# Stable Sexaquark as Dark Matter



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# Stable Sexaquark and other uds Dark Matter



How could we have missed a stable particle made of quarks?

[Hints from Astrophysics]

[Primordial Nucleosynthesis]

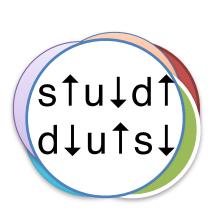
Dark-Matter to Ordinary-Matter ratio

[Detecting S dark matter]

Discovering a stable sexaquark in the lab

### Unique among multi-quark states:

Fermi statistics is compatible with a <u>totally symmetric</u> spatial wave function AND



antisymmetric (singlet) in:

color
flavor
spin
totally symmetric in space

(Most-Attractive Channel)<sup>3</sup>:

S

6-quark, Q=0, B=2

Spin-0, scalar

Flavor singlet

m<sub>S</sub> < 2 GeV???

Same quark content as H-dibaryon\* (Jaffe 1977), but different physics: **not a loosely bound di-Λ!** 

### Why consider $m_S \sim 2 m_p^*$ ? (\*1.876 GeV)

- Light quarks almost massless, i.e. relativistic
  - $m_{u,d} \approx 0$ ,  $m_s = 91$  MeV
- S has same QNs as ground state glueball
  - why not  $m_S \approx m_{glueball} + 180 \text{ MeV} = (1.5-1.7) + 0.18 \text{ GeV} \le 2 \text{ mp}$
- 3 x di-quark mass = 1.2 2 GeV
- m<sub>S</sub> < 2 (m<sub>p</sub> + m<sub>e</sub>): S is absolutely stable
- $m_S > 2 (m_p 8 \text{ MeV})$ : nuclei are stable

Interesting DM candidate

- triple-singlet (color,flavor,spin): MAC, lattice, almost all models  $=> m_S < 2 m_A$
- extensive experimental searches exclude weak-lifetime & m > 2 GeV
- → bound state exists and mass < 2 GeV  $(\tau > \tau_{Univ})$  or stable

# Stable Sexaquark Hypothesis

6	sexa- <sup>[19]</sup>	_	sen- <sup>[20]</sup>	sext-[21]	hex- <sup>[22]</sup>	hexakis- hexaplo- hexad- e.g. hexahedron	hect- <sup>[23]</sup> hectaio-	shat-	
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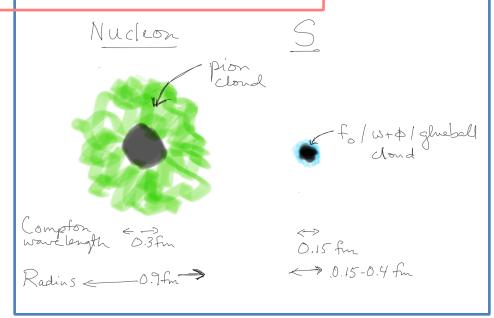
<sup>a</sup> b Sometimes Greek *hexa*- is used in Latin compounds, such as *hexadecimal*, due to taboo avoidance with the English word *sex*.

https://en.wikipedia.org/wiki/Numeral\_prefix

#### **Crucial fact:**

S does not couple to pions => much smaller than usual hadrons => hard to produce with hadrons

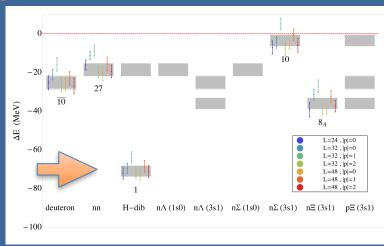
6-quark, Q=0, B=2 *Spin-0, scalar Flavor singlet*m ~ 1.7-2 GeV



### Stable S?

- $\tau > \tau$ Univ
  - $M_S < 2 m_p + 2 m_e = 1877.6 MeV \rightarrow absolutely stable$
  - Ms > 2 mp + 2 BE = 1860 MeV → nuclei absolutely stable
  - higher and lower mass may also work  $\Gamma \sim G_{F^4} \times (wave function overlap)^2$
- Lattice predicts binding (Beane+13)
  - (m<sub>q</sub> = 850 MeV so not realistic)
  - 80 MeV binding

- Experiments exclude decaying S
  - => it must be STABLE ! ;-)



