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Massive stealth fields from deformation method

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Based in the collaboration:

Cristian C. Quinzacara, Paola Meza, Almeira Sampson & M.V. Eur. Phys. J. C (2018) 78. [arXiv:1805.04621]

1 Introduction

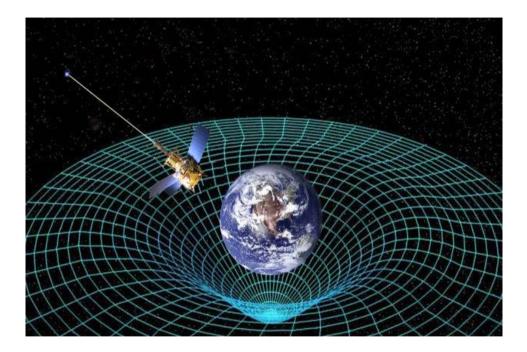
It is generally believed that matter curves the space since given an action principle of pure gravity and matter,

$$S[g,\phi] = S_G[g] + S_M[g,\phi],$$
(1.1)

it happens that the equations of motion of the metric tensor g,

$$\frac{\delta S[g,\phi]}{\delta g^{\mu\nu}} = \sqrt{-g} \left(\underbrace{G_{\mu\nu}[g]}_{\text{Einstein tensor}} - \underbrace{T_{\mu\nu}[g,\phi]}_{\text{Hilbert EM tensor}} \right) = 0.$$
(1.2)

Hence matter curves the spacetime...



However, there are many examples in the literature where this is not the case: Alvarez et al. (2016, 2017), Ayon-Beato et al. (2005, 2006), Ayón-Beato et al. (2013, 2015, 2018), Hassaine (2014), Smolić (2018).

Here the energy momentum tensor vanishes for non-trivial solutions of the matter field equations. In this case, the geometry is not curved in the presence of non-trivial matter fields configuration. These configurations are dubbed *Stealth Fields*.

The purpose of this paper is to **present a method to construct models with massive stealth fields** in arbitrary backgrounds.



The scalar field action is "deformed" into a new action with deformation parameter θ .

The deformed action possesses:

- a massive stealth mode of mass θ^{-1}
- other modes with rescaled effects on the gravity background

2 A toy model

Let f(x) be any function of $x \in \mathbb{R}$ with a saddle point at x = 0,

$$\left. \frac{df}{dx} \right|_{x=0} = 0, \qquad (2.3)$$

and let y(x) another function which possess, for definiteness, two zeros at x = 0 and x = 1,

$$y(0) = 0, \qquad y(1) = 0.$$
 (2.4)

We define the composition of functions F(x) := f(y(x)), which inherits from the parent functions f and y the properties,

$$\left. \frac{dF}{dx} \right|_{x=0} = \left. \frac{dF}{dx} \right|_{x=1} = 0, \qquad (2.5)$$

such that it has two saddle points, at x = 0, 1. It is straightforward to prove this. We can use the chain rule to evaluate dF/dx at x = 0, 1,

$$\left. \frac{dF}{dx} \right|_{x=0,1} = \left. \left(\frac{df(y)}{dy} \frac{dy(x)}{dx} \right) \right|_{x=0,1} = \left. \frac{df(y)}{dy} \right|_{y=0} \left. \frac{dy(x)}{dx} \right|_{x=0,1}$$

Here the $df(y)/dy|_{y=0}$ vanishes because from (2.4) y takes zero-value and from (2.3) the derivative of f vanishes when the argument is zero.

Hence from an arbitrary function f(x) with saddle point at x = 0 (2.3) we can construct another arbitrary function F(x) with saddle points at the kernel of the map $y : \mathbb{R} \to \mathbb{R}$, in this example x = 0, 1.

