## CERN Accelerator School: Introductory Course



## Table of Contents:

1. Introduction ..... 4
1.1. Literature ..... 4
1.2. Bending radius and beam rigidity ..... 6
2. Magnets ..... 9
2.1. General remarks on the calculation of magnetic fields ..... 9
2.2. Particle beam guidance ..... 10
2.3. Particle beam focusing ..... 12
2.4. Correction of chromatic errors ..... 17
2.5. Multipole expansion ..... 19
2.6. Effective field length ..... 21
3. Linear Beam Optics ..... 23
3.1. A quick and simple first approach using geometric optics ..... 23

## CERN Accelerator School: Introductory Course

3.2. Some considerations concerning the equations of motion ..... 28
3.3. Equations of motion in a moving reference system ..... 30
3.4. Matrix formalism ..... 36
3.4.1. Drift space ..... 37
3.4.2. Dipole magnets ..... 37
3.4.3. Quadrupole magnets ..... 40
3.4.4. Particle orbits in a system of magnets ..... 42
3.5. Particle beams and phase space ..... 44
3.3.1. Beam emittance ..... 44
3.5.2. Twiss parameters ..... 46
3.5.3. Beta functions ..... 49
3.5.4. Transformation in phase space ..... 53
4. Circular Accelerators ..... 61
4.1. Weak focusing ..... 61
4.2. Strong focusing ..... 64
4.2.1. Stability criterion ..... 66
4.3. Periodic focusing systems ..... 69
4.3.1. General FODO lattice ..... 69
4.3.2. Periodic beta functions ..... 73
4.4. Transverse beam dynamics ..... 76
4.4.1. Closed orbit ..... 76
4.4.2. Betatron tune ..... 77
4.4.2. Filamentation ..... 78
4.4.3. Normalized coordinates ..... 80
4.4.4. Closed orbit distortions ..... 81
4.4.5. Gradient errors ..... 89
4.4.6. Optical resonances ..... 95
4.5. Beam dynamics with acceleration ..... 99
5. Dynamics with Off Momentum Particles ..... 100
5.1 Dispersion and dispersion functions ..... 100
5.2 Dispersion in circular accelerators ..... 103
5.3. Chromaticity ..... 106
5.4 Path length and momentum compaction ..... 111


## Recommended reading:

- A. Wolski, Beam Dynamics in high energy particle accelerators, Imperial College Press, ISBN 978-1-78326-277-9
- CAS proceedings and references therein


## 1. Introduction

### 1.1. Literature

- S.Y. Lee: Accelerator Physics, $3^{\text {rd }}$ edition, World Scientific, New Yersey 2012, ISBN 978-981-4374-94-1
- Bryant / Johnson: The Principles of Circular Accelerators and Storage Rings, Cambridge University Press, Cambridge 2005, ISBN 978-0-521-61969-1
- Edwards / Syphers: An Introduction to the Physics of High Energy Accelerators, John Wiley \& Sons, New York 1992, ISBN 978-0-471-55163-8
- K. Wille: The physics of particle accelerators, Oxford Univ. Press 2005, Oxford, ISBN 0-19-850550-7
- H- Wiedemann: Particle Accelerator Physics, $4^{\text {th }}$ edition, Springer 2015, Berlin, ISBN 978-3-319-18316-9
- Chao / Tigner: Handbook of Accelerator Physics and Engineering, $2^{\text {nd }}$ edition, World Scientific, Singapore 2013, ISBN 987-4417-17-4


## CERN Accelerator School: Introductory Course

- F. Hinterberger: Physik der Teilchenbeschleuniger und Ionenoptik,

2. Ausgabe, Springer 2008, Berlin, ISBN 978-3-540-75281-3

- K. Wille: Physik der Teilchenbeschleuniger und Synchrotronstrahlungsquellen,

2. überarb. und erw. Ausgabe, Teubner 1996, Stuttgart, ISBN 978-3-519-13087-1

- Rossbach / Schmüser: Basic Course on Accelerator Optics,

CAS $5^{\text {th }}$ general accelerator physics course CERN 94-01


### 1.2. Bending radius and beam rigidity

Particle guidance and focusing based on beam deflection by Lorentz force

$$
\vec{F}=q \cdot(\vec{E}+\vec{v} \times \vec{B})
$$

Ultra-relativistic particles move with speed very close to speed of light!
Impact of magnetic fields is enhanced by enormous factor:

$$
v \approx c \quad \Rightarrow \quad B=1 \text { Tesla } \leftrightarrow E=3 \cdot 10^{8} \mathrm{~V} / \mathrm{m}
$$

Only magnetic fields are used for beam deflection!
Bending radius from balance of forces $\left(m=\gamma_{r} m_{0}\right)$ :

$$
\vec{B} \perp \vec{v}: \quad m \frac{v^{2}}{\rho}=q \cdot v \cdot B \quad p=m v=q \rho B
$$

Leads to the definition of the magnetic rigidity $\boldsymbol{B} \rho$ !
In circular accelerators, the magnetic rigidity defines the momentum of the beam:

$$
\frac{p}{q}=B \rho=1 \mathrm{Tm} \hat{=} \quad p=0.3 \frac{\mathrm{GeV}}{\mathrm{c}}
$$




## 2. Magnets

### 2.1. General remarks on the calculation of magnetic fields


$\Phi=$ const. defines the pole's contour!
Magn. Induction from $\vec{B}=\mu_{0} \mu_{r} \vec{H}$
Taylor Expansion of the Magnetic Field:

$$
B_{y}(x, y)=\underbrace{B_{y}(0, y)}_{\text {Dipoles }}+\underbrace{x \cdot \frac{\partial B_{y}}{\partial x}(0, y)}_{\text {Quadrupoles }}+\underbrace{x^{2} \cdot \frac{\partial^{2} B_{y}}{\partial x^{2}}(0, y)}_{\text {Sextupoles }}+\ldots
$$

## CERN Accelerator School: Introductory Course

### 2.2. Particle beam guidance

Deflection of particles $\rightarrow$ homogenous field: $\quad \vec{B}=B_{0} \cdot \hat{e}_{y}=$ const.
Corresponding magnetic potential:

$$
\Phi(x, y)=-B_{0} \cdot y
$$

defining the pole's profile to be flat and parallel: Dipole Magnets!

$n \cdot I=\oint \vec{H} \cdot d \vec{s}=\int_{\text {gap }} \vec{H}_{0} \cdot d \vec{s}+\int_{\text {yoke }} H_{E}, \quad \mu_{r} \cdot\left|H_{E}\right|=\left|H_{0}\right| \Rightarrow\left|H_{0}\right| \gg\left|H_{E}\right|$

$$
B_{0}=\mu_{0} \frac{n \cdot I}{h}, \quad \text { Curvature: } \kappa=\frac{1}{\rho}=\frac{q}{p} B_{0}=\frac{q \mu_{0}}{p} \frac{n \cdot I}{h}, \quad[\kappa]=\mathrm{m}^{-1}
$$

## CERN Accelerator School: Introductory Course

## Dipole Magnets:



> Iron dominated:
> field determined by
> geometry of poles
> $\quad \rightarrow 2$ flat poles

## Superconducting:

field determined by
geometry of coils

$$
\rightarrow j(\phi) \sim \cos \phi
$$

### 2.3. Particle beam focusing

Restoring force, linearly increasing with increasing distance from the axis:

$$
B_{y}=-g \cdot x, \quad B_{x}=-g \cdot y \quad \text { with } \quad g=\frac{\partial B_{y}}{\partial x}=\frac{\partial B_{x}}{\partial y}=\text { const. }
$$

Corresponding potential: $\quad \Phi(x, y)=g \cdot x \cdot y, \quad$ solves $\quad \vec{\nabla} \cdot \vec{B}=-\Delta \Phi=0$
defining the pole's profile to four hyperbolic poles: Quadrupole Magnets!


The "restoring" force acting on the particles is

$$
\vec{F}=q \cdot(\overrightarrow{\mathrm{v}} \times \vec{B})=q \mathrm{vg} \cdot\left(x \hat{e}_{x}-y \hat{e}_{y}\right)
$$

A quadrupole magnet is therefore focusing only in one plane and defocusing in the other; depending on the sign of $g$.

The $g$-parameter may be related to the current of the coils by evaluating the closed loop integral

$$
n \cdot I=\oint \vec{H} \cdot d \vec{s}=\int_{0}^{1} \vec{H}_{0} \cdot d \vec{s}+\int_{1}^{2} \vec{H}_{E} \cdot d \vec{s}+\int_{2}^{0} \vec{H}_{0} \cdot d \vec{s} \approx \int_{0}^{1} \vec{H}_{0} \cdot d \vec{s}
$$



One obtains with $\vec{H} \cdot d \vec{s}=\frac{g}{\mu_{0}} r \cdot d r$ :

$$
g=\frac{2 \cdot \mu_{0} \cdot n \cdot I}{a^{2}}, \text { normalized: }
$$

Quadrupole Strength

$$
k=\frac{q}{p} g=\frac{2 q \mu_{0}}{p} \frac{n \cdot I}{a^{2}}, \quad[k]=\mathrm{m}^{-2}
$$

The focal length of a thin quadrupole magnet of length $L$ can be derived from the deflection angle $\alpha$ of the particles beam and its relation to the quadrupole strength $k$,


$$
\tan \alpha=\frac{x}{f}
$$

$$
\tan \alpha=\frac{L}{R}=L \cdot \frac{q}{p} B_{y}=\frac{q}{p} g x L=x k L
$$

$\rightarrow$ Gives a better understanding of the quadrupole strength:

$$
\frac{1}{f}=k \cdot L
$$

Here we have assumed the length $L$ to be short compared to the focal length $f$ such that $R$ does not change significantly within the quadrupole magnetic field.

## CERN Accelerator School: Introductory Course

## Strong Focusing:

Light optics:


Magnet optics:
Strong focusing

or
AG focusing

Simplest way: FODO lattice

Detailed discussion later!

## CERN Accelerator School: Introductory Course

## Quadrupole magnets:



> Iron dominated:
> field determined by
> geometry of poles
> $\rightarrow 4$ hyperbolic poles

## Superconducting:

field determined by geometry of coils
$\rightarrow j(\phi) \sim \cos 2 \phi$

## CERN Accelerator School: Introductory Course

### 2.4. Correction of chromatic errors

Quadratic increase of magnetic fields increasing distance from the axis:

$$
B_{y}=\frac{1}{2} g^{\prime} \cdot\left(x^{2}-y^{2}\right) \quad \text { with } \quad g^{\prime}=\frac{\partial^{2} B_{y}}{\partial x^{2}}=\text { const } \text {. }
$$

Corresponding potential: $\quad \Phi(x, y)=\frac{1}{6} g^{\prime}\left(y^{3}-3 x^{2} y\right)$, solves $\vec{\nabla} \cdot \vec{B}=-\Delta \Phi=0$


Sextupole Magnets
Six poles, profile

$$
x(y)= \pm \sqrt{\frac{y^{2}}{3} \pm \frac{2 \Phi_{0}}{g^{\prime} y}}
$$

or using the aperture

$$
\begin{aligned}
a & =\sqrt[3]{6 \Phi_{0} / g^{\prime}} \\
x(y) & = \pm \sqrt{\frac{y^{2}}{3} \pm \frac{a^{3}}{3 y}}
\end{aligned}
$$

The $g$ ' parameter may be related to the current of the coils in the well-known manner:

$$
g^{\prime}=\frac{\partial^{2} B_{y}}{\partial x^{2}}=6 \mu_{0} \frac{n I}{a^{3}}
$$

and we obtain for the transverse magnetic fields:

$$
B_{x}(x, y)=-\frac{\partial \Phi}{\partial x}=g^{\prime} x y \quad \text { and } \quad B_{y}(x, y)=-\frac{\partial \Phi}{\partial y}=\frac{1}{2} g^{\prime}\left(x^{2}-y^{2}\right)
$$



We will therefore expect a coupling of particles motion in the horizontal and vertical plane due to the $\boldsymbol{y}$-dependence of the vertical field.

Normalizing $g$ ' to the particles momentum, we obtain the sextupole strength

$$
m=\frac{q}{p} g^{\prime}=\frac{6 q \mu_{0}}{p} \frac{n I}{a^{3}}, \quad[m]=\mathrm{m}^{-3}
$$

A simple understanding of the action of a sextupole will be given later!

