

Theoretical status of $\bar{B} \rightarrow X_s \gamma$ and $\bar{B} \rightarrow X_s \ell^+ \ell^-$

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1. Introduction
2. The isospin asymmetry in $\bar{B} \rightarrow X_s \gamma$
3. Resolved photons in $\bar{B} \rightarrow X_s \gamma$ and $\bar{B} \rightarrow X_s \ell^+ \ell^-$
4. Cuts in q^2 and M_{X_s}
5. Status of the perturbative calculations
6. Summary

The most important operators for $\bar{B} \rightarrow X_s \gamma$ and $\bar{B} \rightarrow X_s \ell^+ \ell^-$:

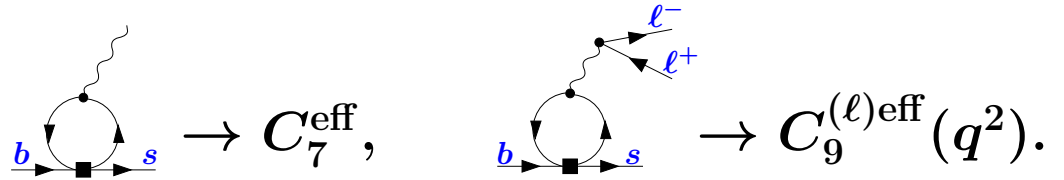
$$Q_7 = \frac{em_b}{16\pi^2} \bar{s}_L \sigma_{\mu\nu} b_R F^{\mu\nu} \quad Q_{9(10)}^{(\ell)} = \frac{e^2}{16\pi^2} [\bar{s}_L \gamma^\alpha b_L] [\bar{\ell} \gamma_\alpha (\gamma_5) \ell]$$

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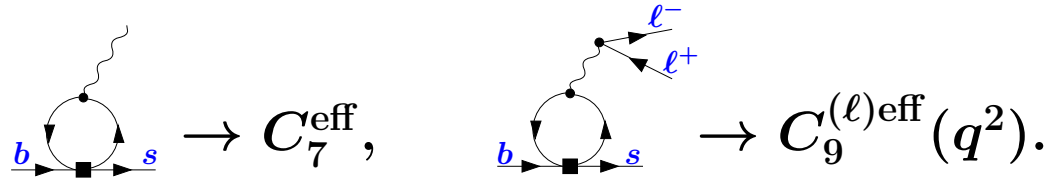


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$$\text{Diagram 1} \rightarrow C_7^{\text{eff}}, \quad \text{Diagram 2} \rightarrow C_9^{(\ell)\text{eff}}(q^2).$$

$[R_K, R_{K^*}, P'_5, \dots \text{ in } B \rightarrow K^{(*)} \ell^+ \ell^-]$ can be explained, e.g., by $C_9^{(\mu)\text{NP}} = -C_{10}^{(\mu)\text{NP}} = (-0.67 \pm 0.16)$.

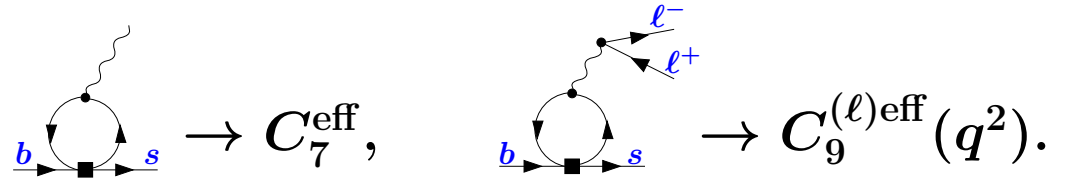
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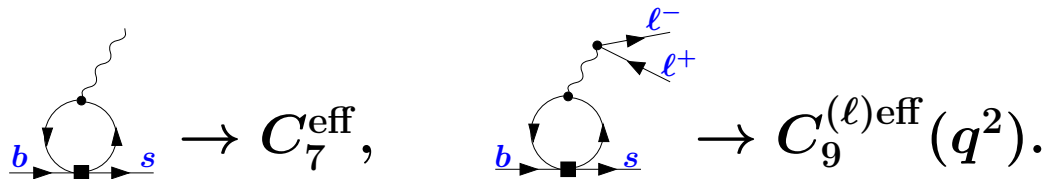
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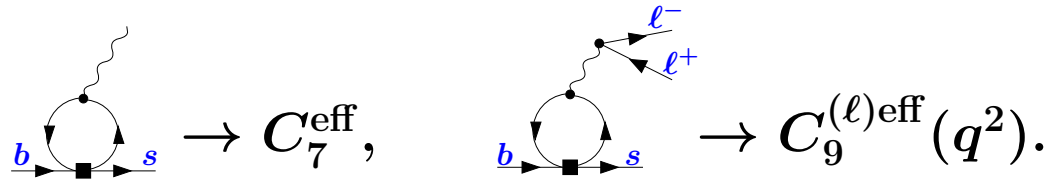
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$$\frac{d^2\Gamma}{dq^2 dz} = \frac{3}{8} [(1 + z^2)H_T(q^2) + 2(1 - z^2)H_L(q^2) + 2zH_A(q^2)], \quad z = \cos[\angle(\vec{p}_{\ell^+}, \vec{p}_B)] \text{ in the dilepton frame}$$

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$|C_{10}^{(\mu)}|$ from the average time-integrated $B_s \rightarrow \mu^+ \mu^-$ branching ratio:

$$\left. \begin{array}{l} \bar{\mathcal{B}}_{s\mu}^{\text{exp}} \times 10^9 = 2.93 \pm 0.46 \\ \bar{\mathcal{B}}_{s\mu}^{\text{SM}} \times 10^9 = 3.54 \pm 0.21 \end{array} \right\} \Rightarrow \left| \frac{C_{10}^{(\mu)}}{C_{10}^{(\mu)\text{SM}}} \right| = 0.91 \pm 0.08 \quad (1.1\sigma)$$

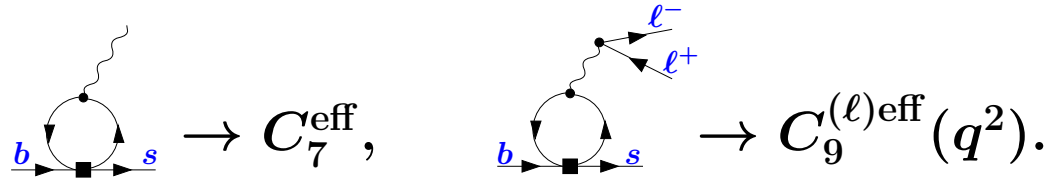
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The same central values would give $> 2\sigma$ with the future $\pm 6\%$ in $\bar{\mathcal{B}}_{s\mu}^{\text{exp}}$ [LHCb].

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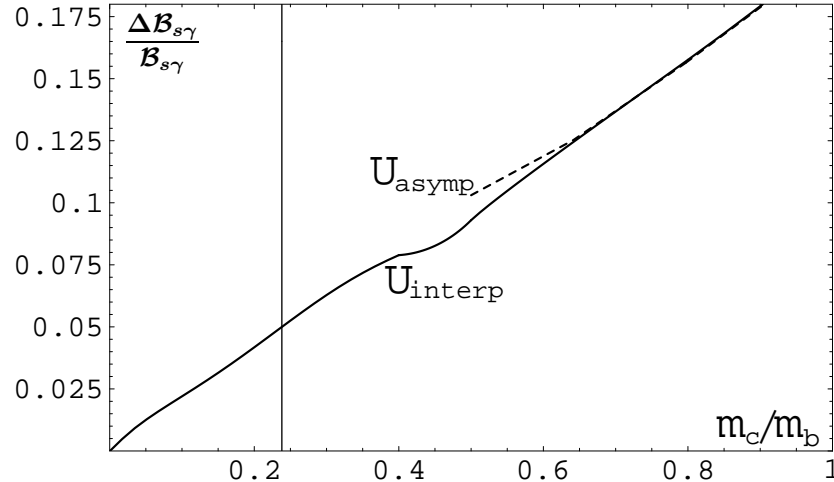
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Effect of the interpolated $\mathcal{O}(\alpha_s^2)$ contribution on the branching ratio:



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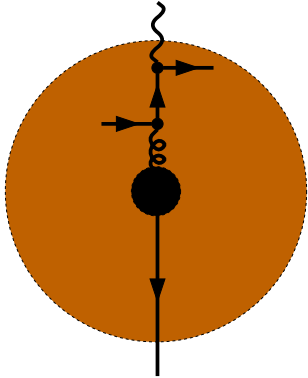
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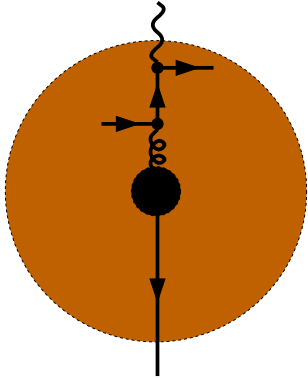
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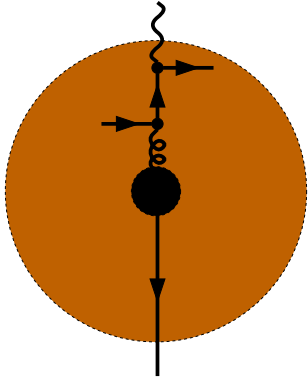
Dominant in Δ_{0-} : $\Gamma[B^- \rightarrow X_s \gamma] \simeq A + BQ_u + CQ_d + DQ_s, \quad \Gamma[\bar{B}^0 \rightarrow X_s \gamma] \simeq A + BQ_d + CQ_u + DQ_s$

