

Connecting B decay anomalies with neutrino mass, dark matter and flavor violation

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Talk based on [arXiv:1806.10146](#)

With Chandan Hati, Ana Teixeira, and Jean Orloff

Introduction

The Standard Model (SM) is a remarkably successful theory of subatomic world but is not a complete description

A partial list of deficiencies of the SM includes:

Neutrino oscillations

Matter-antimatter Asymmetry

Dark Matter candidate

Many theoretical motivations...

Introduction

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Dark Matter candidate

Many theoretical motivations...

No direct evidence of new particle at LHC so far

Several intriguing signs of NP in semileptonic B decays!



Need to look for
beyond the SM Physics

LFU violation in B decays ?

In the SM, couplings of gauge bosons to leptons are independent of lepton flavour

Define lepton flavor universal observable:

$$R_{K^{(*)}} = \frac{\text{BR}(B \rightarrow K^{(*)}\mu^+\mu^-)}{\text{BR}(B \rightarrow K^{(*)}e^+e^-)}$$

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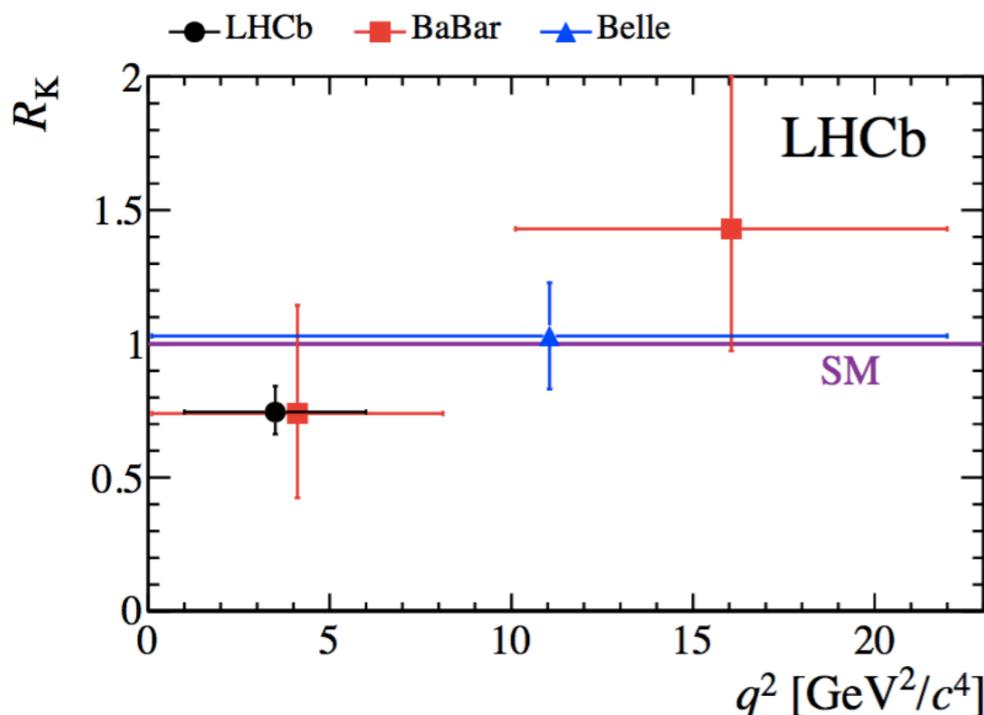
~ 2.5 sigma tension in lepton flavour universality ratio in

$$R_{K,[1-6] \text{ GeV}^2} = 0.745^{+0.090}_{-0.074} \pm 0.036$$

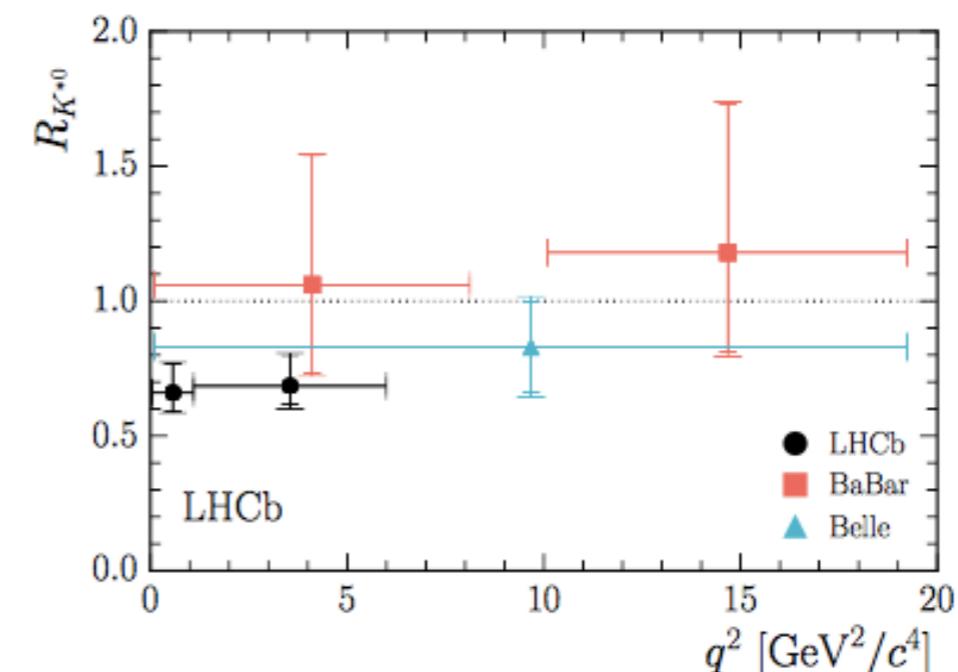
$$R_{K^*[1.1,6]} = 0.69^{+0.11}_{-0.07} \pm 0.05,$$

$$R_{K^*[0.045,1.1]} = 0.66^{+0.11}_{-0.07} \pm 0.03$$

LHCb PRL 113 (2014) 151601



LHCb, JHEP 1708, 055 (2017)



Hints of New Physics ?

Claims of NP are further strengthened by other data on $b \rightarrow s\mu^+\mu^-$

