



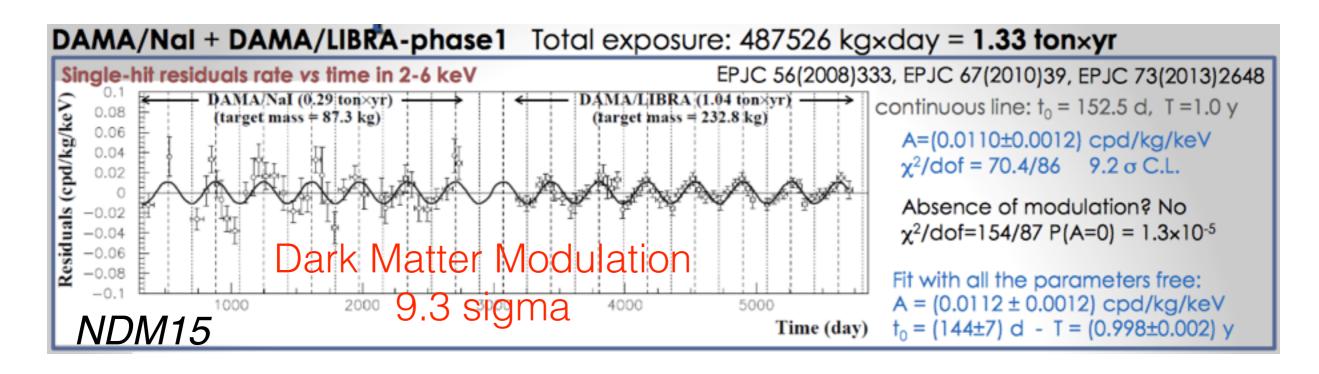
Status of the COSINE-100 experiment

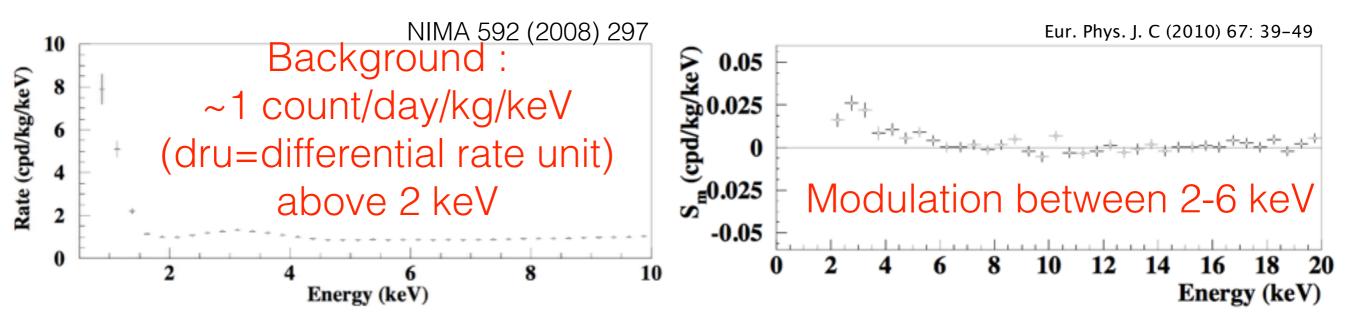


Chang Hyon Ha

On behalf of the COSINE-100 collaboration Center for Underground Physics (CUP), IBS, Korea

Motivation: The DAMA annual modulation signal, to be confirmed with independent measurements by the same NaI(TI) target material





DAMA/LIBRA-phase2

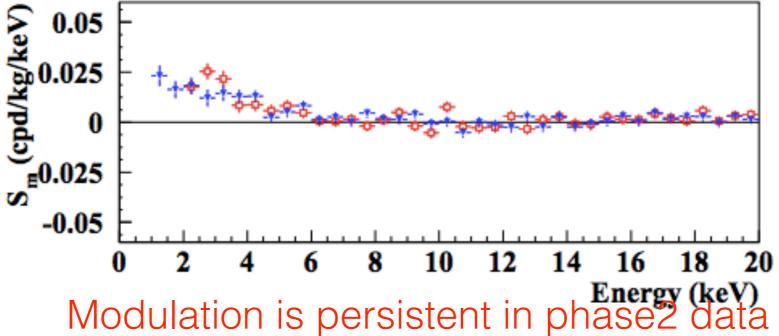
First model independent results from DAMA/LIBRA-phase2

R. Bernabei^{a,b}, P. Belli^{a,b}, A. Bussolotti^b, F. Cappella^{c,d}, V. Caracciolo^e, R. Cerulli^{a,b}, C.J. Dai^f, A. d'Angelo^{c,d}, A. Di Marco^b, H.L. He^f, A. Incicchitti^{c,d}, X.H. Ma^f, A. Mattei^d, V. Merlo^{a,b}, F. Montecchia^{b,g}, X.D. Sheng^f, Z.P. Ye^{f,h}

		A (cpd/kg/keV)	$T = \frac{2\pi}{\omega} \text{ (yr)}$	t_0 (days)	C.L.
DAMA/LIBRA-phase2:					
	1-3 keV	(0.0184 ± 0.0023)	1.0	152.5	8.0σ
	1-6 keV	(0.0105 ± 0.0011)	1.0	152.5	9.5σ
l .	$2-6~{ m keV}$	(0.0095 ± 0.0011)	1.0	152.5	8.6σ
İ	1-3 keV	(0.0184 ± 0.0023)	(1.0000 ± 0.0010)	153 ± 7	8.0σ
	1-6 keV	(0.0106 ± 0.0011)	(0.9993 ± 0.0008)	148 ± 6	9.6σ
	2-6 keV	(0.0096 ± 0.0011)	(0.9989 ± 0.0010)	145 ± 7	8.7σ
DAMA/LIBRA-phase1 + phase2:					
	$2-6~\mathrm{keV}$	(0.0095 ± 0.0008)	1.0	152.5	11.9σ
	2-6 keV	(0.0096 ± 0.0008)	(0.9987 ± 0.0008)	145 ± 5	12.0σ
DAMA/NaI + DAMA/LIBRA-phase1 + phase2:					
	2-6 keV	(0.0102 ± 0.0008)	1.0	152.5	12.8σ
i '	2-6 keV	(0.0103 ± 0.0008)	(0.9987 ± 0.0008)	145 ± 5	12.9σ

