

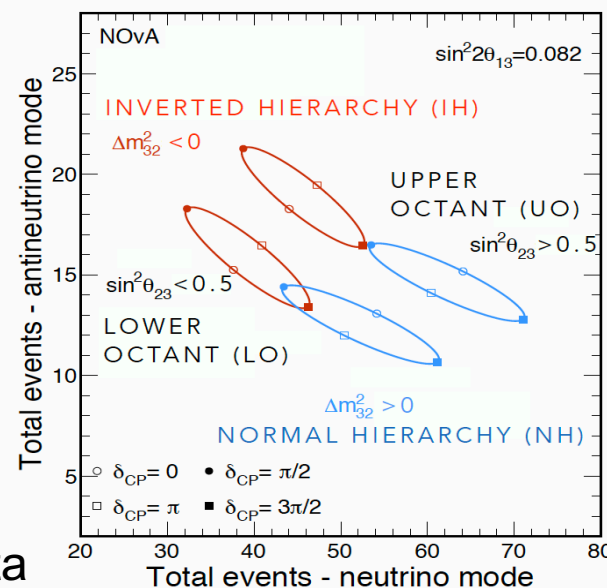
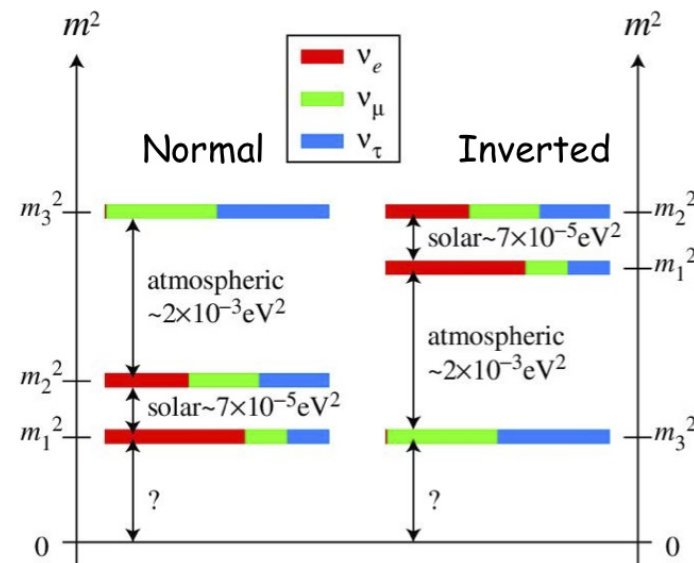
Latest Results from NOvA

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07-07-2018

ICHEP2018, Seoul, Korea

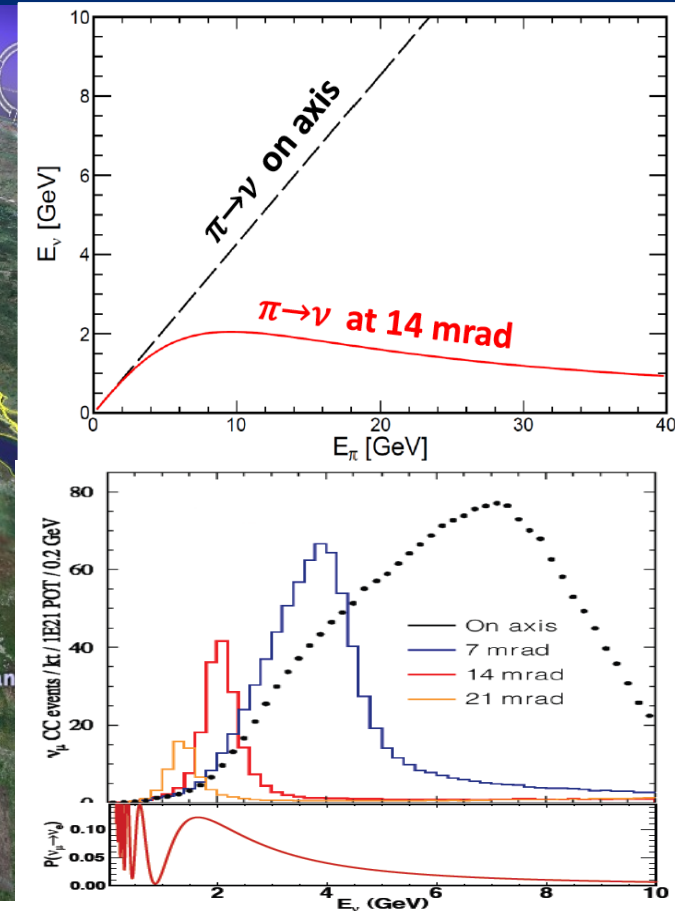
NOvA Physics Goals

- ν_e appearance + ν_μ disappearance
 - **Mass hierarchy**: $m_3 > m_{1,2}$ or $m_{1,2} > m_3$?
 - **CP phase δ_{CP}** : whether neutrinos and antineutrinos behave the same way in oscillation?
 - **Octant of θ_{23}** : Is ν_3 more strongly coupled to ν_τ or ν_μ ? Is θ_{23} exactly 45° ?
- NC disappearance
 - **Sterile neutrino search**: are there other neutrinos beyond the three known active flavors?
- Also, cross sections, exotic phenomena and non-beam physics



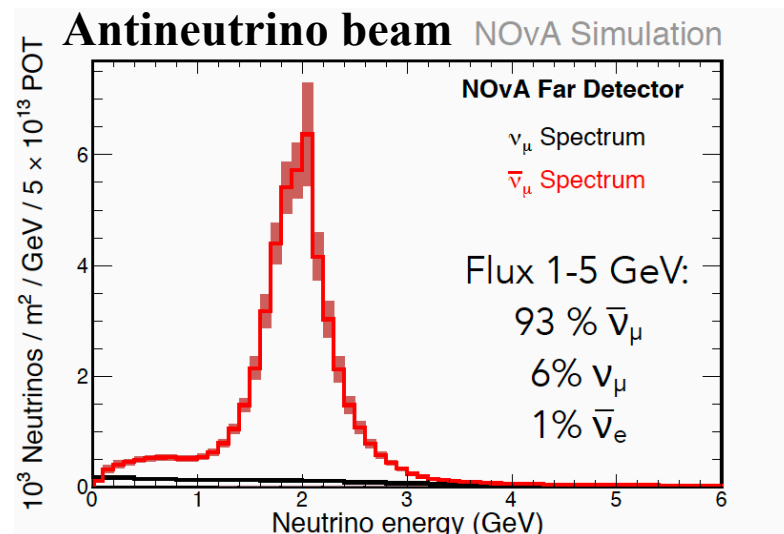
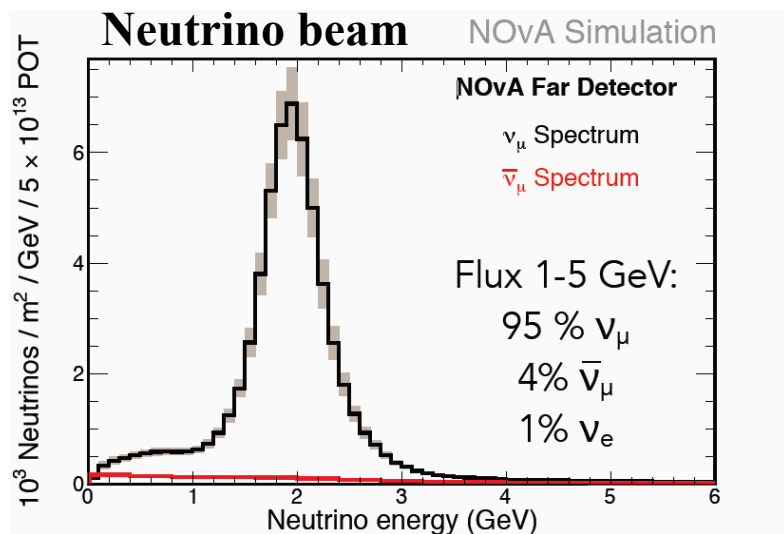
This talk: New results with NOvA's first antineutrino data

NuMI Off-Axis ν_e Appearance Experiment

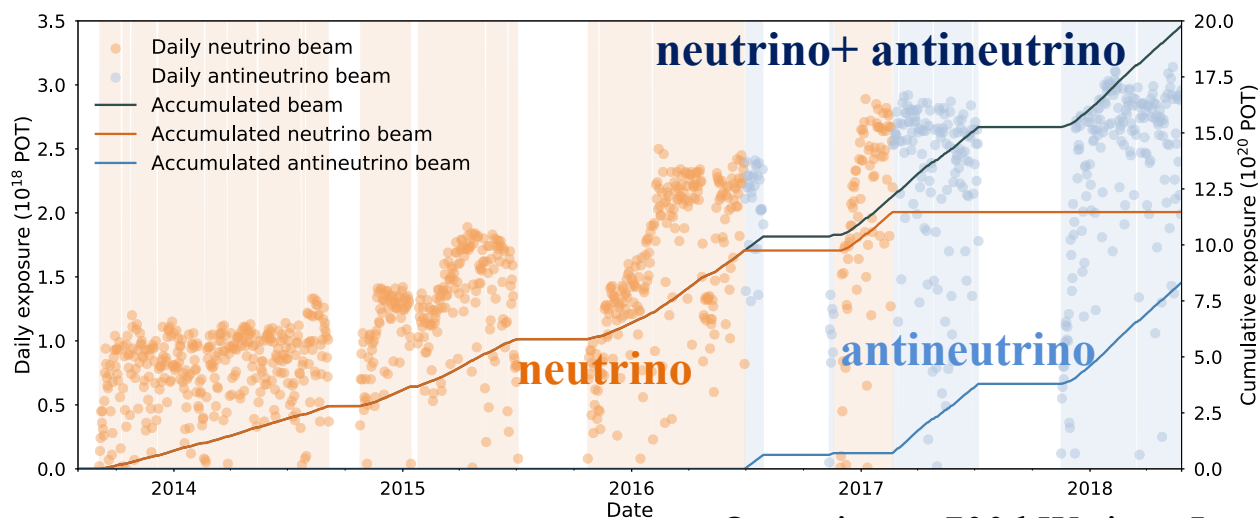


- Upgraded NuMI muon neutrino beam at Fermilab (700 kW design)
- Longest baseline in operation (810 km), large matter effect ($\pm 30\%$), sensitive to mass hierarchy
- Far/Near detector sited 14 mrad off-axis, narrow-band beam around oscillation maximum, small wrong sign components

Beam Performance



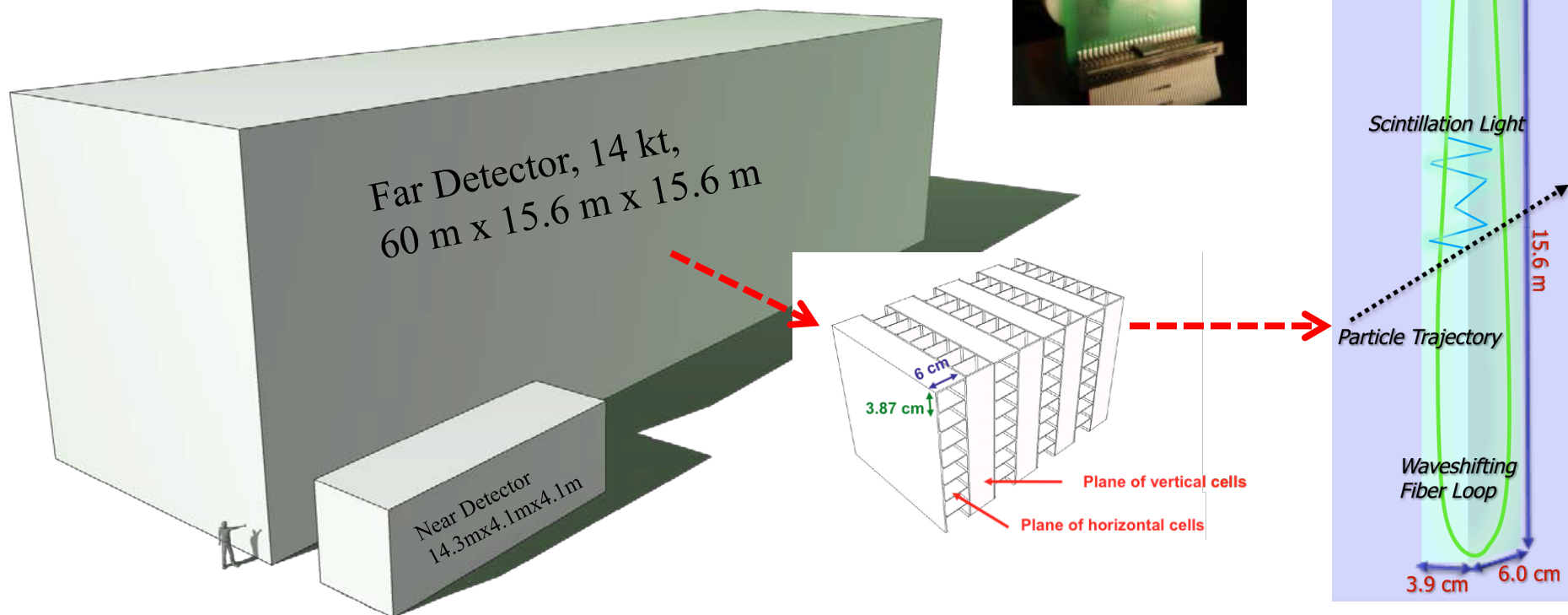
- **Neutrino beam data:** 8.85×10^{20} POT, taken Feb 2014 - Feb 2017
- **First antineutrino data:** 6.9×10^{20} POT, taken Feb 2017 - April 2018



