



Status of EDELWEISS

EDELWEISS-III

New result: ALP limits

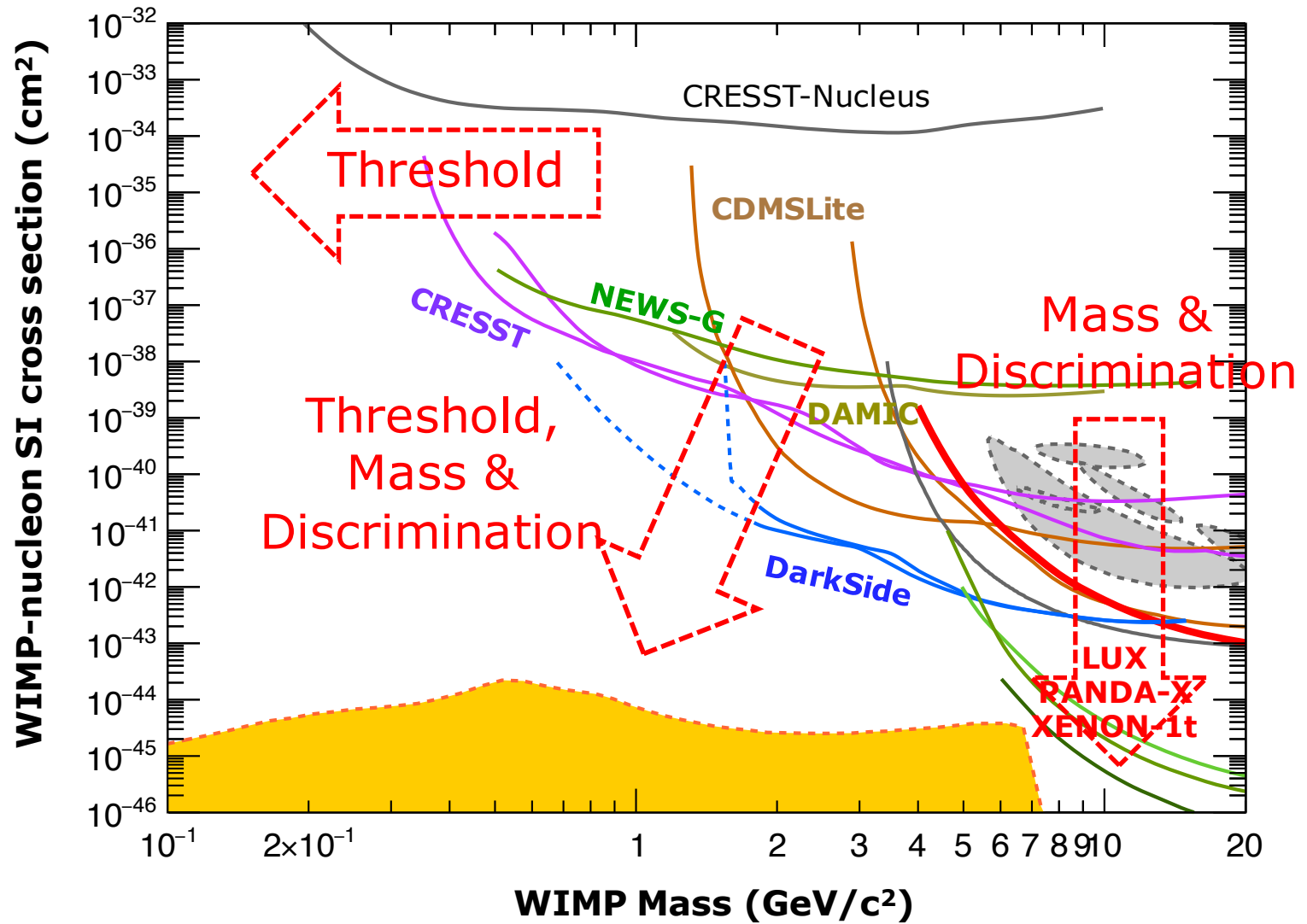
Prospects for discrimination at low energy

EDELWEISS-LT program

New result: Sub-GeV WIMP limit @surface

Jules Gascon
(IPNLyon, Université Lyon 1 + CNRS/IN2P3)

Low-mass WIMP searches



EDELWEISS Low-mass WIMP searches

EDELWEISS-III

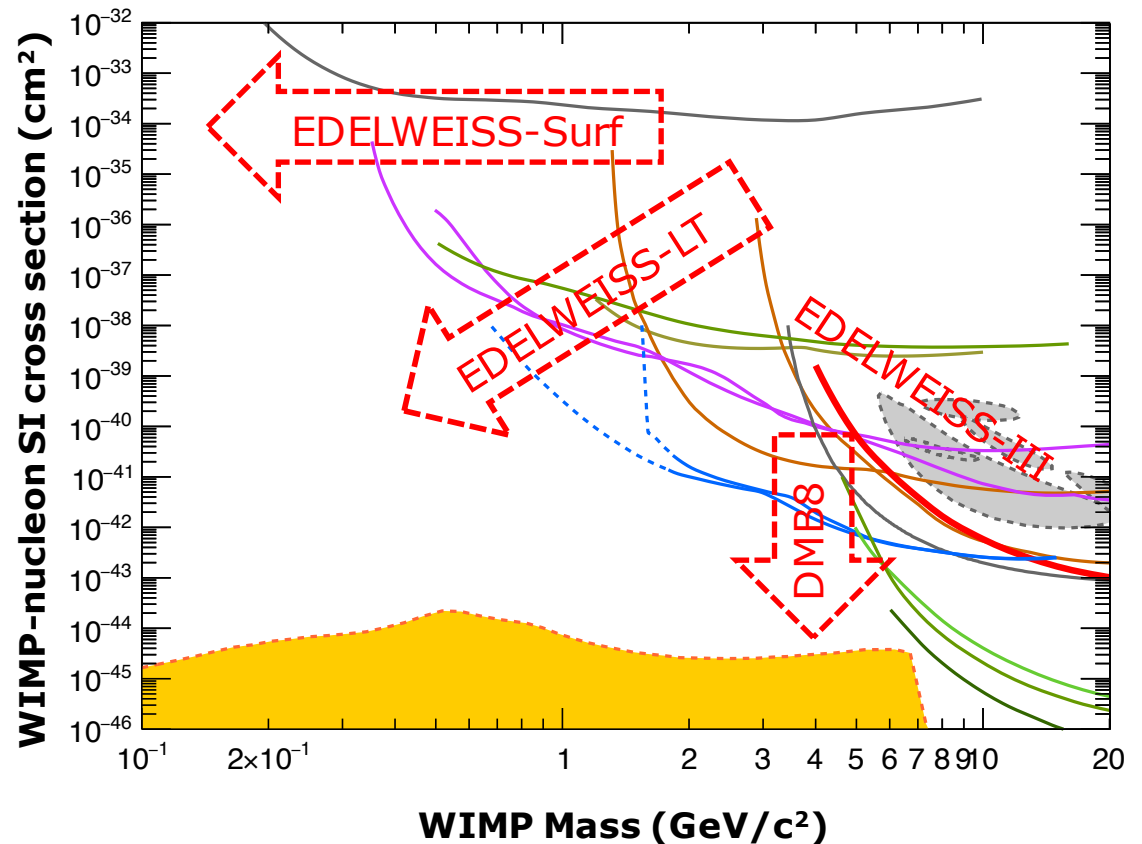
- Exploitation of results with 20 kg array

Ionization R&D

- Improving discrimination to explore the ^8B region with resolution (DMB8)
- Exploring non-WIMP DM with smaller array (MELODI)

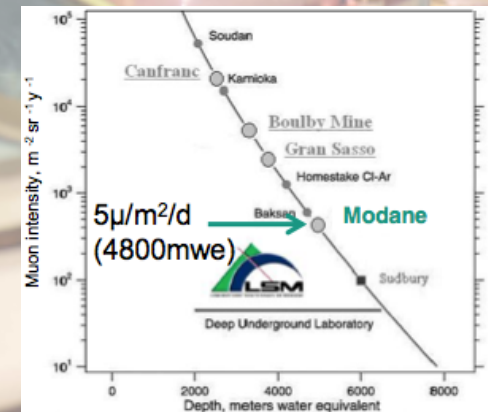
Heat channel R&D

- Improving the heat channel resolution to reach lower WIMP masses
- Above-ground R&D (Surf) and deployment at LSM (LT)



