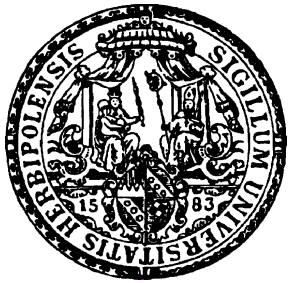
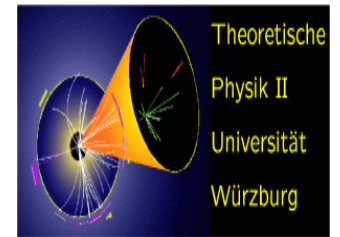


Supersymmetric LFV

Werner Porod



Universität Würzburg



- Lepton flavour violation
- Signals in models with
 - Dirac neutrinos
 - Majorana neutrinos
 - Neutrino masses via R-parity violation
- Conclusions

Neutrinos: tiny masses

$$\Delta m_{atm}^2 \simeq 3 \cdot 10^{-3} \text{ eV}^2$$

$$\Delta m_{sol}^2 \simeq 7 \cdot 10^{-5} \text{ eV}^2$$

$${}^3\text{H decay: } m_\nu \lesssim 2 \text{ eV}$$

Neutrinos: large mixings

$$|\tan \theta_{atm}|^2 \simeq 1$$

$$|\tan \theta_{sol}|^2 \simeq 0.4$$

$$|U_{e3}|^2 \lesssim 0.05$$

strong bounds for charged leptons

$$BR(\mu \rightarrow e\gamma) \lesssim 1.2 \cdot 10^{-11}$$

$$BR(\mu^- \rightarrow e^- e^+ e^-) \lesssim 10^{-12}$$

$$BR(\tau \rightarrow e\gamma) \lesssim 1.1 \cdot 10^{-7}$$

$$BR(\tau \rightarrow \mu\gamma) \lesssim 6.8 \cdot 10^{-8}$$

$$BR(\tau \rightarrow lll') \lesssim O(10^{-8}) \quad (l, l' = e, \mu)$$

$$|d_e| \lesssim 10^{-27} \text{ e cm}, \quad |d_\mu| \lesssim 1.5 \cdot 10^{-18} \text{ e cm}, \quad |d_\tau| \lesssim 1.5 \cdot 10^{-16} \text{ e cm}$$

SUSY contributions to anomalous magnetic moments

$$|\Delta a_e| \leq 10^{-12}, \quad 0 \leq \Delta a_\mu \leq 43 \cdot 10^{-10}, \quad |\Delta a_\tau| \leq 0.058$$

analog to leptons or quarks

$$Y_\nu H \bar{\nu}_L \nu_R \rightarrow Y_\nu v \bar{\nu}_L \nu_R = m_\nu \bar{\nu}_L \nu_R$$

requires $Y_\nu \ll Y_e$

⇒ no impact for future collider experiments

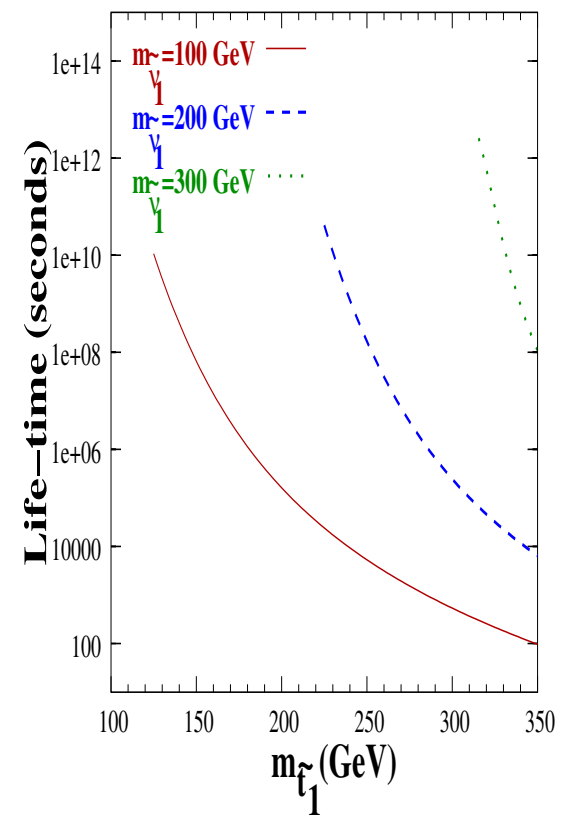
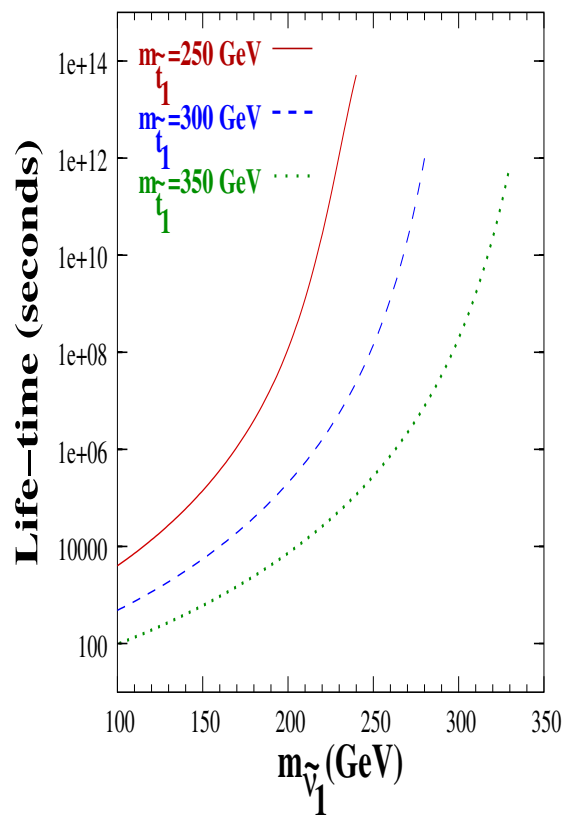
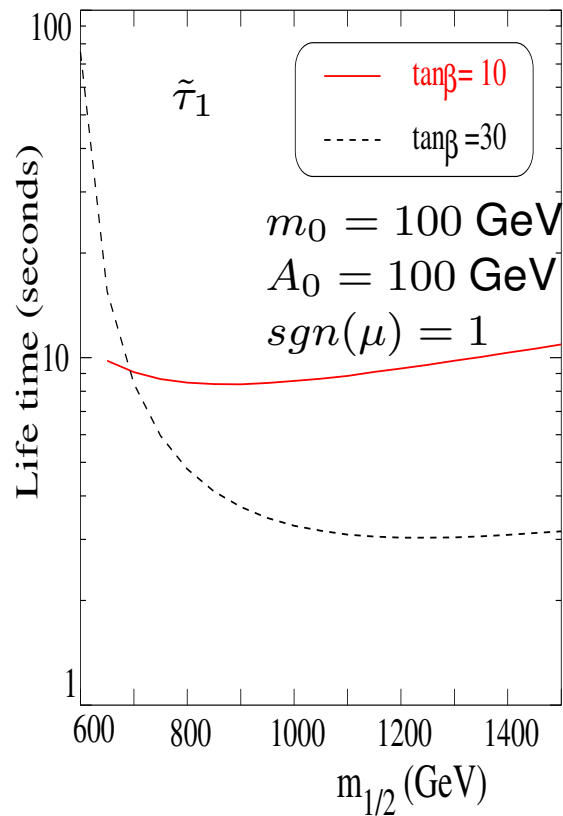
Exception: $\tilde{\nu}_R$ is LSP and thus a candidate for dark matter

⇒ long lived NLSP, e.g. $\tilde{t}_1 \rightarrow l^+ b \tilde{\nu}_R$

$$\frac{dm_{\tilde{\nu}_R}^2}{dt} = \frac{1}{8\pi^2} (Y_\nu A_\nu)^2$$

⇒ e.g. $m_{\tilde{\nu}_R} \simeq m_0$ in mSUGRA, $m_{\tilde{\nu}_R} \simeq 0$ in GMSB

S. Gopalakrishna, A. de Gouvea and W. P., JHEP **0611** (2006) 050



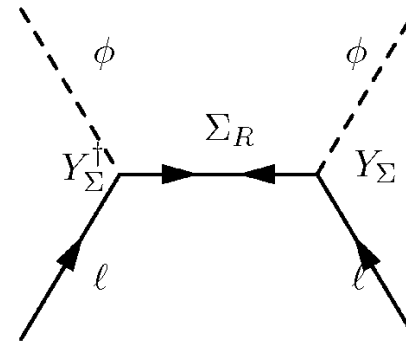
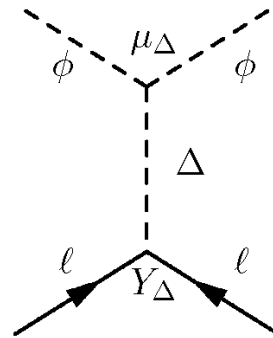
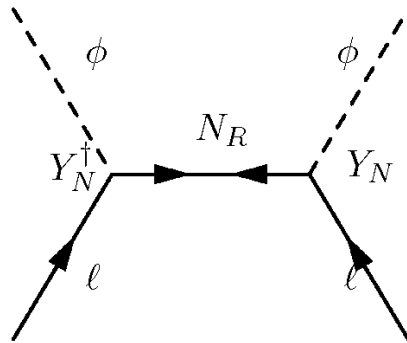
general signature: two long lived particles + multi jets and/or multi lepton

S. K. Gupta, B. Mukhopadhyaya, S. K. Rai, PRD 75 (2007) 075007

D. Choudhury, S. K. Gupta, B. Mukhopadhyaya, PRD 78 (2008) 015023

Neutrino masses due to

$$\frac{f}{\Lambda}(HL)(HL)$$



* P. Minkowski, *Phys. Lett. B* **67** (1977) 421; T. Yanagida, KEK-report 79-18 (1979);
M. Gell-Mann, P. Ramond, R. Slansky, in *Supergravity*, North Holland (1979), p. 315;
R.N. Mohapatra and G. Senjanovic, *Phys. Rev. Lett.* **44** 912 (1980); M. Magg and C. Wetterich,
Phys. Lett. B **94** (1980) 61; G. Lazarides, Q. Shafi and C. Wetterich, *Nucl. Phys. B* **181**
(1981) 287; J. Schechter and J. W. F. Valle, *Phys. Rev. D* **25**, 774 (1982);
R. Foot, H. Lew, X. G. He and G. C. Joshi, *Z. Phys. C* **44** (1989) 441.

Supersymmetry

$$W = Y_e^{ji} \hat{L}_i \hat{H}_d \hat{E}_j^c + Y_\nu^{ji} \hat{L}_i \hat{H}_u \hat{N}_j^c + M_{R_i} \hat{N}_i^c \hat{N}_i^c$$

neutrino masses

$$m_\nu \simeq -(Y_\nu^T v) M_R^{-1} (Y_\nu v) \quad \Rightarrow \quad \hat{m}_\nu = U^T \cdot m_\nu \cdot U$$

convenient parameterization[†]:

$$Y_\nu = \sqrt{2} \frac{i}{v_U} \sqrt{\hat{M}_R} \cdot R \cdot \sqrt{\hat{m}_\nu} \cdot U^\dagger$$

RGE running

$$(\Delta M_{\tilde{L}}^2)_{ij} = -\frac{1}{8\pi^2} (3m_0^2 + A_0^2) (Y_\nu^\dagger L Y_\nu)_{ij}$$

$$(\Delta A_l)_{ij} = -\frac{3}{8\pi^2} A_0 Y_{l_i} (Y_\nu^\dagger L Y_\nu)_{ij}$$

$$(\Delta M_{\tilde{E}}^2)_{ij} = 0$$

$$L_{kl} = \log\left(\frac{M_X}{M_k}\right) \delta_{kl}$$

[†]J. A. Casas and A. Ibarra, Nucl. Phys. **B618**, 171 (2001), [hep-ph/0103065].

$(\Delta M_{\tilde{L}}^2)_{ij}$ and $(\Delta A_l)_{ij}$ induce

$$\begin{aligned} l_j &\rightarrow l_i \gamma, l_i l_k^+ l_r^- \\ \tilde{l}_j &\rightarrow l_i \tilde{\chi}_s^0 \\ \tilde{\chi}_s^0 &\rightarrow l_i \tilde{l}_k \end{aligned}$$

Neglecting L - R mixing:

$$\begin{aligned} Br(l_i \rightarrow l_j \gamma) &\propto \alpha^3 m_{l_i}^5 \frac{|(\delta m_L^2)_{ij}|^2}{\tilde{m}^8} \tan^2 \beta \\ \frac{Br(\tilde{\tau}_2 \rightarrow e + \chi_1^0)}{Br(\tilde{\tau}_2 \rightarrow \mu + \chi_1^0)} &\simeq \left(\frac{(\Delta M_{\tilde{L}}^2)_{13}}{(\Delta M_{\tilde{L}}^2)_{23}} \right)^2 \end{aligned}$$

Moreover, in most of the parameter space

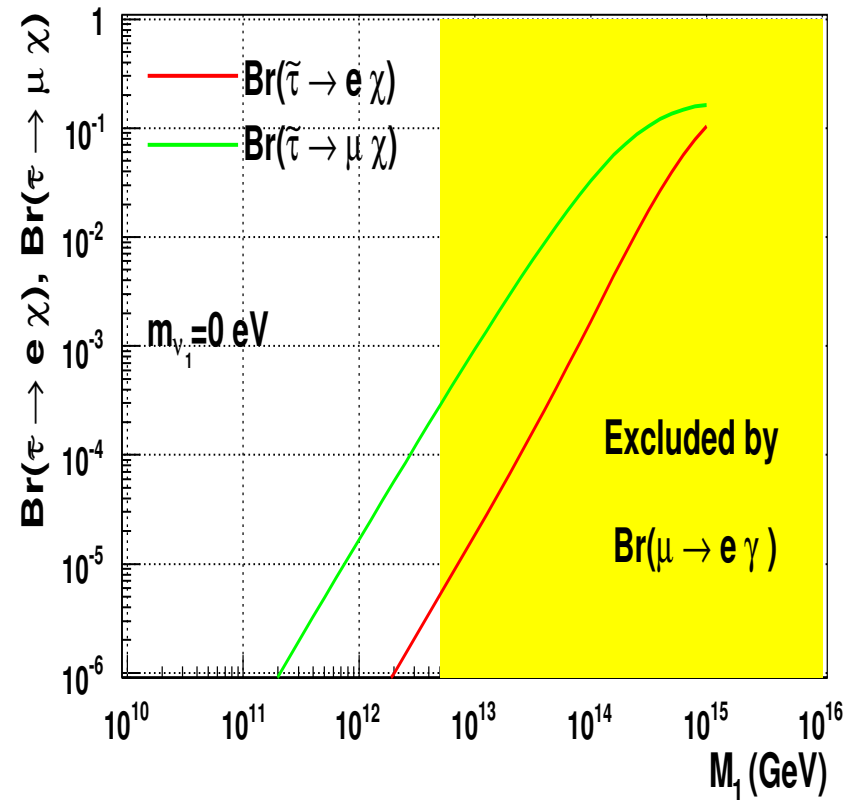
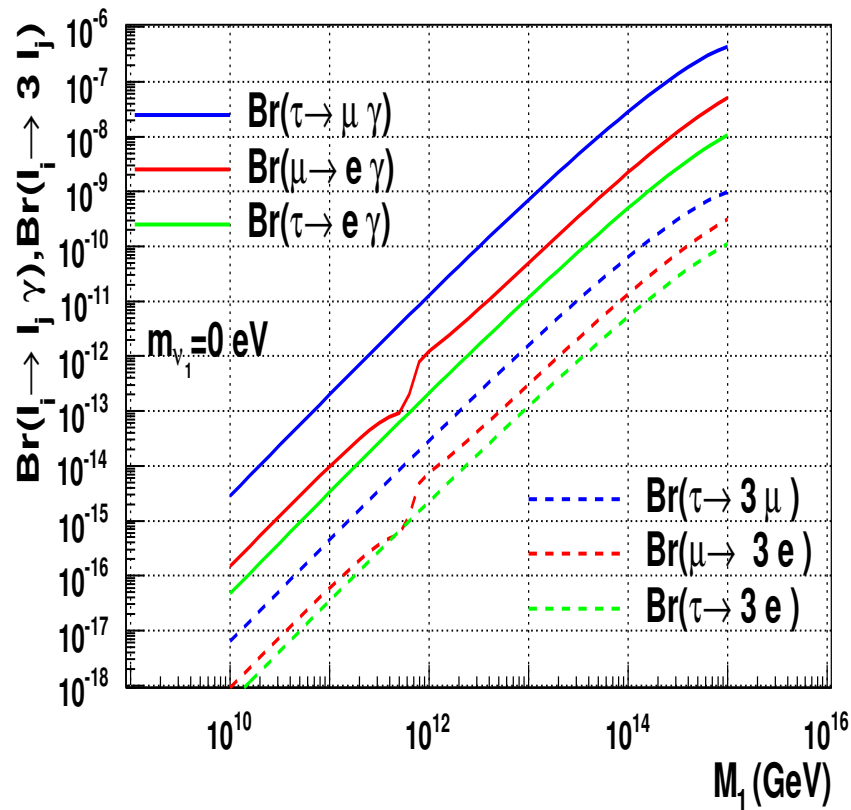
$$\frac{Br(l_i \rightarrow 3l_j)}{Br(l_i \rightarrow l_j + \gamma)} \simeq \frac{\alpha}{3\pi} \left(\log\left(\frac{m_{l_i}^2}{m_{l_j}^2}\right) - \frac{11}{4} \right)$$

