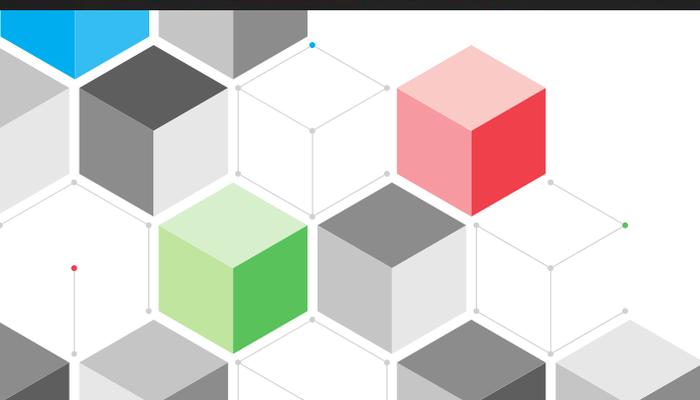




Mihaly VADAI & Harri TOIVONEN
November 14th, 2017

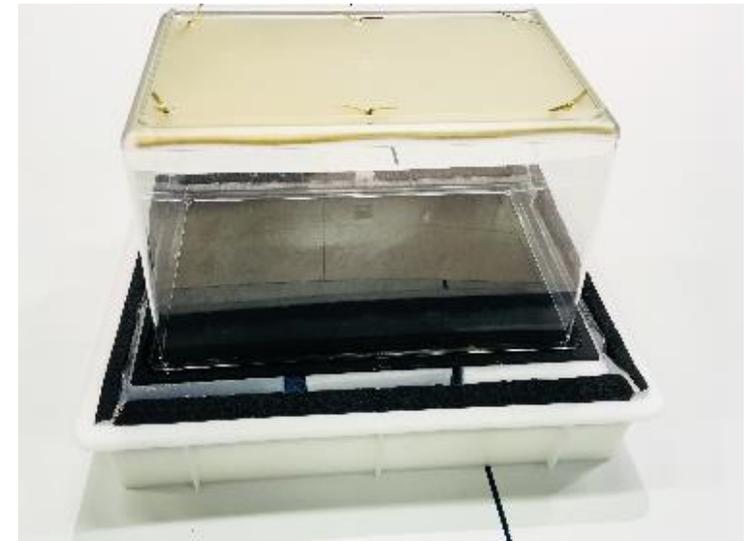


CERN Ideasquare

Activity Night

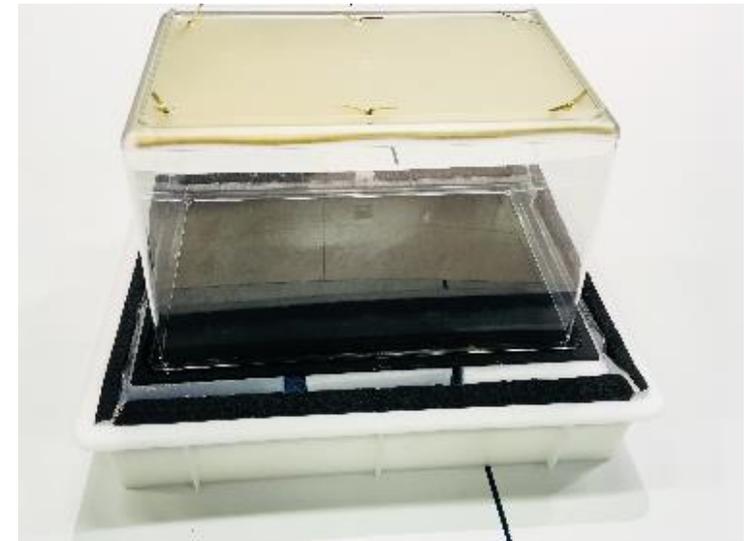
Program for the evening 1/2

- Put together two different kinds of particle detectors
 1. Assembling a MuonHunter – a coincidence detector
 - Open source kit for learning about particle physics, detector design, electronics and programming. Project started in 2014, hardware currently in revision 3.
 2. Assembling a cloud chamber – an “analog” detector
 - Widely used for educational purposes to learn about the presence of charged particles all around us.

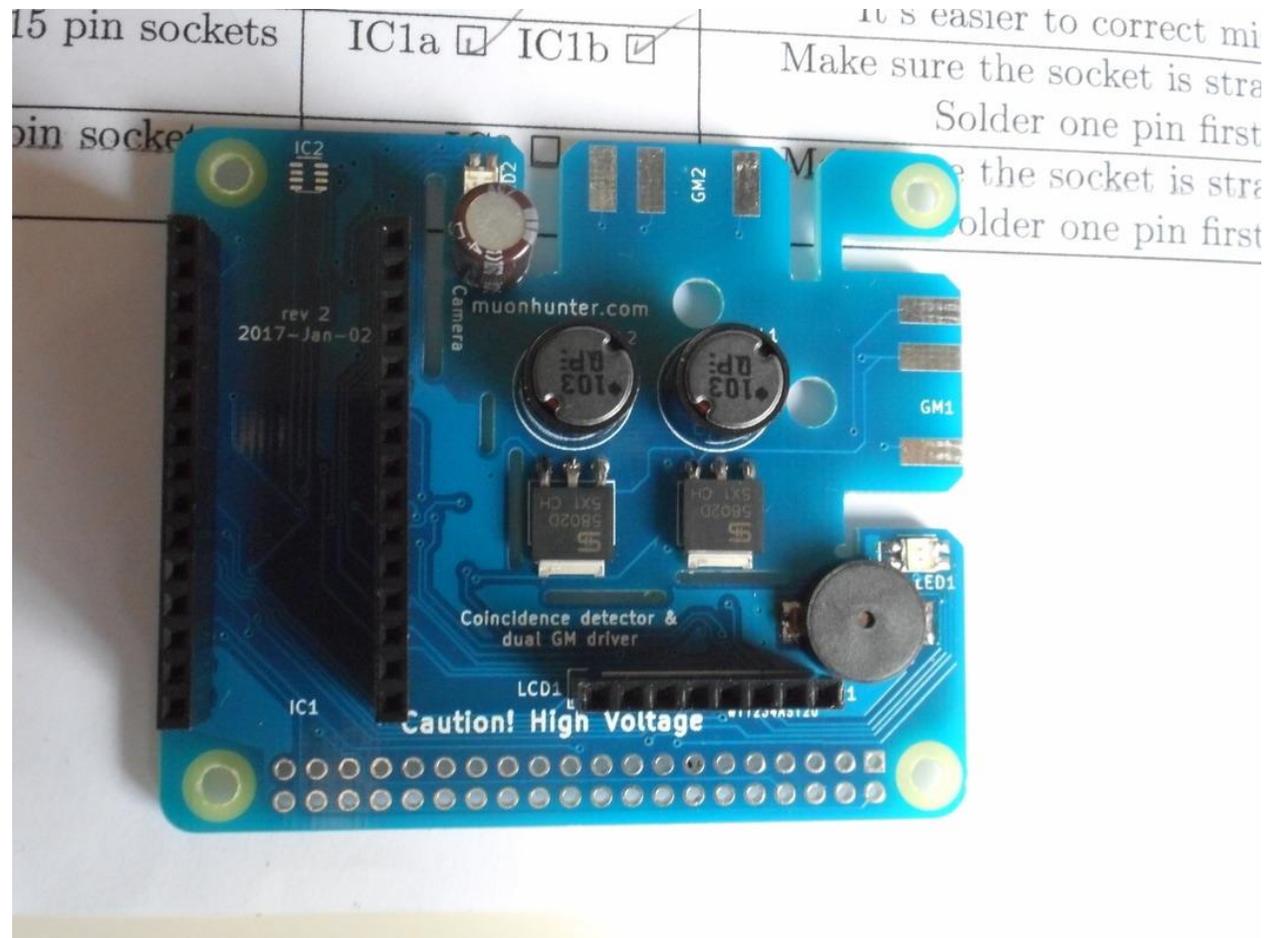


Program for the evening 2/2

- If time allows:
 3. Putting these two detectors together, to form a „hybrid“ detector
 - To take photos of the tracks of charged particles, in a (distantly) similar manner than is done with the large detector experiments at CERN (in a miniature, slightly less accurate way).



MuonHunter assembly & test



Assembly and observation tasks!

- Observe your MuonHunter
- Test whether different positioning angles effect coincidence detection frequency.
- Test whether spacing between GM counters have an effect on coincidence detection frequency.
- Assemble a MuonHunter “selfie machine” (i.e. take a photo of the MuonHunter when coincidence detected (leds flash))

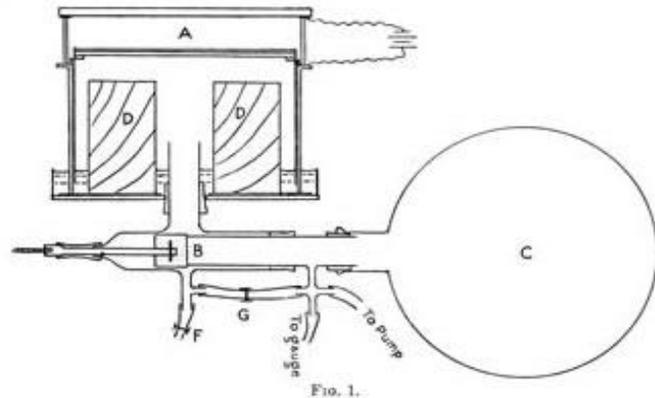
Cloud chamber assembly & observation



Cloud Chambers - History

Charles T. R. Wilson (1869 - 1959)

This Scottish physicist perfected the first (expansion) cloud chamber in 1911 and received the Nobel Prize in 1927.

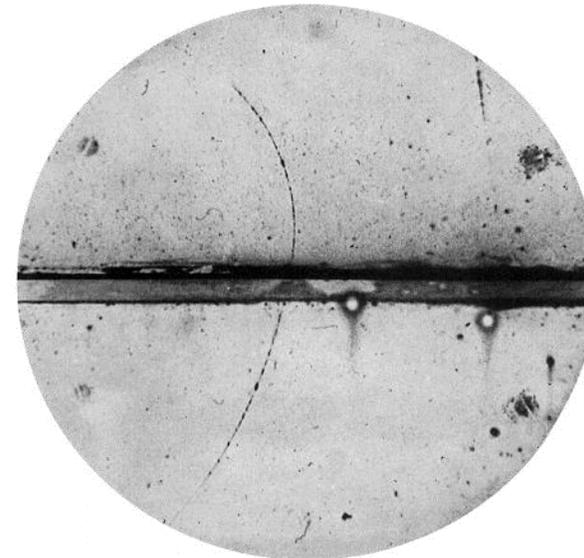


A diagram of Wilson's apparatus. The cylindrical cloud chamber ('A') is 16.5cm across by 3.4cm deep.

C. T. R. WILSON: *On an Expansion Apparatus for Making Visible the Tracks of Ionising Particles in Gases and Some Results Obtained by Its Use.* Proc. R. Soc. Lond. A. 1912 87 277-292 DOI:[10.1098/rspa.1912.0081](https://doi.org/10.1098/rspa.1912.0081)

Carl Anderson (1905 - 1991)

This physicist discovered the positron in 1932 and the muon in 1936 using a cloud chamber. He received the Nobel Prize in 1936.



Carl D. Anderson (1905–1991) - Anderson, Carl D. (1933). "The Positive Electron". *Physical Review* 43 (6): 491–494. DOI:[10.1103/PhysRev.43.491](https://doi.org/10.1103/PhysRev.43.491).