### Conditions on QCD Dark Matter

- √ T<sub>DM</sub> > T<sub>Univ</sub> , cold, neutral
- ✓ primordial nucleosynthesis
- ✓ Particle must not be already excluded
  - accelerator searches
  - exotic isotopes
  - DM searches
  - indirect impacts (heating planets, helioseismology,...)
  - stability of nuclei
  - equation of state of neutron stars (and their stability)
- ✓ Correct relic density (for natural  $m_{DM} \& \sigma_{f.o.}$ )

# Relic Abundance of uds Dark Matter

### Quark-Gluon Plasma → Hadrons @~150 MeV

- Lattice QCD: crossover transition 160-140 MeV
  - T > 160 MeV: u,ū,d,d̄,s,s̄,gluons; NO vacuum condensates
  - T < 140 MeV: pions, kaons, p, $\bar{p}$ , ...; <q $\bar{q}$ > & <GG> condensates
  - Abundance relative to photons (for species in equilibrium):

Baryogenesis ⇒

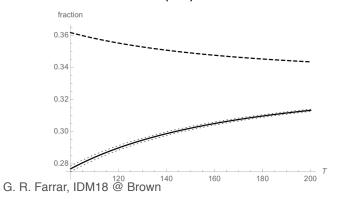
$$\eta_{\rm tot} = \eta(1 + \Omega_{DM}/(y_b\Omega_b)) \approx 4.1 \times 10^{-9}$$

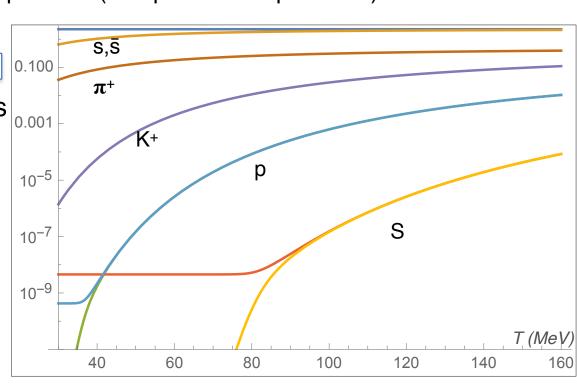
• u,d,s ratio from q masses

$$m_u = 2.118(38) \text{ MeV}$$

$$m_d = 4.690(54) \text{ MeV}$$

$$m_s = 92.52(69) \text{ MeV}$$

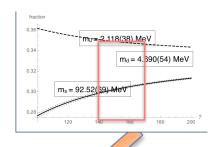


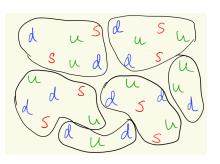


### DM to (left-over) baryon ratio

- Hypothesis: DM has u,d,s in equal numbers
  - sexaquark DM, strange quark nuggets (Witten, 1984)

$$\frac{\Omega_{DM}}{\Omega_b} = \frac{y_b \,\kappa_s \, 3f_s}{1 - \kappa_s \, 3f_s}$$





- $y_b = DM \text{ mass/m_p (mass per unit baryon number)}$
- $f_s$  = fraction of quarks that are s
- 3 f<sub>s</sub> is number uds per unit baryon # ranges from 0.964 to 0.948 as T decreases from 160 MeV to 140 MeV.
- $\kappa_s$  is efficiency of uds → DM (Boltzmann, from hyperon and S masses)

$$\kappa_s(m_S, T) = \frac{1}{1 + (r_{\Lambda,\Lambda} + r_{\Lambda,\Sigma} + 2r_{\Sigma,\Sigma} + 2r_{N,\Xi})} \quad \boxed{r_{1,2} \equiv \exp[-(m_1 + m_2 - m_S)/T]}$$

### $\Omega_{\rm DM} / \Omega_{\rm b}$

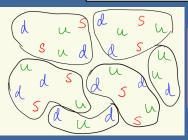
# follows from stat mech, quark masses & temperature of QGP-hadronization transition

0.32

$$\frac{\Omega_{DM}}{\Omega_b} = \frac{y_b \,\kappa_s \, 3f_s}{1 - \kappa_s \, 3f_s}$$

$$\kappa_s(m_S, T) = \frac{1}{1 + (r_{\Lambda, \Lambda} + r_{\Lambda, \Sigma} + 2r_{\Sigma, \Sigma} + 2r_{N, \Xi})}$$

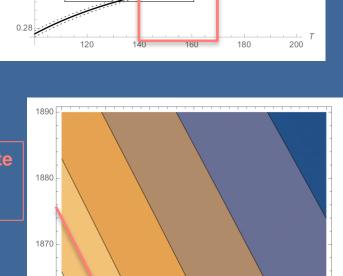
$$r_{1,2} \equiv \exp[-(m_1 + m_2 - m_S)/T]$$



Prediction is both correct AND accurate to ~20% for entire range (uncertainties cancel)

 $\Omega_{\rm DM}$  /  $\Omega_{\rm b}$  = 5.3 ± 0.1

Prediction also applies to strange quark nuggets...



