



## Heavy neutrinos and the search for their Dirac/Majorana nature

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#### Overview



### 2 Searches for heavy Dirac/Majorana neutrinos



• Evidence of neutrino oscillations imply: Neutrinos are massive:

$$egin{aligned} m_1, \ m_2, \ m_3: & \Delta m_{12}^2 \sim 10^{-5} \ {
m eV}^2, & \Delta m_{23}^2 \sim 10^{-3} \ {
m eV}^2. \ & \sum_i m_i \lesssim 0.1 \ {
m eV} \end{aligned}$$
 (Cosmology)

and mix:  $\nu_1, \nu_2, \nu_3$  combinations of  $\nu_e, \nu_\mu, \nu_\tau$ :

$$\sin^2 \theta_{12} \sim 0.3, \quad \sin^2 \theta_{23} \sim 0.5, \quad \sin^2 \theta_{13} \sim 0.02$$

• In the SM, neutrinos are massless (left handed Weyl fermions) Open questions:

- Why neutrinos are so much lighter than all other fermions?
- How is the SM modified to include the massive neutrinos?

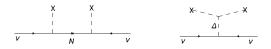
A top down approach:

A fundamental theory that explains  $m_{\nu}$  as a consequence.

- How can the SM be modified to include massive neutrinos?
- $\bullet~$  SM +~ Higgs was already not a satisfactory theory.
- Extended gauge symmetries:
  - new interactions
  - larger fermion representations (extra  $\nu_R$  fields?)
  - composite scalars?
  - other exotics?
- In terms of purely SM fields, the only (effective) term: [S. Weinberg (1979)]

$$\mathcal{L} \sim rac{L\phi L\phi}{\Lambda} \quad \Rightarrow m_
u \sim rac{v^2}{\Lambda}.$$

- Add extra fields to get a renormalizable operator at tree level. Only 3 ways:
  - Type I seesaw:  $(L\phi)$  forms a fermion singlet: extra N.
  - Type II seesaw: (LL) forms a scalar triplet: extra  $\Delta = (\Delta^{++}, \Delta^{+}, \Delta^{0})$
  - Type III seesaw: (L $\phi$ ) forms a fermion triplet: extra  $\Sigma = (\Sigma^+, \Sigma^0, \Sigma^-)$



• More possibilities... masses from loops, RpV SUSY,...

#### Let us focus on Type 1 (extra singlet fermions, N).

• A Majorana mass of the singlet N violates Lepton number (and thus B - L):

 $-\mathcal{L} \sim M_{\Lambda} \ \overline{N^c} N + \dots$ 

with  $M_{\Lambda} \sim \Lambda$  (the scale of the extra gauge and L breaking)

- E.g.: a simple UV completion is the LR symmetric model, where  $N \equiv \nu_R$  is part of a *right handed* doublet.
- ...other gauge groups such as SO(10),  $E_6$ ....

• Consider a single  $N \equiv \nu_R$  (per family) as in the LR symmetric model:

$$-\mathcal{L} = y\overline{L}\phi^{\mathcal{C}}\nu_{R} + \frac{1}{2}M\overline{\nu}_{R}^{c}\nu_{R} + \dots$$

 $\Rightarrow$  After EWSB:

$$-\mathcal{L} = \frac{1}{2} \begin{pmatrix} \bar{\nu}_L, \ \bar{\nu}_R^c \end{pmatrix} \begin{pmatrix} 0 & m_D \\ m_D^T & M \end{pmatrix} \begin{pmatrix} \nu_L^c \\ \nu_R \end{pmatrix} + \dots$$

with  $m_D = yv/\sqrt{2}$ .

• This is the original Type 1 seesaw model: [P. Minkowski (1977), T. Yanagida (1979)] Two Majorana neutrinos,  $\nu_l$  and N, with masses  $m_l \sim \frac{m_D^2}{M}$ ,  $m_N \sim M$ ,

and mixing in the EW currents:  $J^{(-)}_{\mu} = \bar{e} \gamma_{\mu} P_L \Big( V_{el} \ 
u_l + V_{eN} \ N \Big)$  :

$$V_{eN} \sim rac{m_D}{M} ~~ \sim \sqrt{rac{m_I}{M}} ~~ \sim 10^{-5} \sqrt{rac{GeV}{M}}$$

