

Fuzzy Dark Matter and Non-Standard Neutrino Interaction

Jia Liu
Chicago U., EFI

Coauthor: Vedran Brdar, Joachim Kopp, Pascal Prass, Xiao-Ping Wang
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The era for non-canonical DM

- ◆ Standard WIMP
 - ◆ CDM@small scale structure
 - ◆ Null result in direct detection
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- Ultra-light bosonic DM
 - If $m \sim 10^{-22}$ eV, de Broglie wavelength \sim kpc, good for small scale
 - Large occupation number

Connection to SM: photon, lepton, baryon, neutrino (Asher Berlin, Gordan Krnjaic, Pedro A. N. Machado, Lina Necib) etc

Ultralight bosonic DM

- Scalar DM model

$$\mathcal{L}_{\text{scalar}} = \bar{\nu}_L^\alpha i\gamma^\nu \partial \nu_L^\alpha - \frac{1}{2} m_\nu^{\alpha\beta} \overline{(\nu_L^c)^\alpha} \nu_L^\beta - \frac{1}{2} y^{\alpha\beta} \phi \overline{(\nu_L^c)^\alpha} \nu_L^\beta$$

- Vector DM model

$$\mathcal{L}_{\text{vector}} = \bar{\nu}_L^\alpha i\gamma^\nu \partial \nu_L^\alpha - \frac{1}{2} m_\nu^{\alpha\beta} \overline{(\nu_L^c)^\alpha} \nu_L^\beta + gQ^{\alpha\beta} \phi^\mu \bar{\nu}_L^\alpha \gamma_\mu \nu_L^\beta$$

- DM MSW effect for neutrinos
- Solve classic E.O.M ($p^2 = m^2$)
- $m \rightarrow m + y\phi$ $p_\mu \rightarrow p_\mu + gQ\phi_\mu$

Ultralight bosonic DM

- Dispersion relation: $(E_\nu - V_{\text{eff}})^2 = \vec{p}_\nu^2 + m_\nu^2$

- Scalar DM model

$$V_{\text{eff}} = \frac{1}{2E_\nu} \left(\phi (y m_\nu + m_\nu y) + \phi^2 y^2 \right)$$

- Vector DM model

$$V_{\text{eff}} = -\frac{1}{2E_\nu} \left(2(p_\nu \cdot \phi)gQ + g^2 Q^2 \phi^2 \right)$$

- Neutrino effective potential in DM medium

- Vacuum oscillation: $V_{\text{vac}} = \frac{\Delta m^2}{2E_\nu}$

- MSW oscillation at Earth: $V_{\text{MSW}} = \sqrt{2}n_e G_F$

- Schrödinger equation: $H = V_{\text{vac}} + V_{\text{MSW}} + V_{\text{eff}}$

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Done!

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$$H_{\beta\alpha} |\alpha\rangle = i\partial_t |\beta\rangle$$

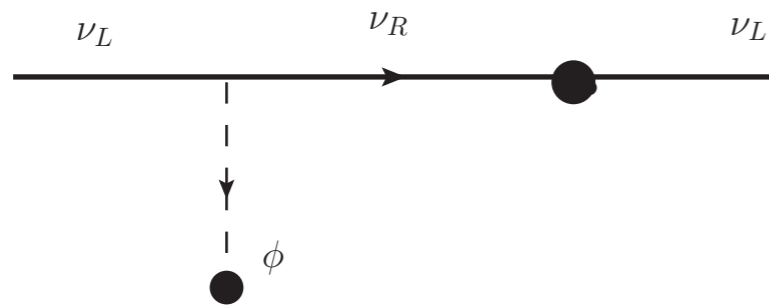
$$P_{\alpha\beta}(t) = |\langle \alpha(t) | \beta(0) \rangle|^2$$

