Status and expected progress of global fits

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Warwick, "Towards the Ultimate Precision in Flavour Physics"

Outline

- Introduction
- Observables in rare B-decays
- Global analyses
- Conclusion

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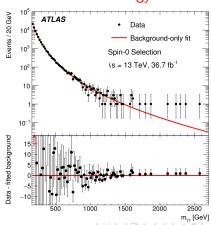
Towards the Ultimate Precision in Flavour Physics

- Precision (exp. & theo.): apart from improving SM, look for NP
- This decade has not been dominated by NP discoveries @ LHC, thus reinforcing the need to scrutinize low-energy obs. for NP

No significant diphoton excess

[ATLAS: Phys. Lett. B 775 (2017) 105-125]

[CMS: Phys. Lett. B 767 (2017) 147]



Why rare decays

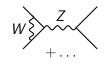
- In the SM, first at one loop (GIM mechanism)
- Privileged way to look for NP manifestations, and shape its structure



etc., not possible at tree



e.g., $K \to \pi \nu \bar{\nu}$ CKM suppressed



e.g., $B_{(s)} \rightarrow \ell^+ \ell^$ helicity suppressed

Less ideal: $s \to d\ell^+\ell^-$ and $c \to u\ell^+\ell^-$ transitions, etc.

Going beyond the SM

- Striking evidences of NP (requiring "less" theo. precision): e.g., LFN violating processes: $\mu \to e \gamma$, etc.
- Emerging deviations

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- Striking evidences of NP (requiring "less" theo. precision): e.g., LFN violating processes: $\mu \to e\gamma$, etc.
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HERE: rare *B*-decays, i.e., $b \rightarrow s(\gamma, \ell^+\ell^-)$ transitions

experimental side: [talks by Thomas Blake, Rafael S. Coutinho, Giampiero Mancinelli]

theoretical aspects: [this talk, and talks by Joaquim Matias, Marco Ciuchini, Patrick H. Owen]

pheno consequences: [talks by Dario Buttazzo, Gudrun Hiller, Diego Guadagnoli]



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$$\begin{split} \langle \big(\gamma^{(*)},\ell^+\ell^-\big) M_{(\lambda)}|\mathcal{H}_{\mathrm{eff}}|\bar{B}\rangle &= \tfrac{G_F\alpha}{\sqrt{2}\pi} V_{tb} V_{ts}^* \big[\big(A_{\nu} + \tfrac{T_{\nu}}{\nu} \big) \bar{u}_{\ell} \gamma^{\nu} v_{\ell} + \tfrac{B_{\nu}}{\nu} \bar{u}_{\ell} \gamma^{\nu} \gamma_5 v_{\ell} \big], \\ M_{(\lambda)} &= \bar{K}, \bar{K}_{\lambda}^*, \ldots, \ q_{\ell\ell}^2 \equiv q^2 \end{split}$$

Short-distances (SD) above $\simeq m_b$, Long-distances (LD) below $\simeq m_b$

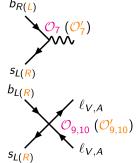
$$\begin{split} \langle (\gamma^{(*)},\ell^+\ell^-) M_{(\lambda)} | \mathcal{H}_{\mathrm{eff}} | \bar{B} \rangle &= \tfrac{G_F \alpha}{\sqrt{2}\pi} V_{tb} V_{ts}^* [(A_{\nu} + \tfrac{T_{\nu}}{\nu}) \bar{u}_{\ell} \gamma^{\nu} v_{\ell} + \tfrac{B_{\nu}}{\nu} \bar{u}_{\ell} \gamma^{\nu} \gamma_5 v_{\ell}], \\ M_{(\lambda)} &= \bar{K}, \bar{K}_{\lambda}^*, \ldots, \ q_{\ell\ell}^2 \equiv q^2 \end{split}$$

Short-distances (SD) above $\simeq m_b$, Long-distances (LD) below $\simeq m_b$

$$\begin{split} A_{\nu} &= -\frac{2m_{b}q^{\rho}}{q^{2}} \, C_{7}^{\mathrm{eff}} \langle M_{(\lambda)} | \bar{s} \sigma_{\nu\rho} P_{R} b | \bar{B} \rangle + C_{9} \langle M_{(\lambda)} | \bar{s} \gamma_{\nu} P_{L} b | \bar{B} \rangle \\ B_{\nu} &= C_{10} \langle M_{(\lambda)} | \bar{s} \gamma_{\nu} P_{L} b | \bar{B} \rangle \end{split}$$

