



清华大学
Tsinghua University



Direct Dark Matter Search with CDEX at CJPL

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Tsinghua University

On behalf of CDEX Collaboration

2018.7.25



BROWN UNIVERSITY



中国锦屏地下实验室

China Jinping Underground Laboratory

清华大学·雅砻江流域水电开发有限公司

OUTLINE



- Introduction to CDEX
- Recent status of CDEX-1 and CDEX-10
- R&D of key technologies
- Future plan of CDEX
- Summary

China Dark matter EXperiment

- Formed in 2009, ~70 scientists and graduate students;
- Direct detection of light DM by P-type Point-Contact Ge detectors.



(new member, Rn issues)

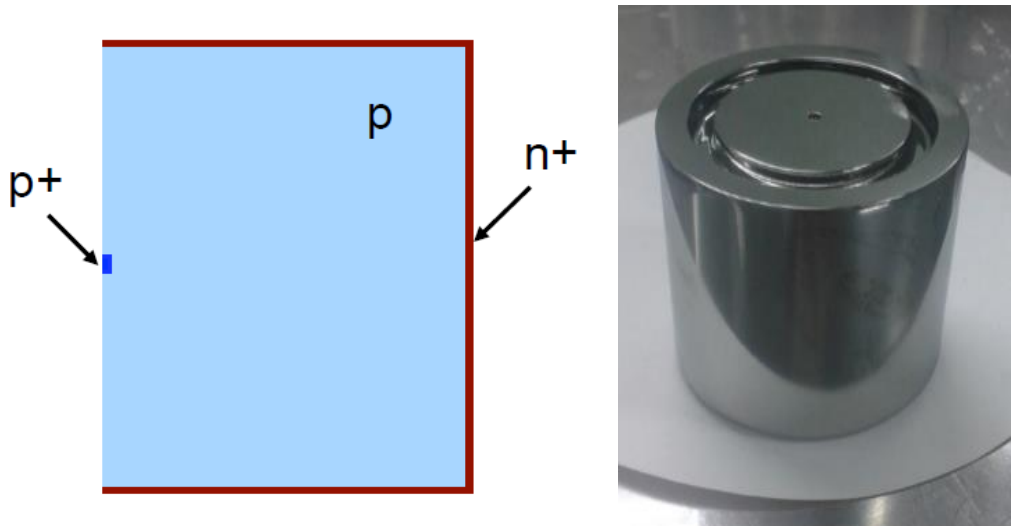


+ TEXONO Collaboration

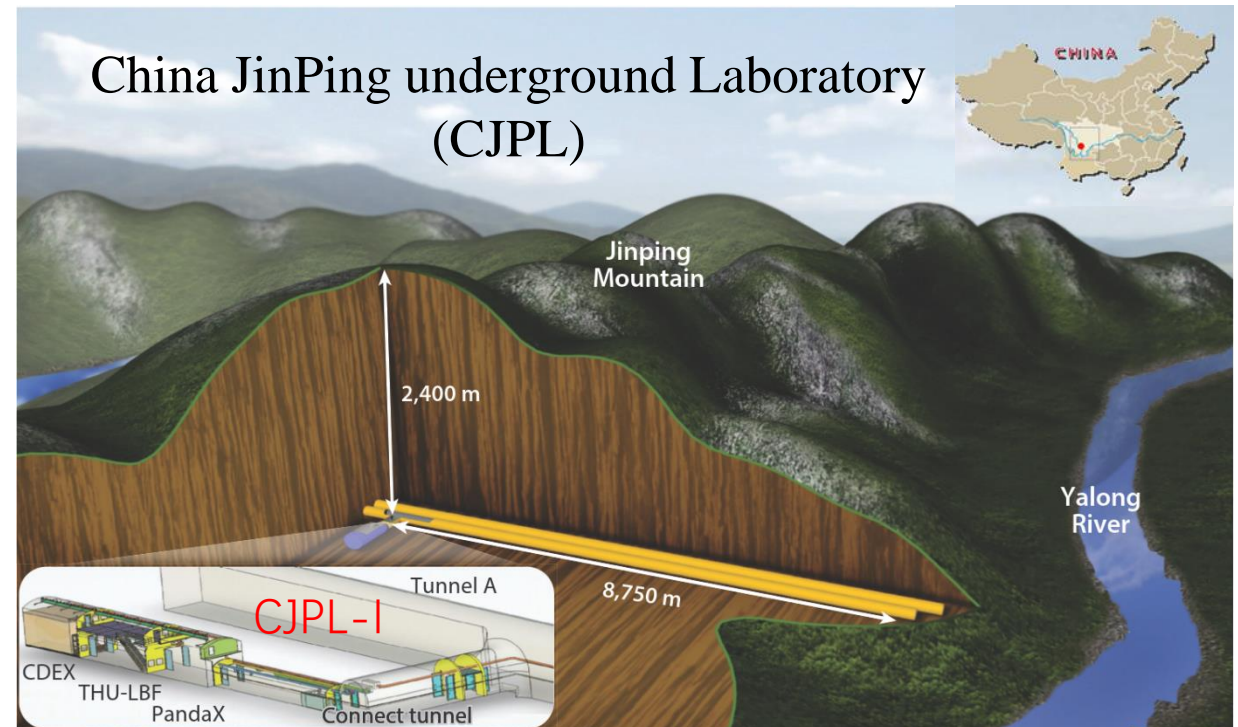
CDEX Stages

- DM detection w/ Ge prepared since 2003 and started in 2005 in Y2L (5g);
- CDEX-1: Development of PPC Ge detector, bkg understanding, since 2011;
- CDEX-10: Performances of Ge array detector immersed in LN_2 , since 2016;
- CDEX-10X: Home-made Ge detector and Ge crystal growth;

P-type Point-Contact(PPC)
Germanium detector

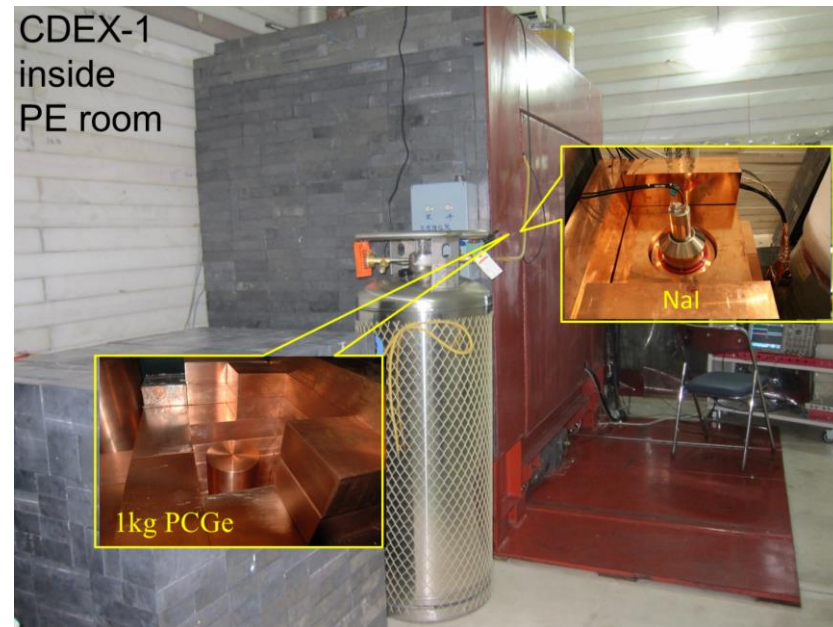
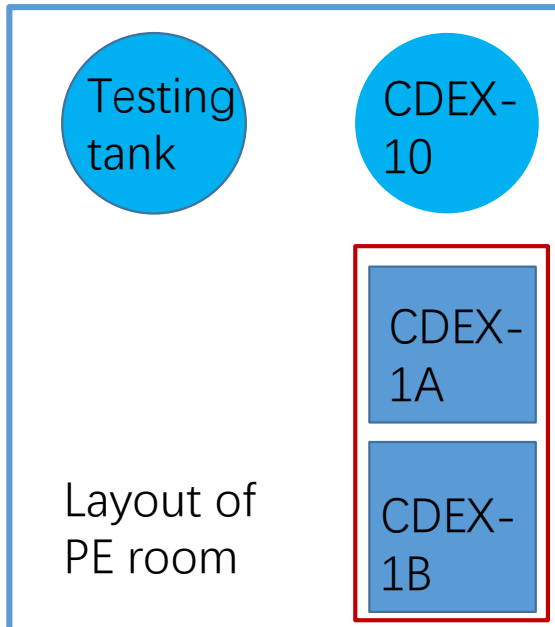


Direct detection of low-mass WIMPs w/ Ge detector at CJPL.

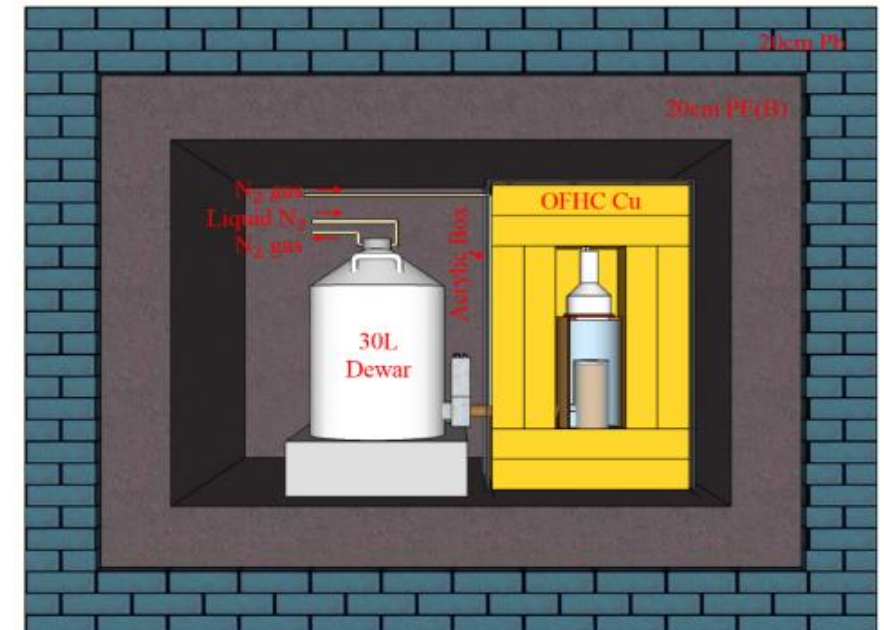


CDEX-1 stage

- 2 sub-stages: CDEX-1A(prototype, 2011)→1B(upgraded, 2013);
- Traditional single-element $\sim 1\text{kg}$ PPC Ge detector;
- Low-bkg Pb&Cu passive shield + NaI veto detector;
- Located in PE room at CJPL-I;



