DARK SHOW IRS: MOTIVATION, SIMULATION, AND TRIGGERING

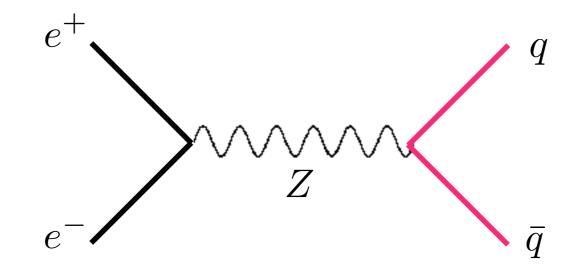
Carleton

DANIEL STOLARSKI

P. Schwaller

PCTS "Triggering on New Physics at the HL-LHC" Jan 16, 2018

OCD SHOWERING



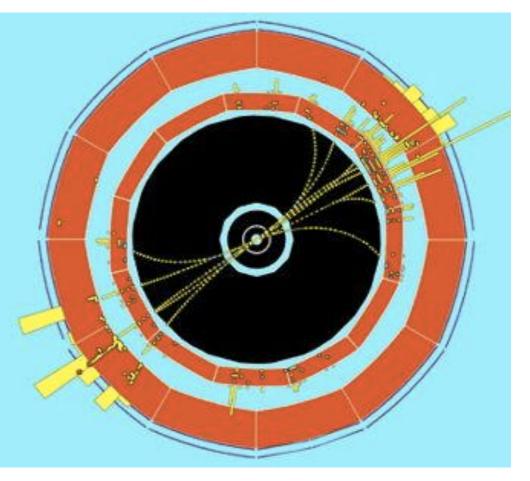
Quarks produced at a lepton collider will shower in a calculable way in QCD.

$$\frac{\mathrm{d}}{\mathrm{d}\log(t/\mu^2)} \int_{x}^{f_q(x,t)} \int_{x}^{q} = \int_{x}^{1} \frac{\mathrm{d}z}{z} \frac{\alpha_s}{2\pi} \int_{f_q(x/z,t)}^{P_{qq}(z)} \int_{x}^{q} + \int_{x}^{1} \frac{\mathrm{d}z}{z} \frac{\alpha_s}{2\pi} \int_{f_g(x/z,t)}^{P_{gq}(z)} \int_{x}^{q} \int_{x}^{1} \frac{\mathrm{d}z}{z} \frac{\alpha_s}{2\pi} \int_{f_q(x/z,t)}^{P_{qq}(z)} \int_{x}^{q} + \int_{x}^{1} \frac{\mathrm{d}z}{z} \frac{\alpha_s}{2\pi} \int_{f_g(x/z,t)}^{P_{gq}(z)} \int_{x}^{q} \int_{x}^{1} \frac{\mathrm{d}z}{z} \frac{\alpha_s}{2\pi} \int_{f_q(x/z,t)}^{P_{qq}(z)} \int_{x}^{q} + \int_{x}^{1} \frac{\mathrm{d}z}{z} \frac{\alpha_s}{2\pi} \int_{f_g(x/z,t)}^{P_{gq}(z)} \int_{x}^{q} \int_{x}^{1} \frac{\mathrm{d}z}{z} \frac{\alpha_s}{2\pi} \int_{x}^{1} \frac{\mathrm{d}$$

OCD SHOWERING

Cartoon picture Hoche, 1411.4085.

