

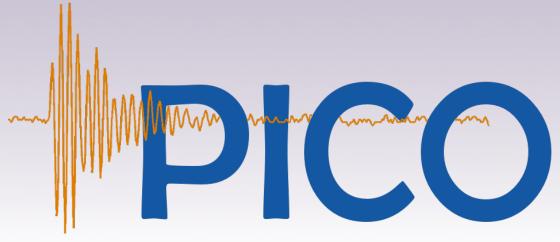


Electron Recoil Bubble Nucleation in PICO Bubble Chambers

Dan Baxter

July 23, 2018

Identification of Dark Matter 2018





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R. Filgas, I. Stekl



J.I. Collar, A. Ortega



S. Chen, M. Laurin, J.-P. Martin, A. Plante, A.E. Robinson, N. Starinski, F. Tardif, D. Tiwari, V. Zacek, C. Wen Chao,



I. Lawson



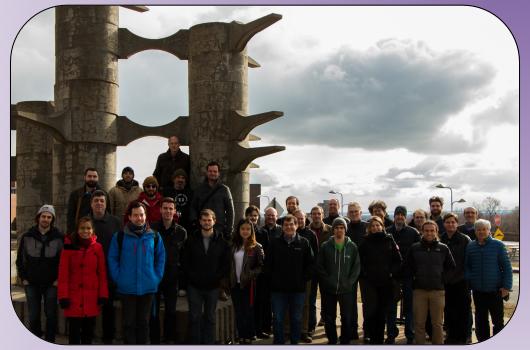
M. Ardid, M. Bou-Cabo, I. Felis



M. Bressler, R. Neilson



P.S. Cooper, M. Crisler, W.H. Lippincott, A. Sonnenschein



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D. Baxter, C.E. Dahl, M. Jin, J. Zhang



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SOUTH BEND

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M. Das, S. Sahoo, S. Seth



D. Maurya, S. Priya, Y. Yan





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I. Arnquist, C. Cowles, C.M. Jackson, B. Loer, K. Wierman



J. Farine, A. Le Blanc, C. Licciardi, O. Scallon, U. Wichoski



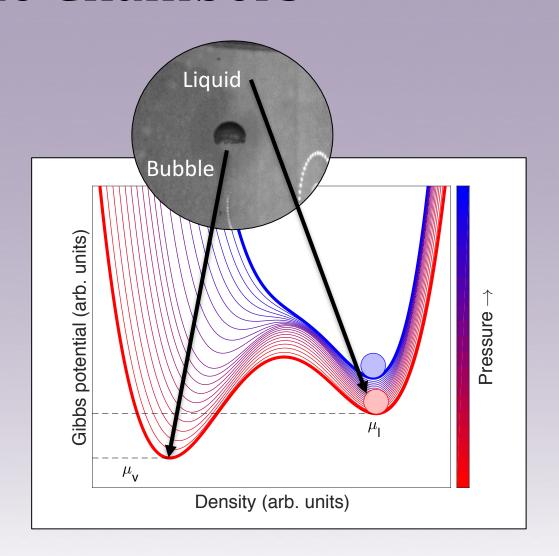
B. Broerman, G. Cao, K. Clark, G. Giroux, C. Hardy, C. Moore, A. Noble





Bubble Chambers

- Superheat a fluid by lowering pressure under constant temperature
- Metastable state awaits energy deposition
- Interaction nucleates small gas pocket that grows
- Trigger on bubble with cameras