### 2.5. Multipole expansion

General treatment by multipole expansion in polar coordinates:

$$
\begin{aligned}
& B_{r}(r, \phi)=B_{0} \sum_{n=1}^{\infty}\left(\frac{r}{\rho}\right)^{n-1} \cdot\left(b_{n} \sin (n \phi)-a_{n} \cos (n \phi)\right) \\
& B_{\phi}(r, \phi)=B_{0} \sum_{n=1}^{\infty}\left(\frac{r}{\rho}\right)^{n-1} \cdot\left(a_{n} \sin (n \phi)+b_{n} \cos (n \phi)\right)
\end{aligned}
$$



Contribution of multipole $n: \quad|B|_{n}=\sqrt{B_{r, n}^{2}+B_{\phi, n}^{2}}=B_{0}\left(\frac{r}{\rho}\right)^{n-1} \sqrt{a_{n}^{2}+b_{n}^{2}}$
Generally: 2 n -pole has $2 \pi / n$ symmetry, $|B|_{n}$ scales with $r^{n-1}$.
$n=1$ : dipole magnet
$n=2$ : quadrupole magnet
$n=3$ : sextupole magnet
$n=4$ : octupole magnet
$n=5$ : decapole magnet

## Classification:

$b_{n} \neq 0$ : "upright" magnets
$a_{n} \neq 0$ : "skew" magnets, rotated by $\pi / n$

## CERN Accelerator School: Introductory Course

## Normal or upright magnets:





Skew or rotated magnets:



from Zolkin, Timofey, Phys.Rev.Accel.Beams 20 (2017) no.4, 043501

### 2.6. Effective field length

The assumption of a constant field distribution along the longitudinal axis $(\partial \vec{B} / \partial s=0)$ is not valid in general due to the fringing fields at the end of the magnets. In order to simplify the calculation of the optics of particle accelerators, an effective field length $l_{\text {eff }}$ of each magnet is usually defined, calculated from the path-integral

$$
\int_{-\infty}^{\infty} \vec{B} \cdot d \vec{s}=\vec{B}_{0} \cdot l_{e f f}
$$


and approximating the real longitudinal field by a rectangular shaped profile.
Note: $l_{\text {eff }}$ differs from the length $L$ of the iron poles, in almost all cases $l_{\text {eff }}>L$.


## 3. Linear Beam Optics

### 3.1. A quick and simple first approach using geometric optics

Reference path = path of a particle moving on the design path:


Use coordinate system fixed to reference particle, moving along the reference path! Horizontal position and angle of a particle given by displacements $x, x^{\prime}$
Considering paraxial optics: $\quad x \ll \rho, \quad x^{\prime}=\tan \alpha \approx \alpha$


Impact of magnets in a very rough approximation:
dipole magnet: drift of length $L_{D}$
quadrupole magnet: thin lens with focal lengths $f_{x}=-\frac{1}{k L_{Q}}, f_{y}=\frac{1}{k L_{Q}}$
Particle positions in horizontal / vertical phase space are changed by matrices:

## CERN Accelerator School: Introductory Course



| Phase space | Position | Drift | Dipole | Quadrupole |
| :---: | :---: | :---: | :---: | :---: |
| horizontal | $\vec{X}(s)=\binom{x}{x^{\prime}}$ | $\mathbf{M}_{d}=\left(\begin{array}{cc}1 & L \\ 0 & 1\end{array}\right)$ | $\mathbf{M}_{D}=\left(\begin{array}{cc}1 & L \\ 0 & 1\end{array}\right)$ | $\mathbf{M}_{Q}=\left(\begin{array}{cc}1 & 0 \\ \pm 1 / f & 1\end{array}\right)$ |


| Phase Space | Position | Drift | Dipole | Quadrupole |
| :---: | :---: | :---: | :---: | :---: |
| vertical | $\vec{Y}(s)=\binom{y}{y^{\prime}}$ | $\mathbf{M}_{d}=\left(\begin{array}{ll}1 & L \\ 0 & 1\end{array}\right)$ | $\mathbf{M}_{D}=\left(\begin{array}{cc}1 & L \\ 0 & 1\end{array}\right)$ | $\mathbf{M}_{Q}=\left(\begin{array}{cc}1 & 0 \\ \mp 1 / f & 1\end{array}\right)$ |

Calculation of single particle trajectories by matrix multiplication, e.g.:



## CERN Accelerator School: Introductory Course

### 3.2. Some considerations concerning the equations of motion

A correct treatment requires solving the equations of motion in a moving reference system. Again we will try a somehow "superficial" approach and look at the forces:

$$
m \ddot{\vec{r}}=m(\beta c)^{2} \vec{r}^{\prime \prime}=\vec{F}=e \cdot(\overrightarrow{\mathrm{v}} \times \vec{B})
$$

using $\quad \dot{\vec{r}}=\frac{d \vec{r}}{d s} \cdot \frac{d s}{d t}=\vec{r}^{\prime} \cdot \dot{s}, \ddot{\vec{r}}=\frac{d \dot{\vec{r}}}{d s} \cdot \frac{d s}{d t} \approx \frac{d \vec{r}^{\prime}}{d s} \cdot \dot{s} \cdot \frac{d s}{d t}=\vec{r}^{\prime \prime} \cdot \dot{s}^{2} \approx\left(\beta_{r} c\right)^{2} \vec{r}^{\prime \prime}(s)$

- quadrupoles: $\quad \vec{F}=m \cdot\left(\beta_{r} c\right)^{2} \cdot k \cdot\left(x \hat{e}_{x}-y \hat{e}_{y}\right)$, remember: $k=\frac{q}{p} g$
- dipoles:

$$
\vec{F}=m \cdot\left(\beta_{r} c\right)^{2} \cdot \frac{1}{\rho} \cdot \frac{\Delta p}{p_{0}} \cdot \hat{e}_{x}, \quad \text { remember: } \quad \frac{1}{\rho}=\frac{q}{p_{0}} B_{0}
$$

(take care: a particle with nominal momentum $p_{0}$ isn't deflected in the moving frame when traversing a dipole, so this contribution to the deflection in the lab frame has to be subtracted: $\vec{F} \rightarrow \vec{F}(p)-\vec{F}\left(p_{0}\right)$, leading to $\left.p \rightarrow \Delta p\right)$

In addition, we have to take care of the geometric focusing in the horizontal plane:

## CERN Accelerator School: Introductory Course

Quadrupole Magnet:

$$
1 / f=k_{\text {Quad }} \cdot L
$$


$\vec{F}_{m a g}=m \cdot\left(\beta_{r} c\right)^{2} \cdot k \cdot\left(x \hat{e}_{x}-y \hat{e}_{y}\right)$

Dipole Magnet:

focusing analogue to quadrupole:

$$
1 / f=k_{D i p} \cdot L
$$

$$
k_{D i p}=-1 / \rho^{2}
$$

Putting everything together, we obtain the famous linear equations of motion:

$$
\begin{gathered}
x^{\prime \prime}(s)+\left(\frac{1}{\rho^{2}(s)}-k(s)\right) \cdot x(s)=\frac{1}{\rho(s)} \frac{\Delta p}{p} \\
y^{\prime \prime}(s)+k(s) \cdot y(s)=0
\end{gathered}
$$

### 3.3. Equations of motion in a moving reference system

Moving orthogonal, right-handed coordinate system ( $\boldsymbol{x}, \boldsymbol{y}, \boldsymbol{s}$ ) that follows a reference particle traveling along its ideal path (design orbit):


We will concentrate on ideal orbits laying within the horizontal plane, therefore

$$
\vec{r}=(\rho+x) \cdot \hat{e}_{x}+y \cdot \hat{e}_{y}, \quad x, y \ll \rho
$$

## CERN Accelerator School: Introductory Course

Using this reference system moving with $\dot{s}=R \cdot \dot{\varphi}$, we obtain the following time derivatives of the coordinate vectors

$$
\begin{aligned}
& \dot{\hat{e}}_{x}=-\dot{\varphi} \hat{e}_{s}=-\frac{\dot{s}}{\rho} \hat{e}_{s} \\
& \dot{\hat{e}}_{s}=+\dot{\varphi} \hat{e}_{x}=+\frac{\dot{s}}{\rho} \hat{e}_{x} \\
& \dot{\hat{e}}_{y}=0
\end{aligned}
$$

and by using $\dot{x}=\frac{d x}{d s} \cdot \frac{d s}{d t}=x^{\prime} \cdot \dot{s}, \dot{y}=\frac{d y}{d s} \cdot \frac{d s}{d t}=y^{\prime} \cdot \dot{s}$, we obtain for the time derivatives:

$$
\begin{gathered}
\dot{\vec{r}}=x^{\prime} \dot{s} \hat{e}_{x}+y^{\prime} \dot{s} \hat{e}_{y}-\left(1+\frac{x}{\rho}\right) \dot{s} \hat{e}_{s} \\
\ddot{\vec{r}}=\{x^{\prime \prime} \dot{s}^{2}-\left(1+\frac{x}{\rho}\right) \frac{\dot{s}^{2}}{\rho}+\underbrace{x^{\prime} \ddot{s}}_{\approx 0}\} \hat{e}_{x}+\{y^{\prime \prime} \dot{s}^{2}+\underbrace{y^{\prime} \ddot{s}}_{\approx 0}\} \hat{e}_{y}+\{-2 x^{\prime} \frac{\dot{s}^{2}}{\rho}-\underbrace{\left(1+\frac{x}{\rho}\right)}_{\approx 0} \ddot{\sim}\} \hat{e}_{s}
\end{gathered}
$$

## Now proceed with several approximations:

- small displacements $x \ll \rho, y \ll \rho, \ddot{s} \approx 0$ (paraxial optics)
- only dipole and quadrupole magnets (linear field changes)
- design orbit lies in a plane (flat accelerator)
- no coupling between motion in hor. and vert. plane (upright magnets)
- small momentum deviations (quasi monochromatic beam)
- in general: no quadratic or higher order terms (linear beam optics)

Magnetic field in terms of strength parameters:

$$
\frac{q}{p_{0}} \vec{B}=k y \hat{e}_{x}+\left\{-\frac{1}{\rho}+k x\right\} \hat{e}_{y},
$$

and from simple geometrical considerations, we may write the particles longitudinal velocity v in terms of the change of the longitudinal coordinate $\dot{s}$ :

## CERN Accelerator School: Introductory Course

$$
\begin{gathered}
\dot{s}=\rho \dot{\varphi}, \quad \mathrm{v}_{s}=(\rho+x) \dot{\varphi} \\
\Rightarrow \quad \mathrm{v}_{s}=\frac{\rho+x}{\rho} \dot{s}=\left(1+\frac{x}{\rho}\right) \dot{s} \\
\mathrm{v}_{x}=x^{\prime} \dot{s}, \quad \mathrm{v}_{y}=y^{\prime} \dot{s}
\end{gathered}
$$

and for $p=\gamma m \cdot \mathrm{v}=p_{0}+\Delta p$ we have

$$
\frac{1}{p} \approx \frac{1}{p_{0}}+\left.\Delta p \frac{\partial(1 / p)}{\partial p}\right|_{p=p_{0}}=\frac{1}{p_{0}}\left(1-\frac{\Delta p}{p_{0}}\right)
$$



The particles are deflected due to the Lorentz force $\gamma m \cdot \ddot{\vec{r}}=q \cdot(\dot{\vec{r}} \times \vec{B})$, thus

$$
\left(\begin{array}{c}
x^{\prime \prime} \dot{s}^{2}-\left(1+\frac{x}{\rho}\right) \frac{\dot{s}^{2}}{\rho} \\
y^{\prime \prime} \dot{s}^{2} \\
-2 x^{\prime} \frac{\dot{s}^{2}}{\rho}
\end{array}\right)=\frac{q}{\gamma m} \cdot\left(\begin{array}{c}
\left(1+\frac{x}{\rho}\right) \dot{s} B_{y} \\
-\left(1+\frac{x}{\rho}\right) \dot{s} B_{x} \\
\dot{s}\left(x^{\prime} B_{y}-y^{\prime} B_{x}\right)
\end{array}\right)
$$

## CERN Accelerator School: Introductory Course

We will concentrate on the transverse planes. With the corresponding multipole strengths and the momentum expansion, we get

$$
\begin{aligned}
x^{\prime \prime}-\left(1+\frac{x}{\rho}\right) \frac{1}{\rho} & =\frac{q}{\dot{s}} \\
\underbrace{\left(\frac{\mathrm{v}}{p}\right)}_{=1 / \gamma m} & \left(1+\frac{x}{\rho}\right) \underbrace{\frac{p_{0}}{q}\left(k x-\frac{1}{\rho}\right)}_{=B_{z}}=\left(1+\frac{x}{\rho}\right)^{2}\left(k x-\frac{1}{\rho}\right)\left(1-\frac{\Delta p}{p_{0}}\right) \\
y^{\prime \prime} & =-\frac{q}{\dot{s}} \underbrace{\left(\frac{\mathrm{v}}{p}\right)}_{=1 / \gamma m}\left(1+\frac{x}{\rho}\right) \underbrace{\frac{p_{0}}{q} k y}_{=B_{x}}=-\left(1+\frac{x}{\rho}\right)^{2} k y\left(1-\frac{\Delta p}{p_{0}}\right)
\end{aligned}
$$

