

Full result for the three-loop static quark potential

V.A. Smirnov

Nuclear Physics Institute of Moscow State University

in collaboration with A.V. Smirnov and M. Steinhauser

- Introduction. Our calculational scheme
- Our techniques:
 - reduction to master integrals by FIRE;
 - evaluating master integrals by Mellin–Barnes representation;
 - numerical checks by FIESTA
- Results

The QCD potential between a heavy static quark and its antiquark in an expansion in α_s and heavy-quark velocity $v \sim \alpha_s$.

Two alternative definitions of the potential:
in terms of a Wilson loop,
in the framework of an effective field theory.

The dynamics of a nonrelativistic quark-antiquark pair involves three well separated scales:
hard $\sim m_q$; soft $\sim m_q v$; ultrasoft $\sim m_q v^2$

Expansion by regions

[M. Beneke & V.A. Smirnov '98]

Four regions are relevant:

hard (energy and momentum $\sim m_q$);

soft (energy and momentum $\sim m_q v$);

potential (energy $\sim m_q v^2$, momentum $\sim m_q v$);

ultrasoft (energy and momentum $\sim m_q v^2$)

‘Integrating out’ the hard modes: QCD \rightarrow NRQCD

‘Integrating out’ the soft modes and potential gluons:

NRQCD \rightarrow pNRQCD.

The static potential as an operator in pNRQCD Lagrangian.

The static colour-singlet potential

$$V(|\mathbf{q}|) = -\frac{4\pi C_F \alpha_s(|\mathbf{q}|)}{q^2} \left[1 + \frac{\alpha_s(|\mathbf{q}|)}{4\pi} a_1 + \left(\frac{\alpha_s(|\mathbf{q}|)}{4\pi} \right)^2 a_2 + \left(\frac{\alpha_s(|\mathbf{q}|)}{4\pi} \right)^3 \left(a_3 + 8\pi^2 C_A^3 \ln \frac{\mu^2}{q^2} \right) + \dots \right]$$

with $C_A = N$ and $C_F = (N^2 - 1)/(2N)$ are the eigenvalues of the quadratic Casimir operators of the adjoint and fundamental representations of the $SU(N)$ colour gauge group, respectively, $T_F = 1/2$ is the index of the fundamental representation, and n_l is the number of light-quark flavours.

One loop

[W. Fischler '77; A. Billoire '80]

$$a_1 = \frac{31}{9}C_A - \frac{20}{9}T_F n_l$$

Two loops

[M. Peter'97; Y. Schröder'99]

$$a_2 = \left[\frac{4343}{162} + 4\pi^2 - \frac{\pi^4}{4} + \frac{22}{3}\zeta(3) \right] C_A^2 - \left[\frac{1798}{81} + \frac{56}{3}\zeta(3) \right] C_A T_F$$
$$- \left[\frac{55}{3} - 16\zeta(3) \right] C_F T_F n_l + \left(\frac{20}{9} T_F n_l \right)^2$$

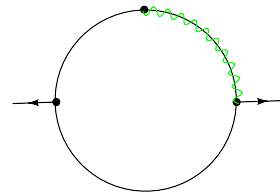
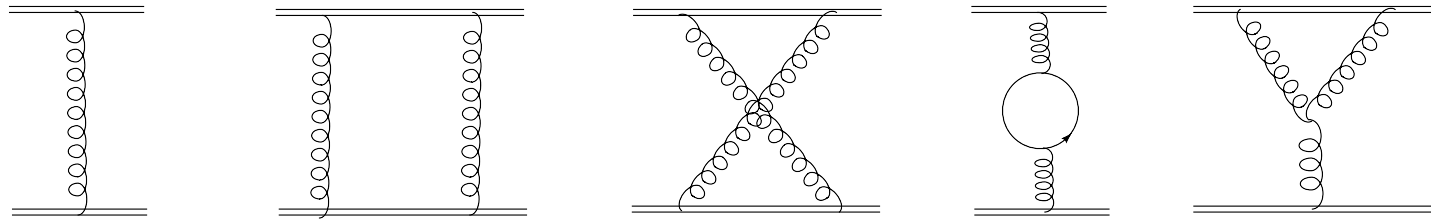
$1/m_q$ -correction

[B.A. Kniehl, A.A. Penin, V.A. Smirnov, and M. Steinhauser'02]

Three loops: $a_3 = a_3^{(3)} n_l^3 + a_3^{(2)} n_l^2 + a_3^{(1)} n_l + a_3^{(0)} = ?$

a_3 is one of a few missing NNNLO ingredients of the non-relativistic dynamics near threshold.

