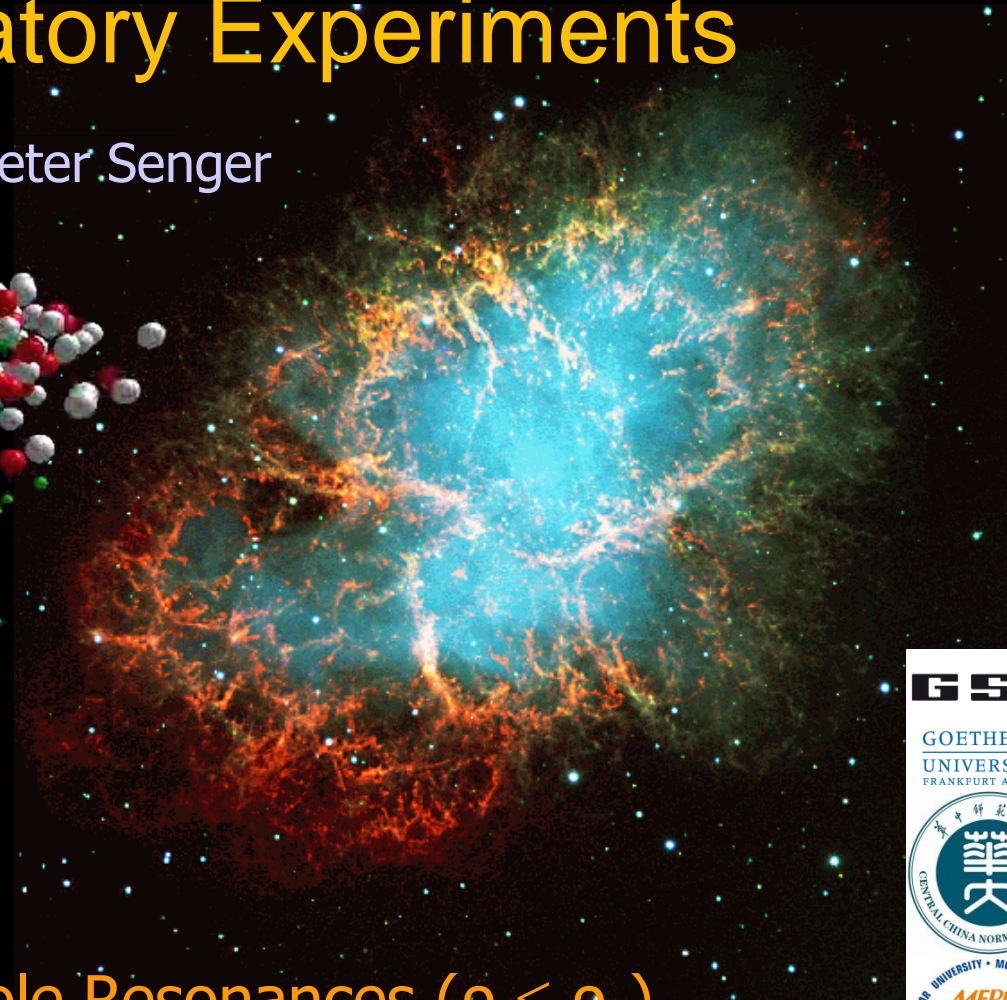
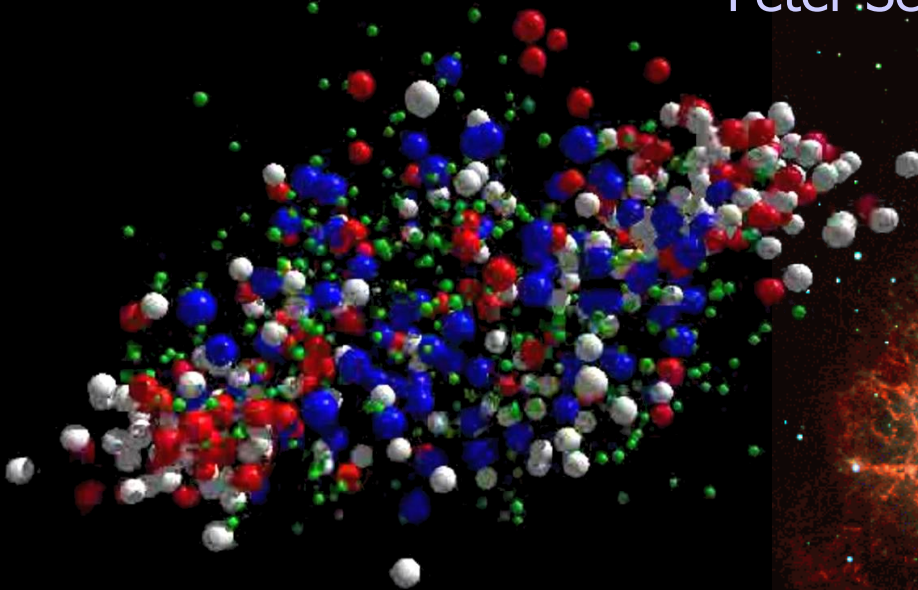


The high-density Equation of State from Laboratory Experiments

Peter Senger



Outline:

- EOS from Giant Monopole Resonances ($\rho \leq \rho_0$)
- EOS from heavy-ion collisions at $\rho < 3 \rho_0$
- Observables sensitive to EOS at $\rho > 3 \rho_0$?

3rd CBM China workshop, April 16 -18, 2018, Yichang, China



Even stronger push for the scientific motivation of FAIR

from P. Giubellino



Neutron Star Mergers



FAIR Research Pillars

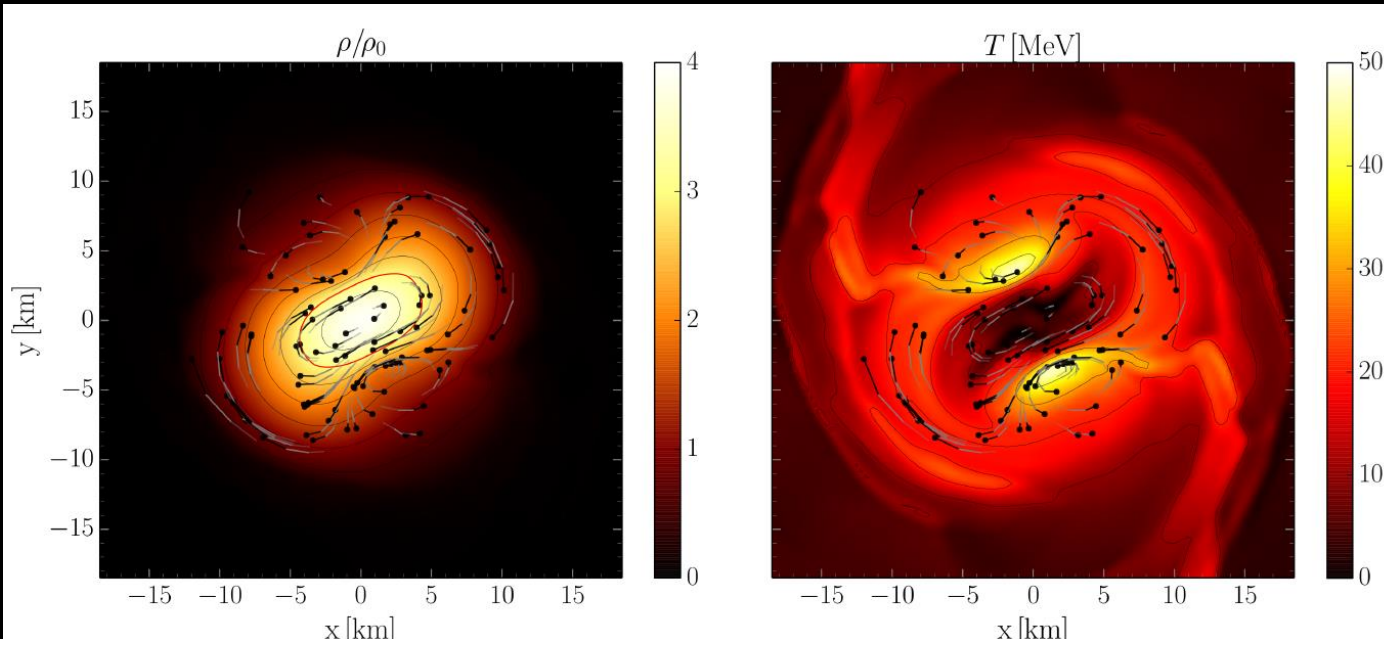
- Equation of State (**Hades, CBM**)
 - Gravitational wave signal
 - Amount of ejecta
- Baryon-Baryon interaction (**PANDA**)
- Exotic neutron-rich nuclei (**NUSTAR**)
 - r-process nucleosynthesis and abundancies of the heaviest elements gold, platinum and beyond
- Plasma and atomic opacities (**APPA**)
 - Kilonova electromagnetic transient

FAIR offers unique opportunities for studying these fundamental questions!

Neutron star mergers and heavy-ion collisions

density

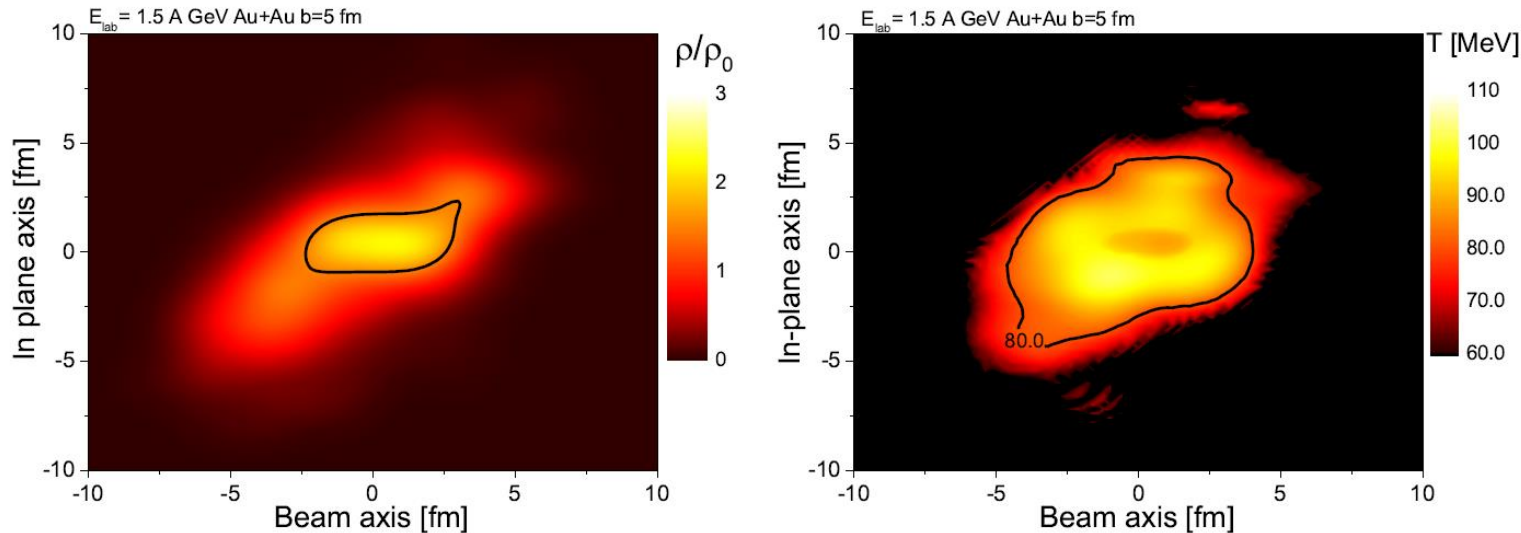
temperature



M. Hanauske et al.,
J. Phys.: Conf. Ser.
878 012031

n-star merger

EOS



Au +Au
1.5A GeV

The equation-of-state of nuclear matter

The nuclear matter equation of state (EOS) describes the relation between density, pressure, temperature, energy, and isospin asymmetry

$$P = \left. \delta E / \delta V \right|_{T=\text{const}}$$

$$V = A / \rho$$

$$\delta V / \delta \rho = -A / \rho^2$$

$$P = \rho^2 \left. \delta(E/A) / \delta \rho \right|_{T=\text{const}}$$

$$E_A(\rho, \delta) = E_A(\rho, 0) + E_{\text{sym}}(\rho) \cdot \delta^2 + O(\delta^4)$$

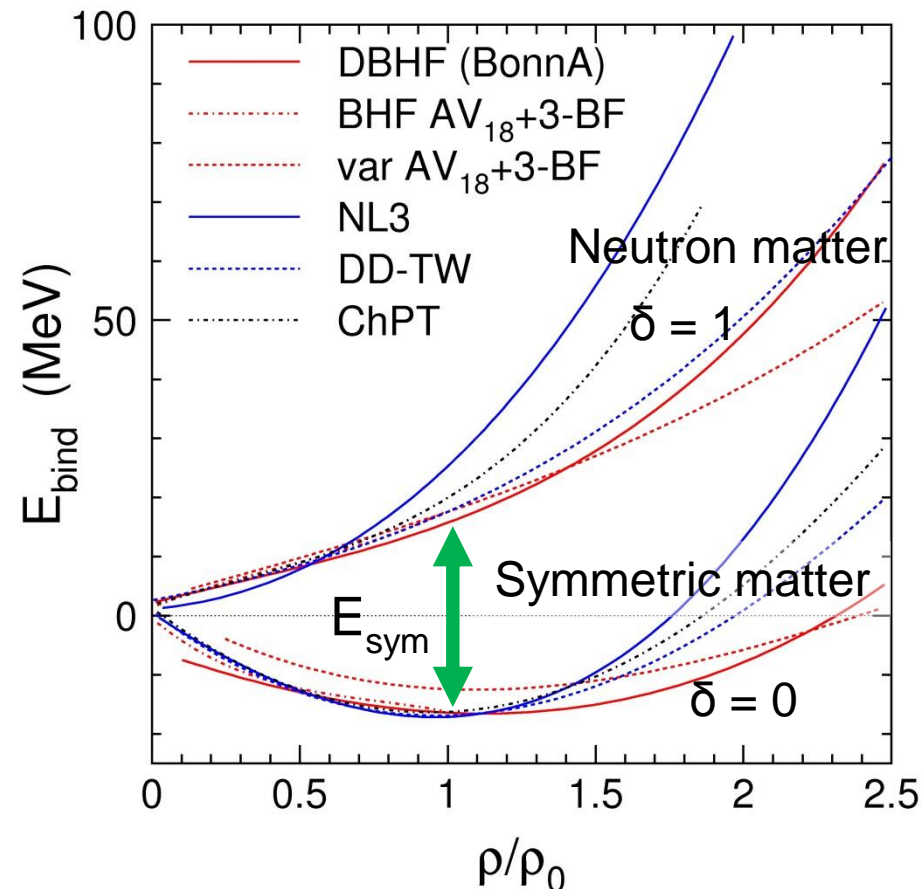
with asymmetry parameter $\delta = (\rho_n - \rho_p) / \rho$

Finite nuclei:

$$E = -a_{\text{vol}} A + a_{\text{surf}} A^{2/3} + a_{\text{sym}} \frac{(N-Z)^2}{A} + a_c \frac{Z^2}{A^{1/3}} + E_{\text{pair}}$$

Bethe-Weizsäcker mass formula

Fuchs and Wolter, EPJA 30 (2006)



The equation-of-state of (symmetric) nuclear matter

$$E_A(\rho, \delta) = E_A(\rho, 0) + E_{\text{sym}}(\rho) \cdot \delta^2 + O(\delta^4)$$

C. Fuchs, Prog. Part. Nucl. Phys. 56 (2006) 1

$T=0$: $E/A = 1/\rho \int U(\rho) d\rho$

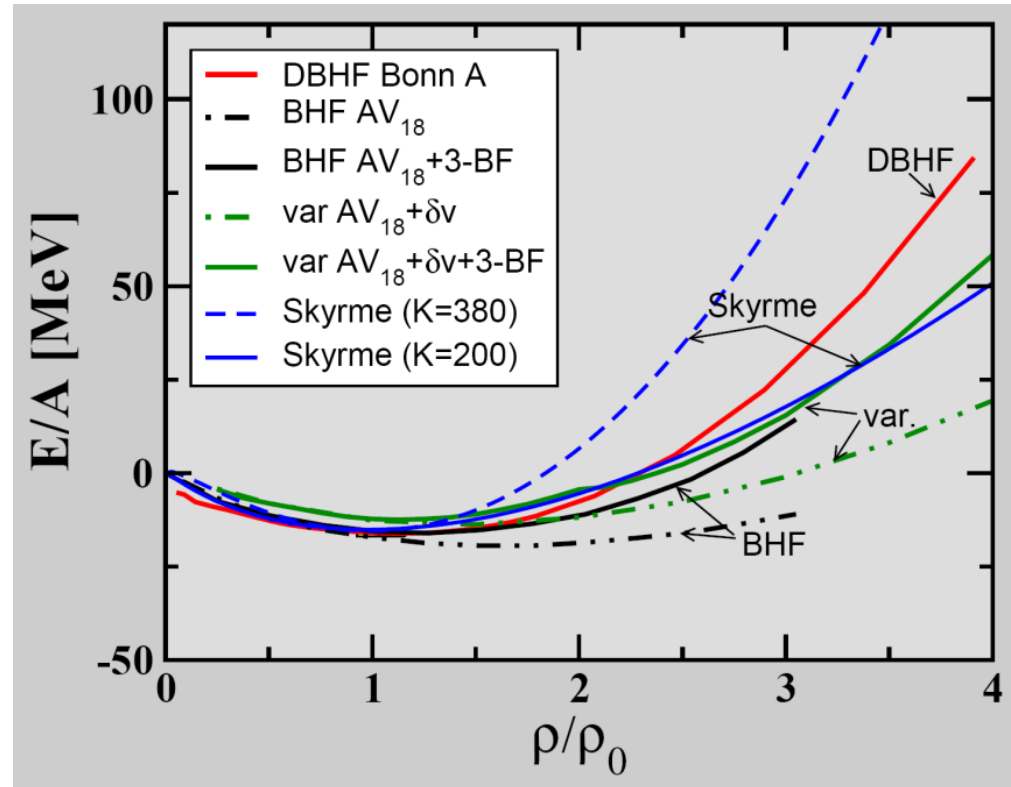
Effective NN-potential:

$$U(\rho) = \alpha\rho + \beta\rho^\gamma$$

➤ $E/A(\rho_0) = -16 \text{ MeV}$

➤ $\delta(E/A)(\rho_0)/\delta\rho = 0$

➤ $K_{\text{nm}} = 9\rho^2 \delta^2(E/A)/\delta\rho^2$
(nuclear incompressibility)