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$$\text{Isospin-averaged decay rate:} \quad \Gamma \simeq A + \frac{1}{2}(B + C)(Q_u + Q_d) + DQ_s \equiv A + \delta\Gamma_{78\text{res}}$$

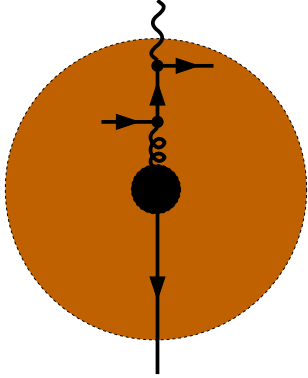
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Q_8-Q_8	$[-0.3, 1.9]\%$	no change	$(0.80 \pm 1.10)\%$	no change
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\Leftarrow Belle Δ_{0-}
[arXiv:1807.04236v4](https://arxiv.org/abs/1807.04236v4)

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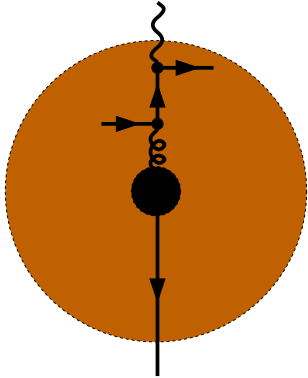
$$\Rightarrow \frac{\delta\Gamma_{78\text{res}}/\Gamma}{\Delta_{0-}} \simeq \frac{(B+C)(Q_u+Q_d)+2DQ_s}{(C-B)(Q_u-Q_d)} = \frac{Q_u+Q_d}{Q_d-Q_u} \left[1 + 2 \frac{D-C}{C-B} \right]$$

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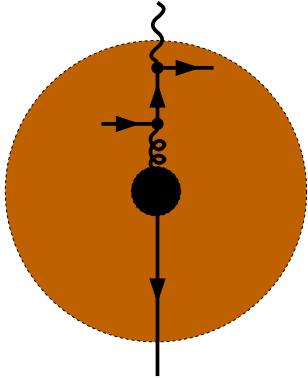
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$$\frac{\delta\Gamma_{78res}}{\Gamma} \simeq -\frac{1}{3}\Delta_{0-} \left[1 + 2 \frac{D-C}{C-B} \right] = -\frac{1}{3}(-0.48 \pm 1.49 \pm 0.97 \pm 1.15)\% \times (1 \pm 0.3) = (0.16 \pm 0.74)\% \quad \boxed{4}$$

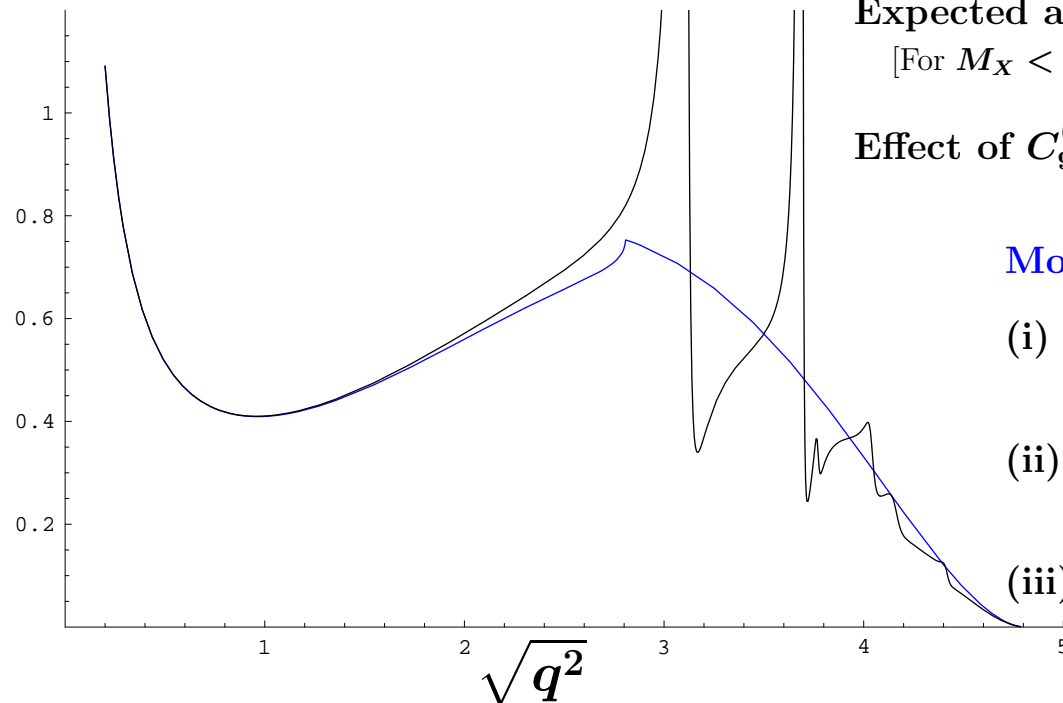
SM predictions for $\mathcal{B}(\bar{B} \rightarrow X_s \ell^+ \ell^-) \times 10^6$ with $q^2 \in [1, 6]$ GeV and no cut on M_{X_s} :

$\left. \begin{array}{l} 1.64 \pm 0.11, \ell = e \\ 1.59 \pm 0.11, \ell = \mu \end{array} \right\}$ parametric and perturbative uncert. only [Huber, Lunghi, MM, Wyler, hep-ph/0512066]

$\left. \begin{array}{l} 1.67 \pm 0.10, \ell = e \\ 1.62 \pm 0.09, \ell = \mu \end{array} \right\}$ param. update + Krüger-Sehgal, [Huber, Hurth, Lunghi, arXiv:1503.04849]
 PLB380 (1996) 199
 PRD55 (1997) 2799

The corresponding semi-inclusive experimental results:

$\left. \begin{array}{l} 1.93_{-0.51}^{+0.54}, \ell = e \\ 0.66_{-0.80}^{+0.88}, \ell = \mu \end{array} \right\} = 1.60_{-0.45}^{+0.48}$, Babar, arXiv:1312.5364, $471 \times 10^6 B\bar{B}$, extrapolated from $M_X < 1.8$ GeV.
 $1.49_{-0.58}^{+0.65}$, Belle, hep-ex/0503044, $152 \times 10^6 B\bar{B}$.



Expected accuracy at Belle II with 50 ab^{-1} : $\sim \pm 6.5\%$.
 [For $M_X < 2.0$ GeV, Table 64 of The Belle II Physics Book, arXiv:1808.10567]

Effect of $C_9^{(\mu)\text{NP}} = -C_{10}^{(\mu)\text{NP}} = -0.67$: $\mathcal{B} \searrow \sim 30\%$.

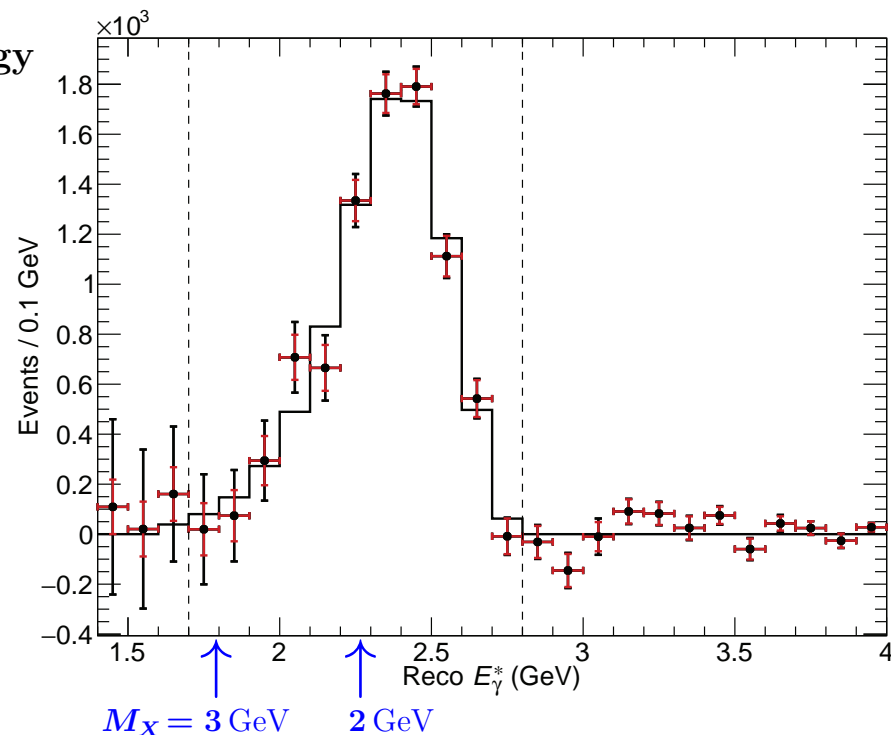
Most recent contributions on the theory side:

- (i) Benzke, Hurth, Turczyk, arXiv:1705.10366, charm loops & resolved photons \rightarrow next slides.
- (ii) de Boer, arXiv:1707.00988, two loops analytically.
- (iii) Huber, Qin, Vos, arXiv:1806.11521, five body & pheno update \rightarrow next speaker.

