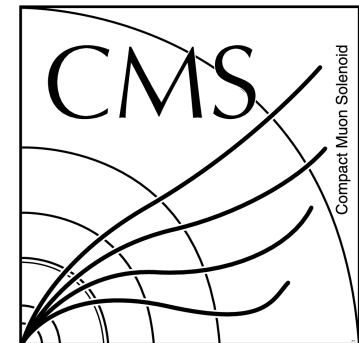


# Triggering DM to LL



Phil Harris (MIT)

input from  
D. Curtin/J. Shelton/  
M. Maccoullough/P. Meade



# Smattering of EXO searches<sup>2</sup>

- $t\bar{t}$ +MET
- $Z(\ell\ell)$ +MET
- monoH( $\gamma\gamma$ )
- monoH(bb)
- monoH( $\tau\tau$ )
- monoH(WW)
- monojet
- mono- $\gamma$
- monotop
- Mono-Leptoquark
- displaced e-mu
- Stopped long lived particle
- Displaced jet
- Disappearing track

- dijets
- low-mass dijets
- Boosted dijet
- B-tagged dijets
- $Z' \rightarrow ee/\mu\mu/W' \rightarrow e/\mu/\tau\nu$
- Black Holes
- LQ( $ee/\tau\tau$ )
- $l^* \rightarrow l\gamma$
- $q^* \rightarrow q\gamma$
- $N \rightarrow lqj$
- Multilepton
- $Z(\ell\ell)\gamma$
- $Z(qq)\gamma$
- $\gamma\gamma$

Color coded  
roughly by types  
of models

# Smattering of EXO searches<sup>3</sup>

- $t\bar{t}$ +MET
- $Z(\ell\ell)$ +MET
- monoH( $\gamma\gamma$ )
- monoH(bb)
- monoH( $\tau\tau$ )
- monoH(WW)
- monojet
- mono- $\gamma$
- monotop
- Mono-Leptoquark

- displaced e-mu
- Stopped long lived particle
- Displaced jet
- Disappearing track

- dijets
- low-mass dijets
- Boosted dijet
- B-tagged dijets
- $Z' \rightarrow ee/\mu\mu/W' \rightarrow e/\mu/\tau\nu$

- Black Holes
- LQ( $ee/\tau\tau$ )
- $l^* \rightarrow l\gamma$
- $q^* \rightarrow q\gamma$
- $N \rightarrow lqj$
- Multilepton

- $Z(\ell\ell)\gamma$
- $Z(qq)\gamma$
- $\gamma\gamma$

# Smattering of EXO searches<sup>4</sup>

- $t\bar{t} + \text{MET}$
- $Z(\text{ll}) + \text{MET}$
- $\text{monoH}(\gamma\gamma)$
- $\text{monoH}(bb)$
- $\text{monoH}(\tau\tau)$
- $\text{monoH}(WW)$
- $\text{monojet}$
- $\text{mono-}\gamma$
- $\text{monotop}$
- $\text{Mono-Leptoquark}$

Dark Matter

- displaced e-mu
- Stopped long lived particle
- Displaced jet
- Disappearing track

Long lived

- dijets
- low-mass dijets
- Boosted dijet
- B-tagged dijets
- $Z' \rightarrow ee/\mu\mu/W' \rightarrow e/\mu/\tau\nu$

Resonance

- Black Holes
- $LQ(ee/\tau\tau)$
- $l^* \rightarrow l\gamma$
- $q^* \rightarrow q\gamma$
- $N \rightarrow lqj$
- Multilepton

Exotic sig.

- $Z(\text{ll})\gamma$
- $Z(qq)\gamma$
- $\gamma\gamma$

ys

# What are the triggers we use?

- Broadly speaking can assign triggers to few cats
- **MET triggers** :  $MET > 150-200$  GeV (L1 limited)
- **Jet triggers** :  $p_T > 400-450$  GeV (L1 to 205 GeV)
- **Photon** :  $p_T > 170$  GeV (L1 to 100 GeV)
- **Lepton triggers** :  $p_T > 24/27$  GeV (L1 20/20)
- **Exotic triggers** : No L1 (comes from above)
  - Displaced object @HLT
  - Out of bunch object

# Trigger Assignment<sup>6</sup>

- ttbar+MET
- Z(II)+MET
- monoH( $\gamma\gamma$ )
- monoH(bb)
- monoH( $\tau\tau$ )
- monoH(WW)
- monojet
- mono- $\gamma$
- monotop
- Mono-Leptoquark

Dark Matter

MET Trigger  
Jet Trigger  
Photon Trigger  
Lepton Trigger  
Exotic Trigger

*MET* and lepton triggers drive dark matter searches  
Likely will be for future

# Smattering of EXO searches<sup>7</sup>

MET Trigger

Jet Trigger

Photon Trigger

Lepton Trigger

Exotic Trigger

The most exotic signatures are not trigger sensitive

...to be honest these will quickly become lower priority

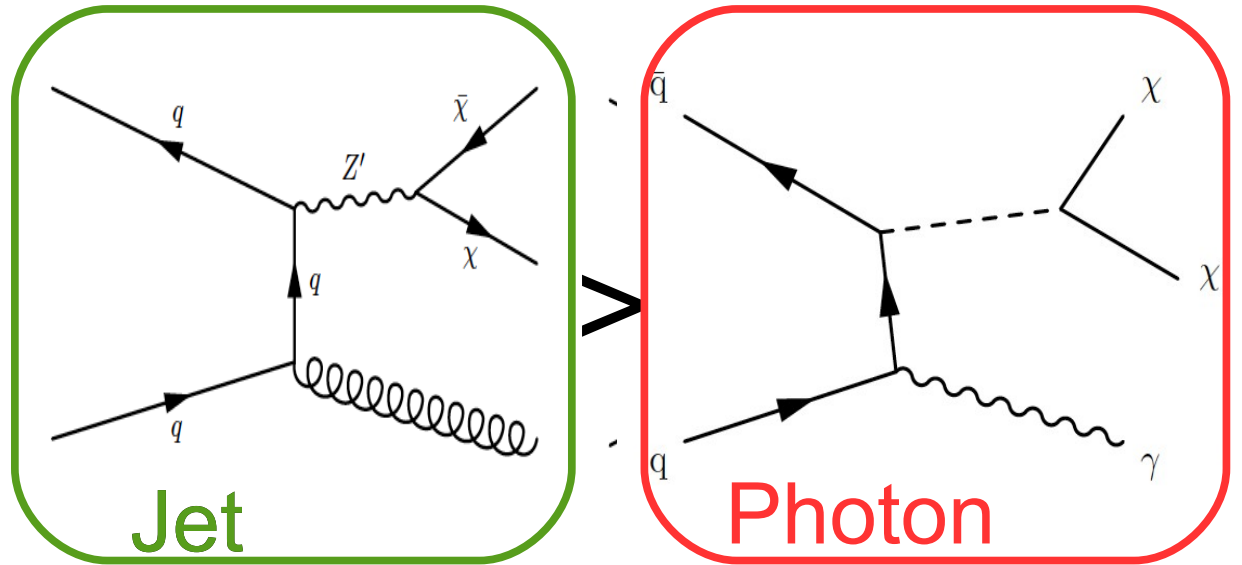
- Black Holes
- LQ( $ee/\tau\tau$ )
- $l^* \rightarrow l\gamma$
- $q^* \rightarrow q\gamma$
- $N \rightarrow lqj$
- Multilepton

Exotic sig.

Lack of  $s^{1/2}$  is a powerful thing

# Smattering of EXO searches<sup>8</sup>

MET Trigger  
Jet Trigger  
Photon Trigger  
Lepton Trigger  
Exotic Trigger



$\gamma$  related searches have been an exciting time....  
but to be honest with exception of  $\gamma\gamma$  for  
Higgs/Resonance they tend to be the 2<sup>nd</sup> string  
when jets/MET fail

- $Z(\ell\ell)\gamma$
- $Z(qq)\gamma$
- $\gamma\gamma$

YS



# Smattering of EXO searches<sup>9</sup>

MET Trigger

Jet Trigger

Photon Trigger

Lepton Trigger

Exotic Trigger

- dijets
- low-mass dijets
- Boosted dijet
- B-tagged dijets
- $Z' \rightarrow ee/\mu\mu/W' \rightarrow e/\mu/\tau\nu$

Resonance

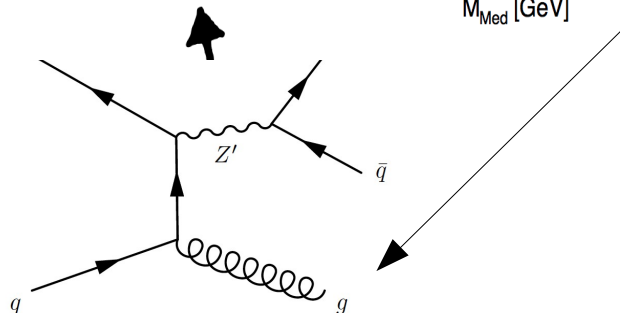
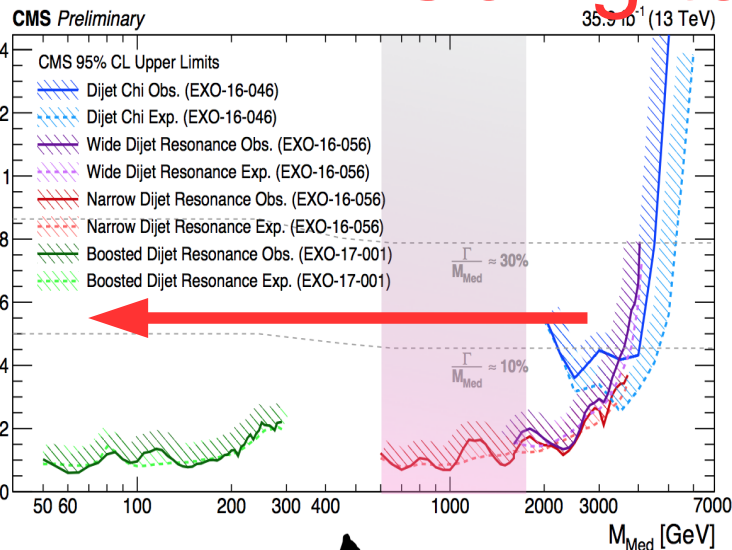
Going to low mass

Data Scouting

Resonances are simple objects  
basically 2 4-vectors  
Store reduced dataset down to L1  
(Use HLT objects with L1 rate)

ISR Tagging

Can use an additional ISR Jet/ $\gamma$  to  
push events over the trigger  
threshold



# Smattering of EXO searches<sup>10</sup>

MET Trigger

Jet Trigger

Photon Trigger

Lepton Trigger

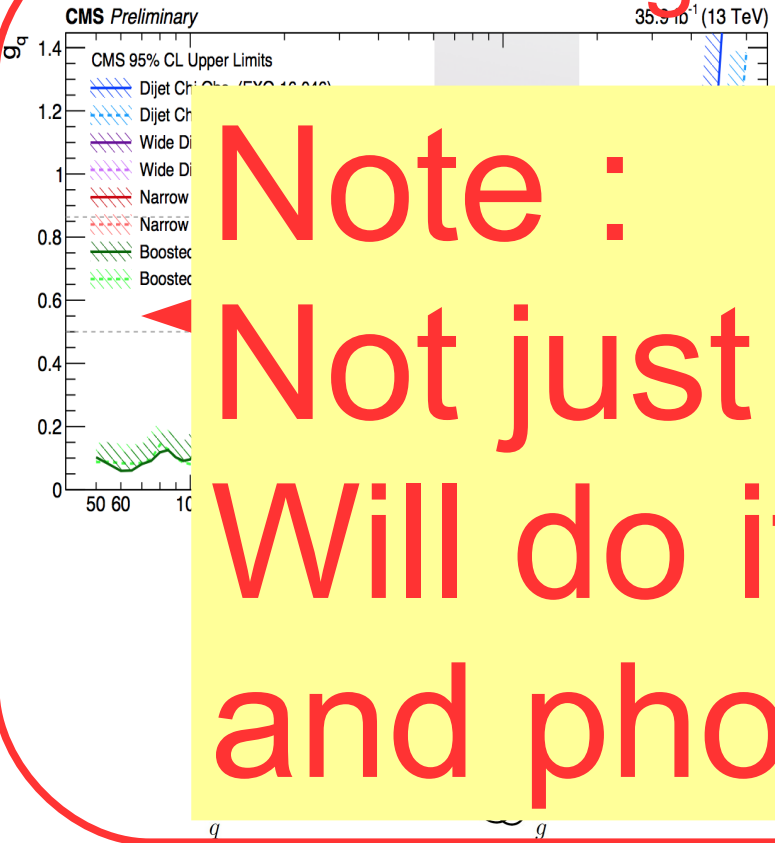
Exotic Trigger

- dijets
- low-mass dijets
- Boosted dijet
- B-tagged dijets
- $Z' \rightarrow ee/\mu\mu/W' \rightarrow e/\mu/\tau\nu$

Resonance

Going to low mass

Data Scouting



**Note :**  
 Not just about jets  
 Will do it for leptons  
 and photons? ....

jects  
 /n to L1  
 (rate)  
 Jet/γ to  
 er

# Smattering of EXO searches<sup>11</sup>

*MET* triggers can drive a lot in upgrade

Hold this thought....

MET Trigger

Jet Trigger

Photon Trigger

Lepton Trigger

Exotic Trigger

Maximizing displaced signatures is something we are just starting to think about

- displaced e-mu
- Stopped long lived particle
- Displaced jet
- Disappearing track

Long lived

- $t\bar{t}$ +MET
- $Z(\ell\ell)$ +MET
- monoH( $\gamma\gamma$ )
- monoH(bb)
- monoH( $\tau\tau$ )
- monoH(WW)
- monojet
- mono- $\gamma$
- monotop
- Mono-Leptoquark

Dark Matter

- displaced e-mu
- Stopped long lived particle
- Displaced jet
- Disappearing track

Long lived

- dijets
- low-mass dijets
- Boosted dijet
- B-tagged dijets
- $Z' \rightarrow ee/\mu\mu/W' \rightarrow e/\mu/\tau\nu$

Resonance

- Black Holes
- LQ( $ee/\tau\tau$ )
- $l^* \rightarrow l\gamma$
- $q^* \rightarrow q\gamma$
- $N \rightarrow lqj$
- Multilepton

Exotic sig.

- $Z(\ell\ell)\gamma$
- $Z(qq)\gamma$
- $\gamma\gamma$

ys

**Tally**

5 MET

6 Jet

5 Photon

11 Lepton

1 Exotic

- ttbar+MET
- Z(H)+MET
- monojet
- monolepton
- monophoton
- mononeutrino
- mono Z
- mono W
- Monoton

- dijets
- low-mass dijets
- high-mass dijets
- monojet
- monolepton
- monophoton
- mononeutrino
- mono Z
- mono W
- Monoton

Excluding exotic sig and

**Worst accept.**

5 Jet

5 MET

4 Photon

7 Lepton

**Best accept.**

Close to even split

- displaced
- Stopped
- Displaced
- Disappearing track

- $\Upsilon\Upsilon$

Resonance

Exotic sig.

ys

**Tally**

5 MET

6 Jet

5 Photon

11 Lepton

1 Exotic



# Back to displaced searches<sup>14</sup>

*MET* triggers can drive a lot in upgrade

Hold this thought....

MET Trigger

Jet Trigger

Photon Trigger

Lepton Trigger

Exotic Trigger

Maximizing displaced signatures is something we are just starting to think about

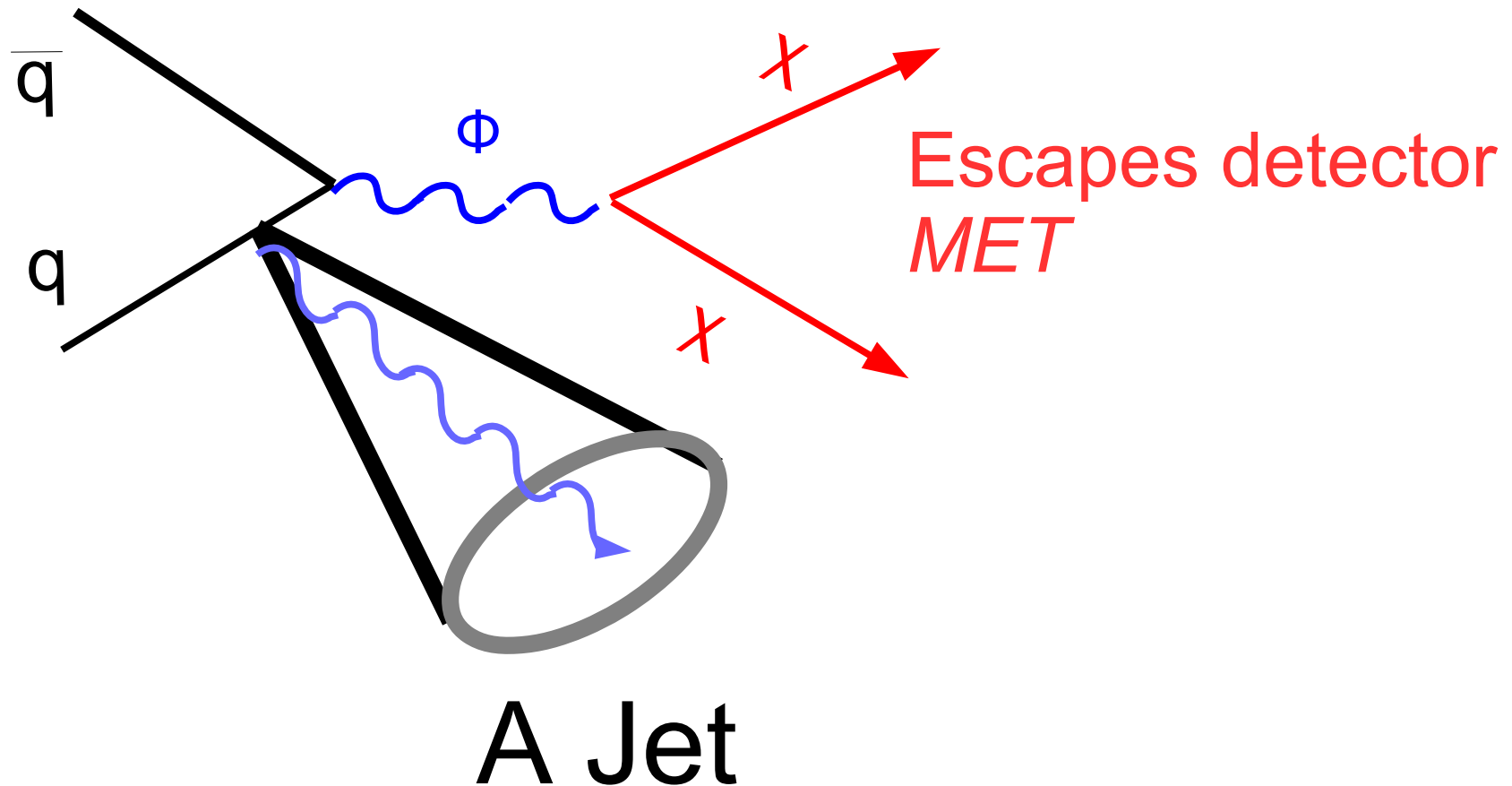
- displaced e-mu
- Stopped long lived particle
- Displaced jet
- Disappearing track

Long lived

# Dark Matter Search

Mono-jet

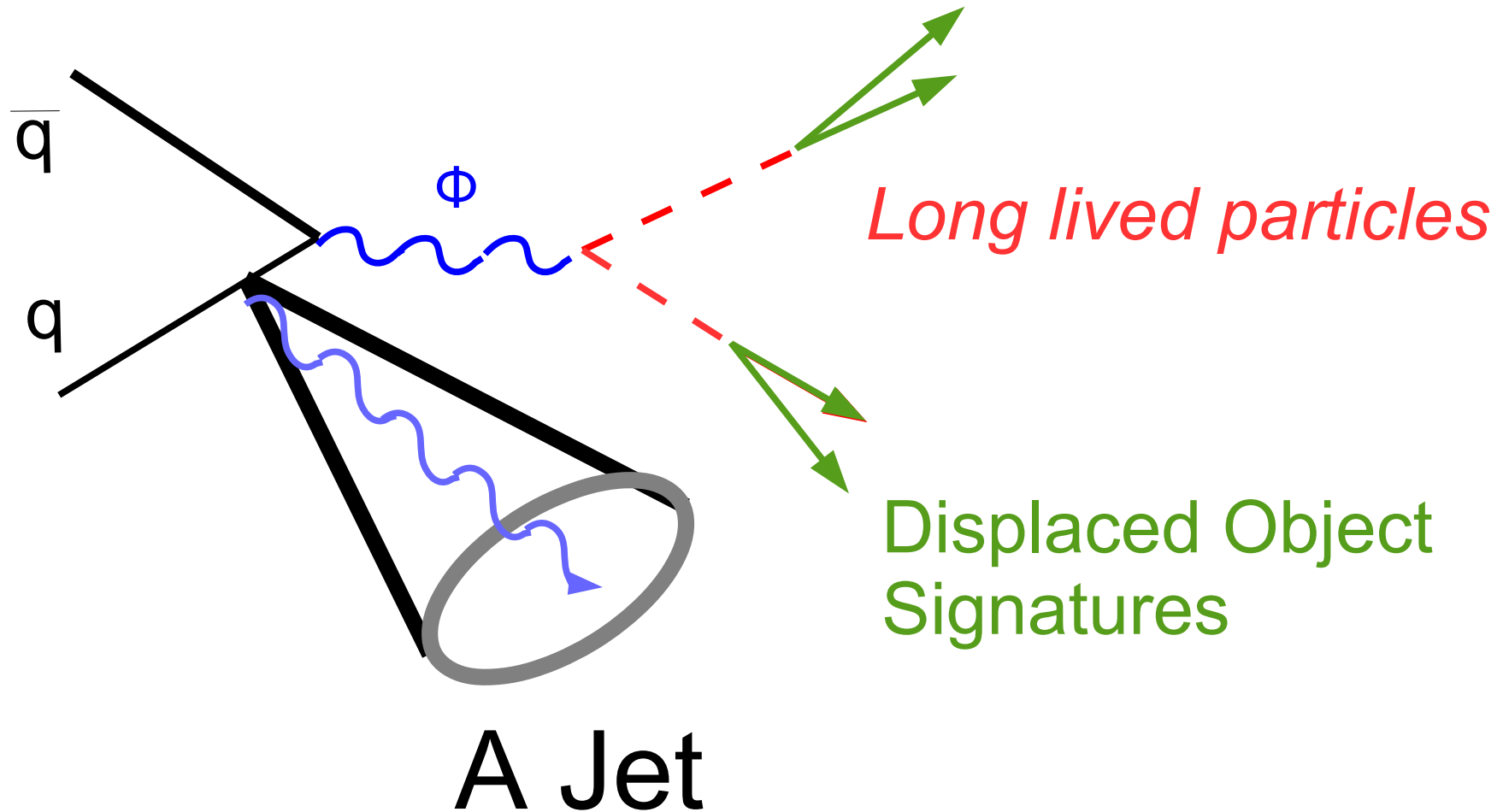
Models : Vector, Axial, Scalar, Pseudoscalar



# Dark Matter Search extended

Mono-jet

Models : Vector, Axial, Scalar, Pseudoscalar

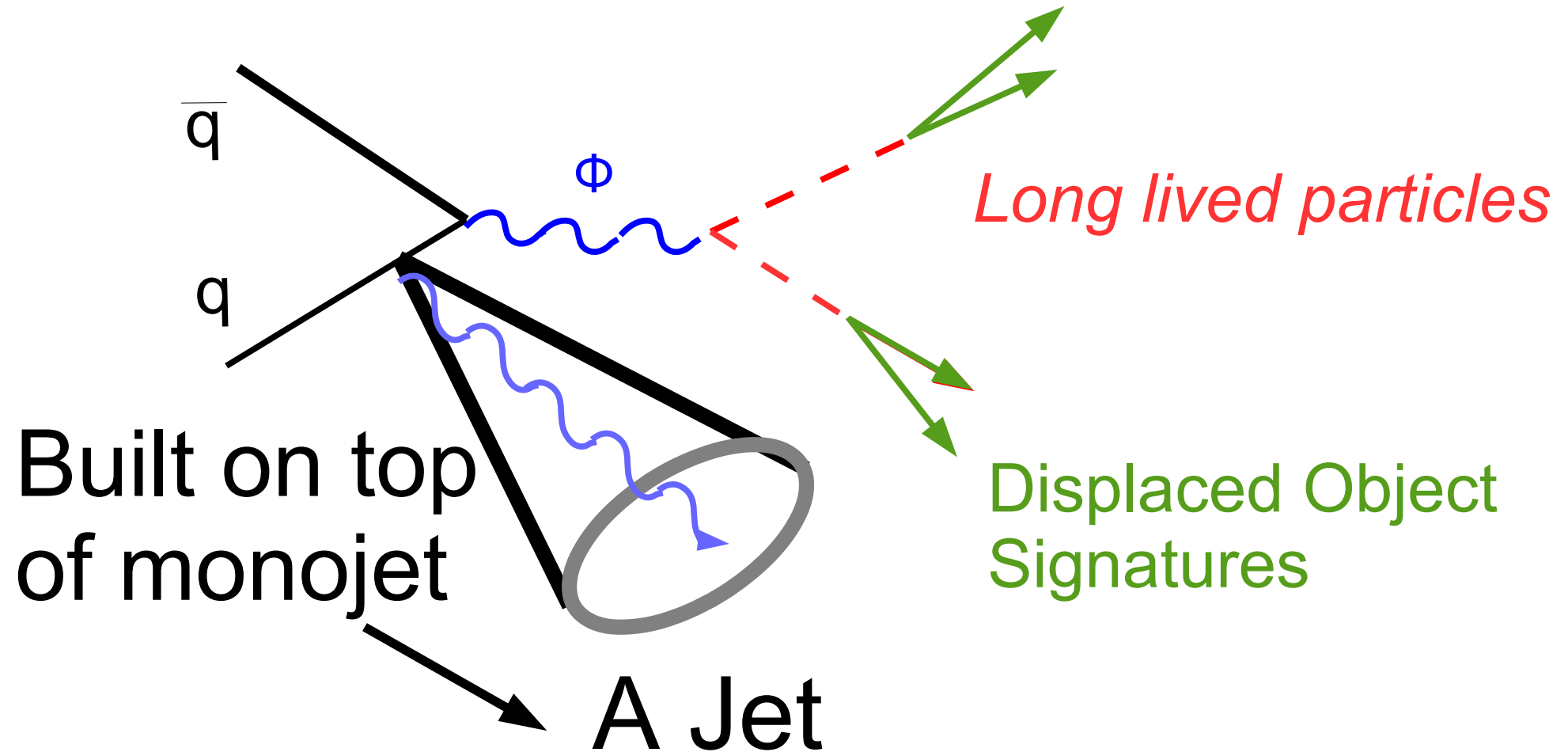




# How Do We Discriminate Models?

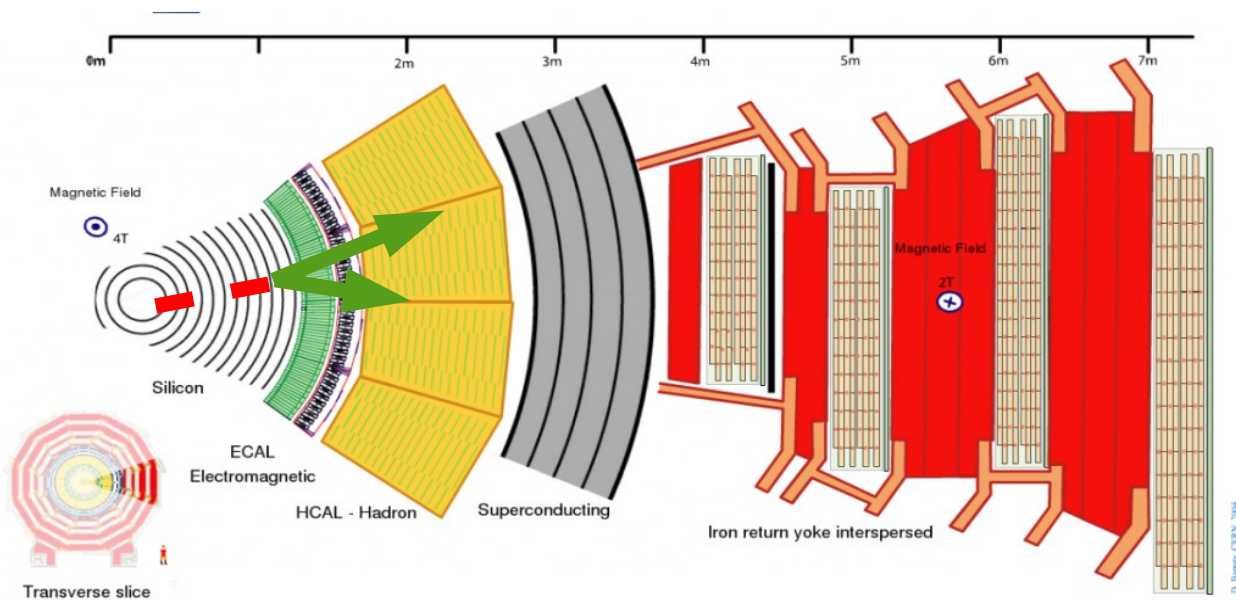
Mono-jet

Models : Vector, Axial, Scalar, Pseudoscalar



# How can we go beyond?

- Tagging a long lived particle



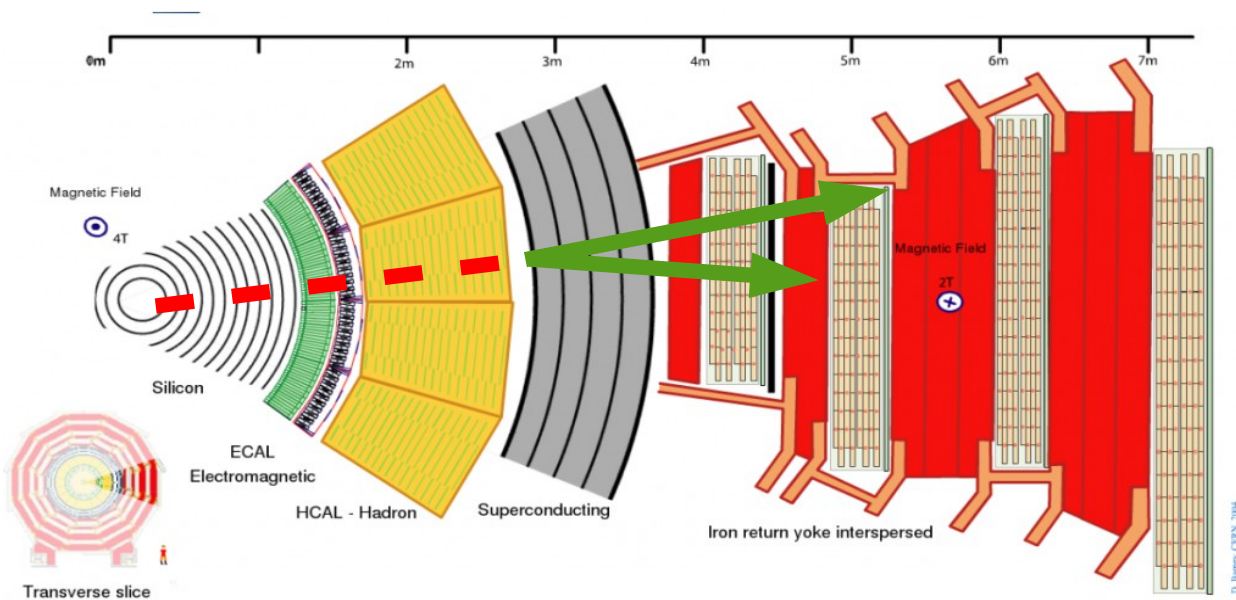
Decay in the tracker :

