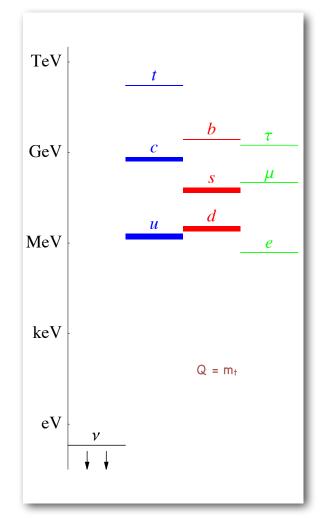
Neutrino masses

Neutrino masses are small

- Neutrino have masses (oscillations)
- Natural scale of fermion masses: $\langle H \rangle \approx 174 \text{ GeV}$
- Why $m_v / \langle H \rangle < 10^{-12}$?
- Must have a different origin than $m_e / \langle H \rangle \approx 0.3 \times 10^{-5}$
 - quantitatively larger hierarchy
 - family independent
 - compelling understanding available



• Reminder: most general mass term with $\psi_1, ..., \psi_n$

$$\frac{m_{ij}}{2}\psi_i\psi_j + \text{h.c.}$$
 (only gauge invariant terms)

$$(\psi_1\psi_2 = \psi_2\psi_1 = \psi_1^{\alpha}\epsilon_{\alpha\beta}\psi_2^{\beta})$$

Neutrino masses in QED + QCD

- Elementary L-handed fermions (1 family)
 d
 d^c
 u
 u^c
 e
 e^c
 v
 - Q -1/3 1/3 2/3 -2/3 -1 1 0

SU(3)_c 3 3* 3 3* 1 1 1

- Most general invariant mass terms: $m_d d^c d + m_u u^c u + m_e e^c e + \frac{m_\nu}{2} \nu \nu$
- Both charged leptons and neutrinos are equally allowed to get a mass term (although of different type)

Neutrino masses in the SM

- d^ci u^ci Li • Elementary L-handed fermions Qi **e**^Ci 3* 3* 1 3 SU(3) 1 SU(2) 2 1 1 2 1 1/6 1/3 -2/3 -1/2 Υ 1
- Most general invariant mass terms: none

• After EWSB
$$m_u = \lambda_u v$$
 $m_{\nu} = 0$ a SM success $m_e = \lambda_e v$

- $m_v = 0$ in the SM is a nice starting point but
- $m_v \neq 0$ needs extra ingredients
- 2 main options:

general compelling understanding of the smallness of m_v

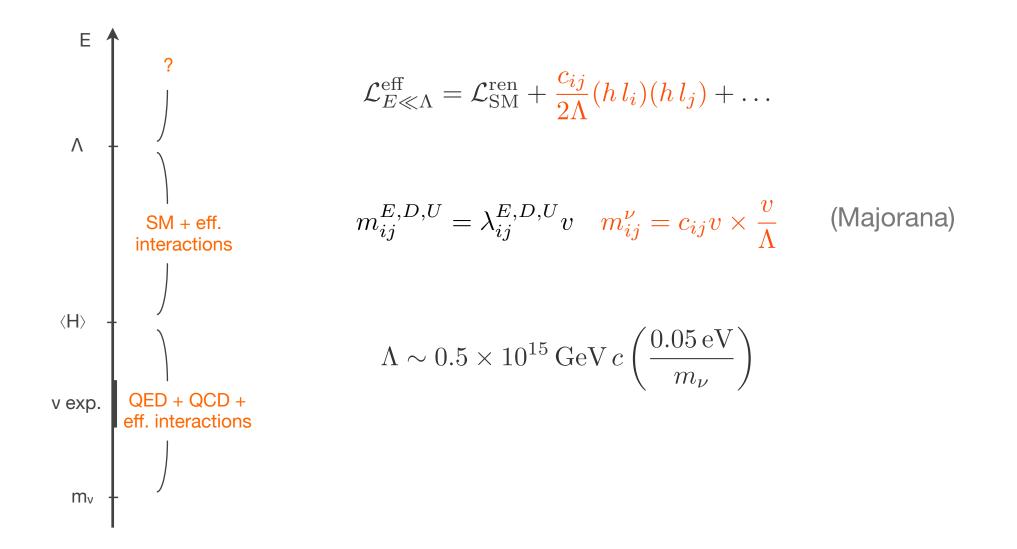
- 1. <u>the new ingredients live at M » Mz (example: see-saw)</u>
- 2. the new ingredients live at $M \leq M_Z$ (example: Dirac neutrinos)