Wilson coefficients: C_7 , C_9 , C_{10} , ... known up to NNLO-QCD

[Huber+'05, Gambino+'03, Gorbahn+'04, Bobeth+'03, Misiak+'06]



$$\langle (\gamma^{(*)}, \ell^+ \ell^-) M_{(\lambda)} | \mathcal{H}_{\mathrm{eff}} | \bar{\mathcal{B}} \rangle = \frac{G_F \alpha}{\sqrt{2} \pi} V_{tb} V_{ts}^* [(A_{\nu} + \frac{T_{\nu}}{\nu}) \bar{u}_{\ell} \gamma^{\nu} v_{\ell} + \frac{B_{\nu}}{\nu} \bar{u}_{\ell} \gamma^{\nu} \gamma_5 v_{\ell}],$$

$$M_{(\lambda)} = \bar{K}, \bar{K}_{\lambda}^*, \dots, \ q_{\ell\ell}^2 \equiv q^2$$

Short-distances (SD) above $\simeq m_b$, Long-distances (LD) below $\simeq m_b$

Non-trivial test of the SM and a comprehensive look into NP

[Altmannshofer, Bobeth, Gambino, Gorbahn, Haisch, Hiller, Huber, Lunghi, Matias, Misiak, Steinhauser, Straub, Virto, ...]

$$\begin{split} \langle (\gamma^{(*)},\ell^+\ell^-) M_{(\lambda)} | \mathcal{H}_{\mathrm{eff}} | \bar{\mathcal{B}} \rangle &= \tfrac{G_F \alpha}{\sqrt{2}\pi} V_{tb} V_{ts}^* [(A_{\nu} + \tfrac{T_{\nu}}{\nu}) \bar{u}_{\ell} \gamma^{\nu} v_{\ell} + \tfrac{B_{\nu}}{\nu} \bar{u}_{\ell} \gamma^{\nu} \gamma_5 v_{\ell}], \\ M_{(\lambda)} &= \bar{K}, \bar{K}_{\lambda}^*, \ldots, \ q_{\ell\ell}^2 \equiv q^2 \end{split}$$

Short-distances (SD) above $\simeq m_b$, Long-distances (LD) below $\simeq m_b$

LD: includes non-local objects

$$T^{\nu} \propto \frac{1}{q^2} \int d^4x \, \mathrm{e}^{iqx} \, \langle M_{(\lambda)} | \mathrm{T}\{j_{\mathrm{e.m.}}^{\nu}(x) \mathcal{H}_{\mathrm{eff}}^{\mathrm{had}}(0)\} | \bar{B} \rangle$$
 helicity (λ) and process dependent cancels the μ dependence of $C_7^{\mathrm{eff}}(\mu)$, $C_9(\mu)$



[Beneke+'01,'04]

[talks by Quim, Marco and Patrick]

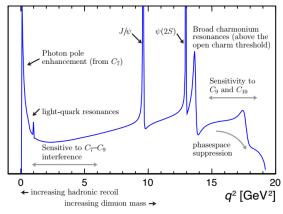
$$C_{9,a}^{tot}(q^2) = C_9^{SM,SD}(q^2) + C_{9,\ell}^{NP} + C_{9,a}^{SM,c\bar{c}}(q^2) \,, \; \ell = e, \mu, \tau \,, \; a = \bot, \parallel, 0$$
 one must not interpret $C_{9,a}^{SM,c\bar{c}}(q^2)$ (i.e., \mathcal{T}^{ν}) as a fake **LFU** $C_{9,\ell}^{NP}$

Theoretical inputs

- Look for sensitivity to SD: avoid J/ψ , $\psi(2S)$ resonances
- High- and low- q^2 rely on different formalisms/techniques for assessing LD effects (QCD factorization, lattice QCD, etc.)

