



NAGOYA UNIVERSITY

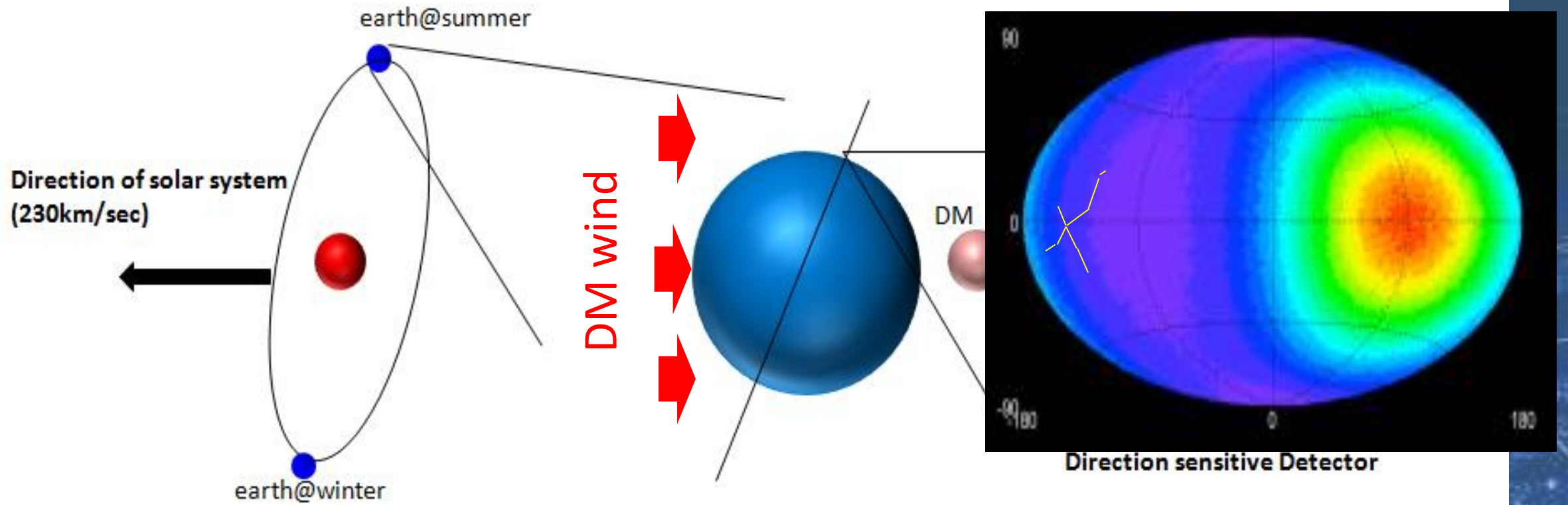


Kobayashi-Maskawa Institute
for the Origin of Particles and the Universe

Direction Sensitive Dark Matter Search with Super-high Resolution Nuclear Emulsion

Tatsuhiro Naka

KMI, Nagoya University@Japan
on behalf of NEWSdm collaboration



Direction sensitive search → new generation dark matter search experiment

- Essentially new systematics, not only annual modulation
- It has 100 times gain statistically for required # of signal to annual modulation search
- Discrimination between neutrino and dark matter using angle information
- Dark Matte Astronomy (dark matter distribution)

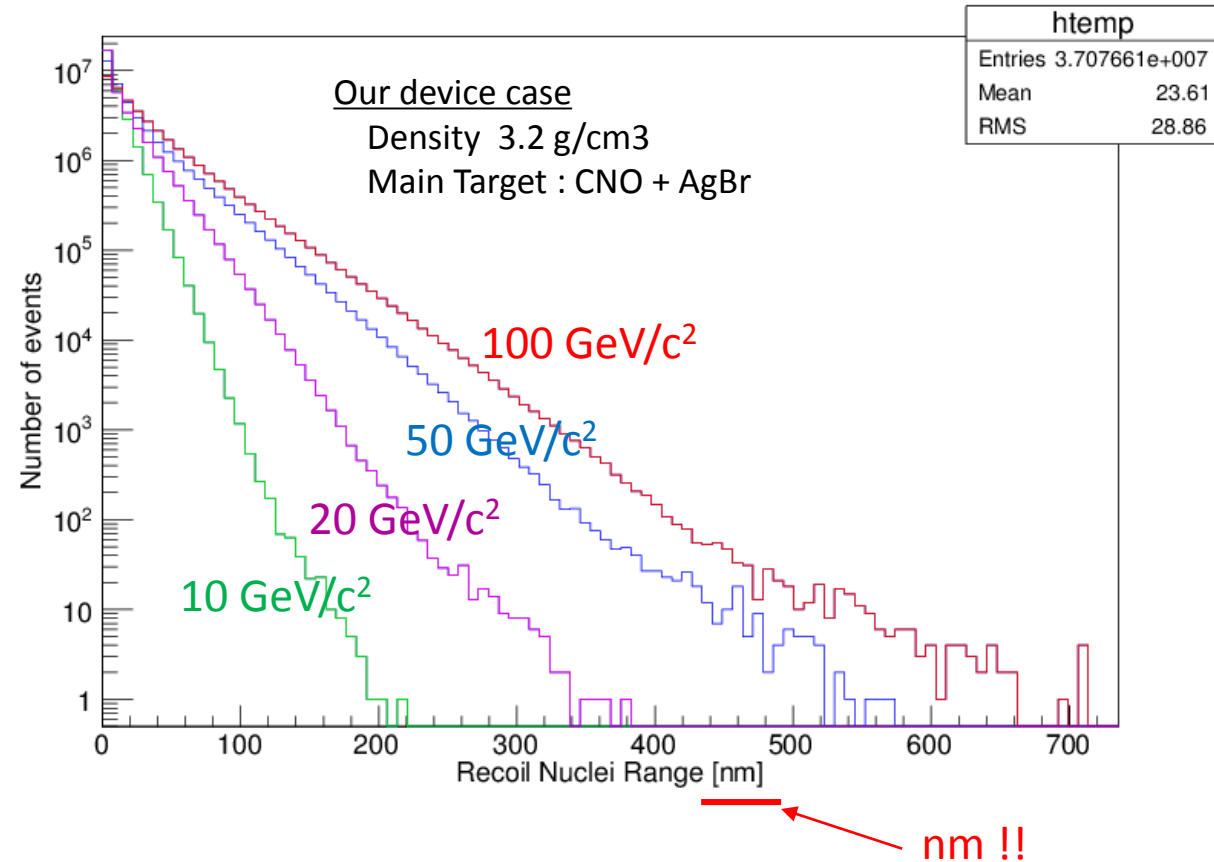
Challenge for Direction Sensitive Dark Matter technologies

Can the solid (or liquid) detector have directional sensitivity to nuclear recoil signal due to WIMPs ? (currently gaseous detector is on studied)

- Track length of recoiled nuclei $< \sim 1 \mu\text{m}$
- Angular dispersion due to straggling $\sim 25\text{deg.}$
- Scalability and low-background

New technical challenge !!

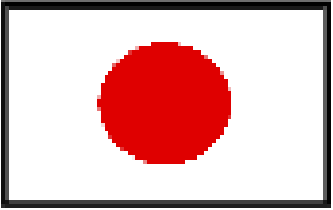





Low mass ($\sim 10 \text{ GeV}/c^2$) search : light target + $< 200 \text{ nm}$ length
High mass ($> 100 \text{ GeV}/c^2$) search : heavy target + $< \sim 700 \text{ nm}$



NEWSdm ~ Nuclear Emulsions for WIMP Search + directional measurement



<http://news-dm.lngs.infn.it>

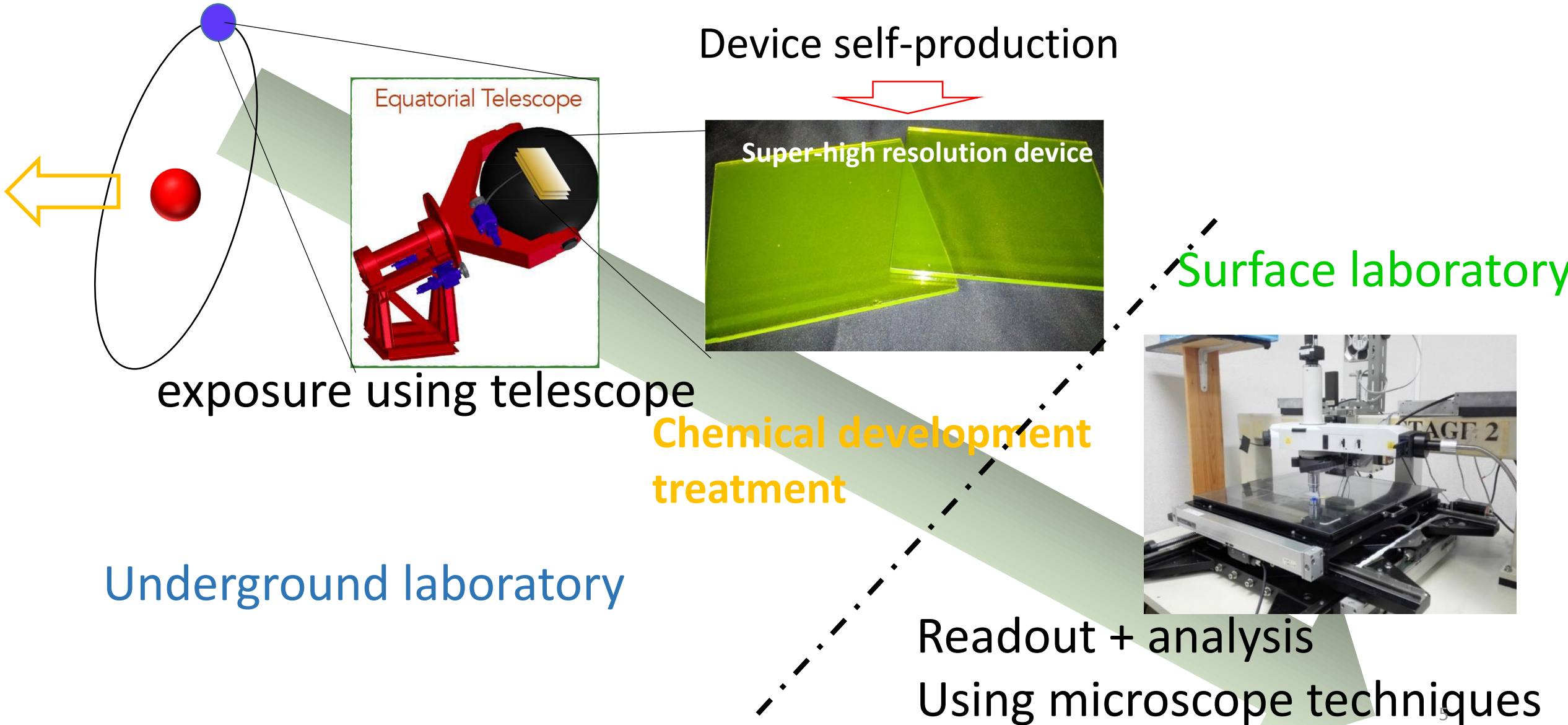
		
Chiba Nagoya		METU Ankara
		
Bari GSSI LNGS Napoli Roma	LPI RAS Moscow JINR Dubna SINP MSU Moscow INR Moscow Yandex School of Data Analysis	Gyeongsang

NEWS: Nuclear Emulsions for WIMP Search
Letter of Intent
(NEWS Collaboration)

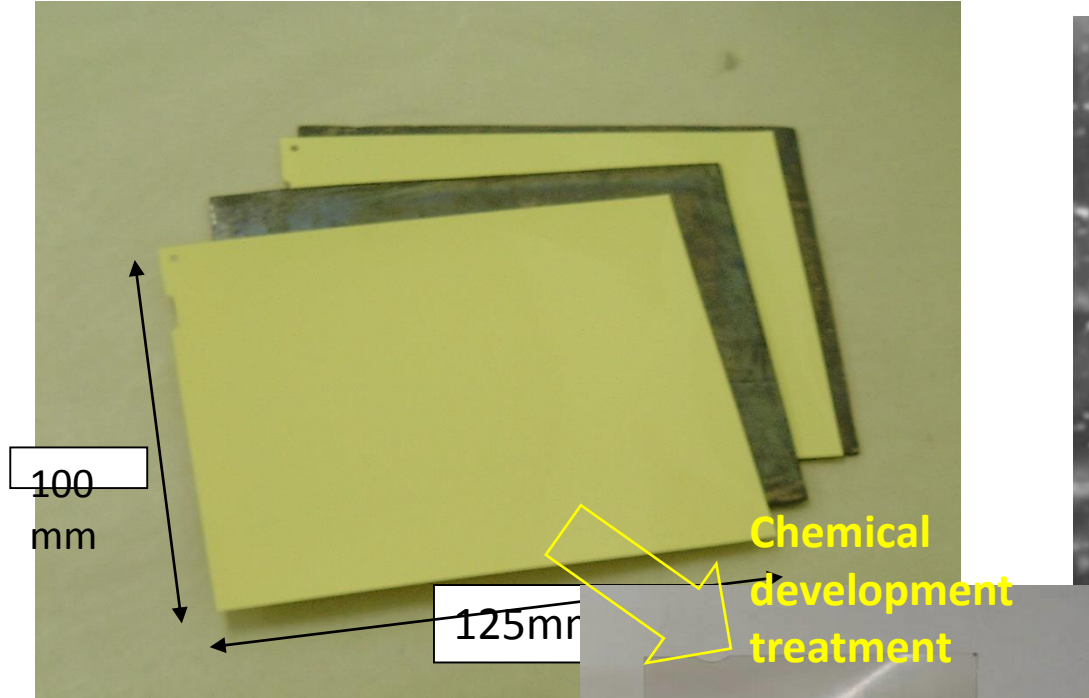
LOI under review by the LNGS science committee

