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Massive stealth fields from deformation method

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Based in the collaboration:

Cristian C. Quinzacara, Paola Meza, Almeira Sampson & M.V. Eur. Phys. J. C (2018) 78. [arXiv:1805.04621]

1 Introduction

It is generally believed that matter curves the space since given an action principle of pure gravity and matter,

$$
S[g, \phi] = S_G[g] + S_M[g, \phi], \tag{1.1}
$$

it happens that the equations of motion of the metric tensor *g*,

$$
\frac{\delta S[g,\phi]}{\delta g^{\mu\nu}} = \sqrt{-g} \left(\underbrace{G_{\mu\nu}[g]}_{\text{Einstein tensor}} - \underbrace{T_{\mu\nu}[g,\phi]}_{\text{Hilbert EM tensor}} \right) = 0. \tag{1.2}
$$

Hence matter curves the spacetime...

However, there are many examples in the literature where this is not the case: Alvarez et al. (2016, 2017), Ayon-Beato et al. (2005, 2006), Ay´on-Beato et al. (2013, 2015, 2018), Hassaine (2014), Smolić (2018).

Here the energy momentum tensor vanishes for non-trivial solutions of the matter field equations. In this case, the geometry is not curved in the presence of non-trivial matter fields configuration. These configurations are dubbed *Stealth Fields.*

The purpose of this paper is to **present a method to construct models with massive stealth fields** in arbitrary backgrounds.

The scalar field action is "deformed" into a new action with deformation parameter *θ*.

The deformed action possesses:

- a massive stealth mode of mass θ^{-1}
- other modes with rescaled effects on the gravity background

2 A toy model

Let $f(x)$ be any function of $x \in \mathbb{R}$ with a saddle point at $x = 0$,

$$
\left. \frac{df}{dx} \right|_{x=0} = 0, \tag{2.3}
$$

and let $y(x)$ another function which possess, for definiteness, two zeros at $x = 0$ and $x = 1$,

$$
y(0) = 0, \t y(1) = 0. \t (2.4)
$$

We define the composition of functions $F(x) := f(y(x))$, which inherits from the parent functions *f* and *y* the properties,

$$
\left. \frac{dF}{dx} \right|_{x=0} = \left. \frac{dF}{dx} \right|_{x=1} = 0, \tag{2.5}
$$

.

such that it has two saddle points, at $x = 0, 1$. It is straightforward to prove this. We can use the chain rule to evaluate dF/dx at $x = 0, 1$,

$$
\frac{dF}{dx}\bigg|_{x=0,1} = \left(\frac{df(y)}{dy}\frac{dy(x)}{dx}\right)\bigg|_{x=0,1} = \frac{df(y)}{dy}\bigg|_{y=0} \frac{dy(x)}{dx}\bigg|_{x=0,1}
$$

Here the $df(y)/dy|_{y=0}$ vanishes because from (2.4) *y* takes zero-value and from (2.3) the derivative of *f* vanishes when the argument is zero.

Hence from an arbitrary function $f(x)$ with saddle point at $x = 0$ (2.3) we can construct another arbitrary function $F(x)$ with saddle points at the **kernel of the map** $y : \mathbb{R} \to \mathbb{R}$, in this example $x = 0, 1$.

3 *θ***-deformation of scalar field theories**

We shall apply now the same logic in the language of functional calculus, with the dictionary:

$$
x \to \phi
$$
, $y(x) \to \phi^{\theta}[g, \phi]$, $f(x) \to S_M[g, \phi]$, $F(f(x)) \to S_M[g, \phi^{\theta}[g, \phi]]$.

Here the scalar field ϕ is the analogous of *x*, ϕ^{θ} is a funcional map from ϕ , $S_M[g, \phi]$ is the action principle of a scalar field (with saddle points) analogous of *f*, and $S_M[g, \phi^{\theta}[g, \phi]]$ is a new action principle obtained from a composition of the functionals S_M and ϕ^{θ} analogous of $F(f(x))$.

The field transformation to be considered is:

$$
\phi^{\theta}[g,\phi] = (1 - \theta^2 \Box)\phi,\tag{3.6}
$$

where θ is a real-valued parameter. Here

$$
\Box \phi := \frac{1}{\sqrt{(-g)}} \partial_\mu (\sqrt{-g} g^{\mu\nu} \partial_\nu \phi) ,
$$

is the Laplace-Beltrami operator acting upon *ϕ*.

