Traversable Wormholes in AdS

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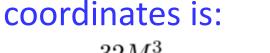
I. Introduction

Wormhole is a vacuum solution of Einstein's equation

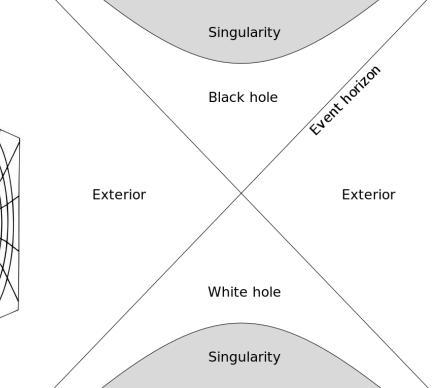
$$G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 0$$

On a time-slice, two interiors of black holes are connected via a tube or "throat" called Einstein-Rosen bridge.

The Maximally Extended Schwarzschild Solution in Kruskal

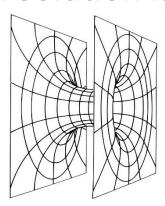


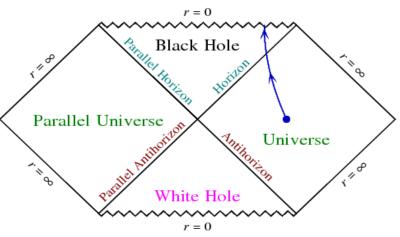
$$ds^{2} = -\frac{32M^{3}}{r}e^{-\frac{r}{2M}}dUdV + r^{2}d\Omega_{2}^{2}$$



The Penrose diagram of the Maximally Extended

Schwarzschild Solution is



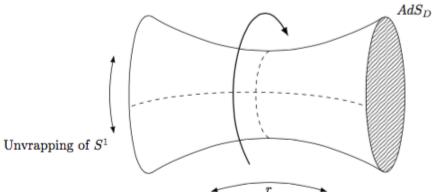


Anti-de Sitter space is a solution of Einstein's equation with negative cosmological constantan

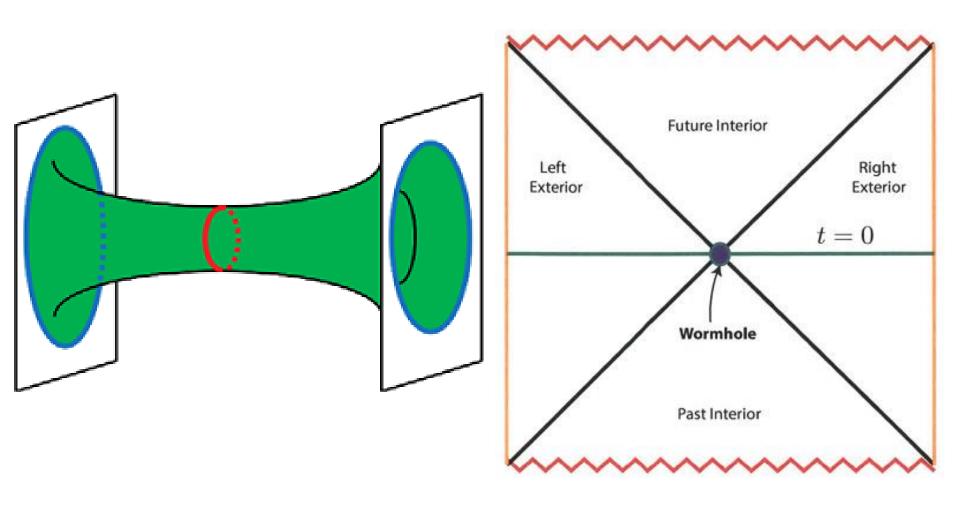
$$G_{\mu\nu} + \Lambda g_{\mu\nu} = 0$$

$$\mathbf{ds}^2 = (1 + r^2/a^2) \, dT^2 - (1 + r^2/a^2)^{-1} \, dr^2 - r^2 (d\theta^2 + \sin^2\theta \, d\phi^2)$$