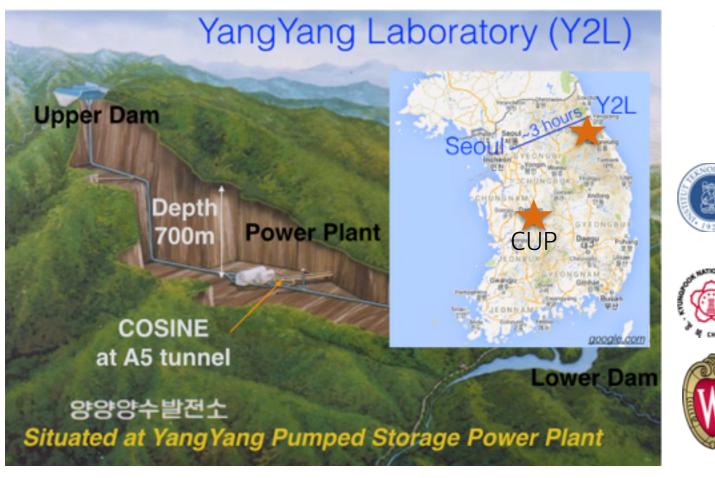
arXiv:1805.10486 (26 May 2018)

Down to 1 keV region



The COSINE-100 experiment

Joint collaboration between KIMS and DM-Ice to search for dark matter interactions in Nal(TI) scintillating crystals.











