take all parameters real

$$U = U_{\text{TBM}} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix}$$

$$R = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

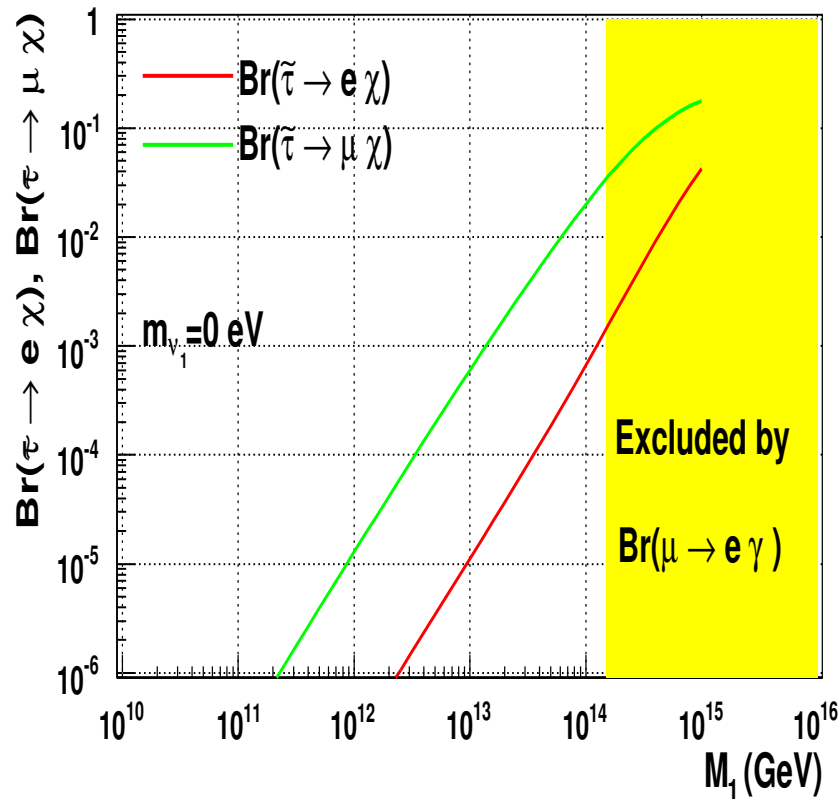
Use 2-loop RGEs and 1-loop corrections including flavour effects



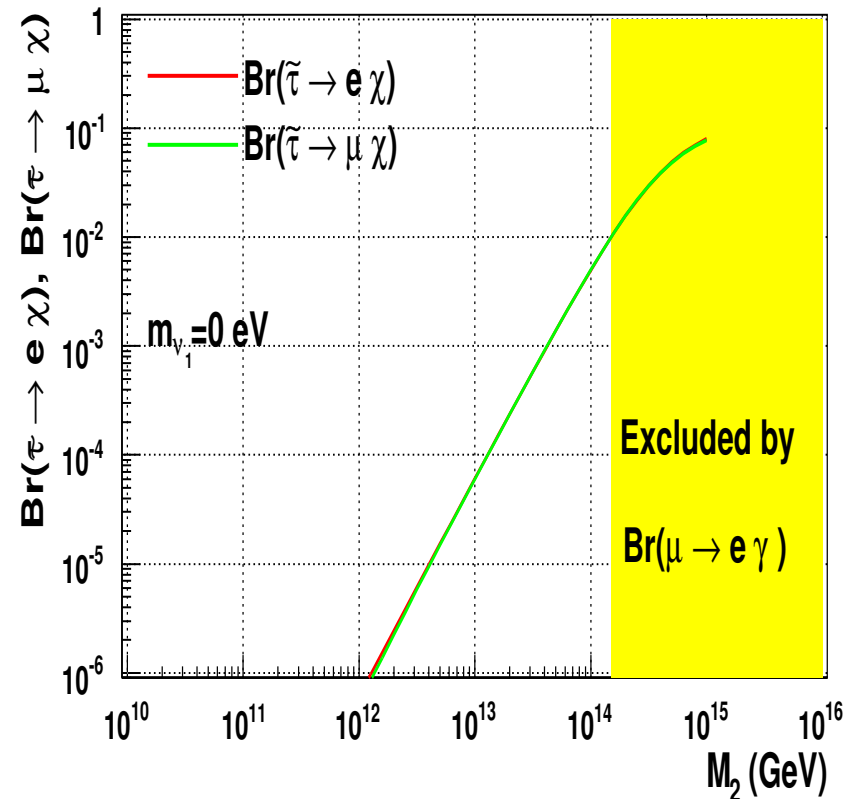
degenerate ν_R

SPS1a' ($M_0 = 70 \text{ GeV}$, $M_{1/2} = 250 \text{ GeV}$, $A_0 = -300 \text{ GeV}$, $\tan \beta = 10$, $\mu > 0$)

M. Hirsch et al. Phys. Rev. D 78 (2008) 013006



degenerate ν_R



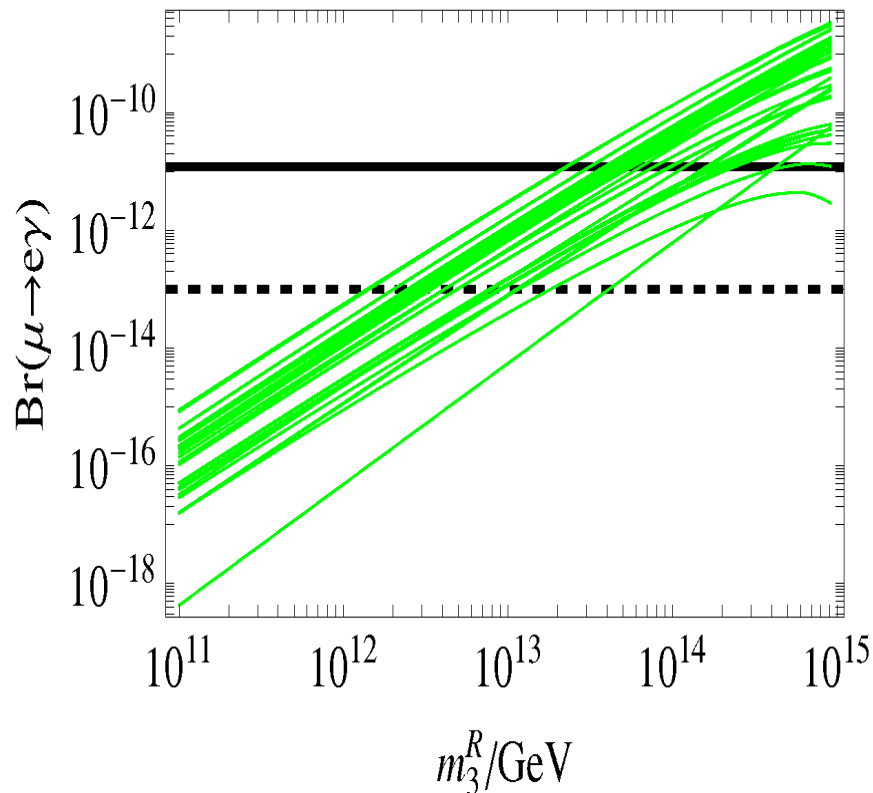
hierarchical ν_R

($M_1 = M_3 = 10^{10} \text{ GeV}$)

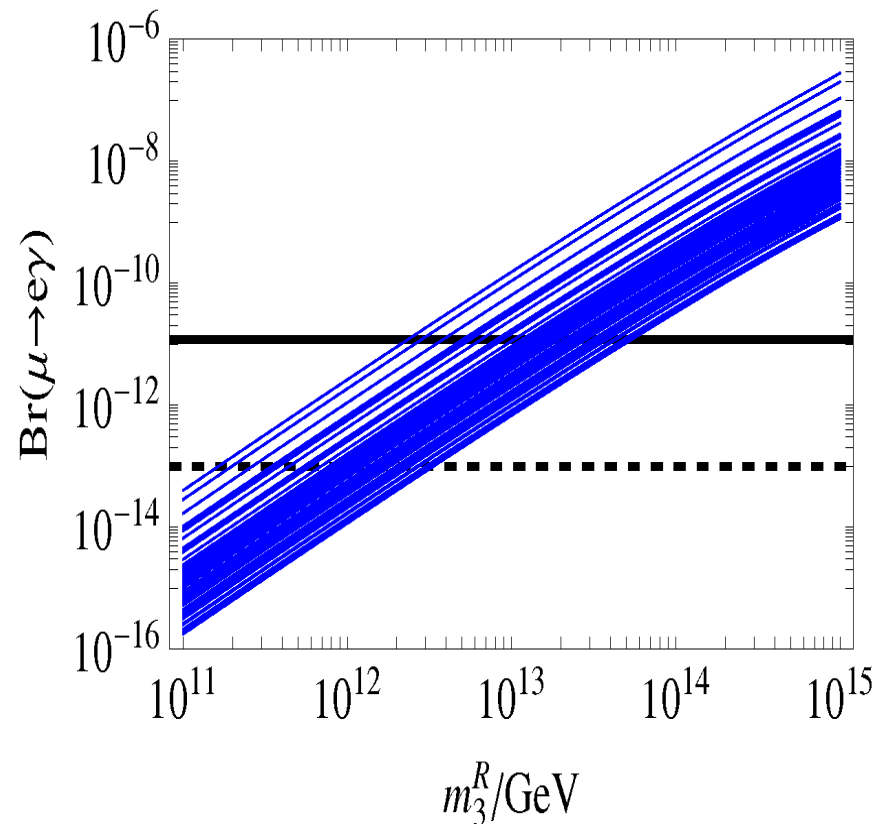
SPS3 ($M_0 = 90 \text{ GeV}, M_{1/2} = 400 \text{ GeV}, A_0 = 0 \text{ GeV}, \tan \beta = 10, \mu > 0$)

M. Hirsch et al. Phys. Rev. D 78 (2008) 013006

Texture models, hierarchical ν_R
real textures



"complexification" of one texture



SPS1a' ($M_0 = 70$ GeV, $M_{1/2} = 250$ GeV, $A_0 = -300$ GeV, $\tan \beta = 10$, $\mu > 0$)

F. Deppisch, F. Plentinger, W. P., R. Rückl, G. Seidl, in preparation

include $SU(2)$ Triplet Higgs

$$W = W_{\text{MSSM}} + \frac{1}{\sqrt{2}} \left(Y_T^{ij} L_i T_1 L_j + \lambda_1 H_1 T_1 H_1 + \lambda_2 H_2 T_2 H_2 \right) + M_T T_1 T_2$$

$$m_\nu = \frac{v_2^2}{2} \frac{\lambda_2}{M_T} Y_T$$

$$\frac{M_T}{\lambda_2} \simeq 10^{15} \text{ GeV} \left(\frac{0.05 \text{ eV}}{m_\nu} \right)$$

Gauge coupling unification \Rightarrow use **15**

$$\mathbf{15} = S + T + Z$$

$$S \sim \left(6, 1, -\frac{2}{3} \right), \quad T \sim (1, 3, 1), \quad Z \sim \left(3, 2, \frac{1}{6} \right)$$

$$W \subset \frac{1}{\sqrt{2}} (Y_T L T_1 L + Y_S d^c S d^c) + Y_Z d^c Z L + Y_d d^c Q H_1 + Y_u u^c Q H_2 + Y_e e^c L H_1$$

$$+ \frac{1}{\sqrt{2}} (\lambda_1 H_1 T_1 H_1 + \lambda_2 H_2 T_2 H_2) + M_T T_1 T_2 + M_Z Z_1 Z_2 + M_S S_1 S_2 + \mu H_1 H_2$$

$$(b_1, b_2, b_3)^{MSSM} = \left(\frac{33}{5}, 1, -3\right)$$

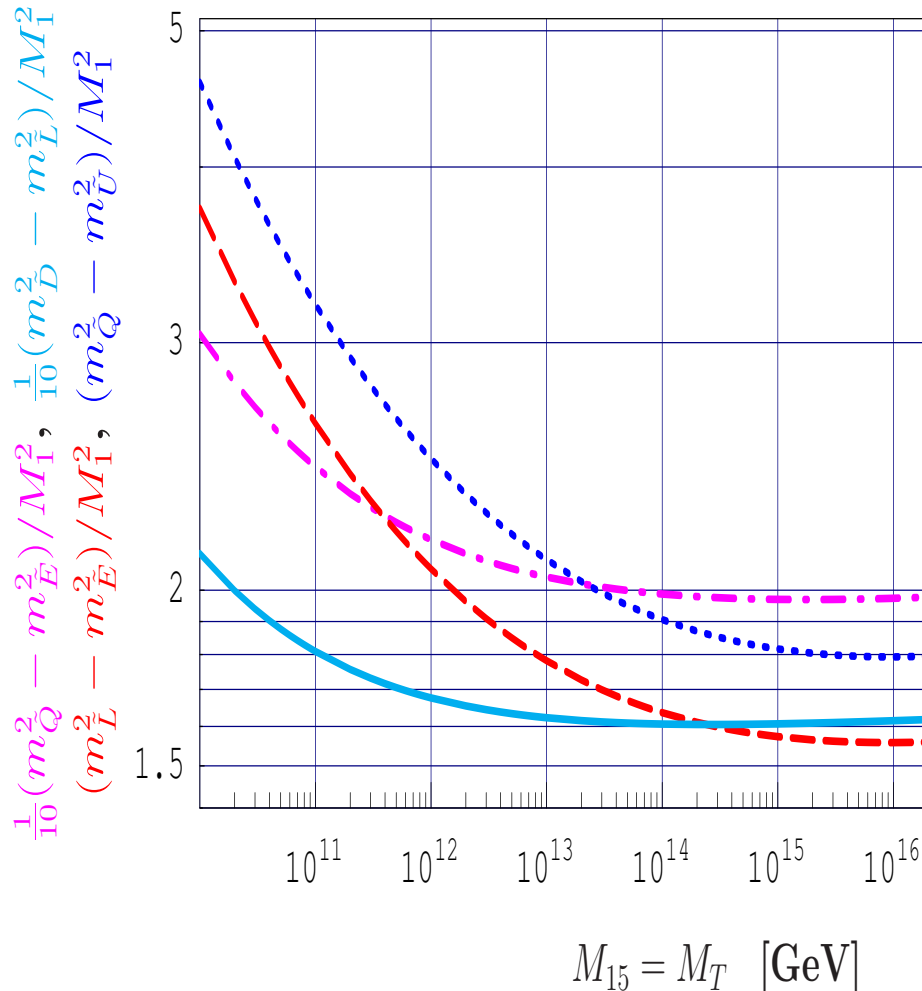
$$(b_1, b_2, b_3)^{T_1+T_2} = \left(\frac{18}{5}, 4, 0\right)$$