Option 1: M » M_Z

Theorem (reminder)

- ★ The effect of any high scale [M » Mz] physics [responsible for neutrino masses] can be described at low E by effective interactions involving only light [SM] dofs and symmetries (no need to know the microscopic theory and dofs) suppressed by M
- Example: SM interactions can be described at E « Mz by effective Fermi interaction involving only light fermions

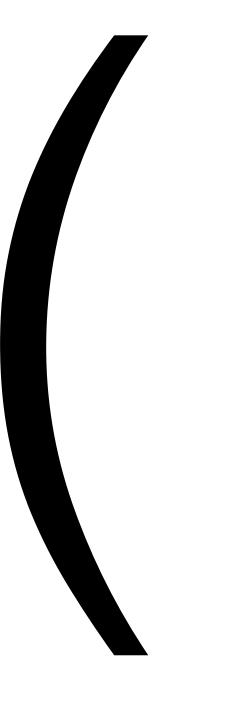
Neutrino masses



- Compelling:
 - An elegant, economical, and model-independent understanding of the smallness of neutrino masses in terms of the heaviness of the scale at which L is violated.
- What makes neutrinos special?
 - They are the only fermions in the SM for which a mass does not arise (after EWSB) from a renormalisable interaction with the Higgs fields. They turn out to be Majorana.

• BUT:

• Could not v have a light v^c partner as all other SM fermions?



"Right-handed" neutrinos

 $\begin{pmatrix} u \\ d \end{pmatrix} \begin{pmatrix} u^c \\ d^c \end{pmatrix} \begin{pmatrix} \nu \\ e \end{pmatrix} \begin{pmatrix} \nu^c \\ e^c \end{pmatrix} \qquad SU(3)SU(3)SU(3)SU(2)SU(2)SU(2)SU(2))W(2)U(1)_{B-L}$

"Right-handed" neutrinos

$$\left(egin{array}{cc} u \ d \end{array}
ight) \qquad egin{array}{cc} u^c & \left(egin{array}{cc}
u \ e \end{array}
ight) \qquad egin{array}{cc}
u^c \ e \end{array}
ight) \qquad egin{array}{cc}
u^c \ e \end{array}
ight) \qquad egin{array}{cc}
u^c \ e \end{array}
ight)$$

 $SU(3)_C \times SU(2)_W \times U(1)_Y$

 $\lambda_{\nu}\nu_{c}LH \to m_{\nu} = \lambda_{\nu}v$

(like the other fermions)

 v_c is a SM singlet and can therefore be heavy

$$\mathcal{L}_{
m HE} \supset -rac{M}{2}
u^c
u^c$$

(unlike the other fermions)

Neutrino masses in the SM + v^c

Elementary L-handed fermions
 Qi
 d^ci
 u^ci
 Li
 e^ci
 V^ci
 SU(3)
 3^{*}
 3^{*}
 1
 1
 1
 SU(2)
 2
 1
 3^{*}
 1
 1
 1
 1

• Most general invariant mass terms:

$$\frac{M_{ij}}{2}\nu_i^c\nu_j^c$$



H

Μ

 \times

vc

Integrate out v^c : L

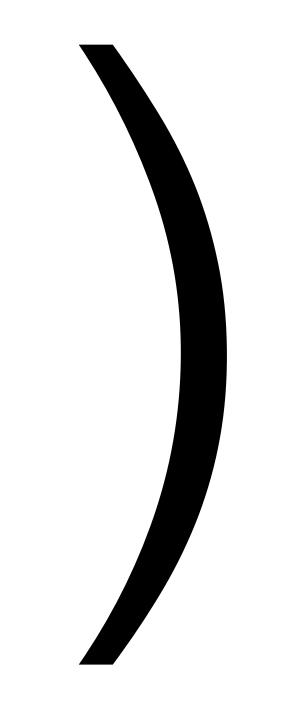
 $\begin{aligned} &\frac{c}{\Lambda}(HL)(HL)\\ &\frac{c}{\Lambda}=-\lambda^T\frac{1}{M}\lambda\\ &\hline m_\nu=-m_{\rm D}^T\frac{1}{M}m_{\rm D} \end{aligned} \qquad {\rm Majorana} \end{aligned}$

Η

exercise

$$\mathcal{L}(L_i, H, \nu_i^c) = \lambda_{ij} \nu_i^c L_j H + \frac{M_{ij}}{2} \nu_i^c \nu_j^c$$