COSINE-100 Construction Timeline

Dec. 2015 Jan. 2016 Feb. 2016

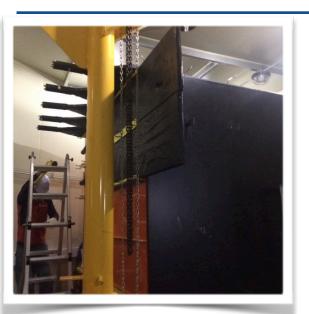








Mar. 2016 Apr. 2016









May. 2016 Jun. 2016 Sep. 2016

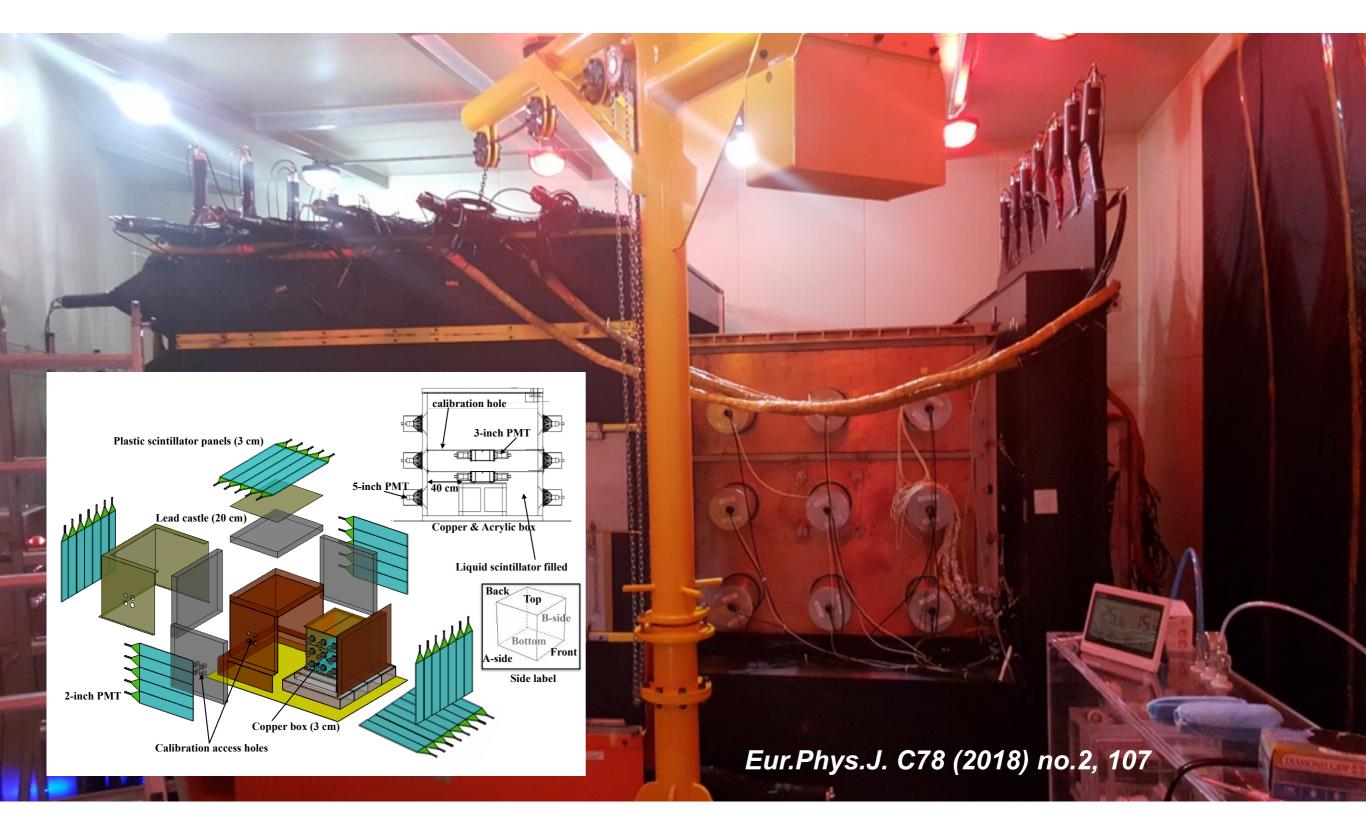




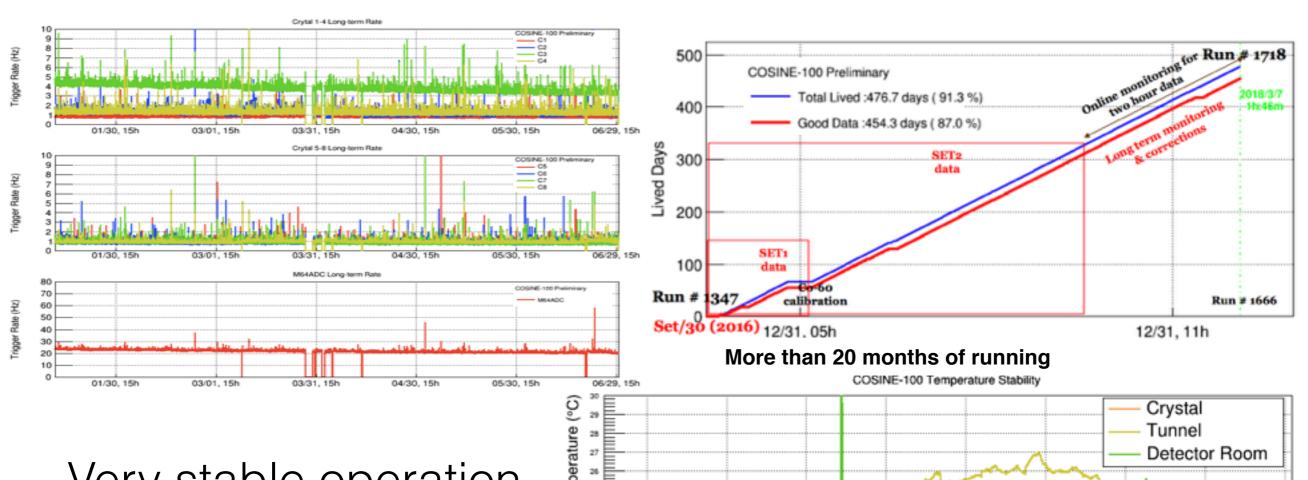




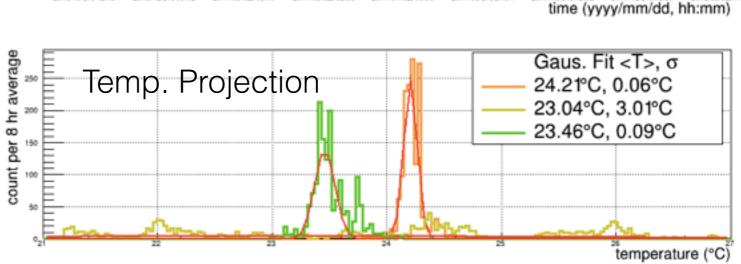
The COSINE-100 detector components



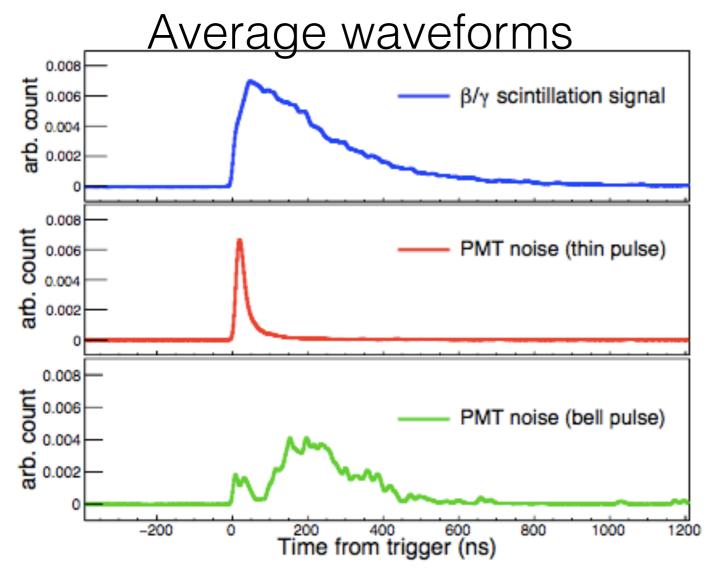
Environmental control/monitoring



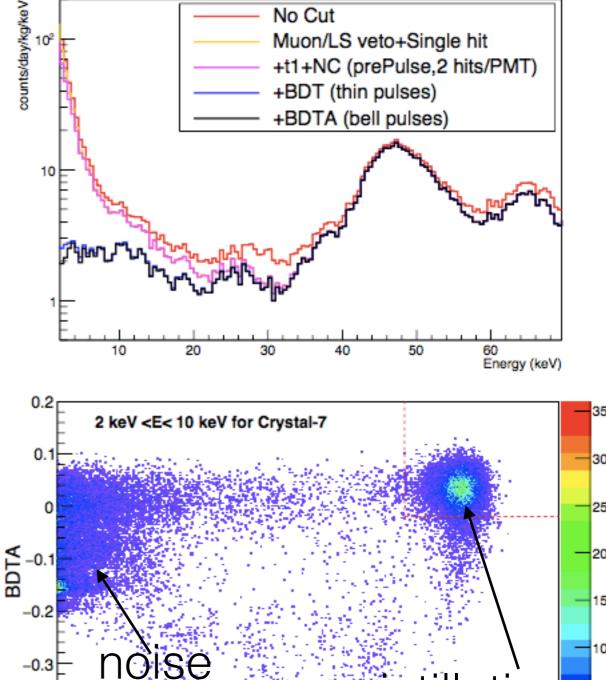
Very stable operation
More than 85%
of physics data.
95% are good runs.
Roughly 20 months
data are collected



PMT noise rejection



With Machine learning technique, we have successfully rejected PMT noise backgrounds.