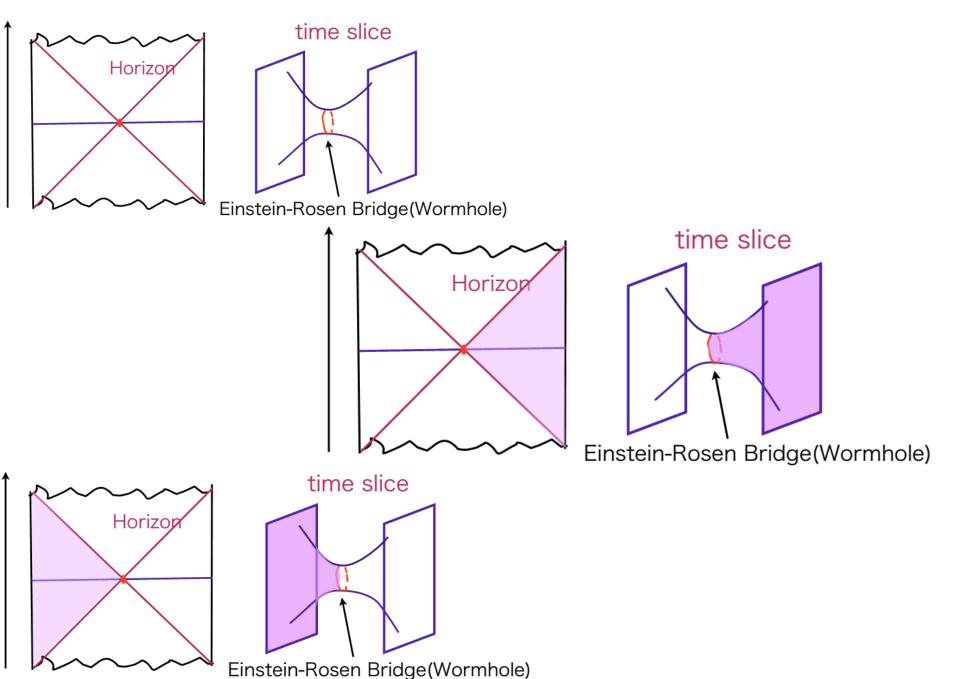
$$3/\Delta = -a^2$$



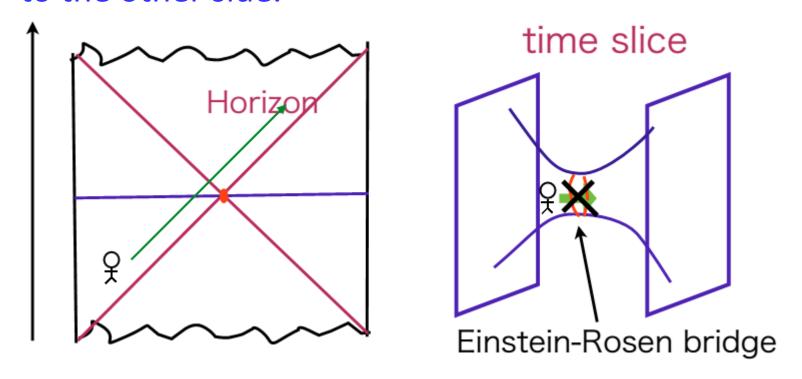
The Penrose diagram of the eternal Anti-de Sitter space in Kruskal coordinates is:



Borrowed from T. Numasawa



Since the Einstein-Rosen bridge is instable an we can not go to the other side.



Traversable wormhole: usually require an exotic matter which violates the Average Null Energy Condition (ANEC).

The ANEC states that along a given infinite null geodesic, the integral of the null component of the stress energy is nonnegative:

$$\int_{-\infty}^{\infty} T_{\mu\nu} k^{\mu} k^{\nu} d\lambda \ge 0$$

 $T_{\mu
u}$ -is the expectation value of the renormalized stress-energy tensor k^{μ} - vector pointing along the direction of the null geodesic

- ANEC in GR ensures that gravity is an attractive force
- ☐ In quantum level ANEC can be violated.

 λ -is an affine parameter along the null geodesic

☐ For all traversable wormhole it is necessary the violation of ANEC. Morris, Thorne, Yurtsever 88

II. Traversable wormhole via double trace deformation Gao-Jefferis-Wall, 16

The Thermofield Double State (TFD)

The thermofield double state is an entangled state in the tensor product of two identical "left" and "right" CFTs:

$$|\Psi_{TFD}\rangle = \frac{1}{\sqrt{Z(\beta)}} \sum_{n} e^{-\frac{\beta E_{n}}{2}} |E_{n}\rangle_{L} \otimes |E_{n}\rangle_{R}$$