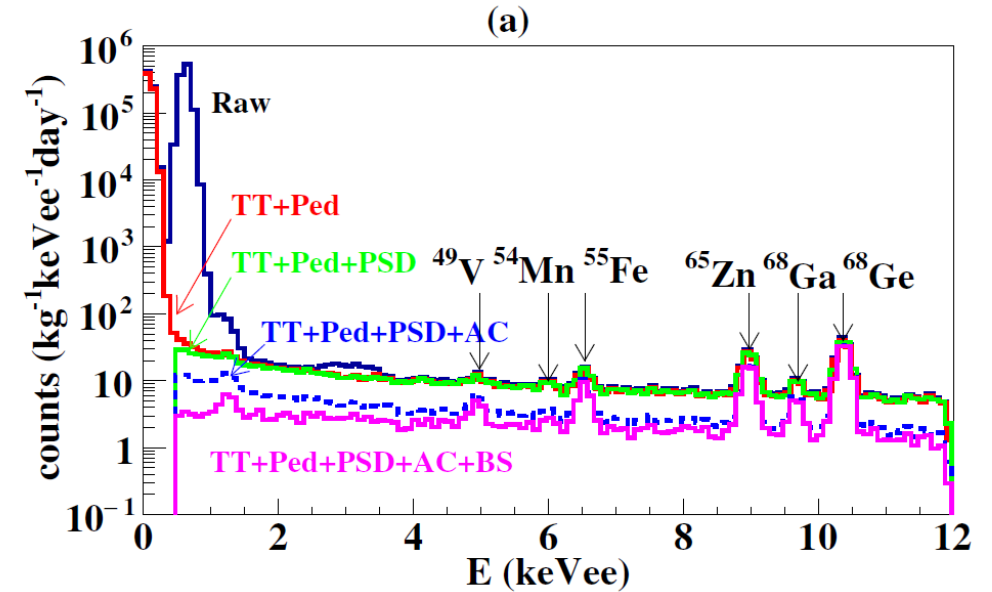
CDEX-1A&B : 1kg PPC Ge \times 2



CDEX-1A Results

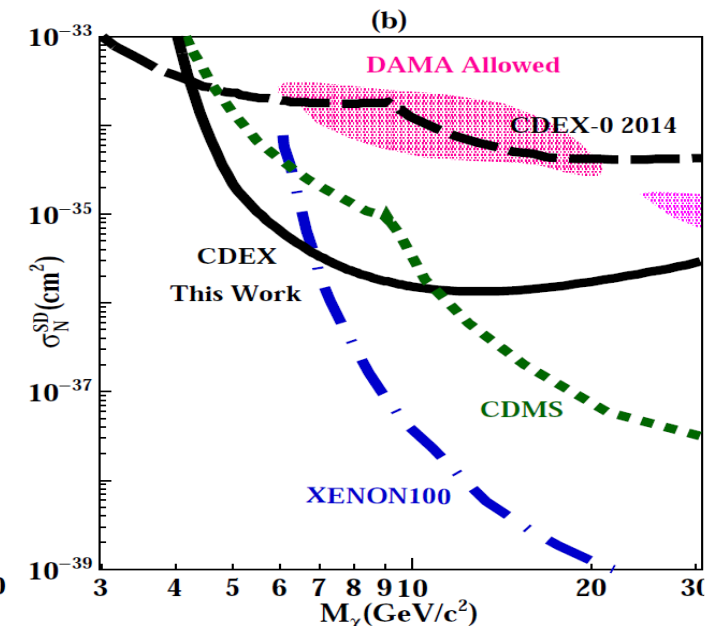
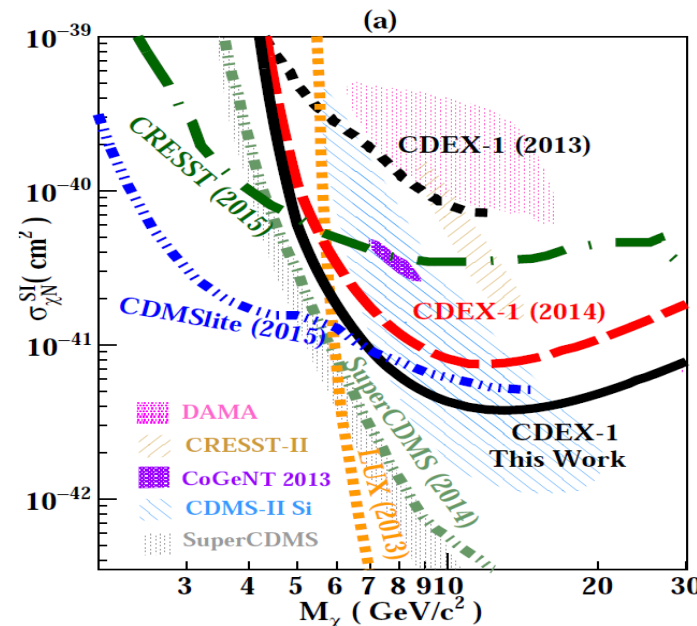


- >500 days run, ~336 d·kg dataset;
- Energy threshold: 475 eVee;
- Bulk/Surface disc. to cut events with slow rise-time and partial charge collection;
- K/L X-rays from Cosmogenic nuclides to trace crystal history;



- SI sensitivity improved;
- SD best below 6 GeV then;

See *PRD93, 092003, 2016*

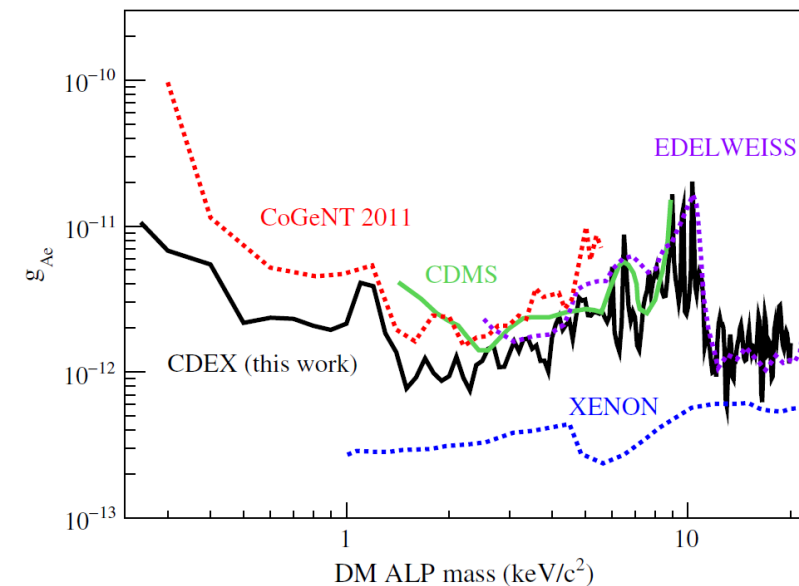
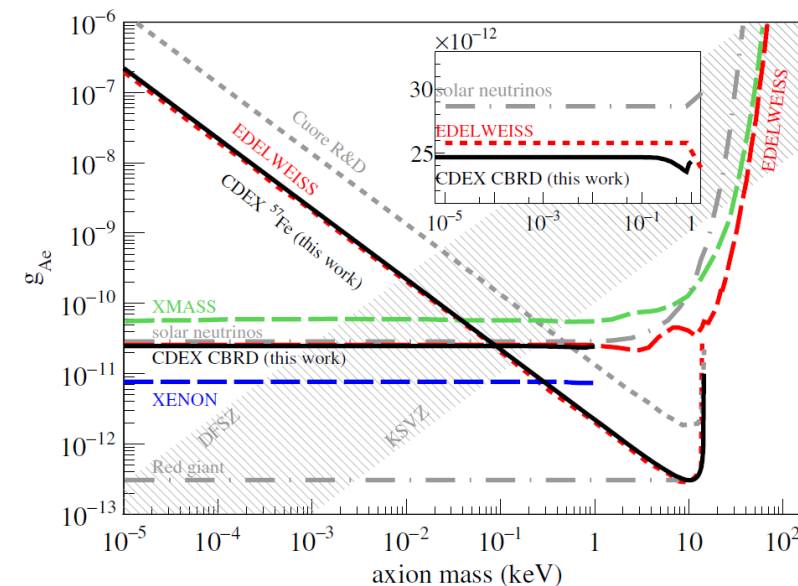


CDEX-1A Results



- Axion (335.6 kg day data)
 - Solar axions : CBRD processes and ^{57}Fe M1 transition;
 - ALPs: more stringent constraint below 1keV;

PRD95, 052006, 2017

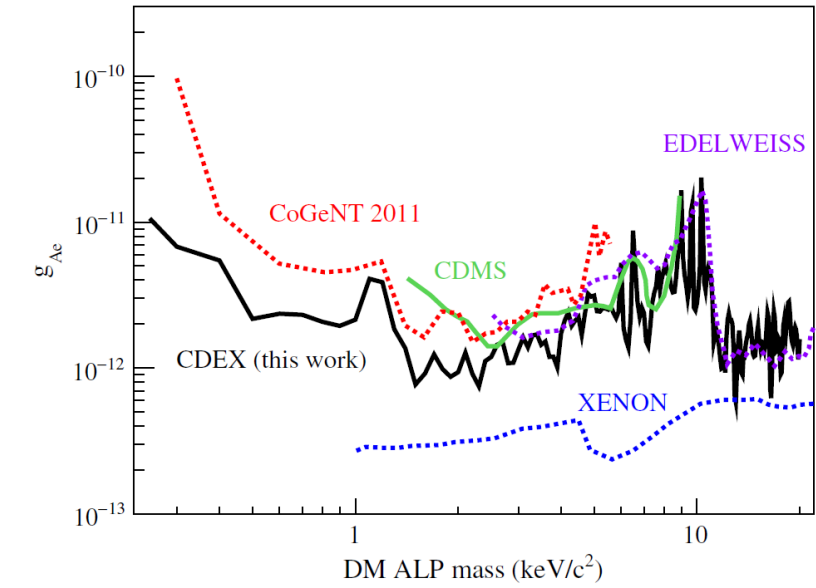
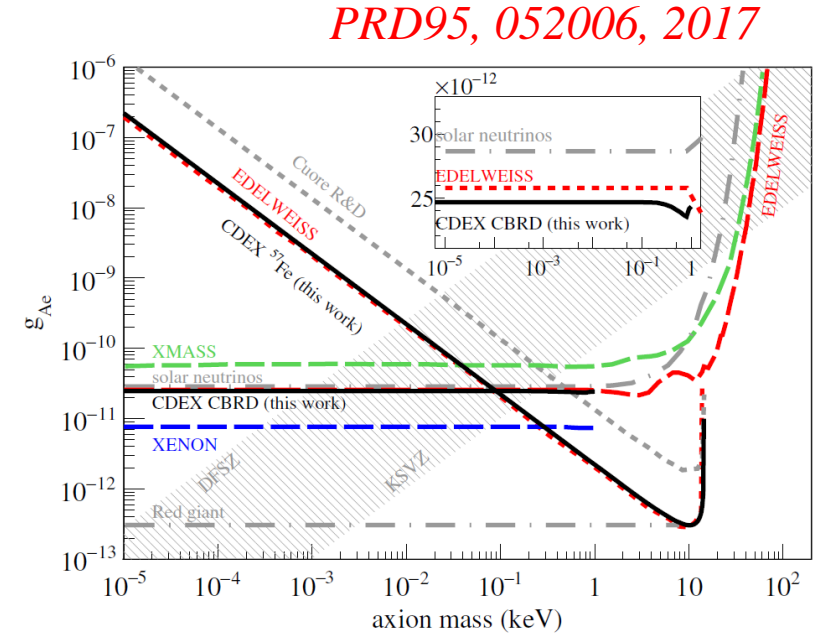
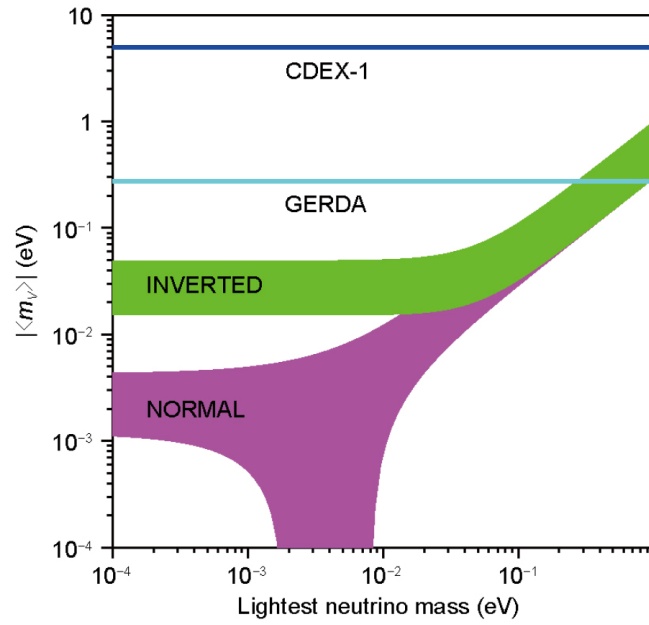
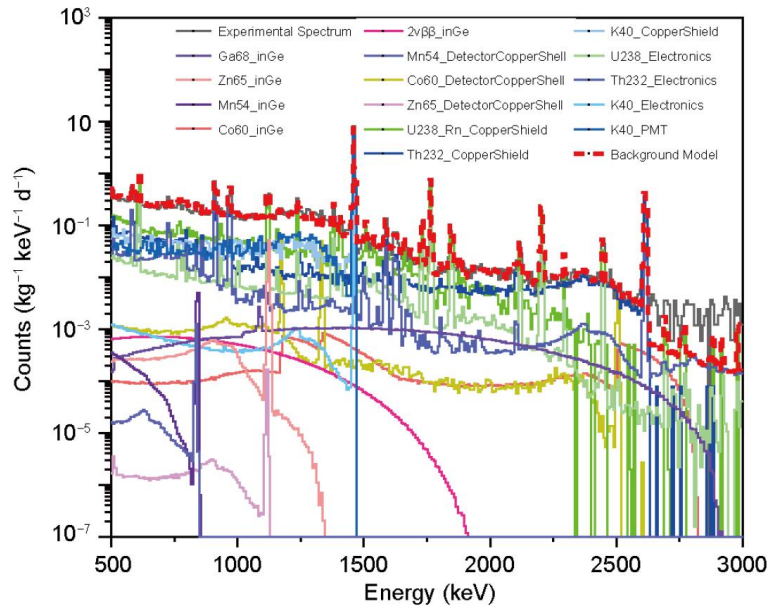


