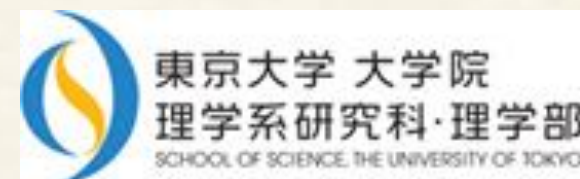




# Long baseline neutrino experiments (accelerator LBL & reactor)

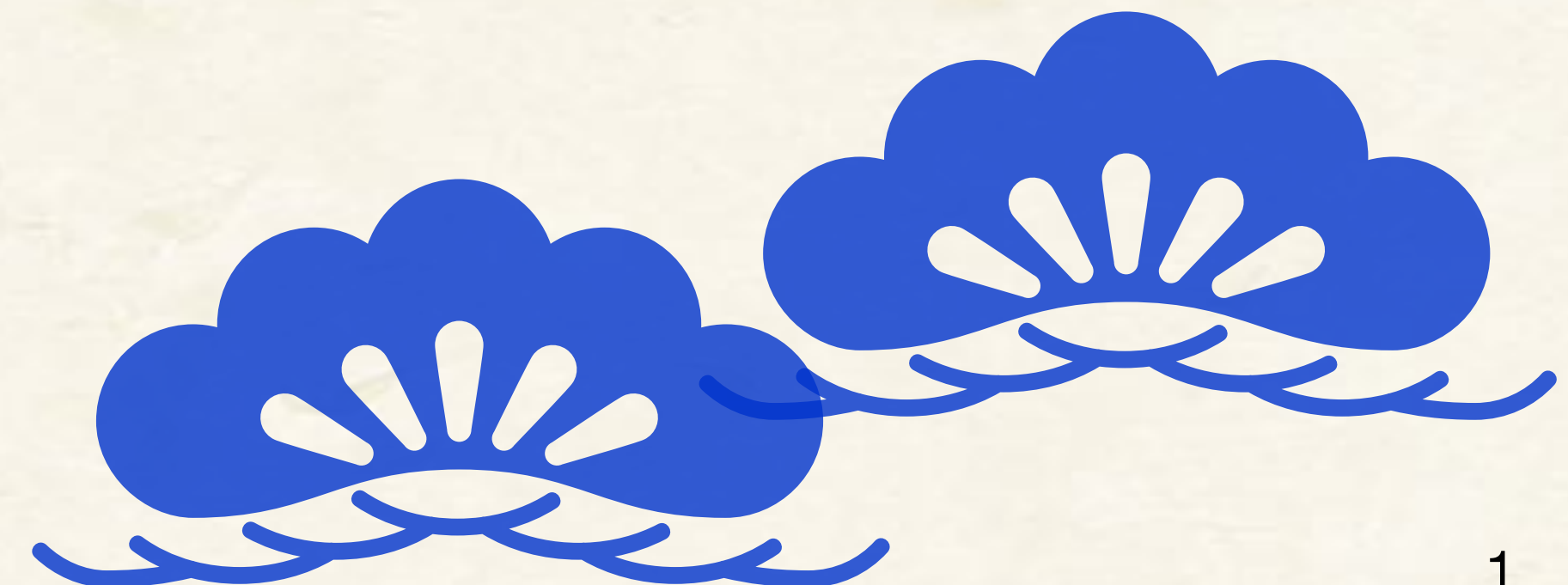
Masashi Yokoyama



Department of Physics, The University of Tokyo



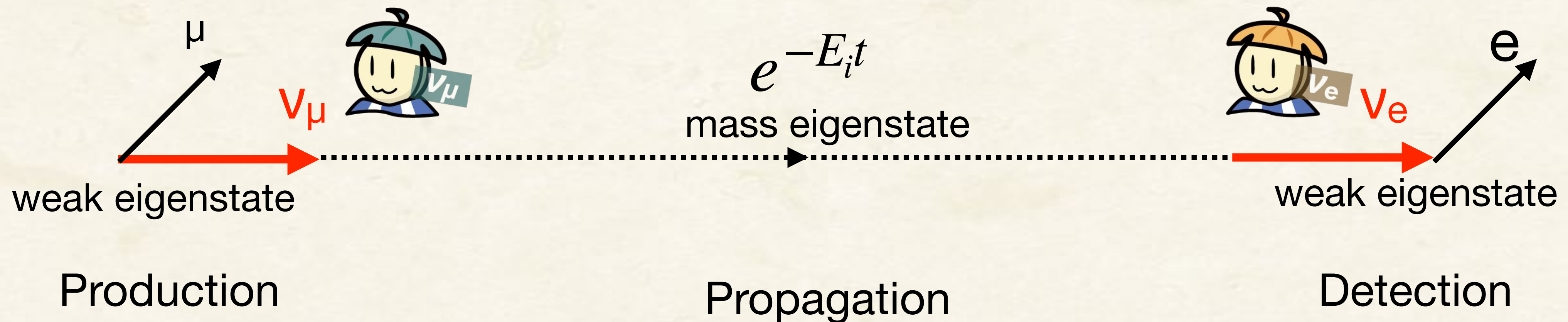
ICHEP2018, Seoul  
July 9, 2018



# Neutrino oscillations

$$|\nu_\alpha\rangle = \sum_{\text{Mass eigenstates } i} U_{\alpha i}^* |\nu_i\rangle$$

Weak eigenstates                      Mass eigenstates



Flavor change during flight

$\Delta m^2$ : mass-squared difference

$$P = \sin^2 2\theta \sin^2 \frac{\Delta m^2 L}{4E} \quad (\text{for 2 flavor})$$

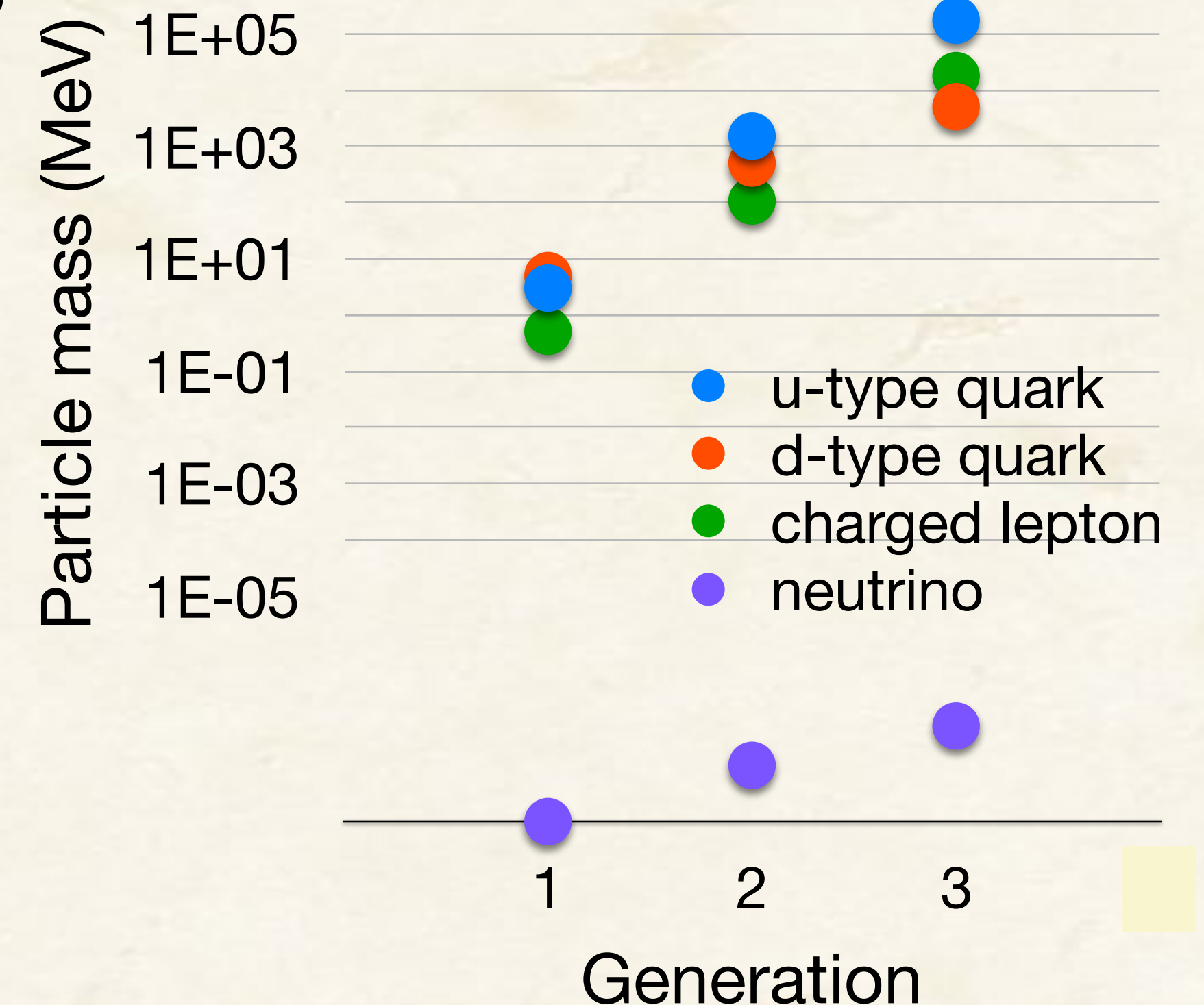
$\theta$ : mixing angle

Quantum effect over very macroscopic length!

# 20 years since its discovery...

We learned a lot about neutrinos through **neutrino oscillation**, but many questions emerged and remains

- ◆ Origin of **tiny mass**
  - ◆ Why mass is much smaller than other fermions?
- ◆ **Large mixing** parameters
  - ◆ Why so different from quarks?
  - ◆ Symmetry behind the pattern?
- ◆ **Mass hierarchy (ordering)**
  - ◆ Which is the heaviest?
- ◆ **CP** violation
  - ◆ Is it violated just as in quarks?
- ◆ Extra neutrino **families**?



Properties of neutrino are considered to be connected with fundamental questions

- ◆ Source of baryon asymmetry of Universe?
- ◆ Very high scale physics? (seesaw?)
- ◆ Origin of generations?

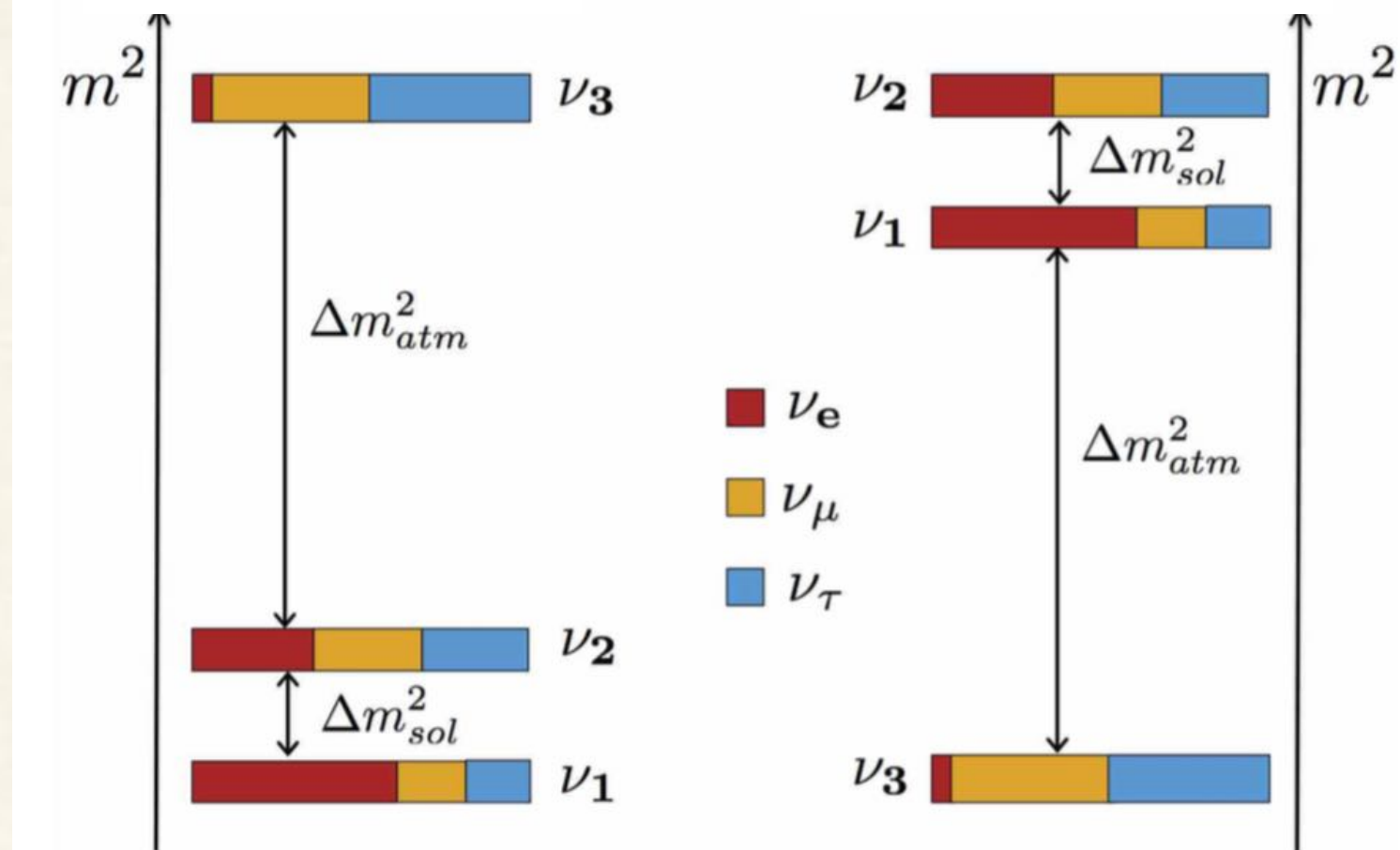
# Oscillation parameter status

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-\delta_{CP}} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{\delta_{CP}} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

M. Tórtola @ NEUTRINO2018

parameter	best fit $\pm 1\sigma$	$3\sigma$ range	
$\Delta m_{21}^2$ [ $10^{-5} \text{eV}^2$ ]	$7.55^{+0.20}_{-0.16}$	7.05–8.14	2.4%
$ \Delta m_{31}^2 $ [ $10^{-3} \text{eV}^2$ ] (NO)	$2.50 \pm 0.03$	2.41–2.60	1.3%
$ \Delta m_{31}^2 $ [ $10^{-3} \text{eV}^2$ ] (IO)	$2.42^{+0.03}_{-0.04}$	2.31–2.51	1.3%
$\sin^2 \theta_{12} / 10^{-1}$	$3.20^{+0.20}_{-0.16}$	2.73–3.79	5.5%
$\sin^2 \theta_{23} / 10^{-1}$ (NO)	$5.47^{+0.20}_{-0.30}$	4.45–5.99	4.7%
$\sin^2 \theta_{23} / 10^{-1}$ (IO)	$5.51^{+0.18}_{-0.30}$	4.53–5.98	4.4%
$\sin^2 \theta_{13} / 10^{-2}$ (NO)	$2.160^{+0.083}_{-0.069}$	1.96–2.41	3.5%
$\sin^2 \theta_{13} / 10^{-2}$ (IO)	$2.220^{+0.074}_{-0.076}$	1.99–2.44	3.5%
$\delta/\pi$ (NO)	$1.32^{+0.21}_{-0.15}$	0.87–1.94	10%
$\delta/\pi$ (IO)	$1.56^{+0.13}_{-0.15}$	1.12–1.94	9%

relative  $1\sigma$  uncertainty



## Current major targets

- More precision measurements
- CP violation
- Mass hierarchy
- $\theta_{23}$  octant ( $\Leftrightarrow 45^\circ$ ?)

<https://globalfit.astroparticles.es/>

deSalas et al, 1708.01186 (May 2018)

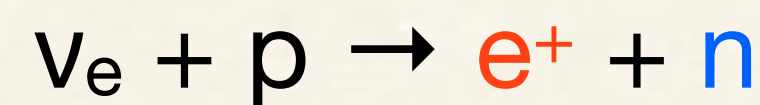
# Reactor $\theta_{13}$ measurement

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta_{13} \sin^2 \left[ \frac{\Delta m_{ee}^2 L}{4E} \right] - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \left[ \frac{\Delta m_{21}^2 L}{4E} \right]$$

..... negligible at ~1 km

At ~km with reactor  $\nu_e$  energy, almost pure  $\sin^2 2\theta_{13}$  measurement

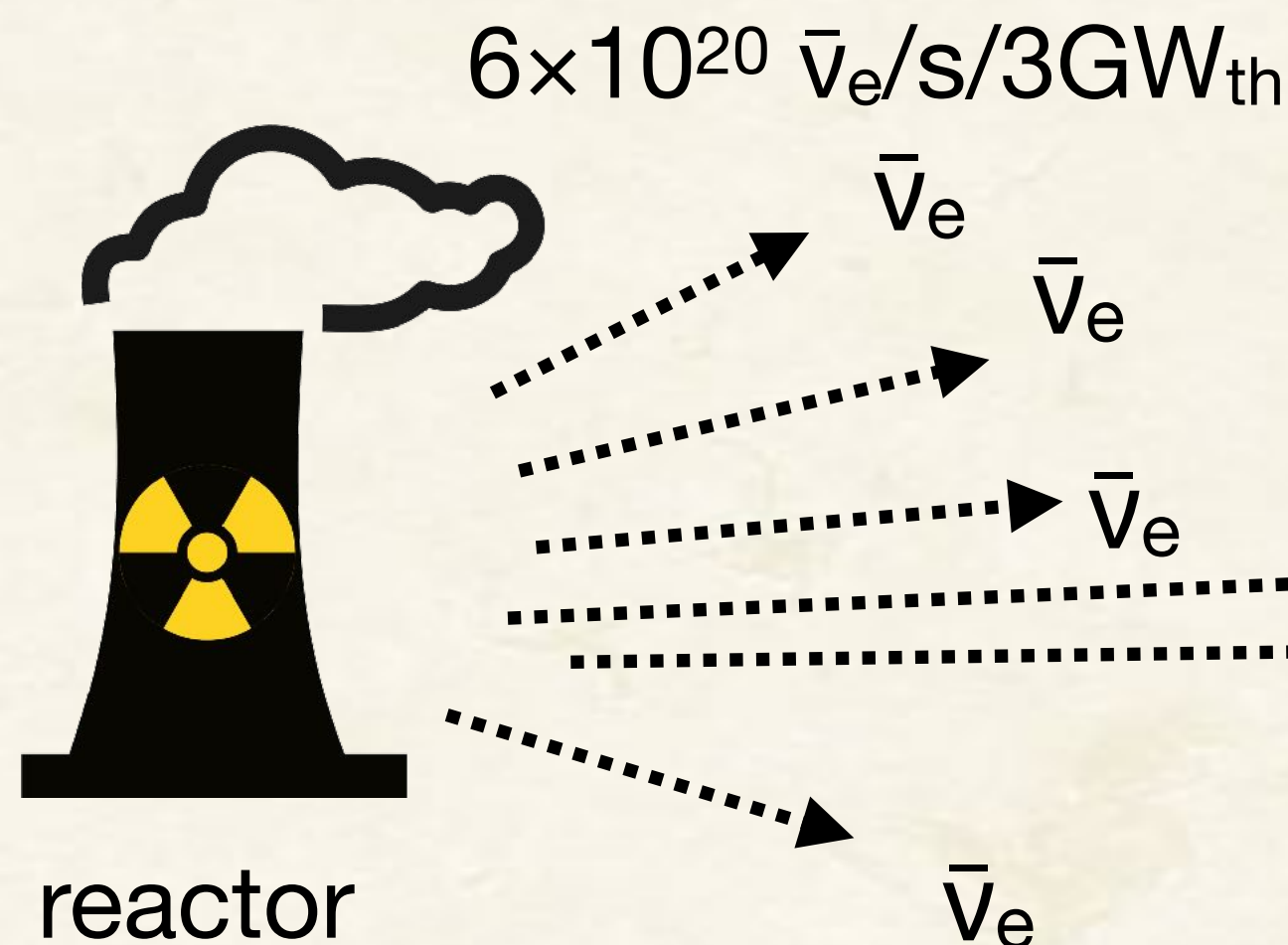
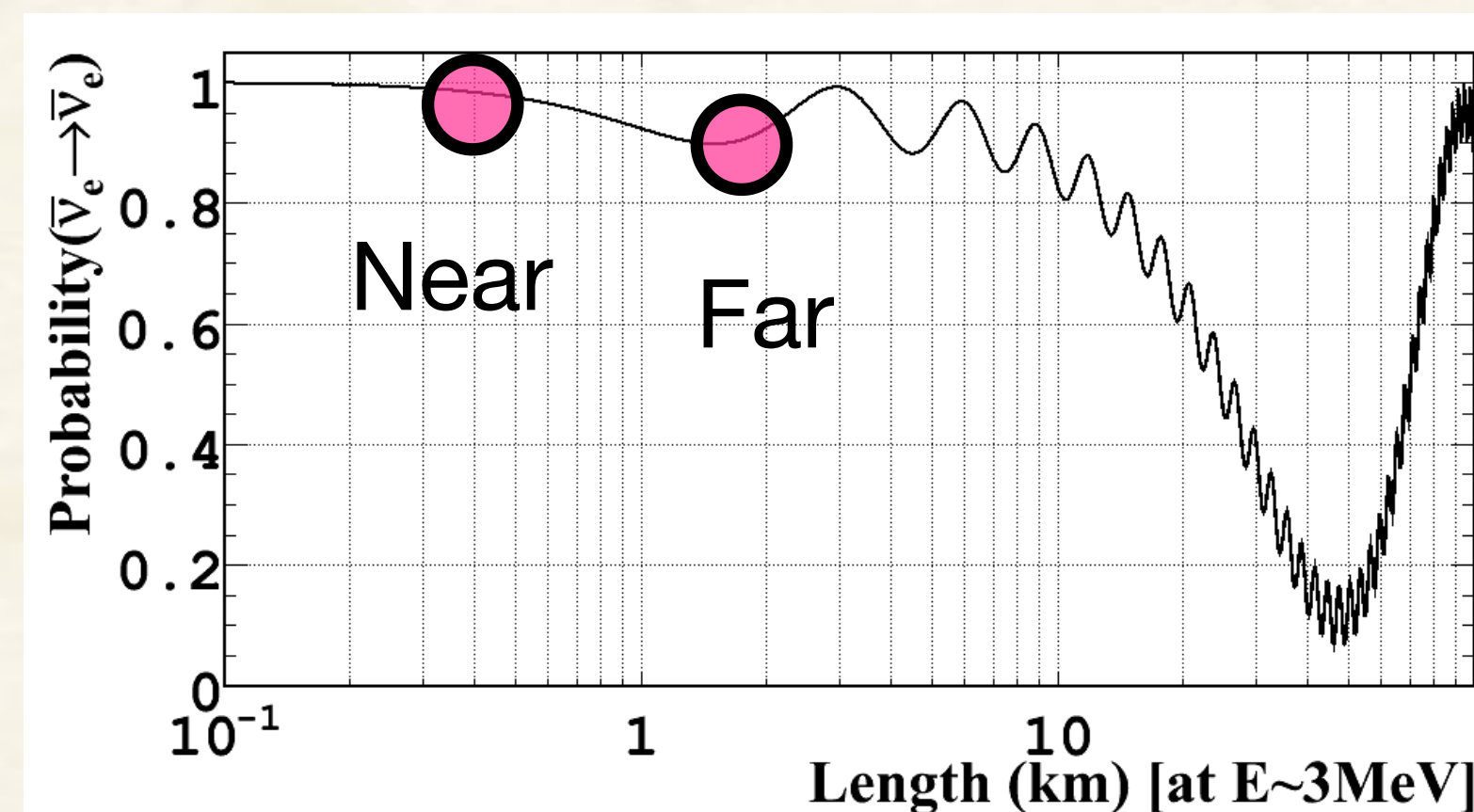
Inverse beta decay for detection (delayed coincidence)



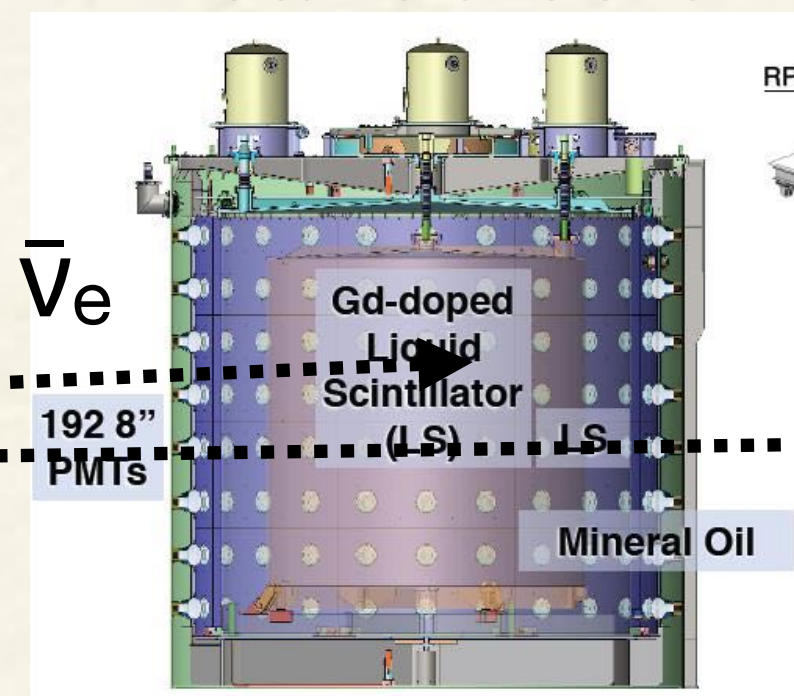
captured on Gd after thermalization

prompt signal

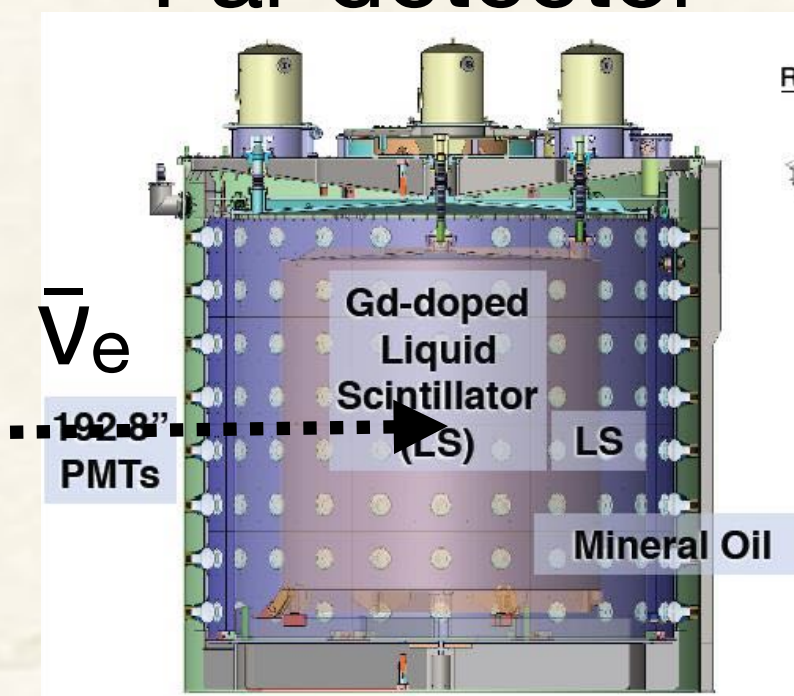
delayed signal



Near detector



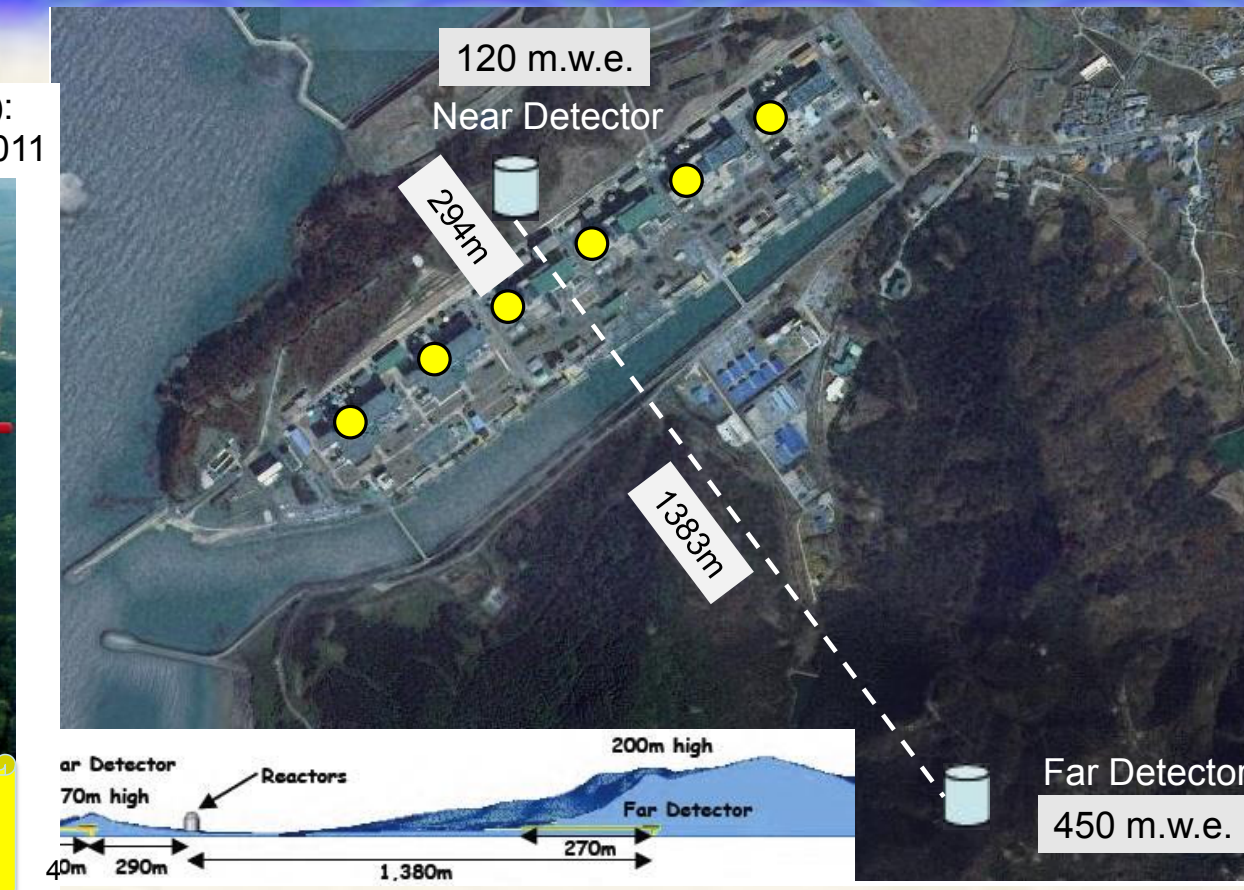
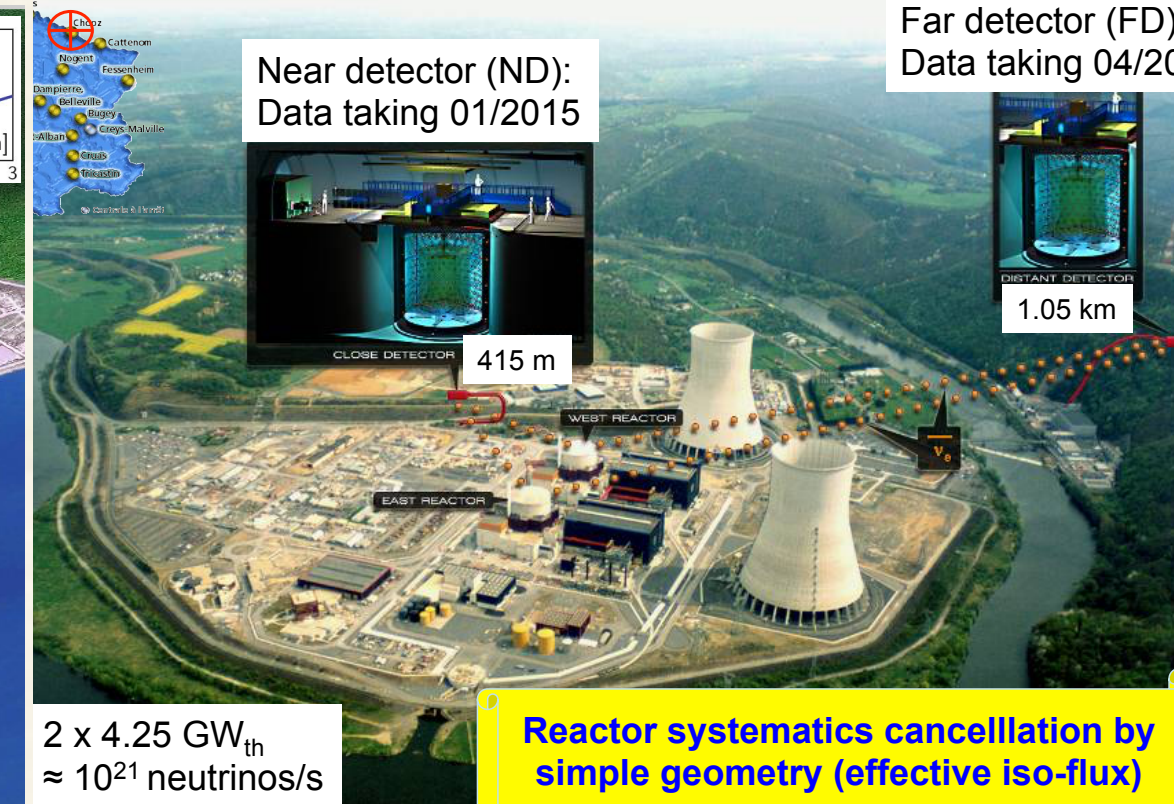
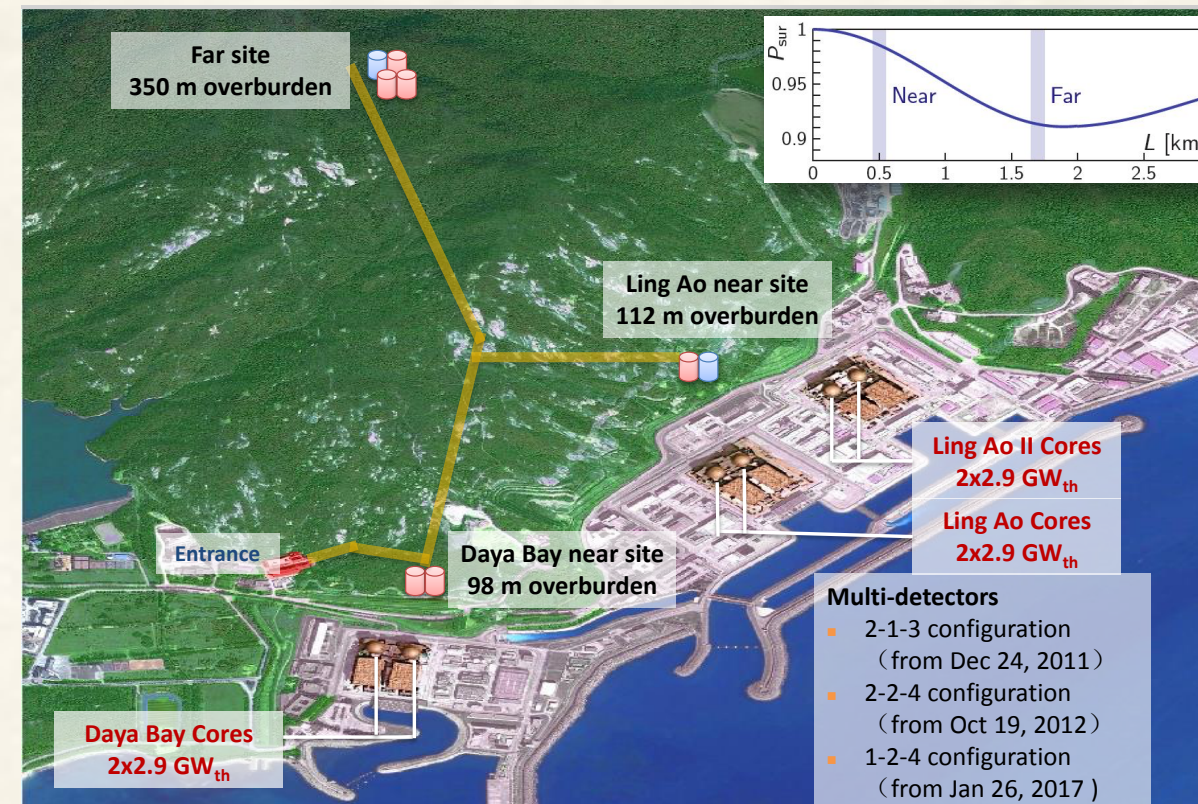
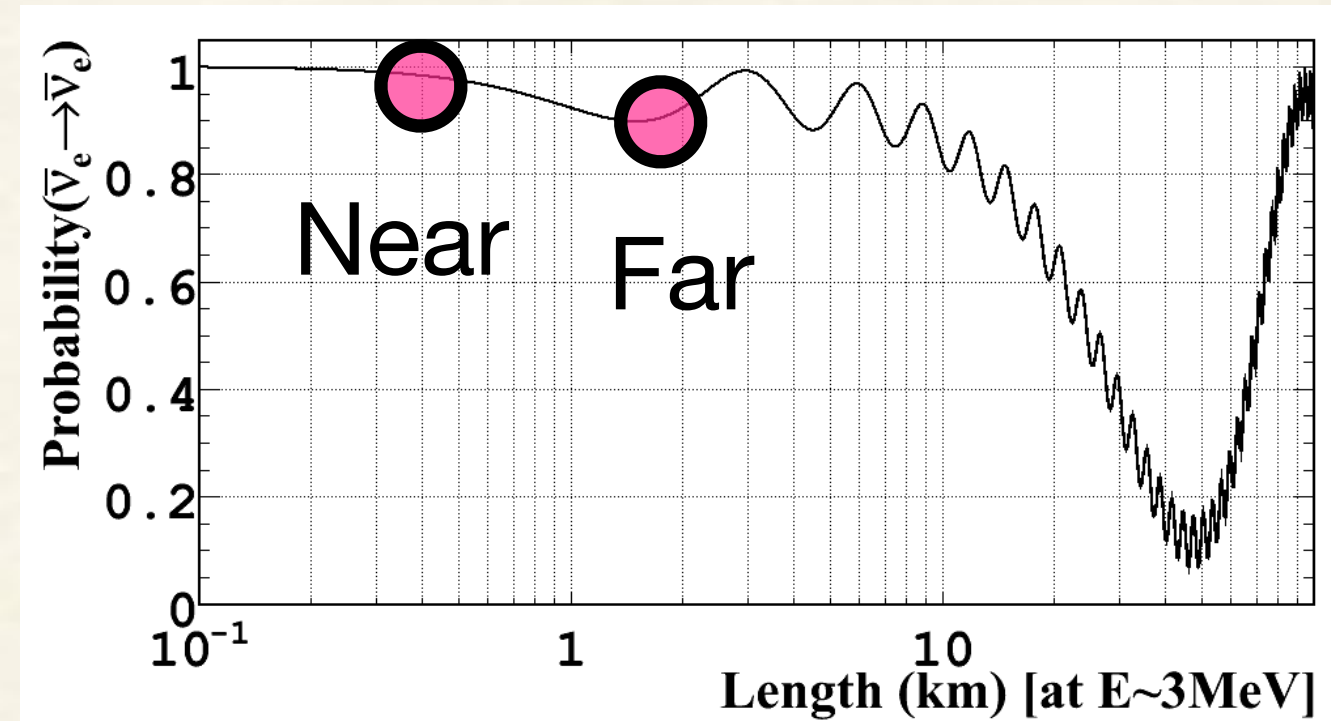
Far detector