- Other Type 1, with more extra neutrinos per family:
  - e.g. low scale seesaw models [R.N. Mohapatra & J.W.F. Valle (1986)]

$$-\mathcal{L} = y\bar{L}\phi^{\mathcal{C}}\nu_{\mathcal{R}} + M \,\,\bar{S}_{L}\nu_{\mathcal{R}} + \frac{1}{2}\mu_{\mathcal{R}}\overline{\nu}_{\mathcal{R}}^{c}\nu_{\mathcal{R}} + \frac{1}{2}\mu_{\mathcal{S}}\bar{S}_{L}S_{L}^{c} + \dots$$

$$\Rightarrow -\mathcal{L} = \frac{1}{2} \begin{pmatrix} \bar{\nu}_L, \ \bar{\nu}_R^c, \ \bar{S}_L \end{pmatrix} \begin{pmatrix} 0 & m_D & 0 \\ m_D^T & \mu_R & M \\ 0 & M^T & \mu_S \end{pmatrix} \begin{pmatrix} \nu_L^c \\ \nu_R \\ S_L^c \end{pmatrix} + \dots$$

- Now the masses:  $m_l \sim \mu_S \frac{m_D^2}{M^2}, \quad m_{N_{1,2}} \sim M, \quad \Delta m_N \sim \mu_S$
- L is conserved and one massless mode if  $\mu_S = \mu_R = 0$ .
- $u_l$ : one light Majorana; the smallness is dictated by a small  $\mu_S$
- $N_{1,2}$ : two heavy Majorana (quasi-Dirac if  $\Delta m_N \lesssim \Gamma_N$ ).
- Mixing is less constrained:  $V_{eN} \sim \frac{m_D}{M} \sim \sqrt{\frac{m_l}{\mu_S}}$
- More seesaw scenarios (linear, double, lobsided, ...) [E. Ma (2009)]

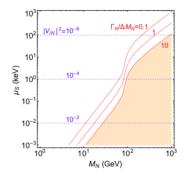
• The decay width also depends on  $m_N$  and  $V_{\ell N}$ :

$$\Gamma_N \sim rac{G_F^2}{10\pi^3} m_N^5 \sum_\ell \left|V_{\ell N}
ight|^2$$

$$\frac{\Gamma_N}{\Delta m_N} \sim \left(\frac{1 \text{keV}}{\mu_S}\right)^2 \left(\frac{m_N}{100 \text{GeV}}\right)^2$$

See figure:

- white zone: N is Majorana.
- red zone: N is Dirac.



from F. Deppisch, P.S.B. Dev & A. Pilaftsis (2015)

Therefore:

- Extra neutrinos can be in a wide range of masses (keV  $\rightarrow$  GUT scale).
- Heavy-to-light mixings also cover a wide theoretical range (there are experimental constraints).
- It is important to determine the Dirac or Majorana character of N's to distinguish between different scenarios.

#### Current bounds on $V_{eN}$ :

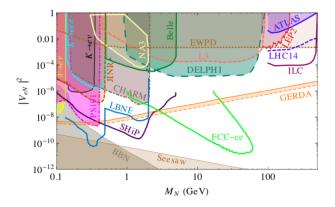


Figure from F. Deppisch, P.S.B. Dev & A. Pilaftsis, New J.Phys. 17, 075019 (2015).

#### Current bounds on $V_{\mu N}$ :

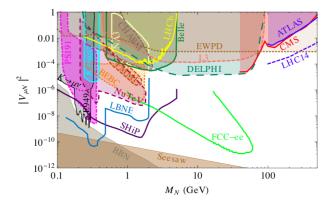


Figure from F. Deppisch, P.S.B. Dev & A. Pilaftsis, New J.Phys. 17, 075019 (2015). ...and more recent bound from LHCb [S. Antusch, E. Cazzato & O. Fisher (2017)].

#### Current bounds on $V_{\tau N}$ :

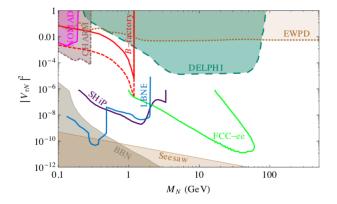
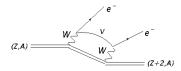


Figure from F. Deppisch, P.S.B. Dev & A. Pilaftsis, New J.Phys. 17, 075019 (2015).

- ♦ Search for heavy sterile neutrinos in processes:
  - Majorana:
    - violation of Lepton Flavor (LFV),
    - violation of Lepton Number (LNV).
  - Dirac:
- violation of LFV only (L# is conserved).
- $\diamond$  Search for Majorana N: best processes are nuclear 0uetaeta decays

...but there can be cancellations even if N are Majorana (e.g. CP phases).



