

# Long-baseline neutrino experiments: status and outlook

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# OUTLINE

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- Status of neutrino oscillation measurements
- Latest results
  - (Largely overlap with T2K and NOvA talks yesterday...)
- Next generation experiments
  - Hyper-K and DUNE
- Caveat
  - This talk focuses on accelerator-based long-baseline expts
  - This talk may be highly biased by my personal view and apologies if your experiment is not covered in this talk

# STATUS OF $\nu$ OSCILLATIONS

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$c_{ij} \equiv \cos\theta_{ij}, \quad s_{ij} \equiv \sin\theta_{ij}$$

- Mixing between all three neutrino flavors has been observed

- $\theta_{12} \sim 34^\circ$

- $\theta_{13} \sim 9^\circ$

- $\theta_{23} \sim 45^\circ$  (maximal?)

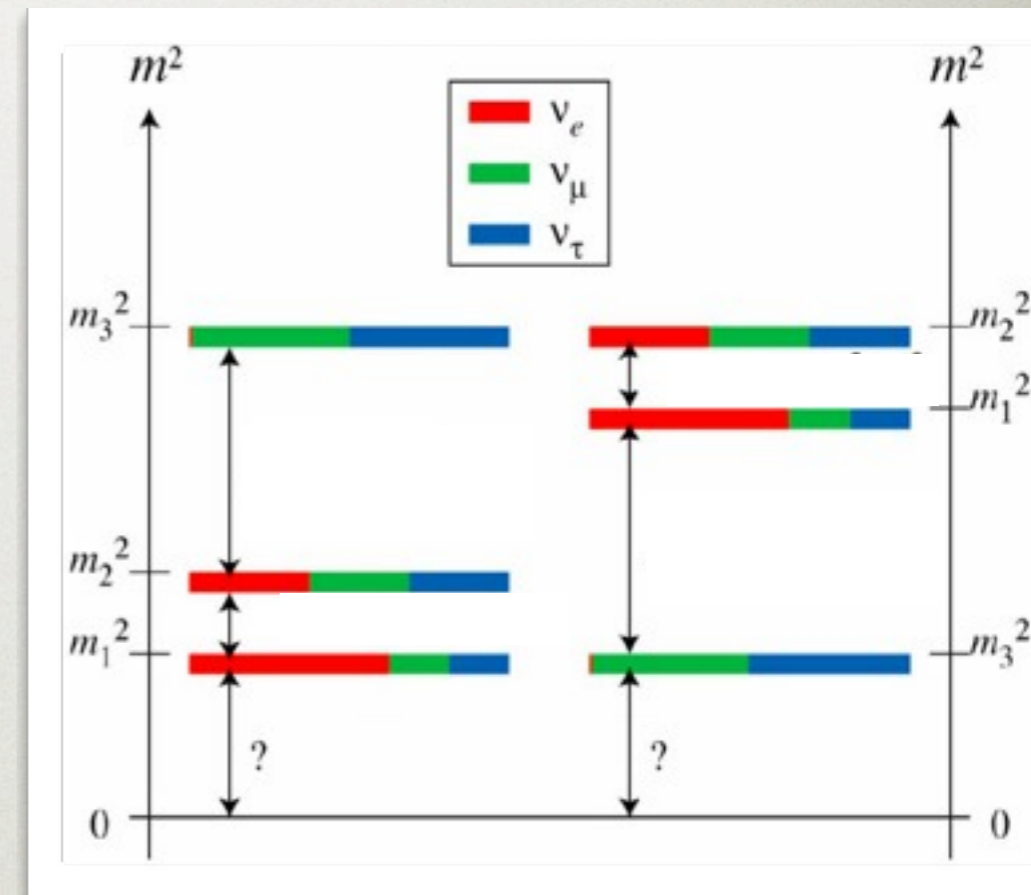
- Two mass differences

- $\Delta m_{21}^2 \sim 7.6 \times 10^{-5} \text{ eV}^2$

- $|\Delta m_{32}^2| \sim 2.4 \times 10^{-3} \text{ eV}^2$  (hierarchy?)

- CP phase  $\delta_{CP}$  remains unknown

- Also need to test “standard” 3-flavor neutrino oscillation paradigm



Normal  
hierarchy  
( $\Delta m_{32}^2 > 0$ )

Inverted  
hierarchy  
( $\Delta m_{32}^2 < 0$ )

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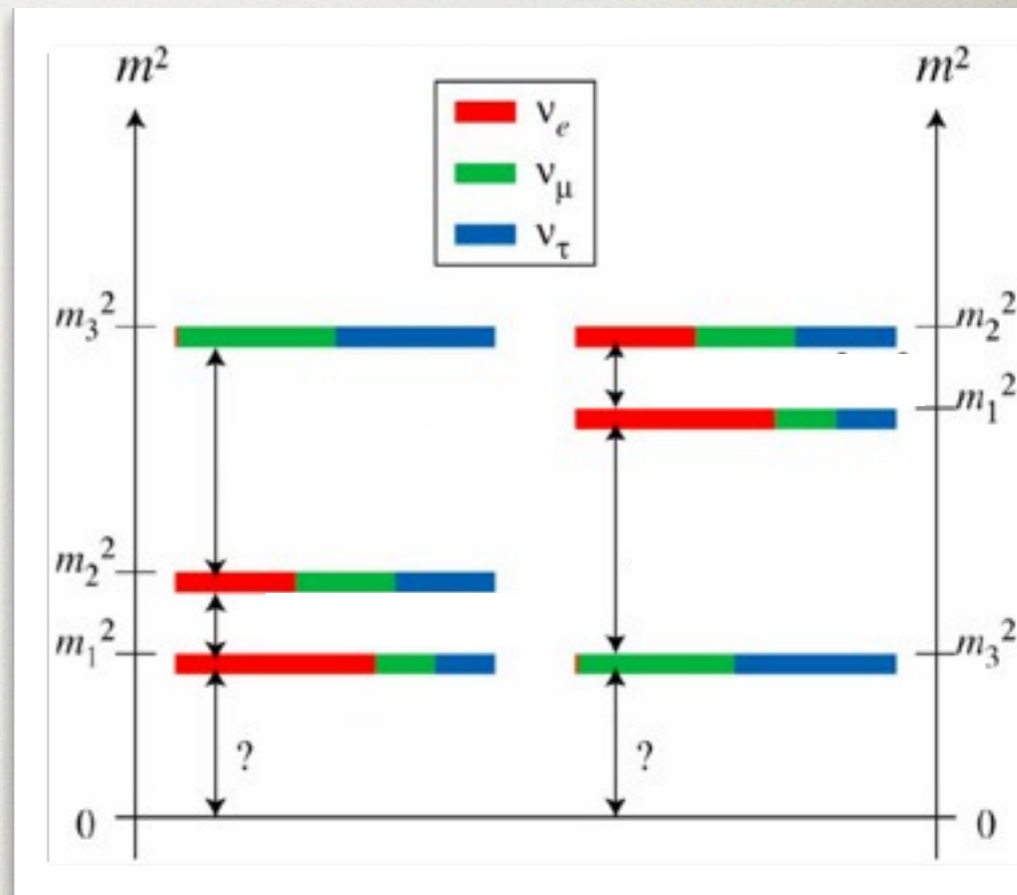
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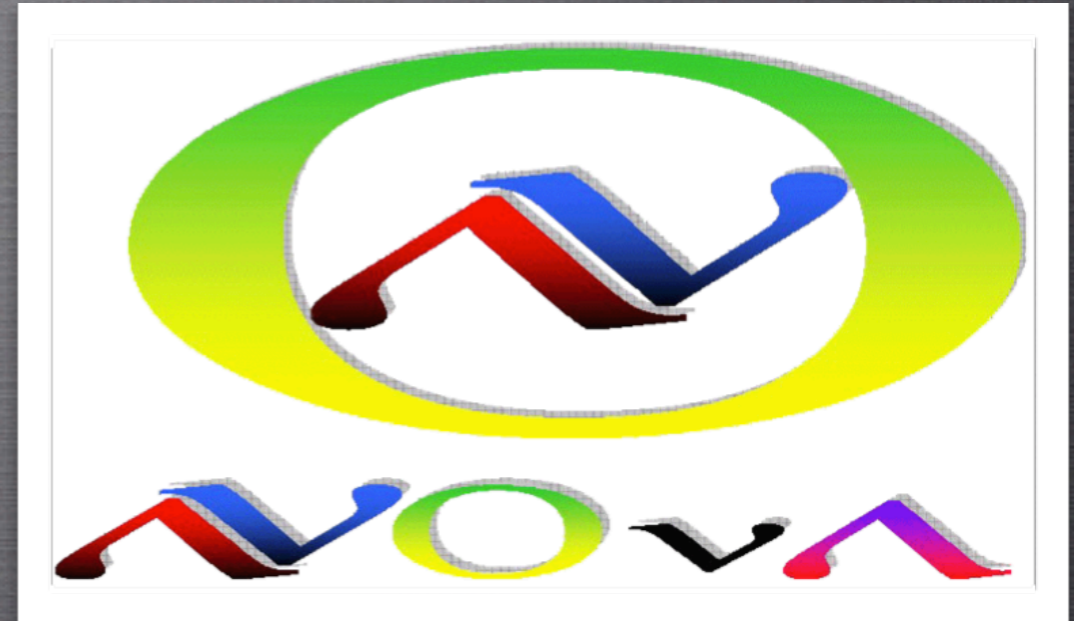
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Normal hierarchy  
( $\Delta m_{32}^2 > 0$ )

Inverted hierarchy  
( $\Delta m_{32}^2 < 0$ )

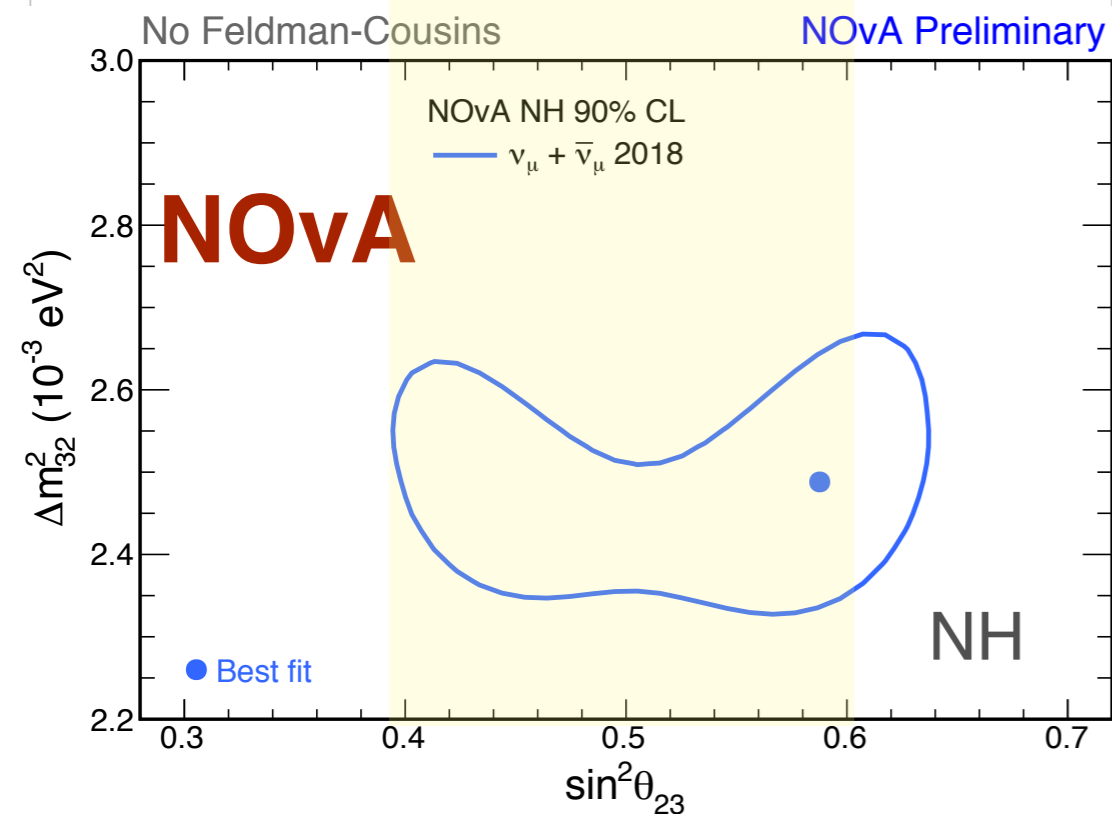
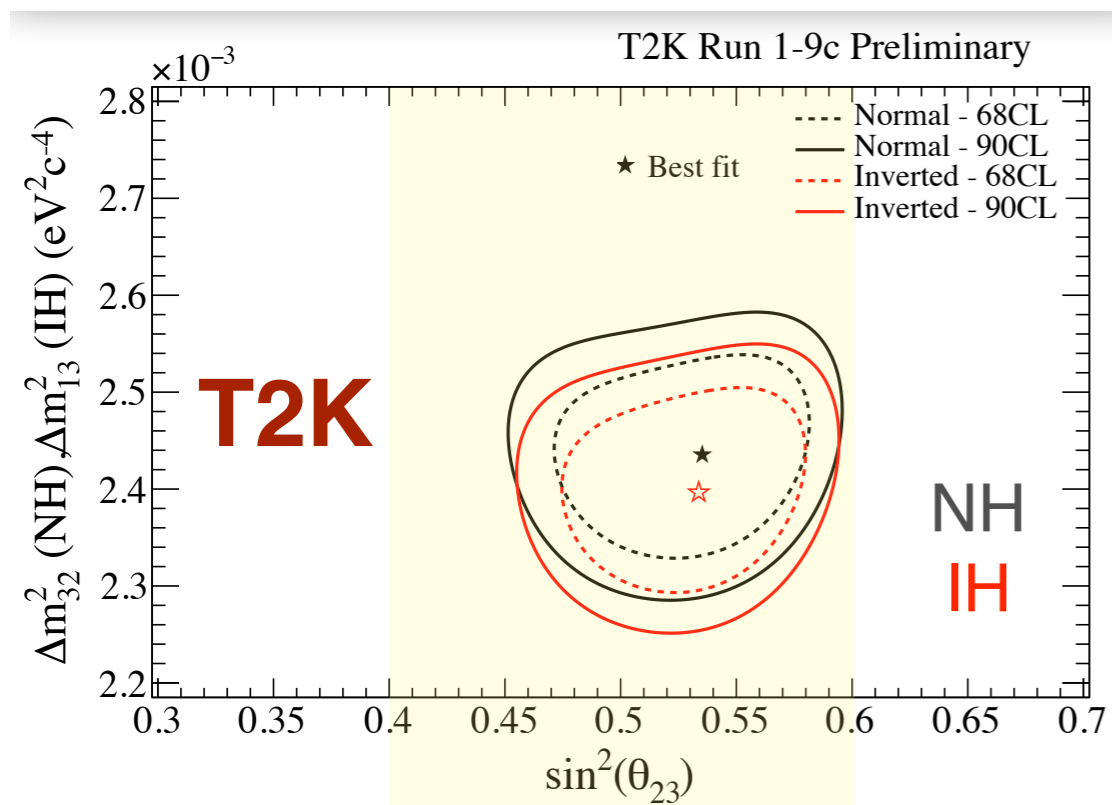


# THE LATEST RESULTS

FROM NEUTRINO 2018  
CONFERENCE

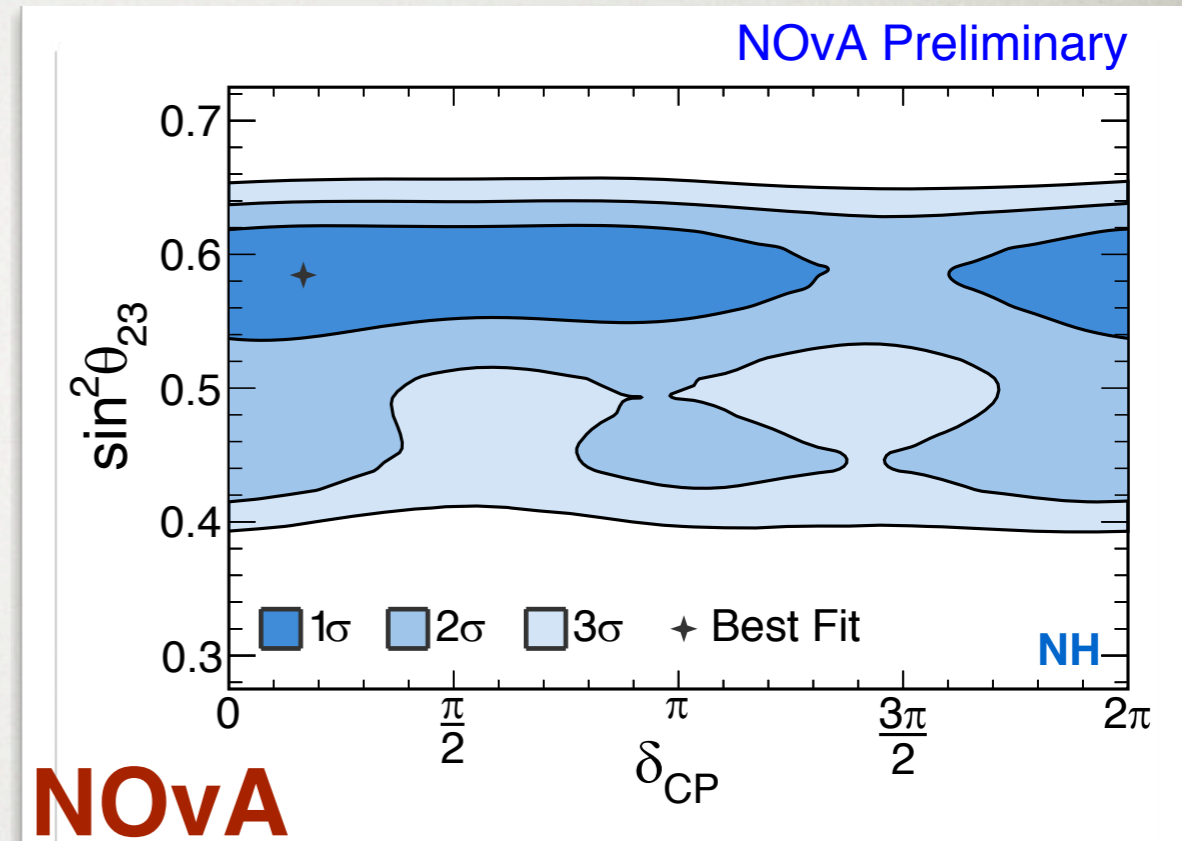
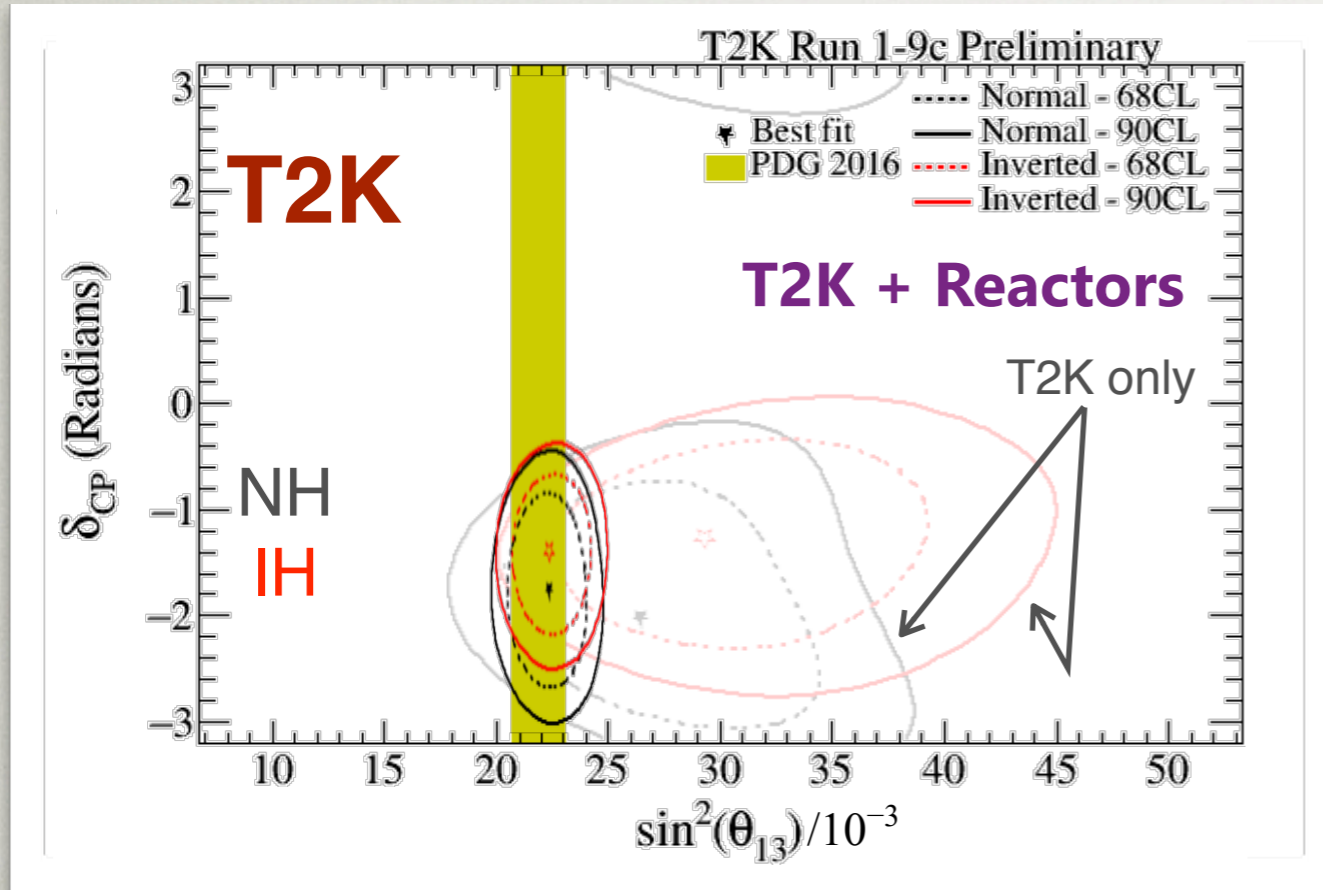
T2K results: M. Wascko's slides at Neutrino 2018  
NOvA results: M. Sanchez's slides at Neutrino 2018