Control systematics by two detector configuration

Long baseline neutrino experiments

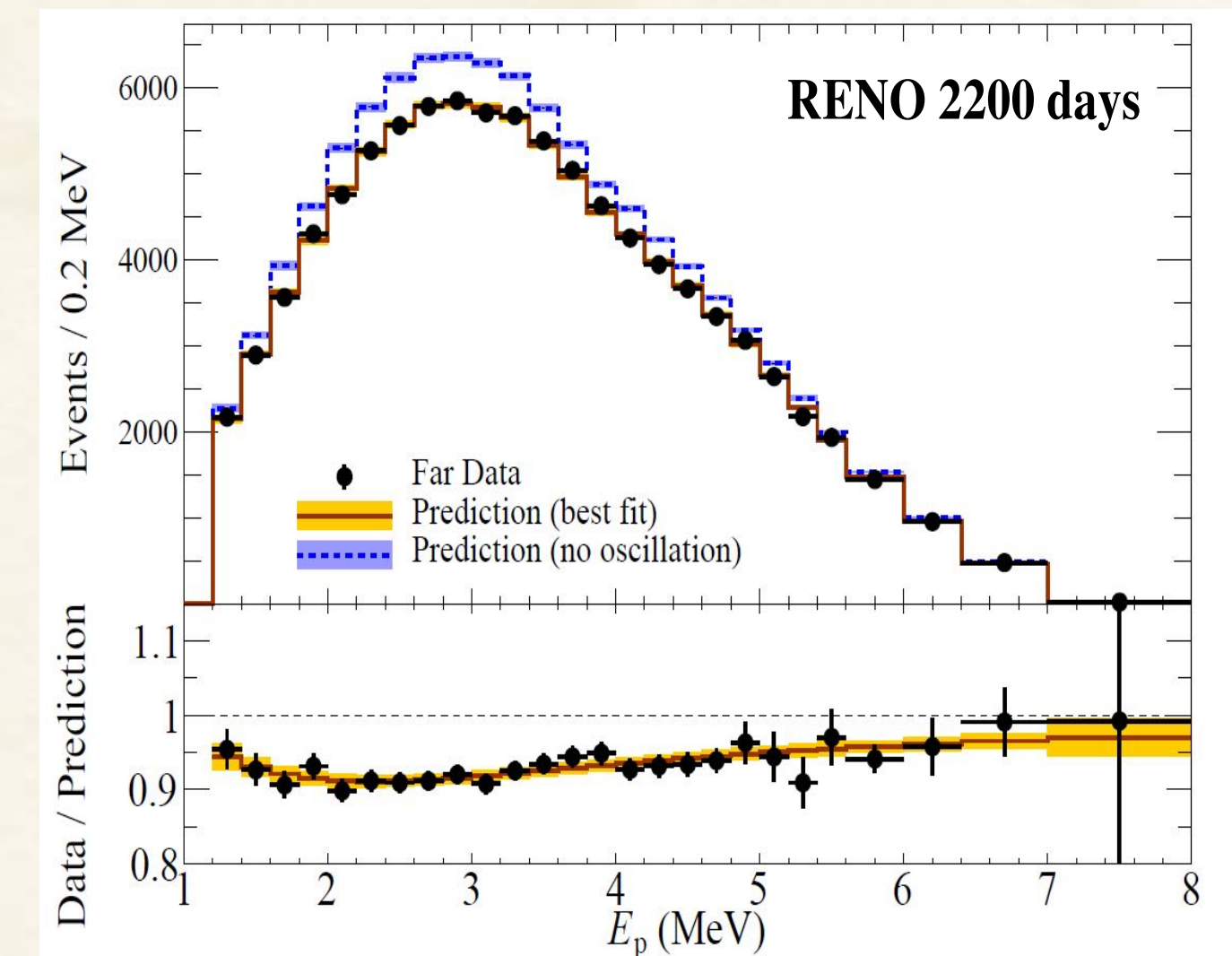
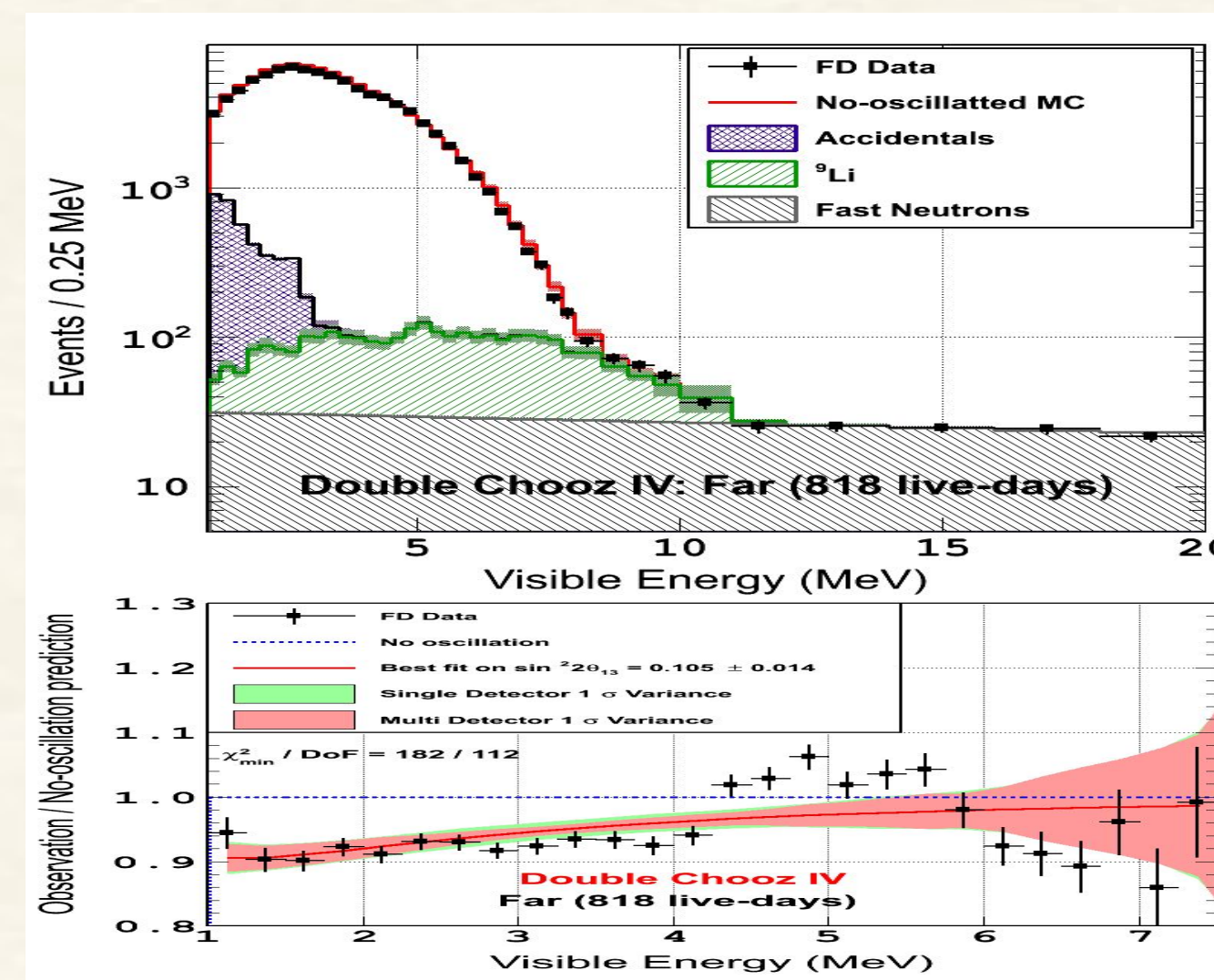
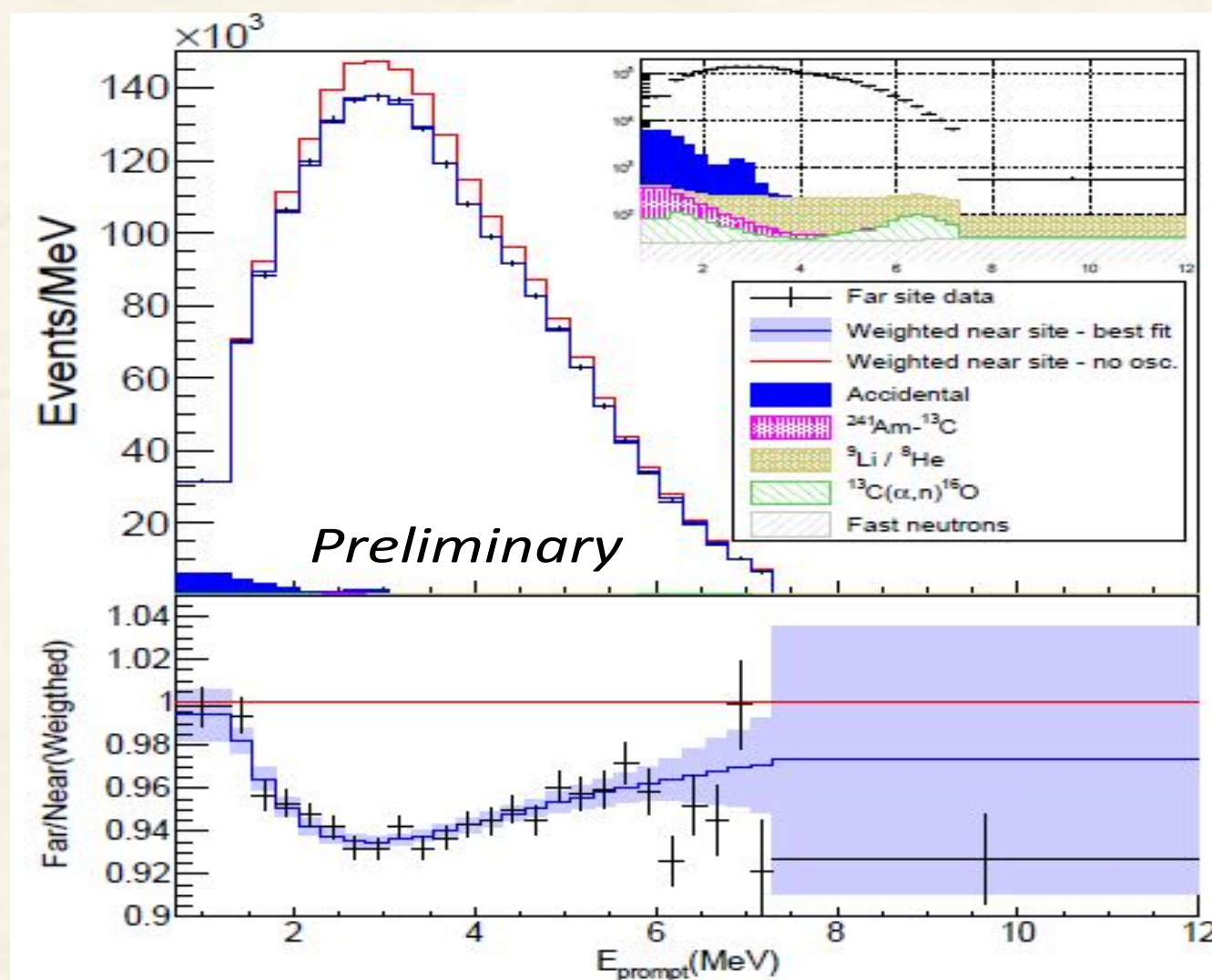
# Reactor $\theta_{13}$ experiments



	Daya Bay (China)	Double Chooz (France)	RENO (South Korea)
Reactor power (GW <sub>th</sub> )	17.4	8.5	16.8
Baseline	470/576/1650	400/1050	409/1444
Overburden near/far (m.w.e.)	250/265/860	80/300	120/450
Gd target mass for far detectors (tons)	80	8.3	16.5

Parallel talks by H.Seo, L.Zhan, A.Stahl, B.Z.Hu, M.C.Chu

# Reactor $\theta_{13}$ : latest results



$$\sin^2 2\theta_{13} = 0.0856 \pm 0.00429$$

$$\sin^2 2\theta_{13} = 0.105 \pm 0.014$$

$$\sin^2 2\theta_{13} = 0.0896 \pm 0.0048 \pm 0.0047$$

$$|\Delta m_{ee}^2| = (2.52 \pm 0.07) \times 10^{-3} eV^2$$

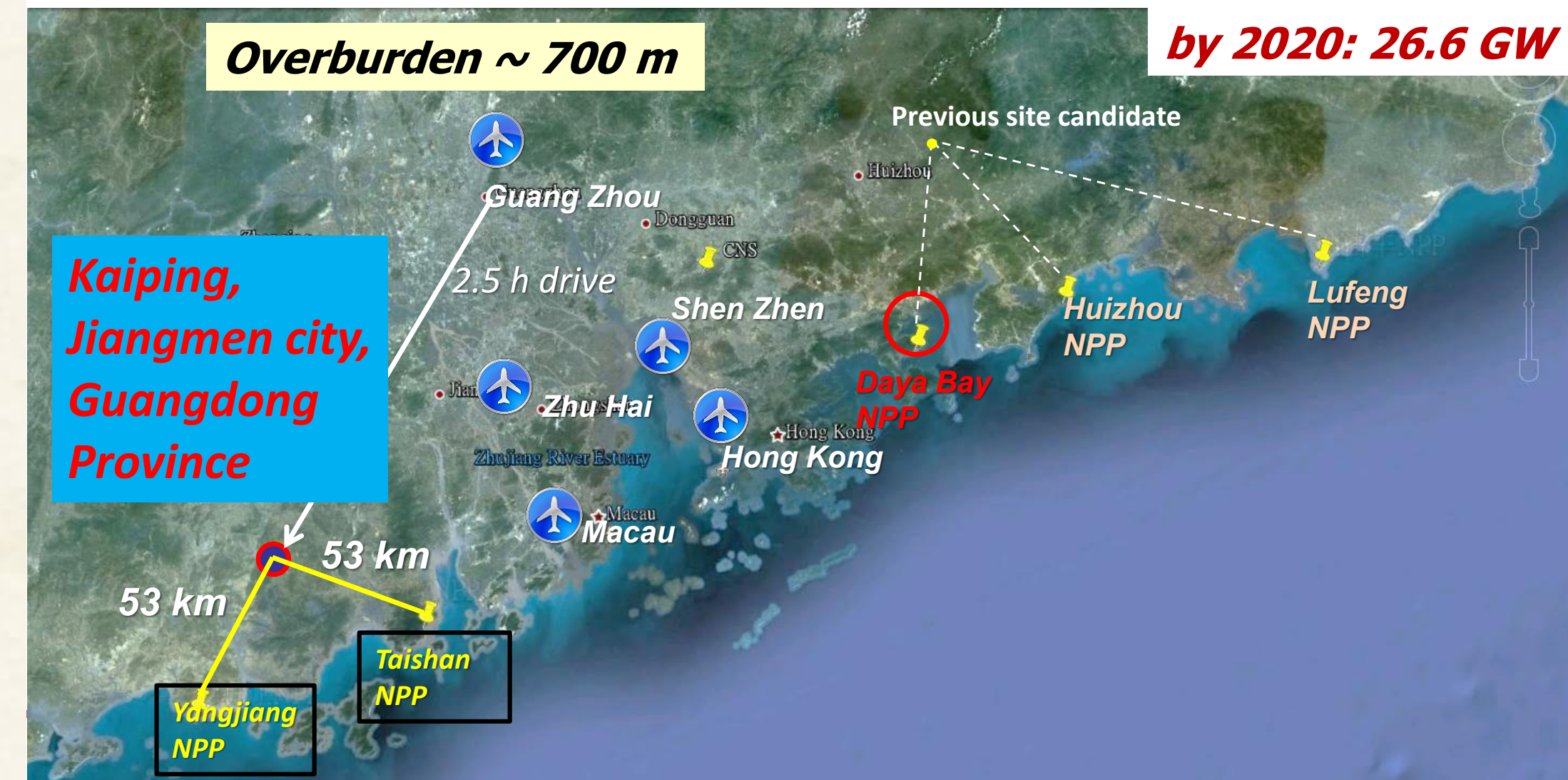
$$|\Delta m_{ee}^2| = (2.68 \pm 0.12 \pm 0.07) \times 10^{-3} eV^2$$

- ◆ Precision measurements of  $\sin^2 2\theta_{13}$  ( $\sim 5\%$ ) and also  $\Delta m^2$  ( $\sim 3\%$ )
- ◆ Many other measurements are also reported (see later talk)
- ◆ Further improvement expected from all experiments in near future

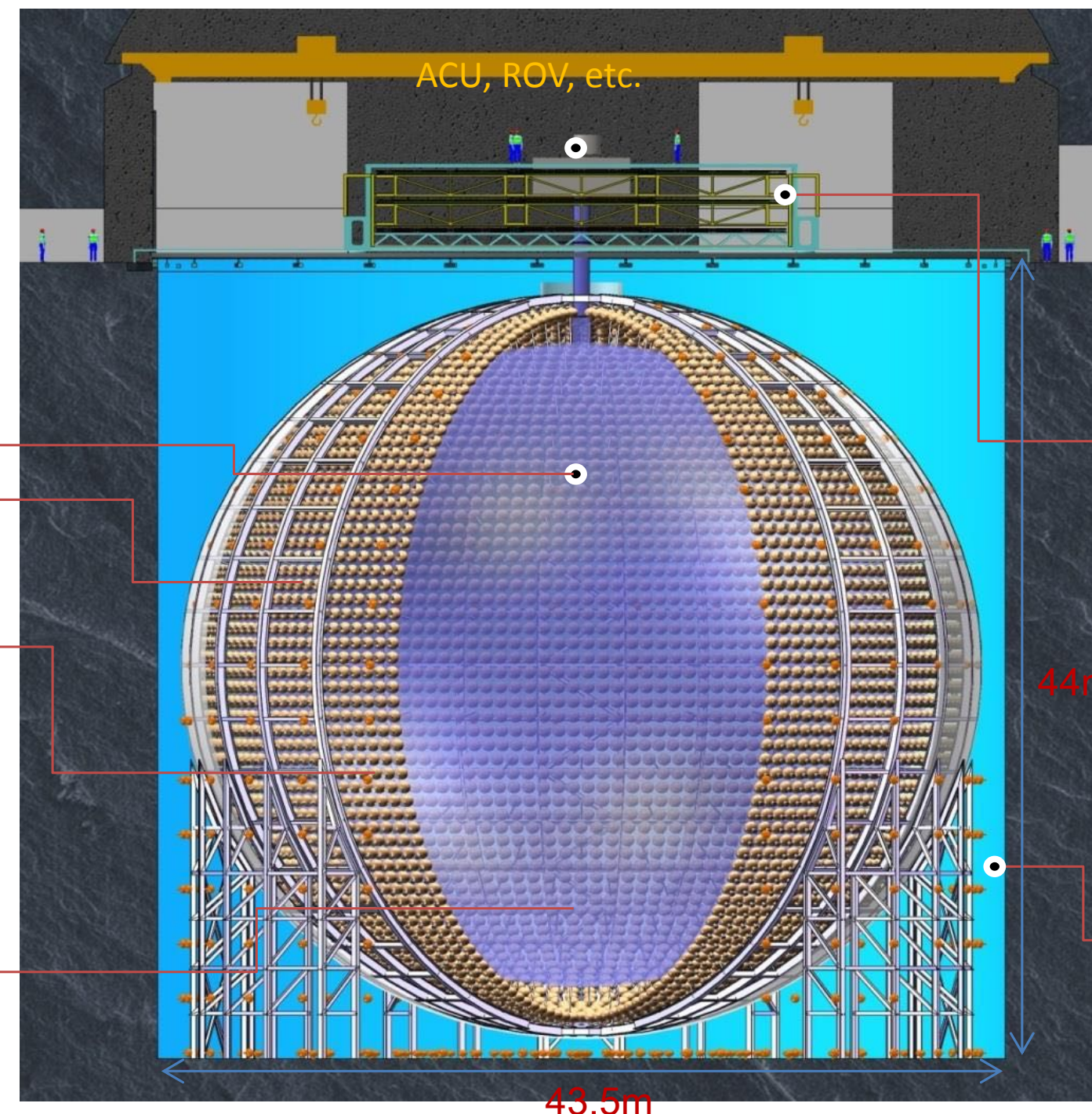
# JUNO

- ◆ **20kton** liquid scintillator detector at 53km from reactors
- ◆ Mass hierarchy determination
- ◆ Precision measurements ( $<1\%$ ) of  $\sin^2\theta_{12}$ ,  $\Delta m^2_{21}$ ,  $|\Delta m^2_{ee}|$
- ◆ Other rich physics potentials

NPP	Daya Bay	Huizhou	Lufeng	Yangjiang	Taishan
Status	Operational	Planned	Planned	Under construction	Under construction
Power	17.4 GW	17.4 GW	17.4 GW	17.4 GW	18.4 GW



## Central detector



## Calibration system

## VETO detector

Top Tracker  
-62 Plastic scintillator walls

Water Cherenkov  
-35 kt high-purity water  
-2000 20" PMTs

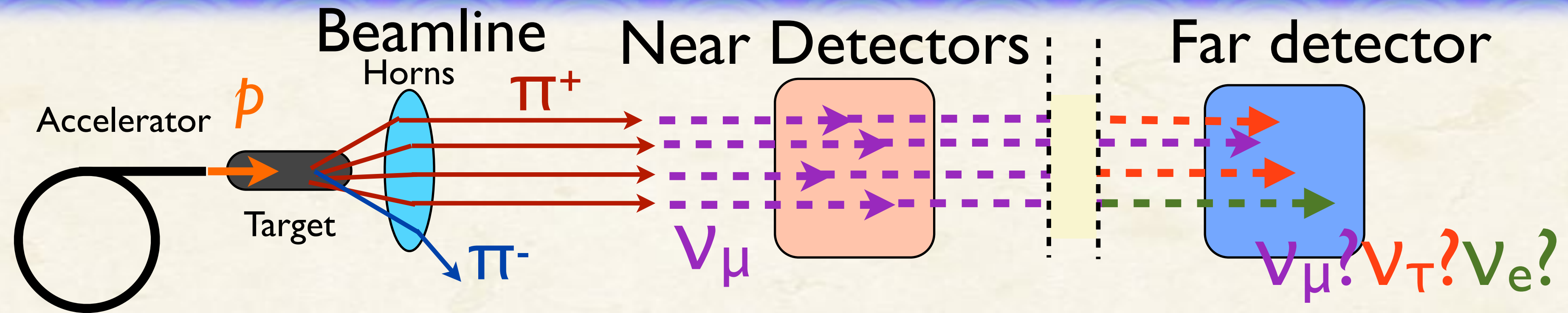
4

- ◆ Civil construction ongoing
- ◆ Detector R&D and fabrication are progressing
- ◆ Aim to start data-taking in 2021

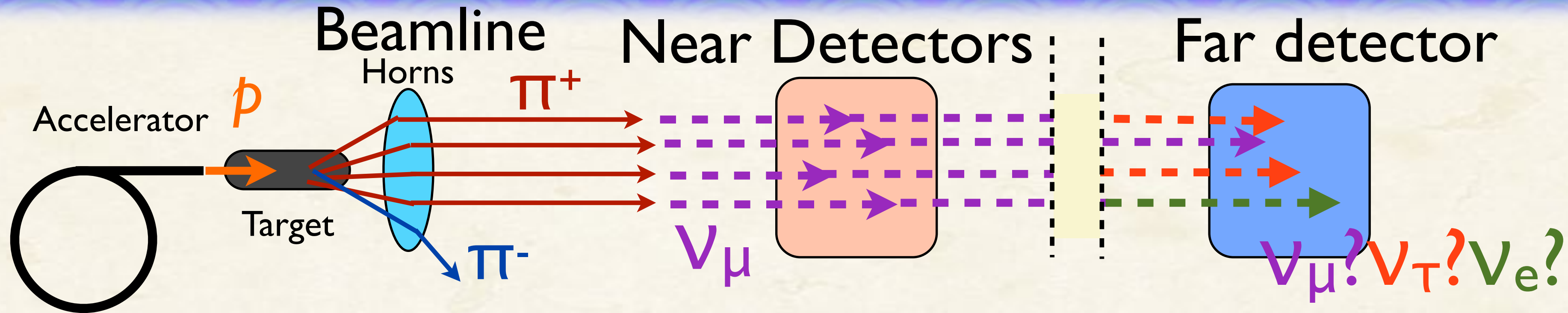
Parallel talks by Q.Zhang, L.Wen, Z.Qin,



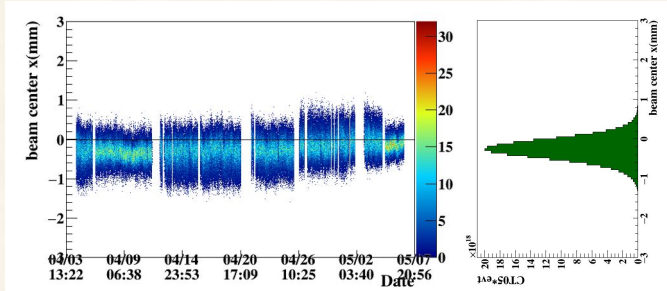
# Accelerator-Based long baseline experiments



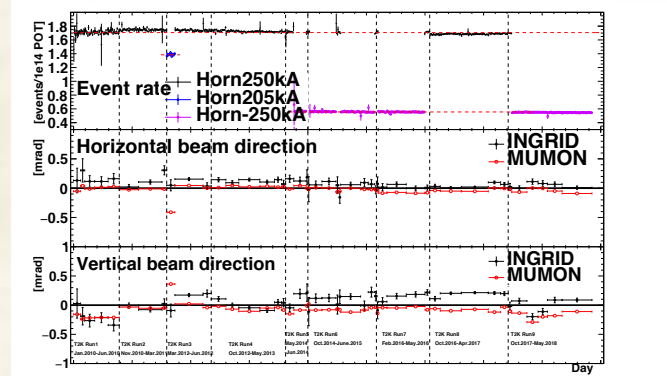
# Accelerator-Based long baseline experiments



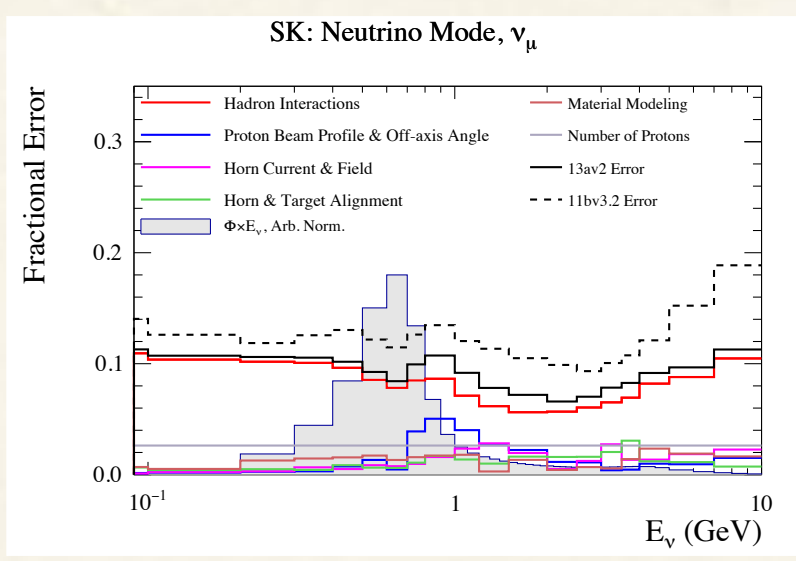
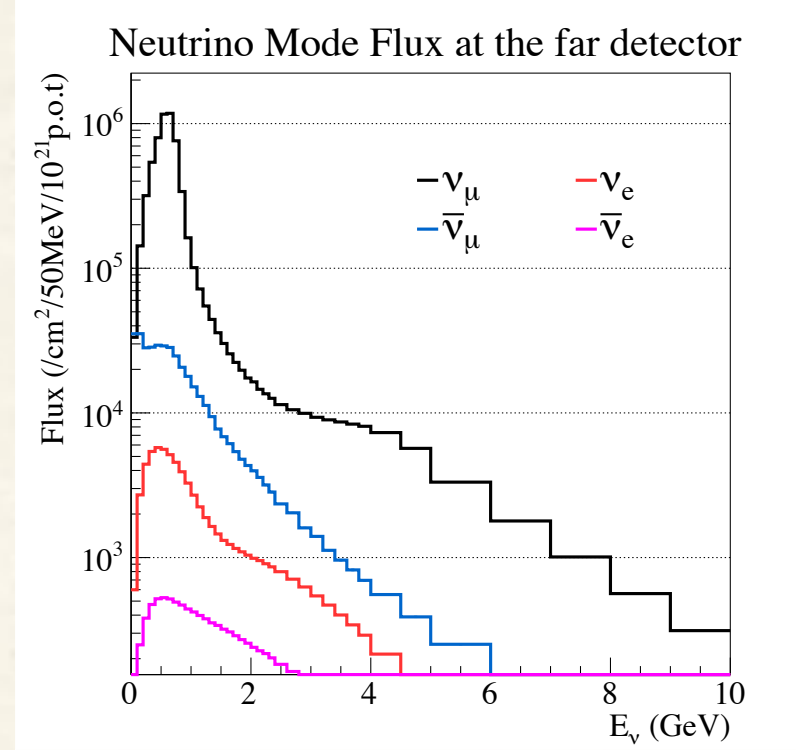
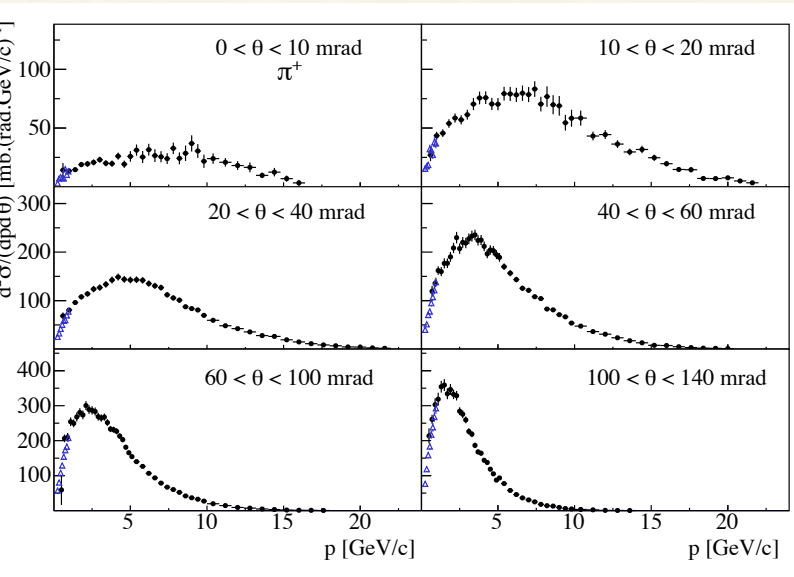
In-situ beam measurements



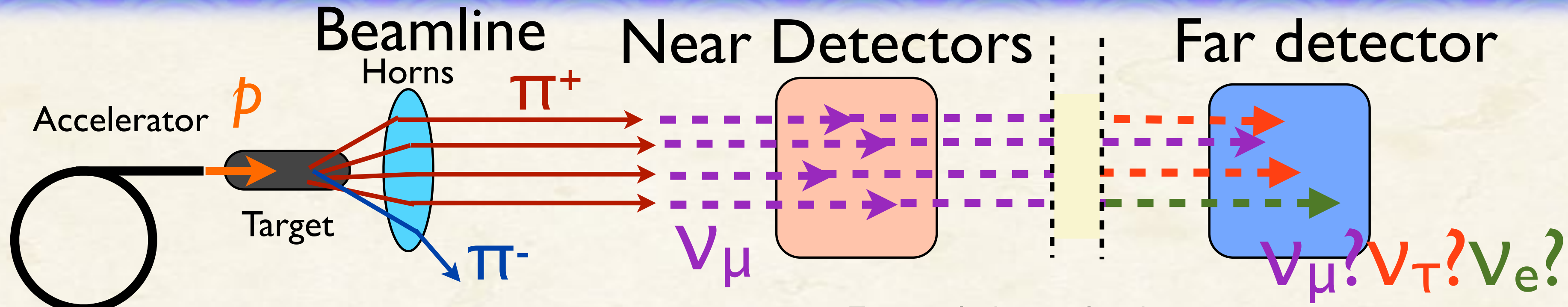
Prediction of neutrino flux, uncertainties and correlation



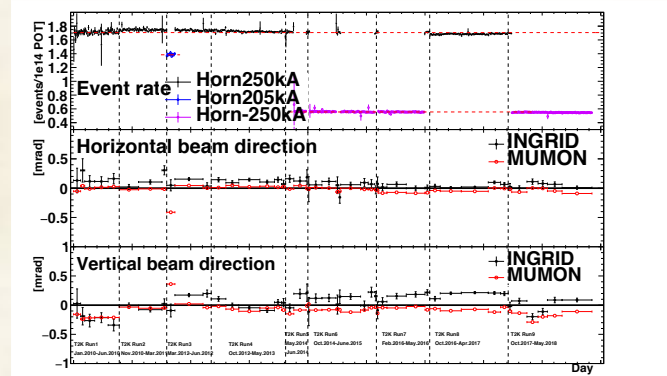
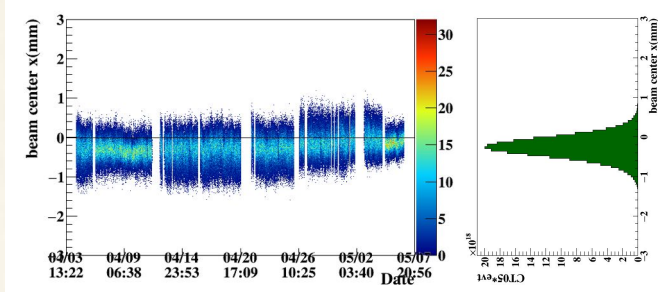
External hadron production measurements



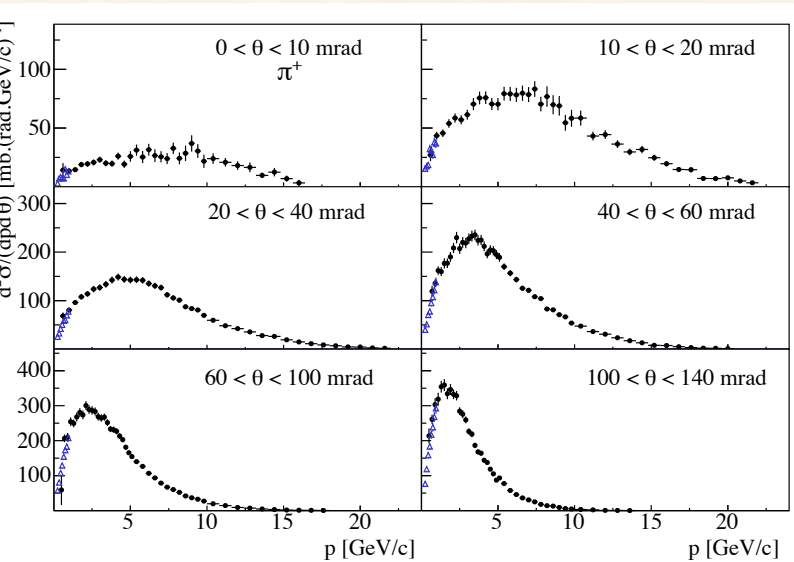
# Accelerator-Based long baseline experiments



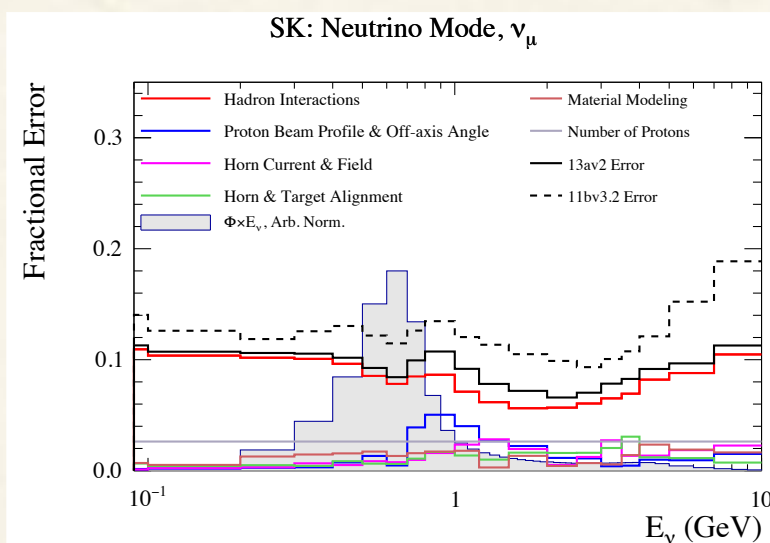
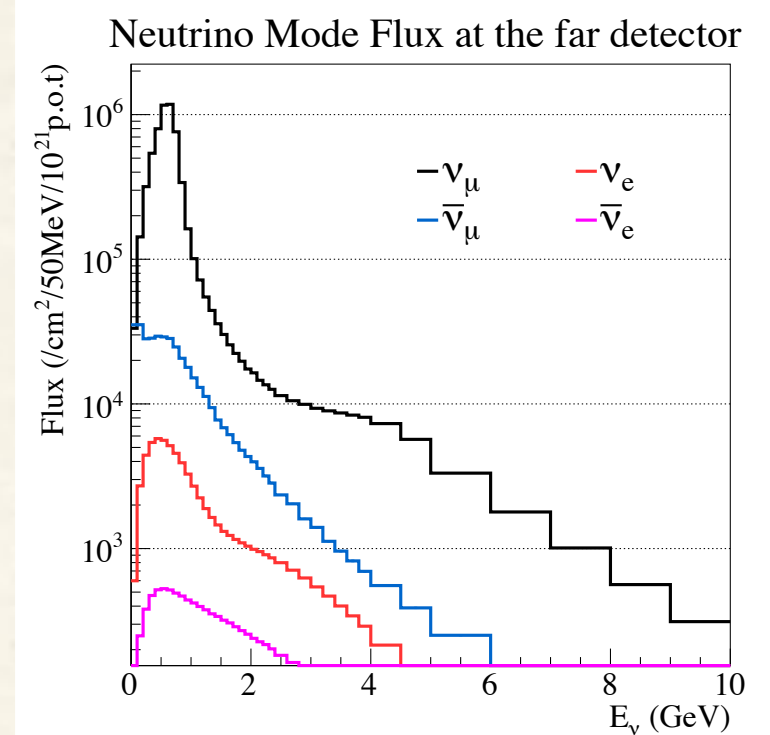
In-situ beam measurements



External hadron production measurements



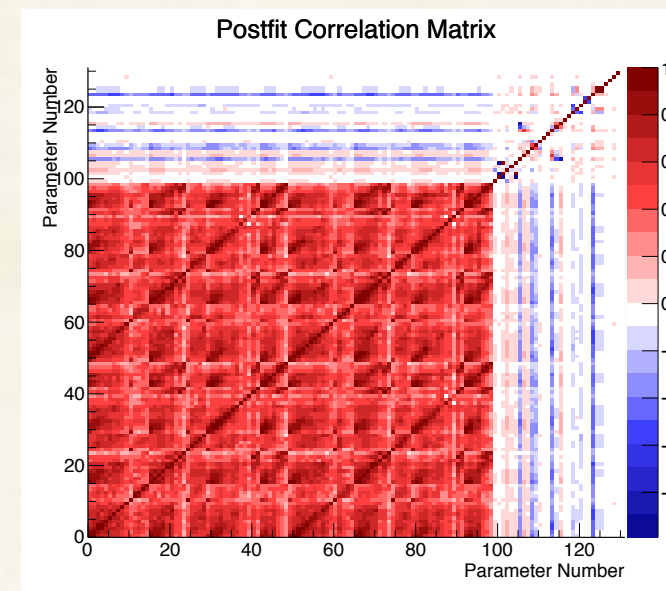
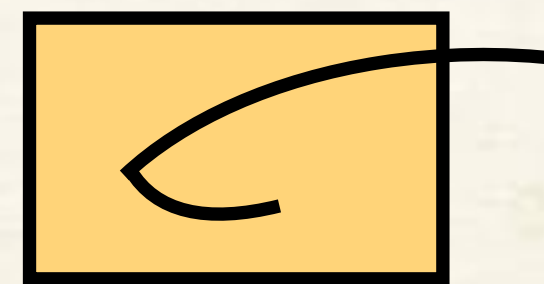
Prediction of neutrino flux, uncertainties and correlation



Extrapolation to far detector

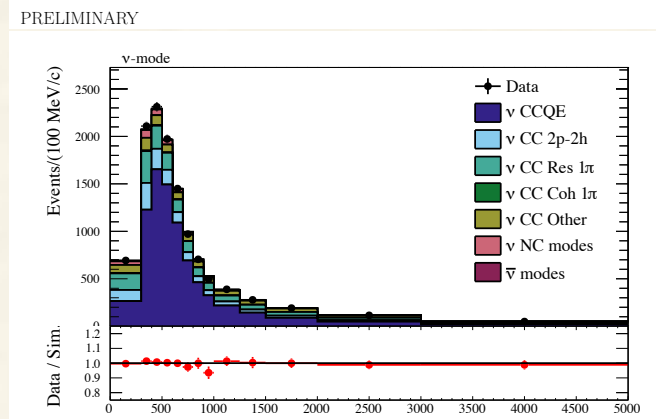
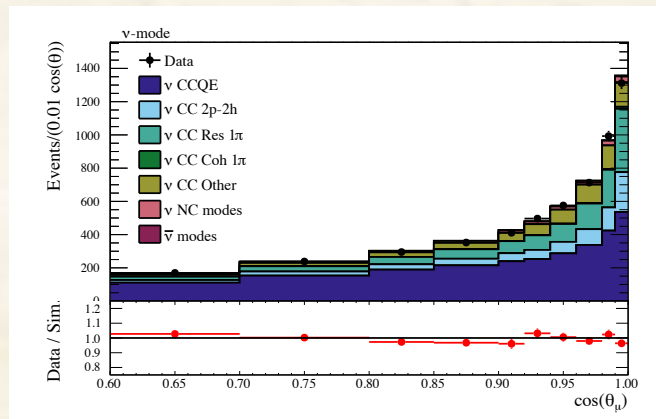


Near detector measurements



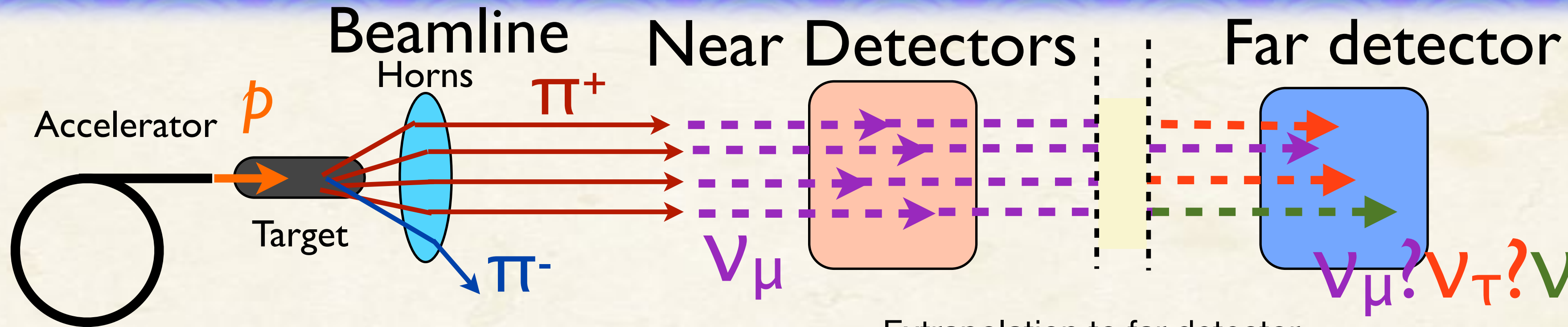
$$\phi_{ND}(E_\nu) \cdot \sigma(E_\nu) \cdot \epsilon_{ND}$$

External cross section measurements and models

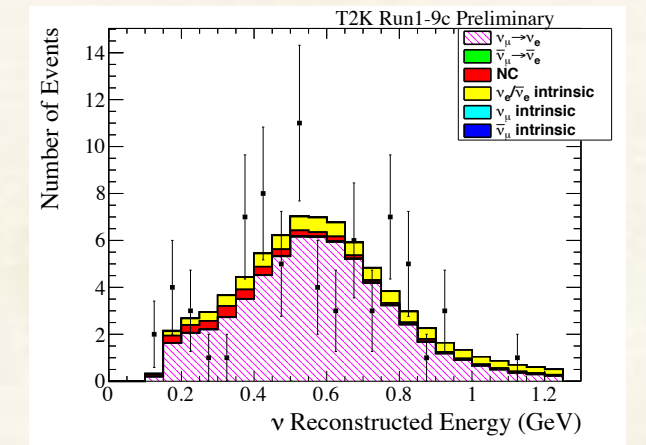


PRELIMINARY

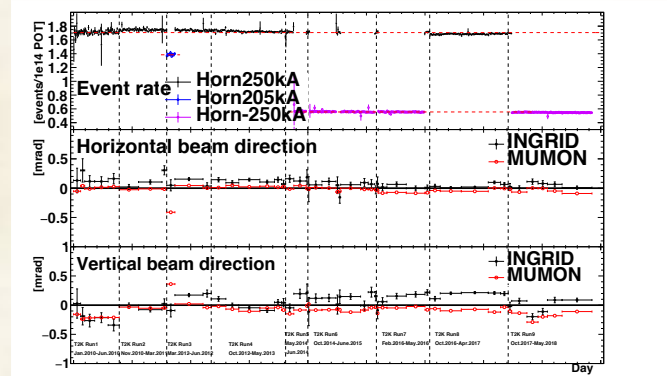
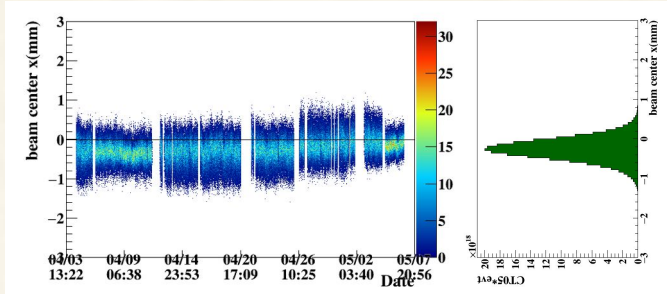
# Accelerator-Based long baseline experiments



