

The Collider-Cosmology Interface III

M.J. Ramsey-Musolf

U Mass Amherst



AMHERST CENTER FOR FUNDAMENTAL INTERACTIONS

Physics at the interface: Energy, Intensity, and Cosmic frontiers

University of Massachusetts Amherst

<http://www.physics.umass.edu/acfi/>

HEP School, Lanzhou
8/1-8/18

Lecture III Goals

- *Introduce the key ideas for relating the electroweak phase transition and WIMP dark matter in the context of electroweak baryogenesis*
- *Provide general considerations regarding dark matter model building*
- *Explain how electroweak baryogenesis works*
- *Discuss simple EWPT-DM scenarios and their collider probes*
- *Invite questions !*

Lecture III Outline

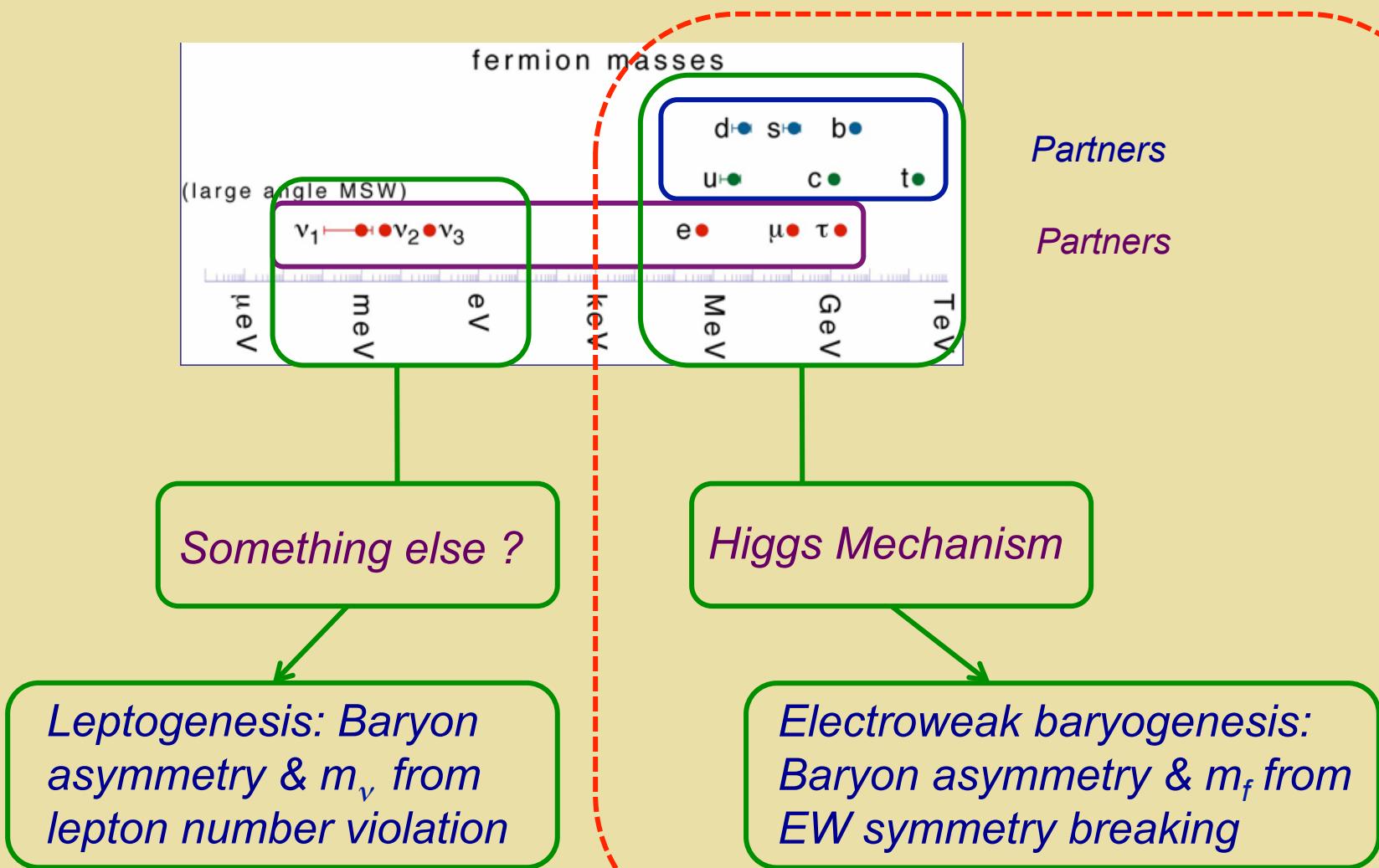
- I. *Electroweak Phase Transition & Dark Matter: Overview*
- II. *WIMP Dark Matter: Context & Properties*
- III. *Electroweak Baryogenesis: How It Works*
- IV. *EWPT & DM: Scenarios & Collider Probes*

Selected References

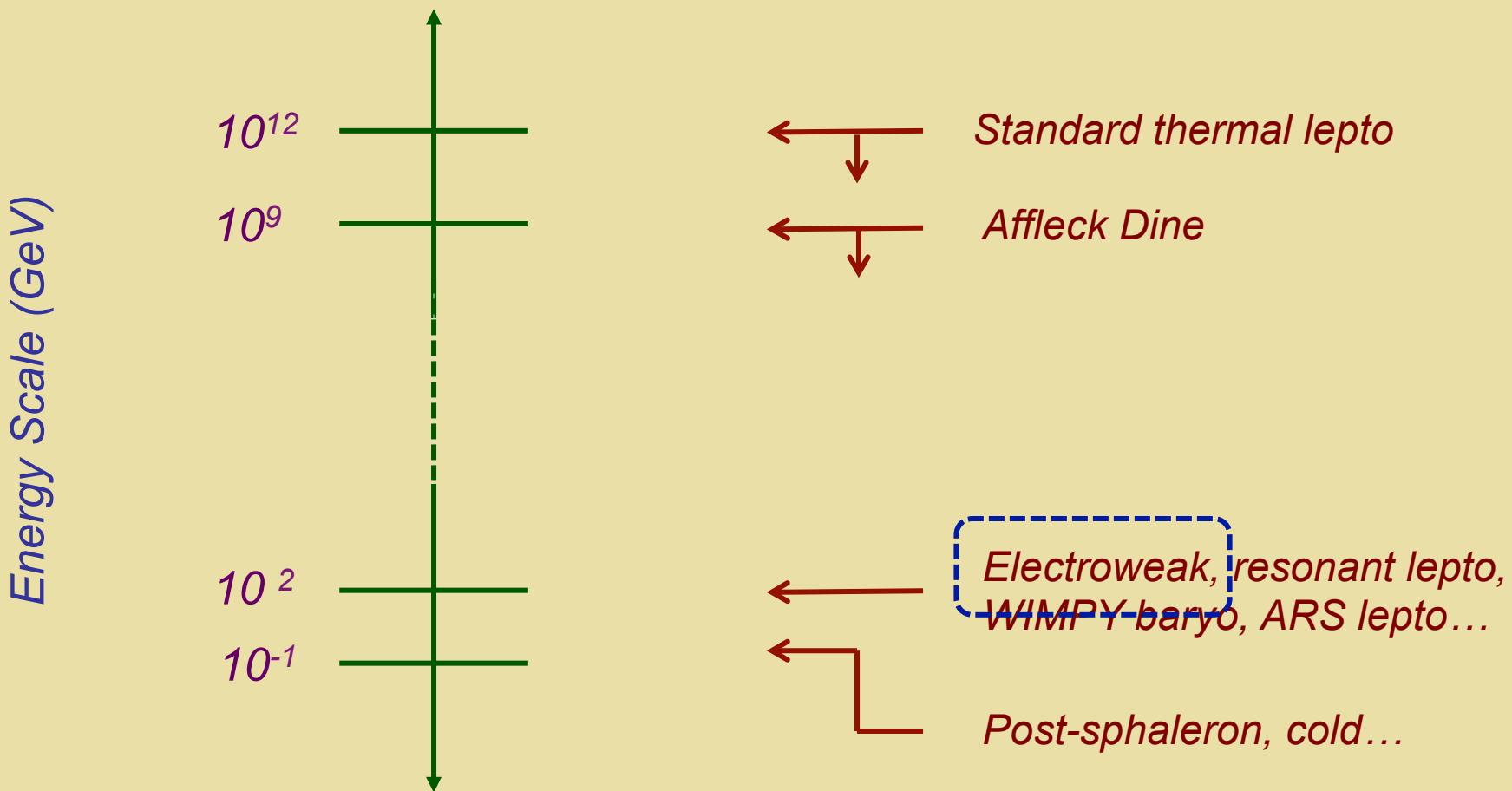
- *M. Quiros, hep-ph/9901312*
- *S. Profumo, MJRM, G. Shaughnessy, 0705.2425 [hep-ph]*
- *H. Patel & MJRM, 1101.4665 [hep-ph]*
- *D. Morrissey & MJRM, 1206.2941 [hep-ph]*
- *K. Asamagan et al, 1604.05324 [hep-ph]*

I. EWPT & Dark Matter: Overview

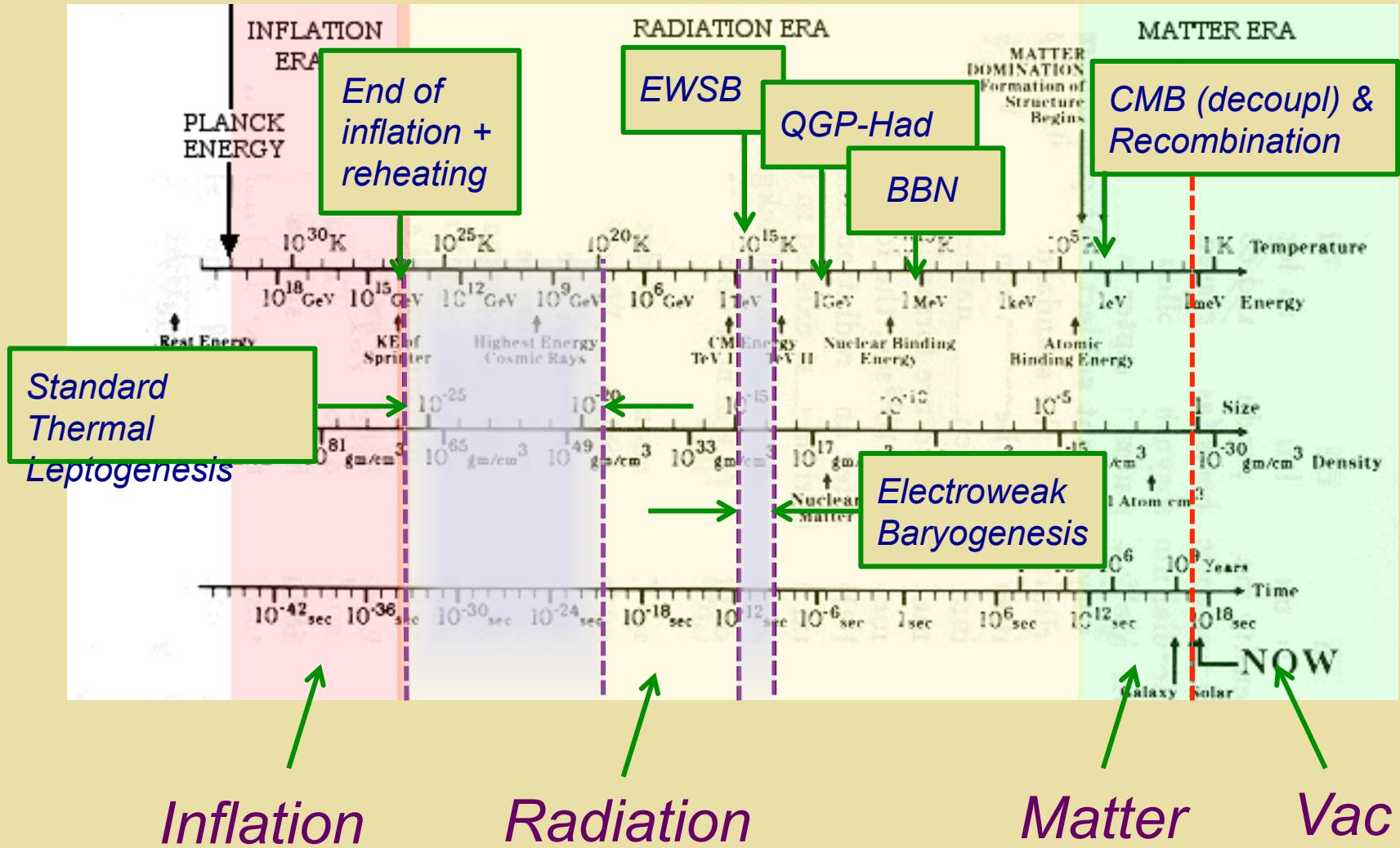
Fermion Masses & Baryon Asymmetry



Baryogenesis Scenarios



Thermal History



Ingredients for Baryogenesis



Scenarios: leptogenesis,
EW baryogenesis, Afflek-
Dine, asymmetric DM, cold
baryogenesis, post-
sphaleron baryogenesis...

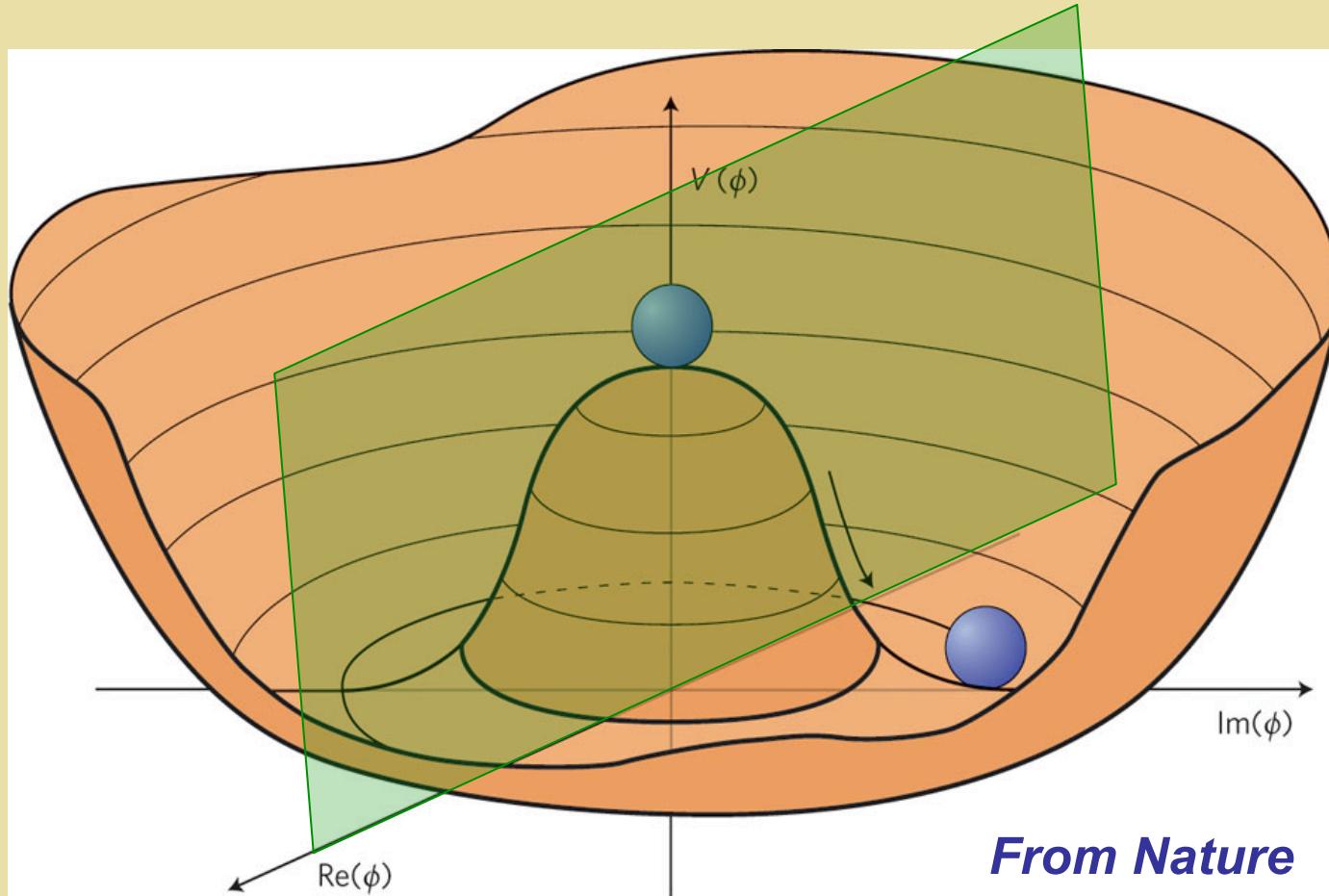
- B violation (sphalerons)
- C & CP violation
- Out-of-equilibrium or CPT violation

Standard Model BSM

✓	✓
✗	✓
✗	✓

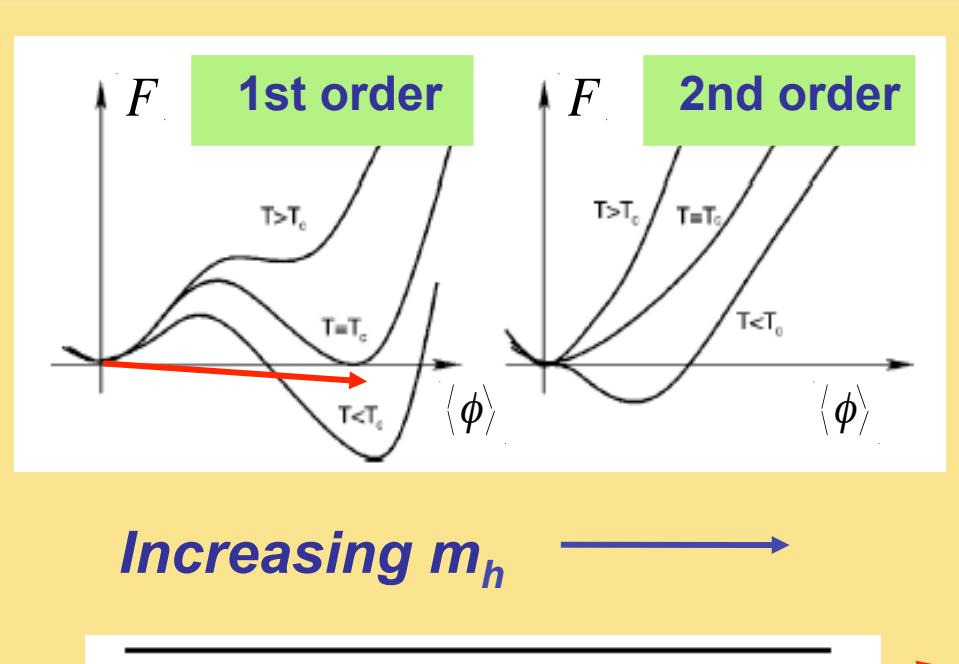
First order EWPT

EWSB: The Scalar Potential

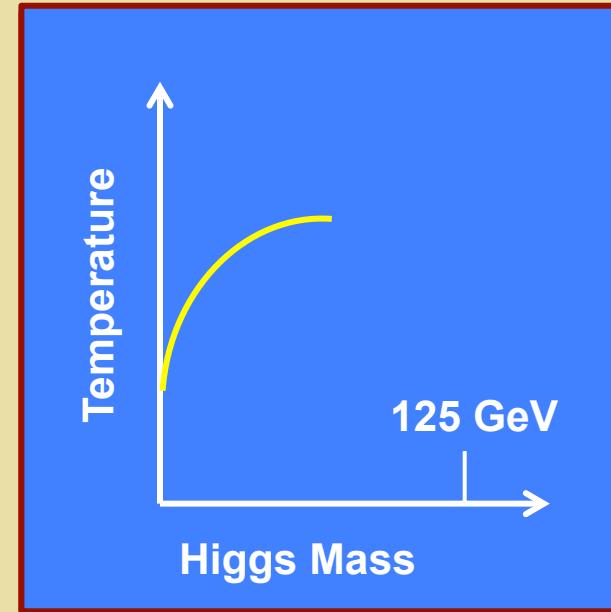


What was the thermal history of EWSB ?

EW Phase Transition: St'd Model



Lattice	Authors	M_h^C (GeV)
4D Isotropic	[76]	80 ± 7
4D Anisotropic	[74]	72.4 ± 1.7
3D Isotropic	[72]	72.3 ± 0.7
3D Isotropic	[70]	72.4 ± 0.9



EW Phase Diagram

How does this picture change in presence of new TeV scale physics ? What is the phase diagram ?

SM EW: Cross over transition

First Order EWPT from BSM Physics

- *Thermal loops involving new bosons*
- *T=0 loops (CW Potential)*
- *Change tree-level vacuum structure*

EWPT & Dark Matter

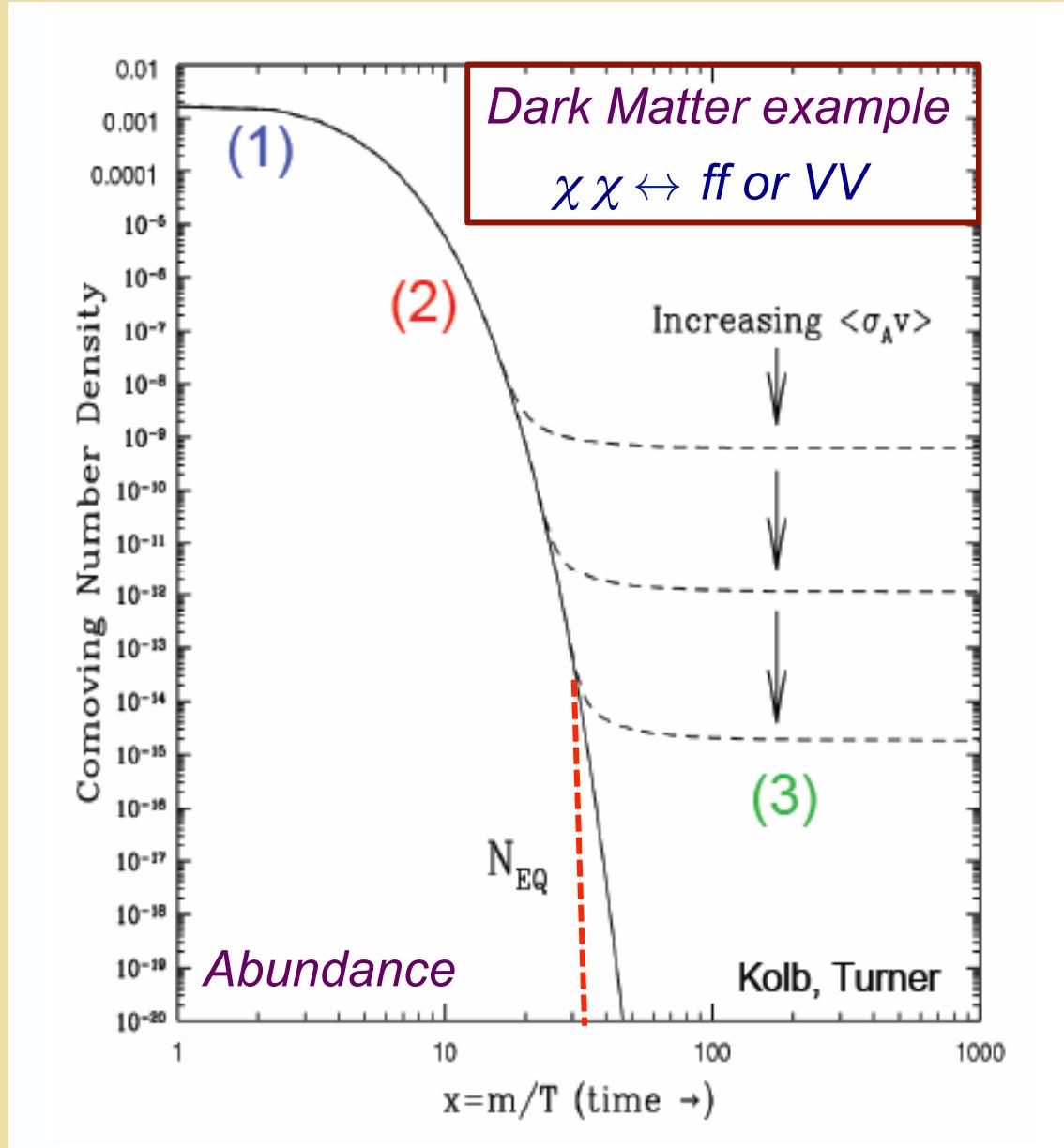
- *Can the BSM particles that catalyze a first order EWPT also be viable dark matter candidates ?*