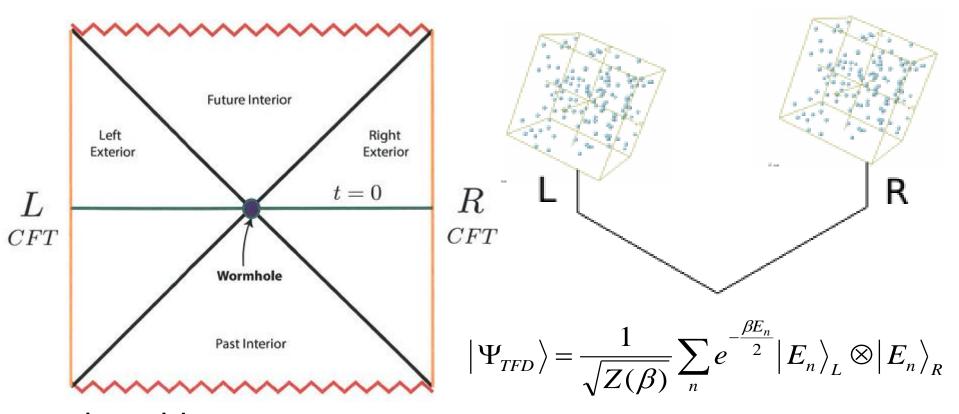
where β is the inverse temperature

sum is over energy eigenstates.

where Z is the partition function

Assuming the AdS/CFT correspondence TFD is holographically dual to an eternal two-sided AdS black hole.

Start with the eternal AdS black hole, dual to the thermofield double state in the decoupled product of two



Israel, Maldacena

Gao, Jafferis and Wall showed that a wormhole in the eternal AdS black hole scenario can be made traversable by turning on a coupling between the left and right boundaries of form

$$\delta H(t_1) = -\int d^{d-1}x_1 h(t_1, x_1)\mathcal{O}_R(t_1, x_1)\mathcal{O}_L(-t_1, x_1)$$

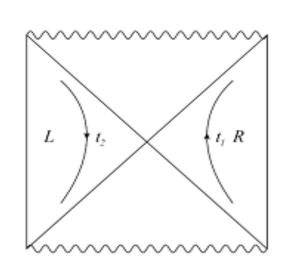
$$h(t_1, x_1) = \begin{cases} h\kappa^{2-2\Delta}, & t_0 \leqslant t_1 \leqslant t_f \\ 0, & \text{otherwise.} \end{cases}$$
 is a small deformation.

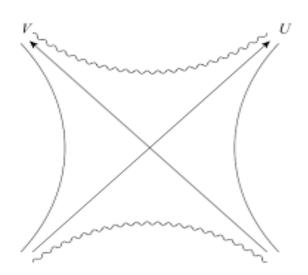
 $\mathcal{O}_{L/R}$ is a scalar operator of dimension $\Delta = \frac{d}{2} - \sqrt{\left(\frac{d}{2}\right)^2 + m^2}$ living in the Left/Right CFT, dual to a bulk scalar field with mass m.

Bañados, Teitelboim Zanalli 92

$$ds^2 = -\frac{r^2 - r_h^2}{\ell^2} dt^2 + \frac{\ell^2}{r^2 - r_h^2} dr^2 + r^2 d\phi^2, \quad \phi \sim \phi + 2\pi$$

Kruskal coordinates: $e^{2r_ht} = -\frac{U}{V}$, $\frac{r}{r_h} = \frac{1-UV}{1+UV}$





The bulk stress tensor associated to the scalar field is

$$T_{\mu\nu} = \partial_{\mu}\Phi\partial_{\nu}\Phi - \frac{1}{2}g_{\mu\nu}g^{\rho\sigma}\partial_{\rho}\Phi\partial_{\sigma}\Phi - \frac{1}{2}g_{\mu\nu}m^{2}\Phi^{2}.$$

They considered correlation function

$$\langle \varphi_R^H(t, r, \phi) \varphi_R^H(t', r', \phi') \rangle$$

First order in contribution in h

$$G_h = i \int_{t_0}^t dt_1 h(t_1) K_{\Delta}(t' + t_1 - i\beta/2) [K_{\Delta}(t - t_1 - i\epsilon) - K_{\Delta}(t - t_1 + i\epsilon)] + (t \leftrightarrow t')$$

Bulk to boundary propagator

$$K_{\Delta}(r, t, \phi; 0, 0) = \frac{r_h^{\Delta}}{2^{\Delta+1}\pi} \left(-\frac{(r^2 - r_h^2)^{1/2}}{r_h} \cosh r_h t + \frac{r}{r_h} \cosh r_h \phi \right)^{-\Delta}$$

