

Delta Gravity: A possible explanation to *Dark Energy* from modified General Relativity

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Standard Cosmological Model: Λ CDM

- ▶ Based on General Relativity:

$$S = \int \left[\frac{1}{2\kappa} (R - 2\Lambda) + \mathcal{L}_M \right] \sqrt{-g} d^4x$$

- ▶ Einstein Field Equations:



$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R + \Lambda g_{\mu\nu} = \frac{8\pi G}{c^4}T_{\mu\nu}$$

Standard Cosmological Model: Λ CDM

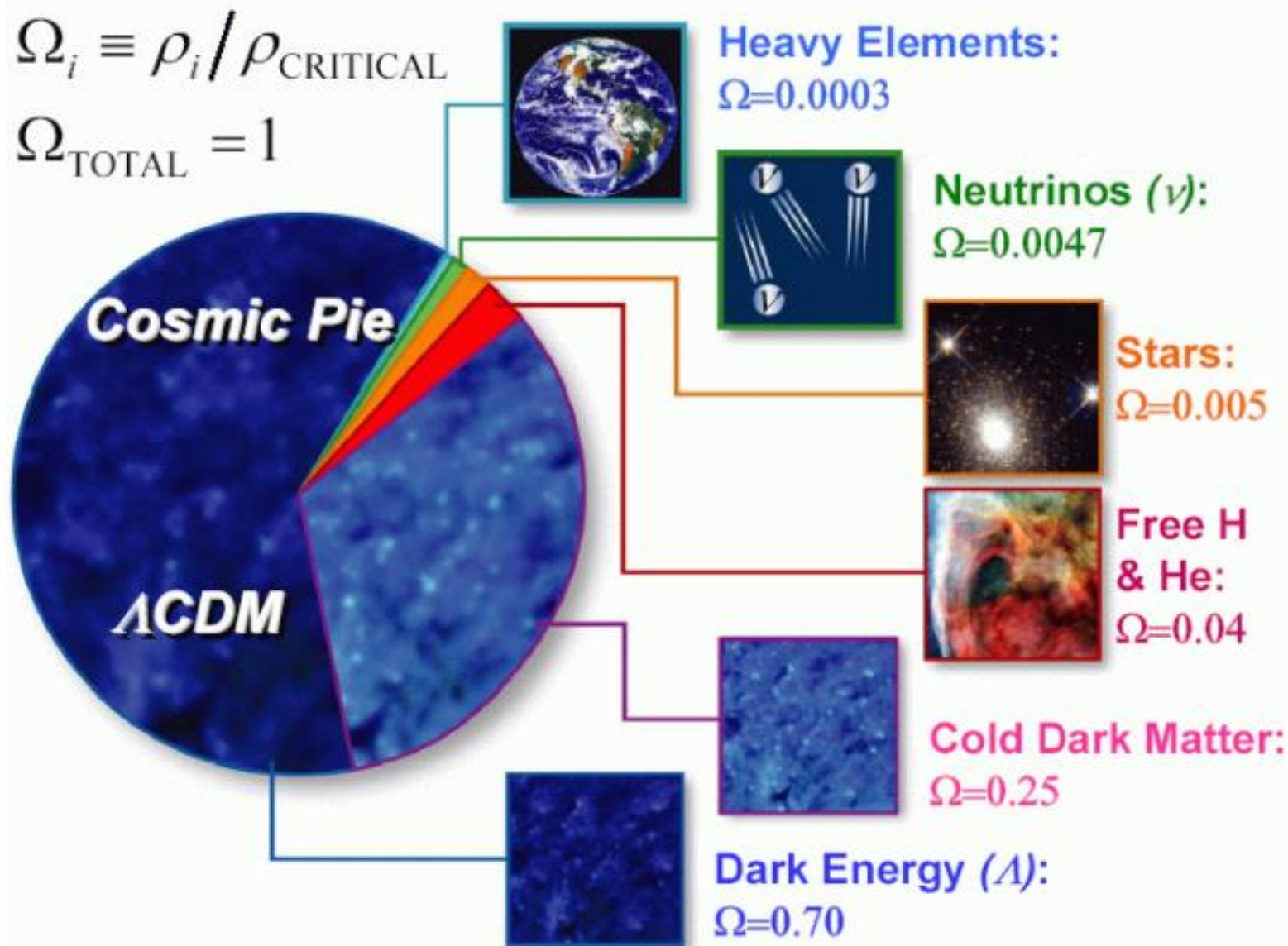
- Friedmann equations:

$$H^2 = \left(\frac{\dot{a}}{a} \right)^2 = \frac{8\pi G\rho}{3} - \frac{k}{a^2} + \frac{\Lambda}{3}$$

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} (\rho + 3p) + \frac{\Lambda}{3}$$

Standard cosmological model: Λ CDM

- ▶ Λ + CDM
 - ▶ Dark Energy
 - ▶ Dark Matter
- ▶ Observational evidence?



