

(In)dependence of various LFV observables in the non-minimal SUSY

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in collaboration with D. Stöckinger and H. Stöckinger-Kim

R-symmetry

- ✦ additional symmetry of the SUSY algebra allowed by the Haag - Łopuszański - Sohnius theorem
- ✦ for N=1 it is a global $U_R(1)$ symmetry under which the SUSY generators are charged
- ✦ implies that the spinorial coordinates are also charged $Q_R(\theta) = 1, \theta \rightarrow e^{i\alpha}\theta$
- ✦ superpotential example

$$\mathcal{L} \ni \int d^2\theta W$$

- ✦ Superpotential is polynomial in fields. For W to transform homogeneously superfields must have definite R-charges

$$e^{i\alpha Q_R} \Phi = e^{i\alpha Q_R} \phi(y) + \sqrt{2}\theta\psi(y) + \theta\theta F(y)$$

- ✦ Similarly one can work out other parts of the Lagrangian

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(we want it to be)
R-invariant

$$\longrightarrow \mathcal{L} \ni \int d^2\theta W$$

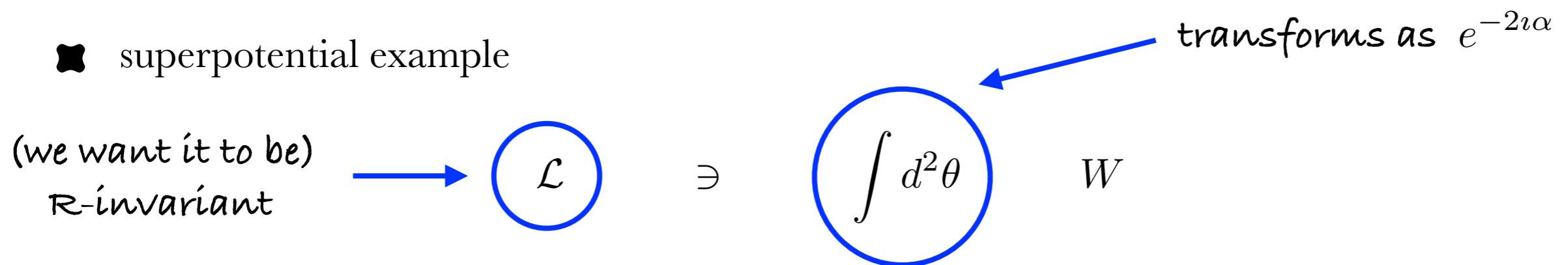
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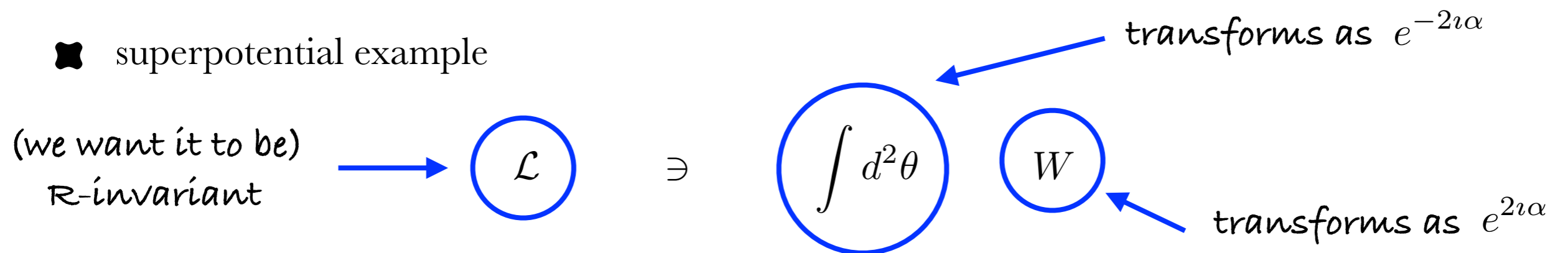
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- ✦ Similarly one can work out other parts of the Lagrangian

Low-energy R-symmetry realization

✘ Different possible models that one can construct

✘ “Natural” choice

$$e^{i\alpha Q_R} \Phi = e^{i\alpha Q_R} \phi(y) + \sqrt{2}\theta\psi(y) + \theta\theta F(y)$$

leptons and quarks	$Q_R = 1$	$Q_R = 1$	$Q_R = 0$
Higgs	$Q_R = 0$	$Q_R = 0$	$Q_R = -1$

✘ Good: no baryon and lepton number violating terms

✘ Bad: No Majorana masses for higgsinos and gauginos

One way to fix it: [Dirac masses](#)
 Minimal R-Symmetric Supersymmetric Standardmodel (MRSSM)
Kribs et.al. arXiv:0712.2039

		$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_R$
Additional fields:	Singlet \hat{S}	1	1	0	0
	Triplet \hat{T}	1	3	0	0
	Octet \hat{O}	8	1	0	0
	R-Higgses \hat{R}_u	1	2	-1/2	2
	\hat{R}_d	1	2	1/2	2

$$W = \mu_d \hat{R}_d \hat{H}_d + \mu_u \hat{R}_u \hat{H}_u$$

$$+ \Lambda_d \hat{R}_d \hat{T} \hat{H}_d + \Lambda_u \hat{R}_u \hat{T} \hat{H}_u + \lambda_d \hat{S} \hat{R}_d \hat{H}_d + \lambda_u \hat{S} \hat{R}_u \hat{H}_u$$

$$- Y_d \hat{d} \hat{q} \hat{H}_d - Y_e \hat{e} \hat{l} \hat{H}_d + Y_u \hat{u} \hat{q} \hat{H}_u$$

MSSM vs. MRSSM

✦ superpotencial

$$\mu \hat{H}_u \hat{H}_d \quad \text{!}$$

$$-Y_d \hat{d} \hat{q} \hat{H}_d - Y_e \hat{e} \hat{l} \hat{H}_d + Y_u \hat{u} \hat{q} \hat{H}_u \quad \text{✓}$$

✦ soft-SUSY breaking terms

B_μ - term ✓

soft scalar masses ✓

Majorana gaugino masses !

A - terms !

✦ superpotencial

$$\mu_d \hat{R}_d \hat{H}_d + \mu_u \hat{R}_u \hat{H}_u$$

$$-Y_d \hat{d} \hat{q} \hat{H}_d - Y_e \hat{e} \hat{l} \hat{H}_d + Y_u \hat{u} \hat{q} \hat{H}_u$$

$$\Lambda_d \hat{R}_d \hat{T} \hat{H}_d + \Lambda_u \hat{R}_u \hat{T} \hat{H}_u + \lambda_d \hat{S} \hat{R}_d \hat{H}_d + \lambda_u \hat{S} \hat{R}_u \hat{H}_u$$

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Kribs, Popitz, Weiter (2008)

MSSM vs. MRSSM

✦ superpotencial

$$\mu \hat{H}_u \hat{H}_d$$



$$-Y_d \hat{d} \hat{q} \hat{H}_d - Y_e \hat{e} \hat{l} \hat{H}_d + Y_u \hat{u} \hat{q} \hat{H}_u$$



✦ soft-SUSY breaking terms

□ B_μ - term



□ soft scalar masses



□ Majorana gaugino masses



□ A - terms



✦ superpotencial

$$\mu_d \hat{R}_d \hat{H}_d + \mu_u \hat{R}_u \hat{H}_u$$

$$-Y_d \hat{d} \hat{q} \hat{H}_d - Y_e \hat{e} \hat{l} \hat{H}_d + Y_u \hat{u} \hat{q} \hat{H}_u$$

$$\Lambda_d \hat{R}_d \hat{T} \hat{H}_d + \Lambda_u \hat{R}_u \hat{T} \hat{H}_u + \lambda_d \hat{S} \hat{R}_d \hat{H}_d + \lambda_u \hat{S} \hat{R}_u \hat{H}_u$$

✦ soft-SUSY breaking terms

□ B_μ -term

□ soft scalar masses

□ Dirac gaugino masses

□ no A-terms

One way to fix it: [Dirac masses](#)

Minimal R-Symmetric Supersymmetric Standardmodel (MRSSM)

Kribs et.al. arXiv:0712.2039

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Kribs, Popitz, Weiter (2008)

Particle content summary: MSSM vs. MRSSM

different number of physical states

completely new states

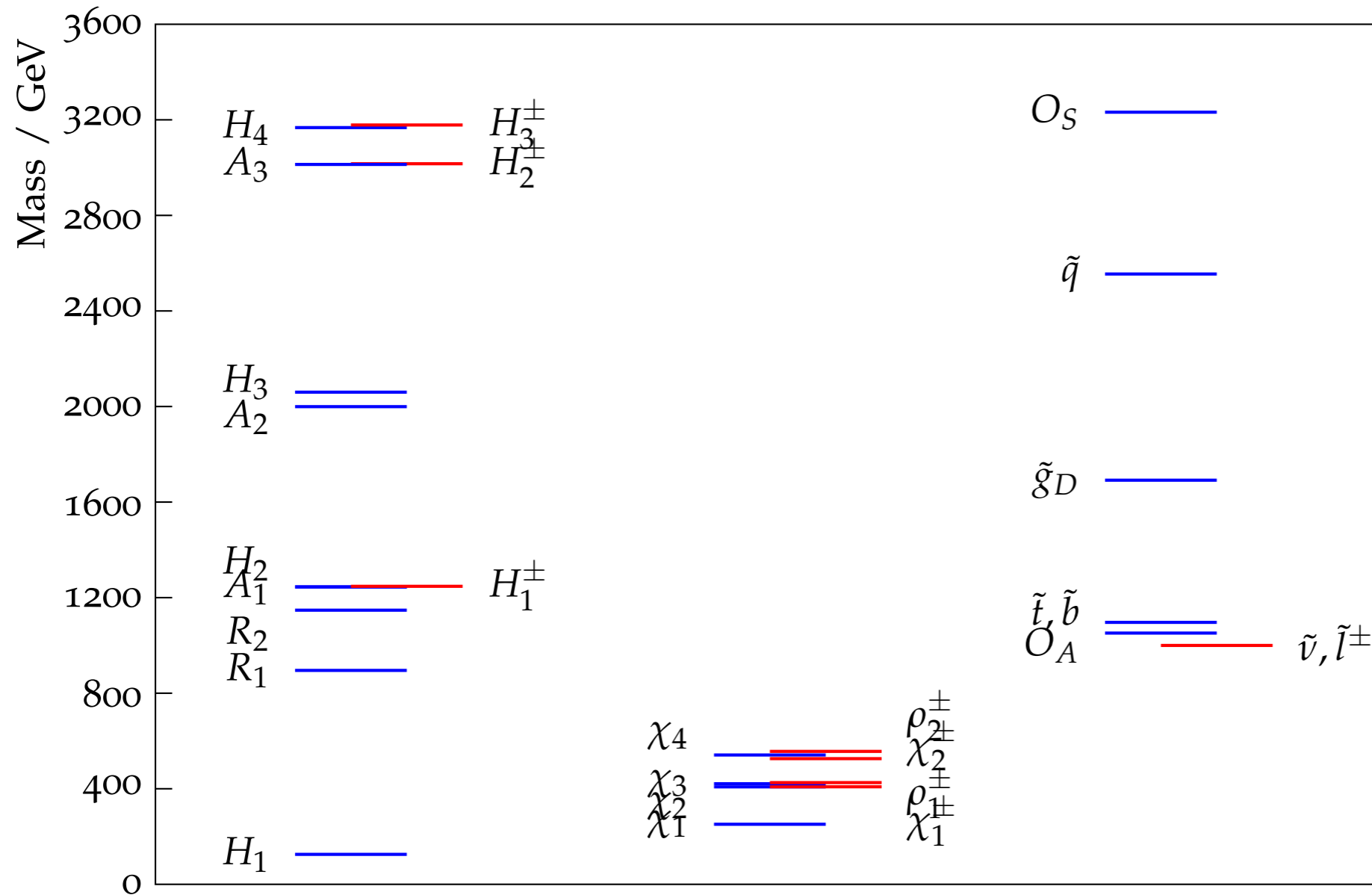
	Higgs			charginos	R-Higgs		sgluon
	CP-even	CP-odd	charged		neutral	charged	
MSSM	2	1	1	2	0	0	0
MRSSM	4	3	3	2+2	2	2	1

	neutralino	gluino
MSSM	4	1
MRSSM	4	1

Majorana fermions

Dirac fermions

Exemplary mass spectrum



arXiv:1410.4791
Higgs mass at 1-loop
level + EWPO

arXiv:1504.05386
2-loop corrections to
Higgs mass

arXiv:1511.09334
DM in light single
scenario

arXiv:1707.04557
NLO SQCD corrections
to squark -
(anti)squark pair
production

Previous and future low energy experiments

✦ As the LHC still sees nothing, we look into low energy experiments:

✦ prospects for g-2 measurement

$$a_{\mu}^{\text{exp}} - a_{\mu}^{\text{SM}} = (28.1 \pm 6.3^{\text{exp}} \pm 3.6^{\text{th}}) \times 10^{-10}$$

$$a_{\mu}^{\text{exp}} - a_{\mu}^{\text{SM}} = (??? \pm 1.6^{\text{exp}} \pm 3.4^{\text{th}}) \times 10^{-10}$$

✦ prospect for $\mu \rightarrow e\gamma$

current: 4.2×10^{-13} (MEG)

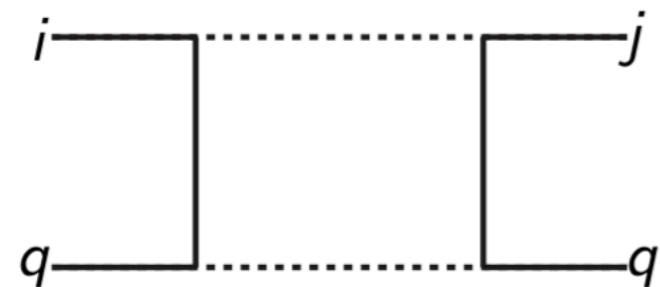
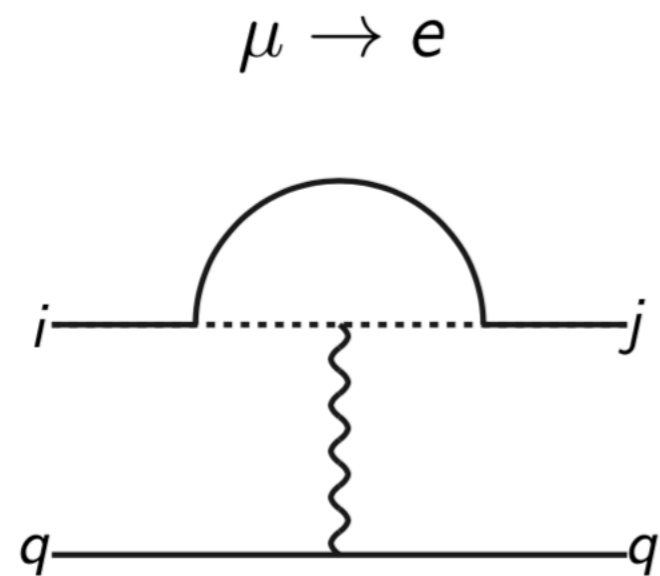
future: $\approx 4 \times 10^{-14}$

✦ prospect for $\mu \rightarrow e$ conversion

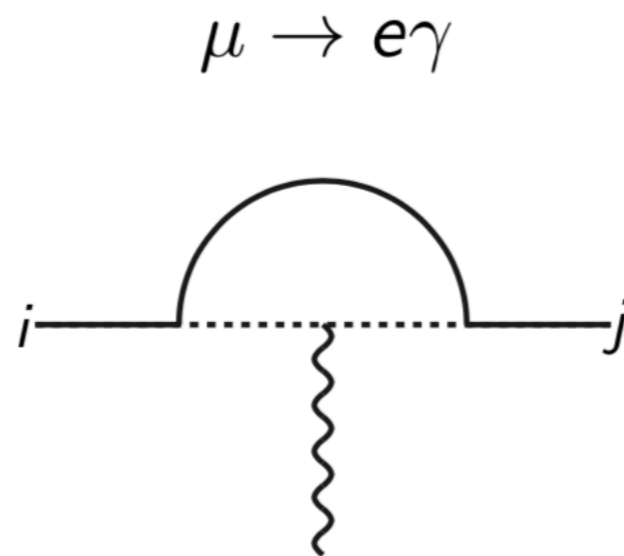
current: 7×10^{-13} (SINDRUM-II)

future: $\approx 10^{-16}$

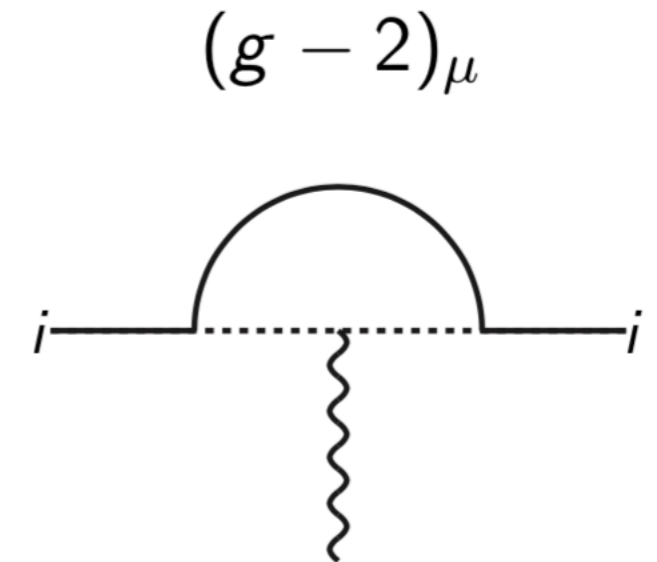
Relation between $(g - 2)_\mu$ and LFV observables



D, A_1^{21}, A_2^{21}



A_2^{21}



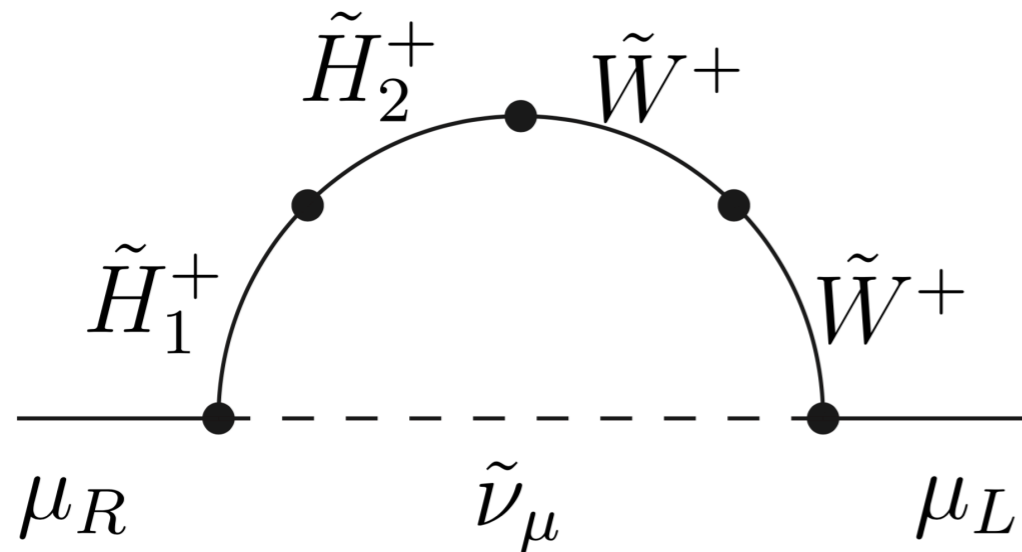
A_2^{22}

each observable requires a dedicated experiment

$(g - 2)_\mu$ in the MSSM

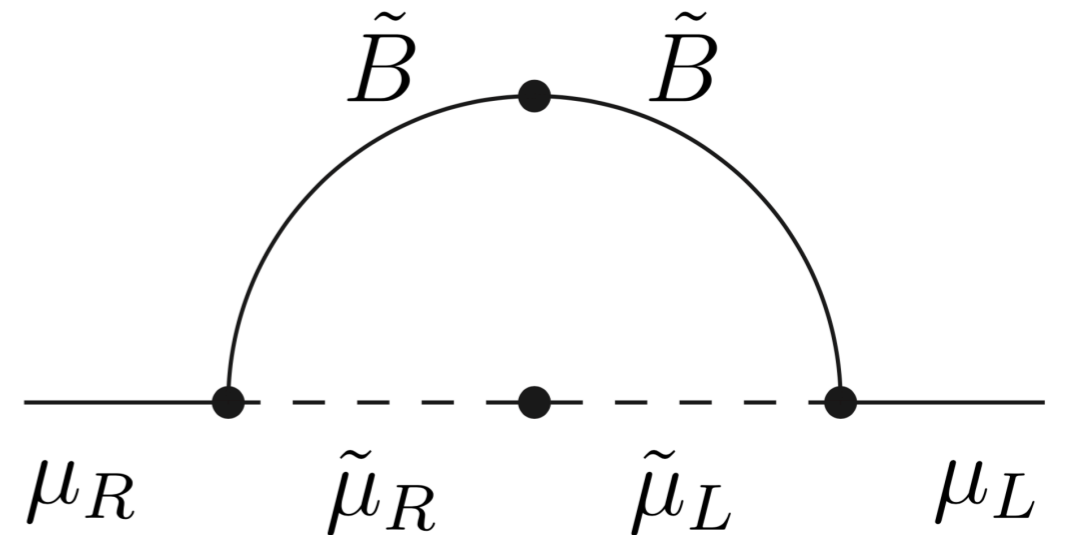
chargino

$$\propto m_\mu^2 \tan \beta \mu M_2$$



neutralino

$$\propto m_\mu^2 \tan \beta \mu M_1$$

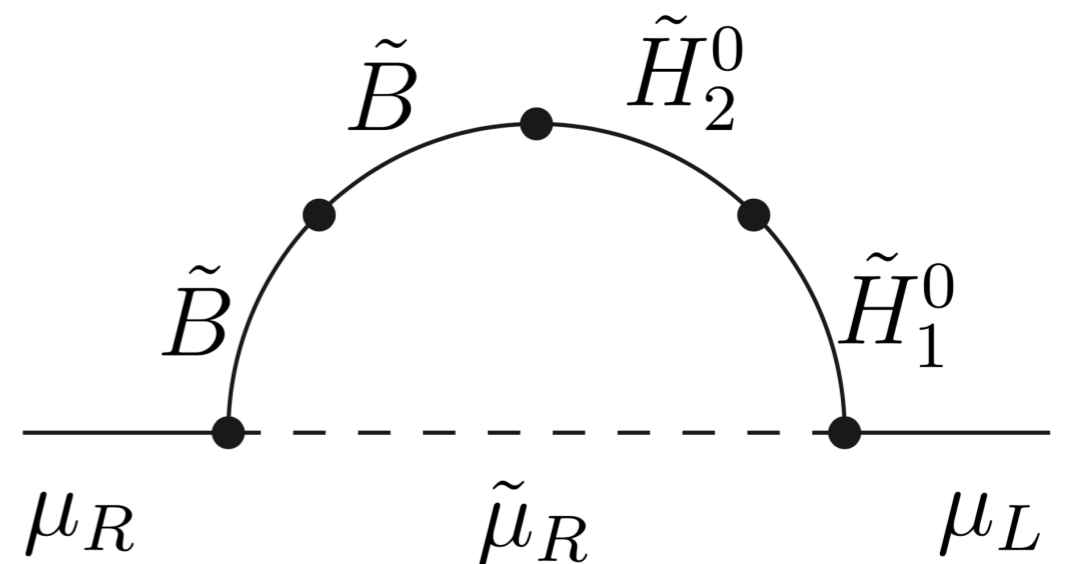


and similarly for $\mu \rightarrow e \gamma$ and $\mu \rightarrow e^-$ - as long as $\tan \beta$ is not very small all considered observables are dominated by the dipole contributions and therefore strongly correlated