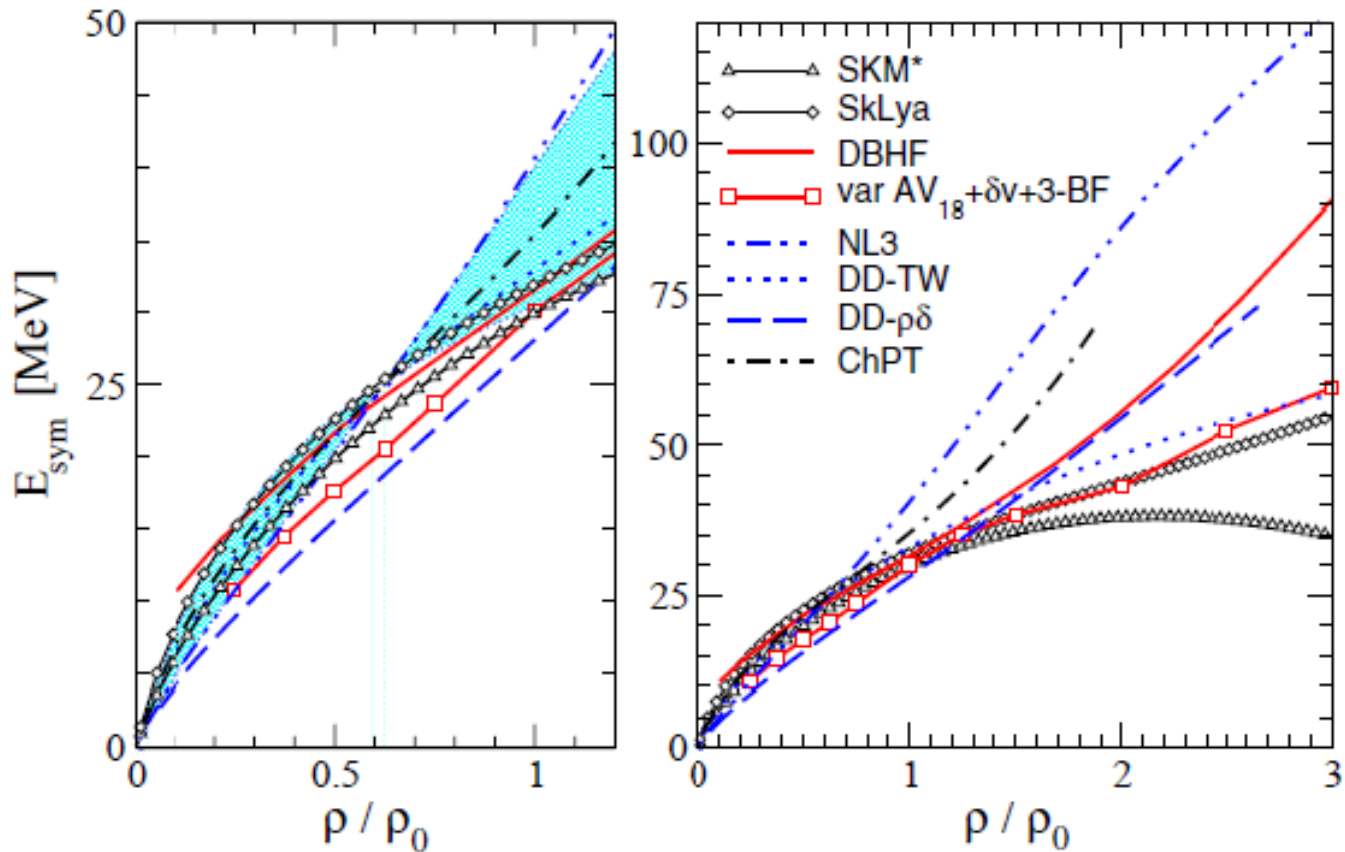
$K_{\text{nm}} = 200 \text{ MeV}$: "soft" EOS

$K_{\text{nm}} = 380 \text{ MeV}$: "stiff" EOS

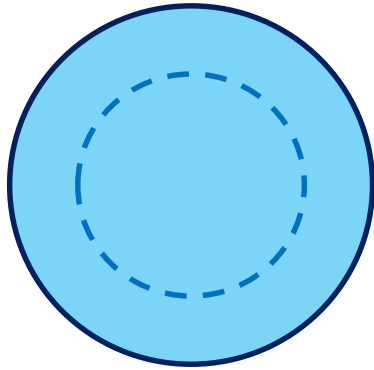
The symmetry energy

$$E_A(\rho, \delta) = E_A(\rho, 0) + E_{\text{sym}}(\rho) \cdot \delta^2 + O(\delta^4)$$

with asymmetry parameter $\delta = (\rho_n - \rho_p) / \rho$



Nuclear Incompressibility from Giant Monopole Resonances ($\rho \leq \rho_0$)



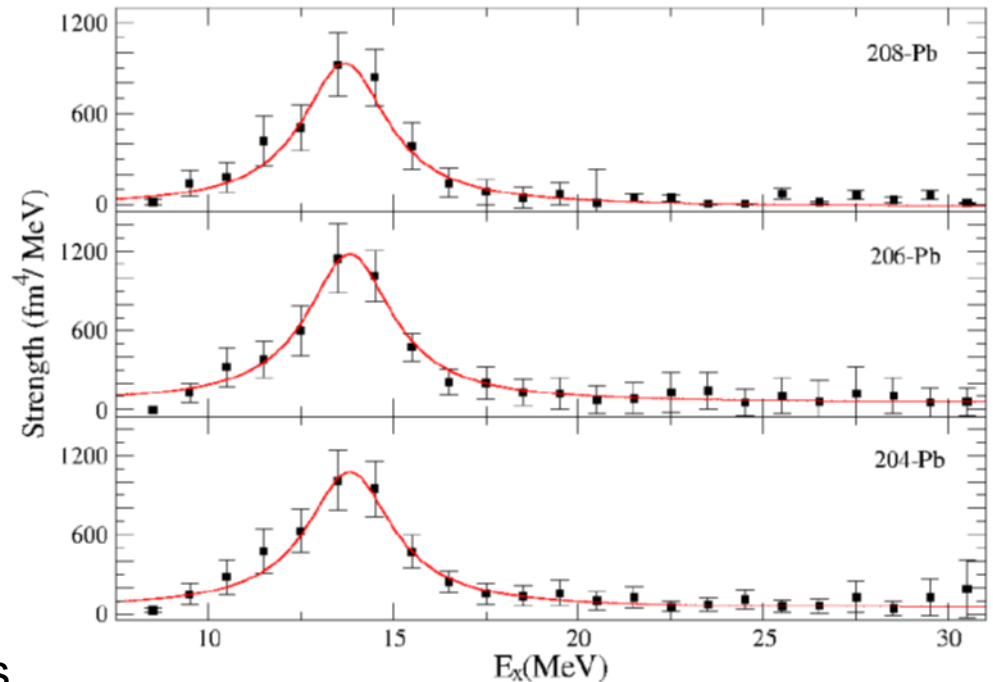
GMR

Energy of the GMR

$$E_{\text{GMR}} = \sqrt{\frac{\hbar^2 K_A}{m \langle r^2 \rangle_0}}$$

K_A : incompressibility of the nucleus

inelastic scattering 400 MeV α + Pb
at very forward angles



U. Garg, Progress of Theoretical Physics Supplement
No. 196, 2012

Nuclear Incompressibility from Giant Monopole Resonances ($\rho \leq \rho_0$)

Nuclear incompressibility:

$$K_{nm} = 9\rho^2 \frac{d^2(E/A)}{d\rho^2} \quad \text{at } \rho = \rho_0$$

Energy of the GMR

$$E_{GMR} = \sqrt{\frac{\hbar^2 K_A}{m \langle r^2 \rangle_0}}$$

K_A : incompressibility of the nucleus

$$K_A = K_{vol} + K_{surf} A^{-1/3} + K_{sym} (N-Z)^2/A^2 + K_{Coul} Z^2/A^{4/3} + \dots \quad (\text{liquid drop model})$$

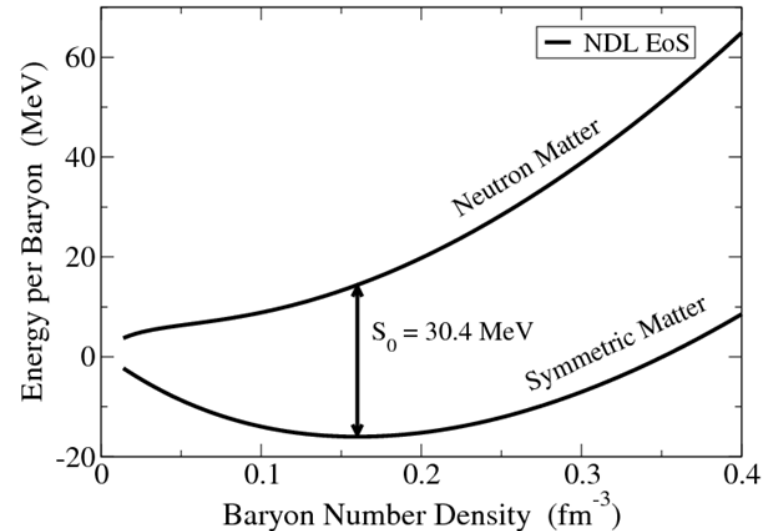
$$K_{nm} = \lim_{A \rightarrow \infty} K_A = K_{vol}$$

400 MeV α + ^{206}Pb :

$$K_{nm} = 220 - 240 \text{ MeV}$$

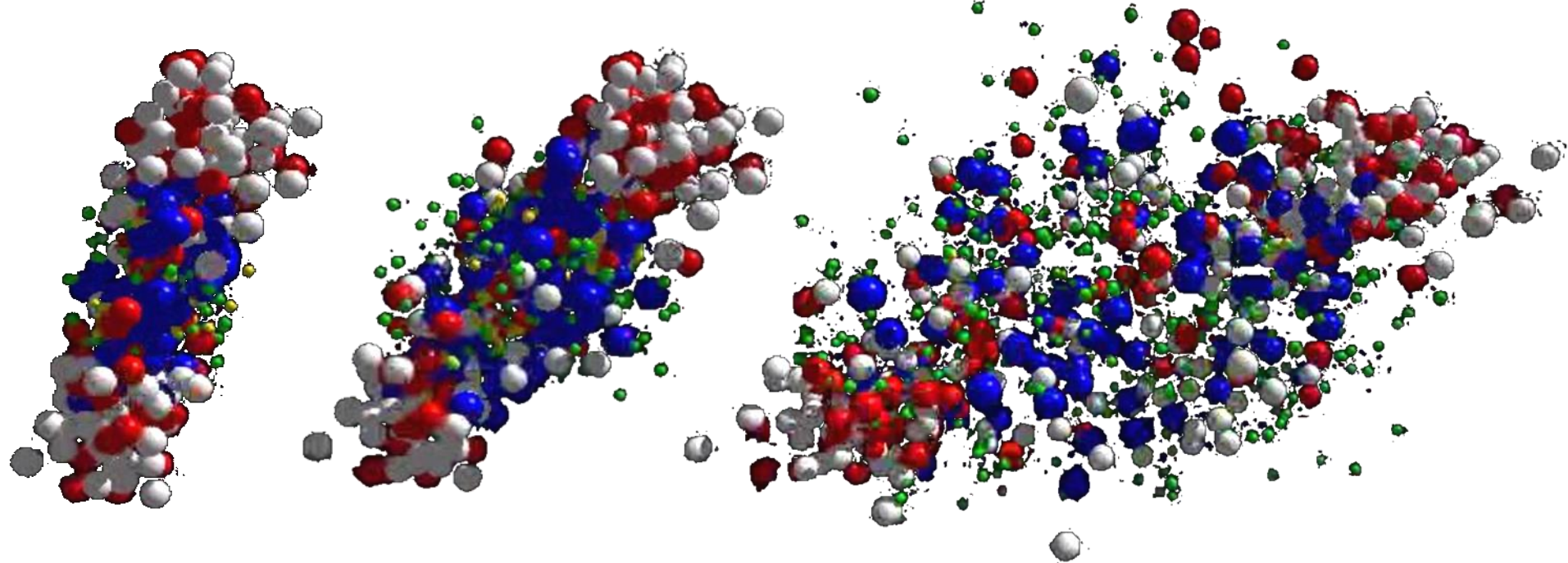
400 MeV α + $^{112-124}\text{Sn}$:

$$K_{sym} = 400 - 600 \text{ MeV}$$



The EOS of nuclear matter at $\rho > \rho_0$

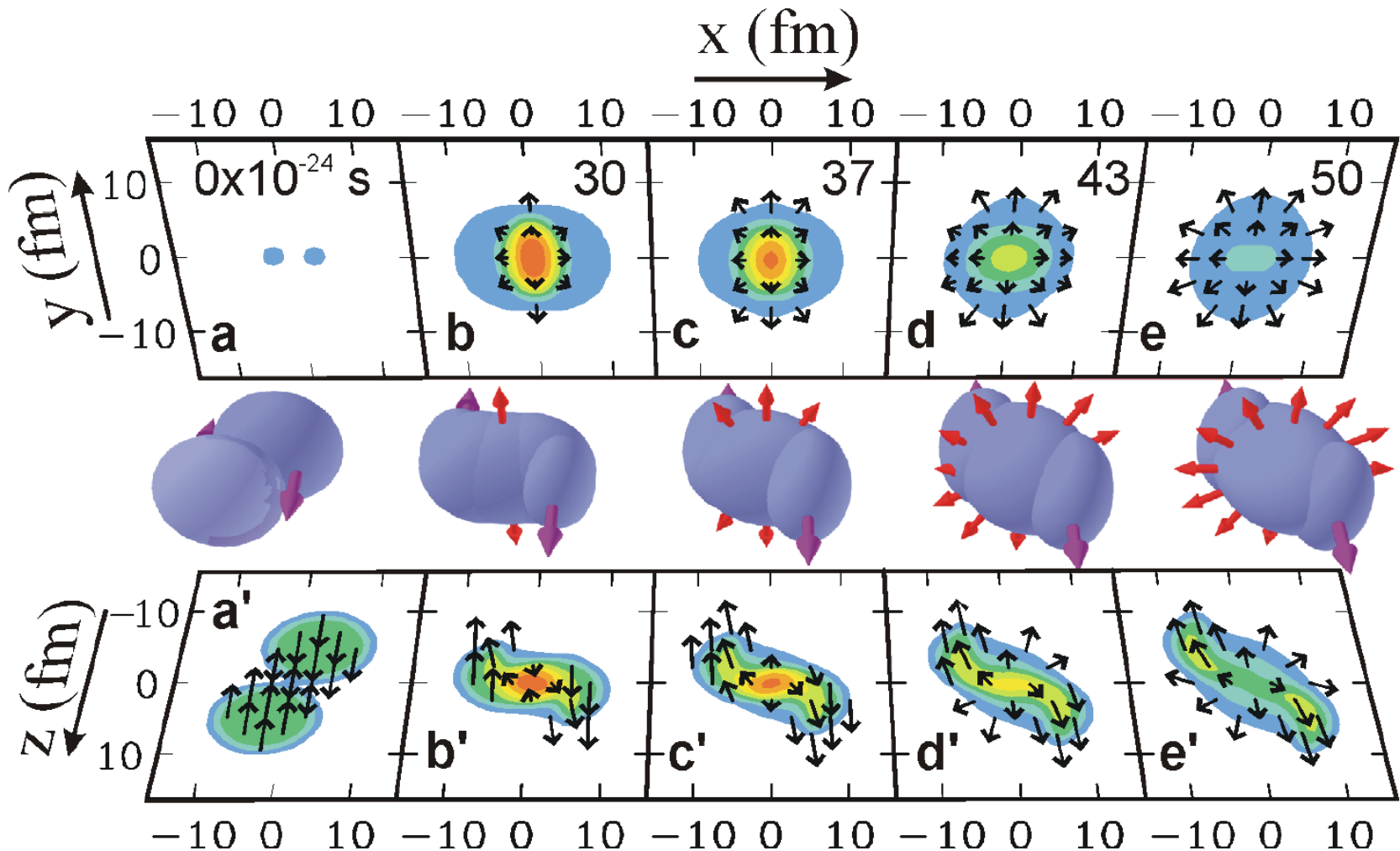
Production of dense nuclear matter in high-energy heavy-ion collisions



Volume about 1000 fm^3 , lifetime 10^{-22} s

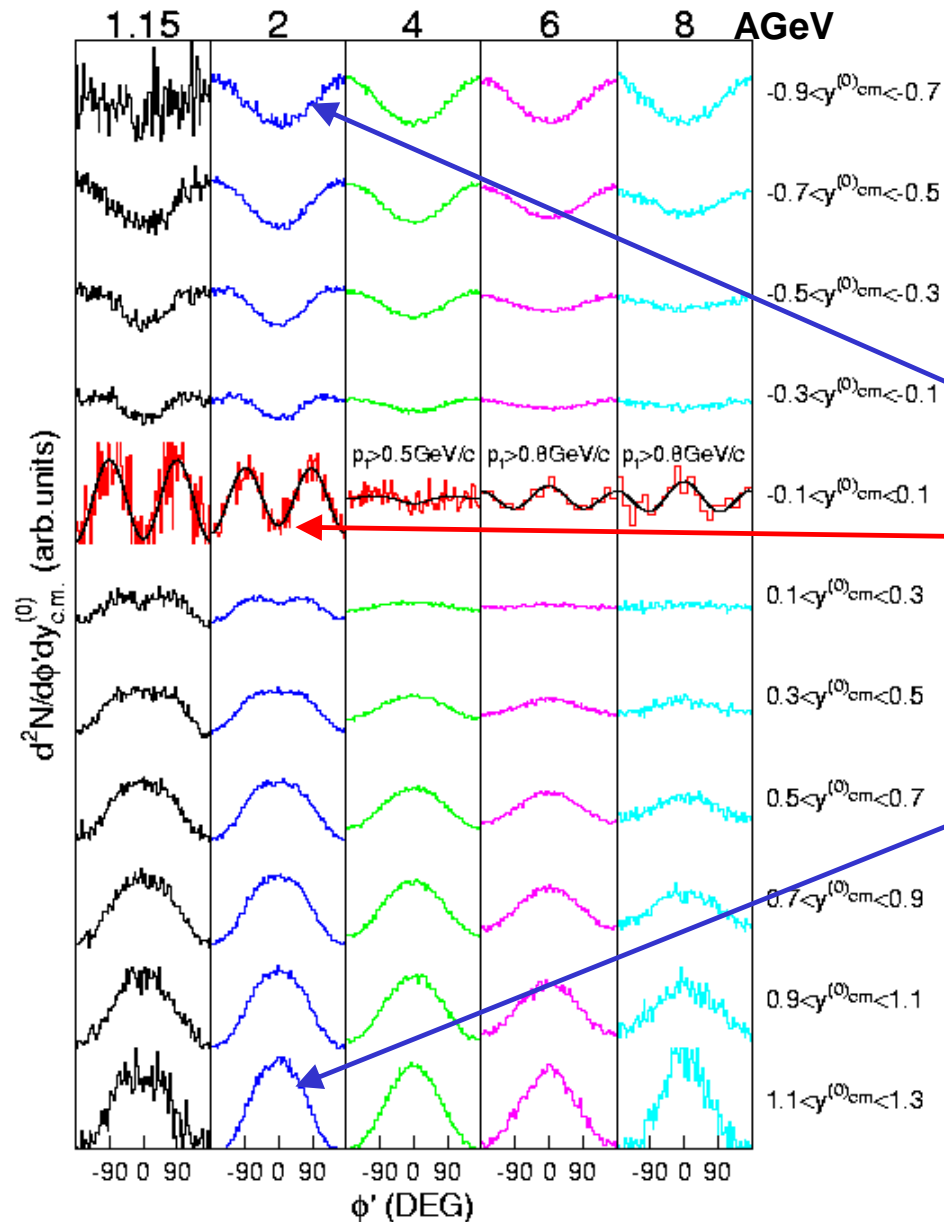
- Finite size, finite lifetime effects?
- Observables:
 - Collective flow
 - Subthreshold particle production