$B \rightarrow K^*\mu^+\mu^-$: Measurement of form factor free observables

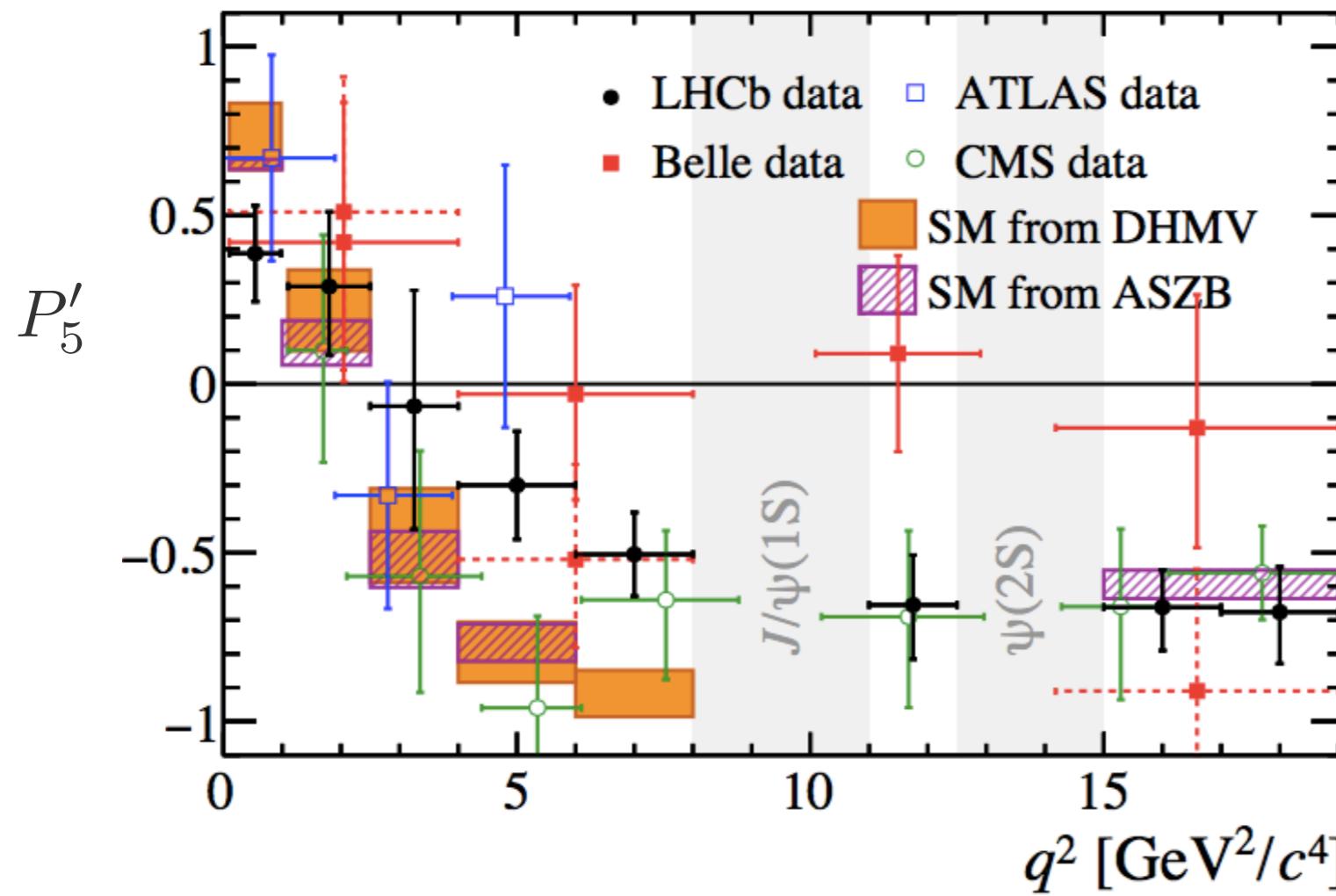


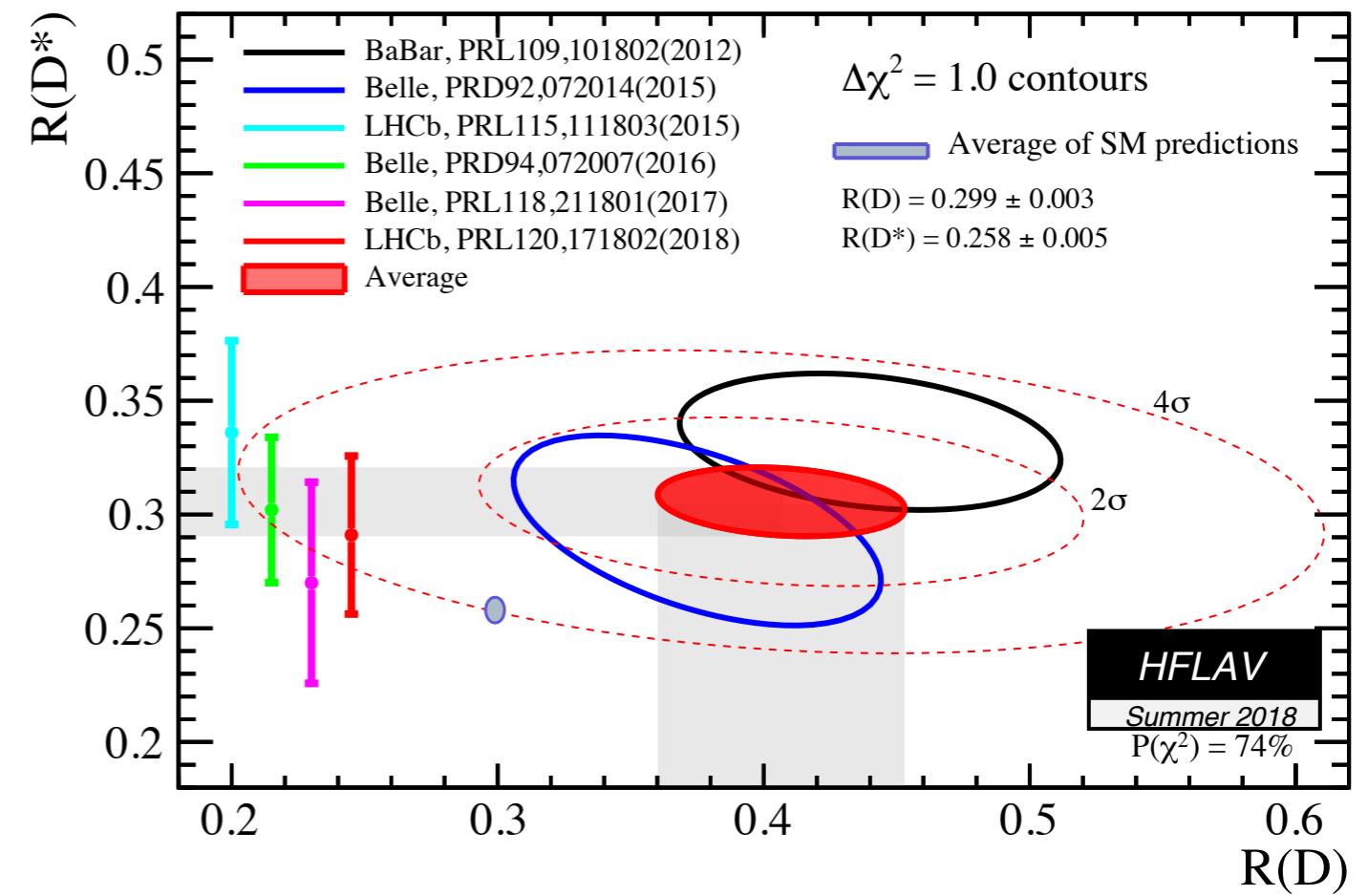
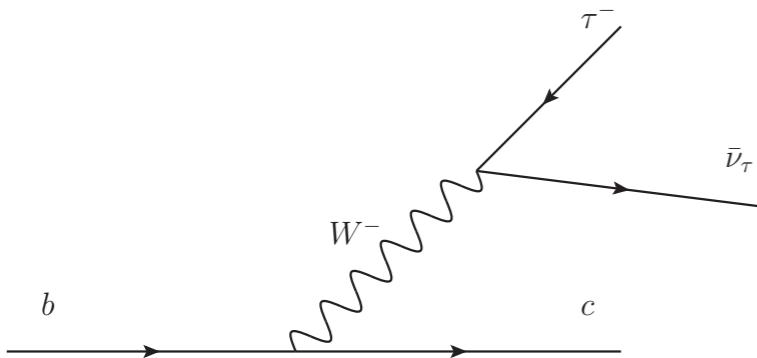
Fig. from T. Gershon's talk,
Moriond 2017

In BR of $B \rightarrow \phi\mu^+\mu^-$, a deviation of 3.5 sigma with respect to SM

LFU violation in B decays ?

~4 sigma deviation from the SM in the ratio

$$R_{D^{(*)}} = \frac{\text{BR}(\bar{B} \rightarrow D^{(*)}\tau\bar{\nu})}{\text{BR}(\bar{B} \rightarrow D^{(*)}\ell\bar{\nu})}; \quad \ell = e, \mu$$



Also, in $B_c^+ \rightarrow J/\psi \tau^+ \nu_\tau$

$$R_{J/\psi} = 0.71 \pm 0.17 \pm 0.18$$

About 2σ above the SM

PRL 120, 121801 (2018)

(Heavy Flavor Averaging group, HFLAV)

The Model

	Field	$SU(3)_C \times SU(2)_L \times U(1)_Y$	Z_2
Fermions	$Q_L \equiv (u, d)_L^T$	(3, 2, 1/6)	1
	u_R	(3, 1, 2/3)	1
	d_R	(3, 1, -1/3)	1
	$\ell_L \equiv (\nu, e)_L^T$	(1, 2, -1/2)	1
	e_R	(1, 1, -1)	1
	Σ_R	(1, 3, 0)	-1
Scalars	H	(1, 2, 1/2)	1
	h_1	(3, 3, -1/3)	1
	h_2	(3, 3, -1/3)	-1

