

# ICHEP2018 SEOUL

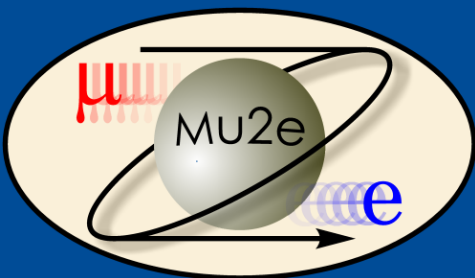
XXXIX INTERNATIONAL CONFERENCE

ON *high Energy* PHYSICS

JULY 4 - 11, 2018

COEX, SEOUL

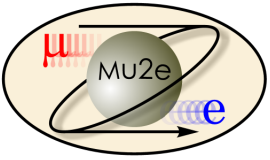
## The Mu2e Tracker



Gianantonio Pezzullo  
Yale University

*on behalf of the Mu2e tracker group*

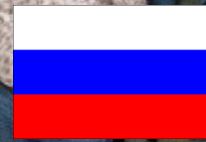




# The Mu2e collaboration



~200 scientists from 35 institutions



Argonne National Laboratory, Boston University, Brookhaven National Laboratory

University of California, Berkeley, University of California, Irvine, California Institute of Technology, City University of New York,

Joint Institute for Nuclear Research, Dubna, Duke University, Fermi National Accelerator Laboratory, Laboratori Nazionali di

Frascati, Helmholtz-Zentrum Dresden-Rossendorf, University of Houston, University of Illinois, INFN Genova, Kansas State

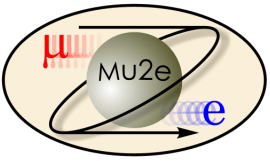
University, Lawrence Berkeley National Laboratory, INFN Lecce and Università del Salento, Lewis University, University of Louisville,

Laboratori Nazionali di Frascati and Università Marconi Roma, University of Minnesota, Muons Inc., Northern Illinois University,

Northwestern University, Novosibirsk State University/Budker Institute of Nuclear Physics, Institute for Nuclear Research, Moscow,

INFN Pisa, Purdue University, Rice University, University of South Alabama, Sun Yat Sen University, University of Virginia, University

of Washington, Yale University

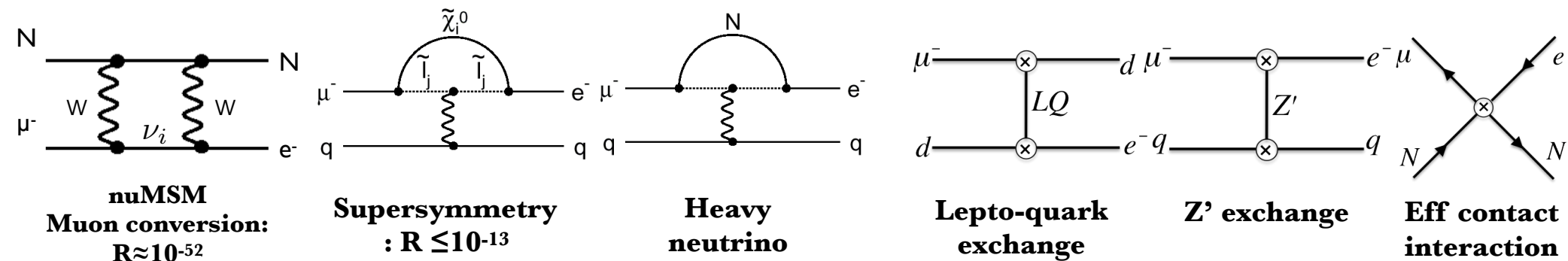


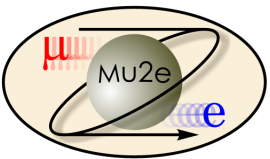
# What is $\mu \rightarrow e$ conversion

- $\mu$  converts to an electron in the presence of a nucleus  $\mu^- N \rightarrow e^- N$

$$E_e = m_\mu c^2 - B_\mu(Z) - C(A) = 104.973 \text{ MeV}$$

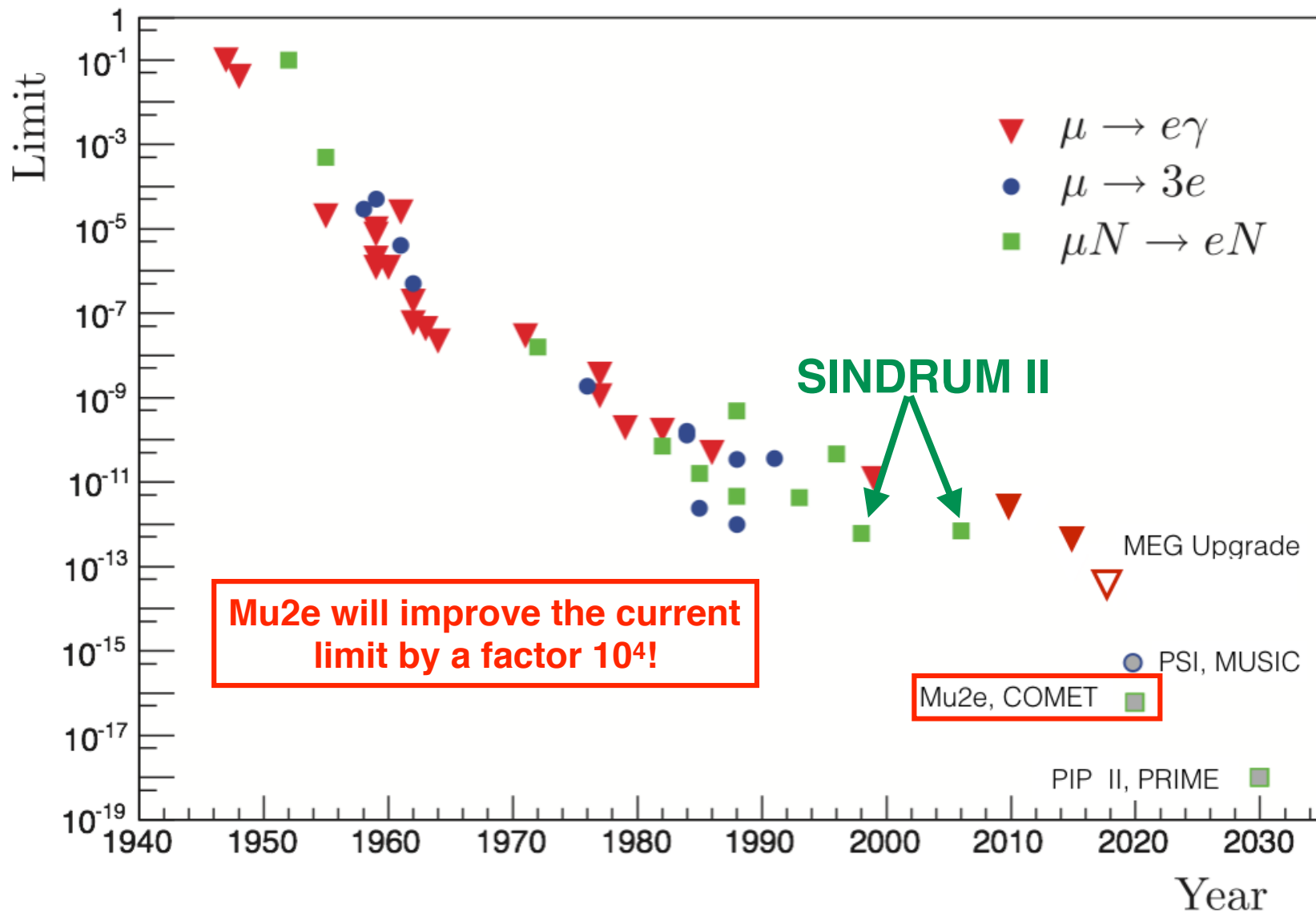
- for Aluminum:  $\begin{cases} B_\mu(Z) \text{ is the muon binding energy (0.48 MeV)} \\ C(A) \text{ is the nuclear recoil energy (0.21 MeV)} \end{cases}$
- $\mu$  conversion in the SM is induced by **neutrino masses and mixing** at a negligible level  $\sim 10^{-52}$
- Many **SM extensions enhance the rate** through mixing in the high energy sector of the theory (other particles in the loop...)





# History of $\mu \rightarrow e$ search

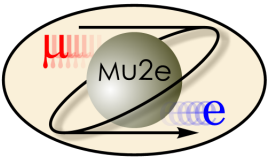
History of  $\mu \rightarrow e\gamma$ ,  $\mu N \rightarrow eN$ , and  $\mu \rightarrow 3e$



R. Bernstein, P. Cooper <https://doi.org/10.1016/j.physrep.2013.07.002>



# Experimental setup

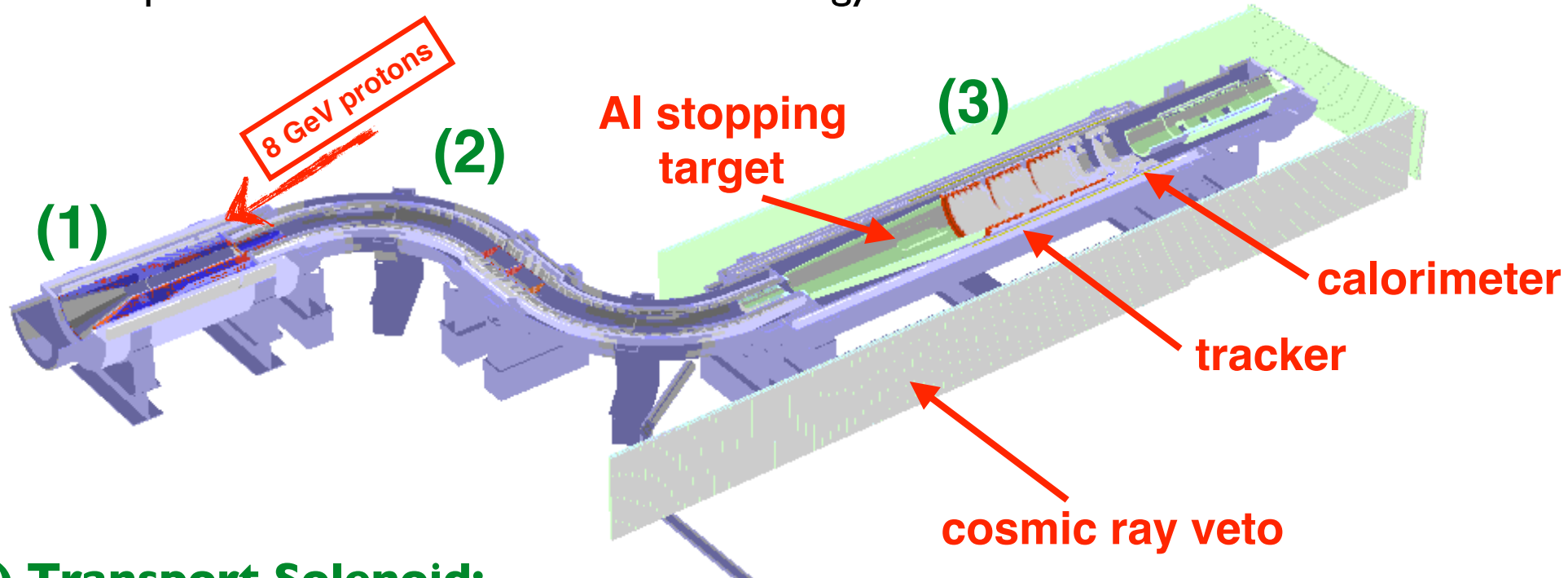


## (1) Production Solenoid:

- ➔ Proton beam strikes target, producing mostly pions
- ➔ Graded magnetic field contains backwards pions/muons and reflects slow forward pions/muons