3 θ -deformation of scalar field theories

We shall apply now the same logic in the language of functional calculus, with the dictionary:

$$x \to \phi, \qquad y(x) \to \phi^{\theta}[g,\phi], \qquad f(x) \to S_M[g,\phi], \qquad F(f(x)) \to S_M[g,\phi^{\theta}[g,\phi]].$$

Here the scalar field ϕ is the analogous of x, ϕ^{θ} is a functional map from ϕ , $S_M[g, \phi]$ is the action principle of a scalar field (with saddle points) analogous of f, and $S_M[g, \phi^{\theta}[g, \phi]]$ is a new action principle obtained from a composition of the functionals S_M and ϕ^{θ} analogous of F(f(x)). The field transformation to be considered is:

$$\phi^{\theta}[g,\phi] = (1-\theta^2 \Box)\phi, \qquad (3.6)$$

where θ is a real-valued parameter. Here

$$\Box \phi := \frac{1}{\sqrt{(-g)}} \partial_{\mu} (\sqrt{-g} g^{\mu\nu} \partial_{\nu} \phi) \,,$$

is the Laplace-Beltrami operator acting upon ϕ .

Therefore the kernel of the deformation map $\phi \to \phi^{\theta}[g, \phi]$ consists of the trivial vacuum $\phi = 0$ and the massive configuration $\phi = \phi_m$ with mass $m = \theta^{-1}$:

$$\phi^{\theta}[g,\phi_m] = (1-\theta^2 \Box)\phi_m = 0, \qquad m = \theta^{-1}, \qquad (3.7)$$

which is equivalent to the Klein-Gordon equation in curved space.

4 Scalar field action deformation

We shall consider theories of the type,

$$S[g,\phi] = S_G[g] + S_M[g,\phi].$$
(4.8)

$$\delta\phi \longrightarrow \Upsilon[g,\phi] := \frac{\delta S_M[g,\phi]}{\delta\phi} = \frac{\delta S[g,\phi]}{\delta\phi} = 0,$$
 (4.9)

Hence the E.o.M are:

$$G_{\mu\nu}[g] - T_{\mu\nu}[g,\phi] = 0, \qquad \Upsilon[g,\phi] = 0, \tag{4.10}$$

Now we produce a "deformed" field theory,

$$S^{\theta}[g,\phi] = S_G[g] + S_M[g,\phi^{\theta}], \qquad \phi^{\theta}[g,\phi] := (1-\theta^2 \Box)\phi \qquad (4.11)$$

with E.o.Ms:

$$G_{\mu\nu}[g] - \widetilde{T}_{\mu\nu}[g,\phi] = 0, \qquad \widetilde{\Upsilon}[g,\phi] = 0, \qquad (4.12)$$

We claim that:

$$\widetilde{\Upsilon}[g,\phi] = 0, \qquad G_{\mu\nu}[g,\phi] = 0, \qquad \widetilde{T}_{\mu\nu}[g] = 0, \qquad (4.13)$$

for the massive mode of mass $m = \theta^{-1}$.

Indeed, in reference Cristian C. Quinzacara, Paola Meza, Almeira Sampson & M. V. Eur. Phys. J. C (2018) 78. [arXiv:1805.04621] we proved that the following statements are true:

- Let $\phi = 0$ (trivial vacuum) be a saddle point of $S_M[g, \phi]$ (i.e. a solution of the equations of motion), then the deformed action $S_M[g, \phi^{\theta}]$ has a saddle point at $\phi^{\theta} = 0$, i.e. for massive ϕ of mass $m = \theta^{-1}$.
- The energy momentum tensor of the massive ϕ ($m = \theta^{-1}$) vanishes (ϕ is a stealth field), hence it does not curve the spacetime.

The proof is very general, and it uses the functional generalization of the chain rule,

$$\frac{\delta F[G[f]]}{\delta f(y)} = \int d^D z \frac{\delta F[G[f]]}{\delta G[f](z)} \frac{\delta G[f](z)}{\delta f(y)}, \qquad (4.14)$$

to be confronted with our toy model.