# LATEST RESULTS: $\theta_{23}$

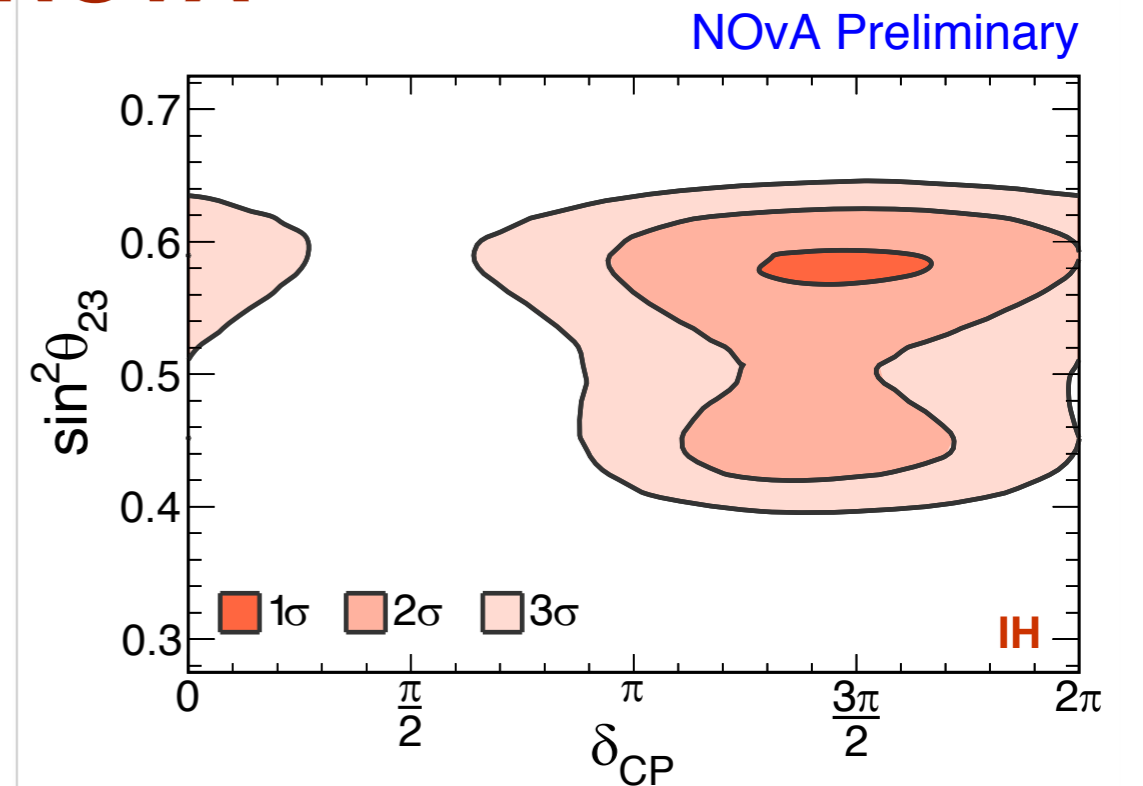


- Results from T2K and NOvA experiments consistent each other
- $0.4 \lesssim \sin^2 \theta_{23} \lesssim 0.6$
- Best fit at upper octant
- $\theta_{23} = 45^\circ$  (maximal mixing) in 90% C.L. allowed region

# LATEST RESULTS: $\delta_{CP}$

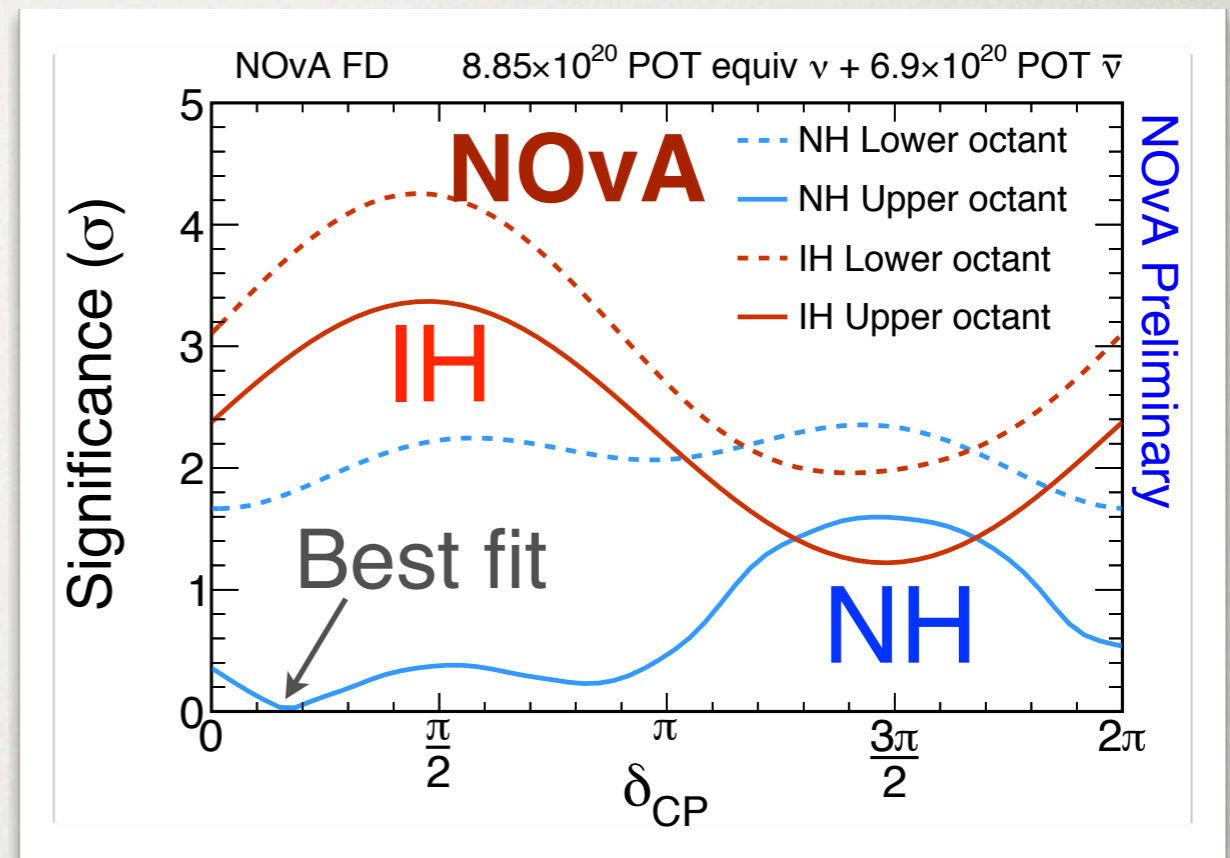
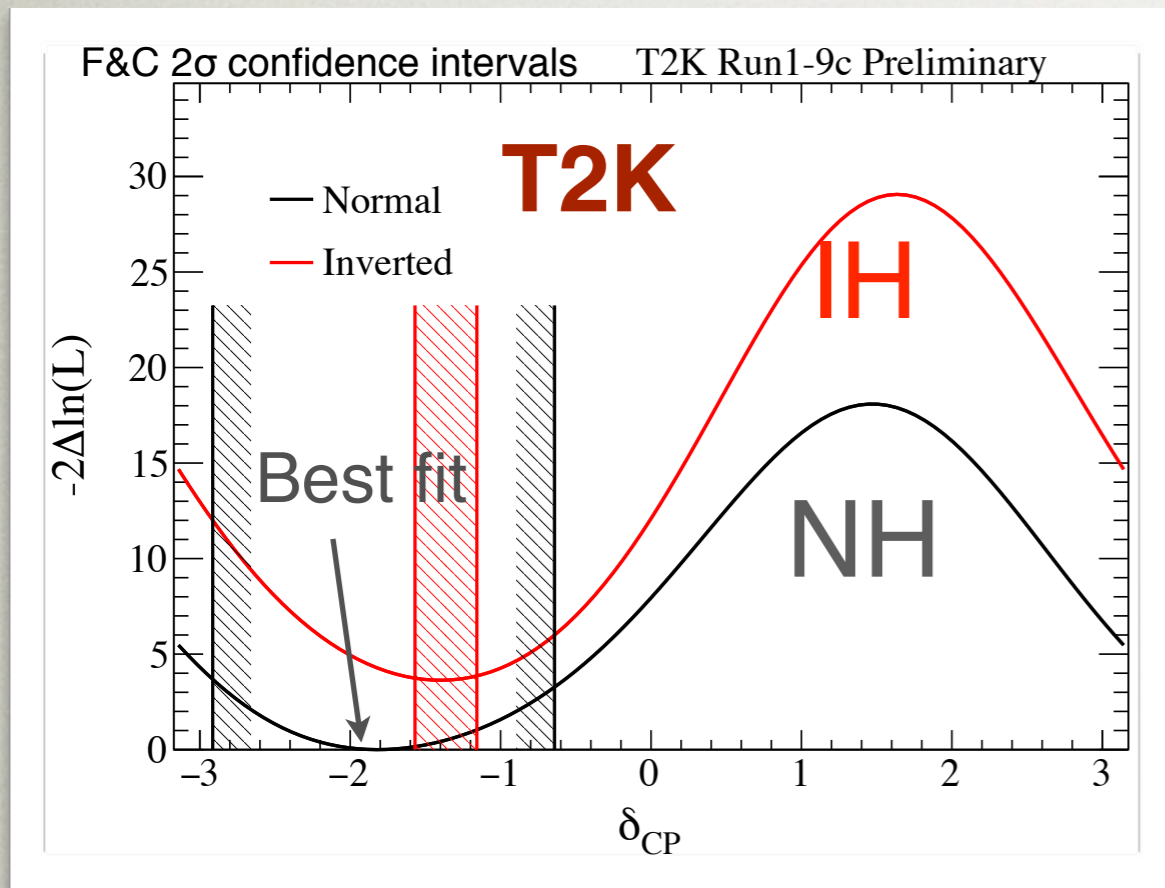


**NOvA**



- T2K results:
  - Exclude  $\delta_{CP}=0$  with  $>2\sigma$  C.L.
  - Stronger constraint than sensitivity
  - Best fit at  $\delta_{CP}\sim-\pi/2$
- NOvA results:
  - Exclude  $\delta_{CP}=+\pi/2$  with  $>3\sigma$  (IH)
  - Best fit at  $\delta_{CP}=0.17\pi$  (NH)

# LATEST RESULTS: MH



## T2K

**Bayesian approach** to extract posterior probability for mass hierarchy and octant

	$\sin^2\theta_{23}\leq 0.5$	$\sin^2\theta_{23}>0.5$	SUM
NH ( $\Delta m^2_{32}>0$ )	0.204	0.684	0.888
IH ( $\Delta m^2_{31}<0$ )	0.023	0.089	0.112
SUM	0.227	0.773	1

- T2K and NOvA both results prefer ( $< 2\sigma$ ):
  - Normal hierarchy
  - Upper octant
- cf. Super-K atmospheric  $\nu$  results prefer NH ( $\sim 2\sigma$ )
  - PRD 97, 072001 (2018)



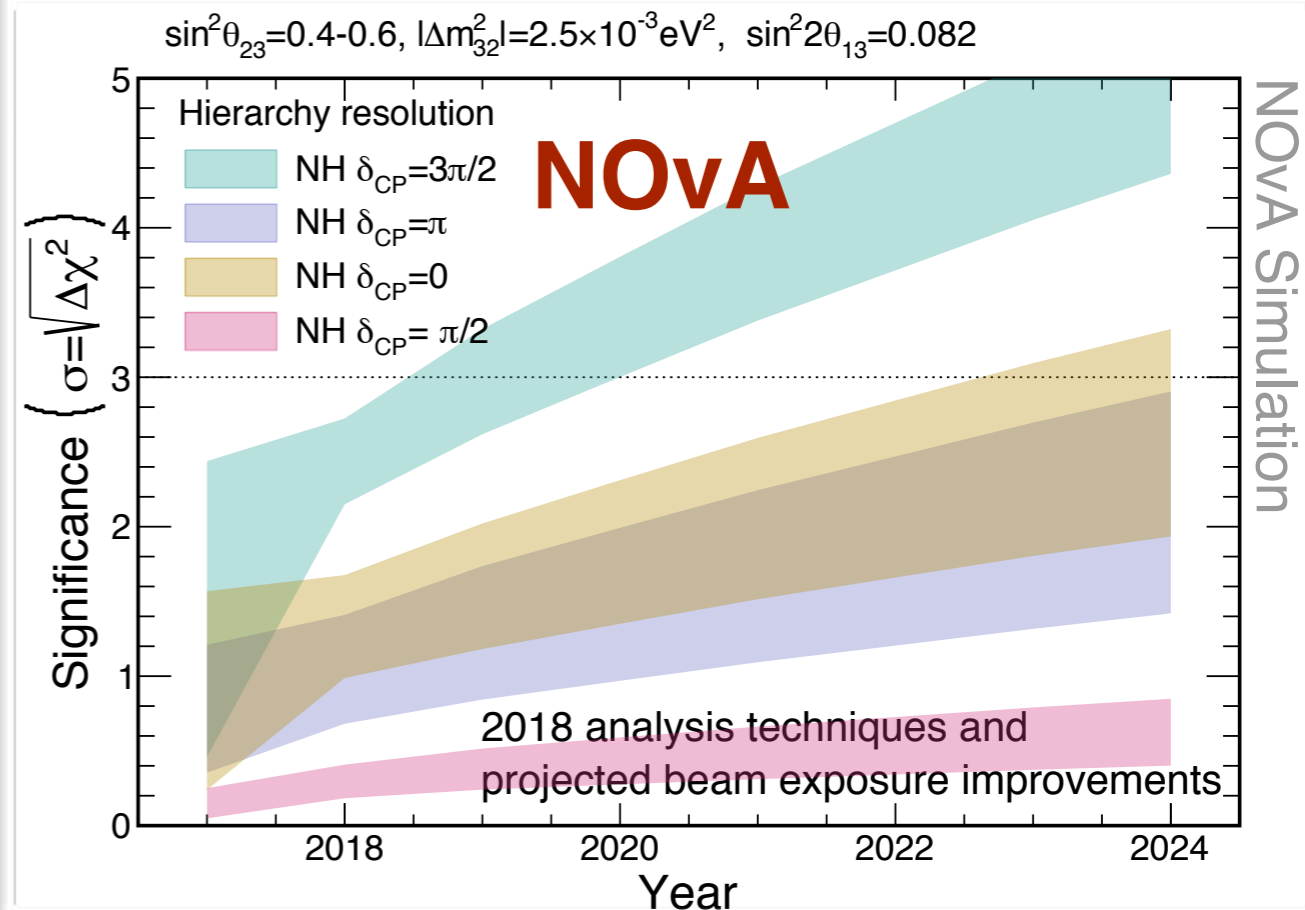
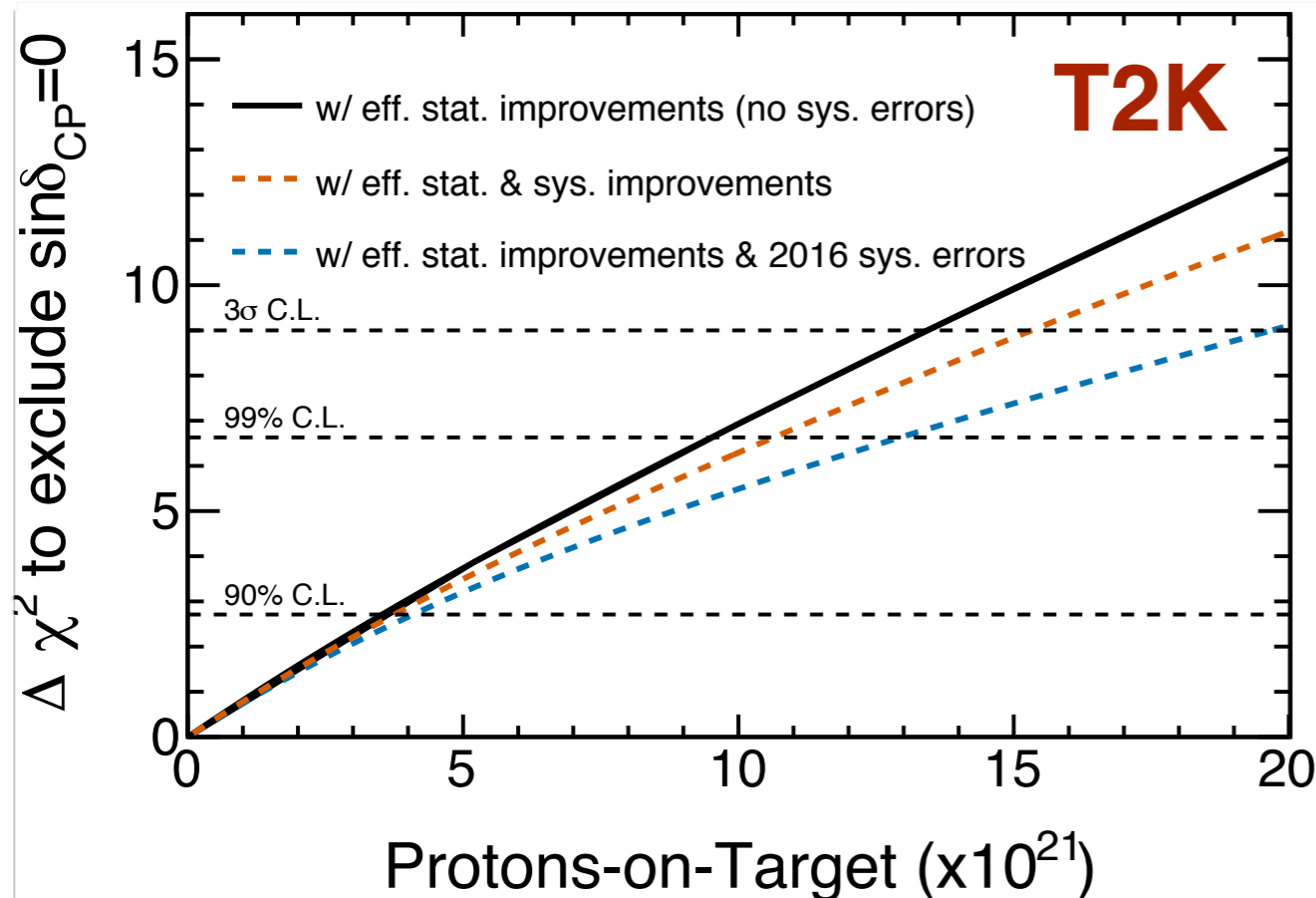
# NEAR FUTURE...