CDEX-1A Results



- Axion (335.6 kg day data)
 - Solar axions : CBRD processes and ^{57}Fe M1 transition;
 - ALPs: more stringent constraint below 1keV;
- $0\nu\beta\beta$ (304 kg·day data)
 - Natural Ge crystal;

$$T_{1/2}^{0\nu} \geq 6.43 \times 10^{22} \text{ yr, } 90\% \text{ C.L.}$$

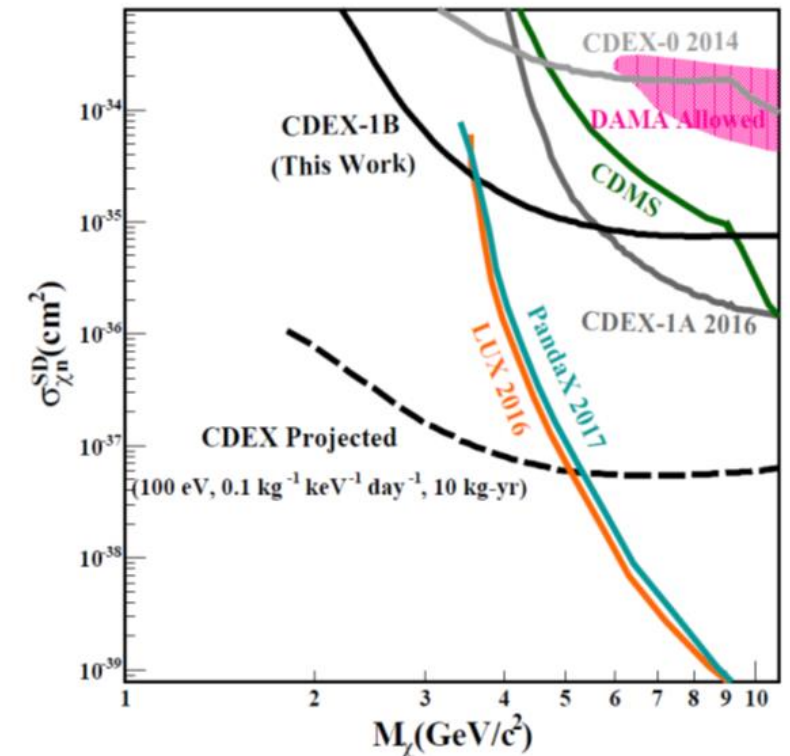
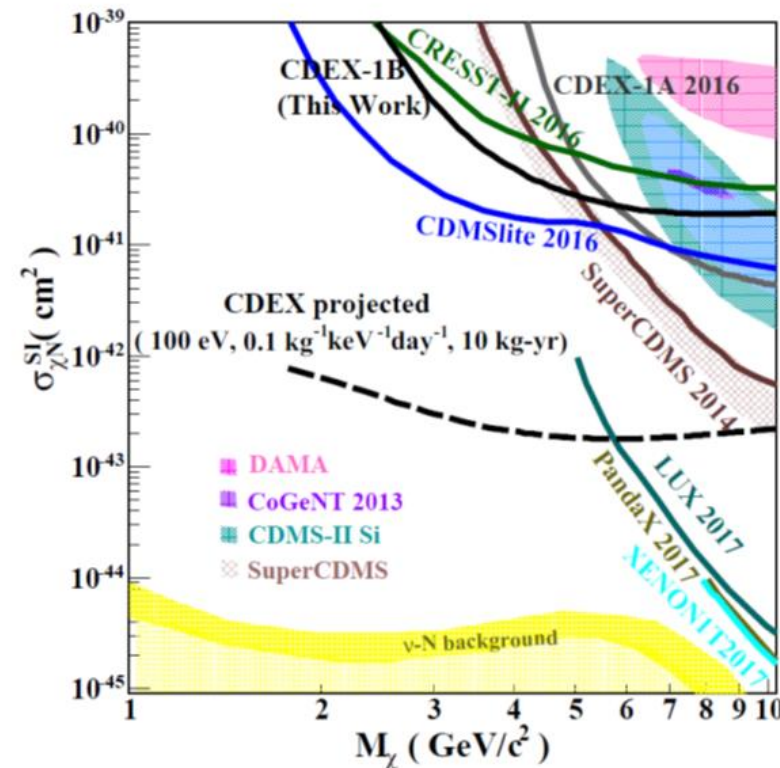


CDEX-1B Results

- Detector upgraded w/ lower JEFT noise and material bkg;
- Run 3.3 years, totally 737.1 kg d exposure;
- Achieving 160 eVee energy threshold;
- Sensitivity improved and extending to $2\text{GeV}/c^2$.

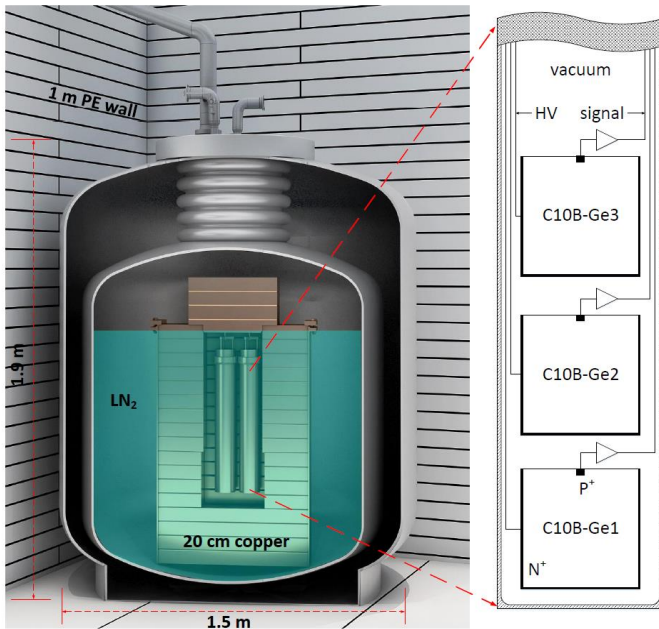
Chinese Physics C 42, 023002, 2018

Detector	FWHM of pulser
CDEX-1A	130 eVee
CDEX-1B	80 eVee

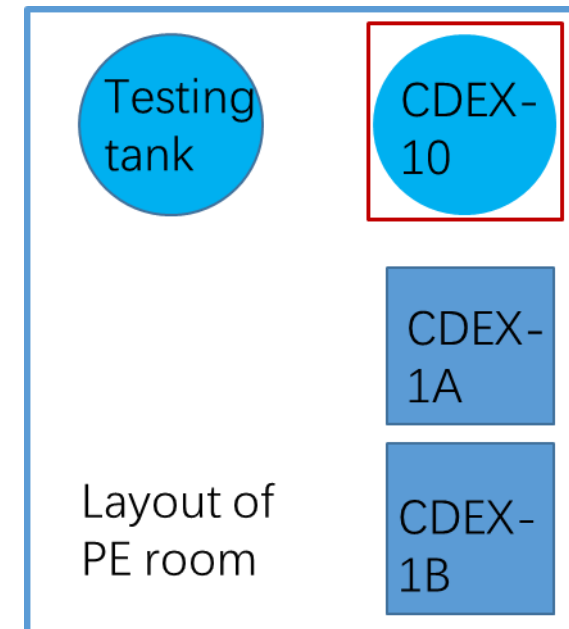


CDEX-10 stage

- Array detectors: 3 strings with 3 det. each, ~10kg total;
- Direct immersion in LN_2 ;
- Prototype system for future hundred-kg to ton scale experiment
 - Light/radio-purer LN_2 replacing heavy shield i.e. Pb/Cu;
 - Arraying technology to scalable capability;



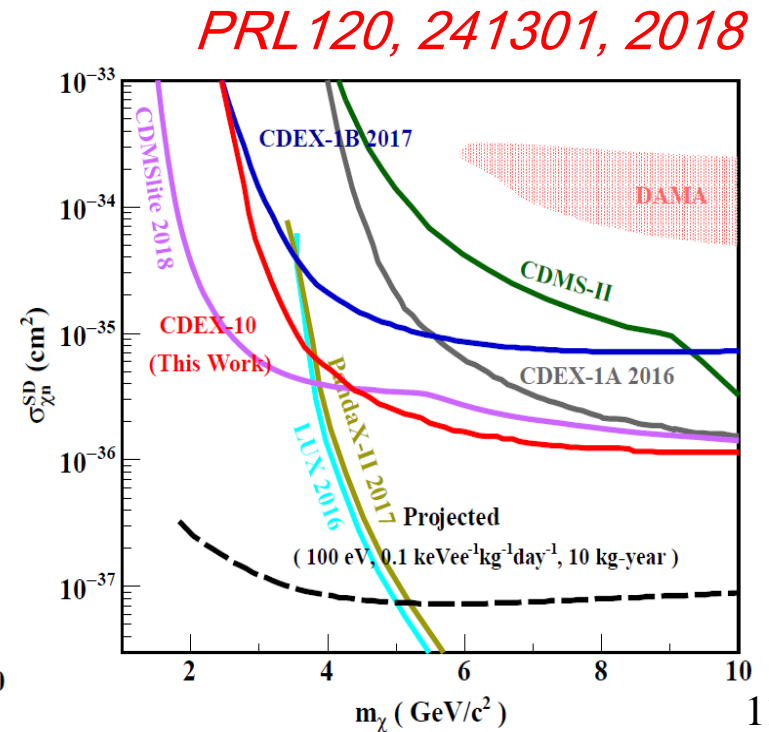
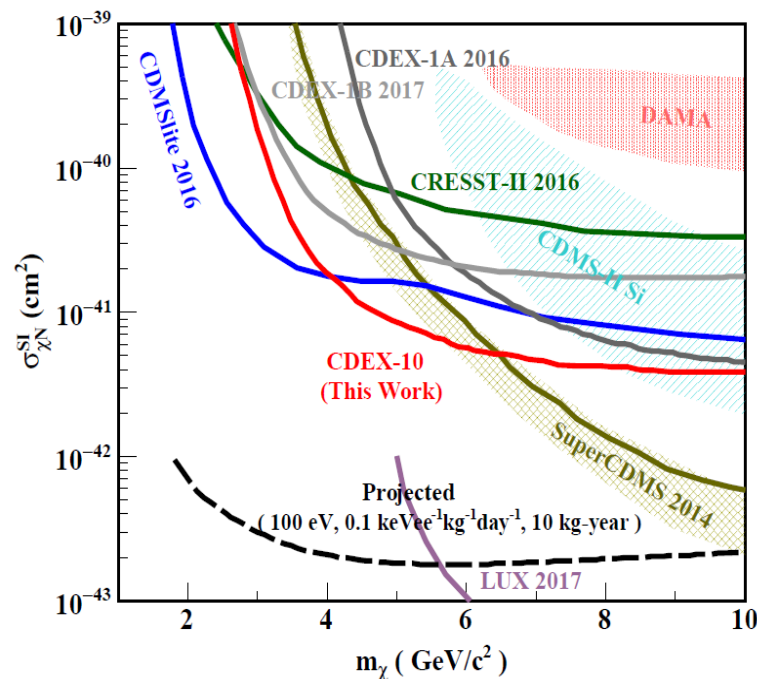
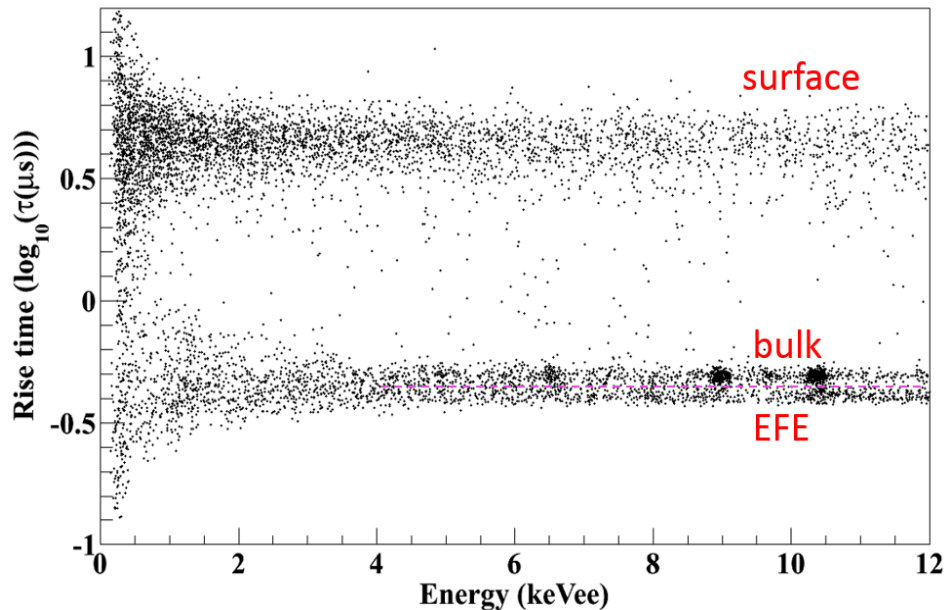
CDEX-10: ~10kg PPC Ge array