- Schrödinger equation: $H = V_{\text{vac}} + V_{\text{MSW}} + V_{\text{eff}}$

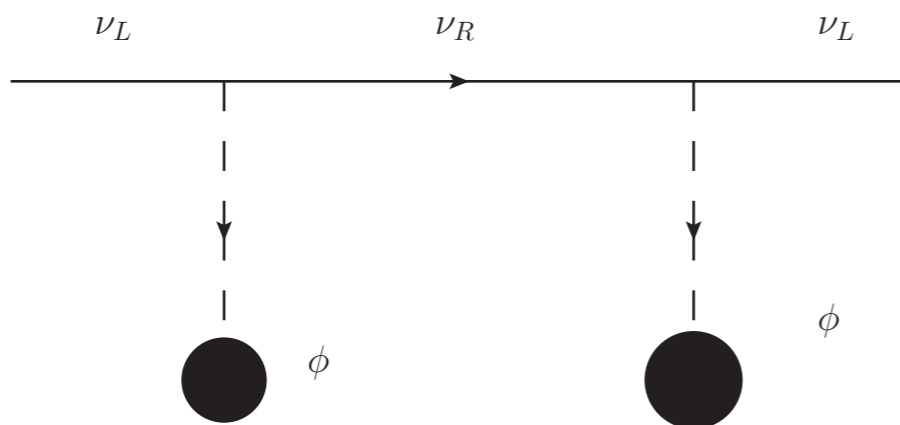
Scalar DM case

$$V_{\text{eff}} = \frac{1}{2E_\nu} \left(\phi (y m_\nu + m_\nu y) + \phi^2 y^2 \right)$$

- Linear term (classic)

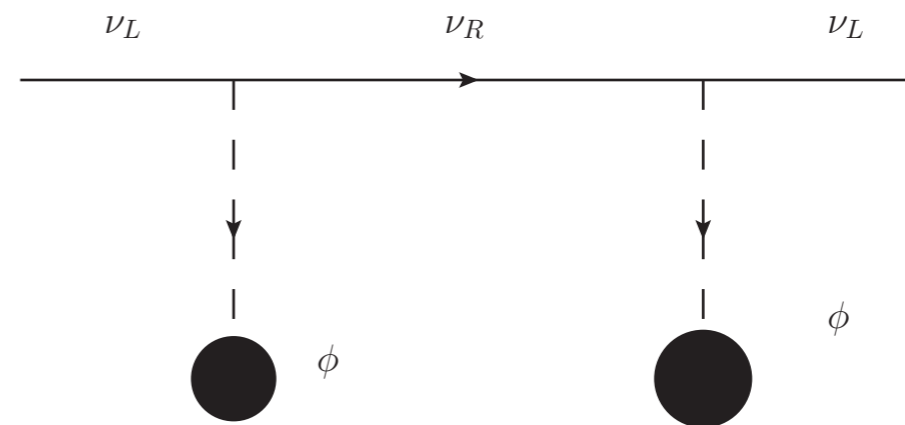
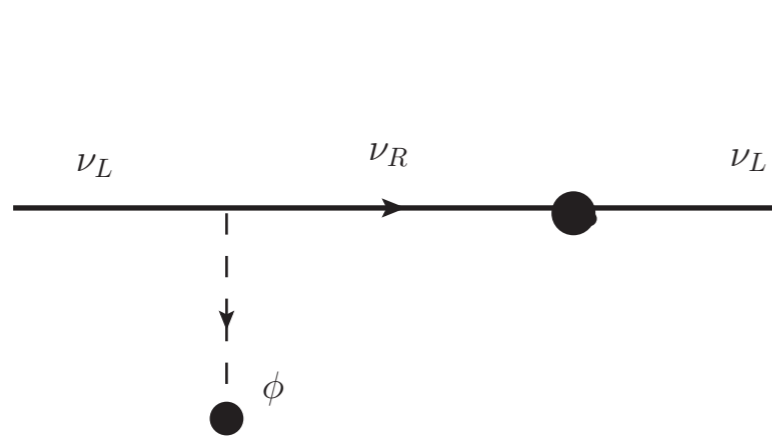


- Quadratic term (classic = QFT forward scattering)



Scalar DM case

$$V_{\text{eff}} = \frac{1}{2E_\nu} \left(\phi (y m_\nu + m_\nu y) + \phi^2 y^2 \right)$$



- DM from Misalignment

$$\partial_t \partial_t \phi + 3H \partial_t \phi + m^2 \phi = 0$$

$$\phi = \phi_0 \cos(m_\phi t) \quad \rho_\phi = \frac{1}{2} m_\phi^2 \phi_0^2$$

Scalar DM case

$$V_{\text{eff}} = \frac{1}{2E_\nu} \left(\phi (y m_\nu + m_\nu y) + \phi^2 y^2 \right)$$