ALEPH event

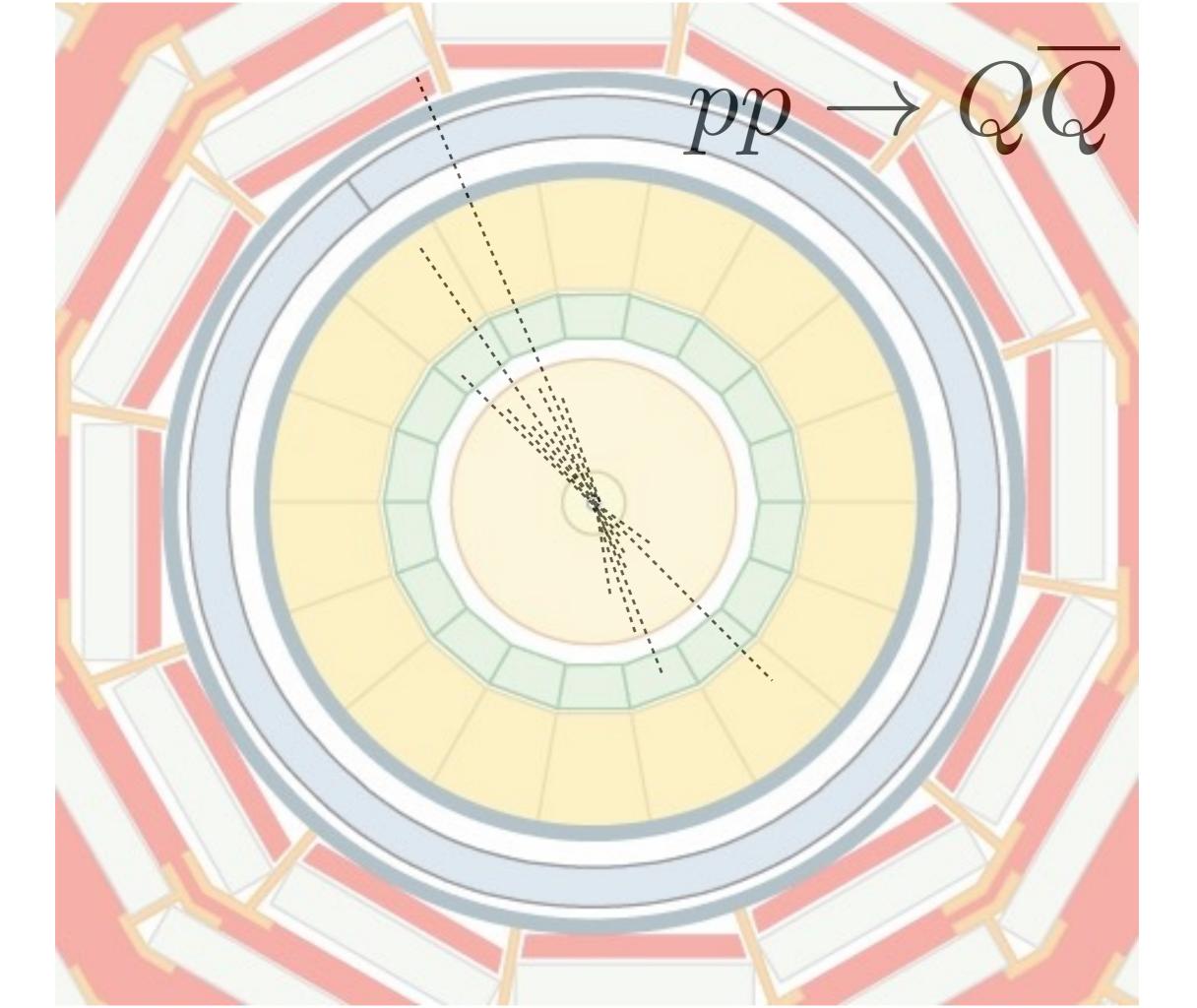


DARK SHOWERS

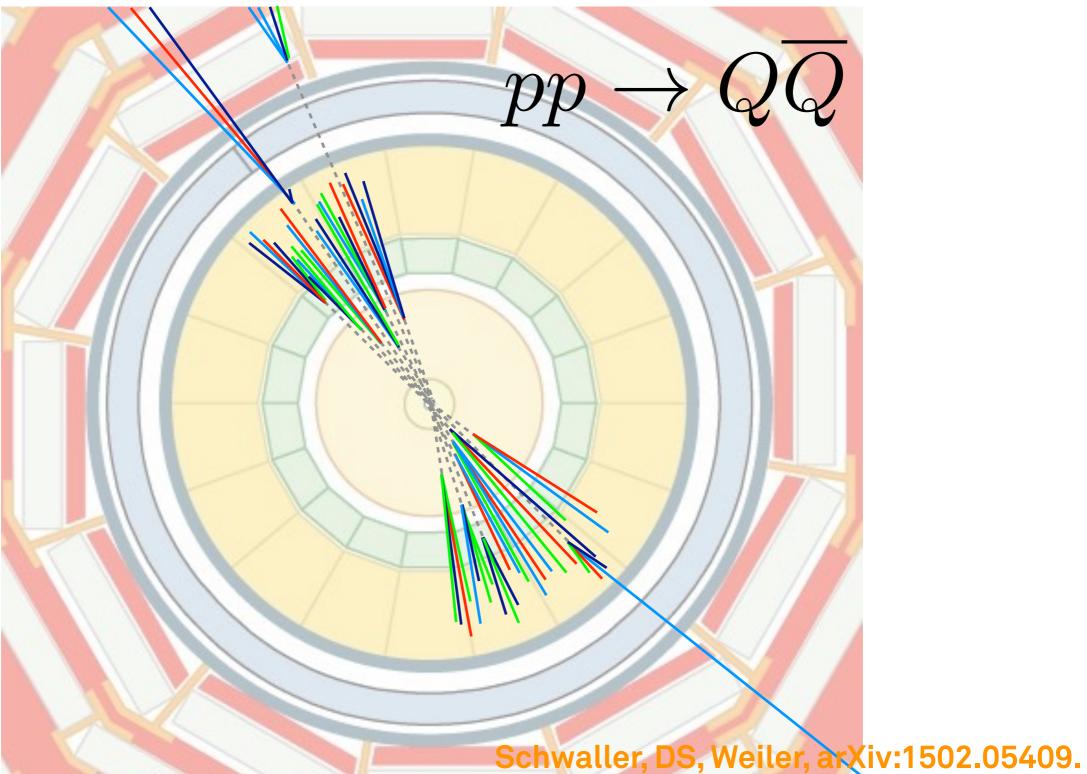
A new confining sector (possibly QCD-like) that:

- Has a relatively low (~ GeV) confining scale.
- All SM particles are neutral under new force.
- All light particles that feel the force are neutral under the SM.

If there is a heavy mediator, then have large multiplicity of BSM particles at LHC. Strassler and Zurek, arXiv:0604261.