$$\text{CR}(\mu \rightarrow e) \propto \alpha \cdot \text{BR}(\mu \rightarrow e \gamma)$$

$$\text{CR}(\mu \rightarrow e) \leq 3 \cdot 10^{-15}$$

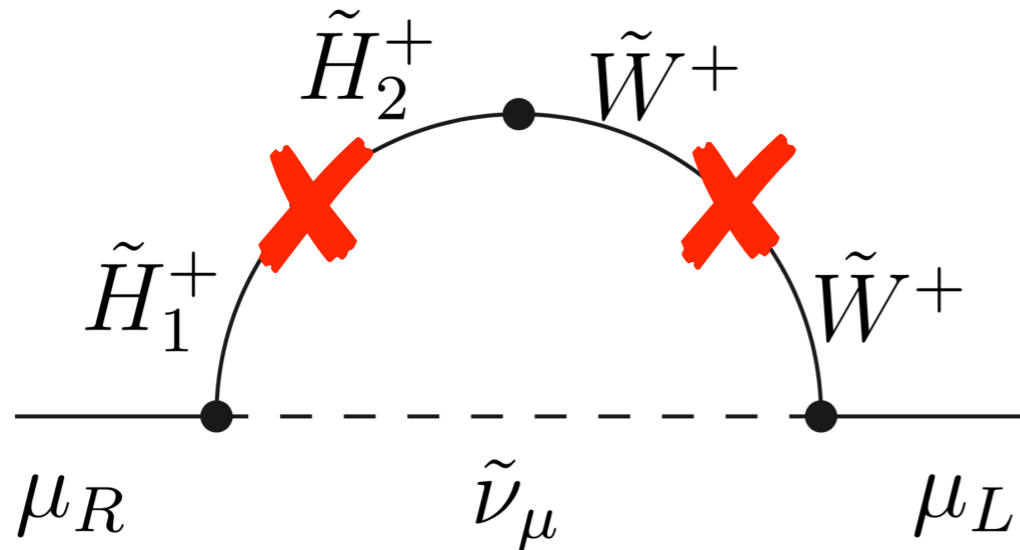
$$\propto m_\mu^2 \tan \beta \mu M_1$$



$(g - 2)_\mu$ in the MRSSM

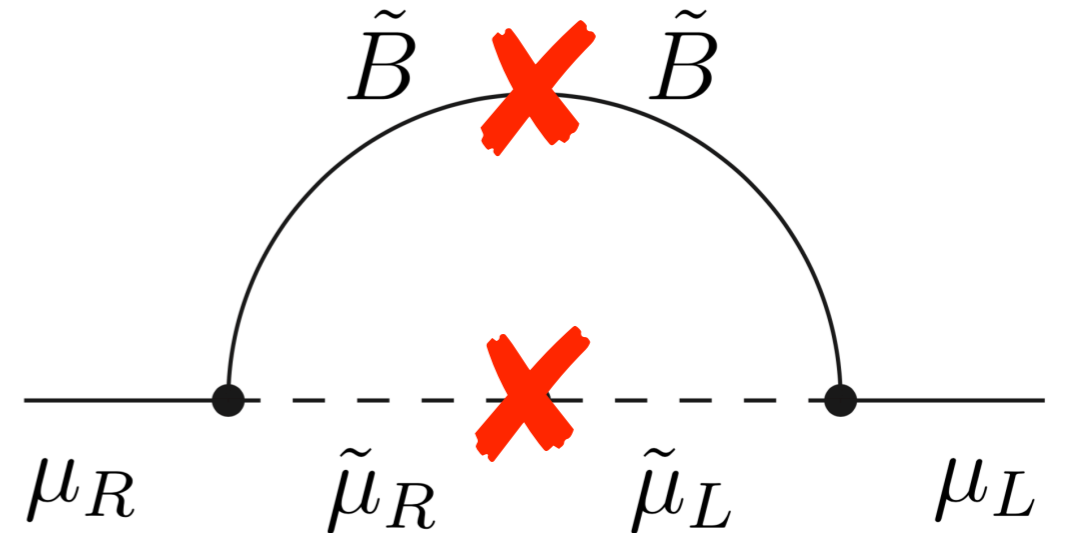
chargino

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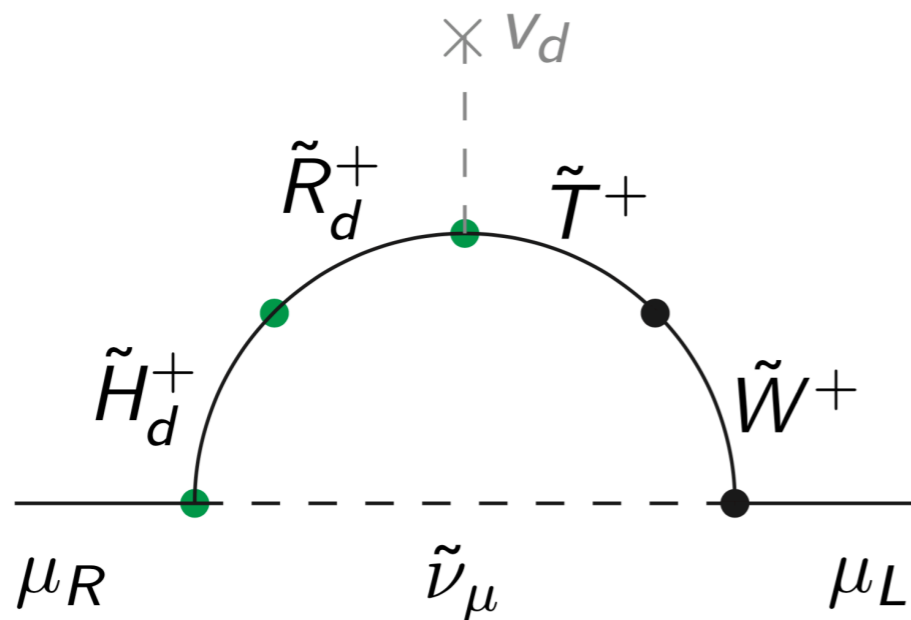