- Naive estimation of constraints: $y\phi \sim m_\nu$
- Local DM density: 0.3 GeV/cm^3 $\rho_\phi = \frac{1}{2} m_\phi^2 \phi_0^2$

$$y/m_\phi \sim \mathcal{O}(eV)$$

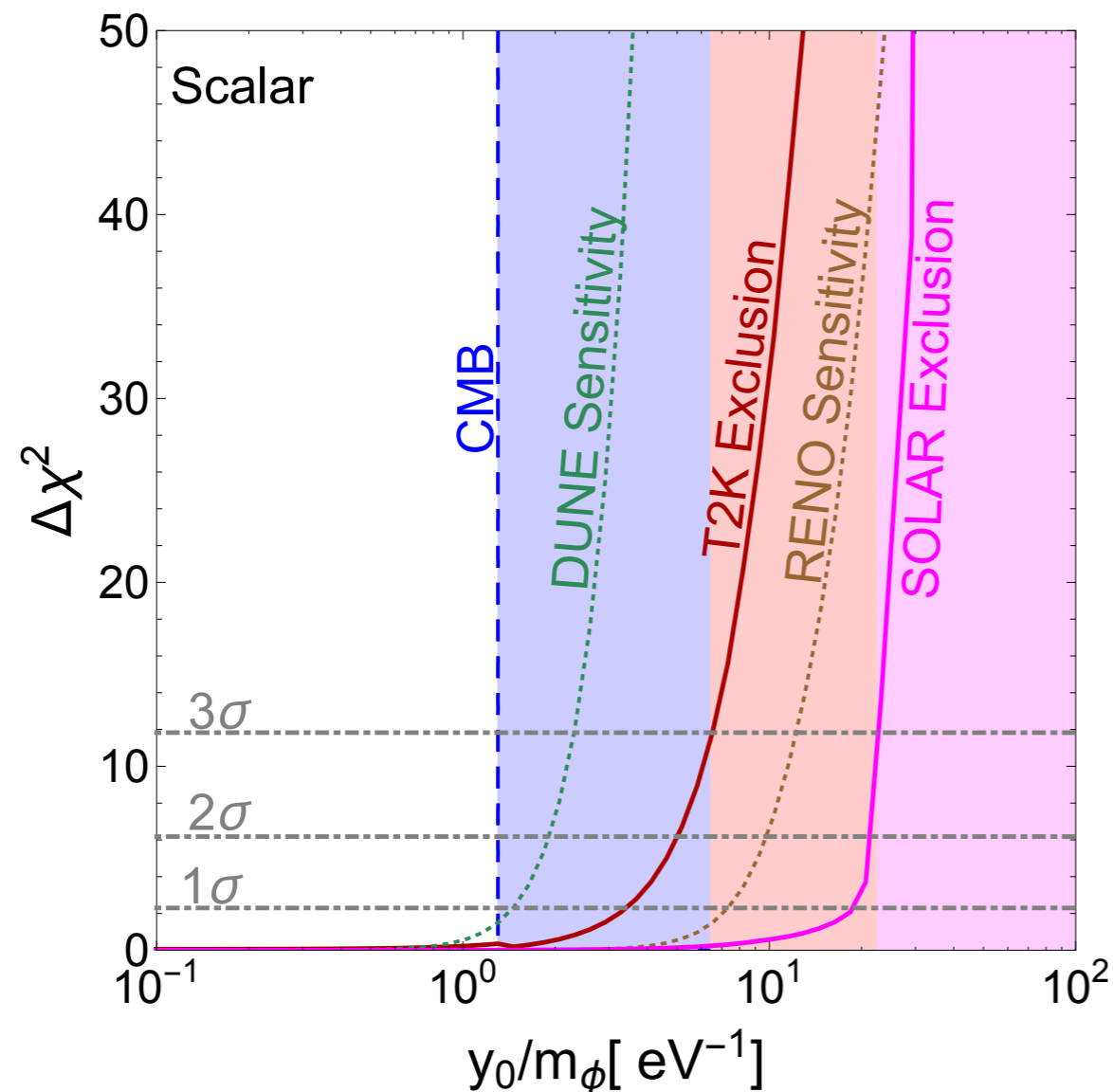
- Expansion on Probability:

$$P_{\alpha\beta} = P_0 + P_1 \cos(m_\phi t) + P_2 \cos^2(m_\phi t)$$

- Modulation effect and Average shift
- Modulation resolution for $1/m > 10 \text{ min}$.

