

# Heavy Neutrinos and (Safe) Jet Vetoes <sup>1</sup>

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**elusives**  
neutrinos, dark matter & dark energy physics



**invisiblesPlus**

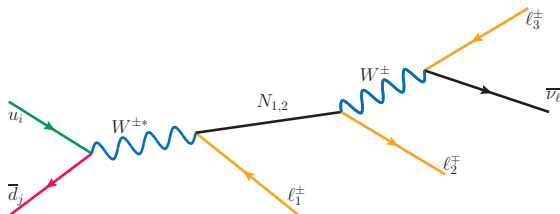
<sup>1</sup>with Silvia Pascoli and Cedric Weiland [1805.09335, 181X.YYYYYY]

<sup>2</sup>IPPP → CP3, Université Catholique de Louvain (Fall '18)

## a short history

2.5 years ago asked if possible to improve LHC searches for leptonic decays of heavy neutrinos,  $N \rightarrow \ell_1 W \rightarrow \ell_1 \ell_2 \nu$

- “improve”  $\neq$  MVA or BDT but a qualitatively new search strategy

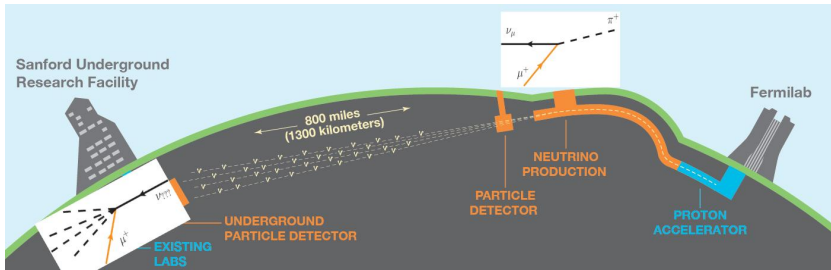


**Why?** new channels ( $W\gamma$  fusion), new technology (automated NLO+PS), unclear lepton number-violating  $\ell_1^{\pm} \ell_2^{\pm} + n_j$  was viable

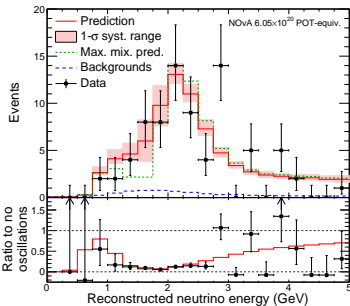
**An idea:** Global QCD activity in heavy  $N$  production different than backgrounds, e.g., fewer central, high- $p_T$  jets.

**The question:** Can jet observables be used to improve heavy  $N$  searches?

## Motivation for new physics from $\nu$ physics



In neutrino fixed-target expts,  $\nu_\mu$  beams from collimated  $\pi^\pm$ , then studied at near and far detectors (reminiscent of early SLAC DIS expts)



Deficit/disappearance of expected  $\nu_\mu$  (+appearance of  $\nu_e/\nu_\tau$ ) interpreted as  $\nu_{\ell_1} \rightarrow \nu_{\text{mass}} \rightarrow \nu_{\ell_2}$  transitions/oscillations [E.g. NO $\nu$ A  $\nu_\mu$  disapp., 1701.05891]

$\Rightarrow \nu$  have mass!



**So, neutrinos have masses  $\lesssim \mathcal{O}(0.1)$  eV.**

**Is this a problem?**

**Yes.**

# Neutrinos Masses and New Physics

To generate Dirac masses for  $\nu$  **like other** SM fermions, we need  $N_R$

$$\begin{aligned}\mathcal{L}_{\nu \text{ Yuk.}} &= -y_\nu \bar{L} \tilde{\Phi} N_R + H.c. = -y_\nu (\bar{\nu}_L \quad \bar{\ell}_L) \begin{pmatrix} \langle \Phi \rangle + h \\ 0 \end{pmatrix} N_R + H.c. \\ &= \underbrace{-y_\nu \langle \Phi \rangle}_{=m_D} \bar{\nu}_L N_R + H.c. + \dots\end{aligned}$$

