Neutrino masses

### Neutrino masses are small

- Neutrino have masses (oscillations)
- Natural scale of fermion masses:  $\langle H \rangle \approx 174$  GeV
- Why  $m_v / \langle H \rangle < 10^{-12}$ ?
- Must have a different origin than m<sub>e</sub> /  $\langle H \rangle \approx 0.3 \times 10^{-5}$ 
	- quantitatively larger hierarchy
	- family independent
	- compelling understanding available



**P** Reminder: most general mass term with  $\psi_1, \dots, \psi_n$ 

$$
\frac{m_{ij}}{2}\psi_i\psi_j + \text{h.c.} \qquad \text{(only gauge invariant terms)}
$$

$$
(\psi_1 \psi_2 = \psi_2 \psi_1 = \psi_1^{\alpha} \epsilon_{\alpha \beta} \psi_2^{\beta})
$$

### Neutrino masses in QED + QCD

- Elementary L-handed fermions (1 family) and do und und entity of the eclion v
	- Q -1/3 1/3 2/3 -2/3 -1 1 0

 $SU(3)_c$  3  $3^*$  3  $3^*$  1 1 1

- Most general invariant mass terms:  $m_d d^c d + m_u u^c u + m_e e^c e +$  $m_\nu$ 2  $\nu\nu$
- Both charged leptons and neutrinos are equally allowed to get a mass term (although of different type)

### Neutrino masses in the SM

- $d^{c_i}$   $u^{c_i}$   $L_i$  $Q_i$ • Elementary L-handed fermions  $e^{c}$  $3 \t 3^* \t 3^* \t 1$  $SU(3)$  $\overline{1}$  $SU(2)$  2 1 1 2  $\mathbf 1$  $1/6$   $1/3$   $-2/3$   $-1/2$ Y  $\vert$  1
- · Most general invariant mass terms: none

$$
m_d = \lambda_d v
$$
  
\n• After EWSB 
$$
m_u = \lambda_u v
$$

$$
m_\nu = 0
$$
 a SM success  
\n
$$
m_e = \lambda_e v
$$

- $m_v = 0$  in the SM is a nice starting point but
- $m_v \neq 0$  needs extra ingredients
- 2 main options:

general compelling understanding of the smallness of m<sub>ν</sub>

- 1. the new ingredients live at  $M \gg M_Z$  (example: see-saw)
- 2. the new ingredients live at  $M \leq M_Z$  (example: Dirac neutrinos)

Option 1 M » Mz

# **Theorem (reminder)**

- **\*** The effect of any high scale [M » Mz] physics [responsible for **neutrino masses] can be described at low E by effective interactions involving only light [SM] dofs and symmetries (no need to know the microscopic theory and dofs) suppressed by M**
- **Example: SM interactions can be described at E « MZ by effective Fermi interaction involving only light fermions**

### Neutrino masses



- Compelling:
	- An elegant, economical, and model-independent understanding of the smallness of neutrino masses in terms of the heaviness of the scale at which L is violated.
- What makes neutrinos special?
	- They are the only fermions in the SM for which a mass does not arise (after EWSB) from a renormalisable interaction with the Higgs fields. They turn out to be Majorana.

• BUT:

• Could not v have a light  $v^c$  partner as all other SM fermions?



## "Right-handed" neutrinos

 $\nu^c$ *ec*  $\left(\begin{array}{c}u^c\u^c\end{array}\right) \left(\begin{array}{c}\nu^c\v^c\end{array}\right) \left(\begin{array}{c}\nu^c\u^c\end{array}\right)$  SU(3) THE SU(3) THE SU(2) THE MEX UNE DIRECTION  $\mathbf{Z}$ *e*  $u^c$   $\Big)$   $\Big)$ *dc*  $u<sub>1</sub>$ *d*  $\sqrt{2}$ 

## "Right-handed" neutrinos

$$
\left(\begin{array}{c}u\\d\end{array}\right) \qquad \begin{array}{c}u^c\\d^c\end{array}\qquad \left(\begin{array}{c}\nu\\e\end{array}\right) \qquad \begin{array}{c} \nu^c\\e^c\end{array}
$$

 $SU(3)_C \times SU(2)_W \times U(1)_Y$ 

 $\lambda_{\nu} \nu_{c} LH \rightarrow m_{\nu} = \lambda_{\nu} v$ 

(like the other fermions)