How well does it work?

- ▶ Hubble constant from SN-Ia vs Planck
- ▶ Is this correct?

The Last H_0 measurement, 2018

NEW PARALLAXES OF GALACTIC CEPHEIDS FROM SPATIALLY SCANNING THE
HUBBLE SPACE TELESCOPE:
IMPLICATIONS FOR THE HUBBLE CONSTANT

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JOHN W. MACKENTY,¹ J. BRADLEY BOWERS,² KELSEY I. CLUBB,⁵ ALEXEI V. FILIPPENKO,^{5,6}
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73.45 ± 1.66 km / (Mpc s)

Parallax measurements of galactic cepheids; the value suggests a discrepancy with CMB measurements at the 3.7σ level. The uncertainty is expected to be reduced to below 1% with the final release of the Gaia catalog.

From CMB data, Planck 2015

Table 4. Parameter 68 % confidence limits for the base Λ CDM model from *Planck* CMB power spectra, in combination with lensing reconstruction (“lensing”) and external data (“ext”, BAO+JLA+ H_0). While we see no evidence that systematic effects in polarization are biasing parameters in the base Λ CDM model, a conservative choice would be to use the parameter values listed in Column 3 (i.e., for TT+lowP+lensing). Nuisance parameters are not listed here for brevity, but can be found in the extensive tables on the Planck Legacy Archive, <http://pla.esac.esa.int/pla>; however, the last three parameters listed here give a summary measure of the total foreground amplitude (in μK^2) at $\ell = 2000$ for the three high- ℓ temperature power spectra used by the likelihood. In all cases the helium mass fraction used is predicted by BBN from the baryon abundance (posterior mean $Y_P \approx 0.2453$, with theoretical uncertainties in the BBN predictions dominating over the *Planck* error on $\Omega_b h^2$). The Hubble constant is given in units of $\text{km s}^{-1} \text{Mpc}^{-1}$, while r_* is in Mpc and wavenumbers are in

error: 3.7σ

Parameter	TT+lowP 68 % limits	TT+lowP+lensing 68 % limits	TT+lowP 68 % limits	TT+lowP+lensing+ext 68 % limits
H_0	67.31 ± 0.96	67.81 ± 0.92	67.90 ± 0.55	67.74 ± 0.46
Ω_Λ	0.685 ± 0.013	0.692 ± 0.012	0.6935 ± 0.0072	0.6911 ± 0.0062
Ω_m	0.315 ± 0.013	0.308 ± 0.012	0.3065 ± 0.0072	0.3089 ± 0.0062
$\Omega_m h^2$	0.1426 ± 0.0020	0.1415 ± 0.0019	0.1413 ± 0.0011	0.14170 ± 0.00097
$\Omega_m h^3$	0.09597 ± 0.00045	0.09591 ± 0.00045	0.09593 ± 0.00045	0.09598 ± 0.00029
σ_8	0.829 ± 0.014	0.8149 ± 0.0093	0.8154 ± 0.0090	0.8159 ± 0.0086
r_*	170.2 ± 0.5	170.2 ± 0.5	170.2 ± 0.5	170.2 ± 0.5
τ	0.089 ± 0.004	0.089 ± 0.004	0.089 ± 0.004	0.089 ± 0.004
n_s	0.9655 ± 0.0062	0.9677 ± 0.0060	0.9681 ± 0.0060	0.9681 ± 0.0060
$\ln 10^{10} A_s$	3.091 ± 0.013	3.091 ± 0.013	3.091 ± 0.013	3.091 ± 0.013
$\sum \mu_i$	0.02227 ± 0.00020	0.02225 ± 0.00016	0.02226 ± 0.00016	0.02230 ± 0.00014
$\sum \nu_i$	0.02227 ± 0.00020	0.02225 ± 0.00016	0.02226 ± 0.00016	0.02230 ± 0.00014
$\sum \xi_i$	0.02227 ± 0.00020	0.02225 ± 0.00016	0.02226 ± 0.00016	0.02230 ± 0.00014