II. WIMP Dark Matter: Context & Properties

Dark Matter Properties

- *Electrically neutral & colorless*
- *Stable on cosmological time scales*
- “*Cold*”
- *Thermal or Non-Thermal*

Thermal Dark Matter



Boltzmann Eqs:

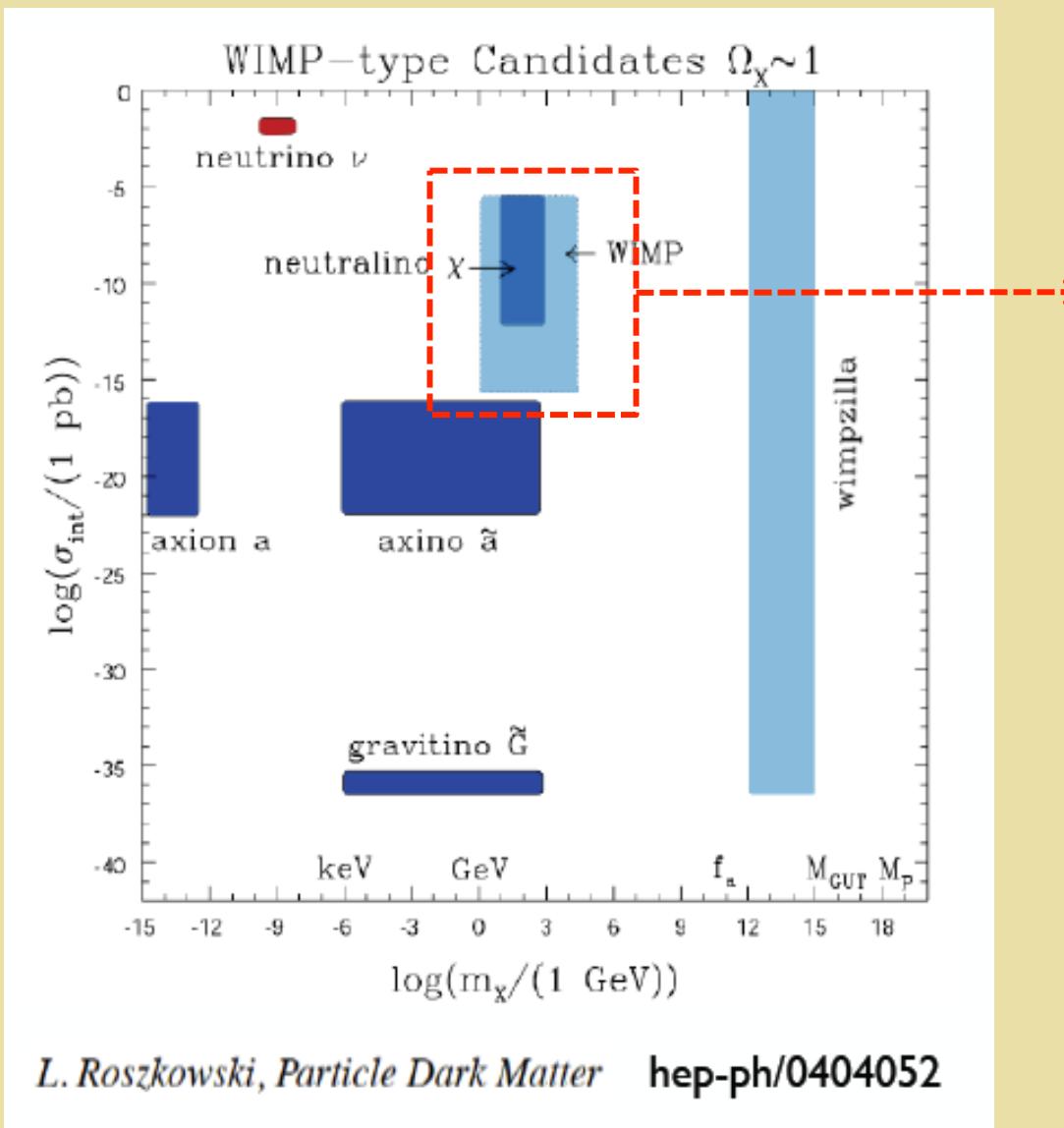
- 1) $N \sim N_{EQ}$
- 2) N starts to depart from N_{EQ}
- 3) N “freezes out” at x_f

$$x_f \sim \mathcal{O}(10) \rightarrow T \sim m / 10$$

Dark Matter Interactions

- *Gravitational*
- *Non-grav interactions w/ St'd Model ?*
- *Non-grav interactions w/ itself ?*

Particle Dark Matter Scenarios



*Weakly Interacting
Massive Particles:
“WIMPS”*

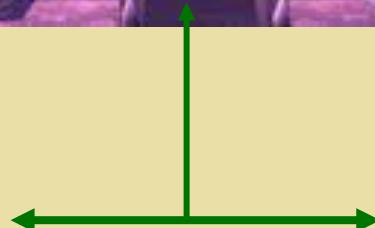
- $\sigma_{\text{INT}} \sim \sigma_{\text{weak}}$
- M_χ : few GeV \rightarrow few TeV

Dark Matter Portals



Standard Model

Hidden Sector

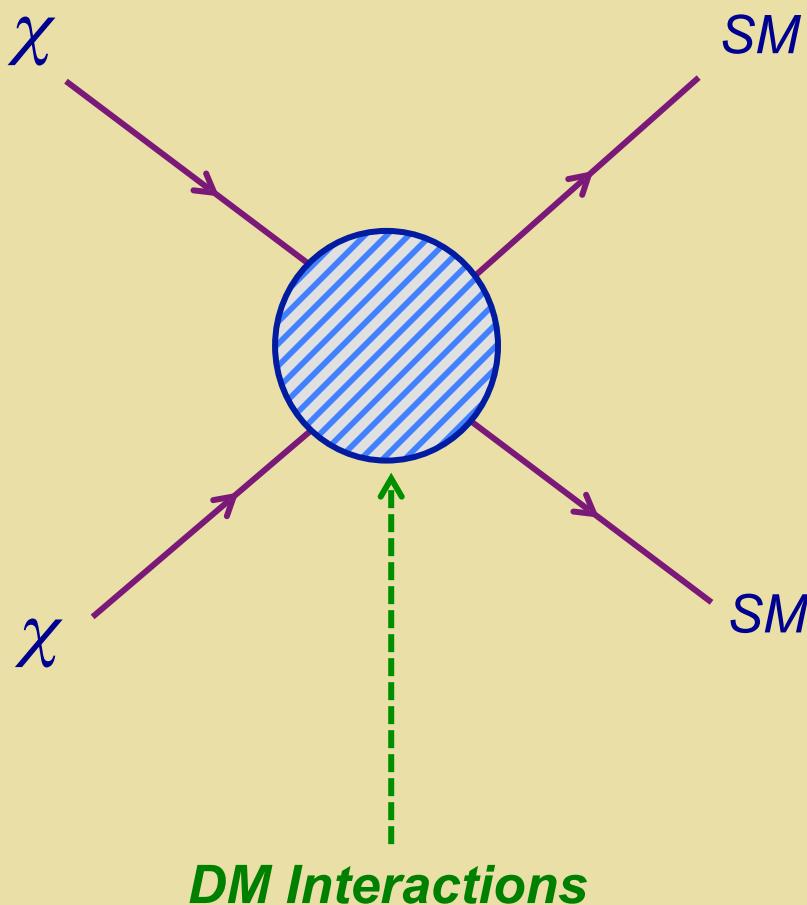


Dark Matter Portals

- *Gauge sector (SUSY neutralinos)*
- *Higgs portal (BSM scalars)*
- *QCD portal (Axion)*
- *Yukawa portal (neutrinos)*

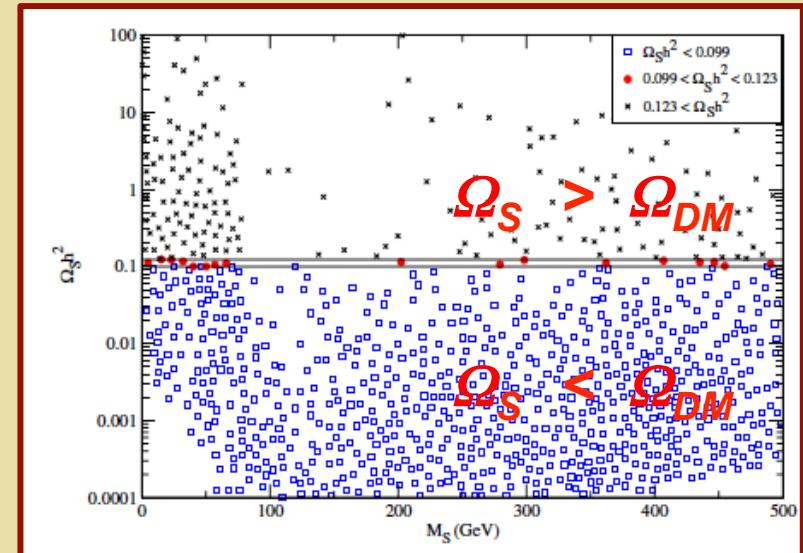
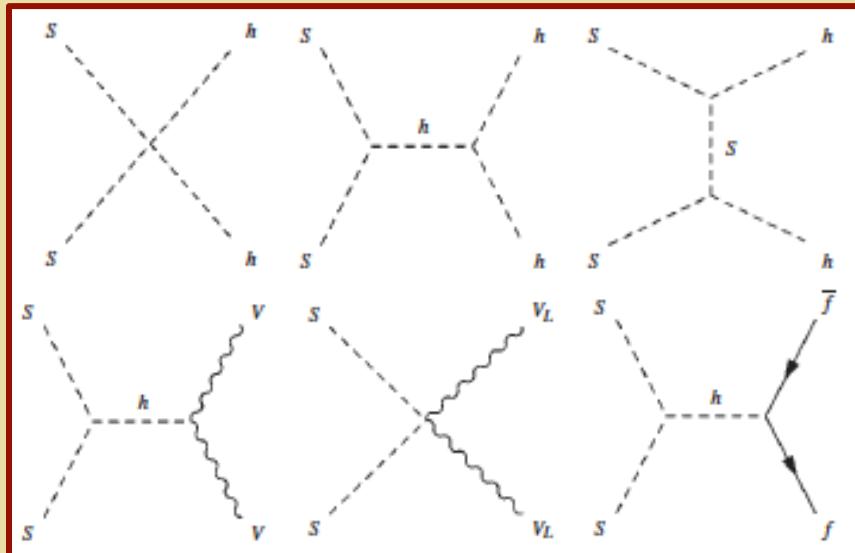
WIMP Dark Matter Probes

*Thermal Abundance & Indirect
Detection*



WIMP DM: Ω_{DM} & Indirect Det

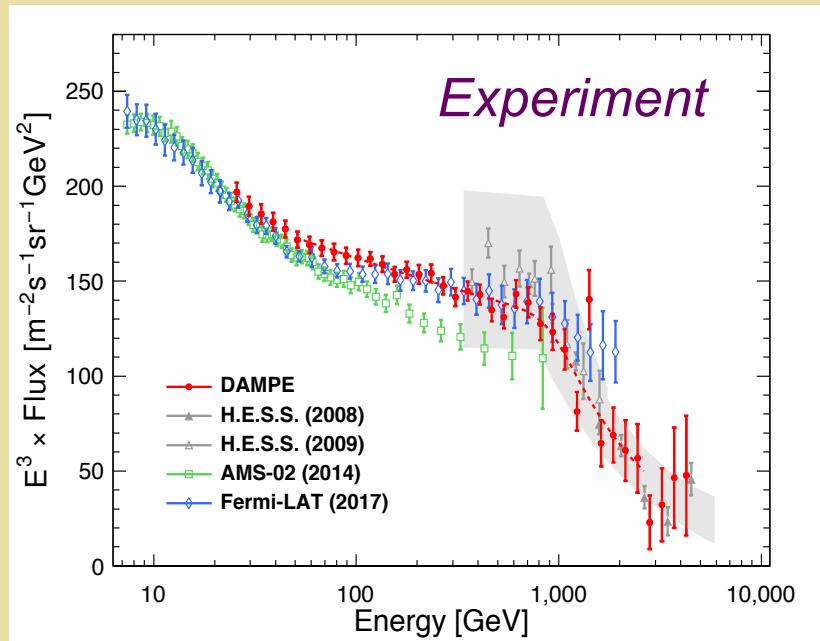
Real singlet extension: “xSM”



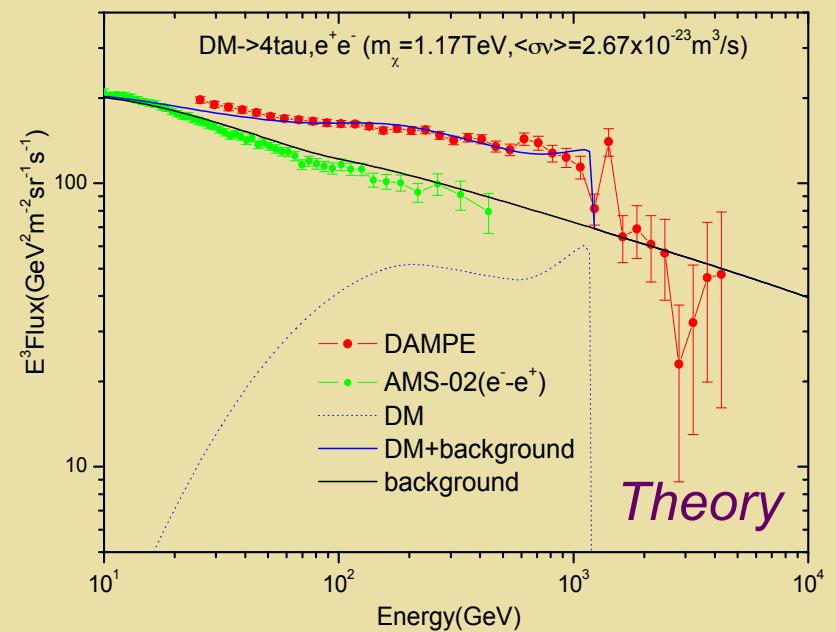
*Barger, Langacker, McCaskey, MJRM,
Shaugnessy 0706.4311 [hep-ph]*

WIMP DM: Ω_{DM} & Indirect Det

Indirect detection



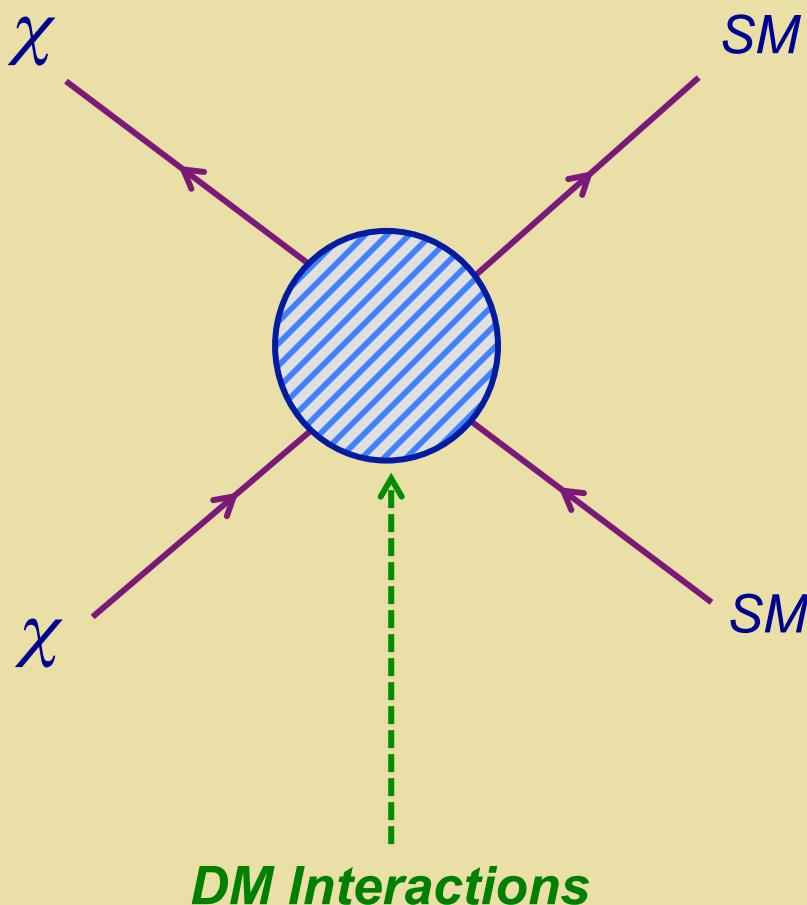
Ambrosi et al (DAMPE Collab)
1711.10981 [astro-ph]



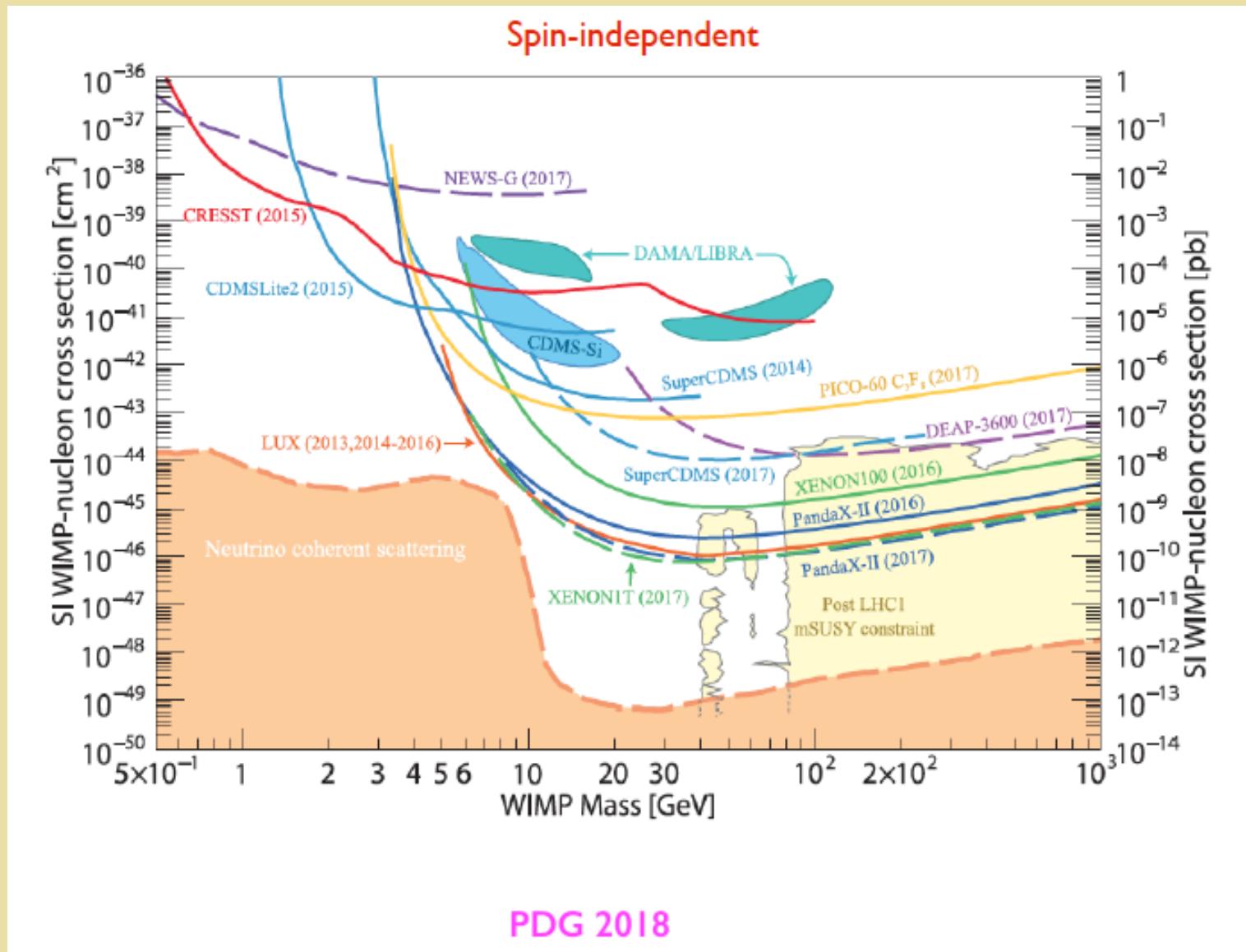
Jin, Yue, Zhang, Chen
1712.00362 [astro-ph]

WIMP Dark Matter Probes

Direct Detection: WIMP-Nucleus scattering

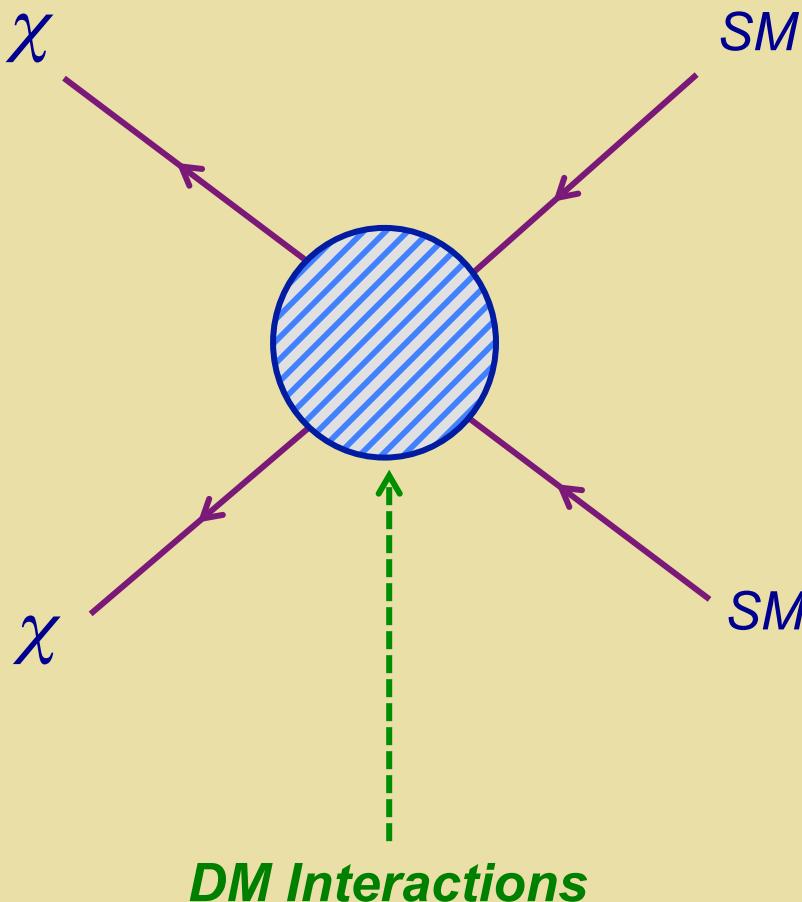


WIMP Dark Matter: Direct Detection



WIMP Dark Matter Probes

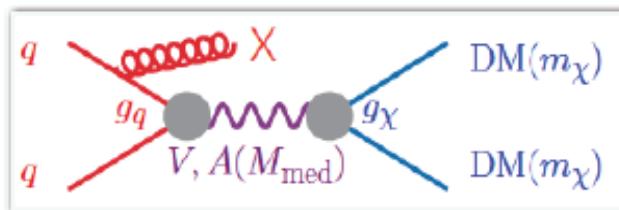
DM Production at Colliders



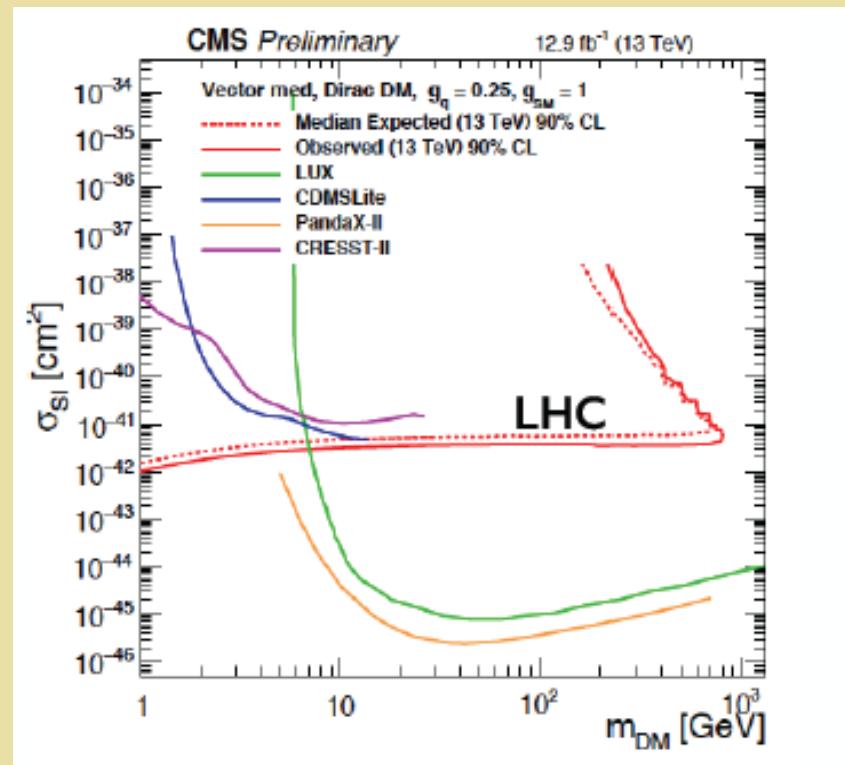
WIMP Dark Matter: Colliders

Signature: $p\bar{p} \rightarrow X + \text{missing } E_T$ (“mono-X” searches)