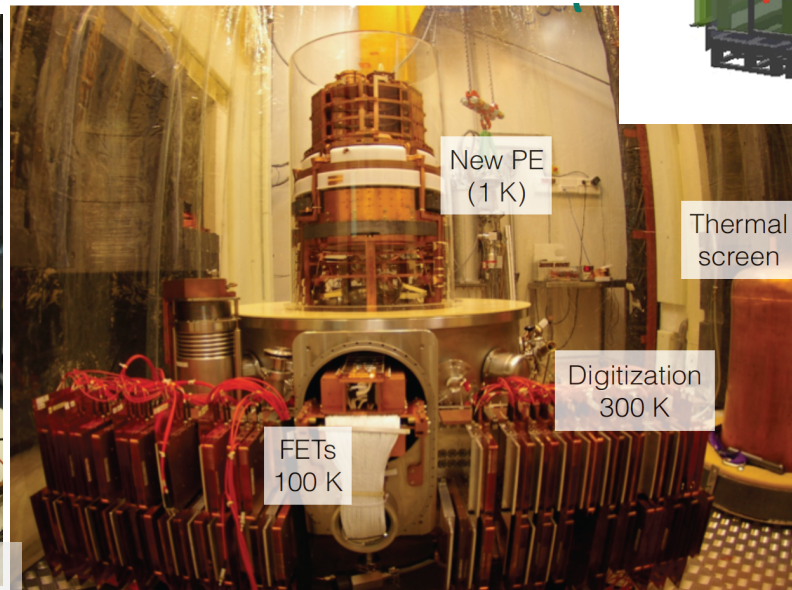
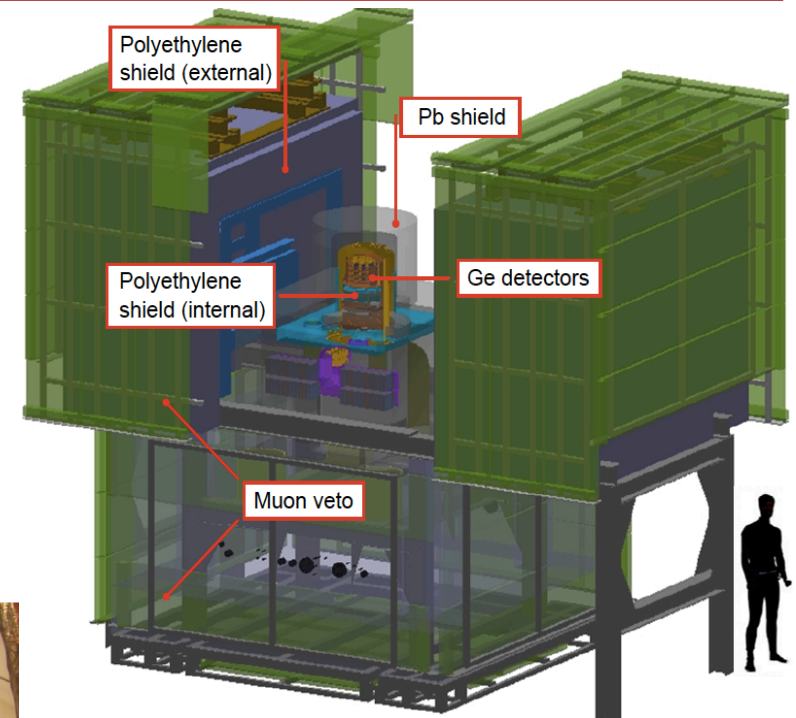
The EDELWEISS-III Experiment

- **Direct detection of WIMPs, germanium target**
- **20 kg Ge total, 870g units**
- **Ionization + Heat**
- **Simple & robust design**
 - Important for scalability to large arrays
 - Initially designed for $>20 \text{ GeV}/c^2$ WIMPs and $\sim 3000 \text{ kgd}$
 - Extended down to $5 \text{ GeV}/c^2$ given achieved resolutions
- **Laboratoire Souterrain de Modane**
 - Deepest in Europe : $5 \mu\text{m}^2/\text{day}$



EDELWEISS Setup

- Clean room + deradonized air
 - Rn monitoring down to few mBq/m³*
- Active muon veto (>98% coverage)
- External (50 cm) + internal polyethylene shielding
 - Thermal neutron monitoring with ³He detector*
- Lead shielding (20 cm, incl. 2 cm Roman lead)
- Selection of radiopure material
- Cryostat can host up to 40 kg detector, at 18 mK



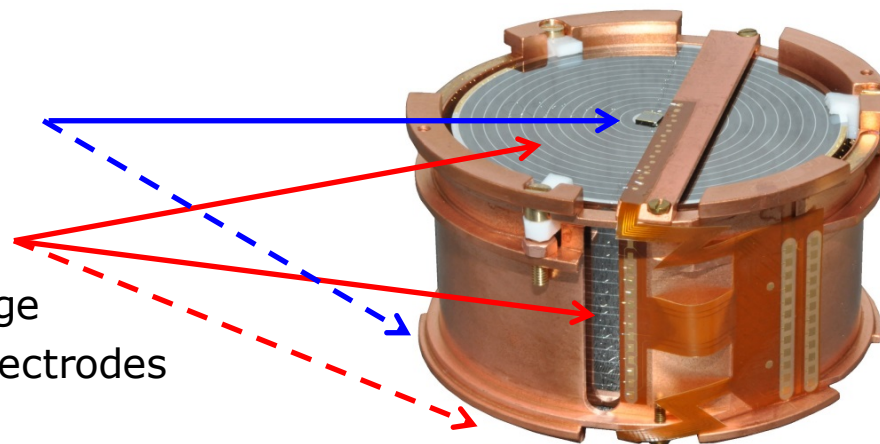
*Performance of the
EDELWEISS-III experiment
for direct dark matter
searches*

[JINST 12 (2017) P08010]

Fully InterDigitized electrode design

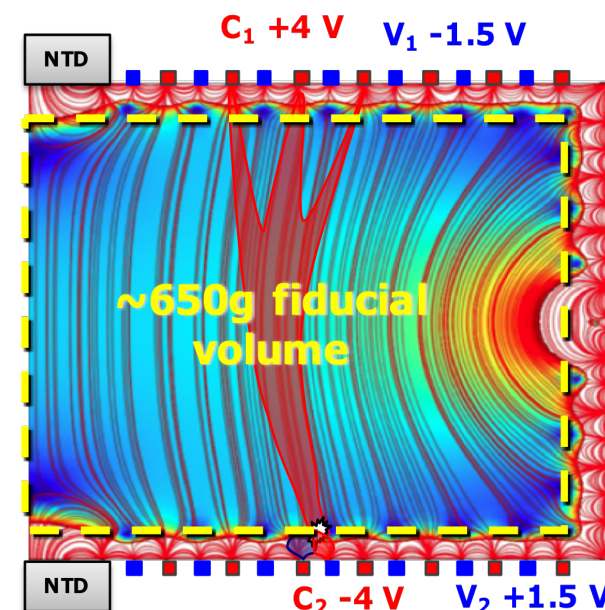
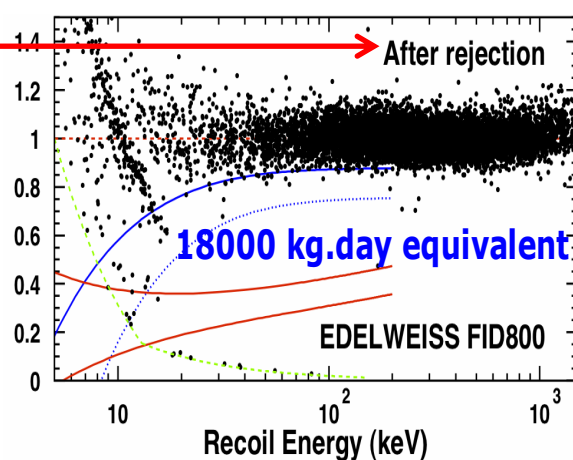
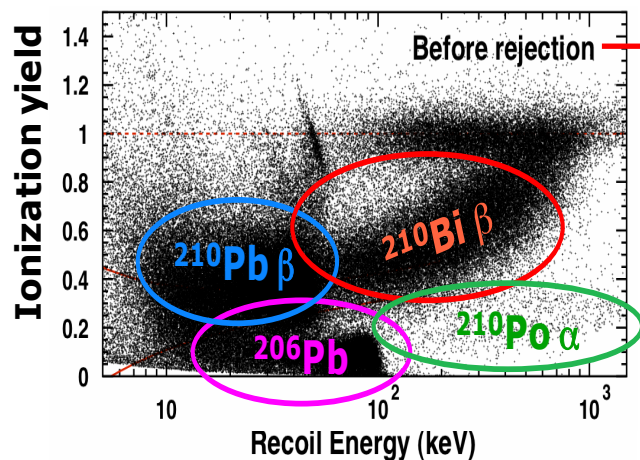
- ~ 870 g detectors ($\phi=70$ h=40 mm)
- 2 GeNTDs heat sensor per detector
- Electrodes: concentric Al rings (2 mm spacing) covering all faces
- XeF_2 surface treatment \rightarrow low leakage current (<1 fA) between adjacent electrodes

J Low Temp Phys (2014) 176: 182-187



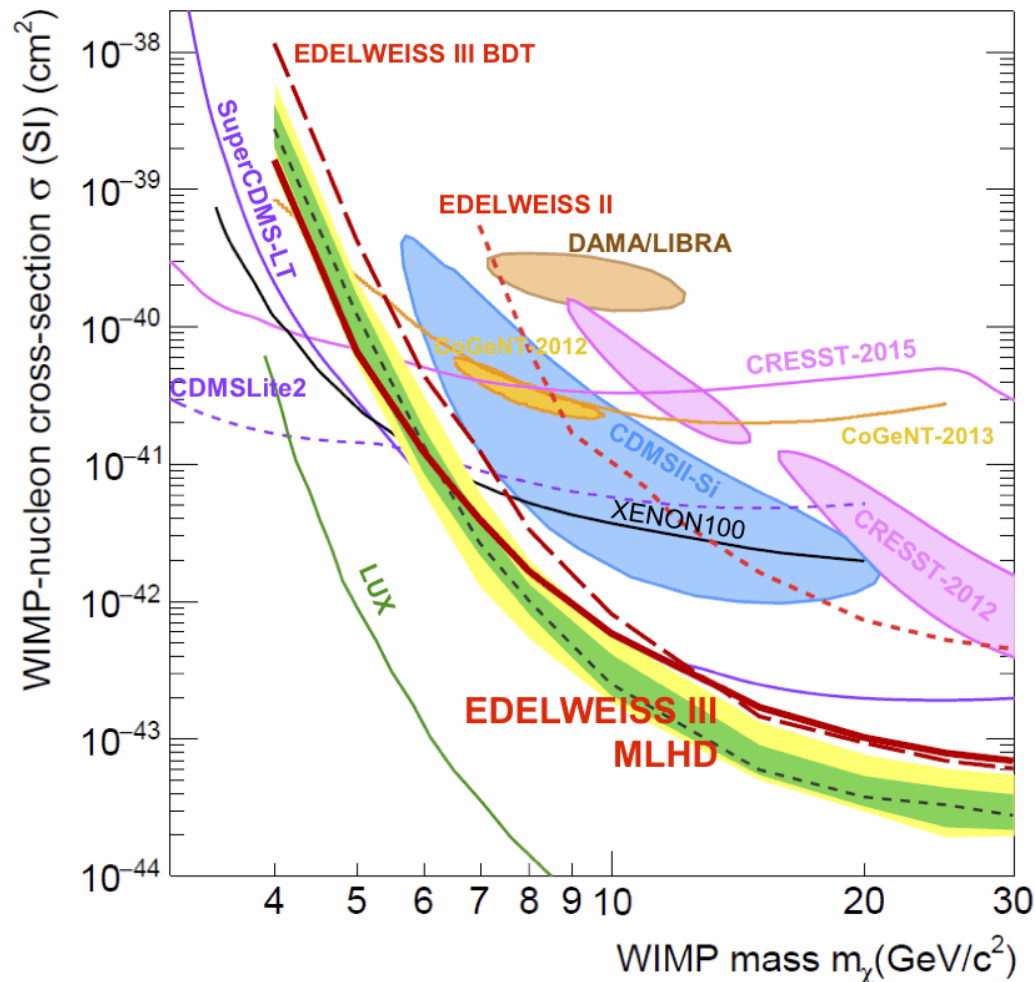
Surface event rejection. *PLB 681 (2009) 305-309*