Planck 2015
 $67.81 \pm 0.92 \text{ Mpc}/(\text{Km s})$

SN-Ia 2018
 $73.45 \pm 1.66 \text{ Mpc}/(\text{Km s})$

Delta Gravity

- ▶ We obtain the new action by extension of the **new symmetry** in EH action

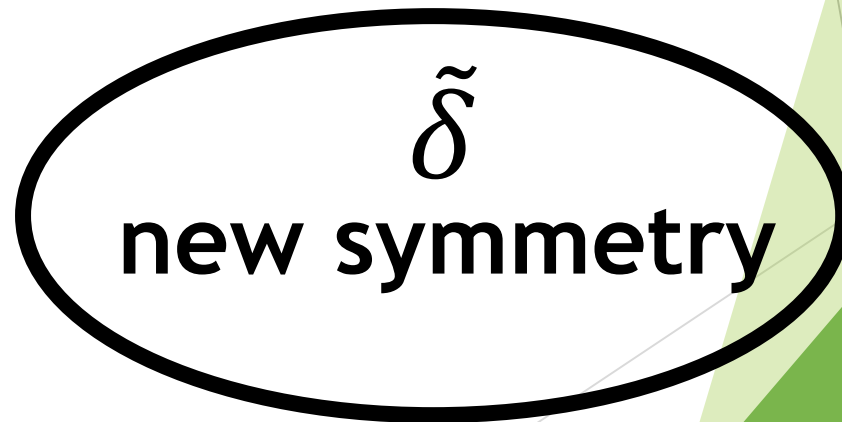
$$S_0 = \int d^4x \sqrt{-g} \left(\frac{R}{2\kappa} + L_M \right) \longrightarrow S = \int d^4x \sqrt{-g} \left(\frac{R}{2\kappa} + L_M - \frac{\kappa_2}{2\kappa} \left(G^{\alpha\beta} - \kappa T^{\alpha\beta} \right) \tilde{g}_{\alpha\beta} + \kappa_2 \tilde{L}_M \right)$$

- ▶ Delta Gauge Theories (extension by the new symmetry).

- ▶ $g_{uv} \rightarrow \tilde{g}_{uv}$

- ▶ We have two new kind of fields:

- ▶ Delta Matter
- ▶ Delta Radiation



Equations of motion:

$$\delta [G_{\mu\nu}] \Rightarrow F^{(\mu\nu)(\alpha\beta)\rho\lambda} D_\rho D_\lambda \tilde{g}_{\alpha\beta} + \frac{1}{2} g^{\mu\nu} R^{\alpha\beta} \tilde{g}_{\alpha\beta} - \frac{1}{2} \tilde{g}^{\mu\nu} R = \kappa \tilde{T}^{\mu\nu}.$$
$$G^{\mu\nu} = \kappa T^{\mu\nu}$$

Conservation rules:

$$\delta (D_\nu T^{\mu\nu}) = 0. \Rightarrow D_\nu \tilde{T}^{\mu\nu} = \frac{1}{2} T^{\alpha\beta} D^\mu \tilde{g}_{\alpha\beta} - \frac{1}{2} T^{\mu\beta} D_\beta \tilde{g}_\alpha^\alpha + D_\beta (\tilde{g}_\alpha^\beta T^{\alpha\mu}).$$
$$D_\nu T^{\mu\nu} = 0$$

Some useful equations

- ▶ The effective metric is:

$$g_{\mu\nu}dx^\mu dx^\nu = -c^2 dt^2 + R^2(t) (dx^2 + dy^2 + dz^2),$$

$$\tilde{g}_{\mu\nu}dx^\mu dx^\nu = -3F_a(t)c^2 dt^2 + F_a(t)R^2(t) (dx^2 + dy^2 + dz^2).$$

$$\mathbf{g}_{\mu\nu} = g_{\mu\nu} + \tilde{g}_{\mu\nu} = -c^2(1 + 3 F_a(t))dt^2 + (1 + F_a(t))R^2(t) (dx^2 + dy^2 + dz^2)$$

- ▶ The normalized effective scale factor :