Scalar DM case

$$V_{\text{eff}} = \frac{1}{2E_\nu} \left(\phi (y m_\nu + m_\nu y) + \phi^2 y^2 \right)$$



- Interaction assumption:

$$y = y_0 m_\nu / (0.1 \text{eV})$$

- CMB constraint on neutrino mass

$$\left. \begin{array}{l} \sum m_\nu < 0.23 \text{ eV} \\ \Omega_\nu h^2 < 0.0025 \end{array} \right\} 95\%, \text{ Planck TT+lowP+lensing+ext.}$$

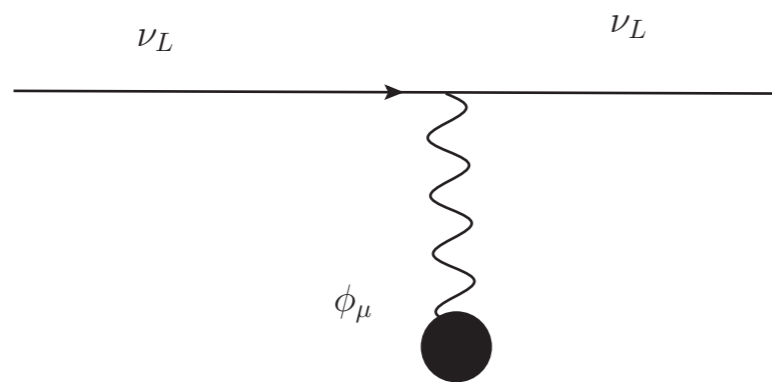
- Complex y
- Pseudo-scalar: axion

Vector DM case

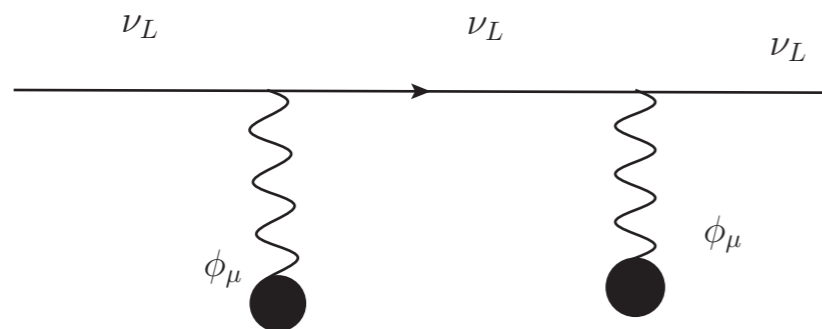
$$V_{\text{eff}} = -\frac{1}{2E_\nu} \left(2(p_\nu \cdot \phi)gQ + g^2 Q^2 \phi^2 \right)$$

- Linear term (classic): Only for fully polarized vector DM $Q_{\mu-\tau} = 0, 1, -1$

$$\phi_\mu = \xi_\mu \phi_0 \cos(m_\phi t)$$



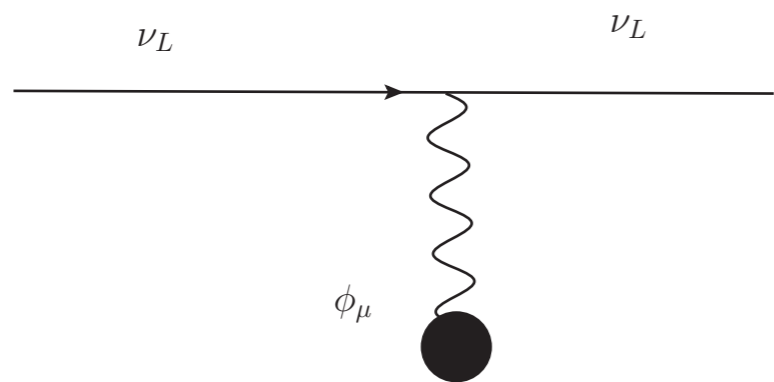
- Quadratic term (classic = QFT forward scattering): both fully polarized or unpolarized