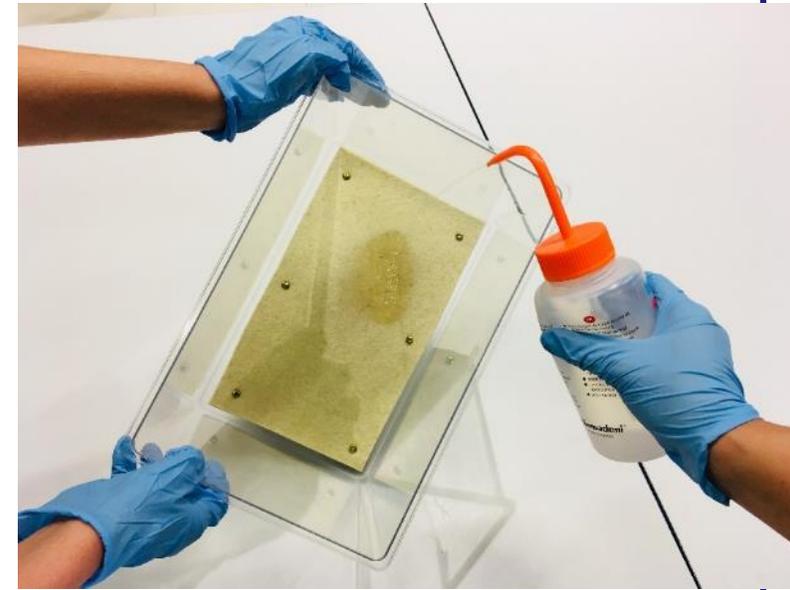
Build your cloud chamber - step by step

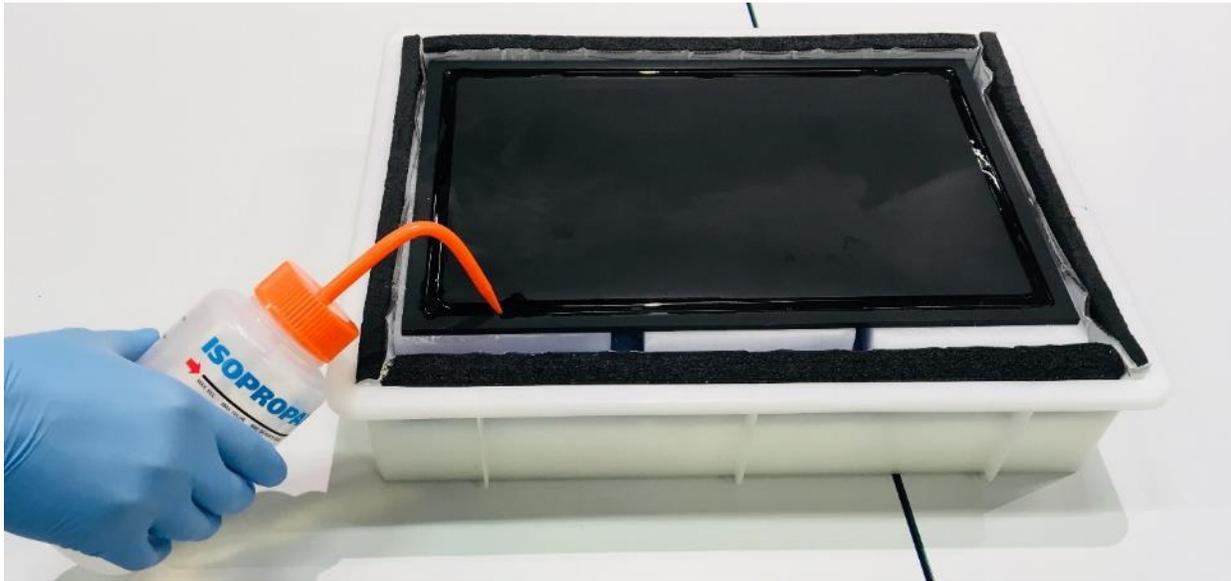


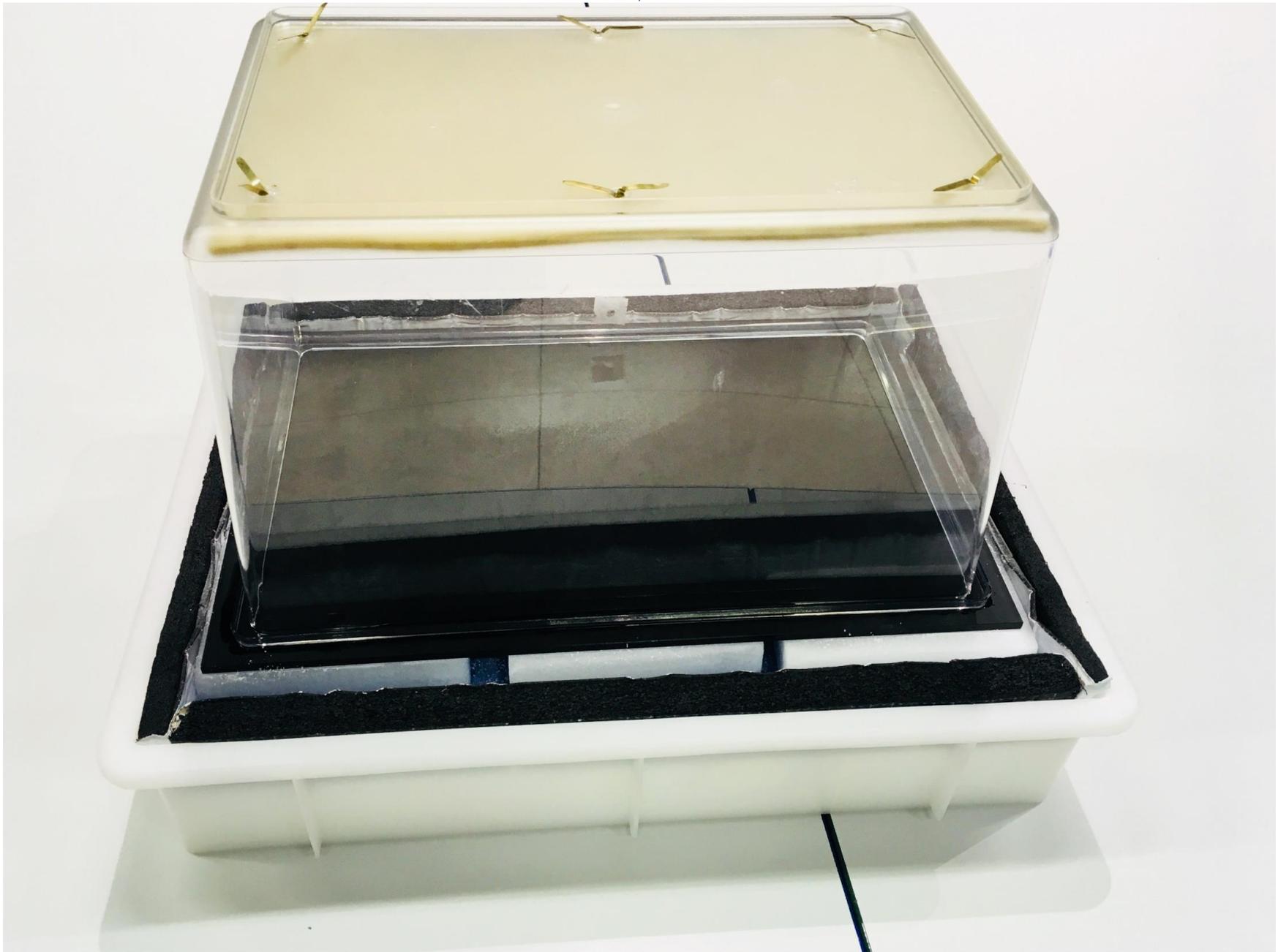
Team Dry ICE



Team ALCOHOL

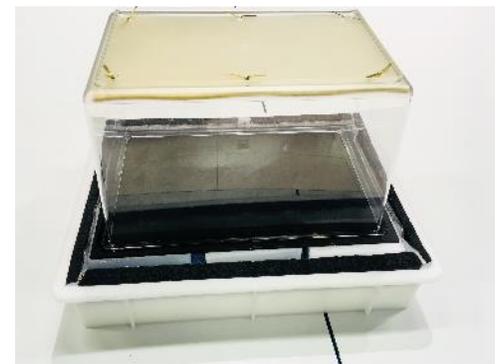
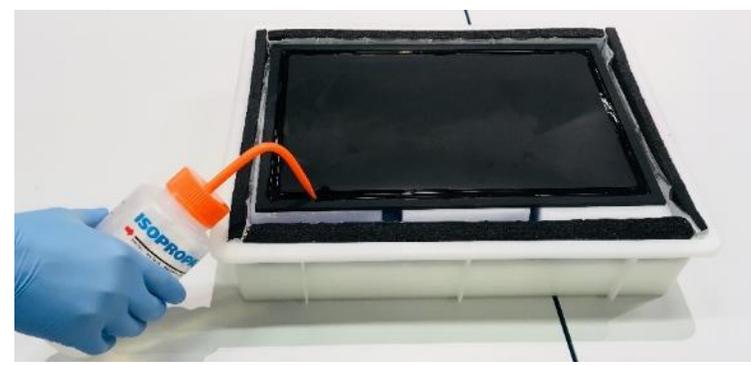
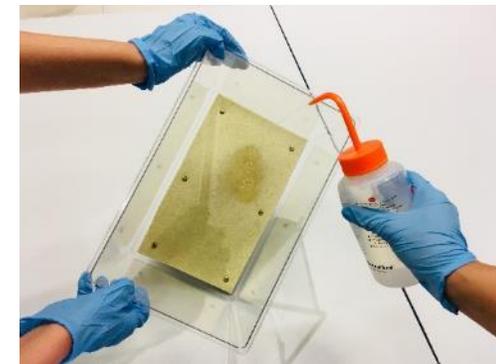






Team Dry ICE

Team ALCOHOL

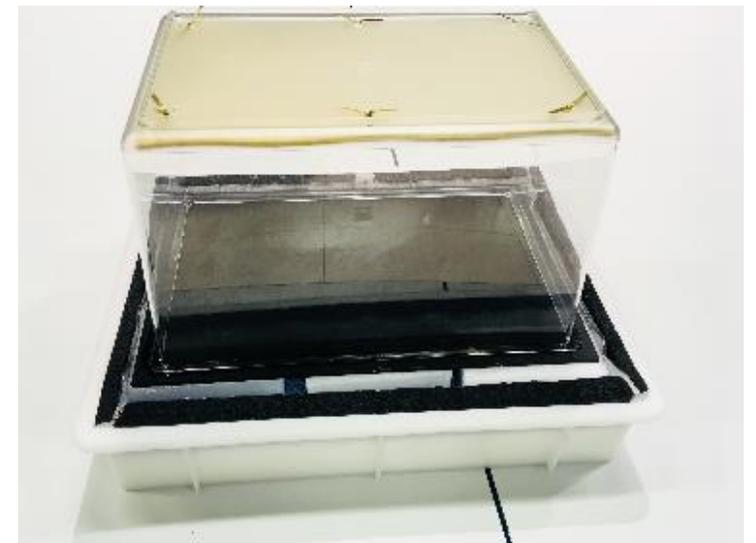


Observation tasks!

- Find the optimal **torch** position and the optimal observation position
- Observe your Cloud Chamber
- Describe visible tracks (shape, length, width, ...)
- Discuss the reason for these tracks
- Count the number of tracks you can see for 1 minute, repeat this measurement 2 times

If there is time..

