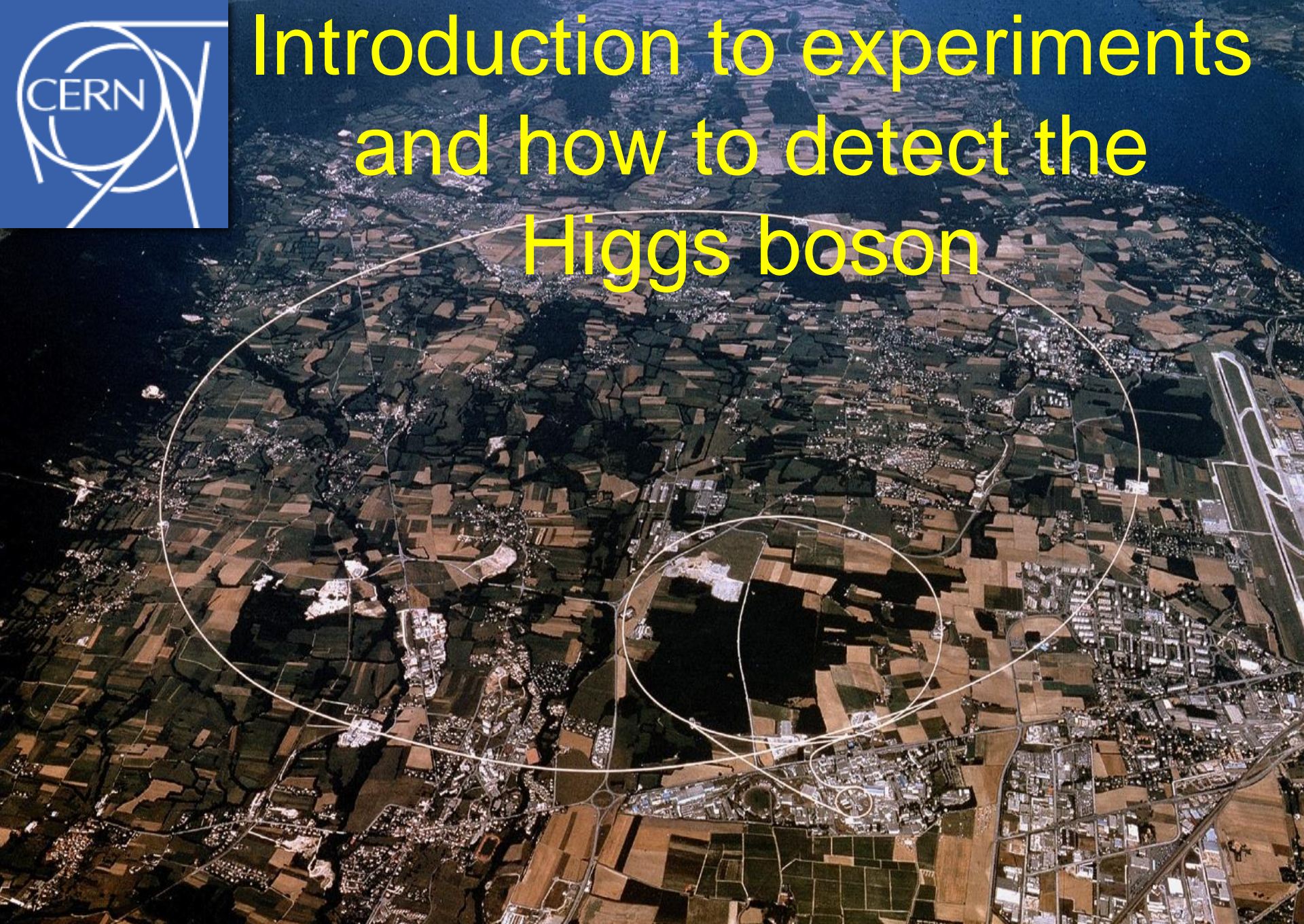




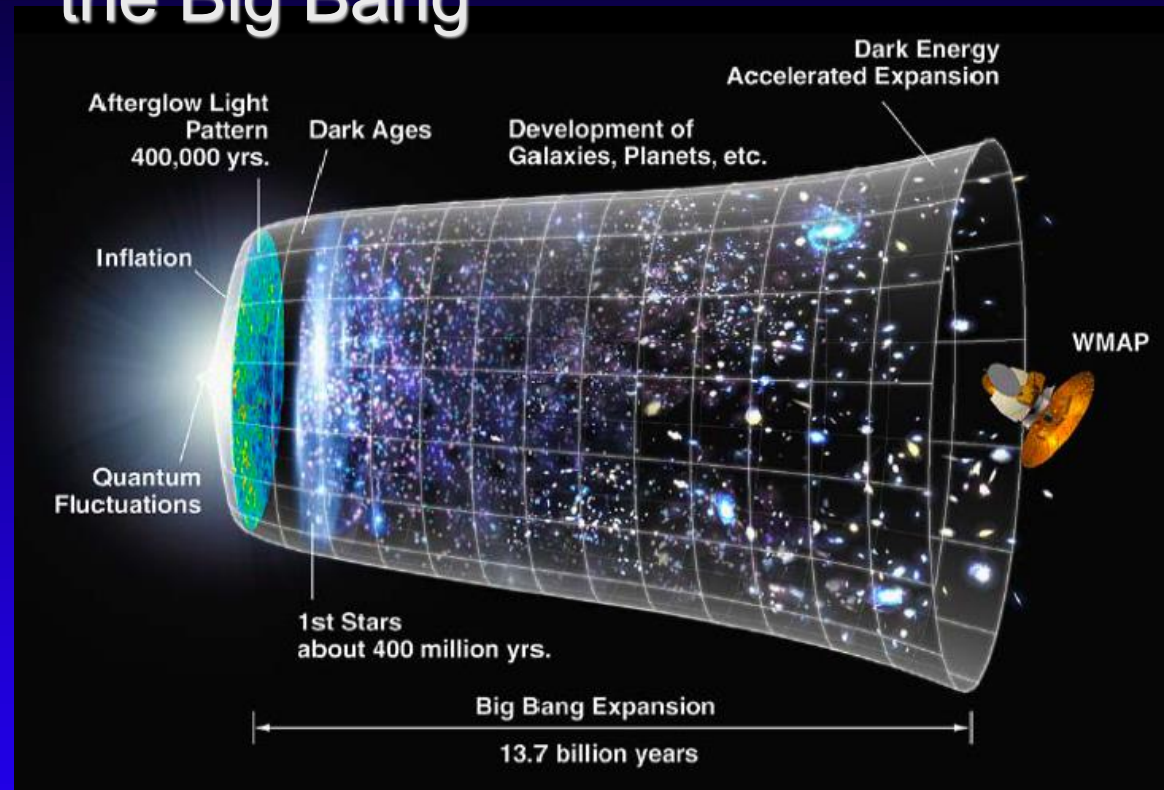
# Introduction to experiments and how to detect the Higgs boson



# Next Scientific Challenge: to understand the very first moments of our Universe after the Big Bang

## Theories .....

- origin of mass
- Dark matter
- extra-dimensions

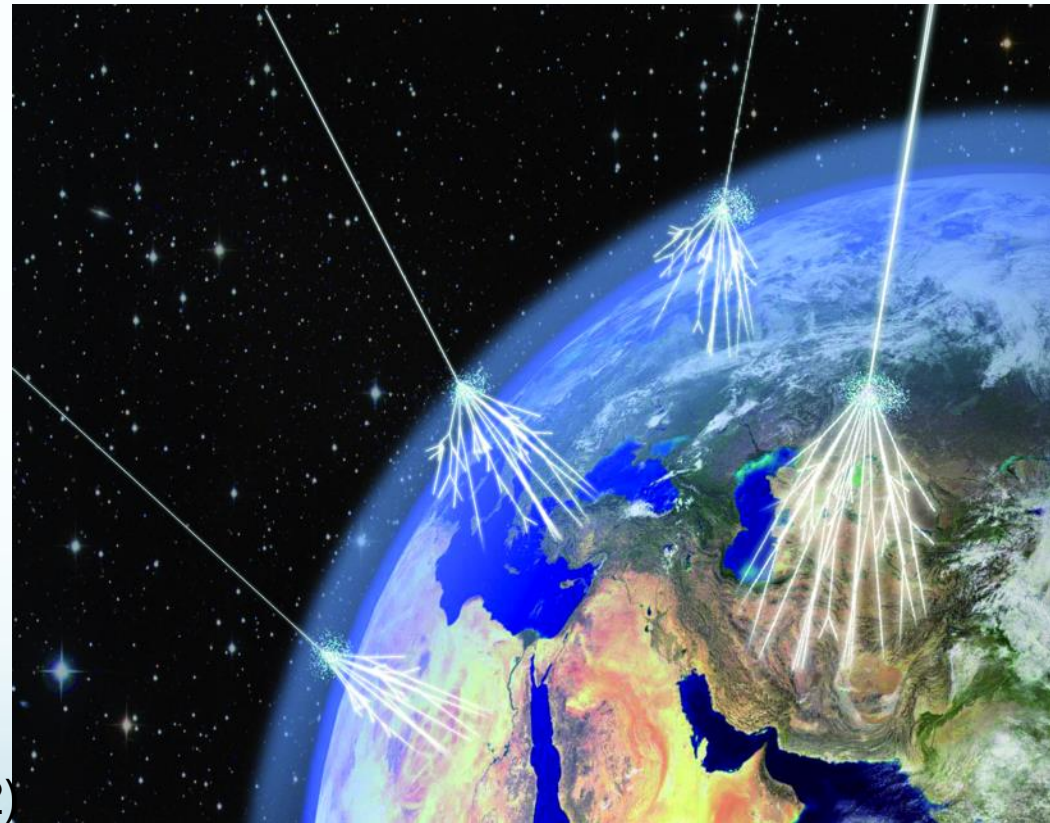


...are **tested experimentally** by reproducing  
conditions  $\sim 10^{-12}$  sec after the Big Bang

Cosmic rays are used to study the performance of the detector. Free of charge! 😊



Hess received the Nobel Prize in Physics in 1936 for his discovery (1912)



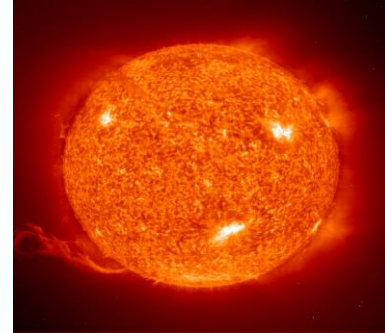
# Each second....

- ....for each square meter
- ~200 particles

- **Where are they coming from ?**
- **What are they?**
- **Which information are they carrying?**

# Where are they from?

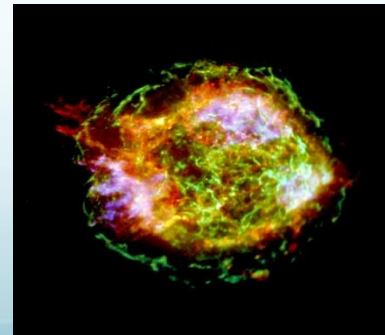
- From the Sun



- From Galaxies



- From Supernovae



# Primary Cosmic Rays...

Primary Cosmic Rays

- **Composition is unclear**

- **From Space**

- **Mostly protons...**

- **They interact in the atmosphere**

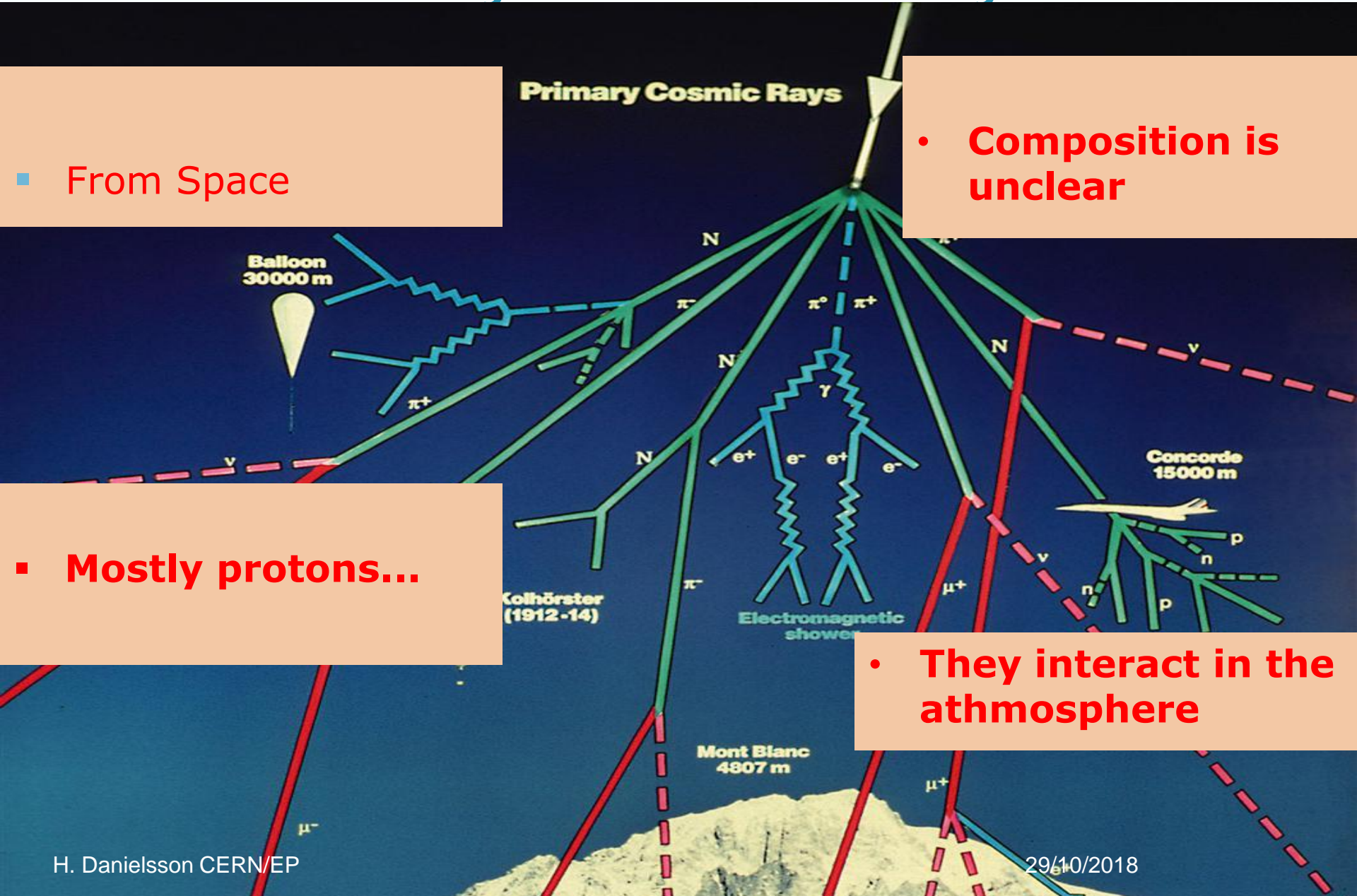
Balloon  
30000 m

Concorde  
15000 m

Kolhörster  
(1912-14)

Mont Blanc  
4807 m

Electromagnetic  
showers



2013 NOBEL PRIZE IN PHYSICS

# François Englert Peter W. Higgs



© The Nobel Foundation. Photo: Lovisa Engblom.

