

Parameter Reconstruction and Estimation of Widths in Coherent Elastic Neutrino-Nucleus Scattering Experiments.

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Introduction

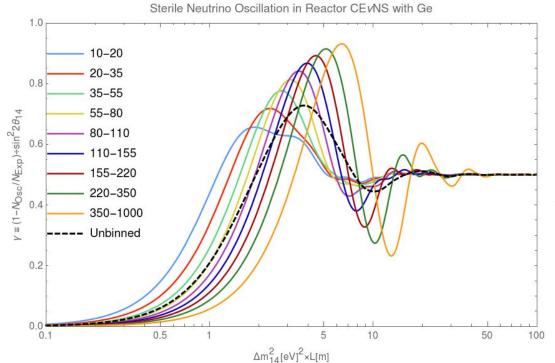
- Short Baseline Sterile Neutrino Searches
- The MIvER Experiment at Texas A&M University
- What are Bayesian and Frequentist Statistics?
- MultiNest What is it? Why use this tool?
- Our Toy Model What is it? What did we do with it? What did it tell us?
- Applications & Future Goals

Github link to the Jupyter Notebook

https://github.com/Andrerg01/Mathematica-Codes



Investigating Sterile Neutrino Hypothesis with Coherent Neutrino-Nucleus Scattering



Coherent Scattering:

Small momentum exchange means large wavelength; Individual nucleons are unresolved

Sterile Neutrino Parameters:

- \bullet Λm^2
- $Sin^2(2\theta)$

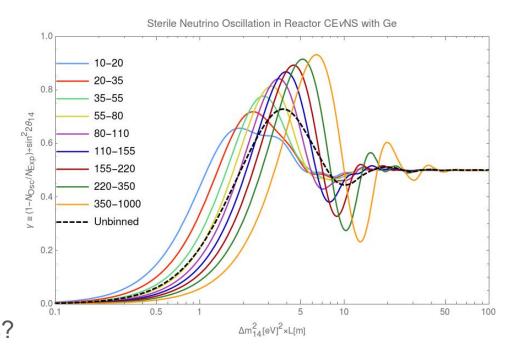
$$P_{(\alpha \to \beta)} = \sin^2 \left[2\theta \right] \times \sin^2 \left[\frac{\Delta m^2 L}{4E_{\nu}} \right]$$



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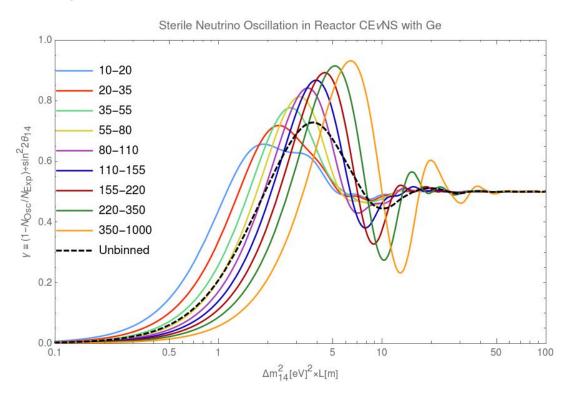
Questions of Interest

- Given a data set, what are the parameters for which the model best reproduces the data?
- For which experimental layout can we optimize the coverage of the parameter space?
- How can we estimate widths for parameter distributions given correlated systematic uncertainties?





Project Goals



- Exclude Parameter Space
- Rule Out Null Hypothesis
- Explore ~1 eV Sterile
 Neutrinos
- Reconstruct Mass Gap and Mixing Angle
- Optimize distance and exposure balance

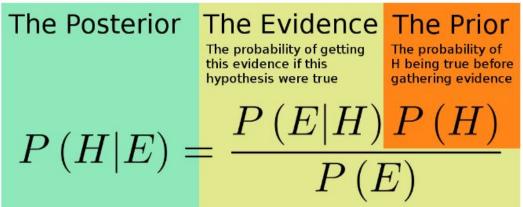


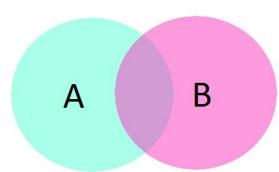
Frequentist Approach v. Bayesian Approach

- If I lose my keys where should I look?
 - Check Random Places?
 - Work from where I start and widen out from there?
 - Check the places where I usually am and where I use them?
 - My Desk?
 - The Car?