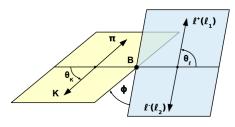
$$Br(B o K^*\ell^+\ell^-)$$
 vs. $q_{\ell\ell}^2\equiv q^2$

[Charles+'99, Beneke+'00, Grinstein+'04, Beylich+'08, Lyon+'14, LHCb '16, Bobeth+'17, Blake+'17]



Observables in $B \to V \ell^+ \ell^-$ decays

 $B \to K^*[\to K\pi]\ell^+\ell^-$ decays: q^2 dependent angular observables



$$\frac{d^{(4)}\Gamma}{da^2d(\cos\theta_{\ell})d(\cos\theta_{\kappa})d\phi} = \frac{9}{32\pi} \times$$

$$\begin{split} \left(\underbrace{I_1^s \sin^2 \theta_K + I_1^c \cos^2 \theta_K + \left(I_2^s \sin^2 \theta_K + I_2^c \cos^2 \theta_K\right) \cos 2\theta_\ell + I_3 \sin^2 \theta_K \sin^2 \theta_\ell \cos 2\phi_\ell + I_4 \sin 2\theta_K \sin 2\theta_\ell \cos \phi + I_5 \sin 2\theta_K \sin 2\theta_\ell \cos \phi + \left(I_6^s \sin^2 \theta_K + I_6^c \cos^2 \theta_K\right) \cos \theta_\ell + I_7 \sin 2\theta_K \sin \theta_\ell \sin \phi + I_8 \sin 2\theta_K \sin 2\theta_\ell \sin \phi + I_9 \sin^2 \theta_K \sin^2 \theta_\ell \sin 2\phi_\ell \right) \end{split}$$

Angular observables

I: functions of the helicity amp. $H^{\lambda}_{V,A}$, $\lambda=0,\pm$ (below, $m_{\ell}\to 0$)

$$\begin{aligned} \frac{I_2^c}{I_2^s} &= -\frac{F}{2} (|H_V^0|^2 + |H_A^0|^2) \\ \frac{I_2^s}{I_2^s} &= \frac{F}{8} (|H_V^+|^2 + |H_V^-|^2) + (V \to A) \end{aligned}$$

"longitudinal" rate "transverse" rate F_L , Br

$$I_6^s = F \text{Re}[H_V^- (H_A^-)^* - H_V^+ (H_A^+)^*]$$

lepton FB asym.

 A_{FB} or P_2

$$\frac{I_4}{I_5} = \frac{F}{4} \text{Re}[(H_V^- + H_V^+)(H_V^0)^*] + (V \to A)
I_5 = \frac{F}{2} \text{Re}[(H_V^- - H_V^+)(H_A^0)^*] + (V \leftrightarrow A)$$

 $P'_{4,5}$

[Descotes+'12,'13]

$$\begin{aligned} I_3 &= -\frac{F}{2} \text{Re}[H_V^+(H_V^-)^*] + (V \to A) \\ I_9 &= \frac{F}{2} \text{Im}[H_V^+(H_V^-)^*] + (V \to A) \end{aligned}$$

"wrong-helicity" (suppressed in SM)

$$(F = rac{\sqrt{\lambda}q^2}{3 imes 2^5\pi^3m_P^3}BF(K^* o K\pi))$$

Some representative observables

- HQE @ high-recoil: $\{V, A_{0,1,2}, T_{1,2,3}\}$ FFs $\rightarrow \{\xi_{\perp}, \xi_{\parallel}\}$ FFs
- Improved sensitivity on FFs

$$\mathcal{B}r(B \to K^*\mu^+\mu^-) = \mathcal{O}(\xi_{\perp}^2, \xi_{\parallel}^2)$$

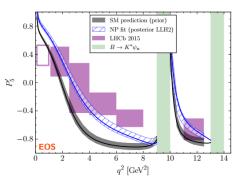
$$F_L, S_i = \mathcal{O}(\xi_{\perp}^2/\xi_{\parallel}^2)$$

$$P_5' \equiv \frac{I_5}{\sqrt{-I_5^s I_5^c}}$$

$$P_5' = P_5'^{\infty} (1 + \mathcal{O}(\alpha_s \xi_{\perp})) + \text{p.c.}$$

Further corrections in α_s and Λ_{QCD}/m_b (= p.c.)