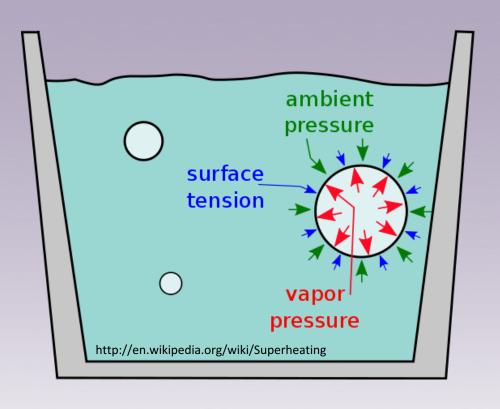
Threshold Determination

 Critical radius set by fluid properties and run conditions

$$P_{v} - P_{l} \ge \frac{2\sigma}{r_{c}}$$

 Critical radius determines energy threshold

$$E_T = E(T, P, fluid)$$





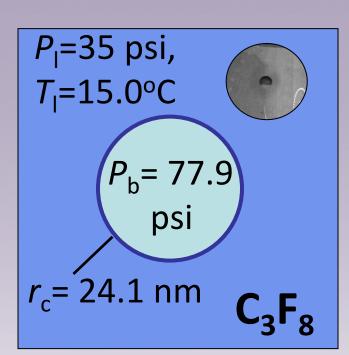


Threshold Calculation

Surface energy, Bulk energy, Reversible Work

$$E_{T} = 4\pi r_{c}^{2} \left(\sigma - T\left(\frac{\partial\sigma}{\partial T}\right)_{\mu}\right)$$
1.57 keV
$$+ \frac{4\pi}{3} r_{c}^{3} \rho_{b} \left(h_{b} - h_{l}\right)$$
1.95 keV
$$- \frac{4\pi}{3} r_{c}^{3} \left(P_{b} - P_{l}\right)$$
-0.15 keV

Heat threshold = 3.37 keV







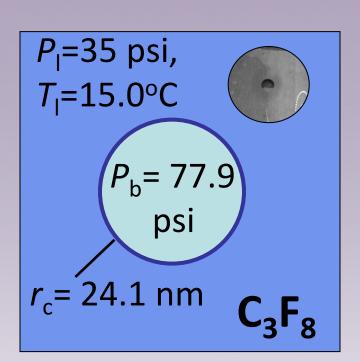
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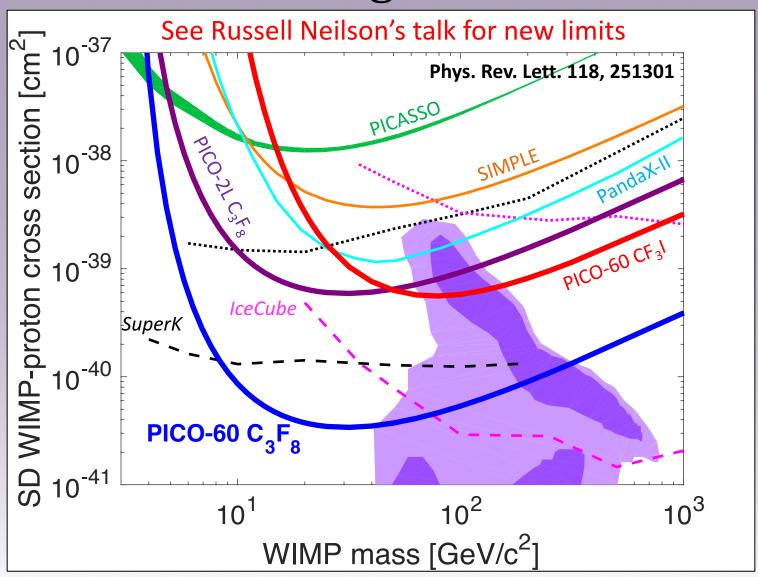
Minimum Work:

$$W = \int_0^{r_c} 4\pi r^2 dr \left(\frac{2\sigma}{r} - (P_b - P_l) \right) = \frac{4}{3}\pi\sigma r_c^2 \approx 50 - 75 \text{ eV}$$