V. Cirigliano

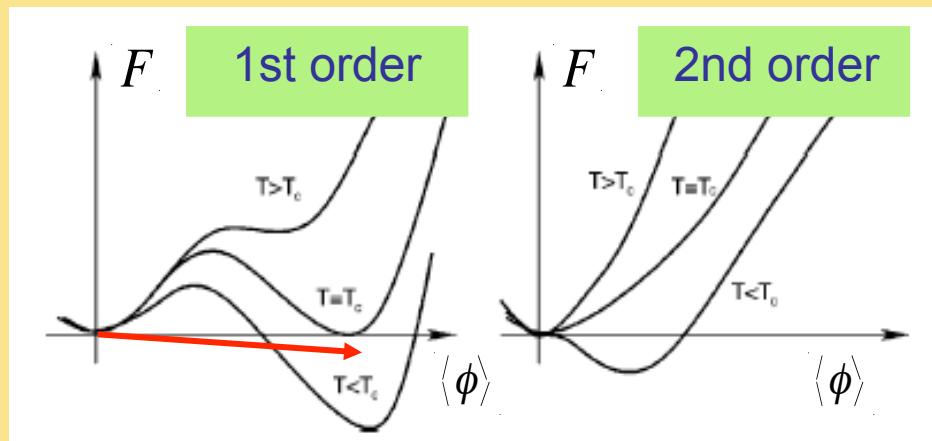


Irreducible background:
 $p\bar{p} \rightarrow Z(v\bar{v}) + \text{jets}$

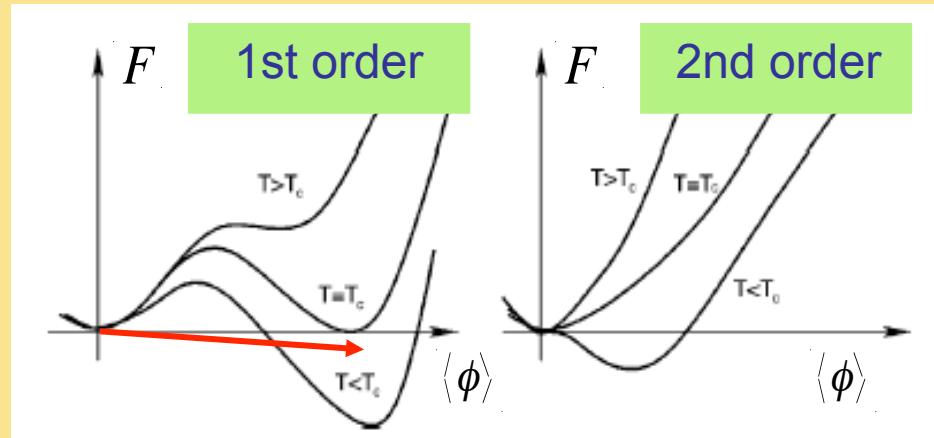


III. Electroweak Baryogenesis: How It Works

EW Phase Transition: New Scalars & CPV

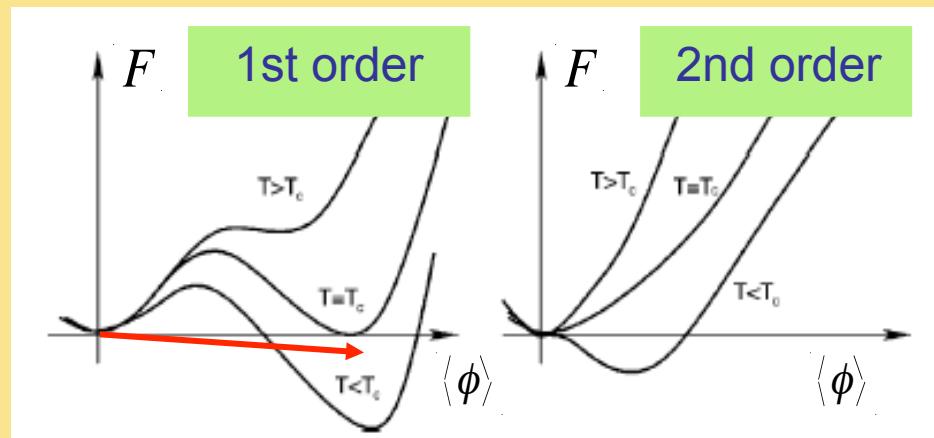


EW Phase Transition: New Scalars & CPV



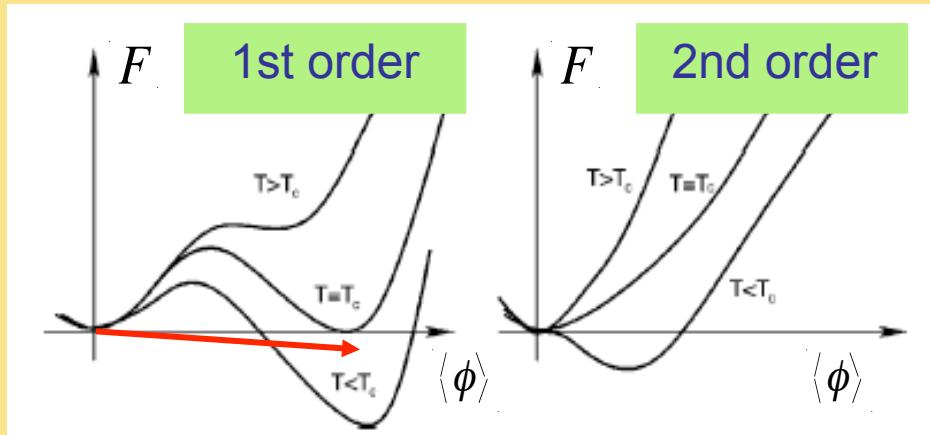
Increasing m_h \longrightarrow

EW Phase Transition: New Scalars & CPV



Increasing m_h \longrightarrow
 \longleftarrow New scalars

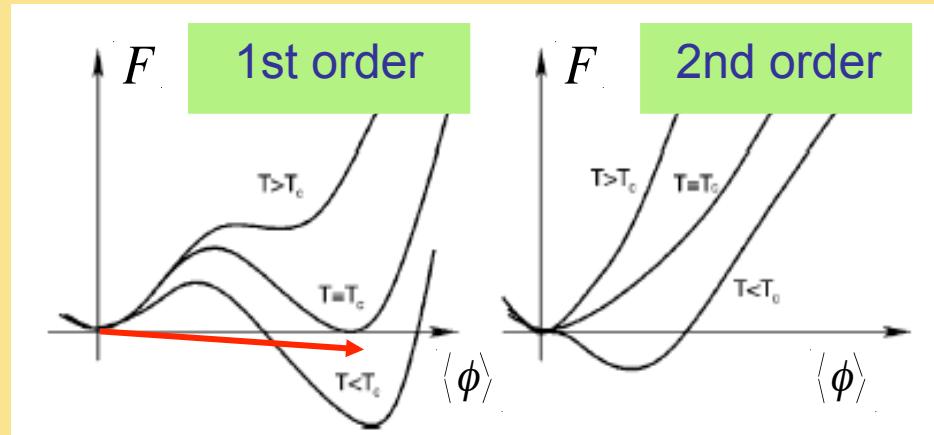
EW Phase Transition: New Scalars & CPV



Increasing m_h \longrightarrow
 \longleftarrow New scalars

- *Loop effects*
- *Tree-level barrier*

EW Phase Transition: New Scalars & CPV

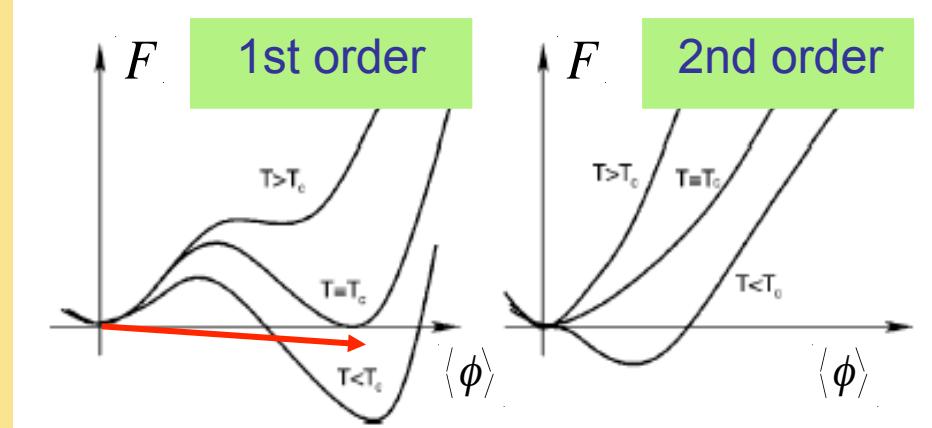


Increasing m_h \longrightarrow
 \longleftarrow New scalars

Baryogenesis
Gravity Waves
Scalar DM
LHC Searches

"Strong" 1st order EWPT

EW Phase Transition: New Scalars & CPV



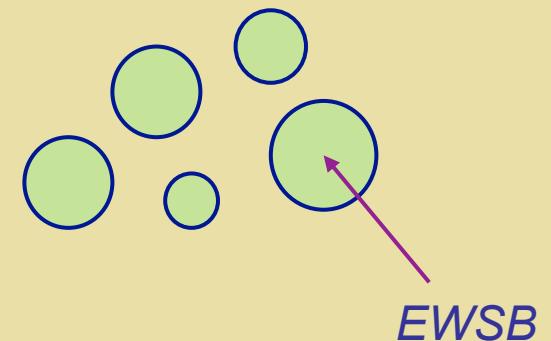
Increasing m_h \longrightarrow

\longleftarrow New scalars

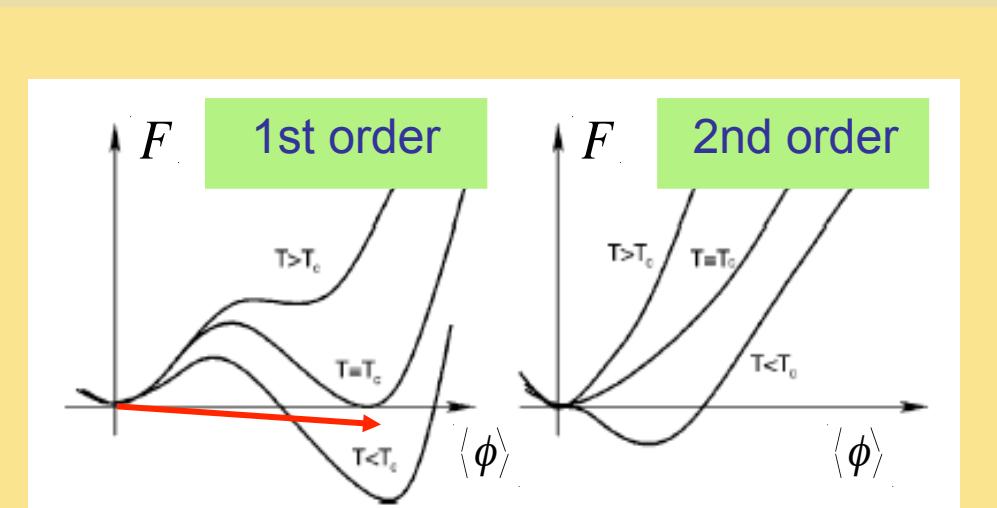
Baryogenesis
Gravity Waves
Scalar DM
LHC Searches

“Strong” 1st order EWPT

Bubble nucleation

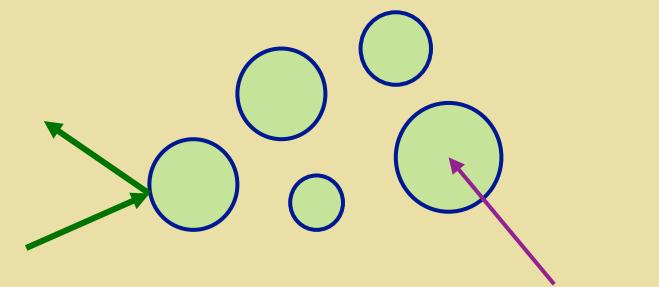
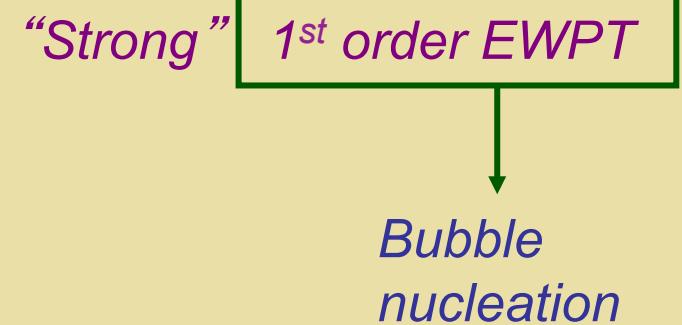


EW Phase Transition: New Scalars & CPV



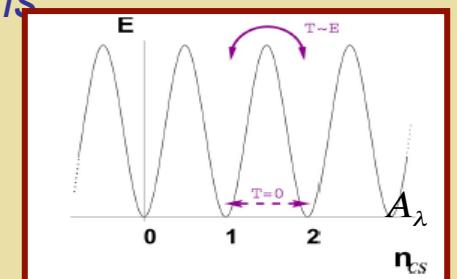
Increasing m_h \longrightarrow
 \longleftarrow New scalars

Baryogenesis
 Gravity Waves
 Scalar DM
 LHC Searches

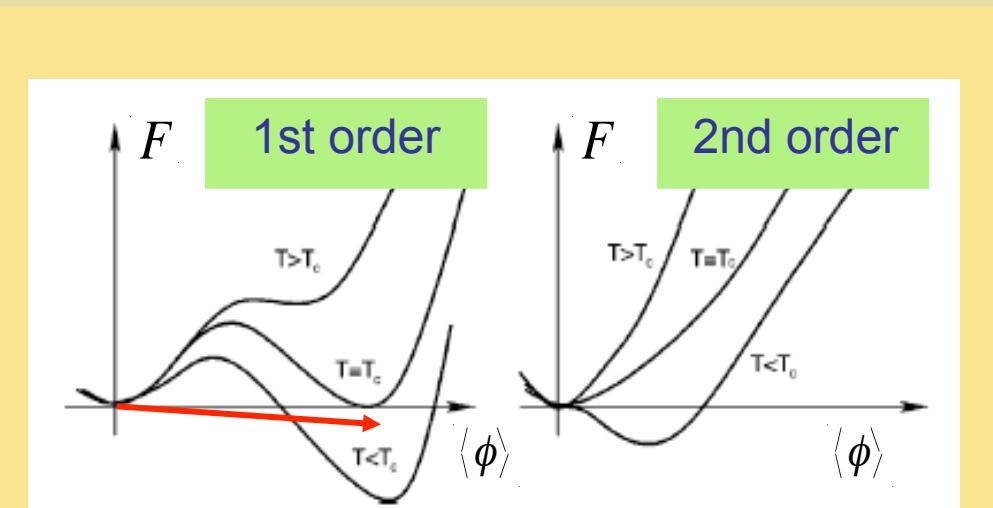


Y_B : CPV &
 EW sphalerons

EWSB

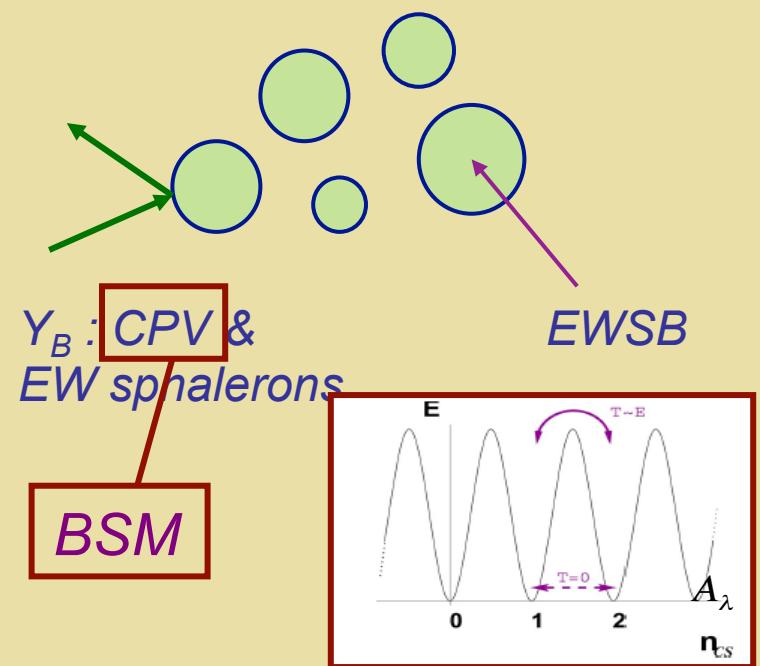
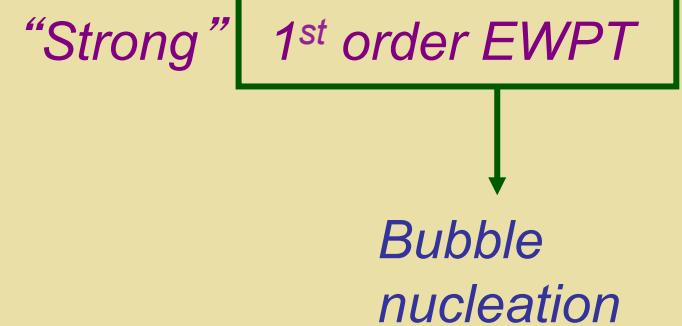


EW Phase Transition: New Scalars & CPV

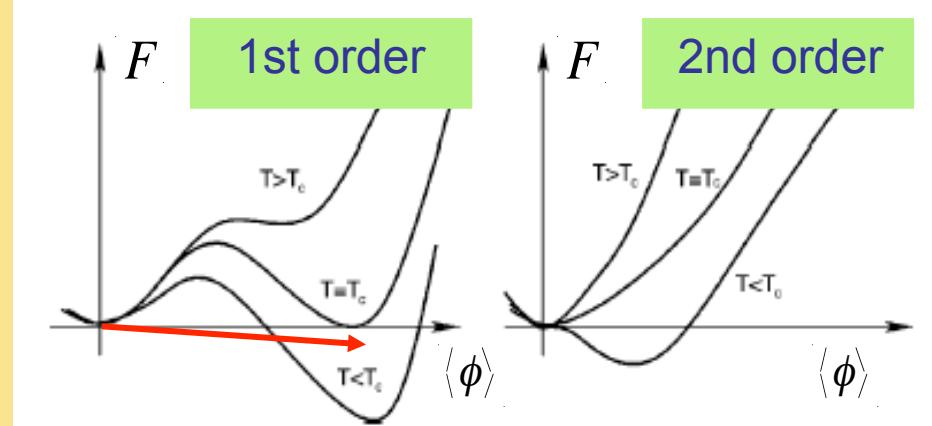


Increasing m_h → ← New scalars

*Baryogenesis
Gravity Waves
Scalar DM
LHC Searches*



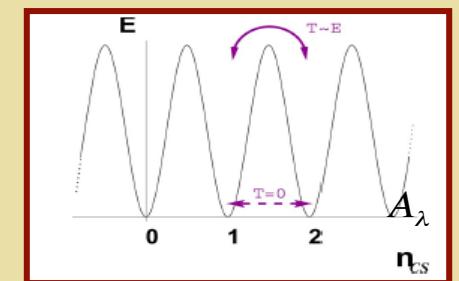
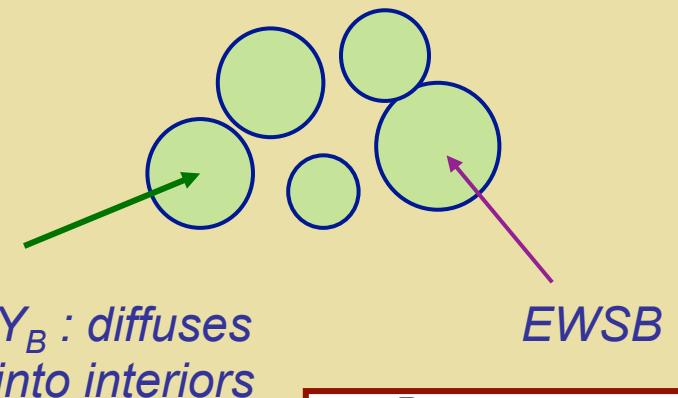
EW Phase Transition: New Scalars & CPV



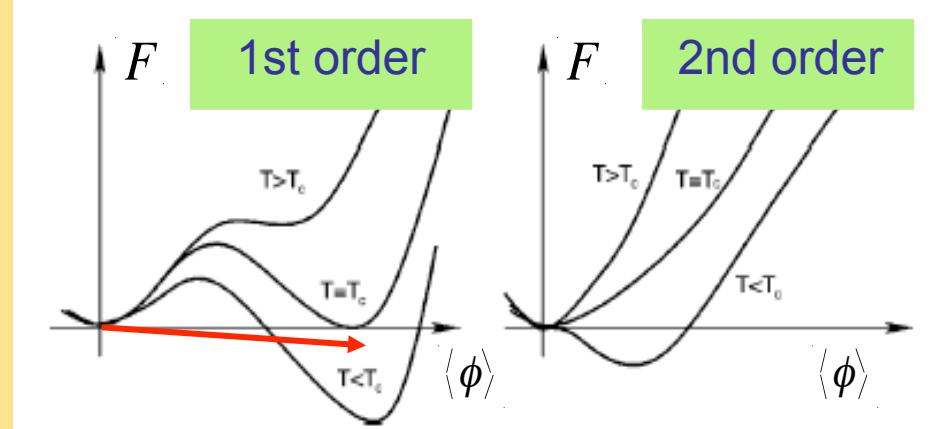
Increasing m_h \longrightarrow
 \longleftarrow New scalars

Baryogenesis
 Gravity Waves
 Scalar DM
 LHC Searches

“Strong” **1st order EWPT**
 ↓
 Bubble nucleation



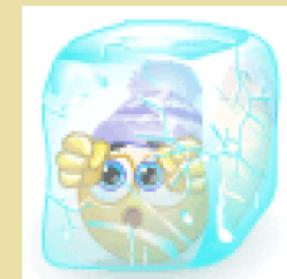
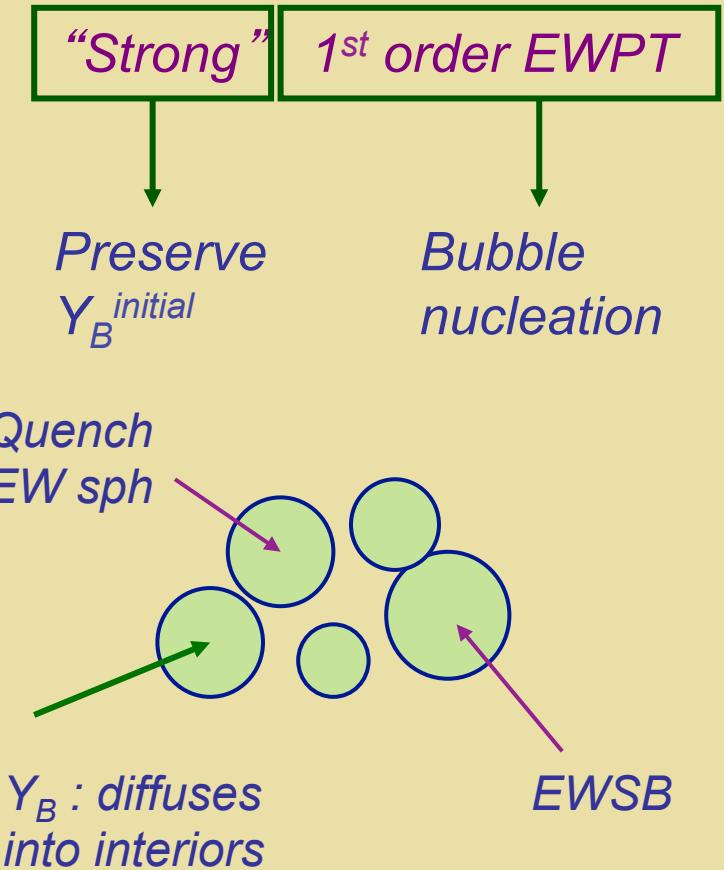
EW Phase Transition: New Scalars & CPV