However,  $N_R^i$  do not exist in the SM, implying  $m_D = 0$

## Significance of Neutrino Oscillations:

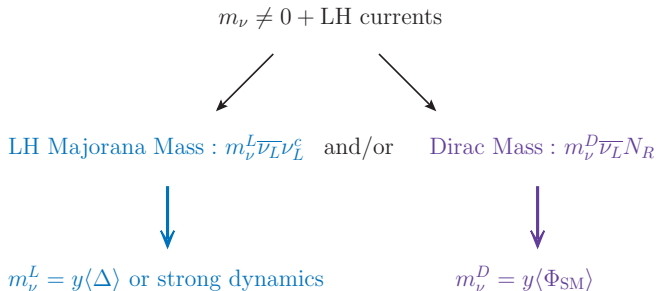
- Neutrino masses  $\implies \mathcal{L}_{\text{Universe}} \neq \mathcal{L}_{\text{SM}} (+\mathcal{L}_{\text{gravity}})$
- Instead,  $\mathcal{L}_{\text{Universe}} \approx \mathcal{L}_{\text{SM}} + \underbrace{\mathcal{L}_{\nu \text{ masses}}}_{\text{BSM physics!}} + \dots$

BSM physics! 

**Neutrino masses  $\implies$  existence of physics beyond the SM!**

# Neutrinos Masses and New Particles?

Nonzero neutrino masses implies new degrees of freedom exist [Ma'98]:



$m_\nu + \text{gauge inv.} + \text{renormalizability} \implies \text{new particles!}$

- No guarantee that new particles are  $N_R$ , e.g. Type II Seesaw
- New particles might be charged under new gauge symmetries, e.g.,  $(N_R, e_R)$  form  $SU(2)_R$  doublet or  $\Delta_L$  is scalar  $SU(2)_L$  triplet



# Collider Connection to Neutrino Mass Models<sup>3</sup>

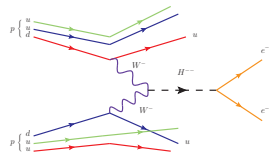
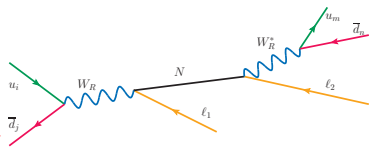
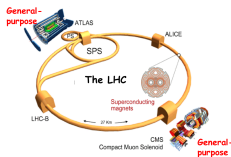
Neutrino mass models (aka Seesaw models) hypothesize new particles of all shapes, spins, charges, and color:

$N$  (Type I),  $T^{0,\pm}$  (Type III),  $Z_{B-L}$ ,  $H_R^{\pm,\pm\pm}$  (Type I+II), ...

**Produced** in  $ee/ep/pp$  collisions through gauge couplings and mixing:

**DY** :  $q\bar{q} \rightarrow \gamma^*/Z^* \rightarrow T^+T^-$  and  $qq' \rightarrow W_R^\pm \rightarrow N\ell^\pm$

**VBF** :  $W^\pm W^\pm \rightarrow H^{\pm\pm}$       **GF** :  $gg \rightarrow h^*/Z^* \rightarrow N\nu_\ell$



**Identification** of Seesaw particles through decays to SM particles

<sup>3</sup>Review on  $\nu$  mass models at colliders, Y. Cai, T. Li, T. Han, [RR \[1711.02180\]](https://arxiv.org/abs/1711.02180)

# Heavy Neutrinos and Colliders

# (Heavy) Neutrino Mixing for Non-experts (1/2)

After EWSB,  $\nu_L$  and  $N_R$  have same quantum numbers  $\implies$  mixing!