- T2K and NOvA experiments both propose to extend their data taking up around 2024~2026
- $\sin\delta_{CP}=0$  can be excluded with  $3\sigma$  C.L. (T2K)
- MH can be determined with  $>3\sigma$  C.L. (NOvA)

[Assume true Mass Hierarchy is 'Normal' and  $\delta_{CP}=-\pi/2$ ]

T2K:  $\sin\delta_{CP}=0$  exclusion sensitivity  
(arXiv:1609.04111v1)

NOvA: Mass Hierarchy sensitivity  
(Neutrino 2018)



# 次世代実験

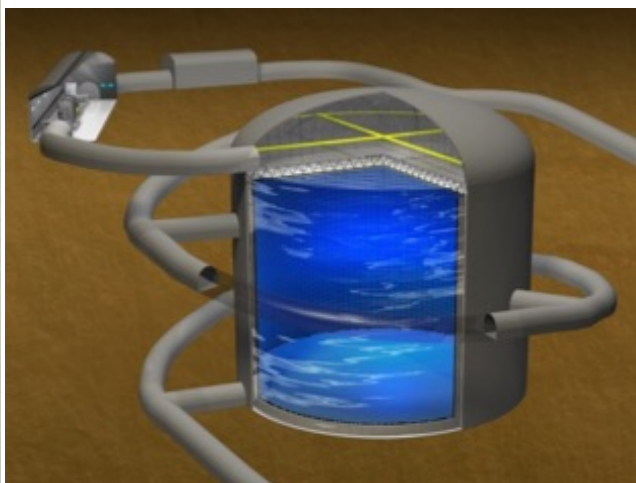
THE NEXT GENERATION

# EXPERIMENTAL STRATEGY

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- Next generation LBL experiments target CP violation (CPV) and Mass Hierarchy
- **Hyper-Kamiokande**
  - **Shorter baseline (295km):**  
Earth matter effect insignificant → Focus on CPV
  - $\nu$  beam: Flux peak at 1st oscillation maximum
  - → **Off-axis narrow band beam**
  - Mass Hierarchy can be determined with atmospheric  $\nu$
- **DUNE/LBNF**
  - **Longer baseline (1300km):**  
Measure matter effect (MH) → Unfold CPV from Earth matter effect through  $\nu$  spectrum shape
  - $\nu$  beam: Cover wide energy range (1st and 2nd maxima)
  - → **On-axis wide band beam**

# NEXT GENERATION LBL EXPTS



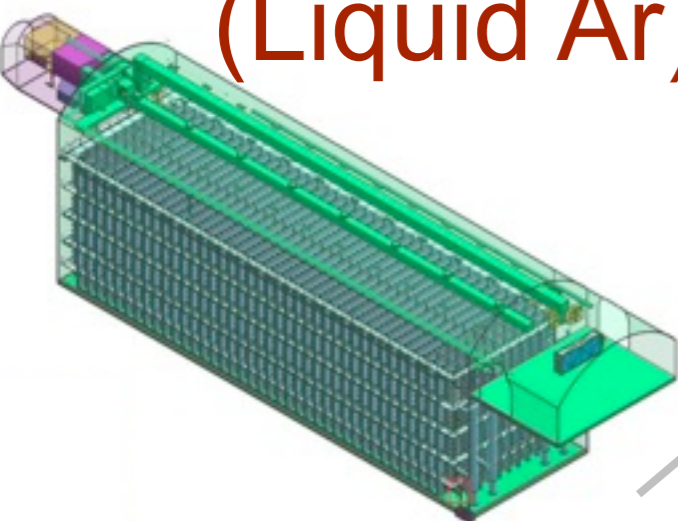
**Hyper-K**  
(Water Č)



**J-PARC**  
Accelerator Complex



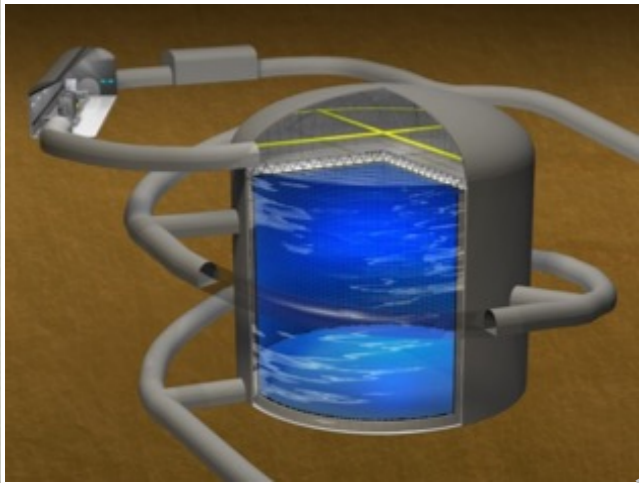
**DUNE**  
(Liquid Ar)



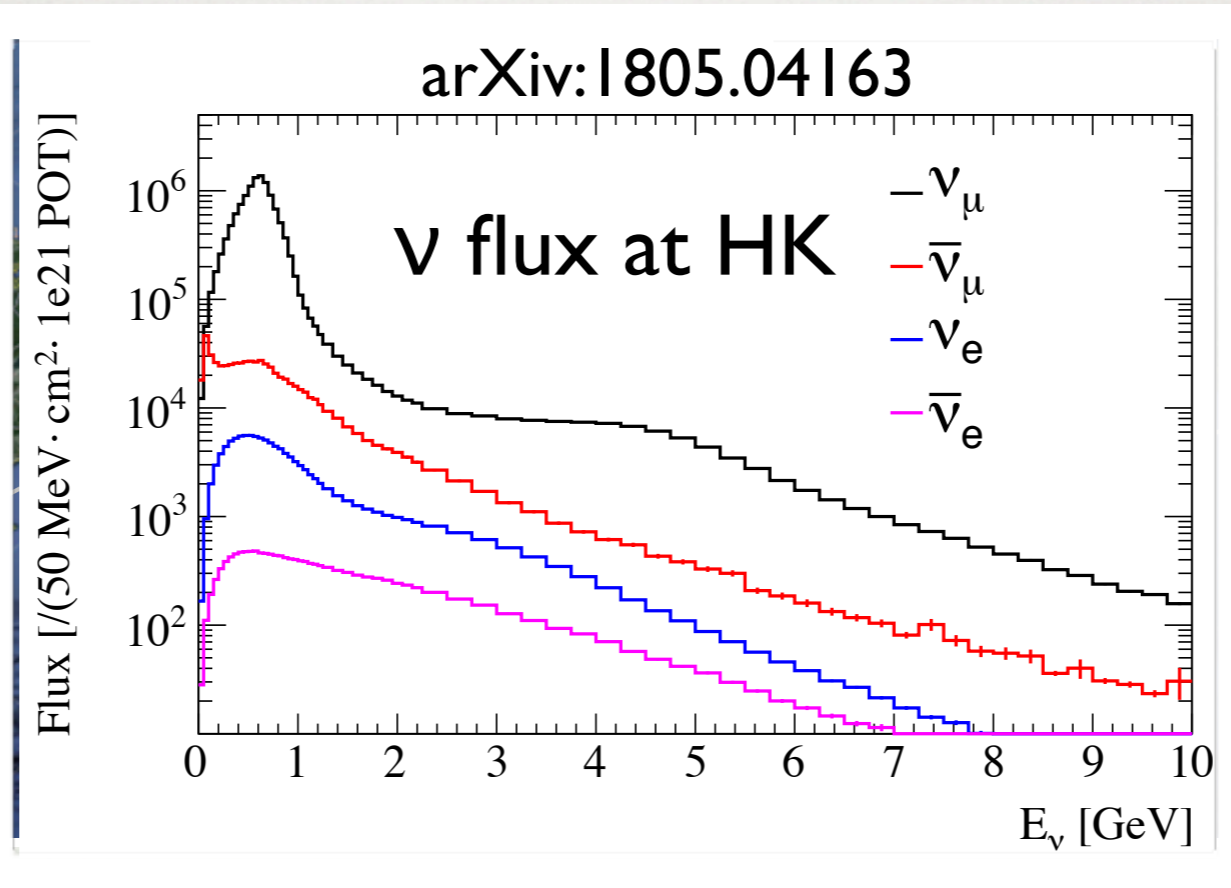
**Fermilab**  
**LBNF**



# NEXT GENERATION LBL EXPTS



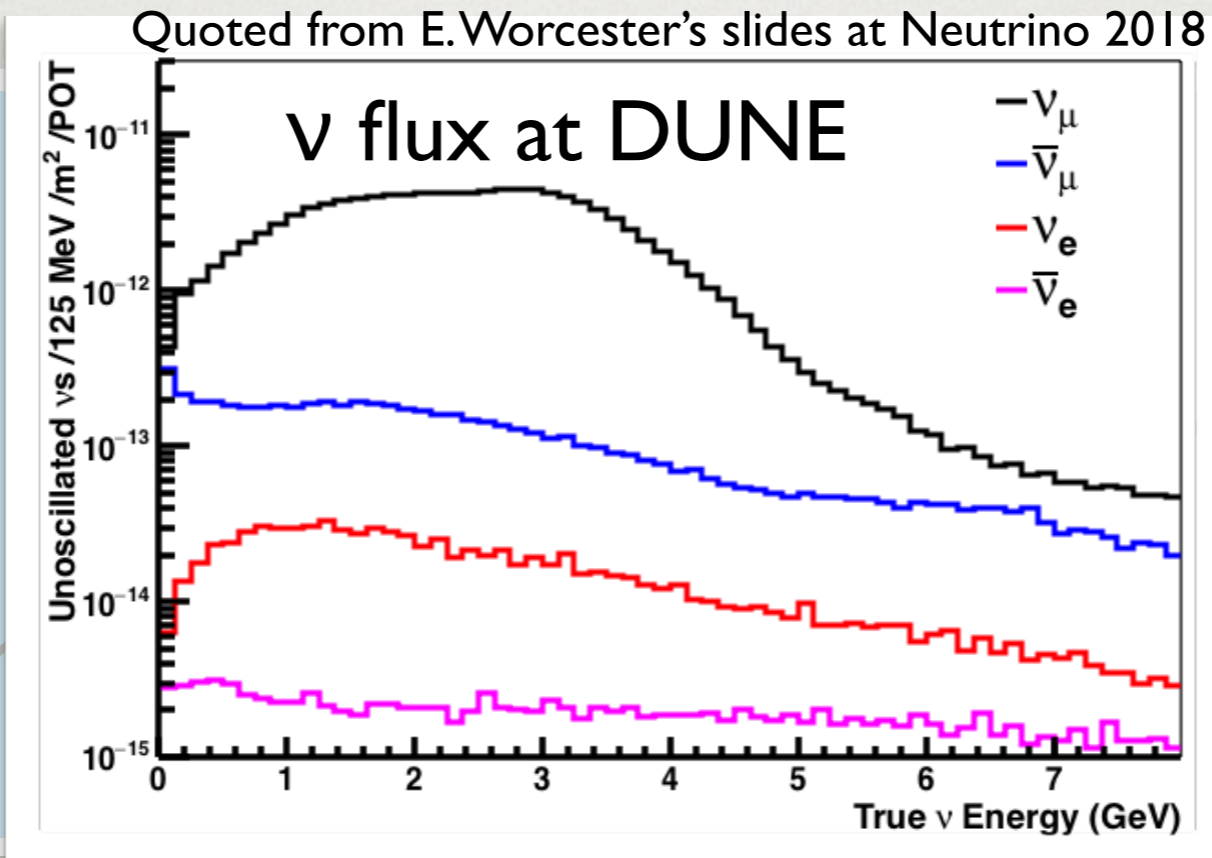
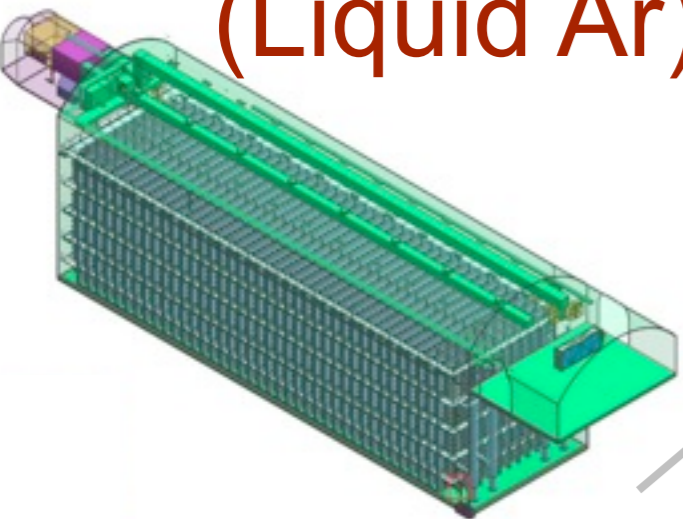
**Hyper-K**  
(Water Č)



**J-PARC**  
Accelerator Complex



**DUNE**  
(Liquid Ar)



**Fermilab**  
**LBNF**



## Super-Kamiokande IV

Run 999999 Sub 0 Event 686

11-11-23:19:16:48

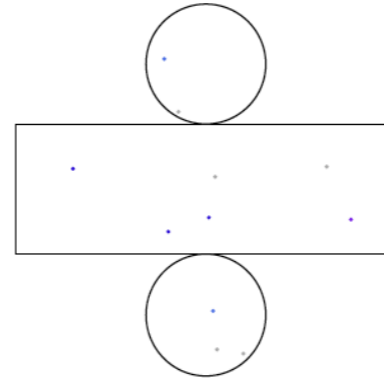
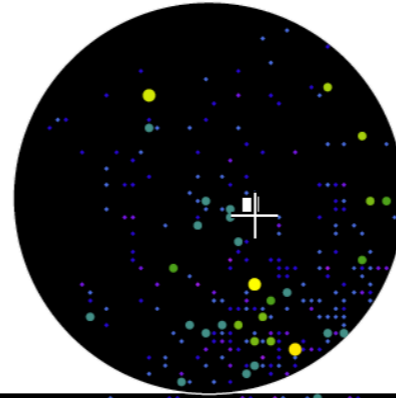
Inner: 2222 hits, 4687 pe

Outer: 6 hits, 6 pe

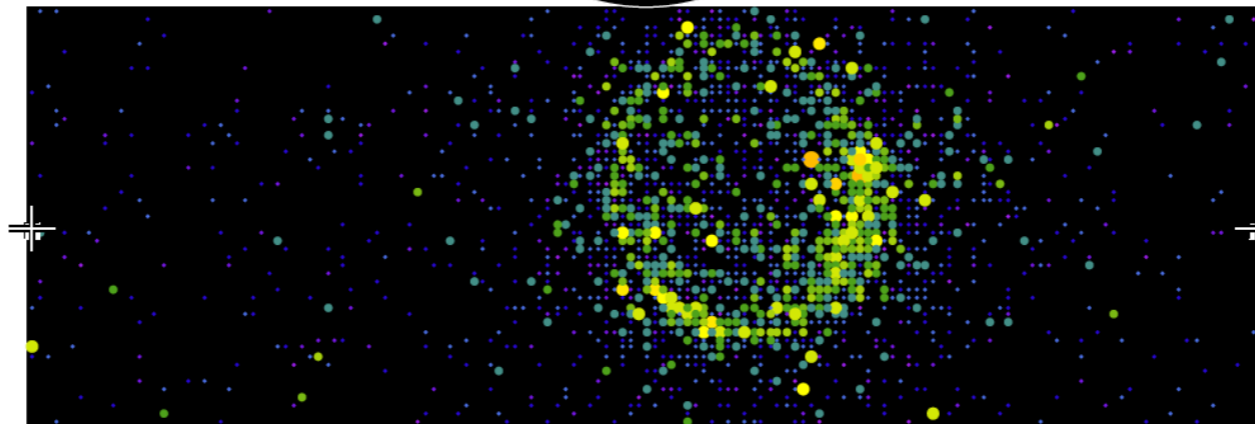
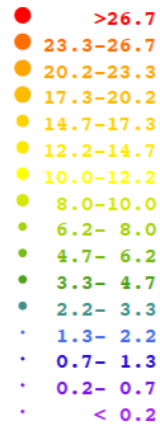
Trigger: 0x07

D\_wall: 1270.2 cm

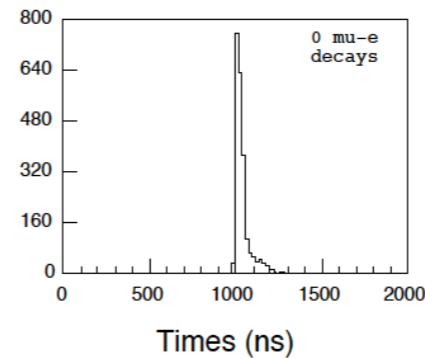
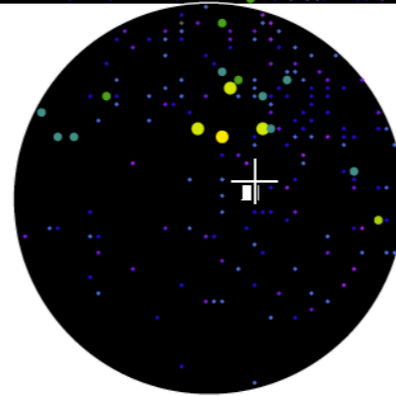
e-like, p = 480.6 MeV/c



### Charge (pe)



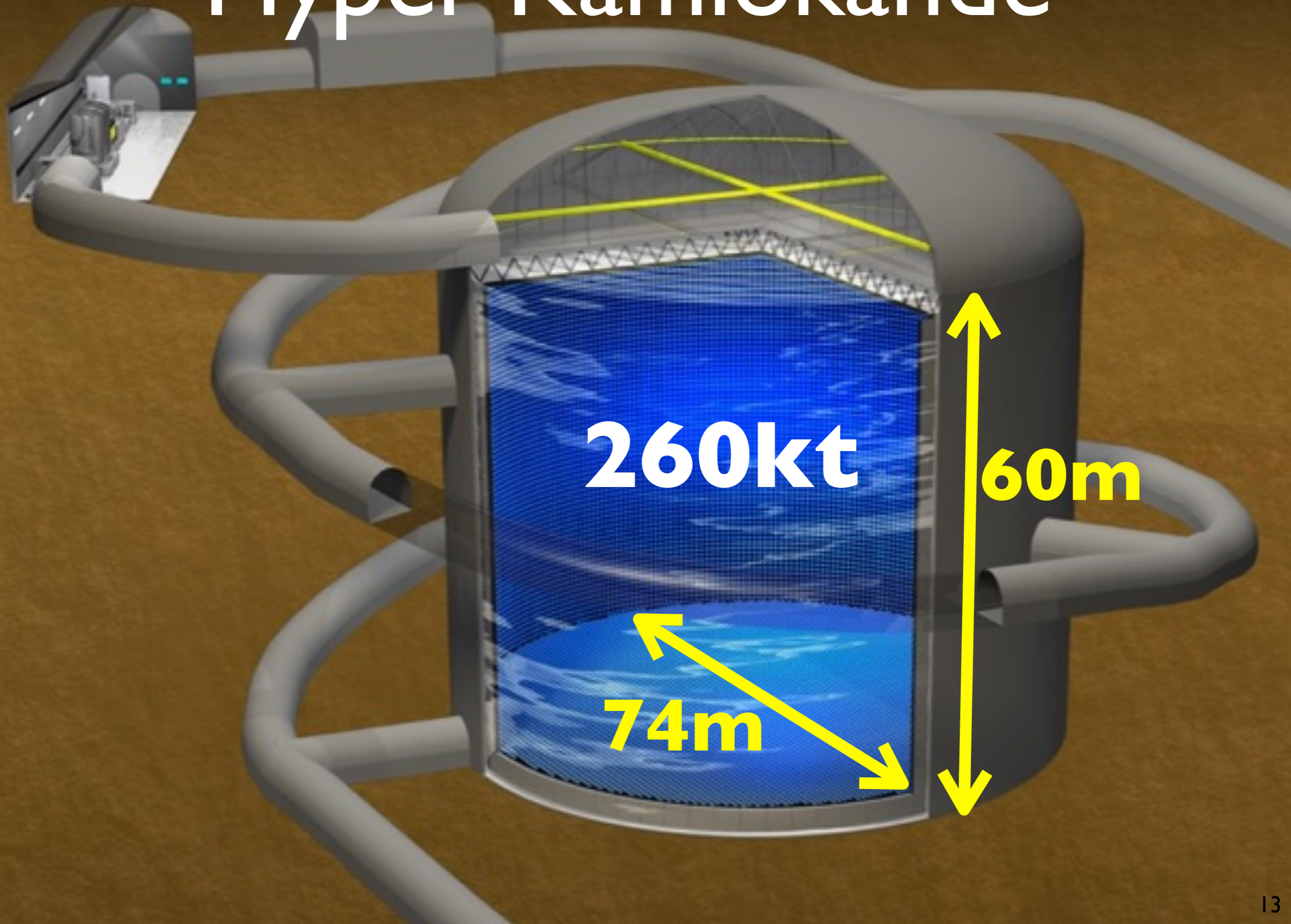
$\nu_e$  CC



# HYPER-KAMIOKANDE

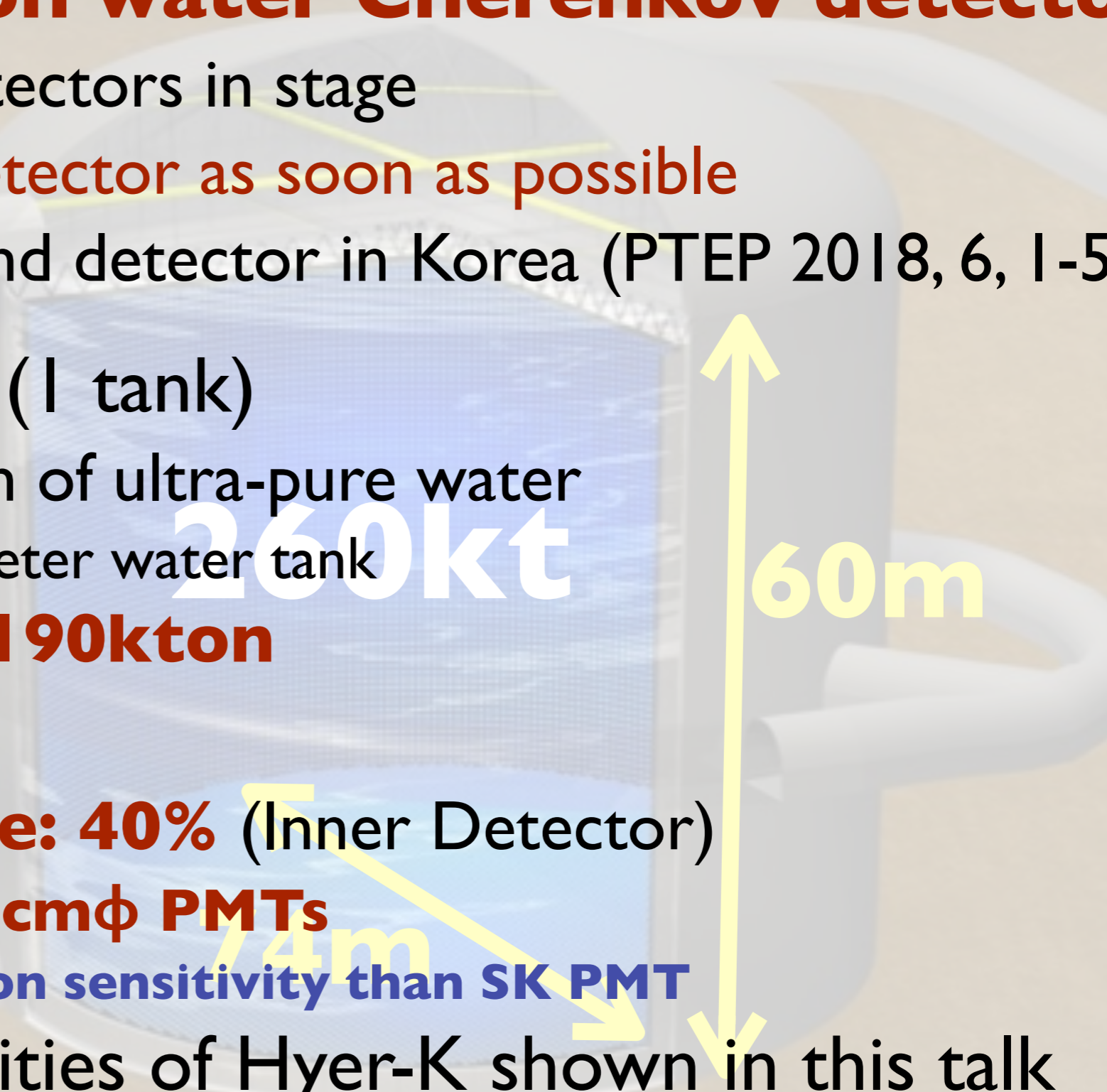
Hyper-K Design Report  
arXiv:1805.04163