Operating at 700 kW since Jan 2017

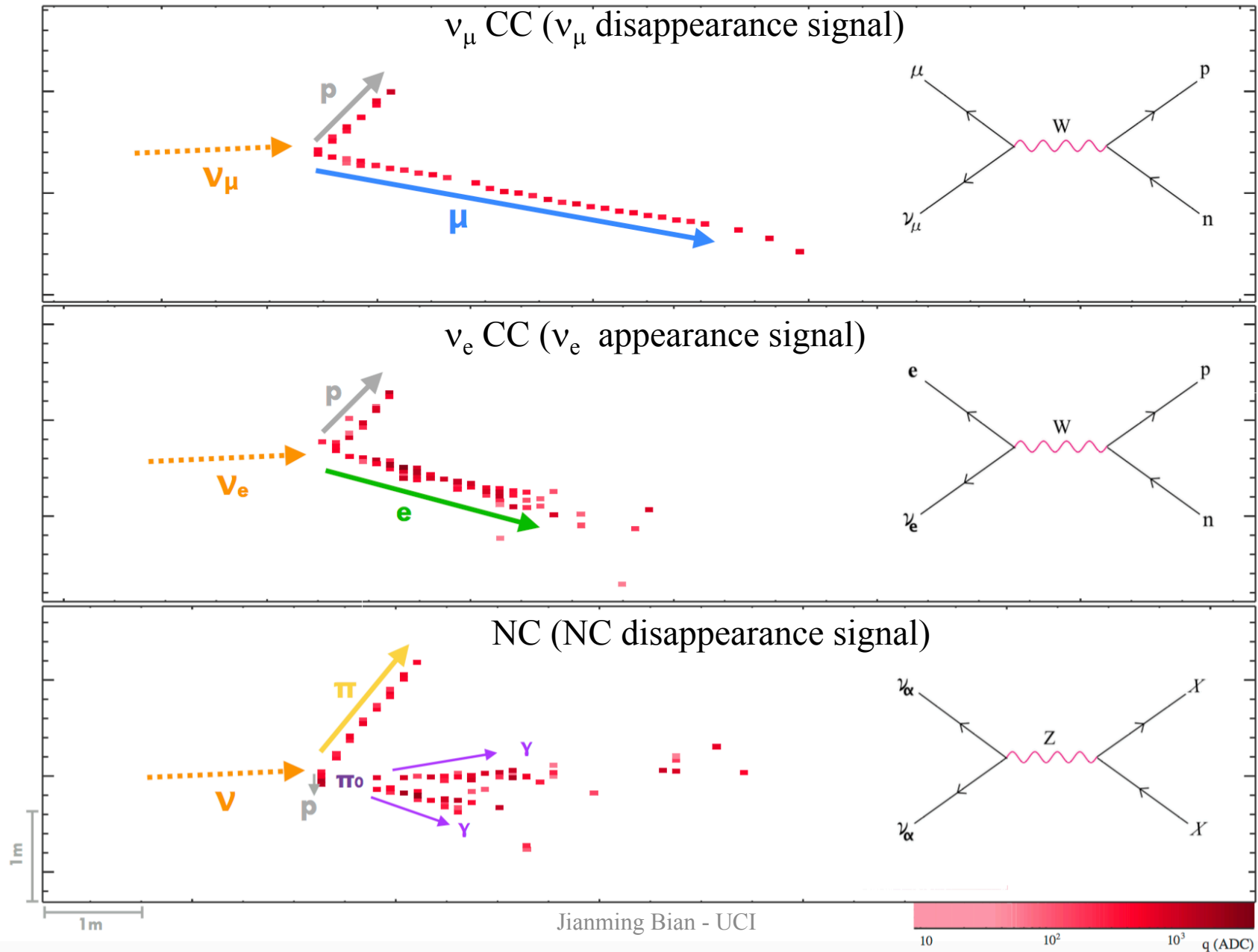
The NOvA Detectors

- 14-kton Far Detector
- 344,064 detector cells
- 0.3-kton functionally identical Near Detector
- 18,432 cells



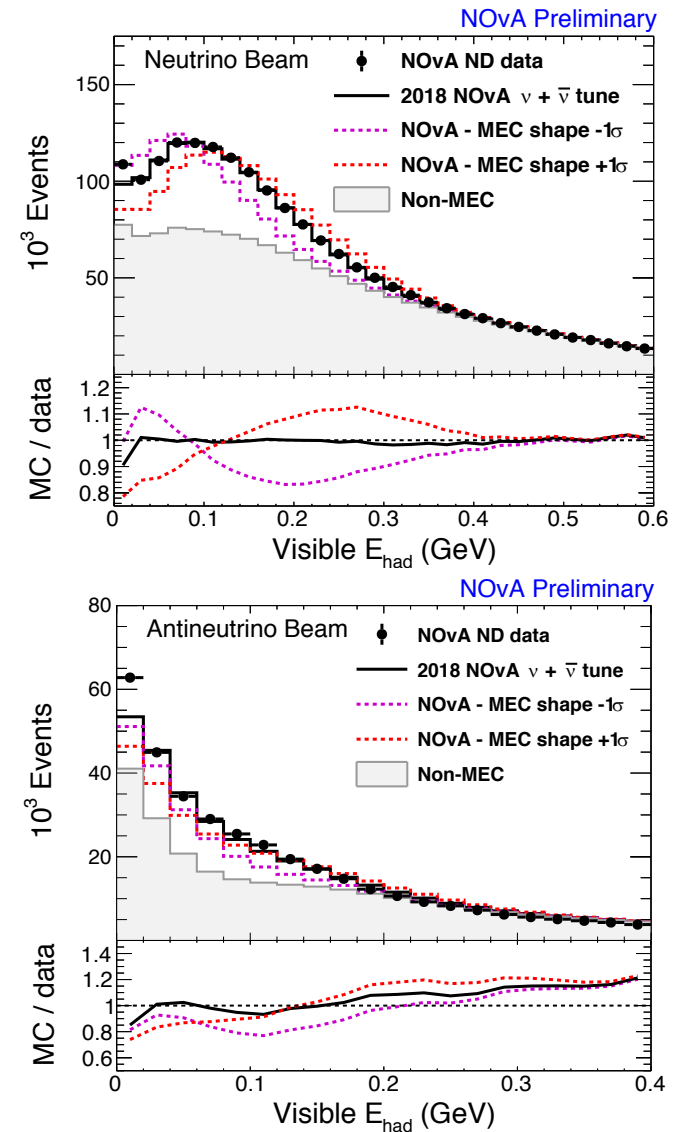
- Composed of PVC modules extruded to form long tube-like cells : 16m long in FD, 4m ND
- Each cell is filled with liquid scintillator and has a loop of wavelength-shifting fiber (WLS) routed to an Avalanche Photodiode (APD)
- Cells arranged in planes, assembled in alternating vertical and horizontal directions
- Low-Z and low-density, each plane just 0.15 X_0 . Great for e^- vs π^0

NOvA Event Topologies



Neutrino Interaction Tuning

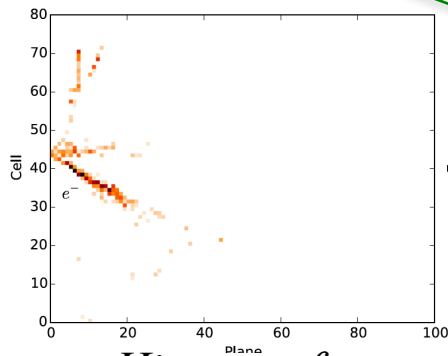
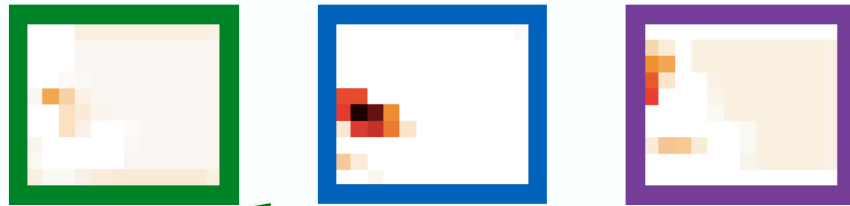
- QE, RES tuned to consider long-range nuclear correlations using València model via work of R. Gran (MINERvA) [<https://arxiv.org/abs/1705.02932>]
- DIS at high invariant mass ($W > 1.7 \text{ GeV}/c^2$) weighted up 10% based on NOvA data
- Empirical MEC (Meson Exchange Current) model for Multi-nucleon ejection (2p2h) [T. Katori, AIP Conf. Proc. 1663, 030001 (2015)], amount tuned in 2D 3-momentum and energy transfers ($q_0 = E_\nu - E_\mu$, $|\mathbf{q}| = |\mathbf{p}_\nu - \mathbf{p}_\mu|$) space to match ND data
- MEC shape systematic estimated by re-fitting using models with QE and RES related systematic shifts



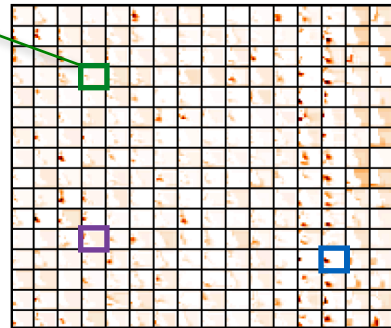
Deep-Learning based PID for ν_e and ν_μ Analyses

- CVN: a convolutional neural network (CNN), based on modern image recognition technology
- Introduce convolution filters to extract features from the hit map for the training of the neural net
- Statistical power equivalent to 30% more exposure than previous ν_e PIDs
- ν_e , ν_μ and NC analyses all use CVN as event selector

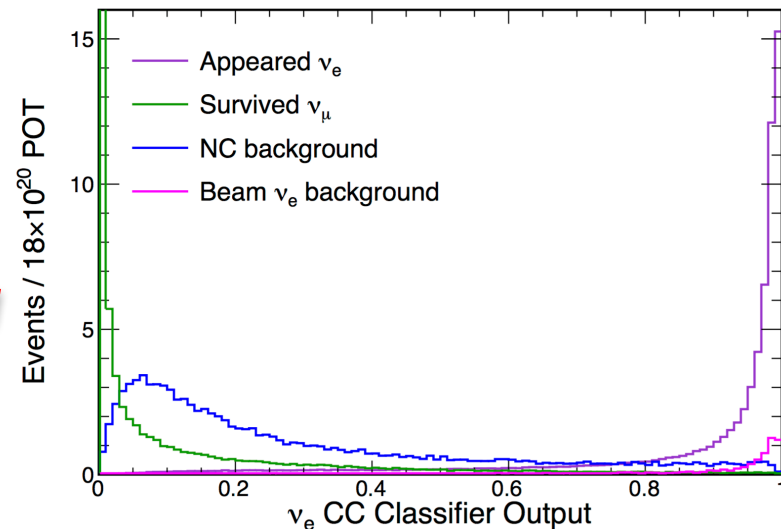
Outputs of convolutional filters (features)



*Hit map of
a ν_e CC event*



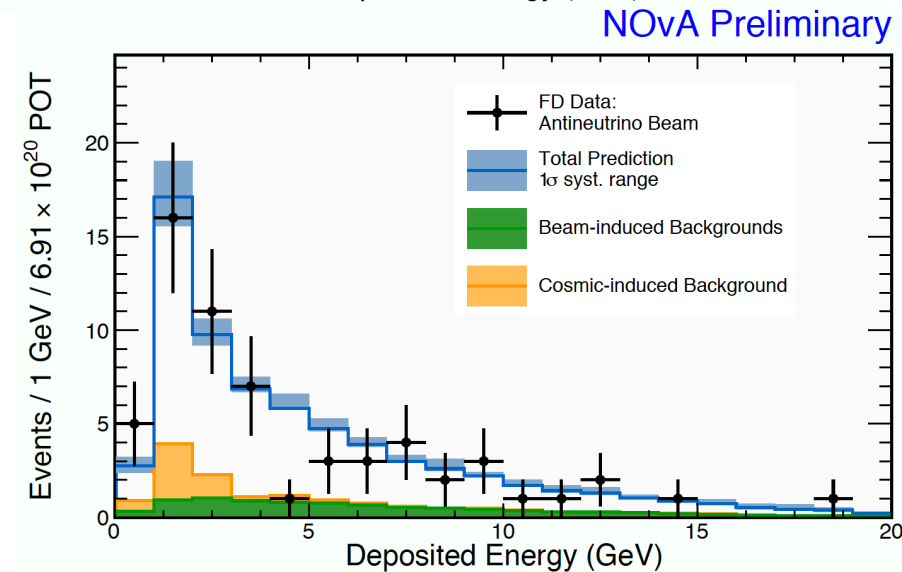
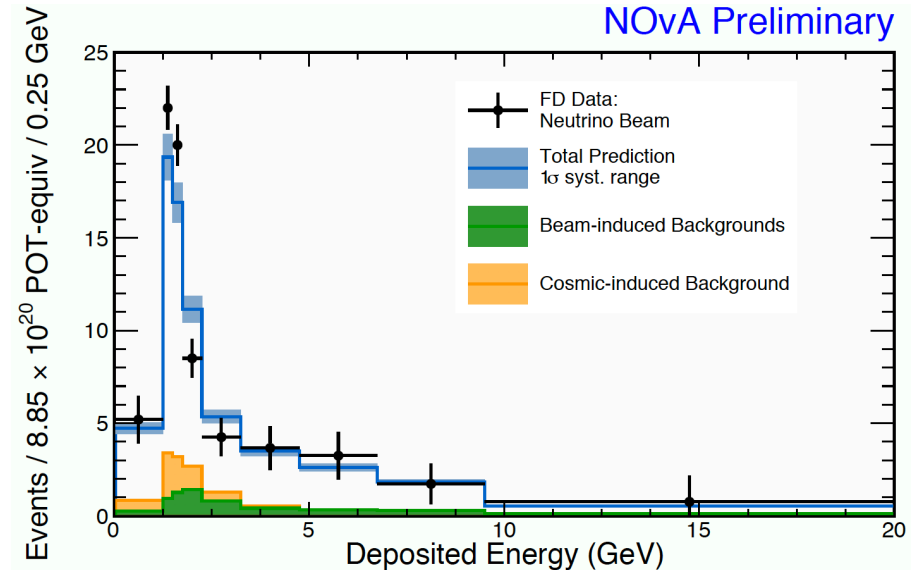
CVN output in the far detector MC



At NOvA, CVN has been extended to single particle ID, energy reconstruction, etc for future analyses

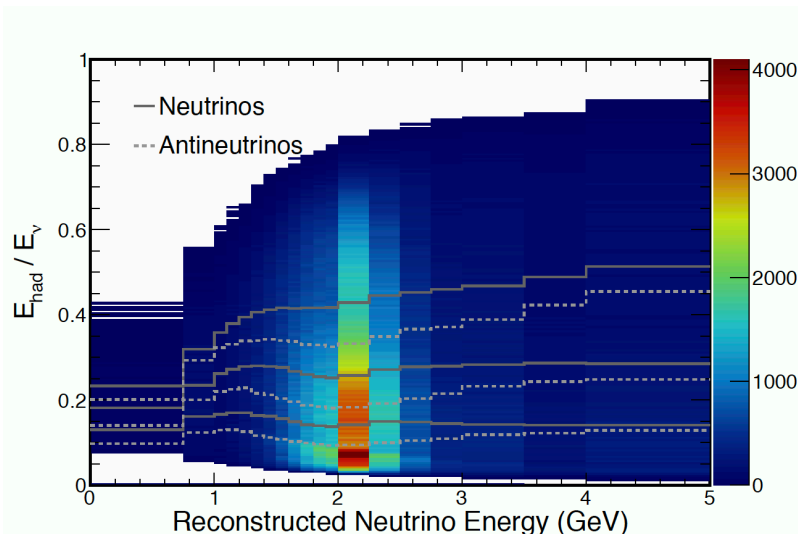
Observed NC events in Far Detector

- FD selection:
 - NC CVN selection applied
 - Additional Deep-learning based cosmic rejection
- Neutrino beam:
 - Observe 201 events, predict 188 ± 13 (syst.) events (38 bkg.)
- Antineutrino beam:
 - Observe 61 events, predict 69 ± 8 (syst.) events (16 bkg.)
- No significant suppression for NC observed, consistent with 3-flavor oscillation