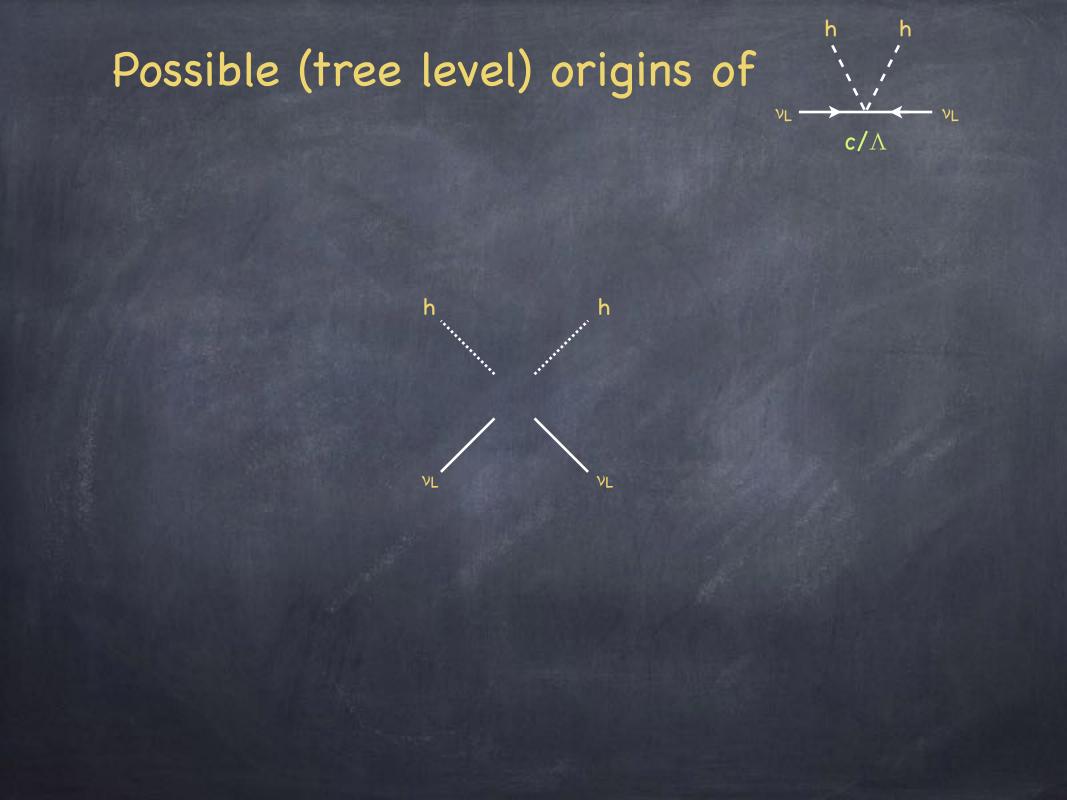
e.o.m. (neglect kinetic term): $\frac{\partial \mathcal{L}}{\partial \nu_i^c} = 0$
 $\mathcal{L}_{\text{eff}}(L_i, H) = ?$



Theorem (reminder)

- ★ The effect of any high scale [M » Mz] physics [responsible for neutrino masses] can be described at low E by effective interactions involving only light [SM] dofs and symmetries (no need to know the microscopic theory and dofs) suppressed by M
- ★ Example: SM interactions can be described at E « M_Z by effective Fermi interaction involving only light fermions
- If the higher € theory is known, the specific form of the NR remnants can be derived
- If the higher E theory is unknown: i) model-independent parameterization of NP, ii) the experimental determination of effective interactions tells us about the microscopic theory
 - e.g.: Fermí Gamow-Teller Sudarshan-Marshak interaction

$$\frac{G_F}{\sqrt{2}}\bar{\psi}_1\Gamma^A\psi_2\bar{\psi}_3\Gamma_A\psi_4 \implies \frac{G_F}{\sqrt{2}}\bar{\psi}_1\gamma^\mu(1-\gamma_5)\psi_2\bar{\psi}_3\gamma_\mu(1-\gamma_5)\psi_4 \implies \mathsf{SM}$$