Dynamics of a semi-central Au+Au collision at 2 AGeV

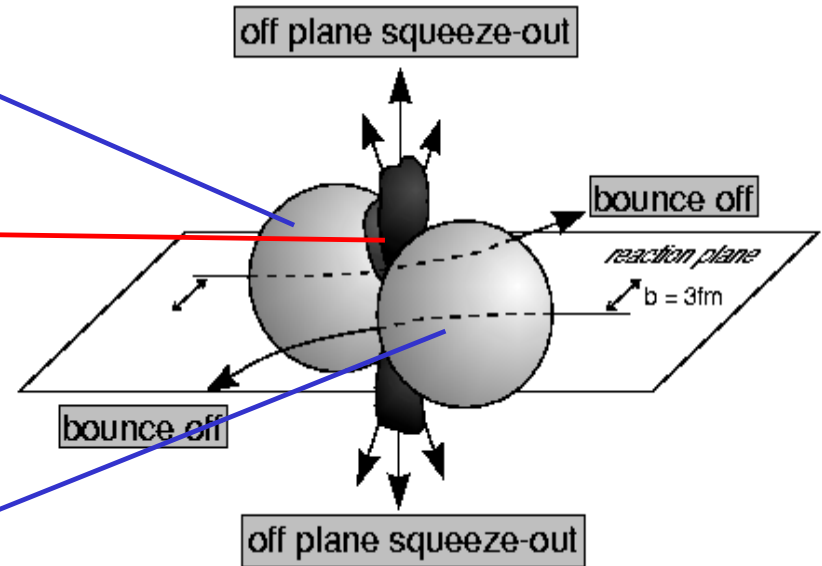


Collective flow of nucleons: driven by pressure gradient

Azimuthal angular distribution of protons in Au+Au collisions at 1.15 - 8 AGeV



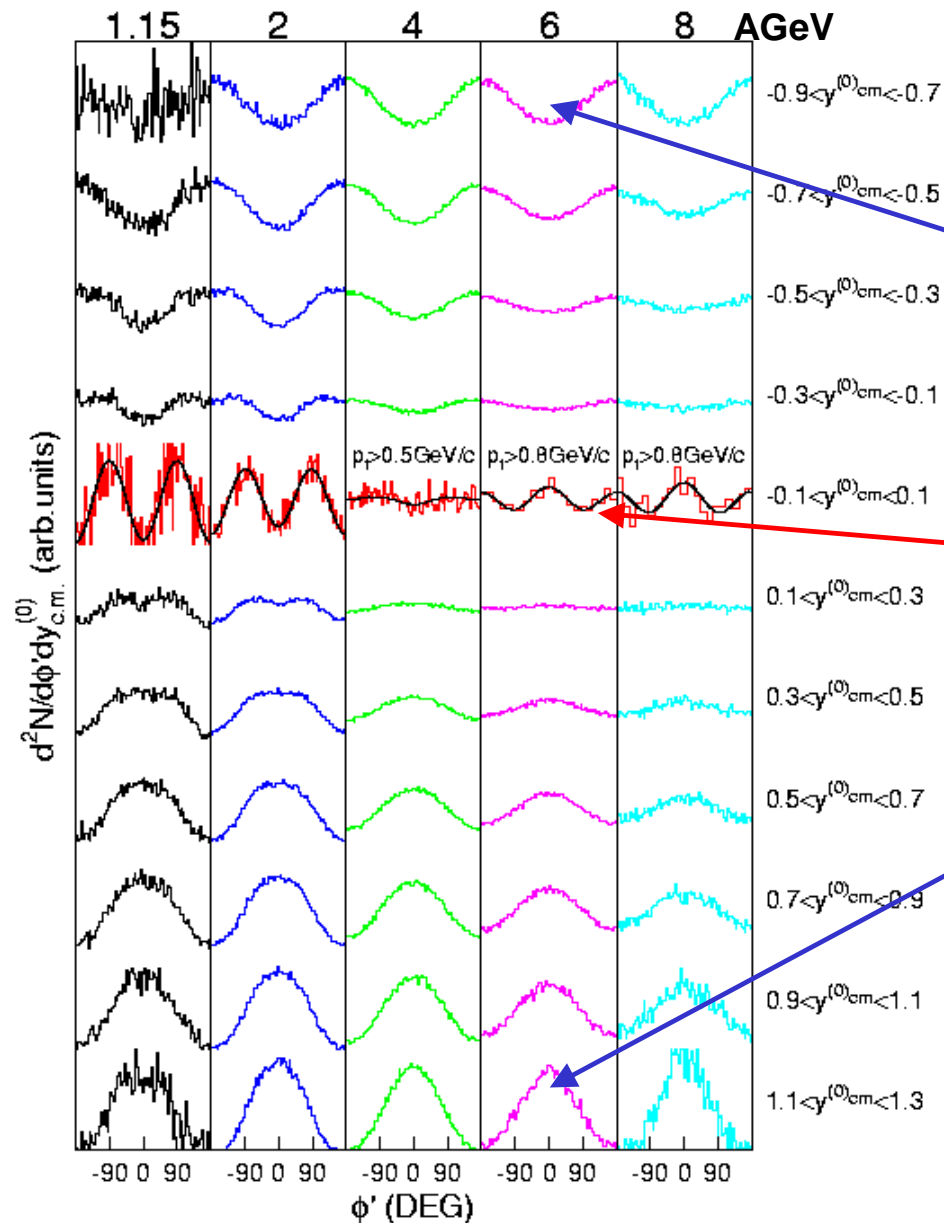
C. Pinkenburg et al., (E895),
Phys. Rev. Lett. 83 (1999) 1295



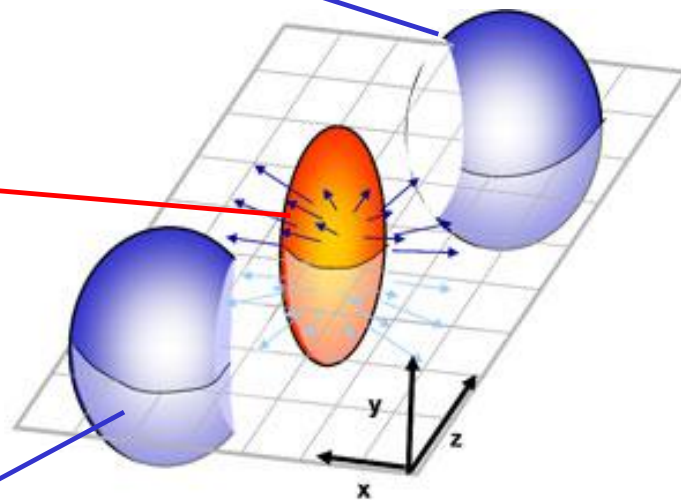
Azimuthal angle distribution:
 $dN/d\Phi \propto (1 + 2v_1 \cos\Phi + 2v_2 \cos 2\Phi)$

Rapidity: $y^{(0)} = y - y_m$
with $y = 0.5 \ln [(E+p_z)/(E-p_z)]$

Azimuthal angular distribution of protons measured in Au+Au collisions at 1.15, 2, 4, 6, 8 AGeV



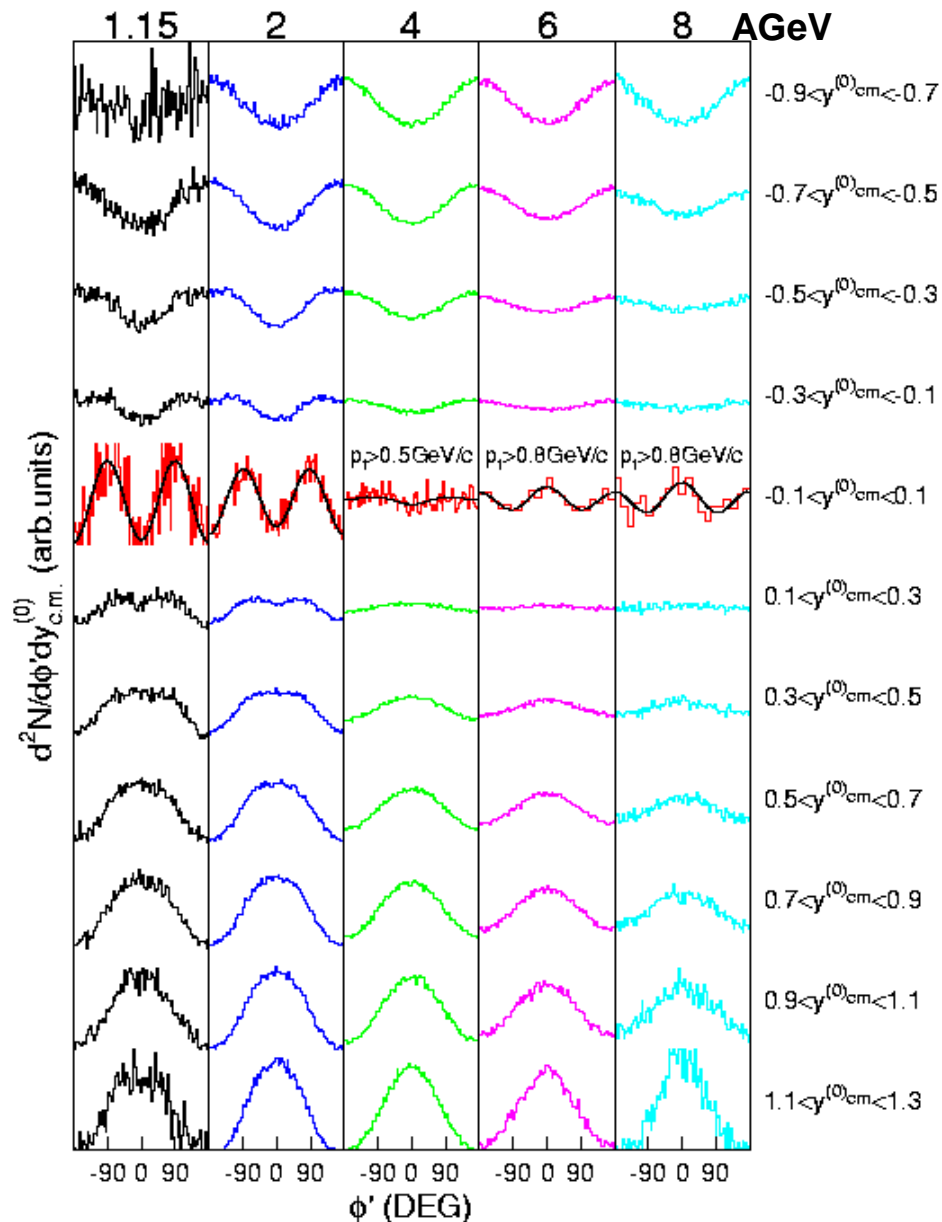
C. Pinkenburg et al., (E895),
Phys. Rev. Lett. 83 (1999) 1295



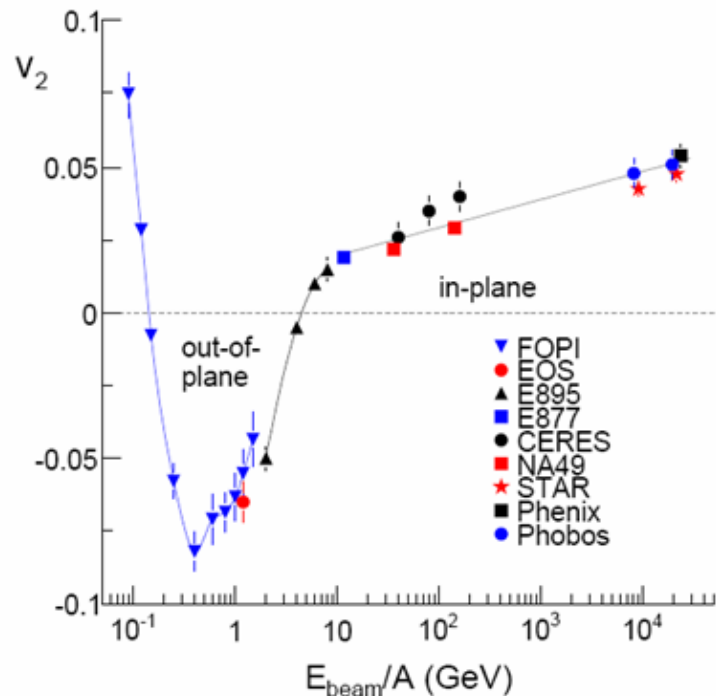
Azimuthal angle distribution:
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Azimuthal angular distribution of protons measured in Au+Au collisions at 1.15, 2, 4, 6, 8 AGeV



C. Pinkenburg et al., (E895),
Phys. Rev. Lett. 83 (1999) 1295



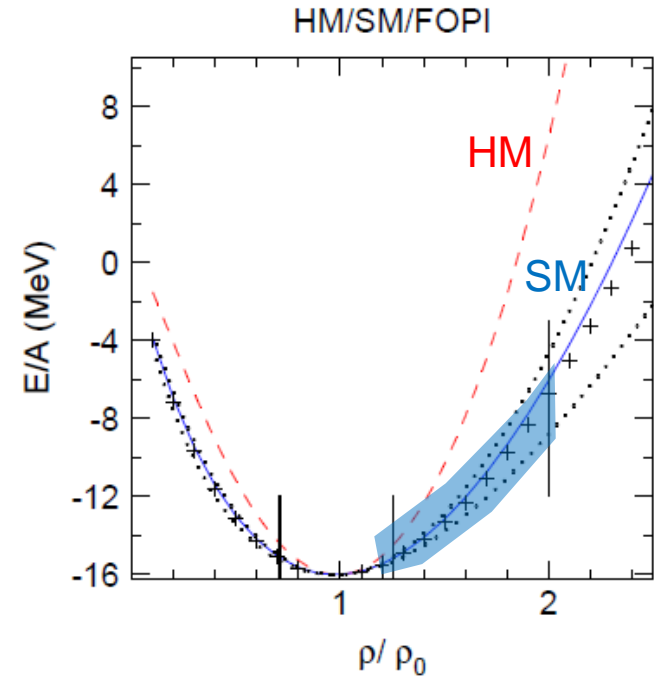
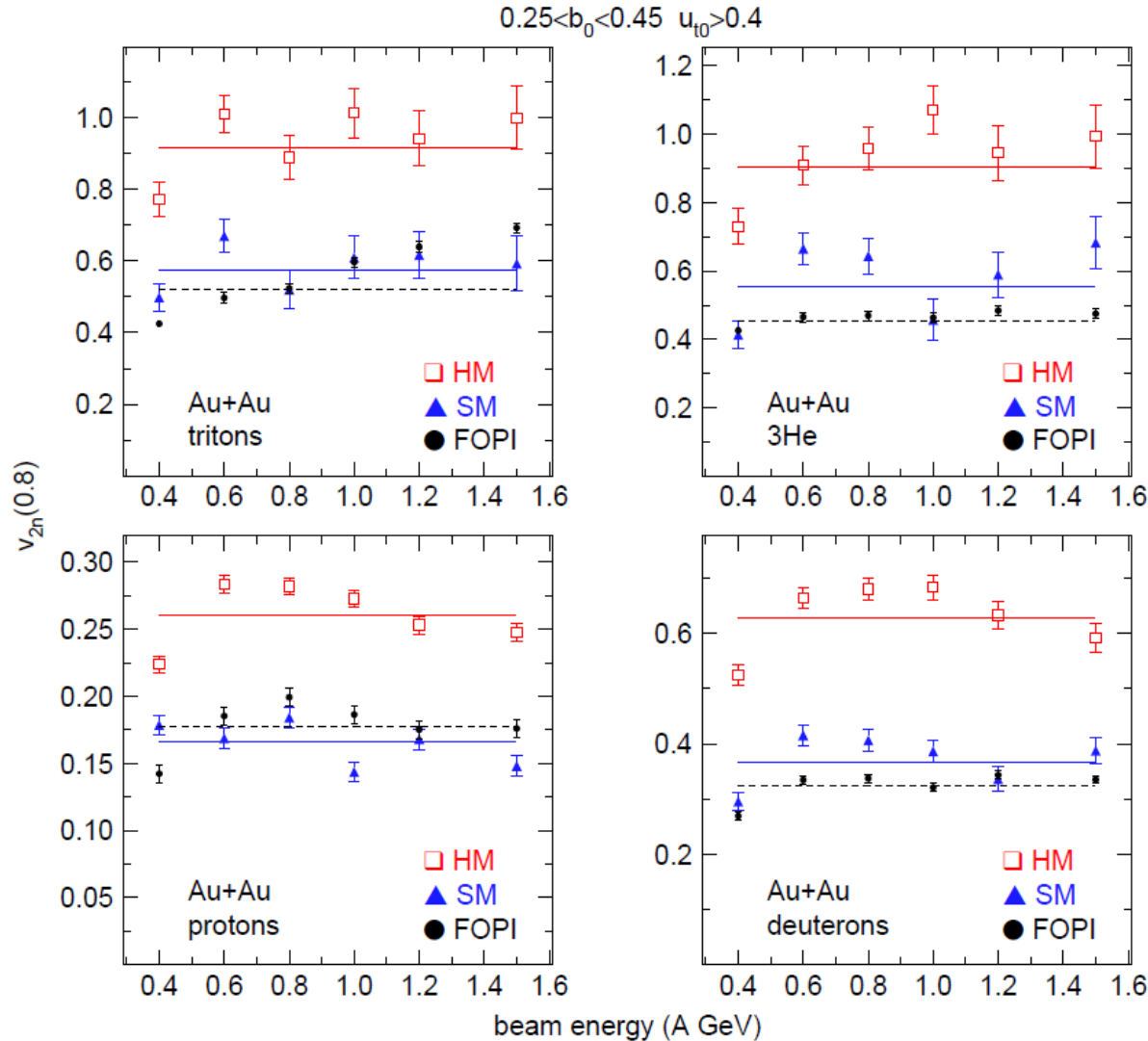
Azimuthal angle distribution:

$$dN/d\Phi \propto (1 + 2v_1 \cos\Phi + 2v_2 \cos 2\Phi)$$

Rapidity: $y^{(0)} = y - y_m$
 with $y = 0.5 \ln [(E+p_z)/(E-p_z)]$

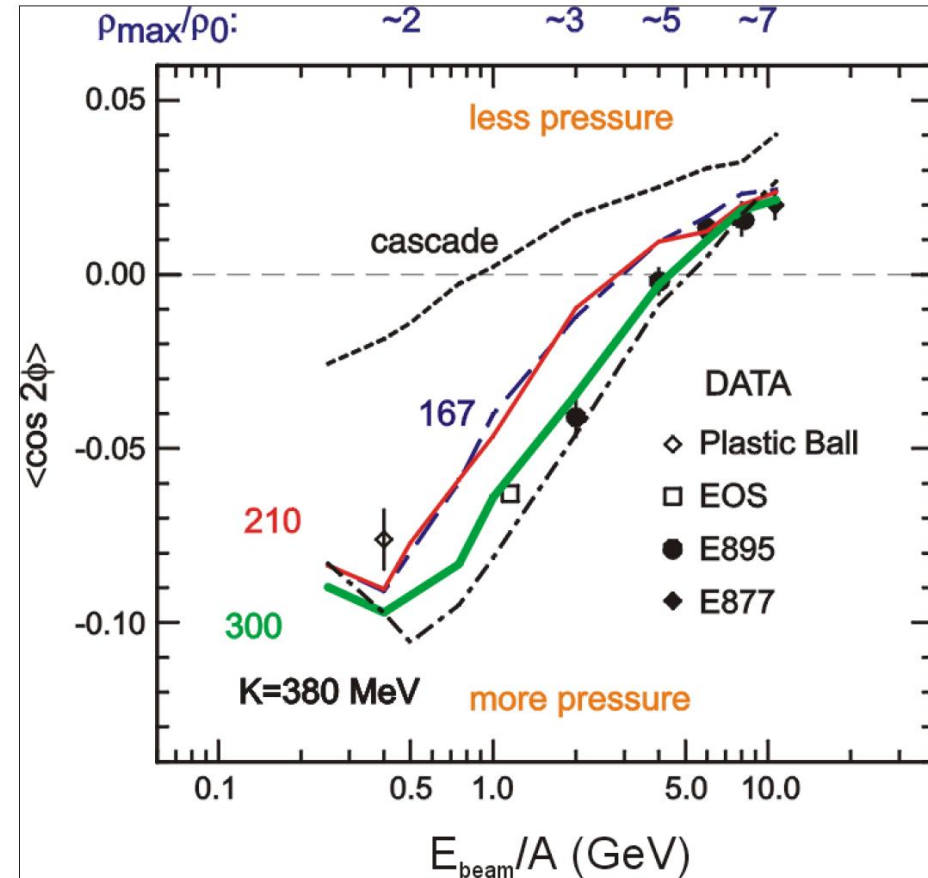
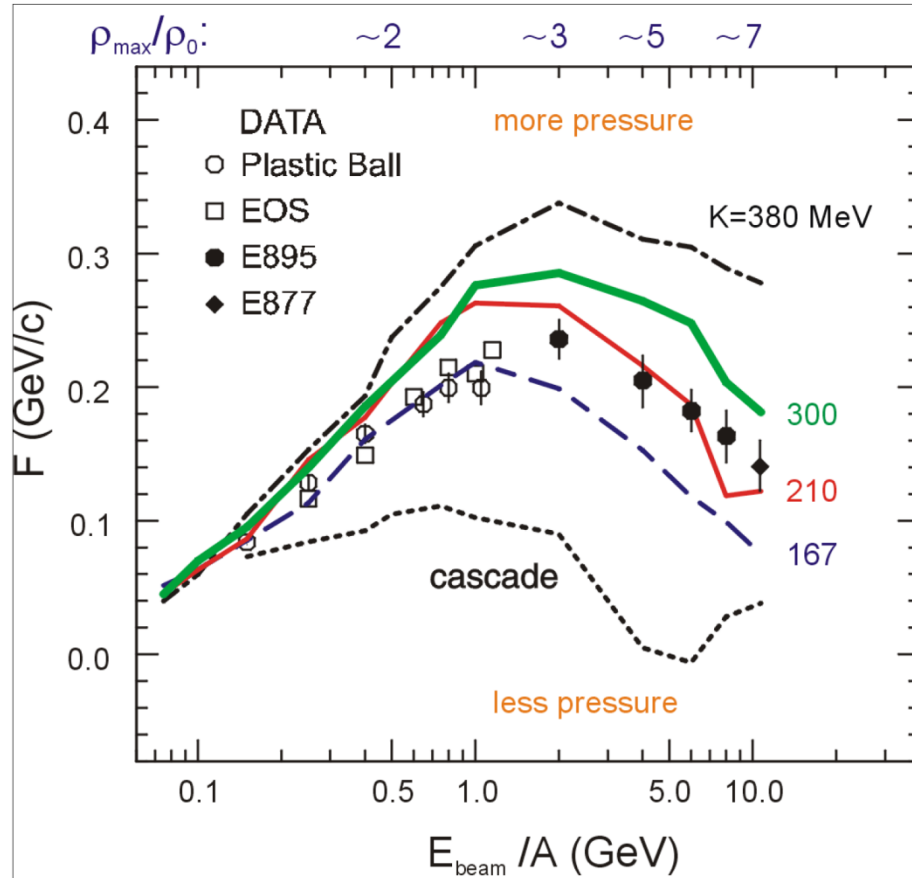
EOS from the elliptic flow of fragments in Au+Au collisions at SIS18 energies ($\rho < 3\rho_0$)