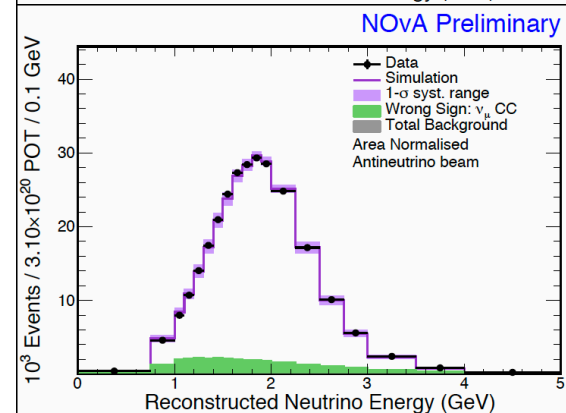
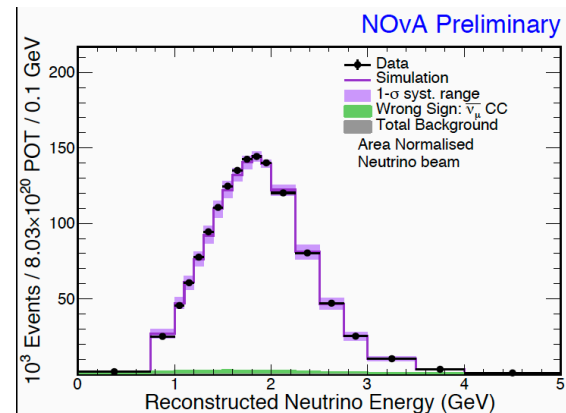


Near Detector Spectrum (ν_μ disappearance)

- Select ν_μ ($\bar{\nu}_\mu$) CC in ND from neutrino (antineutrino) beam, wrong sign contamination 3% (11%)
- $E_\nu = E_\mu + E_{\text{had}}$, data split in 4 equal energy quantiles based on E_{had}/E_ν , resolution varies from 5.8% (5.5%) to 11.7% (10.8%) for neutrino (antineutrino) beam.
- Normalize ND MC to data in each E_ν bin, then extrapolate the 4 quantiles to FD



Reco ν_μ ($\bar{\nu}_\mu$) energy, all Quartiles



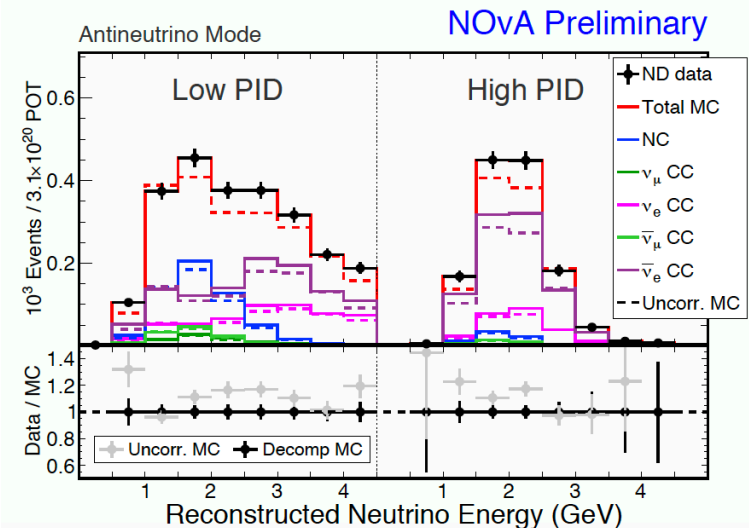
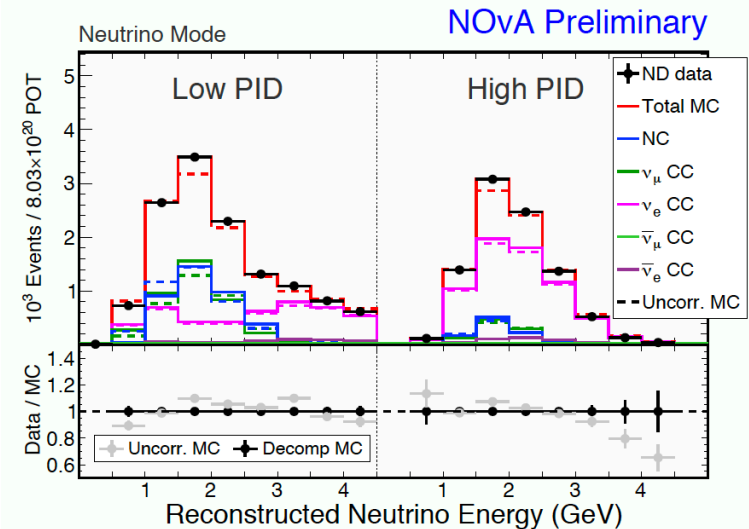
Area-normalized, **shape-only systematics**

Data/MC normalization difference:

1.3% and 0.5% for ν_μ and $\bar{\nu}_\mu$

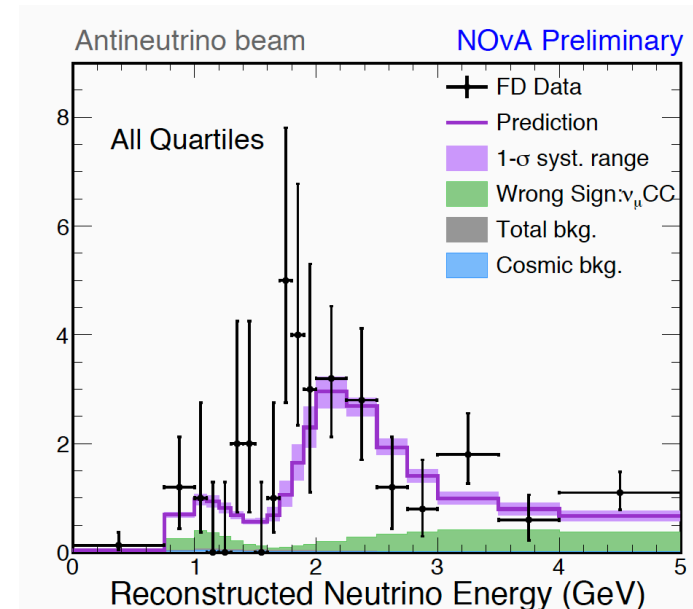
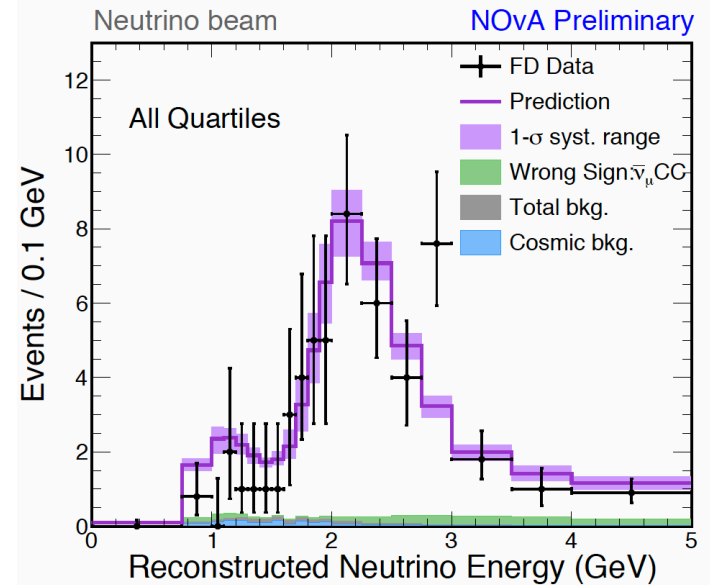
Near Detector Spectrum (ν_e appearance)

- Select ν_e ($\bar{\nu}_e$) CC in ND from neutrino (antineutrino) beam
- $E_\nu = f(E_e, E_{\text{had}})$, data split into low and high particle ID (purity) range
- For neutrino beam:
 - Contained and uncontained ν_μ events constrain the π/K contributions to the beam ν_e 's.
 - Michel electrons constrains NC/ ν_μ CC balance in each E_ν bin
- For antineutrino beam, scale all components evenly to match data
- ND \rightarrow FD extrapolation: Each component propagated independently in energy and PID bins



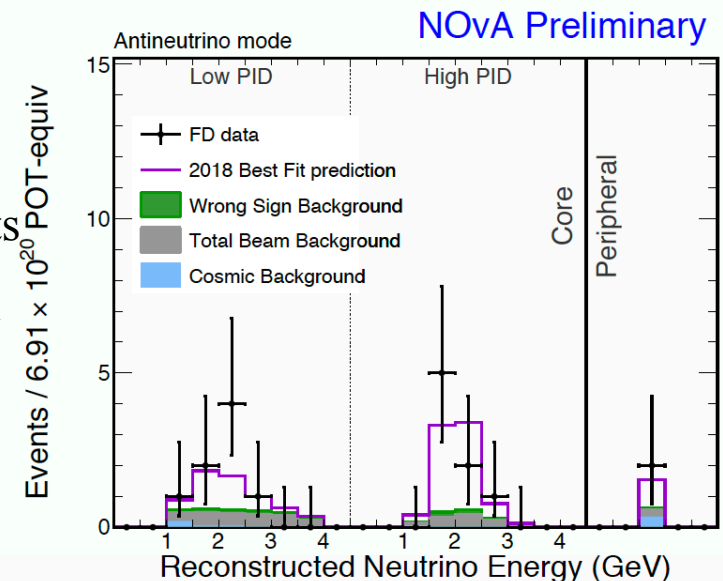
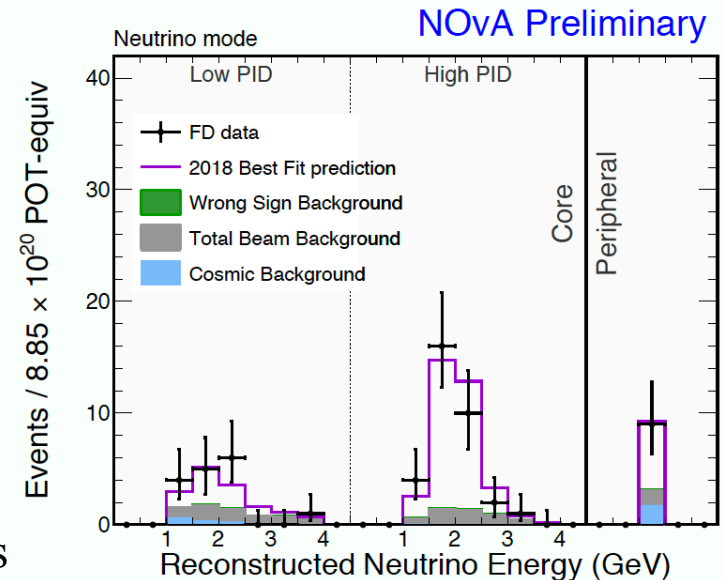
ν_μ Data at Far Detector

- FD selection:
 - Additional BDT to reduce cosmic backgrounds
 - Estimate cosmic background rate from timing sidebands of the NuMI beam triggers and cosmic trigger data
- Neutrino beam:
 - Observe 113 events
 - Expect $730 +38/-49(\text{syst.})$ w/o oscillations
- Antineutrino beam:
 - Observe 65 events
 - Expect $266 +12/-14(\text{syst.})$ w/o oscillations



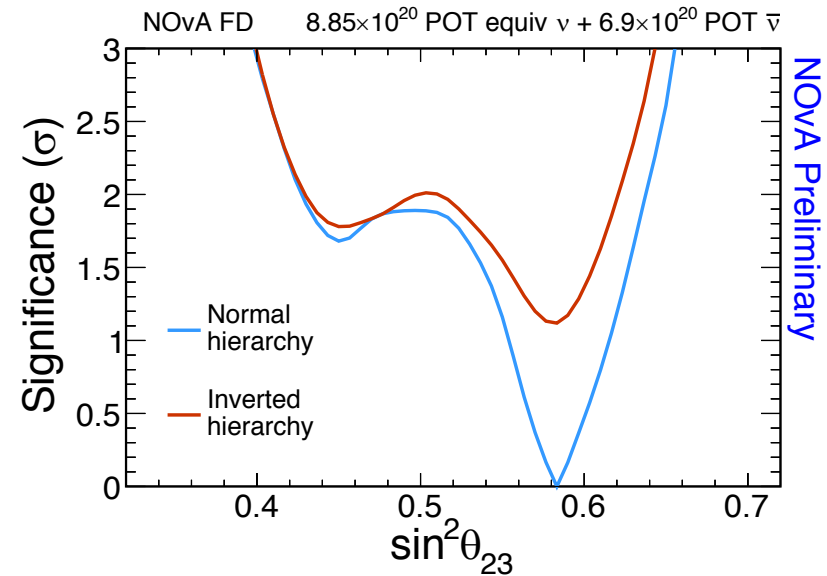
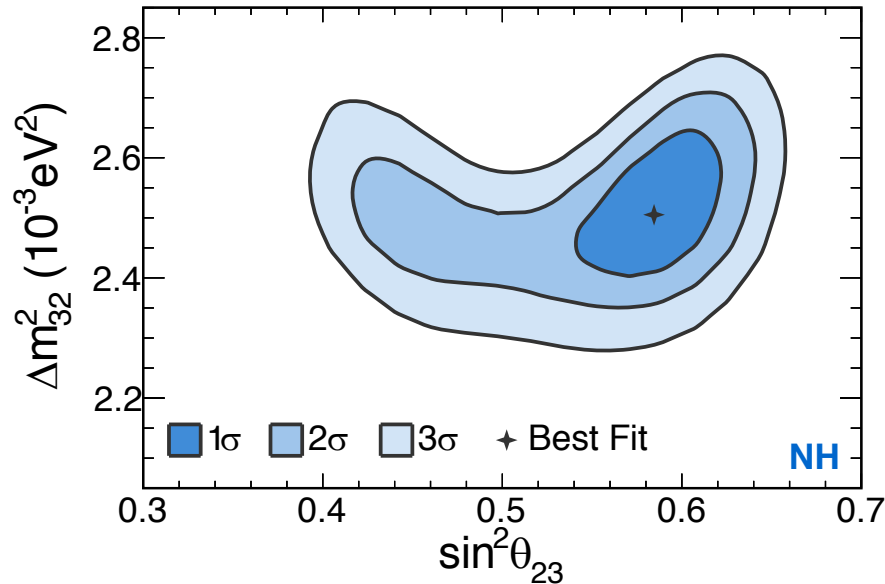
ν_e Data at Far Detector