- ♦ Search for heavy sterile neutrinos at colliders:
  - For  $m_N > M_W$ , preferred channel is  $pp \to \ell^{\pm} \ell^{\pm} j j$  (Majorana N only). (see also talk by O. Fischer on LFV  $e^{\pm} \mu^{\mp} j j$  at future colliders).

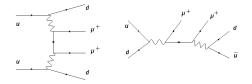


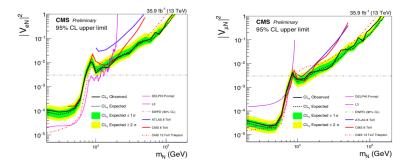
Figure : Left: *N* in the *t*-channel; right: dominant for resonant *N*.

- $\sigma(pp \rightarrow W)$  decreases with  $m_N$ : LHC sensitive up to  $\sim 1$  TeV
- Larger  $\sigma(pp \rightarrow W_R)$  for larger  $m_N$ : stronger bounds for LR sym. model.

Search by CMS at  $\sqrt{s} = 13$  TeV with 35.9  $fb^{-1}$  pp collisions into same-sign dilepton + jets.

[CMS PAS-EXO 17-028 (2018)]

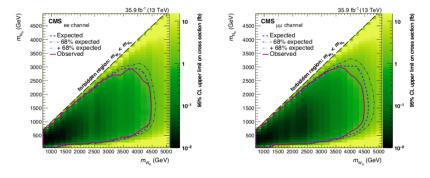
• Found no signal above background: bounds on  $|U_{eN}|$  and  $|U_{\mu N}|$  for  $m_N < 1$  TeV.



Search for  $W_R$  and N by CMS at  $\sqrt{s} = 13$  TeV with 35.9  $fb^{-1}$  $pp \rightarrow W_R \rightarrow \ell N \rightarrow \ell \ell W_R^* \rightarrow \ell \ell j j$ 

[CMS coll., JHEP 1805, 148 (2018)]

• Larger expected rates imply bounds  $M_{W_R} > 4.5 \text{ TeV}$ and  $m_N$  up to 3 TeV.



◊ Search for heavy sterile neutrinos at colliders:

- For  $m_N > M_W$ , preferred channel is  $pp \to \ell^\pm \ell^\pm j j$  (Majorana N only)...
- ...but for  $m_N < M_W$ , jets are softer  $\Rightarrow$  trilepton modes are preferred.

$$p p 
ightarrow e^{\pm} e^{\pm} \mu^{\mp} 
u, \quad p p 
ightarrow \mu^{\pm} \mu^{\pm} e^{\mp} 
u$$

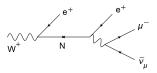
[E. Izaguirre & B. Shuve (2015)]

- Missing final  $\nu$ : cannot test *L* conservation directly.
- Majorana N: sum of two rates (LNV + LNC): e.g.  $W^+ \rightarrow e^+ e^+ \mu^- \bar{\nu}_\mu + W^+ \rightarrow e^+ e^+ \mu^- \nu_e$
- Dirac N:

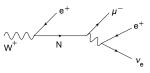
LNC rate only:  $W^+ 
ightarrow e^+ e^+ \mu^- 
u_e$ .

 $W^+ 
ightarrow e^+ e^+ \mu^- 
u$ :

• Lepton Number Violating (LNV) mode: caused by Majorana N only:  $W^+ \rightarrow e^+ e^+ \mu^- \bar{\nu}_{\mu}$ :



• Lepton Number Conserving (LNC) mode: caused by both Majorana and Dirac  $N: W^+ \rightarrow e^+ e^+ \mu^- \nu_e$ :



- How can we tell whether N is Dirac or Majorana?
- Exploit the fact that  $\mu^-$  (the odd lepton) in the N decay is attached to a different current.

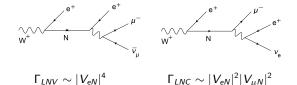
 $W^+ 
ightarrow e^+ e^+ \mu^- 
u$ :

- How to tell a Majorana from a Dirac N:
- Exploit the fact that  $\mu^-$  (the odd lepton) in the N decay is attached to a different current.
- Three possible differences:
  - 1. Different heavy-to-light mixing elements (different rates).
  - 2. Different  $\mu$  energy spectrum.
  - 3. Different  $\mu$  angular distributions.

