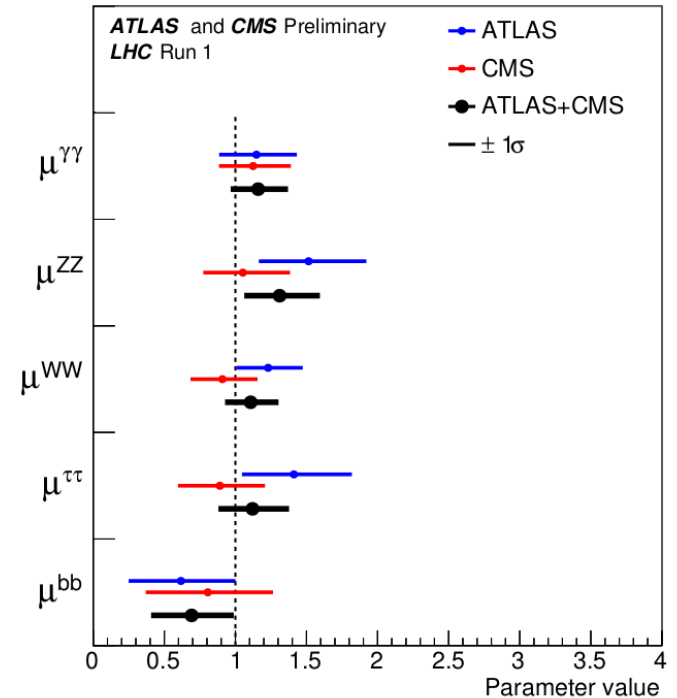
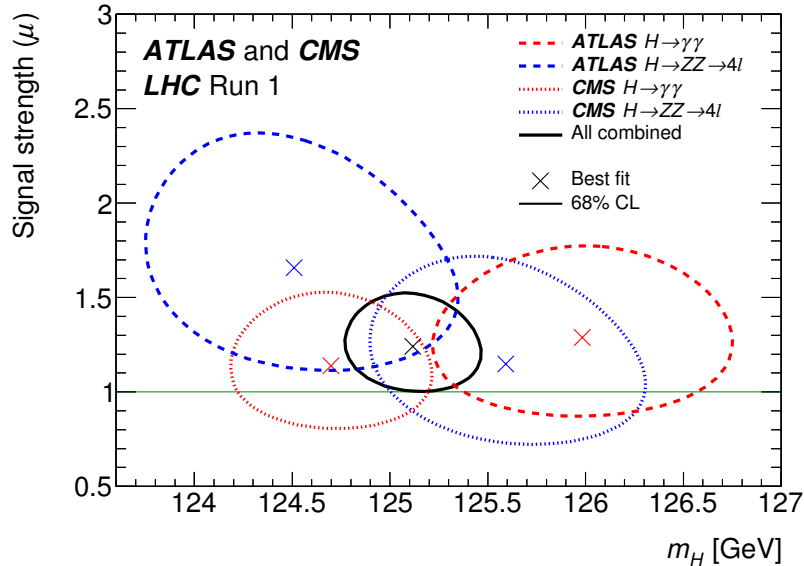


Constraints from LHC-data on Sleptons in scenarios with light ν_R

Werner Porod

Universität Würzburg

in coll. with N. Cerna-Velazco, T. Faber, J. Jones-Perez



$$m_H = 125.09 \pm 0.21 \text{ (stat)} \pm 0.11 \text{ (sys)} \text{ GeV}$$

run 1, PRL **114** (2015) 191803

$$\text{ATLAS: } m_H = 124.98 \pm 0.19 \text{ (stat)} \pm 0.21 \text{ (sys)} \text{ GeV}$$

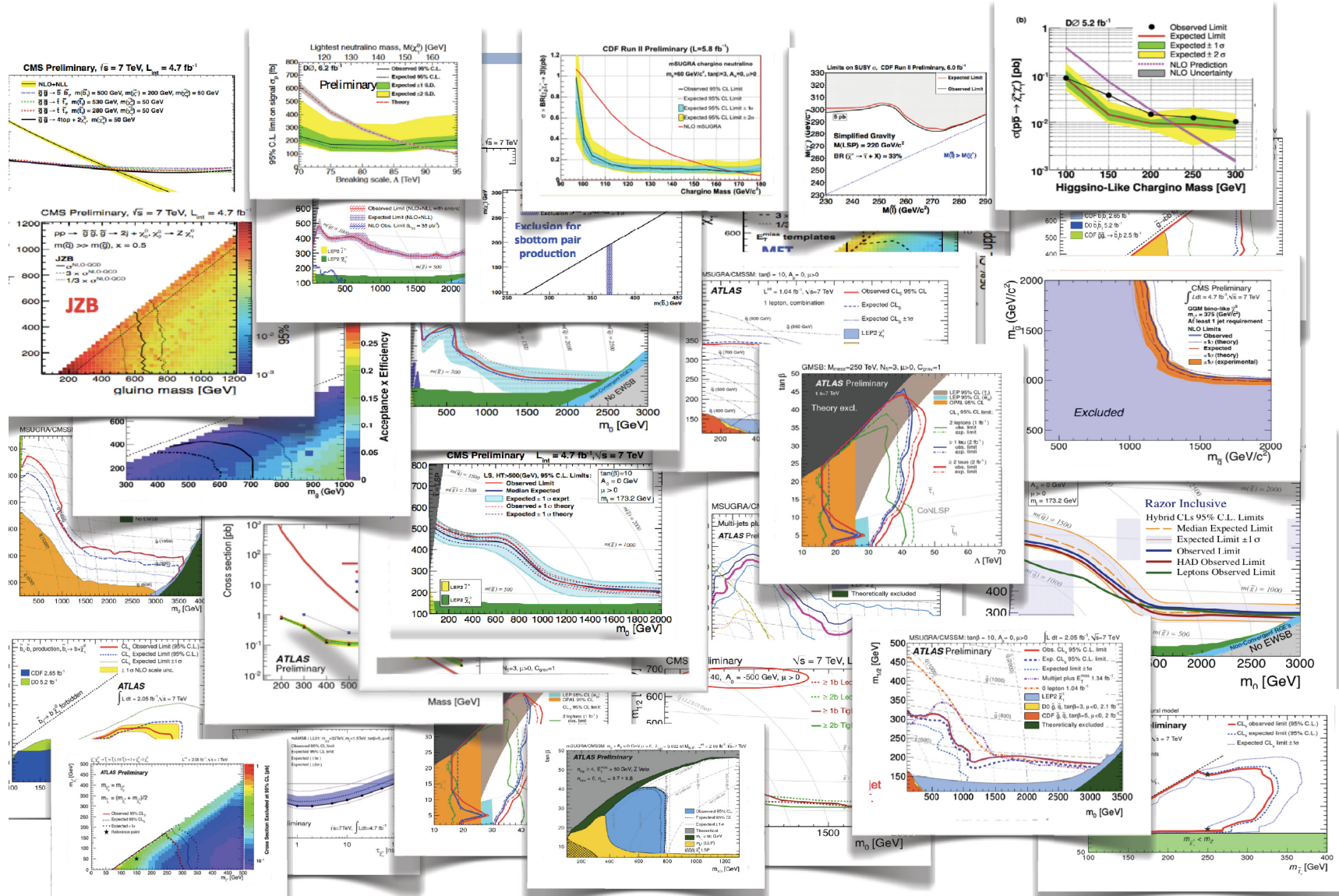
$$\text{CMS: } m_H = 125.26 \pm 0.20 \text{ (stat)} \pm 0.08 \text{ (sys)} \text{ GeV}$$