A. Le Fevre , Y Leifels, W. Reisdorf, J. Aichelin, Ch. Hartnack, arXiv:1501.02546



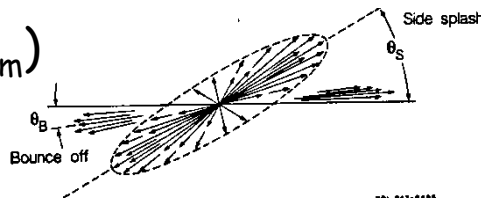
Nuclear incompressibility from collective proton flow

P. Danielewicz, R. Lacey, W.G. Lynch, Science 298 (2002) 1592

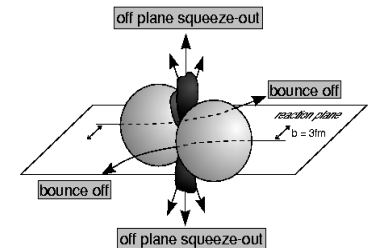


Transverse in-plane flow:

$$F = d(p_x/A)/d(y/y_{cm})$$

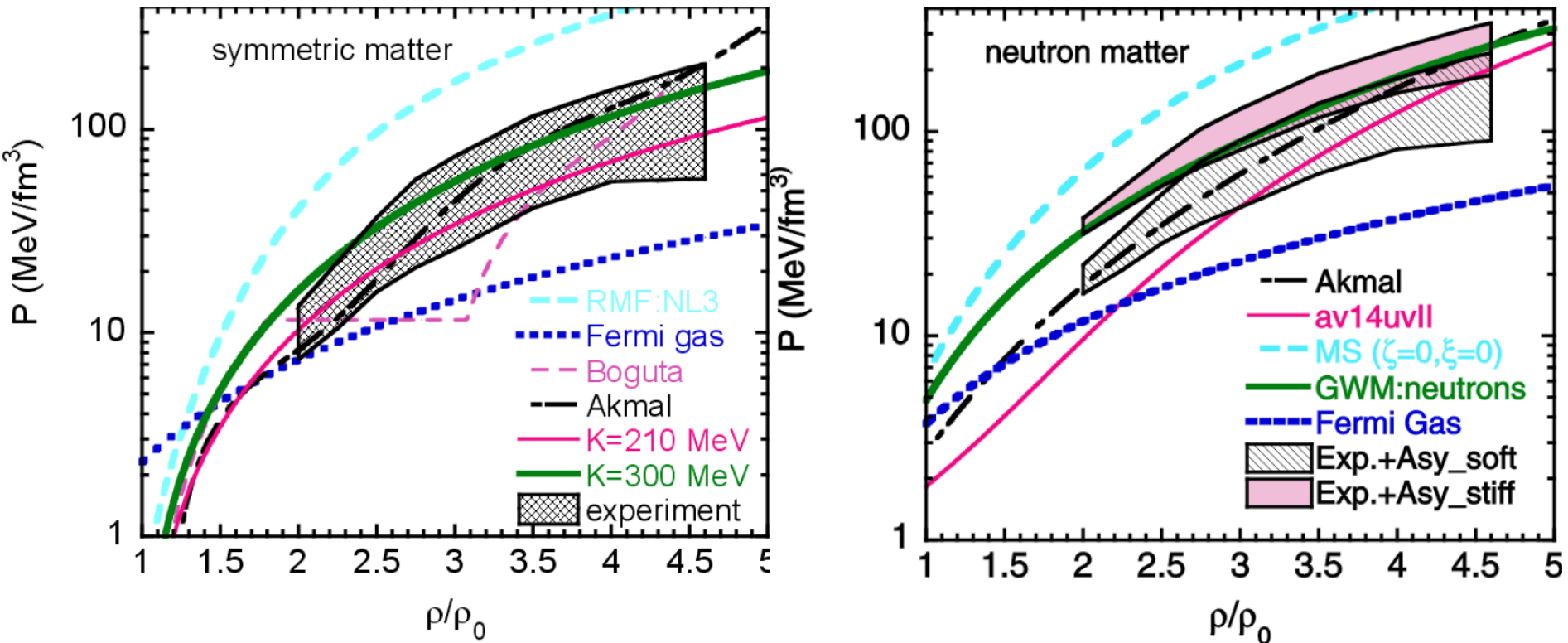


Elliptic flow:

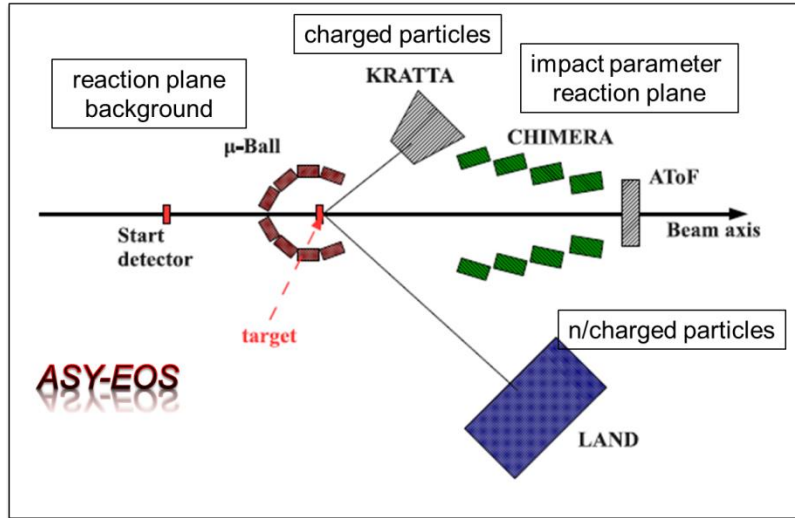


$$dN/d\Phi \propto (1 + 2v_1 \cos\Phi + 2v_2 \cos^2\Phi)$$

Nuclear equation-of-state from collective proton flow

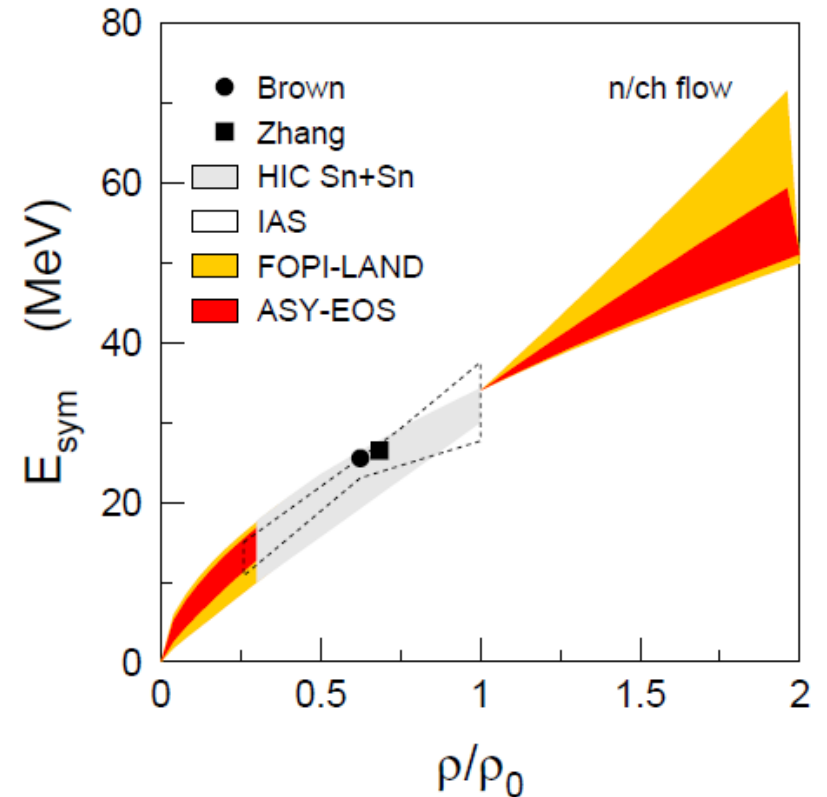
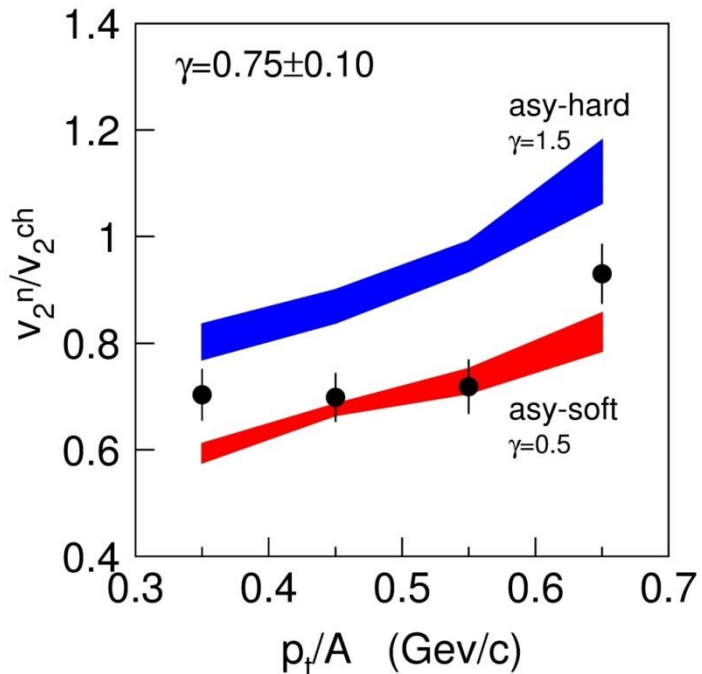


The ASY-EOS experiment at GSI: The symmetry energy at supra-saturation density



P. Russotto et al.,
Phys. Rev. C 94, 034608 (2016)

Measuring the elliptic flow of
charged particles and neutrons
in Au+Au collisions at 400A MeV



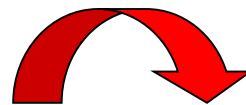
Collective flow in Au+Au collisions and the nuclear matter EOS

Beam energy A GeV	central density	flow observable	incompressibility K
0.4 – 1.5	$\rho = 1 - 3 \rho_0$	v_2 of p, d, t, 3He	≈ 200 MeV
2 – 10	$\rho = 3 - 7 \rho_0$	v_1 of p,	≈ 200 MeV
2 – 10	$\rho = 3 - 7 \rho_0$	v_2 of p,	≈ 300 MeV

Within microscopic transport models the collective flow is sensitive to:

- The nuclear matter equation of state
- In-medium nucleon-nucleon cross sections
- Momentum dependent interactions

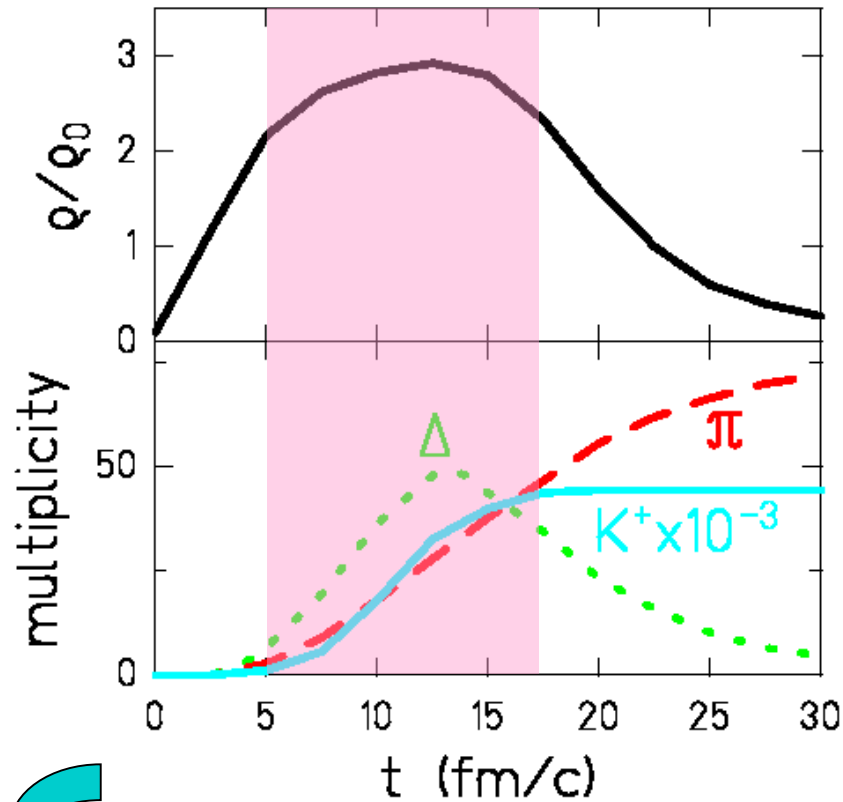
Independent observables?



particle production

Probing the nuclear equation-of-state ($\rho = 1 - 3 \rho_0$) by K^+ meson subthreshold production in HI collisions

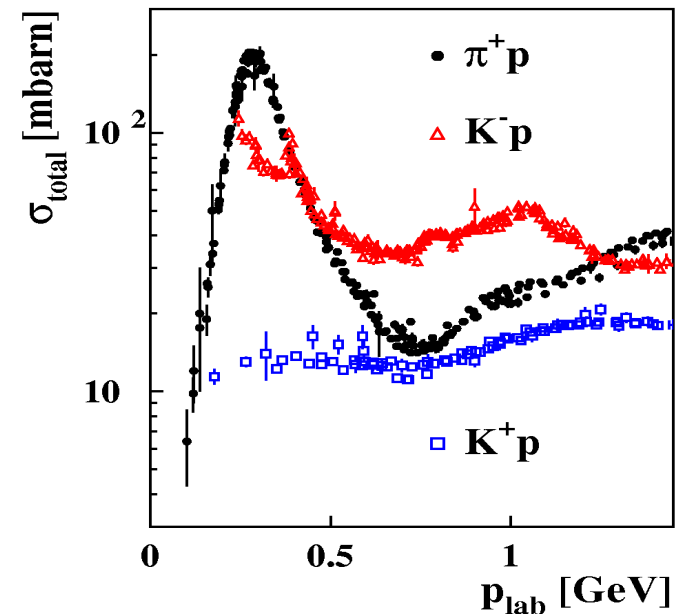
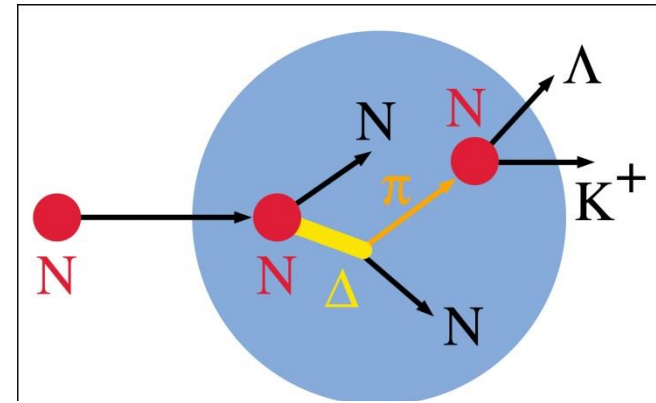
RBUU: Au+Au 1 AGeV, $b=0$ fm



K^+ mesons probe high densities

K^+ mesons scatter elastically only

$pp \rightarrow K^+ \Lambda p$ ($E_{\text{thres}} = 1.6$ GeV)



Probing the nuclear equation-of-state ($\rho = 1 - 3 \rho_0$) by K^+ meson production in C+C and Au+Au collisions

Idea: K^+ yield \propto baryon density \propto compressibility

Transport model (RBUU)

Au+Au at 1 AGeV:

$\kappa = 200 \text{ MeV} \Rightarrow \rho_{\text{max}} \approx 2.9 \rho_0 \Rightarrow K^+ \nearrow$

$\kappa = 380 \text{ MeV} \Rightarrow \rho_{\text{max}} \approx 2.4 \rho_0 \Rightarrow K^+ \searrow$

Reference system C+C:

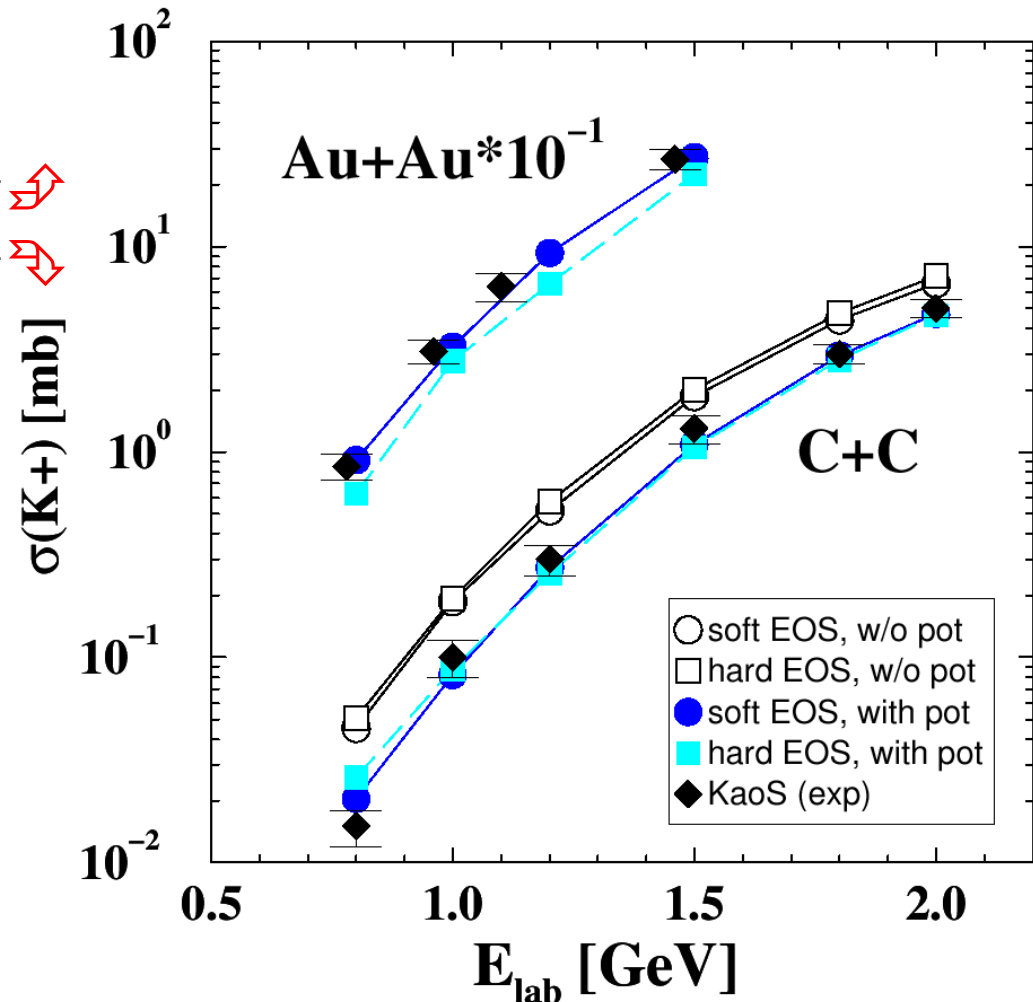
K^+ yield not sensitive to EOS

Experiment:

C. Sturm et al., (KaoS Collaboration),
Phys. Rev. Lett. 86 (2001) 39

Theory:

Ch. Fuchs et al.,
Phys. Rev. Lett. 86 (2001) 1974

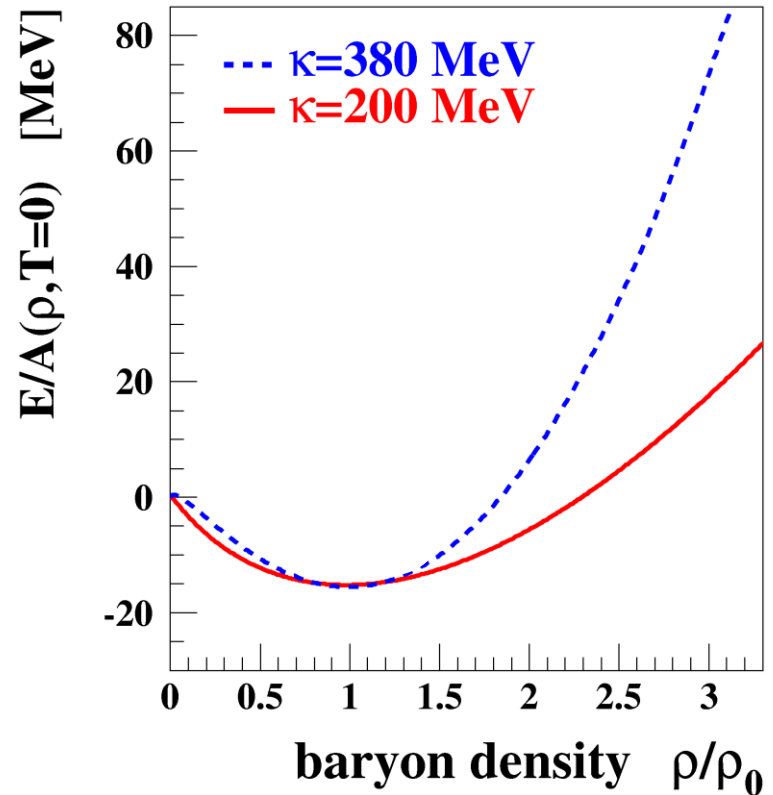
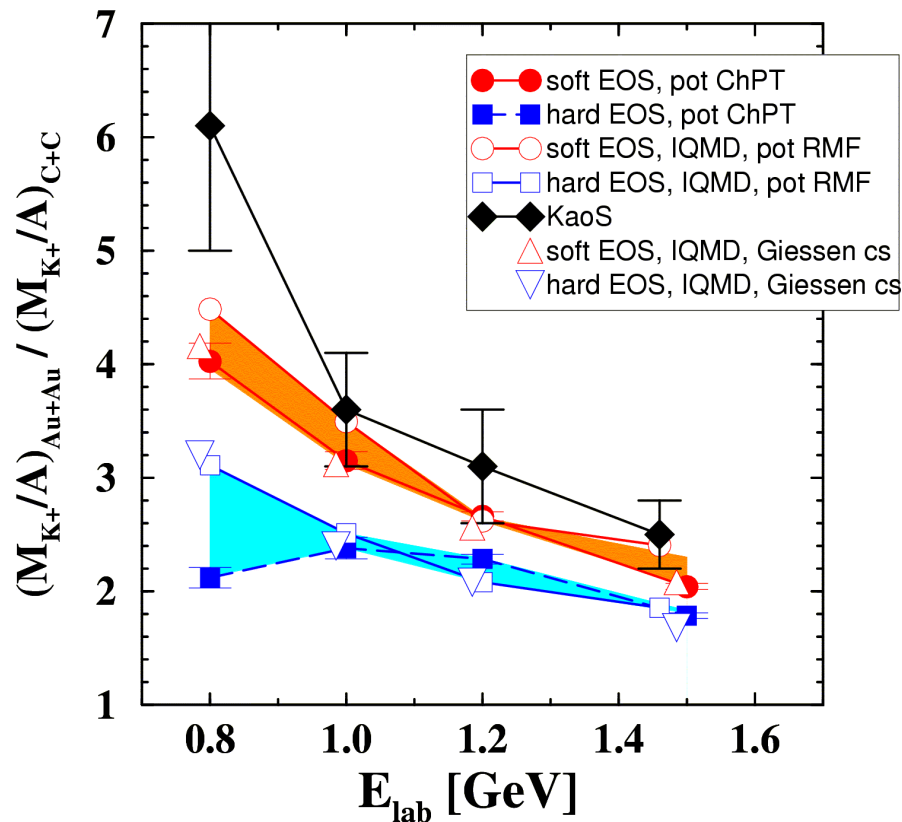


The compressibility of (symmetric) nuclear matter

Experiment: C. Sturm et al., (KaoS Collaboration) Phys. Rev. Lett. 86 (2001) 39

Theory: QMD Ch. Fuchs et al., Phys. Rev. Lett. 86 (2001) 1974

IQMD Ch. Hartnack, J. Aichelin, J. Phys. G 28 (2002) 1649



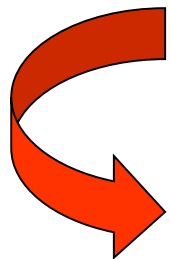
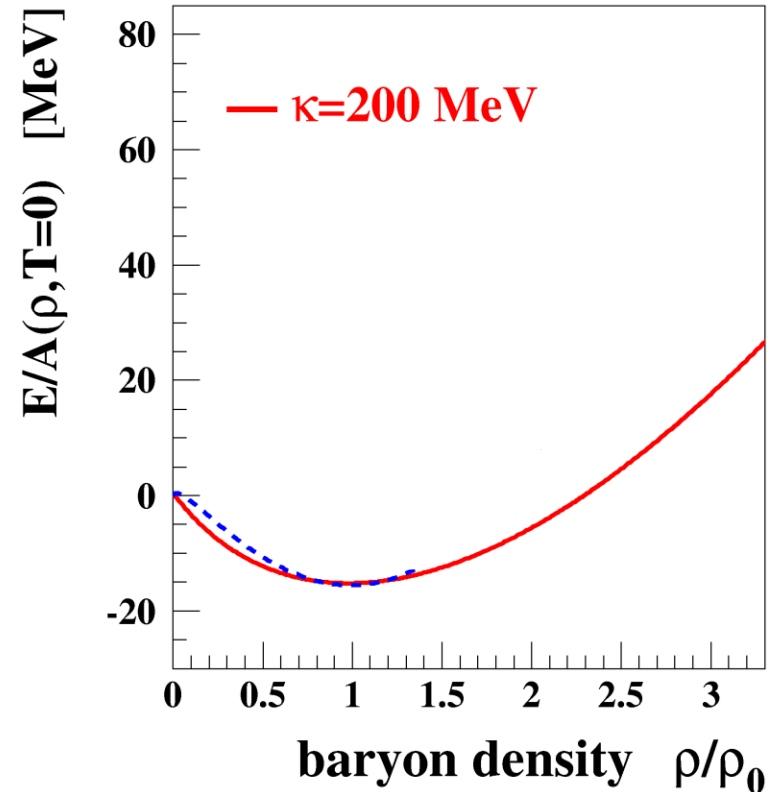
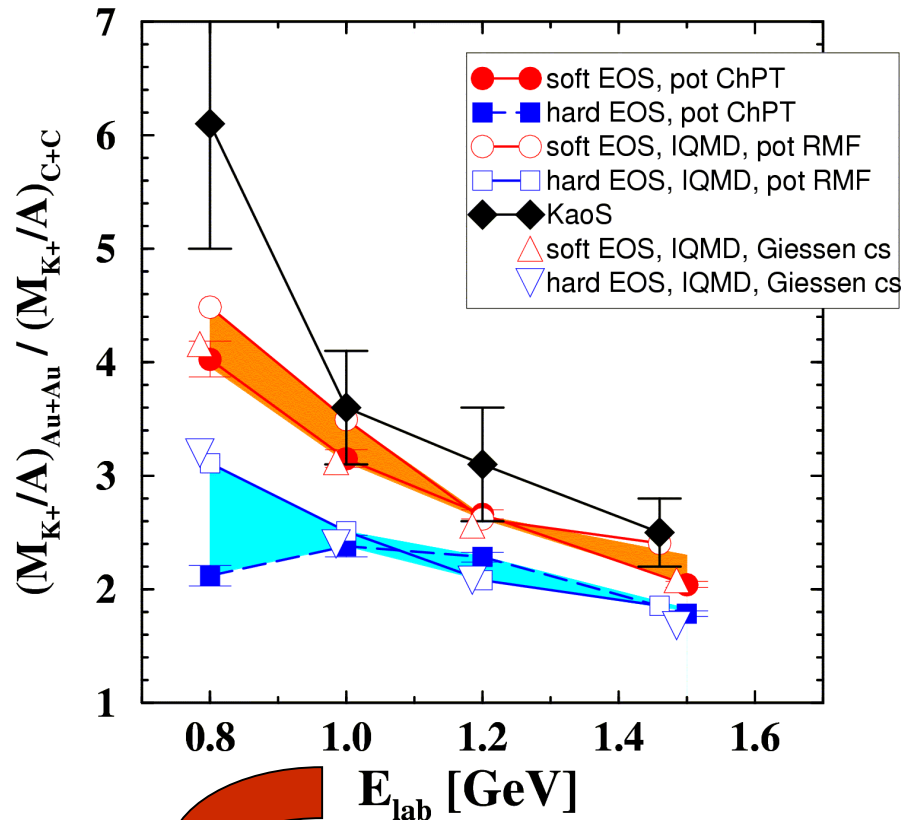
Au/C ratio: cancellation of systematic errors both in experiment and theory

The compressibility of (symmetric) nuclear matter

Experiment: C. Sturm et al., (KaoS Collaboration) Phys. Rev. Lett. 86 (2001) 39

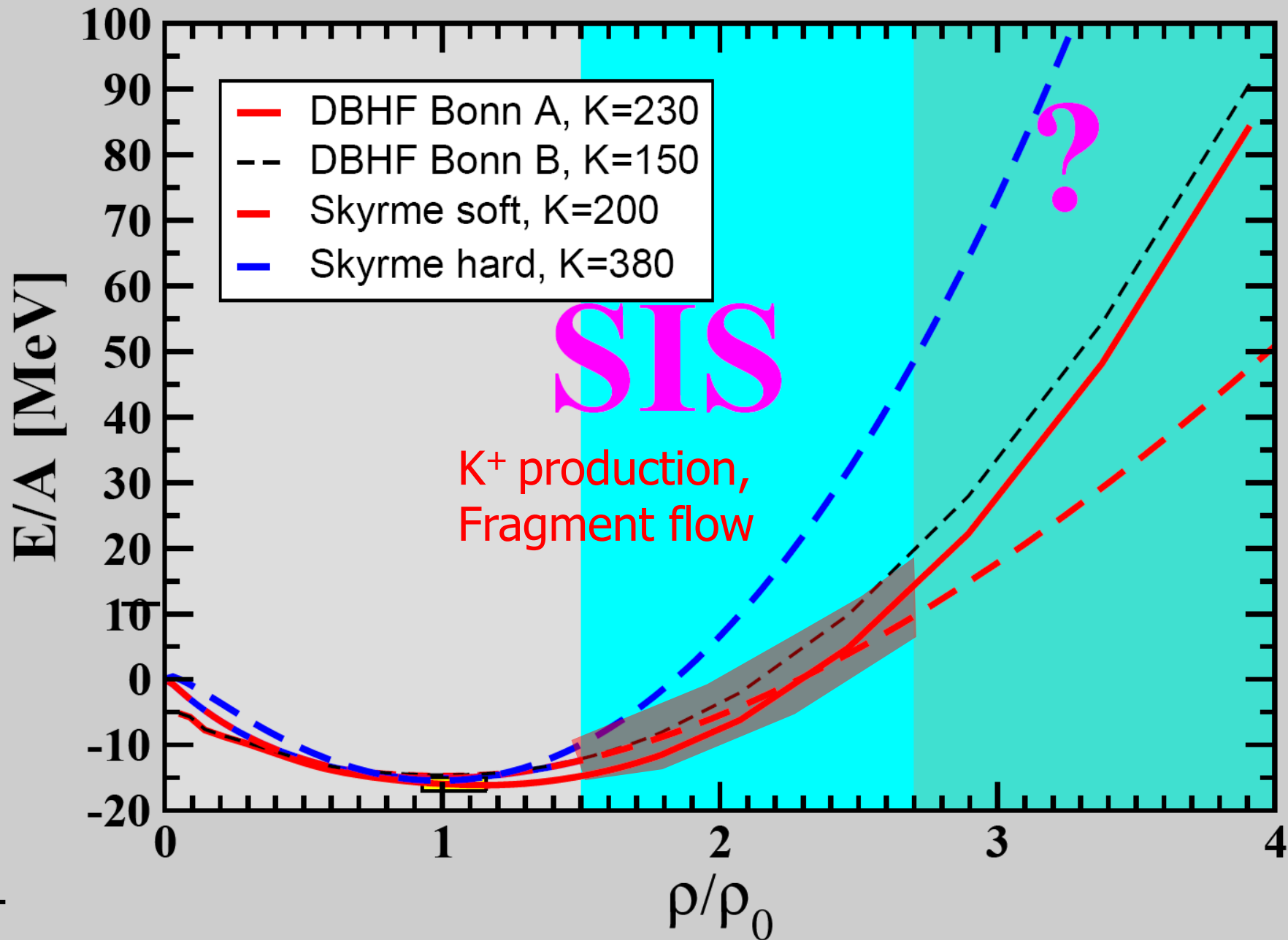
Theory: QMD Ch. Fuchs et al., Phys. Rev. Lett. 86 (2001) 1974