Displaced track signature

No missing transverse energy

# How can we go beyond?

- Tagging a long lived particle

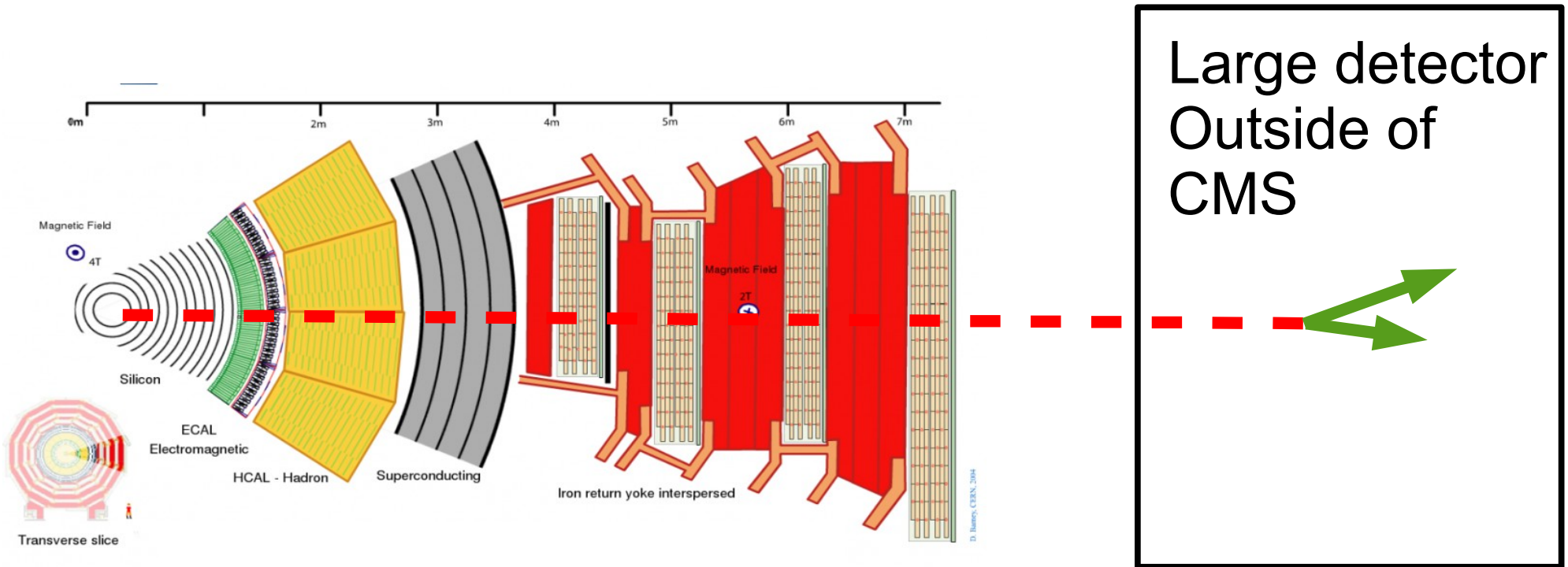


Decay in the Calorimeters : CERN

Calorimeter/Muon signature  
Missing transverse energy

# How can we go beyond?

- Tagging a long lived particle



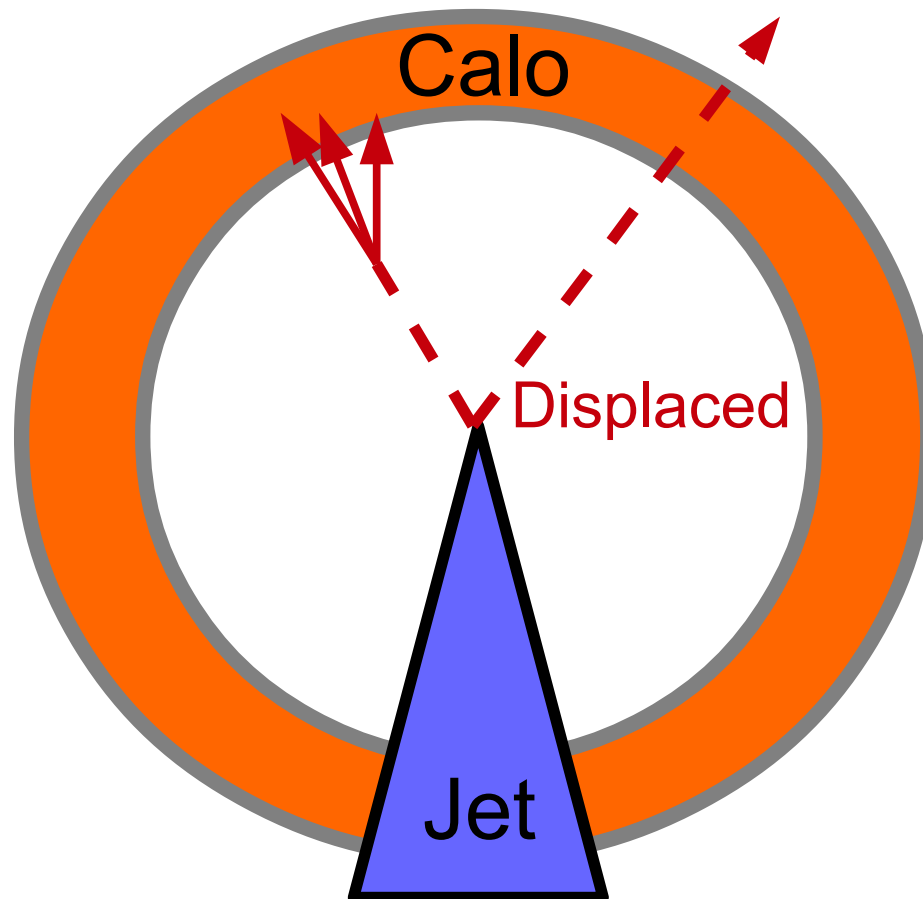
Decay outside :

Missing transverse energy  $\rightarrow$  monojet signature

Few proposals are out like **MATHUSLA**

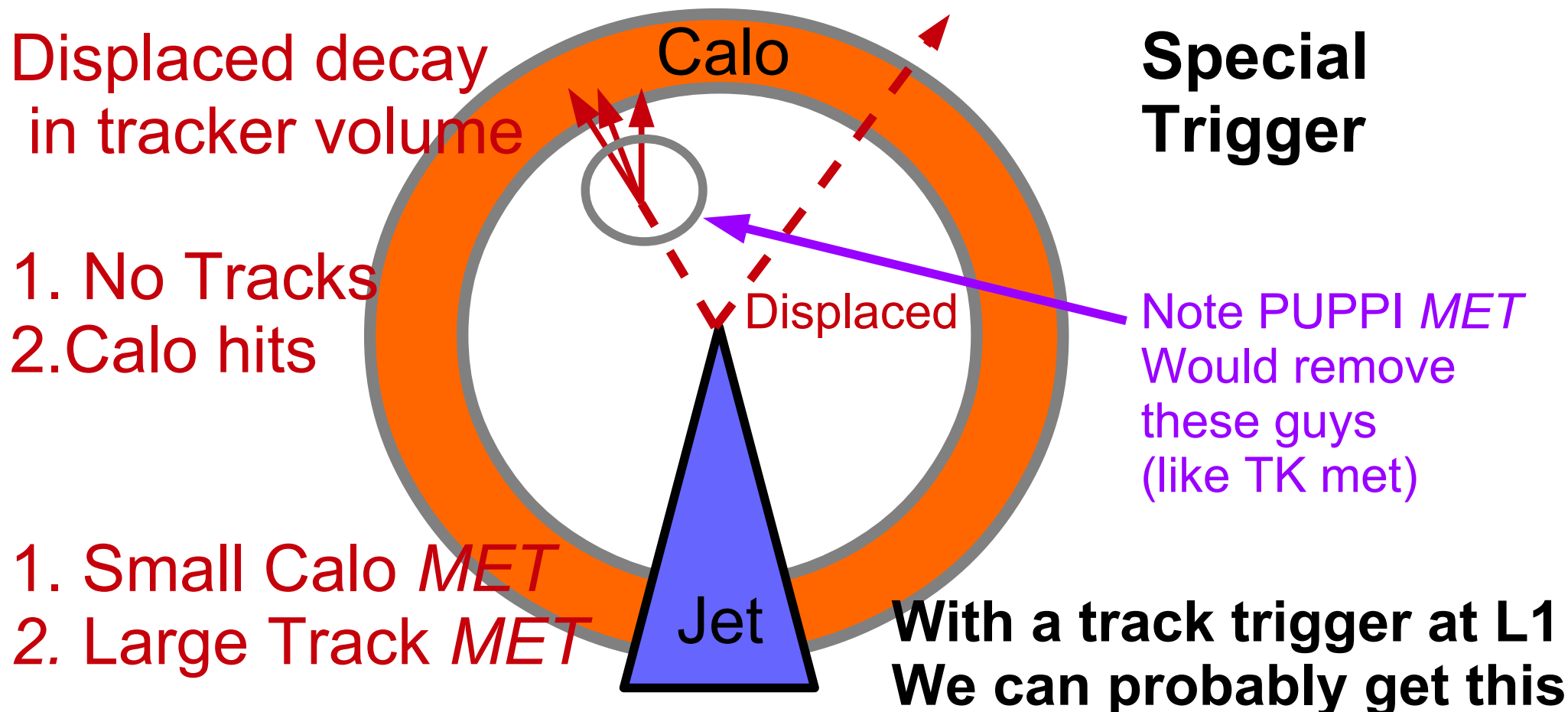
# What about the trigger?

- Displaced tag can reduce the background by  $10^{-3}$
- Displaced objects
  - Depend on whether we can reconstruct them



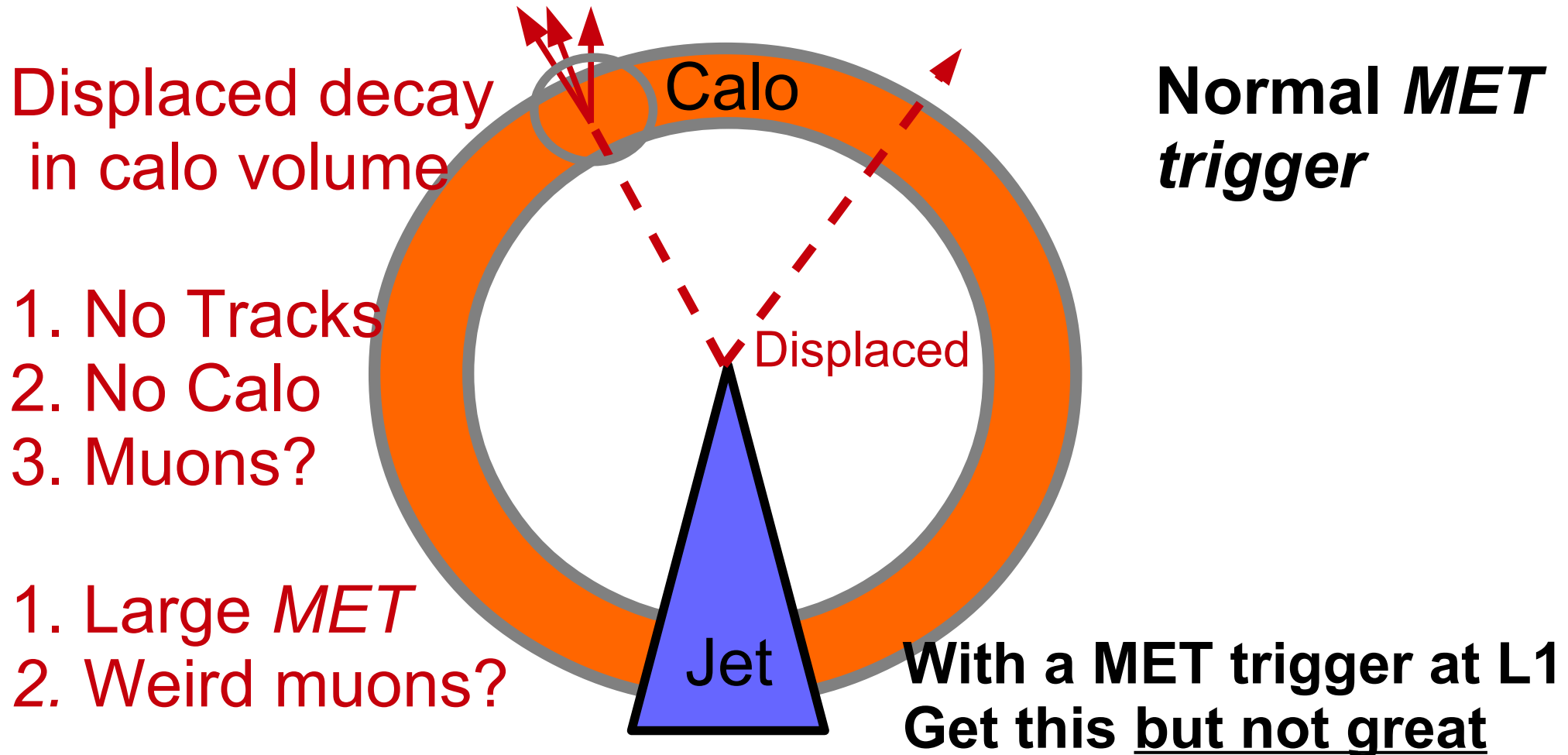
# What about the trigger?

- Displaced tag can reduce the background by  $10^{-3}$
- Displaced objects
  - Depend on whether we can reconstruct them



# What about the trigger?

- Displaced tag can reduce the background by  $10^{-3}$
- Displaced objects
  - Depend on whether we can reconstruct them



1. No Tracks
2. No Calo
3. Muons?

1. Large *MET*
2. Weird muons?

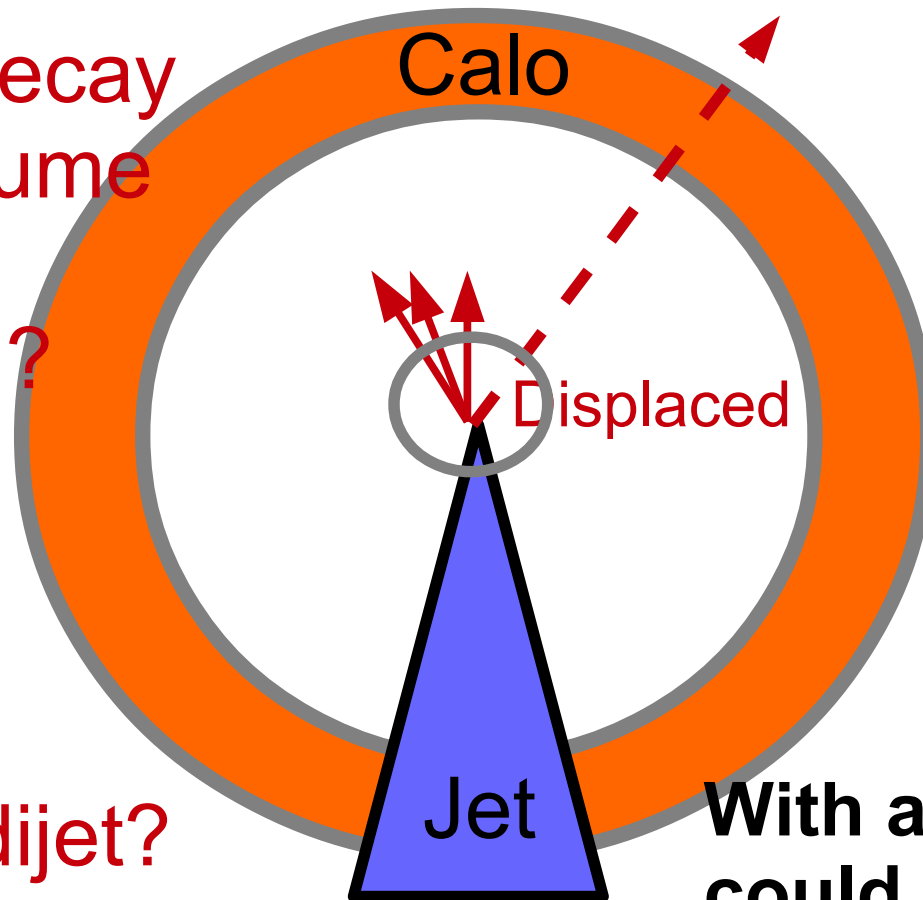
# What about the trigger?

- Displaced tag can reduce the background by  $10^{-3}$
- Displaced objects
  - Depend on whether we can reconstruct them

Displaced decay  
In Pixel volume

0. displaced?
1. Tracks
2. Calo

1. Strange dijet?



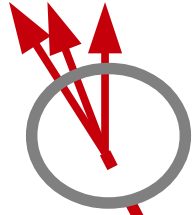
**Displaced at  
L1 does not  
exist now**

**With a displaced L1 we  
could do thi**



# What about the trigger?

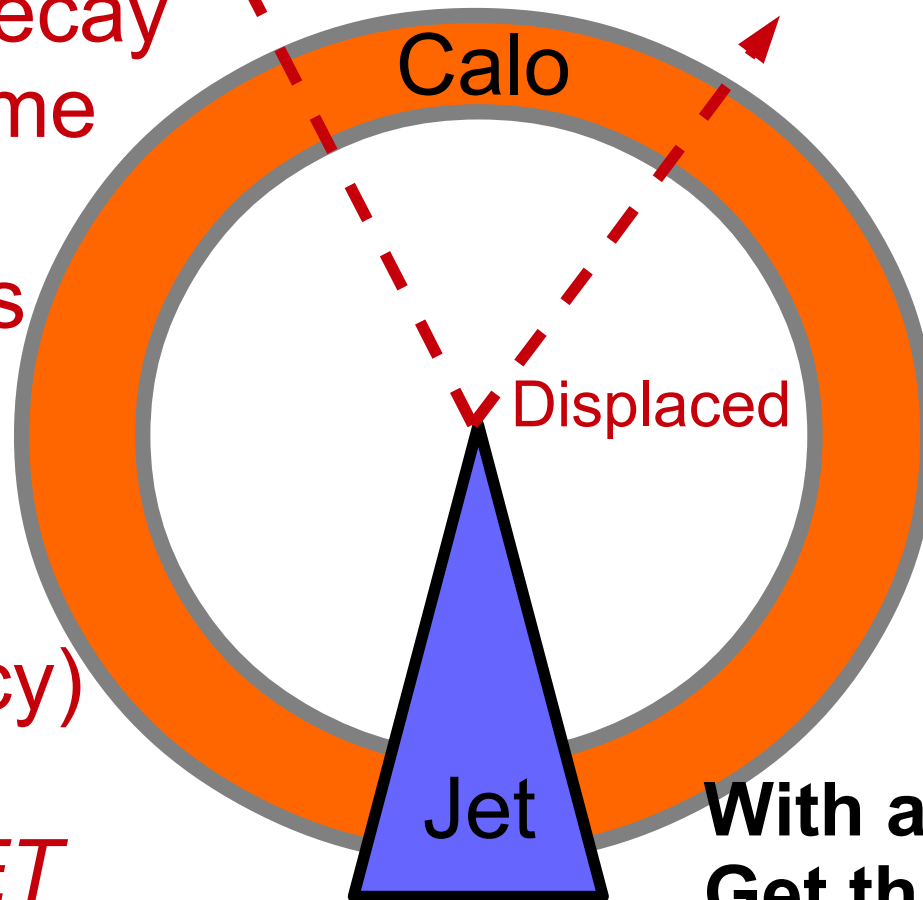
Detector  
 $\mu\text{s}$  away



Displaced decay  
Out of volume

1. No Tracks
2. No Calo
3. External Detector?  
(large latency)

1. Large *MET*



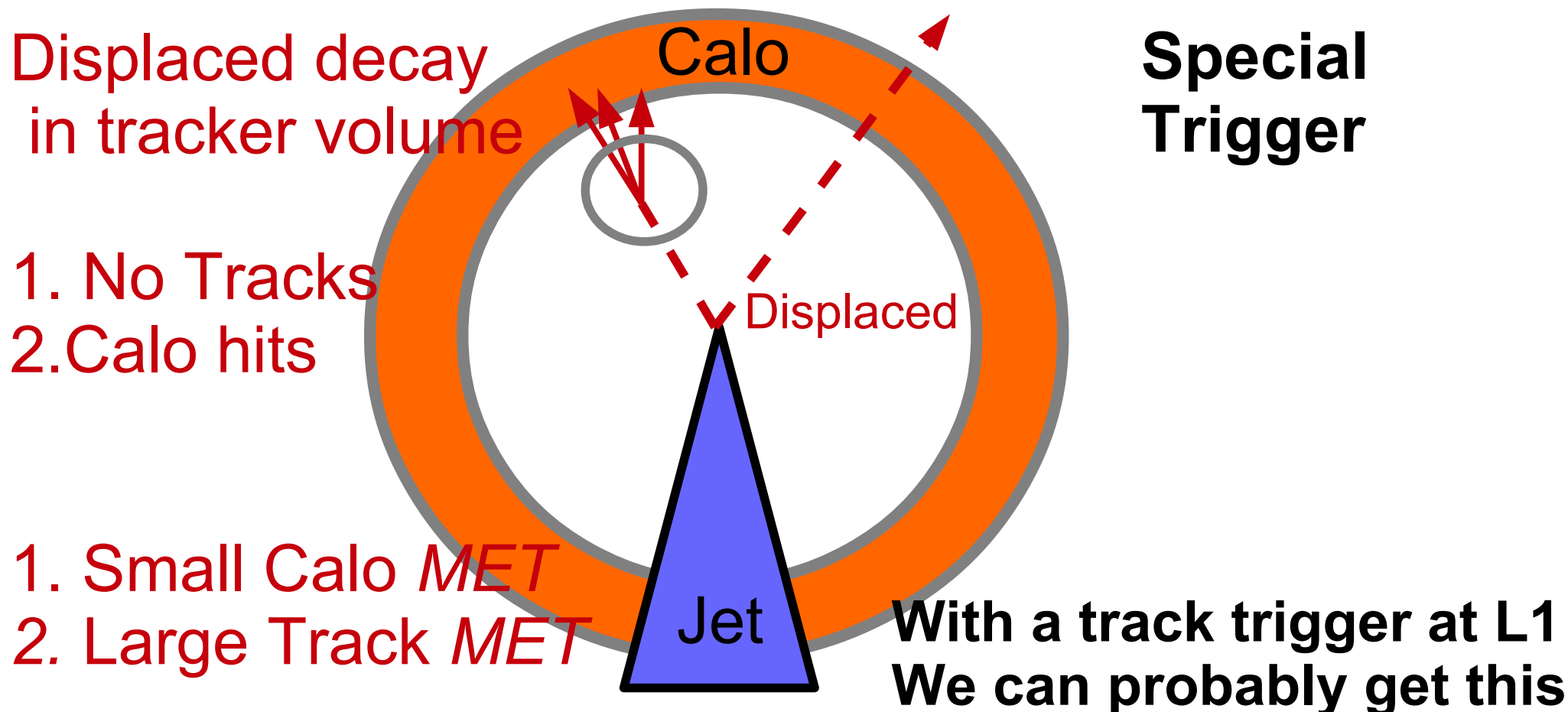
**Normal *MET* trigger**

Contend  
with normal  
L1 rates

**With a MET trigger at L1  
Get this but not great**

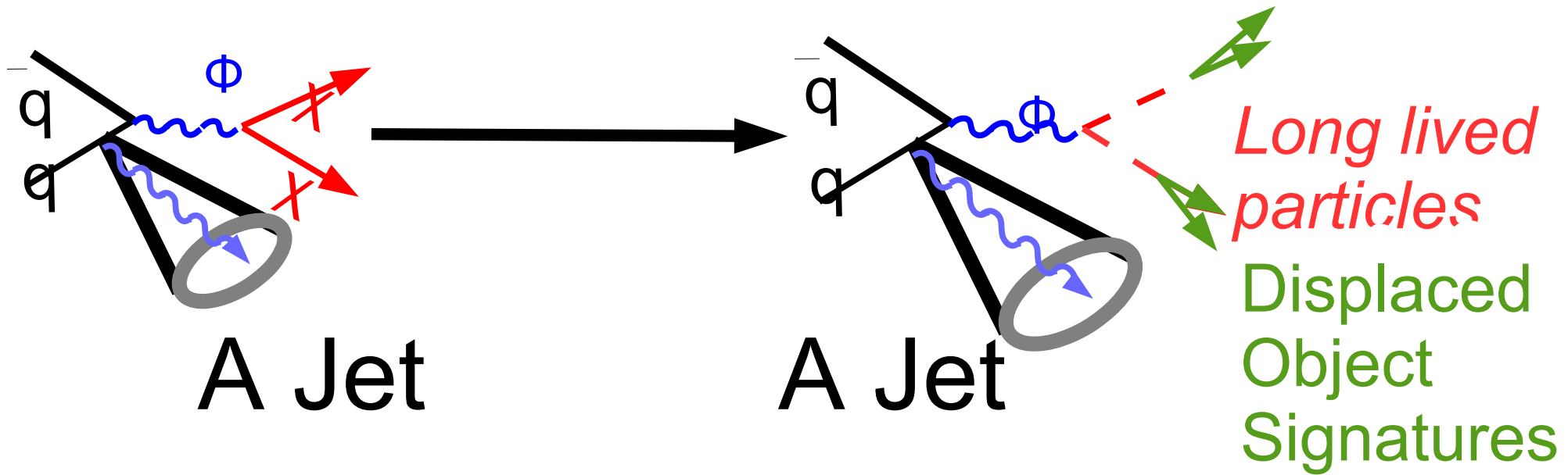
# Current default: *MET* trigger

- Many displaced searches use a *MET* trigger
  - Basically unless there is a displaced lepton
  - **L1 *MET* is high right now** : remains a limiting factor



# Part 2 : Implications of Long Lived Signatures

# Dark Matter Search



From a dark matter perspective:

Transition from a dark matter to an LLP is natural

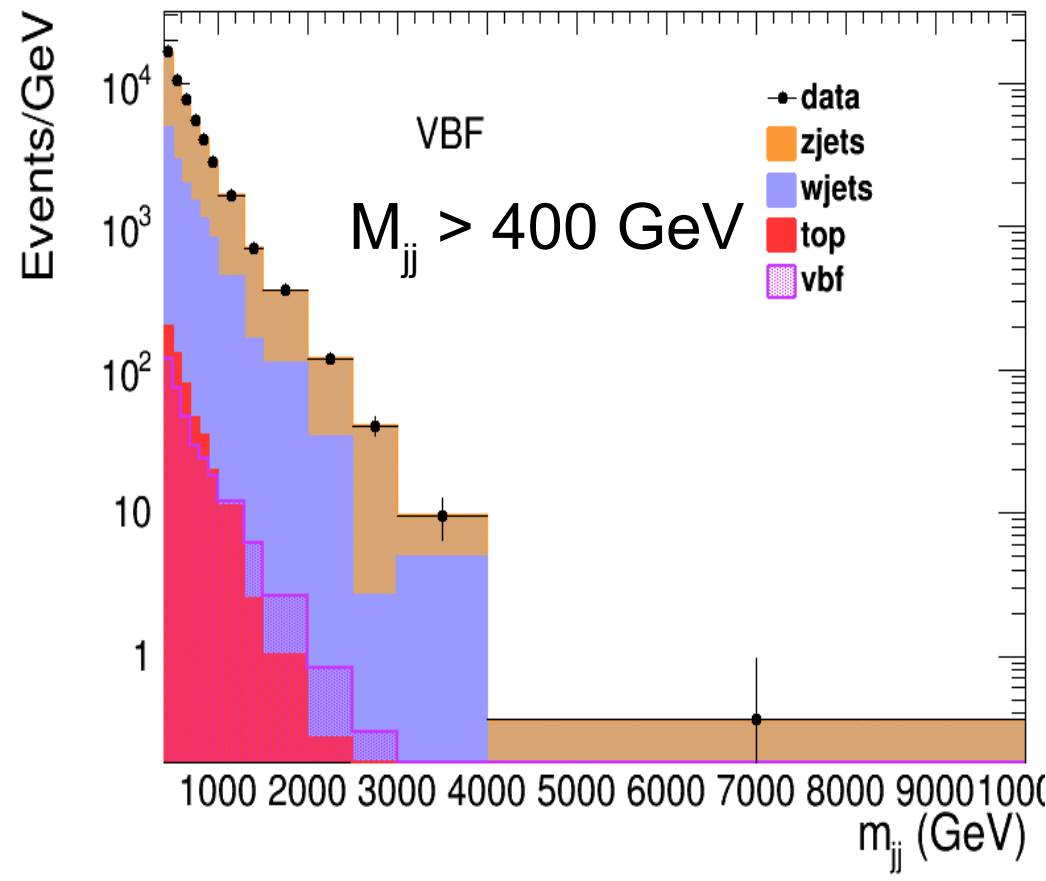
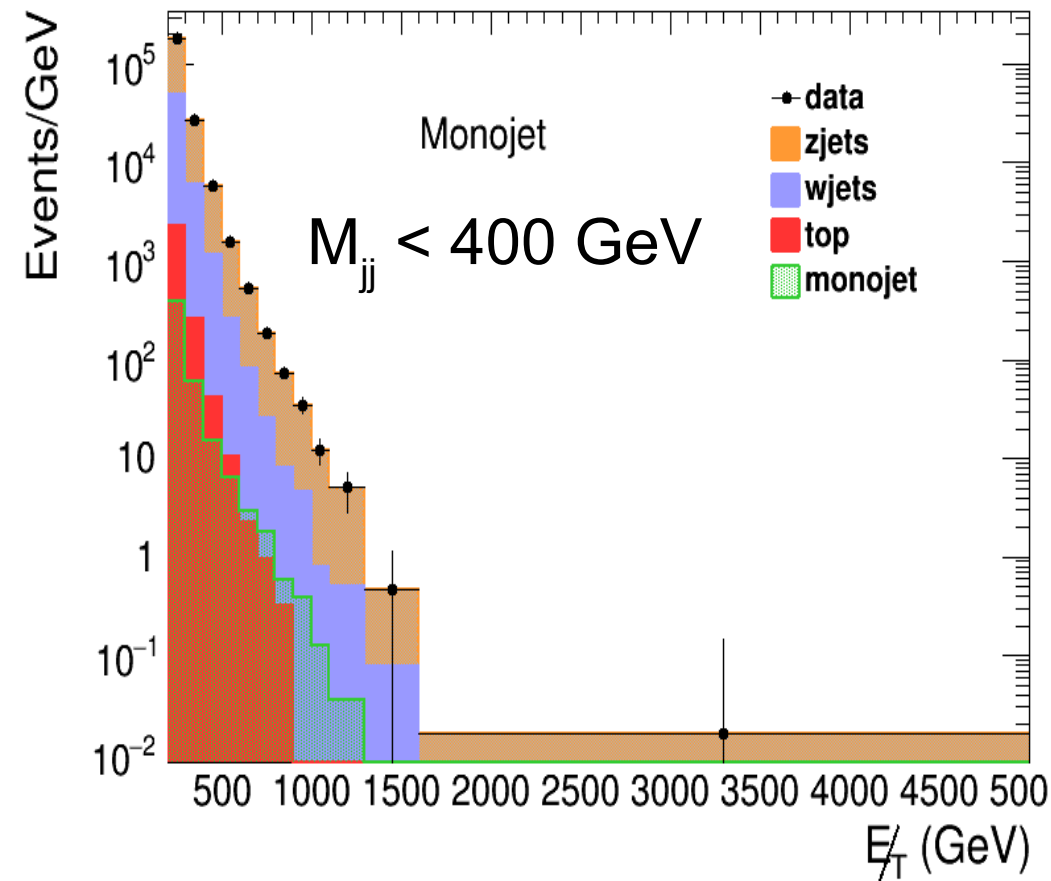
Great way to maintain relic density & bounds

Many ways to extend the dark sector with LLPs

Take current DM searches and understand impact

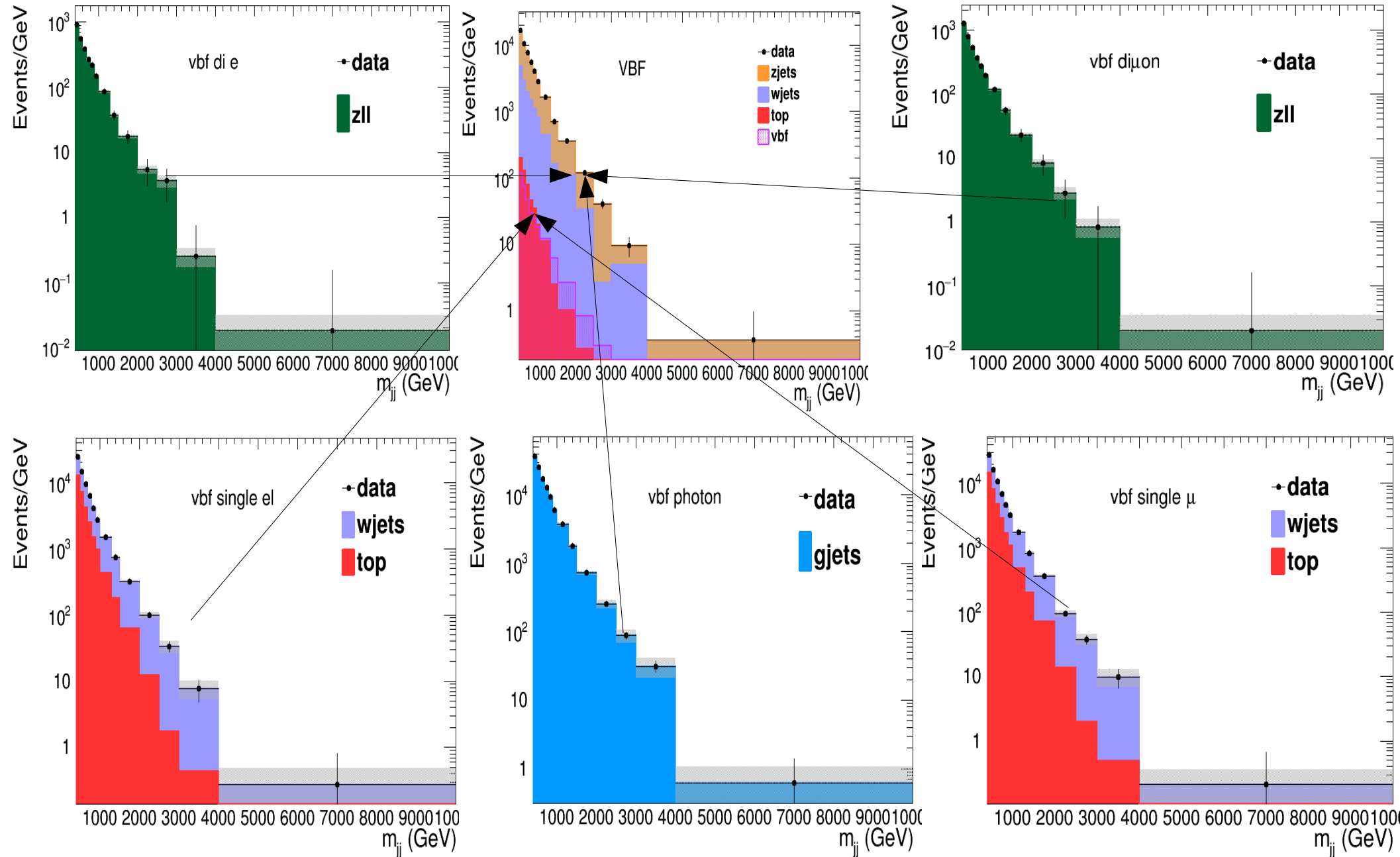
# The Classic Higgs to invisible

- Construct a modern version of Higgs to invisible
  - Dominant signal is VBF
  - Two category fit
  - Pre-selection: No leptons &  $MET > 200$  GeV



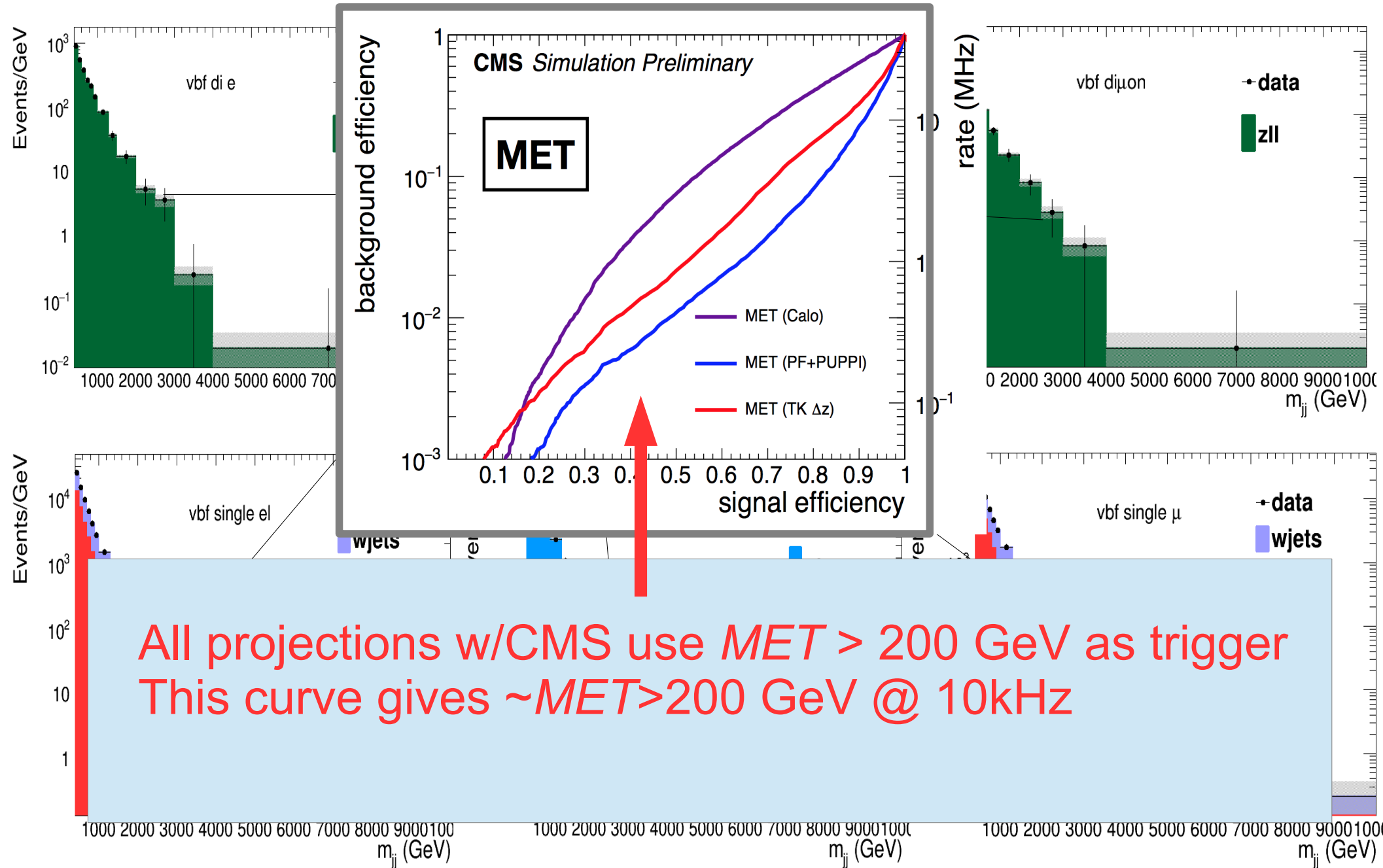
# Fit each bin with a control region

# Full Fit



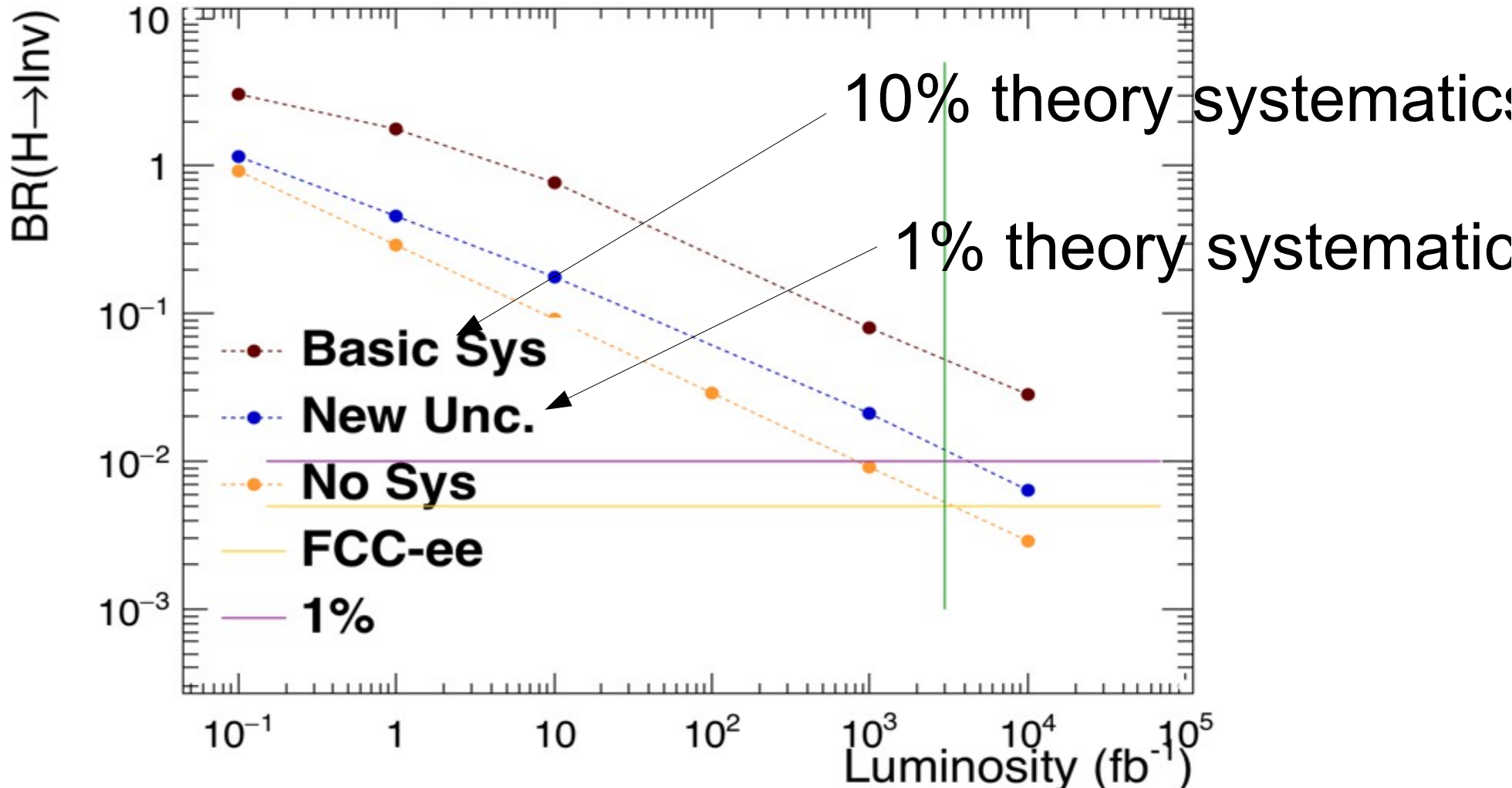
Fit each bin with a control region

Full Fit



# How does it look?

- Higgs Invisible propagated through

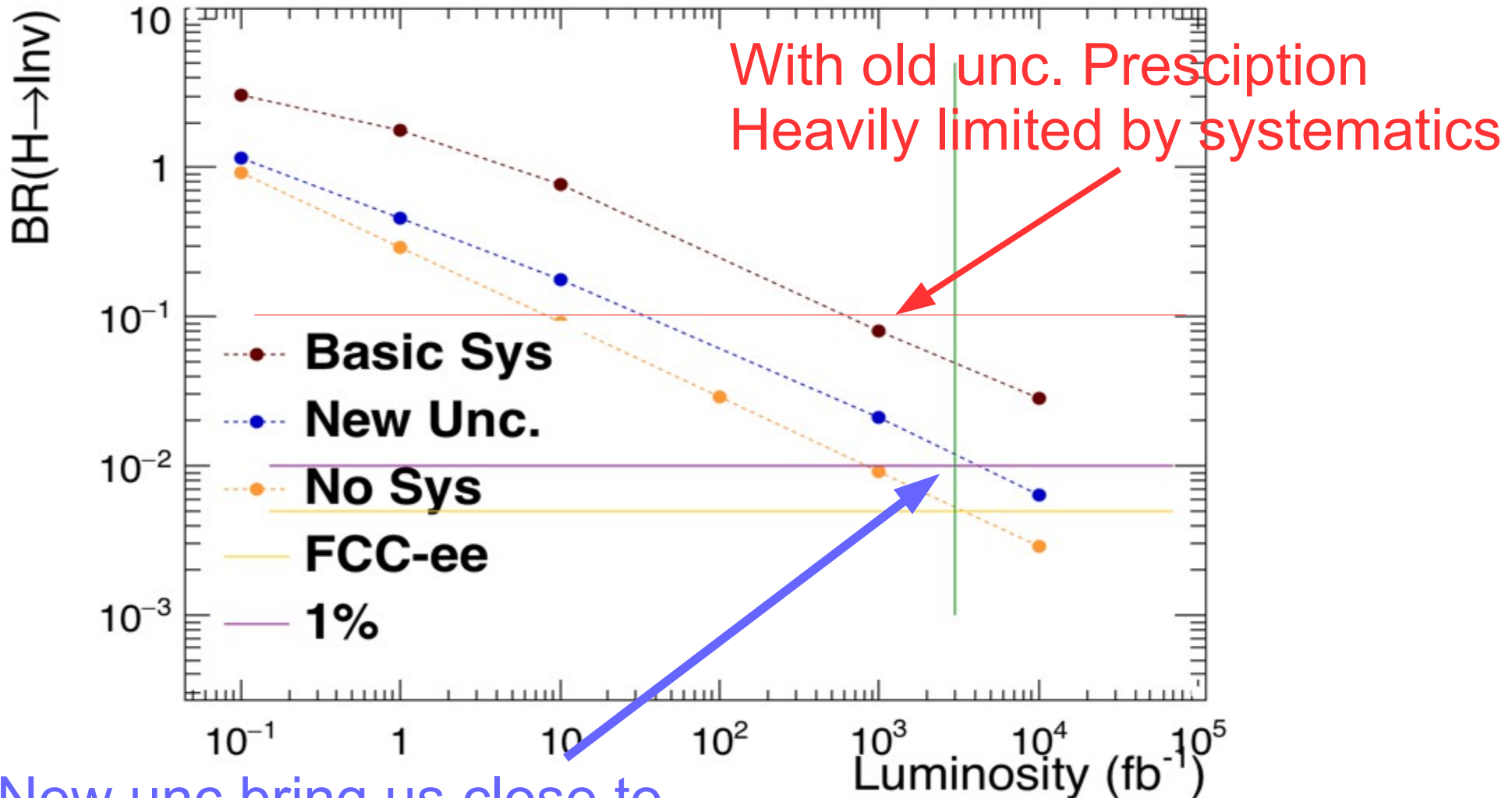


Hit 1% with the full unc. Scheme and  $3 \text{ ab}^{-1}$



# How does it look?

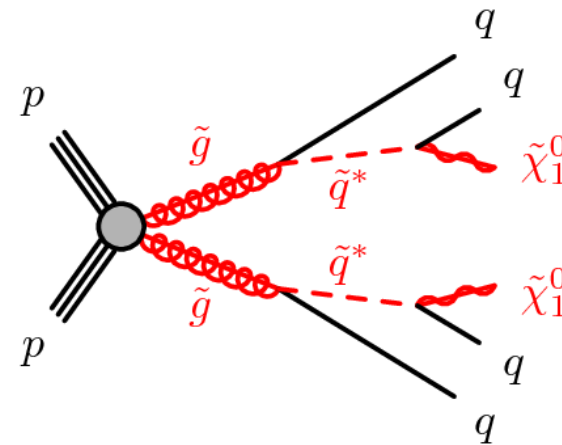
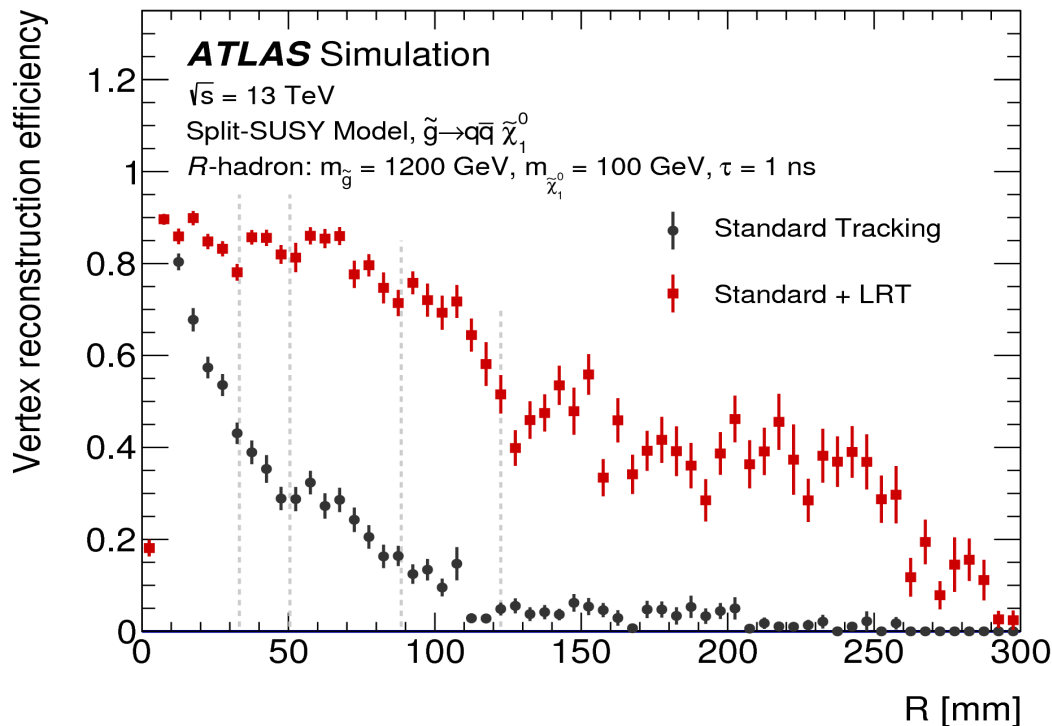
- Higgs Invisible propagated through



New unc bring us close to  
Ultimate sensitivity

# Case #1: a long lived signature

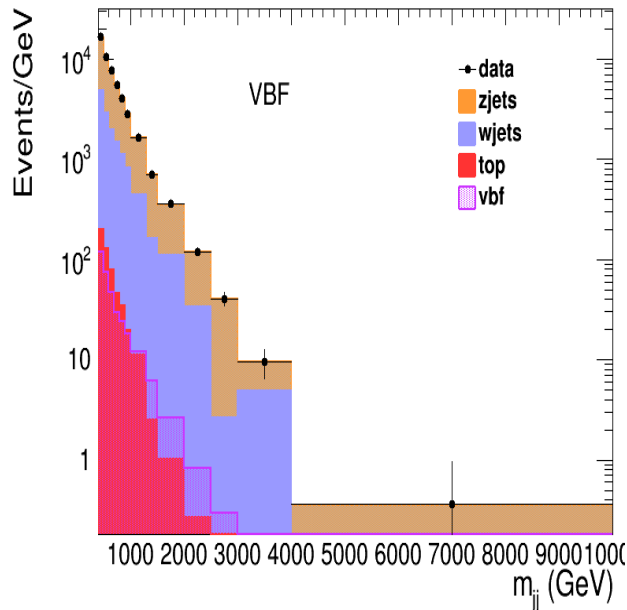
- Follow ATLAS approach  **$MET > 250 \text{ GeV}$** 
  - SUSY-2016-08 :
    - Search for LLP with displaced vertices +  $MET$
  - $MET$  triggered events + displaced vertex tag
    - Cut tight on vertices to have almost 0 bkg events



Requiring a secondary vertex along with 3 tracks and mass  $> 10 \text{ GeV}$  reduces background by 3 orders of magnitude

# Constructing an LL analysis

- Can recast the dark matter searches

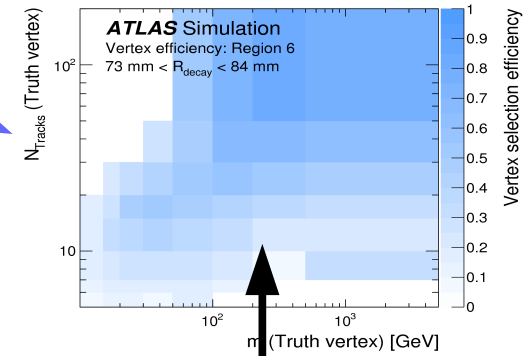
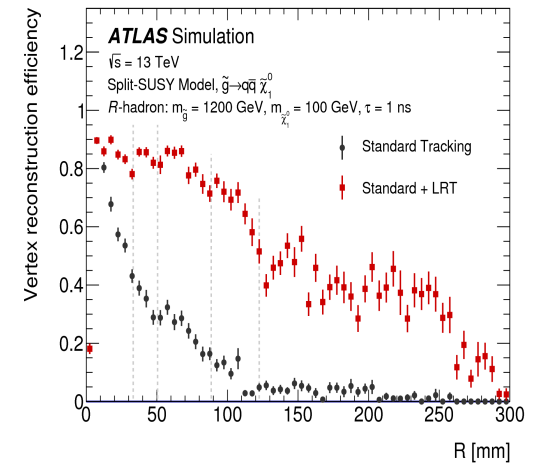


$\times 10^{-2}$

$\times 10^{-4}$

$\times$

Preselection



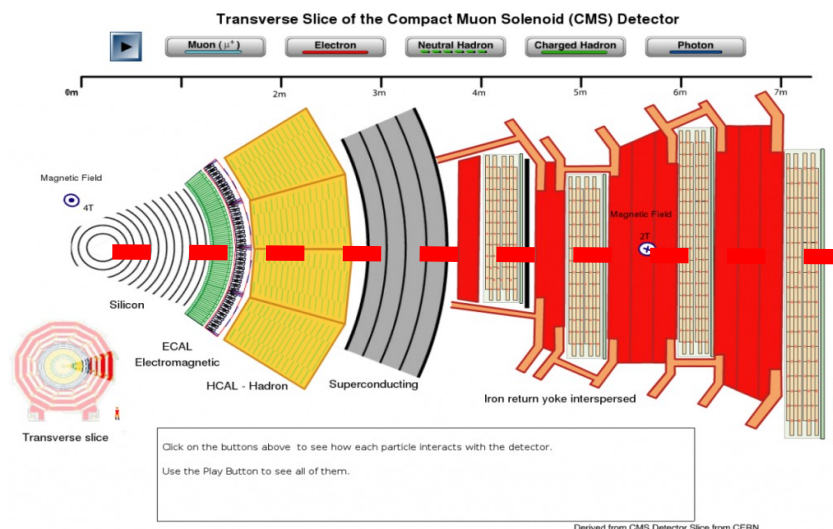
Efficiency map  
given lifetime

Run analysis on reduce background and correct for eff  
(Approximates right scale)