EMERGING JETS



DIFFERENT DARK SHOWERS

Emerging Jets: QCD like, displaced vertices. Schwaller, DS, Weiler, arXiv:1502.05409.

Semivisible jets: missing energy in the jet. Cohen, Lisanti, Lou, arXiv:1503.00009. Cohen, et. al. arXiv:1707.05326.

Soft bombs or SUEP (Soft Unclustered Energy Patterns): spherical distribution of particles. Knapen, Griso, Papucci, Robinson, arXiv:1612.00850.

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???

See also talk by D'Agnolo yesterday.

MOTIVATION

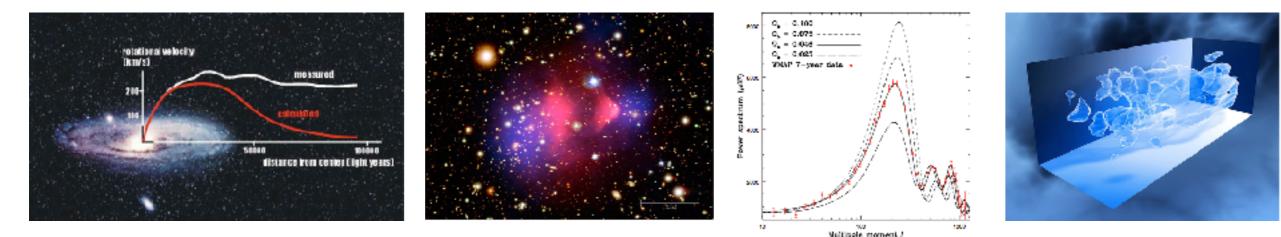
MOTIVATION



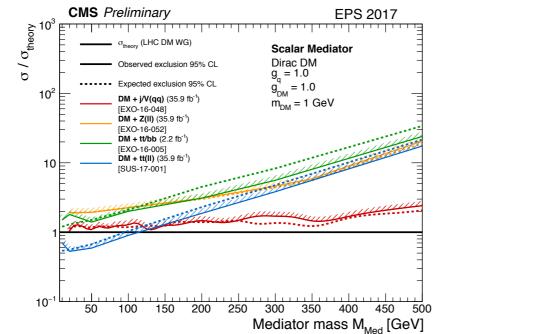
H. Murayama

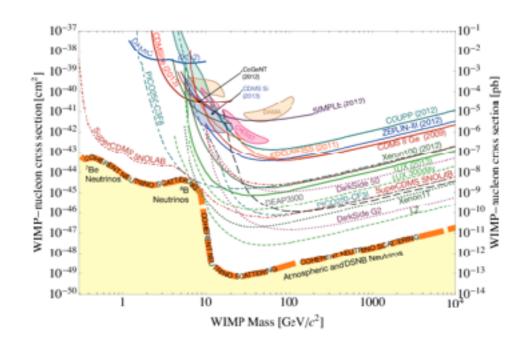
DARK MATTER

We have seen dark matter in the sky.



But not in the lab.





$\Omega_{DM} \simeq 5\Omega_B$

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$\Omega_{DM} = m_{DM} n_{DM} \qquad \qquad \Omega_B = m_p n_B$

$\Omega_{DM} \simeq 5\Omega_B$

Controlled by complicated (known) QCD dynamics $\Omega_B = m_p n_B$

$\Omega_{DM} = m_{DM} n_{DM}$

$\Omega_{DM} \simeq 5\Omega_B$

Controlled by complicated (known) QCD dynamics

 $\Omega_B = \overset{\bigstar}{m_p n_B} \overset{n_B}{\checkmark}$

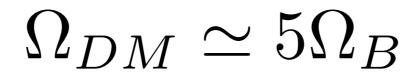
$\Omega_{DM} = m_{DM} n_{DM}$

Unknown dynamics of baryogenesis

$\Omega_{DM} \simeq 5\Omega_B$

Controlled by complicated (known) QCD dynamics $\Omega_B = \dot{m}_p n_B$ $\Omega_{DM} = m_{DM} n_{DM}$ Nussinov, '85. Kaplan '92. Unknown dynamics Kaplan, Luty, Zurek, '09. of baryogenesis Bai and Schwaller, '13.

. . .

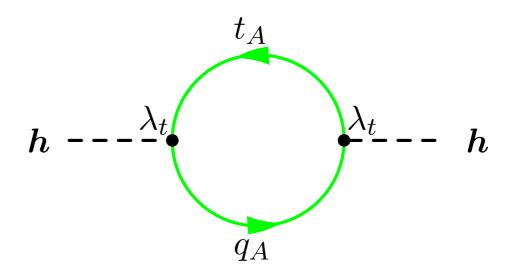


QCD like Controlled by complicated (known) QCD dynamics $\Omega_B = n_p n_B$ $\Omega_{DM} = m_{DM} n_{DM}$ Nussinov, '85. Kaplan '92. Unknown dynamics Kaplan, Luty, Zurek, '09. of baryogenesis Bai and Schwaller, '13.

