## Collider Probes of Axion-like Particles

Andrea Thamm CERN

with Martin Bauer and Matthias Neubert

based on arXiv:1610.00009, 1704.08207, 1708.00443 and work in progress



15 May 2018 Grenoble

## Outline

- Motivation
- ALPs and collider probes
  - Effective Lagrangian
  - Exotic Higgs and ALP decays
  - Probing the ALP parameter space
  - Muon  $(g-2)_{\mu}$
- Conclusions and Outlook



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#### Motivation

- Pseudo-scalars in many extensions of the SM
  - QCD axion solution to strong CP-problem
  - Nambu-Goldstone bosons of a broken symmetry
  - mediators to the dark sector
  - explanations of various anomalies
- Good reason to study them!
- Large regions of parameter space already probed by many different experiments
- We add a region that can be probed through exotic Higgs decays in run 2 of LHC

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• Interactions at dimension-5

[Weinberg: PRL 40 (1978) 223] [Wilczek: PRL 40 (1978) 279] [Georgi, Kaplan, Randall: Phys. Lett.169 B (1986)]

$$\mathcal{L}_{\text{eff}}^{D \leq 5} = \frac{1}{2} \left( \partial_{\mu} a \right) \left( \partial^{\mu} a \right) + \sum_{f} \frac{c_{ff}}{2} \frac{\partial^{\mu} a}{\Lambda} \bar{f} \gamma_{\mu} \gamma_{5} f + g_{s}^{2} C_{GG} \frac{a}{\Lambda} G_{\mu\nu}^{A} \tilde{G}^{\mu\nu,A} + e^{2} C_{\gamma\gamma} \frac{a}{\Lambda} F_{\mu\nu} \tilde{F}^{\mu\nu} + \frac{2e^{2}}{s_{w} c_{w}} C_{\gamma Z} \frac{a}{\Lambda} F_{\mu\nu} \tilde{Z}^{\mu\nu} + \frac{e^{2}}{s_{w}^{2} c_{w}^{2}} C_{ZZ} \frac{a}{\Lambda} Z_{\mu\nu} \tilde{Z}^{\mu\nu}$$

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• Decay into photons, leptons, hadrons

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- Decay into photons, leptons, hadrons
- Higgs interactions at dimension-6 and 7

$$\mathcal{L}_{\text{eff}}^{D \ge 6} = \frac{C_{ah}}{\Lambda^2} \left(\partial_{\mu} a\right) \left(\partial^{\mu} a\right) \phi^{\dagger} \phi + \frac{C_{Zh}^{(7)}}{\Lambda^3} \left(\partial^{\mu} a\right) \left(\phi^{\dagger} i D_{\mu} \phi + \text{h.c.}\right) \phi^{\dagger} \phi + \dots$$

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[Dobrescu, Landsberg, Matchev: 0005308] [Dobrescu, Matchev: 0008192]

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#### Contributions



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• Numerically

 $C_{Zh}^{\text{eff}} \approx C_{Zh}^{(5)} - 0.016 c_{tt} + 0.030 C_{Zh}^{(7)} \left[\frac{1 \text{ TeV}}{\Lambda}\right]^2$ 

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• Decay rate normalised to SM  $\Gamma(h \rightarrow Z\gamma)_{\rm SM} = 6.32 \cdot 10^{-6} {\rm GeV}$ 



• Enhanced rates for this process



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[ATLAS and CMS:1606.02266]

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- From top loop and dim-7:  $Br(h \rightarrow Za) = O(10^{-3})$

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- Interesting final states

$$h \to Za \to Z\gamma\gamma$$

$$\bullet \quad h \to Za \to Zll$$

 $h \to Za \to Z2jets$   $h \to Za \to Z+invisible$ 

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• Dim-6 Higgs portal and loop diagrams

[Dobrescu, Landsberg, Matchev: 0005308] [Dobrescu, Matchev: 0008192] [Chang, Fox, Weiner: 0608310]



$$\begin{aligned} C_{ah}^{\text{eff}} &= C_{ah}(\mu) + \frac{N_c y_t^2}{4\pi^2} c_{tt}^2 \left[ \ln \frac{\mu^2}{m_t^2} - g_1(\tau_{t/h}) \right] - \frac{3\alpha}{2\pi s_w^2} \left( g^2 C_{WW} \right)^2 \left[ \ln \frac{\mu^2}{m_W^2} + \delta_1 - g_2(\tau_{W/h}) \right] \\ &- \frac{3\alpha}{4\pi s_w^2 c_w^2} \left( \frac{g^2}{c_w^2} C_{ZZ} \right)^2 \left[ \ln \frac{\mu^2}{m_Z^2} + \delta_1 - g_2(\tau_{Z/h}) \right] \end{aligned}$$