CDEX-10 First Results



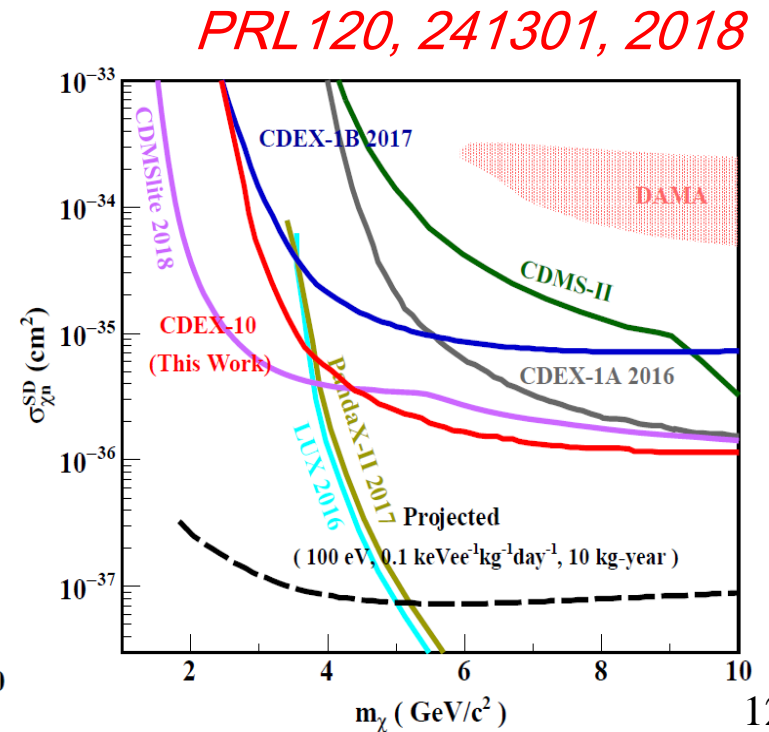
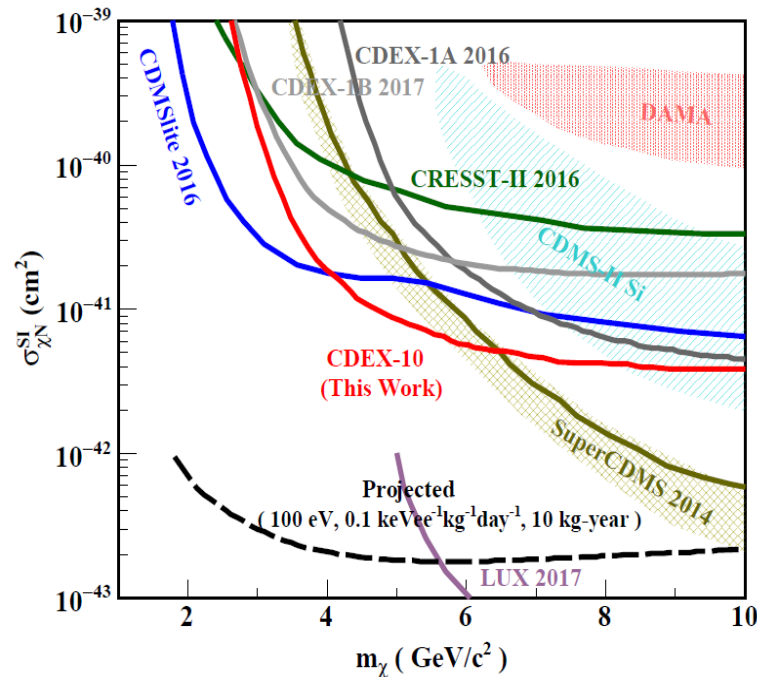
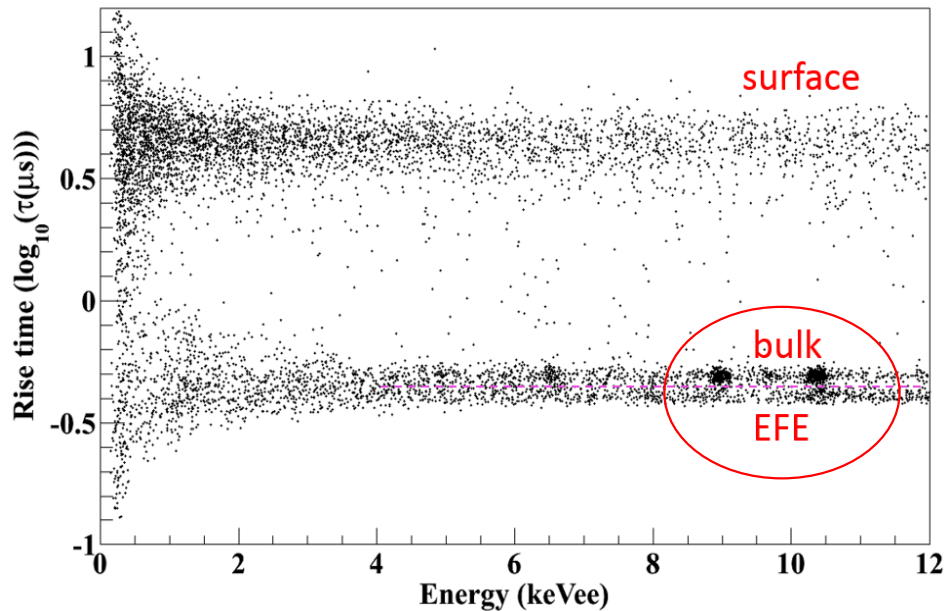
- First results from 102.8 kg day exposure w/ Eth 160eV;
- Bkg level: 2 cpkkd @ 2-4 keV;
- New SI limit on 4-5 GeV/c².



CDEX-10 First Results



- First results from 102.8 kg day exposure w/ Eth 160eV;
- Bkg level: ~ 2 cpkd @ 2-4 keV;
- New SI limit on 4-5 GeV/c²;
- Ultra-fast events observed in bulk band (See L. T. Yang's talk Thu. pm);

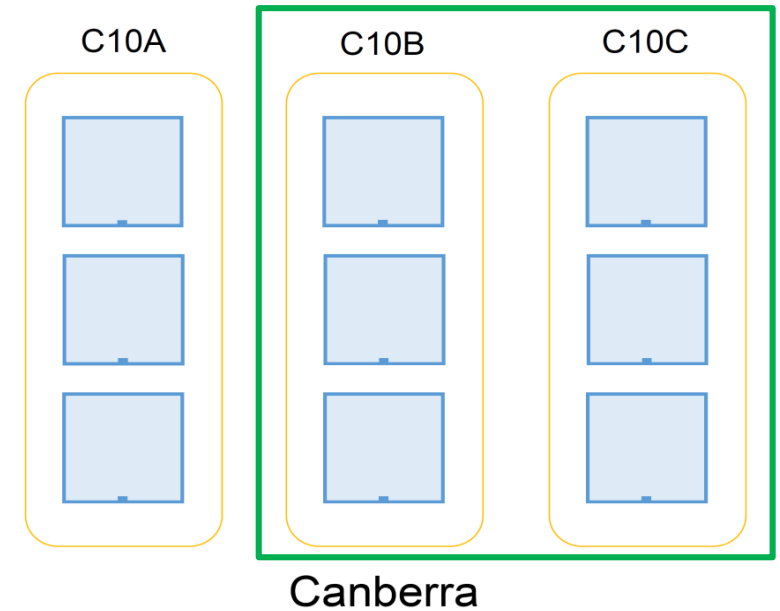


PRL 120, 241301, 2018

CDEX-10 Detectors



- C10A-1 upgrade:
 - new Cu canister;
 - new front-end electronics...
- Testing underway @ CJPL-I



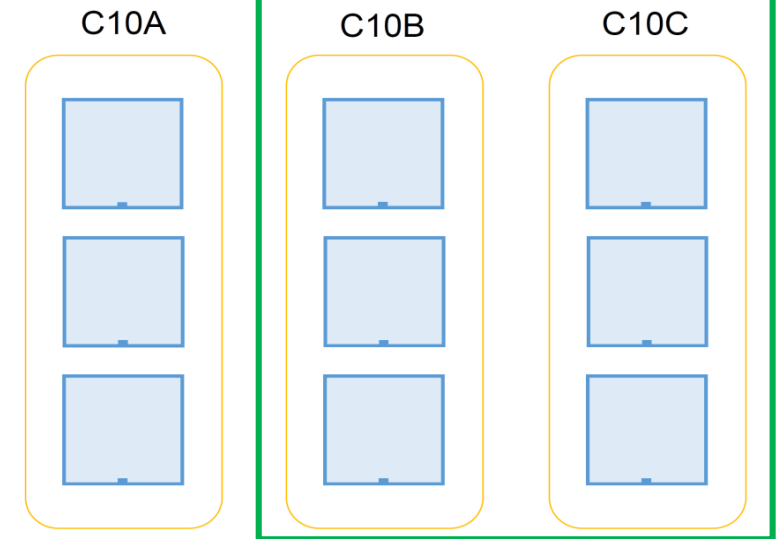
CDEX-10 Detectors



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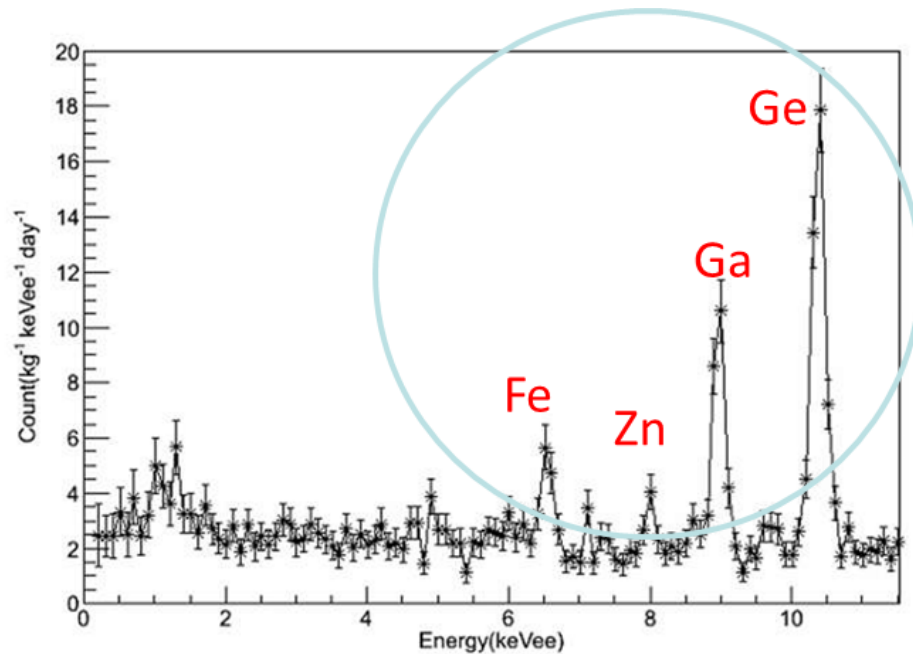
Preliminary

C10A-1 parameters	
Diameter	62mm
Height	62mm
High Voltage	+2500V
FWHM	70eV @Pulser
	430eV @ 122keV

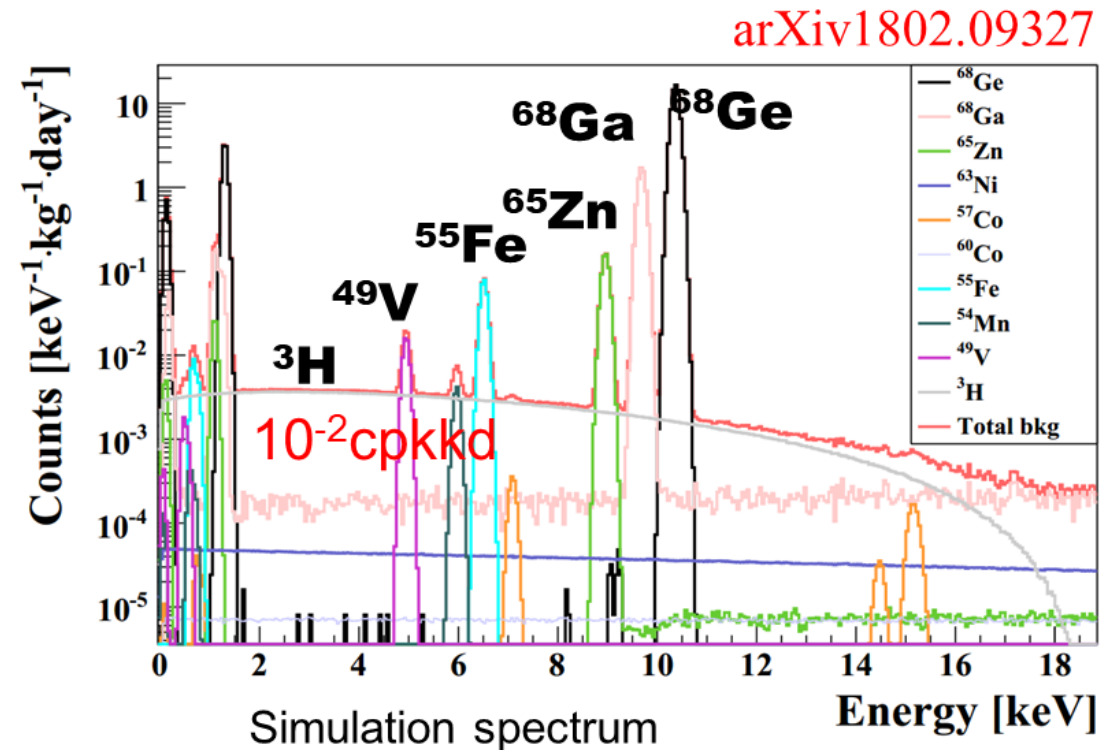


Cosmogenic bkg of Ge crystal

- Long-time ground preparation of detector induces high cosmogenic bkg;
- Based on simulation, 2 months ground fabrication and transportation could decrease the ^3H continuous bkg level to $\sim 10^{-2}\text{cpkkgd}$ @ 2-4 keV.