IQMD Ch. Hartnack, J. Aichelin, J. Phys. G 28 (2002) 1649

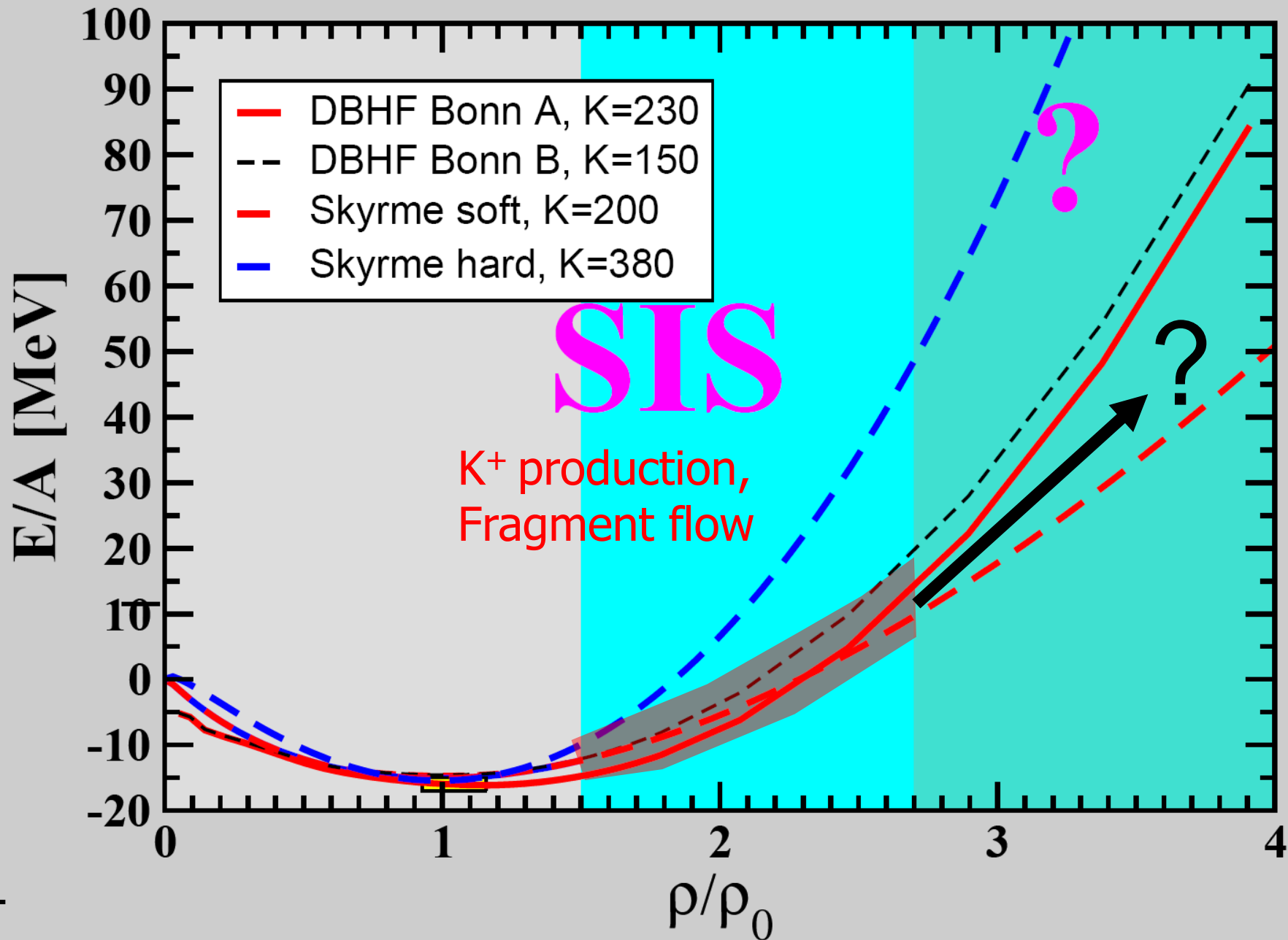


Soft equation-of-state: $\kappa \leq 200$ MeV
Confirmation of flow measurements

nuclear matter EOS



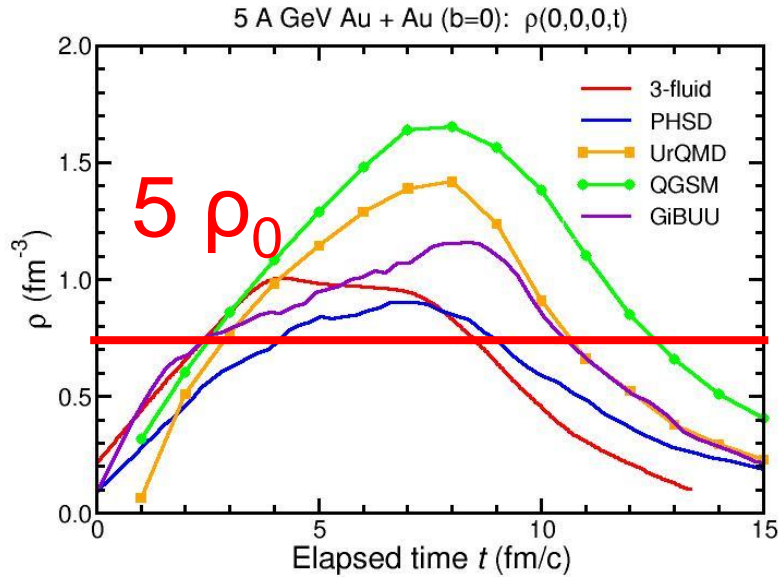
nuclear matter EOS



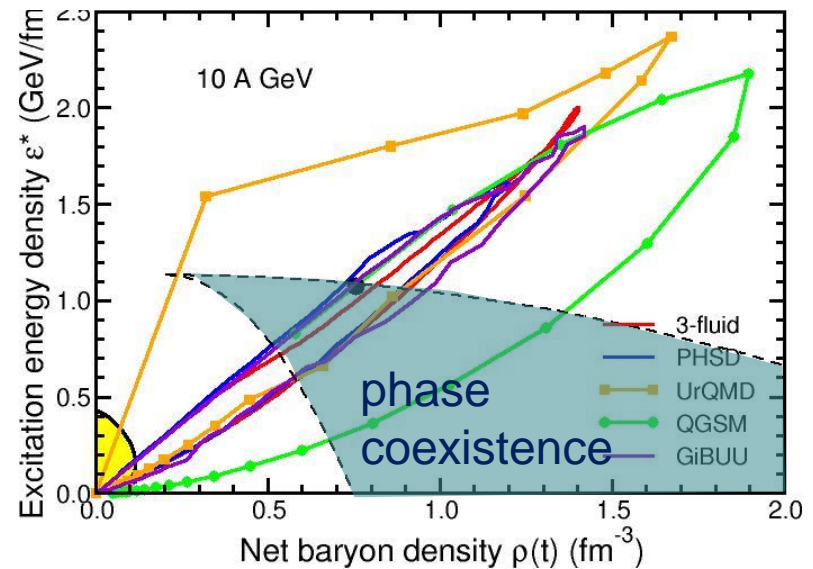
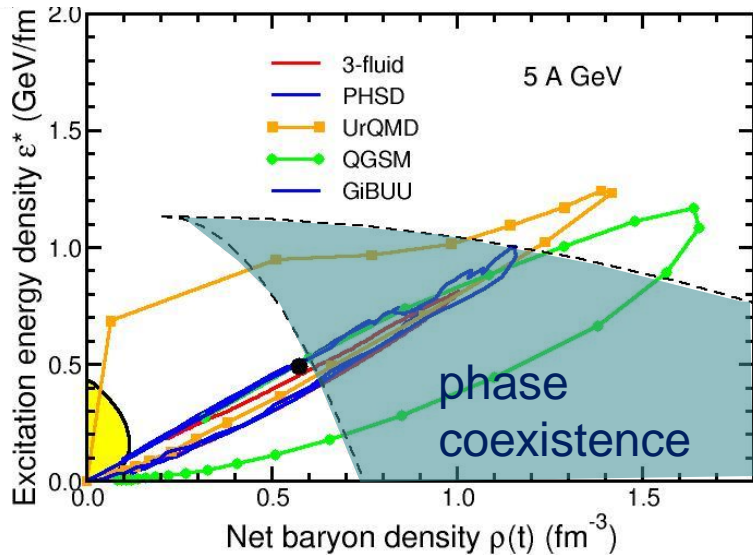
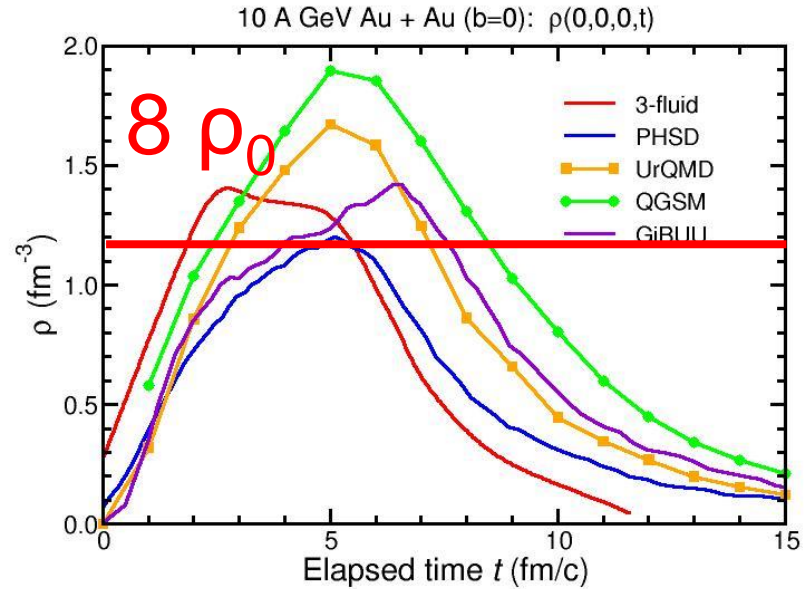
Baryon densities in central Au+Au collisions

I.C. Arsene et al., Phys. Rev. C 75, 24902 (2007)

5 A GeV

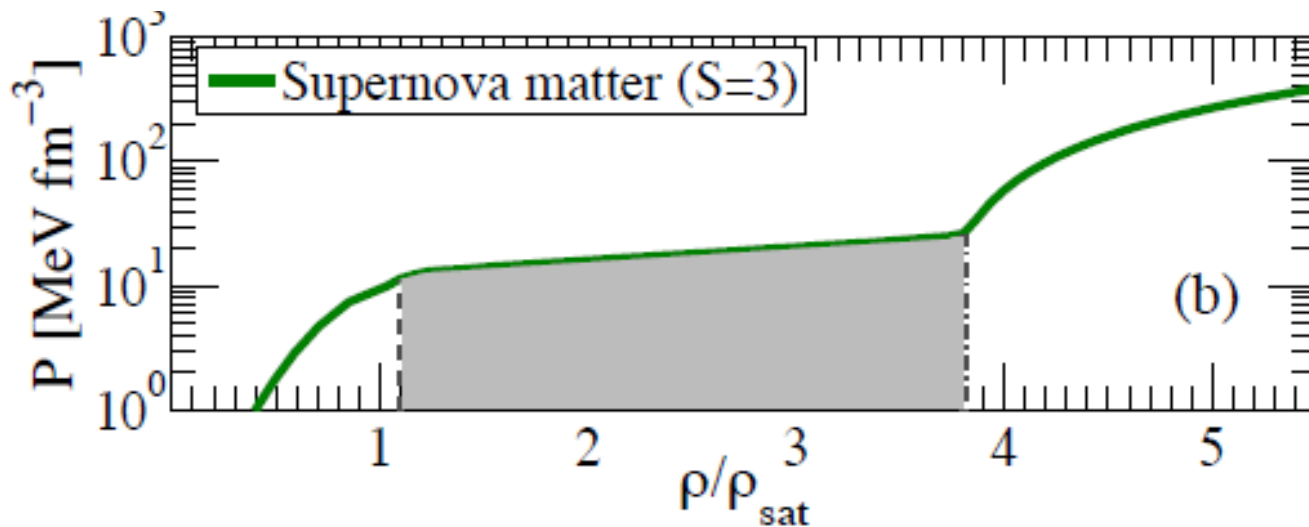
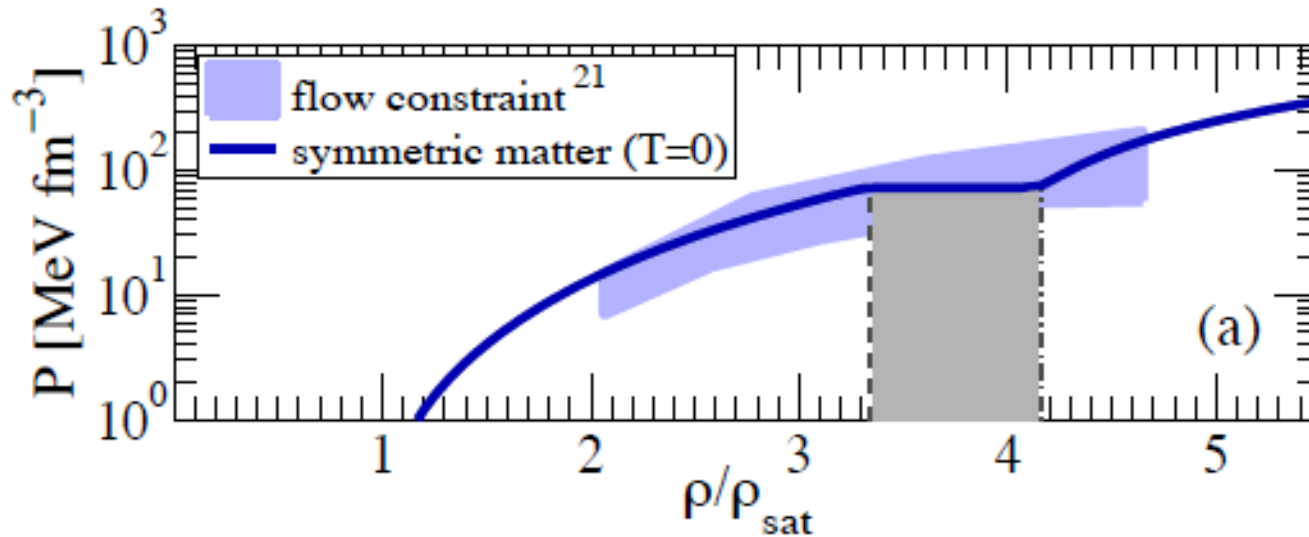


10 A GeV



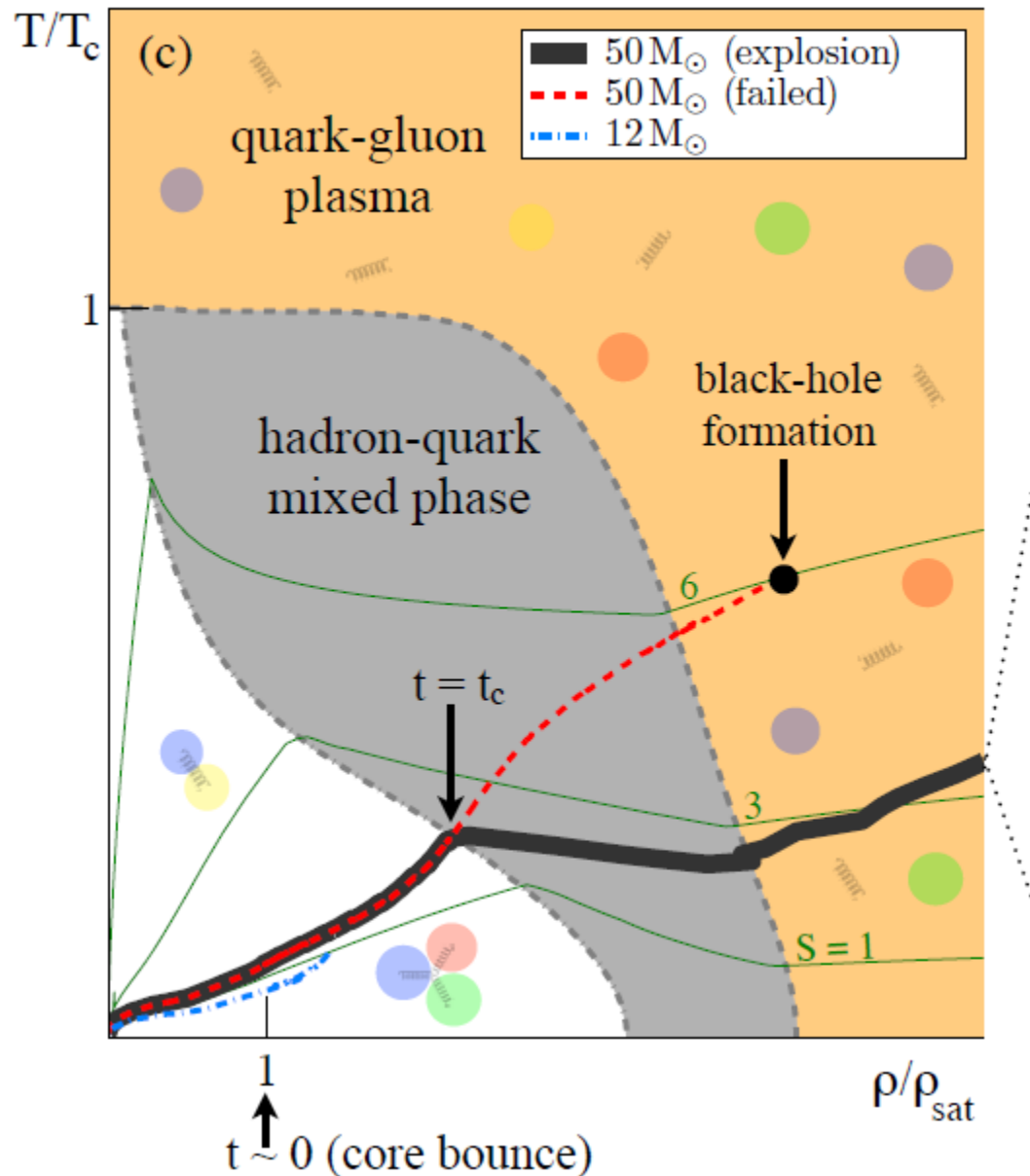
EOS with phase transition

T. Fischer, N. Bastian, M. Wu, S. Typel, T. Klähn, D. Blaschke, arXiv:1712.08788



Supernova explosion of a supermassive star

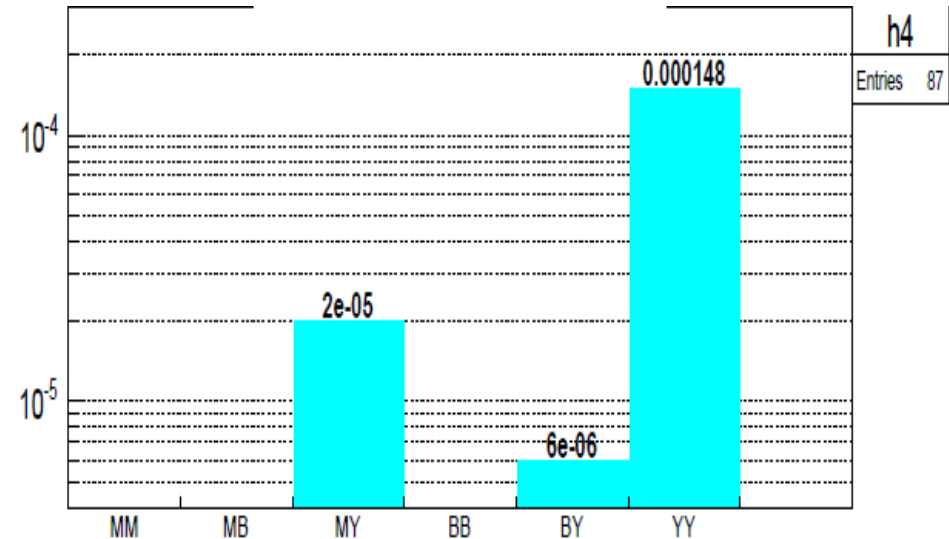
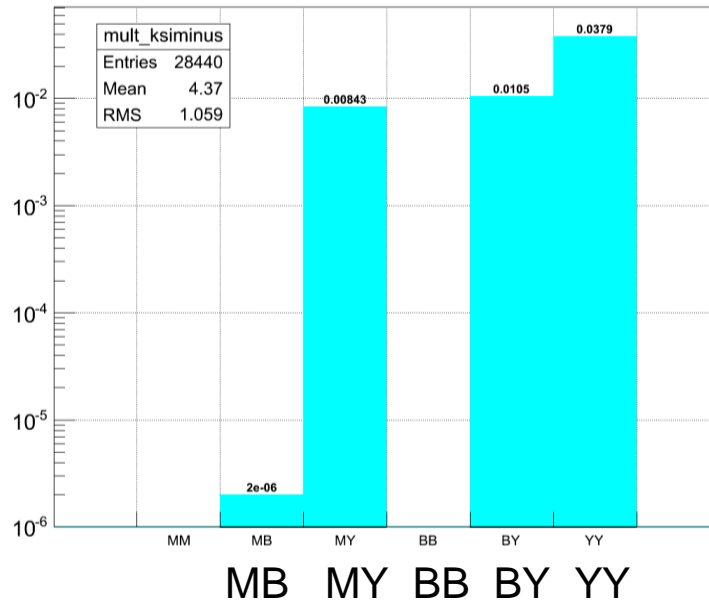
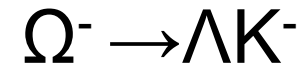
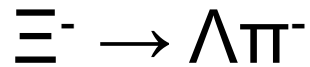
T. Fischer, N. Bastian, M. Wu, S. Typel, T. Klähn, D. Blaschke, arXiv:1712.08788



The equation-of-state of symmetric nuclear matter at neutron star core densities

Observable: multistrange hyperon production at (sub)threshold energies

HYPQGSM calculations for 4 A GeV Au+Au



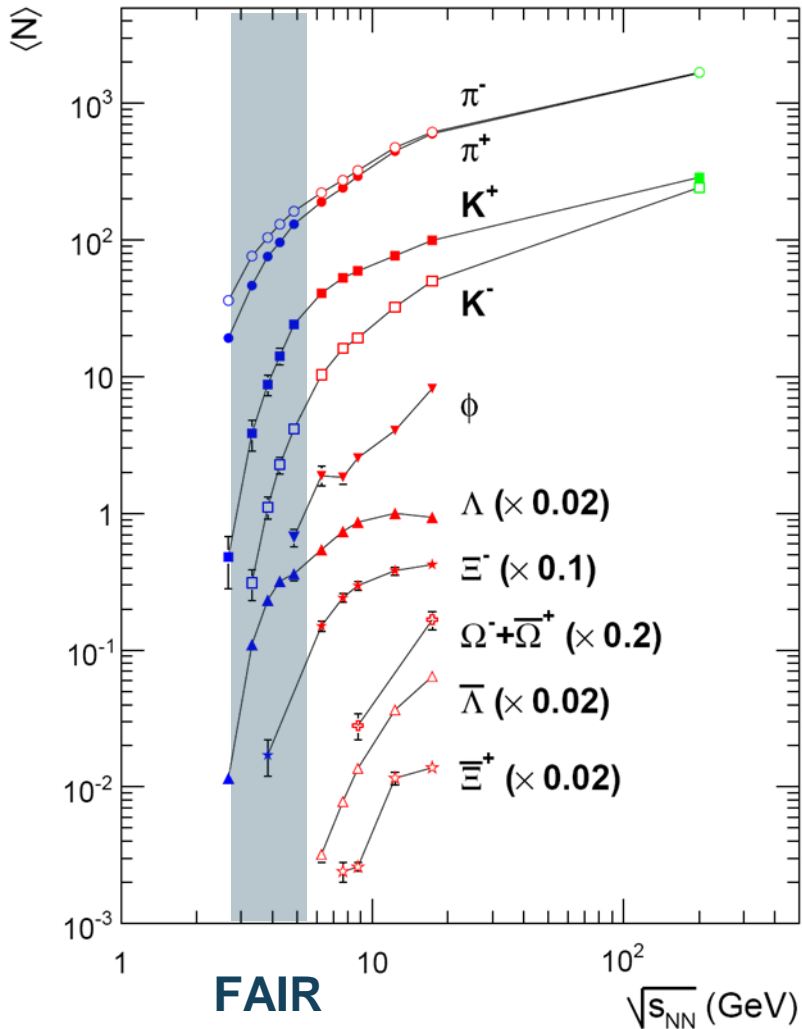
Yu.A. Murin^{†1}, K.K. Gudima^{1,2}, E.I. Litvinenko¹, V.A. Vasendina¹, and A.I. Zinchenko¹

¹Joint Institute for Nuclear Research (JINR), Dubna, Russia; ²Institute of Applied Physics of Academy of Sciences of Moldova, Chisinau, Moldova

Strangeness

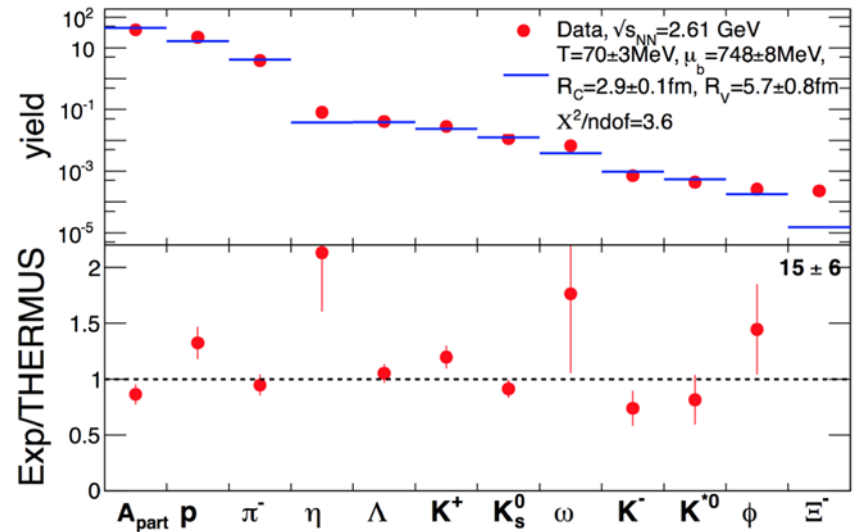
Data situation

Pb+Pb, Au+Au (central)