Possible (tree level) origins of

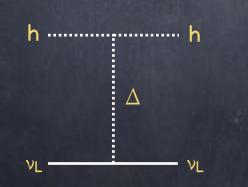
See-saw type I S: SM singlet

See-saw type III T: SU(2)_L triplet, Y = 0 E « M

 $\begin{array}{c}
\mathbf{h} \quad \mathbf{h} \\
\mathbf{\dot{v}} \\
\mathbf{\dot{v}} \\
\mathbf{v}_{L} \\
\mathbf{\dot{v}} \\
\mathbf{\dot$

h

 c/Λ



h

h

νL

VL

M

M

S

S

Т

h

h

νL

νL

See-saw type II T: SU(2)∟ triplet, Y = 1 Option 2: $M \leq M_Z$

Example: Dirac neutrinos

Lepton number is "exactly" conserved: h_{ij} = 0

• Neutrino masses then need an L = -1 neutrino v^c

 $m_{ij}^{N} \nu_{i}^{c} \nu_{j} + m_{ij}^{E} e_{i}^{c} e_{j} + m_{ij}^{D} d_{i}^{c} d_{j} + m_{ij}^{U} u_{i}^{c} u_{j} + \text{h.c.}$

In the SM:

 $\mathcal{L}_{SM}^{\text{flavor}} = \lambda_{ij}^{N} \nu_{i}^{c} L_{j} H + \lambda_{ij}^{E} e_{i}^{c} L_{j} H^{\dagger} + \lambda_{ij}^{U} u_{i}^{c} Q_{j} H + \lambda_{ij}^{D} d_{i}^{c} Q_{j} H^{\dagger} + \text{h.c.}$ $= m_{ij}^{N} \nu_{i}^{c} \nu_{j} + m_{ij}^{E} e_{i}^{c} e_{j} + m_{ij}^{U} u_{i}^{c} u_{j} + m_{ij}^{D} d_{i}^{c} d_{j} + \text{h.c.} + \dots$ $m_{ij}^{N} = \lambda_{ij}^{N} v \quad m_{ij}^{E} = \lambda_{ij}^{E} v \quad m_{ij}^{D} = \lambda_{ij}^{D} v \quad m_{ij}^{U} = \lambda_{ij}^{U} v$

• Needs L and $\lambda^{N} < 10^{-11}$: why?

Low scale lepton number violation

singlet neutrino mass M < v (EW scale)</p>

effective description not sound anymore

• $\lambda \approx 10^{-6} (M/v)^{1/2}$

Low-scale origin of L-violation (1)

TeV-scale see-saw

• v^c with M \approx TeV

• Probe v^c through $\lambda \,
u^c LH$: $m_
u = -m_D^T rac{1}{M} m_D$, $m_D = \lambda \, \langle H
angle$

• M ~ TeV
$$\Rightarrow \lambda = \frac{m_D}{\langle H \rangle} \sim 10^{-6} \Big(\frac{m_\nu}{0.05 \, {\rm eV}} \Big)^{1/2} \Big(\frac{M}{\rm TeV} \Big)^{1/2}$$
too small for LHC

• Unless $\lambda \gg 10^{-7}$ + cancellations in $m_{\nu} = -m_D^T \frac{1}{M} m_D$ (2 or more ν^c 's)

• "magical", e.g.: $m_{\nu} = 0 + \text{corrections}$ if $m_{nj}^D = \alpha_n \beta_j m_0, \quad M_R = \text{Diag}(M_1 \dots M_n), \quad \sum \alpha_n^2 M_n = 0$

natural, e.g.:

 $L_e, L_\mu, L_ au, (
u_R)_1 \equiv N$ have L = 1, $(
u_R)_2 \equiv N'$ has L = -1

Low-scale origin of L-violation (2)

(R_P-violating) supersymmetry

Supersymmetry does not guarantee (accidental) L (or B) conservation, unlike the SM: H_d ≈ L_i

$$W = \lambda_{ij}^U u_i^c Q_j H_u + \lambda_{ij}^D d_i^c Q_j H_d + \lambda_{ij}^E e_i^c L_j H_d + \mu H_u H_d$$
$$+ \lambda_{ijk}^{\prime\prime} u_i^c d_j^c d_k^c + \lambda_{ijk}^{\prime} L_i Q_j d_k^c + \lambda_{ijk} L_i L_j e_k^c + \mu_i H_u L_i$$

 $\mathcal{L}_{\text{soft}} = A_{ij}^U \tilde{u}_i^c \tilde{Q}_j H_u + A_{ij}^D \tilde{d}_i^c \tilde{Q}_j H_d + A_{ij}^E \tilde{e}_i^c \tilde{L}_j H_d + B\mu H_u H_d$ $+ A_{ijk}'' \tilde{u}_i^c \tilde{d}_j^c \tilde{d}_k^c + A_{ijk}' \tilde{L}_i \tilde{Q}_j \tilde{d}_k^c + A_{ijk} \tilde{L}_i \tilde{L}_j \tilde{e}_k^c + (B\mu)_i H_u \tilde{L}_i$