# Hyper-Kamiokande



# Hyper-Kamiokande

- **Next generation water Cherenkov detector**
  - Construct two detectors in stage
  - Realize the first detector as soon as possible
  - An option of second detector in Korea (PTEP 2018, 6, 1-56)
- The first detector (1 tank)
  - Filled with 260kton of ultra-pure water
    - 60m tall x 74 diameter water tank
  - **Fiducial mass: 190kton**
    - ~10 x Super-K
  - **Photo-coverage: 40%** (Inner Detector)
    - 40,000 of **new 50cm $\phi$  PMTs**
      - x2 higher photon sensitivity than SK PMT
- All physics sensitivities of Hyer-K shown in this talk assumes 1 tank

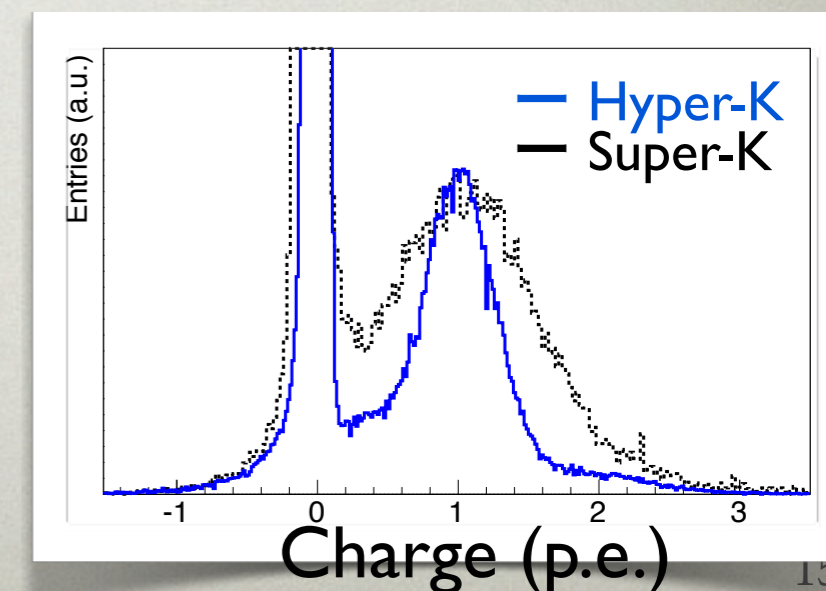
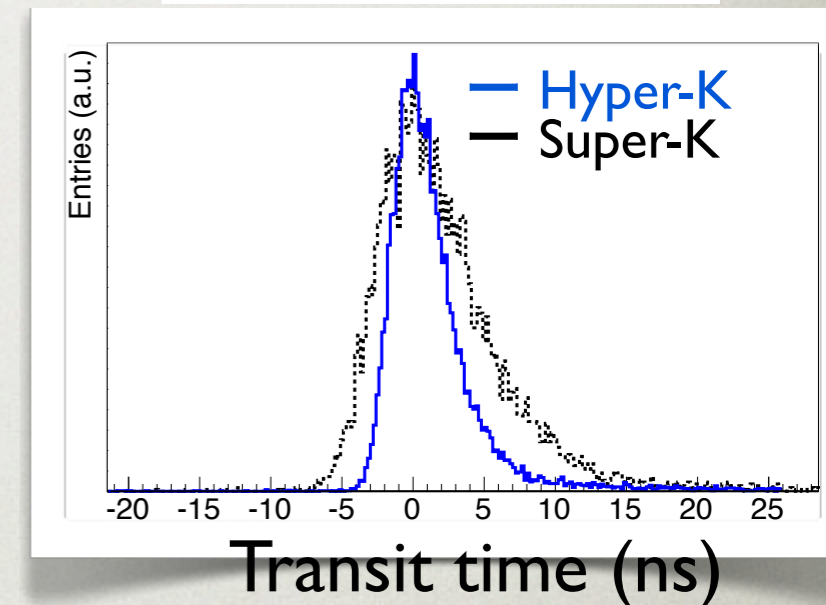
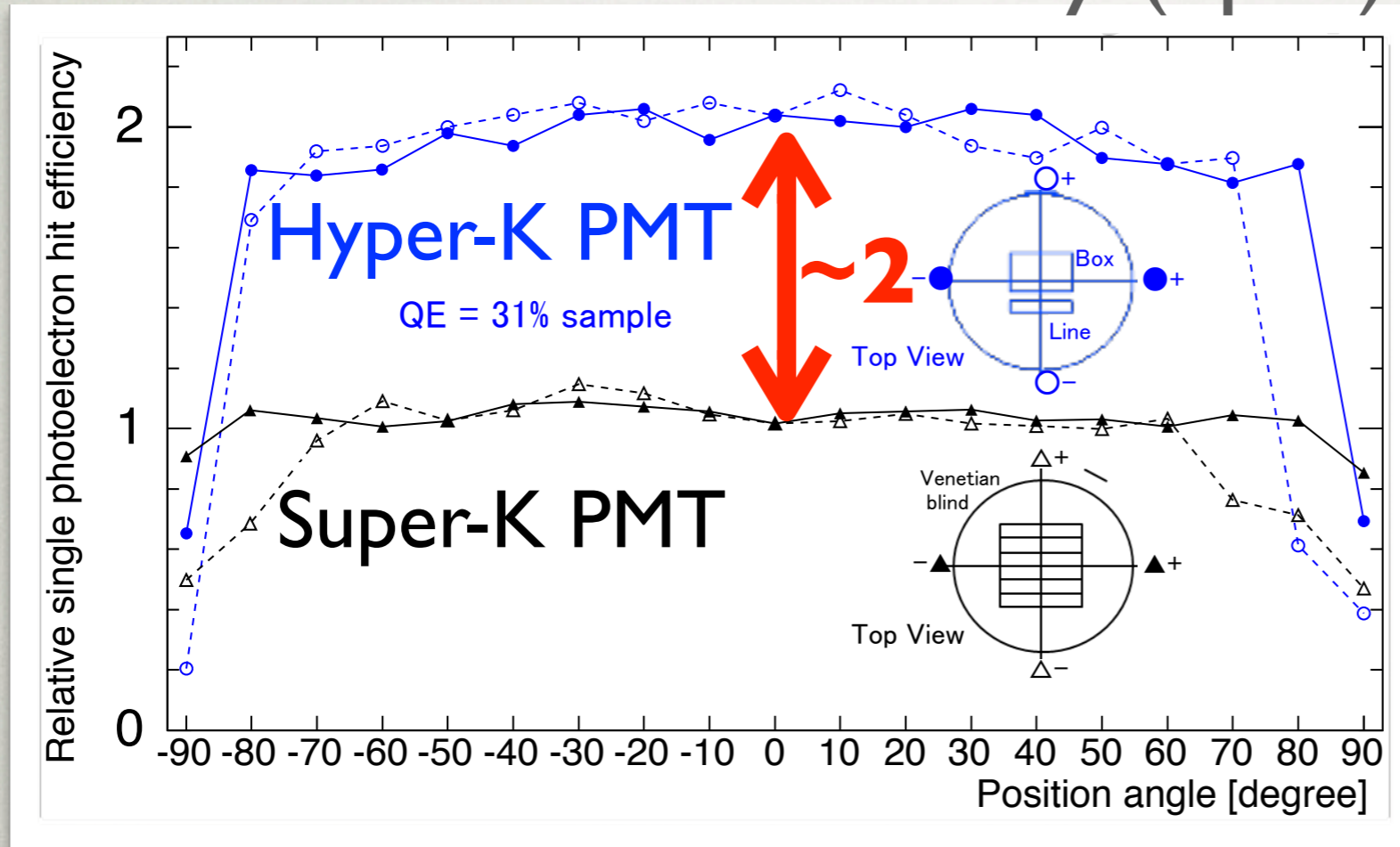




# NEW 50CM $\phi$ PMT FOR HYPER-K

Box & line dynode PMT

Photo-detection efficiency (l p.e.)

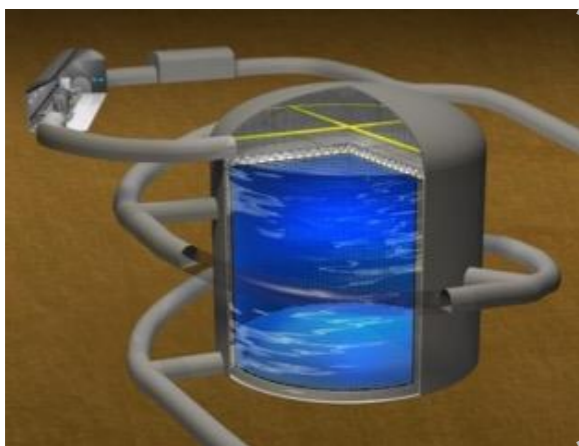
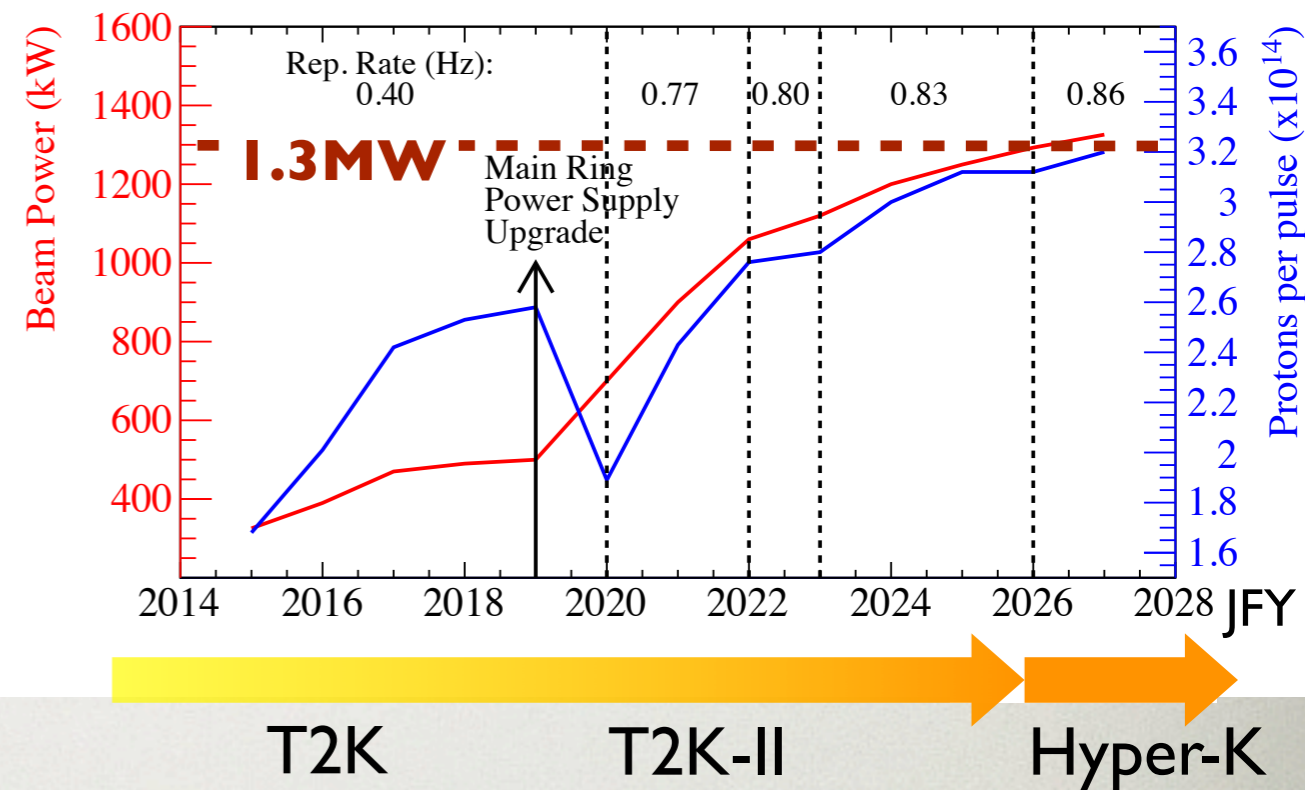


- Twice better photo-detection efficiency than SK PMTs
- Timing resolution (TTS): 1.1ns
  - cf. SK PMT: 2.1ns
- Higher pressure tolerance:  $>80$ m

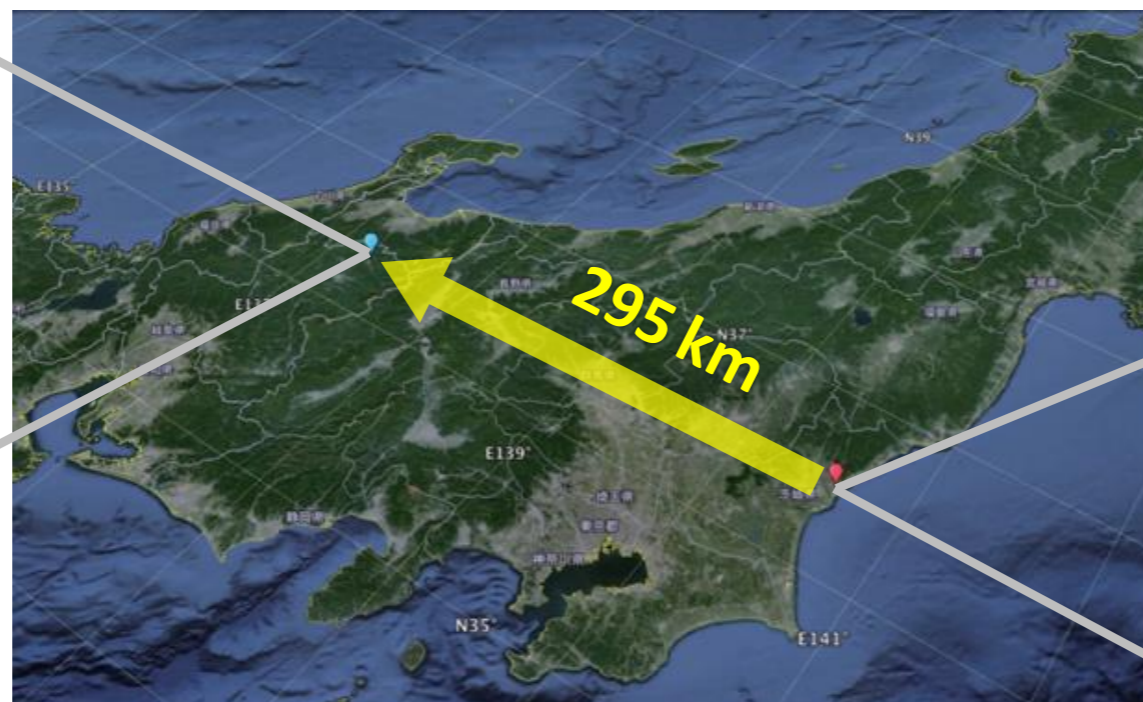
# V BEAM FOR HYPER-K

- High quality & high intensity neutrino beam
- 2.5 deg. off-axis narrow band neutrino beam (same as T2K)
- Beam power: 1.3MW (before Hyper-K begins)
  - KEK Project Implementation Plan: top priority on 'J-PARC upgrade for Hyper-K'
  - cf. Reached ~500kW for T2K

J-PARC MR Fast Extraction Power Projection



Hyper-K



J-PARC Accelerator Complex



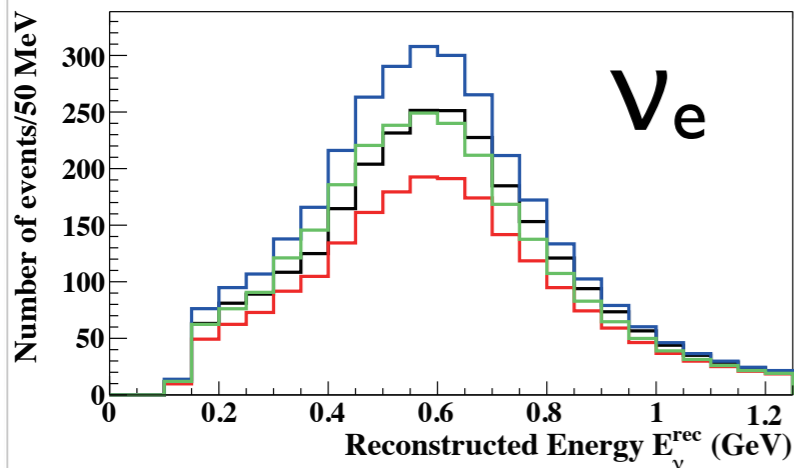
# EXPECTED EVENTS IN HK FOR CPV

Expected # of events in  $\nu_e/\bar{\nu}_e$  appearance  
(after applying  $\nu_e$  selection criteria)

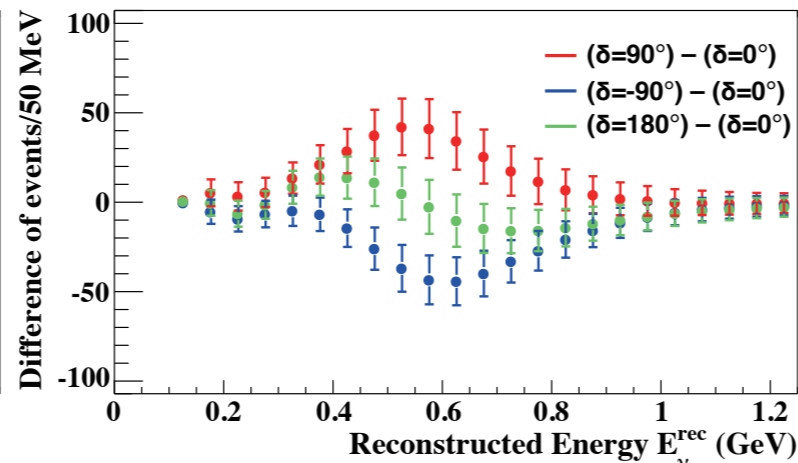
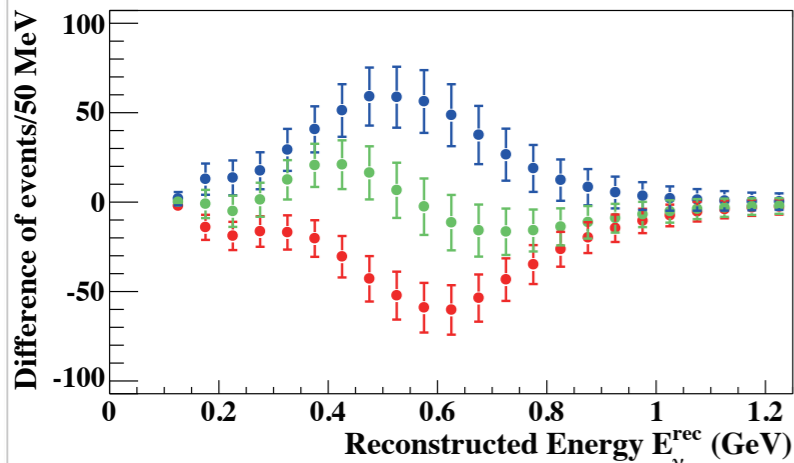
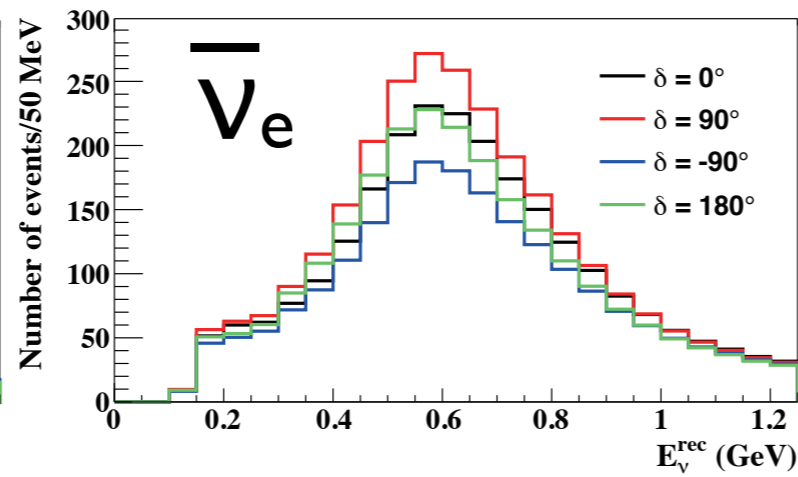
for $\delta_{CP} = 0$	Signal $\nu_\mu \rightarrow \nu_e$ CC	Wrong sign appearance	$\nu_\mu/\bar{\nu}_\mu$ CC	Beam $\nu_e/\bar{\nu}_e$ contamination	NC
$\nu$ beam	1,643	15	7	259	134
$\bar{\nu}$ beam	1,183	206	4	317	196