Neglecting all nonlinear terms in $x, y$, and $\Delta p / p_{0}$, we again obtain the (already well known) linear equations of motion:

$$
\begin{gathered}
x^{\prime \prime}(s)+\left(\frac{1}{\rho^{2}(s)}-k(s)\right) \cdot x(s)=\frac{1}{\rho(s)} \frac{\Delta p}{p} \\
y^{\prime \prime}(s)+k(s) \cdot y(s)=0
\end{gathered}
$$

## Remember:

Can be driven resonantly like a child's swing
$\leftrightarrow$ Parametric oscillator


$$
\omega=\sqrt{\frac{g}{l}} \quad \rightarrow \quad \ddot{\varphi}+\omega^{2}(t) \cdot \varphi=0
$$

### 3.4. Matrix formalism

We will characterize a particles state by a vector built from its relative coordinates:

$$
\left.\left.\left(\begin{array}{c}
x \\
x^{\prime} \\
y \\
y^{\prime}
\end{array}\right)=\left(\begin{array}{c}
\text { horizontal displacement } \\
\text { horizontal angular displacement } \\
\text { vertical displacement } \\
\text { vertical angular displacement }
\end{array}\right) \quad\right\} \begin{array}{l}
\text { hor. phase space } \\
\end{array}\right\} \text { vert. phase space }
$$

and use the matrix formalism to describe particles trajectories: $\vec{X}=\mathbf{M} \cdot \vec{X}_{0}$. In case of upright magnets there will be no coupling of the transverse planes and we can generally write:

$$
\mathbf{M}=\left(\begin{array}{cccc}
r_{11} & r_{12} & 0 & 0 \\
r_{21} & r_{22} & 0 & 0 \\
0 & 0 & r_{33} & r_{34} \\
0 & 0 & r_{43} & r_{44}
\end{array}\right)=\left(\begin{array}{cccc}
\left\langle x \mid x_{0}\right\rangle & \left\langle x \mid x_{0}{ }^{\prime}\right\rangle & 0 & 0 \\
\left\langle x^{\prime} \mid x_{0}\right\rangle & \left\langle x^{\prime} \mid x_{0}{ }^{\prime}\right\rangle & 0 & 0 \\
0 & 0 & \left\langle y \mid y_{0}\right\rangle & \left\langle y \mid y_{0}{ }^{\prime}\right\rangle \\
0 & 0 & \left\langle y^{\prime} \mid y_{0}\right\rangle & \left\langle y^{\prime} \mid y_{0}{ }^{\prime}\right\rangle
\end{array}\right)
$$

Next, we have to derive the matrices for drift, dipole and quadrupole magnets.

## CERN Accelerator School: Introductory Course

### 3.4.1. Drift space

$1 / \rho(s)=k(s)=0$ gives $\quad x^{\prime}(s)=x_{0}{ }^{\prime}=$ const.,$y^{\prime}(s)=y_{0}{ }^{\prime}=$ const.
Thus we get:

$$
\mathbf{M}_{\text {drift }}=\left(\begin{array}{cc}
\begin{array}{|cc|}
\hline 1 & L \\
0 & 1 \\
\hline
\end{array} & \begin{array}{cc}
0 & 0 \\
0 & 0 \\
0 & 0 \\
0 & 0
\end{array} \\
\begin{array}{|cc|}
\hline 1 & L \\
0 & 1 \\
\hline
\end{array}
\end{array}\right)
$$

### 3.4.2. Dipole magnets

Constant bending radius: $k=0$. Homogeneous solution (case $\Delta p / p=0$ ):

$$
x_{h}(s)=a \cdot \cos \frac{s}{\rho}+b \cdot \sin \frac{s}{\rho}
$$

The integration constants $a, b$ are derived from the boundary conditions at $s=0$

$$
x(s=0)=a=x_{0}, \quad x^{\prime}(s=0)=\frac{b}{\rho}=x_{0}^{\prime}
$$

## CERN Accelerator School: Introductory Course

and by defining the bending angle $\varphi=L / \rho$ of the dipole magnet, we obtain :

$$
\begin{gathered}
x(L)=x_{0} \cdot \cos \varphi+\rho \cdot x_{0}{ }^{\prime} \cdot \sin \varphi \\
y(L)=y_{0}+\rho \cdot \varphi \cdot y_{0}{ }^{\prime} \\
\mathbf{M}_{\text {dipole }}=\left(\begin{array}{cc}
\begin{array}{|cc|}
\hline \cos \varphi & \rho \sin \varphi \\
-1 / \rho \cdot \sin \varphi & \cos \varphi
\end{array} & \begin{array}{cc}
0 & 0 \\
0 & 0
\end{array} \\
\begin{array}{cc}
0 & 0 \\
0 & 0
\end{array} & \begin{array}{|cc|}
\hline & \rho \varphi \\
0 & 1 \\
\hline
\end{array}
\end{array}\right)
\end{gathered}
$$

A sector magnet is therefore focusing in the horizontal plane.
Sector- / rectangular dipole magnets and edge focusing:


Rectangular Magnet

The focusing / defocusing effect of the fringe fields (edge focusing) depends on the entrance (exit) angle $\psi$ and may again be described by a linear transformation matrix

$$
\mathbf{M}_{\psi}=\left(\begin{array}{cc}
\begin{array}{|cc|}
\hline 1 & 0 \\
\tan \psi / \rho & 1 \\
\hline
\end{array} & \begin{array}{cc}
0 & 0 \\
0 & 0
\end{array} \\
\begin{array}{cc}
0 & 0
\end{array} & \begin{array}{|cc|}
\hline 1 & 0 \\
-\tan \psi / \rho & 1
\end{array}
\end{array}\right)
$$

We finally obtain with $\psi=\varphi / 2$ and $\mathbf{M}_{\text {rect }}=\mathbf{M}_{\psi} \cdot \mathbf{M}_{\text {dipole }} \cdot \mathbf{M}_{\psi}$

$$
\mathbf{M}_{\text {rect }}=\left(\begin{array}{cc}
\begin{array}{|cc|}
\hline 1 & \rho \sin \varphi \\
0 & 1
\end{array} & \begin{array}{cc}
0 & 0 \\
0 & 0
\end{array} \\
\begin{array}{cc}
0 & 0 \\
0 & 0
\end{array} & \begin{array}{|cc|}
\hline 1-\rho \varphi / f & \rho \varphi \\
\rho \varphi / f^{2}-2 / f & 1-\rho \varphi / f
\end{array}
\end{array}\right)
$$

where we have defined the focal length $f \approx \rho / \tan \psi$ caused by edge (de)focusing.

## A rectangular dipole magnet is therefore focusing in the vertical plane. It acts like a drift space in the horizontal plane!

### 3.4.3. Quadrupole magnets

Assuming a pure quadrupole magnet we set the bending term $1 / \rho=0$. The solution of the equation of motion depends on the sign of the quadrupole strength $k$. For $k<0$ we get the solution of a quadrupole magnet, which is horizontal focusing and vertical defocusing (the case $k>0$ can be treated completely analog):

$$
\begin{gathered}
x(s)=a \cdot \cos (\sqrt{|k|} \cdot s)+b \cdot \sin (\sqrt{|k|} \cdot s) \\
y(s)=c \cdot \cosh (\sqrt{|k|} \cdot s)+d \cdot \sinh (\sqrt{|k|} \cdot s)
\end{gathered}
$$

The integration constants $a, b, c, d$ are derived from the boundary conditions at $s=0$ :

$$
\begin{array}{ll}
x(s=0)=a=x_{0}, & x^{\prime}(s=0)=b=x_{0}{ }^{\prime} \\
y(s=0)=c=y_{0}, & y^{\prime}(s=0)=d=y_{0}{ }^{\prime}
\end{array}
$$

Substituting and building the first derivative, we obtain the transformation matrices for a horizontal focusing (FQ) and a horizontal defocusing (DQ) quadrupole, where we put $\Omega=\sqrt{|k|} \cdot L$ with the quadrupole length $L$ and focal length $1 / f=k L$.

QF $(k<0)$ :

QD $(k>0)$ :

$$
\mathbf{M}_{Q D}=\left(\begin{array}{cc}
\begin{array}{|cc|}
\hline \cosh \Omega & \frac{1}{\sqrt{|k|}} \sinh \Omega \\
\sqrt{|k|} \sinh \Omega & \cosh \Omega
\end{array} & \begin{array}{cc}
0 & 0 \\
0 & 0
\end{array} \\
\begin{array}{cl}
0 & 0 \\
0 & 0
\end{array} & \begin{array}{|cc|}
\hline \cos \Omega & \frac{1}{\sqrt{|k|}} \sin \Omega \\
-\sqrt{|k|} \sin \Omega & \cos \Omega
\end{array}
\end{array}\right) \stackrel{\substack{L \rightarrow 0}}{ }\left(\begin{array}{ccc}
\begin{array}{|ccc}
1 & 0 \\
\frac{1}{f} & 1 \\
\hline
\end{array} & \begin{array}{cc}
0 & 0 \\
0 & 0
\end{array} \\
0 & 0 & \begin{array}{|cc|}
\hline 1 & 0 \\
0 & 0
\end{array} \\
\hline-\frac{1}{f} & 1 \\
\hline
\end{array}\right)
$$

### 3.4.4. Particle orbits in a system of magnets

With the derived matrixes particle trajectories may be calculated for any given arbitrary beam transport line by cutting this beam line into smaller uniform pieces so that $k=$ const. and $R=$ const. in each of these pieces:

Quadrupole

but:


### 3.5. Particle beams and phase space

### 3.3.1. Beam emittance

## Beam = statistical set of points in phase space!

Consider e.g. horizontal phase space, intensity distribution in $x, x^{\prime}$.
Choose origin of the coordinate axes $\hat{e}_{x}$ and $\hat{e}_{x^{\prime}}$ at the barycentre of the points:

$$
\bar{x}=\frac{1}{N} \sum_{i=1}^{N} x_{i}=0, \quad \overline{x^{\prime}}=\frac{1}{N} \sum_{i=1}^{N} x_{i}{ }^{\prime}=0
$$

Interested in variances (rms spread):

$$
\sigma_{x}{ }^{2}=\frac{1}{N} \sum_{i=1}^{N} x_{i}^{2}, \quad \sigma_{x^{\prime}}{ }^{2}=\frac{1}{N} \sum_{i=1}^{N} x_{i}^{\prime 2}
$$

System ( $X, X^{\prime}$ ) which is rotated by $\theta$ :

$$
\frac{\partial \sigma_{X}{ }^{2}}{\partial \theta}=\frac{\partial \sigma_{X}{ }^{2}}{\partial \theta}=0
$$



## CERN Accelerator School: Introductory Course

We will define the spread of the distribution, which is called the emittance $\varepsilon_{\mathrm{x}}$, by

$$
\varepsilon_{x}=\sigma_{X} \cdot \sigma_{X}^{\prime}=\sqrt{\overline{x^{2}} \cdot \overline{x^{\prime 2}}-\overline{x x^{\prime}}}
$$

It is important to note that this is a statistical definition of $\varepsilon$ !
More general, $\varepsilon$ will be defined over the area $\varepsilon=\int x^{\prime} \cdot d x$ !
The emittance can be considered as a statistical mean area:

$$
\varepsilon_{x}=\frac{1}{N} \sqrt{\frac{1}{2} \sum_{i=1}^{N} \sum_{j=1}^{N}\left(x_{i} x_{j}{ }^{\prime}-x_{j} x_{i}{ }^{\prime}\right)^{2}}=\frac{1}{N} \sqrt{2 \sum_{i=1}^{N} \sum_{j=1}^{N} A_{i j}{ }^{2}}
$$

(remember: $2 A_{\Delta}=|\vec{a} \times \vec{b}| \stackrel{a_{3}=b_{3}=0}{=}\left|a_{1} b_{2}-a_{2} b_{1}\right|$ ) where $A_{i j}$ is the area of the triangle $0 P_{i} P_{j}$ and $\varepsilon$ is a measure of the spread of the points around their barycentre.