see e.g. talk by A.-M. Magnan, ALPS 2018

$$(125 \text{ GeV})^2 \simeq m_Z^2 + (86 \text{ GeV})^2 \Rightarrow \text{large corrections within MSSM}$$

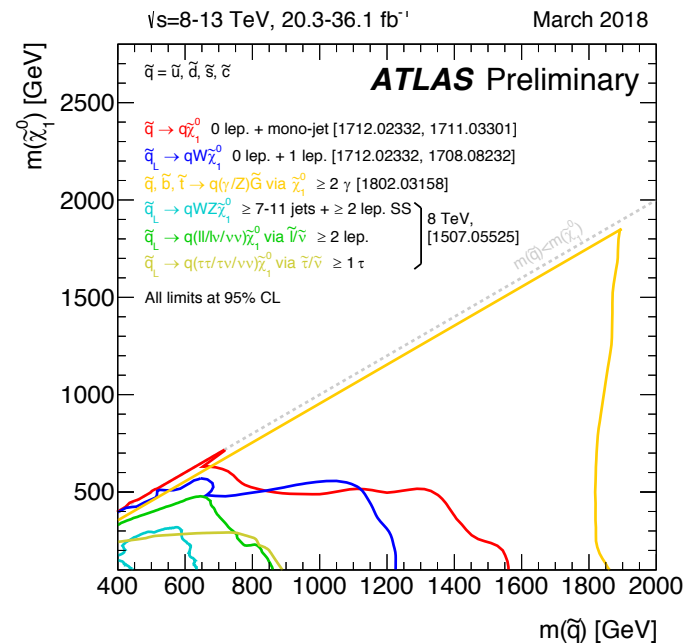
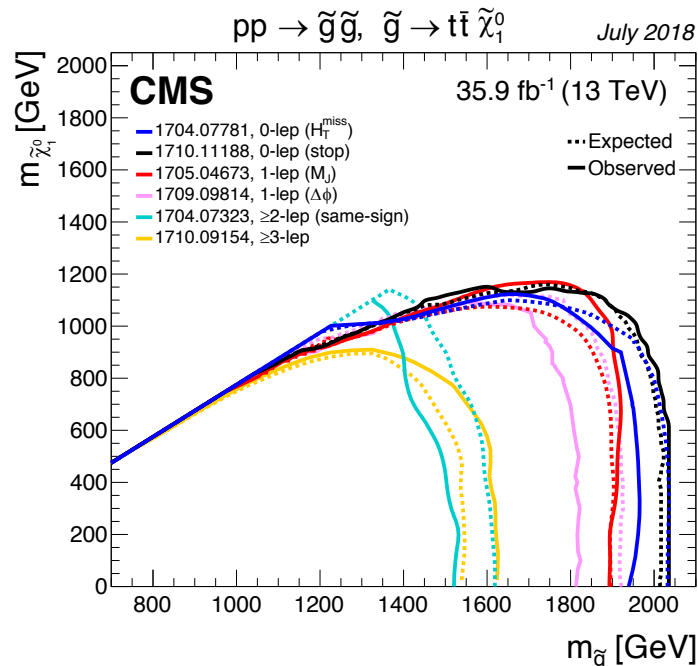
ATLAS-CONF-2015-044

CMS-PAS-HIG-15-002



Summary of generic RPC searches

- In **simplified model approach** (depending on decay mode and/or mass splittings):
 - $M_{\tilde{g}} \approx \mathcal{O}(1 \text{ TeV}) - \mathcal{O}(2 \text{ TeV}) @ 95\% \text{ CL}$
 - $M_{\tilde{q}} \approx \mathcal{O}(0.5 \text{ TeV}) - \mathcal{O}(1.5 \text{ TeV}) @ 95\% \text{ CL}$
 - $M_{\tilde{t}} \approx \mathcal{O}(0.7 \text{ TeV}) - \mathcal{O}(1.1 \text{ TeV}) @ 95\% \text{ CL}$
- } Can be even worse in some corners of simplified model space.



$m_h = 125.2 \text{ GeV} \Rightarrow$ large loop contributions
 \Rightarrow heavy stops and/or large left-right mixing for stops

- GMSB: $m_{\tilde{t}_1} \gtrsim 6 \text{ TeV}$,
 M. A. Ajaib, I. Gogoladze, F. Nasir, Q. Shafi, arXiv:1204.2856
 more complicated models based on P. Meade, N. Seiberg and D. Shih,
 arXiv:0801.3278 \Rightarrow allow additional terms
 e.g. S. Knappen, D. Redigolo, arXiv:1606.07501 $m_{\tilde{t}_1} \simeq m_{\tilde{b}_1} \gtrsim 1 \text{ TeV}$ if
 $M_{\text{mess}} \gtrsim 10^{15} \text{ GeV}$
- CMSSM, NUHM models: $|A_0| \simeq 2m_0$,
 H. Baer, V. Barger and A. Mustafayev, arXiv:1112.3017; M. Kadastik *et al.*,
 arXiv:1112.3647; O. Buchmueller *et al.*, arXiv:1112.3564; J. Cao, Z. Heng, D. Li,
 J. M. Yang, arXiv:1112.4391; L. Aparicio, D. G. Cerdeno, L. E. Ibanez,
 arXiv:1202.0822; J. Ellis, K. A. Olive, arXiv:1202.3262; ...
 CMSSM fit to data P. Bechtle *et al.*, arXiv:1508.05951: best fit point with
 $m_{\tilde{g}}, m_{\tilde{q}} \gtrsim 2 \text{ TeV}$, $m_{\tilde{l}_R} \simeq 600 \text{ GeV}$, $m_{\tilde{\chi}_1^0} \simeq 450 \text{ GeV}$
- general high scale models: $A_0 \simeq -(1-3) \max(M_{1/2}, m_{Q_3}, m_{U_3}) @ M_{GUT}$
 among other cases, details in F. Brümmer, S. Kraml and S. Kulkarni, arXiv:1204.5977

