Exciting the Vacuum

aka Central Exclusive Production

from the ISR (Glueballs) ... to the Tevatron (leptons, photons, hadrons, charmonium)... to the LHC (Higgs, WW, Z?, SUSY?)

 $pp \rightarrow p$ X p where X is a simple system completely measured

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Part 1 :
The vacuum: an experimental viewpoint (hit it!)

Different hammers: electromagnetic, strong and weak Different exposure times: shorter → Planck longer → Age of Universe!

Experiments (observing) with electromagnetic eyes

Observing with strong eyes (looking for glueballs)

Combining the above (photoproduction)

Now at the LHC: (weak eyes) exciting the Higgs field

Part 2 : Central Exclusive π + π - production in CDF (new)

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GAP-X-GAP Physics or Central Exclusive Production

Mike Albrow, Fermilab

>> What is GXG and why is it interesting?

>> X = low mass hadrons (double pomeron exchange, D P E) at ISR (AFS), SPS (Ω), Tevatron (CDF), LHC

>> Other GXG channels: $\gamma + \gamma$, $\gamma + P$ in CDF and CMS & P+P $\rightarrow \chi_c$

>> Future, high M(X) at LHC: X = H, JJ, W+W-

CMS: 2 candidates

Review:

Central Exclusive Particle Production at High Energy Hadron Colliders. M.G. Albrow (Fermilab), T.D. Coughlin (University Coll. London), J.R. Forshaw (Manchester U.). Jun 2010. 64 pp. Published in Prog.Part.Nucl.Phys. 65 (2010) 149-184

arXiv:1006.1289 [hep-ph]

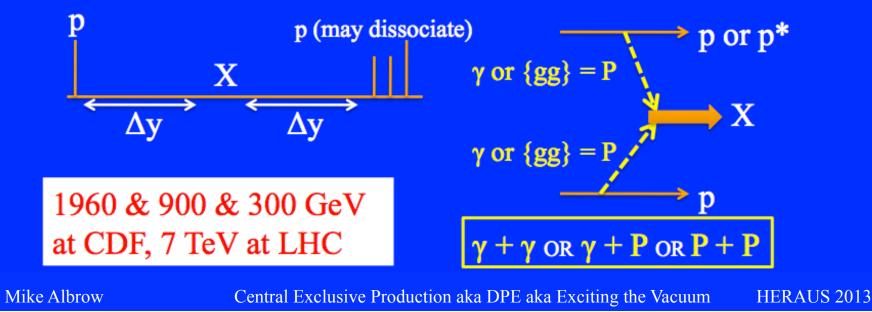
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What is GXG and why is it interesting?

 $G = GAP = Rapidity gap: \Delta y > ~ 4$ with NO HADRONS

X = central particles, fully measured (exclusive, distinct from inclusive) Examples of X, all observed for 1st time in hadron+hadron by CDF: $X = e^+e^-, \mu^+\mu^-, \gamma + \gamma, J/\psi, \psi(2S), \chi_c, Jet+Jet$ New X = 2 – 4 hadrons ($\pi\pi$, & to come: KK, $\rho\rho$, $\phi\phi$, $\Lambda pK(?)$ etc.) Later in CDF: X = high n(ch), ΣE_T , gluon jets (leaders?), charm, etc. Later at LHC: X = more Jet+Jet, W⁺W⁻, Higgs, ...



Three classes of GXG:

2-photon (QED + small QCD corrections ... shadowing or P) Studied in e+e- at LEP, a little in ep at HERA, now also in pp. Photoproduction γ +P at HERA, now also in pp. Double pomeron exchange, only in hadron-hadron (pp/ppbar) Comparing M(X) spectra in different processes helps meson spectroscopy

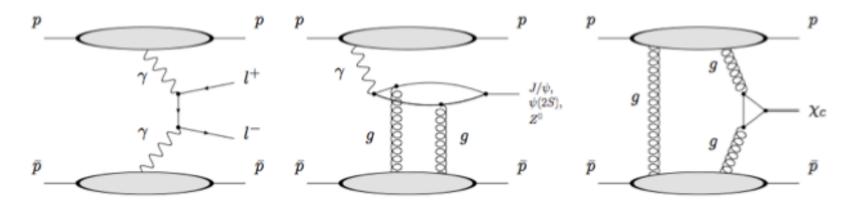


Figure 10: Feynman diagrams for processes contributing to the exclusive di-lepton signal. (a) $\gamma\gamma \rightarrow l^+l^-$, (b) $\gamma I\!\!P \rightarrow J/\psi, \psi(2S), Z^0$, and (c) $I\!\!P I\!\!P \rightarrow \chi_{c0}$.

At low masses, Double Pomeron region, non-perturbative and difficult to calculate. New efforts (Harland-Lang et al, Antoni Szczurek et al.)

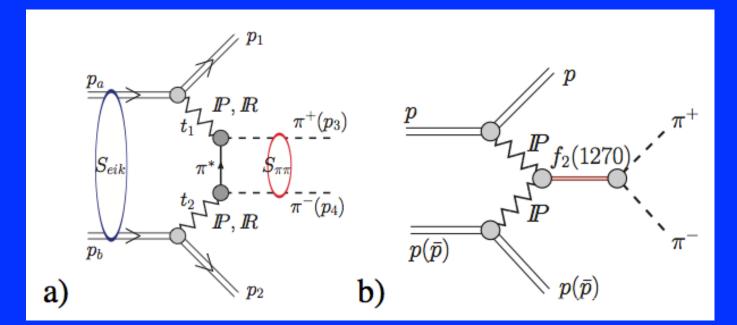
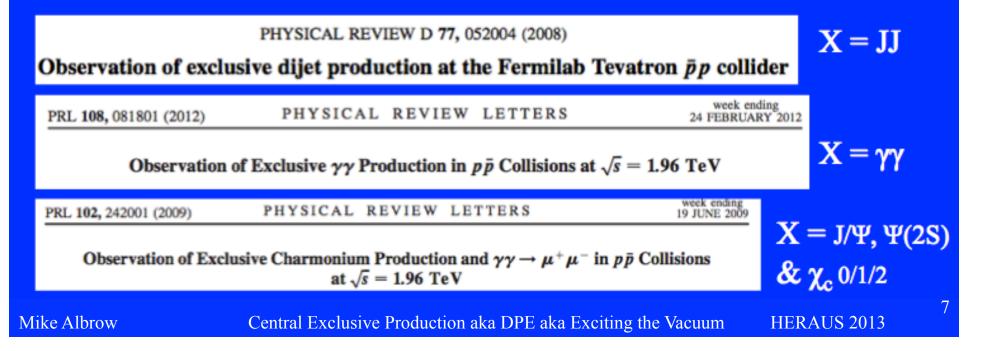


Diagram from P.Lebiedowicz & A.Szurek or $f0(600)/\sigma$ or f0(980) etc. arXiv:12120166

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P + **P**

P = pomeron = strongly interacting color singlet t-channel exchange, in leading order {gg}, at higher Q² gluons and q-qbar pairs evolve in. P + P → X has selection rules: Q = S = C = B = 0 (of course), I^G = 0⁺, J^{PC} = 0⁺⁺ or 2⁺⁺ DOMINANT Interesting for hadron spectroscopy, and for understanding pomeron (> 25% of p-pbar collisions have pomeron exchange!) Higgs boson obeys all selection rules (vacuum) GHG expected at LHC (~2 - 5 fb y) We already did some, and they importantly tested the theory for p + H + p:



ISR: 1st to allow $\Delta y = 3 + 2 + 3$ Low mass (<~ 2 GeV) DPE cross sections at ISR ($\sqrt{s} \le 63$ GeV).

If central state (~ $\pi\pi$) restricted to fixed central y-bin, σ falls with \sqrt{s} . At lower \sqrt{s} Regge exchanges, X includes ρ (I=1) etc., which die out at high \sqrt{s} leaving P exchange.

If Gap widths Δy are fixed at 3, σ rises (more room for X, higher M(X) allowed.

At top ISR energy σ (DPE) ~ 15 µb

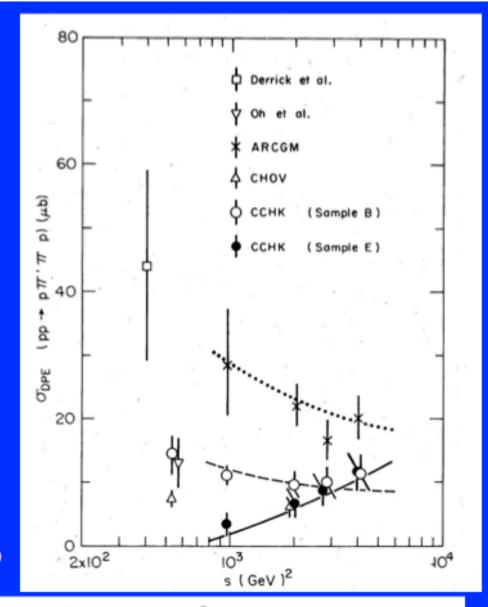


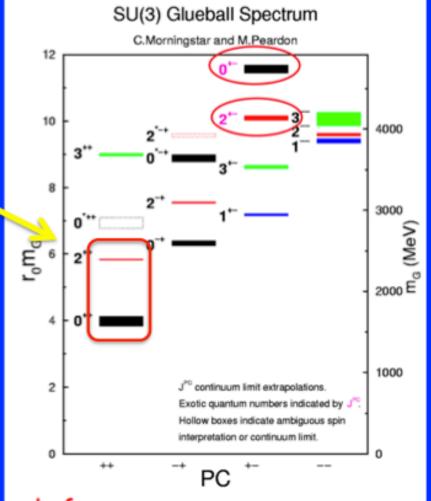
Figure 7: Experimental DIPE cross sections (μ b) versus s (GeV²) at the ISR together with the Regge calculations of Ref [26]. The full circles and the rising solid line are for two gaps with $\Delta y > 3$. The dashed line is for $|y_{\pi}| < 1.0$ and the dotted line for $|y_{\pi}| < 1.5$. Figure from Ref. [26].

Double Pomeron Exchange favours Glueballs or Glue-rich hadrons

Scalars especially are poorly understood: big hole in understanding QCD. Note: no narrow/light glueball: $\Gamma_G > \text{few 100's MeV}$, \rightarrow not an isolated particle. H(125) is ! $\Gamma_H \sim 5 \text{ MeV} \ll M(\pi)$

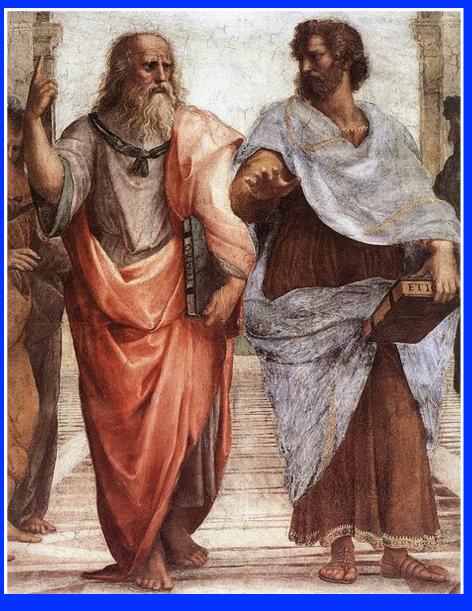
Representative lattice QCD calculation: Note especially lightest expected states 0++ (~1600) and 2++ (~2300) These Q.Nos. could not be hybrids (?)