$$(b_1, b_2, b_3)^{\overline{15}+15} = (7, 7, 7)$$

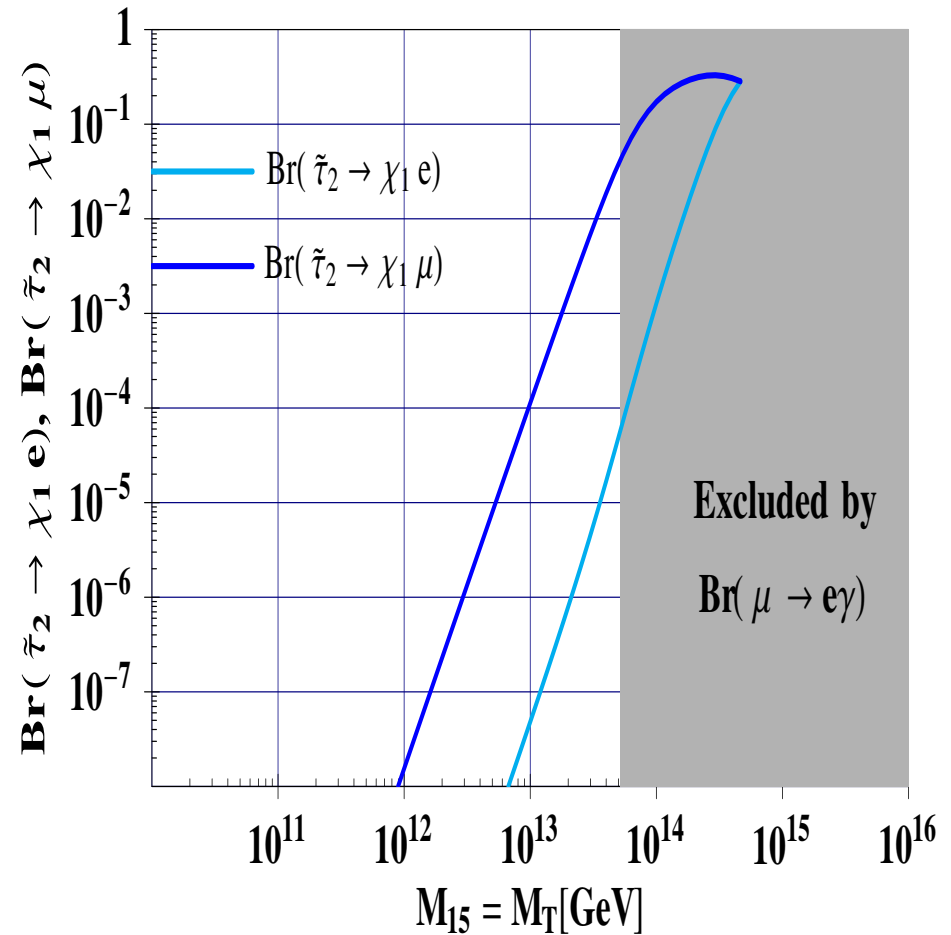
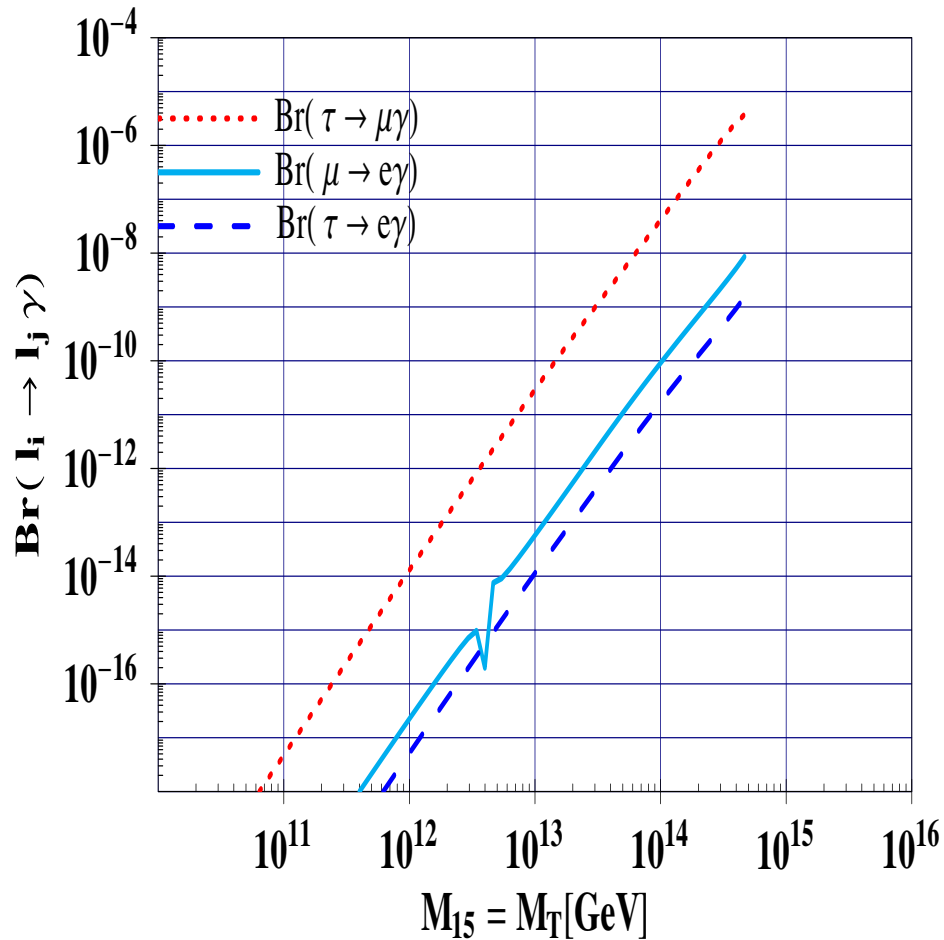
Seesaw I (\simeq MSSM)

$$\frac{m_Q^2 - m_E^2}{M_1^2} \simeq 20, \quad \frac{m_D^2 - m_L^2}{M_1^2} \simeq 18$$

$$\frac{m_L^2 - m_E^2}{M_1^2} \simeq 1.6, \quad \frac{m_Q^2 - m_U^2}{M_1^2} \simeq 1.55$$



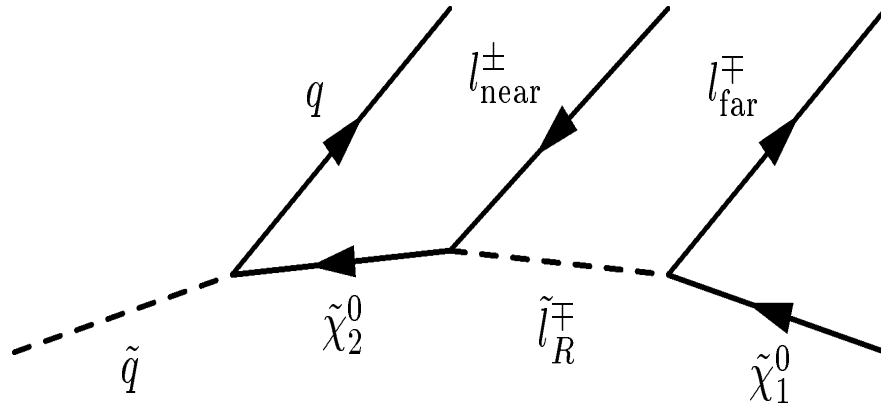
M. R. Buckley, H. Murayama, PRL **97** (2006) 231801; M. Hirsch, S. Kaneko, W. P., PRD **78** (2008) 093004.



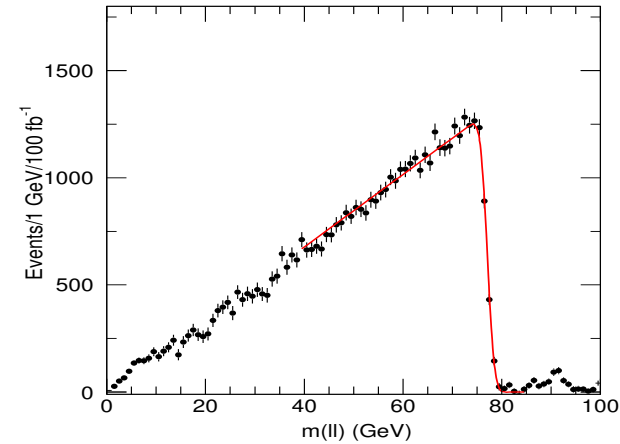
$$\lambda_1 = \lambda_2 = 0.5$$

$$\text{SPS3 } (M_0 = 90 \text{ GeV}, M_{1/2} = 400 \text{ GeV}, A_0 = 0 \text{ GeV}, \tan \beta = 10, \mu > 0)$$

M. Hirsch, S. Kaneko, W. P., Phys. Rev. D 78 (2008) 093004.



G. Polesello

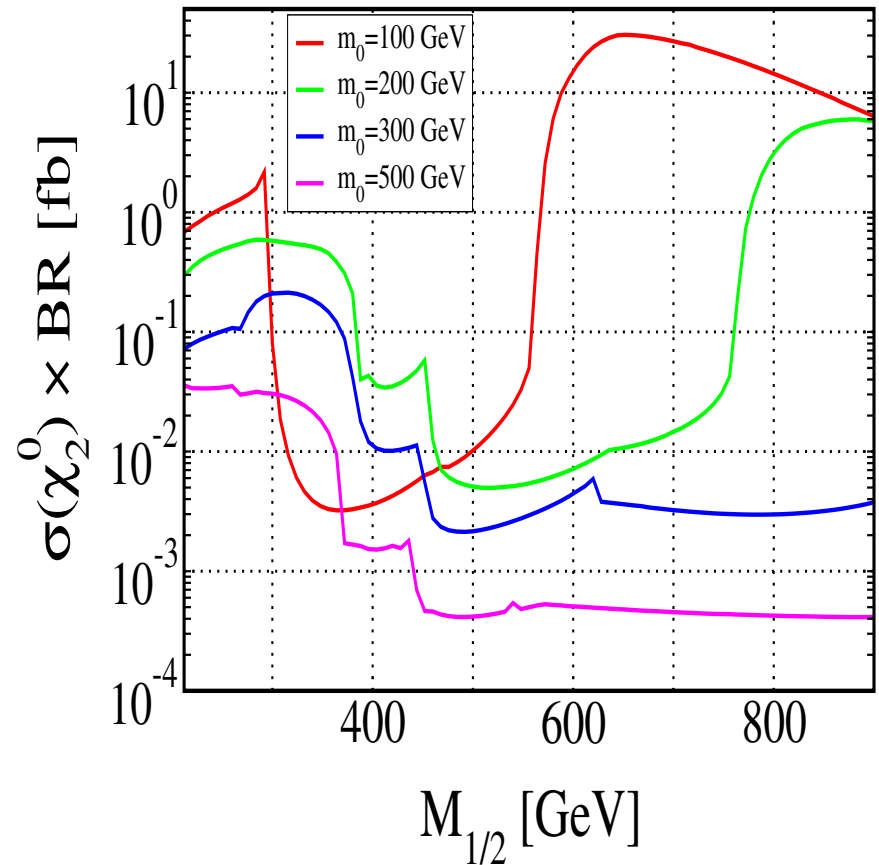
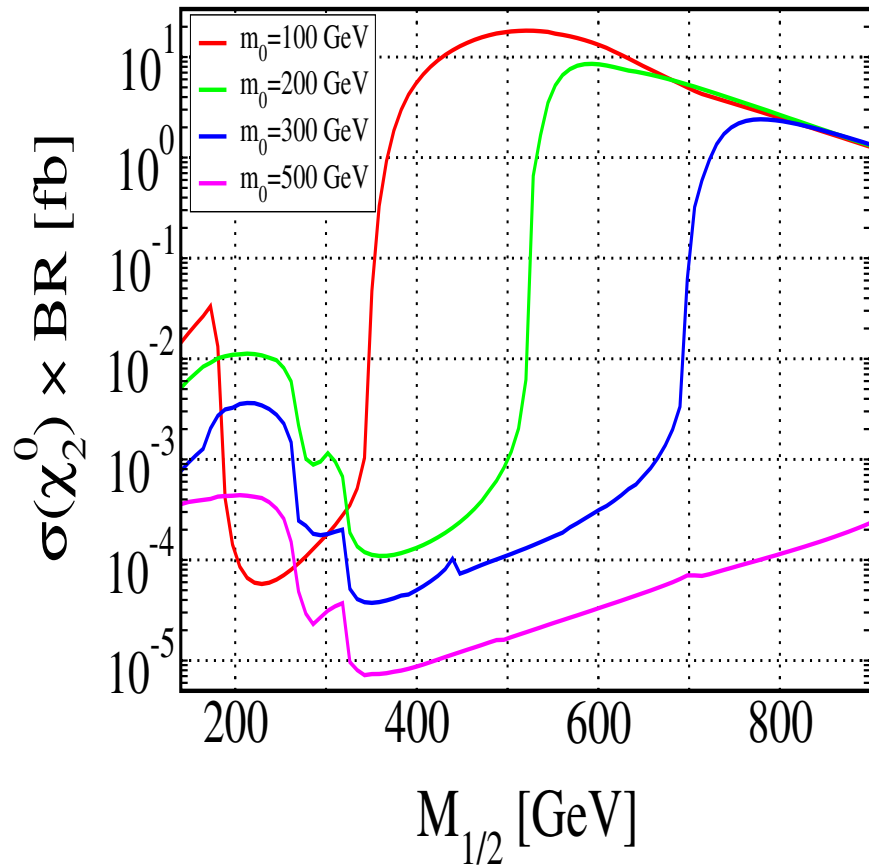


5 kinematical observables depending on 4 SUSY masses

e.g.: $m(ll) = 77.02 \pm 0.05 \pm 0.08$
 \Rightarrow mass determination within 2-5%

