
Second Order Non-Gaussianity - SONG

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Today

A Boltzmann code like CAMB or CLASS, but at second order in cosmological perturbation theory

Advantages of going to second order

- Higher precision
- New physics
 - Non-Gaussianity
 - Mode coupling
 - Scalar-Vector-Tensor coupling
 - New source terms

$$\Delta^{(1)}(\vec{k}) = T^{(1)}(k)\Phi(\vec{k})$$
$$\Delta^{(2)}(\vec{k}) = \int \int d\vec{k}_1 d\vec{k}_2 \delta^3(\vec{k} - \vec{k}_1 - \vec{k}_2) T^{(2)}(k, \vec{k}_1, \vec{k}_2) \Phi(\vec{k}_1) \Phi(\vec{k}_2)$$

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- Source available at <https://github.com/coccoinomane/song>
- Based on CLASS, installation and running the code work the same
- provide .ini and .pre file in CLASS conventions, containing both SONG and CLASS parameters

```
255 include_sachs_wolfe_in_los_1st_order = no
256 include_integrated_sachs_wolfe_in_los_1st_order = no
257
258
259 ***** Second-order line of sight (LOS) sources *****
260
261 use_test_source = no
262
263 include_pure_scattering_in_los_2nd_order = yes
264 include_quad_scattering_in_los_2nd_order = yes
265 include_metric_in_los_2nd_order = yes
266
267 include_quad_metric_in_los_2nd_order = no
268 include_time_delay_in_los_2nd_order = no
269 include_redshift_in_los_2nd_order = no
270 include_lensing_in_los_2nd_order = no
271
272 Abracadabra
273 use_delta_tilde_in_los = yes
274 use_delta_tilde_approx_in_los = no
275
276 Should we include the Sachs-Wolfe and integrated Sachs-Wolfe effects at second
277 order? Note that turning on the SW or the ISW effect will automatically turn
278 off the include_metric_in_los_2nd_order option.
279 include_sachs_wolfe_in_los_2nd_order = no
```

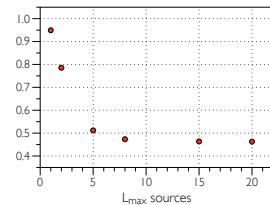
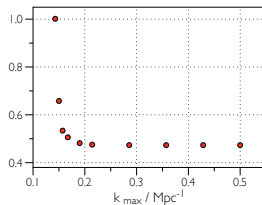
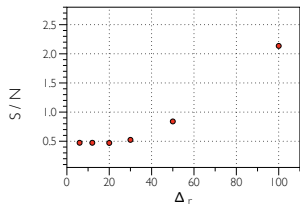
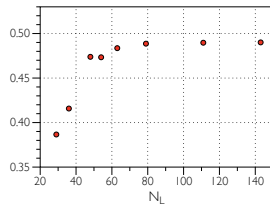
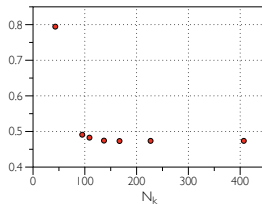
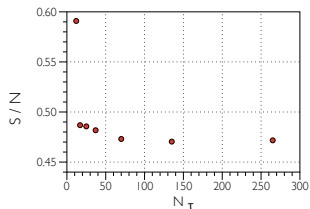
For a run computing intrinsic non-Gaussianity