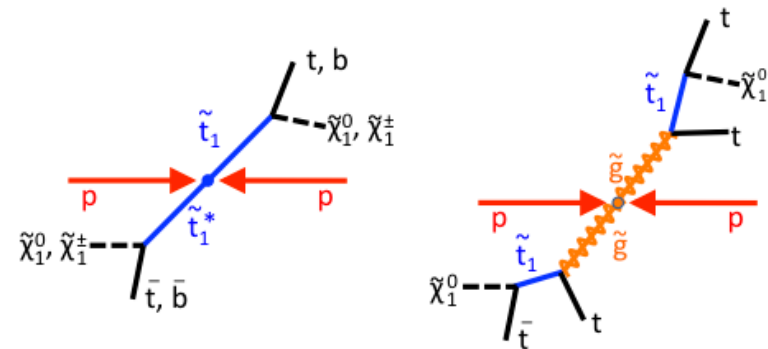
several studies: S. Sekmen et al., arXiv:1109.5119; A. Arbey, M. Battaglia, A. Djouadi, F. Mahmoudi, arXiv:1211.4004; M. Cahill-Rowley, J. Hewett, A. Ismail, T. Rizzo, arXiv:1308.0297 ...

- generic signatures are well known: multi-lepton, multi-jets + missing E_T
- sub-class of general MSSM: 'natural SUSY'
see e.g. M. Papucci, J. T. Ruderman and A. Weiler, arXiv:1110.6926;
H. Baer, V. Barger, P. Huang, A. Mustafayev, X. Tata, arXiv:1207.3343
keep only SUSY particles light needed for 'natural Higgs':

$$\tilde{t}_1, \tilde{b}_1, \tilde{g}, \tilde{\chi}_{1,2}^0 \simeq \tilde{h}_{1,2}^0, \tilde{\chi}_1^+ \simeq \tilde{h}^+$$

$$\Rightarrow 100 \text{ MeV} \lesssim m_{\tilde{\chi}_1^+} - m_{\tilde{\chi}_1^0} \simeq m_{\tilde{\chi}_2^0} - m_{\tilde{\chi}_1^0} \lesssim 5 - 10 \text{ GeV}$$

$$\begin{aligned} \tilde{g} &\rightarrow \tilde{t}_1 t, \tilde{b}_1 b \\ \tilde{t}_1 &\rightarrow t \tilde{\chi}_{1,2}^0, b \tilde{\chi}_1^+, W^+ \tilde{b}_1 \\ \tilde{b}_1 &\rightarrow b \tilde{\chi}_{1,2}^0, t \tilde{\chi}_1^-, W^- \tilde{t}_1 \end{aligned}$$



BRs depend on the nature of \tilde{t}_1 and \tilde{b}_1

Higgsino mass: $\mu + \mu'$ with soft SUSY breaking parameter: $\mathcal{L} = -\mu' \tilde{H}_d \tilde{H}_u$

(G. G. Ross, K. Schmidt-Hoberg and F. Staub, arXiv:1701.03480)

$$\mathcal{W}_{eff} = \mathcal{W}_{MSSM} + \frac{1}{2}(M_R)_{ij} \hat{\nu}_{R,i} \hat{\nu}_{R,j} \\ + (Y_\nu)_{ij} \hat{L}_i \cdot \hat{H}_u \hat{\nu}_{R,j}$$

$$(Y_\nu)_{l5} = \pm (Z_\ell^{\text{NH}})^* \sqrt{\frac{2m_3 M_5}{v_u}} \cosh \gamma_{56} e^{\mp i\theta_{56}}$$

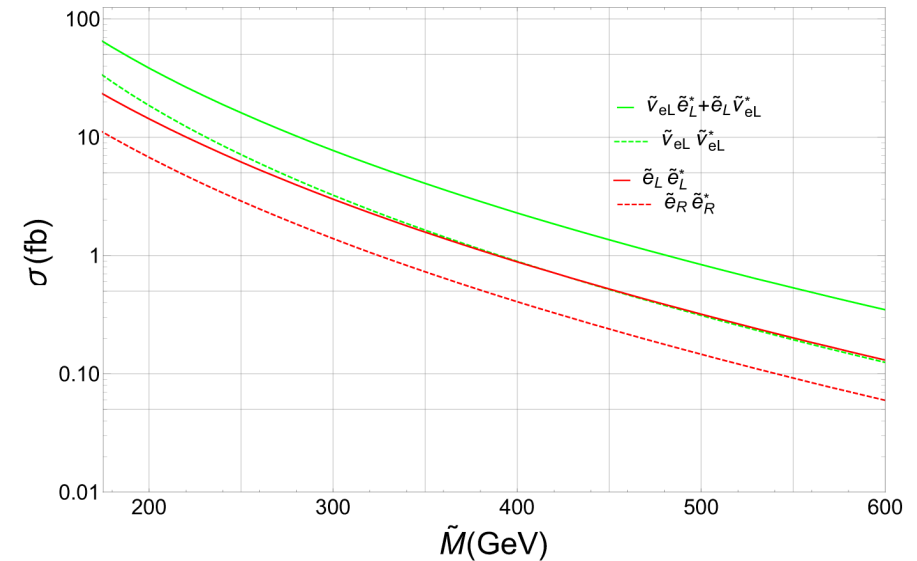
$$(Y_\nu)_{l6} = -i (Z_\ell^{\text{NH}})^* \sqrt{\frac{2m_3 M_6}{v_u}} \cosh \gamma_{56} e^{\mp i\theta_{56}}$$

$$R = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \phi_{56} & \sin \phi_{56} \\ 0 & -\sin \phi_{56} & \cos \phi_{56} \end{pmatrix}$$

$$\phi_{56} \in \mathbb{C}$$

$$m_{\nu_h,i} \simeq M_{i-3}, M_4 = O(\text{keV}), \\ M_5 \simeq M_6 = O(\text{few} - 100 \text{ GeV})$$

search for sleptons



LHC, 13 TeV, tree-level
for searches: \times K-factor 1.17
(B. Fuks et al., arXiv:1304.0790)

dominant decays:

$$\tilde{l}_L \rightarrow l \tilde{\chi}_1^0, \nu \tilde{\chi}_1^-$$

$$\tilde{\nu}_L \rightarrow l^- \tilde{\chi}_1^+, \nu \tilde{\chi}_1^0$$

$$\mathcal{W}_{eff} = \mathcal{W}_{MSSM} + \frac{1}{2} (M_R)_{ij} \hat{\nu}_{R,i} \hat{\nu}_{R,j} + (Y_\nu)_{ij} \hat{L}_i \cdot \hat{H}_u \hat{\nu}_{R,j}$$

$$(Y_\nu)_{l5} = \pm (Z_\ell^{\text{NH}})^* \sqrt{\frac{2m_3 M_5}{v_u}} \cosh \gamma_{56} e^{\mp i\theta_{56}}$$