ν<sub>c</sub> is a SM singlet and can therefore be heavy

$$
{\cal L}_{\rm HE}\supset -\frac{M}{2}\nu^c\nu^c
$$

*L*unlike the other fermions)

### Neutrino masses in the  $SM + v^c$

 $d^c$ i u<sup>c</sup>i Li e<sup>c</sup>i • Elementary L-handed fermions  $Q_i$  $V^C$  $3^*$   $3^*$  1 1  $SU(3)$  $3<sup>7</sup>$  $\mathbf 1$ SU(2) 2 1 1 2 1  $\mathbf 1$  $1/6$   $1/3$   $-2/3$   $-1/2$  1 Y  $\Omega$ 

· Most general invariant mass terms:

$$
\frac{M_{ij}}{2} \nu_i^c \nu_j^c
$$



### exercise

$$
\mathcal{L}(L_i, H, \nu_i^c) = \lambda_{ij} \nu_i^c L_j H + \frac{M_{ij}}{2} \nu_i^c \nu_j^c
$$
  
e.o.m. (neglect kinetic term):  $\frac{\partial \mathcal{L}}{\partial \nu_i^c} = 0$   

$$
\mathcal{L}_{\text{eff}}(L_i, H) = ?
$$



# **Theorem (reminder)**

- **\*** The effect of any high scale [M » Mz] physics [responsible for **neutrino masses] can be described at low E by effective interactions involving only light [SM] dofs and symmetries (no need to know the microscopic theory and dofs) suppressed by M**
- **Example: SM interactions can be described at E « MZ by effective Fermi interaction involving only light fermions**
- **If the higher E theory is known, the specific form of the NR remnants can be derived**
- **If the higher E theory is unknown: i) model-independent parameterization of NP, ii) the experimental determination of effective interactions tells us about the microscopic theory** 
	- **• e.g.: Fermi Gamow-Teller Sudarshan-Marshak interaction**

$$
\frac{G_F}{\sqrt{2}} \bar{\psi}_1 \Gamma^A \psi_2 \bar{\psi}_3 \Gamma_A \psi_4 \longrightarrow \frac{G_F}{\sqrt{2}} \bar{\psi}_1 \gamma^\mu (1 - \gamma_5) \psi_2 \bar{\psi}_3 \gamma_\mu (1 - \gamma_5) \psi_4 \longrightarrow \text{SM}
$$

# Possible (tree level) origins of

 $v_L \longrightarrow V \longrightarrow$ 

h h

 $c/\Lambda$ 

h March 1999

ν<sup>L</sup> ν<sup>L</sup>

h

# Possible (tree level) origins of

See-saw type I S: SM singlet

See-saw type III T:  $SU(2)_L$  triplet,  $Y = 0$  E « M

 $\ket{\overline{\phantom{0}}}$ 

 $v_L \longrightarrow V \longrightarrow$ c/Λ

h h

 $v_L \longrightarrow V \longrightarrow$ 

h h

c/Λ



 $\mathsf T$ 

S S  ${\color{blue}\nu_\mathsf{L}} \begin{picture}(60,6) \put(0,0){\dashbox{0.5}(7.0) \put(0,0){\$ 

 $v_{\mathsf{L}} \longrightarrow \mathsf{L} \longrightarrow \mathsf{L}$  is a set of  $\mathsf{L}$ 

 $\boldsymbol{\times}$ 

M

h h

 $\boldsymbol{\times}$ 

M

h h

See-saw type II T:  $SU(2)_L$  triplet,  $Y = 1$  Option 2:  $M \leq M_Z$ 

### Example: Dirac neutrinos

Lepton number is "exactly" conserved:  $h_{ij} = 0$  $\mathcal{O}$ 

Neutrino masses then need an  $L = -1$  neutrino  $v^c$  $\bullet$ 

 $m_{ij}^{N}\nu_{i}^{c}\nu_{j}+m_{ij}^{E}e_{i}^{c}e_{j}+m_{ij}^{D}d_{i}^{c}d_{j}+m_{ij}^{U}u_{i}^{c}u_{j}+\mathrm{h.c.}$ 

In the SM:

 $m_{ij}^N = \lambda_{ij}^N v \quad m_{ij}^E = \lambda_{ij}^Ev \quad m_{ij}^D = \lambda_{ij}^D v \quad m_{ij}^U = \lambda_{ij}^U v$  $\mathcal{L}^{\text{flavor}}_{\text{SM}} = \lambda^N_{ij} \nu^c_i L_j H + \lambda^E_{ij} e^c_i L_j H^{\dagger} + \lambda^U_{ij} u^c_i Q_j H + \lambda^D_{ij} d^c_i Q_j H^{\dagger} + \text{h.c.}$  $\begin{array}{rcl} \displaystyle =m_{ij}^N\nu_i^c\nu_j &\displaystyle +m_{ij}^Ee_i^ce_j &\displaystyle +m_{ij}^Uu_i^cu_j &\displaystyle +m_{ij}^Dd_i^cd_j &\displaystyle +\mathrm{h.c.}+\dots \end{array}$ 

 $\bullet$  Needs L and  $\lambda^N < 10^{-11}$ : why?

### Low scale lepton number violation

singlet neutrino mass M < v (EW scale)

effective description not sound anymore  $\bullet$ 

 $\phi$  λ ≈ 10<sup>-6</sup> (M/v)<sup>1/2</sup>

### Low-scale origin of L-violation (1)

TeV-scale see-saw  $\bullet$ 

νc with M ≈ TeV

Probe  $v^{\mathsf{c}}$  through  $\lambda \, \nu^c L H \colon \; m_{\nu} = - m_D^T$ 1  $\frac{1}{M}m_D, \quad m_D=\lambda \langle H \rangle$ 

M ~ TeV  $\Rightarrow \lambda =$ too small for LHC *m<sup>D</sup>*  $\langle H \rangle$  $\sim 10^{-6}$  $\left( m_{\nu} \right)$ 0*.*05 eV  $\bigwedge^{1/2}$  *( M* TeV  $\setminus$ <sup>1/2</sup>

Unless  $\lambda$   $\gg$  10<sup>-7</sup> + cancellations in  $\; m_{\nu} = - m_D^T \;$ (2 or more  $v c' s$ ) 1  $\frac{1}{M}m_D$ 

"magical", e.g.:  $m_\nu = 0 + {\rm corrections}$  if  $m_{nj}^D = \alpha_n \beta_j m_0, \quad M_R = \text{Diag}(M_1 \dots M_n), \quad \sum \alpha_n^2 M_n = 0$ *n*

natural, e.g.:

 $L_e, L_\mu, L_\tau, (\nu_R)_1 \equiv N$  have L = 1,  $(\nu_R)_2 \equiv N'$  has L = -1

# Low-scale origin of L-violation (2)

#### (Rp-violating) supersymmetry Ø

Supersymmetry does not guarantee (accidental) L (or B) Ø conservation, unlike the SM:  $H_d \approx L_i$ 

> $\overline{W} = \lambda_{ij}^U u_i^c Q_j H_u + \lambda_{ij}^D d_i^c Q_j H_d + \lambda_{ij}^E e_i^c L_j H_d + \mu H_u H_d$  $+ \lambda_{ijk}^{\prime\prime}u_{i}^{c}d_{j}^{c}d_{k}^{c} + \lambda_{ijk}^{\prime}L_{i}Q_{j}d_{k}^{c} + \lambda_{ijk}L_{i}L_{j}e_{k}^{c} + \mu_{i}H_{u}L_{i}$

 $\mathcal{L}_{\text{soft}} = A^U_{ij} \tilde{u}^c_i \tilde{Q}_j H_u + A^D_{ij} \tilde{d}^c_i \tilde{Q}_j H_d + A^E_{ij} \tilde{e}^c_i \tilde{L}_j H_d + B \mu H_u H_d$  $+A_{ijk}^{\prime\prime}\tilde{u}_{i}^{c}\tilde{d}_{j}^{c}\tilde{d}_{k}^{c}+A_{ijk}^{\prime}\tilde{L}_{i}\tilde{Q}_{j}\tilde{d}_{k}^{c}+A_{ijk}\tilde{L}_{i}\tilde{L}_{j}\tilde{e}_{k}^{c}+(B\mu)_{i}H_{u}\tilde{L}_{i}$ 