Need knowledge: >> Different decay modes (B.R.'s) >> Widths >> Production mechanisms, e.g. P+P favours g-g while γ+γ favours q-qbar



"Nature abhors a vacuum"

Oh, really?



The vacuum \supset all of physics

Plato & Aristotle

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An early experiment on the vacuum: (Otto von Guericke, 1650)





"The vacuum sucks!"

Common wisdom, but it's those tiny atoms pushing in from the outside, and even teams of horses could not beat those tiny atoms! 16 horses against lots of atoms! There's strength in numbers.

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Atoms are much emptier than most people imagine





Even:



Cold fusion anyone?

Better:



Really more like a pinhead in a Welsh house

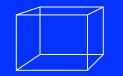


Sorry, • should be much smaller

 10^{-13} cm / 10^{-8} cm ~ 0.2 mm / 20m

How can an experimenter study the vacuum? HIT IT! (How else?)

Box of vacuum:

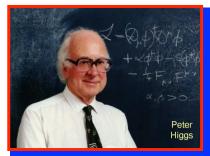


In copper: 5.10^{-15} ... of course its <u>*full*</u> of EM fields.

Matter (nucleons) take 10⁻¹⁸ of volume! (in air)

In pursuit of the perfect vacuum (recipe) :

- 1) Remove all the nucleons and electrons
- 2) Cool down to so close to T = 0 that thermal μ waves don't fit.
- 3) Thick shield (no radioactivity) to kill X-rays and γ -rays.
- 4) Choose a time window when no high energy cosmic rays traverse.
- 5) Damn the neutrinos! Bright to weak eyes, but invisible to EM eyes.
- 6) Pionic and gluonic fields on the walls, but they don't come in far.
- 7) Let the box fall freely in outer space : no gravitational fields.



Are we there yet?

Nice try, but no, for at least 2 reasons



1) Professor Higgs said that a "potential" (usually called a "vacuum expectation value") or scalar field is everywhere. So <u>that's</u> why an electron (or μ) here and in Andromeda have the same mass, and they know what mass to have even in $\gamma\gamma \rightarrow \mu+\mu$. If there's really "nothing", how could they possibly know?

How do we define the vacuum anyway, if not the absence of any stuff? Let's say: the lowest possible energy state $\Delta E.\Delta t > \hbar$

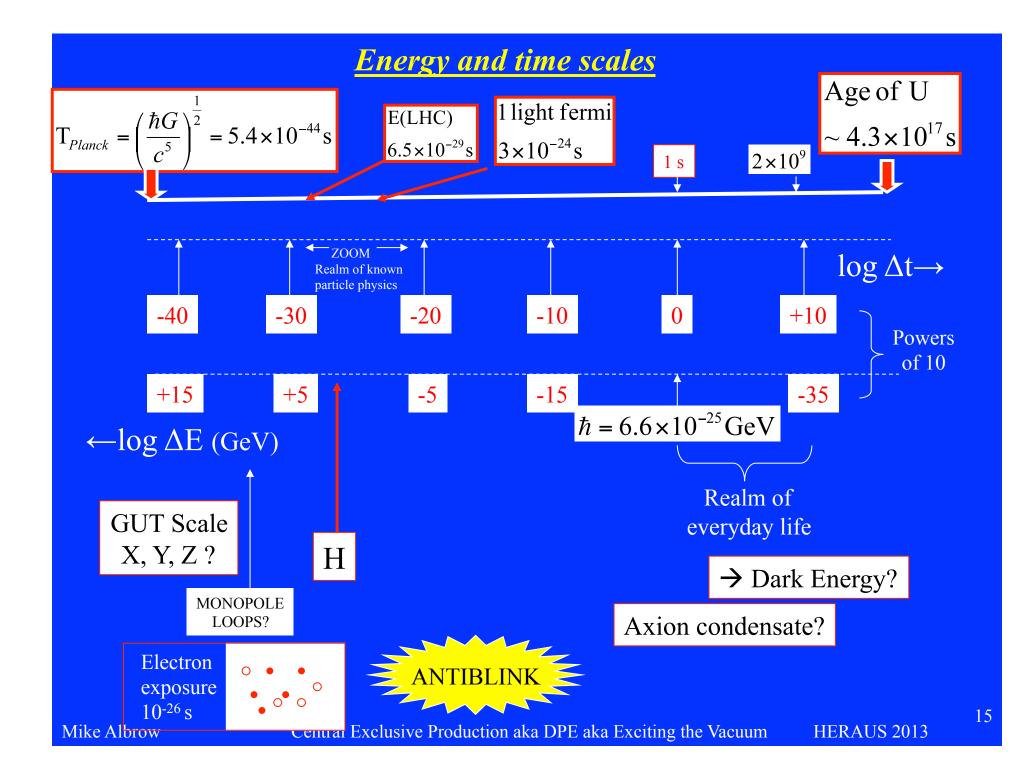
But Professor Heisenberg said that E = 0does not make sense unless you have an ∞ exposure time.

If you antiblink, ANTIBLINK short exposure, high energy fluctuations. Mike Albrow Central Exclusive Production aka DPE aka Exciting the Vacuum



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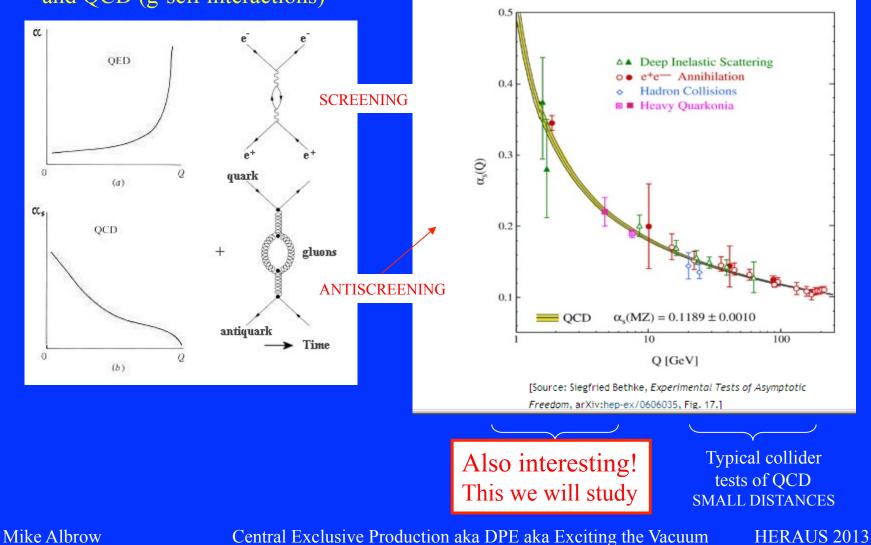
TIRLINI



Energy and time scales : ZOOM IN Realm of known 1 light fermi E(LHC) particle physics 3×10^{-24} s 6.5×10^{-29} s $\log t \rightarrow$ Powers -26s -24s -22s -28s -21s of 10 $\leftarrow \log E (GeV)$ 1 TeV 1 GeV 1 MeV $\mu^+\mu^$ qq $e^+e^$ tt **EM eyes** ←pQCD G →? **Strong eyes** $\sigma(600)$ $\chi_{b0}\,\chi_{c0}$ $q\overline{q}$ Н? tt H? Weak eyes tt

Particles shown have vacuum quantum numbers hadrons $\rightarrow I^{G}J^{PC} = 0^{+}0^{++}$ and can fluctuate up in time Δt (I^G = 0+ if appropriate ... hadrons) Mike Albrow Central Exclusive Production aka DPE aka Exciting the Vacuum HERAUS 2013

Vacuum polarization: opposite effect in QED and QCD (g-self interactions)



Asymptotic Freedom and the Coupling Constant of QCD

Rapidity gaps, with (almost) no Regge theory

The mother of all rapidity gaps Δy Elastic scattering:

$$\Delta y = 2. \ln \frac{\sqrt{s}}{m_{p/e}}$$

$$y_{beam} = \ln \frac{\sqrt{s}}{m_{p/e}}$$

$$y_{beam} = \ln \frac{\sqrt{s}}{m_{p/e}}$$

$$g_{beam} = \ln \frac{\sqrt{s}}{m_{p/e}}$$

Rap-gap cross sections go like: $\sigma(\Delta y) \sim e^{(\alpha(0)-1)\Delta y} \sim e^{(J-1)\Delta y}$

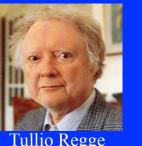
Thus: over large gaps exchange has $J \ge 1$, Q = 0, color singlet. Only 2 possibilities: Photon γ (in ee and pp and ep) Gluon (ep and pp) ... ?**color**!!? Cancel it with >= 1 other gluon. Call it pomeron (gg) $\alpha(0)=1+\varepsilon$ Pomeron has C = +1 (2 gluons OK). C = -1 OK with (ggg) : odderon ... (a) large |t| ? Z obeys the above rules, OK in e+e- but p inevitably break up.

= 18.0 LHC13 pp

Large (>~4) rapidity gaps only possible by (t) exchange of 4-momentum with: No color or charge, and effective spin at t ~ 0 >= 1. J = 1, α(0) >=1.
But (a) we have such large gaps in strong interactions
(b) QCD is THE theory of strong interactions. Unlike QED, there is no elementary (q,g) object with these properties.
(c) In QCD, with Regge theory to describe exchanges of states in the t-channel, only >= 2 g exchange can work.

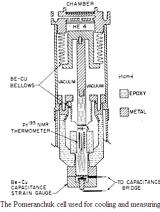
Scattering amplitudes must obey analyticity, crossing symmetry & unitarity. Allow complex angular momenta \rightarrow Regge theory

gg (C = +1) → Pomeranchukon → Pomeron IP ggg (C = -1) → Odderon O (not yet (?) detected. $\alpha < 1$?)



Pomeranchuk cooling cell (1950) → 1 mK

Are you a theorist?