**Example:** In a two-state system, mixing between chiral eigenstates and mass eigenstates is given by unitary transformation/rotation

$$\underbrace{\begin{pmatrix} \nu_L \\ N_R^c \end{pmatrix}}_{\text{chiral basis}} = \underbrace{\begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}}_{\text{mass matrix}} \underbrace{\begin{pmatrix} \nu_1 \\ N_2 \end{pmatrix}}_{\text{mass basis}}$$

Decompose chiral/interaction states into mass states using:

$$|\nu_L\rangle = \cos \theta \underbrace{|\nu_1\rangle}_{\text{light}} + \sin \theta \underbrace{|N_2\rangle}_{\text{heavy}} \stackrel{\theta \ll 1}{\approx} \left(1 - \frac{1}{2}\theta^2\right) |\nu_1\rangle + \theta |N_2\rangle$$

## (Heavy) Neutrino Mixing for Non-experts (2/2)

Realistic neutrino mass models are much more complicated<sup>4</sup>. After diagonalizing mass matrix one gets two limits<sup>5</sup>:

- **High-scale seesaw:**  $\mu_M \gg \langle \Phi_{SM} \rangle \implies m_\nu \sim m_D \left( \frac{m_D}{\mu_M} \right), m_N \sim \mu_M$
- **Low-scale seesaw:**  $\mu_M \ll \langle \Phi_{SM} \rangle \implies m_\nu \sim \mu_M \left( \frac{m_D}{m_R} \right)^2, m_N \sim m_R$

For **discovery purposes**, no need to complicate life. Take agnostic/pheno. approach<sup>6</sup> with generic  $V_{\ell N}$  parametrization and one  $N$  mass eigenstate

- This requires modifying SM interactions with the following<sup>7</sup>:

$$\nu_{\ell L} = \sum_{m=1}^3 U_{\ell m} \nu_m + V_{\ell m'=4} N_{m'=4}$$

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<sup>4</sup>See for example, C. Weiland's thesis, [1311.5860]

<sup>5</sup>Kersten, Smirnov [0705.3221]; Moffat, Pascoli, Weiland [1712.07611]

<sup>6</sup>Appropriate since other Seesaws can mimic Type I pheno. See **RR**, [1703.04669]

<sup>7</sup>Atre, Han, Pascoli, Zhang [0901.3589]

# Heavy Neutrinos Couplings to EW Bosons

Consider left-handed (LH)  $SU(2)_L$  doublets (**gauge basis**):

$$L_{aL} = \begin{pmatrix} \nu_a \\ l_a \end{pmatrix}_L, \quad a = 1, 2, 3.$$

The SM  $W$  chiral coupling to **leptons** in **flavor basis** is given by

$$\mathcal{L}_{CC} = -\frac{g}{\sqrt{2}} W_\mu^+ \sum_{\ell=e}^{\tau} [\bar{\nu}_{\ell L} \gamma^\mu P_L \ell^-] + H.c.$$

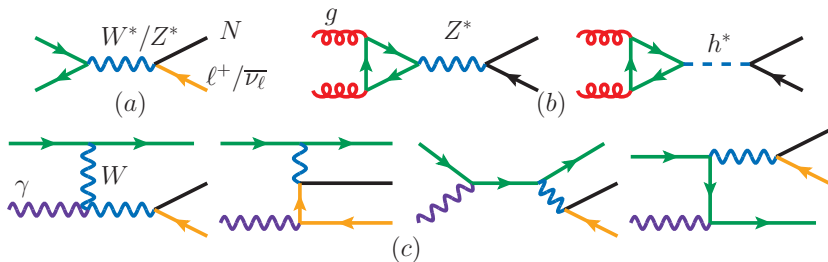
The SM  $W$  chiral coupling to **leptons** in the **mass basis**

$$\mathcal{L}_{CC} = -\frac{g}{\sqrt{2}} W_\mu^+ \sum_{\ell=e}^{\tau} \left[ \sum_{m=1}^3 \bar{\nu}_m U_{m\ell}^* + \bar{N}^c V_{N\ell}^* \right] \gamma^\mu P_L \ell^- + H.c.$$