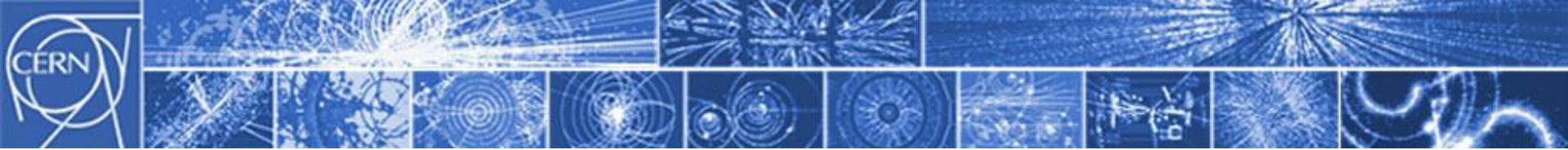


8 October 2013

The Royal Swedish Academy of Sciences has decided to award the Nobel Prize in Physics for 2013 to

**François Englert and Peter Higgs**

*“for the theoretical discovery of a mechanism that contributes to our understanding of the origin of mass of subatomic particles, and which recently was confirmed through the discovery of the predicted fundamental particle, by the ATLAS and CMS experiments at CERN’s Large Hadron Collider”*



# Outline

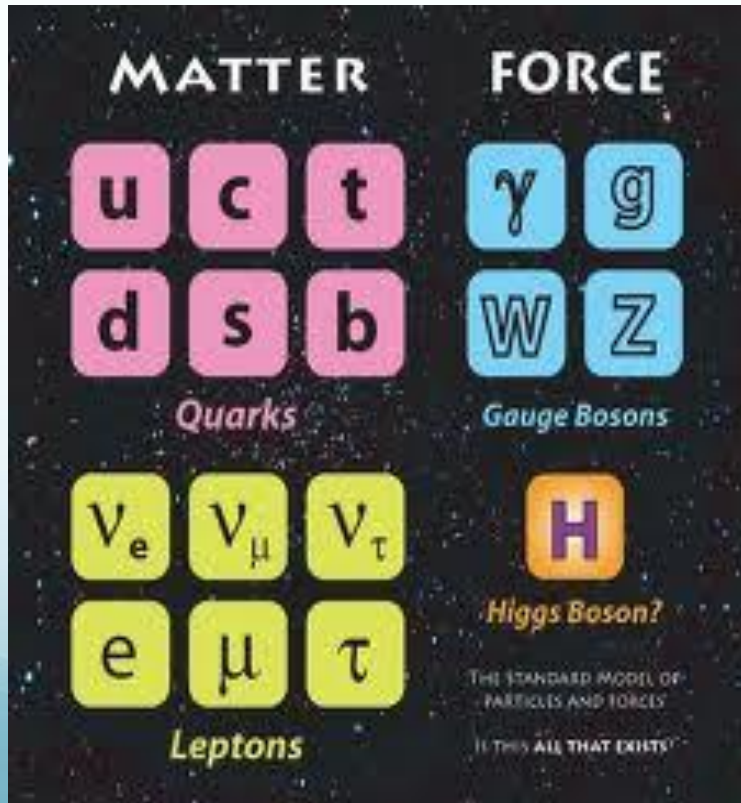
- Introduction
  - Matter, forces, particles
- CERN and the Large Hadron Collider (LHC)
  - The accelerator
  - The detectors
- The Higgs discovery
- Are we finished now ?





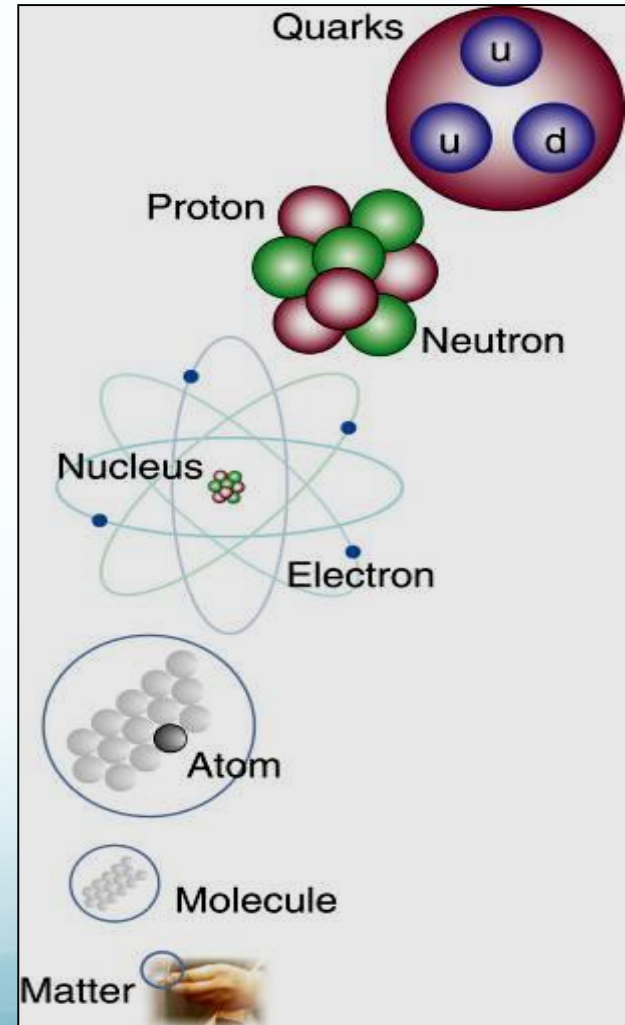
# The Standard Model (1970-90s)

- Matter particles: fermions (1/2 integer spin)
- ‘Force’ particles: bosons (integer spin)
- Higgs field causes electro weak symmetry breaking and gives particles their masses



→ Nucleon level (partons) : binding energy ~98% of the mass

→ Most of the (luminous) mass in the universe comes from QCD confinement energy



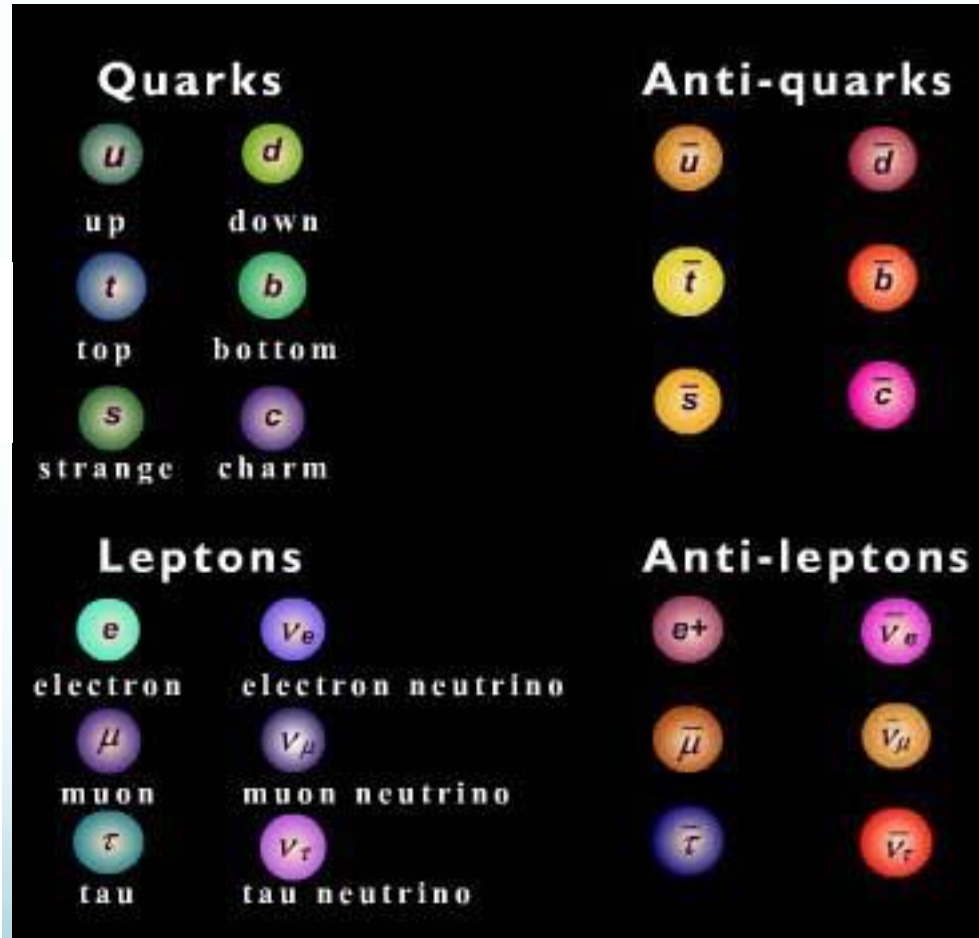
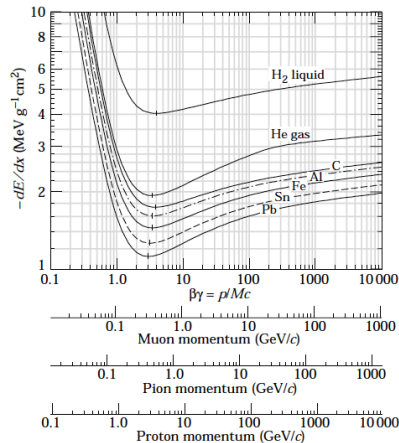
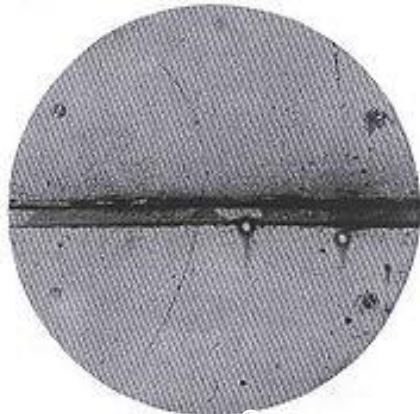
# The Standard Model

- Is a very successful theory and describes the world around us.
- The Standard Model is a discovery in itself
- However, it may explain only a fraction of the universe ( $\sim 5\%$ ) (or something else....)
  - 95% is dark energy and dark matter. What is made of? The search is ongoing...
  - What about super symmetry (SUSY)

# Matter vs antimatter

How does this broken symmetry works ?

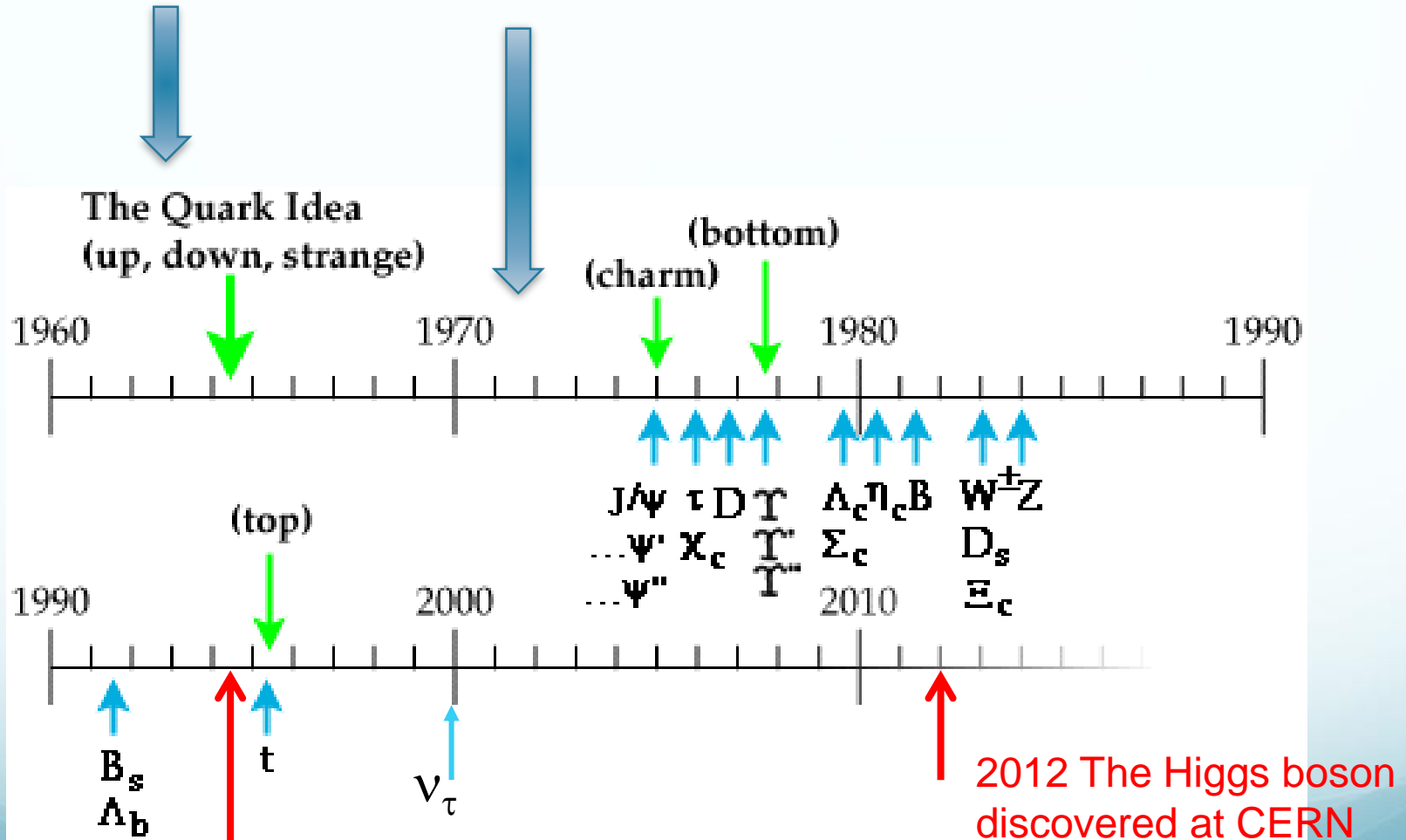
- Einstein  $E=mc^2$
- Paul Dirac  $\left( i\gamma_\mu \frac{\partial}{\partial x_\mu} - m \right) \Psi = 0$
- Carl Anderson discovered the positron in 1932 in a cloud chamber