<https://arxiv.org/abs/1604.04199>

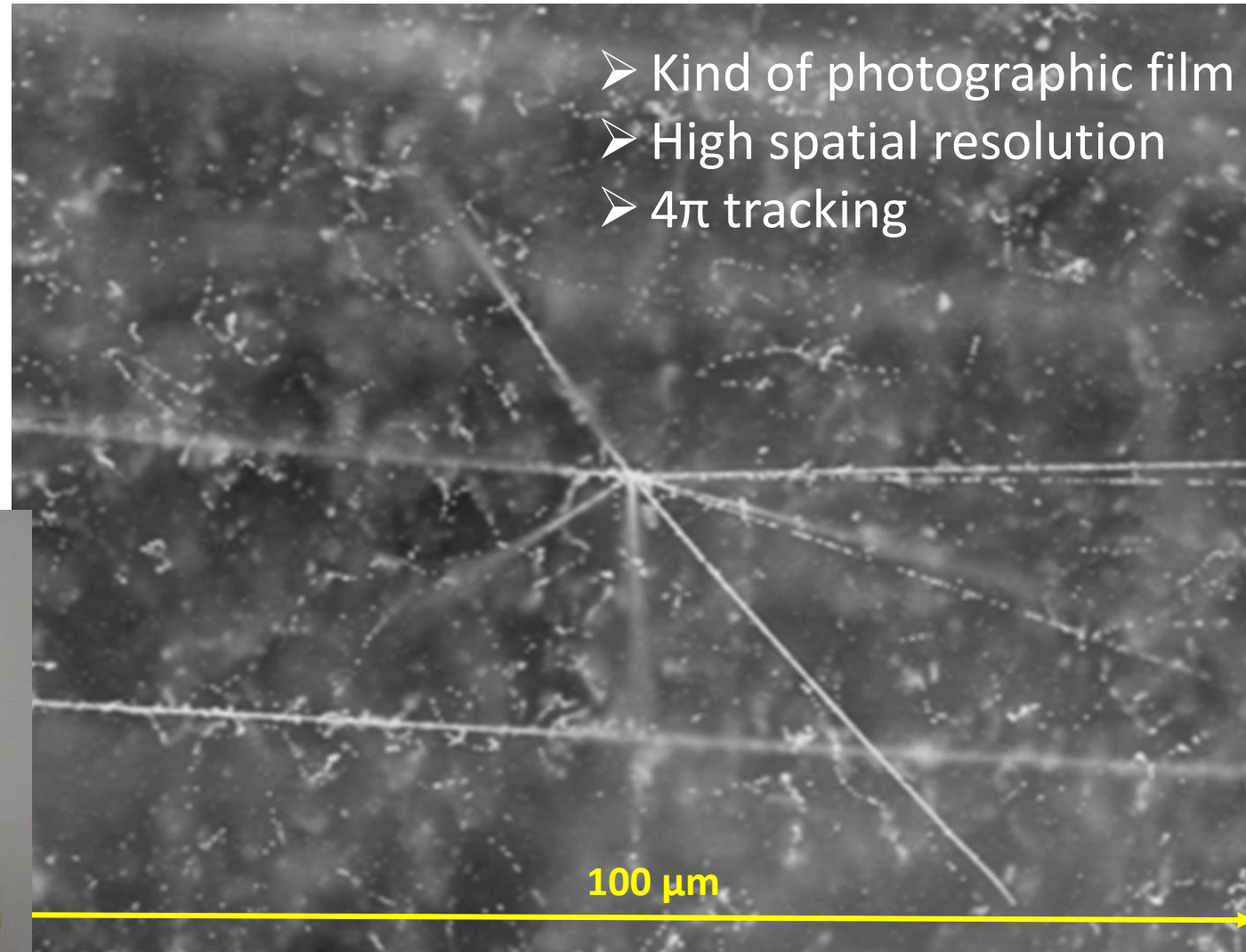
Concept of NEWSdm experiment



Nuclear Emulsion

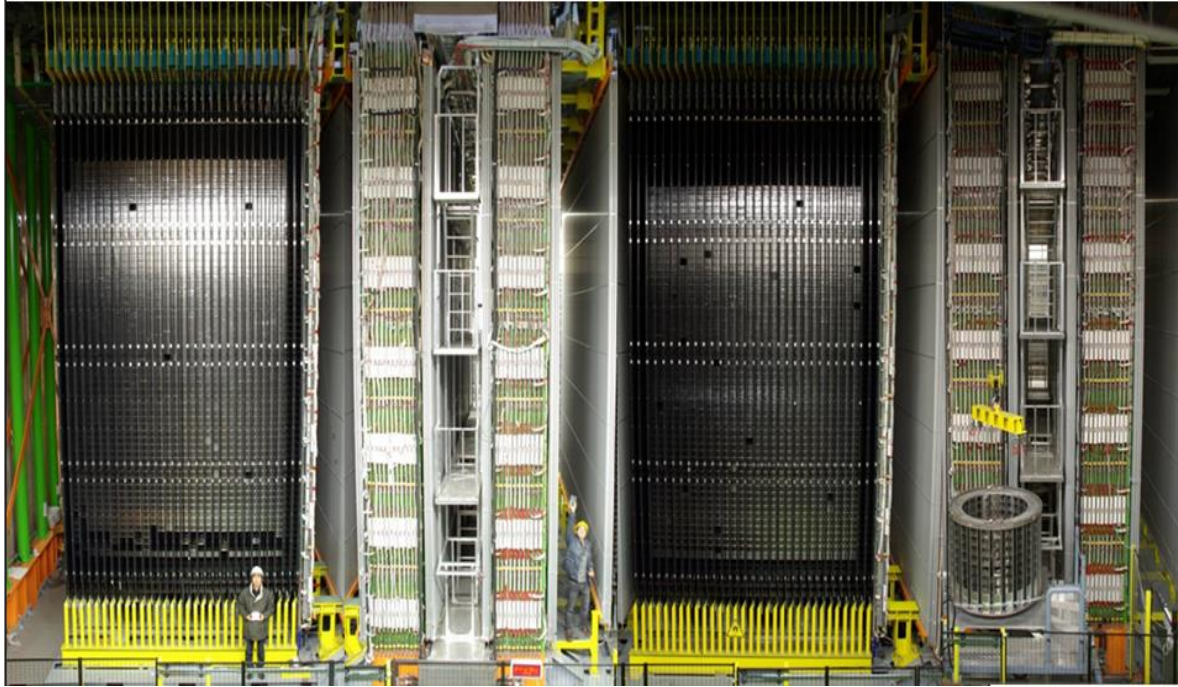


Solid tracking detector :
Density $\sim 3 \text{ g/cm}^3$



Latest nuclear emulsion experiment and readout

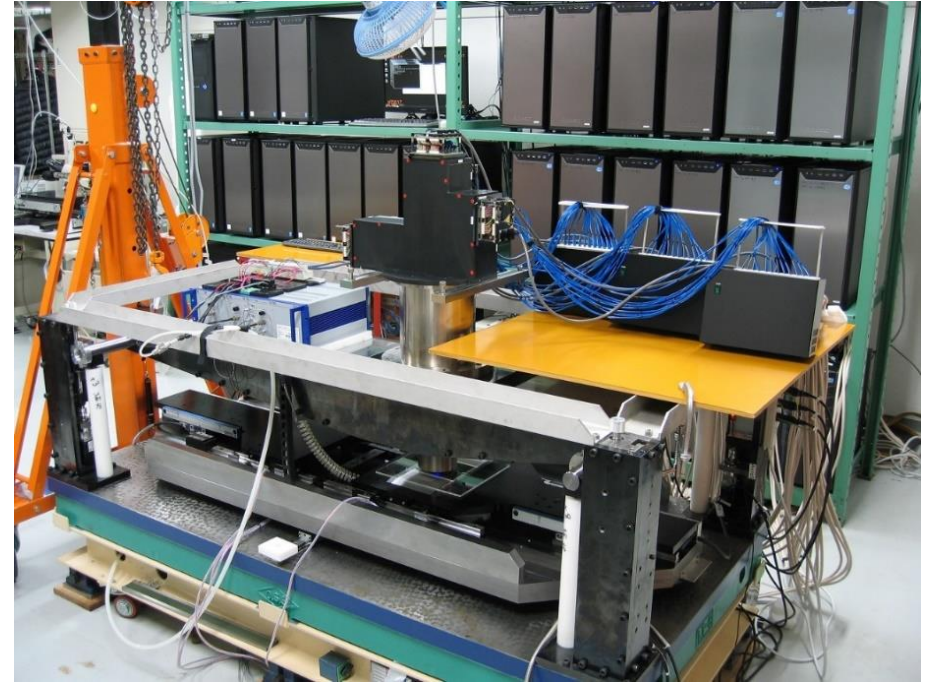
OPERA detector



20 m

Observed neutrino oscillation with
30 ton emulsion detector x 5 years (150 ton·year)
(Emulsions are 20 % volum in this picture)

Current highest speed readout system

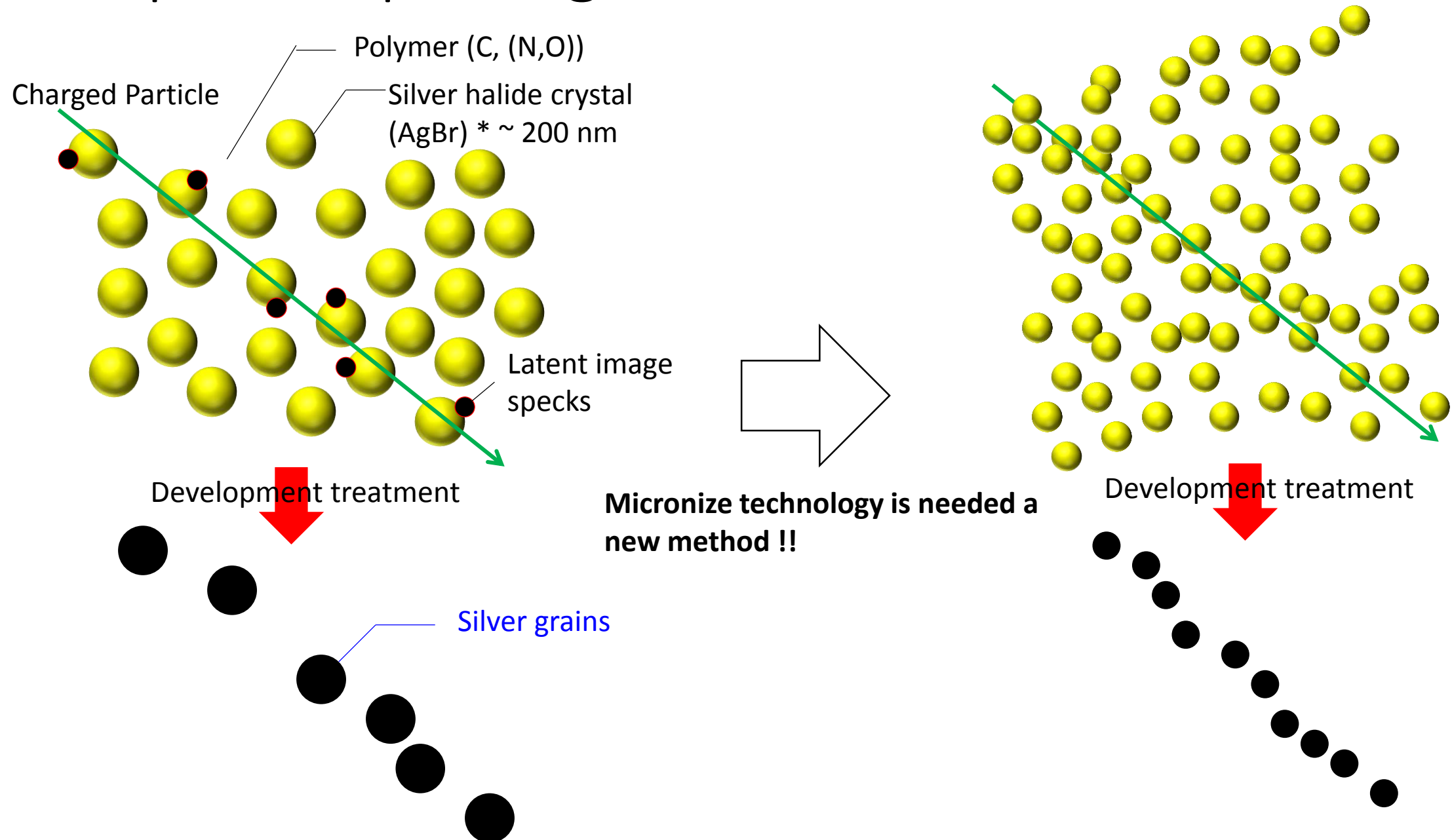


Scanning speed \sim several ton /year

Ref : M. Yoshimoto et al., [arXiv:1704.06814](https://arxiv.org/abs/1704.06814) [physics.ins-det]