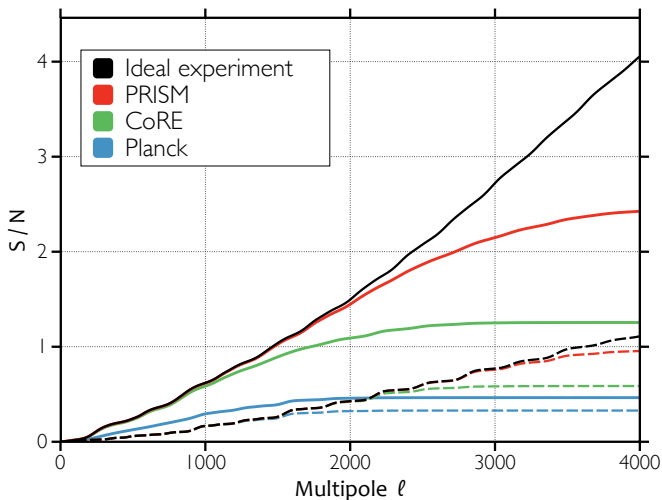
- **perturbations2.c**: Solve the differential system from the deep radiation era, till last scattering
 - Instead of $T(k)$ on a grid in k , we have $T(k, k_1, k_2)$ on a 3-D grid
 - Cannot focus on scalar modes alone
 - Many source terms that are constructed from CLASS output
- **perturbations2.c**: Build the line-of-sight sources
 - Sources cannot be recast into $l = 0$, but exist for all multipoles
- **transfer2.c** and **bessel2.c**: Line-of-sight integration is used to evolve the photon perturbations until today
- **spectra2.c** / **bispectra2.c**: Compute spectrum or bispectrum
 - Integration of bispectra for a completely general shape
- **fisher.c**: Compute the f_{NL} contamination and signal-to-noise

The second order computations are very complicated needing a lengthy code. To reduce the risk of mistakes we performed many test with **SONG**:

- We match many analytical limits for the sub horizon transfer functions
Goroff, Grinstein, et al., 1986; Makino, Sasaki & Suto, 1992; Jain & Bertschinger, 1994; Bernardeau et al., 2002; Bartolo, Matarrese & Riotto 2006; Matarrese, Mollerach & Bruni 1998; Boubeker et al. 2009
- Match to the squeezed limit analytic approximation for the bispectrum (Assumes CDM domination and instantaneous recombination)
Creminelli et al. 2011; Bartolo et al. 2011; Lewis 2012
- Performed detailed convergence and consistence checks
- Code is double checked against independent codes
Beneke, Fidler & Klingmuller 2011; Huang & Vernizzi 2013; Su & Shellard 2013

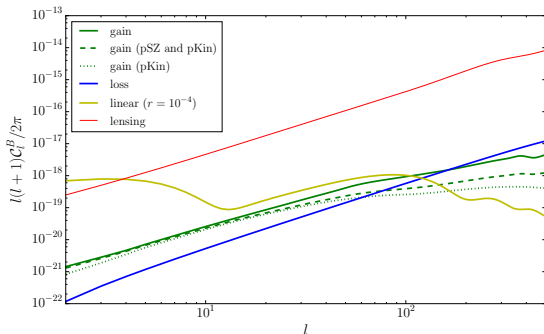
Convergence tests





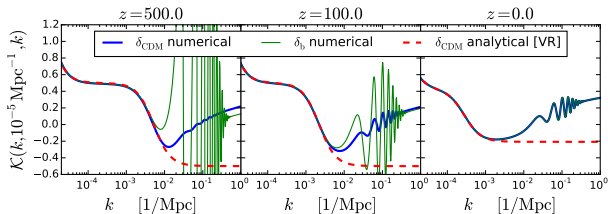
Late effects require an alternative approach as the amount of multipoles becomes impossible to include. Individual treatment is required

- Lensing: expand in deflection angles
- Redshift: modify LOS-sources plus analytical residual
- Reionisation: blurring + low- l residuals

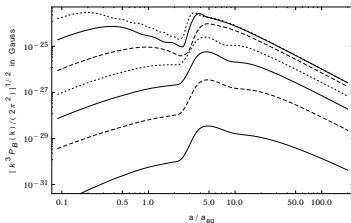


Further Applications

SONG can compute matter perturbations beyond linear perturbation theory



SONG can solve other perturbations such as magnetic fields



- **SONG** is a tool to simulate the Universe in full general relativity to second order in perturbation theory
- Many applications for CMB, but also in other areas of cosmology
- Based on CLASS, inherits the structure and is comparably easy to pick up if you already know CLASS
- Made easy to modify; more than 10.000 lines of comments
- Interested to support further applications and developments
 - Spectral distortions
 - 21-cm
 - Intensity mapping