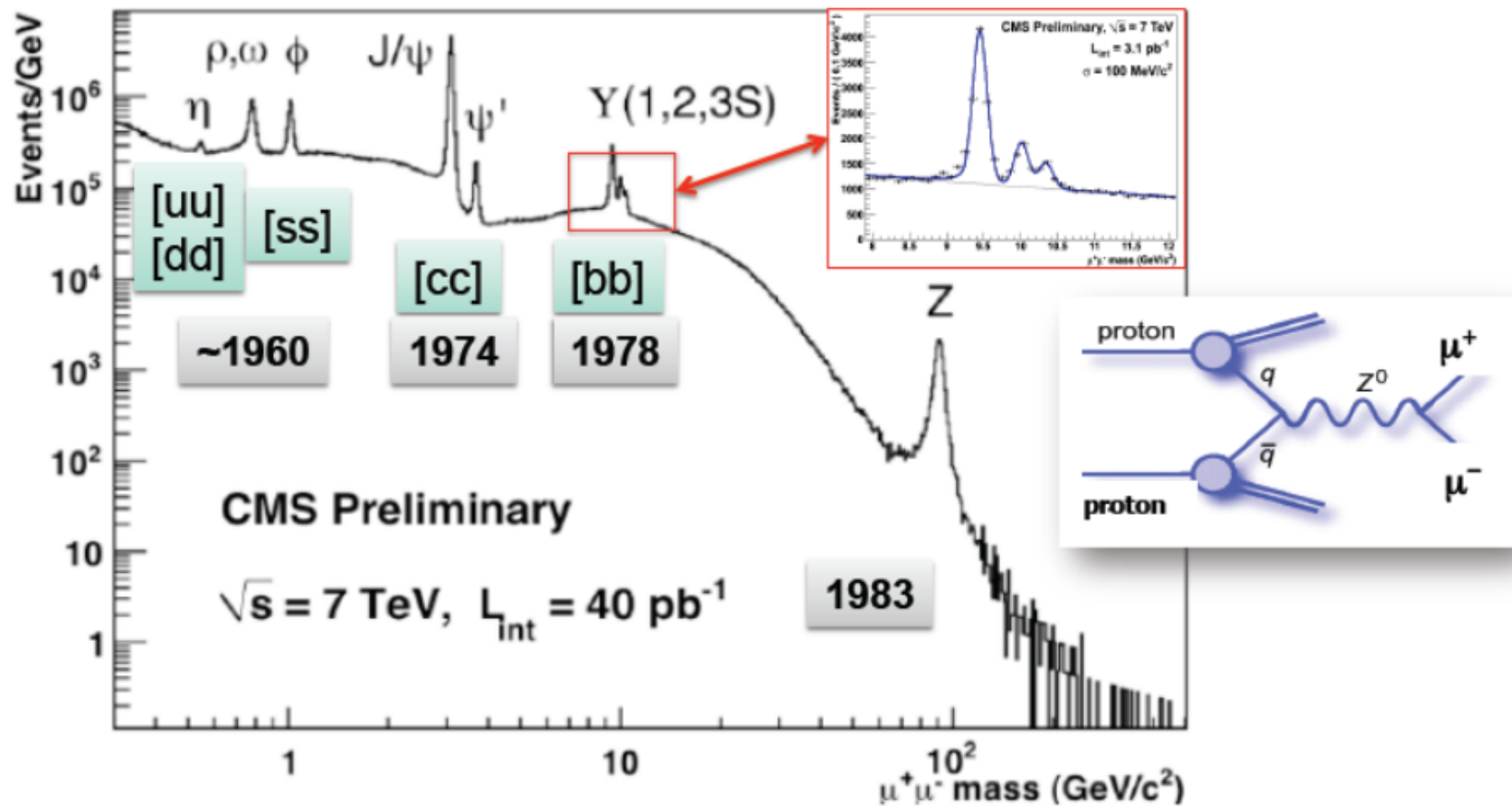
# A bit of history

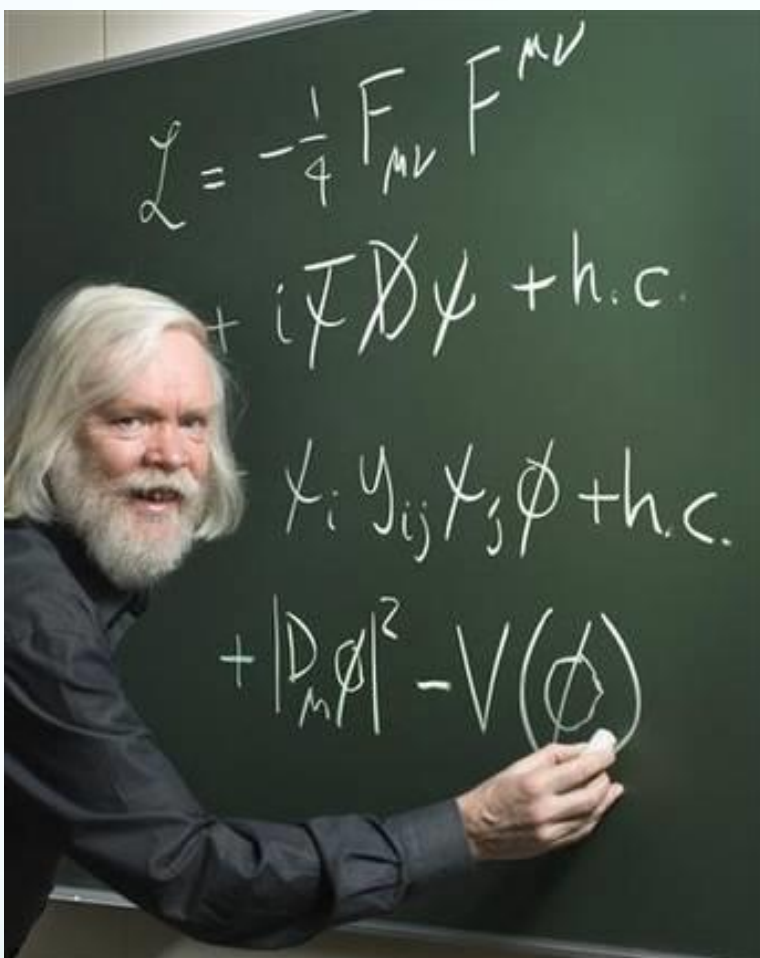
1964 The BEH mechanism

The Standard Model completed



# After 10 min of LHC running: full history of SM





In 1976:

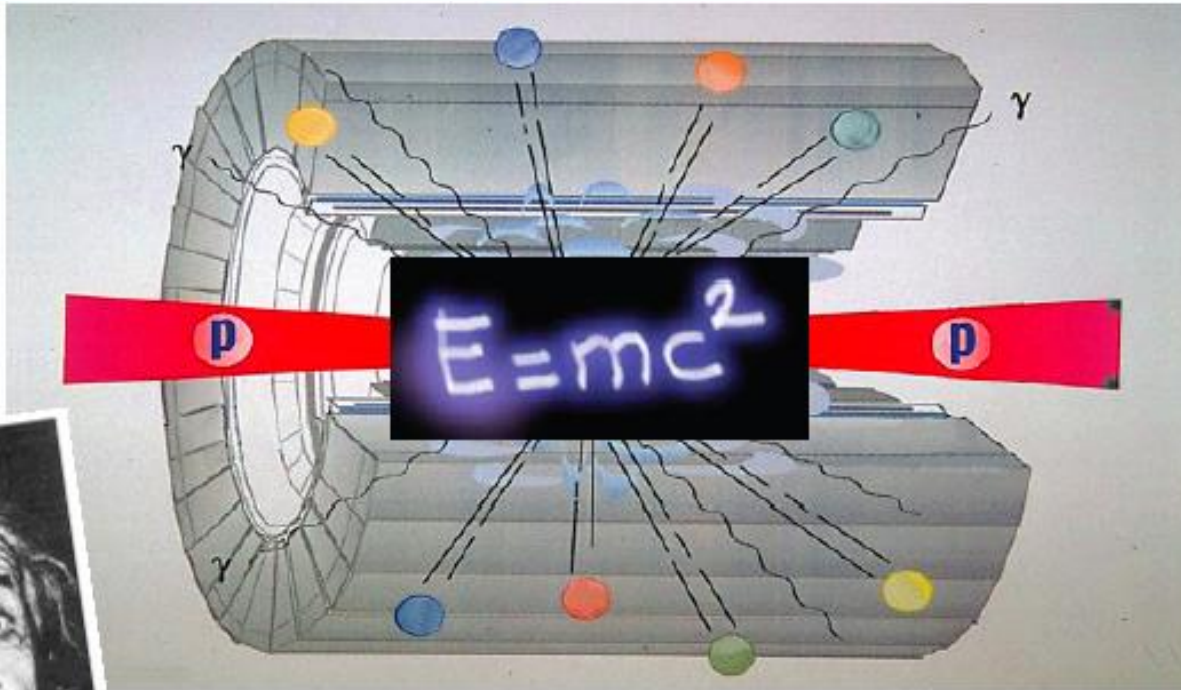
A PHENOMENOLOGICAL PROFILE OF THE HIGGS BOSON

John Ellis, Mary K. Gaillard <sup>\*)</sup> and D.V. Nanopoulos <sup>+) )</sup>  
CERN -- Geneva

*The Roadmap:*

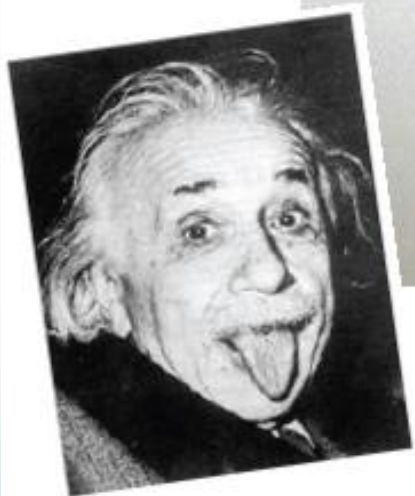
We should perhaps finish with an apology and a caution. We apologize to experimentalists for having no idea what is the mass of the Higgs boson, unlike the case with charm <sup>3),4)</sup> and for not being sure of its couplings to other particles, except that they are probably all very small. For these reasons we do not want to encourage big experimental searches for the Higgs boson, but we do feel that people performing experiments vulnerable to the Higgs boson should know how it may turn up.

# How ?

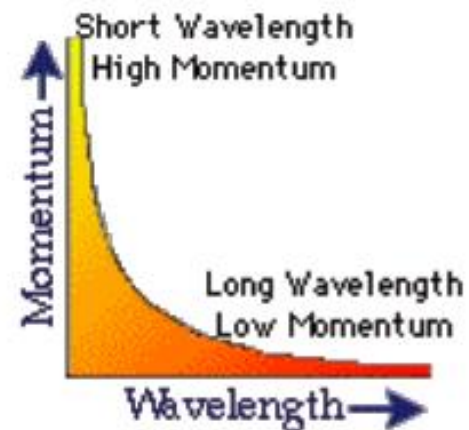


$E=3.5\text{TeV} \rightarrow$   
 $V=99.999996\%$  of  $c$

$$m = \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}}$$

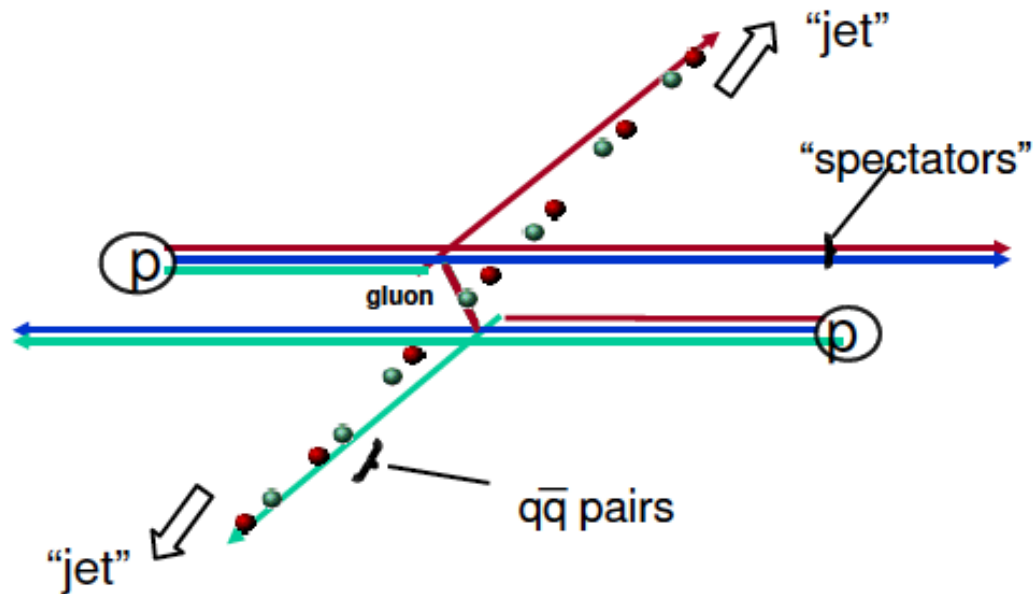
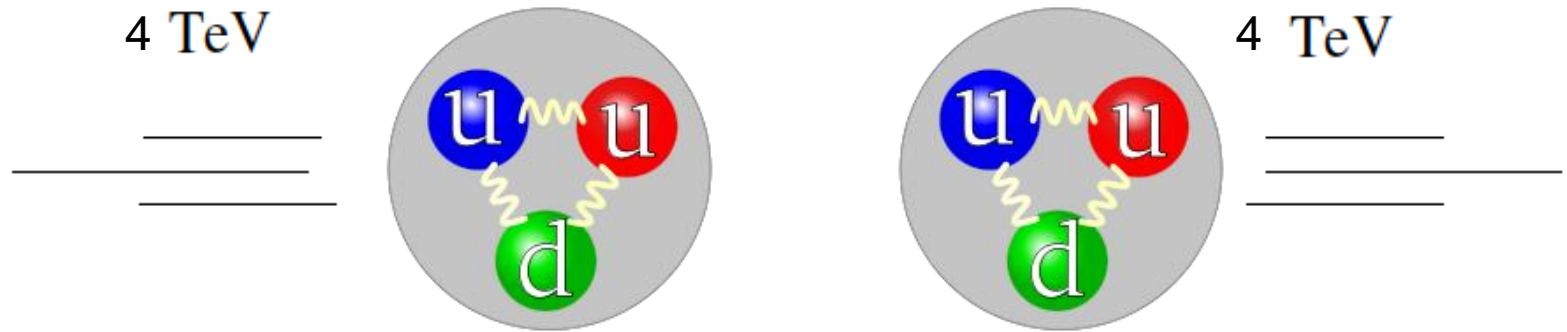


Energy = Matter  
 $E^2 = (m_0 c^2)^2 + (pc)^2$



# Experimental High Energy Physics – detecting particles

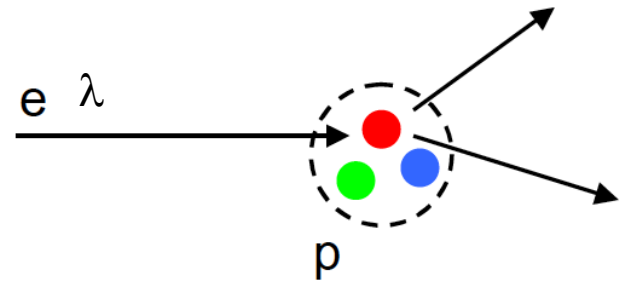
Two Protons collide at high energy  
Large Hadron Collider (LHC) at CERN





## On example: the discovery of the quarks at SLAC in 1968

$$l = \frac{h}{p}, P = 20\text{GeV} \Rightarrow l \gg 10^{-17}\text{m}$$



- The quark model was independently proposed by physicists [Murray Gell-Mann](#) and [George Zweig](#) in 1964.
- Gell-Mann found the quarks in:

“Three quarks for Muster Mark!  
Sure he has not got much of a bark  
And sure any he has it's all beside the mark.”

