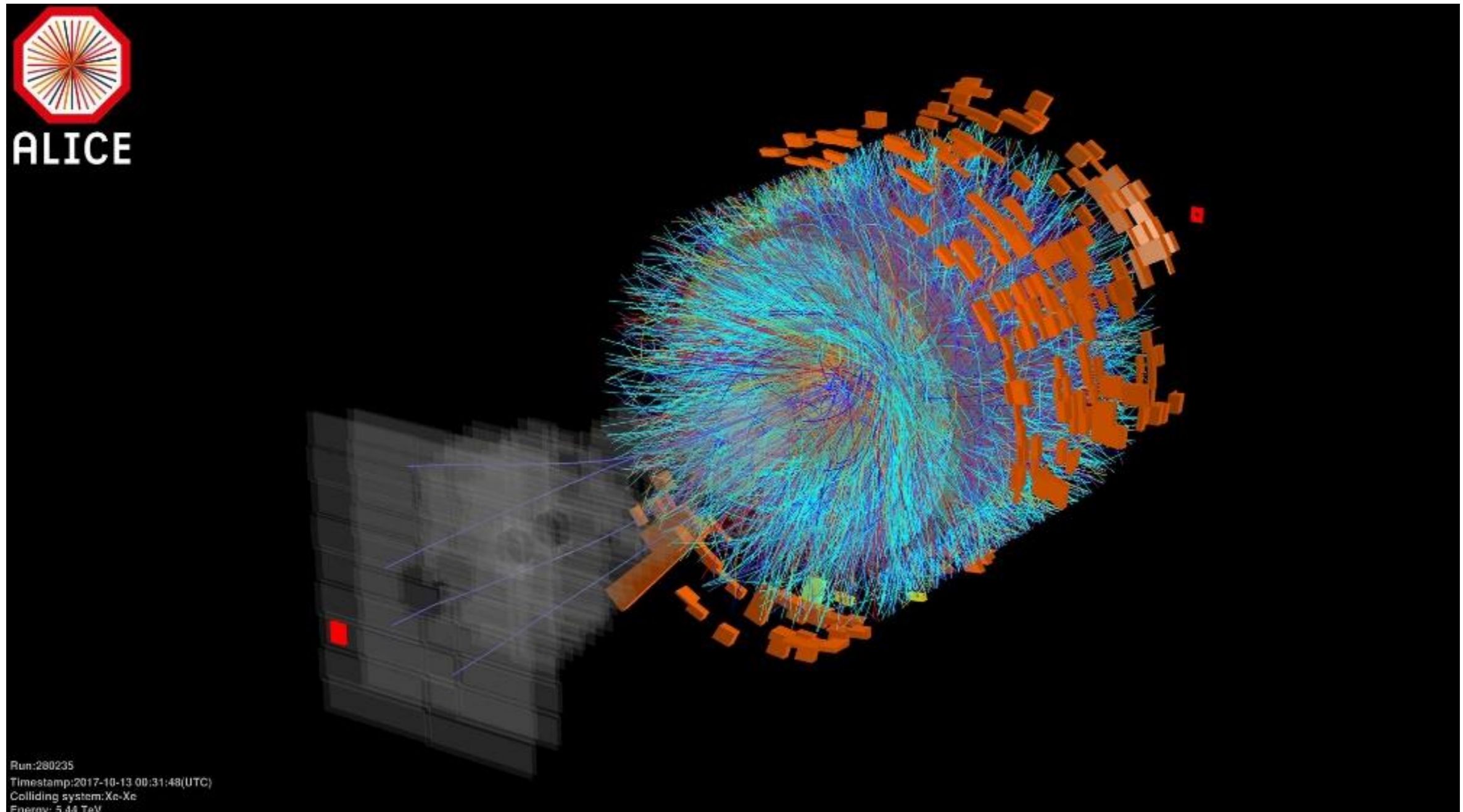


1 H																	2 He
3 Li	4 Be											5 B	6 C	7 N	8 O	9 F	10 Ne
11 Na	12 Mg											13 Al	14 Si	15 P	16 S	17 Cl	18 Ar
19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr
37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe
55 Cs	56 Ba	57 La	* 72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn
87 Fr	88 Ra	89 Ac	* 104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Nh	114 Fl	115 Mc	116 Lv	117 Ts	118 Og
			* 58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu	
			* 90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr	

Light Ions for Jet Probes

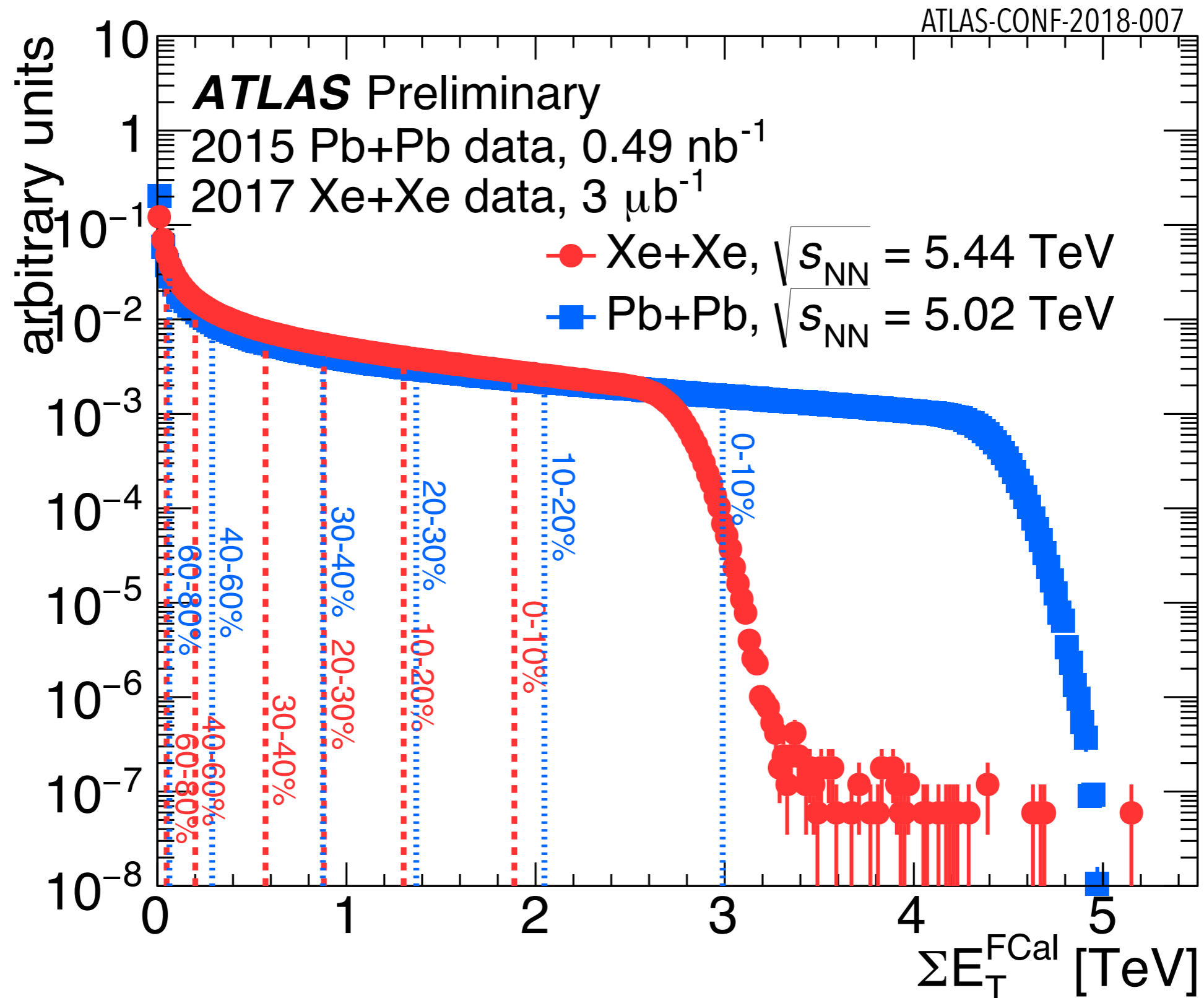
Anne M. Sickles
June 19, 2018

- what did we learn about jets from XeXe collisions so far?
- arguments in favor of high luminosity light ion running
 - increased rates for hard probes
 - understanding the onset of jet quenching
 - decreased systematics from the UE
- questions
 - which ion(s)?
 - how much running is needed?