 $+ \tilde{m}_Q^2 \tilde{Q}^{\dagger} \tilde{Q} + (\tilde{m}_i^2 H_d^{\dagger} \tilde{L}_i + \text{h.c.}) + \text{gaugino masses}$

• L and B violating terms controlled by $R_P = (-1)^{3(B-L)+2s}$

A small R_P breaking:

• induces $(h_{ij}/\Lambda)L_iL_jHH$, with $\Lambda = \widetilde{m}$, $h \leftrightarrow \text{small } R_P$ breaking

makes the LSP unstable (could be any susy partner)

Majorana o Dirac?

Dirac vs Majorana at the particle level

- QFT: each particle p corresponds to an antiparticle p (CPT)
- Particle and antiparticle have same mass, spin (opposite helicities if m = 0), opposite conserved charges
- CPT does not forbid $p = \bar{p}$; however this cannot happen if the particle is charged or, if s \neq 0, massless
- If Q = 0 and m ≠ 0 a fermion can coincide with its antiparticle (Majorana fermion) or be independent of it (Dirac fermion)
- Neutrinos are the only known fermions that can be Majorana or Dirac, all other known fermions are Dirac
- The mass term of a Majorana fermion violates all charges carried by the fermion
- If neutrinos are Majorana, lepton number is violated

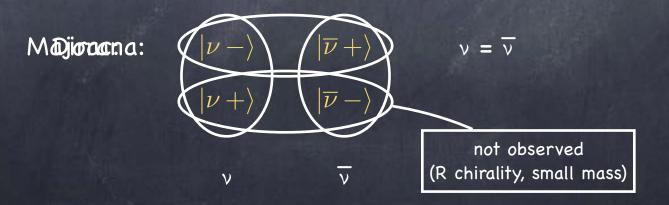
Dirac vs Majorana at the particle level

• m = 0

- the helicity is an invariant of a 1-particle state (1 dof)
- neutrino particles have 1 dof
- CC interactions only produce |
 u
 angle and $|\overline{
 u} +
 angle$ (L chirality)

⊙ m ≠ 0

 helicity depends on the reference frame, one particle states have 2 dofs corresponding to two opposite helicities



Dirac vs Majorana at the field level

 ψ_i "L" fermions

Most general mass term:

 $\frac{m_{ij}}{2}\psi_i\psi_j$

 m_{ij} symmetric

 $\frac{m}{2}\psi\psi$

 ψ

"Majorana" breaks any charge of ψ

example: ψ = ν_L m ≠ 0 allowed by QED+QCD breaks lepton number

Dirac vs Majorana at the field level

 ψ_i "L" fermions

Most general mass term:

$$\frac{m_{ij}}{2}\psi_i\psi_j$$

 m_{ij} symmetric

 $\frac{m_1}{2}\psi\psi + \frac{m_2}{2}\psi^c\psi^c + m\psi^c\psi$

 $|\psi,\psi^c|$

"Majorana"

"Dirac" Q(ψ) + Q(ψ^c) = 0 (all charged SM , fermions)

(e.g. Dirac neutrino mass term: $m \overline{\nu_R} \nu_L$ $\psi^c = \overline{\nu_R} \quad \psi = \nu_L$ needs ν_R does not break lepton number)

(e.g. electron mass term: $m \overline{e_R} e_L$ $\psi^c = \overline{e_R} \quad \psi = e_L$)

Dirac vs Majorana in the interaction

- * The difference shows up only in the $m \neq 0$ case:
 - Dirac (**m** = **0**)

 $\overline{\nu_L}|0\rangle = |\nu - \rangle \qquad \qquad \nu_L|0\rangle = |\overline{\nu} + \rangle$

• Majorana (m = 0)

$$\overline{\nu_L}|0\rangle = |\nu - \rangle \qquad \qquad \nu_L|0\rangle = |\nu + \rangle$$

- In oscillations, once the O(m/E) terms have been neglected:
 - the helicity does not play a role
 - there is no L-violation
 - oscillation formulae are identical for Dirac and Majorana v's

Dirac vs Majorana (particle content)

- * The difference shows up only in the $m \neq 0$ case:
 - Dirac (**m** ≠ **0**)

 $\overline{\nu_L}|0\rangle = |\nu - \rangle + \mathcal{O}(m/E)|\nu + \rangle$

• Majorana ($m \neq 0$)