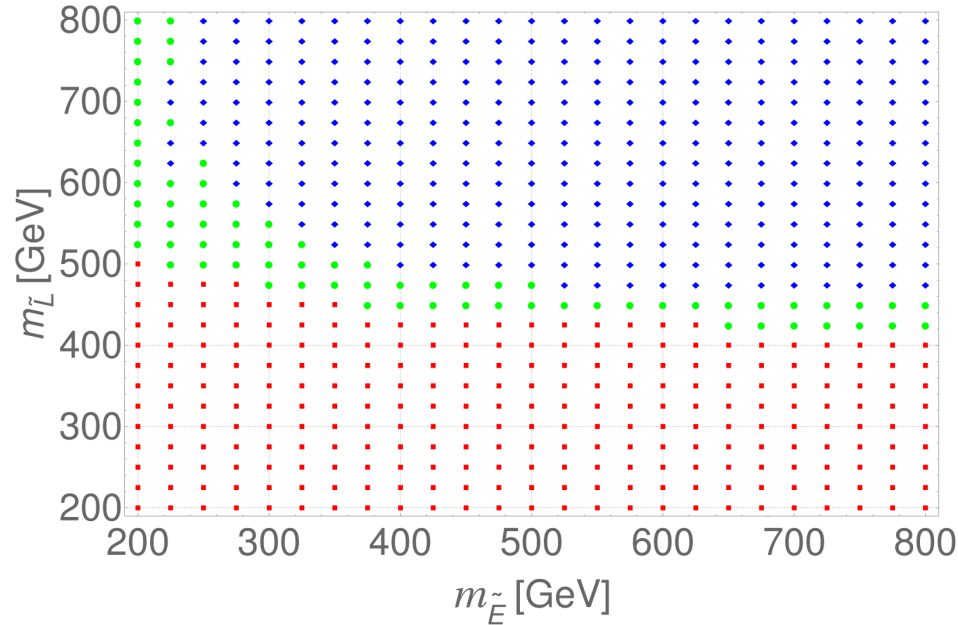
$$(Y_\nu)_{l6} = -i (Z_\ell^{\text{NH}})^* \sqrt{\frac{2m_3 M_6}{v_u}} \cosh \gamma_{56} e^{\mp i\theta_{56}}$$

$$R = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \phi_{56} & \sin \phi_{56} \\ 0 & -\sin \phi_{56} & \cos \phi_{56} \end{pmatrix}$$

