# Neutrino Experiment for Oscillation at Short Baseline

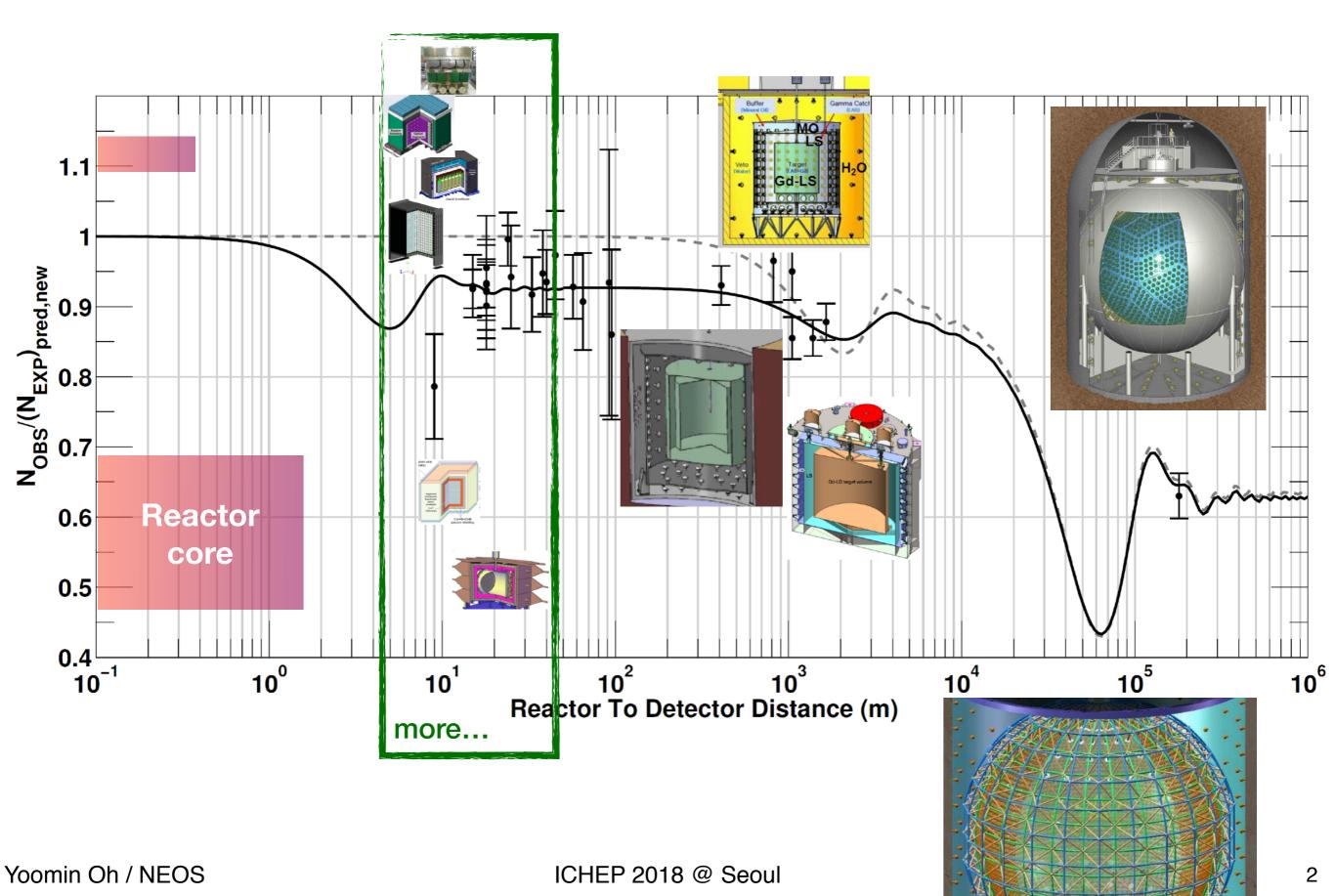
July 4-9 @ ICHEP 2018, Seoul

Yoomin Oh for the NEOS collaboraiton

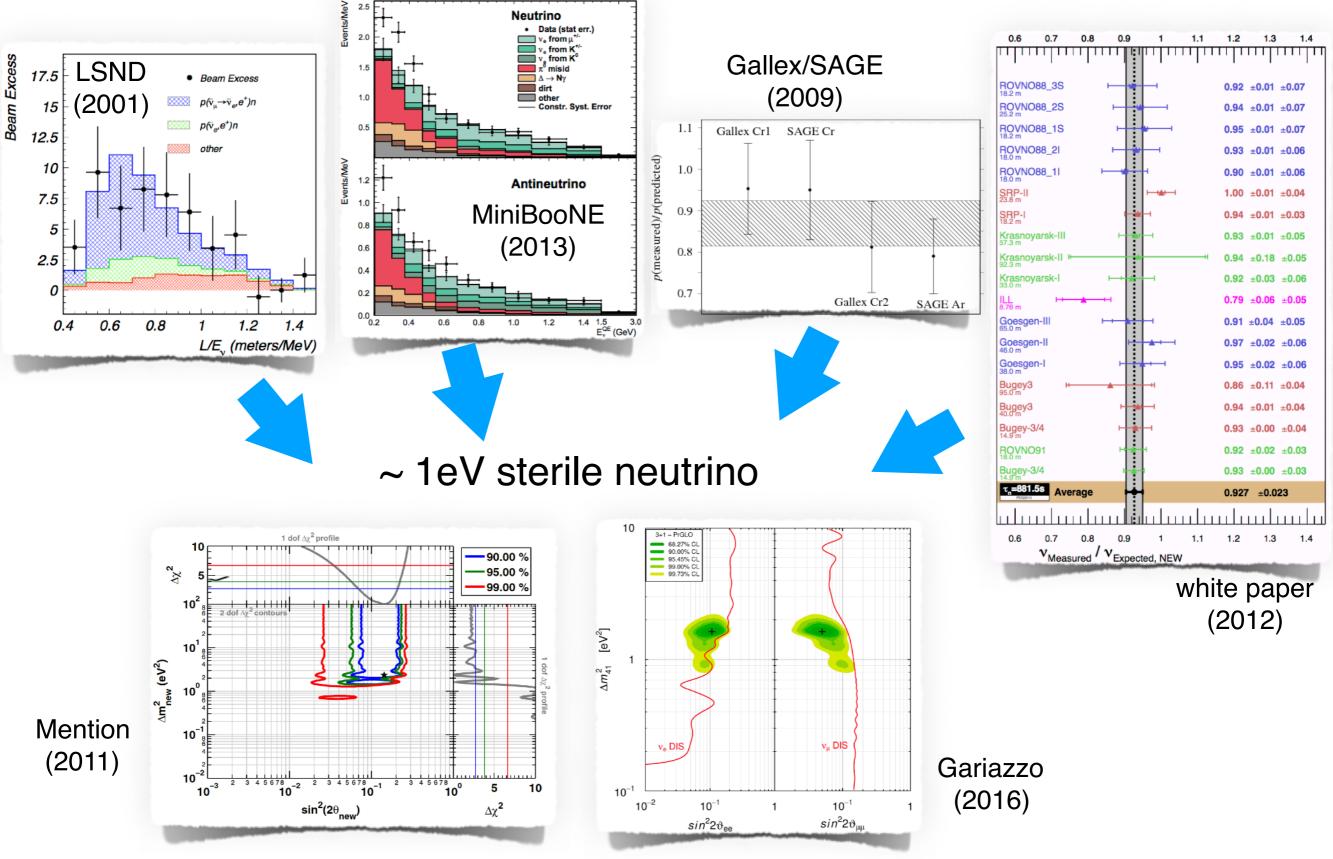




#### **Reactor neutrino experiments**





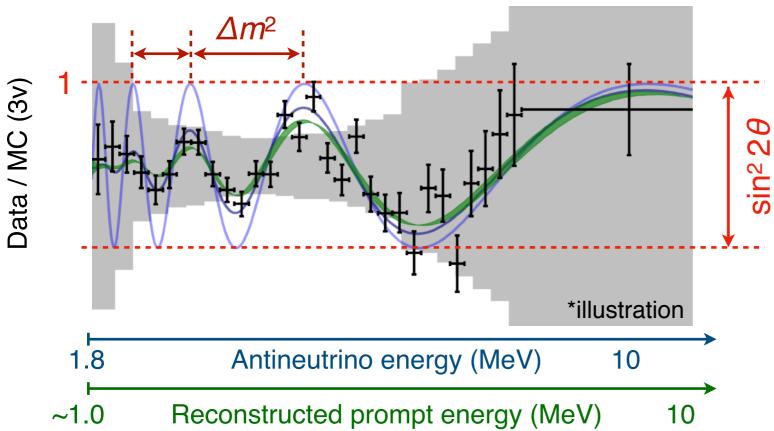


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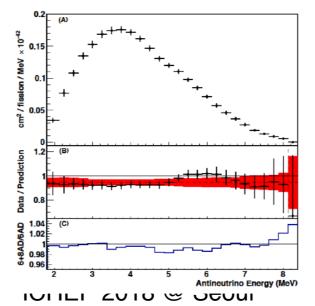
# NEOS Sensitivity to 3+1v oscillation

$$P(\bar{v}_{e} \to \bar{v}_{e}) \sim 1 - \sin^{2}2\theta_{14} \sin^{2}\left(1.27 \frac{\Delta m^{2}L}{E_{v}} \left[\frac{\text{eV}^{2} \cdot \text{m}}{\text{MeV}}\right]\right)$$