 $+ \tilde{m}_Q^2 \tilde{Q}^{\dagger} \tilde{Q} + (\tilde{m}_i^2 H_d^{\dagger} \tilde{L}_i + \text{h.c.}) +$  gaugino masses

L and B violating terms controlled by  $R_P = (-1)^{3(B-L)+2s}$ Ø

A small  $R_P$  breaking: ø

induces  $(h_{ij}/\Lambda)$ L<sub>i</sub>LjHH, with  $\Lambda$  =  $\mathsf{\widetilde{m}},$  h  $\leftrightarrow$  small R<sub>P</sub> breaking

makes the LSP unstable (could be any susy partner)  $\mathcal{O}$ 

### Majorana o Dirac?

### Dirac vs Majorana at the particle level

- -<br>5 QFT: each particle  $p$  corresponds to an antiparticle  $\bar{p}$  (CPT)  $\mathcal{O}$
- Particle and antiparticle have same mass, spin (opposite  $\bullet$ helicities if m = 0), opposite conserved charges
- CPT does not forbid  $p = \bar{p}$ ; however this cannot happen if the particle is charged or, if  $s \neq 0$ , massless  $\overline{a}$
- $\circ$  If Q = 0 and m  $\neq$  0 a fermion can coincide with its antiparticle (Majorana fermion) or be independent of it (Dirac fermion)
- Neutrinos are the only known fermions that can be Majorana or Dirac, all other known fermions are Dirac
- The mass term of a Majorana fermion violates all charges จ carried by the fermion
- If neutrinos are Majorana, lepton number is violated จ

### Dirac vs Majorana at the particle level

#### $m = 0$

- the helicity is an invariant of a 1-particle state (1 dof)
- neutrino particles have 1 dof
- $CC$  interactions only produce  $|\nu-\rangle$  and  $|\overline{\nu}+\rangle$  (L chirality)  $\mathcal{O}$

#### $\circ$  m  $\neq$  0

helicity depends on the reference frame, one particle states have 2 dofs corresponding to two opposite helicities



### Dirac vs Majorana at the field level

 $\psi_i$  "L" fermions

Most general mass term:

 $\frac{m_{ij}}{2}\psi_i\psi_j$ 

*mij* symmetric

*m* 2

 $\psi$ 

"Majorana" breaks any charge of ψ

example:  $ψ = ν<sub>L</sub>$ m ≠ 0 allowed by QED+QCD breaks lepton number

### Dirac vs Majorana at the field level

 $\psi_i$  "L" fermions

Most general mass term:

$$
\frac{m_{ij}}{2}\psi_i\psi_j
$$

*mij* symmetric

*m*<sup>1</sup>  $\frac{1}{2} \psi \psi +$ *m*<sup>2</sup>  $\frac{v_2}{2} \overline{\psi^c}$  $\widetilde{\psi}^c + m \psi^c$  $\psi$ 

 $|\overline{\psi}, \overline{\psi^c}|$ 

"Majorana" "Dirac"  $\overline{Q(\psi)} + \overline{Q(\psi^c)} = 0$ (all charged SM fermions)

(e.g. Dirac neutrino mass term:  $m \overline{\nu_R} \nu_L$  $\psi^c = \overline{\nu_R} \quad \psi = \nu_L$ needs  $\nu_R$ does not break lepton number)

 $(e.g.$  electron mass term:  $m \overline{e_R} e_L$  $\psi^c = \overline{e_R} \quad \psi = e_L$ 

## Dirac vs Majorana in the interaction

- The difference shows up only in the  $m \neq 0$  case:  $\ast$ 
	- Dirac  $(m = 0)$

 $\overline{\nu_L}|0\rangle = |\nu - \rangle$   $\nu_L|0\rangle = |\overline{\nu} + \rangle$ 

• Majorana  $(m = 0)$ 

$$
\overline{\nu_L}|0\rangle = |\nu - \rangle \qquad \qquad \nu_L|0\rangle = |\nu + \rangle
$$

- In oscillations, once the O(m/E) terms have been neglected:  $\ast$ 
	- the helicity does not play a role
	- there is no L-violation
	- oscillation formulae are identical for Dirac and Majorana ν's

# Dirac vs Majorana (particle content)

- The difference shows up only in the  $m \neq 0$  case:  $\ast$ 
	- Dirac  $(m \neq 0)$