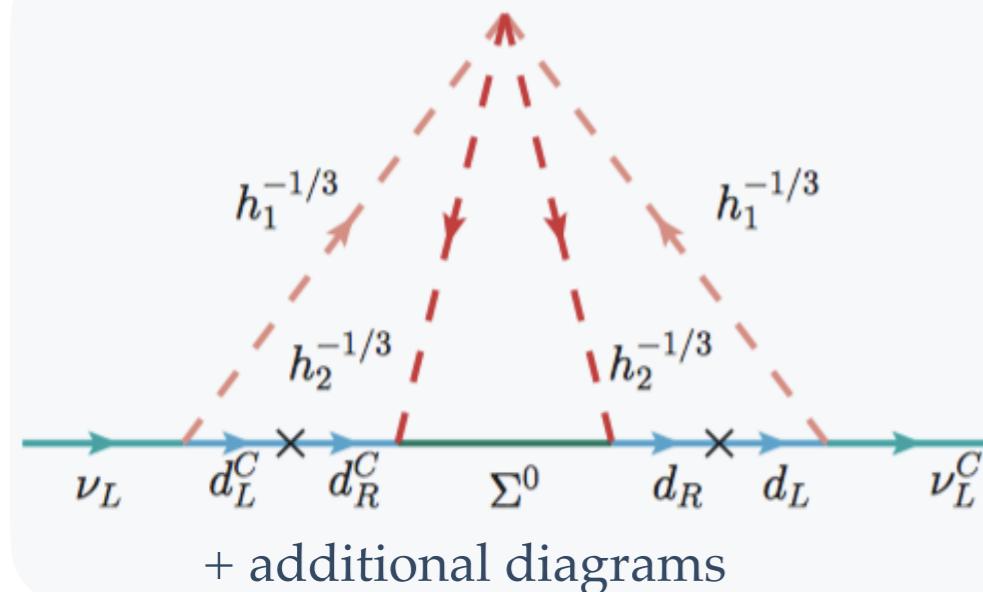
Reinforced by a Z_2 symmetry :

Forbids Type III seesaw,
provides DM candidate

$$\mathcal{L}_{\text{int}}^{h,\Sigma} = y_{ij} \bar{Q}_L^{Ci} \epsilon (\vec{\tau} \cdot \vec{h}_1) L_L^j + \tilde{y}_{ij} \overline{(\vec{\tau} \cdot \vec{\Sigma})_R}^{C i, ab} [\epsilon (\vec{\tau} \cdot \vec{h}_2) \epsilon^T]^{ba} d_R^j$$

$$-\frac{1}{2} \overline{\Sigma^C}^i M_{ij}^\Sigma \Sigma^j - V_{\text{scalar}}^{H,h} + \text{h.c.}$$

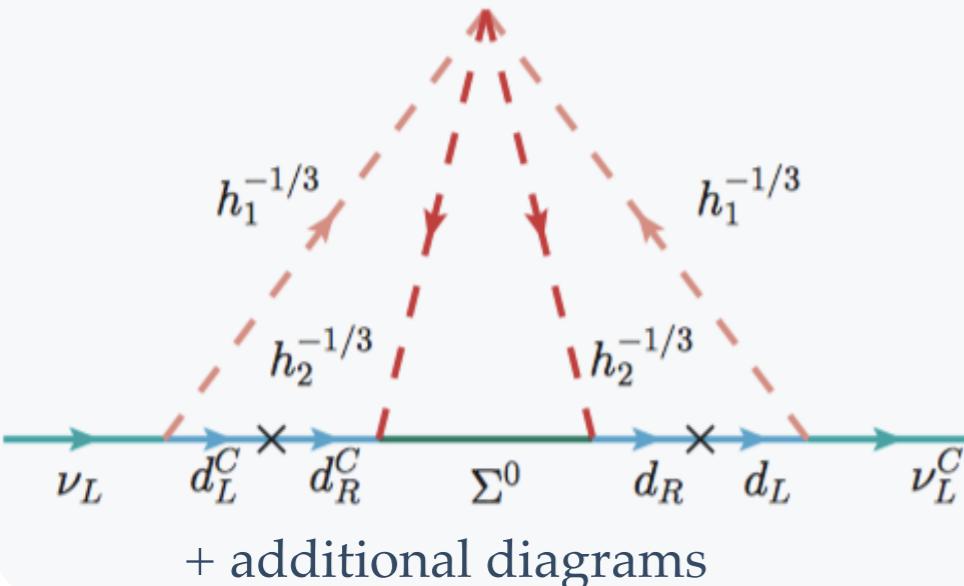
Neutrino mass generation



$$(m_\nu)_{\alpha\beta} = - 30 \frac{\lambda_h}{(4\pi^2)^3 m_{h_2}} y_{\alpha i}^T m_{D_i} \tilde{y}_{ij}^T G \left(\frac{m_{\Sigma_j}^2}{m_{h_2}^2}, \frac{m_{h_1}^2}{m_{h_2}^2} \right) \tilde{y}_{jk} m_{D_k} y_{k\beta}$$

Krauss, Nasri, Trodden, PRD 2003 (with ordinary scalars
with singlet DM in the loop)

Neutrino mass generation



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Modified Casas-Ibarra parametrisation to write \tilde{y} in terms of y :

Casas, Ibarra Nucl. Phys. B 618, 171 (2001)

$$m_\nu^{\text{diag}} = U^T y^T m_D \tilde{y}^T F(\lambda_h, m_\Sigma, m_{h_{1,2}}) \tilde{y} m_D y U$$

U : PMNS matrix

\mathcal{R} : Complex Orthogonal matrix

$$(\sqrt{m_\nu^{\text{diag}}}^{-1} U^T y^T m_D \tilde{y}^T \sqrt{F})(\sqrt{F} \tilde{y} m_D y U \sqrt{m_\nu^{\text{diag}}}^{-1}) = I = \mathcal{R}^T \mathcal{R}$$

$$\tilde{y} = F^{-1/2} \mathcal{R} \sqrt{m_\nu^{\text{diag}}} U^\dagger y^{-1} m_d^{-1}$$

In agreement with Cheung et al, PRD 2017

A Viable Dark Matter Candidate

Recall: Σ_R is odd under Z_2 symmetry

EW radiative corrections : $m_{\Sigma^\pm} - m_{\Sigma^0} \sim 166$ MeV

Cirelli, Fornengo, Strumia
Nucl.Phys.B753, 2006

This renders Σ_0 the lightest stable particle among new states

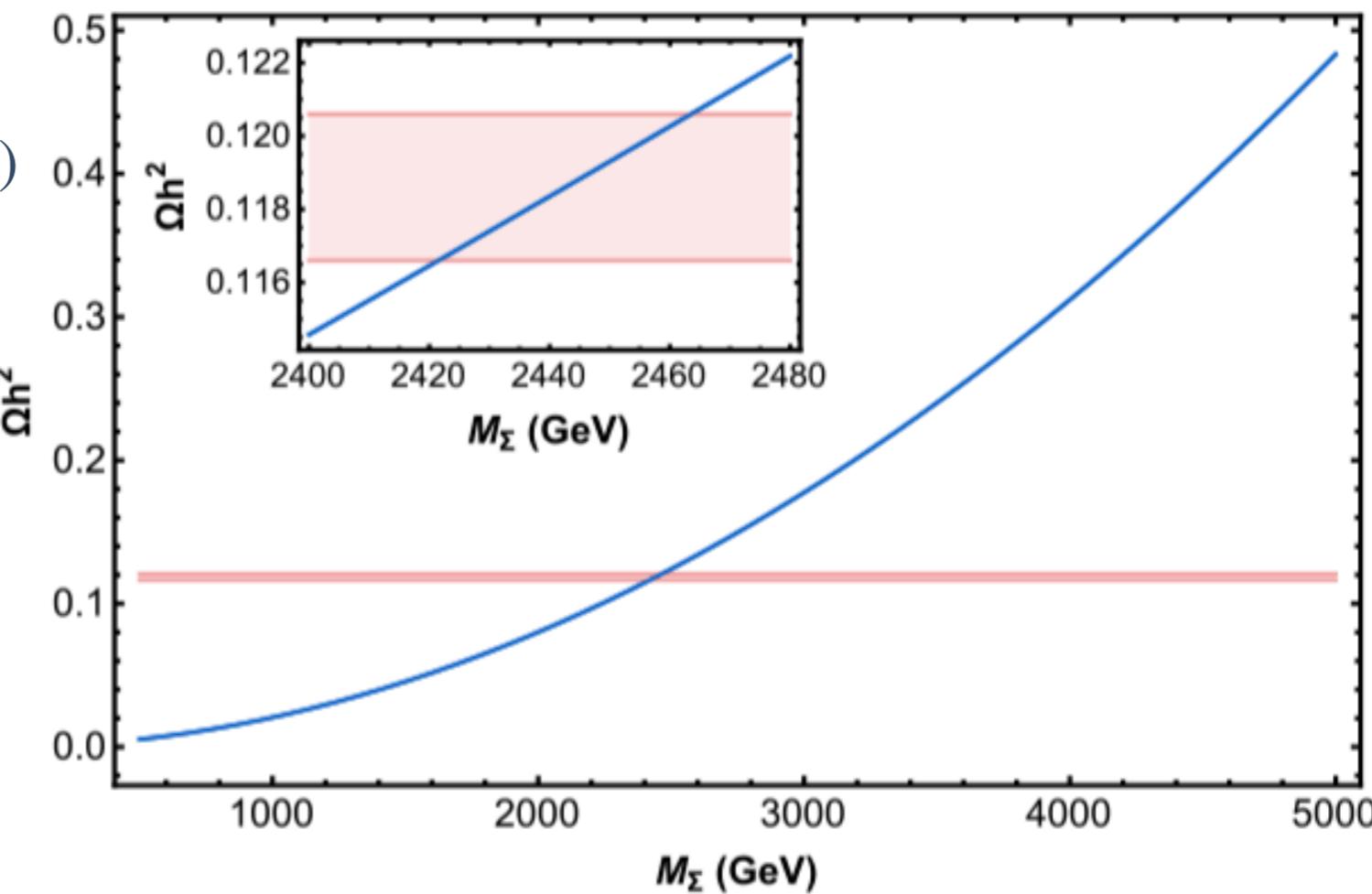
Σ_R coannihilation primarily
via gauge interactions

