# Axion predictions in SO(10) x U(1)<sub>PQ</sub> models

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In collaboration with...



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#### The aim:

Study the properties of the axion –mass, couplings to photons, fermions, nucleons, domain wall number– in grand unified theories with SO(10) gauge groups, and furnished with a global U(1) symmetry so as to solve the strong CP problem. Account for constraints from unification, proton decay, star cooling, black hole superradiance, fermion mass fits.

#### The novelty:

Properties of SO(10) axion had not been studied systematically

Our formalism bridges the gap between the simple UV SO(10)  $\times$  U(1) symmetries and the low-energy description (e.g. couplings to nucleons, domain wall number).

#### The plan:

- Motivation of SO(10) theories and the axion solution to the strong CP problem
- The guts of the axion solution to the strong CP problem
- The GUTs of the axion solution to the strong CP problem

# Motivation

# Why unification?

Group structure

Matter content in each generation

Anomaly cancellation, charge quantization

Hierarchies of masses and mixing angles

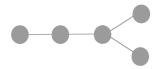
Can the SM model be a low energy effective description of a more predictive theory with a simple gauge group, and fewer representations?

B and L are accidental symmetries in the SM: expect B violation, proton decay!

# Bird's eye view of GUT models

-  $\bullet$   $\cup$  (1) **SM** group is of rank 4: Embed into simple groups of rank 4 or more.

SU(5) Each generation a 5 and  $\overline{10}$ Non-SUSY ruled out by  $\sin^2 \theta_W$  and proton lifetime



**E6** 

SO(10)Each generation a single 16 irrep! (with RH neutrino)Anomaly cancellation automatic

Larger rank allows for **multi-step symmetry breaking** with different chains

A single generation plus Higgses fits in 27 irrep Anomaly cancellation automatic Multi-step breaking, multiple chains

# Why SO(10) x U(1) $_{PQ}$ ?

Minimality and anomaly freedom of matter representations

RH neutrinos automatically included, allowing for leptogenesis

Rich Higgs sector can accommodate axion, inflaton

Multi-step breaking suggests intermediate mass scales which could be tied to leptogenesis and the axion, while playing a role in unification.

Can we build a grand-unified version of SMASH, the model proposed by Ballesteros, Ringwald, Redondo, CT, to account for:

> Predictive inflation Neutrino masses CP problem Dark matter Baryogenesis?

# Why the axion?

Solves the strong CP problem by making the  $\theta$  angle dynamical [Peccei & Quinn, Weinberg, Wilczek].

The axion can be dark matter! [Preskill et all, Abbott and Sikivie, Dine and Fischler]

Need: Spontaneously broken U(1) symmetry, anomalous under QCD: scalar fields giving masses to strongly interacting fermions.

The axion is a combination of scalar phases. Its effective action has a finite symmetry group whose dimension is the domain wall number. Domain walls can overclose the universe it, but if  $N_{DW}$ =1 they can decay. Otherwise one needs to break U(1) (e.g. discrete symmetry)

Axion models classified by

Coupling of axion to gauge bosons (nonderivative, from anomalies)

Axion mass (from QCD effects, fixed by coupling to gluons)

Coupling of axion to fermions, nucleons (derivative)

Domain wall number

# The axion's ID card

$$\mathcal{L}[A]_{\text{eff}} = \frac{1}{2} \partial_{\mu} A \partial^{\mu} A + \partial_{\mu} A \sum_{a} c_{A\psi_{a}} (\psi_{a}^{\dagger} \bar{\sigma}^{\mu} \psi_{a}) + A \sum_{k=1}^{N_{g}} \frac{1}{f_{A,k}} \frac{g_{k}^{2}}{16\pi^{2}} \bar{\text{Tr}} \tilde{F}_{\mu\nu}^{k} F^{k,\mu\nu}$$

- $f_{A,k}$  Coupling of axion to gauge bosons (nonderivative, from anomalies)
- $m_A$  Axion mass (from QCD effects, fixed by coupling to gluons)

$$m_A = 57.0(7) \left(\frac{10^{11} \text{GeV}}{f_{A,3}}\right) \mu \text{eV}.$$
 [Borsanyi et al, di Cortona et al]

 $c_{A\psi_a}$  Coupling of axion to fermions, nucleons (derivative)  $N_{DW}$  Domain wall number

### Axion constraints: dark matter, cooling, superradiance

#### Dark matter:

Oscillating axion field behaves as pressureless dark matter. If U(1) restored after inflation:

 $\Omega_A h^2 \approx 0.12 \Rightarrow 3 \times 10^{10} \,\text{GeV} \lesssim f_A \lesssim 1.2 \times 10^{11} \,\text{GeV} \quad ; \quad 50 \,\mu\text{eV} \lesssim m_A \lesssim 200 \,\mu\text{eV}$ 

#### Star cooling:

Axion coupling to photons allows photon-axion conversion in stars, enhancing energy loss and cooling rate. This gives bounds based on e.g. cooling of helium burning stars in globular clusters, duration of neutrino burst in supernovae explosions [Raffelt,...]

 $g_{aa\gamma} < 0.66 \times 10^{-10}$  [Ayala et al]

#### Superradiance:

An axion scalar field interacting with a rotating black hole has **unstable localized modes** which **extract energy and spin from the black hole**, if the black hole radius is comparable to the scalar's Compton wavelength [Arvanitaki et al]. This would lead to reduced black hole populations for particular values of spin and mass. Leads to

 $f_A < 1 \times 10^{17} \mathrm{GeV}$ 

# What has been done for the axion in GUTs

 $U(1)_{PQ}$  extensions of SO(10) GUTs have been studied before [Lazarides, Kim, Bacj et al, Babu et al, Bertolini et al, Altarelli et al].

A global U(1) can be motivated so as to make the Yukawa sector more predictive It is automatically anomalous and provides an axion solution to the strong CP problem.