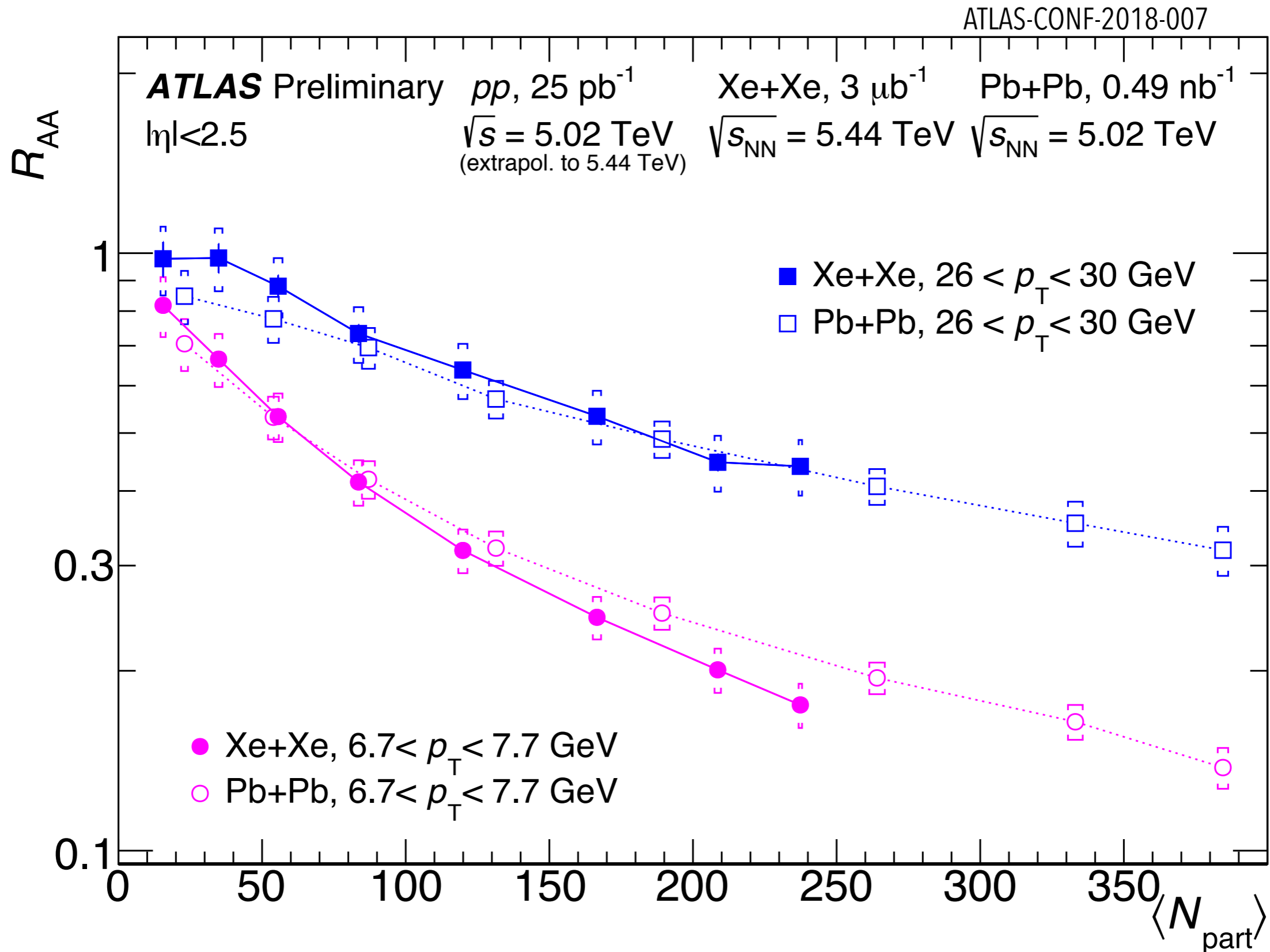


- new XeXe (max $N_{part} \sim 280$) provided a first look at lighter species at the LHC

central XeXe still pretty central



smaller systems \rightarrow smaller quenching

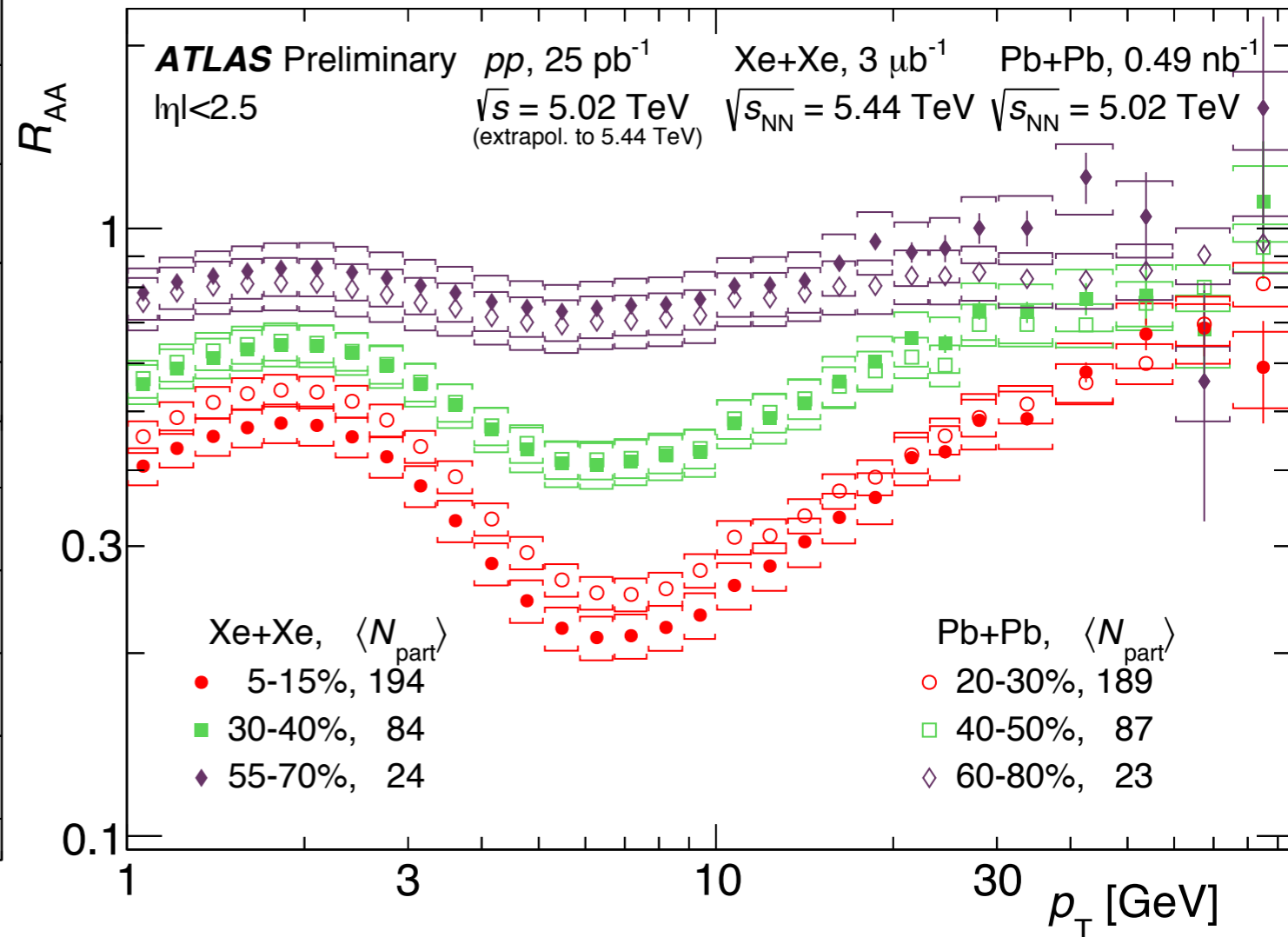
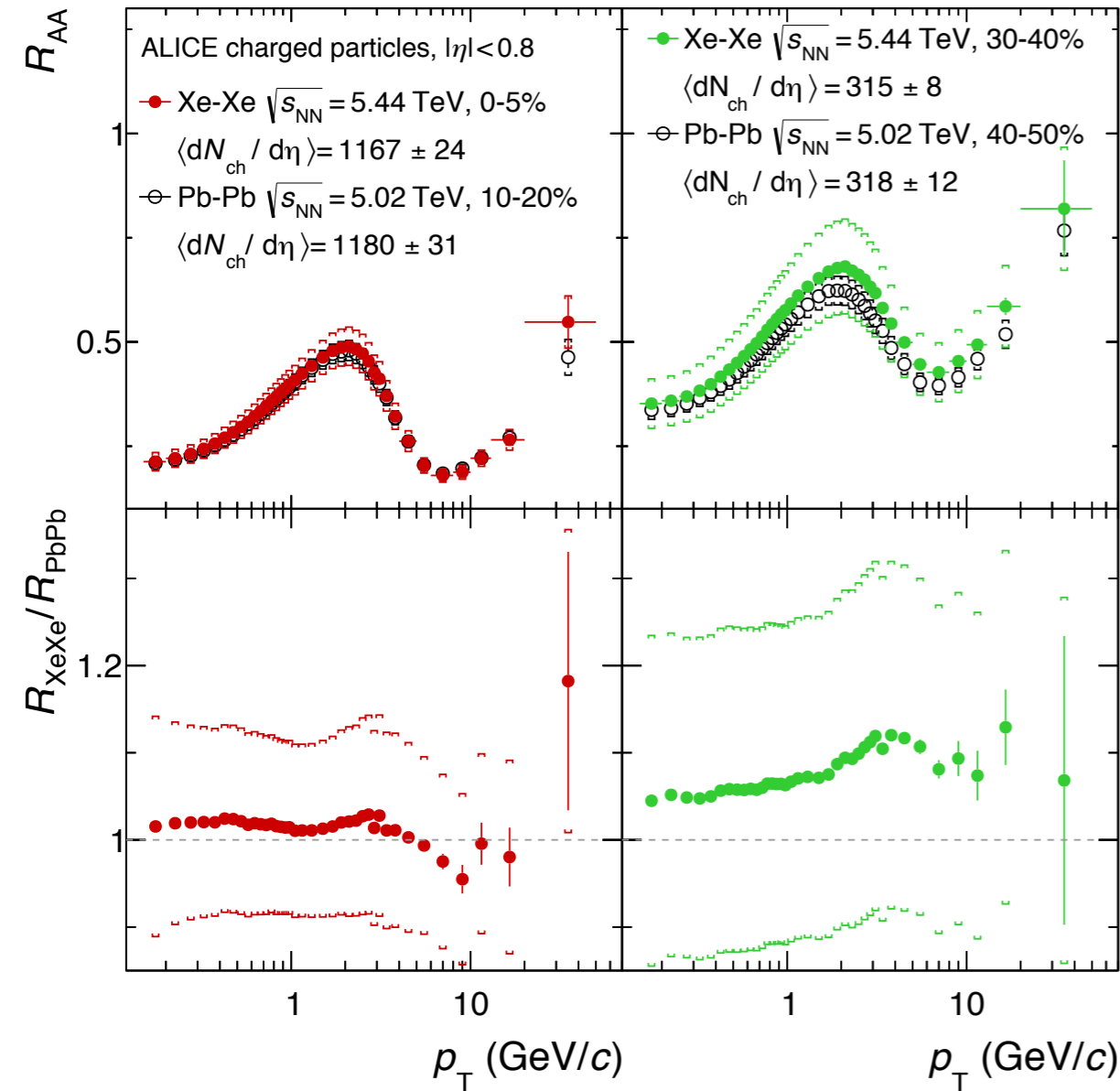