 $\overline{\nu_L}|0\rangle = |\nu -\rangle + \mathcal{O}(m/E)|\nu +\rangle$ 

• Majorana  $(m \neq 0)$ 

 $\overline{\nu_L}|0\rangle = |\nu - \rangle + \mathcal{O}(m/E) |\nu + \rangle$ 

$$
\nu_L|0\rangle = |\overline{\nu} + \rangle + \mathcal{O}\left(m/E\right)|\overline{\nu} - \rangle
$$

$$
\nu_L|0\rangle=|\nu+\rangle{+}\mathcal{O}\left(m/E\right)|\nu-\rangle
$$

- In oscillations, once the O(m/E) terms can be neglected:  $\ast$ 
	- the helicity does not play a role
	- there is no L-violation
	- oscillation formulae are identical for Dirac and Majorana ν's

Ov2ß decay



Violates lepton number Needs Majorana neutrinos

# 0ν2β decay

$$
(A, Z) \rightarrow (A, Z + 2) + 2e^{-}; e.g.: {}^{76}\text{Ge} \rightarrow {}^{76}\text{Se} + 2e^{-}
$$

$$
\Gamma \propto |m_{ee}|^2 \langle Q \rangle^2
$$

**Depends on** 

- **Phases**
- **Nuclear matrix elements**
- **Dirac vs Majorana**

 $m_{ee} = U_{eh}^2 m_h = c_{13}^2 (m_1 c_{12}^2 + m_2 s_{12}^2 e^{2 i \alpha}) + m_3 s_{13}^2 e^{2 i \beta'}$ 



 $|m_{ee}| < O(1) \times 0.2$  eV (Heidelberg-Moscow)  $\rightarrow O(1) \times 0.01$  eV (Genius)



### Neutrino masses are small

### Whatrielse do we known?

- Natural scale of fermion masses:  $\langle H \rangle \approx 174$  GeV
- Why  $m_v / \langle H \rangle < 10^{-12}$ ?
- Must have a different origin than m<sub>e</sub> /  $\langle H \rangle \approx 0.3 \times 10^{-5}$ 
	- quantitatively larger hierarchy
	- family independent
	- compelling understanding available



### Neutrino (lepton) flavour parameters

• Quark sector

$$
m_{ij}^D \overline{d_{iR}} d_{jL} + m_{ij}^U \overline{u_{iR}} u_{jL} = m_{di} \overline{d'_{iR}} d'_{iL} + m_{ui} \overline{u'_{iR}} u'_{iL}
$$

$$
j_{c,\text{had}}^\mu = \overline{u_{iL}} \gamma^\mu d_{iL} = V_{ij} \overline{u'_{iL}} \gamma^\mu d'_{jL}
$$

• Lepton sector, including neutrino masses (Majorana for definitess)

$$
\frac{m_{ij}^{\nu}}{2}\nu_{iL}\nu_{jL} + m_{ij}^{E}\overline{e_{iR}}e_{jL} = \frac{m_{\nu_{i}}}{2}\nu_{iL}'\nu_{iL}' + m_{e_{i}}\overline{e'_{iR}}e'_{iL}
$$

$$
j_{c,\text{lep}}^{\mu} = \overline{\nu}_{iL}\gamma^{\mu}e_{iL} = U_{ij}^{\dagger}\overline{\nu_{iL}'}\gamma^{\mu}e'_{jL}
$$

### Physical flavour parameters in the lepton sector

$$
U = \begin{pmatrix} \frac{m_{\nu_i}}{2} \nu_{iL} \nu_{iL} + m_{e_i} \overline{e_{iR}} e_{iL} & j_{c,\text{lep}}^{\mu \dagger} = U_{ij} \overline{e}_{iL} \gamma^{\mu} \nu_{jL} \\ e^{i\gamma_1} & e^{i\gamma_2} \end{pmatrix} \begin{pmatrix} \text{standard par.} \\ e^{i\alpha} & e^{i\beta} \end{pmatrix}
$$
  