Freeze-out temperature $x_f (\equiv m_\Sigma / T_f)$
is calculated using

$$x_f = \ln \left(\frac{0.038 g_{\text{eff}} M_{\text{Pl}} m_\Sigma \langle \sigma_{\text{eff}} |\bar{\nu}| \rangle}{g_*^{1/2} x_f^{1/2}} \right)$$

$$\Omega h^2 = \frac{1.07 \times 10^9 x_f}{g_*^{1/2} M_{\text{Pl}}(\text{GeV}) I_a}$$

$$I_a = x_f \int_{x_f}^{\infty} x^{-2} a_{\text{eff}} dx$$



$\mathbf{b} \rightarrow \mathbf{s} \ell^+ \ell^-$: \mathbf{R}_K and \mathbf{R}_{K^*}

$$\mathcal{H}_{eff} = -\frac{4G_F}{\sqrt{2}} \lambda_t \sum_i C_i^\ell O_i^\ell + \sum \frac{\mathcal{C}^{NP}}{\Lambda_{NP}^2} \mathcal{O}^{NP}$$

$$O_9^\ell \propto (\bar{s}_L \gamma^\mu b_L)(\bar{\ell} \gamma_\mu \ell), \quad O_{10}^\ell \propto (\bar{s}_L \gamma^\mu b_L)(\bar{\ell} \gamma_\mu \gamma_5 \ell)$$

In SM

$$C_9 \simeq -C_{10} \simeq -4.1$$

In this model, we obtain

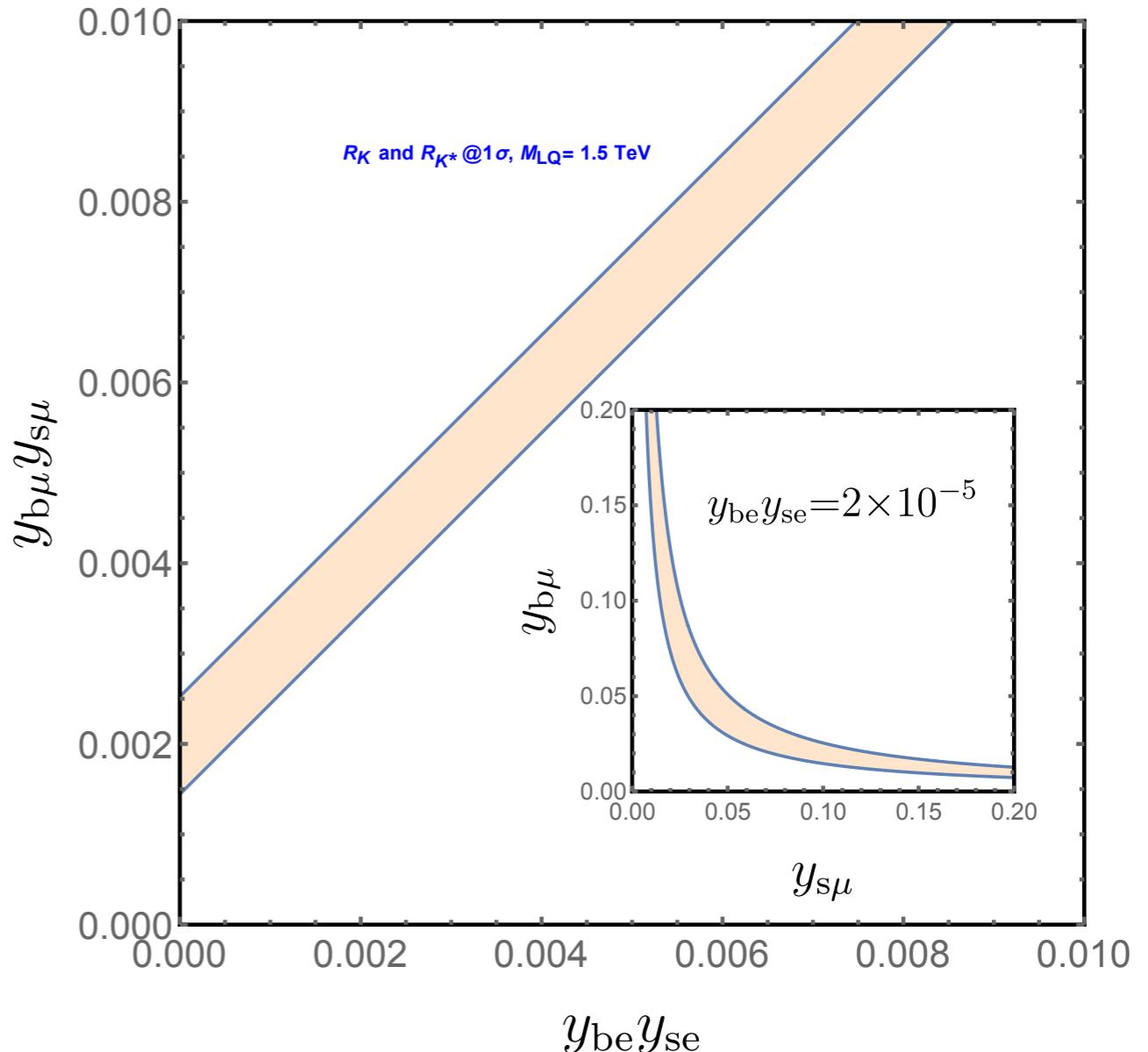
$$C_9^{\ell\ell'} = -C_{10}^{\ell\ell'} = \frac{\pi v^2}{\alpha_e V_{tb} V_{ts}^*} \frac{y_{b\ell'} y_{s\ell}^*}{m_{h_1}^2}$$

Neutral current anomalies can be explained by satisfying:

$$-1.4 \lesssim 2 \operatorname{Re}[C_{9,NP}^{\mu\mu} - C_{9,NP}^{ee}] \lesssim -0.8$$

@ 1σ

Hiller et al 1707.05444, Matias et al 1704.05340



$\mathbf{b} \rightarrow \mathbf{c} \tau\nu : \mathbf{R}_D$ and \mathbf{R}_{D^*}

$$\mathcal{H}_{\text{eff}} = \frac{4 G_F}{\sqrt{2}} V_{jk} \left(U_{\ell i} - \frac{\nu^2}{4 V_{cb} m_{h_1}^2} (yU)_{ki} (V y^*)_{j\ell} \right) \left(\bar{u}_j \gamma^\mu P_L d_k \right) \left(\bar{\ell} \gamma_\mu P_L \nu_i \right) + \text{h.c.}$$

Define

$$\frac{R_{D^{(*)}}}{R_{D^{(*)}, \text{SM}}} = \frac{1 - 2 \operatorname{Re} (x_{c\tau} y_{b\tau})}{1 - 2 \operatorname{Re} (x_{c\mu} y_{b\mu})}$$

with

$$x_{j\ell} = (V y^*)_{j\ell} \left(\nu^2 / 4 V_{cb} m_{h_1}^2 \right)$$

After taking all the relevant flavor constraints into account in our model

$b \rightarrow c\tau\nu$ is **SM-like**

We satisfy muon to electron LFU ratio

$$\frac{R_{D^{(*)}}^{\mu/e}}{R_{D^{(*)}, \text{SM}}^{\mu/e}} = \frac{1 - 2 \operatorname{Re} (x_{c\mu} y_{b\mu})}{1 - 2 \operatorname{Re} (x_{ce} y_{be})}$$

Belle Collab: Glattauer 2015, Abdesselam 2017

$$R_D^{\mu/e, \text{exp}} = 0.995 \pm 0.022 \pm 0.039$$

$$R_{D^*}^{e/\mu, \text{exp}} = 1.04 \pm 0.05 \pm 0.01$$

A way out if anomalies persist : add another scalar LQ (3, 2, 7/6) ?