Single detector experiment:

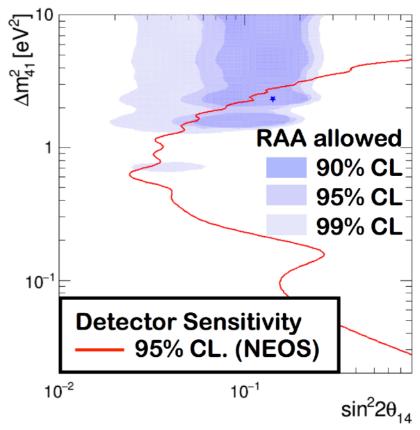
- a decent model spectrum,
- detector simulation



washed out by:
Baseline size of reactor, detector
Detector

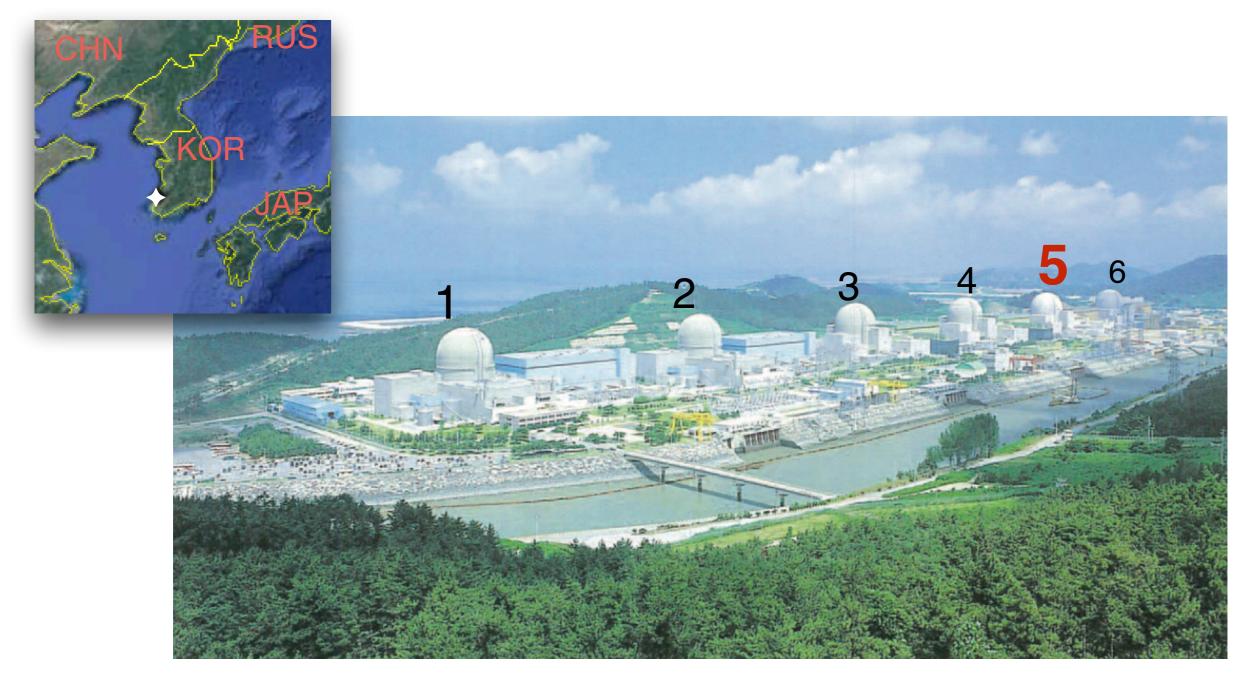
neutrino  $\rightarrow$  prompt energy

- ⊗ Uncertainties in model bin-to-bin (un)correlated
- ⊗ Statistics, background, ...