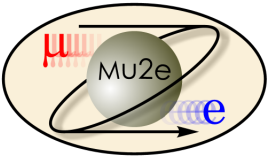
## (3) Detector Solenoid:

- ➔ Capture muons on Al target
- ➔ Graded field “focuses”  $e^-$  in tracker fiducial
- ➔ Measure momentum in tracker and energy in calorimeter



## (2) Transport Solenoid:

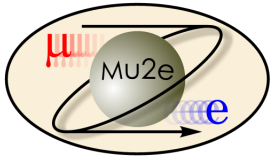
- ➔ Select low momentum, negative muons



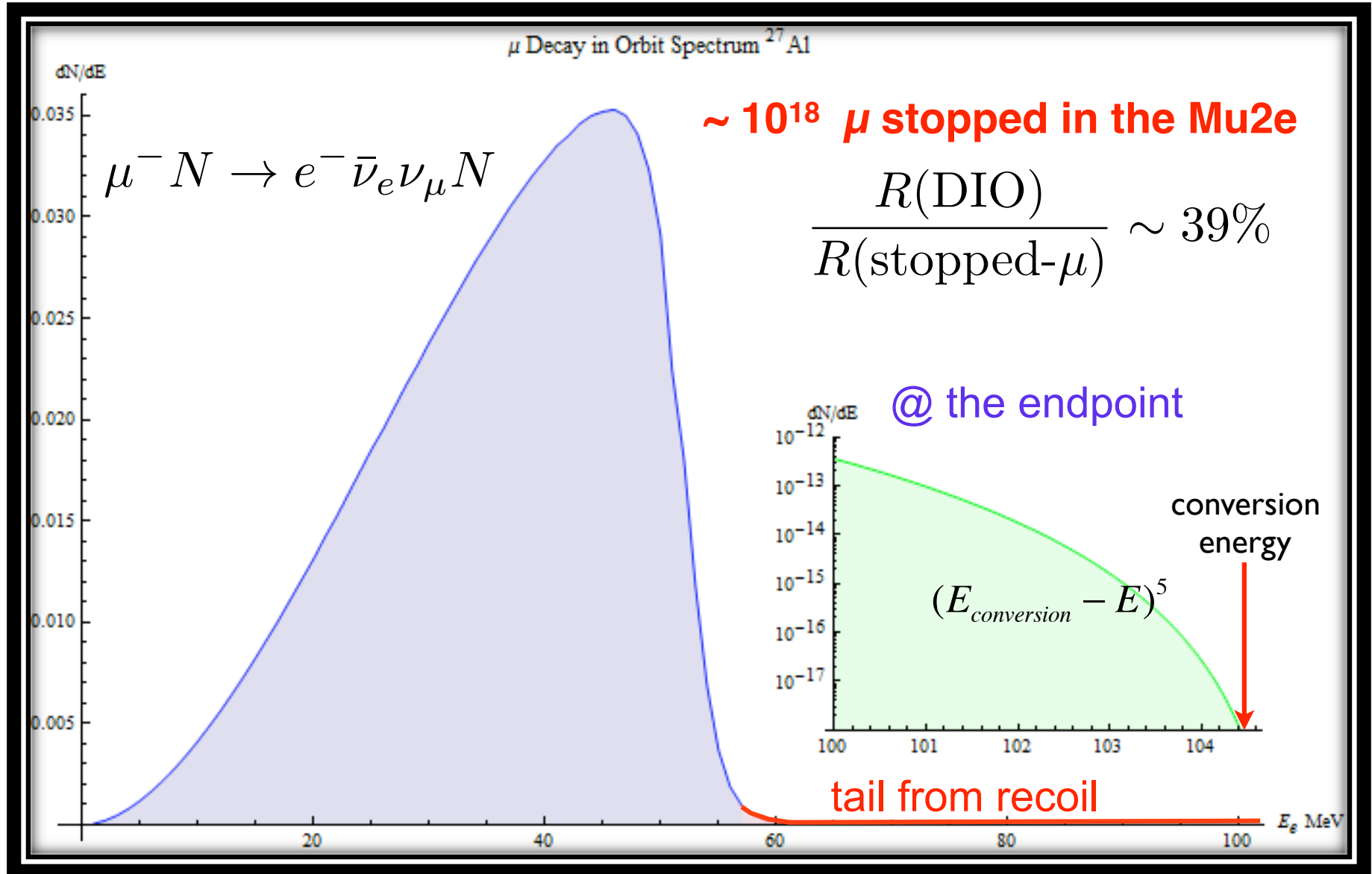
# Backgrounds & design



	Source	Scale with	Solution
Intrinsic	decay-in-orbit	# of stopped- $\mu$	Tracker resolution
Beam	radiative $\pi$ capture	closeness to beam pulse	pulsed beam
Running time	Cosmic ray	live time	veto system & PID

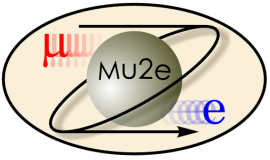


# μ decay-in-orbit (DIO)



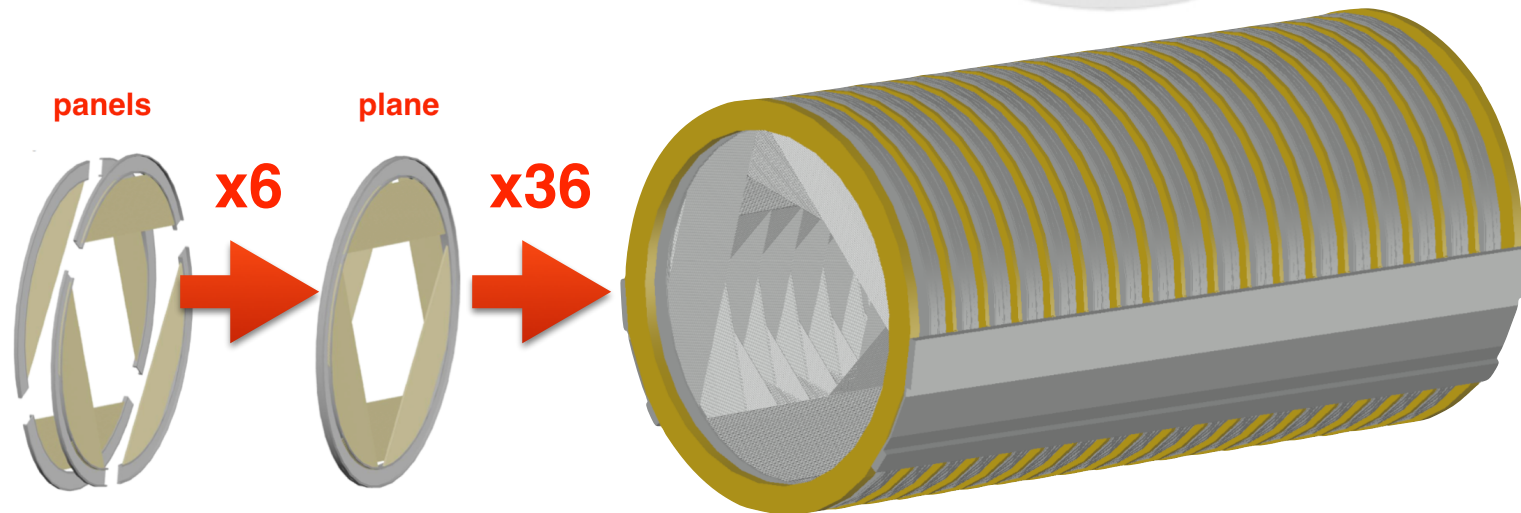
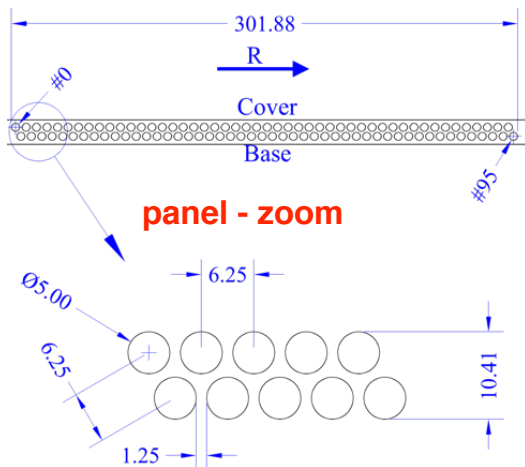
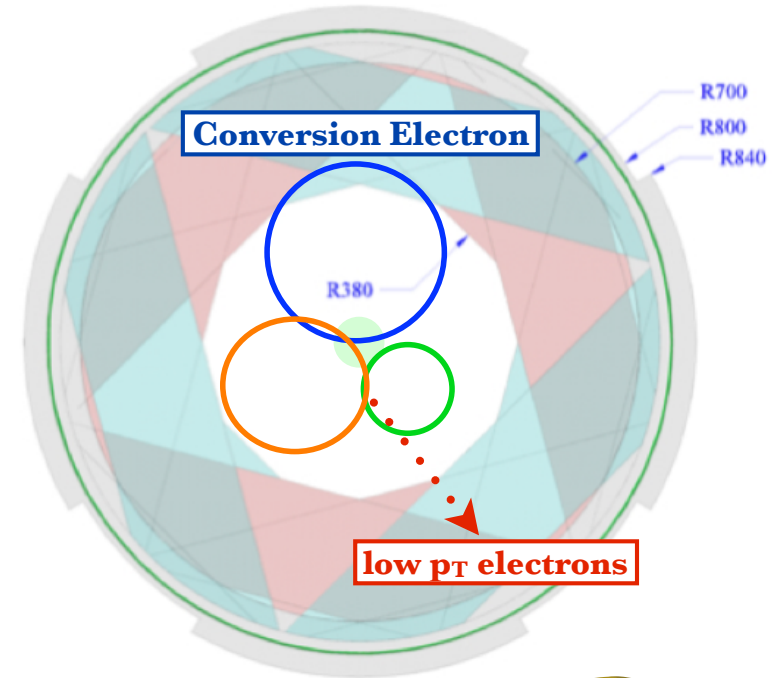
R. Szafron, A. Czarnecki <https://doi.org/10.1016/j.physletb.2015.12.008>





# Tracker design

- 36 double-layer planes equally spaced with straws transverse to the beam
- Inner 38 cm un-instrumented:
  - ✓ blind to beam flash
  - ✓ blind to >99% of the DIO spectrum
- Expected resolution:
  - ✓ ~ 200 keV/c @ 105 MeV



# Straw technology

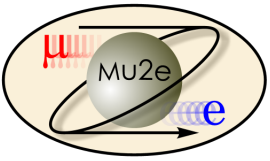
- ~ 20k straws employed in the tracker
- Multiple scattering is the major contributor to  $dp$ 
  - ✓ straw material budget is comparable to the gas
- Straw specs:
  - ✓ 5 mm diameter, 2x6.25  $\mu\text{m}$  Mylar walls Au and Al coated
  - ✓ 25  $\mu\text{m}$  Au-plated W sense wire
  - ✓ 80/20 Ar/CO<sub>2</sub> with HV ~ 1500 V
- Straw length varies from 44 to 114 cm