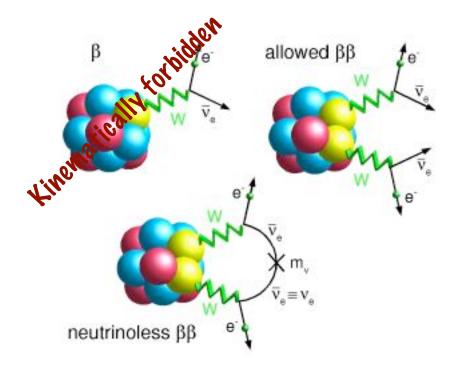
 $\overline{\nu_{L}}|0\rangle = |\nu - \rangle + \mathcal{O}\left(m/E\right)|\nu + \rangle$

$$\nu_L |0\rangle = |\overline{\nu} + \rangle + \mathcal{O}(m/E) |\overline{\nu} - \rangle$$

$$\nu_{L}|0\rangle = |\nu + \rangle + \mathcal{O}(m/E) |\nu - \rangle$$

- In oscillations, once the O(m/E) terms can be neglected:
 - the helicity does not play a role
 - there is no L-violation
 - oscillation formulae are identical for Dirac and Majorana v's

 $0\nu 2\beta$ decay



Violates lepton number Needs Majorana neutrinos

0v2β decay

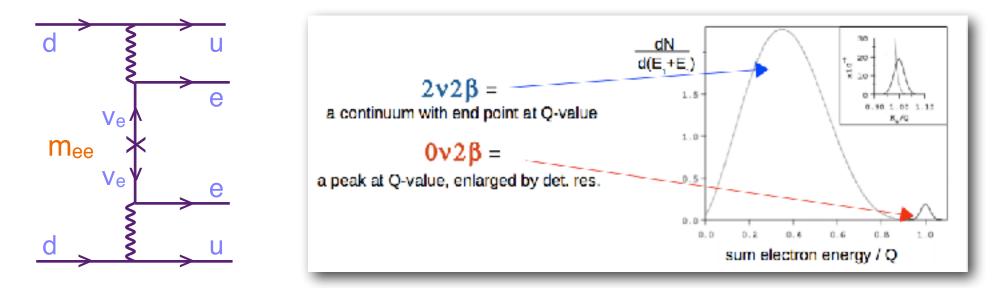
$$(A, Z) \rightarrow (A, Z + 2) + 2e^{-}; \quad \text{e.g.:} \ ^{76}\text{Ge} \rightarrow ^{76}\text{Se} + 2e^{-}$$

 $\Gamma \propto |m_{ee}|^2 \langle Q \rangle^2$

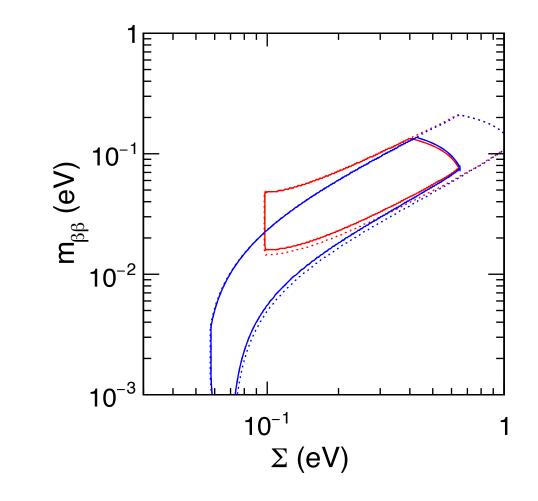
Depends on

- Phases
- Nuclear matrix elements
- Dírac vs Majorana

 $m_{ee} = U_{eh}^2 m_h = c_{13}^2 (m_1 c_{12}^2 + m_2 s_{12}^2 e^{2i\alpha}) + m_3 s_{13}^2 e^{2i\beta'}$



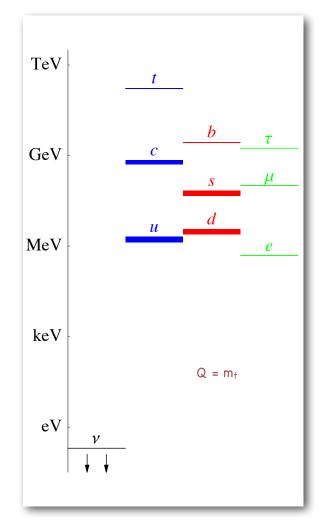
 $|m_{ee}| < \mathcal{O}(1) \times 0.2 \,\mathrm{eV} \,(\mathrm{Heidelberg-Moscow}) \to \mathcal{O}(1) \times 0.01 \,\mathrm{eV} \,(\mathrm{Genius})$