$\implies N$  is **accessible through**  $W/Z/h$  currents

# Heavy Neutrino Production At Hadron Colliders

Heavy  $N$  can be produced through a variety of mechanisms in  $pp$  collisions

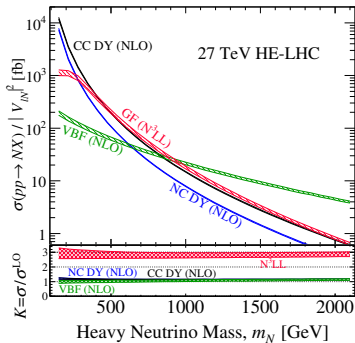
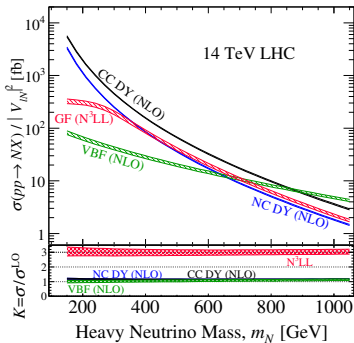


In fact, a resurgence of calculations in recent years<sup>8</sup>

- Clarity needed on (i)  $m_N, \sqrt{s}$  dependence and (ii) conflicting claims  
 $\implies$  more physical collider definitions + public tools [1602.06957]

<sup>8</sup>DY@NLO [\*1509.06375]; GF [1408.0983, \*1602.06957] @NNLL [\*1706.02298]; VBF [1308.2209, \*1411.7305, \*1602.06957]; DY,VBF Automation@NLO [\*1602.06957]; For extensive details, see review: [\*1711.02180]; (\*) = Pittsburgh and/or IPPP

# Across different colliders, wild interplay of PDF and matrix elements



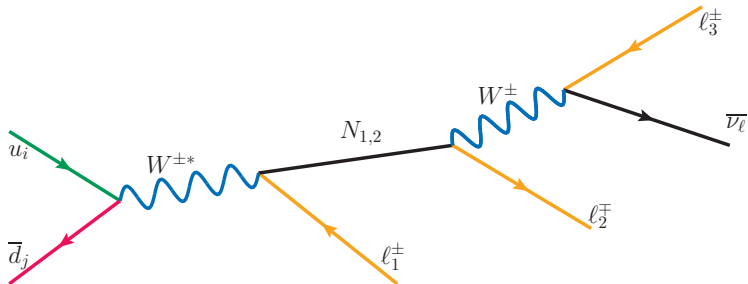
**Plotted:** Normalized production rate ( $\sigma/|V|^2$ ) vs heavy  $N$  mass ( $m_N$ )

- For  $\sqrt{s} \gtrsim 25 - 27$  TeV **GF** greater than **DY** due to **gg** luminosity
- For  $m_N \gtrsim 1 - 2$  TeV, **VBF** dominant due to large Yukawa couplings

**Note:** For  $m_N = 10$  TeV and  $|V_{eN}|^2 \sim 10^{-3}$ , then at 100 TeV, one has  $\mathcal{O}(30)$  VBF events after  $30 \text{ ab}^{-1}$ ! If  $\text{BR} \times \epsilon \times \mathcal{A} \sim \frac{1}{3}$ , then  $\sqrt{N_{\text{Obs.}}} > 3\sigma$

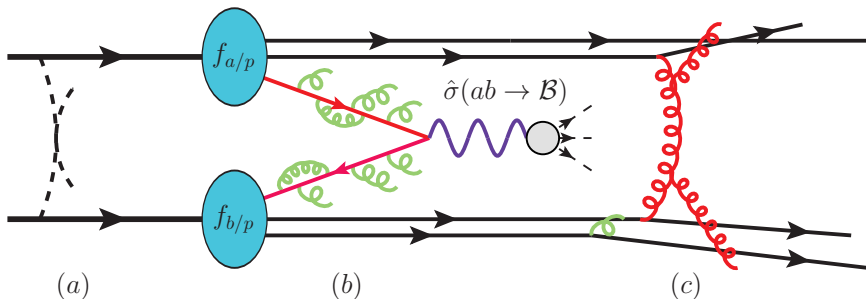
# Heavy Neutrinos and Jet Vetoes:

## The Difficulties of Jet Vetoes





# 1. Sources of Jets



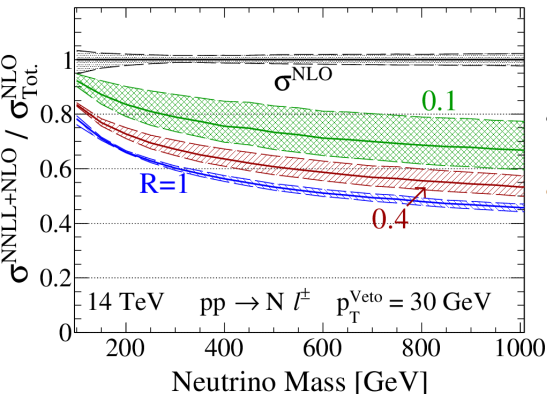
In  $pp$  collisions, hadronic activity has three sources:

- (a) Early, colorless, ultra-collinear/soft exchanges (**suppressed**)
- (b) Hard scattering (**calculable**)
- (c) Secondary/double/multi-parton scattering<sup>9</sup> (...)
  - ▶ Can contaminate jets from (b)

<sup>9</sup>Extremely rich physics. See, e.g., Collins, "Foundations of pQCD" (2011)

## 2. & 3. Poor signal Efficiencies and Uncertainties

For Drell-Yan and other color-singlet processes,  
more/harder QCD radiation (jets!) for systems with larger masses<sup>10</sup>



Prob. for signal to survive veto:

$$\varepsilon = \frac{\sigma^{\text{NLO+NNLL}}(p_T^j < p_T^{\text{Veto}})}{\sigma_{\text{Tot.}}^{\text{NLO}}}$$

$\varepsilon \rightarrow 0$  as  $(p_T^{\text{Veto}} / m_N) \rightarrow 0$

Large uncertainty despite high  
(NLO+NNLL!) precision  
( $R$  = jet radius parameter)

**Paradox:** Relax  $p_T^{\text{Veto}}$  for increasing  $m_N$  but top background jumps!

<sup>10</sup>Eg., Sleptons: F. Tackmann, et al [[1603.03052](#)];  $W'$ : B. Fuks, [RR](#) [[1701.05263](#)]

# Heavy Neutrinos and *Dynamical* Jet Vetoes

**A thought:**<sup>11</sup> *how about  $p_T$  of the leading charged lepton in the event?*

- For  $pp \rightarrow N\ell \rightarrow 3\ell X$ ,  $p_T^\ell \propto m_N \implies$  increases for larger  $m_N$

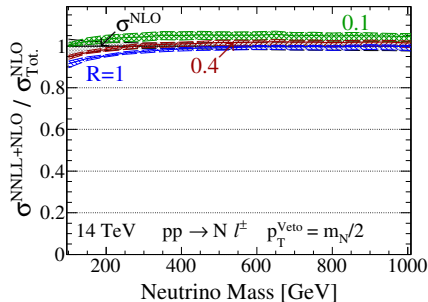
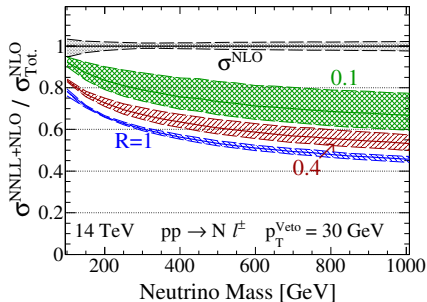
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<sup>11</sup>Disclosure: discovered basis of idea in an unrelated CMS paper on  $WW \rightarrow 0j$