Increasing m_h →

← New scalars

Baryogenesis
Gravity Waves
Scalar DM
LHC Searches

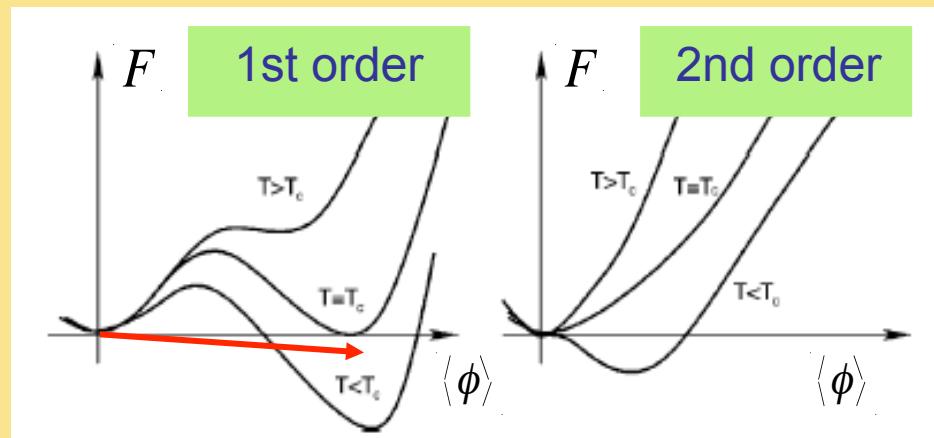


IV. EWPT & DM: Scenarios & Collider Probes

Higgs Portal DM & EWPT



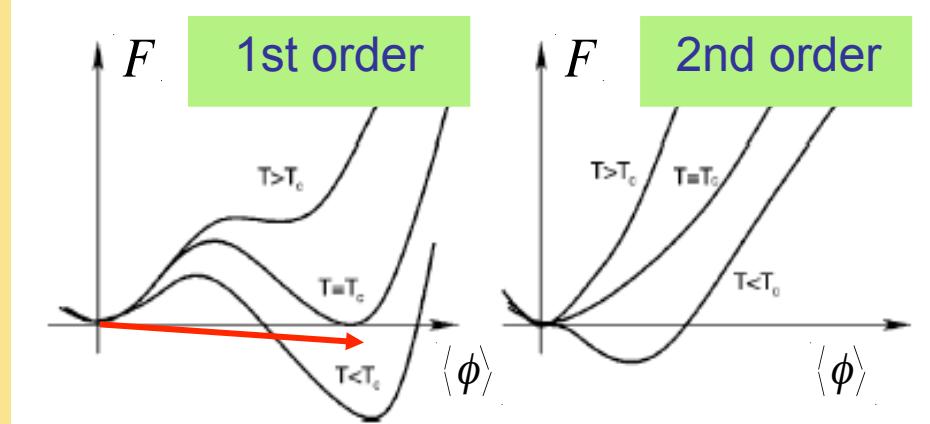
EW Phase Transition: Higgs Portal



Increasing m_h \longrightarrow
 \longleftarrow New scalars

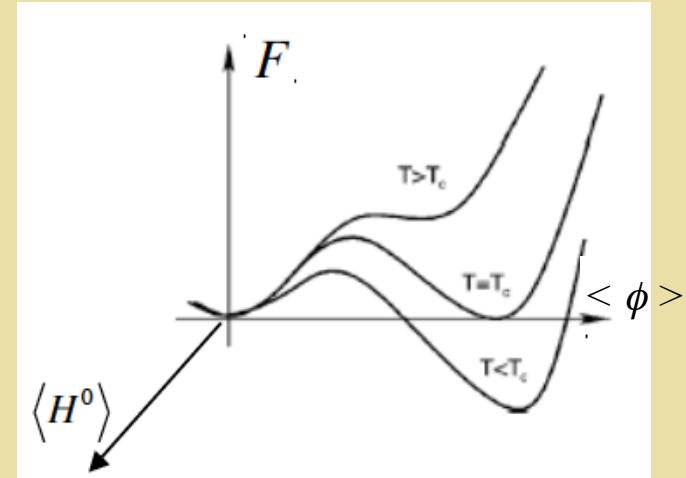
$$\mathcal{O}_4 = \lambda_{\phi H} \phi^\dagger \phi H^\dagger H + \dots$$

EW Phase Transition: Higgs Portal



Increasing m_h \longrightarrow
 \longleftarrow New scalars

$$\mathcal{O}_4 = \lambda_{\phi H} \phi^\dagger \phi H^\dagger H + \dots$$



- Renormalizable
- ϕ : singlet or charged under $SU(2)_L \times U(1)_Y$
- Generic features of full theory (NMSSM, GUTS...)
- More robust vacuum stability
- Novel patterns of SSB

Higgs Portal: Simple Scalar Extensions

Extension	DOF	EWPT	DM
Real singlet: $\cancel{Z_2}$	1	✓	✗
Real singlet: Z_2	1	✓	✓
Complex Singlet	2	✓	✓
EW Multiplets	3+	✓	✓

May be low-energy remnants of UV complete theory & illustrative of generic features

Higgs Portal: Simple Scalar Extensions



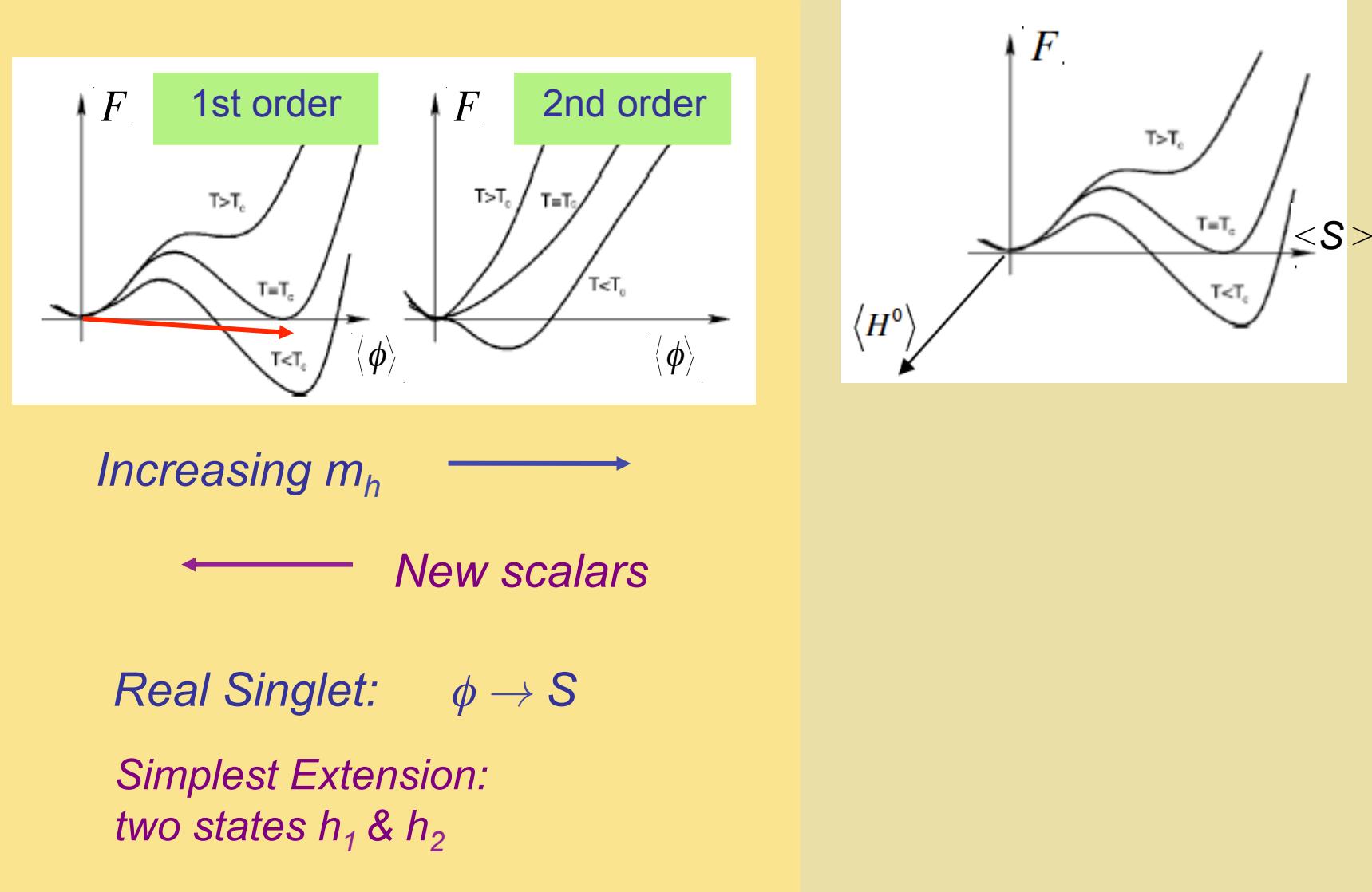
May be low-energy remnants of UV complete theory & illustrative of generic features

Higgs Portal: Simple Scalar Extensions

Extension	DOF	EWPT	DM
Real singlet: $\cancel{Z_2}$	1	✓	✗
Real singlet: Z_2	1	✓	✓
Complex Singlet	2	✓	✓
EW Multiplets	3+	✓	✓

May be low-energy remnants of UV complete theory & illustrative of generic features

EW Phase Transition: New Scalars



Simplest Extension

Standard Model + real singlet scalar

$$V_{\text{HS}} = \frac{a_1}{2} (H^\dagger H) S + \frac{a_2}{2} (H^\dagger H) S^2$$

- *Strong first order EWPT*
- *Two mixed singlet-doublet states*

Simplest Extension

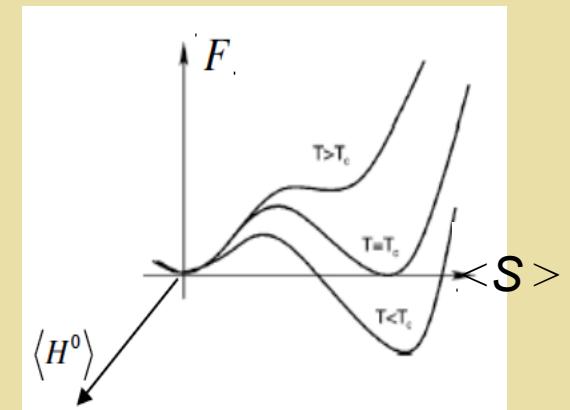
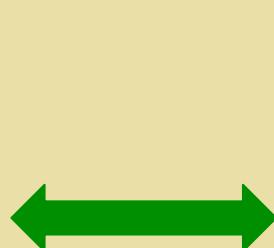
Standard Model + real singlet scalar

$$V_{HS} = \frac{a_1}{2} (H^\dagger H) S + \frac{a_2}{2} (H^\dagger H) S^2$$

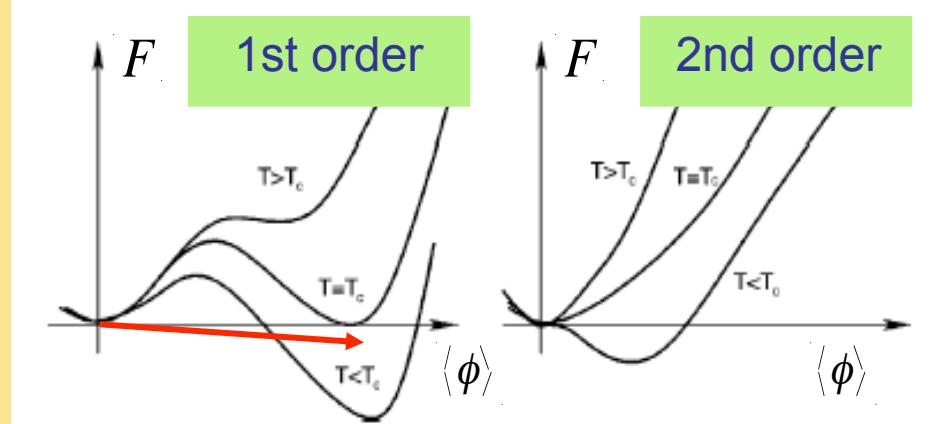
Phenomenology

$$\begin{aligned} h_1 &= \sin \theta \ s + \cos \theta \ h \\ h_2 &= \cos \theta \ s - \sin \theta \ h \end{aligned}$$

$m_{1,2}$; θ ; $h_i h_j h_k$ couplings



EW Phase Transition: New Scalars

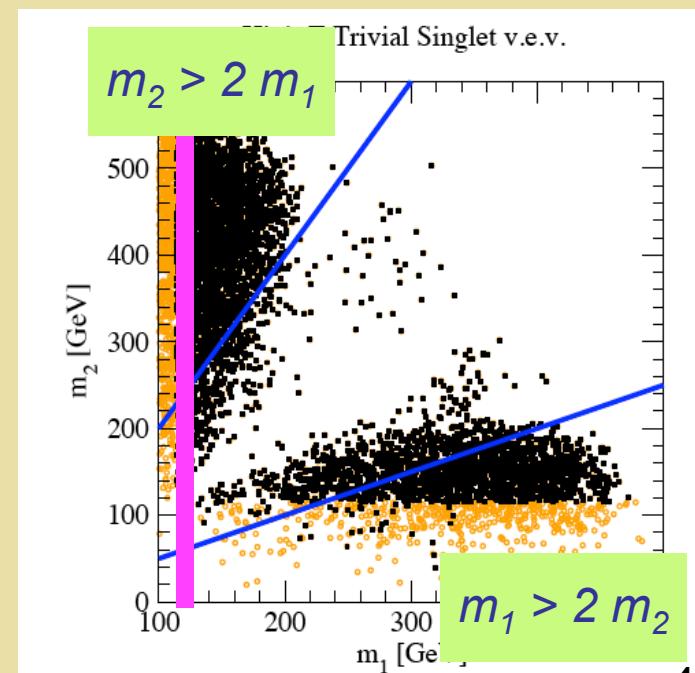
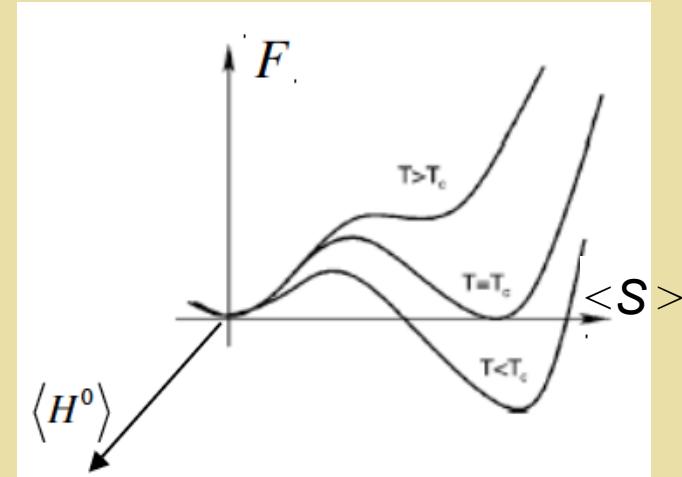


Increasing m_h \longrightarrow

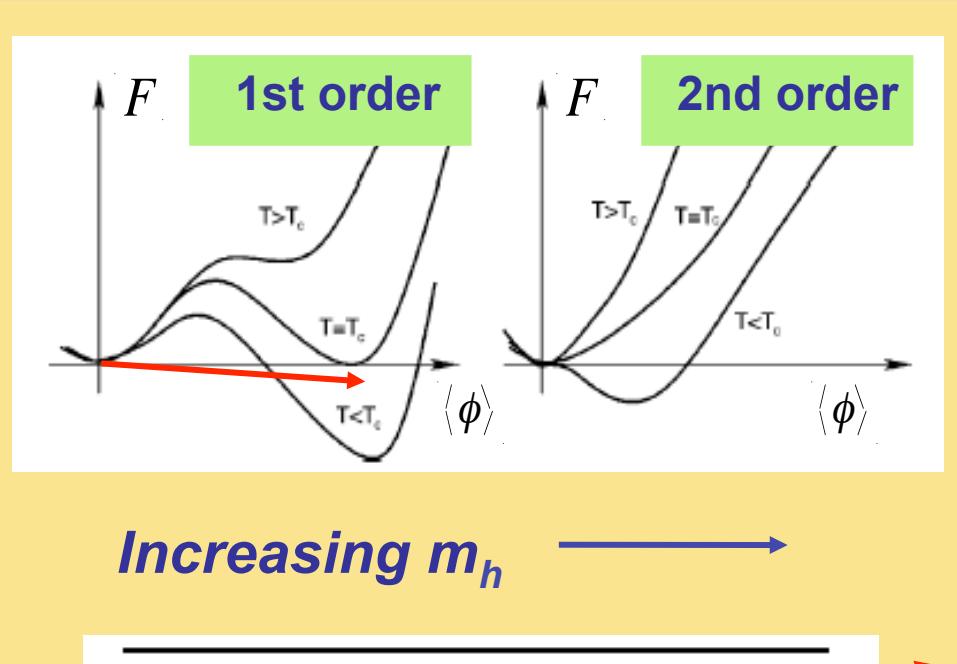
\longleftarrow New scalars

Real Singlet: $\phi \rightarrow S$

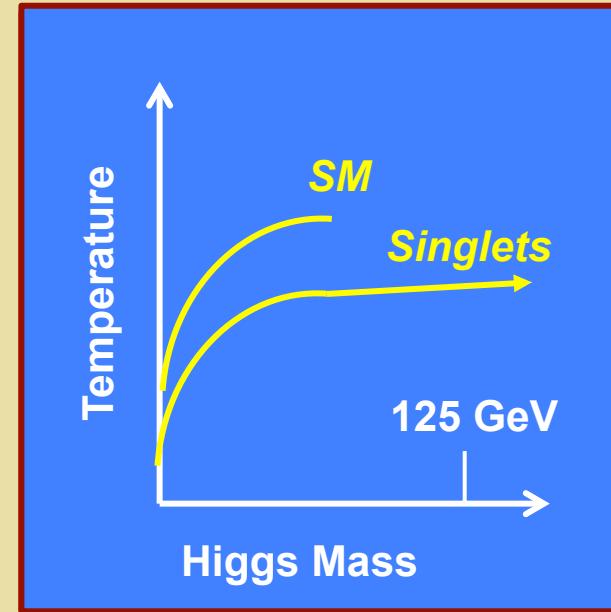
Simplest Extension:
two states h_1 & h_2



EW Phase Transition: Singlet Scalars



Lattice	Authors	M_h^C (GeV)
4D Isotropic	[76]	80 ± 7
4D Anisotropic	[74]	72.4 ± 1.7
3D Isotropic	[72]	72.3 ± 0.7
3D Isotropic	[70]	72.4 ± 0.9

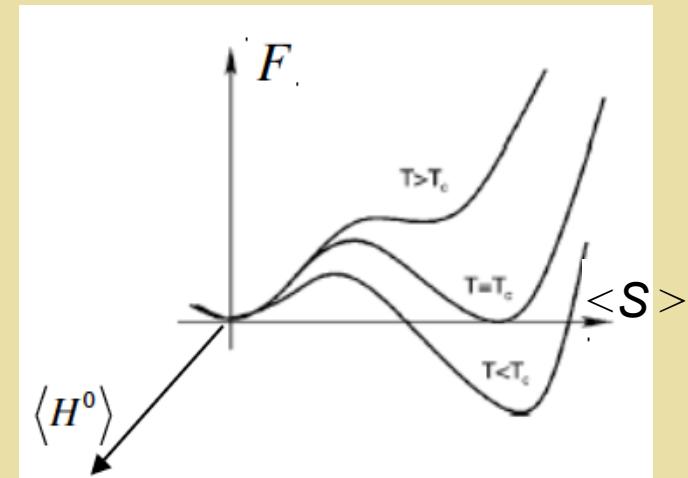
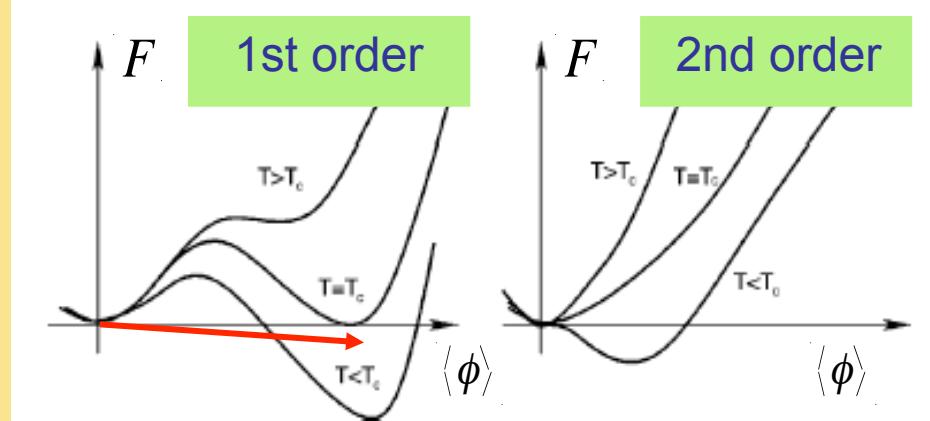


EW Phase Diagram

How does this picture change in presence of new TeV scale physics ? What is the phase diagram ?