neutralino

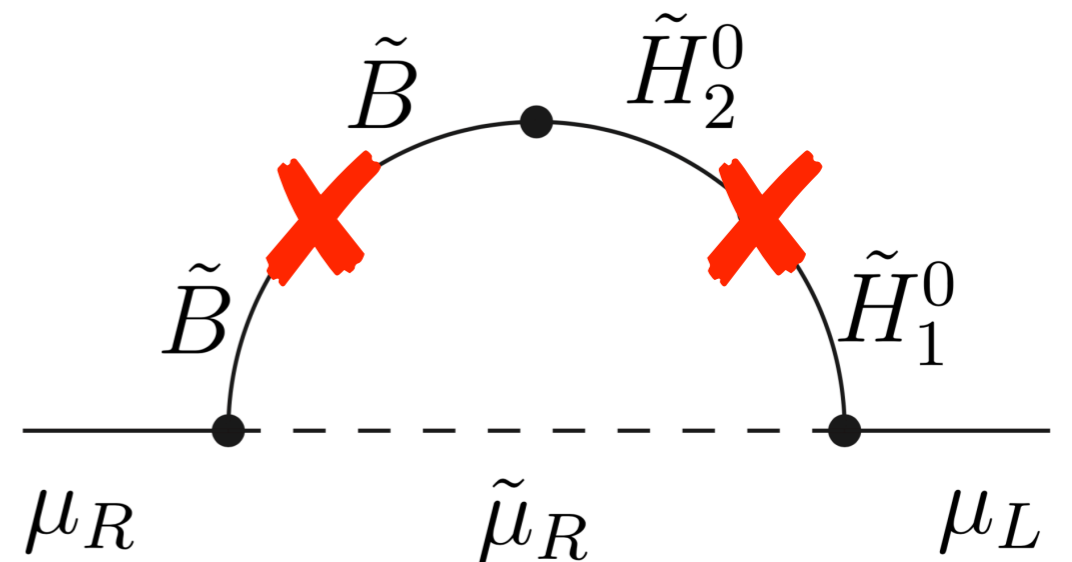
$$\propto m_\mu^2 \tan \beta \mu M_1$$



there is one class of enhanced diagram though

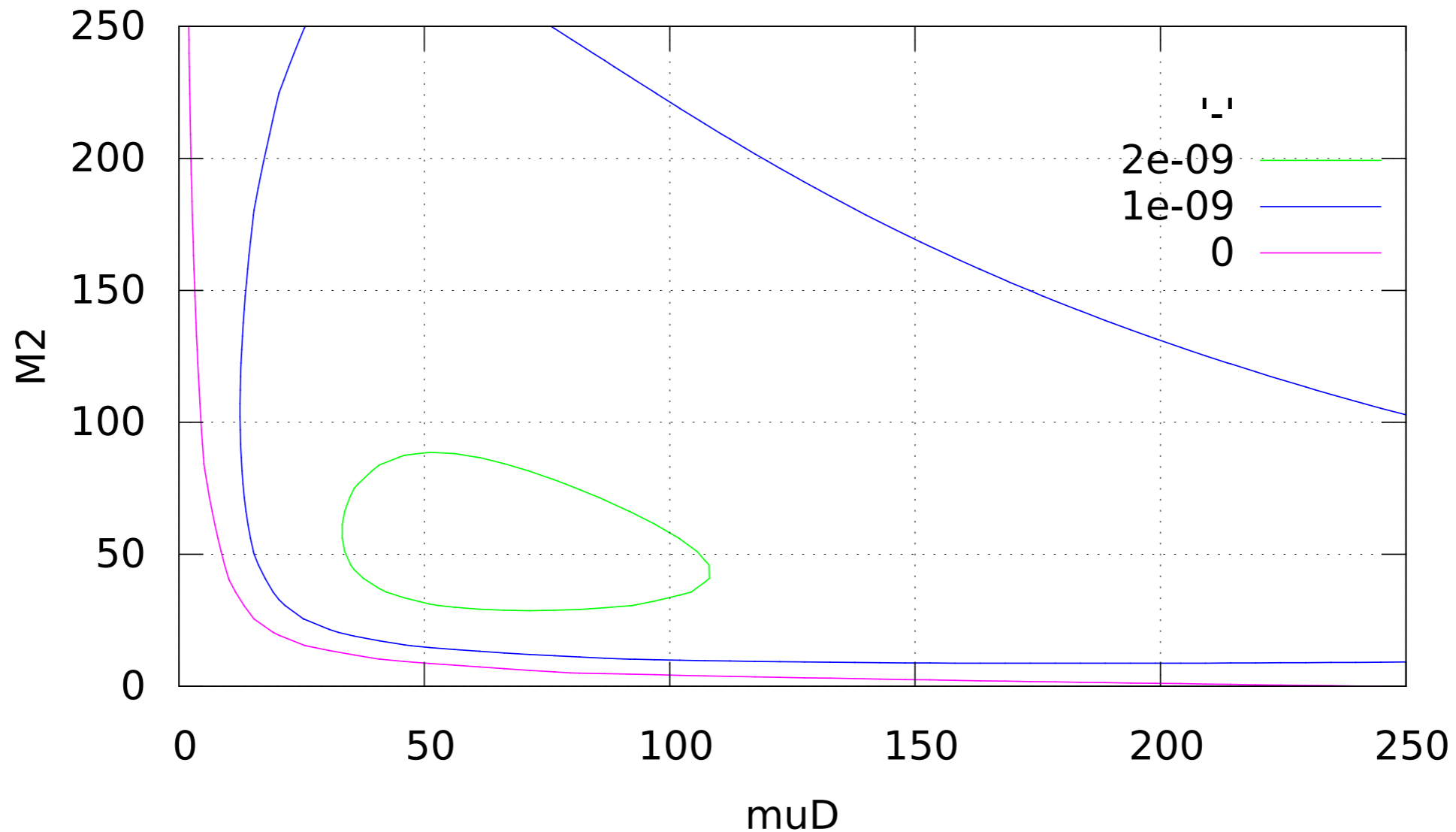


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$(g - 2)_\mu$ in the MRSSM

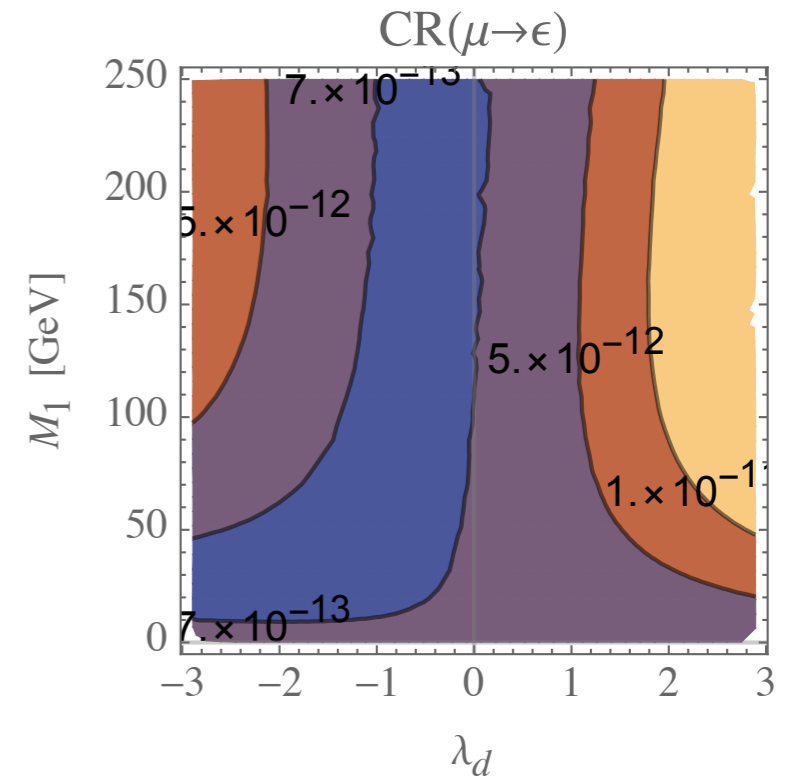
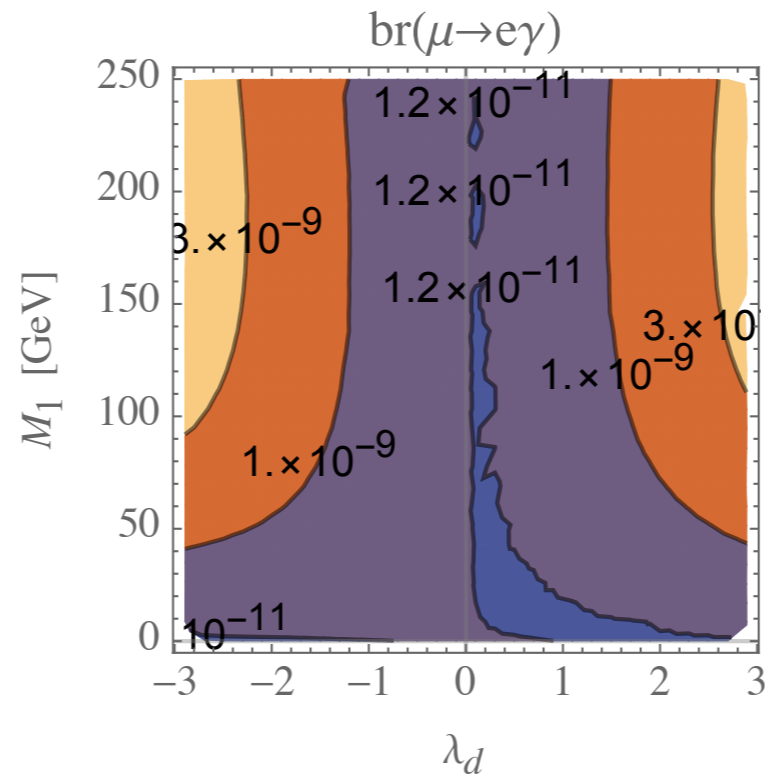
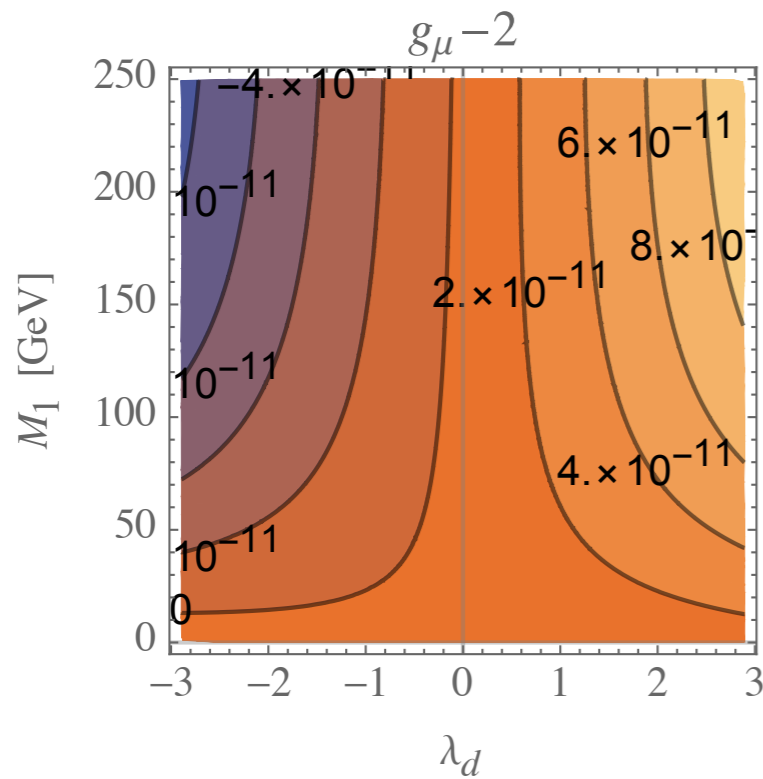
- It is possible to obtain large contribution to $g-2$



- The price to pay are light EW-inos, in tension with experiment

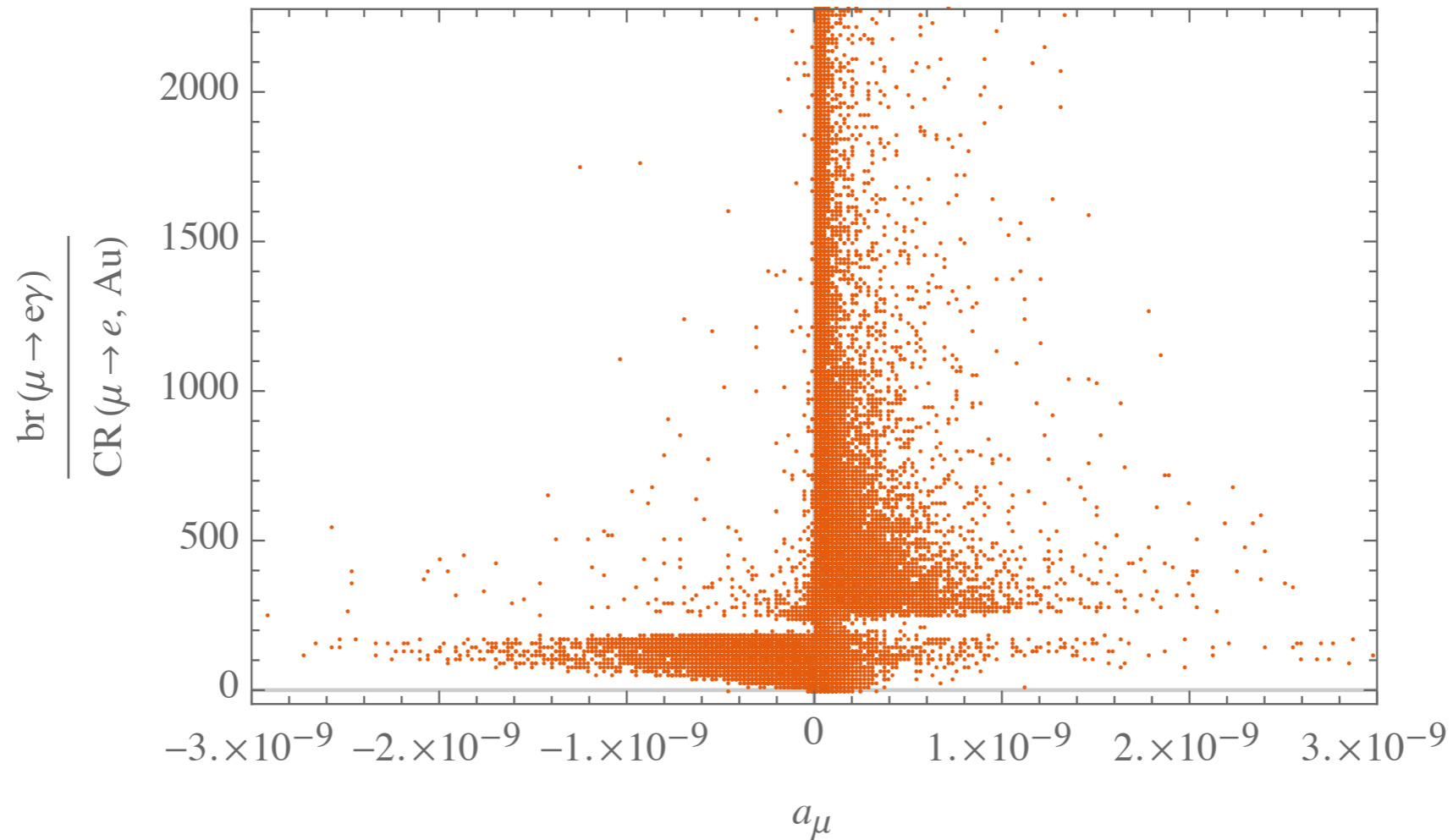
Photonic penguin dominance

- For $|\lambda_d| \gtrsim 1$ the dipoles dominate: $g_{\mu-2}$ scales linearly with λ_d , while $\mu \rightarrow e\gamma$ and $\mu \rightarrow e$ quadratically



- For $|\lambda_d| \gtrsim 1$ the ratio of $\mu \rightarrow e\gamma$ over $\mu \rightarrow e$ is of the order 100, as in the MSSM where $\text{CR}(\mu \rightarrow e) \propto \alpha \cdot \text{BR}(\mu \rightarrow e\gamma)$
- Near $|\lambda_d| \approx 0$ the ratio is of order 10

$$a_\mu \text{ VS. } \frac{\text{br}(\mu \rightarrow e\gamma)}{\text{CR}(\mu \rightarrow e, Au)}$$



- ✦ In the region dominated by the dipoles the $\text{br}(\mu \rightarrow e\gamma) \sim \sin^2 2\theta \cdot a_\mu$
- ✦ In the MRSSM this is a region of $|\lambda_d| \gtrsim 1$, in the MSSM $\tan \beta \gtrsim 5$

Conclusions:

- ✘ Two distinct cases: $|\lambda_d| \approx 0$, $|\lambda_d| > 0$
- ✘ For large $|\lambda_d|$ observables get dominated by photon „penguins” and are strongly correlated
- ✘ Generating sufficient contribution to $g-2$ through large λ_d overshoots LFV observables (unless one fine-tunes the mixing angle)
- ✘ Similar things happen for Λ_d
- ✘ For $|\lambda_d| \approx 0$ the $g-2$ and $\mu \rightarrow e\gamma$ are still correlated but the $\mu \rightarrow e$ conversion rate can be dominated by so-called charge radius, Z-penguin and box contributions
- ✘ It is therefore possible to find a parameter points not excluded by current experimental results, within reach of the next $\mu \rightarrow e$ conversion (but not $\mu \rightarrow e\gamma$) experiment