CDEX-10 background spectrum



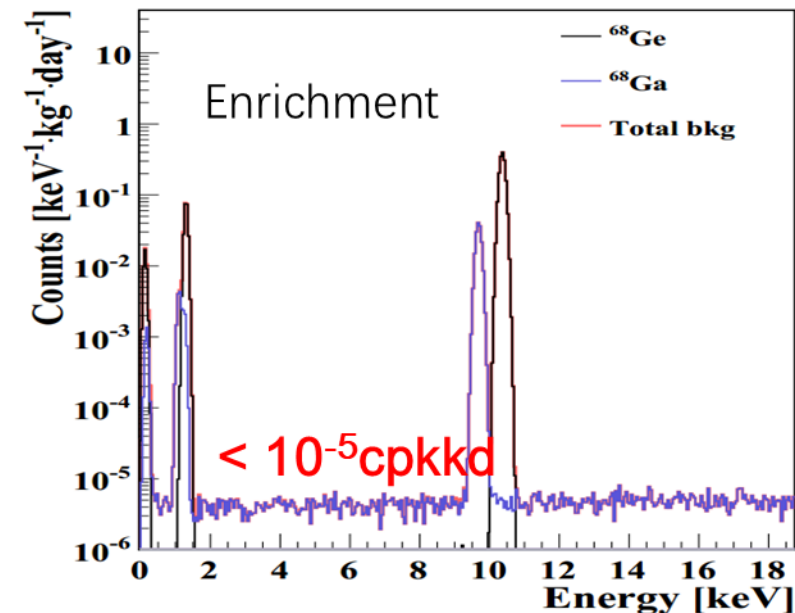
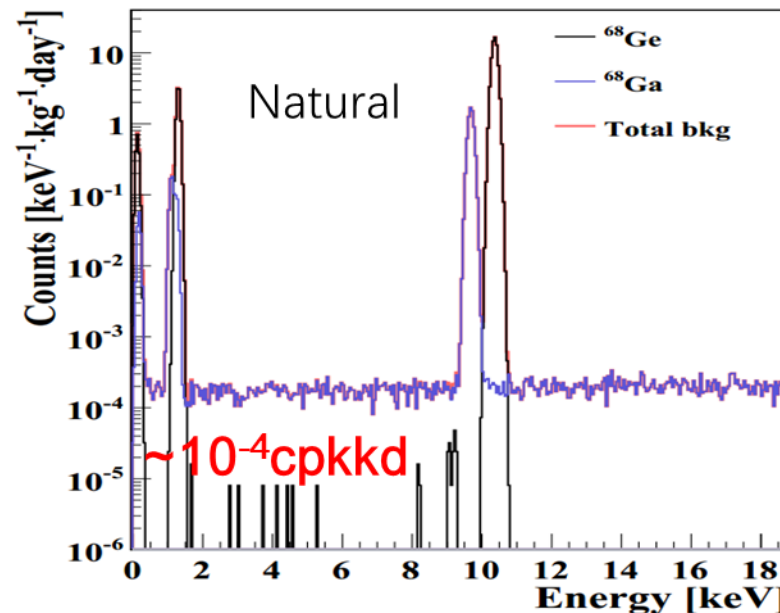
Simulation spectrum

Cosmogenic bkg of Ge crystal



- **Underground** germanium crystal growth and detector fabrication could dramatically decrease the cosmogenic bkg from **non-Ge isotopes**, such as ^3H , ^{65}Zn (... $^{68}\text{Ge}/^{68}\text{Ga}$ left);

Simulation spectrum: UG + 3 years cooling



arXiv:1802.09327

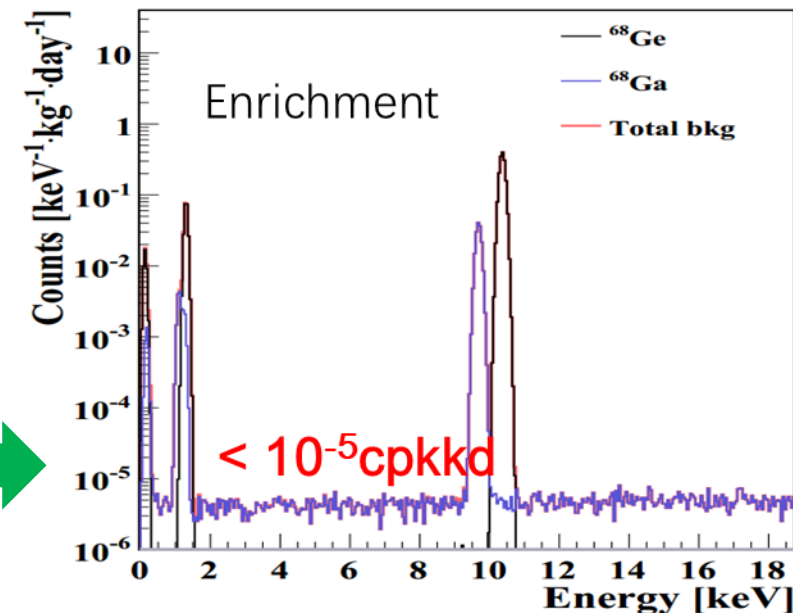
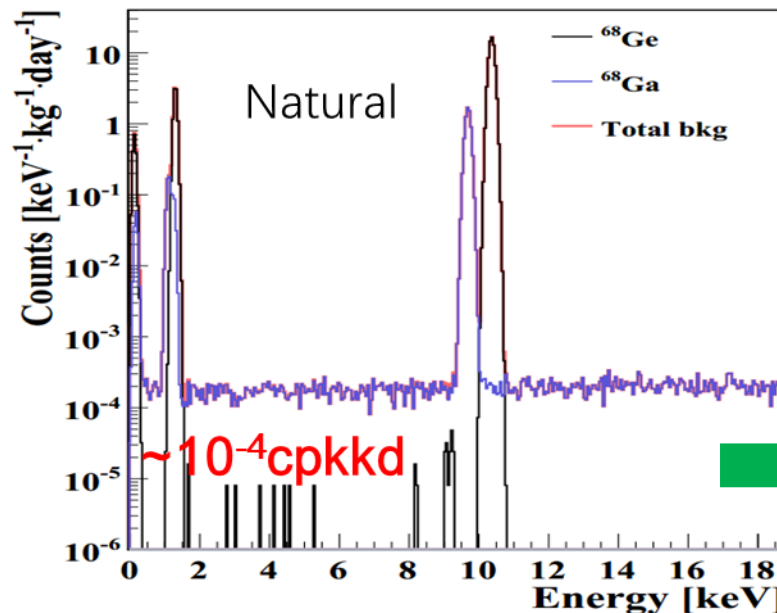
Cosmogenic bkg of Ge crystal



- **Underground** germanium crystal growth and detector fabrication could dramatically decrease the cosmogenic backgrounds from **non-Ge isotopes**, such as ^3H , ^{65}Zn ($^{68}\text{Ge}/^{68}\text{Ga}$ left);
- **^{76}Ge Enriched** Ge material could further help to decrease ^{68}Ge (^{68}Ga).

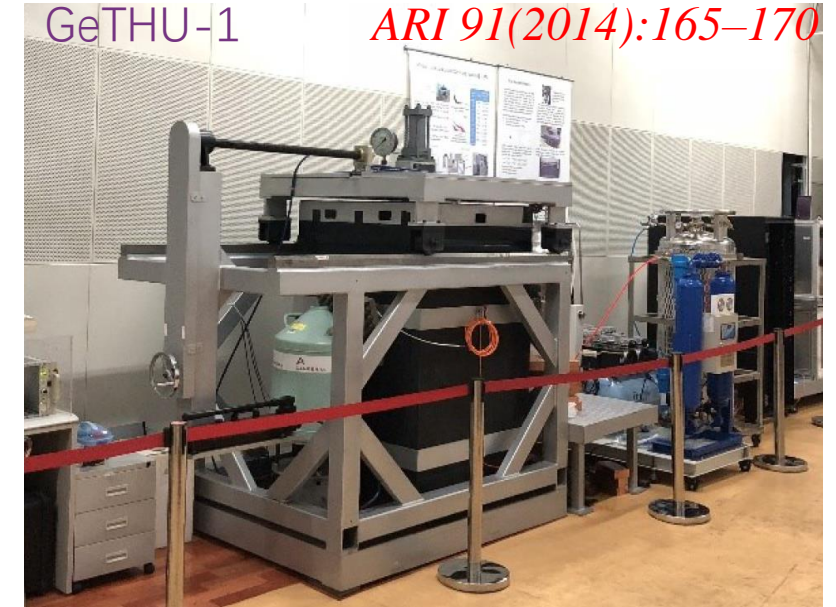
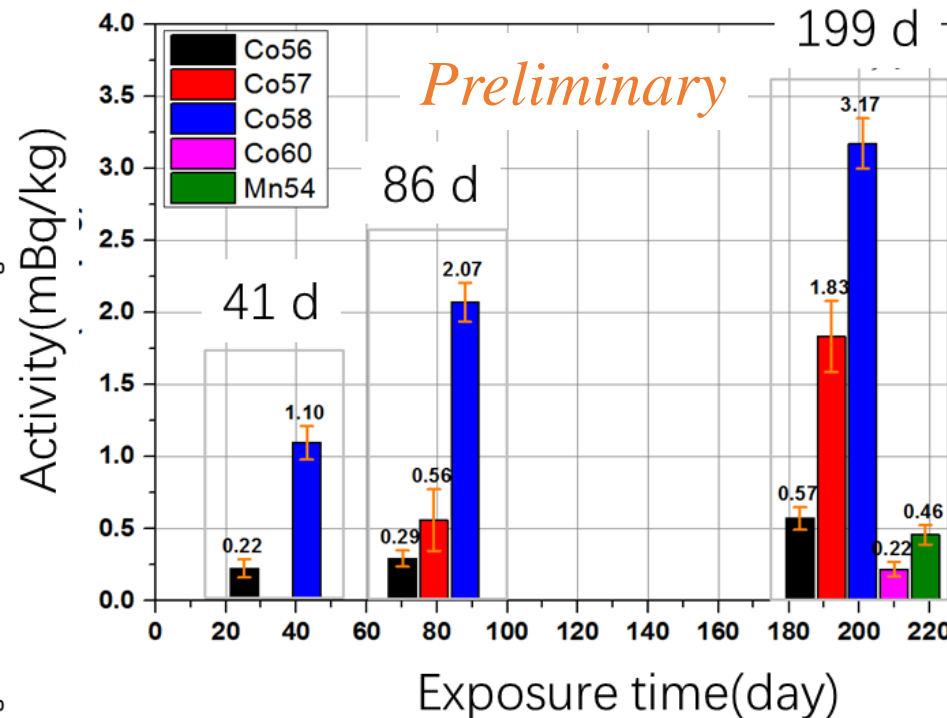
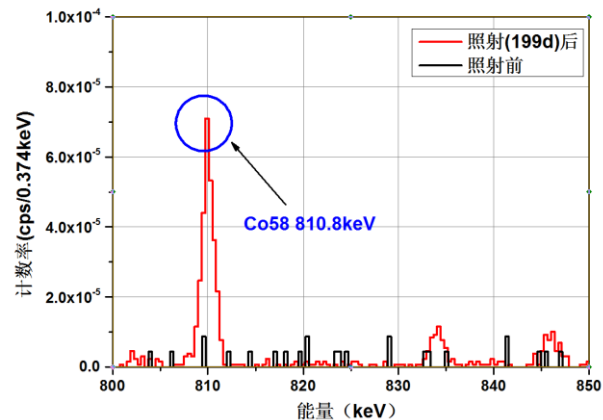
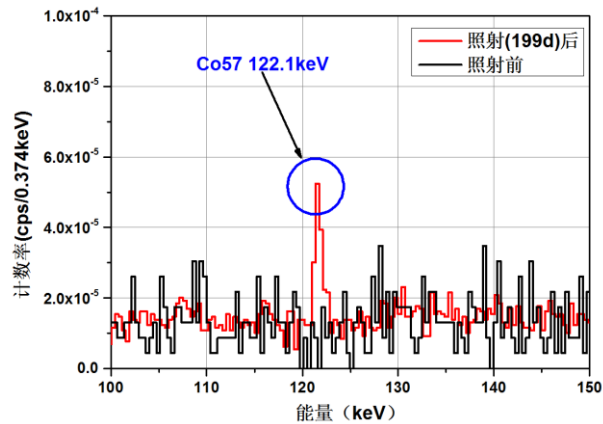
^{76}Ge abundance
7.6% \rightarrow 86.6%

Simulation spectrum: UG + 3 years cooling



Cosmogenic bkg of Copper

- Cosmic shower @ ~2500m altitude outside CJPL;
- Gamma spectrometry by GeTHU-1;
- Data under analysis...and **vs. simulation**