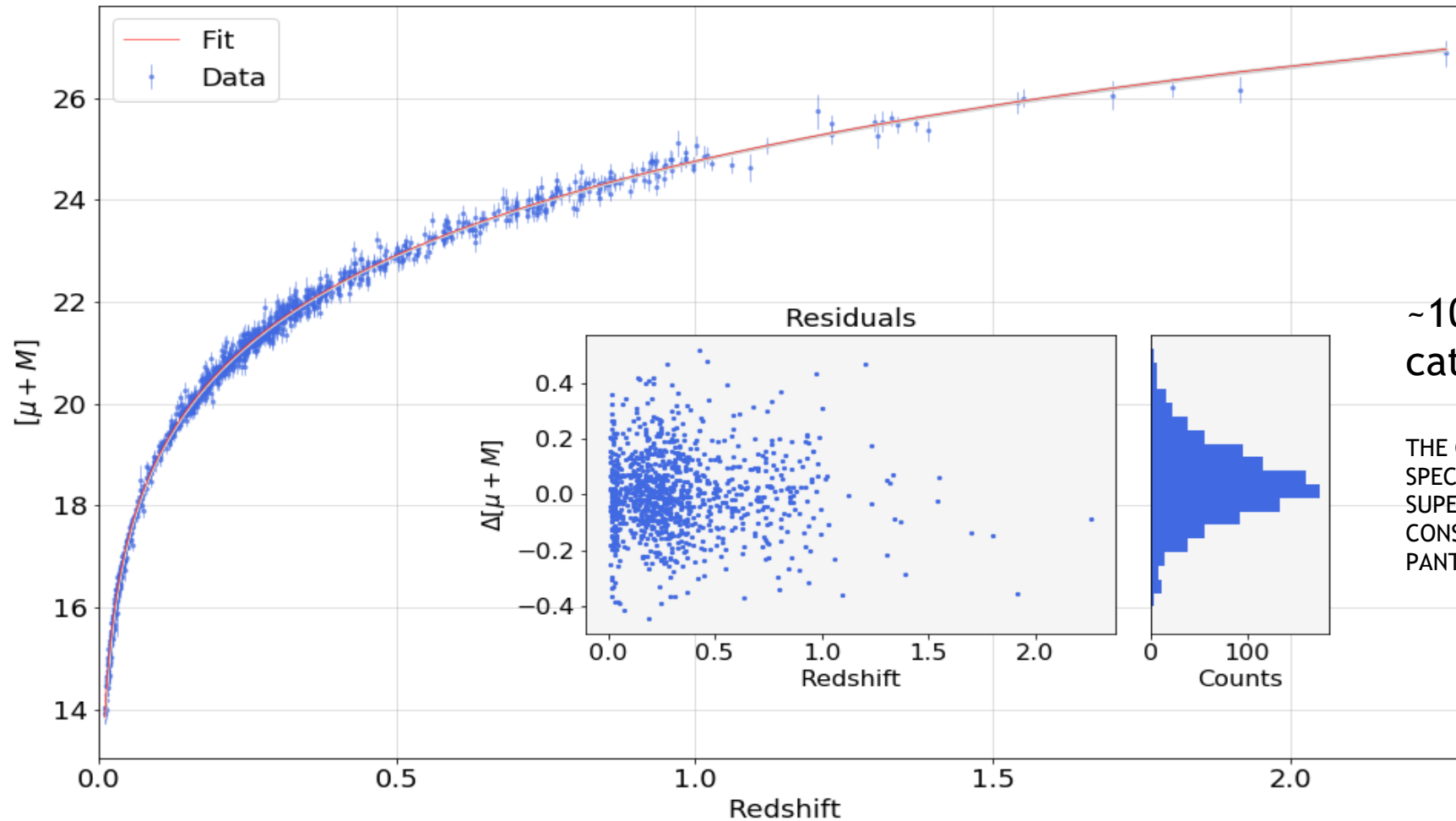
$$Y_{DG}(L_2, C, Y) = \frac{Y}{R_{DG}(t_0)} \sqrt{\frac{1 - L_2 \frac{Y}{3} \sqrt{Y + C}}{1 - L_2 Y \sqrt{Y + C}}}$$

- ▶ To fit SN data:

$$\mu = m - M = 5 \log_{10} \left(\frac{d}{10 \text{ pc}} \right) \quad d_L(z, L_2, C) = c \frac{(1+z)\sqrt{C}}{100\sqrt{h^2\Omega_{r0}}} \int_{Y(t_1)}^1 \frac{Y}{\sqrt{Y+C}} \frac{dY}{Y_{DG}(t)}$$

¿Why Delta Gravity?