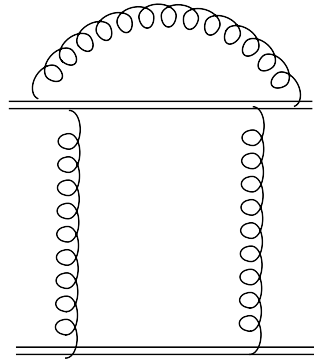
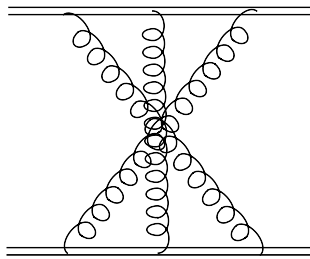
Tree and one-loop approximations:



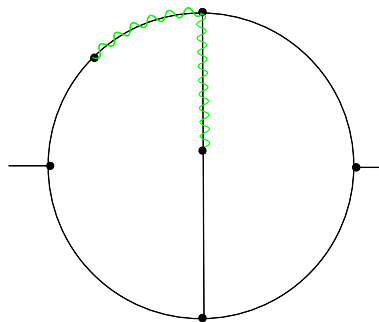
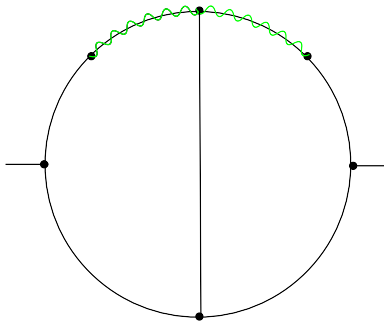
$$F(a_1, a_2, a_3) = \int \frac{d^d k}{(-k^2 - i0)^{a_1} (-(q - k)^2 - i0)^{a_2} (-v \cdot k - i0)^{a_3}}$$

with $q \cdot v = 0$, $v = (1, \vec{0})$

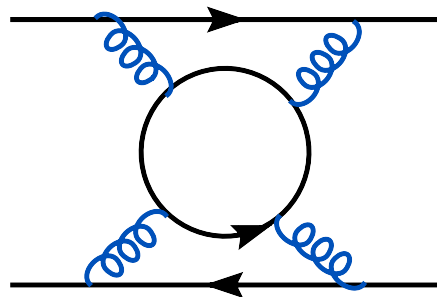
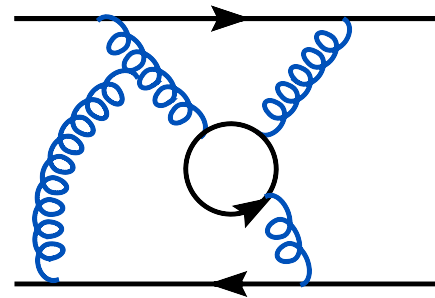
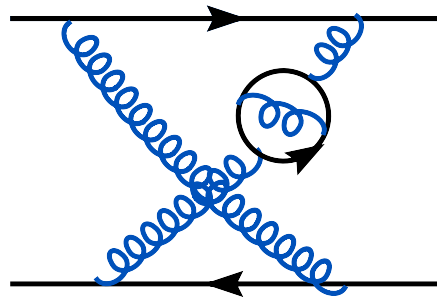
Two loops:

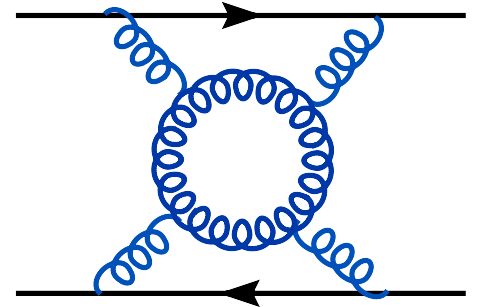
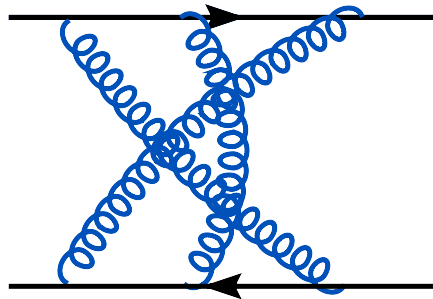
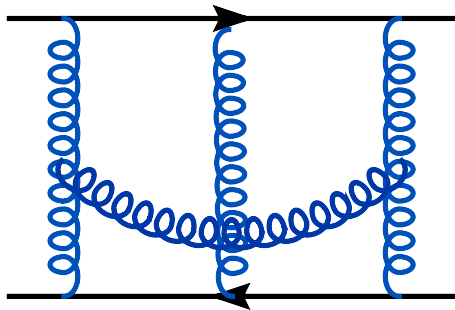