For background suppression

$$N(e^+e^-) + N(\mu^+\mu^-) - N(e^+\mu^-) - N(\mu^+e^-)$$

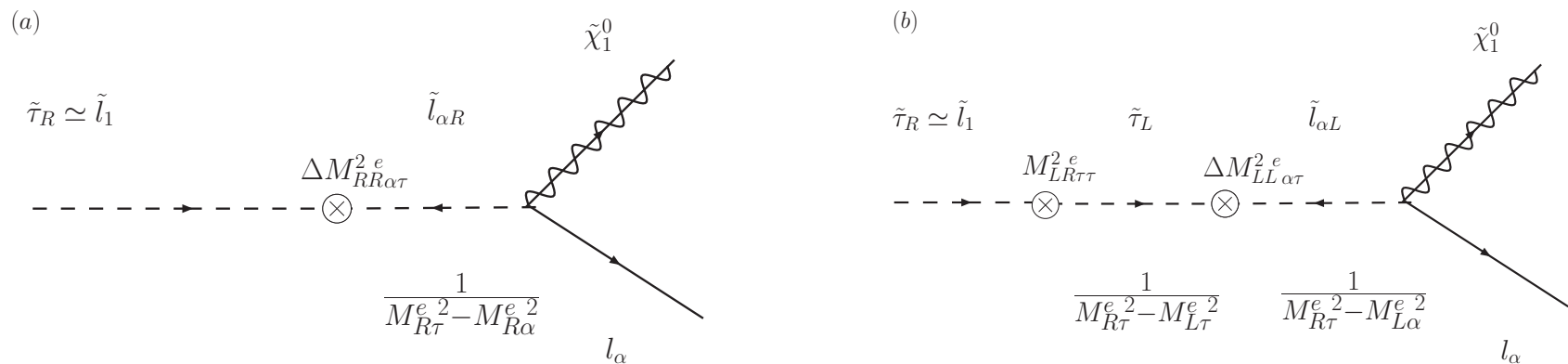
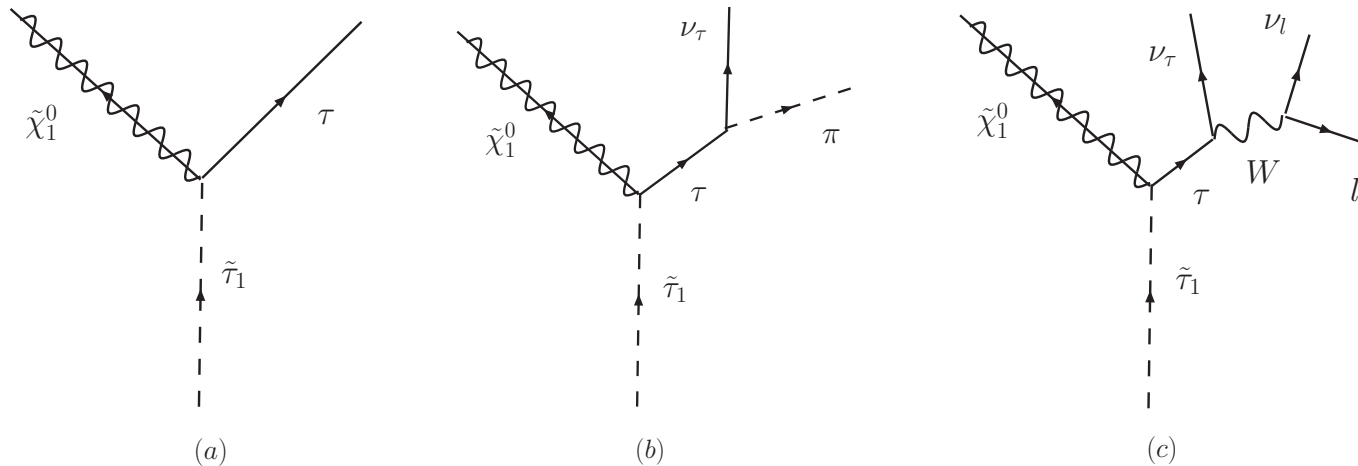


$$\sigma(pp \rightarrow \tilde{\chi}_2^0) \times BR(\chi_2^0 \rightarrow \sum_{i,j} \tilde{l}_i l_j \rightarrow \mu^\pm \tau^\mp \tilde{\chi}_1^0)$$

$$A_0 = 0, \tan \beta = 10, \mu > 0 \text{ (Seesaw II: } \lambda_1 = 0.02, \lambda_2 = 0.5)$$

J.N. Esteves et al., arXiv:0903.1408

mSugra: stau co-annihilation for DM, in particular $m_{\tilde{\tau}_1} - m_{\tilde{\chi}_1^0} \lesssim m_\tau$



$\tilde{\tau}_1$ life times up to 10^4 sec

[†] S. Kaneko et al., arXiv:0811.0703 (hep-ph)

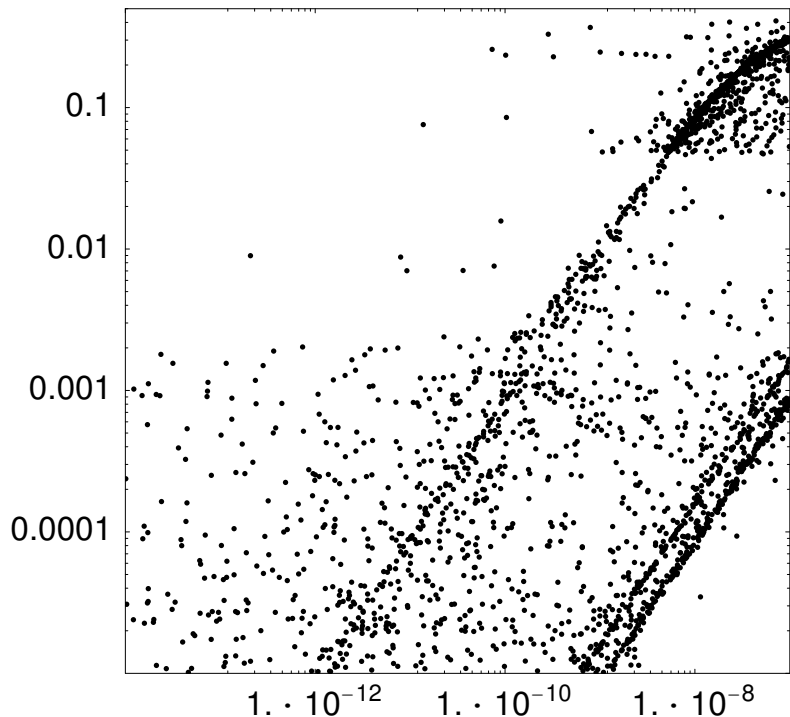
general problem up to now: $m_\nu \simeq 0.1 \text{ eV} \Rightarrow Y^2/M$ fixed
 forbid dim-5 operator, e.g. $Z_3 + \text{NMSSM}^\dagger$

$$\frac{(LH_u)^2 S}{M_6^2}, \quad \frac{(LH_u)^2 S^2}{M_7^3}$$

solves at the same time the μ -problem

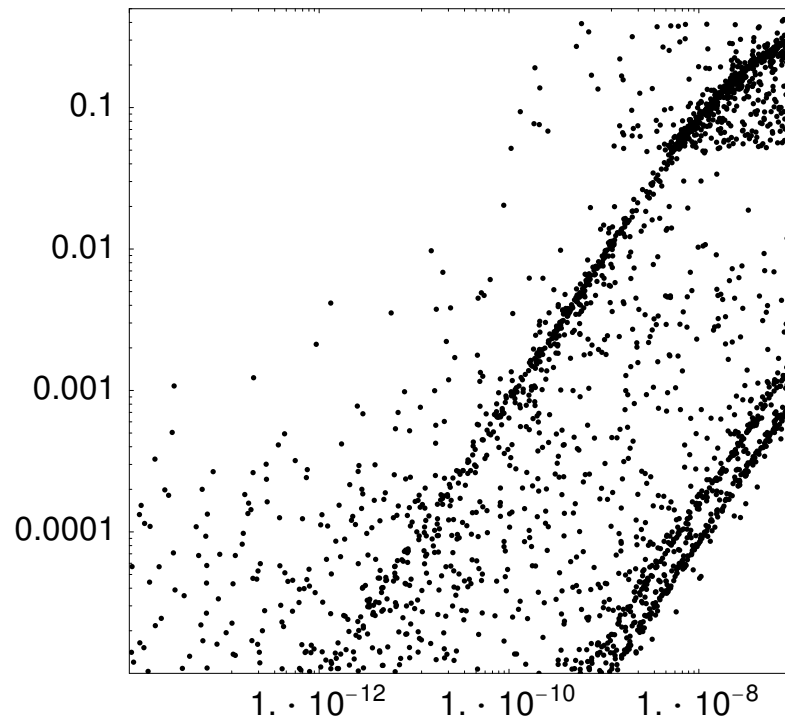
[†] I. Gogoladze, N. Okada, Q. Shafi, arXiv:0809.0703 (hep-ph)

$$\text{BR}(\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 e^\pm \tau^\mp)$$



$$\text{BR}(\tau \rightarrow e\gamma)$$

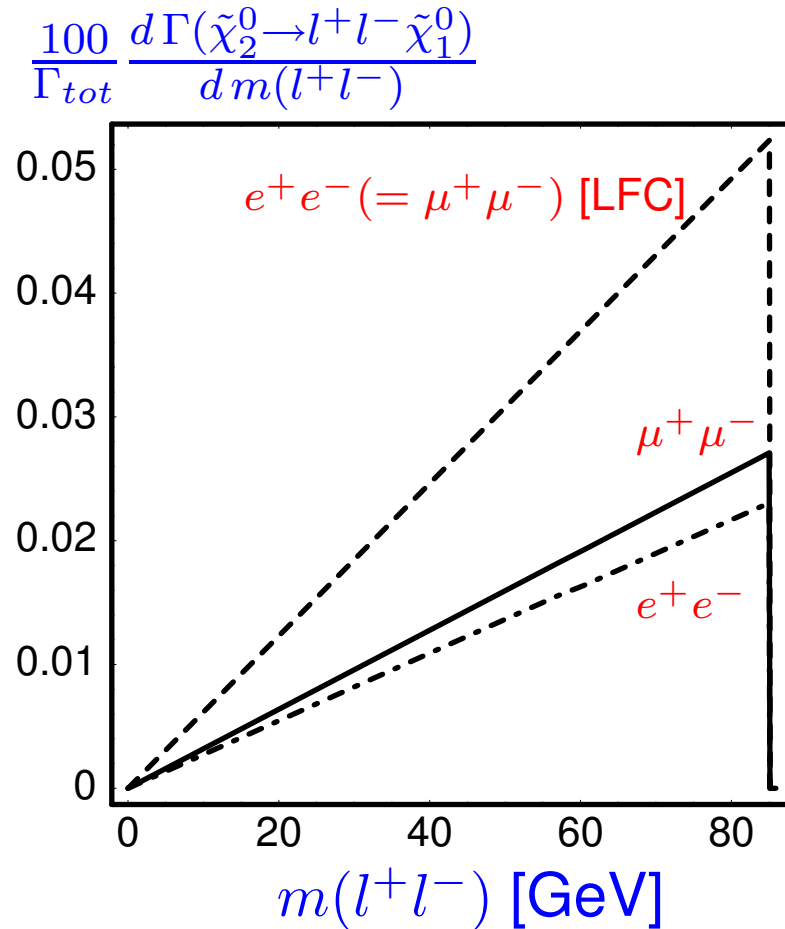
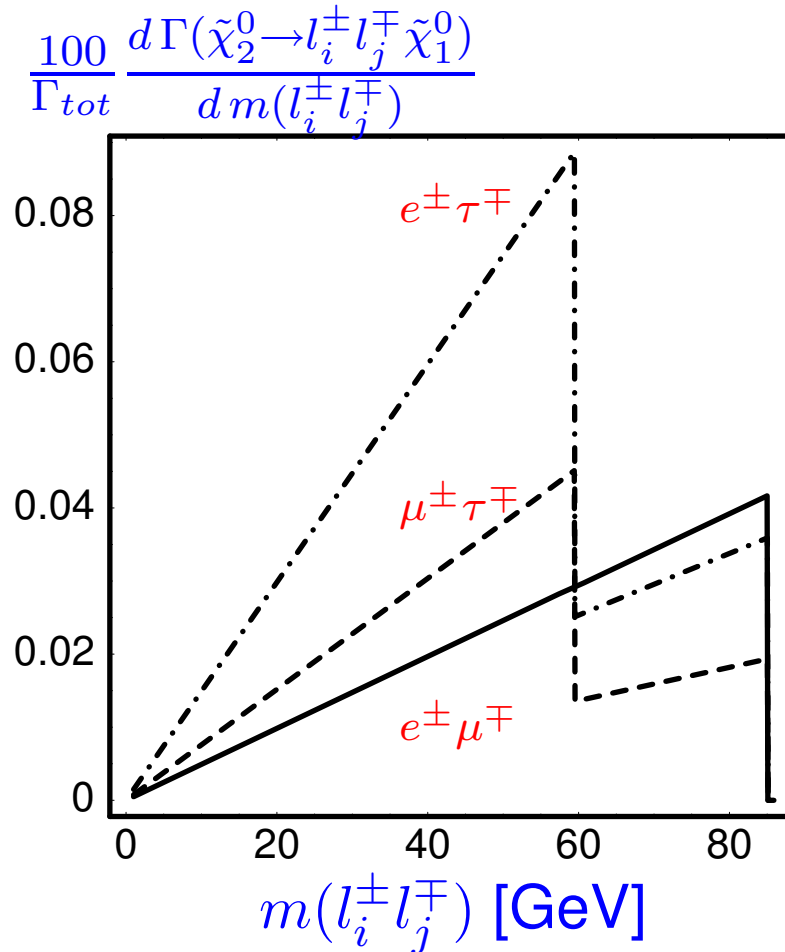
$$\text{BR}(\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 \mu^\pm \tau^\mp)$$



$$\text{BR}(\tau \rightarrow \mu\gamma)$$

Variations around SPS1a

$$(M_0 = 100 \text{ GeV}, M_{1/2} = 250 \text{ GeV}, A_0 = -100 \text{ GeV}, \tan \beta = 10)$$



A. Bartl et al., Eur. Phys. J. C 46 (2006) 783

R-parity: $(-1)^{3(B-L)+2S}$

bilinear R-parity violation:

$$W = W_{MSSM} + \epsilon_i \hat{L}_i \hat{H}_u$$

⇒ mixings between SM and SUSY particles

Gravitino as dark matter*

generic prediction of GMSB : light Gravitino LSP

relic density requires: $100 \text{ eV} \lesssim m_{3/2} \lesssim 1000 \text{ eV}$

NLSP: $\tilde{\chi}_1^0$ oder \tilde{l}_R ($l = e, \mu, \tau$)

*S. Borgani, A. Masiero, M. Yamaguchi, PLB386 (1996) 189

F. Takayama and M. Yamaguchi, PLB 485 (2000) 388

M. Hirsch, W. P., D. Restrepo, JHEP 0503, 062 (2005)

basis $\psi^{0T} = (-i\lambda', -i\lambda^3, \tilde{H}_d^1, \tilde{H}_u^2, \nu_e, \nu_\mu, \nu_\tau)$ we get:

$$M_N = \begin{bmatrix} \mathcal{M}_{\chi^0} & m^T \\ m & 0 \end{bmatrix}$$

$$\mathcal{M}_{\chi^0} = \begin{bmatrix} M_1 & 0 & -\frac{1}{2}g'v_d & \frac{1}{2}g'v_u \\ 0 & M_2 & \frac{1}{2}gv_d & -\frac{1}{2}gv_u \\ -\frac{1}{2}g'v_d & \frac{1}{2}gv_d & 0 & -\mu \\ \frac{1}{2}g'v_u & -\frac{1}{2}gv_u & -\mu & 0 \end{bmatrix}, \quad m = \begin{bmatrix} -\frac{1}{2}g'v_1 & \frac{1}{2}gv_1 & 0 & \epsilon_1 \\ -\frac{1}{2}g'v_2 & \frac{1}{2}gv_2 & 0 & \epsilon_2 \\ -\frac{1}{2}g'v_3 & \frac{1}{2}gv_3 & 0 & \epsilon_3 \end{bmatrix}$$