## CERN Accelerator School: Introductory Course

The area of the "rms"-envelope-ellipse is just $\pi$ times the emittance $\varepsilon$

$$
A=\pi a b=\pi \sigma_{X} \sigma_{X^{\prime}}=\pi \varepsilon_{x}
$$

and its equation with respect to the axes $X$ and $X^{\prime}$ is

$$
\frac{X^{2}}{\sigma_{X}{ }^{2}}+\frac{X^{\prime 2}}{\sigma_{X}{ }^{2}}=1 \quad \Leftrightarrow \quad X^{2} \cdot \sigma_{X^{\prime}}{ }^{2}+X^{\prime 2} \cdot \sigma_{X}{ }^{2}=\varepsilon_{x}^{2}
$$

### 3.5.2. Twiss parameters

By an inverse rotation of angle $-\theta$ in phase space we obtain

$$
\varepsilon_{x}^{2}=x^{2} \cdot \sigma_{x^{\prime}}^{2}-2 x x^{\prime} \cdot \overline{x x^{\prime}}+x^{\prime 2} \cdot \sigma_{x}^{2}=x^{2} \cdot \sigma_{x^{\prime}}^{2}-2 x x^{\prime} \cdot r \sigma_{x} \sigma_{x^{\prime}}+x^{\prime 2} \cdot \sigma_{x}^{2}
$$

where we have defined the correlation coefficient

$$
r=\frac{\overline{x x^{\prime}}}{\sqrt{\overline{x^{2}} \cdot \overline{x^{\prime 2}}}}
$$

It is more or less obvious, that such a correlation term must exist in general.

## CERN Accelerator School: Introductory Course

We may define the so-called Twiss-parameters $\alpha_{x}, \beta_{x}$, and $\gamma_{x}$ such that

$$
\begin{gathered}
\sigma_{x}=\sqrt{\overline{x^{2}}}=\sqrt{\beta_{x} \varepsilon_{x}} \\
\sigma_{x^{\prime}}=\sqrt{\overline{x^{\prime 2}}}=\sqrt{\gamma_{x} \varepsilon_{x}} \\
r \sigma_{x} \sigma_{x^{\prime}}=\overline{x x^{\prime}}=-\alpha_{x} \varepsilon_{x}
\end{gathered}
$$

and the equation of the envelope-ellipse reads in the "conventional" form:

$$
\gamma_{x} x^{2}+2 \alpha_{x} x x^{\prime}+\beta_{x} x^{\prime 2}=\varepsilon_{x}
$$

All the above derived equations appear in identical form for the vertical plane, $x$ has only to be replaced by $y$. In the following, we will skip the index $x$ for reason of simplicity. Please note, that this doesn't imply that emittances and corresponding Twiss parameters are equal in both planes - they are not!

The meaning of the Twiss-parameters can be read off from the graphical representation of the envelope-ellipse:

## CERN Accelerator School: Introductory Course



- $\sqrt{\beta}$ represents the r.m.s. beam-envelope per unit emittance,
- $\sqrt{\gamma}$ represents the r.m.s. beam divergence per unit emittance,
- $\alpha$ is proportional to the correlation between $x$ and $x^{\prime}$.


## CERN Accelerator School: Introductory Course

### 3.5.3. Beta functions

In the following, we will first concentrate on the situation where $\Delta p / p=0$. With
$K_{x}(s)=1 / \rho^{2}(s)-k(s)$ and $K_{y}(s)=k(s)$ the equations of motion read

$$
x^{\prime \prime}(s)+K_{x}(s) \cdot x(s)=0, \quad y^{\prime \prime}(s)+K_{y}(s) \cdot y(s)=0
$$

They describe a transverse oscillation with position dependent amplitude and phase, which is called betatron oscillation. Both transverse planes can be treated similar!

We will therefore concentrate on $x$ and try to solve this equation, making the Ansatz

$$
x(s)=A \cdot u_{x}(s) \cdot \cos \left(\mu_{x}(s)+\varphi_{0}\right)
$$

( $A$ and $\varphi_{0}$ are integration constants, we will skip the index $x$ from now on) and obtain:

$$
\left[u^{\prime \prime}-u \cdot \mu^{\prime 2}+K \cdot u\right] \cdot \cos \left(\mu+\varphi_{0}\right)-\left[2 \cdot u^{\prime} \cdot \mu^{\prime}+u \cdot \mu^{\prime \prime}\right] \sin \left(\mu+\varphi_{0}\right)=0
$$

This relation is valid for any given phase $\mu(s)$ at any given position $s$, therefore

$$
\begin{aligned}
u^{\prime \prime}-u \cdot \mu^{\prime 2}+K \cdot u & =0 \\
2 \cdot u^{\prime} \cdot \mu^{\prime}+u \cdot \mu^{\prime \prime} & =0
\end{aligned}
$$

## CERN Accelerator School: Introductory Course

By integration of the second equation we obtain

$$
\mu(s)=\int_{0}^{s} \frac{d \tilde{s}}{u^{2}(\tilde{s})}
$$

and by using this relation $\quad u^{\prime \prime}-\frac{1}{u^{3}}+K \cdot u=0$.
With the definition of the beta function $\beta(s):=u^{2}(s)$ we derive for the amplitude and phase of the oscillation:

$$
\begin{gathered}
x(s)=A \cdot \sqrt{\beta(s)} \cdot \cos \left(\mu(s)+\varphi_{0}\right) \\
\mu(s)=\int_{0}^{s} \frac{d \tilde{s}}{\beta(\tilde{s})}
\end{gathered}
$$

Building the first derivative and defining $\alpha(s):=-\frac{\beta^{\prime}(s)}{2}$, we obtain

$$
x^{\prime}(s)=-\frac{A}{\sqrt{\beta(s)}}\left\{\alpha(s) \cdot \cos \left(\mu(s)+\varphi_{0}\right)+\sin \left(\mu(s)+\varphi_{0}\right)\right\}
$$

The equation for $x$ can be transformed to

$$
\cos ^{2}\left(\mu+\varphi_{0}\right)=\frac{x^{2}}{A^{2} \cdot \beta}
$$

which can be used in combination with the equation for $\mathrm{x}^{\prime}$ to obtain

$$
\sin ^{2}\left(\mu+\varphi_{0}\right)=\left(\frac{\sqrt{\beta}}{A} \cdot x^{\prime}+\frac{\alpha}{A \sqrt{\beta}} \cdot x\right)^{2}
$$

Using $\cos ^{2}+\sin ^{2}=1$ we derive

$$
\frac{x^{2}}{\beta(s)}+\left(\frac{\alpha(s)}{\sqrt{\beta(s)}} \cdot x+\sqrt{\beta(s)} \cdot x^{\prime}\right)^{2}=A^{2}
$$

which can be transformed by defining $\gamma(s):=\frac{1+\alpha^{2}(s)}{\beta(s)}$ to:

$$
\gamma x^{2}+2 \alpha x x^{\prime}+\beta x^{\prime 2}=A^{2}, \quad \text { where } \quad \frac{1}{\beta(s)}=\mu^{\prime}(s), \quad \alpha=-\frac{\beta^{\prime}}{2}, \quad \gamma=\frac{1+\alpha^{2}}{\beta}
$$

## Note:

Each particle will stay on its own ellipse, which will enclose a constant area in phase space $A$. The amplitude factor $A$ represents the Courant Snyder invariant! The shape of the ellipse is determined by the Twiss parameters $\alpha, \beta, \gamma$ and will change along the magneto-optics system, its area will stay always constant (Rem.: in case of conservative forces and no acceleration). The shape (not the size) of all single particle ellipses are determined by the same Twiss parameters!


## CERN Accelerator School: Introductory Course

### 3.5.4. Transformation in phase space

According to Liouville's theorem, all particles enclosed by an envelope ellipse will stay within that ellipse. The transformation of the horizontal and vertical ellipse parameters along the beam line may be derived from the transport matrixes in the horizontal and vertical plane. Starting at $s=0$, we have for a particle on this ellipse

$$
\gamma_{0} x_{0}{ }^{2}+2 \alpha_{0} x_{0} x_{0}{ }^{\prime}+\beta_{0} x_{0}{ }^{\prime 2}=\varepsilon=\gamma x^{2}+2 \alpha x x^{\prime}+\beta x^{\prime 2}
$$

Any particle trajectory starting at $s=0$ transforms to $s \neq 0$ by

$$
\binom{x}{x^{\prime}}=\left(\begin{array}{ll}
r_{11} & r_{12} \\
r_{21} & r_{22}
\end{array}\right) \cdot\binom{x_{0}}{x_{0}^{\prime}}=\left(\begin{array}{ll}
C(s) & S(s) \\
C^{\prime}(s) & S^{\prime}(s)
\end{array}\right) \cdot\binom{x_{0}}{x_{0}^{\prime}}
$$

which gives for the transformed ellipse equation via

$$
\binom{x_{0}}{x_{0}^{\prime}}=\underbrace{\frac{1}{C S^{\prime}-C^{\prime} S} \cdot\left(\begin{array}{cc}
S^{\prime}(s) & -S(s) \\
-C^{\prime}(s) & C(s)
\end{array}\right)}_{=\mathbf{M}^{-1}} \cdot\binom{x}{x^{\prime}} \stackrel{|\mathbf{M}|=1}{=}\binom{S^{\prime} x-S x^{\prime}}{-C^{\prime} x+C x^{\prime}}
$$

and $\quad x_{0}{ }^{2}=S^{\prime 2} x^{2}-2 S S^{\prime} x x^{\prime}+S^{2} x^{\prime 2}, \quad x_{0}{ }^{2}=C^{\prime 2} x^{2}-2 C C^{\prime} x x^{\prime}+C^{2} x^{\prime 2}, \quad x_{0} x_{0}^{\prime}=\ldots$

$$
\begin{aligned}
& \underbrace{\left(S^{\prime 2} \cdot \gamma_{0}-2 S^{\prime} C^{\prime} \cdot \alpha_{0}+C^{\prime 2} \cdot \beta_{0}\right)}_{=\gamma} \cdot x^{2}+2 \underbrace{\left(-S S^{\prime} \cdot \gamma_{0}+\left(S^{\prime} C+S C^{\prime}\right) \cdot \alpha_{0}-C C^{\prime} \cdot \beta_{0}\right)}_{=\alpha} \cdot x x^{\prime} \\
& +\underbrace{\left(S^{2} \cdot \gamma_{0}-2 S C \cdot \alpha_{0}+C^{2} \cdot \beta_{0}\right)}_{=\beta} \cdot x^{\prime 2}=\varepsilon
\end{aligned}
$$

This gives the transformation of the beam parameters in matrix formulation

$$
\left(\begin{array}{l}
\beta \\
\alpha \\
\gamma
\end{array}\right)=\left(\begin{array}{ccc}
C^{2} & -2 S C & S^{2} \\
-C C^{\prime} & S^{\prime} C+S C^{\prime} & -S S^{\prime} \\
C^{\prime 2} & -2 S^{\prime} C^{\prime} & S^{\prime 2}
\end{array}\right) \cdot\left(\begin{array}{l}
\beta_{0} \\
\alpha_{0} \\
\gamma_{0}
\end{array}\right)
$$

Another useful relation may be obtained by defining the Beta matrix B

$$
\mathbf{B} \equiv\left(\begin{array}{cc}
\beta & -\alpha \\
-\alpha & \gamma
\end{array}\right), \quad|\mathbf{B}|=\beta \gamma-\alpha^{2}=1, \quad \varepsilon \cdot \mathbf{B}=\left(\begin{array}{cc}
\sigma_{x}^{2} & \sigma_{x x^{\prime}} \\
\sigma_{x x^{\prime}} & \sigma_{x^{\prime}}^{2}
\end{array}\right) \equiv \boldsymbol{\Sigma}
$$

The equation of the envelope-ellipse can be transformed to:

$$
\varepsilon={ }^{T} \vec{X}_{0} \cdot \mathbf{B}_{0}{ }^{-1} \cdot \vec{X}_{0}={ }^{T} \vec{X}_{1} \cdot \mathbf{B}_{1}{ }^{-1} \cdot \vec{X}_{1}
$$

where the inverse of the Beta matrix is

## CERN Accelerator School: Introductory Course

$$
\mathbf{B}^{-1}=\left(\begin{array}{ll}
\gamma & \alpha \\
\alpha & \beta
\end{array}\right)
$$

and displacement-vector $\vec{X}$ transforms according to

$$
\vec{X}_{1}=\mathbf{M} \cdot \vec{X}_{0}, \quad{ }^{T} \vec{X}_{1}={ }^{T}\left(\mathbf{M} \cdot \vec{X}_{0}\right)={ }^{T} \vec{X}_{0} \cdot{ }^{T} \mathbf{M}
$$