—James Joyce, *Finnegans Wake*



Center-of-Mass Energy (Nominal)  
14 TeV ?

Center-of-Mass Energy (close to nominal)  
13 TeV

5/2017

Restart in 2015

*LHCb*

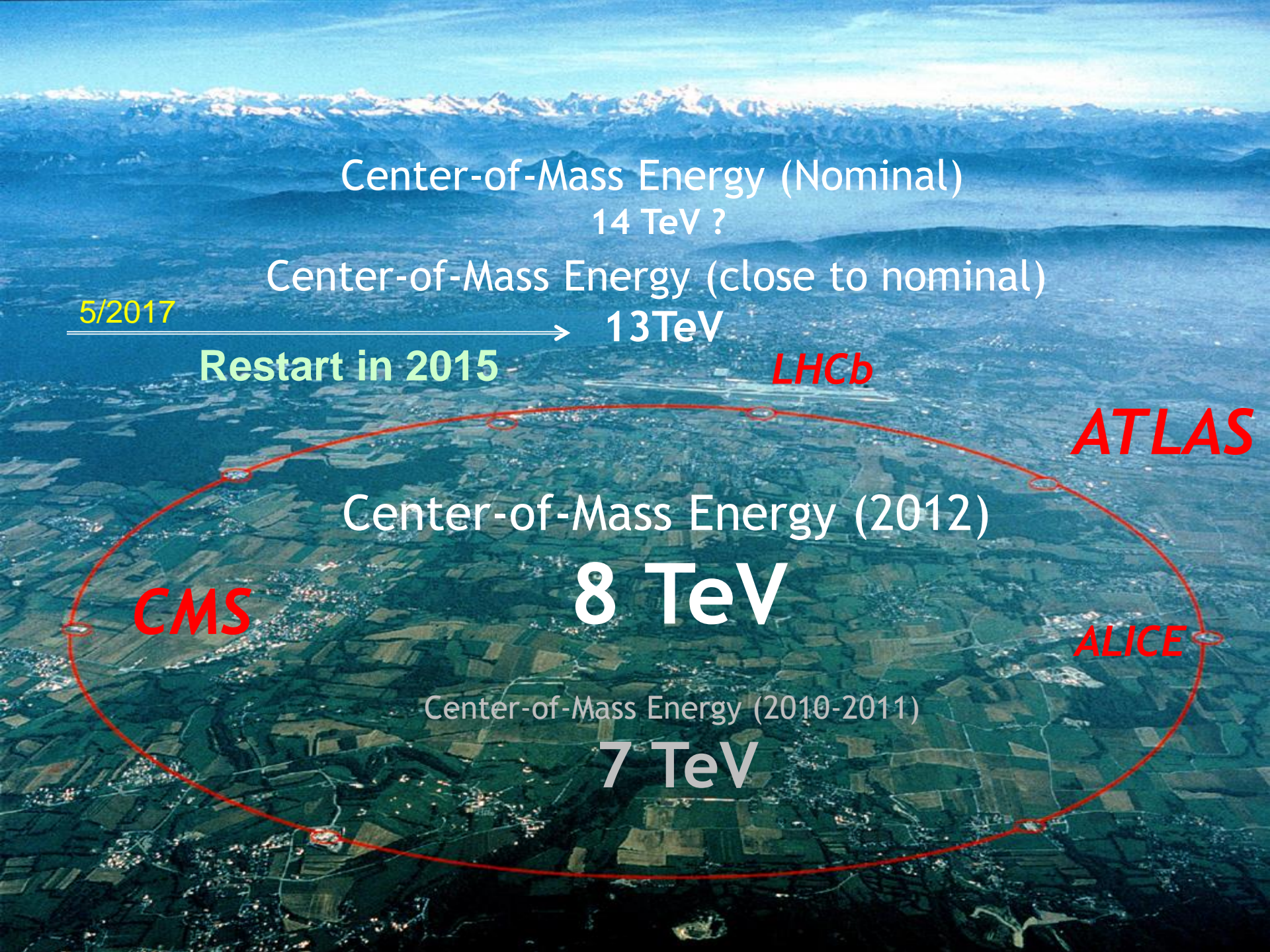
*ATLAS*

Center-of-Mass Energy (2012)  
8 TeV

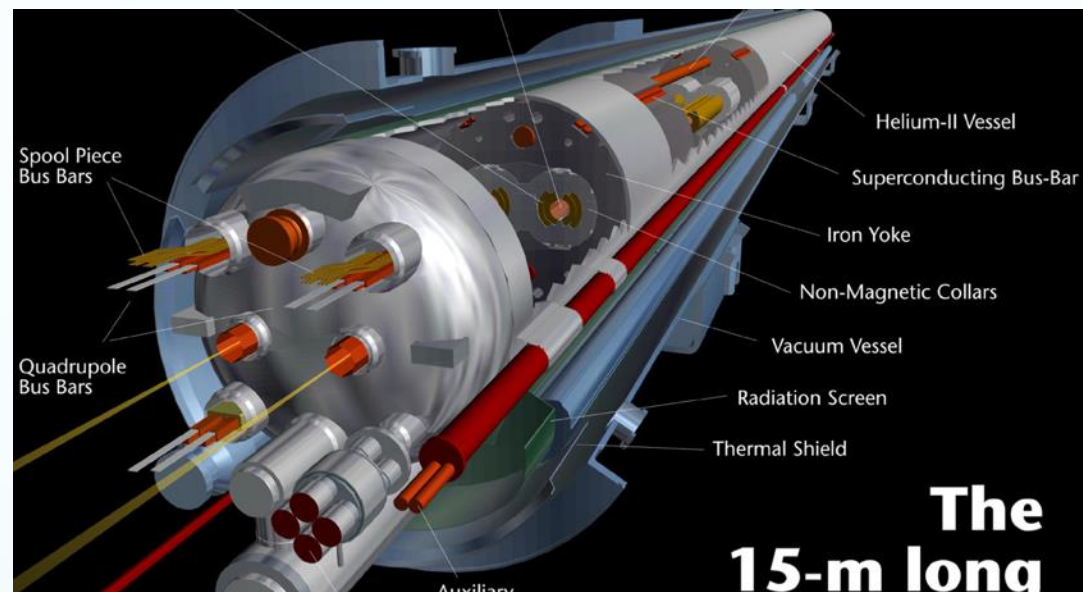
*CMS*

*ALICE*

Center-of-Mass Energy (2010-2011)  
7 TeV

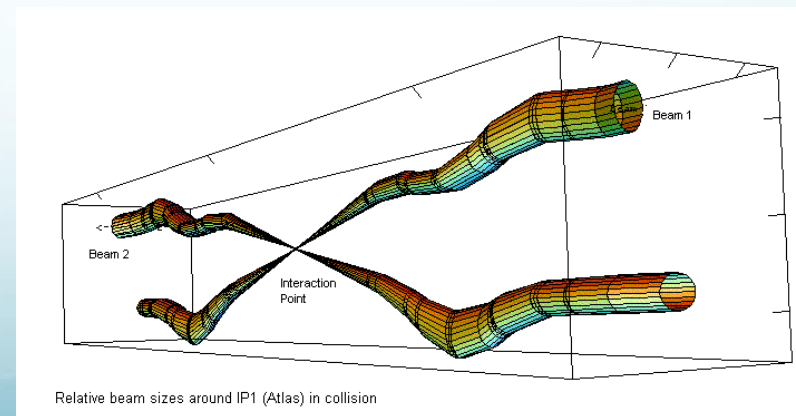


# Large Hadron Collider (LHC)



## • The Accelerator

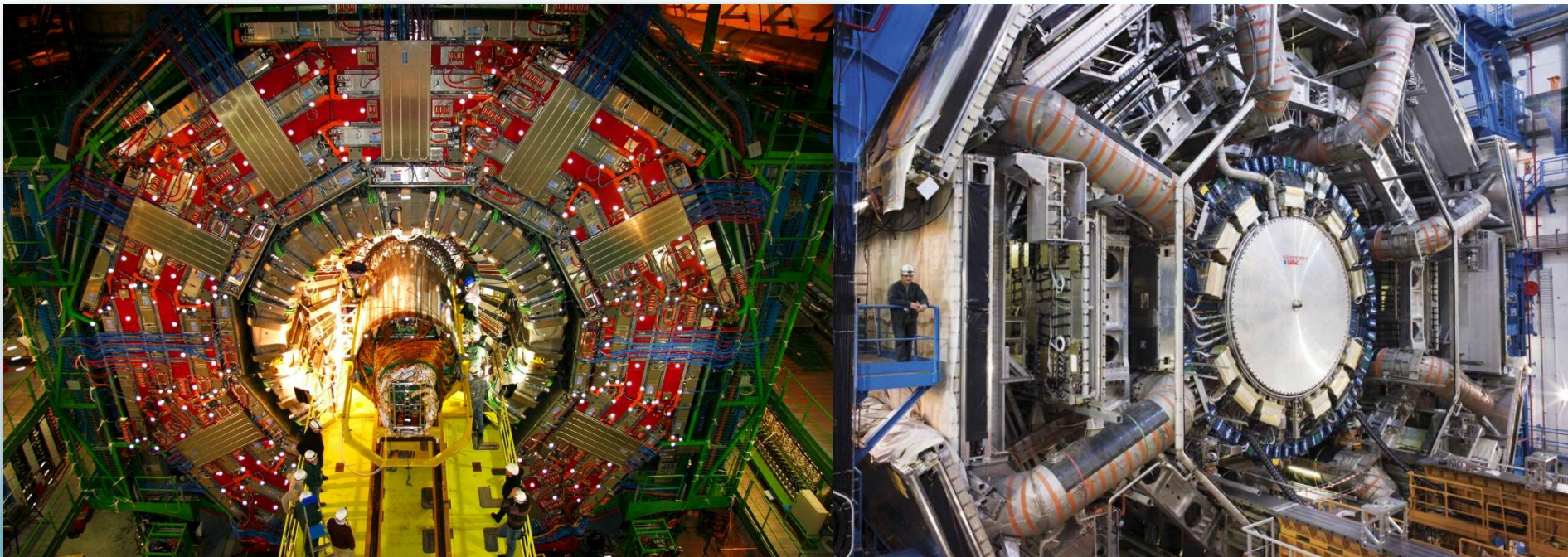
- 100 - 150 m below surface at 1.9 Kelvin in a tunnel 27 km long.
- The protons circulate at a speed of  $\sim 11000$  turns/sec
- There are 2808 bunches
- Collisions at 40 MHz (every 25 ns)
- **600 000 000 collisions per second !**