The need for intermediate scales in SO(10) unification implies that the axion can be dark matter (otherwise absent in non-SUSY SU(5) GUTs).

Axion field constructed in very few cases

Models have been proposed with  $N_{DW} = 1$ , arguing in terms of the UV symmetries.

# ... and what was missing

A systematic identification of axion field and axion decay constant/mass in relation to thresholds/VEVs in the theory

A systematic calculation of couplings to gauge bosons, fermions/nucleons, including low energy effects

An identification of the global symmetry corresponding to the physical axion, which is orthogonal to the transformations associated with the massive gauge bosons

A direct calculation of domain wall number for the above symmetry

Studies of constraints from unification, proton decay, fermion mass fits, stellar cooling, superradiance.

Which experiments can probe GUT axions?

The guts of the strong CP problem

## Ingredients: a global U(1), anomalous, broken

Weyl fermions  $\psi_a$ 

Nonabelian gauge fields

$$A^{k}_{\mu} = A^{k,a}_{\mu}T^{a}, \ k = 1, \dots, N_{g}$$

Complex scalars  $\phi_j$ 

Global U(1):  $\psi_a \to e^{iq_a\alpha}\psi_a$  $\phi_j \to e^{iq_j\alpha}\phi_j$  Current  $J^{\mu} = \sum_{a} q_{a} \psi^{\dagger}_{a} \bar{\sigma}^{\mu} \psi_{a} + i \sum_{j} q_{j} (\partial_{\mu} \phi^{\dagger}_{j} \phi_{j} - \phi^{\dagger}_{j} \partial_{\mu} \phi_{j}),$ 

Anomaly

$$\begin{aligned} \langle \partial_{\mu} J^{\mu} \rangle &= \sum_{k=1}^{N_g} \frac{g_k^2 \hat{N}_k}{16\pi^2} \,\bar{\mathrm{Tr}} \,\tilde{F}_{\mu\nu}^k F^{k,\mu\nu}, \\ \hat{N}_k &= 2 \sum_a q_a T_k(\rho_a) \end{aligned}$$

Symmetry breaking:

$$\phi_j = \frac{1}{\sqrt{2}} (v_j + \rho_j) e^{iA_j/v_j}.$$

The axion/Goldstone:

$$\begin{split} A_i &= \frac{q_i v_i}{f_{\rm PQ}} A + \text{ orthogonal excitations, } f_{\rm PQ} = \sqrt{\sum_j q_j^2 v_j^2} \\ A &= \frac{1}{f_{\rm PQ}} \sum_i q_i v_i A_i \end{split}$$

### From the anomaly to the effective Lagrangian

$$\langle \partial_{\mu} J^{\mu} \rangle = \sum_{k=1}^{N_g} \frac{g_k^2 \hat{N}_k}{16\pi^2} \,\bar{\mathrm{Tr}} \,\tilde{F}^k_{\mu\nu} F^{k,\mu\nu} = f_{\mathrm{PQ}} \Box A + \sum_a q_a \partial_{\mu} (\psi_a^{\dagger} \bar{\sigma}^{\mu} \psi_a)$$

Equivalent to Euler-Lagrange equations from the following effective action [Srednicki]

$$\mathcal{L}[A]_{\text{eff}} = \frac{1}{2} \partial_{\mu} A \partial^{\mu} A + \partial_{\mu} A \sum_{a} \frac{q_{a}}{f_{\text{PQ}}} (\psi_{a}^{\dagger} \bar{\sigma}^{\mu} \psi_{a}) + A \sum_{k=1}^{N_{g}} \frac{1}{f_{A,k}} \frac{g_{k}^{2}}{16\pi^{2}} \,\bar{\text{Tr}} \,\tilde{F}_{\mu\nu}^{k} F^{k,\mu\nu}$$
$$f_{A,k} = \frac{f_{\text{PQ}}}{\hat{N}_{k}},$$

Alternatively, one can obtain Lagrangian by redefining fermion fields [Kim, Dias et al] Redefinition not unique: Physically equivalent Lagrangians that differ by field redefinitions.

Axion couplings to neutral gauge bosons is not affected by this ambiguity (EM unbroken). Therefore  $f_{A3'}f_{AEM}$  only depend on the PQ charges of the scalars

Fermion-axion couplings reflect the PQ charges only in a particular basis; info on PQ charges can become hidden in axion coupling to massive bosons.

### Nonperturbative hadronic/QCD effects

#### Axion coupling to nucleons:

Recovered in a fermionic axial basis in which interactions involve axial currents

$$\partial_{\mu}A\left[\frac{q_{1}}{f_{\mathrm{PQ}}}\psi^{\dagger}\bar{\sigma}^{\mu}\psi + \frac{q_{2}}{f_{\mathrm{PQ}}}\tilde{\psi}^{\dagger}\bar{\sigma}^{\mu}\tilde{\psi}\right] \to -\partial_{\mu}A\,\frac{q_{1}+q_{2}}{2f_{\mathrm{PQ}}}\,\overline{\Psi}\gamma^{\mu}\gamma_{5}\Psi.$$

The Dirac fermions contain Weyl spinors linked by a PQ invariant Yukawa couplings, so  $q_1 + q_2$  is equal to a scalar PQ charge. From chiral perturbation theory at NLO and lattice results [Villadoro et al]

$$\delta \mathcal{L}_{eff} = -\partial_{\mu} A \frac{C_{AN}}{2f_A} \overline{N} \gamma^{\mu} \gamma_5 N - \partial_{\mu} A \frac{C_{AP}}{2f_A} \overline{P} \gamma^{\mu} \gamma_5 P,$$
  
$$C_{AN} = -0.02(3) + 0.41(2) \frac{q_{H_u} f_A}{f_{PQ}} - 0.83(3) \frac{q_{H_d} f_A}{f_{PQ}},$$
  
$$C_{AP} = -0.47(3) - 0.86(3) \frac{q_{H_u} f_A}{f_{PQ}} + 0.44(2) \frac{q_{H_d} f_A}{f_{PQ}}.$$

Fermion couplings in axial basis only depend on scalar PQ charges  $q_i/f_{PQ}$ .