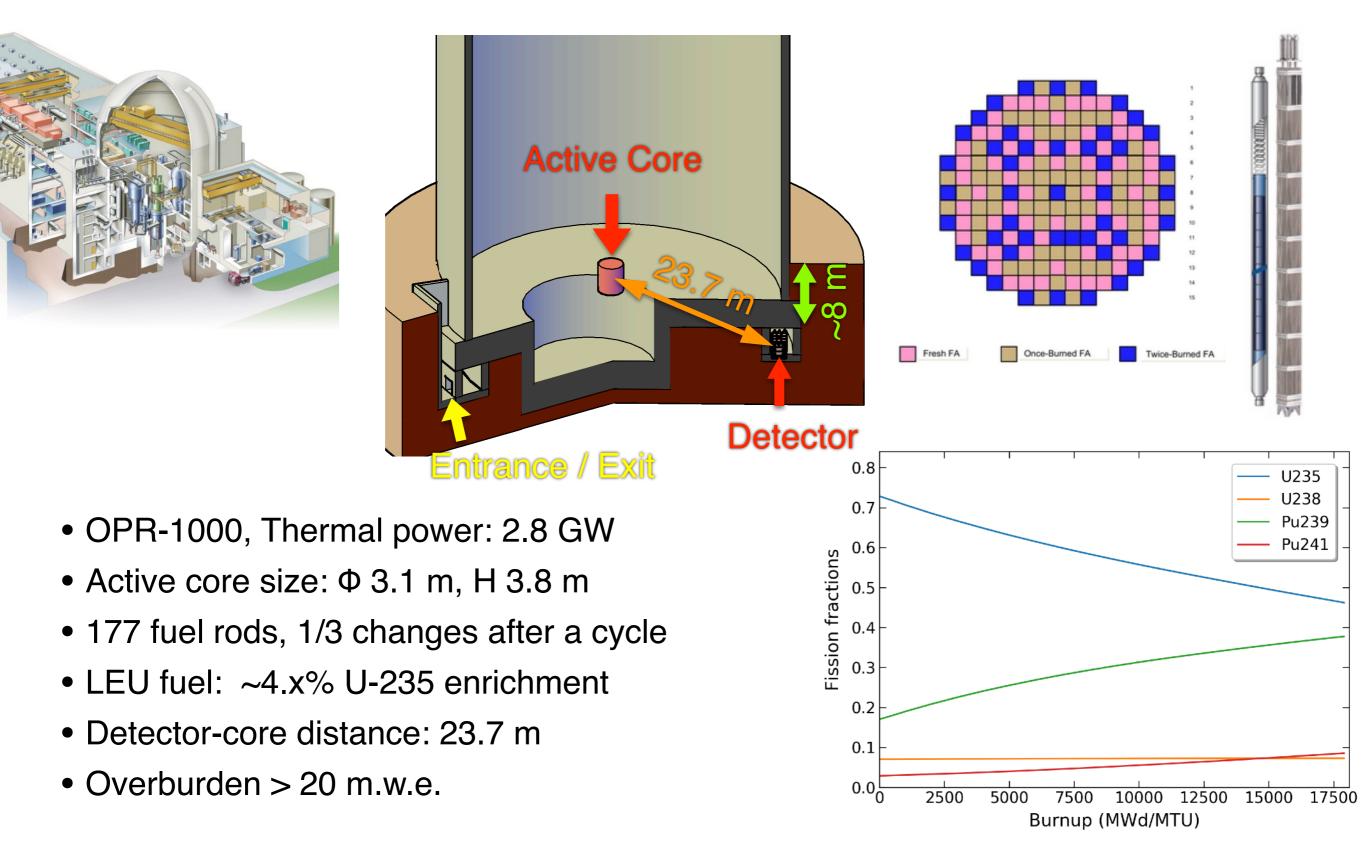
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# Experimental site