-0.2

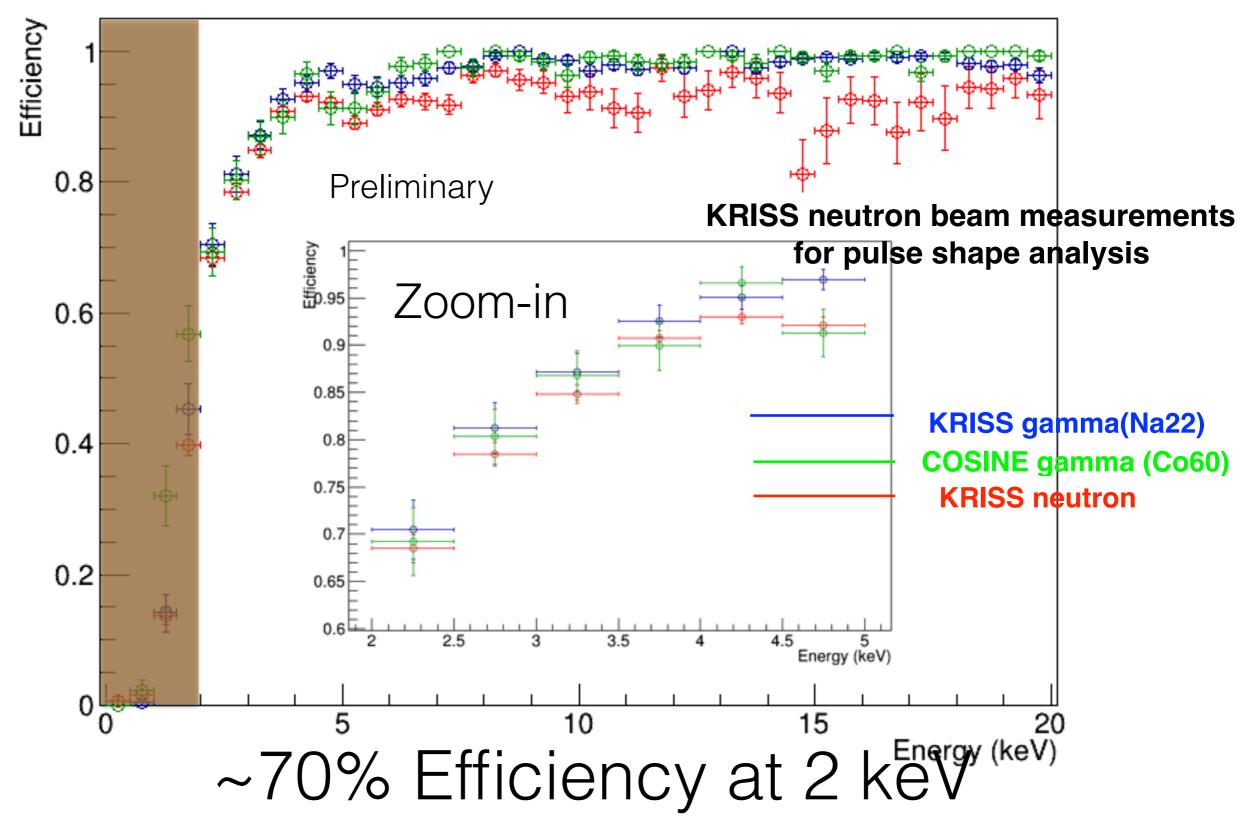
BDT

ICHEP2018, Seoul, July 4-11

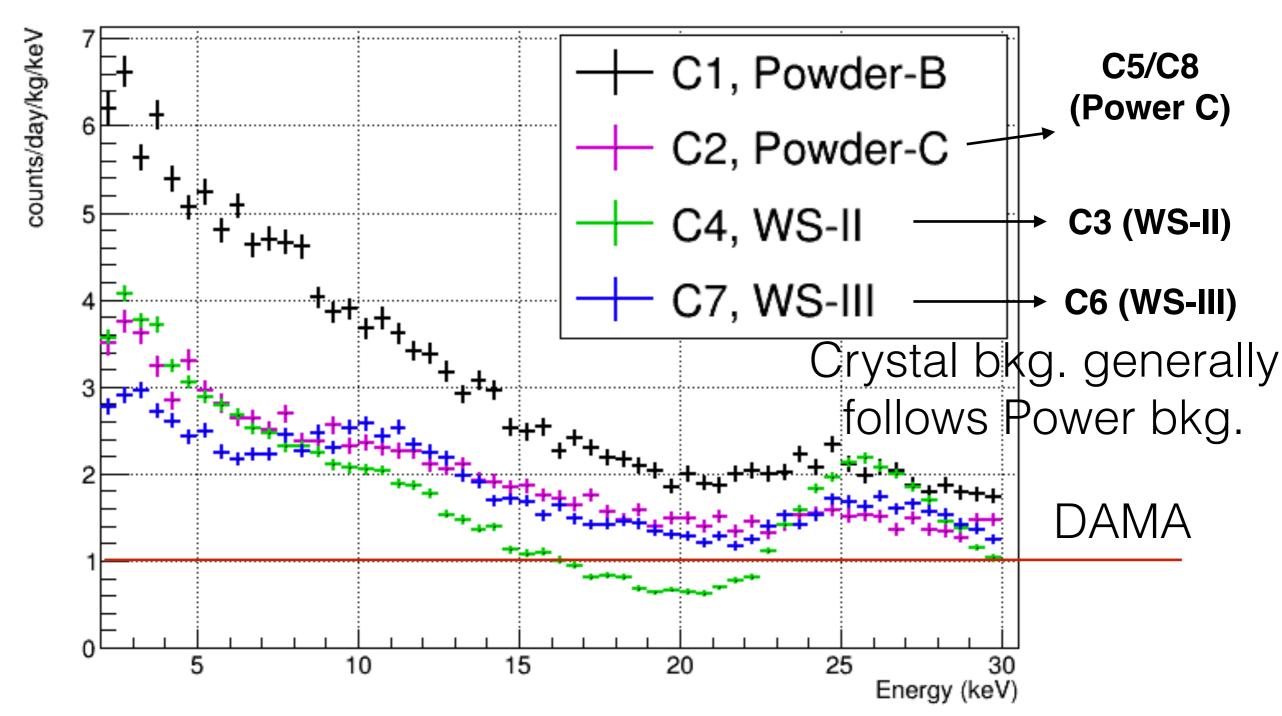
scintillation

-0.4

Analysis Selection Efficiency



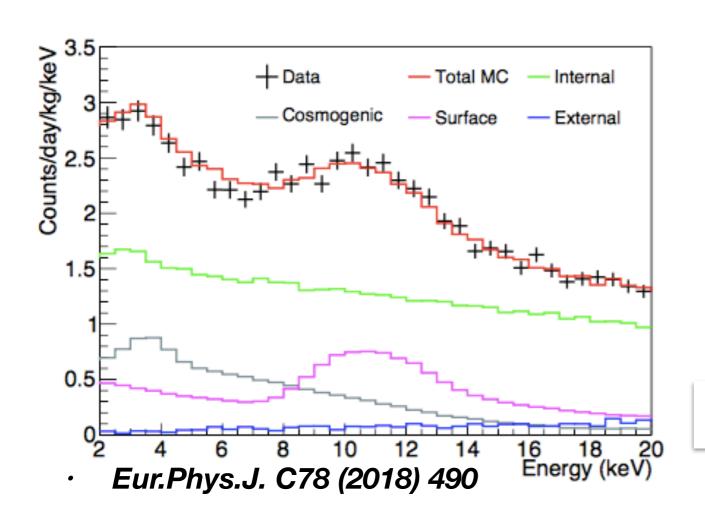
Energy Spectrum (Single-hit spectrum) Efficiency Corrected



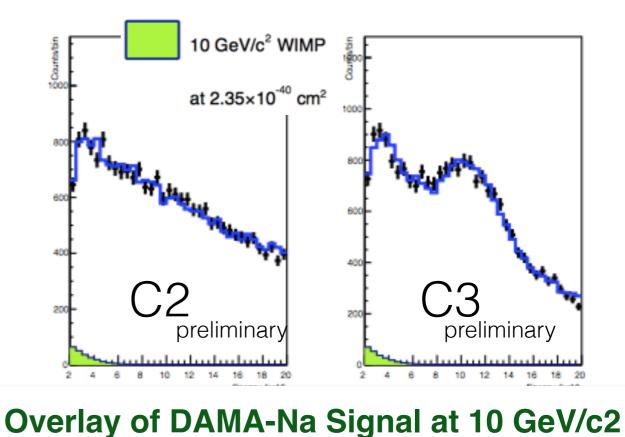
Average count rate at 2-6 keV is 3.5 counts/day/kg/keV

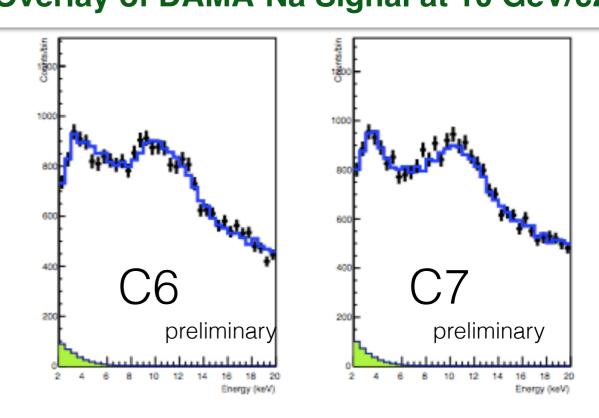
Understanding Background Counts/day/kg/ke Pb210 (internal) K40 + Cd109 External Pb210(Teflon) Th228+Ra226 Cd109 Ra226 Single-Hit Ra226 U238+Ra228 (2-6 keV region not used) 10^{-2} 10 30 60 400 600 800 10001200140016001800 Energy (keV) Counts/day/kg/k + Data Total MC Internal Te121m Cosmogenic Surface External K40 Na22 Ra226 Multiple-Hit Na22 10^{-1} Eur.Phys.J. C78 (2018) 490 10^{-2} 30 60 20 10 50 40 200 400 600 800 10001200140016001800 Energy (keV)