- Putting MuonHunter and Cloud Chamber together, to form a „hybrid“ detector
 - To take photos of the tracks of charged particles, in a (distantly) similar manner than is done with the large detector experiments at CERN (in a miniature, slightly less accurate way).
- Optimal position for the electronics?
Camera? Lights? Can you catch a particle track with the camera?



Additional Material: Air Shower Simulation

Cosmic Ray Air Shower Pictures

by H.-J. Drescher drescher@th.physik.uni-frankfurt.de.

Air showers are cascades of secondary particles induced in the atmosphere by high energy cosmic rays. What you see here is a **visualisation of realistic simulations of these showers**. Of course, not all of the particles in a shower are displayed, there are far too many! The **fraction displayed here is about $1e-6$** , sampled with a **thinning algorithm**.

blue:electrons/positrons

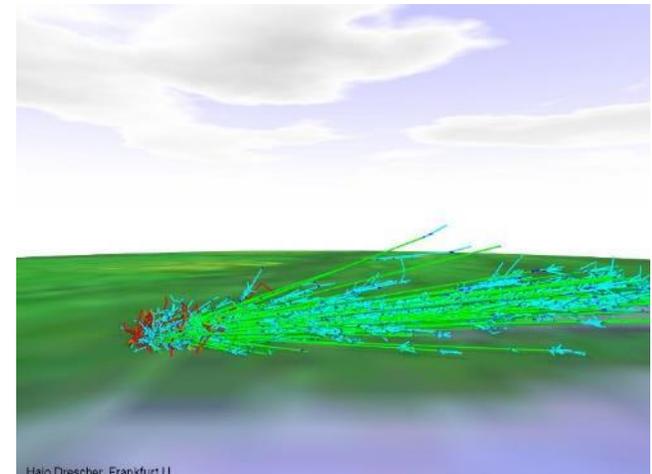
cyan:photons

red:neutrons

orange: protons

gray: mesons

green:muons



<http://th.physik.uni-frankfurt.de/~drescher/CASSIM/>

blue:electrons/positrons

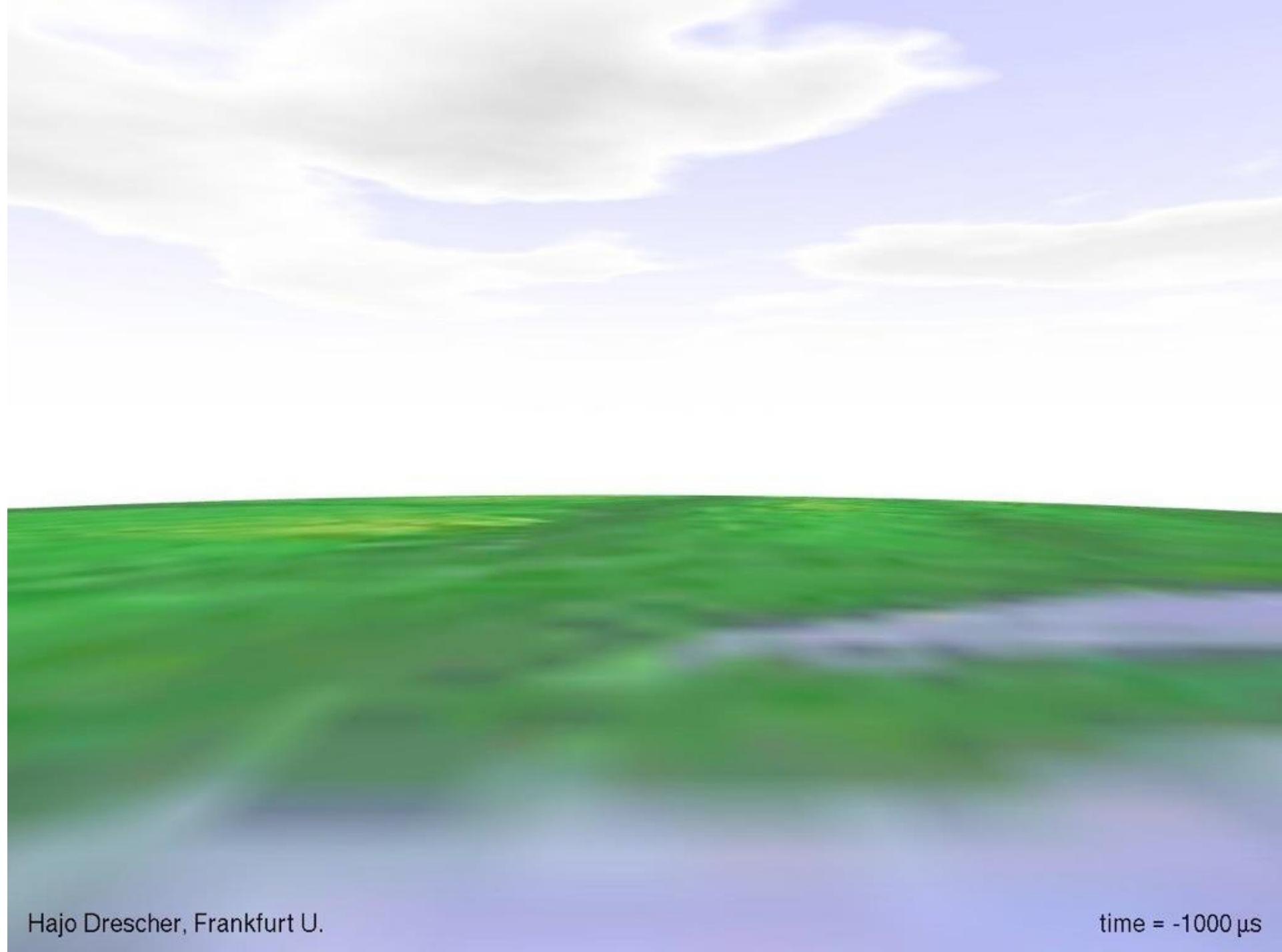
cyan:photons

red:neutrons

orange: protons

gray: mesons

green:muons



blue:electrons/positrons

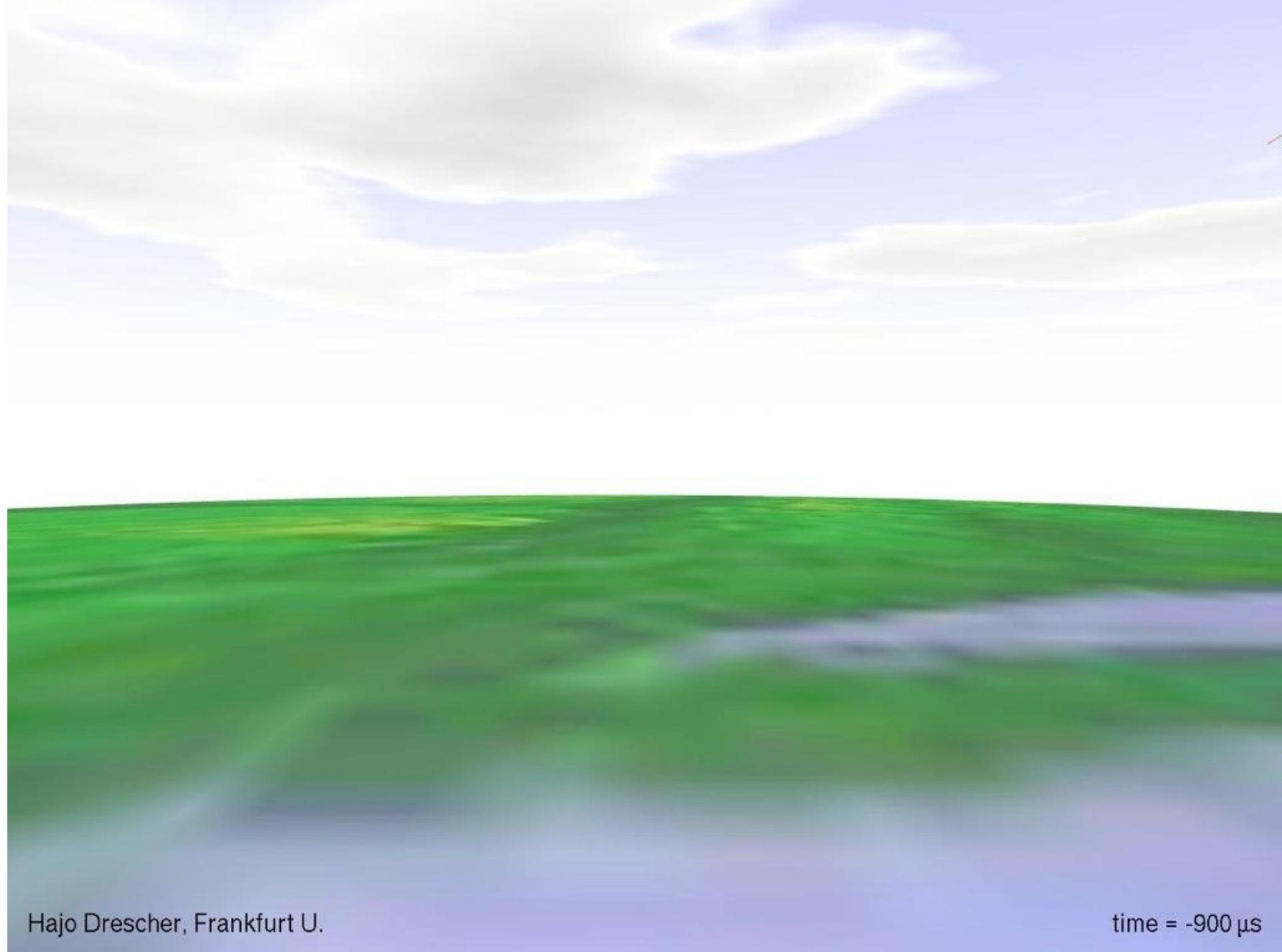
cyan:photons

red:neutrons

orange: protons

gray: mesons

green:muons



blue:electrons/positrons

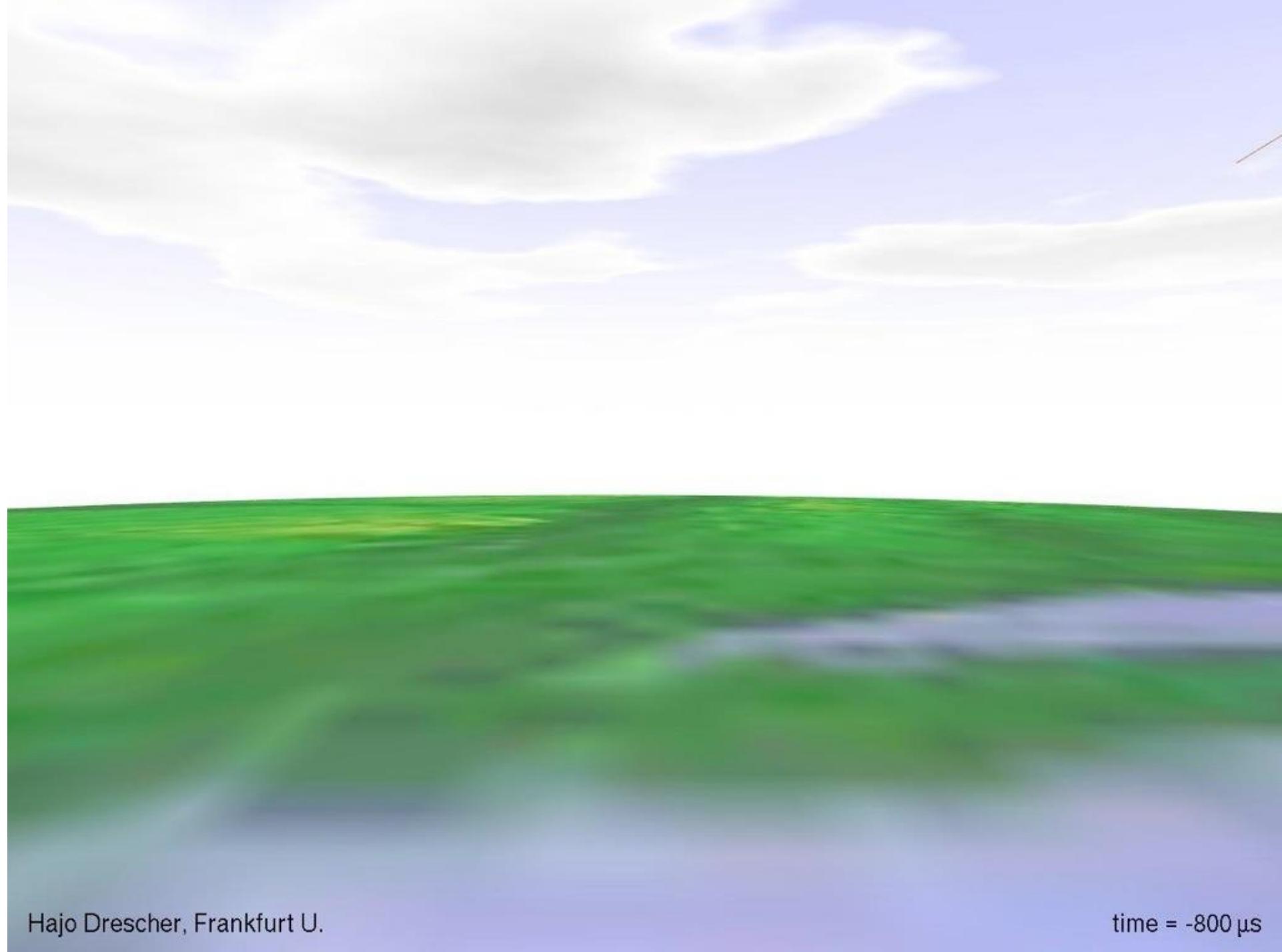
cyan:photons

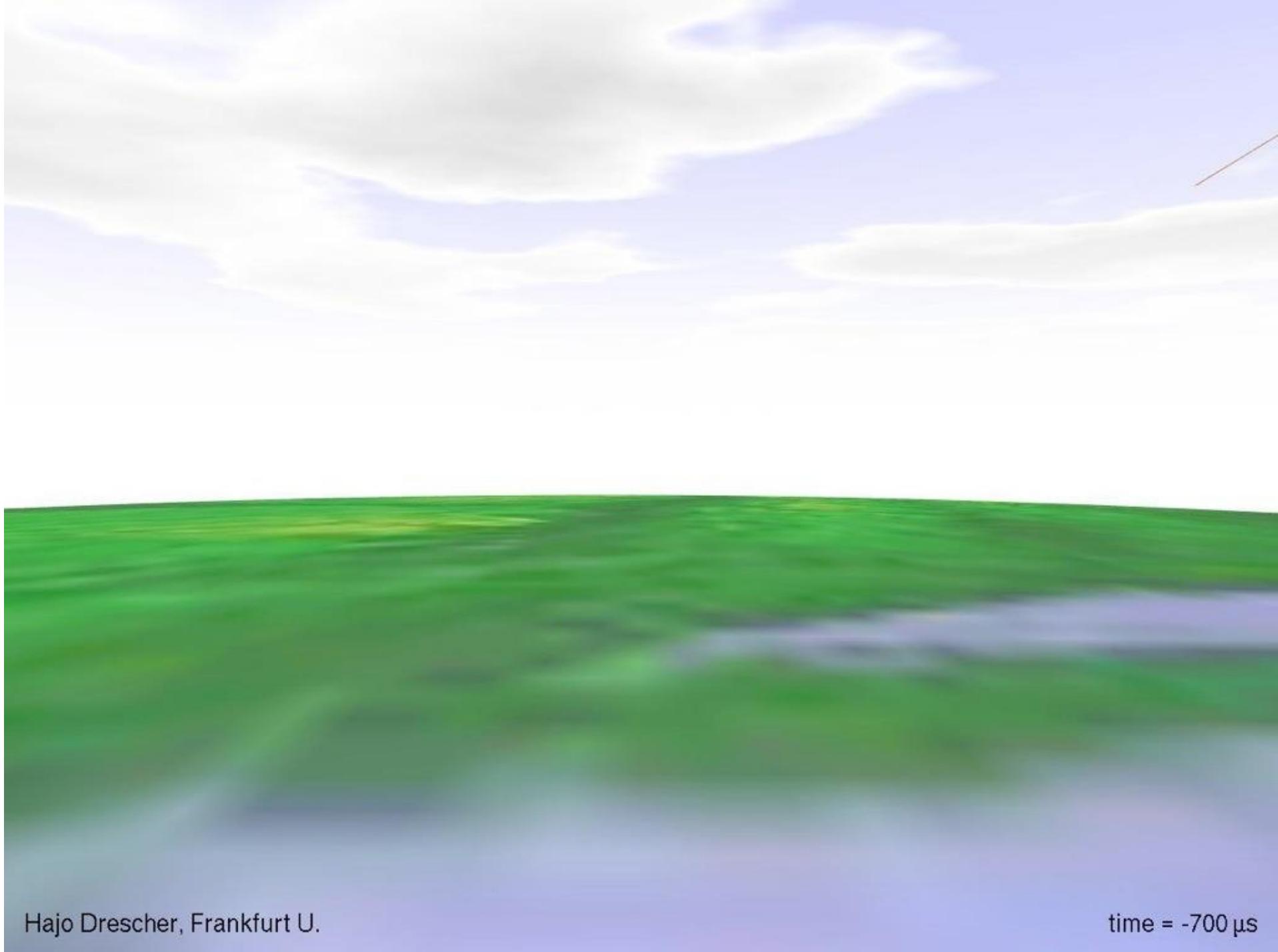
red:neutrons

orange: protons

gray: mesons

green:muons





blue:electrons/positrons

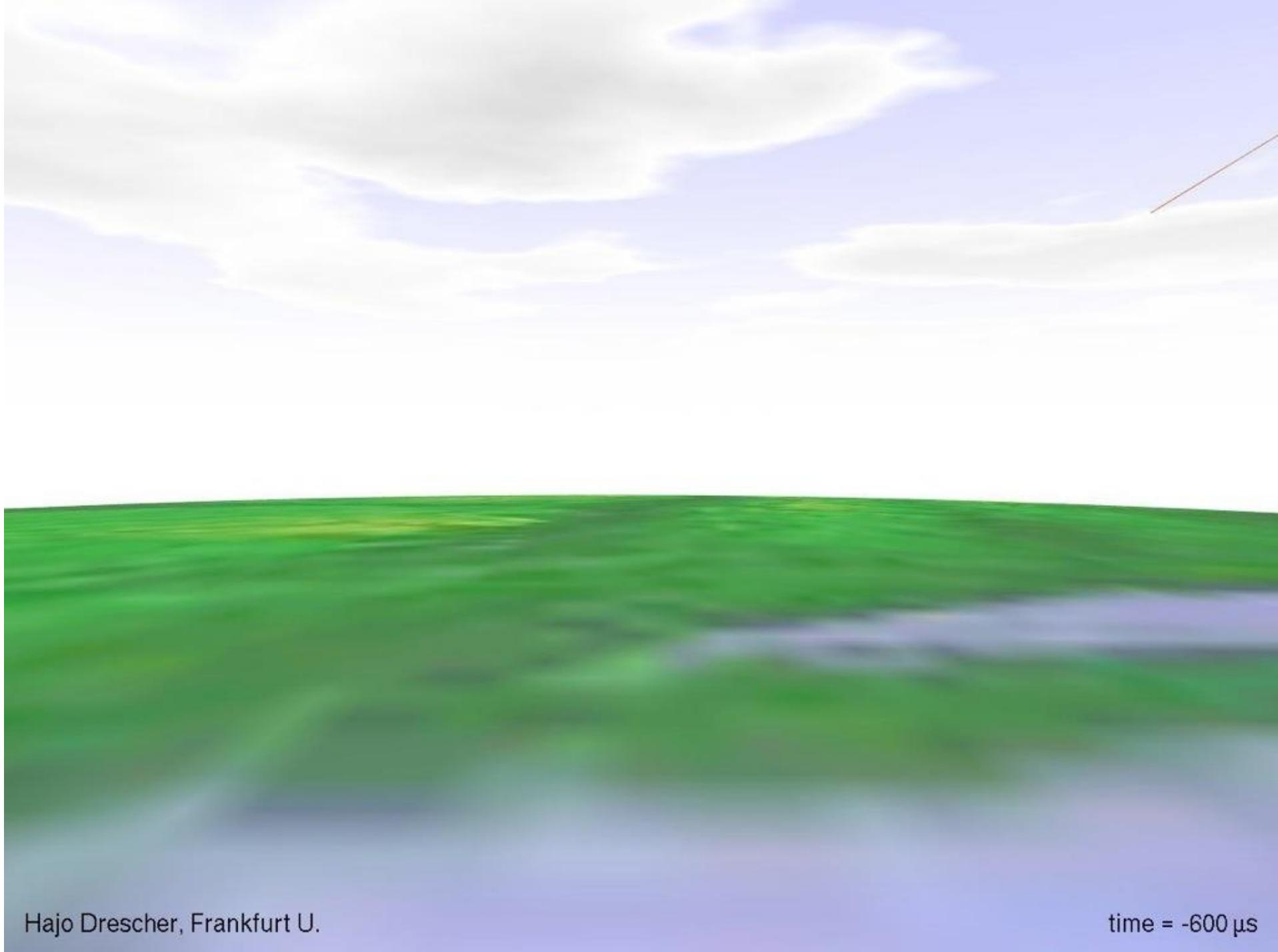
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blue:electrons/positrons