Approximate diagonalization as in usual seesaw mechanism gives

$$m_{\nu,eff} = \frac{M_1 g^2 + M_2 g'^2}{4 \det(\mathcal{M}_{\chi^0})} \begin{pmatrix} \Lambda_1^2 & \Lambda_1 \Lambda_2 & \Lambda_1 \Lambda_3 \\ \Lambda_1 \Lambda_2 & \Lambda_2^2 & \Lambda_2 \Lambda_3 \\ \Lambda_1 \Lambda_3 & \Lambda_2 \Lambda_3 & \Lambda_3^2 \end{pmatrix}, \quad \Lambda_i = \mu v_i + v_d \epsilon_i$$

second ν mass via loops

$$m_{\nu}^{11p} \simeq \frac{1}{16\pi^2} \left(3h_b^2 \sin(2\theta_{\tilde{b}}) m_b \log \frac{m_{\tilde{b}_2}^2}{m_{\tilde{b}_1}^2} + h_{\tau}^2 \sin(2\theta_{\tilde{\tau}}) m_{\tau} \log \frac{m_{\tilde{\tau}_2}^2}{m_{\tilde{\tau}_1}^2} \right) \frac{(\tilde{\epsilon}_1^2 + \tilde{\epsilon}_2^2)}{\mu^2}$$

$$\tilde{\epsilon}_i = V_{ji}^{\nu} \epsilon_j$$

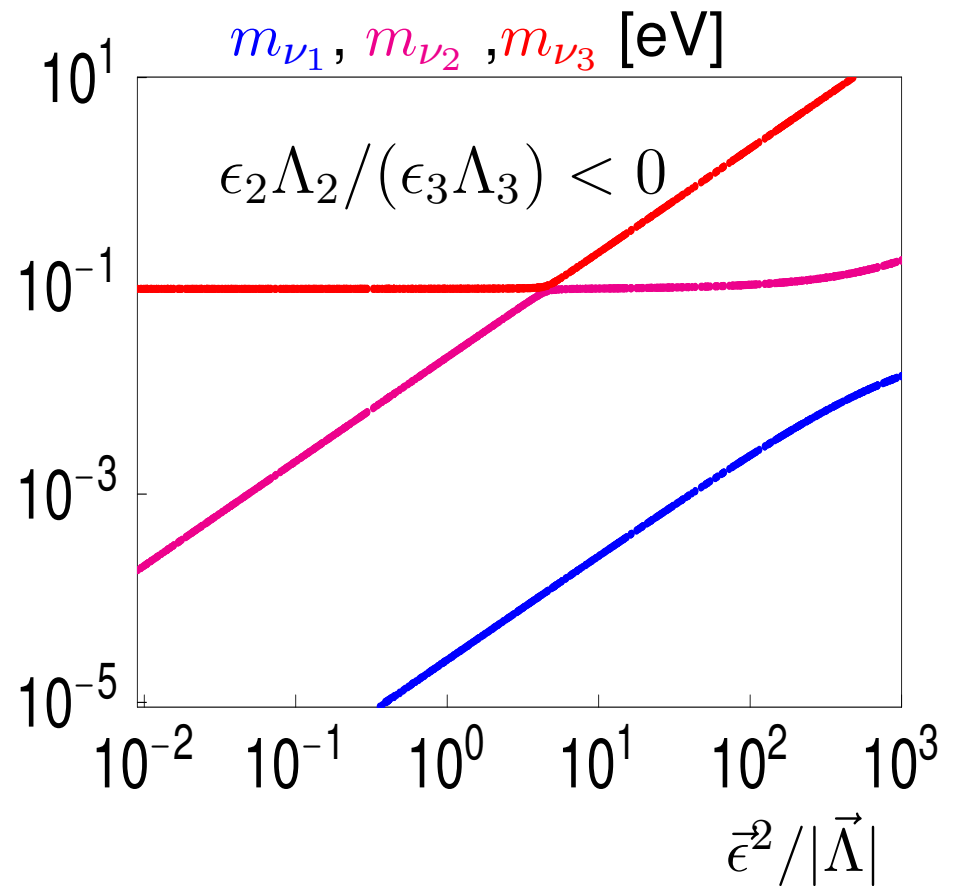
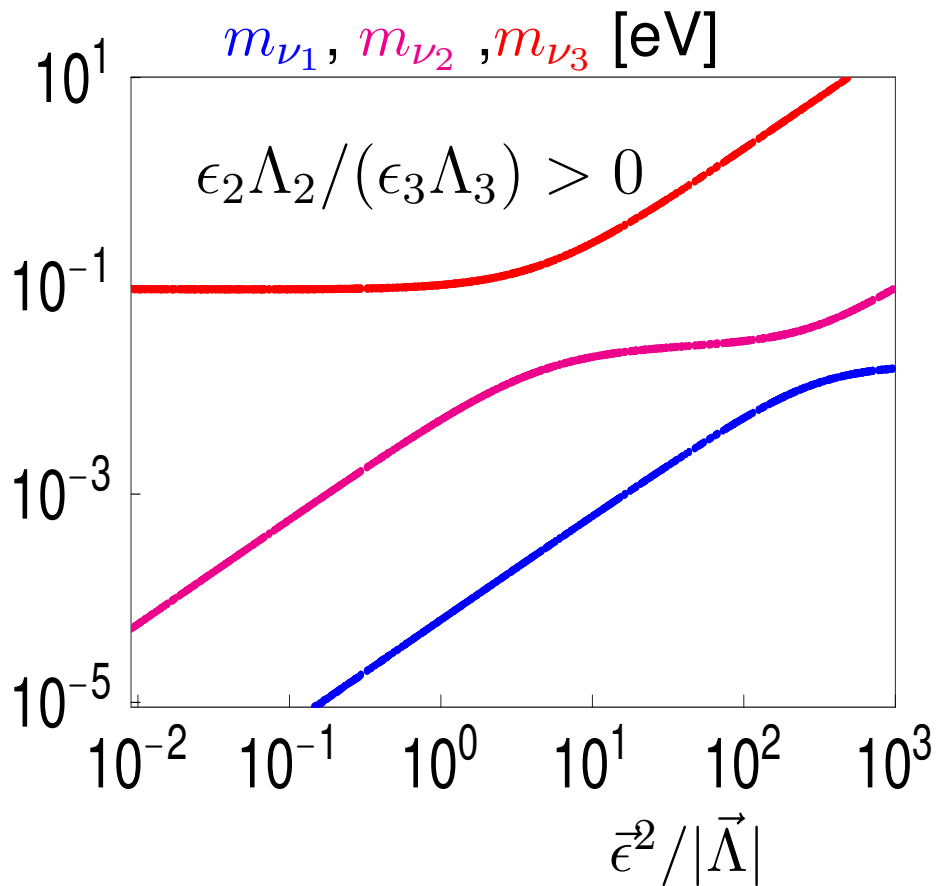
mixing angles

$$\tan^2 \theta_{atm} \simeq \left(\frac{\Lambda_2}{\Lambda_3} \right)^2, \quad U_{e3}^2 \simeq \frac{|\Lambda_1|}{\sqrt{\Lambda_2^2 + \Lambda_3^2}}, \quad \tan^2 \theta_{sol} \simeq \left(\frac{\tilde{\epsilon}_1}{\tilde{\epsilon}_2} \right)^2$$

experimental data require:

$$\frac{|\vec{\Lambda}|}{\sqrt{\det \mathcal{M}_{\tilde{\chi}^0}}} \sim O(10^{-6}), \quad \frac{|\vec{\epsilon}|}{\mu} \sim O(10^{-4})$$

Two examples of neutrino masses as function of $\vec{\epsilon}^2 / |\vec{\Lambda}|$
(other parameters fixed):



dominant modes R-parity violating modes

$$\Gamma(\tilde{\chi}_1^0 \rightarrow W^\pm l_i^\mp) \propto \frac{\Lambda_i^2}{\det \mathcal{M}_{\tilde{\chi}^0}}$$

$$\Gamma(\tilde{\chi}_1^0 \rightarrow \sum_i Z \nu_i) \simeq \frac{1}{2} \sum_i \Gamma(\tilde{\chi}_1^0 \rightarrow W^\pm l_i^\mp)$$

$$\Gamma(\tilde{\chi}_1^0 \rightarrow \nu \tau^+ l_i^-) \propto \frac{\epsilon_i^2}{\mu^2}$$

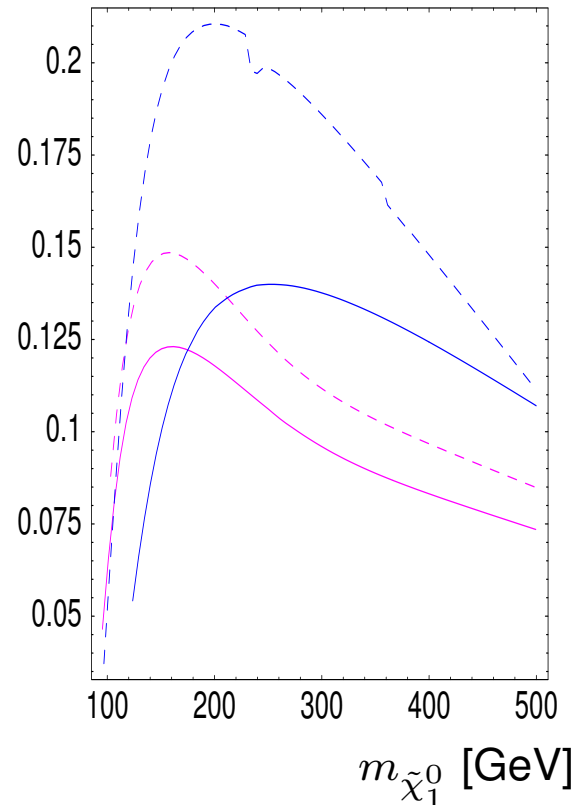
R-parity conserving mode

$$\Gamma(\tilde{\chi}_1^0 \rightarrow \tilde{G} \gamma) \simeq 1.2 \times 10^{-6} \kappa_\gamma^2 \left(\frac{m_{\tilde{\chi}_1^0}}{100 \text{ GeV}} \right)^5 \left(\frac{100 \text{ eV}}{m_{3/2}} \right)^2 \text{ eV}$$

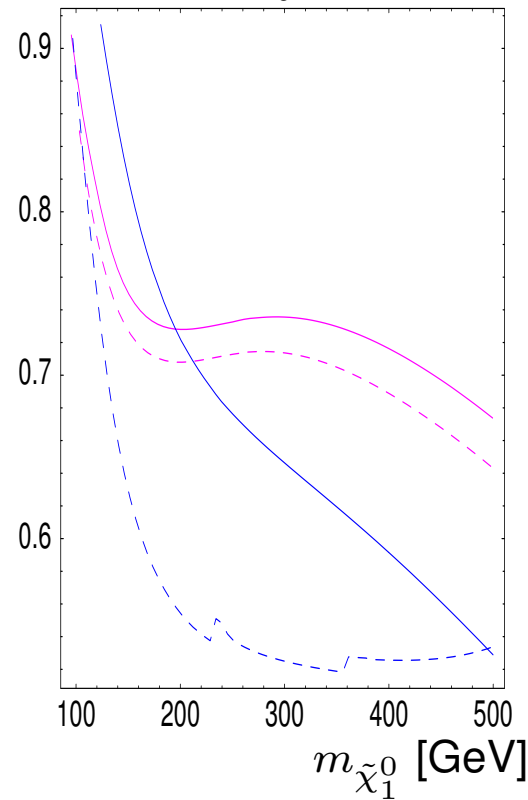
total width

$$\Gamma \simeq (10^{-4} - 10^{-2}) \text{ eV}$$

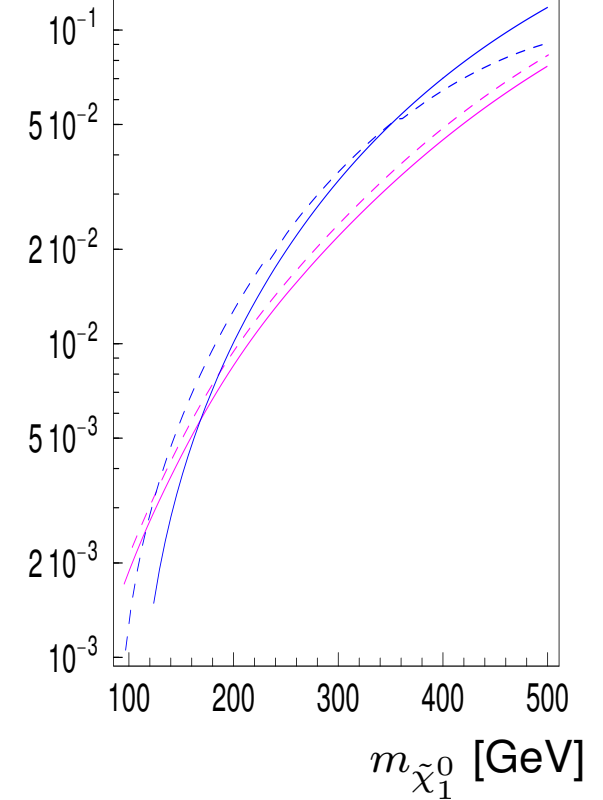
$$\text{BR}(\tilde{\chi}_1^0 \rightarrow \sum_i W l_i)$$



$$\text{BR}(\tilde{\chi}_1^0 \rightarrow \sum_{ij} \nu_i \tau l_j)$$



$$\text{BR}(\tilde{\chi}_1^0 \rightarrow \tilde{G} \gamma)$$



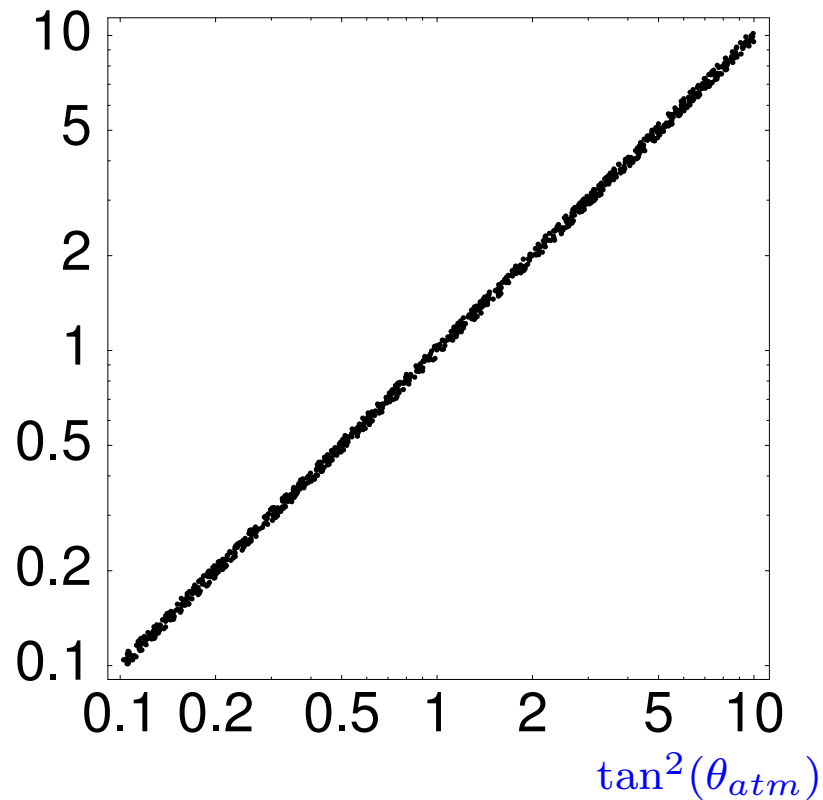
— $\tan \beta = 10, \mu > 0$, - - $\tan \beta = 10, \mu < 0$, — $\tan \beta = 35, \mu > 0$, - - $\tan \beta = 35, \mu < 0$

$m_{3/2} = 100 \text{ eV}, n_5 = 1$

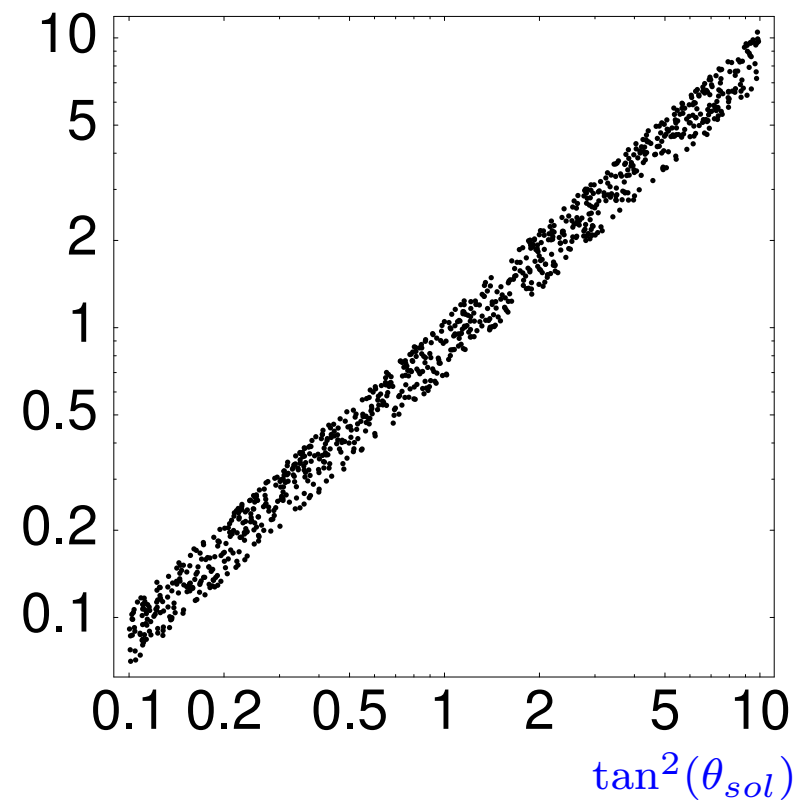
M. Hirsch, W. P. und D. Restrepo, JHEP 0503, 062 (2005)