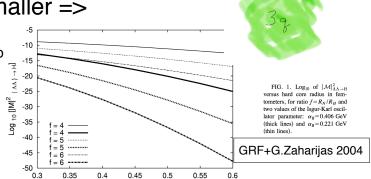
mr=-2.118(38) MeV

 $m_s = 92.52(69) Me$ 

 $m_d = 4.590(54) \text{ MeV}$ 

# S has not been discovered at accelerators because it is <u>elusive</u>

- Many negative searches, but all are inapplicable. They either:
  - looked for H-dibaryon through decays (but S is stable)
  - restricted to mass > 2 GeV (but  $m_S < 2 \text{ GeV}$ )
  - required fast production in S=-2 hypernuclei (but small overlap with baryons)
- Wavefunction overlap with baryons is very small. Extremely rare fluctuation required for  $S \Leftrightarrow \Lambda\Lambda$ ;  $S \Leftrightarrow NN$  is  $G_{F^4}$  smaller =>
  - nuclei can be stable ( $\tau > 10^{29} \, \text{yr}$ ) even for  $m_S > 2 \, m_p$
  - hard to produce in fixed target experiments
- S is similar to (much more copious) neutrons
- Promising accelerator detection strategies



c[fm]

- Apparent lack of baryon number and strangeness conservation:
  - $\cdot \Delta B = \pm 2$  with  $\Delta S = \mp 2$
- Reconstruct missing mass, e.g.:

• 
$$\Upsilon$$
 ->  $\Lambda \Lambda \overline{S}$  (+ pions)  $M_{S^2} = (p_Y - p_{\Lambda 1} - p_{\Lambda 2} - \sum p_{\pi_i})^2$ 

 $N^{\infty}V$ 

### Experimental searches so far

# Looking for Jaffe's H-dibaryon (same QN but assumed to be unstable and r~1 fm)

- Require M > 2 GeV:
  - Gufstafson+ FNAL1976: Beam-dump + tof Limit on production of neutral stable strongly interacting particle with mass > 2 GeV.
  - Carroll+ BNL 1978: No narrow missing mass peak above 2 GeV in pp -> K K X
- Require H-dibaryon decay:
  - Badier+ NA3 1986
  - Bernstein+ FNAL 1988: Limit on production of neutral with 10-8 < τ < 2 x 10-6 s</li>
  - Belz+ BNL 1996: H -/-> Λ n or Σ n [c.f., issue raised by L. Littenberg]
  - Kim+ Belle 2013: no narrow resonance in Y → A p K
- Limits from production in doubly-strange hypernuclei:
  - Ahn+ BNL 2001
  - Takahashi+ KEK 2001

#### Search for Six-Quark States

A. S. Carroll, I-H. Chiang, R. A. Johnson, T. F. Kycia, K. K. Ki,
L. S. Littenberg, and M. D. Marx
Brookhaven National Laboratory, Upton, New York 11973

and

R. Cester, R. C. Webb, and M. S. Witherell Princeton University, Princeton, New Jersey 08540 (Received 26 July 1978)

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## Cosmology & structure formation

- DM-baryon interaction: momentum transfer => slight drag on DM during structure formation
  - Dvorkin, Blum, Kamionkowski (2014), Gluscevic+Boddy (2017), Xu+18
    - Ly-alpha forest:  $\sigma < \sim 10$  mb if v-indept no problem for S
  - Buen-Abad, Marques-Tavares, Schmaltz (2015):
    - momentum transfer helps reconcile H<sub>o</sub> & σ<sub>8</sub>
  - Boring or an opportunity? To be determined...
- S-S self interactions + S-baryon interactions:
  - could have similar benefits as Self Interacting DM
    - core-cusp, "too-big-to-fail" & missing sub-halos problems.