R_{AA} comparisons with system size selection

charged particles

1805.04399

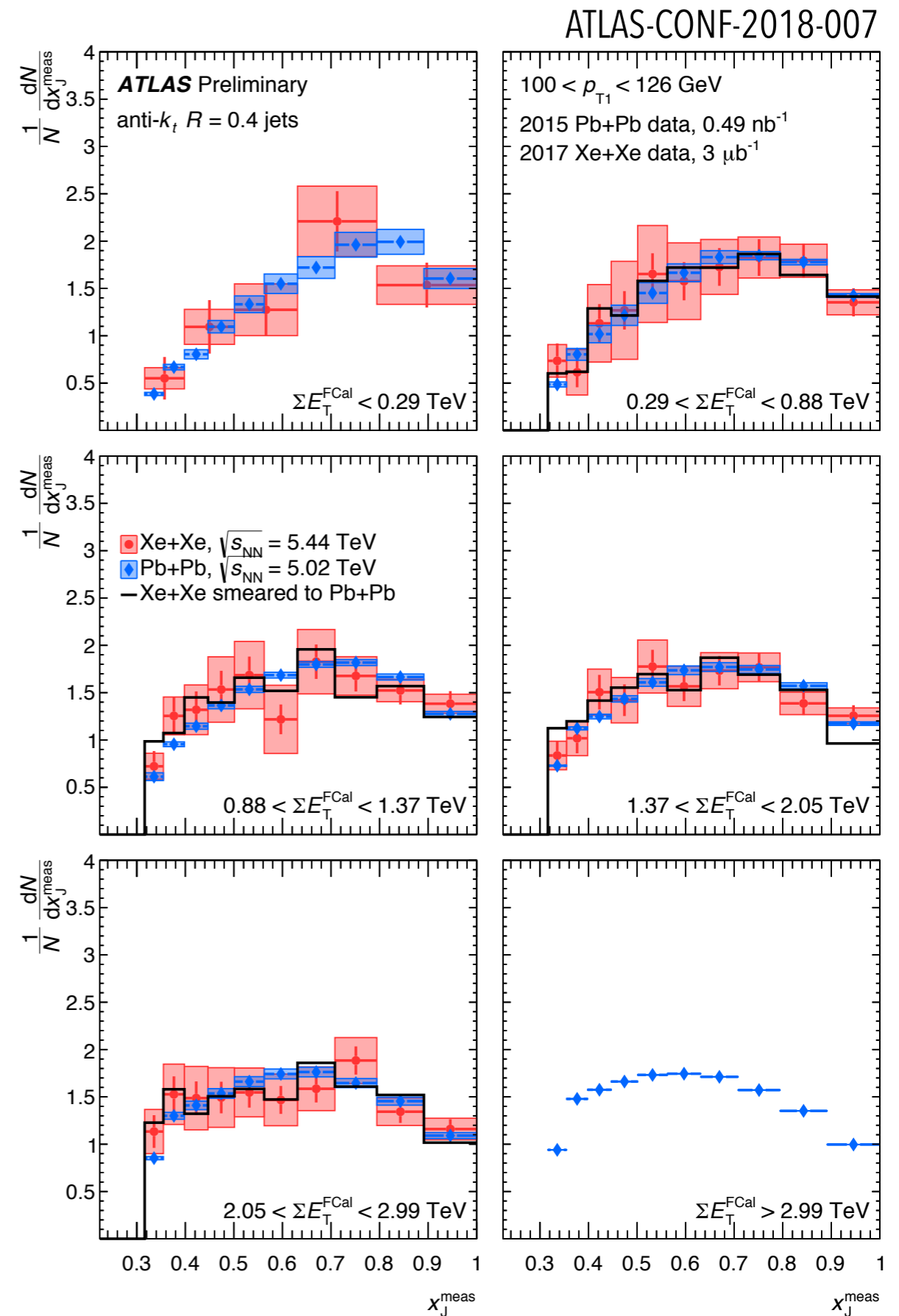
ATLAS-CONF-2018-007



R_{AA} comparable for systems with the similar $dN_{ch}/d\eta$ or N_{part}

dijet asymmetry in XeXe

- XeXe and PbPb have consistent x_J distributions when selected on the same forward ET
- no unfolding done here—additional smearing applied to XeXe to match the UE fluctuations of the PbPb
- with current (large) uncertainties comparing as a function of centrality works as well
- would have been very interesting to see if a very small system (00) would have discriminated between centrality and forward E_T



why lighter ions?