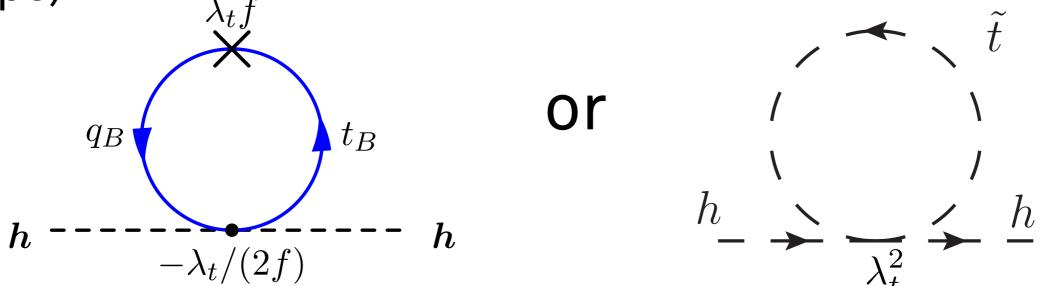
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TWIN HIGGS/FOLDED SUSY

Gauge hierarchy problem:



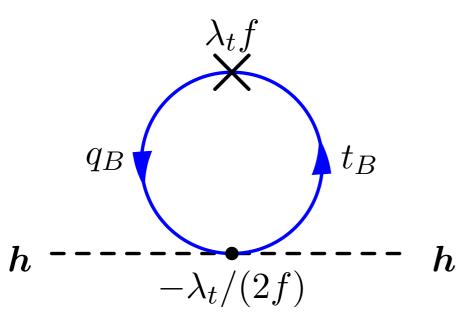
Solved in composite Higgs (SUSY) with top-partners (stops) $\lambda_t f$

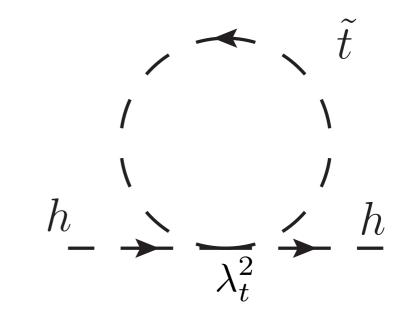


Do these partners need to be coloured?

TWIN HIGGS/FOLDED SUSY

No! But still need factor of 3.





Chacko, Goh, Harnik, hep-ph/0506256. Burdman, Chacko, Goh, Harnik, hep-ph/0609152.

Most models have twin colour which confines around GeV scale (or slightly higher).

SINULATION

SIMULATION

Dark QCD is in Pythia. Can vary number of dark colours and flavours as well as mass scales. Carloni, Sjorstrand, 2010. Carloni, Rathsman, Sjorstrand, 2011. Further updates in versions > 8.2.

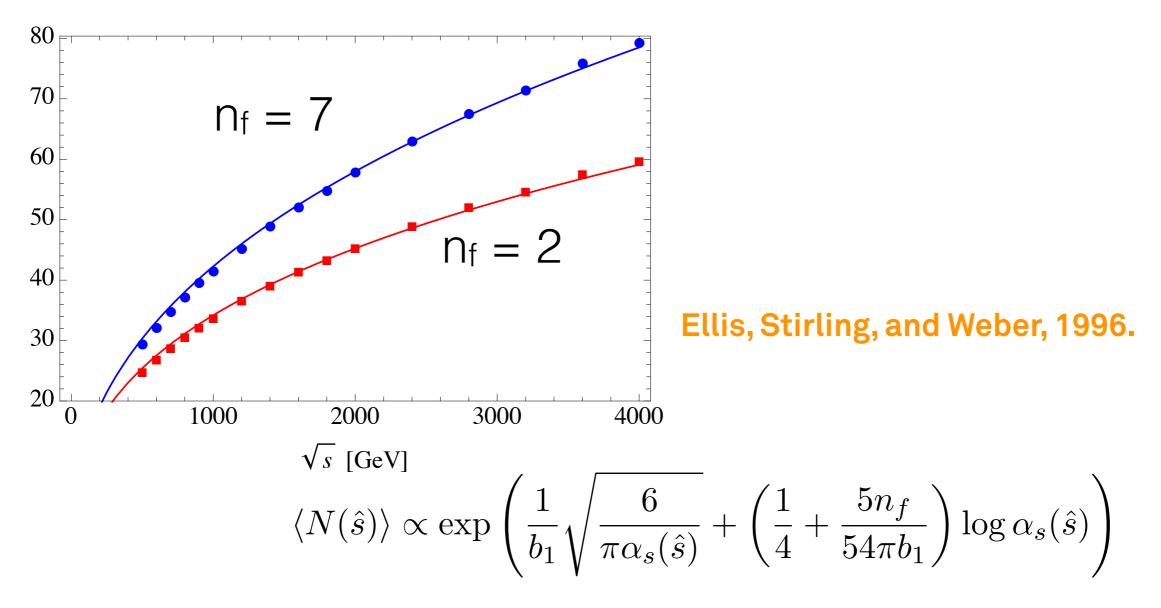
Two mediators also implemented.