HADES: Ar + KCl 1.76 A GeV

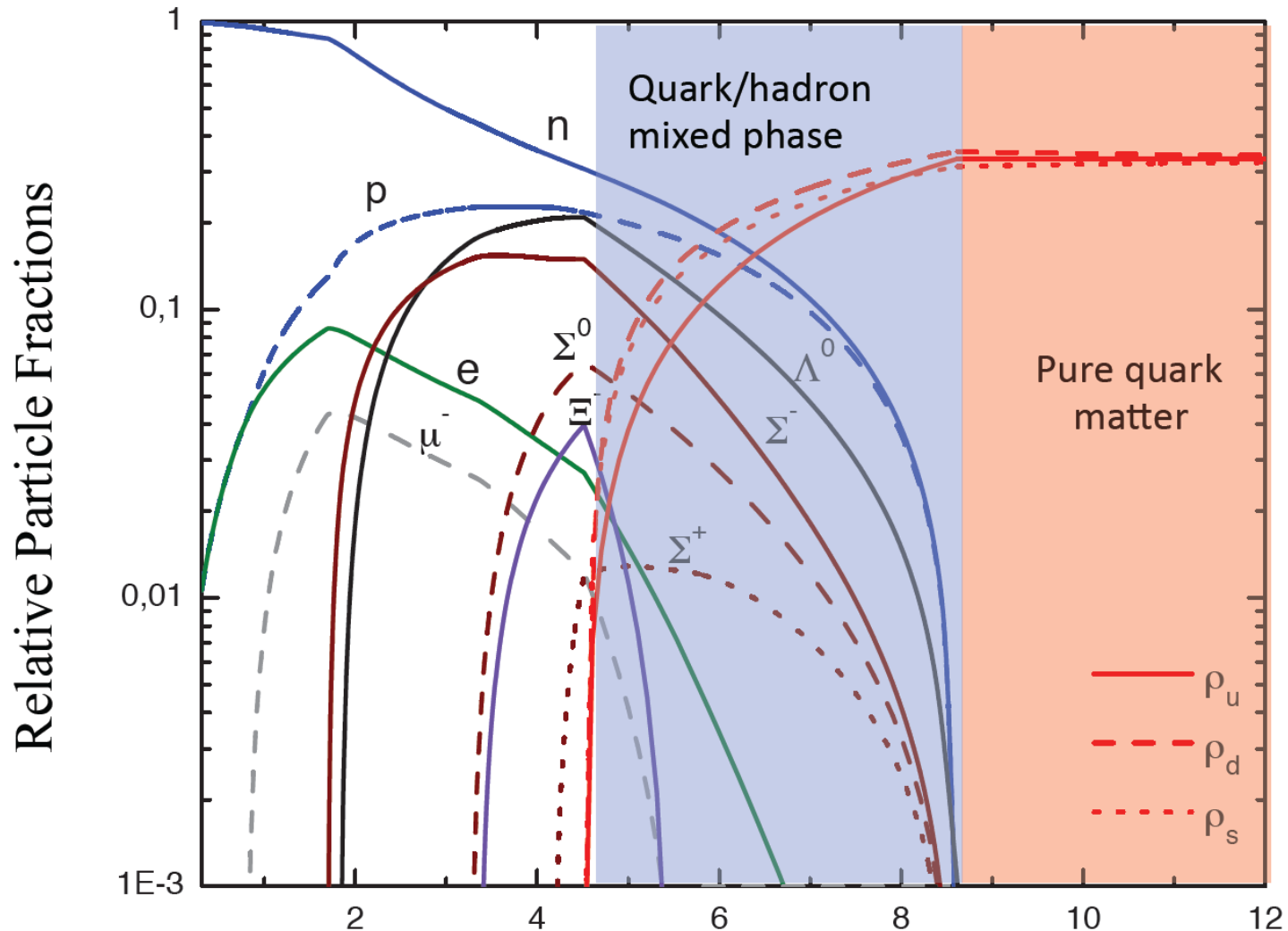
G. Agakishiev et al., arXiv:1512.07070



Degrees of freedom in massive neutron stars?

Equation-of-state: Non-local SU(3) NJL with vector coupling

M. Orsaria, H. Rodrigues, F. Weber, G.A. Contrera, arXiv:1308.1657



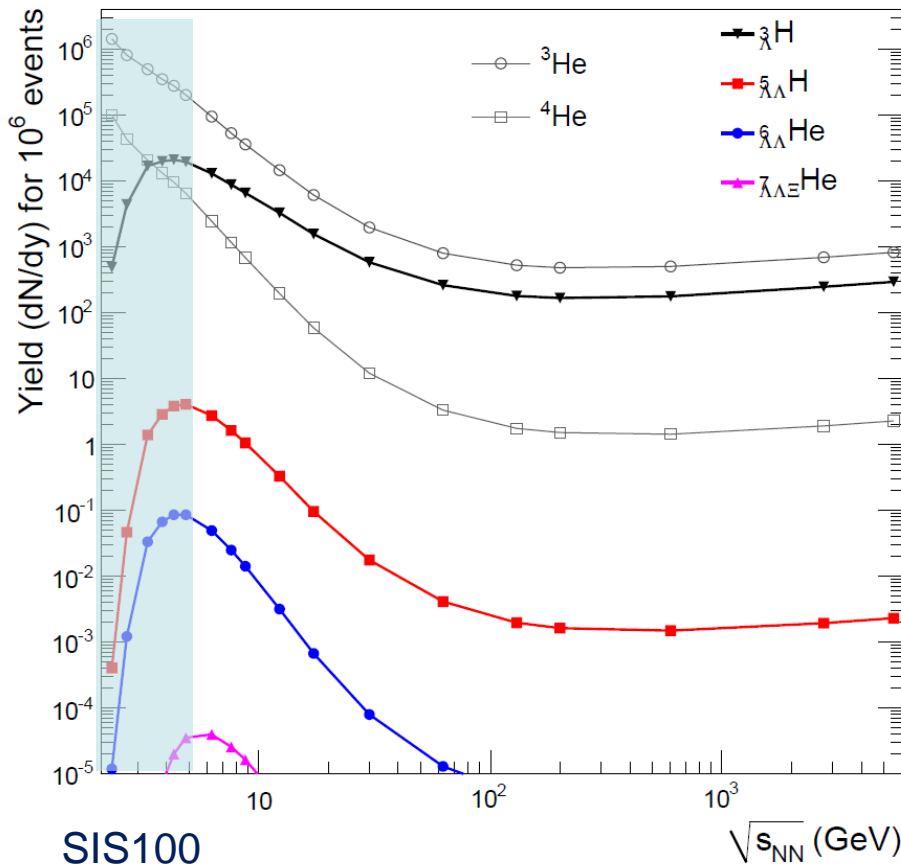
Hyperon puzzle
in neutron stars



N- Λ , Λ - Λ interaction ?

Hyperon production in A+A collisions

N- Λ , Λ - Λ interaction \blacktriangleright (double-) lambda hypernuclei



SIS100

$\sqrt{s_{NN}}$ (GeV)

Double lambda hypernuclei production in central Au+Au collisions at 10 A GeV:

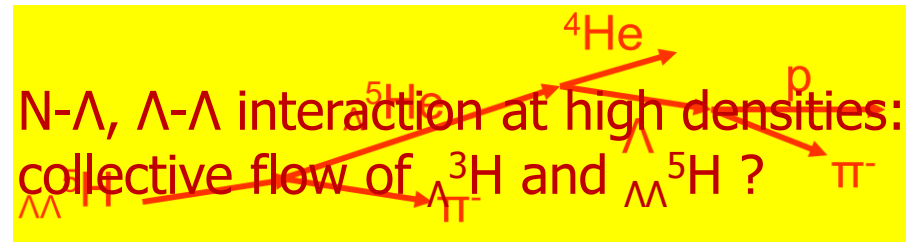
	Multiplicity	Yield in 1 week
$\Lambda\Lambda^5\text{H}$	$5 \cdot 10^{-6}$	3000
$\Lambda\Lambda^6\text{He}$	$1 \cdot 10^{-7}$	60

Assumption for yield calculation:

Reaction Rate 1 MHz

BR 10% (2 sequential weak decays)

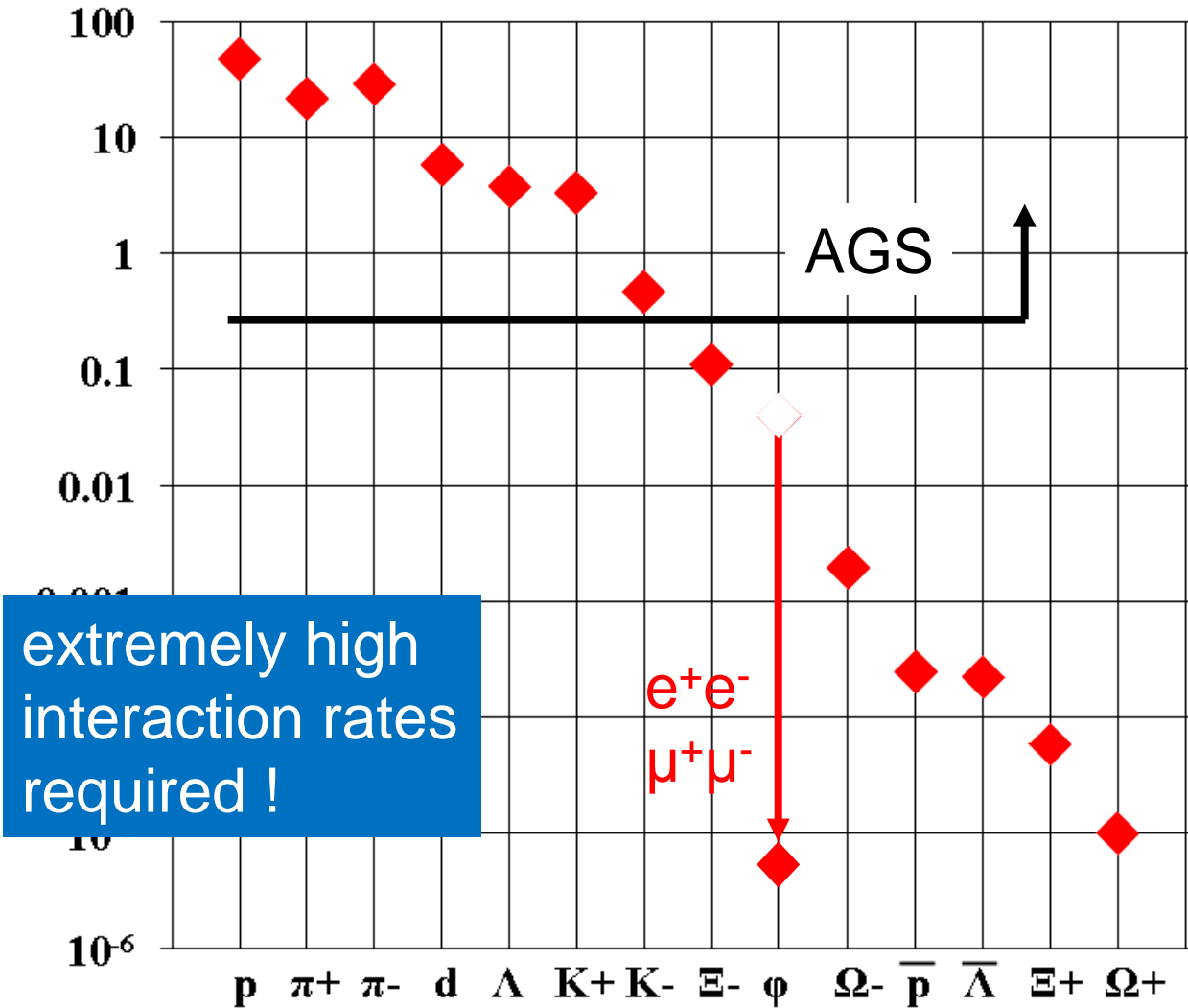
Efficiency 1%



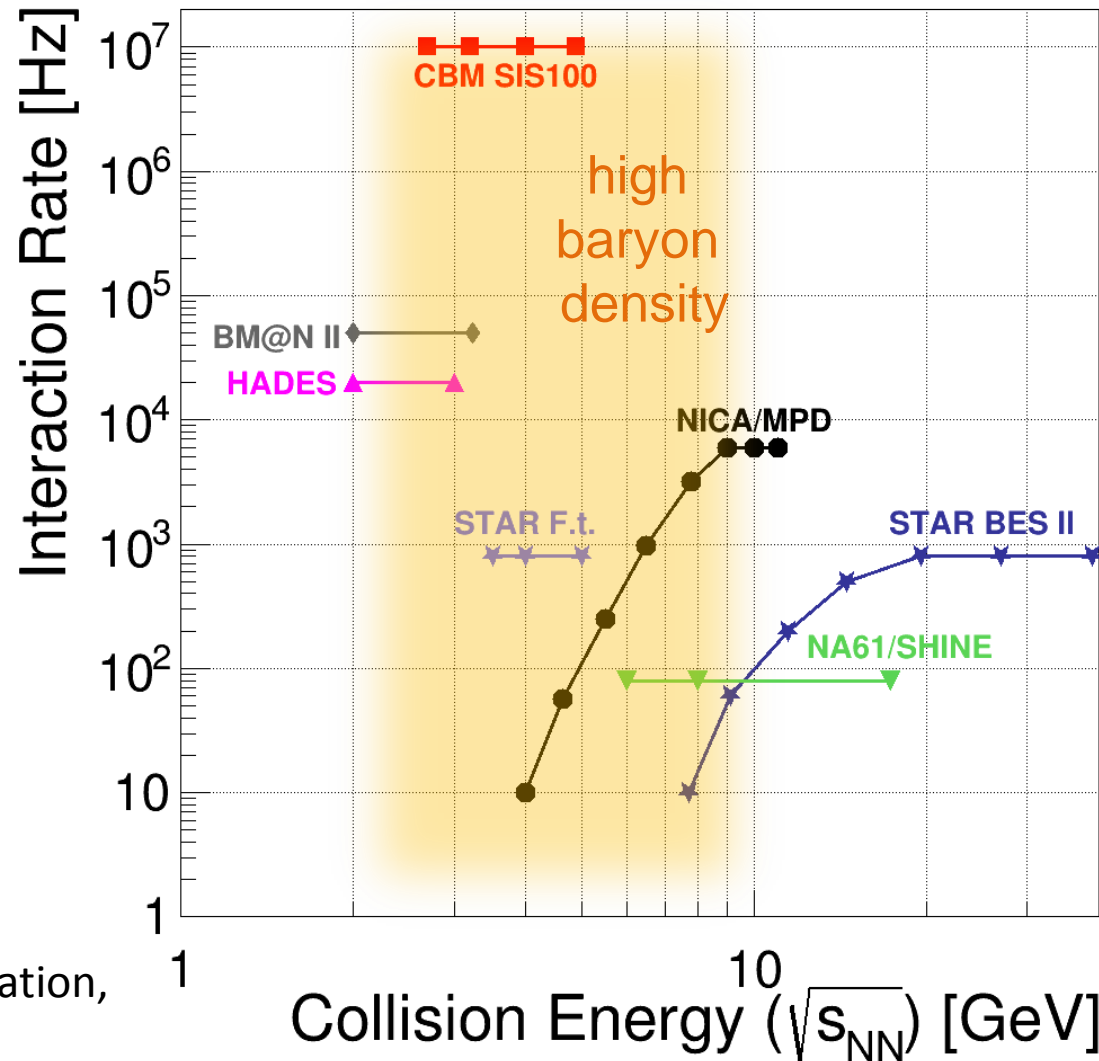
Experimental challenges

Particle yields in central Au+Au 4 A GeV

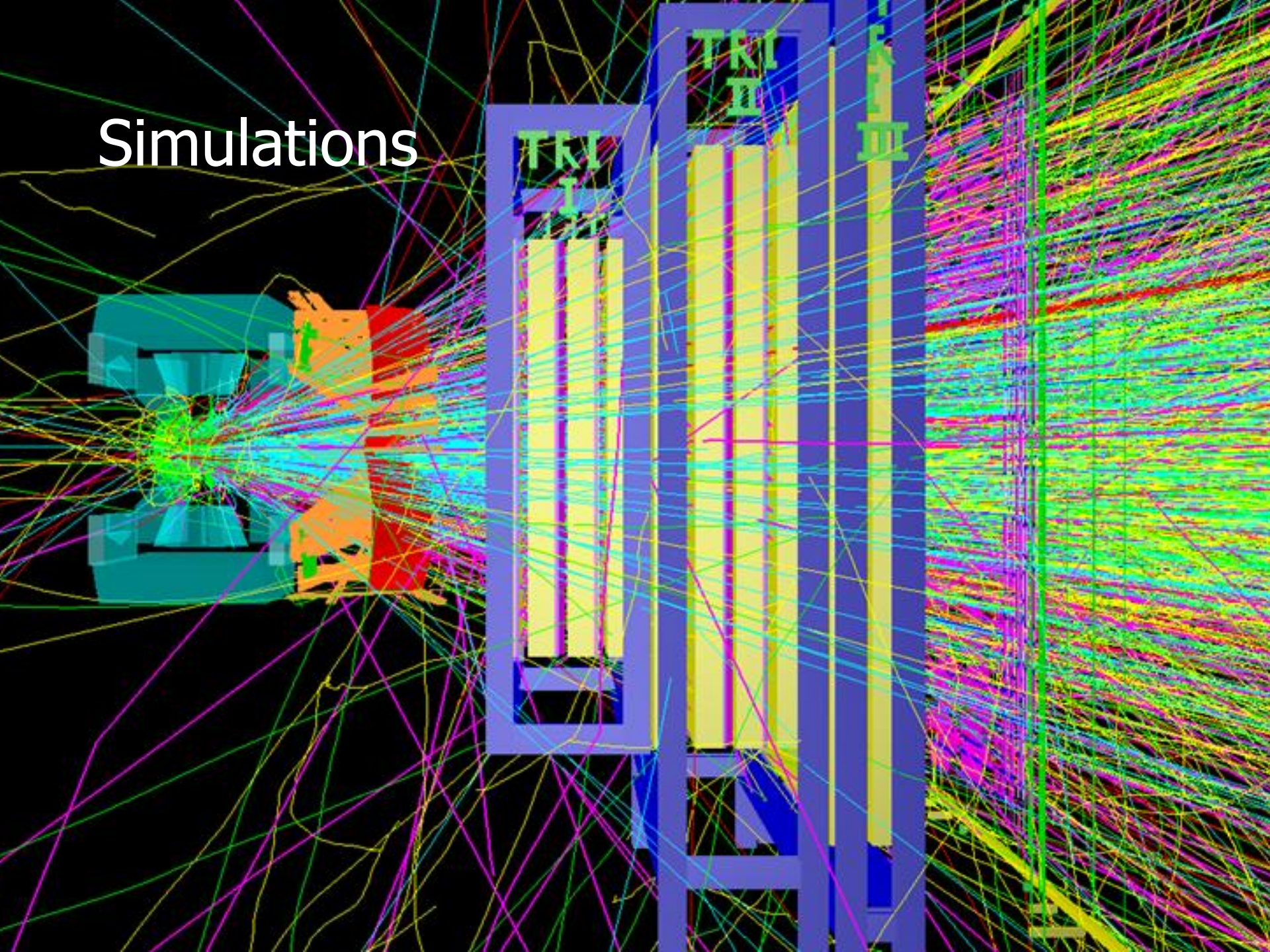
Multiplicity Statistical model, A. Andronic, priv. com.