# Case #2: MATHUSLA signature

- Follow conventional MATHUSLA approach
  - Take the same signal model
  - Assume zero background
  - Count number of MATHUSLA interactions
  - If interactions greater than  $\geq 4$  then exclude

**Compute the probability for this**

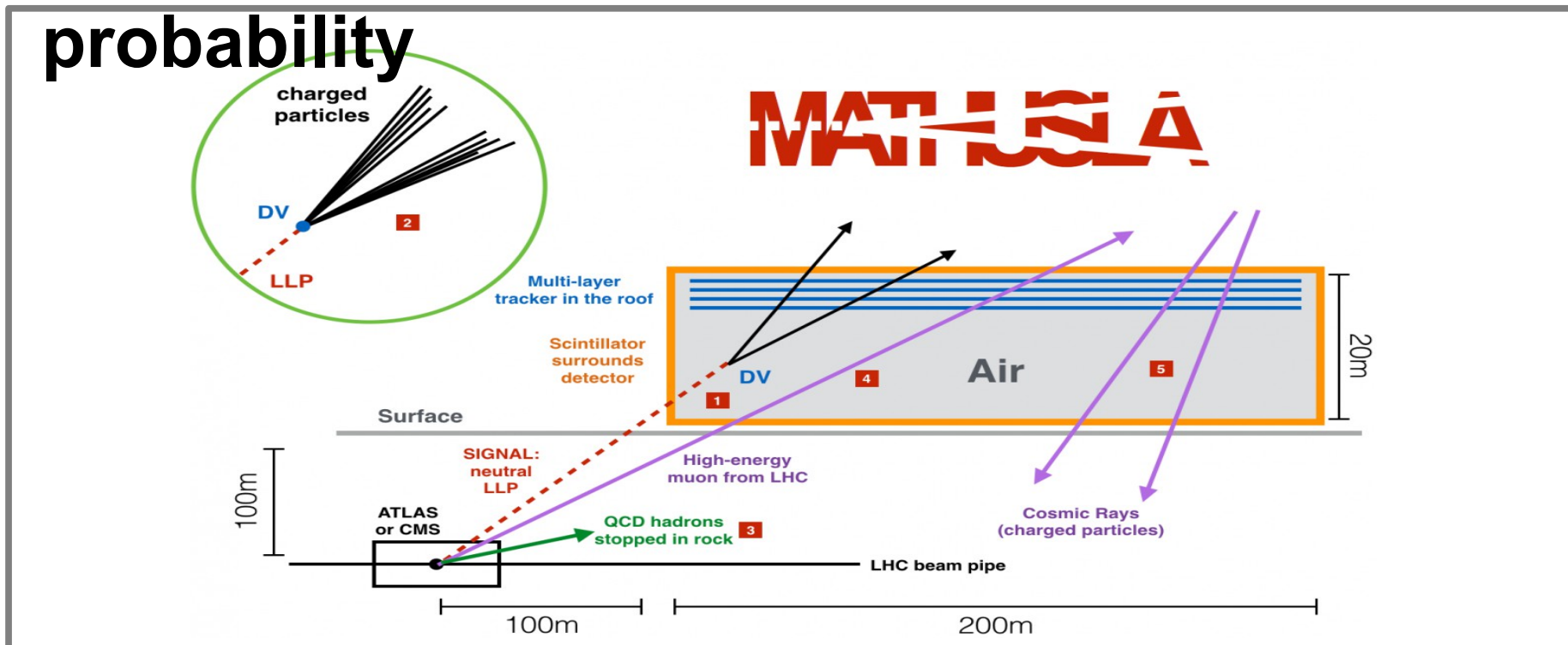


Large  
detector  
Outside  
of CMS

MATHUSLA  
volume

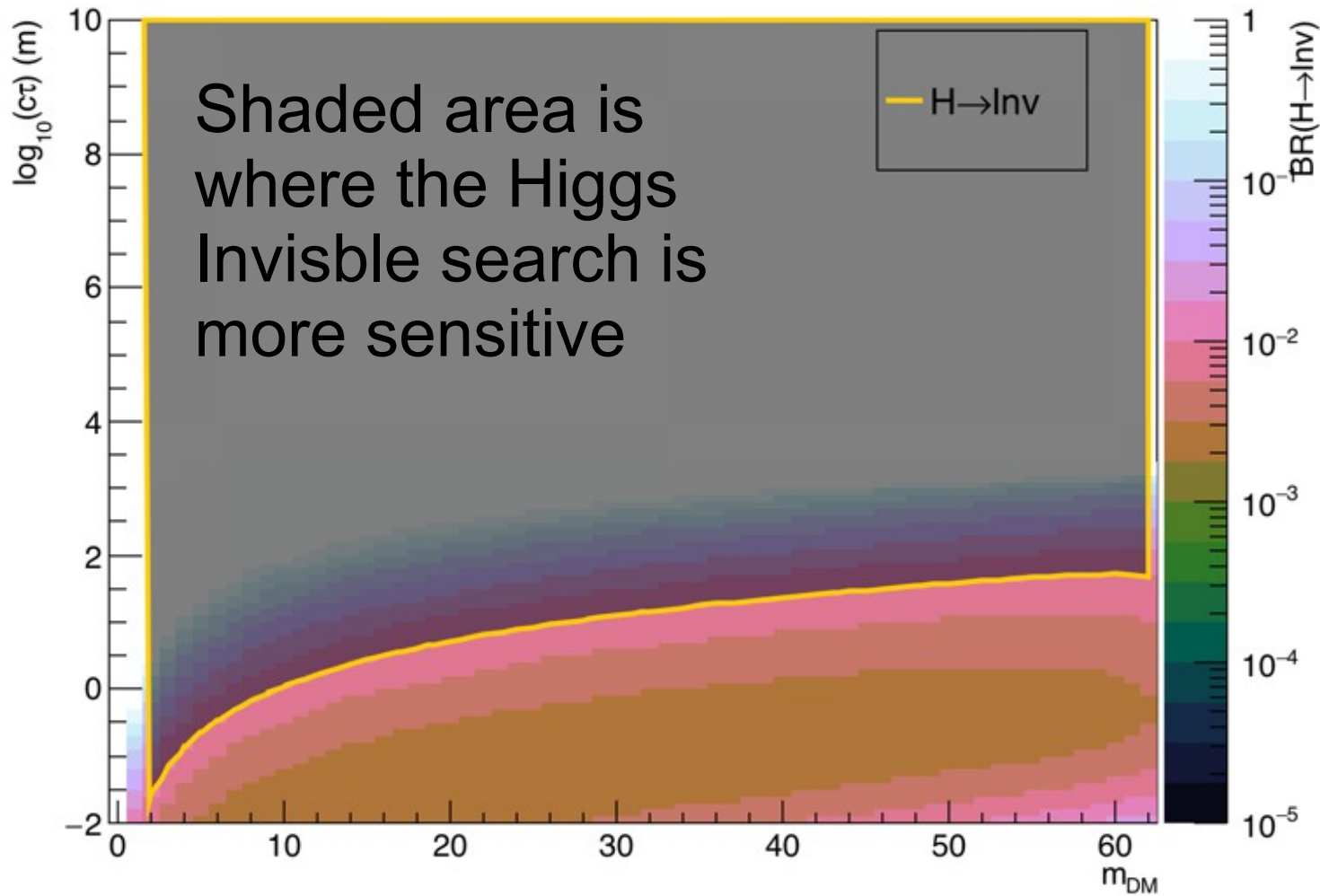
# Case #2: MATHUSLA signature

- Follow conventional MATHUSLA approach
  - Take the same signal model
  - Assume zero background
  - Count number of MATHUSLA interactions
  - If interactions greater than  $\geq 4$  then exclude



# Lifetime Bounds w/Tracker

- Higgs Invisible in with a decay in tracker



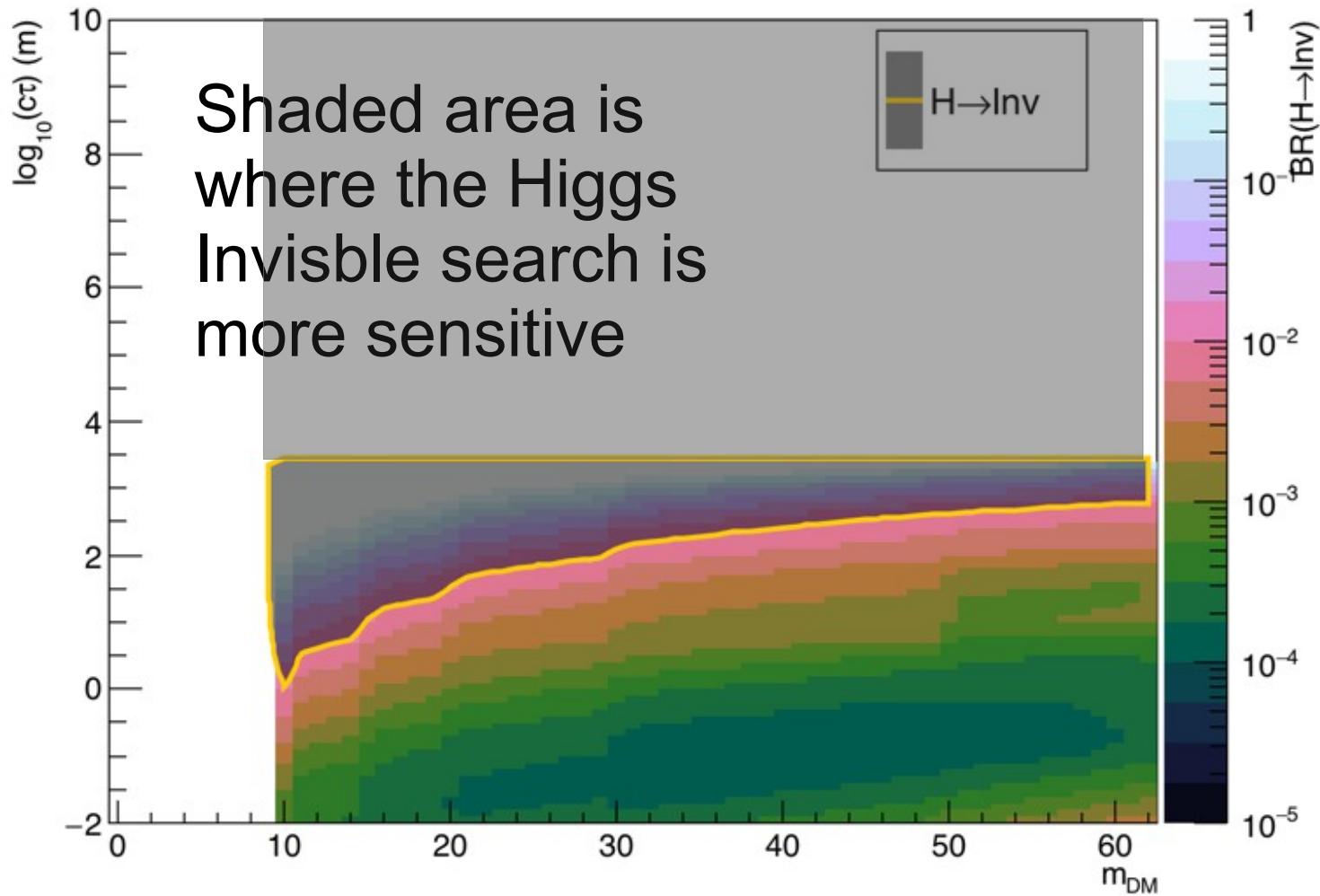
This result is  
With the  
preselection only

$10^{-3}$

Covers the lower range beyond  $H \rightarrow \text{Inv}$  ( $10^{-3}$ )

# Lifetime Bounds w/Tracker

- Higgs Invisible in with a decay in tracker



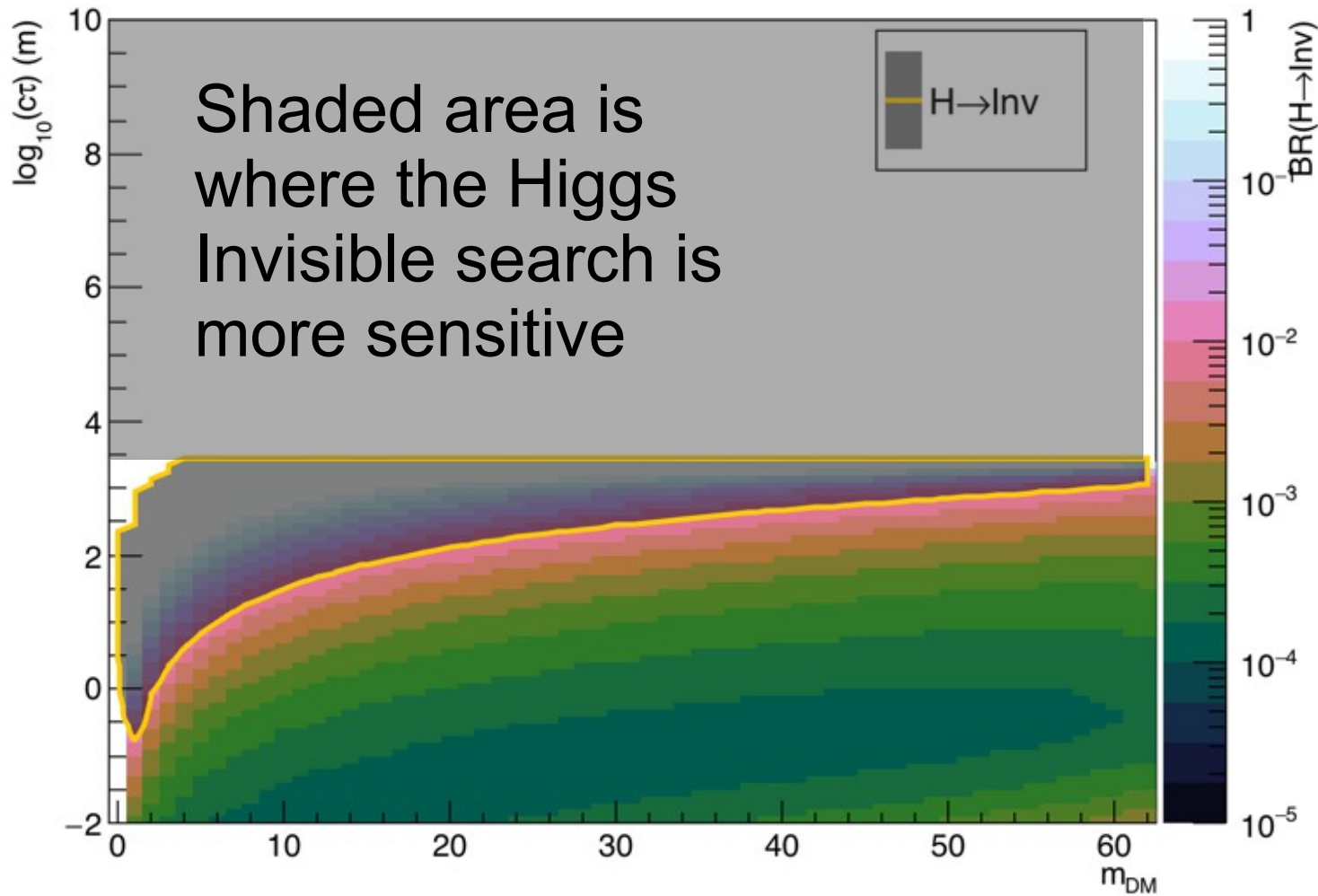
This result is  
With the full  
selection  
& parametrization

$10^{-4}$

Covers the lower range beyond  $H \rightarrow \text{Inv}$  ( $10^{-3}$ )

# Lifetime Bounds w/Tracker

- Higgs Invisible in with a decay in tracker



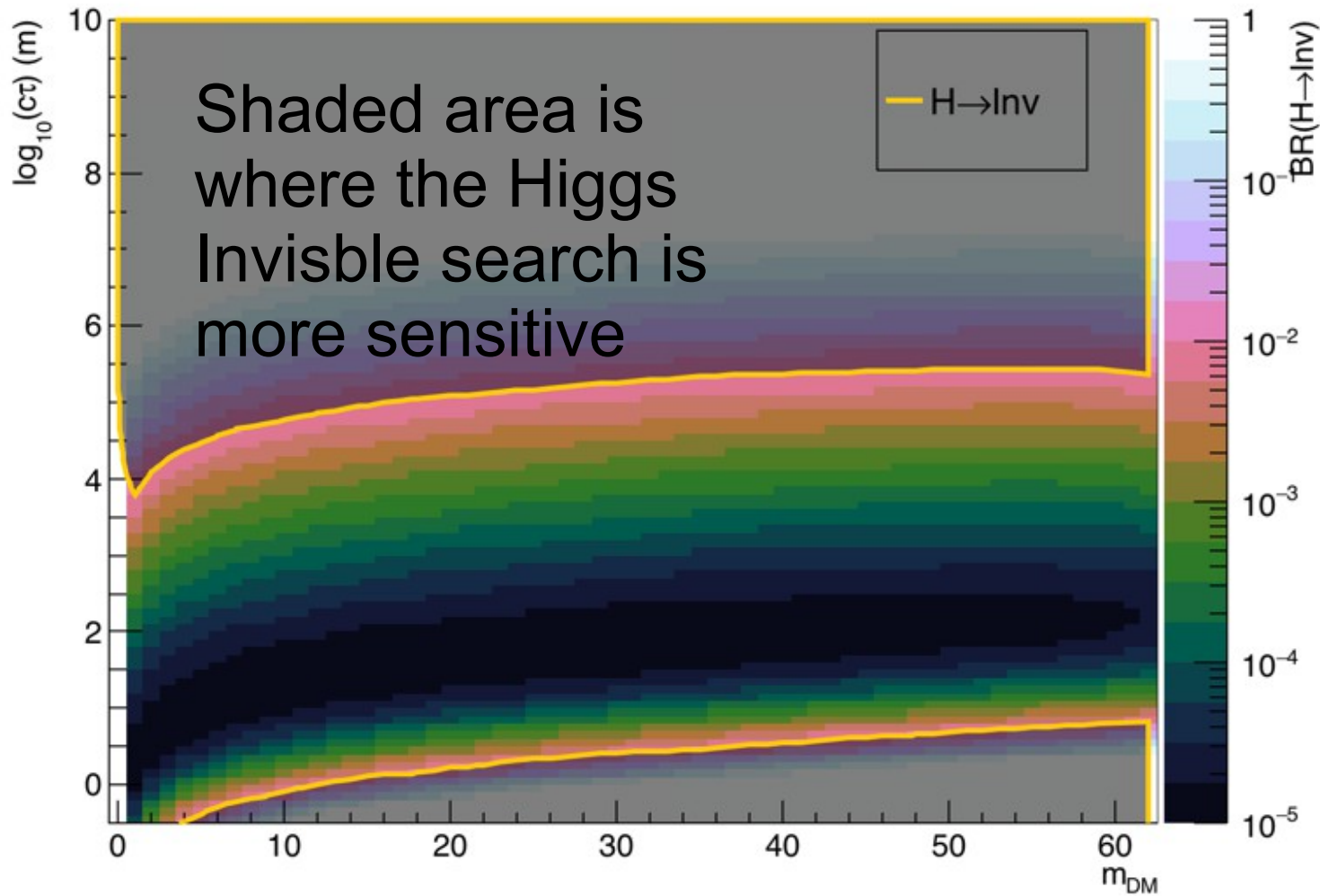
This result is  
With the full  
selection &  
modified  
Parametrization  
That goes to 0  
 $10^{-4}$

Covers the lower range beyond  $H \rightarrow \text{Inv}$  ( $10^{-3}$ )



# Lifetime Bounds w/Tracker

- Higgs Invisible in with a decay in tracker



This result is  
With the  
preselection only

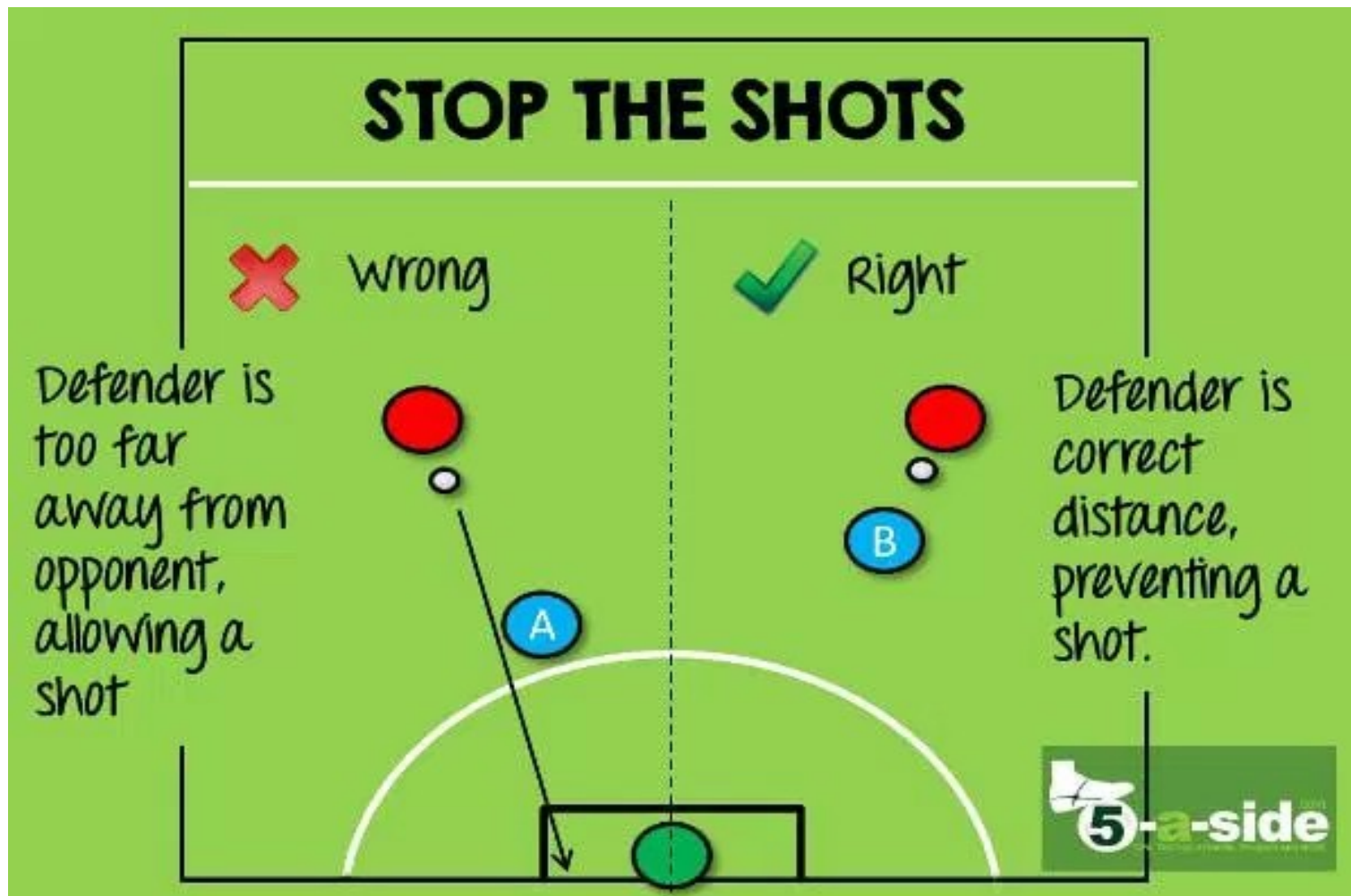
$10^{-5}$

Covers the lower range beyond H→Inv ( $10^{-3}$ )

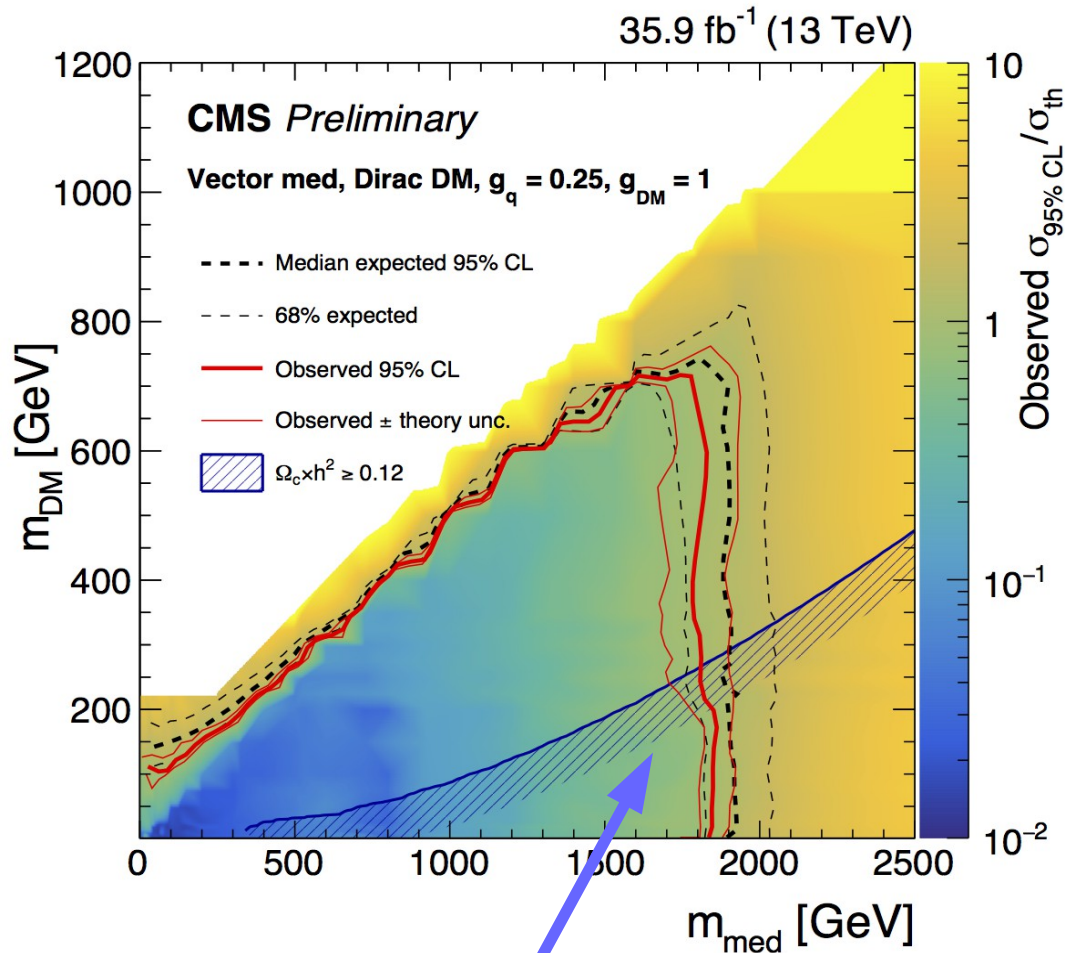
# Fate of the Lifetime

- With higgs to invisible as a benchmark
  - Can probe coefficients down to very small values
  - Detector like MATHUSLA reaches order of mag more
    - What if we have full trigger acceptance for others?
    - In principle acceptance can increase significantly
    - How important are out of volume detectors?
- Triggers for long lived searches possible?
  - Projects listed have been motivated by *MET* trigger
  - Can we build a more sophisticated trigger?
- Is MATHUSLA the appropriate volume?