See Bećirević et al, 1806.05689

Identifying textures of scalar triplet LQ Yukawa

We take phenomenological approach

Parametrize $y_{ij} = a_{ij} \odot \epsilon^{n_{ij}}$

$$y \sim \begin{pmatrix} \epsilon^{n_{11}} & \epsilon^{n_{12}} & \epsilon^{n_{13}} \\ \epsilon^{n_{21}} & \epsilon^{n_{22}} & \epsilon^{n_{23}} \\ \epsilon^{n_{31}} & \epsilon^{n_{32}} & \epsilon^{n_{33}} \end{pmatrix}$$

Requirement of saturating the B -meson anomalies $[R_{K^{(*)}}]$ at 1σ for $m_{h_1} = 1.5$ TeV

$$y_{22} y_{32} \approx 2.1555 \times 10^{-3} \sim \epsilon^{n_{22}+n_{32}} \rightarrow \epsilon^4 \sim 2.1555 \times 10^{-3} \Leftrightarrow \epsilon \approx 0.215$$

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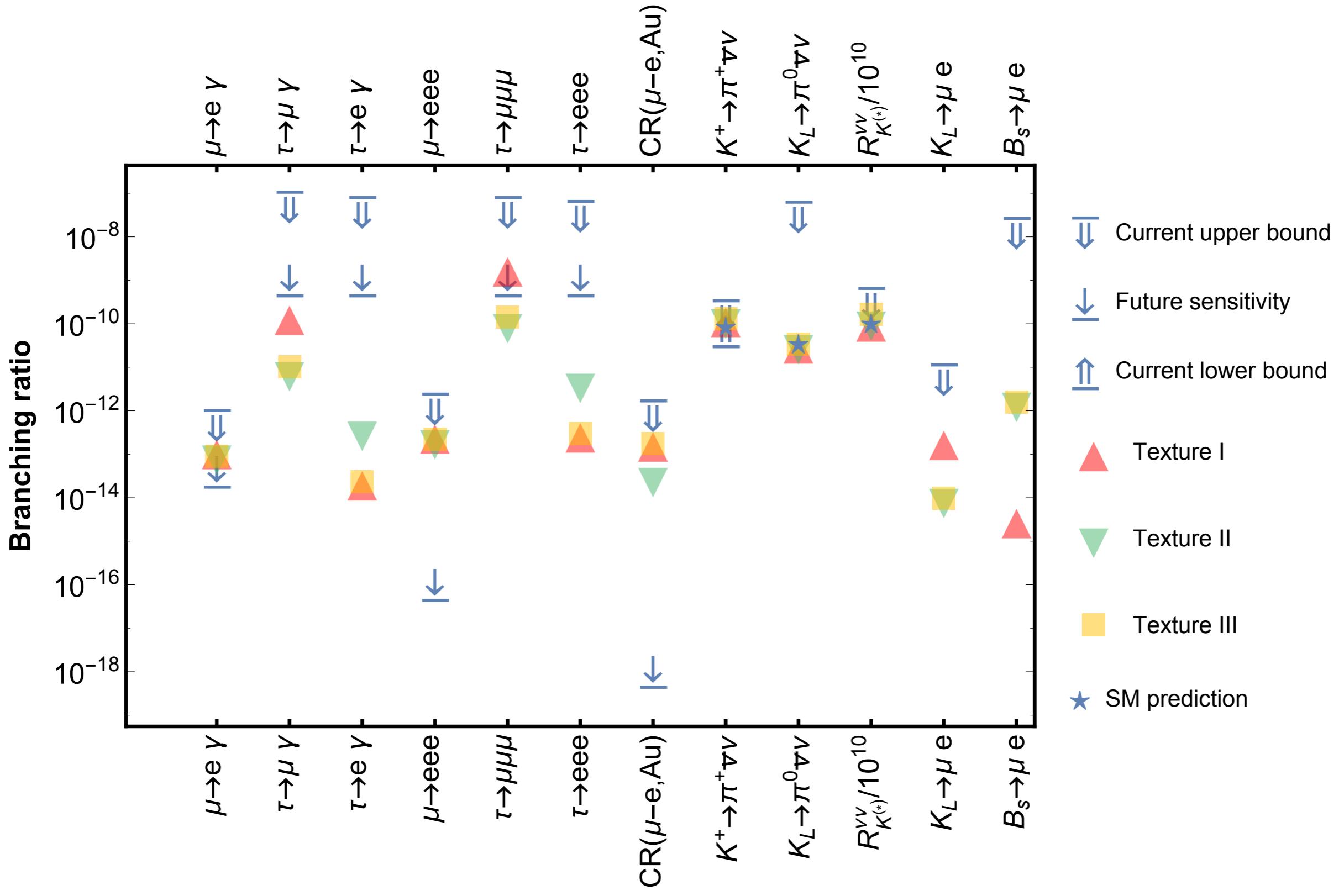
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We find the following allowed textures consistent with flavor violation processes

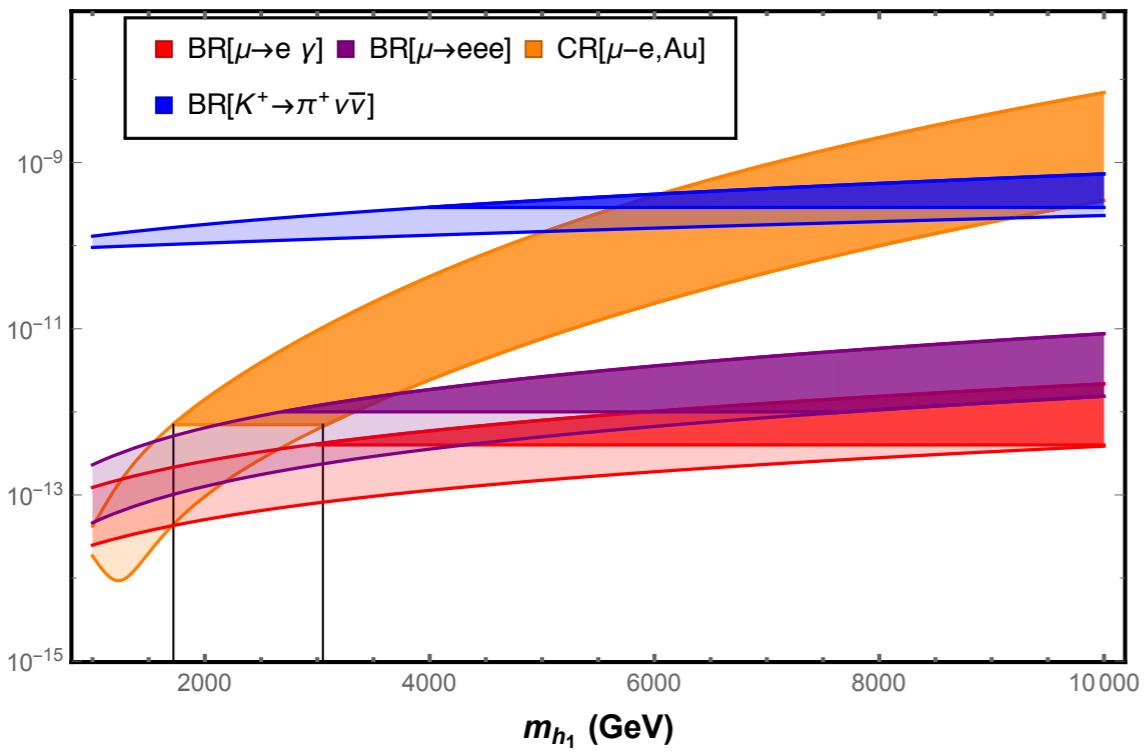
	Type I	Type II	Type III
y	$\begin{pmatrix} \times & \times & \times \\ \times & \epsilon^3 & \times \\ \times & \epsilon & \times \end{pmatrix}$	$\begin{pmatrix} \times & \times & \times \\ \times & \epsilon^2 & \times \\ \times & \epsilon^2 & \times \end{pmatrix}$	$\begin{pmatrix} \times & \times & \times \\ \times & \epsilon & \times \\ \times & \epsilon^3 & \times \end{pmatrix}$
Generic allowed textures	$\begin{pmatrix} \epsilon^4 & \epsilon^{\geq 5} & \epsilon^{\geq 2} \\ \epsilon^{\geq 3} & \epsilon^3 & \epsilon^{\geq 4} \\ \epsilon^{\geq 4} & \epsilon & \epsilon^{\geq 1} \end{pmatrix}$	$\begin{pmatrix} \epsilon^6 & \epsilon^{\geq 4} & \epsilon^{\geq 3} \\ \epsilon^{\geq 5} & \epsilon^2 & \epsilon^{\geq 3} \\ \epsilon^{\geq 3} & \epsilon^2 & \epsilon^{\geq 1} \end{pmatrix}$	$\begin{pmatrix} \epsilon^5 & \epsilon^{\geq 5} & \epsilon^{\geq 4} \\ \epsilon^4 & \epsilon & \epsilon^{\geq 2} \\ \epsilon^{\geq 4} & \epsilon^3 & \epsilon^{\geq 1} \end{pmatrix}$