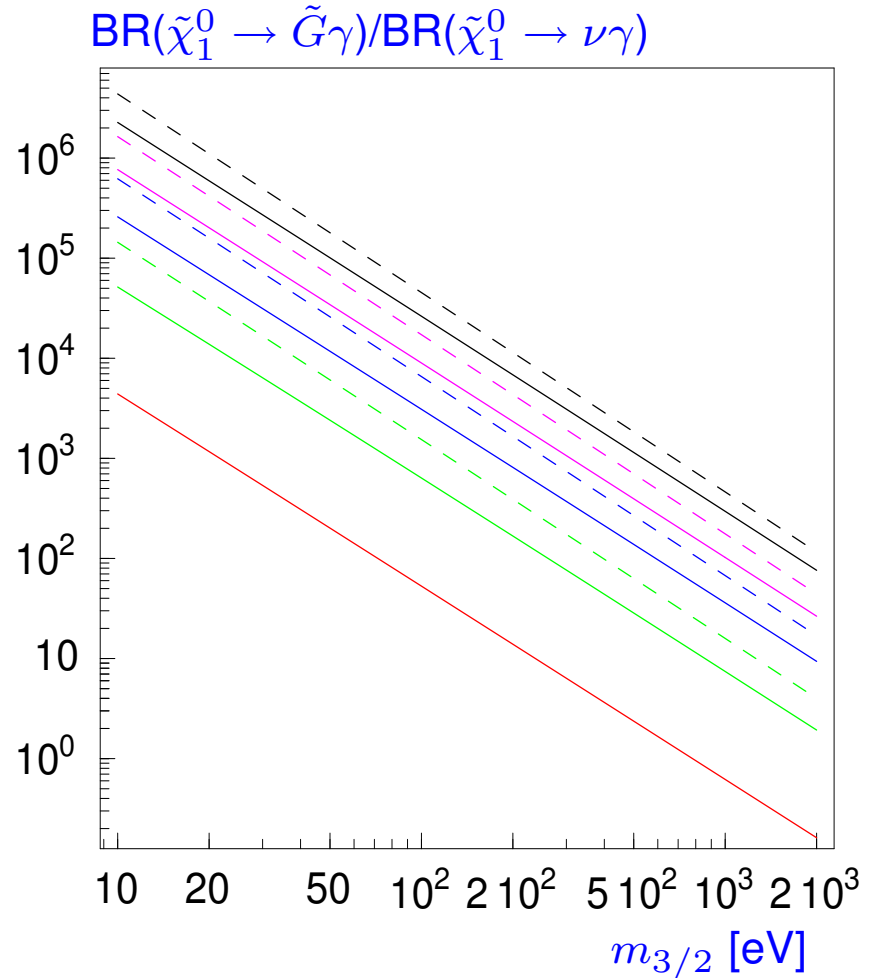
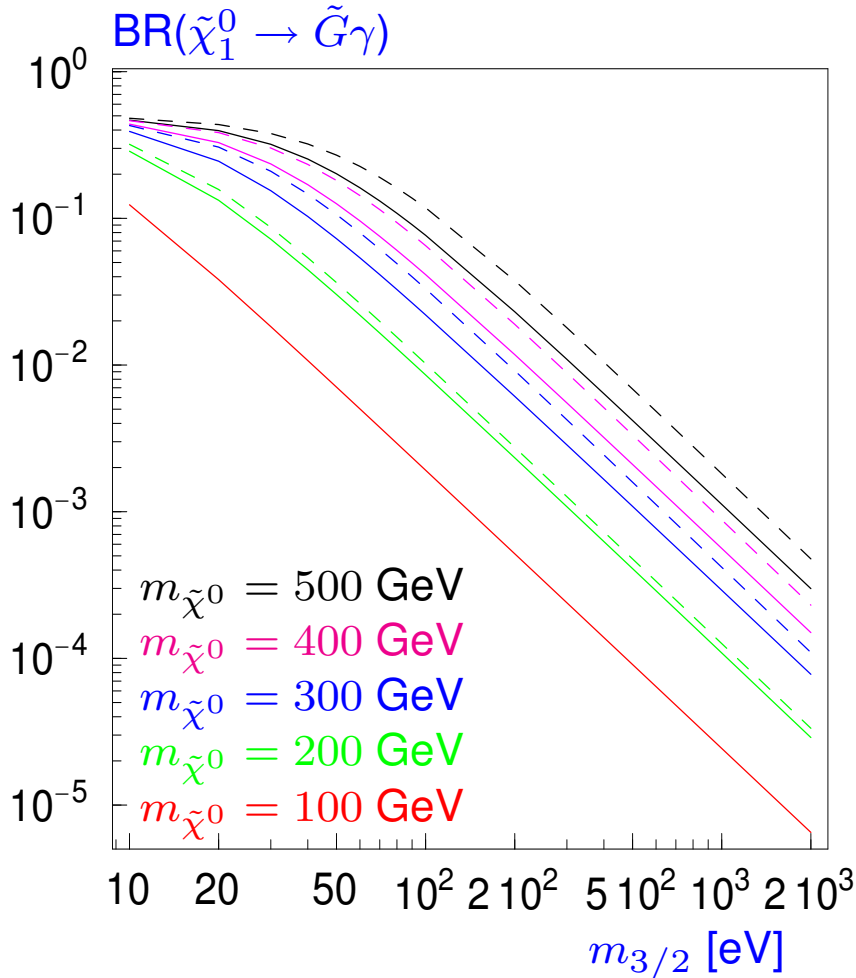
Correlations

$$\text{BR}(\tilde{\chi}_1^0 \rightarrow W\mu) / \text{BR}(\tilde{\chi}_1^0 \rightarrow W\tau)$$



$$\text{BR}(\tilde{\chi}_1^0 \rightarrow \nu e\tau) / \text{BR}(\tilde{\chi}_1^0 \rightarrow \nu\mu\tau)$$





$$n_5 = 1, \tan \beta = 10$$



$$\frac{m_{\tilde{\tau}_1}}{m_{\tilde{\chi}_1^0}} \propto \frac{1}{\sqrt{n_5}}$$

⇒ for $n_5 \geq 3$ hardly points with $\tilde{\chi}_1^0$ LSP



\tilde{l}_R NLSPs: $\text{BR}(l\nu) \gg \text{BR}(l\tilde{G})$



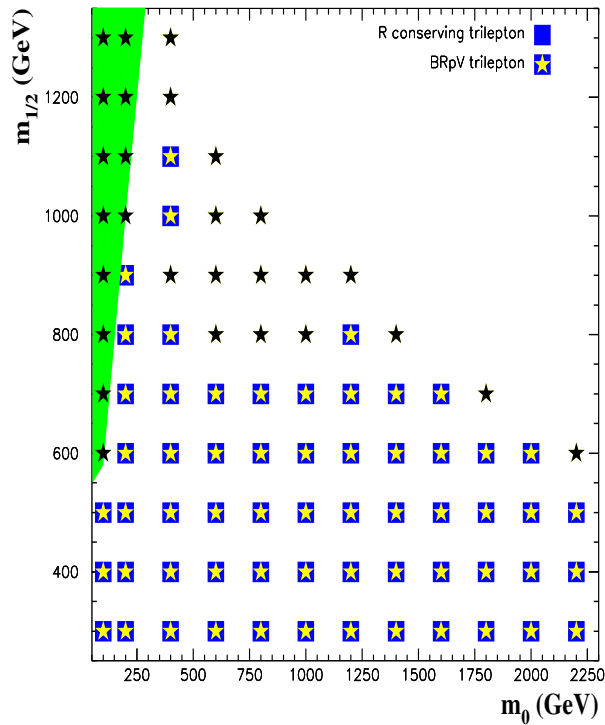
$n_5 = 2$: $\text{BR}(\tilde{G}\gamma)$ reduced by a factor 2-3



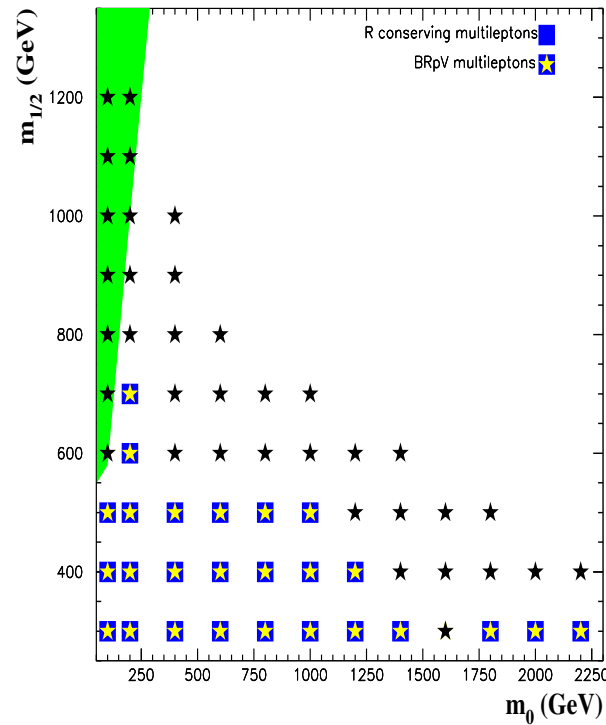
\tilde{G} decays via R-parity violating couplings, however:

$$\Gamma(\tilde{G}) \simeq 3.5 \cdot 10^{-16} \frac{m_\nu [\text{eV}]^3}{0.05 \text{eV}} \frac{m_{3/2}^3}{M_{Pl}^2} \Rightarrow \tau(\tilde{G}) \sim O(10^{31}) \text{Hubbletimes}$$

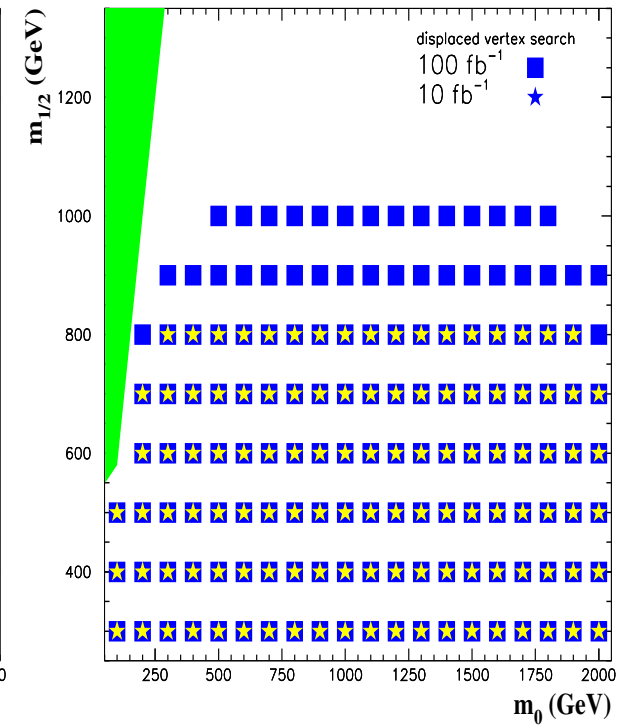
3-lepton channel



multi-lepton channel



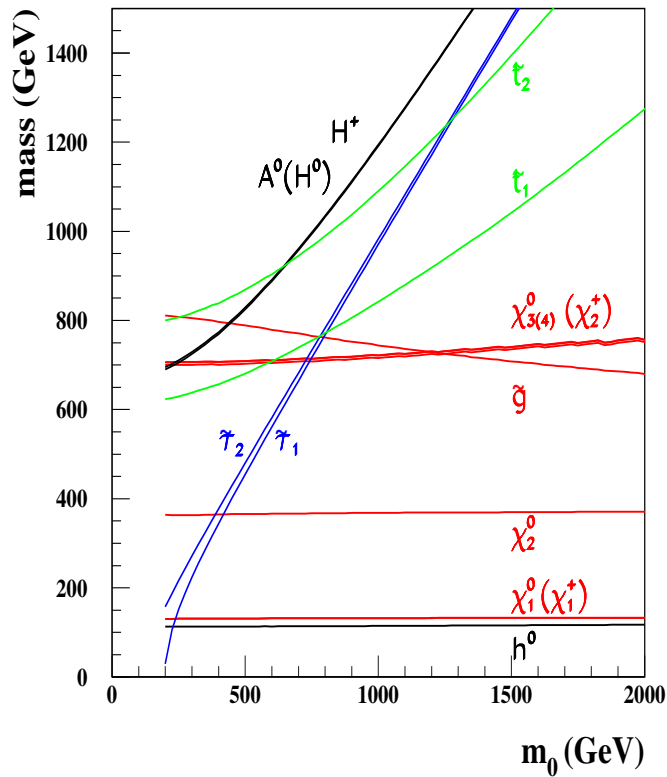
displaced vertex



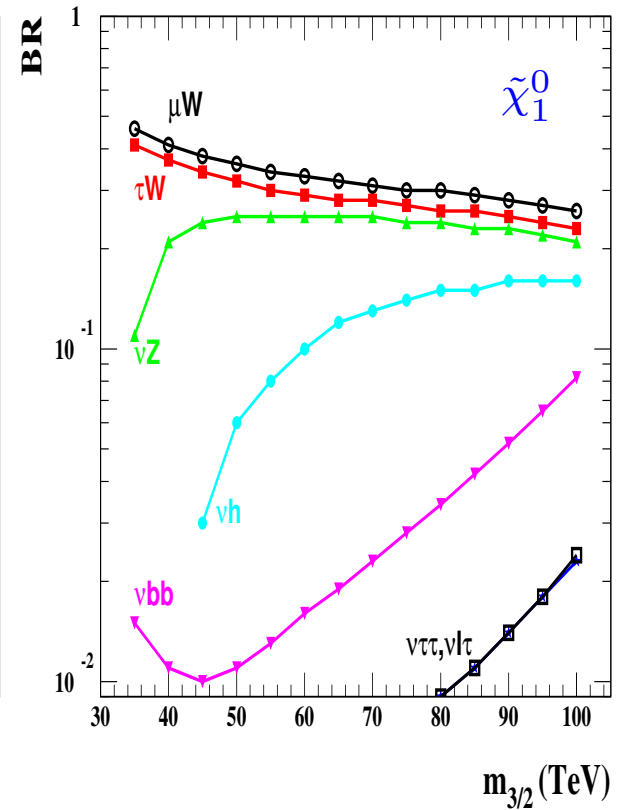
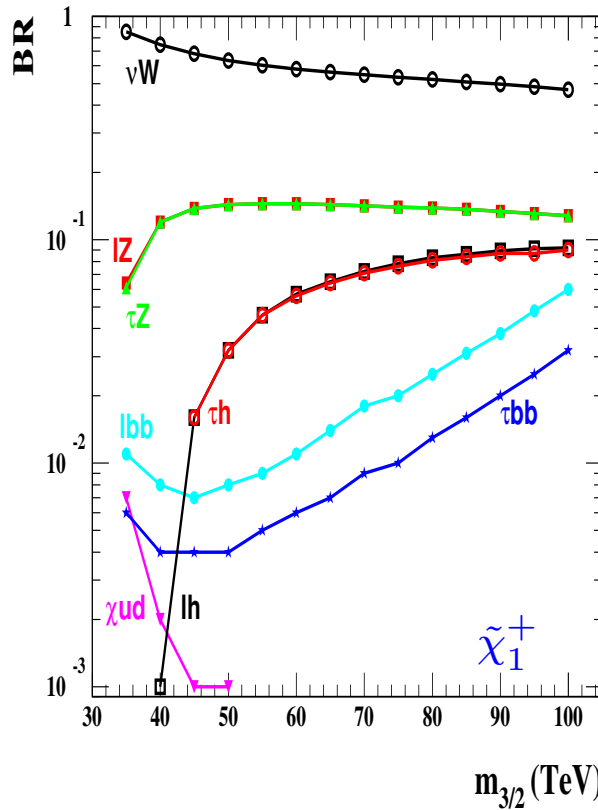
$$L = 100 \text{ fb}^{-1}, A_0 = -100 \text{ GeV}, \tan \beta = 10, \mu > 0$$