## Reconstructed $E_\nu$ spectra

Neutrino mode: Appearance



Antineutrino mode: Appearance



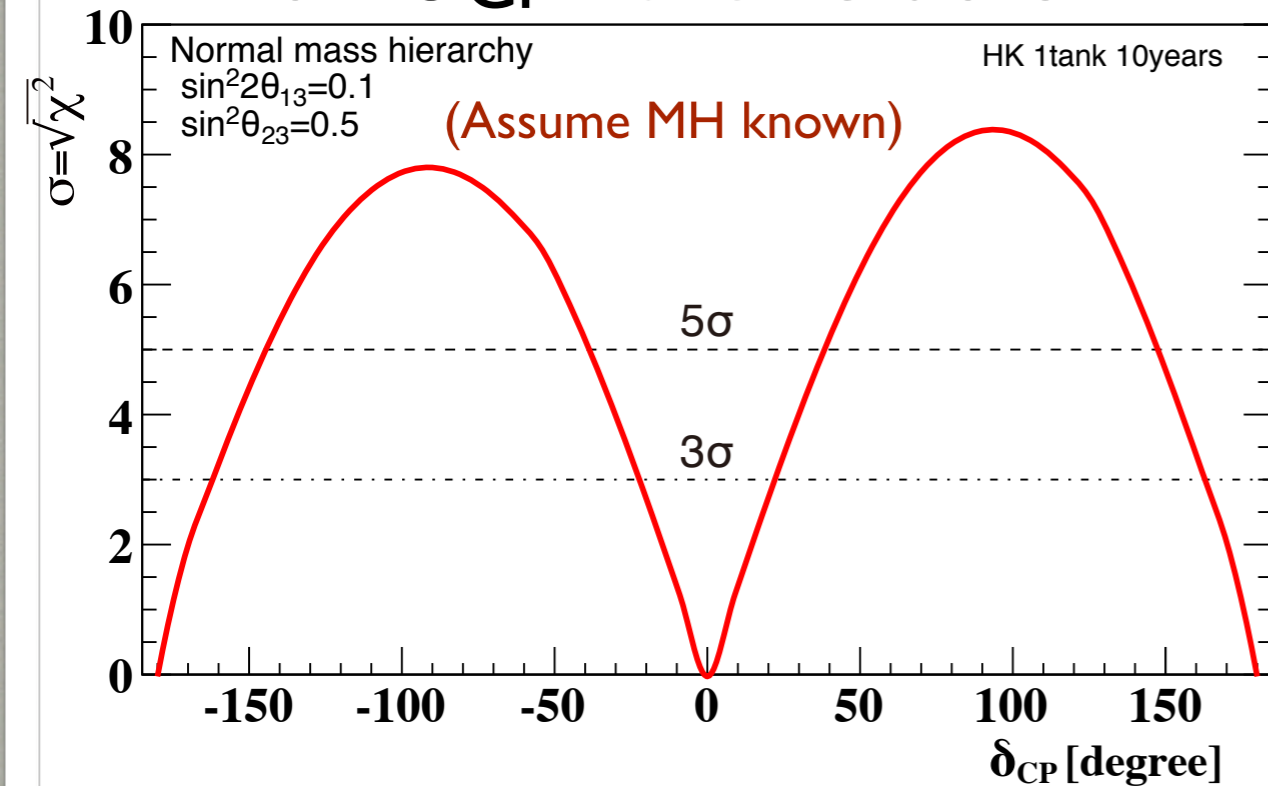
1.3MW x 10 years ( $10^8$  sec),  $\nu:\bar{\nu}=1:3$

Also sensitive to  
New Physics  
(ex. if any additional phase)

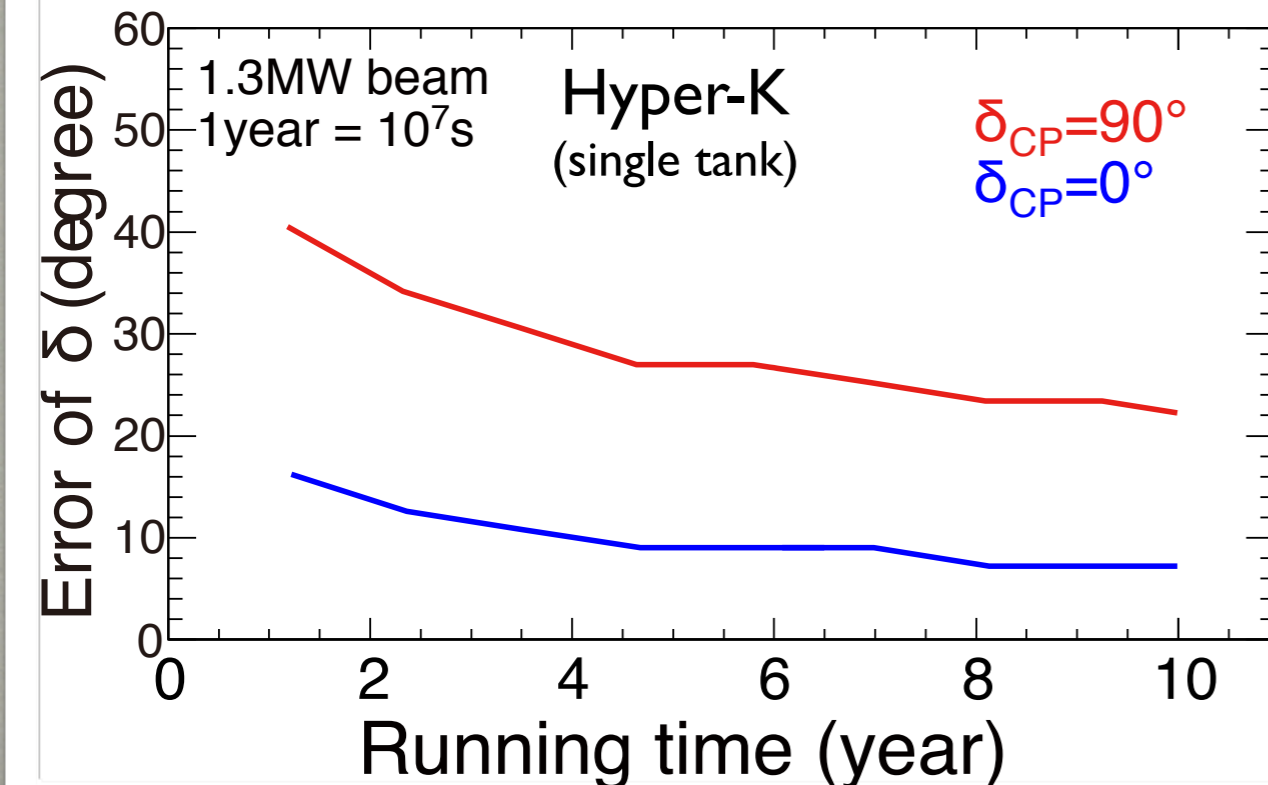
$\sin^2 2\theta_{13}=0.1$   
Normal Hierarchy

# HYPER-K SENSITIVITY FOR CPV

## $\sin\delta_{CP}=0$ exclusion

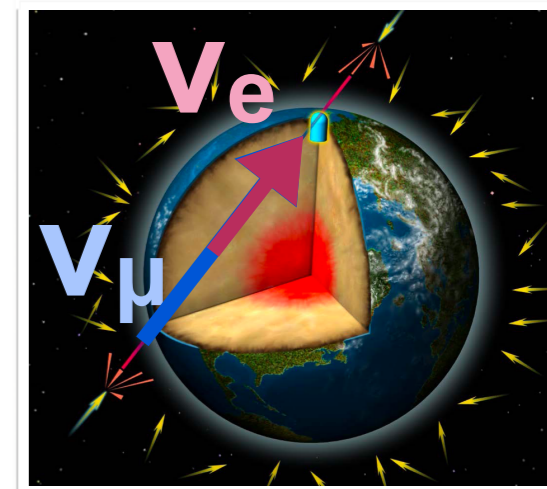
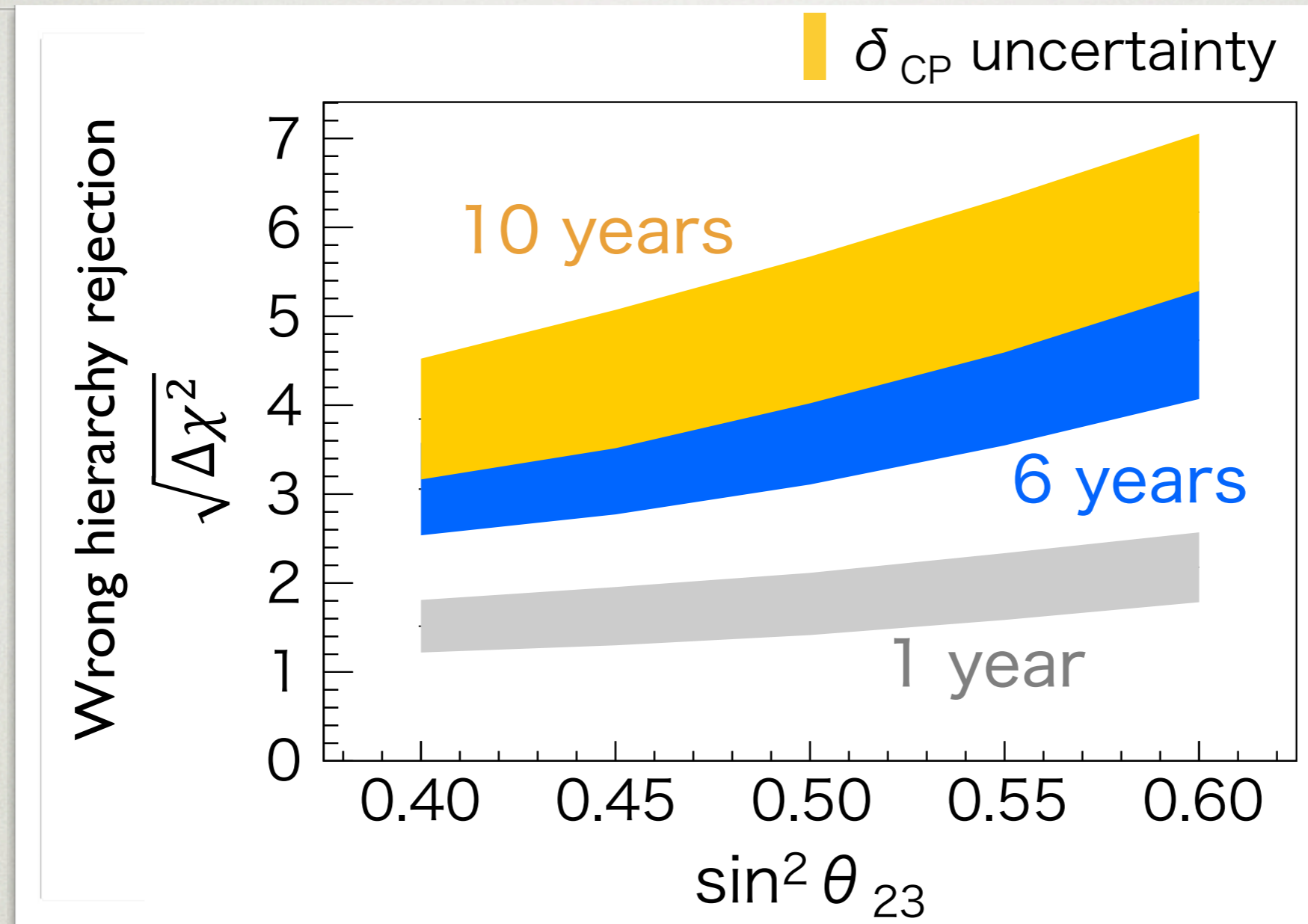


## $\delta_{CP}$ $1\sigma$ error



- $\sin\delta_{CP}=0$  exclusion:
  - $\sim 8\sigma$  for  $\delta_{CP}=-90^\circ$  (T2K best fit)
  - $\sim 80\%$  coverage of  $\delta_{CP}$  parameter space with  $>3\sigma$
- $\delta_{CP}$  resolution:
  - $22^\circ$  at  $\delta_{CP}=\pm 90^\circ$
  - $7^\circ$  at  $\delta_{CP}=0^\circ, 180^\circ$
- Sensitivity studies adopt analysis techniques and systematic uncertainties used in T2K
  - Realistic systematic uncertainties plus expected reduction of error
  - $3\sim 4\%$  syst. err (cf.  $6\sim 7\%$  in T2K)

# MASS HIERARCHY SENSITIVITY IN HK



- Hyper-K can determine Mass Hierarchy in  $\sim 5$  years ( $\sin^2 \theta_{23} = 0.5$ ) using atmospheric  $\nu$ 's, even if MH not determined before Hyper-K era
- cf. Super-K suggests Normal Hierarchy with  $\sim 2\sigma$ 
  - Phy. Rev. D97, 072001 (2018)

# Hyper-K: multi-purpose detector

- **Comprehensive study of  $\nu$  oscillation**

- CPV: 76% of  $\delta$  space w/  $3\sigma$ ,  $<22^\circ$  precision
- MH determination for all  $\delta$  with J-PARC/Atm  $\nu$
- $\theta_{23}$  octant determination at  $|\theta_{23}-45^\circ|>2^\circ$
- $<1\%$  precision of  $\Delta m^2_{32}$
- Test standard  $\nu$  oscillation scenario w/ acc/atm  $\nu$

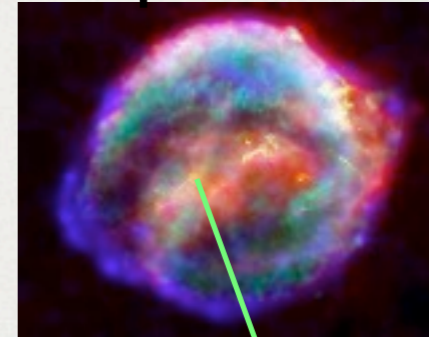
- **Proton decay  $3\sigma$  discovery potential**

- $1 \times 10^{35}$  years for  $p \rightarrow e^+ \pi^0$
- $3 \times 10^{34}$  years for  $p \rightarrow \nu K^+$

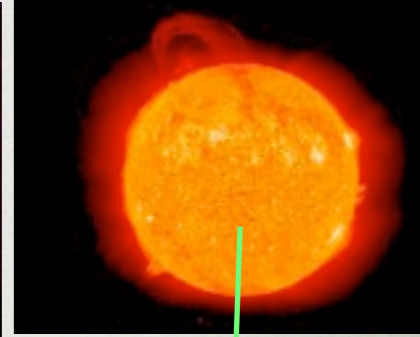
- **Astrophysical neutrino**

- Solar  $\nu$ : test standard matter effect (MSW) model
- Supernova  $\nu$ , supernova relic- $\nu$
- Dark matter neutrinos from Sun, Galaxy, Earth

Supernova



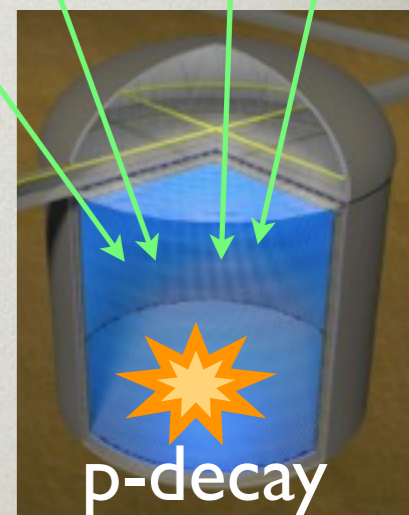
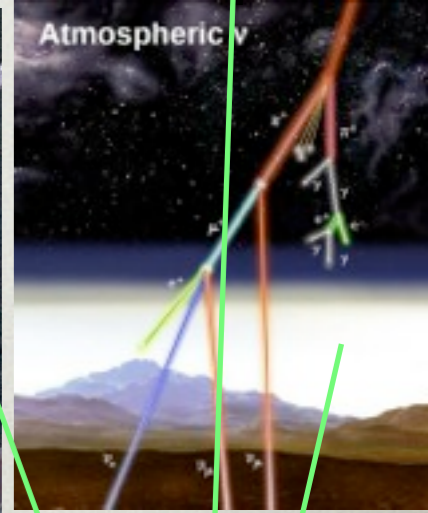
Sun



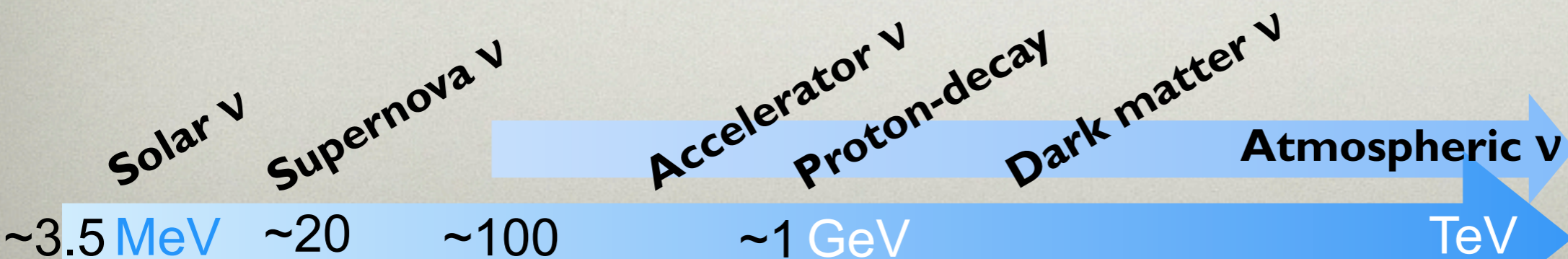
Accelerator



Atmospheric

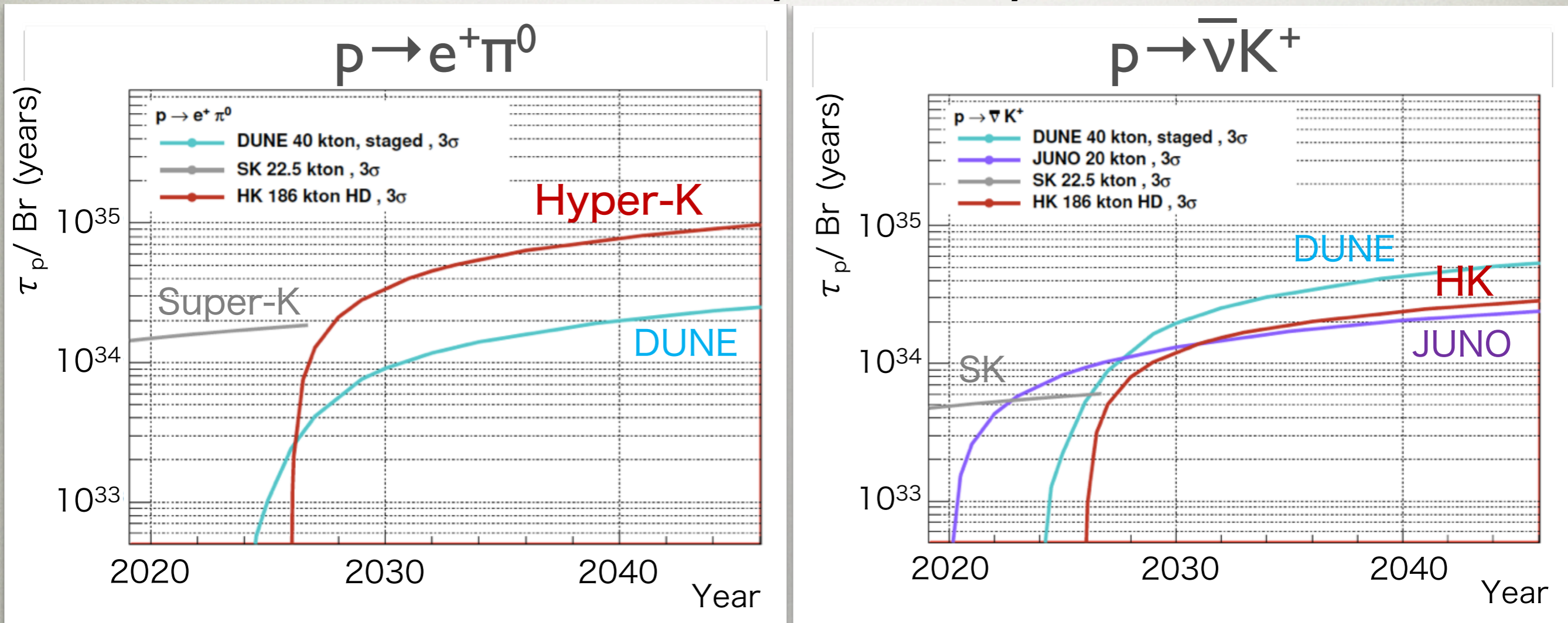


p-decay



# SEARCH FOR PROTON DECAY

## $3\sigma$ discovery sensitivity



- Hyper-K will explore 10 times longer proton-lifetime than current running experiment
  - ex.  $p \rightarrow e^+ \pi^0$ ,  $p \rightarrow \bar{\nu} K^+$
- Many other decay modes can also be searched with an order of magnitude better sensitivity

# Status of Hyper-K

arXiv:1805.04163v1



## Hyper-Kamiokande

Design Report  
(Dated: May 9, 2018)

### Inaugural Symposium (MoU), Jan. 2015



### NNSO Inaugural Ceremony, Oct. 2017





# Status of Hyper-K

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- MEXT (funding agency) selected Hyper-K in the ‘Roadmap 2017’