By inserting $\mathbf{1}=\mathbf{M}^{-1} \cdot \mathbf{M}$, we obtain:

$$
\begin{aligned}
\varepsilon & ={ }^{T} \vec{X}_{0} \cdot{ }^{T} \mathbf{M} \cdot{ }^{T} \mathbf{M}^{-1} \cdot \mathbf{B}_{0}^{-1} \cdot \mathbf{M}^{-1} \cdot \mathbf{M} \cdot \vec{X}_{0} \\
& ={ }^{T}\left(\mathbf{M} \cdot \vec{X}_{0}\right) \cdot\left({ }^{T} \mathbf{M}^{-1} \cdot \mathbf{B}_{0}{ }^{-1} \cdot \mathbf{M}^{-1}\right) \cdot\left(\mathbf{M} \cdot \vec{X}_{0}\right) \\
& ={ }^{T} \vec{X}_{1} \cdot\left(\mathbf{M} \cdot \mathbf{B}_{0} \cdot{ }^{T} \mathbf{M}\right)^{-1} \cdot \vec{X}_{1}
\end{aligned}
$$

and we can read off the transformation of the Beta matrix:

$$
\mathbf{B}_{1}=\mathbf{M} \cdot \mathbf{B}_{0} \cdot{ }^{T} \mathbf{M}
$$

This can e.g. be used to derive the beta-function around a symmetry-point of a trans-fer-line where $\alpha=0$ in a simple way:

## CERN Accelerator School: Introductory Course

$$
\mathbf{B}_{1}(s)=\left(\begin{array}{ll}
1 & s \\
0 & 1
\end{array}\right) \cdot\left(\begin{array}{cc}
\beta_{s y m} & 0 \\
0 & 1 / \beta_{s y m}
\end{array}\right) \cdot\left(\begin{array}{ll}
1 & 0 \\
s & 1
\end{array}\right)=\left(\begin{array}{cc}
\beta_{s y m}+\frac{s^{2}}{\beta_{s y m}} & \frac{s}{\beta_{s y m}} \\
\frac{s}{\beta_{s y m}} & \frac{1}{\beta_{s y m}}
\end{array}\right)
$$

This gives the relations for the beam parameters around a symmetry-point:

$$
\begin{aligned}
& \beta(s)=\beta_{s y m}+\frac{s^{2}}{\beta_{s y m}} \\
& \alpha(s)=-\frac{s}{\beta_{s y m}} \\
& \gamma(s)=\frac{1}{\beta_{s y m}}
\end{aligned}
$$

The corresponding beam size scales with

$$
\sigma_{x}=\sqrt{\varepsilon \cdot \beta(s)}!
$$

Remember: $\quad \sigma_{x}=\sqrt{\varepsilon \cdot \beta(s)}, \quad \sigma_{x}{ }^{\prime}=\sqrt{\varepsilon \cdot \gamma(s)}, \quad$ and therewith:

## CERN Accelerator School: Introductory Course

$$
\sigma(s)=\sigma_{0} \cdot \sqrt{1+\left(\frac{s}{\beta_{0}}\right)^{2}}, \quad \sigma^{\prime}(s)=\frac{\varepsilon}{\sigma_{0}}=\text { const. }
$$

To obtain further insights, we will compare the particle's beam with a Gaussian light beam (TEM ${ }_{00}$ ), characterized by its waste radius $w(\mathrm{~s})$ and Rayleigh length $z_{R}$, in which $w$ is doubled. From diffraction theory, we know (from diffraction integrals):

$$
w(s)=w_{0} \cdot \sqrt{1+\left(\frac{s}{z_{R}}\right)^{2}}, \quad z_{R}=\frac{\pi w_{0}{ }^{2}}{\lambda}, \quad \theta_{\max }=\frac{\lambda}{\pi w_{0}}, \quad I(x, y)=I_{\max } \cdot\left(\frac{w_{0}}{w}\right)^{2} \cdot e^{\frac{-2\left(x^{2}+y^{2}\right)}{w^{2}}}
$$

This indicates:

$$
\beta_{0}=z_{R}=\frac{\pi w_{0}^{2}}{\lambda}, \quad \text { and from } \quad \sigma_{x}^{2}=\frac{\iint x^{2} I(x, y) \cdot d x d y}{\iint I(x, y) \cdot d x d y}=\frac{w^{2}}{4}
$$

and replacing $w=2 \sigma=2 \sqrt{\varepsilon \beta}$ we obtain the important relation:

$$
\Rightarrow \quad 4 \pi \cdot \varepsilon=\lambda
$$

## CERN Accelerator School: Introductory Course




## CERN Accelerator School: Introductory Course

The transformation matrix $\mathbf{M}$ can be derived also from the Twiss parameters. With

$$
\begin{gathered}
x(s)=\sqrt{\varepsilon \beta} \cos \left(\mu+\varphi_{0}\right)=\sqrt{\varepsilon} \cdot \sqrt{\beta} \cdot\left\{\cos \mu \cdot \cos \varphi_{0}-\sin \mu \cdot \sin \varphi_{0}\right\} \\
x^{\prime}(s)=-\frac{\sqrt{\varepsilon}}{\sqrt{\beta}} \cdot\left\{\alpha \cdot\left[\cos \mu \cdot \cos \varphi_{0}-\sin \mu \cdot \sin \varphi_{0}\right]-\sin \mu \cdot \cos \varphi_{0}+\cos \mu \cdot \sin \varphi_{0}\right\}
\end{gathered}
$$

and the starting conditions $x(0)=x_{0}, x^{\prime}(0)=x_{0}{ }^{\prime}, \mu(0)=0$, which transform to

$$
\begin{aligned}
& \cos \varphi_{0}=x_{0} / \sqrt{\varepsilon \beta_{0}} \\
& \sin \varphi_{0}=-1 / \sqrt{\varepsilon}\left(x_{0}^{\prime} \sqrt{\beta_{0}}+\alpha_{0} x_{0} / \sqrt{\beta_{0}}\right)
\end{aligned}
$$

$$
\mathbf{M}(s)=\left(\begin{array}{cc}
\frac{\sqrt{\beta}}{\sqrt{\beta_{0}}}\left(\cos \mu+\alpha_{0} \sin \mu\right) & \sqrt{\beta \beta_{0}} \sin \mu \\
\frac{\alpha_{0}-\alpha}{\sqrt{\beta \beta_{0}}} \cos \mu-\frac{1+\alpha \alpha_{0}}{\sqrt{\beta \beta_{0}}} \sin \mu & \frac{\sqrt{\beta}}{\sqrt{\beta_{0}}}(\cos \mu-\alpha \sin \mu)
\end{array}\right)
$$

## CERN Accelerator School: Introductory Course



## 4. Circular Accelerators

### 4.1. Weak focusing

Beam stability: transverse focusing in both planes!
Equation of motion:

$$
\begin{gathered}
x^{\prime \prime}(s)+\underbrace{\left(\frac{1}{\rho^{2}(s)}-k(s)\right)}_{>0} \cdot x(s)=0 \\
y^{\prime \prime}(s)+\overbrace{k(s)} \cdot y(s)=0
\end{gathered}
$$

Idea: horizontally defocusing $k$ is overcompensated by geometrical focusing!

$$
0<k=-\frac{q}{p} \frac{\partial B_{y}}{\partial x}<\frac{1}{\rho^{2}}
$$

With $p=q \rho B_{0}$, where $B_{0}$ defines the bending field at the design orbit, one obtains

$$
0<n=-\frac{\rho}{B_{0}} \frac{\partial B_{y}}{\partial x}<1 \quad(\text { Steenbeck 1924) }
$$

## CERN Accelerator School: Introductory Course

where we have defined the field index $n$ to

$$
n=k \rho^{2}=-\frac{\rho}{B_{0}} \frac{\partial B_{y}}{\partial \rho} \quad \rightarrow \quad B(r)=B_{0} \cdot\left(\frac{r}{\rho}\right)^{-n}
$$

Thus, a circular accelerator like a synchrotron has to be made of dipole magnets with radially decreasing bending field strength fulling the above derived weak focusing condition.

Particles will oscillate around the reference trajectory with the spatial frequency

$$
\omega_{x}=\sqrt{\frac{1}{\rho^{2}}-k}=\frac{\sqrt{1-n}}{\rho}, \quad \omega_{y}=\sqrt{k}=\frac{\sqrt{n}}{\rho}
$$

The number $Q$ of oscillations per turn of length $L=2 \pi \rho$ will then be

$$
Q_{x}=\frac{1}{2 \pi} \oint \frac{d s}{\beta_{x}}=\sqrt{1-n}<1, \quad Q_{y}=\frac{1}{2 \pi} \oint \frac{d s}{\beta_{y}}=\sqrt{n}<1
$$

## Problem:

> We derive for the constant beta functions $\beta_{x, y}>\rho$
> $\rightarrow$ beam size $\sigma=\sqrt{\varepsilon \beta}$ will increase remarkably with increasing radius!

## CERN Accelerator School: Introductory Course



## CERN Accelerator School: Introductory Course

### 4.2. Strong focusing

Focusing in both planes possible in case of alternating gradient - well know from light optics:


## Magnet optics:

Simplest configuration: FODO lattice, periodic arrangement of identical structures


## CERN Accelerator School: Introductory Course



### 4.2.1. Stability criterion

If $\mathbf{M}(L)$ is the transformation matrix for one periodic cell we will have for $N$ cells:

$$
\mathbf{M}(N \cdot L)=[\mathbf{M}(L)]^{N}
$$

For a full lattice period, we take use of Floquet's theorem. Recalling the equations of motions

$$
\begin{array}{ll}
x^{\prime \prime}(s)+K_{x}(s) \cdot x(s)=0 & \text { with } \quad K_{x}(s)=1 / \rho^{2}(s)-k(s) \\
y^{\prime \prime}(s)+K_{y}(s) \cdot y(s)=0 & \text { with } \quad K_{y}(s)=k(s)
\end{array}
$$

it states (Gaston Floquet, 1847 - 1920) for e.g. $x(s)=A \sqrt{\beta_{x}(s)} \cos \left(\mu_{x}(s)+\varphi_{0}\right)$
If $K(s)$ is periodic, the amplitude function (and therefore $\beta(s)$ ) is periodic as well. In this case we call the DGL Hill's equation (George William Hill 1838 - 1914).

Please note and take care:
Floquet's theorem doesn't state that $\mu(\mathrm{s})$ and therewith $x(\mathrm{~s}), y(\mathrm{~s})$ are periodic as well! This would be an exception! (catastrophic, as we will see later)

Thus we recommend periodic boundary conditions $\beta=\beta_{0}, \alpha=\alpha_{0}$ and obtain, using the Twiss parameter representation of the transfer matrix:

$$
\mathbf{M}=\left(\begin{array}{cc}
\cos \mu+\alpha_{0} \sin \mu & \beta_{0} \sin \mu \\
-\gamma_{0} \sin \mu & \cos \mu-\alpha_{0} \sin \mu
\end{array}\right)
$$

This matrix was first derived by Twiss from general mathematics principles and is called the Twiss matrix (Richard Q. Twiss, 1920 - 2005).