[Qin Qin]

More on resolved photons in $\bar{B} \rightarrow X_s \gamma$ and $\bar{B} \rightarrow X_s l^+ l^-$

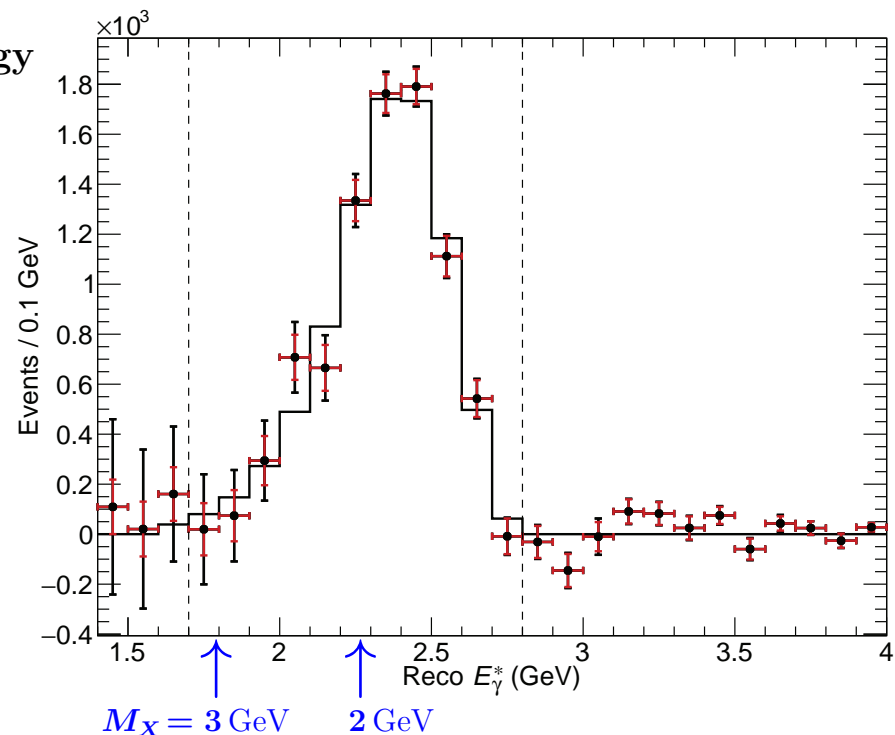
Background-subtracted $\bar{B} \rightarrow X_{s+d} \gamma$ photon energy spectrum in the $\Upsilon(4S)$ rest frame, from Fig. 1 of the Belle analysis in arXiv:1608.02344.



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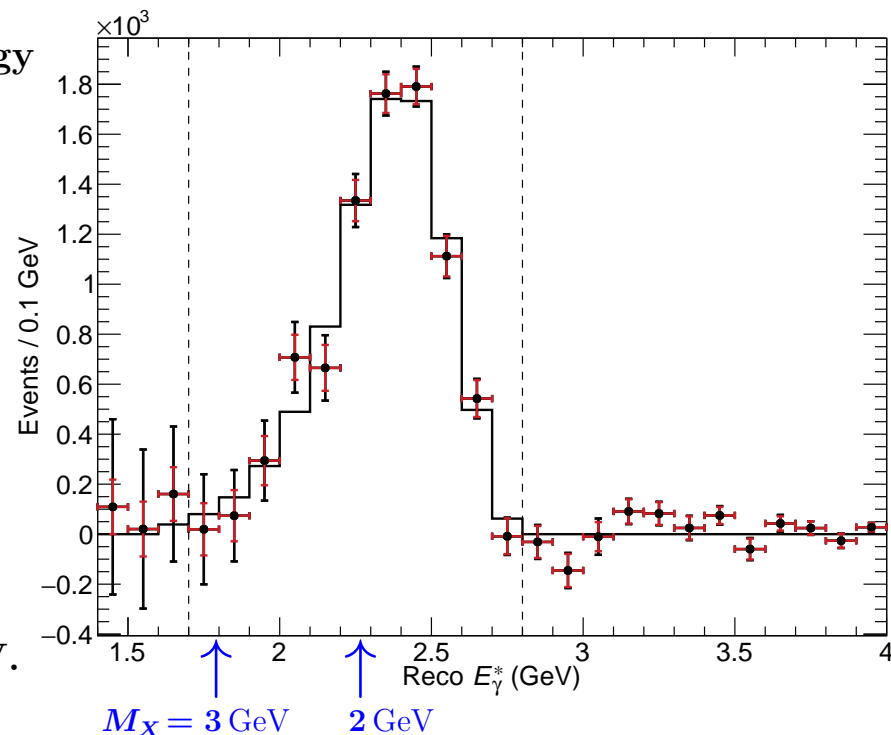
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In the presence of 4-quark ops:

Light quark loops – suppressed by $C_{3,\dots,6}$ or CKM;

Charm loops – factorizable or local if m_c^2 is sufficiently large w.r.t. $m_b \Lambda$. Numerically, $\mathcal{O}(3\%)$ non-fact. effects found in $\mathcal{B}(\bar{B} \rightarrow X_s \gamma)$ and $\mathcal{B}(\bar{B} \rightarrow X_s \ell^+ \ell^-)$ with $q^2 \in [1, 6] \text{ GeV}^2$.

[Buchalla, Isidori, Rey, NPB 511 (1998) 594]



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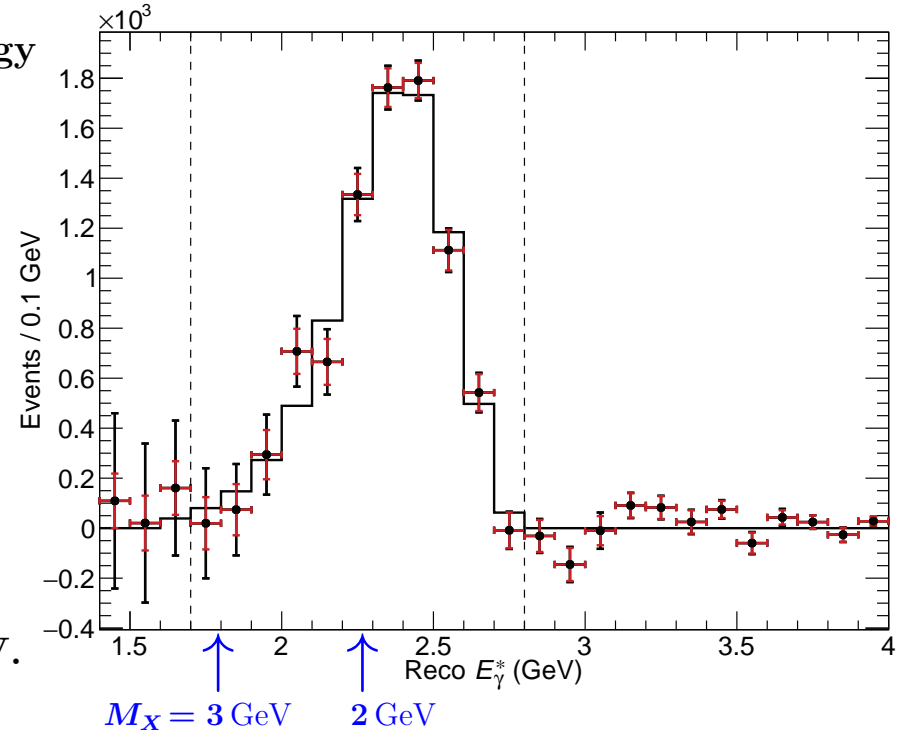
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However, m_c^2 is not sufficiently large \Rightarrow Treat it as $\mathcal{O}(m_b \Lambda)$ and use SCET, so far up to $\mathcal{O}(\Lambda/m_b)$:

For $\mathcal{B}(\bar{B} \rightarrow X_s \gamma)$ with $E_\gamma > 1.6 \text{ GeV}$ [Benzke, Lee, Neubert, Paz, JHEP 1008 (2010) 099] $[-4.8\%, +5.6\%]$ uncert. range.

