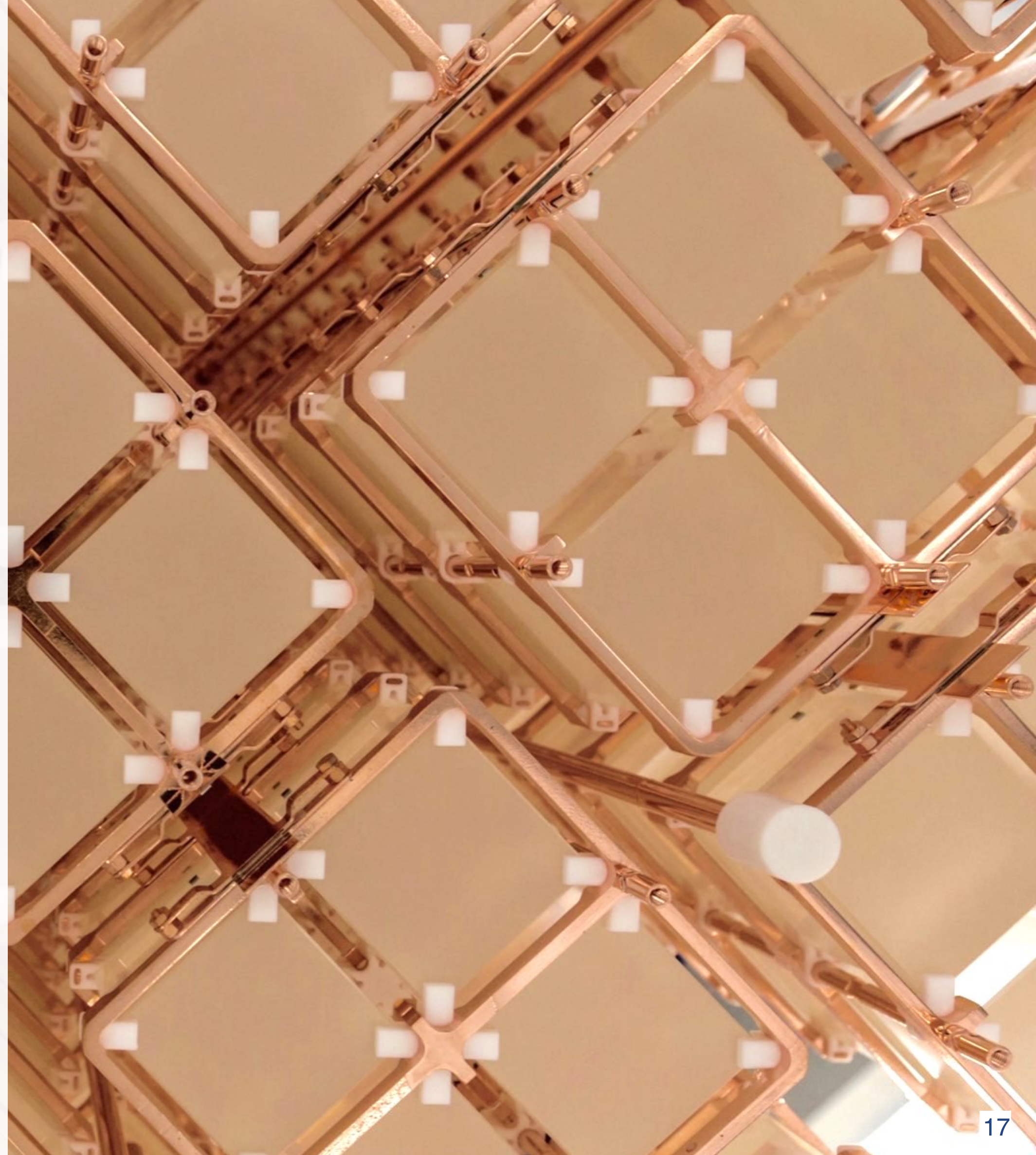
## Partial warm up

- Fix small leak in the cryostat region
- Warmed up to 100 K



# Outline

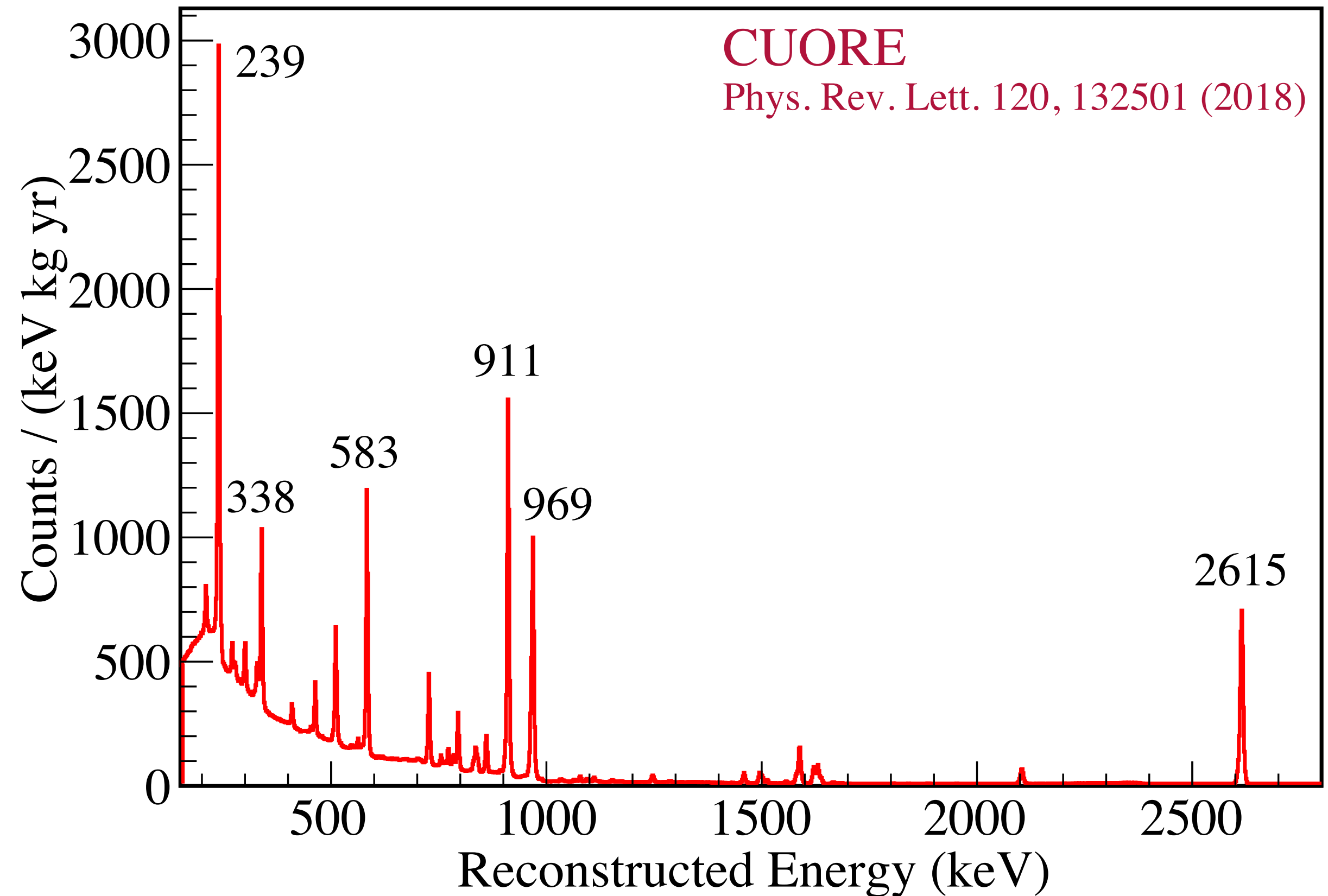
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# Calibration spectrum

- Calibration strings deployed inside the CUORE detector
- Summed energy spectrum of all the CUORE detectors-datasets
- Calibration data used for:
  - energy scale calibration
  - thermal gain stabilisation
  - detector response (line shape) study

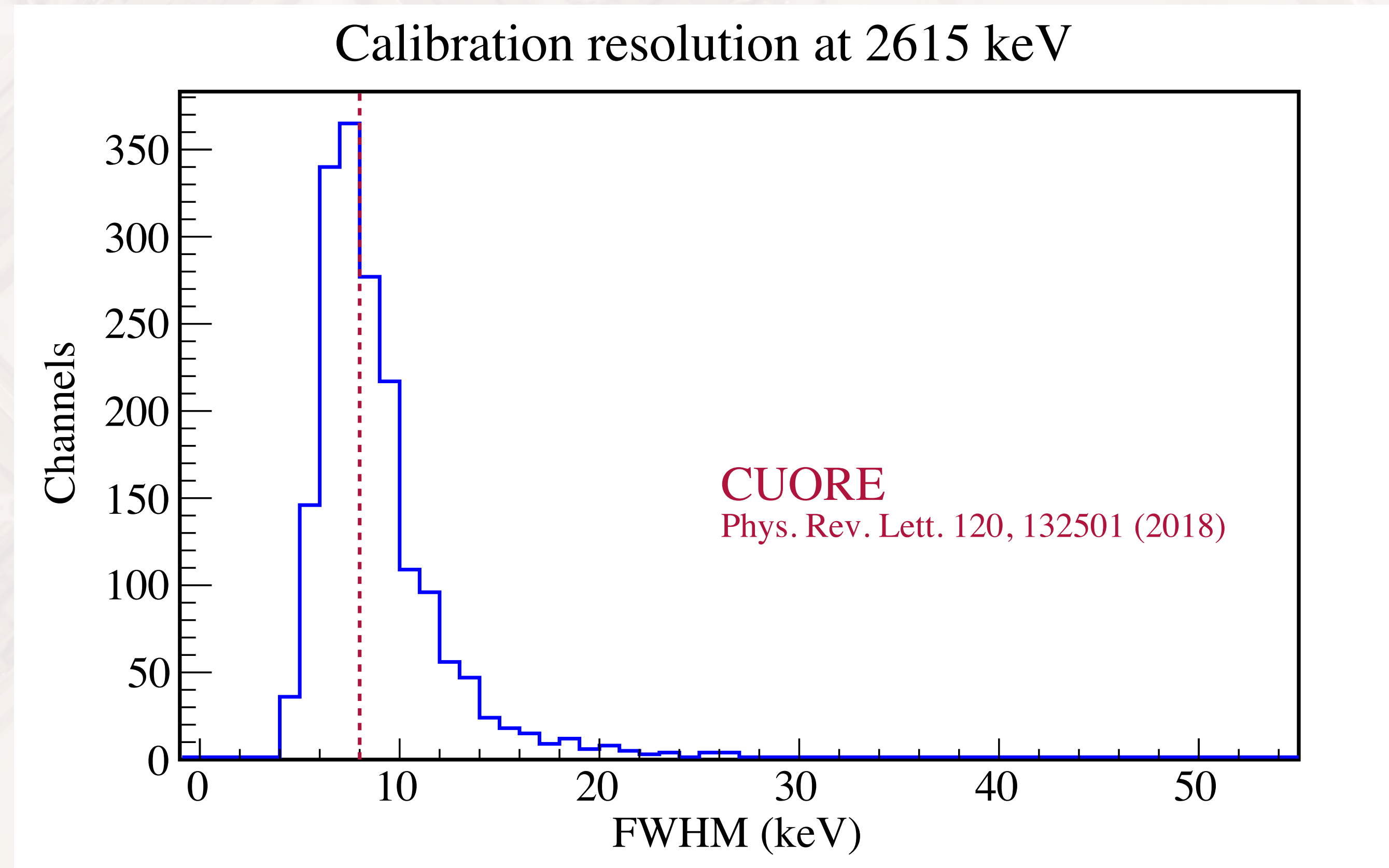
239 keV -  $^{212}\text{Pb}$   
338, 911, 969 keV -  $^{228}\text{Ac}$   
583, 2615 keV -  $^{208}\text{Tl}$