We calculate its eigenvalues from

$$
|\mathbf{M}-\lambda \cdot \mathbf{I}|=\lambda^{2}-\operatorname{Tr}\{\mathbf{M}\} \cdot \lambda+1=0
$$

With $\operatorname{Tr}\{\mathbf{M}\}=2 \cdot \cos \mu$ we obtain

$$
\lambda_{1,2}=\cos \mu \pm i \sin \mu=e^{ \pm i \mu}
$$

We require that the eigenvalues remain finite thus requiring a real betatron phase $\mu$.
This is guaranteed when $|\cos \mu| \leq 1$ and leads to the general stability condition:

$$
|\operatorname{Tr}\{\mathbf{M}\}|=\left|r_{11}+r_{22}\right| \leq 2
$$

## CERN Accelerator School: Introductory Course

And now comes the "clou": Rewriting the Twiss matrix using

$$
\mathbf{J}=\left(\begin{array}{cc}
\alpha & \beta \\
-\gamma & -\alpha
\end{array}\right), \quad \mathbf{J}^{2}=\left(\begin{array}{cc}
-1 & 0 \\
0 & -1
\end{array}\right)=-\mathbf{I}
$$

it can be expressed by

$$
\mathbf{M}=\mathbf{I} \cdot \cos \mu+\mathbf{J} \cdot \sin \mu
$$

Similar to Moivre's formula we get for $N$ equal periods

$$
\mathbf{M}^{N}=(\mathbf{I} \cdot \cos \mu+\mathbf{J} \cdot \sin \mu)^{N}=\mathbf{I} \cdot \cos (N \mu)+\mathbf{J} \cdot \sin (N \mu)
$$

and

$$
\left|\operatorname{Tr}\left\{\mathbf{M}^{N}\right\}\right|=|2 \cdot \cos (N \mu)| \leq 2
$$

Conclusion:
In case of a real betatron phase advance $\mu$, the beam size in a circular accelerator will remain finite (the 100 Mio $\$$ question in the 50 's!). This can easily be proofed by calculating the trace of the one turn matrix:

$$
\operatorname{Tr}\{\mathbf{M}\} \mid \leq 2
$$

### 4.3. Periodic focusing systems

### 4.3.1. General FODO lattice



The FODO geometry can be expressed symbolically by the sequence

$$
\underbrace{1 / 2 \mathrm{QF}, \mathrm{D}, 1 / 2 \mathrm{QD}}_{=\mathbf{M}_{-1 / 2}}, \underbrace{1 / 2 \mathrm{QD}, \mathrm{D}, 1 / 2 \mathrm{QF}}_{=\mathbf{M}_{1 / 2}}
$$

It is sufficient to use the thin lens approximation ( $l_{Q} \ll f$ ). We will set the focal lengths to $f_{2}=2 f_{D}, f_{1}=2 f_{F}$, the drift length to $L$. Defining

$$
1 / f^{*}=1 / f_{1}+1 / f_{2}-L /\left(f_{1} \cdot f_{2}\right)
$$

the transformation matrix of half a FODO cell is

$$
\mathbf{M}_{1 / 2}=\left(\begin{array}{cc}
1 & 0 \\
-1 / f_{2} & 1
\end{array}\right) \cdot\left(\begin{array}{cc}
1 & L \\
0 & 1
\end{array}\right) \cdot\left(\begin{array}{cc}
1 & 0 \\
-1 / f_{1} & 1
\end{array}\right)=\left(\begin{array}{cc}
1-L / f_{1} & L \\
-1 / f^{*} & 1-L / f_{2}
\end{array}\right)
$$

Multiplication with the reverse matrix gives

$$
\mathbf{M}_{\mathrm{FODO}}=\left(\begin{array}{cc}
1-2 L / f^{*} & 2 L \cdot\left(1-L / f_{2}\right) \\
-2 / f^{*} \cdot\left(1-L / f_{1}\right) & 1-2 L / f^{*}
\end{array}\right) \text { and } \quad|\operatorname{Tr}\{\mathbf{M}\}|=\left|2-\frac{4 L}{f^{*}}\right|<2
$$

This is equivalent to $0<\frac{L}{f^{*}}<1$, and defining $u=\frac{L}{f_{1}}, v=\frac{L}{f_{2}}$ we get

$$
0<u+v-u \cdot v<1
$$

from which we derive the boundaries of the stability region

$$
\begin{aligned}
& \left|u_{1}\right|=1, \quad\left|v_{2}\right|=\frac{|u|}{1-|u|} \\
& \left|v_{1}\right|=1, \quad\left|v_{3}\right|=\frac{|u|}{1+|u|}
\end{aligned}
$$

## CERN Accelerator School: Introductory Course

which gives the famous necktie-diagram for thin lens approximation:


In the simple case of equal focusing strengths, we arrive at

$$
\left|f_{1}\right|=\left|f_{2}\right|=\frac{\left|f_{D}\right|}{2}=\frac{\left|f_{F}\right|}{2}=\frac{|f|}{2} \quad \rightarrow \quad\left|\frac{2 L}{f}\right|=\left|\frac{L_{\mathrm{FODO}}}{f}\right|<1
$$

## CERN Accelerator School: Introductory Course

## LHC: Lattice Design

the ARC $90^{\circ}$ FoDo in both planes

equipped with additional corrector coils
MB: main dipole
MQ: main quadrupole
MQT: Trim quadrupole
MQS: Skew trim quadrupole
MO: Lattice octupole (Landau damping)
MSCB: Skew sextupole
Orbit corrector dipoles
MCS: Spool piece sextupole
MCDO: Spool piece 8 / 10 pole
BPM: Beam position monitor + diagnostics
© Bernhard Holzer

## CERN Accelerator School: Introductory Course

### 4.3.2. Periodic beta functions

Periodic solutions of a periodic lattice of period-length $L$ will be

$$
\begin{aligned}
& \beta\left(s_{0}+L\right)=\beta\left(s_{0}\right)=\beta_{0} \\
& \alpha\left(s_{0}+L\right)=\alpha\left(s_{0}\right)=\alpha_{0}
\end{aligned}
$$

Comparing the transfer matrix for one period with its Twiss parameter representation

$$
\mathbf{M}=\left(\begin{array}{ll}
r_{11} & r_{12} \\
r_{21} & r_{22}
\end{array}\right)=\left(\begin{array}{cc}
\cos \mu+\alpha_{0} \sin \mu & \beta_{0} \sin \mu \\
-\gamma_{0} \sin \mu & \cos \mu-\alpha_{0} \sin \mu
\end{array}\right)
$$

we can determine the Twiss parameters at the symmetry points (where $\alpha=0$ !)

$$
\alpha_{0}=0, \quad \beta_{0}=\frac{r_{12}}{\sqrt{1-r_{11}^{2}}}, \quad \gamma_{0}=\frac{-r_{21}}{\sqrt{1-r_{11}^{2}}}, \quad \cos \mu=r_{11}
$$

and transform them to any position $s$ using e.g. the beta matrix formalism

$$
\left(\begin{array}{cc}
\beta & -\alpha \\
-\alpha & \gamma
\end{array}\right)=\mathbf{M}\left(s, s_{0}\right) \cdot\left(\begin{array}{cc}
\beta_{0} & 0 \\
0 & \gamma_{0}
\end{array}\right) \cdot{ }^{T} \mathbf{M}\left(s, s_{0}\right)
$$

thus revealing the development of $\beta(\mathrm{s}), \alpha(\mathrm{s}), \gamma(\mathrm{s})$. Minimum $\langle\beta\rangle$ for $\mu \approx 90^{\circ}$ !

## CERN Accelerator School: Introductory Course

Example: simple model toy ring (taken from Wille):


Choosing $\left|k_{Q F}\right|=\left|k_{Q D}\right|=1.20 \mathrm{~m}$, we can calculate the transfer matrix $M$ and extract the Twiss parameters, obtaining:


## CERN Accelerator School: Introductory Course

### 4.4. Transverse beam dynamics

### 4.4.1. Closed orbit

Remember: In circular accelerators the amplitude function is periodic according to Floquet's theorem and reproduces itself after one turn.

This implies, that the charge center of the beam also moves on a closed trajectory, which is called the closed orbit! The shape of the closed orbit is determined by the magnets and can - due to errors and misalignments - significantly deviate from the design orbit!

Dedicated steerer magnets (small dipoles), which have to be installed around the ring, are used to correct closed orbit deviations.


## CERN Accelerator School: Introductory Course

### 4.4.2. Betatron tune

The betatron tune $Q$ is defined as the number of oscillations per revolution:

$$
Q_{x, y}=\frac{\mu(L)}{2 \pi}=\frac{1}{2 \pi} \cdot \oint \frac{d s}{\beta_{x, y}(s)}
$$

If one regards the phase space at an arbitrarily chosen point, a single particle moves on its phase space ellipse.

The points represents the parameters after $1,2, \ldots 5$ revolutions.


The betatron tune is one of the most important parameter in circular accelerators!

## CERN Accelerator School: Introductory Course

### 4.4.2. Filamentation

If the envelope ellipse $\sigma_{b}$ of the beam is not matched to the ellipse $\sigma_{m}$ of the periodic lattice, it will start to rotate with a phase advance per revolution of $2 \pi Q$


Due to effects of higher order the quadrupole strengths and therefore the phase advance depends on the amplitude (horizontal and vertical displacements). In case of mismatch, the beam phase space distribution starts to filament. After a large number of revolutions, the distribution may be surrounded by a large ellipse of the form of the lattice ellipse.

## CERN Accelerator School: Introductory Course

Example for an unmatched and matched beam (taken from B. Schmidt):

after 20 turns


### 4.4.3. Normalized coordinates

It is useful to transform the oscillatory solution with varying amplitude and frequency $x(s)=\sqrt{\varepsilon_{x} \beta_{x}(s)} \cdot \cos \left(\mu_{x}(s)+\varphi_{0}\right)$ to a solution which looks exactly like that of a harmonic oscillator. We introduce Floquet's coordinates through the transformation:

$$
\begin{aligned}
\psi(s) & =\frac{\mu(s)}{Q} \\
\eta_{y}\left(\psi_{x}\right) & =\frac{x(s)}{\sqrt{\beta_{x}(s)}}
\end{aligned}
$$

The angle $\psi$ advances by $2 \pi$ every revolution. It coincides with $\theta$ at each $\beta^{\text {max }}$ and $\beta^{\text {min }}$ location and does not depart very much from $\theta$ in between.

Using these normalized coordinates, the equation of motion can be transformed to

$$
\frac{d^{2} \eta}{d \psi^{2}}+Q^{2} \eta=0
$$

## CERN Accelerator School: Introductory Course

The solution transforms to

$$
\eta(\psi)=\eta_{0} \cdot \cos (Q \psi+\lambda)
$$

and the phase space ellipse becomes an invariant cycle of radius $\eta_{0}$.
The use of normalized coordinates is convenient in the discussions of perturbations and aberrations:

### 4.4.4. Closed orbit distortions

Let us assume a dipole field error produced by a short dipole which makes a constant angular kick in divergence (from $l=r \cdot \varphi \approx \frac{p}{q(\delta B)} \cdot \delta x^{\prime}=\frac{\rho B}{\delta B} \cdot \delta x^{\prime}$ )

$$
\delta x^{\prime}=\frac{\delta(B l)}{B \rho}
$$

which perturbs the orbit trajectory which elsewhere obeys


$$
\frac{d^{2} \eta}{d \psi^{2}}+Q^{2} \eta=0, \text { with } \quad \eta=\eta_{0} \cos (Q \psi+\lambda)
$$

We choose $\psi=0$ to be diametrically opposite to the kick. Then by symmetry $\lambda=0$ and the disturbed orbit oscillates around the ideal path


Differentiation gives $\frac{d \eta}{d \psi}=-\eta_{0} Q \cdot \sin (Q \psi)=-\eta_{0} Q \cdot \sin (\pi Q)$ at $\psi=\pi$.
With $\frac{d \psi}{d s}=\frac{1}{Q \beta_{0}}$ and $\delta \eta^{\prime}=\frac{1}{\sqrt{\beta_{0}}} \cdot \delta x^{\prime}$, we may relate $\eta_{0}$ to the real kick by

$$
\frac{\delta x^{\prime}}{2}=\sqrt{\beta_{0}} \cdot \frac{\delta \eta^{\prime}}{2}=-\sqrt{\beta_{0}} \cdot \frac{d \eta}{d s}=-\sqrt{\beta_{0}} \cdot \frac{d \eta}{d \psi} \cdot \frac{d \psi}{d s}=\frac{\eta_{0}}{\sqrt{\beta_{0}}} \cdot \sin (\pi Q)
$$

giving

$$
\eta_{0}=\frac{\sqrt{\beta_{0}}}{2 \sin (\pi Q)} \delta x^{\prime}
$$

position s for a field error at $s_{0}$ with $x=\sqrt{\beta} \cdot \eta$ and $\mu(s)-\mu\left(s_{0}\right)+Q \pi=Q \cdot \psi$ :

$$
x_{c}(s)=\sqrt{\beta} \eta_{0} \cos (Q \psi)=\underbrace{\left[\frac{\sqrt{\beta(s) \beta\left(s_{0}\right)}}{2 \sin (\pi Q)} \frac{\delta(B l)}{B \rho}\right]}_{=\text {amplitude at position } s} \cdot \cos \left(\mu(s)-\mu\left(s_{0}\right)+Q \pi\right)
$$

The effect of a random distribution of dipole errors can be estimated from the r.m.s. average, weighted according to the $\beta_{0}$ values of the kicks $\delta x_{i}$ :

$$
x_{c}(s)=\frac{\sqrt{\beta(s)}}{2 \sin (\pi Q)} \cdot \oint_{s} \sqrt{\beta\left(s_{0}\right)} \cdot \frac{\delta B l}{B \rho} \cdot \cos \left(\mu(s)-\mu\left(s_{0}\right)+Q \pi\right) \cdot d s_{0}
$$