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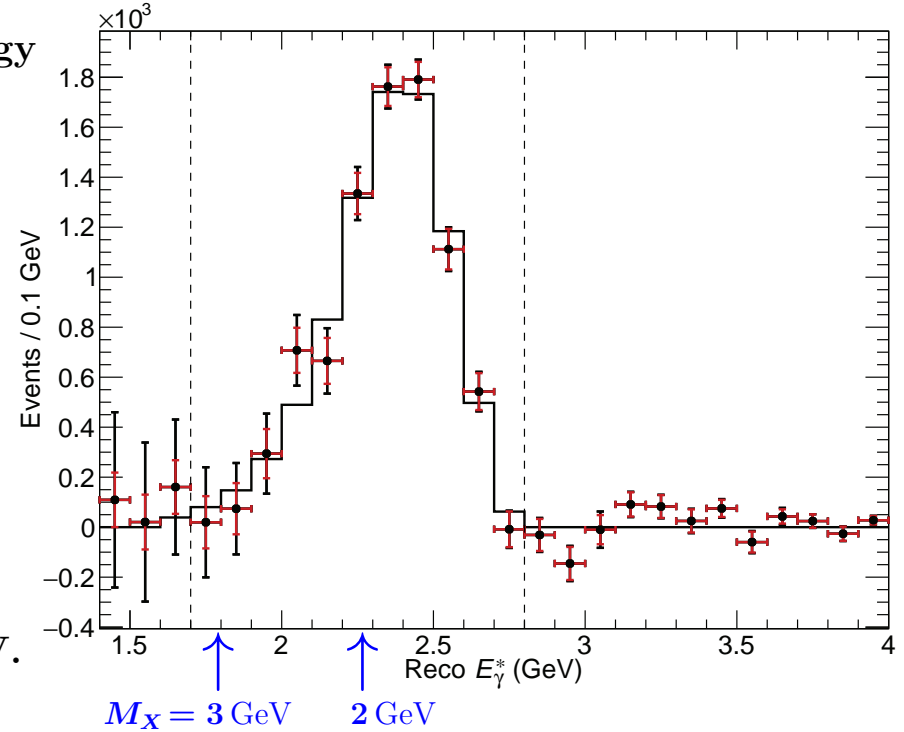
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Corrections not involving Q_7 and Q_8 are of higher order, i.e. $\mathcal{O}\left[\left(\frac{\Lambda}{m_b}\right)^a\right]$ with $a \geq \frac{3}{2}$ and/or $\mathcal{O}\left(\frac{\alpha_s \Lambda}{m_b}\right)$.

That's what we miss using the purely perturbative expression for $|C_9^{\text{eff}}(q^2)|^2$ and the local $1/m_c^2$ effects.

However, the applied SCET power counting works only for small M_X and small q^2 – verte.

SCET power counting in the $\bar{B} \rightarrow X_s \ell^+ \ell^-$ analysis of Benzke, Hurth and Turczyk, JHEP 1710 (2017) 031.

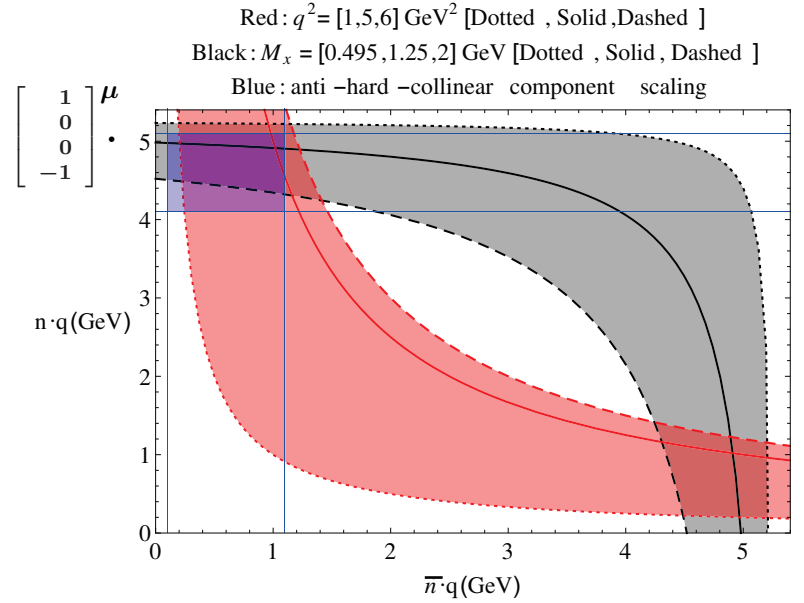
$$\lambda = \frac{\Lambda}{M_B}, \quad M_X \lesssim \sqrt{M_B \Lambda} = M_B \sqrt{\lambda} \sim m_c.$$

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↑
any vector

In the plot $q_\perp = 0 \Rightarrow q^2 = (\bar{n}q)(nq)$.

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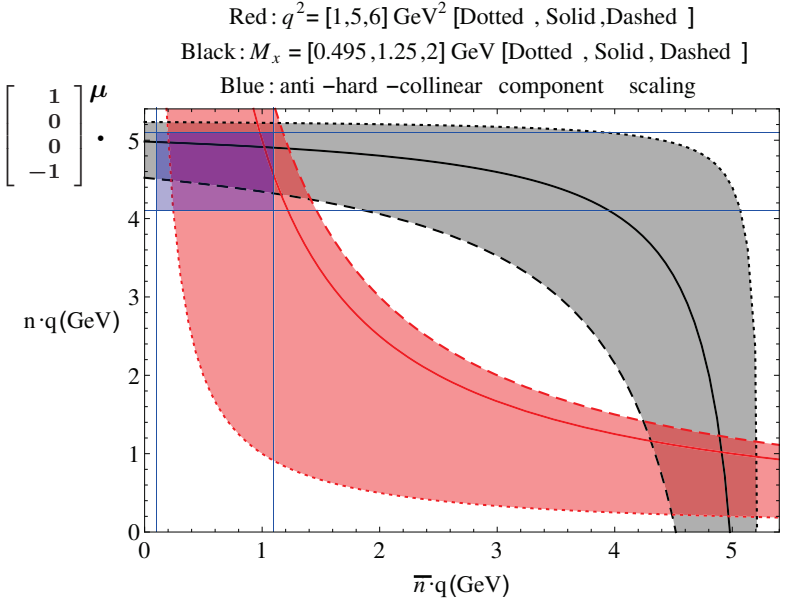
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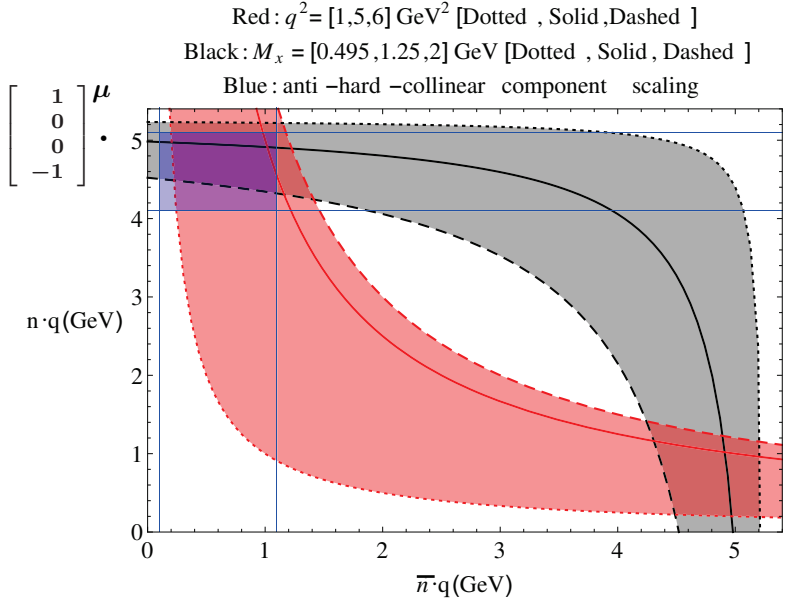
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The factorization formula:

$$d\Gamma = \sum_{n=0}^{\infty} \frac{1}{m_b^n} \sum_i H_i^{(n)} J_i^{(n)} \otimes S_i^{(n)} + \sum_{n=1}^{\infty} \frac{1}{m_b^n} \left[\sum_i H_i^{(n)} J_i^{(n)} \otimes S_i^{(n)} \otimes \bar{J}_i^{(n)} + \sum_i H_i^{(n)} J_i^{(n)} \otimes S_i^{(n)} \otimes \bar{J}_i^{(n)} \otimes \bar{J}_i^{(n)} \right]$$