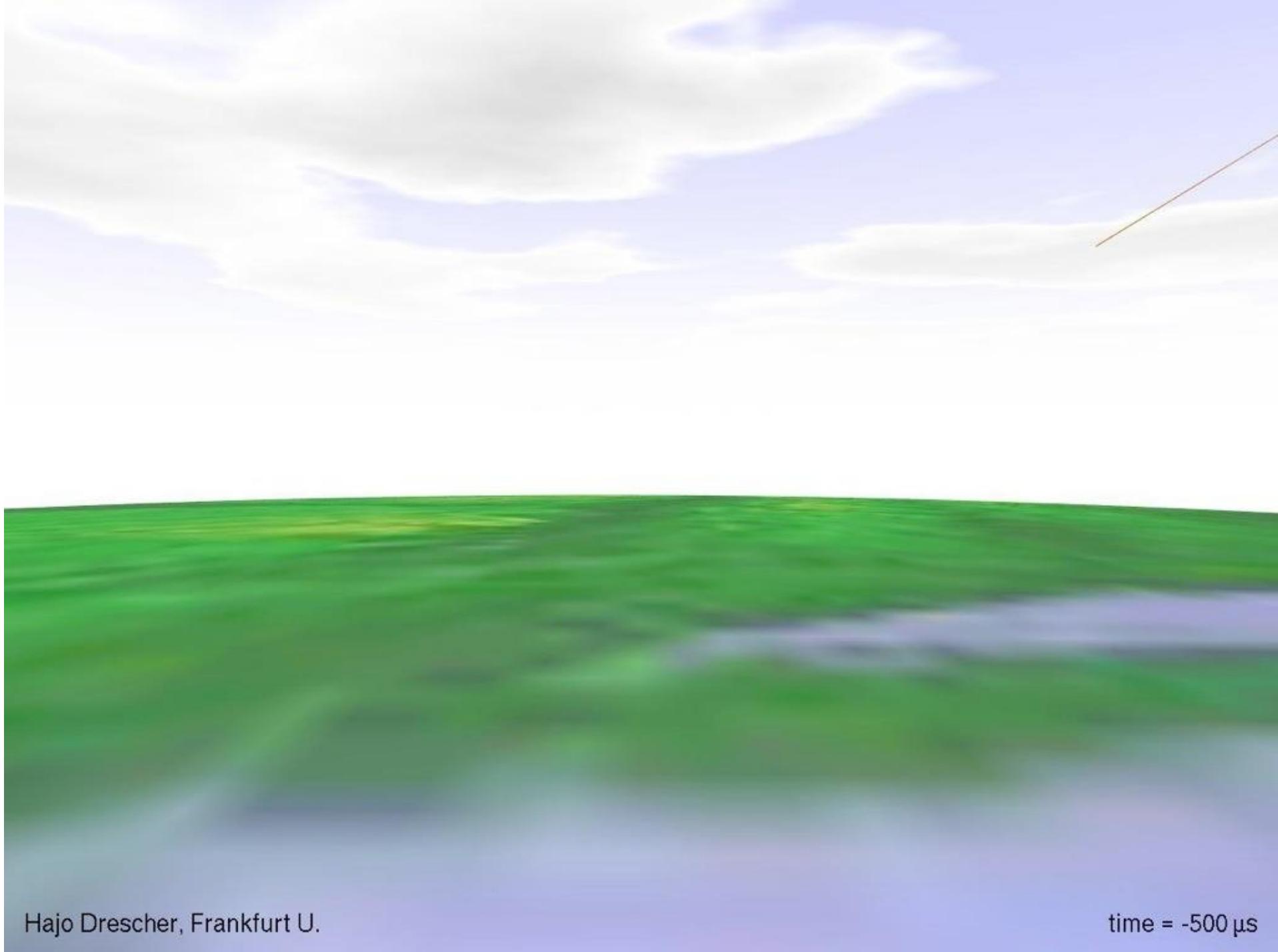
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blue:electrons/positrons