Contributions to various flavor violating processes

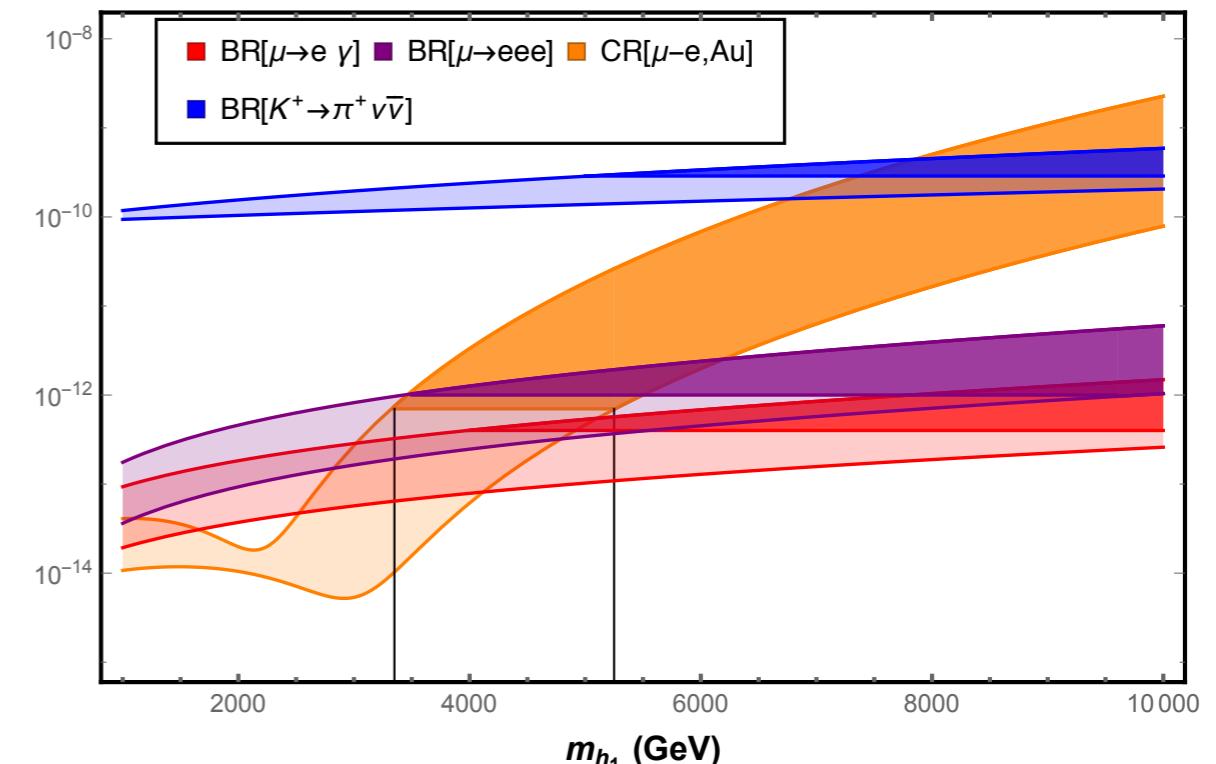


Constraints on LQ masses from flavor violation

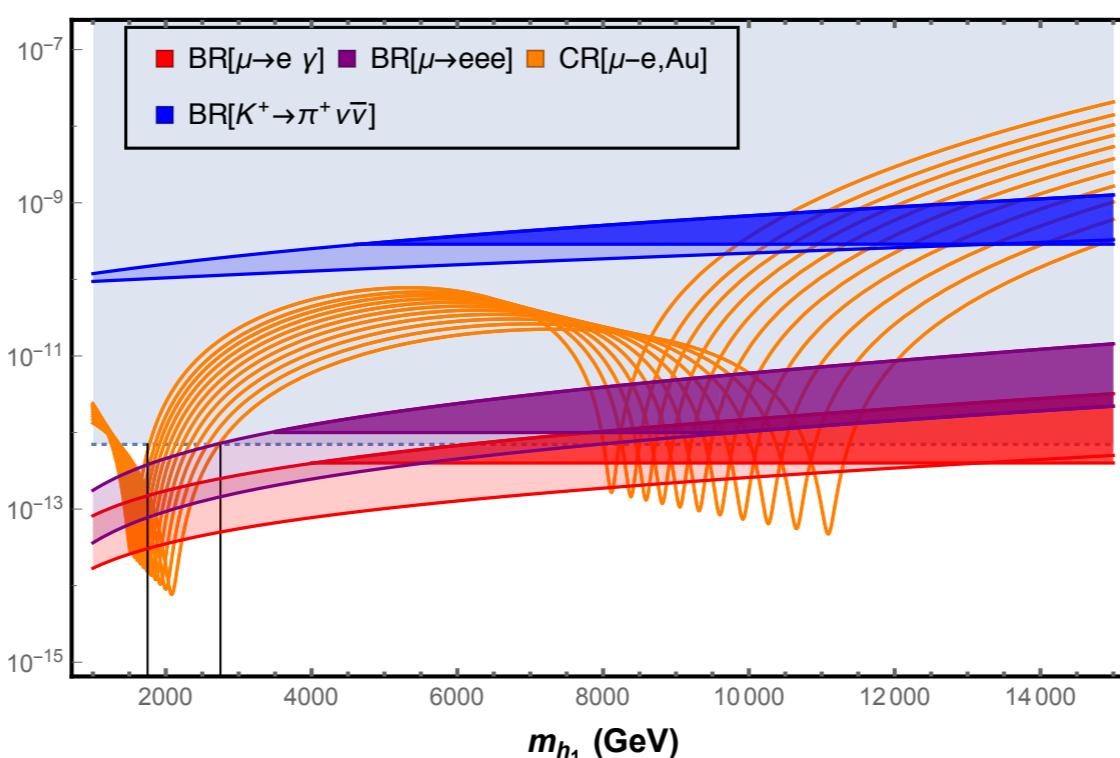
Type I texture



Type II texture



Type III texture



Constraints from neutrino oscillation data

$$\tilde{y} = F^{-1/2} \mathcal{R} \sqrt{m_\nu^{\text{diag}}} U^\dagger y^{-1} m_d^{-1}$$

$m_{\Sigma^{1,2,3}} \sim 2.45, 3.5, 4.5 \text{ TeV}$

$m_{h_1} \sim 1.5 \text{ TeV} \quad m_{h_2} \sim 2.6 \text{ TeV}$

Lightest neutrino mass $\sim 0.001 \text{ eV}$
Global best fit values for other parameters