Neutrino masses are small

Whatrielse doawe (knibuig?s)

- Natural scale of fermion masses: $\langle H \rangle \approx 174 \text{ GeV}$
- Why $m_v / \langle H \rangle < 10^{-12}$?
- Must have a different origin than $m_e / \langle H \rangle \approx 0.3 \times 10^{-5}$
 - quantitatively larger hierarchy
 - family independent
 - compelling understanding available



Neutrino (lepton) flavour parameters

• Quark sector

$$m_{ij}^{D}\overline{d_{iR}}d_{jL} + m_{ij}^{U}\overline{u_{iR}}u_{jL} = \begin{pmatrix} m_{d_i}\overline{d'_{iR}}d'_{iL} + m_{u_i}\overline{u'_{iR}}u'_{iL} \\ j_{c,had}^{\mu} = \overline{u}_{iL}\gamma^{\mu}d_{iL} = V_{ij}\overline{u'_{iL}}\gamma^{\mu}d'_{jL} \end{pmatrix}$$

• Lepton sector, including neutrino masses (Majorana for definitess)

$$\frac{m_{ij}^{\nu}}{2}\nu_{iL}\nu_{jL} + m_{ij}^{E}\overline{e_{iR}}e_{jL} = \frac{m_{\nu_{i}}}{2}\nu_{iL}^{\prime}\nu_{iL}^{\prime} + m_{e_{i}}\overline{e_{iR}^{\prime}}e_{iL}^{\prime}$$
$$j_{\mathrm{c,lep}}^{\mu} = \overline{\nu}_{iL}\gamma^{\mu}e_{iL} = U_{ij}^{\dagger}\overline{\nu_{iL}^{\prime}}\gamma^{\mu}e_{jL}^{\prime}$$

Physical flavour parameters in the lepton sector

$$\frac{m_{\nu_{i}}}{2}\nu_{iL}\nu_{iL} + m_{e_{i}}\overline{e_{iR}}e_{iL} \qquad j_{c,lep}^{\mu\dagger} = U_{ij}\overline{e}_{iL}\gamma^{\mu}\nu_{jL}$$

$$U = \begin{pmatrix} e^{i\gamma_{1}} \\ e^{i\gamma_{2}} \\ e^{i\gamma_{3}} \end{pmatrix} \begin{pmatrix} \text{standard par.} \end{pmatrix} \begin{pmatrix} 1 \\ e^{i\alpha} \\ e^{i\beta} \end{pmatrix}$$

$$\underbrace{(\text{unphysical})}_{physical} \begin{pmatrix} M_{ajorana} \end{pmatrix} \begin{pmatrix} M_{ajorana} \end{pmatrix} \begin{pmatrix} 1 \\ e^{i\alpha} \end{pmatrix} \begin{pmatrix} M_{ajorana} \end{pmatrix} \begin{pmatrix} 1 \\ e^{i\beta} \end{pmatrix} \begin{pmatrix} M_{ajorana} \end{pmatrix} \begin{pmatrix} M_{ajorana} \end{pmatrix} \begin{pmatrix} 1 \\ e^{i\beta} \end{pmatrix} \begin{pmatrix} M_{ajorana} \end{pmatrix} \begin{pmatrix} M_{ajoranaa} \end{pmatrix} \begin{pmatrix} M_{ajoran$$

Physical flavour parameters in the lepton sector

$$\frac{m_{\nu_i}}{2}\nu_{iL}\nu_{iL} + m_{e_i}\overline{e_{iR}}e_{iL} \qquad j_{c,lep}^{\mu\dagger} = U_{ij}\overline{e}_{iL}\gamma^{\mu}\nu_{jL}$$