- FD selection:
 - Add a one-bin peripheral with less stringent containment selection to include more signal
 - Use location dependent BDT and tight PID cuts to recover signal events in this peripheral bin
- Neutrino beam:
 - Observe 58 events, expect 15 background events
 - Background: 11 beam, 3 cosmic and < 1 wrong sign
- Antineutrino beam:
 - Observe 18 events, expect 5.3 background events
 - Background events : 3.5 beam, < 1 cosmic and 1 wrong sign
- $> 4\sigma$ $\bar{\nu}_e$ appearance

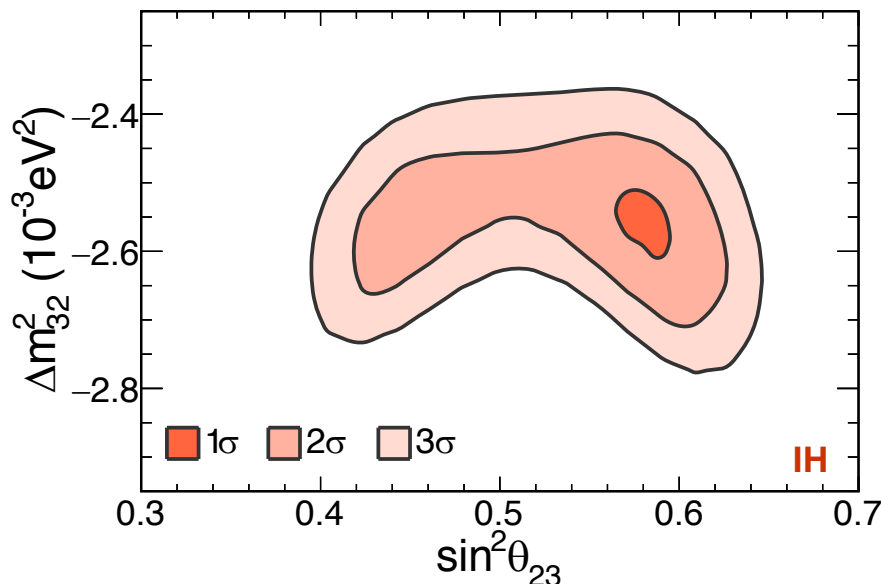


Joint Appearance and Disappearance

NOvA Preliminary



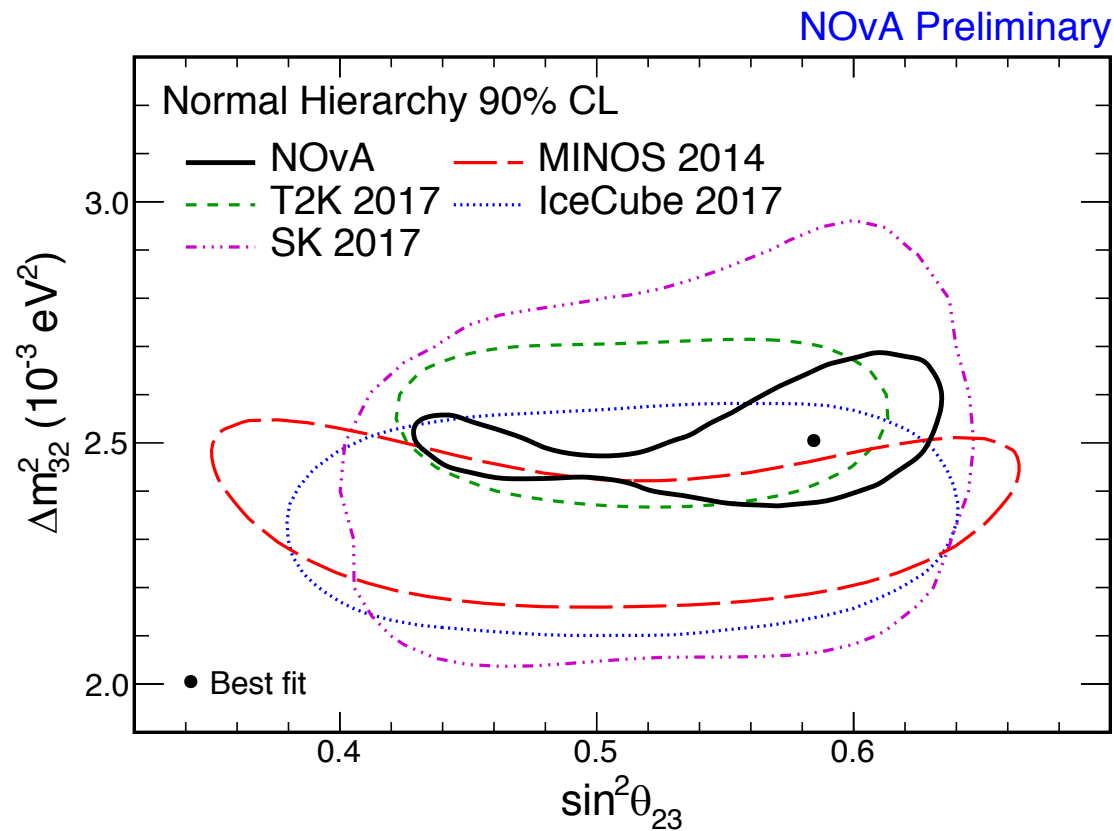
NOvA Preliminary



- Statistically limited, largest systematics for ν_μ and ν_e are calibration and cross-sections.
- Best fit:
 - Normal Hierarchy
 - $\sin^2 \theta_{23} = 0.58 \pm 0.03$ (UO)
 - $\Delta m_{32}^2 = (2.51 + 0.12 - 0.08) \times 10^{-3} \text{eV}^2$
- Prefer Upper Octant

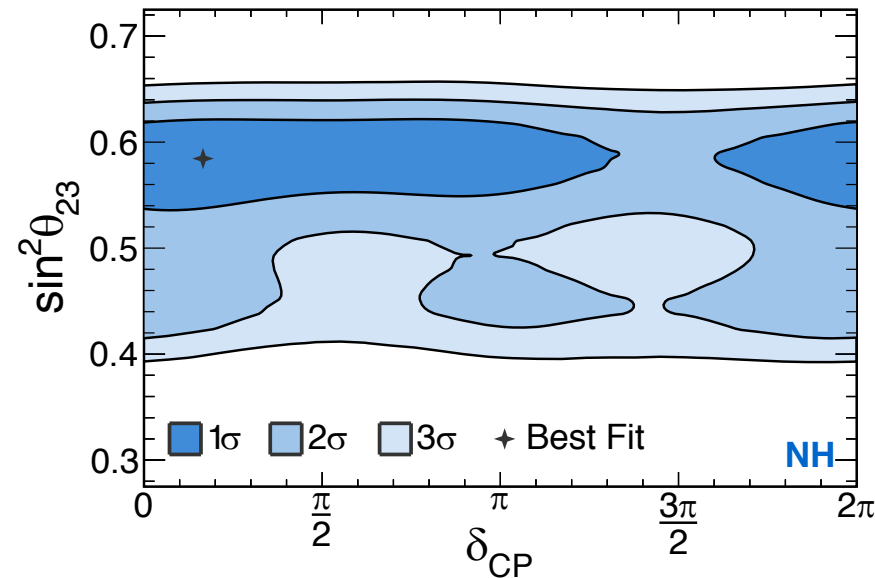
Joint Appearance and Disappearance

NOvA's allowed 90% C.L. regions are compatible to other experiments

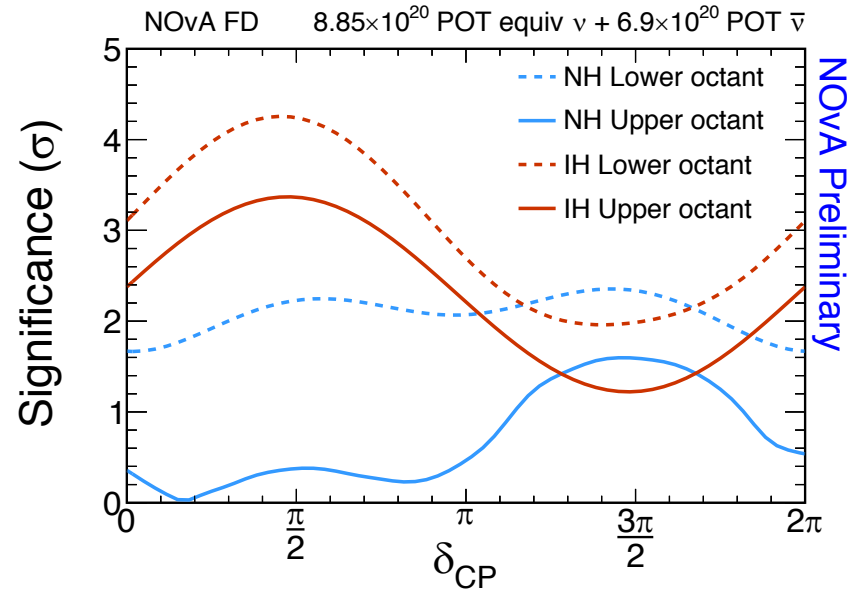
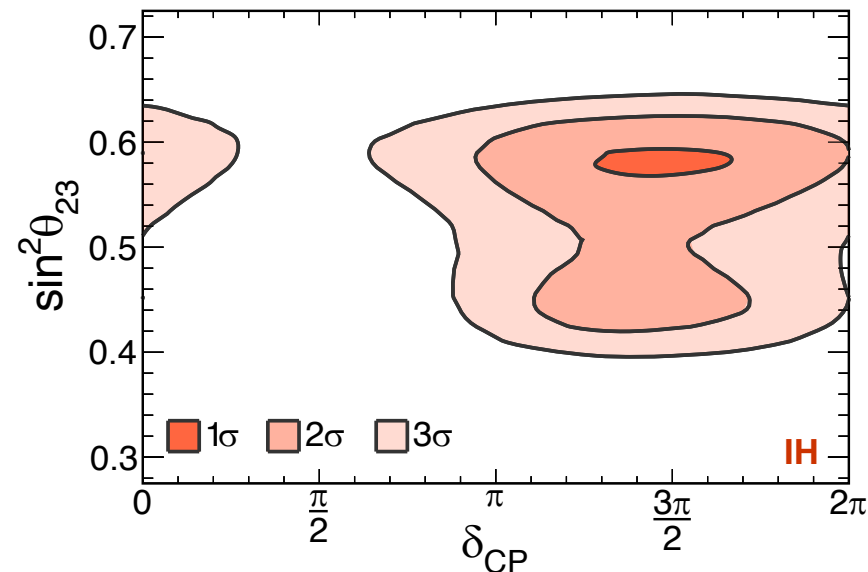


Joint Appearance and Disappearance

NOvA Preliminary



NOvA Preliminary

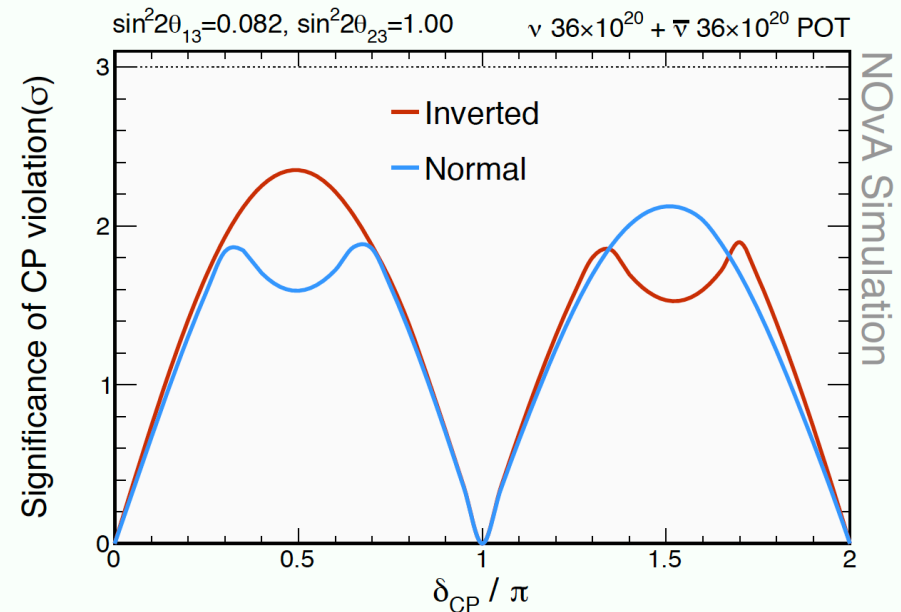
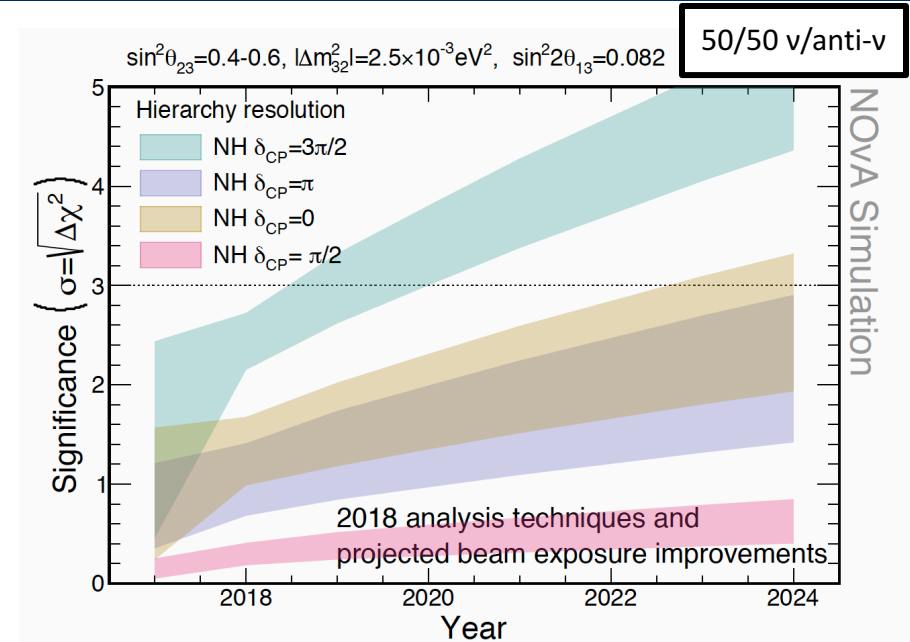