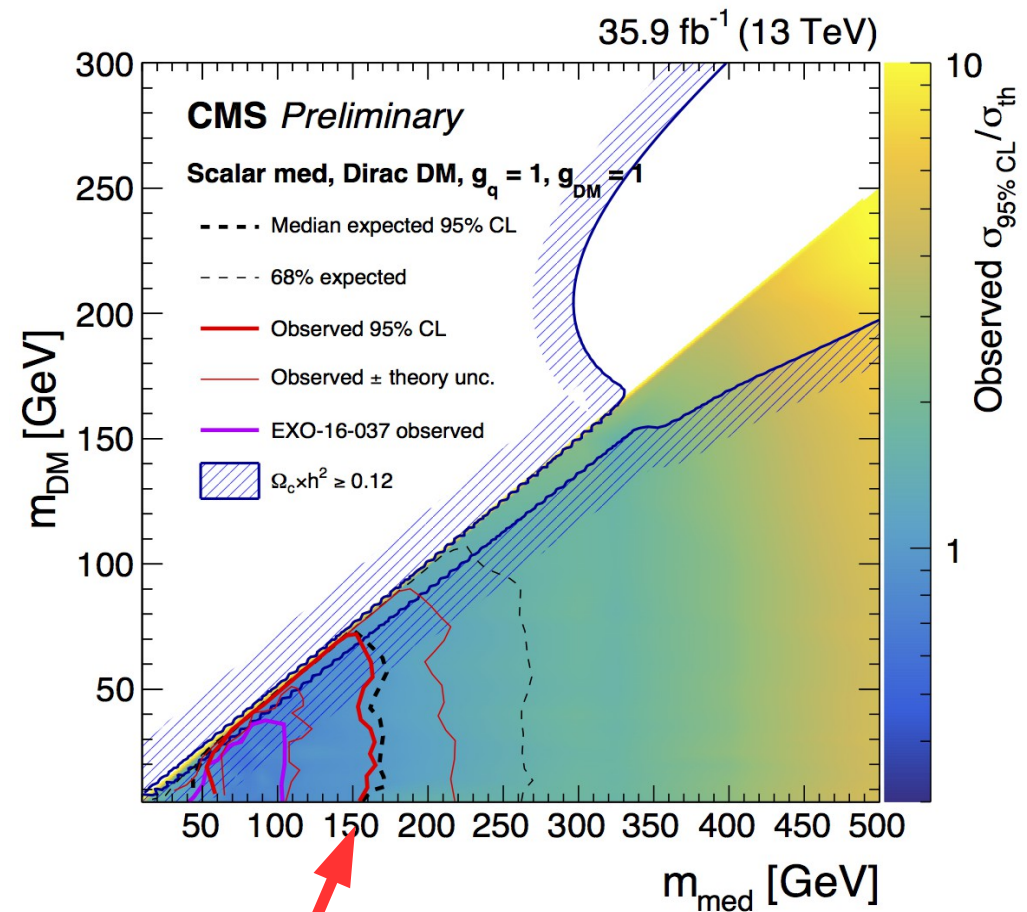
# Where to put a detector?



# What about other DM models?



Vector (Spin 1) has been pushed out far and is well covered by monojet and dijet

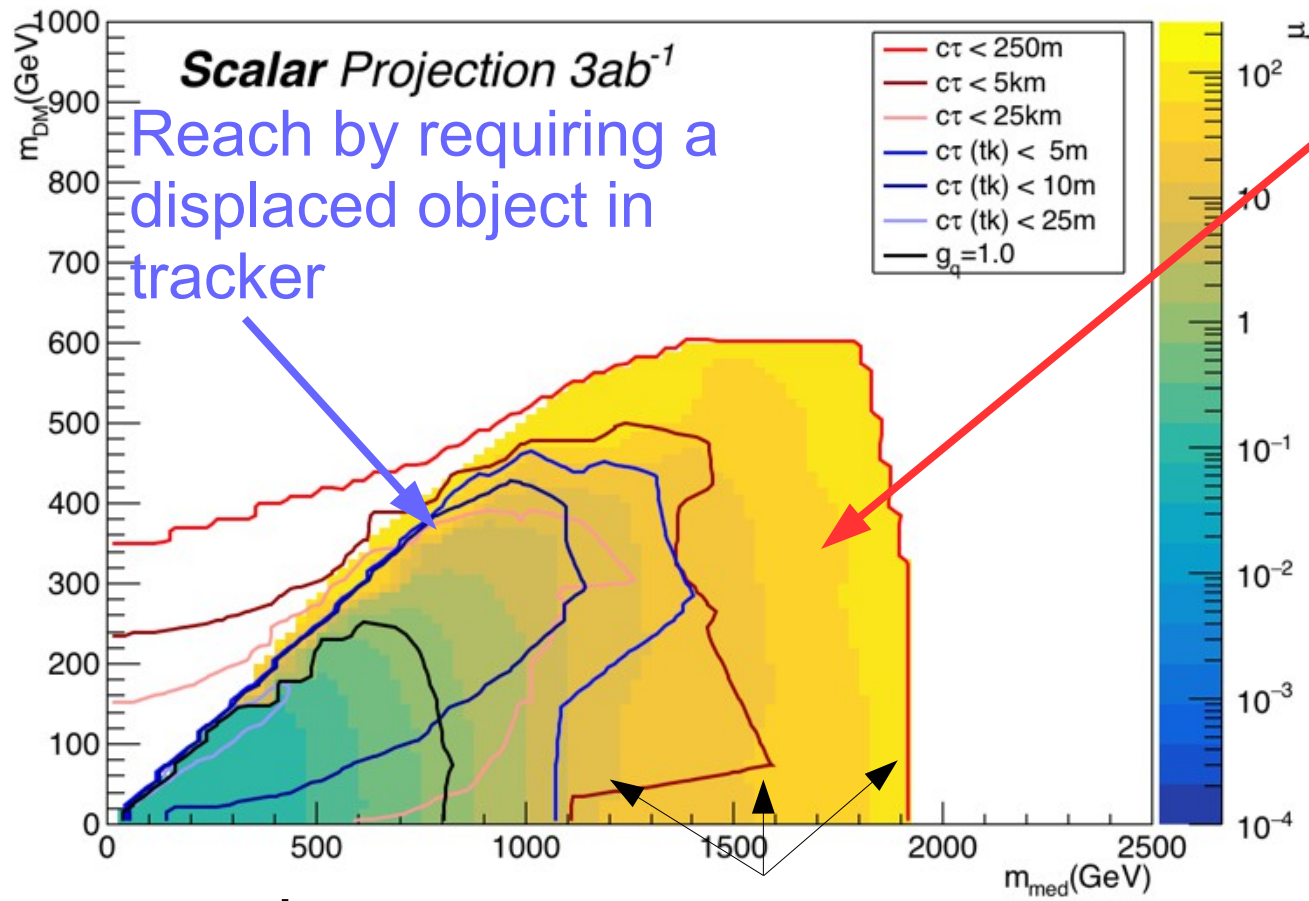


Scalar and pseudoscalar models are just starting to be excluded

# Scalar Simplified models: $3ab^{-1}$

- Scalar Bounds
  - Similar gains for Higgs invisible

Reach by requiring a displaced object in the MATHUSLA detector



In the low mass region

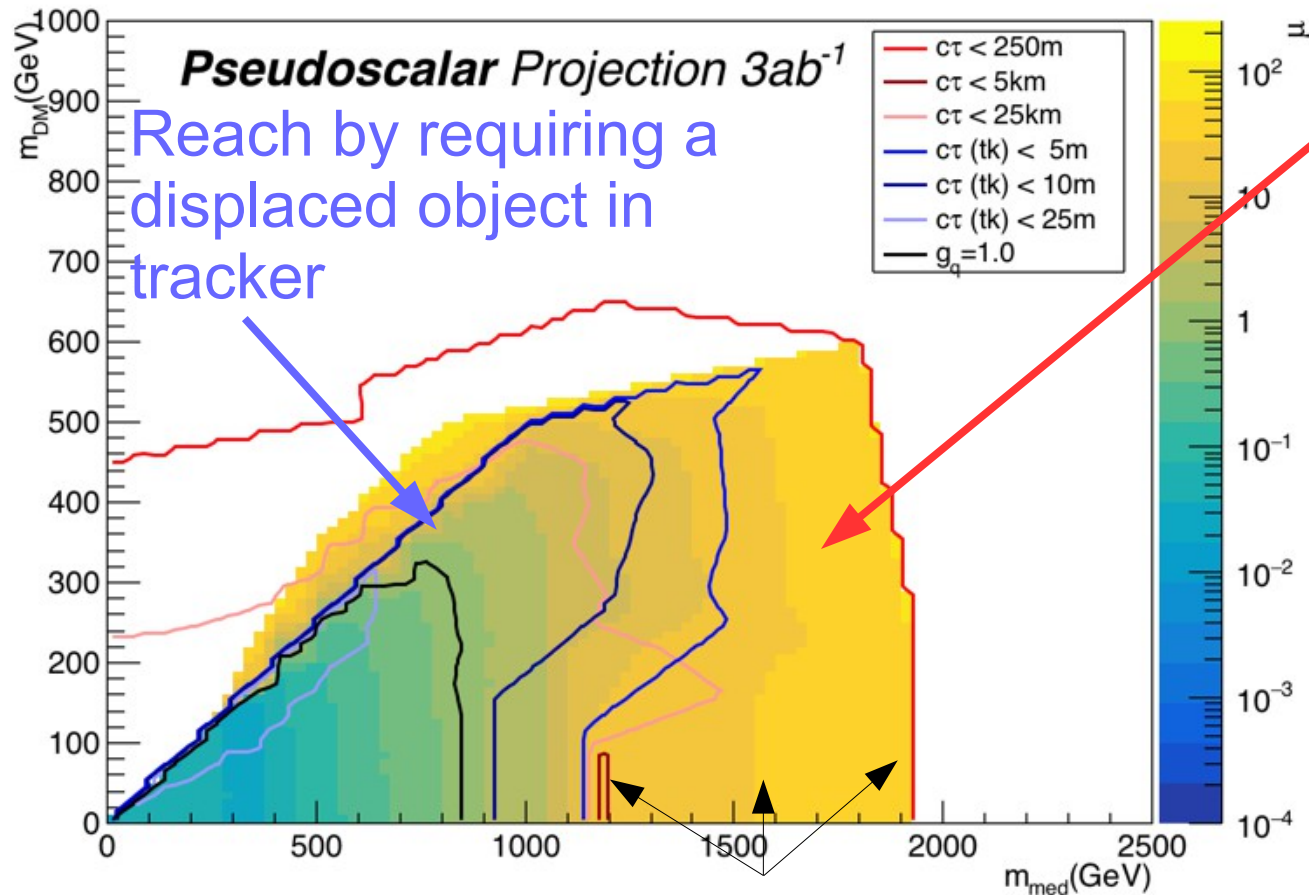
Note: interpolation error these should curve down

gains from LL searches more important (@high mass its less)

# Pseudoscalar models: $3ab^{-1}$

- Pseudoscalar Bounds
  - Stronger bounds than scalar

Reach by requiring a displaced object in the MATHUSLA detector



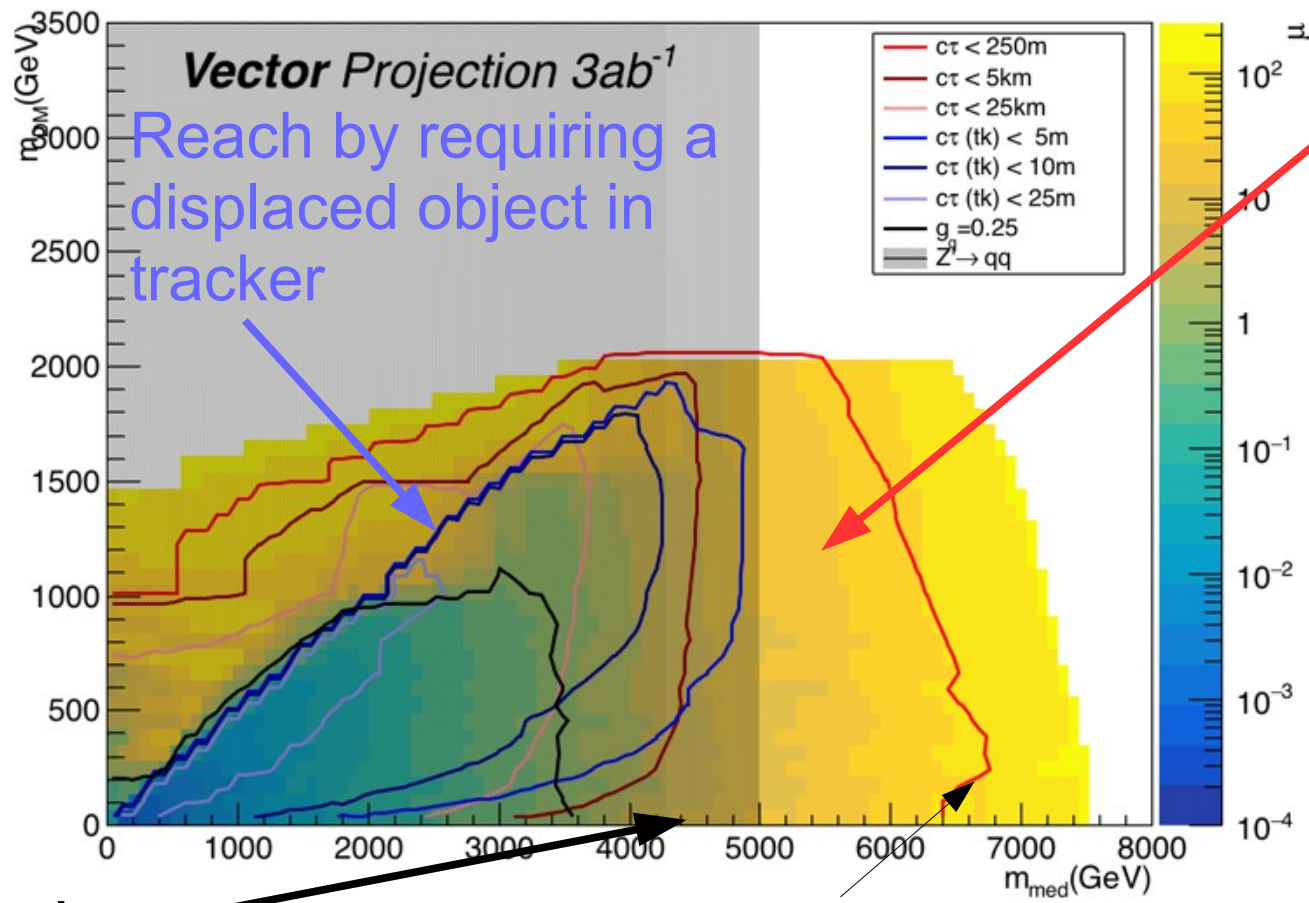
Note: interpolation error these should curve down

Tracker based displaced vertices can push out beyond a TeV

# Vector Simplified models: $3ab^{-1}$

- Vector Bounds
  - Now contend with dijet bounds

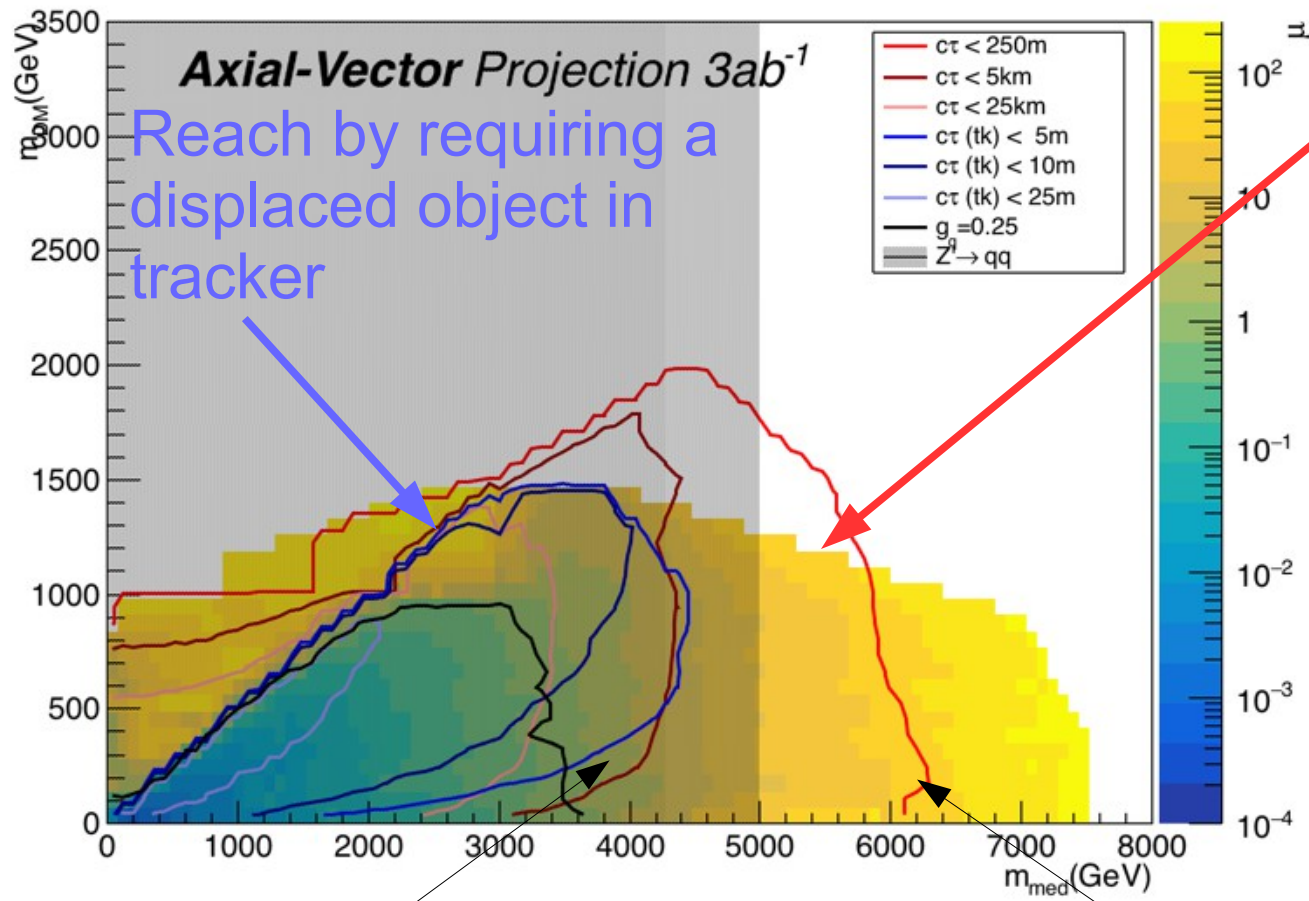
Reach by requiring a displaced object in the MATHUSLA detector



# Axial-Vector Simplified models: $3ab^{-1}$

- Axial-vector Bounds
  - Have to contend with dilepton

Reach by requiring a displaced object in the MATHUSLA detector

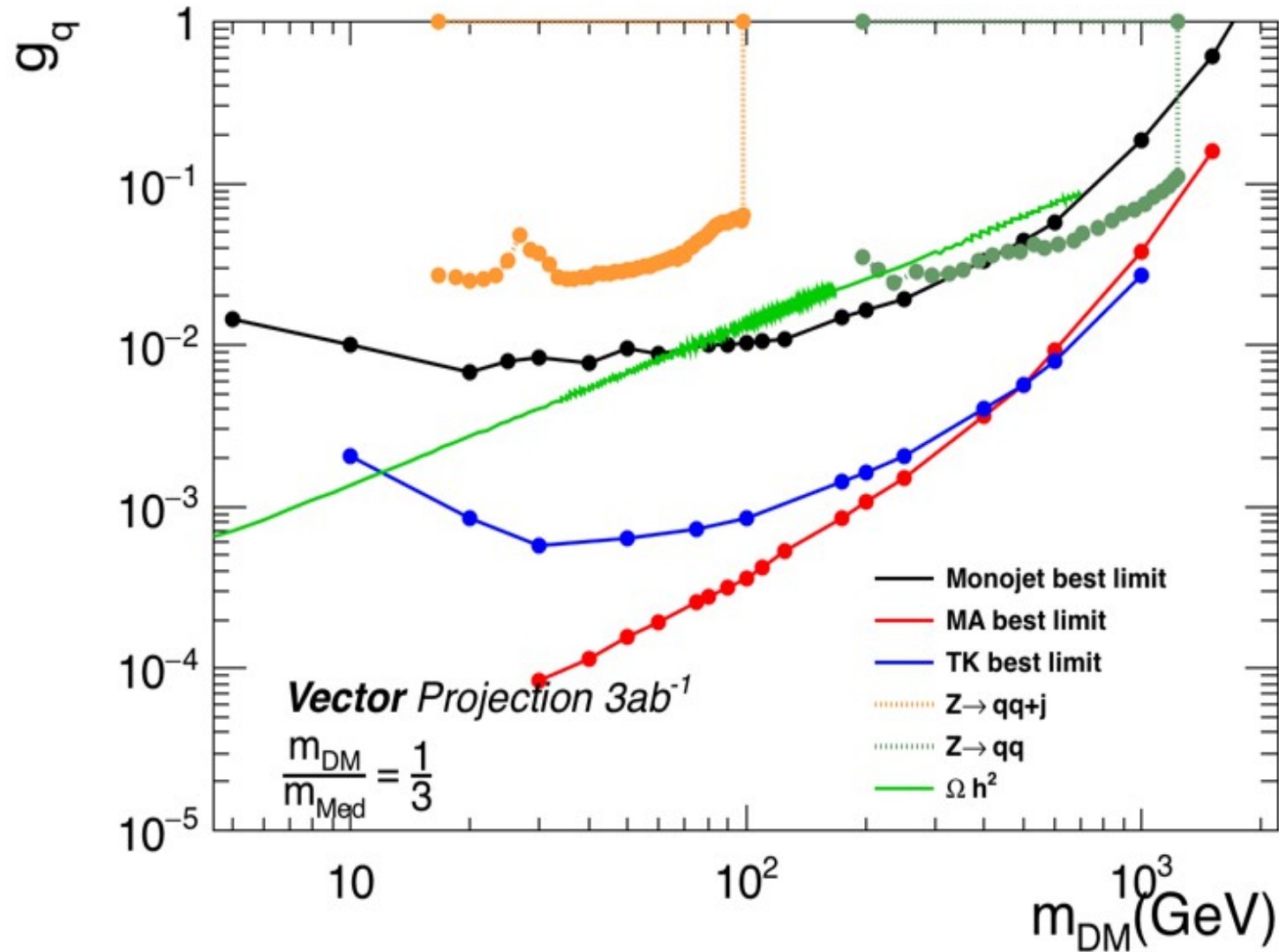


Dijet bounds

Dilepton bounds will go to roughly 6-7TeV

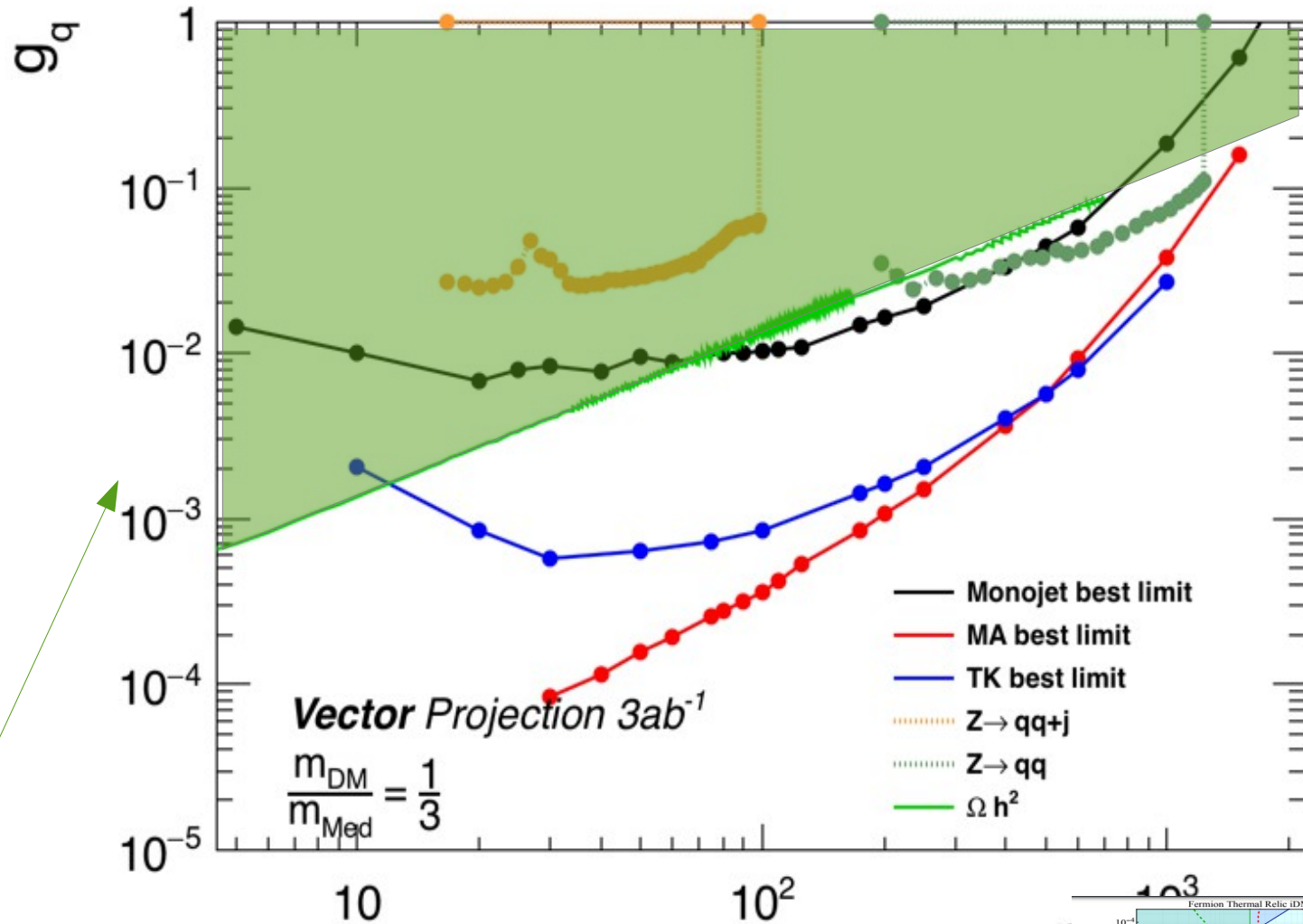


# Coupling Bound

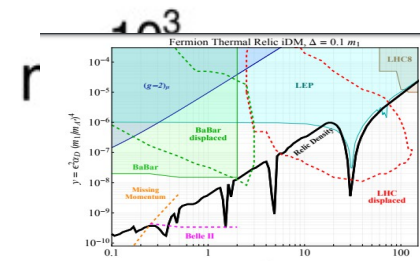


Note with lifetime added  
 Relic bound is roughly the same

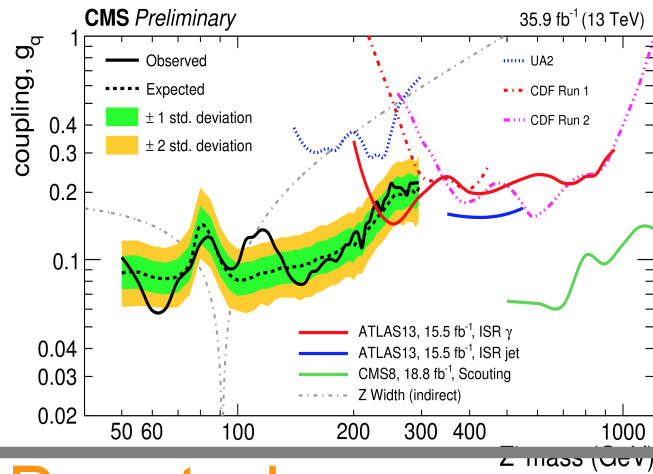
# Coupling Bound



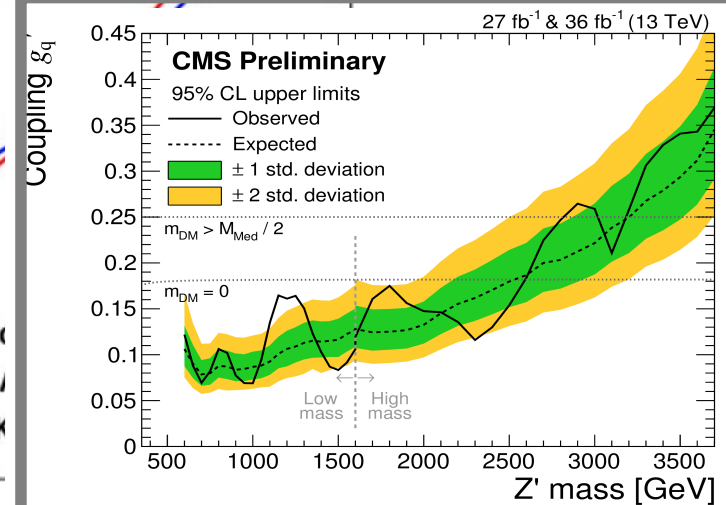
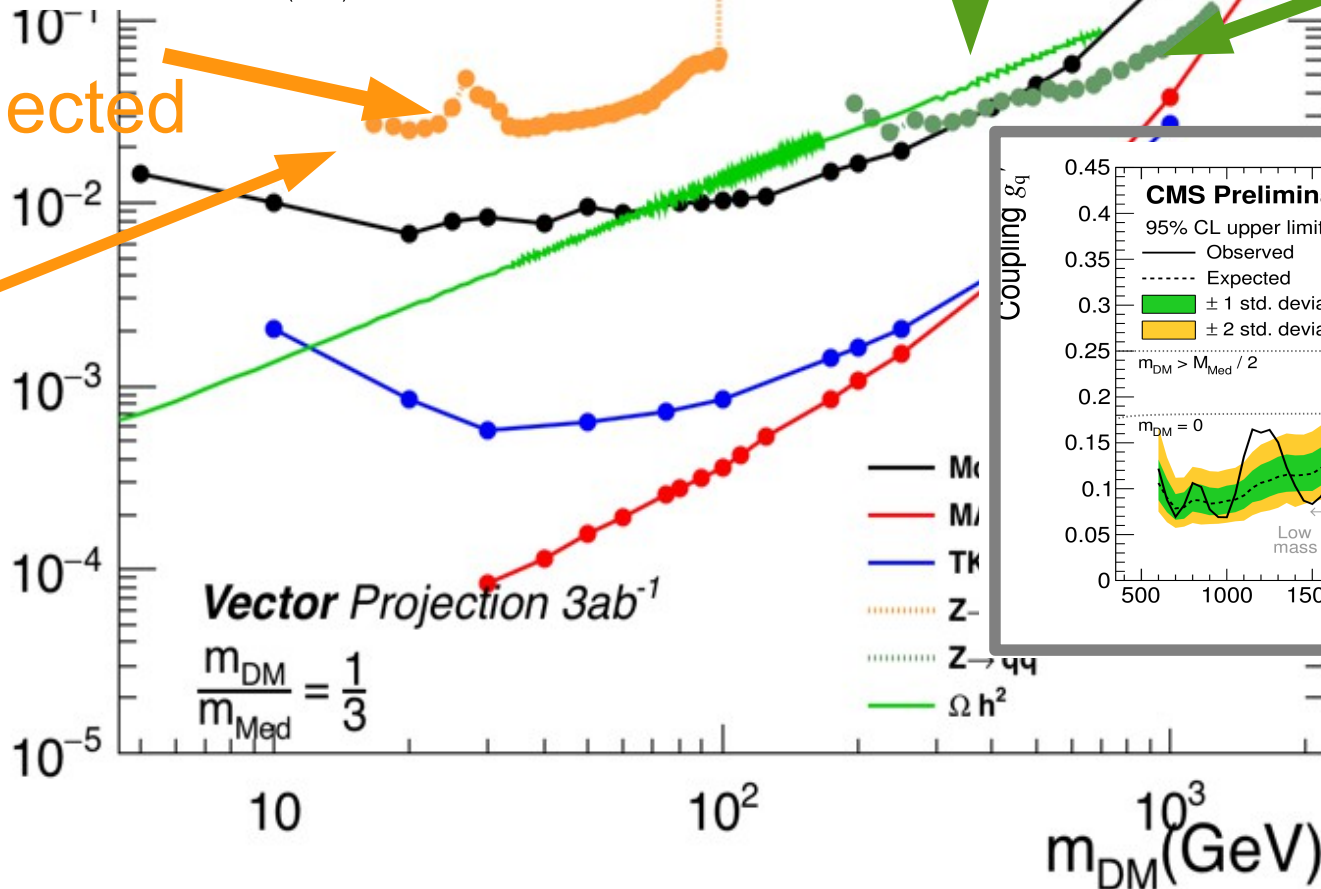
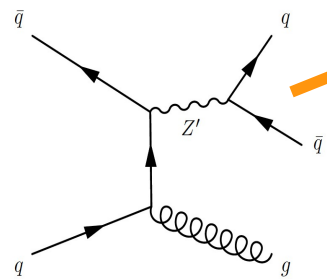
Relic bound below we overclose  
 (see <https://arxiv.org/abs/1508.03050>)