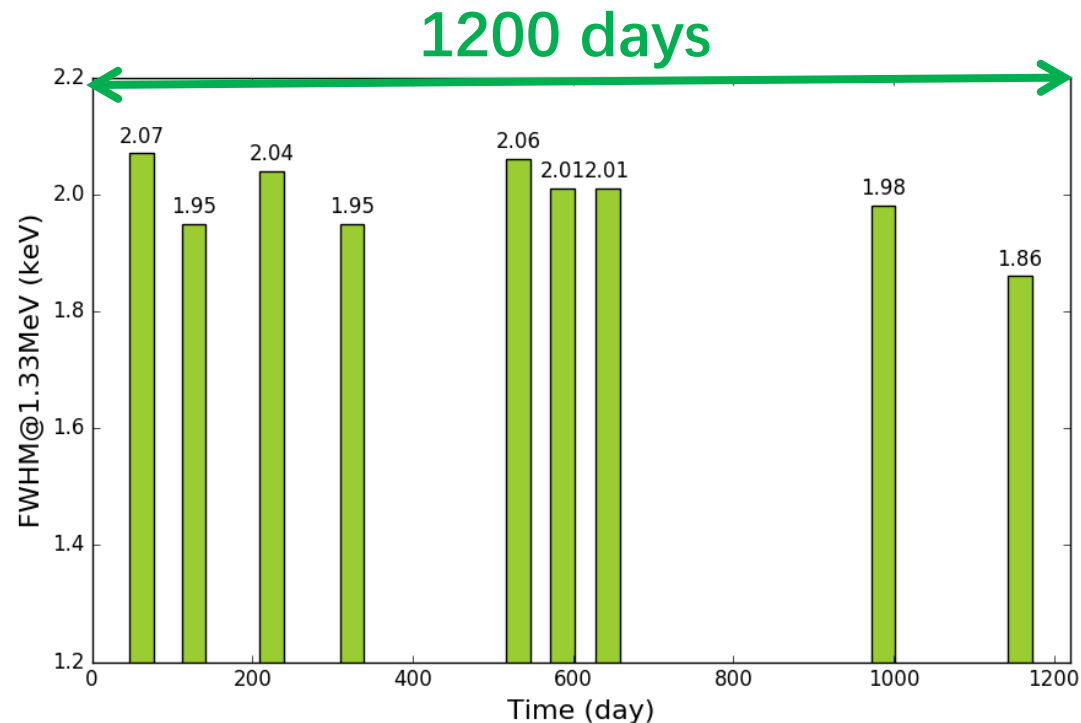


Underground more than 3 years

CDEX10X: Ge detector fabrication



- CDEX10+X home-made Ge detectors;
- Understand & reduce detector intrinsic bkg;
- Various types, ~20 detectors
 - P-type planar/coaxial;
 - P-type point contact/ BEGe;
- Long time stability



Vacuum systems

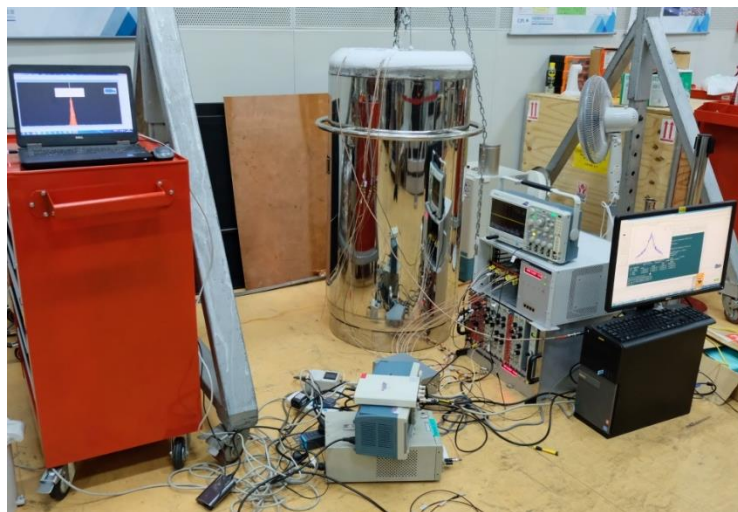
CDEX-10X Detector (T1)



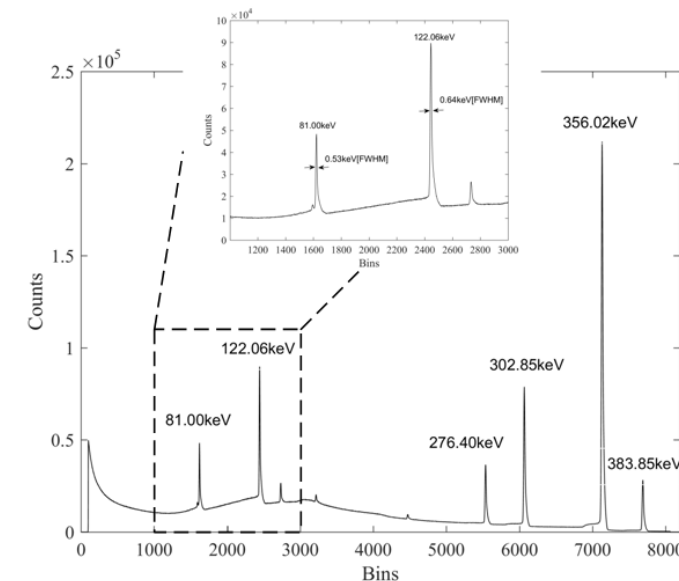
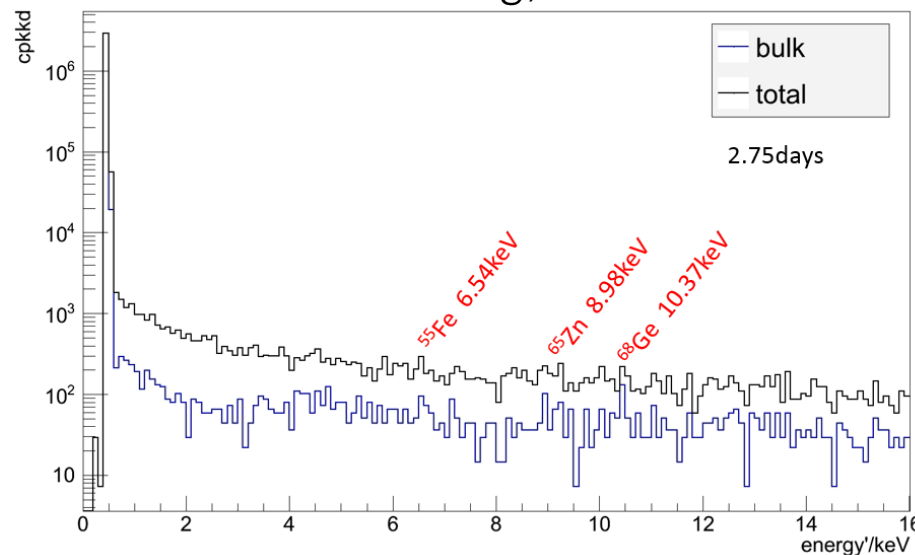
- Commercial Ge crystal + stainless steel canister;
- T1 detector: 500g Ge($\phi 50 \times 50 \text{ mm}$) + CMOS ASIC preAmp;
- Works, and Performance expected;
- Next step to improve bkg, low-noise electronics...



From Y.L. Li, PIRE-GEMADARC Collaboration Summer Meeting, 2018



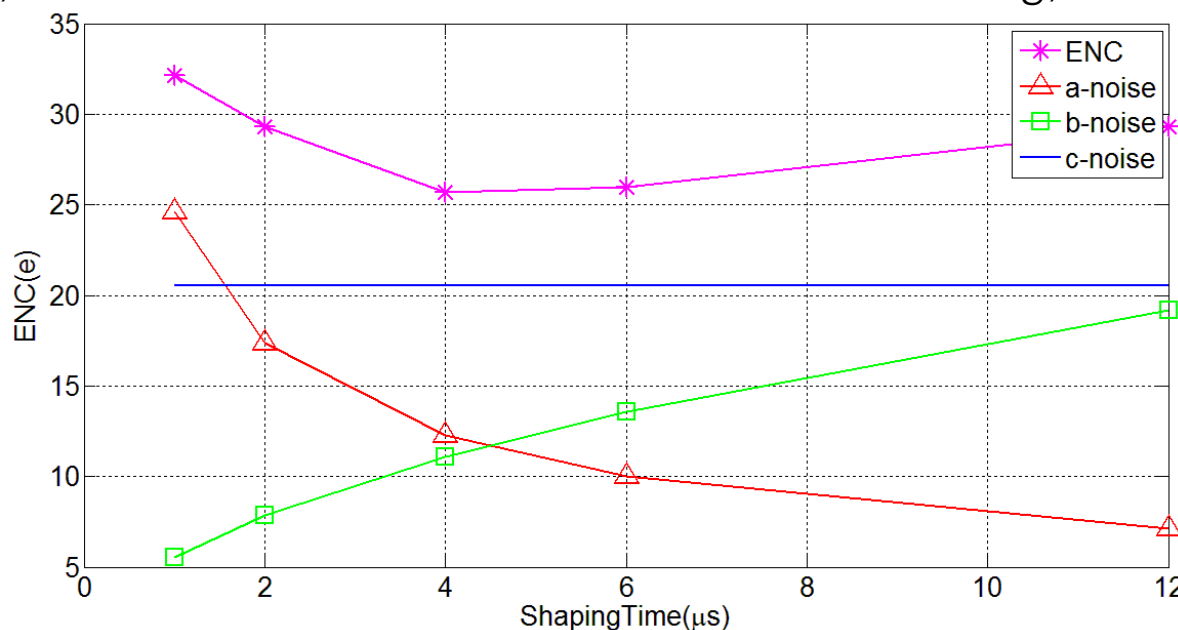
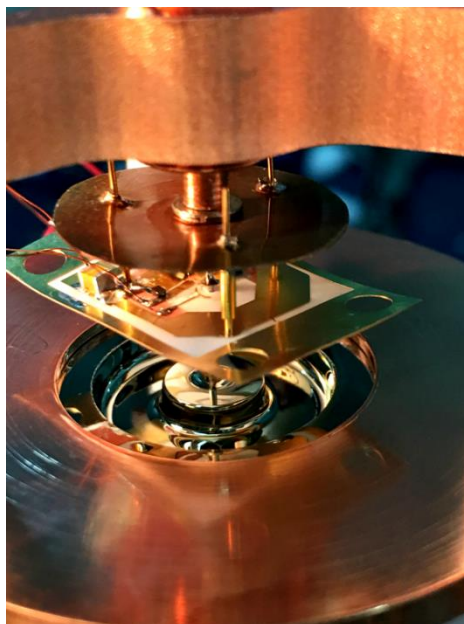
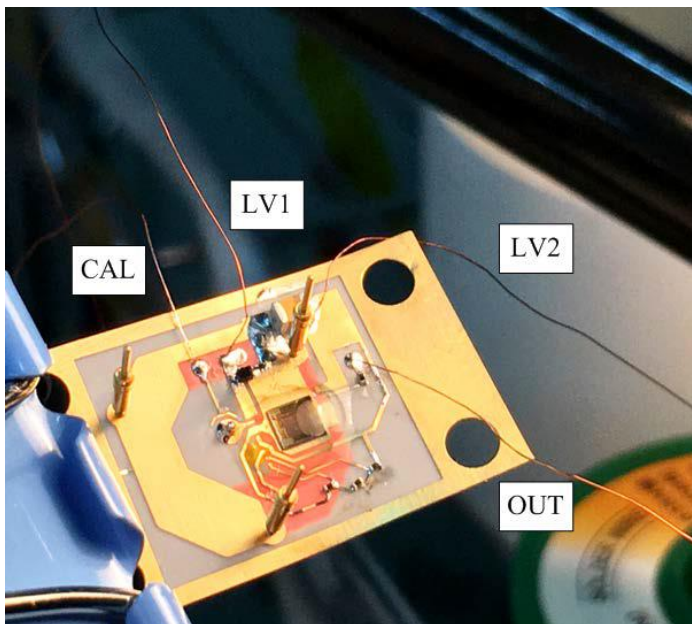
Tested in CJPL-I



FWHM=0.64keV@122keV_Co57

- Light DM search \rightarrow low noise/threshold (low capacity, etc)
- Very close to Ge detectors \rightarrow low bkg (radiopure, low-mass, etc)
- ASIC preamplifier @ 77K
 - PCB material: PTFE (Rogers 4850);
 - ENC $\sim 26e$ ($< 200eV$) w/ $4\mu s$ shaping time, mainly from $1/f$ noise ($\sim 21e$);

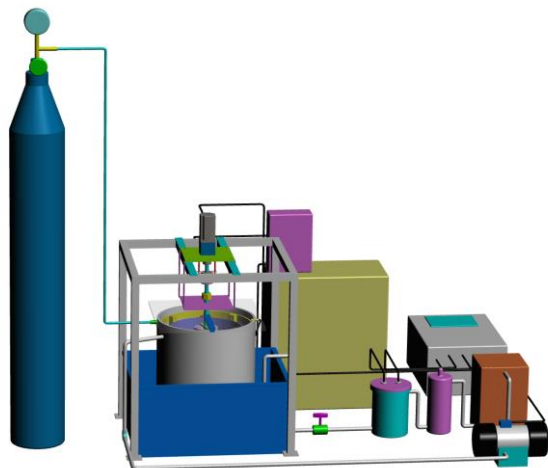
From L. He, PIRE-GEMADARC Collaboration Summer Meeting, 2018