# Energy resolution - Calibration runs

A total of **1811** (92% of live channels) **channels-dataset couples** were used in this analysis; **discarded channels had poor line or pulse shapes, or the energy couldn't be reconstructed accurately**

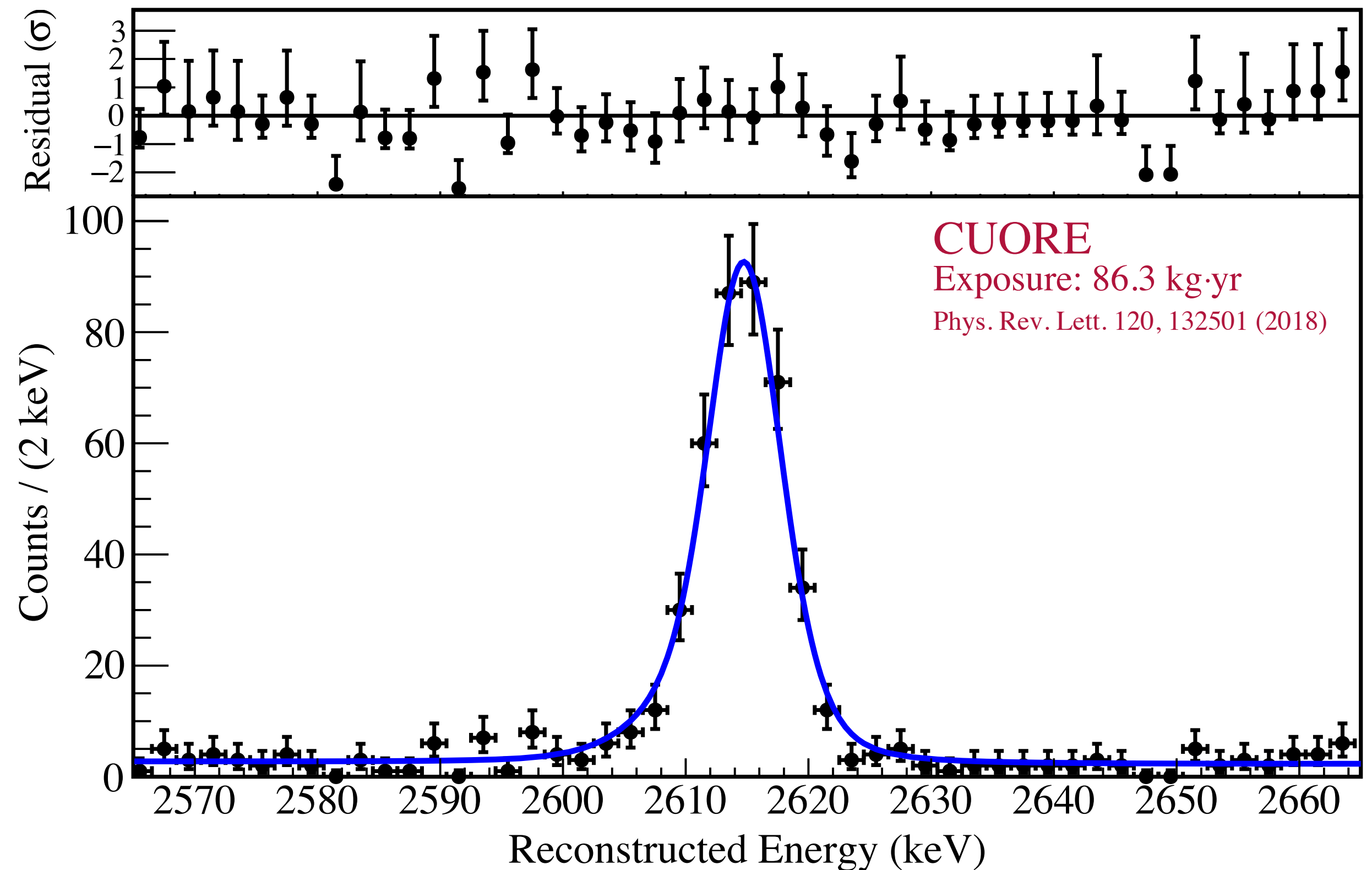
@ 2615 keV  
 ds3018: 9.0 keV FWHM  
 ds3021: 7.4 keV FWHM  
**effective (exposure-weighted):  
 8.0 keV FWHM**



# Energy resolution - Physics runs

A total of **1811** (92% of live channels) **channels-dataset couples** were used in this analysis; **discarded channels had poor line or pulse shapes, or the energy couldn't be reconstructed accurately**

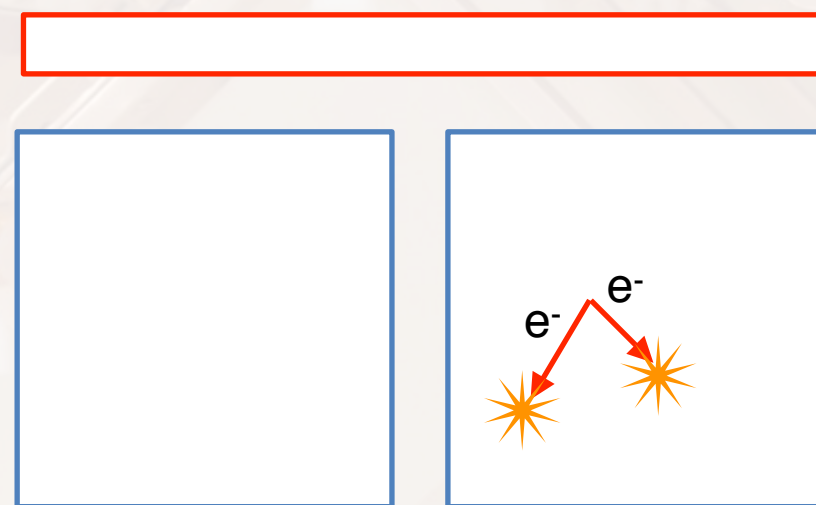
@ Q-value  
 ds3018:  $(8.3 \pm 0.4)$  keV FWHM  
 ds3021:  $(7.4 \pm 0.7)$  keV FWHM  
**effective (exposure-weighted):**  
 **$(7.7 \pm 0.5)$  keV FWHM**



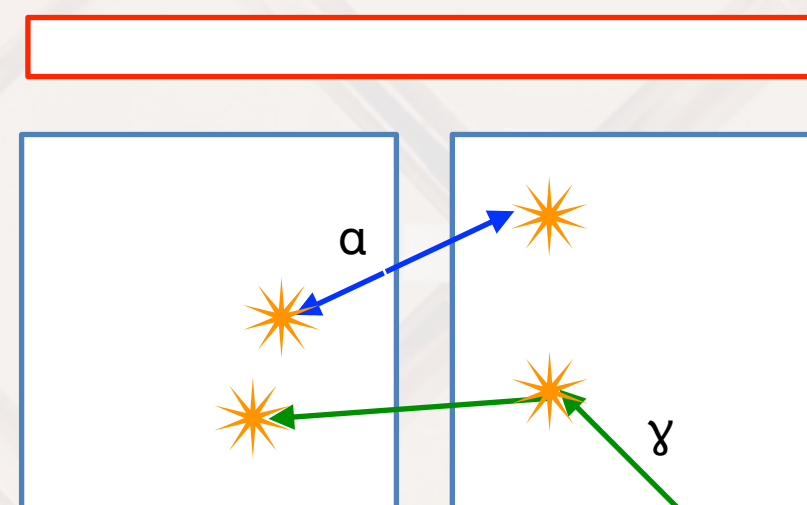
# Event selection

Event selection occurs after periods of low-quality data ( $\sim 1\%$  of the total live time) are removed.

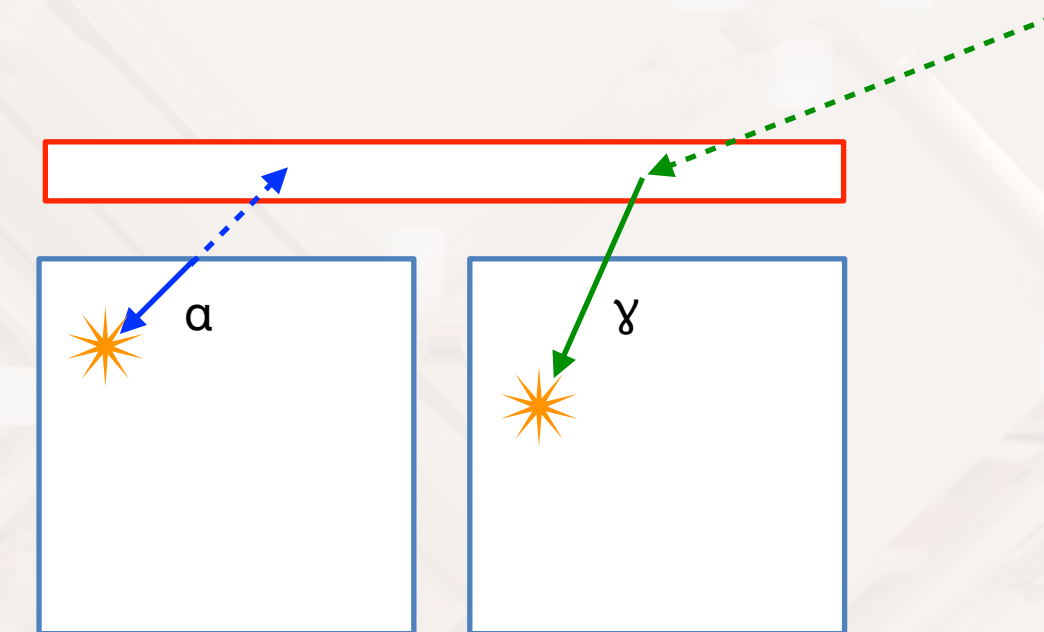
- Base quality cuts (number of pulses in the window, baseline stability, etc...)
- Anti-coincidence