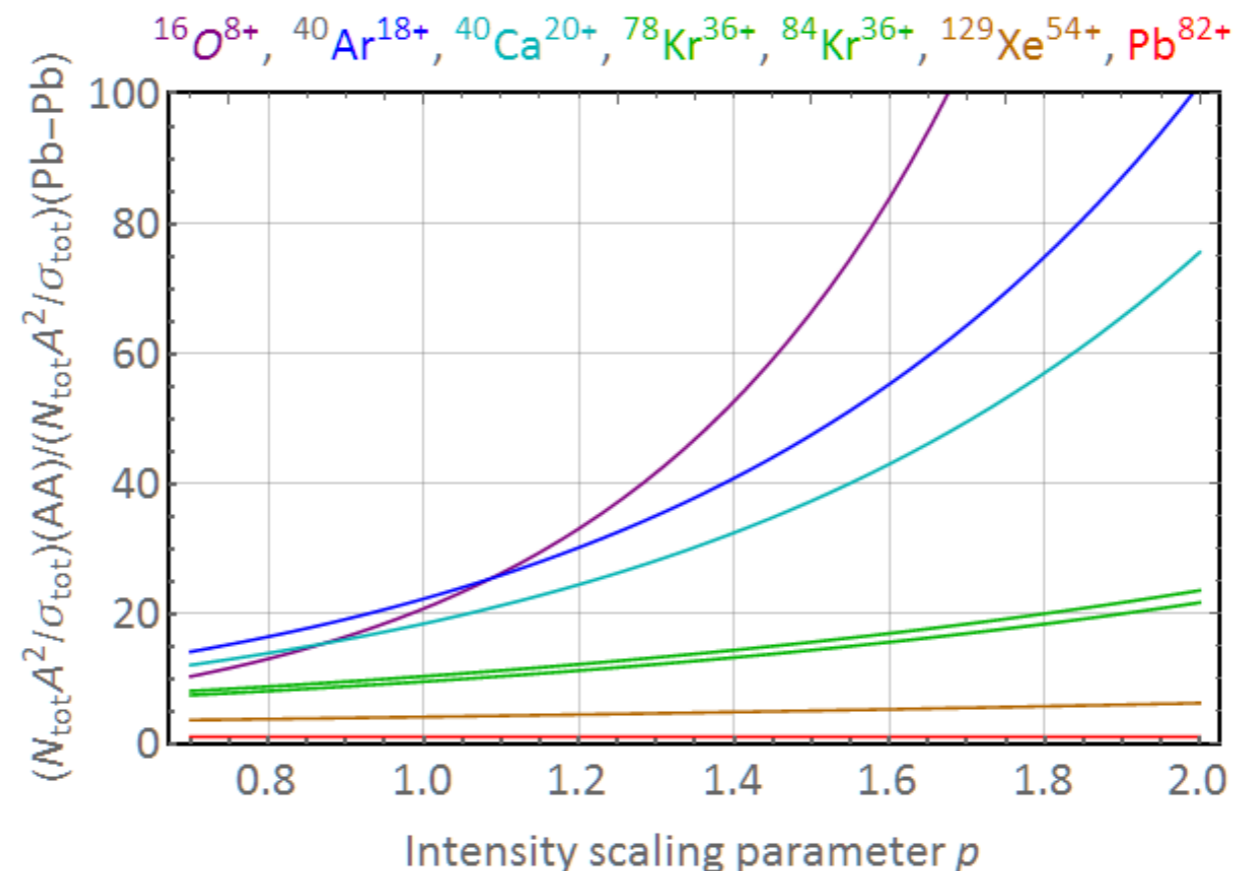
gain in nucleon-nucleon luminosity as a function of A

Gains in ULTIMATE integrated nucleon-nucleon luminosity PER FILL wrt Pb-Pb

This would be on the assumption that a fill would be kept forever until one beam was exhausted (and other loss mechanisms are neglected). Real gain/fill will be less.

In reality, one also gains from longer luminosity lifetime and less time spent refilling the machine.

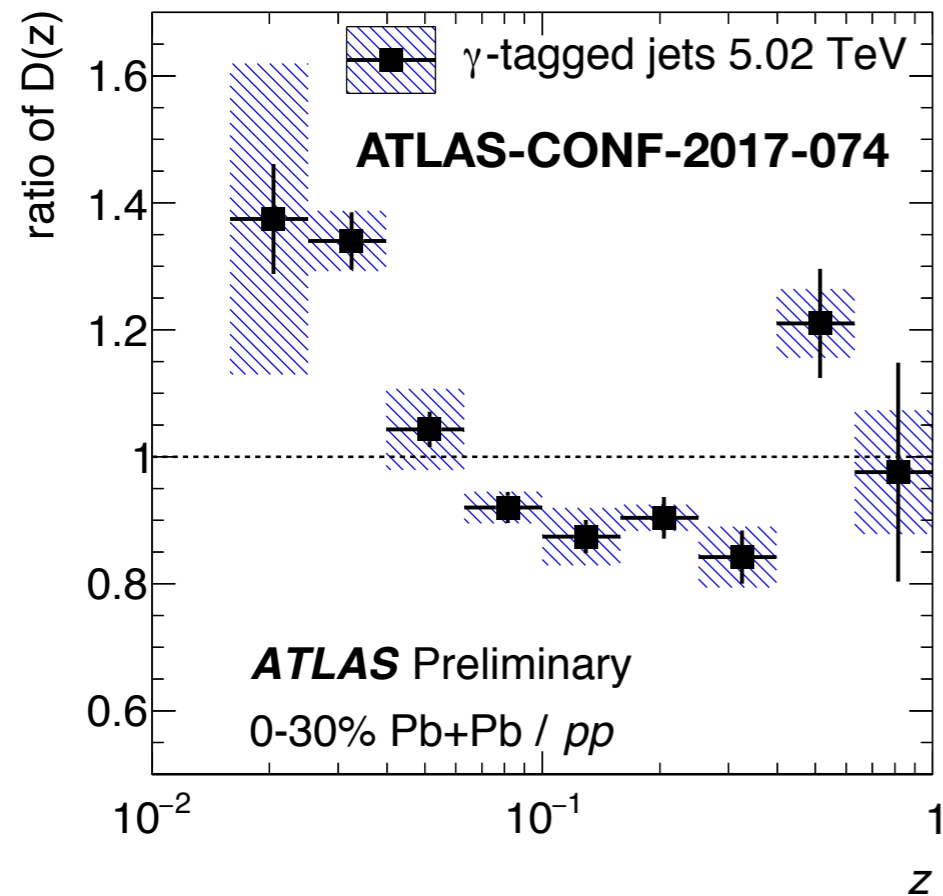
We will try to quantify this better in future.



the smaller the system, the greater the increase in the number of jets
ArAr 20-50x more NN luminosity than PbPb