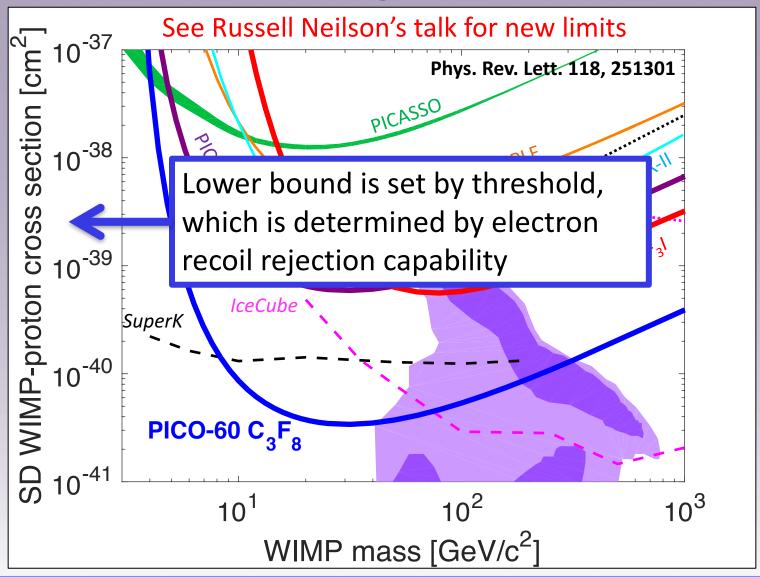
Existing Limits







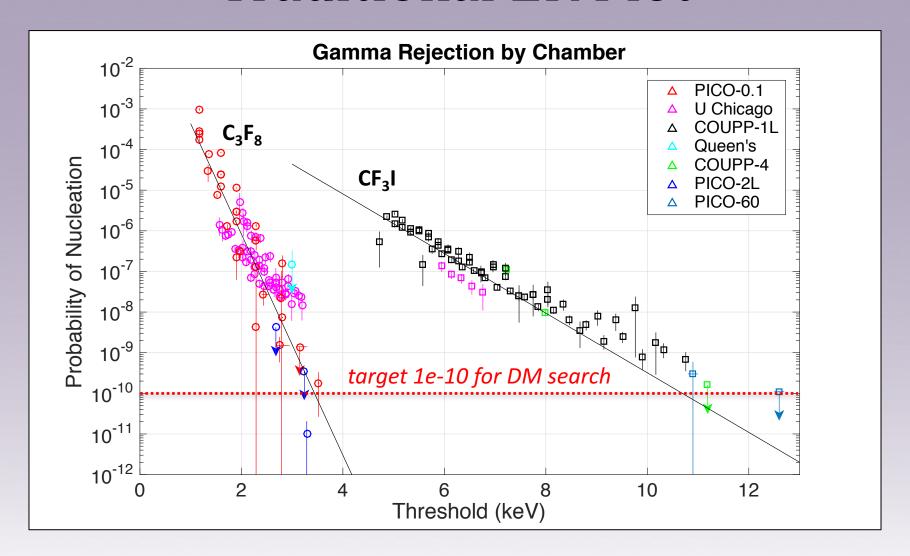
Existing Limits







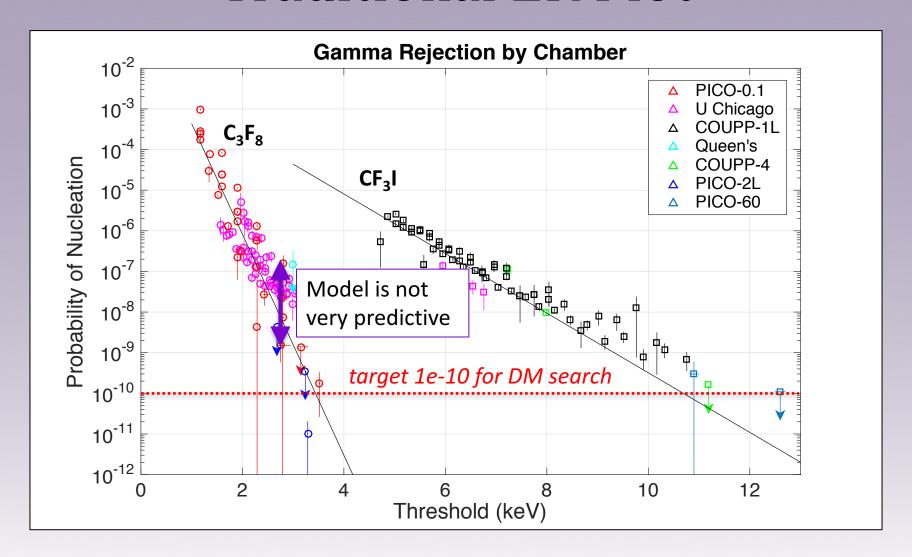
Traditional ER Plot







Traditional ER Plot







Probability of Nucleation

(Historically)

Per what...

y-axis

– Photon Interaction: each photon interaction has some probability to nucleate a bubble

Nucleation by... (scales with)

x-axis

Heat: Seitz Model describes threshold (few keV)