#### Axion coupling to photons:

At low energy, axion field can mix with other pseudo-Goldstones, like the neutral pion. An appropriate field redefiniton removes the mixing and gives for the physical axion

$$\delta \mathcal{L}_{\text{eff}} = \frac{\alpha}{8\pi f_A} \,\delta C_{A\gamma} A \tilde{F}_{\mu\nu} F^{\mu\nu}, \,\delta C_{A\gamma} = -\frac{2}{3} \left(\frac{4m_u + m_d}{m_u + m_d}\right) + \text{higher order} = -1.92(4).$$

## Wherefore art thou a physical Goldstone?

Axion field must be orthogonal to massive gauge bosons. Its associated **physical PQ symmetry** will not be the simple naive one.

 $PQ_{phys}$  is generated by a combination of original PQ and other symmetries  $S_i$ :

$$PQ_{\rm phys} = PQ + \sum_{j} \lambda_j S_j \to A = \sum_{i} c_i A_i, \quad c_i = \frac{q_i^{\rm phys} v_i}{\sqrt{(q^{\rm phys})_i^2 v_i^2}}$$

Unknown  $c_i$  giving charges of PQ<sub>phys</sub> can be obtained by solving masslessness and orthogonality conditions

Art thou massless?  $\mathcal{L} \supset m\left(\sum_{m} d_{m}A_{m}\right)^{2} \Rightarrow \sum_{m} c_{m}d_{m} = 0$ 

#### Art thou orthogonal to massive gauge bosons?

Avoidance of kinetic mixing with gauge bosons implies  $\sum_{mn} v_m T_{mn}^a c_n = 0.$ For a massive U(1) boson with associated scalar charges  $\tilde{q}_m$ :  $\sum_m v_m \tilde{q}_m c_m = 0.$ 

Need at least 2 scalars charged under massive U(1) for them to contain the axion.  $f_A$  of the order of the smallest VEV of fields charged under massive U(1).

### Domain wall number

$$\mathcal{L}_{\text{eff}}[A] \supset A \sum_{k=1}^{N_g} \left( \theta_k + \frac{A\hat{N}_k}{f_{\text{PQ}}} \right) \frac{g_k^2}{16\pi^2} \,\bar{\text{Tr}} \,\tilde{F}_{\mu\nu}^k F^{k,\mu\nu} + \text{derivatives}$$

Axion-Goldstone shifts under PQ<sub>phys</sub>,  $A \rightarrow A + \alpha f_{PQ}$ . The anomalies break shift symmetry in the SU(3) sector to a discrete subset (equivalent to translations  $\theta_3 \rightarrow \theta_3 + 2\pi$ ).

$$S_{\text{phys}}(n): A \to A + \frac{2\pi n}{\hat{N}} f_{\text{PQ}}, \quad n \in \mathbb{Z}.$$

Some of this translations correspond to unphysical rotations of phases  $A_i$  by  $2\pi$ 

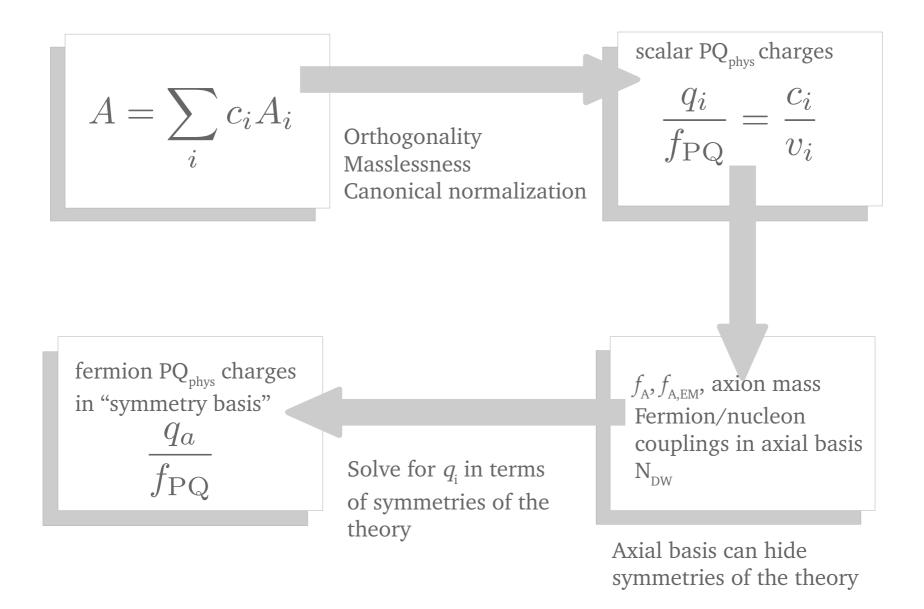
$$P_{\rm phys}(n_i): A \to A + \sum_i \frac{2\pi n_i q_i v_i^2}{f_{\rm PQ}}, \quad n_i \in \mathbb{Z}.$$
$$N_{\rm DW} = \dim \left[\frac{S_{\rm phys}}{P_{\rm phys}}\right] = \text{ min. integer} \left\{\frac{1}{f_A} \sum_i n_i c_i v_i, \, n_i \in \mathbb{Z}\right\} \left( \to \frac{\hat{N}}{\mathrm{MCD}[q_i]}, \text{ rational } q_i \right)$$

Relation to naive UV PQ symmetry (without imposing orthogonality conditions): Have to mod out by the discrete transformations in the center Z of the gauge group!

$$N_{\rm DW} = \dim\left[\frac{S_{\rm phys}}{P_{\rm phys}}\right] = \dim\left[\frac{S}{ZP}\right]$$

Fun fact. DFSZ models have  $N_{DW}$ =3, not the usually quoted 6, because  $Z_{SU(2)}$ = $Z_2$ .