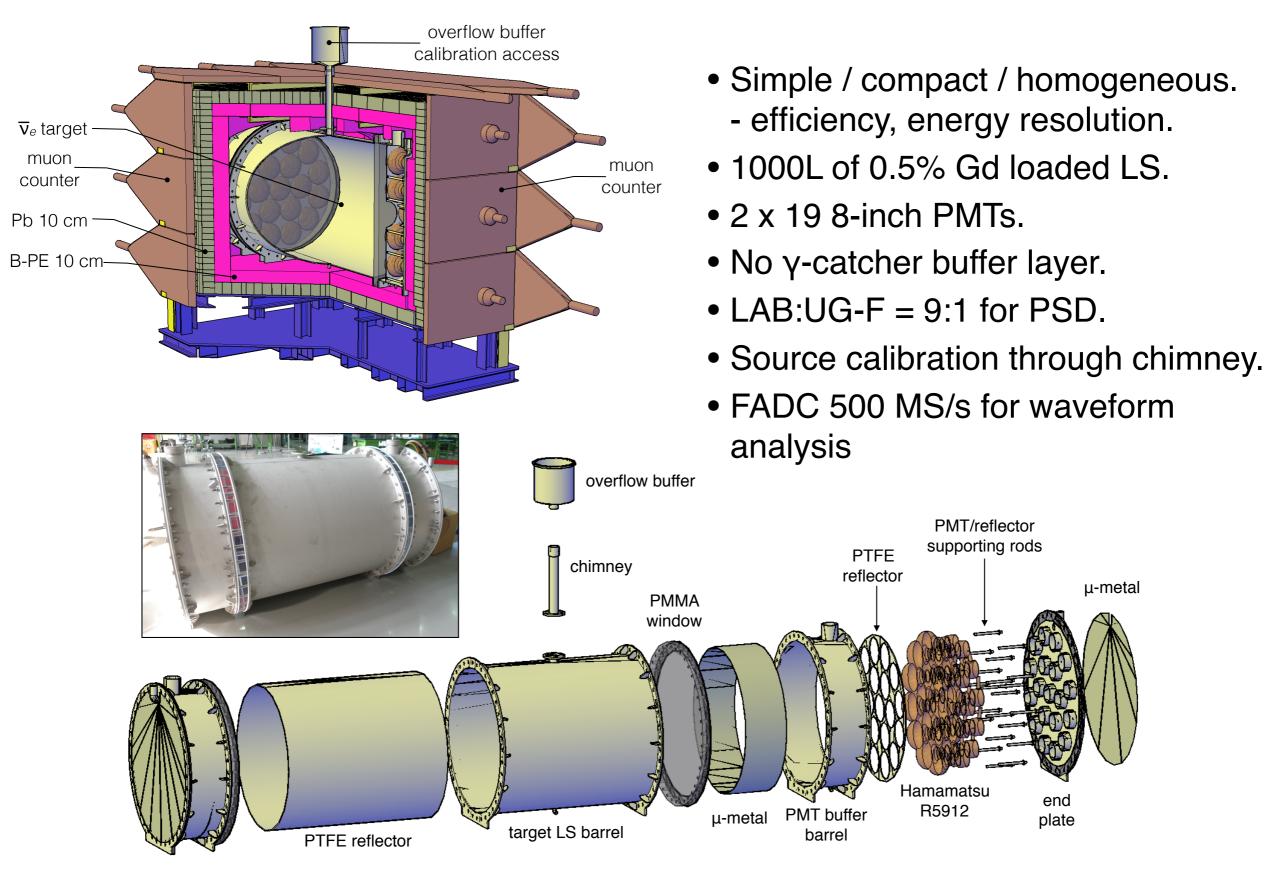


- Hanbit-5 reactor in Yeonggwang (靈光, ghost illumination), Korea
- Distance between neighboring reactor cores: 256 m
- Same reactor complex being used for the RENO experiment