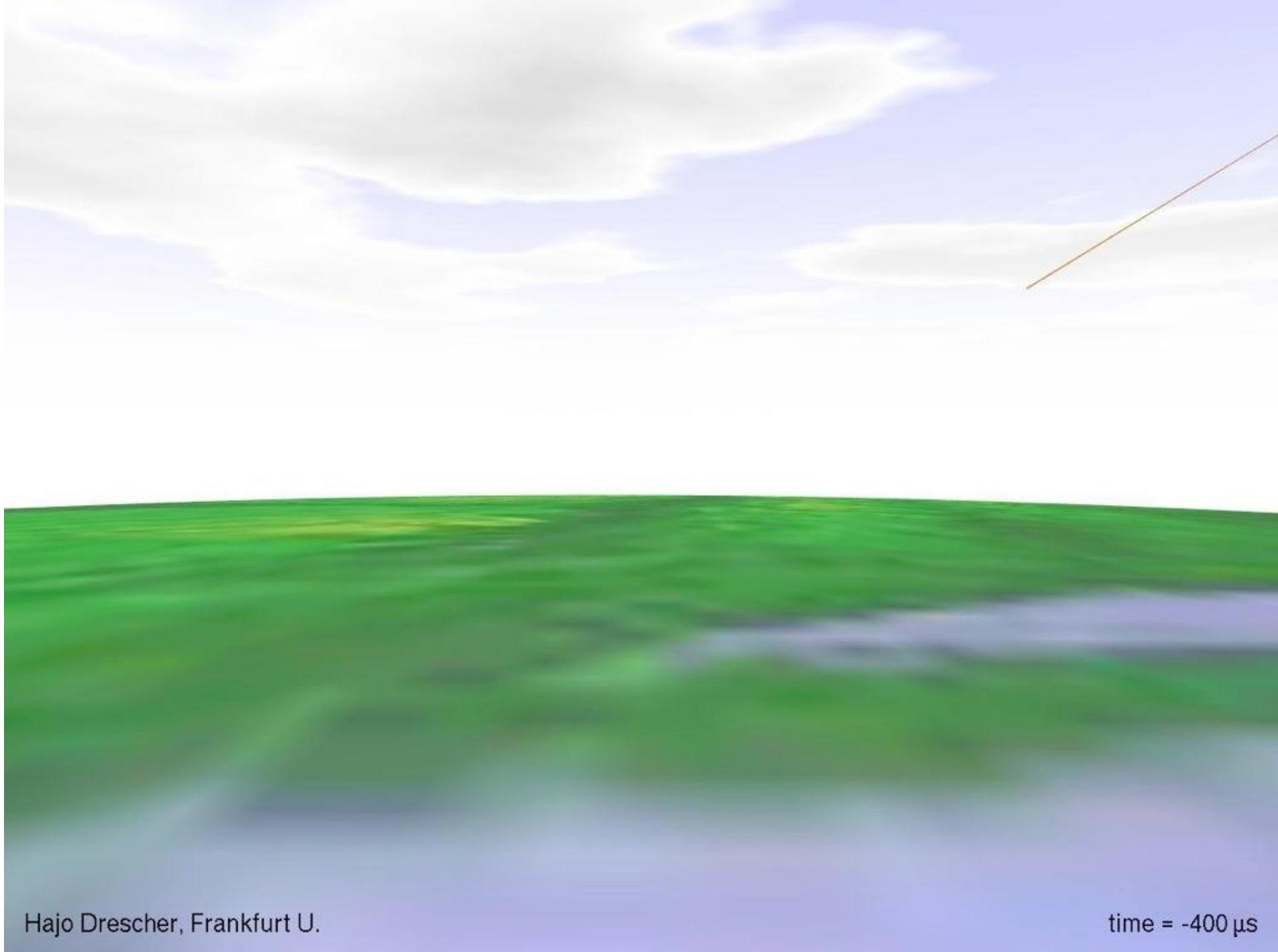
cyan:photons

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blue:electrons/positrons

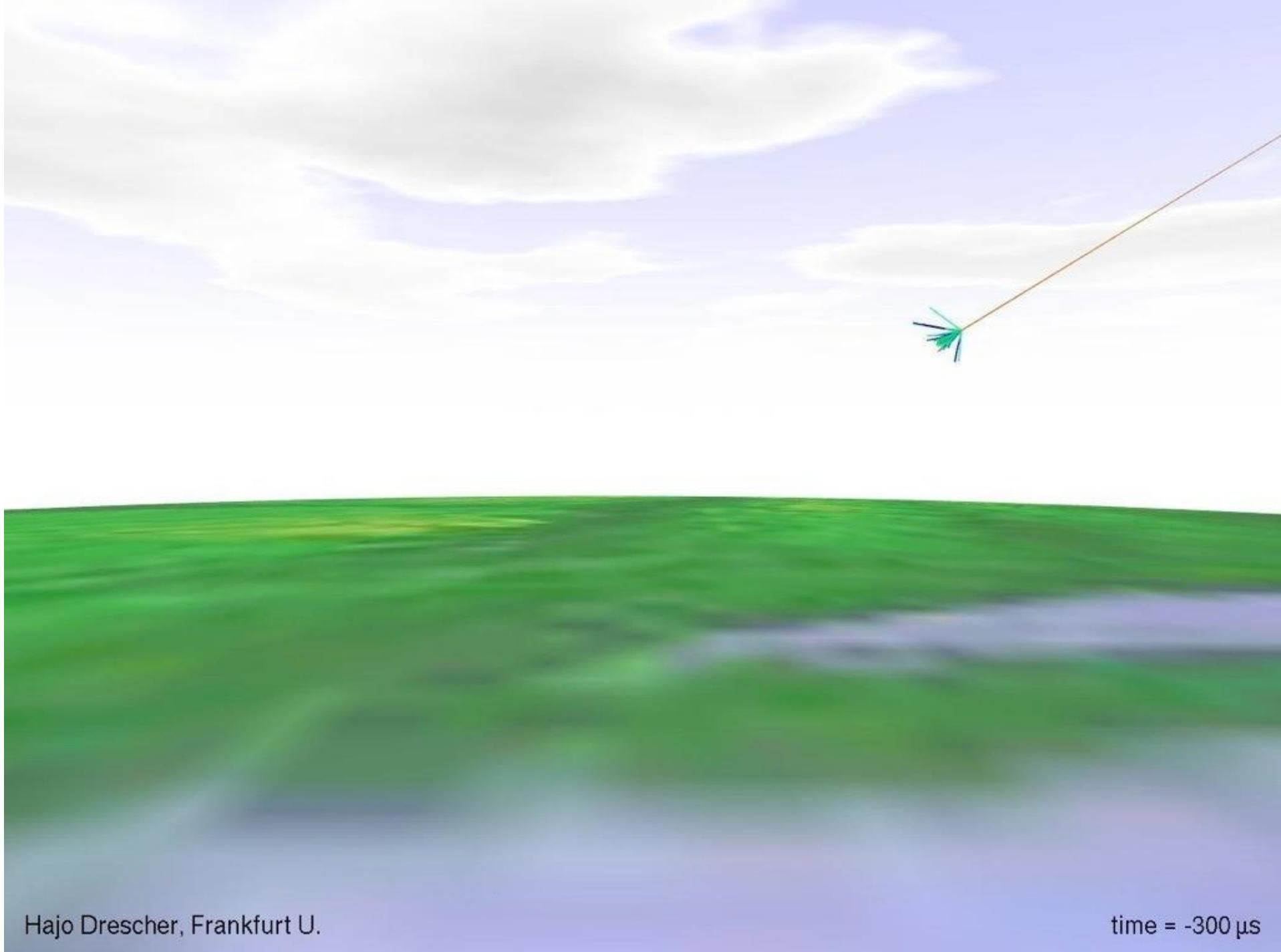
cyan:photons

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blue:electrons/positrons

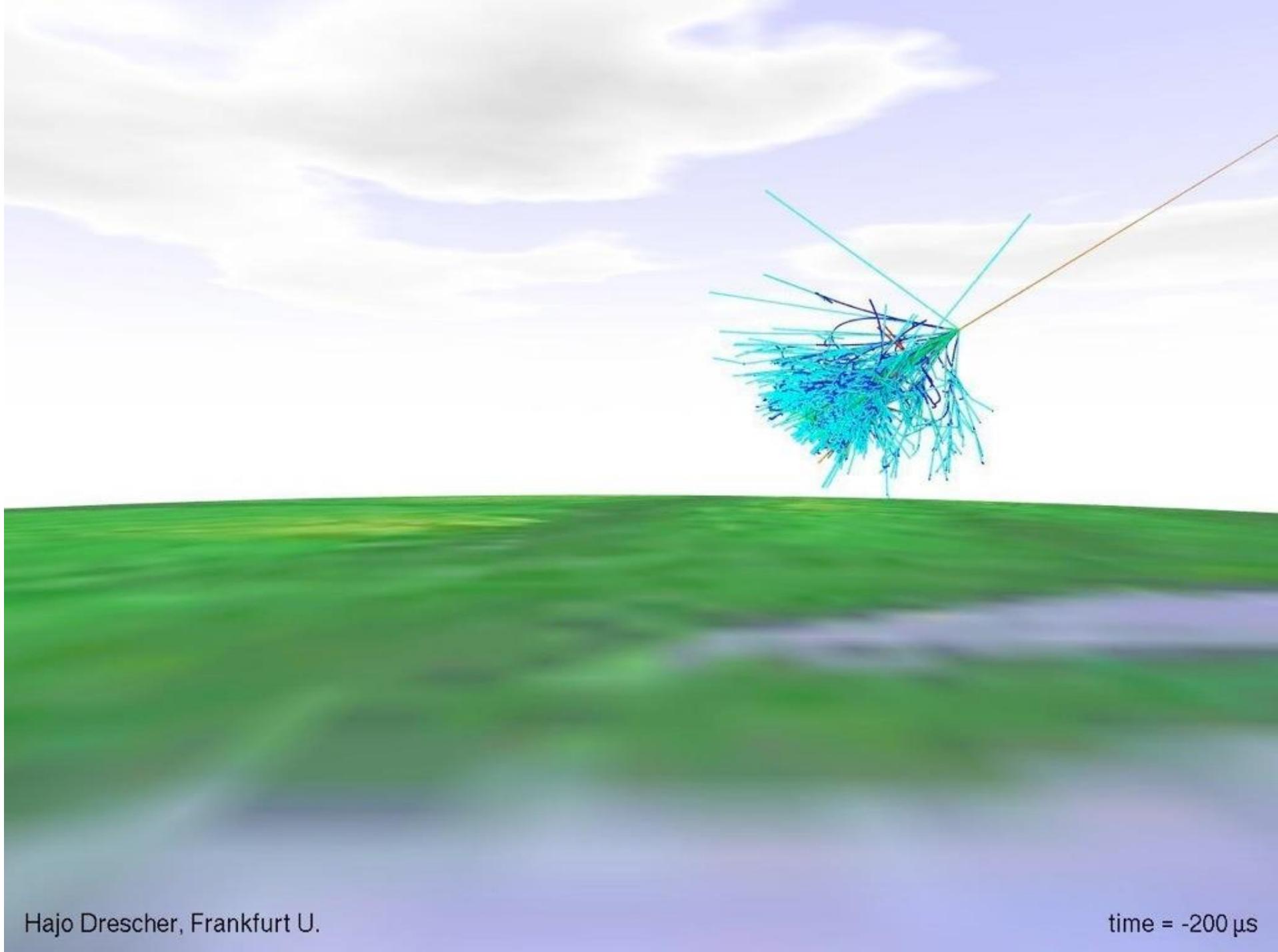
cyan:photons

red:neutrons

orange: protons

gray: mesons

green:muons



blue:electrons/positrons

cyan:photons

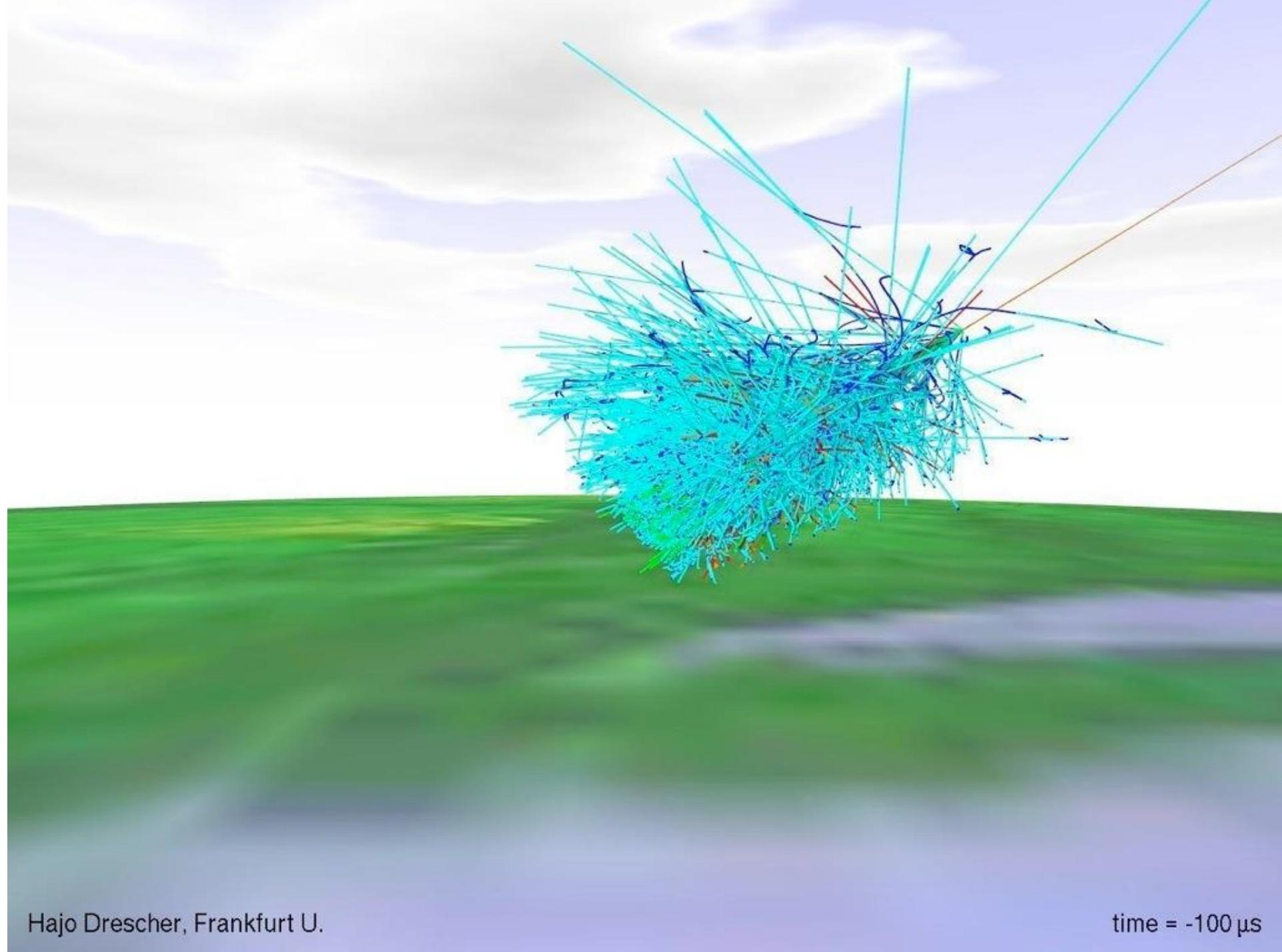
red:neutrons

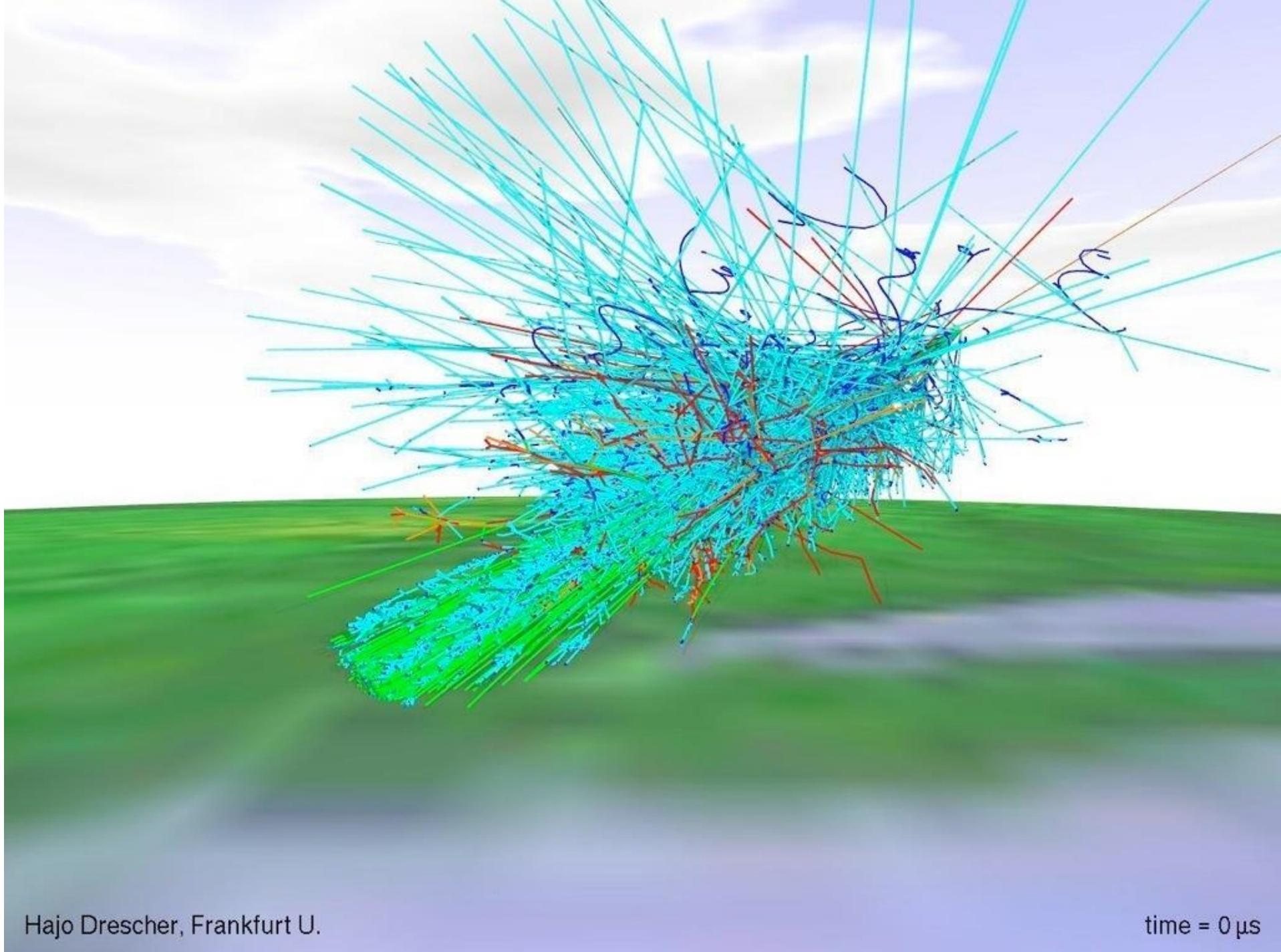
orange: protons

gray: mesons

green:muons

blue:electrons/positrons
cyan:photons
red:neutrons
orange: protons
gray: mesons
green:muons