**straw tube**



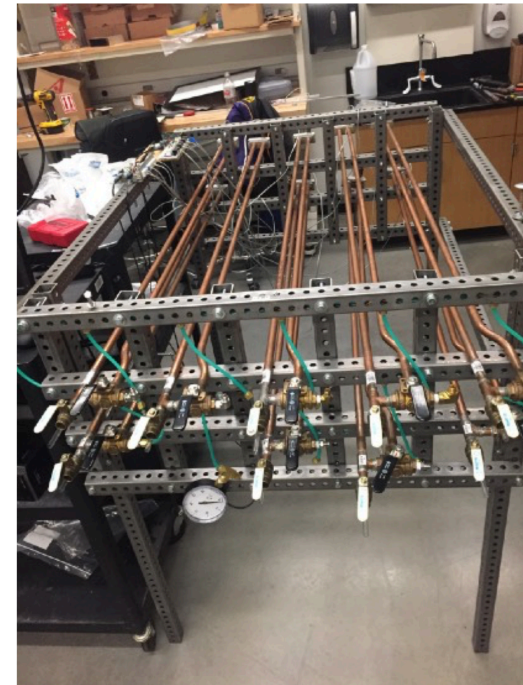
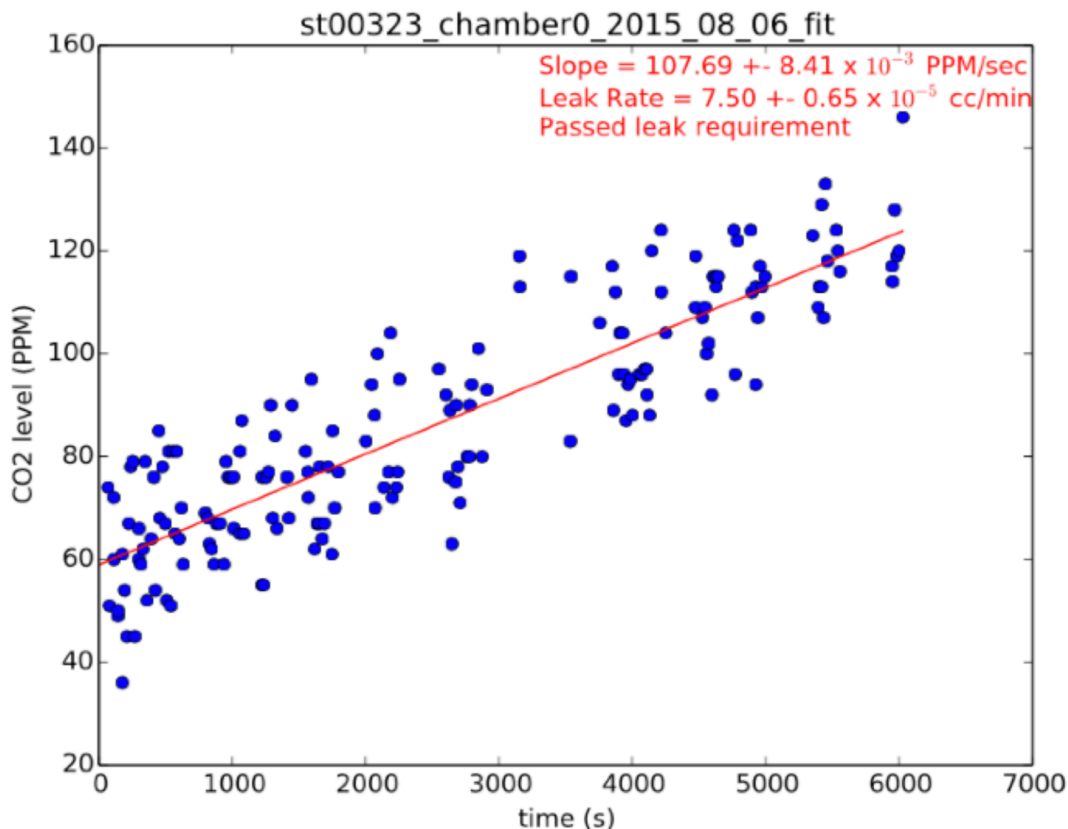
**Mylar roll**





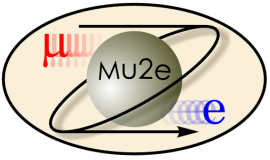
# Vacuum requirements

- There are ~ 20k straws that collectively must leak less than 6 cm<sup>3</sup>/min
  - ✓ total surface > 3e6 cm<sup>2</sup> with ~ 1 atm differential pressure Ar:CO<sub>2</sub>
  - ✓ straw are delicate and we need to test all of them!
- CO<sub>2</sub> permeates Mylar at least x10 faster than Argon
- We measure the CO<sub>2</sub> leak rate for each straw with custom chambers