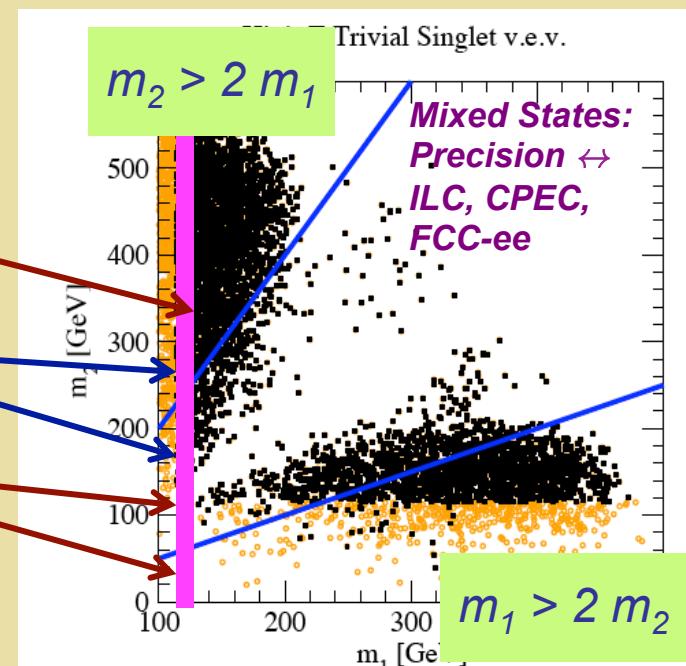
SM EW: Cross over transition

EW Phase Transition: Singlet Scalars

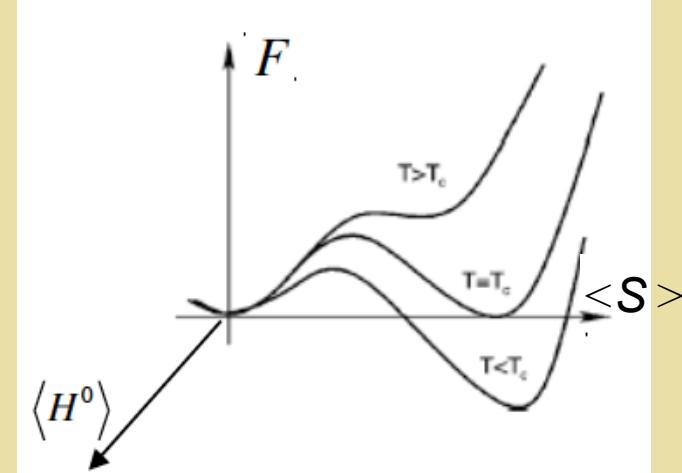
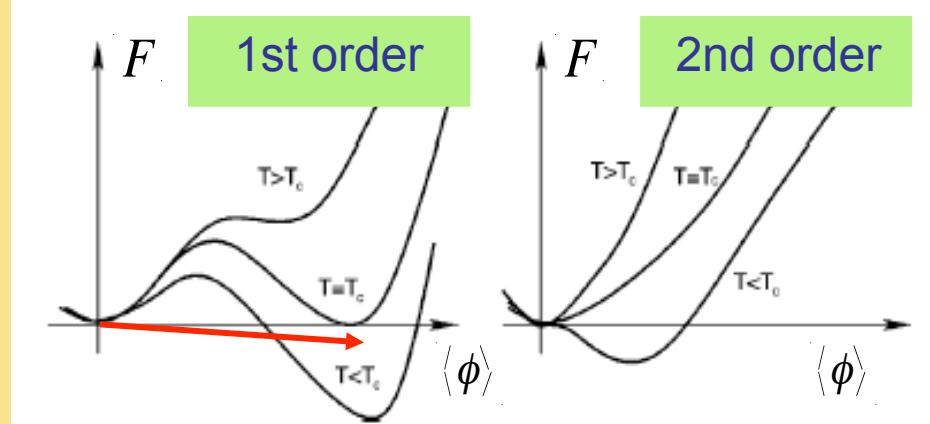


Collider probes

- Resonant di-Higgs production
- Precision Higgs measurements
- Non-resonant di-Higgs & exotic Higgs decays

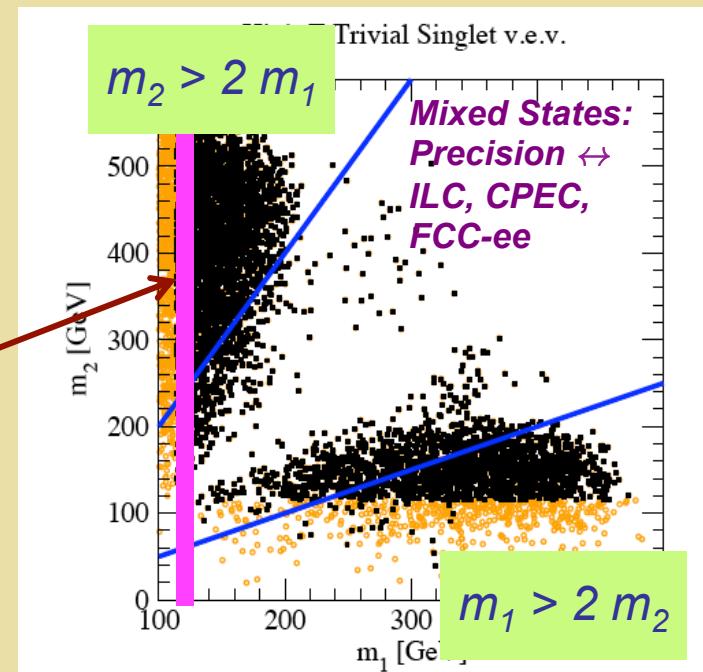
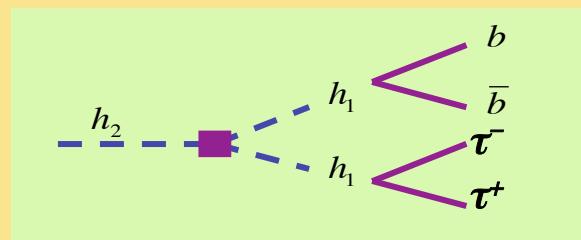


EW Phase Transition: New Scalars



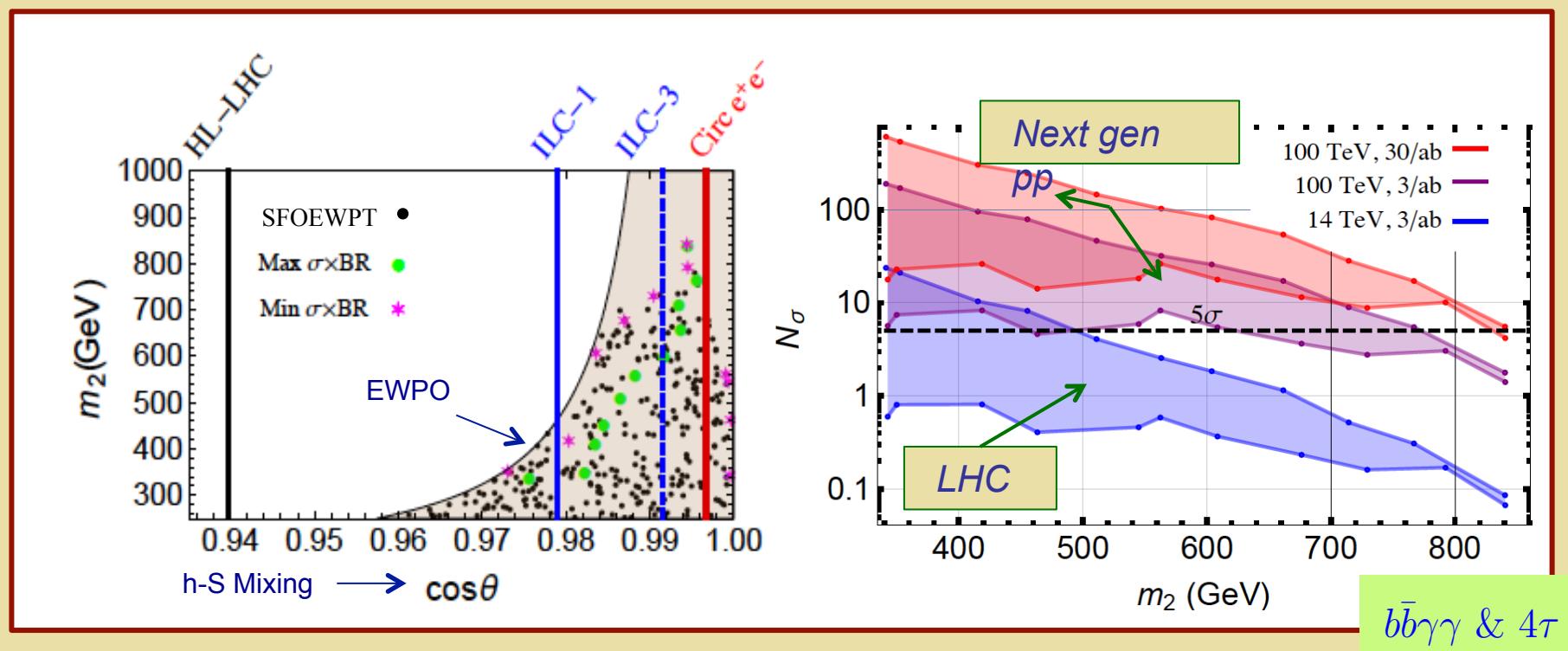
Increasing m_h \longrightarrow

Resonant di-Higgs production



EW Phase Transition: Singlet Scalars

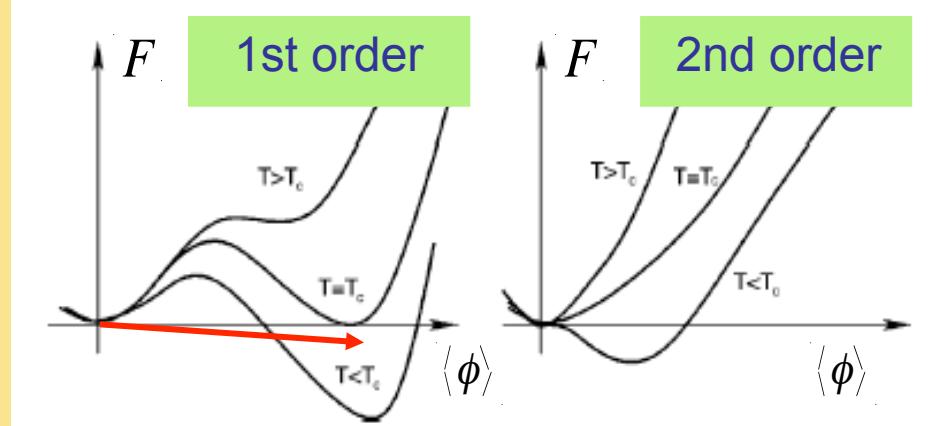
SFOEWPT Benchmarks: Resonant di-Higgs & precision Higgs studies



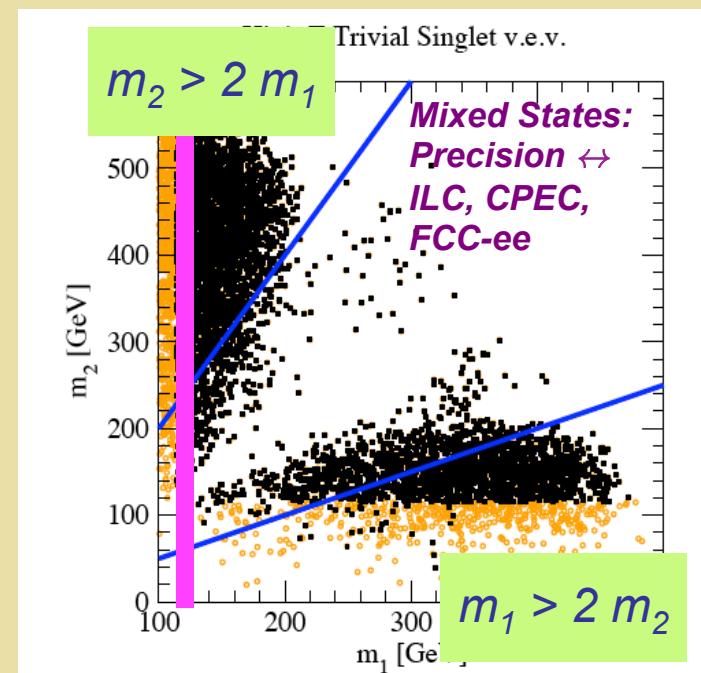
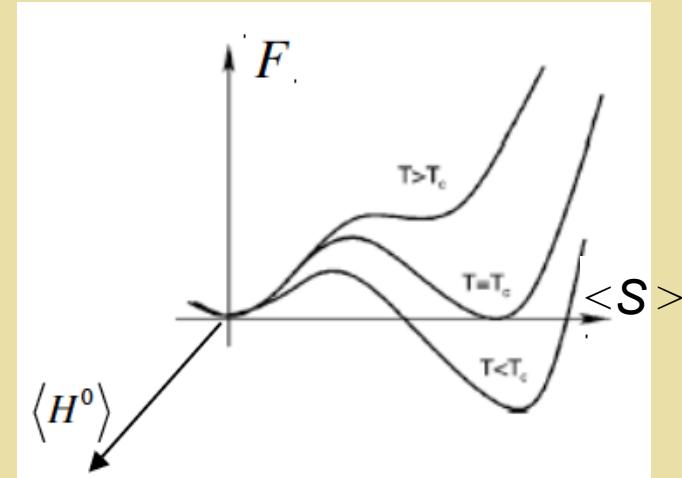
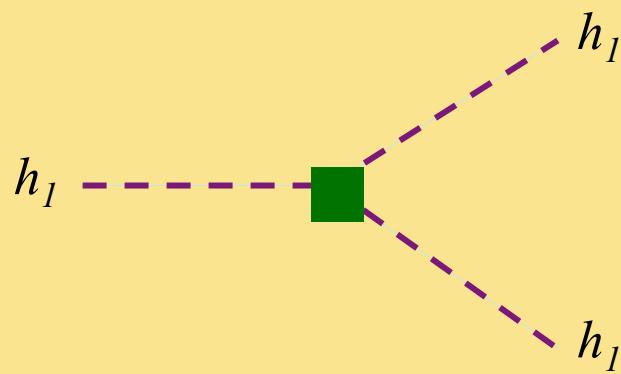
Kotwal, No, R-M, Winslow 1605.06123

See also: Huang et al, 1701.04442

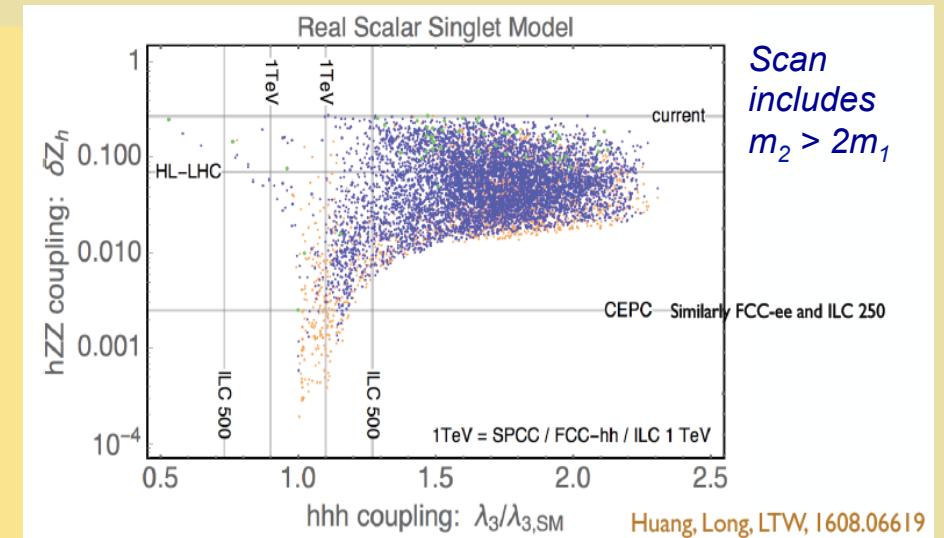
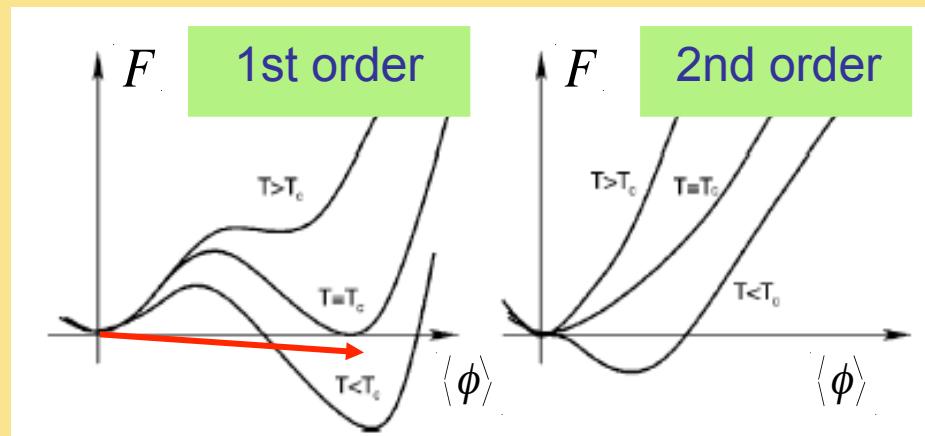
EW Phase Transition: New Scalars



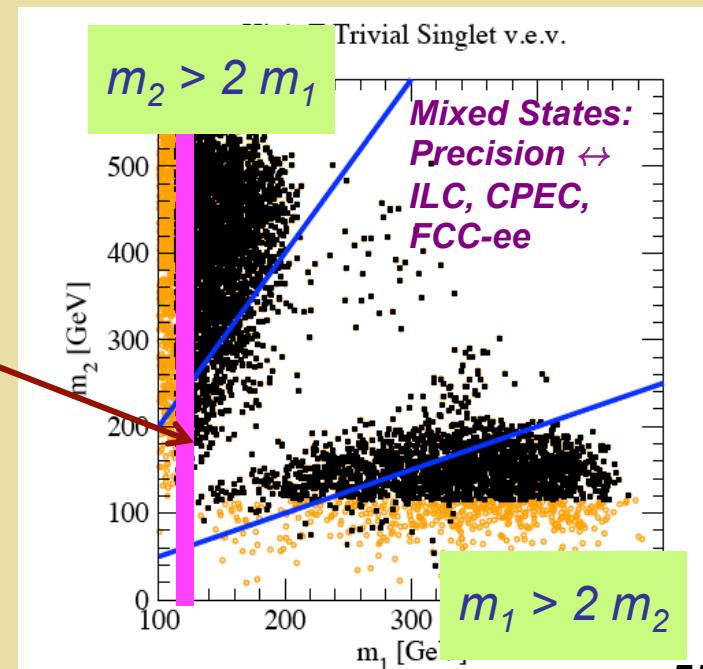
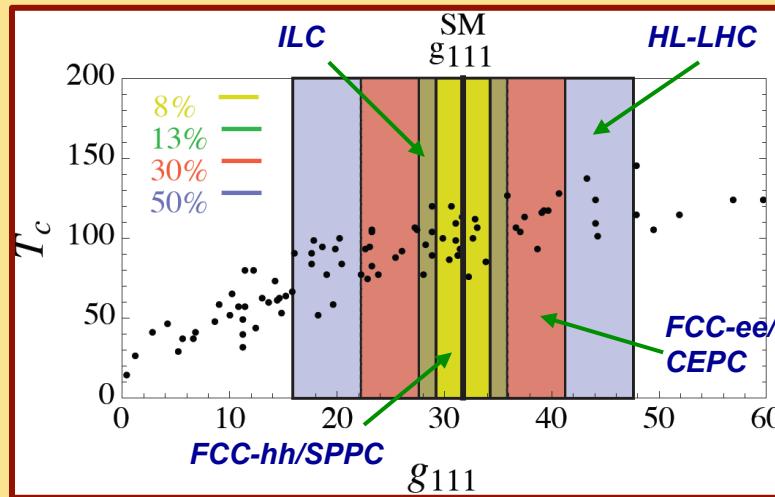
Modified Higgs Self-Coupling



EW Phase Transition: Singlet Scalars

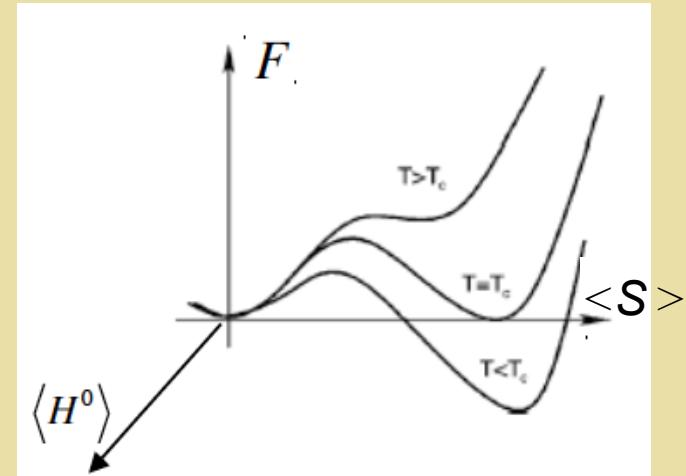
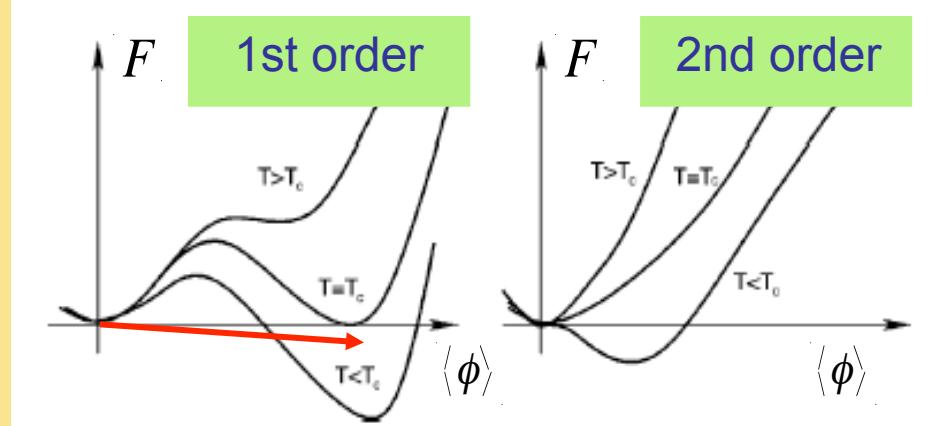


Modified Higgs Self-Coupling

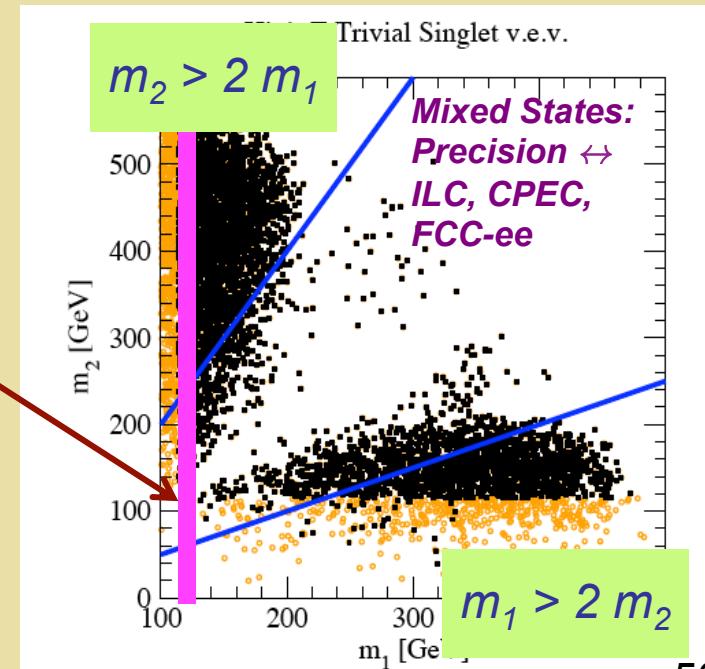
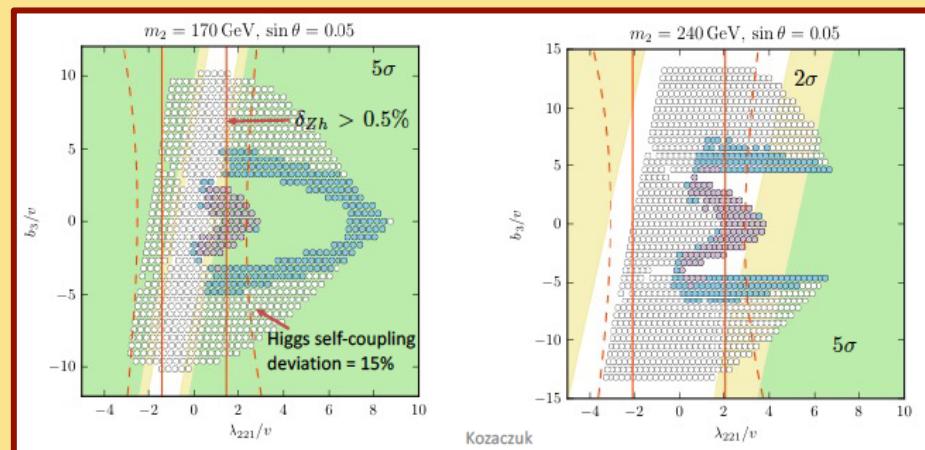


Profumo, R-M, Wainwright, Winslow: 1407.5342;
see also Noble & Perelstein 0711.3018

EW Phase Transition: Singlet Scalars



Singlet-like pair production



Higgs Portal: Simple Scalar Extensions

Extension	DOF	EWPT	DM
Real singlet: $\cancel{Z_2}$	1	✓	✗
Real singlet: Z_2	1	✓	✓
Complex Singlet	2	✓	✓
EW Multiplets	3+	✓	✓

May be low-energy remnants of UV complete theory & illustrative of generic features

The Simplest Extension

DM Scenario

Dark Matter Stability:

- Invariance under $S \rightarrow -S$
- $\langle S \rangle = 0$
- $\theta_{hs} = 0$