blue:electrons/positrons

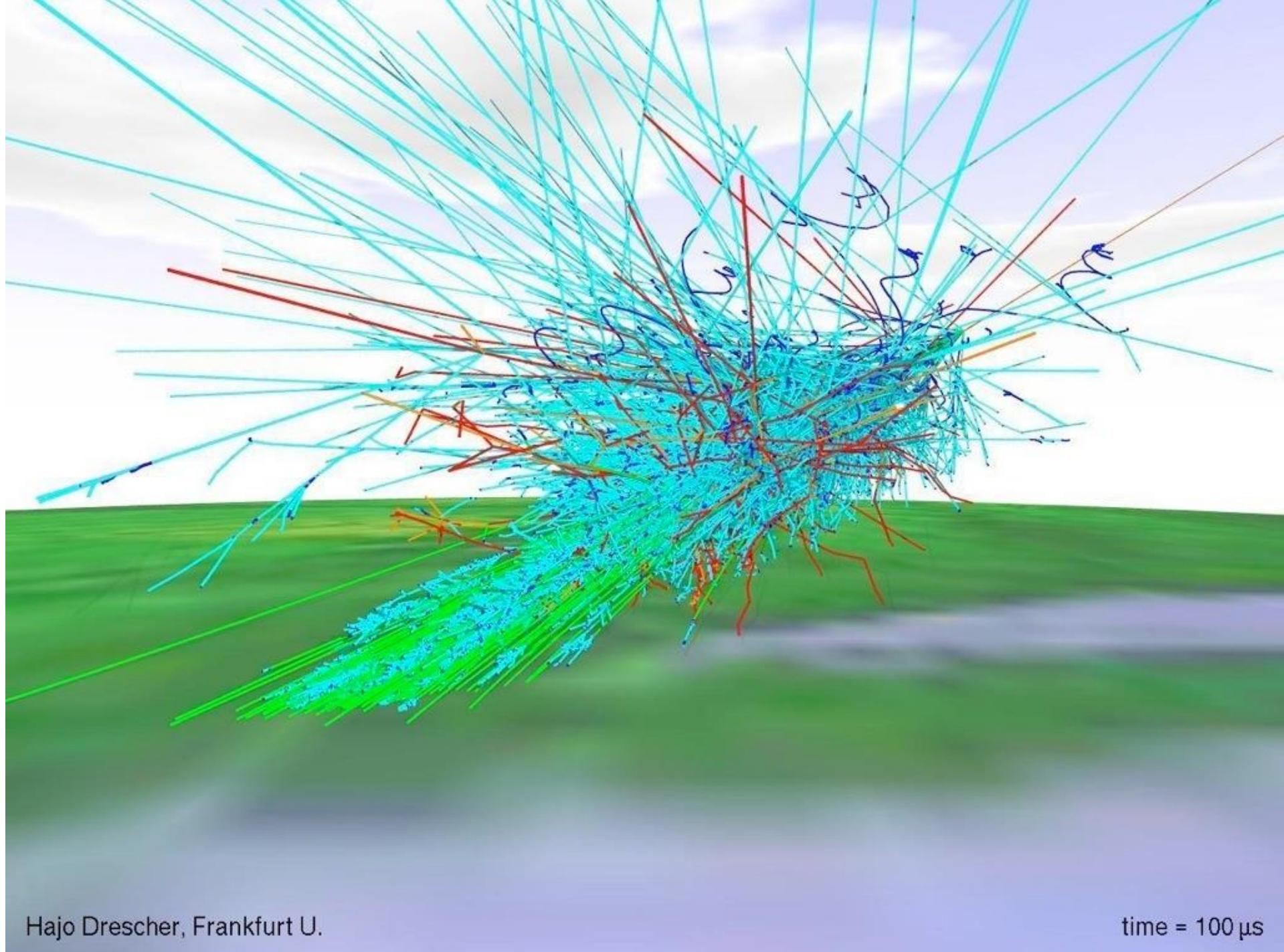
cyan:photons

red:neutrons

orange: protons

gray: mesons

green:muons



blue:electrons/positrons

cyan:photons

red:neutrons

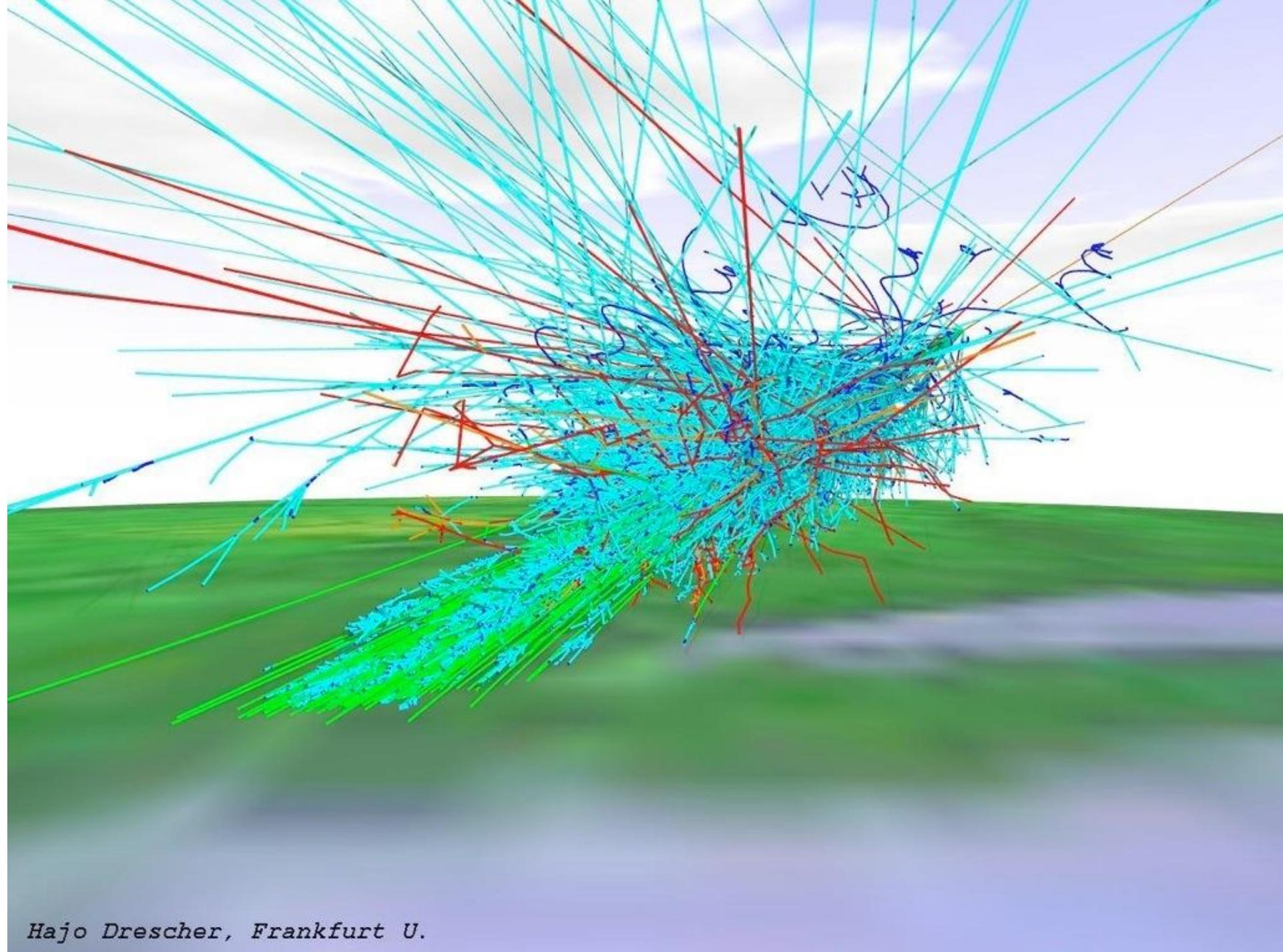
orange: protons

gray: mesons

green:muons

Hajo Drescher, Frankfurt U.