M1 - signal like



M2 - rejected background  
and important information  
for bkg studies

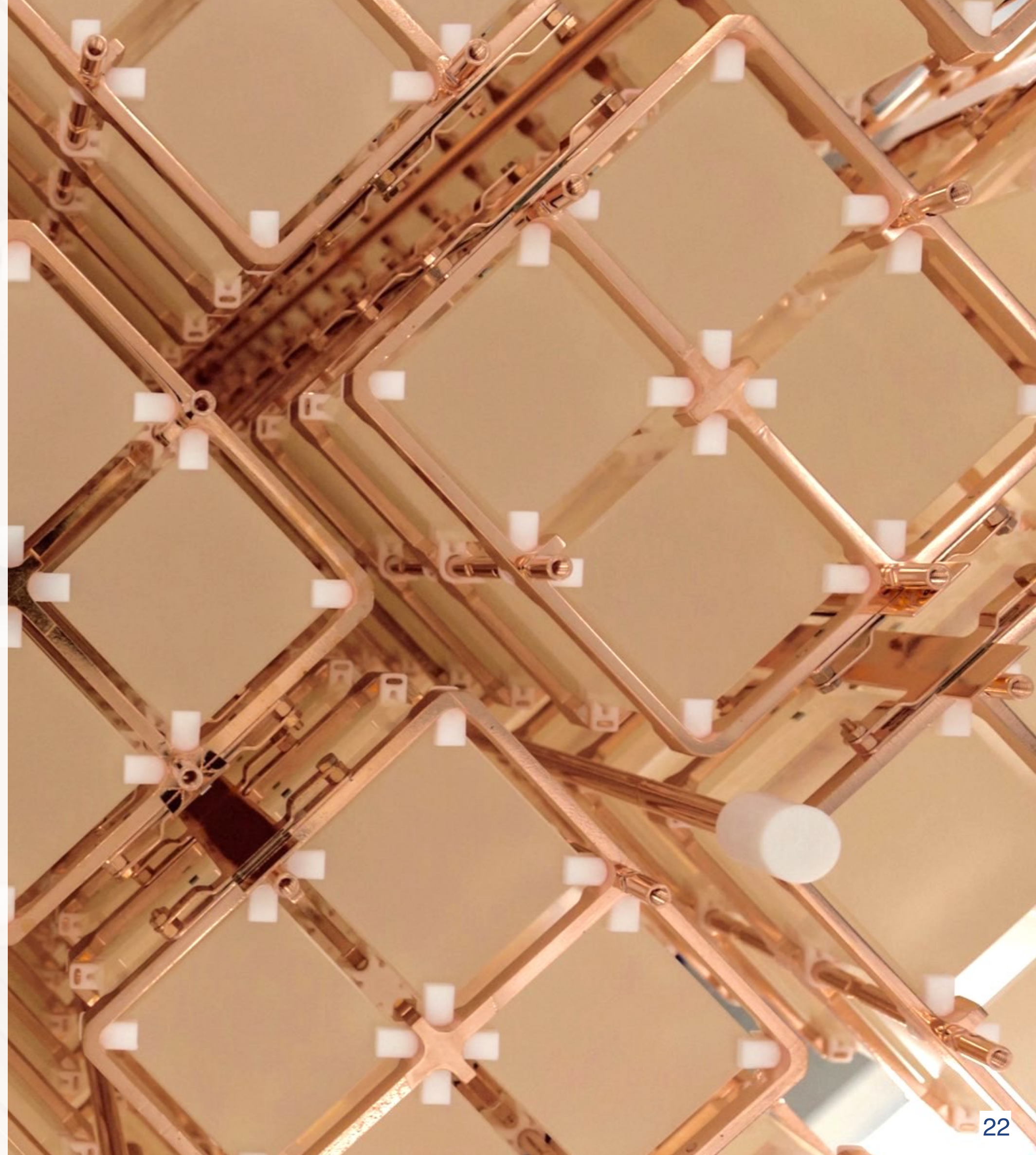


M1 - dangerous background

- Pulse shape analysis: deformed events are not used to build the final spectra to avoid spectral shape distortions (but are still used for anti-coincidence)

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# Fit in the 0νDBD ROI

Region of interest: **2465 to 2575 keV**

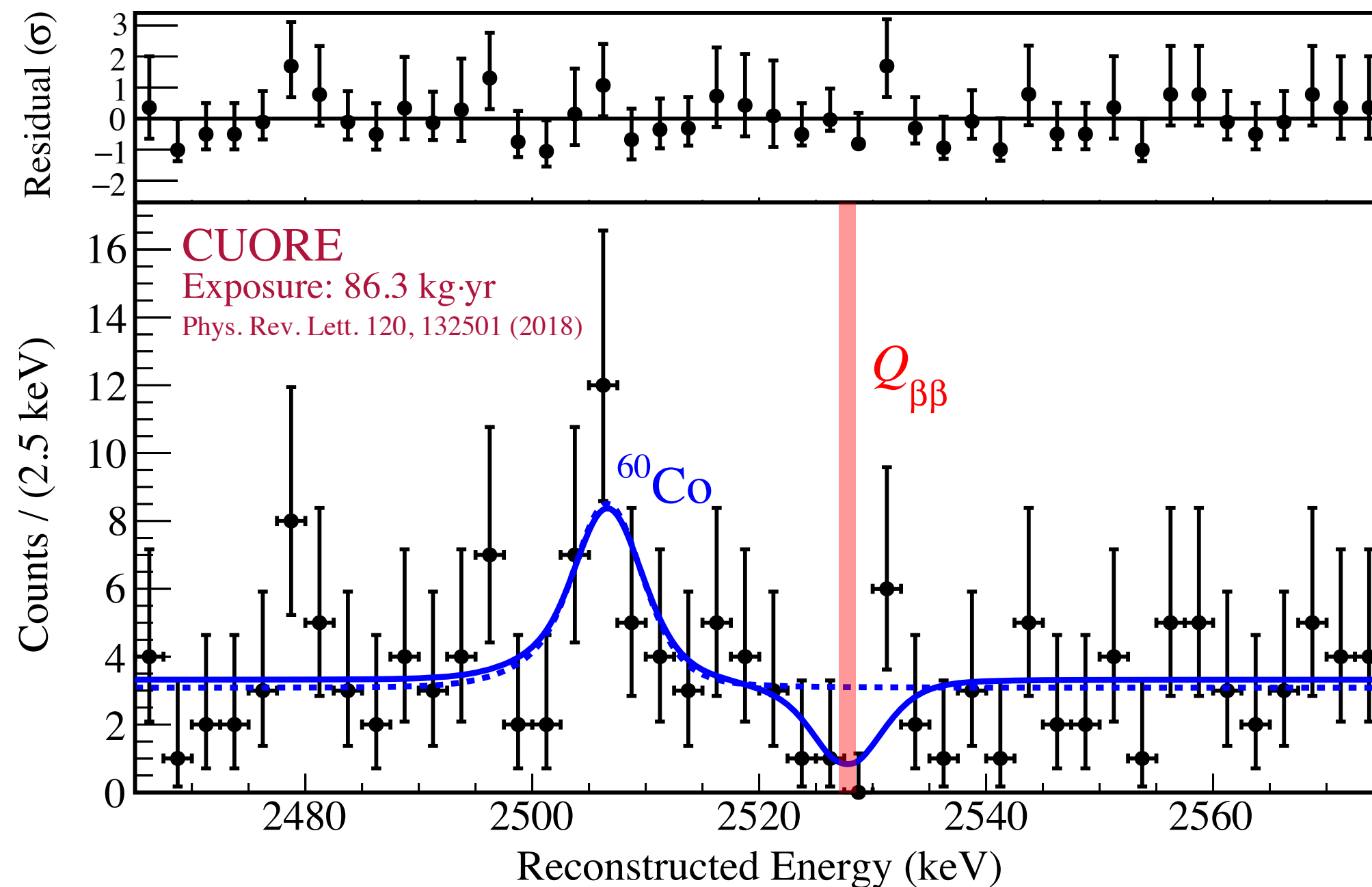
Overall signal efficiency: **(75.7 ± 3.0)% - ds3018**  
**(83.0 ± 2.6)% - ds3021**

Events in the region of interest: **155**

ROI background index: **(1.49<sub>-0.17</sub><sup>+0.18</sup>) × 10<sup>-2</sup> c/(keV · kg · yr)**  
**(1.35<sub>-0.18</sub><sup>+0.20</sup>) × 10<sup>-2</sup> c/(keV · kg · yr)**

Best fit for <sup>60</sup>Co mean: **(2506.4 ± 1.2) keV**

Best fit decay rate: **(-1.0<sub>-0.3</sub><sup>+0.4</sup> (stat.) ± 0.1 (syst.)) × 10<sup>-25</sup> / yr**



**No evidence of signal**

Decay rate limit (90% CL, including systematics):  
**0.51 × 10<sup>-25</sup> / yr**

Half-life limit (90% CL, including systematics):  
**1.3 × 10<sup>25</sup> yr**

Median expected sensitivity:  
**7.0 × 10<sup>24</sup> yr**

CUORE, CUORE-0 and CUORICINO  
 combined 90% C.L. limit is  
**T<sub>0ν</sub> > 1.5 × 10<sup>25</sup> yr**

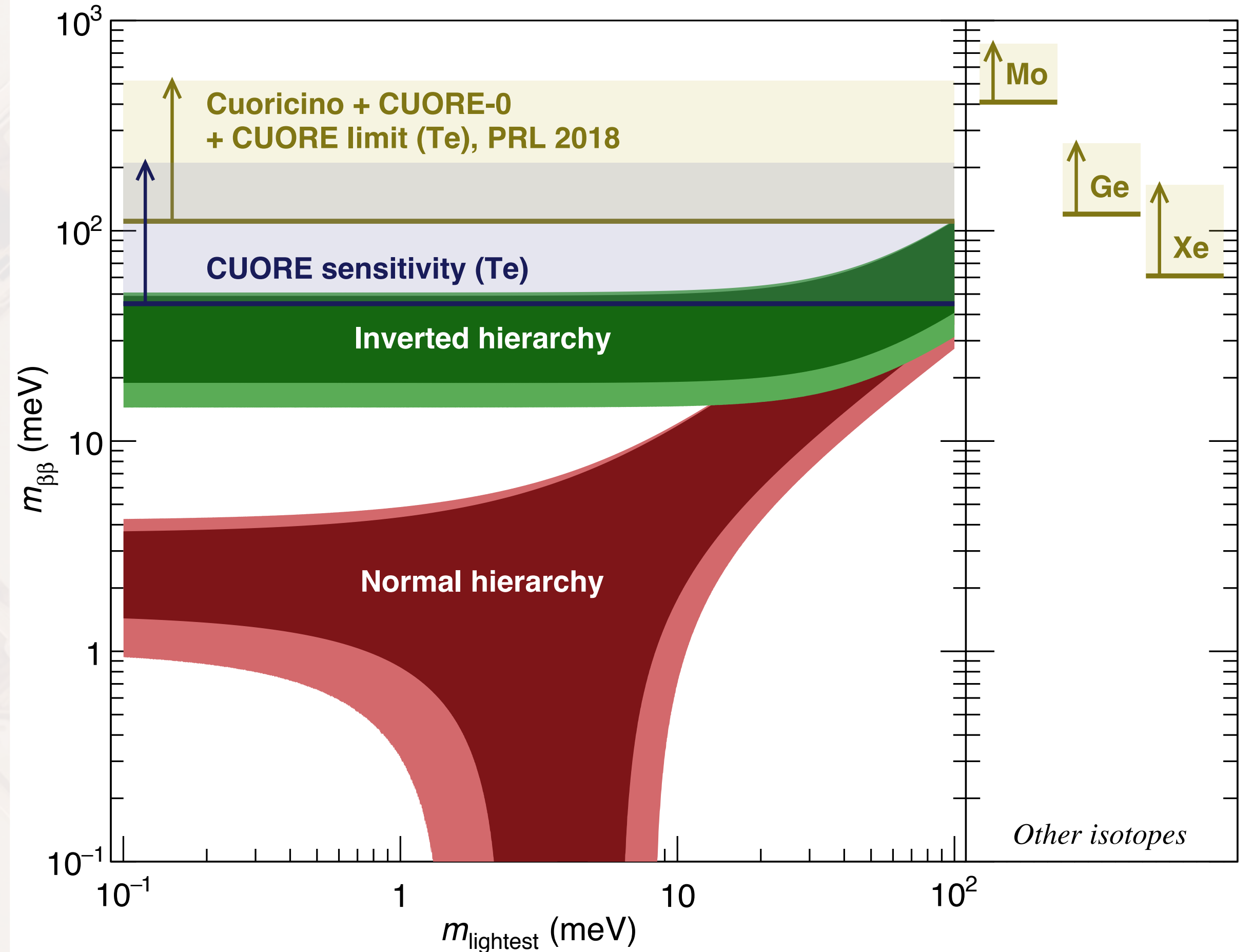
Phys. Rev. Lett. 120, 132501 (2018)

# Combination with previous 0νDBD results

- Total  $^{130}\text{Te}$  exposure
  - 86.3 kg·yr of CUORE
  - 19.75 kg·yr of Cuoricino
  - 9.8 kg·yr of CUORE-0
- The combined 90% C.L. limit is
  - $T_{0\nu} > 1.5 \times 10^{25}$  yr
  - $m_{\beta\beta} < 110\text{-}520$  meV

NME:  
 JHEP02 (2013) 025  
 Nucl. Phys. A 818, 139 (2009)  
 Phys. Rev. C 87, 045501 (2013)  
 Phys. Rev. C 87, 064302 (2014)  
 Phys. Rev. C 91, 034304 (2015)  
 Phys. Rev. C 91, 024613 (2015)  
 Phys. Rev. C 91, 024309 (2015)  
 Phys. Rev. C 91, 024316 (2015)  
 Phys. Rev. Lett. 105, 252503 (2010)  
 Phys. Rev. Lett. 111, 142501 (2013)