Bayesian Statistics





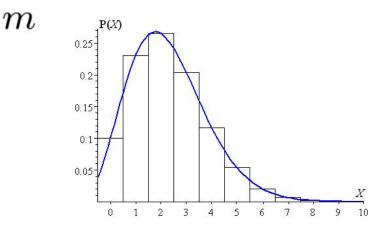
$$P(A|B)P(B) = P(B|A)P(A)$$



Likelihood Functions - Gaussian and Poisson Distributions

$$L_D = P(D|\theta) = \prod P(x_m|\theta) S(x_m)$$

$$P(X) = \frac{\lambda^x e^{-\lambda}}{X!}$$





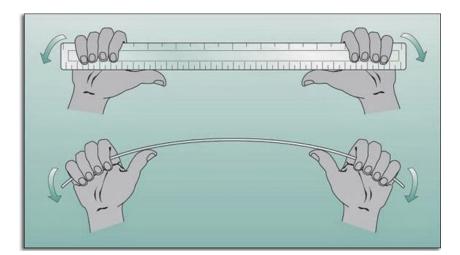
Penalyzing for putting too much strain on our model.

Finding the parameters that most efficiently reflect the model.

Allowing for lots of error allows for an easier fit but becomes meaningless.

Allowing for no error makes it difficult if not impossible to fit and still

meaningless.





What is MultiNest and why use it?

- Allows input of a prior and a likelihood function.
- Returns Parameters of Best Fit and their error.
- Returns posterior and relation between parameters

Model

Likelihood Function Parameters Prior $\rightarrow \boxed{\text{Multinest}} \rightarrow \begin{bmatrix} \text{Parameters of Best Fit} \\ \text{Width of Parameters} \\ \underline{\text{Evidence}} \end{bmatrix}$

D Parameter:

Mean: 12.0110622229

Standard Deviation: 0.200583092612

Most frequent histogram value: 11.9576348919

Correct Value: 12.0

E Parameter:

Mean: 0.120608118532

Standard Deviation: 0.115976420488

Most frequent histogram value: 0.0951341765873

Correct Value: 0.125663706144

Global Log Evidence:

Mean: -42.7401375839

Standard Deviation: 5.1871037815

Most frequent histogram value: -42.5917808219

Importance Global Log Evidence:

Mean: -42.7401375839

Standard Deviation: 5.1871037815

Most frequent histogram value: -42.5917808219

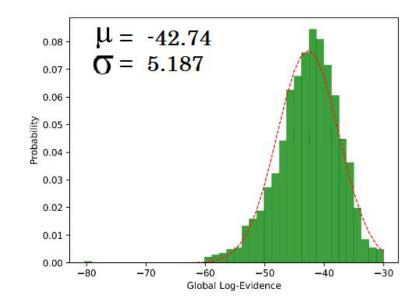


Why MultiNest?

Returns Log Evidence and its Error for the given model.

Evidence can (and will) be used as a tool to compare

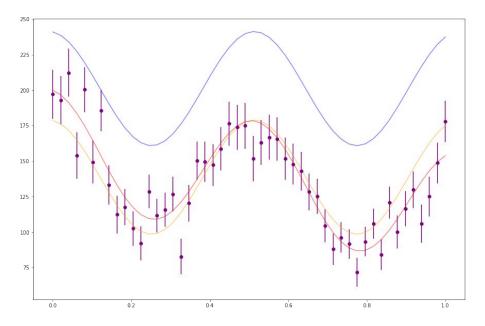
models.