- Statistically limited, largest systematics for ν_μ and ν_e are calibration and cross-sections
- Best fit:
 - Normal Hierarchy
 - $\delta_{CP} = 0.17\pi$
 - $\sin^2 \theta_{23} = 0.58 \pm 0.03$ (UO)
 - $\Delta m_{32}^2 = (2.51 + 0.12 - 0.08) \cdot 10^{-3} \text{ eV}^2$
- Consistent with all δ_{CP} values in NH at $< 1.6\sigma$
- Exclude $\delta = \pi/2$ in IH at $> 3\sigma$
- **Prefer NH at 1.8 σ**

Looking Forward

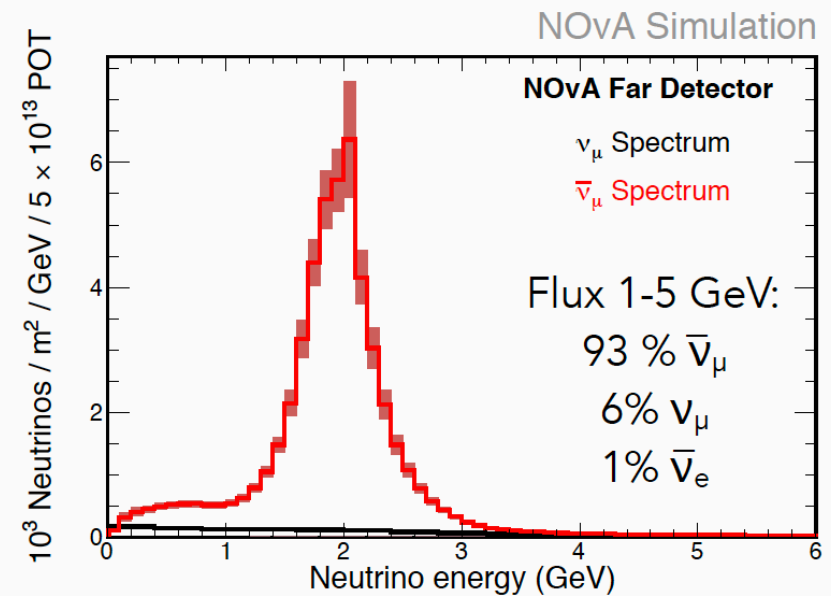
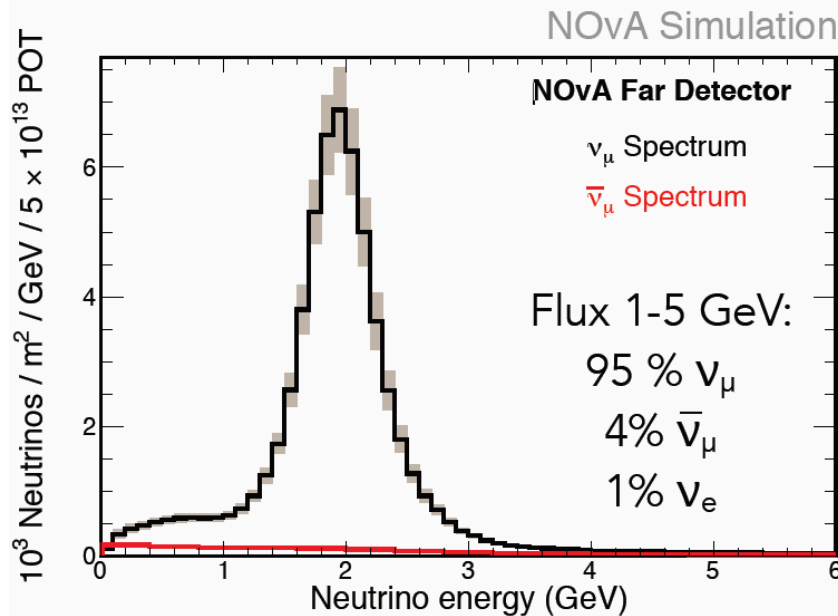
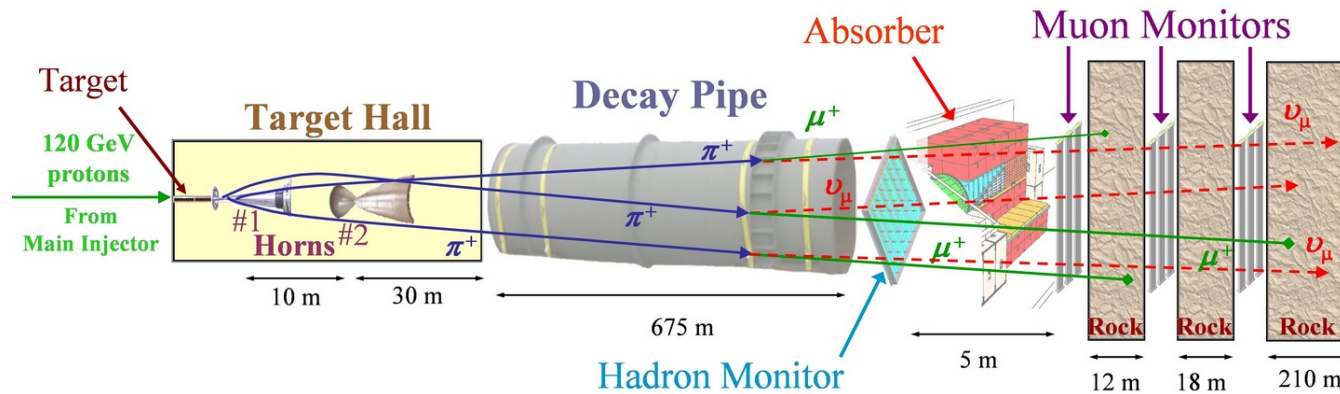
- Taking antineutrino data since 2017, switch back to neutrinos in 2019, run 50% neutrino, 50% anti-neutrino
- Extended running through 2024, test beam program and potential accelerator improvement to enhance ultimate reach
- If $\delta_{CP}=3\pi/2$, 3 σ sensitivity to MH by 2020, $\sim 5 \sigma$ by 2024
- 3 σ to MH for 30-50% (depending on octant) of δ_{CP} range by 2024
- 2+ σ to CP at $\delta_{CP}=3\pi/2$ or $\delta_{CP}=\pi/2$ by 2024

Thank you!

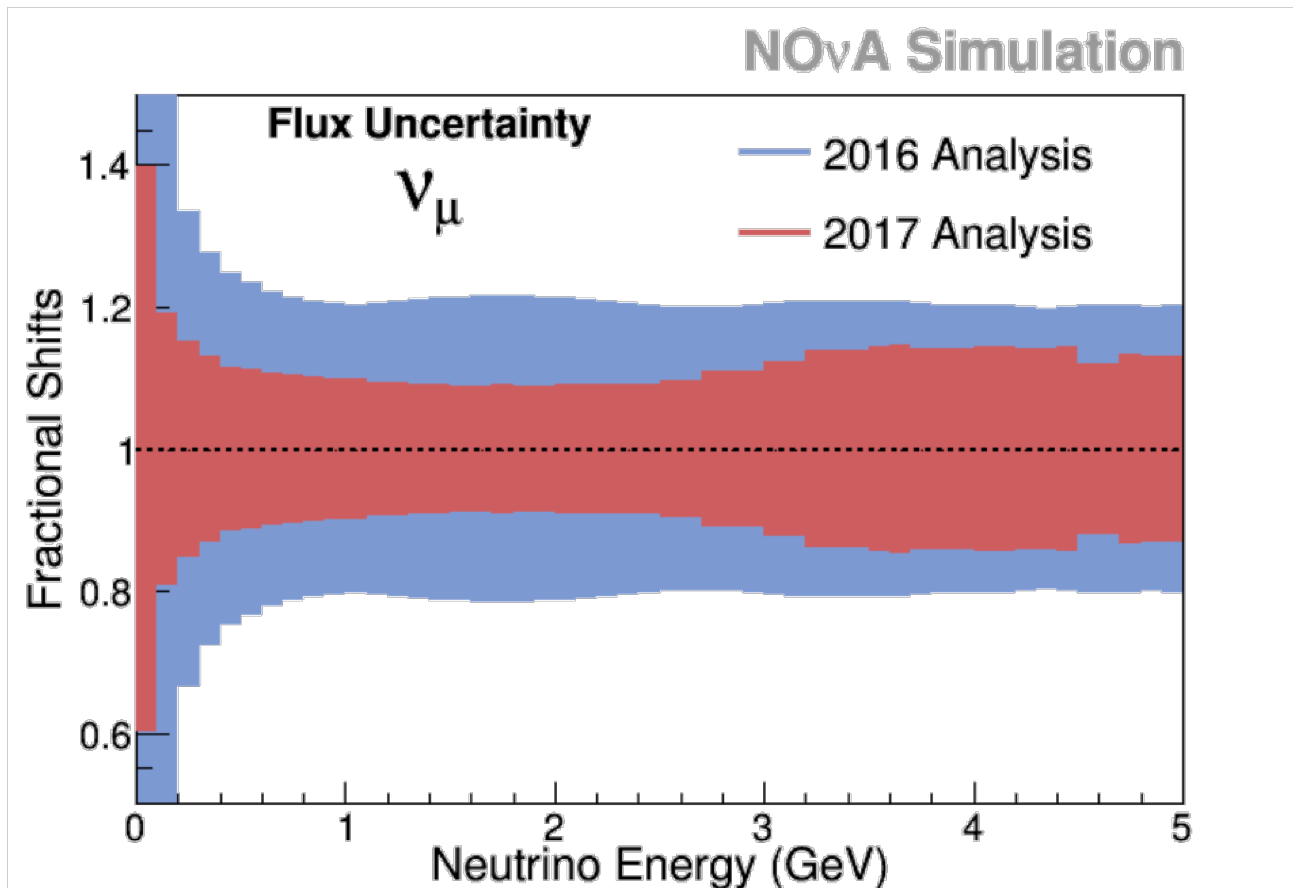


Backup

NuMI Off-Axis ν_e Appearance Experiment



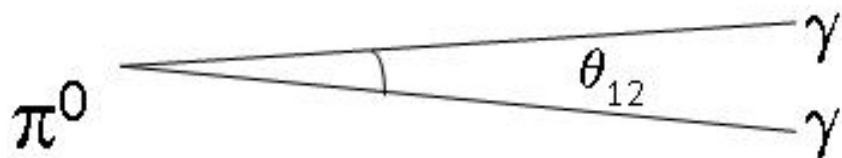
Improved Flux Model



- Package to Predict the Flux (**PPFX**) from MINERvA.
 - Based on thin target hadron production data from NA49 and MIPP.
- Significantly reduced systematic uncertainties.
 - Central values also changed within prior systematics, but not shown here.

Systematic Error in Calibration

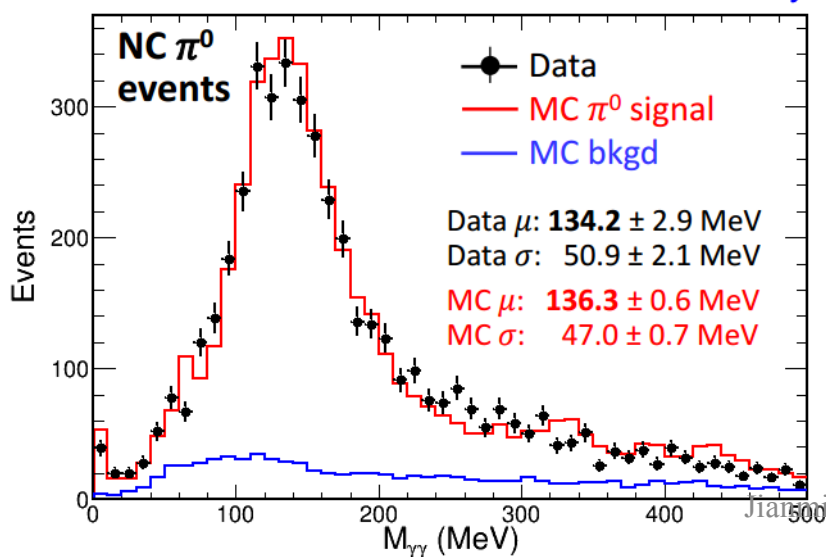
- Our calibration is built on dE/dx from stopping cosmic muons.
- Control samples for calibration uncertainty
 - π^0 mass peak in ND
 - Michel electrons in ND and FD



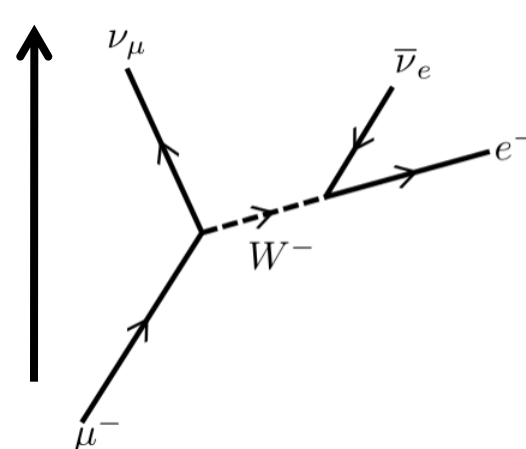
A diagram showing a π^0 particle decaying into two photons (γ). The angle between the two photons is labeled θ_{12} .