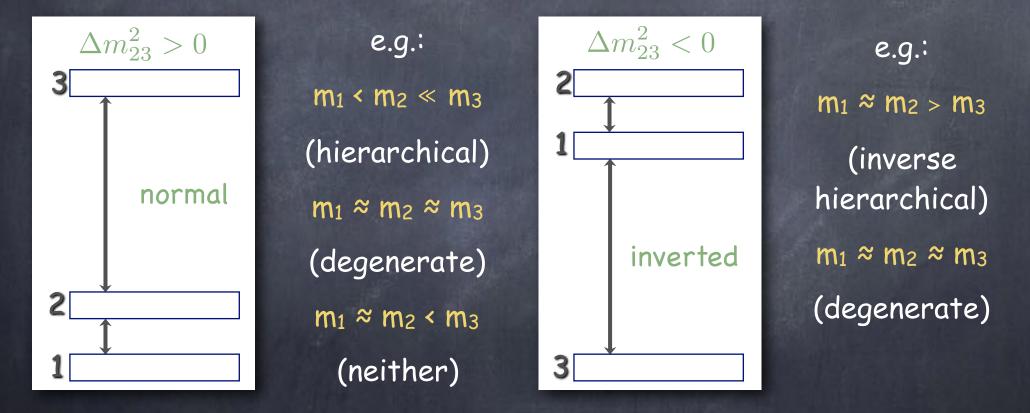
$$m_e, m_\mu, m_\tau, m_{\nu_1}, m_{\nu_2}, m_{\nu_3}, \theta_{23}, \theta_{12}, \theta_{13}, \delta, \alpha, \beta$$

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\sigma} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}s^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}s^{i\delta} & c_{23}c_{13} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha} & 0 \\ 0 & 0 & e^{i\beta} \end{pmatrix}$$

$$0 \le \theta_{23}, \theta_{12}, \theta_{13} \le \frac{\pi}{2}, \quad 0 \le \delta < 2\pi, \quad 0 \le \alpha, \beta < \pi$$

Standard labeling of eigenstates

 $egin{aligned} 0 < \Delta m^2_{12} < |\Delta m^2_{23}| & ext{uniquely defines the labeling} \ \Delta m^2_{12} > 0 & ext{by definition; } \Delta m^2_{23} ext{ can have both signs} & egin{pmatrix} \delta m^2 &\equiv \Delta m^2_{12} \ \Delta m^2 &pprox \Delta m^2_{23} &pprox \Delta m^2_{13} \end{pmatrix} \end{aligned}$





Accessible to oscillations

Charged sector

 $m_{e,\mu, au}$

 δm^2

 $|\Delta m^2|$

 $ext{sign}(\Delta m^2)$ $heta_{12}, heta_{23}, heta_{13}, \delta$ Not accessible to oscillations

 $m_{ ext{lightest}} \ lpha \ eta \ eta$

 $\begin{aligned} (\Delta m_{ij}^2 \equiv m_{\nu_j}^2 - m_{\nu_i}^2) \\ \Delta m_{13}^2 &= \Delta m_{12}^2 + \Delta m_{23}^2 \\ \delta m^2 \equiv \Delta m_{12}^2 \ll |\Delta m_{23}^2| \Rightarrow \Delta m_{23}^2 \approx \Delta m_{13}^2 \approx \Delta m^2 \equiv (\Delta m_{13}^2 + \Delta m_{12}^2)/2 \end{aligned}$

Accessible to oscillations

Not accessible to oscillations

 $m_{
m lightest}$

 α

 β

Charged sector

 $m_{e,\mu, au}$

 δm^2

 $\operatorname{sign}(\Delta m^2)$

 $|\Delta m^2|$

Well known

 $\theta_{12}, \theta_{23}, \theta_{13}, \delta$

Bounds

Hints

Known

$$\begin{split} |\Delta m^2| &\approx 2.4 \times 10^{-3} \,\mathrm{eV}^2 \quad \theta_{23} \sim 45^\circ \\ \delta m^2 &\approx 0.73 \times 10^{-4} \,\mathrm{eV}^2 \quad \theta_{12} \approx 33^\circ \\ \theta_{13} &\approx 8^\circ \end{split}$$

sign $(\Delta m^2) > 0$ preferred at 3σ $\delta \sim 3\pi/2 \pm 15\%$

 $|m_{ee}| \equiv |U_{ei}^2 m_{\nu_i}| < \mathcal{O}(1) \times 0.2 \,\text{eV}$ $(m^{\dagger} m)_{ee} = |U_{ei}|^2 m_{\nu_i}^2 < (2 \,\text{eV})^2$ $\sum m_{\nu_i} \lesssim 0.7 \,\text{eV} \text{ (priors)}$

double-beta decay

beta decay

cosmology

Guidelines for theory:

 $m_{
u_i} \ll 174 \,\mathrm{GeV}$ $heta_{23} \sim 45^\circ (= 45^\circ?)$ $heta_{12} \approx 33^\circ \neq 45^\circ$ $heta_{13} \text{ not so small}$ $\delta m^2 / |\Delta m^2| \approx 0.03 \ll 1$ normal ordering? large CP-violation?

Thank you romanino@sissa.it