F. de Campos et al., JHEP 0805, 048 (2008)

$\tan \beta = 10, \mu > 0$
 $m_{3/2} = 40 \text{ TeV}$



$\tan \beta = 10, \mu > 0, m_0 = 800 \text{ GeV}$



F. de Campos et al., arXiv:0803.4405 (hep-ph)

add singlet \widehat{S} with lepton number

$$\begin{aligned} \mathcal{W} = & h_U^{ij} \widehat{Q}_i \widehat{U}_j \widehat{H}_u + h_D^{ij} \widehat{Q}_i \widehat{D}_j \widehat{H}_d + h_E^{ij} \widehat{L}_i \widehat{E}_j \widehat{H}_d \\ & + h_\nu^i \widehat{L}_i \widehat{\nu}^c \widehat{H}_u - h_0 \widehat{H}_d \widehat{H}_u \widehat{\Phi} + h \widehat{\Phi} \widehat{\nu}^c \widehat{S} + \frac{\lambda}{3!} \widehat{\Phi}^3 \end{aligned}$$

electroweak symmetry breaking $\Rightarrow \langle \widehat{\nu}^c \rangle = v_R / \sqrt{2}$

$$h_\nu^i \widehat{L}_i \widehat{\nu}^c \widehat{H}_u \Rightarrow \frac{v_R h_\nu^i}{\sqrt{2}}$$

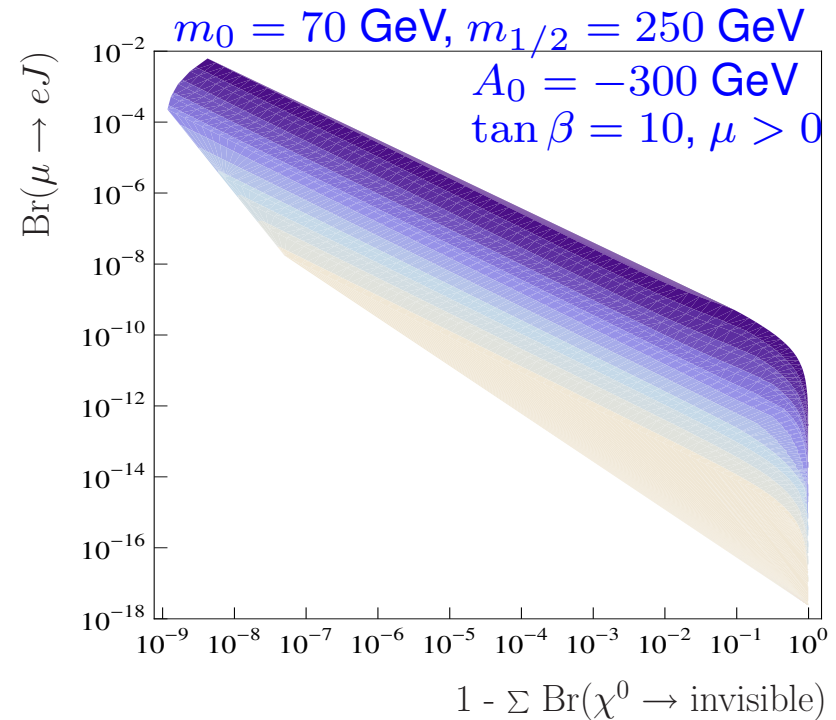
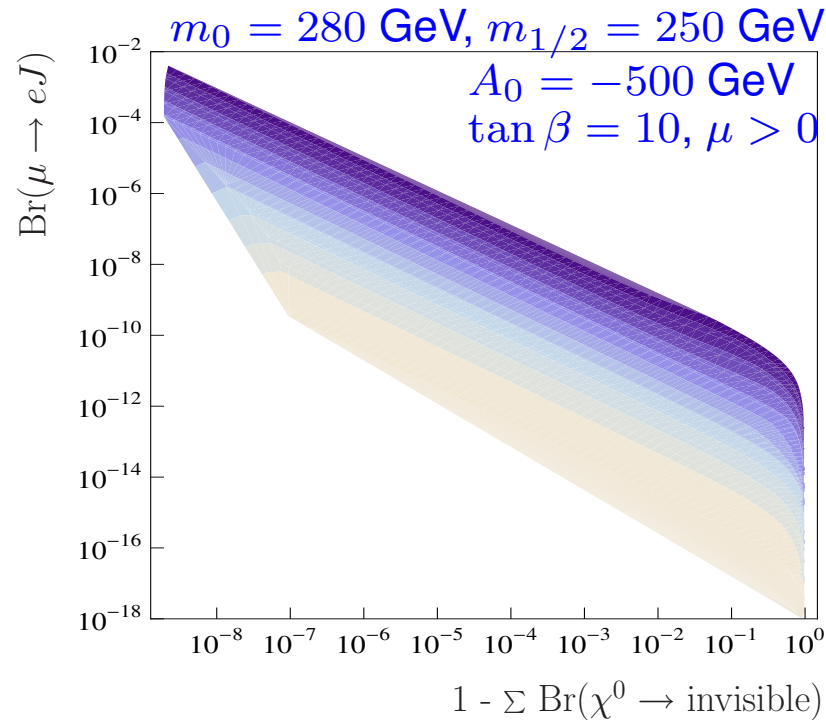
broken lepton number \Rightarrow Goldstone boson called majoron J

$$h^0 \rightarrow JJ$$

M. Hirsch, J. C. Romão, J. W. F. Valle and A. Villanova del Moral, PRD **70**, 073012 (2004),
PRD **73**, 055007 (2006)

induce also

$$\tilde{\chi}_1^0 \rightarrow \nu J \quad , \quad \mu \rightarrow e J$$



M. Hirsch, W. P., PRD **74** (2006) 055003;

M. Hirsch, A. Vicente, J. Meyer, W. P., PRD **79** (2009) 055023.

no majoron J but heavy gauge bosons

extra $U(1)$: additional Z' , \tilde{Z}'

(L. L. Everett, P. Fileviez Perez, S. Spinner PRD **80** (2009) 055007; P. Fileviez Perez, S. Spinner, PD **80** (2009) 015004)

extra $U(1)$ + SUGRA: additional Z' , \tilde{Z}' + exotic states with masses in the TeV range
(R. S. Hundi, S. Pakvasa and X. Tata, PRD **79** (2009) 095011)

Left-right symmetric models: triplet Higgs bosons + higgsinos, W_R , Z_R , \tilde{W}_R , \tilde{Z}_R
K. Huitu and J. Maalampi, *Phys. Lett.* **B344** (1995) 217

Connection to trilinear R -parity violation: rotate (\hat{H}_d, \hat{L}_i) such, that $\epsilon'_i = 0$; gives in leading order of ϵ_i/μ :

$$\lambda'_{ijk} = \frac{\epsilon_i}{\mu} \delta_{jk} Y_{d_k}$$

and

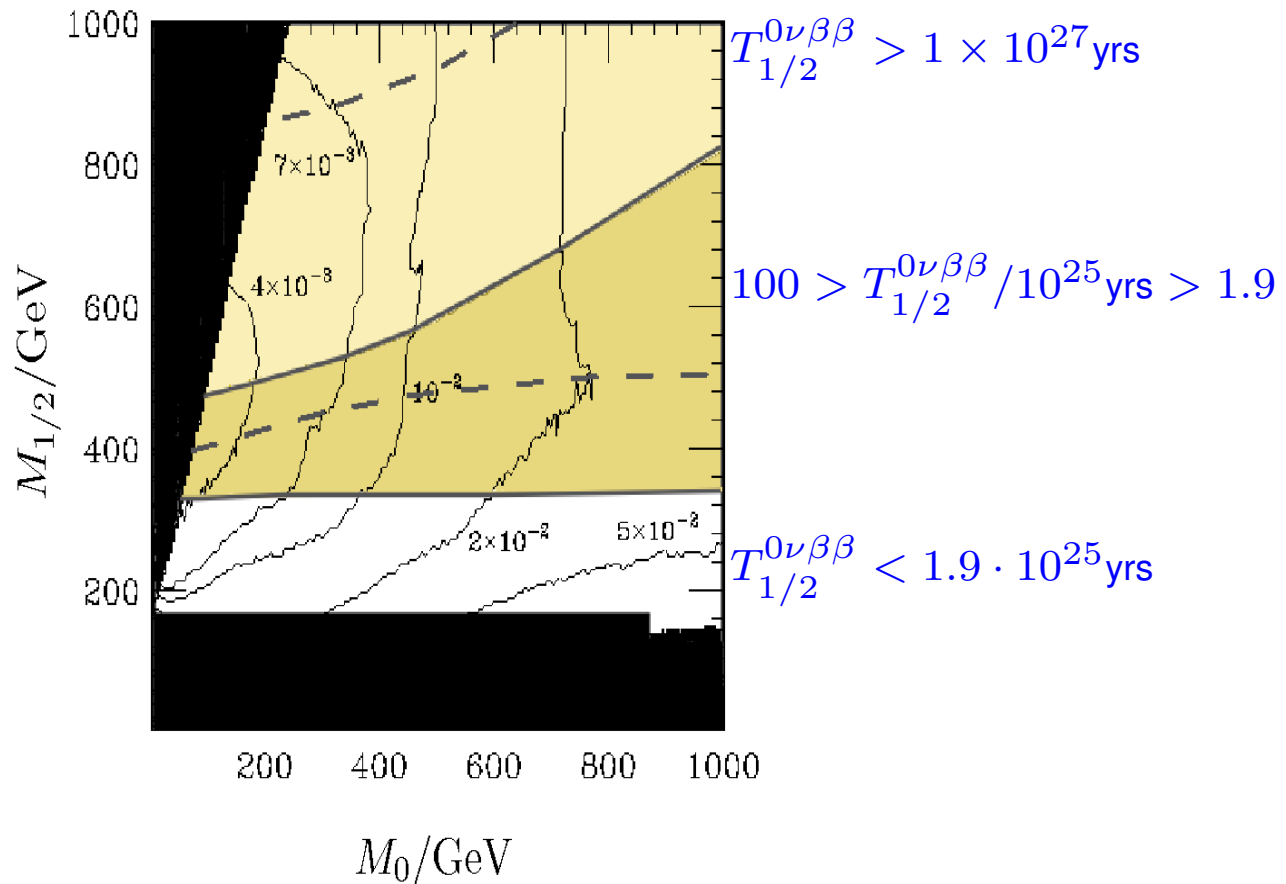
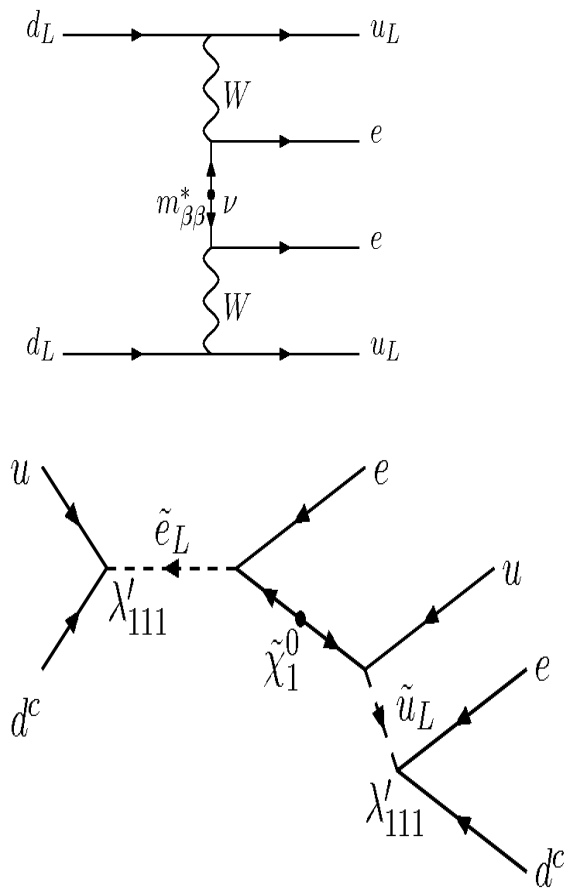
$$\lambda_{121} = Y_e \frac{\epsilon_2}{\mu}, \quad \lambda_{122} = Y_\mu \frac{\epsilon_1}{\mu}, \quad \lambda_{123} = 0$$

$$\lambda_{131} = Y_e \frac{\epsilon_3}{\mu}, \quad \lambda_{132} = 0, \quad \lambda_{133} = Y_\tau \frac{\epsilon_1}{\mu}$$

$$\lambda_{231} = 0, \quad \lambda_{232} = Y_\mu \frac{\epsilon_3}{\mu}, \quad \lambda_{233} = Y_\tau \frac{\epsilon_2}{\mu}$$

$$\lambda_{ijk} = -\lambda_{jik}$$

regions for 5σ discovery



$$T_{1/2}^{0\nu\beta\beta} > 1 \times 10^{27} \text{ yrs}$$

$$100 > T_{1/2}^{0\nu\beta\beta} / 10^{25} \text{ yrs} > 1.9$$

$$T_{1/2}^{0\nu\beta\beta} < 1.9 \cdot 10^{25} \text{ yrs}$$

$L = 10 \text{ fb}^{-1}$ at 14 TeV, $A_0 = 0$, $\tan \beta = 10$, $\mu > 0$

B. C. Allanach, C. H. Kom, H. Päs, HEP 0910 (2009) 026

related work: H. K. Dreiner, P. Richardson, M. H. Seymour, JHEP 0004 (2000) 008

- Dirac neutrinos: displaced vertices if $\tilde{\nu}_R$ LSP, e.g. $\tilde{t}_1 \rightarrow lb\tilde{\nu}_R$
(but NMSSM: $\tilde{t}_1 \rightarrow lb\nu\tilde{\chi}_1^0$)
- Seesaw models:
 - most promising: $\tilde{\tau}_2$ decays
 - very difficult to test at LHC, signals of O(10 fb) or below
 - exceptions: either special kinematics or additional symmetries + NMSSM
 - in case of Seesaw II: different mass ratios

- R-parity violation
 - interesting correlations between ν -physics and LSP decays, testable at LHC
 - displaced vertices
 - Can the model be pinned down?

talk by I. Borjanovic at 'Flavour in the era of LHC', Nov.'05, CERN

$L=100 \text{ fb}^{-1}$

Fit results

| Edge | Nominal Value | Fit Value | Syst. Error Energy Scale | Statistical Error |
|------------------------------------|---------------|-----------|-----------------------------|----------------------|
| $m(ll)^{\text{edge}}$ | 77.077 | 77.024 | 0.08 | 0.05 |
| $m(qll)^{\text{edge}}$ | 431.1 | 431.3 | 4.3 | 2.4 |
| $m(ql)^{\text{edge}}_{\text{min}}$ | 302.1 | 300.8 | 3.0 | 1.5 |
| $m(ql)^{\text{edge}}_{\text{max}}$ | 380.3 | 379.4 | 3.8 | 1.8 |
| $m(qll)^{\text{thres}}$ | 203.0 | 204.6 | 2.0 | 2.8 |