# Coupling Bound



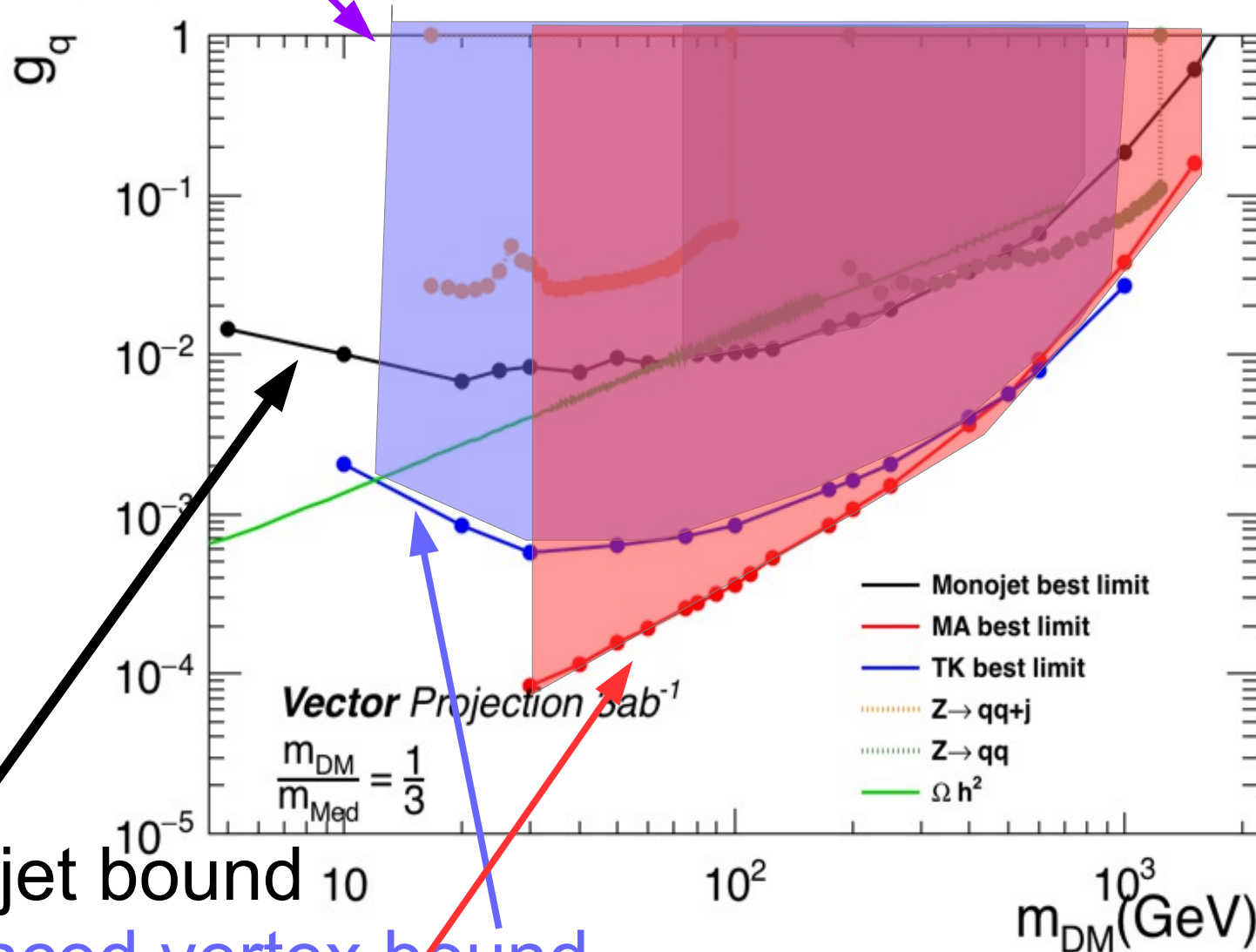
Boosted  
Dijet Projected



Dijet  
Projected

# Coupling Bound

For masses down to 10  
exclude all couplings that  
give relic

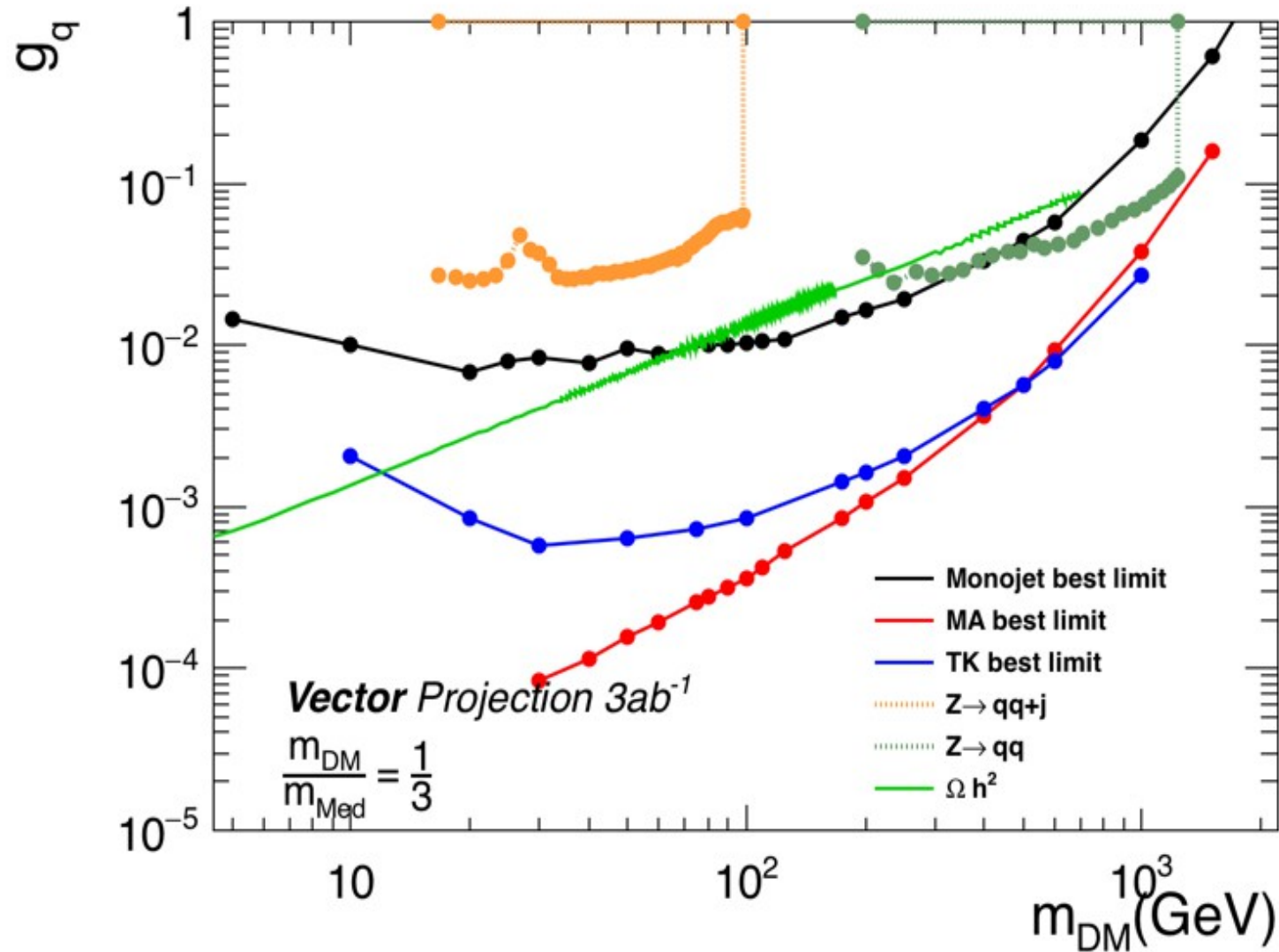


Monojet bound

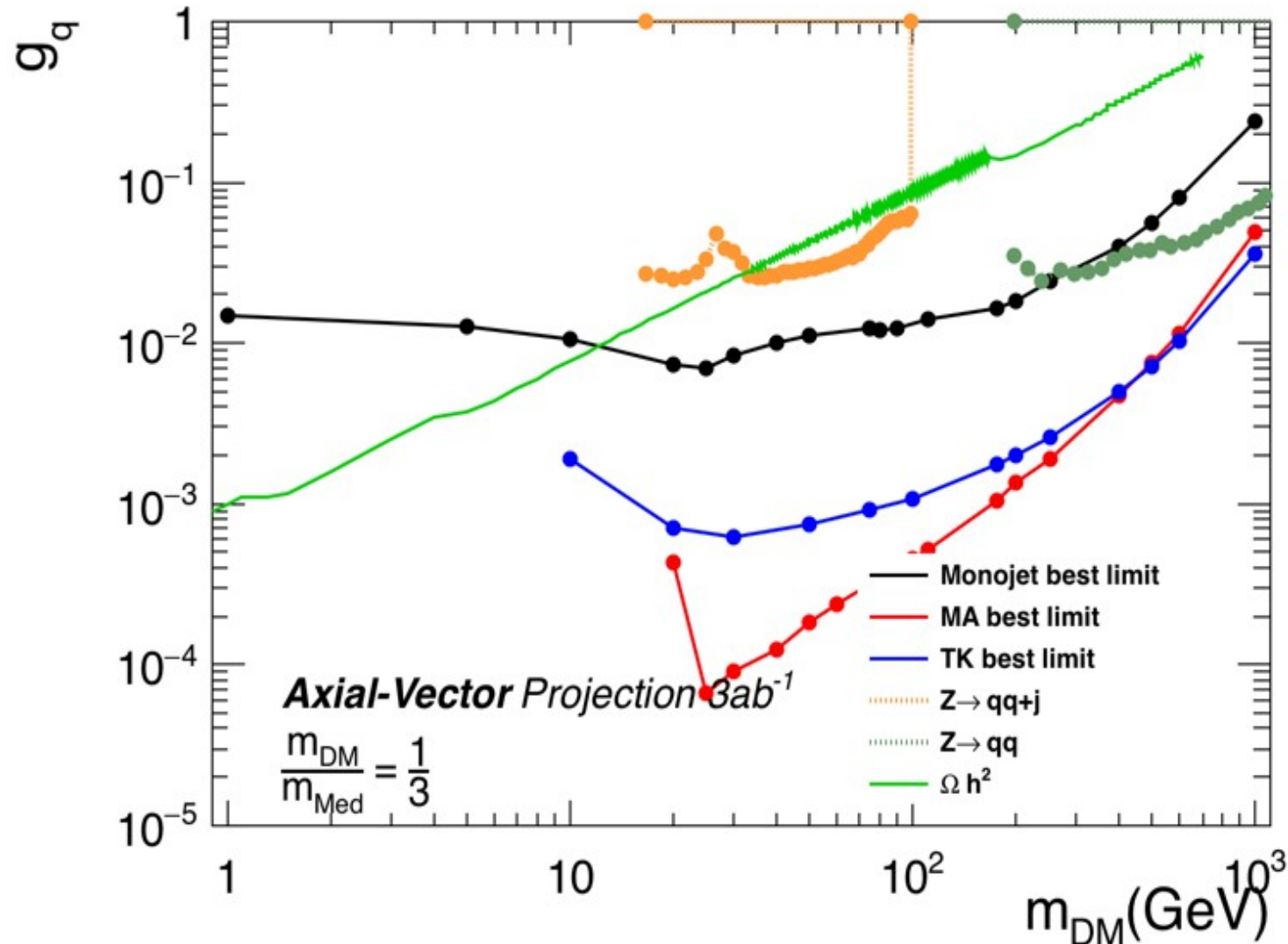
Displaced vertex bound

Mathusla bound

# Vector Coupling Bound

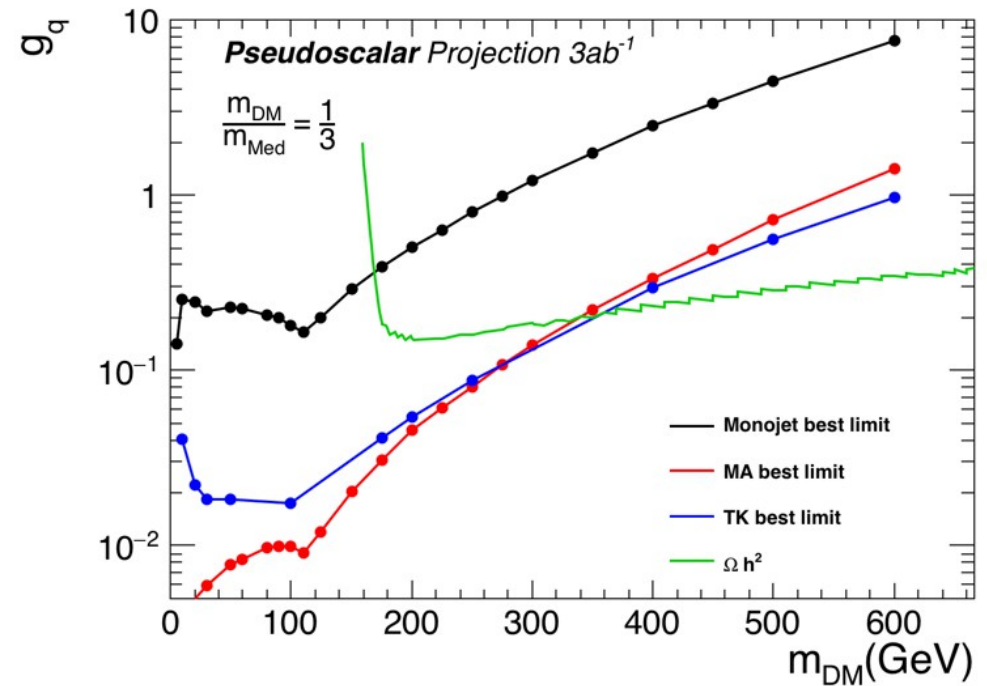
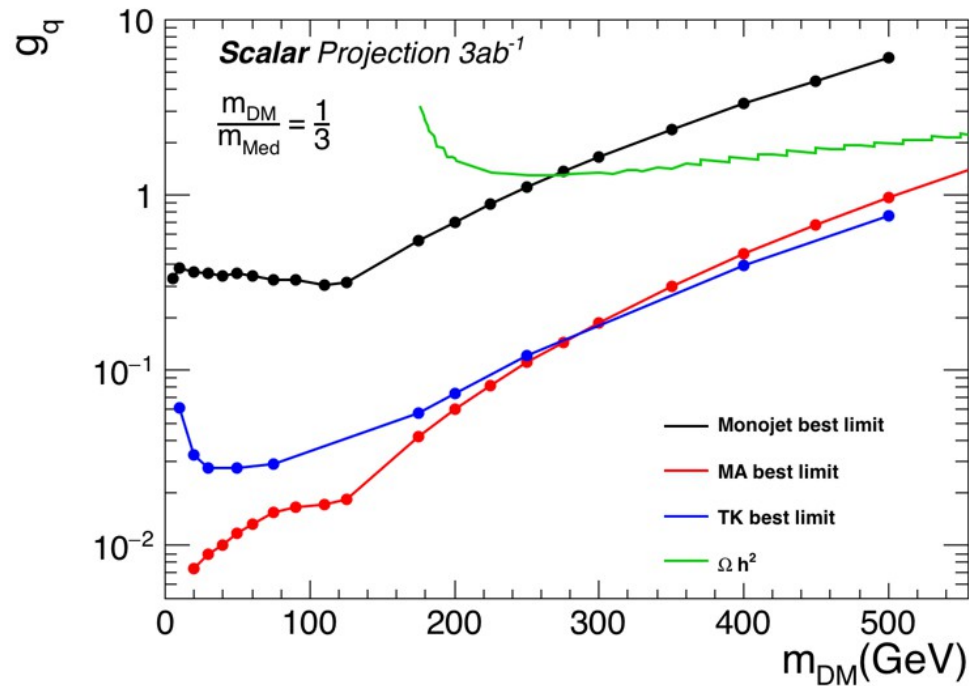


# Axial-Vector Coupling Bound



Bounds are strong for Axial vector coupling

# (Pseudo) Scalar Coupling Bounds



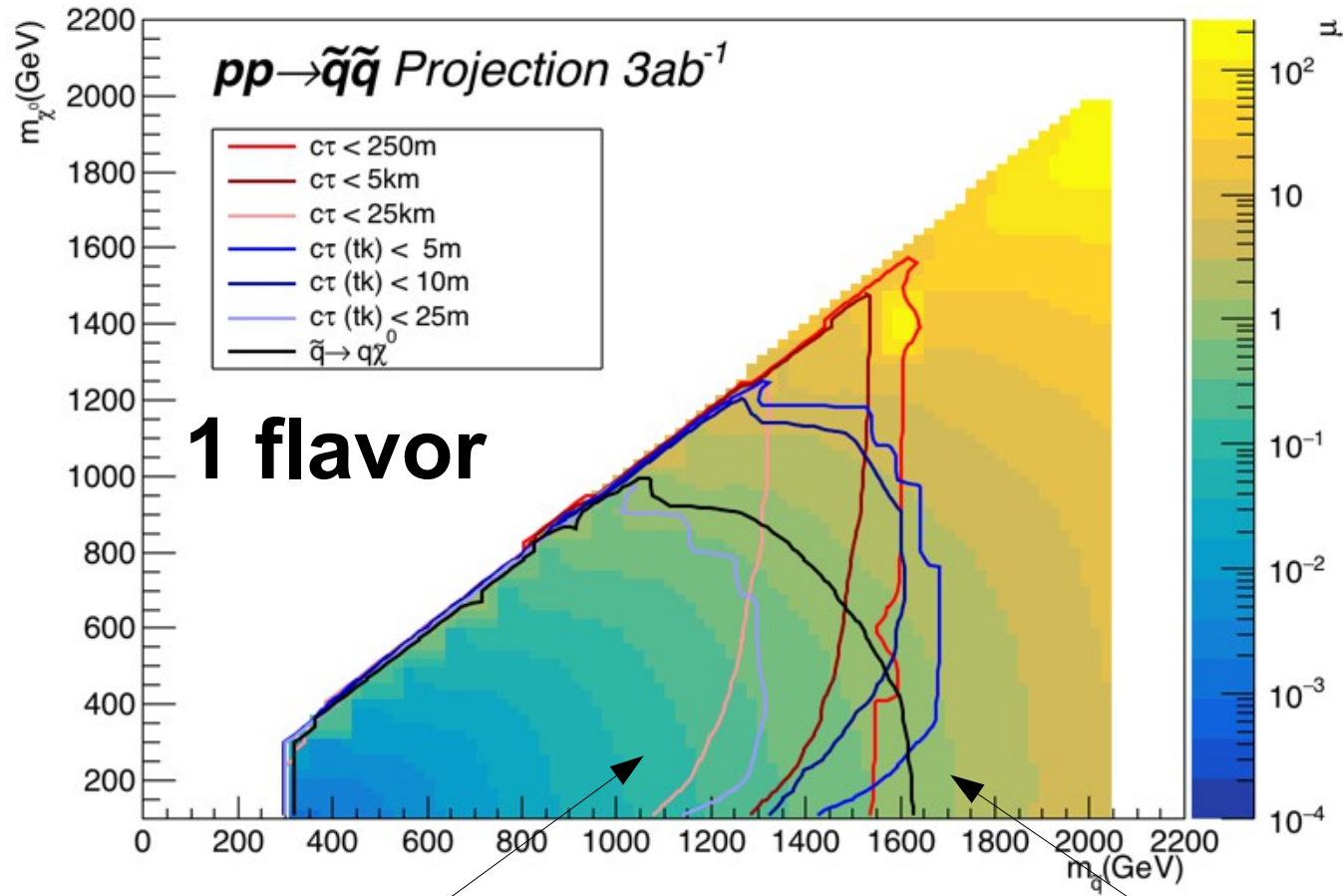
Long lived searches down to small couplings

Full implications need to be investigated

# A SUSY example

- SUSY bounds
  - Lots of heavy objects

With heavy objects its harder to win

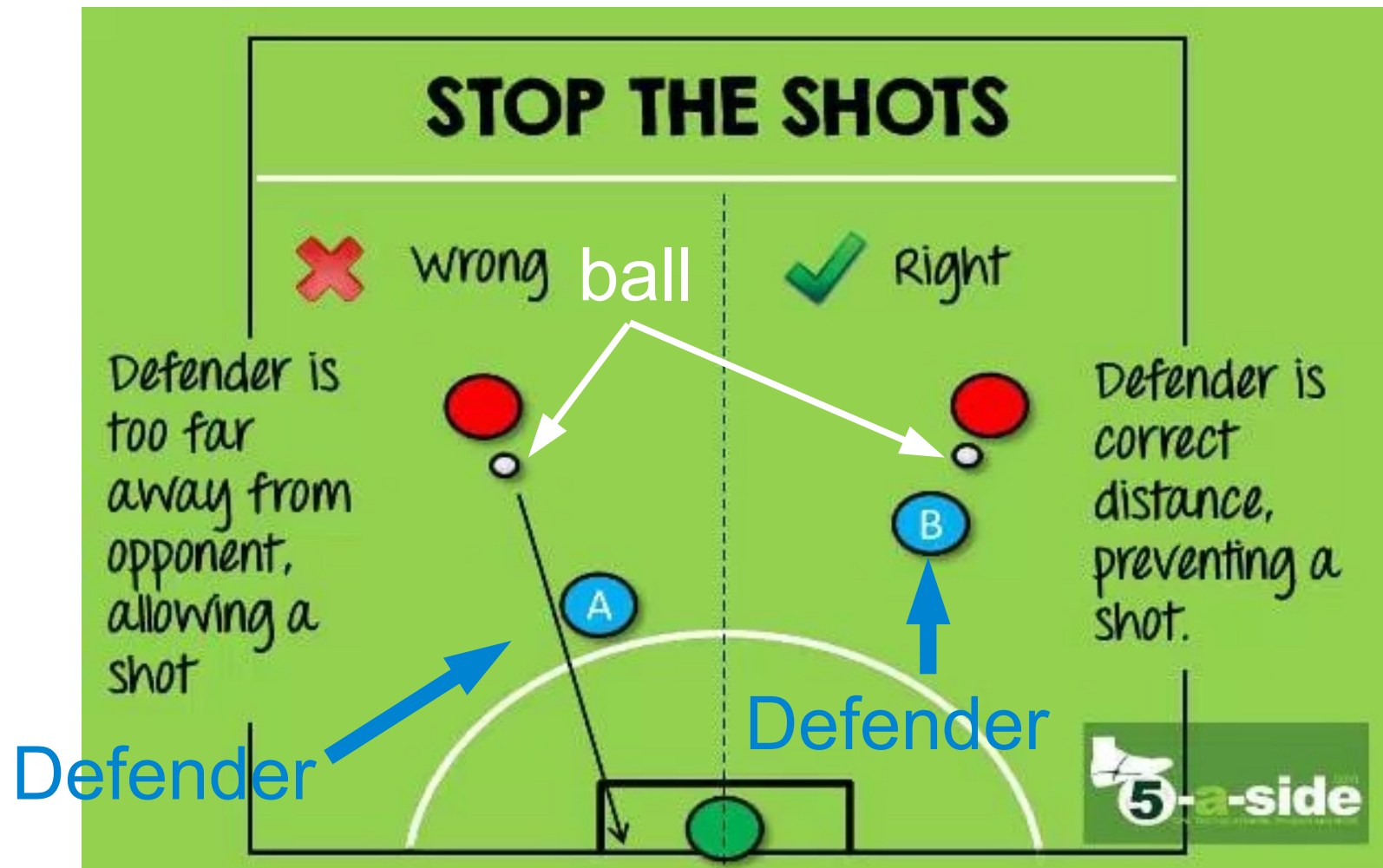


Project CMS bound

LL and MATHUSLA  
bring less



# Put a long lived detect close?



# Conclusions

- In terms of trigger coverage for EXO:
  - Rely heavy on Jets/*MET*/Leptons
  - No supremely dominant trigger (leptons a bit of help)
- Long-lived decays
  - Powerful tool to push down invisible branching ratio
  - Mainstay inside CMS is from the *MET* trigger
    - Can benefit from external detectors to trigger
    - Can potentially benefit from smarter triggers in CMS?
- Are there other possibilities?