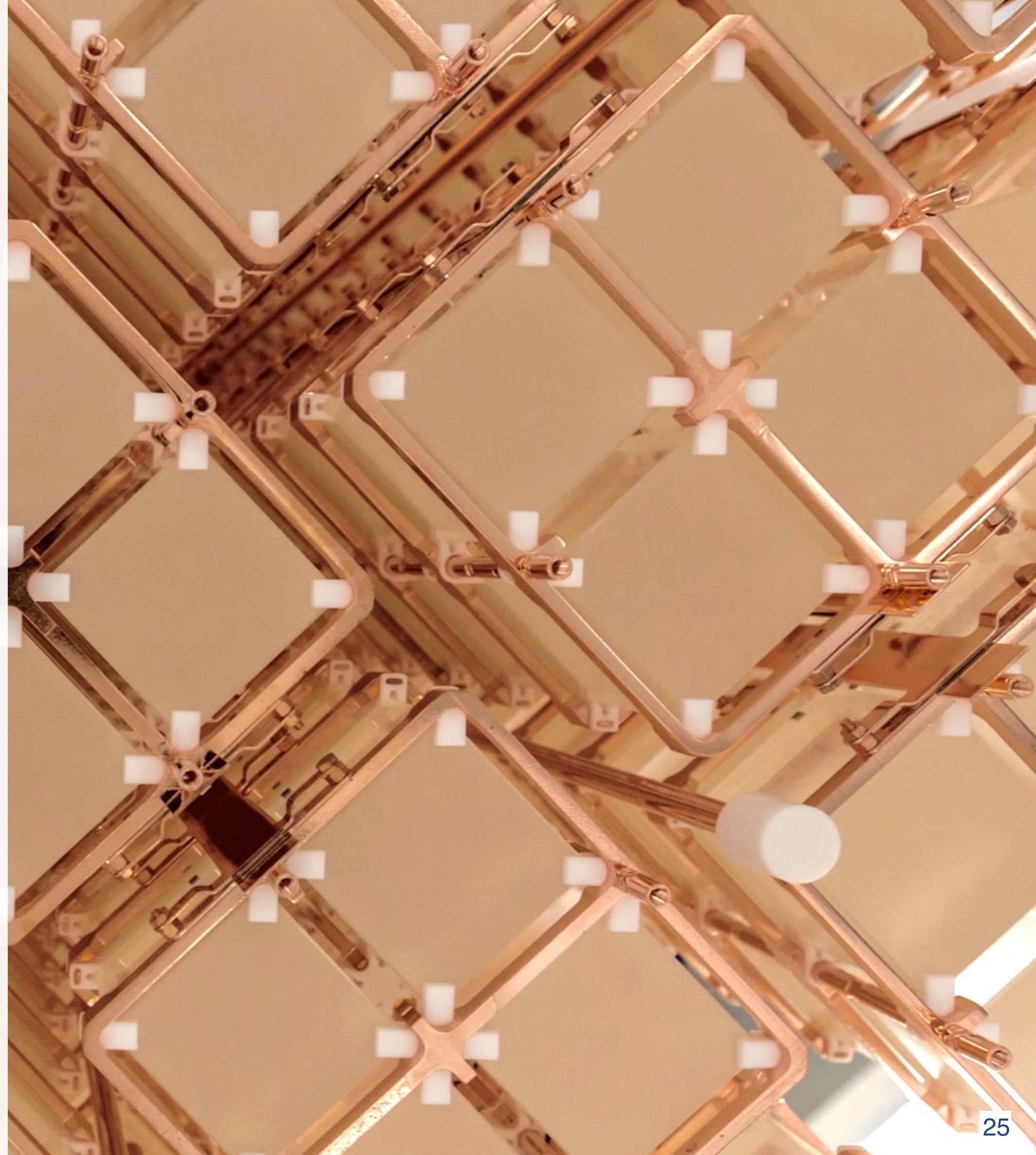
Experiments:  
 $^{130}\text{Te}$ :  $1.5 \times 10^{25}$  yr from PRL 120, 132501 (2018)  
 $^{76}\text{Ge}$ :  $8.0 \times 10^{25}$  yr from PRL 120, 132503 (2018)  
 $^{136}\text{Xe}$ :  $1.1 \times 10^{26}$  yr from Phys. Rev. Lett. 117, 082503 (2016)  
 $^{100}\text{Mo}$ :  $1.1 \times 10^{24}$  yr from Phys. Rev. D 89, 111101 (2014)  
 CUORE sensitivity:  $9.0 \times 10^{25}$  yr





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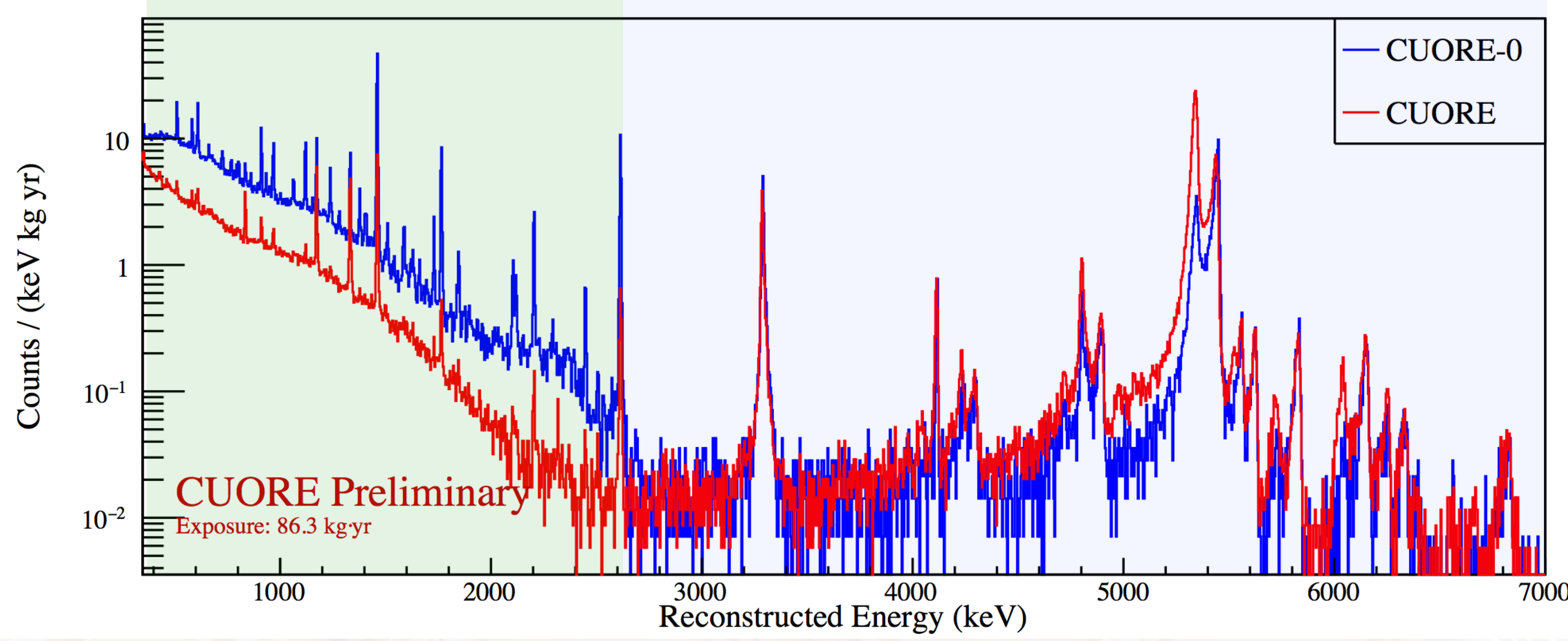


# CUORE Background Model



$\gamma$  background significantly reduced

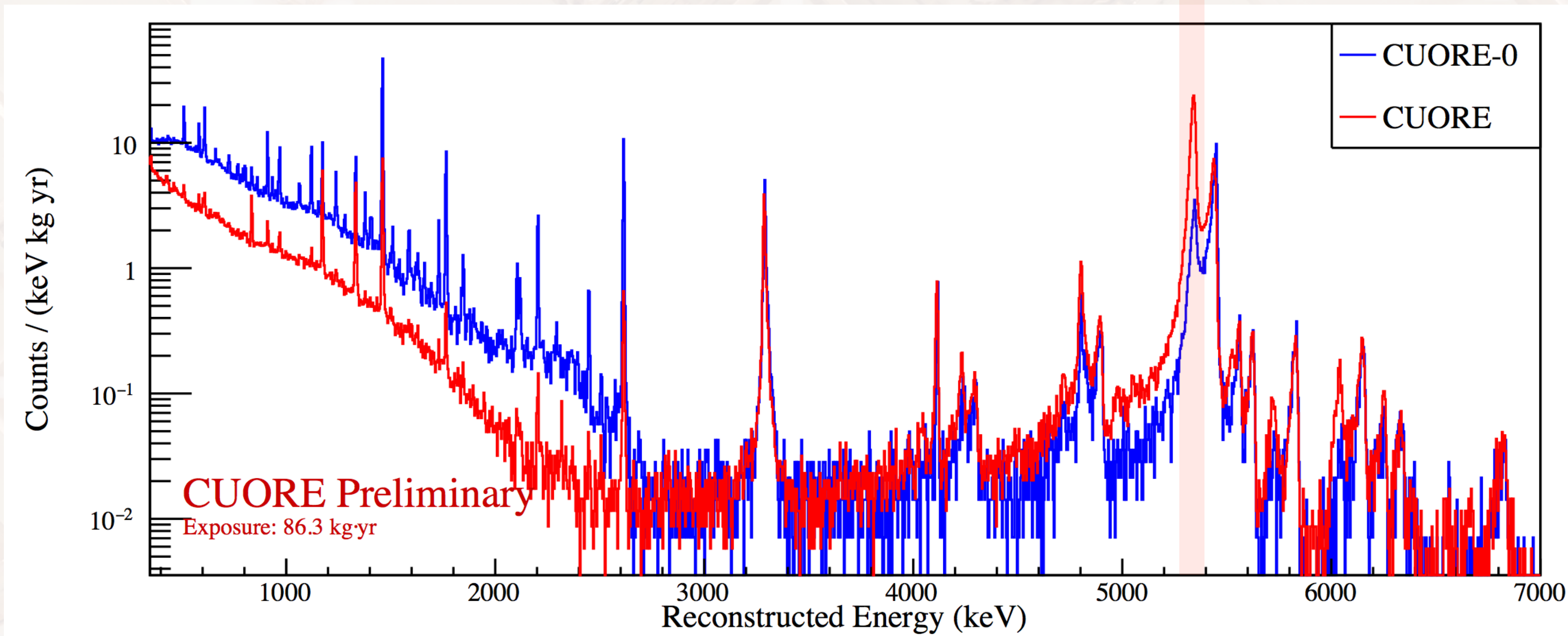
$\alpha$  background mostly consistent with CUORE-0 (as expected)



# CUORE Background Model

$^{210}\text{Po}$  excess still under investigation:

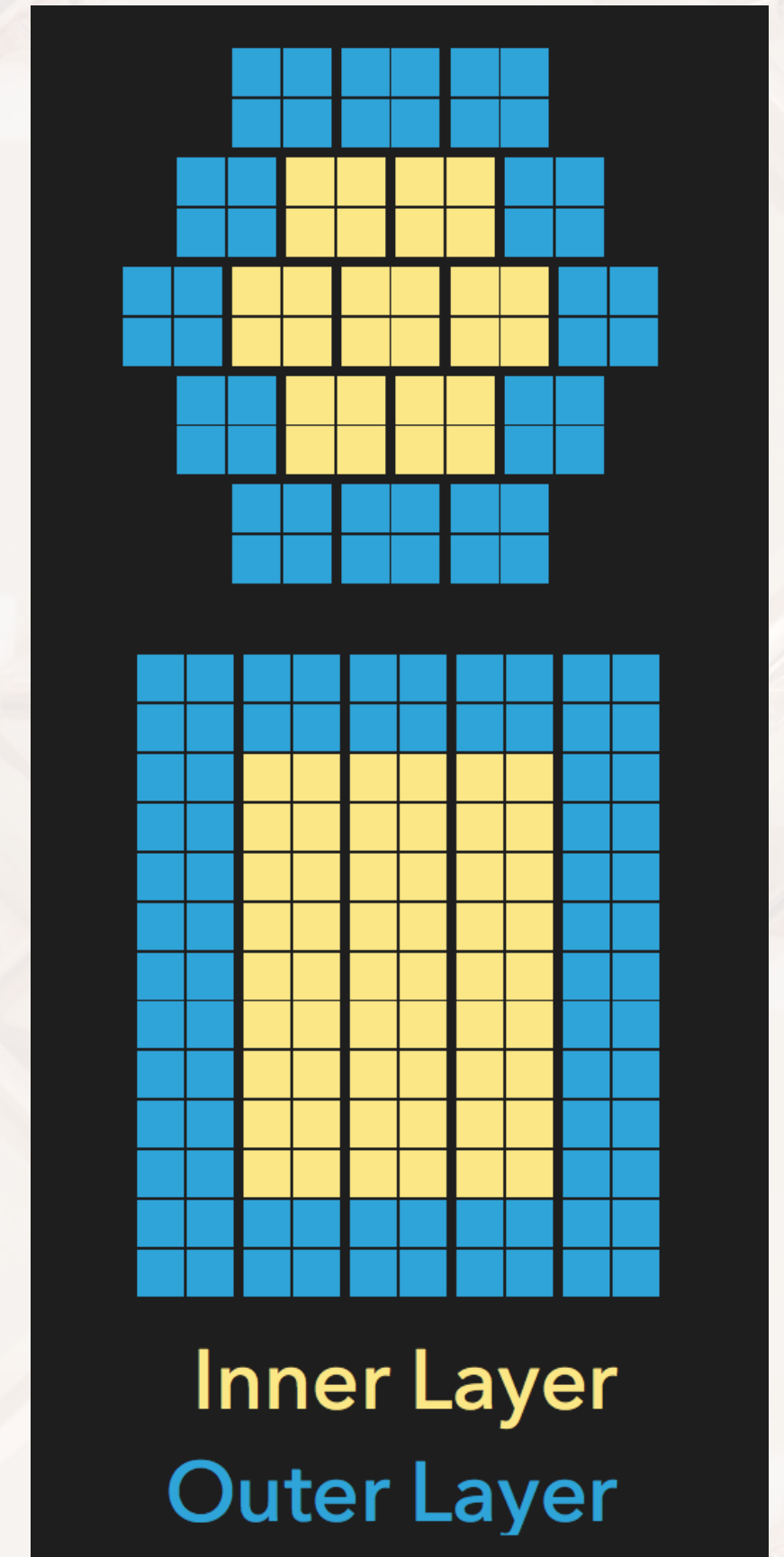
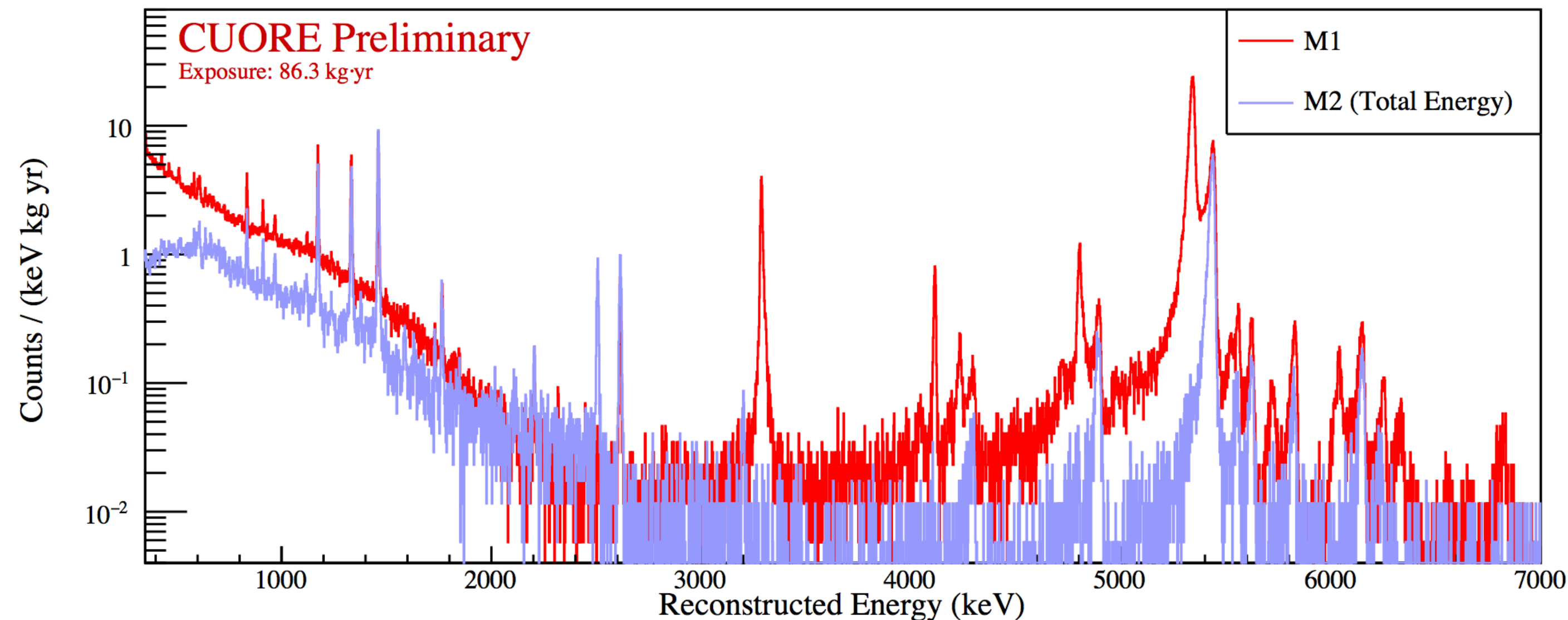
- appears to be from surface contamination in copper around the detector
- estimated contribution to ROI  $\sim 10^{-4}$  ckky ( $\sim 1\%$ )



# CUORE Background Model

## DATA

- 86.3 kg yr of  $\text{TeO}_2$ , same data used for  $0\nu\text{DBD}$  analysis
- split data into inner and outer (2 crystals thick) layers
  - outer layers more sensitive to external backgrounds
- split data into Multiplicity 1, Multiplicity 2 and Multiplicity 2 Sum
  - different multiplicities are sensitive to different types of backgrounds

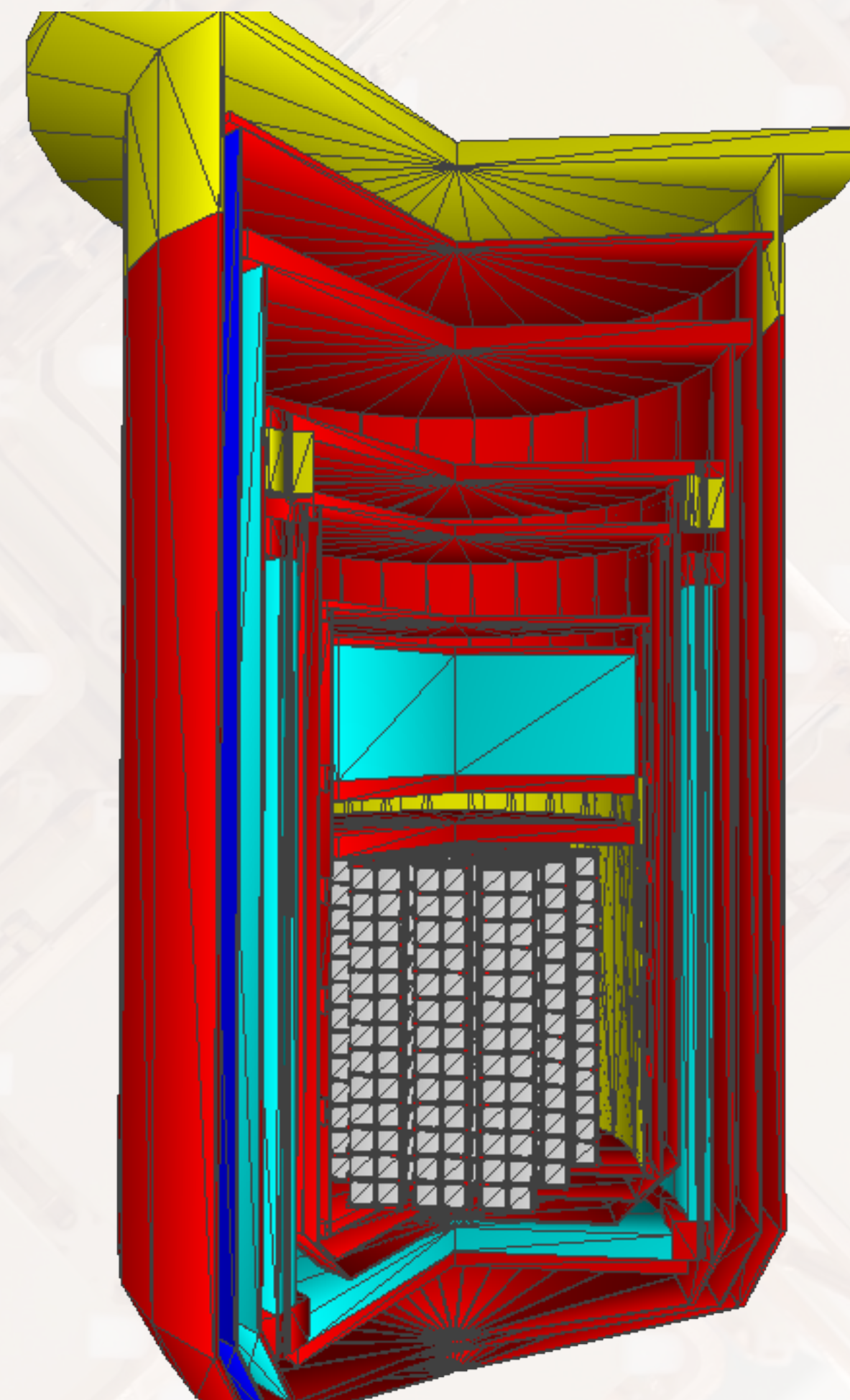


# CUORE Background Model

Volume	Type	Components
TeO <sub>2</sub>	Bulk	$2\nu\beta\beta$ , $^{210}\text{Pb}$ , $^{232}\text{Th}$ , $^{228}\text{Ra}$ - $^{208}\text{Pb}$ , $^{238}\text{U}$ - $^{230}\text{Th}$ , $^{230}\text{Th}$ , $^{226}\text{Ra}$ - $^{210}\text{Pb}$ , $^{40}\text{K}$ , $^{60}\text{Co}$ , $^{125}\text{Sb}$ , $^{190}\text{Pt}$
TeO <sub>2</sub>	Surface (0.01 $\mu\text{m}$ )	$^{232}\text{Th}$ , $^{228}\text{Ra}$ - $^{208}\text{Pb}$ , $^{238}\text{U}$ - $^{230}\text{Th}$ , $^{226}\text{Ra}$ - $^{210}\text{Pb}$ , $^{210}\text{Pb}$
TeO <sub>2</sub>	Surface (1 $\mu\text{m}$ )	$^{210}\text{Pb}$
TeO <sub>2</sub>	Surface (10 $\mu\text{m}$ )	$^{210}\text{Pb}$ , $^{232}\text{Th}$ , $^{238}\text{U}$
CuNOSV	Bulk	$^{232}\text{Th}$ , $^{238}\text{U}$ , $^{40}\text{K}$ , $^{60}\text{Co}$ , $^{54}\text{Mn}$
CuNOSV	Surface (0.01 $\mu\text{m}$ )	$^{210}\text{Pb}$ , $^{232}\text{Th}$ , $^{238}\text{U}$
CuNOSV	Surface (1 $\mu\text{m}$ )	$^{210}\text{Pb}$ , $^{232}\text{Th}$ , $^{238}\text{U}$
CuNOSV	Surface (10 $\mu\text{m}$ )	$^{210}\text{Pb}$ , $^{232}\text{Th}$ , $^{238}\text{U}$
Roman lead	Bulk	$^{232}\text{Th}$ , $^{238}\text{U}$ , $^{108m}\text{Ag}$
Top lead	Bulk	$^{232}\text{Th}$ , $^{238}\text{U}$ , $^{210}\text{Bi}$
Ext. lead	Bulk	$^{210}\text{Bi}$
CuOFE	Bulk	$^{232}\text{Th}$ , $^{238}\text{U}$ , $^{60}\text{Co}$
External	-	Cosmic muons