Mass reconstruction

5 endpoints measurements, 4 unknown masses

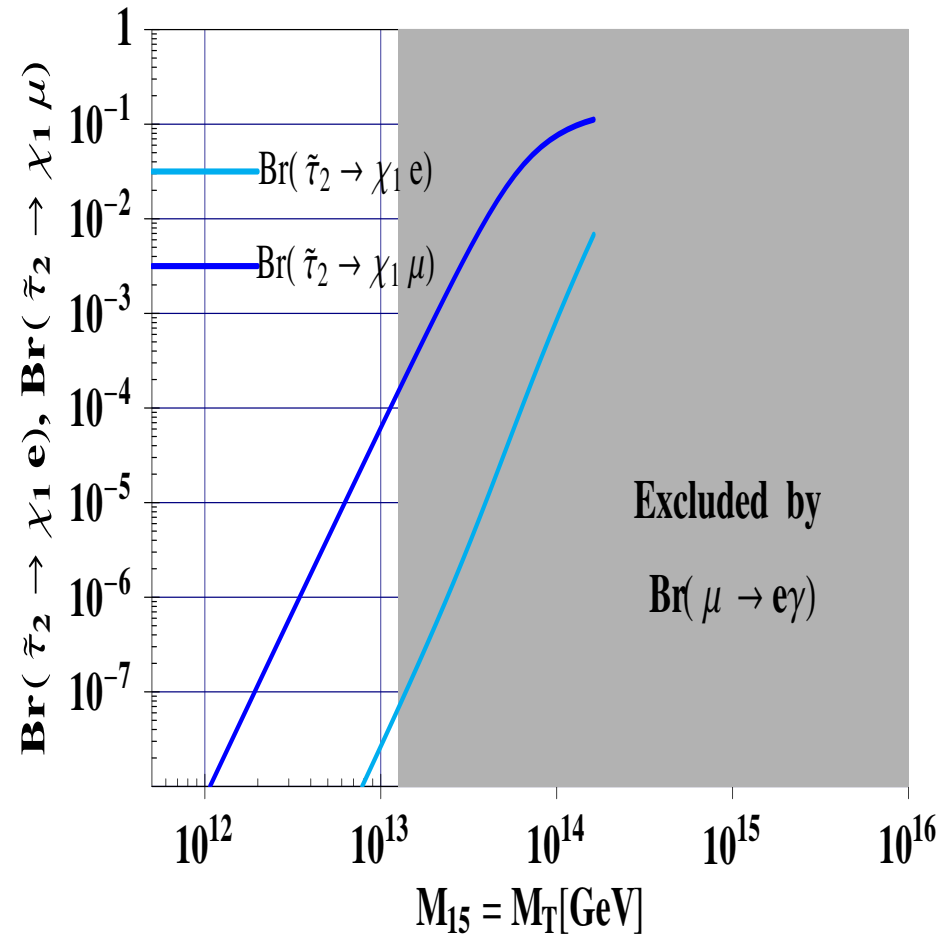
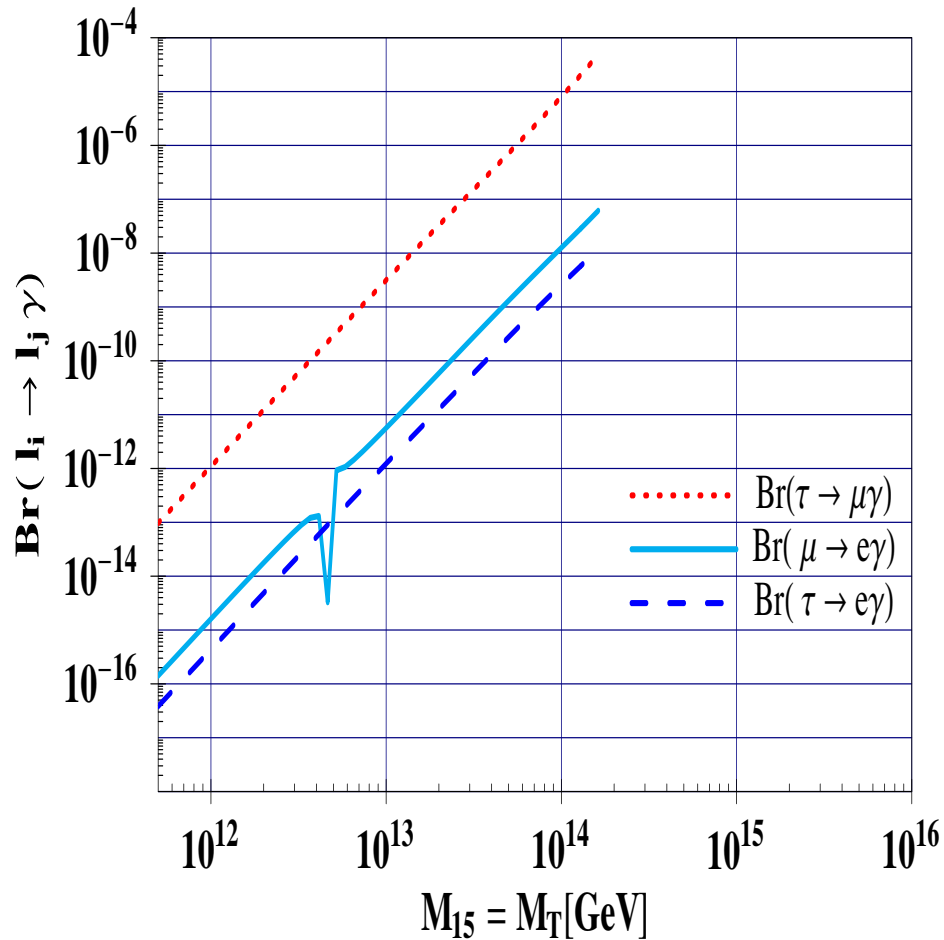
$$\chi^2 = \sum \chi_j^2 = \sum \left[\frac{E_j^{\text{theory}}(\vec{m}) - E_j^{\text{exp}}}{\sigma_j^{\text{exp}}} \right]^2$$

$$E_j^i = E_j^{\text{nom}} + a_j^i \sigma_j^{\text{fit}} + b_j^i \sigma_j^{\text{Escale}}$$

$m(\chi_1^0) = 96 \text{ GeV}$
 $m(l_R) = 143 \text{ GeV}$
 $m(\chi_2^0) = 177 \text{ GeV}$
 $m(q_L) = 540 \text{ GeV}$

$\Delta m(\chi_1^0) = 4.8 \text{ GeV}, \quad \Delta m(\chi_2^0) = 4.7 \text{ GeV},$
 $\Delta m(l_R) = 4.8 \text{ GeV}, \quad \Delta m(q_L) = 8.7 \text{ GeV}$

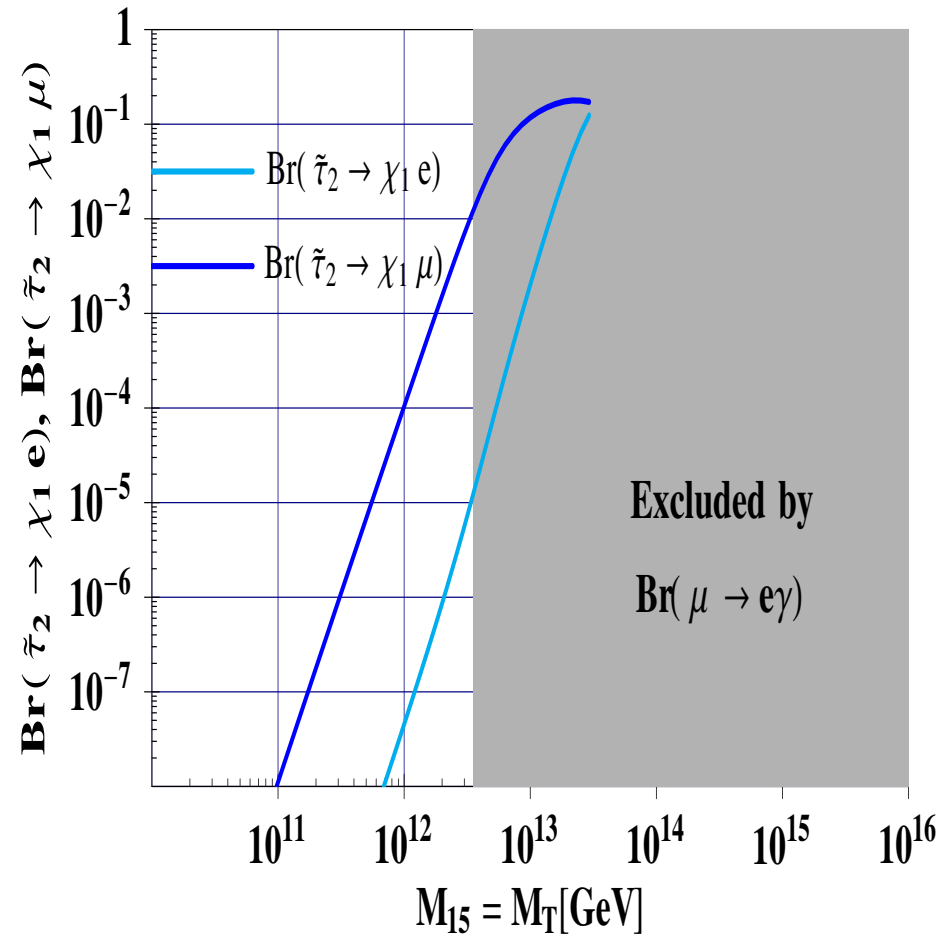
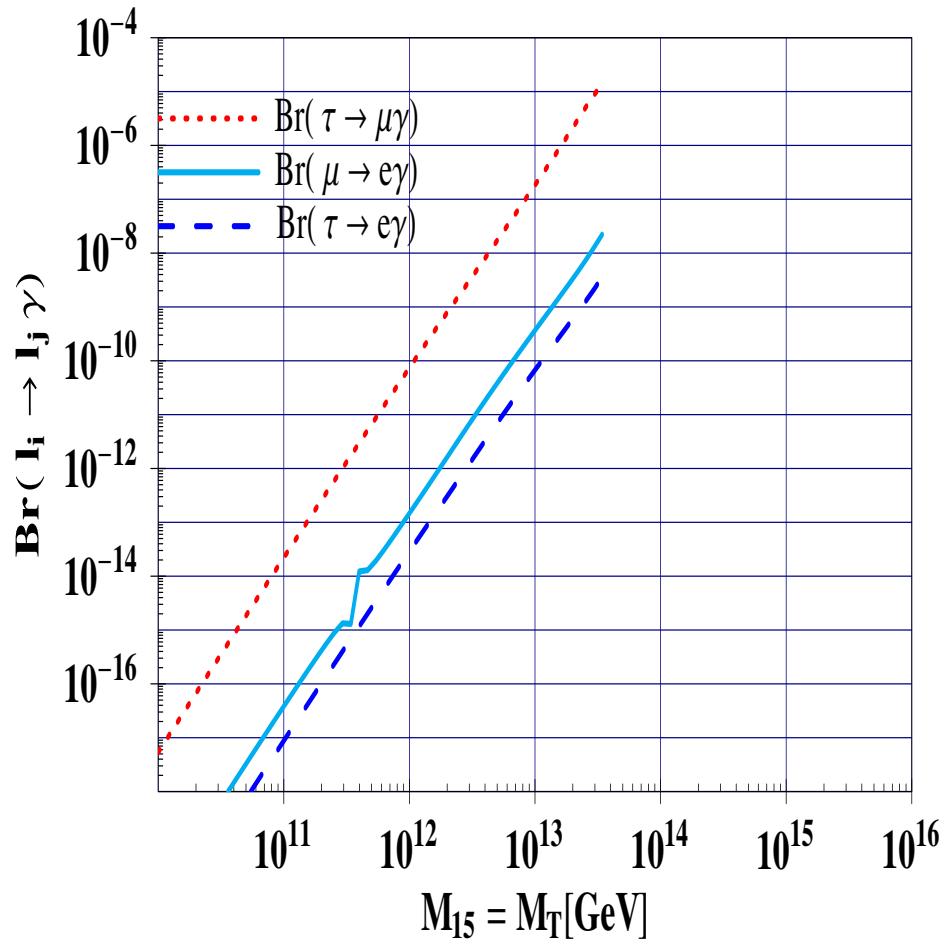
Gjelsten, Lytken, Miller, Osland, Polesello, ATL-PHYS-2004-007



$$\lambda_1 = \lambda_2 = 0.5$$

SPS1a' ($M_0 = 70 \text{ GeV}$, $M_{1/2} = 250 \text{ GeV}$, $A_0 = -300 \text{ GeV}$, $\tan \beta = 10$), $\mu > 0$

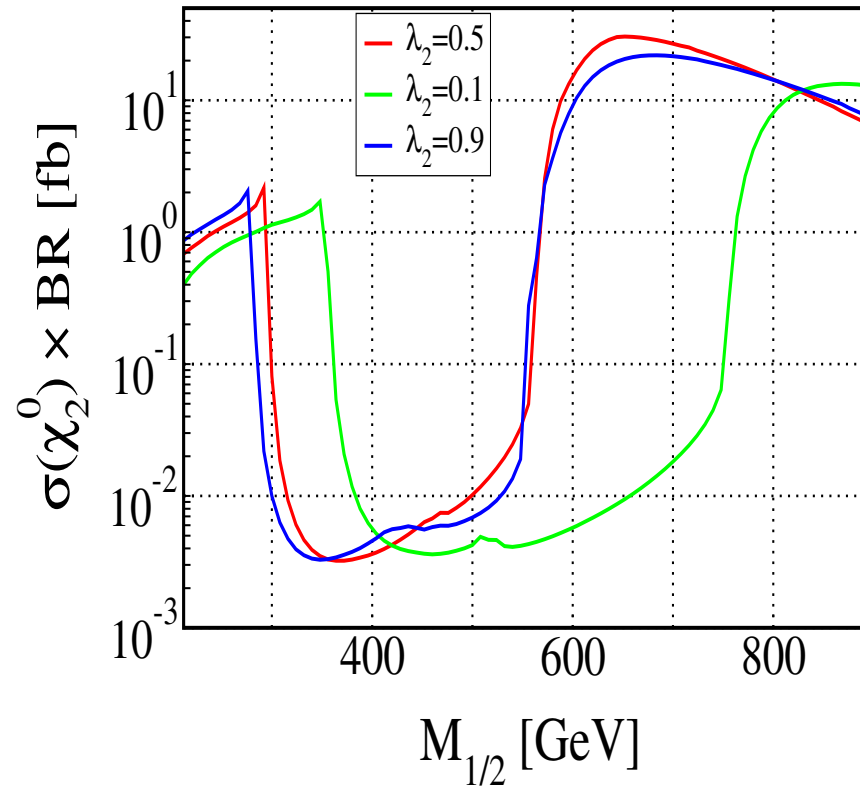
M. Hirsch, S. Kaneko, W. P., Phys. Rev. D 78 (2008) 093004.



$$\lambda_1 = \lambda_2 = 0.05$$

SPS3 ($M_0 = 90 \text{ GeV}$, $M_{1/2} = 400 \text{ GeV}$, $A_0 = 0 \text{ GeV}$, $\tan \beta = 10$), $\mu > 0$

M. Hirsch, S. Kaneko, W. P., Phys. Rev. D 78 (2008) 093004.



$$\sigma(pp \rightarrow \tilde{\chi}_2^0) \times BR(\chi_2^0 \rightarrow \sum_{i,j} \tilde{l}_i l_j \rightarrow \mu^\pm \tau^\mp \tilde{\chi}_1^0)$$

$$m_0 = 100 \text{ GeV } A_0 = 0, \tan \beta = 10, \mu > 0, \lambda_1 = 0.02$$

J.N. Esteves et al., arXiv:0903.1408