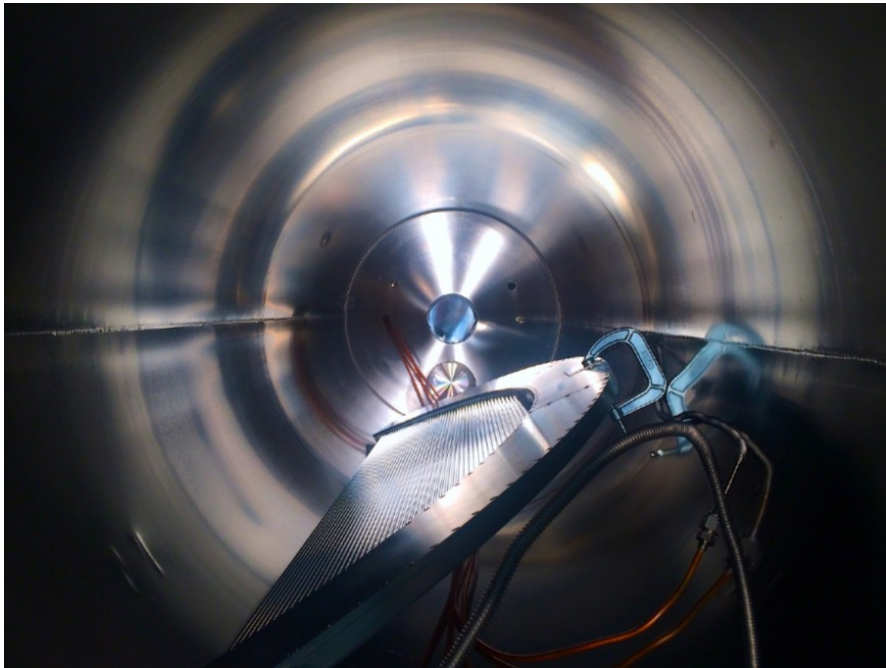
**chamber @ UMN**



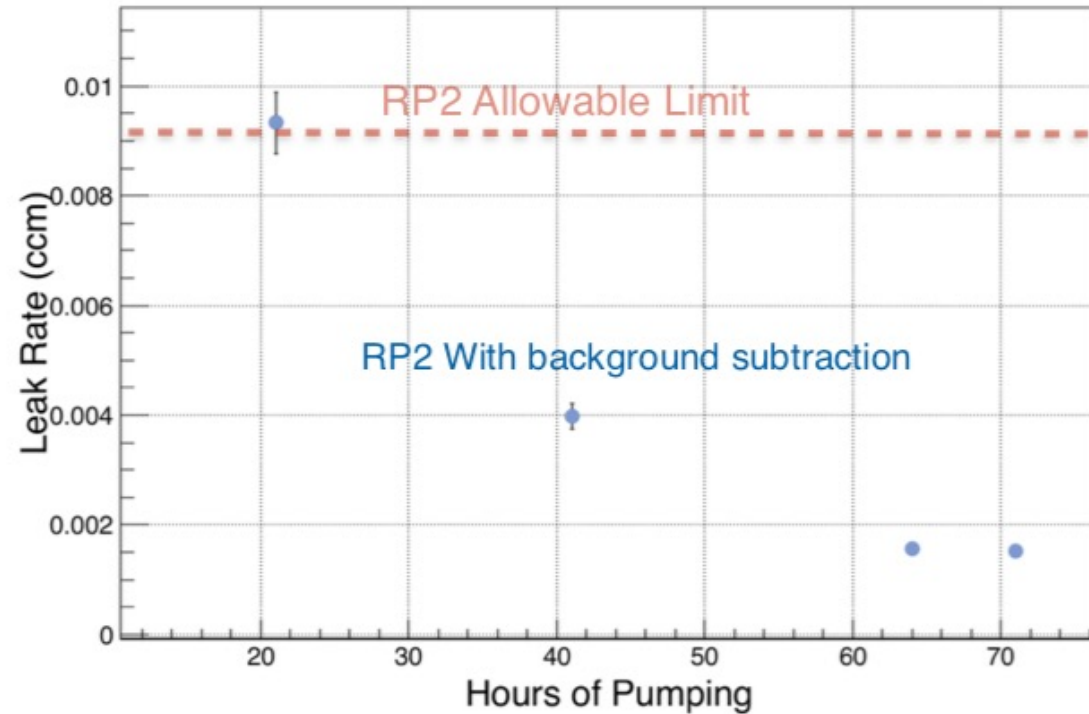


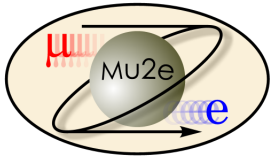
# Vacuum requirements

- Vacuum test on each panel to check  
✓ gas manifold leak + outgassing



**panel vacuum test @ Fermilab**

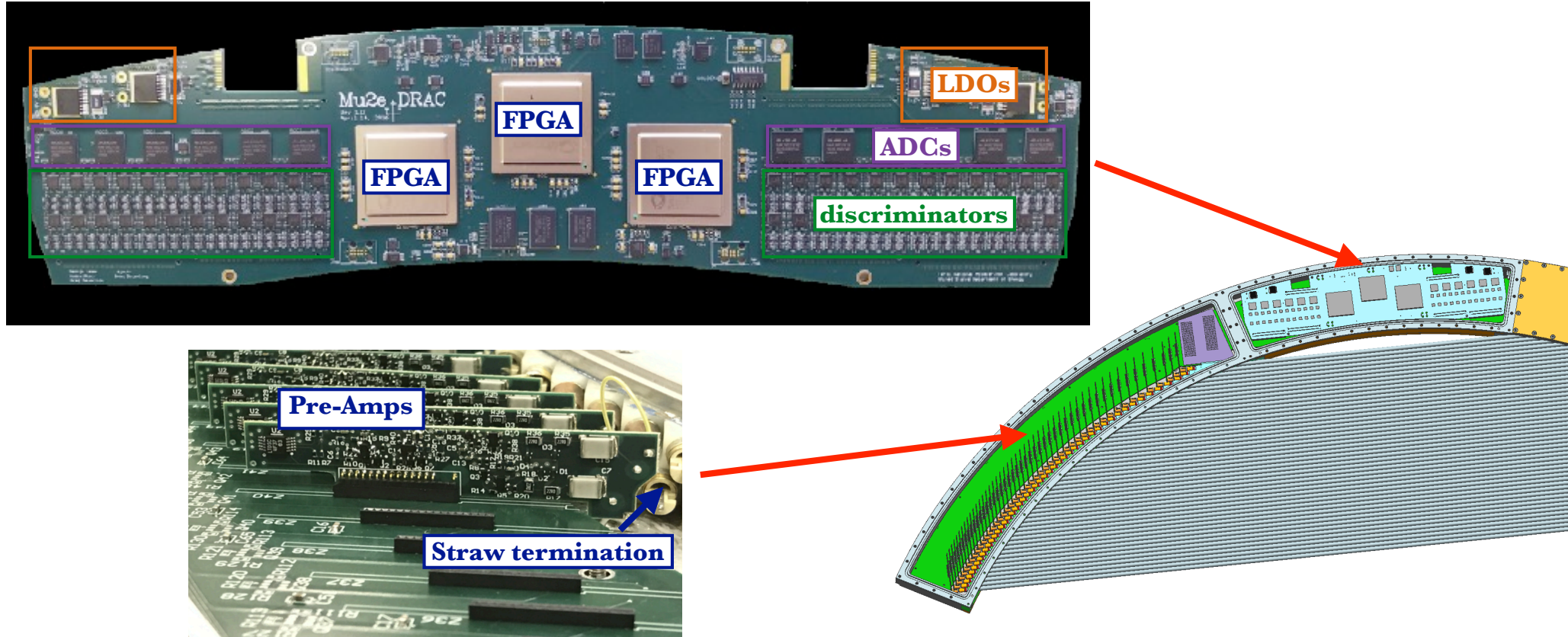


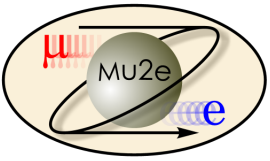


# Electronics



- Outer part of the panel houses the FEE and Digitizer Readout & Assembler Controller (DRAC)
- High hit rate sustainability: 15 kHz/cm<sup>2</sup>

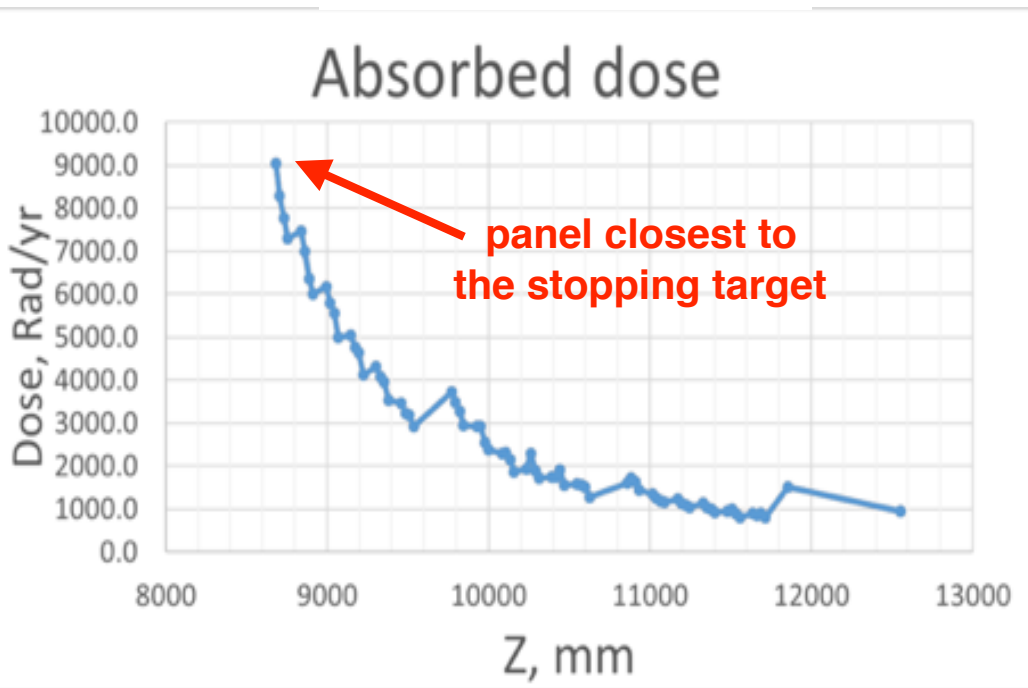




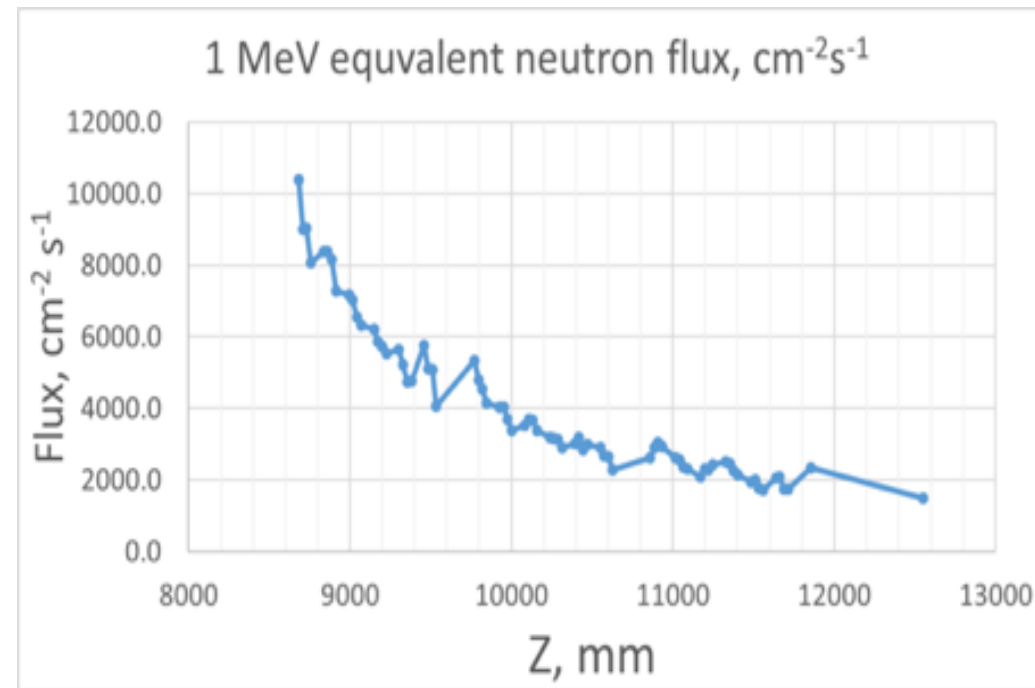
# Radiation Environment

- Detailed Monte Carlo simulations with MARS and GEANT4
- peak dose  $\sim 9$  krad/year in the first plane
- Neutron flux up to  $2 \times 10^{11}$  n<sub>1MeVeq</sub>/cm<sup>2</sup>/year

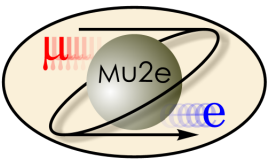
## Total Ionizing Dose (TID)



## Non Ionizing Energy Loss (NIEL)

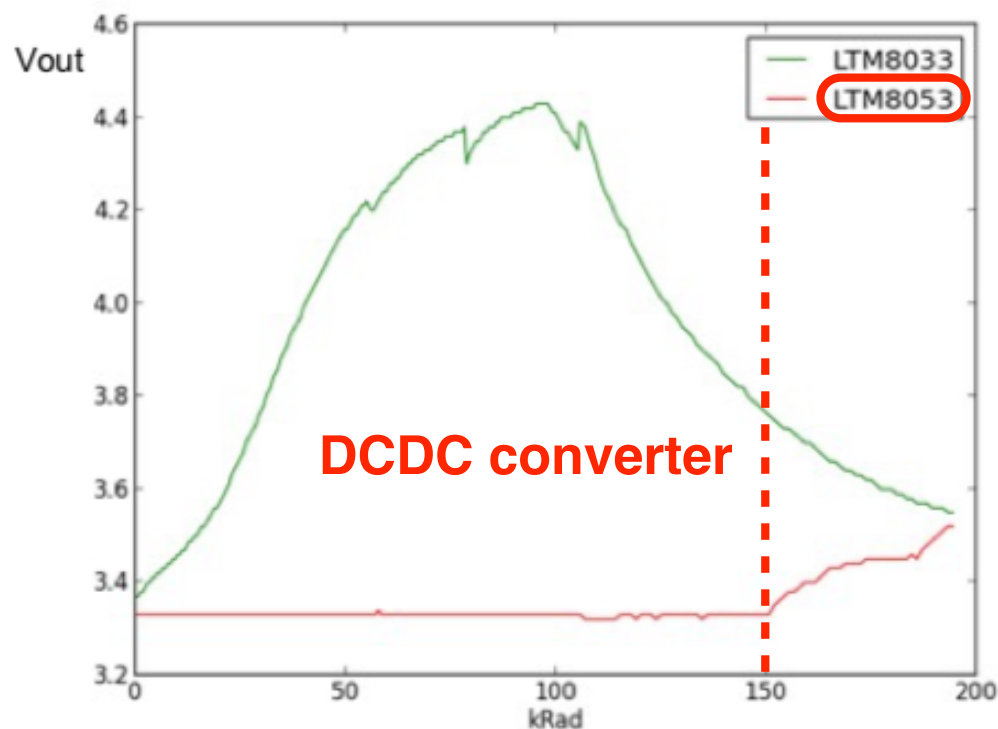
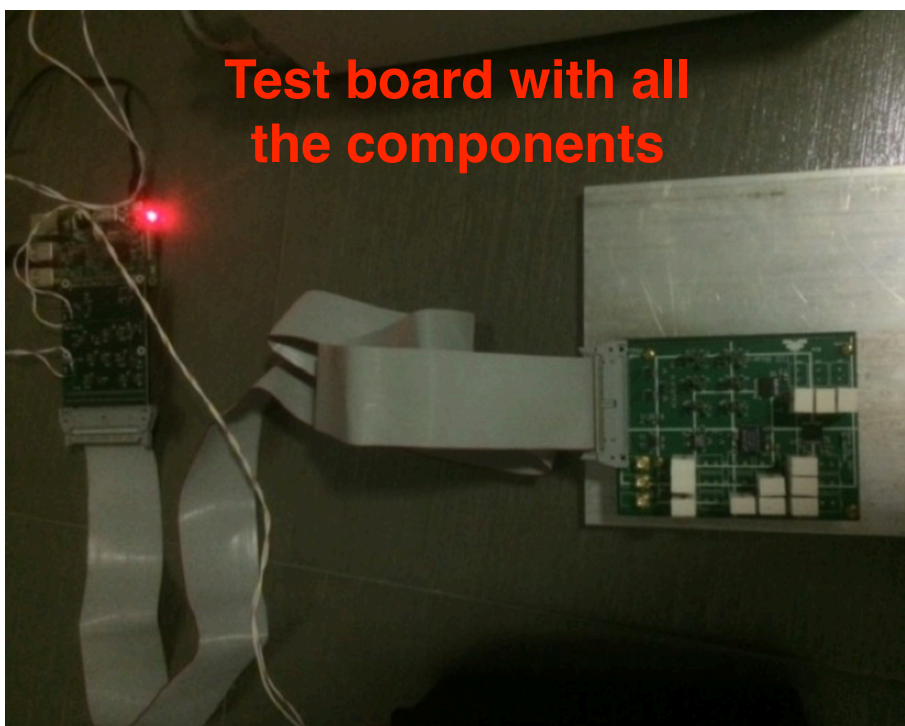