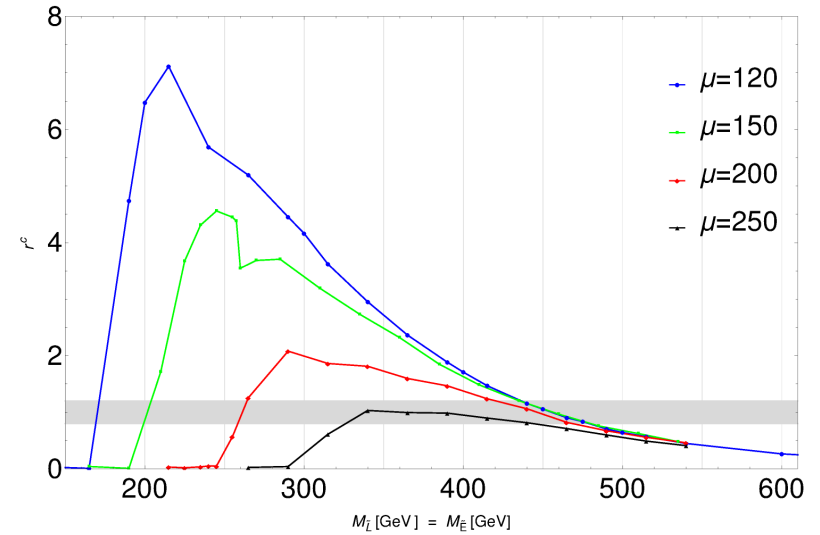
$$\phi_{56} \in \mathbb{C}$$

$$m_{\nu_h, i} \simeq M_{i-3}, M_4 = O(\text{keV}),$$

$$M_5 \simeq M_6 = O(\text{few} - 100 \text{ GeV})$$



$\mu = 120 \text{ GeV}, \tan \beta = 10$



$m_{\tilde{L}} = m_{\tilde{E}}, \tan \beta = 10$

■ excluded, ● ambiguous, ◇ allowed

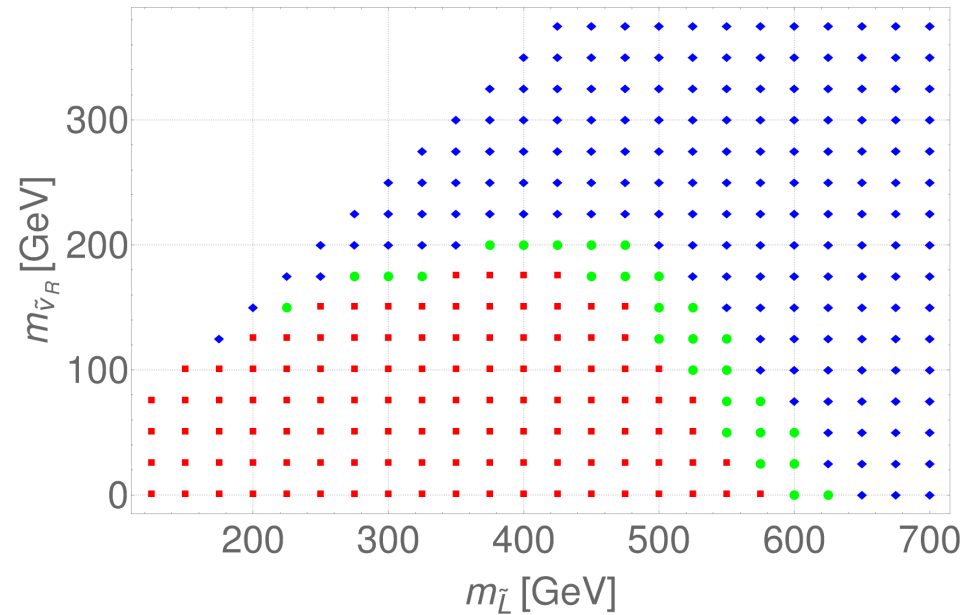
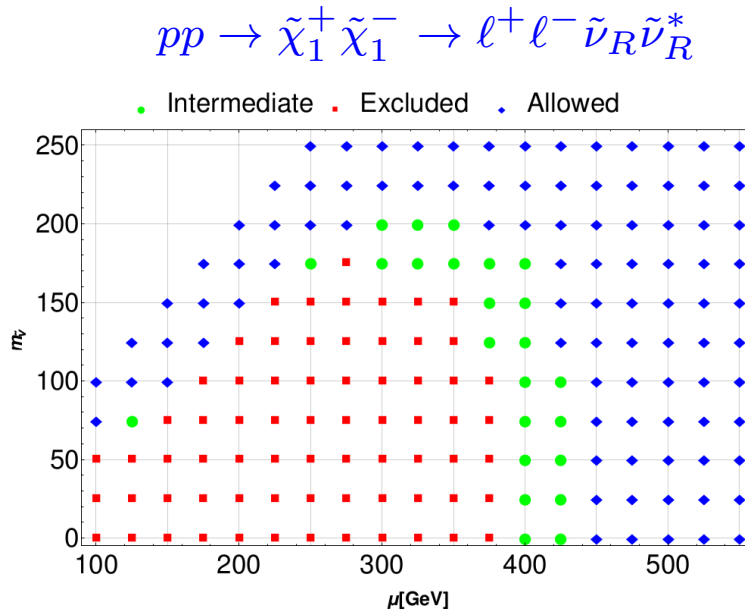
8+13 TeV data (13.9 fb^{-1})

using CheckMATE 2.0

Nh. Cerna-Velazco, Th. Faber, J. Jones-Perez, WP

arXiv:1705.06583

additional constraint



8+13 TeV data (13.9 fb^{-1})

using CheckMATE 2.0

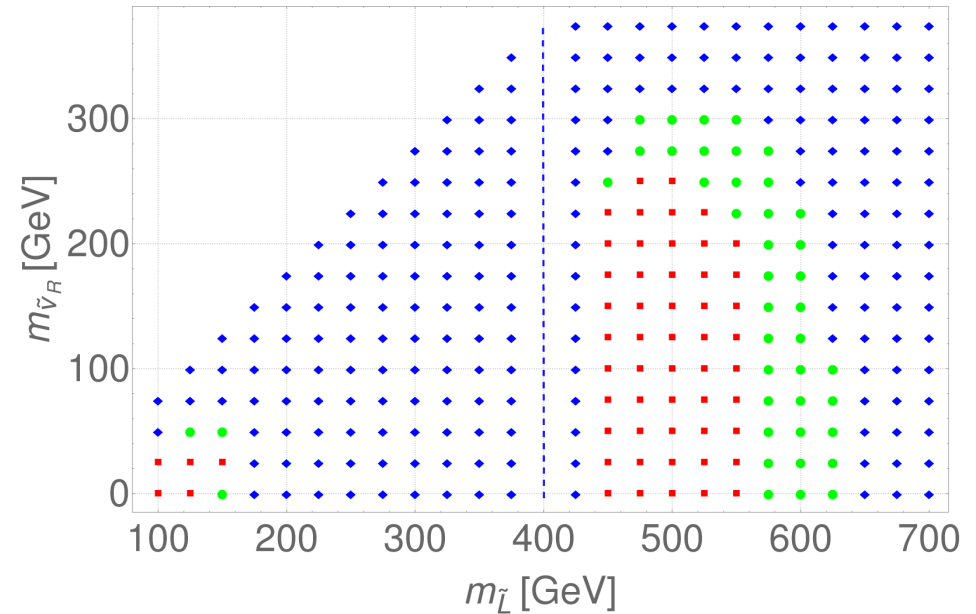
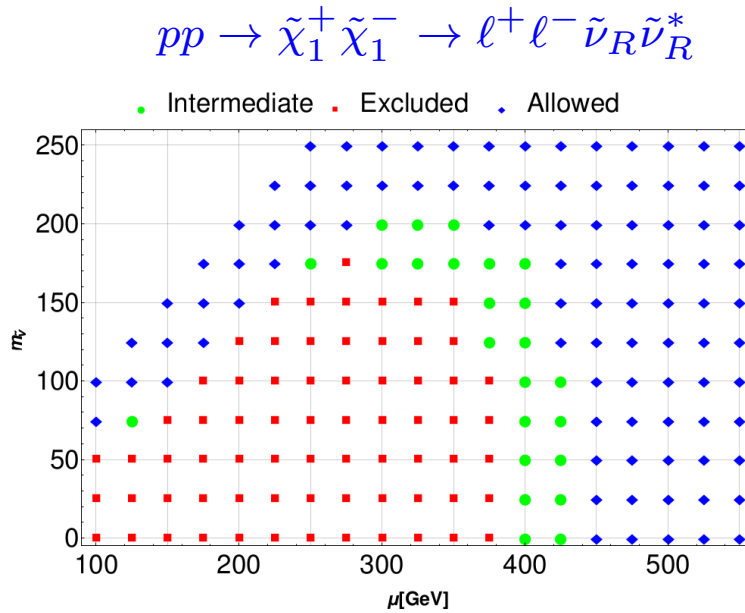
$m_{\nu_R} = 20 \text{ GeV}$

$$\mu = 25 + m_{\tilde{\nu}} < m_{\tilde{t}} \simeq m_{\tilde{L}} = m_{\tilde{E}}$$

$$M_1 = M_2 = 2 \text{ TeV}, \tan \beta = 6$$

Nh. Cerna-Velazco, Th. Faber, J. Jones-Perez, WP arXiv:1705.06583

additional constraint



8+13 TeV data (13.9 fb^{-1})