- Bulk: collection by C_1+C_2 ; V_1+V_2 act as veto
- Surface: charges collected by C_1+V_1 or C_2+V_2



Low-Mass reach of EDELWEISS-III

- Analysis with Boosted Decision Tree [*JCAP05 (2016) 019*]
- Analysis with Profile Likelihood [*EPJC 76 (2016) 548*]



- Improvement by x20 to x150 between 7 and 10 GeV wrt EDELWEISS-II
- Limited by heat-only background: *identification and rejection using the $\sigma=230$ eV resolution on ionization*
- Ionization resolution is key for rejection
- Heat resolution is key for low thresholds

Axion-Like Particle searches

- Starting point: study of electron recoil spectrum of tritium paper [\[Astropart. 91 \(2017\) 51\]](#)

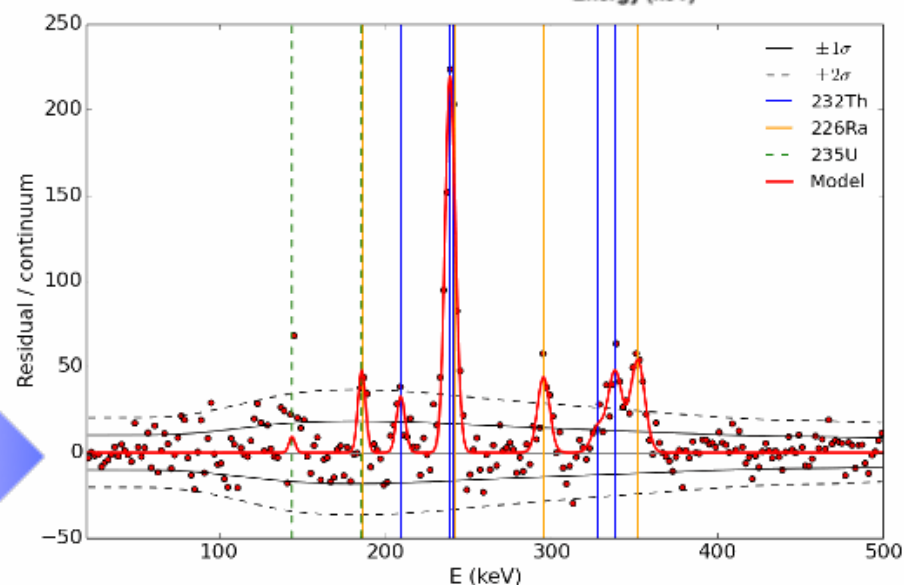
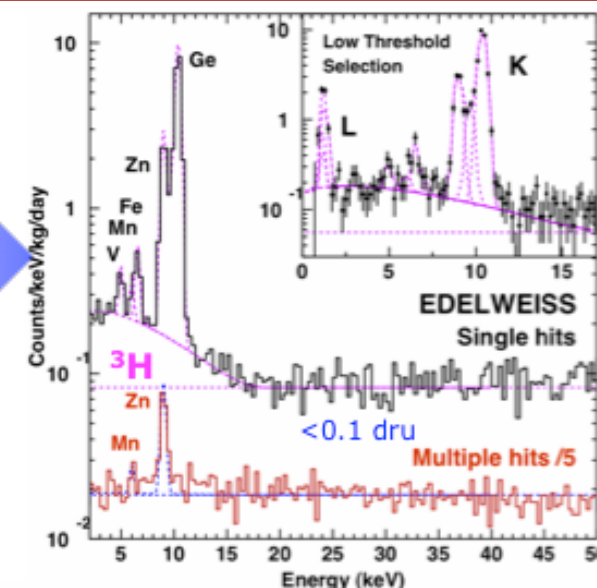
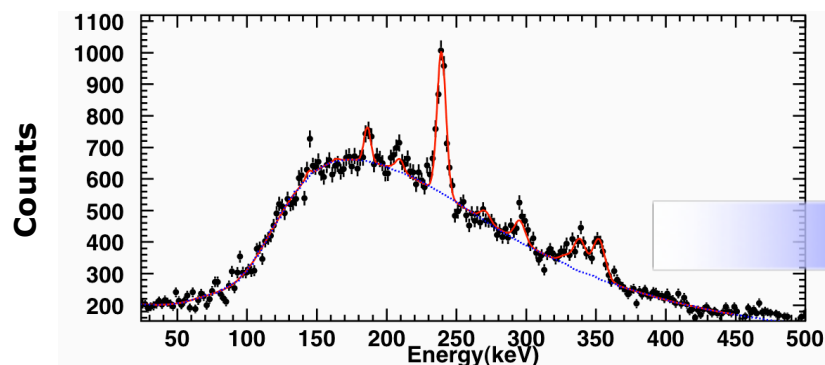
- 1149 kgd with 2 keV_{ee} threshold
- 287 kgd with 0.8 keV_{ee} thresholds

- Analysis extended to higher energy for line search up to 500 keV_{ee}

- Intensities of observed peaks consistent with known Th/U lines

- Resolution:

- Baseline 0.19 keV_{ee}
- Proportional term = 1.2%



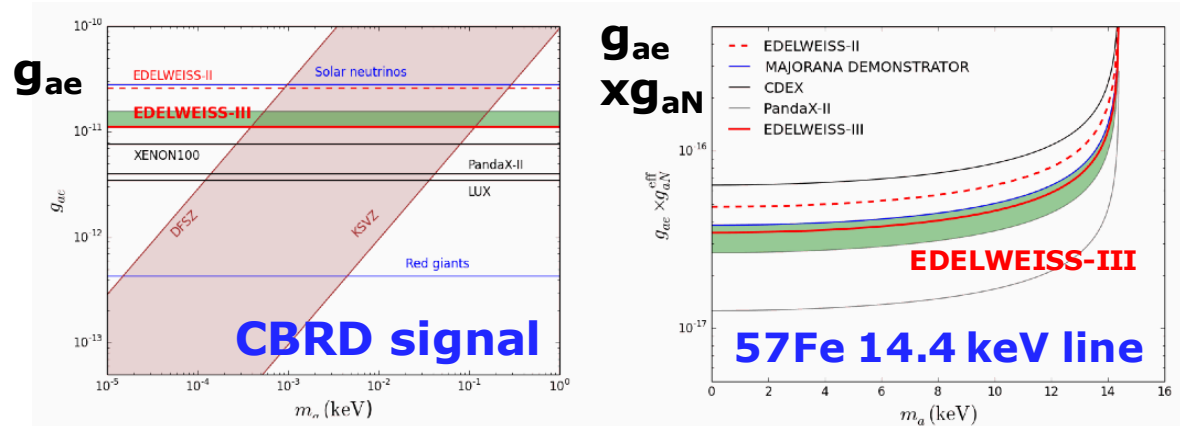
ALP & dark photons results

- Emission of axion/ALPs from the sun

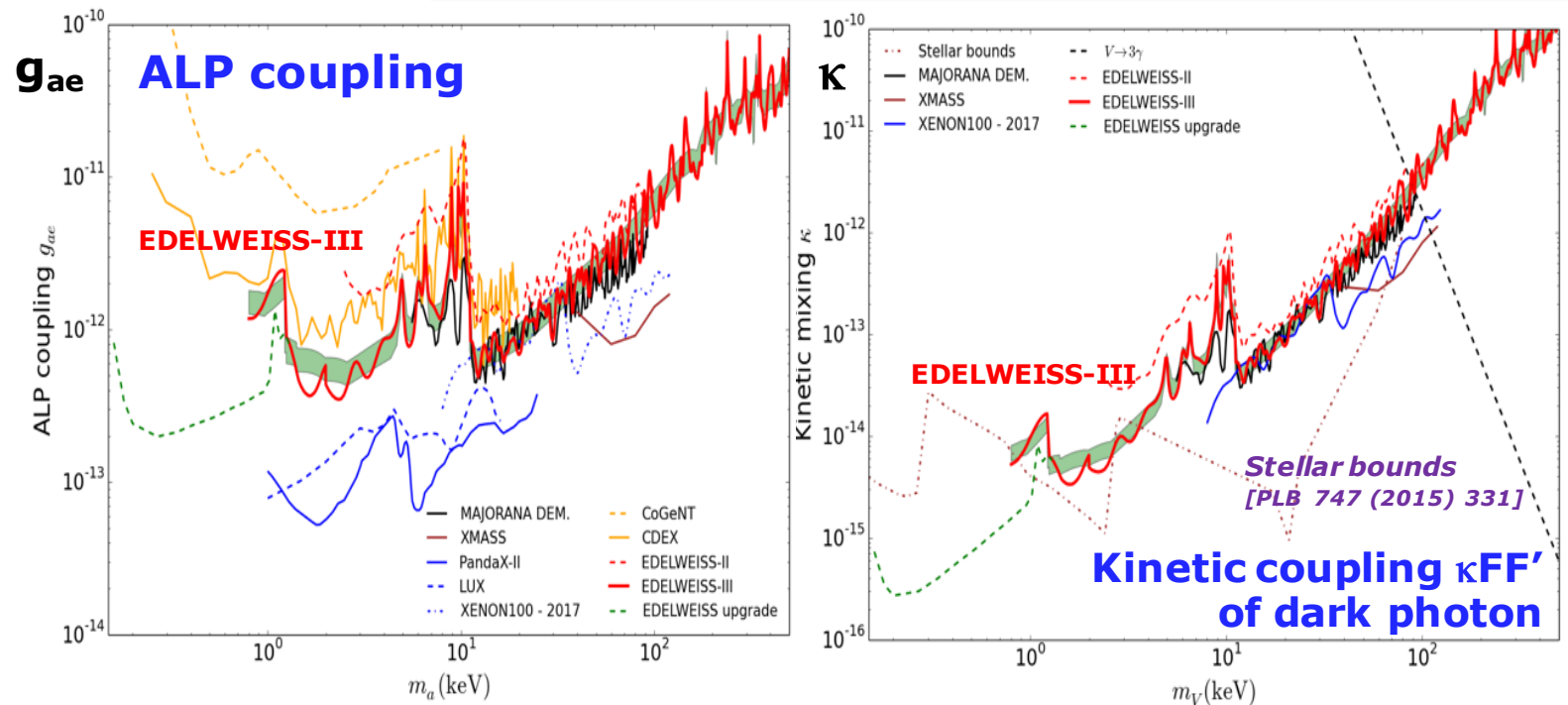
keV-scale Bosonic DM:

- Best Ge-based limits < 6 keV (thanks to surface rejection)
- Start to explore < 1 keV

Emission of axions from the sun



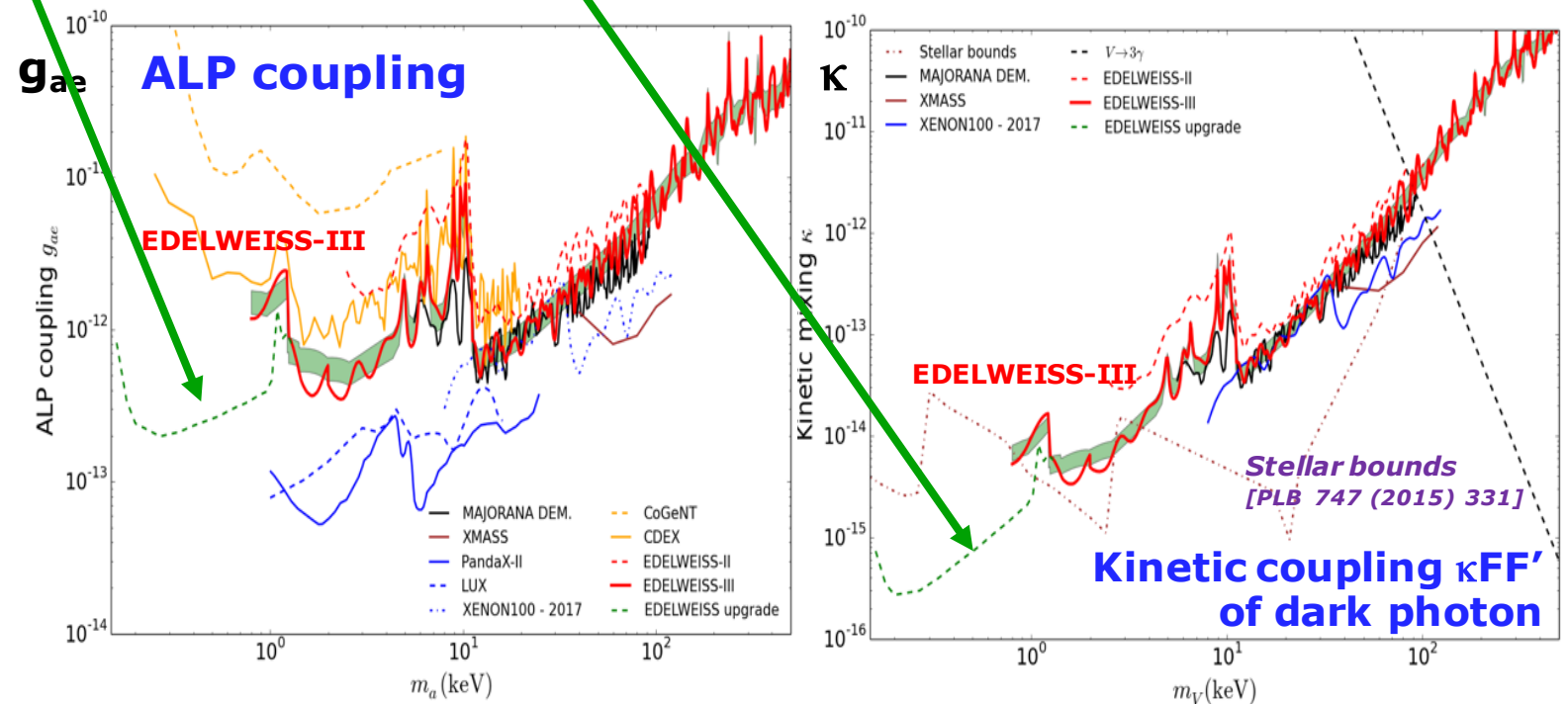
Bosonic DM



Prospective with improved ionization

- Improvements foreseen in the 100 eV – 1 keV region with improved ionization

Bosonic
DM

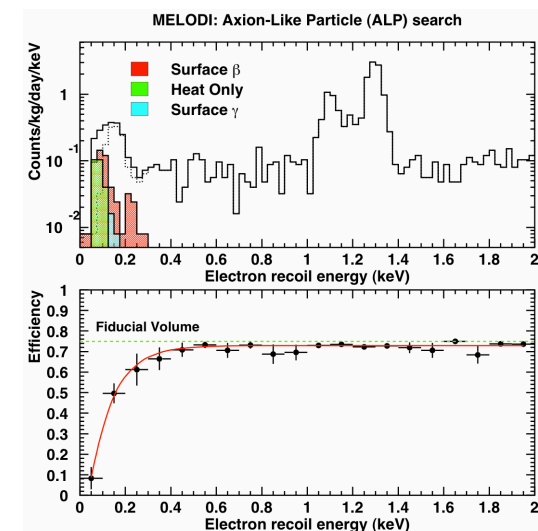
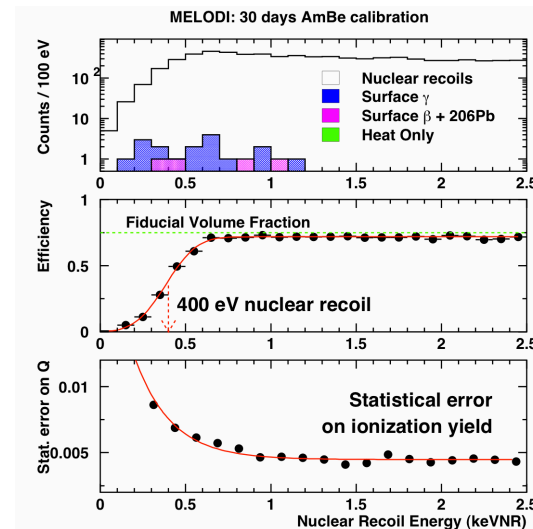
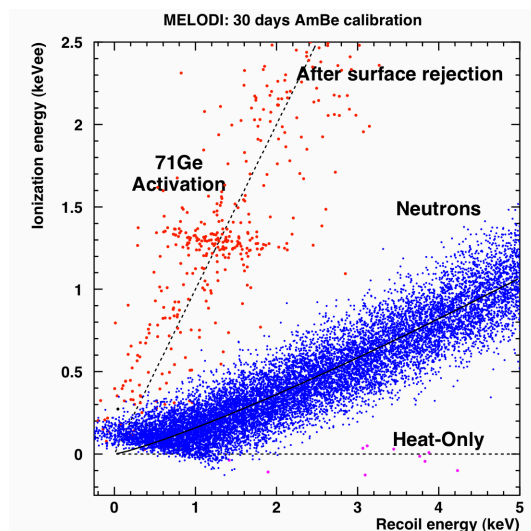
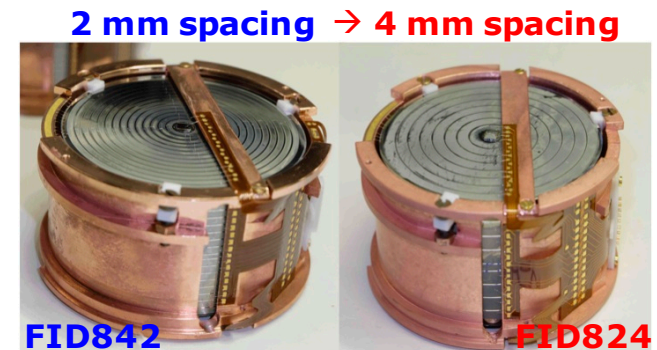