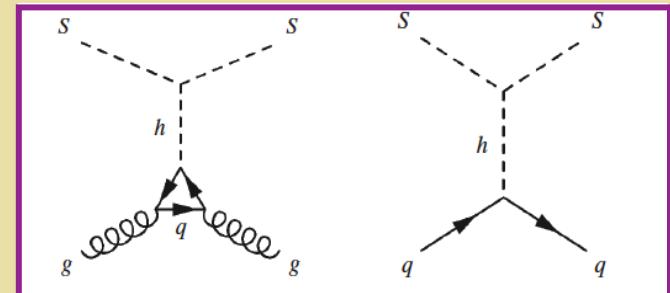
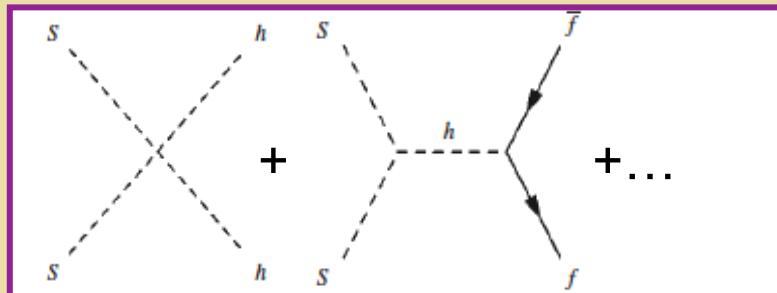
*“ Z_2
Symmetry”*

The Simplest Extension

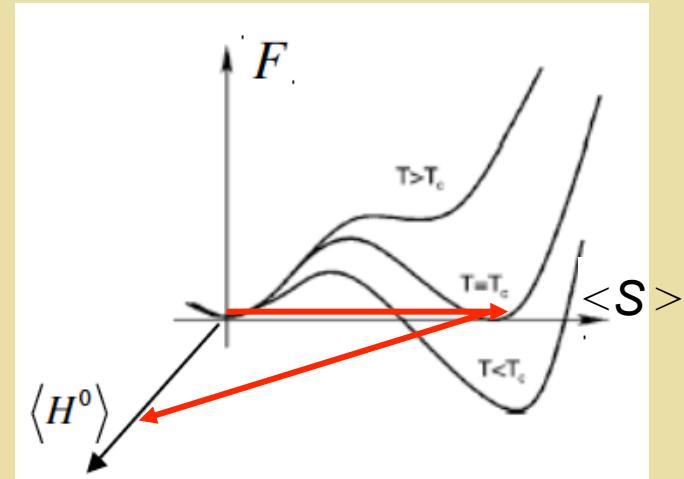
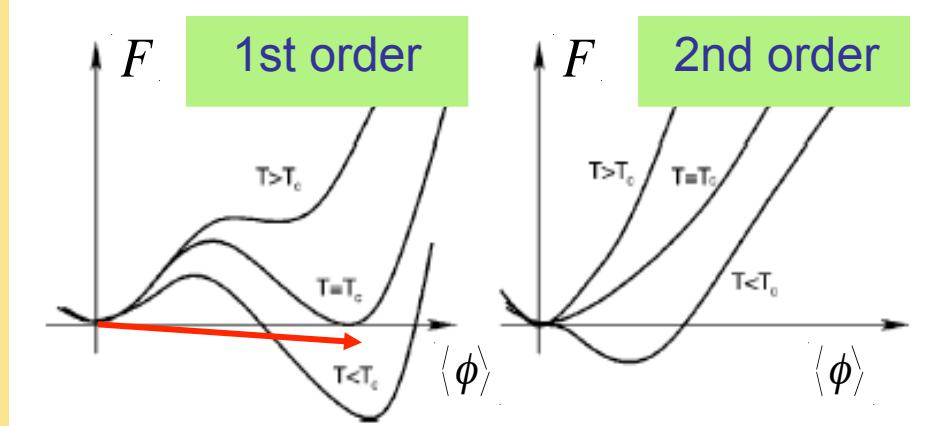
DM Scenario

$$V_{\text{HS}} = + \frac{a_2}{2} \left(H^\dagger H \right) S^2$$

Ω_{DM} & σ_{SI}



EW Phase Transition: Two-Step

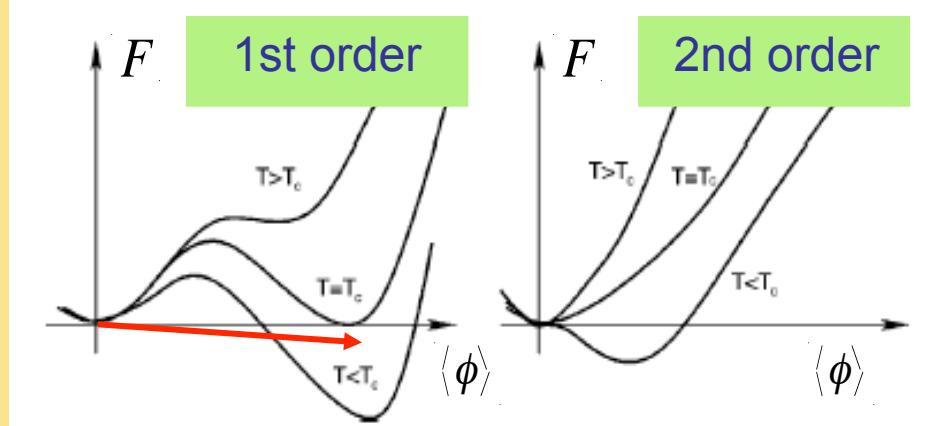


Profumo, R-M, Shaugnessy 2007

Curtain, Meade, Yu: arXiv: 1409.0005

Jiang, Bian, Huang, Shu 1502.07574

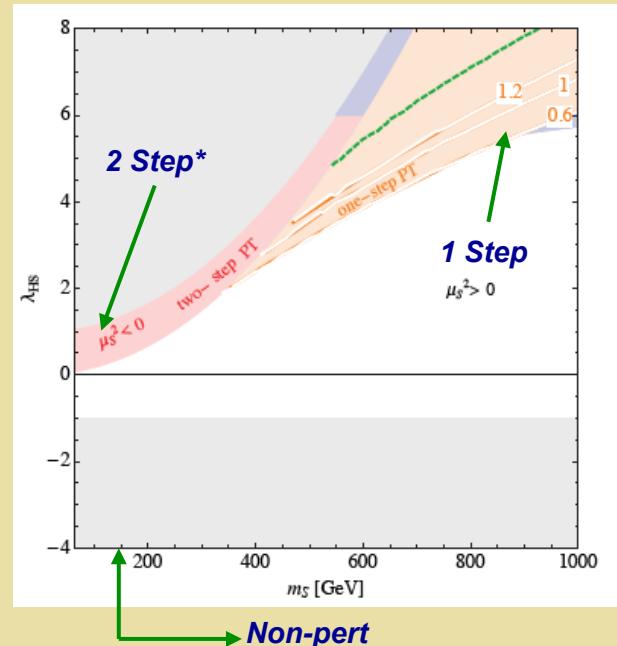
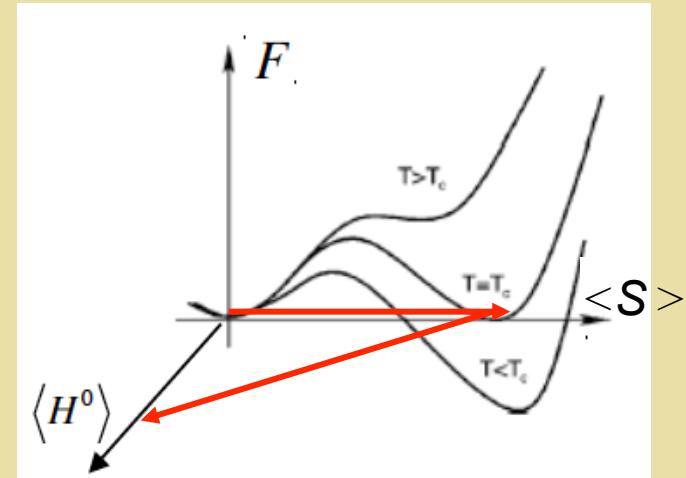
EW Phase Transition: Singlet Scalars



Curtain, Meade, Yu: arXiv: 1409.0005

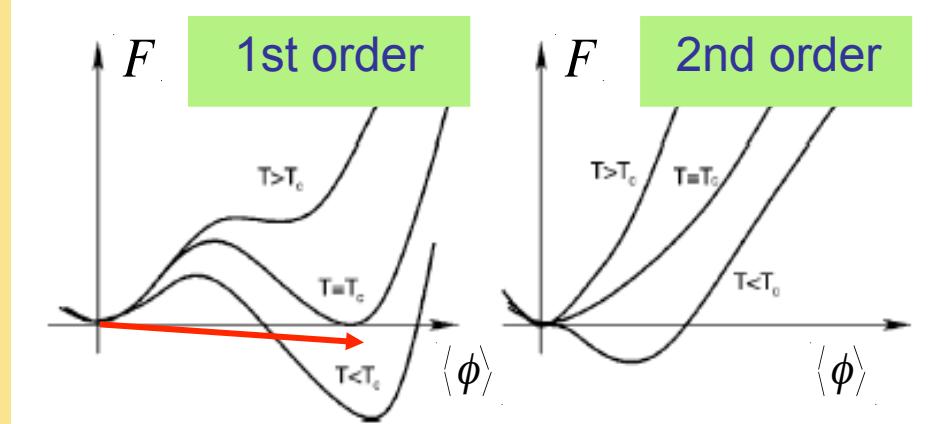
Z_2 symmetric real singlet extension

- Loop-induced 1-step transition
- 2-step transition for $\mu_S^2 < 0$



* Singlet two step: see also Profumo, R-M, Shaugnessy 2007

EW Phase Transition: Singlet Scalars



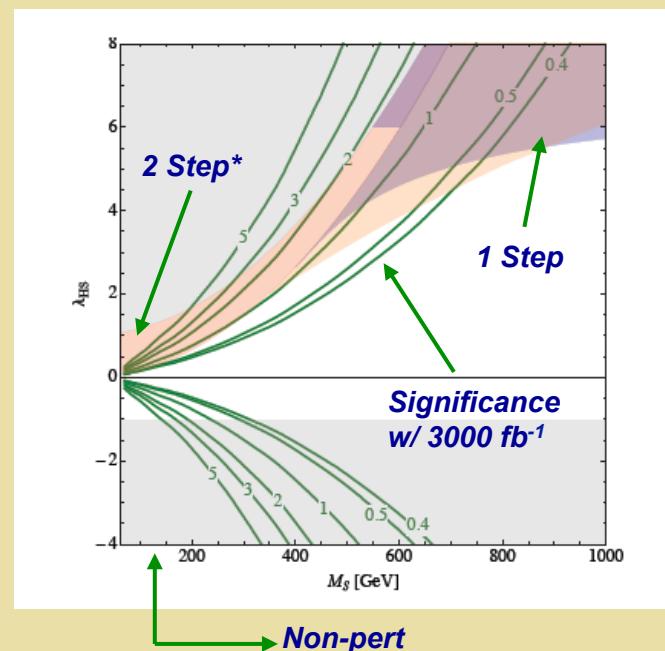
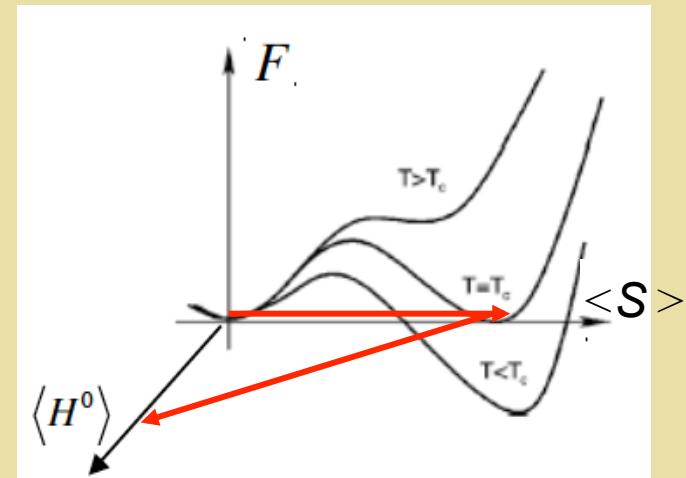
Curtain, Meade, Yu: arXiv: 1409.0005

Z_2 symmetric real singlet extension

- Loop-induced 1-step transition
- 2-step transition for $\mu_S^2 < 0$

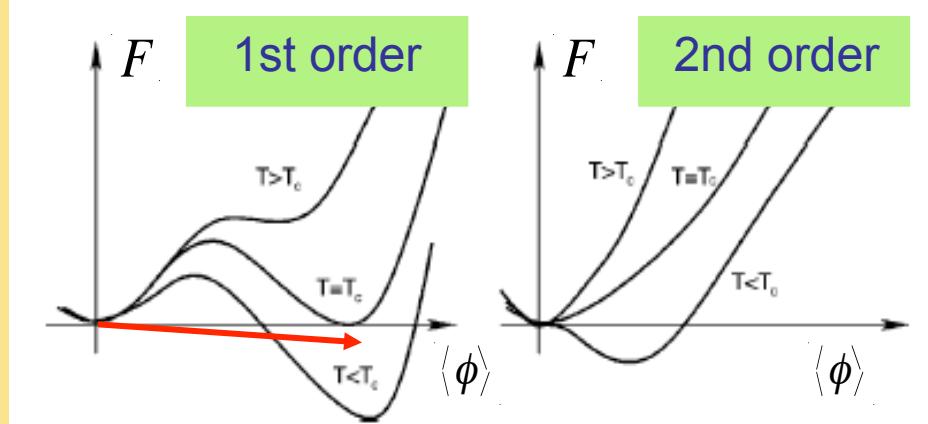
VBF @ 100 TeV pp:

$$pp \rightarrow h jj, h \rightarrow invis$$



* Singlet two step: see also Profumo, R-M, Shaugnessy 2007

EW Phase Transition: DM Direct Detection

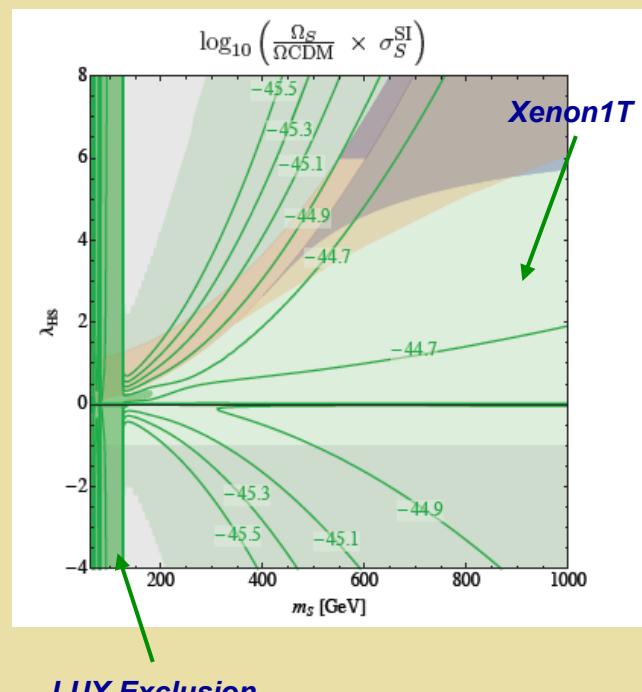
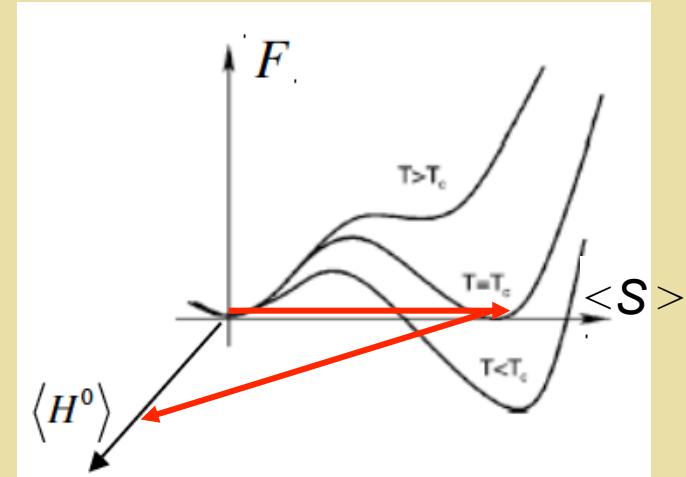


Curtain, Meade, Yu: arXiv: 1409.0005

Z_2 symmetric real singlet extension

- Loop-induced 1-step transition
- 2-step transition for $\mu_S^2 < 0$

Scalar singlet DM: direct detection

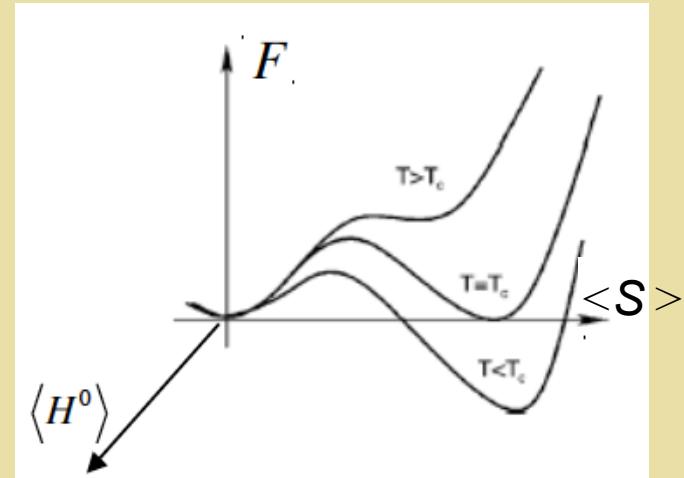
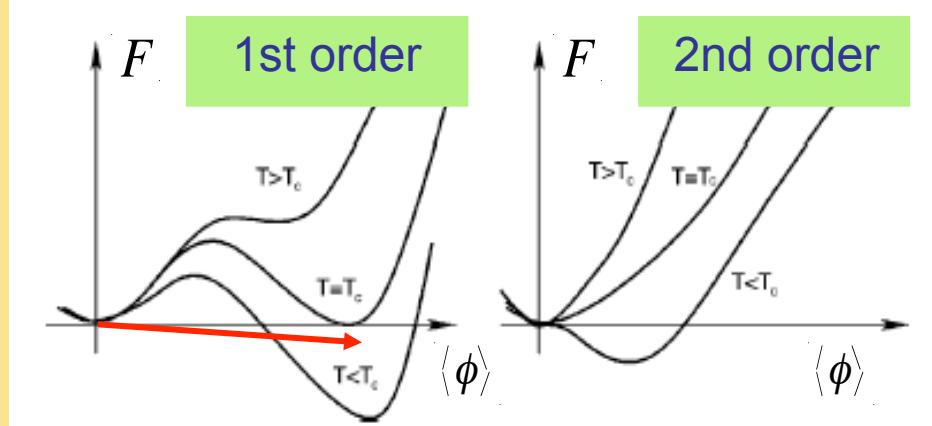


Higgs Portal: Simple Scalar Extensions

Extension	DOF	EWPT	DM
Real singlet: $\cancel{Z_2}$	1	✓	✗
Real singlet: Z_2	1	✓	✓
Complex Singlet	2	✓	✓
EW Multiplets	3+	✓	✓

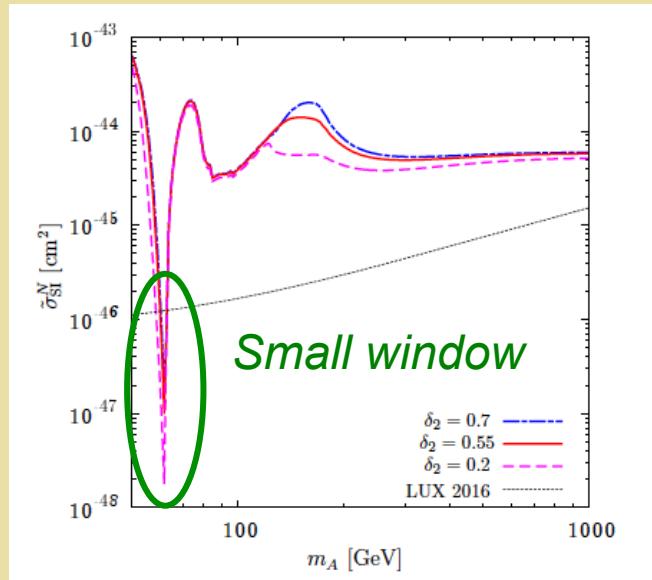
May be low-energy remnants of UV complete theory & illustrative of generic features

EWPT: Complex Singlet Scalars



SFOEWPT-viable parameters

- DM: Spontaneously & softly-broken global $U(1)$
- Possibility of SFOEWPT & DM candidate



Higgs Portal: Simple Scalar Extensions

Extension	DOF	EWPT	DM
Real singlet: $\cancel{Z_2}$	1	✓	✗
Real singlet: Z_2	1	✓	✓
Complex Singlet	2	✓	✓
EW Multiplets	3+	✓	✓

May be low-energy remnants of UV complete theory & illustrative of generic features

Real Triplet

Σ^0 , Σ^+ , Σ^-

$\sim (1, 3, 0)$

Fileviez-Perez, Patel, Wang, R-M:
PRD 79: 055024 (2009); 0811.3957
[hep-ph]

$$V_{H\Sigma} = \frac{a_1}{2} H^\dagger \Sigma H + \frac{a_2}{2} H^\dagger H \text{ Tr } \Sigma^2$$

EWPT: $a_{1,2} \neq 0$ & $\langle \Sigma^0 \rangle \neq 0$

DM & EWPT: $a_1 = 0$ & $\langle \Sigma^0 \rangle = 0$

Small: ρ -param

Real Triplet

$$\Sigma^0, \Sigma^+, \Sigma^-$$

$\sim (1, 3, 0)$

Fileviez-Perez, Patel, Wang, R-M:
PRD 79: 055024 (2009); 0811.3957
[hep-ph]

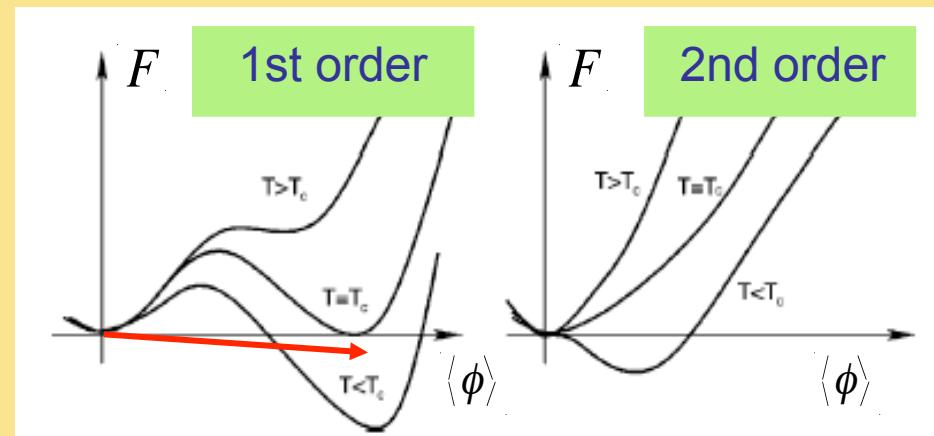
$$V_{H\Sigma} = + \frac{a_2}{2} H^\dagger H \text{ Tr } \Sigma^2$$

EWPT: $a_{1,2} \neq 0$ & $\langle \Sigma^0 \rangle \neq 0$

DM & EWPT: $a_1 = 0$ & $\langle \Sigma^0 \rangle = 0$

DM Stability

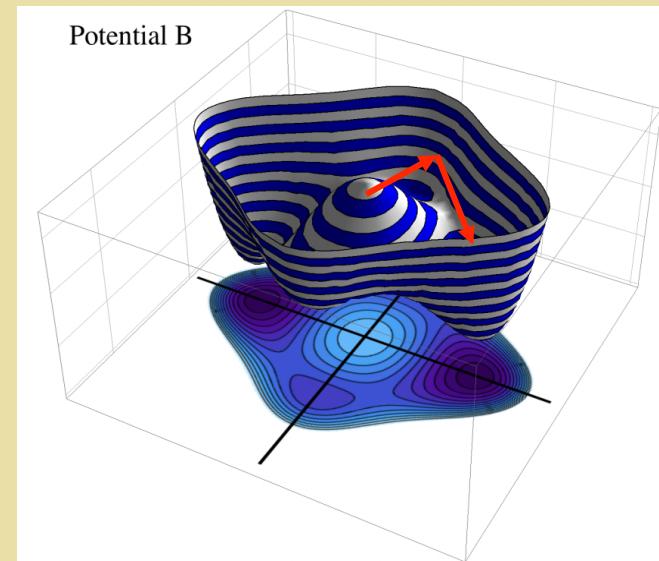
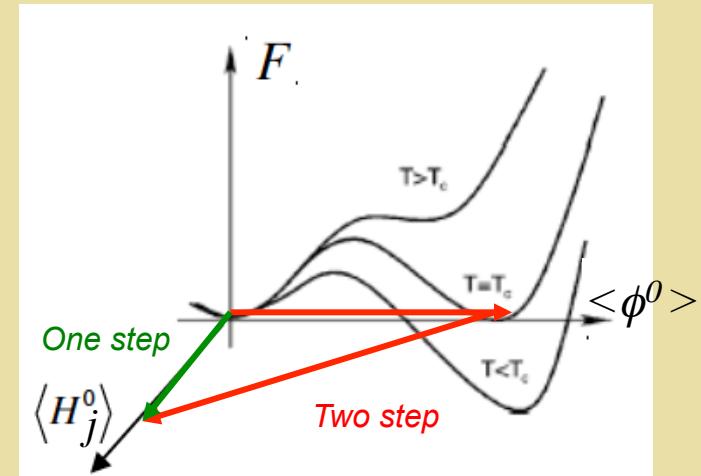
EW Multiplets: Two-Step EWPT