5 Example 1:

The simplest example we can imagine is:

$$\widetilde{S}[g,\phi] = -\frac{M^2}{2} \int d^D x \sqrt{-g} \,\phi^2 \,, \qquad (5.15)$$

The equation of motion for the scalar field is, $\phi = 0$, and its energy momentum tensor vanishes for this solution.

The deformed action is:

$$\widetilde{S}[g,\phi^{\theta}] = -\frac{M^2}{2} \int d^D x \sqrt{-g} (\phi^{\theta}[g,\phi_m])^2 = -\frac{M^2}{2} \int d^D x \sqrt{-g} (\phi^2 - 2\theta^2 \phi \Box \phi + \theta^4 (\Box \phi)^2),$$
(5.16)
where we replaced the original field ϕ by ϕ^{θ} .

This action can be regarded as a degenerated (single-parameter) case of the twoparametric fourth-order action principle analyzed in Hawking and Hertog (2002). The equation of motion for ϕ yields:

$$(1 - \theta^2 \Box)^2 \phi = 0.$$
 (5.17)

Clearly satisfied by massive ϕ of mass θ^{-1} . A direct calculation of the energy momentum tensor produces:

$$\widetilde{T}_{\mu\nu}[g,\phi] = -\frac{1}{4}M^2 g_{\mu\nu} \left((1-\theta^2 \Box)\phi \right)^2 - \frac{1}{2}M^2 \theta^2 g_{\mu\nu} \Box \phi (1-\theta^2 \Box)\phi + \frac{1}{2}M^2 \theta^2 \left(\delta^{\rho}_{\mu} \delta^{\sigma}_{\nu} + \delta^{\rho}_{\nu} \delta^{\sigma}_{\mu} - g_{\mu\nu} g^{\rho\sigma} \right) \nabla_{\rho} \phi \nabla_{\sigma} \left((1-\theta^2 \Box)\phi \right), \quad (5.18)$$

which is evidently zero for the massive $\phi = \phi_m$.

6 Example 2: Deformation of the massive field action

In the case of the scalar field ϕ , with mass M the matter action principle is given by:

$$S_M[g,\phi] = -\frac{1}{2} \int d^D x \sqrt{-g} \left(\nabla^\mu \phi \nabla_\mu \phi + M^2 \phi^2 \right).$$
(6.19)

With E.o.M:

$$\left(\Box - M^2\right)\phi = 0. \tag{6.20}$$

The deformed action principle is

$$S_M[g,\phi] = -\frac{1}{2} \int d^D x \sqrt{-g} \left(\nabla^\mu \phi^\theta \nabla_\mu \phi^\theta + M^2(\phi^\theta)^2 \right), \qquad \phi^\theta[g,\phi] = (1-\theta^2 \Box)\phi.$$
(6.21)

We obtain the E.o.M (from variation with respect to ϕ):

$$\left(1 - \theta^2 \Box\right)^2 \left(\Box - M^2\right) \phi = 0, \qquad (6.22)$$

which has both solutions, of mass θ^{-1} and of mass M.

The energy momentum tensor is given by:

$$\widetilde{T}_{\mu\nu}[g,\phi] = -\frac{1}{2} \left(\frac{1}{2} g_{\mu\nu} \left(\nabla^{\rho} \phi^{\theta} \nabla_{\rho} \phi^{\theta} + M^{2} (\phi^{\theta})^{2} \right) - \nabla_{\mu} \phi^{\theta} \nabla_{\nu} \phi^{\theta} \right) + \frac{1}{2} \theta^{2} g_{\mu\nu} \Box \phi \left(\Box - M^{2} \right) \phi^{\theta} - \frac{1}{2} \theta^{2} \left(\delta^{\rho}_{\mu} \delta^{\sigma}_{\nu} + \delta^{\rho}_{\nu} \delta^{\sigma}_{\mu} - g_{\mu\nu} g^{\rho\sigma} \right) \nabla_{\rho} \phi \nabla_{\sigma} \left(\Box - M^{2} \right) \phi^{\theta} , \quad (6.23)$$

which vanishes for the field of mass θ^{-1} , because $\phi^{\theta} = (1 - \theta^2 \Box)\phi = 0$.