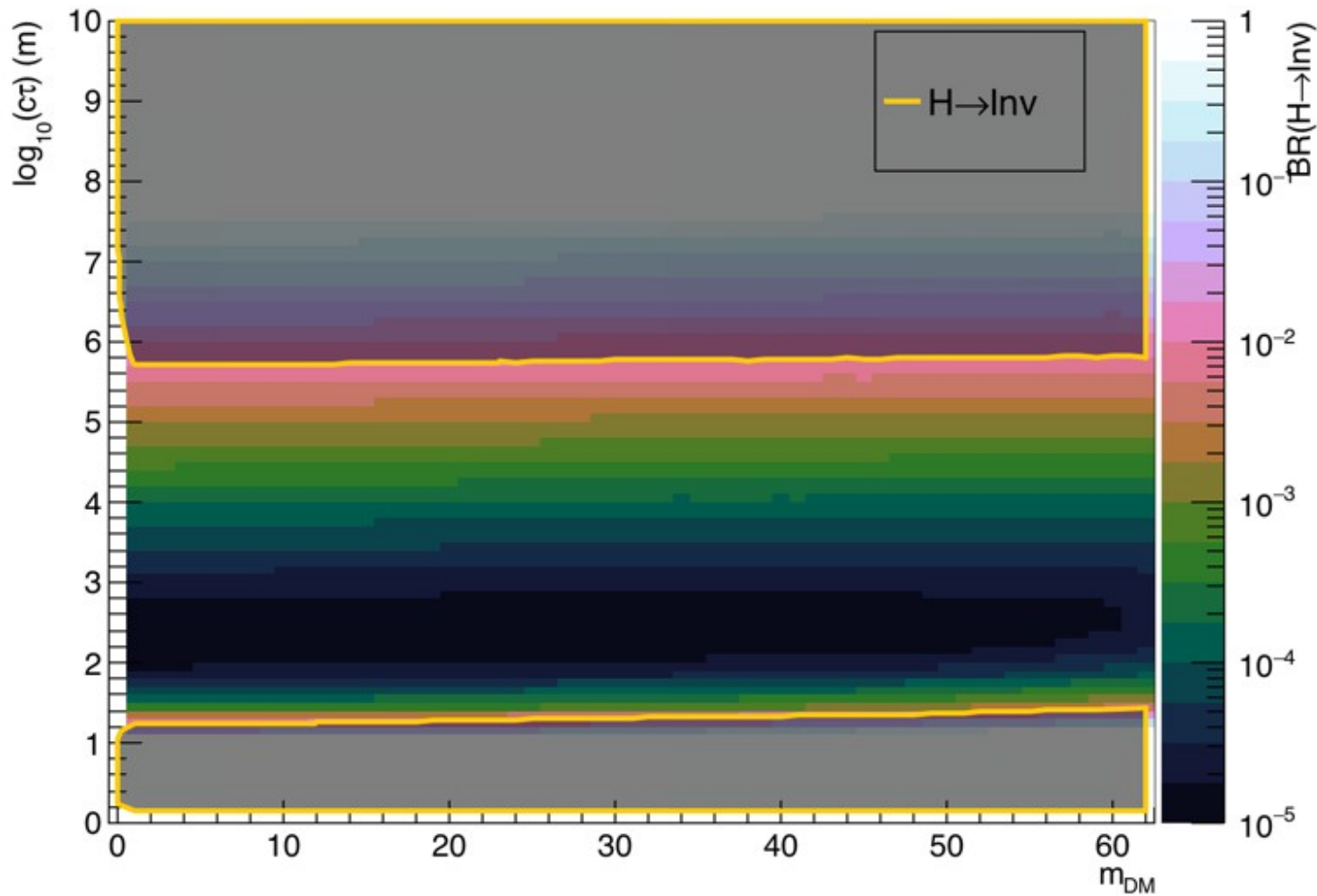
# Backup

# Make Lifetime Bounds

- We can decay the Higgs signal to DM
  - Sample the sphere in the Higgs rest-frame
    - Insert DM particles and boost back
    - Had to do this: was using an undecayed H sample
  - Compute probability of decay in the MATHUSLA vol.
    - Following verbatim the mathematica notebook
    - For the length sampling a finely binned histogram in  $\theta, \varphi$
- Again use a benchmark for  $3ab^{-1}$ 
  - Require 4 events for the exclusion
  - All of this can be switched, but seems pretty standard

# Lifetime Bounds w/MATHUSA

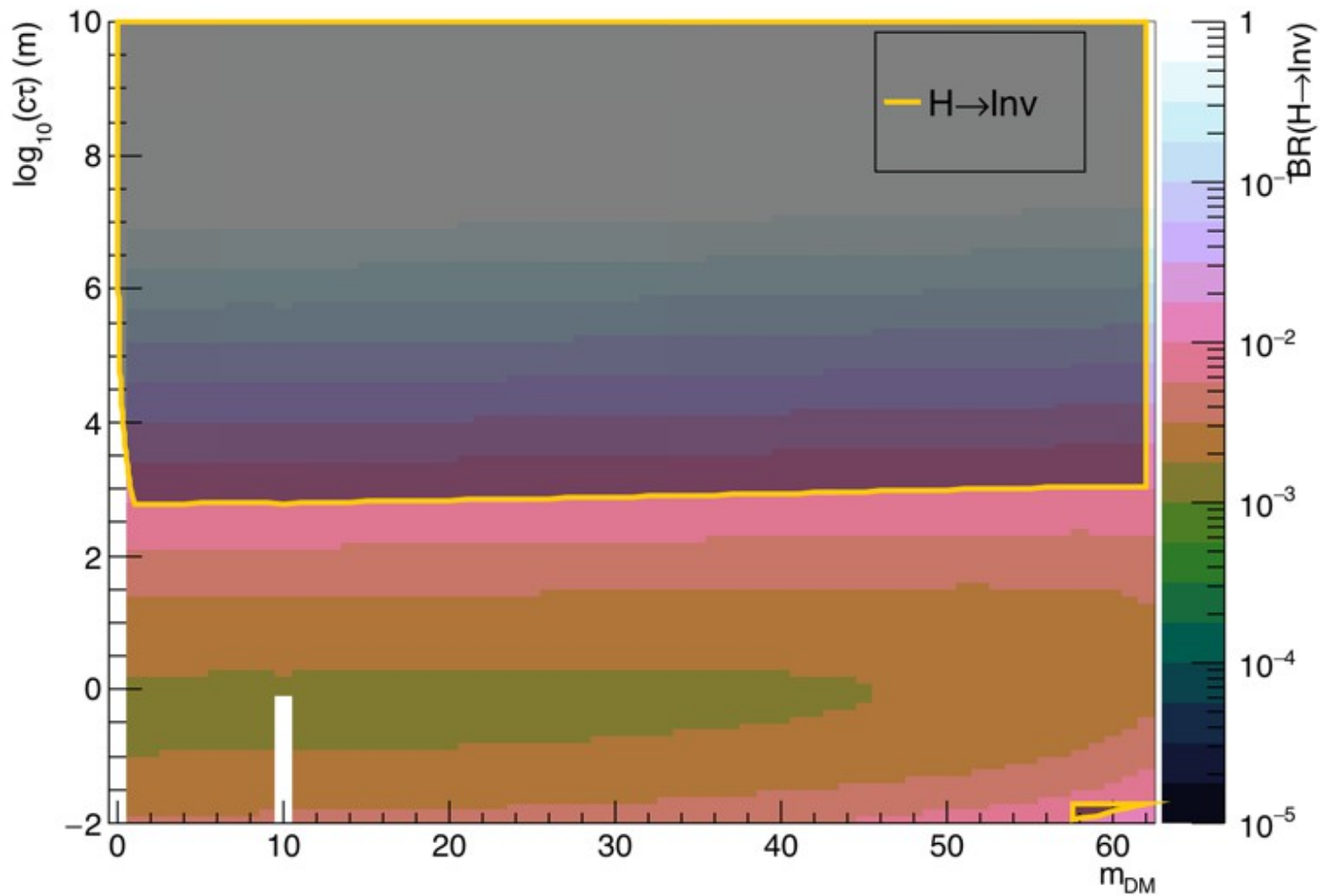
- Higgs Invisible in Mathusla detector



Just as a reminder

# Lifetime Bounds w/Tracker

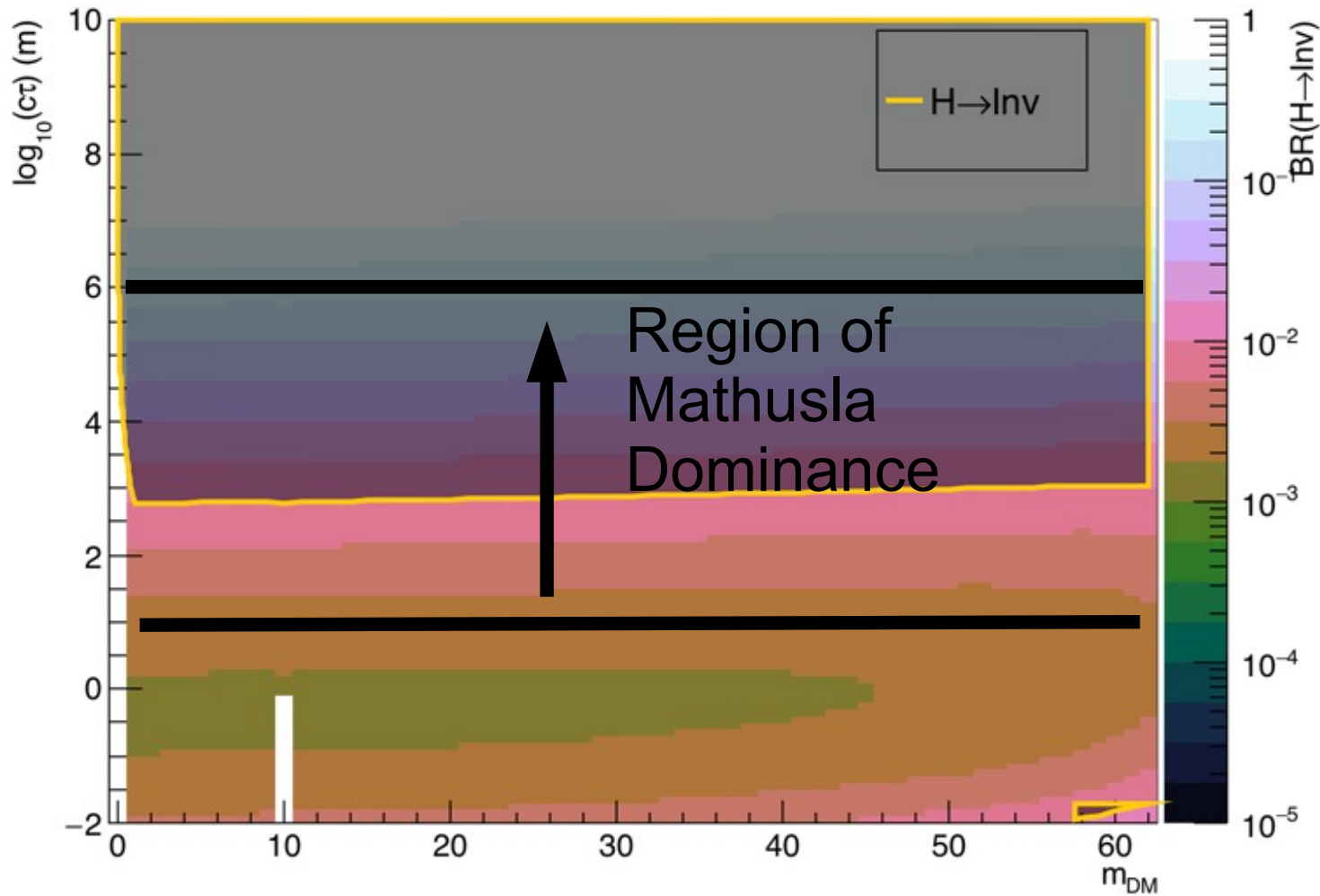
- Higgs Invisible in with a decay in tracker



Covers the lower range beyond  $H \rightarrow Inv$  ( $10^{-3}$ )

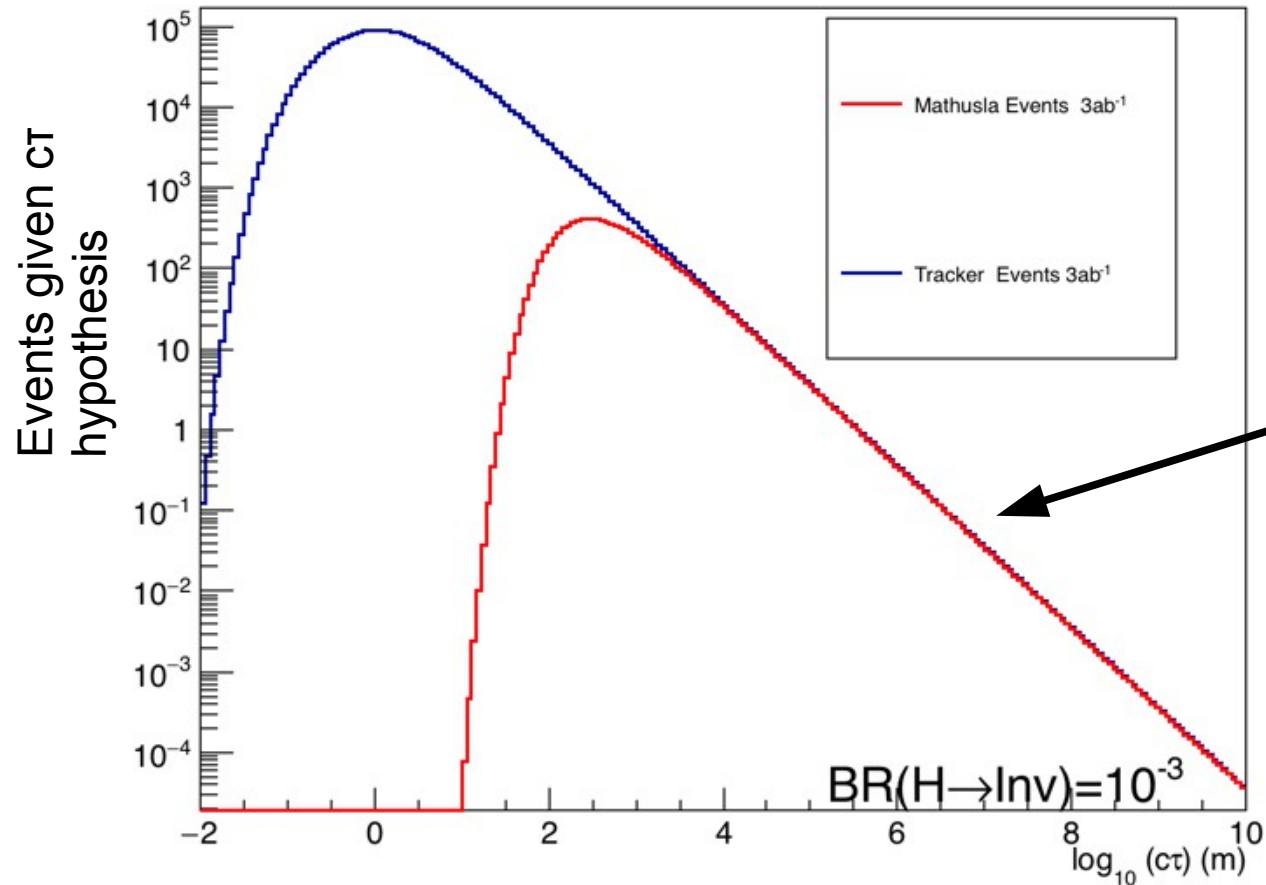
# Lifetime Bounds w/MATHUSA

- Higgs Invisible in with a decay in tracker



Is there a way to motivate lifetimes in this range?

# Visualizing the Acceptance



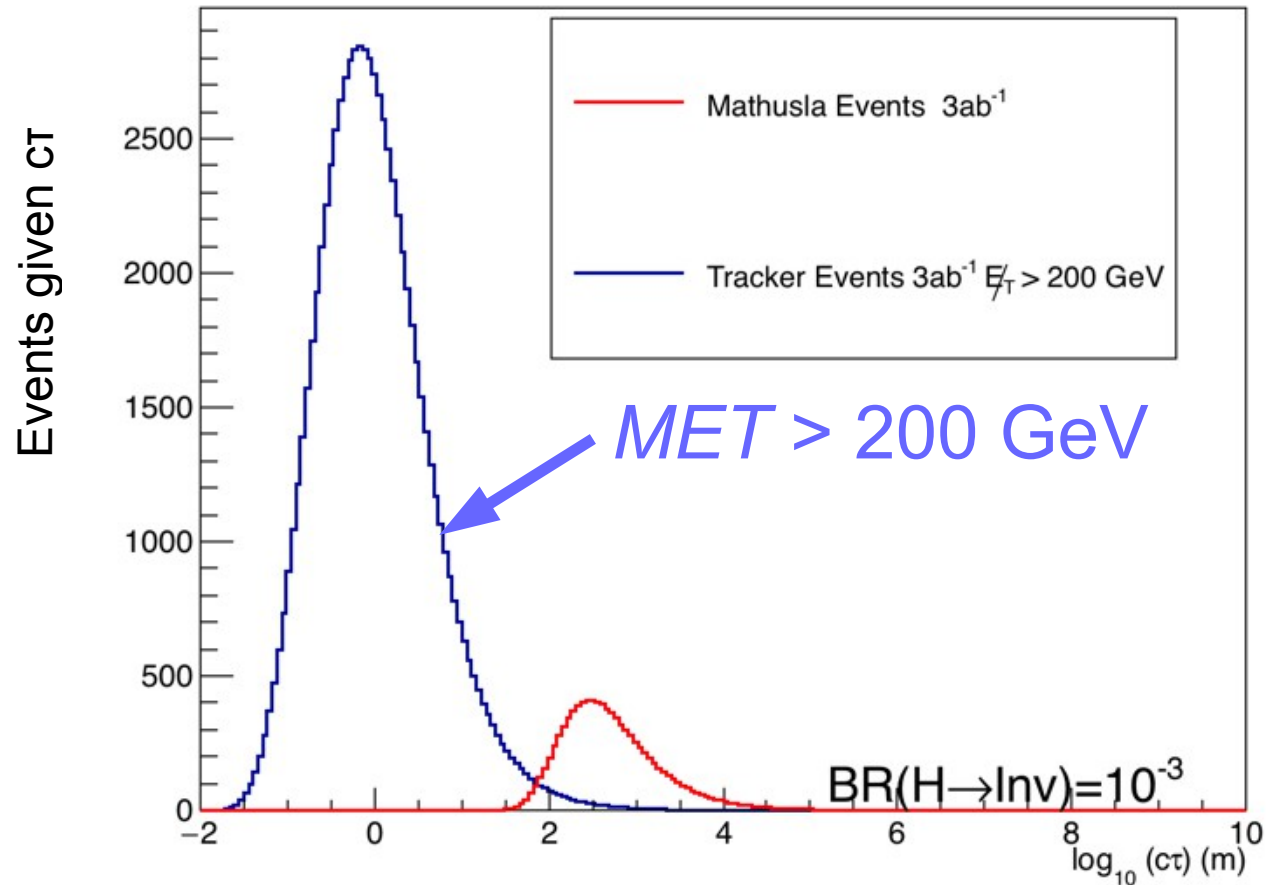
Blue and  
Red aligned

Were these  
dimensions  
purposeful?

- Large yields with  $3/ab$  and  $H \rightarrow \text{Inv}$ 
  - No cuts are applied here

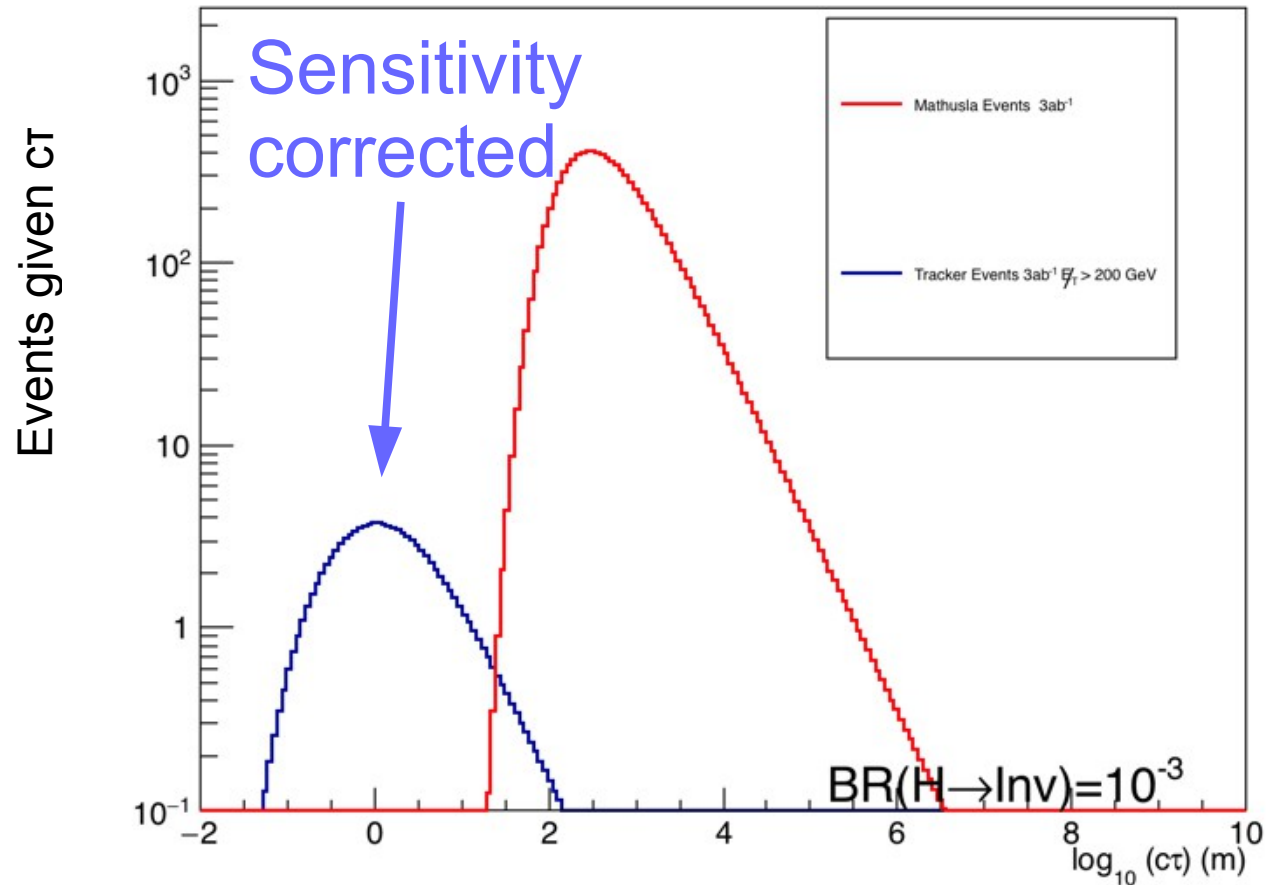


# Visualizing the Acceptance



- Adding a  $MET$  cut to the Tracker
  - Now yield are more comparable

# Visualizing the Acceptance



- Correcting for the sensitivity from the limit
  - ie accounting for bkg
  - Our equivalent yield on no background is 4 events

# Recap

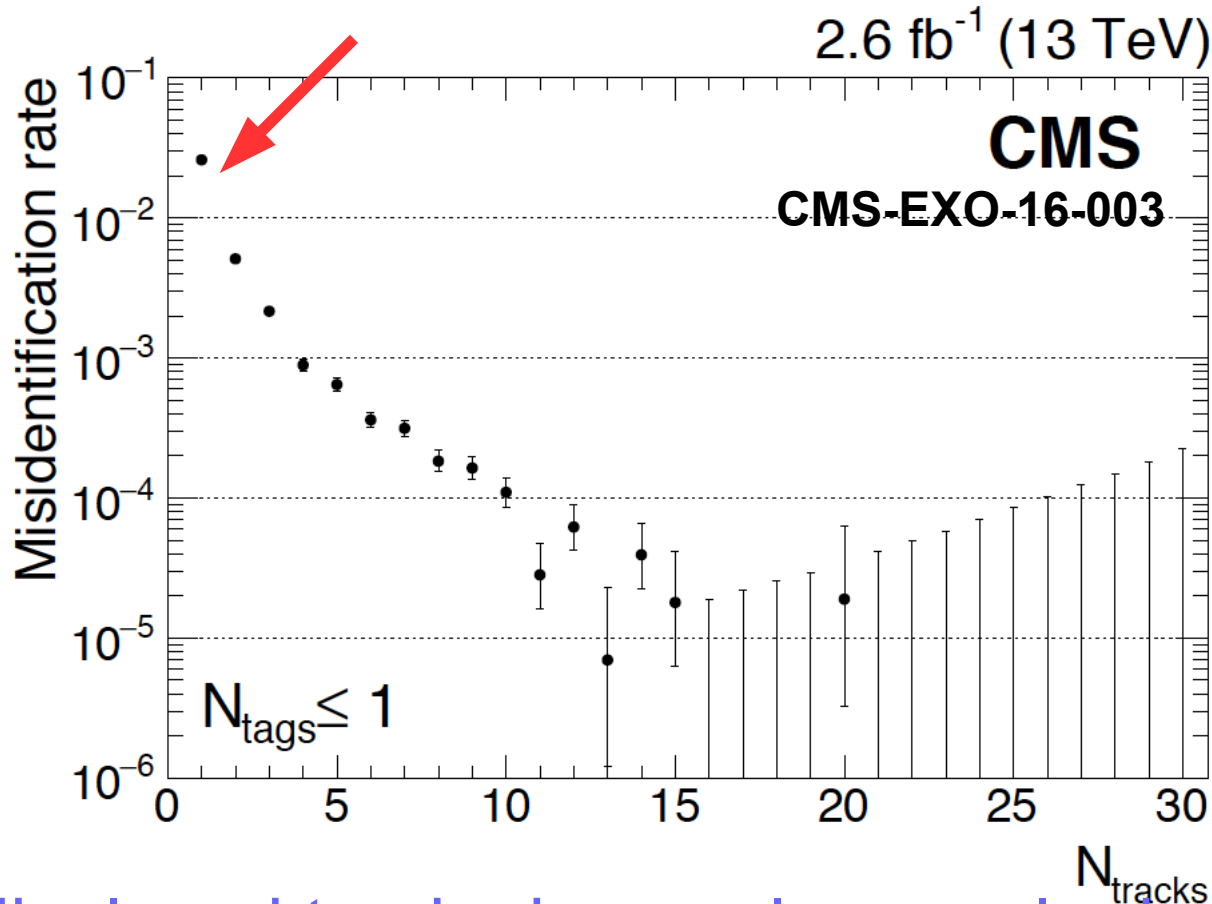
- Higgs invisible with best technology
  - $2\sigma$  exclusion Limit is 1%
  - Limit approaches FCC-ee (0.5%)
  - Using state of the art technology; **not unreasonable**
- Long-lived bounds:
  - Tracking will get to 0.1% (conservatively)
  - MATHUSLA will get to 0.001% (longer lifetimes)
- **Can probe 0.1% with lifetimes up to 10km in CT**

# Part 2: DM Simplified models

# Displaced Object Id

- Given a displaced object :
  - What is the likelihood of a fake id

For a jet w/ a single track jet rate its displaced is 2.5%



Prior is conservative so we take 1% with 100% eff (see later)

For referenced analysis integrated rate was 0.1-0.4% depending on prior

Rate of displaced track given a large calo deposit as prior

# Lepton Veto definition & co

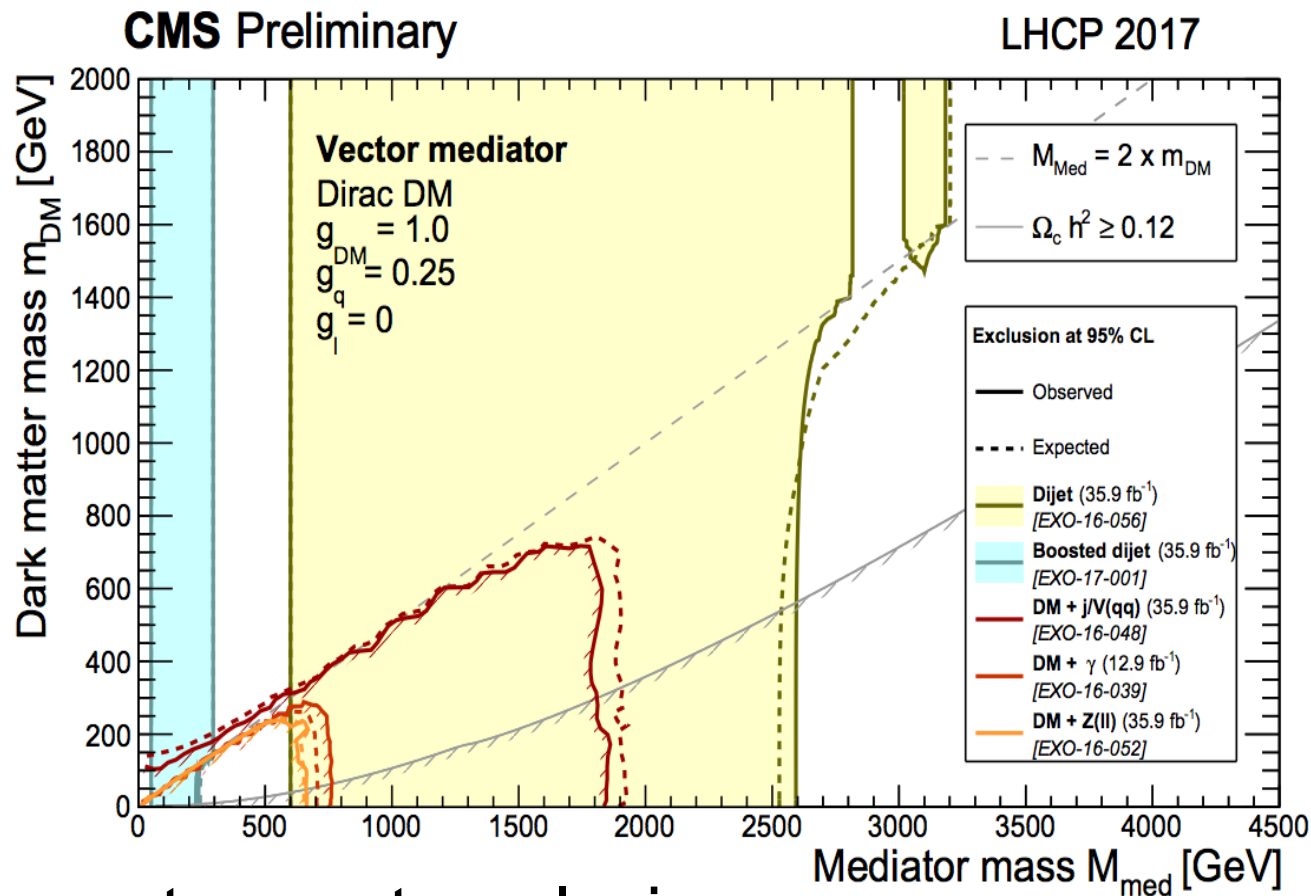
- **Select leptons up to  $|\eta| < 4.0$** 
  - Can do this due to pixel upgrade (tracking to  $|\eta| < 4.0$ )
- Assume a lepton efficiency of :
  - 85% for  $\tau$  leptons  $p_T > 20$  GeV (this is! our current eff)
  - 92% for e leptons  $p_T > 15$  GeV (roughly current)
  - 96% for  $\mu$  leptons  $p_T > 10$  GeV (worse than current)
- Events which pass lepton veto are mainly  $\tau$ 
  - Propagate efficiency unc. as a nuisance (5% for  $\tau$ )
    - Dominant uncertainty is from the  $\tau$  id efficiency
- **Invert lepton selection to make control regions**
  - Require single and dilepton control regions