using CheckMATE 2.0

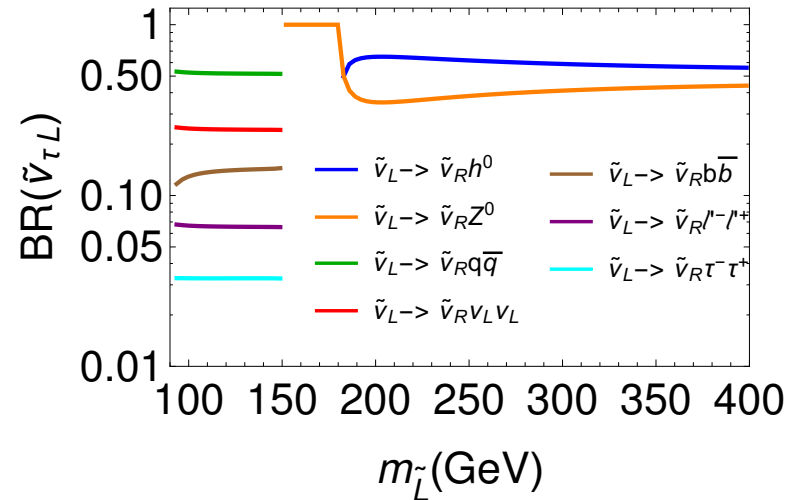
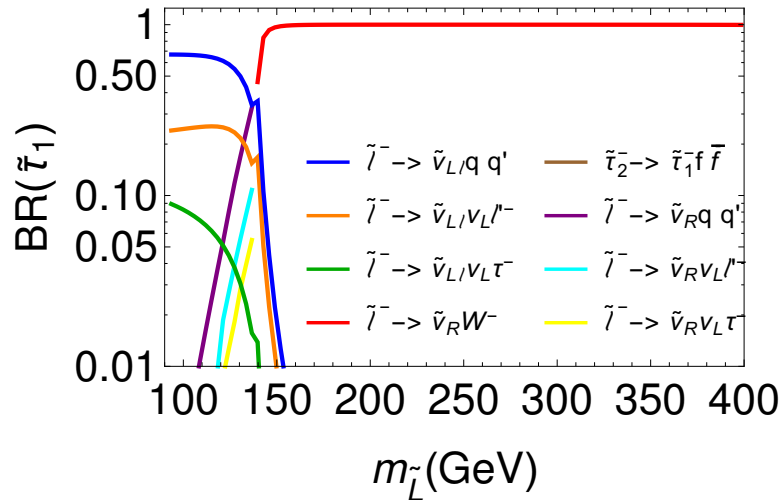
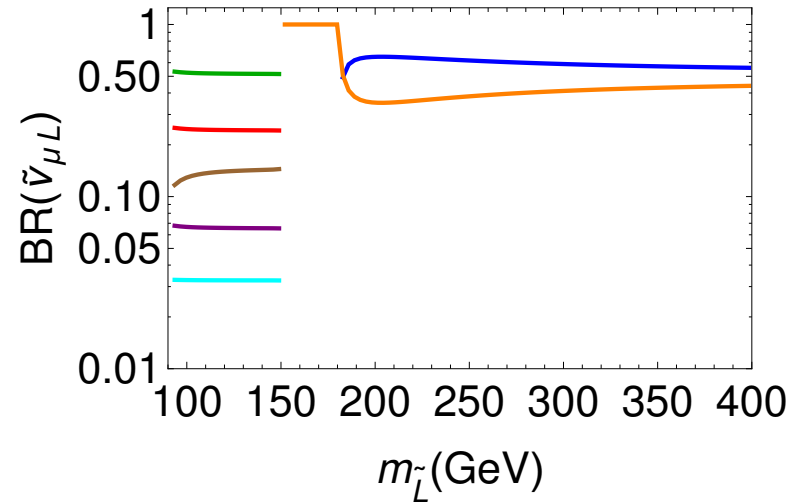
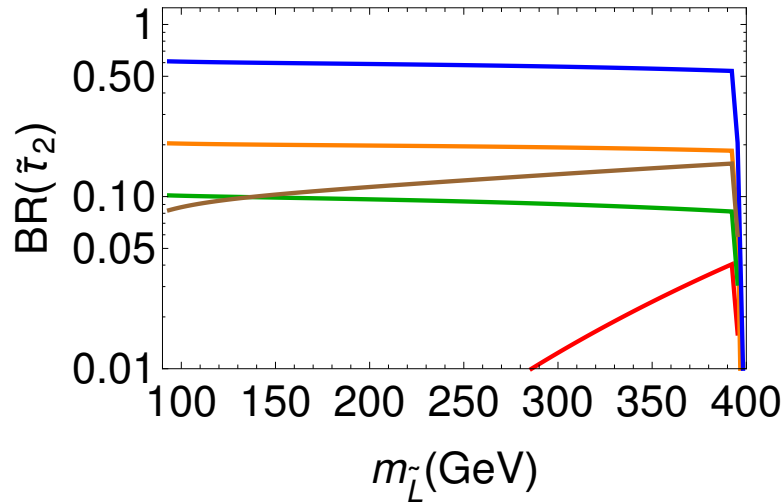
$m_{\nu_R} = 20 \text{ GeV}$