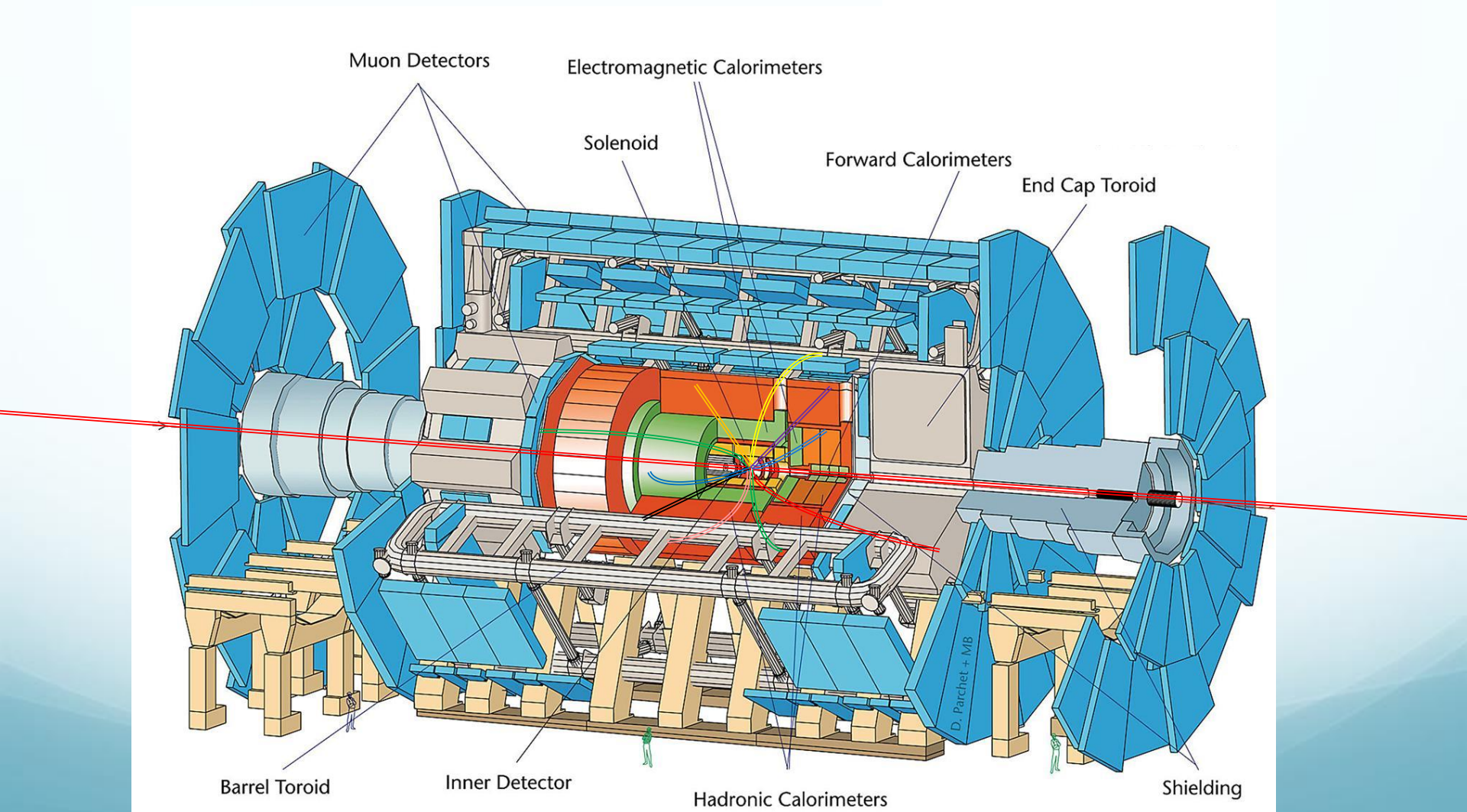
# The experiments

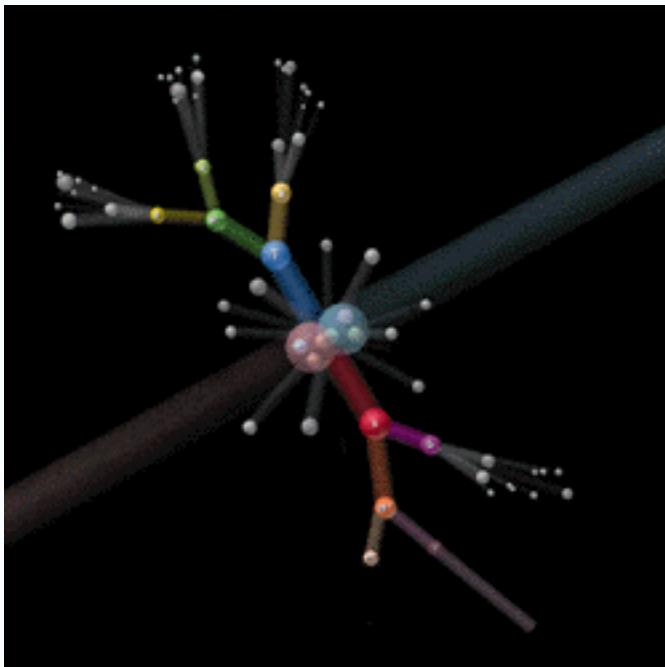
*CMS: heavier than  
the Eiffel Tower*

*ATLAS: as big as a  
5 storey building*



# Största och mest sofistikerade detektorer



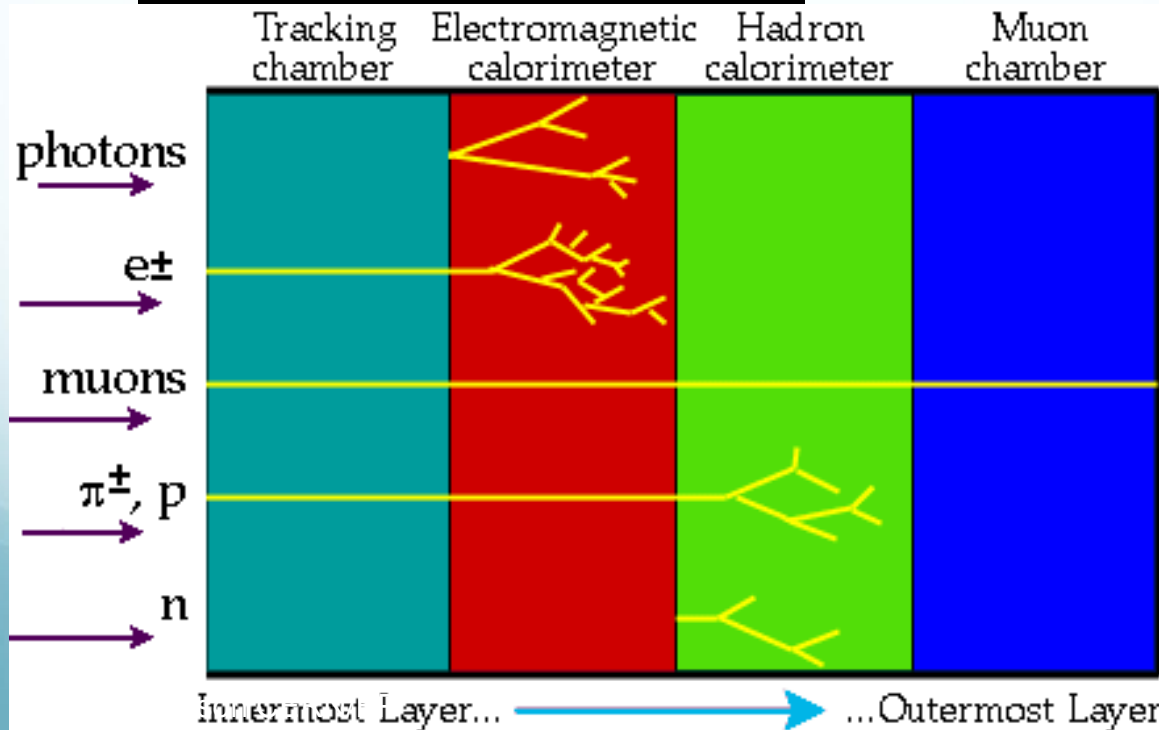


# Principles of Detection

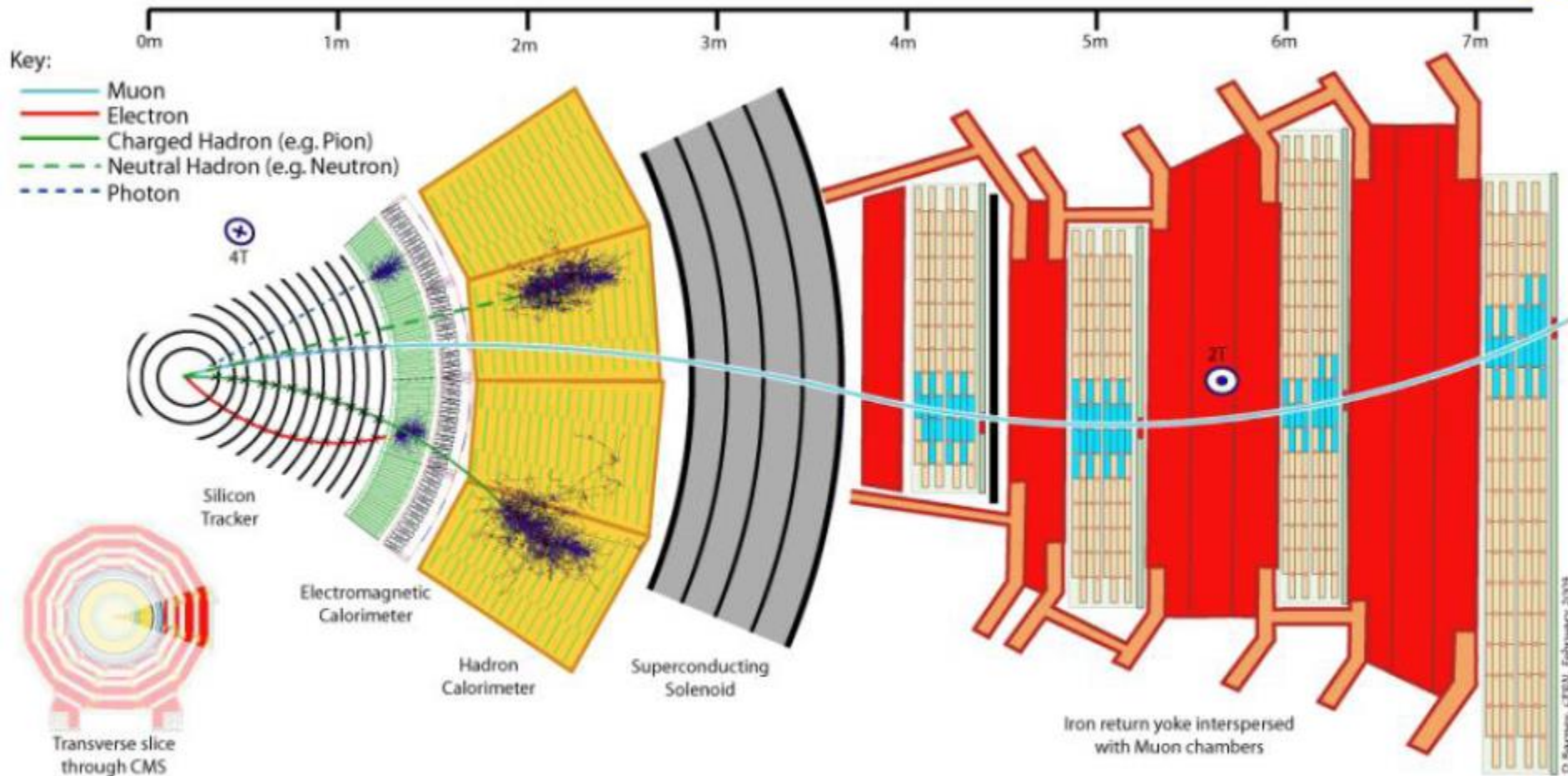


The collision energy condenses into particles ( $e$ ,  $p$ ,  $\pi$ ,  $\mu$ ,  $\gamma$   $K$ ...)

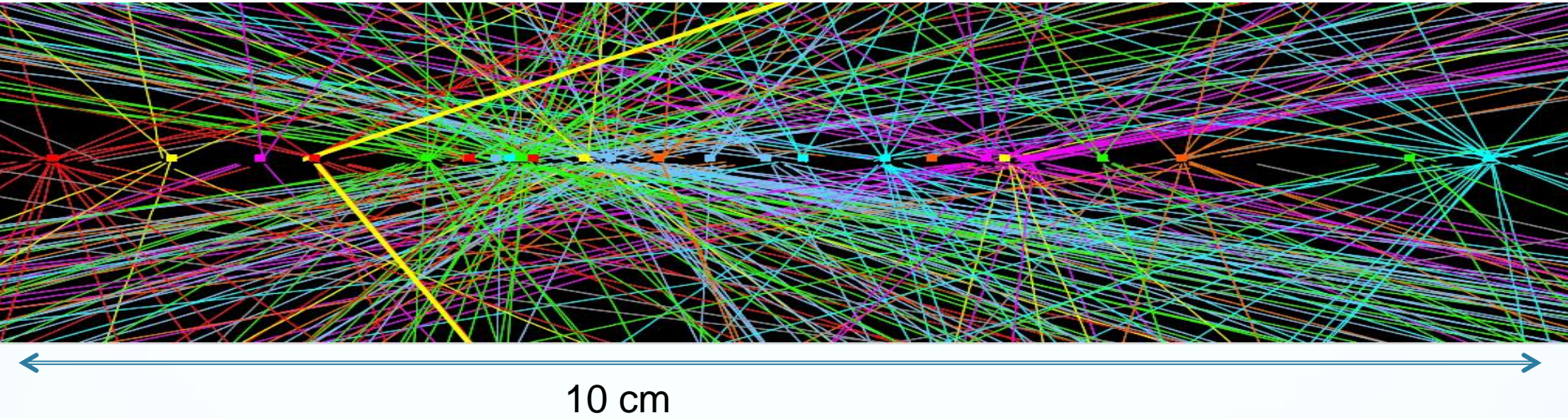
Detectors surrounding the collision point (or *after* in case of fixed target) are sensitive to the passage of energetic particles.



# Partikeldetektorer



# Detector Challenges (Highlights)

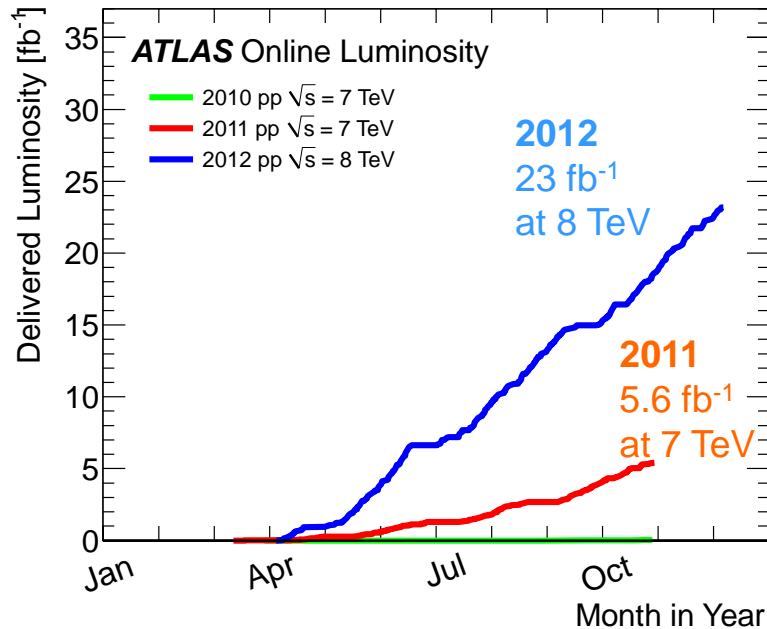


- **Trigger Challenge** : How to select 400 out of  $20 \times 10^6$  events per second while keeping the interesting (including unknown) physics
- **Computing Challenge** : How to reconstruct, store and distribute 400 increasingly complex events per second (over 100 Petabyte per experiment)



# The first LHC run

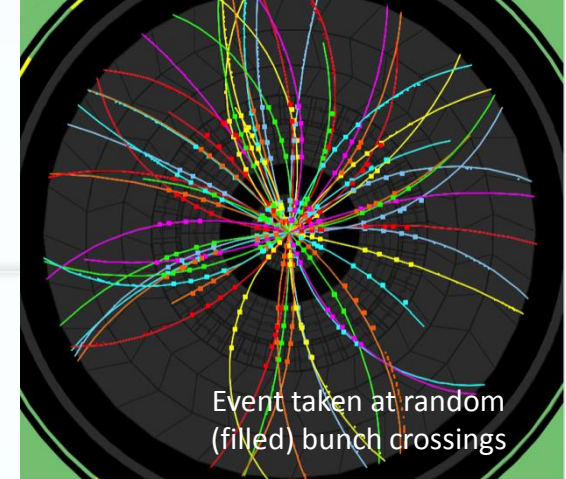
Event rate = luminosity x cross-sections



2010

O(2) Pile-up events

150 ns inter-bunch spacing



2011

O(10) Pile-up events

50 ns inter-bunch spacing

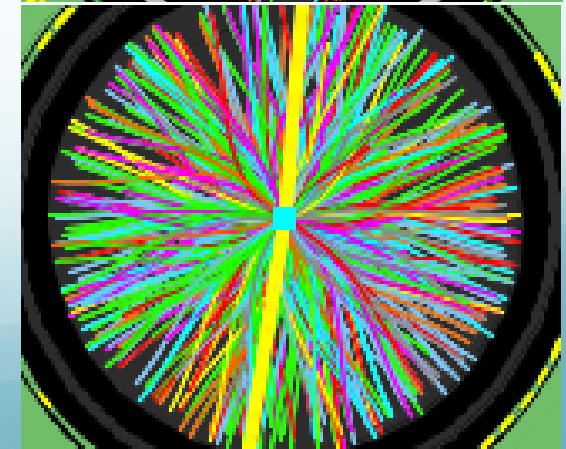


Design value  
(expected to be reached at L=10<sup>34</sup> !)

