## Bayesian analysis and naturalness of (Next-to-)Minimal SUSY Models

D. Harries

(IPNP, Charles University in Prague)

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Contact: harries@ipnp.mff.cuni.cz

#### Naturalness Problems

• Common motivation for SUSY is the hierarchy problem,



- Example of a "naturalness problem"
  - See also: strong CP problem, cosmological constant problem, flatness problem, ...

- Possible characterisation: propensity of model to reproduce data
- Related to fine-tuning, e.g., if observations require a priori unjustified tuning of parameters ⇒ model unnatural
- Prefer models in which fine-tuning not required/reduced?
- E.g., "little hierarchy problem" in MSSM (at tree-level)

$$m_{h_1}^2 \leq m_Z^2 \cos^2 2eta \lesssim (91 \, \operatorname{GeV})^2$$

 $\Rightarrow$  non-minimal SUSY models?

# Quantifying Fine-tuning

#### Typical Motivation

 $<\!$  model name> raises the Higgs mass at tree-level, and therefore is more natural than the MSSM, e.g., in the NMSSM

$$m_{h_1}^2 \lesssim m_Z^2 \cos^2 2eta + rac{\lambda^2 oldsymbol{v}^2}{2} \sin^2 2eta$$

- $\bullet\,$  Justify/check claim that model is more natural than another  $\Rightarrow$  quantify naturalness, i.e., fine-tuning
- Traditionally, construct fine-tuning measure, i.e., calculable function of model parameters that is identified with tuning
- $\bullet\,$  Focus on different features of a model's behaviour  $\Rightarrow\,$  different fine-tuning measure
  - Model fine-tuned according to one measure  $\Rightarrow$  also fine-tuned under a different measure?
- I.e., depends on your definition of fine-tuning ....

### Traditional Tuning Measures

What constitutes "fine-tuning"? Opinions differ.

• Large cancellations [1]?

$$\Delta_{EW} = \max_i rac{2|C_i|}{m_Z^2}, \quad C_1 = -\mu^2, \ C_2 = rac{m_{H_d}^2}{ an^2 eta - 1}, \ldots$$

• Extreme sensitivities [2]?

$$\Delta_{BG} = \max_i \left| rac{\partial \ln m_Z^2}{\partial \ln p_i} 
ight|, \quad p_i \in \{ ext{fundamental parameters}\}$$

- High or low energy contributions ( $\Delta_{EW}$  or  $\Delta_{HS}$ )? Definition of parameters  $p_i$ ? Just  $m_Z$ ?
- Model fine-tuned  $\Leftrightarrow \Delta_{EW/HS/BG/...} >$ ?
- Compare tuning between models?

#### What does $\Delta_{EW/HS/BC/...} > x$ mean for our degree of belief in model?

H. Baer, V. Barger, P. Huang, A. Mustafayev, and X. Tata, Phys. Rev. Lett. 109 (2012) 161802 [arXiv:1207.3343]
 J. R. Ellis, K. Enqvist, D. V. Nanopoulos, and F. Zwirner, Mod. Phys.Lett. A1 (1986) 57; R. Barbieri and G. F. Giudice, Nucl. Phys. B306 (1988) 63

## Traditional Tuning Measures



[arXiv:1309.2984]

## A Bayesian Approach

- Results based on traditional measures depend heavily on choice of measure
- Generally, when we talk about "naturalness" we are really talking about plausibility
  - Unnatural model  $\Leftrightarrow$  implausible model
  - Traditional measures do not have an unambiguous interpretation in this sense
- ullet  $\Rightarrow$  better questions to ask are,

How plausible is a given parameter space point in a model, in light of data?

Which model in a given set is the most plausible, in light of data?