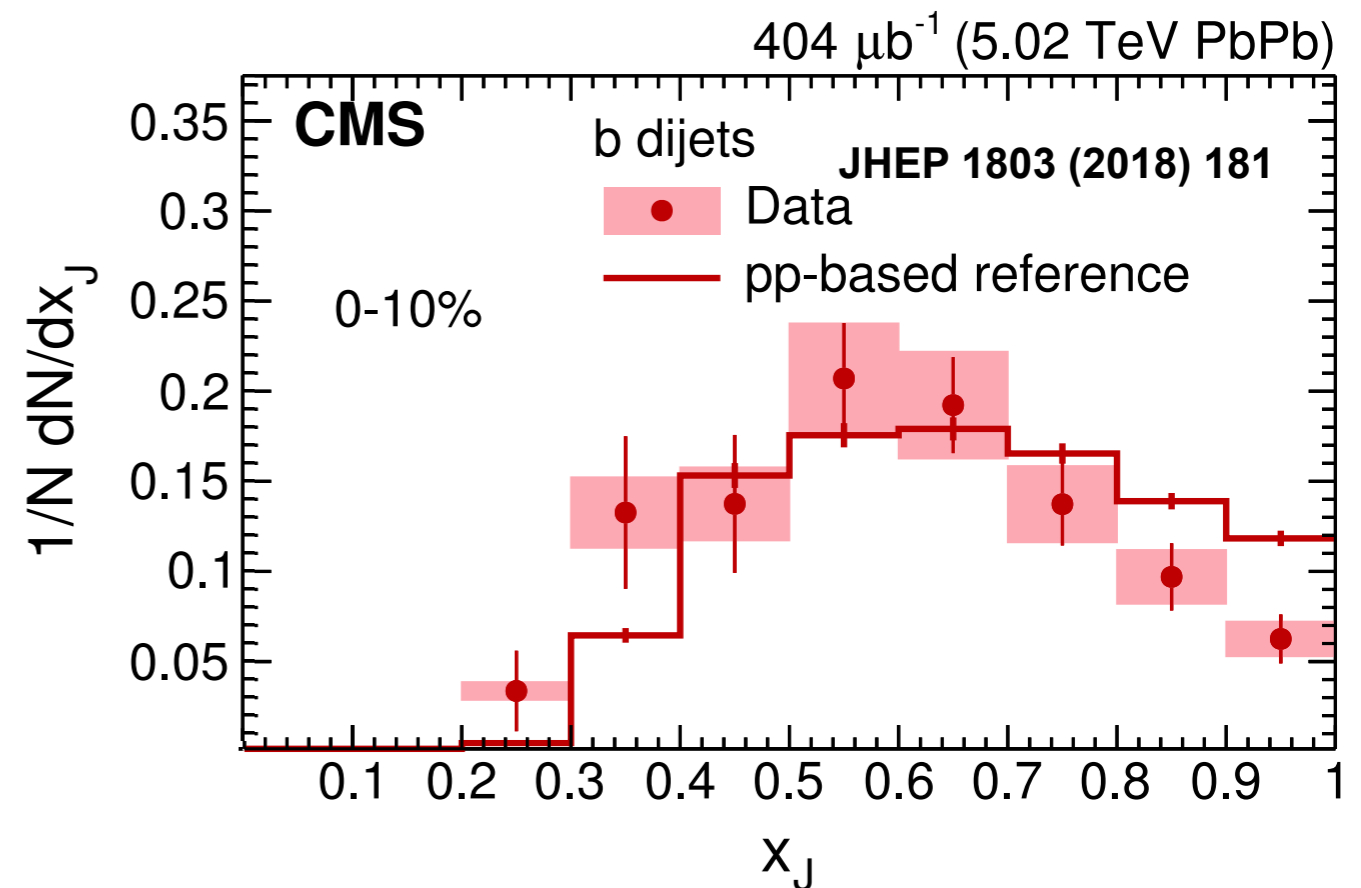
impact of limited statistics

fragmentation functions of jets opposite photons



photon p_T : 79.6-125 GeV
 jet p_T : 63.1-144 GeV

asymmetry of b-jets

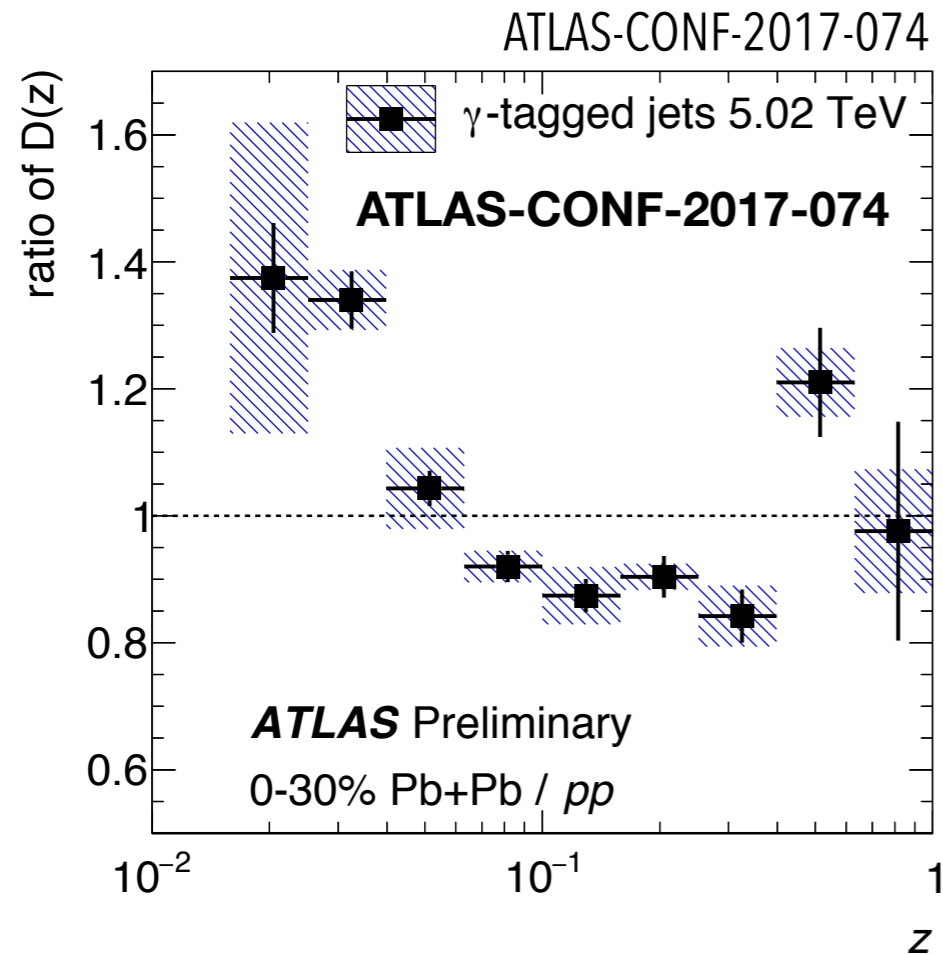


jet $p_{T,1}$: > 100 GeV
 jet $p_{T,2}$: > 40 GeV

non-inclusive jet observables—statistics limitations preclude looking at p_T / centrality dependence and comparison to corresponding inclusive jet observables

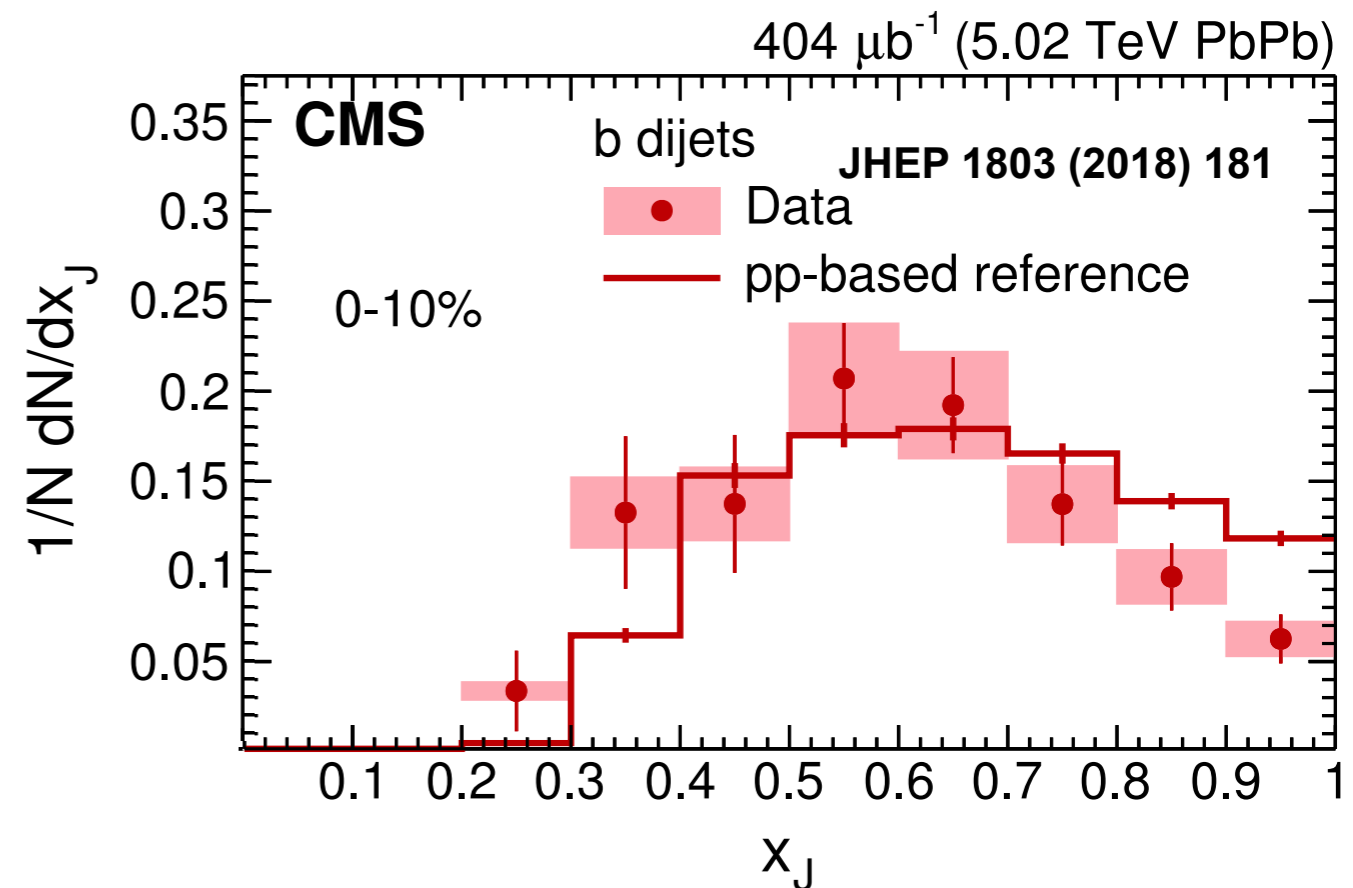
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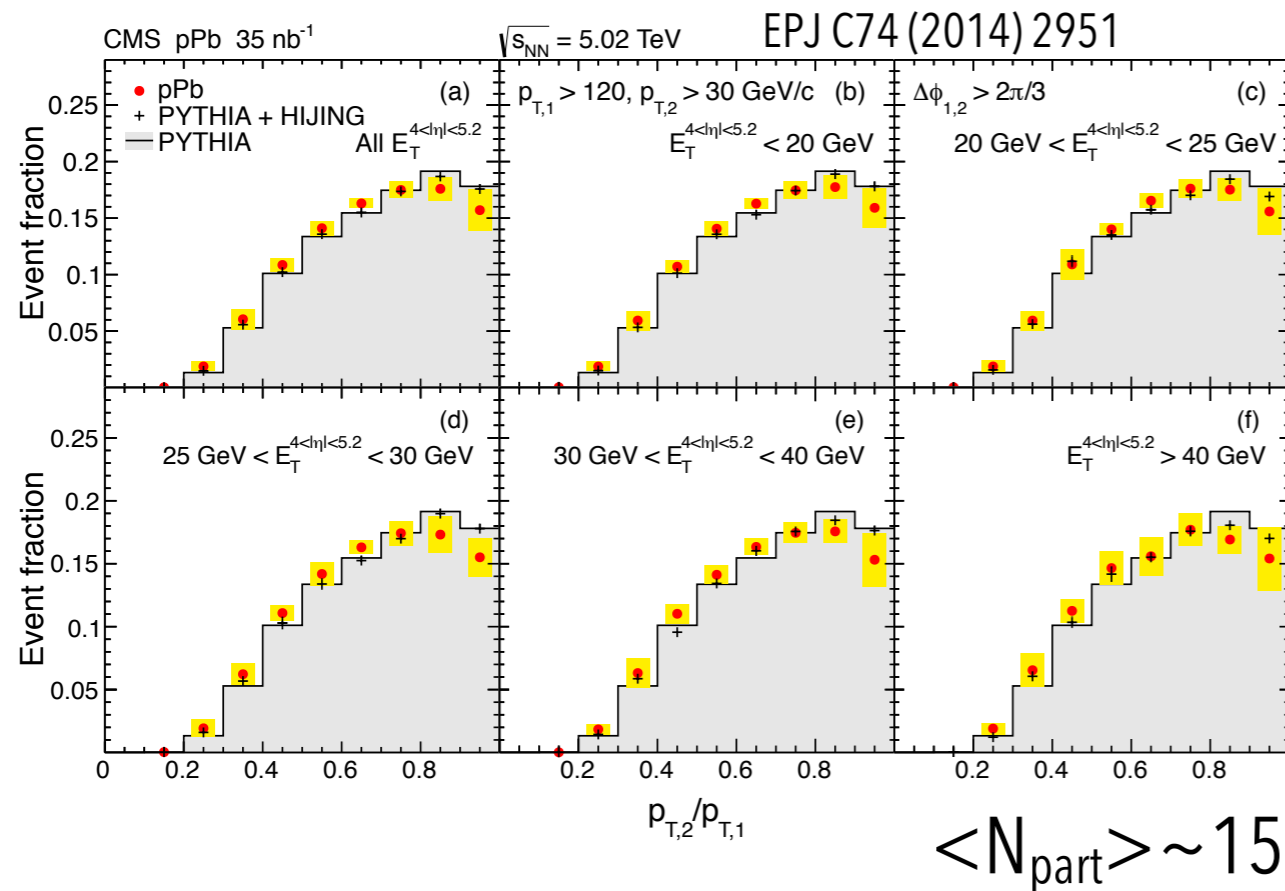