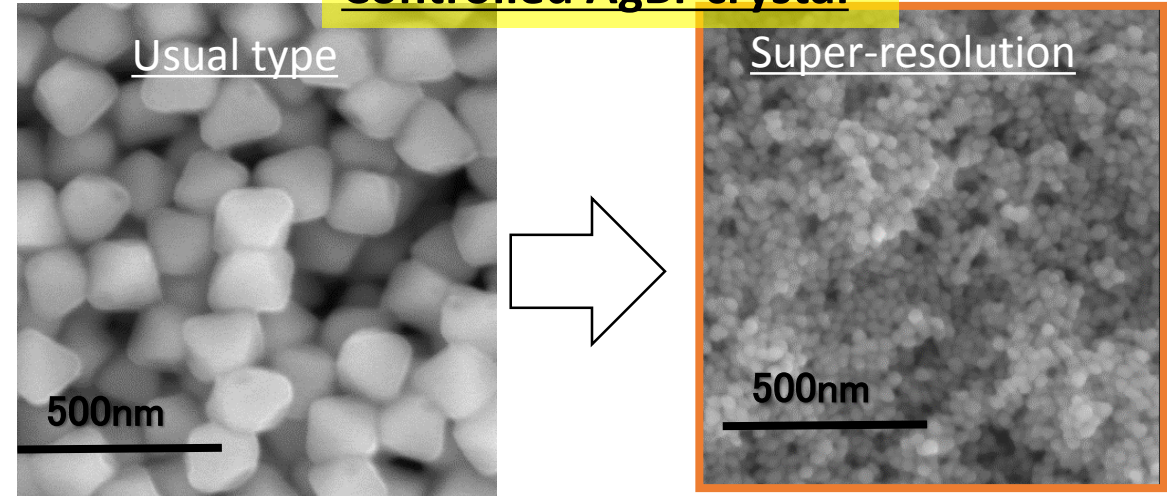
It has good potential for ton scale experiment !

Concept of super-high resolution

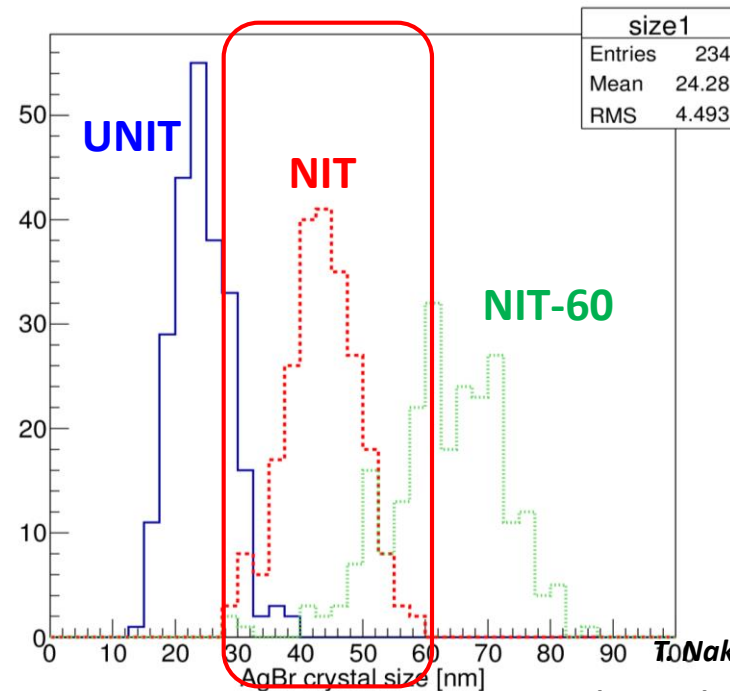


Self-production of Nano Imaging Tracker(NIT)

Controlled AgBr crystal



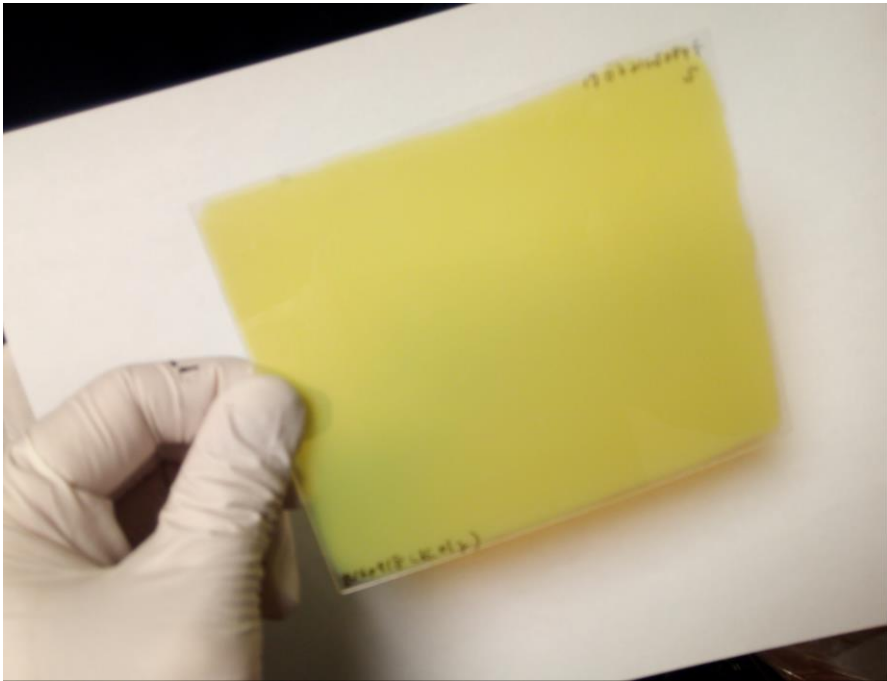
- Production time : 4-5 hours /batch
 - One butch : ~ 100 g (+ 300 g)
(there are 2 type machines)
- ⇒ kg scale production is possible using this machine.



Current standard Device :
Nano Imaging Tracker [NIT]
crystal size : 44 nm

Finest grain emulsion :
Ultra-NIT [UNIT]
crystal size : 25 nm

prototype NIT film for dark matter experiment



Elemental composition of NIT

For high-mass DM s

For low-mass DM

	Mass fraction	Atomic Fraction
Ag	0.44	0.10
Br	0.32	0.10
I	0.019	0.004
C	0.101	0.214
O	0.074	0.118
N	0.027	0.049
H	0.016	0.410
S, Na + others	~ 0.001	~ 0.001

◆ Intrinsic radioactivity :

U-238	Th-232	K-40	Ag-110m	C-14
27	6	35	(~400)	24000

[mBq/kg]

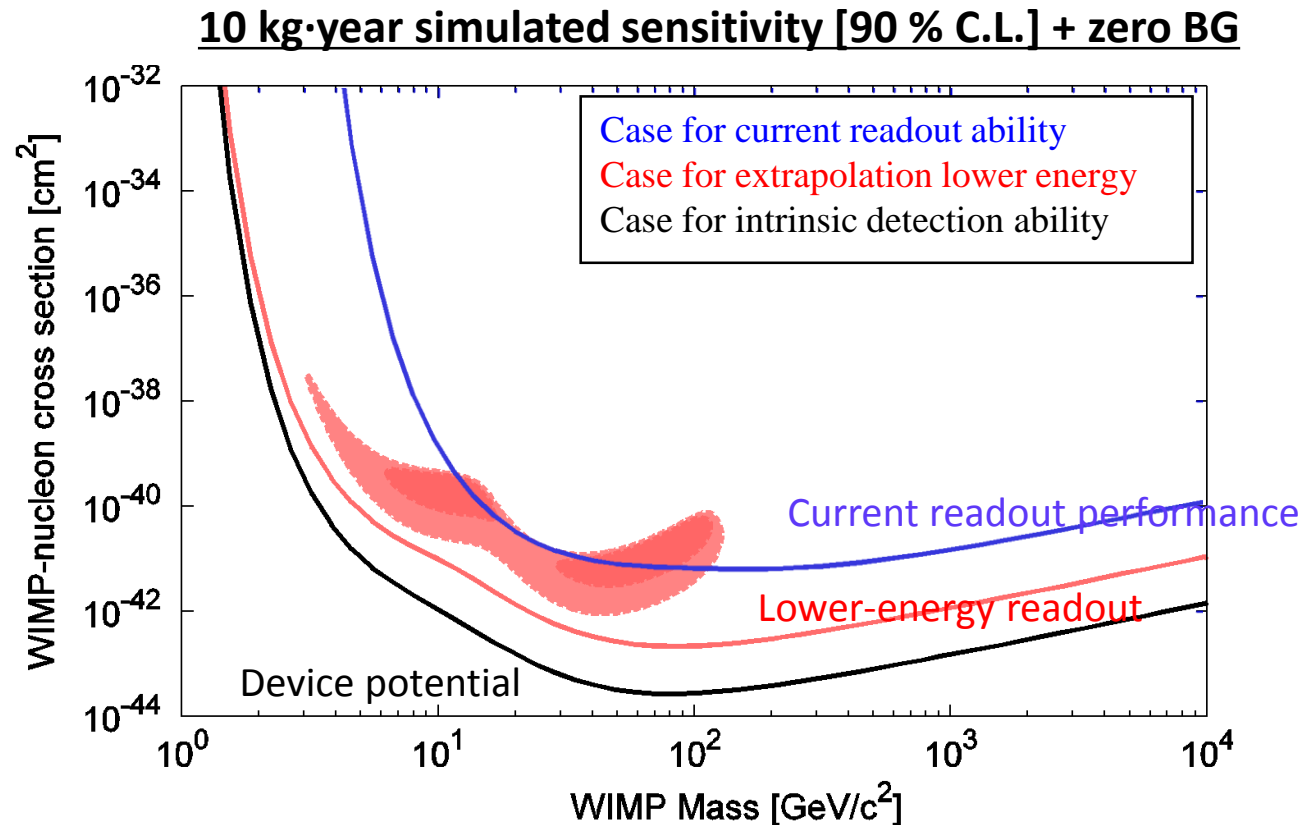
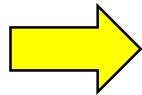
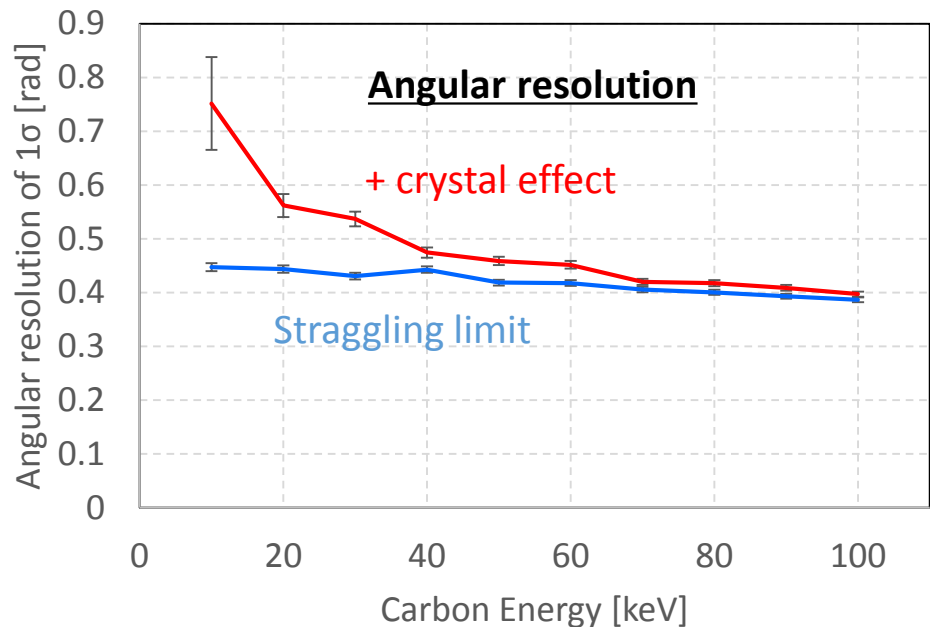
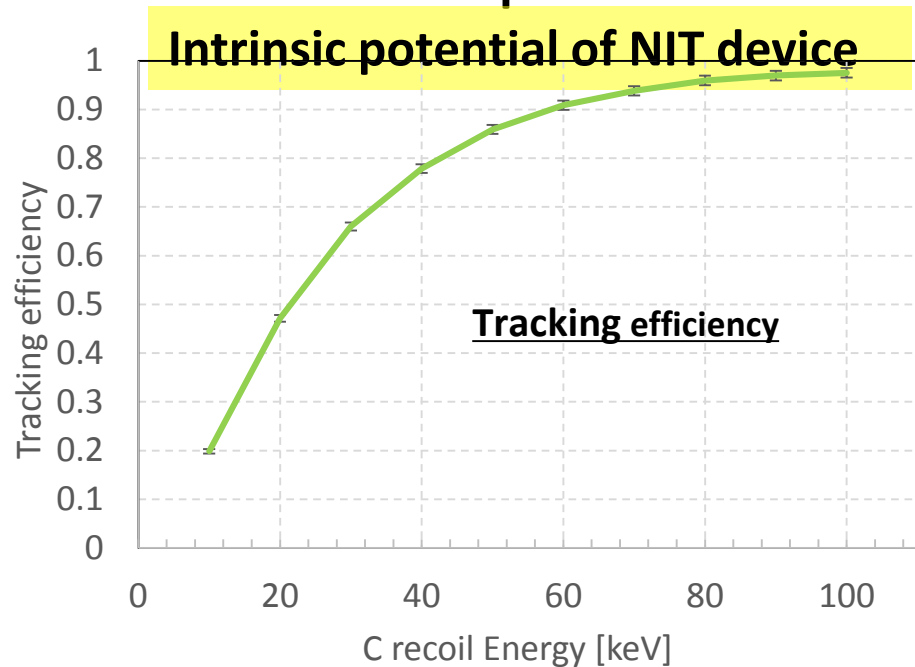
◆ Intrinsic neutron emission:

~ 1.2 /kg/y (by SOURCE simulation)

⇒ ~ **0.1 /kg/y (> 100 nm nuclear recoil)**

Detail shown in *Astropart. Phys.* 80 (2016)16-21

NIT device potential



NIT detector / CNO sensitive / no Bkg no directionality
 Simulation limit is “energy > 5 keV for all atoms (SRIM limit)”
 & “Sensitivity > 0.1 % (Simulation statistics limit; 10 event)”

Device potential : 10 keV of C recoil (> ~ 10% eff. and 45° angl. Res.)

Low-velocity ion tracking

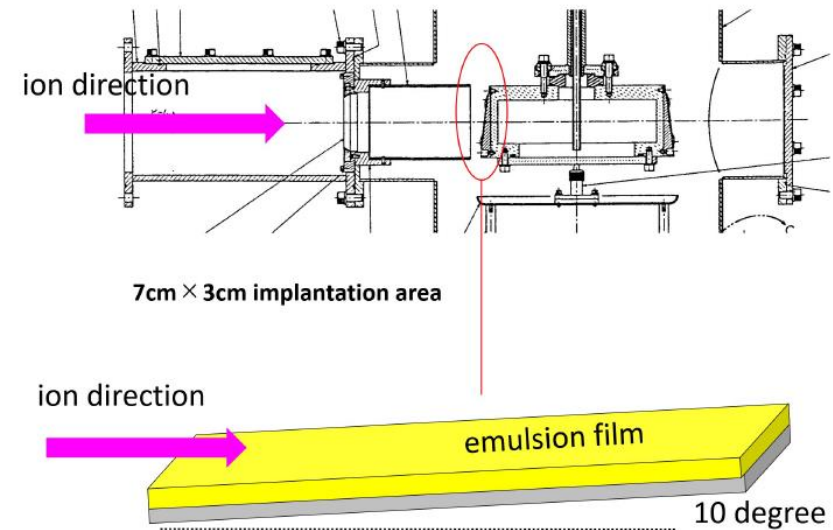
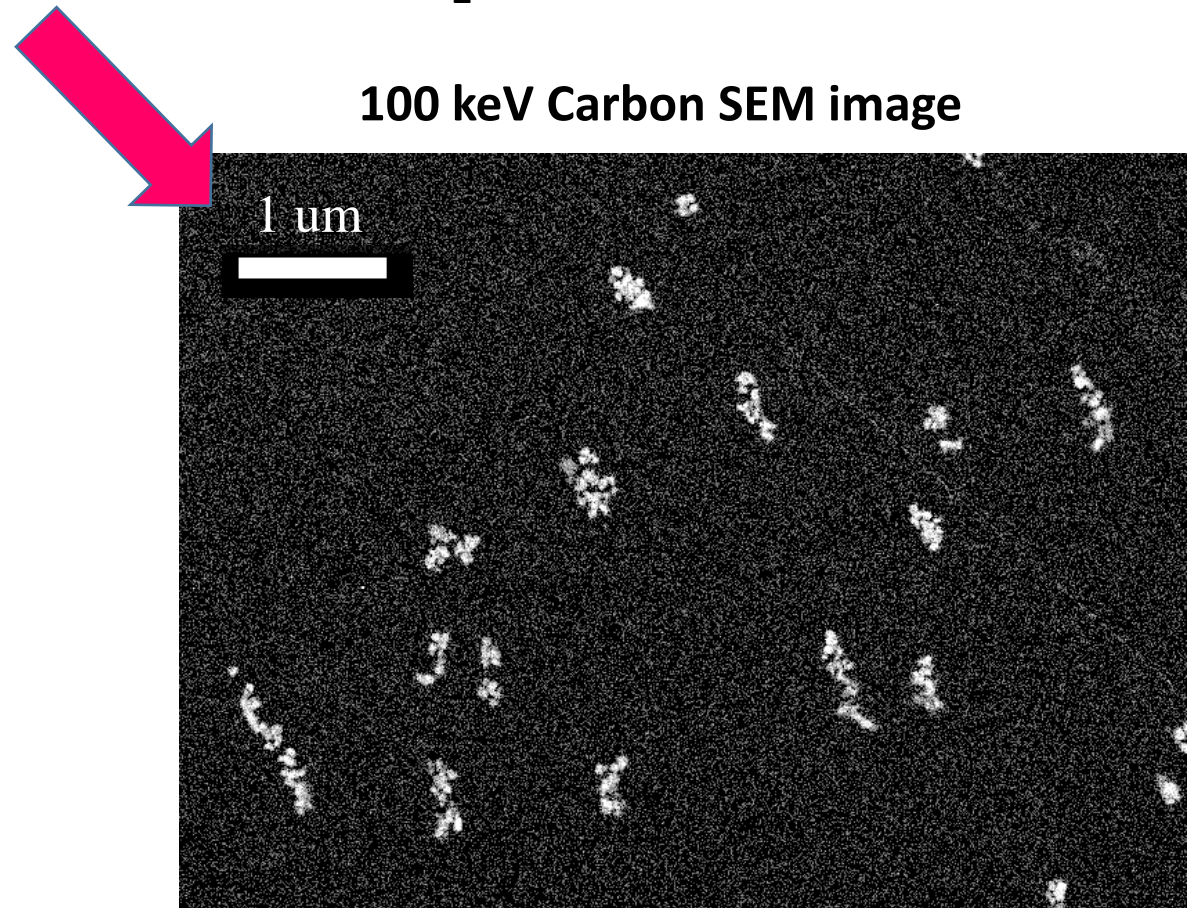
Can use ion implantation as calibration source

- Mono energy (± 0.1 keV)
- Good direction uniformity (<10 mrad)
- Now, C from CO_2 , Ar, Kr (but other various ion is possible)