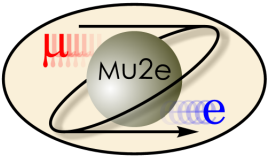




# Electronics vs radiation

- PolarFire FPGA from Microsemi (tested up to 500 krad)
- rad hard Optical transceiver VTRx from CERN (tested up to 1 Mrad)
- tests campaign to qualify other major components:
  - ✓ TID tests @ local radiotherapy clinic up to 200 krad
  - ✓ NIEL tests @ UC Davis McClellan reactor up to  $2 \times 10^{14}$  n<sub>1MeVeq</sub>/cm<sup>2</sup>

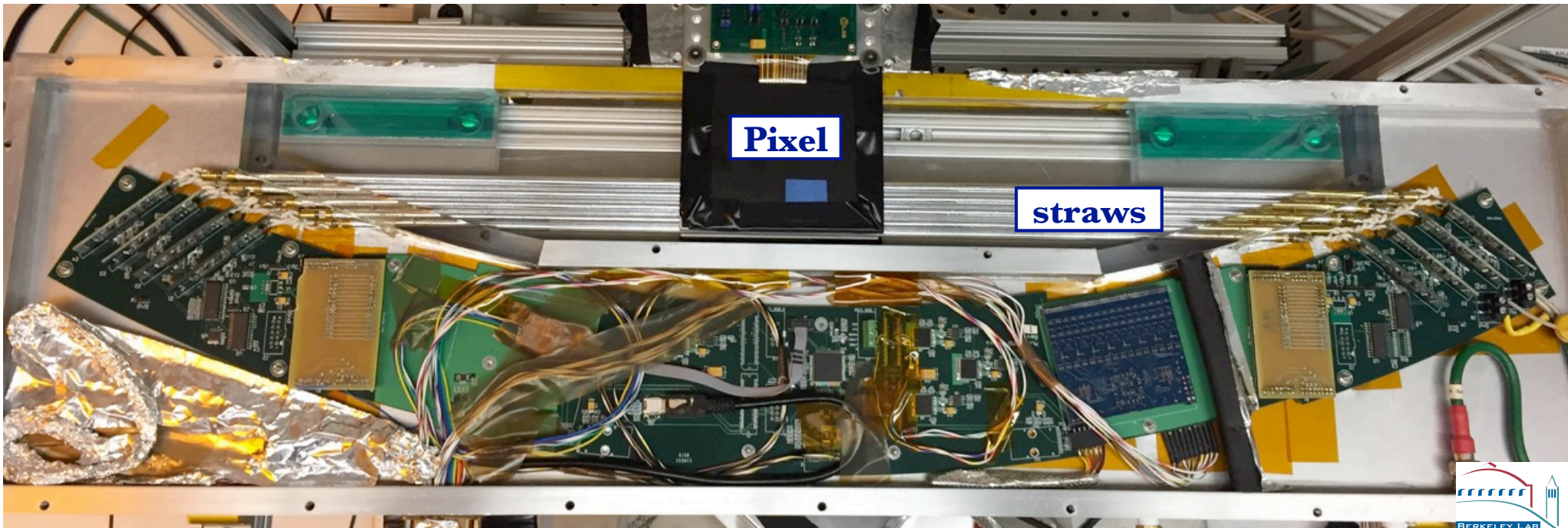




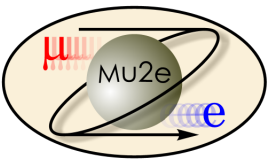
# Prototype



- Small panel prototype with 8 straws used to test performance @ LBL
- Test for cross talk using LBL 88" cyclotron
- Test with Cosmic rays and radioactive sources to measure performance

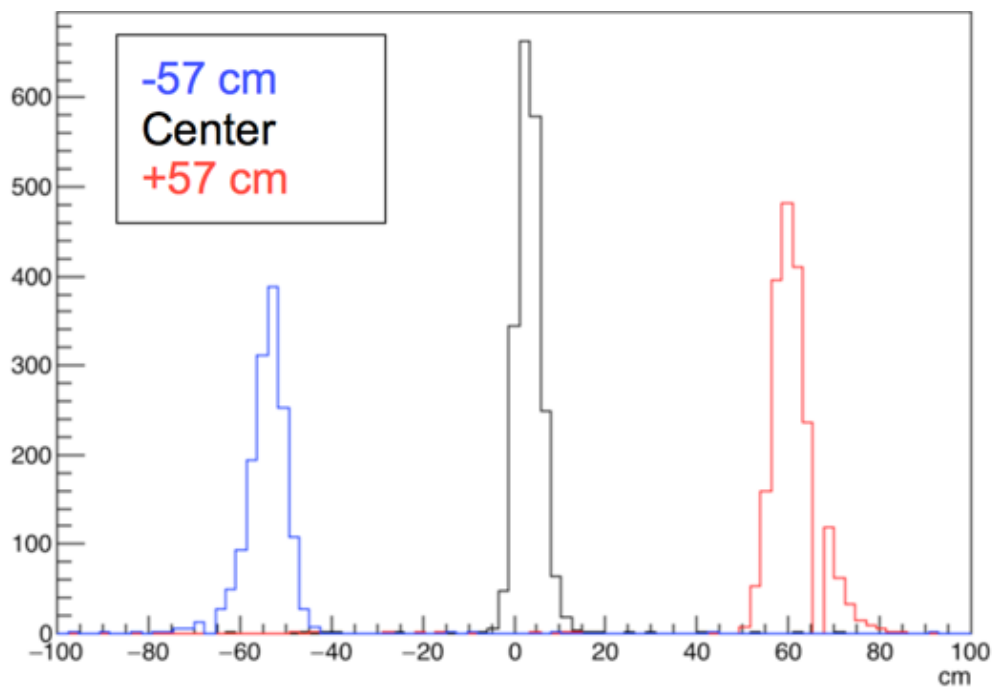


# Prototype - Results

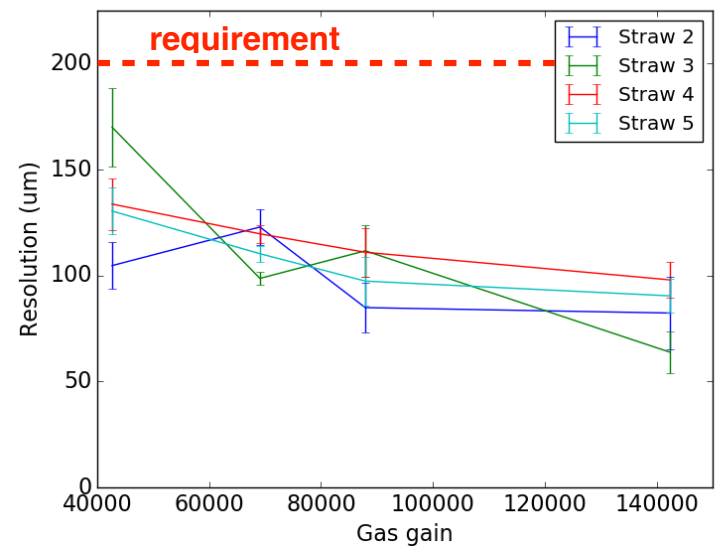


- Single hit efficiency and resolution meet requirements

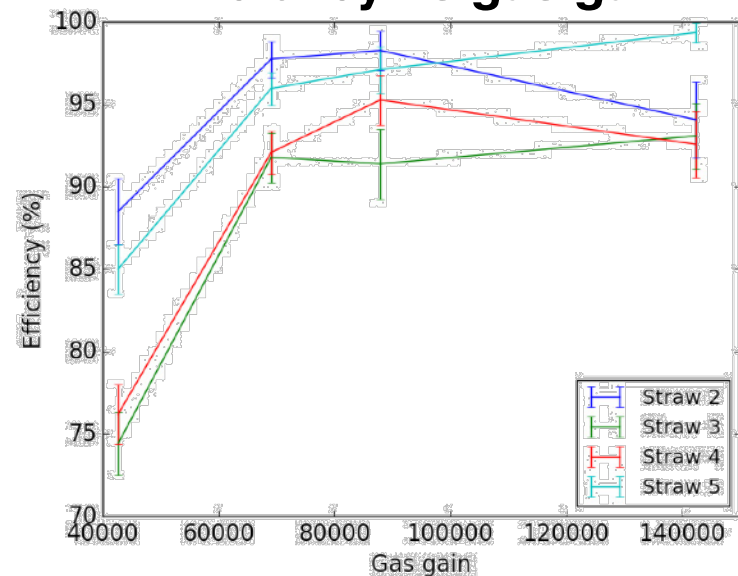
## Position measurement vs source location