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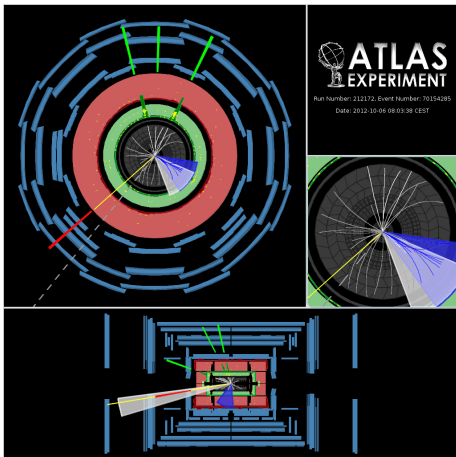
**QCD** uncertainties *shrink* since 2-scale problem converted into 1-scale

- less precise predictions, e.g., LL/parton shower, now more reliable

**Question:** What about backgrounds?

<sup>11</sup>Disclosure: discovered basis of idea in an unrelated CMS paper on  $WW \rightarrow 0j$

# Top Quark Background vs Dynamical Jet Vetoes



$pp \rightarrow t\bar{t}Z \rightarrow 1\mu + 3e + 2j_b + \cancel{E}_T$   
candidate event [1509.05276]

Classic kinematics:

- $m_{ee} = 93$  GeV
- $\cancel{E}_T = 57$  GeV

Typically,

- $p_T^{e1} \sim \frac{m_t}{4} \left(1 + \frac{M_W^2}{m_t^2}\right) \sim 50$  GeV
  - $p_T^{e3} \sim \frac{M_Z}{2} \sim 45$  GeV
  - $p_T^{b1} \sim \frac{m_t}{2} \left(1 - \frac{M_W^2}{m_t^2}\right) \sim 60$  GeV
- $p_T^{b1} > p_T^{\ell1} \implies$  event vetoed!

Setting  $p_T^{\text{Veto}}$  on event-by-event basis to  $p_T^{\ell1}$  can eliminate top quark background **without**  $b$ -jet tagging!

# Jet Vetoes and SM Backgrounds

Associated Top Quark Production:  $pp \rightarrow t\bar{t}l\bar{l}, t\bar{t}l\nu, tq\bar{l}l$

- Typical  $p_T$  of lepton from top:  $p_T^\ell \sim \frac{m_t}{4} \left(1 + \frac{M_W^2}{m_t^2}\right) \sim 50$  GeV
- Typical  $p_T$  of  $b$  from top:  $p_T^b \sim \frac{m_t}{2} \left(1 - \frac{M_W^2}{m_t^2}\right) \sim 65$  GeV
- $p_T^\ell < p_T^b \implies$  top events vetoed without need of  $b$ -tagging

# Jet Vetoes and SM Backgrounds

Associated Top Quark Production:  $pp \rightarrow t\bar{t}l\ell, t\bar{t}l\nu, tqll$

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Fake Leptons:

- Low- $p_T$  jet (in  $t\bar{t}$  events) identified as  $e^\pm$  or  $\tau^\pm$
- Low- $p_T$  charged  $\ell^\pm$  from weak decays of hadrons (in  $t\bar{t}$  events)
- Color conservation  $\implies$  second jet with comparable  $p_T$  likely exist

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Electroweak Production:  $pp \rightarrow 4l, 3l\nu, WWW, WWl\bar{l}$

- Jet veto  $\implies$  EW bosons at rest since no recoil
- Typical  $S_T \equiv \sum_\ell |\vec{p}_T^\ell|$  for  $3W$  or  $WZ$ :  $S_T \sim \frac{3M_W}{2} \sim 120 - 130$  GeV
- Typical  $S_T$  for heavy  $N$ :  $S_T \sim \frac{m_N}{3} + \frac{m_N}{2} + \frac{m_N}{4} = \frac{13m_N}{12}$