# The axion hunting flow



# The GUTs of the strong CP problem

### What to expect from GUT symmetry

 $\mathcal{L}_{\rm GUT} \supset \bar{\mathrm{Tr}}_{\rm GUT} \frac{g^2 \theta_{\rm GUT}}{32\pi^2} \epsilon^{\mu\nu\rho\sigma} F^{\rm GUT}_{\mu\nu} F^{\rm GUT}_{\rho\sigma} = \bar{\mathrm{Tr}}_3 \frac{g^2 \theta_{\rm GUT}}{32\pi^2} \epsilon^{\mu\nu\rho\sigma} F^3_{\mu\nu} F^3_{\rho\sigma} + \dots$ 

Only one fundamental  $\theta$  angle. Only one  $\theta_{phys}$ .

The GUT axion must solve the CP problem in all subgroups!

$$\mathcal{L}_{\rm GUT}[A] \supset \frac{A}{f_A} \frac{g^2}{16\pi^2} \,\bar{\rm Tr}_{\rm GUT} \,\tilde{F}^{\rm GUT}_{\mu\nu} F^{\rm GUT,\mu\nu} = \frac{A}{f_A} \frac{g^2}{16\pi^2} \,\bar{\rm Tr}_3 \,\tilde{F}^3_{\mu\nu} F^{3,\mu\nu} + \dots$$

GUT symmetry explicit —  $\blacktriangleright$  all  $\theta_k$  and all  $f_{Ak}$  will be the same [modulo normalization factors]

Field redefinitions can break the GUT symmetry and give different  $\theta_k$ ,  $f_{Ak}$ , yet with unique  $\theta_{phys}$ 

Eg: axial basis where the axion couplings to nucleons are calculated. In such basis  $f_{A2}$  deviates from  $f_{A3}$  and the GUT symmetry is hidden, but could in principle be uncovered if all couplings could be measured...

PQ<sub>phys</sub> should be a combination of symmetries of GUT theory! Possibly involving Cartan of SO(10)

### Relevant SO(10) representations

Each generation comes come in spinorial 16 representation

| SO(10) | $3_C 2_L 1_Y$                                    |
|--------|--|
| $16_F$ | $(3, 2, \frac{1}{6}) := q$                       |
|        | $(1,2,-\frac{1}{2}) := l$                        |
|        | $\left(\overline{3}, 1, \frac{1}{3}\right) := d$ |
|        | (1,1,1) := e                                     |
|        | $\left(\bar{3}, 1, -\frac{2}{3}\right) := u$     |
|        | (1,1,0) := n                                     |

 $16_F \times 16_F = 10_H + 120_H + 126_H$ 

Scalar VEVs can break the rank 5 group SO(10). Which scalar representations?

 $SO(10) \xrightarrow{M_U - 210_H} 4_C 2_L 2_R \xrightarrow{M_{BL} - \overline{126}_H} 3_C 2_L 1_Y \xrightarrow{M_Z - 10_H, 126_H} 3_C 1_{em}$  $SO(10) \xrightarrow{M_U - 210_H} 4_C 2_L 2_R \xrightarrow{M_{PQ} - 45_H} 4_C 2_L 1_R \longrightarrow 3_C 2_L 1_Y \xrightarrow{M_Z - 10_H, 126_H} 3_C 1_{em}$  $Why a U(1) symmetry? \quad \mathcal{L}_Y = 16_F \left( Y_{10} 10_H + \tilde{Y}_{10} 10_H^* + Y_{126} \overline{126}_H \right) 16_F + h.c.$ 

Can make it more predictive by imposing global U(1) which forbids  $\tilde{Y}_{10}$ . Anomalous, chiral  $16_F \rightarrow 16_F e^{i\alpha}, \quad 10_H \rightarrow 10_H e^{-2i\alpha}, \quad \overline{126}_H \rightarrow \overline{126}_H e^{-2i\alpha}, \quad \text{PQ symmetry of axion!}$ 

# SO(10) models

#### Simplest model: (ruled out) electroweak axion

Model with above PQ charges has  $f_A$  at the electroweak scale due to orthogonality conditions!

Way out: need more fields with large VEVs charged under PQ!

GUT scale axion (no axion DM in post-inflationary scenarios)

 $N_{DW} = \frac{\hat{N}}{\mathrm{MCD}[q_i]}, \text{ rational } q_i$ 

Intermediate scale (additional  $45_{H}$ ): Avoid GUT sale  $f_{Ak}$  by not charging 210 under PQ.

|             | $16_F \to 16_F e^{i\alpha},$                           |   |
|-------------|--|---|
|             | $10_H \to 10_H e^{-2i\alpha},$                         | $\hat{N}_{\text{GUT}} = 2 \times 3 \text{ gens} \times (\text{ index of } 16_F = 2) = 12$ |
| Model 2.1 : | $\overline{126}_H \to \overline{126}_H e^{-2i\alpha},$ | Rational scalar PQ charges with MCD 2.  |
|             | $210_H \to 210_H,$                                     | Center of SO(10): $Z_2$   |
|             | $45_H \to 45_H e^{4i\alpha}$                           | Expect $N_{\rm DW} = 3$   |

## SO(10) models

Additional  $45_{H}$  with 2 heavy fermion multiplets in  $10_{F}$ : Add extra heavy fermions in  $10_{F}$  to change GUT anomaly coefficient and get  $N_{DW} = 1!$  [Lazarides].