Underground e-forming copper and Assay



- Prototype setup for underground e-forming Cu production
 - Cathode mandrel: 316L stainless steel, $\phi 95 \times 380 \text{ mm}$;
 - Plating bath: PE, $\phi 400 \times 500 \text{ mm}$;
 - Goal: Majorana copper, U/Th content $\sim O(0.1 \mu\text{Bq/kg})$;
- Test run in Tsinghua U. and moved to CJPL;
- U/Th Analysis by ICP-MS
 - preConcentration procedure testing... , blank sensitivity $\sim 10^{-13} \text{ g/g}$



E-forming setup



optimized electrical parameters

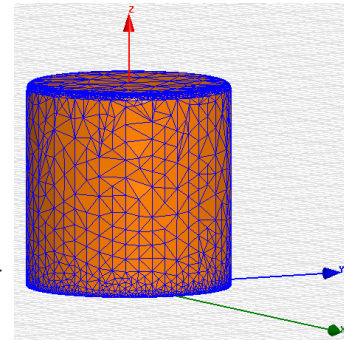
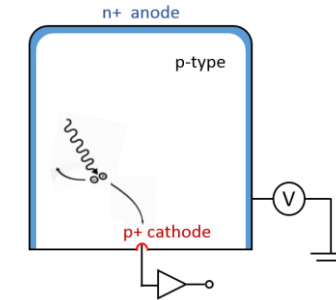


ICP-MS

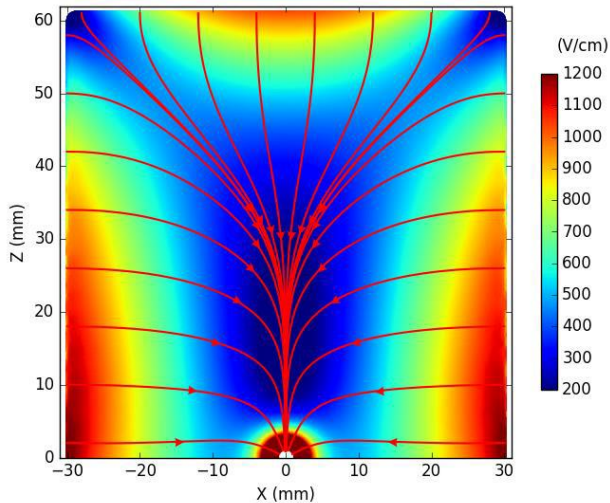
Detector simulation



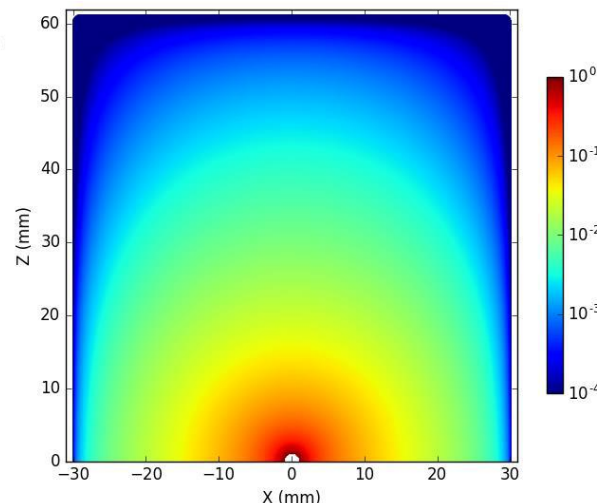
- Optimize detector design on charge collection & bkg reduction;
- Understand physics behind experimental phenomena;
- Pulses generated by
 - Geant4: interaction & energy deposition;
 - ICC package: Induced Charge/Current signal (Shockley-Ramo theorem);
 - Both PPC Ge and BEGe detector;



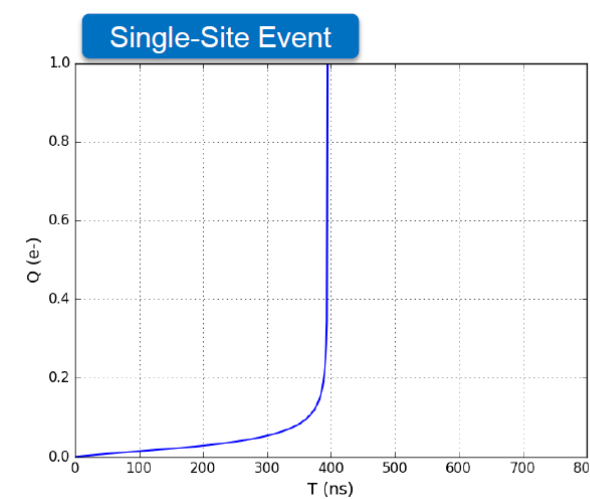
Finite element
analysis
of e-field



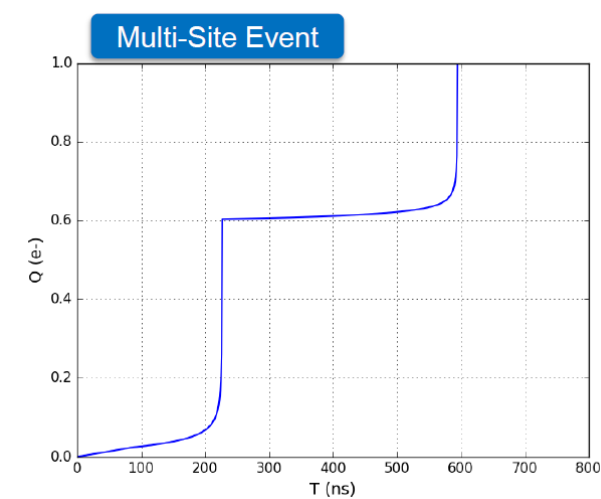
Electric field



Weighting potential



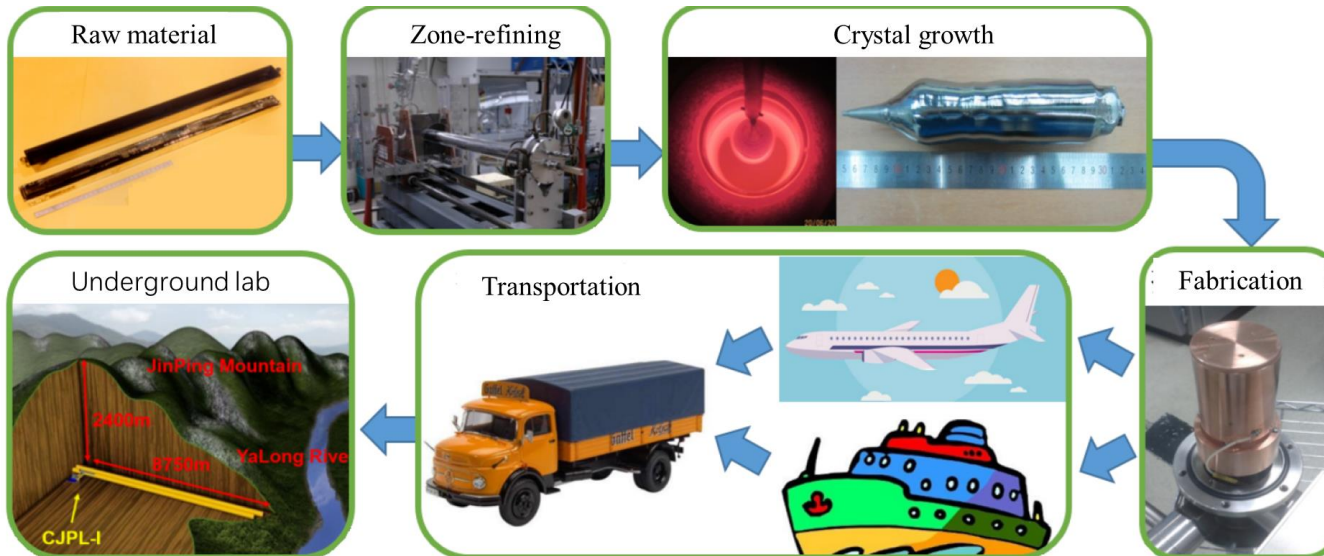
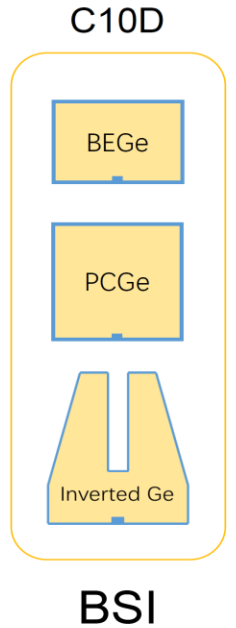
Simulated pulses



Future Plan - Detector



- New detectors cooperated with companies
 - 3kg from BSI, 2kg from ORTEC, planned 5kg from CANBERRA/ORTEC;
 - Particular control of detector fabrication process above ground;

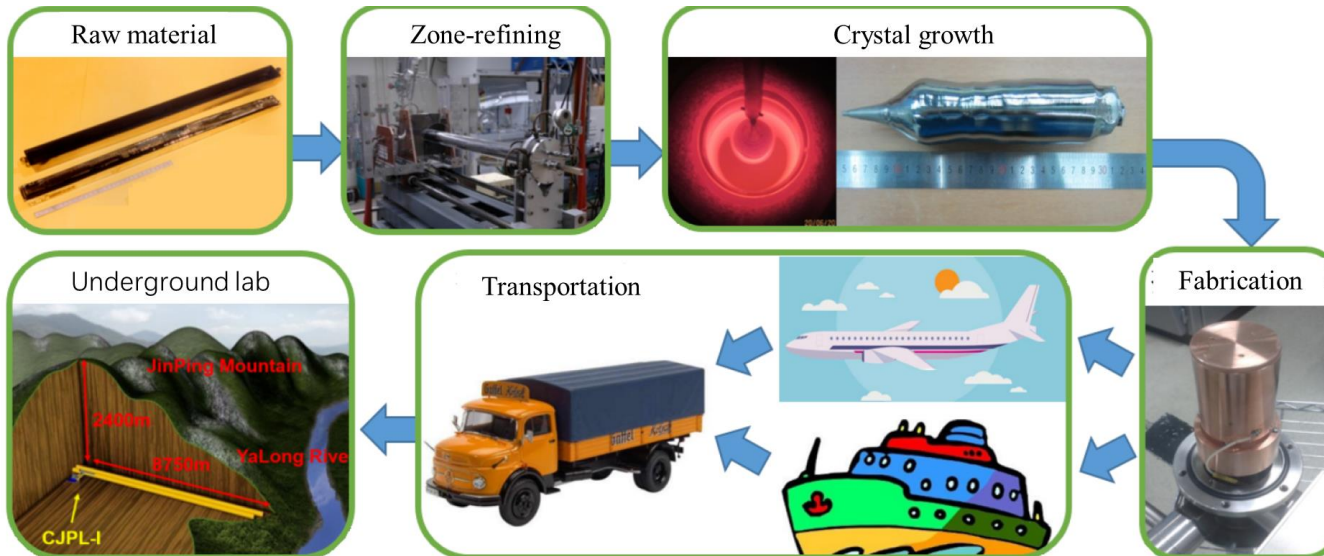
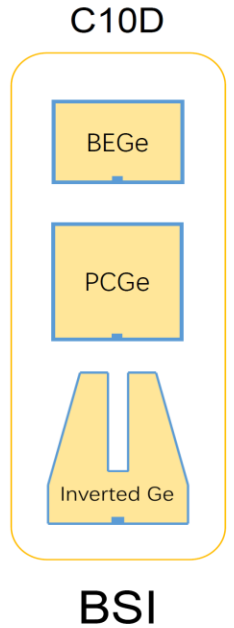


Detector production: 45days +
Ground transportation: 60 days +
Underground cooling: 180days →
Cosmogenic bkg: 0.03cpkkg(sim.).

Future Plan - Detector



- New detectors cooperated with companies
 - 3kg from BSI, 2kg from ORTEC, planned 5kg from CANBERRA/ORTEC;
 - Particular control of detector fabrication process above ground;
- Home-made detectors
 - Improve T1 w/ low bkg material and low noise electronics;
 - Setup the facility for underground fabrication and testing;



Detector production: 45days +
Ground transportation: 60 days +
Underground cooling: 180days →

Cosmogenic bkg: 0.03cpkkg(sim.).