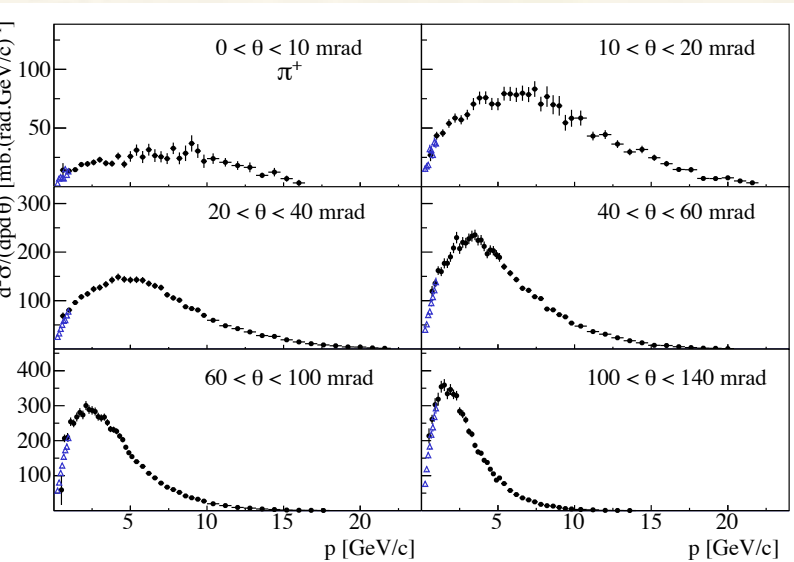
Far detector data



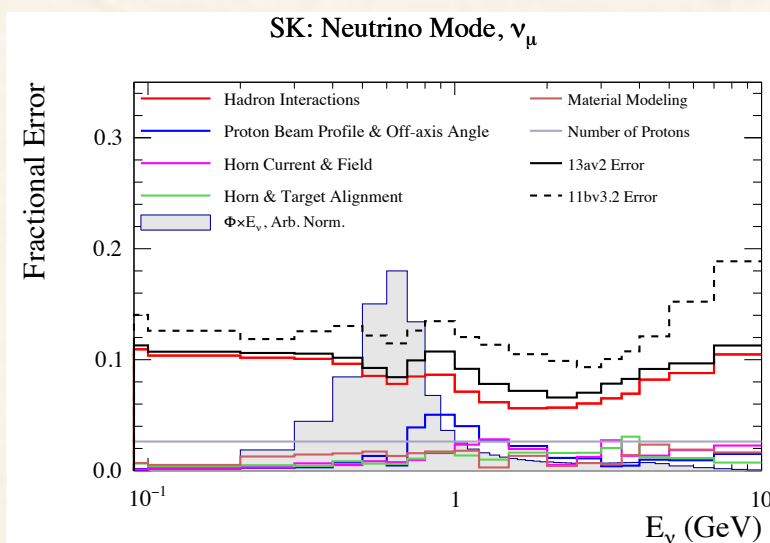
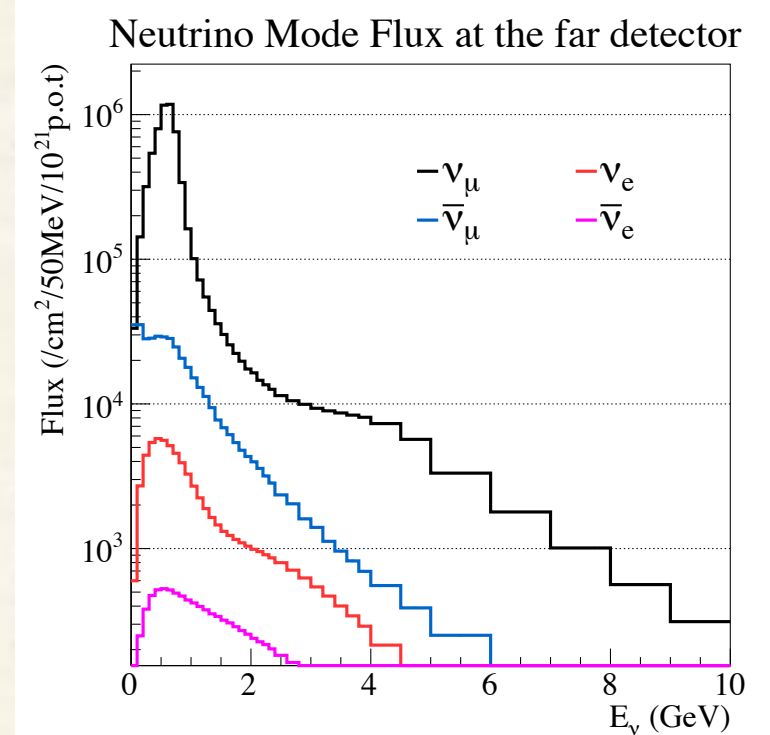
In-situ beam measurements



External hadron production measurements

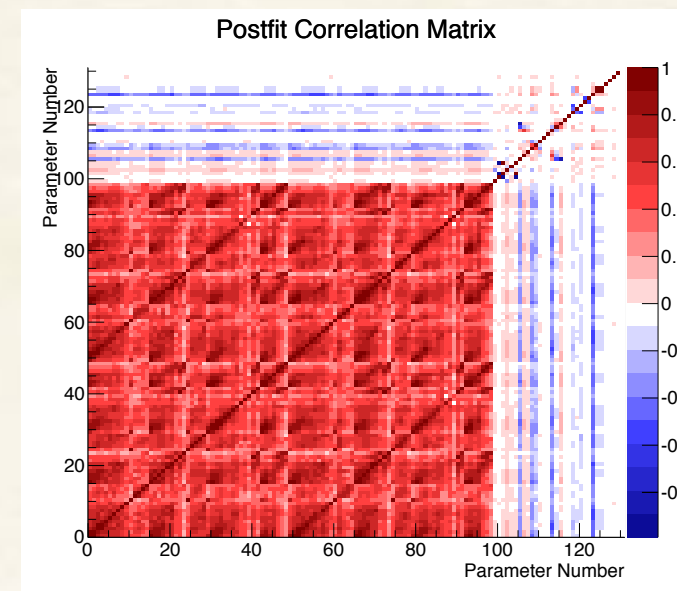
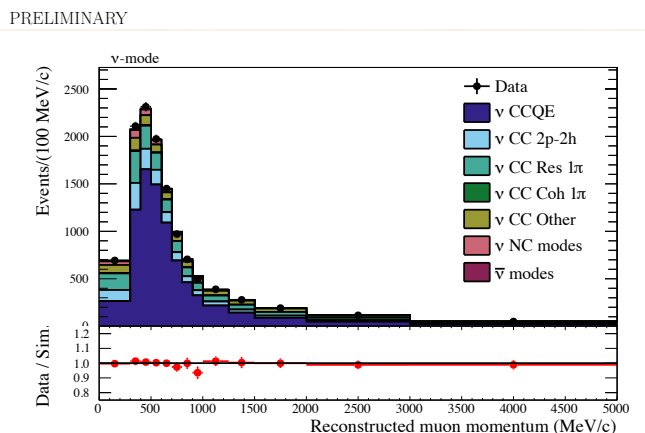
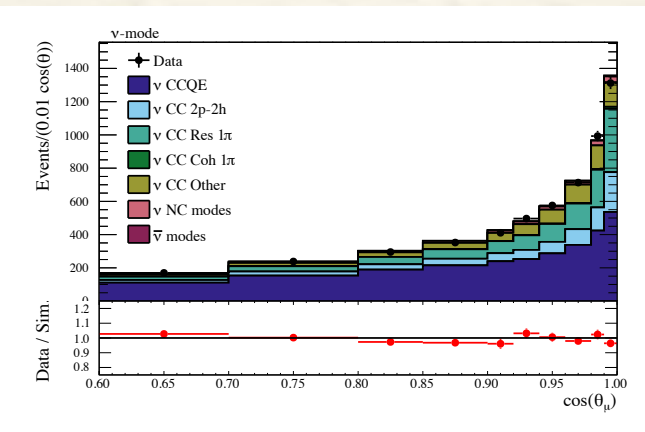
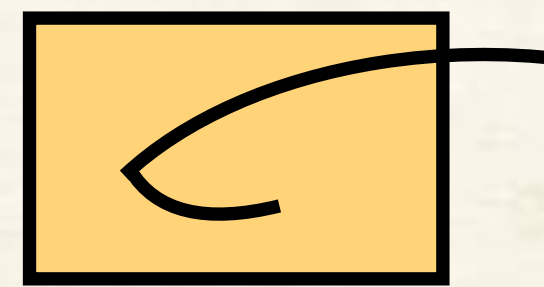


Prediction of neutrino flux, uncertainties and correlation



Extrapolation to far detector

Near detector measurements



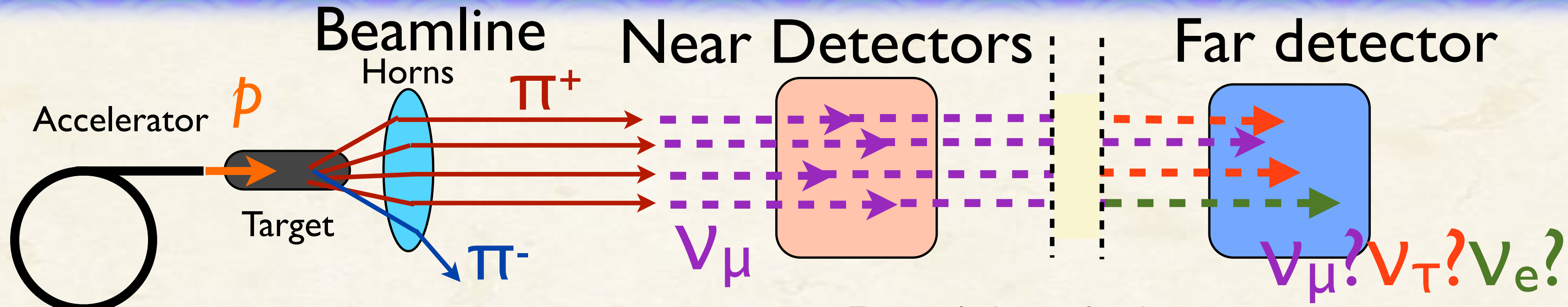
$$\phi_{ND}(E_\nu) \cdot \sigma(E_\nu) \cdot \epsilon_{ND}$$

Prediction at far detector

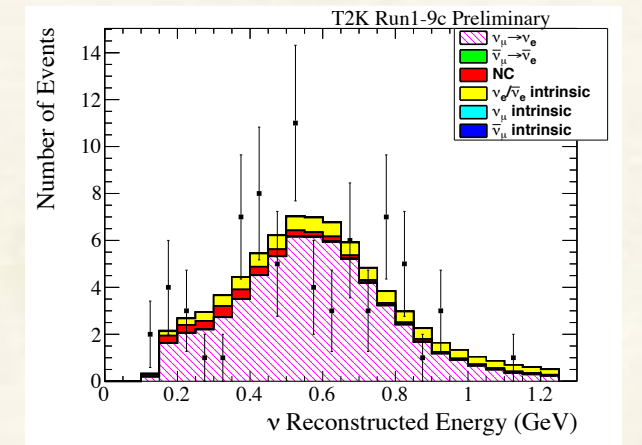
$$\phi_{FD}(E_\nu) \cdot \sigma(E_\nu) \cdot \epsilon_{FD} \cdot P_{osc}(E_\nu)$$

External cross section measurements and models

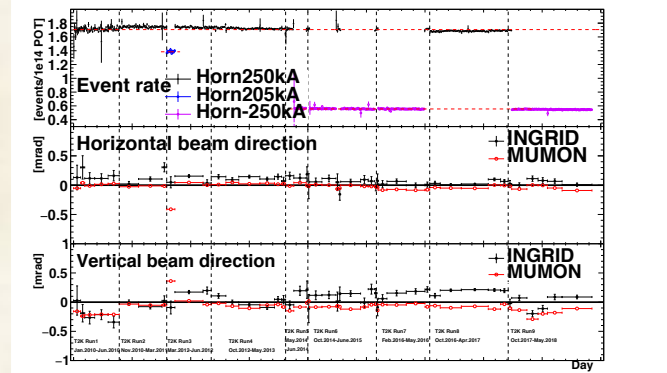
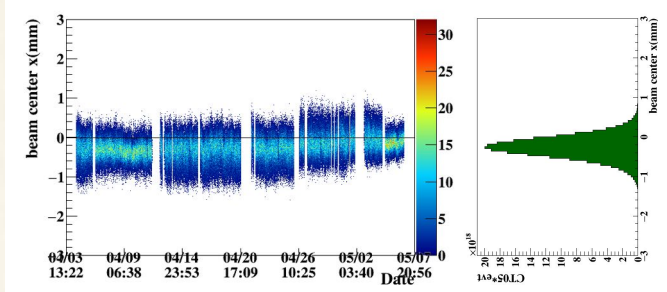
# Accelerator-Based long baseline experiments



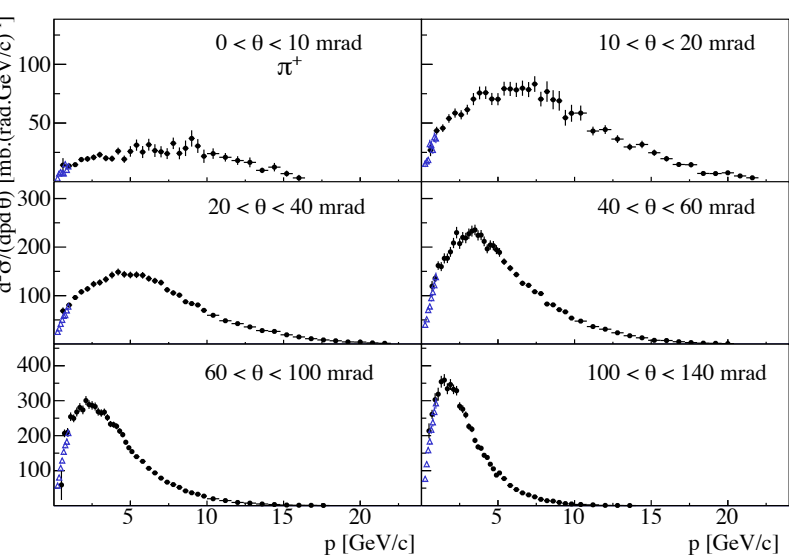
Far detector data



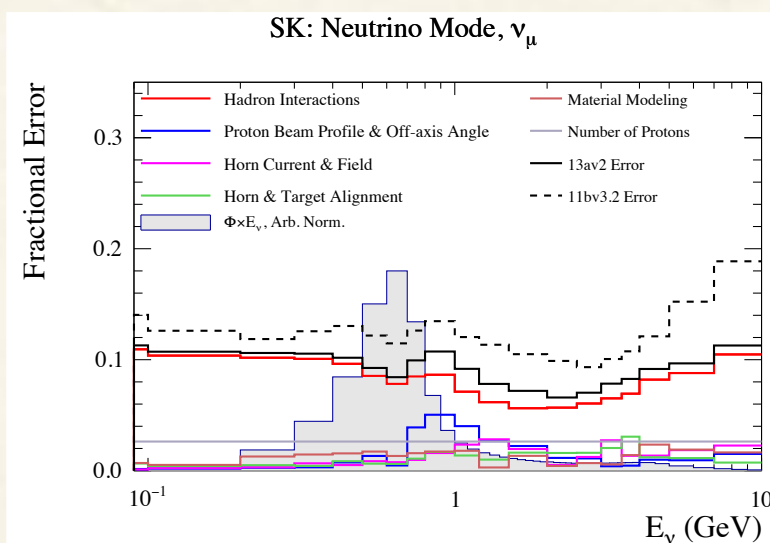
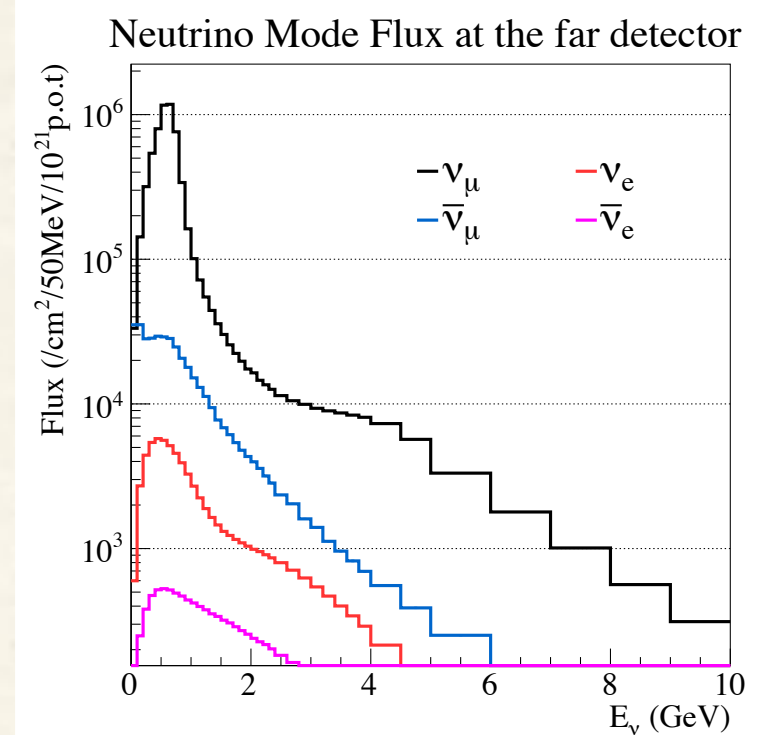
In-situ beam measurements



External hadron production measurements

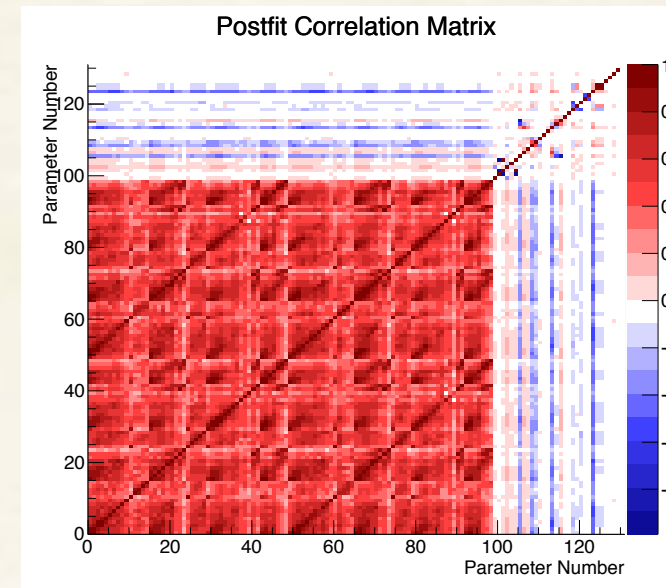
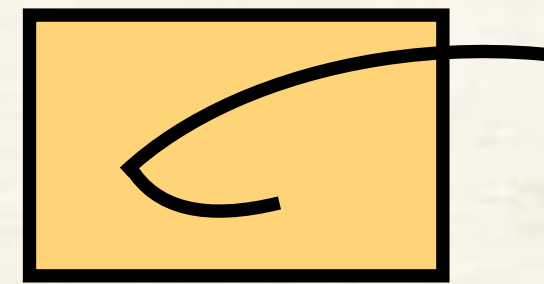


Prediction of neutrino flux, uncertainties and correlation



Extrapolation to far detector

Near detector measurements



$$\phi_{ND}(E_\nu) \cdot \sigma(E_\nu) \cdot \epsilon_{ND}$$

Prediction at far detector

$$\phi_{FD}(E_\nu) \cdot \sigma(E_\nu) \cdot \epsilon_{FD} \cdot P_{osc}(E_\nu)$$

External cross section measurements and models

Oscillation parameters

# Oscillations in accelerator LBL experiments

$\nu_\mu$  “disappearance”

$$P(\nu_\mu \rightarrow \nu_\mu) \simeq 1 - (\cos^4 \theta_{13} \sin^2 2\theta_{23} - \sin^2 2\theta_{13} \sin^2 \theta_{23}) \sin^2 \left[ \frac{\Delta m_{32}^2 L}{4E} \right]$$

$\sin^2 2\theta_{23}$  from the leading term

$\nu_e$  “appearance”

$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2[(A-1)\Delta]}{(A-1)^2}$$

$$\alpha = \frac{\Delta m_{21}^2}{\Delta m_{32}^2} \ll 1 \quad \Delta = \frac{\Delta m_{32}^2 L}{4E} \quad A = 2\sqrt{2} G_F N_e \frac{E_\nu}{\Delta m_{32}^2}$$

$$\mp \alpha \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \sin \delta_{\text{CP}} \sin \Delta \frac{\sin A\Delta}{A} \frac{\sin[(1-A)\Delta]}{1-A}$$

M.Freund, Phys.Rev. D64 (2001) 053003

$$+ \alpha \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \delta_{\text{CP}} \cos \Delta \frac{\sin A\Delta}{A} \frac{\sin[(1-A)\Delta]}{1-A} + O(\alpha^2)$$

The leading term dependent on  $\sin^2 2\theta_{13}$ ,  $\delta_{\text{CP}}$  and mass hierarchy from sub-leading terms

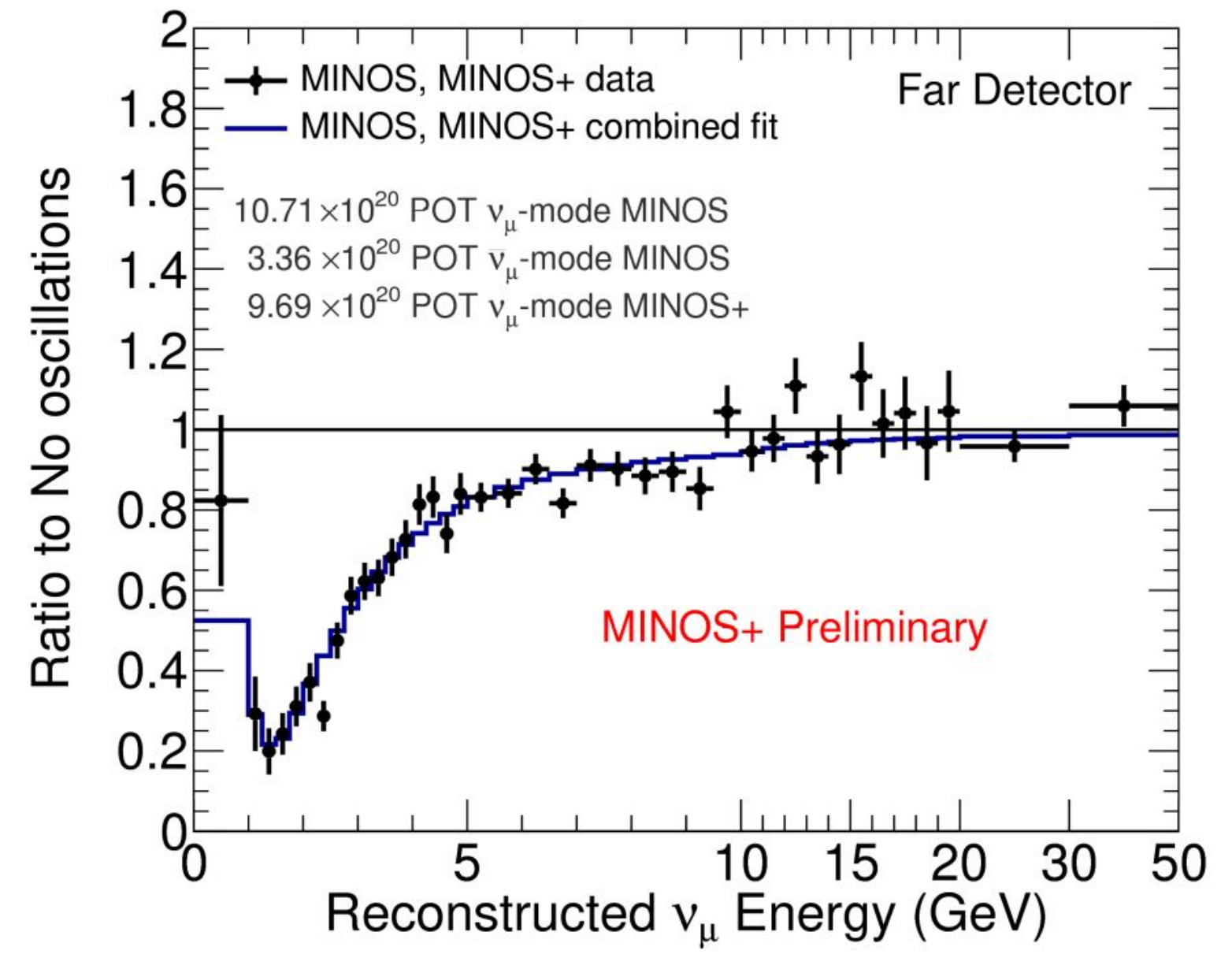
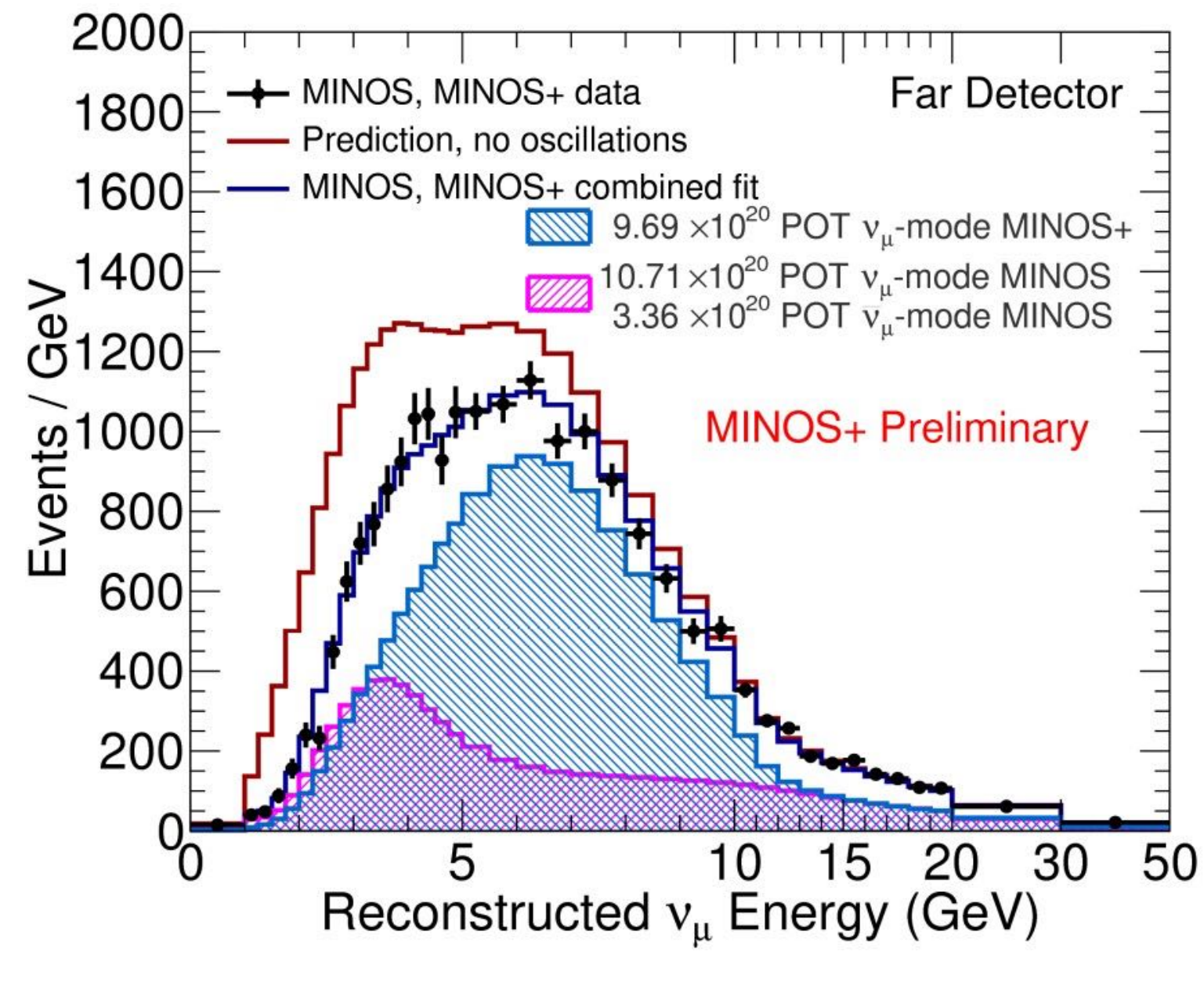
# Final results from MINOS/MINOS+



2005-2016



Baseline:  
730km  
Peak energy:  
~3GeV (MINOS)  
~7GeV (MINOS+)  
+atmospheric  $\nu$

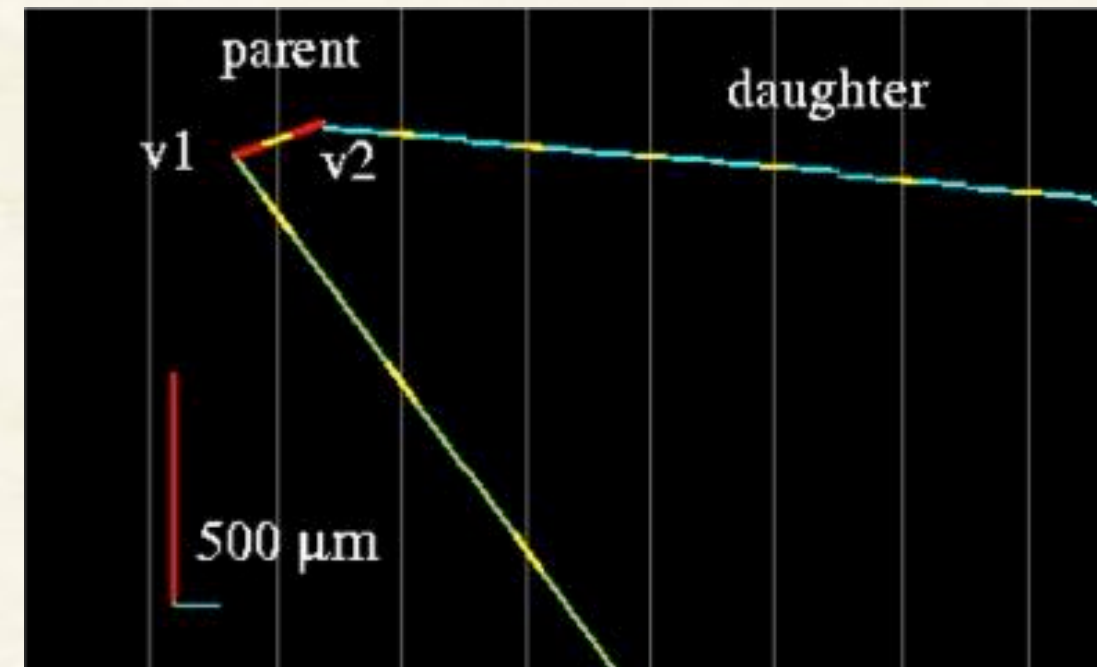
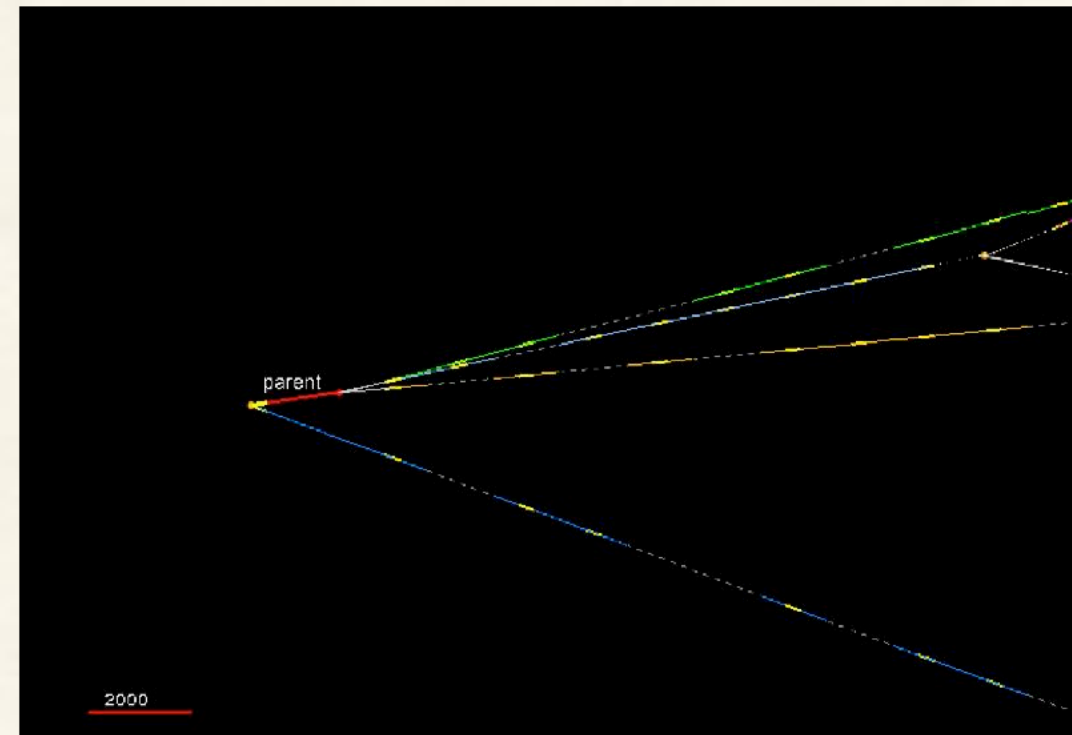
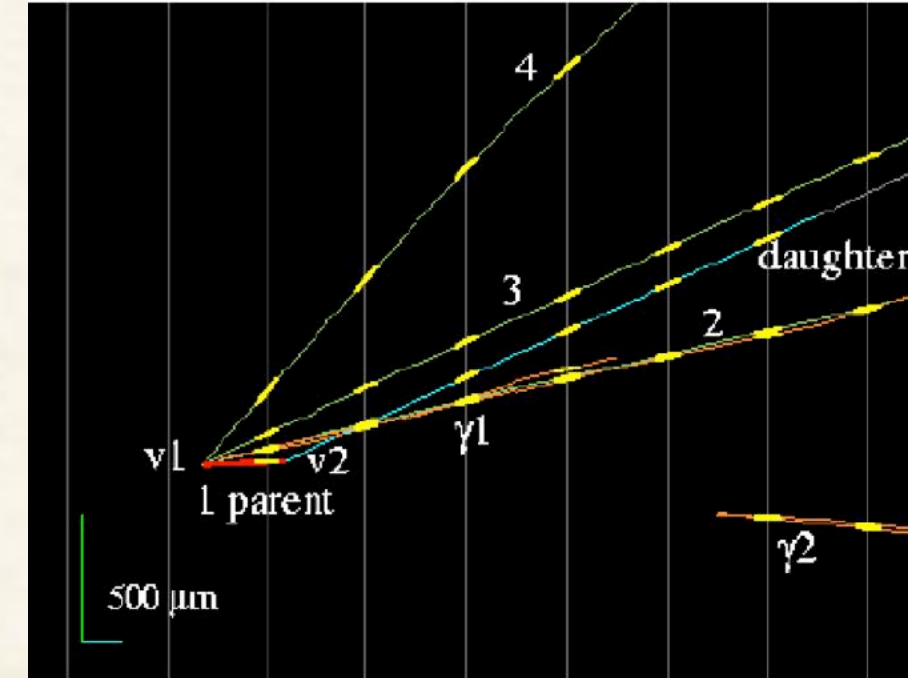
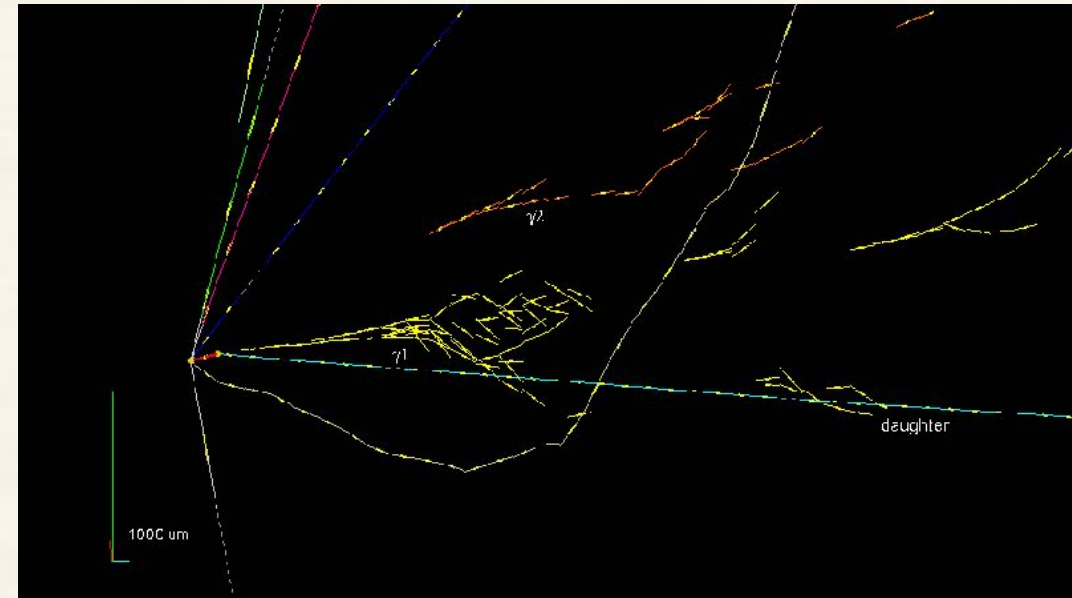


$ \Delta m^2_{32} $	$(2.28-2.55) \times 10^{-3} \text{ eV}^2$ (NH)	$\sin^2\theta_{23}$	0.36-0.65
	$(2.33-2.60) \times 10^{-3} \text{ eV}^2$ (IH)		(90%CL)

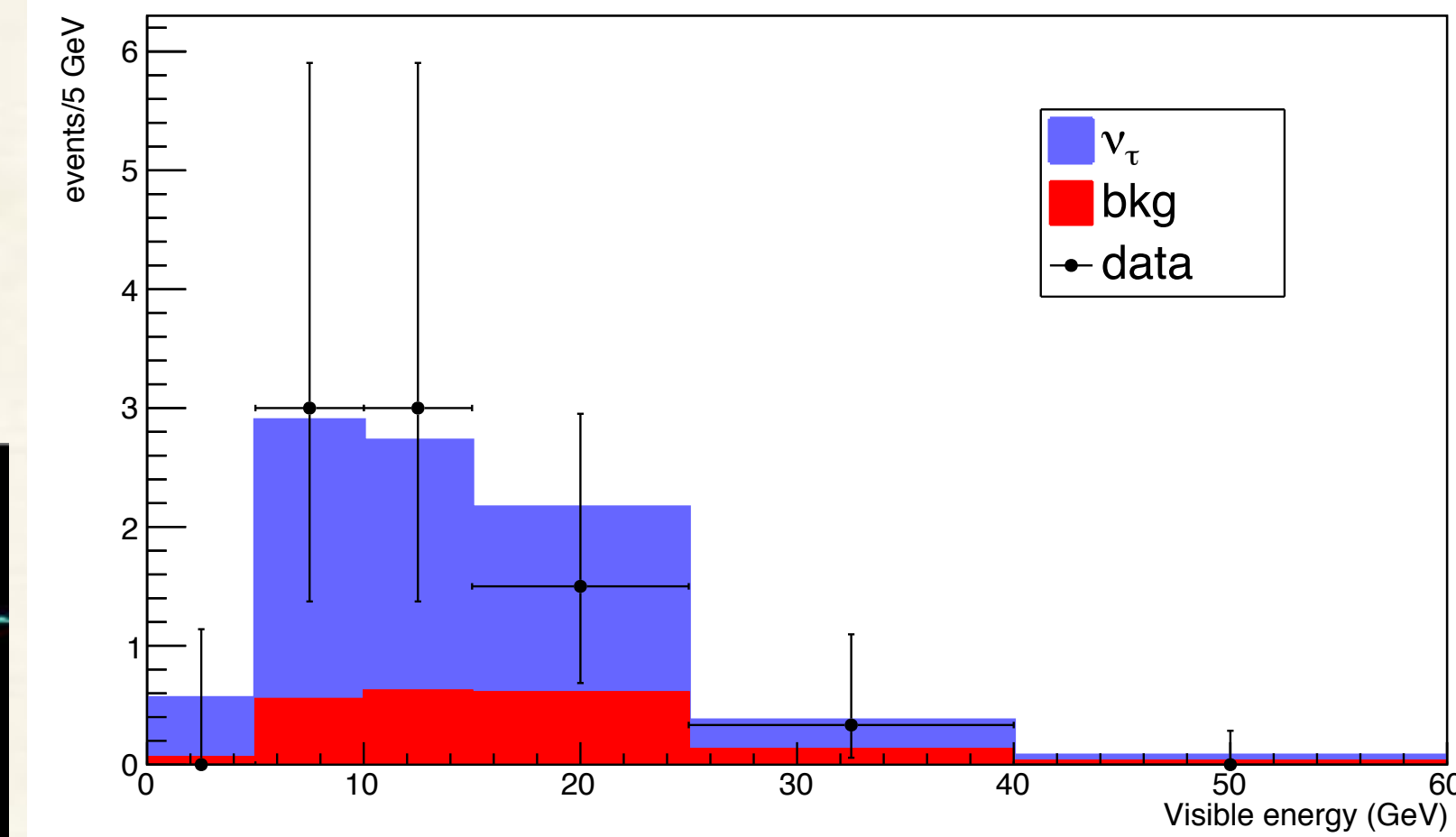
- ◆ Precise measurements of  $\theta_{23}$  and  $\Delta m^2_{32}$
- ◆ Consistency with three flavor prediction tightly constrains alternate oscillations hypotheses