Side view of ion

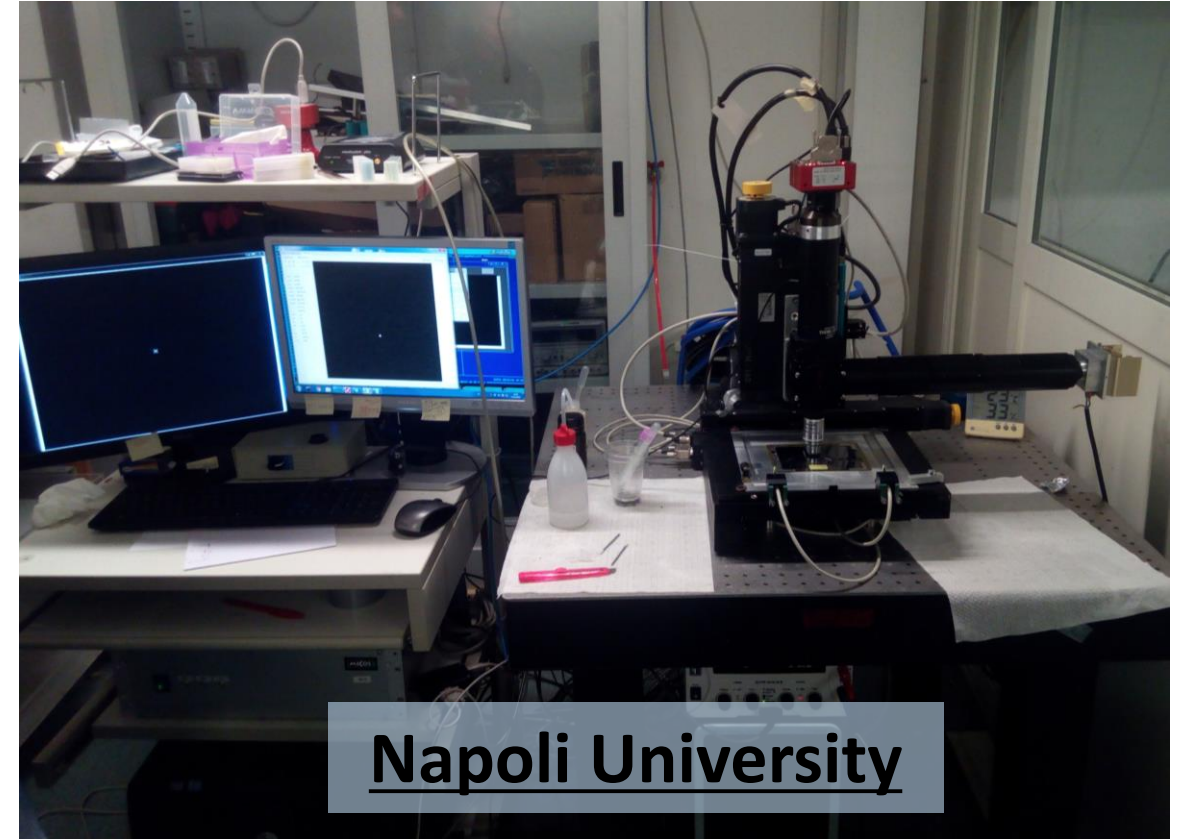
100 keV Carbon SEM image



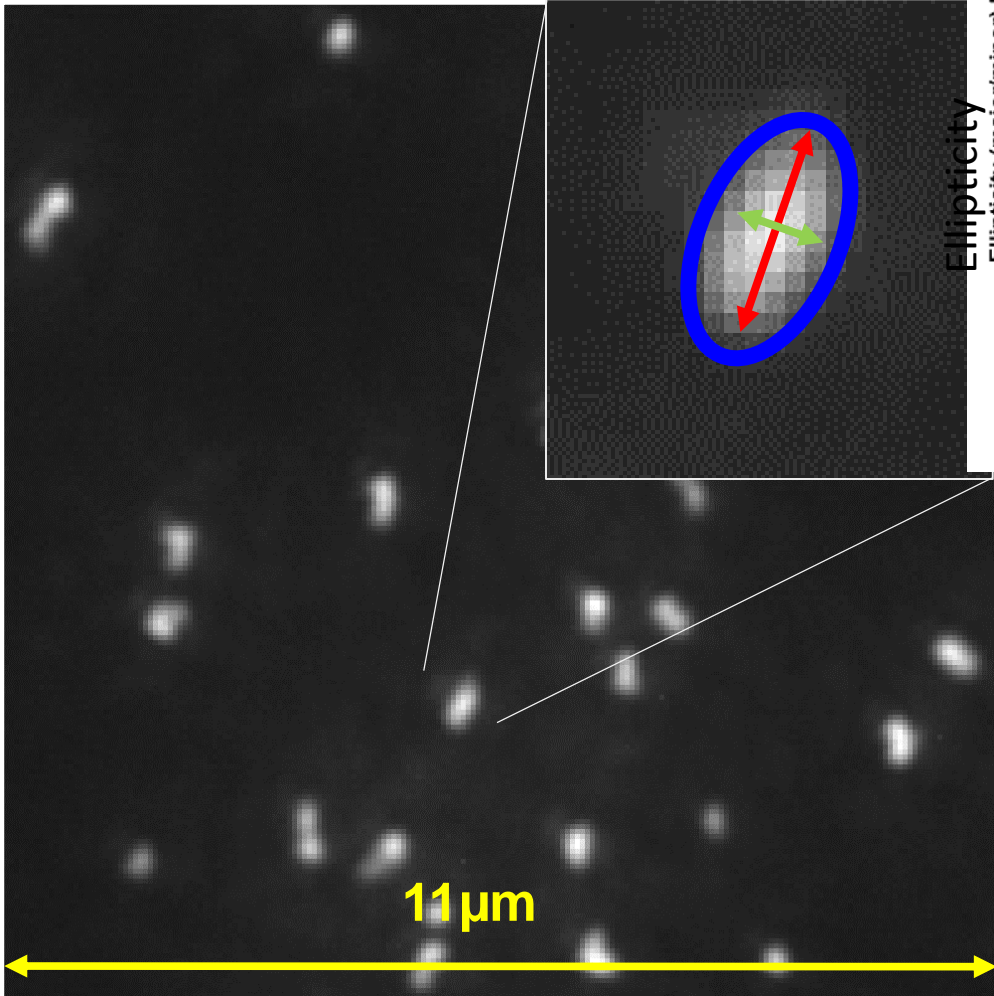
AgBr crystal has good sensitivity about Carbon ($\sim 100\%$ efficiency)

Development of New Readout System for nano-tracking

Prototype R&D system @Nagoya and Napoli



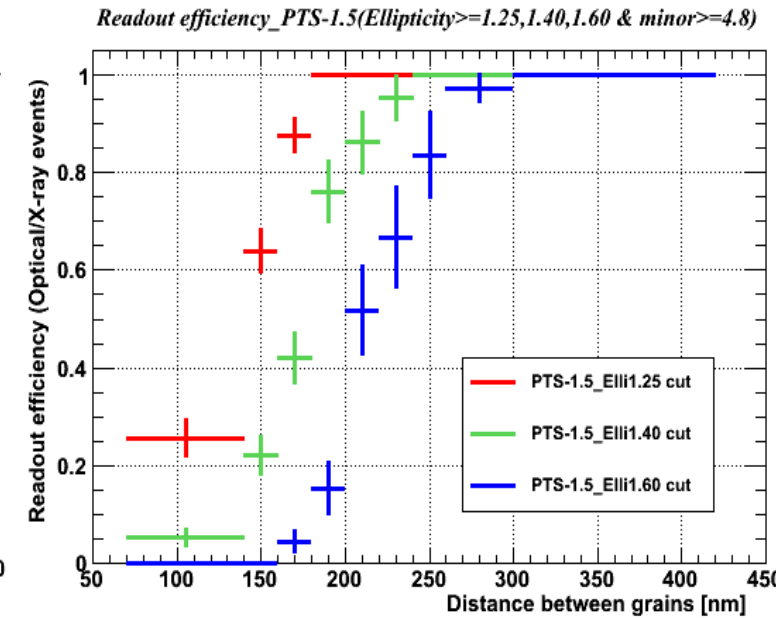
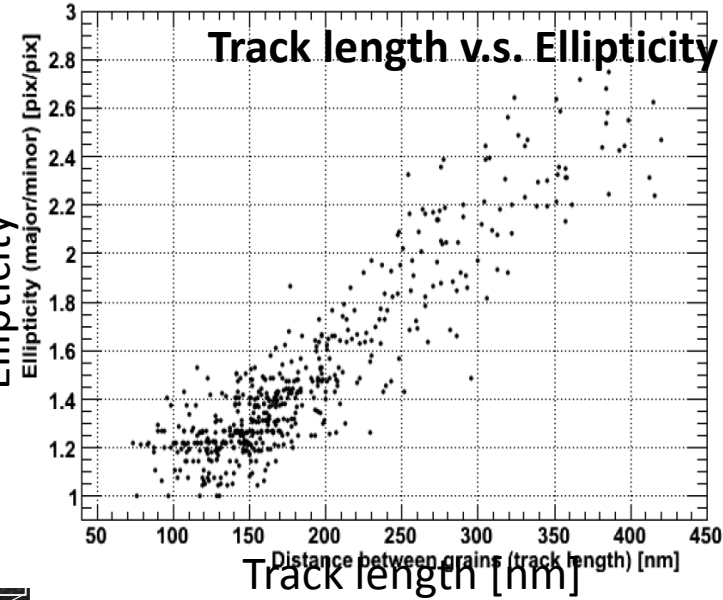
Candidate selection method using epi-illuminated optical microscope



K. Kimura and T. Naka, Nucl. Inst. Meth. A 680 (2012) 12-17

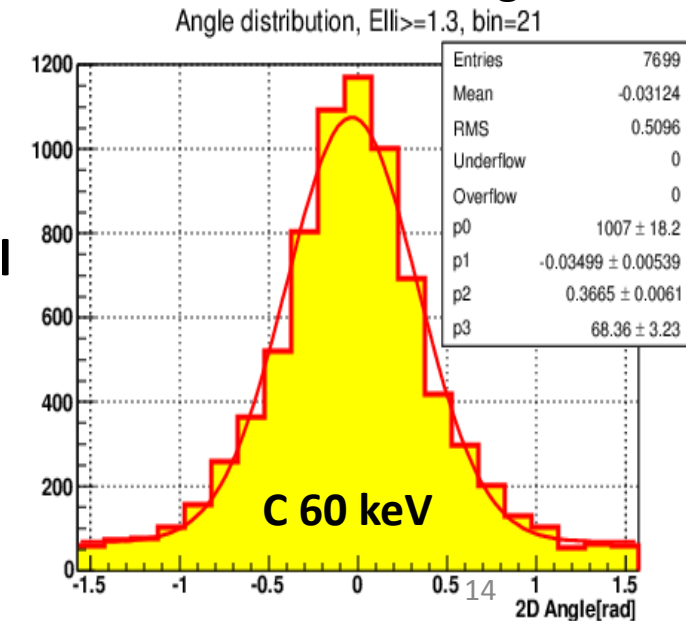
T. Katsuragawa et al, JINST 12 T04002 (2017)

Performance using only elliptical shape analysis

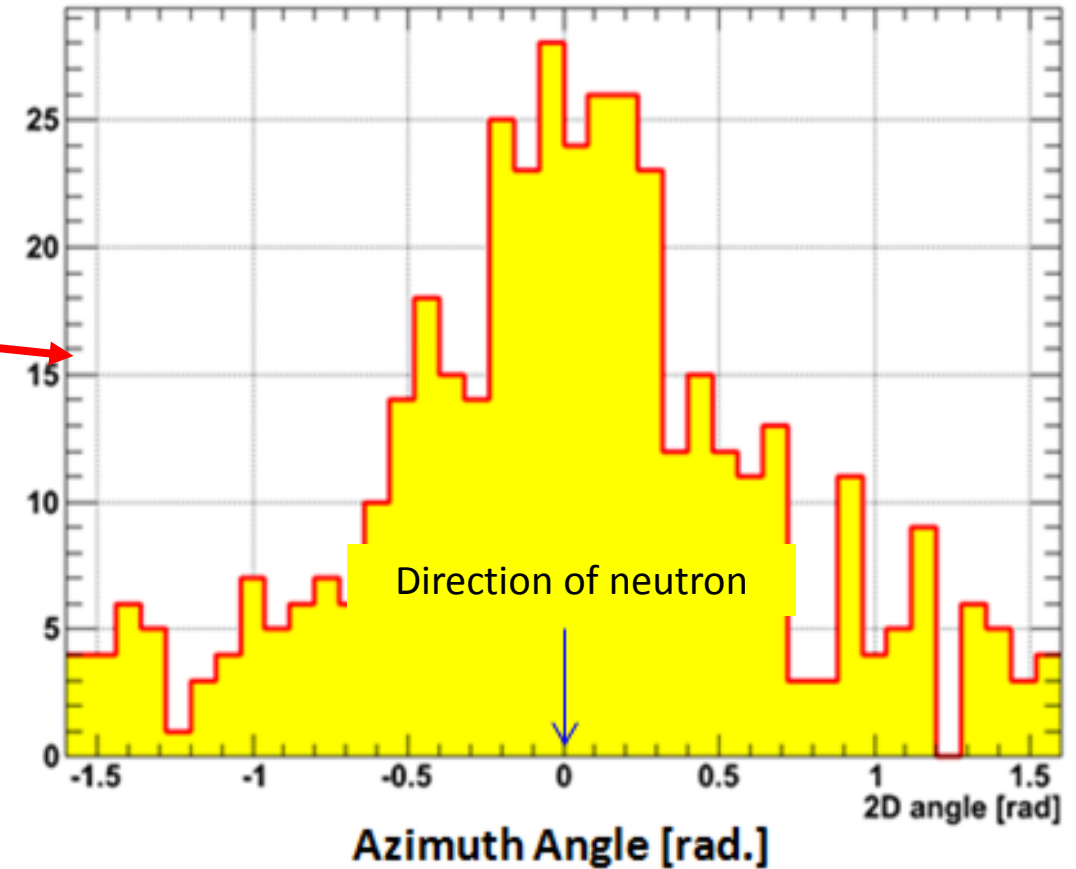
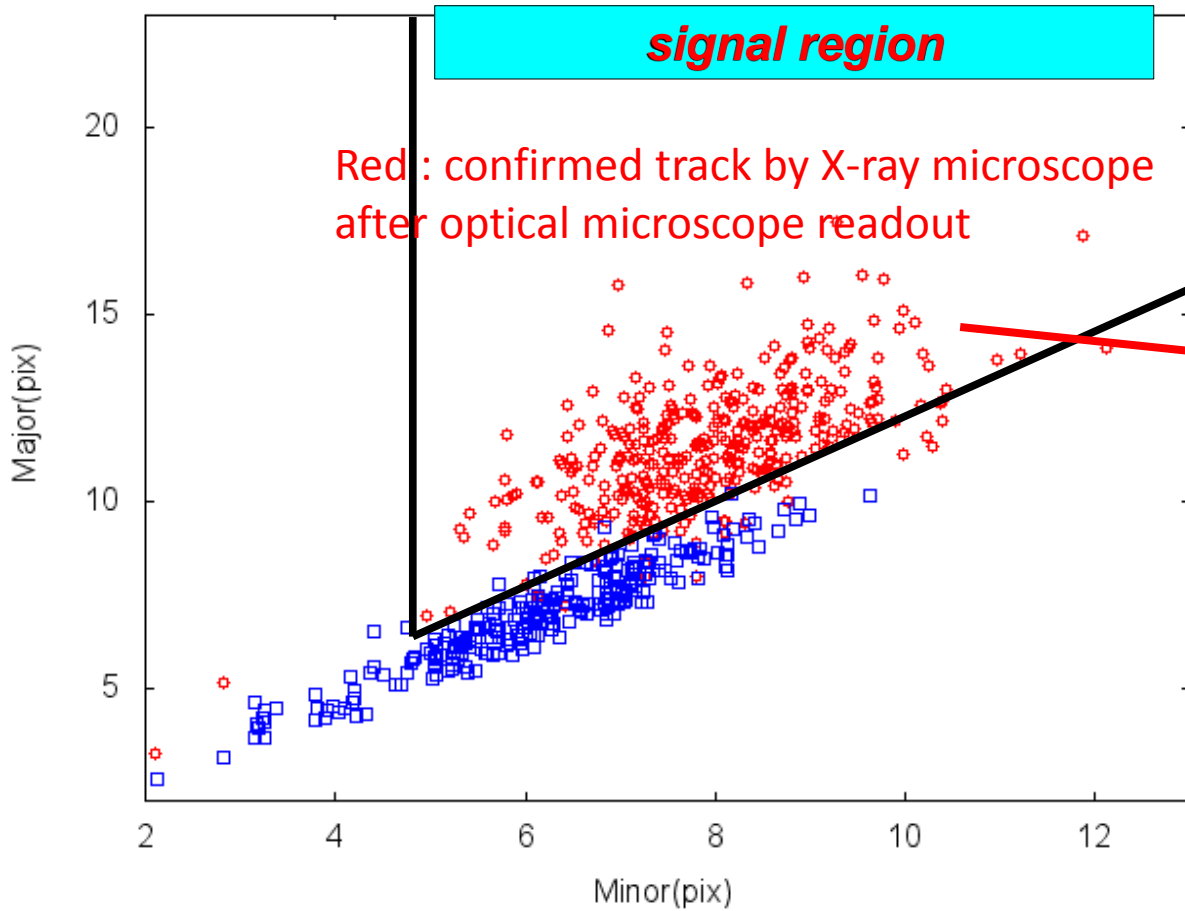