jet $p_{T,1}$: > 100 GeV
 jet $p_{T,2}$: > 40 GeV

- the planned 10 nb⁻¹ of PbPb will improve the statistics of all observables
- however, new observables can be exploited with larger luminosity gains
- b-tagged substructure (1801.00008), top (PRL 120 232301), ...

where is the onset of jet quenching?

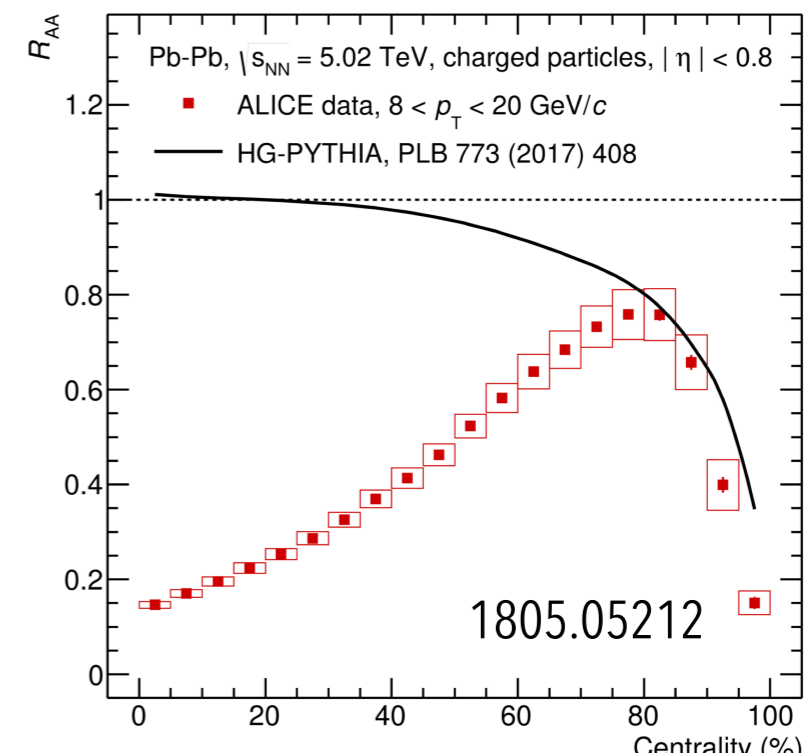
understanding the onset of jet quenching is one of the most important outstanding questions