### Galaxies & Clusters

### DM-gas scattering provides a source of heating, useful for

- Milky Way's extended hot gas halo 2 x10<sup>6</sup> K
- Quenching star formation
- Avoiding "cooling flow catastrophe" in X-ray clusters

# Key points to take home

- There may a tightly bound 6-quark state S= uuddss
  - Unique, symmetric structure ⇒ other hadrons don't provide guidance
    - mass is not driven by chiral symmetry breaking (unlike baryons)
    - constituent quark model probably completely misleading
  - If  $M_S < 2 m_p + 2 m_e$ , S is absolutely stable
- If S is stable, its an excellent Dark Matter candidate
  - Relic abundance is natural. EXPLAINS 7Li Discrepancy in BBN and Dark Matter to baryon ratio
  - Usual WIMP detection strategy isn't applicable.
  - May reconcile tension in H<sub>0</sub> &  $\sigma_8$  and explain astrophysics puzzles ("quenching", core-cusp, DM rotation curves...)
- S may be waiting to be discovered in existing Y-decays or LHC experiments... mass can be accurately measured in Y-decay exclusive final states.
- SDM will be challenging to detect, but not impossible. Astrophysical and cosmological effects may allow it to be constrained, excluded or confirmed.

# Backup Slides

## Sexaquark Discovery Strategy

Apparent lack of B and S conservation:

$$\cdot \Delta B = \pm 2 + \Delta S = \mp 2$$

- Reconstruct missing mass, e.g.:
  - $\Upsilon \rightarrow \Lambda \Lambda \bar{S}$  (+ pions)  $M_{S^2} = (p_Y p_{\Lambda 1} p_{\Lambda 2} \sum p_{\pi_i})^2$
  - LHC:  $\overline{S} + N \rightarrow \overline{\Lambda} K^+$   $M_{S^2} = (p_{\overline{\Lambda}} + p_K p_N)^2$
- Snolab nuclei: pn -> S e+ $\nu$  G<sub>F</sub><sup>4</sup>  $\tau$  > 10+29 yr

(+ pions)

~ 0.03 fm

~106V

- Y is localized source of ggg
  - $\Rightarrow$  production of S is (relatively) enhanced
- Many x 10<sup>8</sup> events collected (CLEO, Babar, Belle)
  - detectors pretty hermetic, have good mass resolution, O(10 MeV)
  - $\Lambda$  decays quickly to  $p\pi^-$  so easy to ID.  $c\tau = 8$  cm
- Can MEASURE m<sub>S</sub> via missing mass
- Very clean
  - Main bkg is K<sub>S</sub> K<sub>S</sub> K<sub>L</sub> K<sub>L</sub> (+ pions)
    - K<sub>S</sub>'s mis-ID'd as Λ's and K<sub>L</sub>'s escaping before decay: negligible for Belle
      - rare and can model accurately
      - K<sub>S</sub> K<sub>S</sub> K<sub>L</sub> K<sub>L</sub> (+ pions) *is measurable*, from K<sup>+</sup> K<sup>+</sup> K<sup>-</sup> (+ pions)
  - "Conspiracy" of missed particles producing  $\Delta B = \pm 2$ ,  $\Delta S = \mp 2$  very hard

Background does not have narrow peak in missing mass!

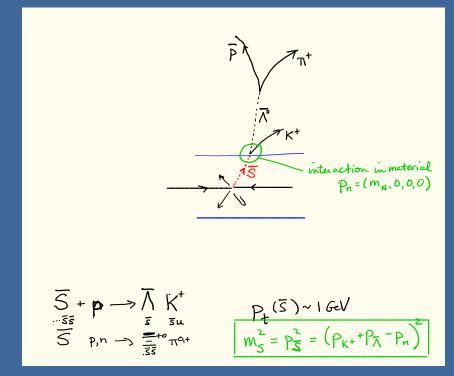
# LHC

- · Hadronic collisions: low production rate due to small wfn overlap
- Find a needle in a haystack (10<sup>11</sup> recorded events; potential for trigger >>> more
- Statistical examination of correlation between

$$\Delta B = \pm 2$$
,  $\Delta S = \mp 2$ 

· S annihilation in tracker, tag by

$$\overline{\Xi}^{+,o} \rightarrow \overline{\Lambda} \ \pi^{+,o} \ (c_{\tau} = 5\gamma \ cm) \ \overline{\Lambda} \rightarrow \overline{p} \ \pi^{+}$$
 or  $\overline{\Lambda} \ K^{+}$ 



2nd exponential in scattering-length distribution of n-like interactions, due to S