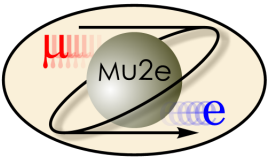
## Drift resolution vs gas gain



## Efficiency vs gas gain

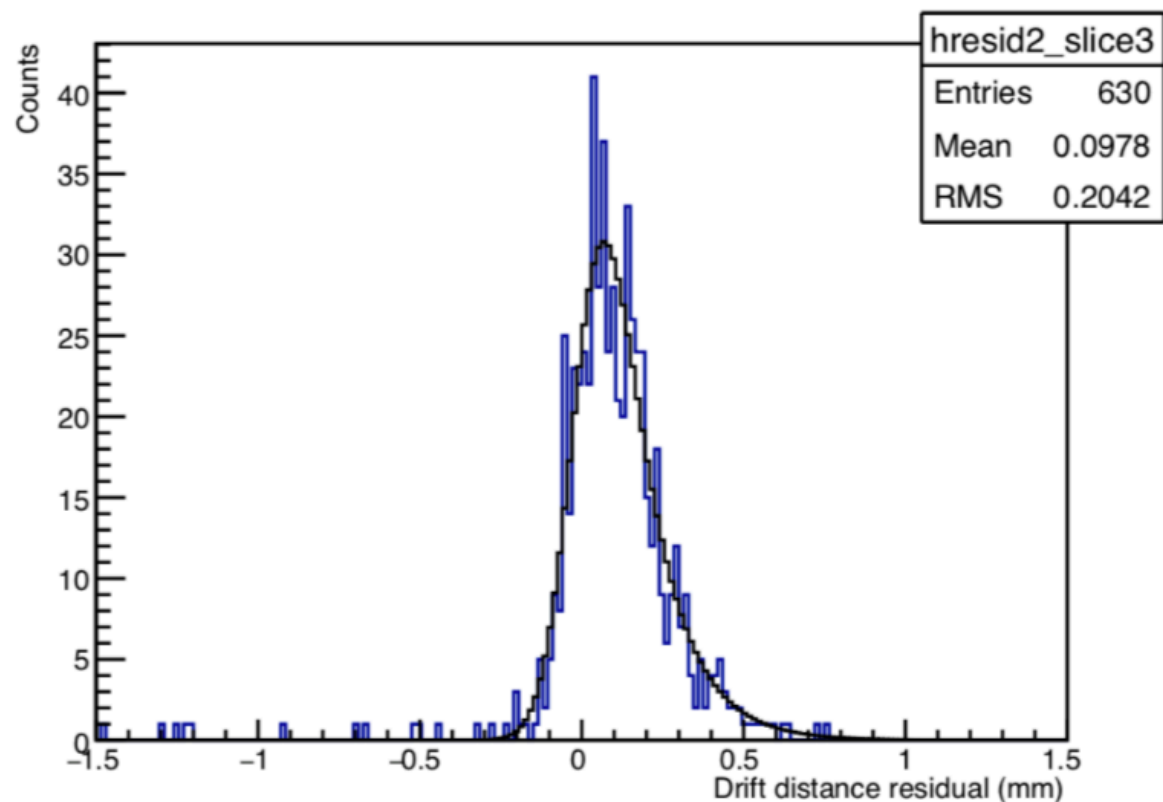


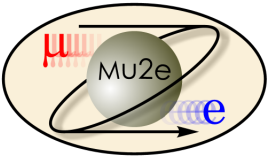




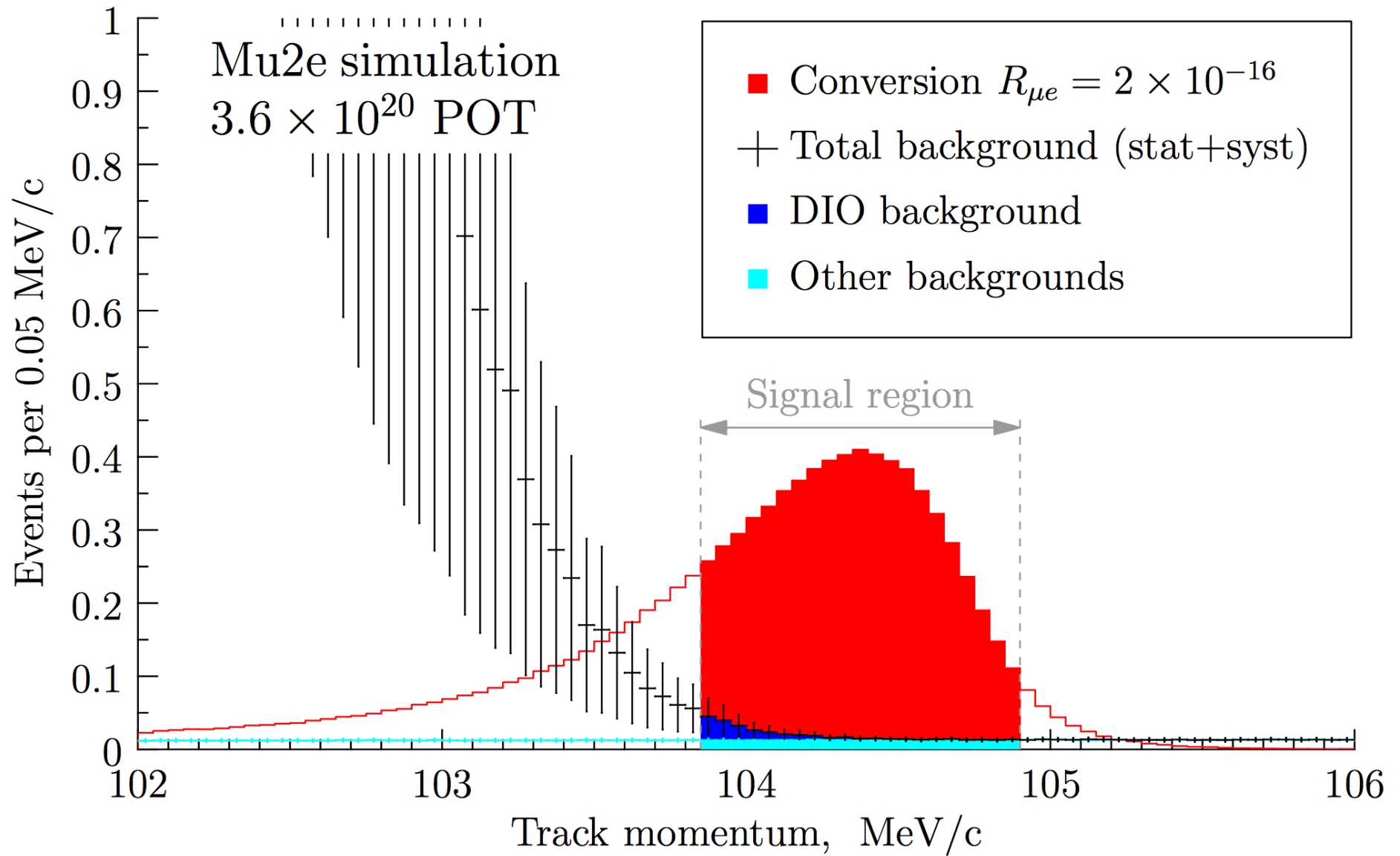
# Prototypes

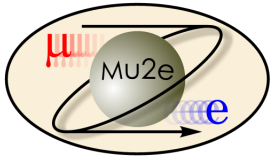
- straw resolution measured for each of the 8 straws using cosmic rays
- Simulation implements detailed model of physical and electronics reposes to G4 energy deposition
- ✓ Clusters → Drift → Current pulse → Voltage waveform → Digitization
- ✓ Tuned to literature and prototype





# Mu2e sensitivity





# Panel production



- Panel assembly @ Univ of Minnesota (with Fermilab and U. Houston)
- Full plane assembled at Fermilab
- QA of panel components @ CUNY, Duke and LBL/UC Berkeley



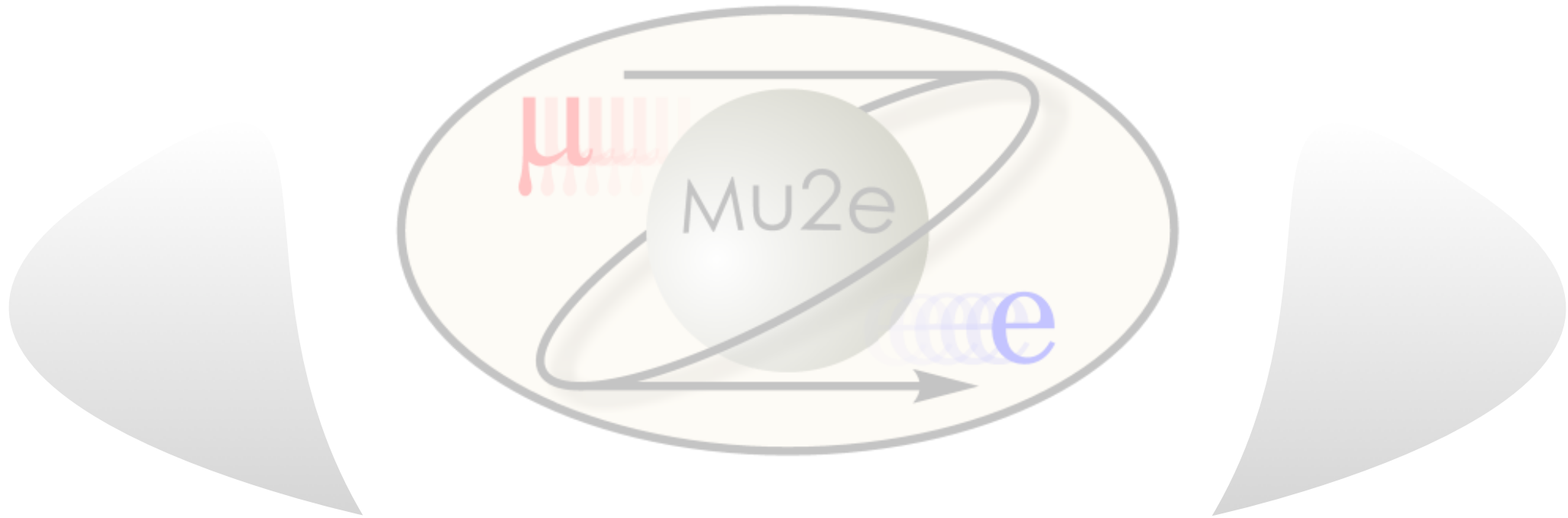


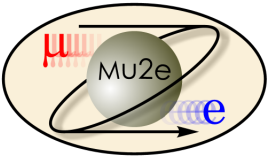
# Summary

- Mu2e will improve the sensitivity by four orders of magnitude
- The Mu2e tracker has been designed to:
  - ✓ provide accurate reconstruction of  $e^-$  momentum
  - ✓ operate in a very harsh environment (high  $\gamma$  and neutron dose)
- First 3 panels built and already under test:
  - ✓ good agreement of data with our Monte Carlo simulation
- QA process and plane assembly started