no evidence seen for jet quenching in pPb collisions from ALICE, ATLAS or CMS

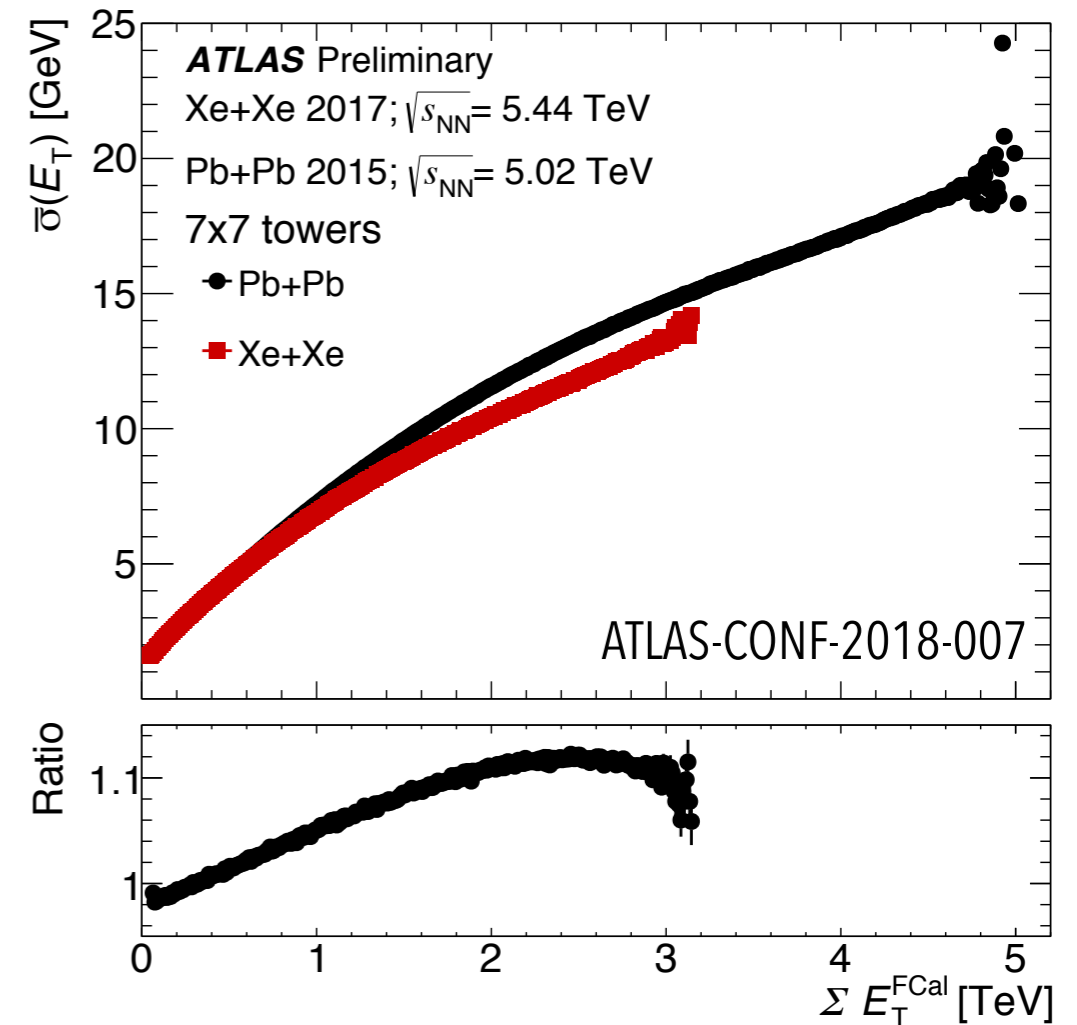
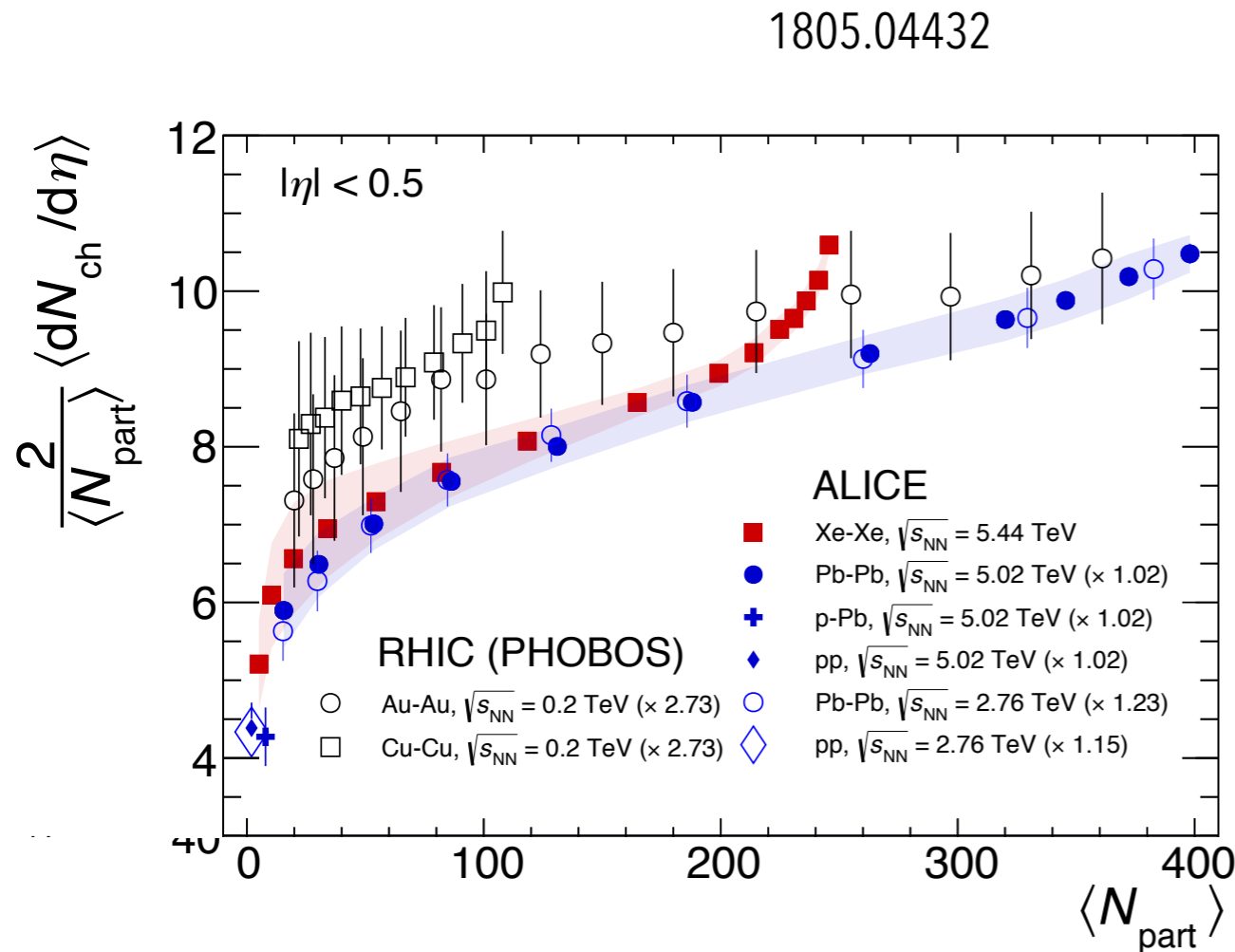
looking for jet quenching in peripheral PbPb collisions complicated by large TAA uncertainties and selection biases (ALICE QM18)

light ions (both Ar & O) have small N_{part} in central collisions providing excellent systems to study the onset of jet quenching



measurement of the UE & fluctuations in XeXe & PbPb

mean UE fluctuations over the size of an $R = 0.4$ jet



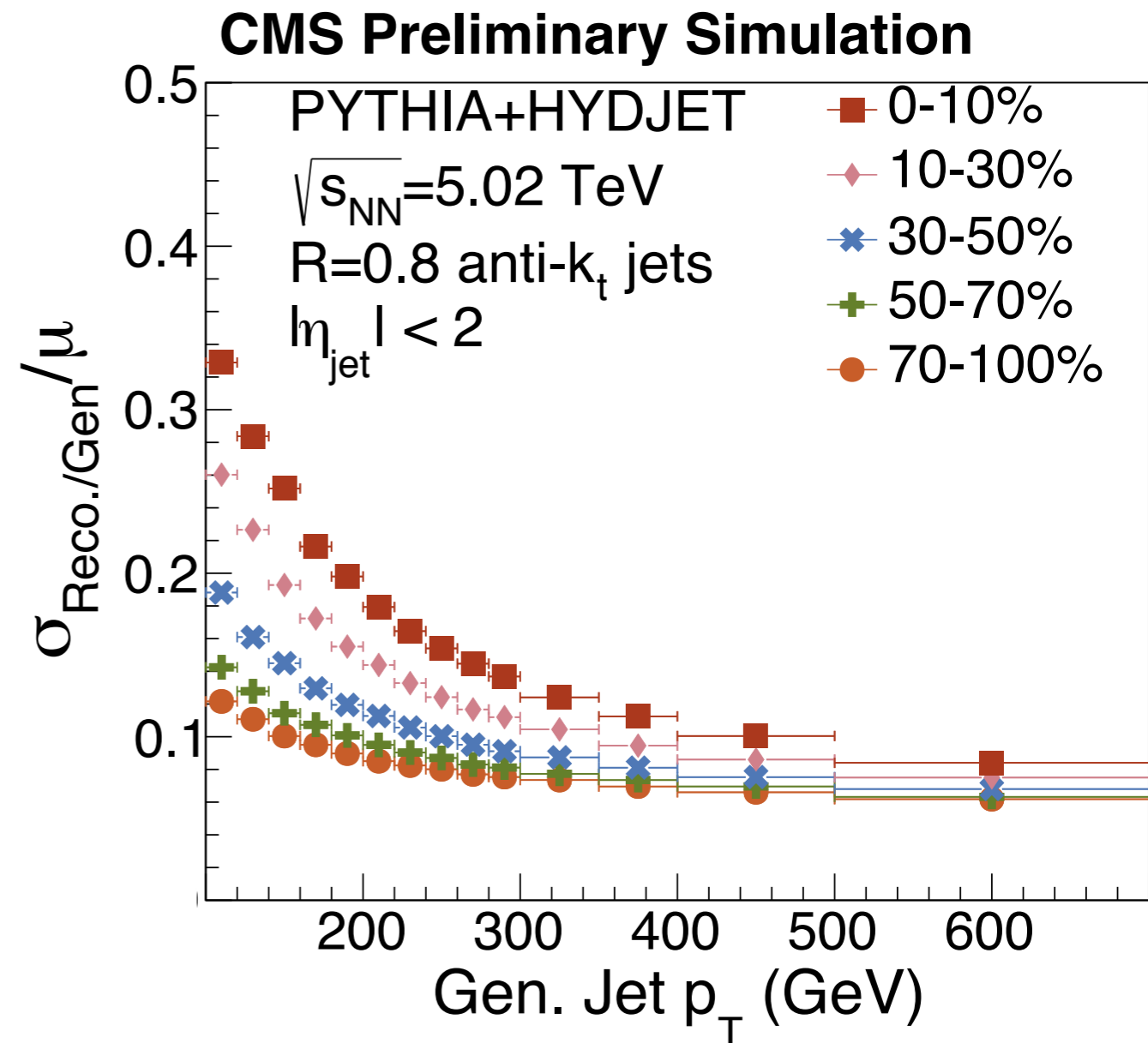
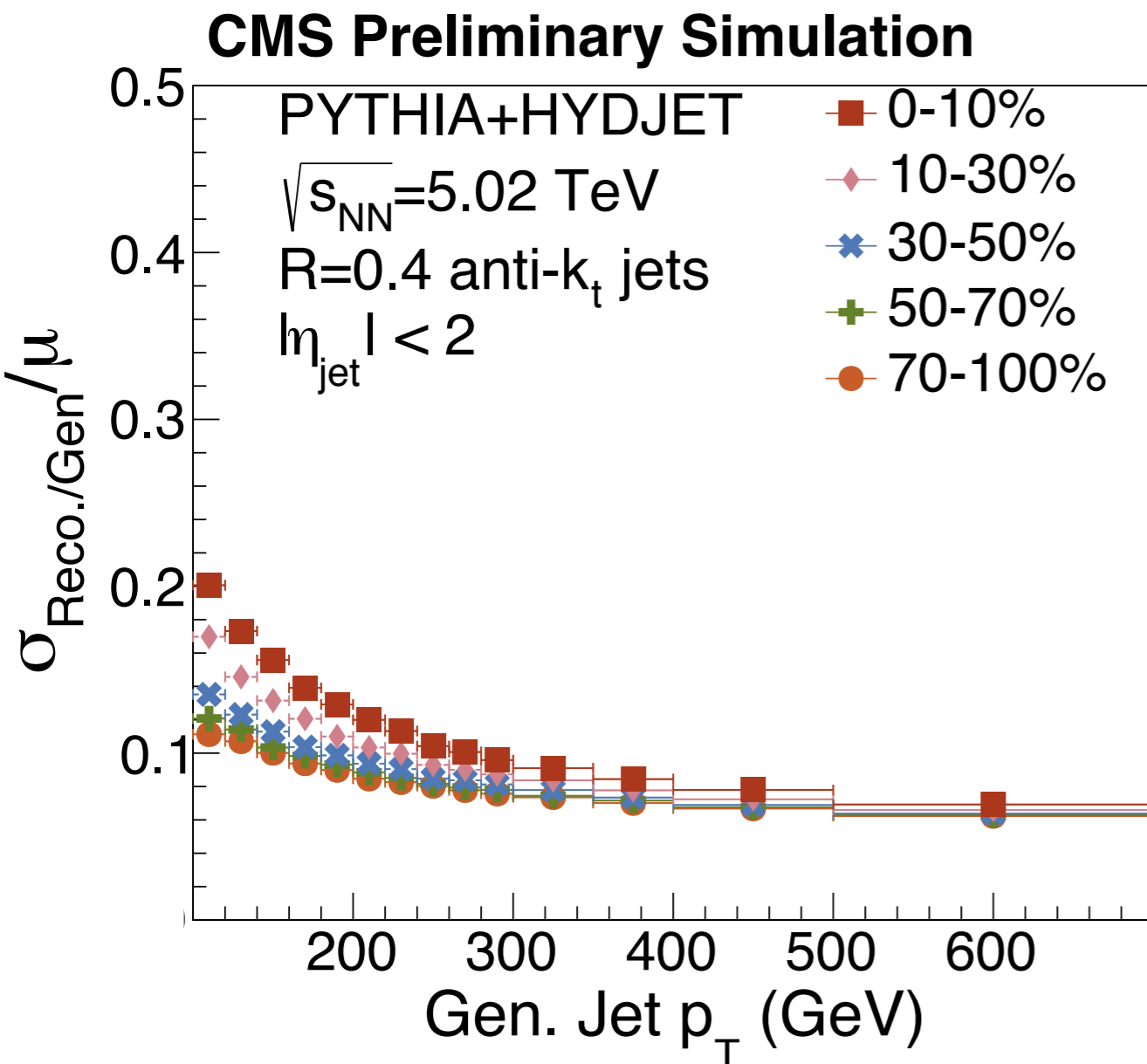
UE in central collisions $\sim 2x$ smaller in central XeXe than in PbPb