• Scalar
$$pp \to \Phi \Phi^{\dagger} \to \bar{q} Q_d \ \overline{Q}_d q$$

• Vector $pp \to Z_d \to Q_d \; Q_d$

VAL DATION

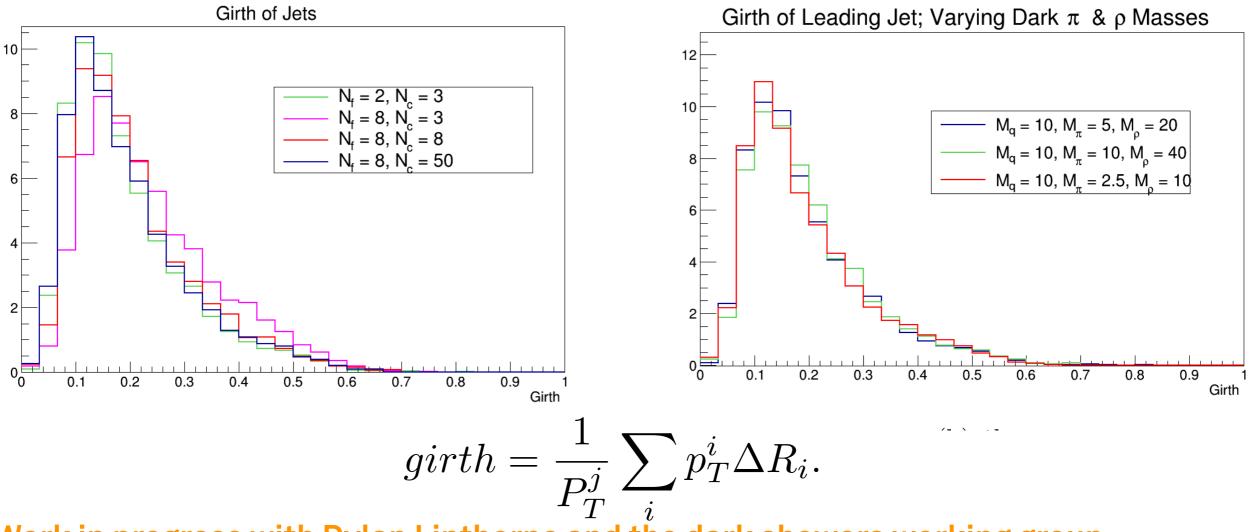
Check to see if simulation makes sense by 4000 looking at average particle/multiplicity.



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DARK JET SHAPES

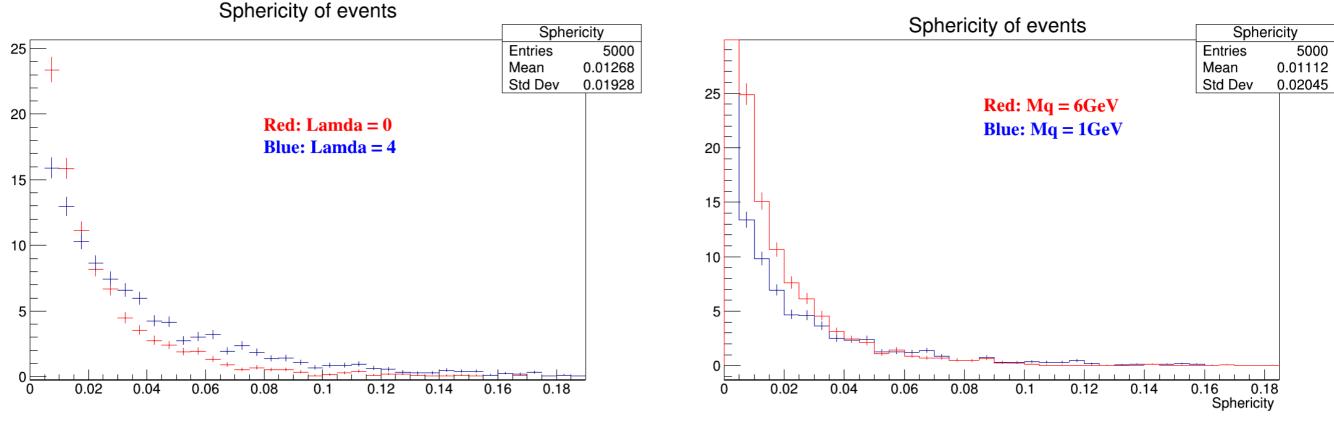
Dark shapes do not change much varying parameters in Pythia.



Work in progress with Dylan Linthorne and the dark showers working group.

DARK JET SHAPES

Dark shapes do not change much varying parameters in Pythia.



Sphericity

Work in progress with Dylan Linthorne and the dark showers working group.

SOFT BOMBS

Soft bombs have approximately isotropic particle distribution.

Knapen, Griso, Papucci, Robinson, arXiv:1612.00850. Possibly more in talk by Simon next.

Can be simulated with a thermal distribution.

$\frac{dN}{d^3\mathbf{p}} \sim \exp\left\{-\sqrt{\mathbf{p}^2 + m^2}/T\right\}$

INTERPOLATION?

SUEPs^{*} to Jets: Parameterizing the Theory

Cari Cesarotti

Harvard University

In Collaboration with Matt Reece, Matt Strassler LLP Trieste, October 20 2017

*Soft unclustered energy patterns

INTERPOLATION?