The stress tensor

$$T_{UU} = \lim_{U' \to U} \partial_U \partial_{U'} G_h(U, U').$$

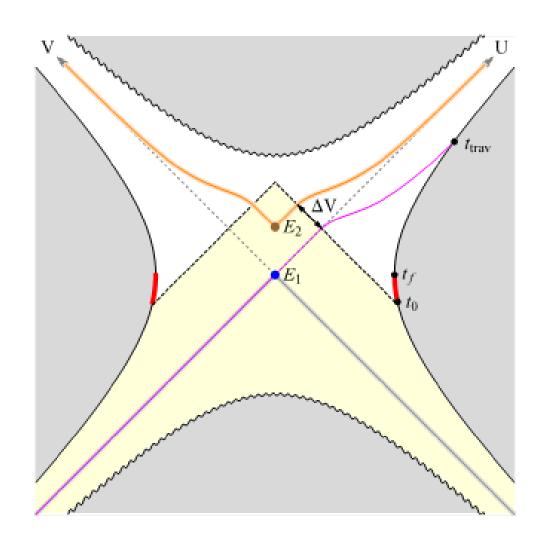
The linearized Einstein equation for the UU component for the fluctuations evaluated at V=0 gives

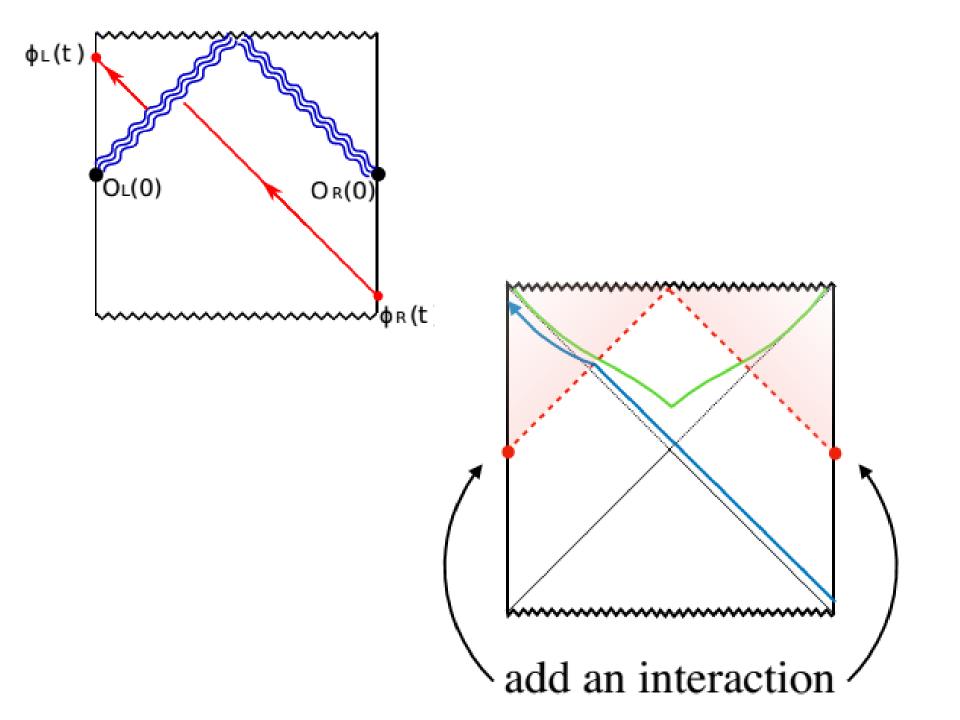
$$\int dU \left(\frac{\kappa}{2r_+} h_{UU} - \frac{r_- \partial_x h_{UU}}{r_+^2} - \frac{\partial_x^2 h_{UU}}{2r_+^2} \right) = 8\pi G_N \int dU T_{UU}$$

The null geodesics at the horizon caused by the interaction is

$$V(U) \sim \int dU T_{UU}$$

$$\Delta V \sim \frac{h \, G_N}{R^{D-2}}$$





III. Final Comments

Rotaing E. Caceres, A. Misobuchi

Charged

In collaboration with: E. Caceres, C. Rivera and A. Misobuchi

$$\left|\Psi_{TFD}\right\rangle = \frac{1}{\sqrt{Z(\beta,\mu)}} \sum_{n} e^{-\frac{\beta(E_{n}+\mu Q_{n})}{2}} \left|E_{n},Q_{n}\right\rangle_{L} \otimes \left|E_{n},-Q_{n}\right\rangle_{R}$$