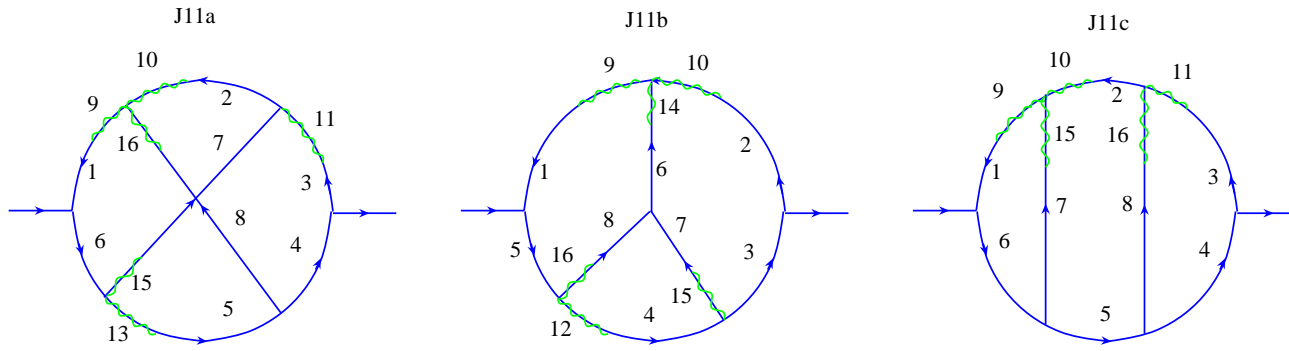
+ ...