2012

O(20) Pile-up events

50 ns inter-bunch spacing



# The detection of the Higgs boson

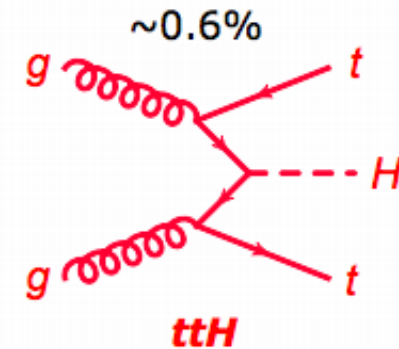
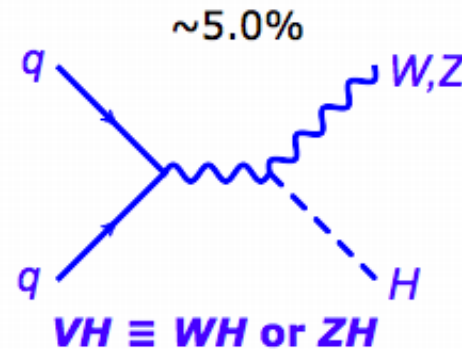
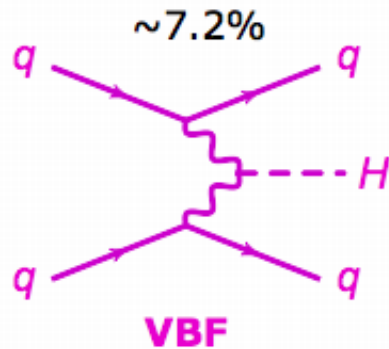
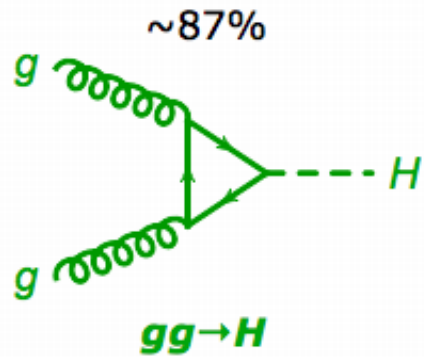
# Higgs production

Vector boson fusion  
VBF

Top-antitop fusion  
ttH

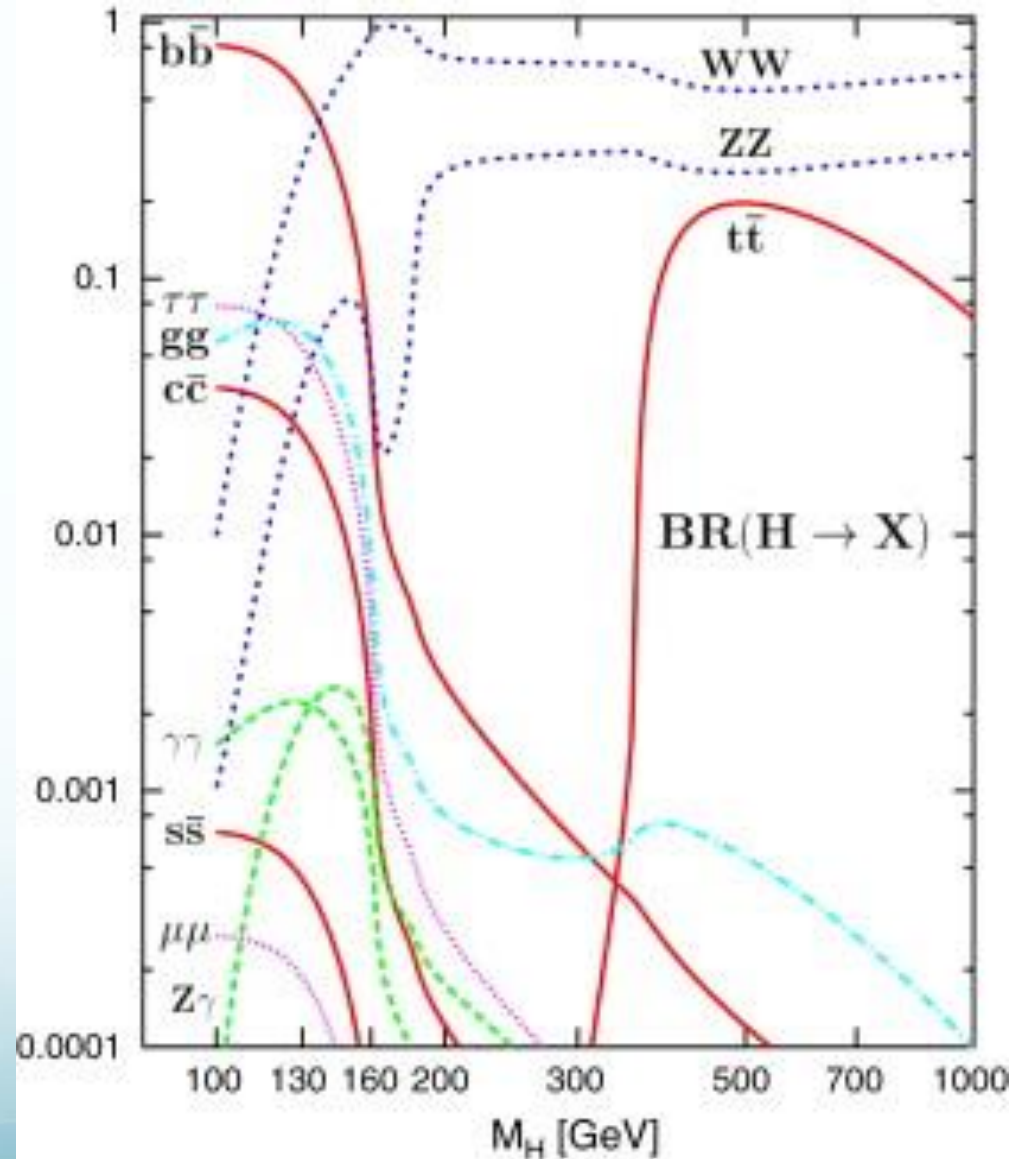
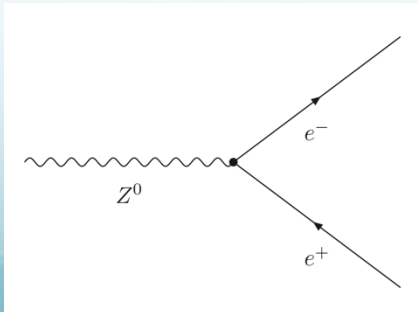
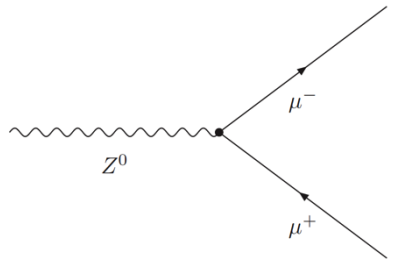
Gluon-gluon fusion  
 $gg \rightarrow H$

Higgs strahlung  
VH



# Detect Higgs by decay products

- Variety of decay channels
- Massive particles more likely
- Difficult to detect from background
- Life time is  $1.56 \times 10^{-22}$  s (!)  
(predicted in the Standard Model)
- $\gamma\gamma$  is clean, but rare



# Media and Press Relations



News

Media visits

Events

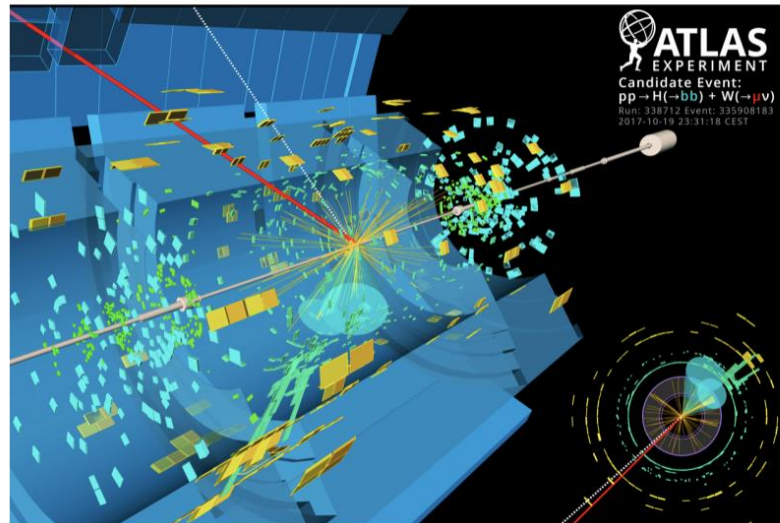
Photos & Videos

Other resources

Contact us

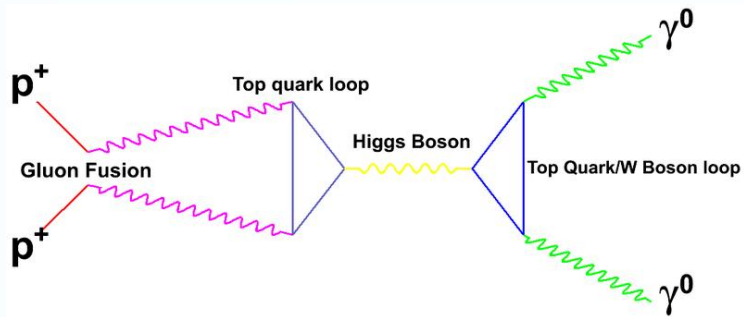
## Long-sought decay of Higgs boson observed

28 Aug 2018

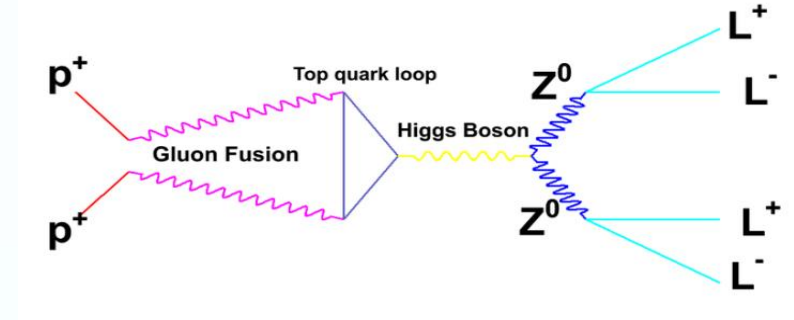


*An ATLAS candidate event for the Higgs boson (H) decaying to two bottom quarks (b), in association with a W boson decaying to a muon ( $\mu$ ) and a neutrino ( $\nu$ ). Image : ATLAS/CERN.*

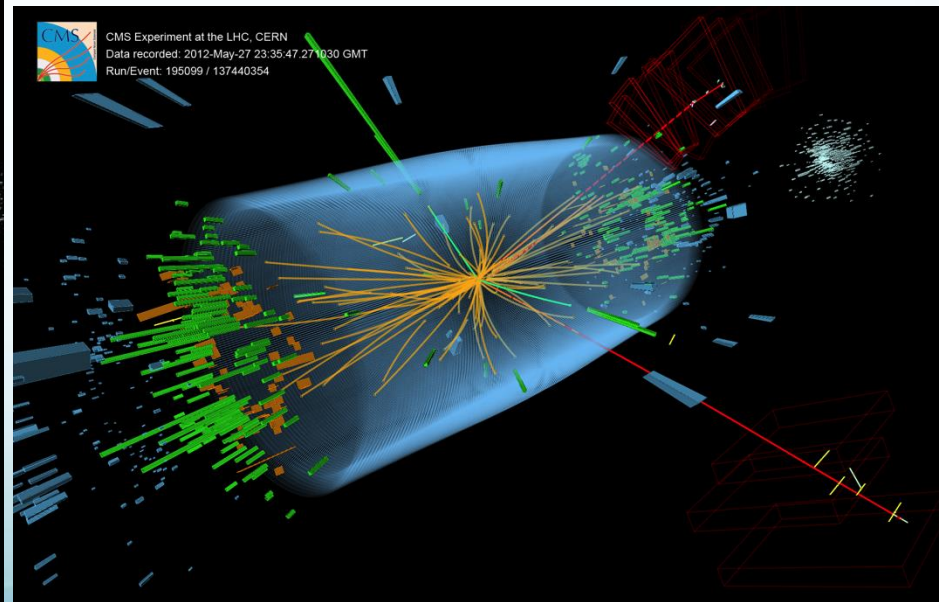
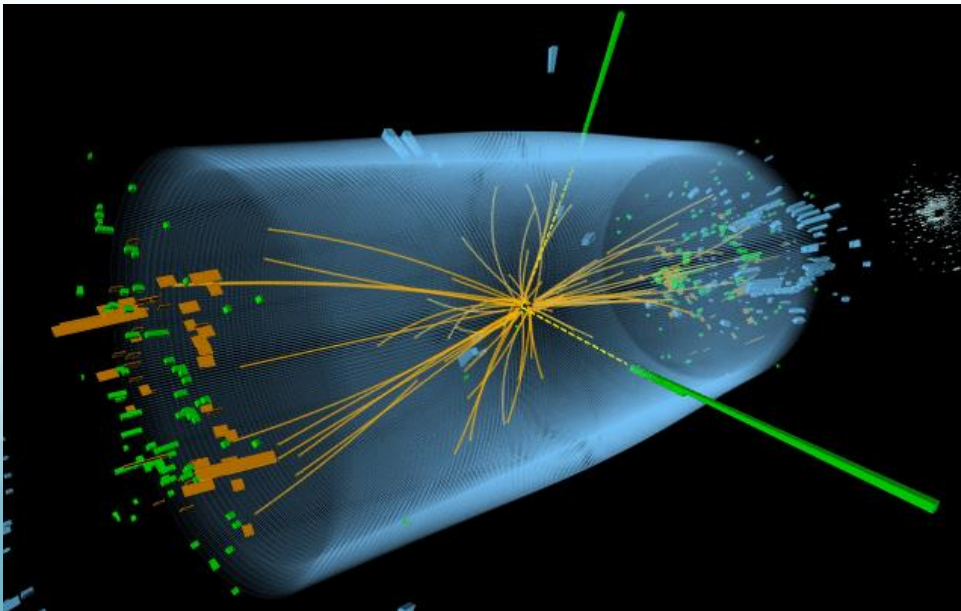
# April-July 2012: $8\text{ TeV}$ , $5.8\text{ fb}^{-1}$