# Final $\nu_\tau$ results from OPERA

2008-2012



PRL **120** (2018) 211801

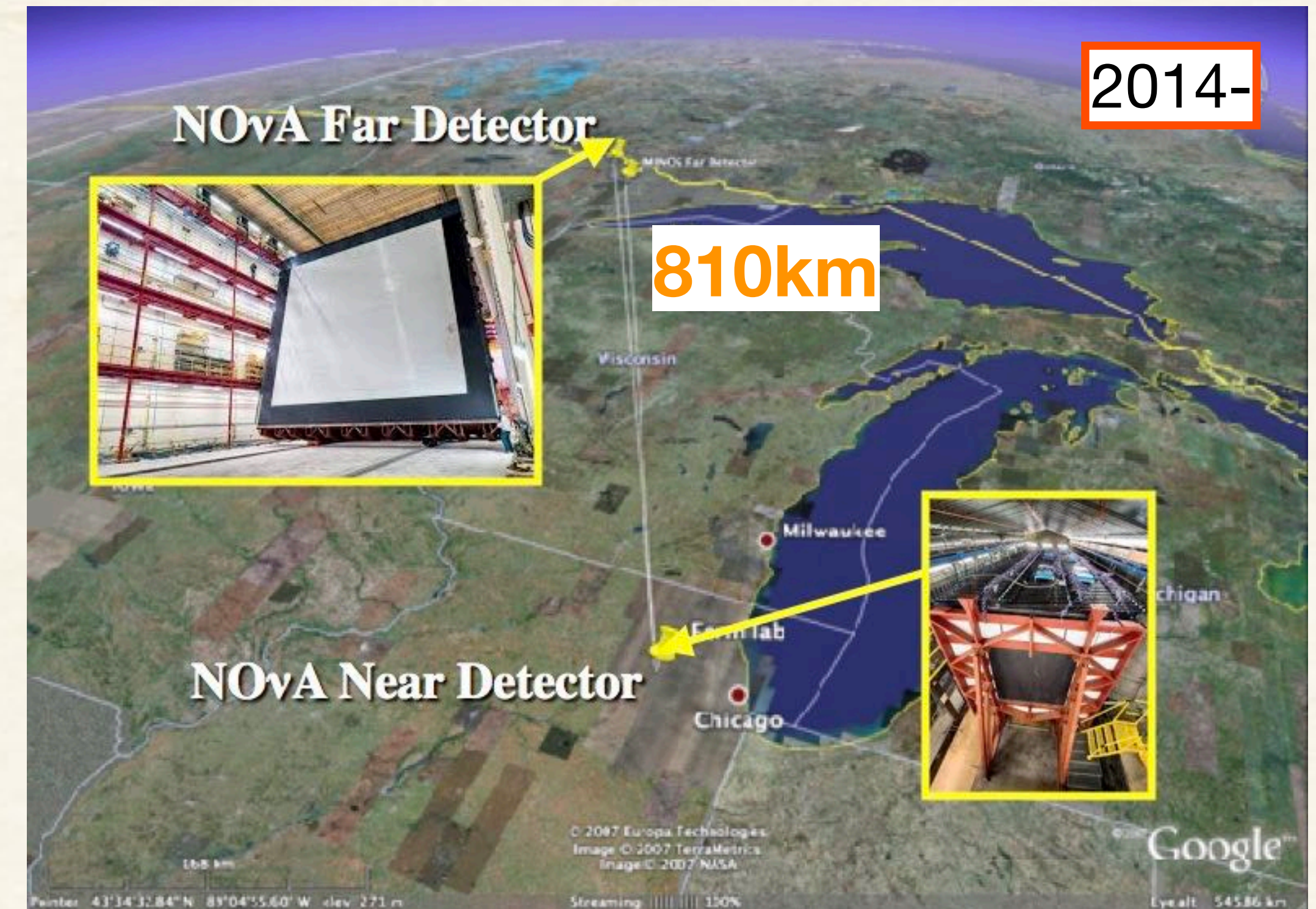
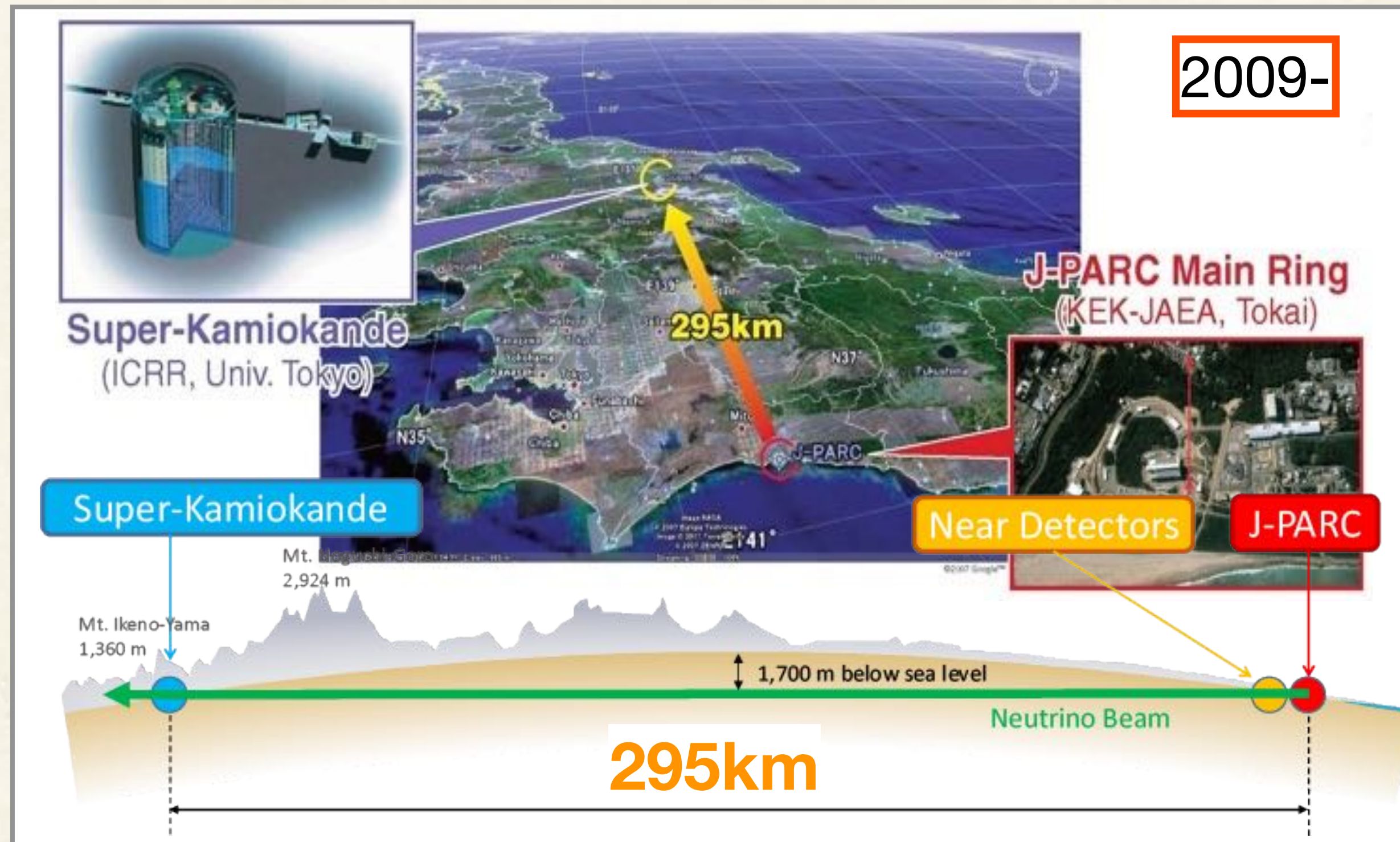


$\Delta m^2$  consistent with disappearance measurements

- Observation of  $\nu_\tau$  interaction using a huge emulsion-based detector
- **10  $\nu_\tau$  candidates observed**
  - $2.0 \pm 0.4$  BG expected
  - $6.1\sigma$  significance of  $\nu_\tau$  appearance



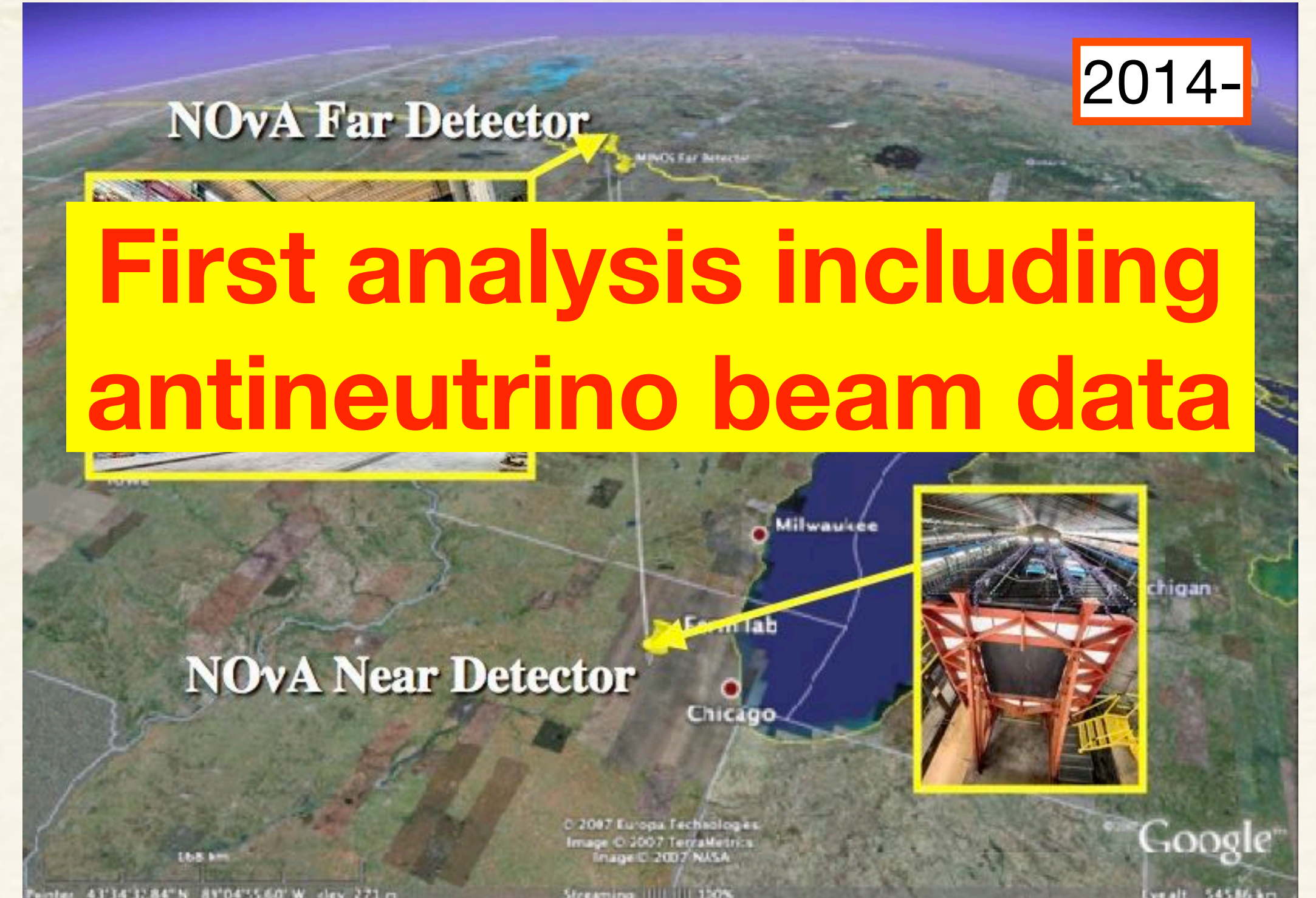
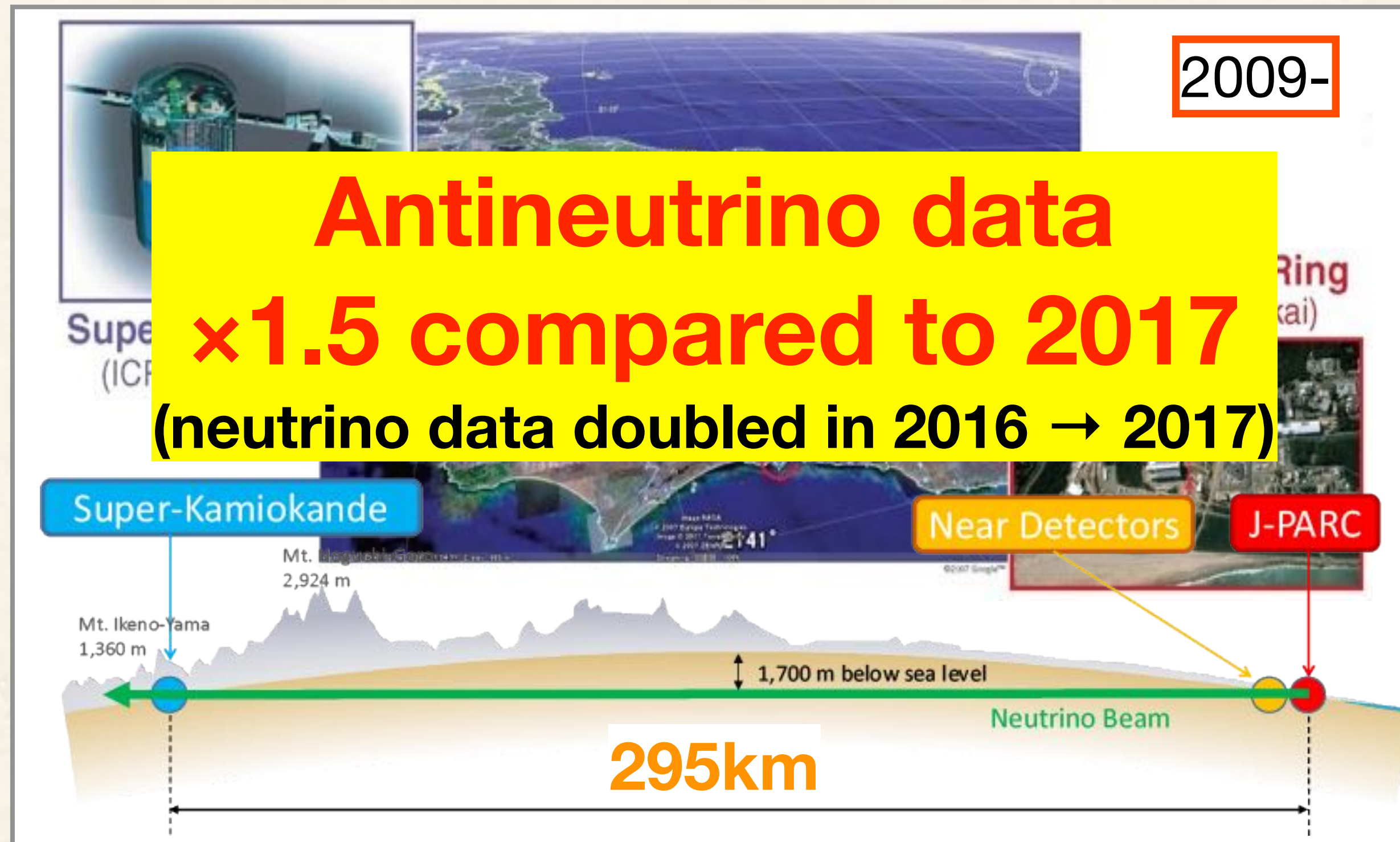
# T2K and NOvA



- ◆ Different baselines — different effect from matter effect (and possibly others not dependent on  $L/E$ )
- ◆ T2K has a shorter baseline, purer effect of CPV
- ◆ NOvA has a longer baseline, more matter effect and sensitivity to mass hierarchy
- ◆ Can provide complementary information

Parallel talks by P.Litchfield and J.Bian

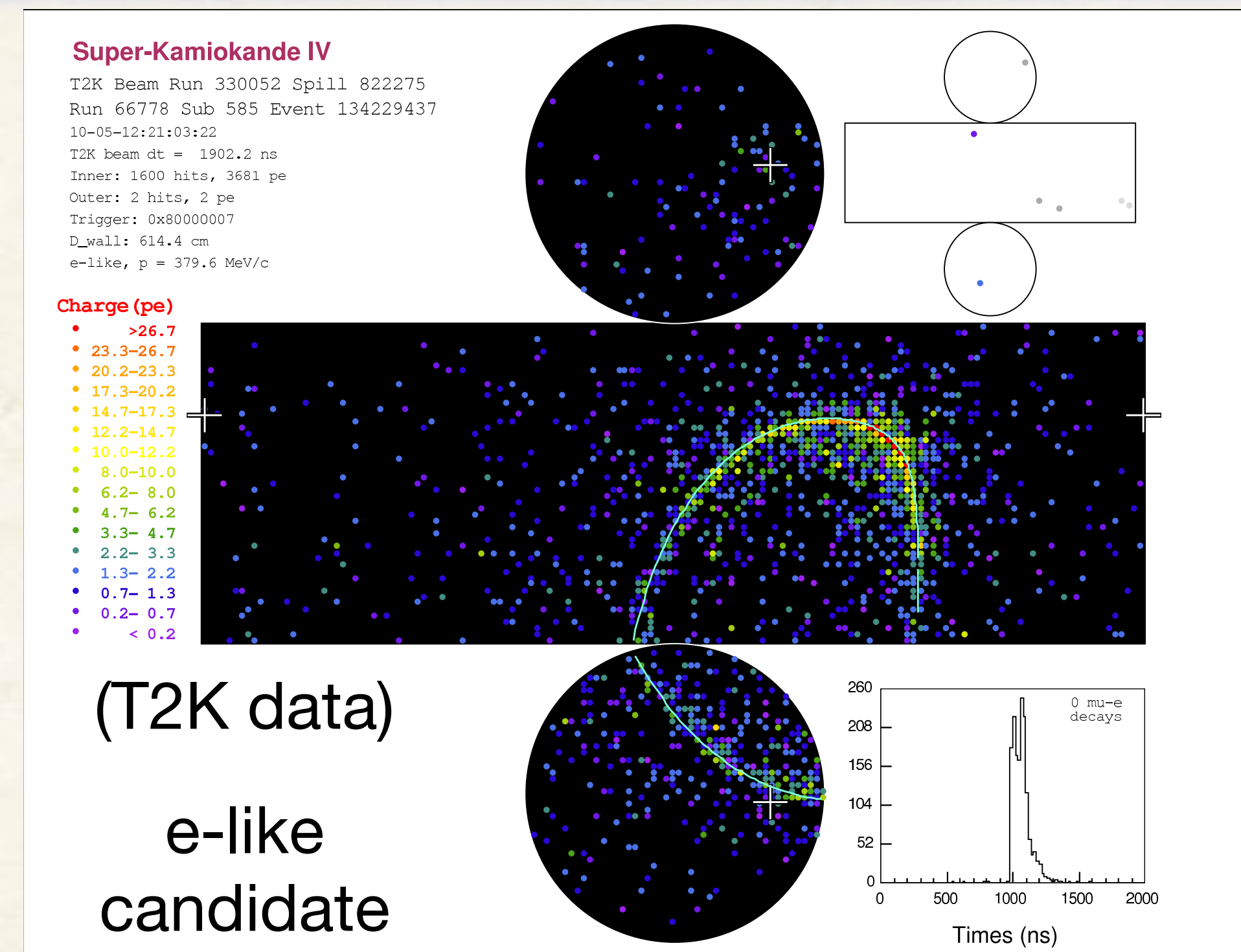
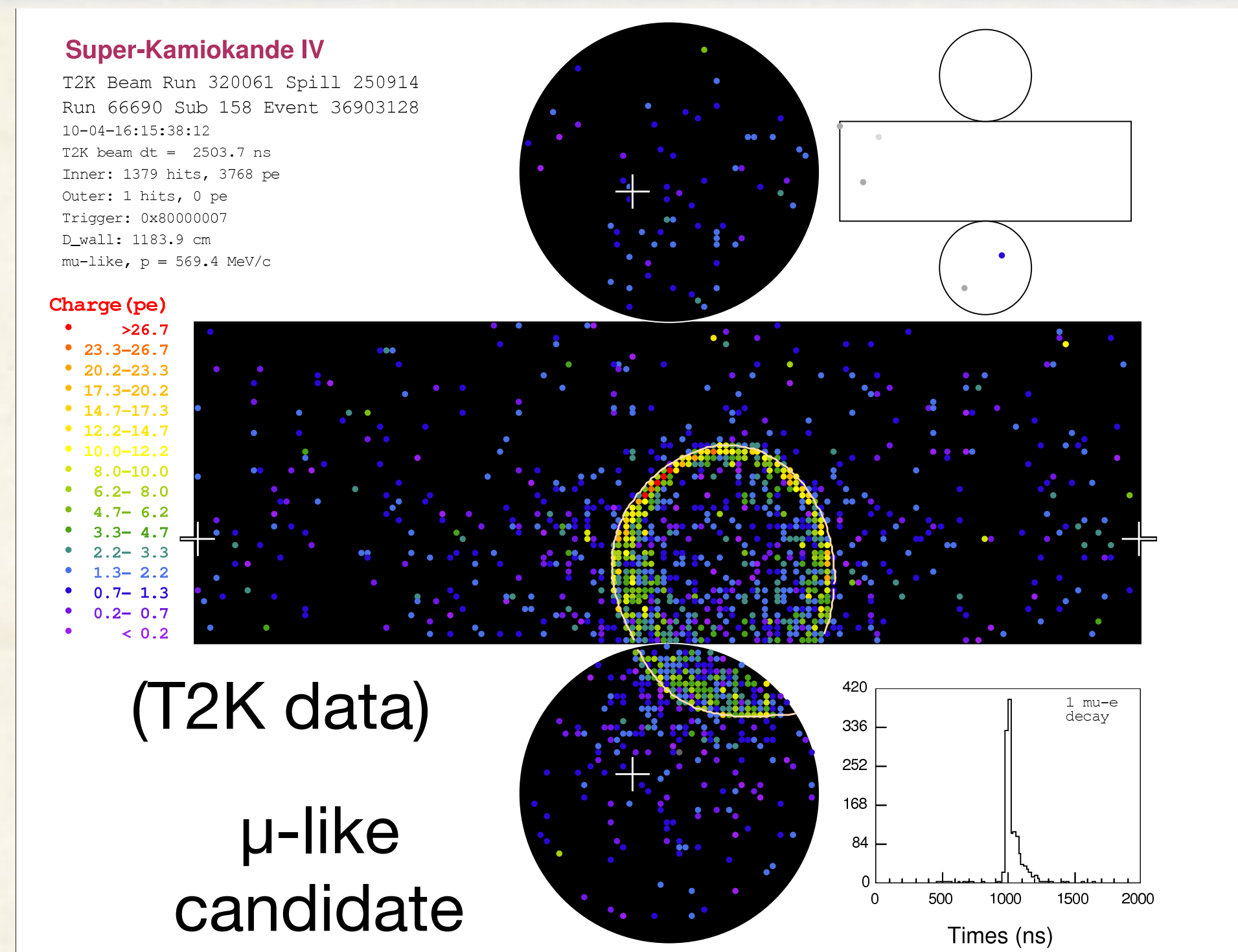
# T2K and NOvA



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Parallel talks by P.Litchfield and J.Bian

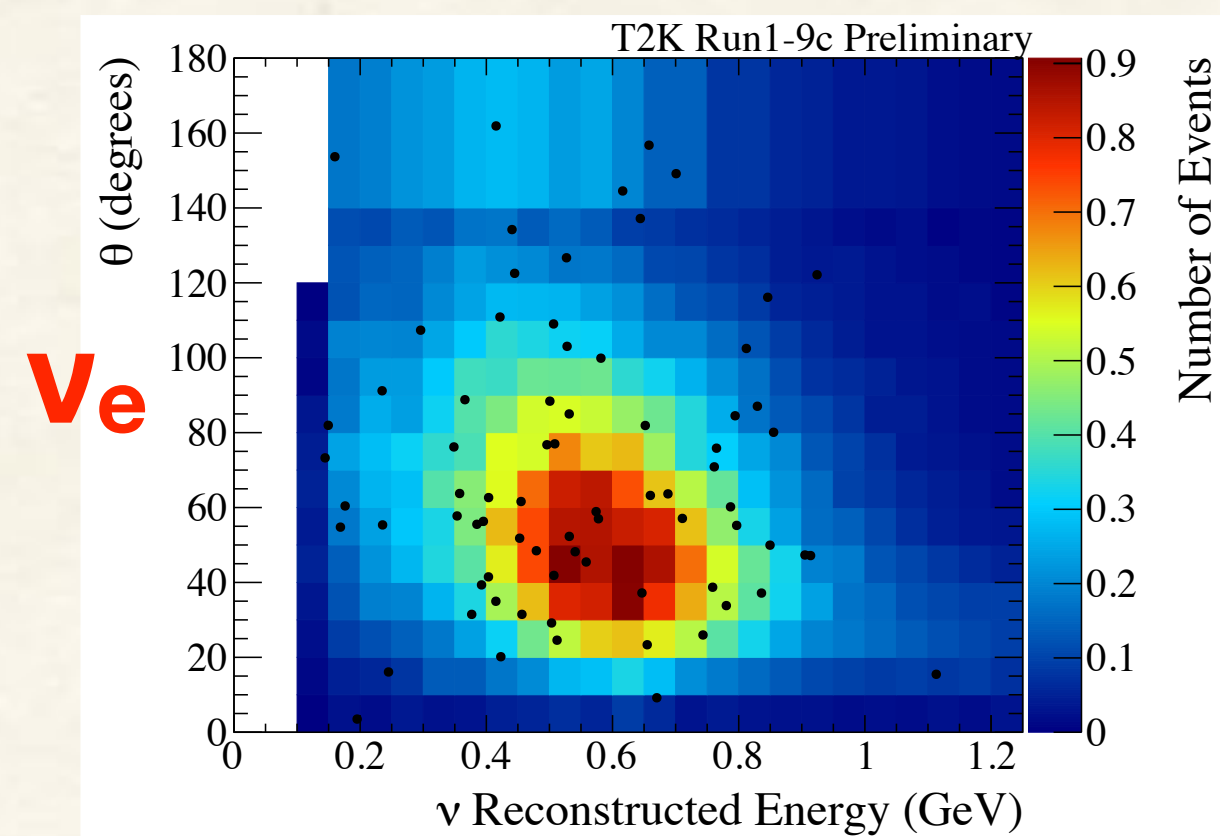
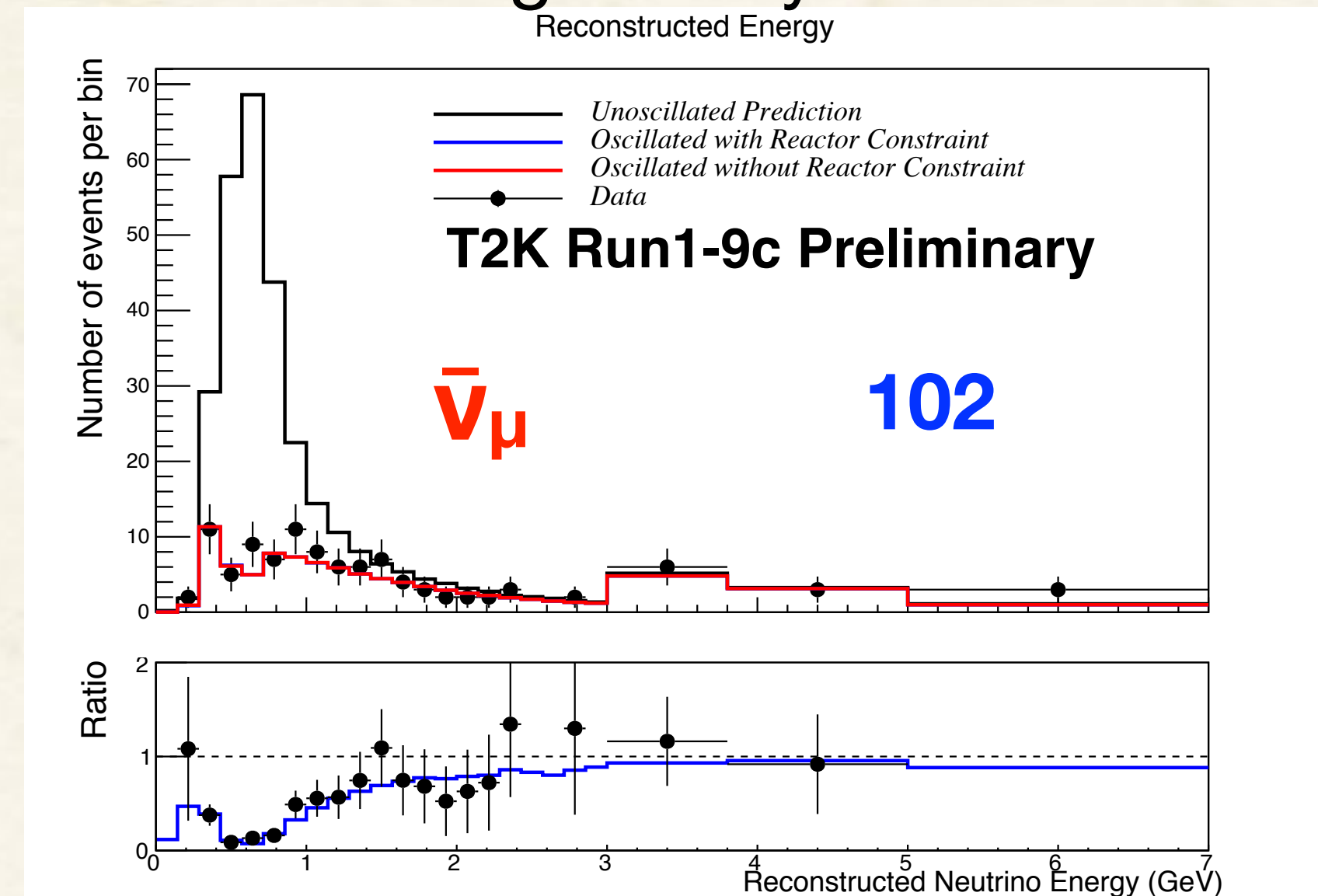
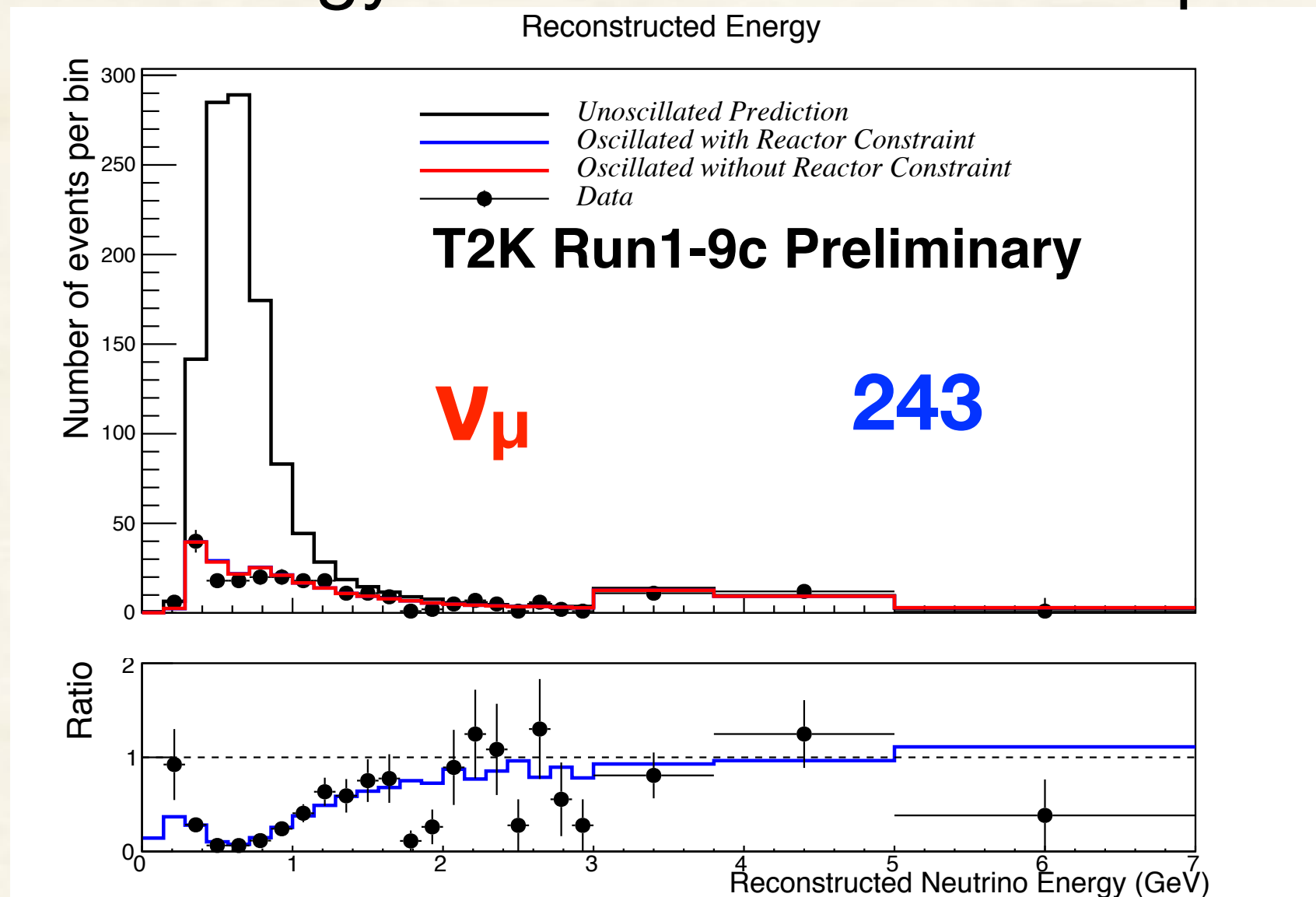
# T2K: reconstruction and analysis



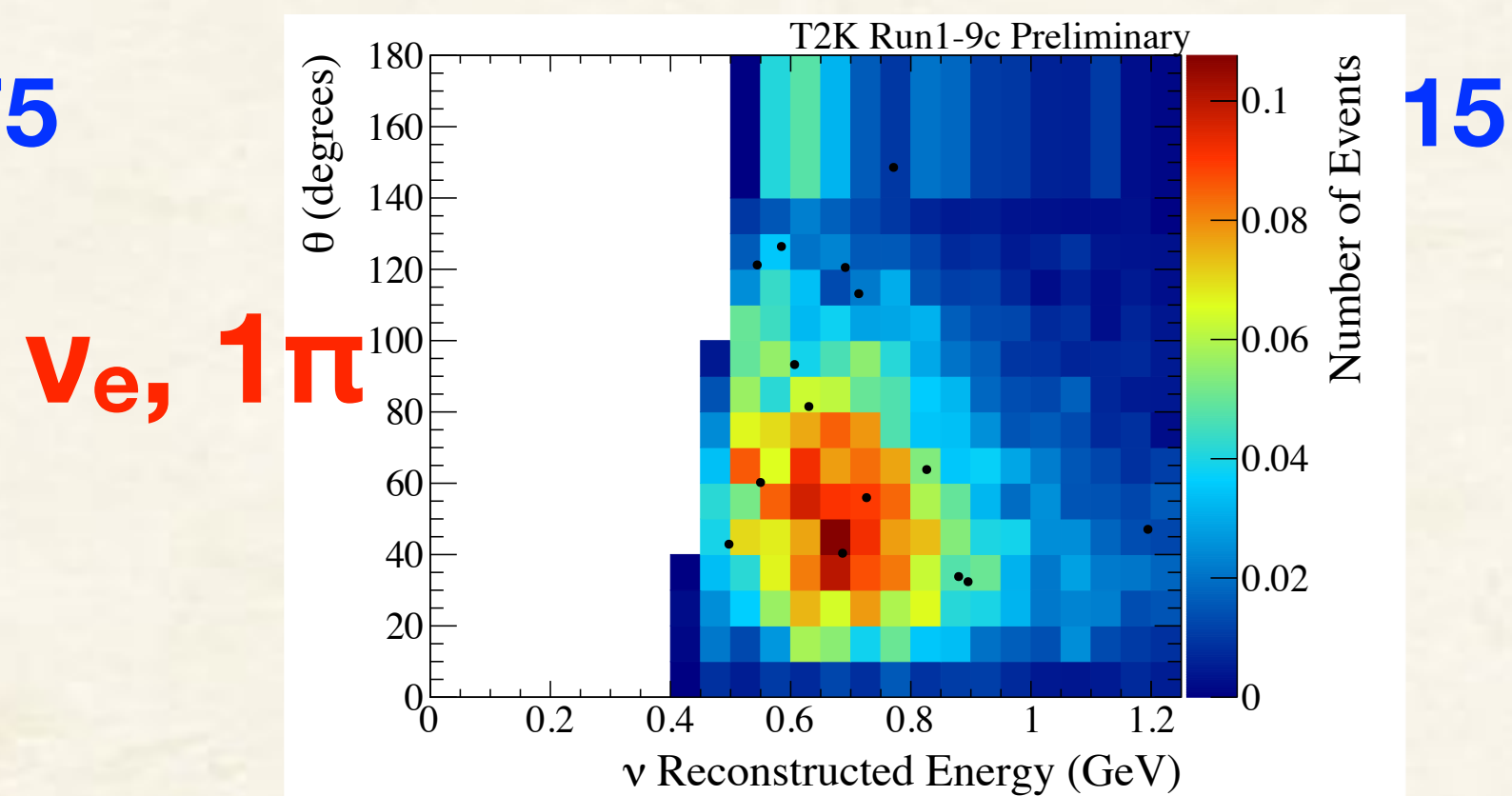
- ◆ **Improved reconstruction algorithm applied since 2017**
  - ◆ Maximize likelihood based on charge and time from all PMTs
  - ◆ Improved signal efficiency and purity
  - ◆ Optimized fiducial selection to increase statistics (+~20%)
- ◆ Improvements in cross section modeling

# T2K: far detector (SK) data

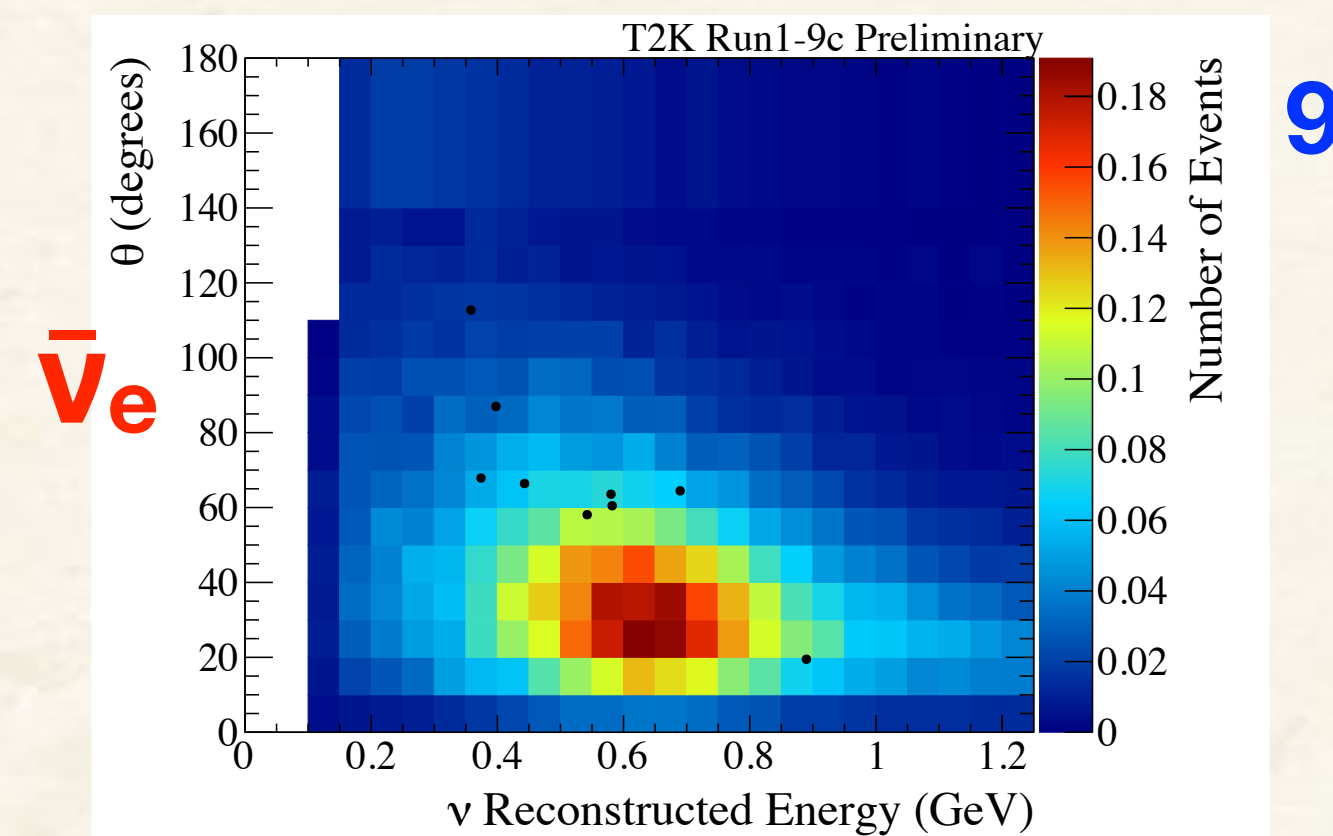
Energy reconstructed from lepton kinematics assuming 2 body interaction



$\nu$  beam, 1 ring e, no decay-e

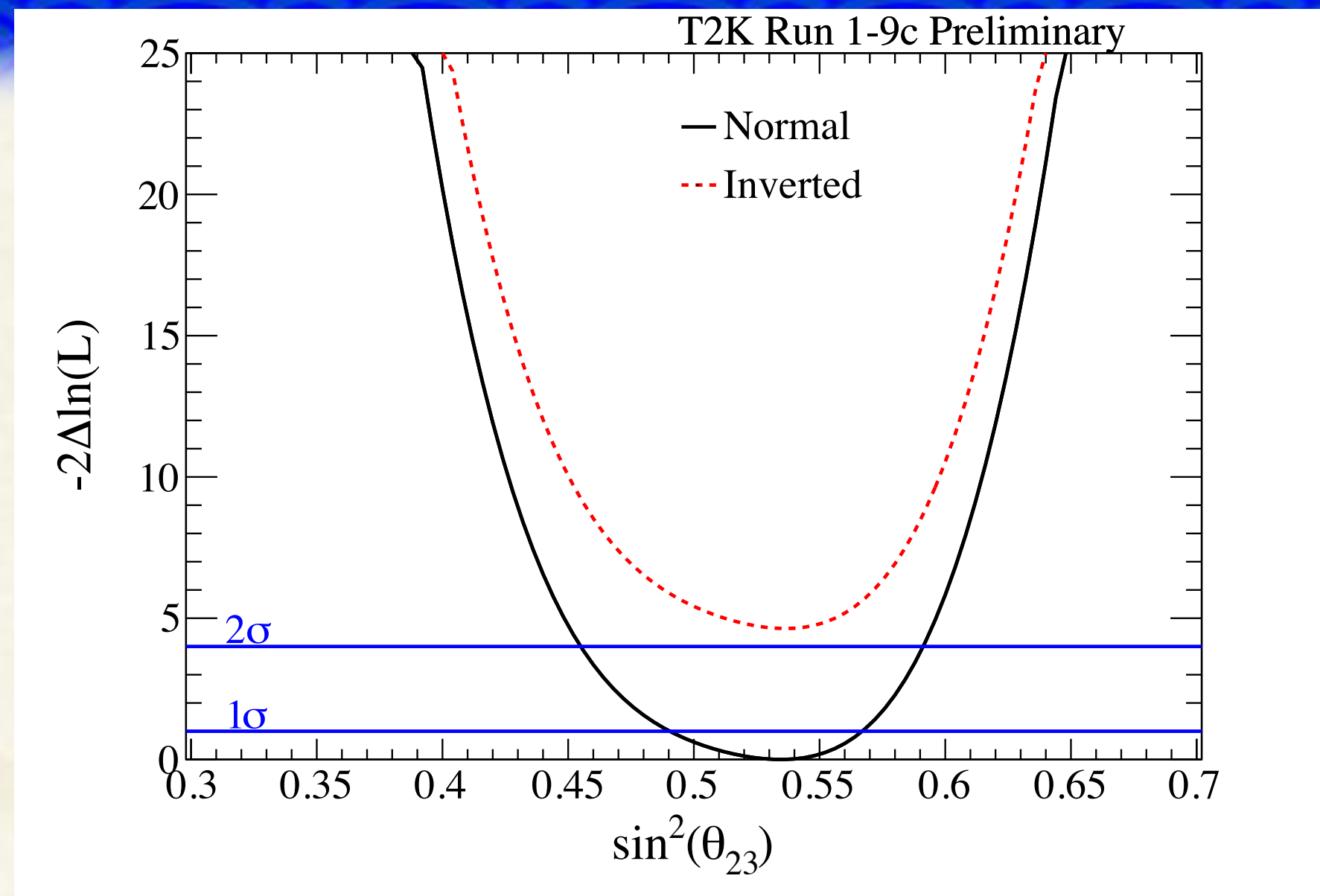


$\nu$  beam, 1 ring e, 1 decay-e



$\bar{\nu}$  beam, 1 ring e, no decay-e

# T2K: $\Delta m_{32}^2$ and $\theta_{23}$



Joint fit of all samples

Normal hierarchy:

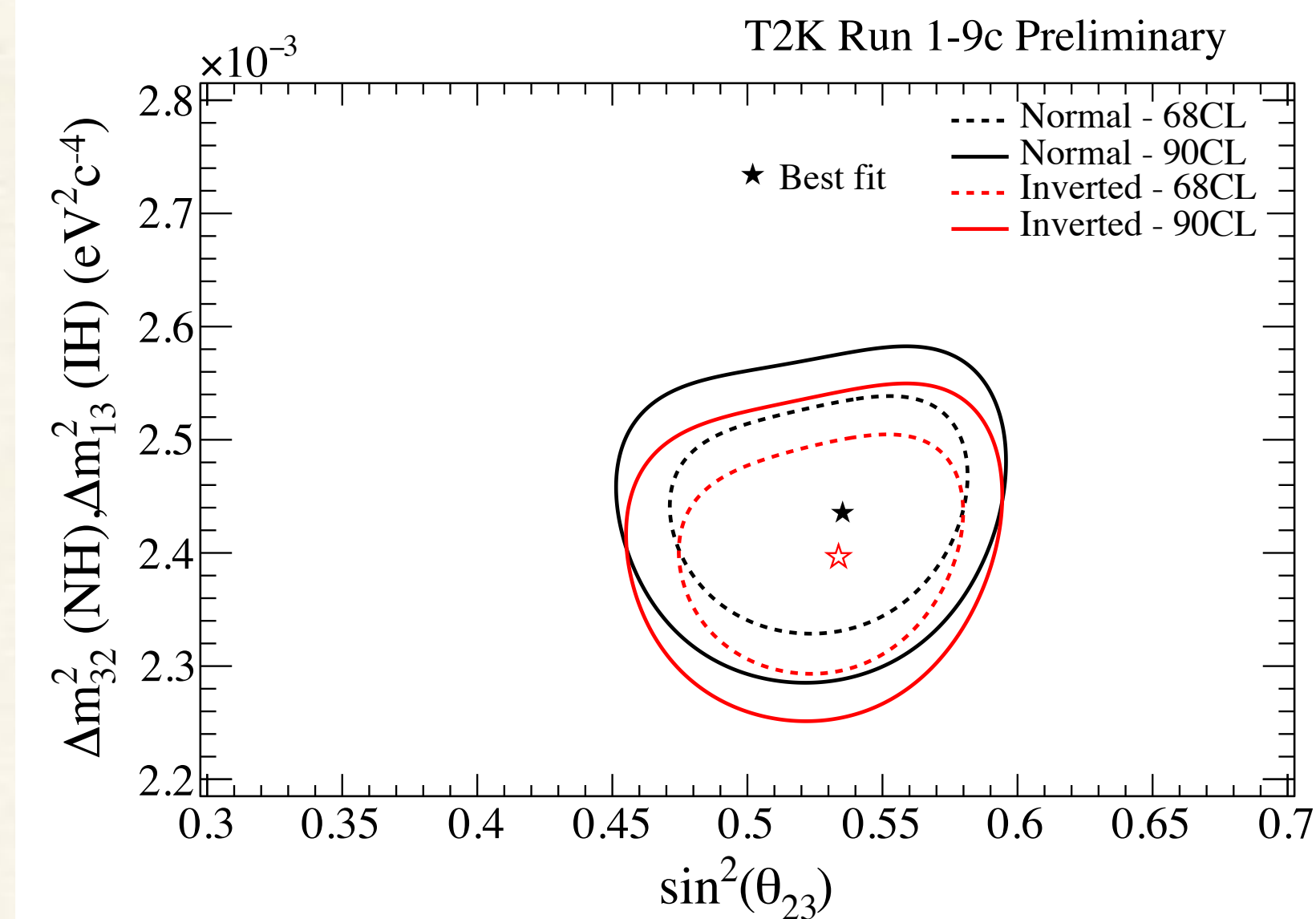
$$\sin^2 \theta_{23} = 0.536^{+0.031}_{-0.046}$$

$$\Delta m_{32}^2 = (2.434 \pm 0.064) \times 10^{-3} \text{eV}^2$$

Inverted hierarchy:

$$\sin^2 \theta_{23} = 0.536^{+0.031}_{-0.041}$$

$$\Delta m_{13}^2 = (2.410^{+0.062}_{-0.063}) \times 10^{-3} \text{eV}^2$$



Posterior probabilities based on a Bayesian analysis

	$\sin^2\theta_{23}<0.5$	$\sin^2\theta_{23}>0.5$	Sum
NH	0.204	0.684	0.888
IH	0.023	0.089	0.112
Sum	0.227	0.773	1

Bayes factor NH/IH = 7.9

# $\bar{\nu}_e$ appearance search in T2K

- ◆ Dedicated searches performed by hypothesis testing:  
PMNS  $\bar{\nu}_e$  appearance ( $\beta=1$ ) and no  $\bar{\nu}_e$  appearance ( $\beta=0$ )

Expected events: 11.8 for  $\beta=1$   
6.5 for  $\beta=0$   
Observed: 9 events

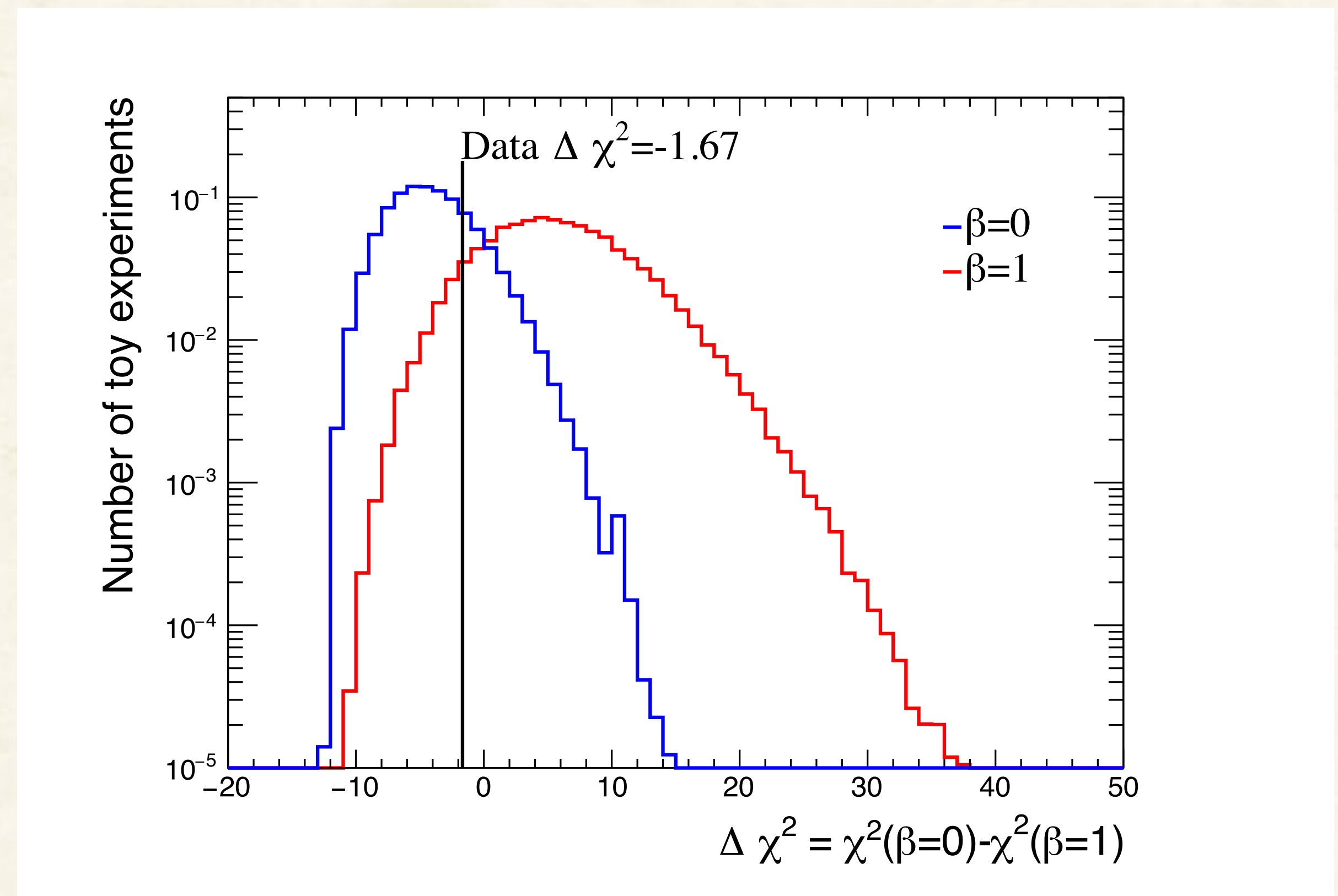
Construct  $\Delta\chi^2$  with rate+shape ( $p, \theta$ )

other oscillation parameters constrained from  
T2K data other than  $\bar{\nu}_e$  sample

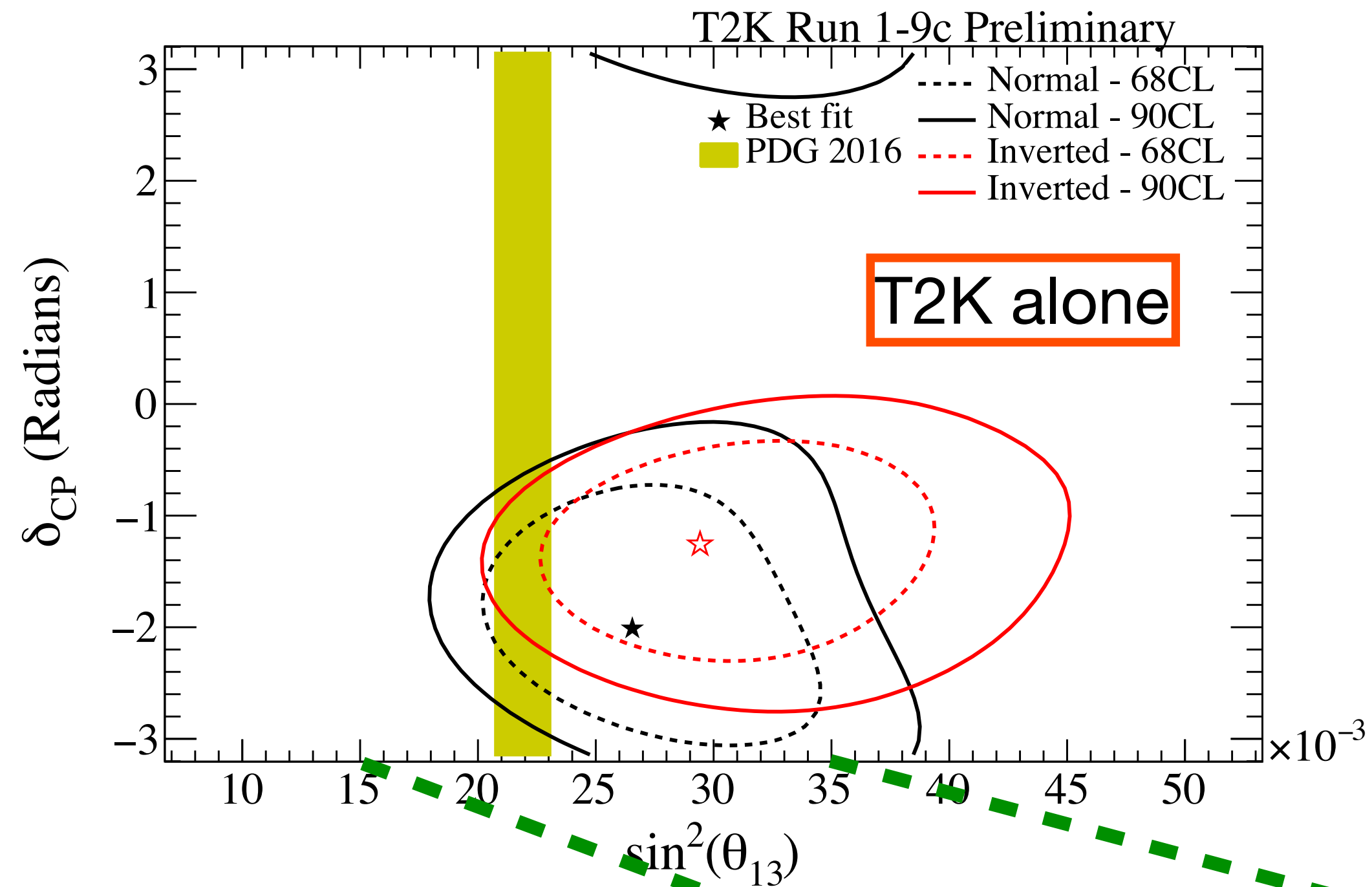
Obtained p-values:

PMNS appearance ( $\beta=1$ ):  $p=0.0867$   
No appearance ( $\beta=0$ ):  $p=0.233$

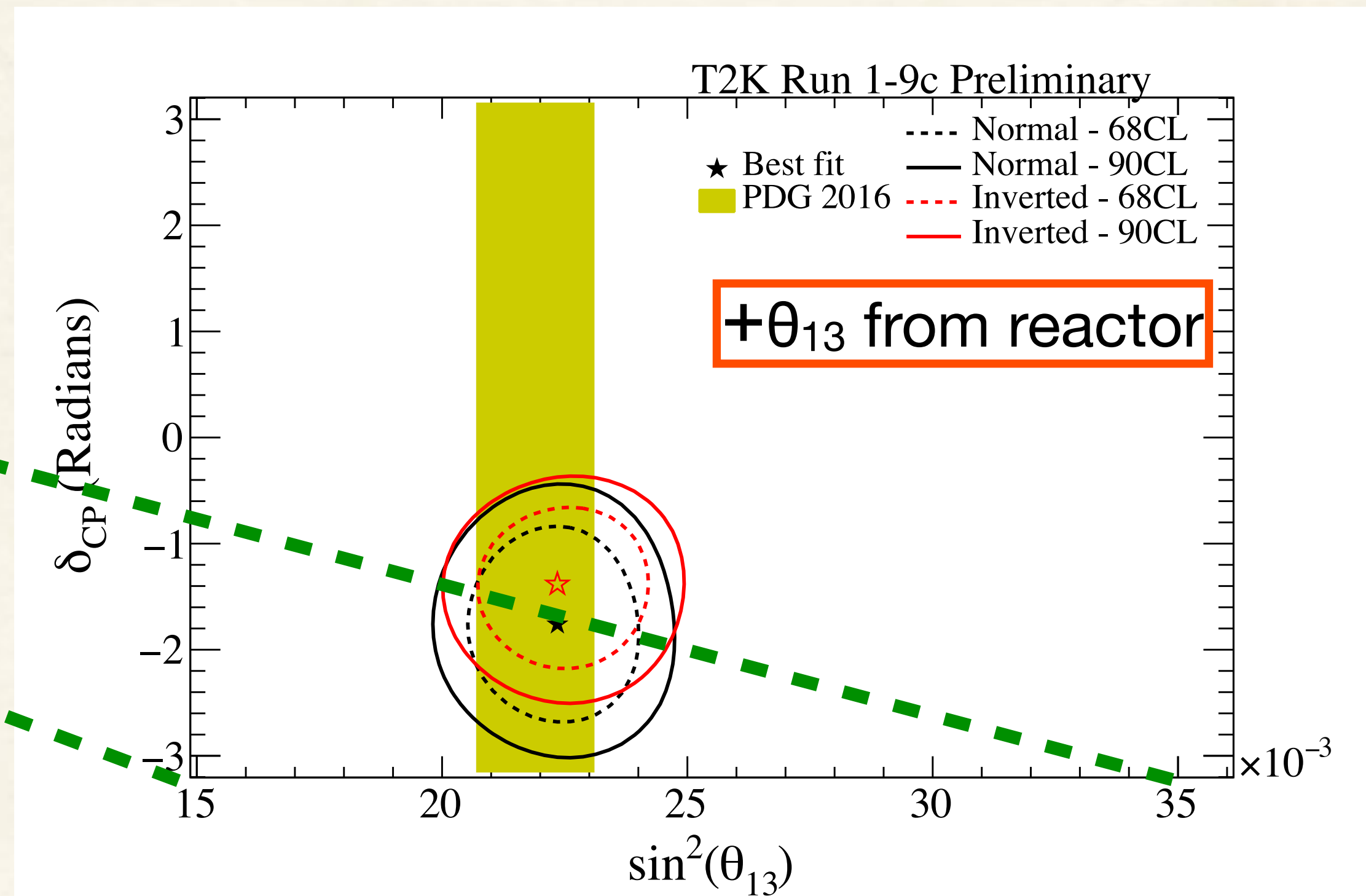
**No strong statistical conclusion from T2K data yet**



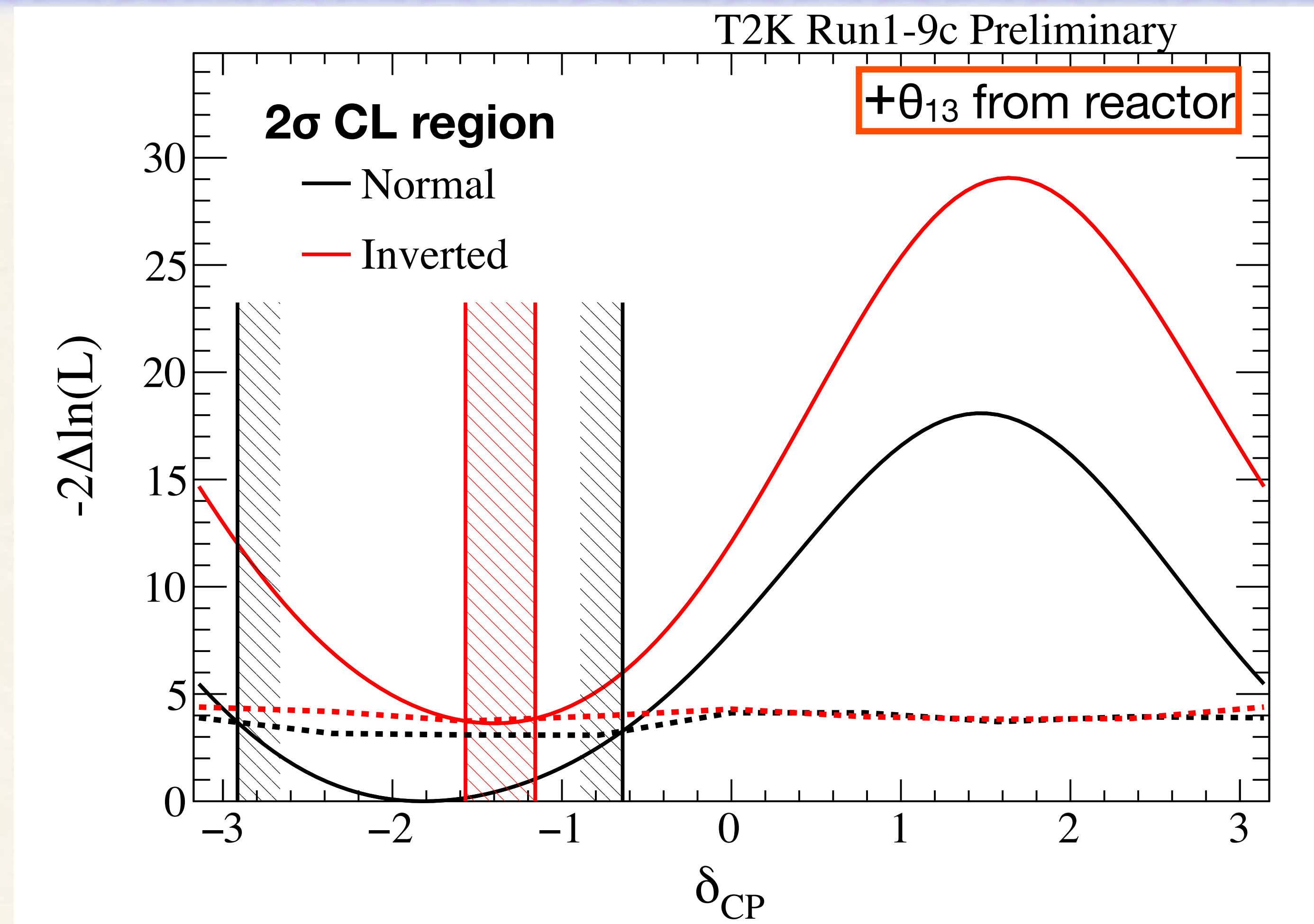
# T2K: $\theta_{13}$ and $\delta_{CP}$



- ◆ **Constraint on  $\delta_{CP}$  with T2K data alone**
- ◆ **Tighter constraint with  $\theta_{13}$  value from reactor**



# T2K: constraint on $\delta_{CP}$

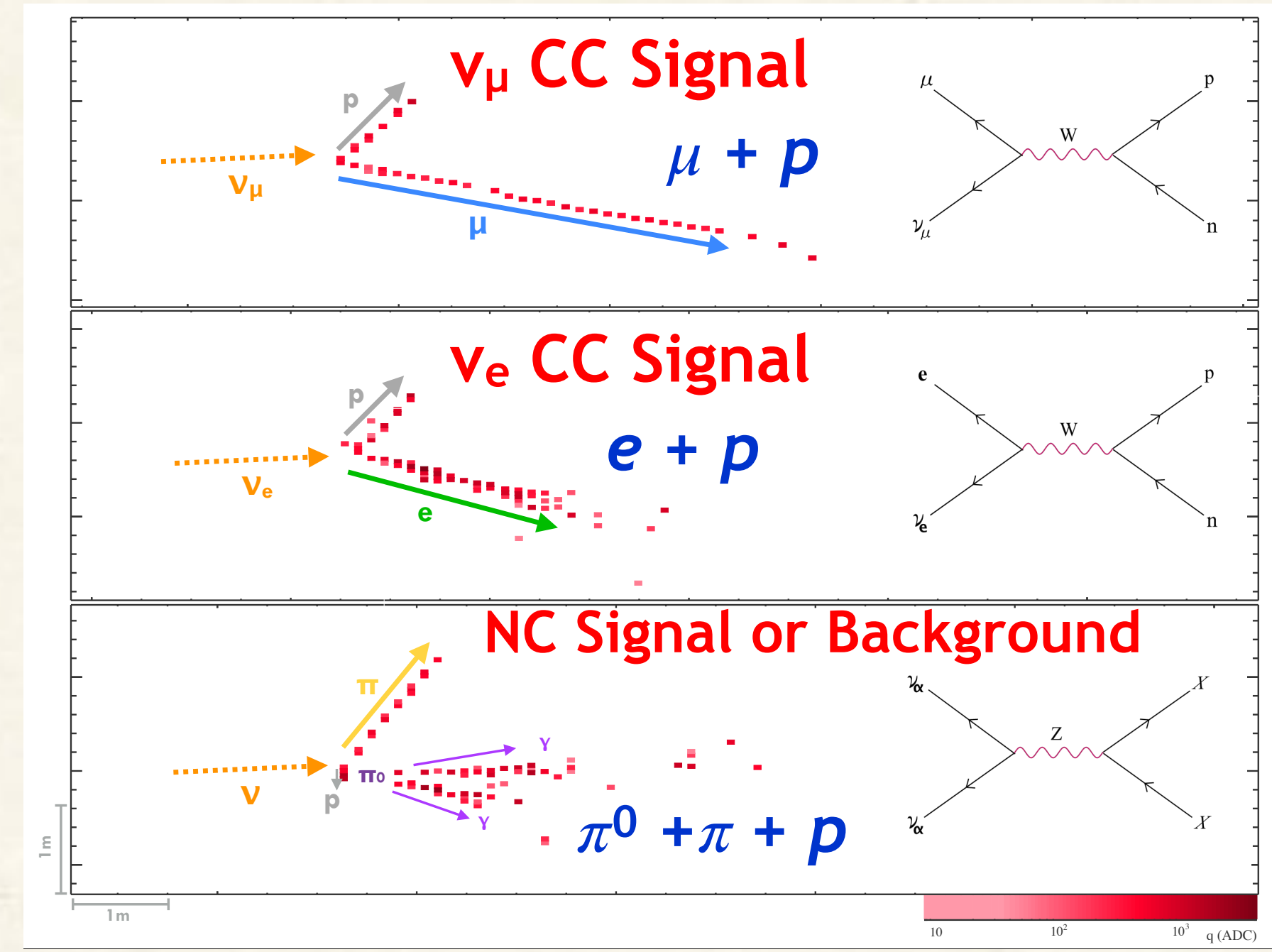
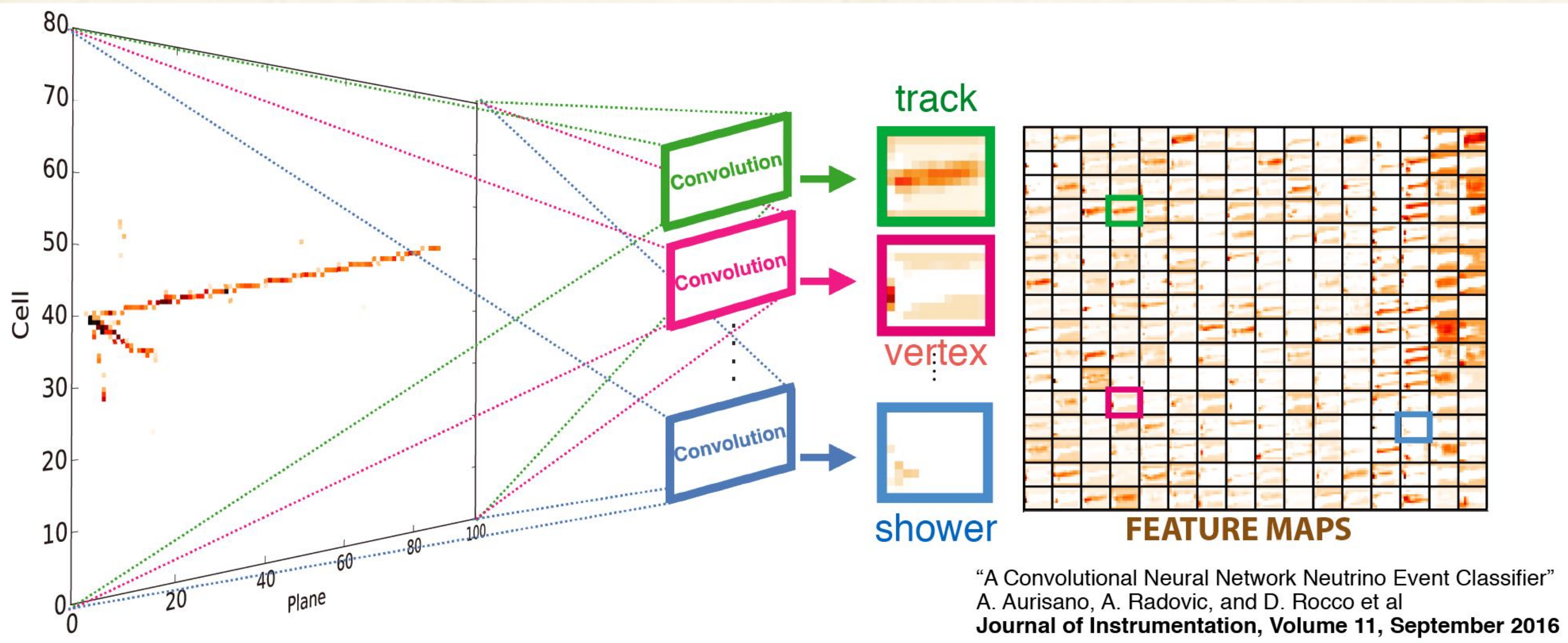


**$\sin\delta_{CP}=0$  ( $\delta=0, \pi$ ) outside of  $2\sigma$  CL region**  
**First hint of CP violation in the lepton sector!**



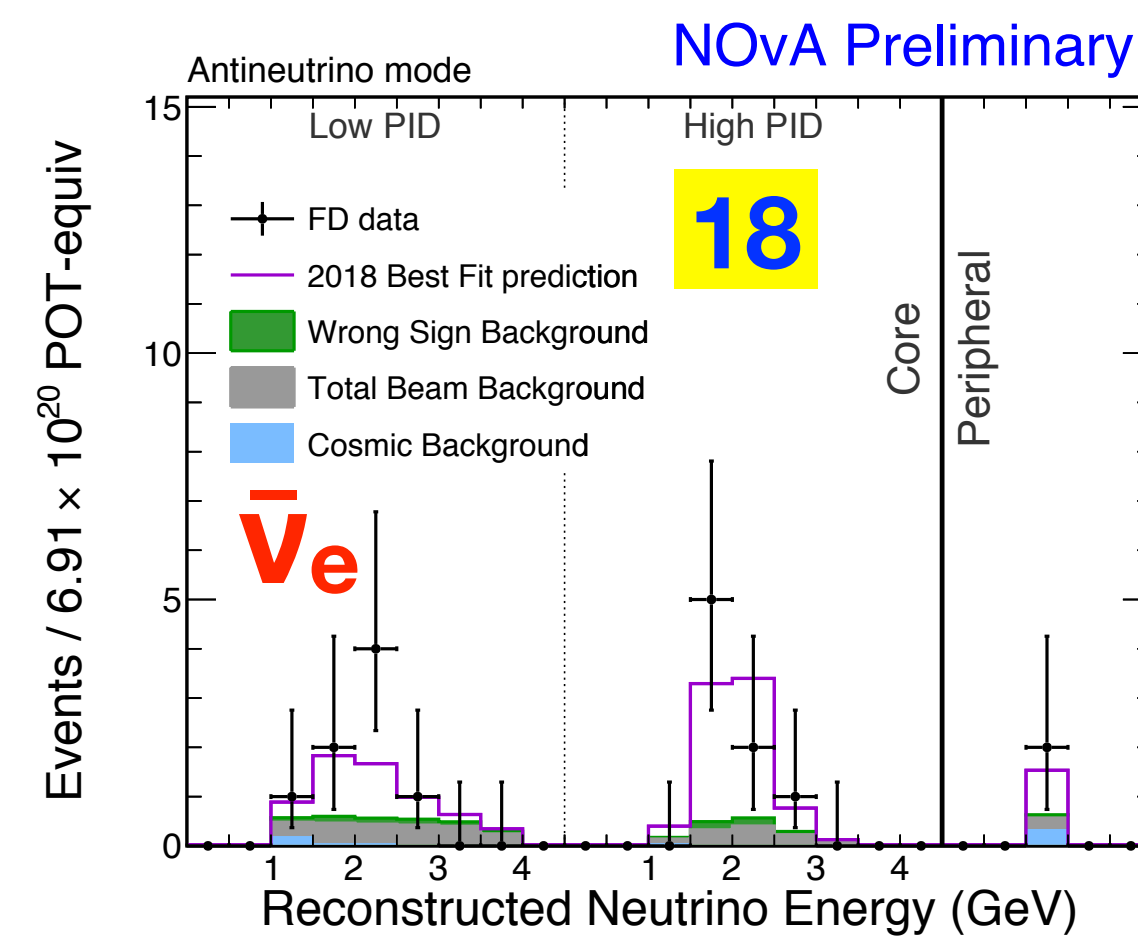
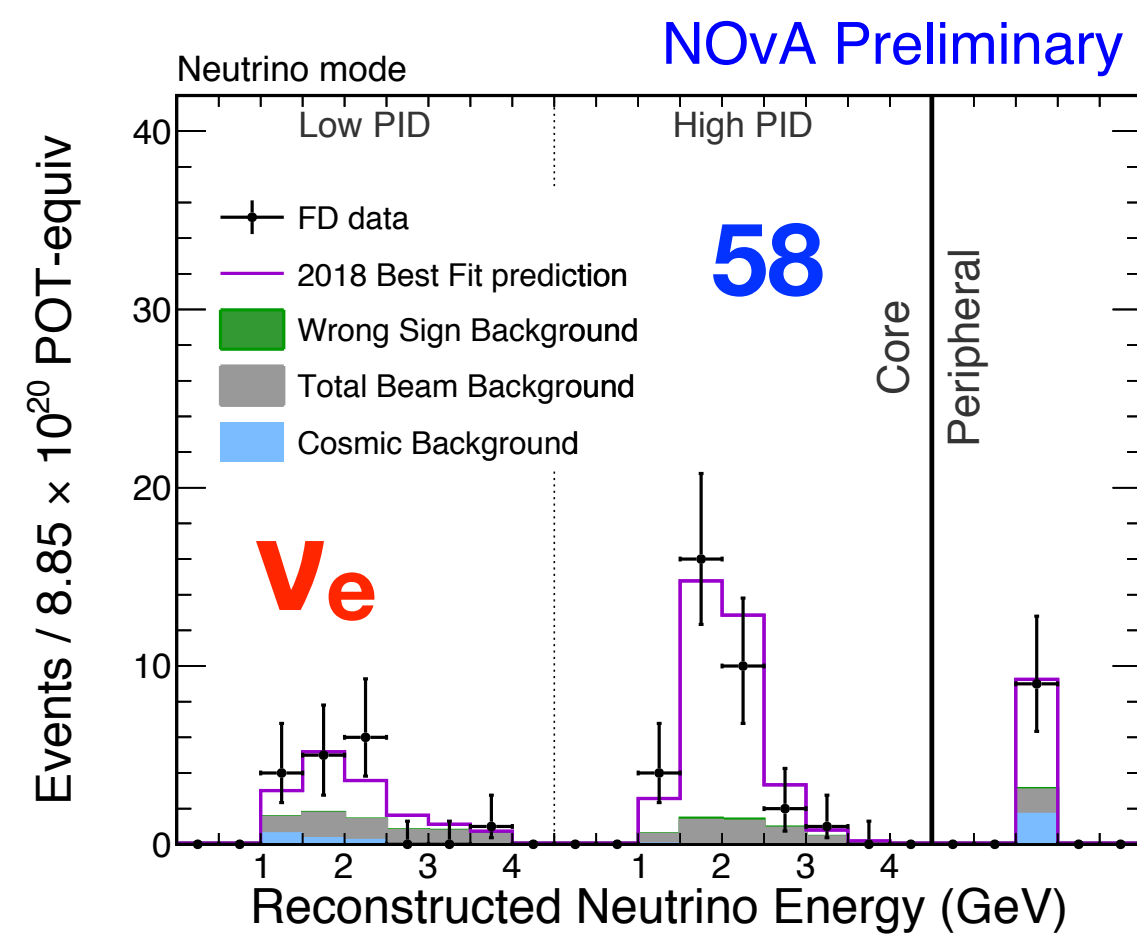
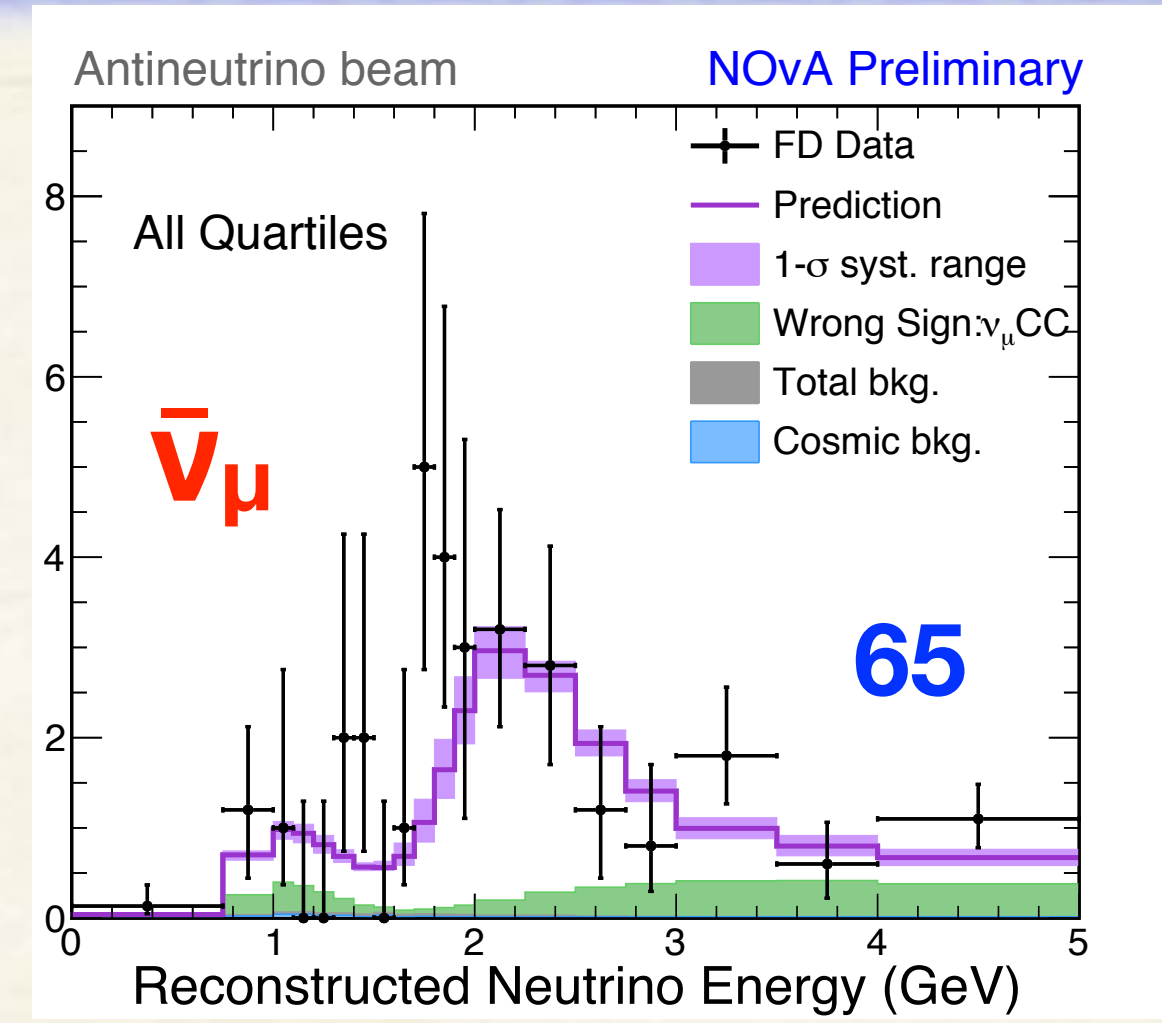
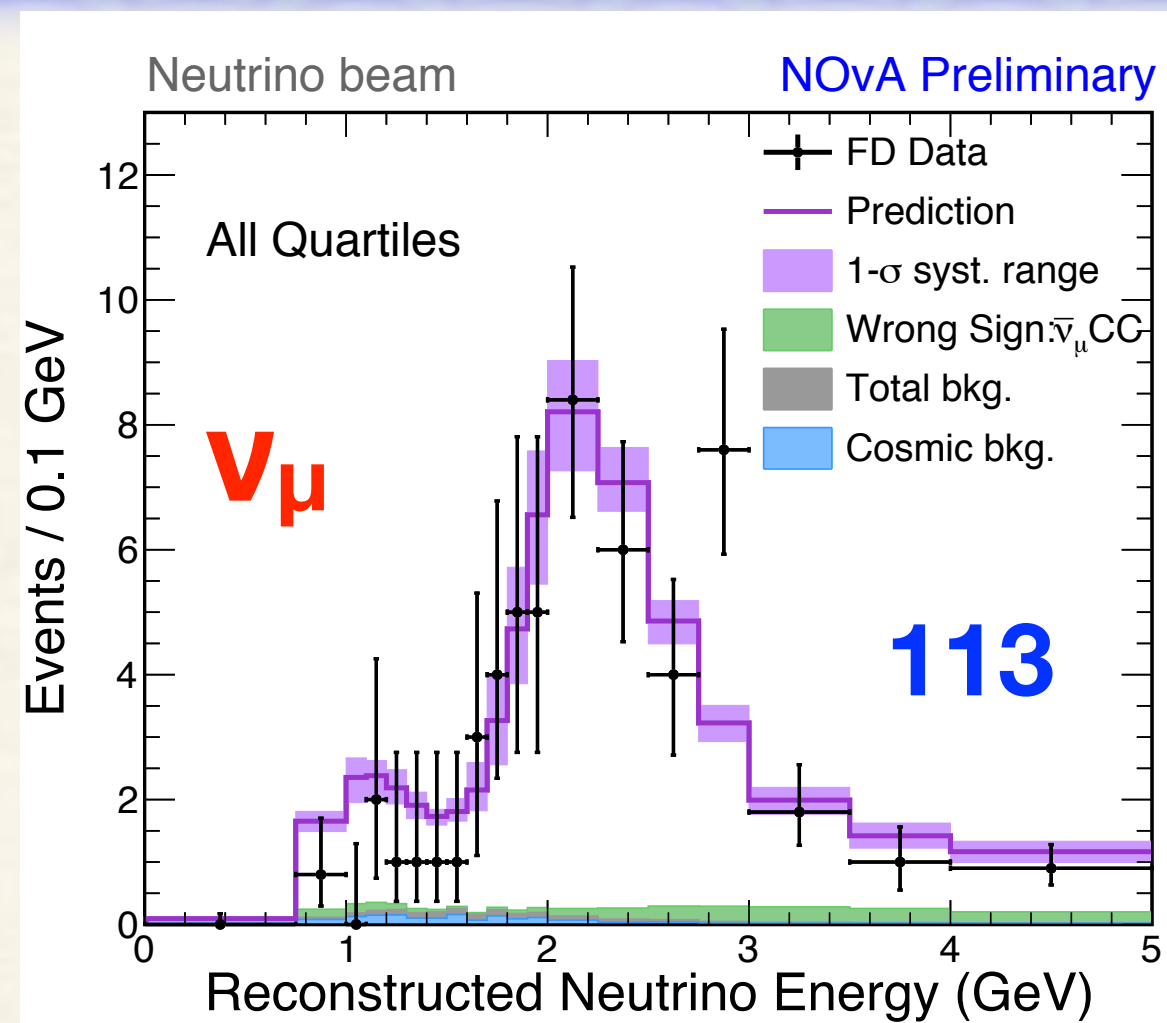
# NOvA event classification

- Pioneering the use of **Convolutional Neural Networks** in neutrino experiment
- Treating cells as pixels and charge as color, extract features from data
- Improved classifier used for 2018 analysis
  - with separate training for the neutrino and antineutrino beams