1. Different heavy-to-light mixing elements (different rates).

[C.D., C.S. Kim, K. Wang, J. Zhang (2016)]

• Consider e.g.  $W^+ \rightarrow e^+ e^+ \mu^- \nu$ :



• If N is Dirac:

$$\begin{array}{l} \Gamma(e^+e^+\mu^-) \sim |U_{eN}|^2 |U_{\mu N}|^2 \\ \Gamma(\mu^+\mu^+e^-) \sim |U_{eN}|^2 |U_{\mu N}|^2 \end{array} \Rightarrow \Gamma(e^+e^+\mu^-) = \Gamma(\mu^+\mu^+e^-) \end{array}$$

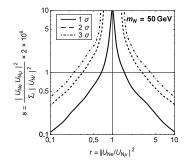
• If *N* is Majorana:

$$\begin{array}{ll} \Gamma(e^+e^+\mu^-) \sim |V_{eN}|^4 + |V_{eN}|^2 |V_{\mu N}|^2 & \Rightarrow \Gamma(e^+e^+\mu^-) \neq \Gamma(\mu^+\mu^+e^-) \\ \Gamma(\mu^+\mu^+e^-) \sim |V_{\mu N}|^4 + |V_{eN}|^2 |V_{\mu N}|^2 & \text{if } |V_{eN}| \neq |V_{\mu N}| \end{array}$$

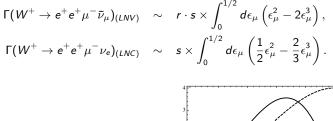
- Different heavy-to-light mixing elements (different rates). [C.D., C.S. Kim, K. Wang, J. Zhang (2016)]
  - $|V_{eN}| \neq |V_{\mu N}|$  ... but how different must they be?
  - Simulations for 14 TeV *pp* collisions with 3000 fb<sup>-1</sup>:

Figure: confidence levels for excluding the Dirac case given a Majorana scenario, for  $m_N = 50$  GeV.

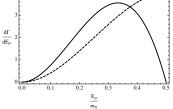
• For larger s (more events),  $r \equiv \frac{|V_{eN}|^2}{|V_{\mu N}|^2} \text{ can be closer to } 1.$ 



- 2. Different  $\mu$  energy spectrum in  $W^+ \rightarrow e^+ e^+ \mu^- \nu$ . [C.D. & C.S. Kim, PRD 92, 093009 (2015)]
  - Muon energy in the N rest frame:  $\epsilon_{\mu} = E_{\mu}/m_N$
  - The spectra:



Muon spectrum: - LNV (solid); - LNC (dashed);

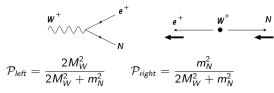


3. Different  $\mu$  angular distributions.

[C. Arbelaez, C.D., I. Schmidt & J.C. Vasquez (2018)]

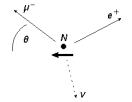
 $W^+ 
ightarrow {
m e}^+ {
m N}$ , followed by  $N 
ightarrow {
m e}^+ \mu^- ar{
u}_\mu$  or  $N 
ightarrow \mu^- {
m e}^+ 
u_e$  :

• N produced mainly with left handed helicity in  $W^+ \rightarrow e^+ N$ :

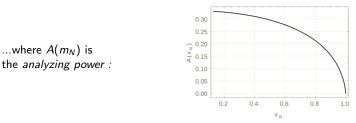


• Polarized N decays have different angular distributions:

$$\frac{d\Gamma(N_{pol} \to e^+ \mu^- \bar{\nu}_{\mu})}{d\cos\theta} \sim (1 - \cos\theta),$$
$$\frac{d\Gamma(N_{pol} \to \mu^- e^+ \nu_e)}{d\cos\theta} \sim (1 + A(m_N)\cos\theta)$$



3. Different  $\mu$  angular distributions.



Different forward-backward asymmetries of the muon in the N frame, depending on whether N is Dirac or Majorana:

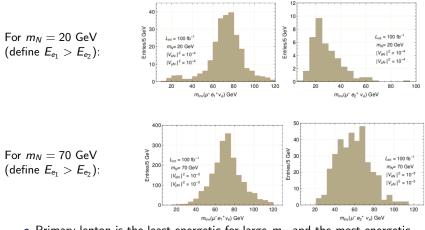
$$\mathcal{A}_{FB} = rac{\mathcal{N}(\cos heta > 0) - \mathcal{N}(\cos heta < 0)}{\mathcal{N}(\cos heta > 0) - \mathcal{N}(\cos heta < 0)}.$$

• Dirac:  $A_{FB} = -1/6$  independent of mixings  $|V_{\ell N}|$ .

• Majorana: depends on 
$$|V_{\ell N}|$$
.  
If  $|V_{eN}| > |V_{\mu N}|$ ,  $A_{FB} = -1/2$  ( $\rightarrow -1/6$  as  $m_N \rightarrow M_W$ ).  
If  $|V_{eN}| < |V_{\mu N}|$ , use the mode  $\mu^{\pm} \mu^{\pm} e^{\mp}$  instead.