time = 100 μ s



blue:electrons/positrons

cyan:photons

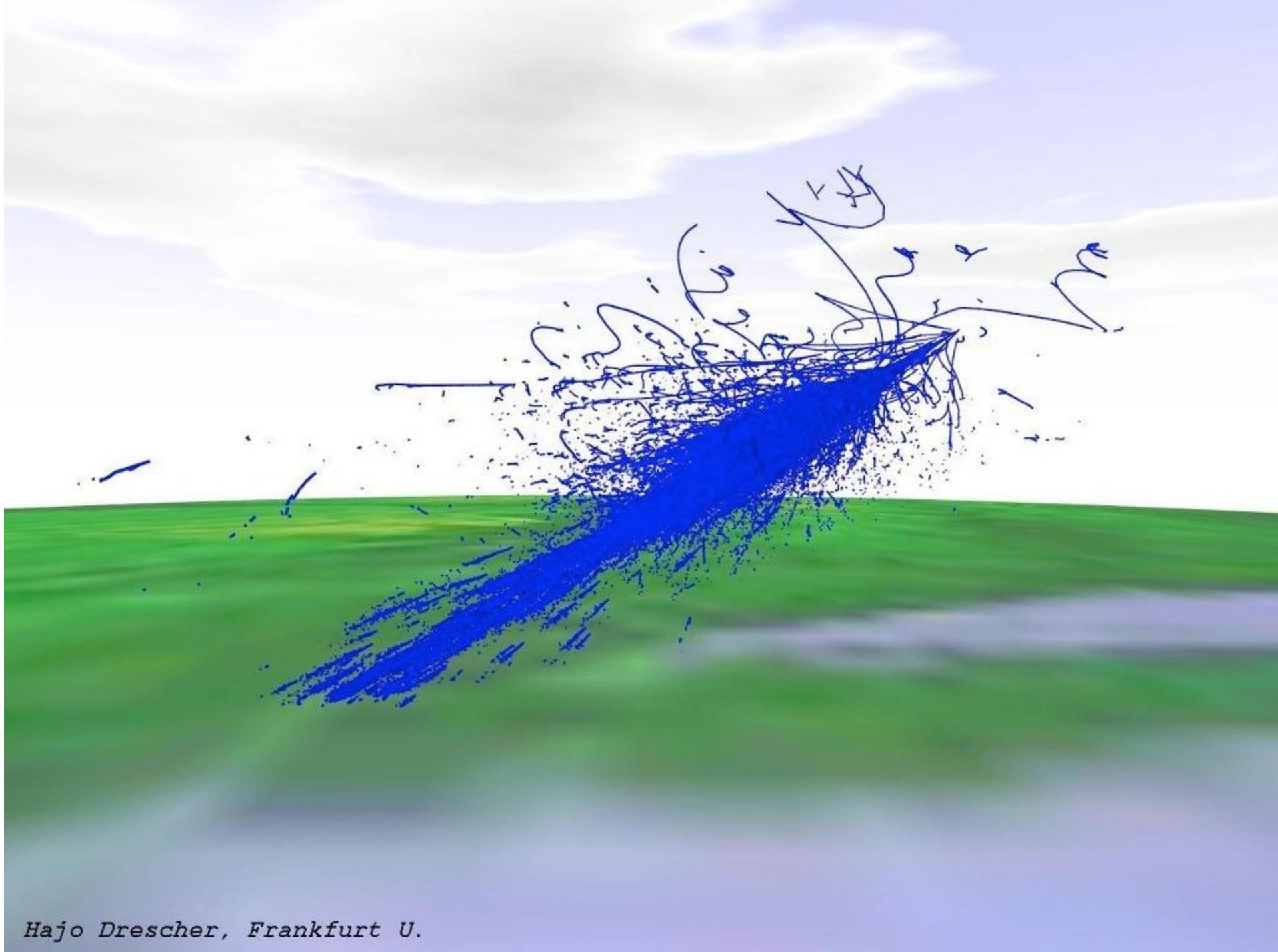
red:neutrons

orange: protons

gray: mesons

green:muons

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blue:electrons/positrons

cyan:photons

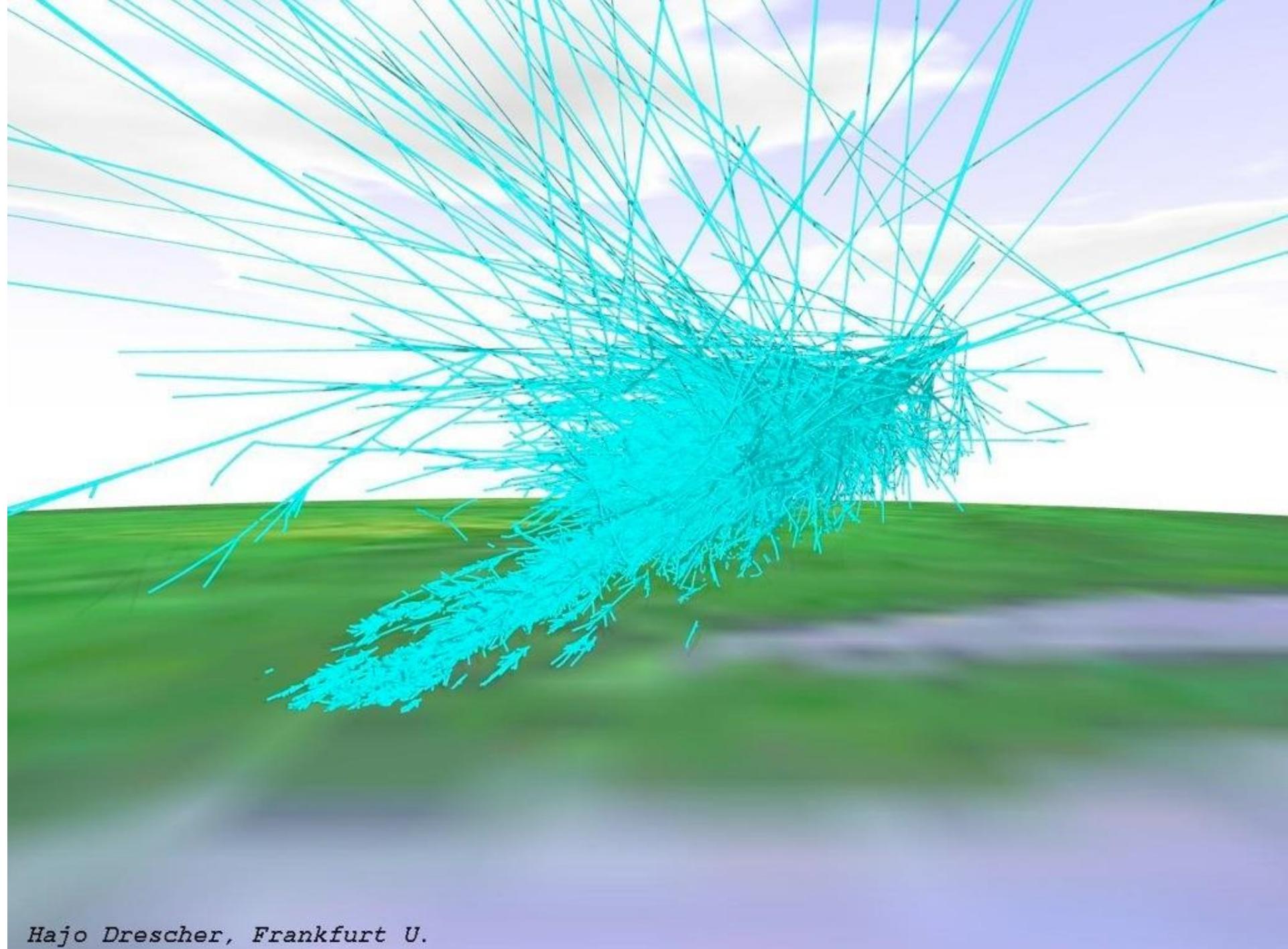
red:neutrons

orange: protons

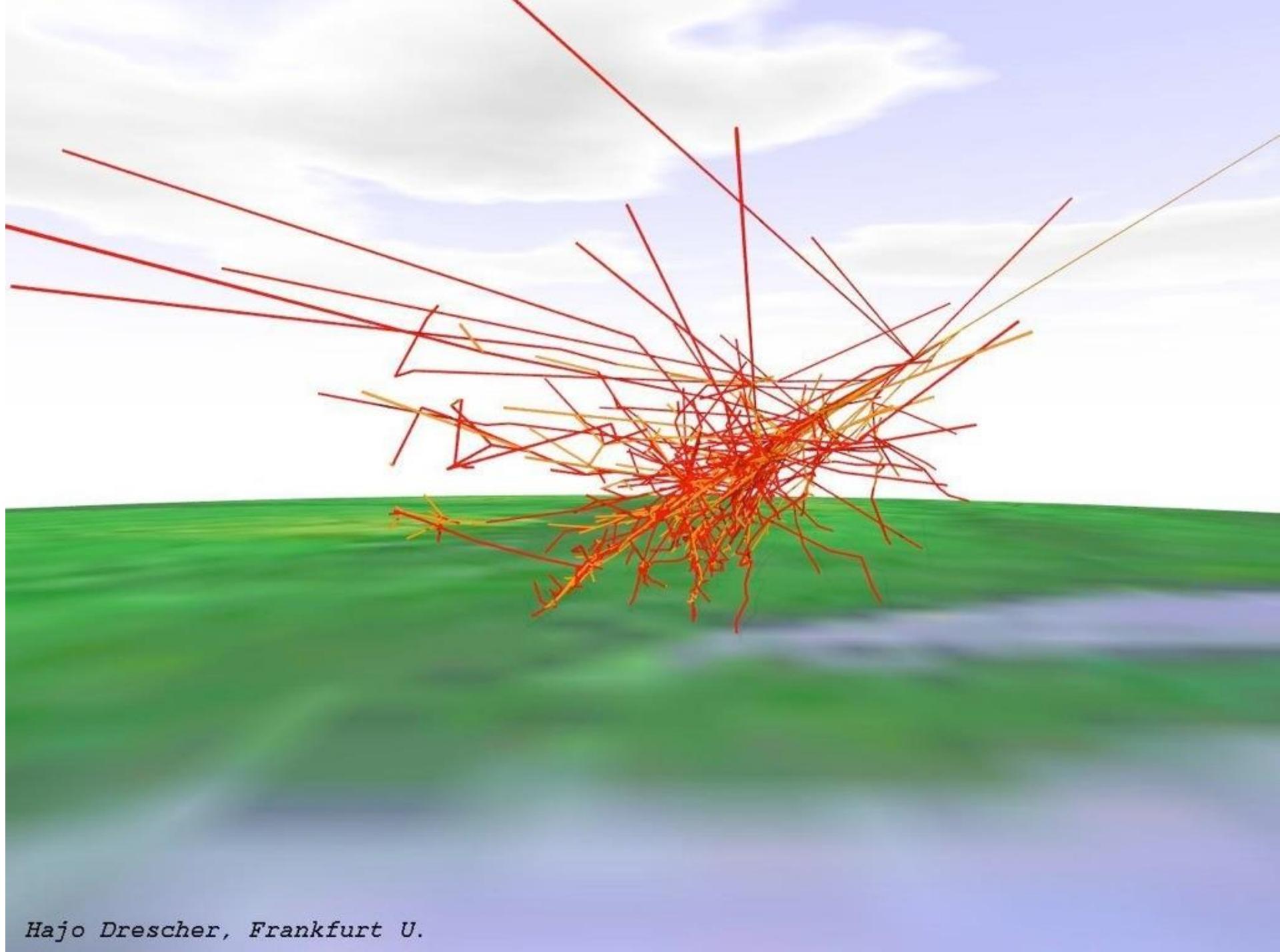
gray: mesons

green:muons

blue:electrons/positrons
cyan:photons
red:neutrons
orange: protons
gray: mesons
green:muons



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blue:electrons/positrons

cyan:photons

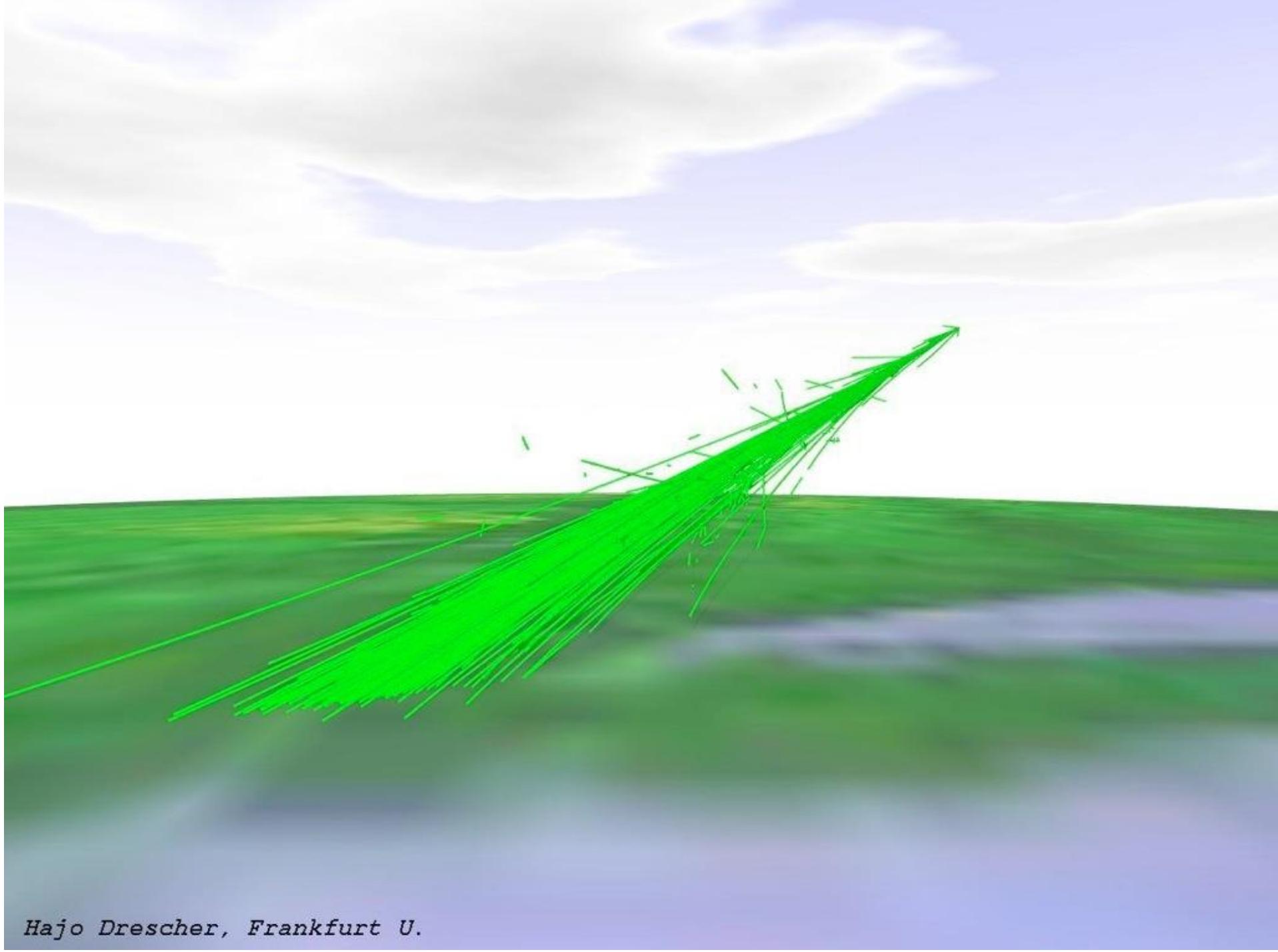
red:neutrons

orange: protons

gray: mesons

green:muons

Hajo Drescher, Frankfurt U.



blue:electrons/positrons

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red:neutrons

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gray: mesons

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Hajo Drescher, Frankfurt U.