Using matrix algebra, the displacement of the closed orbit at the position of the field error can be calculated from the displacement just before and after the kick element:

## CERN Accelerator School: Introductory Course

$$
\binom{x_{c, 0}}{x_{c, 0} 0^{\prime}-\delta x^{\prime}}=\mathbf{M} \cdot\binom{x_{c, 0}}{x_{c, 0}{ }^{\prime}}=\left(\begin{array}{cc}
\cos \mu+\alpha_{0} \sin \mu & \beta_{0} \sin \mu \\
-\gamma_{0} \sin \mu & \cos \mu-\alpha_{0} \sin \mu
\end{array}\right) \cdot\binom{x_{c, 0}}{x_{c, 0}{ }^{\prime}}
$$

with $\mu=2 \pi Q$, giving
$x_{c, 0}=\frac{\beta_{0} \delta x^{\prime}}{2 \sin (\pi Q)} \cos (\pi Q)$
$x_{c, 0}{ }^{\prime}=\frac{\delta x^{\prime}}{2 \sin (\pi Q)}\left[\sin (\pi Q)-\alpha_{0} \cos (\pi Q)\right]$


The closed orbit displacement $x_{c}(s)$ is calculated from $\vec{x}_{c}(s)=\mathbf{M}\left(s_{0}, s\right) \cdot \vec{x}_{c, 0}$ :

$$
\binom{x_{c}(s)}{x_{c}^{\prime}(s)}=\left(\begin{array}{cc}
\sqrt{\frac{\beta(s)}{\beta_{0}}}\left(\cos \mu+\alpha_{0} \sin \mu\right) & \sqrt{\beta(s) \beta_{0}} \sin \mu \\
-\frac{1+\alpha(s) \alpha_{0}}{\sqrt{\beta(s) \beta_{0}}} \sin \mu+\frac{1-\alpha(s) \alpha_{0}}{\sqrt{\beta(s) \beta_{0}}} \cos \mu & \sqrt{\frac{\beta(s)}{\beta_{0}}}\left(\cos \mu-\alpha_{0} \sin \mu\right)
\end{array}\right) \cdot\binom{x_{c, 0}}{x_{c, 0}^{\prime}}
$$

$$
x_{c}(s)=\sqrt{\beta} \eta_{0} \cos (Q \psi)=\frac{\sqrt{\beta(s) \beta\left(s_{0}\right)}}{2 \sin (\pi Q)} \frac{\delta(B l)}{B \rho} \cdot \cos \left(\mu(s)-\mu\left(s_{0}\right)+Q \pi\right)
$$

Closed orbit distortions: uncorrected and corrected


## CERN Accelerator School: Introductory Course

## LHC Operation: Beam Commissioning

First turn steering "by sector:"
aOne beam at the time
aBeam through 1 sector ( $1 / 8$ ring),
correct trajectory, open collimator and move on.

YASP DV LHCRING / INJ-TEST-NB / beam 2


## 

FT - P $450.12 \mathrm{GeV} / \mathrm{C}$ - Fill \# 830 INJPROT - 10/09/08 15-01-58
85



## Dipole error and integer tune:



## CERN Accelerator School: Introductory Course

## Quadrupole error and half integer tune:



### 4.4.5. Gradient errors

Consider a small gradient error which affects a quadrupole at position $s$ in the lattice of a circular accelerator. Translated to matrix algebra, we have to multiply a perturbation matrix

$$
\boldsymbol{\delta} \mathbf{Q}(s)=\left(\begin{array}{cc}
1 & 0 \\
-\delta k(s) \cdot d s & 1
\end{array}\right)
$$

with the unperturbed matrix for one circle staring at $\mathrm{s}\left(\right.$ where $\left.\alpha(s)=\alpha_{0}, \beta(s)=\beta_{0}, \gamma(s)=\gamma_{0}\right)$

$$
\mathbf{M}_{0}=\left(\begin{array}{cc}
\cos \mu_{0}+\alpha_{0} \sin \mu_{0} & \beta_{0} \sin \mu_{0} \\
-\gamma_{0} \sin \mu_{0} & \cos \mu_{0}-\alpha_{0} \sin \mu_{0}
\end{array}\right)
$$


giving:

$$
\begin{aligned}
\tilde{\mathbf{M}}(s) & =\boldsymbol{\delta} \mathbf{Q}(s) \cdot \mathbf{M}_{0} \\
& =\left(\begin{array}{cc}
\cos \mu_{0}+\alpha_{0} \sin \mu_{0} & \beta_{0} \sin \mu_{0} \\
-\delta k d s\left(\cos \mu_{0}+\alpha_{0} \sin \mu_{0}\right)-\gamma_{0} \sin \mu_{0} & -\delta k d s \beta_{0} \sin \mu_{0}+\cos \mu_{0}-\alpha_{0} \sin \mu_{0}
\end{array}\right)
\end{aligned}
$$

## CERN Accelerator School: Introductory Course

From $\frac{1}{2} \operatorname{Tr}\{\tilde{\mathbf{M}}\}=\cos \mu$ we can calculate the change in $\cos \mu$ :

$$
\begin{gathered}
\Delta(\cos \mu)=-\Delta \mu \cdot \sin \mu_{0}=-\frac{1}{2} \sin \mu_{0} \beta_{0} \delta k d s \\
2 \pi \Delta Q=\Delta \mu=\frac{1}{2} \beta(s) \delta k(s) d s
\end{gathered}
$$

Integrating over the length of the quadrupol perturbation, one obtains

$$
\Delta Q=\frac{1}{4 \pi} \oint \beta(s) \delta k(s) d s
$$

Effect on beta function:


## CERN Accelerator School: Introductory Course

A gradient error will not influence the closed orbit but the betatron function of the lattice. In order to calculate the betatron amplitude modulation, we have to determine the single turn transport matrix starting at a given observer position s , introducing a small gradient perturbation at position $s_{0}$ :

$$
\tilde{\mathbf{M}}_{s}=\mathbf{M}\left(s, s_{0}\right) \cdot \boldsymbol{\delta} \mathbf{Q}\left(s_{0}\right) \cdot \mathbf{M}\left(s_{0}, s\right)=\left(\begin{array}{ll}
b_{11} & b_{12} \\
b_{21} & b_{22}
\end{array}\right) \cdot\left(\begin{array}{cc}
1 & 0 \\
-\delta k d s_{0} & 1
\end{array}\right) \cdot\left(\begin{array}{ll}
a_{11} & a_{12} \\
a_{21} & a_{22}
\end{array}\right)
$$

It is only necessary to evaluate the element $\tilde{r}_{12}$ which is

$$
\tilde{r}_{12}=b_{11} a_{12}+b_{12}\left(-\delta k d s \cdot a_{12}+a_{22}\right)=r_{12}-\delta k d s_{0} \cdot a_{12} b_{12}
$$

where $r_{12}$ from the unperturbed matrix found by putting $\delta k d s_{0}=0$. Thus the variation in the $r_{12}$ term due to the perturbation is

$$
\begin{aligned}
& \Delta\left[\beta(s) \sin \left(2 \pi Q_{0}\right)\right]=-\delta k d s_{0} \beta(s) \beta\left(s_{0}\right) \cdot \sin \left(\mu(s)-\mu\left(s_{0}\right)\right) \cdot \sin \left(\mu\left(s_{0}\right)-\mu(s)\right) \\
& =-\delta k d s_{0} \beta(s) \beta\left(s_{0}\right) \cdot \sin \left(\mu(s)-\mu\left(s_{0}\right)\right) \cdot \sin \left[2 \pi Q_{0}-\left(\mu(s)-\mu\left(s_{0}\right)\right)\right]
\end{aligned}
$$

Using $\sin \alpha \cdot \sin \beta=\frac{1}{2}[\cos (\alpha-\beta)-\cos (\alpha+\beta)]$ the left-hand and right-hand sides can be expanded to give

$$
\begin{aligned}
\Delta \beta(s) \sin \left(2 \pi Q_{0}\right) & +\beta(s) \cdot \underbrace{2 \pi \Delta Q}_{\uparrow} \cdot \cos \left(2 \pi Q_{0}\right)= \\
& \overbrace{\frac{1}{2} \delta k d s_{0} \beta\left(s_{0}\right)}^{\downarrow} \beta(s)\left\{\cos \left(2 \pi Q_{0}\right)-\cos \left[2\left(\mu(s)-\mu\left(s_{0}\right)-\pi Q_{0}\right)\right]\right\}
\end{aligned}
$$

This leaves the final expression for the betatron amplitude modulation (the so called beta-beating):

$$
\Delta \beta(s)=\frac{\beta(s)}{2 \sin \left(2 \pi Q_{0}\right)} \cdot \oint_{s} \delta k\left(s_{0}\right) \beta\left(s_{0}\right) \cos \left[2\left(\mu(s)-\mu\left(s_{0}\right)-\pi Q_{0}\right)\right] \cdot d s_{0}
$$



## Ideal World:



## Single Quadrupole Error:




### 4.4.6. Optical resonances

Dipole errors will give a large closed orbit displacement when the tune is close to an Gradient errors will produce an average tune shift $\Delta Q$ and an amplitude modulation of the beta function which will explode for half integer Q values.

These phenomena are called resonances. Due to the turn by turn modulation of the tune, there exist regions of instability called stop bands around the resonance conditions. The width of these stop bands are given by the tune modulation amplitude.

These effects can be studied best when regarding the normalized phase space, where the particles ellipses transform to circles:

## CERN Accelerator School: Introductory Course



No average tune shift
Tune modulation amplitude $d Q$

Gradient Errors:


Average tune shift $\Delta Q=\frac{1}{4 \pi} \beta \delta(k l)$
Tune modulation amplitude $d Q=\Delta Q$

Any particle whose unperturbed $Q$ lies in the stop band width $d Q$ will lock into resonance and is lost.

We may generalize and give a list of resonances and their driving multipoles:

## CERN Accelerator School: Introductory Course

## resonance type

integer resonance: $Q=n$
half-integer resonance $2 \cdot Q=n$
third-integer resonance $3 \cdot Q=n$
driving multipole
dipole errors
quadrupole errors
sextupole errors


Due to betatron coupling, perturbations may depend on the betatron amplitude in both planes. These coupling terms lead to the generalized resonance condition

$$
j \cdot Q_{x}+k \cdot Q_{z}=N
$$

where $j+k$ indicates the order of the resonance. The circle represents the tune on the energy ramp of ELSA.

## CERN Accelerator School: Introductory Course

## Example: LHC

Courtesy: R. Steinhagen


Tune stability requirements: $\Delta Q<0.001$ vs exp. Drifts $\sim 0.06$
Note: need to stay much further of resonances due to finite tune width (chromaticity, momentum spread), space charge, beam-beam, etc., and finite width of stop bands.

## CERN Accelerator School: Introductory Course

### 4.5. Beam dynamics with acceleration

Phase space in accelerator physics $\neq$ phase space in classical mechanics:

$$
\text { coordinates } x, x^{\prime} \quad \leftrightarrow \quad \text { canonical coordinates } x, p_{x}
$$

$$
p_{x}=m \cdot \dot{x}=m \cdot \dot{s} \cdot x^{\prime} \approx p_{0} \cdot x^{\prime}=\beta_{r} \gamma_{r} \cdot m_{0} c \cdot x^{\prime} \quad \rightarrow \quad \beta_{r} \gamma_{r} \cdot x^{\prime}=\text { const. }
$$

Beam acceleration (momentum increase) causes compression of $x^{\prime}$ axis and therewith decrease of the beam emittance, which is called adiabatic damping:

$\rightarrow$ Define normalized emittance, which is conserved: $\varepsilon_{n}=\beta_{r} \gamma_{r} \cdot \varepsilon$

## 5. Dynamics with Off Momentum Particles

We will come back to the equation of motion, now explicitly treating the momentum dependent right hand side, depending on the relative momentum deviation $\delta=\Delta p / p_{0}$

$$
\begin{gathered}
x^{\prime \prime}(s)+\left(\frac{1}{\rho^{2}(s)}-k(s)\right) \cdot x(s)=\frac{1}{\rho(s)} \frac{\Delta p}{p} \\
y^{\prime \prime}(s)+k(s) \cdot y(s)=0
\end{gathered}
$$

Since the dynamics of off momentum particles is only affected in the horizontal plane, we will restrict the treatment to 1 D including the momentum dependence.