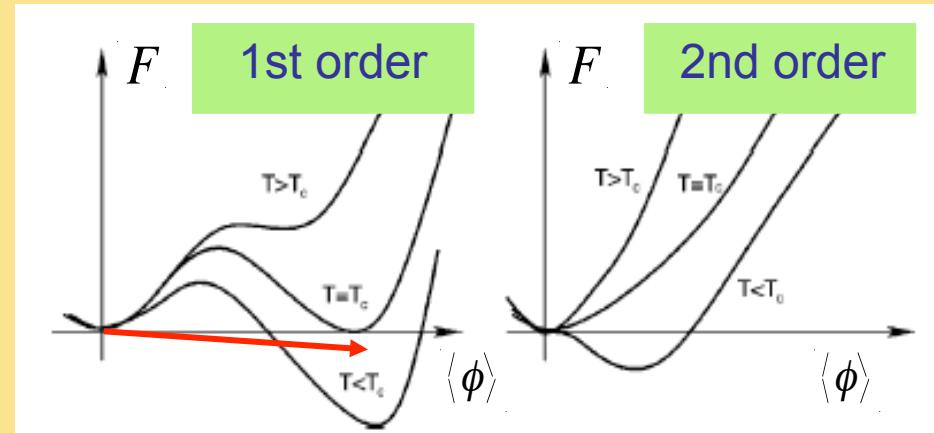
Increasing m_h \longrightarrow

\longleftarrow New scalars

- Step 1: thermal loops
- Step 2: tree-level barrier



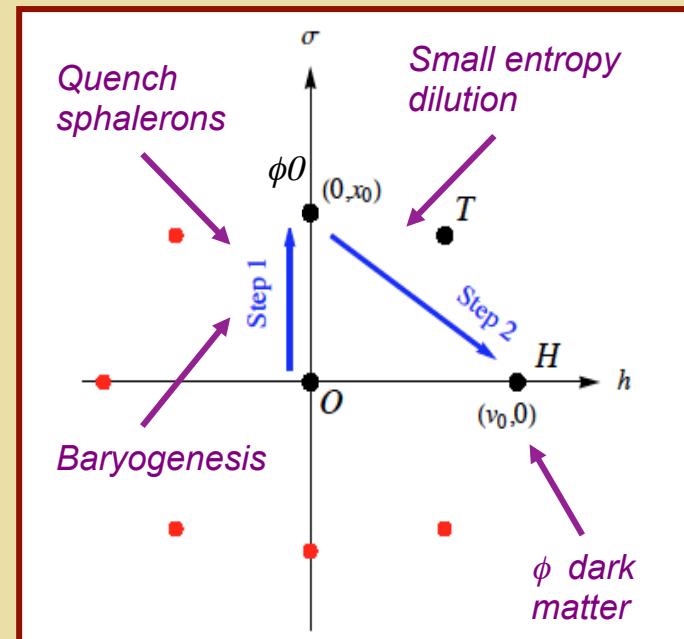
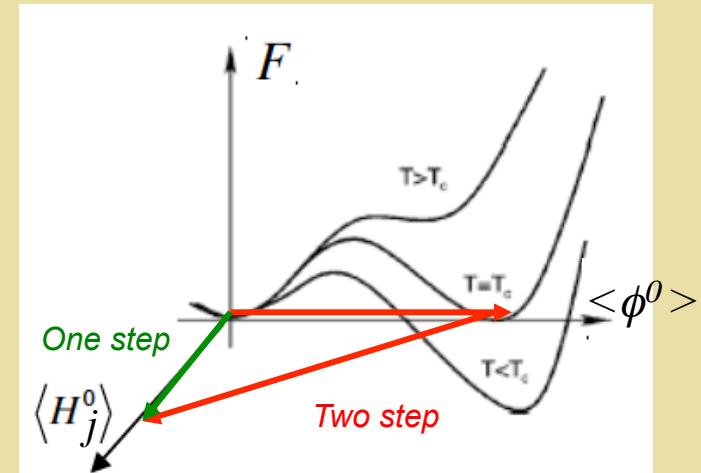
EW Multiplets: Two-Step EWPT



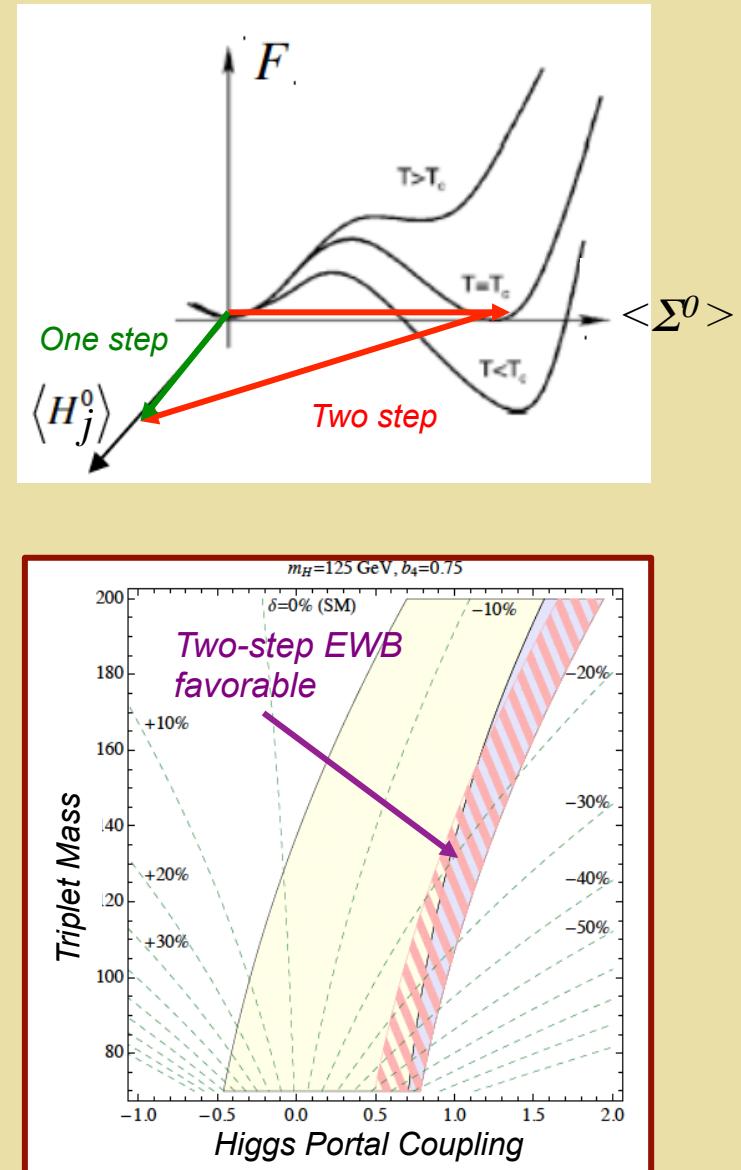
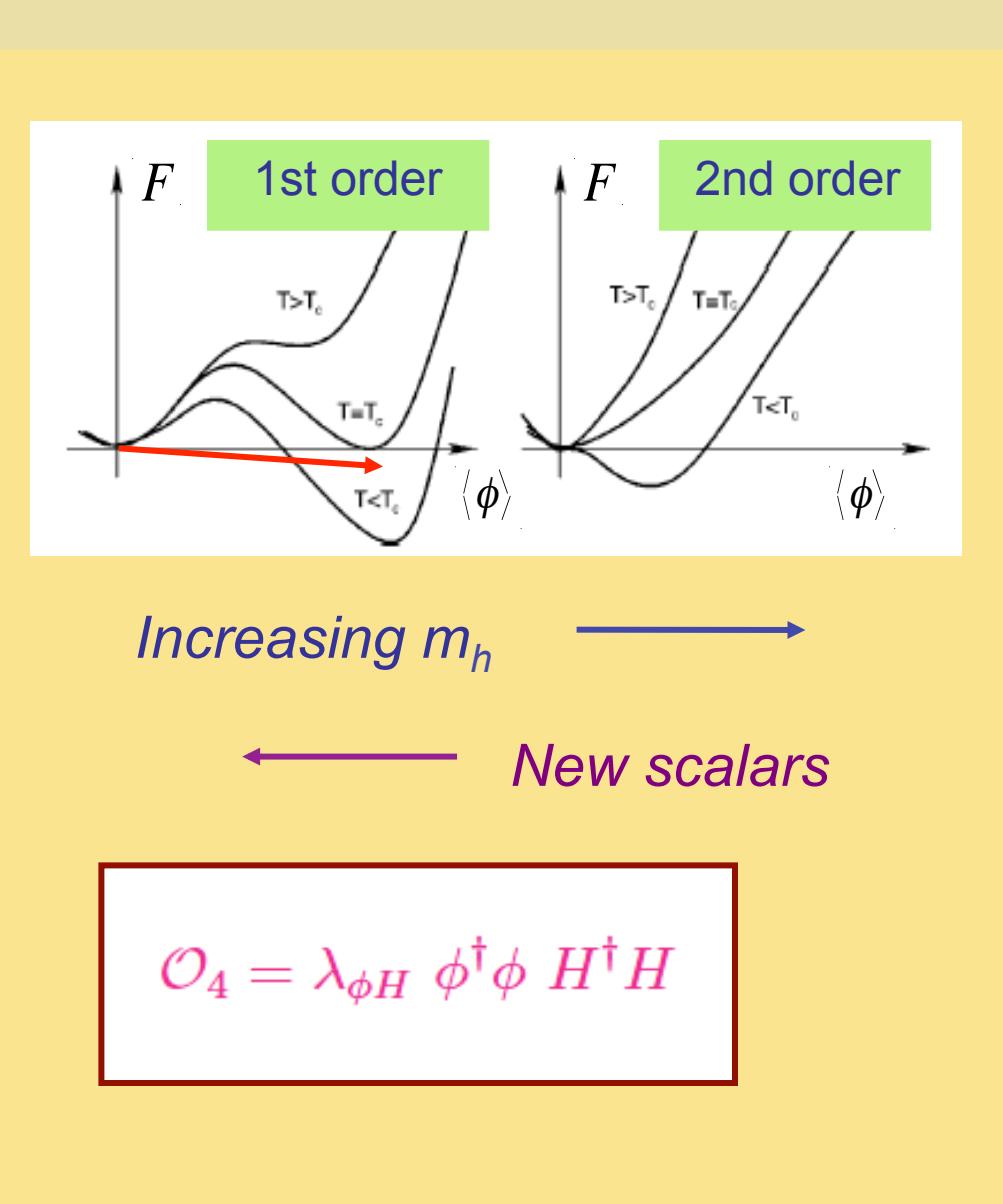
Increasing m_h \longrightarrow

\longleftarrow New scalars

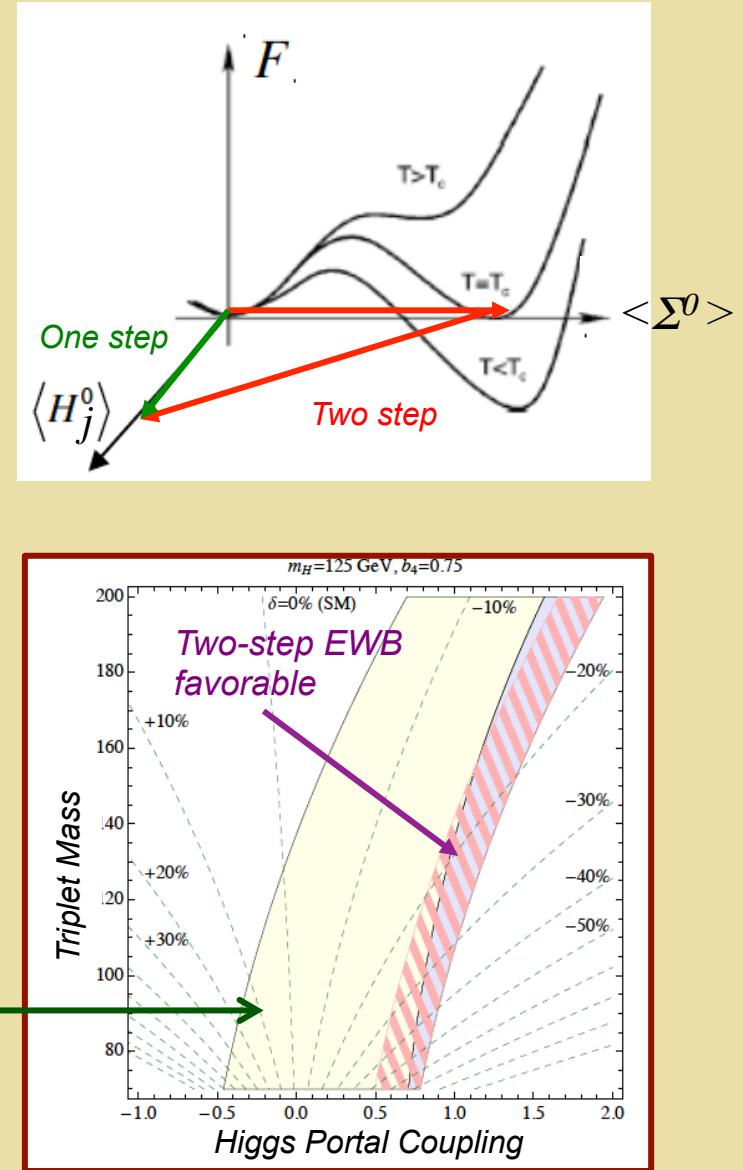
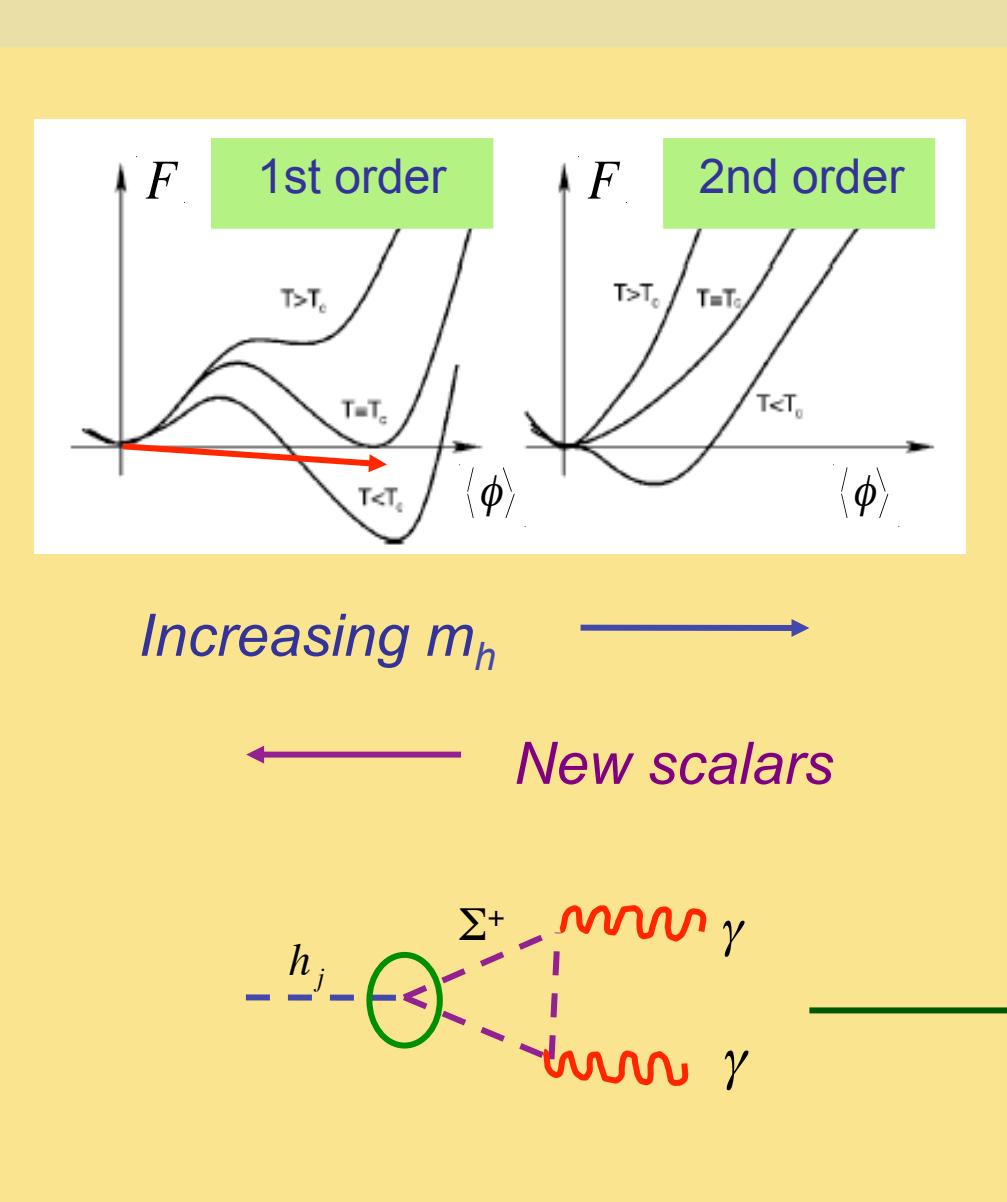
- Step 1: thermal loops
- Step 2: tree-level barrier



EW Multiplets: Real Triplet



EW Multiplets: Real Triplet



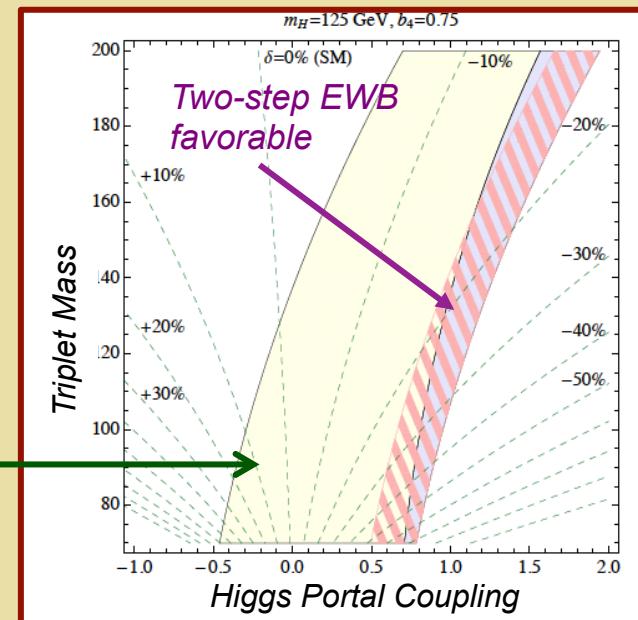
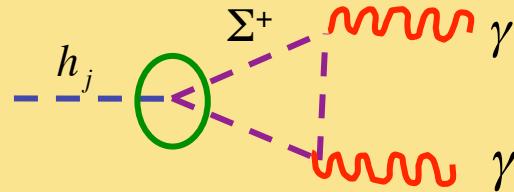
EW Multiplets: Two-Step EWPT

Using $\text{BR}(H \rightarrow ZZ^*)$ from FCC-ee (known at $\sim 0.3\%$ from $\delta g_{HZZ} \sim 0.15\%$), production ratios $\sigma(H \rightarrow XY)/\sigma(H \rightarrow ZZ^*)$ for $p_T > 100 \text{ GeV}$ return the following stat precision on the **absolute value** of rare BRs

M. Mangano

$\delta \text{ BR}$	$\gamma\gamma$	$Z\gamma$	$\mu\mu$	FCC-ee: < 2% on $\delta_{H\gamma\gamma}$
	$\sim 0.5\%$	$\sim 1\%$	$\sim 1\%$	

Increasing m_h →
← New scalars



Lecture III Key Ideas

- *Role of 1st order EWPT in electroweak baryogenesis*
- *Thermal WIMP dark matter: relic density, indirect & direct detection signatures, collider probes*
- *Higgs portal EWPT & dark matter: simplest singlet & electroweak multiplet scalar sector extensions; Z₂ symmetry; multi-step EW symmetry breaking*
- *Collider implications: resonant di-Higgs production; modified Higgs self-coupling, Higgs signal strengths (mixing), Higgs di-photon decay*

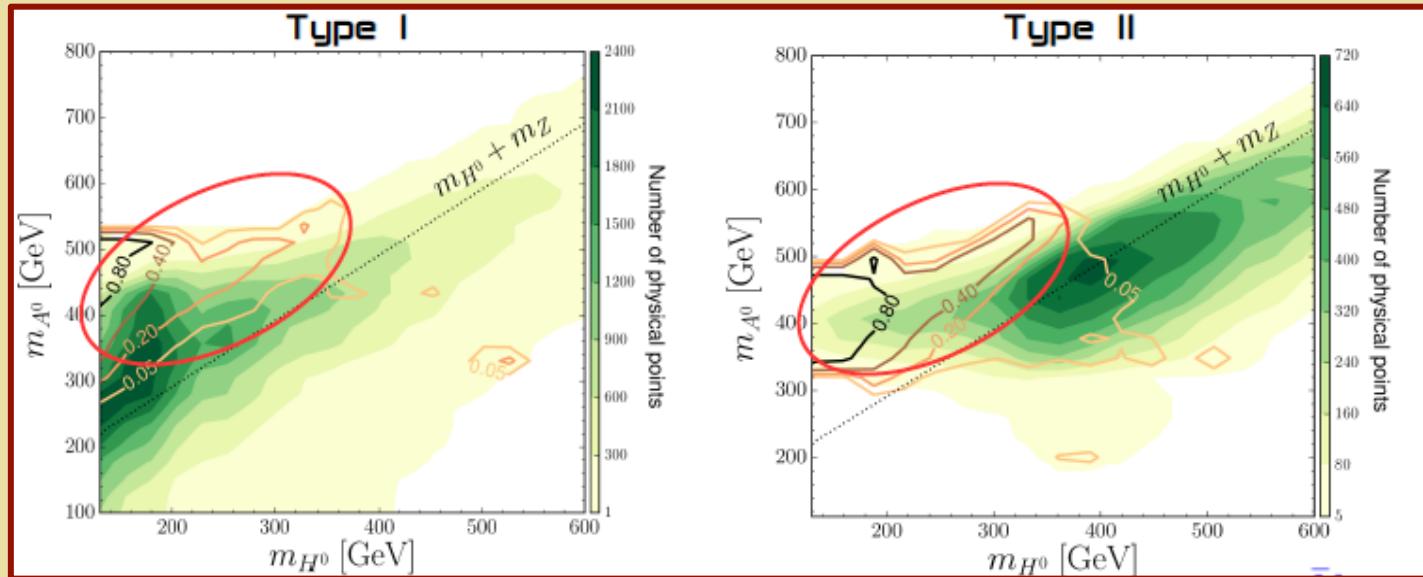
Research Opportunities

- *Finite T , out-of-equilibrium quantum field theory*
- *Model building*
- *Collider phenomenology*
- *Inter-frontier connections: high-energy, cosmic,
& high sensitivity*