at the same forward energy XeXe has up to $\sim 10\%$ smaller UE fluctuations than PbPb

effects of the UE and its fluctuations decrease with system size—
improving precision of measurements

JER for large R jets in PbPb

example: light ions would improve the feasibility of large R jets in central collisions

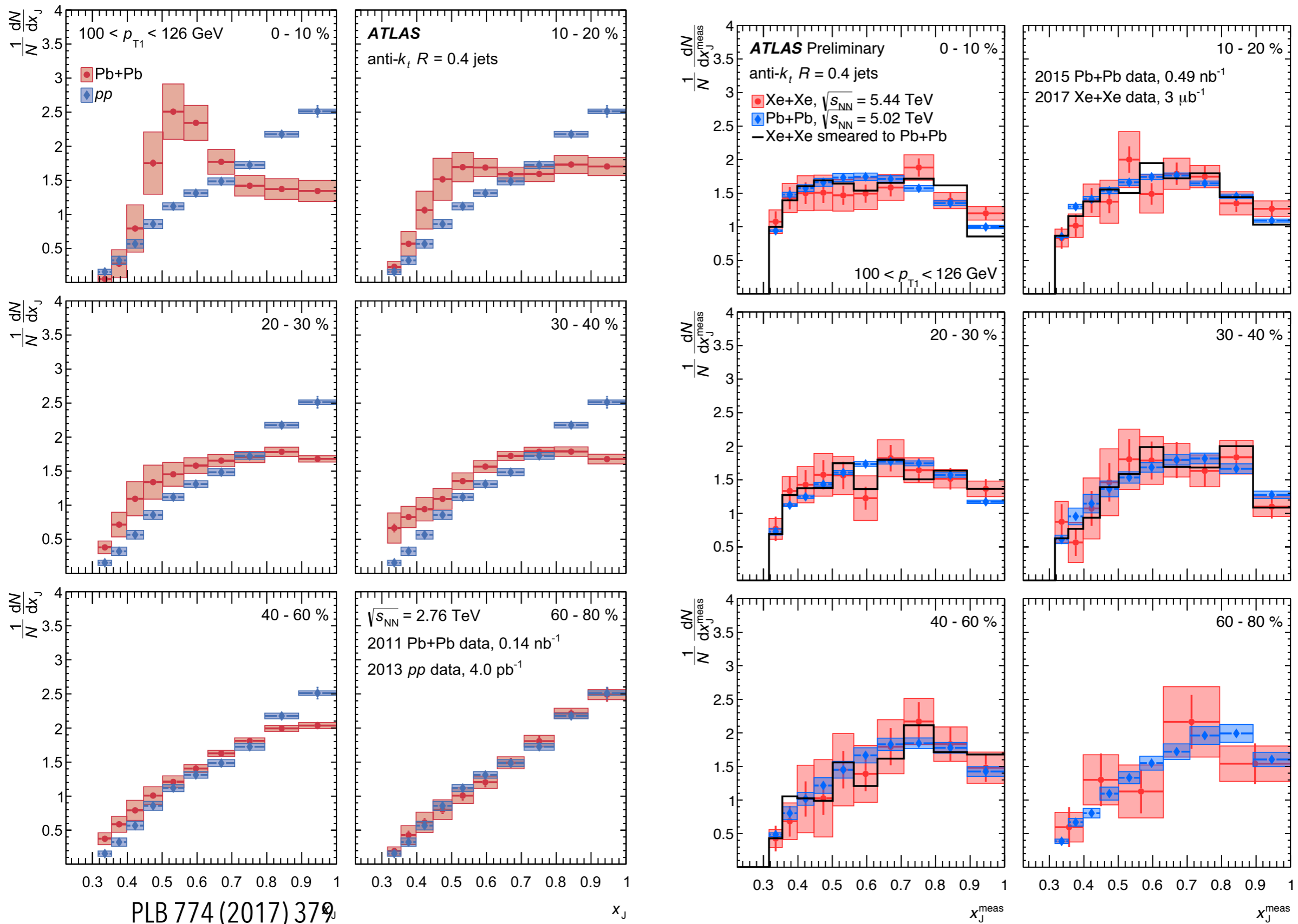


which ions and how much running?

- ions:
 - trade off between luminosity increases and medium effects
 - we have seen the Xe+Xe collisions have significant quenching
 - for onset of quenching probably want very light ions; O+O, Ar+Ar
- how much luminosity:
 - if we were sure that we could see all the Pb+Pb effects with light ions there would be no reason to go back to Pb+Pb
 - strawman proposal: a sample of Ar+Ar early in Run 3 which has \sim the same number of jets as 2015 Pb+Pb data
 - greatly improve understanding of quenching in small systems
 - understand the balance between luminosity and medium effects for future planning

- XeXe data shows significant jet quenching based on first results
- arguments in favor of high luminosity light ion running
 - increased rates for hard probes
 - understanding the onset of jet quenching
 - decreased systematics from the UE
- light ions are promising looking toward Run3 & 4
 - it would be very useful to have a large sample of Ar+Ar collisions to
 - study the onset of quenching
 - understand the balance between luminosity and medium effects for future planning

backup



jet fragmentation & medium response