WIMP Search, 59.5 days of Data

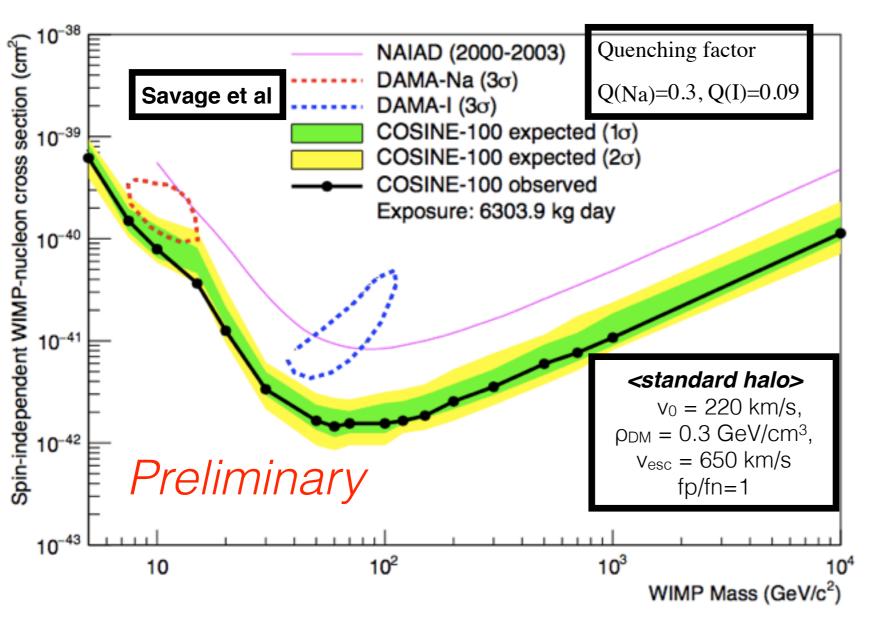


With bkg. understanding, 8 single-hit spectra are fit simultaneously with an assumed WIMP signal (SHM as described in Savage et al., Journal of cosmology and astrophysics), Note that bkg. understanding consideration from Kudryavtsev et al. Astropart. Phys. 33 (2010)91



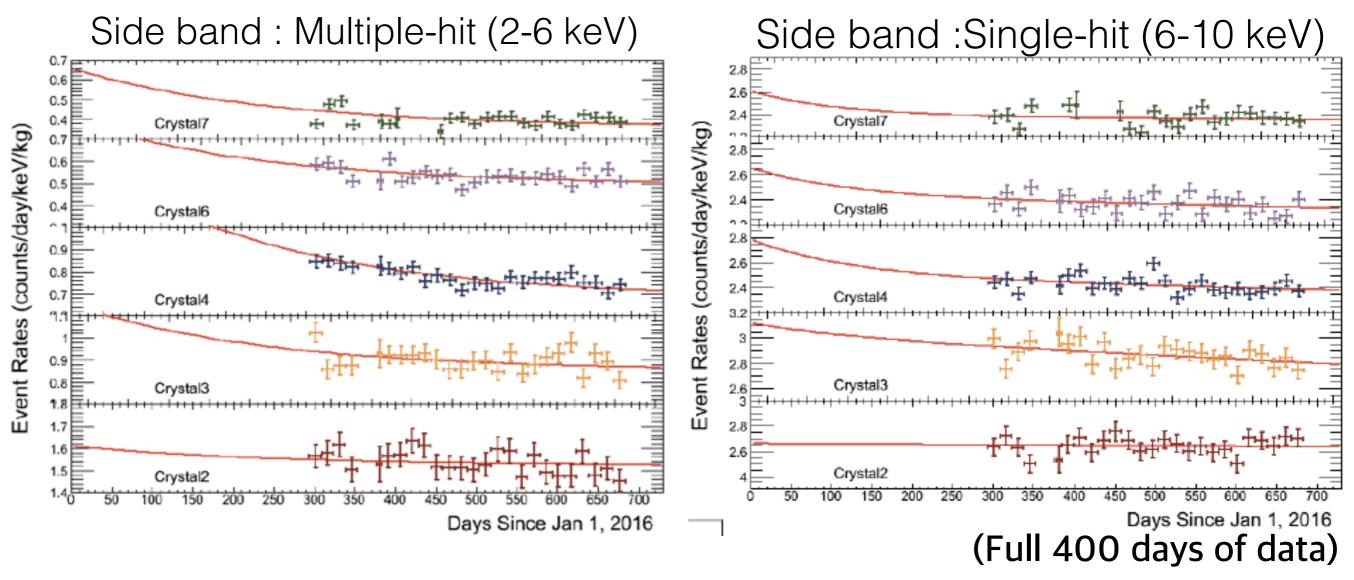


Spin independent WIMP-nucleon cross section limit (59.5 days of the COSINE-100 data)



- Spectrum with known sources of backgrounds
- COSINE-100 excludes DAMA/LIBRA-phase1's signal as spinindependent WIMP with Standard Halo Model in NaI(TI)
- Consistent with null results from other direct detect experiments with different target medium