Backup

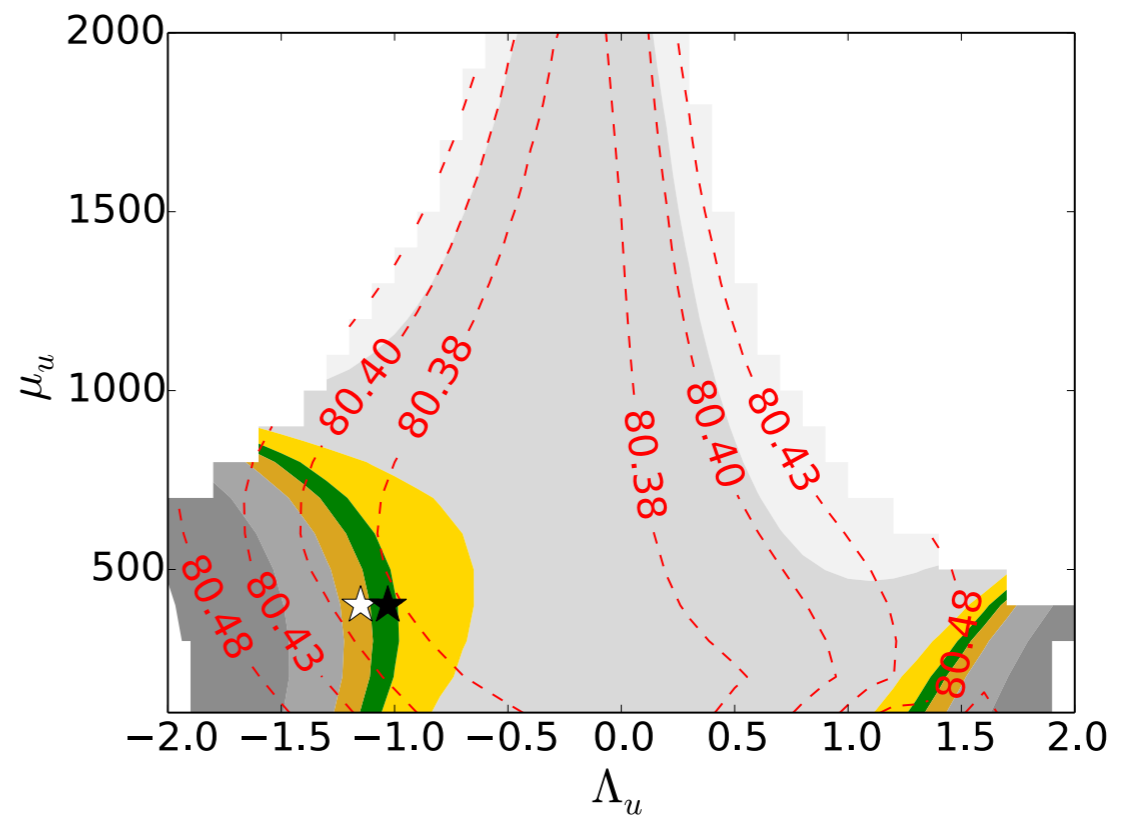
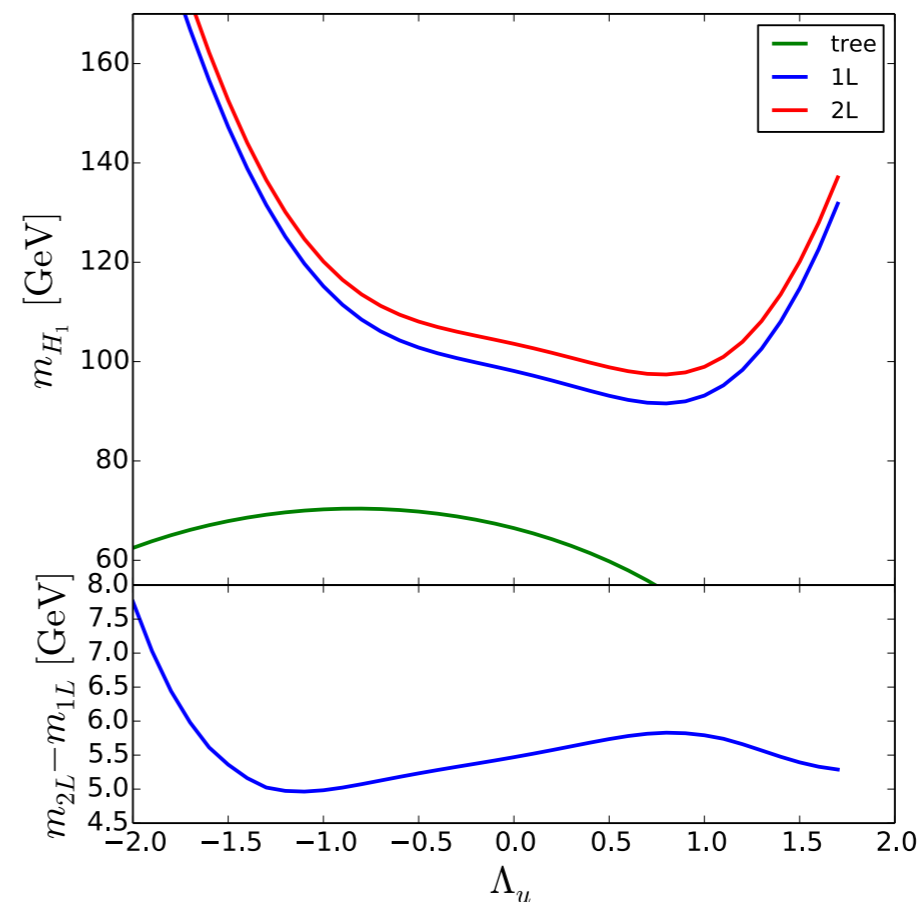
EW sector of the MRSSM (status)

- The SM-like Higgs boson mass in the MRSSM has been calculated including full 1-loop and leading 2-loop corrections^{1,2}
- Impact of EWPO was analyzed¹
- MRSSM can predicts correct dark matter relic density while being in agreement with dark matter direct detection bounds³
- Its EW signatures were checked against available 7 and 8 TeV data³

1. P. Dießner, J. Kalinowski, W. Kotlarski and D. Stöckinger, JHEP **1412** (2014) 124

2. P. Dießner, J. Kalinowski, W. Kotlarski and D. Stöckinger, Adv. High Energy Phys. **2015** (2015) 760729

3. P. Dießner, J. Kalinowski, W. Kotlarski and D. Stöckinger, JHEP **1603** (2016) 007

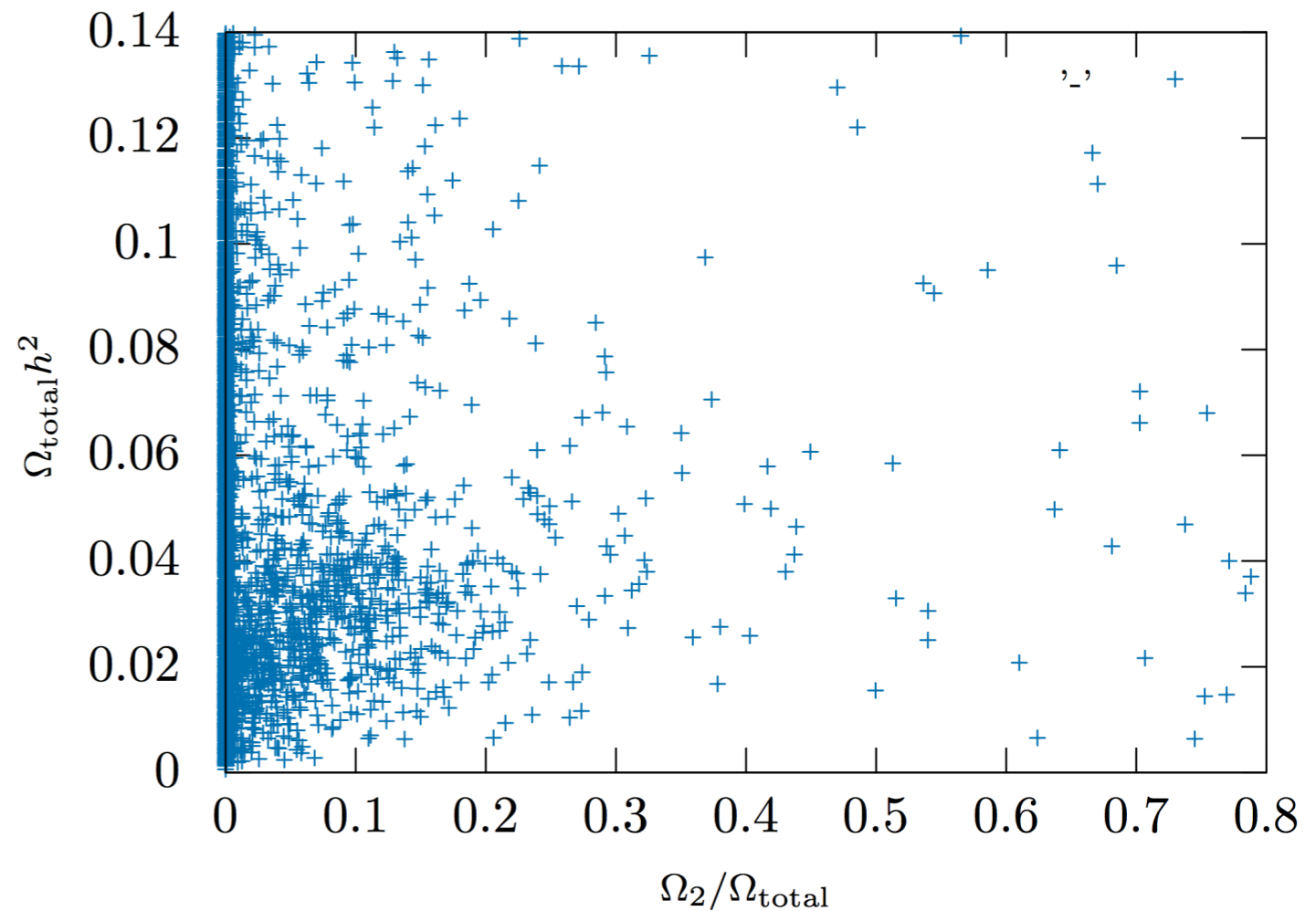


2 component dark matter

- consider scenarios where the lightest particle with $R=1$ is neutralino or sneutrino with mass m_{LSP1}
- if $m_{R_1^0} < 2m_{\text{LSP1}}$, lightest neutral R -Higgs is also stable
- two SUSY dark matter candidates with relic densities Ω_1 and Ω_2
- requirements
 - $\Omega_{\text{total}}h^2 \equiv (\Omega_1 + \Omega_2)h^2 \approx 0.11$
 - substantial fraction $\Omega_2/\Omega_{\text{total}}$
- (for now) best points are not collinear friendly:

$$m_{\tilde{\chi}_1^0} = 367 \text{ GeV}$$

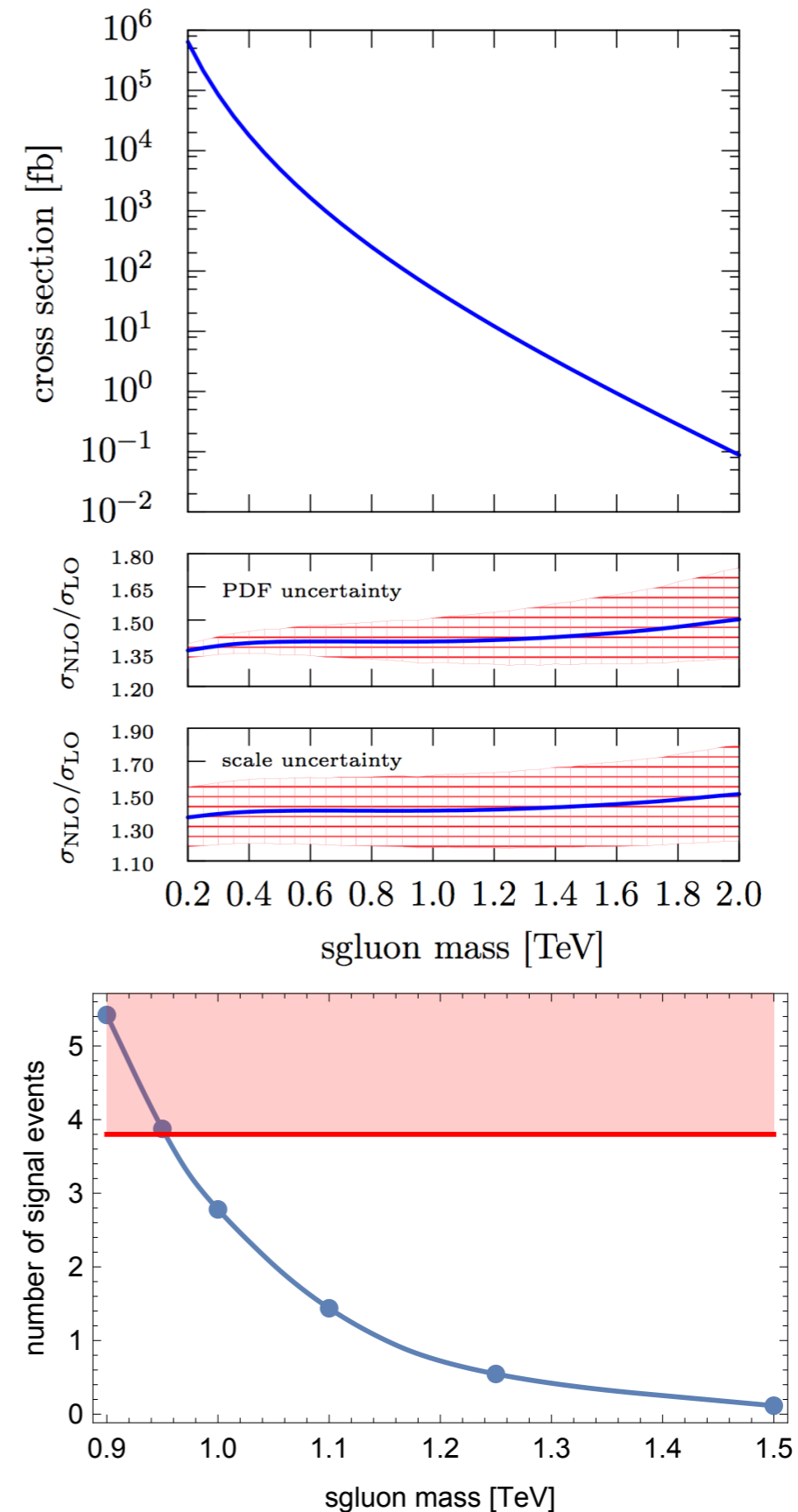
$$m_{R_1^0} = 571 \text{ GeV}$$



Sgluon pair production at 13 TeV LHC

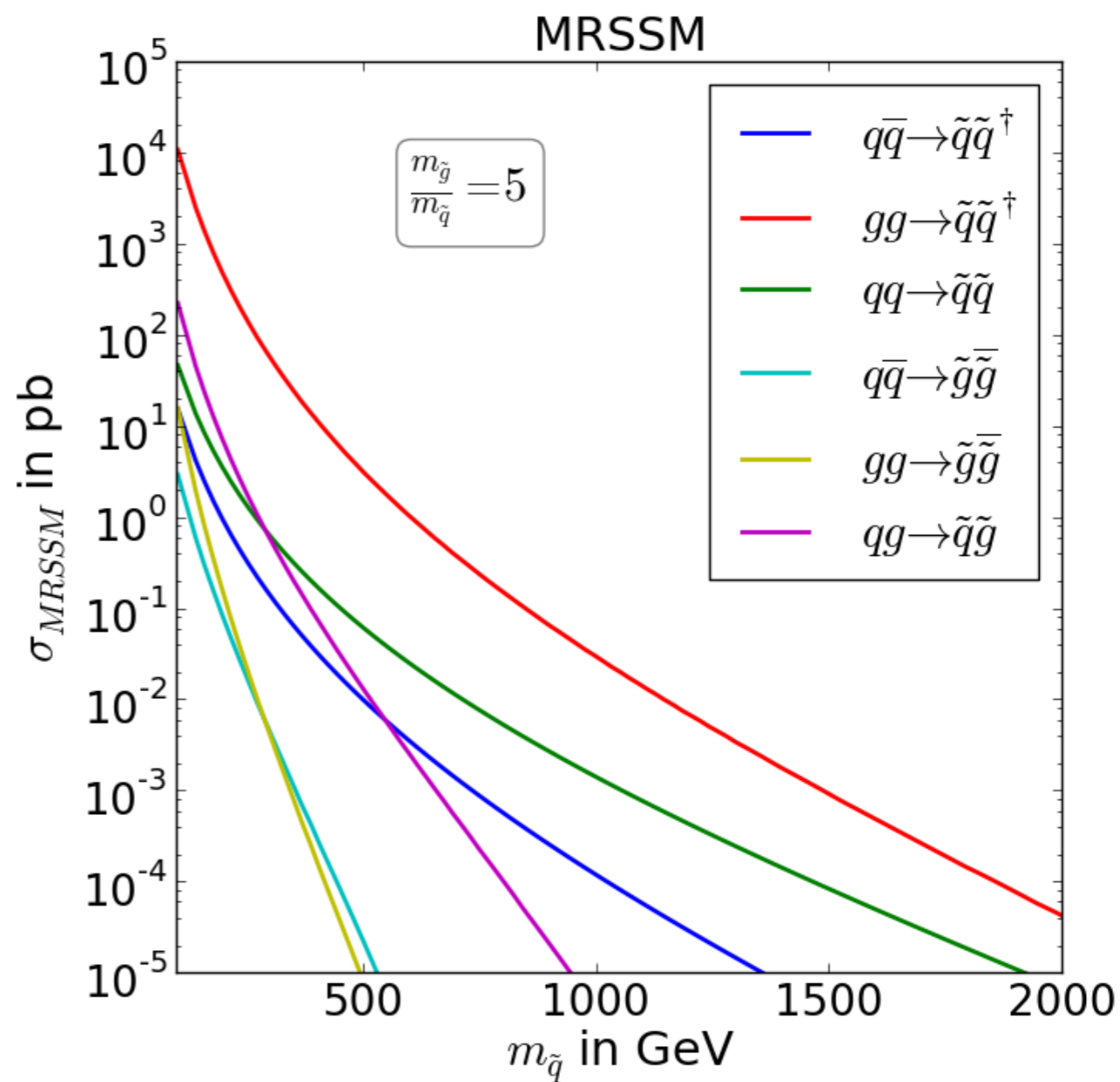
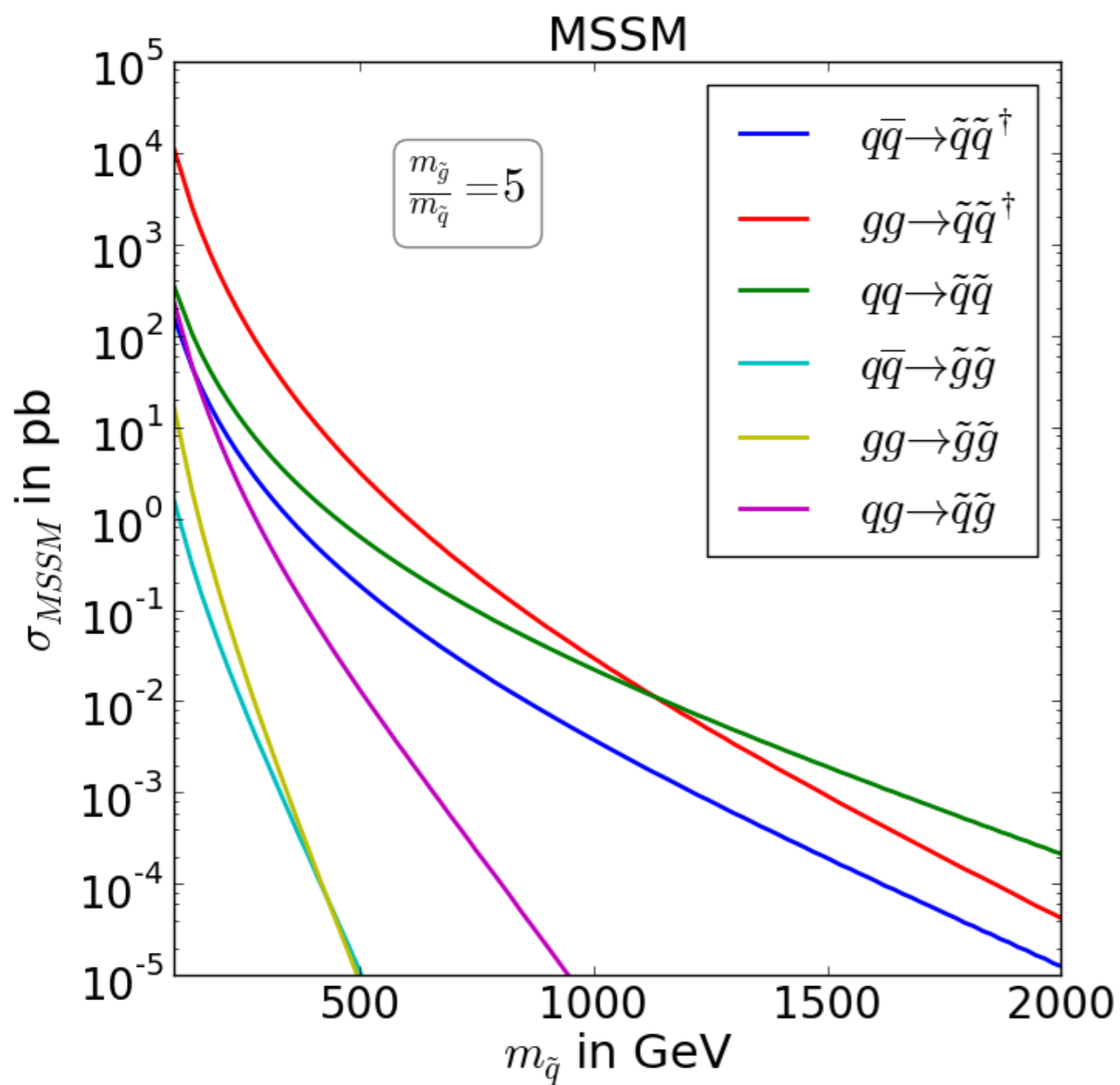
W. Kotlarski [arXiv:1608.00915]

- Analysis of the sgluon pair production with subsequent decay into $t\bar{t}$ pairs. Recasting ATLAS search in the same-sign lepton channel using 3.2/fb of integrated luminosity
- Signal simulated at NLO using MadGraph5_aMC@NLO + FeynRules + NLOCT and matched to parton shower in the MC@NLO scheme
- Detector response parametrized using Delphes3
- Analysis validated on background processes $t\bar{t}l^+l^-$, $t\bar{t}l^\pm\nu$
- Mass of pair produced real sgluons decaying with $\text{BR}(O \rightarrow t\bar{t}) = 1$ excluded up to 950 GeV