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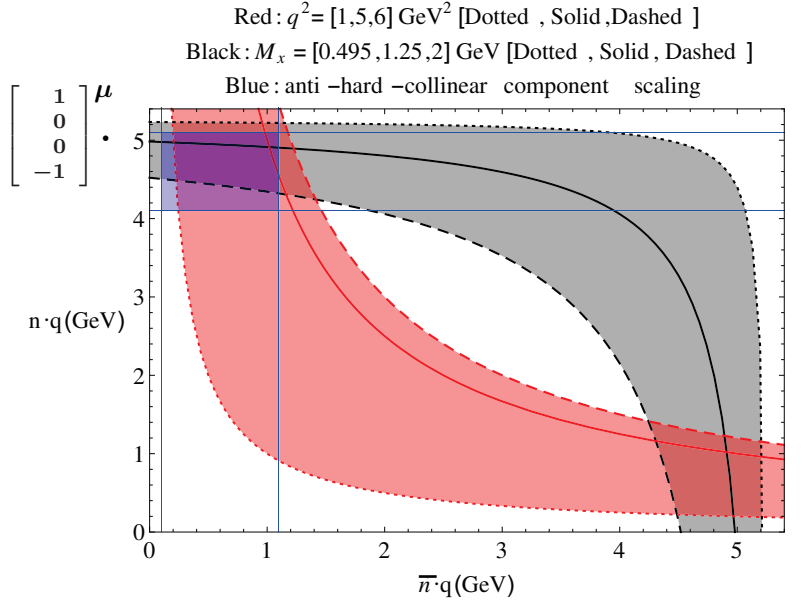
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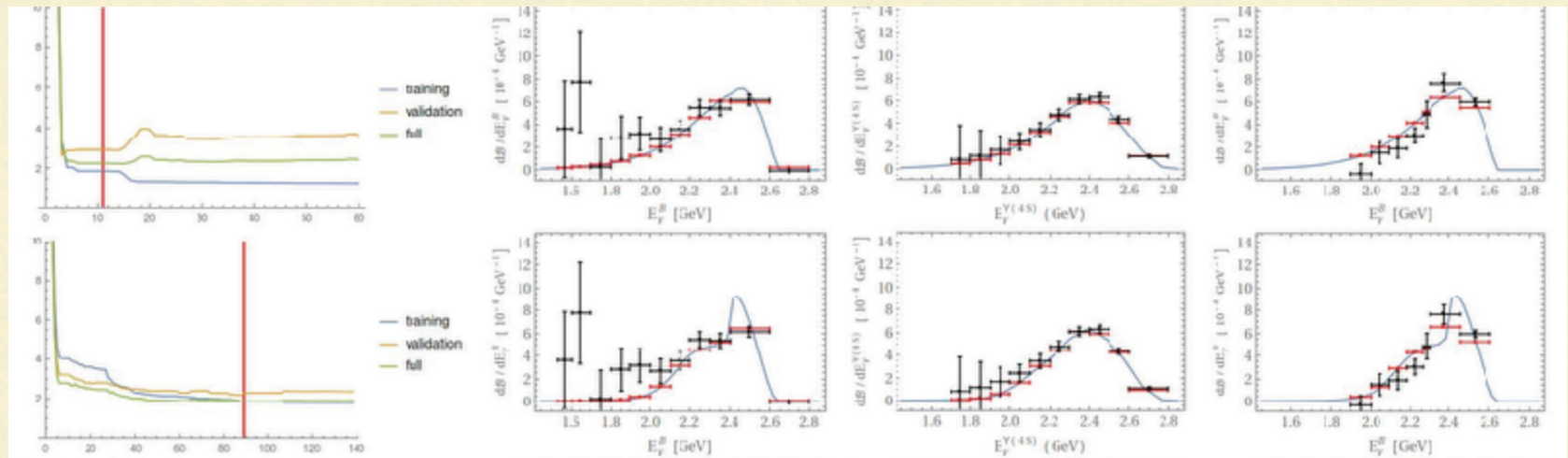
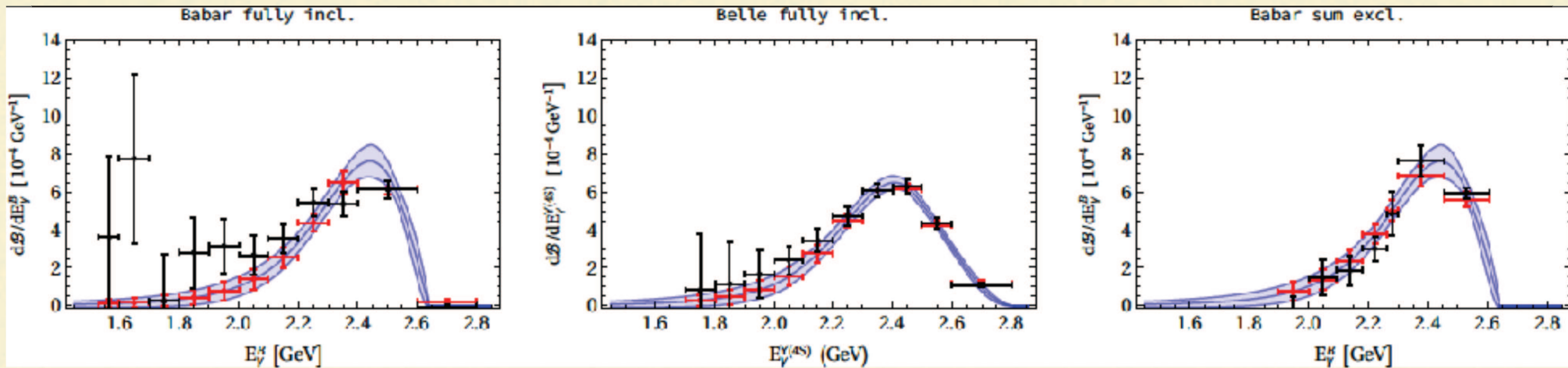
Remarks:

1. Not proven. Contradictions observed in the $Q_8 - Q_8$ case, claimed to be phenomenologically irrelevant.
2. Relates unknowns to unknowns. Models of soft functions needed (constraints available).
3. Corrections beyond $\mathcal{O}(\Lambda/M_B)$ are likely to be relevant because $|C_{9,10}/C_7| \sim 13$ (work in progress) [BHT].
4. Other observables – after including the above corrections.
5. Different power counting than in the previous SCET analyses where no "resolved photons" were included [Lee, Stewart, PRD74 (2006) 014005], [Bell, Beneke, Huber, Li, NPB843 (2011) 143].



The $b \rightarrow sy$ spectrum

Gambino, Lunghi, Misiak, Schacht
in progress



Up-to-date theoretical description of spectrum to get leading SF and reliable extrapolation to low cuts.

NNLO QCD corrections to $\bar{B} \rightarrow X_s \gamma$

The relevant perturbative quantity $P(E_0)$:

$$\frac{\Gamma[b \rightarrow X_s \gamma]_{E_\gamma > E_0}}{\Gamma[b \rightarrow X_u e \bar{\nu}]} = \left| \frac{V_{ts}^* V_{tb}}{V_{ub}} \right|^2 \frac{6\alpha_{\text{em}}}{\pi} \underbrace{\sum_{i,j} C_i(\mu_b) C_j(\mu_b) K_{ij}}_{P(E_0)}$$

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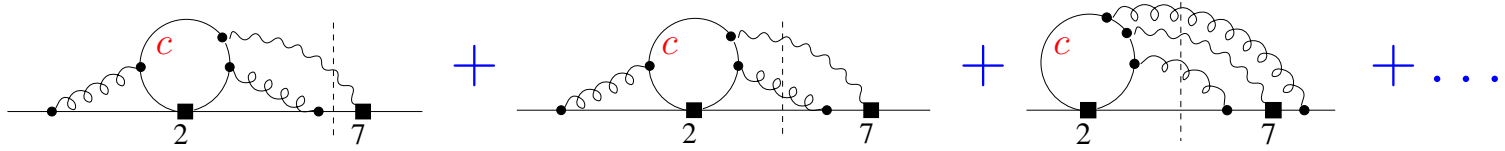
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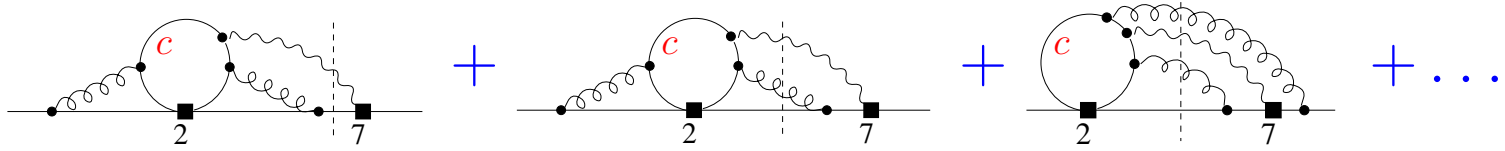
Most important at the NNLO: $K_{77}^{(2)}$, $K_{27}^{(2)}$ and $K_{17}^{(2)}$.

They depend on $\frac{\mu_b}{m_b}$, $\delta = 1 - \frac{2E_0}{m_b}$ and $z = \frac{m_c^2}{m_b^2}$.