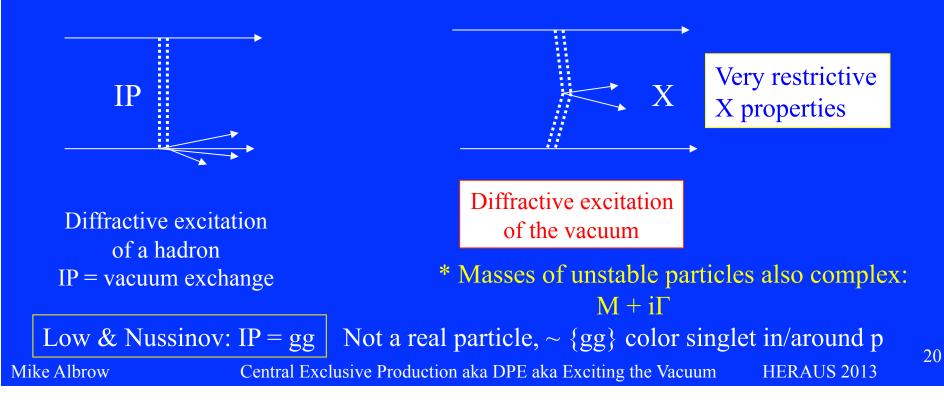
Isaak Pomeranchuk 1913 - 1966 HERAUS 2013

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Pre 1970: Strong interactions described by Regge Theory. Reggeons, pomerons, ...

Post ~1974: QCD rules! But hadrons @ low Q² on the outside looking in. Make spacetime a lattice (Wilson) : Lattice Gauge Theory "Pomeron" not invited to the party. ... "the P-word" Double Pomeron Exchange (DPE) double yuk! Yet: it's fundamental strong interaction physics, at large distances. New phenomena invited.



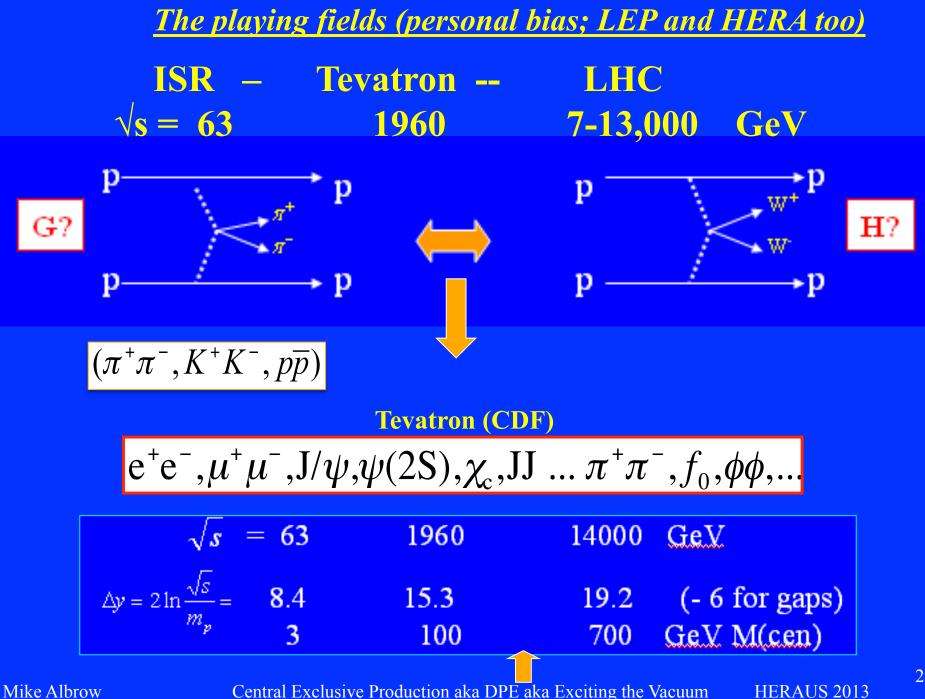
The **QCD** vacuum is a very complicated matter, made of strongly interacting quark and gluon fields, and one can hardly understand the hadrons without understanding the vacuum first. (Ed Shuryak)

The vacuum has phases revealed under extreme conditions:

Quark gluon plasma (or perfect liquid?) at high temperature **Color superconductor** at high density

→numerical simulations : Lattice Gauge Theory LGT

→Semiclassical methods : Fundamental topological solutions of Yang-Mills equations called instantons (these are not particles)



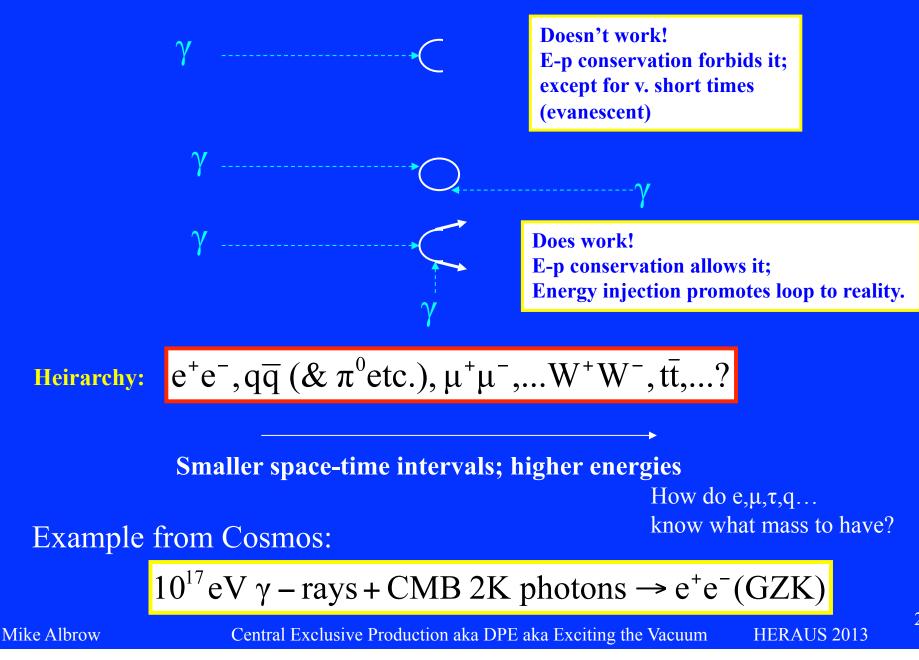
Axial Field Spectrometer (ISR)

CMS (LHC)

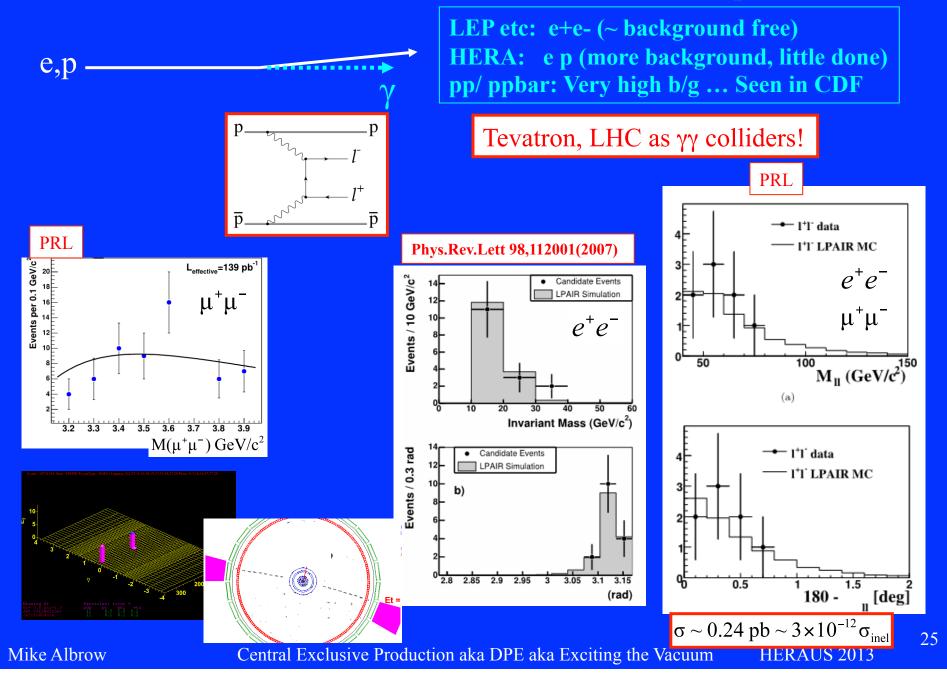


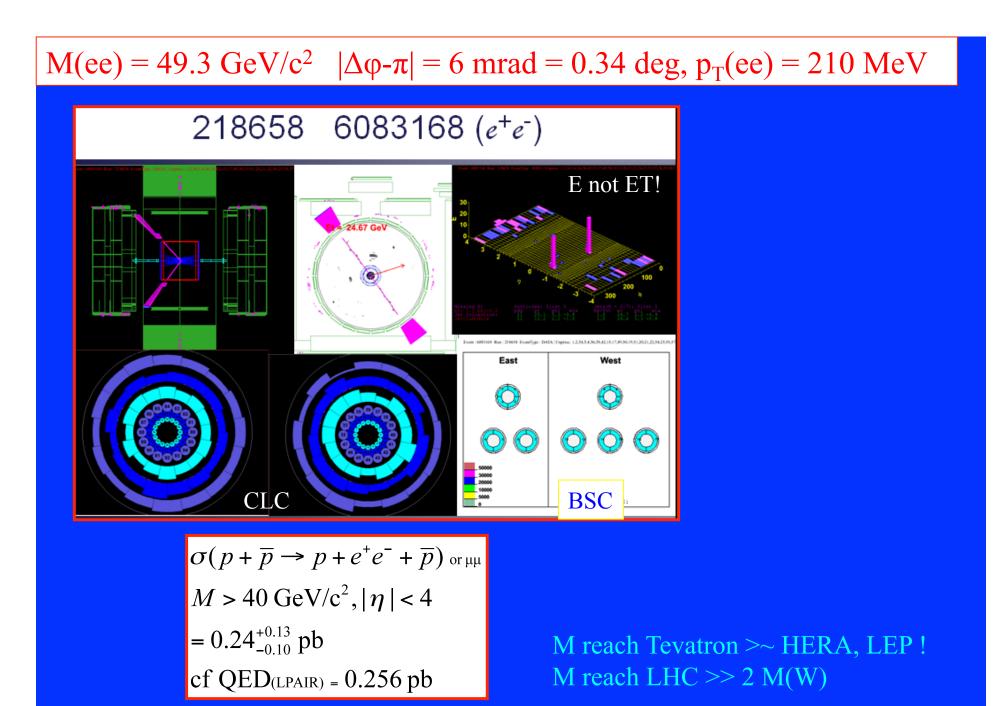
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Exciting the vacuum with photons



Photon "beams" radiated from electrons and protons





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All CDF measurements agree with QED: So what?