How to do it

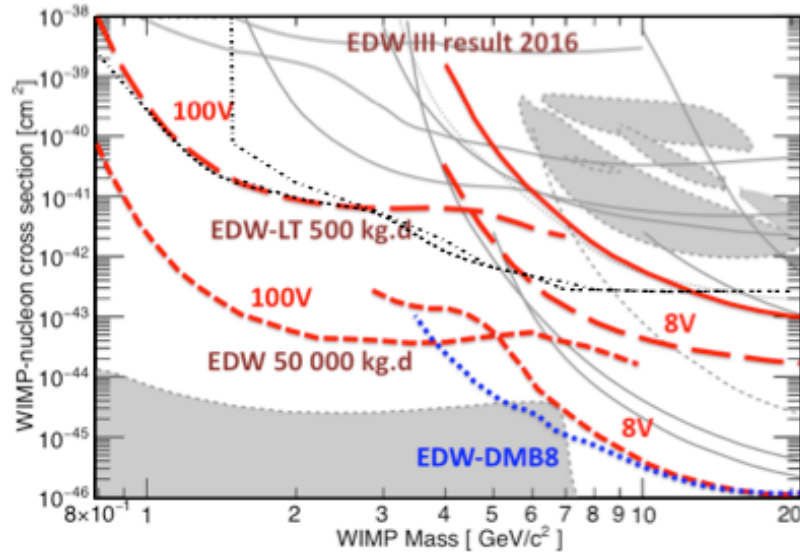
- Cold front-end: replace JFET @100K with HEMT (High Electron Mobility Transistor) @4K
- Can be operated at 4K: shorter cabling -> reduced capacitance -> better signal/noise
- Successful HEMT amplifier with sub-100 eV resolution operated on a CDMS-II detector
[A. Phipps et al., arXiv:1611.09712]
- EDELWEISS electrode design with lower capacitance:
2 → 4 mm spacing already achieved. Goal: reach 50 eV_{ee}.

First applications with prototype detector: MELODI

- FID detector with <40 pF + <50 eV_{ee} ion. resolution
- **Precision nuclear recoil ionization yield measurement**
- **Axion-Like Particle Dark Matter searches with 100 eV - 1 keV masses**



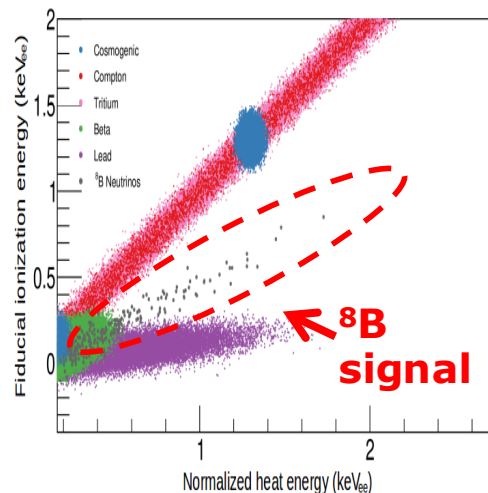
Ionization improvements: EDELWEISS-DMB8



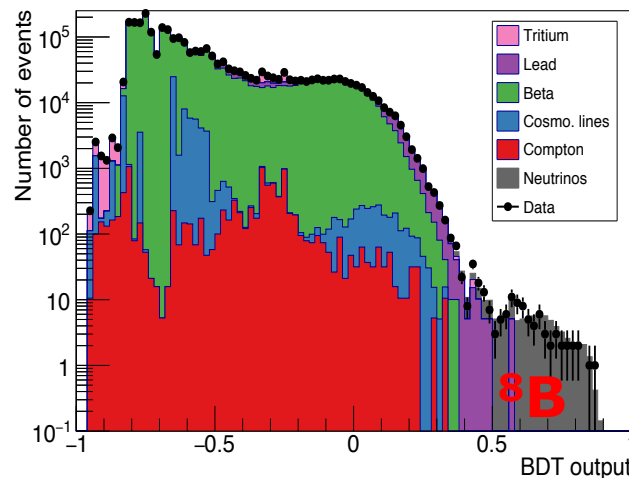
Coherent nuclear scattering from solar ^8B neutrinos mimic a ~ 6 GeV WIMP with $\sigma \sim 4.4 \times 10^{-45} \text{ cm}^2$

Probe with discrimination and resolution ($\sim 10\%$) a bkg that will soon become relevant for WIMP searches near $6 \text{ GeV}/c^2$.

[PRD 97 (2018) 022003]



1 000 kg.y



EDELWEISS-DMB8:

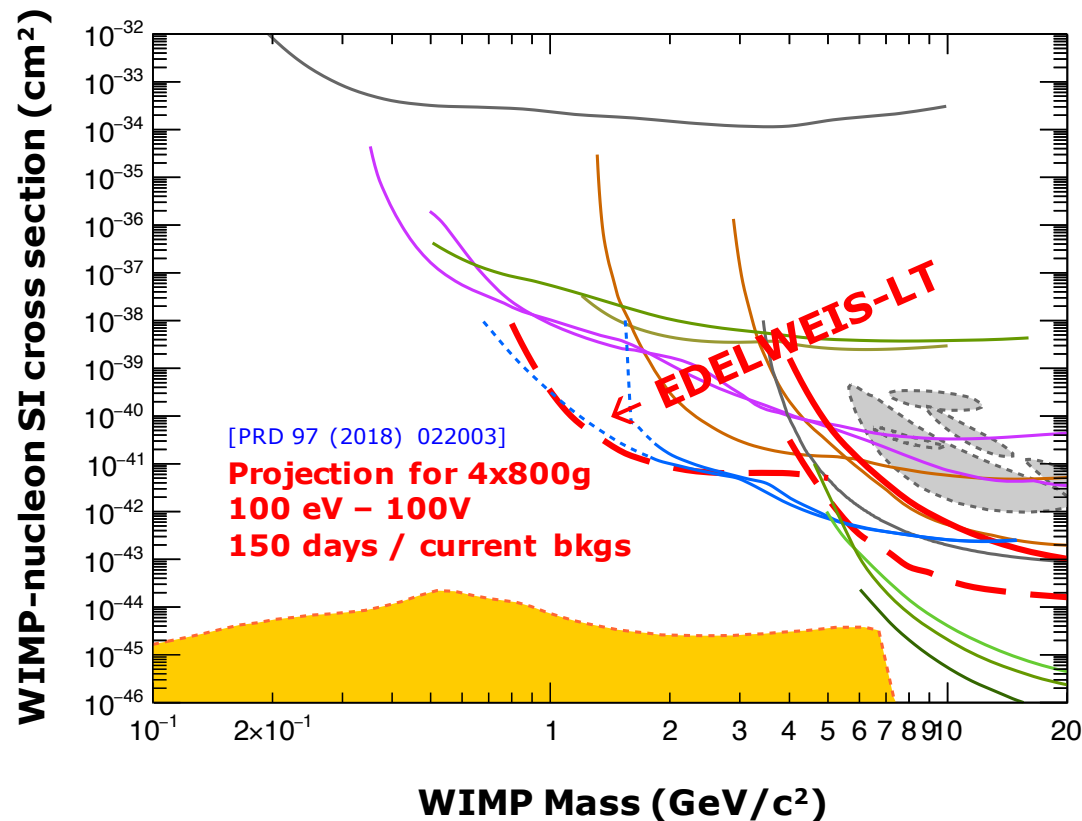
Operation of a 200 kg array @8V (with nuclear recoil discrimination + $<10\%$ energy resolution) in the improved background environment of SuperCDMS @ SNOLAB

Probing the region of the coherent scattering of ^8B solar ν 's with resolution and discrimination

Prospects for GeV-range masses

- Complete study based on present measured backgrounds and resolutions vs possible improvements [PRD 97 (2018) 022003]

Lower thresholds

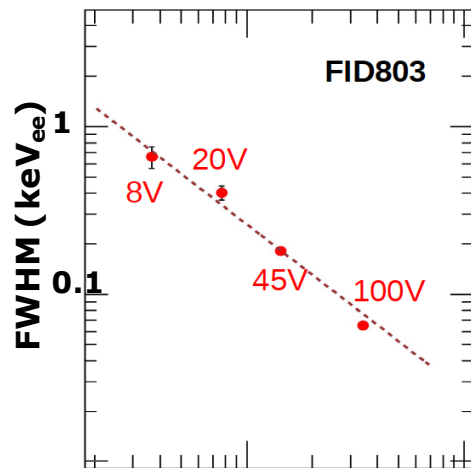
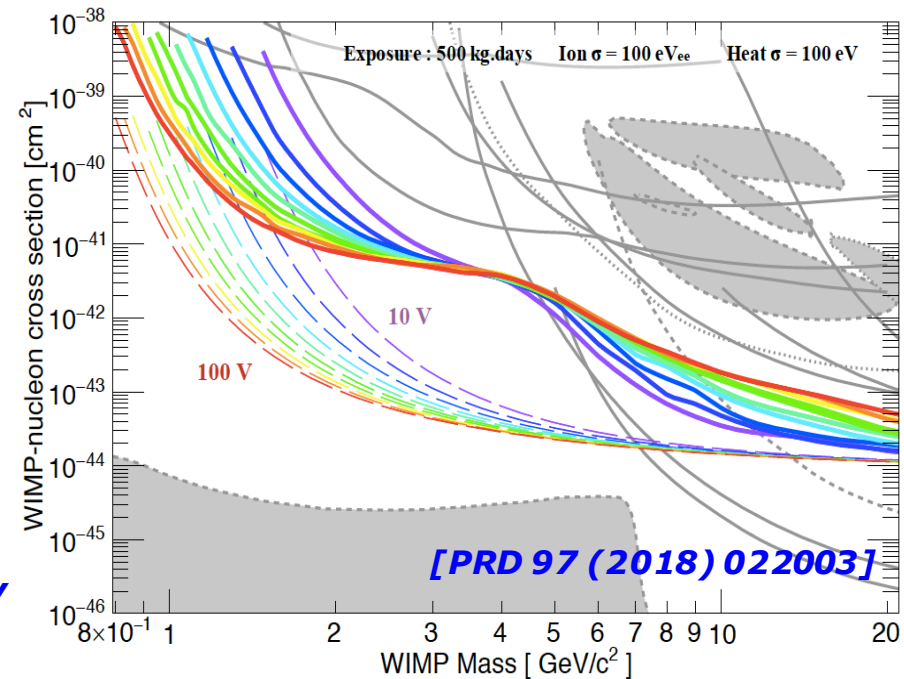