#### Some difficulties:

• need to determine the N rest frame: need to identify the primary positron.



• Primary lepton is the least energetic for large  $m_N$  and the most energetic for smaller  $m_N$ .

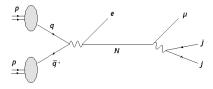
4. (there is always a forth way) put everything together: MVA (multivariate analysis)

[C.D., C.S. Kim, Kechen Wang (2017)]

- Simulations for 14 TeV pp collisions with integrated luminosity of 3000 fb<sup>-1</sup>.
- Majorana-Dirac distinction can be reached at  $5\sigma$  significance for  $m_N$  between 20 GeV and 75 GeV, provided:
  - $\bullet |V_{eN}|^2 \sim |V_{\mu N}|^2 \gtrsim 10^{-5}$
  - $|V_{eN}|^2$  or  $|V_{\mu N}|^2\gtrsim 10^{-6}$  and the other is considerably smaller.
- For  $m_N \lesssim 15$  GeV the N production and decay vertices are spatially separated in the detector.

This vertex displacement drastically helps on the kinematics and on BKG reduction.

- $\diamond$  Search for  $m_N < 15$  GeV at colliders: displaced vertices.
  - $\bullet$  For  $m_N \lesssim 15~\text{GeV}$  , N lives long. i.e. decay vertex visibly displaced from production vertex:



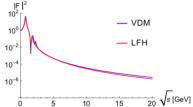
• New bounds on  $|V_{\mu N}|^2$  in the range  $m_N = 4.5 - 10$  GeV from LHCB [S. Antusch, E. Cazzato & O. Fischer, Phys.Lett. B774, 114 (2017)].

- $\diamond$  Search for  $m_N < 15$  GeV at colliders: exclusive semileptonic modes.
  - Semileptonic modes may be difficult to detect for low  $m_N$  (low  $p_t$  jets lost in bkgnd).
  - Exclusive semileptonic with displaced vertices may help: try

$$N \to \ell^{\pm} \pi^{\mp}, \quad N \to \ell^{\pm} \pi^{\mp} \pi^{0}, \quad N \to \ell^{\pm} \pi^{\pm} \pi^{\mp} \pi^{\mp}.$$

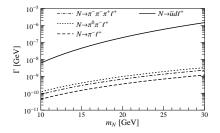
[C.D., C.S. Kim, N. Neill & X-B. Yuan (2017)]

For 2-pion and 3-pion modes, need for factors. VDM for 2-pion and 3-pion. [BaBar coll., PRD 86, 032013 (2012)] and Light Front Holography (LFH) for 2-pion [S. Brodsky et al, P. Rept. 584, 1 (2015)] .



 $\diamond$  Search for  $m_N < 15$  GeV at colliders: exclusive semileptonic modes.

- Single pion mode is suppressed by  $f_{\pi}/m_N$ . -Two-pion and three-pion modes are comparable.



- Can 2-pion and 3-pion modes be seen at the LHC?
- 2-pion mode has one neutral pion (two-photon signal???)
- 3-pion signal has no neutrals, but it is a bit smaller.
- Must study the observability of modes:  $pp \rightarrow W \rightarrow \ell \ell (n\pi)$ .
- For a sample of ~ 10<sup>9</sup> W's by the end of LHC Run 2, ATLAS / CMS should be able to put new limits  $|V_{\ell N}|^2 \sim 2 \times 10^{-6}$  in the mass range 5 GeV <  $m_N$  < 15 GeV.
- Study of particle id. and background rejection (in progress).

#### Summary

- Neutrino masses lead us to physics beyond the Standard Model.
- Most explanations for the very small neutrino masses imply the existence of extra neutrinos, either Majorana or Dirac ...or in between.
- It is important to know the Dirac / Majorana character of neutrinos to distinguish between theoretical scenarios.
- Here we focus on searches for heavy neutrinos in high energy collisions:
- Semileptonic modes:  $\ell^{\pm}\ell^{\pm}jj$ ,  $e^{\pm}\mu^{\mp}jj$ .
- Trilepton modes  $e^{\pm}e^{\pm}\mu^{\mp}\nu$ ,  $\mu^{\pm}\mu^{\pm}e^{\mp}\nu$ .
- Dirac/Majorana character difficult to detect in trileptons ... ... but it can be done.
- For  $m_N < 15$  GeV, displaced vertices due to long living N helps in background reduction (can use semileptonic modes again).



Thank you