- 1) It shows we know how to select rare exclusive events in hadron-hadron environment
- 2) No other h-h cross section is so well known theoretically except Coulomb elastic. Possible Luminosity calibration at LHC (but Van der Meer v.good)
- Outgoing p-momenta extremely well-known (limited by beam spread). Calibrate forward proton spectrometers.
- 4) Practice for other $\gamma\gamma$ collisions at LHC:

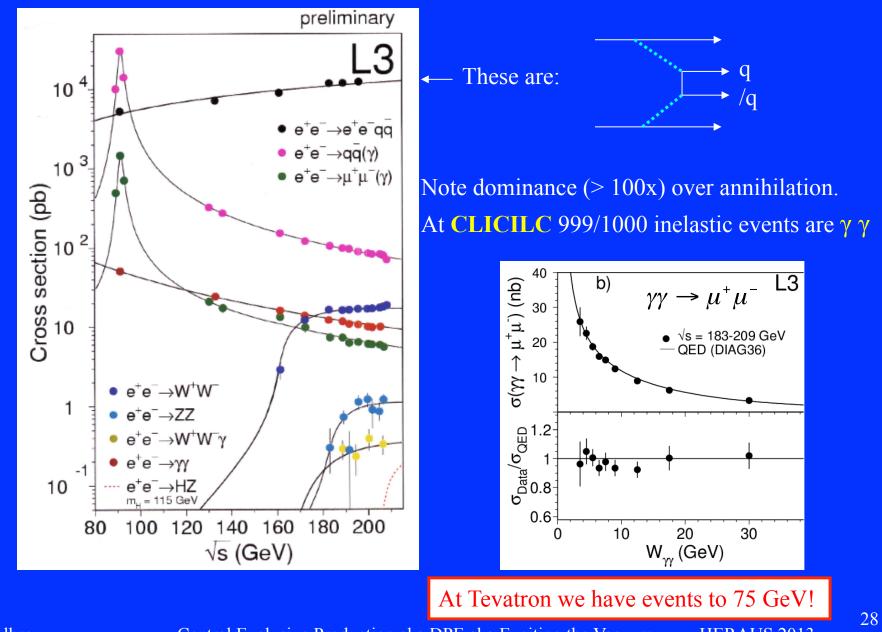
Note: These are present in the vacuum if you look quickly enough, and charged so they are visible to "light", rather EM, probes.

4400 events in 500 pb^{-1} with

 $\gamma\gamma \rightarrow W^+W^-, \widetilde{l}^+\widetilde{l}^-, \dots$

 $M(\mu^+\mu^-) > 10 \text{ GeV and } |\eta| < 2$

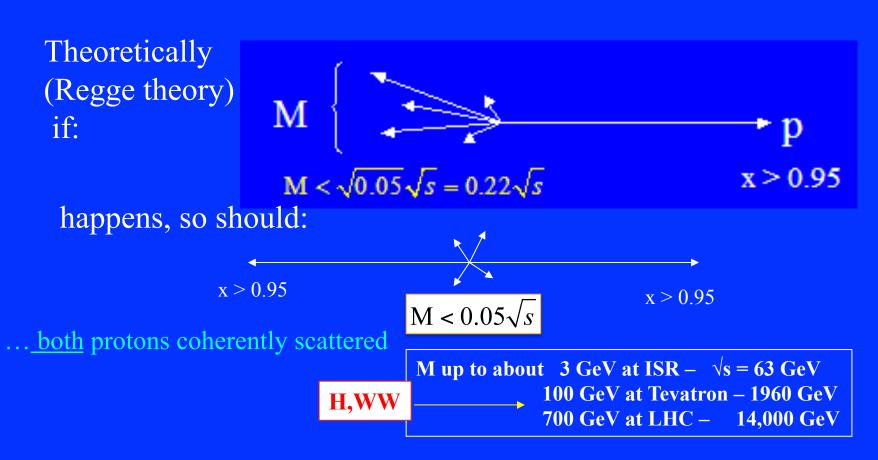
Exciting the vacuum with electromagnetic eyes



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Central Diffractive Excitation



Optical theorem: total cross section = *Im*(forward scattering amplitude)

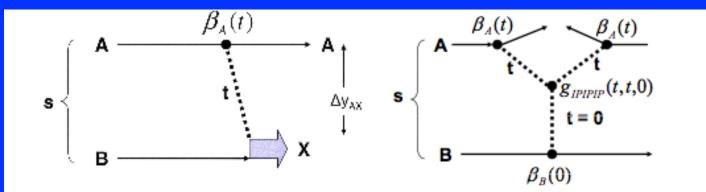


Figure 2: (a) Diffractive excitation of particle B to a state of mass M_X by pomeron exchange. (b) The corresponding cut diagram in the limit of large M_X .

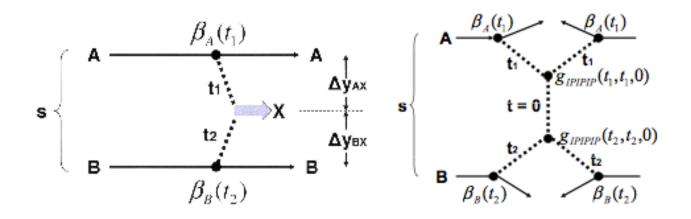


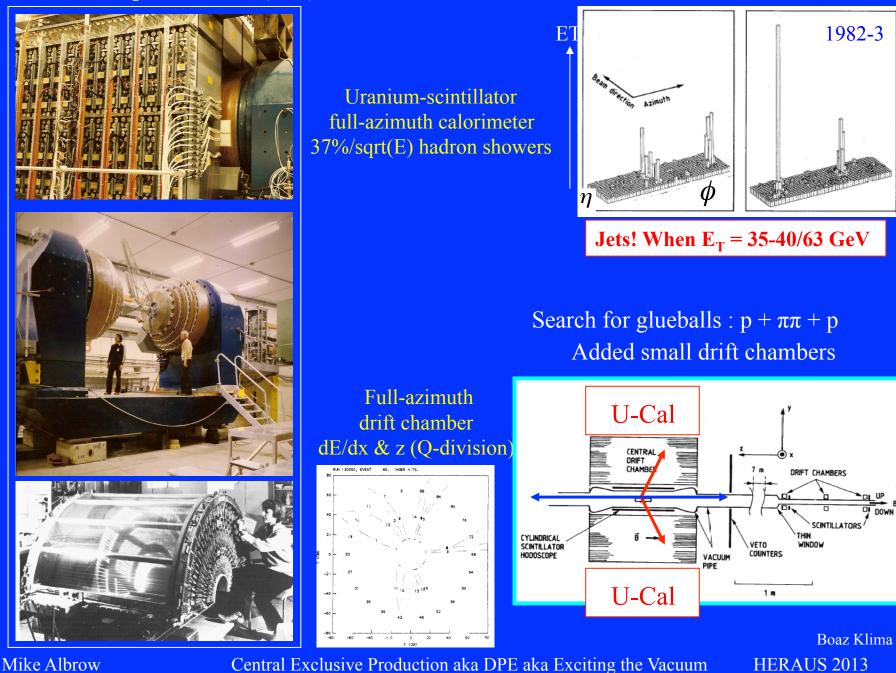
Figure 3: (a) Diagram for double pomeron exchange. (b) The corresponding cut diagram in the limit of large M_X .

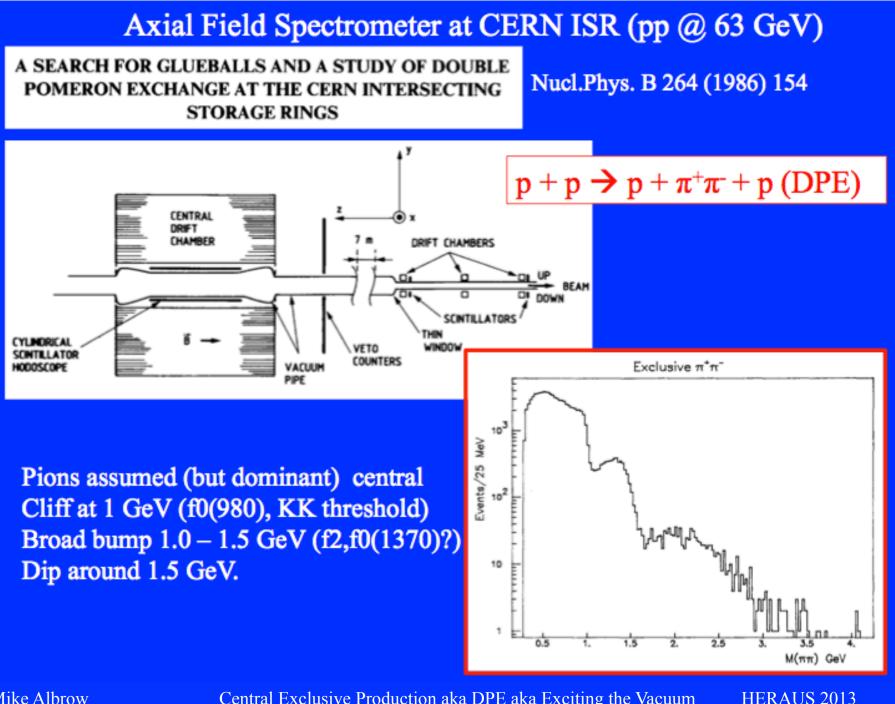
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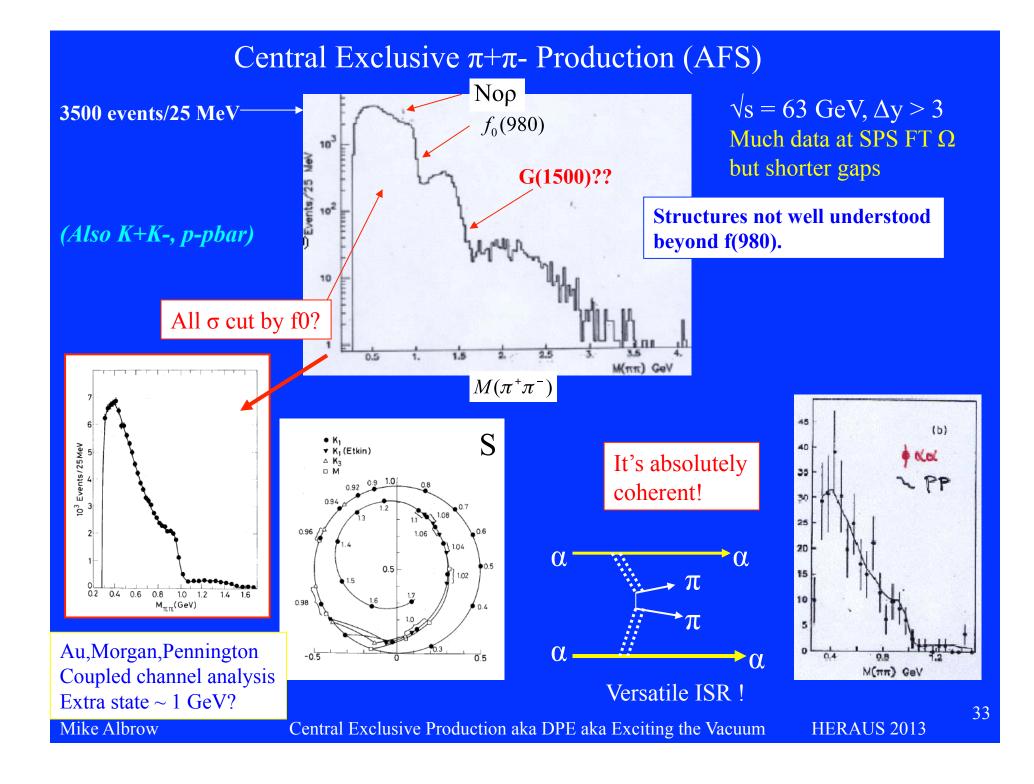
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Axial Field Spectrometer (ISR)