1. Use of Luke-Neganov boost to lower thresholds (up to 100V bias)
2. Improve heat resolution, objective of 800g detectors:
 $\sigma_{\text{phonon}} = 500 \text{ eV} \rightarrow 100 \text{ eV}$
(50 eV resolution already achieved on 200 g detector)
3. Reduction x100 of heat-only background

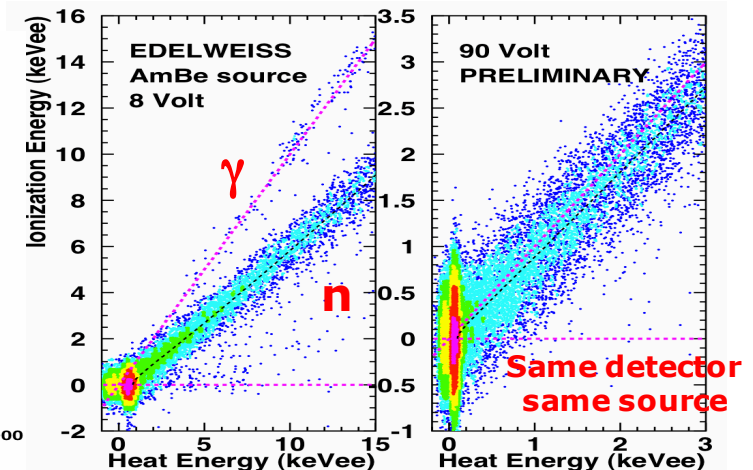
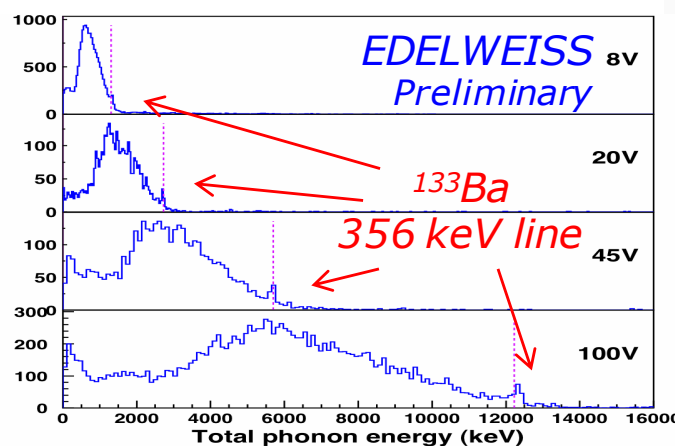
EDELWEISS-LT: Luke-Neganov boost

Heat thresholds can be improved by applying larger bias voltages

- Heat signal boosted by Neganov-Luke effect (\sim Joule heating, factor $[1 + V_{\text{bias}}/3]$)
- Loss of ionization-based bkg discrimination: method benefits low-mass searches only
 $\rightarrow 10^{-41} \text{ cm}^2$ with 500 kgd and current bkg
- ✓ **100V bias already achieved**
- ✓ **Observe nucl. recoils down to $\sim 0.1 \text{ keV}_{\text{ee}}$**
- ✓ **Full ion.+heat readout possible at any V**
- ✓ **First WIMP Data@100V analysis underway**

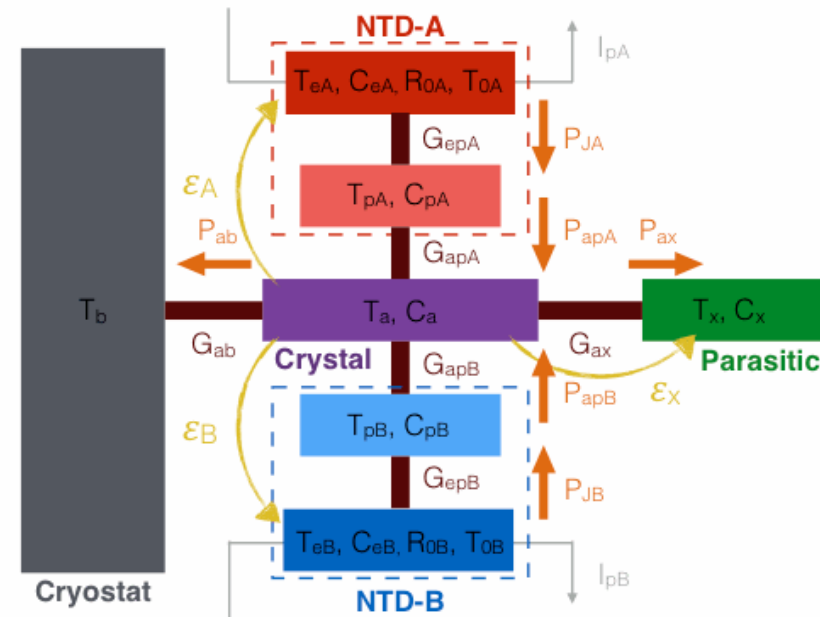


Luke boost = $1 + V/3$

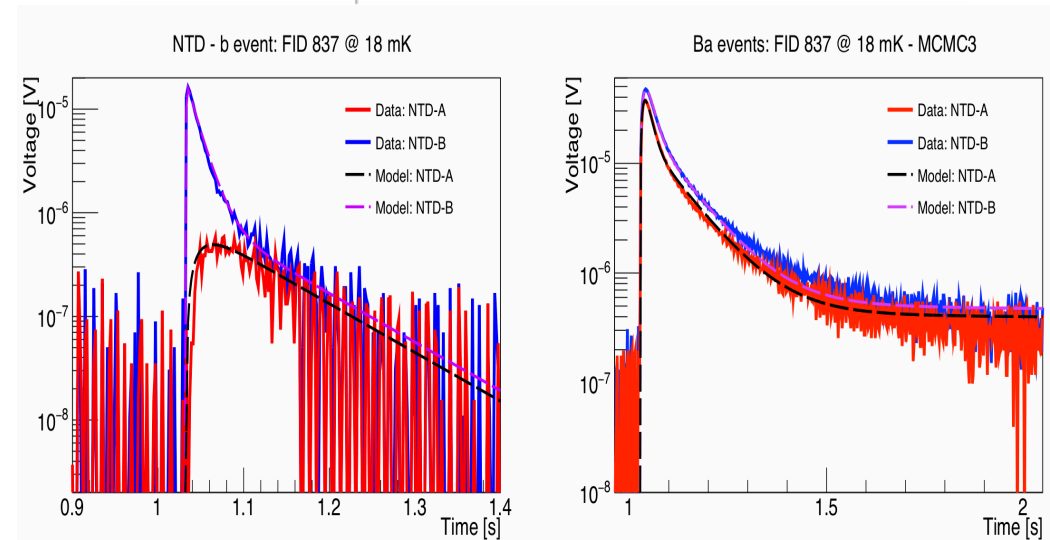


EDELWEISS-LT: heat sensor improvement

- Goal: $\sigma_{\text{phonon}} = 100 \text{ eV}$
baseline resolution on phonon signal
- Best achieved so far
on 800g: $\sigma_{\text{phonon}} = 300 \text{ eV}$
- Thermal modeling of signal, verified with dedicated measurements:
 - Identification of sensitivity to ballistic phonons
 - Identification of parasitic heat capacity
- **Also: Explore performance of high-impedance NbSi TES sensitive to athermal phonons**