$\mu = 400 \text{ GeV}, m_{\tilde{l}} \simeq m_{\tilde{L}} = m_{\tilde{E}}$

$M_1 = M_2 = 2 \text{ TeV}, \tan \beta = 6$

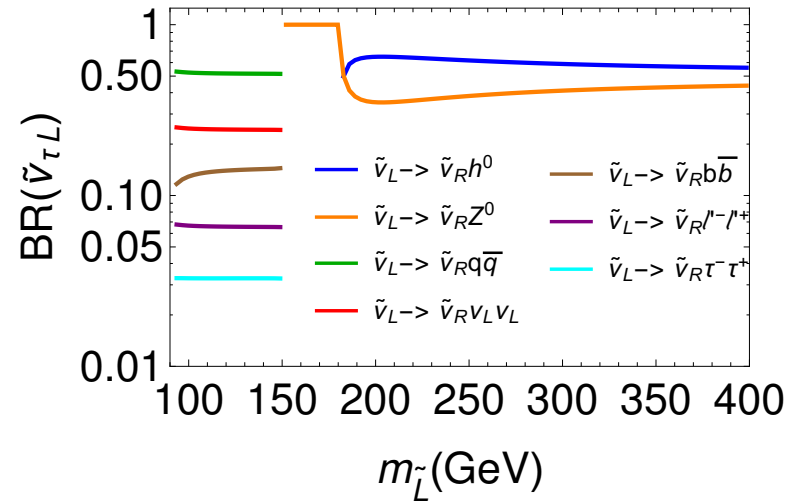
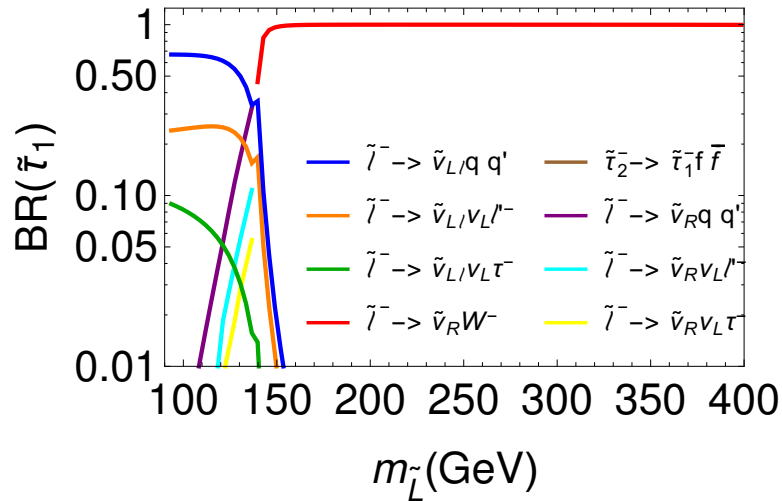
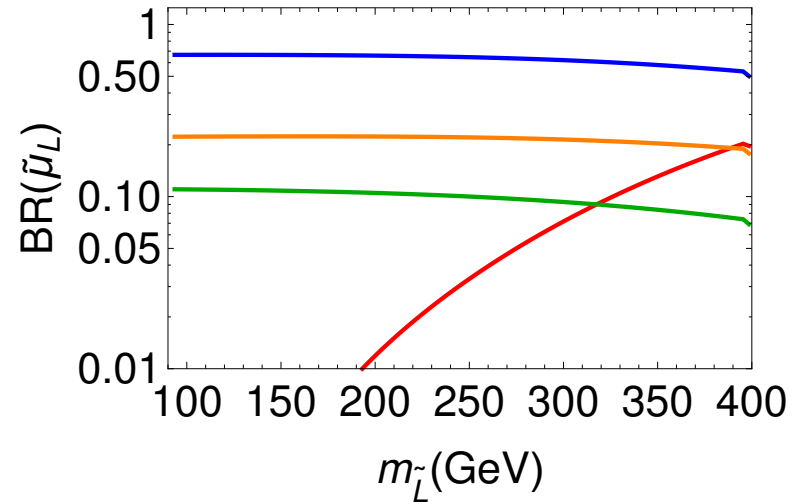
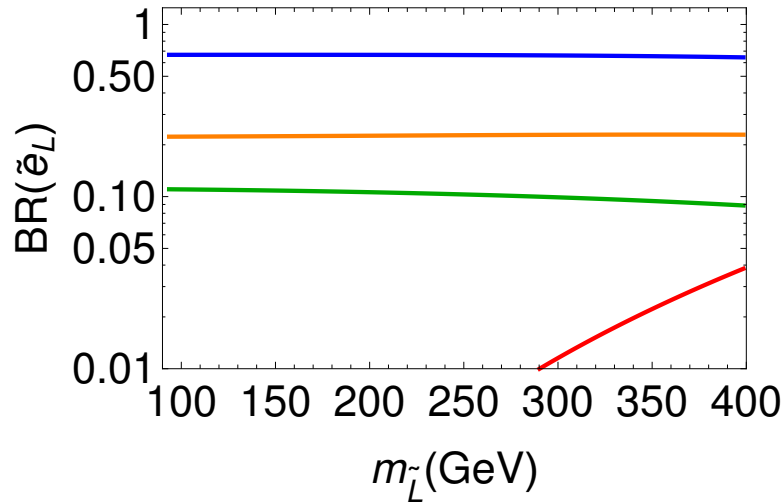
Nh. Cerna-Velazco, Th. Faber, J. Jones-Perez, WP arXiv:1705.06583

for $\mu = 400 \text{ GeV} > m_{\tilde{L}} = m_{\tilde{E}}, \tan \beta = 6, M_1, M_2 \geq 500 \text{ GeV}$

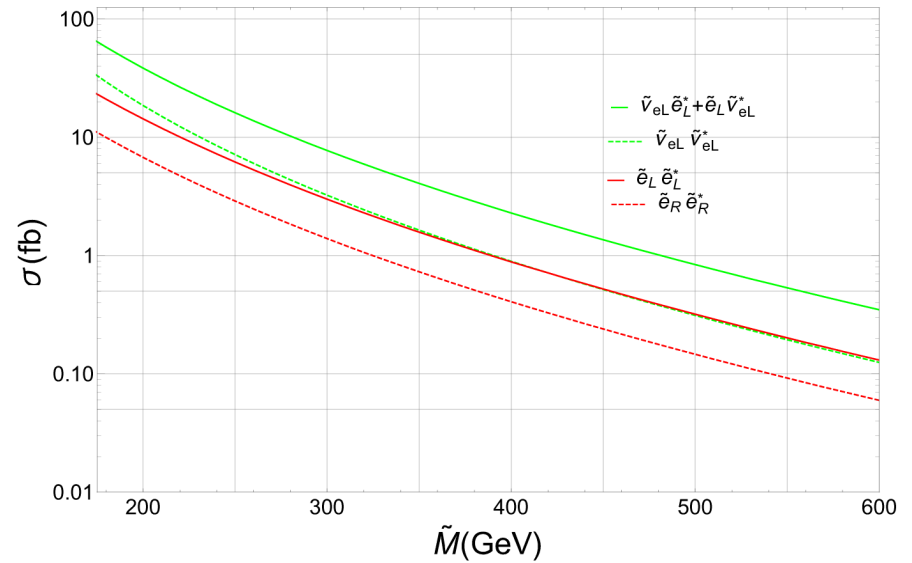
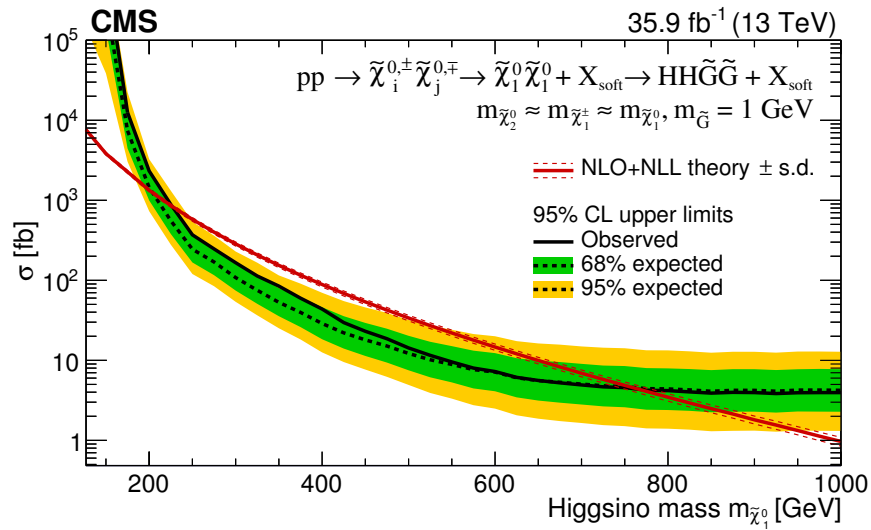