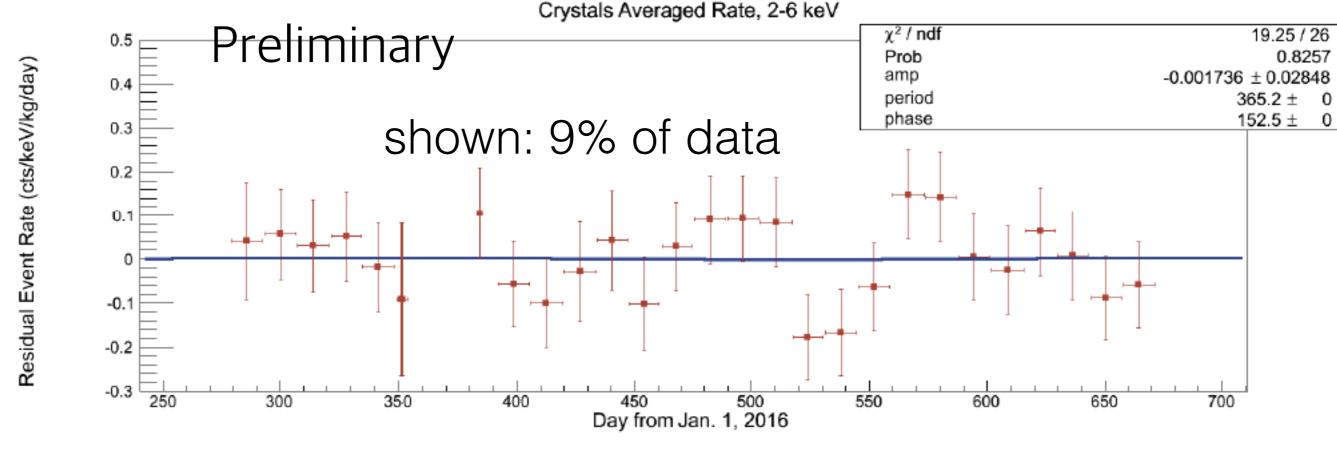
Annual Modulation Analysis



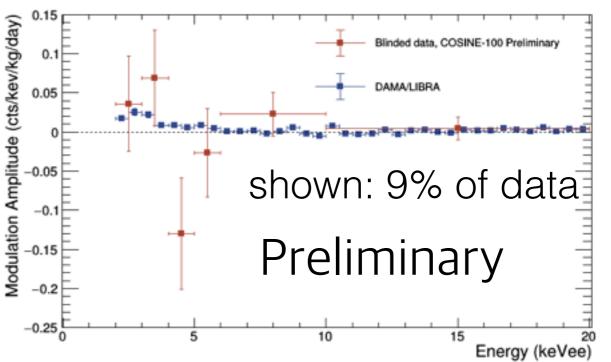
Crystal 1, 5, and 8 are excluded in this analysis due to low light yield and excessive PMT noise

Side band data fits well with simple exponential models built from the known cosmogenic components

Annual modulation analysis: Preliminary, Blinded

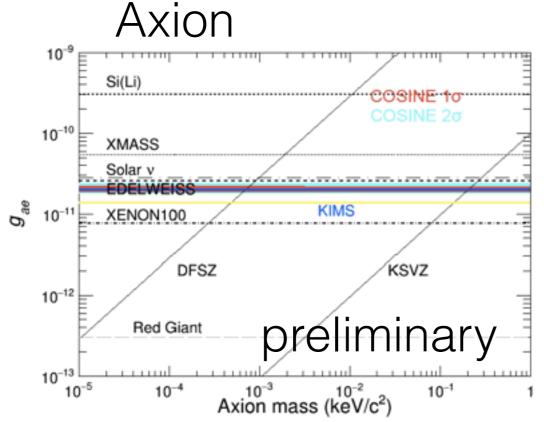


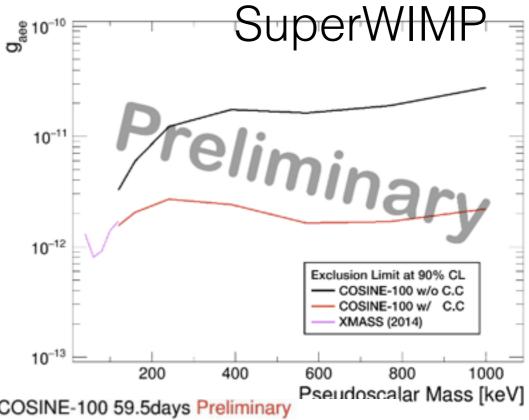
unblinding soon. Stay tuned!