[Jäger+'12.'14. Hurth+'15.'17. Capdevila+'17]



Analyticity of $h_{\lambda}(\propto T \cdot \varepsilon)$ Theo. model (LCSR) $q^2 \lesssim 0$ Exp. data below J/ψ

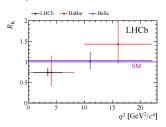
[Bobeth+'17; Blake+'17] [talk by Patrick]

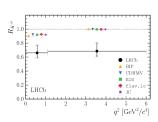
LFU violation testing observables

ullet SM: universal gauge couplings, small lepton masses (w.r.t. $\sqrt{q^2}$)

$$R_{K^{(*)}}[q_1^2,q_2^2] = rac{\int_{q_1^2}^{q_2^2} dq^2 Br(B o K^{(*)}\mu^+\mu^-)}{\int_{q_1^2}^{q_2^2} dq^2 Br(B o K^{(*)}e^+e^-)} \stackrel{SM}{\simeq} 1$$

- In the SM, $R_{K^{(*)}}$ largely independent of unc. (for large q^2)
- (tiny) Correction induced by final-state photon radiation





Another promising observable $Q_i \equiv P'^{\mu}_i - P'^{e}_i$, i = 4,5

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Testing the SM and looking for NP

 $\mathcal{H}_{\mathrm{eff}}$: shifts to C_7^ℓ , C_9^ℓ , C_{10}^ℓ , $C_7^{\prime\ell}$, $C_9^{\prime\ell}$, $C_{10}^{\prime\ell}$, etc.

(some) Available observables (LHCb, Belle, ATLAS, CMS)

- $B \to K^* \mu \mu$ ($P_{1,2}$, $P'_{4,5,6,8}$, F_L , Br), also R_{K^*} , $Q_{4,5}$
- $B_s \rightarrow \phi \mu \mu \ (P_1, P'_{4.6}, F_L)$
- $B^+ \to K^+ \mu \mu$, $B^0 \to K^0 \ell \ell$ (Br), $\ell = e, \mu$: R_K
- $B \rightarrow X_s \gamma$, $B \rightarrow X_s \mu \mu$ (Br); $B_s \rightarrow \mu \mu$ (Br)
- $B^0 \to K^{*0} \gamma$ (A_I , $S_{K^* \gamma}$), $B^+ \to K^{*+} \gamma$, $B_s \to \phi \gamma$

Different groups

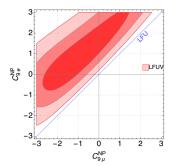
- ≠ statistical approaches (frequentist, Bayesian, etc.),
- \neq angular observables (e.g., P_i vs. S_i),

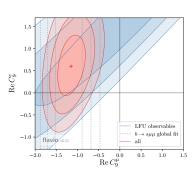
- [cf. Hurth+'17]
- form factor inputs ([LCSR, low-q²: Khodjamirian+'10] VS. [fit LCSR + lattice: Bharucha+'16]),
- \neq treatment of had. effects ([Khodjamirian+'10], $h_{\lambda}(q^2) \simeq h_{\lambda}^{(0)} + \frac{q^2}{(1~{
 m GeV})^2} h_{\lambda}^{(1)} + \ldots$)

LFU violating data

- ullet $R_{\mathcal{K}^{(*)}}^{exp}$ substantially below $R_{\mathcal{K}^{(*)}}^{\mathrm{SM}} \simeq 1$
- Extend the SM and fit for $\delta C_9^e \neq \delta C_9^\mu$
- Tension with SM/LFU picture $\delta C_9^e = \delta C_9^\mu$ of $\sim 3\sigma$

(hypothesis testing)

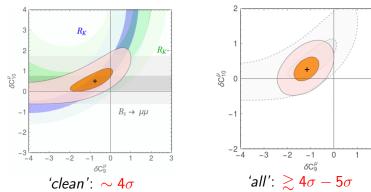




[Specific obs. in plots: Capdevila+'17, Altmannshofer+'17; see also Ciuchini+'17, D'Amico+'17, Geng+'17, Hiller+'17]

Full data set

- Add up $B_s \to \mu^+ \mu^-$ ('clean'); add up ang. obs. ('all')
- Test for NP effects in the μ sector, e.g., δC_9^μ vs. $\delta C_{10}^\mu \to \text{central values are} \sim 20\%$ of the SM $C_{9,10}$
- Similar numerics for 1D fits; adding $C_9^{'\mu}$, etc. also possible