Seems that you can simulate the interpolation using an extra dimensional model and the AdS/CFT correspondence.

See talk by Cesarotti at Trieste workshop for more details.

https://indico.cern.ch/event/649760/

TRGGERING

EASY WAYS TO TRIGGER

Easy ways to trigger on dark showers:

- Lots of energy: use H_T or multi jet trigger.
- Missing energy.

Will almost always get one of these if mediator mass is large (SUEP is an interesting exception).

• Isolated leptons (difficult if multiplicity is large).

SINPLE ENTITIENT



Can do simple trigger study using known trigger thresholds.

Pythia hidden valley.

Current thresholds, not high-lumi.

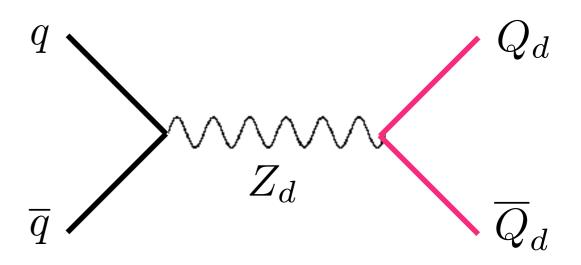
ATLAS not CMS. 🛃



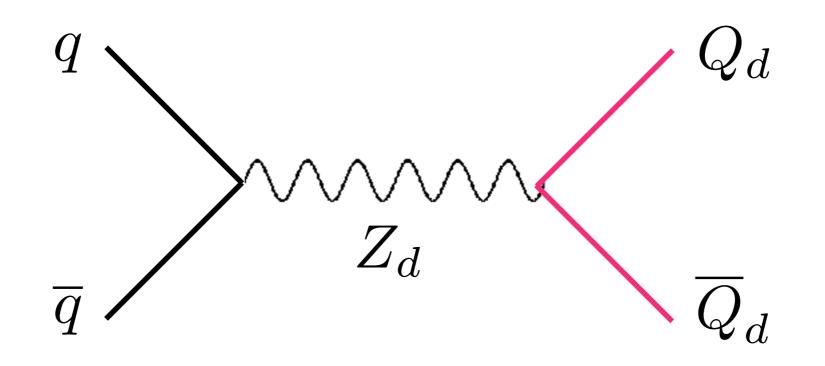
The ATLAS Collaboration

During 2015 the ATLAS experiment recorded 3.8 fb^{-1} of proton–proton collision data at a centre-of-mass energy of 13 TeV. The ATLAS trigger system is a crucial component of the experiment, responsible for selecting events of interest at a recording rate of approximately 1 kHz from up to 40 MHz of collisions. This paper presents a short overview of the changes to the trigger and data acquisition systems during the first long shutdown of the LHC and shows the performance of the trigger system and its components based on the 2015 proton–proton collision data.

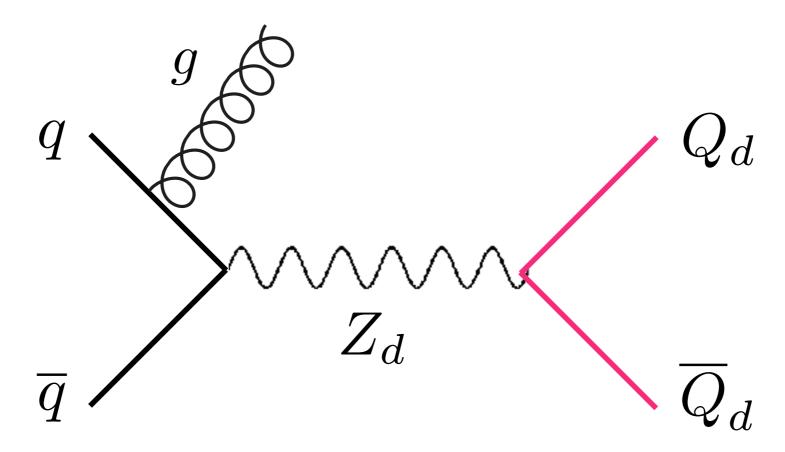
arXiv:1611.09661



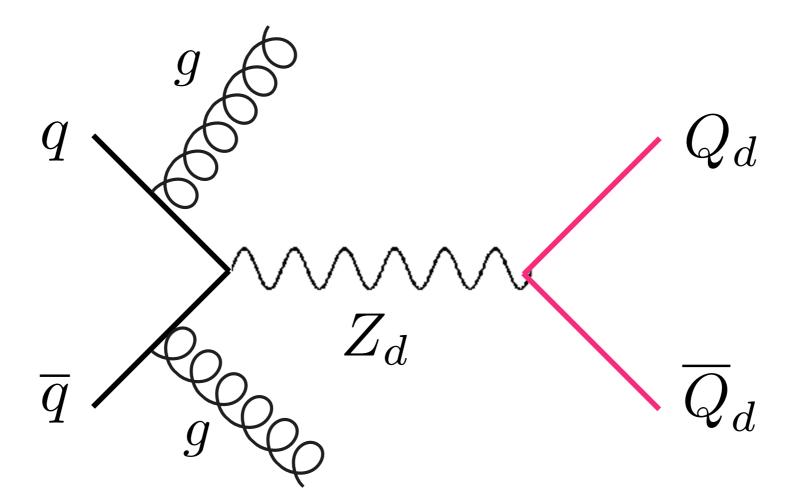
For any dark shower, ISR is always there.