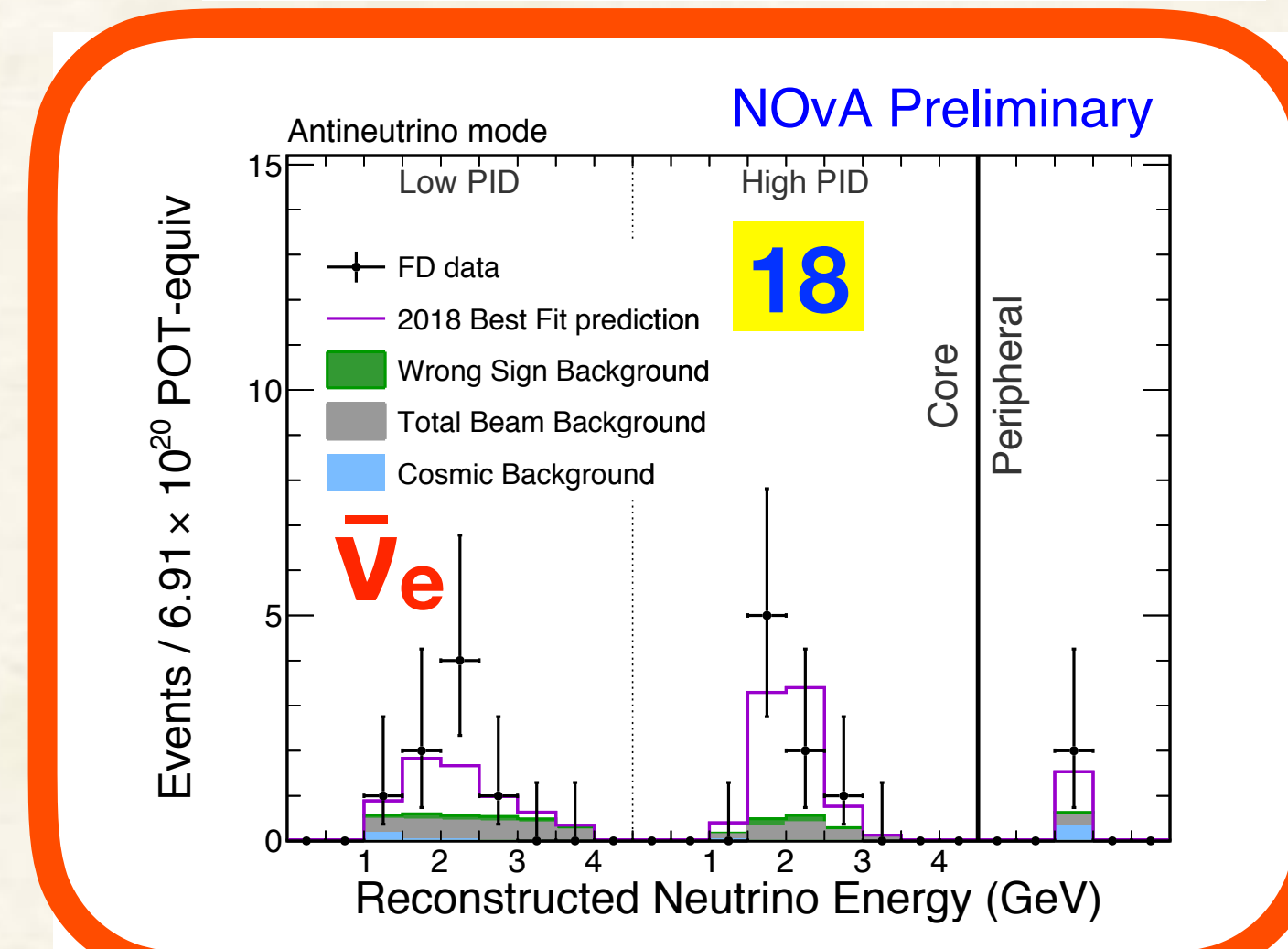
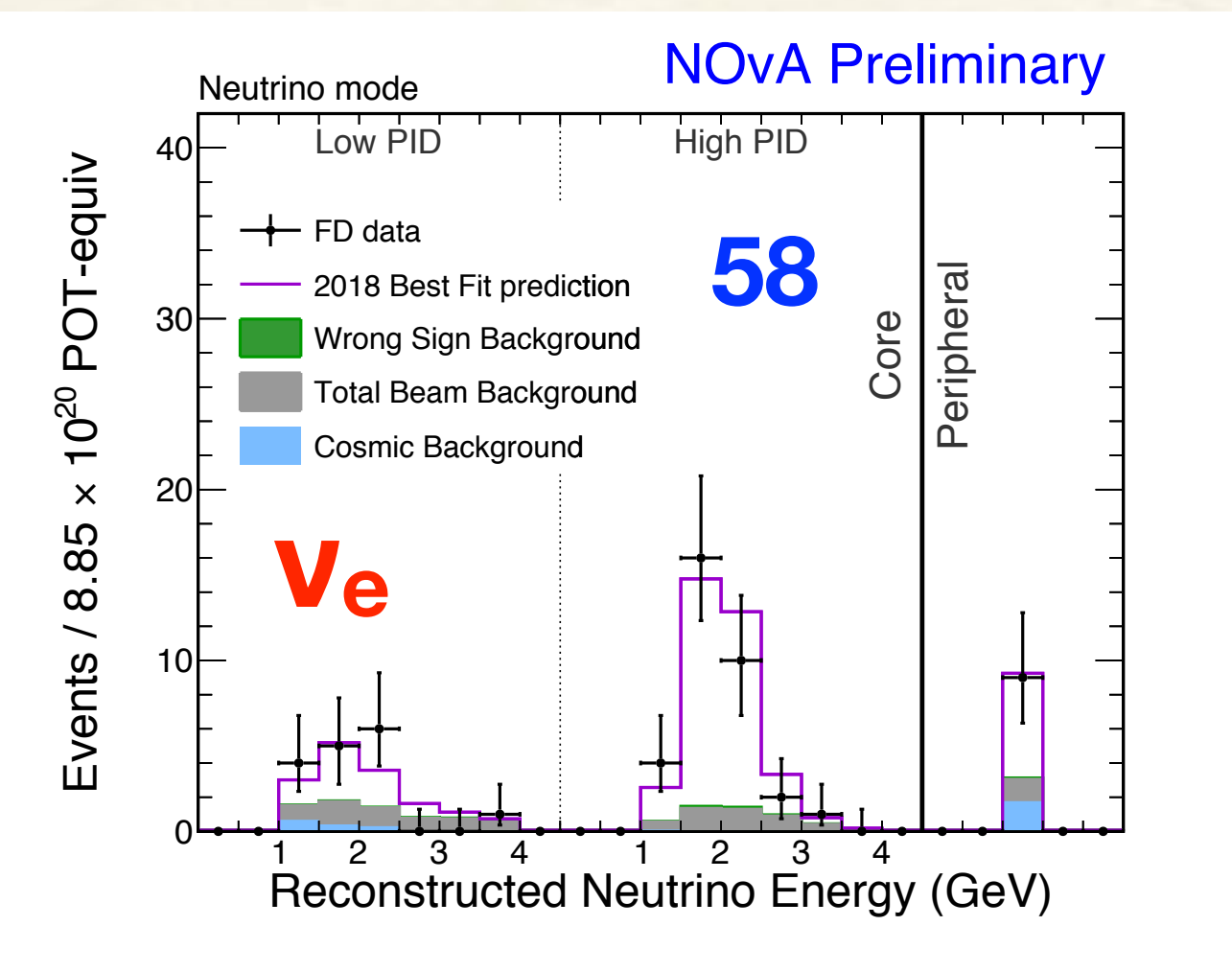
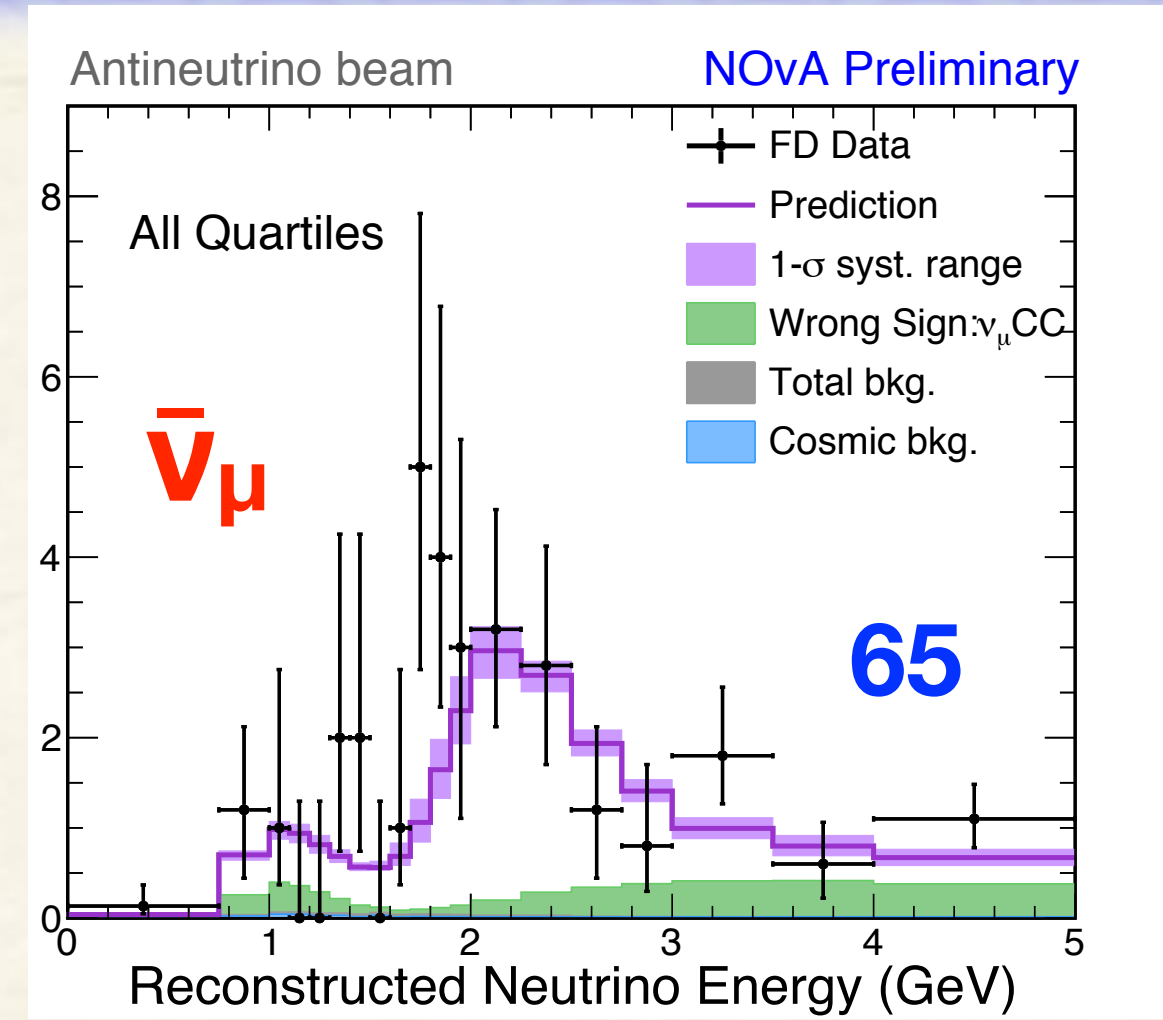
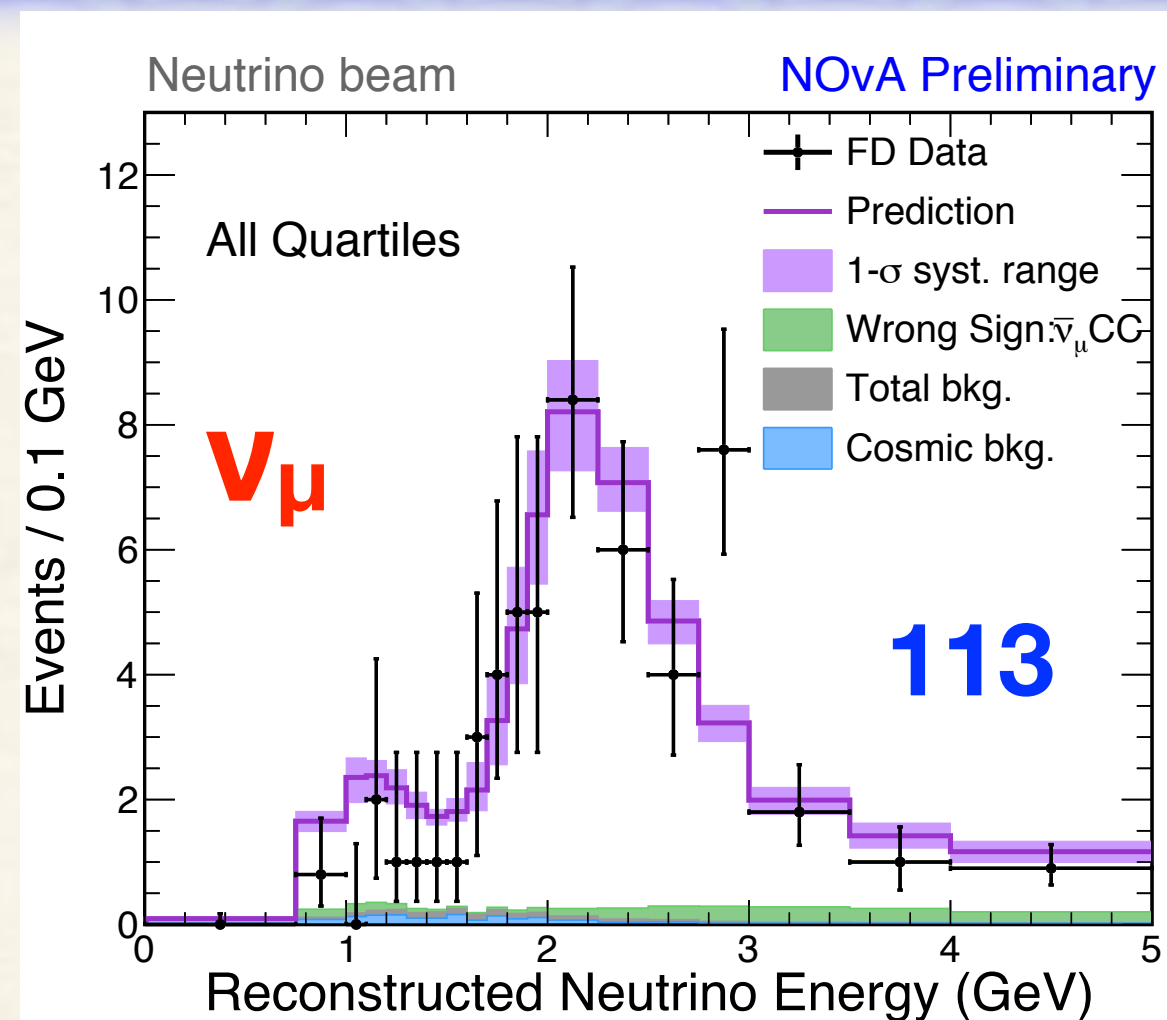


- Neutrino energy reconstruction from  $E_l$  (range[ $\mu$ ]/calorimetric[e]) +  $E_{had}$  (calorimetric)

# NOvA FD data

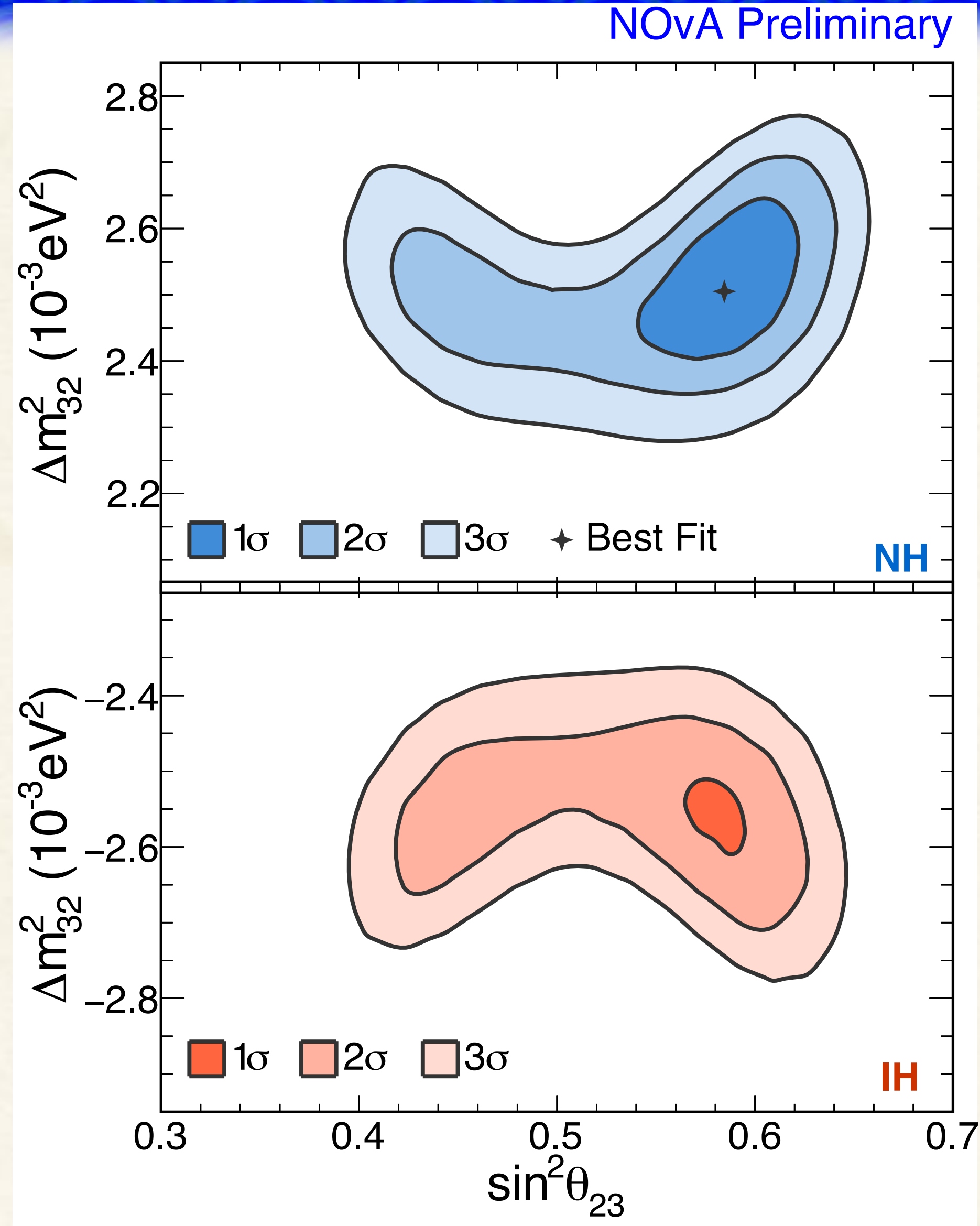


# NOvA FD data



**>4 $\sigma$  evidence of  $\bar{\nu}_e$  appearance!**

# NOvA: $\Delta m_{32}^2$ and $\theta_{23}$

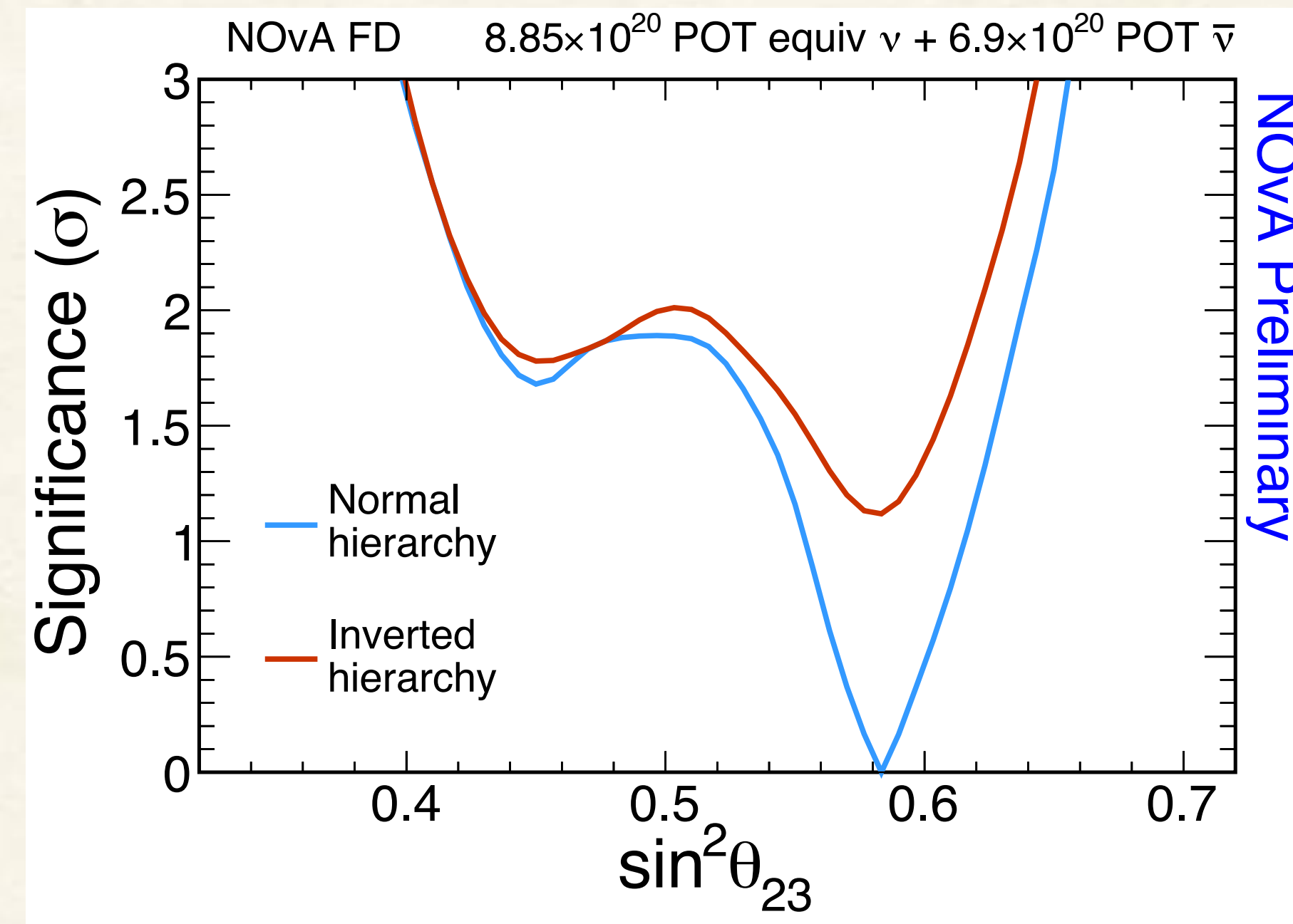


- Results from joint fit of  $\nu_\mu$  and  $\nu_e$

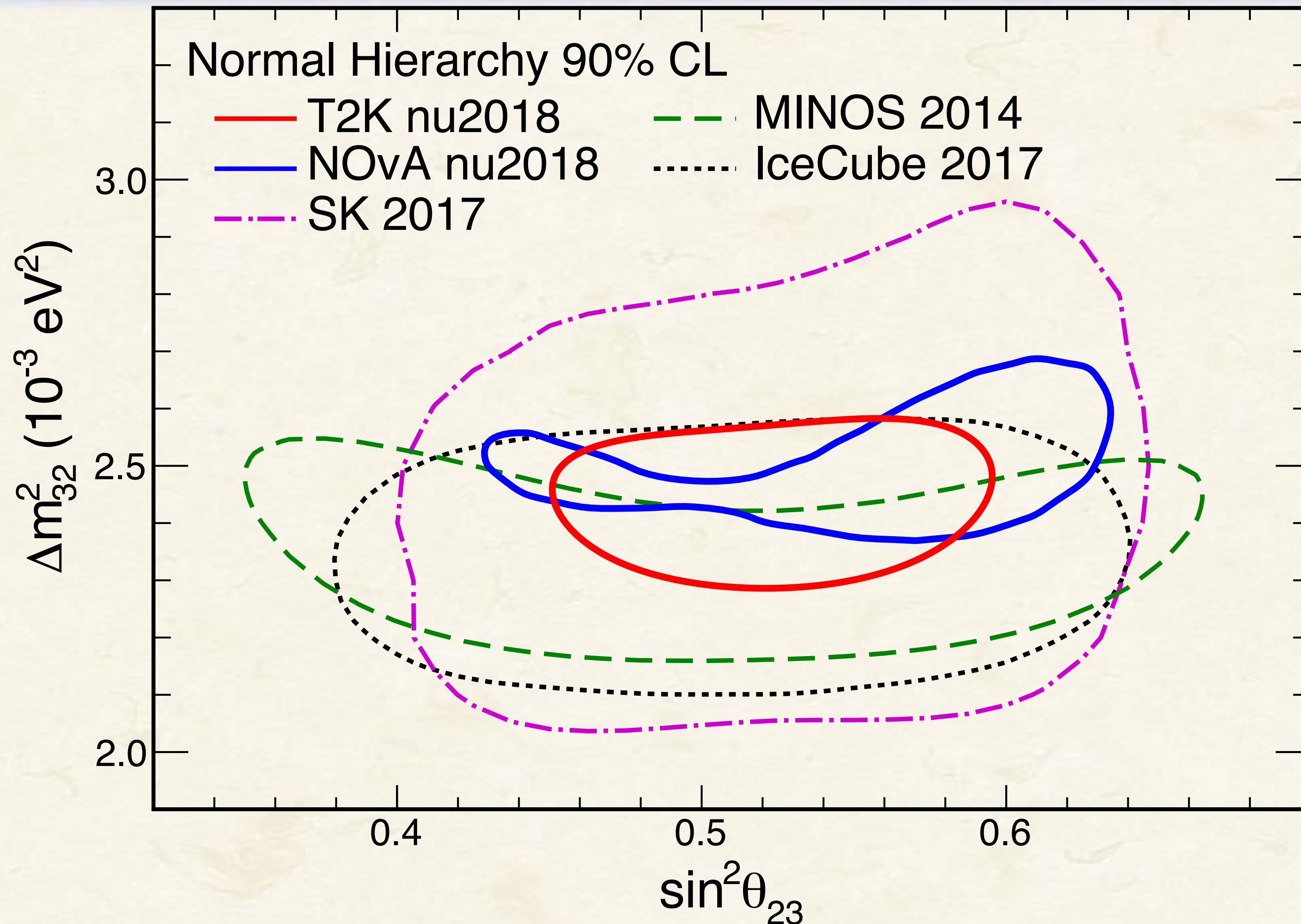
$$\sin^2 \theta_{23} = 0.58 \pm 0.03$$

$$\Delta m_{32}^2 = (2.51^{+0.12}_{-0.08}) \times 10^{-3} \text{eV}^2$$

**Prefer non-maximal at  $1.8\sigma$**   
**Exclude lower octant at similar level**



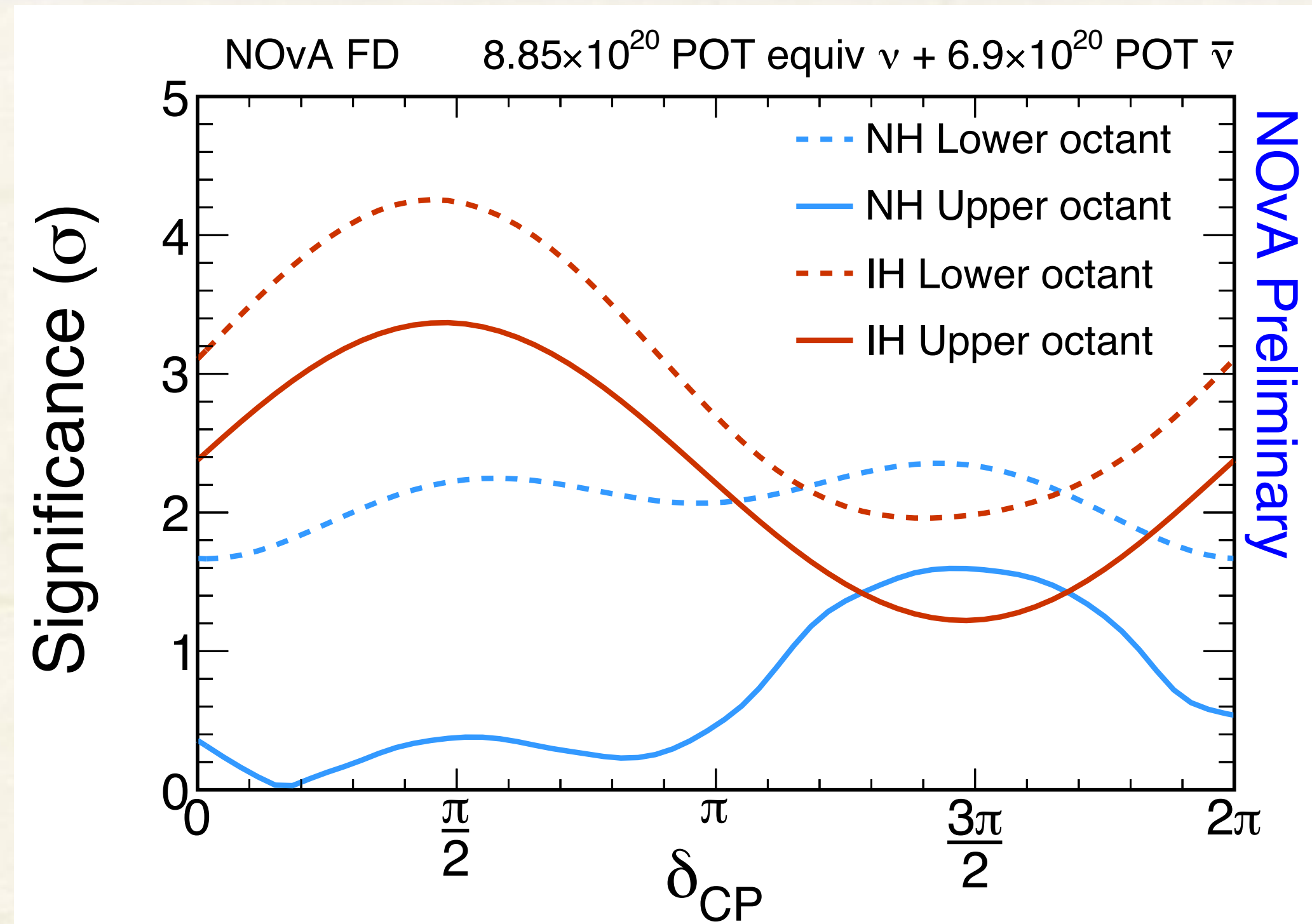
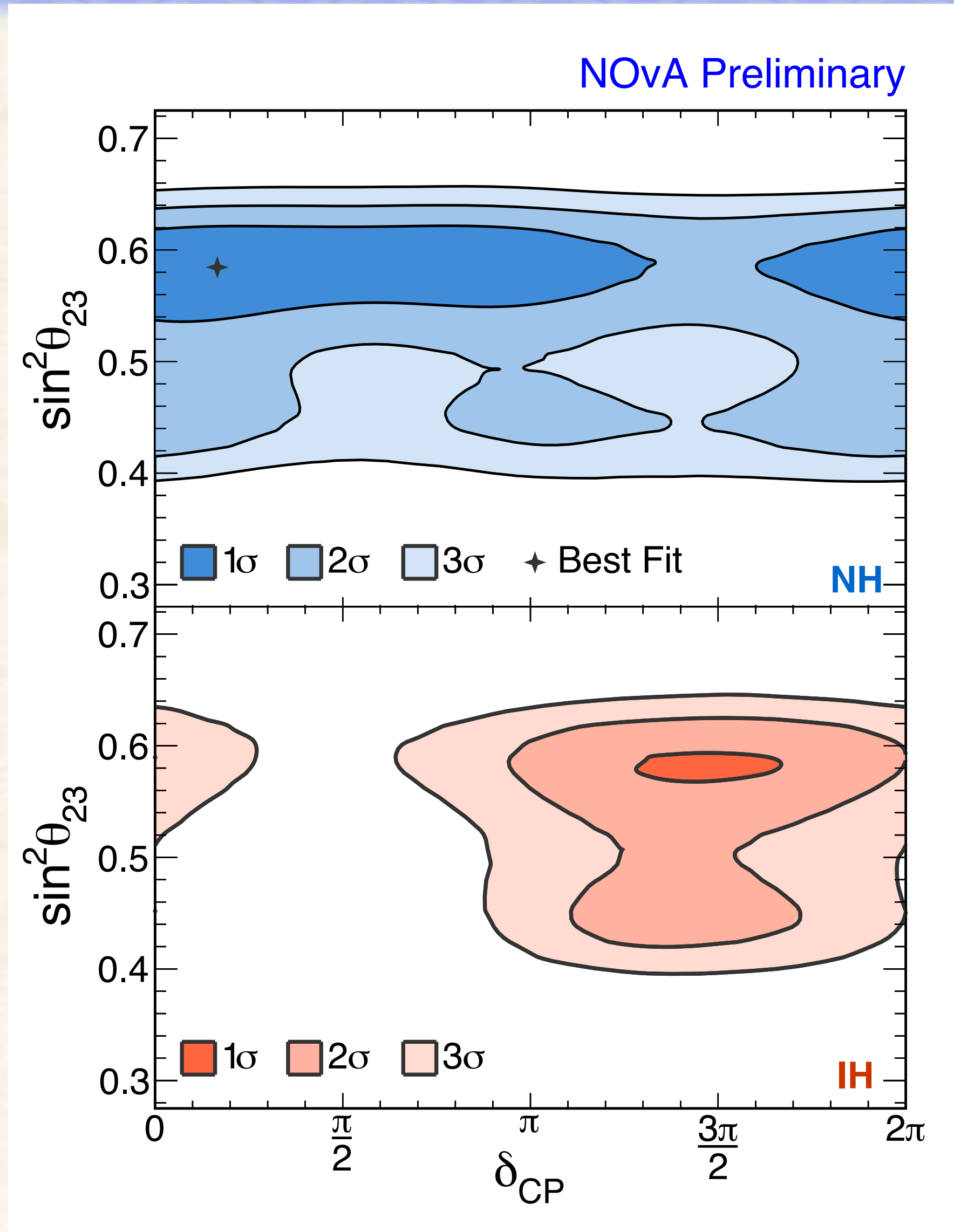
# Comparison: $\Delta m^2_{32}$ — $\theta_{23}$



$\sin^2\theta_{23}$ : <10%  
 $|\Delta m^2|$ : 3-4%

NH preferred by  
T2K, NOvA, SK

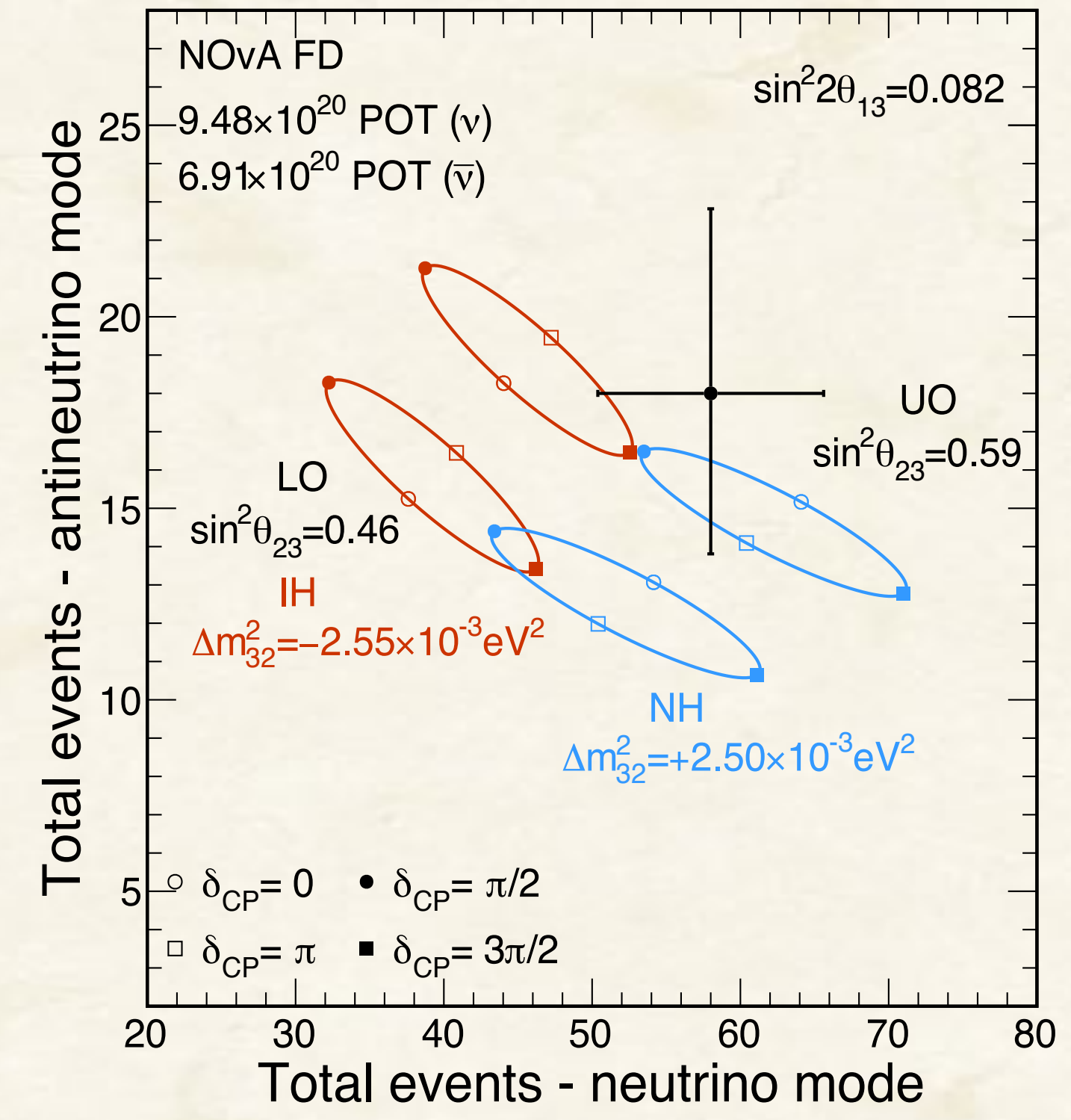
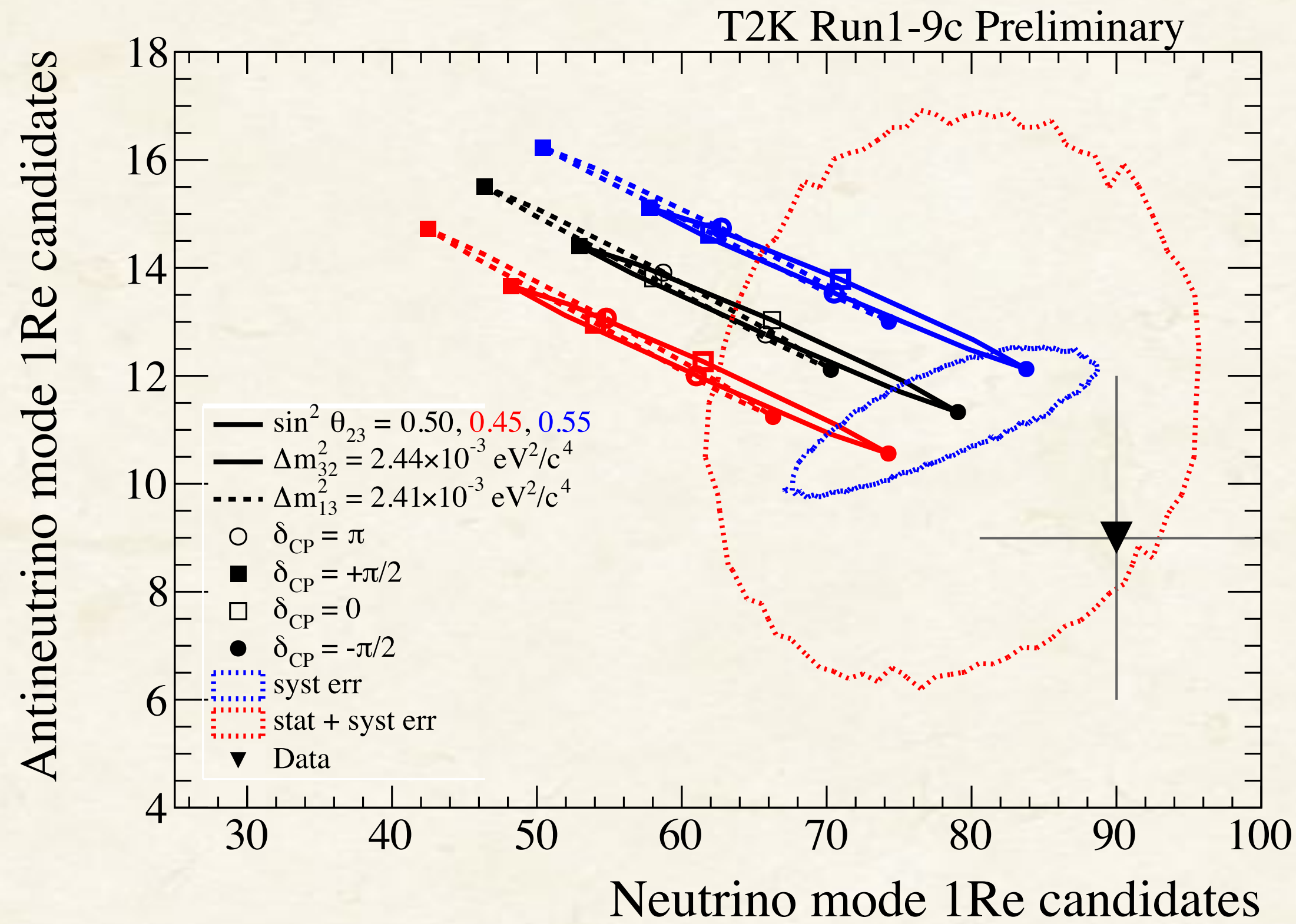
# NOvA: $\delta_{CP}$ and mass hierarchy