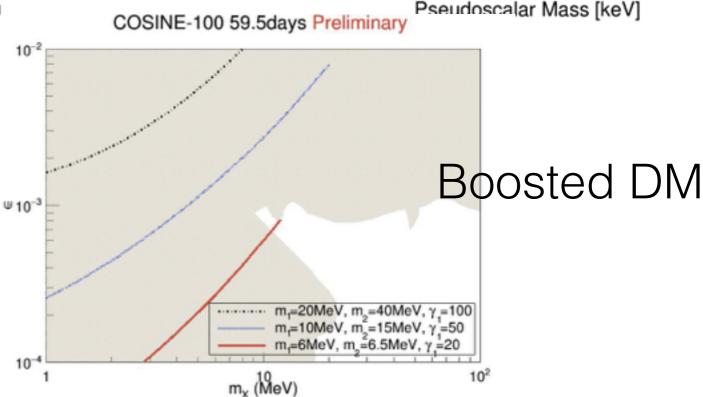
Other searches

16



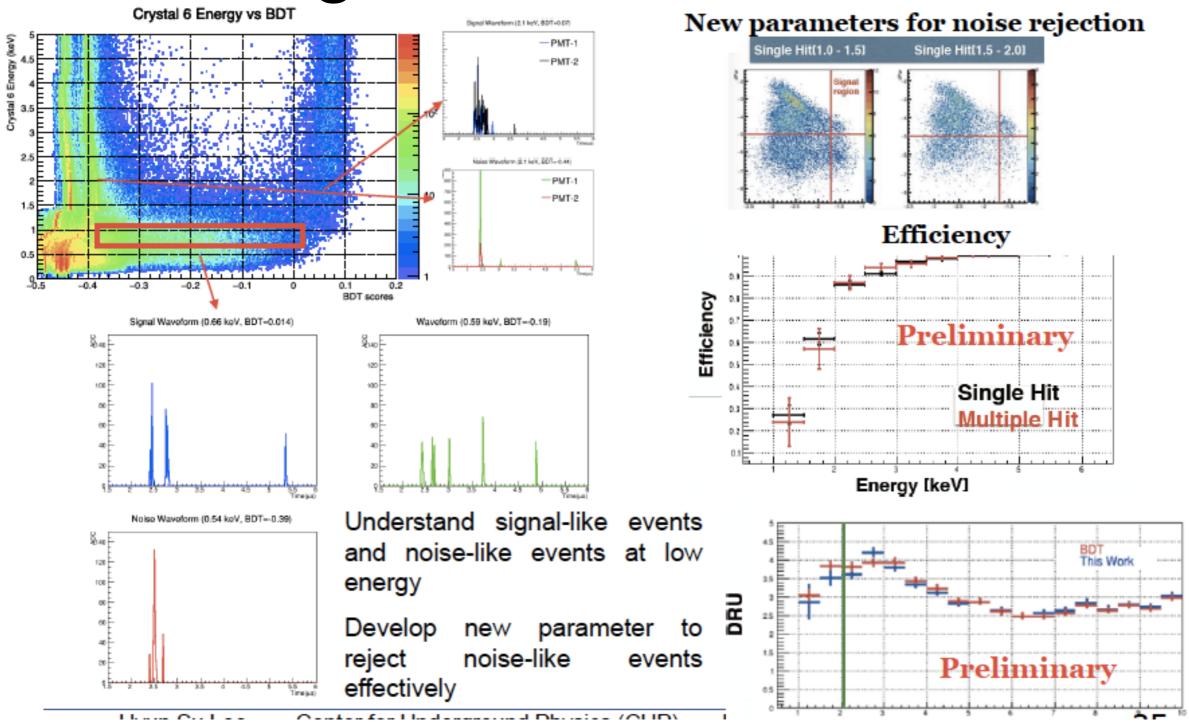


From the background understanding,
Other interesting searches are actively on-going



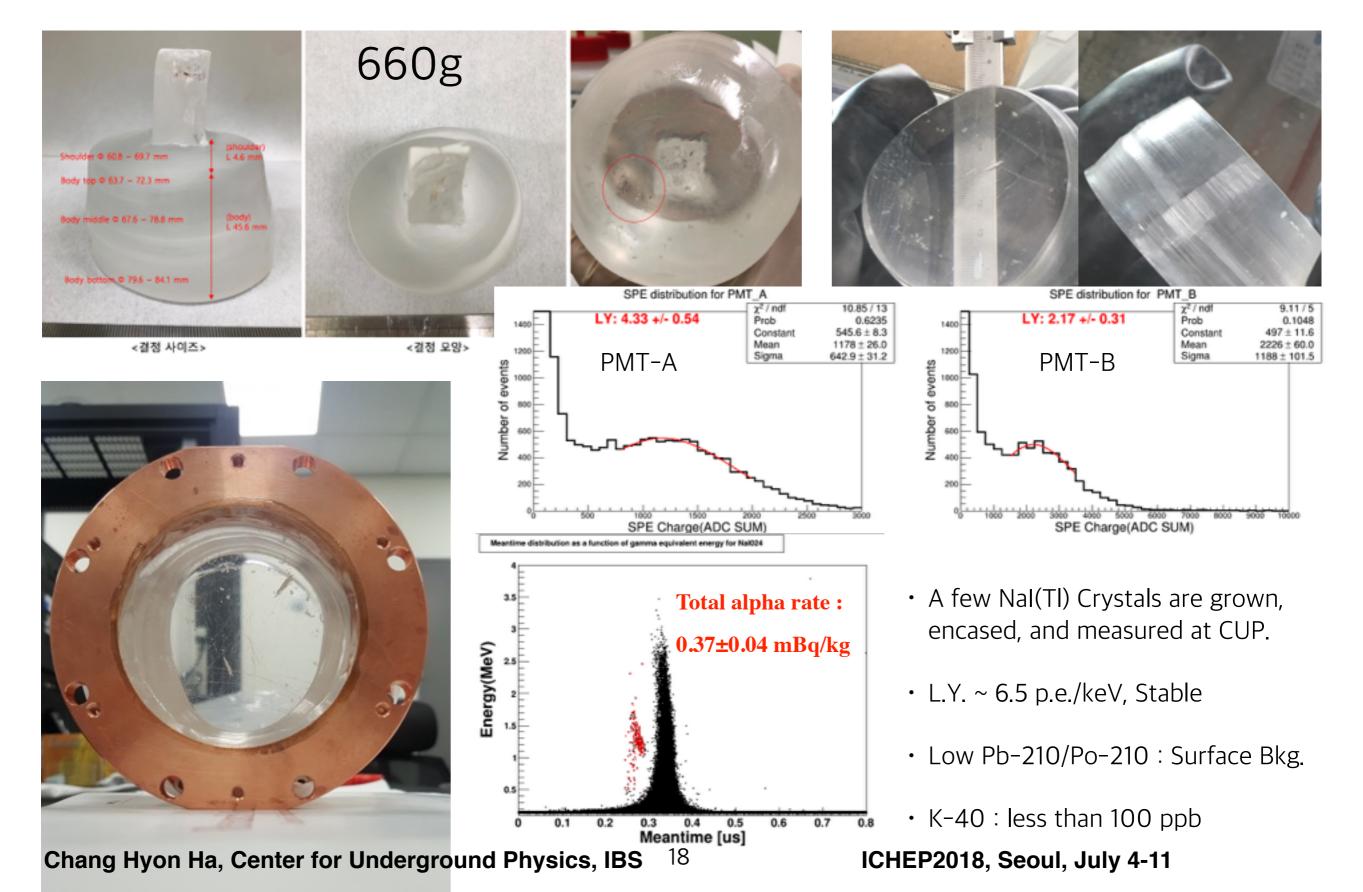
ICHEP2018, Seoul, July 4-11

Going to 1 keV threshold



In near future, we expect to minimize this noise as much as possible

Growing low radioactive NaI(TI) Crystals at CUP



Summary & Outlook

- The COSINE-100 experiment was installed at Y2L and runs smoothly for 20+ months.
- In COSINE-100, we have reached on average 3.5 counts/day/kg/keV with 2 keV thresholds.
- COSINE-100 confirms that DAMA's modulation signal cannot be from standard WIMP & SHM with NaI(TI).
- Modulation analysis is on-going.
- Analysis with lower threshold is underway
- Much progress made in developing the capabilities to grow and encapsulate radio-pure NaI(Tl) crystals at IBS-CUP