\nunphysical  
\n9 = 3 + 3 + 1 + 2

### Physical flavour parameters in the lepton sector

$$
\frac{m_{\nu_i}}{2} \nu_{iL} \nu_{iL} + m_{e_i} \overline{e_{iR}} e_{iL} \qquad j_{c,\text{lep}}^{\mu \dagger} = U_{ij} \overline{e}_{iL} \gamma^{\mu} \nu_{jL}
$$

$$
m_e, m_\mu, m_\tau, m_{\nu_1}, m_{\nu_2}, m_{\nu_3}, \theta_{23}, \theta_{12}, \theta_{13}, \delta, \alpha, \beta
$$

$$
U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23}-c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23}-s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23}-c_{12}c_{23}s_{13}s^{i\delta} & -c_{12}s_{23}-s_{12}c_{23}s_{13}s^{i\delta} & c_{23}c_{13} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha} & 0 \\ 0 & 0 & e^{i\beta} \end{pmatrix}
$$

$$
0\leq \theta_{23}, \theta_{12}, \theta_{13}\leq \frac{\pi}{2}, \quad 0\leq \delta < 2\pi, \quad 0\leq \alpha, \beta < \pi
$$

### Standard labeling of eigenstates

 $0 < \Delta m^2_{12} < |\Delta m^2_{23}|$  uniquely defines the labeling  $\Delta m^2_{12} > 0$  by definition;  $\Delta m^2_{23}$  can have both signs  $\delta m^2 \equiv \Delta m^2_{12}$  $\delta m^2 \equiv \Delta m^2_{12}$ <br>  $\Delta m^2 \approx \Delta m^2_{23} \approx \Delta m^2_{13}$ 





Accessible to oscillations

Charged sector

 $\delta m^2$ 

 $|m_{e,\mu,\tau}|$ 

 $|\Delta m^2|$  $sign(\Delta m^2)$  $\theta_{12}, \theta_{23}, \theta_{13}, \delta$  Not accessible to oscillations

*m*lightest  $\alpha$  $\beta$ 

 $(\Delta m_{ij}^2 \equiv m_{\nu_j}^2 - m_{\nu_i}^2)$  $\Delta m_{13}^2 = \Delta m_{12}^2 + \Delta m_{23}^2$  $\delta m^2 \equiv \Delta m_{12}^2 \ll |\Delta m_{23}^2| \Rightarrow \Delta m_{23}^2 \approx \Delta m_{13}^2 \approx \Delta m^2 \equiv (\Delta m_{13}^2 + \Delta m_{12}^2)/2$ 

Accessible to oscillations

> Not accessible to oscillations

 $m_{\rm lightest}$ 

 $\alpha$ 

 $\beta$ 

Charged sector

 $|m_{e,\mu,\tau}|$ 

 $\delta m^2$  $\overline{|\Delta m^2|}$ 

 $sign(\Delta m^2)$ 

Well known

 $\overline{\theta_{12}, \theta_{23}, \theta_{13}}, \overline{\delta}$ 

**Bounds** 

**Hints** 

Known

$$
|\Delta m^2| \approx 2.4 \times 10^{-3} \,\text{eV}^2 \quad \theta_{23} \sim 45^\circ
$$

$$
\delta m^2 \approx 0.73 \times 10^{-4} \,\text{eV}^2 \quad \theta_{12} \approx 33^\circ
$$

$$
\theta_{13} \approx 8^\circ
$$

 $sign(\Delta m^2) > 0$  preferred at  $3\sigma$  $\delta \sim 3\pi/2 \ \pm 15\%$ 

 $|m_{ee}| \equiv |U_{ei}^2 m_{\nu_i}| < \mathcal{O}(1) \times 0.2 \,\text{eV}$  $(m^{\dagger}m)_{ee} = |U_{ei}|^2 m_{\nu_i}^2 < (2 \,\text{eV})^2$  $\sum m_{\nu_i} \lesssim 0.7$  eV (priors)

double-beta decay

beta decay

cosmology

#### Guidelines for theory:

 $m_{\nu_i} \ll 174 \,\text{GeV}$  $\theta_{23} \sim 45^{\circ} (= 45^{\circ}$ ?)  $\theta_{12} \approx 33^\circ \neq 45^\circ$  $\theta_{13}$  not so small  $\delta m^2/|\Delta m^2| \approx 0.03 \ll 1$ normal ordering? large CP-violation?

### Thank you [romanino@sissa.it](mailto:romanino@sissa.it)