◆ Best fit: Normal Hierarchy,  $\delta_{CP} = 0.17\pi$

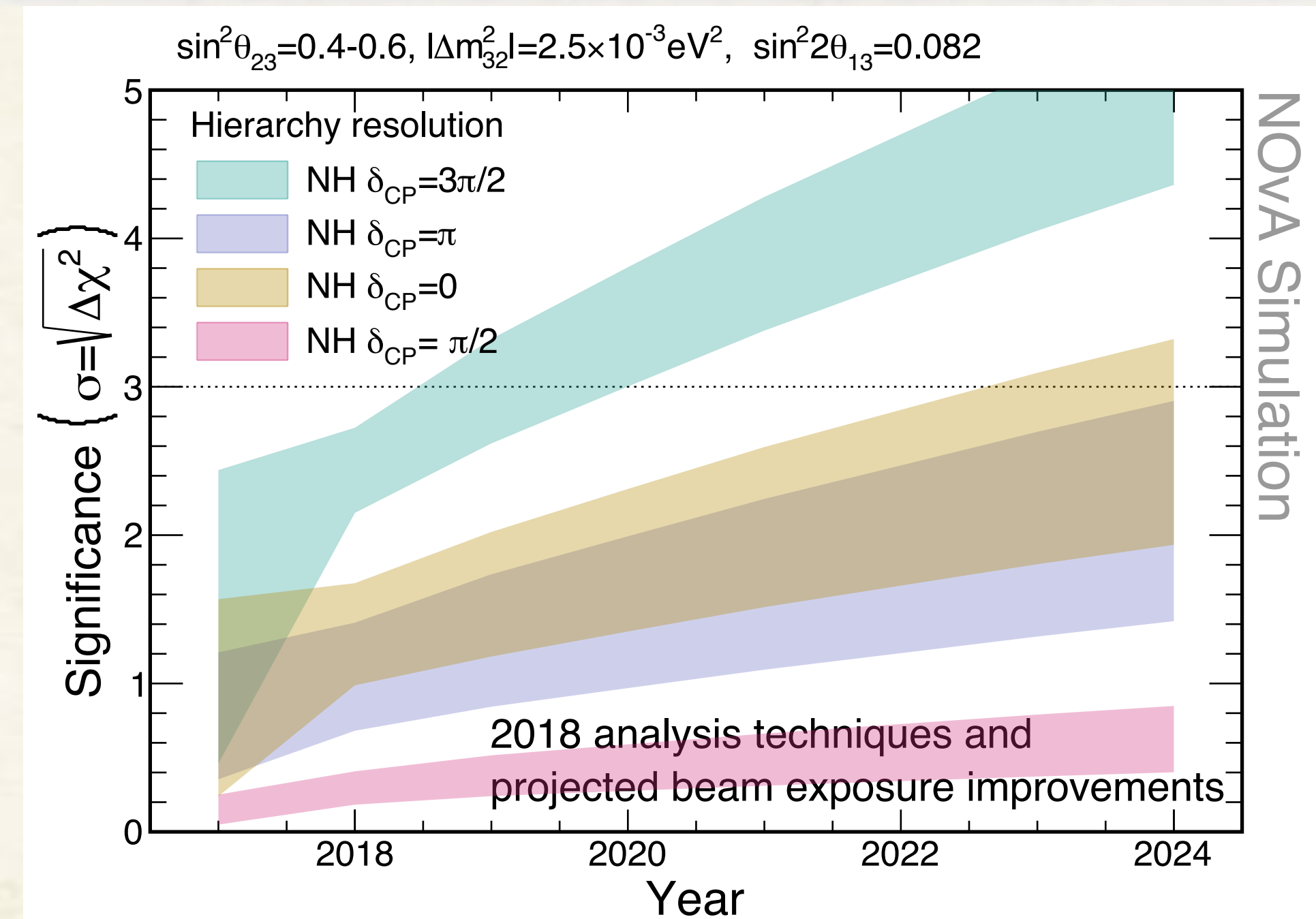
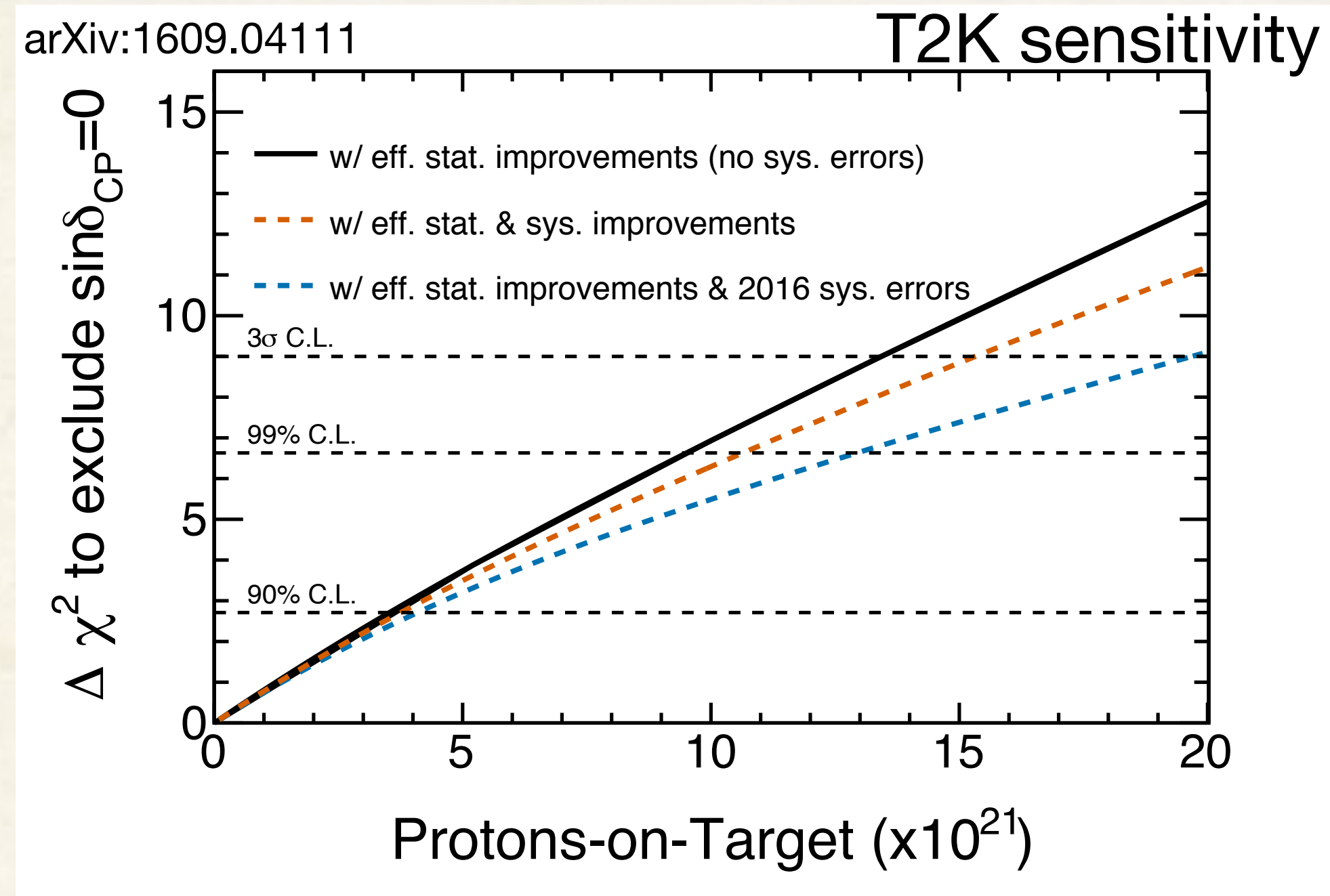
**Prefer NH by 1.8 $\sigma$**

**Exclude  $\delta_{CP} = \pi/2$  in the IH at  $>3\sigma$**

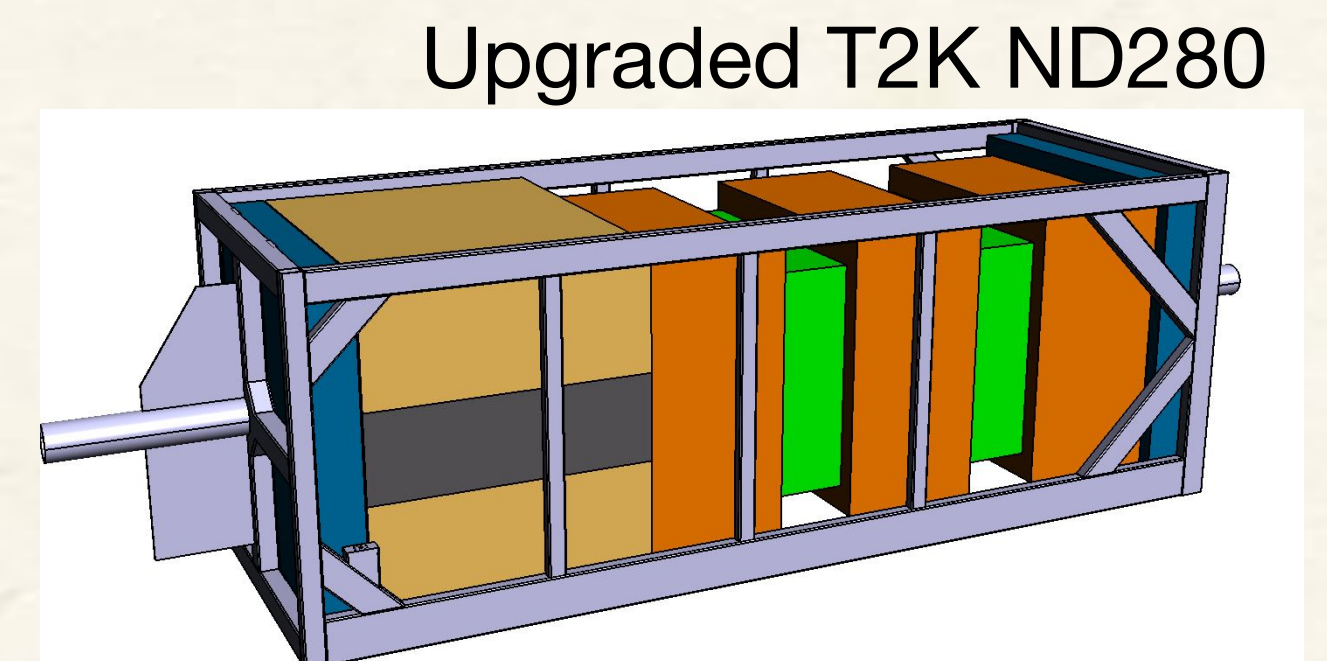


**Beginning of lepton CPV era!**

# The excitement continues...



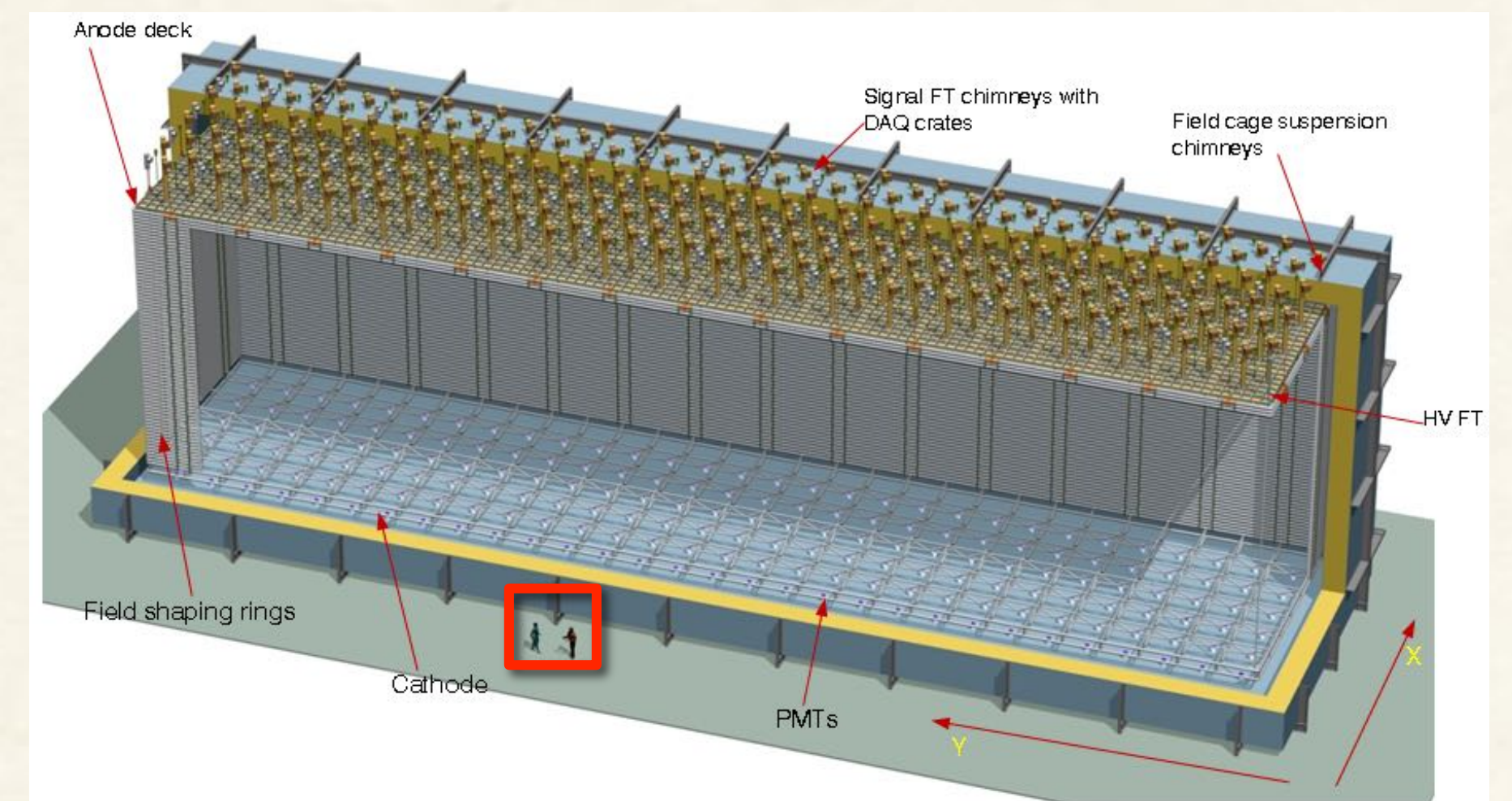
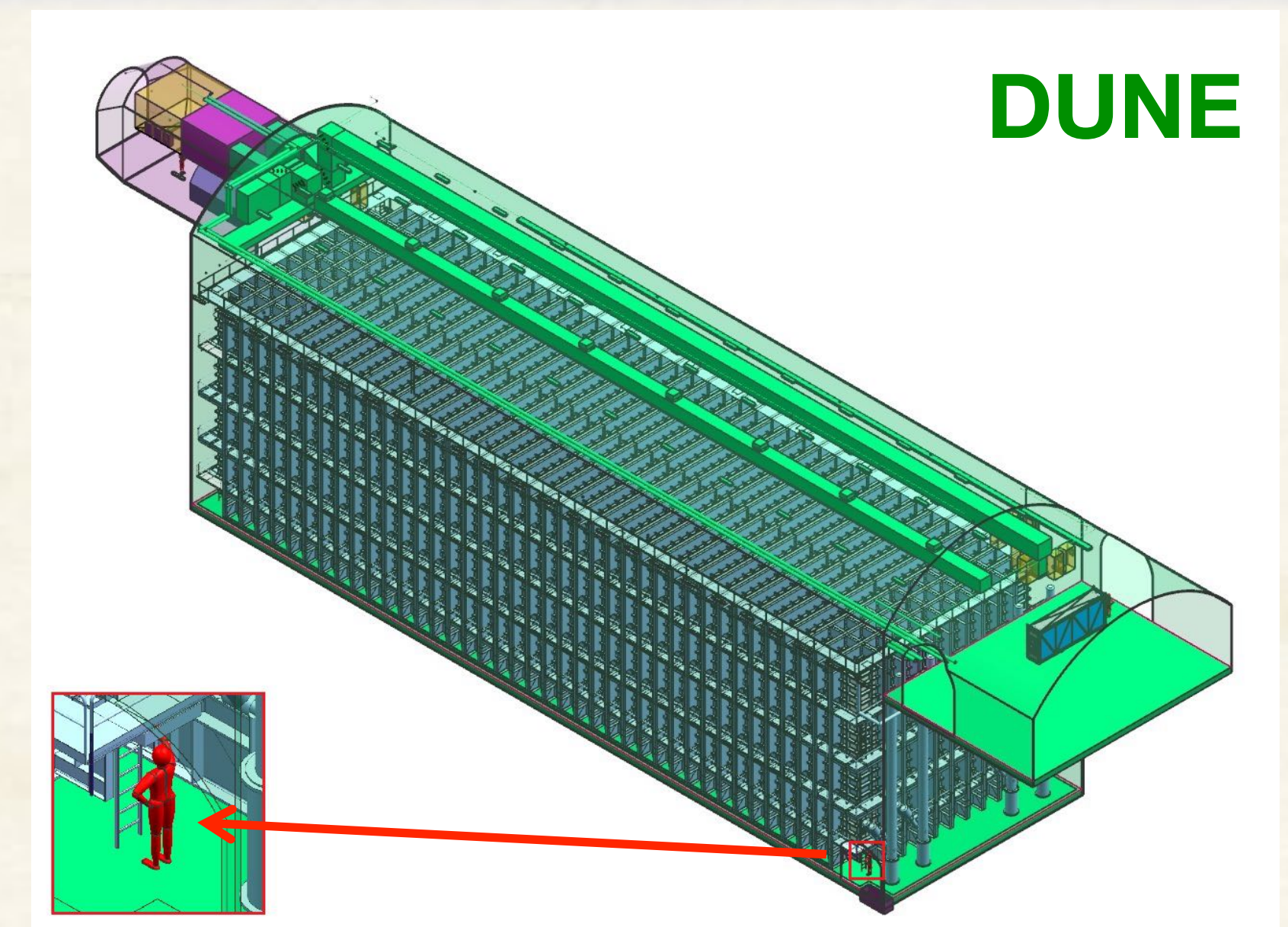
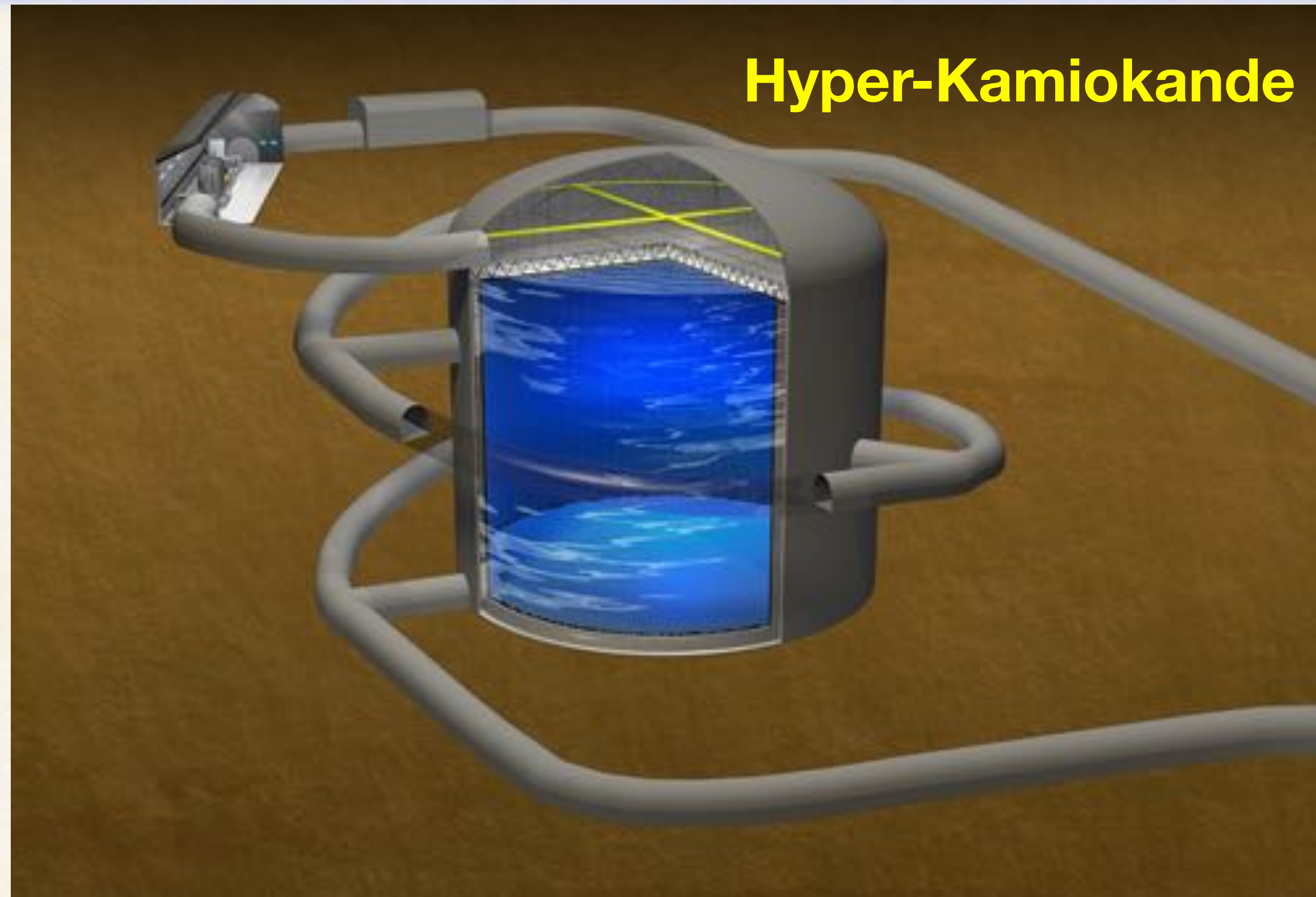
- Both experiments envision to further enhance their capabilities
- T2K: beam power increase (1.3MW) and ND280 upgrade
- NOvA: accelerator improvement (0.9MW) and test beam
- Analysis improvements
- Good prospects for mass hierarchy, CP violation, ...



Parallel talk by K.Iwamoto



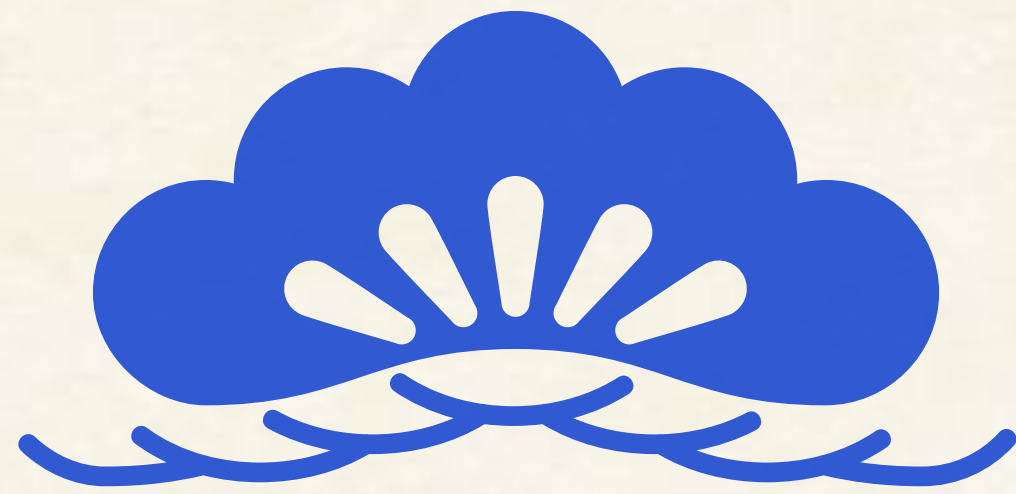
# ...to the next generation!



See later talk by Prof. Jae Yu

# Conclusions

- ◆ Rapid and steady progress in neutrino oscillation physics
- ◆ **MINOS/MINOS+** and **OPERA** final results strongly support the three flavor scenario
- ◆ **Daya Bay/Double Chooz/RENO** continue to improve precision of  $\theta_{13}$ , **JUNO** is coming
- ◆ **T2K** and **NOvA** explore CP and mass hierarchy with neutrino and antineutrino beams
- ◆ Excitement will continue and grow. Stay tuned for more results!



**Additional material**



# JUNO

## 3.1 Mass Hierarchy

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□ Oscillation probability is independent of CP phase and  $\theta_{23}$   
(Reactor neutrinos)

$$P_{ee}(L/E) = 1 - P_{21} - P_{31} - P_{32}$$

$$P_{21} = \cos^4(\theta_{13}) \sin^2(2\theta_{12}) \sin^2(\Delta_{21})$$

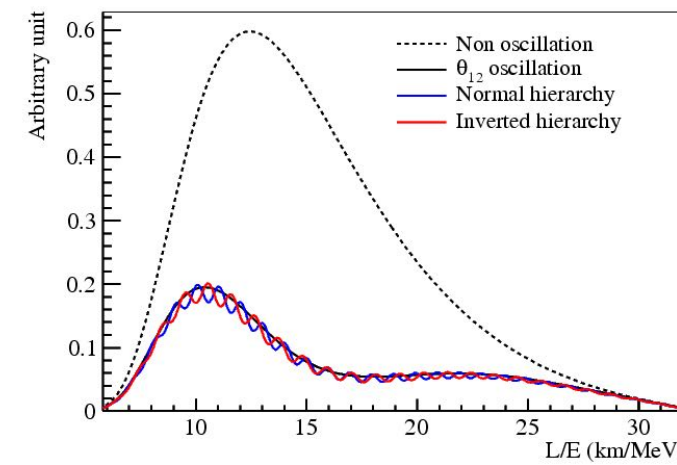
$$P_{31} = \cos^2(\theta_{12}) \sin^2(2\theta_{13}) \sin^2(\Delta_{31})$$

$$P_{32} = \sin^2(\theta_{12}) \sin^2(2\theta_{13}) \sin^2(\Delta_{32})$$

$$P_{ee} = 1 - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2(\Delta_{21}) - \sin^2 2\theta_{13} \sin^2(|\Delta_{31}|) - \sin^2 \theta_{12} \sin^2 2\theta_{13} \sin^2(\Delta_{21}) \cos(2|\Delta_{31}|)$$

+ NH  
- IH

$$\pm \frac{\sin^2 \theta_{12}}{2} \sin^2 2\theta_{13} \sin(2\Delta_{21}) \sin(2|\Delta_{31}|)$$



- The big suppression is the “solar” oscillation ( $\Delta m^2_{12}, \sin^2 \theta_{12}$ )
- “Large” value of  $\theta_{13}$  is crucial



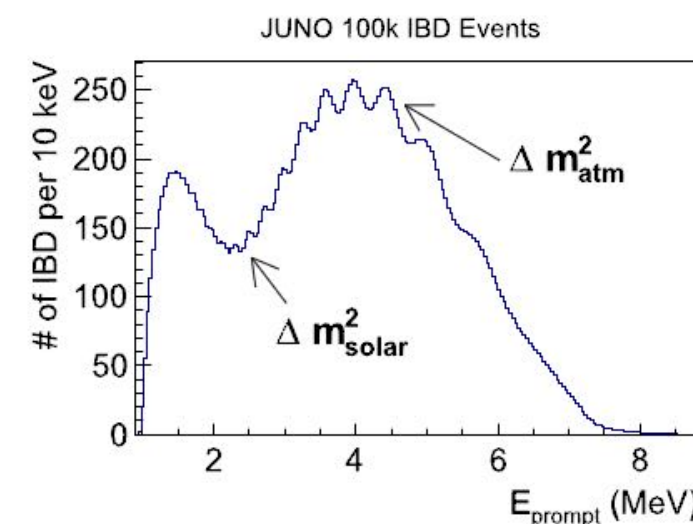
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## 3.2 Measurement of Oscillation Parameters

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Due to good energy resolution and proper baseline, JUNO can help to:

- Improve precisions of three parameters ( $\Delta m^2_{21}, \Delta m^2_{ee}$  and  $\sin^2 \theta_{12}$ ) to **sub-percent level**, several times improvement compared with current precision.
- Probe the unitarity of  $U_{PMNS}$  to  $\sim 1\%$  level

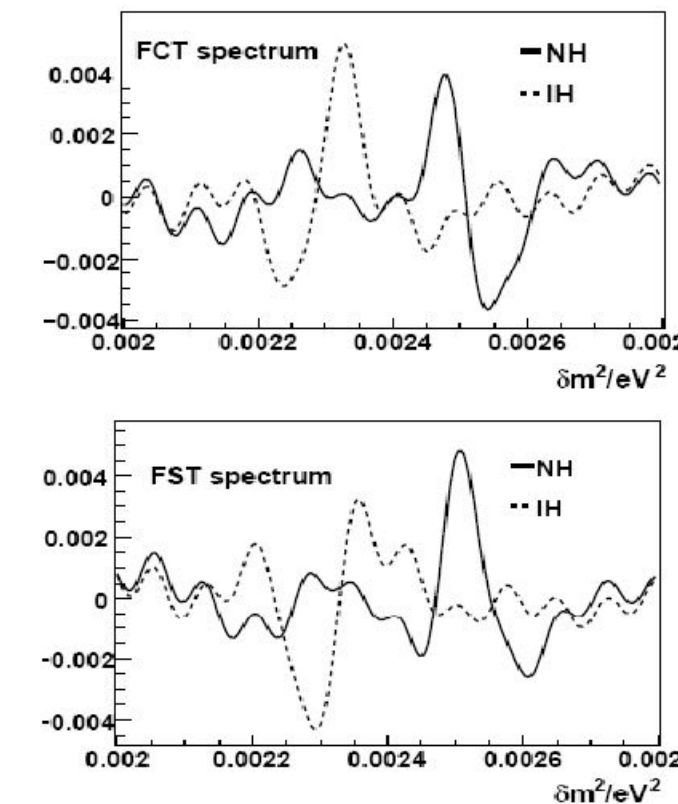


	Nominal	+B2B (1%)	+BG	+EL (1%)	+NL (1%)
$\sin^2 \theta_{12}$	0.54%	0.60%	0.62%	0.64%	0.67%
$\Delta m^2_{21}$	0.24%	0.27%	0.29%	0.44%	0.59%
$ \Delta m^2_{ee} $	0.27%	0.31%	0.31%	0.35%	0.44%

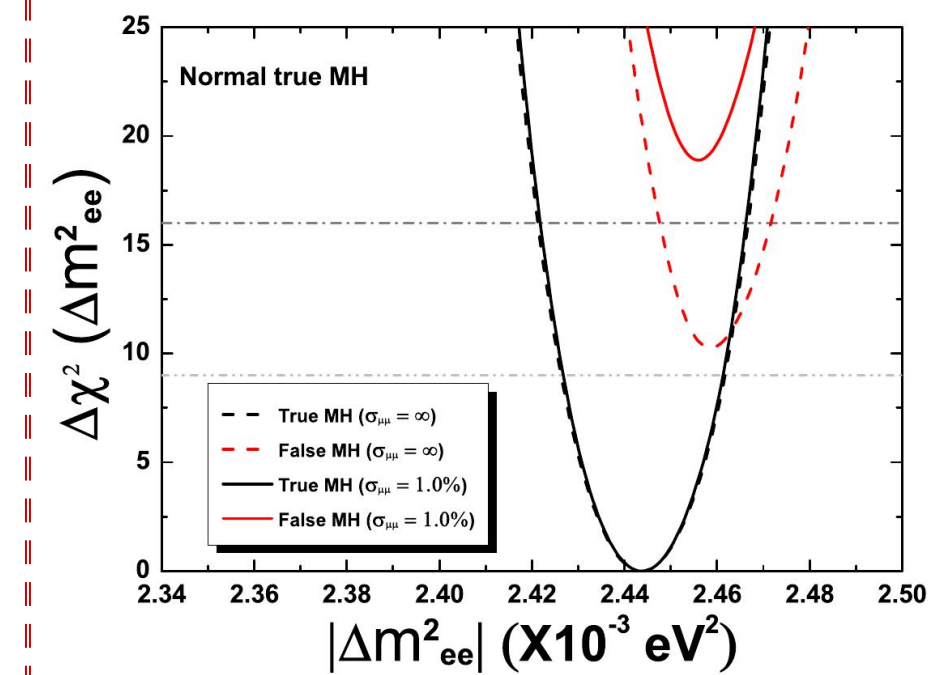


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- No pre-condition of  $\Delta m^2_{32}$
- Only depends on shape but not absolute peak position



□ Sensitivity with 100k events ( $\sim 6$  yrs)

- ✓ No constraint:  $\overline{\Delta \chi^2} > 9$
- ✓ With 1% constraint:  $\overline{\Delta \chi^2} > 16$

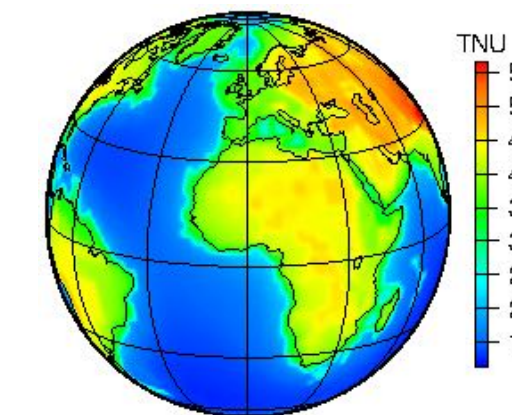


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## 3.3 Neutrino Astrophysics and Others

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- Neutrinos from the Earth escape freely and bring the information about U, Th and K abundances and their distributions
- Due to its largest LS size, the expected geo-neutrino rate in JUNO is  $\sim 1.1/\text{day}$ .
- Within the 1<sup>st</sup> year, JUNO will record more geo-neutrino events than all other detectors



➢ JUNO will be the most precise experiment for geo-neutrino study. In the meanwhile, JUNO is also attractive for other neutrino astrophysics, such as supernova neutrinos, diffuse supernova neutrinos, solar neutrinos and atmospheric neutrinos.

Beside these, additional physics is also rich in JUNO

- Sterile neutrinos
- Dark matter searches
- Proton decay
- Other exotic searches



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# Neutrino beams and long baseline experiments

Accelerator	Experiment	Baseline	Beam power	Years
KEK-PS (KEK)	K2K	250km	5kW	1999-2004
Main Injector (Fermilab)	MINOS(+)	730km	400kW+	2005-2016
SPS (CERN)	OPERA / ICARUS	730km	510kW	2008-2012
J-PARC MR (J-PARC/KEK)	T2K	295km	500kW (design:750kW)	2009-
Main Injector (Fermilab)	NOvA	810km	700kW	2014-

# T2K: data and predictions

Sample	Predicted ( $\sin^2\theta_{23}=0.528$ )				Observed
	$\delta=0$	$\delta=+\pi/2$	$\delta=\pi$	$\delta=-\pi/2$	
$\nu$ beam, 1R $\mu$	268.2	268.5	268.9	268.5	<b>243</b>
$\bar{\nu}$ beam, 1R $\mu$	95.3	95.5	95.8	95.5	<b>102</b>
$\nu$ beam, 1Re 0 decay-e	61.6	50.1	62.2	73.8	<b>75</b>
$\nu$ beam, 1Re 1 decay-e	6.0	4.9	5.8	6.9	<b>15</b>
$\bar{\nu}$ beam, 1Re 0 decay-e	13.4	14.9	13.3	11.8	<b>9</b>

$$\sin^2\theta_{23} = 0.528$$

$$\sin^2\theta_{13} = 0.0219$$

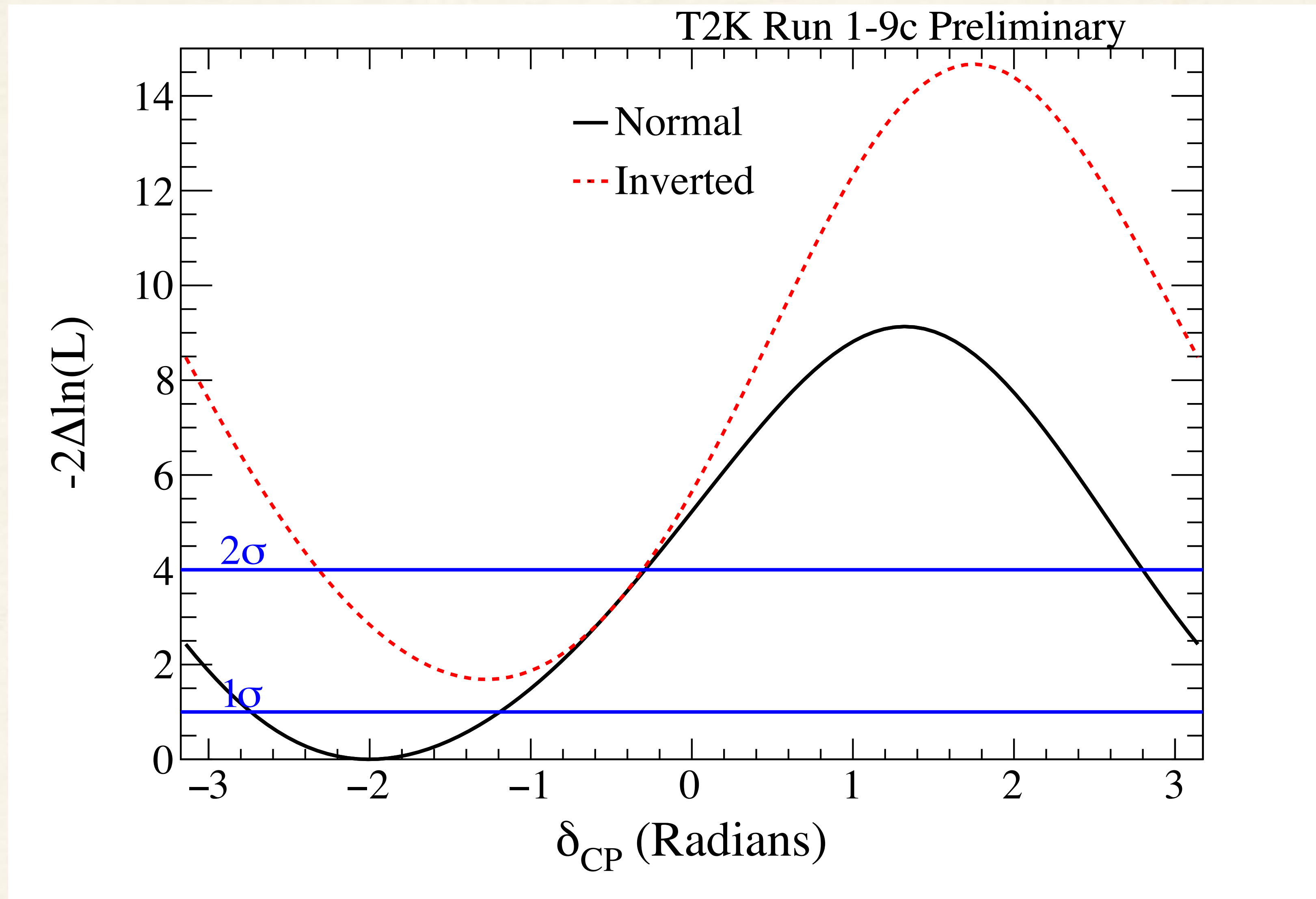
$$\Delta m^2_{32} = 2.5 \times 10^{-3} \text{ eV}^2$$

**Consistent with maximal  $\nu_\mu$  disappearance**  
**Prefer large CP violation ( $\delta_{CP} \sim \pi/2$ )**

# T2K systematic error

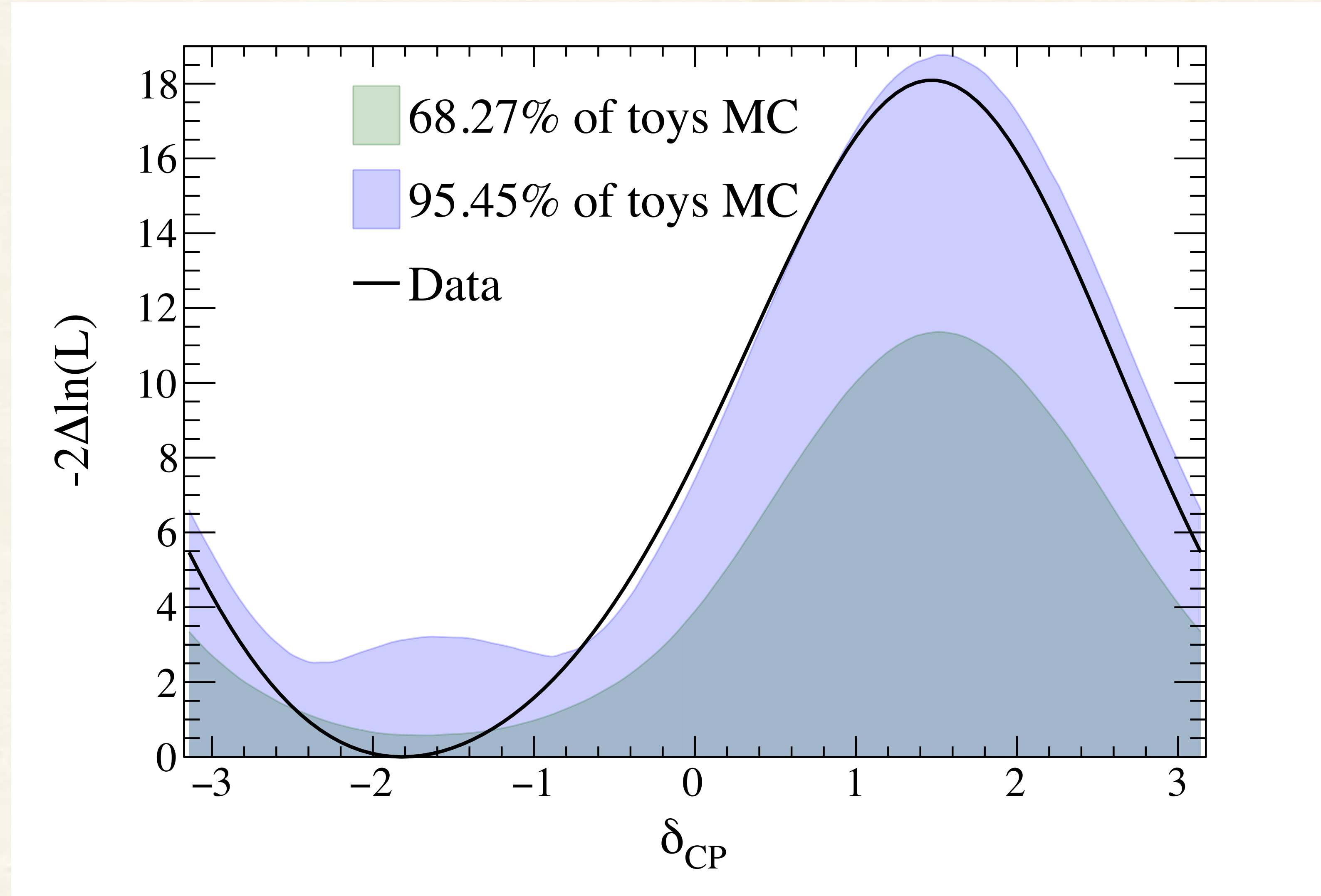
Error source	1-Ring $\mu$		1-Ring $e$			
	FHC	RHC	FHC	RHC	FHC 1 d.e.	FHC/RHC
SK Detector	2.40	2.01	2.83	3.79	13.16	1.47
SK FSI+SI+PN	2.20	1.98	3.02	2.31	11.44	1.58
Flux + Xsec constrained	2.88	2.68	3.02	2.86	3.82	2.31
$E_b$	2.43	1.73	7.26	3.66	3.01	3.74
$\sigma(\nu_e)/\sigma(\bar{\nu}_e)$	0.00	0.00	2.63	1.46	2.62	3.03
NC1 $\gamma$	0.00	0.00	1.07	2.58	0.33	1.49
NC Other	0.25	0.25	0.14	0.33	0.99	0.18
Osc	0.03	0.03	3.86	3.60	3.77	0.79
All Systematics	4.91	4.28	8.81	7.03	18.32	5.87
All with osc	4.91	4.28	9.60	7.87	18.65	5.93