Towards complete $K_{17}^{(2)}$ and $K_{27}^{(2)}$ for arbitrary m_c [MM, A. Rehman, M. Steinhauser, ...]
in progress

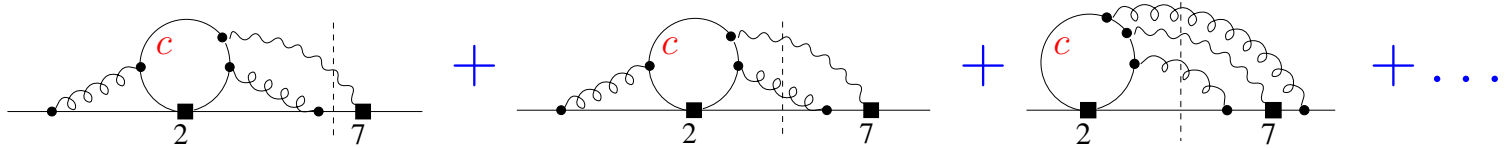


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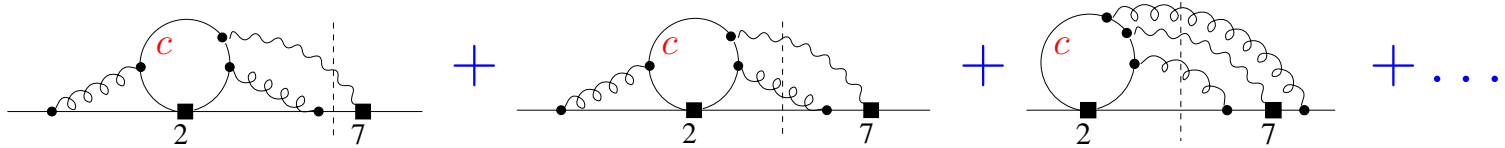
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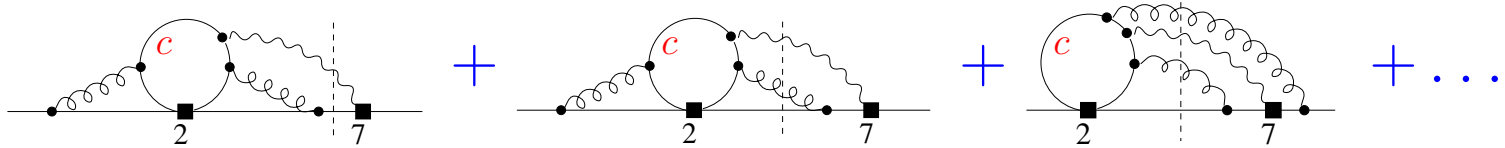
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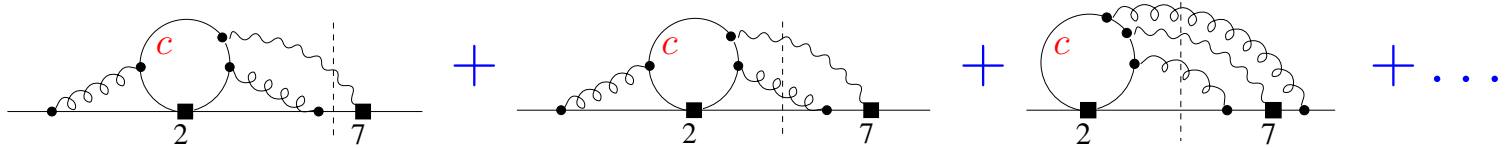
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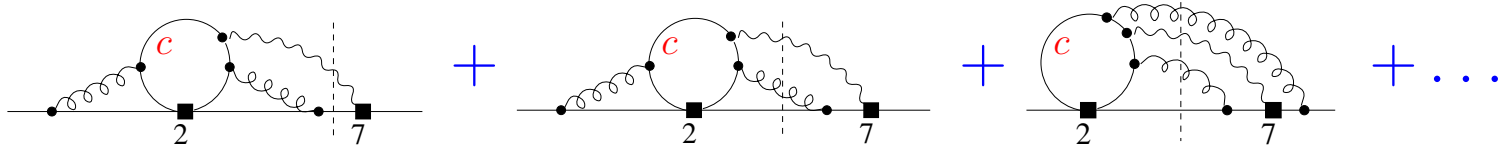
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6. Solving the system (*) numerically [A.C. Hindmarch, <http://www.netlib.org/odepack>] along an ellipse in the complex z plane. Doing so along several different ellipses allows us to estimate the numerical error.

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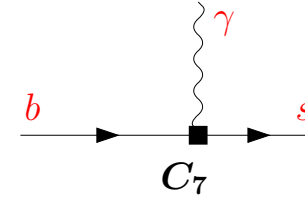
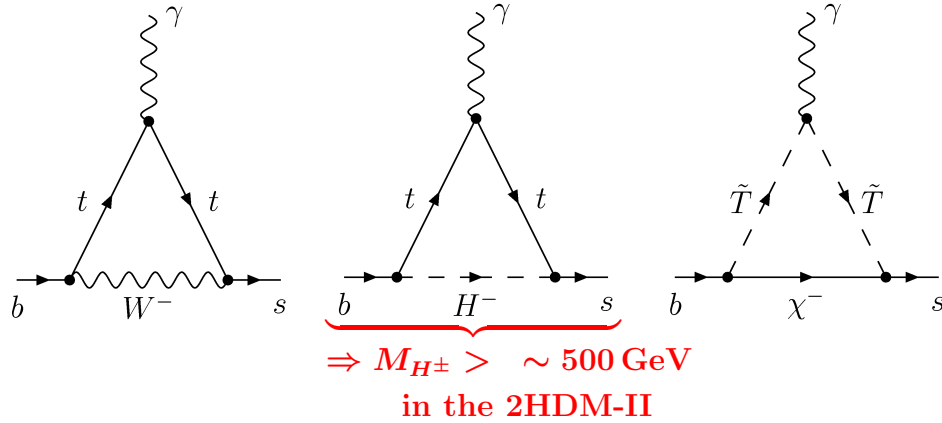
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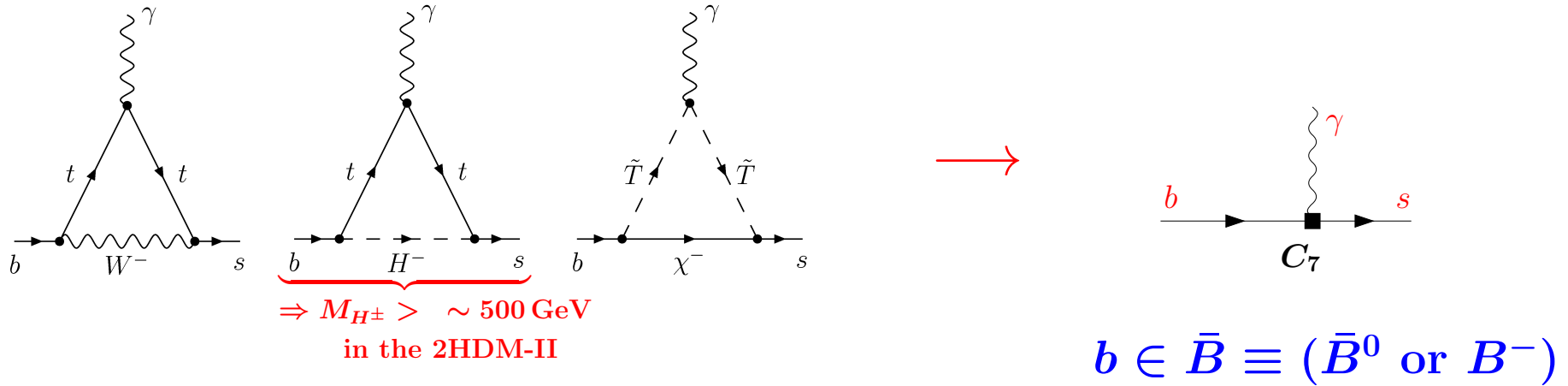
BACKUP SLIDES

Information on electroweak-scale physics in the $b \rightarrow s\gamma$ transition is encoded in an effective low-energy local interaction:



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The inclusive $\bar{B} \rightarrow X_s \gamma$ decay rate for $E_\gamma > E_0$ is well approximated by the corresponding perturbative decay rate of the b -quark:

$$\Gamma(\bar{B} \rightarrow X_s \gamma) = \Gamma(b \rightarrow X_s^p \gamma) + \left(\begin{array}{c} \text{non-perturbative effects} \\ (3 \pm 5)\% \end{array} \right)$$

[G. Buchalla, G. Isidori and S.-J. Rey, Nucl. Phys. B511 (1998) 594]
 [M. Benzke, S.J. Lee, M. Neubert and G. Paz, JHEP 1008 (2010) 099]
 (BLNP)

provided E_0 is large ($E_0 \sim m_b/2$)