### Hanbit-5 reactor and tendon gallery



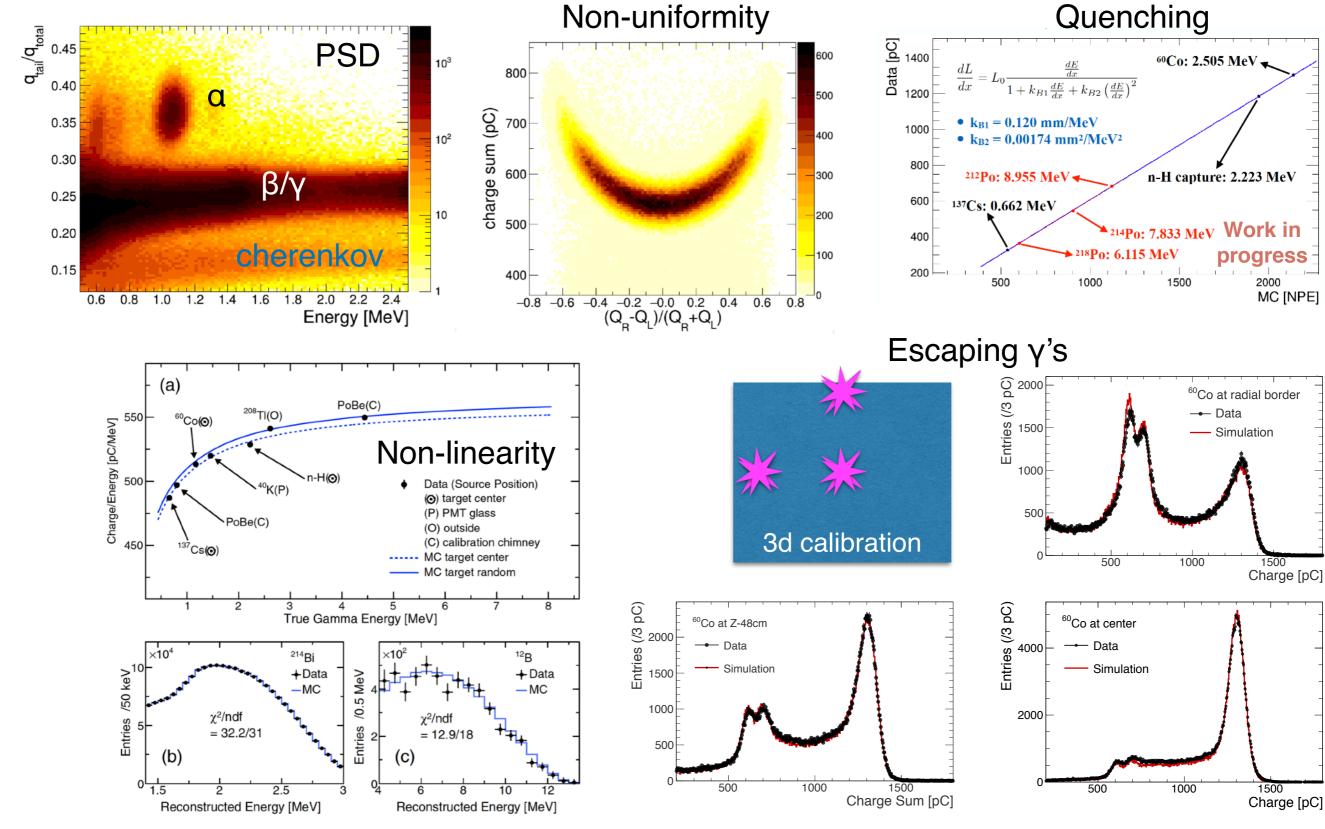
## Detector design & construction



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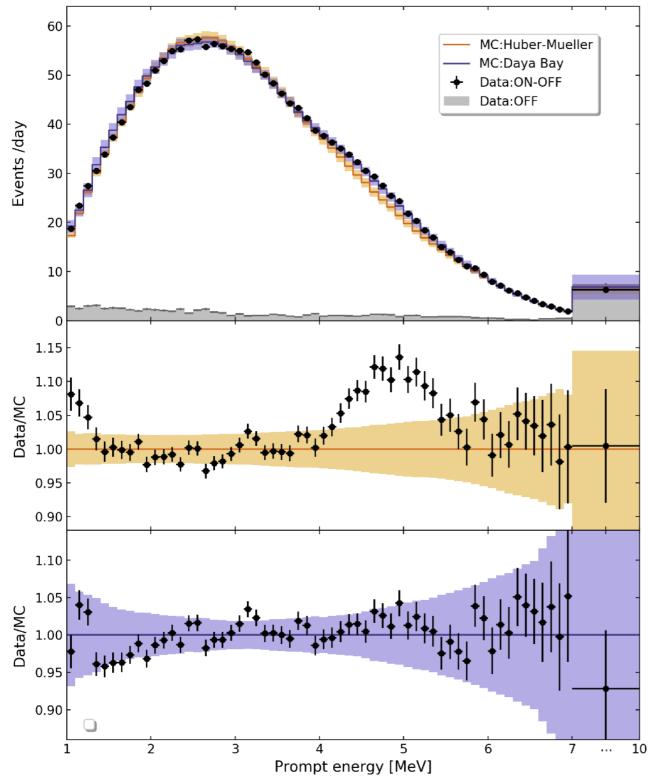
#### **Detector responses & simulation**



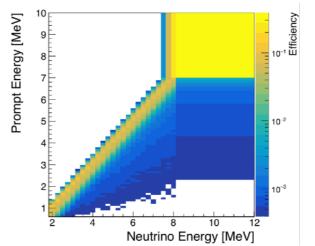
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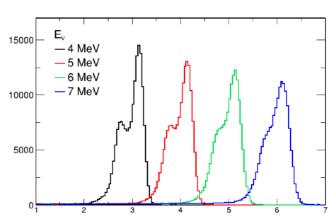
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## Prompt energy spectrum



- 1976 IBD candidates/day (on), 85 /day (off); S/N~23.
- Negligible background change between on-off periods.
- Spectral anomaly observed (vs Huber-Mueller normalization).
- Small structural fluctuations.
- Reference spectra generated using detector response matrix,
  - energy escape due to detector size