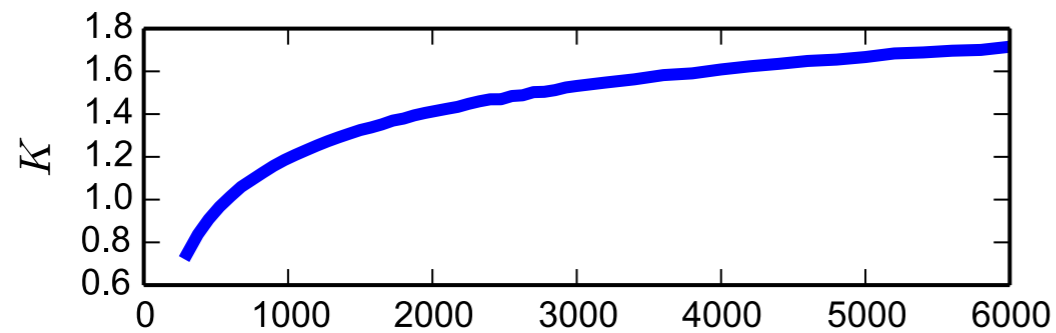
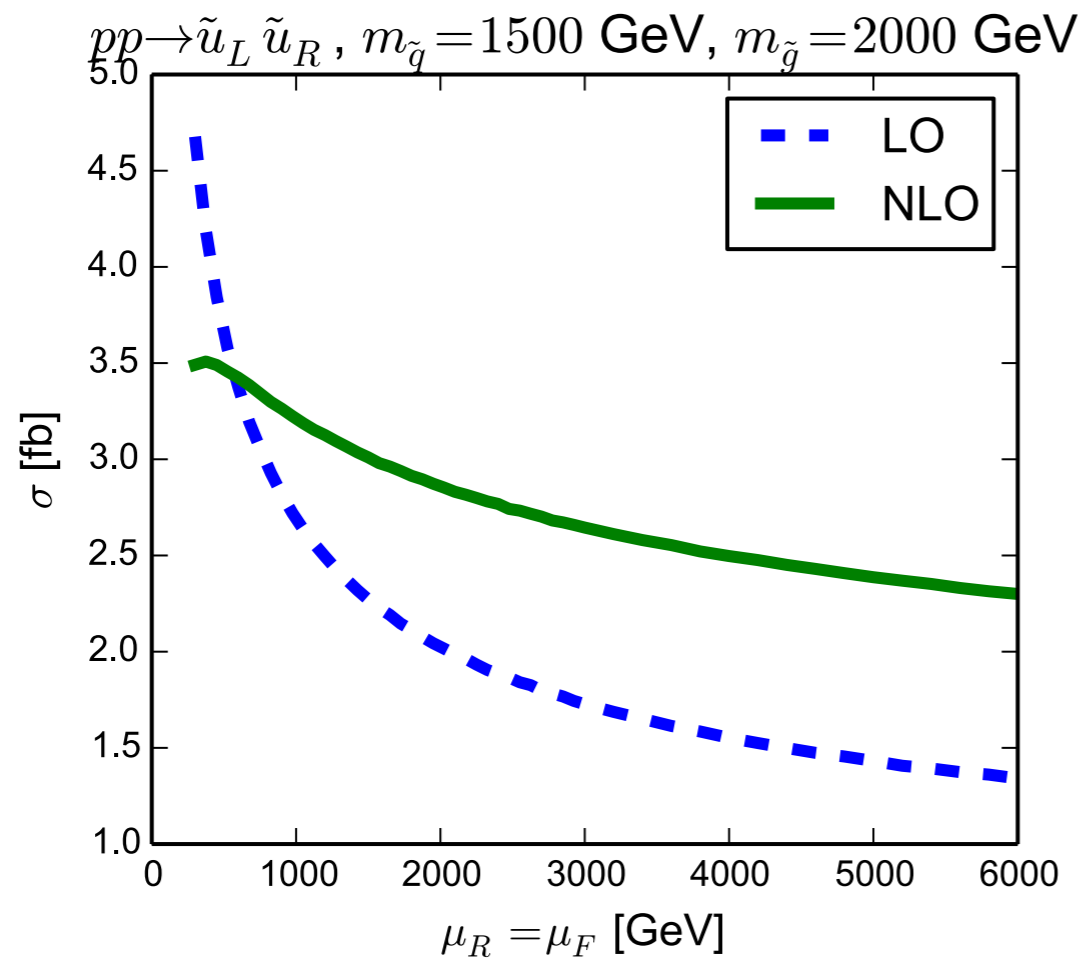


Leading order analysis

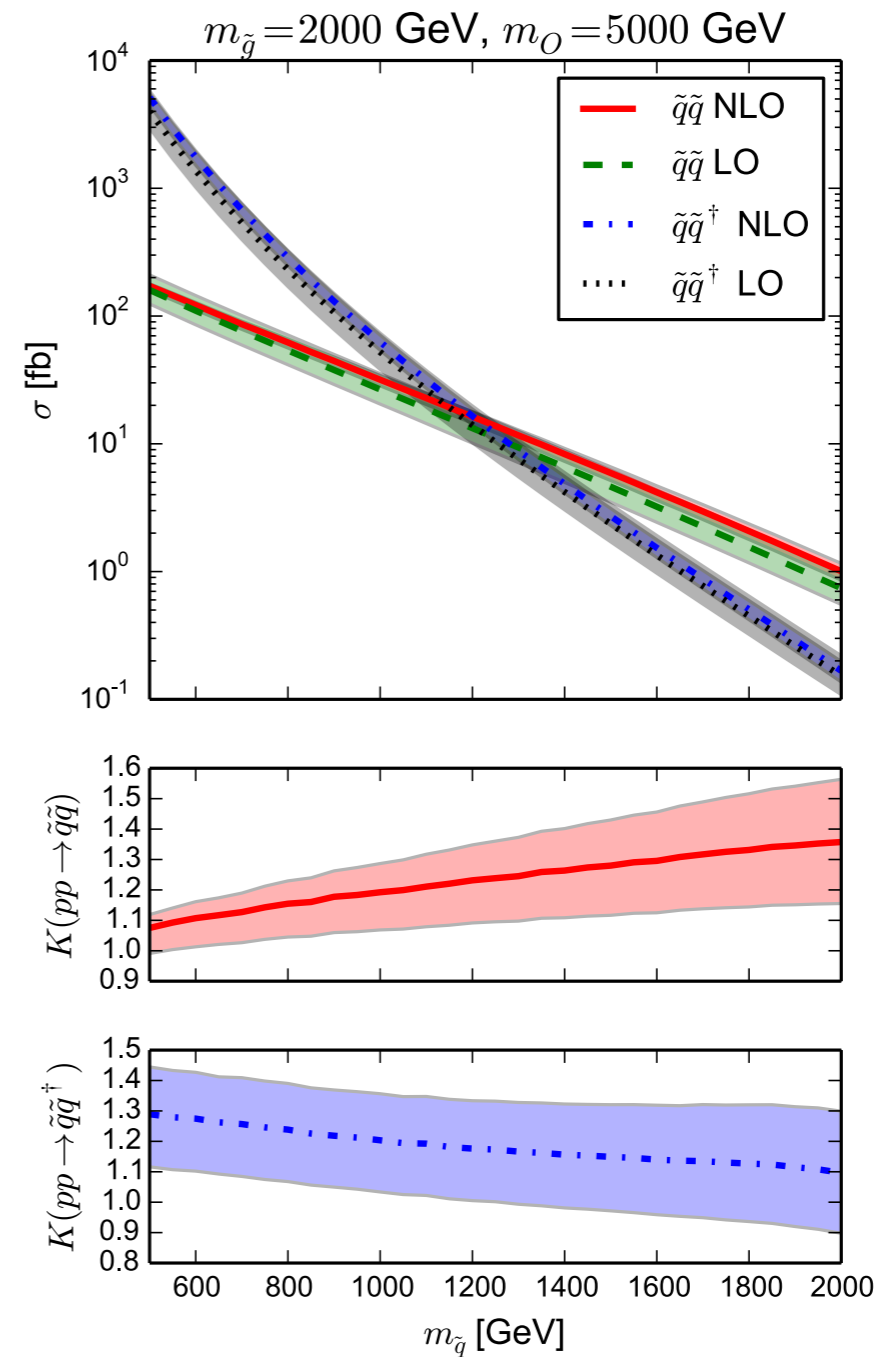
LO cross-sections for sparticle production at the LHC at $\sqrt{s} = 13\text{TeV}$



NLO improvements



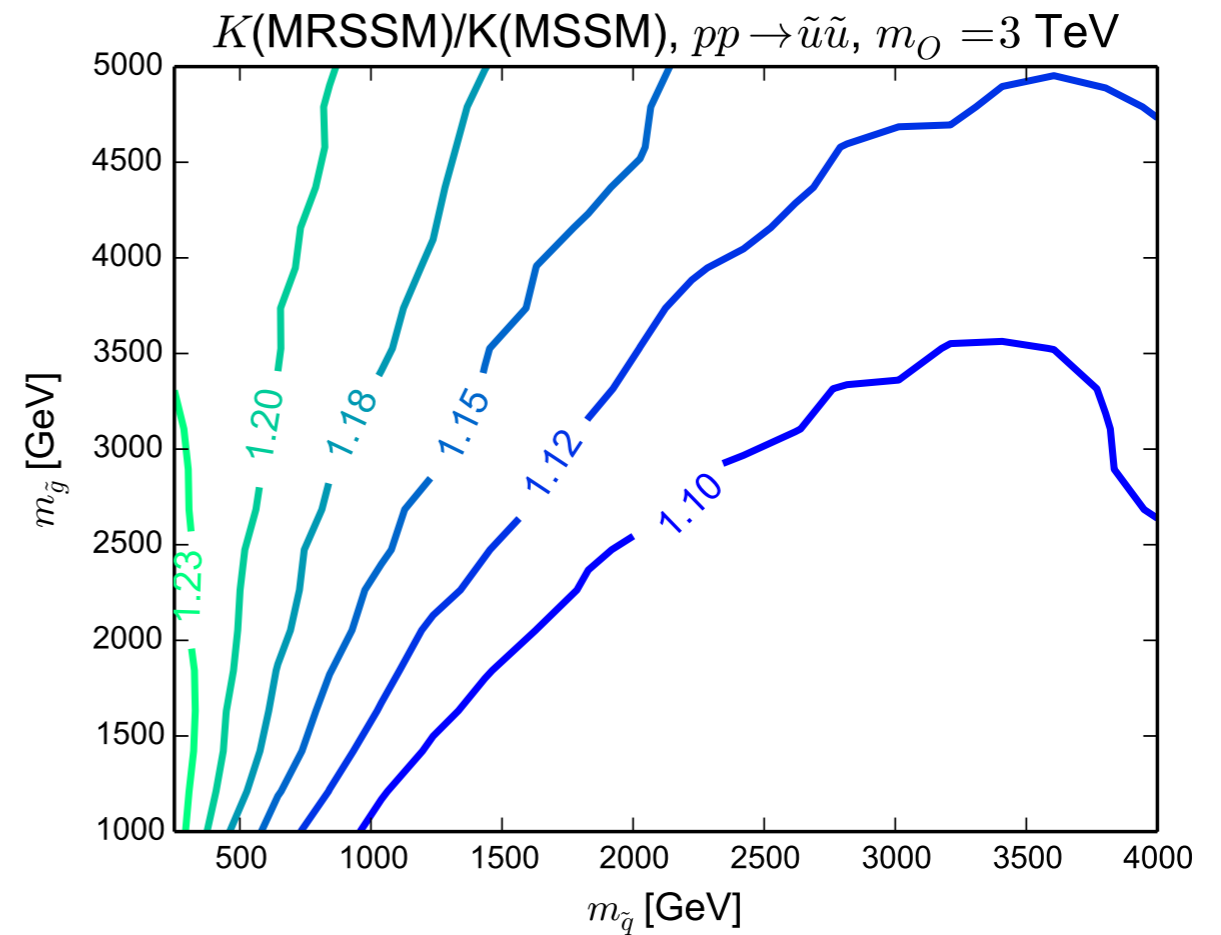
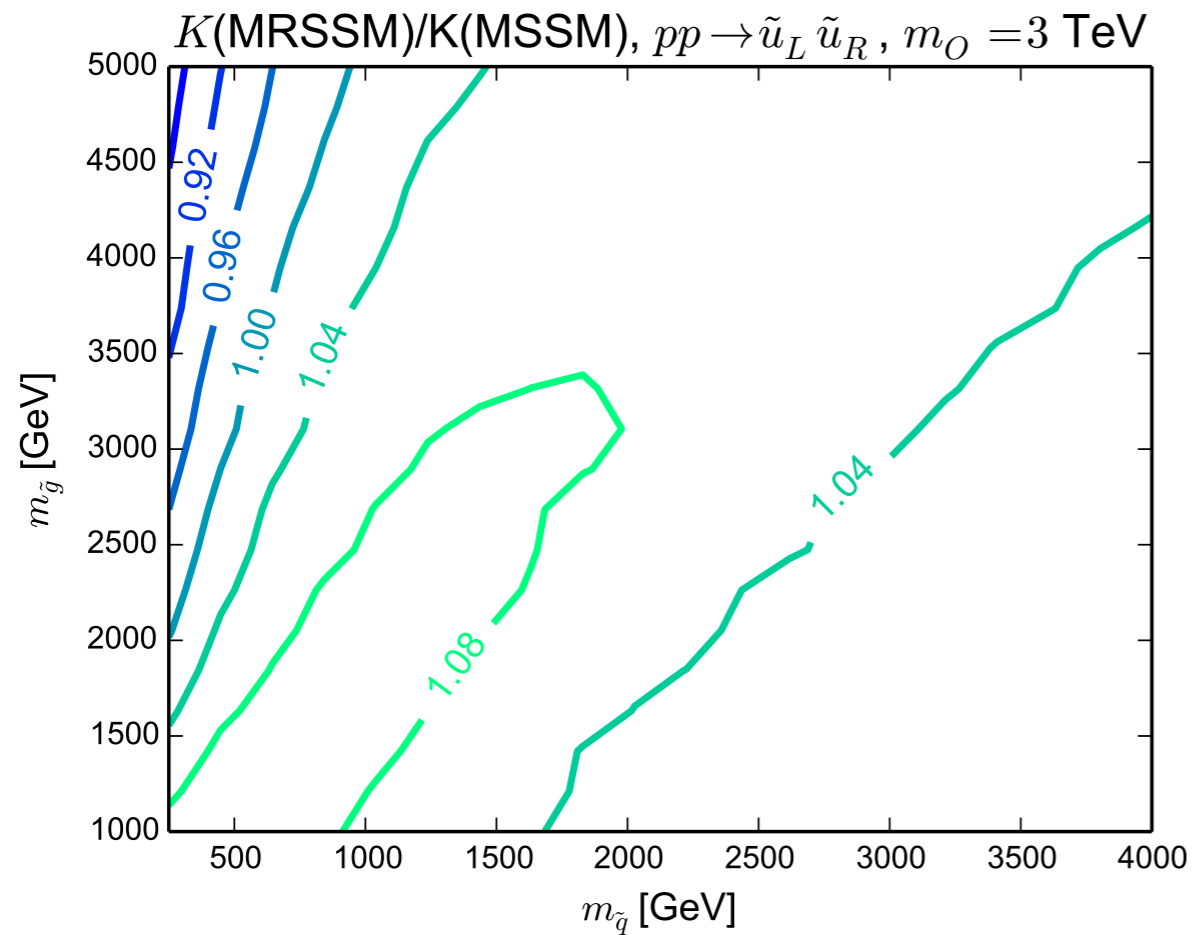
reduction of theoretical uncertainty