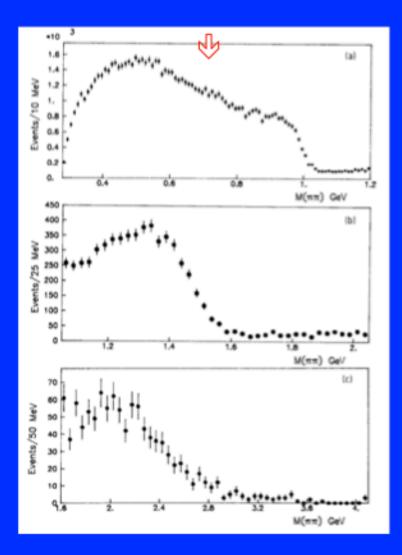




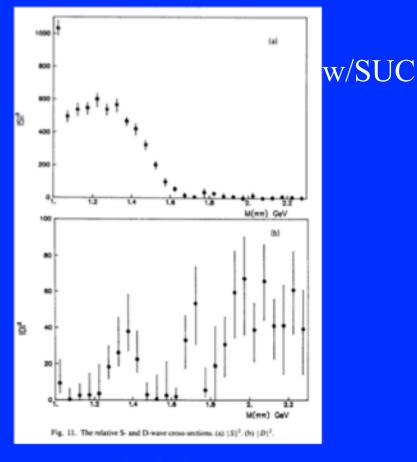
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Same AFS data, linear scale No $\rho(770)$ (I = 1)

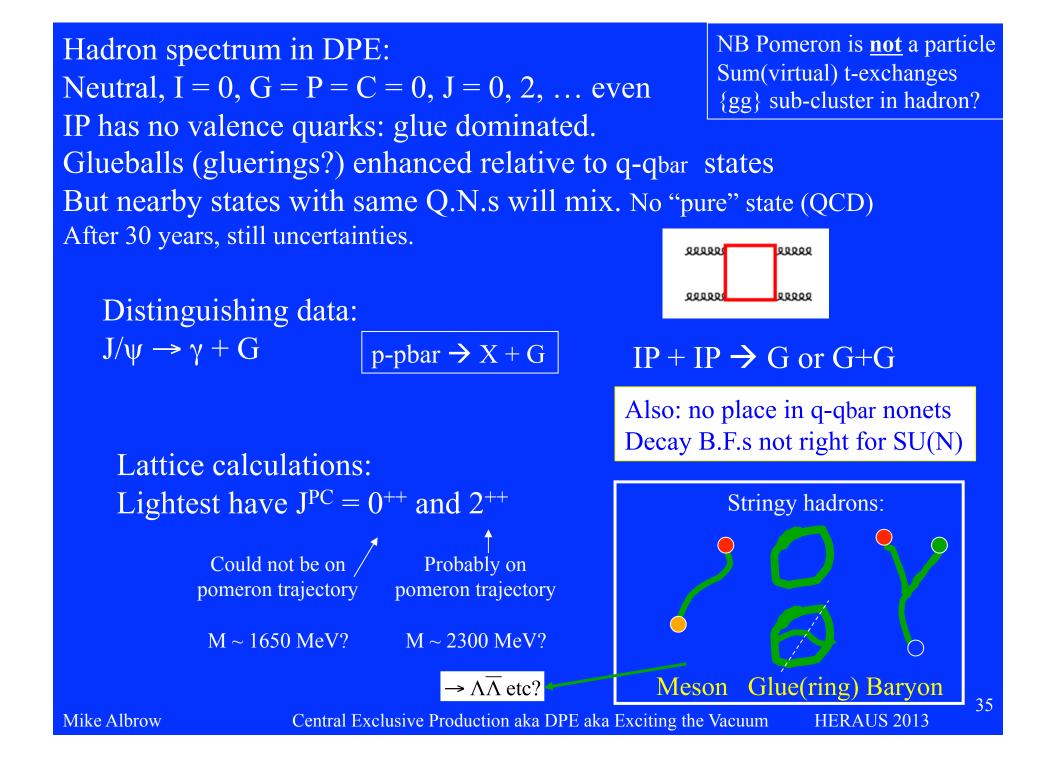


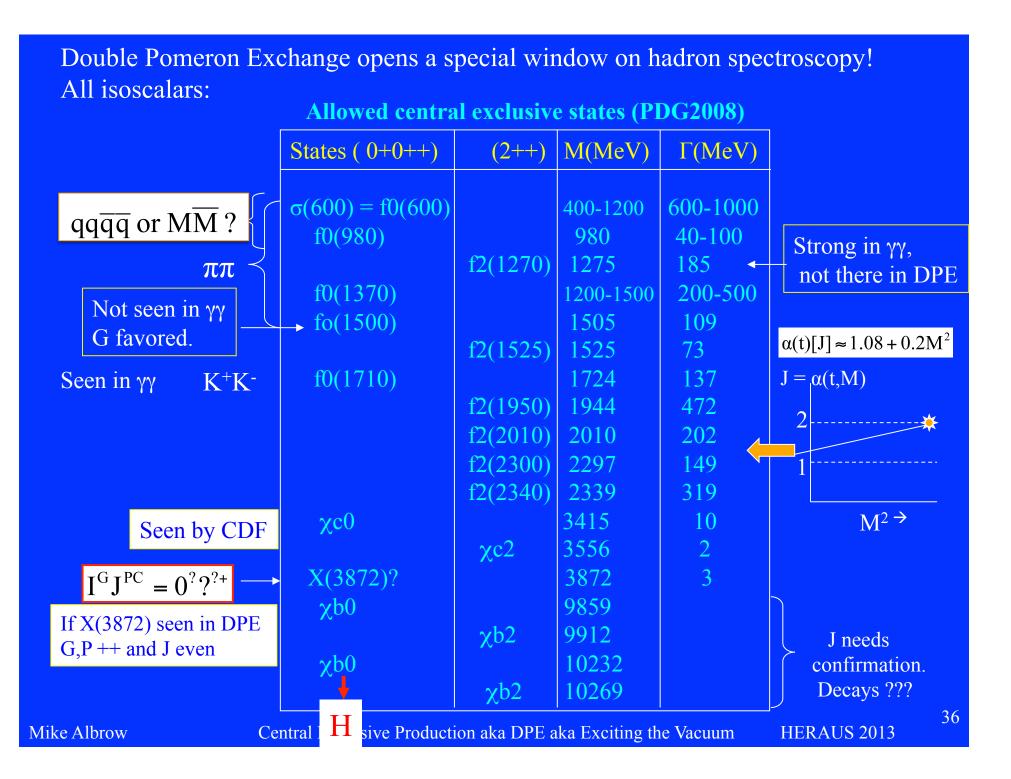
Partial wave analysis by bin above 1 GeV. Mostly S-wave (as < 1 GeV). Small D-wave structures ... $f_2(1270), f_2'(1525)?, f_2(2130)?$



Note: GXG includes t = 0. Here p-seen, |t| > min value, different.

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Name	$M(MeV/c^2)$	$\Gamma(MeV)$	$I^G J^{PC}$	$\pi\pi$	$K\bar{K}$	Other modes
$f_0(600)/\sigma$	400-1200	600-1000	0+0++	~ 100	-	-
$f_0(980)$	980 ± 10	40-100	0+0++	dominant	seen	-
$f_2(1270)$	1275.1 ± 1.2	185 ± 3	0^+2^{++}	$84.8^{+2.4}_{-1.2}$	4.6 ± 0.4	$2\pi^+2\pi^-$ 2.8%
$f_0(1370)$	1200-1500	150-250	0+0++	seen	seen	$\rho\rho$ dominant
$f_0(1500)$	1505 ± 6	109 ± 7	0+0++	34.9 ± 2.3	$8.6{\pm}1.0$	$4\pi 49.5 \pm 3.3$
$f_2'(1525)$	1525 ± 5	76 ± 10	0^+2^{++}	0.8 ± 0.2	88.7 ± 2.2	$\eta\eta \ 10.4{\pm}2.2$
$f_0(1710)$	1720 ± 6	135 ± 8	0+0++	seen	seen	$\eta\eta$ seen
$f_2(1950)$	$1944{\pm}12$	472 ± 18	0^+2^{++}	seen	seen	$\eta\eta$ seen
$f_2(2010)$	$2011\pm \sim 70$	$202\pm \sim 70$	0^+2^{++}	-	seen	$\phi\phi$ seen
$f_2(2300)$	2297 ± 28	149 ± 41	0^+2^{++}	-	seen	$\phi\phi$ seen
$f_2(2340)$	2339 ± 55	$319\substack{+81 \\ -69}$	0^+2^{++}	-	-	$\phi\phi$ seen
$f_6(2510)$	2465 ± 50	$255\pm~40$	$0^{+}6^{++}$	$6.0{\pm}1.0$	-	

Table 5: Light quark meson states allowed in DIPE. Branching fractions are in %. (PDG 2010)

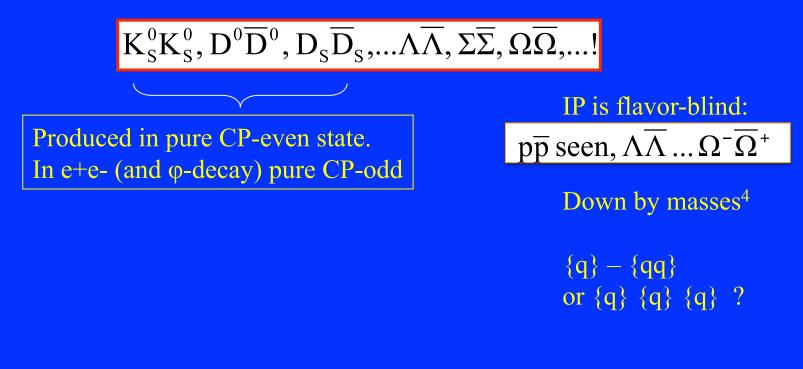
With large samples of DPE data and multi-channels identified we could clean this up! CDF now: 350,000 π + π -, to come: 4 π , K+K-. K*K*, $\eta\eta$, n'n', $\rho\rho$, $\Phi\Phi$, etc. One week of low P-U running (μ ~1) at LHC with special triggers – **fantastic!!** 2015?