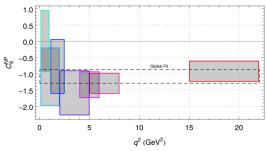


 $[Specific obs. \ in plots: \ Geng+'17; \ see \ also \ Altmannshofer+'17, \ Capdevila+'17, \ Ciuchini+'17, \ \underline{D}'Amico+'17]) \\ @> \bigcirc \\$

Correlation of LFU NP and hadronic effects

$$C_{9,a}^{tot}(q^2) = C_9^{SM,SD}(q^2) + C_9^{NP} + C_{9,a}^{SM,c\bar{c}}(q^2), \ a = \perp, \parallel, 0$$

- No clear indication of q^2 dependence, that would favor $C_{9,a}^{SM,c\bar{c}}$ thus allowing at the moment to interpret δC_9 as $C_9^{\rm NP}$
- ullet Debate still open, but $C_{9,a}^{SM,car c}$ cannot accommodate $R_{K^{(*)}}$



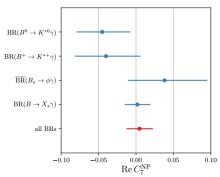
Only $b \to s\mu\mu$; [2, 5]: Pull_{SM} = 2.5, [4, 6]: Pull_{SM} = 3.1, [5, 8]: Pull_{SM} = 3.5, etc.

[talks by Quim and Marco] [Plot from: Descotes-Genon+'16; see also: Ciuchini+'15,'16, Chobanova+'17]

Radiative decays

- Dipole operator: cannot fit $R_{K^{(*)}}^{exp}$
- C_7 vs. C_7' fit: offers complementary picture of possible NP
- Also, four-quark operators $b \to sc\bar{c}$

[Jäger+'17]



Close horizon

LHCb and Belle II

year		2012	2020	2024	2030
LHCb	\mathcal{L} [fb ⁻¹] 3	8	22	50
	$\mathrm{n}(bar{b})$	0.3×10^{12}	1.1×10^{12}	37×10^{12}	87×10^{12}
	\sqrt{s}	$7/8\mathrm{TeV}$	$13\mathrm{TeV}$	$14\mathrm{TeV}$	$14\mathrm{TeV}$
Belle (II	\mathcal{L} [ab ⁻¹] 0.7	5	50	-
	$n(B\bar{B})$	0.1×10^{10}	0.54×10^{10}	5.4×10^{10}	-
	\sqrt{s}	$10.58\mathrm{GeV}$	$10.58\mathrm{GeV}$	$10.58\mathrm{GeV}$	-

 $[\mathsf{Albrecht} + '17 \; (\mathsf{and} \; \mathsf{refs.} \; \mathsf{therein})]$

- More data on the already measured channels
- New channels (with different backgrounds) &
 new observables ⇒ test the consistency of the LFUV picture

LHCb and Belle II - FCNC

[Albrecht+'17 (and refs. therein)]

Belle II

Observable	q^2 interval	Measurement	Extrapolations	
		$0.7 \mathrm{ab^{-1}}$	$5 {\rm ab}^{-1}$	$50 \rm ab^{-1}$
R(K)	$1.0 < q^2 < 6.0 \mathrm{GeV}^2$	-	11%	3.6%
R(K)	$q^2 > 14.4 \text{GeV}^2$	-	12%	3.6%
$R(K^*)$	$1.1 < q^2 < 6.0 \mathrm{GeV}^2$	-	10%	3.2%
$R(K^*)$	$q^2 > 14.4 \mathrm{GeV}^2$	-	9.2%	2.8%

LHCb

Observable	q^2 interval	Measurement	Extrapolations		
		3fb^{-1}	$8 {\rm fb^{-1}}$	$22 {\rm fb}^{-1}$	$50 {\rm fb^{-1}}$
$R(\phi)$	$1.0 < q^2 < 6.0 \mathrm{GeV}^2$	-	0.159	0.086	0.056
$R(\phi)$	$15.0 < q^2 < 19.0 \text{GeV}^2$	-	0.137	0.074	0.048
R(K)	$1.0 < q^2 < 6.0 \text{GeV}^2$	$0.745^{+0.090}_{-0.074} \pm 0.036$ [17]	0.046	0.025	0.016
R(K)	$15.0 < q^2 < 22.0 \text{GeV}^2$	-	0.043	0.023	0.015
$R(K^*)$	$0.045 < q^2 < 1.1 \text{GeV}^2$	$0.66^{+0.11}_{-0.07} \pm 0.03$ [18]	0.048	0.026	0.017
$R(K^*)$	$1.1 < q^2 < 6.0 \text{GeV}^2$	$0.69^{+0.11}_{-0.07} \pm 0.05$ [18]	0.053	0.028	0.019
$R(K^*)$	$15.0 < q^2 < 19.0 \text{GeV}^2$	-0.07	0.061	0.033	0.021

 \sim 2 % stat.