Vector DM case

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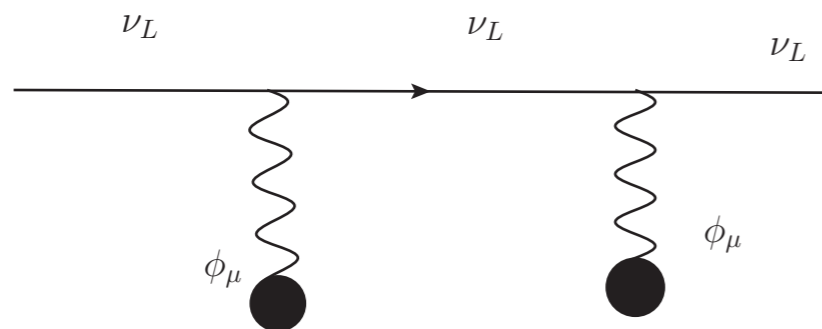
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$$\phi_\mu = \xi_\mu \phi_0 \cos(m_\phi t)$$

$$E_\nu g \phi_0 \sim m_\nu^2$$

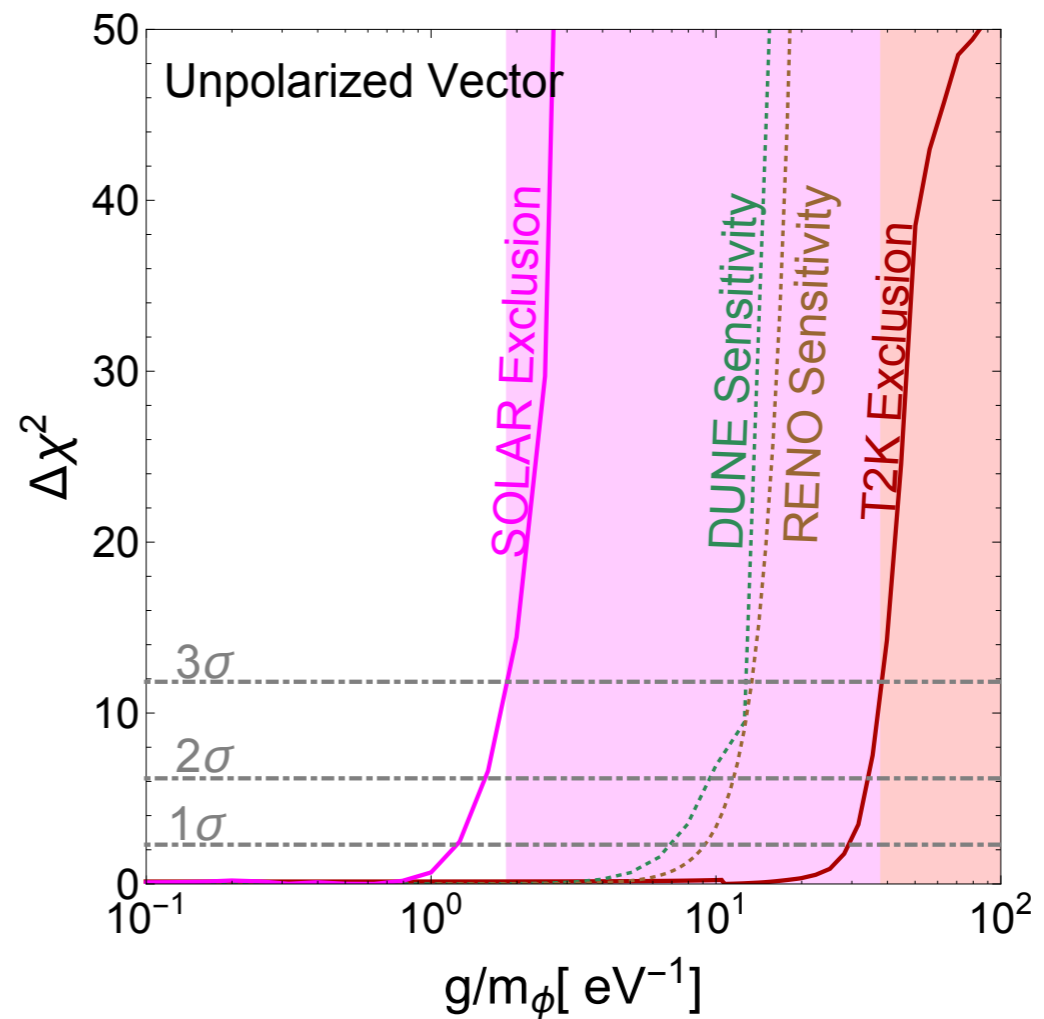
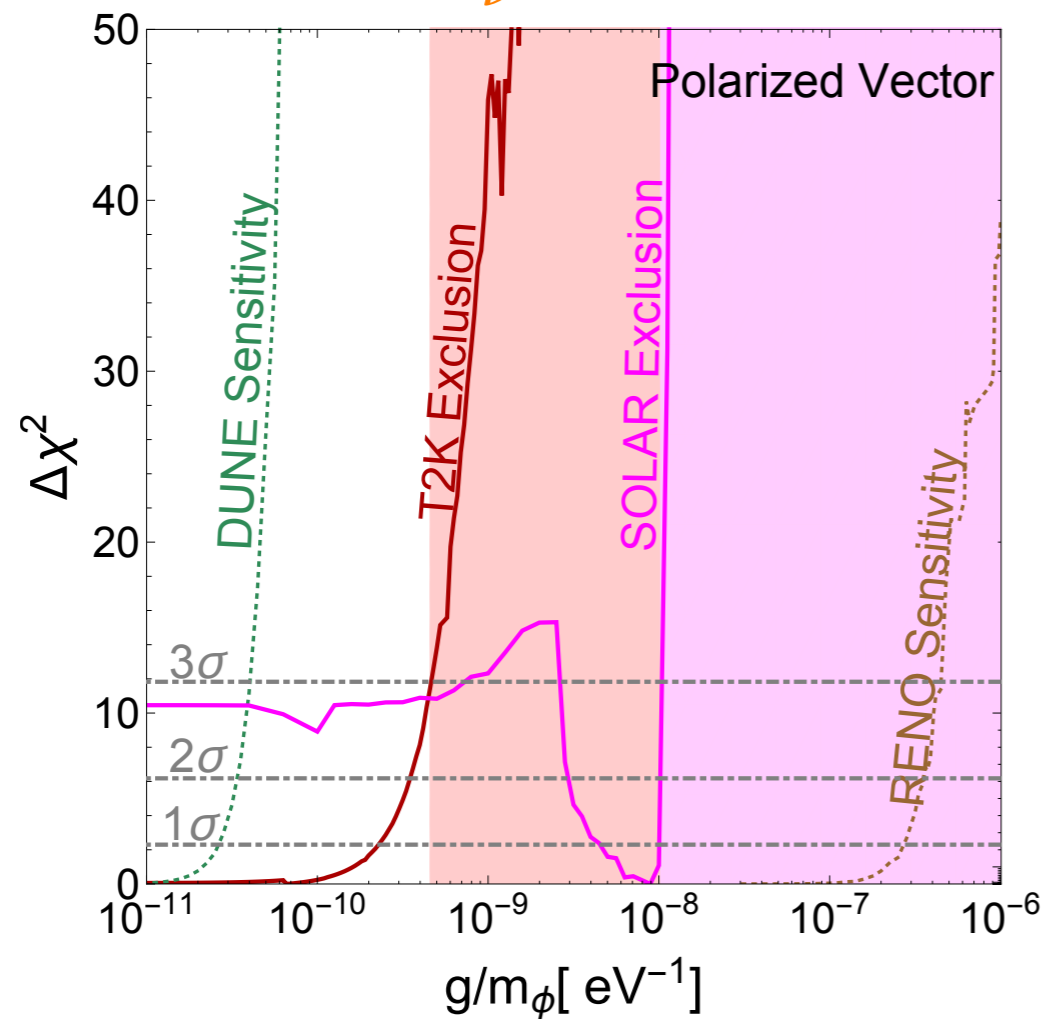
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$$(g\phi_0)^2 \sim m_\nu^2$$

Vector DM case

$$V_{\text{eff}} = -\frac{1}{2E_\nu} \left(2(p_\nu \cdot \phi)gQ + g^2 Q^2 \phi^2 \right)$$



- Some thoughts: $m \sim 10^{-22}$ eV, $g < 10^{-30}$!!!!

Ultralight DM Summary

- We introduce ultralight bosonic DM and non-standard interaction with neutrinos
- Stringent constraints from oscillation:
 - Scalar, unpolarized Vector: $y/m < \sim \text{eV}$
 - Polarized Vector: $y/m < \sim 10^{-10} \text{ eV}$
- Linear term apply when classic field description holds, $\# \gg 1$ ($m < 10^{-16} \text{ eV}$, if neutrino coherent length $L \sim 10^{-10} \text{ m}$)
- Quadratic term always applies (QFT forward scattering), valid for any DM mass