$$m_{\pi^0}^2 = 2E_{\gamma 1} E_{\gamma 2} (1 - \cos \theta_{12})$$

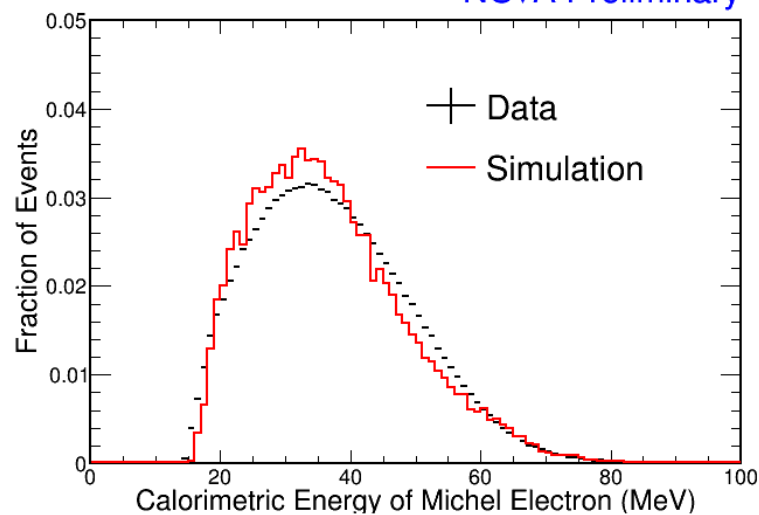
NOvA Preliminary



Michel electrons
from muon decays

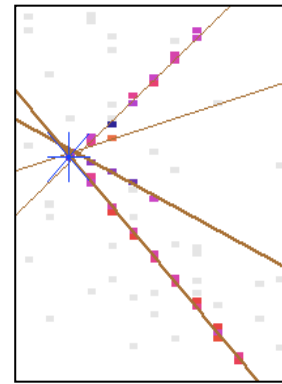
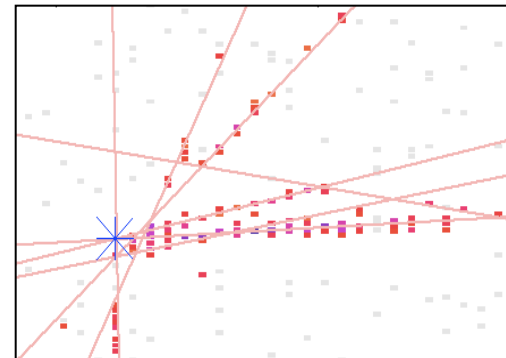
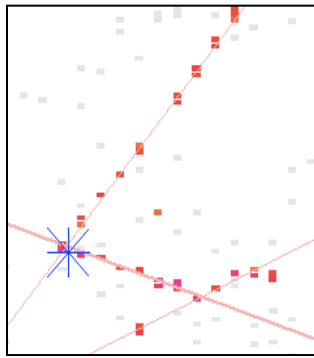


NOvA Preliminary

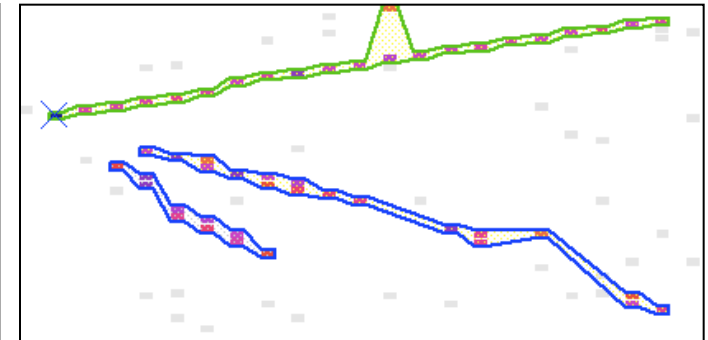
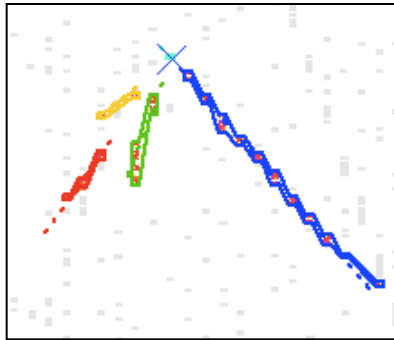


Prong/track Reconstruction

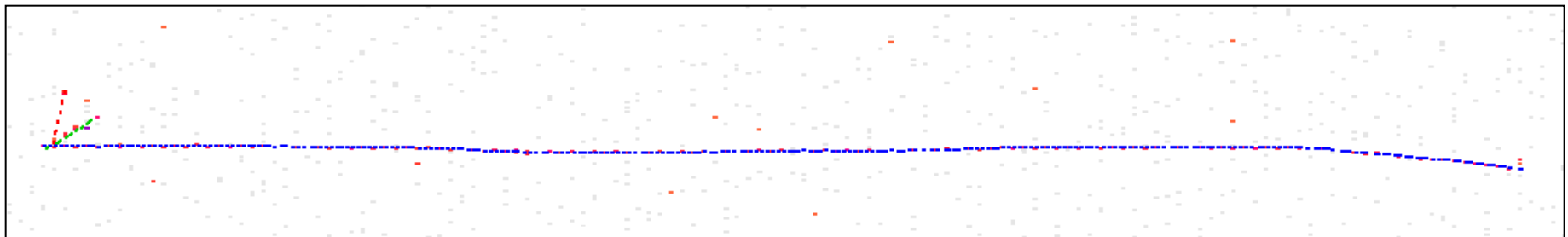
Vertexing: Find lines of energy depositions with Hough transform. Then determine the vertex that all lines converge to



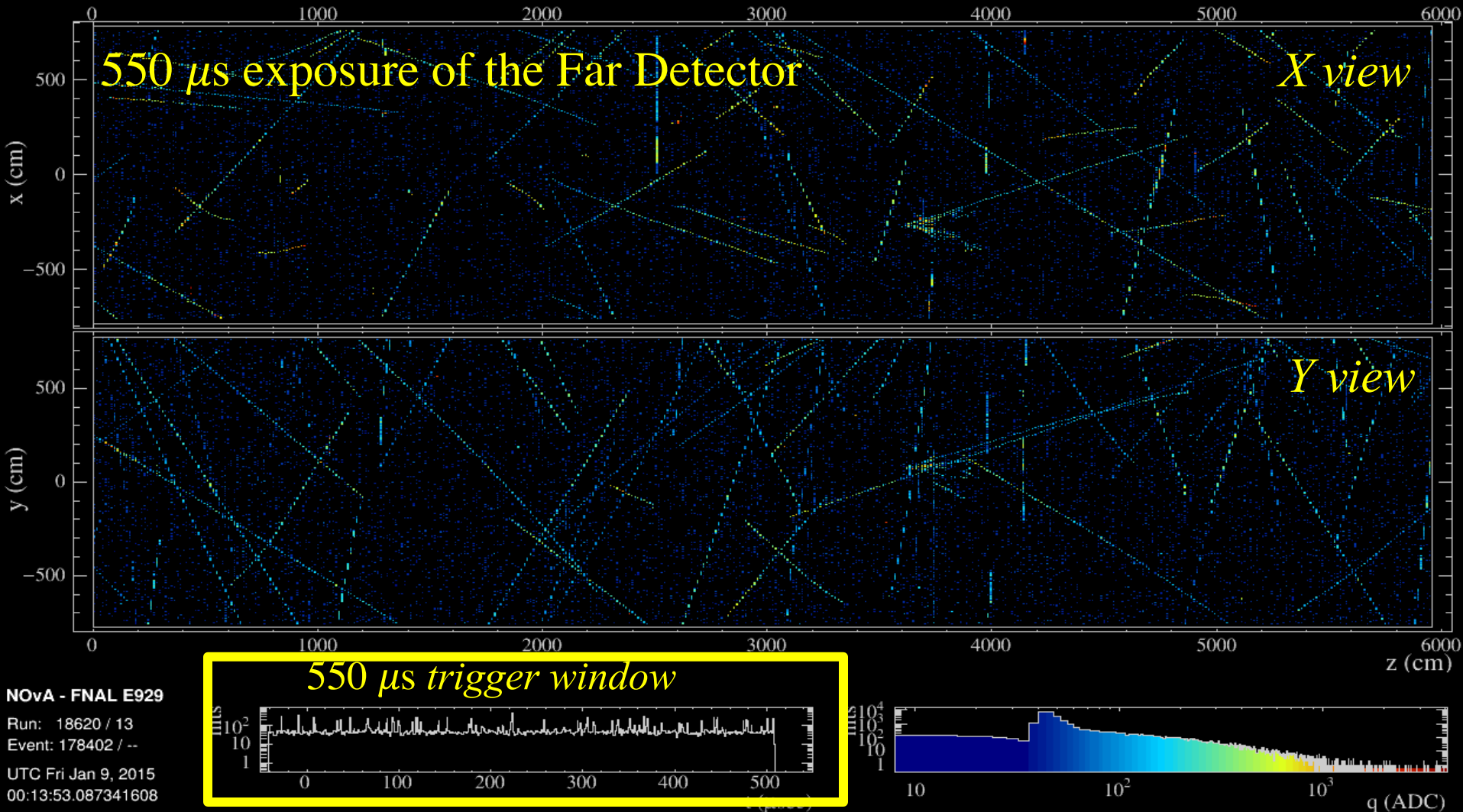
Shower Clustering: Based on the vertex and the lines, showers are reconstructed by angular clustering



Tracking: Trace particle trajectories with **Kalman filter** tracker (below). Also have a **cosmic ray tracker** that reconstructs cosmic tracks with high speed.