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Central Exclusive Production (DPE) of hadrons

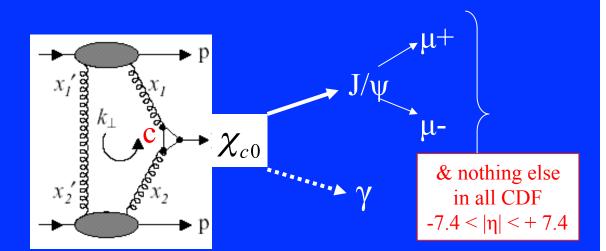
Higher energy is better, larger Δy (>5 at Tevatron)GTeV!Tevatron and LHC both good for this physics (but R.I.P. Tevatron)(LHC is overkill in terms of energy, but need low pile-up runs & triggers)CDF is an excellent detector: especially tracking, vertexing and particle ID

Establishing the spectroscopy of scalar and tensor states is one program. But there's much more!



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<u>CDF measurement exclusive $\chi c \rightarrow J/\psi + \gamma \rightarrow \mu + \mu - \gamma$ </u>



Added to CDF: Beam Shower Counters BSC: $5.2 < |\eta| < 7.4$ (5.9) Scintillator paddles tightly wrapped around beam pipes. Detect showers produced in beam pipes if p or p dissociate. e.g. $p \rightarrow p \pi \pi$ 8 + 10 counters (4 + 4)

If these are all empty, p and p did not dissociate (or BSC inefficient, could estimate from data) but went down beam pipe with small (<~ 1 GeV/c) transverse momentum.

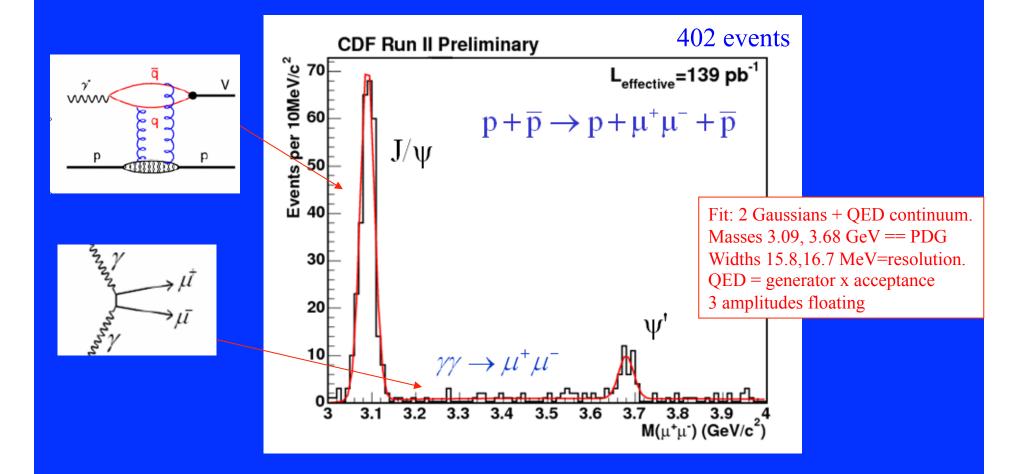


Mike AI (size greatly exaggerated!) clusive Production aka DPE aka Exciting the Vacuum

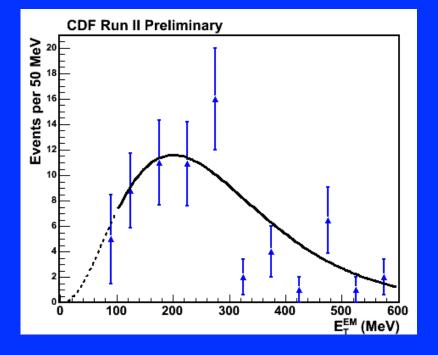
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Observation of Exclusive Charmonium Production and $\gamma\gamma \rightarrow \mu^+\mu^-$ in $p\bar{p}$ Collisions at $\sqrt{s}=1.96$ TeV.

$$p + \overline{p} \rightarrow p + \mu^+ \mu^- + \overline{p}$$



Now allow photons: EmEt spectrum with J/y mass cut:



J/ ψ have photons : 286 \rightarrow 352 ψ (2S) do not :

 $\chi_c \rightarrow J/\psi + \gamma$

Empirical functional form

65 events above 80 MeV cut. 3 events below (estimated from fit) \rightarrow 4% background under J/psi \rightarrow # $\chi_c = 65 + /-8$

MC also estimates only few % of $\chi_c \rightarrow J/\psi + \gamma$ under the cut But CDFSIMulation not reliable for such low E_T

Cannot resolve three χ_c states ... really measured $\Sigma B_i \sigma_i$ hadronic decays can resolve.

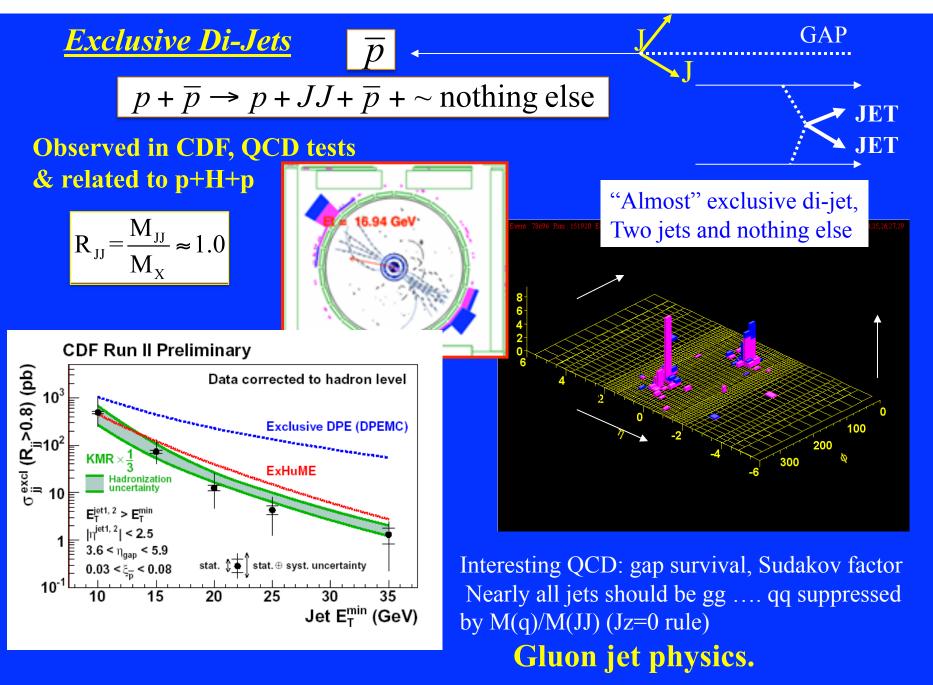
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State	$\chi_{c0}(3415)$	$\chi_{c1}(3511)$	$\chi_{c2}(3556)$
IG JPC			
10	0+0++	0+1++	0+2++
Mass(MeV):	3414.76 ± 0.35	3510.66 ± 0.07	3556.20 ± 0.09
Width (MeV):	10.4 ± 0.7	0.89 ± 0.05	2.06 ± 0.12
BF(Channel)			
$J/\psi + \gamma$	1.16 ± 0.08	35.6 ± 1.9	20.2 ± 1.0
Above with $J/\psi \to \mu^+\mu^-$	0.077	0.021	0.012
$\pi^+\pi^-\pi^+\pi^-$	2.27 ± 0.19	0.76 ± 0.26	1.11 ± 0.11
$\pi^+\pi^-K^+K^-$	1.80 ± 0.15	0.45 ± 0.10	0.92 ± 0.11
$3(\pi^+\pi^-)$	1.20 ± 0.18	0.58 ± 0.14	0.86 ± 0.18
$\pi^+\pi^-$	$0.56 {\pm} 0.03$	< 0.1	0.159 ± 0.009
K^+K^-	0.60 ± 0.03	<0.1	0.11±0.008
$\pi^+\pi^-K^0_sK^0_s$	0.58 ± 0.11	< 0.1	0.92 ± 0.11
Above with $K_S^0 \to \pi^+\pi^-$	0.27 ± 0.05	< 0.1	0.43 ± 0.05
$K^+K^-K^+K^-$	0.28 ± 0.03	0.06 ± 0.01	0.18 ± 0.02
$\pi^+\pi^-par{p}$	$0.21{\pm}0.07$	< 0.1	0.13 ± 0.03
Total %	7.2	1.9	4.7

Table 9: Branching fractions (BF in %) of χ_c states, for decays to all charged particles with BF > 0.1%.

Table 8: Upper limits on χ_{c0} cross sections.

State:	$\chi_{c0} ightarrow \pi^+\pi^-$	$\chi_{c0} ightarrow K^+K^-$	
Background (est.)	722.9	940.0	
Events in window	754	951	
90% CL upperlimit (events)	69.6	59.2	
Acceptance	24.2%	21.8%	
$d\sigma/dy _{y=0}$, 90% CL UL	21.4 ± 4.2 (syst.) nb	18.9 ± 3.8 (syst.) nb	



Forward Shower Counters, FSC, installed on both sides of CMS $|\eta| \sim 6 - 8$ (simple counters, otherwise blind region) Mainly useful as rapidity gap detectors in low – pile-up runs





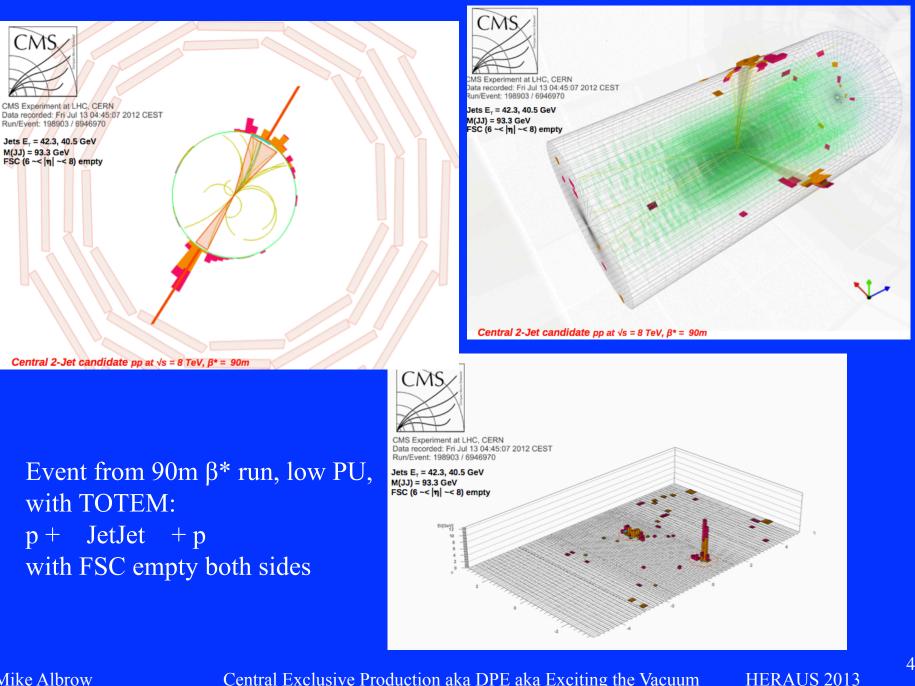
3rd station at 114 m after warm dipoles Following from BSC counters in CDF