Scan for \tilde{y} complying with flavor constraints

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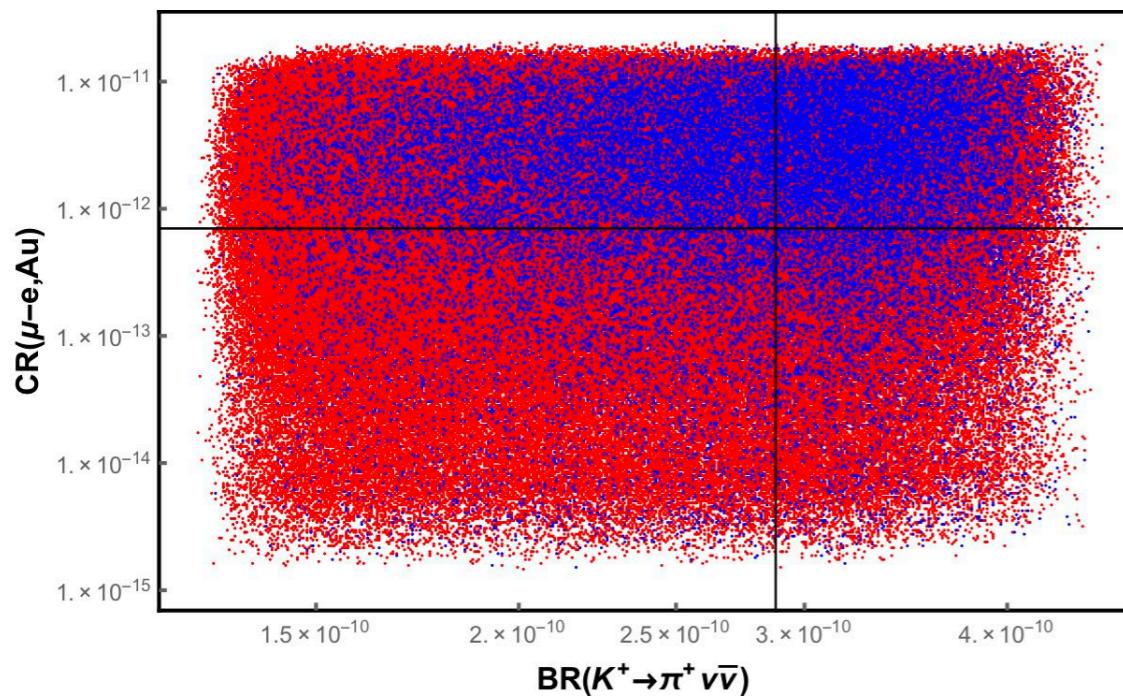
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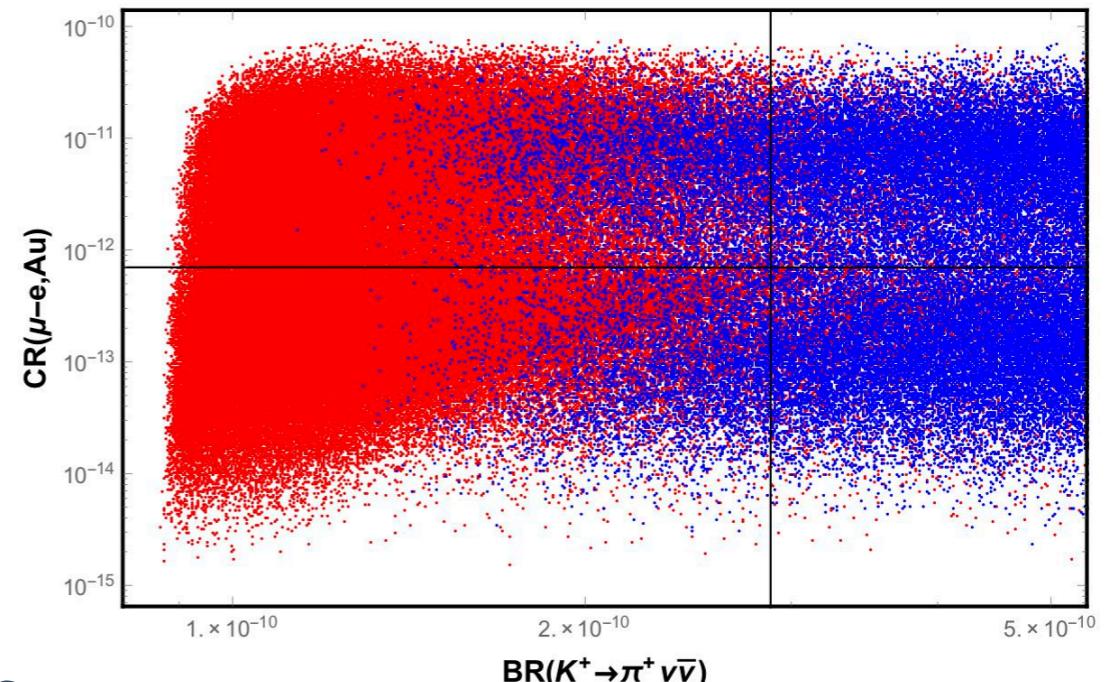
Lightest neutrino mass ~ 0.001 eV
 Global best fit values for other parameters

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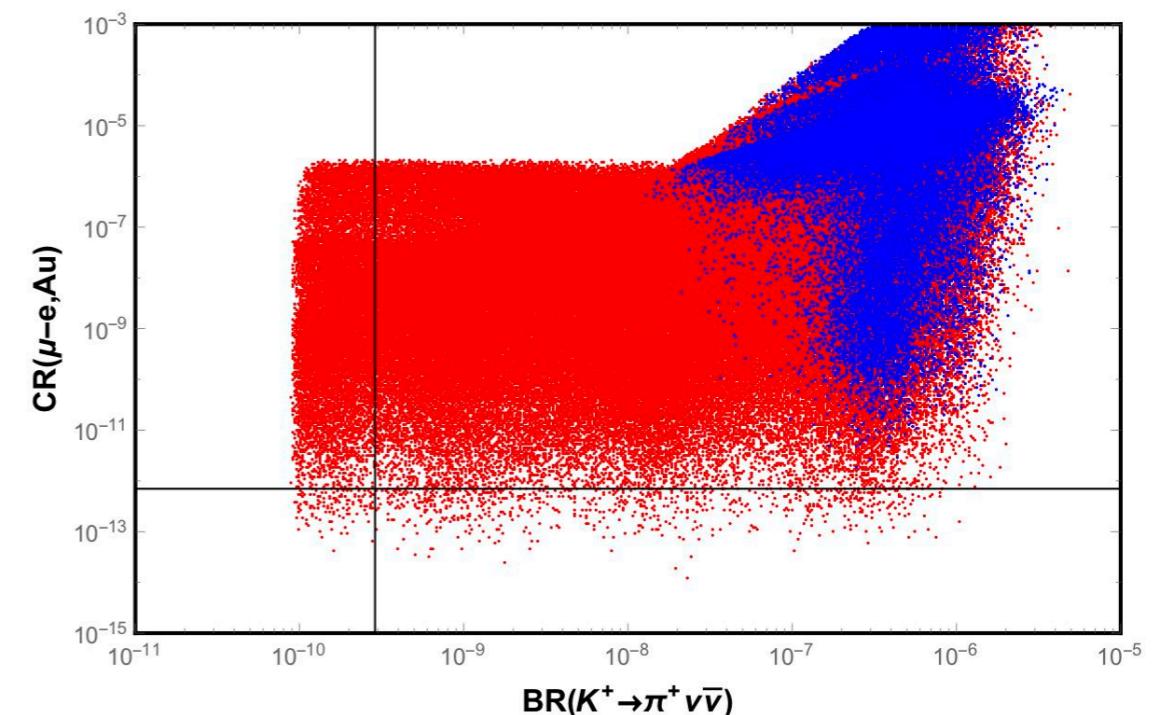
Type II texture



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Type III texture



Conclusions

The SM extension via (2) scalar LQ + (3) Triplet Majorana fermions

- ✓ Addresses neutrino mass generation
- ✓ Explains anomalies in R_K and R_{K^*}
- ✓ Has a viable dark matter candidate accounting for correct relic density

Satisfy all relevant flavor constraints : meson decays, meson-antimeson oscillations, cLFV decays

$\mu - e$ and $K \rightarrow \pi \nu \bar{\nu}$ are the most constraining observables for this class of models

Does not account for anomalies in charged current $b \rightarrow c$; possibility of accommodating these via extended the LQ sector