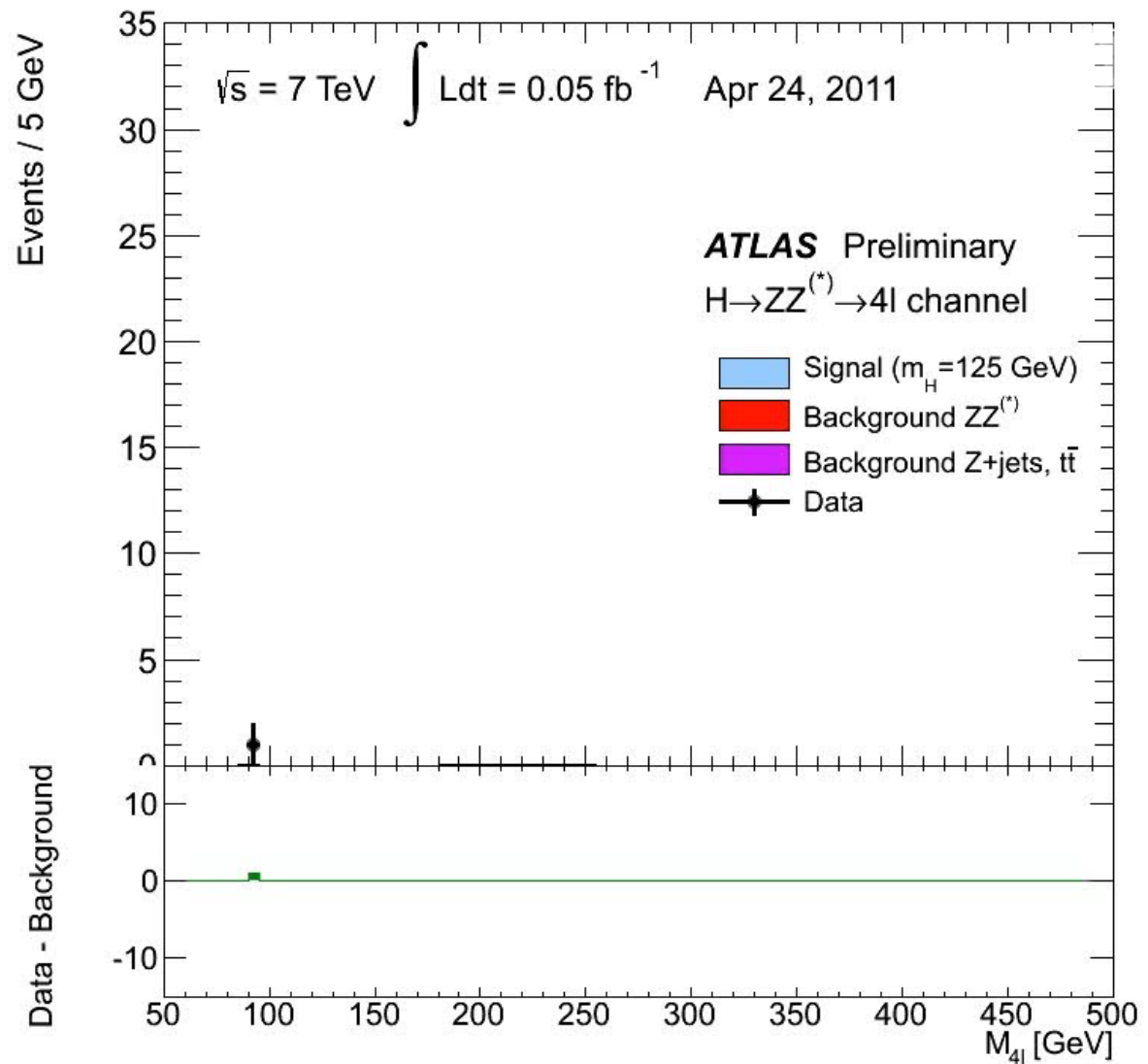
*Measure energy of photons emitted*



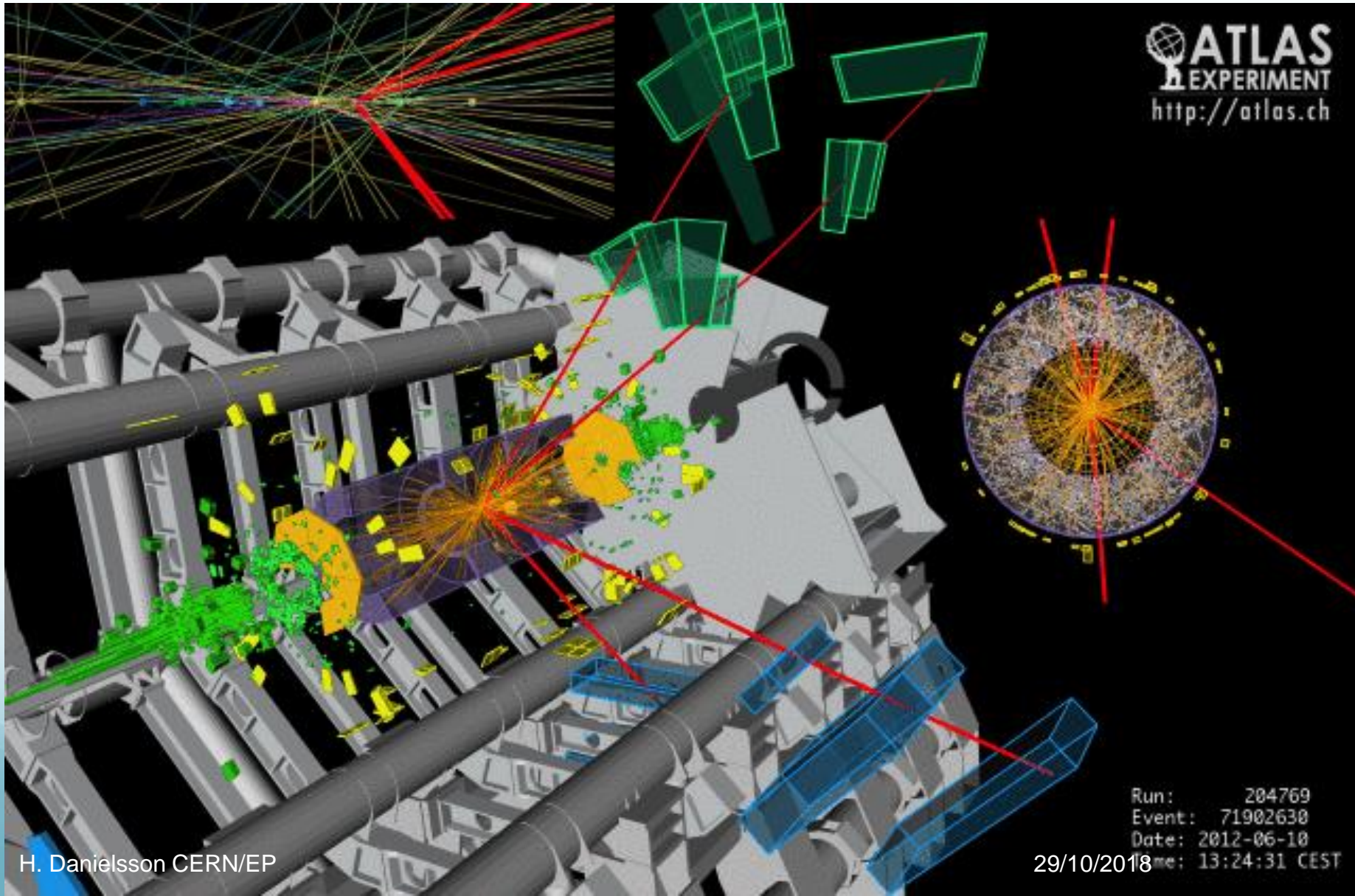
*Measure decay products of Z bosons*



# H → 4l

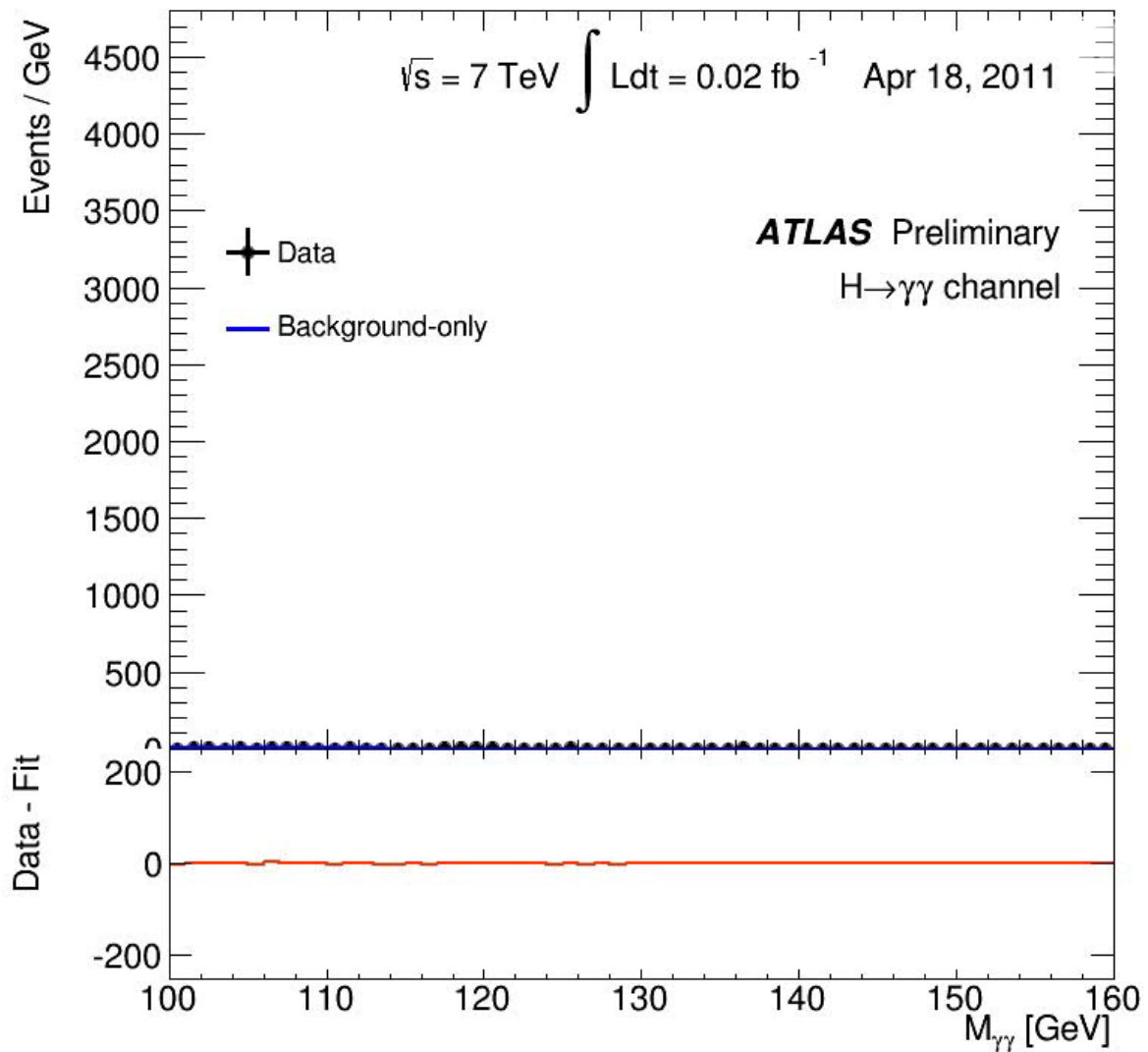


# Higgs events $H \rightarrow 4l$ (muons)

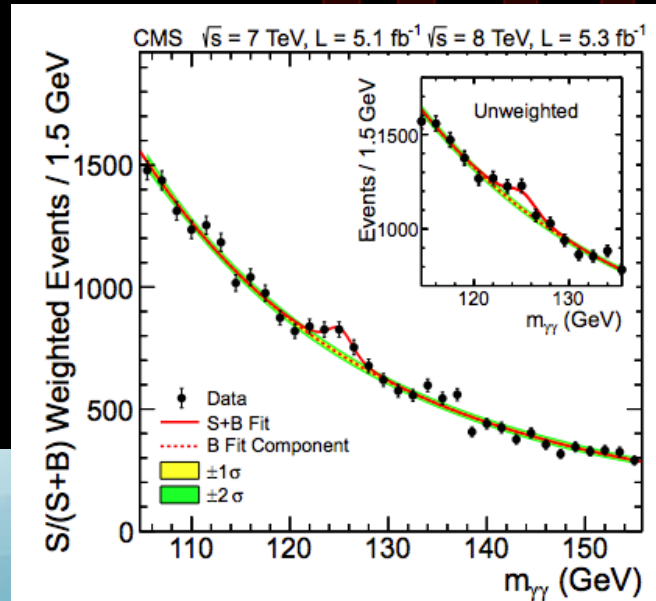
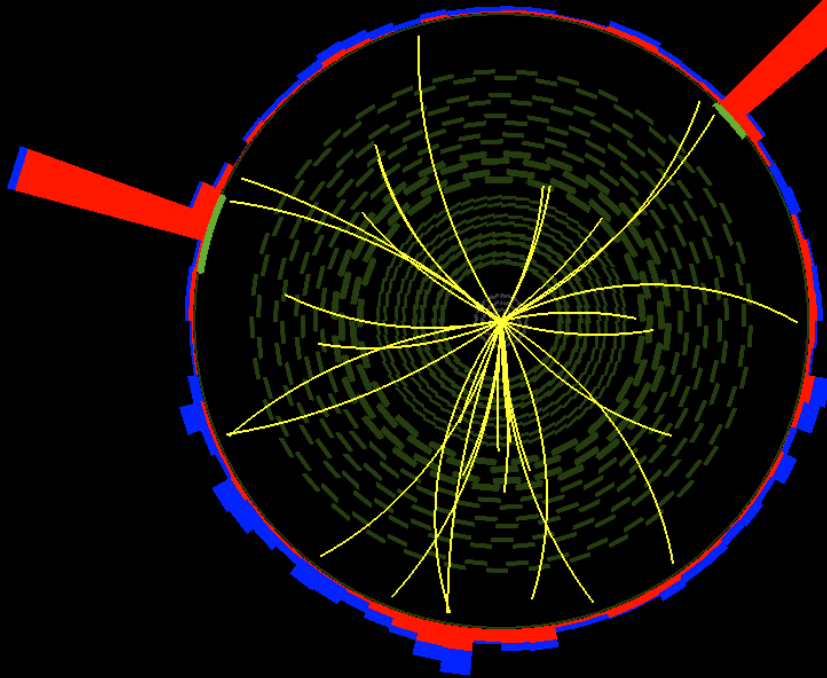