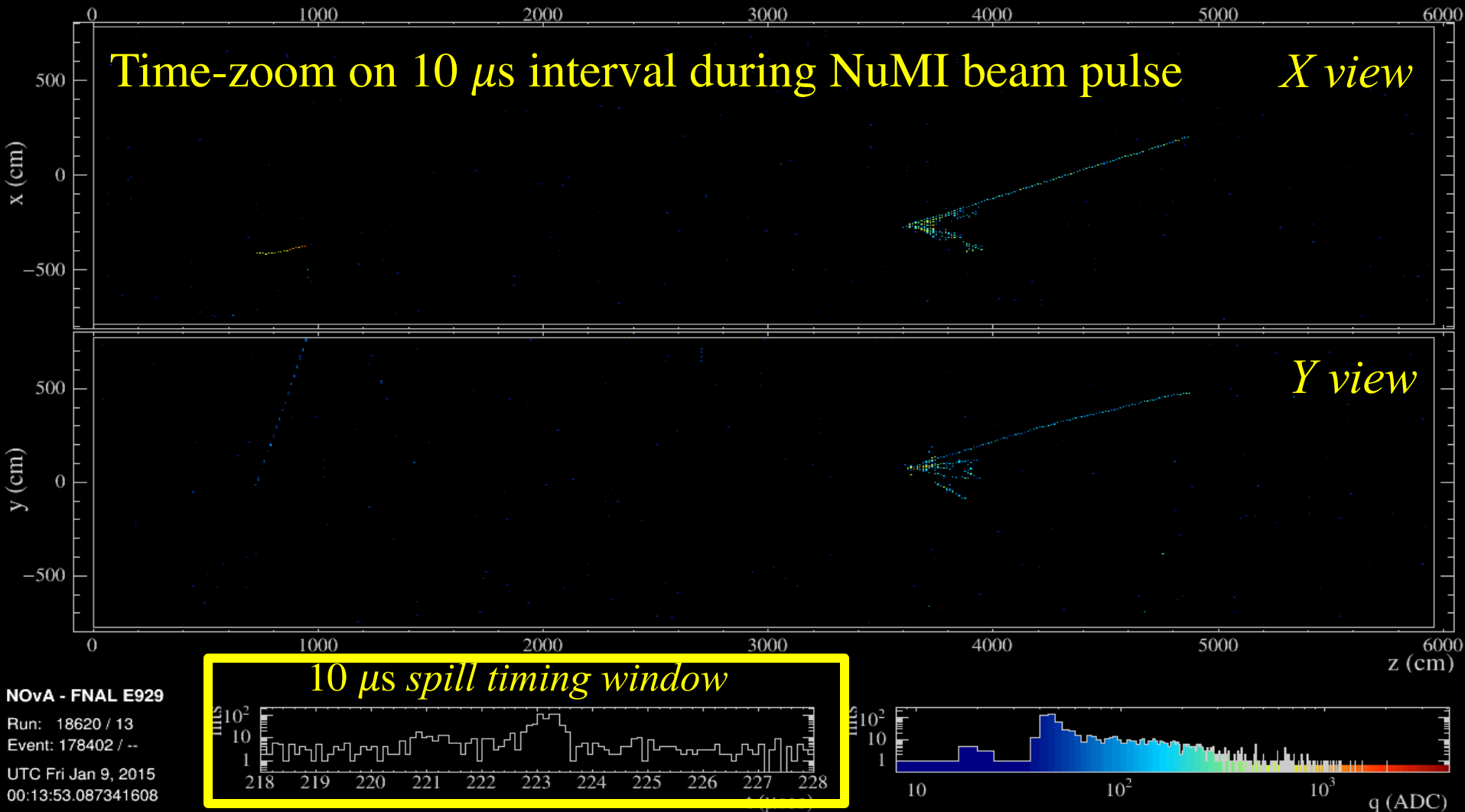


Event clustering



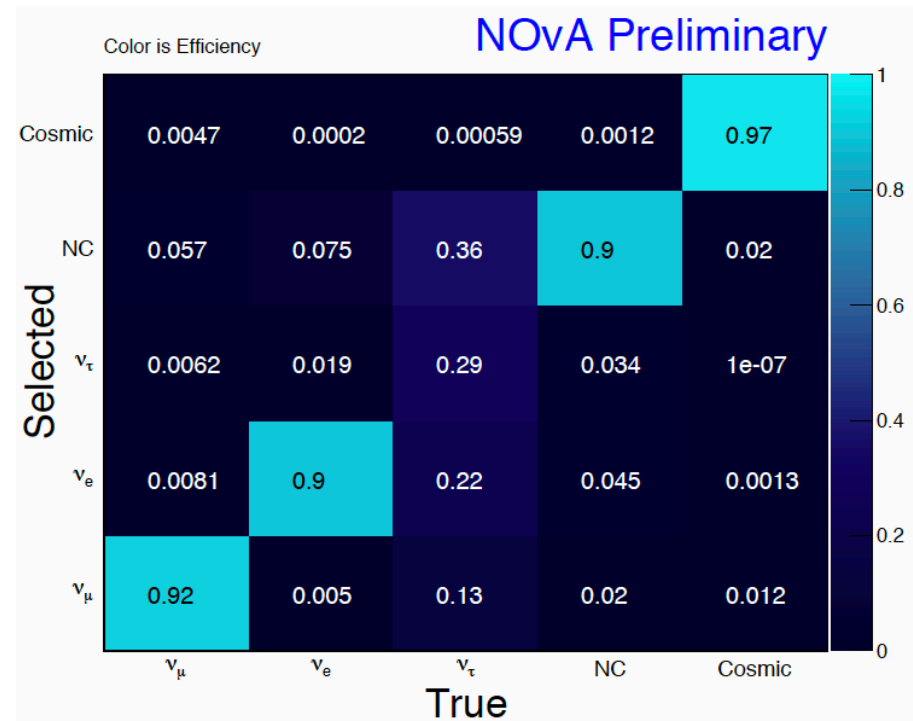
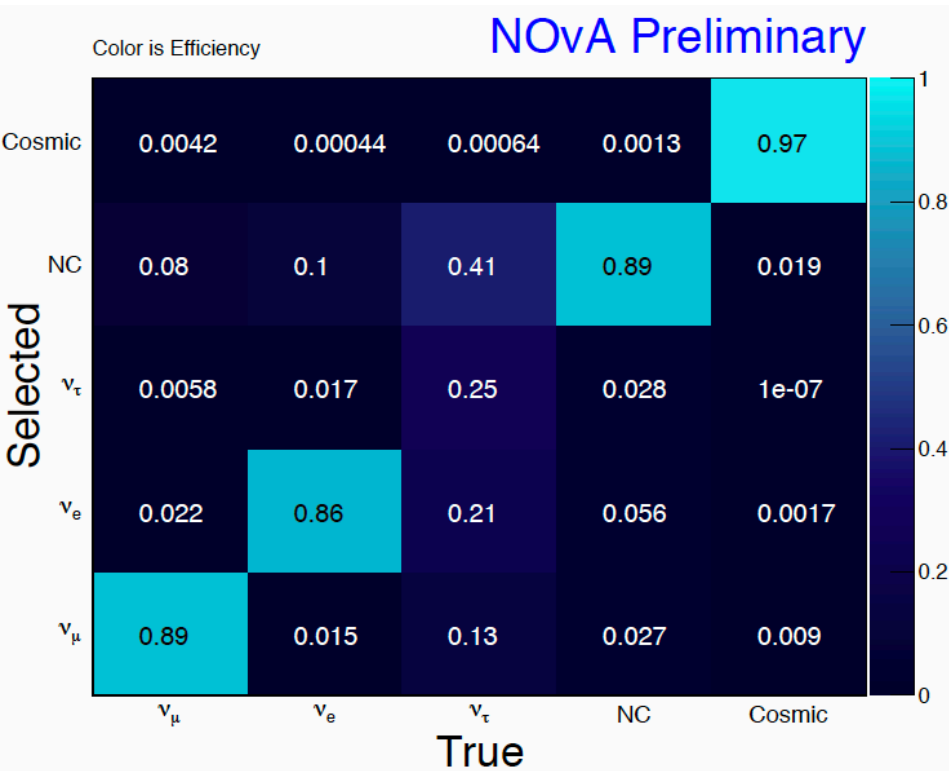
- Because NOvA is on surface, hits in a trigger window are a combination of cosmic and beam events.
- First step in reconstruction is to cluster hits by space-time coincidence to separate neutrino hits and cosmic hits.

Event clustering



Event clusters that contain neutrino interactions can be correctly selected in the neutrino spill timing window

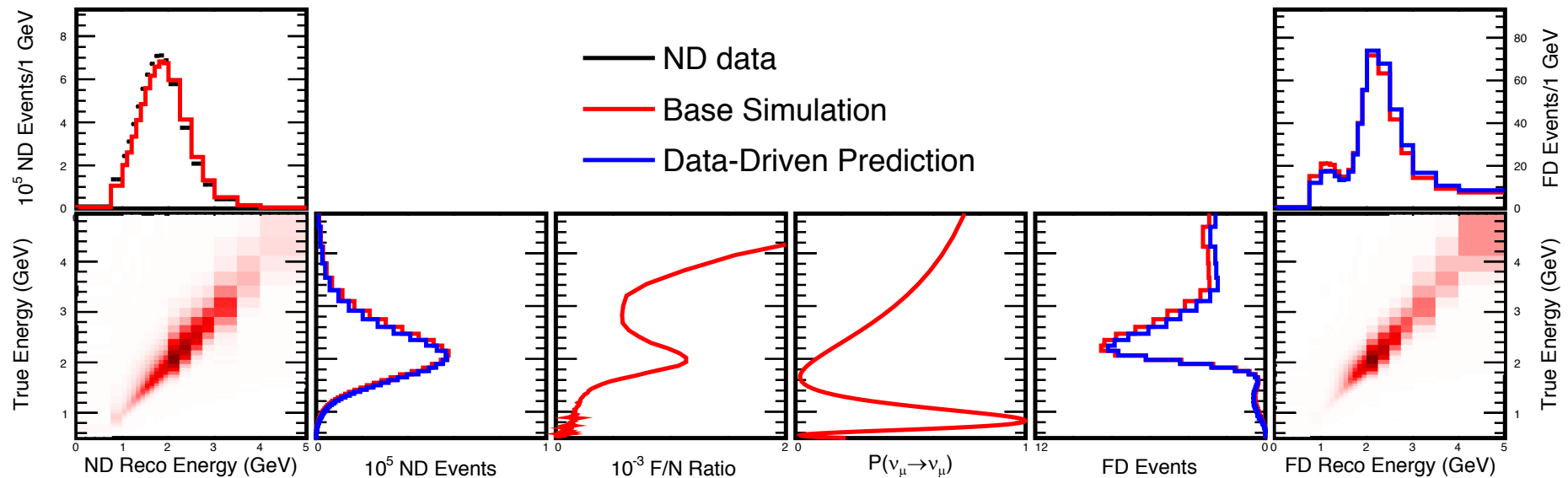
PID efficiencies



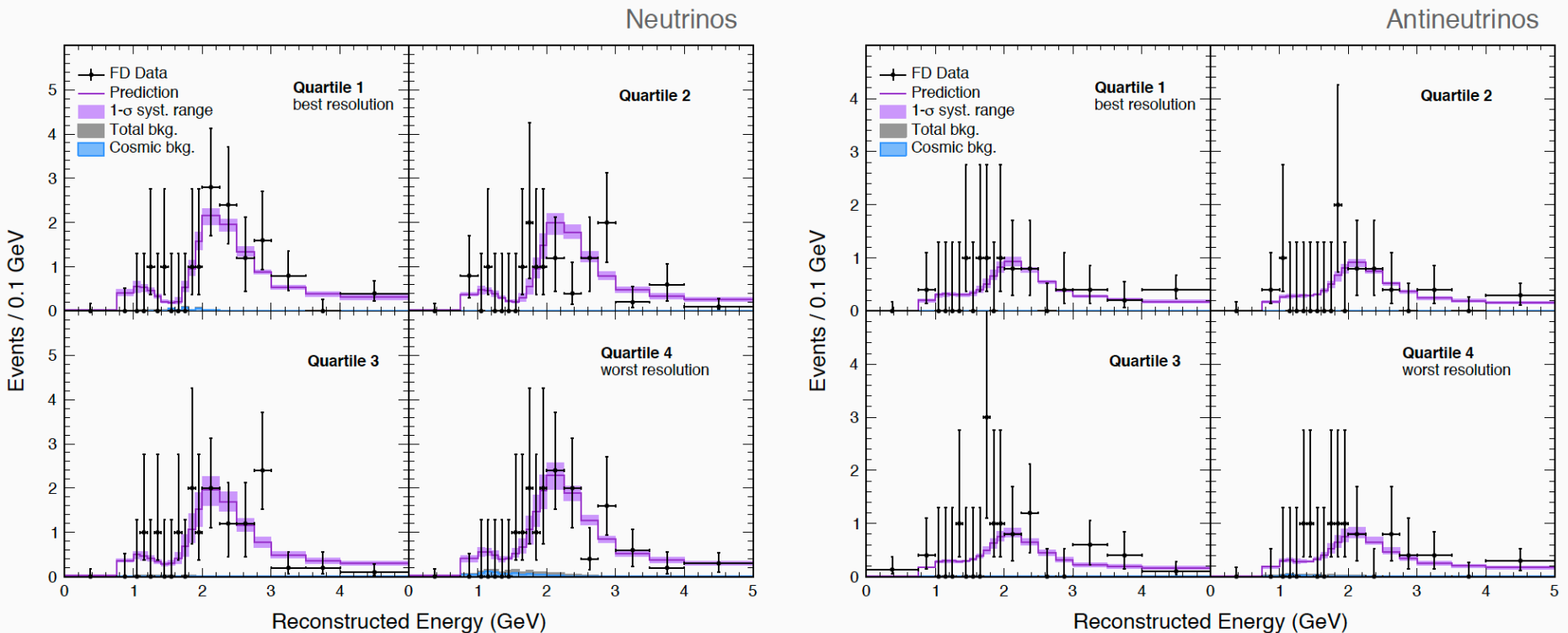
Analysis Strategy

- Separate ν_μ/ν_e CC interactions from beam backgrounds
- Extrapolate observed ND spectrum to FD, reject cosmic rays in FD, make FD unoscillated prediction
- Measure shapes and yields of ν_μ/ν_e in energy/PID bins in the FD to determine oscillation parameters