## Results<sup>12</sup>

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<sup>12</sup>with Silvia Pascoli and Cedric Weiland [1805.09335, 180X.YYYYY]

Jet vetoes are nonstandard selection cuts and make MC generation tricky

- **Need:** reliable description of *leading* jet at all  $p_T$  for signal (color-singlet) and background
- **Need:** “jets” (resummation/parton shower + jet definition)  
⇒ cannot apply veto at same time as other cuts
- **Need:** inclusive samples since bkg include  $\ell^\pm$  outside fid. volume

**Moto:** “We start at NLO”

- Event Generation: HeavyN@NLO UFO<sup>13</sup> + MadGraph5\_aMC@NLO
  - ▶ Bare-bones gen-level cuts on leptons + MadSpin for decay
- Shower: Pythia8.2 (w/ QED shower + recoil + Monash\* Tune)
- Particle-level Reco (lhe output)<sup>14</sup>: MadAnalysis5 +  $R = 1$  anti- $k_T$
- Smearing + offline analysis: private ROOT code

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<sup>13</sup>C. Degrande, O. Mattelaer, **RR**, J. Turner [[1602.06957](#)]

Available from FeynRules database: [feynrules.irmp.ucl.ac.be/wiki/HeavyN](http://feynrules.irmp.ucl.ac.be/wiki/HeavyN)

<sup>14</sup>See  $W'$ +jet veto analysis, Fuks, **RR** [[1701.05263](#)]

# Flavor Hypothesis and Signal Definition

As a benchmark flavor mixing scenario we set:

$$|V_{e4}| = |V_{\tau 4}| \neq 0 \quad \text{and} \quad |V_{\mu 4}| = 0$$

Two complementary<sup>15</sup> signal processes ( $\ell_X = e, \mu, \tau_h$ ):

**Signal I:**  $pp \rightarrow \tau_h^+ \tau_h^- \ell_X + \text{MET}$     and    **Signal II:**  $pp \rightarrow \tau_h^\pm e^\mp \ell_X + \text{MET}$

**Selection Cuts:** Standard ID requirements and  $m_{2\ell,3\ell}$  cuts

**Nonstandard Cuts:**

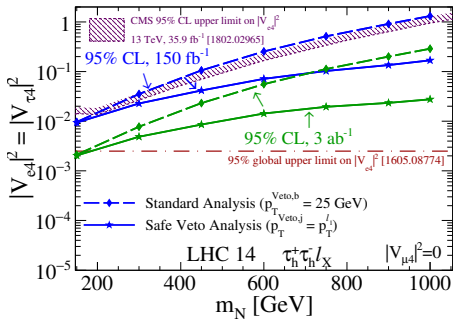
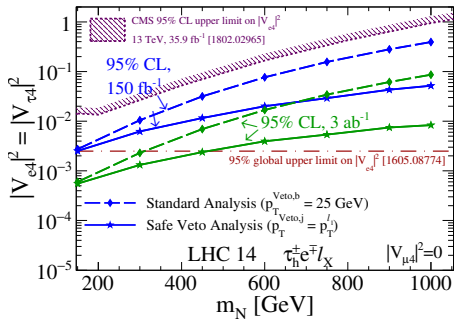
- Require  $p_T^{j_1} < p_T^{\ell_1}$  (jet veto) and  $S_T^\ell > 120$  GeV
- Given  $m_N$  hypothesis, cut on closest multi-body transverse mass  $\tilde{M}$

$$\begin{aligned} \tilde{M}_{T,i}^2 &= \left[ \sqrt{p_T^2(\ell^{\text{OS}}) + m_{\ell^{\text{OS}}}^2} + \sqrt{p_T^2(\ell_i^{\text{SS}}, \vec{p}_T) + M_W^2} \right]^2 \\ &\quad - \left[ \vec{p}_T(\ell^{\text{OS}}, \ell_i^{\text{SS}}) + \vec{p}_T \right]^2, \quad i = 1, 2. \end{aligned}$$