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- Science Council of Japan selected Hyper-K as one of the top priority large-scale projects in ‘Master Plan 2017’
- MEXT (funding agency) selected Hyper-K in the ‘Roadmap 2017’
- U.Tokyo launched ‘Next-Generation Neutrino Science Organization’ (NNSO) for Hyper-K construction

arXiv:1805.04163v1



## Hyper-Kamiokande

Design Report  
(Dated: May 9, 2018)

Inaugural Symposium (MoU), Jan. 2015



NNSO Inaugural Ceremony, Oct. 2017



# Status of Hyper-K

- ‘Hyper-K Design Report’ released
  - arXiv:1805.04163, KEK, ICRR preprints
- Two host institutes: U.Tokyo/ICRR and KEK/IPNS (MoU for Hyper-K)
- Science Council of Japan selected Hyper-K as one of the top priority large-scale projects in ‘Master Plan 2017’
- MEXT (funding agency) selected Hyper-K in the ‘Roadmap 2017’
- U.Tokyo launched ‘Next-Generation Neutrino Science Organization’ (NNSO) for Hyper-K construction
- U.Tokyo making all efforts to get Hyper-K funded, aiming to begin the detector construction in JFY2019 and begin the operation in JFY2026

arXiv:1805.04163v1



## Hyper-Kamiokande

Design Report  
(Dated: May 9, 2018)

Inaugural Symposium (MoU), Jan. 2015



NNSO Inaugural Ceremony, Oct. 2017



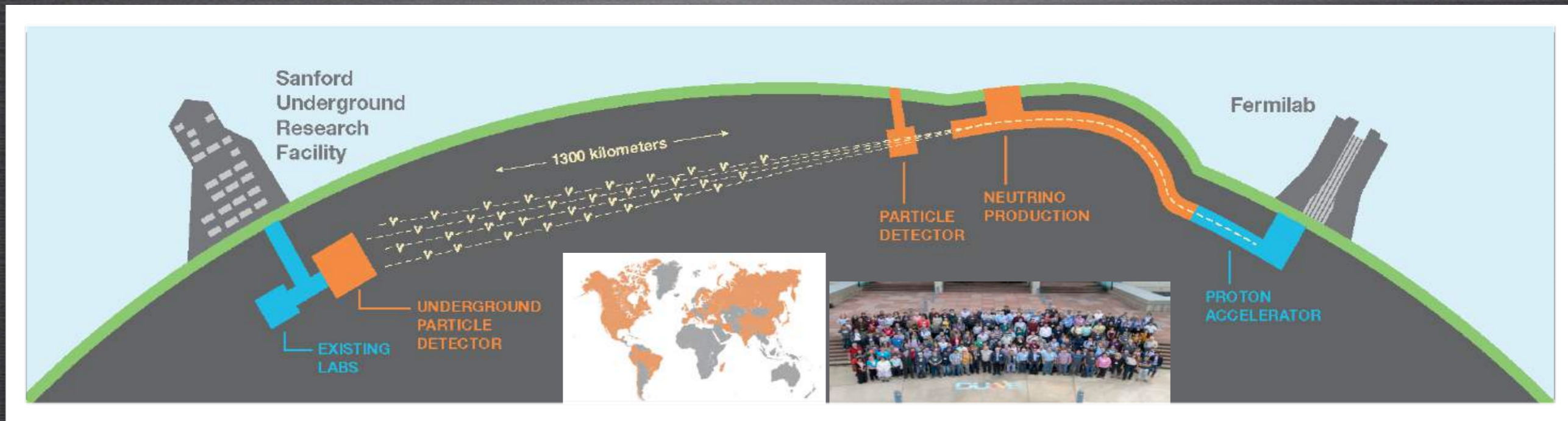
# Hyper-K proto-collaboration



- ~300 collaborators
- 73 institutions from 15 countries
- ~75% of collaborators from abroad

**Open for new collaborators**



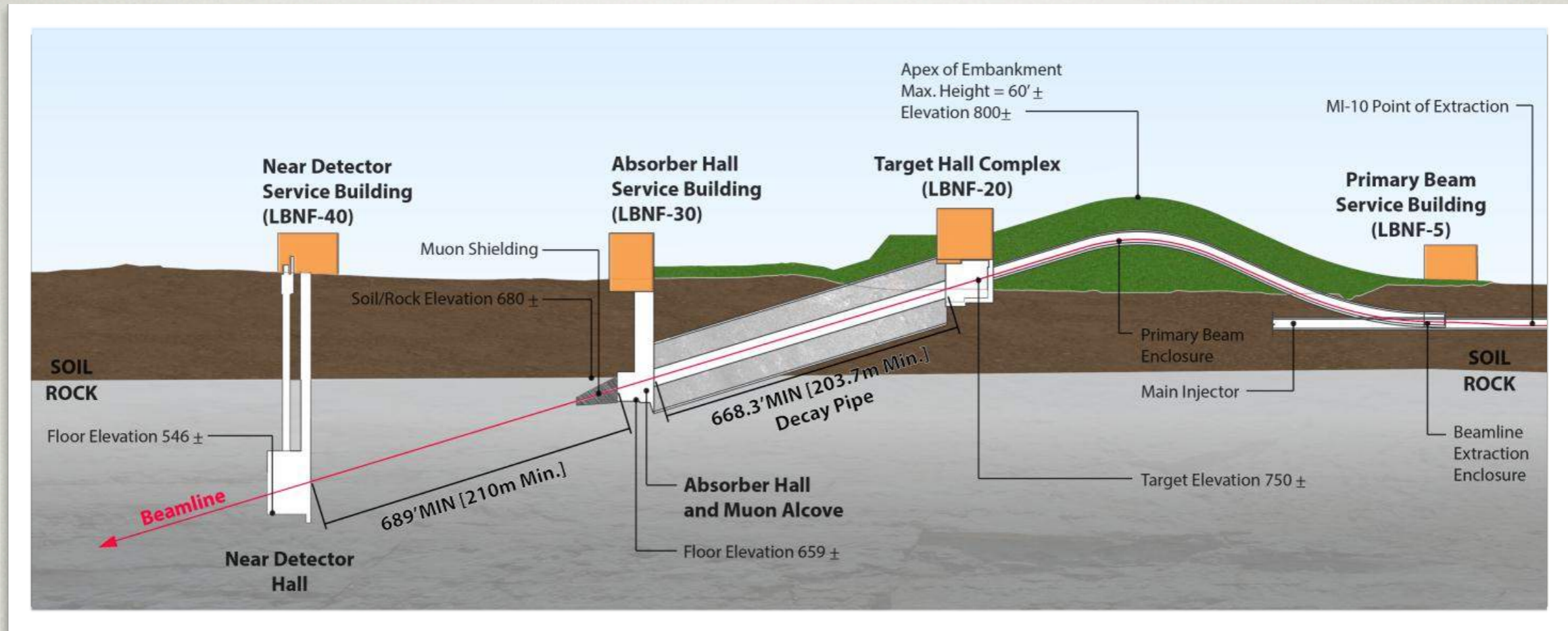


# DUNE/LBNF

E. Worcester's slides at Neutrino 2018  
J. Bian's slides at ICHEP 2018



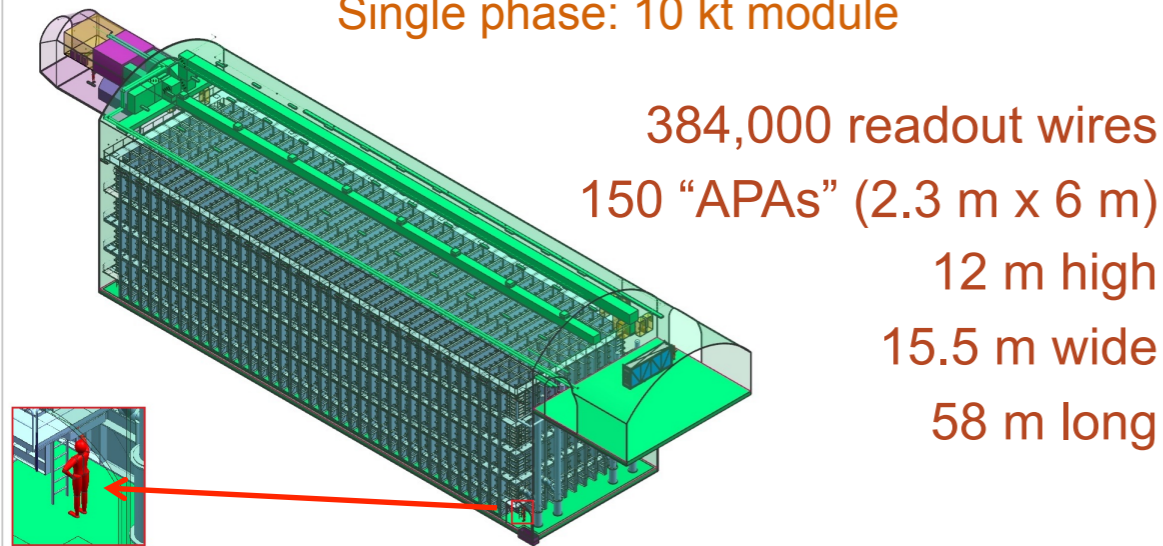
# LBNF BEAM



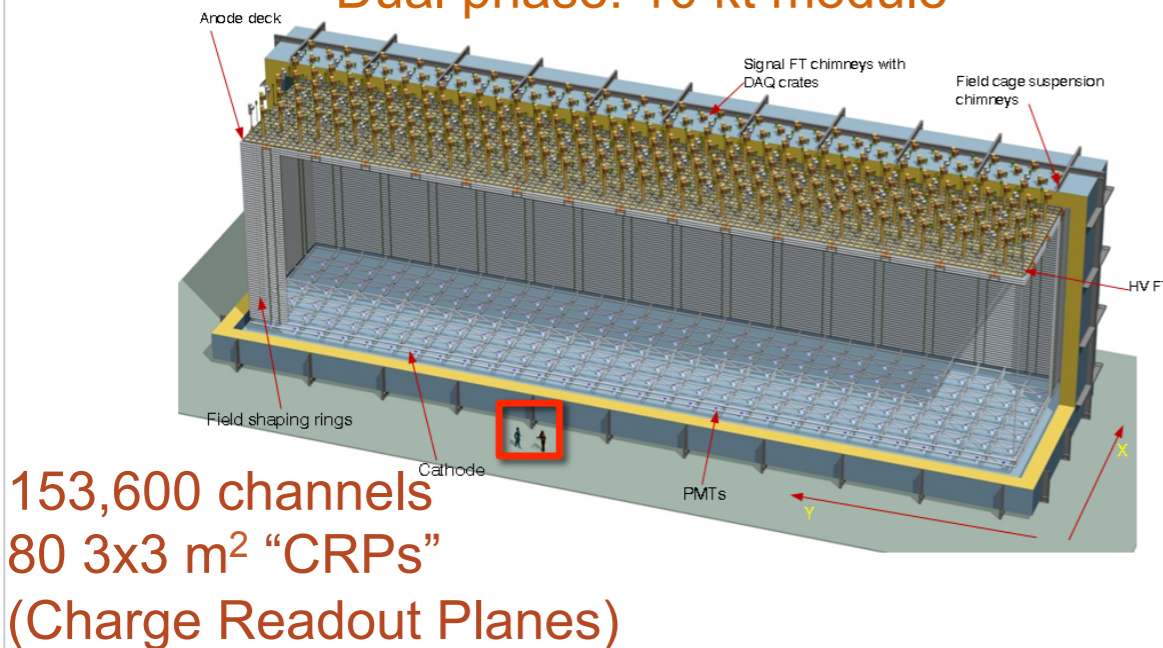
- 60-120 GeV proton beam
- 1.2 MW proton beam power @80GeV
  - cf. NuMI achieved 700kW for NOvA
  - Upgradeable to 2.4 MW
  - DUNE sensitivity studies adopt the beam power upgrade
- Reference design similar to NuMI, optimized to improve sensitivity to oscillation measurement

# DUNE DETECTOR

Single phase: 10 kt module



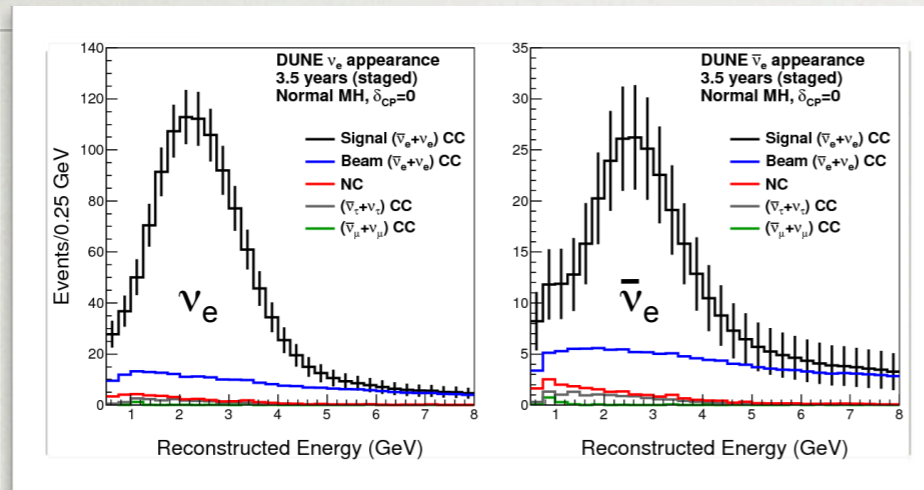
Dual phase: 10 kt module



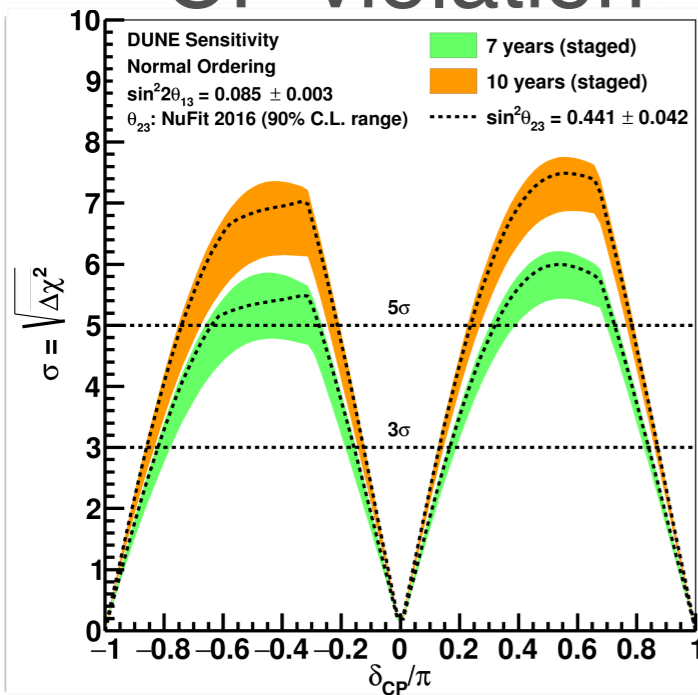
- Large underground Liquid Argon (LAr) Time Projection Chamber
- Single- and dual-phase LAr TPC
- Total fiducial mass of 40 kton
  - 10 kton × 4 modules
- Detector modules installed in stages
  - Start with 20kton (FV) and add other 20kton in 4 years
  - 1st modules single-phase
- Two prototype TPCs under construction for a test-beam at CERN (770ton LAr each)

# DUNE SENSITIVITY FOR CPV AND MH

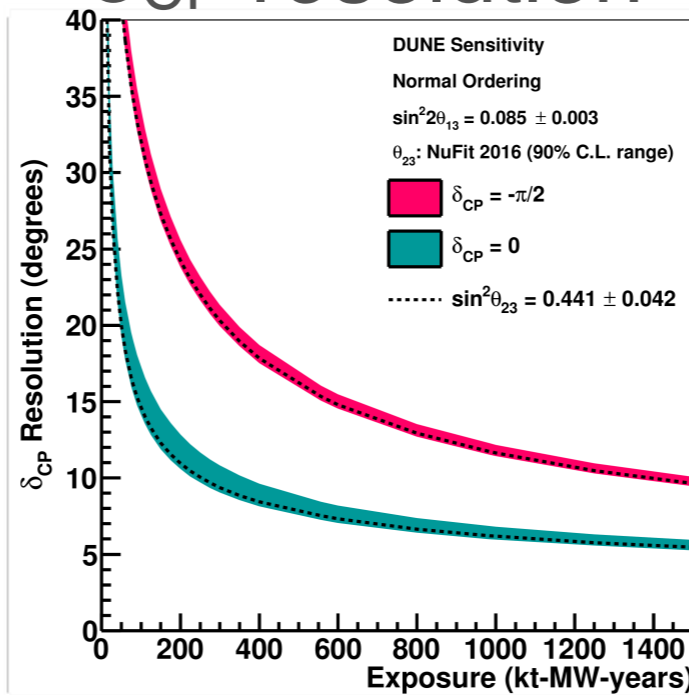
Order 1000  $\nu_e$  appearance events in  $\sim 10$  years of equal running in  $\nu$  and  $\bar{\nu}$  mode



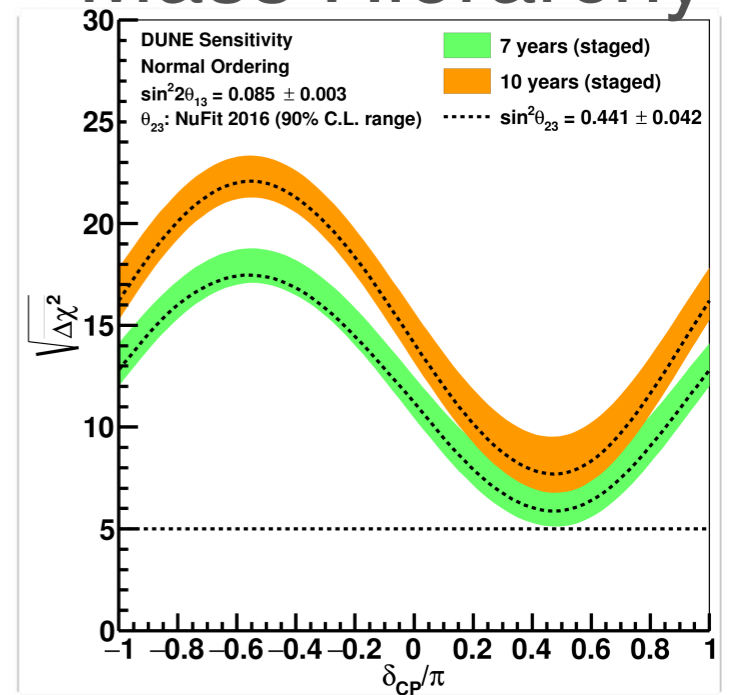
## CP violation



## $\delta_{CP}$ resolution



## Mass Hierarchy



- Mass Hierarchy determination at  $>5\sigma$
- 75% coverage for  $3\sigma$  CPV discovery
- $7^\circ \sim 15^\circ$   $\delta_{CP}$  resolution

- \* Efficiency tuned using hand scan results
- \* Syst. uncertainties approximated using normalization uncertainties

# Status/plan of DUNE

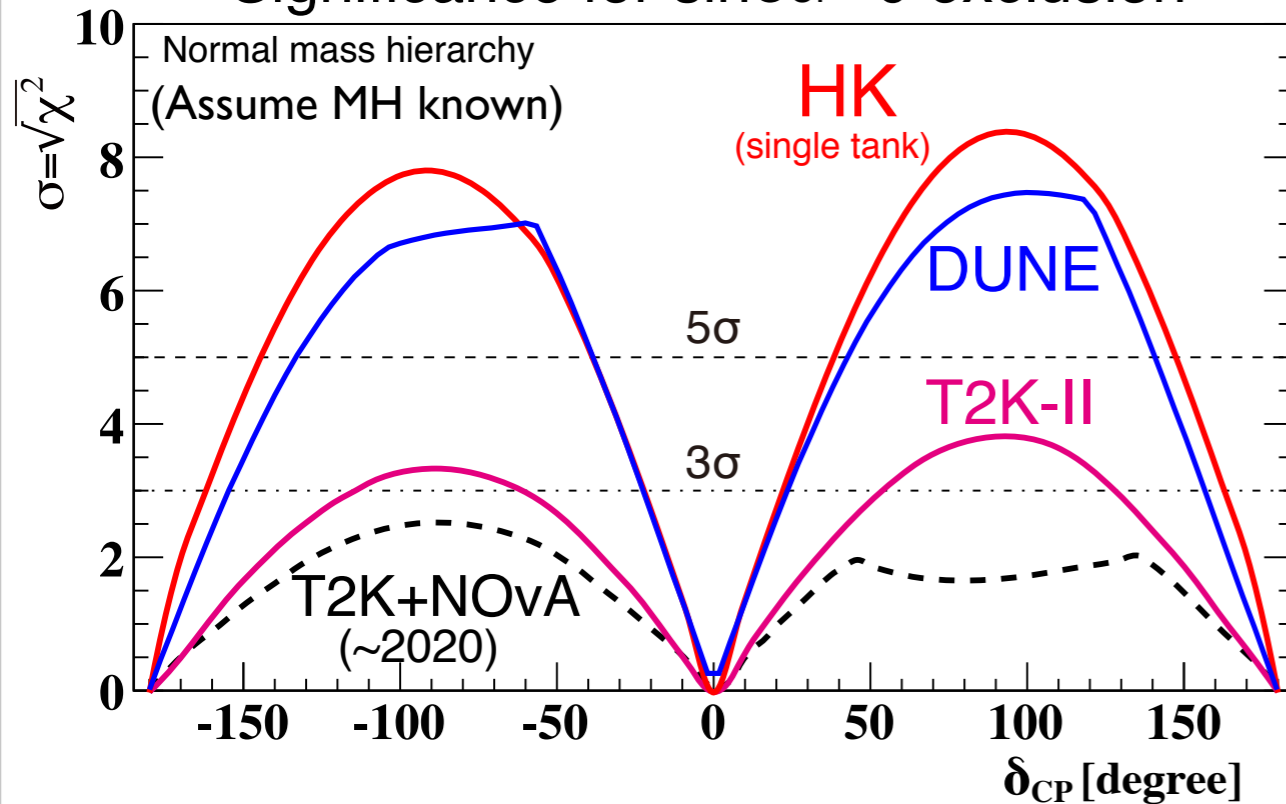
Quoted from J. Bian's slides at ICHEP 2018