• Rigorous logical framework exists for answering these questions: Bayesian statistics

#### Naturalness Priors

- Bayesian framework automatically captures intuition about "naturalness"
- Bayes' theorem applied to model *M*, parameters *x*, "observables" *O*:

$$p(\mathbf{x}|\text{data}, M) = \frac{p(\text{data}|\mathbf{x}, M)p(\mathbf{x}|M)}{p(\text{data}|M)}, \quad p(M|\text{data}) = \frac{p(\text{data}|M)p(M)}{p(\text{data})}$$
$$p(\text{data}|M) = \int d^n x_i \ p(\text{data}|\mathbf{x}, M)p(\mathbf{x}|M), \quad p_{\text{eff.}}(x_j, \ldots) = \int dO_i \ p(\mathbf{0}, \mathbf{x}'|M, O_i^{\text{exp.}})$$

• Reparameterise in favour of  $\boldsymbol{O}$  and remaining parameters  $\boldsymbol{x}'$ , e.g.,

$$p(\mathbf{x}|M) = \mathcal{J}p(\mathbf{0}, \mathbf{x}'|M)$$

 $\Rightarrow$  evidence and effective naturalness priors suppressed by Jacobian  $\Delta_J \sim \mathcal{J}$  [3],

$$\Delta_J = \left| \det \frac{\partial \ln O_i}{\partial \ln x_j} \right|$$

[3] D. Kim, P. Athron, C. Balázs, B. Farmer, and E. Hutchison, Phys. Rev. D 90 (2014) 055008 [arXiv:1312.4150]

#### Example: the SM Hierarchy Problem

Toy model 
$$\Rightarrow M_Z^2 = -rac{ar{g}^2}{8\lambda}(\mu^2 + \Lambda_{NP}^2)$$



• Traditional BG measure applied to, e.g., cut-off  $\Lambda^2_{NP},$ 

$$\Delta_{\Lambda^2_{NP}}=rac{ar{g}^2}{8\lambda}rac{\Lambda^2_{NP}}{M^2_Z}$$

 $\Rightarrow$  large tuning for  $\Lambda^2_{NP} \gg M^2_Z$  but interpretation unclear

• Compute posterior for  $\Lambda^2_{NP}$  using

$$p_{\text{eff.}}(\Lambda^2,\lambda) \propto \left| rac{\partial M_Z}{\partial \mu^2} 
ight|_{\mu=\mu_Z}^{-1} p(\Lambda^2,\mu_Z^2,\lambda|SM)$$

 $\Rightarrow$  captures traditional intuition about tuning

• But also has a well-defined probabilistic interpretation

#### Models Considered

• CMSSM as reference model with parameters  $\mu_0, B_0 \mu_0, \, {\rm sign} \, \mu, \, {\rm and}$ 

$$m_Q^2 = m_{u^c}^2 = \ldots = m_{H_d}^2 = m_{H_u}^2 = m_0^2,$$
  
 $M_1 = M_2 = M_3 = m_{1/2},$   
 $A_u = A_d = A_e = A_0$ 

- Naturalness priors from  $\{|\mu_0|, B_0\mu_0\} \rightarrow \{M_Z^2, \tan\beta\}$
- Semi-constrained  $\mathbb{Z}_3$ -NMSSM,

$$\widehat{W}_{NMSSM} = \widehat{W}_{MSSM}\Big|_{\mu=0} + \lambda \hat{S} \hat{H}_d \cdot \hat{H}_u + \frac{1}{3}\kappa \hat{S}^3,$$
  
new parameters  $\lambda_0$ ,  $\kappa_0$ ,  $m_{S_0}^2$ , sign $(\lambda \langle S \rangle)$ ,  $A_\lambda$ ,  $A_\kappa$ 

• Trade  $\{\lambda_0, \kappa_0, m_{S_0}^2\} \rightarrow \{\lambda, M_Z^2, \tan\beta\}$ 