Therefore the kernel of the deformation map $\phi \to \phi^{\theta}[g, \phi]$ consists of the trivial vacuum $\phi = 0$ and the massive configuration $\phi = \phi_m$ with mass $m = \theta^{-1}$:

$$
\phi^{\theta}[g,\phi_m] = (1 - \theta^2 \square)\phi_m = 0, \qquad m = \theta^{-1}, \qquad (3.7)
$$

which is equivalent to the Klein-Gordon equation in curved space.

4 Scalar field action deformation

We shall consider theories of the type,

$$
S[g, \phi] = S_G[g] + S_M[g, \phi].
$$
\n(4.8)

$$
\delta\phi \qquad \rightarrow \qquad \Upsilon[g,\phi] := \frac{\delta S_M[g,\phi]}{\delta\phi} = \frac{\delta S[g,\phi]}{\delta\phi} = 0, \tag{4.9}
$$

Hence the E.o.M are:

$$
G_{\mu\nu}[g] - T_{\mu\nu}[g,\phi] = 0, \qquad \Upsilon[g,\phi] = 0,
$$
\n(4.10)

Now we produce a "deformed" field theory,

$$
S^{\theta}[g,\phi] = S_G[g] + S_M[g,\phi^{\theta}], \qquad \phi^{\theta}[g,\phi] := (1 - \theta^2 \square)\phi \qquad (4.11)
$$

with E.o.Ms:

$$
G_{\mu\nu}[g] - \widetilde{T}_{\mu\nu}[g,\phi] = 0, \qquad \widetilde{\Upsilon}[g,\phi] = 0,
$$
\n(4.12)

We claim that:

$$
\widetilde{\Upsilon}[g,\phi] = 0, \qquad G_{\mu\nu}[g,\phi] = 0, \qquad \widetilde{T}_{\mu\nu}[g] = 0, \qquad (4.13)
$$

for the massive mode of mass $m = \theta^{-1}$.

Indeed, in reference **Cristian C. Quinzacara, Paola Meza, Almeira Sampson & M. V. Eur. Phys. J. C (2018) 78. [arXiv:1805.04621]** we proved that the following statements are true:

- Let $\phi = 0$ (trivial vacuum) be a saddle point of $S_M[g, \phi]$ (i.e. a solution of the equations of motion), then the deformed action $S_M[g, \phi^{\theta}]$ has a saddle point at $\phi^{\theta} = 0$, i.e. for massive ϕ of mass $m = \theta^{-1}$.
- The energy momentum tensor of the massive ϕ ($m = \theta^{-1}$) vanishes (ϕ is a stealth field), hence it does not curve the spacetime.

The proof is very general, and it uses the functional generalization of the chain rule,

$$
\frac{\delta F[G[f]]}{\delta f(y)} = \int d^D z \frac{\delta F[G[f]]}{\delta G[f](z)} \frac{\delta G[f](z)}{\delta f(y)},\tag{4.14}
$$

to be confronted with our toy model.

5 Example 1:

The simplest example we can imagine is:

$$
\widetilde{S}[g,\phi] = -\frac{M^2}{2} \int d^D x \sqrt{-g} \phi^2, \qquad (5.15)
$$

The equation of motion for the scalar field is, $\phi = 0$, and its energy momentum tensor vanishes for this solution.

The deformed action is:

$$
\widetilde{S}[g,\phi^{\theta}] = -\frac{M^2}{2} \int d^D x \sqrt{-g} (\phi^{\theta}[g,\phi_m])^2 = -\frac{M^2}{2} \int d^D x \sqrt{-g} (\phi^2 - 2\theta^2 \phi \Box \phi + \theta^4 (\Box \phi)^2), \qquad \phi^{\theta} \tag{5.16}
$$

where we replaced the original field ϕ by ϕ^{θ} .

This action can be regarded as a degenerated (single-parameter) case of the twoparametric fourth-order action principle analyzed in Hawking and Hertog (2002).

The equation of motion for *ϕ* yields:

$$
(1 - \theta^2 \square)^2 \phi = 0. \tag{5.17}
$$

Clearly satisfied by massive ϕ of mass θ^{-1} . A direct calculation of the energy momentum tensor produces:

$$
\widetilde{T}_{\mu\nu}[g,\phi] = -\frac{1}{4}M^2 g_{\mu\nu}\left((1-\theta^2 \Box)\phi\right)^2 - \frac{1}{2}M^2\theta^2 g_{\mu\nu}\Box\phi(1-\theta^2 \Box)\phi \n+ \frac{1}{2}M^2\theta^2\left(\delta^{\rho}_{\mu}\delta^{\sigma}_{\nu} + \delta^{\rho}_{\nu}\delta^{\sigma}_{\mu} - g_{\mu\nu}g^{\rho\sigma}\right)\nabla_{\rho}\phi\nabla_{\sigma}\left((1-\theta^2 \Box)\phi\right), (5.18)
$$

which is evidently zero for the massive $\phi = \phi_m$.