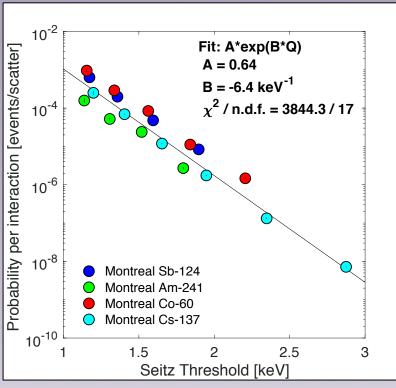


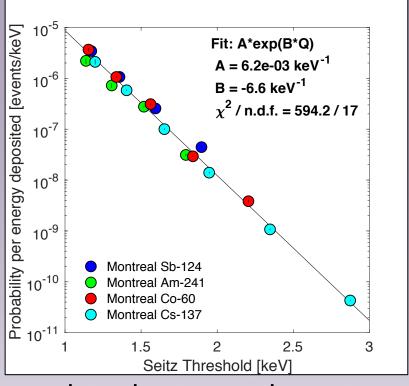
Probability of Nucleation

- Per what...
 - Photon Interaction: each photon interaction has some probability to nucleate a bubble
 - Track length: each delta electron has some probability to nucleate a bubble
 - (Scales with energy deposited)
- Nucleation by... (scales with)
 - Heat: Seitz Model describes threshold (few keV)
 - Cavitation: Only minimum work to create bubble is required (less than 100 eV)









- Only statistical error bars are shown, but there are almost no systematic errors between these data
- Nucleation by delta electrons (~total energy deposited) was actually put forward to describe classical bubble chamber nucleation (Tenner 1963)

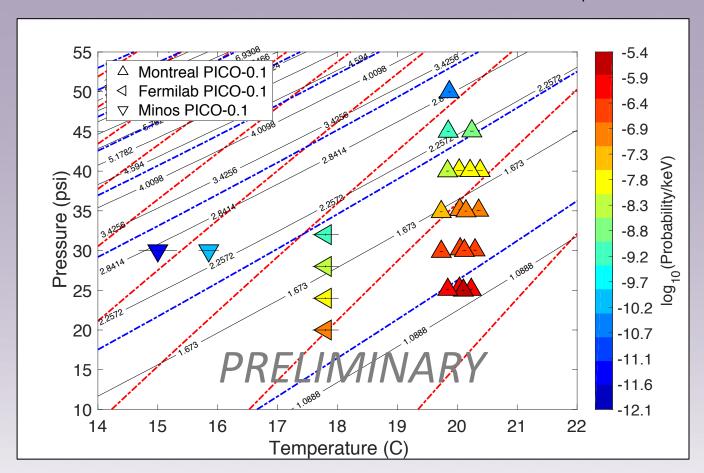


Northwester Nucleation by delta electrons

Q = Seitz Threshold

W = Minimum Work

 r_1 = Critical Length Scale



Contours of Constant

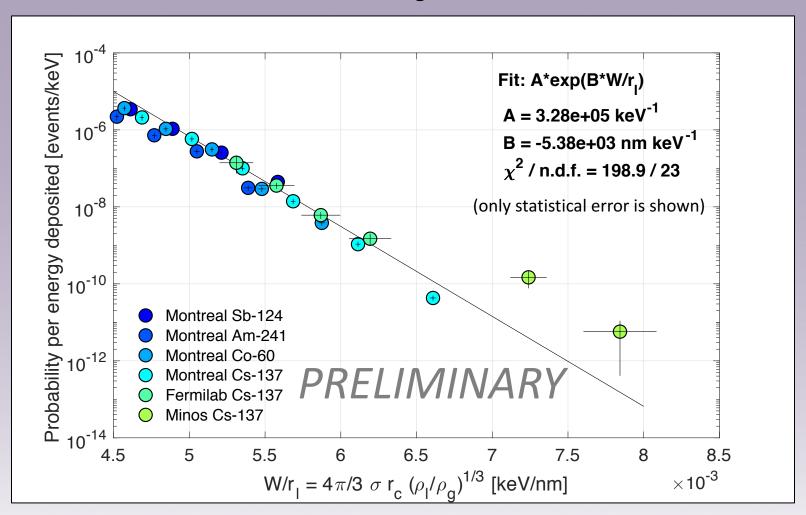
- Black: Q
- Blue: Q/r_I
- Red: W/r_I
- something else?

Constant minimum work per radius is best fit





Nucleation by cavitation

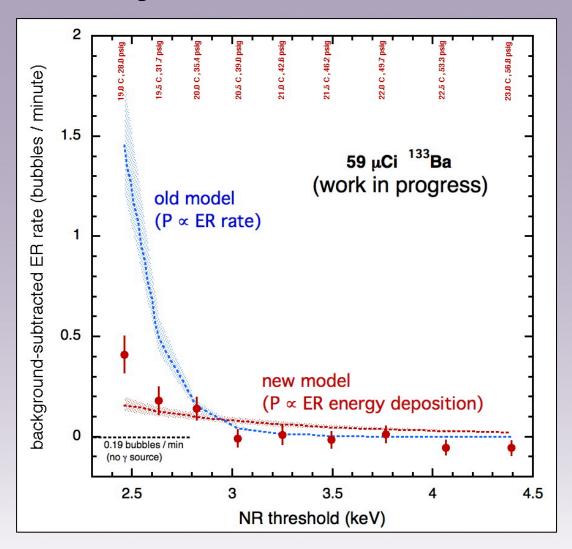






Nucleation by cavitation

- Beginning to test at UofChicago with the GUNTER chamber (first results)
- Initial results favor cavitation threshold by delta electrons