Current microscope has the potential to select > 100 nm length tracks

Direction sensitive eff. :
 ~30 % @60 keV
 (Currently limited by the optical condition \Rightarrow to be upgrade)
Angular resolution :
 ~30 deg. @60 keV
 \Rightarrow good compatibility to expectation by simulation



Demonstration of direction sensitive nuclear recoil detection due to 14.8 MeV neutrons



Mostly detected target was Br recoil [< 200 keV]

Now on preparing CNO recoil demonstration due to 565 keV (Li-p nuclear fission reaction)

Study of higher level event selection technique

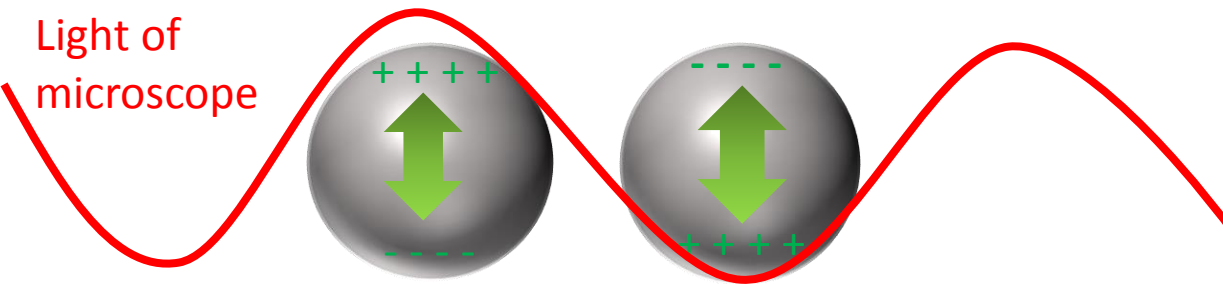
- For Signal/Background discrimination
- For lower energy threshold

- Super-resolution microscopy
- X-ray microscope (TN et al., Rev Sci Instrum. 2015 86(7):073701)
- Machine learning
- Phase-difference imaging
- etc.

**Cutting-edge technologies
for microscopy and image-
processing**

Localized Surface Plasmon Resonance (LSPR)

Localized Surface Plasmon Resonance



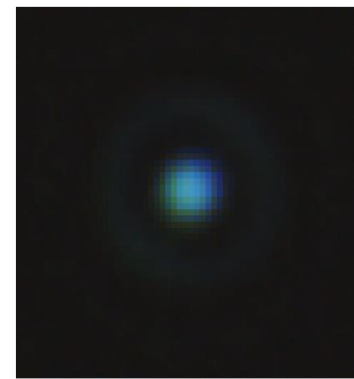
silver nano particle

$$p = 4\pi\epsilon_m a^3 \frac{\epsilon_1(\lambda) - \epsilon_m(\lambda)}{\epsilon_1(\lambda) + 2\epsilon_m(\lambda)} E_0$$

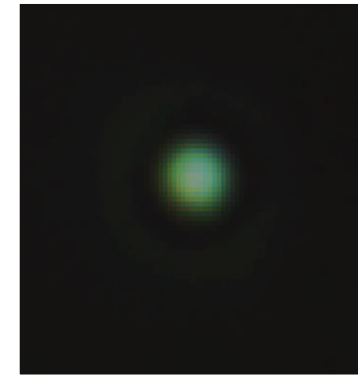
$$\epsilon_1(\lambda_l) + 2\epsilon_m(\lambda_l) \approx 0$$

Plasmon Resonance condition

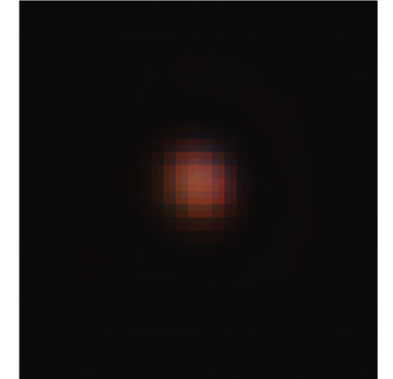
Silver-nano particle



40 nm



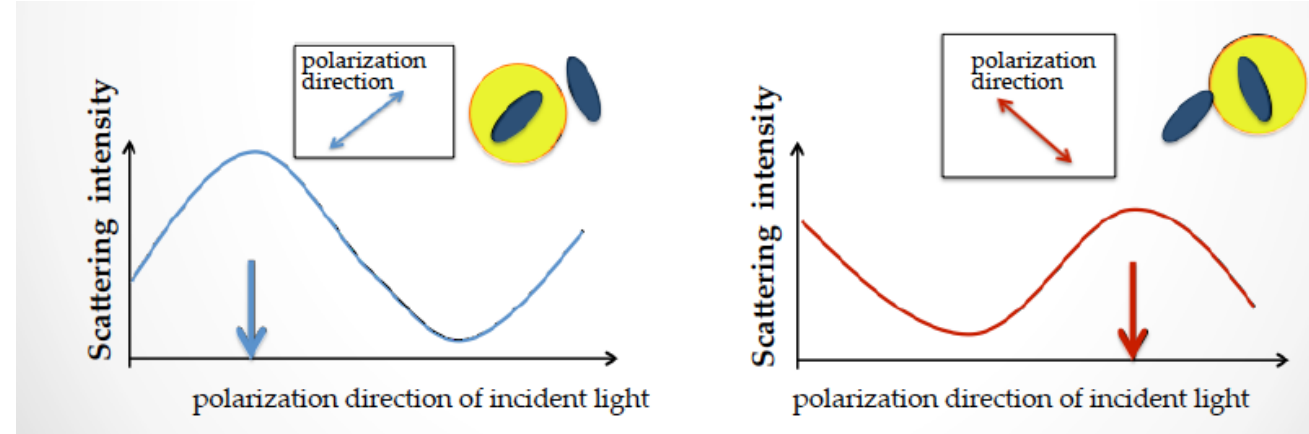
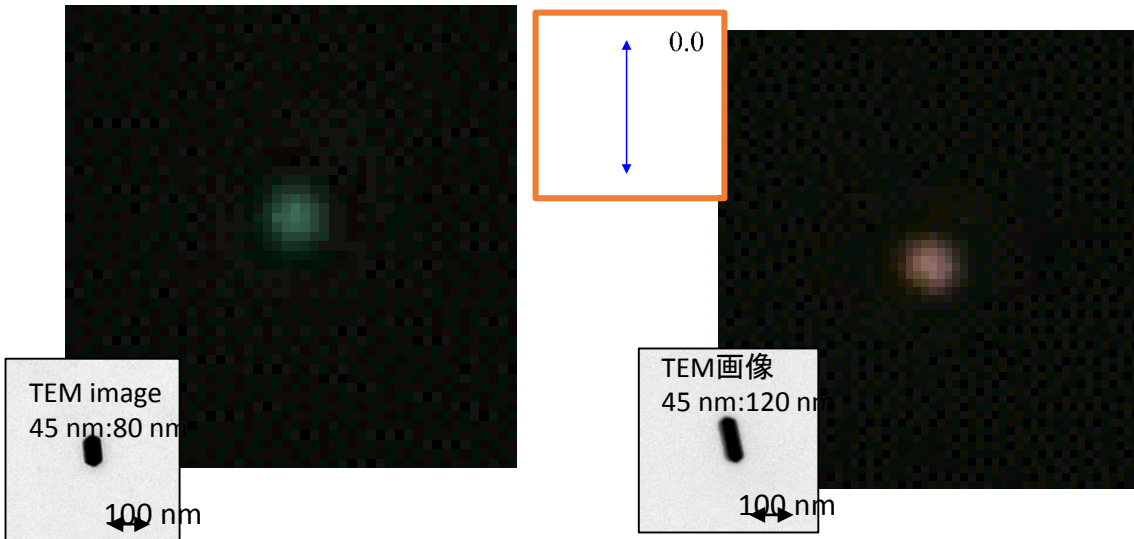
80 nm



120 nm

Recoiled proton track
due to neutron

Concept of super-resolution microscopy using LSPR



By combination both the wavelength shift and rotation of polarized angle, non-diffractive condition is realized.
 optical resolution $\hat{=}$ position accuracy
 (~ 10 nm or less)

- Dipole moment for non-spherical Ag nano particle depends on the polarization angle
- Resonance wavelength is shifted by that.