### 5.1 Dispersion and dispersion functions

A particular solution for a non-vanishing $\delta=\Delta p / p$ is $x_{i h}(s)=\rho \cdot \delta$. Recalling the solution of the homogenous equation, this gives:

$$
x(s)=x_{h}(s)+x_{i h}(s)=a \cdot \cos (s / \rho)+b \cdot \sin (s / \rho)+\rho \cdot \delta
$$

## CERN Accelerator School: Introductory Course

The integration constants $a, b$ are again derived from the boundary conditions at $s=0$, but now the inhomogenous solution has to be included:

$$
x(s=0)=a+\rho \cdot \delta=x_{0}, \quad x^{\prime}(s=0)=\frac{b}{\rho}=x_{0}^{\prime},
$$

and by defining the bending angle $\varphi=L / \rho$ of the dipole magnet, we obtain :

$$
\begin{aligned}
& x(L)=x_{0} \cdot \cos \varphi+\rho \cdot x_{0}^{\prime} \cdot \sin \varphi+\rho(1-\cos \varphi) \cdot \delta \\
& x^{\prime}(L)=-x_{0} / \rho \cdot \sin \varphi+x_{0}^{\prime} \cdot \cos \varphi+\sin \varphi \cdot \delta
\end{aligned}
$$

This can be easily implemented in the matrix formalism by adding a $3^{\text {rd }}$ component to the particle's position vector dealing with the actual relative momentum deviation compared to the reference particle:

$$
\vec{X}=\left(\begin{array}{c}
x \\
x^{\prime} \\
\delta
\end{array}\right) \quad \mathrm{M}_{\text {dipole }}=\left(\begin{array}{cc}
\begin{array}{|cc|}
\hline \cos \varphi & \rho \sin \varphi \\
-1 / \rho \sin \varphi & \cos \varphi
\end{array} & \rho(1-\cos \varphi) \\
\sin \varphi \\
0 & 0
\end{array}\right.
$$

## CERN Accelerator School: Introductory Course

First neglecting the dependence of the quadrupole strength $k$ on the actual particle's momentum, the quadrupole transfer matrices remain "unchanged":
$\mathrm{M}_{\mathrm{QF}}=\left(\begin{array}{ccc}\begin{array}{|cc|}\hline \cos \Omega & \sqrt{|k|} \sin \Omega \\ -1 / \sqrt{|k|} \sin \Omega & \cos \Omega\end{array} & 0 \\ 0 & 0 & 1\end{array}\right)$

$$
\mathrm{M}_{\mathrm{QD}}=\left(\begin{array}{ccc}
\begin{array}{|cc|}
\hline \cosh \Omega & \sqrt{|k|} \sinh \Omega \\
1 / \sqrt{|k|} \sinh \Omega & \cosh \Omega
\end{array} & 0 \\
0 & 0 & 1
\end{array}\right)
$$

## Important:

Whereas a quadrupole magnet will not directly cause an impact on the particle's trajectory, a dipole magnet creates a (horizontal) dispersion:

$$
D=r_{16}=\rho(1-\cos \varphi), \quad D^{\prime}=r_{26}=\sin \varphi
$$

The dispersion represents the offset due to a relative momentum deviation $\Delta p / p=1$.
In general, we have: $\quad x(s)=x_{h}(s)+x_{D}(s)=x(s)+D(s) \cdot \frac{\Delta p}{p}$
Here, $D(\mathrm{~s})$ is the dispersion function, a solution of the equation of motion for $\delta=1$.

## But now take care:

Due to $x(s)=x_{h}(s)+x_{D}(s)$, we will observe a change of the dispersion orbit $x_{D}(s)$ when passing a dipole magnet or a quadrupole magnet!!

Both dipole and quadrupole magnets therefore will modify an existing dispersion according to

$$
\left(\begin{array}{c}
D(s) \\
D^{\prime}(s) \\
1
\end{array}\right)=\underbrace{\left(\begin{array}{ccc}
\begin{array}{|cc|}
\hline r_{11} & r_{12} \\
r_{21} & r_{22}
\end{array} & r_{16} \\
r_{26} \\
0 & 0 & 1
\end{array}\right)}_{\mathbf{M}_{\text {dipole }}} \cdot\left(\begin{array}{c}
D_{0} \\
D_{0}^{\prime} \\
1
\end{array}\right) \quad\left(\begin{array}{c}
D(s) \\
D^{\prime}(s) \\
1
\end{array}\right)=\underbrace{\left(\begin{array}{|cc|}
\begin{array}{|cc|}
\hline r_{11} & r_{12} \\
r_{21} & r_{22} \\
\hline & 0
\end{array} & 0 \\
0
\end{array}\right)}_{\mathbf{M}_{\text {quadrupole }}} \cdot\left(\begin{array}{c}
D_{0} \\
D_{0}^{\prime} \\
1
\end{array}\right)
$$

### 5.2 Dispersion in circular accelerators

In a periodic lattice, the dispersion function has - as well as the beta function - to fulfill periodic boundary conditions:

$$
D\left(s_{0}+L\right)=D\left(s_{0}\right)
$$

Thus the dispersion function can obtained from applying the $3 \times 3$ transport matrix $\mathbf{M}_{3}$ for a full period

$$
\begin{gathered}
\left(\begin{array}{c}
D_{0} \\
D_{0}{ }^{\prime} \\
1
\end{array}\right)=\mathbf{M}_{3} \cdot\left(\begin{array}{c}
D_{0} \\
D_{0}{ }^{\prime} \\
1
\end{array}\right)=\left(\begin{array}{ccc}
r_{11} & r_{12} & r_{13} \\
r_{21} & r_{22} & r_{23} \\
0 & 0 & 1
\end{array}\right) \cdot\left(\begin{array}{c}
D_{0} \\
D_{0}{ }^{\prime} \\
1
\end{array}\right) \\
D_{0}{ }^{\prime}=\frac{r_{21} r_{13}+r_{23}\left(1-r_{11}\right)}{2-r_{11}-r_{22}} \\
D_{0}=\frac{r_{12} D_{0}{ }^{\prime}+r_{13}}{1-r_{11}}
\end{gathered}
$$

yielding:
which for a symmetry point, where $D_{0}{ }^{\prime}=0$, simplifies to

$$
D_{0}^{\text {sym }}=\frac{r_{13}}{1-r_{11}}
$$

Applying this to our model toy synchrotron, we can derive the dispersion function which is plotted in blue:


Please note that the total beam width is given by $\sigma_{x}=\sqrt{\varepsilon_{x} \beta_{x}+\left(D_{x} \delta\right)^{2}}$ !

## CERN Accelerator School: Introductory Course

### 5.3. Chromaticity

The variation of tunes is called chromaticity and is defined by the factor $\xi$ in

$$
\Delta Q_{x, y}=\xi_{x, y} \cdot \frac{\Delta p}{p_{0}}
$$

We distinguish between natural chromaticity created by the chromatic aberration of quadrupole magnets and perturbations derived from non-linear perturbations in the particles
 trajectories (e.g. produced by sextupole magnets).

## Natural Chromaticity:

The quadrupole strength scales with the particles momentum:

$$
\Delta k=-k \cdot \Delta p / p_{0}
$$

and the tune shift can therefore be calculated from:

$$
\Delta Q_{x, y}=\underbrace{-\frac{1}{4 \pi} \int \beta_{x, y}(\tilde{s}) \cdot k_{x, y}(\tilde{s}) \cdot d \tilde{s}}_{=\xi_{x, y}} \cdot \frac{\Delta p}{p_{0}}
$$

Chromaticity produced by sextupoles:
A beam of particles moving on a dispersion orbit through a sextupole magnet is "focused" by the nonlinear field due to horizontal displacement $x=D \cdot \frac{\Delta p}{p_{0}}$. We can derived a position dependent focusing strength from

$$
\frac{q}{p} \vec{B}_{\text {sext }}=m x y \hat{e}_{x}+\frac{1}{2} m\left(x^{2}-y^{2}\right) \hat{e}_{y}
$$

giving a dispersion dependent $k_{x}$ and $k_{z}$ to:

$$
\begin{aligned}
& k_{x}=\frac{q}{p} \cdot \frac{\partial B_{y}}{\partial x}=m \cdot x=m \cdot D \cdot \frac{\Delta p}{p_{0}} \\
& k_{y}=\frac{q}{p} \cdot \frac{\partial B_{x}}{\partial y}=m \cdot x=m \cdot D \cdot \frac{\Delta p}{p_{0}}
\end{aligned}
$$

This adds to the natural chromaticity and gives in total:

$$
\xi_{x, y}=-\frac{1}{4 \pi} \int\left[k_{x, y}(\tilde{s})-m(\tilde{s}) D(\tilde{s})\right] \cdot \beta_{x, y}(\tilde{s}) d \tilde{s}
$$

In order to avoid a large tune spread, chromaticity has to be corrected by the use of additional sextupole magnets right after focusing and defocusing quadrupoles where the horizontal dispersion does not vanish:


This correction will have an influence on the stability of the beam and the maximum aperture given by nonlinear effects (so called dynamic aperture):

## CERN Accelerator School: Introductory Course



The dynamic aperture can be calculated from a tracking of the particles orbit through the accelerator where the nonlinear effect of sextupole magnets has to be treated as step by step correction in linear beam matrix optics:


The orbit vector is transformed from $s_{0}$ to $s_{1}$ by matrix transformation

$$
\vec{X}_{1}=\mathbf{M}_{1} \cdot \vec{X}_{0}
$$

A sextupole of length $l$ will produce an angular kick in the horizontal and vertical or-
bit of

$$
\begin{aligned}
& \Delta x_{1}^{\prime}=\frac{1}{2} m l \cdot\left(x_{1}^{2}-y_{1}^{2}\right) \\
& \Delta y_{1}^{\prime}=m l \cdot x_{1} y_{1}
\end{aligned}
$$

which gives an orbit vector right after the sextupole of

$$
\vec{X}_{2}=\left(\begin{array}{c}
x_{1} \\
x_{1}^{\prime}+\Delta x_{1}^{\prime} \\
y_{1} \\
y_{1}^{\prime}+\Delta y_{1}^{\prime}
\end{array}\right)
$$

By this method a randomly chosen distribution of start vectors $\vec{X}_{0}$ is tracked through the accelerator for many revolutions and the resulting dynamic aperture is derived from the phase space representation.

### 5.4 Path length and momentum compaction

The path length of a particle with horizontal orbit displacement $x_{D}$ is influenced by the curved sections of the beam line. The total path length is therefore given by

$$
L=\int r d \varphi=\int_{s_{0}}^{s}\left[\frac{\rho(\tilde{s})+x_{D}(\tilde{s})}{\rho(\tilde{s})}\right] d \tilde{s}=L_{0}+\int_{s_{0}}^{s} \frac{x_{D}(\tilde{s})}{\rho(\tilde{s})} d \tilde{s}
$$

With a given relative momentum deviation $\delta=\Delta p / p$, we have $x_{D}(s)=D(s) \cdot \delta$ and obtain the deviation $\Delta L=L-L_{0}$ from the ideal path length

$$
\Delta L=\delta \int_{s_{0}}^{s} \frac{D(\tilde{s})}{\rho(\tilde{s})} d \tilde{s}
$$

This variation is determined by the momentum compaction factor $\alpha_{c}$, defined for a circular accelerator by

$$
\alpha_{c}=\frac{\Delta L / L_{0}}{\Delta p / p}=\frac{1}{L_{0}} \cdot \oint_{L_{0}} \frac{D(\tilde{s})}{\rho(\tilde{s})} d \tilde{s}
$$



The travel time is given by $\tau=L /\left(\beta_{r} c\right)$, and its relative variation is obtained from the logarithmic differentiation (using $\delta=\frac{\Delta p}{p}=\frac{1}{p} \frac{\partial p}{\partial \beta} \Delta \beta=\frac{\Delta \beta \cdot m_{0} c}{p} \cdot \frac{\partial(\beta \gamma)}{\partial \beta}=\frac{\Delta \beta}{\beta} \cdot \gamma^{2}$ )

$$
\Delta \ln \tau=\frac{\Delta \tau}{\tau}=\frac{\Delta L}{L}-\frac{\Delta \beta_{r}}{\beta_{r}}=\left(\alpha_{c}-\frac{1}{\gamma_{r}^{2}}\right) \cdot \delta=-\eta \cdot \delta
$$

where we have defined the slip factor $\eta$ by

$$
\eta=\frac{1}{\gamma_{r}^{2}}-\alpha_{c}
$$

The momentum compaction factor therefore characterizes a critical energy

$$
\gamma_{t r}=\frac{1}{\sqrt{\alpha_{c}}}
$$

which is called the transition energy, where the slip factor vanishes. In this case, all particles will have - to first order independent from their individual momentum - the same revolution frequency. The (catastrophic) consequences will be treated in the lecture on longitudinal beam dynamics.