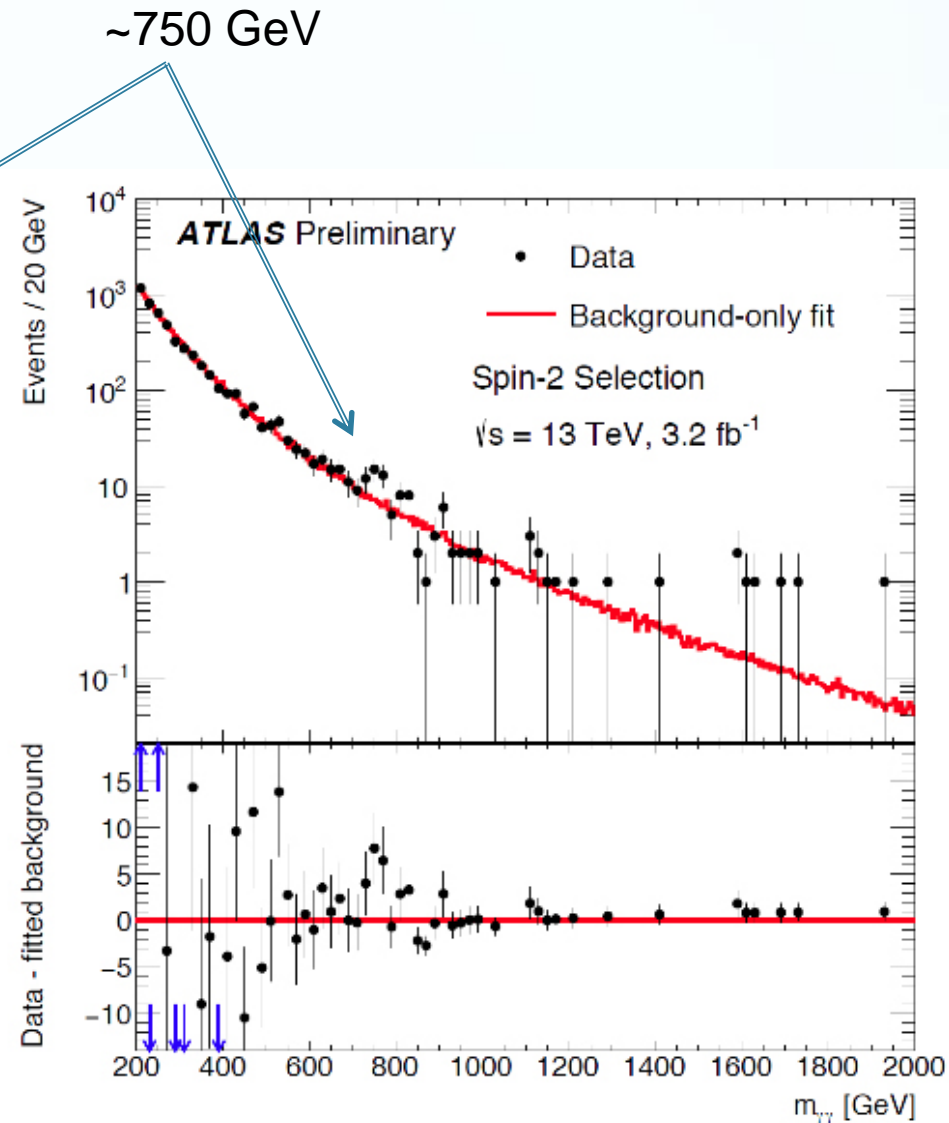
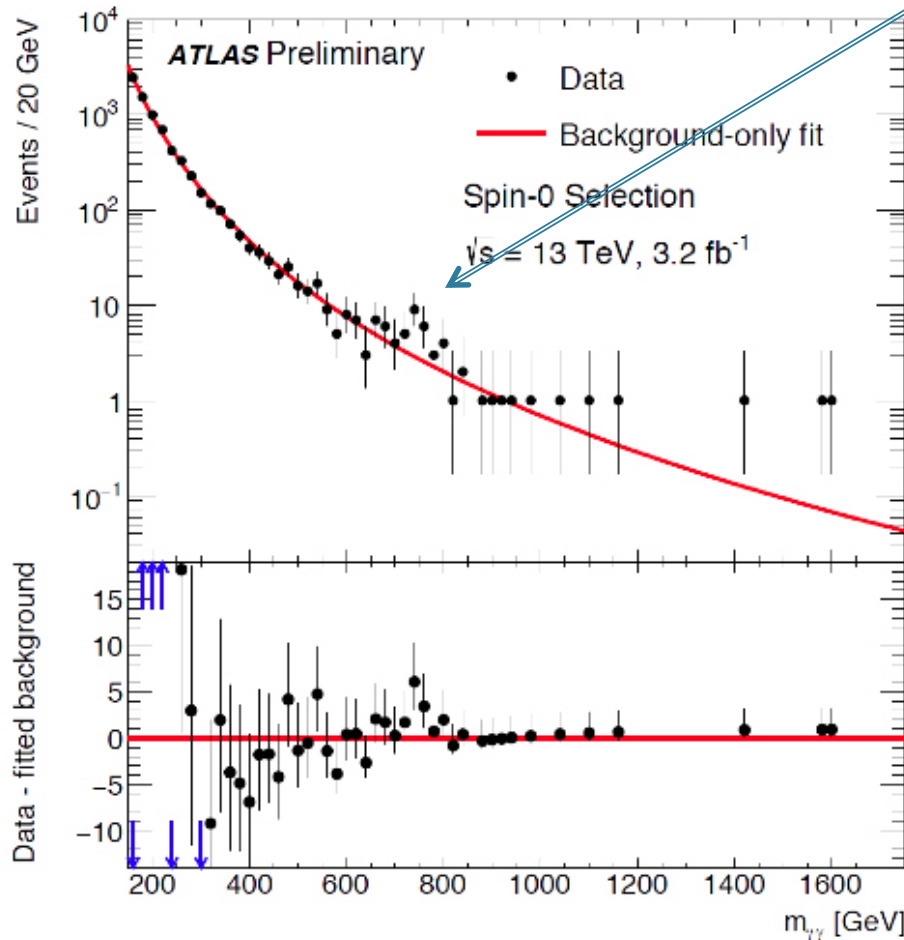
$$H \rightarrow \gamma\gamma$$



# From CMS Higgs $\rightarrow \gamma\gamma$



# Searching beyond the Standard Model with photon pairs



<https://atlas.cern/updates/physics-briefing/searching-beyond-standard-model-photon-pairs>

# Future (after LHC): FCC?

Higgs + Nothing Else@LHC?

A Fine-tuning of  
~ 1% for weak scale

Nima Arkani-Hamed

Reference:

<http://indico.cern.ch/event/282344/contributions/1630763/attachments/519399/716598/FCCtalk.pdf>

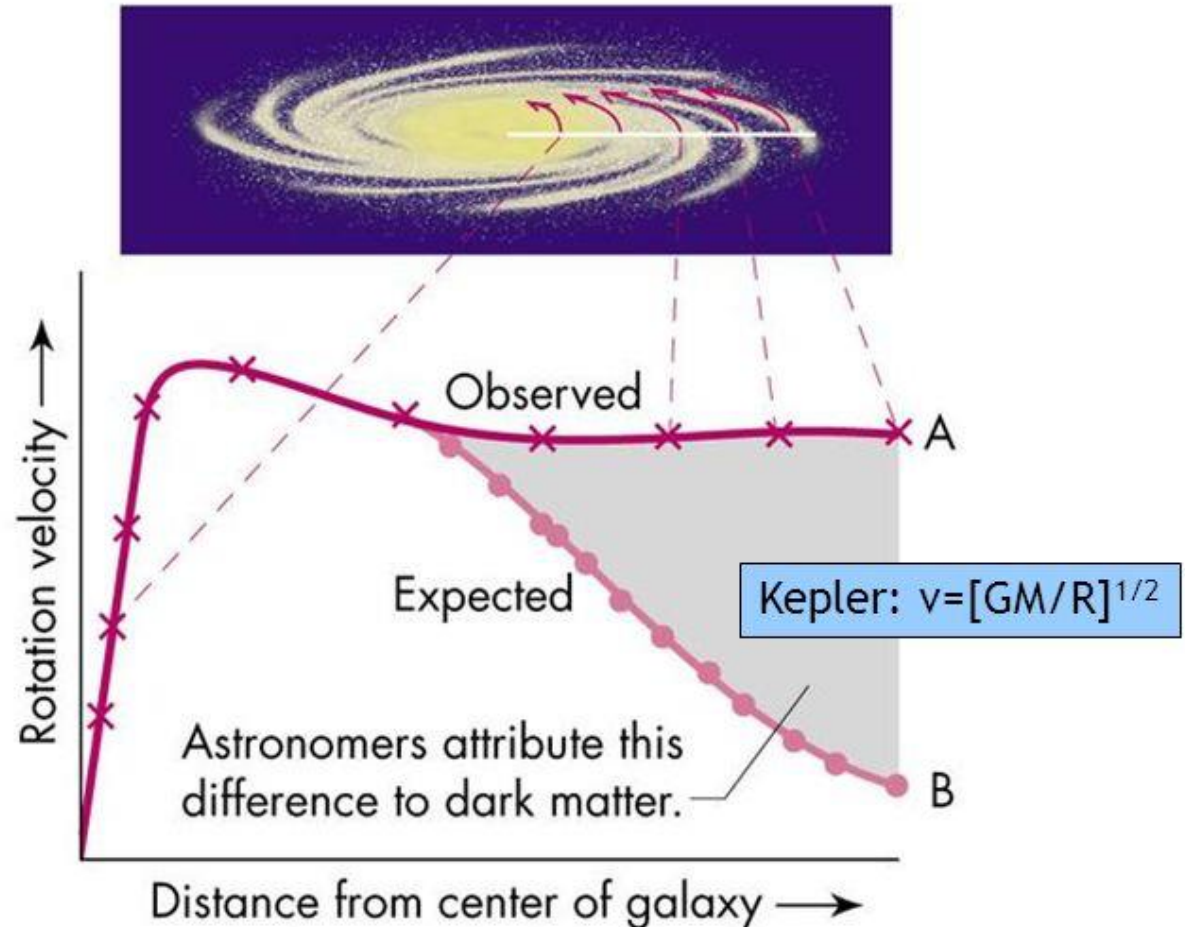
# Dark Matter? Dark Energy?



- **Dark Matter** is invisible matter, it does not emit light. Its evidence comes from the study of the motion of galaxies and groups of galaxies
- **Dark Energy** is the term introduced to justify the acceleration of the Universe expansion (is it equivalent to Einstein's cosmological constant)

# Potential Wells are much deeper than can be explained with visible matter

We have measured this for many years on galactic scales



But Where Is Everybody?



Nima Arkani-Hamed

# Modified Newtonian Dynamics (MOND) as an alternative to dark matter !

## Who is right ?

>A new theory of gravity

>Experiments:

- In space or on the ground
- Accelerators (CERN)

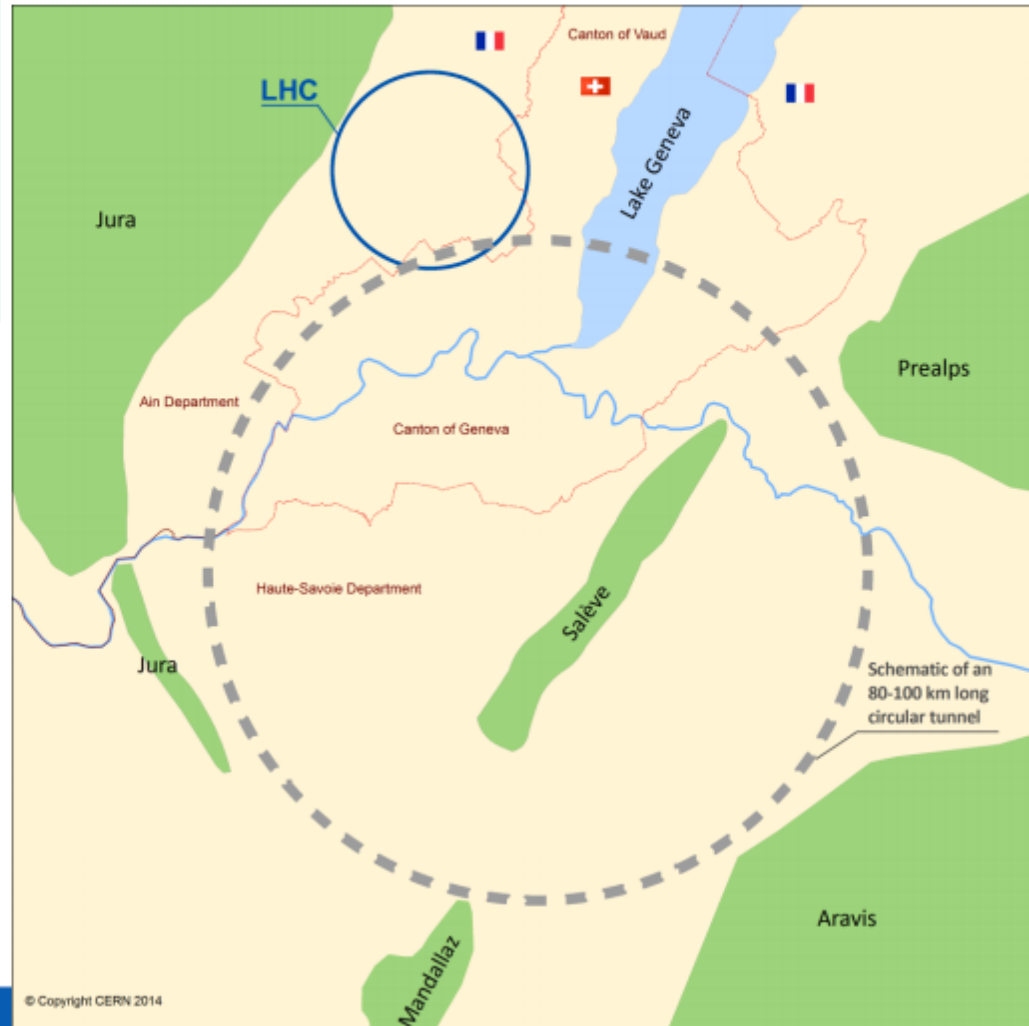




# 80-100 km tunnel infrastructure in Geneva area – design driven by pp-collider requirements (FCC-hh) with possibility of e<sup>+</sup>-e<sup>-</sup> (FCC-ee) and p-e (FCC-he)

## FCC (Future Circular Colliders) CDR and cost review for the next ESU (2018) (including injectors)

16 T ⇒ 100 TeV in 100 km  
20 T ⇒ 100 TeV in 80 km



© Copyright CERN 2014

# Literature

- CERN Academic Training  
<http://indico.cern.ch/conferenceDisplay.py?confId=266737>
- CERN ATLAS  
<http://www.atlas.ch/HiggsResources/>
- <https://www.newscientist.com/article/mg22229670-400-forget-dark-matter-embrace-my-mond-theory-instead/>
- Youtube!

# Spares