## MONTE CARLO

- ~60 independent simulations of sources/location in the setup
- full radioactive chains and single isotopes
- the different energy spectra (inner/outer, M1/M2, M2sum) of each source/location are generated with Geant4 based simulation implementing a detailed geometry of the setup (detector, cryostat, shields)



## DATA

- 86.3 kg yr of  $\text{TeO}_2$ , same data used for 0vDBD analysis
- split data into inner and outer (2 crystals thick) layers
  - outer layers more sensitive to external backgrounds
- split data into Multiplicity 1, Multiplicity 2 and Multiplicity 2 Sum
  - different multiplicities are sensitive to different types of backgrounds

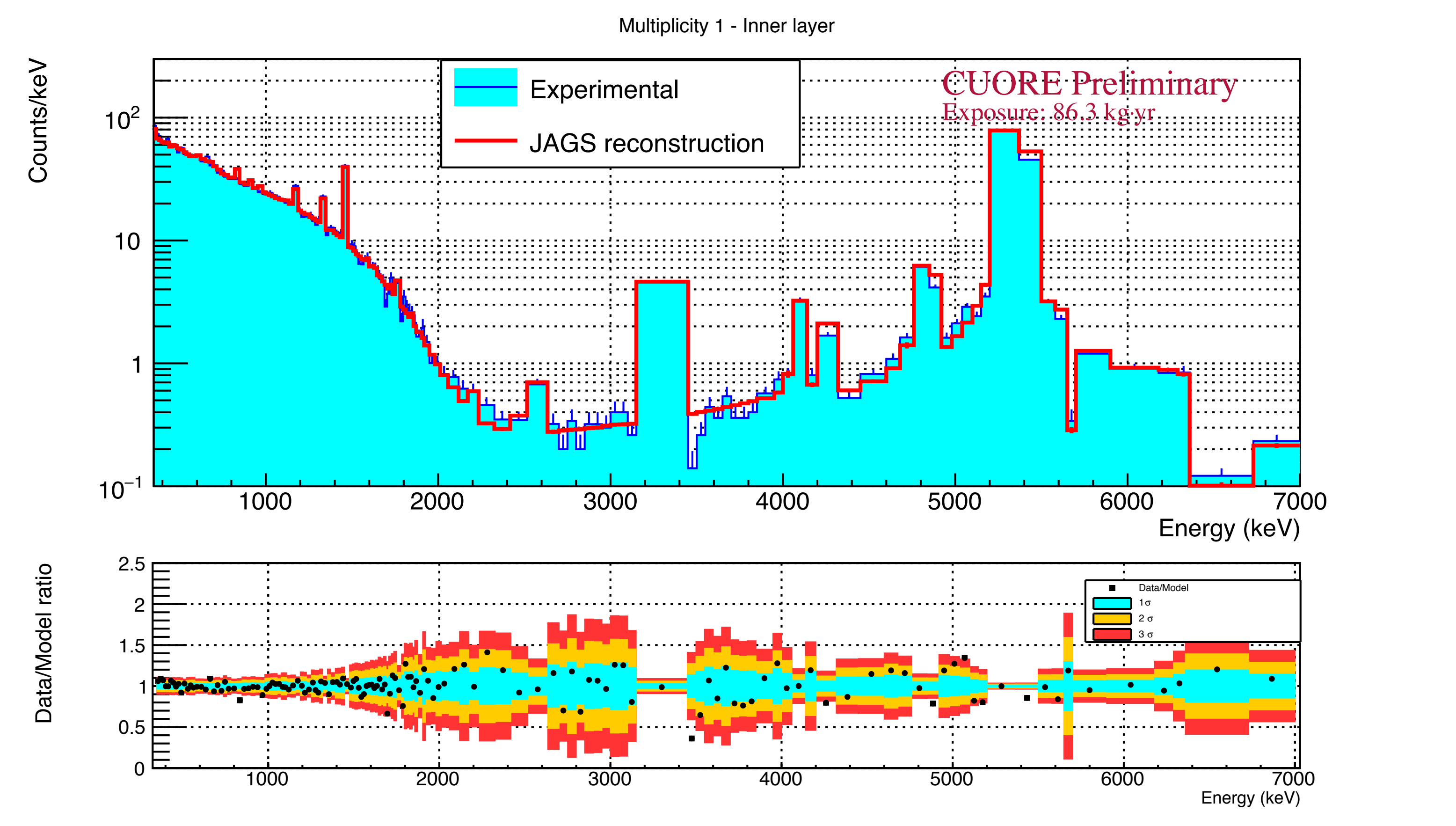
## BAYESIAN FIT

- Assign to each ingredient (source/location) a normalisation factor
- Assign a prior to each normalisation factor from material screening, assays and cosmogenic analysis
- Fit the model to the data and sample the posteriors with MCMC Gibbs sampler

## MONTE CARLO

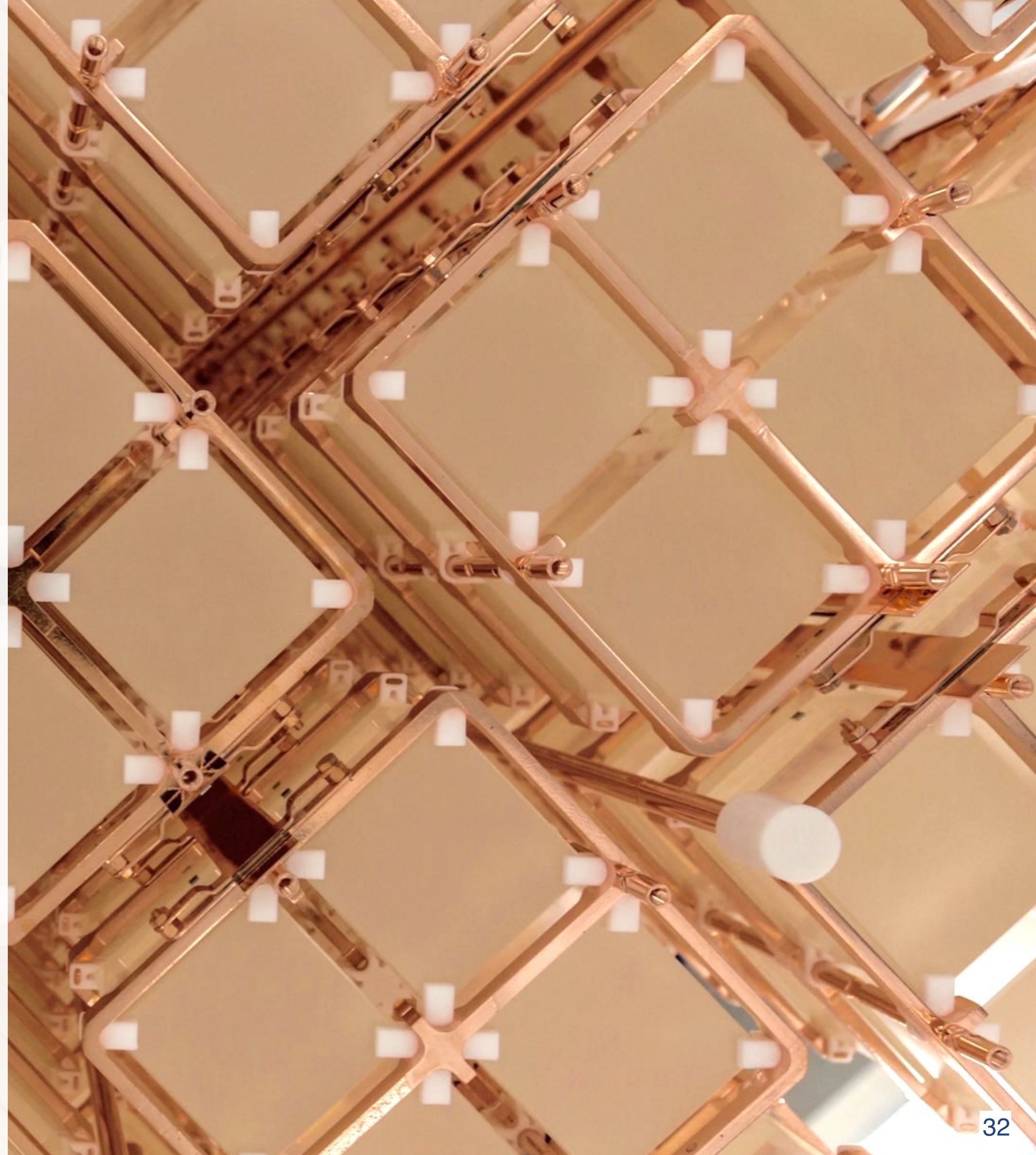
- ~60 independent simulations of sources/location in the setup
- full radioactive chains and single isotopes
- the different energy spectra (inner/outer, M1/M2, M2sum) of each source/location are generated with Geant4 based simulation implementing a detailed geometry of the setup (detector, cryostat, shields)