Generation of diagrams by QGRAPH





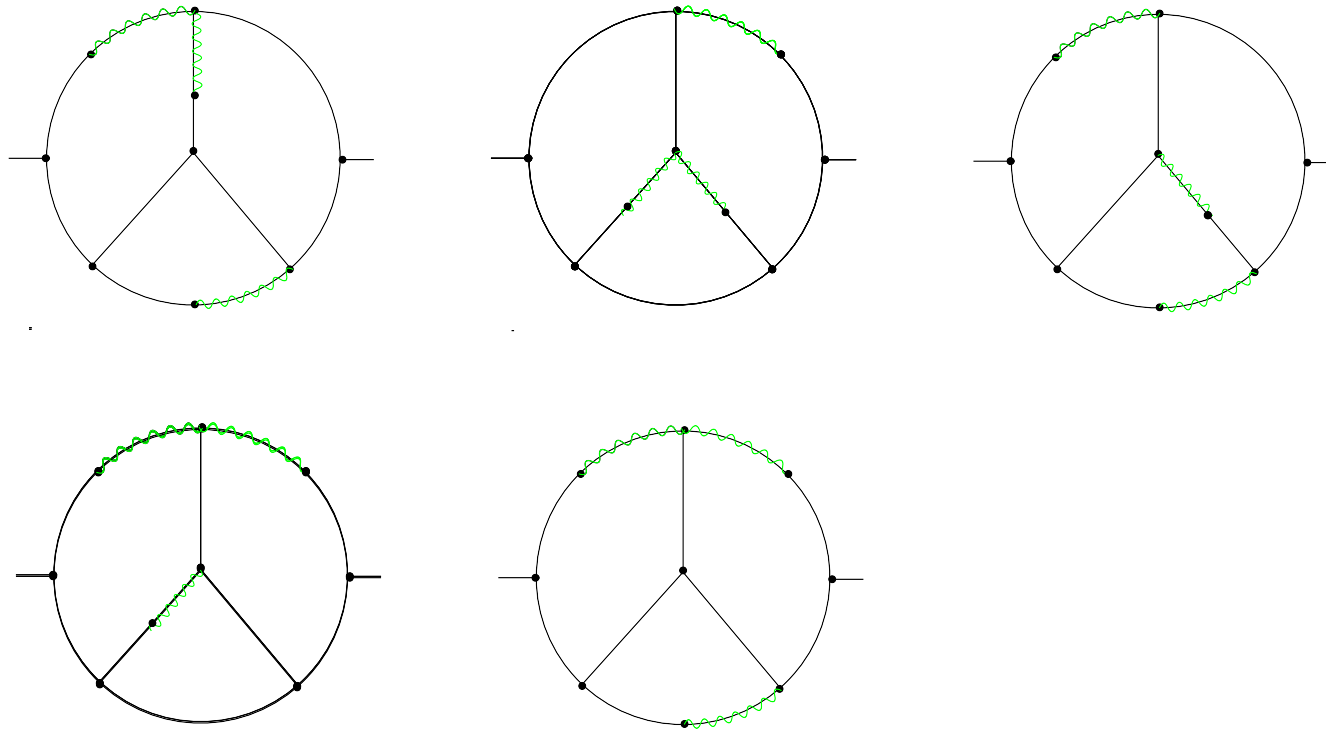


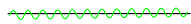
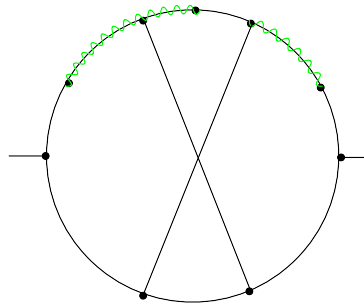
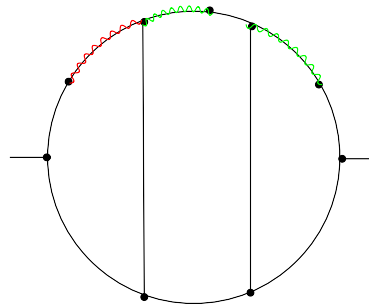
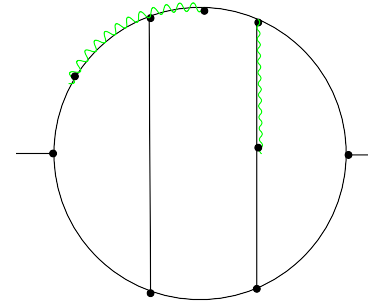
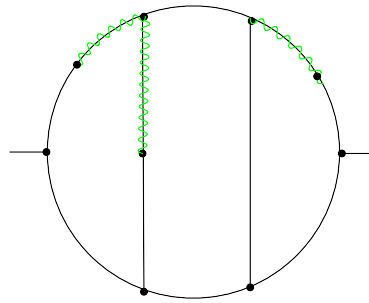
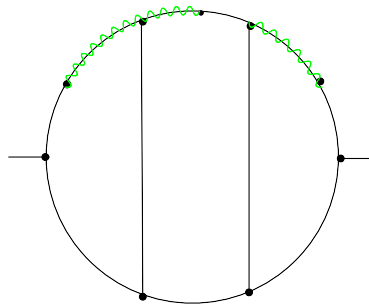
$$\begin{aligned}
& J_{11}^a(n_1, n_2, n_3, n_4, n_5, n_6, n_7, n_8; n_0; n_9, n_{10}, n_{11}, n_{13}, n_{15}, n_{16}; s_9, s_{10}, s_{11}, s_{13}, s_{15}, s_{16}) \\
&= \iiint \frac{(-p_0^2)^{-n_0} dk dl dr}{(-p_1^2)^{n_1} (-p_2^2)^{n_2} (-p_3^2)^{n_3} (-p_4^2)^{n_4} (-p_5^2)^{n_5} (-p_6^2)^{n_6} (-p_7^2)^{n_7} (-p_8^2)^{n_8}} \\
&\quad \times \frac{1}{(-v \cdot p_1 - s_9 i 0)^{n_9} (-v \cdot p_2 - s_{10} i 0)^{n_{10}} (-v \cdot p_3 - s_{11} i 0)^{n_{11}} (-v \cdot p_5 - s_{13} i 0)^{n_{13}} (-v \cdot p_7 - s_{15} i 0)^{n_{15}} (-v \cdot p_8 - s_{16} i 0)^{n_{16}}} \\
&\quad p_0 = k - r, p_1 = k, p_2 = l, p_3 = r, p_4 = q + r, p_5 = q + k - l + r, p_6 = q + k, p_7 = l - r, p_8 = k - l
\end{aligned}$$