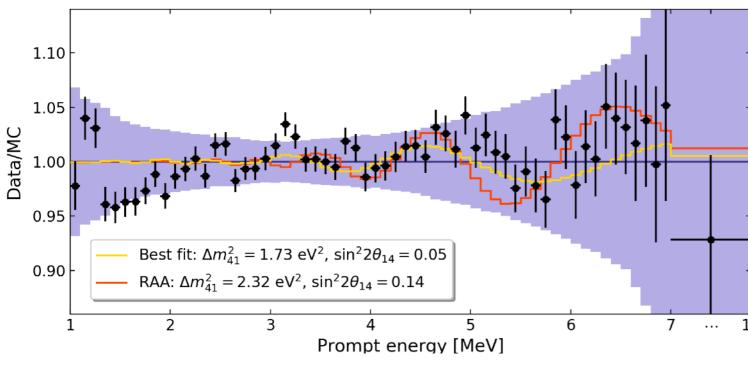


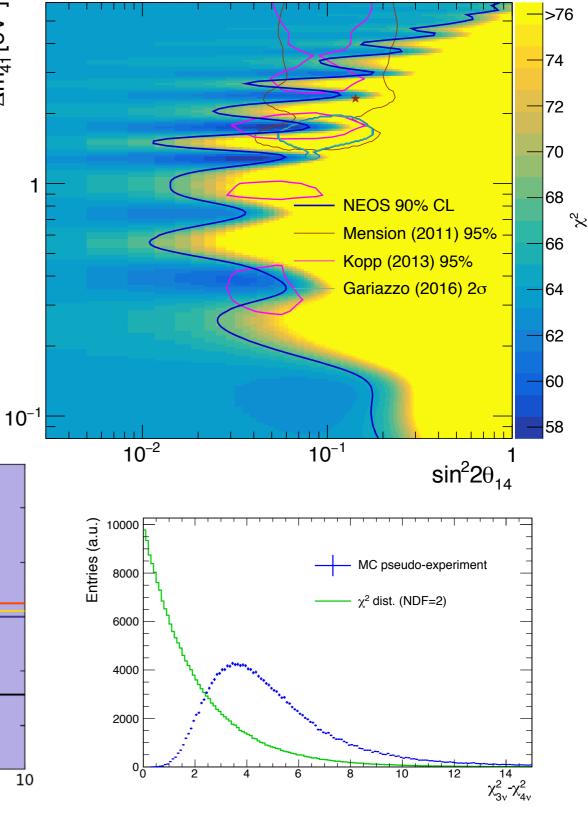


### Active-to-sterile oscillation

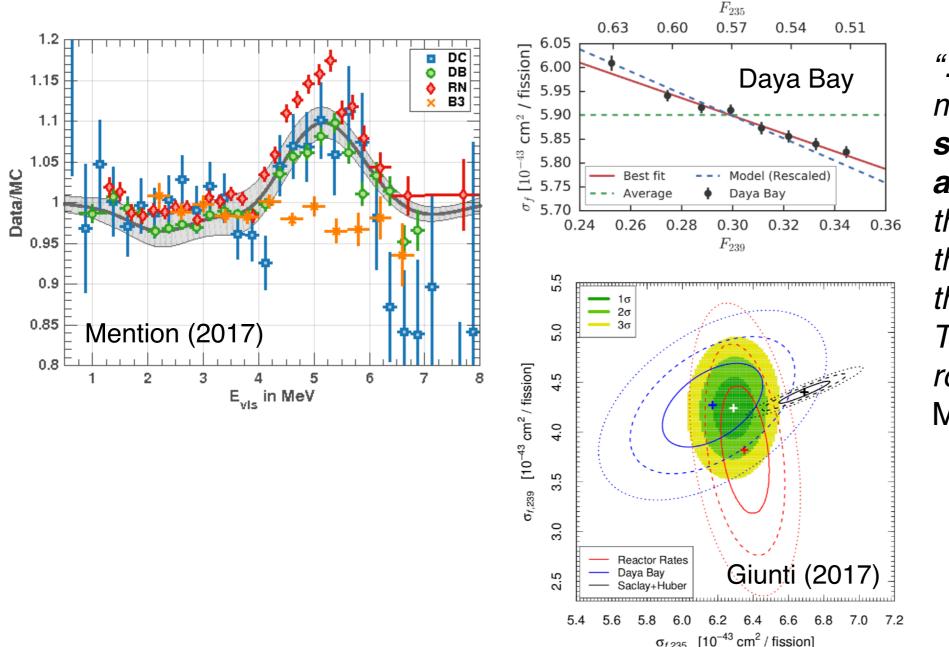
 $\Delta m^2_{41}$  [eV<sup>2</sup>]

- Normalized with the Daya Bay shape
- Best fits at: (1.73 eV<sup>2</sup>, 0.05), (1.30 eV<sup>2</sup>, 0.04) with  $\chi^2(3v) \cdot \chi^2(4v) = 6.5$ , p-value = 0.22
- Not a definitive answer for 3+1v oscillation or fine structures in reactor v spectrum.