## BAYESIAN FIT



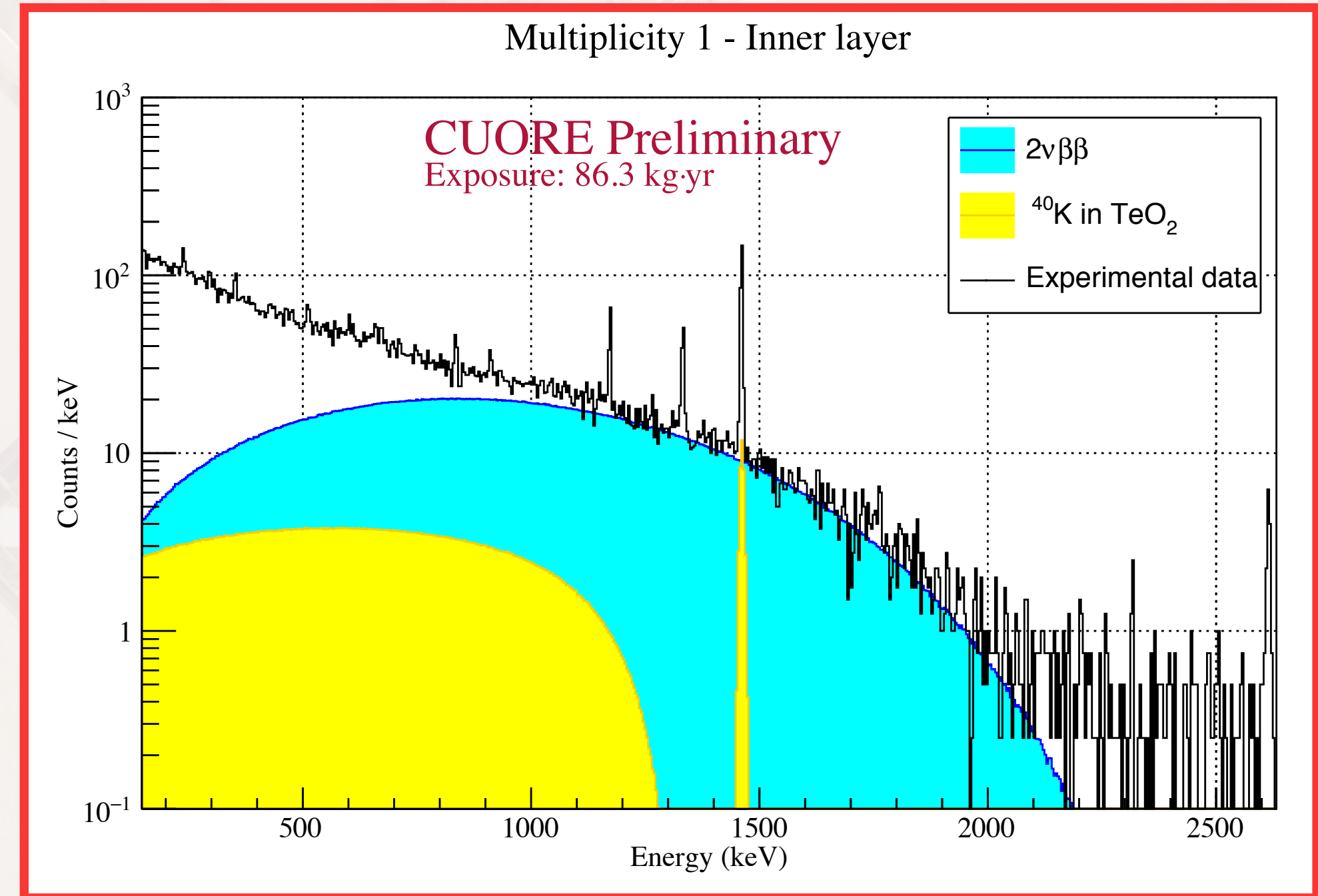
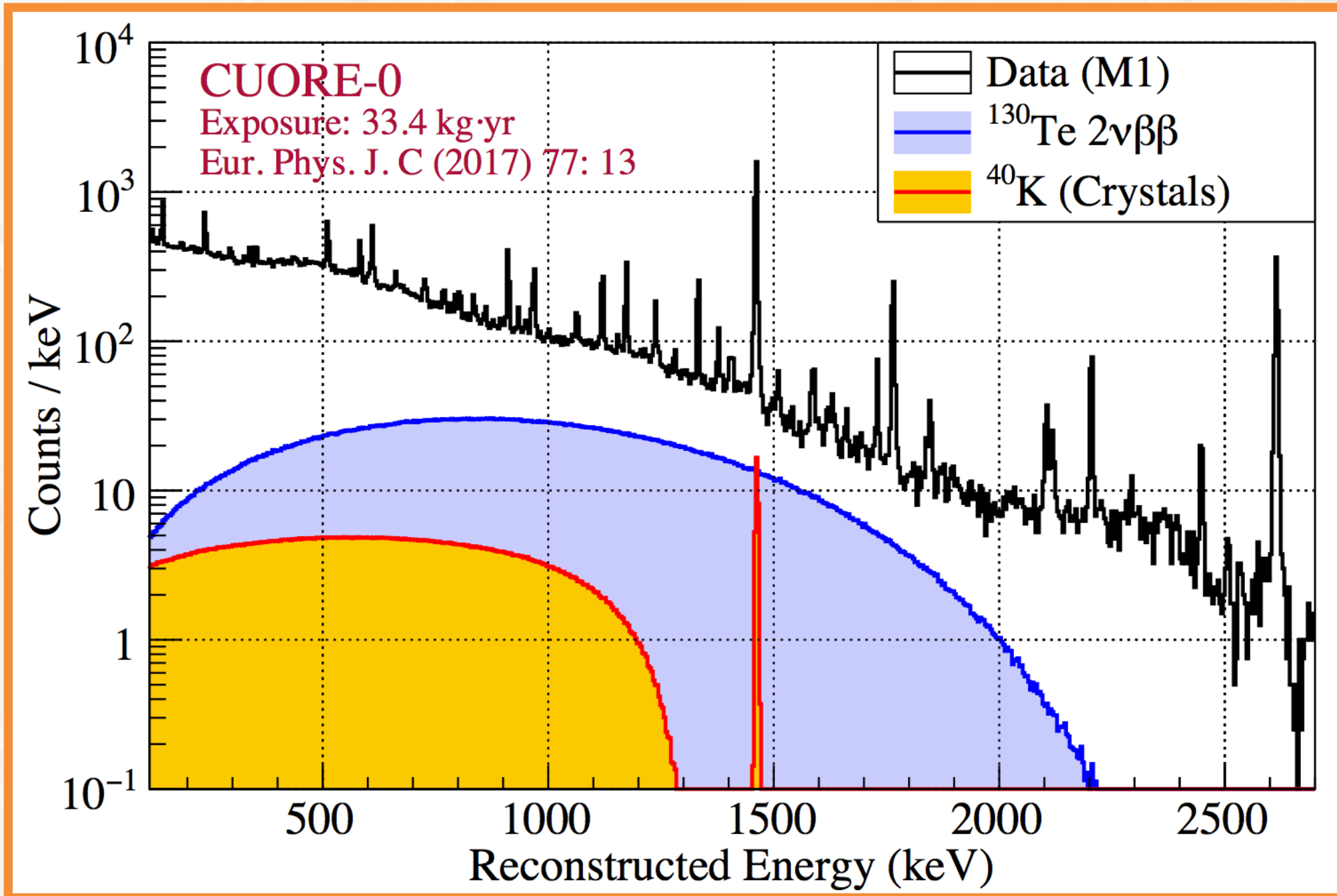
# Outline

- TeO<sub>2</sub> and thermal detectors for neutrino-less DBD
- CUORE setup
  - Location
  - Cryogenic system
  - Calibration system
  - Shielding
- CUORE roadmap
- Analysis procedures
  - Calibration and detector response
  - Event selection and topology
- **Physics results**
  - $0\nu\beta\beta$  and lepton number violation
  - background model
  - $2\nu\beta\beta$
- Conclusions and outlook





2νDBD spectrum is one of the ingredients in the background model



**CUORE-0:**

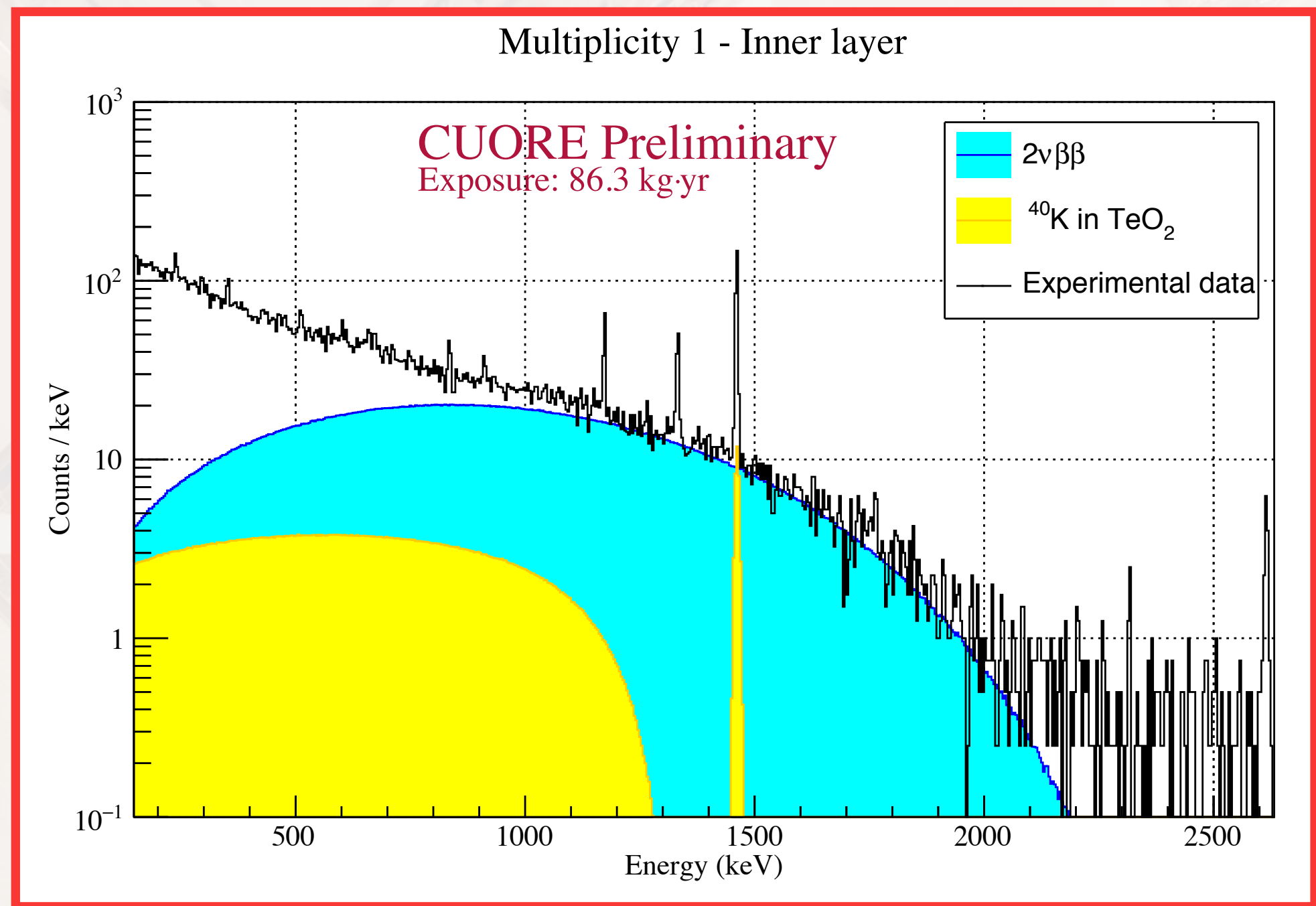
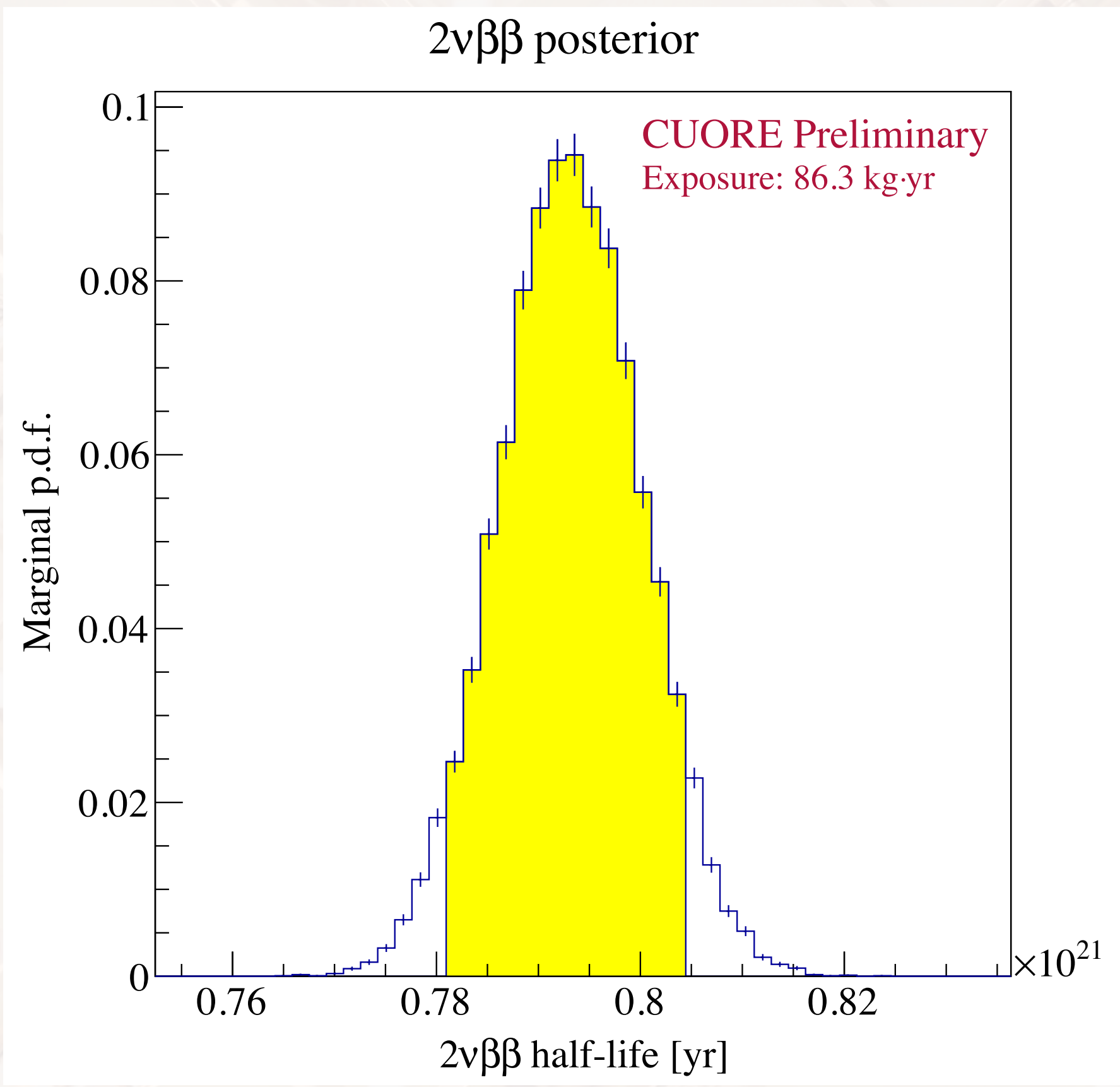
~ 20% of counts in 1-2 MeV region