$$\begin{aligned}
& J_{11}^b(n_1, n_2, n_3, n_4, n_5, n_6, n_7, n_8; n_0; n_9, n_{10}, n_{12}, n_{14}, n_{15}, n_{16}; s_9, s_{10}, s_{12}, s_{14}, s_{15}, s_{16}) \\
&= \iiint \frac{(-p_0^2)^{-n_0} dk dl dr}{(-p_1^2)^{n_1} (-p_2^2)^{n_2} (-p_3^2)^{n_3} (-p_4^2)^{n_4} (-p_5^2)^{n_5} (-p_6^2)^{n_6} (-p_7^2)^{n_7} (-p_8^2)^{n_8}} \\
&\quad \times \frac{1}{(-v \cdot p_1 - s_9 i 0)^{n_9} (-v \cdot p_2 - s_{10} i 0)^{n_{10}} (-v \cdot p_4 - s_{12} i 0)^{n_{12}} (-v \cdot p_6 - s_{14} i 0)^{n_{14}} (-v \cdot p_7 - s_{15} i 0)^{n_{15}} (-v \cdot p_8 - s_{16} i 0)^{n_{16}}} \\
&\quad p_0 = l - q, p_1 = k - q, p_2 = r - q, p_3 = r, p_4 = l, p_5 = k, p_6 = k - r, p_7 = l - r, p_8 = k - l
\end{aligned}$$

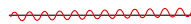
$$\begin{aligned}
& J_{11}^c(n_1, n_2, n_3, n_4, n_5, n_6, n_7, n_8; n_0; n_9, n_{10}, n_{11}, n_{15}, n_{16}; s_9, s_{10}, s_{11}, s_{15}, s_{16}) \\
&= \iiint \frac{(-p_0^2)^{-n_0} dk dl dr}{(-p_1^2)^{n_1} (-p_2^2)^{n_2} (-p_3^2)^{n_3} (-p_4^2)^{n_4} (-p_5^2)^{n_5} (-p_6^2)^{n_6} (-p_7^2)^{n_7} (-p_8^2)^{n_8}} \\
&\quad \times \frac{1}{(-v \cdot p_1 - s_9 i 0)^{n_9} (-v \cdot p_2 - s_{10} i 0)^{n_{10}} (-v \cdot p_3 - s_{11} i 0)^{n_{11}} (-v \cdot p_7 - s_{15} i 0)^{n_{15}} (-v \cdot p_8 - s_{16} i 0)^{n_{16}}} \\
&\quad p_0 = k - r, p_1 = k, p_2 = l, p_3 = r, p_4 = q + r, p_5 = q + l, p_6 = q + k, p_7 = k - l, p_8 = l - r
\end{aligned}$$

Various classes of Feynman integrals with 12 indices:





$$\frac{1}{(-v \cdot k - i0)^a}$$

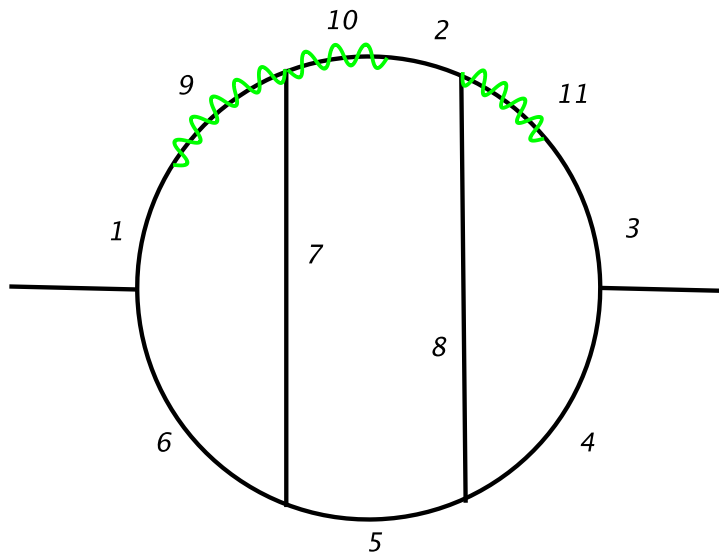


$$\frac{1}{(-v \cdot k + i0)^a}$$



$$\frac{1}{(-k^2 - i0)^a}$$

For example,



with the numerator chosen as $-(k - r)^2$ ^{-a₁₂}

Our techniques

IBP [K.G. Chetyrkin & F.V. Tkachov'81]