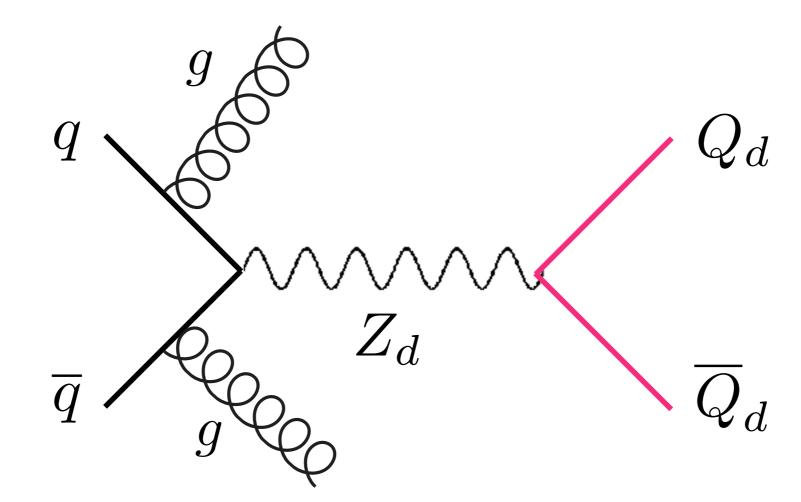
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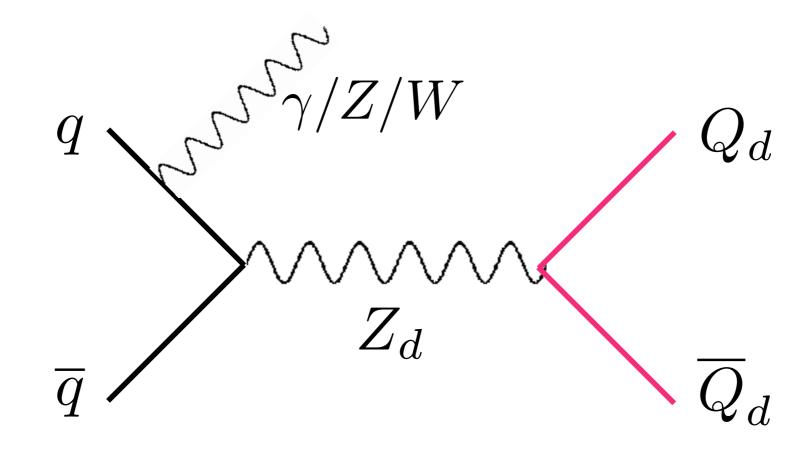


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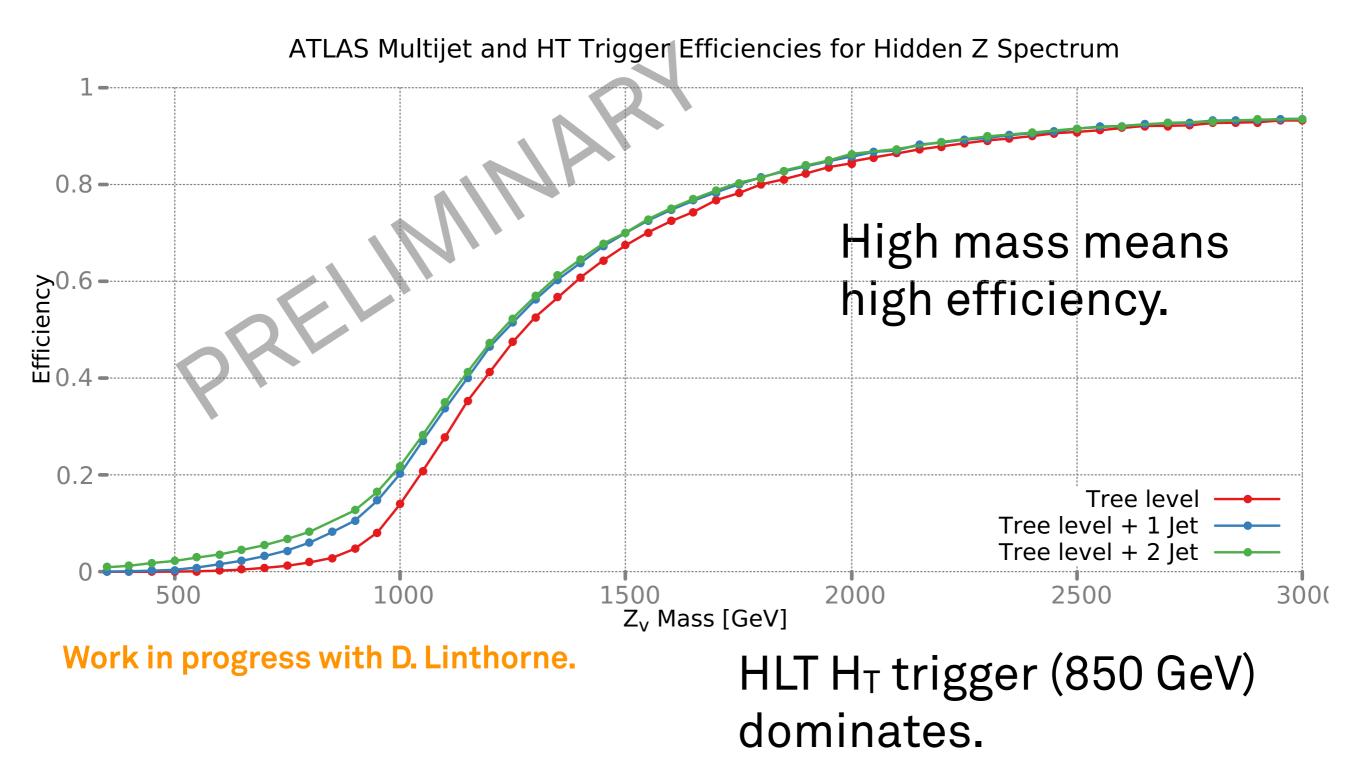
Strategy of mono-X searches.