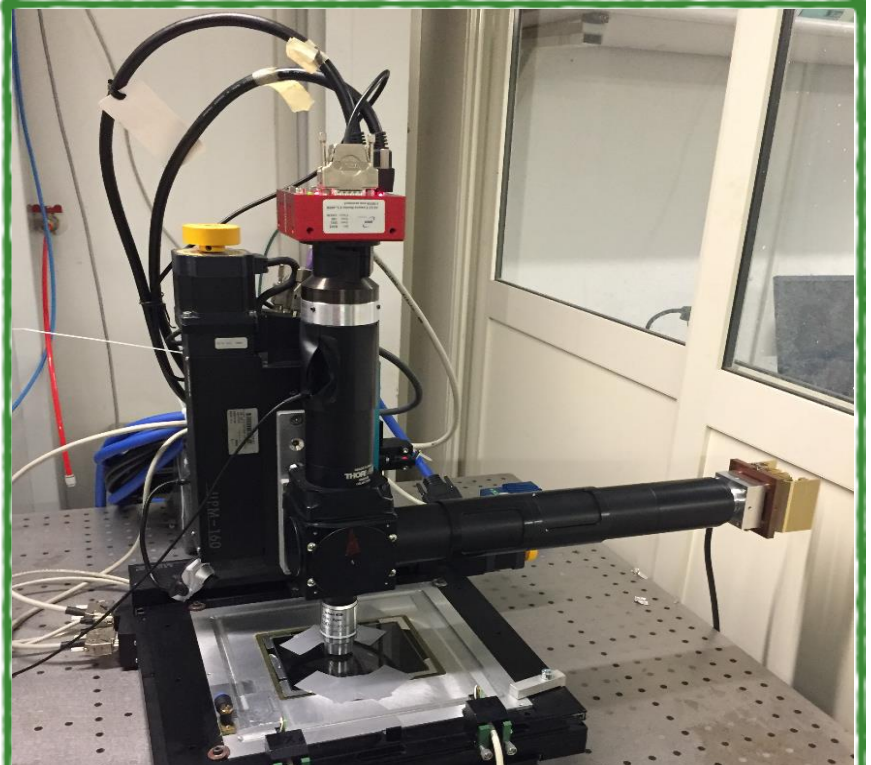


Electron microscope image

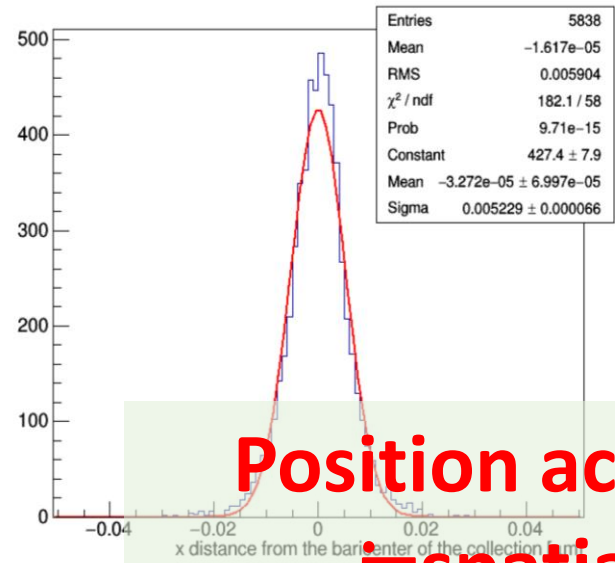
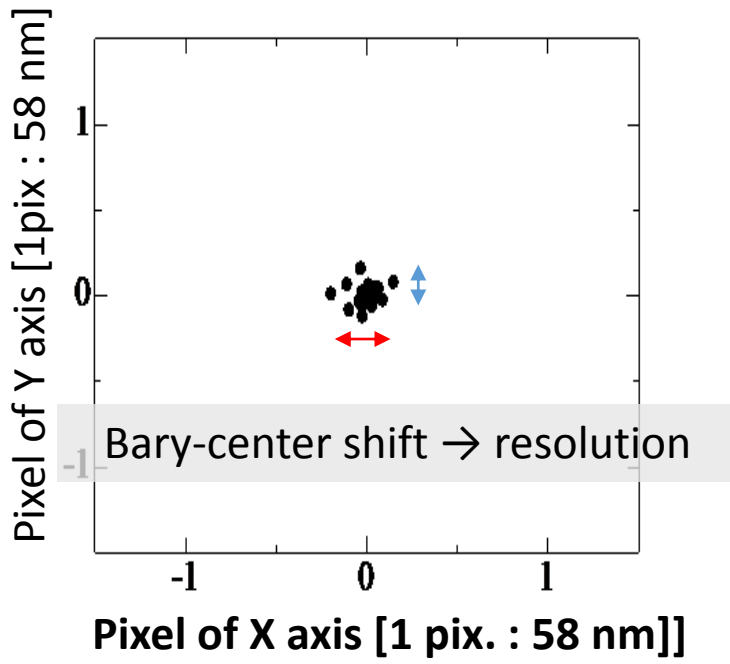
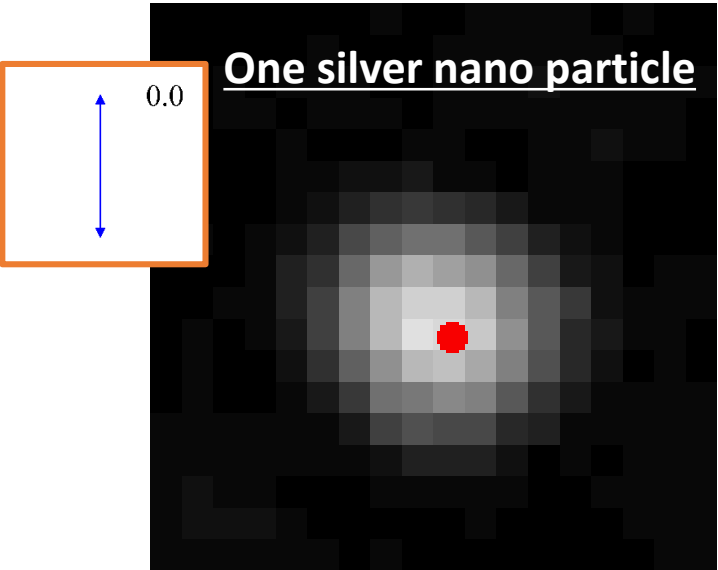
Developed silver grain such as the nuclear recoil has very complicated silver filament structure

\Rightarrow It's very good to realize above concept

Resolution calibration



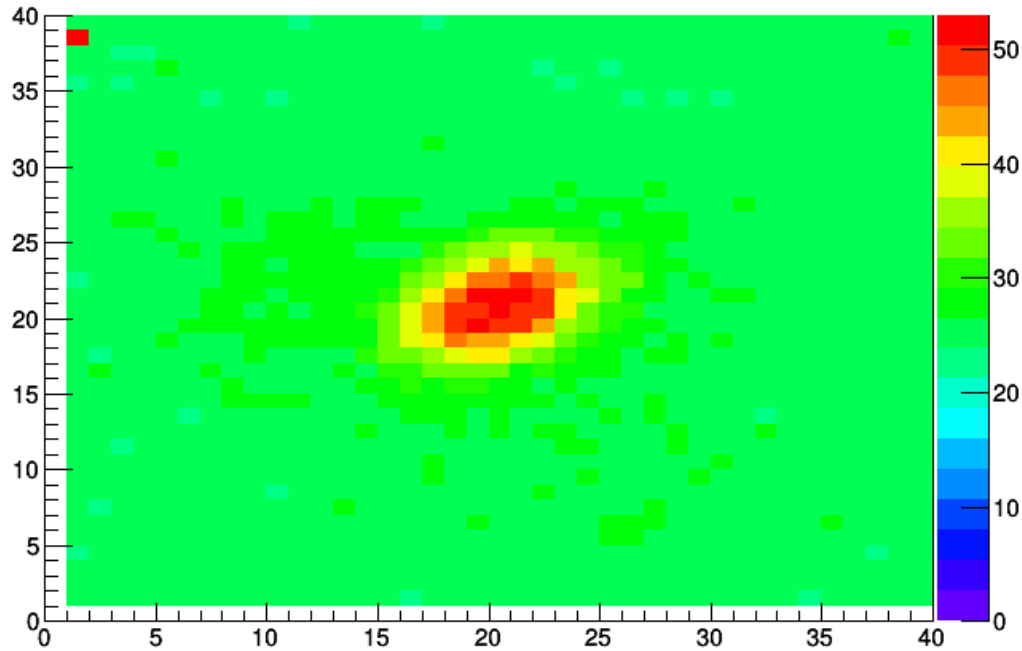
- Automatic plasmon analysis microscope system
- Liquid crystal rotater
 - suppression of stage vibration
 - (▪ color imaging)
 - (▪ combine machine learning)



Position accuracy ~ 5 nm!!
≡ spatial resolution

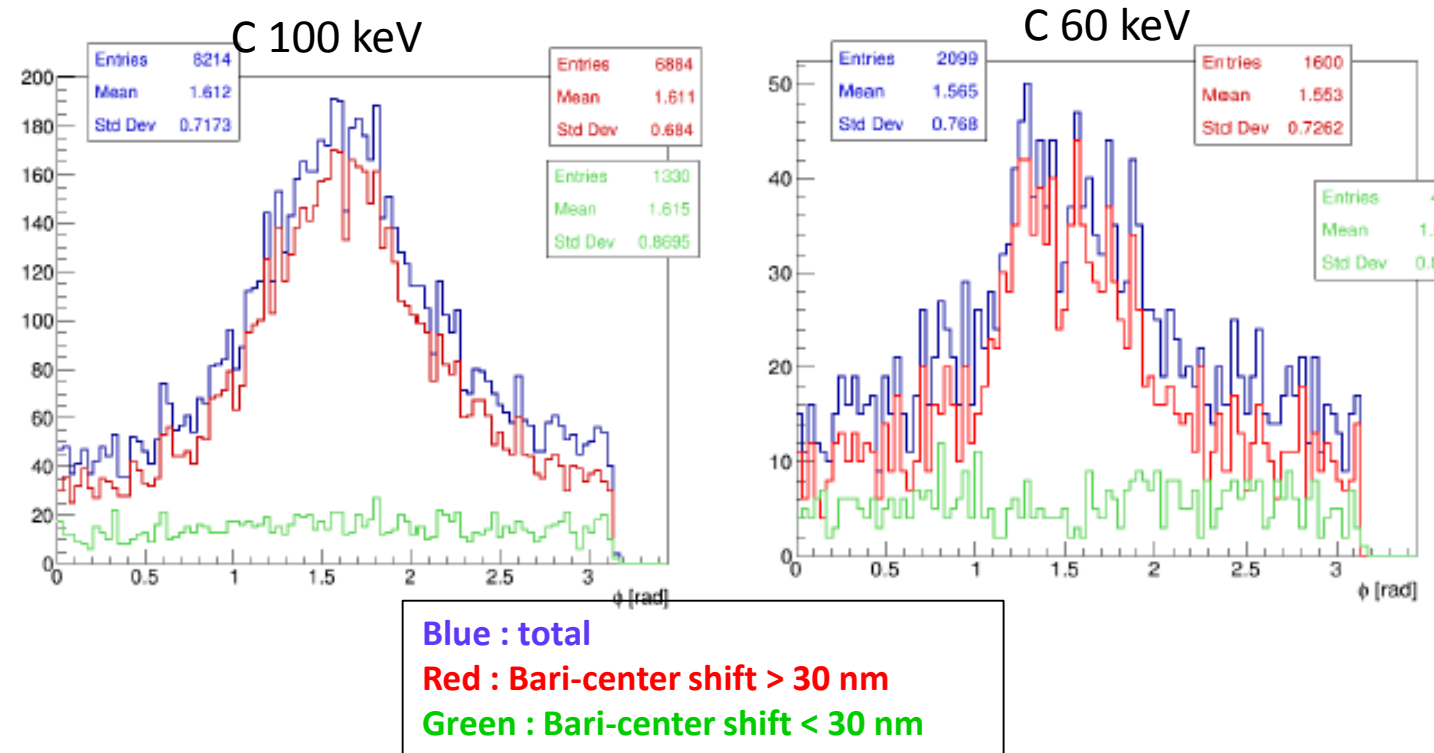
Readout study using the plasmonics information

cl 3474 in frame 140 at xy: -4.46 11.04



Shift of barycenter is important information for nano-scale structure

Direction sensitivity using plasmon analysis

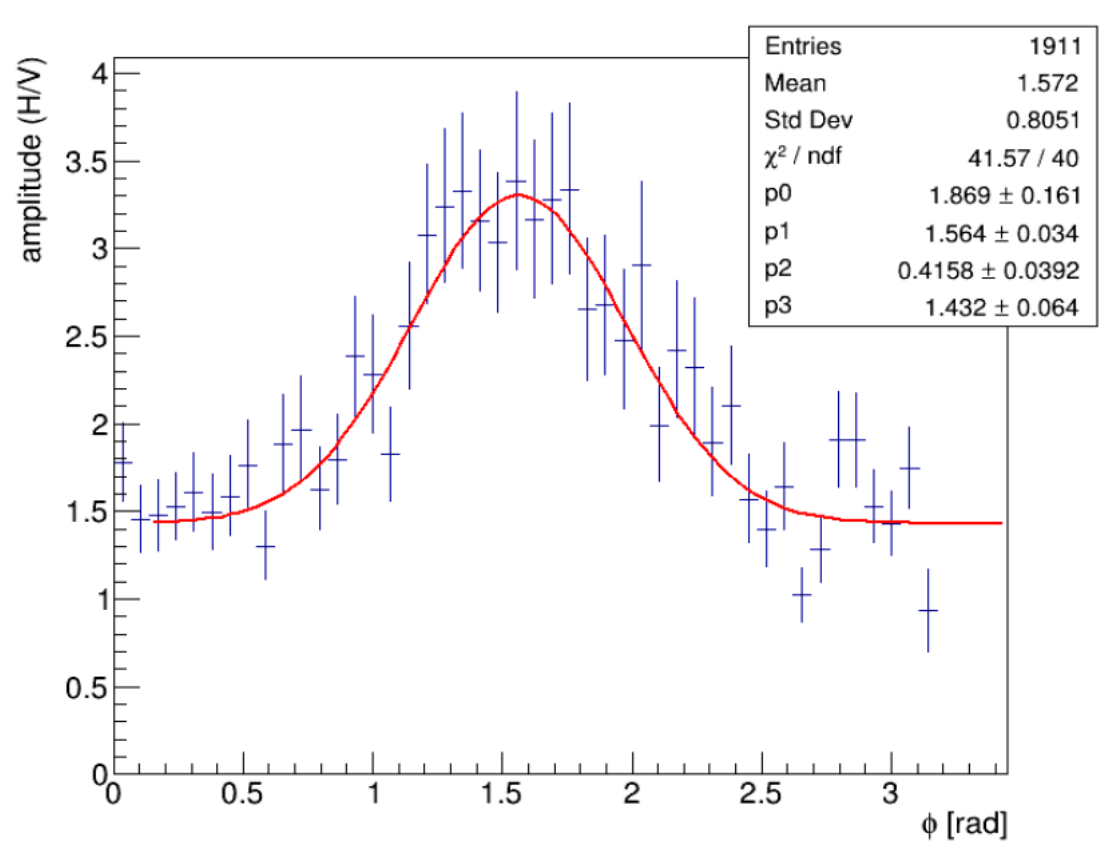


Demonstration of the direction sensitivity have been done .