The whole problem of the evaluation of a given family of Feynman integrals \rightarrow

- constructing a reduction procedure using IBP
- evaluating master integrals

Solving reduction problems algorithmically:

- Laporta's algorithm

[S. Laporta & E. Remiddi'96; S. Laporta'00; T. Gehrman & E. Remiddi'01]

A public version:

AIR

[C. Anastasiou & A. Lazopoulos'04]

Private versions

[T. Gehrman & E. Remiddi, M. Czakon, Y. Schröder, A. Pak, C. Sturm, P. Marquard & D. Seidel, V. Velizhanin, ...]

- Baikov's method

[Baikov'96-09]

- Gröbner bases

[O.V. Tarasov'98, A.V. Smirnov & V.A. Smirnov'05]

- Lee's approach

[R.N. Lee'08]

A combination of Laporta's algorithm and Gröbner bases:
FIRE = Feynman Integrals REduction [A.V. Smirnov'08]
(implemented in `Mathematica`)

<http://science.sander.su>

or

<http://www-ttp.particle.uni-karlsruhe.de/~asmirnov>

Evaluating master integrals

~ 100 master integrals in the whole problem

Mellin-Barnes representation

[V.A. Smirnov'99, J.B Tausk'99]

$$\frac{1}{(X+Y)^\lambda} = \frac{1}{\Gamma(\lambda)} \frac{1}{2\pi i} \int_{-i\infty}^{+i\infty} dz \frac{Y^z}{X^{\lambda+z}} \Gamma(\lambda+z) \Gamma(-z) .$$

- Derive an MB representation
- Resolve the singularity structure in ϵ
- Expand in a Laurent series in ϵ and evaluate expanded MB integrals

Evaluating Feynman Integrals (STMP 211);

Feynman Integrals Calculus (Springer 2006)

For planar diagrams, **AMBRE** can be used

[J. Gluza, K. Kajda & T. Riemann'07]

MB tools at <http://projects.hepforge.org/mbtools/>:

MB.m [M. Czakon]

[M. Czakon'05]

MBresolve.m [A. Smirnov]

[A. & V. Smirnovs'09]

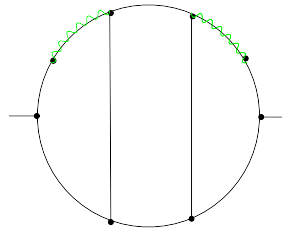
MBasymptotics.m [M. Czakon]

barnesroutines.m [D. Kosower]

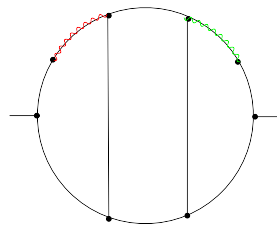
(applying Barnes lemmas automatically)

$$\frac{1}{2\pi i} \int_{-i\infty}^{+i\infty} dz \Gamma(\lambda_1 + z) \Gamma(\lambda_2 + z) \Gamma(\lambda_3 - z) \Gamma(\lambda_4 - z)$$
$$= \frac{\Gamma(\lambda_1 + \lambda_3) \Gamma(\lambda_1 + \lambda_4) \Gamma(\lambda_2 + \lambda_3) \Gamma(\lambda_2 + \lambda_4)}{\Gamma(\lambda_1 + \lambda_2 + \lambda_3 + \lambda_4)}$$

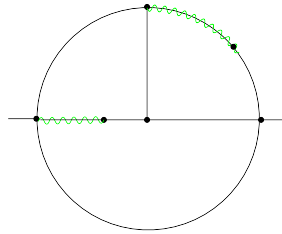
For example,



$$\frac{(i\pi^{d/2})^3}{(q^2)^3 v^2} \left[\frac{56\pi^4}{135\epsilon} + \frac{112\pi^4}{135} + \frac{16\pi^2\zeta(3)}{9} + \frac{8\zeta(5)}{3} + O(\epsilon) \right]$$



$$\frac{(i\pi^{d/2})^3}{(q^2)^3 v^2} \left[-\frac{64\pi^4}{135\epsilon} - \frac{128\pi^4}{135} - \frac{32\pi^2\zeta(3)}{9} + \frac{8\zeta(5)}{3} + O(\epsilon) \right]$$



$$\frac{(i\pi^{d/2})^3}{q^2 v^2} \left[\frac{32\pi^4}{135\epsilon} - \frac{128\pi^4}{135} + \frac{88\pi^2\zeta(3)}{9} + \frac{188\zeta(5)}{3} + O(\epsilon) \right]$$