Hence ϕ of mass θ^{-1} is **stealth**.

Now we analyze the energy-momentum for ϕ_M . First note that:

$$\phi^{\theta}[g,\phi_M] := (1-\theta^2 \Box)\phi_M = \lambda \phi_M, \qquad \lambda := 1 - \frac{M^2}{m^2} = 1 - M^2 \theta^2, \qquad (6.24)$$

Hence the energy-momentum tensor for this solution is (6.23) given by:

$$\widetilde{T}_{\mu\nu}[g,\phi_M] = -\frac{1}{2}\lambda^2 \left(\frac{1}{2}g_{\mu\nu} \left(\nabla^\rho \phi_M \nabla_\rho \phi_M + M^2 \phi_M^2\right) - \nabla_\mu \phi_M \nabla_\nu \phi_M\right) = \lambda^2 T_{\mu\nu}[g,\phi_M],$$
(6.25)

where $T_{\mu\nu}[g, \phi_M]$ is the energy momentum tensor provided by the standard massive field action (6.21).

Therefore, the energy-momentum tensor of the original gravity-matter system is rescaled by factor λ^2 in the deformed theory.

This can be interpreted also as a rescaling of the Newton coupling constant by a factor, $G_{\mathbb{N}} \to \lambda^2 G_{\mathbb{N}}$, in the standard nomenclature.

 \rightarrow Hence the mass of the stealth field (equivalently the deformation parameter) can be used to smooth or amplify the effects of the massive field of mass M on the gravitational background.

7 Overview and remarks

- We can construct a wide class of scalar field action principles in curved space which admits massive stealth configurations.
- The existence of stealth configurations may produce cosmological effects, by means of the energy-momentum tensor rescaling of regular matter fields.

Thank you !

References

- Abigail Alvarez, Cuauhtemoc Campuzano, Miguel Cruz, Efraín Rojas, and Joel Saavedra. Stealths on (1 + 1)-dimensional dilatonic gravity. *Gen. Rel. Grav.*, 48 (12):165, 2016. doi: 10.1007/s10714-016-2158-7.
- Abigail Alvarez, Cuauhtemoc Campuzano, Efraín Rojas, and Joel Saavedra. Gravitational Stealths on dilatonic (1 + 1)-D black hole. In Proceedings, 14th Marcel Grossmann Meeting on Recent Developments in Theoretical and Experimental General Relativity, Astrophysics, and Relativistic Field Theories (MG14) (In 4 Volumes): Rome, Italy, July 12-18, 2015, volume 3, pages 2723-2726, 2017. doi: 10.1142/9789813226609_0335. URL http://inspirehep.net/record/1641468/ files/9789813226609_0335.htm.

- Eloy Ayon-Beato, Alberto Garcia, Alfredo Macias, and Jose M. Perez-Sanchez. Note on scalar fields nonminimally coupled to (2+1) gravity. *Phys. Lett.*, B495:164–168, 2000. doi: 10.1016/S0370-2693(00)01241-7.
- Eloy Ayon-Beato, Cristian Martinez, Ricardo Troncoso, and Jorge Zanelli. Gravitational Cheshire effect: Nonminimally coupled scalar fields may not curve spacetime. *Phys. Rev.*, D71:104037, 2005. doi: 10.1103/PhysRevD.71.104037.
- Eloy Ayon-Beato, Cristian Martinez, and Jorge Zanelli. Stealth scalar field over-flying a (2+1) black hole. *Gen. Rel. Grav.*, 38:145–152, 2006. doi: 10.1007/s10714-005-0213-x.
- Eloy Ayón-Beato, Alberto A. García, P. Isaac Ramírez-Baca, and César A. Terrero-Escalante. Conformal stealth for any standard cosmology. *Phys. Rev.*, D88(6): 063523, 2013. doi: 10.1103/PhysRevD.88.063523.