right figure summed over flavors

shift of cross-sections

Comparison with the MSSM

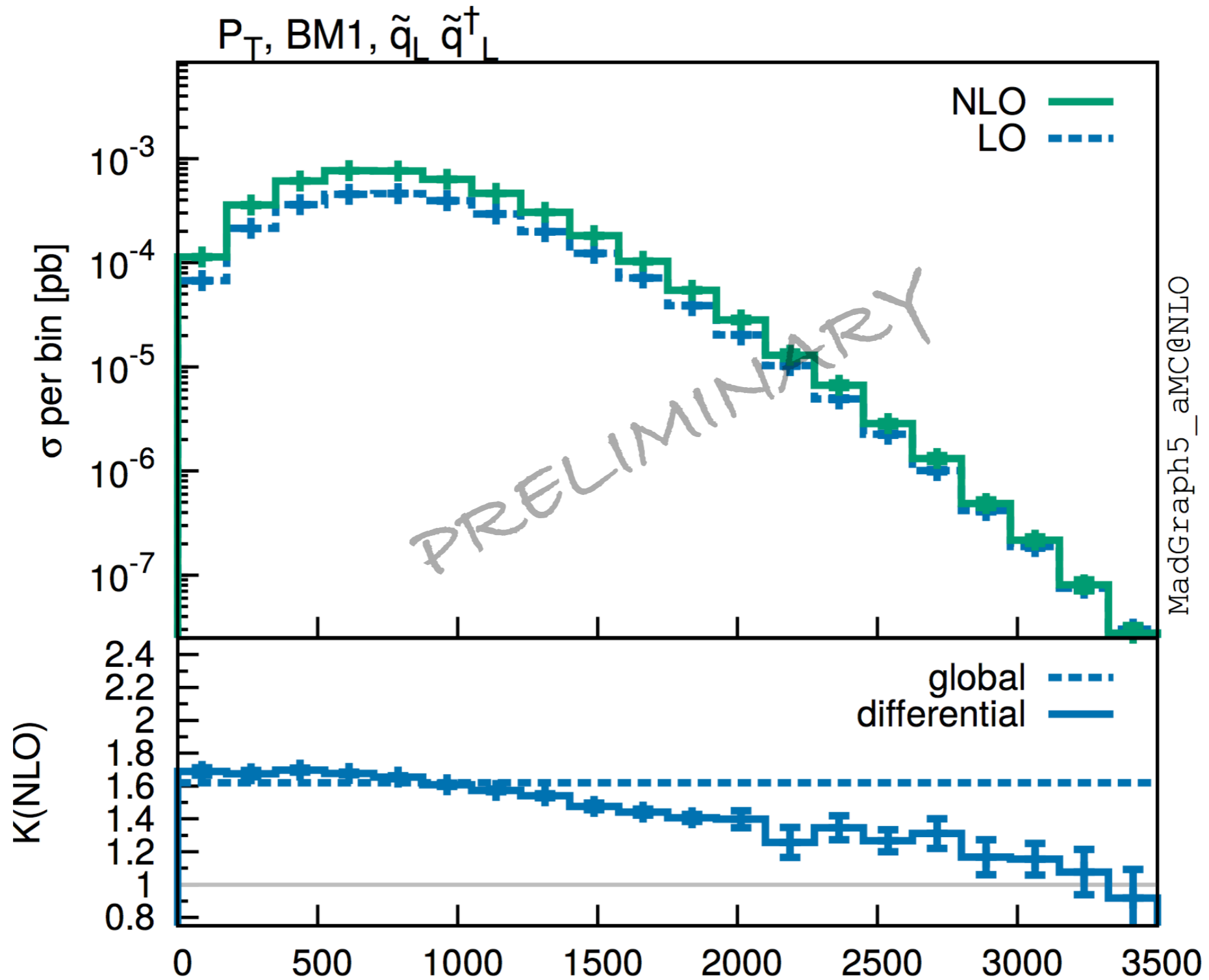


✖ Two possible definitions of K-factors:

* unsummed over L- and R-squarks

* summed

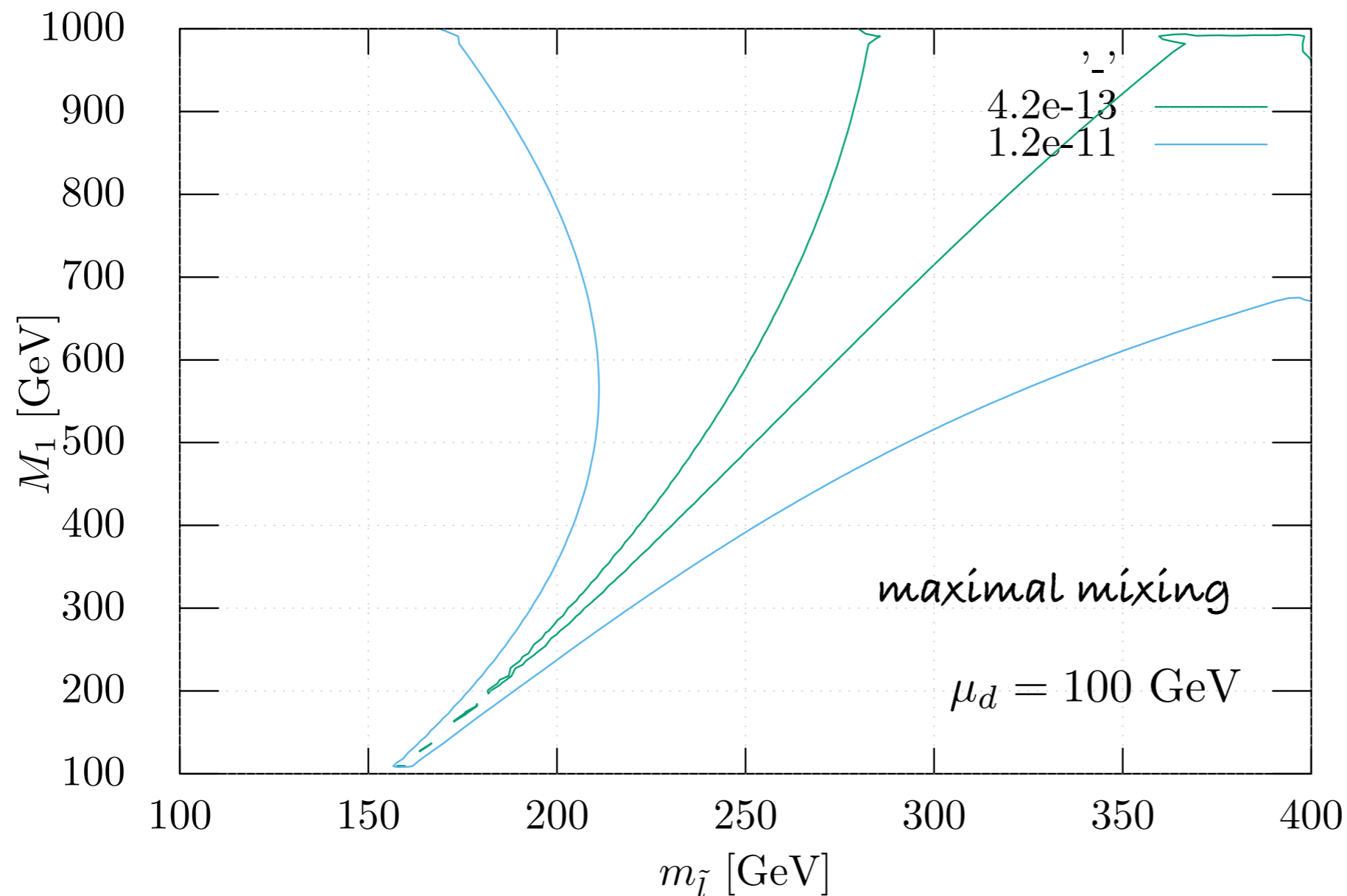
Differential distributions



$\mu \rightarrow e\gamma$ in the MRSSM

- first analysis performed by Fok and Kribs [*Phys. Rev. D* 82, 035010 (2010)]
- simplifying assumptions: $M_2, \mu_u \rightarrow \infty$, only 2 neutralinos containing \tilde{B}, \tilde{H}_d contribute

$$m_{\tilde{l}_2} = \frac{3}{2}m_{\tilde{l}}$$



new MEG results

old MEG results