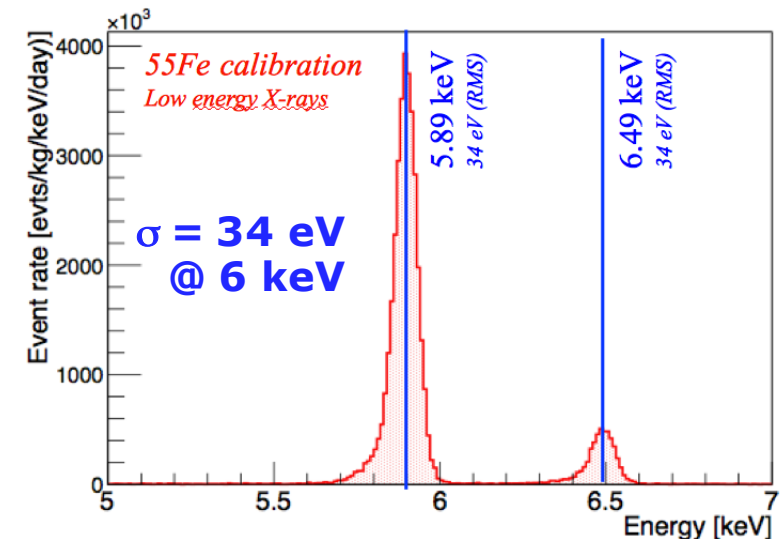
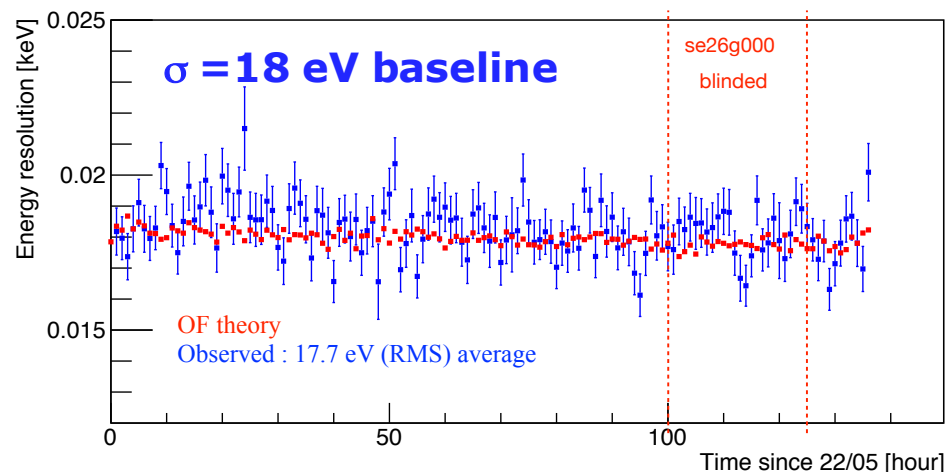
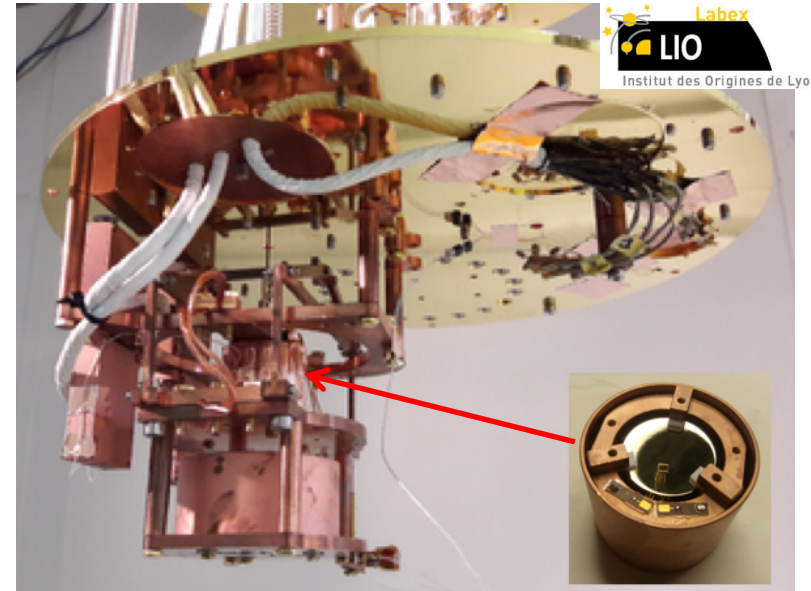


J. Billard et al., JLT(2016)184:299



Resolution improvements on a 32g detector

- R&D with 32 g combined with the objective of testing the above-ground sensitivity to sub-GeV WIMPs
- *Optimized* NTD heat sensor on a 32g crystal, no electrodes (i.e. $1 \text{ keV} = 1 \text{ keV}_{\text{NR}}$)
- Kept at 17 mK in IPNL low-vibration dilution fridge [ArXiv:1803.03463]
- Stable $\sigma = 18 \text{ eV}$ baseline resolution
- One day blinded for WIMP search in [0-2] keV region

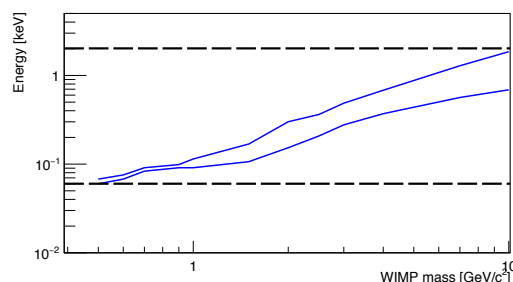


Data analysis

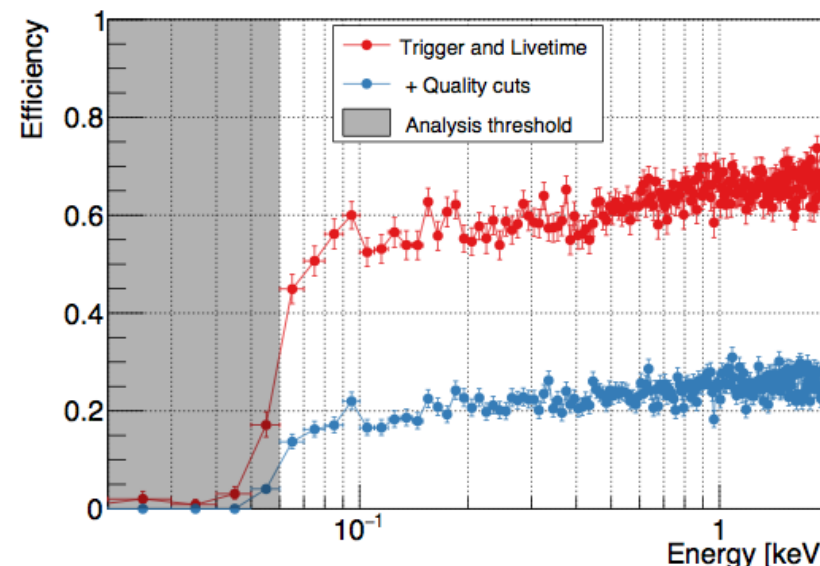
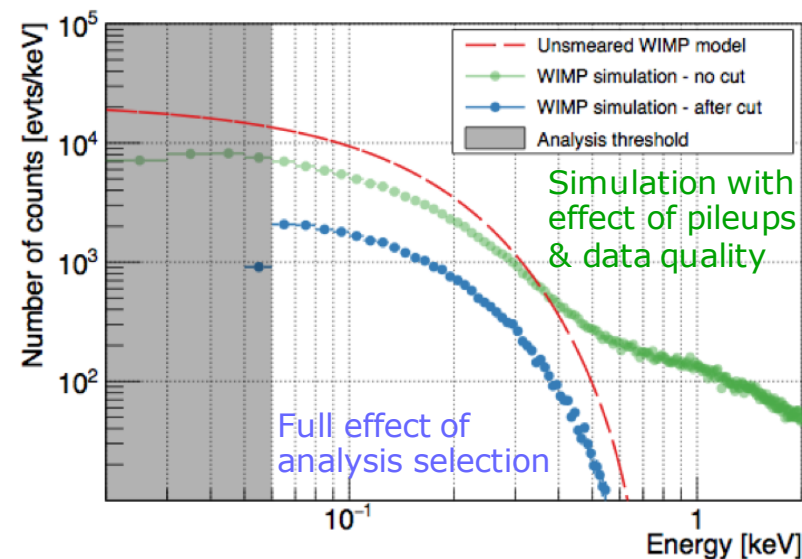
- Continuous streaming of 400 Hz data
- Optimal filtering
- Energy-dependent efficiency *and dead-time* studied by adding fake WIMP pulses to the recorded streams

→ **60 eV analysis threshold**

- Background model based on the 4.7 day of unblinded data
- Poisson limit with no background subtraction: all events within range counted as WIMP candidates
- Mass-dependent range optimized using simulations of bkg + WIMP



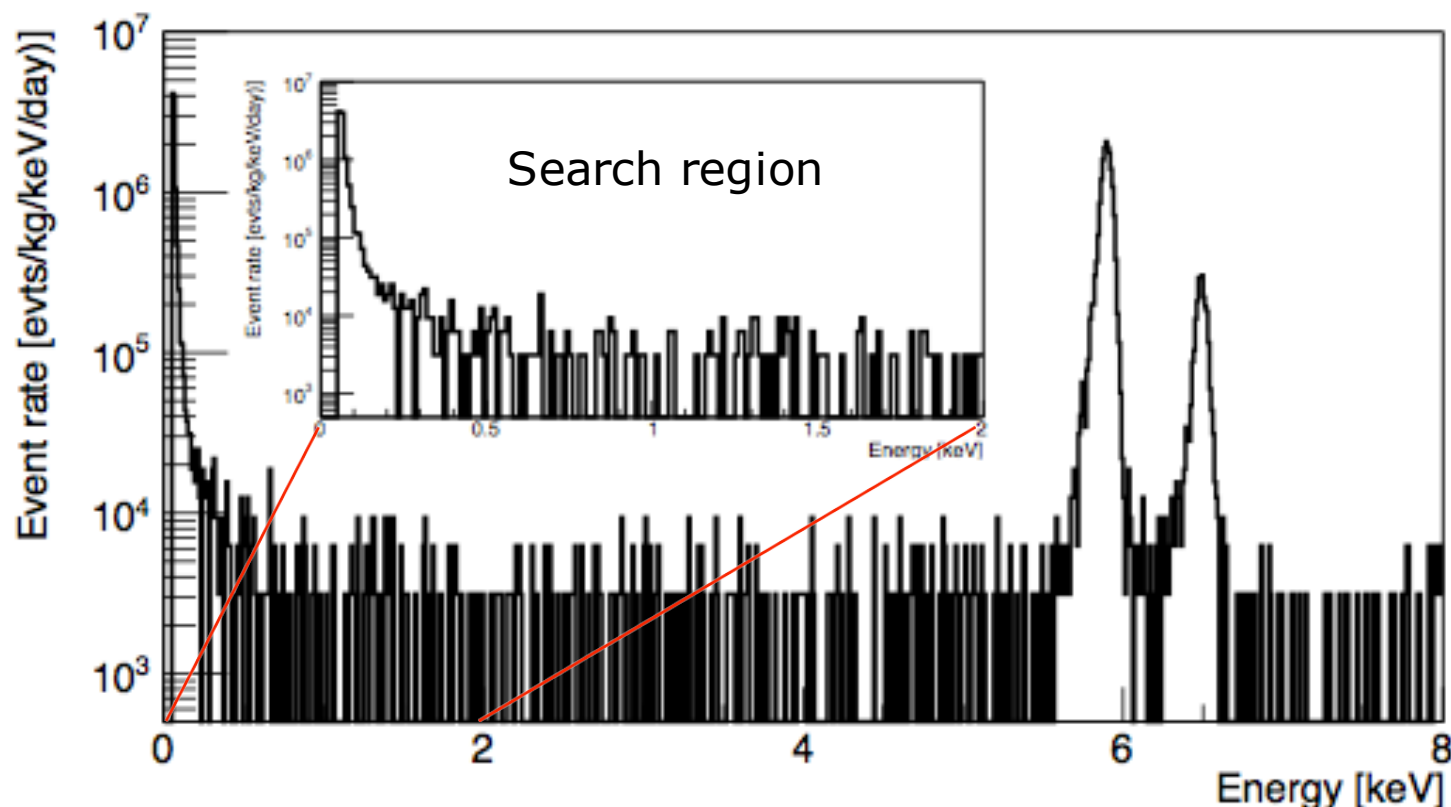
*Pre-defined
search range
vs WIMP mass*



Unblinding the data

- No surprise:

blinded day = carbon copy of preceding + following days

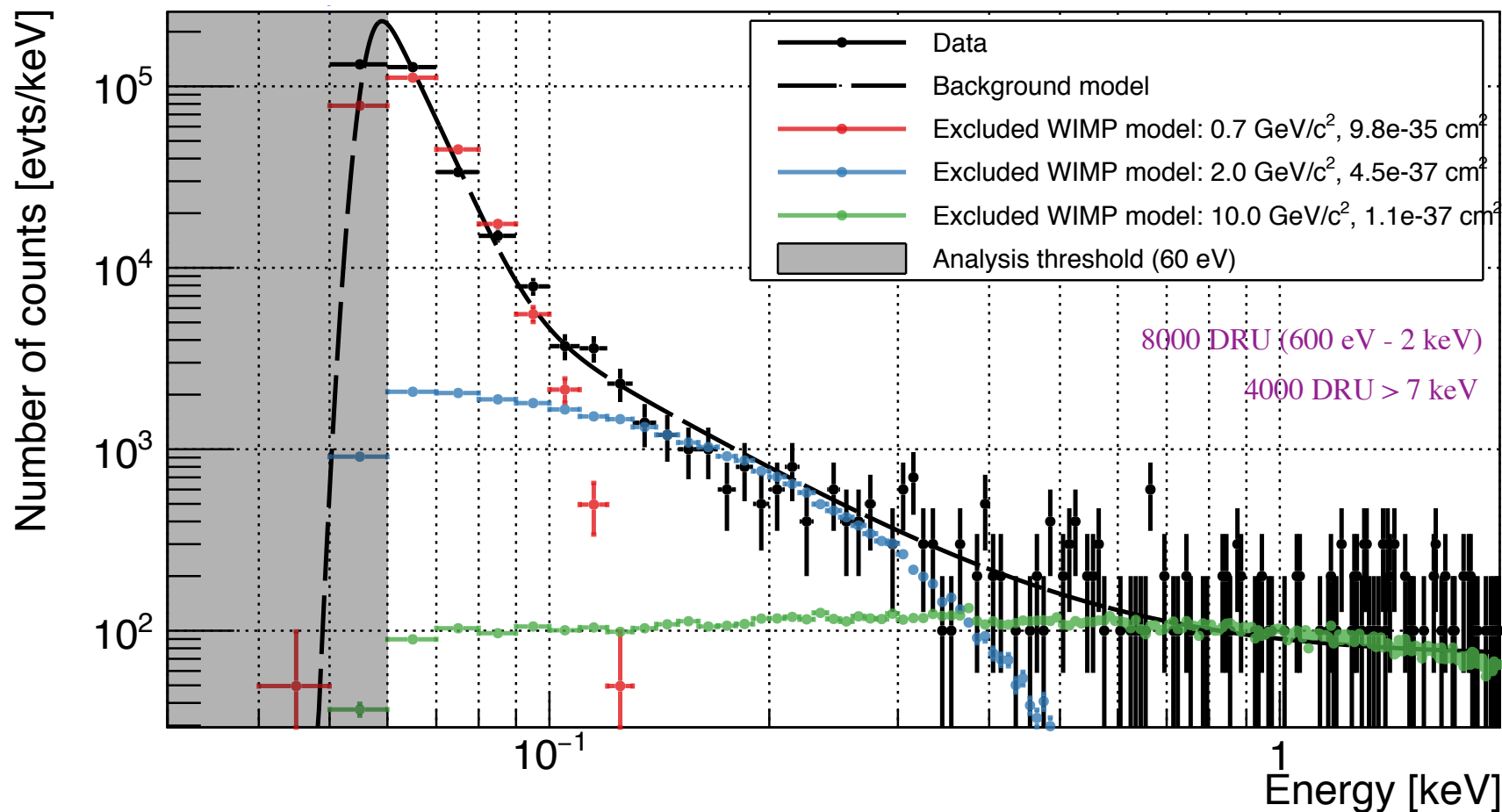


8000 DRUs 0.6-2.0 keV

4000 DRUs > 7 keV

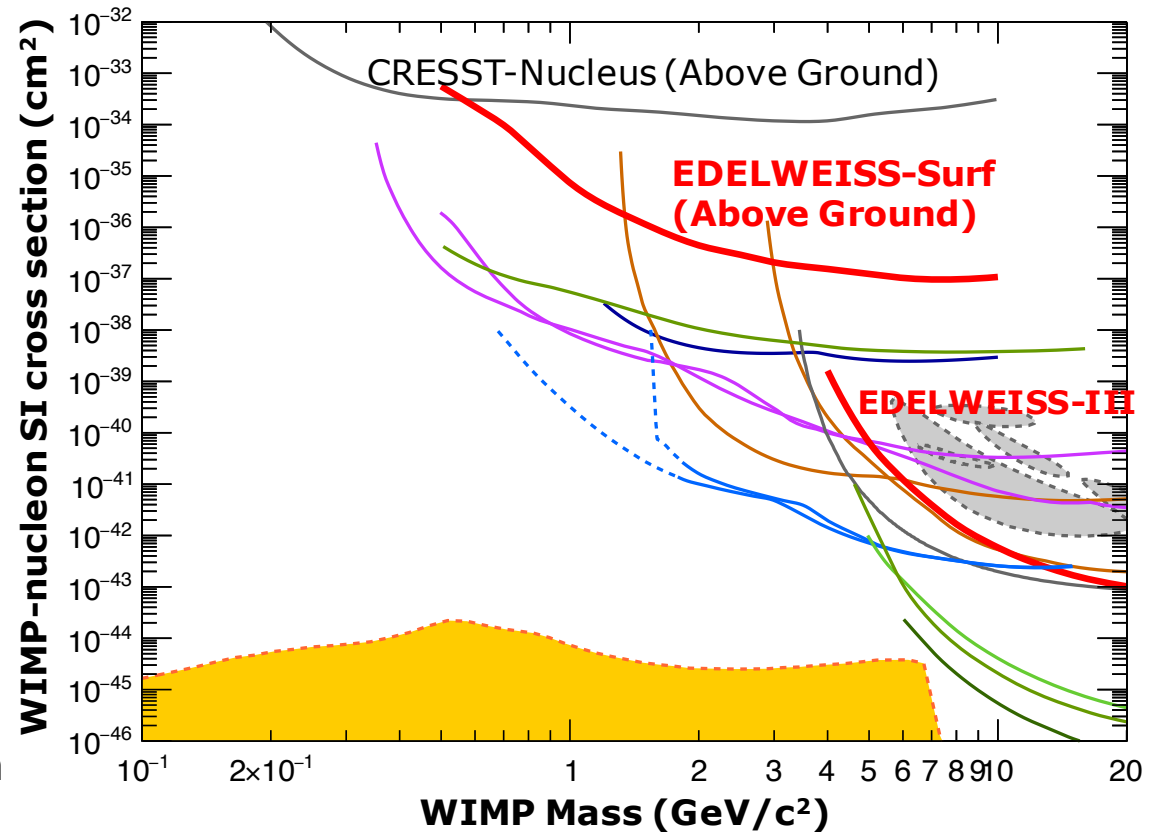
90%CL limit on WIMP signals

- Find maximal WIMP rate compatible with total number of counts observed in the pre-defined windows



Surface limit

- Achieved resolution on a smaller detector exceeds by x5 the original LT goal with 800 g detectors
- Best above-ground limit down to 600 MeV/c²: SIMP
- First sub-GeV limit with Ge, down to 500 MeV/c²
- Opens the way for the 0.1 – 1 GeV/c² range
- Small detectors with lower thresholds to be combined with expertise acquired on HV: threshold reduction by factor $(1+V_{\text{bias}}/3)$ in keV_{ee}



Conclusions

- EDELWEISS-III : large detectors with rejection
 - Excellent rejection performance, but not competitive with large Ar/Xe detectors above 6 GeV
 - Exploitation of excellent surface event rejections from FID to get best Ge ALPs limits, and enter the sub-keV range
- EDELWEISS-MELODI : develop large detector with FID design
 - Exploring non-WIMP DM with prototype: ALPs in the 0.1-1 keV range, and direct sensitivity to quenching at $\sim 0.4 \text{ keV}_{\text{NR}}$
 - Building block for a much larger search experiment (DMB8, suited for future SNOLAB program) addressing specifically the region where the DM signal needs to be spectrally resolved from a solar ^8B neutrino signal
- Prospects in the sub-GeV-WIMP range: beyond EDELWEISS-LT
 - **Going beyond original [PRD] goal: $100 \text{ eV} \rightarrow 18 \text{ eV}$ (wrt $\sim 500 \text{ eV edw3}$)**
 - **Best surface limit for WIMPs above $0.6 \text{ GeV}/c^2$**
 - **Resolutions achieved on 32g detectors + expertise acquired in HV runs open the way for the 0.1-1 GeV range**

EDELWEISS-III collaboration



CNRS/IN2P3



CNRS/IN2P3



CNRS/INP



LABORATOIRE
DE PHOTONIQUE
ET DE NANOSTRUCTURES

CNRS



Irfu

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