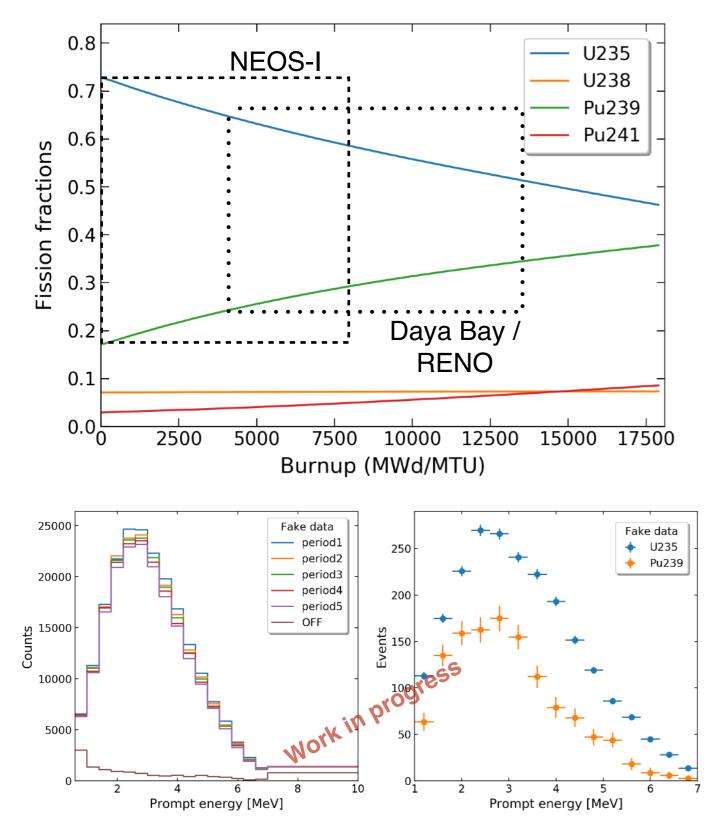
# **Recent Issues**



"... sterile neutrino models fail to simultaneously account for all the  $v_e \rightarrow v_e$  data, the  $v_{\mu} \rightarrow v_e$  data and the  $v_{\mu} \rightarrow v_{\mu}$  data. This conclusion is robust;" Maltoni @ v2018

+ fine structures in the flux spectrum > requires high resolution detector?

# NEOS phase-II measurement



- Starts in September 2018:
  - one full burnup cycle (~500 days), plus two maintenance (off) periods (>100 days, contributions from spent fuels?)
  - expecting ~1 M IBD events.
  - same reactor, same detector.
- Spectrum evolution with fuel component changes.
  - from a single LEU core,
  - absolute spectrum measurement,
  - decomposing spectra for different elements.

## Detector refurbishment and preparation



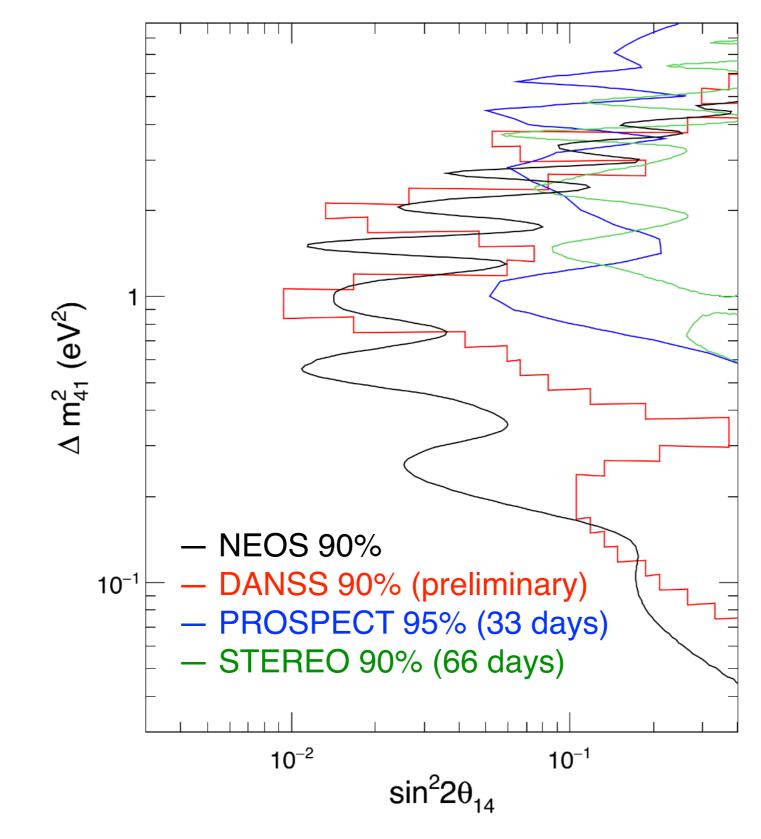
- To make more stable detector for longer data taking.
- Producing fresh Gd-LS and its QC.
- Upgrading slow control/monitoring system temperature / PMT gain.
- MC simulation upgrade, e.g. GEANT4 versions, n-Gd data.

# Summary

- IBD prompt energy spectral shape has been successfully measured, using 2.8 GWt commercial reactor at 24 m distance.
- No strong sign of active-to-sterile neutrino oscillation for  $\Delta m^2 \sim 1 \text{ eV}^2$ ,  $\sin_2 2\theta \sim 0.1$ .
- Similar spectral anomaly observed as mid-baseline experiments.
- Measurement will be resumed soon to see the evolution of the reactor neutrino flux/spectrum according to the fuel component changes.

#### Thank you.

# **B1.** Recent limits



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