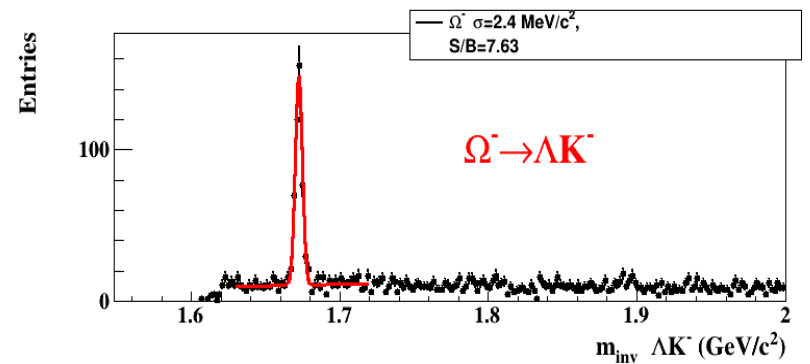
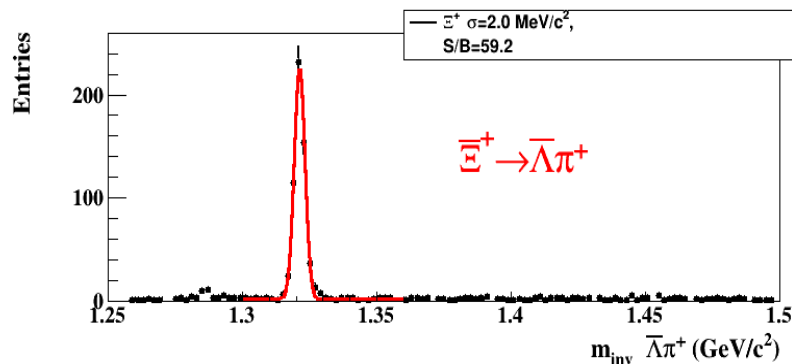
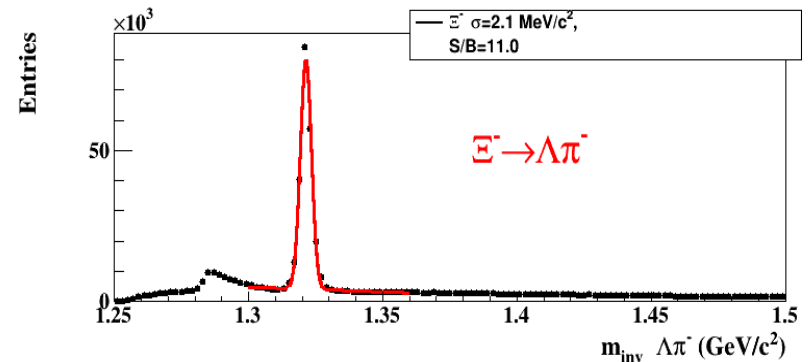
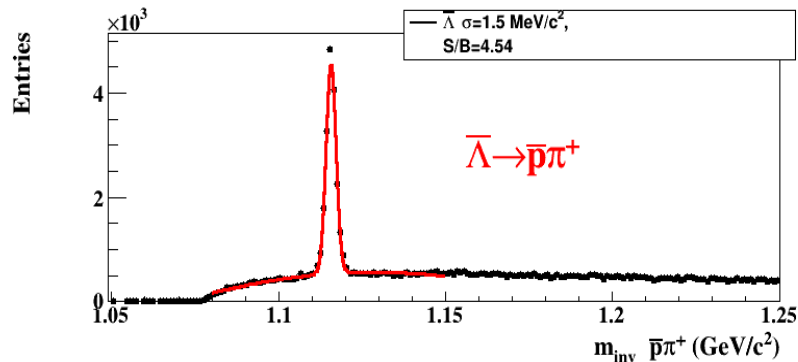
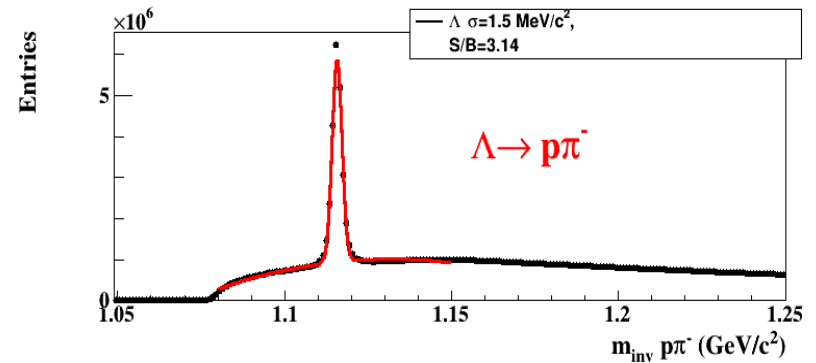
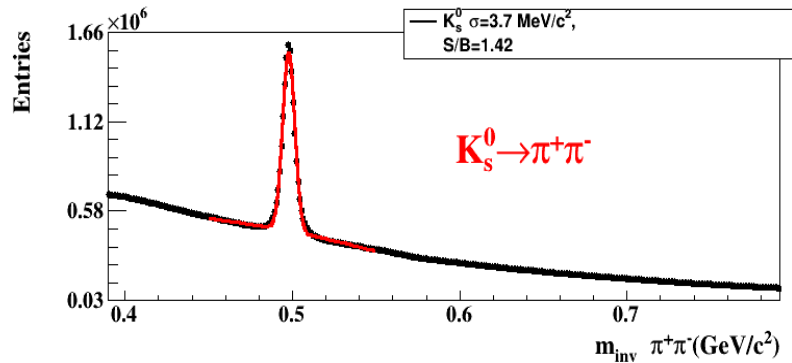
The landscape of experiments exploring dense QCD matter



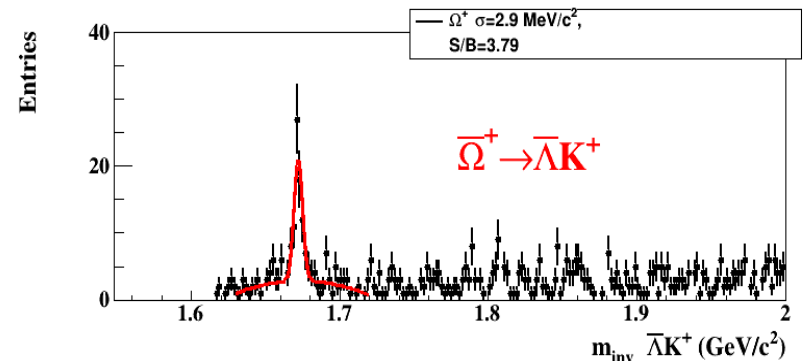
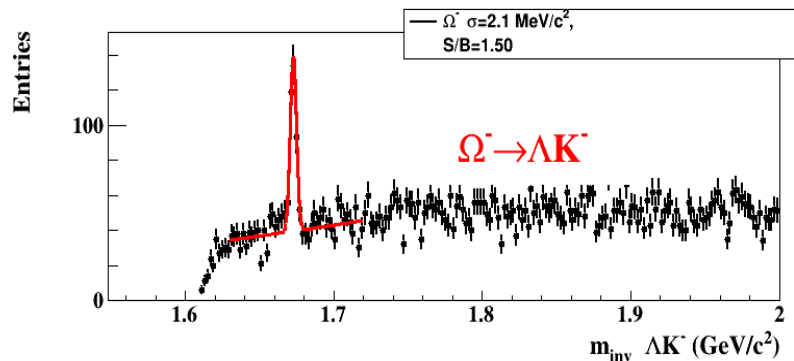
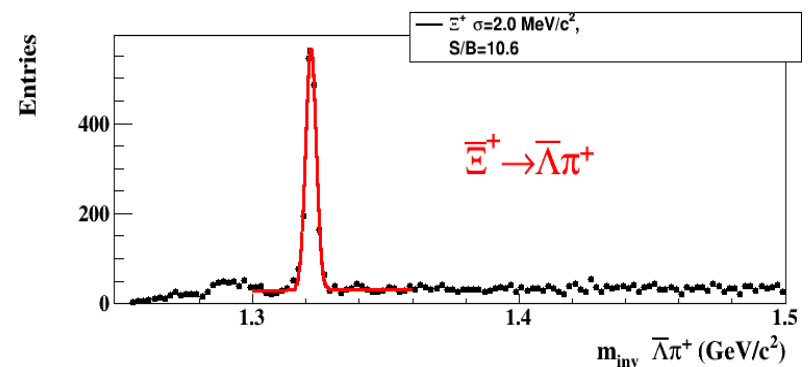
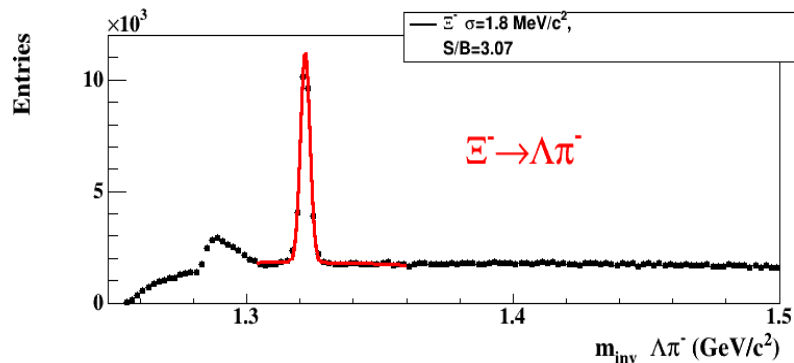
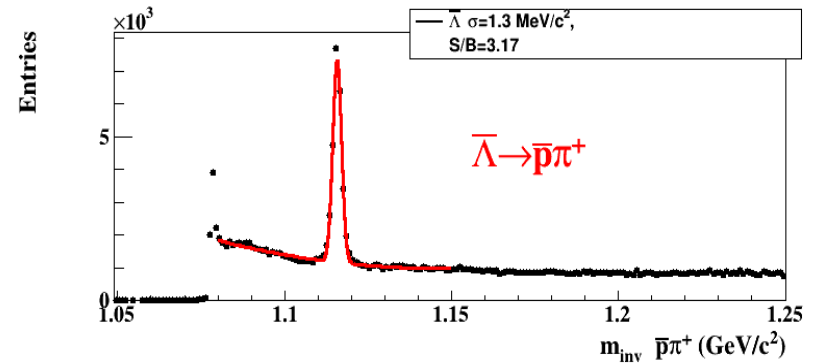
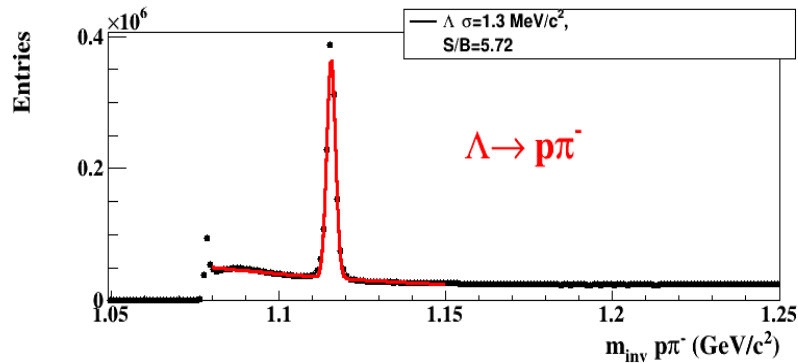
Simulations



Simulations using CBM KF Particle Finder 5M central AuAu collisions 10A GeV (Lab)



Real data analyzed with CBM KF Particle Finder STAR 4.4 M mbias AuAu collisions, $\sqrt{s_{NN}}=7.7$ GeV



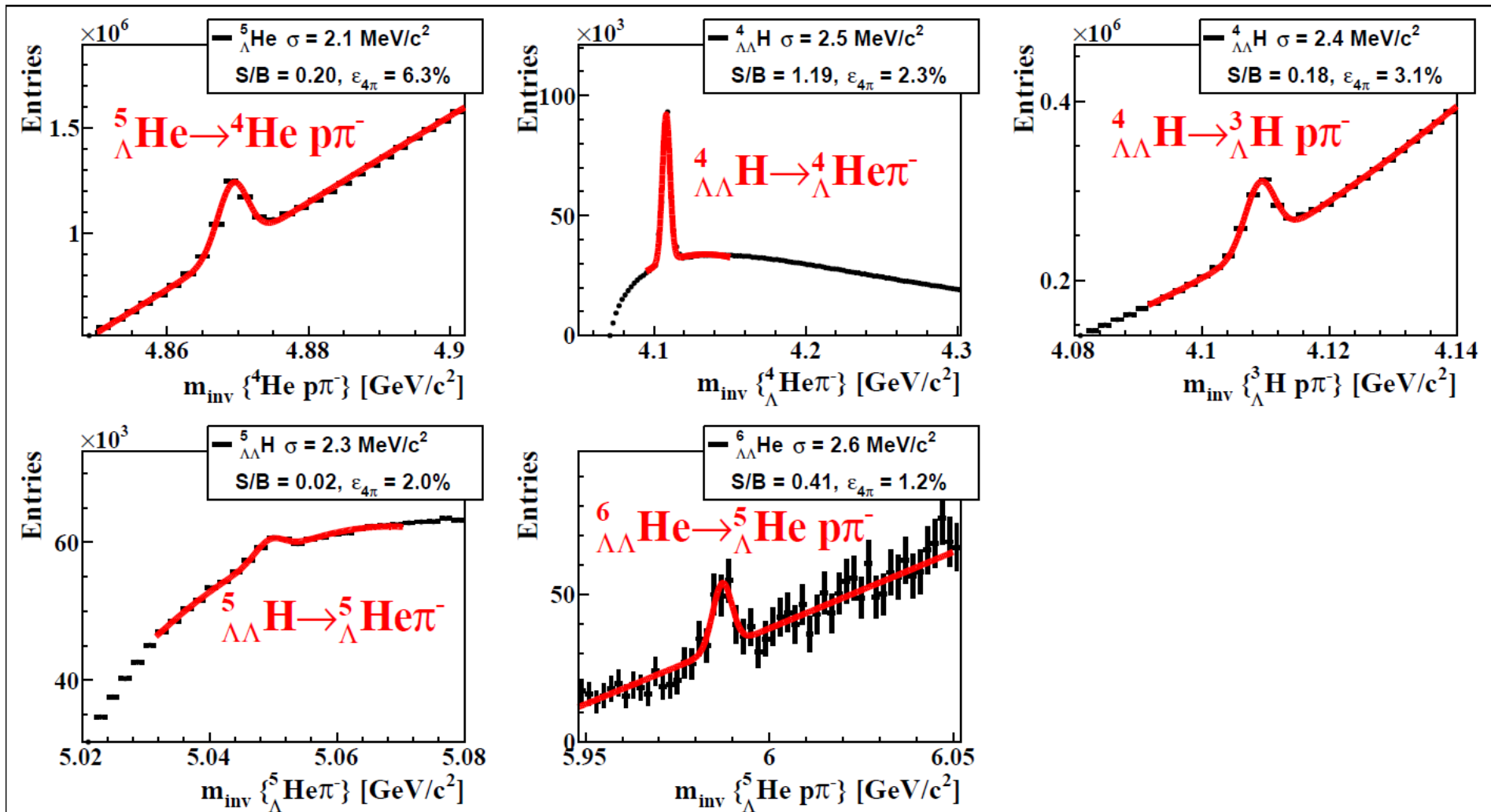
Hyperon yields at CBM

4 A GeV min. bias Au+Au collisions,
multiplicities from statistical model,
Reaction rate $10^6/s$

Particle	$E_{\text{thr}} \text{ NN}$ GeV	Multiplicity min.bias	ϵ %	Yield/week min. bias
Ξ^-	3.7	$2.5 \cdot 10^{-2}$	3	$4.5 \cdot 10^8$
Ω^-	6.9	$5 \cdot 10^{-4}$	3	$9 \cdot 10^6$
Anti- Λ	7.1	$5 \cdot 10^{-5}$	15	$9 \cdot 10^5$
Ξ^+	9.0	$1.5 \cdot 10^{-5}$	3	$2.7 \cdot 10^5$
Ω^+	12.7	$2.5 \cdot 10^{-6}$	3	$4.5 \cdot 10^4$

Simulation and reconstruction

Hypernuclei in central Au+Au 10 AGeV



Conclusions

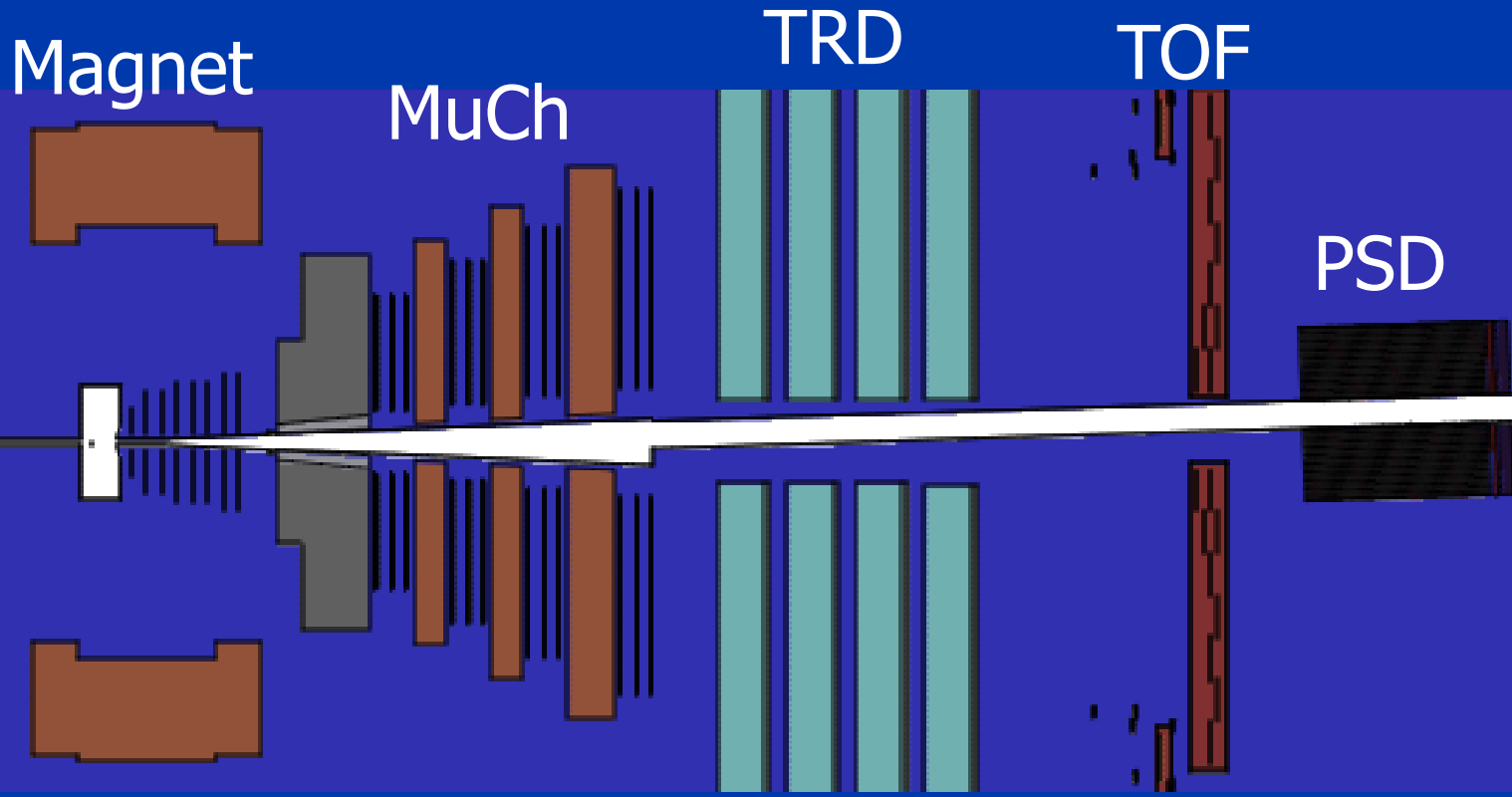
The CBM physics program includes:

- The exploration of the high density nuclear matter EOS by measuring excitation functions of:
 - the collective flow of identified particles
 - multi-strange hyperon production

- The study of $\Lambda\Lambda$ and ΛN interactions by measuring yields, livetime, and collective flow of (double) Λ hypernuclei

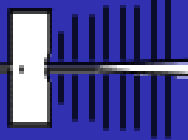
The goal is to produce multi-differential data with unprecedented statistics in heavy-ion collisions at beam energies from 2 - 14 A GeV (Au beam up to 11 A GeV)

Measurement of E_{sym} with CBM ?



Measurement of E_{sym} with CBM ?

Magnet



LAND

TOF

PSD

LAND