Nh. Cerna-Velazco, Th. Faber, J. Jones-Perez, WP arXiv:1705.06583

for $\mu = 400 \text{ GeV} > m_{\tilde{L}} = m_{\tilde{E}}, \tan \beta = 6, M_1, M_2 \geq 500 \text{ GeV}$



Nh. Cerna-Velazco, Th. Faber, J. Jones-Perez, WP arXiv:1705.06583



CMS arXiv:1709.04896

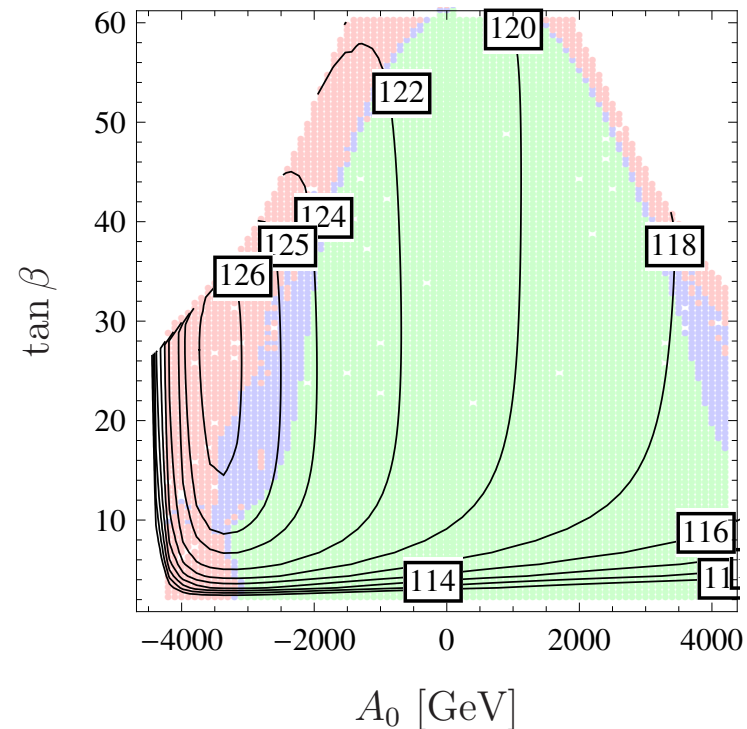
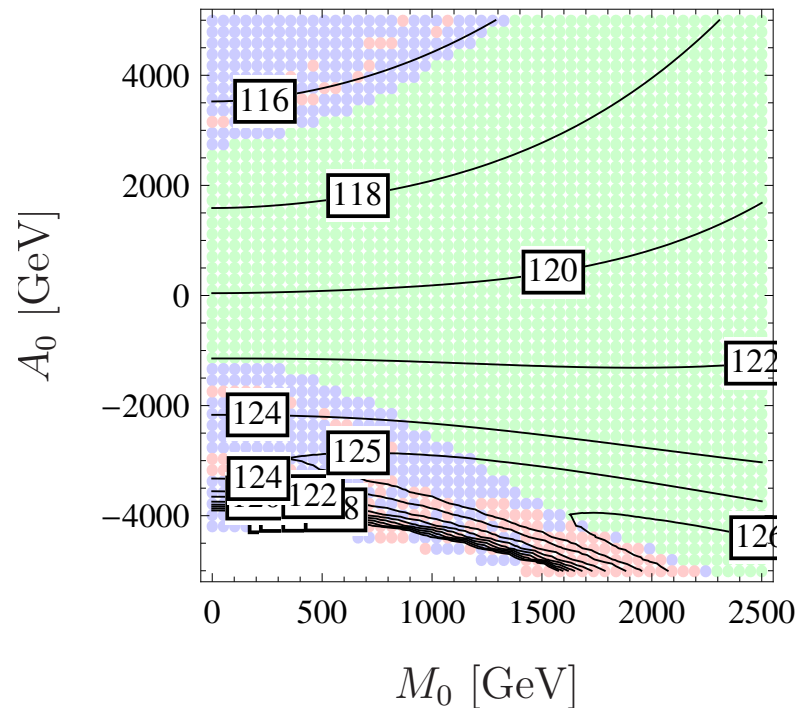
see also talk by Reina Camacho Toro

- ‘Natural SUSY’: take only those states light which contribute to EWSB: $\tilde{h}^{0,\pm}, \tilde{t}_1, \tilde{g}, \tilde{b}_i$
disadvantage: cannot explain dark matter relic density
- consider scenarios with light ν_R and a minimal inverse seesaw
 $\Rightarrow \tilde{\nu}_R$ LSP potential DM-candidate, no direct DM constraint apply
- $m_{\tilde{h}^+} \lesssim 400$ GeV excluded if $m_{\tilde{h}^+} - m_{\tilde{\nu}_R} \gtrsim 150$ GeV
- 2-body decays dominate: $m_{\tilde{E}} \gtrsim 200$ GeV & $m_{\tilde{L}} \gtrsim 600$ GeV if sufficiently hard leptons in the final state
- 3-body decays dominate: \tilde{l}_L with masses of about 200 GeV still allowed

Outlook:

- include newer analyses, in particular those containing b -jets with somewhat compressed spectra

- SUSY models contain many scalars \Rightarrow complicated potential
- usually some parameters (μ, B) are chosen to obtain correct EWSB
- does not exclude the existence of other minima breaking charge and/or color!



$$M_{1/2} = 1 \text{ TeV}, \tan \beta = 10, \mu > 0$$

$$M_{1/2} = M_0 = 1 \text{ TeV}$$

J.E. Camargo-Molina, B. O'Leary, W.P., F. Staub, arXiv:1309.7212

Constraints from Z -width, invisible width

$$\left| 1 - \sum_{ij=1, i \leq j}^3 \left| \sum_{k=1}^3 U_{ik}^\nu U_{jk}^{\nu,*} \right|^2 \right| < 0.009$$

dominant decays

$$\nu_j \rightarrow W^\pm l^\mp$$

$$\nu_j \rightarrow Z \nu_i$$

$$\nu_j \rightarrow h_k \nu_i$$

if $m_\nu > m_h$

$$BR(\nu_j \rightarrow W^\pm l^\mp) : BR(\nu_j \rightarrow Z \nu_i) : BR(\nu_j \rightarrow h_k \nu_i) \simeq 0.5 : 0.25 : 0.25$$