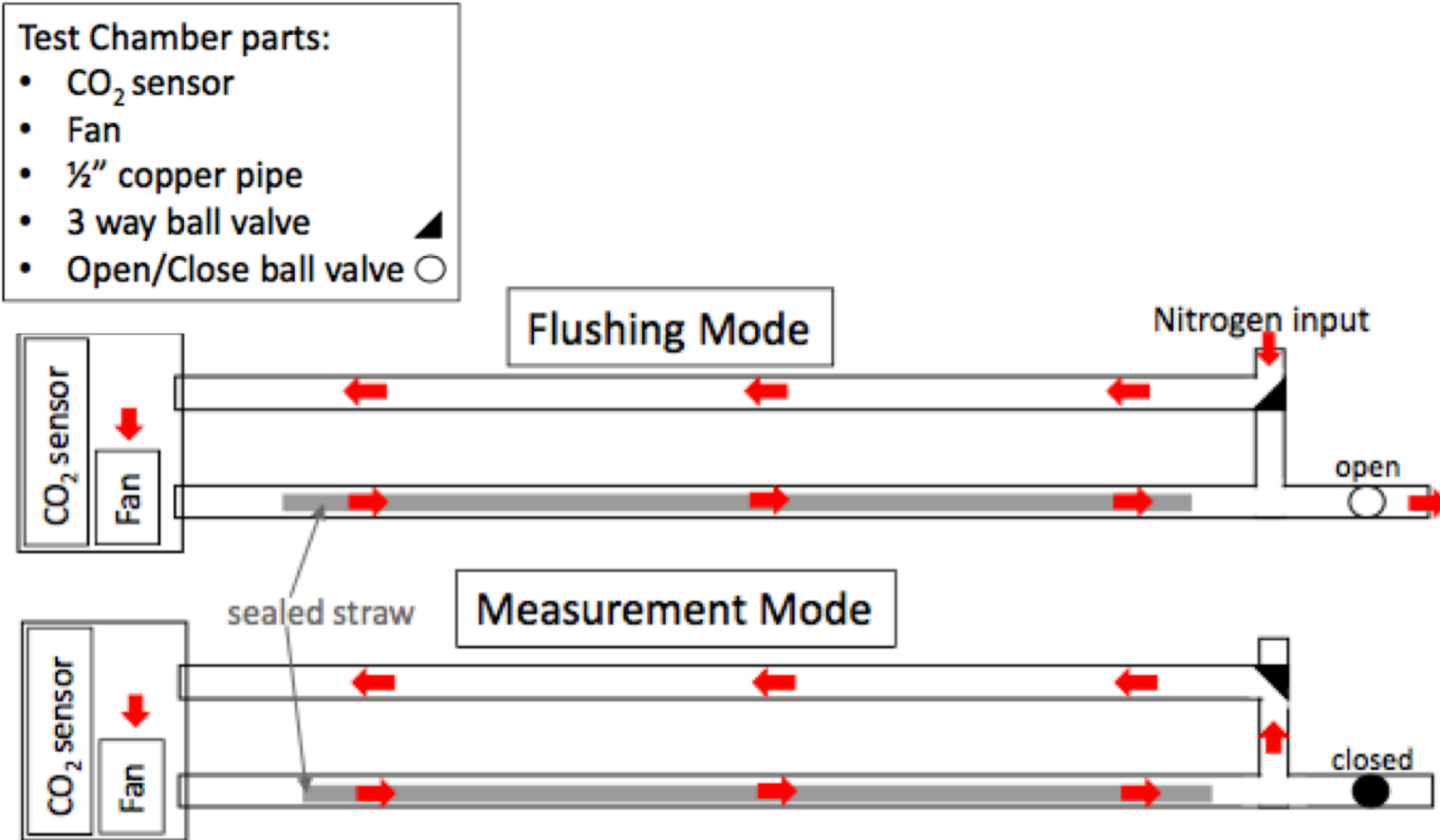


# backup slides

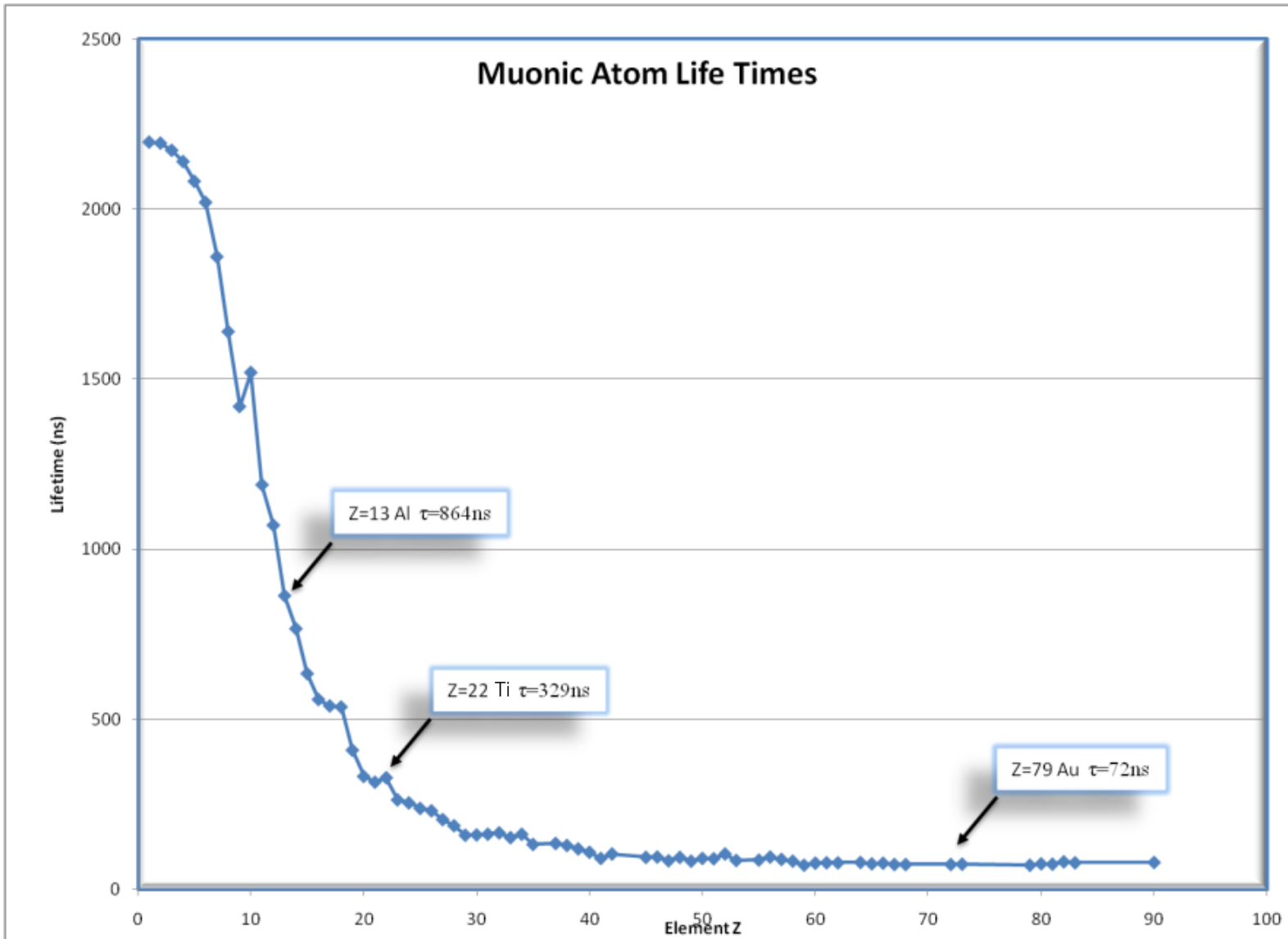
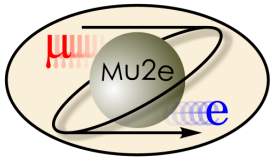


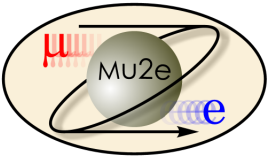


# Straw leak chamber



# Muonic atom life times

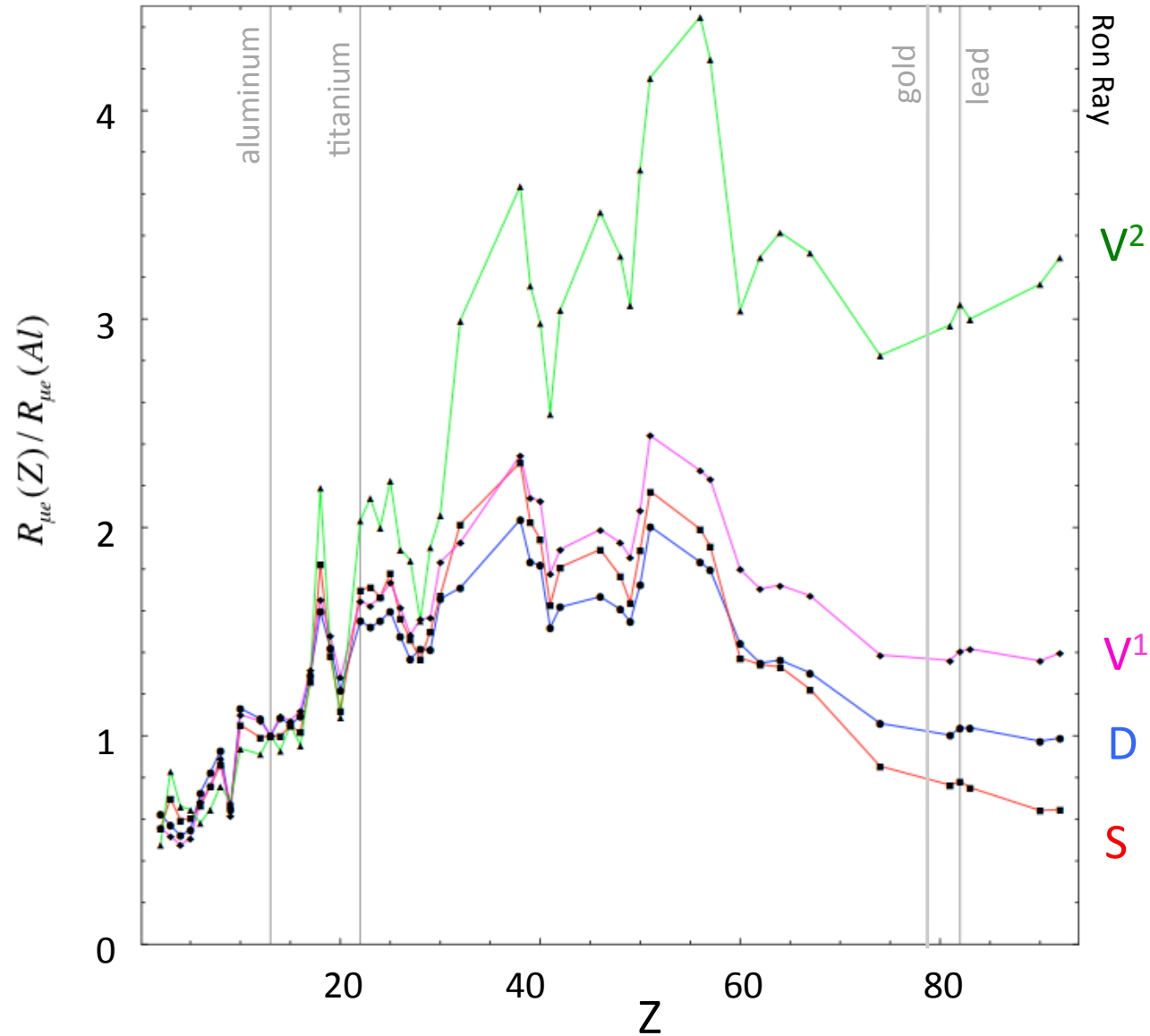




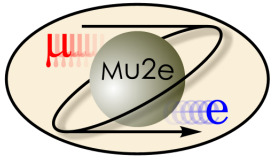
# $R_{\mu e}$ rate vs $Z$



V. Cirigliano et al., *phys. Rev.* **D80** 013002 (2009)







# Mu2e sensitivity



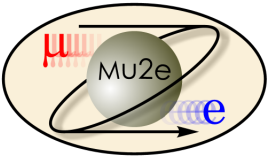
W. Altmannshofer, A.J.Buras, S.Gori, P.Paradisi, D.M.Straub