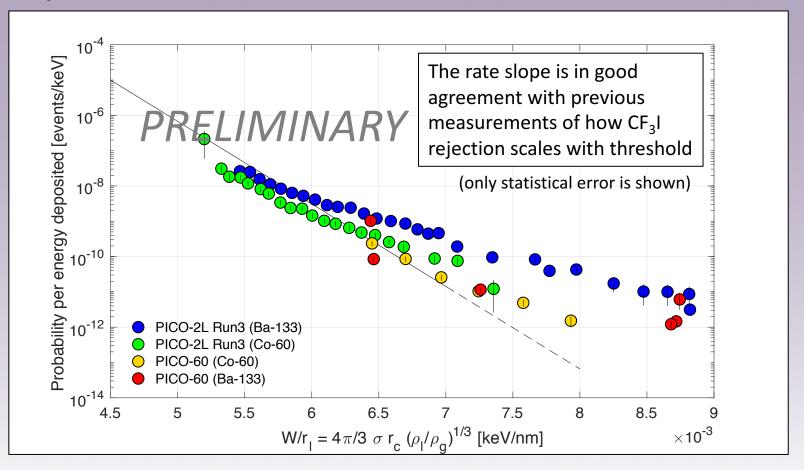






Anomalous Behavior

 In the C₃F₈ runs of both PICO-60 and PICO-2L Run 3, we expect some residual iodine







Auger Cascades

- An Auger cascade occurs when an inner shell vacancy results in the emission of many lowenergy x-rays and electrons originating from the same atom
- In the presence of heavy contamination, binding energy release through an Auger cascade is:
 - Over the heat threshold (for K-shell iodine, 30keV)
 - 2. Local on the scale of the critical radius



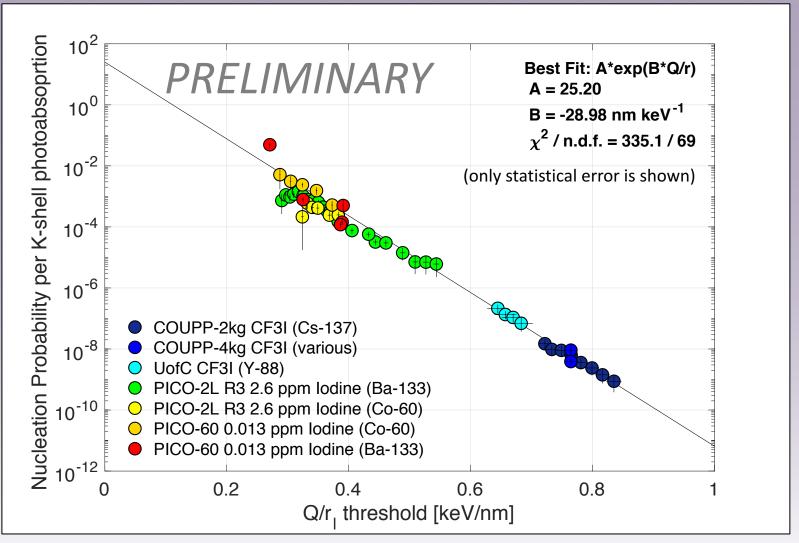


Auger Cascades

- For our existing dark matter results, only contamination at <ppmv levels
 - In situ calibrations still account for this
- Could be present in any chamber which previously saw iodine (CF₃I)
 - PICO-60: first run was with CF₃I -> expect residual cross-contamination
 - PICO-2L: third run filled with C_3F_8 from an old, empty CF_3I bottle -> expect cross-contamination
- Should also be the dominant nucleation mechanism in CF₃I -> constrains normalization











Assay Results

- Assays are in order-of-magnitude agreement with the best fit values
- We are still working on correctly determining the error in each of these measurements