- 2017: Far site construction begins
- 2018: Start to operate full-scale protoDUNE at CERN
- 2019: DUNE Technical Design Report ready for funding agencies
- 2019: Main Cavern Excavation
- 2020: Far Detector fabrication facilities ready
- 2022: Start to install FD modules
- 2024: Two FD modules (20kton) operational
- 2026: Beam on with two FD modules

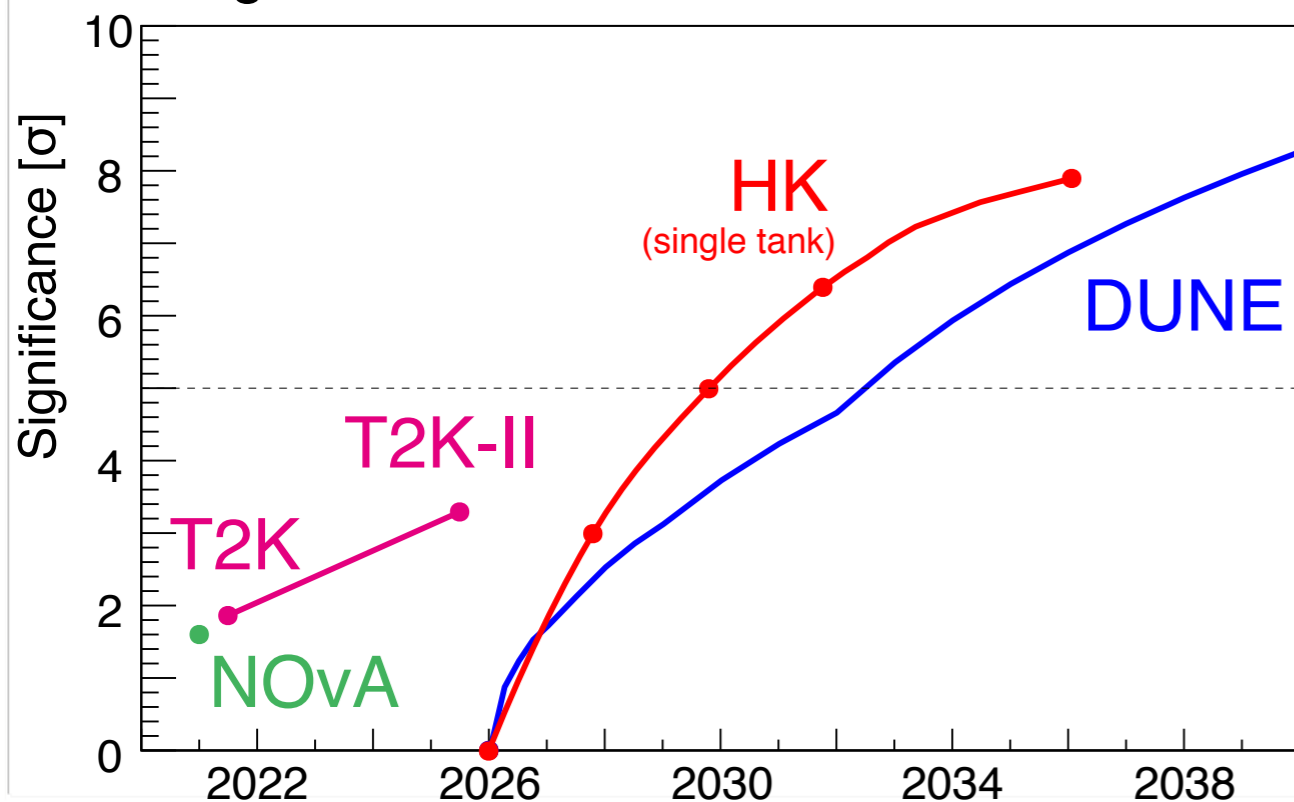


# SUMMARY OF CPV SENSITIVITY

Significance for  $\sin\delta_{CP}=0$  exclusion



CPV significance for  $\delta_{CP}=-90^\circ$ , normal hierarchy



- Significance for  $\sin\delta=0$  exclusion
- Hyper-K and DUNE both 5 $\sigma$  sensitivity near  $\delta=-90^\circ$  after 10ys operation
- Next generation expts have way better sensitivity than current running expts

- Hyper-K DR: [arXiv:1805.04163](https://arxiv.org/abs/1805.04163)
- DUNE CDR: [arXiv:1512.06148](https://arxiv.org/abs/1512.06148)

# SUMMARY

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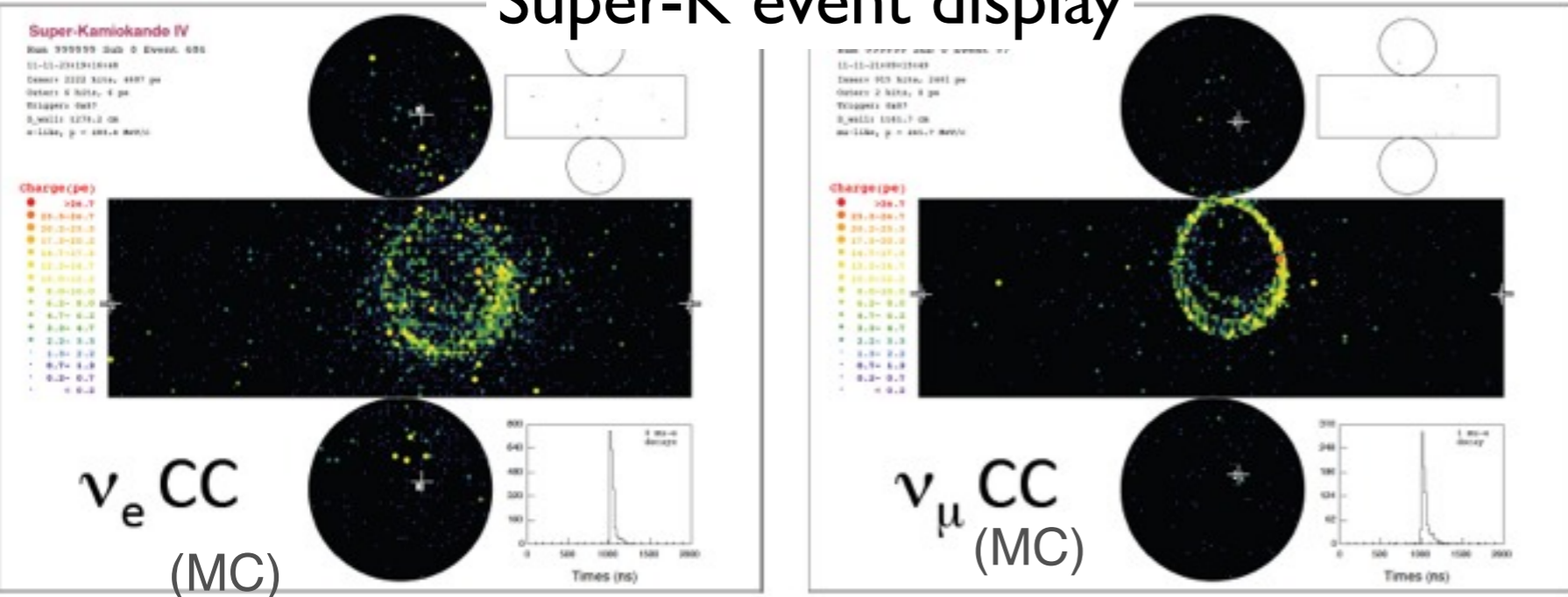
- Current running LBL experiments (T2K, NOvA) playing a major role to improve our understanding of neutrino oscillation
  - Maybe, T2K and NOvA are finding a hint of neutrino CP violation and Mass Hierarchy?
  - Non-accelerator-based experiments (Super-K, IceCUBE, ORCA, JUNO, INO, ...) also have a great potential for Mass Hierarchy determination
- Next generation LBL experiments, Hyper-K and DUNE, aim to reveal a full picture of  $\nu$  oscillation
  - Primary targets: CPV and Mass Hierarchy
  - Aim to start data taking with  $\nu$  beam in 2026
  - Significant complementarily
- Wide physics topics, many discovery potential
  - Proton decay, astrophysical neutrino, ...

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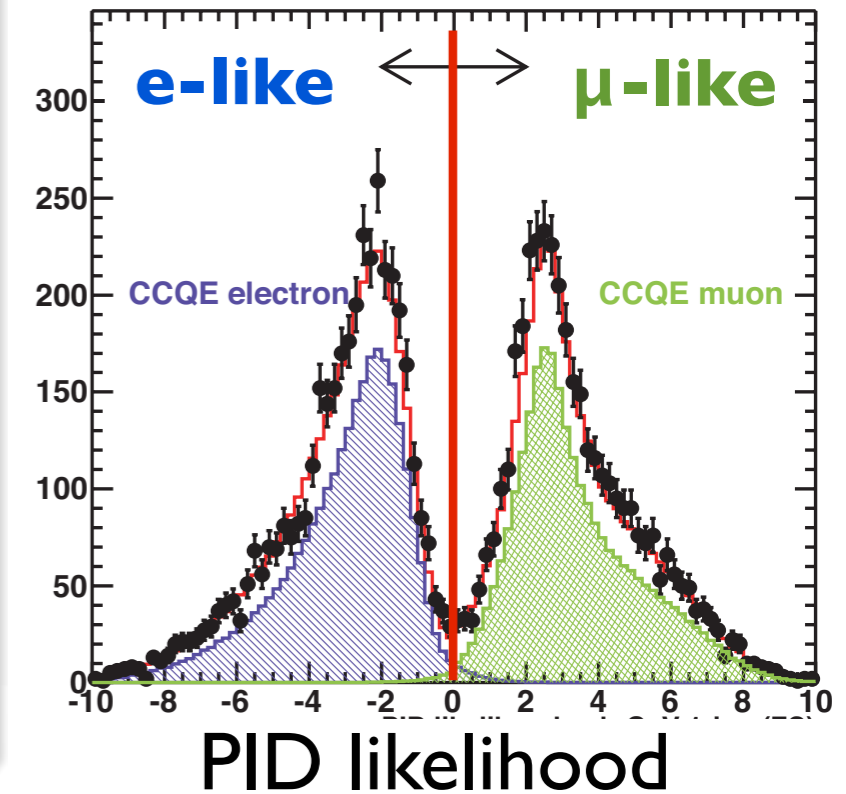
**BACK UP**

# Detector performance

## Super-K event display

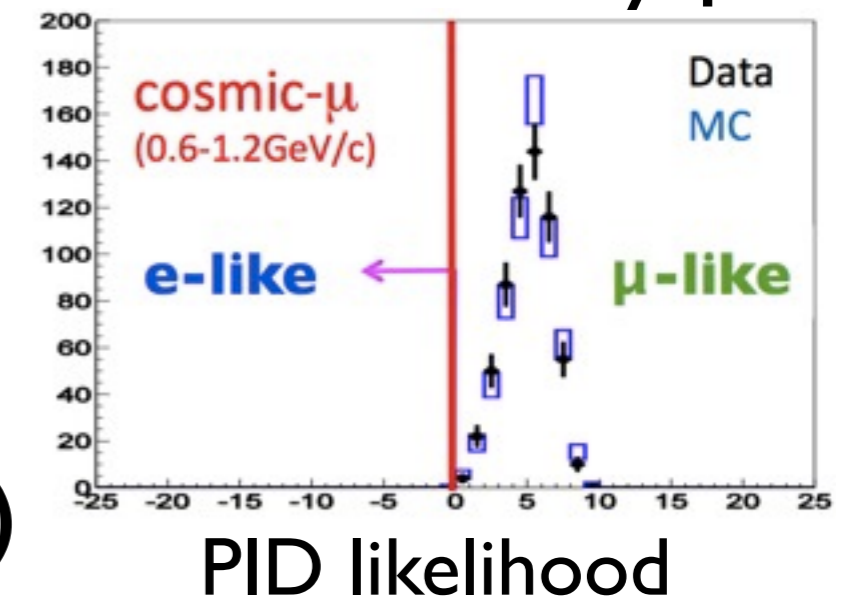


## SK atmospheric $\nu$



- **Large mass** ( $\sim 10 \times$  Super-K FV)
- Statistics is always critical
- **Excellent particle ID (e/ $\mu$ )**
  - Mis-identification  $< 1\%$
- Energy resolution e/ $\mu \sim 3\%$
- Quasi-elastic is dominant (sub-GeV)
  - Clean one-ring event

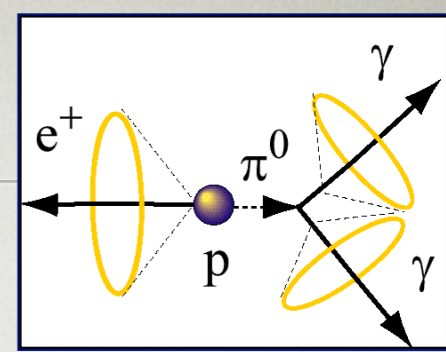
## SK cosmic-ray $\mu$



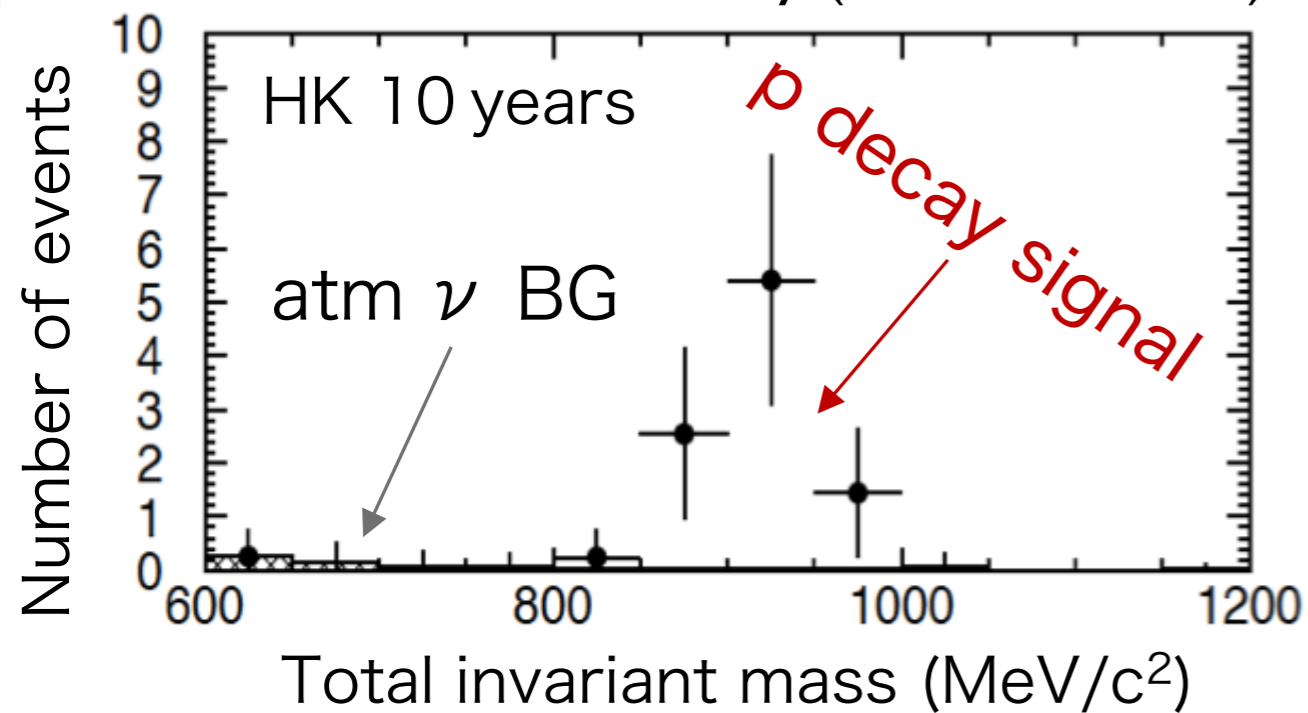
mis-PID:  
 Data:  $0.00 \pm 0.16$  (stat.)%  
 MC :  $0.10 \pm 0.10$  (stat.)%



# $p \rightarrow e^+ \pi^0$ search in Hyper-K

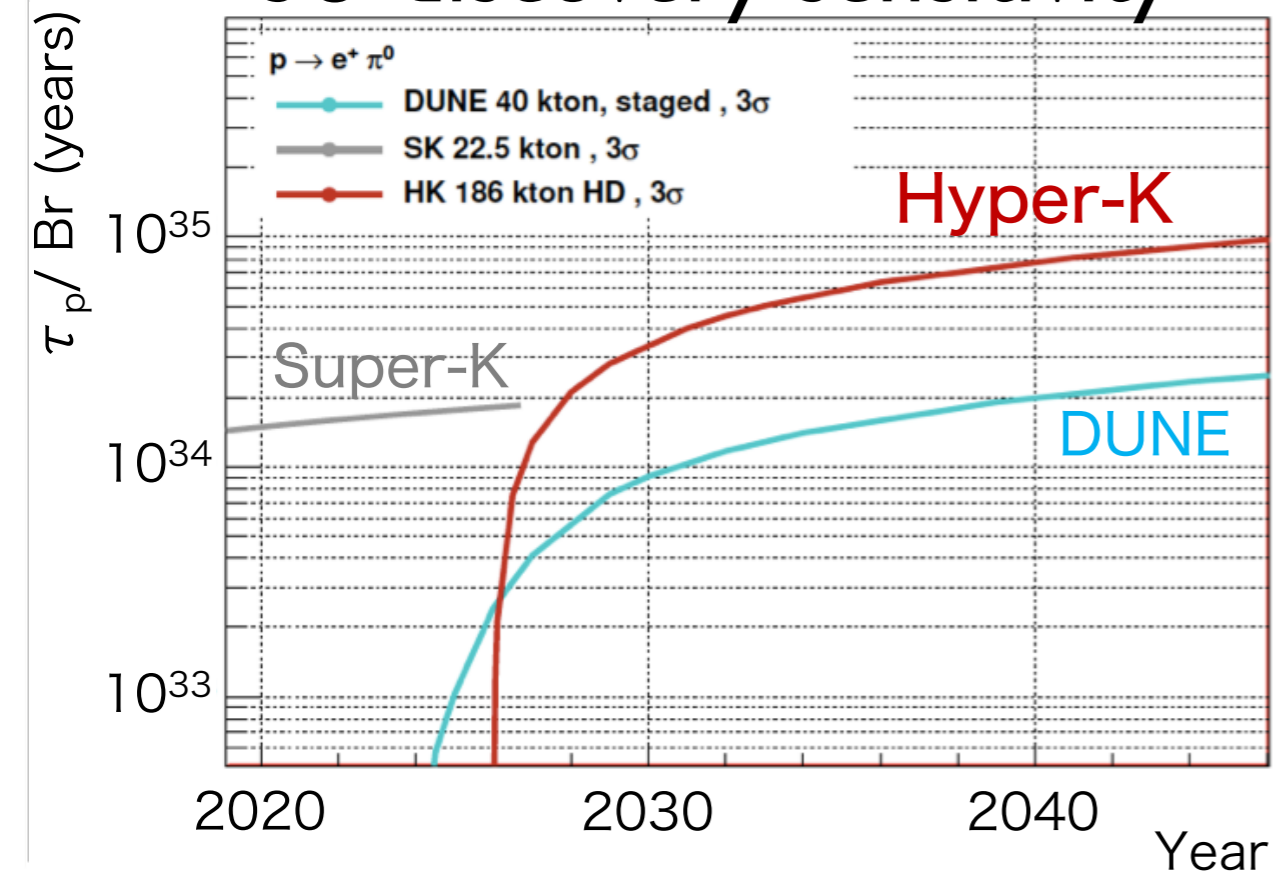


Assume  $\tau/\text{Br} = 1.7 \times 10^{34} \text{y}$  (SK 90%CL limit)