★★★★ = Discovery Sensitivity

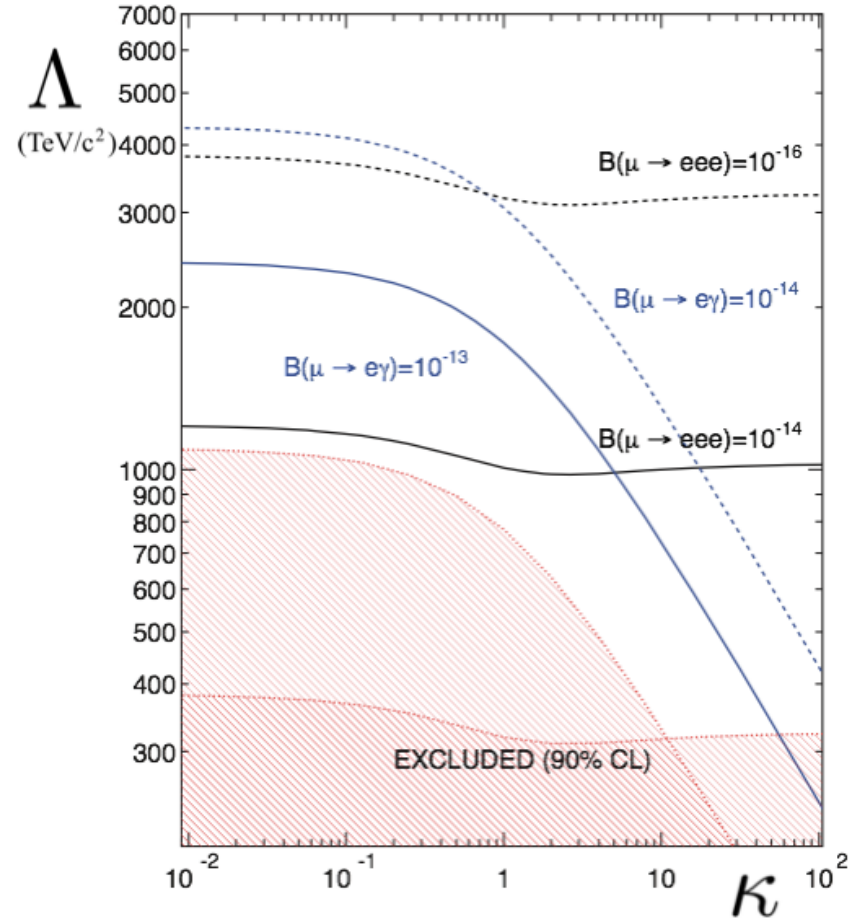
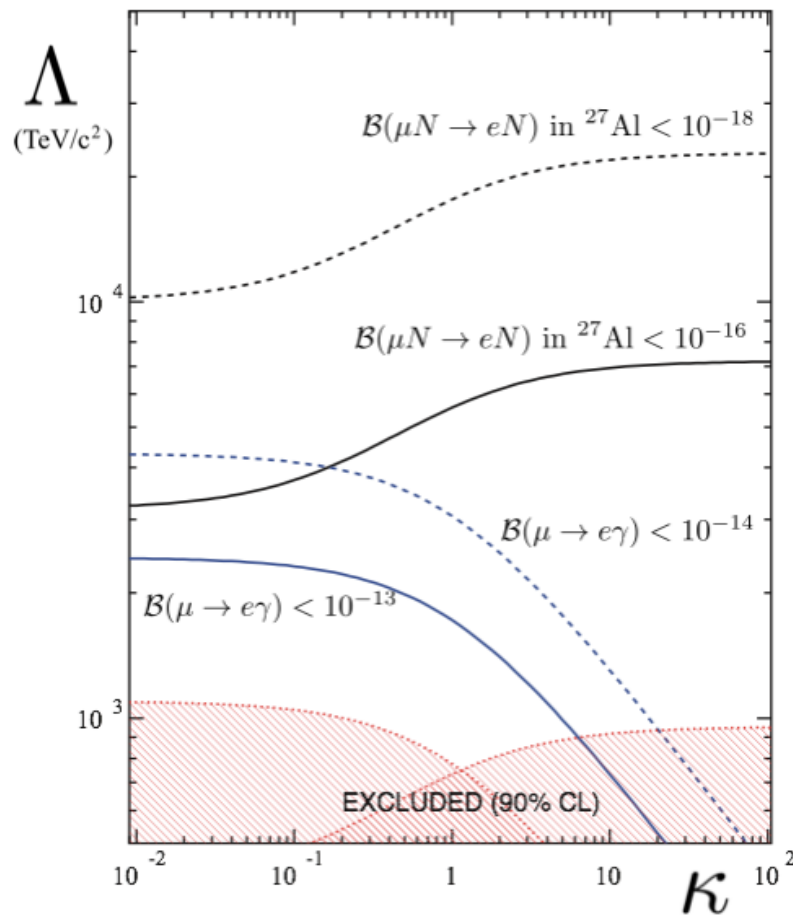
	AC	RVV2	AKM	$\delta LL$	FBMSSM	LHT	RS
$D^0 - \bar{D}^0$	★★★	★	★	★	★	★★★	?
$\epsilon_K$	★	★★★	★★★	★	★	★★	★★★
$S_{\psi\phi}$	★★★	★★★	★★★	★	★	★★★	★★★
$S_{\phi K_S}$	★★★	★★	★	★★★	★★★	★	?
$A_{CP}(B \rightarrow X_s \gamma)$	★	★	★	★★★	★★★	★	?
$A_{7,8}(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★★★	★★★	★★	?
$A_9(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★	★	★	?
$B \rightarrow K^{(*)} \nu \bar{\nu}$	★	★	★	★	★	★	★
$B_s \rightarrow \mu^+ \mu^-$	★★★	★★★	★★★	★★★	★★★	★	★
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	★	★	★	★	★	★★★	★★★
$K_L \rightarrow \pi^0 \nu \bar{\nu}$	★	★	★	★	★	★★★	★★★
$\mu \rightarrow e \gamma$	★★★	★★★	★★★	★★★	★★★	★★★	★★★
$\tau \rightarrow \mu \gamma$	★★★	★★★	★	★★★	★★★	★★★	★★★
$\mu + N \rightarrow e + N$	★★★	★★★	★★★	★★★	★★★	★★★	★★★
$d_n$	★★★	★★★	★★★	★★	★★★	★	★★★
$d_e$	★★★	★★★	★★	★	★★★	★	★★★
$(g-2)_\mu$	★★★	★★★	★★	★★★	★★★	★	?

arXiv:0909.1333[hep-ph]

Table 8: “DNA” of flavour physics effects for the most interesting observables in a selection of SUSY and non-SUSY models ★★★ signals large effects, ★★ visible but small effects and ★ implies that the given model does not predict sizable effects in that observable.



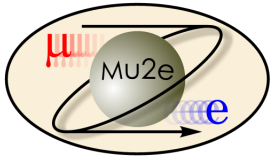
# Model independent Lagrangian



$$L_{\text{CLFV}} = \frac{m_\mu}{(\kappa + 1)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{\kappa}{(\kappa + 1)\Lambda^2} \bar{\mu}_L \gamma_\mu e_L (\bar{e} \gamma^\mu e)$$

“dipole term”

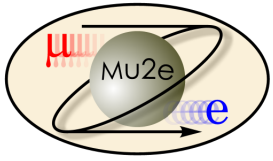
“contact term”



# CLFV limits I



Process	Upper limit
$\mu^+ \rightarrow e^+ \gamma$	$< 5.7 \times 10^{-13}$
$\mu^+ \rightarrow e^+ e^- e^+$	$< 1.0 \times 10^{-12}$
$\mu^- \text{Ti} \rightarrow e^- \text{Ti}$	$< 1.7 \times 10^{-12}$
$\mu^- \text{Au} \rightarrow e^- \text{Au}$	$< 7 \times 10^{-13}$
$\mu^+ e^- \rightarrow \mu^- e^+$	$< 3.0 \times 10^{-13}$
$\tau \rightarrow e \gamma$	$< 3.3 \times 10^{-8}$
$\tau^- \rightarrow \mu \gamma$	$< 4.4 \times 10^{-8}$
$\tau^- \rightarrow e^- e^+ e^-$	$< 2.7 \times 10^{-8}$
$\tau^- \rightarrow \mu^- \mu^+ \mu^-$	$< 2.1 \times 10^{-8}$
$\tau^- \rightarrow e^- \mu^+ \mu^-$	$< 2.7 \times 10^{-8}$
$\tau^- \rightarrow \mu^- e^+ e^-$	$< 1.8 \times 10^{-8}$
$\tau^- \rightarrow e^+ \mu^- \mu^-$	$< 1.7 \times 10^{-8}$
$\tau^- \rightarrow \mu^+ e^- e^-$	$< 1.5 \times 10^{-8}$



# CLFV limits 2

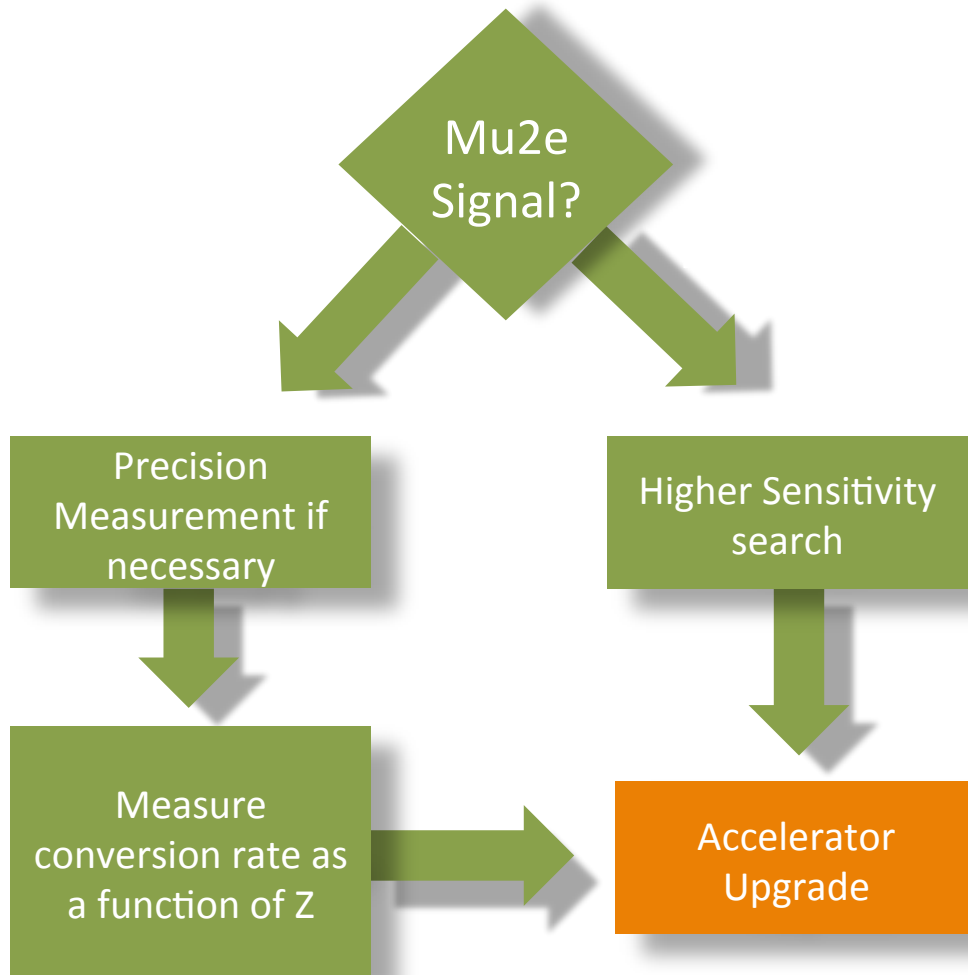


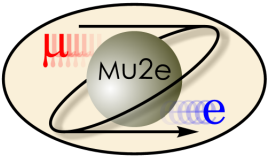
Process	Upper limit
$\pi^0 \rightarrow \mu e$	$< 8.6 \times 10^{-9}$
$K_L^0 \rightarrow \mu e$	$< 4.7 \times 10^{-12}$
$K^+ \rightarrow \pi^+ \mu^+ e^-$	$< 2.1 \times 10^{-10}$
$K_L^0 \rightarrow \pi^0 \mu^+ e^-$	$< 4.4 \times 10^{-10}$
$Z^0 \rightarrow \mu e$	$< 1.7 \times 10^{-6}$
$Z^0 \rightarrow \tau e$	$< 9.8 \times 10^{-6}$
$Z^0 \rightarrow \tau \mu$	$< 1.2 \times 10^{-6}$



# Mu2e signal?

- A next-generation Mu2e experiment makes sense in all scenarios:
  - ✓ Push sensitivity or
  - ✓ Study underlying new physics
  - ✓ Will need more protons  
upgrade accelerator
  - ✓ **Snowmass** white paper,  
arXiv:1802.02599





# Mu2e detector hall



North face

**put solenoid pictures here!!!**



East face



DS bay