# **Dilepton Results from STAR BES-I**

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## Outline

Introduction and motivation

- Dielectron production from RHIC BES-I
- ♦ Future prospects
- ♦ Summary

#### Introduction: EM penetrating probes



Do not participate in strong interactions. Bring undistorted information as where produced. Penetrate medium properties.

**Challenge**: Time-space integrated from every stages. Continuum at IMR.

Drell-Yan

IMR

**HMR** 

Heavy quark correlation

Heavy quarkonia production

QGP thermal radiation

#### Introduction: in-medium modifications



Low mass excess was observed in previous experiments.

Vacuum  $\rho$  unable to describe data.

Rule out Dropping-Mass Scenario (Brown-Rho).

Good agreement with broadening of  $\rho$  spectral function (Rapp-Wambach).

### Introduction: hadronic vs partonic





LMR: inversed slopes show mass dependence
-- hadronic process dominate, radial flow
IMR: no indication of mass dependence
-- thermal radiation from partonic phase
Energy dependence of the slopes could be sensitive to the medium dynamics.

### Dilepton at RHIC top energy



*LM* excess observed consistent with  $\rho$  in-medium modification - possible link to chiral symmetry restoration.

11/04/18

Yifei Zhang / USTC

### What about low energies?



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### **RHIC BES program**



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### **Dilepton excess spectra in BES-I**



### **Dilepton excess spectra in BES-I**



Consistent with  $\rho$  in-medium modification.

### **Dilepton excess spectra in BES-I**





AuAu@19.6,200: STAR, PLB750 64 2015 AuAu@27,39,62&UU@193: S. Yang, Quark Matter 2015 InIn@17.3: NA60, EPJ C59 607 2009 Theory: R. Rapp, PRC 63 (2001) 054907

Consistent with ρ in-medium modification. Weak collision energy dependence => Leptons are blindly emitted in HG + QGP.

#### Excess yield and medium life time



### IMR signal / background



### $M_{T}$ slope at IMR



### $M_T$ slope at IMR



#### Possible observation at phase transition?



Both  $T_{eff}$  and its slope  $\kappa$  in medium are significant higher than the system w/o medium.

Phase transition could happen if the  $T_{eff}$  increases dramatically or the sign of its slope  $\kappa$  changes from negative to positive.

### Projection of Run 18&19



- With 2-3 weeks data taken, the low mass statistics will be significantly improved.
- > Possible access for QGP radiation component at IMR.

### **BES Phase II**

- ♦ Electron cooling will provide increased luminosity.
- $\Rightarrow$  **iTPC** + HFT + MTD upgrades
- $\diamond$  Enables increased statistics for the BES energies
- ♦ Statistics enriched data for rare probes, especially for dilepton measurements.



#### Proposed energies for BES-II (Years 2019-2020):

√S <sub>NN</sub> (GeV)	7.7	9.1	11.5	14.5	19.6
$\mu_{B}$ (MeV)	420	370	315	250	205
BES II (MEvts)	100	160	230	300	400

### Projection with iTPC for BES-II



- Systematically study dielectron continuum from 7.7 19.6 GeV.
- Inner Time Projection Chamber (iTPC) upgrade: reduce uncertainties.
- Quantify different models.
- Study total baryon density effect at lower energies from BES-II.

### **Opportunity at BES-II**



Possible chance for QGP radiation at IMR.

Phase transition, QGP turn-off signature, baryon density dependence.

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### Excess of dielectron at very low $p_T$ at RHIC STAR



Significant excess with respect to hadronic cocktail in peripheral Au + Au and U + U collisions!

Excess observed over the whole measured mass region!

#### Excess of dielectron at NICA & CBM



Considerable production rate and softer mass spectrum at NICA & CBM

### Summary

♦ Dielectron mass spectra in 19.6 - 200 GeV Au+Au were measured by RHIC-STAR.

Low mass enhancement was observed and can be well described by model calculations with broadening ρ mass spectra function for all collision energies and systems at RHIC and SPS.

♦ The normalized excess yield is proportional to the medium life time from 17.3 to 200 GeV Au+Au collisions and 193 GeV U+U collisions.

♦ STAR future Runs and upgrades enable further exploration of the dilepton continuum.

- -- QGP thermal radiation
- -- Correlated charm modifications

Thank you!