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 $C_{ah}^{\text{eff}} \approx C_{ah}(\Lambda) + 0.173 c_{tt}^2 - 0.0025 \left( C_{WW}^2 + C_{ZZ}^2 \right)$ 

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$$\Gamma(h \to aa) = \frac{v^2 m_h^3}{32\pi\Lambda^4} \left| C_{ah}^{\text{eff}} \right|^2 \left( 1 - \frac{2m_a^2}{m_h^2} \right)^2 \sqrt{1 - \frac{4m_a^2}{m_h^2}}$$

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- For  $\operatorname{Br}(h \to aa) = 0.1 \operatorname{need} |C_{ah}| / \Lambda^2 \approx 0.62 \operatorname{TeV}^{-2}$
- From top-loop only:  $Br(h \to aa) = 0.01$  for  $|c_{tt}|/\Lambda \approx 1.04 \text{ TeV}^{-1}$

## Exotic Higgs Decays $h \rightarrow aa$

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- Interesting final states
  - $h \to aa \to \gamma \gamma \gamma \gamma$   $h \to aa \to 4 jets$
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#### ALP decays

• Assuming effective Wilson coefficients to be 1





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## Detecting ALPs in $h \to Za$

• Average decay length perpendicular to beam axis

$$L_a^{\perp}(\theta) = \sin \theta \, \frac{\beta_a \gamma_a}{\Gamma_a} = \sin \theta \sqrt{\gamma_a^2 - 1} \, \frac{\operatorname{Br}(a \to X\bar{X})}{\Gamma(a \to X\bar{X})}$$

• Fraction of ALPs decaying before travelling a certain distance

$$f_{\text{det}} = \int_{0}^{\pi/2} d\theta \sin \theta \left( 1 - e^{-L_{\text{det}}/L_{a}^{\perp}(\theta)} \right)$$
  
Decay into photons  
before EM calorimeter  
 $L_{\text{det}} = 1.5 \text{ m}$   
Decay into electrons  
before inner tracker  
 $L_{\text{det}} = 2 \text{ cm}$ 

• Effective branching ratios

 $\operatorname{Br}(h \to Za \to \ell^+ \ell^- X\bar{X})\big|_{\operatorname{eff}} = \operatorname{Br}(h \to Za) \times \operatorname{Br}(a \to X\bar{X}) f_{\operatorname{dec}} \operatorname{Br}(Z \to \ell^+ \ell^-)$ 

• Constraints on ALP mass and coupling to photons



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[Bauer, Neubert, Thamm: 1704.08207]

• Constraints on ALP mass and coupling to photons



• Large hierarchy in couplings can be plausible



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- Large hierarchy in couplings can be plausible
- Integrating out the top





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• Current bounds on  $h \rightarrow Za$ 



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[Bauer, Neubert, Thamm: 1708.00443]

#### • Current bounds on $h \rightarrow aa$



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#### Muon $(g-2)_{\mu}$

• Persistent  $3\sigma$  deviation

[Particle Data Group 2016]

$$a_{\mu}^{\exp} - a_{\mu}^{SM} = (288 \pm 63 \pm 49) \cdot 10^{-11}$$

• ALP can account for discrepancy



[Haber, Kane, Sterling: Nucl. Phys. B 161 (1979)] [Chang, Chang, Chou, Keung: 0009292] [Marciano, Masiero, Paradisi, Passera: 1607.010122]



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#### Conclusions

- Rare Higgs decays provide a powerful way to probe the existence of ALPs with masses between 30 MeV and 60 GeV and couplings suppressed by the 1 - 100 TeV scale
- Connection to low-energy physics probes such as  $(g-2)_{\mu}$

• Reach in  $Z \to \gamma a$ 



• Constraints on ALP mass and coupling to photons



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Combination of cosmological bounds (measurement of Neff, primordial deuterium abundance, modification to BB nucleosynthesis, distortion of CMB, diffuse photon background,...)

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SLAC: measurement of photon - nuclei scattering (Primakoff effect)

Axion helioscopes: Tokyo Axion Helioscope and CERN axion solar telescope

Energy loss of stars constrained by ratio of red giant to younger stars of horizontal branch

Absence of photon burst after supernova

Combination of cosmological bounds (measurement of Neff, primordial deuterium abundance, modification to BB nucleosynthesis, distortion of CMB, diffuse photon background,...)

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## SHiP expected reach

• Fixed target facility at CERN SPS (Search for Hidden Particles)



<sup>[</sup>Alekhin et al.: 1504.04855]