ND \rightarrow FD extrapolation for ν_μ disappearance

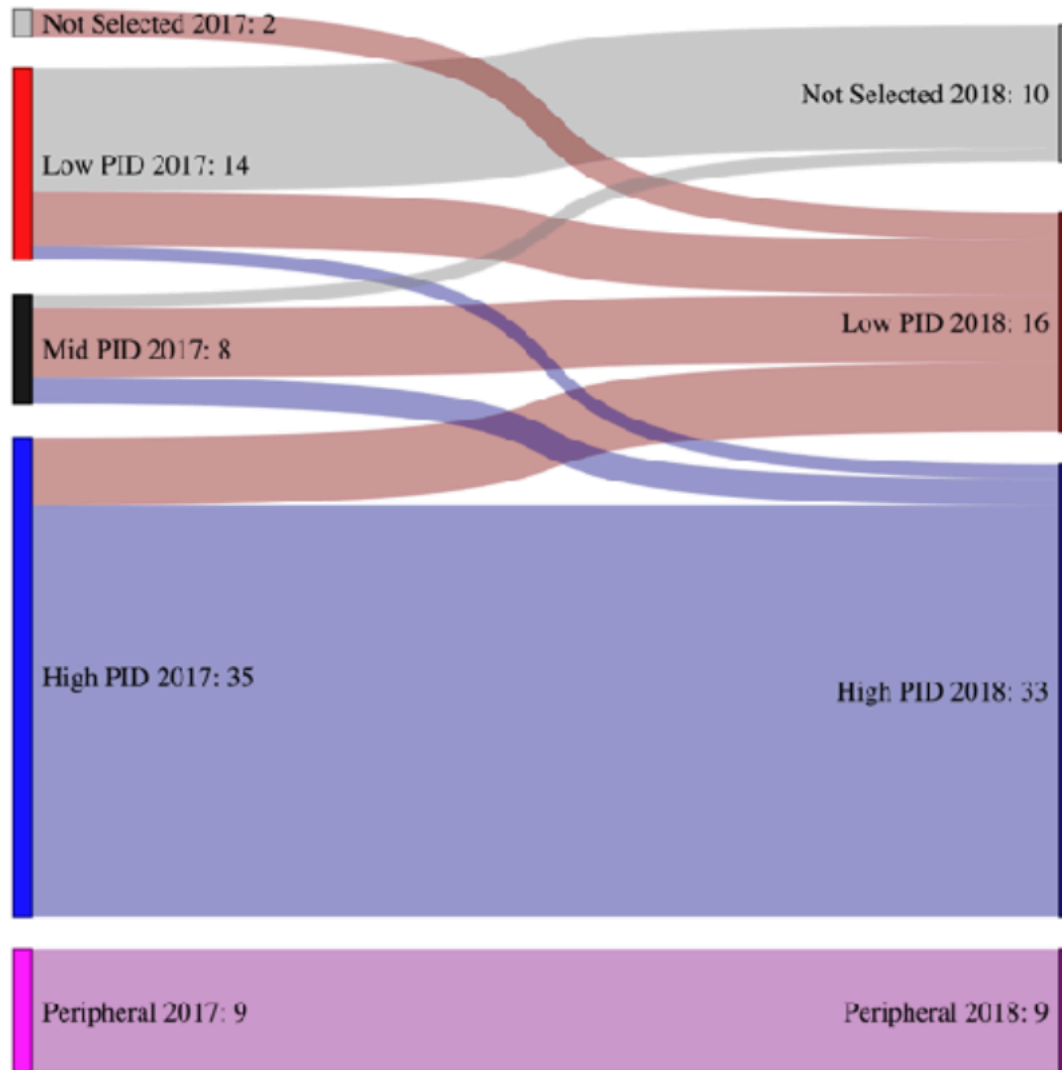


FD ν_μ Events in 4 Quartiles



2017/2018 RHC ν_e FD Data

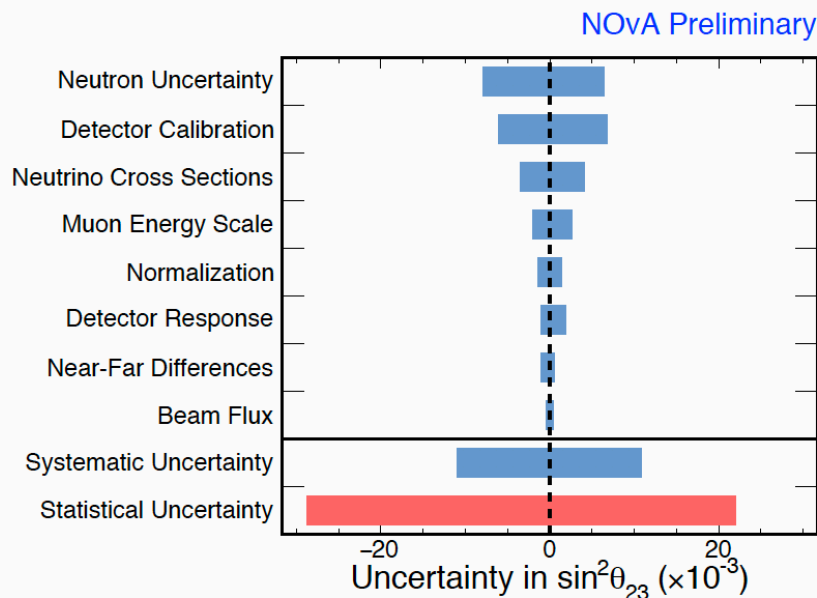
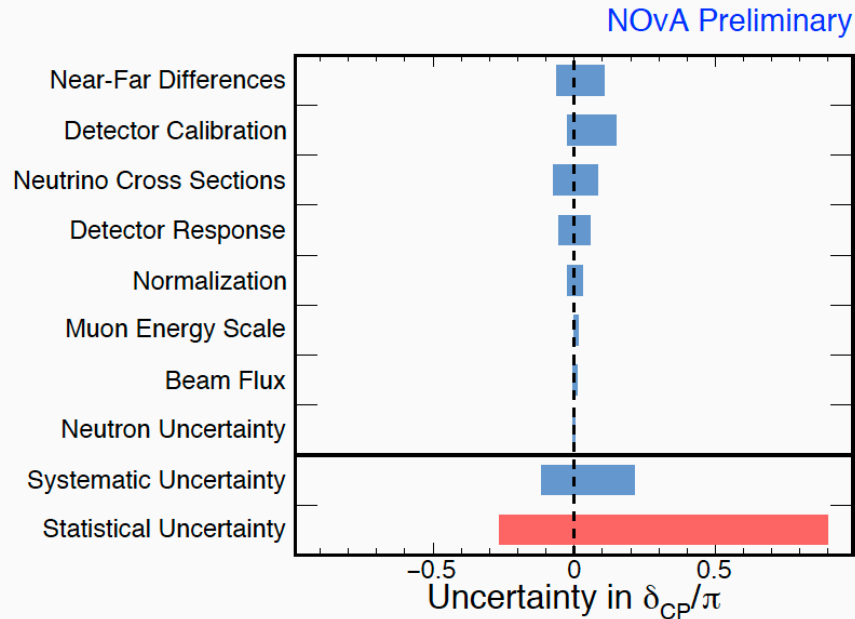
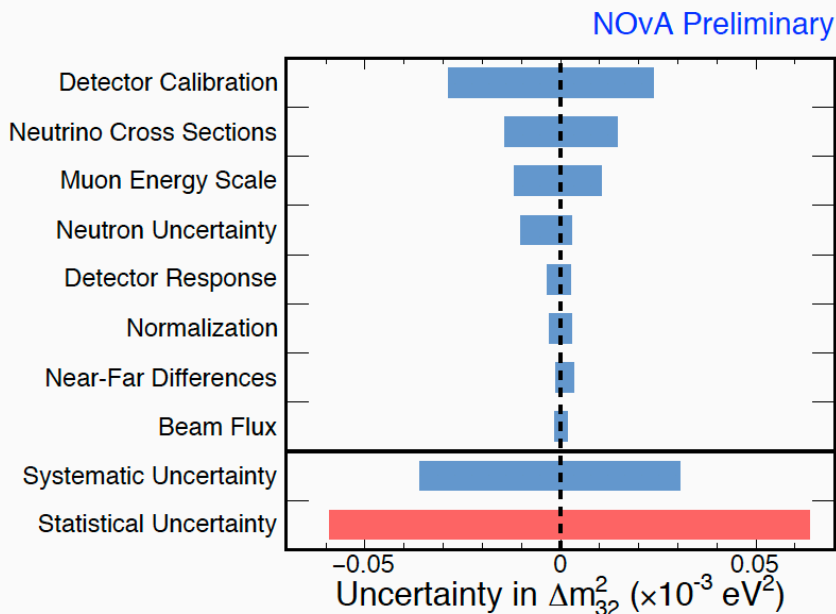
66 FD data
events in 2017
analysis



58 FD data
events in 2017
analysis

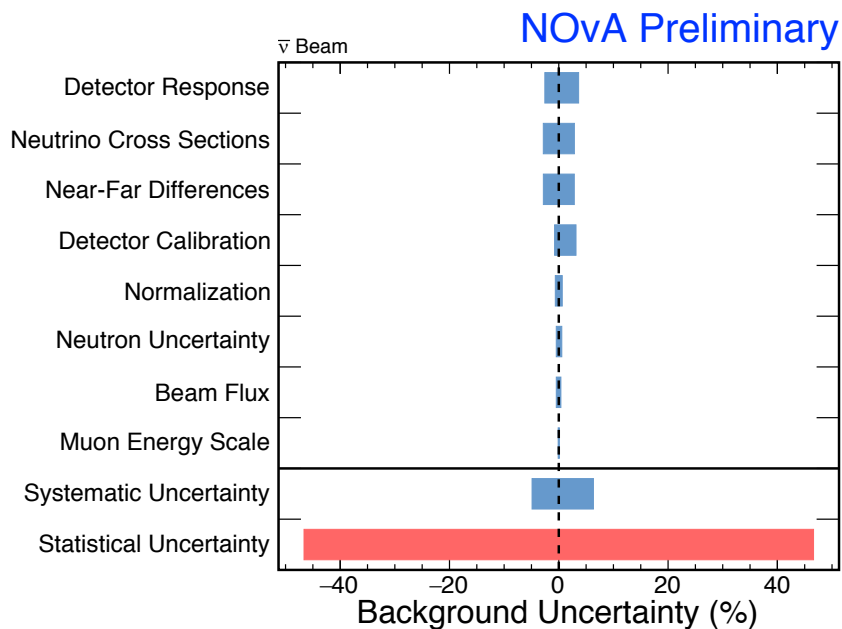
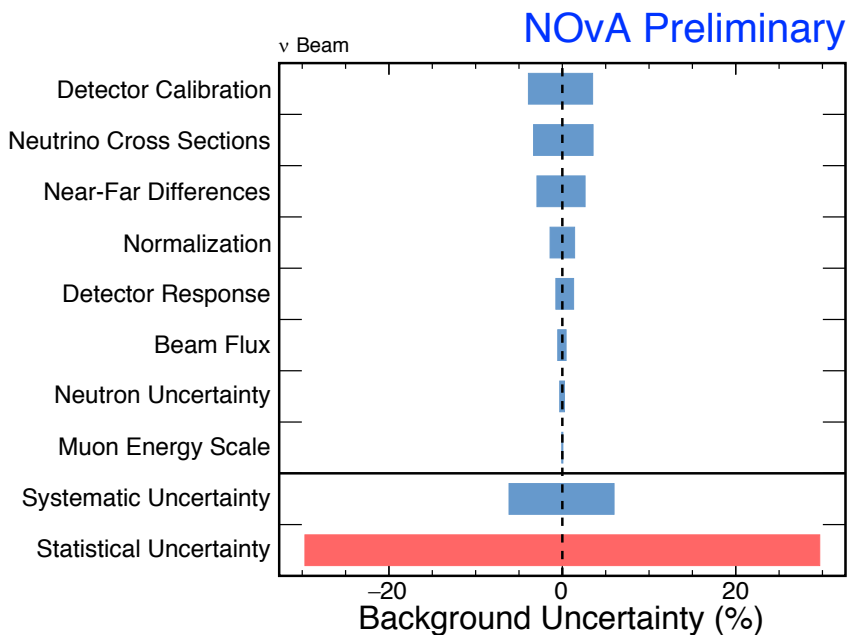
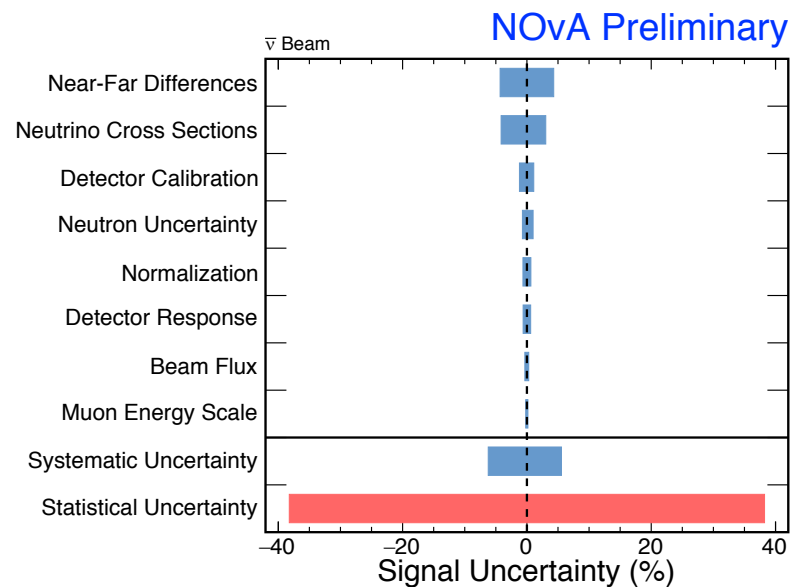
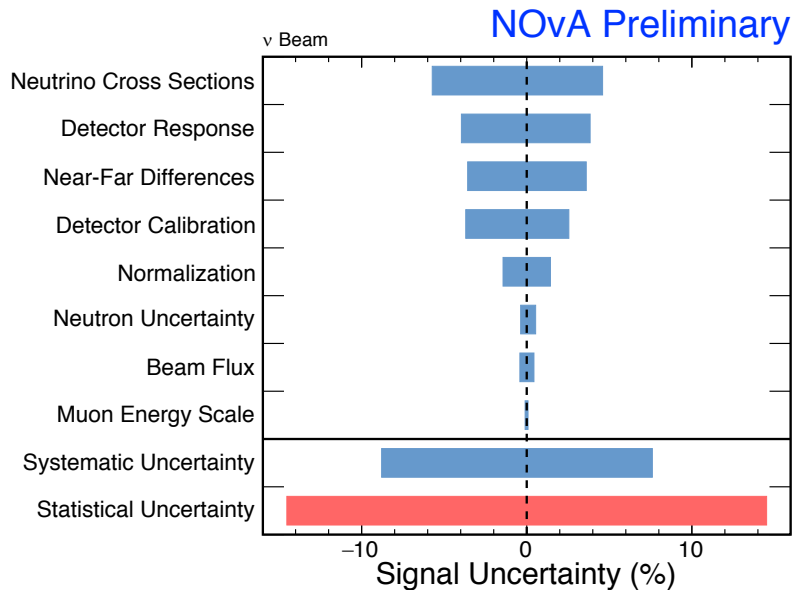
Change in data events after retraining of PID, new training improved bkg rejection

Systematic Uncertainties (ν_μ)



- Largest systematics for ν_μ and ν_e are **calibration** and **cross-sections**.
- Both analyses are statistically limited.

Systematic Uncertainties (ν_e)

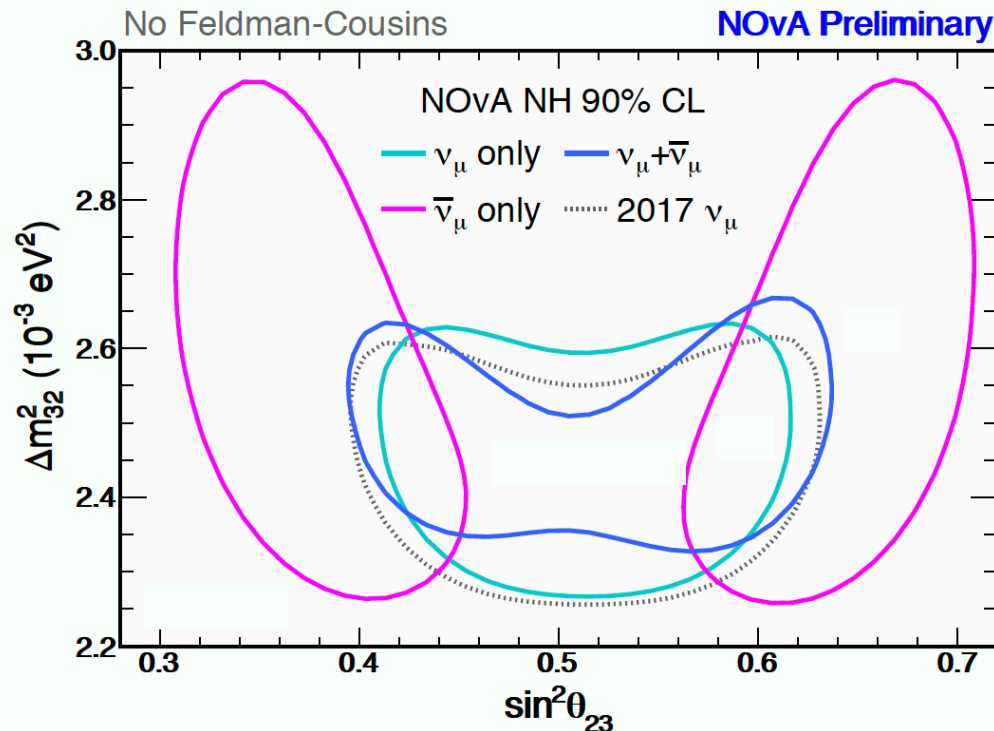


Systematic Uncertainties (Joint Fit)

| Source of Uncertainty | $\sin^2\theta_{23} (\times 10^{-3})$ | δ_{CP}/π | $\Delta m_{32}^2 (\times 10^{-3} \text{ eV}^2)$ |
|-------------------------|--------------------------------------|-------------------|---|
| Beam Flux | +0.42 / -0.48 | +0.0088 / -0.0048 | +0.0016 / -0.0015 |
| Detector Calibration | +6.9 / -6.1 | +0.15 / -0.023 | +0.024 / -0.029 |
| Detector Response | +1.9 / -0.99 | +0.055 / -0.054 | +0.0027 / -0.0034 |
| Muon Energy Scale | +2.6 / -2.1 | +0.015 / -0.0026 | +0.01 / -0.012 |
| Near-Far Differences | +0.56 / -1.1 | +0.11 / -0.064 | +0.0033 / -0.0013 |
| Neutrino Cross Sections | +4.2 / -3.5 | +0.085 / -0.072 | +0.015 / -0.014 |
| Neutron Uncertainty | +6.4 / -7.9 | +0.002 / -0.0052 | +0.0028 / -0.01 |
| Normalization | +1.4 / -1.5 | +0.031 / -0.024 | +0.0029 / -0.0027 |
| Systematic Uncertainty | +9.6 / -11 | +0.21 / -0.11 | +0.032 / -0.035 |
| Statistical Uncertainty | +22 / -29 | +0.9 / -0.27 | +0.064 / -0.059 |

ν_μ appearance fit

- The combined data of neutrino and antineutrino beams are fitted assuming CPT invariance.
- We observe 113 events and expect 126 at this combined best fit for the neutrino beam mode and observe 65 events and expect 52 at the best fit in antineutrino beam mode.
- If fit separately, the antineutrino beam mode prefers a more non-maximal solution than the neutrino beam mode. However the χ^2 s are consistent with the combined fit oscillation parameters with $p > 4\%$.



Looking Forward