# Strategy

- Run same fit on standard DM simplified models
  - Use same categorization ( $MET$  and  $m_{jj}$ )
    - Again split by  $m_{jj} > 400$  no reoptimization
  - Category splitting gives slightly better limits to monojet
- Do this for :
  - Spin 1 mediators : Vector and Axial-vector couplings
  - Spin 0 mediators : Scalar and Pseudoscalar
    - I skip the details, since I hope they are known
- Allow the DM candS from mediators to decay
  - Scan lifetime for the TK and M(athulsa) scenarios

# Vector Simplified models

- Current Mass Bounds



With most recent analysis :  
 Monojet reach is 2 TeV for spin 1  
 Dijet reach is 2.6 TeV for spin 1



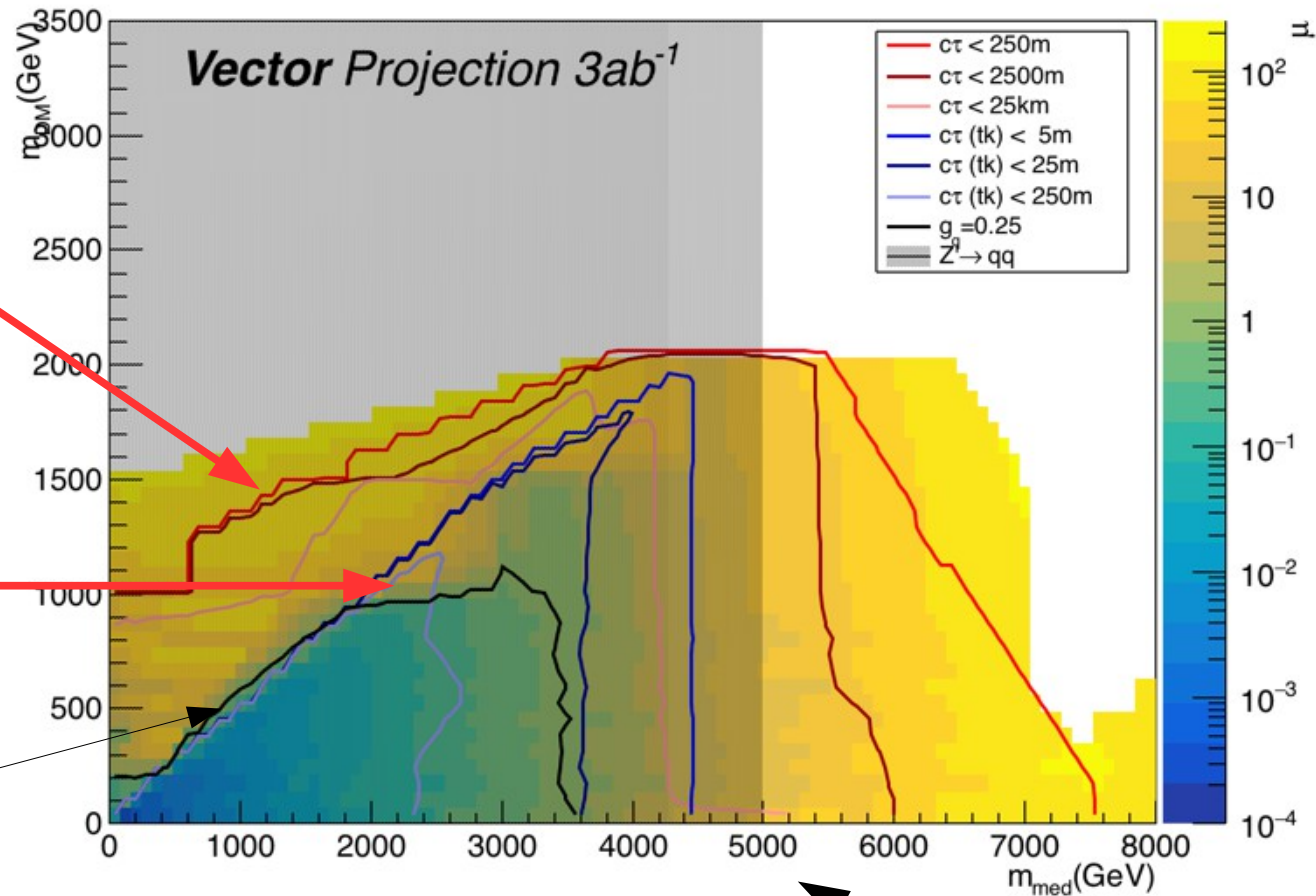
# Vector Simplified models

- Projected bounds to full luminosity

Interpolation is not smooth  
Can make it look nicer

Monojet bound

Dijet bounds beat tracker decays  
Mathusla wins at very high mass



Expect mass reach for dijet to go to 5 TeV

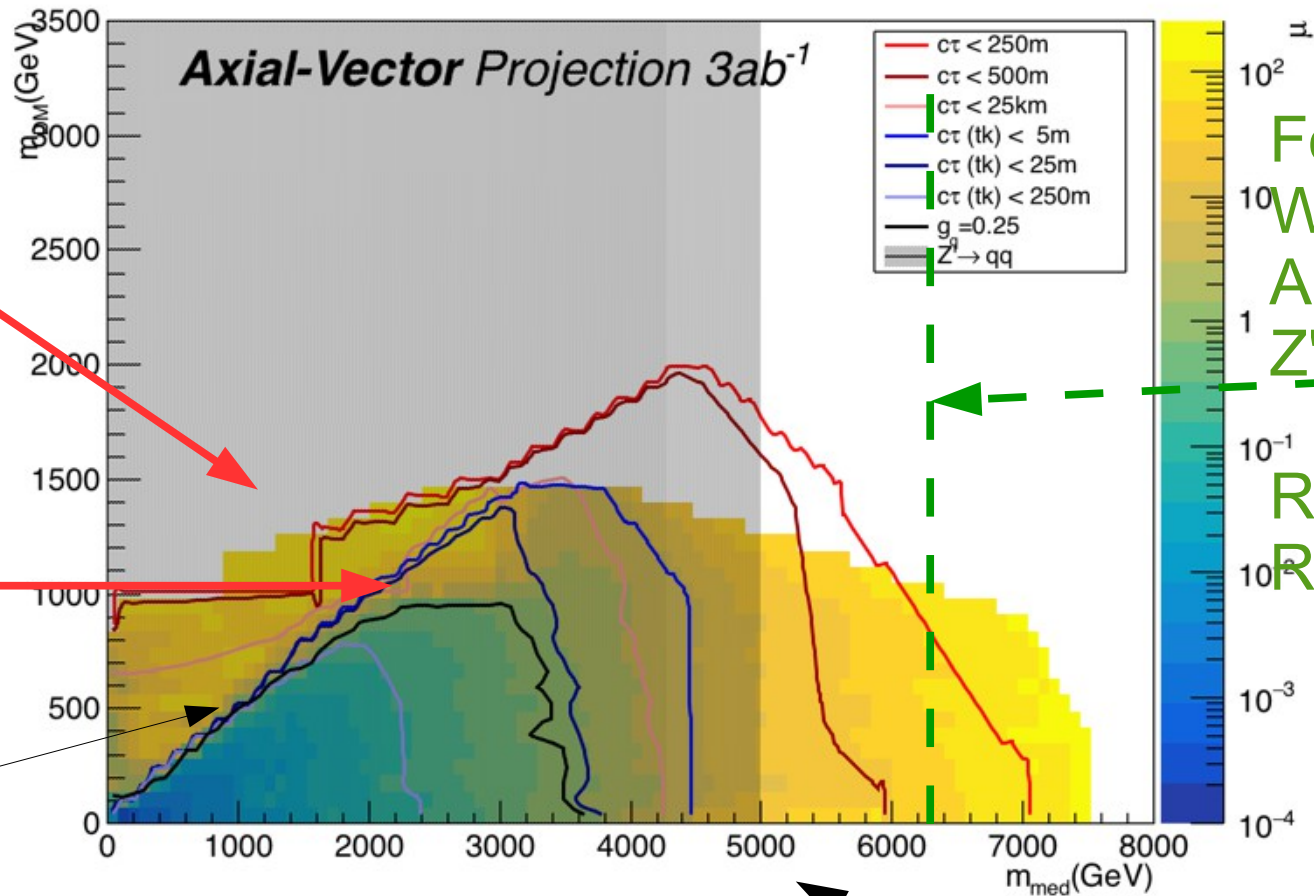
# Axial-Vector Simplified models

- Now with an Axial vector mediator

Interpolation is not smooth  
Can make it look nicer

Monojet bound

Dijet bounds beat tracker decays  
Mathusla wins at very high mass



For Axial-vector  
Would need to  
Add high mass  
 $Z'$  to leptons

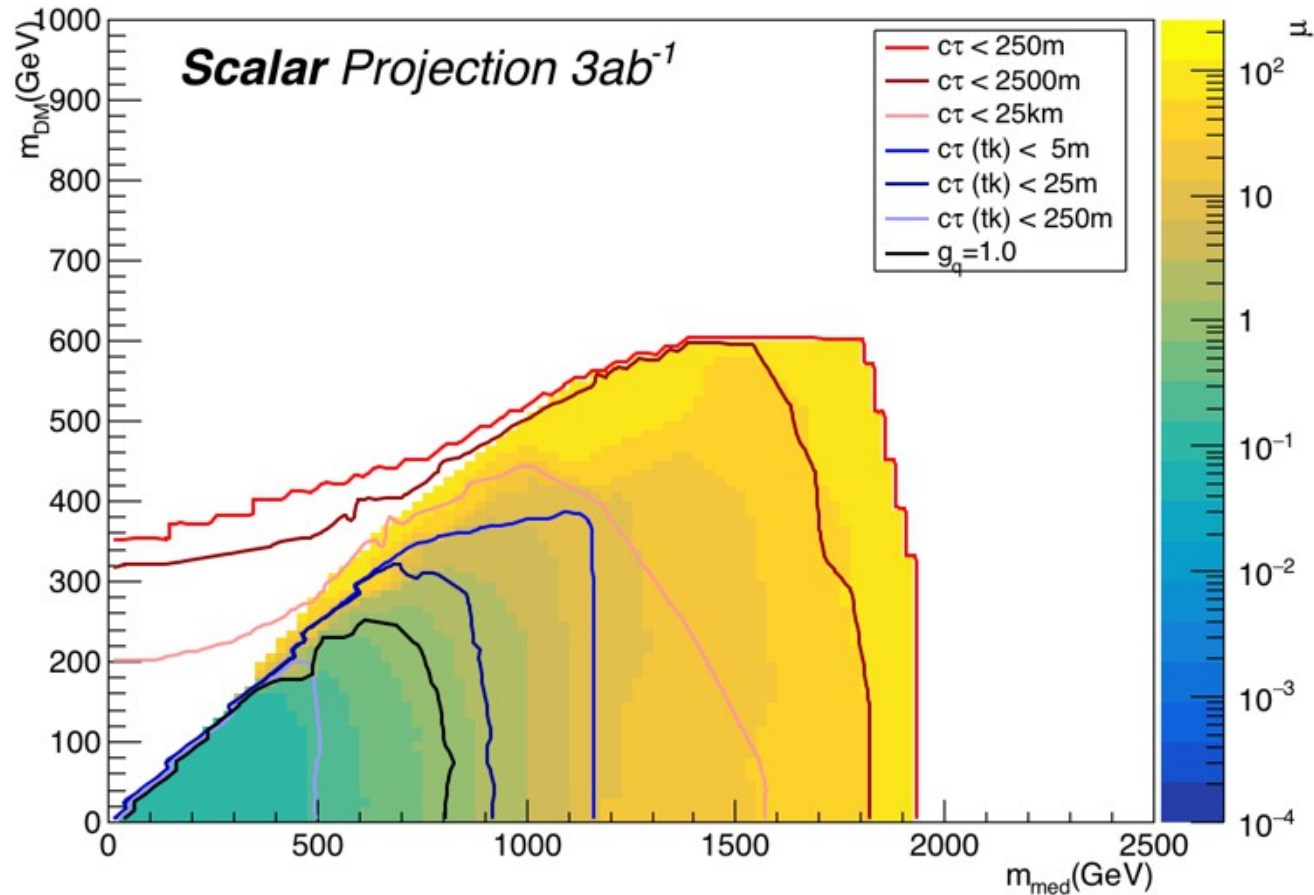
Reach is  
Roughly 7 TeV

Expect mass reach for  
dijet to go to 5 TeV

# Scalar Simplified models

- Scalar Bounds

- This is a very nice plot (note Z axis always  $\mu$  for monojet)



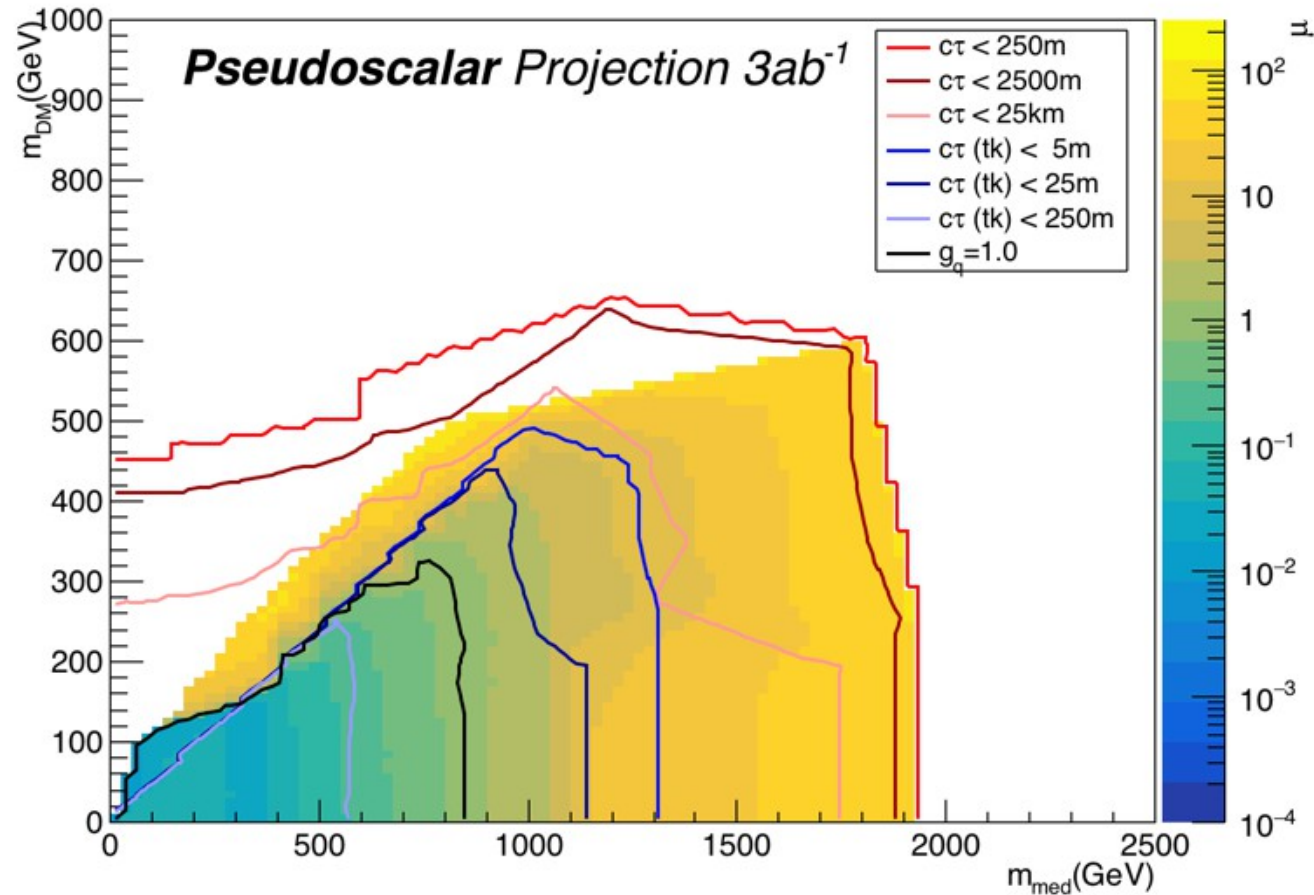
In the low mass region

gains from LL searches are more important

# Pseudoscalar Simplified models

- Pseudoscalar bounds
  - Slightly larger than scalar bounds

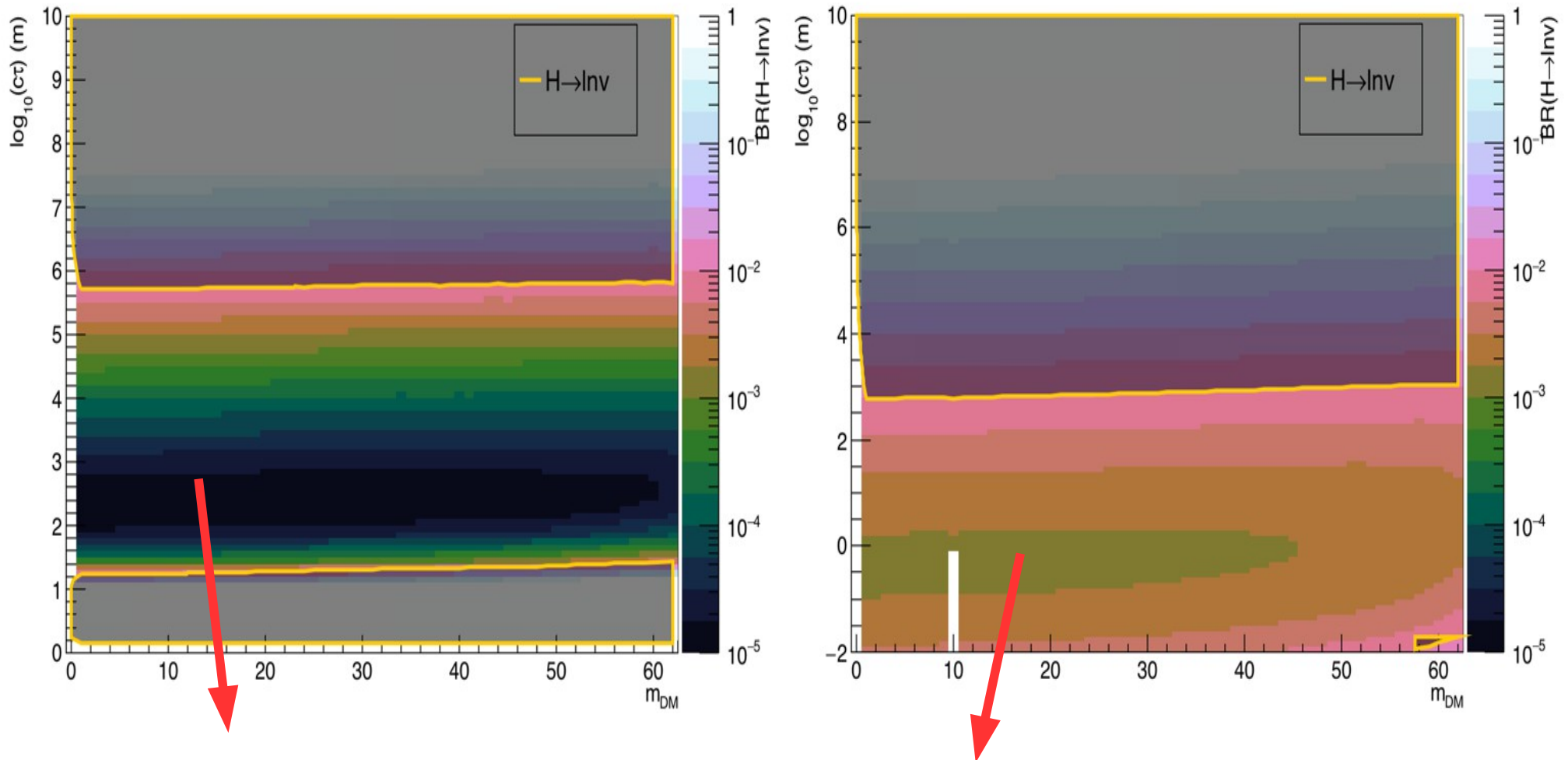
Some more  
interpolation issues  
can be smoothed



Both spin 0 mediators give similar bounds to Higgs invisible bounds, now projected in mass

# Translating to Coupling

- We can also translate our bounds to coupling



Take the best bound from each displaced experiment  
 Translate this to a limit on the coupling

# Translating to Coupling

- Coupling translation

For  $\mu < 1$  limit scales with  $g_q^2$

For  $g_q < 0.25$  width is dominated by  $gq$

This means :

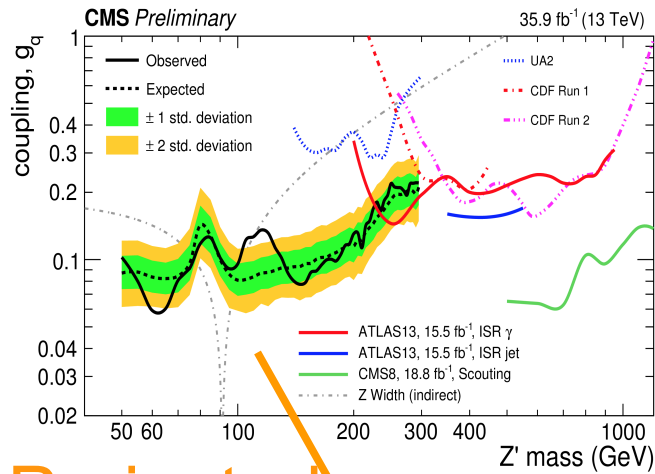
$$\mu = (g_q / 0.25)^2$$

And we solve for  $g_q$

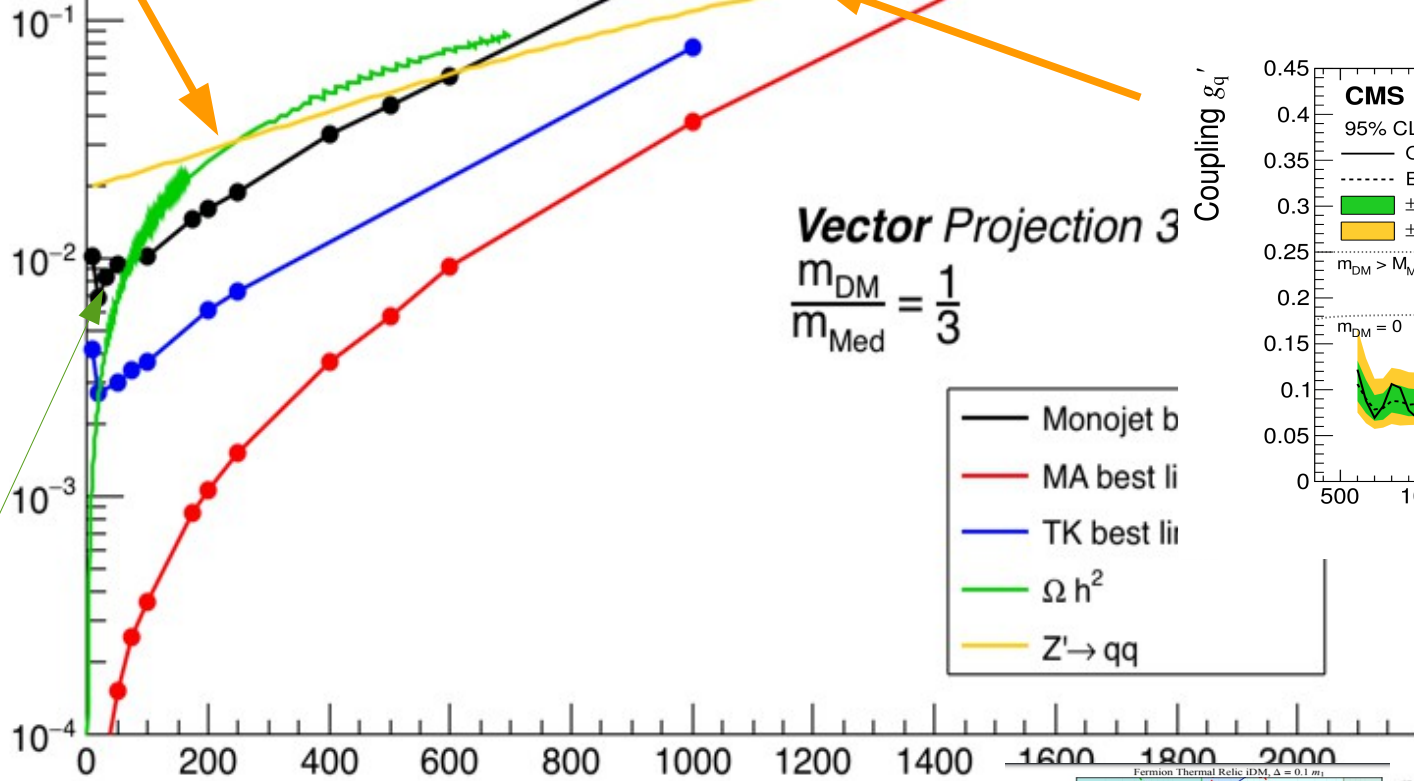
Also applies for  $g_q = 1$  in a scalar model

On top add relic bound given just simplified model  
Solve for minimum coupling that doesn't over produce

# Vector Coupling Bound



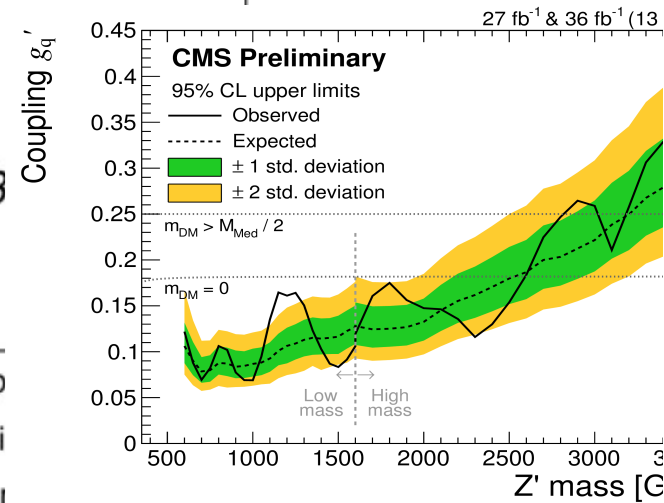
Projected



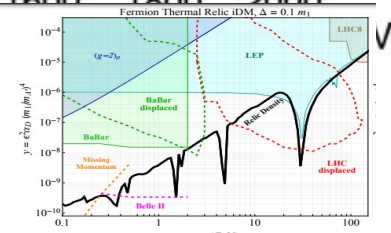
Vector Projection 3

$$\frac{m_{DM}}{m_{Med}} = \frac{1}{3}$$

Projected



Relic bound below we overclose  
(see <https://arxiv.org/abs/1508.03050>)



# Simplified Model Summary

- For spin 1 mediators:
  - Visible bounds from dijet are competitive with LL
  - For low masses LL can reduce coupling sensitivity
  - Probe relic down to 10 GeV in mediator mass
  - Expect to cover spin 1 almost completely
- For spin 0 mediators:
  - LL adds a very powerful component
  - LL opens interesting region for heavy mediators
- It might be good to tie these to real models?