 $\begin{array}{ll} 16_F \rightarrow 16_F e^{i\alpha}, \\ 10_H \rightarrow 10_H e^{-2i\alpha}, \\ \overline{126}_H \rightarrow \overline{126}_H e^{-2i\alpha}, \\ 210_H \rightarrow 210_H, \\ 45_H \rightarrow 45_H e^{4i\alpha}, \\ 10_F \rightarrow 10_F e^{-2i\alpha} \end{array} \begin{array}{ll} \hat{N}_{\rm GUT} = 2 \times (3 \ {\rm gens}) \times (1{\rm PQ}) \times (\ {\rm ind} \ {\rm of} \ 16_F = 2) \\ + 2 \times (2 \ {\rm heavy} \ {\rm F}) \times (-2{\rm PQ}) \times (\ {\rm ind} \ {\rm of} \ 10_F = 1) = 4 \\ {\rm Rational \ scalar \ PQ \ charges \ with \ MCD \ 2.} \\ {\rm Center \ of \ SO(10): \ Z_2} \\ {\rm Expect} \ N_{\rm DW} = 1 \end{array}$ 

Models with decay constants independent of gauge symmetry breaking

$$\begin{array}{ll} 16_F \rightarrow 16_F e^{i\alpha}, & 16_F \rightarrow 16_F e^{i\alpha}, \\ 10_H \rightarrow 10_H e^{-2i\alpha}, & 10_H \rightarrow 10_H e^{-2i\alpha}, \\ \hline 126_H \rightarrow \overline{126}_H e^{-2i\alpha}, & \mathbf{Model \ 3.2:} & \overline{126}_H \rightarrow \overline{126}_H e^{-2i\alpha}, \\ 210_H \rightarrow 210_H, & 210_H , \\ S \rightarrow S e^{4i\alpha} & S \rightarrow S e^{4i\alpha}, \\ 10_F \rightarrow 10_F e^{-2i\alpha}, \end{array}$$

### Example case: axion and couplings in 2.2

Scalar components getting VEVS, in terms of decompositions under  $PS=SU(4)xSU(2)_L xSU(2)_R$ . SM arises from PS as follows:

$$SU(4) \supset SU(3) \times U(1)_{B-L} \qquad Y = U(1)_R + \frac{1}{2}U(1)_{B-L}$$
  
$$SU(2)_R \supset U(1)_R$$

| SO(10)             | $4_C 2_L 2_R$ | $3_C 2_L 1_Y$          | $3_C 1_{\rm em}$                   | scale        | VEV               |
|--------------------|---------------|------------------------|------------------------------------|--------------|-------------------|
| $10_H$             | (1, 2, 2)     | $(1,2,\frac{1}{2})$    | $(1,0) =: H_u$                     | $M_Z$        | $v_u^{10}$        |
|                    |               | $(1, 2, -\frac{1}{2})$ | $(1,0) =: H_d$                     | $M_Z$        | $v_d^{10}$        |
| $45_H$             | (1, 1, 3)     | (1, 1, 0)              | $(1,0) := \boldsymbol{\sigma}$     | $M_{\rm PQ}$ | $v_{\mathrm{PQ}}$ |
| $\overline{126}_H$ | (10, 1, 3)    | (1, 1, 0)              | $(1,0) := \Delta_R$                | $M_{\rm BL}$ | $v_{\rm BL}$      |
|                    | (15, 2, 2)    | $(1, 2, \frac{1}{2})$  | $(1,0) := \Sigma_{\boldsymbol{u}}$ | $M_Z$        | $v_{u}^{126}$     |
|                    |               | $(1, 2, -\frac{1}{2})$ | $(1,0) := \Sigma_d$                | $M_Z$        | $v_d^{126}$       |
| $210_{H}$          | (1, 1, 1)     | (1, 1, 0)              | $(1,0) := \phi$                    | $M_{ m U}$   | $v_{\mathrm{U}}$  |

Axion is a combination of phases of the first 6 scalars (because 210 has no PQ charge):

$$\phi_1 \equiv \Sigma_u, \phi_2 \equiv \Sigma_d, \phi_3 \equiv H_u, \phi_4 \equiv H_d, \phi_5 \equiv \Delta_R, \phi_6 \equiv \sigma,$$

$$\phi_j = \frac{1}{\sqrt{2}} (v_j + \rho_j) e^{iA_j/v_j}.$$

$$A = \sum_{i} c_i A_i$$

### Example case: axion and couplings in 2.2

All equations from orthogonality/constraints compatible! (need 5 eqs. and normalization condition)

$$A = -\frac{(A_4v_4 + A_2v_2)(v_3^2 + v_1^2) + (A_3v_3 + A_1v_1)(v_4^2 + v_2^2) - A_6v_6v^2}{\sqrt{v^2((v_4^2 + v_2^2)(v_3^2 + v_1^2) + v_6^2v^2)}}, \quad v^2 \equiv \sum_{i=1}^4 v_i^2$$

From  $c_i$  to scalar PQ<sub>phys</sub> charges:  $\frac{q_i}{f_{PQ}} = \frac{c_i}{v_i}$ 

$$PQ_{phys} = s_1 PQ + s_2 U(1)_R + s_3 U(1)_{B-L}.$$

$$\frac{s_1}{f_{PQ}} = \frac{v}{4\sqrt{(v_1^2 + v_3^2)(v_2^2 + v_4^2) + v^2 v_6^2}}, \quad \frac{s_2}{f_{PQ}} = -\frac{v_1^2 - v_2^2 + v_3^2 - v_4^2}{v\sqrt{(v_1^2 + v_3^2)(v_2^2 + v_4^2) + v^2 v_6^2}},$$
$$\frac{s_3}{f_{PQ}} = \frac{v_1^2 - 3v_2^2 + v_3^2 - 3v_4^2}{4v\sqrt{(v_1^2 + v_3^2)(v_2^2 + v_4^2) + v^2 v_6^2}}$$

From this we can get  $PQ_{phys}$  of Weyl fermions and from the latter  $f_{Ak}$ . All  $f_{Ak}$  are equal modulo hypercharge normalization, as follows from GUT symmetry!