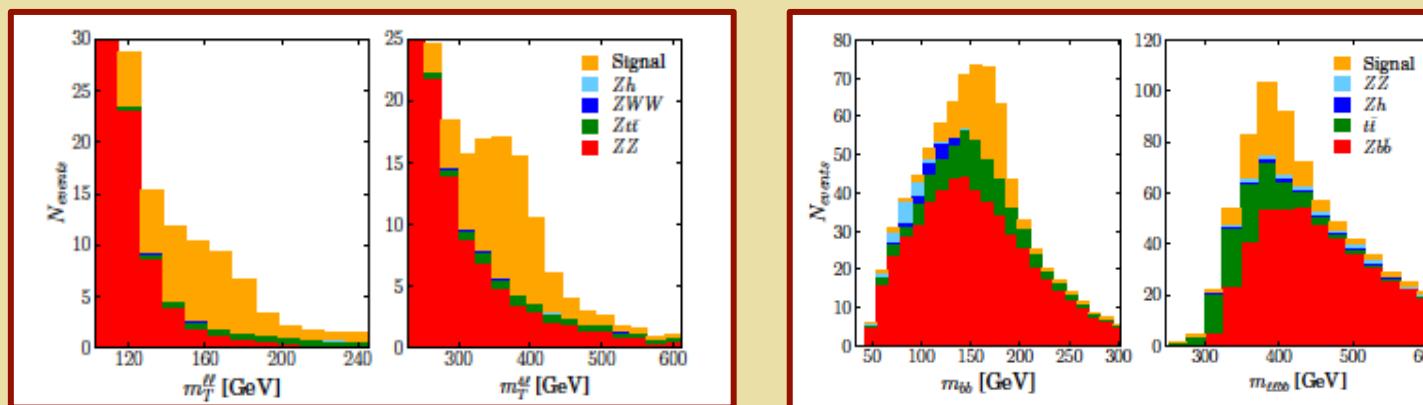
謝謝

Back Up Slides

EW Multiplets: 2HDM



$A^0 \rightarrow Z H^0$
Signature



SFOEWPT
testable w/
LHC

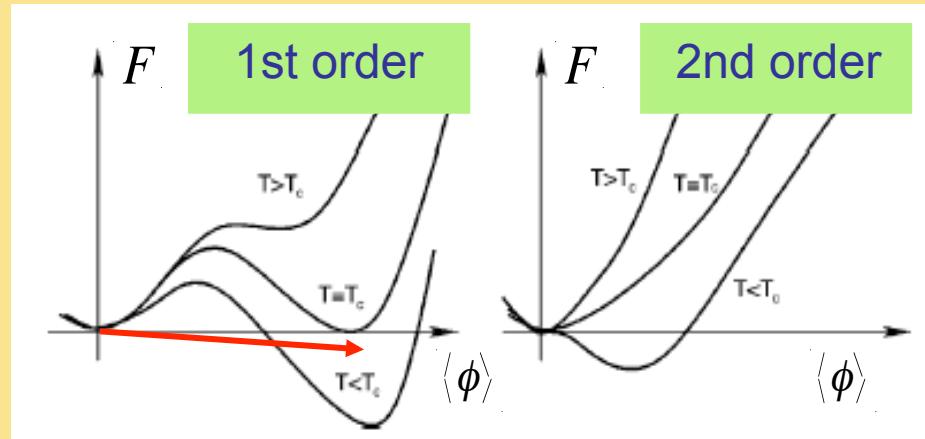
$bb \parallel : 5\sigma$ for 40 fb^{-1}

$4l + \text{MET}: 5\sigma$ for 60 fb^{-1}

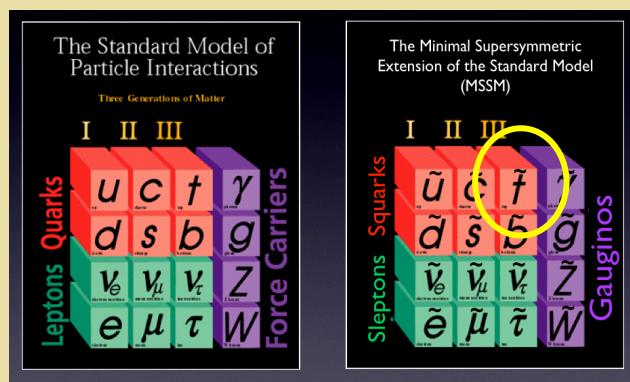
Dorsch et al

See S. Su talk

EW Phase Transition: SUSY

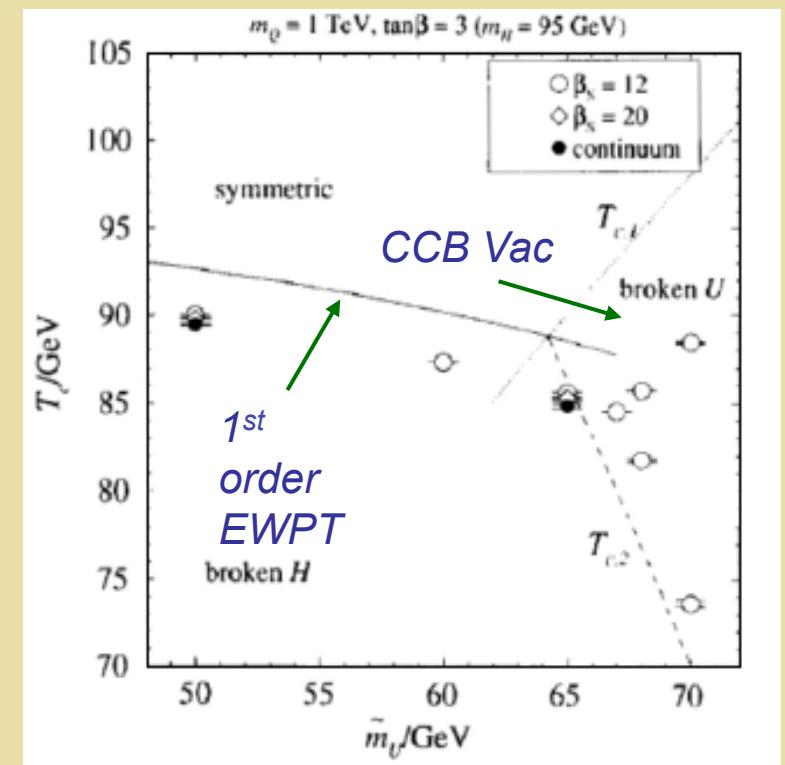


Increasing m_h →
← New scalars



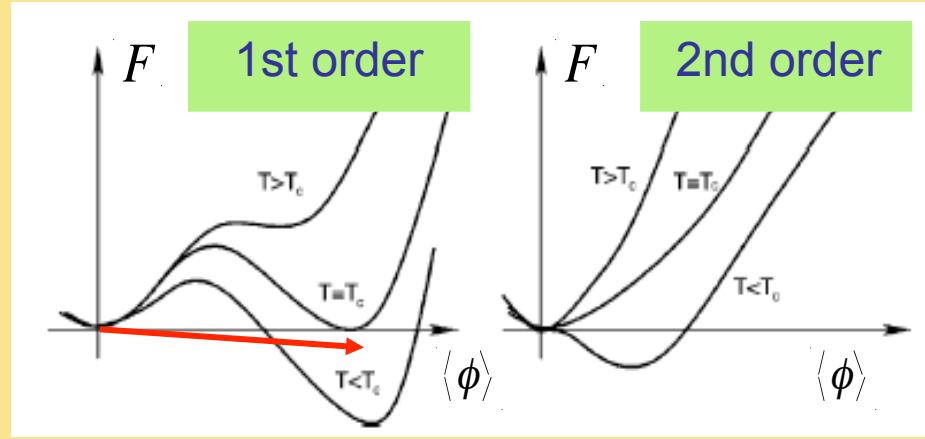
MSSM: Light Stop Scenario

Lattice: Laine, Rummukainen

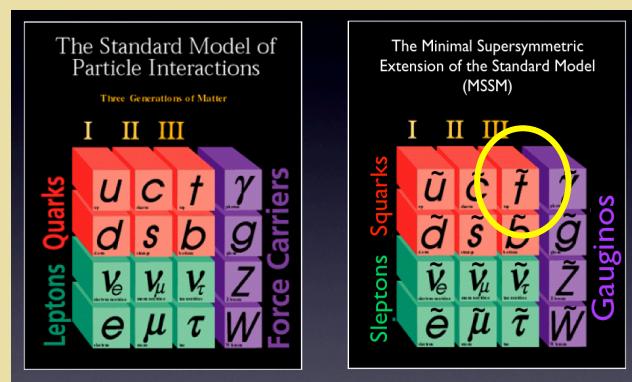


Decreasing RH stop mass →

EW Phase Transition: SUSY

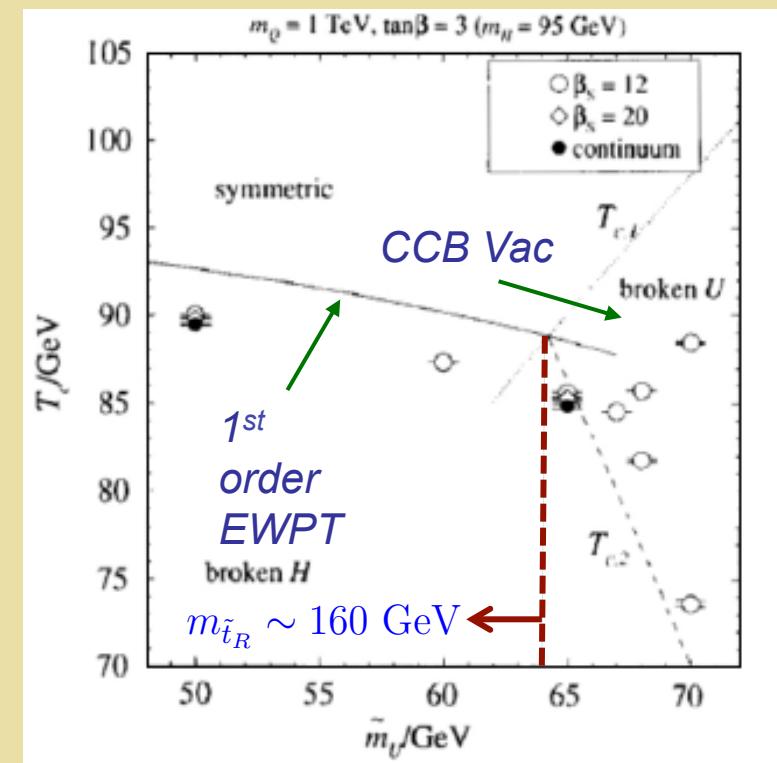


Increasing m_h →
← New scalars



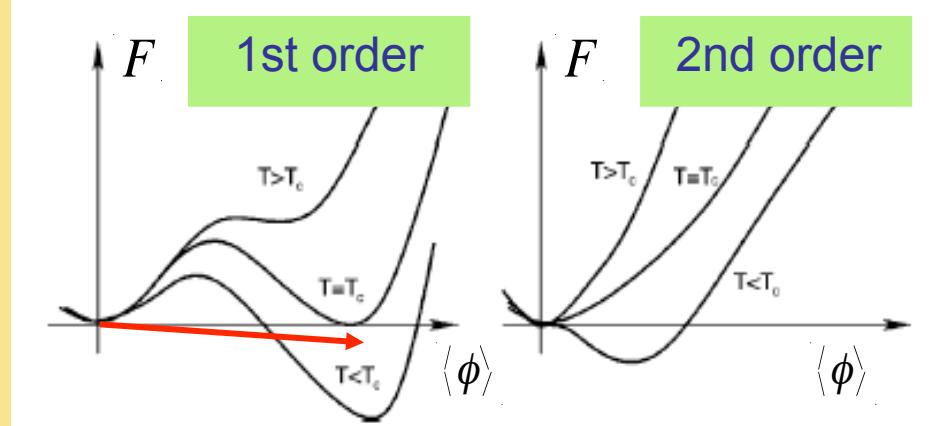
MSSM: Light Stop Scenario

Lattice: Laine, Rummukainen



Decreasing RH stop mass →

EW Phase Transition: SUSY

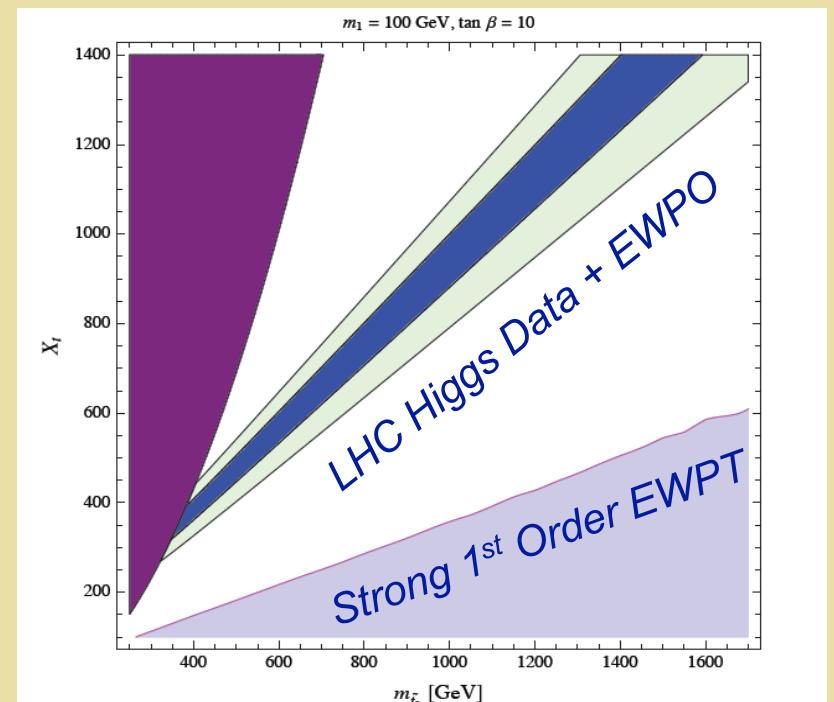


Increasing m_h →
← New scalars

Light RH stops also affect
Higgs properties

Curtin, Jaiswal, Meade 1203.2932

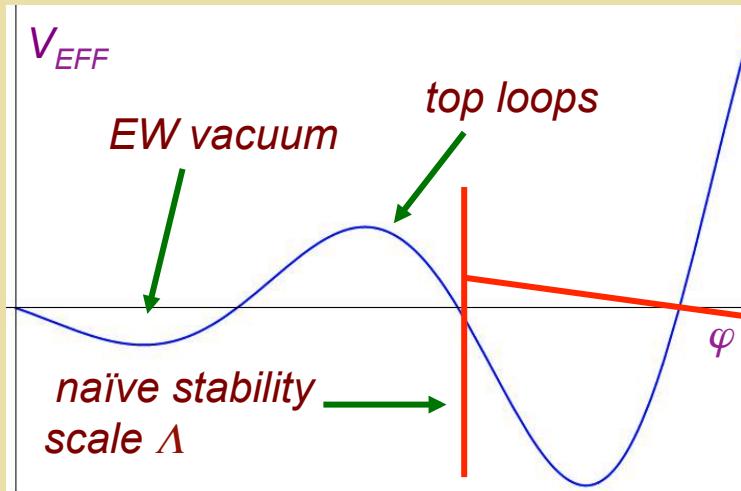
$$MSSM + \delta\lambda_4 (H_u^\dagger H_u)^2$$



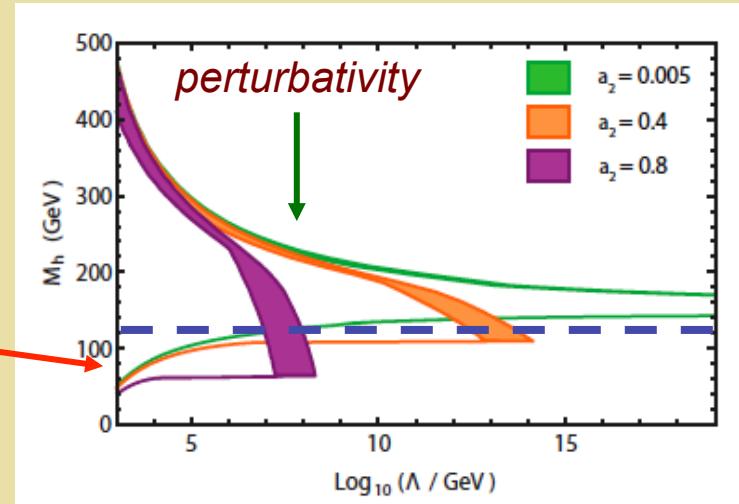
Katz, Perelstein, R-M,
Winslow 1509.02934

Vacuum Stability & Perturbativity

Preserving EW Min



“Funnel plot”



m_H ?

$$\beta_\lambda = \frac{1}{16\pi^2} \left(4\lambda^2 + 12a_2^2 - 36y_t^4 + 12\lambda y_t^2 - 9\lambda g^2 - 3\lambda g'^2 + \frac{9}{4}g'^4 + \frac{9}{2}g^2 g'^2 + \frac{27}{4}g^4 \right)$$

DM-H coupling

top loops

Temperature Dependence of $V(\phi)$

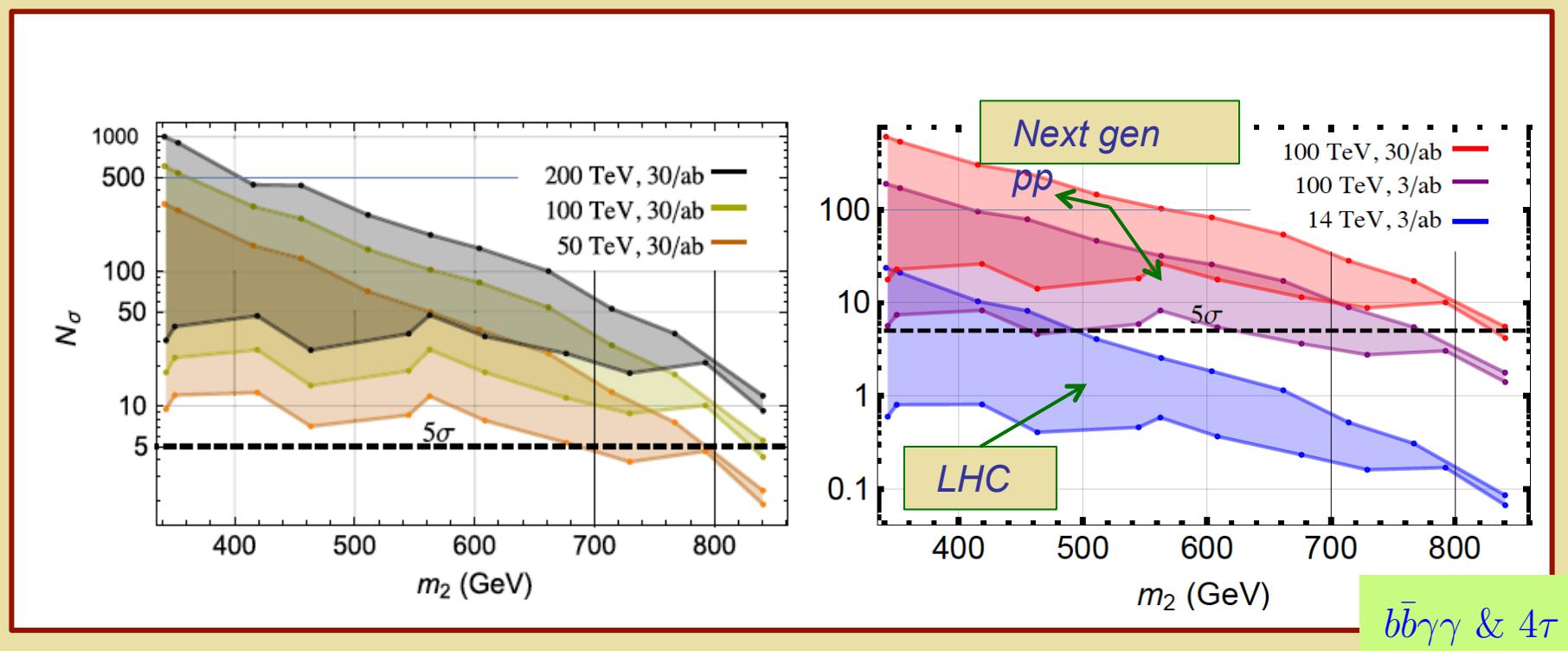
Standard Model Effective Potential: High- T expansion

$$V_{\text{eff}}(\phi, T) = D(T^2 - T_0^2)\phi^2 - ET\phi^3 + \frac{\bar{\lambda}}{4}\phi^4 + \dots$$

- *Potential is gauge-dependent*
- T_C is gauge-independent: Nielsen identities
- Detailed discussion: H. Patel & MJRM:
[1101.4665 \[hep-ph\]](https://arxiv.org/abs/1101.4665)

EW Phase Transition: Singlet Scalars

SFOEWPT Benchmarks: Resonant di-Higgs

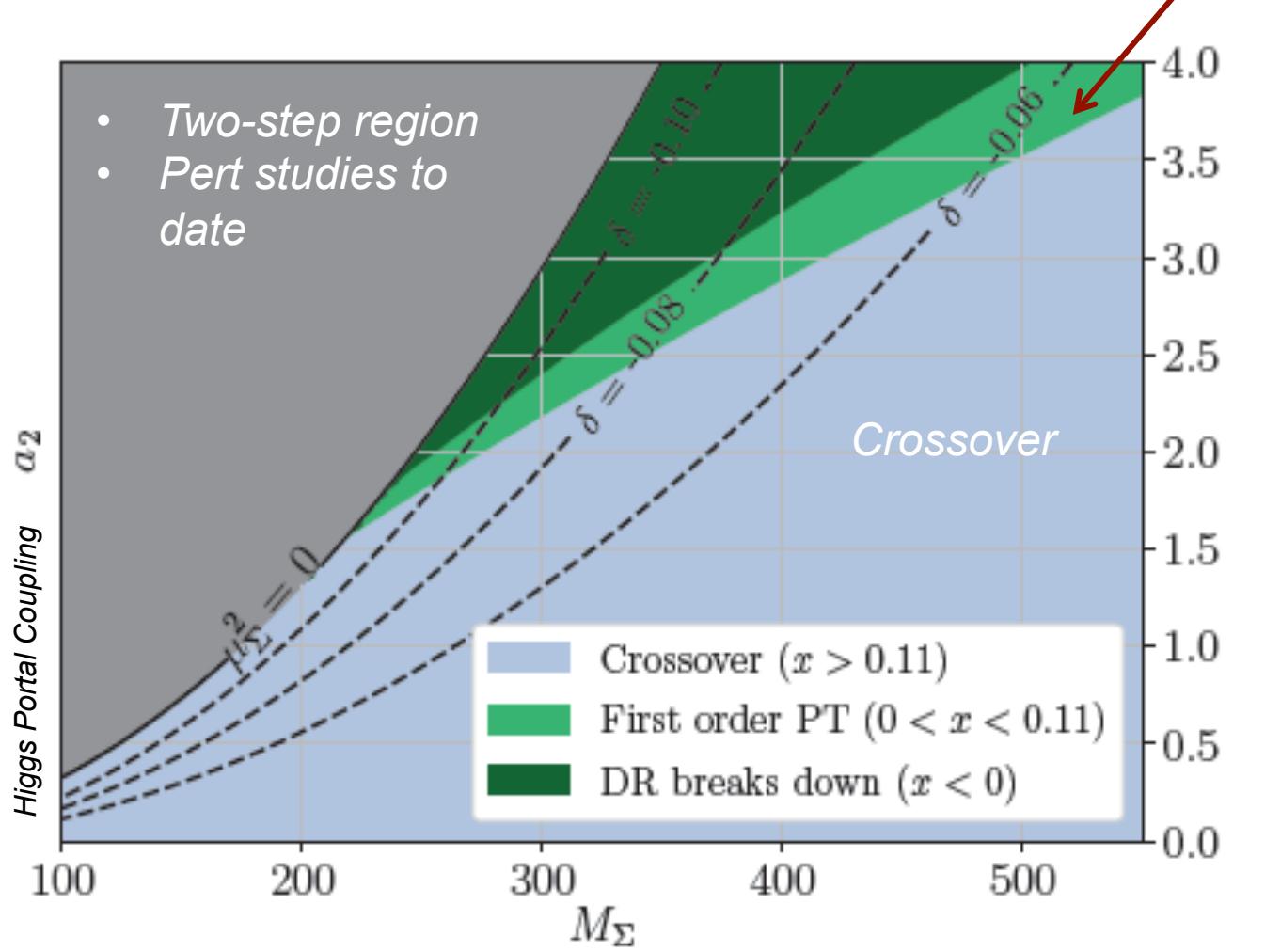


Kotwal, No, R-M, Winslow 1605.06123

See also: Huang et al, 1701.04442

Real Triplet & EWPT

FOEWPT



Real Triplet & EWPT