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Thank you

Scalar potential : Backup

$$\begin{aligned}
V(H, h_1, h_2) = & \mu_H^2 H^\dagger H + \frac{1}{2} \lambda_H |H^\dagger H|^2 + \mu_{h_1}^2 \text{Tr}[h_1^\dagger h_1] + \mu_{h_2}^2 \text{Tr}[h_2^\dagger h_2] + \\
& + \frac{1}{8} \lambda_{h_1} [\text{Tr}(h_1^\dagger h_1)]^2 + \frac{1}{8} \lambda_{h_2} [\text{Tr}(h_2^\dagger h_2)]^2 + \frac{1}{4} \lambda'_{h_1} \text{Tr}[(h_1^\dagger h_1)]^2 + \frac{1}{4} \lambda'_{h_2} \text{Tr}[(h_2^\dagger h_2)]^2 + \\
& + \frac{1}{2} \lambda_{Hh_1} (H^\dagger H) \text{Tr}[h_1^\dagger h_1] + \frac{1}{2} \lambda'_{Hh_1} \sum_{i=1}^3 (H^\dagger \tau_i H) \text{Tr}[h_1^\dagger \tau_i h_1] + \\
& + \frac{1}{2} \lambda_{Hh_2} (H^\dagger H) \text{Tr}[h_2^\dagger h_2] + \frac{1}{2} \lambda'_{Hh_2} \sum_{i=1}^3 (H^\dagger \tau_i H) \text{Tr}[h_2^\dagger \tau_i h_2] + \\
& + \frac{1}{4} \lambda_h \text{Tr}[h_1^\dagger h_2]^2 + \frac{1}{8} \lambda'_h [\text{Tr}(h_1^\dagger h_2)]^2 + \frac{1}{4} \lambda''_h \text{Tr}[h_1^\dagger h_1] \text{Tr}[h_2^\dagger h_2] + \text{H.c.}
\end{aligned}$$

$$\frac{\lambda_h}{4} \text{Tr}(h_1^\dagger h_2 h_1^\dagger h_2) = \frac{\lambda_h}{2} h_1^{-1/3} h_2^{1/3} h_1^{-1/3} h_2^{1/3} - \lambda_h h_1^{-1/3} h_2^{-2/3} h_1^{-1/3} h_2^{4/3}$$

Interaction Lagrangian

$$\begin{aligned}
\mathcal{L}_{\text{int}}^{h,\Sigma} = & -y_{ij} \bar{d}_L^C{}^i h_1^{1/3} \nu_L^j - \sqrt{2} y_{ij} \bar{d}_L^C{}^i h_1^{4/3} e_L^j + \sqrt{2} y_{ij} \bar{u}_L^C{}^i h_1^{-2/3} \nu_L^j - y_{ij} \bar{u}_L^C{}^i h_1^{1/3} e_L^j \\
& - 2 \tilde{y}_{ij} \overline{\Sigma^0}{}_R^C{}^i h_2^{1/3} d_R^j - 2 \tilde{y}_{ij} \overline{\Sigma^+}{}_R^C{}^i h_2^{-2/3} d_R^j - 2 \tilde{y}_{ij} \overline{\Sigma^-}{}_R^C{}^i h_2^{4/3} d_R^j + \text{H.c.}
\end{aligned}$$

Flavor constraints : Backup

Process	Observables	SM Prediction	Experimental data
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	$\text{BR}(K^+ \rightarrow \pi^+ \nu \bar{\nu})$	$(8.4 \pm 1.0) \times 10^{-11}$ [Buras et.al. 2015]	$17.3_{-10.5}^{+11.5} \times 10^{-11}$ [E949] $< 11 \times 10^{-10}$ [Na62]
$K_L \rightarrow \pi^0 \nu \bar{\nu}$	$\text{BR}(K_L \rightarrow \pi^0 \nu \bar{\nu})$	$(3.4 \pm 0.6) \times 10^{-11}$ ["]	$\leq 2.6 \times 10^{-8}$ [E391a]
$B \rightarrow K^{(*)} \nu \bar{\nu}$	$R_K^{\nu \nu}, R_{K^*}^{\nu \nu}$	$R_{K^{(*)}}^{\nu \nu} = 1$	$R_K^{\nu \nu} < 3.9$ [Belle] $R_{K^*}^{\nu \nu} < 2.7$ [Belle]
$B_s^0 - \bar{B}_s^0$ oscillation	mixing parameters	$C_{B_s} = 1$ $\phi_{B_s} = 0$	$C_{B_s} = 1.070 \pm 0.088$ [UTfit] $\phi_{B_s} = (0.054 \pm 0.951)^\circ$ [UTfit]
$K_L \rightarrow \mu e$	$\text{Br}(K_L \rightarrow \mu e)$	—	$< 4.7 \times 10^{-12}$ [PDG]
$B_s \rightarrow \mu e$	$\text{Br}(B_s \rightarrow \mu e)$	—	$< 1.1 \times 10^{-8}$ [PDG]

This list is not exhaustive

cLFV processes : Backup

LFV Process	Current experimental bound bound (Experiment)	Future sensitivity sensitivity (Experiment)
$\text{Br}(\mu \rightarrow e\gamma)$	$< 4.2 \times 10^{-13}$ (MEG)	4×10^{-14} (MEG II)
$\text{Br}(\tau \rightarrow \mu\gamma)$	$< 4.4 \times 10^{-8}$ (BaBar)	10^{-9} (Super B)
$\text{Br}(\tau \rightarrow e\gamma)$	$< 3.3 \times 10^{-8}$ (BaBar)	10^{-9} (Super B)
$\text{Br}(\mu \rightarrow 3e)$	$< 1.0 \times 10^{-12}$ (SINDRUM)	$10^{-15(16)}$ (Mu3e)
$\text{Br}(\tau \rightarrow 3\mu)$	$< 4.4 \times 10^{-8}$ (BaBar)	10^{-9} (Super B)
$\text{Cr}(\mu \rightarrow e, N)$	$< 7 \times 10^{-13}$ (Au) (SINDRUM)	10^{-14} (DeeMe) $10^{-15(-17)}$ (COMET) 3×10^{-17} (Mu2e) 10^{-18} (PRISM/PRIME)

Dark Matter : Backup

s-channel

$$\Sigma^0 \Sigma^\pm \rightarrow W^\pm \rightarrow W^\pm W^0, W^\pm H, \bar{f}f'$$

$$\Sigma^+ \Sigma^- \rightarrow W^0 \rightarrow W^+ W^-, W^0 H, \bar{f}f$$

t-channel

$$\Sigma^0 \Sigma^0 \rightarrow W^+ W^- \quad \Sigma^0 \Sigma^\pm \rightarrow W^\pm W^0$$

$$\Sigma^\pm \Sigma^\pm \rightarrow W^\pm W^\pm$$

$$\Sigma^+ \Sigma^- \rightarrow W^0 W^0 (W^+ W^-)$$

Thermally averaged eff. cross section

$$\begin{aligned} \langle \sigma_{\text{eff}} |\bar{v}| \rangle = & \frac{g_0^2}{g_{\text{eff}}^2} \sigma(\Sigma^0 \Sigma^0) |\bar{v}| + 4 \frac{g_0 g_\pm}{g_{\text{eff}}^2} \sigma(\Sigma^0 \Sigma^\pm) |\bar{v}| \left(1 + \frac{\Delta_{m_\Sigma}}{m_\Sigma}\right)^{3/2} \exp\left(-\frac{\Delta_{m_\Sigma}}{m_\Sigma} x_f\right) + \\ & + \frac{g_\pm^2}{g_{\text{eff}}^2} [2 \sigma(\Sigma^+ \Sigma^-) |\bar{v}| + 2 \sigma(\Sigma^\pm \Sigma^\pm) |\bar{v}|] \left(1 + \frac{\Delta_{m_\Sigma}}{m_\Sigma}\right)^3 \exp\left(-2 \frac{\Delta_{m_\Sigma}}{m_\Sigma} x_f\right) \end{aligned}$$

$$x_f = \ln \left(\frac{0.038 g_{\text{eff}} M_{\text{Pl}} m_\Sigma \langle \sigma_{\text{eff}} |\bar{v}| \rangle}{g_*^{1/2} x_f^{1/2}} \right)$$

$$\Omega h^2 = \frac{1.07 \times 10^9 x_f}{g_*^{1/2} M_{\text{Pl}} (\text{GeV}) I_a}$$

$$g_{\text{eff}} = g_0 + 2 g_\pm \left(1 + \frac{\Delta_{m_\Sigma}}{m_\Sigma}\right)^{3/2} \exp\left(-\frac{\Delta_{m_\Sigma}}{m_\Sigma} x_f\right)$$