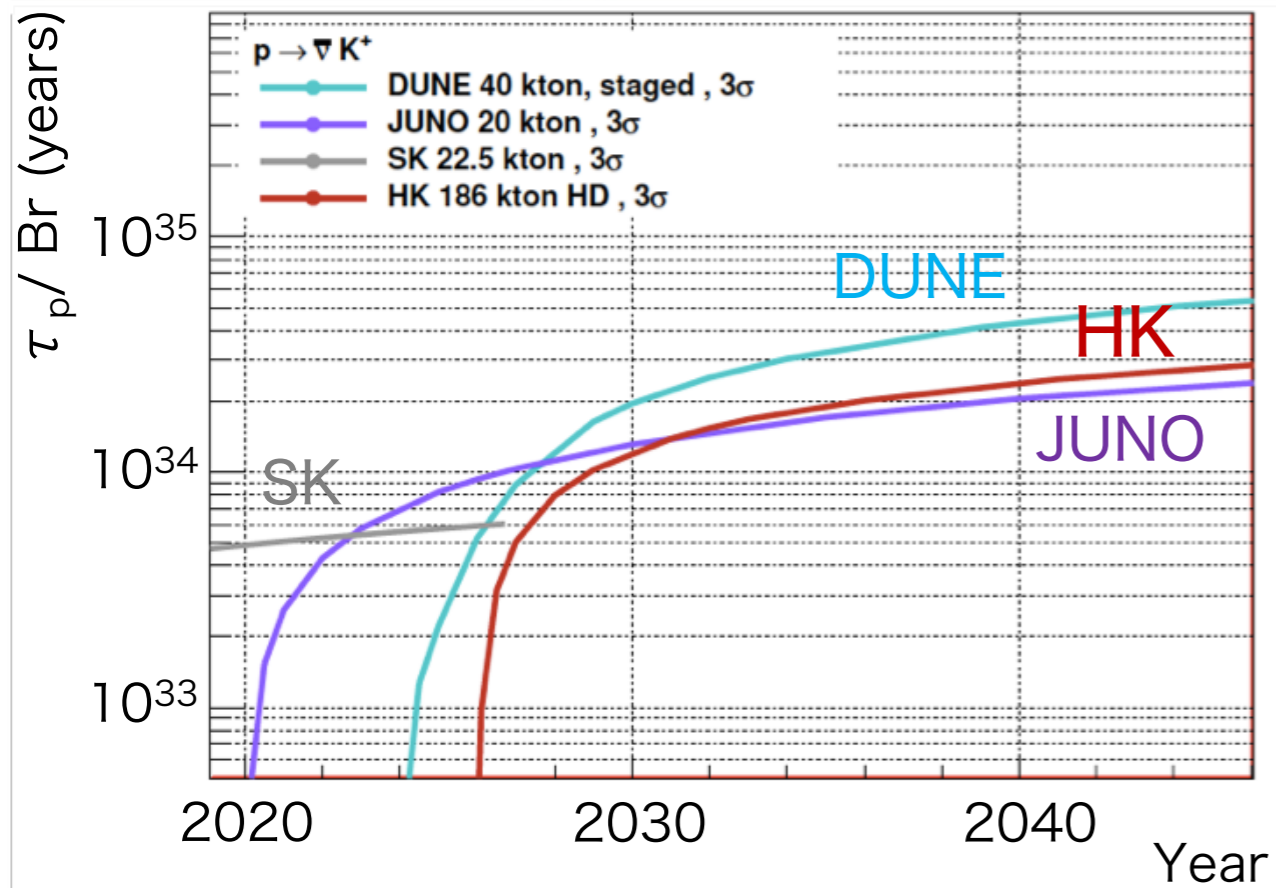
- “Background free” meas. of proton decay
- 0.06 Bkg events / Mt·year
- Bkg atm- $\nu$  events are largely reduced by ‘neutron-tag’: eff.~70% with new PMT
- $n+p \rightarrow d+\gamma$  (2.2MeV)

## 3 $\sigma$ discovery sensitivity



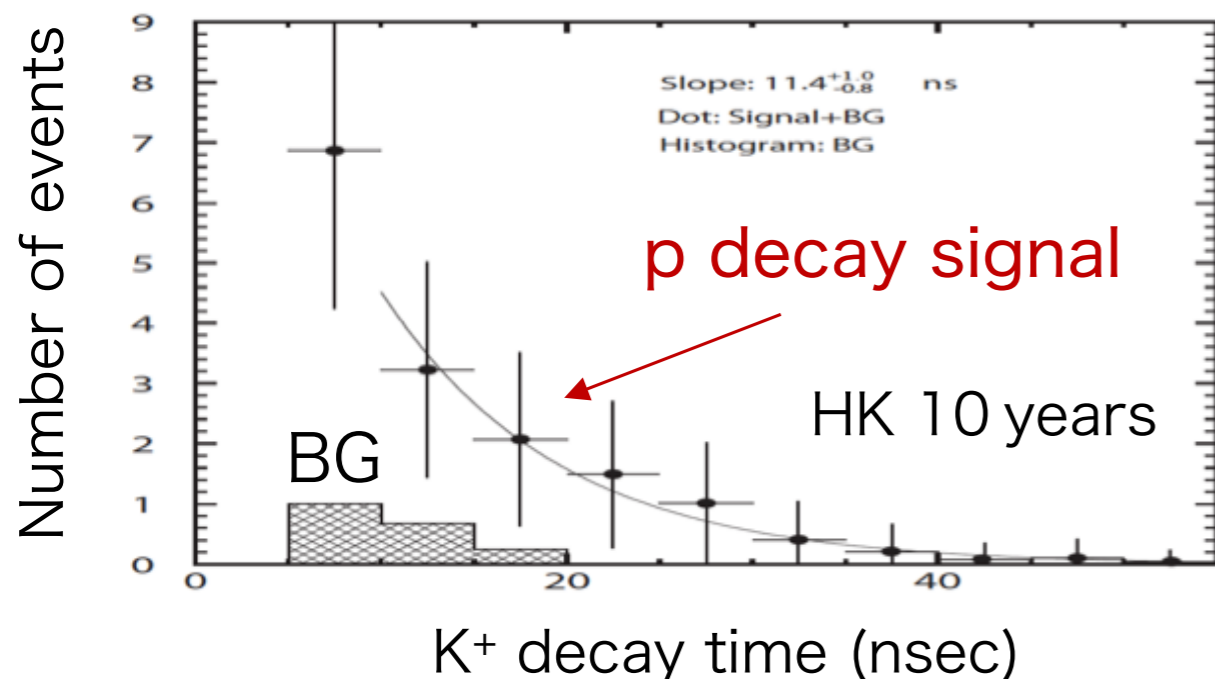
- Great discovery potential
- 3 $\sigma$  discovery sensitivity reaches  $\tau_p/\text{Br} = 10^{35}$  years

# $p \rightarrow \bar{\nu} K^+$ search in Hyper-K

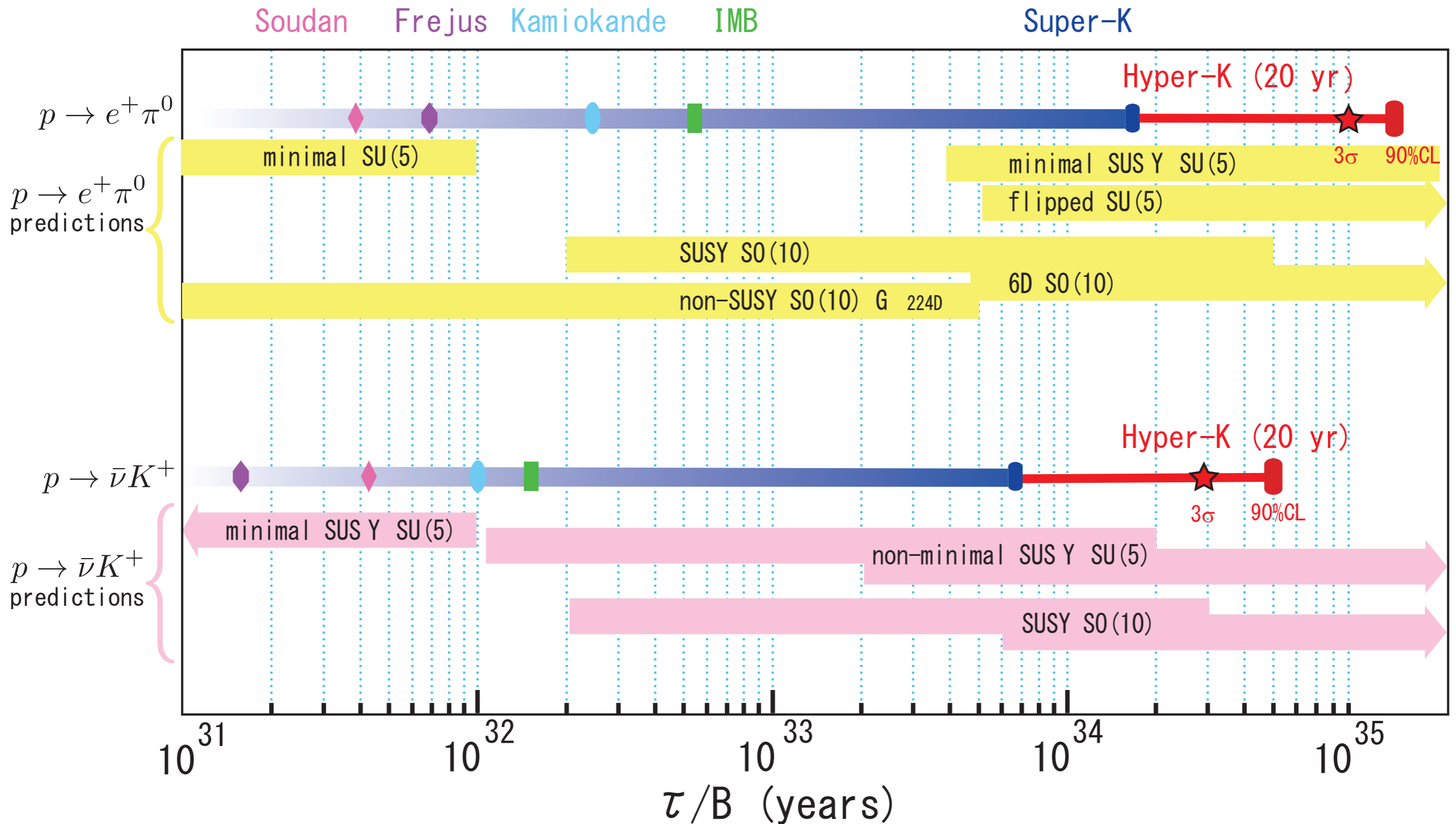


- Identify  $K^+$  by its decaying products
- $K^+ \rightarrow \mu + \nu$  (Br: 64%)
  - 236 MeV/c  $\mu^+$
  - de-excitation  $\gamma$  from  $^{15}O^*$  (6 MeV  $\gamma$ )
- $K^+ \rightarrow \pi^+ \pi^0$  (Br: 21%)
  - 205 MeV/c  $\pi^0$  &  $\pi^+$  back-to-back
- **New PMT improves signal and background efficiencies**
- **Other decay modes,  $I + \omega$ ,  $\rho$ ,  $\eta$ ,  $\chi$  improved than SK**

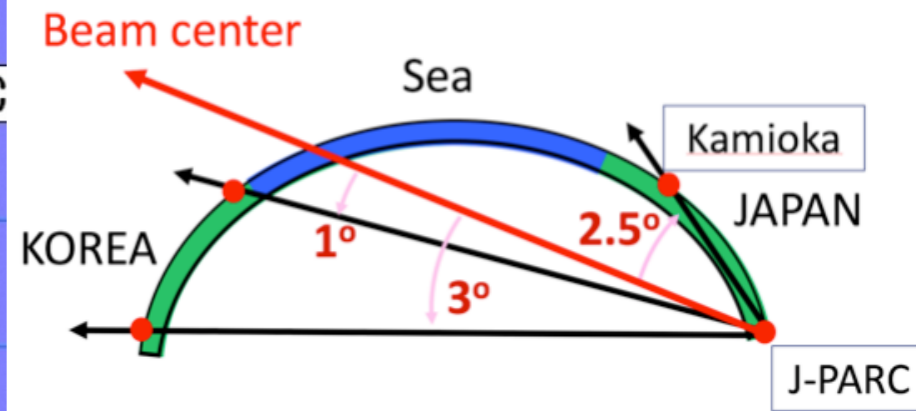
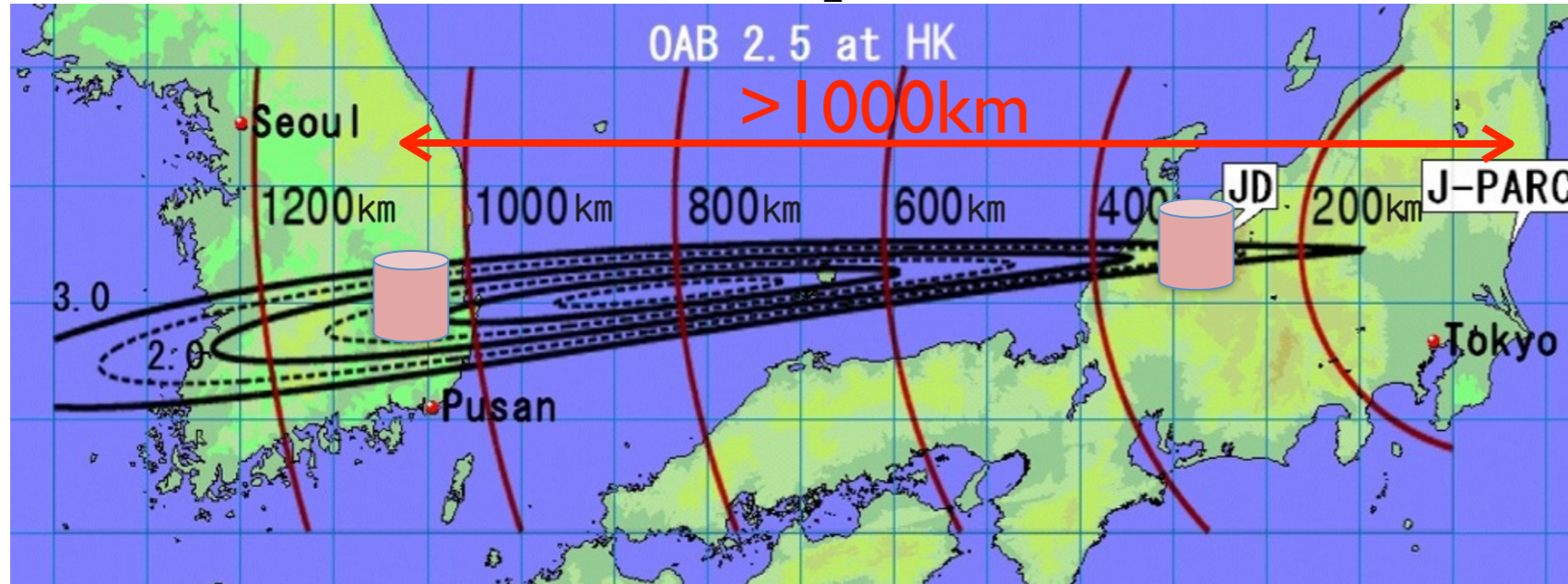
Assume  $\tau/Br = 6.6 \times 10^{33} y$  (SK 90%CL limit)



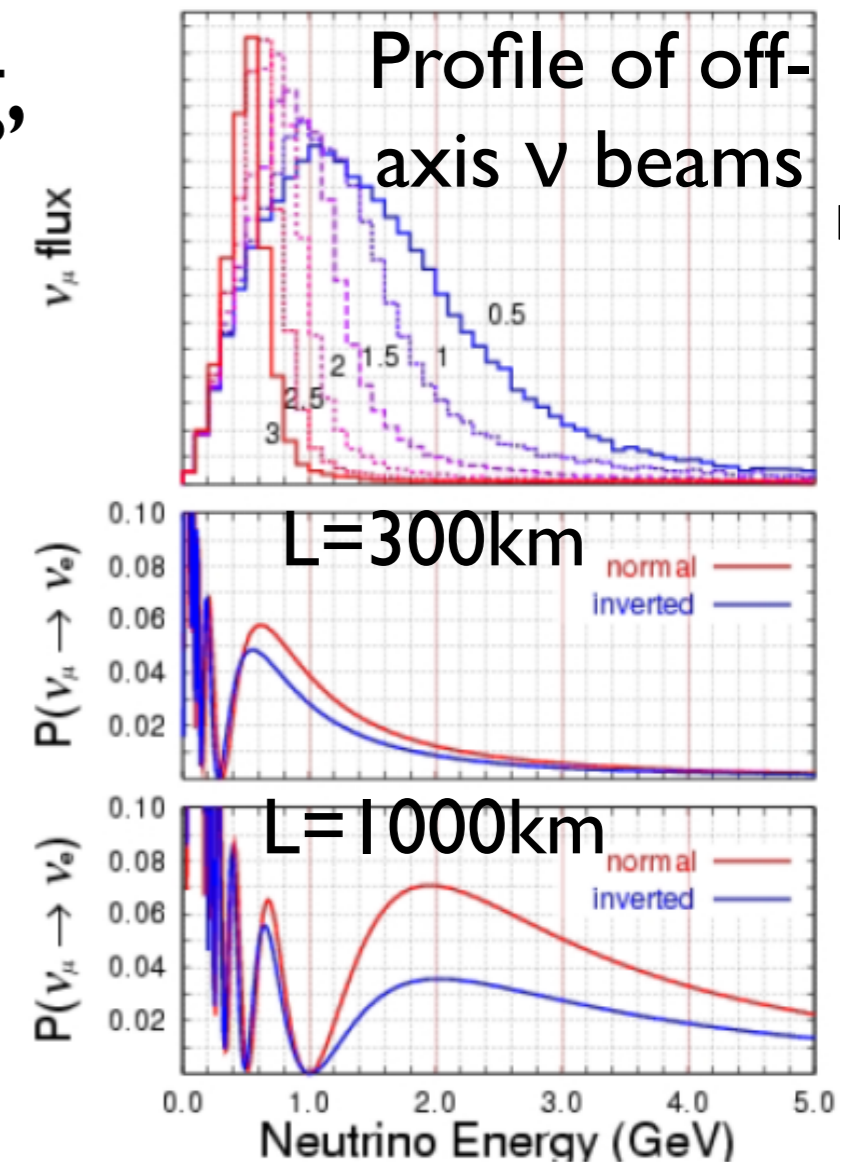
# Predictions & experiments for p-decay



# Korean option for 2nd tank



- Given the two tank design with staging, may benefit from building the second tank at a different baseline
- Option for second tank in Korea is being considered
- Advantage of 2nd detector in Korea:
  - CP effect at second oscillation maximum
  - Mass-hierarchy sensitivity to complement the measurement with atmospheric  $\nu$

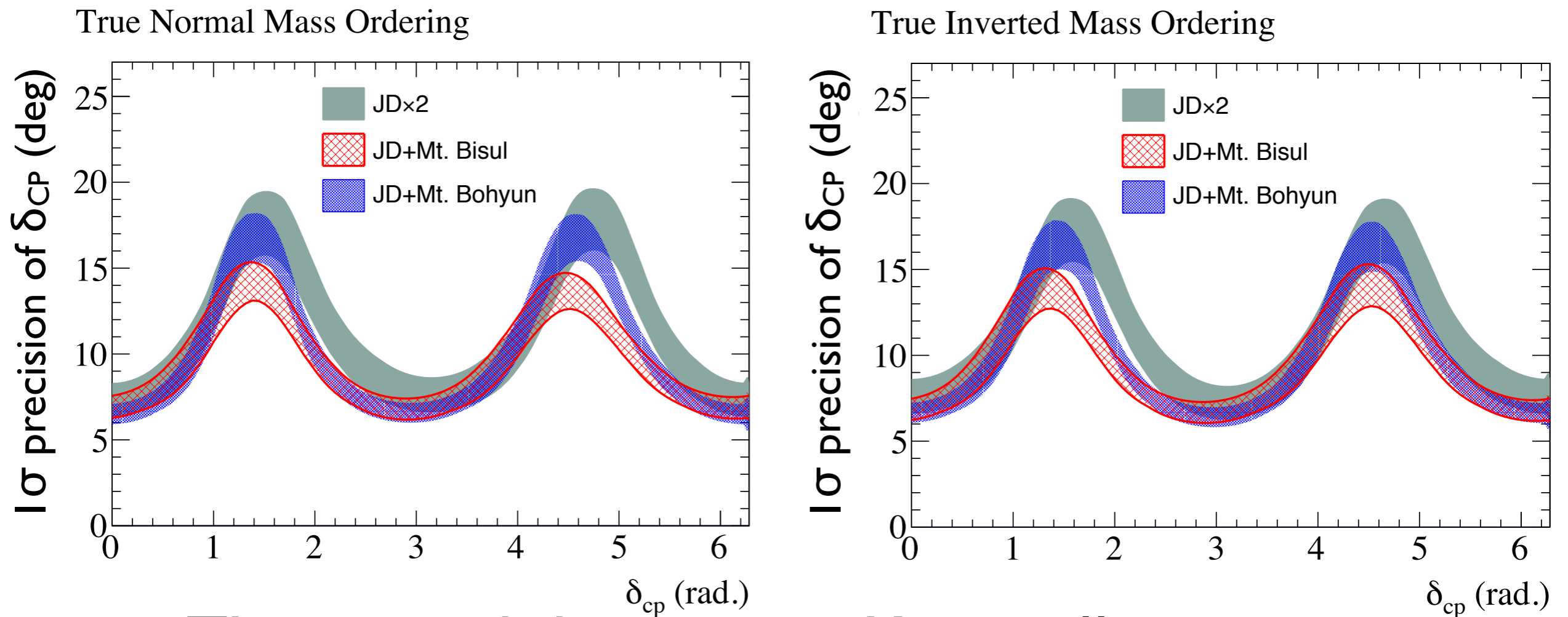


# T2HKK: $\delta_{CP}$ resolution

arXiv:1611.06118

■ Two detectors in Japan  
■ Second detector in Korea, two colours for two different candidate sites

Band indicates range  
 $0.4 < \sin^2 \theta_{23} < 0.6$



The second detector in Korea allows us to better measure the CP-phase, compared with both detectors in Japan

# Intermediate Water Č detector

- Intermediate water Č detector proposed
  - Locate 1~2km downstream of J-PARC neutrino beam
  - J-PARC E61 (NuPRISM)
- Off-axis spanning ( $1^\circ \sim 4^\circ$ ) to probe neutrino energy vs. final state kinematics relationship
- Increase  $\nu_e$  purity at higher angles: larger contribution from Kaon
  - Electron neutrino cross-section
  - Aim to  $\sigma(\nu_e)/\sigma(\nu_\mu)$  cross-section ratio error down to 2~3%