CMSSM	
$\begin{matrix} m_0 \\ m_{1/2} \\ A_0 \\  \mu_0  \\ B_0 \mu_0 \end{matrix}$	$\begin{array}{l} \mbox{Log, 1 GeV} - 20 \mbox{TeV} \\ \mbox{Log, 1 GeV} - 15 \mbox{TeV} \\ \mbox{Log, 100 GeV} <  A  \leq 20 \mbox{TeV}, \\ \mbox{Flat, }  A  \leq 100 \mbox{GeV} \\ \mbox{Log, 100 GeV} - 20 \mbox{TeV} \\ \mbox{Log, (100 GeV)}^2 - (20 \mbox{TeV})^2 \end{array}$
sign $\mu$	$\pm 1$ with equal probability
NMSSM	
$\lambda_0 \ \kappa_0 \ m_{S_0} \ A_\lambda \ A_\kappa$	Log, $10^{-6}-1$ Log for $10^{-10}< \kappa <1$ Same as $m_0$ Same as $A_0$ Same as $A_0$

# Low Fine-tuning and Credible Regions







- High posterior density  $\Leftrightarrow$ low fine-tuning (according to  $\Delta_{BG}$ )
- $m_{h_1} \approx 125 \text{ GeV not}$ required  $\Rightarrow$  weak-scale soft parameters preferred

# Impact of $m_h \approx 125$ GeV





- Credible intervals shifted  $\sim$  2 orders of magnitude in  $m_0$ ,  $m_{1/2}$ 
  - Expected: LHC  $\Rightarrow$  SUSY implausible below  $\sim$  1 TeV
- Still find most plausible regions ⇔ lowest tunings consistent with data
- Both approaches ⇒ focus-point region is preferred

# Comparing Fine-tuning Measures





- Can use  $\Delta_J$  as a standalone tuning measure
- Qualitative agreement between  $\Delta_{BG}$ ,  $\Delta_{EW}$ ,  $\Delta_J$
- But meaningful results ⇒ ideally should compute full posterior densities or evidences

- Some differences, e.g.,  $\Delta_J$  shows much stronger preference for focus point than  $\Delta_{BG}$
- Aside: differing definitions of measures ⇒ direct comparison of values not reasonable

## Posterior Distributions





- Posterior densities reflect traditional intuition
- Pre-Higgs densities  $\Rightarrow$  low  $m_{\rm SUSY} \lesssim 1 \,{\rm TeV}$  most plausible ("natural")
- Inclusion of  $m_h \Rightarrow m_{\rm SUSY} \gtrsim 4 \,{\rm TeV}$  required, TeV scale soft parameters

- Bayesian approach ⇒ systematically update degree of belief given new data
- Ambiguous when using approach based on  $\Delta_{EW/BG/\dots}$

## Summary

- Model comparison based on traditional tuning measures is ill-defined
- $\bullet\,$  More relevant to ask which parameter values/models are most plausible given observations  $\Rightarrow\,$  utilitise Bayesian methods
- Bayesian approach automatically incorporates traditional intuitions about fine-tuning
- Effective naturalness priors resemble traditional measures but have a rigorous probabilistic interpretation
- Qualitative agreement between traditional measures and naturalness priors, credible intervals in the CMSSM and semi-constrained NMSSM
- In principle, can go further: have all ingredients to compare model evidences ⇒ statistically meanginful model comparisons

#### Thank you for listening!

#### Additional Slides

## Minimum Tuning Values

	Mz		$M_Z$ and $m_h pprox$ 125 GeV	
	CMSSM	NMSSM	CMSSM	NMSSM
$\Delta_J  _{M_{\text{CUT}}}$	$3 imes 10^{-9}$	$2  imes 10^{-10}$	0.004	$8 imes 10^{-7}$
$\Delta_J \Big _{m_{\text{SUSY}}}$	$6 imes 10^{-7}$	$2 imes 10^{-10}$	0.005	$8 imes 10^{-7}$
$\Delta_{EW}$	0.3	0.3	48.7	47.4
$\Delta_{BG}$	0.1	0.2	451.9	133.2

- Note: numbers should be compared keeping in mind differing definitions of measures
- $\Delta_J \big|_{m_{SUSY}} \Rightarrow$  contributions to Jacobian from RG running omitted
- All measures increase once *m<sub>h</sub>* included
- Note  $\Delta_{EW}$  similar for both models, while  $\Delta_{BG}$ ,  $\Delta_J$  smaller for NMSSM