- Eloy Ayón-Beato, Mokhtar Hassaïne, and María Montserrat Juárez-Aubry. Stealths on Anisotropic Holographic Backgrounds. 2015.
- Eloy Ayón-Beato, P. Isaac Ramírez-Baca, and César A. Terrero-Escalante. Cosmological stealths with nonconformal couplings. *Phys. Rev.*, D97(4):043505, 2018. doi: 10.1103/PhysRevD.97.043505.
- Carl M. Bender and Philip D. Mannheim. No-ghost theorem for the fourth-order derivative Pais-Uhlenbeck oscillator model. *Phys. Rev. Lett.*, 100:110402, 2008. doi: 10.1103/PhysRevLett.100.110402.
- Cuauhtemoc Campuzano, Víctor H. Cárdenas, and Ramón Herrera. Mimicking the LCDM model with Stealths. *Eur. Phys. J.*, C76(12):698, 2016. doi: 10.1140/epjc/s10052-016-4546-2.

- F. J. de Urries and J. Julve. Ostrogradski formalism for higher derivative scalar field theories. J. Phys., A31:6949–6964, 1998. doi: 10.1088/0305-4470/31/33/006.
- Valerio Faraoni and Andres F. Zambrano Moreno. Are stealth scalar fields stable? *Phys. Rev.*, D81:124050, 2010. doi: 10.1103/PhysRevD.81.124050.
- Jack Gegenberg, Cristian Martinez, and Ricardo Troncoso. A Finite action for three-dimensional gravity with a minimally coupled scalar field. *Phys. Rev.*, D67: 084007, 2003. doi: 10.1103/PhysRevD.67.084007.
- Mokhtar Hassaine. Analogies between self-duality and stealth matter source. J. Phys., A39:8675–8680, 2006. doi: 10.1088/0305-4470/39/27/008.
- Mokhtar Hassaine. Rotating AdS black hole stealth solution in D=3 dimensions. *Phys. Rev.*, D89(4):044009, 2014. doi: 10.1103/PhysRevD.89.044009.

- S. W. Hawking and Thomas Hertog. Living with ghosts. Phys. Rev., D65:103515, 2002. doi: 10.1103/PhysRevD.65.103515.
- Marc Henneaux, Cristian Martinez, Ricardo Troncoso, and Jorge Zanelli. Black holes and asymptotics of 2+1 gravity coupled to a scalar field. *Phys. Rev.*, D65: 104007, 2002. doi: 10.1103/PhysRevD.65.104007.
- Hideki Maeda and Kei-ichi Maeda. Creation of the universe with a stealth scalar field. *Phys. Rev.*, D86:124045, 2012. doi: 10.1103/PhysRevD.86.124045.
- Cristian Martinez, Juan Pablo Staforelli, and Ricardo Troncoso. Topological black holes dressed with a conformally coupled scalar field and electric charge. *Phys. Rev.*, D74:044028, 2006. doi: 10.1103/PhysRevD.74.044028.
- Julio Oliva and Mauricio Valenzuela. Topological self-dual vacua of deformed gauge theories. *JHEP*, 09:152, 2014. doi: 10.1007/JHEP09(2014)152.

- Ivica Smolić. Spacetimes dressed with stealth electromagnetic fields. Phys. Rev., D97(8):084041, 2018. doi: 10.1103/PhysRevD.97.084041.
- P. K. Townsend, K. Pilch, and P. van Nieuwenhuizen. Selfduality in Odd Dimensions. *Phys. Lett.*, 136B:38, 1984. doi: 10.1016/0370-2693(84)91753-2,10.1016/0370-2693(84)92051-3. [Addendum: Phys. Lett.137B,443(1984)].