2nd station at 85 m

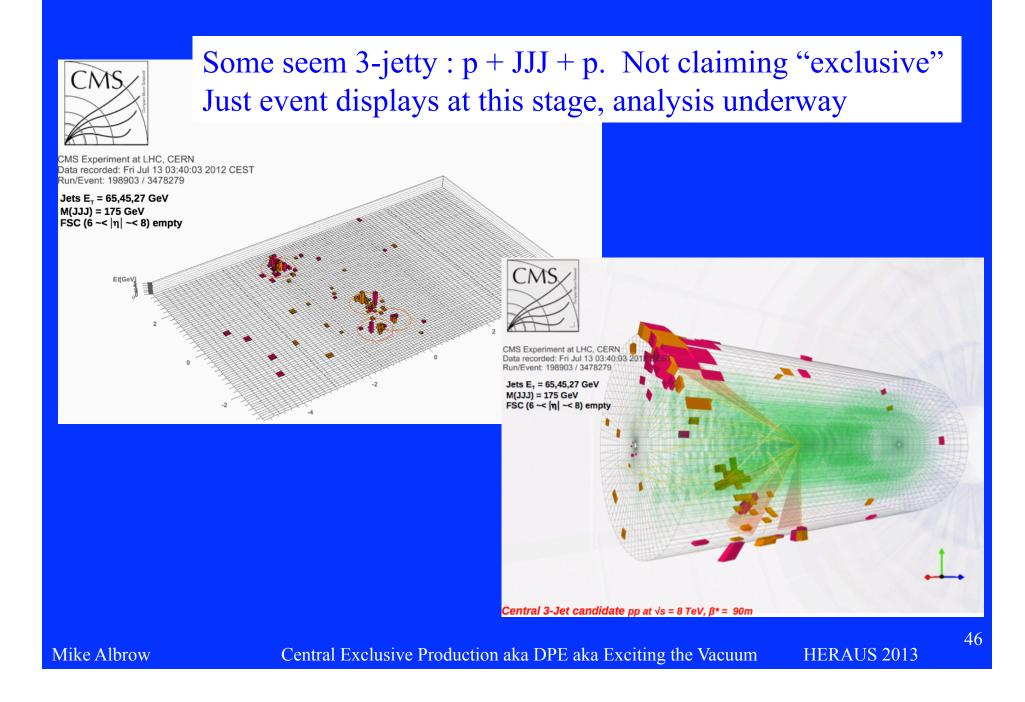
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Central Exclusive Production aka DPE aka Exciting the Vacuum

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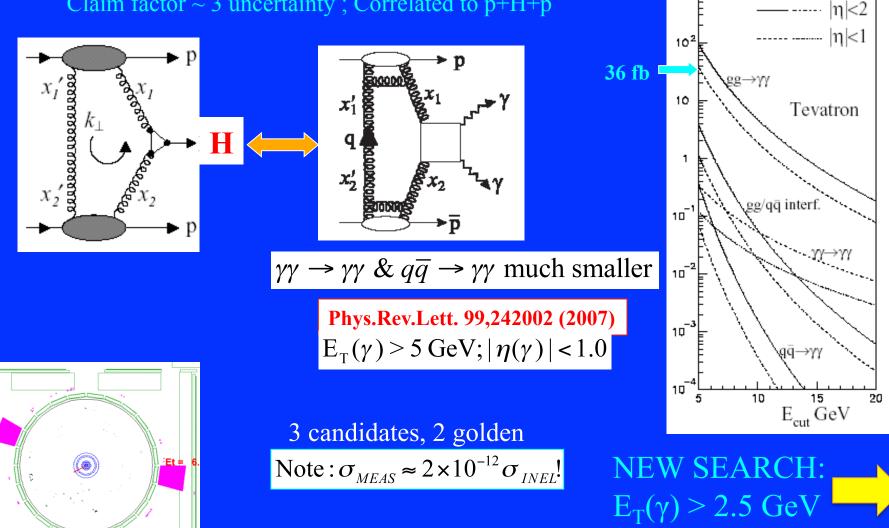
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Exclusive 2-Photon Production in CDF

Khoze, Martin and Ryskin, hep-ph/0111078, Eur.Phys.J. C23: 311 (2002) KMR+Stirling hep-ph/0409037

Claim factor ~ 3 uncertainty ; Correlated to p+H+p



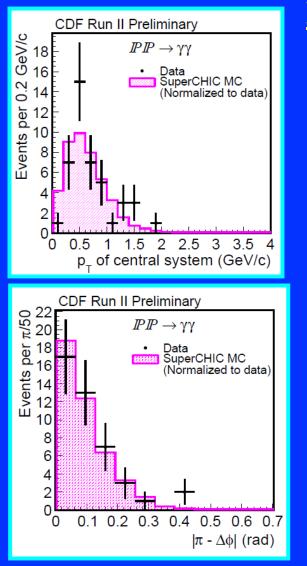
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Tevatron

 $10^{3} \frac{\sigma_{\gamma\gamma}(E_T > E_{cut})}{E_T}$ fb

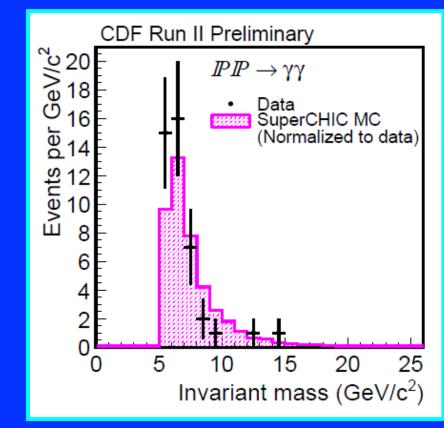
NEW SEARCH: $E_T(\gamma) > 2.5 \text{ GeV}$

43 candidate events. $\pi^0\pi^0$ background fit = 0, < 30% at 95% cl

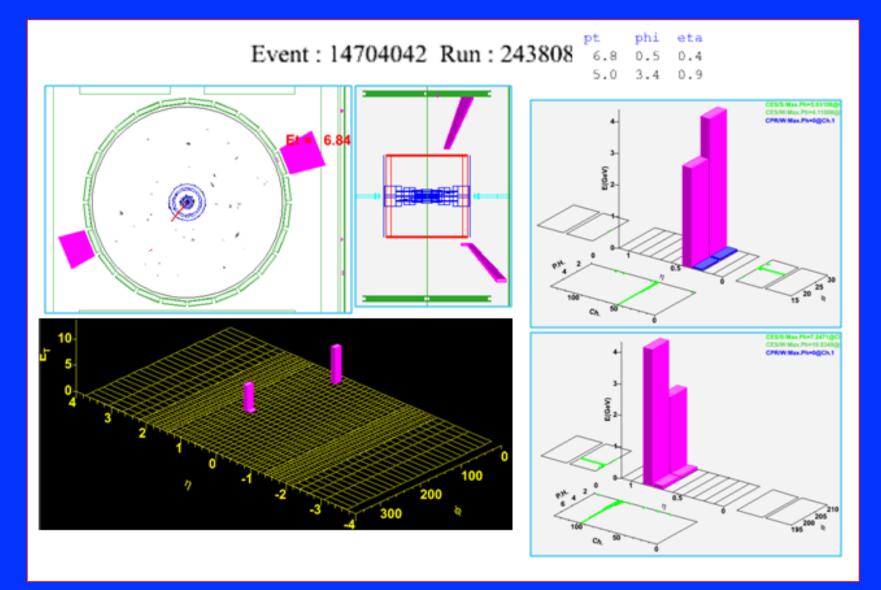


Distributions of photon pairs

Normalization to equal area (shape comparison)



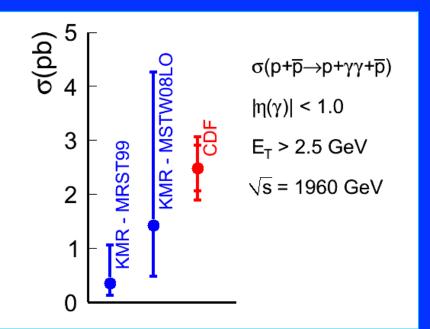
Example of yy event



$p + \overline{p} \rightarrow p + \gamma \gamma + \overline{p} via IP + IP (QCD)$

43 candidate events. $\pi^0\pi^0$ background fit = 0, < 30% at 95% cl

$\sigma_{\gamma\gamma}^{ \eta <1,E_T>2.5\text{GeV}}$	=	$2.48\pm0.42(\text{stat})\pm0.41(\text{sys})\text{pb}$
$\sigma_{ ext{SuperCHIC}}^{ \eta <1, E_T>2.5 ext{GeV}}$	=	$1.42^{ imes 3}_{\div 3}$ pb
$\sigma^{ \eta <1,E_T>2.5\mathrm{GeV}}_{\mathrm{SuperCHIC}(\mathrm{MRST99})}$	=	$0.35^{ imes 3}_{\div 3}$ pb

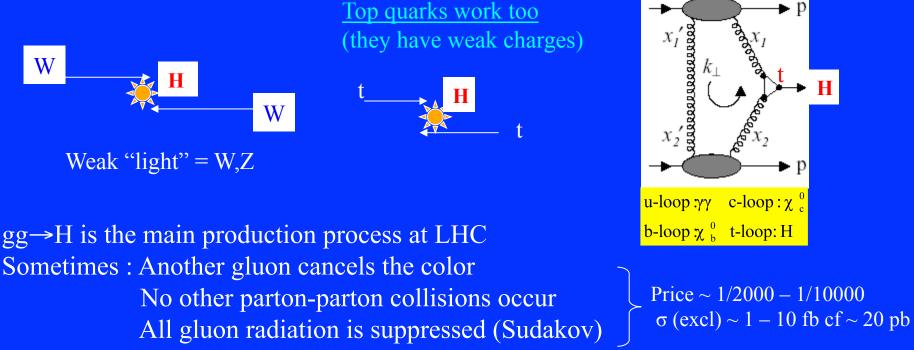


Implies that p + H + phappens at LHC (if H) with $\sigma(SMH120) \sim 3-10$ fb

Exclusive χ_c and JJ agree too

The Ultimate Vacuum Excitation

Above the χb , the only "known" heavier particle with vacuum Q.Nos. is the Higgs. Vacuum is everywhere Higgs field. ANTIBLINK equickly enough and with weak eyes you can "see" them. Hit the vacuum hard with a pair of weak probes, and you can promote them Virtual \rightarrow Real



But if you see them, and measure both p, powerful benefits

$$M_{\text{CEN}} = \sqrt{(p_1 + p_2 - p_3 - p_4)^2} \implies \sigma(M_H) \approx 2 \text{ GeV per event}$$