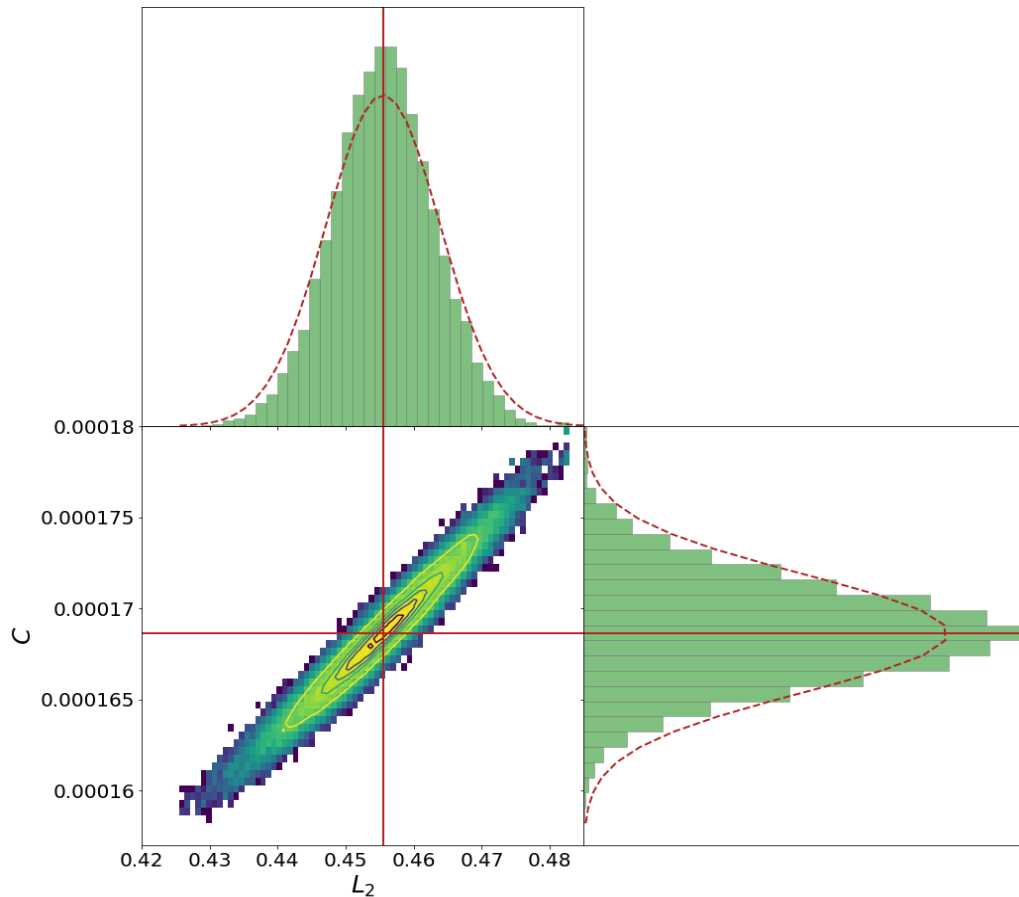
► **DG PREDICTS ACCELERATING UNIVERSE WITHOUT DARK ENERGY!!!!!!**



~1000 SN data. Most updated catalog, 2018 (Scolnic & Jones)

THE COMPLETE LIGHT-CURVE SAMPLE OF SPECTROSCOPICALLY CONFIRMED TYPE IA SUPERNOVAE FROM PAN-STARRS1 AND COSMOLOGICAL CONSTRAINTS FROM THE COMBINED PANTHEON SAMPLE

Results from MCMC (Markov Chain Monte Carlo):