$$f_{A,3c} = f_{A,2L} = \frac{5}{3} f_{A,Y} = \sqrt{\frac{v_6^2 v^2 + (v_4^2 + v_2^2)(v_3^2 + v_1^2)}{v^2}} \sim v_{\sigma}.$$
  
We also get  $N_{\rm DW} = 1$  from the scalar PQ<sub>phys</sub> charges!  $N_{\rm DW} = \min$ . integer  $\left\{ \frac{1}{f_A} \sum_i n_i c_i v_i, n_i \in \mathbb{Z} \right\}$ 

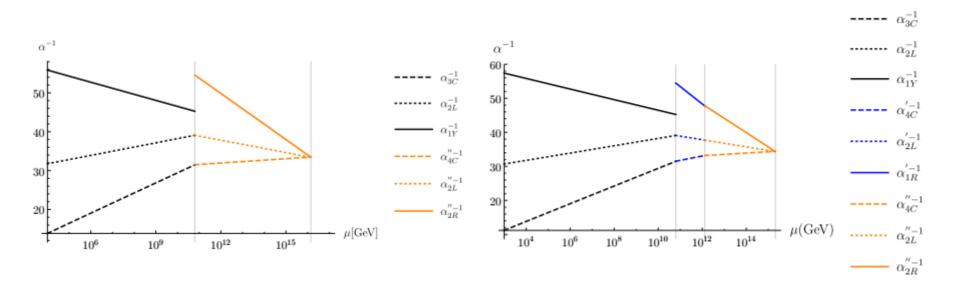
## Unification constraints

2 loop RG equations, ignoring Yukawas. 3 thresholds scales  $M_{PQ}, M_{BL}, M_{U}$ 

#### Full 1 loop threshold corrections

**Extended survival hypothesis [del Aguila & Ibañez]:** RG at a given scale only includes scalar multiplets which acquire a VEV at lower scales, with the exception of  $\Sigma_u$ ,  $\Sigma_d$ , which are assumed to decouple at a scale  $M_{BL}$  in order to give rise to a low-energy 2HDM limit [Babu & Khan]

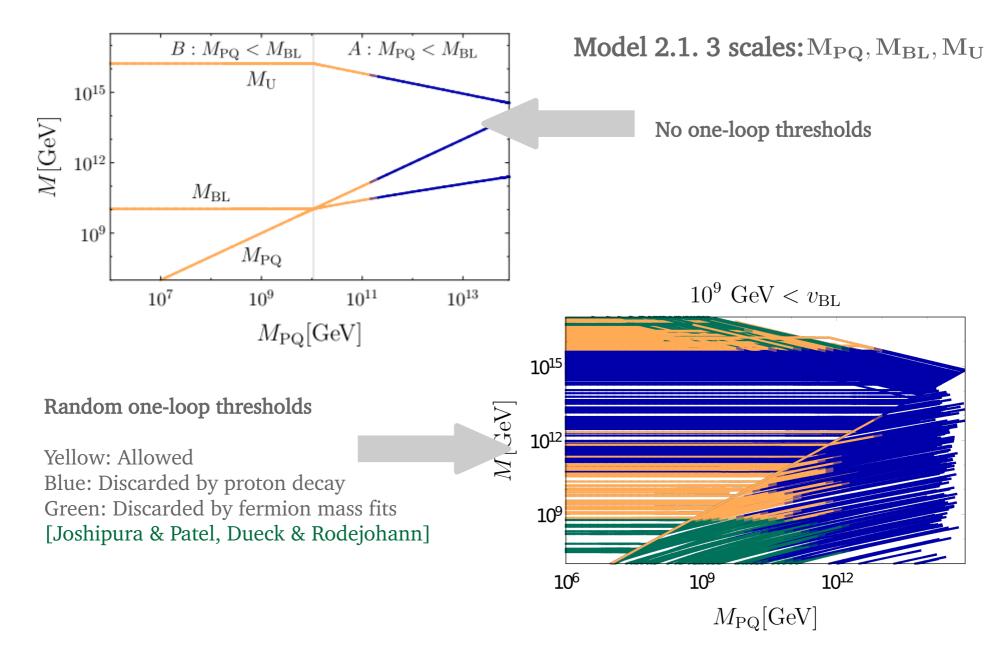
At thresholds we assume that the scalars and fermions which decouple take masses in the range of 1/10-10 times the threshold mass scale.