Transcendentality level 5 (and, sometimes, 6) is needed.
The constants of level 6 that we encounter:

$$\zeta(6), \zeta(4) \ln^2 2, \zeta(3)^2, \zeta(3)\zeta(2) \ln 2, \zeta(2) \ln^4 2, \\ \ln^6 2, \text{Li}_6\left(\frac{1}{2}\right), \text{Li}_4\left(\frac{1}{2}\right) \ln^2 2, \dots s_6 = S(\{-5, -1\}, \infty)$$

PSLQ

[H.R.P. Ferguson & D.H. Bailey'91]

Only three constants are not known analytically.

Numerical checks by FIESTA [A.V. Smirnov & M.N. Tentyukov'08]

(Feynman Integral Evaluation by a Sector decomposition Approach)

A tool to evaluate Feynman integrals numerically by sector decompositions.

[T. Binoth & G. Heinrich'00]

Strategies that are guaranteed to terminate

[C. Bogner & S. Weinzierl'07]

A, B, C, X

Strategy S

[A.V. Smirnov & M.N. Tentyukov'08]

<http://science.sander.su>

<http://www-ttp.particle.uni-karlsruhe.de/~asmirnov>

FIESTA 2

[A. & V. Smirnovs & M.N. Tentyukov'09]

- Expanding integrals automatically in limits of momenta and masses (with the use of Mellin–Barnes representation)

[M.C. Bergère, C. de Calan and A.P.C. Malbouisson,'78; K. Pohlmeier,'82; V. Pilipp'08]

- complete parallelization (even to multiple computers)
- high-precision arithmetics
- new integrators
- Speer sectors as a strategy

[A. & V. Smirnovs'09]

Results

$$a_3 = a_3^{(3)} n_l^3 + a_3^{(2)} n_l^2 + a_3^{(1)} n_l + a_3^{(0)},$$

$$a_3^{(3)}, a_3^{(2)}, a_3^{(1)}$$

[A.V. Smirnov, V.A. Smirnov, and M. Steinhauser'08]

$$a_3^{(3)} = - \left(\frac{20}{9} \right)^3 T_F^3,$$

$$a_3^{(2)} = \left(\frac{12541}{243} + \frac{368\zeta(3)}{3} + \frac{64\pi^4}{135} \right) C_A T_F^2,$$
$$+ \left(\frac{14002}{81} - \frac{416\zeta(3)}{3} \right) C_F T_F^2,$$

$$\begin{aligned}
a_3^{(1)} = & (-709.717) C_A^2 T_F + \left(-\frac{71281}{162} + 264\zeta(3) + 80\zeta(5) \right) C_A C_F T_F \\
& + \left(\frac{286}{9} + \frac{296\zeta(3)}{3} - 160\zeta(5) \right) C_F^2 T_F + (-56.83(1)) \frac{d_F^{abcd} d_F^{abcd}}{N_A},
\end{aligned}$$

$$C_A = N_c, \quad C_F = \frac{N_c^2 - 1}{2N_c}, \quad T_F = \frac{1}{2},$$

$$\frac{d_F^{abcd} d_F^{abcd}}{N_A} = \frac{18 - 6N_c^2 + N_c^4}{96N_c^2}.$$

UV poles in ϵ are removed by renormalization.

The remaining simple pole is canceled by the pole of the US contribution:

$$\frac{4\pi C_F \alpha_s}{|\vec{q}|^2} \left(\frac{\alpha_s}{4\pi}\right)^3 16\pi^2 C_A^3 \left(\frac{1}{6\epsilon} + \dots\right) + \dots$$

$$a_3^{(0)} = (\dots) C_A^3 + (-136.39(6)) \frac{d_F^{abcd} d_A^{abcd}}{N_A},$$

$$\frac{d_F^{abcd} d_A^{abcd}}{N_A} = \frac{N_c^3 + 6N_c}{48}$$

An essential cancelation of the n_l^1 and n_l^0 contributions is expected.

Some additional checks are in progress.

to be completed