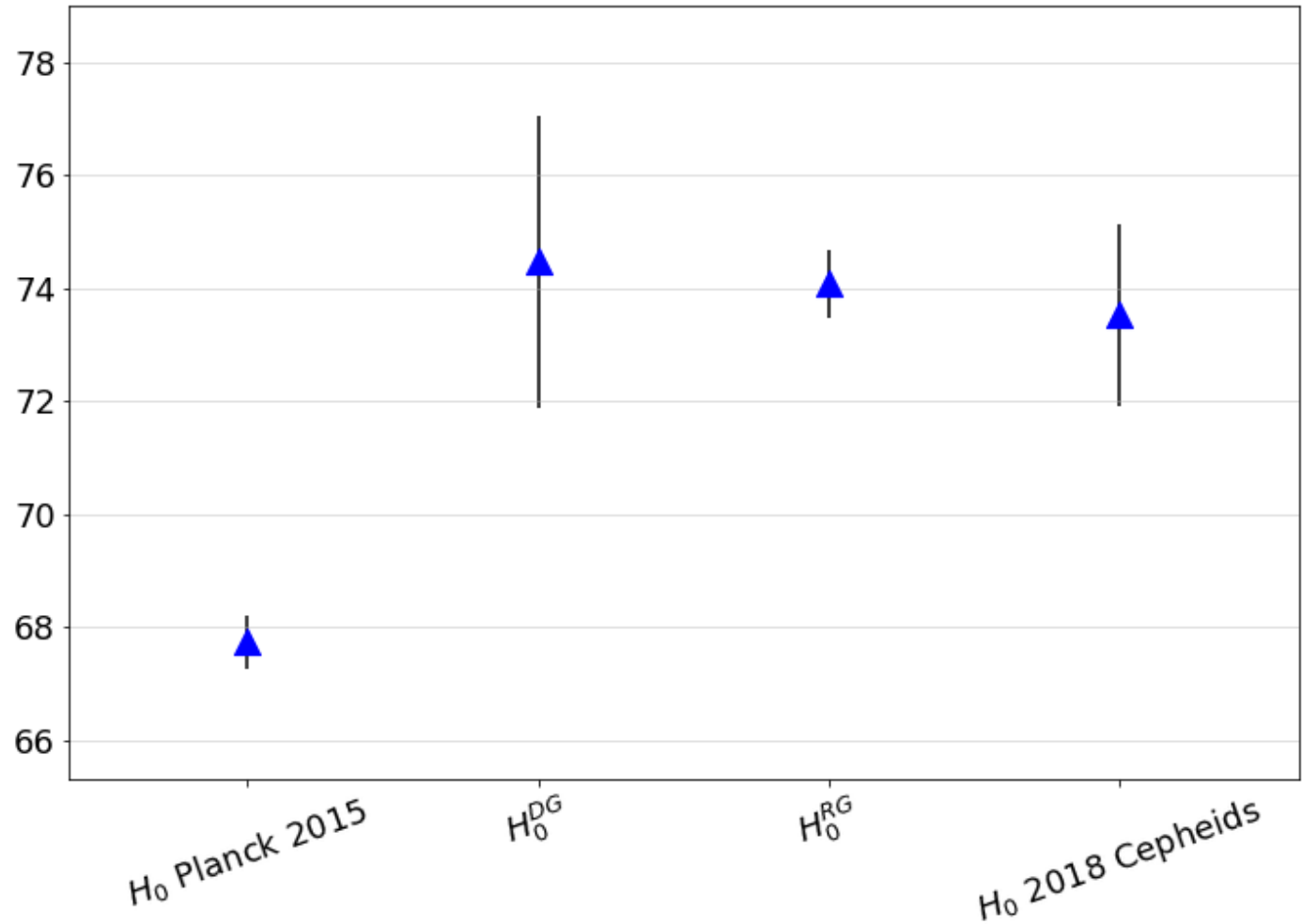
WITH THE TWO FREE
PARAMETERS FITTED:

L_2 and C

WE CAN PREDICT COSMOLOGICAL
INTERESTING VALUES, LIKE H_0

Good H_0 Value in Delta Gravity

- ▶ H_0 is higher than Λ CDM with Planck data
- ▶ It is in concordance with the last independent-model measurement
- ▶ The acceleration is naturally produced from the geometry of the Model, and it is not imposed using a constant, like Λ



This is not the end...

- ▶ We are working on obtain the CMB Power Spectrum
- ▶ We want to fit the CMB and compare with SN-Ia data

Thanks!

<https://arxiv.org/abs/1704.02888>

$\tilde{\delta}$ Gravity, $\tilde{\delta}$ matter and the accelerated expansion of the Universe

Jorge Alfaro, Pablo González

<https://arxiv.org/abs/1811.05828>

An accelerating Universe without Λ in concordance with the last H_0 measured value

Jorge Alfaro, Marco San Martin, Joaquin Sureda