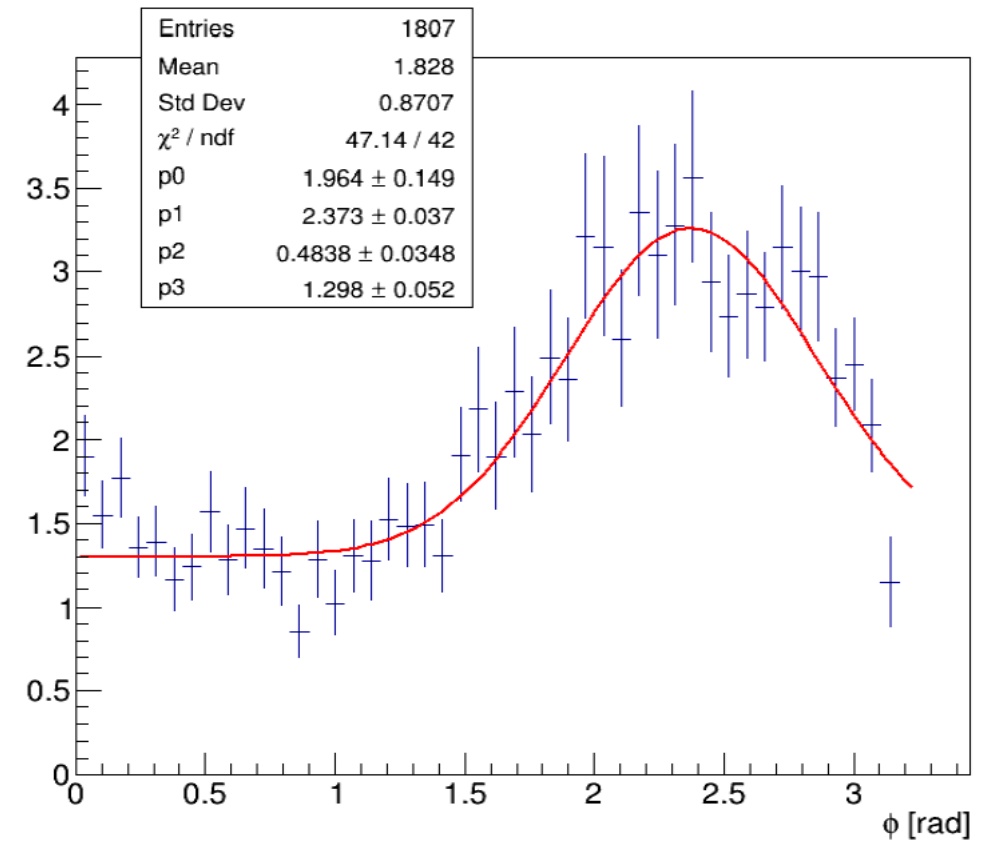
Preliminary

Direction sensitivity of low-energy C ion [30 keV]

Expected : 90° -> Measured : 90°



Expected : 135° -> Measured : 136°



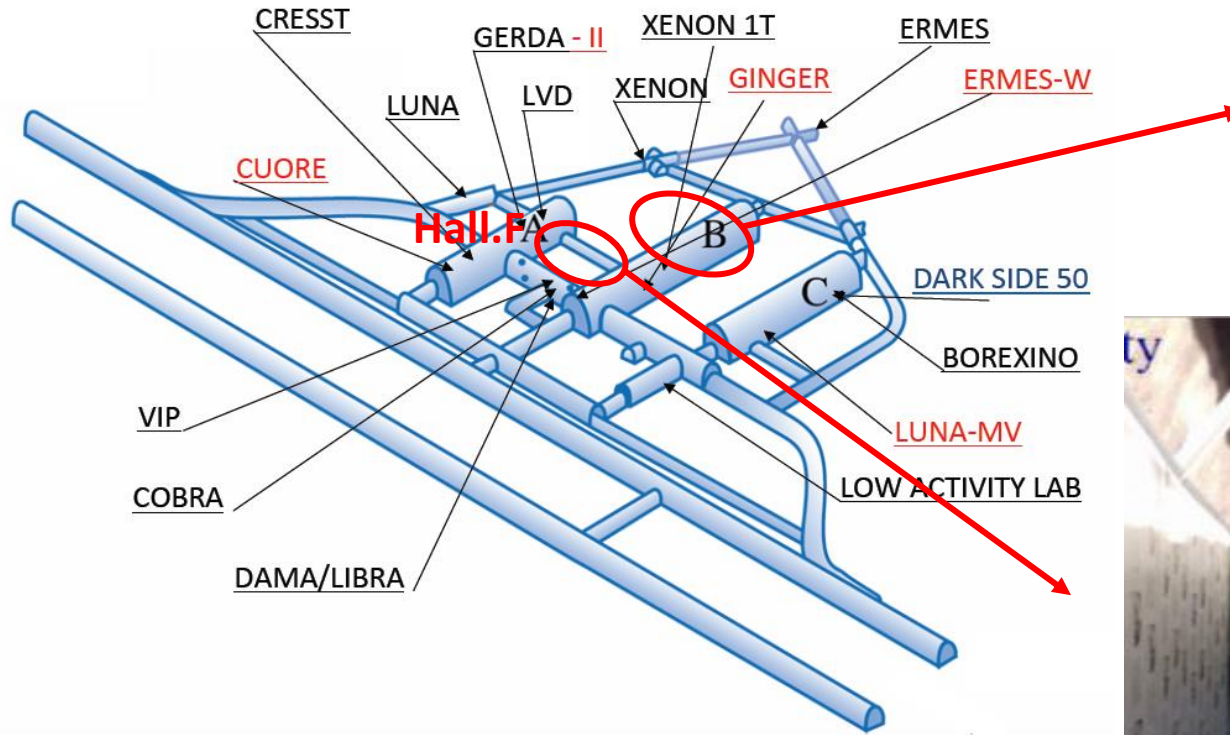
Indication :: low-mass dark matter (< 10 GeV/c²) can be seen with the direction !

Background studying

	Main source	Technologies	Expected rejection power or event rate
Physical BG			
Electrons	C-14 β Environment gamma	Crystal temperature dependence (<i>M. Kimura et al., NIM A 845 (2017) 373</i>) Crystal sensitivity control Image and plasmonic analysis	(> 10^6 or more rejection power (< $O(1)$ /kg/day)) *now on studying
		Synthetic Polymer	> 10^3 or more
Neutron	Intrinsic (α, n)	-	$\sim 3 \times 10^{-4}$ /kg/day or less Astropart. Phys. 80 (2016)16-21
	Environment	Water shield	< $1E-4$ /kg/day
Cosmic-ray	Recoiled nuclei	Coincidence with MIP sensitive emulsion	*on studying using simulation
	Spallation neutron	(under studying with simulation)	($\sim O(10^{-4})$)/kg/day * now on study)
Nonphysical BG			
Contaminated dust	(under studying)	Clean room Plasmonic analysis and image processing Machine learning Chemical treatment	Under studying (at least > 10^6 or more, in principle it should not be background)

Test experiment Site

Gran Sasso underground laboratory, Italy



Device Production and handling facility



Now on constructing the device production facility at Hall.F, LNGS

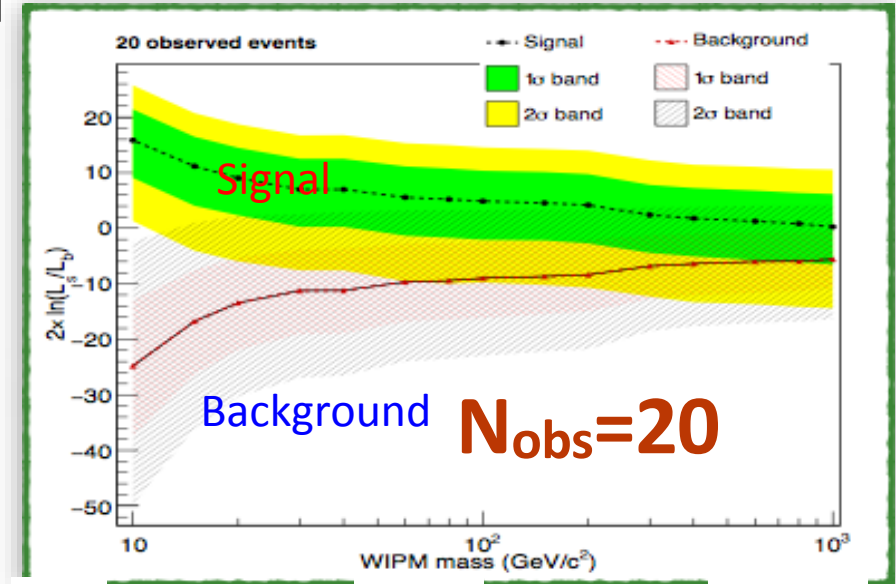
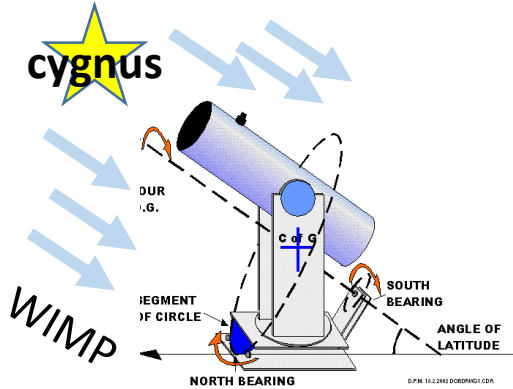
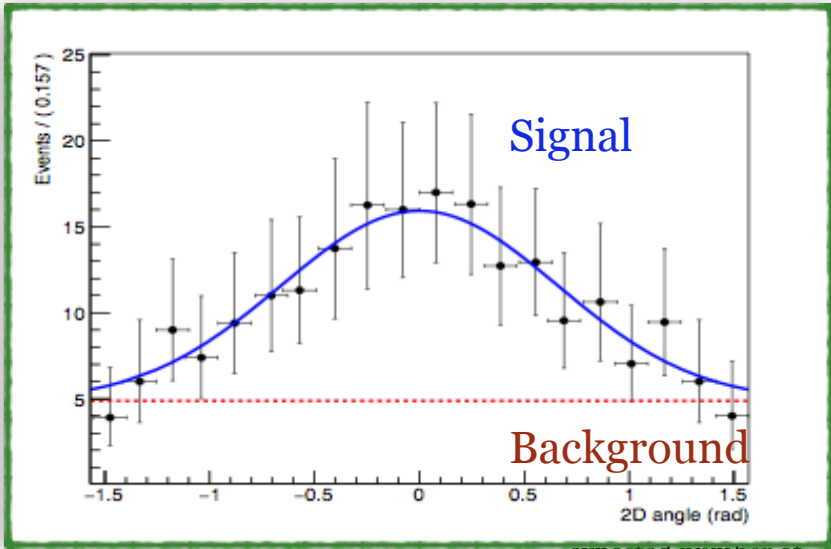
➡ Operation will be started in this October

Conclusion

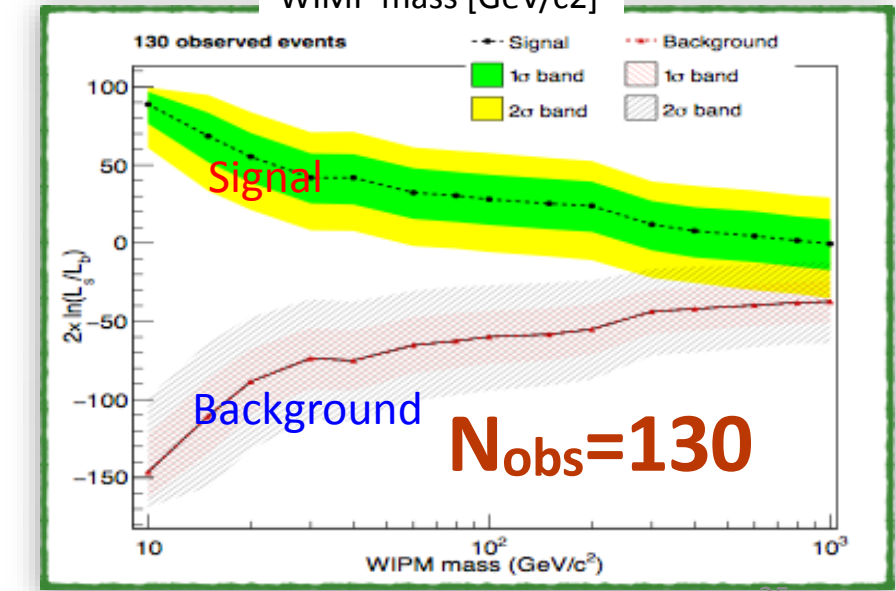
- **Directional sensitive search is new methodology to obtain new information for direct dark matter search**
- **Super-fine grained nuclear emulsion (Nano Imaging Tracker : NIT) have been firstly demonstrated to be able to detect the nuclear recoil as track.**
- **NEWSdm project is now on going as international experiment toward directional dark matter search in the LNGS**
 - 2018 : underground facility construction at LNGS**
 - 2019 : TDR will be prepared and start the small scale experiment**
 - 2020 - : preparation and run for larger scale experiment of ~ kg scale**
- **Background study and new readout technologies are now on progress**

Potential of Directional Sensitive Search

arXiv:1705.00613 [astro-ph.CO]



WIMP mass [GeV/c²]



WIMP mass [GeV/c²]

expected number of WIMP events

expected number of background events

signal pdf

background pdf

$$\mathcal{L}(\sigma_{\chi-n}, R_b) = \frac{e^{-(\mu_{\chi} + \mu_b)}}{N!} \times \prod_{i=1}^N [\mu_{\chi} f_{\chi}(\vec{q}_i; t_i) + \mu_b f_b(\vec{q}_i)]$$

total number of observed events

set of observables

Direction information : Several 10 events

Gain of 100 times

Annual modulation : Several 1000 events