For any dark shower, ISR is always there.

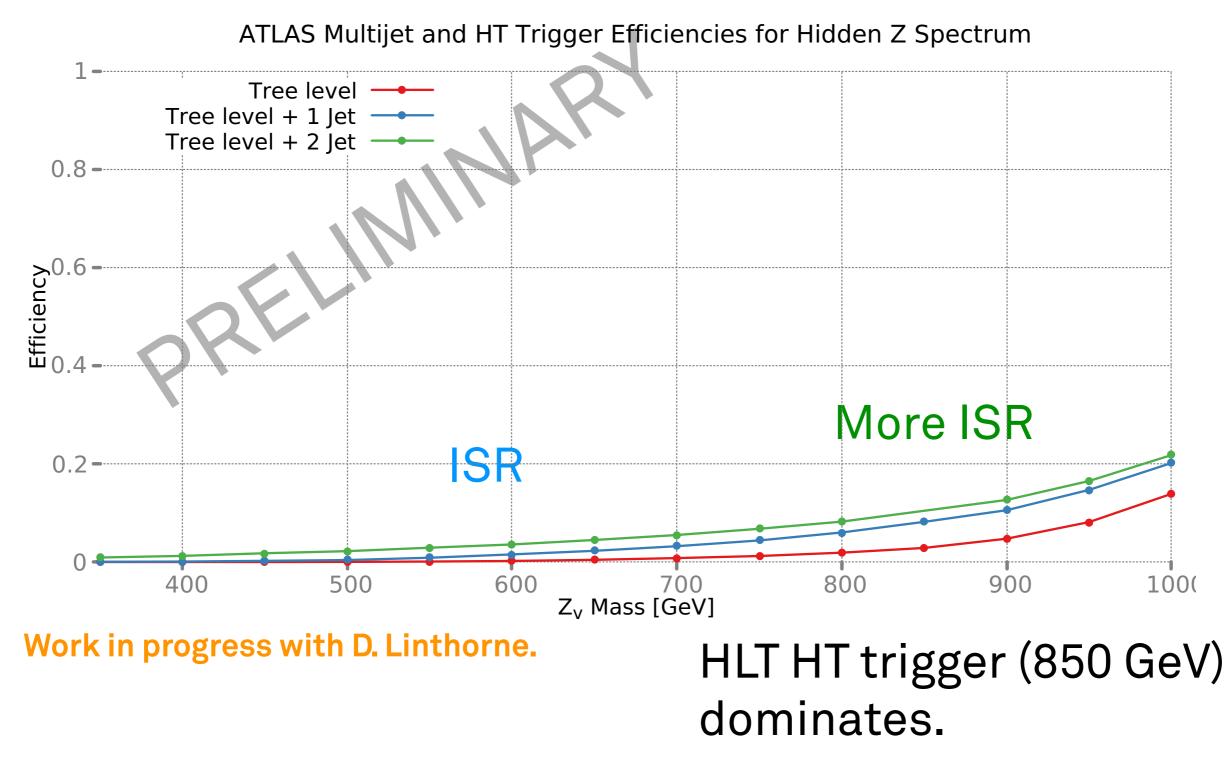


Strategy of mono-X searches.

TRIGGER EFFICIENCY



TRIGGER EFFICIENCY



ESSONS

Some showering models can naively be impossible to trigger.

Even for worst case model, can get a few percent trigger efficiency with ISR jet.

Worthwhile to simulate ISR (using usual matching procedure) to can get more coverage.

ONGOING WORK

To do:

- See if you can do better with multiple (vanilla) triggers.
- Check electroweak ISR.
- Compare to dedicated triggers.
- Interplay between L1 and HLT.

Stay tuned!

DETAILS

	Model A
Λ_d	$10 \mathrm{GeV}$
m_V	20 GeV
m_{π_d}	$5 \mathrm{GeV}$
$c au_{\pi_d}$	150 mm

Trigger menu (or):

- HT > 850 (pT > 50)
- 3 jets > 175
- 4 jets > 85
- 5 jets > 60
- 6 jets > 45

Delta function detector