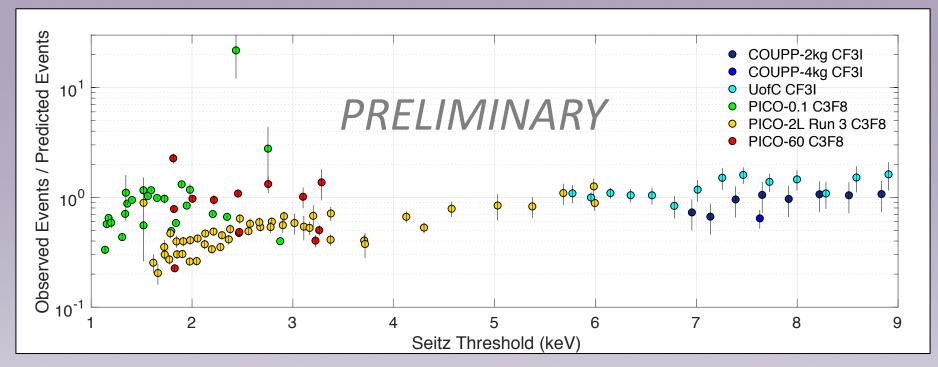
Sample	Source	ICP-MS (PNNL)	GC-MS (PNNL)	Best Fit
PICO-2L Run 3*	Old CF3I storage	0.64 ppmv	0.96 ppmv	2.6 ppmv
PICO-60 Run 2	Run 1 w/ CF3I	0.020 ppmv	0.09 ppmv	0.013 ppmv

^{*}Note: The PICO-2L sample sat in a plastic collection back for months before being analyzed. Thus these assays should be treated as lower bounds on the concentration in the detector itself.





Putting it together



- We now have a model that is reasonably predictive to an order of magnitude
- In C₃F₈, data points towards nucleation by cavitation through delta electrons
- Heavy contamination (like iodine) allows for more efficient nucleation through Auger cascades
 - Even in PICO-60, iodine contamination is under control and understood
 - Future chambers from PICO-40L and beyond will NOT have previous exposure to CF_3I





Probability of Nucleation

In pure C₃F₈

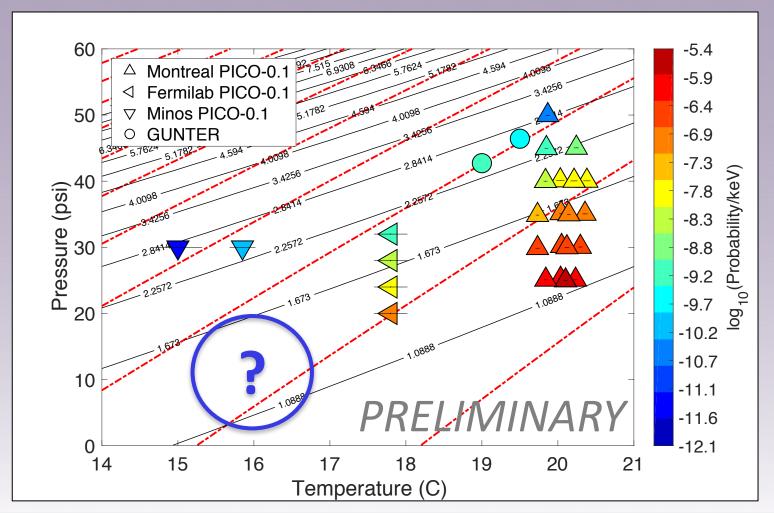
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What does this mean?

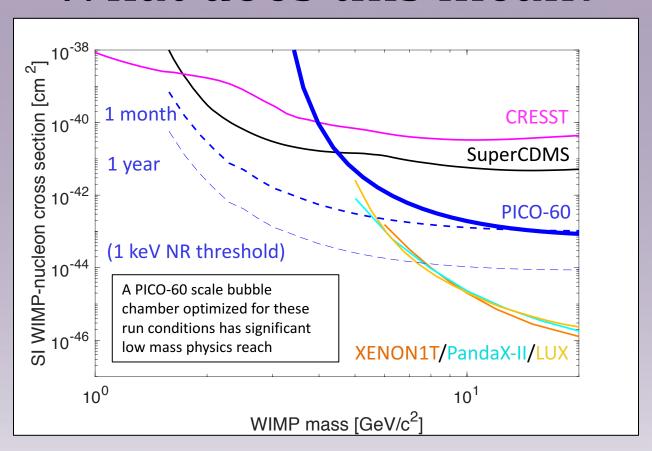
Black: NR (heat) Threshold Red: ER (cavitation) Threshold







What does this mean?



Different threshold for ER nucleation allows <u>lower</u> NR thresholds with same ER rejection

Prefer operation at lower pressure & temperature





Paper coming soon* to an arXiv near you...