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 $J,C,P,\Gamma?$

FP420 : Forward Protons 420m & 220m downstream of CMS & ATLAS



CMS: Inner Vacuum Tank insertion

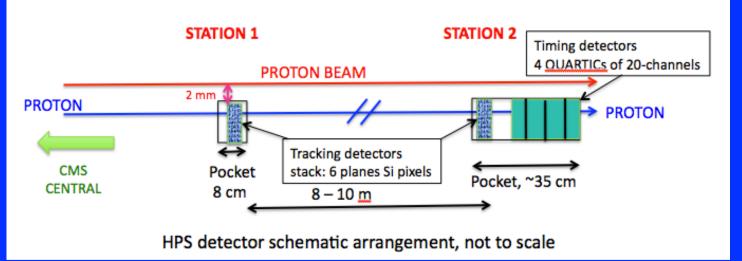
 $\gamma\gamma \rightarrow WW; H \rightarrow WW, H \rightarrow ZZ$ $H \rightarrow b\overline{b}; BSM \rightarrow WW, SUSY etc$



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At 240 m (& 220 m) the beam pipe is exposed, so it's relatively easy to install

the detectors:



At 420 m missing magnet, pipe straight but cold. It requires a cryogenic by-pass Needed for p + H(126) + p

> Looking downstream (CMS behind you)

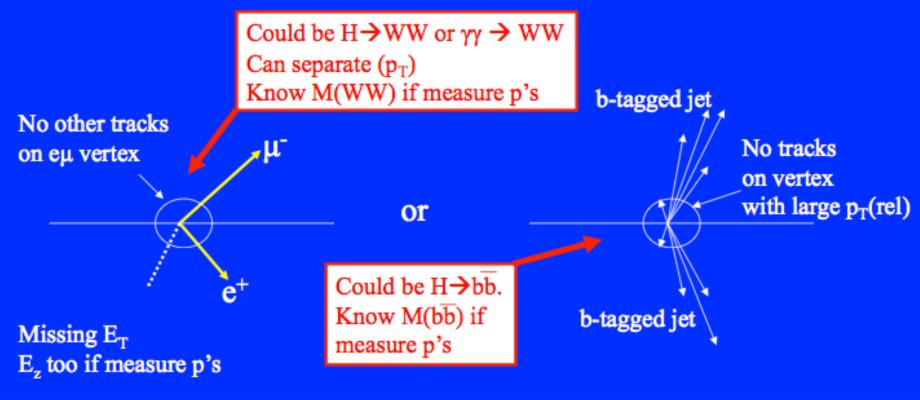


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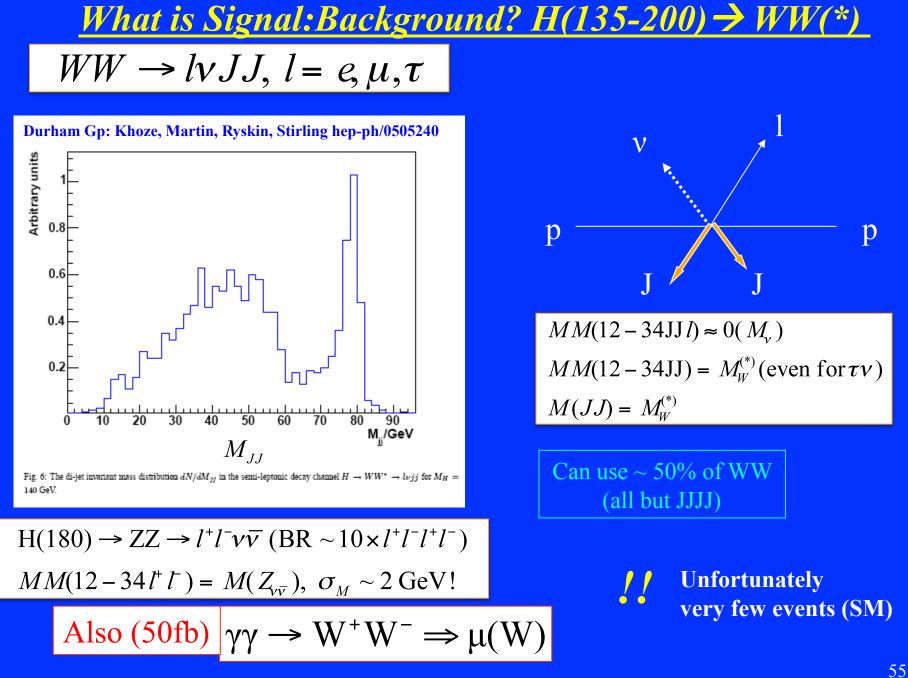
The cleanest, simplest inelastic pp collisions.

Class 3 Interactions at the LHC: "Inelastic, with no hadrons produced"

Consider WW + nothing (p's go down pipe, small p_T) $\sigma (\gamma \gamma \rightarrow W^+W^-) \sim 50 \text{ fb} \dots \text{ or } H + \text{ nothing}$



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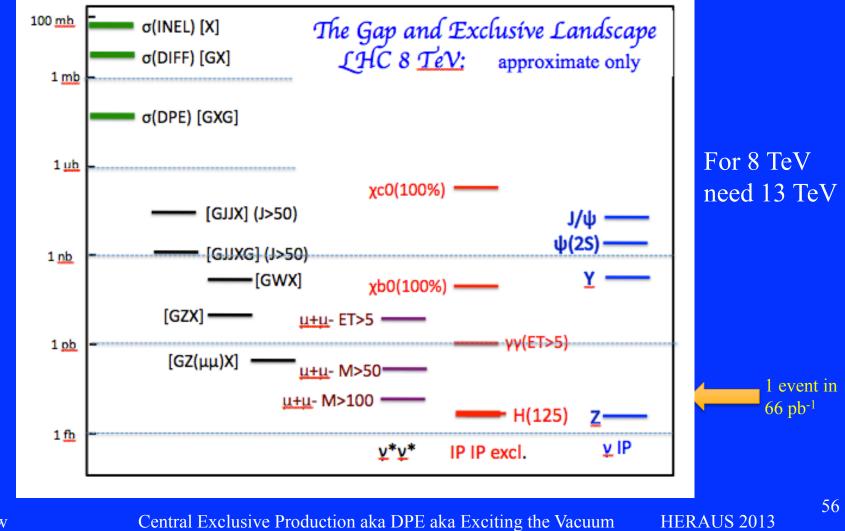
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What could one do in a low pile-up (<n/x> = 1) week?

1 week of $\langle n/x \rangle = 1$ running: 120 hours x 3600 secs x 2800 bunches x 11,625 O/s = 1.4 10¹³ X/week 1X == 80 mb \rightarrow 1.8 10¹¹ mb⁻¹/week = 180 pb⁻¹ / week \rightarrow 66 pb⁻¹ of no-PU events If for $\gamma\gamma$ (sa) PU \sim 5 can be used, \sim 0.8 fb⁻¹ gets \sim 1000 $\gamma\gamma$ events



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<u>Summary</u>

The vacuum (QCD, Electroweak, ..) is teeming with states, from e+e- loops (even v-loops if you have weak eyes) to H (maybe).

If you truly understand the vacuum you'll understand all of physics.

Theoretically amazing; experimentally we can excite it with a pair of hammers

Electromagnetic probes: $\gamma\gamma \rightarrow$ charged pairs $e^+e^-, \mu^+\mu^-, \tau^+\tau^-, \dots, W^+W^-, \tilde{l}\tilde{l}$? Strong probes: IP IP \rightarrow Neutral states (G?) & Pairs $\sigma, f_0, G, \chi_c, \chi_b, D\overline{D}, \Omega\overline{\Omega}$... Weak probes: WW \rightarrow H, etc... $e^+e^- \rightarrow \nu + W^+W^- + \nu \rightarrow \nu + H + \nu$

There is much more to be done: Tevatron data and especially LHC (low & high PU)

Use the LHC as a $\gamma\gamma$ and gg collider with tiny but high precision detectors (present and future)

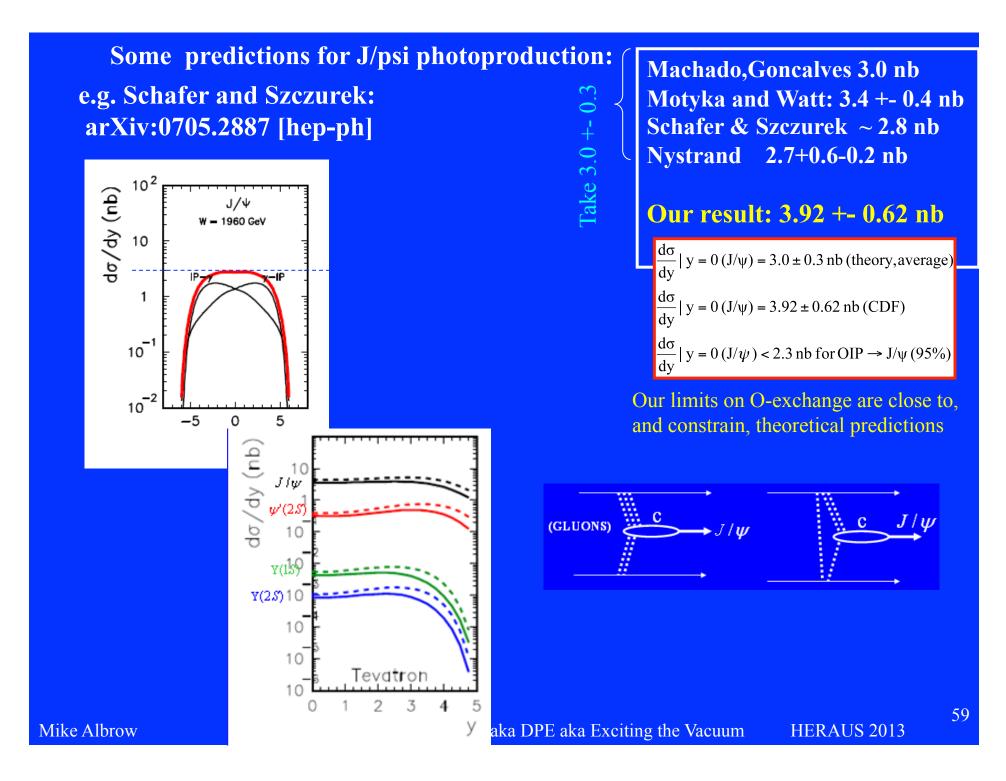
Thank you

Overflow slides :

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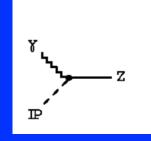
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Exclusive Z production : CDF Search

Allowed in SM (like V) but $\sigma \sim 0.3$ fb (Motyka+Watt)

 $p \xrightarrow{\gamma^* \vee q} Z^0 \xrightarrow{Z^0} \overline{p}$ $\overline{p} \xrightarrow{p} \overline{p}$ Could be enhanced by BSM loops?



Interesting?! γ-IP-Z eff.coupling. ZOOM IN to see how!

 $2.2/\text{fb}: 318\text{K} e+e-\&\mu+\mu-M > 40 \text{ GeV}; 183\text{K} \text{ in Z window } 82-98 \text{ GeV}$ Require no other interaction, no additional tracks, all calorimeters in noise (E)