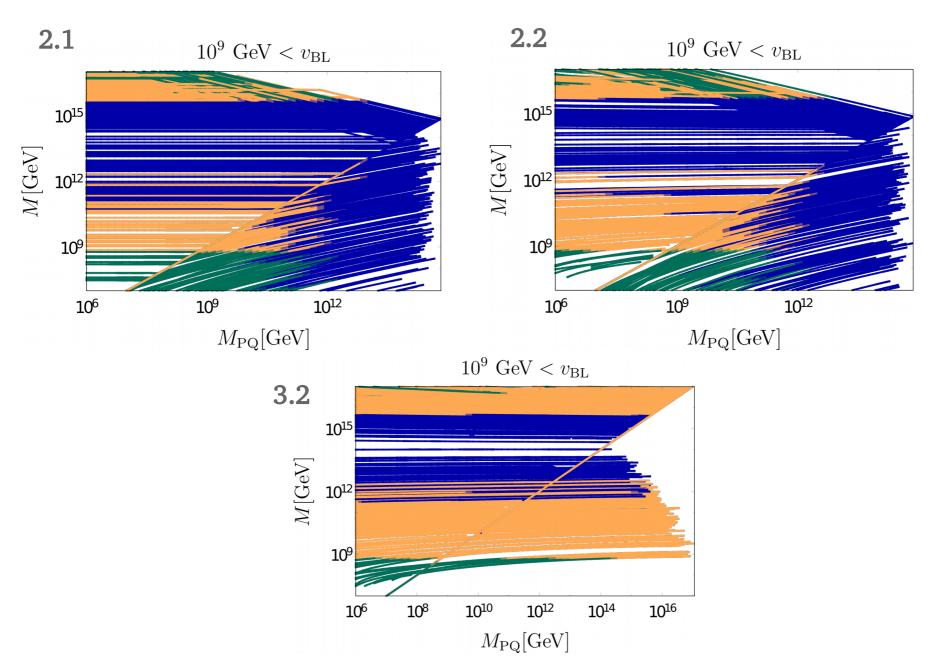
Model 1. 2 step. No loop thresholds

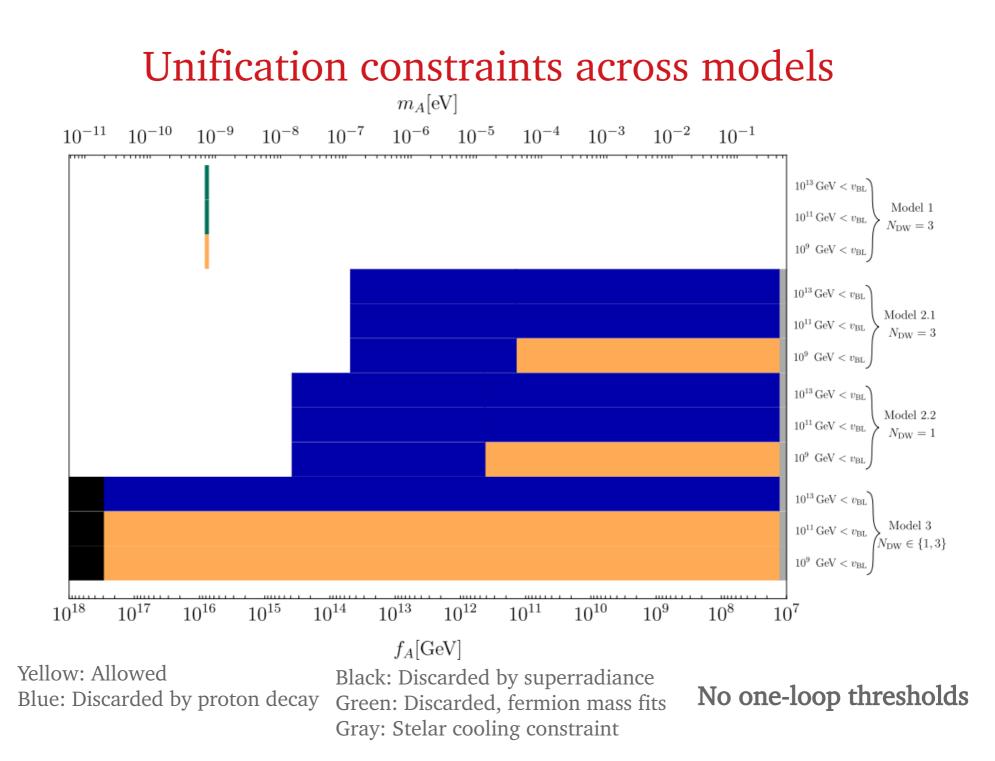
Model 2.1. 3 step. No loop thresholds

# Unification constraints: impact of 1 loop thresholds

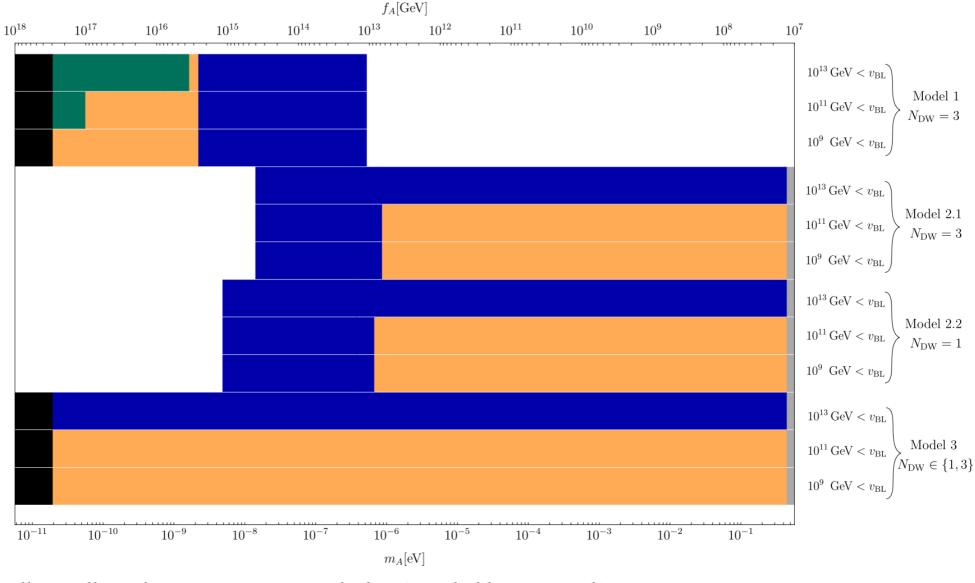


# Unification constraints: impact of 1 loop thresholds



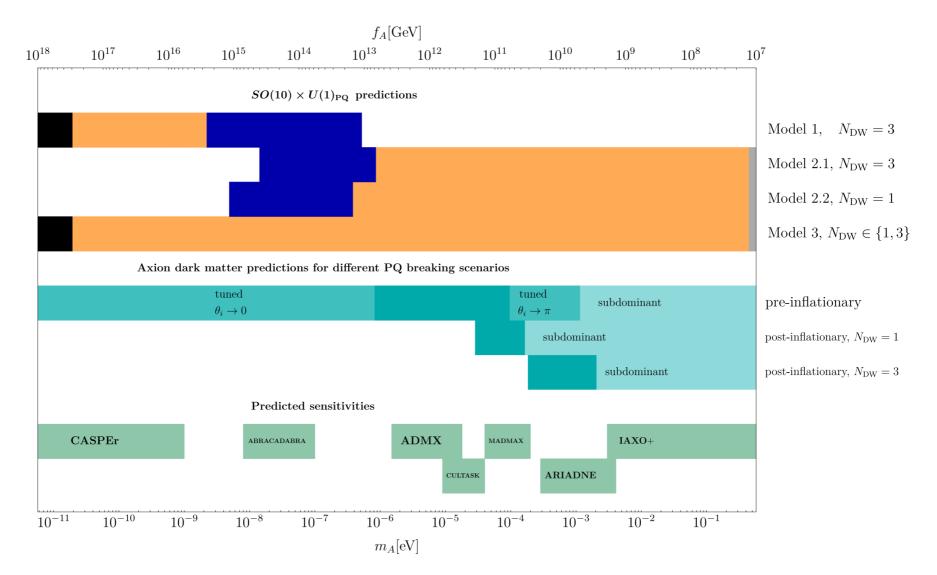


# Unification constraints across models



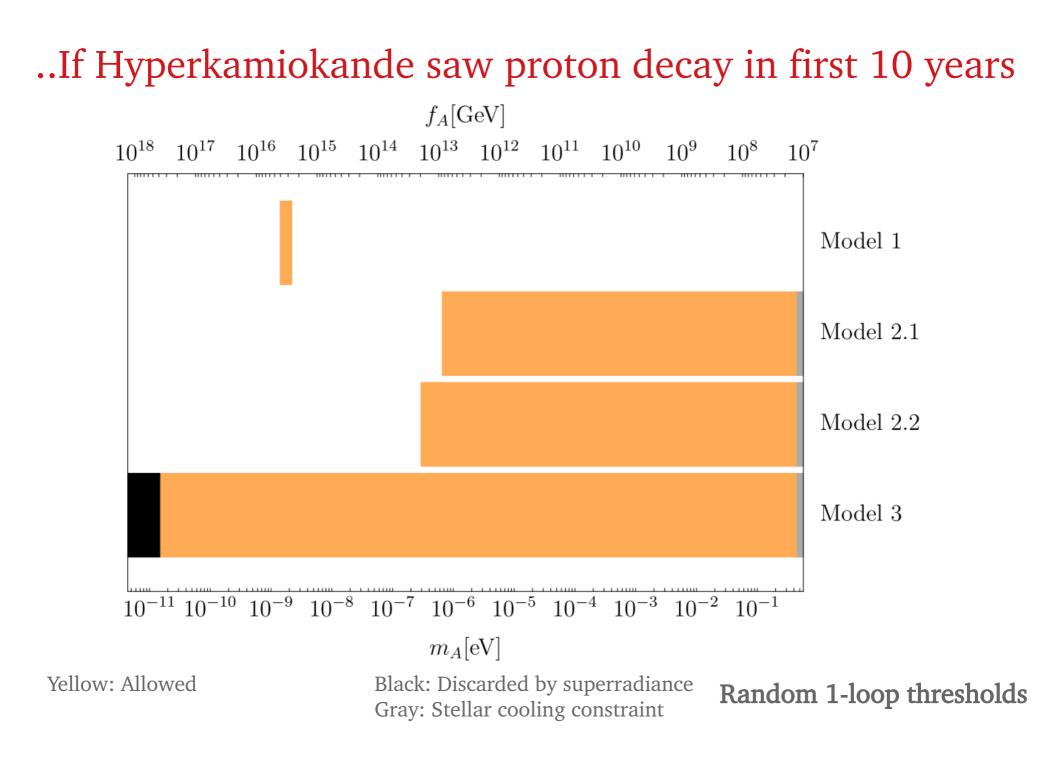
Yellow: Allowed Blue: Discarded by proton decay Black: Discarded by superradiance Green: Discarded, fermion mass fits Gray: Stellar cooling constraint

# Summary: Unification constraints



Yellow: Allowed Blue: Discarded by proton decay Black: Discarded by superradiance Green: Discarded, fermion mass fits Gray: Stellar cooling constraint

Random 1-loop thresholds



### Summary of axion couplings

In axial basis (in which GUT symmetry is not manifest)

$$\mathcal{L} = \frac{1}{2} \partial_{\mu} A \partial^{\mu} A - \frac{1}{2} m_A^2 A^2 + \frac{\alpha}{8\pi} \frac{C_{A\gamma}}{f_A} A F_{\mu\nu} \tilde{F}^{\mu\nu} - \frac{1}{2} \frac{C_{Af}}{f_A} \partial_{\mu} A \overline{\Psi}_f \gamma^{\mu} \gamma_5 \Psi_f ,$$