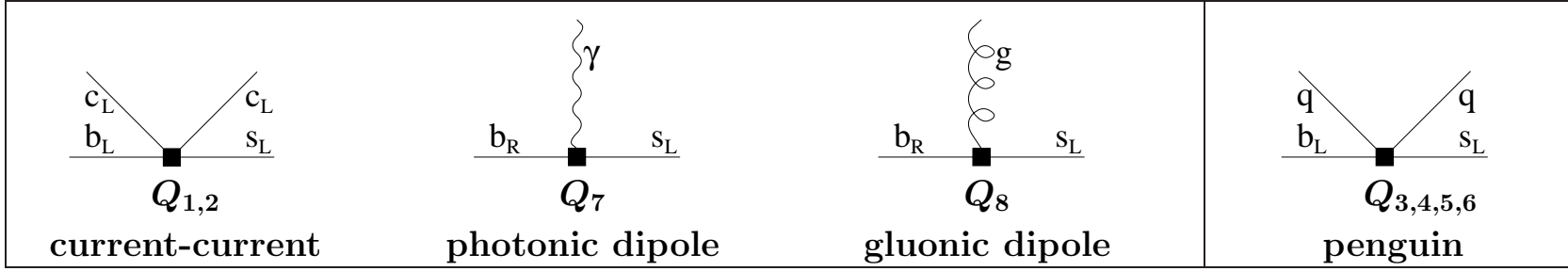
but not too close to the endpoint ($m_b - 2E_0 \gg \Lambda_{\text{QCD}}$).

Conventionally, $E_0 = 1.6 \text{ GeV} \simeq m_b/3$ is chosen.

Decoupling of $W, Z, t, H^0 \Rightarrow$ effective weak interaction Lagrangian:

$$L_{\text{weak}} \sim \sum_i C_i Q_i$$

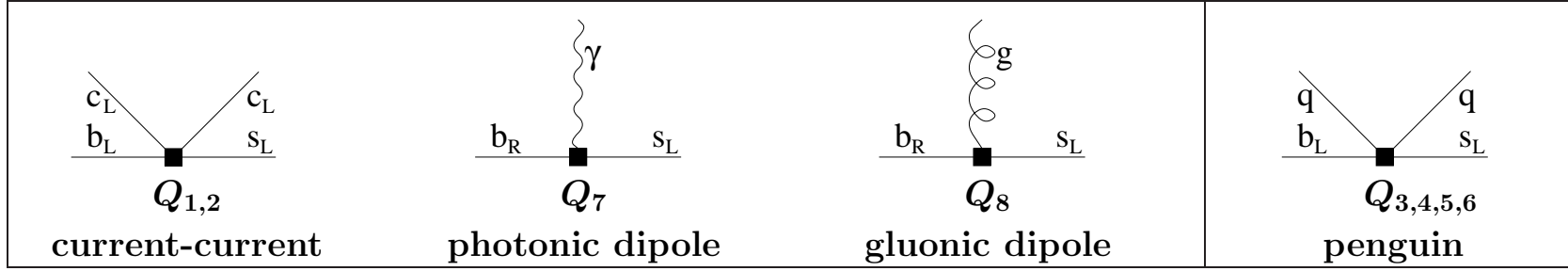
Eight operators Q_i matter for $\mathcal{B}_{s\gamma}^{\text{SM}}$ when the NLO EW and/or CKM-suppressed effects are neglected:



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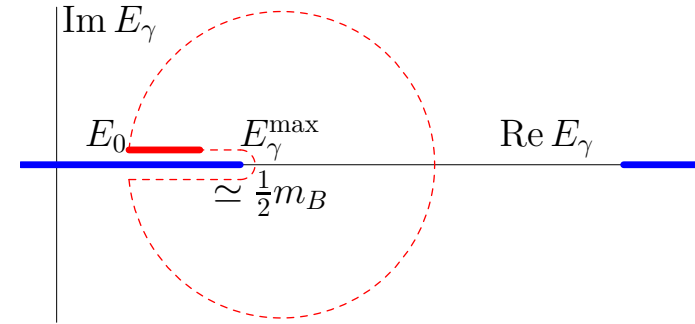


$$\Gamma(\bar{B} \rightarrow X_s \gamma)_{E_\gamma > E_0} = |C_7(\mu_b)|^2 \Gamma_{77}(E_0) + (\text{other}) \quad (\mu_b \sim m_b/2)$$

Optical theorem:

$$\frac{d\Gamma_{77}}{dE_\gamma} \sim \text{Im} \left\{ \bar{B} \begin{array}{c} \gamma \\ \nearrow \quad \searrow \\ q \quad q \end{array} \begin{array}{c} \gamma \\ \nearrow \quad \searrow \\ q \quad q \end{array} \bar{B} \right\} \equiv \text{Im} A$$

Integrating the amplitude A over E_γ :



OPE on the ring \Rightarrow Non-perturbative corrections to $\Gamma_{77}(E_0)$ form a series in $\frac{\Lambda_{\text{QCD}}}{m_b}$ and α_s that begins with

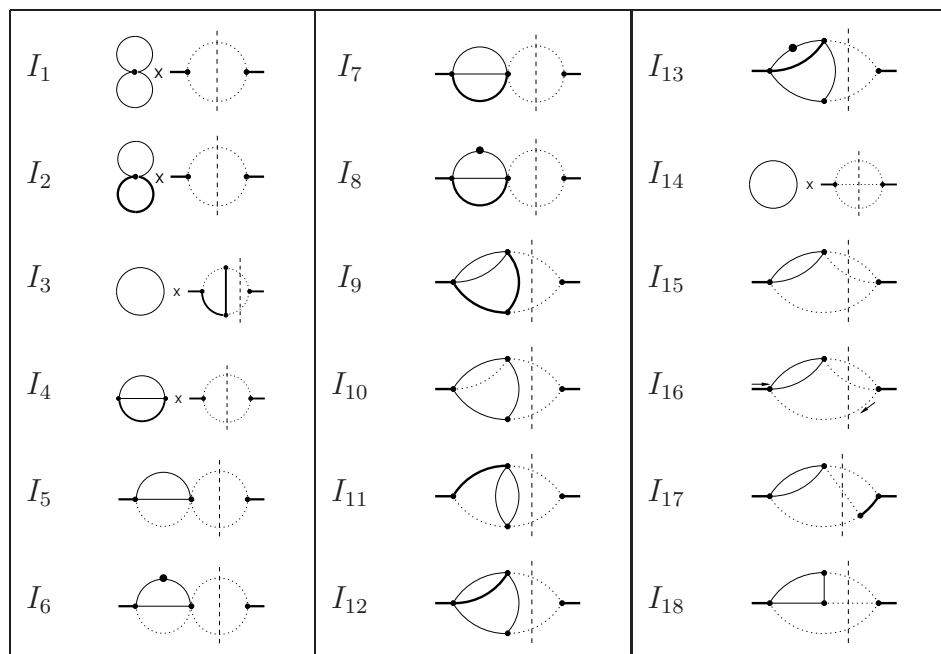
$$\frac{\mu_\pi^2}{m_b^2}, \frac{\mu_G^2}{m_b^2}, \frac{\rho_D^3}{m_b^3}, \frac{\rho_{LS}^3}{m_b^3}, \dots; \frac{\alpha_s \mu_\pi^2}{(m_b - 2E_0)^2}, \frac{\alpha_s \mu_G^2}{m_b(m_b - 2E_0)}; \dots,$$

where $\mu_\pi, \mu_G, \rho_D, \rho_{LS} = \mathcal{O}(\Lambda_{\text{QCD}})$ are extracted from the semileptonic $\bar{B} \rightarrow X_c e \bar{\nu}$ spectra and the $B-B^*$ mass difference.