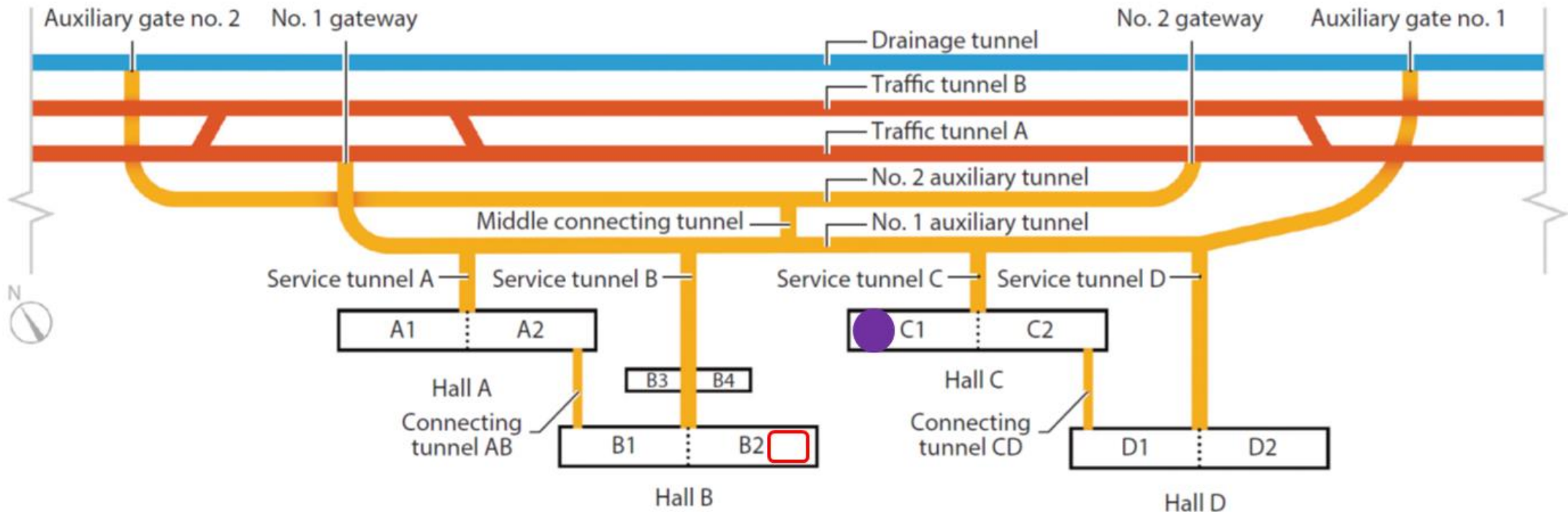
Future Plan - Lab



- CJPL-I to CJPL-II

- Volume: 4000 to 300,000 m³;
- 1 main hall (6.5x6.5x42m) to 4 main halls (14x14x130m each);
- 2 additional pits for next generation CDEX and PandaX;

Y. Yang's talk yesterday



Future Plan - Lab



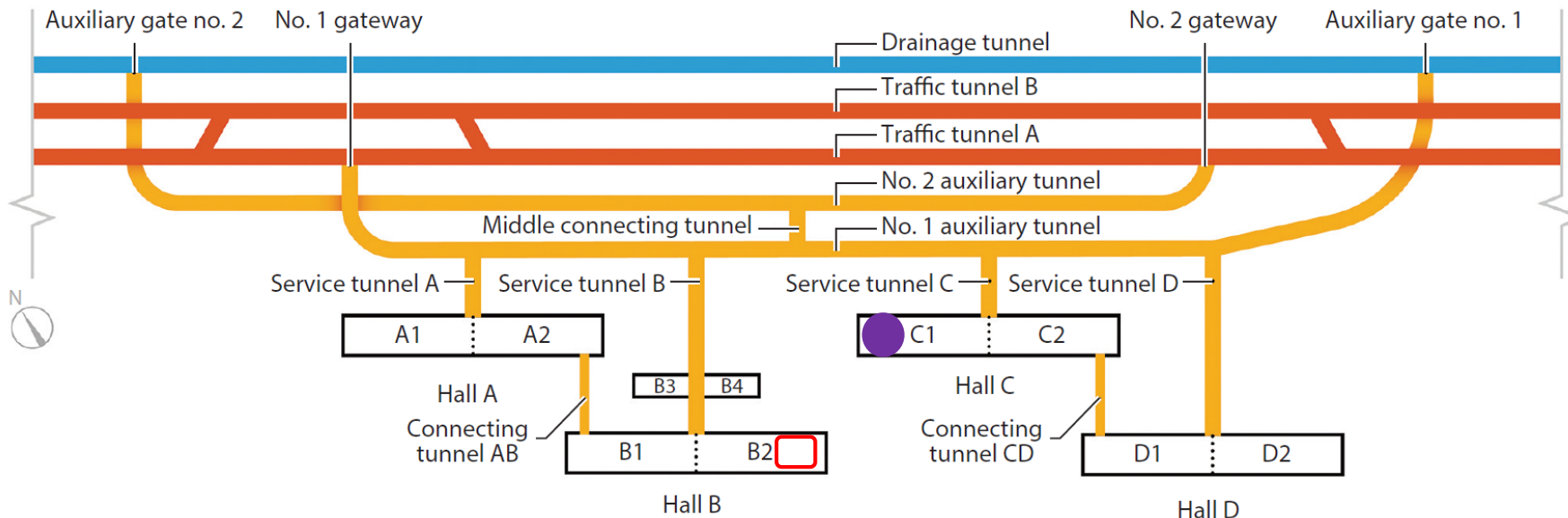
- CJPL-I to CJPL-II

- Volume: 4000 to 300,000 m³;
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- 2 additional pits for next generation CDEX and PandaX;

- CJPL-II status

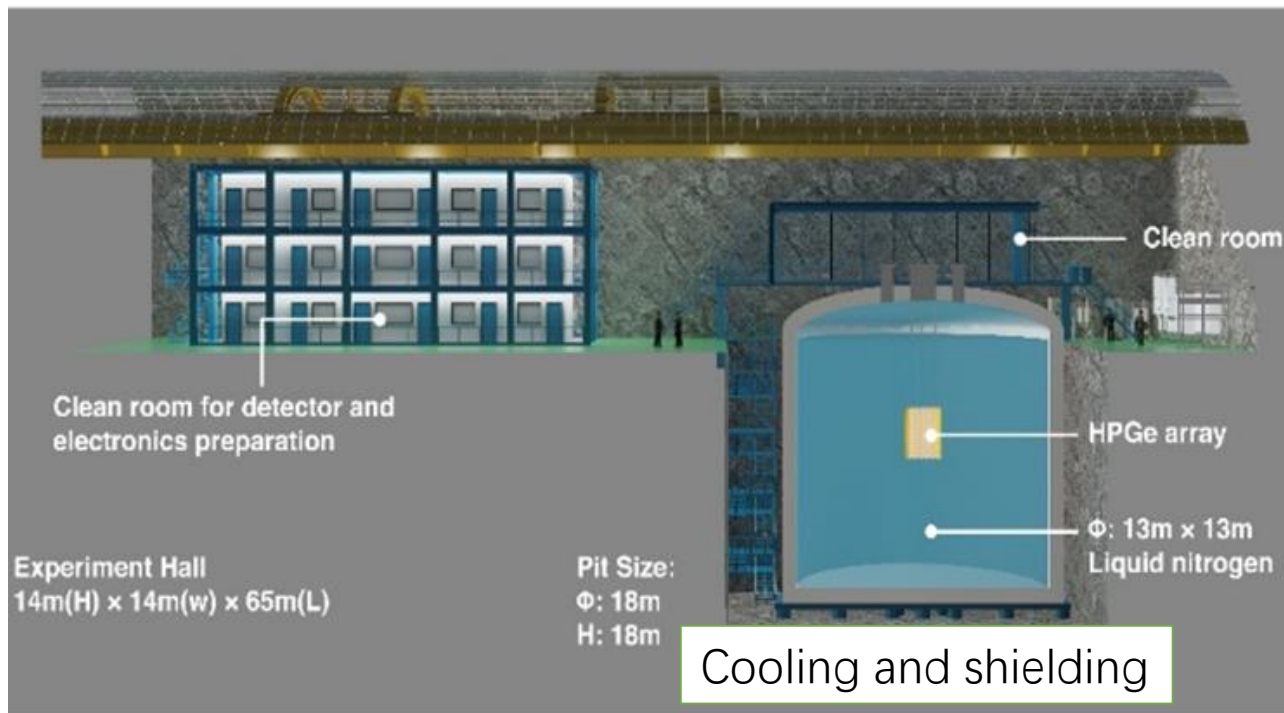
- Civil engineering from Dec. 2014 to May 2016;
- Ventilation system: 3 nine-km-long PE pipes till Jun. 2018;

Y. Yang's talk yesterday



Future Plan - CDEX

- CDEX10X moves to a 1700m³ LN₂ tank (φ13x13m) located in the pit;
- Construction of LN₂ tank will kick off in Aug. 2018;
- 10+X kg detectors direct-immersion and then operation in LN₂ in 2019;
- CDEX-100 stage under technical design, report expected by the end of 2018.

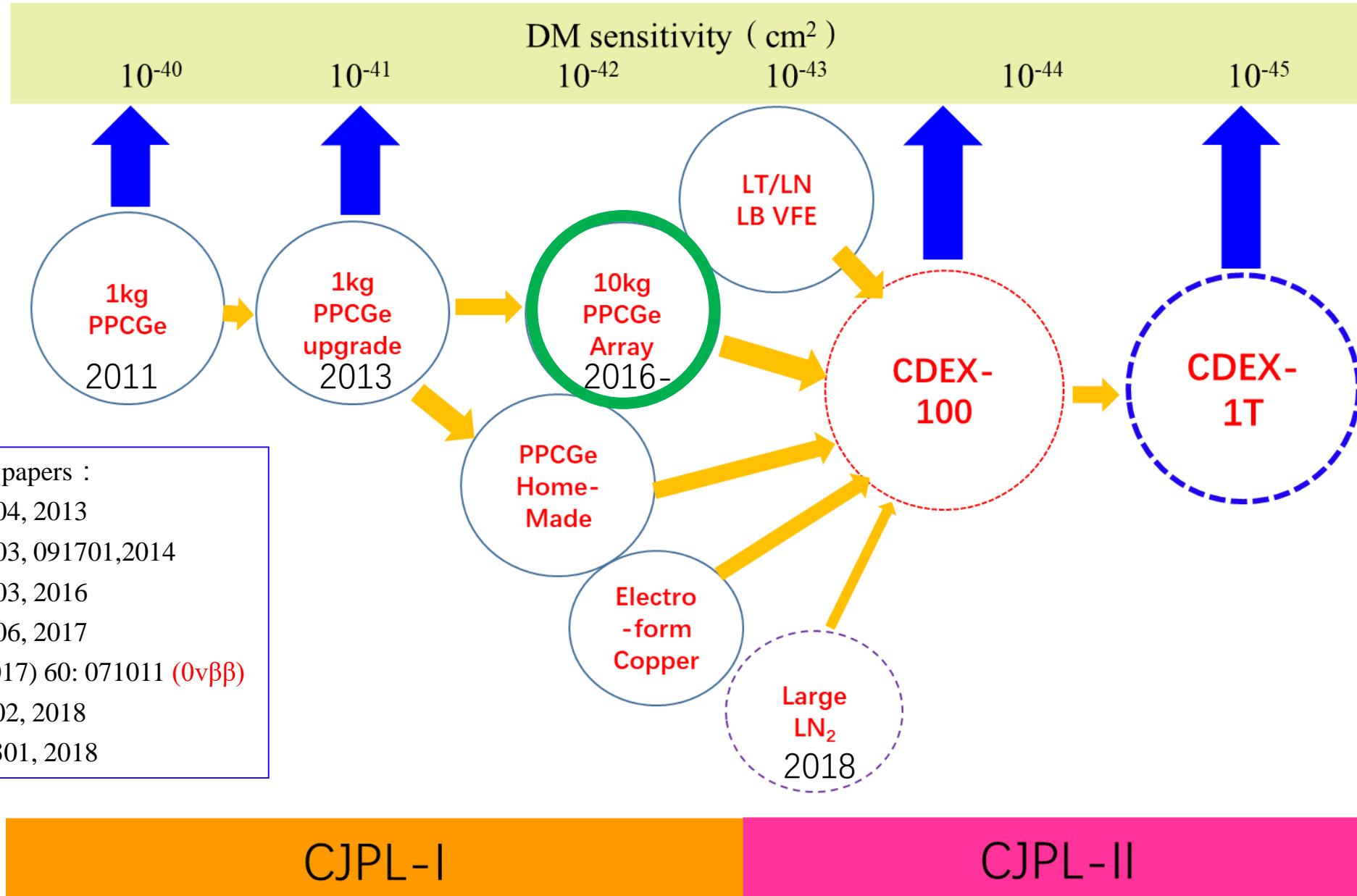


Summary



- CDEX: unique advantages of PPC Ge detectors for light DM search at CJPL;
- New SI limit $8 \times 10^{-42} \text{cm}^2$ at 4-5 GeV by CDEX-10 first results;
- New site in Hall C1 of CJPL-II project;
- Easy scalability and lower bkg expected w/ new large cryo-tank;
- Home-made Ge detector, FE electronics, crystal growth, UG copper e-forming ongoing...
- More detectors on the way w/ particular consideration and control of cosmogenic bkg.
- Other physics: Axion, $0\nu\beta\beta$,...

CDEX Roadmap



CDEX physical papers :

- ✓PRD88, 052004, 2013
- ✓PRD90, 032003, 091701, 2014
- ✓PRD93, 092003, 2016
- ✓PRD95, 052006, 2017
- ✓Sci. China (2017) 60: 071011 ($0\nu\beta\beta$)
- ✓CPC42, 023002, 2018
- ✓PRL120, 241301, 2018

Thanks for your attention!
Welcome to CJPL!