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<sup>15</sup>BR( $\tau/W \rightarrow eX$ ) are well-measured  $\implies$  can falsify no-LFV hypothesis if measured

# Results for 14 TeV LHC: $e\tau$ Scenario



**Plotted:** LHC 14 sensitivity to active-sterile neutrino mixing (coupling) vs heavy neutrino mass

- Dash = standard search<sup>16</sup> with  $b$ -jet veto (13 TeV CMS for  $e/\mu$ )
- Solid = “improved” analysis with special type of jet veto

Improved sensitivity up to  $10 - 11 \times$  with  $\mathcal{L} = 3 \text{ ab}^{-1}$ .

<sup>16</sup>More aggressive cuts on charged leptons: e.g.,  $p_T^{\ell_1} > 55 \text{ GeV}$ ,  $m_{3\ell} \geq 80 \text{ GeV}$

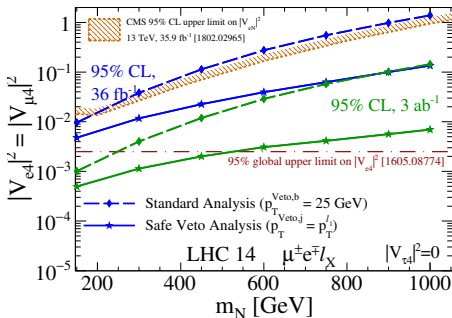
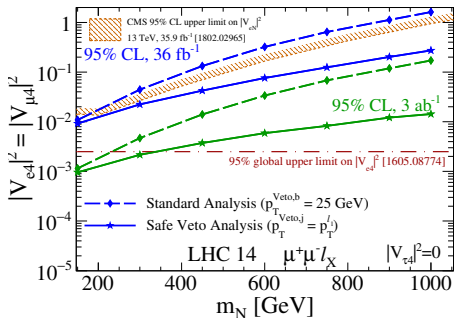
# More at 14 TeV LHC: $e\mu$ Scenario

Benchmark flavor mixing scenario II:

$$|V_{e4}| = |V_{\mu 4}| \neq 0 \quad \text{and} \quad |V_{\tau 4}| = 0$$

Two complementary signal processes ( $\ell_X = e, \mu, \tau_h$ ):

**Signal I:**  $pp \rightarrow \mu^+ \mu^- \ell_X + \text{MET}$     and    **Signal II:**  $pp \rightarrow \mu^\pm e^\mp \ell_X + \text{MET}$



Again, improved sensitivity  $> 10\times$  with  $\mathcal{L} = 3 \text{ ab}^{-1}$ .

# Summary

Heavy neutrinos remain one of the best (but not the only!) explanations for tiny neutrino masses

**Idea:** *We have developed a new approach to searches for heavy  $N$  at  $pp$  colliders, one based on an unusual jet veto scheme ( $p_T^{\text{Veto}} = p_T^{\ell_1}$ )*

- New veto scheme reveals  $> 90 - 95\%$  signal acceptance with little-to-no dependence on  $m_N$  (contrary to previous veto schemes)
- Substantial reduction in QCD theory uncertainty at NLO+NNLL(Veto)  $\implies$  less need for high-precision resummation
- Redesigned search analysis with better reduction of background  $\implies$  Improved LHC sensitivity by up to  $10\times$  over LHC's lifetime

**Remember:** *"The LHC is planned to run over the next 20 years, with several stops scheduled for upgrades and maintenance work" [press.cern]*

- High-Luminosity LHC and Belle II goals:  $3\text{-}5 \text{ ab}^{-1}$  and  $50 \text{ ab}^{-1}$
- Premature to claim "nightmare scenario" (SM Higgs + nothing else)

The logo for IPP3 is centered on the slide. It consists of the letters 'IP' in a large, blue, serif font, followed by a smaller blue '3'. The text is enclosed within a light blue oval border. A decorative wavy line extends horizontally from the left and right sides of the oval.

**Thank you.**