- Create "Perfect" data set.
- Add "Background Noise"
 - Shift vertically
 - Tilt curve
 - "Smudge" data points.
- "Simulate" N Data Points

Model: A+Bx+C*Cos(Dx+E)
5 Parameters
Uniform Priors
Gaussian Likelihood Function





- Run MultiNest.
- Record Output
- Generate new data set
- Repeat

```
A = 154.07 \pm 4.09

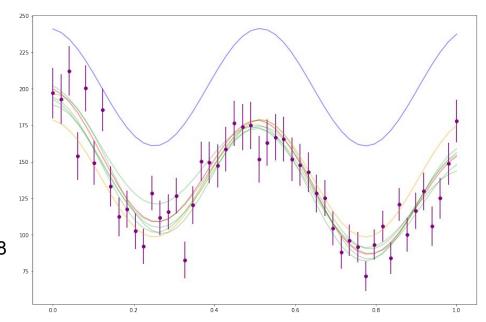
B = -36.45 \pm 6.87

C = 39.20 \pm 2.68

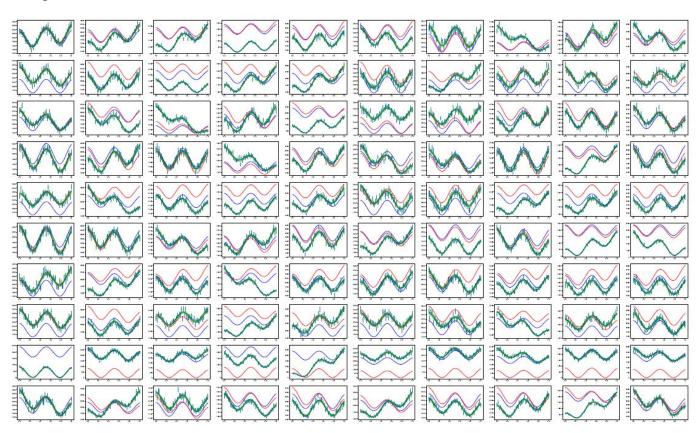
D = 11.98 \pm 0.21

E = 0.09 \pm 0.11

Log - Evidence = -48.29 \pm 0.28
```

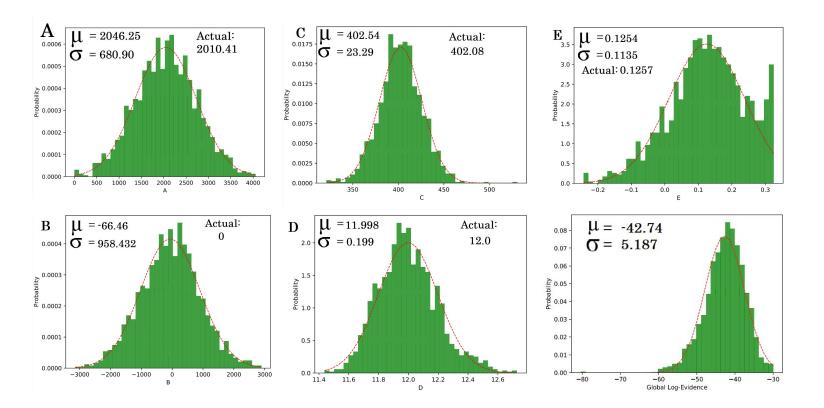








A+Bx+C*Cos(Dx+E)

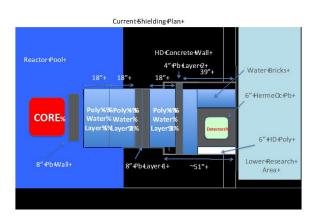




The MIvER Experiment - Texas A&M University

- Reactor Facility offers superior prospects for Short Baseline Neutrino studies
 - Megawatt Class Nuclear Reactor
 - Has a Flux on the order of 10¹² Neutrinos / cm² / s at 1 m
 - Moveable Core
 - Experimental baseline as short as 3 m with full shielding
 - Ultra low threshold detectors allow high sensitivity to CEvNS process.





1/14/77 Nov 14th MINER Meeting 3 11/14/17+ Nov-44th-MINER-Meeting



Summary and Future Works

- MultiNest presents much potential to be used as a method for fitting parameters and comparing models.
- By interpolating numerically calculated functions where the parameter simply scale the axis, we can use multinest to fit parameters and examine the evidence for each model - likelihood function considered.
- By examining the output we can determine the exposure time / distance from the detector (among other external parameters) that will produce the most meaningful results.



Thank you!