# $\delta_{CP}$ with T2K alone



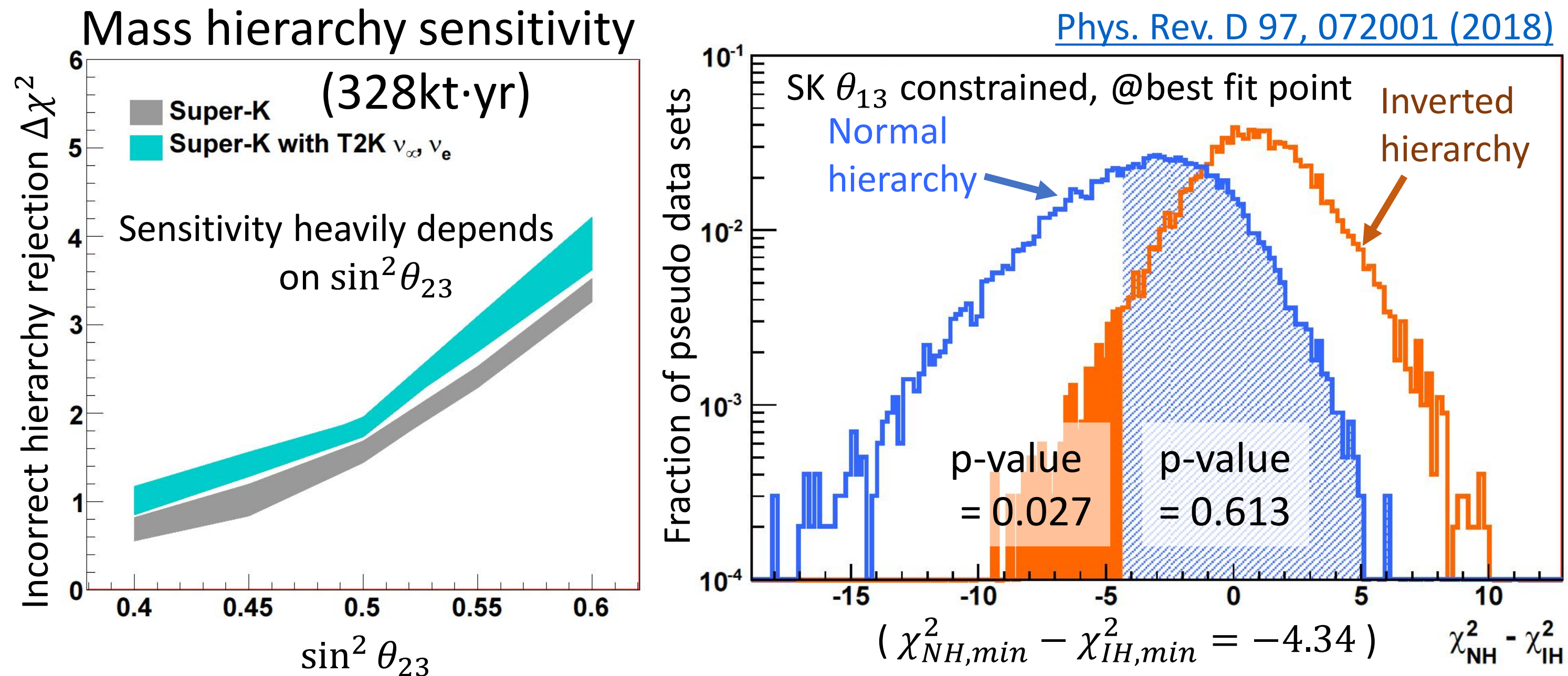


# T2K sensitivity and data result



# Mass hierarchy from Super-K

## Determination of hierarchy determination



Estimate p-values using pseudo-data

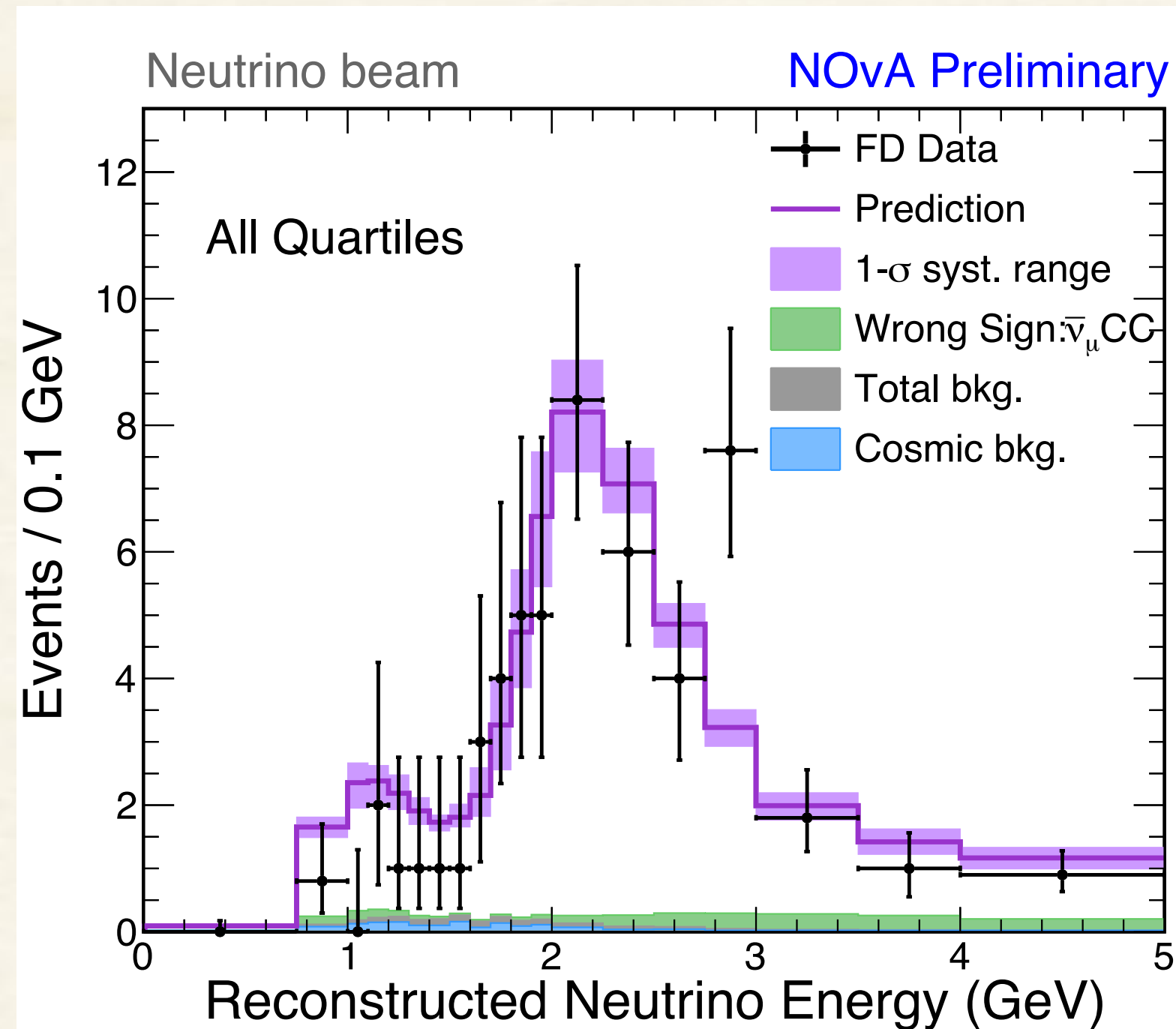
for the smallest and largest  $\sin^2 \theta_{23}$ .

Hypothesis test  $\sim$  CL<sub>s</sub> method :  $CL_s(\text{IH rejection}) \equiv \frac{p_0(\text{IH})}{1-p_0(\text{NH})}$

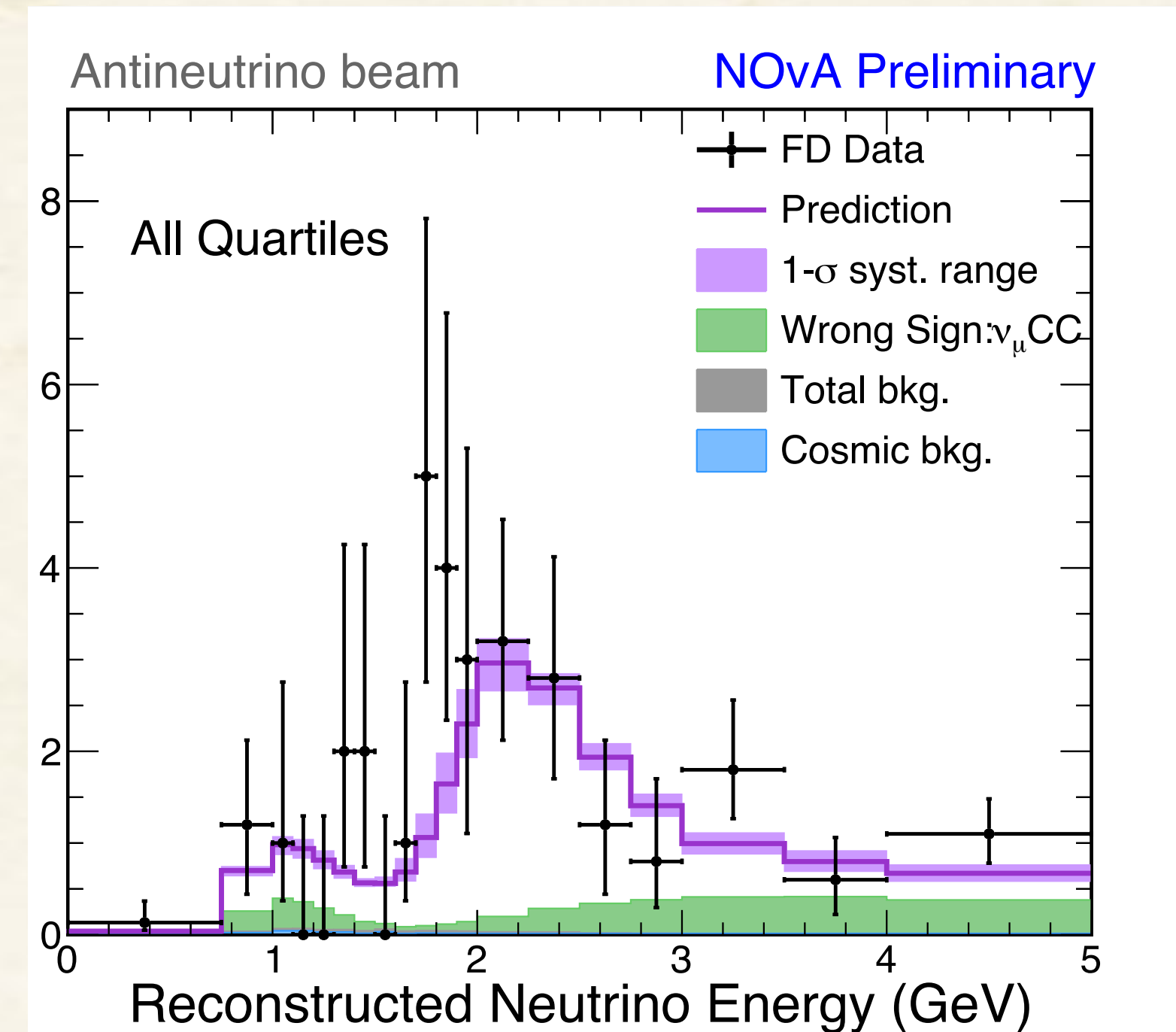
Normal hierarchy is favored  $\rightarrow$

SK only	80.6 ~ 96.7%
SK + T2Kmodel	91.5 ~ 94.7%

# NOvA: $\nu_\mu$ and $\bar{\nu}_\mu$ data

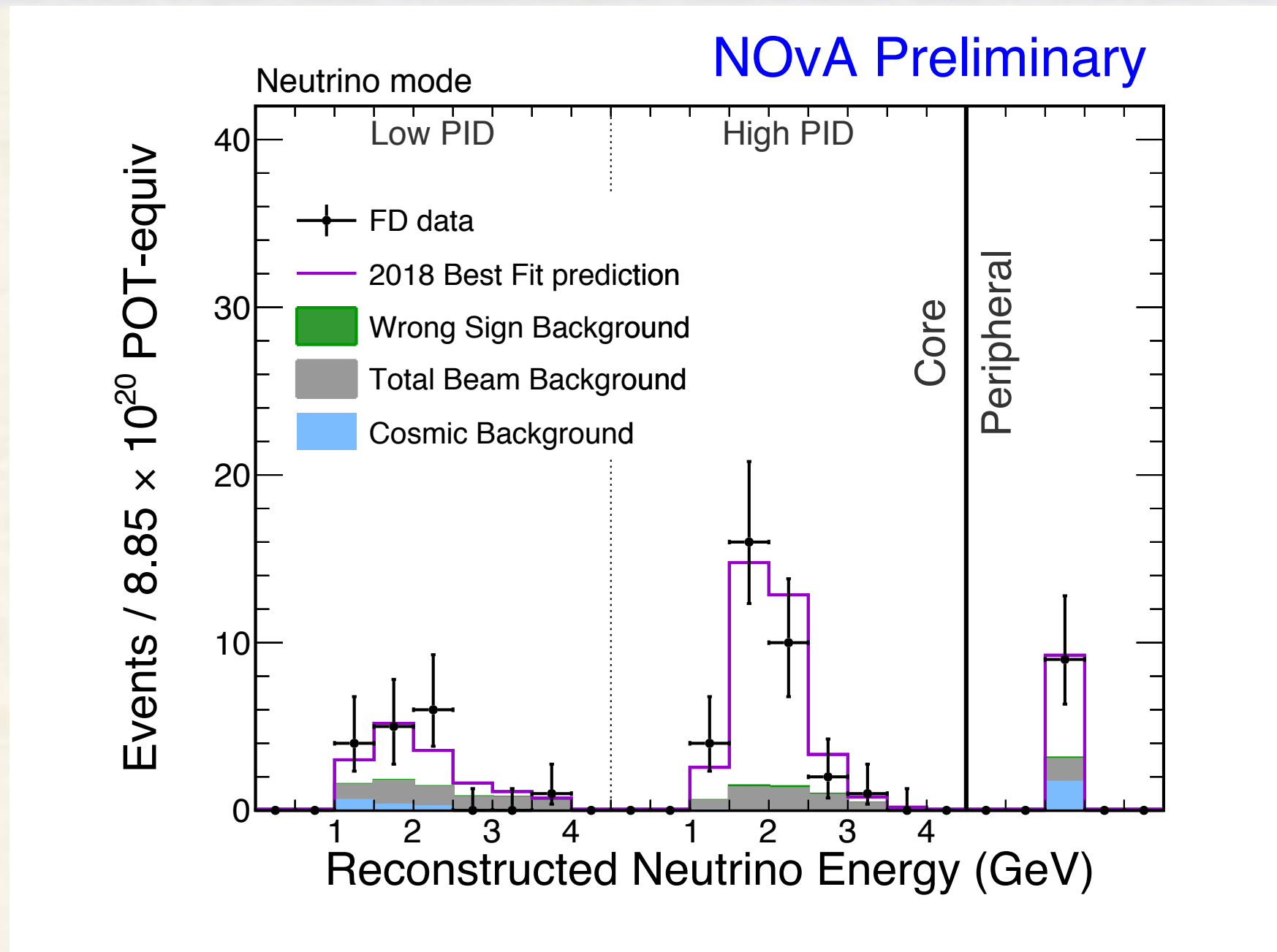


Observed	113
Best fit prediction	121
Beam BG	1.2
Cosmic BG	2.1
Unoscillated	730 $+38/-49(\text{syst})$

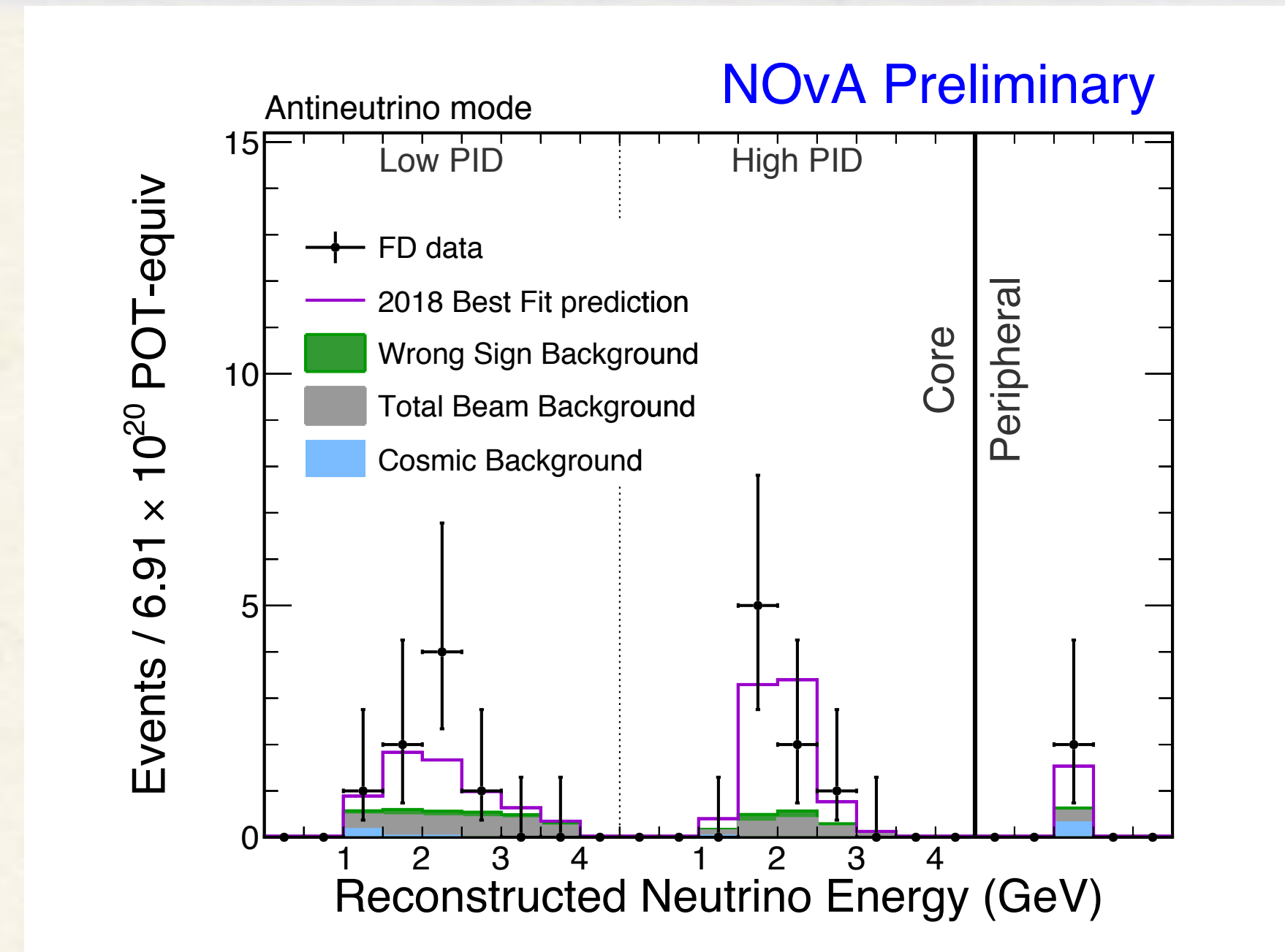


Observed	65
Best fit prediction	50
Beam BG	0.6
Cosmic BG	0.5
Unoscillated	266 $+12/-14(\text{syst})$

# NOvA: $\nu_e$ and $\bar{\nu}_e$ data



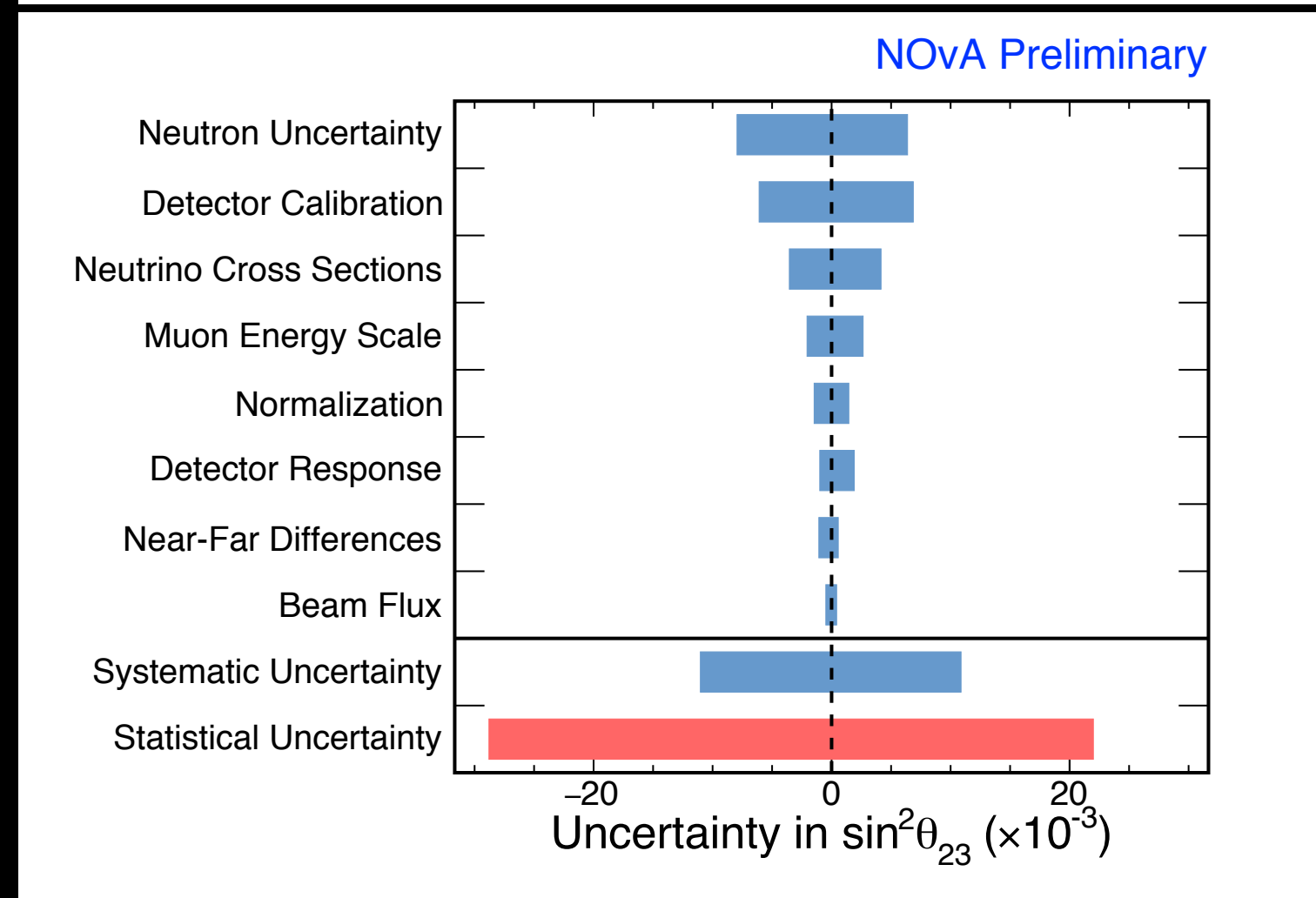
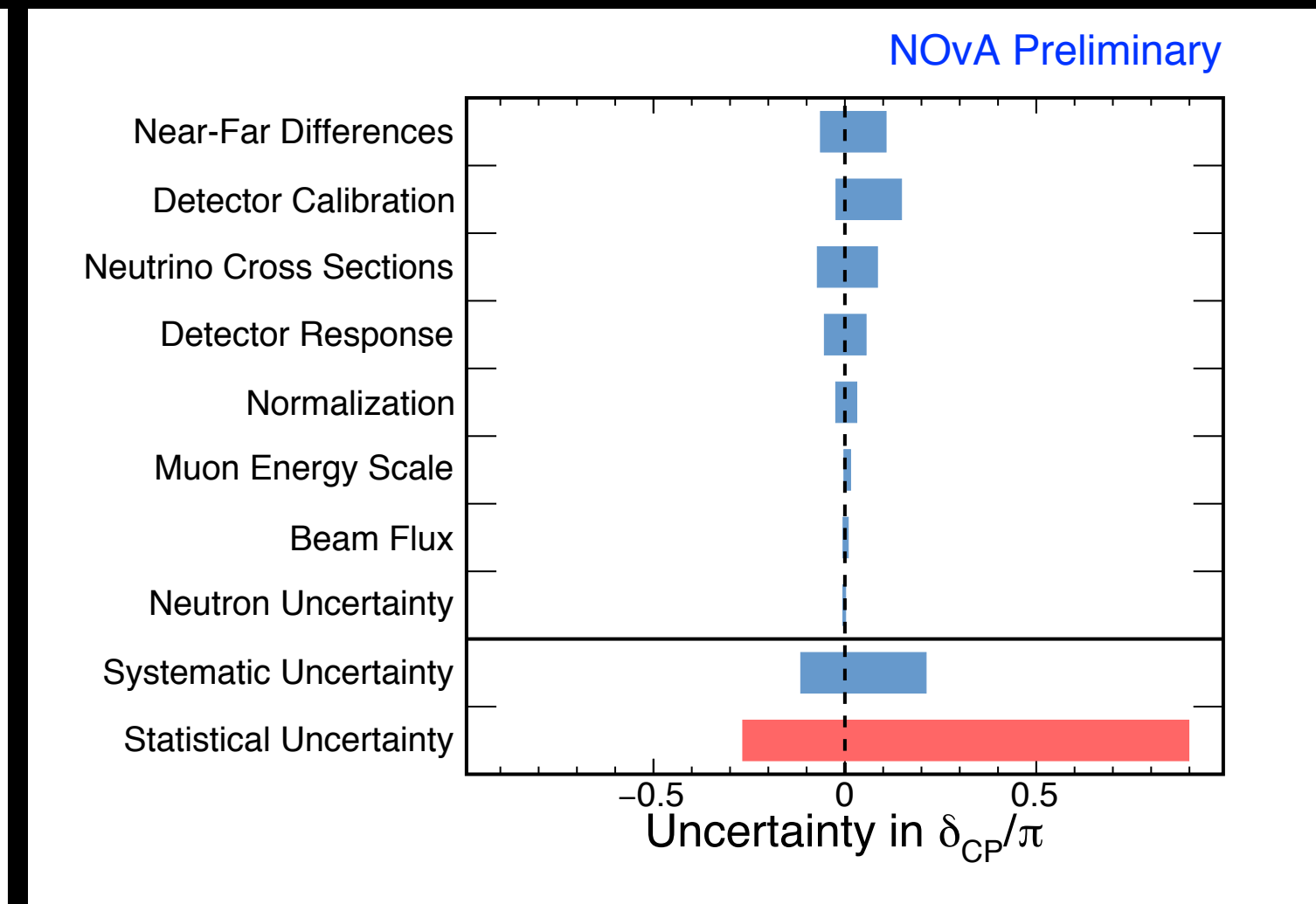
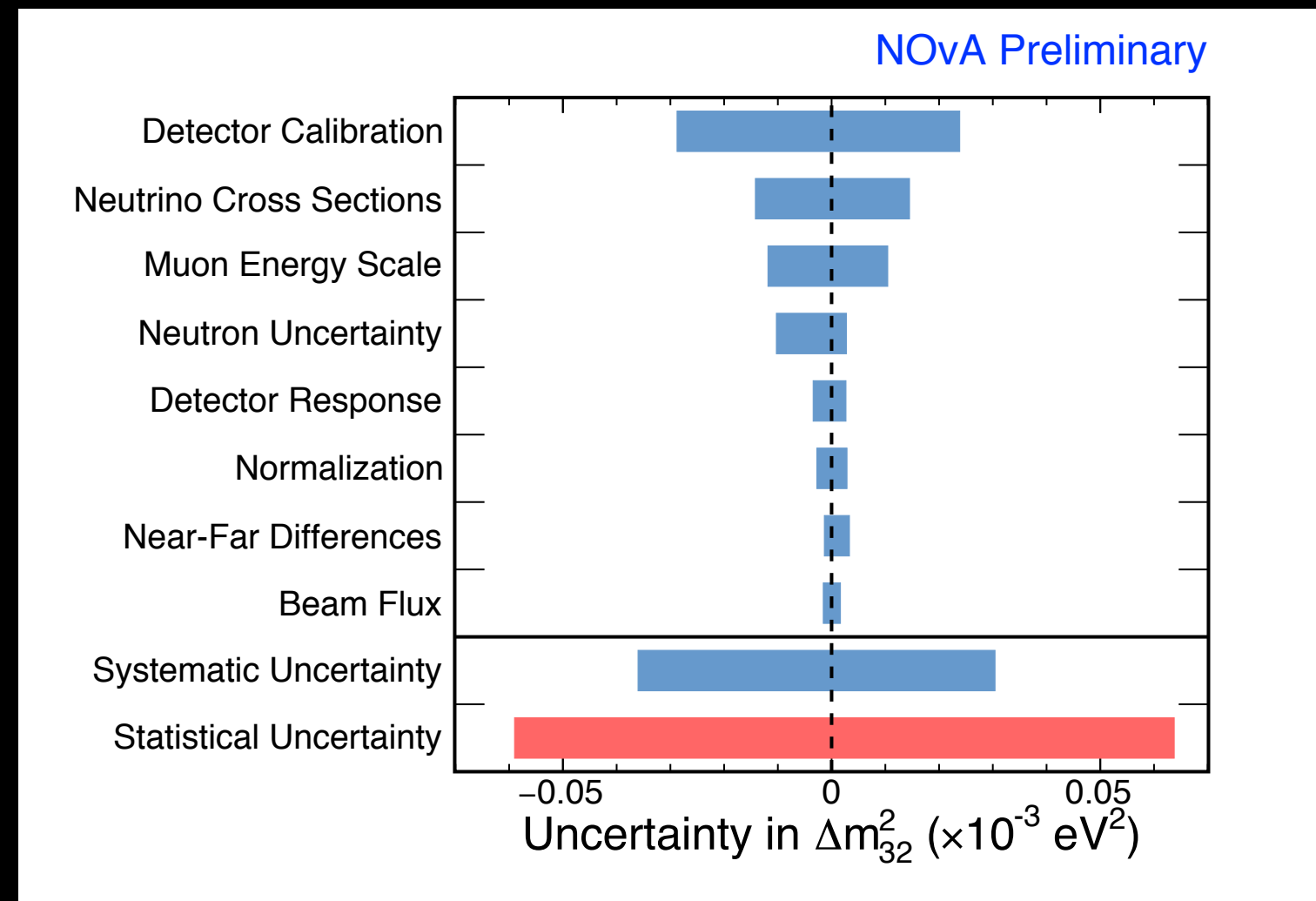
Observed	58
Prediction	30-75
Wrong-sign	0.3-1.0
Beam BG	11.1
Cosmic BG	3.3
Total BG	14.7-15.4



Observed	18
Prediction	10-22
Wrong-sign	0.5-1.5
Beam BG	3.5
Cosmic BG	0.7
Total BG	4.7-5.7

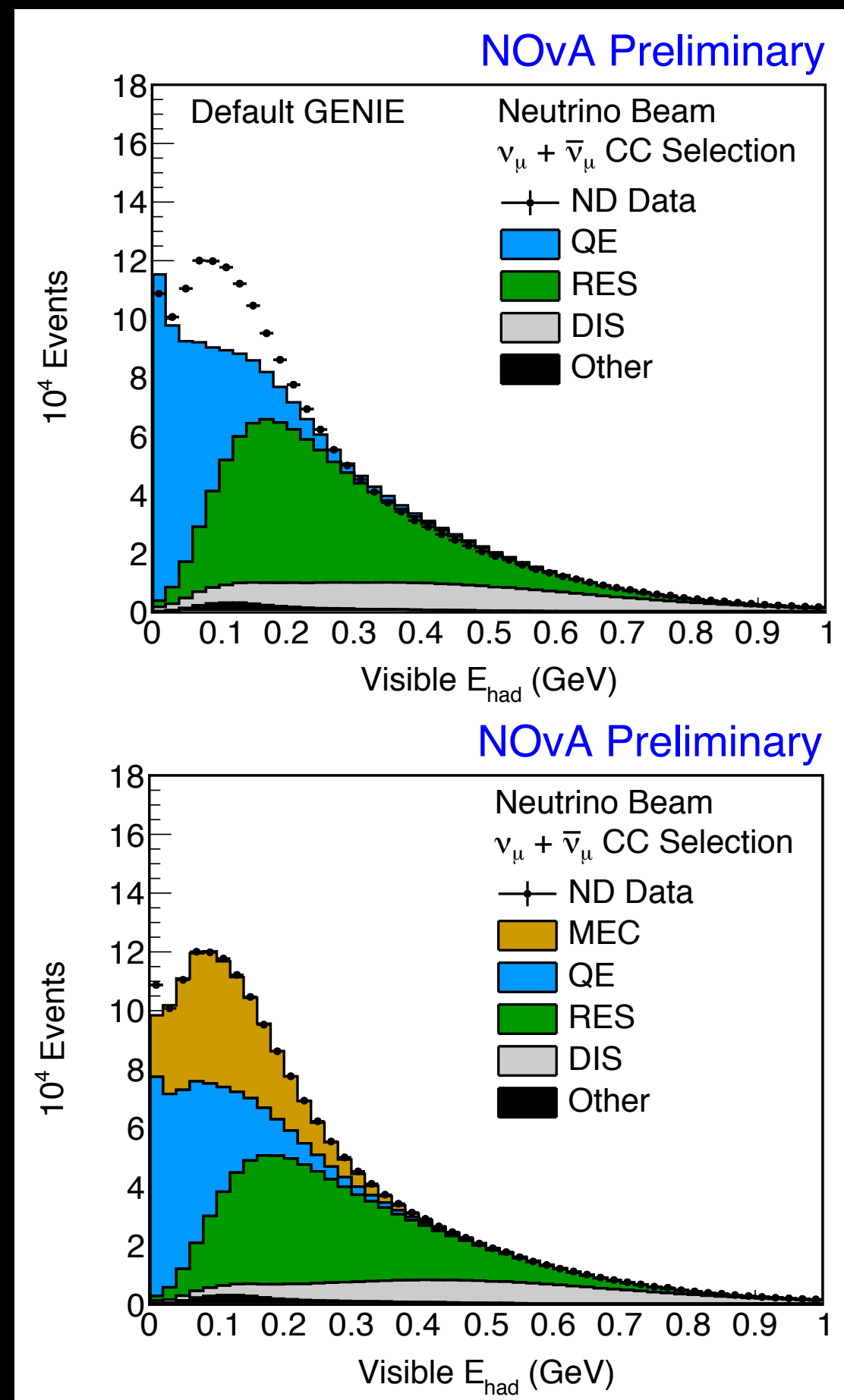
**>4 $\sigma$  evidence of  $\bar{\nu}_e$  appearance!**

# NOvA systematic uncertainties



- Improved systematic uncertainties. We are still statistics limited but calibration and cross sections are the largest uncertainties.
- Our upcoming test beam program will address the calibration and detector response uncertainties.

## NEUTRINO INTERACTION TUNING



Mayly Sanchez - ISU

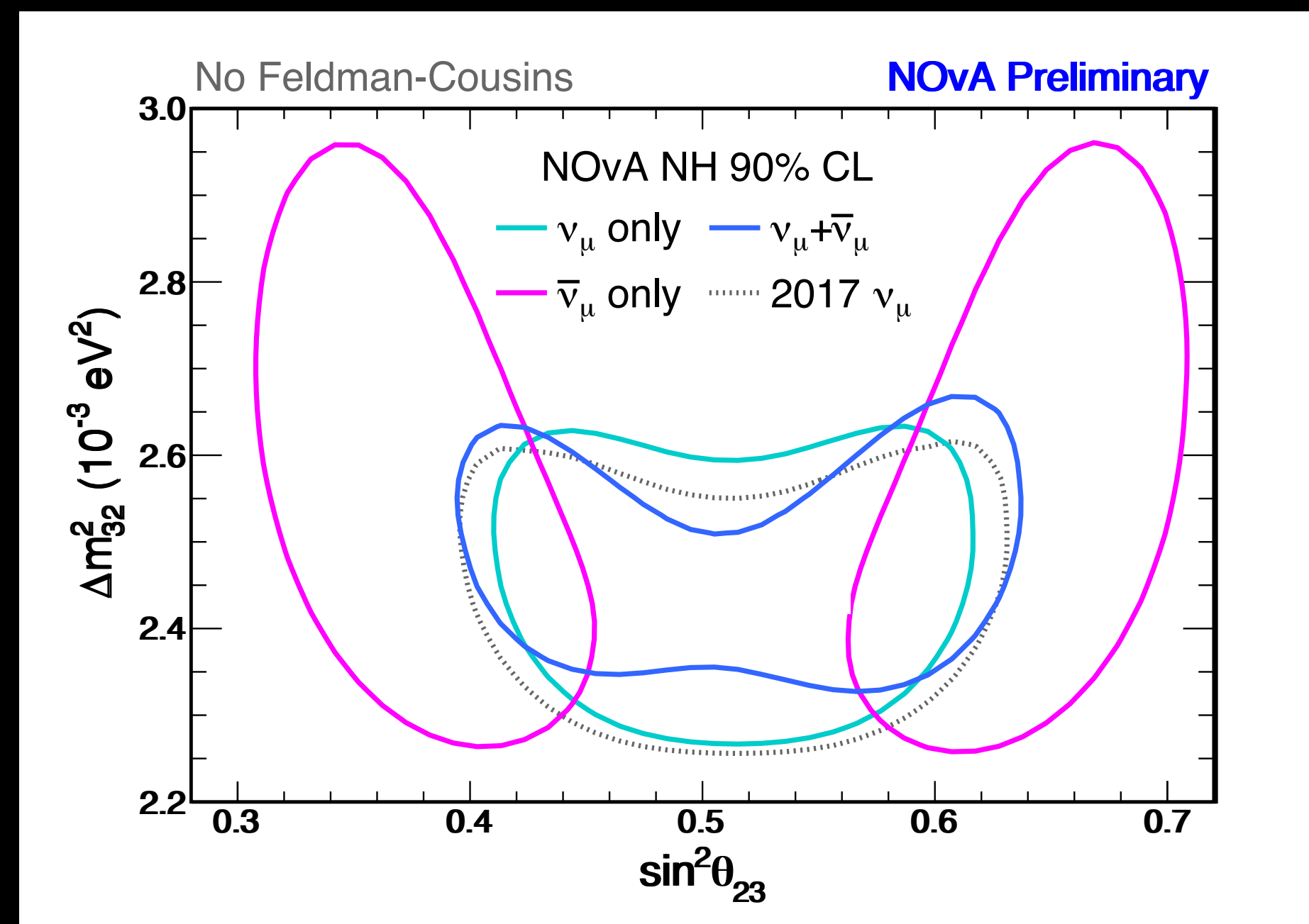
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- The tuning is done independently for the neutrino vs antineutrino beam samples.
- Various corrections and tunings are applied:
  - Correct quasielastic component to account for effect of long-range nuclear correlations using model of València group via work of R. Gran (MINERvA) [<https://arxiv.org/abs/1705.02932>]
  - Apply same long-range effect as for QE to resonant baryon production as well. Nonresonant inelastic scattering (DIS) at high invariant mass ( $W > 1.7 \text{ GeV}/c^2$ ) weighted up 10% **based on NOvA data**.
  - Introduce custom tuning of GENIE "Empirical MEC" [T. Katori, AIP Conf. Proc. 1663, 030001 (2015)] **based on NOvA ND data** to account for multinucleon knockout (2p2h).

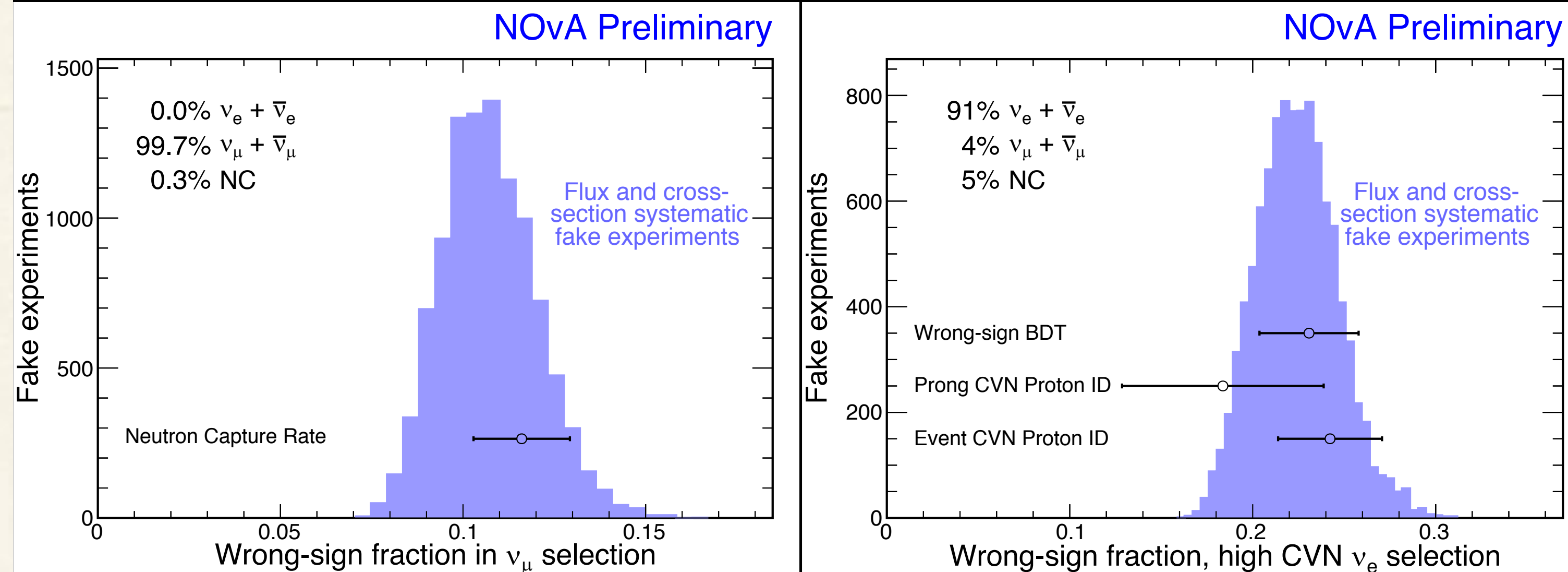
# NOvA

## MUON NEUTRINO DISAPPEARANCE

- The combined data of neutrino and antineutrino beams are fitted assuming CPT invariance.
- We observe 113 events and expect 126 at this combined best fit for the neutrino beam mode and observe 65 events and expect 52 at the best fit in antineutrino beam mode.
- If fit separately, the antineutrino beam mode prefers a more non-maximal solution than the neutrino beam mode. However the  $\chi^2$ s are consistent with the combined fit oscillation parameters with  $p > 4\%$ .



# WHAT IS DIFFERENT IN ANTINEUTRINOS? WRONG-SIGN CONTAMINATION



- 11% wrong-sign in the  $\nu_\mu$  ND sample background
  - Consistent with data-based cross-check using neutron captures.
- 22% (32%) in the  $\nu_e$  ND background in the high (low) PID bin
  - Consistent with data-based cross-checks using identified protons and event kinematics.
- •~10% systematic uncertainty from flux and cross section
- Does not include uncertainties from detector effects.



# PREDICTING THE FAR DETECTOR OBSERVATIONS

- The neutrino spectrum is measured at the ND (before oscillations), this is a combination of neutrino flux, cross section and efficiency.
  - Estimate the underlying true energy distribution of selected ND events.
- The measured spectrum is used to make a prediction of the expectation at the FD
  - Multiplying the true energy distribution by the Far/Near Ratio, applying oscillation probabilities and then converting to a predicted reconstructed energy distribution
- Since NOvA has functionally similar Near and Far Detectors the flux combined with the cross sections uncertainties largely cancel.