$$C_{A\gamma} = \frac{8}{3} - 1.92(4), \qquad C_{Ae} = \frac{1}{N_{\rm DW}} \sin^2 \beta,$$
  

$$C_{Ap} = -0.47(3) + \frac{3}{N_{\rm DW}} [0.29 \cos^2 \beta - 0.15 \sin^2 \beta \pm 0.02],$$
  

$$C_{An} = -0.02(3) + \frac{3}{N_{\rm DW}} [-0.14 \cos^2 \beta + 0.28 \sin^2 \beta \pm 0.02],$$

$$\tan^2 \beta = ((v_u^{10})^2 + (v_u^{126})^2) / ((v_d^{10})^2 + (v_d^{126})^2)$$

DFSZ recovered for  $N_{\rm DW}$ =3

## Conclusions

We identified the axion field, obtained its couplings to gauge bosons, elementary fermions and nucleons, and computed the domain wall number corresponding to the physical PQ symmetry, in several SO(10) models.

Our formalism bridges the gap between UV and IR symmetries. Accounting for orthogonality with respect to massive gauge bosons, we identified the physical PQ symmetry as a combination of UV symmetries.

We clarified issues pertaining to fermion field redefinitions.

We studied in detail constraints from unification, superradiance, cooling, and fermion masses.

The axion in these models can be probed by upcoming experiments!