6 Example 2: Deformation of the massive field action

In the case of the scalar field ϕ , with mass *M* the matter action principle is given by:

$$
S_M[g,\phi] = -\frac{1}{2} \int d^D x \sqrt{-g} \left(\nabla^\mu \phi \nabla_\mu \phi + M^2 \phi^2 \right). \tag{6.19}
$$

With E.o.M:

$$
\left(\Box - M^2\right)\phi = 0. \tag{6.20}
$$

The deformed action principle is

$$
S_M[g,\phi] = -\frac{1}{2} \int d^D x \sqrt{-g} \left(\nabla^\mu \phi^\theta \nabla_\mu \phi^\theta + M^2 (\phi^\theta)^2 \right) , \qquad \phi^\theta[g,\phi] = (1 - \theta^2 \Box) \phi.
$$
\n(6.21)

We obtain the E.o.M (from variation with respect to ϕ):

$$
\left(1 - \theta^2 \Box\right)^2 \left(\Box - M^2\right) \phi = 0, \qquad (6.22)
$$

which has both solutions, of mass θ^{-1} and of mass M.

The energy momentum tensor is given by:

$$
\widetilde{T}_{\mu\nu}[g,\phi] = -\frac{1}{2} \left(\frac{1}{2} g_{\mu\nu} \left(\nabla^{\rho} \phi^{\theta} \nabla_{\rho} \phi^{\theta} + M^{2} (\phi^{\theta})^{2} \right) - \nabla_{\mu} \phi^{\theta} \nabla_{\nu} \phi^{\theta} \right) + \frac{1}{2} \theta^{2} g_{\mu\nu} \Box \phi \left(\Box - M^{2} \right) \phi^{\theta} \n- \frac{1}{2} \theta^{2} \left(\delta^{\rho}_{\mu} \delta^{\sigma}_{\nu} + \delta^{\rho}_{\nu} \delta^{\sigma}_{\mu} - g_{\mu\nu} g^{\rho\sigma} \right) \nabla_{\rho} \phi \nabla_{\sigma} \left(\Box - M^{2} \right) \phi^{\theta}, \quad (6.23)
$$

which vanishes for the field of mass θ^{-1} , because $\phi^{\theta} = (1 - \theta^2) \phi = 0$.

Hence ϕ of mass θ^{-1} is **stealth**.

Now we analyze the energy-momentum for ϕ_M . First note that:

$$
\phi^{\theta}[g,\phi_M] := (1 - \theta^2 \Box)\phi_M = \lambda \phi_M, \qquad \lambda := 1 - \frac{M^2}{m^2} = 1 - M^2 \theta^2, \qquad (6.24)
$$

Hence the energy-momentum tensor for this solution is (6.23) given by:

$$
\widetilde{T}_{\mu\nu}[g,\phi_M] = -\frac{1}{2}\lambda^2 \left(\frac{1}{2}g_{\mu\nu}\left(\nabla^\rho\phi_M\,\nabla_\rho\phi_M + M^2\phi_M^2\right) - \nabla_\mu\phi_M\,\nabla_\nu\phi_M\right) = \lambda^2\,T_{\mu\nu}[g,\phi_M]\,,\tag{6.25}
$$

where $T_{\mu\nu}[g,\phi_M]$ is the energy momentum tensor provided by the standard massive field action (6.21).

Therefore, the energy-momentum tensor of the original gravity-matter system is rescaled by factor λ^2 in the deformed theory.

This can be interpreted also as a rescaling of the Newton coupling constant by a factor, $G_{\text{N}} \to \lambda^2 G_{\text{N}}$, in the standard nomenclature.

 \rightarrow Hence the mass of the stealth field (equivalently the deformation parameter) can be used to smooth or amplify the effects of the massive field of mass *M* on the gravitational background.

7 Overview and remarks

- We can construct a wide class of scalar field action principles in curved space which admits massive stealth configurations.
- The existence of stealth configurations may produce cosmological effects, by means of the energy-momentum tensor rescaling of regular matter fields.

Thank you !

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