- lower background
- self-shielding
- anti-coincidence efficiency

**CUORE:**

~ 100% of counts in 1-2 MeV region

2νDBD spectrum is one of the ingredients in the background model



**CUORE:**  
**~ 100% of counts in 1-2 MeV region**

2νDBD spectrum is one of the ingredients in the background model

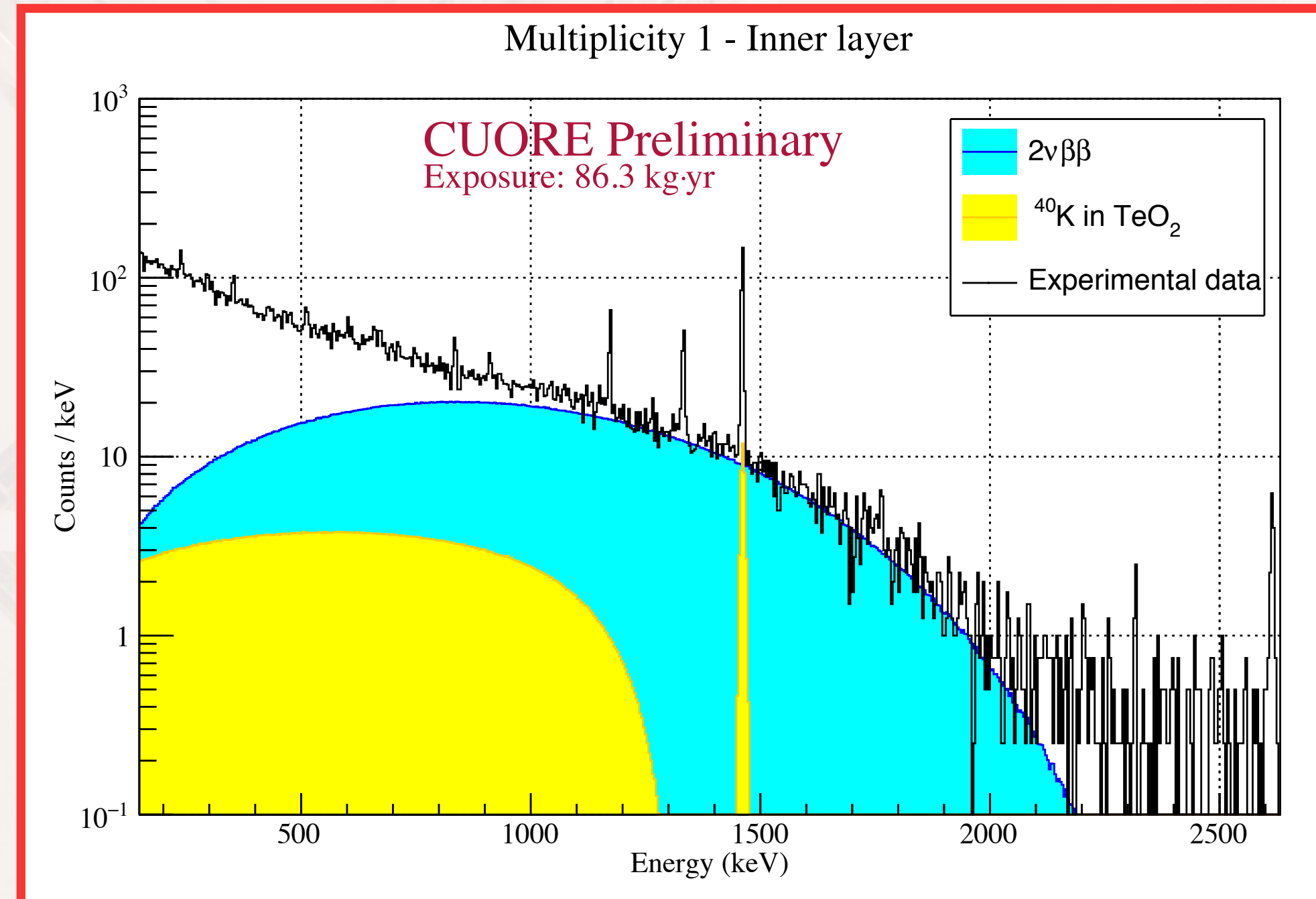
$$\Gamma_{1/2}^{2\nu} = [8.7 \pm 0.1(\text{stat.}) \pm 0.2(\text{syst.})] \times 10^{-22} \text{yr}^{-1}$$

$$T_{1/2}^{2\nu} = [7.9 \pm 0.1(\text{stat.}) \pm 0.2(\text{syst.})] \times 10^{20} \text{yr}$$

$$\text{CUORE} - 0 : T_{1/2}^{2\nu} = [8.2 \pm 0.2(\text{stat.}) \pm 0.6(\text{syst.})] \times 10^{20} \text{yr}$$

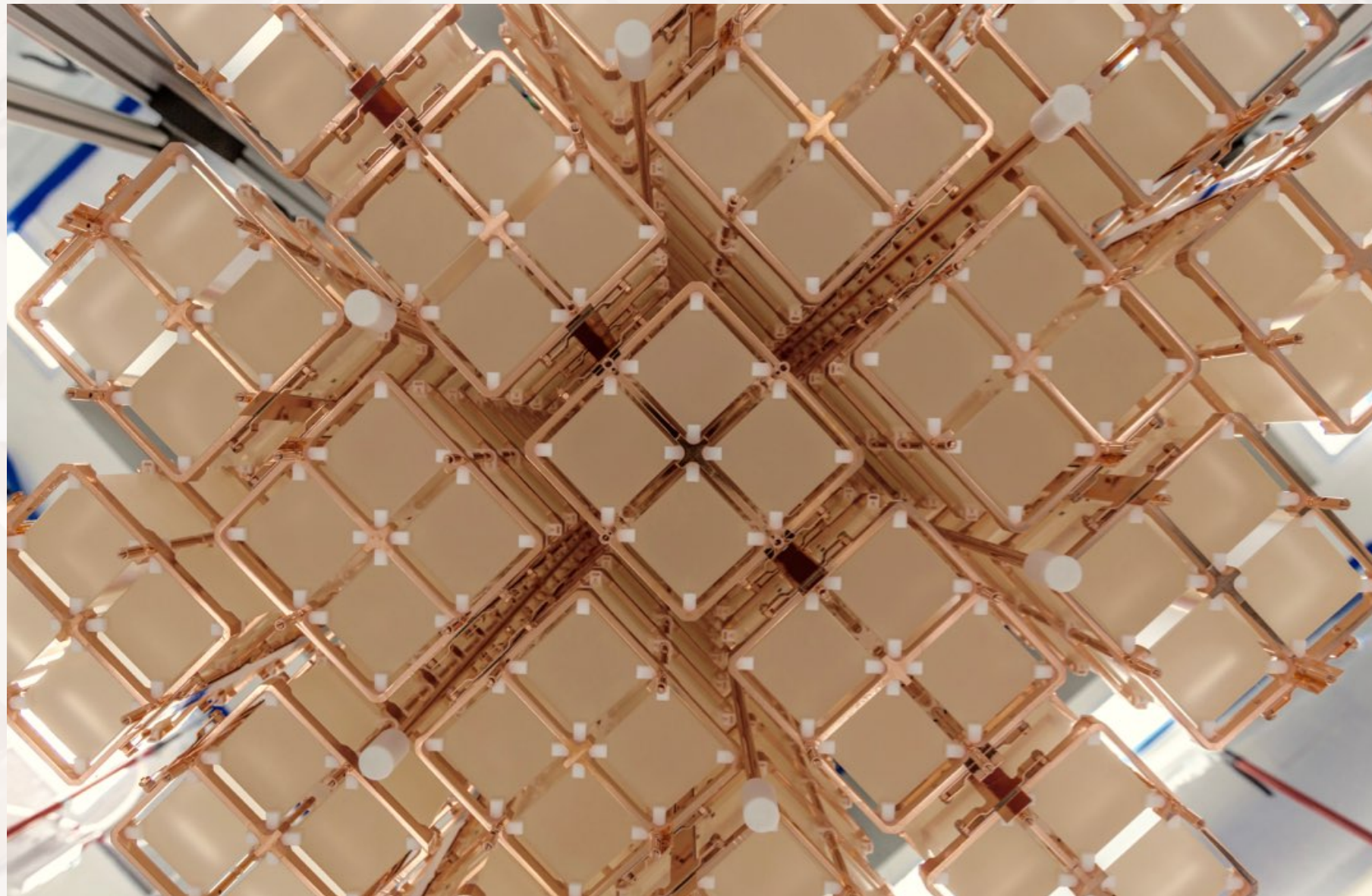
$$\text{NEMO} - 3 : T_{1/2}^{2\nu} = [7.0 \pm 0.9(\text{stat.}) \pm 1.1(\text{syst.})] \times 10^{20} \text{yr}$$

- Systematic uncertainty dominated by uncertainty on contaminant location and uniformity
  - studied by repeating the fit with different geometric splitting of the data
  - more data → finer geometric splitting of the array → reduce this component
- Fit is independent of energy threshold over the range 100 - 750 keV
- Used M2 data (purer sample of particle events) to improve the systematic uncertainty on the selection efficiency by one order of magnitude w.r.t. PRL 2018 data release



**CUORE:**  
**~ 100% of counts in 1-2 MeV region**

- With the first two datasets CUORE have:
  - accumulated a total exposure of almost 100 kg·y
  - Invaluable operational experience
  - collected important information on detector performance, noise, resolutions, background levels
  - pushed for the first time the limit on neutrino-less double beta decay half life of  $^{130}\text{Te}$  beyond  $10^{25}$  years
  - performed the most precise measure of two-neutrino half life to date



- The largest and most complex cryogenic experiment is taking physics data
- The first analysis efforts were focused on the neutrino-less and two-neutrino double beta decay of  $^{130}\text{Te}$  to GS
- Physics results on more processes are on their way:
  - Majoron emission, CPTV
  - Dark matter, axions
  - $\beta\beta$  to excited states,  $\beta^+/\text{EC}$  decays
- With an unprecedented amount of data, CUORE is the best tool to study and model the backgrounds for the next generation experiments