The same method has been applied to the 3-loop counterterm diagrams

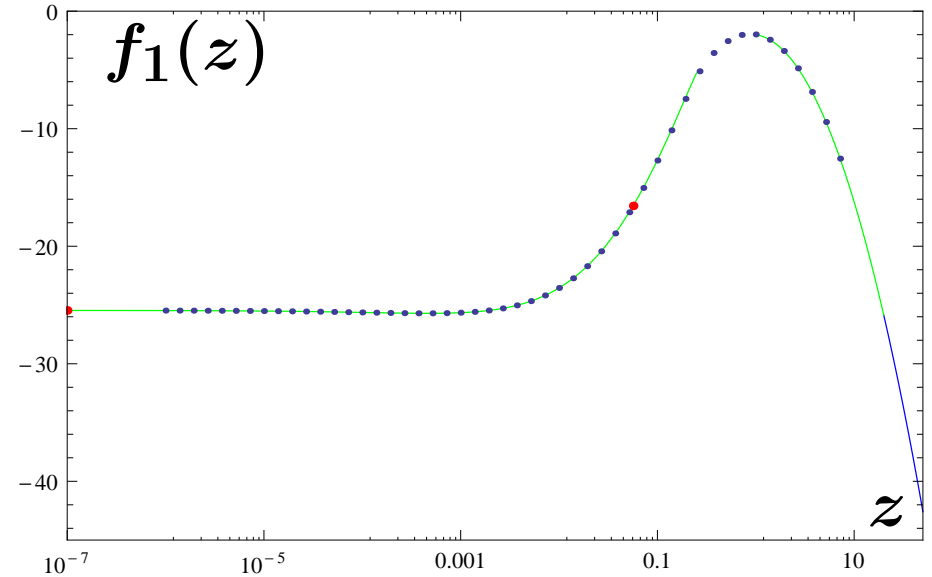
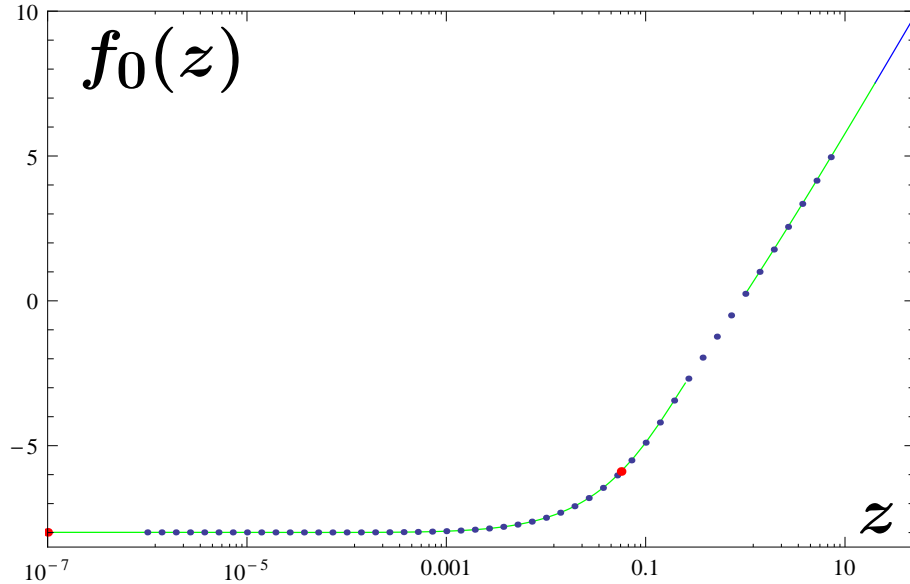
[MM, A. Rehman, M. Steinhauser, PLB 770 (2017) 431]

Master integrals:



Results for the bare NLO contributions up to $\mathcal{O}(\epsilon)$:

$$\hat{G}_{27}^{(1)2P} = -\frac{92}{81\epsilon} + f_0(z) + \epsilon f_1(z) \xrightarrow{z \rightarrow 0} -\frac{92}{81\epsilon} - \frac{1942}{243} + \epsilon \left(-\frac{26231}{729} + \frac{259}{243}\pi^2 \right)$$



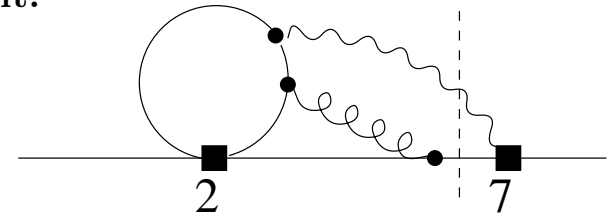
Dots: solutions to the differential equations and/or the exact $z \rightarrow 0$ limit.

Lines: large- and small- z asymptotic expansions

Small- z expansions of $\hat{G}_{27}^{(1)2P}$:

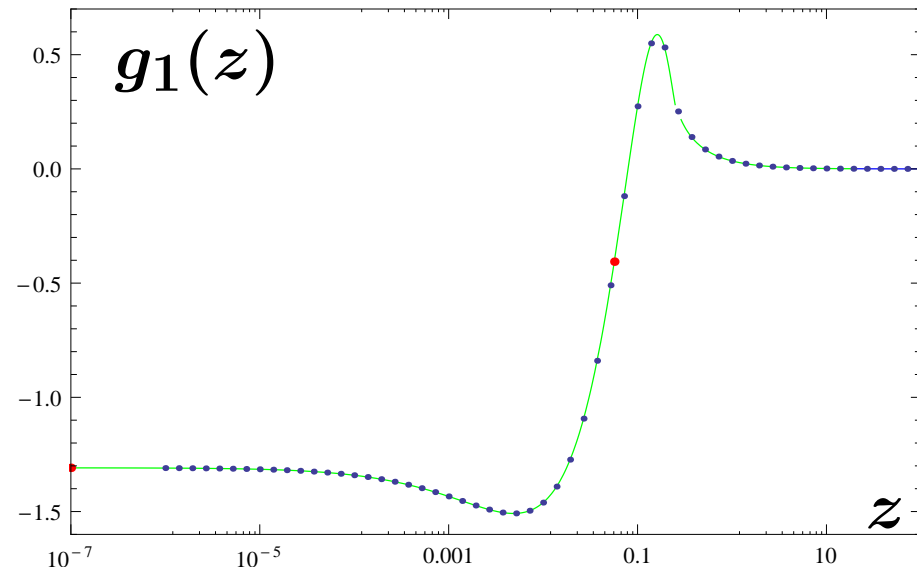
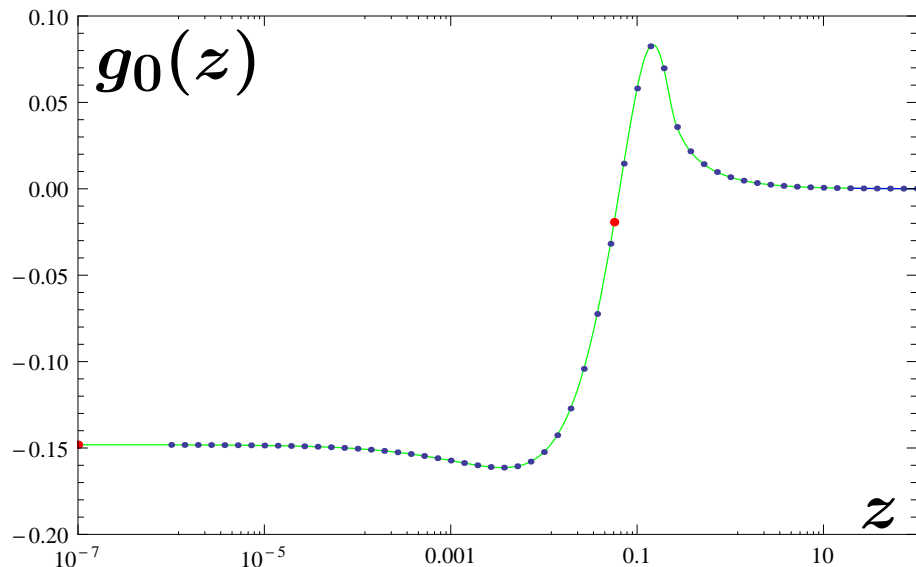
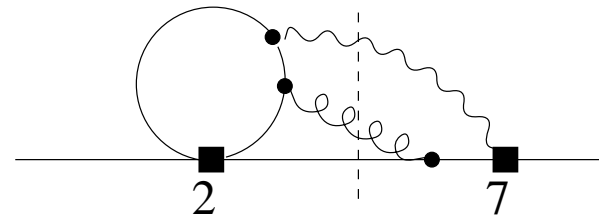
f_0 from C. Greub, T. Hurth, D. Wyler, hep-ph/9602281, hep-ph/9603404,
A. J. Buras, A. Czarnecki, MM, J. Urban, hep-ph/0105160,

f_1 from H.M. Asatrian, C. Greub, A. Hovhannisyanyan, T. Hurth and V. Poghosyan, hep-ph/0505068.



Analogous results for the 3-body final state contributions ($\delta = 1$):

$$\hat{G}_{27}^{(1)3P} = g_0(z) + \epsilon g_1(z) \xrightarrow{z \rightarrow 0} -\frac{4}{27} - \frac{106}{81}\epsilon$$



Dots: solutions to the differential equations and/or the exact $z \rightarrow 0$ limit.

Lines: exact result for g_0 , as well as large- and small- z asymptotic expansions for g_1 .

$$g_0(z) = \begin{cases} -\frac{4}{27} - \frac{14}{9}z + \frac{8}{3}z^2 + \frac{8}{3}z(1-2z) s L + \frac{16}{9}z(6z^2 - 4z + 1) \left(\frac{\pi^2}{4} - L^2\right), & \text{for } z \leq \frac{1}{4}, \\ -\frac{4}{27} - \frac{14}{9}z + \frac{8}{3}z^2 + \frac{8}{3}z(1-2z) t A + \frac{16}{9}z(6z^2 - 4z + 1) A^2, & \text{for } z > \frac{1}{4}, \end{cases}$$

where $s = \sqrt{1-4z}$, $L = \ln(1+s) - \frac{1}{2} \ln 4z$, $t = \sqrt{4z-1}$, and $A = \arctan(1/t)$.