Combined Belle II and LHCb should be able to **establish** $\gg 5\sigma$ in $R_{K(*)}$

I HCb and Belle II - FCNC

- LFU Violating obs. $P_5^{\prime\mu} P_5^{\prime e}$ by LHCb, and Belle II
- $B_a \rightarrow \mu\mu$: e.g., discovery of $B_d \rightarrow \mu\mu$ by CMS (>2030)
- $b \to d\ell\ell$: e.g., $\frac{\mathcal{B}(B^+ \to \pi^+ \mu \mu)}{\mathcal{B}(B^+ \to \pi^+ ee)}$ by LHCb (300 fb⁻¹)
- $B_a^0 \overline{B}_a^0$ mixing

[talks by Zoltan, Vincenzo]

	Sensitivities of	modes	with $\nu \overline{\nu}$	in the	final st	$_{ m tate}$
Ob 1-1	D-11- 0	71 -1-	-1 D-11	- TT F .	₋ 1 ₋ -1	D -1

Observables	Belle 0.71 ab^{-1}	Belle II 5 ab^{-1}	Belle II 50 ab^{-1}
$\mathcal{B}(B^+ \to K^+ \nu \bar{\nu})$	< 450%	30%	11%
${\cal B}(B^0 o K^{*0} uar u)$	< 180%	26%	9.6%
$\mathcal{B}(B^+ \to K^{*+} \nu \bar{\nu})$	< 420%	25%	9.3%
$f_L(B^0 o K^{*0} uar u)$	_	_	0.079
$f_L(B^+ \to K^{*+} \nu \bar{\nu})$	_	-	0.077
$\mathcal{B}(B^0 \to \nu \bar{\nu}) \times 10^6$	< 14	< 5.0	< 1.5

B2TiP Report (in progress)

 $[\tau]$'s in the final state: talk by Giampierol

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Summary

- Laboratory for SM/QCD and NP
- Interesting/surprising indications of NP, tensions in $b \to c \ell \nu$ transitions at a similar level
- New techniques being implemented/improved to address long-distance effects
- Promising future with LHCb and Belle (II)

Thanks!

[Illustrative figures from Gratrex+'15, Blake+'16]

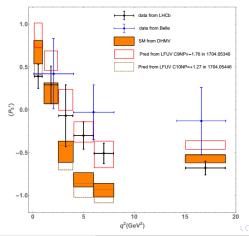
(and apologies for omissions in the references)

Consistency among experimental data

- Indirect information out of LFU violation agrees with full fit
- \rightarrow Coherent picture of a large set of observables,

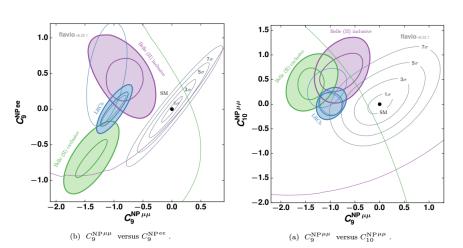
of different categories

For **illustration** only: Inputs: $R_{K^{(*)}}$, $Q_{4,5}$ $C_9^\mu = -1.76$ (red) $[C_{10}^\mu = 1.27$ (brown), C_{10}^e equiv. to SM (orange)]



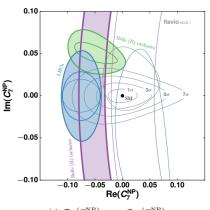
Prospects for the future - 1

[Albrecht+'17 (and refs. therein)]



Prospects for the future - 2

[Albrecht+'17 (and refs. therein)]



(c) $\mathcal{R}e\left(C_7^{\mathrm{NP}}\right)$ versus $\mathcal{I}m\left(C_7^{\mathrm{NP}}\right)$.

